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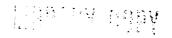
RESEARCH REPORTS - 1980 NASA / ASEE SUMMER FACULTY FELLOWSHIP PROGRAM

University of Alabama University, Alabama

and

University of Alabama in Huntsville Huntsville, Alabama

October 1980



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Prepared for

NASA - George C. Marshall Space Flight Center Marshall Space Flight Center, Alabama 35812

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This document is a collection of	of reports on t	he research condu	ucted by the p	articipants in
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Marshall Space Flight Center (MSFC) for the s	ixteenth consecut	tive year. Th	e 1980 program
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N81-11977

RESEARCH REPORTS

1980 NASA/ASEE SUMMER FACULTY FELLOWSHIP PROGRAM

George C. Marshall Space Flight Center The University of Alabama and The University of Alabama in Huntsville

EDITORS:

Dr. B. F. Barfield, Professor of Mechanical Engineering The University of Alabama

Mr. Marion I. Kent, Office of University Affairs Marshall Space Flight Center

Dr. James Dozier, Director, Research and Technology Office Marshall Space Flight Center

Dr. Gerald Karr, Professor of Mechanical Engineering University of Alabama in Huntsville

NASA CR - 161511

PREFACE

For the sixteenth consecutive year, a NASA/ASEE Summer Faculty Fellowship Research Program was conducted at the Marshall Space Flight Center (MSFC). The program was conducted by The University of Alabama and MSFC during the period May 26, 1980 through August 1, 1980. The program was operated under the auspices of the American Society for Engineering Education (ASEE). The program at MSFC, as well as those at five other NASA Centers, was sponsored and funded by the Office of University Affairs, NASA Headquarters, Washington, D.C. The basic objectives of the programs, which are in the seventeenth year of operation nationally, are:

- a. To further the professional knowledge of qualified engineering and science faculty members;
- b. To stimulate an exchange of ideas between participants and NASA:
- c. To enrich and refresh the research and teaching activities of participants; instutions; and,
- d. To contribute to the research objectives of the NASA Centers.

The Faculty Fellows spent ten weeks at MSFC engaged in a research project commensurate with their interests and background and worked in collaboration with a NASA/MSFC Colleague. This document is a compilation of Fellow's reports on their research during the Summer of 1980. University of Alabama Report No. BER-259-94 presents the Co-Directors' report on the administrative operations of the Research Fellowship Program. Further information can be obtained by contacting any of the editors.

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

1980

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

DEVELOPMENT OF A FACILITY FOR LARGE-SCALE TESTING OF MULTISTORY BUILDINGS

Prepared By:

Daniel P. Abrams, Ph.D.

Academic Rank:

Assistant Professor

University of Colorado

University and Department:

NASA/MSFC: (Laboratory) (Division)

MSFC Counterpart:

Date:

Contract No.:

at Boulder Civil,Environmental and Architectural Engineering

Space Sciences Atmospheric Sciences

Dr. Nicholas C. Costes

August 1, 1980

NGT-01-002-099 (University of Alabama)

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Development of a Facility for Large-Scale Testing of Multistory Buildings

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Daniel P. Abrams, Ph. D. Assistant Professor of Civil Engineering University of Colorado Boulder, Colorado

Abstract

Current experimental research pertaining to response of structures subjected to lateral loads such as strong winds or earthquake motions is limited to either tests of single structural components or small-scale (approximately one-tenth) multistory structures. The feasibility of developing a facility where large-scale multistory structures could be loaded to failure is discussed.

The test facility would consist of a series of hydraulic actuators mounted on reaction frames currently in use at the Marshall Space Flight Center for structural testing of spacecraft components. The actuators would be controlled from signals computed by an on-line analysis of measured data. This method of loading could be used to simulate inertial forces resisted by a structure behaving in the nonlinear range of response subjected to motion at the base or to impulses along the height as would occur during a strong earthquake or wind.

Development and utilization of the test facility would occur in two phases. In the first phase, the testing system would be developed and verified for loading a structure within a single plane. The second phase would utilize the loading system by testing a half-scale ten-story reinforced concrete building 45 feet tall. Further development of the test facility may include the capability of loading individual floor levels to simulate translational and rotational inertia in three directions.

ACKNOWLEDGEMENTS

The writer expresses his appreciation to the National Aeronautics and Space Administration, and the University of Alabama and the American Society of Engineering Education for funding and conducting the Summer Faculty Fellowship Program at the Marshall Space Flight Center under the direction of Marion Kent of NASA and Dr. Robert Barfield of the University of Alabama.

Special gratitude is extended to Dr. Nicholas Costes who provided many creative ideas and stimulating conversations, and was responsible for introducing the writer to numerous helpful individuals at Marshall.

The writer wishes to thank Dr. George McDonough for his support of the project and George Shofner, Charles Watson, Frank Vinz and Gerry Waggoner for their technical advice and constructive comments.

Credit is given to Willa Eslick for her patience in typing the paper.

INTRODUCTION

"Earthquakes have caused, and can cause in the future, enormous loss of life, injury, destruction of property, and economic and social disruption."

This statement quoted from the Earthquake Hazards Reduction Act of 1977 [1] demonstrates the importance for safer structures in regions of high seismicity. More specifically, the primary objective of the Act passed by Congress is

"The development of technologically and economically feasible design and construction methods and procedures to make new and existing structures, in areas of seismic risk, earthquake resistant..."

In accordance with this objective, several investigators have examined both analytically and experimentally the response of structures subjected to strong ground motions. Of the experimental studies, investigations have been limited to response of one of two general specimen types: large-scale single components of more complex structures, or small-scale (on the order of one-twelfth) multistory structures. Large-scale multistory structures have not been tested to destruction because costs of developing a loading system and constructing a test structure were thought previously to be prohibitive.

This paper presents the feasibility of developing a facility where largescale multistory structures could be loaded to failure. The proposed test facility would consist of a series of hydraulic actuators mounted on reaction frames currently in use at the Marshall Space Flight Center (MSFC) for structural testing of spacecraft components. Development and utilization of the test facility is discussed in terms of the cost-value relationship of using large-scale test structures for earthquake engineering research.

OBJECTIVES

Results of an experimental study using a large-scale test structure could be the necessary evidence required to transfer existing research findings to standards for building earthquake-resistant structures. Moreover, tests of large-scale structures may help discover and develop new methods of designing and constructing safer structures. Measured response of a largescale multistory structure would:

(1) verify the capability of the testing system to simulate inertial loads resulting from a programmed impulse;

(2) for the first time, provide comprehensive data for an indepth study of both the overall and local response of a multistory structure subjected to a strong ground motion;

(3) verify currently used small-scale modeling procedures thus extend experimental research of multistory structures using shaking tables;

(4) provide a data base for testing newly developed analytical models used for calculating nonlinear dynamic response of multistory structures;

(5) serve as a "proof test" of recent improvements in earthquakeresistant design procedures developed with mathematical and small-scale physical models;

(6) provide data to evaluate methods of rehabilitation of existing and damaged structures such as epoxy grout injections or masonry in-fill panels.

The objective of the study presented in this paper is to examine the feasibility of developing a large-scale testing facility and to outline a test program utilizing the facility to illustrate these attributes of large-scale testing.

A hypothetical structure (Fig. 1) has been chosen as an example of a typical test article. The structure is a one-half scale reinforced concrete building consisting of two three-bay frames coupled to a slender shear wall at ten levels with floor slabs poured monolithically. The configuration was chosen to replicate a similar configuration previously used with small-scale models [2, 3, 4]. The scale was chosen so that sizes of members and reinforcement would be representative of actual construction. It is planned to develop a testing system to load this structure laterally at all ten levels from a single direction.

OUTLINE OF TEST PROGRAM

An outline of the test program is presented in Fig. 2. The program would consist of three stages. In the first stage the testing system would be developed using existing loading equipment at MSFC and a computer analysis that would be developed. The second stage would consist of verifying operation of the testing system by constructing and loading a replica of a smallscale (approximately 8-ft. tall) ten-story structure tested previously on a shaking table [2]. The loading procedure which would consist of a series of slowly applied loading reversals simulating inertial forces could be validated by comparing displacement response of the structure tested on a shaking table with that of the structure loaded laterally with actuators. If agreement can be made, then the loading system would be developed further to test the 45-ft tall structure shown in Fig. 1 using the same load control system and on-line computer analysis with larger size actuators and reaction structure.

It is planned to test the 45-ft tall half-scale structure using the same input base - acceleration history as used for the 8-ft tall structure. In this way, scaling relationships may be examined for modeling reinforced concrete at small scales. After the structure has been tested, it can be repaired and loaded once again to investigate methods of rehabilitation. The test structure then can be used as a loading system for investigating the behavior of masonry sub-panels by repairing the structure once again, in filling each frame panel with masonry and retesting. In this manner, each of sixty frame panels, or masonry elements, will be subjected to a different ratio of shear to normal stress thus providing a larger population of data than currently exists in all of the past masonry research.

LOADING SYSTEM

The proposed loading system consists of a series of hydraulic actuators mounted to an existing reactor frame used presently for testing portions of the liquid oxygen tanks of the Space Shuttle. The reactor structure (Fig. 3) which is located within the structural test tower (Fig. 4) of the load test annex of building 4619 has sufficient lateral-load capacity to support the 20 actuators (two per level, Table 1) required to load the 45-ft tall test structure. Minor modifications of the reaction structure would be necessary to support individual actuators. In addition, the reaction structure could be used to support actuators mounted in transverse and vertical directions for further development of the loading facility.

A contingency plan could also be developed to use the existing test tower (Fig. 4) if removal of the reaction frame were necessary for other reasons. The test area measures $51 \times 49 \times 47$ meters high. A double trolley crane with a total capacity of 54 metric tons is accessible for erection of the test structure. Further detailed information about the facility is given in Reference 5.

The existing load control system located in building 4619 has been cited [6] by the Japanese as being a unique system for controlling several actuators simultaneously. A closed-loop control system using a PDP-11 computer monitors response of a specimen at intervals of 1/60th of a second and adjusts the load to attain a preselected specimen response. This method of control is particularly useful if it is desired to displace a structure through a specified deflected shape. Actuators loading a structure at several levels can be programmed to displace prescribed amounts despite interaction among actuators which has troubled investigators previously. The capability of the load-control system to coordinate simultaneous actuator movement is proven by numerous experiments including a recent one where several actuators were used to load in three directions an ortho-grid assembly used for part of the floor system in Space lab. The proposed loading facility would utilize ten of fifty-eight channels of the existing load-control system.

The actuators would be controlled from signals generated by an on-line analysis of measured data to simulate inertial loads resulting from a prescribed impulse. For a standard time increment of the impulse duration, each actuator would displace an amount calculated from previous force and displacement measurements and the prescribed forcing function. The differential equation to be satisfied for each floor level is:

 $\dot{mx} + F_r = p(t)$ (1) where m is the mass of the floor and lumped portions of the columns and wall, \ddot{x} is the acceleration of the mass, F_r is the restoring force of the structure, and p(t) is the prescribed forcing function. Using finite differences, the acceleration may be related to the displacement over a time increment by

$$\ddot{\mathbf{x}}_{i} = (\mathbf{x}_{i+1} - 2\mathbf{x}_{i} + \mathbf{x}_{i-1}) / (\Delta t)^{2}$$
(2)

where Δt denotes the time interval and i refers to the particular time increment. The new displacement, x_{i+1} , that controls the actuator movement, may be calculated from \ddot{x}_i which is determined from equation (1)

$$\mathbf{m}_{i} = \mathbf{p}(t)_{i} - \mathbf{F}_{ri}$$

For an impulse at the base of a multistory structure, p(t) becomes $m_{x_0}^{*}$ where

xo is the input acceleration at base. Any prescribed ground motion including translational, rocking or torsional motions may be input to the online analysis to load the fixed-base structure "pseudo dynamically."

The restoring force, F_r , is measured continuously during the loading which is particularly important for a nonlinearly behaving structure because the restoring force or stiffness of the structure is constantly changing and sometimes unpredictable. For a nonlinear dynamics problem the fact that the inertial forces are dependent on the stiffness, and the stiffness is dependent on the loading history, justifies the on-line analysis as an irreplaceable technique for simulating inertial loads. Because the pseudo-dynamic test method is not velocity dependent, the actuators may be operated as slow as the hydraulic power supply requires. Some investigators have even advocated that slowing the loading process down to a nearly static rate is good for observing propagation of cracks and other types of progressive damage. It should be noted that the loading method does not account for effects of viscous, or velocity dependent damping, and strain-rate effects. However, it is reasonable to assume that most of the energy dissipation in reinforced concrete structures behaving within the nonlinear range of response is attributable to hysteretic damping which is not velocity dependent. Furthermore, strain-rate effects for velocities corresponding to the fundamantal mode and possibly the second mode of vibration of prototype structures have been shown to be negligible [7].

EXPERIMENTAL PROCEDURES

To economize on formwork the test structure (Fig. 1) would be erected from planar wall and frame elements cast in the horizontal plane and lifted to the vertical position. Reinforcement at base of columns and wall would be anchored in concrete blocks prestressed to the floor of the test area to prevent welding of reinforcing bars. Floor slabs could then be cast to complete erection of the structure. A minor amount of disassembly of the reaction structure may be necessary for lifting and placement of the structure.

Instrumentation would consist of two displacement and two load transducers at each level of the structure. Additionally, strain gages would be placed on reinforcing bars to detect internal forces in members and to infer bond characteristics between concrete and reinforcing bars. An estimated maximum of 200 channels of data would be a small percent of the 6000 channel Structural Test Data Acquisition System (STDAS) available at MSFC.

The test structure would be subjected to a series of lateral loads simulating response to the ground motion measured at El Centro, California, in the 1940 Imperial Valley earthquake (NS component). The intensity of the input motion would be scaled according to an analysis used for design of the structure so that similar levels of damage would be attained as for the smallscale structures. After the first test run, or loading, the structure would be loaded corresponding to an input motion of twice the intensity to further investigate strength and stiffness deterioration and energy dissipation characteristics. This pattern may be repeated until destruction occurs. At that time, the structure may be repaired using epoxy grout injections in the cracked concrete and tested once again. If the rehabilitation process can regain a sufficient amount of the original stiffness, the structure can be repaired once again and then strengthened with concrete-block masonry placed in each frame panel. Instrumentation can be installed to measure normal and shear stresses on each masonry panel when the structure is loaded once again.

In addition to the pseudo-dynamic testing, a series of low-amplitude free and forced vibration tests can be run to determine dynamic characteristics such as modal frequencies, shapes, and damping factors.

COST ESTIMATE

Approximate costs of developing a large-scale testing facility to load a 45-ft tall ten-story structure are itemized below. Costs include verification of the pseudo-dynamic loading system, and construction of the reinforced concrete test structure. Cost estimates have been compiled by engineers at MSFC who would be directly involved with the proposed work.

(1) Verify pseudo-dynamic testing method. Costs include testing an 8-ft tall model structure with the newly developed test method.

Reaction structure	\$5,000
Loading system	\$20,000
Instrumentation	-
Construction of 8-ft model	\$8,000
	\$33,000
Man months	30

(2) Develop large-scale testing facility. Costs include modifying existing structural test equipment and purchasing needed actuators, hardware, cables, hydraulic lines and miscellaneous items.

Reaction structure	\$30,000	
Loading system	\$65,000	
Man months	\$95,000 35	

(3) Constructing and testing 45-ft tall structure

Construction of structure Instrumentation Disassembly	\$85,000 \$7,000 \$10,000	
	\$102,000	
Man months	25	
Total	\$230,000	
Man months	90	

(4)

The overall cost of the project appears to be justifiable when compared with a similar attempt in Japan [6] and a previous feasibility study [8], each costing well over twenty million dollars.

EXTENDED DEVELOPMENT OF THE TEST FACILITY

Development of the loading system could be extended further in several steps to result ultimately in a loading system that could simulate translational inertia of each floor slab in three directions in addition to rotary inertia about three axes. Base motions input to the on-line analysis could then contain three translational components: two horizontal and one vertical, and three rotational components: torsion and rocking about each horizontal axis.

The existing reaction structure, hydraulic power supply, and loadcontrol system could be utilized. Two additional horizontal actuators, and two additional vertical actuators would be required per floor level.

Operation of the loading system could be verified by constructing several test structures at a small scale in duplicate. One of each pair of structures would be tested on the six-degree-of-freedom shaking table located in building 4663. The other structure of the pair would be tested using the newly developed loading system. Comparing response of each structure would verify the reliability of the pseudo-dynamic test method for multiaxial loading.

Furthermore, the existing structural test tower (Fig. 4) located in the local test annex of building 4619 is sufficiently tall and strong to accommodate a test structure up to 100-ft tall. A twenty-story half-scale, or ten-story full-scale, building could be tested with the additional cost primarily attributable to the purchase and support of additional actuators.

SUMMARY AND CONCLUSIONS

The feasibility of developing a facility for testing large-scale multistory buildings to failure has been presented. The facility would consist of a series of hydraulic actuators mounted to existing structural test equipment at the Marshall Space Flight Center. Displacement of each actuator would be controlled by an on-line analysis of measured data to simulate motion of a multistory structure subjected to a prescribed series of impulses input at the base as would occur during an earthquake. Construction, instrumentation, and test procedures are outlined for testing a 45-ft tall reinforced concrete structure to illustrate the worth of large-scale testing and the approximate costs.

The role of a large-scale testing facility for earthquake engineering research has been shown to be consistent with the objectives of the Earthquake Hazards Reduction Act of 1977. Development of such a facility at the MSFC using existing structural test equipment appears to be feasible technically and justifiable economically.

REFERENCES

- "Earthquake Hazards Reduction Act of 1977", Congressional Record, Vol. 123, October 7, 1977.
- Abrams, D. P. and M. A. Sozen, "Experimental Study of Frame-Wall Interaction in Reinforced Concrete Structures Subjected to Strong Earthquake Motions", Civil Engineering Studies, Structural Research Series, No. 460, University of Illinois, Urbana, May 1979.
- 3. Healey, T. J., and M. A. Sozen, "Experimental Study of the Dynamic Response of a Ten-Story Reinforced Concrete Frame with a Tall First Story", Civil Engineering Studies, Structural Research Series, No. 450, University of Illinois, Urbana, August 1978.
- 4. Moehle, J. P., and M. A. Sozen, "Earthquake-Simulation Tests of a Ten-Story Reinforced Concrete Frame with a Discontinued First-Level Beam", Civil Engineering Studies, Structural Research Series No. 451, University of Illinois, Urbana, August 1978.
- 5. "Potential Utilization of the NASA/George C. Marshall Space Flight Center in Earthquake Engineering Research", Earthquake Engineering Research Institute, Berkeley, California, December 1979.
- U. S. Japan Planning Group, "Recommendations for a U.S. Japan Cooperative Research Program Utilizing Large-Scale Testing Facilities", Earthquake Engineering Research Center, University of California, Berkeley, Report No. EERC -79/26, September 1979, pg. 92.
- 7. Staffier, S. R. and M. A. Sozen, "Effect of Strain Rate on Yield Stress of Model Reinforcement," Civil Engineering Studies, Structural Research Series No. 415, University of Illinois, Urbana, February 1975.
- Penzien, J., J. Bouwkamp, R. W. Clough, and D. Rea, "Feasibility Study Large-Scale Earthquake Simulation Facility," Earthquake Engineering Research Center, University of California, Berkeley, Report No. EERC-67/1, September 1967.

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Table 1

Actuator Requirements (2 per level)

Level	<u>Capacity</u> (kips)	Bore (inches)	<u>Stroke</u> (inches)
10	35	5	48
9	35	5	48
8	35	5	38
7	35	5	38
6	25	4	29
5	25	4	29
4	25	4	19
3	25	4	19
2	15	3	10
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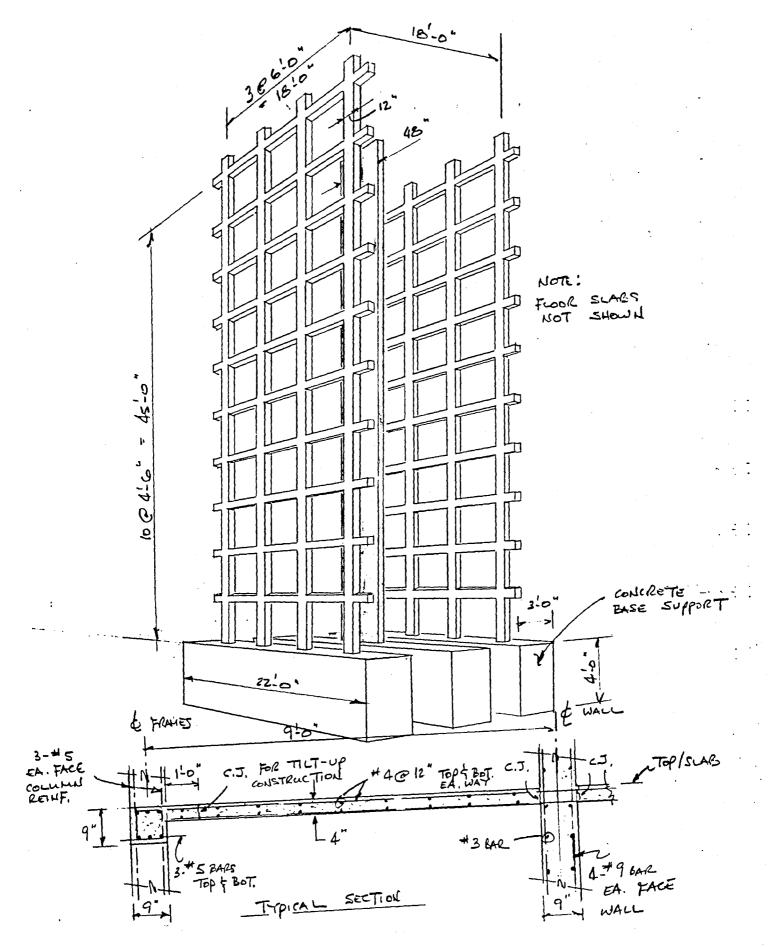
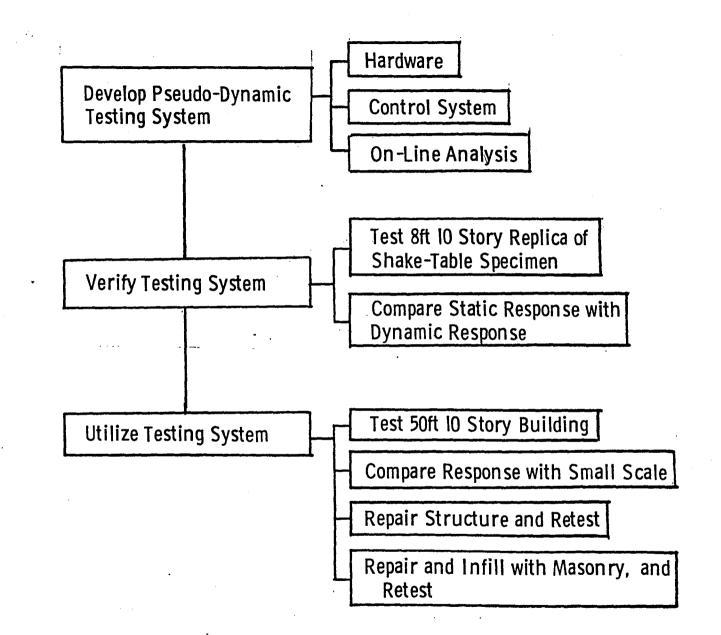
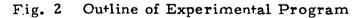


Fig. 1 Test Structure





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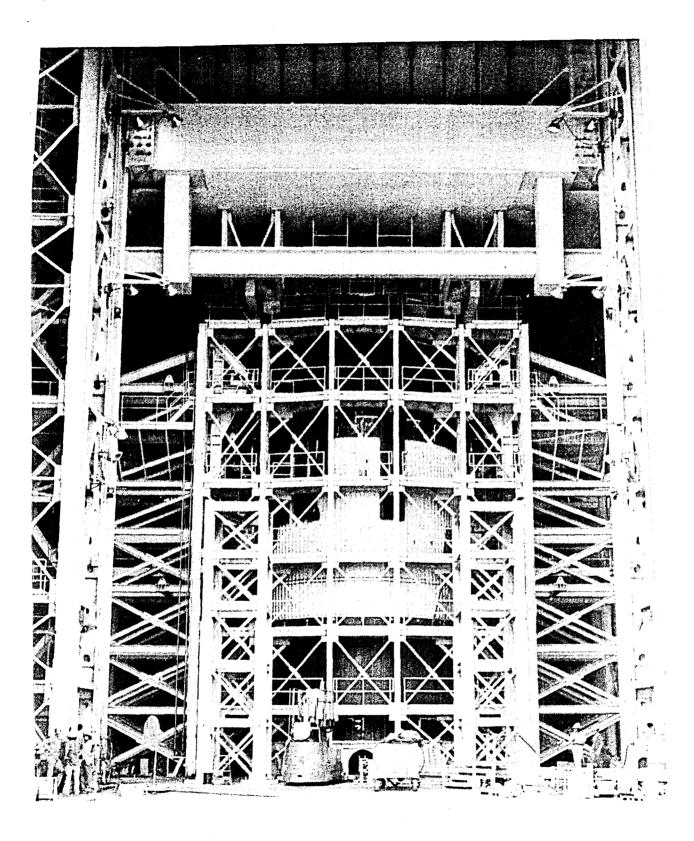


Fig. 3 Reaction Structure

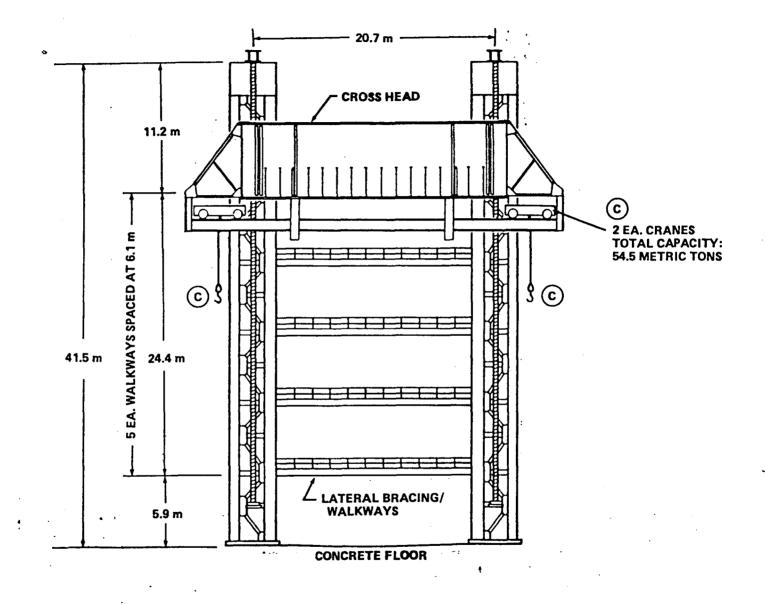


Fig. 4 Structural Test Tower

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

MSFC PERSONNEL MANAGEMENT TASKS: RECRUITMENT AND ORIENTATION OF NEW EMPLOYEES

Prepared By:

Thomas A. Brindley, Ph. D.

Academic Rank:

Associate Professor

University and Department:

The University of Alabama in Huntsville Department of Education

NASA/MSFC: Office Division

MSFC Counterpart:

Date:

Contract No.:

Personnel Office Personnel Development

James R. Johnson

August 1, 1980

NGT-01-002-099 (University of Alabama)

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MSFC PERSONNEL MANAGEMENT TASKS: RECRUITMENT AND ORIENTATION OF NEW EMPLOYEES

ΒY

Thomas A. Brindley, Ph. D. Associate Professor of Education The University of Alabama in Huntsville Huntsville, Alabama

ABSTRACT

In order to encourage highly motivated young students to learn about NASA and consider it for a career, a formal program will be initiated whereby selected students can work on a voluntary basis at Marshall Space Flight Center. These students will work under the guidance and supervision of expert NASA employees. The first task has been to develop the working plan and procedures for this program, called Student Volunteer Service Program, in the writing of MSFC official guidelines, the Marshall Management Instruction (the MMI) which is a binding document that defines policy and establishes procedures and guidelines. Particular considerations written into the MMI after numerous consultations, interviews, and discussions about a satisfactory policy, include: arrangements to be made between the student, the school authorities, and concerned MSFC employees; management of the work assignments; and procedures for the student's welfare and safety.

The second task has been the development of a Recruitment Brochure for the attraction of new employees, especially scientists and engineers. Since private industry has certain advantages in hiring, particularly in offering higher initial salaries, it is imperative that NASA stress the valid reasons why working at MSFC would be of great benefit to the prospective employee. Thus, the brochure needs to stress aspects of ongoing NASA projects which show how various disciplines of science and engineering apply.

The third task assigned has been to develop a plan called Orientation of New Employees.¹ The main intention is to establish a procedure whereby new employees can best become acquainted in their new work with a minimum of stress involved in coming to a new organization. This plan attempts to schedule various experiences that will enable the new employee to understand not only the particular work situation for which the employee is assigned, but to show the overall mission and operations of the Marshall Space Flight Center. The plan is to indicate the types of orientation believed necessary to aid the new employee in adjusting to work at MSFC.

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INTRODUCTION

The Personnel Development Division of the Personnel Office requested that three tasks be accomplished by this NASA/ASEE Summer Faculty Fellow. The tasks were to investigate the needs and ramifications of several aspects of recruiting and orientation of employees and to write up a study or report on each of the subjects, taking each project as far as it could be carried in the ten weeks. Specifically, these management tasks were:

- 1) the orientation of new employees;
- 2) a recruiting brochure for the hiring of new scientists and engineers; and,
- 3) a student volunteer service program.

These three tasks have been completed and are reported here in respective sections.

The research undertaken involved the reading of all materials pertinent to past actions on the subjects and numerous interviews, consultations, and discussions with concerned MSFC employees about satisfactory policies related to these tasks. The tasks also demanded that this writer understand as much as possible about NASA and MSFC in a holistic, comprehensive way. Since the author had little previous acquaintance with MSFC in either a technological sense or in an organizational management sense, it was necessary to learn about NASA/MSFC in a short time so that the tasks could be begun. Thus, in effect this writer experienced his own orientation program.

VOLUNTARY STUDENT SERVICE PROGRAM

In 1978, the MSFC Personnel Office in response to various university and school requests in the Huntsville area initiated inquiries into the feasibility of forming a plan to describe means whereby exceptional students could work at MSFC on a voluntary basis. The initial exchange of memorandum and letters regarding the subject were described as the "Worksite Experience Program."

Laboratory supervisors were asked if they had needs in their organizations for visiting students. Several responded that they would be interested in allowing students to work in their laboratories. The Associate Director of the Center and the Director of Science and Engineering were especially interested in pursuing this program.

Coincidentally and unrelated to local efforts, guidelines were being discussed at the national level to provide for voluntary student services. The Civil Service Reform Act of 1978, allowed agencies to devise memoranda or official guidelines which would establish workable programs of this nature. MSFC temporarily delayed pursuing this further until the assignment for this NASA/ASEE Fellow.

Until the Civil Service Reform Act of 1978, there were no provisions for agencies of the Federal government to accept voluntary or gratuitous services from interested persons or parties. Thus, visiting individual researchers and interested students could find no means to offer services to a Federal agency because no employment status existed for them. There were no arrangements or agreements to allow students to take part in government programs gratuitously. Thus, a talented high school student could not work at MSFC in a laboratory where the youngster could gain in knowledge and experience. Such an opportunity would also encourage the youngster to think about future employment at NASA after completing college work.

The first task assigned and completed was to write a proposed Marshall Management Instruction (the MMI) which is a binding document that defines policy and establishes procedures and guidelines. The MMI sets up a formal mechanism whereby selected, highly motivated young students can work on a voluntary basis at MSFC. The following study plan, following the format of the Marshall Management Instruction, describes the Voluntary Student Service Program.

STUDY PLAN: THE MARSHALL MANAGEMENT INSTRUCTION (DRAFT)

SUBJECT: Student Volunteer Service Program

1. PURPOSE

- a. To provide a means whereby qualified students can volunteer their services to the Center in order to gain experience and knowledge.
- b. To establish Center policy and procedures and to identify responsibilities in the management of the Student Volunteers.
- c. To establish guidelines for the Student Volunteers while they are working at the Center.

2. SCOPE

This instruction is applicable to all elements of MSFC.

3. REFERENCES

- a. FPM Chapter 308, Subchapter 7.
- b. FPM Letter 308-16, November 8, 1979.

4. POLICY

- a. High School students who show outstanding promise in their school work or school activities as determined by their school teachers and administrators in all areas of school study, especially, but not limited to science subjects, may wish to receive additional knowledge and expertise and be provided advanced opportunities to learn through on-the-job experience. These students will be allowed to work in a nonpay status on a voluntary basis at the Marshall Space Flight Center during a specified time and under the direction of a designated NASA employee.
- b. Each student assignment will be established and conducted through written agreement between the Marshall Center and the educational institution or with the organization so

designated by schools or boards of education.*

- c. The participating student is not to be considered a Federal employee for any purposes other than injury compensation or laws related to the Federal Tort Claims Act. Service is not creditable for leave accrual or any other employee benefits.
- d. Assignments will be tailored individually at the chief discretion of the project officer (Mentor).

5. DEFINITIONS

- a. The definitions of "volunteer service," "student," "half-time student," and "agreement" are those contained in FPM Chapter 308, Subchapter 7 and FPM Letter 308-16.
- b. Mentor -- the MSFC employee responsible for the student's work, welfare, and learning experiences, who is in charge of the student's assignments, scheduling, reporting, and evaluation, and who supervises the student.

6. BACKGROUND

The authority for the acceptance of unconditional gifts and services donated to the agency is contained in the National Aeronautics and Space Act, Section 203 (c)(4). Based upon this authority the Center has been developing a program to allow participation in Center activities by students who are willing to donate their time and effort, in response to interest expressed by local educational institutions. The Civil Service Reform Act of 1978 authorized the acceptance of volunteer service and the Office of Personnel Management implemented regulations outlining procedures for the development of such programs government-wide.

7. GENERAL PROVISIONS

- a. Responsibilities for the student's work at the Center shall be mutually entrusted to the Mentor. Participation in the program is conditioned upon the student's consent by written agreement to certain specified guidelines. Although students are not considered Federal employees, they will be allowed to work alongside and at the direction of Federal employees at the Center.
- b. This program is limited to students <u>enrolled</u> not less than half time in high school, trade school, technical or vocational institution, or junior college, college or university.
- * Agreements and forms, as attachments, are not included here.

- c. Students receive no compensation.
- d. The student offers service (comes to work at NASA) with permission of sponsoring school or institution where student is enrolled.
- e. Nominations for participation will be made by the sponsoring school or institution where the student is enrolled. Selection shall be within the sole discretion of the Center.
- f. The student will not replace or displace any Center employee or staff a vacancy.
- g. The student will not be used to replace or substitute for routine and/or clerical duties while on assignment.
- h. The student will not be allowed to enroll or partake in MSFC training courses.
- i. Students will be at least 16 years old and must be at least in their junior year in High School. Students must be citizens of the United States. Participation shall not be based upon race, color, religion, sex, national origin, age, mental or physical handicap, marital status, or political affiliations, as prohibited by law, rule, or regulation.

8. **RESPONSIBILITIES AND PROCEDURES**

- a. The academic institution will be responsible for:
 - Requesting in writing to the MSFC Center Director placement of nominated students in designated disciplines.
 - (2) Recommending students who have demonstrated a serious interest in gaining work experience at the Center and who have demonstrated outstanding curricular and/or extracurricular achievements as deemed appropriate by the academic institution.
 - (3) Executing and complying with the terms of the Memorandum of Understanding with MSFC.
- b. The MSFC Personnel Office will be responsible for:
 - (1). Selecting participants from those nominated by their educational institution.
 - (2) Identifying work locations.
 - (3) Appointing a Mentor to supervise the student during the student's assignment at MSFC.
 - (4) Negotiating appropriate agreements.
 - (5) Documenting volunteer service on Standard Form 50 (SF-50), "Notification of Personnel Action."
 - (6) Establishing an Official Personnel Folder (OPF).
 - (7) Evaluating the program.
 - (8) Submitting any required reports.
 - (9) Managing a one week orientation program to acquaint the student with the overall NASA-MSFC mission.

- c. The MSFC Security Office will be responsible for:
 - Providing proper authorization of the student's special status while assigned or visiting the Center.
 - (2) Acquainting the student with necessary security measures.
 - (3) Limiting the student only to non-security materials and areas.
- d. The MSFC Mentor will be responsible for:
 - (1) Supervising the student's work and learning experiences.
 - (2) Supervising the student's assignments and schedules.
 - (3) Assigning the student a definite worksite and monitoring the student's activities.
 - (4) Evaluating the student's performance and submitting any necessary reports to the Personnel Office.
 - (5) Maintaining informal relationships with responsible parties of the student's academic institution.
- e. The student will be responsible for:
 - (1) Complying with rules and regulations governing the internal operations of the Center. Any infractions of responsibilities may be cause for termination of the program.
 - (2) Being at the work location during the hours specified in the work experience Memorandum of Understanding or in other designated places upon approval of the Mentor.
 - (3) Following the authority of the assigned Mentor.
 - (4) Contacting the Mentor if absence from work is necessary.
 - (5) Caring for and properly maintaining all equipment and facilities used.
 - (6) Providing transportation to and from the MSFC worksite.
- f. Mutual agreements between parties:
 - A written agreement, called the Memorandum of Understanding, will be signed by the Director of the Personnel Office and the representative of the participating institution. The agreement will explain the general assignments, time of assignment, and any special requirement desired by the parties.
 - (2) A "Notification of Personnel Action," Form 50, will be completed and signed by Department Head and Applicant.
 - (3) An "Application For Use of Research Facility, Equipment, or Computation Service" will be completed and signed by designated persons.
 - (4) A "Memorandum Agreement Relative to an Unconditional Donation of Services and Acceptance Thereof," will be completed and signed by designated persons.

THE RECRUITMENT BROCHURE

The second management task has been the development of a recruitment brochure for the attraction of new employees, especially scientists and engineers. In 1968, the first reduction in force (RIF) affected MSFC. From time to time thereafter various reductions of employees occurred through 1979, so that the work force which was once over 7000 in the heyday of the Apollo-Saturn missions fell by Summer, 1980, to 3600 employees. Coupled with this, there are over 1500 employees who have reached twenty years of service and are eligible for retirement. The rate of employee attrition will certainly increase in the next five years unless new people can be brought in. Although there has been some hiring in critical skill areas over the last twelve years, there is abundant need to recruit and hire scientists and engineers at MSFC to maintain a strong and active work force. Even though there is a 3561 end of fiscal year 1980 employment ceiling at MSFC imposed by Congressional budgetary considerations, new hiring is very necessary.

Recruitment teams are sent to selected universities each year. There are numerous kinds of materials, but there is no special handout available for recruiters to utilize which is aimed definitely at attracting employees to this center to work on the kinds of projects in process for which MSFC is renowned, and to praise the advantages of living in the Huntsville, Alabama area. A recruitment brochure is deemed necessary to enhance the search for new employees of critical skills in science and engineering.

There were several important considerations that were discussed in the numerous interviews and inquiries taken by this researcher in order to determine a final plan for the nature of the brochure. It was determined that the particular engineering and scientific skills in which a prospective employee has been educated should be shown in their direct relationship to the kinds of projects and activities being developed at this particular center. A prospective employee should feel that his or her university preparation and past experience will be properly utilized at MSFC. The employee has to feel that there are opportunities to work on significant projects, to acquire a sense of mission and purpose, to see one's work in meaningful relation to the whole, and to work with materials and processes that enhance one's skills and draw upon one's best talents. The new employee has to be shown that MSFC allows for personal and professional growth and development and that MSFC provides abundant opportunities for advancement and recognition. Since private industry has certain advantages in hiring, particularly in offering higher initial salaries, a recruiting brochure must show how the Federal pay and

promotion scale and the MSFC organization allow an employee to gain in pay and status rather rapidly over the years. New employees are concerned about the area in which they live. Thus, the many advantages of the good life in the Tennessee Valley should be described.

The actual development of the brochure itself included a number of steps. The researcher had to ascertain what the needs are at MSFC and what kind of employees are being sought. The researcher had to learn what MSFC offers that would be of concern and interest to possible new employees. What especially attracts recent college graduates and/or experienced hands? Can NASA/MSFC attract candidates of the highest quality in the face of considerable competition from other governmental organizations and from private industry?

Various federal agencies, other NASA centers, and some aerospace companies were contacted about their own recruiting efforts. After this information was gathered, the researcher had to discuss the plan of the brochure with various supervisors, chiefly the head of recruitment.

Then the researcher made several plans, called thumbnail sketches. Priorities of arrangement and order were determined. The final sketch was presented to the graphics department to make a draft mockup. This is the stage reached by the end of the seventh week. Thereafter, each draft must then be developed further and refined. Appropriate pictures must be acquired and adequate text must be written. Then a final mockup is made by the graphics department. This is presented to various directors for approval and then submitted to the printers for the final edition.

The organizational pattern of the draft mockup recruitment brochure as completed follows in this order. First, an attractive cover is planned to show the chief ongoing project for which MSFC is mainly responsible -- the Space Shuttle itself. At the beginning of the booklet the Director of the Center introduces MSFC to prospective employees. The Center itself is shown and the geographic relationship of MSFC to other field centers is represented. Chief accomplishments of NASA in the past, in the present, and planned for the future are pictorially shown. Then the main missions and projects of the Marshall Center are depicted. These include the Space Shuttle, the Spacelab Payloads, various important scientific laboratories, and some of the spin-offs, materials processing in space, and special assigned tasks at MSFC, such as solar heating and cooling. Integrated into the portrayal of these various missions are references indicating how the education and experience of qualified engineers and scientists work on each of the projects in particular. Then some pages describe briefly how future employees will benefit by

working at MSFC. Several pages show how employees have special advantages and opportunities at MSFC in advancement, career development, pay and status, training programs, and professional enhancement. Finally, the attractive surroundings of the Tennessee Valley and the numerous educational, cultural, and social advantages of this area of America are portrayed. The sense of living in a multicultural, pleasant, and attractive community is stressed.

The various thumbnail drafts and the first mockup from graphics are already completed at the time of this writing. However, the bulk of the draft brochure precludes it being inserted as a part of this report. The first draft mockup brochure serves as a springboard for its final development and incorporates and organizes the many conceptual ideas into a meaningful plan for further action. The ongoing continuation of the recruitment brochure lies in the hands of the Personnel Director.

ORIENTATION OF NEW EMPLOYEES

The third task completed has been to develop a study plan called the Orientation of New Employees. The main purpose is to establish a procedure indicating guidelines whereby new employees can best become acquainted in their new work with a minimum of stress involved in working in a new organization.

Presently very little is being done at MSFC in a formal manner to orient new employees about the ways of the Center. There is a one time meeting given by a representative of the Personnel Office, as presented by law in the guidelines of the Office of Personnel Management to describe the necessary work conditions and benefits of employees. This is not a true orientation since it is limited to a description of procedural matters relative to all NASA employees.

The Director and supervisors of the Personnel Office recognize the need for a completely developed orientation plan. But there are differing opinions among them and among various supervisors and employees throughout MSFC about the merits and the type of orientation that would be necessary. Informal inquiries taken among several new employees also show a division of opinion about the nature and need of an orientation program for the recently hired.

Some feel that orientation efforts are a waste of time, others feel they are beneficial. Some believe orientation meetings should be brief and narrowly construed. Others feel that meetings should be very comprehensive of the entire MSFC operations beyond one's own workplace. Others feel that there is no need to learn about areas unrelated to one's work. Some feel that an orientation schedule should occur within the first few days of a new person's hire or, at most, but a week long. Others feel that a proper orientation cannot be hurried or overloaded in a short time frame, but should be extended over several months. There are many concerns about the nature of an orientation program. Does it encourage better morale in the work force? Does it enhance the sense of mission? Does it build cooperative efforts?

The main argument seems to be whether a new employee needs to become acquainted with the general workings of the new organization, or whether the new employee needs to learn only about one's own particular job and only needs to learn about the wider ramifications of the organization as the person's contacts develop from the job site. Although this might be an individual matter, there seems to be a broad difference in approach between generalists and particularists based upon one's own research training and education. Some scientists and engineers want to work alone and without much to do with others. Some scientists and engineers wish to see how their own discipline relates with others and wish to ascertain how their own particular skills might be relevant to other parts of the whole center.

Management, on the other hand, has to determine what is best for the new employee and what kind of orientation would lead to the best satisfaction of the missions and purposes of the center as a whole. There are conflicting views here, too. Some personnel supervisors feel that the integrated nature of much of the work at NASA demands a general understanding of MSFC as a whole and of different projects and laboratories other than one's own specialty or assigned area. Other supervisors feel that an extensive orientation program would be too time consuming on the parts of both the participants and the presenters and would be too costly.

The orientation plan devised by this researcher incorporates many approaches to meet expected needs of new employees to help them adjust to work at MSFC. This plan attempts to schedule various experiences that will enable the new employee to understand not only the particular work situation in which the employee is assigned, but to show the overall mission and operations of the Marshall Space Flight Center. The study plan which follows here is devised to allow for maximum opportunities, but it may be pared down or condensed if so desired, as the plan moves up the ladder in the approval process.

STUDY PLAN: ORIENTATION OF NEW EMPLOYEES

PROBLEM:

Since 1968, when MSFC's first reduction in force took place, NASA-MSFC has been in a reduced hiring mode. As a result, new employee orientation during this time has been based primarily upon satisfying general information needs. Much of a new person's orientation is conducted informally at the work site by the employee's immediate supervisor. The Personnel Office, on the other hand, gives the new employee basic procedural information concerning OPM regulatory requirements and benefits. This is usually done on a one-to-one basis and later in a short general meeting for those brought on-board that week.

In the last year, NASA-MSFC hiring of new employees has increased. Yet, no plan has been developed to acquaint these new employees with the overall mission and purposes of NASA or with a proper perspective in seeing the relationship of one's own work to the whole. This has generated a need for a more comprehensive, in-depth orientation of new employees.

OBJECTIVES:

The objectives of a plan for the orientation of new employees are:

- 1. To inform each employee about the basic requirements of adjusting to the new organization. This includes acquiring a knowledge of benefits, hospitalization, security, NASA-MSFC regulations and various other procedural and administrative matters.
- 2. To provide a method whereby the employee is thoroughly oriented to the work situation. Presently, the employee is acquainted with the work situation in an informal manner.
- 3. To provide a general understanding of the overall NASA mission and to show how the various organizations work together at MSFC to achieve national goals. Not only will the employee begin to sense the importance of NASA-MSFC as a whole, but also will learn how the new employee's particular work fits into a wider framework.

This study deals with these objectives and outlines the structure and content of a proposed orientation plan that will satisfy the needs of the new employees and be beneficial to NASA.

PROPOSAL:

The "Orientation Plan for New Employees" will consist of three phases as follows:

- 1. Administrative Orientation -- procedures, responsibilities and rights, guidelines and benefits.
- 2. Work Site Orientation -- the particular tasks of the employee at the immediate place of work.
- 3. NASA-MSFC Orientation -- the overall perspective and acquaintanceship of the mission of MSFC as a whole.

PHASE ONE: ADMINISTRATIVE ORIENTATION

1. Personnel Staffing Specialist Meeting

The very first acquaintanceship of the new employee upon being hired will be a scheduled visit with an assigned Personnel Staffing Specialist (PSS) on an informal, one-to-one, personal basis. The most immediate concerns of the new employee, such as housing, locating main facilities of the Center, work hours, health plans, and pay procedures, will be discussed. The main attempt here is to give the new employee a feeling that he or she is welcome and has a place to find information immediately. A designated number of new employees will be assigned to each of the Personnel Staffing Specialists. Even new employees who have worked with other federal agencies would be expected to participate in this orientation. The PSS will "follow through" with each of the new employees over the course of the entire three phase orientation period. From time to time the PSS will contact the newcomer to inquire about the person's welfare at MSFC. The PSS should give the newcomer a contact phone number on the first visit. Also, the new employee will be given a packet, which includes necessary information about MSFC and Huntsville, employee unions, equal opportunity, and regulations, to keep and study.

2. Personnel Administrative General Meeting

At a designated time, preferably at the beginning of every two week pay-period in order to keep in phase with payroll records, the new employees hired up to that date will meet as a group for a general orientation conducted, as presently done, by the Personnel Office. The necessary items scheduled for discussion will follow the "New Employee Orientation Checklist," MSFC Form 3688 (April, 1968), which is signed upon completion by both the employee and Staffing Officer and kept in the employee's 201 File. Discussion of any particular points will be encouraged and details will be provided as needed by the new employees.

PHASE TWO: WORK SITE ORIENTATION

It is of utmost importance that the new employee learn to feel at home in his/her own work environment as soon as possible. The adjustment to a new situation will produce some frustrations, but it need not be stressful if proper considerations are given. The new employee will be under the authority of the immediate supervisor. This person should pay great attention to the psychological welfare of the newcomer. Introduction to all fellow employees, an opportunity to see others at work and to see who is responsible to whom, a review of work procedures, and an acquaintance with the location of key facilities and equipment are some of the concerns necessary to the well being of the new employee. This acquaintanceship period takes time, cannot be rushed, and should include some free time to digest the new information. In essence, one is learning some of the informal ways of the organization.

During this time the Personnel Staffing Specialist will be in contact to review progress and handle problems. This is in line with regular personnel procedures which require a ninety day follow-up. Sometimes, several other orientation activities (of Phase Three, below) will be provided during the time the new employee is on the job in this early acquaintaceship period. Such activities might include the bus tour of MSFC or selected lectures and/or interviews by NASA specialists on particular topics. No additional paperwork is required by the supervisor other than the checklist now used routinely by the supervisor in reviewing the completion of special tasks in this part of the orientation. The intent of work site orientation is to keep procedures as natural and informal as possible.

PHASE THREE: NASA-MSFC ORIENTATION

1. The major purpose of this phase is to assure that new employees are given opportunities to see their own particular work in relation to the whole of MSFC and of the entire NASA agency. It is important to the overall mission of NASA that integration and cooperation of efforts be understood by all employees not only better to achieve the goals of NASA, but to enhance esprit de corps. Also, an understanding of the interrelationships of NASA's many projects often leads to a better feeling about one's own work and aids an individual in one's own personal and professional advancement.

- 2. Since the new employee will now be at work on the job site, this phase of the orientation must be scheduled so as to cause minimum disruption to the person's work schedule. It is proposed that this general orientation be staggered over a number of weeks to allow the new employee sufficient time to digest the new information and to prevent overburdening the offices of those who present materials by taking time out of their regular workload.
- 3. Administration of all aspects of Phase Three, including scheduling and monitoring, will be under the management of the Personnel Office.
- 4. All new employees should be involved together as a group in a one day common orientation. In the morning a bus tour should take the newly hired to various locations of MSFC. They should visit the main projects, test stands, certain laboratory facilities, museum mockups, and special projects, such as the Neutral Buoyancy Facility.

In the afternoon a general meeting should be arranged in the Tenth Floor Conference Room of Building 4200. At this time various selected Center management officials will introduce themselves and conduct a presentation on various important general MSFC topics. Such topics might include a general discussion of the Space Shuttle or of the importance of coordinating payloads.

A representative of the Office of the Director or one of the Executive Staff Offices should give an introductory address called the Center Overview, followed by other senior management speakers, who can lend an aura of importance to the new hiree's appointment into the MSFC workforce. This Conference Room meeting should include as many new hirees as possible at the same time.

- 5. Up to this time the orientation is the same for all types and categories of employees: scientists and engineers, administrative and professional, technicians, clerical and operational. At this point the orientation is divided into two modules. Each module applies to different categories of new personnel. The difference in orientation lies in the degree of depth and scope. The two modules are:
 - 1) Scientists and Engineers, Business and Professional, and

Technicians (NCC 200/700, 600, 300)

- 2) Clerical (NCC 500).
- New employees of both categories (modules) should visit certain basic projects, labs, or offices in common and as a large group. These include:
 - 1) selected science and selected engineering laboratories;
 - 2) key projects, such as Shuttle propulsion.
- 7. The module plan could be modified or compressed if the plan involves too much time of the employees and of the supervisors by cutting down on the staggered sessions and by having the new employees meet together in larger groups.

Module One: Scientists, Engineering, Technicians, Business and Professionals

- On a bi-weekly basis a certain number of new hirees will meet as a group to interface with each other, to share experiences, and occasionally to listen to guest speakers about special projects. The purpose of this meeting is to provide for some identifiable group loyalty and cohesion to disparate members of the MSFC work force in order to build up a sense of common solidarity across NASA as a whole, to build a sense of community and shared associations, and to provide for cross-fertilization of ideas. (Precedence for this has already been established by the NASA/ASEE Summer Faculty weekly "fellowship" meetings).
- 2. An opportunity for new employees to meet key personnel can be developed by inviting such people to these bi-weekly meetings to converse with the group, or to schedule specially arranged times when small numbers of hirees can visit in selected offices.
- 3. Rather fundamental kinds of programs, projects, and interfaces ought to be included in the itinerary. These might include:
 - Advanced Systems, planning layout
 Examples: (1) Description of all future projects;
 (2) Coal Casification;
 - (2) Coal Gasification;
 - (3) Nuclear Waste Disposal;
 - (4) Solar Energy Use.
 - b. Space Shuttle operations
 Examples: (1) Configuration Change Board of Solid Rocket

Boosters in Building 4610;

- (2) Telecon Level II PRCB in Building 4202, Room 411.
- c. Payload and Spacelab Projects
 - (1) Overview.
- d. Science and Engineering Laboratories
 - Examples: (1) Metallurgical analysis and use of electron microscopes;
 - (2) Cell separation in zero gravity
- e. MSFC Administrative Support
 - Examples: (1) Personnel Training;
 - (2) Procurement;
 - (3) Plans and Analysis Office;
 - (4) Overall administration.

Module Two: Clerical

- 1. Since the new employees of this group will not have had extensive, formal scientific and engineering education, the presentations should be given in less complicated technical detail. However, the general nature of the various projects and tasks indicated in Module One should be of interest to this group. Thus, a schedule should be developed in the Personnel Office for this module following the same pattern.
- 2. Scheduling for both modules will be done in accordance with new hiring activity and with the concern to distribute presentations among many different people and offices.
- 3. A special session for Module Two would be a general, overall tour and explanation of a selected number of key installations. This would be different from the usual NASA tour. This would encompass such office functions as procurement, personnel training, computer services, and the public affairs office. Selected scientific laboratories would be included on the itinerary. This tour would be planned to encompass an entire morning.

ORIENTATION PLAN RECOMMENDATIONS:

- 1. That a comprehensive, in-depth new employee orientation be implemented at MSFC.
- 2. That the administration and coordination of the Orientation of New Employees be the responsibility of the Personnel Office.
- 3. That an MMI be published identifying Center policy and assigning duties and responsibilities.

CONCLUSIONS AND RECOMMENDATIONS

The Personnel Office of Marshall Space Flight Center assigned three management tasks to this researcher. All three were completed and the results were presented to the Personnel Director. A meeting of selected supervisors was called so that the researcher could orally present a status report on these topics. They expressed satisfaction with the projects and the work done. Final suggestions were made, although constant surveillance and assistance by these same people had been given continuously during the entire time of investigation and writing.

This report has described the nature of the three tasks and the substance of their results. Since each of the tasks is a topic that is ongiong in nature, the final products will not be ready until all reviewers and authorities give approval. Undoubtedly some changes and revisions will be made. Nevertheless, the basic structure, the main outline and format, and the conceptual plans have been determined at this point.

There was no shortcut in the undertaking of these tasks. First, a thorough orientation and grounding into the nature of the scientific and engineering projects which make up the mission of NASA-MSFC and the nature of the administration and organization of MSFC had to be obtained by this researcher. Then, many different viewpoints and opinions had to be gathered and discussed about the specific nature of the tasks. Materials relevant to the tasks were read and analyzed, but these usually consisted of memos, letters, and similar works previously done. Since most of these materials are on file in various offices at MSFC, they do not form a descriptive bibliography or a list of references, and, consequently, are not stated in this report. Finally, each draft of each task had to be reviewed by an authority on the subject before the next revision could be written. Most often the suggestions of several people were incorporated into the next stage of the writings.

It is recommended that these three projects be continued immediately until their final approval is given by the Director of the Center. A Marshall Management Instruction (MMI) on the Student Volunteer Service Program should be completed in final form following the model presented here. The Recruitment Brochure should follow the general theme as worked upon to this point. Pictures should be selected and minimal text written. A comprehensive Orientation Plan for New Employees should be implemented at MSFC and an MMI should be written to identify Center policy and guidelines. All three of these tasks should be administered by the Personnel Office of MSFC.

1980 NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

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THE EFFECT OF TIME DELAY IN THE 6 DOF MOTION SYSTEM AND ITS COMPENSATION

Richard Brown, Ph.D.
Associate Professor
University of Alabama University, AL Department of Mathematics
Data Systems Laboratory Engineering Computers Division Simulation Systems Branch
Frank L. Vinz
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ANALYSIS AND IMPROVEMENT OF CONTACT DYNAMICS SIMULATION USING THE SIX DEGREE OF FREEDOM MOTION SYSTEM

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Richard C. Brown, Ph.D. Associate Professor of Mathematics University of Alabama, Tuscaloosa, Alabama

ABSTRACT

The 6DOF motion system has been employed for closed loop simulation of the contact dynamics involved in the docking of two satellites (for example TRS and Skylab). Under most initial conditions however the system gives inaccurate results. Typically in the simulation if docking fails to occur the rebound velocity of the TRS exceeds its approach velocity, thus generating an artificial and unacceptable gain of momentum in the system.

It has long been well known that the presence of a time lag can generate instability in closed loop systems. To see if the time lag (about 200 ms.) in the 6 DOF system can satisfactorily account for its errors, a simple computer mode of the system incorporating this lag and other revelant parameters has been implemented. The model in fact does predict all the observed rebound errors with reasonable precision.

Also, since it is probable that the time delay can only be reduced but not eliminated, various compensation schemes have been devised. Two give adequate correction for delays under 100 ms for a non-preloaded probe. More powerful schemes may be adequate for the present 200 ms. environment and the presence of a preload.

Our study relies heavily on the mathematical theory of second order differential equations with retarded arguments and standard tools of numerical analysis.

ACKNOWLEDGEMENT

I want to express my thanks and appreciation to the members of the Simulation Systems Branch - especially F. L. Vinz, David Allen, J. S. Spear and Bert Gover for many helpful discussions concerning the 6 DOF system.

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NOMENCLATURE

ተ	-	Time lag of the 6 DOF system					
۵t	-	Sample frame time of the AD converter between the sensor and the EAI 8900					
F(t)	-	Sensor force signal					
۵F	-	Change in force signal in time step					
r	- 1	Probe compression					
η.	-	Maximum probe compression					
K	-	Hooke's constant of probe					
М	-	Mass of probe					
е	-	Base of the natural logarithmns					
v	-	Initial velocity of probe					
vţ	-	Observed rebound velocity of probe					
∀(ŧ)	-	Velocity of probe					
a(t)	-	Acceleration of probe					
IVP	-	Initial value problem					
w,²	-	K/M					
Fp	-	Preload of probe					
X.		n'+ Fo/K					
×T	-	×(t-T) + +					
*	-	Convolution, $f = q := \int f(e-a) g(s) ds = \int f(s) g(e-s) ds$					

F_t - Force threshold
 λ(t) [a,b] - Characteristic function based on the interval [a,b]
 h - Step in a digital simulation

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M	-	Step parameter ; 🕆 = h.M
∂(h [#])	-	Error proportional to h^{k}
∨ _n	-	v (nh)
qin	-	a(nh)
ve	-	Compensated version of ${f v}$
x _c	-	Compensated version of x

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I INTRODUCTION

The LINK Six-Degree-of-Freedom (6 DOF) Motion System has been employed by the Marshall Space Flight Center for the physical simulation of the contact forces and dynamics in the docking of two satellites - for example the Teleoperator Retrieval System (TRS) and Skylab. A similar device was developed at the Johnson Space Center in Houston for contact dynamics simulation prior to the Apollo-Soyuz mission of 1975.

In this section we give an informal description of the system and its performance. Some illustrations, photographs, and values of some key parameters may be found in the Appendix. See in particular the block diagram, Fig.10.

A docking probe assembly is placed on the platform of the LINK 6 DOF Motion System. The platform is free to move with any general motion within its limits of travel. Above the probe is a docking collar or other target which is fixed. The software of the EAI 8900 hybrid computer is capable of simulating a set of initial conditions which might include velocities, accelerations, and rotations preliminary to docking.

Since, as mentioned above, the target is stationary, this information is calculated relative to a coordinate system fixed in the target. The initial conditions are translated by the CAI digital computer into commands to change lengths of the six legs supporting the 6 DOF platform. These commands are executed by the hydraulic actuators in the legs causing the 6 DOF system and the probe to perform motions appropriate to the initial conditions.

When the probe is in contact with the target, sensors, also placed on the table, record the forces and torques encountered by the probe. In the open loop mode this information is recorded on strip charts for later use. If desired software can be developed so that the probe simulates dynamic interaction with This option has so far been rejected since the the target. resulting motion would not be generated by the forces actually encountered in the contact phase. A more useful choice is closed loop simulation where the forces are used in a feedback loop through the EAI 8900 and the CAI computers to control the motion of the probe itself. Basically two integrations of the hybrid subsystem give velocity and position functions of the TRS relative to the target. As in the case of the initial conditions the position function is converted into leg length commands. The result is that the probe is supposed to move as if it was freely responding to its own contact dynamics. This mode when perfected should be a valuable tool in the testing and design of docking hardware.

In preliminary testing of the system, the closed loop simulation has been radically simplified so that probe-target interaction is "head on;" only (vertical) translational motion is accepted. Also no thrusts firings of the chase vehicle are allowed. Table 1 presents results of some of these tests.

Obviously apart from the unacceptable level of noise, there is an artificial and unwarranted gain of momentum in the system. The analysis and attempted resolution of this problem is the central concern of the present study.

We close this section with a brief survey of the rest of the paper. In Section II a simple mathematical model based on the time delay inherent in the system is developed. It is also shown how the model can be modified so that it is "noisy." Section III developes several methods to compensate for the delay. The last two sections contain respectively data and graphs illustrating our results and recommendations for further action. II THE EFFECT OF TIME LAG ON THE 6 DOF CLOSED LOOP SIMULATION

It has long been known that the presence of time delay can affect the stability of a closed loop feedback system - see for example.

Minorsky [10], [1] or Callendar, Hartree and Porter [4]. It has been verified that a time delay defined as the time it takes for the legs to complete a response to a signal from the force sensor, exists in the 6 DOF system.

This delay (let us call it " π ") is approximately 200 ms; the various contributions to it by hardware subsystems are shown in the block diagram Fig.10.

A simple physical argument shows how $\boldsymbol{\tau}$ might explain the high rebound velocities of table 1. Suppose at time t the probe has reached maximum compression $\boldsymbol{\bar{n}}$ and should turn around. The command to do this is executed $\boldsymbol{i} + \boldsymbol{\tau}$ later; in the meantime, the probe has continued to move and overshot $\boldsymbol{\bar{n}}$. Therefore during a time period of length $\boldsymbol{\tau}$ additional force has been communicated to the software resulting in a gain in energy. If the probe did not detach, $\boldsymbol{\tau}$ sec. after reaching its rest position the overshot would be repeated and one might expect a sequence of oscillations of increasing amplitude.

2.1 A Mathematical Model - Let us now try to make the above argument mathematically precise. Suppose in the interval $[+, \pm + \Delta \pm]$ a change in the force signal ΔF is sampled. When the EAI 8900 computes a change of compression $\Delta \eta$ corresponding to ΔF , it does this by digitally solving the equation \Im and computing the solution at t + Δ t

$$(2.1) \qquad M\ddot{n} = F(t)$$

Now $\eta(t+\Delta t)$ should equal $\eta(t) + \Delta h$. However Δh is not executed until the actuators have finished their response at $t+\tau$. Thus $\eta(t+\tau)$ corresponds to $F(t) - \eta(t)$ equivalently corresponds to $F(t+\tau)$. It follows - (τ includes the sample interval or "frame time" Δt)that the system behave as if it were solving the equation

(2.2) $M\ddot{n} = F(t-\tau)$

Thus we could model the effect of delay by using (2.2) instead of (2.1) in a computer simulation.

In case the probe is an ideal spring F = -kn and (2.2) becomes

(2.3) $M\ddot{\eta} + K \eta(t-\tau) = 0$

(2.2) and (2.3) are examples of retarded differential equations. Because of their importance to control theory applications, such equations have been intensively studied [3], [6], [7] in recent years. The resulting theory has many similarities with the classical theory of differential equations but also important differences. In particular it can be shown that if

TC 2/w.e

then "practically all" solutions oscillate with exponentially increasing amplitude (R. D. Driver, private communication, also cf. Driver [6]).

Also unlike differential equations, a correctly formulated initial value problem (IVP) for (2.2) - (2.3) requires initial conditions on an initial interval rather than just at a point. For example, consider the first τ seconds after contact: The probe continues to compress oblivious of contact since the legs have not yet moved in response to F(+). Thus

$$\eta(t) = N_t$$

(2.4)

 $\dot{\eta}(t) = N_0, 0 \leq 1 \leq T$.

These replace the point conditions of a differential equation IVP.

2.2 A Few Refinements - Essentially equations (2.1) and (2.2) will constitute the mathematical model of the 6 DOF system in our report. Before solving it we discuss some possible improvements (only the most significant of which have been implemented).

2.2.1 Probe Nonlinearity - As is shown in Fig. 2, the probe is not an ideal spring. It is both nonlinear and preloaded. The nonlinearity can be modeled by an equation of the form

$$M\ddot{n} + K(n(t-\tau)) = 0$$

For the graph of k(u) see Fig. . To approximate K(n) one might choose a piecewise linear function, a least squares parabolic fit, or a spline. None of these choices is difficult to implement on the computer. However, since K(n) is nearly constant in the neighborhood of 50-60 lbs/ft, its variation is probably not a source of major error, we will omit this feature of the spring from our model.

Of much greater significance is the existence of a preload (F_p) of approximately 69 lbs. By energy considerations a preloaded spring is equivalent to a compressed longer spring having the same Hook's constant; the fictitious compression is F_P/K . This suggests that we define $\chi(t) := \eta(t) + F_P/K$ and solve

(2.5)

 $\dot{X} + \omega_0^2 \mathcal{X}_{+} = 0$ $\chi(t) = \chi_0 + \mathcal{N}_0 t$ $\dot{\chi}(t) = \mathcal{N}_0 \qquad 0 \leq t \leq \tau$

 $X_{o} := F_{P}/K$ $w_{o}^{2} := K/M$ $X_{D} := X(t-T)$

It will turn out that the preload magnifies the effect of a pure delay.

In this report we will also consider an ideal probe with no preload and K=1200 lbs/ft., since this value has been used in a recent analog simulation [1] of the effect of delays on spring motion in the 6-DOF system.

-2.2.2 Filters - As the block diagram Fig. 10 shows there are several filters in the system. They contribute to the total delay and probably somewhat distort the signal. In [1] it has been shown that the addition of a low pass filter with transfer function </arc gives an additional destablizing effect.

In a digital simulation the presence of such a filter changes (2.3) to the equation

$$M\ddot{n} + K - e(t) + n(t - \tau) = 0$$

(2.6)

e' `

 $e(4) := 1 - e^{-4}$

In general the presence of several filters with constants $c_1 \cdots c_n$ gives the equation

(2.5)
$$M\ddot{n} + K R_1 + R_2 + \cdots + R_n + N(t-\tau)$$

where

The presence of convolutions can be handled by nested integration loops in a digital simulation. Since the parameters of the filters actually in the system are unknown we ignore their effect except in sofar as they contribute to the delay. Our results compared with tests suggest that this is not a crucial omission.

2.2.3 Force Threshold - In addition to the filters a force threshold (F_{4}) is used to reduce noise. This means that F(t) is accepted if and only if its absolute value exceeds F_{4} i.e.,

$$F(t) = \begin{cases} F(t) & \text{if } |F(t)| \gg F_{t} \\ 0 & \text{otherwise} \end{cases}$$

Thus the IVP (2.5) becomes

(2.6)
$$M\ddot{x} + K\lambda(x) = N_0 + F_0/K$$

(2.6) $\chi(4) = N_0 + F_0/K$

$$\dot{x}(t) = N_{\bullet}$$
, $o \leq t \leq \gamma$

....

Fortunately (2.6) is equivalent to a simpler IVP where F_{\pm} plays a role like the preload. The probe must travel F_{\pm}/K before force due to compression is "accepted." The elapsed time is $\bar{t} = F_{\pm}/Kv_{\pm}$: If we regard \bar{t} as time zero, (2.6) becomes

$$\ddot{x} + \omega_0^2 x_{\tau} = 6$$

X(+) = Xo+ No+

(2.7)

$$\dot{X}(t) = N_0$$

 $X_0 = (F_P + F_{tr}) / K$
 $N(t) = X(t) - F_0 / K$

Note that the system should be very sensitive to F_t . If F_t is large and v_0 is small the only force affecting the probe might be that of the bottoming spring. For example if K=60, $F_t=50$ $X_0 > 14.4$ in which is greater than the maximum compression of the probe. In the absence of delay, even threshold of 30 lbs. would produce a severe distortion. When delay is present its destablizing effects are greatly magnified. In the case of a spring with K = 1200 and no preload the effects thought not so severe are still noticable.

The presence of a force threshold also causes some ambiguity in time measuresments during the simulation. Suppose t is the time during rebound when $K_{\Pi} = F_{t}$. Since F(t) is measured directly on the strip chart the simulation period is $E-F_{t}/Kv_{0}$, $t_{F} + F_{t}/Kv_{0}$. The period in which the forces are producing a response in the legs is $[\tau , t_{F} + \tau]$. Currently the exit velocity is measured at $t_{F} + \tau$, time $t_{F} + \tau + F_{t}/Kv_{0}$ from contact. To avoid confusion one needs to keep these various times clearly in mind especially when F_{t} is high and/or K, V is low.

2.2.4 Variable Delay - The IVP (2.6) is a continuous realization of a discutty sampled system. The AD and DA converters (1), (2) (see Fig. 10) sample the signal at discrete intervals Atapproximately 31-35 ms. If all the converters where synchronous little would be lost since we in fact solve the I V P in steps on the computer. However, the converters (3), (4) sample at intervals of 25 ms. This means that periodically (3) samples the same signal twice, producing an additional 25 ms delay. Since they do not start in synchrony (they start where they stopped) the effect varies from run to run. We have put a "noisy" delay into the model. The results suggest that the asynchronization explains some of the noise in the data.

2.2.5 Noise from the table - Since the force sensors are in the table, they may feel some of the inertial forces in the table itself. Again forces due to this source are executed seconds later. This suggest that we consider the neutral to second order equation.

$$(2.8) \qquad \qquad \ddot{X} - \epsilon \underbrace{M}_{r} \ddot{X}_{r} + \omega_{0}^{2} x = 0$$

with the same initial conditions of (2.7) where ϵ is a noise parameter given by a random number generator and M₁ is the mass of the table + probe = 323.5. Preliminary tests (see Appender T) show that this complication is indeed a potent source of noise.

2.3 Solving the Model

2.3.1 Formulas in closed form - Suppose we linearise (2.7) by taking the first two terms of the Taylor expansion, We obtain

(2.9)

 $\ddot{x} + \omega_0^2 x - \omega_0^2 \tau \dot{x} = 0$ $x(0) = X_0$ $\dot{x}(0) = N_0$

If $\omega, \gamma << j$ and $\chi_{o} = 0$ (no preload, no threshold) an approximate solution to (2.9) is

$$X(t) = - \sum_{n=1}^{n} \sum_{j=1}^{n} \sum_{j=1}^{n} w_{i}t$$

Hence

$$(2.10) \qquad \forall \mathbf{r} \geq \mathbf{v}_{\mathbf{r}} \mathbf{e}^{\mathsf{H}_{\mathbf{w}}} \mathbf{r}$$

Neither of these formulas are accurate for much beyond a few seconds. However, (2.9) gives an error of under 3% for $\gamma \leq 2$. If there is a force threshold the formula

(2.11)
$$V_{F} = V_{0} e^{\frac{\pi}{2} W_{0} T} (V_{0} + F_{t} T/_{2M})$$

yields an error < 15% provided F_t < 120 lbs.

An exact closed form solution to (2.7) or 2.8) can be given by an inductive procedure known as the "Method of Steps." Suppose \sim and \times are known on the interval [(k-i)+, k+]. Then setting .

+ = KT+S

we have for $k + \leq t \leq (k+1)$.

(2.12)

$$N(t) = N(kT) - w_0^2 \int_{kT}^{t} X(s) ds$$

$$X(t) = X(kT) + \int_{kT}^{t} V(u) du$$

Since the initial functions are simple the solution may be calculated in closed form, step by step. Thus for $\tau < t \leq 2\tau$

$$N(+) = N_0 + \frac{1}{2} W_0^2 V_0 (+-\tau)^2$$

$$X(4) = X_0 + V_0 T + \frac{1}{3!} W_0^2 V_0 (t-T)^3$$

etc.

Input

2.3.2 Numerical Integration- The Method of Steps gives a solution in the form of an infinite series. Since the result tells us nothing about the properties of the solution, nothing is lost if we numercially integrate (2.7). The following algorithm which uses the corrected trapezoid rule (IISI P.74) has $O(H^4)$ discretization error and has been implemented on the Pict Sigma 5 computer and the writers TISQ.

K, Vo, Xo, Ft, Fp, M, K, T, N

$$X_0 \leftarrow X_0 + (F_{+} + F_{0})/K$$

For i= 1 to N+1

$$X_{i} = X_{i-1} + V_{0} H$$
$$V_{i} = V_{0}$$

Continue

$$V_{i+1} = V_i - W_a^2 (H_2(X_{i+1} - H_i + X_{i-H_i}) + H_{12}^2 (V_{i-H_i} - V_{i+1-H_i})$$

Write Vir,

$$X_{i+1} = X_i + \frac{H_2}{V} (V_{i+1} V_{i+1}) - \frac{\omega_0^2 H^2}{12} (X_{i+1-k} - X_{i-k})$$
$$\eta_{i+1} = X_{i+1} - F_{\mu}/K$$

Write Miti

Continue

2.4 Results - Figures (2)-(7) give the results of several runs with different initial conditions and parameters. Note that the results are consistent with tests of the 6 DOF system and also agree almost perfectly with the results of the analog simulation [1].

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III COMPENSATING FOR DELAY

It is hoped that most of the delay can be eliminated by suitable modifications in the hardware. Nervertheless even delays as low as 50 ms can cause noticeable instability. Therefore in this section we develop several algorithms that in some cases can produce near normal performance in spite of delay. Test results for some of the methods described below are given in Section IV.

3.1 Taylor Approximations

Since

(3.1)

$$X_{T} = X - T\dot{X} + \frac{T^{2}\ddot{X}}{2} + \cdots + (-1)^{h} \frac{T^{h}}{h!} X^{(h)}_{(A)} + T < d < 4$$

we can truncate the series and solve for an approximation x_c to x_{-i,e_1}

$$X_{c} \approx X_{+} + T \dot{X} - T^{2} \ddot{X} + \cdots (-1) \frac{1}{2} T^{n-1} X^{(n-1)}$$

with error $\mathcal{O}(\tau^*, x^*)$. We then solve.

$$\ddot{X} + w_0^2 X_c = 0$$

 $X(r) = X_0$

h = x - Fp/K

instead of (2.1).

3.1.1 Runge Kutta - In theory (3.1) can be solved for arbitrarily high n by the Runge Kutta algorithm without knowledge of high derivatives. For example if n = 4 (3.1) can be converted to the system:

$$\begin{pmatrix} X_{1} \\ X_{2} \\ X_{3} \end{pmatrix}^{l} = \begin{pmatrix} O & I & O \\ O & O & I \\ E_{T} & C_{1} & C_{2} \end{pmatrix} \begin{pmatrix} X_{1} \\ X_{2} \\ X_{3} \end{pmatrix}$$
$$\begin{pmatrix} X_{1} \\ X_{2} \\ X_{3} \end{pmatrix} (O) = \begin{pmatrix} X_{0} \\ V_{0} \\ O \end{pmatrix}$$

where

$$E_{T} := backwards chift
C_{1} := - 3!/T^{2}
C_{2} := \frac{3}{T} \left(\frac{3}{T^{2}} w_{0}^{2} + 1 \right)
X_{1} := X^{(1-1)} \quad i = 1, 2, 3$$

Writing (3.2) in the abbreviated form

and applying the Runge Kutta algorithm.

we obtain

$$X_{i+1} = X_i + H\left(\frac{K_{i}i}{6} + \frac{K_{2}i}{3} + \frac{K_{3}i}{3} + \frac{K_{4}i}{6}\right)$$

where

$$k_{1i} = X_i + \frac{H_2}{2} A Y_i$$

$$k_{2i} = X_i + \frac{H_2}{2} A k_{1i}$$

$$k_{3i} = X_i + \frac{H_2}{2} A k_{2i}$$

$$k_{4i} = X_i + \frac{H_2}{2} A k_{3i}$$

The difference between the above version of Runge Kutta and the conventional one for differential equations is that (because of the shift operation) k_i ; requires $q_{i-m} \cdot k_{2i}$, requires k_{i-m} etc. However, the effect this deviation has on the stability and accuracy of the algorithm is unknown and will require additional investigation. Higher order versions of Runge Kutta are also possible. They can be used alone or combined with equivalent Adams predictor corrector methods. If they work quite high order Taylor approximations can be easily generated.

3.1.2 The Algorithm "FIX*T4" - The following direct approach to (3.1) has been implemented and found to work well. We write.

....

$$\ddot{x}_{+} = \ddot{x} - \tau x^{(iv)}$$
$$x_{+}^{iv} = x^{iv}$$

By back substitution the present values of the derivatives can be eliminated. We finally obtain

$$X_{c} = X_{T} + \tau \dot{X}_{T} + \frac{T^{c}}{2} \ddot{X}_{T} + \frac{T^{3}}{6} \ddot{X}_{T} + \frac{T^{9}}{24} \times \frac{W}{7} + O(\tau^{5}, \chi^{(5)})$$

We can truncate the series obtaining second, thind on fourth order

The 4th order version, however, is as easy to implement as the others and more accurate

The algorithm is $O(h^{q}; \times^{n})$

<u>Remark</u> - The corrected trapezoid rule and the two estimates of $\ddot{x}_i - q_i, q_i$ are necessary to maintain accuracy. Warning - In a 6 DOF simulation

$$Q_{i+1} = Q_{i+1}^{c} = F_{i+1} = F(i+1) \Delta t$$

(The frame time $4\pm$ is the step). Also all delayed terms *, *, *, * etc. are the present values of *, * etc., generated in the hybrid (recall that the backward shift introduced into the algorithm simulates the loop circuit time in the 6 DOF system) With this in mind, we write the algorithm in a form suitable for implementation on the 6 DOF. Let $V_{i+1}, *_{i+1}$ be the velocity and position produced by the hybrid from F by integration. Let $V_{i+1}, *_{i+1}$ be the compensated versions. Then the main loop in (3.3) looks like

For n = m + 1 until satisfied:

 V_{int} = Integration of F

 x_{ini} = Integration of v

$$V_{i+1}^{c} = V_{i} - \omega_{0}^{z} \left\{ \frac{H_{2}(X_{i+1} X_{i+1}) + \frac{H_{12}^{2}(V_{i-1} - V_{i+1})}{T_{2}(V_{i-1} - V_{i}) + \frac{T_{3}^{2}}{T_{6}^{2}}(F_{i+1} - F_{i}) \right\}$$

$$X_{i_{11}}^{c} = X_{i} + \frac{H}{2} (V_{i} + V_{i_{11}}^{c}) + \frac{H^{2}}{2} (F_{i} - F_{i+1})$$

 $V_{i} \in V_{i_{11}}^{c}$ $X_{i} \leftarrow X_{i_{11}}^{c}$

(i.e., these values replace V_i, x_i in the next iteration of the loop)

Continue

In the remainder of the paper all algorithms will be written the form (3.3) instead of (3.4) and it will be understood that terms $*_r$ and their digital form $*_{i-m}$ etc. refer to present not past values of variables calculated in the hybrid.

3.2 Polynomial Extrapolation - Porter and Stoneman [12] (also see [13] have proposed the use of Newton-Gregory interpolating polynomials to compute X_c. Specifically they fit a polynomial to X_{n+1}-m, X_{n+1}-(m+4), ... X_{n+1}-(m+4) and extrapolate to compute X_c = X_{n+1}. Thus by the Newton-Gregory backwards interpolation formula

where

$$\nabla X_{n+1} - m = X_{n+1-m} - X_{n+1-(m+1)}$$

$$\nabla^{2} X_{n+1-m} = \nabla X_{n+1-m} - \nabla n_{n+1-(m+1)} \quad \text{efc}$$

No higher than a cubic should be taken. The procedure might begin with a 4th order Taylor method to compute enough x values beyond $[\circ, \uparrow]$ to fit the cubic. For example, the first cubic could be passed through $X_{m+1}, \dots X_{m+4}$, or through $X_{m}, \dots X_{m+3}$ etc. The best choice would be found experimentally. Since interpolating polynomials are extremely sensitive to noise, we suggest substituting a least squares polynomial of low degree based on as many previous points as possible. If we write the polynomial in the form.

$$P_{v}(x) = a_{v} + a_{1} x + \cdots + a_{v} x^{v}$$

and set $\overline{a} = (a_0, a_1, \dots a_n)$ \overline{a} is the solution of the system.

(3.6) Cā = F

where C is the V+lxV+l matrix with general component

$$c_{ij} = \sum_{\substack{p=m}\\p=m}^{p} f_{a}^{i+j-2}$$

where

$$\overline{h} = \begin{pmatrix} \sum_{i=1}^{e} x_{i} \\ \sum_{j=1}^{e} t_{j} x_{j} \\ \sum_{i=1}^{e} t_{i} x_{i} \\ \sum_{j=1}^{e} t_{i} x_{j} \\ e^{-m} \end{pmatrix}, \quad t_{i} := PH$$

and p is a preassigned integer.

Remark - matrix system (3.6) becomes poorly conditioned for large V; most machines will give meaningless results at V = 7 or 8. We recommend double precision arithmetic and V < 4.

Because velocities are available at the same time as the position function, anther choice would be to use this information and require that the polynomial pass through X_{n-m}, V_{n-m} ... etc. For formulas for this type of interpolating polynomial see [5] p62ff. The error is $O(H^{m+1})$ where m is the degree of the polynomial (e.g., 5 if three points are used.)

Remarks

- (1) Although error terms for polynomial extrapolation schemes are similar to Taylor methods, their performance may be superior (particularly for the least squares version) since more information is used and sensitivity to noise is reduced.
- (2) Warning Neither method (or any other known to the author) will make the system stable for more than a short time. The high derivatives present in the error terms ruin the approximation as

3.3 Transfer Function Approximation - The Laplace transform of \times_{τ} is $e^{\tau_4} \downarrow_{x}$. Therefore, it is tempting to find an approximation P(s) to e^{τ_5} and define

$$X_e R X = L^{-1}(P(4)e^{-Ts} IX)$$

= $J^{-1}(P(4)) = X_T$

Thus we solve.

(3.7)
$$\ddot{X} = -\omega_0^2 f'(P(q)) \wedge X_{\tau}$$

The advantage of this method is that if P(s) is properly chosen, (3.7) can be solved on an analog computer much more easily than it can be on a digital machine. This technique provided P(s) is "close" to e^{75} can be shown to be as good as polynomial extrapolation (see [13] p32ff).

Some Approximations for era

3.3.1 <u>Chebyshev polynomials</u> lead to approximations which are much more accurate than Taylor polynomial approximations of the same degree.

Thus

(3.8)	e ^{T1} ?	. 994792 + . 997396 T4 + . 541667 (T4) ² + . 177083(T4) ³
(3.9)	(. 99 4571 7 . 9 67368 TS + . 5 42991 (F1) ² 7 · 177347 (F1) ⁸
(3.10)	ŝ	. 994571 + . 99736874 + . 542991(72) + . 177 347(rs)

See Atkinson [2] Ch 4

The error of each of these approximations is less than $c \times 10^{-3}$ if $|\tau s| < 1$

Tests based on (3.10) have shown that it is generally superior to a fourth order Taylor approximation when there is no threshold and no preload. For a description of the resulting algorithm "FIX*C" see p below.

3.2.2 Pade Approximations - It is possible to approximate accurately by rational functions of equal degree in the numerator and denominator. The functions are called Pade approximations. Inverting them offers the advantage that numerical or analog differentation is avoided. Also Pade approximations are much more accurate than Taylor series approximations of equal degree. For example the 2nd degree Pade

(3.11)
$$\frac{1 + \frac{TS}{2} + (ts)^{2}}{1 - \frac{TS}{2} + (ts)^{2}}$$

coincides with the first five terms of the Taylor series of the sixth term is $(T^{5})^{5}/144$ which is close to $(T^{5})^{5}/120$. Hence this approximation is $O_{(T^{5})}^{5}$.

The third degree

(3.12)
$$\frac{1 + r_{s}}{2} + \frac{(r_{s})^{2}}{10} + \frac{(r_{s})^{3}}{120} + \frac{(r_{s})^{3}}{120$$

is even better, it concides with the first seven terms of the Taylors series. Hence, if our previous discussion is correct these approximations should be similar to cubic and six degree polynomial extrapolation. Pade approximation can be further improved using Chebychev polynomials (See Ralson [17] p 296) Thus

$$(3.13) \quad e^{\tau t} = \frac{2353 + 1176 \tau 5 + 142(\tau 5)^2}{2153 - 1176 \tau 5 + 142(\tau 5)^2}$$

The error of the approximation for ITSICI is 0-10-4

Inversion of Padé approximations and appeal to (3.7) result in complicated integro-differential equations. For example the second degree Pade yields

Unfortunately all Pade approximations for have poles in the half plane. We conjecture that because of the short contact time this source of instability will not be of major imporance. <u>3.4 Richardson Extrapolation</u> - We finally mention a technique which can "boost" the degree of approximation when used in conjunction with the previous techniques; that is if the error term of a given method is $O(H^*)$ we will obtain by modifying the method an error term $O(H^{**})$

Suppose

(3.14)
$$X_{c}(H) = X + C(H^{4}) + O(H^{(4+7)})$$

Having the step gives

$$(3.15) \qquad X_{c}\left(\frac{H}{2}\right) = X + \frac{c}{2}\left(\frac{H}{2}\right) \rightarrow O\left(\frac{H^{H+1}}{2}\right)$$

(The notation " X_c " indicates the dependence of the compensated position on the step h). Multiplying (3.14) by 2⁴ and subtracting 3.13 gives

$$2^{*} x_{c} (\frac{H_{2}}{2}) - x_{c} (H) \geq (2^{*} - 1) \times + O(H^{**})$$

Therefore

.

$$x = \frac{2^{*} x_{c}(\frac{1}{2}) - x_{c}(H)}{2^{*} - 1} + O(H^{**})$$

IV GRAPHS AND TABLES

KEY TO GRAPHS

In figures 2 and 3 the vertical scale is 1 cm = .01 ft. while in figures 4 and 5 it is 1 cm = .001 ft. In fig. 7 the vertical scale is 1 cm = .01 ft/sec.

In all position vs time or velocity vs time graphs the horizontal scale is 1 cm = .1 sec.

"1" signifies theoretical performance with the given parameters.

"2" signifies uncompensated performance.

"3" is FIX*C correction.

The dashed line in Figs. 6,7 is FIX*T4 correction.

"4" in Fig. 2 is performance with a thirty pound force threshold.

Fig. 2 shows the distorting effect of the force threshold (at two levels) with attempted correction.

Fig. 3 represents theoretical, uncompensated and compensated performance of an Apollo type probe with typical parameters.

Fig. 4 contrasts the performance and compensation for a preloaded vs non-preloaded probe.

Fig. 5 is similar to Fig. 3 but the probe has no preload.

6 DOF CONTACT DYNAMICS TEST RESULT (8/27-28/79)

Test involved vertical motion only and a .75" plywood target simulating the docking collar.

Ι.	V. .05 .05 .05 .05	V_{4} .1927 .1907 .1971 .1730 $V_{5} = .1884$	8/29/79 F _t = 30 lbs.
II.	V ₆ .10 .10 .10 .10	V_{f} .1965 .2122 .2245 .2203 $\overline{V}_{f} = .2134$	8/17/79 F _t = 30 lbs
III.	Ve .1 .2 .2	V4 .2216 .3826 .4101	8/28/79 F _t = 35

TABLE 1

THE EFFECT OF FORCE THRESHOLD ON REBOUND PARAMETERS

١

V =.1 ft/sec. K=1200 lbs/ft M=310.56 Slugs **r** =.15 sec

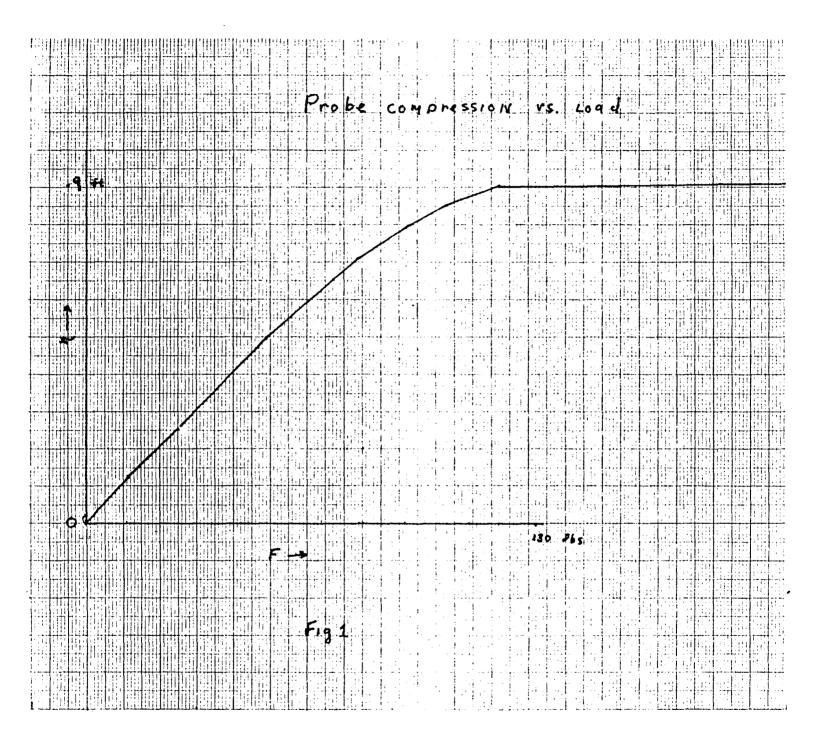
F _t	h - model	V ₄ - formula	V ₄ - Analog Simulation	V _f - Model
0	.0637	.159	.157	.157
20	.0668	.1666	-	.162
30	.0701	.1704	.165	.170
40	.0746	.1742	-	.171
60	.855	.1818	.182	.182
70	.978	.1850	-	.188
80	.984	.1896	-	.1968
90	.102	.1934		.1974
100	.112	.1973	·	.198
120	.127	.2048	.22	.223
240	.22	.2510	.312	.332

PROBE DEFLECTION PROFILE

η (ft)	F(lbs)
0	69
.125	74.8
.25	81
.417	89.3
.5	93.75
.625	101
.708	105.2
.792	112.5
.854	118
.88	121
.9	125
.911	130

For $\eta_{7}.912$ there is a bottoming spring with K = 480,000 lbs/ft

TABLE 3



COMPARATIVE CORRECTION BY 4TH ORDER TAYLOR METHOD (FIX*T4) AND CHEBYCHEV APPROXIMATION (FIX*C)

I F_p=O K=1200 LBS/FT M=311 SLUGS

	THEORETICAL	UN	CORRECTED		FI	<u>X*T4</u>		FI	<u>X*C2</u>	
A $7 = .$.05 .02 .1 .05 .2 .10 .5 .25 1 .51	t 5 1.6 1 1.6 2 1.6 5 1.6	V: .0582 .117 .233 .58 1.16	b .0274 .053 .11 .274 .544	- t 1.65 1.65 1.65 1.65 1.65	V; .052 .1 .2 .3 1.00	n .05258 .051 .1 .256 .511	t 1.65 1.65 1.65 1.65 1.65	V; .049 .1 .2 .499 .997	7 .0256 .051 .1024 .255 .512	£ 1.7 1.6 1.65 1.65 1.65
П-В ? =.1		V4 .068 .135 .268 .67 1.33	T .029 .0589 .118 .29 .59	4 1.7 1.7 1.7 1.8 1.8	v; .048 .0943 .2 .48 1.00	n .0257 .052 .2 .257 .52	+ 1.85 1.8 1.7 1.85 1.7	<pre>V₄ .0499 .999 .195 .488 .976</pre>	й 0259 052 1 26 52	t 1.7 1.7 1.75 1.75 1.75
C 7 =.2		<pre>V↓ .0903 .180 .261 .902 1.81</pre>	й .0342 .0684 .137 .342 .684	t 1.9 1.8 1.8 1.8 1.8	V; .0976 .101 .190 .427 .95	1 .0274 .0576 .116 .276 .551	t 1/85 1.8 1.85 1.85 1.85	V ₄ .0471 .098 .188 .471 .98	7 .0272 .0545 .108 .272 .54	t 1.8 1.8 1.85 1.8 1.8

TABLE 4

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t.

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COMPARATIVE CORRECTION BY 4TH ORDER TAYLOR METHOD (FIX*T4) AND CHEBYCHEV APPROXIMATION (FIX*C)

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II F =69 lbs K=60lbs/ft M =311 SLUGS

	THEORE	TICAL	UNC	ORRECTED		FIX	<u>*T4</u>		FIX	<u>*C2</u>	
$A T = .$ $V_{0} \overline{h}$ $.05 .00$ $.1 .02$ $.2 .08$ $.5 .46$	55 23 69 1	t •45 •9 •25 •6	<pre>√; .0725 .125 .233 .536 1.04</pre>	h .008 .027 .096 .491 1.44	t .6 1.05 1.9 3.70 5.05	<pre>V_f .0725 .12 .22 .53 1.06</pre>	т .0083 .0272 .096 .49 1.45	4 .6 1.025 1.85 3.7 5.3	V₄ .0718 .135 .23 .528 1.02	% .0081 2.73 .0967 .488 1.43	.6 1.1 1.9 3.7 5.1
III-26 B ← = .	1		V∉ .128 .170 .269 .582 1.13	Ъ .01 .0295 .106 .514 1.48	t 1.1 1.3 2.1 3.9 3.4	<pre>√+ .0949 .162 .268 .574 1.06</pre>	й .0106 .0225 .105 .5 1.45	t .75 1.1 2.1 3.9 5.3	• 094 • 147 • 24 • 55 1• 04	Т .011 .032 .106 .506 1.45	.75 1.2 2.0 3.8 5.2
C 7 = .	2		V ₄ .129 .184 .288 .61 1.15	n .0155 .0422 .126 .56 1.57	+ 1 1.45 .225 4.0 5.3	V 128 182 285 595 1.09	т 0155 0419 124 54 1.5	t 1 1.45 2.25 4 5.35	<pre>√↓ .129 .173 .28 .574 1.09</pre>	h .015 .0412 .122 .54 1.49	t 1 1.4 2.25 3.9 5.35

 $F_{p} = 69 \text{ LBS } K= 60 \text{ LBS/FT} M = 311 \text{ SLUGS}$ $F_{\downarrow} = 30 \text{ LBS}.$ Ι $V_{f} = .163$ $V_{a} = .05$ II $V_{h} = .1$ V. = .227 III $V_{o} = .2$ $V_{f} = .33$

IV Ft = 35165

V4 = . 343 $V_0 = \cdot 2$

TABLE 6

,

APPROXIMATE FORMULAS FOR A DELAYED SPRING

Assuming $\omega_{\bullet} \tau_{2} << i$ the following approximate formulas may be derived for the IVP:

 $\ddot{n} + w_0^2 n_r = 0$ $n(t) = v_0 t \quad 0 \le t \le r$ $\dot{n}(t) = v_0$

$$h(t) = N_0 / w_0 \exp(w_0^2 \tau t/z) \sin w_0 t$$

$$\overline{n} = V_0 / w_0 \exp(\pi \tau w_0 / 4)$$

$$V_F = V_0 \exp(\pi \tau w_0 / 2)$$

These formulas give relative errors of under 3%. If a force threshold is present then

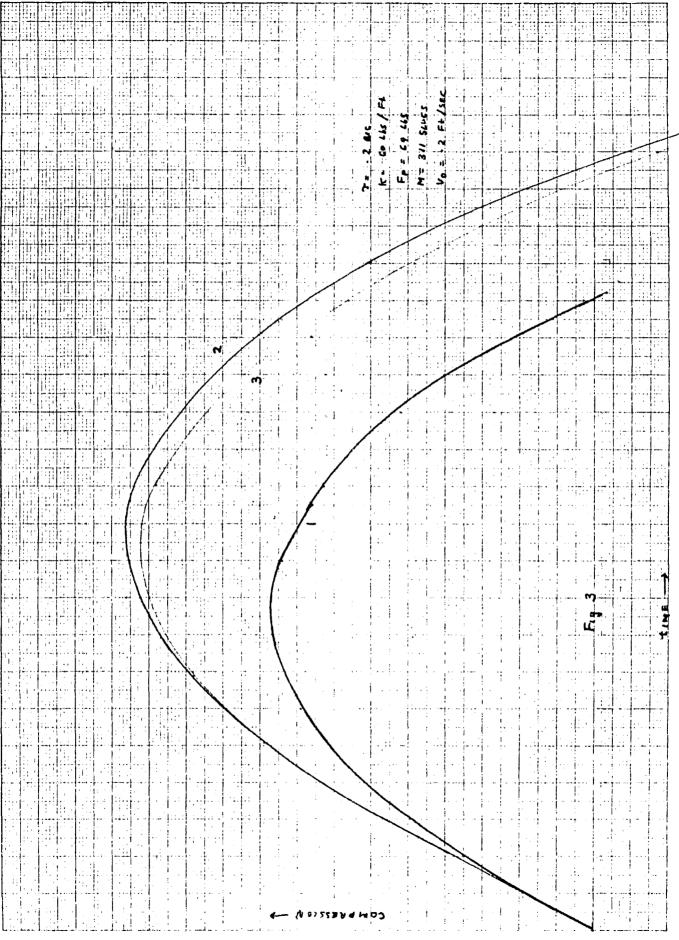
$$V_{F} \approx \exp\left(\frac{\pi}{2} w_{0} r\right) \left(V_{0} + F_{r} r/2M\right)$$

TABLE 7

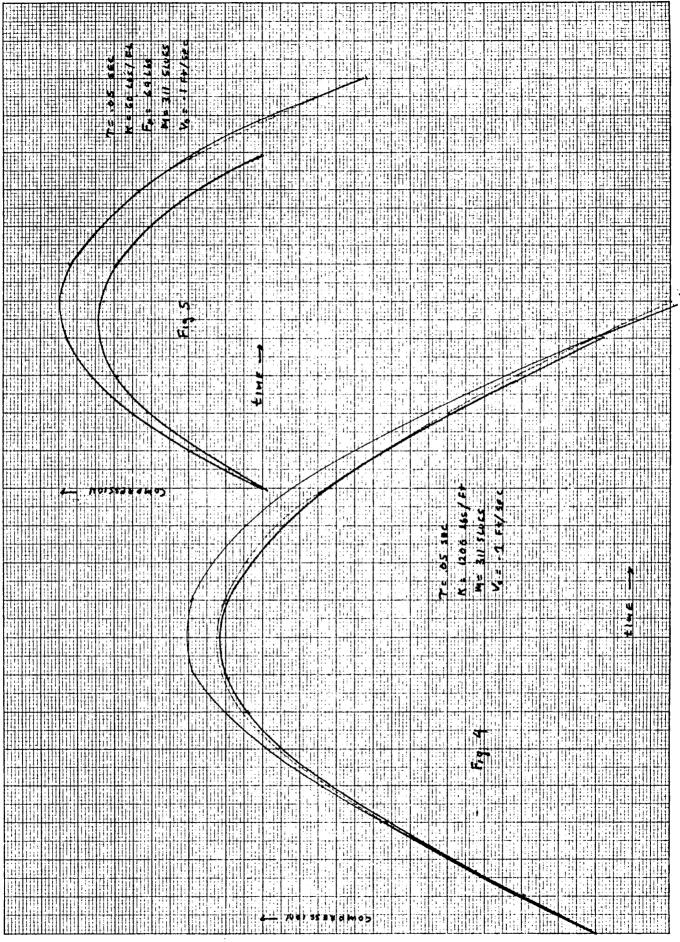
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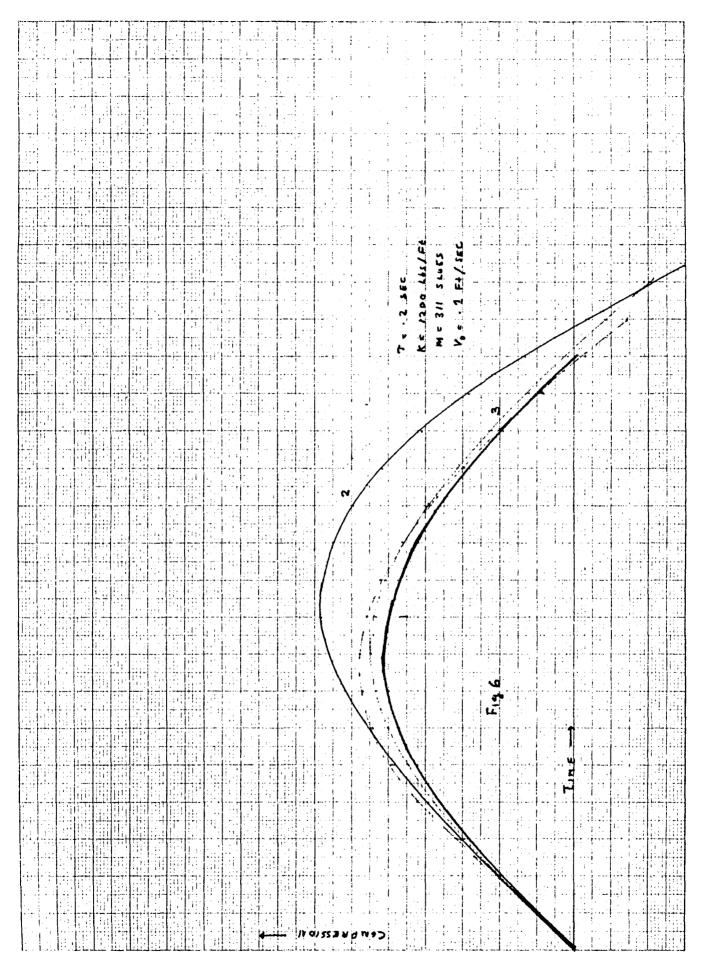
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V CONCLUSIONS AND RECOMMENDATIONS

The second order delay equation model developed in this paper agrees with a recent analog simulation [1] of the effect of delay. It is also reasonably consistent (See Table 6 with the observed behavior (up to noise) of the 6 DOF system when V_{\bullet} = .1 or .2. The most disagreement happens at V_{\bullet} = .05 when our simulation gives a rebound velocity about 13% below the mean of the measurements in Table 1.

The model predicts that for a pure spring with no preload or force threshold V_{4} /V, depends only on r and and not on V_{5} . V_{6} can also be computed quite accurately by a simple formula without numerically integrating the equation in the model. The use of а increase the effect of delay. The results are threshold relatively most severe at high threshold levels and low initial velocities. For thresholds under 100 lbs and delays less than .2 can be predicted with reasonable accuracy by sec V. an approximate formula. Verification of these aspects of the model in hardware tests would strengthen the hypothesis that delay is the predominate factor in the observed errors in the system.

These tests and the isolation of other factors - such as phase shifts due to the filters-should be possible when the noise of the system is reduced. Some of this noise is probably caused by the sensitivity of the force sensors on the table to inertial effects, and by variable delay due to the aysynchronization of the AD/DA converters (see 2.2.4-2.2.5 above). There may also be another source of apparent noise: for the parameters chosen to model the probe and a force threshold the rebound velocity changes rapidly towards the end of the simulation. If the rebound velocity is measured at $t_{\pm} + \gamma$ (cf. 2.2.3) an error of \pm Δ t in the time of measurement can produce a large relative change in V_s . For example when $V_1 = .2 V_s$ is .343; 033 sec later it is 354; 066 sec later it is.38, 066 sec before it is .32. Hence, slight differences in time of final measurement in repeated runs can cause variations close to the order of magnitude of those observed.

Every effort ought to be made to reduce noise and delay. Several means to this end are suggested in [1]. We agree with these recommendations. Since a delay as low as 50 ms still causes considerable errors and this may be close to an achievable lower limit, means of compensating for delay need to be implemented.

In this study both a fourth order Taylor method and a Chebychev polynomial approximation to the transfer function ℓ^{15} gave good results for for an unpreloaded spring when no force threshold is used. The velocity profile correction in this case was excellent; penetration errors were reduced to under 10% (most of the time less than 5%). The Chebychev algorithm "FIX*C" was the most accurate.

We regret however that neither algorithm works in the case of a preloaded spring, when the parameters model the probe tested in Table 1. Indeed its is remarkable that they hardly perturb the uncompensated position time profiles even when $\tau < .05$ sec. Their effects on an unpreloaded spring with a force threshold are more complicated but also unsatisfactory.

If these options are not satisfactory, a deeper mathematical analysis needs to be made on whether compensation of a preloaded system with delay is possible. On this topic the following rough observation seems appropriate. Since the difference between the compensated and the theoretically correct simple harmonic system is $O(\tau^n, X^{(n)})$ where π is the solution of the compensated system an n is the order of the method, the derivative $x^{(n)}$ may grow faster beyond $X= \times_c$ - than τ^n goes to zero. If this phenomenon occurs for all realistic compensation schemes it is doubtful that the 6 DOF system can simulate contact dynamics unless the delay is reduced far below 50 ms.

REFERENCES

- 1. Allen, D. W., W. M. "An Investigation of the Performance Gover, L. G. Thomas, of the Simulation Equipment used in and F. L. Vinz, Contact Dynamics," manuscript.
- 2. Atkinson, K.E., <u>An Introduction to Numerical Analysis</u>, John Wiley, New York, 1978
 - Differential Difference Equations, Academic Press, 1963

Elementary Numerical Analysis,

Equations, 21 (1976), 148-165.

Introduction to the Theory and

Press, New York, 1973.

3rd Ed., McGraw-Hill, New York, 1980.

Small Delays," Journal of Differential

Application of Differential Equations

Equations, Springer Verlag, New York,

with Deviating Arguments., Academic

Theory of Functional Differential

Methods for Solving Engineering Problems Using Analog Computers, McGraw Hill, New York, 1964

Retarded Actions," J. Appl. Mech. 9

Nonlinear Oscillations, Chap. 21, Van Nostrand, Princeton, 1962.

of Pulse-monitored Servo Systems," Proco IEE, 97 pt II (1950)m 597.

"Self-excited Oscillations in

"A New Approach to the Design

Dynamical Systems Possessing

"Linear Differential Systems with

- Callender, A., "Time-lag in a Control System," D. R. Hartree, Phil. Trans. Roy Soc. London, A. Porter, A235 (1936), 415-444.
- 5. Conte, S. D. and C. de Boor

Bellman, R. and

K. L. Cooke,

3.

4.

- 6. Driver R. D.,
- 7. El'sgol'ts, L.E. and S. B. Norkin,
- 8. Hale, J.,
- 9. Levine, L.,
- 10. Minorsky, N.,
- 11.
- 12. Porter, A. and F. Stoneman,
- 13. Ragazzini, J.R. <u>Sampled-Data Control Systems</u>, and G. F. Franklin <u>McGraw-Hill</u>, New York, 1958.

1977.

(1942), GS-71.

14. Ralston, A.,

A First Course in Numerical Analysis, McGraw-Hill, New York, 1965.

15. Wolfe, M.A.,

A First Course in Numerical Analysis, Van Nostrand, London, 1972.

APPENDIX I

INERTIAL NOISE AND THE FORCE SENSOR

The following is an attempted simulation of the effects of noise caused by the inertial forces of the table affecting the sensor. We integrate equation (2.8) (see § 2.2.5), thus representing a "shaking" of the sensor at each step in the movement of the table. Here the "noise parameter" ϵ is a uniformly distributed random variable taking values between 0 and either .5 or 1. The mass of the table is about 12 slugs so that the mass of the table and probe is 323 slugs. Note that the rebound velocities with the model exceed those observed. Parameters: F = 69 lbs K= 60 lbs/ft $\gamma = .2$ sec.

I. € =.5

A. $V_6 = .05$

C. $V_{0} = .2$

V4			Vr
.1778		•	.395
.204			.4012
.2001			.383
.1967			.398

B V₀ =.1

V, .2773 .2525 .269

.2649

II. € =1

A. V. =.05

C. $V_{0} = .2$

Vt	Vŧ
.2248	. 4434
.2186	.46
.2356	.4537
.2433	.4376

B. $V_{p} = .1$

V	1	
•	3100	
•	3429	
	3828	
	3298	

Ⅲ-38

APPENDIX II

LIST OF PROGRAMS

All programs are in BASIC completely interactive and implemented on MFC's Xerox Sigma 5 system.

Name	Purpose
LAG*R	Simulates the basic delay model.
PSP*I	Theoretical spring with arbitrary initial conditions
FIX*C	Chebychev polynomial correction
FIX*T4	Fourth order Taylor method correction
LAG*N	Simulates the delay model with inertial noise

THE ALGORITHM "FIX*C"

Let $C_i := .994526$ $C_2 := .995682$ $C_2 := .543081$ $C_4 :+ .179519$

Set

 $x_{1} = c_{1} x_{1} + c_{2} x_{2} + c_{3} x_{2} + c_{4} x_{7}^{2}$

The following algorithm solves the IVP

 $X'' + w_0^2 X_c = 0$ $X(t) = X_0, \dot{X}(t) = V_0, \quad 0 \in t \leq T$

For i= 1 to m

$$V_{i} := V_{0}$$

$$X_{i} := i + V_{0} + X_{0}$$

Continue

For i = m+1 until satisfied

$$V_{i+1} = V_i - W_0^{2} \begin{cases} \frac{4}{2} C_i \left((X_{i-m} + X_{i+1-m}) + \frac{4}{2} (V_{i-m} - V_{i+1-m}) + C_2 (X_{i+1-m} - X_{i-m}) + C_3 (V_{i+1-m} - V_{i-m}) \end{cases}$$

+ $C_4 \left((Q_{i+1-m} - Q_{i-m}) + C_3 (V_{i+1-m} - V_{i-m}) + C_4 (Q_{i+1-m} - Q_{i-m}) \right)$

III−40

$$Q_{i+1} = - w_0^2 X_{i+1-m}$$

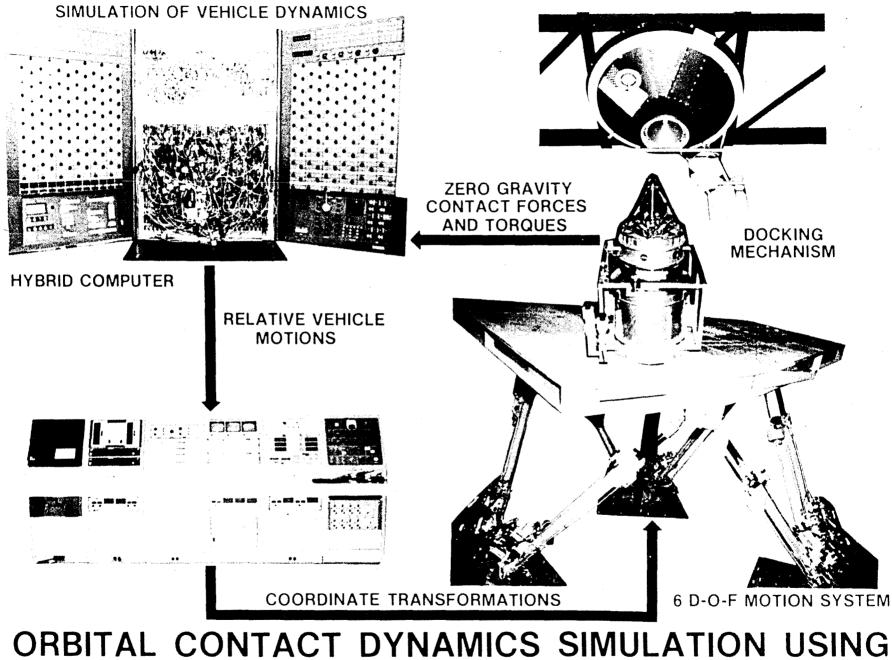
 $Q_{i+1}^c = - w_0^2 (X_{i+1-m} + V_{i+1-m})$

$$X_{i+1} = X_i + H/2 (V_i + V_{i+1}) + H^3/_{12} Q_i^c - (Q_{i+1}^c)$$

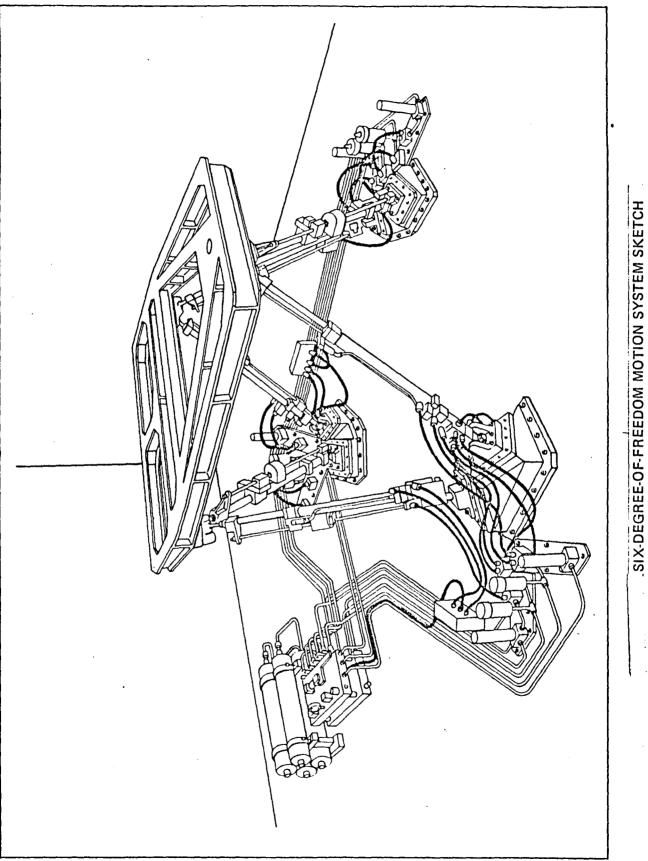
Continue

APPENDIX IV

ILLUSTRATIONS OF THE 6 DOF MOTION SYSTEM



SIX DEGREE-OF-FREEDOM MOTION SYSTEM



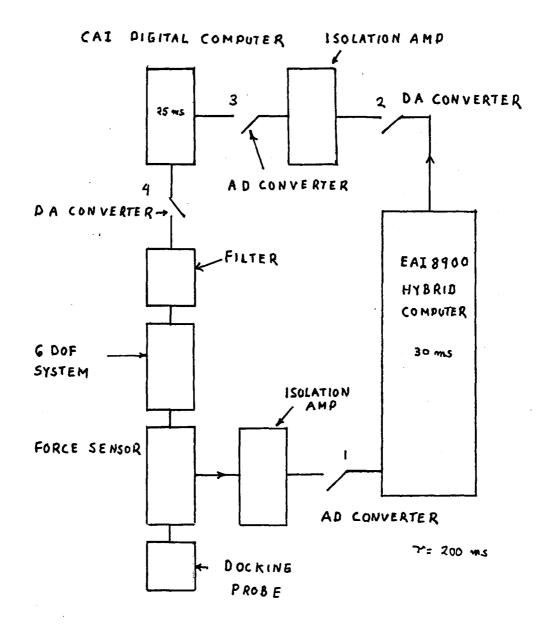


Fig 10

	POSITION	RATE	ACCELE	RATION
			NO LOAD	23,000 LB. LOAD
PITCH	+30°, -20°	± 15°/SEC.	$\pm 6.5 \text{ RAD/SEC}^2$	± 2 RAD/SEC ²
ROLL	± 22°	± 15°/SEC.	\pm 7.0 RAD/SEC ²	$+ 1.6 \\ - 2.0$ RAD/SEC ²
YAW	± 32°	± 15°/SEC	± 6.0 RAD/SEC ²	± 2.0 RAD/SEC ²
VERTICAL	39 IN. UP, 30 IN. DOWN	± 24 IN./SEC.	± 1.6g	± 1.0g
LATERAL	±48 INCHES	± 24 IN./SEC.	± 2.4 g	± 0.6 g
LONGITUDINAL	± 48 INCHES	± 24 IN. /SEC.	± 2.0 g	± 0.6 g

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PERFORMANCE OF SIX-DEGREE-OF-FREEDOM MOTION SYSTEM . .

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1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

PILOT SIGNALS FOR LARGE ACTIVE RETRODIRECTIVE ARRAYS

Prepared By:	C. H. Chan, Ph.D.
Academic Rank:	Professor of Physics
Unviersity and Department:	University of Alabama at Huntsville Department of Physics

NASA/MSFC (Laboratory) (Division) (Branch)

.

MSFC Counterpart:

Date:

Contract No.:

Preliminary Design Office Subsystems Design Division Electrical System Branch

Woolsey Finnell III

August 1, 1980

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PILOT SIGNALS FOR LARGE ACTIVE RETRODIRECTIVE ARRAYS

BY

C. H. Chan, Ph.D.

Professor of Physics

University of Alabama at Huntsville

Huntsville, Alabama

ABSTRACT

It has been suggested that for large active retrodirective arrays, as in the solar power system, a two-tone uplink pilot signal with frequencies symmetrically situated around the downlink frequency be used in order to reduce ionospheric biases and to lower the cost since a two-tone receiver is economically much cheaper than a single-tone phase-locked receiver. Unfortunately such a system now faces the following well-known difficulties: (i) the π -ambiguity, (ii) a large phase difference between the downlink and uplink signals.

We show in this report how the π -ambiguity can be easily removed by using a two-tone uplink signal with both frequencies situated at one side of the downlink frequency, and the phase difference can be greatly reduced with a three-tone or a four-tone uplink pilot signal.

ACKNOWLEDGMENT

I wish to thank the NASA/ASEE Summer Faculty Fellowship Program and its directors for providing the opportunity to participate in the Summer (1980) research project. Appreciation is extended to Mr. W. Finnell III for his support of my work this summer. Special thanks are to Mr. R. H. Durrett and Mr. R. A. Inman of EC-33 for their many enlightening discussions.

I. Introduction

It has been suggested¹ that active retrodirective arrays^{2,3} would be particularly suitable as solar power satellite's antennas⁴ because they are inherently failsafe. The active retrodirective array works on the so-called phase-conjugation principle. It electronically points a microwave beam back at the apparent source of an incident pilot signal. Retrodirectivity is achieved by retransmitting from each element of the array a signal whose phase is the conjugation of that received by the element. In the satellite power system, the pilot source on ground may be situated at the center of a large rectenna and the retrodirective array is the space antenna in geosynchronous orbit.

Retrodirectivity can be most easily achieved if the uplink signal and the downlink beam have the same frequency. But due to input-output isolation problems, the uplink frequency is either upshifted or downshifted from the downlink frequency, a phase-locked receiver is used to achieve phase conjugation. When the uplink and downlink frequencies are different and because the ionosphere and transmission lines are dispersive, the conjugated uplink phase is no longer exactly equal to the downlink phase and the beam coherence at the rectenna can be lost. The downlink beam then points to wrong directions and this is known as beam squint.

A two-tone pilot uplink signal with frequencies symmetrically situated around the downlink frequency has subsequently been suggested. The two-tone uplink signal circumvents the beam squint problem. It reduces ionospheric biases and biases due to the dispersion of the transmission line. It also lowers the cost since a two-tone receiver is much

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cheaper than a single-tone phase-locked receiver. But, it introduces a new problem, known as the π -ambiguity, and the old problem that the conjugated uplink phase could still differ from the downlink phase by an intolerable amount remains. Raytheon, Boeing⁶, and Rockwell⁷ have all made further suggestions to remedy the problems. Their solutions not only are complicated and require much extra hardware in the already very complicated phase conjugation circuitry, but certain problems still remain.

In this report new designs of two-tone and multi-tone uplink signals with frequencies situated at one side of the downlink frequency are suggested. This method removes the above mentioned difficulties and does not require extra components in the phase conjugation circuit. We shall review in our next section the basic principle of phase conjugation and where the problems are in a symmetrically situated two-tone uplink signal. In section III we show how these problems can be circumvented with our new designs.

II. Difficulties With Symmetrically Situated Two-Tone Uplink Signal

A retrodirective array electronically transmits a microwave beam back to the apparent source of a coherent pilot signal. The beam radiated by self-phasing antenna may or may not be coherent across the aperture but it is coherent when it arrives back at the source. Retrodirectivity is the result of phase conjugation of the pilot signal received by each element of the array. Let the phase of the pilot signal of angular frequency ω received by the kth element of the array at time t be $\phi_k(t) = \omega(t - r_k/c)$ where r_k is the distance from the kth element to the source. We define the conjugate of ϕ_k to be

$$\phi_k^*$$
 (t) = $\omega(t + r_k/c) + \phi_0$

where ϕ_0 is an arbitrary phase offset but is constant over the entire array. The phase of the beam received from the kth element by a receiver located at the pilot source (r = 0) is, at time t,

$$\phi_{k}(t,0) = \omega(t + \frac{r_{k}}{c} - \frac{r_{k}}{c}) + \phi_{0} = \omega t + \phi_{0}$$

Thus the contributions to the field at r = 0 from various elements of the array are all in phase at that point.

In the above simple example the uplink frequency was chosen to be the same as the downlink frequency. This restriction is neither necessary nor desirable and is usually avoided because of input-output isolation problems. When these two frequencies are different, a phase-locked receiver is used. Retrodirectivity can still be achieved - provided that the propagating medium is nondispersive.

Due to the fact that single-tone phase-locked receivers are expensive and the ionosphere and transmission lines are dispersive, a two-tone uplink signal with frequencies symmetrically situated around the down-link

frequency was suggested and the average of the phases of the uplink tones be taken as a good estimate of the phase at the downlink frequency. Such a system lowers the cost and removes partially the difference between the uplink and downlink phases but it also introduces a new problem known as the π -ambiguity. We shall review here where these problems are. We shall use the ionosphere as an example to study the effect on the phase conjugation due to the dispersive property of the propagating medium.

The dispersion relation for ionosphere with $\omega > \omega_{p}$ is

$$k = \frac{\omega}{c} \left(1 - \frac{\omega \rho^2}{2\omega^2}\right)$$
 (1)

where c is the speed of light in vacuum, k the wave number, ω the angular frequency and ω_p is the plasma frequency. The plasma frequency ω_p is related to the electron density as

$$\omega_{\rm p}^{2} = \frac{{\rm Ne}^{2}}{\varepsilon_{\rm o}m}$$
(2)

where ε_0 is the vacuum permittivity, e the charge of the electron, m the mass of the electron and N is the volume electron density. From Eqs. (1) and (2)

$$k = \frac{2\pi f}{c} - \frac{Ne^2}{4\pi \varepsilon_0 fcm}$$

where $f = \omega/2\pi$ is the frequency. For a beam to transverse a path length L, the total phase change is

$$\phi = \int k \, ds$$

$$= \frac{2\pi f L}{c} - \frac{e^2}{4\pi \varepsilon_0} \frac{1}{f cm} \int N \, ds$$

$$= A f - \frac{B}{f} \qquad (3)$$

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where $A = \frac{2\pi L}{c}$, $B = \frac{e^2}{4\pi\epsilon_0} \frac{1}{c m} \int N ds$, and $\int N ds$ is the columnal electron density.

With uplink frequencies situated around the downlink frequencies, we have $f_1 = f_D - \Delta f$ and $f_2 = f_D + \Delta f$ where $f_D = 2.45$ GHz is the downlink frequency. Using the notation of Eq. (3)

$$\phi(f_1) = Af_1 - \frac{B}{f_1} = \phi_1$$

$$\phi(f_2) = Af_2 - \frac{B}{f_2} = \phi_2$$

and

then the average of the two phases is

$$\overline{\phi} = \frac{1}{2} (\phi_1 + \phi_2)$$

$$= A f_D - \frac{B}{2} \left(\frac{1}{f_D - \Delta f} + \frac{1}{f_D + \Delta f} \right)$$

$$= A f_D - \frac{B}{f_D} (1 + \varepsilon^2 + \varepsilon^4 \dots)$$
(4)

where $\varepsilon = \Delta f/f_D$ is a small number. We may compare this phase $\overline{\phi}$ with the downlink phase

$$\phi_{\rm D} = {\rm Af}_{\rm D} - {\rm f}_{\rm D}^{\rm B}$$

e.g. their difference is of the order

$$\Delta \phi = | \phi_{\rm D} - \overline{\phi} | \approx \frac{B\varepsilon^2}{f_{\rm D}}$$

To estimate this difference, we shall assume

N ds =
$$5 \times 10^{17} \text{ electrons/m}^2$$

a rather large value for $\int N \, ds$ but taking into account for the possible worst condition. With $\Delta f = 50$ MHz, $\Delta \phi$ is estimated to be about 4°, which is not too small.

When we took the average of phases ϕ_1 and ϕ_2 in Eq. (4), there could introduce an ambiguity known as the π -ambiguity. Let

$$\phi_1 + \phi_2 = K(2\pi) + \Delta$$

where $0 \leq \Delta < 2\pi$ and K is a positive, zero or negative integer. Hence

 $\overline{\phi}$ = K π + $\frac{\Delta}{2}$

In performing the phase average, the KT term could get lost. For K even, no damage is done. For K odd a π error is introduced and one would conjugate the wrong phase.

In order to remove these two difficulties, especially the π -ambiguity, Raytheon⁵, Boeing⁶, and Rockwell⁷ have all made suggestions. Their solutions are very complicated and usually require a lot of hardware in the receiving and phase conjugation circuitries with much added cost. Furthermore, their soultions do not solve the problem completely. We shall show in our next section several simple solutions which circumvent the above mentioned difficulties and do not add extra costs.

III. New Designs of Pilot Beam System

In this section, several uplink designs are proposed. The first simple design avoids the π -ambiguity and the rest are improved versions of the first one. They are all free from the π -ambiguity but reduce the phase difference $\Delta \phi$ to various orders.

(i) This is also a two-tone uplink, but the two frequencies are both on one side of the downlink frequency with

$$f_{1} = f_{D} - \Delta f$$

and $f_{2} = f_{D} - 2\Delta f$ (5)

where f_D is the downlink frequency. We now let $\overline{\phi} = 2\phi (f_1) - \phi (f_2)$ to be the estimation of the downlink phase. With the notation of Eq. (3)

$$\overline{\phi} = Af_D - B \left(\frac{2}{f_D - \Delta f} - \frac{1}{f_D - 2\Delta f} \right)$$
$$= Af_D - \frac{B}{f_D} \left(1 - 2\varepsilon^2 + \dots \right)$$

where again $\varepsilon = \Delta f/f_D$ is a small number. In this simple design the π -ambiguity is removed since no division of the phase is used anywhere. The combined phase $\overline{\phi}$ is also a good estimate of the downlink phase ϕ_D . Their difference $\Delta \phi = |\phi_D - \overline{\phi}|$ is

$$\Delta \phi = \frac{2 B}{f_{\rm D}} \varepsilon^2 \tag{6}$$

which is of the same order as the one with two uplink frequencies situated symmetrically around the downlink frequency and is about 8° when the same values for $\int N \, ds$ and Δf are used as in the last

section. We also note that in this simple design no extra components are required in the receiving circuitry.

(ii) This is an improved version of the first design. It requires three uplink tones but it also greatly improves the accuracy in estimating the downlink phase $\phi_{\rm p}$. The three frequencies are

$$f_1 = f_D - \Delta f$$

$$f_2 = f_D - 2\Delta f$$
and
$$f_3 = f_D - 3\Delta f$$

We now let

$$\overline{\phi} = 3\phi (f_1) - 3\phi (f_2) + \phi (f_3)$$
(7)

to be the estimation of the downlink phase. With the notation of Eq. (3)

$$\overline{\phi} = Af_D - B \quad \left(\frac{3}{f_D - \Delta f} - \frac{3}{f_D - 2\Delta f} + \frac{1}{f_D - 3\Delta f}\right)$$
$$= Af_D - \frac{B}{f_D} \quad \left(1 + 6\varepsilon^3 + \ldots\right)$$

where again $\varepsilon = \Delta f / f_D$. In this design, there is no π -ambiguity as before and the difference between ϕ_D and $\overline{\phi}$ is reduced by an extra factor ε , e.g. the difference $\Delta \phi$ is now

$$\Delta \phi = \frac{6B}{f_{\rm D}} \varepsilon^3 \tag{8}$$

With $\Delta f = 50$ MHz and $\int N \, ds = 5 \times 10^{17} \, \text{electrons/m}^2$, this difference $\Delta \phi$ is only about 0.5°, which is samll.

(iii) This version can be used in the event we would like to have an even smaller $\Delta \phi$ or we would like to use a larger Δf , which

would otherwise result in a too large $\Delta \phi$ even with a three-tone uplink design. These new specifications can be achieved at the expense of adding a fourth tone. The four frequencies are

$$f_{1} = f_{D} - \Delta f$$

$$f_{2} = f_{D} - 2\Delta f$$

$$f_{3} = f_{D} - 3\Delta f$$

$$f_{4} = f_{D} - 4\Delta f$$

and we now let

$$\overline{\phi} = 4 \phi(f_1) - 6 \phi(f_2) + 4 \phi(f_3) - \phi(f_4)$$
(9)
= $\Lambda f_D - B \left(\frac{4}{f_D - \Lambda f} - \frac{6}{f_D - 2\Lambda f} + \frac{4}{f_D - 3\Lambda f} - \frac{1}{f_D - 4\Lambda f} \right)$
= $\Lambda f_D - \frac{B}{f_D} \left(1 - 24 \epsilon^4 + \dots \right)$

In this design, the phase difference is reduced further by a factor $\epsilon,$ e.g. $\Delta \varphi$ is only

$$\Delta \phi = \frac{24B}{f_D} \epsilon^4$$
 (10)

With $\Delta f = 50$ MHz and $\int N \, ds = 5 \times 10^{17}$ electrons/m², this difference is only about 0.04°. Even with $\Delta f = 100$ MHz, this difference is only 0.66° which is still very small.

In all these three designs the uplink frequencies are all on the lower side of the downlink frequency and they are all equally spaced. These are not the only choices. One can use all frequencies on the upper side of the downlink frequency and they need not be equally spaced

either. As a simple example one may very well have a two-tone uplink with

$$f' = f_{D} + 2\Delta f$$

$$f'' = f_{D} + 3\Delta f$$
(11)

and $\overline{\phi} = 3 \phi(f') - 2 \phi(f'')$. It will work just as well.

f

So far we have used the ionosphere as an example to show how the dispersion of the transmission medium could introduce sizable biases between the uplink and downlink phases. For a two-tone uplink signal, this phase difference is about 8°. We suggested one way to suppress it is to use a three-tone uplink. However, it turns out that if the biases were purely due to the ionosphere, it is really not necessary to use a three-tone uplink. This is because even though the phase difference between $\overline{\varphi}$ and φ_n in our two-tone uplink design is large, its variation from subarray to subarray will be decimal if not infinitesimal. To corroborate more on this statement, from Eq. (6) and the definition of B, the difference between the uplink and downlink phases from the pilot to the kth subarray for a two-tone uplink is

$$\Delta \phi_{\mathbf{k}} = \frac{2(\Delta \mathbf{f})^2}{\mathbf{f}_{\mathbf{D}}^3} \frac{\varepsilon^2}{4\pi\varepsilon_{\mathbf{O}}} \frac{1}{\mathrm{cm}} \int \mathbf{N} \, \mathrm{ds}_{\mathbf{k}}$$
(12)

where $\left| N ds_k \right|$ is the columnal electron density along the path from the pilot to the kth subarray. Since the horizontal dimension of the region of the ionosphere that would be transversed by the pilot signal to any subarray is very small, typically of the order less than 100 meters, the transverse variation of the total electron along a path of several hundred kilometers within a tube of diameter less than 100 meters would be very small. Hence even though $\Delta \phi_{\mathbf{k}}$ is estimated to be about 8°, its variation is at least several orders less. Then with

the use of the phase from one of the subarrays as the reference phase, this difference $\Delta \phi_k$ can be subtracted out and the remainder be treated as a constant phase offset which has no effect on the retrodirective beam. In this sense, though the ionosphere is dispersive, it does not cause any problem, and a two-tone uplink is sufficient.

On the other hand, with the use of "central phasing" we cannot avoid the extra path length of transmission lines for some subarrays. Since these lines are not dispersionless, they will introduce a sizable phase difference. This phase difference can also be estimated. If we assume that this transmission line is a wave guide, its dispersion relation is well known

$$k = \frac{2\pi}{c} \sqrt{f^2 - f_{\lambda}^2}$$
(13)

where f_{λ} is the cut-off frequency. Hence the phase for any link is $\phi(f) = \frac{2\pi \ell}{c} \sqrt{f^2 - f_{\lambda}^2} \qquad (14)$

where ℓ is length of the transmission line. Just as before the phase differences for various uplinks can be calculated. For a two-tone, uplink with $\overline{\phi} = 2 \phi(f_1) - \phi(f_2)$, $\ell = 500$ meters and $f_{\lambda} = \frac{1}{2} f_{D}$, where f_1 and f_2 are given as in Eq. (5), $\Delta \phi$ is

$$\Delta \phi = |\phi_{\rm p} - \overline{\phi}| = 256^{\circ}$$

Similarly for a three-tone uplink with $\overline{\phi} = 3 \phi(f_1) - 3 \phi(f_2) + \phi (f_3)$, one obtains $\Delta \phi = 23^{\circ}$

For a four-tone uplink with $\overline{\phi} = 4 \phi(f_1) - 6 \phi(f_2) + 4 \phi(f_3) - \phi(f_4)$, $\Delta \phi$ is further reduced to

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As we see in this example, the phase difference for a two-tone uplink is very large but this number is greatly reduced in a three-tone or four-tone uplink. These multi-tone uplink signals can be used as a useful alternative method to suppress biases due to the dispersion of the transmission line and the medium.

IV. Conclusion

In our last section we illustrated how simple designs can be used to eliminate the π -ambiguity and reduce the ionospheric biases and biases from dispersive transmission lines. We also note that none of our designs require extra components in the receiving circuitry. All one is required to do is to obtain $\overline{\phi}$, which can be achieved rather easily, and simply conjugate it and use it as the phase of the downlink signal leaving the space antenna.

It is also important to remember that we are here to design a pilot beam system as simple as possible with the phase received by the array as close to the downlink phase as possible. We are not asked to and it is not necessary to determine the ionospheric electron density as required in some other designs.

Lastly our designs of pilot beam can be implemented easily in any large retrodirective arrays. Their advantages are (i) avoiding using phase-locked receiver (ii) free from phase ambiguity (iii) greatly reducing biases due to dispersion of the transmission line and medium (iv) very simple to be constructed.

It will be extremely interesting to have such a system built and tested in the very near future.

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References

- R. C. Chernoff, "Large Active Retrodirective Array for Space Application", IEEE Transaction on Antennas and Propagation, Vol. 27, PP 489-496, July, 1979.
- M. I. Skolnik and D. D. King, "Self-Phasing Array Antennas", IEEE Transaction on Antennas and Propagation, Vol. 12, pp 142-149, March, 1964.
- B. A. Sichelstiel, W. M. Waters and T. A. Wild, "Self-focusing Array Research Model", IEEE Transaction on Antennas and Propagation, Vol. 12, pp 150-154, March, 1964.
- P. E. Glaser, "Power From the Sun: Its Future", Science, Vol. 162, pp 857-861, Nov. 22, 1968.
- Raytheon Report No. ER79-4032, "Solar Power Satellite (SPS) Pilot Beam and Communication Link Subsystem Investigation Study - Phase 1", Jan 31, 1979, Contract No. NAS8-33157.
- Boeing Report No. D180-24635-1, "Microwave Power Transmission System", 1978.
- A. K. Nandi and C. Y. Tomita, "Ionospheric Effects in Active Retrodirective Array and Mitigating System Design," NASA, Solar Power Satellite, Workshop on Microwave Power Transmission and Reception, Jan. 15 - 18, 1980.

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Prepared by:

Jack H. Davis, Ph.D.

Academic Rank:

University and Department:

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Associate Professor

The University of Alabama in Huntsville Department of Physics

Materials and Processes Engineering Physics Physical Sciences

Ann F. Whitaker

August 11, 1980

NGT-01-002-099 University of Alabama

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A. Abstract

An experimental project was undertaken to ion plate by electron beam evaporation Al films onto 4340 steel substrates using (and at the same time troubleshooting) the custom built V.T.A. 7375 electron beam ion plating system. A careful recent literature and commercial vendor survey indicates possible means (reported herein) of improving the trouble plagued V.T.A. system.

Acknowledgements

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C, Introduction

Holliday¹² gave the following concise introduction to ion plating.

Since first being reported in the literature in 1963 (ref. 1), ion plating has progressed to the point that today it is used in several commercial processes including the aluminum coating of fasteners used in the manufacture of aircraft (ref. 2). Ion plating has several positive qualities. Probably the most important of which are outstanding film adhesion and deposition on all sides of the substrate (including coverage into cavities) whereas normal vacuum deposition gives the usual line-of-sight coverage. Other qualities obtainable, depending on the material deposited, include exceptional corrosion resistance, high film purity, fine grain structure, very low-coefficient of friction, and improved mechanical properties of metals (ref. 3-5).

Several articles have been published describing the ion-plating process (e.g., ref. 6 and 7). Basically the process consists of two phases. In the initial cleaning phase the object to be plated (substrate) is made the cathode of a dc inert gas discharge or plasma. Some of the inert gas atoms are ionized and accelerated towards the substrate. The bombardment of the substrate by these ions having high kinetic energy produces a clean surface preparatory to the actual plating process. In the plating phase the material to be plated is evaporated from the anode while maintaining the inert gas discharge. Some of the coating material atoms are also ionized and accelerated towards the substrate. These ions follow the electric field lines which terminate on all sides of the biased substrate and thereby help coat not only the front but all sides of the substrate. However, since the degree of ionization in many plasmas is very low (0.1-2%), it has been suggested that this mechanism is probably secondary to gas scattering of neutral atoms in contributing to the high throwing power of ion plating (ref. 8). Thus the energies of the impinging particles range from that obtained from ions accelerated by the potential difference between the anode and cathode to the thermal energy of unionized atoms. These factors, along with the heating of the substrate surface due to its continued bombardment, produce the graded-fused interface, which provides the superior adherence and improved mechanical properties characteristic of ion plating.

D. Objective

A Materials and Processes Laboratory objective is to have the capability to coat a wide variety of mechanical parts with a host of good-bonding protective thick (up to 40μ) films. Ion plating is well suited for providing good film bonding.

Electron beam evaporation of source materials provides a technique applicable to the widest range of materials, even the refractory metals which are vital to both high temperature rocket engines and to high temperature superconductors. However, most electron beam evaporation ion plating systems were not off the shelf but were custom built and debugged.

The immediate objective was to debug and put into smooth operation the trouble plagued (for its five year life) electron beam evaporation ion plating system Model 7375 custom built by Vacuum Technology Associates (now out of business).

E. Experimental Hardware

The ion plating apparatus used, V.T.A. Model 7375, was custom built by Vacuum Technology Associates and consisted of three main components: (1) The vacuum system, a Varian NRC 3117, attained an ultimate pressure of 10^{-6} torr using a 6" diffusion pump. (2) The electron beam gun system was a Sloan Multihearth 270° gun powered by a Sloan Model Five/Ten Power Supply. (3) The high voltage supply for the substrate bias was a 10 KV/500 ma unit made by Vacuum Technology Associates.

A typical ion plating apparatus is shown in Fig. 1.

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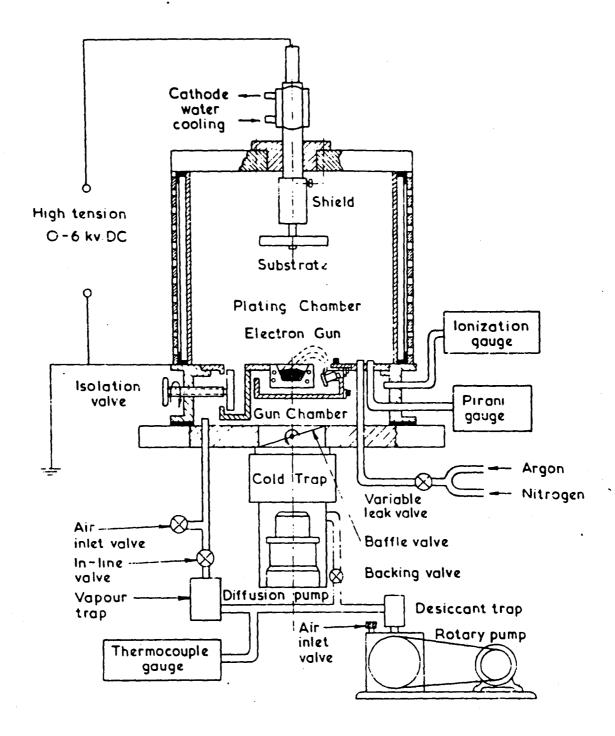


Figure 1. Ion Plating Apparatus with an Electron Beam Gun (D. G. Teer 1977)

F.1. Experimental Improvements Made

a. A brass cathode sample holder was removed from the system since brass was sputtercoating the substrate during precleaning and during ion plating. The chance of Zn contamination is thus reduced.

b. A cylindrical (5½ in. diameter by 3 in. long) grounded skirt extension was constructed which provides two inches of extra shielding to the top of the sample's support. This could reduce some of the above brass sputtering and efficiently concentrate the discharge near the sample substrate.

c. Due to modifications, the position of the sample holding rectangle may now be continuously located by a screw adjustment from direct contact with the cooled cathode plate to a position 5" lower.

d. Greater plating rates are anticipated as minimum sample substrate to hearth distances from about 4" to about 2". Some concern existed about reduced throwing power with the sample close; however, Hollidayll found no reduction in throwing power as hearth-substrate distances are reduced to from 5" to 3". The mean-free-path is in the several mm range so the transport is diffusion controlled.

F.2. Discussion of Recent Literature

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a. In a recent (Oct. '79) review article on ion plating sources, T. Spalvins¹⁶ mentions the following advantages of Wan's¹ hollow cylinder electron beam source:

- (1) It operates in the plasma.
- (2) It does not have the hot filament or the filament related problems.
- (3) It is easy to design and construct.

The main disadvantage is some contamination due to the cylinder sputtering into the substrate.

b. Ion plating by radio frequency induction evaporation is mentioned by Spalvins¹⁰. The source metal is placed in a ceramic crucible which is centered in the inducation coil. Frequency reduction by a factor of 1/6 to 75 MHz prevented coil arcing in the plasma.

- c. Rules for ion plating: 16
 - (1) Ideal conditions for ion plating include:
 - (a) Edge of cathode dark space is as close as possible to the boat.
 - (b) Dark space is as wide as possible.
 - (2) A cathode to anode distance of 15 cm is adequate to establish a glow discharge.
 - (3) A metal screen is used to coat plastics and insulators.
 - (4) Precleaning (surface oxide removal) is usually complete when the discharge current flattens out with time. It may take one hour for stainless steel.
 - (5) Bare wire should not be used to make high voltage connections according to Mattox.⁶
- d. References 19 and 20 are listed because they are recent.
- e. Discussion

The only reason which comes to mind for having a wide dark space is that the high electric field is only in the dark space and not in the plasma. The larger the dark space volume the greater the region from which to draw ions into the cathode substrate. If the boat to dark volume boundary distance is so reduced (the diffusion transport distance is small), the plating rate will be greater.

The metal shield stack (3" high x 4.5" diamter) is now recognized to have a detrimental effect since it acts as a Faraday (shield) cage around the hearth zeroing the E field needed to pull the positive ions to the substrate.

Even with the above stack shield removed, the top of the crucible is surrounded on 3 sides by a $1\frac{1}{4}$ " high by 3" wide magnetized rectangle having 3 sides $1\frac{1}{4}$ " above the top of the crucible (about 2" above the crucible bottom). This shielding could reduce the local electric field near the melted source enough to dramatically reduce the coating rate. The Sloan 270° evaporation hearth and gun were designed for vacuum evaporation and not for ion plating; thus, the irregular top profile is not unexpected. The shielding rectangle arms are magnetized and could not be ground down or removed without a magnetic compensation. Perhaps a large external Helmholtz coil could hopefully compensate and permit removal.

For both ion plating systems shown in ref. 9 and 10 the crucible cup top is flush with the flat baffle plate. Likewise in Teer's²¹ ion plating unit, nothing is shown higher than the hearth. The top of the hearth is flush with the flat baffle plate. See Fig. 1. The Heinz¹³ hearth is only slightly shielded by the horizontal gun.

The sparce $1/2 \mu$ coating on sample #11 could be due to excessive shielding of the source.

F.3. Vendor Search

On June 3rd letters requesting complete brochures with prices on electron beam evaporation ion plating systems and components were sent to all 16 vendors listed under Vacuum Coating Equipment on the <u>Science Instrumentation Index</u>. Not that a new system purchase is expected but that new components might be acquired for a much-needed upgrading of the V.T.A. system.

Denton Vacuum offered a complete unit with a DEG-801 Gun on the DIP-1 ion plating unit.

Hughes uses extensively, but rarely builds, ion plating units for customers. Mr. Cristy at 213/648-2345, Ext. 84369, may however be of help.

Nanotech* literature and prices just arrived listing a wide range of ion plating units with accessories. In fact they offer an option much like NASA's V.T.A. system except the e-beam gun is a Sloan 8 kW or 12 kW (costing about \$31,000) instead of our Sloan 5 kW. Unfortunately no gun mounting drawings were sent. Sloan no longer lists the 5 kW (Five/Ten) power supply. Since the V.T.A. system hearth is rated 12 kW it may be underpowered.

Inflicon Leybold-Heraeus responded indicating no ion plating systems for sale.

F.4. Films Plated

a. e-beam generated films.

All samples were #4340 hot roll steel disk one inch in diameter by 1/4 inch thick. Samples #3, #4, and #11 were

*Prestwich, Manchester M258WD England Phone: 061-773 8514

Sample #	Coating Date	BISBIC (µ Hg) P	Back Front (µm) Al Film Thickness	(in) Elect. Spacing	Shield Used Hearth? Cathode?	(mg) Δm	Film Material	Front u" Surf. Sm Subst.	r.m.s. oothness Film	KV (ma)	Time (min)	Power (kW)	Comments
3	6 June	5	8 <u>+</u> 2	3-4	Not Cathode		Al Marz Grade	100	100	.3 (4)	-	-	Numbers listed are estimates
4	12 June		2.5	5-6	Hearth & Cathode	1	Brass Al	100	100	1 (200)	-	-	
11	25 June	24	<u>0</u> .5	- 8	Cathode Top Collar Hearth	.3	Marz Al	5	5	1,5	17 (10- 35)	.5 to 2	
14	31 July	30	8	2	None	20	Marz Al	5	40	2 (20-5)	20		Resistance evap.
12	l Aug	30	12	2	None	14	Marz Al	5	30	3.5 (10- 20)	60		Resistance evap.

TABLE I (Accuracy ± 30%)

ion coated with Al during trial runs of the V.T.A. 7375 system. Samples #3 and #4 had surface roughness of ~ 100μ while #11 (12, 13 and 14) were polished to a ~ 5μ " smooth finish on one side. Film thickness on samples 3, 4, and 11 were ~ 8, 2.5 and .5 μ m respectively. Some direct evaporation was possible on #3 where the recorded pressure was in the 5μ range.

Sample #3 was mounted on a bare unshielded 1/2" rod. Sample #4 used the 5" cathode and grounded shield but was spaced away by about 3" by the brass interface. Sample #11 was flat mounted against the cooled cathode (to avoid heat induced blackening) and had an extra 2" of grounding skirt. Samples #4 and #11 also had a grounding cup shield around the hearth. Sample #11 had further shielding in the form of a brim ring around the cup top.

b. While the V.T.A. e-beam evaporation system was being repaired, the Denton resistance heated filaments were used to ion plate 8μ and 12μ Al coatings on samples #14 and #12, respectively. The unshielded sample cathode was screw mounted to a vertical 1/4" steel rod and positioned 2" above the filament.

c. Film surface roughness is several times greater than substrate roughness on samples #12 and #14.

d. The Al was M.R.C. Marz grade while the crucibles were Carbon EB-9 Union 76 (POCO).

F.5. Discussion

a. Samples #3, #4, and #11 were coated under exploratory conditions with the e-beam power being gradually increased. The cathode voltage was adjusted to prevent arcing in the chamber. The above variations plus the fact that only three films were coated suggests that only tentative conclusions should be drawn. The worst conditions (lack of shielding and close substrate source distance) gave the greatest film thickness.

b. Holliday^{11 & 12} and this author observed that their ion plated films were rather dull. Jones, Griffith, and Williams²³, ion plating on plastic, also found reflectivities reduced relative to vacuum coating, by about 2.5%, 10%, 18% for Au, Al, and Cu, respectively. Could the trace contaminants always present²² from e-beam evaporation hearth liners in vacuum plating Al account for the decrease reflectivity? Perhaps ion impact induced surface roughness makes the surface less specular (smooth). If one found that plasma cleaning on a 100% reflecting mirror surface causes loss of reflectivity in a lab experiment then the dullness in our films could be explained. The μ sized column film structure is common in ion plating and thus could easily account for the depressed reflectivities. c. The trouble with the V.T.A. system may be inefficiency due to lack of optimum adjustment and that troubles arise as increased high powers are used to compensate for the inefficiency. The Al was easily melted with .6 kW power (only 12% of the rated full power of 5 kW). The highest power recorded was 2.2 kW which was held for the last 5 minutes before the system blinked out 25 June. The Al melt temperature was considerably above the melting point. G.1. Suggested improvements which may be incorporated into the V.T.A. Model 7375 electron beam evaporation ion plating system are listed:

a. Baffle plate improvements

During the past two years previous prospective V.T.A. system users¹² designed a complex aluminum baffle plate using a small orifice which does an excellent job of providing the correct macroscope gauge pressure, both for the plasma above and for the filament chamber below the plate. The following further improvements are recommended for this baffle to lower further the filament pressure.

(1) Remachine this baffle totally from stainless steel instead of aluminum to match more closely the refractory materials (tungsten and molybdenum) in the Sloan e-beam gun. The present Al readily melts and evaporates around the Al orifice casting a direct shadow in the lower high vacuum region onto the gun and onto the high voltage electrodes causing shorts and arcs.

(2) Place stainless steel shim stock over the present Al baffle in the heat sensitive areas. This has been done.

(3) Bore a 1" diameter hole in the Al plate centered on the orifice hole. Machine a set of 7 stainless steel seats with orifice sizes 1/8, 3/16, 1/4, 5/16, 3/8, 7/16, 1/2. The smaller sizes would suffice for material which requires no sweep and would provide a higher vacuum around the filament.

(4) Use only a very small quantity of vacuum grease on the O-rings. Use only enough to give a wet look but not enough for the grease to be visible.

(5) A rectangular orifice may enhance the electron beam current while maximizing the gas flow impedance.

(6) The above rectangles could be continuously adjustable slits. If made of a bimetallic strip the adjustment would be automatic.

(7) Add extra fins and shields around the e-beam path to shield the exposed insulators from unintentional e-beam induced evaporation from around the orifice.

(8) The orifice could be placed with its axis horizontal at the top of the beam's arch, thus causing the gas to jet 180° away from the filament and anode, instead of the present 90° jet. A secondary advantage would be that more of the beam's path would be in the higher vacuum. A disadvantage would be increased coating around the orifice. (9) Again the orifice is only an ~{inch from the filament while the ion gauge is about 10" away. A gauge should be placed within an inch of the orifice with its short tube pointed toward the orifice jet to get the best indication of the filament pressure.

b. Improving other parts of the V.T.A. system other than the baffle plate:

(1) Place low melting point low vapor pressure materials on parts of the vacuum system for overheating indicators for sensitive components such as magnetic coils which must operate under 400 °C.

(2) Purchase at least all the parts listed on the suggested spare parts inventory (cost about \$2000).

(3) Add a shutter to shield initial volatile outgassing impurities from the substrate during the high vacuum soak of the crucible and source.

(4) In view of the manpower shortage connect a 4-pen recorder to the ion plating system to monitor the cathode voltage and current and the electron beam (emission) current. The fourth pen could monitor such quantities as sample temperature by thermocouple or magnet current.

(5) Dr. Arthur Nunes suggested using a focusing current ring (or an electrostatic ring) around the orifice to contact the beam to a finer thread to minimize orifice material evaporation.

(6) Apply the magnetic field sweep above the orifice so as to keep the orifice size as small as possible.

(7) Use a double hearth liner crucible stack with the inner crucible cut off to provide a lower target which should, because of greater insulation, rise to a higher temperature with less power to provide a higher plating rate.

(8) Entering the Ar leak from the bottom (instead of from the top) would cause some upflow which could enhance the transport of metal to the substrate increasing the deposition rate.

(9) Even with the cathode in its lowest position it is still ~ 6" from the hearth. Extending the substrate to the 2" to 3" range would produce a good enhancement in the deposition rate at no loss of throwing power according to Prof. Holliday's experimental data.¹¹ Suggested optimum distance for best ion plating is about six inches between substrate and source so some loss of adhesion may be expected. (10) Switch the two baffle O-rings to Viton to assure better high temperature performance.

(11) The V.T.A. 15 KV 500 ma cathode supply was found inoperative due to both a blown fuse and a blown milliampmeter. An 8 amp fuse was found necessary to withstand the current surges during critical arcing while providing some meter protection. This recently added fuse holder should be upgraded to a 15 KV rating.

G.2. Major additions to the V.T.A. Model 7375 ion plating system.

a. Acquire a "Balzers" e-beam gun and hearth which measures the <u>positive</u> ion current which is "proportional to the ion plating rate." Geometry dependent coating rates with ion plating are difficult to determine compared to vacuum evaporation.

b. Purchase an identical Sloan e-beam evaporation unit so that one unit may be refurbished while the other unit is in operation.

c. Use boronnitride-titanium diboride ribbon evaporator boats with resistance heating of Al as done by Walley and Cross¹⁴ in the article "An Aluminum Evaporation Source for Ion Plating." The electrodes must be water cooled and conducting graphite paper provides better electrical contact to the electrodes. This BN-TB₂ is cited in the Handbook of Thin Film Technology, page 1-47, as a crucible material sold under the name (HDA Composite Ceramic) by Union Carbide, New York, NY.

G.3. Conclusions

The V.T.A. system shows promise in that it operated for about a month with no major problems. The V.T.A. low coating rate will probably improve when the excess electrostatic shielding is reduced. However for quick ion plating of Al, Denton resistance evaporation unit is much more reliable.

REFERENCES

- 1. Mattox, D. M., "Film Deposition Using Accelerated Ions," Rept. No. SC-DR-281-63, Sandia Corp., (Nov., 1963); Mattox, D. M., Electrochem. Technol., <u>2</u> 295 (1964).
 - "Versatile Aluminum Coating Controls Corrosion," Iron Age, (February 9, 1976).
 - 3. Spalvins, T., "The New Applications of Sputtering and Ion Plating," NASA TMX-73551 (1977).
 - 4. Chambers, D. L. and Carmichael, D. C., Research/Development, <u>22</u>, 32 (1971).
 - 5. Spalvins, T., J. Am. Soc. Lub. Eng., 27, 40 (1971).
 - 6. Mattox, D. M., Proceedings of the Conference on Sputtering and Ion Plating, NASA SP-5111, NASA Lewis Research Center, Cleveland, Ohio, p. 41 (1972).
 - 7. Mattox, D. M., J. Vac. Sci. Technol., 10, 47 (1973).
 - 8. Aisenburg, S. and Chabot, R. W., J. Vac. Sci. Technol., <u>10</u>, 194 (1973).
 - Smith, H. R., "Current Developments in Ion Plating," Society of Vacuum Coaters Conference, Detroit, Michigan, March 23, 1972.
- 10. Harker, H. R. and Hill, R. J., "Deposition of Multicomponent Phases by Ion Plating," Journal of Vacuum Science and Technology, <u>v.9</u>, #6, p. 1395 (1972).
- 11. Holliday, R. J., "Ion Plating Investigation," BER Report No. 224-94, The University of Alabama, p. X111 (Sept. 1978).
- 12. Holliday, R. J., "Ion Plating Studies for High Temperature Applications," University of Alabama in Huntsville, Report No. 227, p. X111 (Oct. 1979).
- 13. Heinz, B. and Kithel, G., "Ion Plating with Electron Beam Evaporation," Conference on Ion Plating and Allied Techniques, p. 73, Edinburgh, 1977.
- 14. Walley, P. P. and Cross, K. B., Ibid., p. 82.
- 15. Maissel, L. I. and Glang, R., Handbook of Thin Film Technology, 1970, McGraw Hill, New York.

- Spalvins, T., Survey of Ion Plating Sources, Amer. Vac. Soc. 26th National Symp., New York, NY, (Oct. 2-5, 1979).
- 17. Wan, C. J., Chambers, D. L., and Carmichael, D. C., J. Vac. Sci. Technology 8, Vm. 99 - Vm. 104 (1971)
- 18. Spalvins, T. and Brainard, W. A., NASA TMX-3330 (1976).
- 19. Masatoshi, Y., et. al., "Studies on Ion Plating Process" for National Symposium and Exhibition, San Francisco, CA, Society for the Advancement of Material and Process Engineering, (May 1979).
- 20. Cappelli, P. G., "Coating Processes," Proceedings of the Conference on High Temperature Alloys for Gas Turbines, p. 177 of Applied Science Publishers (1978).
- 21. Teers, D. G., "The Ion Plating Technique," p. 13, op. cit. Ref. 13.
- 22. Graper, E. B., "Evaporation (Vacuum) of Al from an Electron Beam Source," Journal of Vacuum Science and Technology, g #1, 33 (1971).
- 23. Jones, K., Griffith, A. J., and Williams, E. W., p. 115, op. cit. Ref. 13.

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1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

STATISTICAL WIND PROFILE GUST MODEL

Prepared by:

Deva C. Doss, Ph.D.

Academic Rank:

Professor

University and Department:

The University of Alabama in Huntsville Department of Mathematics

NASA/MSFC: (Laboratory) (Division)

Space Sciences Atmospheric Sciences

MSFC Counterpart:

Date:

Contract No.:

O. E. Smith

August 1, 1980

NGT-01-002-099 (University of Alabama)

STATISTICAL WIND PROFILE GUST MODEL

ΒY

Deva C. Doss, Ph.D. Professor of Mathematics The University of Alabama in Huntsville Huntsville, Alabama

ABSTRACT

The purpose of this study is to develop a statistical wind profile gust model for the Space Transportation Operations and Trade Studies by using 1800 Jimsphere wind profile data collected at Cape Kennedy during 1965-1972. Wind profiles from the surface to 20 km in component form, i.e., zonal and meridional are processed through the digital filters of different wave length ranges based on the Martin-Graham cosine rolloff model. The residuals obtained from the filtering processes form the data base for the statistical analysis.

For each wind component the gust and gust length at a specified reference altitude in a residual profile are defined. A two-parameter gamma probability marginal distribution seems to fit well the component gust amplitude, and the gust length when redefined. The problem of finding an appropriate bivariate joint distribution of the gust amplitude and length remains to be solved.

The probability distribution of the modulus of the gust amplitudes has been derived under the assumption that they are independently distributed as gamma variates. It seems to fit the observed data.

INTRODUCTION

Accurate wind profile measurements from the surface of the earth to 20 km are made possible by the Jimsphere-radar system. This system consists of a 2 m diameter, constant-volume balloon and an AN/FPS'-16 or equivalent high precision radar which tracks the position of the balloon. The position data collected by the system are smoothened over a 25 m interval. The resulting data are differenced over 50 m intervals to produce wind profile data points.

This study utilizes 1800 wind profile data in component form collected at Cape Kennedy during 1965-1972. They are processed through the digital filters of different wave length ranges based on the Martin-Graham cosine rolloff model. The residuals obtained from the filtering processes form the data base for the statistical analysis (see Figure 1).

1.1 Definitions

Adelfang [1] in a recent report defines gust amplitude and length for a wind component in a residual profile at a specified reference altitude H_0 . For instance in Figure 2 the gust amplitude (gust) for the zonal wind component u' is defined as the maximum value of |u'| in the vicinity of altitude H_0 with like sign to u' at H_0 . The gust length L_u is defined as the altitude difference of the zero crossings on either side of the gust, i.e.,

$$L = H_2 - H_1 \tag{1}$$

where H_1 = altitude of the first zero crossing for the downward scan,

 H_2 = altitude of the first zero crossing for the <u>upward</u> scan.

Similarly, the gust amplitude and length are defined for the meridional gust component v'.

1.2 Purpose of the Study

One of the objectives of this study is to develop a theoretical probability gust model for each component as well as a joint model for both the components. Then the probability distribution of the modulus of the gust amplitudes can be derived.

GUST ANALYSIS

In statistical literature it is not uncommon to fit a gamma probability

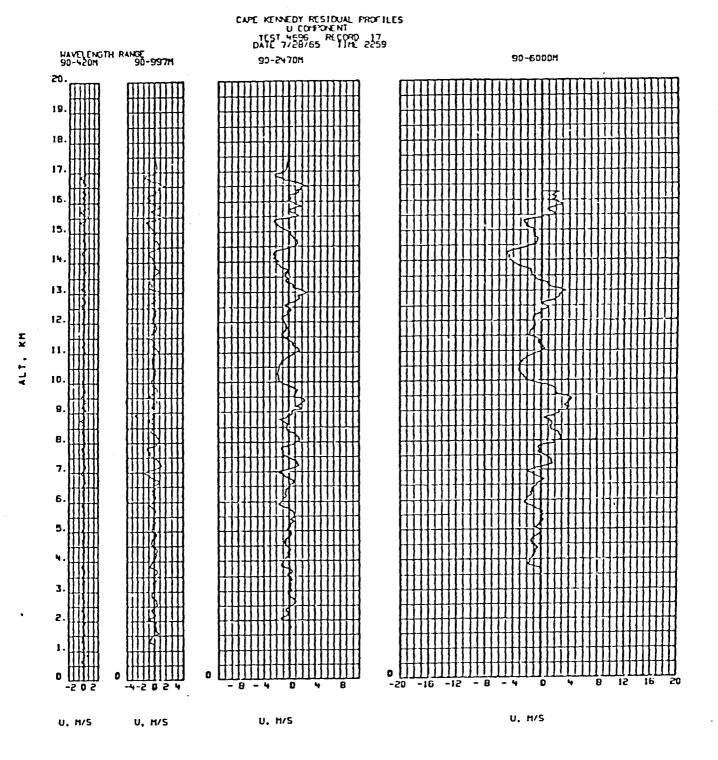
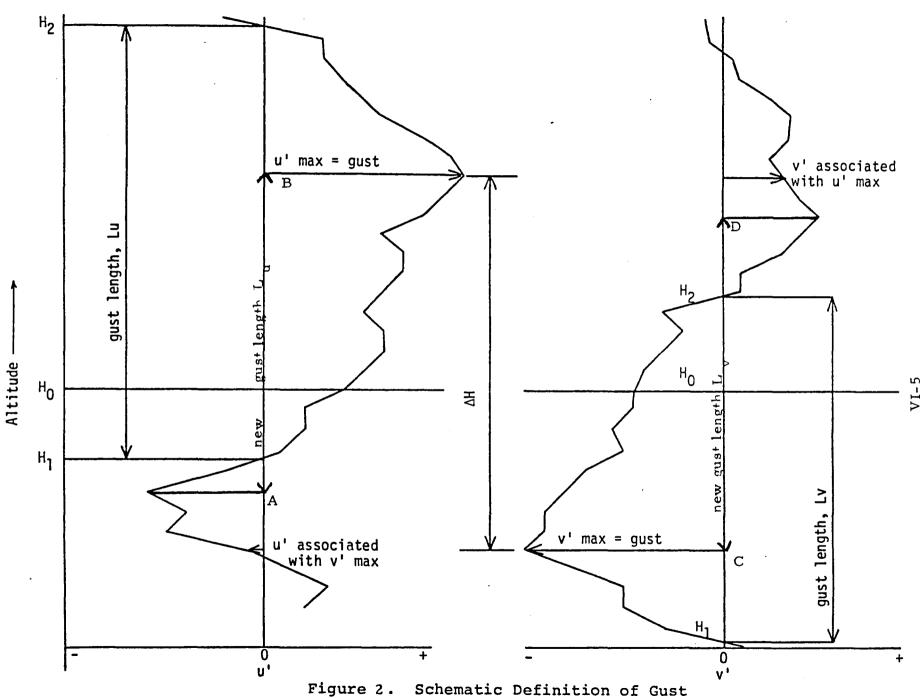


Figure 1. Cape Kennedy Residual Profiles

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model to a climatological measurement. However, when more than one random variable is involved, it is not yet clear whether or not a particular species of several multivariate gamma probability distributions is adequate to describe the behavior of random variables under consideration. This difficulty stems from the fact that any given univariate marginal distributions do not uniquely determine their joint probability distribution. If one marginal of a component of a random vector happens to be different from that of another component, e.g., gamma and beta, then the complexity of reconstructing their joint distribution increases.

2.1 Gust Amplitude

Adelfang [1] has shown that the gust amplitude of any wind component at a given reference altitude for filters of different wave length ranges seems to be distributed as a two-parameter gamma distribution and the fit seems to be good. A gamma probability density of a random variable x is here defined as \checkmark \checkmark \checkmark

$$f(x) = \beta x^{-1} \exp(-\beta x) / \Gamma(Y), \quad x \ge 0, \quad (2)$$

where scale parameter β and shape parameter \forall are positive and estimated by the maximum likelihood procedure as in Thom [3]

$$\hat{\beta} = \hat{\gamma} / \bar{x}$$
(3)

and

$$= \left\{ 1 + (1 + 4A/3)^{\frac{1}{2}} \right\} / 4A_{3}$$
(4)

$$A = ln(\bar{x}) - \overline{ln(x)}, \qquad (5)$$

2.2 Gust Length

Although the gamma and truncated normal probability distributions seem to fit fairly the data on gust length, a pronounced bimodality appears to persist in the observed frequency distribution of the gust length. Therefore, there are two options to pursue:

- 1. To fit a mixture of two distributions
- 2. To redefine the gust length and hopefully avoid bimodality.

Since the first option involves computation of more parameters and analytical complexity, a redefinition of the gust length is in order. Instead of defining the gust length as the distance between two successive zero crossings, it can be defined as the length of the interval containing the reference altitude H_0 , whose end points form successive extrema of the wind profile. In Figure 2, the distance from A to B is the new gust length for u' and the distance from C to D is the new gust length for v'.

It is encouraging to find that newly defined gust length seems to fit a gamma probability distribution without exhibiting any bimodality. It is premature at present to accept this observation without further investigation.

2.3 Joint Distribution of Gust Amplitude and Length

Once it is assured that the gust length follows a gamma model a search for an appropriate bivariate gamma distribution should be made for the gust amplitude and length. Three bivariate gamma distributions are given in Mardia [2], p. 88. When the shape parameters of gamma marginals are almost equal, the first bivariate distribution involving a modified Bessel function may be considered. Because of the restriction, namely, $y \ge x$ imposed on the third distribution, it cannot be employed in fitting the gust amplitude and length.

2.4 The Modulus of the Gust Amplitude

Since it is found that the correlation between the gust amplitudes of the components u' and v' is less than 0.2, the maximum of modulus

$$R = \{ (u')^2 + (v')^2 \}^{1/2}$$
(6)

can be derived from the gamma marginal of the gust amplitudes under the assumption that they are independently distributed. In fact, the probability density of R is:

$$\frac{1}{4} \frac{\beta_{1}^{Y_{1}} \beta_{2}^{Y_{2}} R^{Y_{1}+Y_{2}}}{\Gamma(Y_{1}) \Gamma(Y_{2})} \sum_{n=0}^{\infty} \frac{(-\beta_{1}R)^{n}}{n!} \sum_{m=0}^{\infty} \frac{(-\beta_{2}R)^{m} \Gamma(Y_{1}+n) \Gamma(Y_{2}+n)}{m!} \Gamma(Y_{1}+Y_{2}+n+m+2)$$
(7)

where all parameters β_1 , β_2 , γ_1 , γ_2 are positive. This has to be validated by the data.

2.5 Sum of Two Adjacent Amplitudes

Denote the amplitude of the zonal component at the reference altitude H_0 by |u'| and the amplitude just above by $|u'_2|$. Then it has been found for a set of data that the correlation between them is in the vicinity of 0.5. However,

the observed data on their sum $|u'_1| + |u'_2|$ seems to fit a gamma distribution. (Similar observation may be true also for the other component.) It raises a theoretical question: When is the sum of two correlated gamma random variates distributed as a gamma?

CONCLUSIONS AND RECOMMENDATIONS

This report presents several unanswered questions to which answers have to be sought. It is, with reasonable assurance, seen that the gust amplitude follows a two-parameter gamma distribution the probability model of the newly defined gust length seems promisingly to be a gamma distribution.

The joint distribution of the gust amplitude and length, and the distribution of the modulus of gust amplitudes need to be explored. In fact finding the joint distribution of the gust amplitudes and length, of both the components should be the ultimate goal. One realizes that there are many loopholes in this investigation, but it is hoped that by this time next year the statistical model for the gust is resolved.

REFERENCES

- 1. Adelfang, S. I., "Vector Wind Profile Gust Model." Computer Sciences Corporation Final Report for NASA/MSFC under Contract NAS8-33433, April 9, 1980.
- 2. Mardia, K. V., "Families of Bivariate Distributions." Griffin's Statistical Monographs and Courses. Hafner Publishing Co., Darrien, Conn. 1970
- 3. Thom, H. C. S., "Some Methods of Climatological Analysis." WMO Technical Note 81, WMO-MO 199. TP. 103, 1966

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1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

UPGRADING THE FREE FLYING RENDEZVOUS AND DOCKING SIMULATOR AND THE ORBITAL SERVICER SYSTEM

Prepared By: Robert M. Eastman, Ph.D., P.E Professor University of Missouri - Columbia Department of Industrial Engineering NASA/MSFC: Electronics and Control (Laboratory) Guidance, Control & Instrumentation (Division) Control Mechanisms (Branch)

Kenneth R. Fernandez

July 25, 1980

NGT-01-002-099 (University of Alabama) .

Academic Rank:

University and Department

MSFC Counterpart:

Date:

Contract No.

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UPGRADING THE FREE FLYING RENDEZVOUS AND DOCKING SIMULATOR AND THE ORBITAL SERVICER SYSTEM

ΒY

Robert M. Eastman, PH.D., P.E. Professor of Industrial Engineering University of Missouri Columbia, Missouri

ABSTRACT

The objective of the research was to make recommendations for upgrading two teleoperator/robotics test and simulation systems based upon a review of latest technology advances in the involved disciplines. The Free Flying Rendezvous and Docking Simulator has two vehicles, the Free Flying Mobility Unit and the Free Flying Target Assembly, operating on air pads on the air bearing epoxy flat floor. The Orbital Servicer System has a six degrees-of-freedom manipulator arm for transferring modules between the Orbital Servicer and a prepared satellite.

Based on the research, a new second-generation Free Flying Mobility Unit is It will add a sixth degree-of-freedom and incorporate other recommended. improvements which will greatly expand the Center's capability to perform evaluation tests and demonstrations of advanced systems concepts for rendezvous and docking in support of the Teleoperator Maneuvering System (TMS) Program. Current plans for the TMS call for a technology readiness in 1985 for placement and retrieval of a satellite/payload in a predetermined orbit with the TMS moving from and returning to the Shuttle Orbiter. The second generation Free Flying Mobility Unit will incorporate TMS design features which will enable it to perform realistic simulations for evaluating rendezvous concepts, ranging sensors, lighting and video systems, and docking mechanisms. At a later time it will incorporate advanced technology features for future TMS missions and will be invaluable in defining the requirements and evaluating promising concepts for automated rendezvous and docking.

The Orbital Servicer System provides the capability for testing and demonstrating concepts for on - orbit servicing of compatibly designed satellites/payloads. The TMS will be the transporting vehicle for the servicer. The manipulator arm of the Orbital Servicer System is presently computer controlled in the trajectory portion of the module transfer operation. The ultimate objective is to fully automate its operation having the manipulator arm move from its stowed position, latch on to a mock-up faulty module, remove the module and stow it in the base of the servicer, latch on to a replacement module, place it in the satellite, and having the manipulator arm return to its stowed position. This will require additional capabilities in several fields basic to robotics research. These include sensors, artificial intelligence, image analysis, communications, computer programming, pattern recognition, kinematics and manipulator design. It is recommended that the Electronics and Control Laboratory move to acquire the basic competencies in robotics necessary to achieve the major objective of full automation.

ACKNOWLEDGMENTS

My gratitude to the many NASA people who helped with the research and contributed to our pleasant summer in Huntsville.

Kenneth R. Fernandez as my NASA MSFC counterpart helped with explanation and demonstration of the robotics facilities and test programs. John B. E. Chase was generous with his editorial help, arrangements, and explanation of the factors and issues involved. Ed Guerin and Dave Henderson (Essex Corporation) explained the Free Flying Rendezvous and Docking Simulator and arranged for me to observe the tests. Don Scott and Keith Clark demonstrated the Orbital Servicer System. Charles Cornelius provided logistic support and organizational enlightenment. F. Brooks Moore and Eugene Smith furnished the management support without which no research can prosper. Other members of the Electronics and Control Laboratory helped with information and support.

Mary F. Smith typed and retyped the manuscript and helped in numerous other ways. Eloise Meyer, Bettye Morey and Gayle Dodson were gracious and helpful. Marion Kent's advice was most useful and appreciated.

Dr. Robert Barfield, Project Director, handled the contract substance assisted by Barbara Skaggs. The University of Alabama in Huntsville faculty were gracious hosts socially and professionally.

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INTRODUCTION

The Free Flying Rendezvous and Docking Simulator consists of the Free Flying Mobility Unit and the Free Flying Target Assembly both of which operate on the air bearing epoxy flat floor. The two vehicles are mounted on air pads which provide frictionless motion in one plane for simulating translational movement in space. Simulated propulsion is provided by compressed air jets under the control of the remotely located operator. The simulator is illustrated in Figures 1, 2, and 3 and described in reference (1).

The simulator is an engineering tool used to define, test, evaluate and demonstrate the rendezvous and docking techniques and procedures, and the components and systems designs needed for rendezvous and docking a remotely controlled teleoperator vehicle with a satellite payload. The data recorded include air consumed (measure of fuel consumption), time to dock, docking success, lighting patterns and camera positions. The evaluation and test data influence the selection of the most promising concepts and designs and the refinement of them for potential use in future projects.

The Orbital Servicer System is an engineering model of an on-orbit maintenance system. It contains full-scale typical satellite replacement modules which are exchanged by the six-degree-of-freedom manipulator It also includes the interfacing control system for the manipulator arm. The Servicer System model is used to test, evaluate, and demonstrate arm. the on-orbit maintenance concept and to define design features that would enhance its performance. At present (1980), the operation is partially automated. A computer program controls the teleoperator arm trajectory movement until the module is close to its destination. Then control is transferred to the human operator who completes the module transfer. He operates from a remote location using video cameras and manual controls. The servicer is illustrated in Figure 4 and described in references (2) and (3).

One of the key questions concerning the Orbital Servicer is whether the teleoperator arm can be fully automated to eliminate the human operator (man-in-the-loop). There are advantages and disadvantages to each alternative and strong advocates on each side. A major purpose of the Orbital Servicer System is testing the feasibility of the fully-automated transfer.

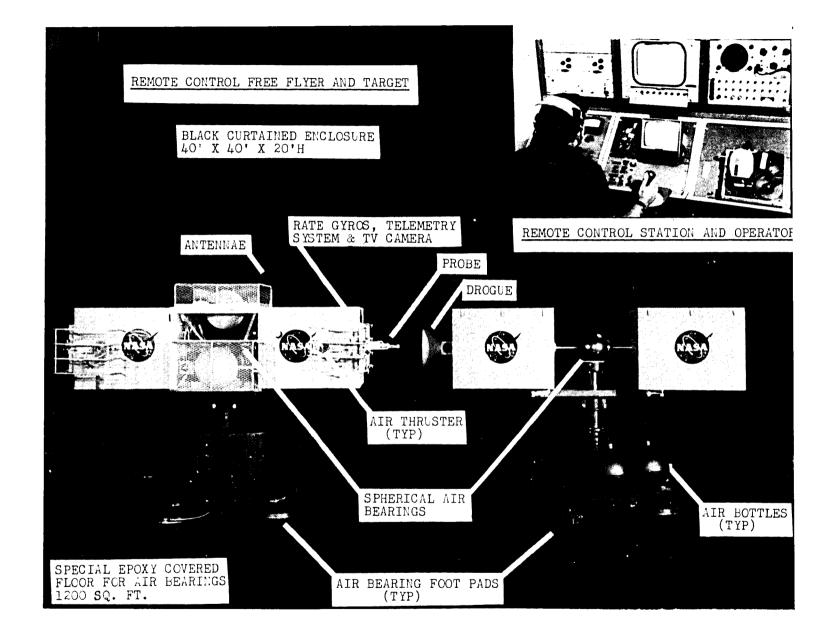


FIGURE 1. FREE FLYING RENDEZVOUS AND DOCKING SIMULATOR

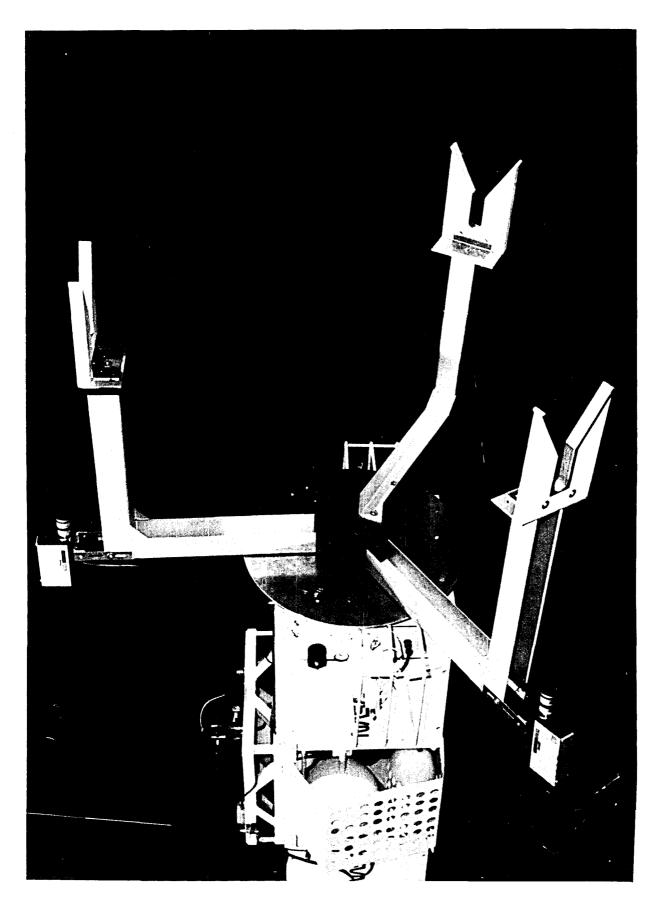


FIGURE 2. FREE FLYING MOBILITY UNIT

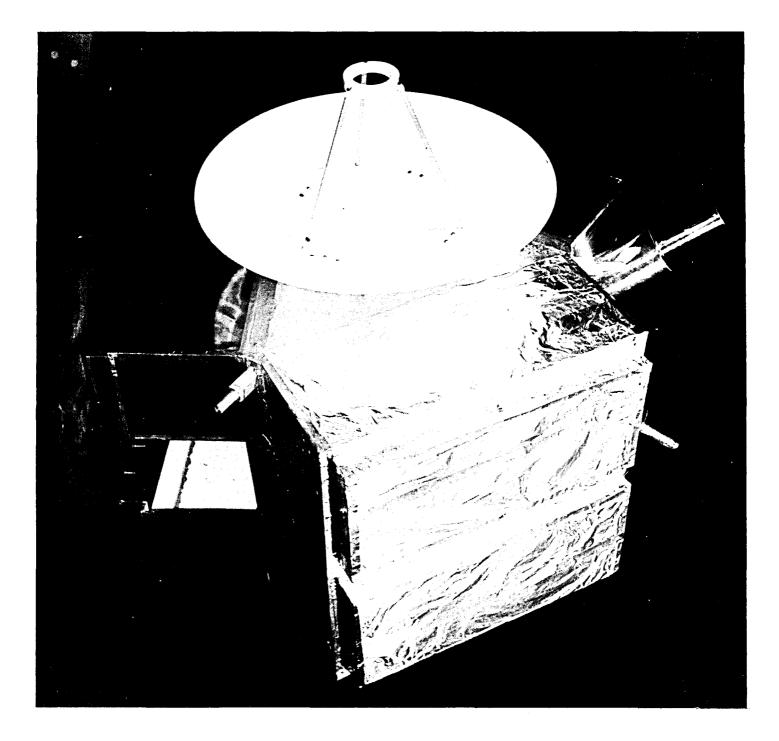


FIGURE 3. FREE FLYING TARGET ASSEMBLY

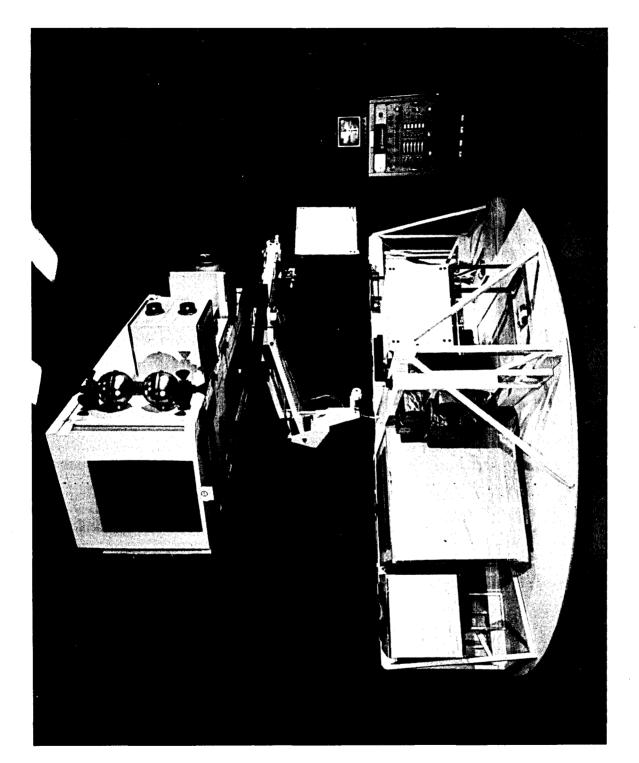


FIGURE 4. ORBITAL SERVICER SYSTEM

OBJECTIVES

The first research objective is to recommend design improvement for the Free Flying Rendezvous and Docking Simulator based upon a review of latest developments in associated disciplines. The subsidiary objectives are:

1. A 6 DOF Free Flying Mobility Unit which could be either a modification of the existing 5 DOF vehicle or a new vehicle.

2. Automatic docking capability.

3. Alleviation or elimination of some current limitations including:

(a) Reliability and maintenance problems with existing equipment.

(b) Physical balance of the current equipment, i.e., stability in pitch and roll.

The second research objective is to recommend design improvements in Orbital Servicer System based on review of latest developments in appropriate disciplines. Subsidiary objectives are:

1. Automated fast servicing of a class of satellites by robotic arm from the Orbital Servicer vehicle.

2. To review the proposed new electrical self-aligning connector design.

3. To suggest improved methods of rack positioning.

METHOD OF INVESTIGATION

The following methods of investigation were used:

1. Literature search.

2. Material and information from the NSF Robotics Research Workshop held April 15-17, 1980, at the University of Rhode Island and from personnel at the McDonnell Aircraft Company (part of McDonnell Douglas Corporation in St. Louis, MO).

3. Observation of the Free Flying Rendezvous and Docking Simulator and the Orbital Servicer System.

- 4. Discussions with MSFC personnel.
- 5. Sketches and calculations.
- 6. Review of previous and associated work.

SYSTEMS INTEGRATION

NASA Headquarters has assigned MSFC the mission of "Systems Integration" or the co-ordination and interfacing of teleoperator/robotics sub-systems and components into an effective total system to support future NASA programs. The term "System Integration" also connotes optimizing the effectiveness of the system while using scarce resources efficiently.

The two lab facilities, the Free Flying Rendezvous and Docking Simulator and the Orbital Servicer System, are essential for evaluating and demonstrating the integration of the functions and the system performance in a simulated space mission. The full-scale dimension provides the experimental setting necessary for accurate test and evaluation. Scale models can be helpful but there are too many possibilities for discrepancies between actual mission systems and part-scale simulators to permit complete reliance on the part-scale model. Computer modeling and simulation can be efficient and time-saving; however, their results require full-scale physical verification because of the possibilities of conceptual errors and data omissions and inadequacies.

BASIC ROBOTICS RESEARCH OVERVIEW

This section will be devoted to a discussion, necessarily brief, of several fields of basic robotics research pertinent to the main objectives of full automation of the Orbital Servicer Teleoperator System and to upgrading the Free Flying Rendezvous and Docking Simulator.

The following research areas are fundamental to advanced teleoperator and robotics development:

1. Artificial intelligence

2. Sensors

- 3. Kinematics
- 4. Control
- 5. Communications

- 6. Manipulator design
- 7. Image analysis
- 8. Locomotion

ARTIFICIAL INTELLIGENCE

Artificial intelligence is the branch of computer science which studies the way machines can replicate human actions requiring intelligence. It involves receiving and evaluating environmental data and making a decision based upon the data and a set of decision criteria. At an advanced level, the machine can learn from experience and adjust its behavior accordingly. Artificial intelligence is necessary for any teleoperator or robot which has an operation which is not completely and undeviatingly predetermined.

Systems have been developed which will perform at least one but never all of the following: (4)

(1) Plans can be generated at multiple levels of detail.

(2) Plans can be viewed as partially ordered sequences of actions with respect to time.

(3) Each action is expected to produce a single state change characterized by a single primary effect.

(4) A plan is not generated at all unless the planner determines that it will be totally successful in meeting all specified goals.

(5) Simple plans requiring information gathering can be generated.

(6) Unsophisticated techniques for dynamic repair of unsuccessful plans during execution have been developed.

(7) Plans can be used to control robot devices with simple use of sensory feedback and systematic replanning.

Major research problem areas in intelligent robots are: (4)

1. Planning for parallel execution

2. Planning for information gathering

3. Planning for planning

4. Learning: Using a plan data base

5. Interactive planning

6. Dynamic plan repair

7. Distributed robotics (i.e., disbots)

a. Hierarchal planning

- b. Control
- c. Basis
- d. Communication
- e. Plan execution

The central problem of robotic intelligence in the real world is that a robot may have a large number of alternative actions, but the user does not want to be burdened with specifying them and the computer is too busy to compute them anyhow (5). -----"Intelligence and decision making" usually mean the ability of a robot to sense some aspect of its environment and take alternative courses of action.

SENSORS

A sensor is a device which collects data from the environment for transmission to the robotics command system. An important research issue in robotics is sensory control in which the sensor is the first link in the chain which finishes with the completed robot action. Fast and accurate sensing is essential to reaching the goal of fully automated space robotics (6).

Among the research areas to be addressed are:

- 1. More capable visual processing
- 2. Faster visual processing
- 3. Better dynamic models of arms
- 4. Tactile and force sensors with desirable characteristics
- 5. Better understanding of the role of compliance

Visual sensing is important in space applications. Some needed improvements are (7):

- 1. Higher resolution, dimensional density and precision
- 2. Improved pixel (i.e., picture element) quality
- 3. Color discrimination
- 4. Better lenses
- 5. Better cameras
- 6. Increased robustness
- 7. Controlled illumination
- 8. Software development
- 9. Range and proximity sensing

Sensor types which have negligible usefulness for space functions are tactile, force, acoustic and temperature.

Radar as a sensor has advantages and disadvantages. The distances involved may be too short to permit radar to function properly. Even more difficult is the problem of interfering reflections from structural components of the system. These reflections will prevent the system from receiving clear and accurate readings on the position of any system component.

However, some current and research in radar may facilitate using radar in fully automated orbital servicer. It has the capability of rapid location of an object and determination of its velocity and direction of movement provided the reflection problem can be solved.

Video cameras produce a good picture for the human operator. The major problem is automated image analysis of the picture. (Image analysis is the conversion of visual pictures such as video to a set of impulses which can be processed and interpreted by a digital computer). The human operator is much better at analysing the video picture and taking appropriate action than is an automated system.

Lasers may be the best sensors for space robotics applications. Much additional research is needed. Light-beam sensors (photoelectric cells) can not produce the detailed image needed. Acoustic sensors lack the medium to transmit sound wave. Contact sensors would require delicate, easily-damaged antennas and would probably be unable to gather enough environmental data to be useful.

Sensors present several robotic control problems such as data filtering and smoothing, multidirectional sensors and data acquisition. Control structures, particularly hierarchical, and control logic need further work. The man-machine interface requires additional development (12).

KINEMATICS

Kinematics and mechanical design refer to the static and dynamic relationships of the physical components of the teleoperator manipulator system. Examples are mass, velocity, momentum, strength and deflection. Correct handling these factors is vital to successful teleoperator and robot operation. Several important research areas remain.

Kinematic preformance can be either time-based (velocity and acceleration) or geometric-based (space and orientation). Further work is needed in both areas (8). This effort should get away from position-byposition analysis to a broader knowledge which eliminates the need for the detailed position approach. A general solution to the optimal number of degrees-of-freedom has yet to be resolved. More DOF produce greater flexibility and versatility; fewer DOF produce a system easier to design and control, cheaper to build and generally more reliable.

Some robotic configurations are more efficient than others. However, the knowledge about these configurations is empirical rather than theoretical. Among the research questions to be resolved are:

1. The optimal manipulator configuration for flexibility, dexterity and reliability.

2. Actuator placement and control.

3. Basic design theory for fluid and electrical actuators.

4. Increasing the number of design parameters which can be rationally chosen given the present state of the art.

5. Elimination or minimization of elastic vibration due to system resonance.

6. Coping with large inertial forces at high speeds.

Another practical design problem is that robot and teleoperator performance fall short of theoretical optimum. Some causes are mechanical interference by structural members, limited rotational range and the mass of components particularly the motors (8).

Speed of operation of robots is vitally important to industrial applications, but less so to space applications. High-speed dynamic effects become critical yet little is known definitively about these effects. Present manipulator configuration design and trajectory are not the best for high speed operation. More needs to be done (9). Three difficulties with high-speed operation are:

1. Non-linear dynamic effects

2. Structural flexibility

3. Discrete time increments resulting from micro-computer controls produce undesirable effects.

Some modifications of manipulator design may be necessary to make the design compatible with automated operation. Present plans are based on the assumption that full automation can be achieved by rewriting the control computer program. This may or may not be a valid assumption. At the very least, the manipulator design should be reviewed for compatibility with full automated control.

CONTROL

Control theory is the branch of engineering and science which deals with monitoring and direction of the teleoperator to achieve the desired result. Control theory has two facets:

1. Systems study and modeling and (2) measuring (sensing) the states of the system. System and task modeling are essential to advanced robot control development. Modeling can be defined as quantitative descriptions of the system and tasks in measurable, calculable and controllable terms.

Among the research issues are sensitivity studies of the model, robot control in feedback schemes and multi-unit robotic systems (10).

Four assumptions of robotics control theory are:

1. Robots are complex articulated mechanical devices designed to perform operations normally requiring human skills.

2. Robots are equipped with sensors capable of measuring internal and external states.

3. Robots are equipped with computers.

4. Humans can communicate with robots.

Paul discusses control applied to robotic operations in space. He postulates 6 DOF with either velocity or force but not both specified along each DOF. He also discusses kinematic control and Jacobian control. Each method has advantages (11).

Graupe and Saridis hold that technological advances of the last fifty years depended upon development of system theoretic methodologies (14). Intelligent control uses the digital computer along with advanced techniques of system theory to produce a unified engineering approach. Based on this, control systems for robotic tasks and systems have been developed for prosthesis and orthotic devices and for general purpose robots. They propose a hierarchally intelligent control approach distributed according to "decreasing precision with increasing intelligence." The levels would be:

- 1. Organization
- 2. Co-ordination
- 3. Hardware Control

Present computer-based systems for robot control require the program to direct all movements and actions of the robot. A specially-designed computer language may be used although some programs are in a standard language such as Fortran. The program can include provision for receiving external stimuli and directing action based on the stimuli received. However, there is no provision for stimuli not in the program or for unprogrammed reactions no matter how desirable (15). Libraries of sub-routines can be developed since the programs can be independent os specific locations. The number and usefulness of sub-routines handling contingencies, errors and emergencies has increased.

The major deficiencies of the second-generation programmed robotic systems are (15):

1. The programs are difficult to debug and modify.

2. Achieving the final position with maximum accuracy is inefficient.

3. It is difficult to foresee all possible emergencies and contingencies. This is particularly important in the Space Shuttle - Orbiter Servicer system. In the next (third) generation of computer robotics control systems, the robot will formulate an appropriate plan of action based on goal directions received from the operator. The robot will have greater artificial intelligence capabilities and much more data than it needs with the present computer program motion direction system. Research on this system is now underway (1980).

The elementary "Pick-and-place" robotic systems are relatively easy to design, program and implement. Many industrial robots are in this category. However, they are unable to handle effectively and reliably events and environments which deviate even slightly from the environment for which the program was designed. However, as the program is modified to handle environmental variations, the complexity of the problem domain and the resulting program increase many-fold (13).

COMMUNICATIONS

Communications will be a major robotics design problem. Data impulses from sensors must be transmitted to the control computer and instructions sent from the control center to the teleoperator components. Another major R&T issue is the computer languages and control algorithms needed for full automation. This is inextricably interconnected to the determination and optimization of the desired operating characteristics.

IMAGE ANALYSIS

The role of image analysis will be the identification and location of the components of the shuttle and the servicer. These components include:

1. Structural components of the Shuttle.

2. Structural components of the Servicer.

3. Alignment of the space vehicle (Servicer and satellite) during docking.

4. The modules at the beginning, during, and of the end of the transfer process.

5. Alignment of the rack which will receive the module.

6. Any component out of its regular position to the extent it might cause damage.

7. Any stray object in a position to cause a collision.

Image analysis is a particularly difficult field in robotics research and application. The processes by which a human operator's mind receives signals from his sensors, and evaluates, interprets and identifies the object seen, its location and its orientation are incredibly complex. Performing these processes automatically has been done only in simple situations with great difficulty.

Basic research in image analysis is going forward. Further progress is essential to the full automation of the Orbital Servicer Teleoperator.

LOCOMOTION

Locomotion is defined as movement of the robot over an irregular terrain surface. This contrasts with space travel in which the robot has no terrain contact. At a cursory glance, locomotion has little to do with the NASA Space Shuttle and Orbital Servicer missions. However, some locomotion research results have application to the simulation of rendezvous and docking operations. Furthermore, locomotion could become important to the MSFC robotics efforts if the mission is changed to include mobile robots for space applications.

THE FREE FLYING RENDEZVOUS AND DOCKING SIMULATOR

The present Free Flying Rendezvous and Docking Simulator uses aging equipment of limited capabilities. The two vehicles were assembled partly from surplus and scrounged components. They require considerable maintenance with increased down time and decreased capability. Further, the data output is limited in comparison to the data potential of a simulator with better equipment and instrumentation.

Second - Generation Free Flying Mobility Unit

A major research and policy question is "Should NASA build a new, second-generation Free Flying Mobility Unit or should it rebuild and improve the present vehicle?" In either case, the vehicle should incorporate a sixth degrees-of-freedom (6 DOF), balance stability in the pitch, roll and Z axes, increased reliability and more data acquisition instrumentation.

The general specifications for a second-generation Free Flying Mobility Unit are:

- 1. Incorporate six degrees- of-freedom.
- 2. Maintaining the vehicle balance during docking maneuvers.
- 3. Minimize weight without interfering with operating characteristics.
- 4. Be compatible with existing equipment and existing data.
- 5. Provide for installing alternative docking mechanisms.
- 6. Have on-board computer capabilities.
- 7. Provides for lighting and camera alternatives.
- 8. Have good maintainability and reliability characteristics.
- 9. Have the necessary communication and control capabilities.
- 10. Permit modification for automated docking maneuvers.

Advantages of a new second-generation unit are:

1. The desired 6th DOF can be incorporated into the design of the new vehicle more easily than into the existing one.

2. Some faults of the present design can be eliminated in the new vehicle design.

3. The reliability of a new vehicle would be much better.

4. Maintenance costs should be lower at least in the early years of its service life.

5. Vehicle balance can be incorporated into the design and operation.

6. Provision for alternative docking mechanisms can be installed.

7. Additional features such as improved lighting and vision can be incorporated.

Disadvantages of a new second-generation unit are:

1. The cost will be considerable. It may be more difficult to obtain the fund appropriation for a new vehicle than to rebuild the present unit with NASA personnel.

2. The procurement lead-time will probably be a minimum of three years between the date of the approval and the date the new vehicle goes into full service.

3. Unforeseen bugs and normal new model difficulties will arise.

Alternative Design Concepts for a Second-Generation Free Flying Mobility Unit

The following alternative design concepts were considered:

1. A vehicle suspended from a mobile overhead track-mounted vehicle.

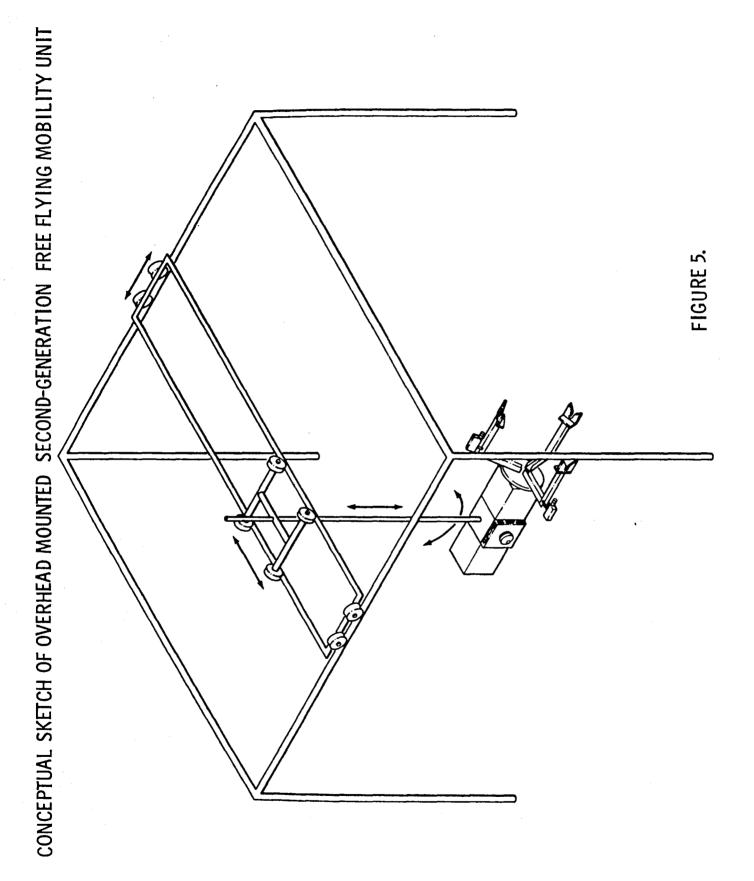
2. A new floor mounted vehicle with its mass concentrated towards the center of gravity.

3. A floor-mounted vehicle with automated computer - controlled balancing.

4. A vehicle mounted on top of a two-way track system above the floor level. (Dr. Campell's concept) This would have some advantages in stability and simpler motion design. However, the mounting might cause interference problems and possible damage (16).

Overhead Track-Mounted Vehicle

The concept of an overhead track vehicle with the docking mechanism suspended underneath is illustrated in Figure 5. The track and carrier will permit locating the mechanism in the X and Y directions. An actuator operated



column or beam would raise or lower the mechanism on the Z axis. Pitch yaw and roll would be provided by movement about the lower column end. Communications could be by direct electrical connection rather than by radio.

The advantages of suspending the docking vehicle from an overhead track are:

1. The sixth degree-of-freedom will be easier to incorporate.

2. Some of the technology is known and well-established.

3. Some components may be completely or partially "off-the-shelf." This will reduce the lead-time and possibilities of trouble with new technology.

4. There will be fewer interference problems such as those with the bases of the two present vehicles.

A major disadvantage is:

1. Maintaining vehicle balance through all DOF movements. This will require either a complex mechanical system or a sophisticated computer control program.

The research and design problems include:

- 1. Adequate building structure which may be difficult to fund.
- 2. Kinematics and deflection.
- 3. Mass and momentum.
- 4. The control system.

Floor-Mounted Free Flying Mobility Unit

A floor-mounted Free Flying Mobility Unit with computer-controlled balancing is a possibility. A sixth degree-of-freedom could be added. This would have the advantages of floor mounting while providing solutions to some major problems with the present model.

The major difficulty with this concept is the design of the control system and the computer programming. Both would be complex and expensive. Further, modifications to the design or latching mechanisms would require extensive redesign, reprogramming and modification.

A second-generation floor-mounted Free Flying Mobility Unit would be similar in some respects to the present model. The design concept is illustrated in Figure 6. A major change is the use of the three actuators to control the plane of the mechanism. This will provide more possibilities of movement and improve the units balance. Some of the other components can be relocated nearer the center of gravity to improve the balance. Other improvements can be incorporated.

CONCEPTUAL SKETCH FLOOR MOUNTED SECOND-GENERATION FREE FLYING MOBILITY UNIT

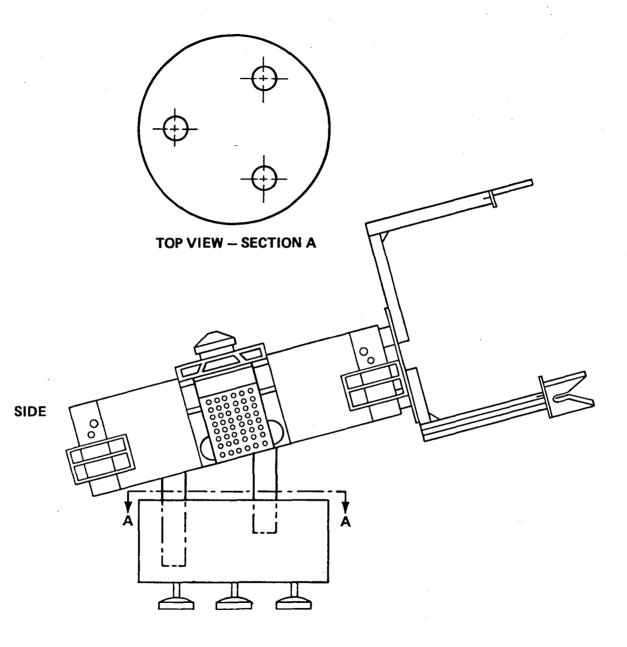


FIGURE 6.

Floor-Mounted Unit with Automated Computer-Controlled Balance

Computer balance control will have the following advantages:

1. Proper design of the balancing system will hold down the mass of the vehicle.

2. The computers can be located remotely from the vehicle. This will hold down the complexity and mass of the vehicle itself.

3. Although trained computer personnel are presently scarce, the number in training who will become available in the near future is increasing substantially.

Major difficulties with Mechanical Balance Control are:

1. Increase in the vehicle mass. This may cause difficulties in operation. The higher mass will increase the impact forces and may also aggrevate some imbalance situations.

2. Personnel capable of designing, building and maintaining mechanical devices are scarce and will become almost unavailable within a few years.

Display System Definition

The present display system for the Free Flying Rendezvous and Docking Simulator needs a thorough review to determine whether it meets current requirements and is optimum for the test facility mission.

FULLY-AUTOMATED ORBITAL SERVICER SYSTEM

The objective of full automation of the Orbital Servicer will require a closely coordinated research and test effort involving several advanced robotics disciplines. The result of successful effort will be the elimination of the manin-the-loop requirement of the present system.

Full automation of the Orbital Servicer Module transfer operation involves complex difficult research issues. It is not enough to extend the present computer controlled trajectory to cover the final movement of the module into the rack and to complete the latching operation. Functions performed by the man-in-loop must be incorporated into the automation. These functions include collision avoidance, path adjustment, coordination of the teleoperator movement and others.

In this discussion, the following assumptions are made:

1. The precision and accuracy, capabilities of the present Orbital Servicer robotic system are inadequate for efficient operation and collision avoidance in the full automation mode.

2. The major configuration features of the present Space Shuttle, Orbital Servicer and TMS will remain substantially unchanged.

Among the manipulator design features which should be reviewed are:

- 1. Mass of the components and the assembly.
- 2. Oscillation and deflection.
- 3. Placement of sensors, controls, motors and similar parts.
- 4. Collision and damage avoidance features.

Application of Basic Robotics Research to the Orbital Servicer System

In the section "Basic Robotics Research Overview," several basic knowledge areas are discussed. In this section, application to the Orbital Servicer System will be covered.

The first major difference between the present partially automated, partially manual mode and the fully automated mode is the need for substantially increased precision and accuracy in the final approach, contact and latching. Extending the present computer control program will not provide enough accuracy to complete the final mating and latching successfully. In order to attain this accuracy, knowledge from several basic fields must be added.

The first addition will be sensors to determine accurately the position of the module with respect to the rack. The predicted accuracy with the present system is about $\pm 1/2$ inch. This is insufficient to insure proper completion of the maneuver.

A necessary function of the sensors will be to detect deviations from the planned or expected movements and locations. The human operator can recognize situations which deviate from plan and take action to abort or adjust the teleoperator arm action. If there is no man-in-the-loop, the teleoperator must be designed to recognize the situation and to take appropriate action.

The first situation the sensors must recognize is misplacement, deviation or absence of the rack, module or other system component. The other is the presence of an obstruction, accidental or not with which the teleoperator would collide if the programmed paths are continued.

At the time of this report (Summer 1980), it is infeasible to specify the type and technical details of the needed sensors. Several types are mentioned in the previous section.

After the sensors have collected data from the environment, image analysis is necessary to interpret the data for action by the control system. Data impulses received by the sensors must be converted into an information format upon which a decision can be based. This has turned out to be one of the most difficult fields in robotics research and application.

The human's ability to analyze and interpret data collected by his sensory organs is unbelievably complex and difficult to replicate. Yet the teleoperator system must perform this function if it is to carry out its mission. After the teleoperator system has received and analyzed the environmental data decisions on the teleoperator action must be made by the system without human intervention. This is the domain of artificial intelligence, another basic field in robotics research.

Communications are essential to the teleoperator system operation. Data on the environment and the system must be transmitted to the analysis and control sub-system. Commands resulting from decisions must be transmitted to the functioning sub-systems and components. Accuracy, reliability, speed and other characteristics are important to mission success and efficient operation.

Control function provides the machine direction and monitors of the of the teleoperator system. Again replicating the process of human control in the fully-automated teleoperator system is an exceedingly complex and difficult task. However, modern control theory research has made considerable progress which will be helpful in this effort.

The kinematics and mechanical design of the teleoperator system and its components will require further research and technology. For example, each additional joint adds degrees of freedom. The number of degrees-of-freedom in the teleoperator exceeds the number which can be analyzed kinematically at the present level of knowledge. Further basic research is needed. Other problems concern⁻⁻ the mass, momentum and impact characteristics of the teleoperator system and its components.

End effector design is an field of robotics research important to the Orbital Servicer Teleoperator system. Improvements are needed in the latching mechanism and in the latching and unlatching operations. One such improvement is computer control of the latching and unlatching as the module moves into and out of the rack.

Locomotion, a basic field of robotics research at some institutions, has little relevance to the current mission of neither the Orbital Servicer System nor the Free Flying Rendezvous and Docking Simulator. However, should the missions change, locomotion might become essential to mission success.

Resolver Feedback

The use of resolver feedback for manipulator joint position sensing should be investigated. The current method is subject to inaccuracies due to nonlinearities in the potentiometers, grounding and noise, and variation in the reference voltages. The very accurate resolver feedback is especially needed for the successful implementation of the current dead-reckoning control system.

ADDITIONAL ORBITAL SERVICER SYSTEM PROBLEMS

In addition to full automation of the Orbital Servicer TMS, there are three subsidiary problems in the assignment.

- 1. Latching mechanism operation.
- 2. Rack attachment.
- 3. Electrical self-aligning connector.

Latching Mechanism Operation

The mechanism for latching and unlatching the modules needs improvement in operation and in coordination with the full automation trajectory mode.

The following constraints must be met:

- 1. Automated operation.
- 2. Positive latching.
- 3. No damage to the module or the mechanism of the manipulator arm.

4. No major modification of the present latching mechanism, teleoperator end effectors, components or the existing control program.

One alternative solution is a mechanical device actuated by contact or visual sensors. As the device takes over motion control, the computer control would be released. The mechanism would complete the latching operation. There are a number of difficult problems with this alternative such as the sensing device, powering the latching device and operational difficulties with the mechanism and teleoperator.

A better solution appears to be computer operation and control of the latching operation from start to completion. The coordination problems will be less than with mixed mode control. Full computer control will produce weight savings over mechanical latching devices. In the future, it will be easier to acquire personnel with the necessary computer and electronic skills than personnel with mechanical design, manufacturing and maintenance skills.

Rack Attachment

The present method of attaching the racks to the satellite and to the Orbital Servicer allows some movement and variation in the rack position. This condition creates little difficulty with the manual mode of completing the trajectory. However, it produces major problems for full automation. Variation in the rack position makes completion of the trajectory by dead reckoning infeasible since this operating mode requires that the final rack position be known within close tolerances. If trajectory completion requires complex systems design using image analysis, artificial intelligence, pattern recognition, and other knowledge areas, variable rack position will add substantially to the difficulty, complexity and expense of full automation.

The constraints on the solution are:

1. No major design change or modification of the module, teleoperator rack or present computer control program.

2. Accuracy of the position must be such that automated teleoperator functioning is feasible.

3. No collision, damage, binding or interference be permitted.

Four Alternate Solutions are:

1. Attach the racks positively and solidly to the Servicer and the Satellite by welding, metal fasteners or similar method.

2. Spring load from one or both sides to center the rack in the desired position.

3. Use artificial vision and feedback with the computer control program to position the module regardless of the location of the rack.

4. Redesign the rack to provide wide entry space tapered to fit the module closely. The best solution appears to be fixed, accurate location of the racks. Full automation will require complex devices and programs to complete the trajectory successfully. Fixed rack positions will hold down the complexity and difficulties of full automation.

Electrical Self-Aligning Connector

Avoiding damage to electrical connectors during transfer of a module between Satellite and Orbiter Servicer is a serious problem. If a pin is bent or the connector is out of alignment, the connection may fail or the connector be bent and damaged.

The proposed solution is an electrical self-aligning connector, NASA Invention Disclosure Case No. MFS - 25211. The plug and receptacle are conical with annular conductive rings spring-loaded for positive contact. This connector does not require precise alignment and is not as liable to damage as the pin-type connector.

It is recommended that the proposed electrical self-aligning connector NASA Invention Disclosure Case No. MFS - 25211 be built and tested. If tests are successful, the design should be incorporated into the TMS and other programs as appropriate.

COMPUTER RESOURCES FOR ROBOTICS

Although MSFC is an acknowledged leader in modern computer facilities, continuous acquisition of new computer technology and upgrading of present computer resources is essential to support the robotics activities.

The robotics and teleoperator efforts should have dedicated computer resources in addition to the general computer resources available at MSFC. The dedicated computers will avoid access delays, provide immediate interface and permit efficient programming and output.

Computers themselves are improving steadily. Storage capacity, operating speed, reliability, flexibility and computing power are steadily increasing while size, weight and cost are going down. There are limits to possible improvements. (The exponential growth curve cannot go on forever.) However, the limits do not appear on the near-time horizon.

Among the dedicated computer facilities that will be needed are:

1. On-board operating and control computers for the Free Flying Mobility Unit in the Free Flying Rendezvous and Docking Simulator. This will provide for maintaining the balance, recording data, and other functions in which the human test subject need not be involved.

2. Graphics software and hardware for both robotics simulators. This may interface with the major computers to gain increased capacity and allow verification through simulation prior to hardware implementation.

3. Full automation of the Orbiter Servicer System will require substantial increase in computer capacity and power in order to perform the necessary command control, image analysis, and other activities.

COMPUTER MODELING

Computer models and simulations can be useful even though they cannot replace full-scale physical mock-ups and simulators. A computer model can be developed, tested and implemented more quickly and less expensively than can a full-scale simulator. The information gained can be useful in design development.

Computer modeling is developing a considerable body of basic knowledge, practical experience, applications and published articles. The topic is included in many current Engineering, Math and Science education programs. The number of professionals with education and training in computer modeling and simulation is steadily increasing. Competence in this area will become increasingly available within NASA and from its contractors.

Computer modeling and simulation does have limits. The validity and reliability of the model output depend upon the extent to which the model represents reality and upon the quality of input data. The right parameters and the correct representations of their behavior must be incorporated in the model. In order to keep the model to manageable size, simplifying assumptions and short cuts are used. These may introduce unacceptable errors. The input data quality is vital to good results. Sometimes there is inadequate communication between the mathematical modeler and the designer-operator of the physical system. The result may be failure to interpret and apply the results properly and to detect anomalies in the model and its output. A major problem in modeling robotics system is the representation of the physical components and the automated sensory - based interpretation of the data. Another is the present method of representing all trajectories as a set of straight-line elements. This requires a large number of steps and is expensive. A third is incorporation of five or more degrees of freedom; most current modules are limited to three (15).

An important application of robotic programming models is the development of "goal domain" programs. This permits the programmer to specify the goal states of the robot while the robot selects the appropriate actions to achieve the desired goal state (17).

Task specification definition needs to be refined and improved. Too great detail in task definition reduces the scope of adaptive programming to meet variations and contingencies. On the other hand, present state of the art in communications, artificial intelligence, computer languages and artificial vision is inadequate to support goal-oriented task definition in NASA's robotics efforts(13).

CONCLUSIONS AND RECOMMENDATIONS

The major conclusions from the research on robotics in the Free Flying Rendezvous and Docking Simulator and the Orbital Servicer System are:

1. These facilities perform an important service in support of NASA's Space Missions.

2. These facilities should be upgraded and improved.

3. MSFC should move to acquire professional competence in research areas basic to robotics research and technology.

The main recommendations are summarized below. Detailed discussion is the body of this report.

1. A second-generation Free Flying Mobility Unit be acquired. This would provide six degrees of freedom and other improvements infeasible with the present model. The present unit might be kept if its role is complementary to the new unit.

2. The Orbital Servicer System be fully automated.

3. MSFC move to acquire competence in areas necessary to robotics teleoperator development. These include artificial intellignece, sensors, image analysis, pattern recognition, control, communication and computer science applied to robotics.

4. Enhanced computer facilities be acquired for the robotics and teleoperator activities. Some can be additions to Center-wide computer resources; others must be dedicated to the robotics and teleoperator work.

5. Procurement and test of the new electrical self-aligning connector be implemented.

6. Other recommendations are contained in the body of the report. These include rack attachment, latching mechanisms and others.

REFERENCES

- Shields, N. Jr., Slaughter, P.H., Bry, R. G., and Henderson, D. E.: "Earth Orbital Teleoperator Systems Evaluation. 1978 Year End Report." Essex Corporation, Contract NAS8-31848, Feb. 1, 1979.
- Fernandez, K. R., "Application of a Computer Controlled Robot to Remote Equipment Maintenance". Accepted for Iascon '80, Cincinnati, OH, Sept. 1980.
- 3. DeRocher, B. et al., "Integrated Orbital Servicing Study Follow-On", Final Progress Review. Martin-Marietta, TRW, Contract NAS8-30820 SA-5, 9, June 1978.
- Sacerdoti, E. D., "Plan Generation and Execution for Robotics". Proceedings NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- Grossman, D. D.; "Robotic Intelligence and Decision Making: Real World Aspects". Proceedings NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- 6. Agin, G. J., "Issues Involving Sensory Control". NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- Nitzan, D., "Assessment of Robotic Sensors". Proceedings NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- 8. Roth, B., "Kinematic Design for Robotics". Proceedings NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- Dubowsky, S., "Dynamics for Manipulation: Areas of Future Research". NSF Workshop for Robotics Research, University of Rhode Island, April 15-17, 1980.
- Bejczy, A. K., "Demands that Robotics Systems Place on Control Theory". NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- Paul, R. P.: Present Research approaches to Control in Robotics Systems. NSF Workshop on Robotics Research, April 15-17, 1980.
- Graupe, D. and Saridis, G. N., "Principles of Intelligent Control for Robotics Prosthetics & Orthotics". Proceedings, NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.

 Ruoff, C. F., "Future Languages needed for Cooperating Robots and Operating Systems for Robots". Proceedings, NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.

х

- 14. Shimano, B., "Levels of Communication and Programming Methods". NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- Wesley, M. A., High Level Languages and Modeling. NSF Workshop on Robotics Research, University of Rhode Island, April 15-17, 1980.
- 16. Campbell, R. A., Unpublished memorandum, MSFC.

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MATERIAL AND PROCESS CONSTRAINTS FOR A FLAT INTERFACE IN THE BRIDGMAN-STOCKBARGER TECH-NIQUE

Prepared By:

Academic Rank:

University and Department

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Larry M. Foster, Ph.D.

Assistant Professor

University of Alabama in Huntsville Mathematics Department

Space Sciences Space Processing Semiconductors and Devices

R. J. Naumann

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MATERIAL AND PROCESS CONSTRAINTS FOR A FLAT INTERFACE IN THE BRIDGMAN-STOCKBARGER TECHNIQUE

By

Larry M. Foster, Ph.D. Assistant Professor of Mathematics University of Alabama in Huntsville Huntsville, Alabama

ABSTRACT

The functional relationships between various material and process parameters necessary for a flat interface in the Bridgman-Stockbarger technique are given. In particular, two boundary condition cases of

(1) $\nabla^2 T = P_{\mathbf{g}} \frac{\partial T}{\partial x}$, x < 0 and 0 < r < 1(2) $\nabla^2 T = P_{\mathbf{g}} \frac{\partial T}{\partial x}$, x > 0 and 0 < r < 1(3) $-K_{\mathbf{g}} \frac{\partial T}{\partial x} \Big|_{\mathbf{x}=0^-} + K_{\mathbf{g}} \frac{\partial T}{\partial \mathbf{x}} \Big|_{\mathbf{x}=0^+} + L = 0$

are considered.

Case 1:
$$T(x,1) = \begin{cases} T_c, x>0 \\ T_h, x<0 \end{cases}$$

Case 2: $\frac{\partial T}{\partial r}(x,1) = \begin{cases} -B_sT, x>0 \\ -B_g(T-T_i), x<0 \end{cases}$

Necessary interval bounds on T_1 and T_h are given assuming a flat melt-solid interface at x = 0. Approximations of T (x, r) and related error estimates are also given.

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NOMENCLATURE

- P liquid Péclet number
- P_s solid Peclet number
- T_c cooler temperature
- T_{μ} heater temperature
- Q material melting point
- K_0 liquid thermal conductivity
- K solid thermal conductivity
- \mathcal{L} crystal growth rate \cdot liquid density \cdot latent heat
- B_s solid region Biot number
- B₉ liquid region Biot number
- T₁ scaled ambient temperature for liquid region
- 'T temperature

INTRODUCTION

The partial differential equations defining cyrstal growth by the Bridgman - Stockbarger technique [2] are

$$\nabla^2 T = P_1 \frac{\partial T}{\partial x}, x < 0$$

and

$$\nabla^2 T = P_s \frac{\partial T}{\partial x}, x > 0$$

where P_{i} and P_{s} are the liquid and solid P_{eclet} numbers respectively, T is temperature and x is distance from a flat solid - melt interface. To conserve energy at the interface,

$$K_{\chi} = \frac{\partial x}{\partial x} \Big|_{x=0^{-}} + K_{s} = \frac{\partial x}{\partial x} \Big|_{x=0^{+}} + \chi = 0$$

where K, and K_s are the respective liquid and solid thermal conductivities and \mathcal{L} is the product of the crystal growth rate, the melt density and the latent heat of solidification. Two heating mechanisms are investigated; first considered is conductive heating defined by

$$T(x,1) = T_{c}, x > 0$$

$$\Gamma(\mathbf{x},1) = \mathbf{T}_{\mathbf{n}} , \mathbf{x} < \mathbf{0}$$

followed by Newton heating

$$\frac{\partial T}{\partial r}(x, 1) = -B_s T, x > 0$$

$$\frac{\partial T}{\partial r}(x, 1) = -B_t (T - T_t), x < 0$$

where T_c and T_{μ} are the respective cooler and heater temperatures, B_s and B_{μ} are the solid and melt Biot numbers respectively, and T_{μ} is the ambient temperature surrounding melt region. The ambient temperature for the solid region has been scaled to 0.

The assumption of a flat melt-solid interface forces K_s , K_s , T_c , T_{μ} , B_s , B_{μ} , T_{μ} , P_s , P_{μ} and Q, the material melting temperature, to be functionally dependent. This dependence, the solutions of the above partial differential equations, and various error estimates are given in the next section followed by several numerical experiments. The final section is devoted to conclusions and recommendations.

SOLUTIONS AND APPROXIMATIONS

The solutions of Problems I-II and their computable approximations are given in this section. Also given are the functional relationships between the material and process parameters P_{g} , P_{s} , K_{t} , K_{s} , T_{c} , T_{μ} , B_{s} , B_{g} , and T_{i} necessary for a solid-melt interface at x = 0. We begin with two elementary but fundamental facts.

Fact 1. If (1) $\{a_m\}$ alternates (2) $|a_{m+1}| \le |a_m|$ (3) $a_m \to 0$ as $m \to \infty$ then (1) $\sum_{i=1}^{m} a_m$ converges (2) $|\sum_{i=1}^{N} a_m - \sum_{i=1}^{m} a_m| \le |a_{N+1}|$

Fact 2. For f(x) sufficiently smooth,

$$f'(x) = \frac{f(x + h) - f(x)}{h} + O(h)$$

Problem I.

We wish to solve, in cylindrical coordinates,

(1)
$$\nabla^2 T = P_1 \frac{\partial T}{\partial x}$$
, $x < 0$, $0 < r < 1$

(2)
$$\nabla^2 T = P_3 \frac{\partial T}{\partial x}$$
, $x > 0$, $0 < r < 1$

subject to

(3)
$$T(x, 1) = T_{c}, x > 0$$

(4)
$$T(x,1) = T_{H}, x < 0$$

(5)
$$-K_{\lambda} \frac{\partial T}{\partial x} + K_{3} \frac{\partial T}{\partial x} + \chi = 0$$

where χ = crystal growth rate • melt density • latent heat

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Condition (5) implies a solid-melt interface at x = 0, i.e.,

(6)
$$T(O, r) = Q$$
.

Unfortunately, conditons (3) - (6) overpose (1) and (2) and generally, overposed problems are impossible to solve. However, in this case, the overposing of (1) and (2) will give constraints on the material and process parameters P_{I} , P_{S} , T_{c} , Q, T_{H} , K_{I} , K_{s} , and $\boldsymbol{\chi}$.

By separation of variables, (1) and (2) with conditions (3), (4) and (6) have solutions [1]

(7) T(x, r) =

$$T_{c} + \sum_{i}^{\infty} \frac{2(Q-T_{c})}{\beta_{m} T_{i}(\beta_{m})} J_{o}(\beta_{m}r) E_{xp} \left[(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{m}^{2}}) \frac{x}{2} \right],$$

$$X > O$$
 and $O < r < |$

$$(8) \qquad T(x,r) =$$

$$T_{H} + \sum_{i}^{\infty} \frac{2(Q-T_{H})}{\beta_{m} J_{i}(\beta_{m})} J_{o}(\beta_{m} r) E_{XP} \left[(P_{J} + \sqrt{P_{J}^{2} + 4\beta_{m}^{2}}) \frac{x}{2} \right],$$

X < 0 and 0 < r < 1

where $\{\beta_m\}$ is the increasing sequence of all positive roots of J. We next show that Q, K, K, T, T, T, X, P, and P, cannot be arbitrary if (5) is to hold. In fact, (5) will imply that Q, K, K, T, T, T, X, P, and P, satisfy a functional inequality.

<u>Remark 1:</u> An intuitive approach to enlisting (5) is to differentiate (7) and (8) with respect to x and insert the resulting series into (5). But this would require $\sum_{i=1}^{\infty} J_{i}^{-1}(\beta_{m})$ to converge, an impossibility. Hence, we will approximate the partial derivatives in (5).

By Fact 2, (5) and (6) imply

(9)
$$-K_{j} \frac{Q - T(-h, 0)}{h} + K_{s} \frac{T(h, 0) - Q}{h}$$

+ $\chi = O(h)$

Fact 3. Let $\{\beta_n\}$ be the increasing sequence of all nonnegative roots of $J_o(x)$. Then

$$\left\{ \begin{array}{c} \frac{1}{\beta_{m} J_{i}(\beta_{m})} \end{array} \right\} \quad \text{alternates in sign,} \quad \left| \frac{1}{\beta_{m+1} J_{i}(\beta_{m+1})} \right| \leq \\ \left| \frac{1}{\beta_{m} J_{i}(\beta_{m})} \right| \quad \text{and} \quad \frac{1}{\beta_{m} J_{i}(\beta_{m})} \longrightarrow 0 \\ \end{array} \right\}$$

Combining Facts 1 and 3 with (7), (8), and (9),

$$\begin{cases} 10 \\ \chi_{H}^{(10)} & = (K_{J} + K_{S})Q + K_{S} \left\{ T_{c} + 2(Q - T_{c}) \sum_{1}^{M} \frac{e^{(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{m}^{2}})\frac{h}{2}}}{\beta_{m} \tau_{i}(\beta_{m})} \right\} \\ & + K_{J} \left\{ T_{H}^{} + 2(Q - T_{H}) \sum_{1}^{M} \frac{e^{(P_{J}^{} + \sqrt{P_{J}^{2} + 4\beta_{m}^{2}})\frac{-h}{2}}}{\beta_{m} \tau_{i}(\beta_{m})} \right\} \\ & = O(h^{2}) + \left[K_{S} 2(Q - T_{c}) \frac{e^{(P_{s}^{} - \sqrt{P_{s}^{2} + 4\beta_{M+i}^{2}})\frac{h}{2}}}{\beta_{M+i} \tau_{i} (\beta_{M+i})} \right] \\ & + \left[K_{J} 2(Q - T_{H}) \frac{e^{(P_{J}^{} + \sqrt{P_{J}^{2} + 4\beta_{M+i}^{2}})\frac{-h}{2}}}{\beta_{M+i} \tau_{i} (\beta_{M+i})} \right]$$

In fact, for computational purposes, if P $_{S}$ \ll 1 and P \ll 1, then the right hand side of (10) can be replaced by

$$O(h^{2}) + 2K_{s}(Q-T_{c}) \frac{e^{(P_{s}-2\beta_{M+1})\frac{h}{2}}}{|\beta_{M+1}J_{1}(\beta_{M+1})|} + 2K_{1}(T_{H}-Q) \frac{e^{(P_{1}+2\beta_{M+1})\frac{-h}{2}}}{|\beta_{M+1}J_{1}(\beta_{M+1})|}$$

(11)

<u>Remark 2:</u> (10) forces \mathcal{X} , K_{g} , K_{g} , Q, T_{g} , T_{H} , P_{g} and P_{g} to be functionally dependent. For computional use, we may set

(12)

$$\begin{aligned}
\mathcal{L}_{h} - (K_{I} + K_{s})Q + K_{s} \left[T_{c} + 2(Q - T_{c}) \sum_{i}^{M} \frac{e^{(P_{s} - \sqrt{P_{s}^{2}} + 4\beta_{m}^{2})\frac{h}{2}}}{\beta_{m} J_{i}(\beta_{m})} \right] \\
+ K_{J} \left[T_{H} + 2(Q - T_{H}) \sum_{i}^{M} \frac{e^{(P_{J} + \sqrt{P_{J}^{2}} + 4\beta_{m}^{2})\frac{-h}{2}}}{\beta_{m} J_{i}(\beta_{m})} \right] \\
= 0
\end{aligned}$$

for h sufficiently small and M sufficiently large. This gives an equation linear in χ , K_s, K_s, Q, T_e, T_H and nonlinear in P_s and P_s. In particular, (12) gives T_H in terms of χ , K_s, Q, T_e, P_s and P_s. Furthermore, if (10) is violated, (1) and (2) cannot have a solution satisfying (3) - (6).

Remark 3: Since
$$f(x + h) - f(x) = f'(x + \frac{h}{2}) + O(h^2)$$
,

it may appear profitable to replace

 $\frac{T(h,0)-Q}{h} \quad \text{and} \quad \frac{Q-T(-h,0)}{h} \quad \text{in (9) by}$

 $\frac{\partial T}{\partial x}(\frac{h}{2},0)$ and $\frac{\partial T}{\partial x}(\frac{-h}{2},0)$ respectively. However,

$$\frac{\partial T}{\partial x}\left(\frac{h}{2},0\right) = 2\left(Q-T_{c}\right)\sum_{i}^{\infty}\frac{P_{s}-\sqrt{P_{s}^{2}+4\beta_{m}^{2}}}{\beta_{m}J_{i}\left(\beta_{m}\right)}e^{\left(P_{s}-\sqrt{P_{s}^{2}+4\beta_{m}^{2}}\right)\frac{h}{4}}$$

converges much slower than the estimates in (10).

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<u>Remark 4:</u> An interval estimate of linear parameters χ , K_{1} , K_{5} , T_{c} and T_{H} can be given. For example, suppose χ , K_{1} , K_{5} , T_{c} , P_{5} , P_{1} and Q are fixed. Then (10) implies

(12)

$$O(h^{2}) - K_{s} 2(Q - T_{c}) \frac{C}{|\beta_{M+1} J_{1}(\beta_{M+1})|}$$

$$- \lambda h + (K_{1} + K_{s})Q + \lambda Q K_{1} \frac{e}{|\beta_{M+1} J_{1}(\beta_{M+1})|}$$

$$-K_{s}\left(T_{c} + 2(Q - T_{c})\sum_{1}^{M} \frac{(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{m}^{2}})\frac{h}{2}}{\beta_{m} J_{i}(\beta_{m})}\right)$$

$$-K_{1} 2Q \sum_{1}^{M} \frac{e^{(P_{1} + \sqrt{P_{1}^{2} + 4\beta_{m}^{2}})\frac{-h}{2}}}{\beta_{m} J_{1}(\beta_{m})} \leq$$

$$K_{1}T_{H}\left(1-2\sum_{i}^{M}\frac{e^{\left(P_{1}^{2}+\sqrt{P_{1}^{2}+4\beta_{m}^{2}}\right)\frac{-h}{2}}}{\beta_{m}J_{i}(\beta_{m})} + 2\frac{(P_{1}^{2}+\sqrt{P_{1}^{2}+4\beta_{M+i}^{2}})\frac{-h}{2}}{\left|\beta_{M+i}J_{i}(\beta_{M+i})\right|}\right)$$

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and
(13)
$$K_{I}T_{H}\left(1-2\sum_{i}^{M}\frac{e^{(P_{I}+\sqrt{P_{I}^{2}+4\beta_{m}^{2}})\frac{-h}{2}}}{\beta_{m}J_{i}(\beta_{m})}-2\frac{(P_{I}+\sqrt{P_{I}^{2}+4\beta_{H+i}^{2}})\frac{-h}{2}}{|\beta_{M+i}J_{i}(\beta_{M+i})|}\right)$$

$$\leq O(h^{2}) + 2K_{s}(Q - T_{c}) - \frac{(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{M+1}^{2}})^{\frac{h}{2}}}{|\beta_{M+1} J_{1}(\beta_{M+1})|}$$

$$-2K_{J}Q - \frac{e^{(P_{g} + \sqrt{P_{g}^{2} + 4\beta_{M+1}^{2}})\frac{-h}{2}}}{|\beta_{M+1} J_{i}(\beta_{M+1})|} - 2h + (K_{g} + K_{s})Q$$

$$-K_{s}\left(T_{c} + 2(Q - T_{c})\sum_{i}^{M} \frac{(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{m}^{2}})\frac{h}{2}}{\beta_{m} J_{i}(\beta_{m})}\right)$$

$$-K_{1} 2Q \sum_{1}^{M} \frac{(P_{1} + \sqrt{P_{1}^{2} + 4\beta_{m}^{2}})^{-h}}{\beta_{m} J_{1} (\beta_{m})}$$

Solving (12) and (13) for $\mathbf{T}_{\boldsymbol{y}}$,

$$T_{1}(M, h) + O(h^{2}) \leq T_{H} \leq T_{2}(M, h) + O(h^{2})$$

For practical purposes, set

(14)
$$T_1(M,h) \le T_H \le T_2(M,h)$$

Since we only know bounds on T₁ (from (14)), for computational use, approximate $T_{H,APP} = \frac{T_2 (M, h) + T_1 (M, h)}{2}$. If T_H is replaced by $T_{H,RPP}$ in (8), then T by

the resulting error is

(15)
$$\left| T_{H} - T_{H, APP} - 2 \sum_{i}^{\infty} \frac{T_{H} - T_{H, APP}}{\beta_{m} J_{i} (\beta_{m})} J_{o} (\beta_{m} r) e^{\left(\frac{P_{A} + \sqrt{P_{i}^{2} + 4\beta_{m}^{2}}\right)\frac{X}{2}}{\beta_{m} J_{i} (\beta_{m})} \right|$$
which has an upper bound of

(16) $(T_{\lambda}(M,h) - T_{i}(M,h)) \left| \frac{1}{\lambda} - \sum_{i}^{m} \frac{J_{o}(\beta_{m}r)}{\beta_{m}J_{i}(\beta_{m})} e^{(P_{\lambda} + VP_{\lambda} + 4\beta_{m}^{-})\frac{1}{\lambda}} \right|$

Problem II.

We wish to solve, in cylindrical coordinates,

(17)
$$\nabla^2 T = \int_{x}^{0} \frac{\partial T}{\partial x}$$
, x < 0 and 0 < r < 1
(18) $\nabla^2 T = \int_{y}^{0} \frac{\partial T}{\partial x}$, X > 0 and 0 < r < 1

subject to

(19)
$$\frac{\partial T}{\partial r}(x, 1) = -B_s T$$
, x >0

(20)
$$\frac{\partial I}{\partial r}(x, 1) = -B_{\lambda}(T-T_{\lambda}), x < 0$$

$$(21) - K_{l} \frac{\partial T}{\partial x} \Big|_{x=0^{-}} + K_{s} \frac{\partial T}{\partial x} \Big|_{x=0^{+}} + \mathcal{L} = 0$$

where \mathcal{L} = crystal rate of growth \cdot liquid density \cdot latent heat, $B_{,} \ge 0$ and $B_s \ge 0$. Condition (21) implies a solid-melt interface at x = 0, i.e.,

(22)
$$T(0, r) = Q$$

As in Problem I, conditions (19) - (22) overpose (17) and (18). This dilemma can be resolved only if the material and process parameters P_{s} , P_{s} , Q , T_{i} , K, , K, and χ are functionally dependent.

Separating variables, (17) and (18) subject to (19), (20) and (22) have solutions [1]

$$T(x, r) = \sum_{j}^{\infty} \frac{2QB_{j}J_{o}(\beta_{m}r)}{(\beta_{m}^{2} + B_{j}^{2})J_{o}(\beta_{m})} \stackrel{(P_{s} - VP_{s}^{2} + 4\beta_{m}^{2})}{(\beta_{m}^{2} + \beta_{s}^{2})J_{o}(\beta_{m})} e^{(P_{s} - VP_{s}^{2} + 4\beta_{m}^{2})\frac{X}{2}}, X > 0$$

(24)

$$T(s, r) = T_{i} + \sum_{1}^{\infty} \frac{2(Q - T_{i}) B_{g} J_{o}(\beta_{m} r)}{(\alpha_{m}^{2} + B_{g}^{2}) J_{o}(\alpha_{m})} e^{(P_{g} + \sqrt{P_{g}^{2}} + 4\beta_{m}^{2})\frac{r}{2}}, x < 0$$

where $\{\alpha_{m}\}$ and $\{\beta_{m}\}$ are the increasing nonnegative sequences of all roots of (25) $B_{s} \ \overline{J}_{o} (\beta_{m}) = \beta_{m} \ \overline{J}_{i} (\beta_{m})$ (26) $B_{s} \ \overline{J}_{o} (\alpha_{m}) = \alpha_{m} \ \overline{J}_{i} (\alpha_{m})$ <u>Fact 4:</u> $\left\{ \left(\beta_{m}^{2} + \beta_{s}^{2}\right) \ \overline{J}_{o} (\beta_{m}) \right\}$ alternates in sign and $\left| \frac{1}{(\beta_{m}^{2} + \beta_{s}^{2}) \ \overline{J}_{o} (\beta_{m})} \right|$ decreases to 0. A similiar result holds for $\left\{ \frac{1}{(\alpha_{m}^{2} + \beta_{s}^{2}) \ \overline{J}_{o} (\alpha_{m})} \right\}$

As in Problem I, (21) and Fact 2 imply

(27)

$$-K_{1} \frac{Q - T(-h, 0)}{h} + K_{s} \frac{T(h, 0) - Q}{h} + L = O(h)$$

Combining Facts 1 and 4 with (23), (24) and (27),

(28)

$$J_{h} - Q(K_{s} + K_{g}) + K_{s} \sum_{i}^{M} \frac{2 B_{s} Q}{(\beta_{m}^{2} + \beta_{s}^{2}) T_{o}(\beta_{m})} e^{(P_{s} - \sqrt{P_{s}^{2} + \gamma \beta_{m}^{2}})\frac{h}{2}}$$

+
$$K_{g}\left\{T_{i} + \sum_{i}^{M} \frac{2(Q-T_{i})B_{g}}{(\alpha_{m}^{2}+B_{g}^{2})J_{o}(\alpha_{m})}e^{\left(P_{g}+\sqrt{P_{g}^{2}+4\alpha_{m}^{2}}\right)\frac{-h}{2}}\right\}$$

$$\leq O(h^{2}) + \left| \frac{2QB_{s}K_{s}}{(\beta_{M+i}^{2} + B_{s}^{2})J_{o}(\beta_{M+i})} e^{(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{M+i}^{2}})\frac{h}{2}} \right|$$

+
$$\left| \frac{2(Q-T_i)B_{g}K_{g}}{(\alpha_{m+1}^{2}+B_{g}^{2})J_{g}(\alpha_{m+1})} e^{\left(P_{g}+\sqrt{P_{g}^{2}+4\alpha_{m+1}^{2}}\right)-\frac{h}{2}} \right|$$

As in Problem I, (28) implies various interval constraints on Q, K_s , K_k , T_1 , P_s , P_k , B_s and B_k . For illustration, consider T_1 . Because we have transformed (in practice) the Newton cooling condition (19) to imply a 0 ambient temperature around the rod for x>0, $0 < Q < T_1$. (28) then implies

(29)

$$O(h^{2}) + Q(K_{s} + K_{g}) - 2QK_{g}B_{g}\sum_{i}^{M} \frac{(P_{g} + \sqrt{P_{g}^{2} + 4\alpha_{m}^{2}})^{-h}}{(\alpha_{m}^{2} + B_{g}^{2})^{-J}\sigma(\alpha_{m})}$$

$$-2K_{s}B_{s}Q\sum_{i}^{M}\frac{e^{(P_{s}-\sqrt{P_{s}^{2}}+4\beta_{m}^{2})\frac{h}{2}}}{(\beta_{m}^{2}+B_{s}^{2})J_{o}(\beta_{m})} -2h$$

+
$$\frac{2B_{s}QK_{s}e^{(P_{s} - \sqrt{P_{s}^{2} + 4\beta_{M+1}^{2}})\frac{h}{2}}{(B_{M+1}^{2} + B_{s}^{2})[J_{o}(\beta_{M+1}^{2})]} - \frac{2B_{l}QK_{l}e^{(P_{l} + \sqrt{P_{l}^{2} + 4\alpha_{M+1}^{2}})\frac{-h}{2}}{(\alpha_{M+1}^{2} + B_{l}^{2})[J_{o}(\alpha_{M+1})]}$$

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$$\geq T_{i} K_{g} \left\{ 1 - 2 B_{g} \sum_{i}^{M} \frac{e^{(P_{g} + \sqrt{P_{g}^{2} + 4 \alpha_{m}^{2}}) \frac{-h}{2}}}{(\alpha_{m}^{2} + B_{g}^{2}) J_{o} (\alpha_{m})} - 2 B_{g} \frac{e^{(P_{g} + \sqrt{P_{g}^{2} + 4 \alpha_{m+1}^{2}}) \frac{-h}{2}}}{(\alpha_{m+1}^{2} + B_{g}^{2}) |J_{o} (\beta_{m+1})|} \right\}$$

and

(30)

$$O(h^{2}) - \frac{2 B_{s} K_{s} Q e}{(\beta_{M+1}^{2} + B_{s}^{2}) | \overline{J}_{o}(\beta_{M+1})|} +$$

$$\frac{2 Q B_{g} K_{g} c}{(\alpha_{M+1}^{2} + B_{g}^{2}) | J_{o}(\beta_{M+1})|} + Q(K_{s} + K_{g})$$

$$-2K_{I}B_{I}Q\sum_{i}^{M}\frac{e^{(P_{I}+\sqrt{P_{I}^{2}+4\alpha_{m}^{2}})^{-\frac{h}{2}}}{(\alpha_{m}^{2}+B_{I}^{2})J_{o}(\alpha_{m})}$$

$$-2K_{s}B_{s}Q\sum_{i}^{M}\frac{\frac{(P_{s}-\sqrt{P_{s}^{2}+4\beta_{m}^{2}})^{\frac{h}{2}}}{(\beta_{m}^{2}+B_{s}^{2})\overline{J_{o}}(\beta_{m})} - Lh$$

$$\leq T_{1}K_{1} \left\{ 1 - 2B_{1}\sum_{i}^{M} \frac{(P_{1} + \sqrt{P_{1}^{2} + 4\alpha_{m}^{2}})^{\frac{-h}{2}}}{(\alpha_{m}^{2} + B_{1}^{2})J_{0}(\alpha_{m})} \right\}$$

+
$$2 B_{g} \frac{(P_{g} + \sqrt{P_{g}^{2} + 4 \alpha_{M+1}^{1}}) - \frac{h}{2}}{(\alpha_{M+1}^{2} + B_{g}^{2}) | \overline{J}_{o}(\alpha_{M+1}) |}$$

Solving (29) and (30) for ${\rm T}_{\rm j}$,

$$T_{4}^{LOW}(M,h) + O(h^{2}) \leq T_{1} \leq T_{1}^{H/GH}(M,h) + O(h^{2})$$

For computational purposes, let

(31)
$$T_{1}^{LOW}(M,h) \leftarrow T_{i} \leftarrow T_{i}^{HIGH}(M,h)$$

(32)

$$T_{l, APPROX} = (T_{l}^{LOW}(M, h) + T_{l}^{HIGH}(M, h)) / 2$$

The error committed by replacing T_{l} in (24) by $T_{l,APPRox}$ is no more than (33)

$$\left(\mathsf{T}_{I}^{HIGH}(\mathsf{M},\mathsf{h})-\mathsf{T}_{I}^{LOW}(\mathsf{M},\mathsf{h})\right)\left|\frac{1}{2}-\sum_{i}^{\infty}\frac{B_{i}J_{o}(\alpha_{m}r)}{\left(\alpha_{m}^{2}+B_{i}^{2}\right)J_{o}(\alpha_{m})}C^{\left(P_{i}+\sqrt{P_{i}^{2}}+4\alpha_{m}^{2}\right)\frac{X}{2}}\right|$$

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<u>Remark 5:</u> It is possible to substitute (23) and (24) into (21) and solve for T_1 . However, this would require computing

(34)

$$\sum_{1}^{\infty} \frac{P_{s} - \sqrt{P_{s}^{2} + 4\beta_{m}^{2}}}{(\beta_{m}^{2} + \beta_{s}^{2}) J_{o}(\beta_{m})}$$

and

(35)

$$\sum_{i}^{\infty} \frac{P_{i} + \sqrt{P_{i}^{2} + 4\alpha_{m}^{2}}}{(\alpha_{m}^{2} + B_{i}^{2}) J_{o}(\alpha_{m})}$$

Although (34) and (35) conditionally converge, the sums in (28) - (30) and (33) converge faster by several orders of magnitude.

<u>Remark 6:</u> Let $B_s = B_l = P_s = P_l = 0 = \chi$ and fix 0 < h << 1.

Letting M = 39, (10) implies $Q = \frac{K_s T_c + K_s T_{\mu}}{K_s + K_s}$

for Problem 1 and (28) gives

$$Q = \frac{K_{I} T_{i}}{K_{S} + K_{I}}$$

for Problem 2.

NUMERICAL EXAMPLES

The results of several numerical experiments are given in this section.

Example 1: In Figure 1 and Table 1 we illustrate the dependence on M of the interval estimate (14). Let $T_c = 400$, Q = 700, $P_s = P_g = 0.001$, $\chi = 0.01$, h = 0.1, $K_s = 0.0032$ and $K_g = 0.002$.

м	Interval Estimate of T _H
9	[1391.34, 1476.13]
14	[1202.19, 1214.26]
19	[1180.56, 1182.74]
24	[1176.34, 1176.54]
29	[1176.34, 1176.54]
34	[1175.92, 1175.93]
39	[1175.90, 1175.90]



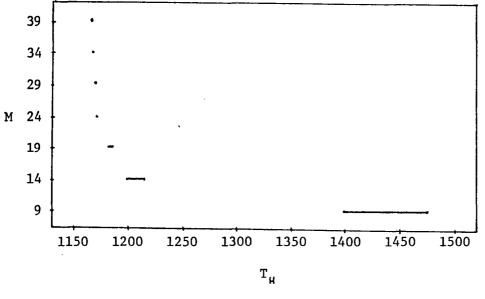


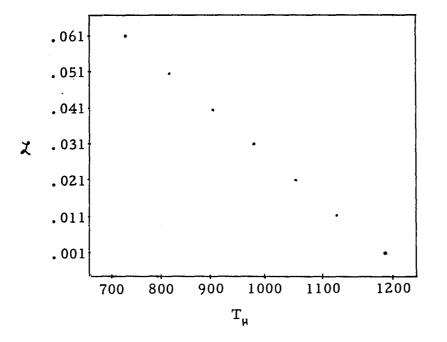
Figure 1

<u>Example 2</u>: In Table 2 and Figure 2 we illustrate the sensitivity of the interval estimate (14) to changes in \mathcal{X} . Let $K_{g} = K_{s} = P_{s} = P_{s} = 0.001$,

h = 0.1, $T_c = 200.0$, Q = 700.0 and M = 39.

Z	Interval Estimate of T _H
.001	[1192.386, 1192.389]
.011	[1116.514, 1116.517] [1040.642, 1040.645]
.031	[964.770, 964.773] [888.898, 888.901]
.051	[813.026, 813.029] [737.154, 737.157]







CONCLUSIONS AND RECOMMENDATIONS

Equations (10) and (28) imply that the material and process parameters cannot be arbitrary if a flat melt-solid interface is to be achieved. The sensitivity of one parameter to changes in another parameter (Example 2) suggests great care should be taken in obtaining the various material parameters.

The relationships between M and h that minimize the interval widths in (14) and (31) remain open questions. Further investigations of Problems I and II could incorporate adiabatic boundary zones near the melt-solid interface and the influences of various container configurations.

REFERENCES

- 1. Carslaw, H. S. and Jaeger, J. C., <u>Conduction of Heat in Solids</u>, 2nd Ed., Oxford Press, 1959.
- 2. Chang, C. E., and Wilcox, W. R., "Control of Interface Shape in the Vertical Bridgman-Stockbarger Technique", Journal of Crystal Growth, Vol. 21, pp. 135-140, 1974.

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

A RANKING ALGORITHM FOR SPACELAB

CREW AND EXPERIMENT SCHEDULING

Prepared by: Robert D. Grone and Frank H. Mathis

Academic Ranks: Assistant Professor

Institution and Department: Auburn University Department of Mathematics

NASA/MSFC:

Division: Mission Analysis

Branch: Mission Integration

MSFC Counterparts: Jerry D. Weiler and William C. Askew

Contract No: NGT-01-002-099 (University of Alabama)

A RANKING ALGORITHM FOR SPACELAB CREW AND EXPERIMENT SCHEDULING

by

Robert D. Grone Assistant Professor Department of Mathematics Auburn University Auburn, Alabama 36849

and

Frank H. Mathis Assistant Professor Department of Mathematics Auburn University Auburn, Alabama 36849

ABSTRACT

This report is concerned with the problem of obtaining an optimal or near optimal schedule for scientific experiments to be performed on Spacelab missions. The design criteria of a method for solving this problem is two-fold. First, the method must produce a schedule which satisfies all material and resource constraints while maximizing a function assigning a grade to each schedule. Secondly, the method should reduce the amount of preflight manpower and computer time currently necessary to achieve this goal and be rapid and flexible enough to allow real-time support of rescheduling problems arising due to launch delays, equipment malfunction, or other factors.

In this paper the current capabilities in this regard are examined and a method of ranking experiments in order of difficulty is developed to support the existing software. Experimental data is obtained from applying this method to the sets of experiments corresponding to Spacelab missions 1, 2 and 3. Finally, suggestions are made concerning desirable modifications and features of second generation software currently being developed for this problem.

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INTRODUCTION

The problem under consideration is that of scheduling a set of experiments to be performed on a given Spacelab mission. Through a preliminary analysis, which is not discussed here, a set of compatible experiments is compiled which are hopefully to be included in a specific mission. The experiments are then converted into model sheets by the principal investigator, who is responsible for the experiment design, and a NASA project engineer familiar with the format and requirements of the model sheets. The model sheet divides each experiment into a sequence of steps, each of which has certain requirements in terms of crew, equipment, and energy usage; and may in addition require that the Spacelab be in a certain position or configuration, or have certain targets available. Examples of targets are planets, communication satellites, and data receiving facilities on the earth's surface. If the experiment is to be performed more than once the number of performances desired is included on the model sheet. A typical Spacelab mission lasts for seven days and involves six crewmembers.

Extremely sophisticated software is available to assist in scheduling of the experiments. The current program is designated by TLP for timeline program. This program takes the model sheets in some order and "front loads" them in this order. By <u>front load</u>, it is meant that the model is scheduled at the earliest possible time which satisfies all the model constraints and which does not conflict with the requirements of previously scheduled models. If it is impossible to schedule a certain model at any time the program then goes to the next model in the sequence. This program is currently being upgraded to a program denoted by ESP for experiment scheduling program, which has the capability of either front loading or back loading any given model, and which is to supersede TLP sometime in 1981.

Even with the extremely advanced and effective software available, the problem of scheduling the experiments in an optimal way currently requires an enormous amount of pre-flight man-hours. By an optimal schedule, we roughly mean one that includes the maximal number of experiments and performances, and which makes maximal use of the available resources. Since each set of models designated for inclusion in a specific mission are only roughly sorted for compatibility, it is frequently impossible to include every performance of every experiment. The method previously employed has been to feed TLP various random orderings of the model sheets and have the project engineers examine the characteristics of the resulting schedules. From these examinations other orderings are suggested and fed to TLP. After months of such trial and error, the best resulting schedule is then modified and edited by hand by a group of NASA personnel familiar with the characteristics of the given mission. The current schedule for Spacelab mission one has been under development for five years. It is clear that such a time expenditure is not only economically prohibitive, but is unacceptable in light of the fact that the Spacelab program eventually envisions flying several missions a year.

OBJECTIVES

The main object of this paper is to provide a method of ranking the model sheets in order of difficulty. By scheduling the most difficult experiments first it is thought that an optimal schedule will result. In support of this idea a grading function is used which assigns a higher grade to schedules which are in some sense better. The problem is then to find a ranking function of the models which provides an ordering which results in a schedule which maximizes the grading function. Since there are 218 model sheets corresponding to Spacelab mission one, an exhaustive search is clearly out of the question.

The remainder of this paper is devoted to the development and examination of such a ranking function. While the method developed did not result in a schedule superior to the schedule hand produced over a five-year period for Spacelab mission one, it did quickly produce schedule for Spacelab missions two and three, which were superior to the best existing schedules for these missions.

These experimental results are described and suggestions are made on the basis of these for possible modifications or additions to the existing software.

THE RANKING ALGORITHM

The problem is now to produce a ranking algorithm which produces an ordering of the models which produces an optimal schedule in terms of a grading function which is described in the results section. While this has certain aspects of a linear programming problem which might be treated with either the simplex method or the new Khachiam method, it is unclear if it is possible to convert this to a linear problem. The method we exhibit is similar to the one described by Garey, Graham, Johnson, and Yao in J. Comb. Theory (A) 21 (1976), 257-298.

We now proceed to describe the method used to rank the models in decreasing order of difficulty. The steps involved are: the construction of a nonnegative real vector corresponding to a single performance of each model, the normalization of these vectors, and the evaluation of the difficulty function on the normalized vectors. The models (vectors) are then arranged in decreasing order of difficulty and this ordering is fed to TLP to produce a schedule.

The vector includes a component corresponding to crew time required. The length of each step is multiplied by the number of crewmen required for the step, and the sum of these terms over all steps in the model is computed to give a figure in crewman-minutes.

The vector includes components for up to 10 resources. On Spacelab mission 1 a total of 4 resources were included: power, high rate multiplex, computer memory, and data transmission. The components for each of these resources were computed by multiplying the length of each step times the level usage of the resource and then summing over all steps.

The vector includes up to 40 components corresponding to various pieces of <u>equipment</u>. As in the case of crew and resources, the number of equipment-minutes is computed by summing the number of minutes a certain piece of equipment is required over all steps of the model. On Spacelab mission 1, a total of 38 pieces of equipment are considered as components of the vector. The vector includes components corresponding to attitude and maneuvers. If a step requires that the Orbiter be performing a certain maneuver the length of the step is multiplied by 9, and if a step requires that no maneuvers are to be performed during its execution the length of the step is multiplied by 1. The sum over all steps of the model then produces the component corresponding to maneuvers. The attitude (Orbiter orientation) component is computed in an analogous fashion.

The vector includes up to 72 components corresponding to <u>targets</u>; Spacelab mission 1 having a total of 50 such components. Once again, a figure in target-minutes is produced for each target and model by summing the number of minutes each target is required by a particular model over all steps of that model.

The vector includes two components referred to as Δ -start and start-time. For the ith model, let M_1 and m_1 denote the maximum and minimum respective times at which the performance of the first step of the model may be initiated. The Δ -start component is

$$\Delta_{i} = L - (M_{i} - m_{i}),$$

where L is the total length of the mission. This component is higher for models which are less flexible with respect to possible starting times. The start-time component is

$$S_i = L - M_i$$
,

and gives higher values to models which must be started earlier in the mission. The start-time figure is included in view of the front-loading nature of TLP which causes the time and resources at the earlier part of the mission to be used up first.

The next step in the ranking algorithms is the <u>normalization</u> of the vectors corresponding to the models. Let v_{ij} denote the jth component of the vector corresponding to the ith model. Hence, if there are m models and n vector components then

$$V = [v_{ij}],$$

is an m-by-n nonnegative real matrix. Spacelab mission 1 corresponds in this fashion to a 218-by-95 matrix. The scaling is obtained by dividing the jth component of each

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vector by the sum of the jth components of all the vectors. That is: we normalize the matrix V to a column-stochastic matrix by dividing each column by its column sum. The targets constitute an exception to this rule. The target time component is normalized by dividing by the total length the target is available during the mission rather than the time the target is required by all the models.

The next step in the ranking algorithm is to evaluate the difficulty function at each model (row of V). The programmer has the option of assigning nonnegative weights; w_j, j=1,...,n, to each of the components of a vector (or alternatively, columns of V). If the ith model has corresponding vector $v_i = (v_{i1}, \ldots, v_{in})$, the <u>difficulty of the ith</u> model is defined by

$$d_{i} = \sqrt{p_{i}} (v_{i1}w_{1} + \ldots + v_{in}w_{n}),$$

where p_i is the number of performances of the ith model requested. The capability of adjusting the weights allows a scheduler to emphasize or de-emphasize certain aspects of the mission deemed more critical or less critical to produce various rankings of the models.

The above described procedure of ranking the models in order is referred to as the <u>model-sort</u> method. A slightly different method referred to as <u>performance-sort</u> was also implemented and tested on SpaceTab missions 1, 2, and 3. The idea behind the performance-sort is to treat each performance of an experiment as a separate model. With the model sort program, all possible performances of a given model are scheduled consecutively. With the performance sort program, each performance of a model corresponds to a vector. For example: SpaceTab 1 lists 935 performances of the 218 models. Hence, when using model-sort, a 218-by-95 matrix is envisioned; and when using performance-sort a 935-by-95 matrix is envisioned. If v_i is the vector previously described which corresponds to a single performance of the ith model, and the ith model requires p_i performances, the difficulty level of the kth performance of the ith model is

 $\sqrt{k} (v_{i1}w_1 + ... + v_{in}w_n); k=1,...,p_i.$

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Another feature of the ranking method is the selection of <u>fixed models</u>. This allows the programmer to specify a certain subset of the models or performances which will be scheduled first. For example, on Spacelab mission 1, such items as "crewman sleep" and "crewman lunch" have model sheets corresponding to them. These models are scheduled first and then the remaining models are ranked using either the model-sort or performance-sort method. On Spacelab mission 1 a total of 16 models were fixed.

Another feature of the algorithm is the <u>model weight</u> capability. This is in some sense dual to the vector normalization in that a row is multiplied by a nonnegative scalar in the model weight feature. This allows the programmer to assign weights to any or all of the models which will change the corresponding vector. This option lets the operator increase the difficulty function of a model which is considered particularly important for political or other reasons not reflected in the difficulty function. The authors used this feature to schedule models which were not included in an otherwise excellent schedule. Successive doubling of the weight of a certain model insures that it will appear higher and higher in the ordering until it is the first model scheduled, if necessary.

RESULTS

We proceed to discuss the results obtained using the previously described methods. The schedule produced due to some choice of sorting technique, weights, fixed models, and model weights is evaluated by means of a grading function which is dependent on four parameters: percentage of crew time, percentage of experiment time, percentage of performances, and percentage of models. The average of these four percentages yields the grade of the corresponding schedule. All four parameters are obtained by comparing amount scheduled to amount requested. In this connection, it should be mentioned that the experiment and crew time requested is the minimum experiment time for experiments whose steps have variable length. Hence, it is possible for the percentage score on experiment and crew time, and the grade, to be over 100. Another fact worth noting is that is possible for a fraction of a model to be performed in some cases. For example, if 3 out of 4 steps in a given model were performed this would be counted The percentage of models reflects the as .75 of a model. extent to which experiment on the mission is performed at least once, and the percentage of performances measures the extent to which all requested experiments (including multiplicities) were scheduled.

For each of Spacelab missions 1, 2, and 3 we include tabulated data comparing the effects of varying sorting technique, weights, fixed models, and model weights. For each mission we include for comparison the best schedule which had been previously produced by the personnel at Marshall Space Flight Center. Their method was to experiment with various orderings, feed them to TLP, obtain some reasonably good schedules, and then to invoke software referred to as TLE (timeline editor) to then shift and fit experiments. The TLE affords the capability of removing and rescheduling a single performance from an existing schedule.

Spacelab mission 1 has undergone the longest development time, and we were unable to exceed the best previous score by our methods. Mission 1 is particularly complicated due to the fact that the set of models chosen for inclusion were selected more for political considerations of being included on the first mission than for compatibility of the various models. It is thought that subsequent missions will be designed around a particular theme such as atmospheric research, materials processing, space physics, biological sciences, or astronomy. On Spacelab mission 1 it was found, somewhat unexpectedly, that the model sort technique worked best. On Spacelab missions 2 and 3, the performance sort worked best and produced schedules superior to the previous best. The overall best results were obtained by using the weights which had produced the schedule with maximal score and then increasing the model weights of those models which failed to schedule. On these missions every model was scheduled for at least one performance. The following three tables summarize the results obtained on these three missions, and compares these to the previous best schedule.

SPACELAB 1

MODELS - 218 -- PERFORMANCES - 935

MIN. EXP. TIME - 4869.71 HRS. -- MIN. CREW TIME - 800.62 HRS.

WEIGHTING	TYPE OF SORT	MODELS	PERFORMANCES	EXP, TIME	CREW TIME	GRADE
ALL EQUAL	MODEL	183.69	858.61	4 <u>9</u> 39.35	756.25	92 . 99
ALL EQUAL	PERFORMANCE	184.51	851.22	4847.83	757.56	92.46
TARGETS = 10, Δ -START = 1	MODEL	193.90	884.39	4664.38	762.87	93.64
PREVIOUS BEST		203,03	897.94	4866,52	751,04	95.73
			(1,1,1,1,1,1,1,1,1,1,1,1,1,1,1,1,1,1,1,			

.

SPACELAB 2

MODELS - 168 -- PERFORMANCES - 1992

MIN. EXP. TIME - 2887.64 HRS. -- MIN. CREW TIME - 647.44 HRS

WEIGHTING	TYPE O	FSORT	MODELS	PERFORMANCES	EXP. TIME	CREW TIME	GRADE
ALL EQUAL EXCEPT ∆-START = 0	MODEL	1	154.54	1931.09	3804.08	682.90	106.54
	PERFOR	MANCE	159.50	1971.09	3813.67	693.83	108.28
CREW, RESOUR TARGETS, WINDOW SIZE ONLY. (EQUIPMENT AND START TIME = 0)	ES, PERFOR	MANCE]	162.15	1974.09	3820.48	700.44	109.03
ABOVE WITH MODEL WEIGH ADJUSTED	PERFOR	MANCE	L64.15	1 <u>9</u> 76.09	3821.98	702.48	109.44
PREVIOUS BE	ST	1	L61.82	1976.65	3808.10	658.08	107.27
					`		

SPACELAB 3

MODELS - 61 -- PERFORMANCES - 592

.

MIN. EXP. TIME - 2507.48 HRS. -- MIN. CREW TIME - 569.02

· · · · · · · · · · · · · · · · · · ·						
WEIGHTING	TYPE OF SORT	MODELS	PERFORMANCES	EXP. TIME	CREW TIME	GRADE
ALL EQUAL	MODEL	57.03	519.00	2476.10	562.52	94.69
ALL EQUAL	PERFORMANCE	57.70	519.00	2480.03	562.55	<u>9</u> 5.00
TARGET AND Δ-START ONLY	PERFORAMCNE	58.90	520.00	2478.37	565.72	95.66
ABOVE WITH MODEL WEIGHTS ADJUSTED	PERFORMANCE	59.77	521.00	2480.97	568.75	96.22
PREVIOUS BEST		57.03	519.00	2484.29	565.94 `	94.92
	-					

CONCLUSIONS

In this section, we indicate some avenues of approach suggested by our investigations. These suggestions arose largely on an intuitive basis, and are in no particular order.

It is certain that experimentation with weights is necessary to implement this algorithm. It appeared that no one weighting would produce the best results for all Spacelab missions. Perhaps a scheme could be developed to give an approximation to the optimal weighting by considering the resource profiles of the specific mission.

The difficulty function could be modified. The difficulty function discussed in the body of this paper was motivated by the sum norm on nonnegative vectors. That is, if $v = (v_1, \ldots, v_n)$ is a nonnegative real vector, the sum norm of v is defined via

$$v_1^{+\dots+v_n}$$

We also experimented with a grading function related to the supremum norm

The results obtained were equal or slightly less than the results described in the previous section. It is rather remarkable that a greater difference was not observed. Further experimentation is possible. Perhaps the Euclidean norm might be investigated.

The grading function could be modified. For example, weights could be assigned to the four parameters involved. Also, the grading function does not take into consideration the evenness of the scheduling distribution. It would be better, for example, if the crew were not completely utilized during the first half of the mission and have a lot of unused time during the second half. Perhaps some mechanism might be devised to give preference to schedules whose resource utilization levels are held fairly constant over the duration of the mission.

An examination of what we shall term a dynamic sort scheme was done. In the method described in this paper no consideration is taken of the fact that a model may become more difficult to schedule if models preceding it deplete the resources it requires. We implemented the following iterative technique to circumvent this phenomena. A re-examination of the normalization procedure employed shows that most components of the vector were normalized by the total amount required by all the models in that component. The target components were normalized by the total available target time rather than the total target time requested. We also experimented with normalizing all components by the amount available corresponding to each component. With this normalization scheme we implemented the following dynamic sort option. In this scheme we would find the most dificult model and then re-normalize the remaining vectors by the amounts remaining of the various resources, etc. This iterative technique has the effect of dynamically increasing the weighting factors on those components which are being depleted fastest. We then assigned various weighting factors and computed the corresponding orderings with and without the dynamic sort The ordering produced with the dynamic sort option option. consistently produced a schedule inferior to the one produced without using the dynamic sort. Even in view of these results, we feel that investigation along these lines might yield superior techniques.

All of the previous discussions have focused on possible ways to improve the method of ordering the models to feed to TLP. Our final suggestion arises from limitations imposed by the front-loading nature of TLP, and involve factors internal to TLP rather than external. With the front-loading scheme. a model is scheduled at the earliest possible time. This results in a schedule which has no flexibility during the first part of the mission. It is fairly certain that this period in the mission is very likely to be plagued with correction and adjustment factors, and perhaps space sickness on the part of the crew. Also, the pieces of equipment are scheduled to begin operation at maximum capacity and then level off. This makes the execution of the mission highly dependent on the Spacelab being able to operate at full capacity immediately. The implementation of a front- or back-loading option on ESP is largely motivated by a desire to avoid this type of scheduling. It would be advantageous to schedule each successive model at the "best" possible time rather than the earliest or latest. "Best" might

be defined by scheduling the model at the time during the mission which keeps the overall resource-available profile the flattest over the duration of the mission. If "best" turns out to be impossible to define, it is possible that even a random choice among possible scheduling times for a model in the sequence might result in a superior scheduling method.

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Marshall Space Flight Center

The University of Alabama

Continuous Flow Electrophoresis:

The Effect of Sample Concentration on Throughput and Resolution in an Upward Flowing System

Prepared By:

Academic Rank:

University and Department:

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No:

Thomas S. Jandebeur, Ph.D.

Assistant Professor

Athens State College Department of Biology

Space Sciences Space Processing Separation Processes

R. S. Snyder, Ph.D.

August 1, 1980

NGT-01-002-099 (University of Alabama) Continuous Flow Electrophoresis:

The Effect of Sample Concentration on Throughput and Resolution in an Upward Flowing System

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Thomas S. Jandebeur, Ph.D. Assistant Professor of Biology Athens State College Athens, Alabama

ABSTRACT

The principle of continuous flow electrophoresis is as follows: a suspension of sample material is introduced as a fine, steady stream into a vertical, flowing "curtain" of electrolyte. A horizontal d.c. electrical field is applied to the curtain. Where a mixture of particles/cells having different electrophoretic mobilities is analyzed, each type particle/cell will assume a different, collectable band position in the curtain electrolyte.

The fundamental behavior of the continuous electrophoretic process in the Beckman Continuous Particle Electrophoresis (CPE) System, with flow in the downward direction, has recently been investigated. The purpose of this research is to determine the effect of sample concentration on throughput and resolution in the CPE System with flow in an upward direction. To maintain upward flow of the curtain buffer and sample stream, and to collect same, certain modifications to the CPE System were made; these also are described.

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Thank you, one and all.

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INTRODUCTION

As developed by Strickler⁵, Hannig² and their coworkers, continuous flow electrophoresis is a technique for separating small amounts of biological material. In this process, a suspension of sample is introduced as a fine, steady stream into a vertical, flowing "curtain" of electrolyte. A horizontal d.c. field is then applied to the curtain. Where a mixture of particles/ cells having different electrophoretic mobilities is analyzed, each type particle/cell will assume a different, collectable band position in the curtain electrolyte/buffer.

The fluid containing the particles/cells to be separated moves through a thin chamber. The depth of the chamber is dictated by a need to suppress temperature gradients within the chamber and by a need to reduce chamber wall effects, both of which generate undesirable flow characteristics. As a result of these considerations, so-called "narrow gap" (0.5 mm - 3.0 mm) chambers lack sufficient sample capacity and/or resolving power, particularly with respect to separating cell populations.

OBJECTIVES

Though it is unlikely to improve resolution, increasing particle/ cell concentration in the sample should increase sample throughput. However, higher cell concentrations lead to a density mismatch between sample stream and curtain buffer. Small amounts of sedimenting sample, diluted by the curtain buffer subsequently are rapidly displaced within the curtain. This results in broadening of the sample stream and ultimately in poorer resolution at higher sample concentrations.

Tests conducted by McDonnell-Douglas⁶ demonstrated upflow machines to be much more sensitive to the described density mismatch than downflow machines, further limiting sample stream concentration, resolution and throughput of material.

The objective of this study, then, is to determine the limits to throughput and resolution effected by sample concentration in the Beckman Continuous Particle Electrophoresis (CPE) System with flow in the upward direction. The data may then be compared to similar data exacted from the same CPE System, but with flow in the downward direction.

To maintain upward flow of the curtain buffer and sample stream, and to collect the sample, certain modifications to the CPE System were necessary; these will be described.

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MATERIALS AND METHODS

To regulate and maintain upward flow of the curtain buffer and sample stream, a 0-140 volt output powerstat was used in tandem with a variable speed fluid pump. Dilute R-1 buffer' cooled to approximately 4°C was pumped through the curtain flow meter at approximately 28cc/min. with the powerstat set at 90 volts.

Sample injection was accomplished using a Sage Instruments Model 355 Syringe Pump. The sample, contained in a 3 cc syringe fitted with a 21 gauge needle and connected to the sample inlet tube by a short piece of PVC tubing, was injected into the electrophoretic cell at an approximate rate of 1 ml/hr (range 1/100 at 24.3% flow). To prevent sedimentation of the sample contained in the syringe, a small stirring bar was placed within the syringe. A magnetic stirrer supported by the syringe pump housing was then inverted over the syringe.

A prototype of the collection system depicted by Figures 5 and 6 facilitated sample collection. The CPE System collecting tubes were connected to the collection bench by 10 inch lengths of 17 gauge PVC tubing fitted with 18 gauge needles at the tubing's distal end. The needle-containing tubes were seated in the 48 hole array of the collection bench. A 48 testtube fraction collector was placed below the collection bench to receive the eluted sample.

The d.c. field was generated by a Hewlett Packard Model 712B power supply; 65 volts/cm constituted the field for each experimental separation.

Concentrated R-1 buffer was used in the electrode rinse and coolant loops. Dilute R-1 buffer (800 ml R-1 concentrate added to 1.2 liters of deionized water) was used as curtain buffer.

To establish the effect of concentration on throughput and resolution in this CPE system, fixed turkey and cow red blood cells were used. These cells are stable, well-characterized standards, having different electrophoretic mobilities; further, they are morphologically distinguishable⁴.

DATA ANALYSIS

Two data sets were collected; the first data set consisting of seven "runs" at sample concentrations ranging from 5×10^7 cells/ml to 2×10^9 cells/ml (Table 1; Figure 3); the second data set consisting of eight runs at the aforementioned seven concentrations plus one run at a sample concentration of 1.5×10^9 cells/ ml. A 5-minute stabilizing flow followed by 10 minutes of sample collection without (zero) and with (field) voltage applied constitutes a "run". Aliquots of the collected sample were examined microscopically; recovered cells were counted using an Improved Neubauer hemocytometer. Percent recoveries for each cell type and for the combined cells were then determined (Table 1). Based on the total number of recovered cells of each type and of the combined cells, percent relative abundances of the recovered cells were determined. This data, plotted for both data sets, is presented in Figures 1 and 2 as percent recovery of fixed cells vs. concentration, and in Figures 3 and 4 as percent relative abundance of recovered fixed cells vs. tube number.

DISCUSSION OF RESULTS

Examination of Table 1 (1st data set) and Figure 1 reveals that, in general, more cells are recovered at lower concentrations than are recovered at higher concentrations, either with (field) or without (zero) voltage applied. This may be explained as a sedimentation effect; i.e., higher cell concentrations engender a density mismatch between sample stream and curtain buffer. Small amounts of sedimenting sample, diluted by the curtain buffer subsequently are rapidly displaced within the curtain. This results in broadening of the sample stream, "spreading" of cells over a broader range of sample collection test tubes (see higher concentrations Tables 2 and 3 and higher concentrations Figures 3 and 4) and ultimately to poorer sample resolution. This interpretation is supported by the lowered recovery of the larger, heavier turkey cells at higher concentrations (Figure 1) and by the enhanced recovery of the smaller, lighter cow cells at higher concentrations (also Figure 1). In part, the apparent increased recovery of cow cells at higher concentrations is, perhaps, due to examining and counting a greater number of sample collection test tubes. In general, the data of Table 1 (2nd data set) and Figure 2 also fit this interpretation, but with notable exception: the turkey cells do not sediment at higher concentrations. It is, perhaps, significant that the electrophoretic cell was coated with 1% bovine serum albumin, the electrode membranes and electrode rinse replaced, and the buffer temperatures lowered to approximately 4°C between collection of the first and second data sets. These measures, taken to improve CPE chamber performance, may have inadvertently altered the previously noted turkey cell sedimentation behavior.

Tables 2 and 3 and Figures 3 and 4 reveal "spreading" of the sample over a widening range of sample collection test tubes with increasing concentration. This is especially evident at the higher concentrations $(1.5 \times 10^9 \text{ and } 2 \times 10^9 \text{ cells/ml})$ at which the sample stream was observed to "fall back" on the inlet tube; periodically, "clouds" of the sample billowed upward in the buffer flow apparently being transported more rapidly than cells remaining in the sample stream.

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CONCLUSIONS

With particular reference to Table 3 and Figure 4, data collected <u>after</u> treating the electrophoretic cell as described in Discussion of Results, and with voltage applied, it appears that maximum resolution in the Beckman CPE System, with flow in an upward direction, is achieved at concentrations ranging from $2 \times 10^{\circ}$ cells/ml to $8 \times 10^{\circ}$ cells/ml. The widest peak separation is at $2 \times 10^{\circ}$ cells/ml; however, the sharpest peaks and least overlap between cell populations is at $8 \times 10^{\circ}$ cells/ml.

Apparently as a result of improved electrophoresis cell performance due to coating the chamber with bovine serum albumin, changing the electrode membranes and rinse, and lowering buffer temperatures, sedimentation effects attending to higher concentrations are diminished.

As measured by recovery of fixed cells, throughput is greatest at lower sample concentrations (ranging from 5×10^7 to 2×10^8 cells/ml in this study), diminished at moderate concentrations (4 x 10⁸ to 1 x 10⁹ cells/ml) and increasing at higher concentrations (1.5 x 10⁹ to 2 x 10⁹ cells/ml) although not increasing to the higher recovery levels noted for lower sample concentrations.

Throughput as measured by recovery of fixed cells is, then, diminished at the concentrations judged most likely to yield satisfactory resolution. The "tradeoff" appears to be as follows: improved recovery/throughput at the expense of resolution.

REFERENCES

- 1. <u>Continuous Particle Electrophoresis System</u>, Beckman Instruments, Inc., Scientific Instruments Division, Fullerton, California; October 1967.
- Hannig, K., "Preparative Electrophoresis," in <u>Electrophoresis</u>: Theory, <u>Methods and Applications</u>, Vol. II, Ed., M. Bier, Academic Press (1967).
- Saville, D.A. and Ostrach, S., "The Fluid Mechanics of Continuous Flow Electrophoresis," Final Report on Contract NASA-31399, Department of Chemical Engineering, Princeton University (1978).
- 4. Seamon, G.V.F., "Electrophoretic Separator Experiments," Final Report on Contract NASA-32609, University of Oregon Health Sciences Center (1979).
- Strickler, A., "Continuous Particle Electrophoresis: A New Analytical and Preparative Capability," Sep. Sci., 2(3):335(1967).
- Weiss, R.A., "Investigation of the Free Flow Electrophoretic Process", Final Report on Contract NASA-33200, McDonnell-Douglas Astronautics Co., St. Louis, Missouri (1979).

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			Field		Zero				
Concentration	Data		% Recovery			% Recovery			
(Cells/ML)	Set	Comb.	Turkey	Cow	Comb.	Turkey	Cow		
5×10^{7}	1	87.22	77.4	98.04	92.88	108.36	77.4		
	2	76.38	64.06	89.7	98.56	118.28	78.84		
1×10^{8}	1	65.79	69.66	61.92	74.82	98.04	51.6		
	2	68.99	81.32	56.67	70.22	86.24	54.2		
2×10^8	1	74.18	83.84	64.5	81.27	81.28	81.28		
	2	69.61	80.08	59.14	53.59	46.82	60.36		
4×10^8	1	55.15	57.4	52.88	53.86	50.96	56.76		
	2	43.12	35.72	50.52	43.12	37.58	48.66		
8 x 10 ⁸	1	48.86	45.8	51.92	42.73	42.24	43.22		
	2	48.05	53.9	42.2	30.18	32.64	27.72		
1×10^9	1	50.31	37.66	62.96	39.22	28.9	49.44		
	2	29.81	15.28	44.36	48.17	28.08	68.26		
1.5 x 10 ⁹	2	54.7	71.95	37.45	27.6	31.7	23.48		
2 x 10 ⁹	1	53.66	26.44	80.88	41.34	29.02	53.66		
Ĩ	2	58.27	60.98	55.56	54.45	53.1	55.8		

TABLE 1.RECOVERY OF FIXED CELLS AT INDICATED CONCENTRATION WITH (FIELD) AND
WITHOUT (ZERO) VOLTAGE APPLIED FOR FIRST (1) AND SECOND (2) DATA SETS

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TABLE 2.	PERCENT RELATIVE ABUNDANCE OF RECOVERED CELLS BY TUBE NUMBER AT INDICATED CONCENTRATION (CELLS/ML.) WI	TH
	(FIELD) AND WITHOUT (ZERO) VOLTAGE APPLIED FIRST DATA SET	

Tube Turkey Cow Turkey Cow Turkey	k 10 ⁸	4 x	10 ⁸
	Cow	Turkey	Cow
# Zero Field Zero Field Zero Field Zero Field Zero Field	Zero Field	Zero Field	Zero Field
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c c} & 0 \\ 2.00 \\ 3.08 \\ 0 \\ 29.23 \\ 24.62 \\ 15.38 \\ 0 \\ 31.75 \\ 56.92 \\ 9.23 \\ 0 \\ \end{array}$	0 31.46 61.80 6.74 0 1.27 35.44 58.23 5.06 0	$\begin{array}{c} & 0 \\ 1.22 \\ 4.88 \\ 51.22 \\ 35.37 \\ 0 \\ 7.32 \\ 35.23 \\ 60.23 \\ 4.54 \\ 0 \end{array}$

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TABLE 2: PERCENT RELATIVE ABUNDANCE OF RECOVERED CELLS BY TUBE NUMBER AT INDICATED CONCENTRATION (CELLS/M.L.) (Cont'd) WITH (FIELD) AND WITHOUT (ZERO) VOLTAGE APPLIED FIRST DATA SET

	1	8 x	108		1	1 x :	109		1.5	x 10 ⁹		2 x 1	.09	
Tube		key		W	Tur	key	Co		Turkey	Cow		key	Cov	
Į.₫	Zero	Field	Zero	Field	Zero	Field	Zero	Field	Zero Field	Zero Field	Zero	Field	Zero	Field
14 15 16 17 18 19 20 21 22 23 24 25 26 27 28 29 30 31 32 33 45 36 37 38 39 40	0 1.53 12.21 61.83 24.43 0	0 8.45 42.96 42.25 6.33 0	0 8.21 63.43 28.36 0	0 0.62 0.62 6.83 55.28 36.02 0	0 0.89 0.89 31.25 64.29 2.68 0	0 2.74 13.70 31.51 30.82 14.38 6.16 0.68 0		0 0.82 14.34 31.97 32.79 16.80 3.28 0	Did not collect this concentration first data set	Did not collect this concentration first data set	$\begin{array}{c} 0\\ 0.44\\ 0\\ 1.33\\ 1.33\\ 0.89\\ 0.44\\ 0.44\\ 0\\ 0\\ 0\\ 0.44\\ 1.33\\ 0.89\\ 2.67\\ 9.78\\ 31.56\\ 35.11\\ 10.22\\ 2.22\\ 0.44\\ 0\\ 0.44\\ 0\\ 0.44\\ 0\\ 0\end{array}$	0 0.98 0.49 0.49 0.49 1.46 5.85 10.73 29.27 20.98 17.56 7.32 2.93 0.98 0	$\begin{array}{c} 0\\ 0.24\\ 0.48\\ 1.20\\ 0.48\\ 1.20\\ 0.96\\ 0.24\\ 0.24\\ 0\\ 0.96\\ 0.72\\ 0\\ 1.44\\ 3.36\\ 6.49\\ 30.05\\ 34.13\\ 11.54\\ 2.64\\ 2.16\\ 0\\ 0.24\\ 1.20\\ -\end{array}$	0.32 0.16 0.96 7.18 25.52 24.88 14.20 1.91 0.32 0 0.16 0

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 Table 3.
 PERCENT RELATIVE ABUNDANCE OF RECOVERED CELLS BY TUBE NUMBER AT INDICATED

 CONCENTRATION (CELLS/ML.) WITH (FIELD) AND WITHOUT (ZERO) VOLTAGE APPLIED:

 SECOND DATA SET.

			5 x 10 ⁷				x 10 ⁸			2	x 10 ⁸			4	x 10 ⁸	
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27		46.15		0	İ	42.42		8.70		21.54		2.08		36.21		0
28		23.08		61, 11		36.36		17.39		18.46		31.25		48.28		3.66
29		0		27.78		9.09		34.78		0		47.92		13.79		39.02
30				11.11		3.03		39.13				16.67		0		46.34
81			0	0	5.71		18.18		0	Į	0	2.08				10.98
2	0	Į	4.17		45.71	1	50.00		7.89		6.12	-	0		1.27	0
3	56.25		50.00		42.86		31.82		36.84		18.37		16.40		16.46	
4	37,5	· .	45.83		2.86		0	· ·	36.84		53.06		70.49		59.49	
85 86	6.25		0		0	ł		1	18.42		22.45		13.11		22.78	
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		4.72 83.03 - 12.26	2.29 14.86 62.86 15.43 4.57 0	0 8.89 68.89 - 20.00 2.22	5.84 65.69 22.63 5.11 0 0.73 -	0 1.75 22.81 62.28 12.28 0.88	11. 29 51. 62 24. 19 12. 90 0	2.89 26.35 57.04 12.64 1.08	0 0.56 0.25 27.53 35.39 32.58 1.12 0.56 0	0 10.35 1.04 0.52 0 56.48 26.42 1.04 2.07 2.07	0.68 0.46 0 0.91 2.28 5.48 20.78 26.94 18.95 17.35 4.57 1.37 0 0.23 0	0 3.50 27.27 11.89 3.50 37.76 13.29 2.10 0 0.70	0.88 0.44 0 1.75 11.84 23.25 26.75 29.39 3.07 2.19 0 0.44 0	0.23 0.23 3.02 10.90 25.99 34.34 19.95 5.10	0.20 0.81 0.40 0.81 1.62 4.24 15.35 23.03 45.66 7.68 0.20 0	0.44 2.43 12.58 25.39 30.46 22.52 6.18	0.22 1.33 12.42 45.23 29.49 9.76 1.55

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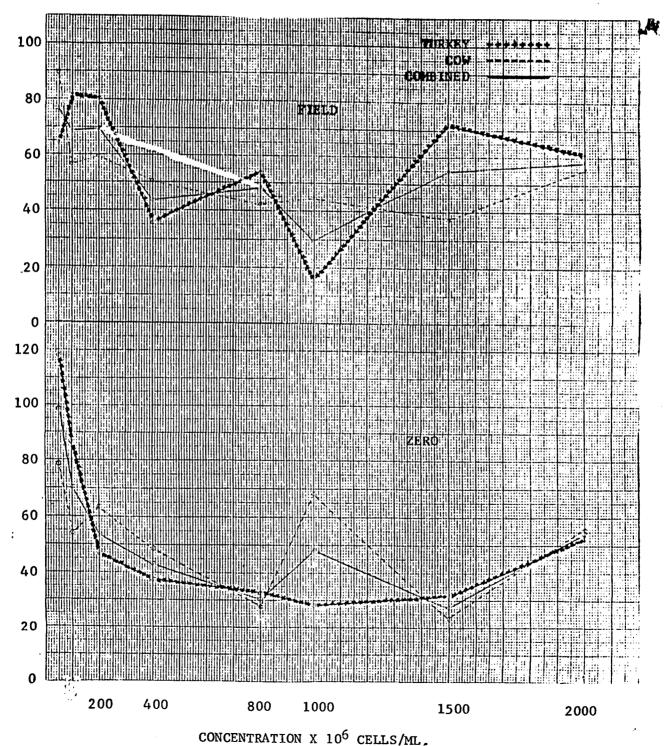
PERCENT RECOVERY.

Figure 1. Percent Recovery of Fixed Cells vs. Concentration with (Field) and without (Zero) Voltage Applied: First Data Set.



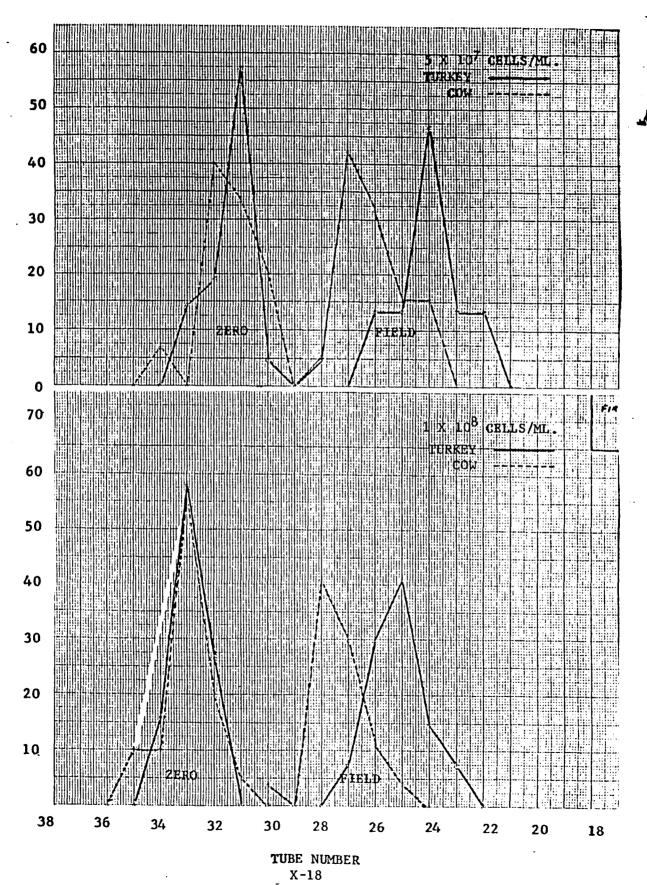
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Percent Recovery of Fixed Cells vs. Concentration with (Field) and without (Zero) Voltage Applied: Second Data Set. Figure 2.



PERCENT RECOVERY

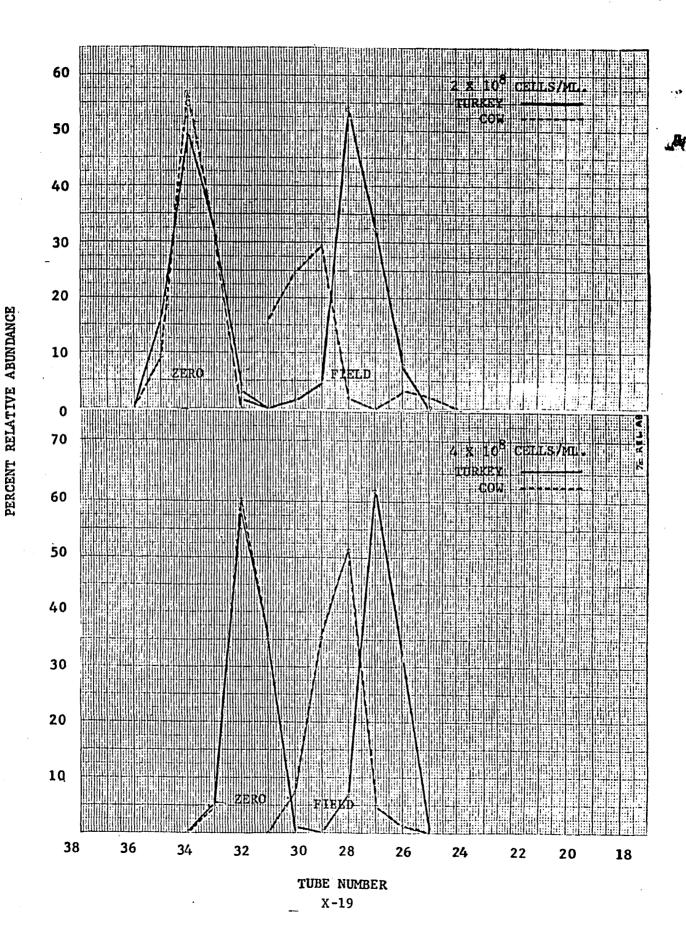
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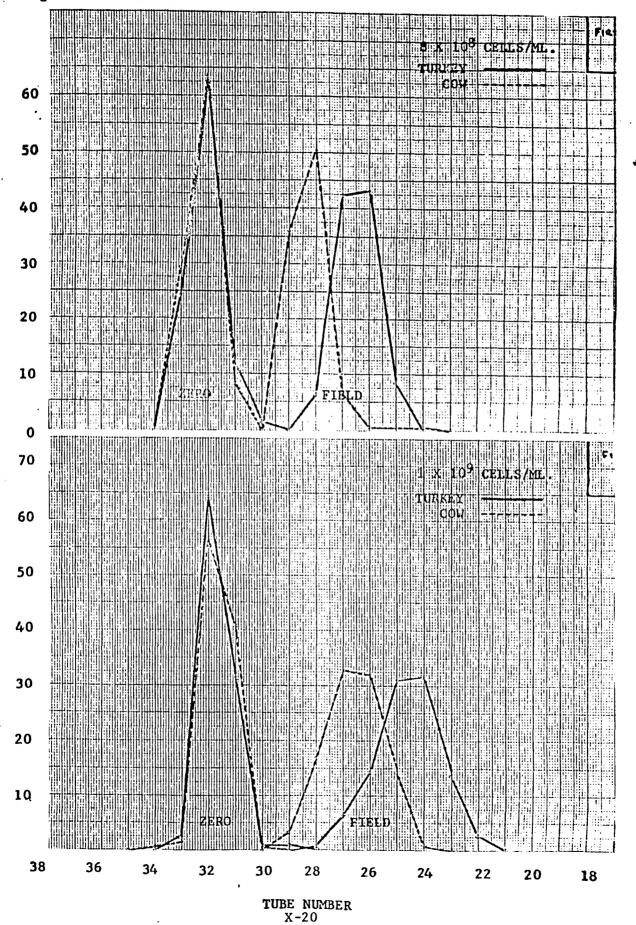
Percent Relative Abundance of Recovered Fixed Cells vs. Figure 3. Tube Number at Indicated Concentration With (Field) and Without (Zero) Voltage Applied: First Data Set.

PERCENT RELATIVE ABUNDANCE





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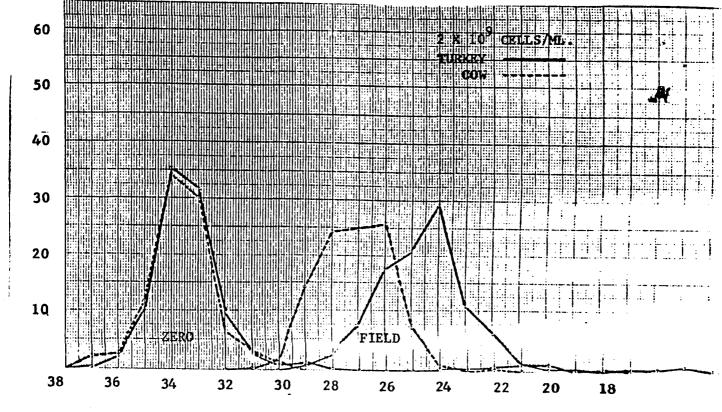
PERCENT RELATIVE ABUNDANCE

Figure 3. Continued

Figure 3. Continued

PERCENT RELATIVE ABUNDANCE

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TUBE NUMBER

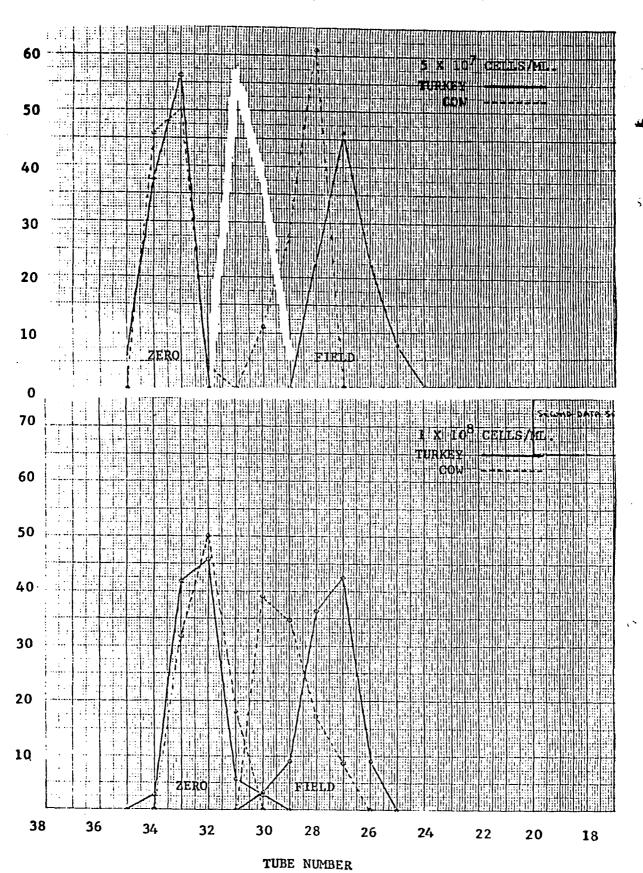
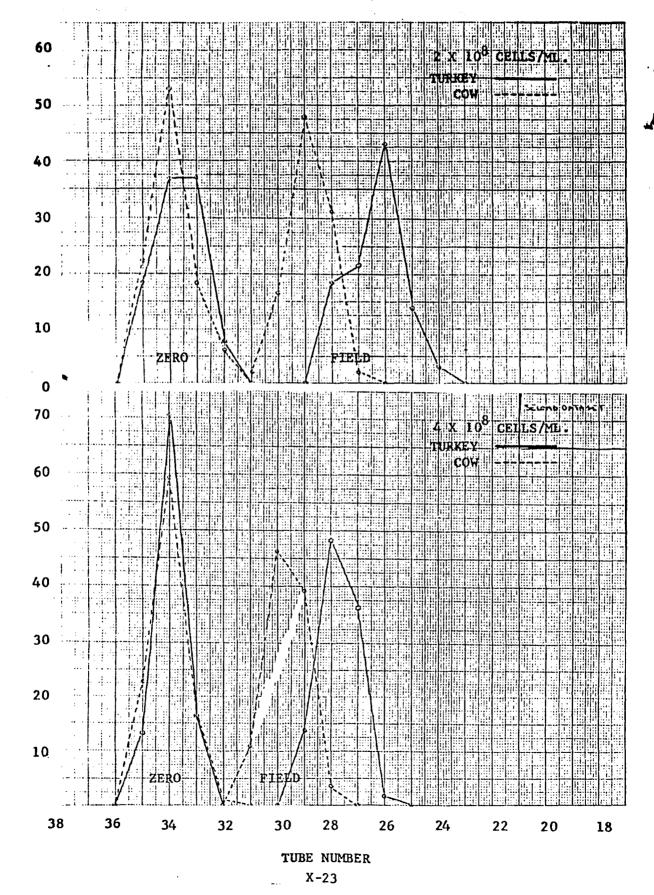


Figure 4. Percent Relative Abundance of Recovered Fixed Cells vs. Tube Number at Indicated Concentration With (Field) and Without (Zero) Voltage Applied: Second Data Set.

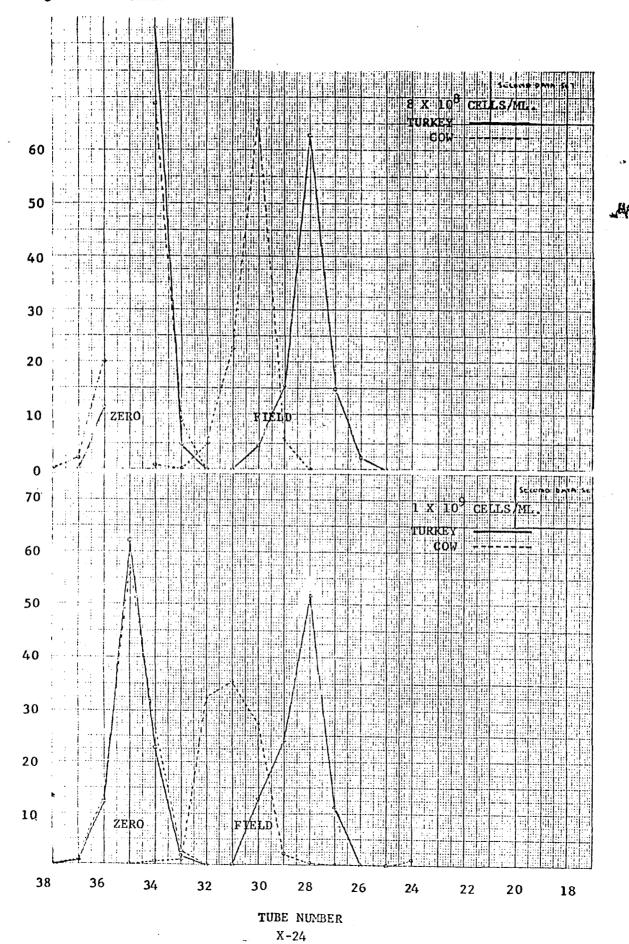
PERCENT RELATIVE ABUNDANCE

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PERCENT RELATIVE ABUNDANCE

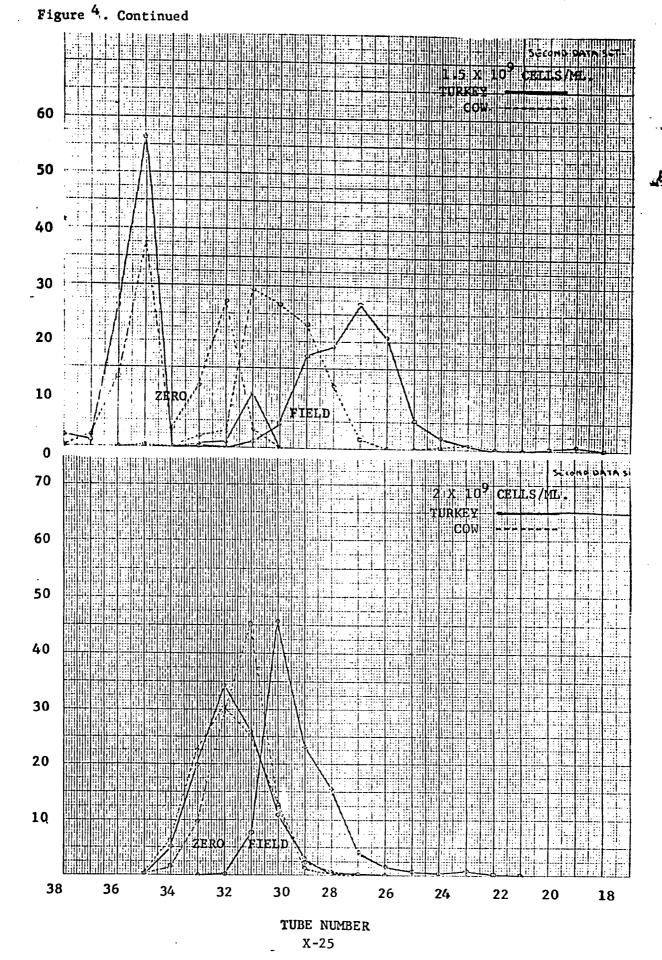
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PERCENT RELATIVE ABUNDANCE

Figure 4. Continued



PERCENT RELATIVE ABUNDANCE

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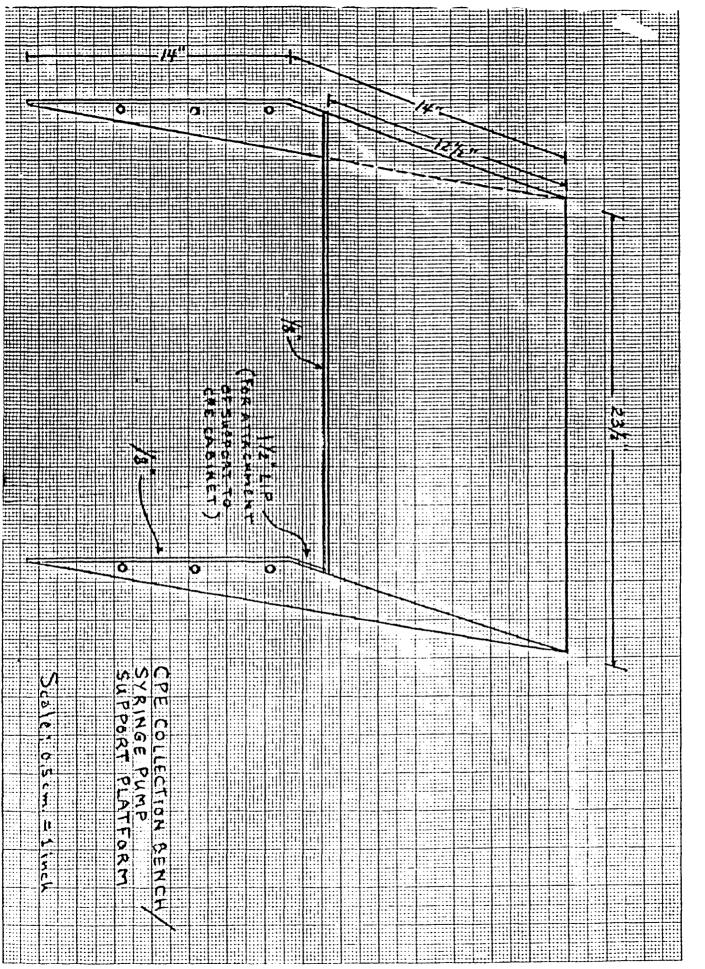


Figure 6. Continuous Particle Electrophoresis (CPE) System: Collection Bench/Syringe Pump Support Platform. 1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

MODELING A MAINTENANCE SIMULATION OF THE GEOSYNCHRONOUS PLATFORM

Prepared by:		A. F. Kleiner, Jr.
Academic Rank	:	Professor
University and	Department:	Drake University Department of Mathematics and Computer Science
NASA/MSFC:	(Office) (Division) (Branch)	Preliminary Design Mission Analysis and Integration Operations Analysis
MSFC Counterp	part:	Dr. J. Steincamp
Date:		August 1, 1980
Contract No. :		NGT-01-002-099 (University of Alabama)

MODELING A MAINTENANCE SIMULATION OF THE GEOSYNCHRONOUS PLATFORM

by

A. F. Kleiner, Jr. Professor of Mathematics Drake University Des Moines, Iowa

ABSTRACT

With the advent of the Space Transportation System it will be possible to service equipment in space. In particular the effect of a maintenance capability for the Geosynchronous Platform (GSP) needs to be examined. Because of the size and complexity of the GSP, analytic results in this area are beyond reach. Hence a simulation study comparing various maintenance routines appears to be the most fruitful approach. The purpose of this report is to outline the basic model which will be used to conduct such a study. Information is included on the computer model and maintenance. A sample system is also modeled.

Acknowledgements

The 1980 NASA/ASEE program has provided me with many benefits both professional and personal. I would like to thank the following people for their assistance this summer: Jim Steincamp, my counterpart, for providing me with an interesting problem and the time to understand it; John Cole, Chief of the Operations Analysis Branch for providing support for my work; Dr. Barfield and Marion Kent for the running of the program; Ronnie Barnes, John Springer and Hank Stoll, fellow participants, for a stimulating carpool and friendship for 8 weeks; Dawn Walton, Carol Byler and Janet Coker for typing this report, and most importantly Mary, Alex and Mimi for making a home away from home for the summer.

1. Introduction

With the advent of the space transportation system NASA will be able to implement a new approach to the reliability of space missions. Rather than build for the mission life from the start it will be possible to visit spacecraft for the purpose of servicing the equipment. The Geosynchronous Platform (GSP) provides a particularly good example of the possibilities of this service for two reasons: (1) GSP will have an extended mission life, 10 to 16 years, requiring heavy redundance costs if maintenance is not available and (2) GSP will serve the communications industry with projected increases in demand over the course of the mission, thus the ability to increase capability is a valuable option.

An analysis of service for GSP is to be carried out to determine if maintenance is a viable concept and to aid in the scheduling of such if it proves feasible. One approach to this problem is to seek an analytic solution by techniques in decision analysis, such as in [1]. The size of this problem however precludes such an analysis (see for example the "small" sample system in Chapter 3). A numerical approach to the same problem can be developed by means of simulation.

Since the problem of maintenance will arise in other settings besides GSP it is useful to develop a general simulation method. Such a method should allow for the incorporation of special features which arise in individual problems such as the need to increase capacity that is an important part of the GSP. This method should be able to accept different mission types and different maintenance policies and produce data which indicate which policies are better.

This report is a preliminary one on work done to develop such a general method. The method is based on earlier work of Dr. J. Steincamp of the Marshall Space Flight Center. Specific details of the method may be found in [3].

2. Objectives

The immediate objective is to construct a simulation model which can be used for an analysis of maintenance for the GSP. It is required to find out if maintenance is an economic approach and if so how it should be carried out.

A more general objective is to develop a method which can answer the above questions for many different space missions. The method is to accept as input the particular system and be capable of running the system through a mission life with different maintenance policies.

This report contains an overview of such a method. The system model is described and illustrated, the basic concepts of a simulation pass are detailed and sections on failures and maintenance are included.

3. The System Model

3.1 Introduction

An operating, physical system may be considered from many points of view. For the purpose of a maintenance simulation two such points of view are required. When a piece of equipment fails it must be replaced, so its location and accessibility are important. Hence a geometric or physical model of the system which details the replacement relationships among the parts is necessary. More over, in order to carry out a program of part failures on the system, it is necessary to have a model of the system's functional or operational mode.

The operation of the system requires that a physical part fail, the operational effects of this failure be carried out and a determination of the need for maintenance of the system be checked. Hence the physical and operational models must be interrelated. The interrelation will be carried out at the most elementary level of the system by identifying the individual parts which fail, called items, with the lowest level operations carried out by the system, called basic functions. Hence forth the dual nature of these items/basic functions will be emphasized by the name IBF. Thus an IBF can be considered either as a specific physical part which supports a basic function of the system or as a basic operation which is performed by some physical item.

3.2 The Physical Model

Inspace fine tuning of systems will not be available in the foreseeable future. Rather, when a part fails some package which contains the failed part will be replaced. This package may contain other parts which have failed as well as good ones. Such a package is called a module. More generally, a module is a replaceable or stockable unit which may contain both items and other modules (sub modules). Each item belongs to same, smallest module. A module which contains only items is called a simple module, one which contains only other modules is called pure and a module which is neither simple nor pure is called mixed. The smallest module which can be used to replace the failed item is said to contain the item.

In a physical system with planned redundancy there will be many instances of duplication. Identical parts will be used in various places and essentially similar but distinct modules will be used for replacement purposes. Although duplicate items will have the same failure distributions, the randomness of the actual failures requires that each physical part be individually represented in the computer model. As a result replacement must be performed and recorded for each module rather than on the classes of similar types. On the other hand it is desirable to take advantage of the similarities among the items and modules for efficient input of the system. For instance if only separate types of items and modules must be entered in the computer a good deal of labor can be saved. Also the identification of these duplicates can streamline ordering, packing and inventory procedures and costs. This identification is established by the concept of generic items and modules. The essential features of each type of item and module are entered into the machine and these types are referred to as generic. The specific items are then identified with the basic functions of the operational model to form the actual IBF's which form the base of the system. It is the IBF's which actually fail according to the failure rule prescribed for each generic item. The actual physical instances of the modules are created by the computer as copies of the appropriate generic module. Records of maintenance are kept for specific modules.

One further feature regarding the physical model of the system is useful. When an IBF fails there are several possibilities for replacement. The IBF may already be scheduled for replacement because it is contained in some module which is to be replaced as a previous failure or because its containing module is a submodule of same module scheduled for replacement. In this case the failure is noted but no adjustment need be made to the list of modules to be replaced. On the other hand if the module containing the IBF was not previously scheduled for replacement it is necessary to add it to the list of modules to be replaced. Moreover any module previously scheduled for replacement whose replacement will be included with the new module should be removed from this list. This procedure is accomplished by distinguishing between kitted and packed modules. A packed module is any module which is scheduled for replacement, either of its own right or because it is contained in some such module. A module is kitted only if it is to be replaced in its own right. Thus, if a module not on the packed list must be replaced, it is to be added to the kitted list, any module which will be replaced with it is removed from the kitted list and all such modules are put on the packed list.

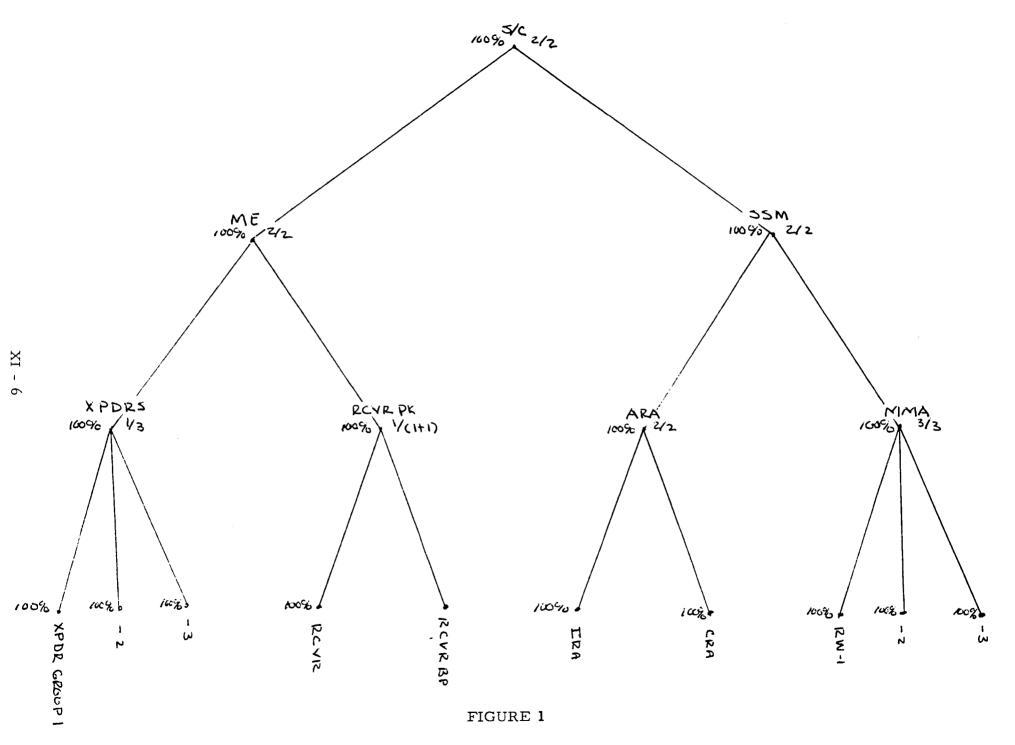
3.3 The Operational Model

Each system has one overall mission or function to perform. This function is accomplished by subfunctions each of which contributes something to the operation of the whole system. Usually, each subfunction can be divided further into functions whose composition is the subfunction. This process of dividing functions into lower level functions until functions which are performed by the basic items of the physical model are reached. These functions, which are called basic functions, are identified with the physical items that perform them and form the IBF's upon which the system rests. Thus the operational mode of the system modeled by a treegraph. The nodes correspond to the various functions with the root node representing the overall system function and the terminal nodes the IBF's.

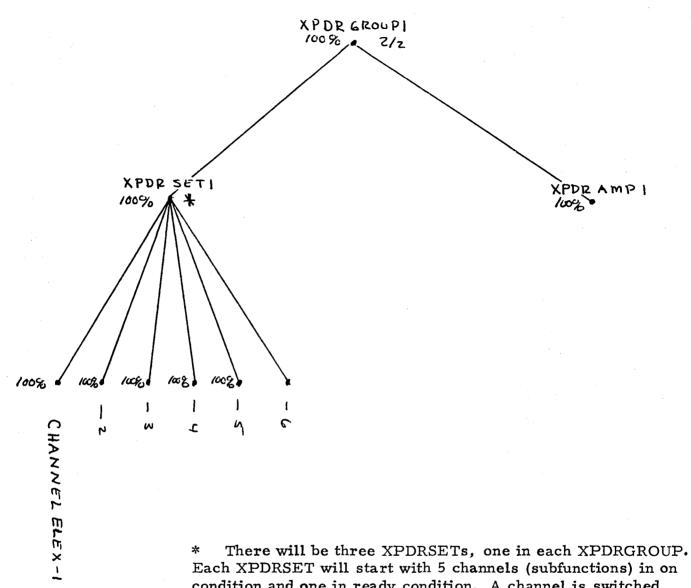
Each function has a set of attributes which are used for defining the system state, processing failures and keeping records. The tree like structure of the system is defined by the attributes chief and base. The chief of a function is the immediate superior of that function in the tree while the base is the set of immediate subordinates of the function. The duty cycle of a function is that proportion of its chief's operating time which it must operate while the condition indicates if the function exists, is online, in standby or has failed. For the purpose of implementing failures, several attributes such as figure of merit, reserve and a user defined index are useful (these are illustrated in the section on failures). Attributes such as operating life, on line time, number of times failed and time to first failure are useful for record keeping and statistics gathering. Each basic function also has a set of failure distributions which are used to cause failure events (see the section on failures).

3.4 A Sample Model

Some of the preceeding concepts are illustrated by the sample model with function tree as displayed in Figures 1-4 and module definitions as in Table 1. This model is for a simplified GSP type system with 119 IBF's, 5 generic modules and 200 functions. Note that in this model several of the modules contain only one IBF and since there is no nesting of modules the distinction between kitted and packed modules does not arise. Each node of the function tree is identified by name and initial duty cycle modifier. Failure rules are included for the non basic functions. The failure distributions for the IBF's are indicated in Table 2.



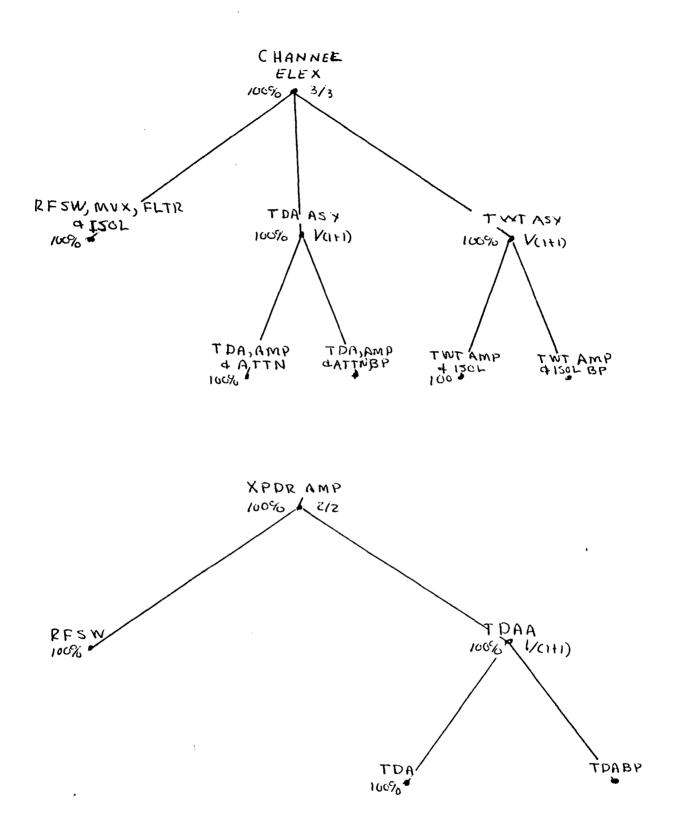
PARTIAL FUNCTION TREE - SAMPLE SYSTEM



condition and one in ready condition. A channel is switched from ready to on if (a) one of the other 5 subfunctions of its chief fails or (b) a channel fails in some other XPDRSET and no backup is left in that set. The XPDRSET will stay on as long as at least one of its channels is on.

FIGURE 2

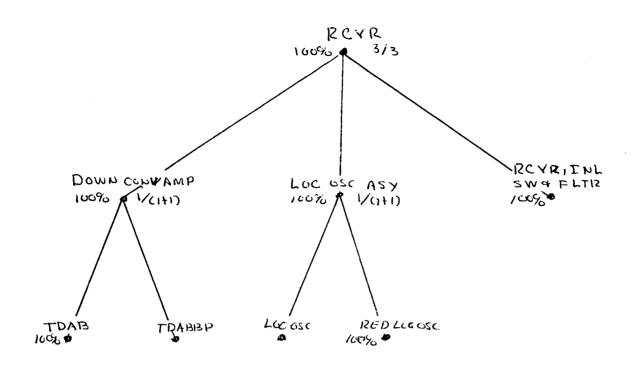
XPDR GROUP TREE - SAMPLE SYSTEM

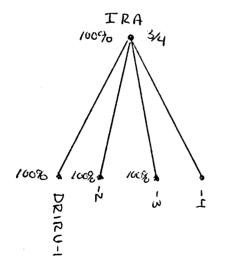


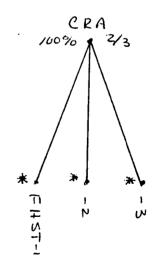


XPDR SET DETAILS - SAMPLE SYSTEM

.







* duty cycle is 66% until first failure then 100%.

FIGURE 4

DETAILS - SAMPLE SYSTEM

MODULE	NUMBER NEEDED	WEIGHT (Kgs)
XPDR GROUP	3	5.2
RCVR PK	1	2.1
IRA	1	3.4
FHST	4	1.2
RW	3	5.1

Payload Capacity of tug is 7.0 Kgs.

Minimum Capability of system is to be 15 channels over a 10 year life.

Desired statistics are

- 1. time average of on line capacity (# of transponders)
- 2. time to first maintenance
- 3. number of times maintenance scheduled due to weight
- 4. number of times maintenance scheduled due to capability
- 5. sample mean and variance for the randum variables
 - a. pass duration
 - b. tug payload weight

TABLE 1

SAMPLE SYSTEM DEFINITIONS

NAME	NUMBER NEEDED	FAILURE DISTRIBUTION
TDA, AMP & ATTN	36	Expon,.488, .0488
TWT, AMP & ISOL	36	Expon, 2.795, .2795*
RFSW, MVX, FLTR & IS	OL 18	Expon, .811, .0811
TDA	6	Expon, .4, .04
TDAB	4	Expon, 1.811, .1181
LOC OSC	4	Expon, 1.6, .16
DRIRU	4	Expon, 15.0
RFSW	3	Expon, .025
FHST	3	Expon, 4.411, .4411
RW	3	Expon, 15.0
RCVR, INL, SW & FLTR	2	Expon, .095,

* Also a drop dead time of 26298 hours

TABLE 2

IBF DEFINITIONS FOR SAMPLE SYSTEM

4. The Simulation Model

4.1 Introduction

As previously indicated, the complexity of this problem suggests an analysis by simulation. For this purpose, the operation of the system across time is best modeled by a discrete event approach with two basic events - failure and maintenance of the system. Various languages are available for such a simulation. In particular, the language SLAM with its discrete event capability and its built-in use of FORTRAN is appropriate in this context (See [2]).

Each overall simulation run will consist of introducing a particular model of the physical system, together with a maintenance policy, demand function and mission lifetime. The system is then run through many passes, each pass corresponding to one mission and the model is re-initialized before each pass. Statistics are compiled at the end of each pass and after the last pass a report is printed. Items of interest typically include the time to first maintenance, total number of maintenance trips for each pass, average capability of the system, etc. The remainder of this section will treat PASS, the subroutine package which completes each simulation pass. The problem of loading the system model is a significant one and has not yet been fully developed - see the section on recommendations and conclusions.

4.2 <u>PASS</u>

The PASS routine sets the system in initial condition, schedules failures of IBF's, carries out failures, schedules maintenance, when required, performs maintenance, increases capability in response to demand, terminates the mission and keeps records.

Several assumptions about the computer model need to be detailed. 'The model is to be capable of increasing the system capability in response to demand. This is done by building the model to its full size at the beginning and introducing a condition DOES NOT EXIST for excess equipment. This equipment is then "brought into existence" during maintenance missions as it is needed. Additionally, some procedure must be chosen for terminating a pass. The standard procedure for this method will be to discontinue maintenance when the mission horizon is reached and then to terminate the mission when the system figure of merit drops to some cutoff point.

Table 3 gives the basic flow ideas of PASS and its major subroutines. These ideas are amplified in the pages following Table 3. PASS

INITIALIZE SAMPLE TIME TO FAIL FAIL MAINT DEMAND MAXTIME STAT

FAIL

MAINT

STORE

DEMAND

IMPACT STAT TIME TO FAIL SMAINT

HEAL MODIFIED FAIL ADDON SAMPLE TIME TO FAIL STAT TURN ON UPDATE SDEMAND TIME TO FAIL STAT

TABLE 3

FLOW IDEAS FOR PASS AND ITS MAJOR SUBROUTINES

INITIALIZE	Sets the model into its initial condition at the start of each pass, each IBF is set to its proper starting condition, duty cycles are computed.
SAMPLE	Input is a set of IBF's, output is the time to fail in on-line (TTFO) and standby (TTFS) modes for each IBF in the input set.
TIME TO FAIL	Input is a set of IBF's, output is the time to next failure for this IBF (see section on Failures for details).
FAIL	An event subroutine which carries out the details of a failure (see Table 3 and below).
MAINT	An event subroutine which carries the details of maintenance (see Table 3 and below).
DEMAND	An event subroutine which carries out the details of "creating" new capacity for the GSP (see Table 3 and below).
MAXTIME	When the mission horizon is reached this routine terminates all maintenance and sets a test for mission termination.
STAT	A general routine which uses the SLAM statistics collection capability to record desired statistics.
IMPACT	Uses actions to carry out the effect of failures on the function trees (see section on Failures).
SMAINT	Uses a given maintenance policy to test the need for maintenance and to schedule maintenance when required.
STORE	Keeps the current records of those failed IBF's which will not be replaced during the current maintenance.
HEAL	Similar to initialize, however, all functions, including those which have been added in response to demand, are set to their initial state.
MODIFIED FAIL	A reduced version of FAIL, this routine refails all IBF whose records have been saved in STORE, no records are kept and no new failure times are required.

ADD ON	All equipment which is "delivered" on the current maintenance mission is "brought into existence" and statistics gathering is started.
TURN ON	New equipment is brought on-line in response to increasing demand.
UPDATE	The system definition which is used to HEAL the system for maintenance is brought up to date.
SDEMAND	The known demand curve is used to schedule the next DEMAND.

.

5. Failures

5.1 <u>IBF Failures</u>

In general, each IBF has a failure distribution with known parameters. Sample time to fail are drawn from this distribution for both on-line (TTFO) and standby (TTFS) modes. The proportion of on-line time (i.e., the duty cycle), \triangleleft , is then used to project the time to fail in each mode. The Red Life Method of failure is assumed here (see [4, page 16]) so that the minimum of these prorated times is used as the time to next failure (TTF) of the IBF. Thus,

TTF = MIN
$$\left\{ TTFO/a , TTFS/(1-a) \right\}$$
.

Variations of this rule occur when the IBF will not fail in standby mode, TTFS =, when it is on, d = 1, or off, d = 0, over an entire interval or when a drop dead time (DDT) requires that the item must fail after a fixed time. These variations are handled by adding or dropping terms in the preceeding formula.

I some other IBF fails first, say after ΔT time units, then the failure times are adjusted by:

TTFO = TTFO -
$$\mathcal{A}(\Delta T)$$
, TTFS = TTFS - (1- \mathcal{A})(ΔT), DDT = DDT- (ΔT)

where \mathbf{d} is the duty cycle of the preceding period. The computation then proceeds as before.

5.2 Failure Propagation

When an IBF fails, a message announcing this failure must be transmitted to its chief. The chief must then react to this message, perhaps issuing orders to its subordinates or send messages further up the function tree. This process must be repeated at each function which receives a message until all effects of the original failure are carried out.

These messages are delivered by means of actions. An action is an internally generated message which includes an ORIGIN (the originating function), a TARGET (the receiving function) and a CMD/CAUSE (a command to change state if the action is directed to a subordinate, a cause if the action is directed to a chief).

Several assumptions are made concerning actions. First among these is the assumption that no function will issue a command to a subordinate unless the subordinate is capable of carrying out the command. Another is that downward actions are always carried out before an upward action. (Actions are filed in a last in - first out set called ACTS).

5.3 A Standard Failure Rule

Let F be a function with a n/((2+m)) failure rule. That is, F has (2+n)identical subordinates and requires that n be in an up condition for F to be up. Further, in the initial state \mathfrak{L} of the subordinates are up ($\mathfrak{L} \ge \mathfrak{n}$) and m have condition ready. The first m failures have no effect on the operation of F, the next **l** -n failures each cause a decrease in the figure of merit of F while the $(\mathbf{Q} + \mathbf{m} - \mathbf{n} + 1)$ th failure causes F to fail. The attributes INDEX, FOM and **RESERVE** are used as follows: INDEX is a user-defined index which assigns to each subordinate of F an integer such that in the initial state the first λ subordinates are up and the last m are ready; FOM is the figure of merit of F, if F is up FOM(F) is the number of subordinates which are on, otherwise FOM(F)=0; RESERVE is a pointer which indicates the lowest index of a subordinate in a ready state, when no such subordinates remain, RESERVE is set to zero. Thus, if FOM(F) \geq N, and S is a subordinate of F, then INDEX(S) is less than RESERVE(F) if COND(S) = UP and greater than or equal to RESERVE(F) if COND(S) = READY. The function F can receive two kinds of notices: a failure notice from a subordinate or a notice to go to a ready state from its chief. The next several tables outline the flow ideas in these two cases.

5.4 A Special Failure Rule

In the sample system of section 3, the XPDRSET's have a special failure rule. Each XPDRSET starts with six channels, five on-line and one standby. When a first channel fails, the standby, if still available, is turned on. When a second fails, the XPDRSET forwards an action to its chief asking that a channel in some other set be turned on if any standbys are left in the system. This can be implemented by using the figure of merit for each XPDRSET to be the number of channels on-line and the reserve to be the number in standby mode. The same meanings are attached to the attributes for each XPDR GROUP while the attributes of ME are the sum of the attributes of all these XPDR GROUPS.

* a failure message has been received by F from subordinate S

IF FOM(F)=0 THEN RETURN	* this failure occurred when F was not up
ELSE IF FOM(F)=N THEN GO TO FAIL	* this caused F to fail
ELSE	* this caused r to fall
IF RESERVE(F)=0 THEN FOM(F)=FOM(F)-1	* no standby left, F does not fail
ELSE	W no standby left, F does not lan
IF INDEX(S) $>$ RESERVE(F) THEN RETURN	* a stanby beyond the pointer failed
ELSE	· a standy beyond the pointer raned
IF THERE IS A T IN BASE(F) WITH	* at least two standbys before the failure
COND(T) = READY AND INDEX(T)	
\$ RESERVE(F) THEN	
IF RESERVE(F) < INDEX(S) THEN CREATE	* failure was on-line, turn on first standby
\$ AN ACTION AS FOLLOWS	
TARGET (ACTION) = $RESERVE(F)$	
ORIGIN (ACTION) = F	
CMD (ACTION) = UP	
FILE ACTION IN ACTS	
RESERVE(F) = INDEX OF FIRST MEMBER	* move pointer to next standby
\$ OF BASE WITH COND = READY	
ELSE	* RESERVE(F) is only standby
IF RESERVE(F) \lt INDEX(S) THEN CREATE	* failure wa s on- line, turn on reserve
\$ AN ACTION AS FOLLOWS	
TARGET (ACTION = $RESERVE(F)$	
ORIGIN (ACTION) = UP	
FILE ACTION IN ACTS	
RESERVE(F) = 0	* no reserves left
RETURN	

STANDARD FAILURE RULE

* the FAIL part of the standard failure rule

FAIL:

COND(F) = FAIL FOM(F) = 0 CREATE AN ACTION AS FOLLOWS TARGET (ACTION) = CHIEF(F) ORIGIN (ACTION) = F CAUSE (ACTION) = FAIL FILE ACTION IN ACTS FOR EACH S IN BASE(F) WITH COND = UP CREATE AN ACTION AS FOLLOWS TARGET (ACTION) = S ORIGIN (ACTION) = F CMD (ACTION) = READY FILE ACTION IN ACTS * adjust attributes of F

* send action to chief

* send action to all subfunctions which are up

RETURN

TABLE 5

FAIL

\$

* F is sent an action to go to ready state

```
COND(F) = READY
FOM(F) = 0
FOR EACH S IN BASE(F) WITH COND = UP
$ CREATE AN ACTION AS FOLLOWS
TARGET (ACTION) = S
ORIGIN (ACTION) = F
CMD (ACTION) = READY
FILE ACTION IN ACTS
RETURN
```

* adjust attribute of F

* send action to all subordinates which are up

TABLE 6

DOWNWARD ACTIONS FOR STANDARD FAILURE RULE

6. Maintenance Policies

6.1 Introduction

It is assumed that demand, D, for the GSP capacity will increase as a function of time and that the demand is known in advance. The minimum acceptable capacity, MC, is typically of the form D + constant where the constant is a "cushion" designed to avoid penalty costs. The nominal capacity, NC, is a desirable level for ordinary purposes. As time increases NC - MC will typically decrease (refer to Figure 5).

At the initial time of the mission, capacity will be NC (0). As time increases failures will occur. Some of these failures will be in the support system and will not cause any decrease in capacity, others will cause such a decrease. For example, in Figure 5 the broken line gives the actual capacity of the system (AC). Capacity failures occur at points 1 and 3, subsystem failures at 2 and 4.

The intent of the simulation is to minimize the cost of maintenance. In order to do this, two sources of cost are considered: (1) the cost of the delivery system, and (2) the inventory cost of delivering capacity before it is needed. The major cost of the delivery system is independent of payload size, thus a trip with less than full payload is inefficient. There exists one or more size limits on the payload (e.g., weight, volume) and these are grouped under the heading size limit. When the first of these is reached, the payload is full. It is never possible to fly with a payload in which the size limit is exceeded.

The inventory cost is kept small by keeping the capacity between the minimal and nominal capacity. Because penalties occur when the capacity is below the minimum the requirement AC \rightarrow MC will be enforced. Thus, whenever the payload is full or the actual capacity drops to the minimum, a service mission is scheduled. Since it is desirable to keep the payload as large as possible and the capacity as close to nominal capacity as possible, two different policies arise. It is assumed that there exists a supply list which contains, in the order needed, the items necessary to bring the mission to nominal capacity.

6.2 Size Constrained Policy

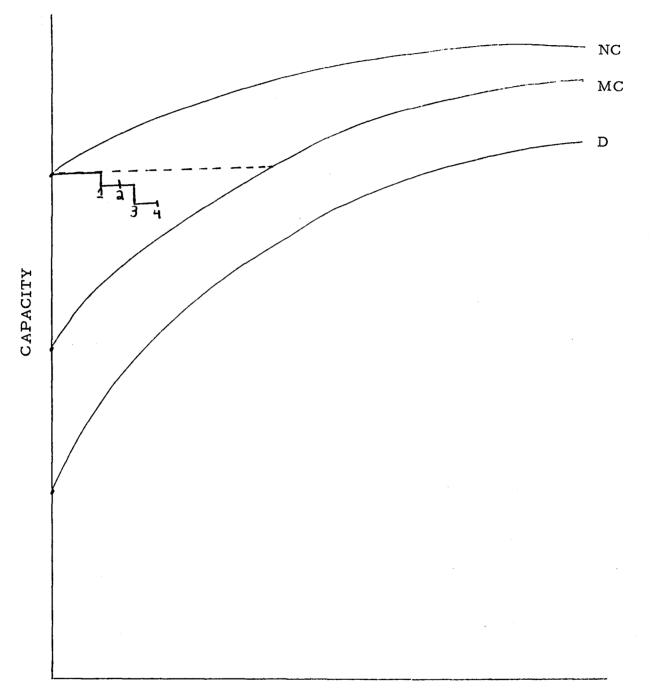
When a failure occurs the proper replacement parts are loaded and a test is made to determine if the size limit has been reached. If so, the service mission is launched. If the size limit has not been reached, the current value of AC is compared to MC. If AC = MC all available space is filled with modules from the supply list and the mission is launched. Under this policy, a service mission always flys a full payload but no relation between AC and NC can be drawn.

6.3 Capacity Constrained Policy

When a failure occurs the proper replacement parts are loaded and so are those items on the supply list which will bring the capacity up to NC at the time of failure. Size and capacity tests are then checked to see if a mission should be launched at the current time. If a size limit is not reached before the actual capacity drops to MC, the launch then takes place when AC = MC. This approach always brings the actual capacity of the system up to the nominal capacity but may produce less than full payloads.

6.4 An Additional Problem

Since failures come at distinct points in time with discrete jumps in capacity (and corresponding payload) it may be that at some point the payload is below the size limit and AC is above MC, however, at the next failure either AC is less than MC or the payload is greater than the size limit. Since minimal capacity is demand plus a cushion, the mission should be flown at once with no change of plans if AC falls below MC. In the case that payload has exceeded the size limit, a problem arises since the size limit is an absolute constraint, the mission cannot be flown if it is violated. In this case, the procedure will be to remove the last package(s) which caused the limit to be exceeded, mark these packages as first on the next service mission and fly the current mission.



TIME

FIGURE 5 CAPACITY CURVES

7. Conclusions and Recommendations

The problem of maintenance of space missions is a serious one and the method outlined herein is a useful tool for approaching this problem. The particular problem of GSP maintenance is a significant one and will be a good test of the method.

Development of tools to solve the GSP problem, starting with the sample system of Chapter 3, will provide significant progress toward the general method.

Among the problems which arise in the implementation of the method one in particular deserves special study. Just as many modules have the same features and can be built into the computer model by the generic module concept, so to many functions one duplicates. A concept of generic function would be useful in this respect (one estimate is that the final GSP model will have 30 identical XPDR GROUPs rather than the 3 of the simple system). In the general case it cannot be assumed that identical basic functions will belong to the same module types. Thus any concept of generic function must allow for the placement of specific IBF's in different modules.

It is recommended that work on this problem continue in two directions. First that a simulation package actually be built and run for GSP and second that work on the general method continue beyond the GSP version of this problem.

REFERENCES

- [1] C. Derman, Finite State Markovian Decision Processes, Academic Press, (New York) 1970.
- [2] A. A. B. Pritsker and C. D. Pegden, Introduction to Simulation and SLAM, John Wiley and Sons (New York) 1979.
- [3] J. Steincamp, 1.2 M XRO OPSIM-, Unpublished Notes, Marshall Space Flight Center.
- [4] S. Sugihara and H. Wong, System Availability Analysis Aerospace Report No. TOR-0066 (5129-01) -17, 1970.

1980 NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM MARSHALL FLIGHT CENTER THE UNIVERSITY OF ALABAMA

AXIALLY SYMMETRIC MOTION OF A ROTATING STRATIFIED FLUID

PREPARED BY:

Benjamin J. Martin, Ph.D.

ACADEMIC RANK:

Professor

COLLEGE AND DEPARTMENT:

Atlanta University Mathematics

NASA/MSFC:

(OFFICE) (DIVISION) (BRANCH) Space Sciences Laboratory Atmospheric Sciences Fluid Dynamics

MSFC COUNTERPART:

DATE:

CONTRACT NO:

Dr. William Fowlis

July 29, 1980

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AXIALLY SYMMETRIC MOTION OF A ROTATING STRATIFIED FLUID

By

Benjamin J. Martin, Ph.D. Professor of Mathematics Atlanta University Atlanta, Georgia

ABSTRACT

A spherical model of the general circulation of the Earth's atmosphere is planned for Spacelab. In order to provide adequate support for the experiment, analytical and numerical studies and simulations must be done. Therefore, an analysis of an earlier effort by Joseph Pedlosky in the area of spherical modeling of geophysical fluid flow is examined. Possible extension of his work is considered and alternative approaches suggested. Recommendations are made for future investigations.

ACKNOWLEDGEMENTS

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NOMENCLATURE

\mathcal{L}	magnitude of the angular velocity vector
R	radius of inner sphere
D	distance between inner and outer spheres
\$	kinematic viscosity
vo	velocity scale factor
$E = \frac{v}{3 \rho 0^2}$	Ekman number $(= \frac{1}{2} E_{\vee})$

$$\varepsilon = \frac{V_0}{2\Omega R}$$

ΔT temperature difference between inner and outer spheres

 α coefficient of thermal expansion

X thermal diffusivity

g gravitational constant

 $\beta = \frac{q_{\alpha} D S \Gamma}{4 \Omega^2 R^2} \text{ thermal Rossby number } (= \frac{1}{4} S)$

T temperature (scaled)

p pressure (scaled)

P density

 $S = \mathcal{V}_R$ aspect rates

u eastward velocity (scaled)

✓ northward velocity (scaled)

ω vertical velocity (scaled)

$$\mu \equiv \sqrt{\frac{1}{2}} | \rho \dot{\rho} \theta |$$

$$\rho g h \theta \equiv \begin{cases} 1, \theta > 0 \\ -1, \theta < 0 \end{cases}$$

$$q' \equiv \frac{d \varphi}{d \theta}$$

.

•

1. INTRODUCTION.

When the Space Shuttle becomes operational some of the flights will take into orbit a laboratory known as Spacelab. A principal use of Spacelab will be the exploitation of the low gravity environment of the shuttle for scientific and technological investigations. In particular, it will be possible to perform geophysical fluid flow model experiments in true spherical geometry. A radial dielectric body force, which is analogous to gravity, can be achieved over a volume of liquid held between two concentric spheres. However on the Earth's surface the dielectric force cannot be made large enough to dominate the effect of terrestrial gravity. Hence the low-gravity environment of Spacelab is the ideal location for these experiments.

In order to adequately design the experiment for Spacelab, preliminary theoretical and numerical work must be done. ^Specifically, the scientific design program for the Atmospheric General Circulation Experiment (AGCE) requires that some baroclinic stability studies in spherical geometry be tackled. This can be done only after a basic state has been established. Simplified models must be used since otherwise the equations are nonseparable.

The particular task here is the extension of some previous work done in spherical geometry by Joseph Pedlosky to the specific case of interest in AGCE. This involves changing the boundary conditions under which the basic equations are to be solved. For simplicity the linear, axially symmetric steady state solution is sought.

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The general governing equations for an incompressible fluid in

dimensional form are

The Momentum Equation

Momentum Equation

$$\frac{\partial}{\partial t} \nabla + (\nabla \cdot \nabla) \nabla + 2 \nabla \times \nabla = -\frac{1}{2} \nabla T + \frac{1}{2} + \nabla T + \frac{1}{2} \nabla T$$

The Thermodynamic Equation

$$\frac{\partial}{\partial t} \widetilde{\tau} + (\overline{v} \cdot \overline{v}) \widetilde{\tau} = u \overline{v}^{3} \widetilde{\tau}$$

The Continuity Equation

and The Equation of State for a liquid

$$\widehat{\rho} = f_{c} \left(1 - \alpha \left(\widehat{\tau} - \tau_{0} \right) \right)$$

Thus we have six equations in six unknowns

$$\overline{\mathsf{V}} = (\widetilde{\mathsf{V}}_r, \widetilde{\mathsf{V}}_{\vartheta}, \widetilde{\mathsf{V}}_{\vartheta}), \ \widetilde{\mathsf{T}}, \ \widetilde{\mathsf{P}}, \ \text{and} \ \widetilde{\mathsf{P}}.$$

Note also that the equations are non-linear. For the model under consideration the steady-state axially-symmetric case is of interest so that the operators are two-dimensional involving only r and e in spherical geometry:

$$(\overline{V} \cdot \nabla) \overline{V} + 3 \overline{\Lambda} \times \overline{V} = - \frac{1}{P} \nabla \widetilde{P} + \overline{q} + V \nabla^2 \overline{V}$$
$$(\overline{V} \cdot \nabla) \widetilde{T} = \mathcal{X} \nabla^2 \widetilde{T}$$
$$\nabla \cdot \overline{V} = 0$$
$$\widetilde{P} = P_0 (1 - \alpha (\widehat{T} - \overline{T}_0)).$$

Introducing non-dimensional variables

$$\widetilde{r} = R + SR Z, S = \frac{D}{R}$$

$$\widetilde{V}_{r} = S V_{0} \omega(e, Z)$$

$$\widetilde{V}_{\theta} = V_{0} V(\theta, Z)$$

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$$\widetilde{V}\varphi = V_0 u(\theta, \overline{z})$$

$$\widetilde{T} = T_0 + \Delta T \overline{z} + (\Im \Omega R V_0 / g \alpha D) \overline{T}(\theta, \overline{z})$$

$$\widetilde{P} = P_0 - f_0 g D \overline{z} + \frac{1}{2} g \alpha D \Delta T f_0 \overline{z}^2 + 2\Omega V_0 R f_0 P(\theta, \overline{z})$$

the equations become

$$\begin{split} \mathcal{E}\left[\omega\frac{\partial u}{\partial z} + v\frac{\partial u}{\partial \theta} + \delta u\omega - u \cdot v \cdot \overline{ton \theta}\right] + (\omega \delta tod \theta - v' \cdot \Delta u \cdot \theta) \\ &= S^{2} \mathbb{E}\left[\frac{1}{tod \theta} \frac{\partial}{\partial \theta} \left(\frac{\partial u}{\partial \theta} \cdot \cos \theta\right) - \frac{u}{tod \theta}\right] + \mathbb{E}\left[\frac{\partial^{2} u}{\partial z^{2}} + \partial \delta \frac{\partial u}{\partial z}\right] \\ \mathcal{E}\left[\omega\frac{\partial v}{\partial z} + v'\frac{\partial v}{\partial \theta} + S v'\omega + u^{2} \cdot ton \theta\right] + u \cdot \Delta u \cdot \theta = -\frac{\partial p}{\partial \theta} + \partial S^{3} \mathbb{E}\left[\frac{\partial w}{\partial \theta} + S^{3} \mathbb{E}\left[\frac{$$

$$\frac{\partial w}{\partial z}$$
 + 2 S w + $\frac{1}{\cos \theta} \frac{\partial}{\partial \theta} (V \cos \theta) = 0$

For $\boldsymbol{\mathcal{E}}$ sufficiently small the above equations can be linearized

$$\begin{split} & Sul (use - V ain e = S^{2} E \left[\frac{1}{(use \frac{1}{2}e)} \frac{1}{2e} \left(\frac{1}{2e} e^{\frac{1}{2}e} \left(\frac{1}{2e} e^{\frac{1}{2}e} e^{\frac$$

These are the equations of motion interms of non-dimensional variables which are appropriate for linear, axially-symmetric, steadystate motion of an incompressible fluid in spherical geometry.

3. THE SOLUTION OF THE SYSTEM.

The system that is used here is the system as derived and simplified by Pedlosky [1] with some obvious corrections and a few notational changes

$$-V \Delta u \dot{u} \theta + S w \cos \theta = S^{2} E \left[\frac{1}{\cos \theta} \frac{2}{\partial \theta} \left(\frac{2u}{\partial \theta} \cos \theta \right) - \frac{u}{\cos \theta} \right] + E \left[\frac{3^{2}u}{32^{2}} + 3S \frac{3u}{32} \right] u \Delta u \dot{u} \theta = -\frac{3T}{3\theta} + S^{2} E \left[\frac{1}{\cos \theta} \frac{2}{3\theta} \left(\frac{2v}{3\theta} \cos \theta \right) - \frac{v}{\cos \theta} \right] + E \left[\frac{3^{2}v}{32^{2}} + 3S \frac{3v}{32} \right]$$

$$-Su \cos \theta = -\frac{\partial P}{\partial z} + T + S^{4} E \left[\frac{\partial}{\partial x \theta} \frac{\partial}{\partial \theta} \left(\frac{\partial w}{\partial \theta} \cos \theta \right) \right] \\ + S^{3} E \left[\frac{\partial^{2} w}{\partial z^{2}} + \partial S \frac{\partial w}{\partial z} \right]$$

$$\beta \omega = \frac{S^{2}E}{\sigma} \frac{1}{\omega 0} \frac{\partial}{\partial \theta} \left(\frac{\partial F}{\partial \theta} \cos \theta \right) + \frac{E}{\sigma} \left[\frac{\partial^{2}F}{\partial z^{2}} + \partial S \frac{\partial F}{\partial z} \right]$$

$$\frac{\partial \omega}{\partial z} + \partial S \omega + \frac{1}{\omega 0} \frac{\partial}{\partial \theta} \left(V \cos \theta \right) = 0.$$

The boundary conditions are

$$(u, v, w) = (0, 0, 0)$$
 on $z=0$ and $z=1$
 $T(0, 0) = H_{1}(0)$
 $T(0, 1) = H_{T}(0)$

In order to solve the system with reasonable effort, the solution is decomposed into three parts--the interior motion (the I-subscripted variables), the lower boundary layer corrections (the L-subscripted variables), and the upper boundary layer corrections. Therefore

$$\begin{split} & u(\theta, \xi) = u_{I}(\theta, \xi) + u_{T}(\theta, \chi) + u_{L}(\theta, q) \\ & \forall (\theta, \xi) = E V_{I}(\theta, \xi) + V_{T}(\theta, \chi) + V_{L}(\theta, q) \\ & u(\theta, \xi) = E^{V_{2}} \left[w_{I}(\theta, \xi) + w_{T}(\theta, \chi) + w_{L}(\theta, q) \right] \\ & T(\theta, \xi) = P_{I}(\theta, \xi) + S E^{V_{2}} \left[P_{T}(\theta, \chi) + P_{L}(\theta, q) \right] \\ & T(\theta, \xi) = T_{I}(\theta, \xi) + E^{V_{2}} \beta \left[T_{T}(\theta, \chi) + T_{L}(\theta, q) \right] \end{split}$$

where $Y = (1-2) E^{\frac{1}{2}}$ and $Y = 2 E^{\frac{1}{2}}$. Since the L and T subscripted variables are signicant only near the boundary, it is required that they tend to zero for large Y and χ .

The equations for the lowest order correction functions in the respective boundary layers are

Lower Boundary Layer

$$-V_{L} \sin \theta = \frac{\partial^{3} u_{L}}{\partial y^{2}} \qquad (3.1)$$

$$u_{L} \sin \theta = \frac{\partial^{3} v_{L}}{\partial y^{2}} \qquad (3.3)$$

$$-S u_{L} \cos \theta = -S \frac{\partial P_{L}}{\partial y} + E^{\frac{N_{2}}{2}} \beta \overline{\Gamma}_{L} \qquad (3.3)$$

$$w_{L} = \frac{1}{2} \frac{\partial T_{L}}{\partial y^{2}} \qquad (34)$$

$$\frac{\partial w_{L}}{\partial y} = -\frac{1}{\cos 2} \frac{\partial}{\partial \theta} \left(V_{L} \cos 2 \right) \quad (3.5)$$

Upper Boundary Layer

$$-V_{T} \sin \theta = \frac{\partial^{2} u_{T}}{\partial x^{2}}$$

$$u_{L} \sin \theta = \frac{\partial^{2} V_{T}}{\partial x^{2}}$$

$$-S u_{T} \cos \theta = S \frac{\partial P_{T}}{\partial x} + E^{V_{2}} \beta T_{T}$$

$$w_{T} = \frac{1}{\partial} \frac{\partial^{2} T_{T}}{\partial x^{2}}$$

$$\frac{\partial w_{T}}{\partial y} = \frac{1}{\cos \theta} \frac{\partial}{\partial \theta} (V_{T} \cos \theta)$$

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Similarly equations for the interior solution are obtained to lowest

order in E

$$V_{I} \Delta w_{i} \Theta = \frac{\partial^{2} u_{I}}{\partial z^{2}} \qquad (3.6)$$

$$U_{I} \Delta w_{i} \Theta = -\frac{\partial P_{I}}{\partial \Theta} \qquad (3.7)$$

$$T_{I} = \frac{\partial P_{I}}{\partial Z} \qquad (3.8)$$

$$\partial T \beta E^{-V_{Z}} \omega_{I} = \frac{\partial^{2} T_{I}}{\partial z^{2}} \qquad (3.4)$$

$$\frac{\partial \omega_{I}}{\partial z} = 0 \qquad (3.10)$$

where, following Pedlosky, we have assumed $\zeta < \varepsilon^{t_2}$ and have dropped all terms involving powers of ζ . The Boundary Layer Corrections.

The system of equations for the lower boundary layer corrections can be easily solved. First the coupled equations (3.1) and (3.2) are solved. This allows the solution of equation (3.5), which then permits the solution of equation (3.6).

$$\begin{aligned} U_{L}(\theta,q) &= -U_{I}(\theta,c) \in \frac{\mu^{2}}{6} & (e^{\mu}) \\ V_{L}(\theta,q) &= U_{I}(\theta,c) \in \frac{\mu^{2}}{6} & (e^{\mu}) \\ & (e^{\mu}) = \frac{1}{696} \xrightarrow{2}{30} \left[\frac{U_{I}(e,c)}{6} & p_{I}me^{2} e^{-\mu^{2}} \\ & (e^{\mu}) e^{\mu} e^{2\mu} \\ & T_{L}(\theta,q) = \frac{1}{600} \xrightarrow{2}{30} \left[\frac{U_{I}(e,c)}{6} & p_{I}me^{2} e^{-\mu^{2}} \\ & (e^{\mu}) e^{\mu} e^{2\mu} \\ & (e^{\mu}) e^{2\mu} e^{2\mu} \\ & (e^{\mu}) e^{2$$

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where the boundary conditions have been used to get

$$\begin{aligned} u_{I}(\theta, o) + u_{L}(\theta, o) &= c \\ EV_{L}(\theta, o) + V_{L}(\theta, c) \approx U_{L}(\theta, o) &= 0 \\ T_{I}(\theta, o) + E^{V_{2}} \beta T_{L}(\theta, c) \approx T_{I}(\theta, c) &= H_{L}(\theta) \quad (3.11) \\ \omega_{I}(\theta, o) + \omega_{L}(\theta, c) &= c \end{aligned}$$

2

$$\omega_{I}(e,o) = -\frac{1}{2lore} \frac{\partial}{\partial e} \left(\frac{\omega_{I}(b,o)}{\mu} \frac{\Delta q_{II}e^{-ce_{I}}e^{-c}}{\mu} \right) (3.12)$$

Similarly, for the upper boundary layer corrections we have $U_{T}(\theta, \chi) = -e^{-\mu\chi} U_{I}(\theta, 1) \quad (0, \mu\chi)$ $U_{T}(\theta, \chi) = e^{-\mu\chi} U_{I}(\theta, 1) \quad (0, \mu\chi) \quad (0, \chi) = -\frac{1}{600} \quad \frac{3}{30} \left[\frac{U_{I}(\theta, 1) \quad xq_{M}\theta}{\sqrt{3}} e^{-\mu\chi} \quad (0, \chi) = -\frac{1}{600} \quad \frac{3}{30} \left[\frac{U_{I}(\theta, 1) \quad xq_{M}\theta}{\sqrt{3}} e^{-\mu\chi} \quad (0, \chi) = -\frac{1}{600} \quad \frac{3}{30} \left[\frac{U_{I}(\theta, 1) \quad xq_{M}\theta}{\sqrt{3}} e^{-\mu\chi} \quad (0, \chi) = -\frac{1}{600} \quad \frac{3}{30} \left[\frac{U_{I}(\theta, 1) \quad xq_{M}\theta}{\sqrt{3}} e^{-\mu\chi} \quad (0, \chi) = -\frac{1}{600} \quad \frac{3}{30} \left[\frac{U_{I}(\theta, 1) \quad xq_{M}\theta}{\sqrt{3}} e^{-\mu\chi} \quad (0, \chi) = -\frac{1}{600} \quad \frac{3}{30} \left[\frac{U_{I}(\theta, 1) \quad xq_{M}\theta}{\sqrt{3}} e^{-\mu\chi} \quad (0, \chi) = -\frac{1}{600} \quad (0, \chi) = -$

The Interior Variables.

Equation (3.10) implies that $\omega_{\underline{r}}$ is independent of $\underline{2}$ therefore $\omega_{\underline{r}}(\theta, 0) = \omega_{\underline{r}}(\theta, 1)$. Hence $\mathcal{U}_{\underline{r}}(\theta, 0) = \mathcal{U}_{\underline{r}}(\theta, 0)$. Assuming $\mathcal{U}_{\underline{r}}(\theta, 0) = \omega_{\underline{r}}(\theta, 1)$ to be known the system can be solved beginning with equation (3.9) working back to equation (3.6):

$$T_{I}(0,z) = \sigma \beta E^{-V_{2}} \omega_{I}(0) (z^{2} - z) + H_{I}(0) z - H_{L}(0) (z^{-1})$$

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$$U_{I}(\theta, z) = -\frac{1}{2\omega_{t}\theta} \left[\sigma_{B} E^{-1/2} \omega_{I}'(\theta) \left(\frac{4z^{3} - 6z^{2} + 1}{12} + H_{r}'(\theta) \frac{\theta z^{3} - 1}{4} \right) - H_{L}'(\theta) \left(\frac{3z^{3} - 4z + 1}{4} \right) \right]$$

$$-H_{L}'(\theta) \left(\frac{3z^{3} - 4z + 1}{4} \right) \left[\frac{1}{4} + H_{r}'(\theta) - H_{L}'(\theta) \right]$$

$$V_{I}(\theta, z) = \frac{1}{2\omega_{t}^{3}\theta} \left[\sigma_{B} E^{-1/2} \omega_{I}'(\theta) \left(\frac{3z - 1}{4} + H_{r}'(\theta) - H_{L}'(\theta) \right) \right]$$

(Note: $\mathcal{P}(\theta, z)$ is not of interest except as needed to find the other dynamic variables and thus has been omitted.)

The system is completely solved when $U_{I}(\theta,c)$ is known (or, equivalently, $W_{I}(\theta)$). $W_{I}(\theta)$ has not yet been found but a second order differential equation for W_{I} can be derived $W_{I}(\theta) = \frac{\Delta q_{L} \theta}{2 \cos \theta} \frac{d}{d\theta} \left\{ \begin{bmatrix} -\frac{1}{12} \sigma \beta e^{-hL} \frac{d}{d\theta} W_{I}(\theta) - \frac{1}{4} (H_{T}'(\theta) + H_{L}'(\theta)) \end{bmatrix} \frac{\cos \theta}{\mu \sin \theta} \right\}.$

This equation has not yet been solved.

4. CONCLUSIONS AND RECOMMENDATIONS.

(a) The technique of Pedlosky is a pertubation technique, however this is disguised in the analysis. The net result is that the order of approximation is different for each variable. It is recommended therefore that each variable be explicitly represented as a perturbation, e.g., $\phi = \phi_0 + E^{k_2} \phi_1 + E \phi_3 + \cdots$, and all solutions be found to the same order of approximation if possible or as nearly as is possible.

(b) Some of the assumptions of smallness made by Pedlosky are too stringent for AGCE. As many as possible should be relaxed, in particular, $\zeta < < E^{\frac{h_2}{2}}$. (c) The author derived the equations of motion for spherical geometry. The resulting equations differ non-trivially from those of Pedlosky. Unless these discrepancies can be adequately explained, all future work should be done using the author's equations.

(d) Some consideration should be given to the non-linear equations, possibly using the technique of Rahm and Walin [2].

REFERENCES

- Pedlosky, J., "Axially symmetric motion of a stratified rotating fluid in a spherical annulus of narrow gap", <u>J. Fluid Mechanics</u>, Vol. 36, part 2, 1969.
- 2. Rahm, Lars and Walin, Gosta, "On thermally forced stratified rotating fluids", J. Fluid Mechanics, Vol. 97, part 4, 1980.
- 3. Reid, William H., editor, <u>Mathematical Problems in the Geophysical</u> Sciences, American Mathematical Society (Providence, Rhode Island: 1971).
- 4. Fein, Jay S., <u>Boundary Layers in Homogeneous and Stratified</u> <u>Rotating Fluids</u>, The University Presses of Florida (Gainesville, Fla: 1978).

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

A RANKING ALGORITHM FOR SPACELAB CREW AND EXPERIMENT SCHEDULING

Prepared by: Frank H. Mathis and Robert D. Grone

Refer to Section IX for report.

1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

PAYLOAD CREW TRAINING COMPLEX (PCTC) UTILIZATION AND TRAINING PLAN STATUS REPORT

Prepared By:

Michael R. Self, Ed.D.

Academic Rank:

Assistant Research Professor

University and Department:

University of Alabama in Huntsville Developmental Learning Program

NASA/MSFC: (Laboratory) (Division) (Branch)

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PAYLOAD CREW TRAINING COMPLEX (PCTC) UTILIZATION AND TRAINING PLAN

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Michael R. Self, Ed.D. Assistant Research Professor University of Alabama-Huntsville Huntsville, AL

ABSTRACT

The first section describes the physical facilities that comprise the Payload Crew Training Complex (PCTC). The PCTC is located at MSFC, Huntsville, Alabama, and contains the Host Simulator; Experiment Simulators; Spacelab Aft Flight Deck, Experiment Pallet, and Experiment Rack mockups; the Simulation Director's Console; Payload Operations Control Center; classrooms; and supporting soft- and hardware.

The second section of the report establishes the parameters of a training philosophy for Payload Crew Training at the PCTC. Major parameters include: (1) The intent of training is not to be diagnostic, to select the "best" Payload Crew, but should be enhancing to the development of high levels of proficiency in every Payload Crew trainee; (2) The intelligence of each Payload Crew trainee is obviously to be respected. This implies that training at the PCTC, while highly procedural and sophisticated should tend to be less rigid than similar kinds of traditional high technology training; and (3) The function of trainee evaluation will be to determine how well training is being conducted. From the cooperative debriefings between the Payload Crew and PCTC trainers, should come ideas for improving the training sequence.

Section three discusses the development of the training plan. Included are discussions of preassessment, skill identification, procedures, material development, and evaluation options.

INTRODUCTION

NASA finds itself on the threshold of a new era in space travel, namely repeated flights with a reuseable Space Transportation System (STS), i.e., the Space Shuttle. Additionally, they have diversified their payload configurations to include a larger portion of the scientific community not traditionally associated with NASA. For the first time too, the European Space Agency (ESA) has primary development responsibility for the Shuttle payload. The result is a new learning opportunity in the co-development of space missions and requires new thinking in training concepts.

One of the major distinctions between past NASA manned flights and this new era has been the changes in training philosophy and requirements. In the past everyone that has flown space missions has been a career astronaut with a scientific or technical background with flight training and experience. Additionally, all of the training tended to be centralized under NASA auspices. Yet in the presently scheduled flights, i.e., Spacelabs 1, 2, and 3, the Principle Investigators (PI) have primary responsibility for training the Payload Crew (PC) in how to conduct the PI's experiments. The PI's are research oriented, typically in a university setting, and literally distributed around the world.

These factors have precipitated the need for decentralized training. The effect of the move to decentralized training has been to: 1) reduce substantially the cost of training to NASA, since NASA does not have to reproduce all of the PI's equipment for training, and 2) add impetus to assuring overall flight competence prior to the actual flight via systems integration training. Concern with this second component is part of the issue addressed in this document. Another facet of the issue is how to best use the Payload Crew Training Complex (PCTC).

The PCTC was established to train the PC on those experiments that require interface with the onboard Spacelab computer, i.e., the Command and Data Management System (CDMS). In establishing the facility many questions arose as to how to achieve maximum utilization from the facility. Some of those questions are addressed herein.

OBJECTIVES

The concern then has been to examine the parameters associated with decentralized training; which, as mentioned previously is a situation unique to the Space Shuttle/Spacelab missions, but a concept that appears to be one that NASA plans to pursue at least for the foreseeable future. The objectives then that derive from this goal are, at present, threefold:

- 1) To describe the PCTC and its training capabilities.
- 2) To define a training philosophy that might guide the PCTC trainers in developing a training plan.
- 3) To describe a procedure for the development and implementation of a training plan for training the PC at the PCTC.

What makes this report a status report is that training has not yet begun, so there are no evaluations of actual training procedures discussed. Furthermore, decisions on the best training approach are still in a deliberative state due to the fact that all of the experiments requiring CDMS interface, while known, have not been broken down into the step-by-step procedures level. This breakdown is basic to skill identification and the subsequent training procedures development.

PCTC

The PCTC is located at MSFC, Huntsville, Alabama. The facility houses the Host Simulator (HS); Experiment Simulators (ES); the Spacelab Aft Flight Deck (AFD), Experiment Pallet, and Experiment Rack mockups; a Simulation Director's Console (SDC), a Payload Operations Control Center (POCC) console; classrooms; and supporting soft- and hardware.

The HS is the computer system that provides the capability to simulate the command, communications, control, and display functions of the Spacelab CDMS. This capability is necessary to execute experiment-unique simulations. It has the following features:

- 1) Two Data Display Systems (DDS) which can be simultaneously operated in support of one mission phase.
- 2) Reconfiguration and turnaround for the support of one training mission to another can be accomplished in eight hours.
- 3) Drives the SDC and POCC consoles for training and to approximate flight conditions.
- 4) Is modular in design and does not preclude expansion in size or computing speed.
- 5) Can simultaneously simulate multiple experiments, consistent with CDMS operating capabilities and provides standard analog-to-digital and digital-to-analog services for interfacing with simple experiment-unique control and display panels.
- 6) Allows for operation in a real time mode so that display time and command responses are commensurate with actual flight conditions.
- 7) Simulates symbolic pictures of the earth, sun, and sky scenes in a standard TV format for training in experiment pointing commands.
- 8) Protects against inadvertent damage to hardware and/or software by any operator or trainee.
- 9) Allows for environmental inputs to experiment models such as orbital state vector, vehicle attitude, and ephemeris of heavenly bodies.
- 10) Allows for detection and accumulation of characteristic data, such as PC errors, response times, and operating times.

The ES's include the hardware and software necessary to simulate experiments through the HS during training. A Data Display Unit (DDU) showing simulated data for experiment operations originates from either an instrument or Experiment Computer Applications Software (ECAS) simulation software package.

Experiment-unique command data is initiated through the DDS-Keyboard (KB). The CDMS is entered via KB command, item, type, and/or function keys and acted upon, and thus generates the appropriate response within the simulation software. Each ES will send data and accept commands from certain key simulated control and display panels located in the experiment racks onboard the Spacelab mockup. Field of view imagery and alphanumeric data for display on payload-unique Cathode Ray Tubes (CRT's) are provided from the ES. Simulated POCC ground command inputs are available to the crew via the ES. The ES accepts and processes commands from the SDC during simulation. These commands control such simulation circumstances as simulation modes, control sequencing, fault initiation, and time cycling.

The SDC enables the Simulation Director and Assistant Simulation Director to monitor, control, and modify the simulation operation. It is composed of four racks containing two computer terminals, a hard copy machine, four large black and white Closed Circuit TV (CCTV) monitors, and a work surface which encompasses the CCTV camera controls. DDU monitors in rack 4 will display the same DDU displays as the trainee sees in the Spacelab module. Cameras for the TV monitors will be used thus: two inside the Spacelab module, one for the AFD and one for general surveillance of the high bay area. CCTV cameras are controllable in pitch and azimuth from hand controllers on the SDC work surface. Three video tape recorders may be used to record data from any one of several video sources: four CCTV cameras, four DDU's, and three Spacelab CCTV's used with scene generation. Hard copies of data seen on any of the CRT's are available by merely pressing a button on the hard copier after selecting the appropriate CRT.

The PCTC mockups include the Spacelab Module Interior, Shuttle AFD Interior, and associated equipment and pallets for part task activities with the capability for supporting payload crew training for any single Spacelab mission and the capability for reconfiguration to any payload.

TRAINING PHILOSOPHY

A training philosophy is nothing more than an acknowledgement of the guiding ideals within which training will hopefully be implemented. The components of that philosophy may not actually be formally stated again, but they should be reflected in how the training is actually conducted. Philosophy statements describe thoughts or events that typically are extremely difficult to quantify for measurement, yet are important enough to be considered. The training philosophy for the utilization of the PCTC hopefully incorporates the following parameters:

- The intent of training is not to be diagnostic to select the "best" PC, but should facilitate the development of high levels of proficiency in every PC trainee.
- 2) The intelligence of each PC trainee is obviously to be respected. This implies that training at the PCTC, while highly procedural and sophisticated should tend to be less rigid than similar kinds of traditional high technology training.
- 3) The function of trainee evaluation will be to determine how well training is being conducted. From the cooperative debriefings between the PC and PCTC trainers should come ideas for improving the training sequence to help assure the highest quality of training and the maximum utilization of the PCTC's capabilities.

Specifically, those philosophical considerations include:

- 1) In order to proceed in a logical training sequence, critical skills and critical experiment sequences should be identified to better prepare the PC for the time when he actually encounters them.
- 2) There is a need for PC assessment upon initiation of training at the PCTC. This assessment is necessary for the PCTC trainers to correctly determine how well they are conducting the training. Initial assessment should probably involve assessment of the PC's skills in utilizing the DDU/KB, understanding the Experiment Computer Operating System (ECOS), and communication skills.
- 3) While individual PC decision making ability is a major concern, it is not planned to "assess" those skills during PCTC training. Instead, the goal is to provide training experiences to facilitate those skills. For example, while PCTC training is highly procedural, in selecting the contingencies for use in training, some of the training scenarios are speculative in nature. They have been developed using the best information available, but in some cases the exact flow of the experiment will not be encountered until the actual flight. Therefore, a training goal should be for the PC to remain cognizant of the overall experiment goal(s). The value of this is to be able to periodically step back from the minutia of experiment procedures and conceptually think through what is supposed to be happening. This should enhance the likehood that the PC would make the best decision when, in flight, he actually encounters that option point in the experiment flow.
- 4) It follows from the above statement that a large segment of learning will not occur until the actual flight because of the sheer uniqueness of space flight; therefore, optimum operations proficiency may be approximated but never achieved. This becomes important in setting achievement criteria for training.

- 5) Training should proceed from the simple to the complex, both in skill acquisition and in experiment training. There is some thought about starting with the more complex because of their criticality and the need for ample training time. However, it is conversely believed that if the PC trainee is allowed to progress from simple to complex, and if an adequate amount of time is allowed for learning each experiment regardless of its complexity, that the PC trainee will be developing the necessary skills to more easily handle the complex experiment when he gets to it. For example, by taking the simple experiments first the trainee is getting familiar and comfortable with the DDU/KB, experiment layout, ECOS, various ECAS/DEP, and the language in which the procedures are written. Then he gathers experience in nominal operations, simple off-nominals, and simple contingencies. All the while he is developing his skill to recognize subtle differences on the DDU displays, learns to discriminate messages, and is enhancing his intuitive knowledge of all the systems. Then when he encounters the more complex experiment, his prior knowledge and experiment experience actually aid in reducing some of the complexity factors of the experiment.
- 6) There should be sufficient training opportunities in the training schedule to allow the PC to work in the various possible dyads that might actually be chosen to fly the various Spacelab missions. Again this is not done to gather information for selection purposes (since the Investigators Working Group (IWG) will actually make the choices), but rather to give the dyads themselves the opportunity to better learn each others style of operation and idiosyncrasies prior to the flight crew selection.
- 7) The focus of training at the PCTC is on CDMS controlled experiments. Non-CDMS experiments will only be addressed as they mesh into the timeline training, i.e., while the PC is conducting a CDMS supported experiment, his operations timeline might call for him to initiate a life sciences experiment, hence, he might walk over to the lifesciences rack and flip a dummy switch and then resume his CDMS procedures.

TRAINING

The intent of this section is to discuss how the PCTC will actually be utilized during training. At present, however, all of the information necessary to make such a plan is incomplete. Therefore, this section represents an estimation of how that training might progress.

Traditionally, training is approached as the compilation of its sequential parts. That sequence is usually thought to be assessment, objectives, materials, procedures, and evaluation. That is generally how training has been approached in the present endeaver. Assessment in this situation refers to pretraining assessment. Three questions have been explored here; 1) why have pretraining assessment, 2) what type assessment is mandated, and 3) who will do the assessment? The why was partially addressed in the previous philosophy section; specifically, the PCTC trainers need to determine how well they are conducting the training. In order to achieve that goal, it will be necessary to determine where the PC trainee is in his training when he arrives at the PCTC. This will allow the PCTC trainers to assess how the trainee has progressed during training and then devise ways to better conduct future training. The second aspect of the why question deals with determining entry level abilities at skills basic to CDMS operation. At present, five areas have been tentatively identified as being basic to all experiment operations, and as such will comprise the initial assessment/training period at the PCTC. The five areas are:

- 1. Keyboard operation; skill and location
- 2. CDMS entry; recognizing different DDU displays, functions, etc.
- 3. ECOS operations
- 4. Communications; transmitting acurately and quickly with POCC
- 5. Timeline operations; getting the total picture of when everything is supposed to run during the mission.

The type of assessment could be group or individual discussions or written exams to determine overall understanding of the various areas. This assessment could preceed, run concurrent with, or follow PC trainee demonstration of skill proficiency in those areas requiring manipulative skills. At present, our thinking is that a group of PCTC trainers could do the assessment, supported with statistical records supplied by the HS on key stroke rate, error, and ability to call up various pages and experiments.

The objectives of training would include the development of high operations proficiency and overall experiment competence. Operations proficiency would include the PC's ability to perform the five areas listed previously, their ability to follow experiment procedures in a timely fashion, and the identification of critical sequences in the experiment operations that require special attention, which is partially a function of the PC's decision making abilities.

Once proficiency in the five basic areas has been acquired, two objectives become primary. The first deals with the PC's successful mastery of experiment procedures. The second deals with developing proficiency in the skills necessary to carry out the first objective. These "critical" skills would be above and beyond those skills mastered during training in the five basic areas. At present, however, these skills have not been identified. This condition exists for two reasons. The first is that the procedures for all the experiments have not been written, which precludes looking at them to identify critical sequences. The second reason is that skill identification has not been prioritized; that is, at present some feel there is questionable doubt as to whether such skills exist, and if they do, they would be developed when the procedures are learned and practiced during training.

At issue too is a philosophical question as to whether training, and hence mission operations, is simple or complex. Is it simple in that all the PC has to do is follow the written procedures which will caution for upcoming critical sequences. The notion here is that all the PC needs to do is develop very high proficiency at knowing and implementing the already spelled out procedures, and then relying on the POCC if questions arise. And while it may be true that the POCC may understand each experiment in more depth than the PC, the POCC has not necessarily developed the procedural knowledge to the degree that the PC has and may be of little substantive help to the PC in some unforseen situations.

The other view is that good scientific enquiry requires complex thinking skills. In this view the PC has the capability to conduct the experiment as he best sees fit, due to the fact that at that instant he possesses the most advanced conceptual understanding of the entire mission. Scientific enquiry demands sharp decision making skills, and as such these should be addressed during training at the PCTC. If this complex view is ascribed to, then training should afford direct or indirect opportunities to make decisions based on prior knowledge of and training in critical sequences or contingencies in each experiment.

Skill identification can start by asking and answering the following questions, pulled from Ely (1962):

"In the main, the definition of critical task characteristics is most effectively accomplished by considering the consequences of a particular activity if it were accomplished poorly or required excessive time for completion. Asking the following questions with respect to the various tasks is helpful for determining criticality.

Would below-minimum performance:

- o lead to an accident?
- o result directly in mission failure?
- o be impossible to remedy within time constraints or not at all?
- o be difficult to detect because of inadequate information feedback?
- o recur over time in such a way as to produce a cumulative effect?
- o contribute a large proportion of time to the total time required for some larger and critical function?

An affirmative reply to at least one of the above questions implies that the activity in question should be further considered in determining the critical behaviors for measurement."

In order to approach this issue, a skill matrix is being compiled from the available experiment procedures. Actually, two matrices are being developed. Both are modifications of existing skill charts (Rabideau, 1964; and Smode, Gruber, and Ely, 1962). Both are felt will yield useful information and will hopefully validate each other (the original charts and their modifications are in Appendix A). In the final evaluation, the skill matrices might very well show that training is in fact "simple" as described previously, but then at least there would be more support for that position.

Following skill identification, training procedures can be more fully articulated using the present training definitions. These definitions exist for each experiment and are described in three levels, from A to C. The descriptions progress from a general description (A), to typical off-nominals (B), and finally to step-by-step procedures (C). These skills should then be incorporated into the level A and B experiment descriptions. The level A, B, and C descriptions could then be utilized additionally by; 1) sending out the A and B descriptions as part of the PC's pre-arrival training materials, and 2) using all three levels during training to brief each experiment from and then to train from. "A" descriptions should reflect time of training, the skills addressed, and where the training will take place. "B" descriptions should identify the training personnel, the possible off-nominals, and critical experiment stages; i.e., where contingencies exist or where the experiment could be aborted or drastically effected via operator error. "C" descriptions represent the training guide the PC might follow during training. "C" descriptions list nominal and off-nominal procedures.

After initiation to the experiment during the pretraining briefing and exposure to experiment procedures, but before going to the mockup for practice, the PC should be able to demonstrate the following:

- What is the PC's knowledge of the experiment; communicated verbally?
- 2) What is the PC's knowledge of the possible off-nominals for this experiment and where will they occur?
- 3) What are the PC's perceptions of his anticipated responses to those off-nominals?
- 4) Can the PC verbally walk through the experiment's nominal run?

Evaluation criteria follow from skill identification and should be fully planned for prior to the implementation of training. At this point, four recommendations are made for training:

- 1) Pertinent PCTC trainers should go through the PC training program prior to any PC trainee to debug and improve on the system.
- 2) If the modified Ely chart for skill identification is validated, then his three instrument and criteria measures in Appendix B can be used to develop PCTC evaluation criteria and instruments.
- 3) Additionally, video tapes, PC record books, debriefing questionnaires, and data graphs of performance can be collected to trace training progress
- 4) Two additional types of measures have been located in our review that have application to the current evaluation needs. One is an effectiveness rating scale:

		Decision Speed		
Extremely Rapid	Very Fast	Acceptable	Slow	Very Slow

Decision Quality

			1	1
Outstanding	Superior	Adequate	Somewhat Adequate	Poor

The second is an assessment tool of procedural operations:

	<u>Task Step</u>	Error & Consequences
1.	Turn power switch to ON position	
2.	Read external temperature	
3.		
4.	etc.	
5.		

The final area addressed in this report is the matter of developing a training schedule. This can not be developed at present until the skill identification is completed. Additional parameters in formulating a schedule would be:

- 1) A need to determine the total time for experiment training, so that a training schedule can be developed.
- 2) How much time is needed for the dyads to work in each of the possible pairs.
- 3) Number of actual training locations within the PCTC for the PC to be trained at.

APPENDIX A

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Modified skill identification charts used to identify PC skills for training definition.

	#- +nevy 2 eax	001	002	003	50	208		5 I 3	14	16	2	0 19	20	021	22	23	74	2.7	29	33	3 4	00
	, Types of Tasks	0	0	0	9	2	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	~~~~
ļ	Follow Routine, Pre-established Checkout Sequence		_	_	4									<u> </u>			<u> </u>	\				
	Position a Set of Discrete/Serial Controls		_		-+-				$\left - \right $			-1						† —	<u>ii</u>	;i		
ł	Obtain Information From Dynamic Displays											{}						<u>ا</u>			{	
ļ	Adjust Instrument or Other Equipment Rem	-+			+					\vdash		\vdash					i	1	17			
	Track																	{	1			
	Communicate Information Utilize Status or Reference Data									-+		<u>}</u> −¦						<u> </u>	+;			
1	Utilize Status or Reference Data		-		- -							-+						┥	<u></u> 	·		
	Directly Observe External Events				-+-					-+		┢─╁					;					
	Compute				-+				-1			í-†	~~ †				·	i-	1		j	
	Select Action From Among Alternative Choices	\vdash	=		-4-							-+						1=	÷1.73	rere it		-745-1
	Follow Established Procedures in Inspection, Calibration, and/or Checkout.				_				-ine			-	-20	÷ .			~~~	ļ				
	Acquire Information From Among Multiple Sources	 			-+-		·					+		j				†				
.[Make Multiple Control Actions				-+-							i-t						1				
Ľ.	Locate Malfunctions in Complex Equipment											\vdash							1-1			
A aldinoo	Remedy, Repair, Fabricate, and/or Assemble Complex Equipment.				<u> </u>																	
	Curry Out Computational/Mensurement Sequences for Determining Status																		ļ_			
	Make Decisions Based on Multiple Information Sources				_						<u> </u>					•		-				
														•								
5					T			_						·								

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Modified from Smode, Gruber, and Ely, 1962

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	Processes	Activities	_ Specific Bebariors Expt. # (frequency)
			Detects
		•	Inspects
		_	Observes
		Searching for and Receiving	Reads
		Information	Receives
			Scans
	Perceptual Processes		Surveys
			Discriminates
		Identifying Objects, Actions, Events	- Identifies
			Locates
	•	•	
			Calculates Codes
		•	
	•	[Information Processing	Computes
			Interpolates
7	-		Itemizes Tabulates
10			
	Mediational Processes	_	Translates
Modified			Analyzes
ed		· ·	Calculates
ι.			Chooses
rom	·	Problem Solving and Decision Making	- Compares
ä			Computes
Ra		•	Estimates
Rabideau			Plans
id			[Advises
62			Answers
, n			Communicates
ب_ر		•	Directs
1964	Communication Processes -	·	- Indicates
4			Informs
			Instructs
		•	Requests
			Transmits
			La clivates
	•	•	Closes
			Connects
		Simple/Discrete	Disconnects
		· ·	Joins
	•		Moves
	Motor Processes		- Presses
			Sets
••••			Adjusts
			Aligns
		Complex Continuous	- Regulates

	(Based on	Sı	noc	łe,	G	rubc	r,	and	E	ly,	196	2)										
	Types of Tasks	Error Amplitude	Error Frequency	Error Content Analysis	Error Change Over Time	Number of Control Responses and Manipulations	Number of Communications	Number of Personnel Interactions	Number of Diagnostic Checks	Number of Errors	Number of Out-of-Tolerance Conditions	Number of Observations	Number of Verbal or Written Reports	Number of Requests for Information	Per cent of Activities Accomplished	Measures of Achieved Reliability	Measures of Achieved Maintainability	Equipment Failure Rate	Cumulative Response Output	Proficiency Test (Written) Scores	Magnitude Achieved-Steady- State Value	Magnitude Achieved- Changing Value
	Follow Routine, Pre-established Checkout Sequence			x			-		x	X	X				X					X		
	Position a Set of Discrete/Serial Controls	T	XX	XXX		X			-	X												
	Obtain Information From Dynamic Displays	X		X						X		X									X	X
	Adjust Instrument or Other Equipment Item	X		XXXXXXXXX		X				X												
믭	Track	X	X	X	X	X				$\overline{\mathbf{X}}$									X		X	X
Simple	Communicate Information			X			X	Х		X			X	X								
S	Utilize Status or Reference Data	X		X			_			X									-	X	X	X
	Directly Observe External Events	X	×××	X						X		X									X	X
	Compute	Ī	X	X						X										X		
	Select Action From Among Alternative Choices		X	\overline{X}						X										X		
1	Follow Established Procedures in Inspection, Calibration, and/or Checkout.	·	X						×	×	×				×					×		
	Acquire Information From Among Multiple Sources	X		X						X		X		X								X
	Make Multiple Control Actions	X	X	X	X	×				X												
oles	Locate Malfunctions in Complex Equipment			X					X	X	X					X	X			X		
Compley		×	×	X		×			×	×							×					
	Carry Out Computational/Measurement Sequences for Determining Status	×	×	×						×			×							×	X	×
•	Make Decisions Based on Multiple Information Sources	X	×	×						×										×		

Application of Task Data for Error, Frequency, and Achievement Measures • (Based on Smode, Gruber, and Ely, 1962)

TABLE 8-2B

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Application of Task Data for Latency and Accuracy	Mcasures
(Based on Smode, Gruber, and Ely, 1962)	

TABLE 8-2A

								_																			
	Types of Tasks	Times of Alexandres	Movement Reaction Time	Time to Initiate Correction	TimetoInitiateActivityFollow- ingCompletionofPriorActivity	S ad	Time to Identity Stimulus	Time to Complete Message. Decision, Control Adjustment	Time to Reach Criterion Value	Time Spent in Activity	Per Cent Time on Target	Accuracy in Identifying	Displays Internal to System Accuracy in Identifying	to 5,13	tance, Direction, Speed, Time	Detection of Change in Stimulus Over Time	Detection of Trend Based on Multiple Related Events	(Ibservation Correctness	Recognition of Signal in Noise	Recognition of Out-of- Tolerance Condition	Accuracy in Control Positioning or Tuol Usage	Accuracy in Reading Displays	Accuracy in Symbol-Usace, Decision-Makine, Computing			Trucking Accuracy	Communications Accuracy
	Follow Routine Pre-Established Checkout Sequence	Ι	Γ				\Box			X	5		रा			X		X		X	X	x			X		
	Position a Set of Discrete/Serial Controls							X		X		$\left \right\rangle$	<					Γ	\Box		X				X		
	Obtain Information From Dynamic Displays	\mathbf{b}	<				X			X	Τ	>	$\langle \rangle$	<u><</u>	X	Х	_	X	X	X		X					
	Adjust Instrument or Other Equipment Item							X	\times	×											Х						
ľ	Track		X	X					X	X	X				X	X					X				X	X	
Simple	Communicate Information							X		×																	X
°	Utilize Status or Reference Data									$\overline{\mathbf{X}}$																	
	Directly Observe External Events .	þ	<			X	X			×			>	<u><</u>	X	Х		X	X								
	Compute								X	X	Τ					_							Х				
	Select Action From Alternative Choices							X		X													Х	X			
	Follow Established Procedures in Inspection, Calibration and/or Checkout	T			×		Π	x	×	×		>	<					×		×	X	×	X		Γ		
	Acquire Information From Among Multiple Sources	Т	7-			X	Π			X			()	<		X	X	X	X			X	Х			Т	
	Make Multiple Control Actions	Ţ	Т	X	X	<u> </u>	П	X		x		T						Γ			X			Γ	X		
- E	Locate Malfunctions in Complex Equipment	>	<u> </u>	X						X	-1	77						X		X			Х		\Box	Π	
Complex		Ι					$\left[\right]$	X	×	×											×						
_	Carry out Computational/Measurement Sequences for Determining Status									×					×	X							×			Ţ	
	Make Decisions Based on Multiple Information Sources			×	×	×	\Box	×		×			Γ			_]	×			l			×	ž	!!	1	ļ

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TABLE 8-2C	
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Application of Task Data for Consumption, Physiological and Behavioral Judgment Measures (Based on Smode, Gruber, and Ely, 1962)

	Types of Tasks	Types of Measures	Fuel/Eaergy Conservation	Units Consumed in Activity Accomplishment	Rate of Consumption	Operator/Crew Physiological Condition	Operator/Crew Behavioral Condition	Rating of Operator/Crew Performance Adequacy	Rating of Task Performance Adequacy	Estimation of Amount or Degree of Behavior Displayed	Analysis of Operator/Crew Behavior Characteristics	Omission of Relevant Behavior	Occurrence of Nonrelevant Behavior	Description of Out-of- Tolerance Condition	Interview Content Analysis	Self-Report of Experiences	Peer, Self, or Supervisor Ratings
	Follow Routine Pre-established Check- out Sequence	·						×	×	×	×	X	×		×	X	×
	Position a Set of Discrete/Serial Controls							X	X	X	X	X	X		X	X	X
	Obtain Information From Dynamic Displays	_						X	X	X	X	X	X		X	X	X
	Adjust Instrument or Other Equipment Item							X	Х	X	X	X	X		X	X	X
Simple	Track							X	X	X	X	X	X		X	X	X
B	Communicate Information							X	X	X	X	X	X		X	X	X
<i>w</i>	Utilize Status or Reference Data							X	Х	X	Χ.	X	X		X	X	×
	Directly Observe External Events							X	. Х	X	X	X	X		X	X	×
	Compute							X	X	X	X	X	X		Х	X	X
	Select Action from Among Alternate Choices							X	X	X	X	X	X		X	X	X
	Follow Established Procedures in Inspection, Calibration, and/or Checkout							×	×	×	×	×	X		×	×	×
	Acquire Information from Among Multiple Sources.							×	×	×	×	Х	×		×	×	×
ĸ	Make Multiple Control Actions	_						<u> </u>	_X_	<u>×</u>	X	X	<u> </u>		X	X	X
ple	Locate Malfunctions in Complex Equipment	_						×	X	X	X	×	X	X	X	Х	×
Complex				×	X			×	×	×	×	×	×		×	×	×
	Carry Out Computional/Measurement Sequences for Determining Status							×	×	×	×	×	×]	×	×	<u>×</u>
	Make Decisions Based on Multiple Information Sources							×	×	X	_ ×	_×[<u>×</u>]		×Į	×Į	-]

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REFERENCES

- Cream, B. W., F. T. Eggemeier, and G. A. Klein. A strategy for the development of training devices. <u>Human Factors</u>, Vol. 20, April 1978, p. 145-158
- Grodsky, M. A. The use of full scale mission simulation for the assessment of complex operator performance. <u>Human Factors</u>, Vol. 9, p. 341-348, 1968
- Johnson, S. L. Retention and transfer of training on a procedural task interaction of training strategy and cognitive style. TLSP: Final Scientific Report, 1 October 1976 - 30 September 1977. 79N12704#
- Meister, David. <u>Human Factors in Operational System Testing</u>: <u>A Manual</u> <u>of Procedures</u>. San Diego: Navy Personnel Research & Development Center, 1978
- Meister, D. and G. Rabideau. <u>Human Factors Evaluation in System</u> Development. New York: John Wiley & Sons, Inc., 1965
- Rabideau, G. Field Measurement of Human Performance in Man/Machine Systems. Human Factors, Vol. 6, December 1964, p. 653-671
- Smode, A. F., A. Gruber, and J. H. Ely. <u>The Measurement of Advanced</u> <u>Flight Vehicle Crew Proficiency in Synthetic Ground Environments</u>. <u>Report MRL-TDR-62-2</u>, Behavioral Science Laboratory, Aerospace Medical Division, Wright-Patterson AFB, February 1962
- Tolcoh, M. A. Human Decision Making. <u>Human Factors Engineering Seminar</u>. Stanford, Conn.: Dunlap & Associates, Inc., 1960
- Wood, M. E. Improved crew member training through multi-media instruction. Proceedings from The Role of Man in Navigation, July 1-3, 1970, p. 477-485, 71A10524#

1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

SOME METHODS IN HIGH ENERGY COSMIC RAY MEASUREMENT

Prepared By:

Julian B. Shand, Jr., Ph.D.

Academic Rank:

Professor

University and Department:

NASA/MSFC (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Berry College Department of Physics

Space Sciences Space Physics Astrophysics

T. A. Parnell

August 1, 1980

NGT-01-002-099 (University of Alabama)

SOME METHODS IN HIGH ENERGY COSMIC RAY MEASUREMENT

By

Julian B. Shand, Jr., Ph.D. Professor of Physics Berry College Mt. Berry, Georgia

ABSTRACT

The response of ion chamber detectors to cosmic rays of energy a few GeV/nucleon and up shows a logarithmic rise with energy. The energy resolution has not been accurately defined. In Ref. 1 a calculation technique has been outlined which gives useful results for cosmic ray protons. The present study has applied the method to much heavier particles and higher energies.

In addition, a balloon flight will soon detect cosmic ray showers. A detector of acoustic waves produced by the showers will be carried, as will emulsion detectors. The present study has involved calibration of a flat plastic scintillation plate which will act as a trigger for the acoustic detector, and will determine the position of exit of the showers in order to expedite visual examination of the emulsions.

INTRODUCTION

Cosmic rays have been investigated since their discovery in 1911. Their flux in space is nearly independent of time and direction. They consist of nuclei, electrons, and electromagnetic radiation. The nuclei are mostly protons (around 90%) and helium (around 9%) but heavy nuclei are present also, being over represented as compared to solar abundances. The energy spectrum extends to the greatest energies known, reaching as much as 10^{20} eV or 1 joule.

The origin of cosmic rays is a mystery, as is their acceleration to ultra high energy. Exploding galaxies, supernovae, pulsars, white dwarfs, and other exotic types of objects have been proposed as sources, but none have been clearly identified.

Measurement of the energy, knowledge of the energy resolution of the measuring devices, and identification of the various nuclear species are important for an understanding of the phenomena. A logarithmic rise in ionization of gas in ion chamber detectors, as cosmic ray energy increases, has long been observed, for energy a few GeV/nucleon and up. The energy resolution has not been accurately defined. In Ref. 1 a procedure has been outlined for calculating resolution for the case of particles of charge 1. This type information is needed also for the heavy nuclei in cosmic rays, and work on this problem is described in the following section.

Flat plates of plastic scintillator can serve several functions in a balloon-borne cosmic ray experiment. Rays passing through a plate produce light in the plate, which can be collected and measured by photomultiplier tubes. The signals from these tubes can indicate the passage of a "shower" of particles produced by collision of a high-energy cosmic ray with a nucleus. These signals can thus be used as a trigger for another detector designed specifically for measuring the energy of such showers. In addition, the relative strengths of the signals from the photomultiplier tubes can be used to determine the place of exit of the particles. This information facilitates examination of nuclear emulsions located next to the plastic scintillator, in order to find tracks in the emulsion created by the particles passing through.

OBJECTIVES

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Two main objectives were pursued. The first was to calculate the average energy actually deposited in an ion chamber by an ultra-high energy particle of large mass and charge. The method was to extend a calculational scheme already applied successfully to particles of charge 1.

The second objective was to calibrate a plate of plastic scintillator for measurement of the position of a cosmic ray shower passing through it. The method of calibration was to inject pulses of light at known positions on the plate and record the responses of photomultiplier tubes at the corner of the plate.

BODY OF REPORT

On Relativistic Rise in Ionization

Relativistic charged particles passing through a gasfilled chamber lose energy mainly by excitation and ionization of the gas atoms (Ref. 4). Experiment shows a slow logarithmic increase in energy loss as particle energy increases. This is called the relativistic rise. The rise finally stops and energy loss becomes constant in what is called the Fermi plateau. Examination of the relativistic rise and knowledge of the energy resolution can allow one to determine the energy of the particle. This requires that charge be measured independently by other detectors through which the high energy particles pass. Typically a Cerenkov counter utilizing solid radiators such as lucite plastic is used for charge identification in the ultrarelativistic energy range. Thus for cosmic ray work the particular kind of nucleus striking the detector can be found (Ref. 5).

One way to look at the interaction between an incident particle and the electrons of an absorber is by representing the electromagnetic field of the particle by two terms (see Fano, Ref. 2). One, called the longitudinal term, is merely a static Coulomb interaction. The second term, called the transverse term, involves emission and reabsorption of virtual photons. Fano has treated the interaction in the lowest order (Born) relativistic approximation. The longitudinal interaction is effectively constant at high velocities. The transverse term, however, increases in strength as velocity increases. This is expected since the transverse field increases proportional to γ , where

$$\gamma = (1 - \beta^2)^{-\frac{1}{2}}$$
 and $\beta = v/c$.

This relativistic rise eventually stops because of polarization of the absorbing medium.

Fano found expressions for differential cross sections for ionization. By integrating over energy he found energy loss. This approach does not agree well with experiment for the case of thin gas absorbers (Ref. 1,3) because the energy loss occurs in a relatively few discrete collisions. Cobb et al. (Ref. 1) have used Fano's results up to a point but instead of integrating over energy have used Monte Carlo techniques and calculated energy losses in discrete collisions many times to get an average energy loss, a spectrum of energy loss, and the amount of relativistic rise. The results are in good agreement with experiment for incident particles of charge 1. This author has adapted Cobb's method, which came from Fano, to the case of particles of large charge. Several difficulties were encountered:

A. Cobb's use of the letter E in expressions was thought to be ambiguous: it could mean energy transfer to an electron, which would include the binding energy of the electron, as Fano seems to use it: or it could mean the kinetic energy of the electron after collision, as is suggested by Cobb's referring to the Rutherford cross-section as being

$$\frac{2\pi}{\mathrm{mv}^2} \frac{1}{\mathrm{E}^2}$$

In this expression Rutherford did mean E as kinetic energy of the electron.

B. Cobb's expression for the longitudinal differential cross-section for ionization included the factor (with units 1/energy)

$$\delta(E-E_i)$$

where E; is the electron binding energy. This has the form of the Dirac delta function, which has meaning when integrated. Since integration is not used in Cobb's method, it is not clear how to use it.

C. Cobb's expression for the dielectric constant ε of the absorber gas contains the frequency ω . This seems ambiguous: earlier in his paper the same variable had been used as the frequency of a virtual photon. However, Fano wrote that the ω in the dielectric constant expression is a root of the equation

$$c^2q^2-\varepsilon(\omega)=o$$

where c is the speed of light and q is momentum transfer to the electron. Since Cobb's treatment follows Fano, an uncertainty in interpretation is introduced.

Consultation with some experts on energy loss and a search of the literature have shed little light on these difficulties. Communication with the authors (Fano, Cobb et al.) will be attempted to resolve the ambiguities. This author has written computer code to calculate cross sections for ionization, using various combinations of interpretations, in order to be able to distinguish the correct ones. The work is still in progress.

Calibrating A Plastic Scintillator for Position Determination

In September 1980, a balloon flight will carry several cosmic ray experiments. Events of the highest energy will be of particular interest. Nuclear emulsions, X-ray film, and calorimetry will measure characteristics of these events. A plastic scintillator can be configured to determine the position of exit of cosmic rays. The rays cause the plastic to emit light. Most of the light is trapped in the thin plate of plastic by total internal reflection. When photomultiplier tubes are optically coupled to the corners of the plate, the tubes measure the amount of light reaching them. The amount varies inversely with the distance to the event, roughly. Several factors complicate the tube response, including edge effects, scratches, crazing of the plastic, inhomogeneity of the scintillating material, and probably others.

Preliminary calibration of the scintillator was done by this author. The method required a light-tight box. The scintillator plate, which measured 44" by 44" by 1/4", lay in the box on white paper which had a grid of horizontal and vertical lines drawn on it, with 10 cm spacing. A photomultiplier tube was attached with optical coupling grease to a specified position at one corner of the plastic plate. The photomultiplier tube was a 3", ten stage EMI, type 9708KB. The output of the tube was amplified and referred to a multichannel analyzer which displayed pulse heights and allowed a determination of average pulse height.

Light pulses were obtained by sending electric pulses at frequency 10Khz to a light-emitting diode (LED). The LED was mounted at the top of a shiny brass tube whose open end fit into a clear plastic disk. The disk, with LED, could be stuck to the plastic scintillator with optical grease and accurately located above an intersection of grid lines on the paper underneath. Thus positions of photomultiplier tube and LED were known to about ½ cm.

After pulses had been injected for 90 seconds and the average pulse height ascertained, the disk with the LED was removed and the plate wiped with dry cotton material and then with the same kind of material moistened with pure ethyl alcohol. All other procedures tried left some residue of film which scattered light out of the plate. After the cleaning, the LED disk was regreased and repositioned. In this way a map was generated. Upon completion, the photomultiplier tube was moved to another corner and another map generated. Only one photomultiplier tube was ready at the time, so 3 other "dead" tubes were kept at the 3 other corners to absorb light as would happen in the final setup.

Checks were made of the reproducibility of the reading from a particular point. Deviations ranged from 0 to 12%, averaging about 5%.

A sample of experimental results is shown here. The positions refer to a Cartesian coordinate with the plastic scintillator plate defining the first quadrant, and the photomultiplier tube centered at x = 5.9 cm, y = 5.9 cm. Pulse heights along an edge and along an approximate diagonal are given. Positions are not integral multiples of 10 cm because visualization is easier than in the system actually used.

Position (cm) Position х У Pulse height y Pulse height х 21.8 10 21.8 20 1640 2740 30 10 1260 31.8 720 31.8 740 41.8 40 410 41.8 10 50 51.8 10 450 51.8 280 61.8 10 300 61.8 60 190 240 71.8 70 180 71.8 10 81.8 10 180 81.8 80 160 91.8 90 130 91.8 10 130 101.8 100 120 101.8 10 110

At several positions in the scintillator pulse heights had local maxima or minima. Some of these could be identified with defects mentioned in the first paragraph under calibrating.

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CONCLUSIONS AND RECOMMENDATIONS

Concerning relativistic rise in ionization, an attempt should be made to locate the authors of Ref. 1 and correspond with them to resolve the difficulties of interpretation of their paper. In addition, the computer trial and error attempts to find the correct interpretation should be continued in case all else fails.

Calibration of the plastic scintillator as a determiner of position has been done. The experimental error indicates a positional error of nearly 5 cm can be expected. Since the calibration was done with light injected from outside, while the actual use will have light generated inside the plastic, a check of the numbers should be made with all 4 photomultiplier tubes in place, mounted in final form. Energetic radiation which will cause the plastic to scintillate should be injected from outside.

REFERENCES

- Cobb, J. H., W. W. M. Allison and J. N. Bunch, "The Ionization Loss of Relativistic Charged Particles in Thin Gas Samples and Its Use for Particle Identification," Nuclear Instruments and Methods 133, 1976.
- 2. U. Fano, "Penetration of Protons, α Particles, Mesons," Annual Review of Nuclear Science 13, 1963.
- Crispin, A., and G. N. Fowler, "Density Effect in the Ionization Energy Loss of Fast Charged Particles in Matter," Reviews of Modern Physics, Vol. 12, No. 3, July 1970.
- Jackson, J. D., Classical Electrodynamics, Wiley, New York, N. Y., 1962.
- 5. Gregory, J. C., and T. A. Parnell, "A Measurement of the Relativistic Rise in Xenon Filled Ionization Chambers for Cosmic Ray Iron," <u>Papers of the Sixteenth International</u> Cosmic Ray Conference, Vol. 12, 1979.

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1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

PHOTOCHEMISTRY OF MONODENTATE AND BIDENTATE CARBONATO COMPLEXES OF RHODIUM (III)

Prepared By: Academic Rank: Assistant Professor University and Department: Colgate University Department of Chemistry NASA/MSFC: (Laboratory) (Division) (Branch) _ _ _ Robert J. Naumann MSFC Counterpart: Date: August 2, 1980

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Peter S. Sheridan

by

Peter S. Sheridan Department of Chemistry Colgate University Hamilton, New York 13446

ABSTRACT

A scheme for the photochemical fixation of water (eq 1) has been proposed.

$$2 \quad H_2 0 \quad \xrightarrow{h \ v} \quad 2H_2 \quad + \quad 0_2 \tag{1}$$

Fuel cells which generate electrical energy using the reverse of eq 1 are used in NASA spacecraft, so the solar-induced regeneration of H_2 and O_2 would allow the same materials to be recycled through the fuel cells, dramatically increasing energy output per pound of fuel carried aloft.

The proposed scheme involved a 5-step reaction sequence; the first step involves the 2-electron reduction of a metal by a coordinated carbonate ligand, with corresponding oxidation of the carbonate to CO_2 and O_2 .

Ligand field photolysis of trans- $[Rh(en)_2 (H_20)(CO_3)] ClO_4 (I)$, and $[Rh(en)_2 (CO_3)] ClO_4 (II)$ have been studied in the solid state and in aqueous solution at various pH values. Both salts are photoinert in the solid phase, but are quite photoreactive in aqueous solution. In solution, the monodentate ion (I) undergoes efficient isomerization to a mixture of cis and trans - $[Rh(en)_2 (H_20)(CO_3)]^+$, presumably with water exchange. A minor pH increase upon photolysis is evidence of inefficient carbonate $(CO_3 =)$ release, with formation of $[Rh(en)_2 (H_20)_2]^{-3+}$.

In contrast, aqueous solutions of the bidentate carbonato complex, (II) undergo efficient pH decrease upon ligand field photolysis. Changes in the electronic spectrum (200-500 nm) and pH changes indicate that the desired redox is occurring (eq 2)

$$\left[\operatorname{Rh}(\operatorname{en})_{2} (\operatorname{CO}_{3}) \right]^{+} \xrightarrow{\operatorname{hv}}_{\operatorname{H}_{2}0} \longrightarrow \left[\operatorname{Rh}(\operatorname{en})_{2} \right]^{+} + \operatorname{CO}_{2} + \frac{1}{2} \operatorname{O}_{2} (2)$$

The pH increase is due to the aqueous behavior of CO_2 (eq 3).

$$CO_2 + H_2O \longrightarrow H^+ + HCO_3$$
 (3)

A positive test for released carbonate ion supports this assignment. Work is proceeding to confirm this new photochemical reaction, with emphasis on gaining understanding of the exact nature of the rhodium photoproduct, the details of the photochemical mechanism, and to test the generality of photoredox reactions of bidentate carbonato complexes.

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NOMENCLATURE

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LMCT	-	ligand-to-metal-charge-transfer				
en	-	ethylenediamine $(H_2NCH_2CH_2NH_2)$				
K	-	equilibrium constant (K = 4.3×10^{-7} for hydrolysis of CO ₂)				
		$CO_2 + H_2O \rightleftharpoons H^+ + HCO_3^-$				
h0	-	implies the reaction conditions included the incidence of light				
nm	-	nanometer				
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INTROPHCTION

The photochemical behavior of Rh(III) complexes has attracted considerable interest¹, as a variety of photoinduced aquations and stereochemical rearrangements have been observed. While exceptions exist, Rh(III) amine complexes characteristically undergo efficient photoinduced ligand loss/aquation, with rearrangement to a symmetric trans configuration. There is little evidence of any photoinduced reduction of the Rh(III) center, even upon photolysis of LMCT bands in such complexes as $[Rh(NH_3)_5I]^{2+2}$; the instability of the Rh(II) oxidation state and the photoreactivity of the ligand field excited states lead to a predominantly ligand field excitedstate chemistry for Rh(III) amines.

Our work on the photochemistry of rhodium complexes has led to a reaction sequence which, on paper, represents the photochemical cleavage of water into H_2 and O_2 . It is a five step reaction sequence (eq4-8); the net reaction is the desired water fixation.

- $L_{4}Rh(H_{2}0)(CO_{3})^{+} \longrightarrow L_{4}Rh^{+} + CO_{2} + \frac{1}{2}O_{2} + H_{2}O \qquad (4)$ $L_{4}Rh^{+}H_{3}O + \longrightarrow L_{4}Rh(H)(H_{2}O)^{2+} \qquad (5)$
- $L_4Rh(H)$ (H₂0) ²⁺ + H₂0 \longrightarrow $L_4Rh(H_20)$ (0H) ²⁺ +H⁺ + H⁻ (6)
- $H^- + H^+ \longrightarrow H_2$ (7)

$$L_4Rh(H_20) (0H)^{2+} + CO_2 \longrightarrow L_4(Rh(H_20) (CO_3)^{+} + H^{+}$$
 (8)

net reaction
$$H_20 \longrightarrow H_2 + \frac{1}{2} \frac{0}{2}$$
 (1)

The first step is the photoinduced oxidation of a d^6 metal center (Rh(III) here) by a carbonate ligand, generating 0_2 , $C0_2$ and a coordinatively unsaturated d^8 complex (eq 4). Oxidative-addition of H_30^+ then regenerates the Rh(III) center, creating a hydride ligand (eq 5). Photoinduced (or thermal) labilization of this pseudohalide ligand (eq 6) would lead to instantaneous H₂ generation (eq 7). Under the proper pH conditions, reaction of the photogenerated CO₂ (eq 1) and the aquohydroxo complex (eq 8) would regenerate the carbonato starting material (eq 8).

This reaction sequence has many advantages over other currently studied water-photolysis schemes:

a) Each of the steps is either a well-known reaction (eq $_7$ and $_8$) or is a reaction for which closely analogous reactions have been studied (eq 4, 5, and $_6$).

b) No radical intermediates are involved, as the 2-electron transfer of the $d^6 \text{-} d^8$ transition metal system allows complementarity.

c) The scheme is a self-contained, catalytic cycle, with only water consumed and H_2 and 0_2 generated.

d) The scheme is quite flexible, as many variables can be adjusted.

A major difficulty with this scheme is the first step, as there is little precedence for such a photoinduced two-electron transfer. An additional problem is that reports on the photochemistry of carbonato complexes are surprisingly rare.

An apparent two-electron transfer was observed upon irradiation of alcohol solutions of some carbonato complexes of $Pt(II)^3$. In the presence of nucleo-philes (X = C₂H₄, P ϕ_3 , CO, ϕ C \equiv C ϕ) photolysis of $PtL_2(CO_3)$ (L = P ϕ_3 , As ϕ_3 or P(CH₃) ϕ_2) generates PtL_2X_n (N = 2 for X = CO; n = 1 for bulkier ligands), consistent with the in situ formation of PtL_2 . Carbon dioxide was released during the photolysis (one mole per mole of Pt) and alcohols added to the photolysis solutions could be oxidized to aldehydes and ketones, confirming that the carbonato ligand was the two-electron reducing agent.

Such apparent two-electron transfers have also been observed upon photolysis of some oxalato complexes. Some $PtL_2(ox)$ complexes, similar to the carbonato complexes just discussed, photolyze to give CO_2 and similar Pt(0) photoproducts.⁴ Irradiation of $Pd(NH_3)_2(ox)$ was reported⁵ to generate CO_2 and a reactive species identified as $Pd(NH_3)_2$, which decomposed to give Pd black. Poulenc⁶ noted that the yellow $Rh(py)_3(ox)Cl$ solid turned red in sunlight, and Gillard et. $a1^7$.have shown the reaction involves another photoinduced two-electron transfer

 $Rh(py)_3(ox)C1 \longrightarrow Rh(py)_3C1 + 2 CO_2$

Work at the Space Sciences Laboratory this summer focused on this first step, the photolysis of Rh(III) carbonato complexes. Two complexes were chosed for study - trans - $[Rh(en)_2 (H_20) (CO_3)]^+$ and $[Rh(en)_2 (CO_3)]^+$ - as they allowed a study of the affect of monodentate vs. bidentate coordination of the carbonate ion on the resulting photochemistry.

OBJECTIVES

Research was directed toward an understanding of the following questions:

- 1. What is the photochemical behavior of Rh(III)-carbonato complexes?
- 2. Do the charge-transfer and ligand field states of these Rh(III) complexes display distinct photochemical reactivity, or does intersystem crossing lead to a common photoreactive state?
- 3. What affect does monodentate vs. bidentate coordination of the carbonato ligand have on the resulting photochemistry?
- 4. Are such complexes photosensitive in the solid state, or is an aqueous solution necessary for photoreactivity? In solution, does the pH affect the photochemical reaction.

The objective of the work was to determine if photoinduced two-electron reduction of a Rh(III) center could be induced by photolysis of the appropriate Rh(III)-carbonato complex.

EXPERIMENTAL

A. Preparation of Complexes

trans-Aquochlorobis(ethylenediamine)rhodium(III) Perchlorate

The trans- $[Rh(en)_2 (H_20) (CO_3)]^+$ ion was prepared by the method of Rerek⁸. A sample of trans- $[Rh(en)_2 (C1)_2] NO_3 .H_20^9$ was anated with a five fold excess of NaN₃ (under reflux) and the desired trans- $[Rh(en)_2 (N_3)_2]^+$ was crystallized as the BF₄ salT upon addition of NaBF₄. This diazido ion was dissolved in dilute HClO₄, and NaNO₂ was added; this generated NO⁺, destroying the coordinated azide.

 $N_3 \rightarrow N_2^0(g) + N_2(g)$

At a neutral pH (adjusted with LiOH) the resulting trans- $[Rh(en)_2 (OH)(H_20)]^{2+}$ was precipitated as the perchlorate salt. A solution of this ion was acidified at pH 1 with HClO₄, and NaHCO₃ slowly added until a pH of 5.80 was reached. The perchlorate salt was precipitated by cooling in ice, followed by dropwise addition of ethanol. The absorption spectrum had λ_{max} at 347 nm ($\epsilon = 105$), consistent with published data.

Carbonatobis(ethylenediamine)rhodium(III) Perchlorate

 $(Rh(en)_2CO_3] C1O_4$ was prepared by the method of Harris et.al.¹⁰ It displayed a single peak in the electronic spectrum at 327 nm (ε =250); this absorption coefficient is less than the literature value, but any impurities were shown not to interfere in the photochemistry. There were no detectable changes in the electronic absorption spectrum of this ion in solutions between pH 1 and pH 13, clearly ruling out significant contamination of either of the monodentate ions, <u>cis</u> or <u>trans</u>- (Rh(en)₂ (H₂O) CO₃)⁺.

Barium chloride solution

Aqueous solutions of barium chloride were prepared by dissolving $BaCO_3(s)$ in 1 M HCl_(aq) until effervescence ceased, and a pH of 7 (+1) was reached. The resulting solution was filtered and kept in a sealed flask to avoid CO_2 contamination. It was boiled for ca. 20 minutes before use.

Other reagents were used as supplied by the manufacturers.

B. Photolyses

Photolyses were performed on either solid samples of aqueous solutions of the complexes. Solid samples were finely ground with mortar and pestle, and then stuck on a piece of green tape (1 cm x 3 cm). This tape was suspended in a 1 cm fused silica cell (Beckman), connected to a microvolumeter (vide infra).

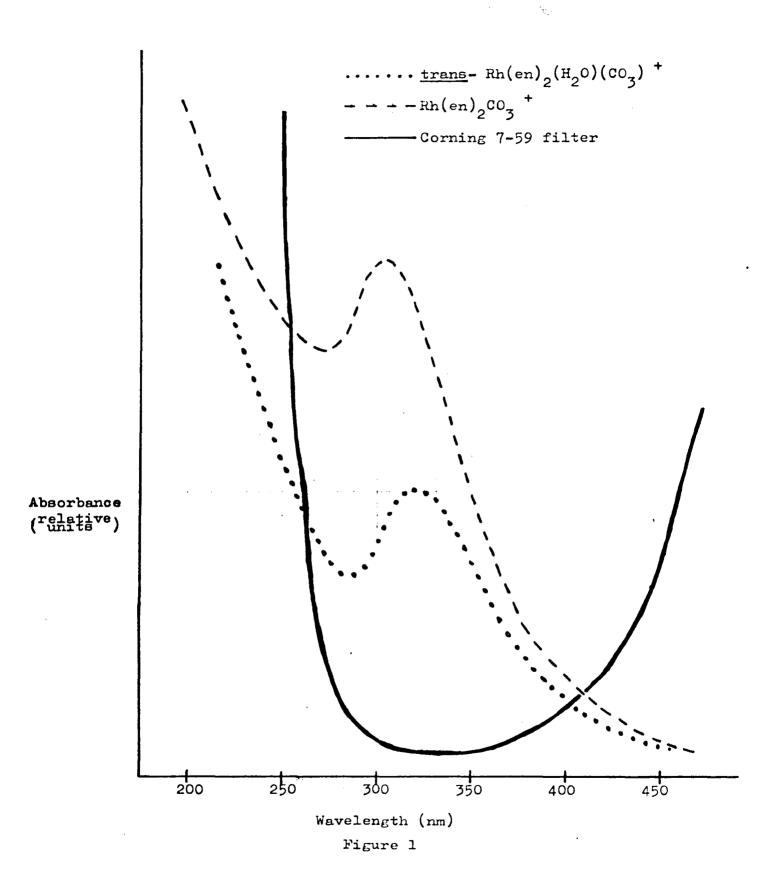
For the solution studies, a sample of the desired sample was dissolved in the appropriate aqueous solution at room temperature, and placed in a 1 cm fused silica cell (Beckman). Electronic spectra were recorded with the sample in the same cell.

C. Equipment

The irradiation source was a 1000 watt Mercury-Xenon Lamp (Hanovia Manufacturing Corporation) with compatible power supply and starter (Schoeffel LPS 255 power supply and Model 359 igniter). Photolyses were performed at 28.5 amps at ²² V. Use of the compatible Schoeffel high-intensity monochromator did not allow sufficient intensity to pass to the sample, so wavelength selection was affected with a Corning colored glass filter (C-7-59). This filter effectively absorbed all radiation below 280 nm, but allowed irradiation of the ligand field bands. The absorption spectra of the C-7-59 filter along with the spectra of the two complexes studied is given in Fig. 1.

Electronic spectra were recorded on a Beckman Acta CII spectrophotometer with an external 10 inch recorder (Beckman Model 1005). The pH measurements were done with a PHM 64 Research pH Meter (Radiometer/Copenhagen) with a GK 2321C Combined Electrode. Samples averaged about 3.0 ml for pH measurements.

A microvolumeter was prepared after a design of $tevenson^{10}$. A 1 ml graduated pipet (graduated in 0.01 ml units) was used, and the evolution of a gas was detected by monitoring the flow of water bubbles down the pipet. Photolysis of solid samples of $[Rh(en)_2(N_3)_2]BF_4$, which releases N₂, was used to check the sensitivity of the microvolumeter.



A. <u>Trans-</u> $[Rh(en)_2 (H_20) (CO_3)]^+$

1. Results

In the solid state, $trans-[Rh(en)_2 (CO_3) (H_2O)] ClO_4$ is photoinert. Prolonged photolyses (2 hrs.) with both filtered and unfiltered light did not cause any detectable release of gas. When the photolyzed solid was dissolved in water, it had an electronic spectrum indistinguishable from that of an unphotolyzed sample.

In aqueous solution, and in the absence of light, $(Rh(en)_2 (H_20)(CO_3)]^+$ is relatively unreactive. At room temperature and at a solution pH between 6 and 10, there is a minor pH increase and a minor increase in the intensity of the ligand field band. This reaction is presumably the thermal aquation of the carbonate ion, but it was slow enough to have no measurable impact on the photochemical reactions.

Upon photolysis, the electronic spectrum undergoes pronounced changes (Fig. 2). Note that photolysis causes:

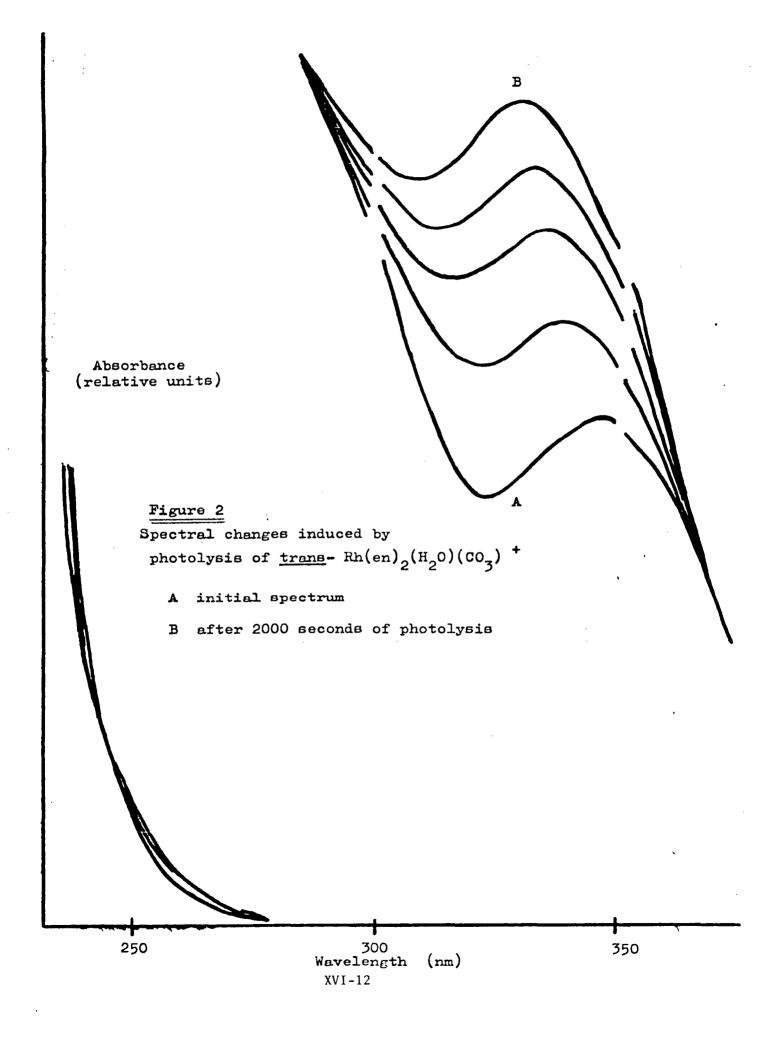
- a) an increase in intensity of the ligand field band,
- b) a blue shift of the peak maximum from 347 nm to about 332 nm,
- c) virtually no shift in the position of the low-energy edge of the charge-transfer band.

There appears to be an isosbestic point near 242 nm, although the spectra are nearly parallel at that point, so the presence of an isobestic is difficult to determine.

In addition to the spectral changes observed in Fig 2., photolysis causes the release of a base into solution. Data from a typical experiment are given in Table 1, where the pH changes observed in an unphotolyzed blank and photolyzed sample are listed. (The pH changes in Table 1 were recorded from the same experiment monitored in Fig 2.).

After extended photolysis, only minor spectral changes could be photoinduced. To determine the isomeric composition of this solution, 2 ml of conc HCl was added to a 3 ml sample of photoproduct, and refluxed for 1 hour. The resulting solution had an electronic spectrum of a mixture of <u>cis</u> and <u>trans</u> - $[Rh(en)_2Cl_2]^+$ (<u>ca</u>. 63% cis and 37% trans).

The specific experiment described in Fig 2 and Table 1 was done with unfiltered light, but the same results were obtained using filtered light. The reaction was approximately 5 times more rapid using the more intense unfiltered irradiation, but the nature of the spectral and pH changes were independent of excitation wavelength.



total		
photolysis time	sample	blank
(seconds)	рН	рН
Initial	7.032	7.060
100	7.110	7.070
200	7.184	7.081
300	7.245	7.096
450	7.325	7.129
600	7.376	7.136
800	7.435	7.151
1000	7.470	7.184
1300	7.504	7.216
2000	7.540	7.220

TABLE 1

Typical pH changes induced by Photolysis of Trans- $Rh(en)_2$ (H₂0)(CO₃) +

TABLE 2

total photolysis time (seconds)	samp1e pH	blank pH
Initial	9.905	9.916
100	9.411	9.862
200	8.557	9.789
300	7.729	9.721
400	7.375	9.616
550	7.120	9.508
705	7.002	9.380
900	6.911	9.277
1300	6.691	9.185
2200	6.495	9.086

Typical pH changes induced by Photolysis of $Rh(en)_2$ (CO₃) +

The available results force the conclusion that photolysis of trans- $[Rh(en)_2 (CO_3)(H_20)]^+$ causes isomerization (eq⁹).

 $\frac{\text{trans-}[Rh(en)_2 (H_20) (CO_3)]}{Rh(en)_2 (H_20) (CO_3)]} \stackrel{+}{\text{ion has an absorption maximum at 332 nm, at}} (9)$ The <u>cis-</u> Rh(en)₂ (H₂0) (CO₃)] + ion has an absorption maximum at 332 nm, at a significantly higher absorption coefficient than the <u>trans</u> isomer, consistent with the blue shift and increased absorption observed upon photolysis. The lack of change in the charge-transfer edge is also significant, as the intense charge-transfer peak is a LMCT band with carbonate ligand as the donor; the LMCT bands for the analogous [Rh(en)₂ (H₂0)₂] + ions begin at significantly higher energies. As photolysis causes no shift in that edge, it is clear that significant amounts of carbonate ion are not labilized during the photolysis.

The pH changes are not consistent with this proposed reaction, however, as the reaction of eq 9 should not lead to any pH changes. Calculation of the amount of OH⁻ released after extensive photolysis shows that only a very minor amount of the Rh complex is releasing base to the solution. For example, during the experiment described in Fig 2 and Table 1, 1000 seconds of photolysis causes spectral changes indicating that about 80% of the starting material has undergone photoinduced isomerization. After those same 1000 sec., only about .0026% of the starting material has released an OH⁻ into solution. We conclude that any pH changes are due to a minor side reaction which is not related to the dominant photoisomerization. The pH increase is apparent as soon as photolysis begins, suggesting that whatever reaction causes the pH increase, it is a primary photochemical process. The rate of pH increase remains relatively constant throughout the photolysis, even when little starting material remains to be photolyzed; this implies that there is also an inefficient base-releasing secondary photochemical reaction. Whether these pH changes are due to inefficient release of carbonate ion or an ethylenedimine chelate is not known at this time.

Further support for the proposed photoisomerization comes from the anation studies. Refluxing the photoproduct in HCl leads to the acid-catalyzed aquation of the carbonate ion, which would be followed by chloride anation of the diaquo ions. Since hydrolysis of the carbonate ion occurs without metal-oxygen bond cleavage, and anations at Rh(III) centers are also known to be stereoretentive, the composition of the resulting mixture of <u>cis</u> and <u>trans-</u> [Rh(en)₂ Cl₂] + would reflect the stereochemical composition of the <u>cis</u> and trans- [Kh(en)₂ (H₂0) (CO₃)] + photoproduct mixture.

The best value for this composition - 63% cis and 37% trans-must be considered a preliminary result; further work will be done to determine whether photolysis of trans- $[Rh(en)_2 (H_20) (CO_3)]^+$ generates a photostationary state of that composition. Such a result, which available information supports, would imply that the dominate photochemical reaction of cis- $[Rh(en)_2 (H_20) (CO_3)]^+$ is isomerization back to the trans_isomer.

The reactions described above occurred upon irradiation with filtered (ligand field) and unfiltered (ligand field and charge-transfer) light. This suggests a single photoreactive state; since no evidence of photoinduced redox behavior was observed, a ligand field photoreactive state is indicated.

(Presumably, photoinduced water exchange accompanies isomerization, although the necessary isotope exchange studies have not been done to verify this assumption). There is no evidence of CO_2 release, which would suggest a photoredox process, even upon excitation of the LMCT bands. The rate of photoreaction was markedly increased when the LMCT bands were irradiated, implying that intersystem crossing, from the initially populated charge-transfer states to the photoreactive ligand field state, is efficient.

In summary, the photochemical behavior of the monodentate carbonate complexes is reminiscent of the behavior of analogous haloamine complexes of Rh(III). The LMCT states are photoinert, while the ligand field states are photoreactive, leading to ligand labilization and aquation, with concomitant stereochemical rearrangement about the Rh(III) center.

B. $[(Rh(en)_2 (CO_3)]^+$

1. Results

In the solid state, $[Rh(en)_2 (CO_3)]$ ClO₄ is photoinert. Prolonged photolysis (2 hrs) with both filtered and unfiltered light did not cause any detectable release of gas, and when a sample so photolyzed was dissolved in water, it had an absorption spectrum indistinguishable from that of an unphotolyzed sample.

In aqueous solution, and in the absence of light, $[Rh(en)_2 (CO_3)]^+$ is very inert. At room temperature and at a solution pH between 6 and 10, there was no detectable change in the absorption spectrum for over 48 hours. No pH changes, beyond those observed in the blank due to absorption of atmospheric CO₂, were observed in aqueous solutions of $[Rh(en)_2 (CO_3)]^+$.

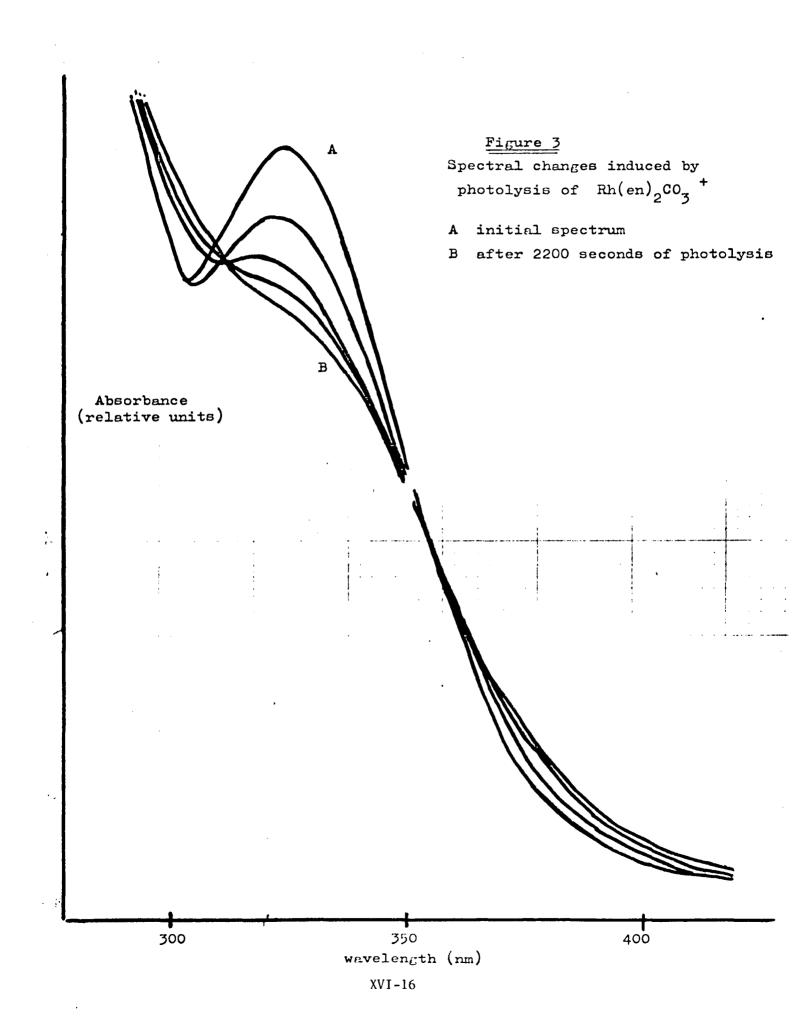
Upon photolysis, the electronic spectrum undergoes pronounced changes (Fig. 3). Note that photolysis causes:

- a) a decrease in intensity of the ligand field band, with virtually no change in wavelength maximum
- b) a minor shift in the low-energy edge of the charge-transfer band to higher energy.

Isosbestic points are not observed in the spectrum during the initial stages of the photolysis, but an isosbestic point does develop at about 314 nm after the initial stages of the photolysis are complete.

In addition to the spectral changes observed in Fig 3, photolysis causes a marked decrease in the solution pH. Data from a typical experiment are given in Table 2. The pH changes recorded in Table 2 were observed from the same experiment monitored in Fig 3.

The presence of photoproduced carbonate ion in solution was checked by photolyzing a sample of $[Rh(en)_2 (CO_3)]^+$ in a solution containing Ba² +



ions; formation of $BaCO_{3(S)}$ would signal the presence of CO_3 in solution. (A stream of $N_{2(g)}$ was bubbled through a slightly acidic solution of $[Rh(en)_2 - CO_3]^+$ for 30 min. This solution was made basic (pH 10) with $LiOH_{(S)}$, and bubbling was continued for another 15 min. This solution volume was doubled with the addition of a freshly boiled $BaCl_2$ solution (0.5 M), with N_2 bubbling continued throughout these manipulations. Any $BaCO_{3(S)}$ which formed was allowed to precipitate and a clear solution was put into a closed, fused silica cell for photolysis. The formation of a precipitate was monitored visually, or by recording the absorption.) Photolysis with unfiltered light rapidly led to precipitate formation (presumably $BaCO_{3(S)}$), which did not occur in a blank sample held in the dark. Photolysis with the filtered light did not lead to immediate precipitate formation, but prolonged photolysis did lead to a measurable cloudiness. We attribute this to the less efficient formation of $BaCO_{3(S)}$.

In an effort to determine the stereochemical composition of the photoproduct, a sample of $[Rh(en)_2CO_3]^+$ was photolyzed with unfiltered radiation for 3200 sec., and the photoproduct solution was then diluted with an equal volume of conc HC1. The solution immediately became yellow, and displayed a broad absorption band with local maxima at about 355 nm and 325 nm. Refluxing this solution for about 1 hour led to a yellow solution without a peak at 355 nm, but a more intense, clean peak centered at 331 nm; this peak had an absorption coefficient of about 3 x 10 ${}^{3}M {}^{-1}cm^{-1}$.

2. Discussion

The photochemical reactivity of $[Rh(en)_2 CO_3]^+$ has not yet been clearly elucidated; additional work is necessary to fully explain the phenomena observed. The following hypothesis, therefore must still be considered tentative.

Direct labilization of carbonate or ethylenediamine ligands can be ruled out as primary photochemical processes, as such reactions would lead to pH increases, not the observed pH decrease. Photoinduced isomerization, as observed for the monodentate complex (vide supra), may be occurring, but cannot account for the dramatic pH decrease; in addition, the photoinduced spectral changes do not support the idea of an isomerization to some other Rh(III) tetramine complex. The observed increase in acidity, along with the detection of free $CO_3^{=}$ in solution, strongly implies that CO_2 is a primary photoproduct. The H⁺ (aq) and $CO_3^{=}$ (aq) result from the rapid hydrolysis of the acidic oxide.

 $CO_2(aq) + H_2O \longrightarrow H^+ + HCO_3^- \longrightarrow 2H^+ + CO_3^+$

Independent detection of H^+ and CO_3^- in the photolyzed solutions supports the supposition of photogeneration of CO_2 .

Two possible routes for CO_2 release seem likely:

- a) Decarboxylation of coordinated CO_3^{-1} to generate CO_2 and OH^{-1}
- b) oxidation of coordinated $CO_3 = to$ give CO_2 and $\frac{1}{2}O_2^2$. In case a), observed for the thermal reaction of the analogous CO(III) complex, the following reaction sequence would be anticipated:

$$[Rh(en)_{2}CO_{3}]^{+} + H^{+} \longrightarrow |Rh(en)_{2}CO_{3}H]^{2+} H_{2}O + [Rh(en)_{2}CO_{3}H]^{2+} \longrightarrow |Rh(en)_{2}(H_{2}O)CO_{3}H]^{2+} [Rh(en)_{2}(H_{2}O)CO_{3}H]^{2+} \longrightarrow |Rh(en)_{2}(H_{2}O)(OH)]^{2+} + CO_{2} Net reaction H_{3}O^{+} + [Rh(en)_{2}CO_{3}]^{+} \longrightarrow [Rh(en)_{2}(H_{2}O)(OH)]^{2+} + CO_{2}$$

This process does release the weak acid, CO_2 , but consumes the strong acid H_3O + . Thus, the observed pH decrease would not be anticipated if such a process were occurring. In addition, reaction of the photoproduct of such a sequence with HCl would lead to <u>cis</u> or <u>trans-</u> [Rh(en)₂Cl₂] +; the observed product, with a single, intense peak at 331 nm is neither of the [Rh(en)₂ Cl₂] + isomers.

Possibility b), the two electron oxidation of the carbonate ligand, is more consistent with the available information (and is the more fascinating photochemical reaction). Assuming concomitant two-electron reduction of the Rh(III) center, a reaction of the form (eq 10)

$$\left[\operatorname{Rh}(\operatorname{en})_{2}\operatorname{CO}_{3} \right]^{\dagger} \longrightarrow \left[\operatorname{Rh}(\operatorname{en})_{2} \right]^{\dagger} + \operatorname{CO}_{2} + {}^{1}_{2}\operatorname{O}_{2}$$
(10)

would be occurring.

Reaction 10 is consistent with a variety of observations. Plots of "moles of H⁺ released" as a function of photolysis time showed two distinct patterns, depending upon the initial pH (Fig. 4). At pH 6, the rapid initial generation of H⁺ becomes less pronounced as the photolysis proceeds; at pH 10, the initial photolysis period generates only minor amounts of H⁺, followed by a more rapid rate of H⁺ generation. This is consistent with the release of a weak acid, as at pH 10 the $[HCO_3 -] / [CO_2]$ ratio would be four orders of magnitude larger than at pH 6.

$$H_20 + CO_2 \implies HCO_3^- + H^+$$
(11)

$$K = \frac{[H^+][HCO_3^-]}{[CO_2]}$$
(12)

$$\frac{K}{[H^+]} = \frac{\left[HCO_3^{-}\right]}{[CO_2]}$$
(13)

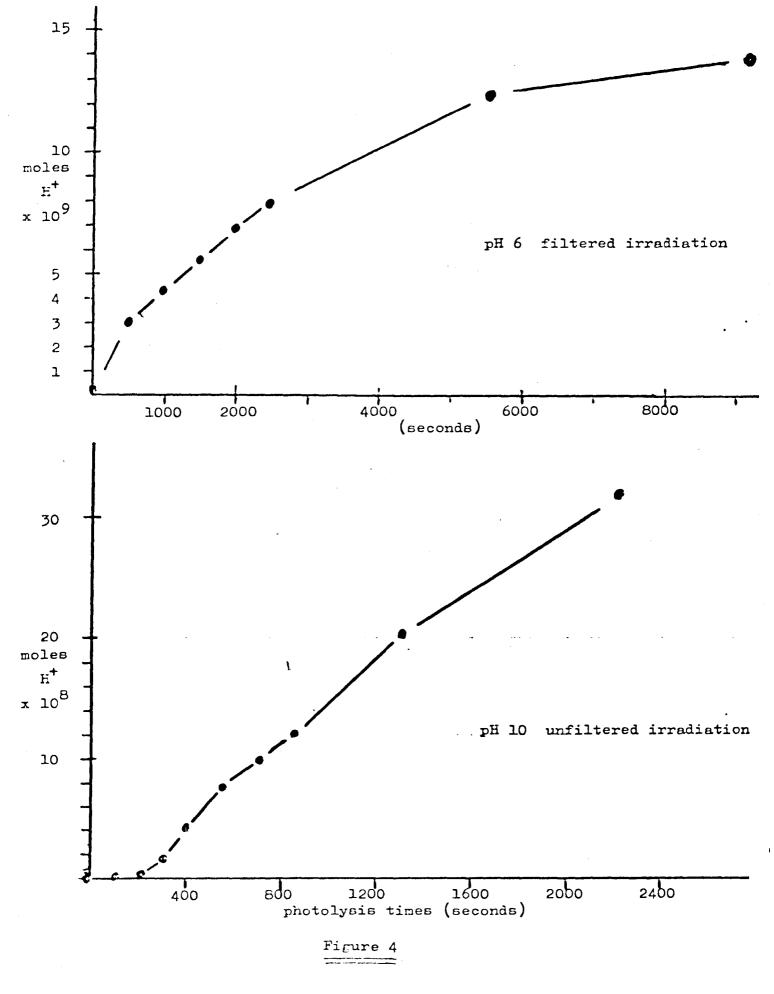
Such a large concentration of bicarbonate ion would act as a buffer, resisting the generation of H^+ ions upon the addition of CO_2 to the solution. As shown in Fig. 4, H^+ increase was initially quite slow, until the buffering action of the bicarbonate ion was overcome. At pll 6, the bicarbonate ion concentration would be less formidable, and little buffering action would be apparent. If we consider that for reaction 11,

$$\begin{bmatrix} H^{+} \end{bmatrix} = \begin{bmatrix} HCO_{3}^{-} \end{bmatrix} \text{ and } \begin{bmatrix} CO_{2} \end{bmatrix} = \begin{bmatrix} CO_{2} \end{bmatrix} \text{ initial } - \begin{bmatrix} H^{+} \end{bmatrix}$$

$$K = \frac{|H^{+}|^{2}}{\begin{bmatrix} CO_{2} \end{bmatrix}} = \frac{|H^{+}|^{2}}{\begin{bmatrix} CO_{2} \end{bmatrix} \text{ initial } - \begin{bmatrix} H^{+} \end{bmatrix}}$$

$$XVI-18$$

we can write



Solving for H^+ , one obtains:

$$[H^+] = \left\{ K \left(\begin{bmatrix} CO_2 \end{bmatrix} \right] \text{ initial} + \frac{{}^{1_4}K^2}{2} - {}^{1_2}K \quad (14) \right\}$$

Evaluating for H^+ at increasing levels of $[\text{CO}_2]$ initial (representing the CO_2 gnerated by photolysis) gives a curve which closely models the pH 6 curve of Fig. 4.

After 2200 seconds of unfiltered irradiation of $[Rh(en)_2CO_3]^{T}$, the spectrum shows that a significant amount of the starting material has photochemically reacted. The data from Table 2 can be used to show that the H⁺ released at that time is equivalent to about 10% of the total Rh(III) concentration. This, again, is consistent with the photogeneration of a weak acid (CO₂). A more quantitative assessment is not yet possible, as the exact nature of the Rh -containing photoproduct is not known.

The nature of the Rh-containing photoproduct is a major mystery remaining after this summer's work. The observed spectral changes do not suggest the formation of any obvious Rh(III) tetraamine species, and refluxing the photoproduct in HCl leads to an intensely yellow species which has, as yet, escaped identification. The spectral changes are consistent with the formation of the same product observed upon BH₄⁻ reduction of $cis - [Rh(en)_2Cl_2]^+$, suggestive of a two-electron reduction of the Rh(III) center, but further work would be necessary to corroborate this supposition.

In summary, work this summer has led us to the tantalizing hypothesis that ligand field and charge-transfer photolysis of a bidentate carbonato complex of Rh(III) leads to the two-electron oxidation of the carbonate ligand, with complementary reduction of the Rh(III) center. This result, if borne out by further studies, would represent a new type of photochemical reaction, and would encourage further work on the water-photolysis scheme (eq 4-8).

1. Conclusions

a) The monodentate carbonato complex, <u>trans-</u> $[Rh(en)_2 (H_{20})(CO_3)]^+$ efficiently undergoes photoisomerization:

 $\underbrace{\operatorname{trans-}\left[\operatorname{Rh}(\operatorname{en})_2 (\operatorname{H}_20)(\operatorname{CO}_3)\right]^+ \xrightarrow{\operatorname{hv}} \operatorname{cis-}\left[\operatorname{Rh}(\operatorname{en})_2 (\operatorname{H}_20)(\operatorname{CO}_3)\right]^+ }_{\text{There is no evidence of electron-transfer upon either ligand field or charge-transfer excitation. A very inefficient base-releasing photochemical reaction also occurs, but its nature has not been characterized. Preliminary evidence implies that the cis-aquocarbonato photoproduct undergoes photoinduced isomerization back to <math display="block"> \underbrace{\operatorname{trans-}}_{\text{IRh}(\operatorname{en})_2} (\operatorname{H}_20)(\operatorname{CO}_3) \right]^+.$

b) Ligand field and charge-transfer photolysis of the bidentate carbonato complex, $Rh(en)_2 CO_3$ + causes the release of CO_2 into solution by a reaction tentatively described as:

 $\left[\operatorname{Rh}(\operatorname{en})_2 \operatorname{CO}_3 \right]^+ \xrightarrow{h \, v} \left[\operatorname{Rh}(\operatorname{en})_2 \right]^+ + \operatorname{CO}_2 + \frac{1}{2} \operatorname{O}_2$

If subsequent studies corroborate the occurrence of a two-electron transfer, it will represent a new class of photochemical reaction. Such a reaction is the crucial first step in a water-photolysis scheme (eq 4-8), so confirmation of an efficient two-electron transfer, induced by a single photon, will open a new area of transition metal photochemistry.

2. Recommendations

This summer's study has shown that a promising area of transition metal photochemistry can be found in the study of bidentate carbonate complexes. The nature of the Rh-containing photoproduct in the photolysis of $[Rh(en)_2 CO_3]$ must receive immediate attention, and the design of future experiments depends on the results of that study. Available evidence strongly suggests that a two-electron redox process is involved, so the next steps must include:

a) a study to determine the factors which maximize the efficiency (quantum yield) of this redox reaction.

b) Studies must be initiated to increase the amount of the solar spectrum absorbed by the photoreactive complex. A logical approach would be to look at some substituted bipyridyl ligands or other aromatic ligands with extensive absorption in the visible and near-IR portions of the spectrum. The use of intensely colored sensitizers should also be studied.

c) Parallel studies should explore the detailed chemistry of the watersoluble complexes of Rh(I). The nature of the four-coordinate d^8 complexes, and the proposed d^6 hydrides must be determined. The thermal and photochemical properties of a variety of such hydride complexes must be studied to determine the conditions which would lead to a photoreactive hydride complex.

In summary, the discovery of a two-electron redox reaction, induced by a single photon, suggests that it may be possible to photochemically generate H₂ from water without the complications inherent in the hydrogen atom radical. This should encourage further study of a carbonate-based system for the photochemical cleavage of water.

REFERENCES

- a) Adamson, A. W.; Fleischauer, P.S., <u>Concepts in Inorganic</u> <u>Photochemistry</u> (eds.), 1975, Wiley-Interscience, N.Y.: (Chapter 5, by Ford, P.C.; Hintze, R.E.; Petersen, J.D.)
 - b) Endicott, J.F.; Wong, C-L,; Inoue, T.; Natarajan, P., <u>Inorg. Chem.</u>, 1979, <u>18</u>,450
 - c) Ford, P.C.; Skibsted, L.H., Strauss, D., <u>Inorg. Chem.</u>, 1979, <u>18</u>, 3171.
 - d) Clark, S.; Petersen, J.D., Inorg. Chem., 1979, 18, 3394.
 - e) Martins, E.; Sheridan, P.S.; Inorg. Chem., 1978, 17, 2822, 3631.
- 2. Kelly, T. C.; Endicott, J.F., J. Phys. Chem., 1972, 76, 1937.
- 3. Blake, D. M.; Mersecchi, R., J. Chem. Soc., D., (Chem. Comm.), 1971, 1045.
- 4. Blake, D. M.; Nyman, C. J., J. Amer. Chem. Soc., 1970, 92, 5359.
- 5. Mann, F. G.; Crowfoot, D. C.; Gattiker, D. C.; Wooster, N., <u>J. Chem. Soc.</u>, 1935, 1642.
- 6. Poulenc, P., Ann. Chim. (France), 1935, 4, 634.
- 7. Addison, W. W.; Gillard, R.D.; Sheridan, P.S.; Tipping, L.R.H., J. Chem. Soc., Dalton, 1974, 709.
- 8. Johnson, S.A.; Basolo, F., Inorg. Chem., 1962, 1, 925.
- 9. Rerek, Mark, Master's Thesis, State University of New York at Binghamton, August, 1980.
- Palmer, D.A.; vanEldik, R.; Kelm, H.; Harris, G.M., <u>Inorg. Chem</u>. 1980, 19, 1009.

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

CALCULATIONS AND MEASUREMENTS IN SUPPORT OF THE TRIGLYCINE SULFATE ZERO GRAVITY GROWTH EXPERIMENT

Prepared By:

John M. Springer, Ph.D..

Academic Rank:

Research Associate

University and Department:

NASA/MSFC: (Laboratory)

(Division) (Branch) Fisk University Department of Physics

Space Sciences Materials Processing Solid State

MSFC Counterpart:

Roger Kroes, Ph.D.

Date:

August 1, 1980

Contract No.:

NGT-01-002-099 (University of Alabama)

CALCULATIONS AND MEASUREMENTS IN SUPPORT OF THE TRIGLYCINE SULFATE ZERO GROWTH GRAVITY EXPERIMENT

By

John M. Springer, Ph.D. Research Associate in Physics Fisk University Nashville, Tennessee

ABSTRACT

The solution growth of crystals of triglycine sulfate (TGS) in a zero gravity environment is a materials processing experiment planned for an early Spacelab mission. This experiment has the twin objectives of both studying the growth process by optical means and also returning the grown crystals to earth for characterization.

This paper describes two areas of investigation which are required for maximum utilization of the flight experiment results. The first area studied is the use of an ultramicroscope to characterize structural defects in the grown TGS crystals. Mie theory has been applied to the light scattering expected from idealized crystal defects in order to predict the sensitivity of ultramicroscopy in the detection of small defects.

The second area of investigation has involved the measurement of optical properties of TGS solutions in water. This data is needed in order to interpret interferometric studies of mass transport through the TGS solution during the growth phase of the flight experiment. In particular, refeactive index data has been obtained for a wide range of temperatures and concentrations of TGS solutions. Absorbance measurements in the ultraviolet have also been made for the same samples.

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INTRODUCTION

One of the materials processing experiments that will be flown aboard an early Spacelab mission is the solution growth of triglycine sulfate (TGS). This experiment has two objectives, one being the observation of the crystal growth process itself by optical techniques such as Schlieren photography and interferometry, the other being the return of the zero-gravity grown crystals to earth for testing and comparison with their earth-grown counterparts.

This report describes two studies performed which are related to the TGS growth experiment. The first part discusses the use of a light scattering technique called ultramicroscopy to characterize triglycine sulfate crystals. The theory of Mie for the scattering of light by spheres is used to indicate the sensitivity of the ultramicroscope in the detection of volume defects. Although based on an idealization of the actual defects that would be present in a real crystal, this analysis is useful in analyzing the effectiveness of this technique.

The second part of the report presents data on the optical properties of solutions of TGS in water. This information on the refractive index and optical absorbance of TGS in solution is required for the analysis of the Schlieren and interferometric data which will be obtained during the crystal growth experiment on Spacelab 3.

I. Crystal Characterization by Ultramicroscopy

A. General

Crystal characterization involves describing a crystal's structure and composition in such a way that those of its properties of relevance to the study at hand are well enough known to predict in principle the behavior of the crystal. One of the important features of a crystal are its departures from the perfect array of atoms by which its structure is commonly described. These departures from ideality, called defects, exist in every real crystal and range from the simple absence of a single atom from the lattice to dislocations involving the entire crystal.

The scattering of electromagnetic radiation incident on a crystal is one tool for detecting defects. The scattering and diffraction of X-rays yields information on structural features whose size is on the order of the lattice spacing. In the case of optically transparent erystals, the scattering of visible light can also reveal structural defects in a crystal, though not with the detail of X-ray topography. For a crystal whose optical properties are germane to its application, however, the measurement of visible light scattering is essential for its proper characterization.

One method being implemented to study TGS crystals by light scattering is that of the ultramicroscope. In this technique, illustrated in Figure 1, a beam of light focussed to a small diameter is used to illuminate the crystal. As the light passes through the crystal, it is scattered by inhomogeneities in the crystal's refractive index, such as are caused by certain types of defects. In order to observe the light scattered from the defects, a microscope is used to collect the light scattered at right angles to the beam. Small, isolated defects show up as points of light in the dark surrounding field of view. Regions of a crystal where the defect density is quite high will show nebulous or diffuse scattering regions. As described by Vand,¹ a small He-Ne laser makes an excellent source for the ultramicroscope. This method has already been used to study the relation between defects in TGS and its dielectric properties.²

A diagram of an ultramicroscope being constructed for the study of TGS is shown in Figure 2. This ultramicroscope will use photoelectric rather than photographic detection of scattering centers, and will be computer driven so a three dimensional map of the light scattering defect may be constructed.^{*} In the remainder of this section of the report, an analysis is described in which the expected sensitivity of this ultramicroscope to defects is calculated. This sensitivity will be seen to depend on defect size and relative effective refractive index.

* This is being constructed at Fisk University with the support of NASA Grant #NSG-8060.

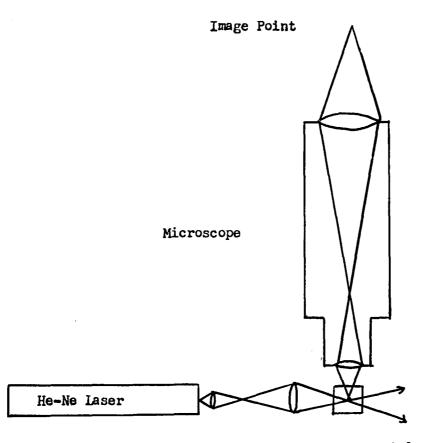




FIGURE 1: Schematic diagram of ultramicroscope used to observe defect scattering.

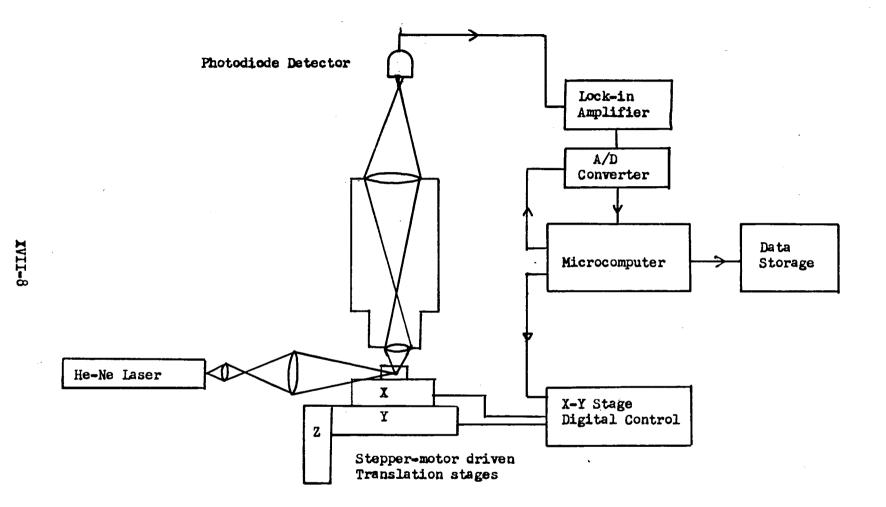


FIGURE 2: Automated ultramicroscope

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B. Discussion

The model "defect" which has been used in this study is that of a homogeneous sphere of diameter D and refractive index n. This spherical defect is assumed to be surrounded by the undisturbed TGS lattice. The absorption coefficient of the defect is assumed to be zero, so the refractive index is a real number. It is illuminated in the ultramicroscope by a plane electromagnetic wave of wavelength 632.8 nm in air.

The theory of scattering of electromagnetic waves by spheres is a classical problem, and was solved analytically by Mie.³ His solutions involve infinite sums of complex functions, but several sets of tables are available in the literature which make the evaluation of the solutions for particular cases straightforward, if somewhat tedious.^{4,5,0} Computer programs for this purpose have also been written, and would be useful if a large number of cases must be calculated.

Figure 3 defines the geometry of the problem. According to the treatment by Boll, the scattered light intensity is represented by two intensity functions i_1 and i_2 , given by

$$i_{1} = \frac{\lambda^{2}}{g_{\pi^{2}}} \left| \sum_{n=1}^{\infty} \left\{ A_{h} \frac{dP_{n}(x)}{dx} + B_{h} \left[x \frac{dP_{n}(x)}{dx} - (1-x^{2}) \frac{d^{2}P_{n}(x)}{dx^{2}} \right] \right\} \right|^{2} (1)$$

and

$$i_{2} = \frac{\lambda^{2}}{\beta_{\pi^{2}}} \left| \sum_{n=1}^{\infty} \left\{ A_{n} \left[x \frac{dP_{n}(x)}{dx} - (1-x^{2}) \frac{d^{2}P_{n}(x)}{dx^{2}} \right] + B_{n} \frac{dP_{n}(x)}{dx} \right\} \right|^{2} (z)$$

where $P_n(x)$ is a Legendre polynomial of order n, with $x = -\cos(\theta)$. Θ is the scattering angle from the incident direction. The coefficients A_n and B_n are complex Ricatti-Bessel and Hankel functions, and for the values to be considered here can be obtained from tables.⁴ These coefficients depend on only two physically relevant parameters for their evaluation, a size parameter.

$$\mathcal{L} = \frac{\pi \mathbf{D}}{\lambda} \tag{3}$$

where D = the scattering particle diameter

 λ = the wavelength of the incident light in the surrounding medium

and m, the refractive index of the scattering particle relative to the surrounding medium.

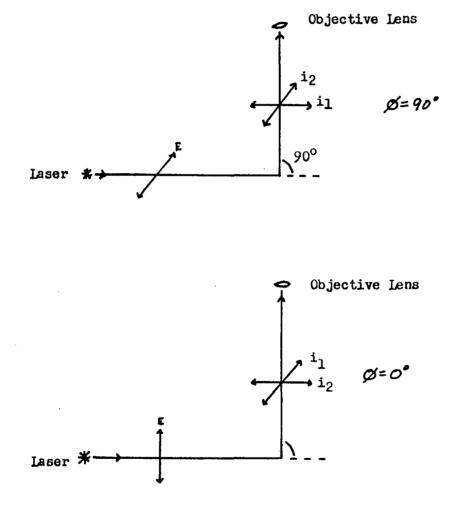


FIGURE 3: Orientation of the intensity functions for right angle scattering.

For plane-polarized incident light, the intensity function i_1 is the component of scattered light polarized perpendicular to the electric vector of the incident beam, and i_2 represents the component of scattered light polarized parallel to the plane formed by the electric vector of the incident beam and its direction of propagation. The resulting scattered intensity per unit incident intensity and unit solid angle is for any direction θ , \emptyset given by

$$\mathbf{I} = \frac{\lambda^2}{4\pi^2} \left[(\cos^2 \phi) \mathbf{i}_1 + (\sin^2 \phi) \mathbf{i}_2 \right]$$
(4)

In the ultramicroscope, the scattered light collected by the objective lens is nominally at a scattering angle $\theta = 90^{\circ}$. If the incident light on the crystal is polarized with its electric vector perpendicular to the plane formed by the incident and scattered light directions of propagation, then $\emptyset = 90^{\circ}$ also, so the total scattered intensity per unit solid angle is

$$I_{n} = I_{0} \frac{\lambda i_{1}}{8\pi^{2}}$$
(5)

where I_0 is the incident intensity (W/m²).

A typical objective used on the microscope will have a numerical appearature of 0.25. Taking into account the refraction that occurs at the surface of the TGS crystal as the scattered light passes into the air, an objective lens of N. A. 0.25 will collect a cone of light with a half-angle of 9° in the crystal, corresponding to a solid angle of 0.00612. So for this case, the scattered power entering the objective lens will be

$$P_{s} = I_{o} \frac{\lambda}{9n^{2}} (6.12 \times 10^{-3}) i_{2}$$
 (6)

In the ultramicroscope the detector is a photodiode with an active area which is larger than the image of the smaller defects with which this analysis is concerned, so the scattered light power collected by the objective lens will all fall on the detector when the defect is centered in its field of view. P_s , therefore, is the relevant value which must be calculated and compared with the sensitivity of the detector and its amplifier.

C. Calculations

The relevant physical parameters of a defect are its diameter D and refractive index n. TGS in solution at its saturation point at 25° has a refractive index of approximately 1.38. The crystalline solid has three principal indices which range from 1.484 to 1.584.⁸ Values of the relative refractive index m could be expected to vary typically from 0.6 for a void with absolute index 1 to 0.93 for a volume defect consisting of uncrystallized solution. The actual values of m could of course be close to 1 for small distortions of the lattice. Values of the size parameter which are available in Boll's? tables range upward from 1 to 120. Since only small defects are of interest in finding the limit of detectable scattering, values of from 1-10 were used in the calculations. In triglycine sulfate, this corresponds to a diameter of from approximately 1.2 x 10-7 meters to 1.2 x 10⁻⁰ meters when the wavelength of the incident light is 632.8nm.

The laser which will be used is a 5mW He-Ne laser, which can be focussed to a 20 x 10^{-0} meter diameter in the crystal. This gives an average beam intensity of $I_0 = 1.6 \times 10^7 \text{ W/m}^2$. Using a refractive index of 1.58 for the crystalline TGS, the wavelength of the laser light in the crystal is $\lambda = 400$ nm. These values give a scattered power from equation 6 of

$$P_{g} = I_{0} \frac{\lambda^{2} (6.12 \times 10^{-3})}{8\pi^{2}} i_{z} = (1.98 \times 10^{-10} \text{Watts}) i_{z}$$
(7)

Representative values of the scattered power collected by the ultramicroscope for several values of and m are shown in Table 1. The effective scattering defect size D is also given.

An evaluation of these values in terms of photoelectric detection can be made by comparison with the noise figure for a photodiode which is being evaluated in the ultramicroscope. The photodiode, a model PN-40A from United Detector Technology, has a noise equivalent power (NEP) of 10⁻¹¹ Watts. For spherical defects with a relative refractive index near unity (0.93), defects of this type would have to exceed 300nm to be above the noise level of the detector. For voids, where the refractive index would be around 0.6, a defect somewhat small (100nm) would scatter sufficiently to be observed.

TABLE	1
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Power scattered at 90° from spherical defects for plane polarized incident beam of wavelength 400 nm and intensity 1.6 x 10^7 W/m². Power is in Watts.

Defect Diameter		Relative Refractive Index m					
d	(nm)	0.6	0.7	0.75	0.8	0.9	0.93
1	127	6.2x10-14	2.9x10-15	1.5x10 ⁻¹⁵	6.6x10 ⁻¹⁶	4.7x10 ⁻¹⁷	1.2x10 ⁻¹⁷
2	250	2.7x10 ⁻¹²	1.2x10 ⁻¹²	7.5x10 ⁻¹³	3.2x10 ⁻¹³	2.8x10 ⁻¹⁴	7.5x10 ⁻¹⁵
3	381					4.1x10 ⁻¹⁴	1.1x10 ⁻¹³
4	508					4.9x10 ⁻¹⁴	1.3x10 ⁻¹³
5	635					8.9x10 ⁻¹⁴	3.2x10 ⁻¹³
10	127	,				2.0x10 ⁻¹²	1.6x10 ⁻¹²

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II. Optical Measurements

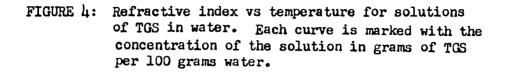
Refractive index and ultraviolet absorbance measurements were made on solutions of triglycine sulfate in distilled water at concentrations ranging from 0.25 grams-TGS/100 grams-H₂O to 32 grams-TGS/100 grams-H₂O. Samples of various concentrations were obtained by making an initial solution of 32g/100g, dividing this in half and diluting one part to make a 16g/100g solution, and then repeating this process to form a series of samples each one-half as concentrated as its predecessor. The resulting solutions were stored in sealed bottles. The samples were weighed during their preparation to a precision of ± 0.01 grams.

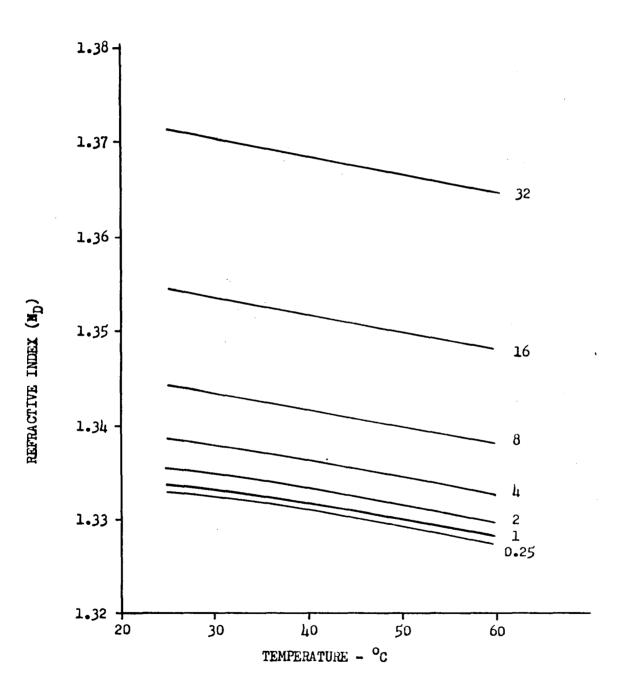
Refractive index measurements were made using a Billingham and Stanley Model 60/ED precision Abbe refractometer. The sample and prism temperature was controlled by water circulated from a water bath in which the temperature was regulated to ± 0.01 degree centigrade. The actual refractometer head temperature versus bath temperature was determined by calibration to compensate for heat loss in the connecting tubing. Temperature measurement was made using a Cole-Parmer Model 8502-25 thermistor thermometer, and a Yellow Springs Inc. Model 703 thermistor probe. A sodium lamp was used to illuminate the refractometer, so the values obtained are for the sodium D₁ line, 589.6 nm. Values calculated from the critical angle measurements using the refractometer are given in Table 2. Graphs of n_d vs temperature for various concentrations and n_d vs concentration for different temperatures are shown in Figures 4 and 5.

Transmission curves in the range of 200 - 300 nm were measured for the same samples used in the refractive index study. The curves were recorded on a Model 701 Heath/Schlumberger single-beam spectrophotometer. All measurements were made for a 10 mm path length. This data will be used in experimentally determining mass diffusion coefficients for TGS in water. For this purpose, a graph of the wavelengths at which the present transmission of the solutions reaches 50% is useful. This graph is shown in Figure 6.

TABLE 2: Refractive Index Data for Triglycine Sulfate Solution $-(n_d)$

Concentration			Solution	n Tempera	ature - 1	Degree Ce	elsius	
g/100g H ₂ 0	24.98	29.66	34.36	39.28	43.95	48.75	53.45	58.23
0.25	1.33290	1.33236	1.33170	1.33100	1.33018	1.32938		
0.50	1.33312	1.33252	1.33191	1.33114	1.33052	1.32952	1.32858	1.32762
1.00	1.33381	1.33326	1.33259	1.33184	1.33108	1.33020	1.32929	
2.00	1.33541	1.33476	1.33411	1.33337	1.33253	1.33176	1.33085	1.32989
4.00	1.33853	1.33791	1.33721	1.33643	1.33560	1.33471	1.33380	1.33292
8.00	1.34408	1.34343	1.34274	1.34190	1.34101	1.34009	1.33918	1.33819
16.00	1.35430	1.35352	1.35270	1.35188	1.35011	1.35011	1.34917	1.34815
32.00	1.37116	1.37038	1.36948	1.36879	1.36779	1.36681	1.36590	1.36494





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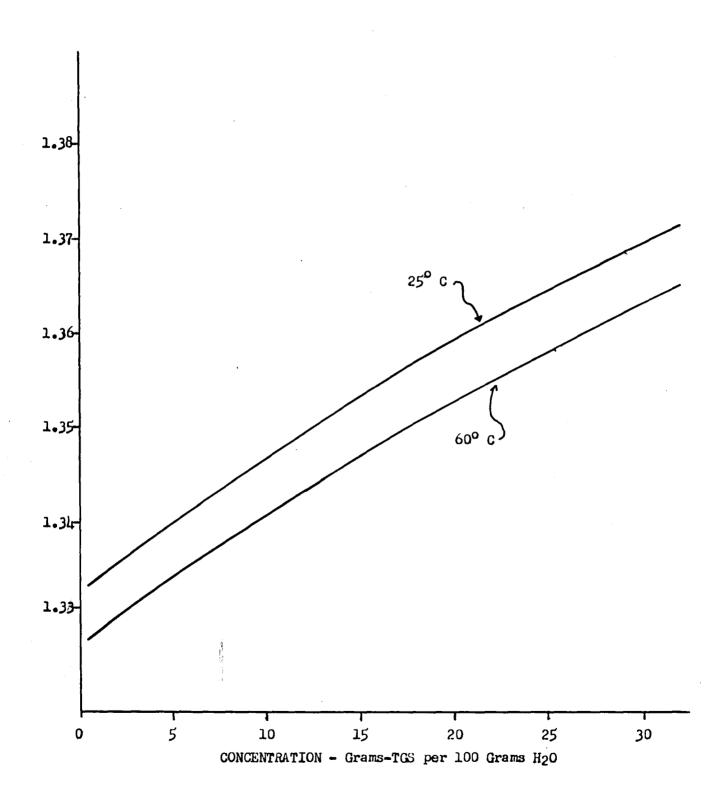
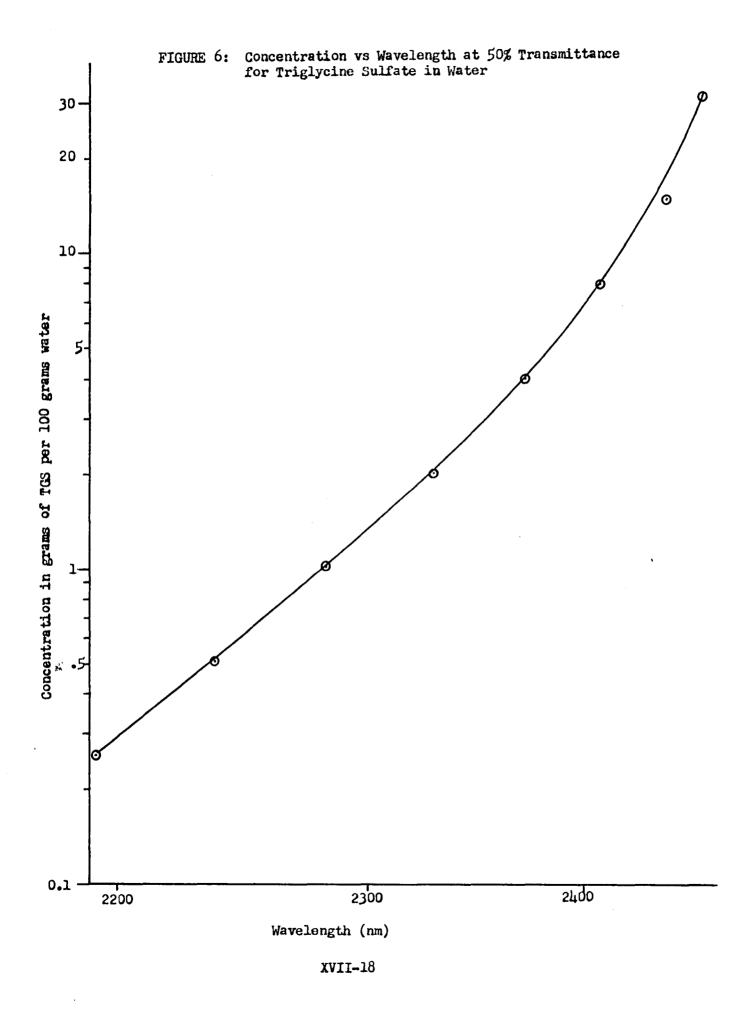


FIGURE 5: Refractive index vs concentration for solutions of TGS in water.

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REFERENCES

- 1. Vand, V. and K. Vedan, "The Laser as Light Source for Ultramicroscopy and Light Scattering by Imperfections in Crystals" J. Appl. Phys. <u>37</u>, 2551-2557, 1966.
- 2. Mohler, R. and W. White, "Characterization of Volume Defects in Triglycine Sulfate by Laser Light Scattering", J. Crystal Growth 28, 240-248, 1975.
- 3. Mie, G., Ann. Physik, 25, 377(1908).
- 4. Deirmendjian, D., Electromagnetic Scattering on Spherical Polydispersions, American Elsevier Publishing Co., New York, N. Y., 1969.
- 5. Che, C., Clark, G. C., and S. W. Churchill, <u>Tables of Angular</u> <u>Distribution Coefficients for Light-Scattering By Spheres</u>, University of Michigan Press, Ann Arbor, Michigan, 1957.
- 6. Denman, H. H., et. al, Angular Scattering Functions for Spheres, Wayne State University Press, Detroit, Michigan, 1966.
- 7. Boll, R. H., Leacock, J. A., Clark, G. C., and S. W. Churchill, <u>Tables of Light-Scattering Functions</u>, The University of Michigan Press, <u>Ann Arbor, Michigan, 1958.</u>
- 8. Dion, I. M., et. al., Acta Crystallographica, 12, 259-264, 1959.

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

FLUID DYNAMIC ANALYSIS OF THE SPACE SHUTTLE MAIN ENGINE HIGH PRESSURE OXIDIZER TURBOPUMP SLINGER SEAL

Prepared By:

James S. Voss, Captain

Academic Rank:

University and Department

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Assistant Professor

United States Military Academy Department of Mechanics

Structures and Propulsion Propulsion Turbomachinery and Combustion Devices

Loren Gross

August 1, 1980

NGT-01-002-099 (University of Alabama)

FLUID DYNAMIC ANALYSIS OF THE SPACE SHUTTLE MAIN ENGINE HIGH PRESSURE OXIDIZER TURBOPUMP

SLINGER SEAL

By

Captain James S. Voss Assistant Professor of Mechanics United States Military Academy West Point, New York

ABSTRACT

An analytical study has been conducted to clarify the details of the flow on the bladed side of a centrifugal type dynamic shaft seal utilized to contain liquid oxygen in the Space Shuttle Main Engine high pressure oxidizer turbopump.

The governing equations are solved to predict the pressure and temperature gradients and to aid in investigating the nature of the liquid-vapor interface.

Recommendations for design improvements are discussed.

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NOMENCLATURE

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a	-	Axial clearance between seal and housing (in)
b	₩.	Blade height (in)
С	-	Constant pressure specific heat (BTU/1bm- ^O R)
c _m	-	Moment coefficient (one side), $\tau / \chi_{\rho \omega}^2 r^5$
C _{ms}	-	Smooth disk moment coefficient (one side)
c _{mv}	-	Vaned disk moment coefficient
Ec	-	Eckert Number, $\omega^2 R^2 / CT_0$
h	-	Enthalpy (BTU/1bm)
k	-	Thermal conductivity (BTU/fthr ^O R)
к	-	Fluid angular velocity ratio, eta / ω
m	-	Mass flow rate (1bm/sec)
Р	-	Pressure (lbf/in ²)
Ē	-	Dimensionless pressure, $p/\rho \omega^2 R^2$
Pr	-	Prandtl number, Cµ/k
r	-	Radial coordinate, radial distance (in)
r	-	Dimensionless radial coordinate, r/R
R	-	Disk radius (in)
Rint	-	Liquid-vapor interface radius (in)
Re	-	Reynolds number, $\omega R^2 / \nu$
Т	-	Temperature (^O R)
Ŧ	-	Dimensionless temperature, T/T _o
т _о	-	Seal fluid flow inlet temperature (^O R)
u	-	Radial velocity component (ft/sec)
ū	-	Dimensionless radial velocity component, u/R ω

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v	-	Tangential velocity component (ft/sec)
v	-	Dimensionless tangential velocity component, v/R ω
W	-	Axial velocity component (ft/sec)
Ŵ	-	Dimenisonless axial velocity component, w/R ω
z	-	Axial coordinate, axial distance (in)
ī	-	Dimensionless axial coordinate, z/R
β	-	Angular velocity of fluid (rad/sec)
μ	-	Viscosity (1bm/ft-sec)
ν	-	Kinematic viscosity (ft ² /sec)
ρ	-	Density (1bm/ft ³)
τ	-	Torque (ft-1bf)
ω	-	Angular velocity of disk (rad/sec)
1	,2 -	Subscripts denoting radial location on disk

INTRODUCTION

In high pressure oxidizer turbopumps it is necessary to have effective sealing systems to prevent contact between the turbine hot gas and the oxidizer being pumped. The centrifugal dynamic shaft seal, or slinger seal, can be a valuable part of such a sealing system.

In a slinger seal, a disk shaped impeller with vanes on one side and the other side smooth, rotates with the shaft to be sealed inside a close fitting, but non-contacting housing, as shown in Fig. 1. The liquid to be sealed is centrifuged to the outer periphery of the rotating disk, but because the vaned side has a steeper pressure gradient [1] there will be a resulting net pressure difference across the seal which limits the fluid flow. A gas introduced to the system or vaporized liquid will form an interface with the liquid on the vaned side of the seal.

The primary seal in the Space Shuttle Main Engine high pressure oxidizer turbopump uses this type of slinger with a downstream labyrinth as shown in Fig. 2. The actual seal is modeled by the simplified geometry indicated by Fig. 3.

Previous investigations have primarily been concerned with the more general topic of rotating disks. The theoretical work on free rotating disks done by von Karmán [2], which was corrected by Cochran [3], and the analysis of the enclosed rotating disk by Schultz-Grunow [4] are fundamental to many of the later studies. Daily and Nece, with their experimental work on enclosed rotating disks [5], [6], [7], laid a base for much of the theoretical [8], [9] as well as experimental [10], [11], [12] work that followed.

A limited number of researchers have studied the utilization of enclosed rotating disks as centrifugal seals, with most of their work being experimental or the correlation of experimental data with simple models [13], [14], [15], [16], [17], [18]. A notable exception is the theoretical analysis of smooth slinger seals by Rosenthal and Reshotko [19], [20].

OBJECTIVES

The purpose of this paper is to analyze the dynamics of the fluid on the vaned side of a slinger seal. The governing equations are established and reduced to a form solvable by numerical methods. Time constraints prevent the complete solution by matching of the boundary conditions and numerical analysis so this will be done as a future task.

Some of the important thermodynamic design considerations for this type seal are the radial pressure distribution, temperature distribution, and the liquid-vapor equilibrium interface. A solution is found for the radial pressure distribution and in order to look at the liquid-vapor interface a simplified analysis is used to obtain a restrictive temperature distribution.

DEVELOPMENT OF EQUATIONS

The fluid will be treated as an incompressible, Newtonian fluid where body forces and radiation effects are negligible. Temperature variations will generally be small so the viscosity, thermal conductivity and specific heats can be treated at constants for a particular temperature range [21]. The flow is steady and axisymmetric. Using these assumptions and the usual notation for rotational flow, the governing equations are:

Continuity

$$\frac{\partial u}{\partial r} + \frac{\partial w}{\partial r} = 0 \tag{1}$$

Radial Momentum

$$\mathbf{u} \frac{\partial \mathbf{u}}{\partial \mathbf{r}} - \frac{\mathbf{v}^2}{\mathbf{r}} + \mathbf{w} \frac{\partial \mathbf{u}}{\partial z} = -\frac{1}{\varrho} \frac{\partial \mathbf{p}}{\partial \mathbf{r}} + \nu \left[\frac{\partial^2 \mathbf{u}}{\partial \mathbf{r}^2} + \frac{\partial}{\partial \mathbf{r}} \left(\frac{\mathbf{u}}{\mathbf{r}} \right) + \frac{\partial^2 \mathbf{u}}{\partial z^2} \right]$$
(2)

Tangential Momentum

$$\mathbf{u} \quad \frac{\partial \mathbf{v}}{\partial \mathbf{r}} + \frac{\mathbf{u} \mathbf{v}}{\mathbf{r}} + \frac{\partial \mathbf{v}}{\partial z} = \mathbf{v} \left[\frac{\partial^2 \mathbf{v}}{\partial r^2} + \frac{\partial}{\partial r} \left(\frac{\mathbf{v}}{\mathbf{r}} \right) + \frac{\partial^2 \mathbf{v}}{\partial z^2} \right]$$
(3)

Axial Momentum

$$u \frac{\partial w}{\partial r} + w \frac{\partial w}{\partial z} = -\frac{1}{\varrho} \frac{\partial p}{\partial z} + \nu \left[\frac{\partial^2 w}{\partial r^2} + \frac{1}{r} \frac{\partial w}{\partial r} + \frac{\partial^2 w}{\partial z^2} \right]$$
(4)

Energy

$$\rho C \left(u \quad \frac{\partial T}{\partial r} + w \quad \frac{\partial T}{\partial z} \right) = k \left(\frac{\partial^2 T}{\partial r^2} + \frac{1}{r} \quad \frac{\partial T}{\partial r} + \frac{\partial^2 T}{\partial z^2} \right) +$$

$$\mu \left[2 \left(\frac{\partial u}{\partial r} \right)^2 + 2 \left(\frac{u}{r} \right)^2 + 2 \left(\frac{\partial w}{\partial z} \right)^2 + \left(\frac{\partial v}{\partial z} \right)^2 + \left(\frac{\partial u}{\partial z} + \frac{\partial w}{\partial r} \right)^2 + \left(\frac{\partial v}{\partial r} - \frac{v}{r} \right)^2 \right]$$
(5)

Equations (1)-(5) can be non-dimensionalized by using the following substitutions:

$$\overline{r} = \frac{r}{R}, \overline{Z} = \frac{Z}{R}, \overline{u} = \frac{u}{\omega R}, \overline{v} = \frac{v}{\omega R}, \overline{w} = \frac{w}{\omega R}, \overline{P} = \frac{P}{\rho \omega^2 R^2}, \overline{T} = \frac{T}{T_0}$$

where T_0 is the fluid inlet temperature.

Using these expressions the differential equations become in dimensionless form:

Continuity

$$\frac{\partial \overline{u}}{\partial \overline{r}} + \frac{\overline{u}}{\overline{r}} + \frac{\partial \overline{w}}{\partial \overline{z}} = 0$$
(6)

Radial Momentum

$$\overline{\mathbf{u}} \quad \frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{r}}} \quad -\frac{\overline{\mathbf{v}}^2}{\overline{\mathbf{r}}} + \overline{\mathbf{w}} \quad \frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{z}}} = -\frac{\partial \overline{\mathbf{p}}}{\partial \overline{\mathbf{r}}} + \frac{1}{\mathrm{Re}} \left[\frac{\partial^2 \overline{\mathbf{u}}}{\partial \overline{\mathbf{r}}^2} + \frac{1}{\overline{\mathbf{r}}} \frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{r}}} - \frac{\overline{\mathbf{u}}}{\overline{\mathbf{r}}} + \frac{\partial^2 \overline{\mathbf{w}}}{\partial \overline{\mathbf{z}}^2} \right]$$
(7)

Tangential Momentum

$$\vec{u} \quad \frac{\partial \vec{v}}{\partial \vec{r}} + \frac{u \vec{v}}{\vec{r}} + \vec{w} \quad \frac{\partial \vec{v}}{\partial \vec{z}} = \frac{1}{Re} \left[\frac{\partial^2 \vec{v}}{\partial \vec{r}^2} + \frac{1}{\vec{r}} \frac{\partial \vec{v}}{\partial \vec{r}} - \frac{\vec{v}}{\vec{r}^2} + \frac{\partial^2 \vec{v}}{\partial \vec{z}^2} \right]$$
(8)

Axial Momentum

$$\overline{\mathbf{u}} \quad \frac{\partial \overline{\mathbf{w}}}{\partial \overline{\mathbf{r}}} + \overline{\mathbf{w}} \quad \frac{\partial \overline{\mathbf{w}}}{\partial \overline{\mathbf{z}}} = - \quad \frac{\partial \overline{\mathbf{P}}}{\partial \overline{\mathbf{z}}} + \frac{1}{\mathsf{Re}} \left[\frac{\partial^2 \overline{\mathbf{w}}}{\partial \overline{\mathbf{r}}^2} + \frac{1}{\mathsf{r}} \quad \frac{\partial \overline{\mathbf{w}}}{\partial \overline{\mathbf{r}}} + \frac{\partial^2 \overline{\mathbf{w}}}{\partial \overline{\mathbf{z}}^2} \right] \tag{9}$$

Energy

$$\overline{\mathbf{u}} \quad \frac{\partial \overline{\mathbf{T}}}{\partial \overline{\mathbf{r}}} + \overline{\mathbf{w}} \quad \frac{\partial \overline{\mathbf{T}}}{\partial \overline{\mathbf{z}}} = \frac{1}{\Pr \operatorname{Re}} \left(\frac{\partial^2 \overline{\mathbf{T}}}{\partial \overline{\mathbf{r}}^2} + \frac{1}{\overline{\mathbf{r}}} \quad \frac{\partial \overline{\mathbf{T}}}{\partial \overline{\mathbf{r}}} + \frac{\partial^2 \overline{\mathbf{T}}}{\partial \overline{\mathbf{z}}^2} \right) + \frac{\mathbf{E} \operatorname{c}}{\operatorname{Re}} \left[2 \left(\frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{r}}} \right)^2 + 2 \frac{\overline{\mathbf{u}}^2}{\overline{\mathbf{r}}^2} + 2 \left(\frac{\partial \overline{\mathbf{w}}}{\partial \overline{\mathbf{z}}} \right)^2 + \left(\frac{\partial \overline{\mathbf{v}}}{\partial \overline{\mathbf{z}}} \right)^2 + \left(\frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{z}}} \right)^2 + \frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{z}}} \partial \overline{\mathbf{r}}} + \frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{z}}} \partial \overline{\mathbf{r}}} + \frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{z}}} \partial \overline{\mathbf{r}}} \right) + \left(\frac{\partial \overline{\mathbf{w}}}{\partial \overline{\mathbf{z}}} \right)^2 + \left(\frac{\partial \overline{\mathbf{u}}}{\partial \overline{\mathbf{u}}} \right)^2 + \left(\frac{\partial \overline{\mathbf{u}}}$$

where

$$Re = \frac{\omega R^2}{\nu}, Pr = \frac{\mu C}{k}, Ec = \frac{\omega^2 R^2}{C T_0}$$
(11)

Equations (6)-(9) can be solved independent of equation (10) and are to be done by numerical analysis. The resulting flow field solution can then be used to solve equation (10) for the temperature distribution. The entire solution of these equations will be completed at a later time.

RADIAL PRESSURE DISTRIBUTION

The sealing capability of a slinger will be determined by the radial pressure gradient. The exact nature of this gradient will be quite complex due to the complicated flow across the disk. As an example, see reference 22 for a detailed look at secondary flows in a centrifugal impeller. Other interrelated factors are the presence of a liquid-vapor interface [18], pressure variation between vanes [23], and fluid throughflow [24], although the experimental results of Due [25] indicated that there was no significant dependance on flow coefficient.

A simplified approach is necessary so it is assumed that once steady state operation is reached the fluid in any thin annulus moves as a solid mass rotating about its own axis with angular velocity K ω where K is the ratio of the fluid angular velocity, β , to the disk angular velocity, ω . The radial velocity will be small compared to the tangential velocity and will be neglected, so the radial momentum equation (2) becomes:

$$\frac{1}{r} \frac{1}{v^2} = \frac{1}{\rho} \frac{\partial p}{\partial r}$$
(12)

Since the pressure is a function of the radius only, this equation reduces to an ordinary differential equation. Then substituting $K\omega r$ for the tangential velocity results in:

$$\frac{dp}{dr} = \frac{\rho (K \omega r)^2}{r}$$
(13)

Integrating between any two radial points on the face of the slinger yields:

$$\Delta p = p_2 - p_1 = \frac{\rho}{2} \kappa^2 \omega^2 (r_2^2 - r_1^2)$$
(14)

Other investigators [5], [12], [19], have obtained the same result by different approaches and their close correlation with their experimental results verifies its general applicability.

The value of K must be determined to solve for the pressure gradient, but K will be influenced by several variables including geometry of the seal, solidity, clearances, Reynolds number and radius. Since the fluid between the vanes on a slinger will rotate essentially with the slinger angular velocity, $K \approx l$, but the fluid in the gap between the vanes and the housing will rotate at some lesser angular velocity. The regions could be considered separately but in reference 25 Due established equations for K by correlating experimental data to equation (14) using a statistical regression analysis. The applicable equation from Due is:

$$K = 0.951 - 0.282 \frac{r}{R} + 0.2137 \frac{b}{a+b} + 0.175(b)$$
(15)

where R is the seal radius, a is the seal to housing gap width, b is the blade height, and r is the particular radial distance. Average values for the subject seal are: a = 0.08 in., b = 0.25 in., and R = 2.24 in. Using these values in equations (14) and (15) the radial pressure distribution can be determined.

TEMPERATURE DISTRIBUTION

Due to the rotation of the slinger, energy is imparted to the fluid, manifested as a temperature rise and eventually as a phase change from liquid to vapor. This phase change causes a drastic change in the density of the fluid and a resulting smaller radial pressure change. To account for this effect, the radial temperature distirbution must be known.

A simple control volume analysis will be used to look at the enthalpy increase from one radial plane to another. Knowing the enthalpy and pressure defines the state at a point so the temperature, density or any other required properties can be found using a standard computer code [27] or from tables or charts [28]. The process will be assumed to be steady state, steady flow [26] and the radial distance between inlet and exit planes will be small so the change in kinetic and potential energies will be negligible. Heat transfer to the walls will be considered negligible. Continuity requires the mass flow at the inlet to equal that of the exit and with the other conditions as above, the one dimensional energy equation will be:

$$\mathbf{\mathring{m}h}_{1} = \mathbf{\mathring{m}h}_{2} + \mathbf{\mathring{W}}$$
(16)

where \dot{m} is the fluid mass flow through the seal, h is the enthalpy at a particular radius and \dot{W} is the shaft power absorbed by the fluid.

$$\dot{\mathbf{W}} = -\tau \omega \tag{17}$$

with torque being defined for one side of a disk $\begin{bmatrix} 5 \end{bmatrix}$, $\begin{bmatrix} 12 \end{bmatrix}$ as:

$$\tau = \frac{C_{\rm m} \rho \omega^2 (r_2^5 - r_1^5)}{4}$$
(18)

so the change in enthalpy becomes:

$$h_2 - h_1 = \frac{C_m \rho \omega^3 (r_2^5 - r_1^5)}{4 \dot{m}}$$
 (19)

The moment coefficient C_m must be determined. For the smooth side of the seal the value of C_{ms} is found to be 0.0008 for $R_e = 10^8$ using the empirical

equations from reference 8. For this seal the flow is between regimes 3 and 4 as defined by Daily and Nece $\begin{bmatrix} 5 \end{bmatrix}$ so an average value of C_{ms} from the two applicable equations is used. The disk is thin so no contributions due to the tip are considered.

For the vaned side of the disk the moment coefficient will be greater and will be approximated by the method derived by Thew $\lceil 15 \rceil$ where

$$C_{mv} = C_{ms} + 1.2 \frac{b+a}{R} (I-K)^2 \left[\left(\frac{a+b}{0.8 a} \right) - 1 \right]^2 \left[1 - \left(\frac{R_{int}}{R} \right)^4 \right]$$
 (20)

The pressure and enthalpy distributions can now be calculated, yielding any other thermophysical properties needed at a radial point. The velocity ratio is found to be essentially constant for the smooth side, K = 0.44, but varies with radius on the vaned side so C_m also varies as shown. The mass flow of 0.13 lbm/sec and slinger angular velocity of 2932 rad/sec are obtained from engine balance data [30]. On the vaned side it is necessary to assume a value of R_{int} , determine the point at which the fluid begins vaporizing, and use this as the new assumed value of R_{int} , continuing until the assumed and determined values converge. For this seal a radius of 2.2 in. is where vaporization, therefore the interface begins.

Using the property values from reference 29 (their station 20) as being at the root of the smooth side of the slinger (r = 1.7 in.) the radial property variations are calculated and summarized in Table 1.

CONCLUSIONS AND RECOMMENDATIONS

A thorough literature survey has been completed with the most applicable work listed beginning on page XVIII-16. A set of non-dimensional partial differential equations have been established which describe the flow on the face of a slinger seal. The pressure and temperature distributions for the Space Shuttle Main Engine high pressure oxidizer turbopump primary slinger seal have been described and the location of the liquid-vapor interface determined.

The radial pressure distribution equation (14) should provide good results based on the experimental work confirming it, but the analysis of the temperature variation is approximate at best and too many simplifying assumptions are made to trust its validity. The smooth side values for C_m are sufficiently grounded with experimental work but the equation for C_{mv} has not been investigated enough to verify its generality. Due to these factors and the nature of the inital data used from reference 29, it is believed that this simplified analysis can be used only to predict tendencies.

Further work needs to be done on vaned slinger seals. Solution of the governing equations and matching of all boundary conditions should be done to get a complete picture of the flow involved and to predict more accurate pressure and temperature distributions. An experimental program should be conducted to determine actual pressure and temperature values to correlate with the analytical work. No previous experimental work could be found through the literature survey which deals with liquid oxygen, high rotational speeds as found in modern turbo-pumps, or the energy transfer involved with the production of a liquid-vapor interface.

Recommendations based on the literature survey that could improve the sealing performance of the Space Shuttle Main Engine high pressure oxidizer turbopump slinger are:

- include circumferential tangs on the vaned side of the seal with matching grooves in the housing to stabilize the liquid-vapor interface, thereby reducing leakage [14].
- increase the number of vanes to 24 which has been established as optimum [14].
- utilize a non-wetting coating on the housing opposite the vaned side of the seal to reduce secondary flow that enhances leakage. [20].

REFERENCES

- Thew, M. T. and M. G. Saunders, "The Hydrodynamic Disk Seal", Proc. 3rd Int. Conf. on Fluid Sealing, British Hydromechanics Research Assoc., Paper H5, April, 1967.
- Kármán, Th. von, "Uber laminare und turbulent Reibung", Zeitschrift angew. Math. Mech., Vol. 1, 233-252, 1921.
- 3. Cochran, W. G., "The flow due to a rotating disk", Proc. Cambridge Phil. Soc. 30, 365-375, 1934.
- 4. Schultz-Grunow, F., "Der Reibungswiderstand Rotierender Scheiben in Gehausen", Zeitschrift angew Math. Mech., Vol. 15, 191, 1935.
- Daily, J. W. and R. D. Nece, "Roughness and Chamber Dimensional Effects on Induced Flow and Frictional Resistance of Enclosed Rotating Disks", MIT Hydrodynamics Laboratory Technical Report No. 27, May 1958.
- 6. Daily, J. W. and R. D. Nece, "Chamber Dimensional Effects on Induced Flow and Frictional Resistance of Enclosed Rotating Disks", Journal of Basic Engineering, Trans. ASME, Series D, Vol. 82, 217-232, 1960.
- 7. Daily, J. W. and R. D. Nece, "Roughness Effects on Frictional Resistance of Enclosed Rotating Disks", Journal of Basic Engineering, Trans. ASME, Series D, Vol. 82, 553-562, 1960.
- 8. Ketola, H. N. and J. M. McGrew, "Theory of the Partially Wetted Rotating Disk", Proceedings of the Third International Conference on Fluid Sealing, British Hydromechanics Research Association, Paper H4, April 3-4, 1967.
- 9. Ketola, H. N. and J. M. McGrew, "Pressure, Frictional Resistance, and Flow Characteristics of the Partially Wetted Rotating Disk", Journal of Lubricating Technology, Trans. ASME, Series F, Vol. 90, 395-404, 1968.
- Due, H. F. and W. C. Keathly, "A Study of Pressure Prediction Methods for Radial Flow Impellers", Pratt and Whitney Aircraft Report FR-952, 13 April 1964.
- Due, H. F. and W. E. Young, "Investigation of Pressure Prediction Methods for Radial Flow Impellers, Phase II", Pratt and Whitney Aircraft Report FR-1276, 8 March 1965.
- Due, H. F. and W. E. Young, "Investigation of Pressure Prediction Methods for Radial Flow Impellers, Phase III", Pratt and Whitney Aircraft Report FR-1716, 4 Feb. 1966.

- McGrew, J. M. and H. N Ketola, "Non-contacting Dynamic Seals", General Electric Report on "Study of Dynamic and Static Seals for Liquid Rocket Engines, Final Report, Jan. 1, 1965-Sept. 1, 1965".
- 14. Wood, G. M., D. V. Manfredi and J. E. Cygnor, "Centrifugal Types of Dynamic Shaft Seals", ASME Paper 63-WA-167, 1963.
- 15. Thew, M. T., "Further Experiments on the Hydrodynamic Disk Seal", Proc. 4th Int. Conf. on Fluid Sealing, British Hydro. Res. Assoc., Paper 39, April 1969.
- 16. McGrew, J. M. and H. N. Ketola, "Non-contacting Dynamic Seals", General Electric Report on "Study of Dynamic and Static Seals for Liquid Rocket Engines, Final Report, Apr. 1, 1964-Oct. 1, 1964."
- Thew, M. T., "The Labyrinth/Hydrodynamic Disk Seal (LHDS)", Proc. 8th Int. Conf. on Fluid Sealing, British Hydromechanics Research Assoc., Paper Bl, Sept. 1978.
- 18. Kraev, M. V., "Nature of Liquid-Gas Interface on Bladed Face of a Rotating Disk", Soviet Aeronautics, Vol. 19, No. 1, 116-118, 1976.
- 19. Rosenthal, R. L. and E. Reshotko, "Laminar Flow Analysis of a Plane Disk Slinger Seal", Case Institute of Technology FTAS/TR-66-14, Aug. 1966.
- Reshotko, E. and R. L. Rosenthal, "Fluid Dynamic Considerations in the Design of Slinger Seals", Lubrication Engineering, Vol. 24, 303-314, July 1968.
- 21. Schlichting, H., "Boundary Layer Theory", 6th ed., McGraw-Hill, New York, 1968.
- 22. Eckardt, D., "Detailed Flow Investigations Within a High Speed Centrifugal Compressor Impeller", Journal of Fluids Engineering, Sept. 1976.
- 23. Wislicenus, G. F., "Fluid Mechanics of Turbomachinery", 2d ed., Dover, New York, 1965.
- 24. Daily, J. W., W. D. Ernst, and V. V. Asbedian, "Enclosed Rotating Disks with Superposed Throughflow: Mean Steady and Periodic Unsteady Characteristics of Induced Flow", MIT Hydrodynamics Laboratory Report No. 64, April 1964.
- 25. Due, H. F., "An Empirical Method for Calculating Radial Pressure Distribution on Rotating Disks", Journal of Engineering For Power, April 1966.

- 26. Van Wylen, G. J., and R. E. Sonntag, "Fundamentals of Classical Thermodynamics", 2d ed., SI Version, John Wiley and Sons, Inc., New York, 1976.
- Hendricks, R. C., et al., "Gasp A Computer Code for Calculating the Thermodynamic and Transport Properties for Ten Fluids: Parahydrogen, Helium, Neon, Methane, Nitrogen, Carbon Monoxide, Oxygen, Fluorine, Argon, and Carbon Dioxide", NASA TN D-7808, Feb. 1975.
- 28. "ASRDI Oxygen Technology Survey, Vol. I, Thermophysical Properties", NASA SP 3071, 1972.
- 29. Anderson, P. G., et al., "Fluid Flow Analysis of the SSME High Pressure Oxidizer Turbopump", Lockheed Missile and Space Company, Inc., HREC TR D568277, Aug. 1978.
- 30. "Rocketdyne SSME FMOF Engine Balance", Nov. 1978.

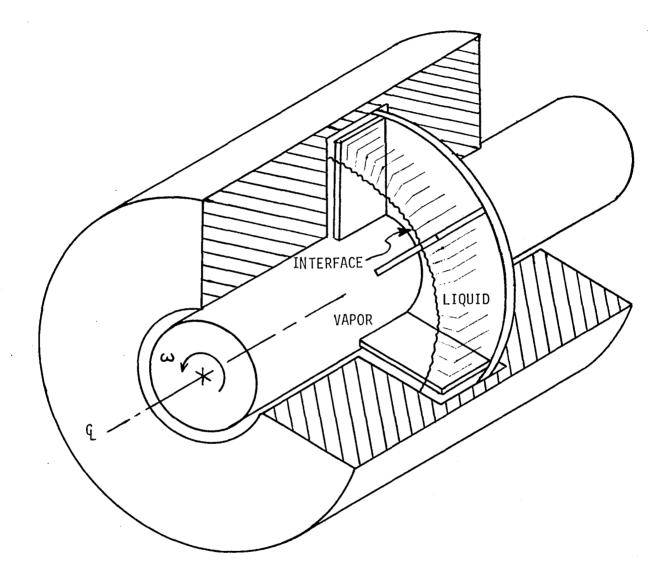


Figure 1. Cutaway view of slinger seal model

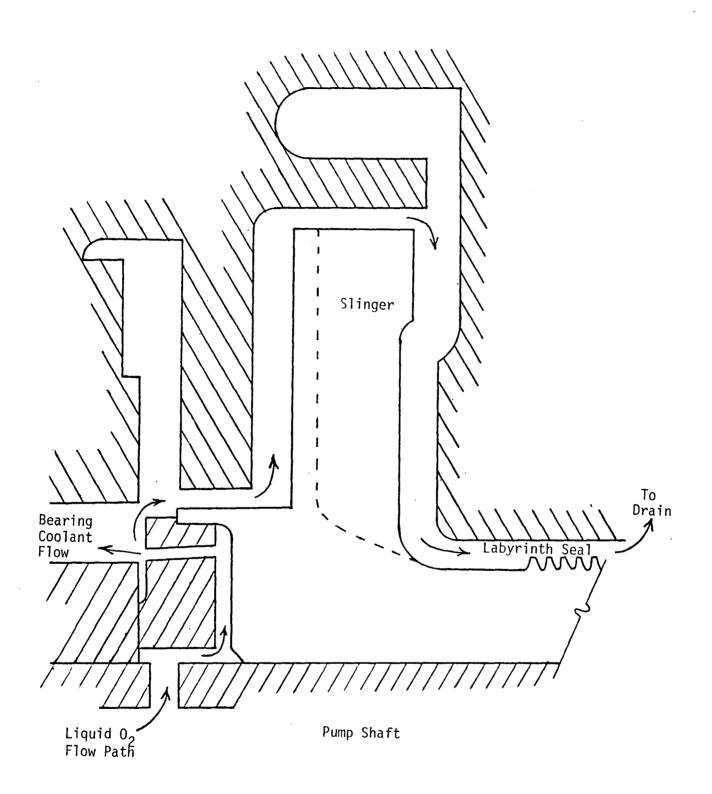


Figure 2. Actual slinger configuration

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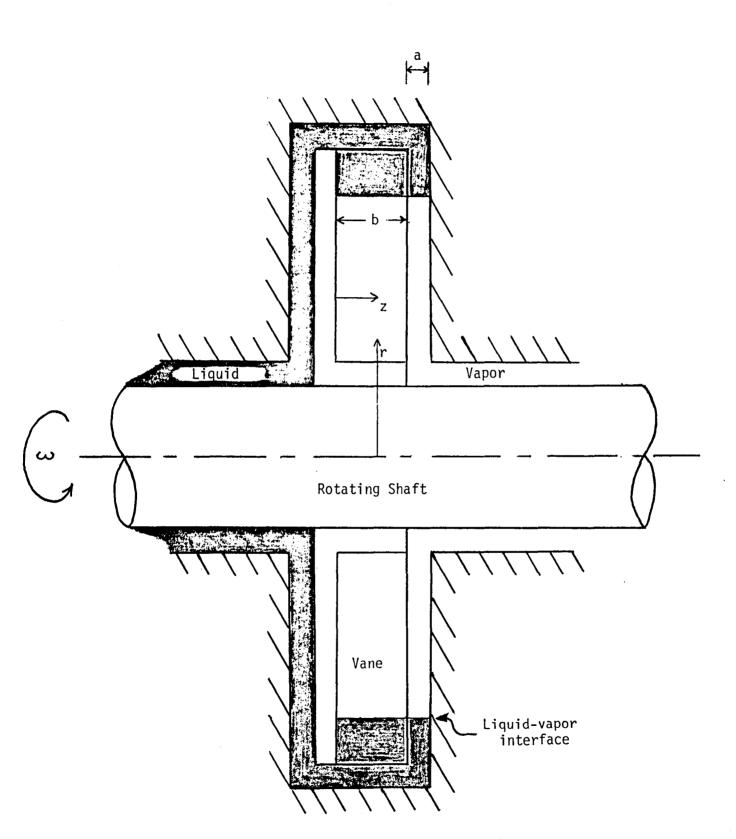


Figure 3. Slinger model with notation.

TABLE 1

SLINGER RADIAL PROPERTY DISTRIBUTIONS

Radius	Velocity	Moment	Enthalpy Pressure		Temperature	Density	
(in.)	Ratio	Coefficient	(BTU/1bm) (1bf/in ²)		(^O R)	(1bm/ft ³)	
1.7	.44	.0008	-34.3	334	216	61.2	
1.8	.44	.0008	-32.5	361	219	60.5	
1.9	.44	.0008	-30.3	388	223	59	
2.0	.44	.0008	-27.5	416	230	58	
2.1	.44	.0008	-24.3	446	237	56	
2.2	.44	.0008	-20.6 476		245	54	
2.24	.44	.0008	-18.9 488		247	53	

Smooth Side

Vaned Side

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Velocity Ratio	Moment Coefficient	Enthalpy (BTU/lbm)	Pressure (lbf/in. ²)	Temperature (^O R)	Density (1bm/ft ³)	Quality
		-18.9	488	247	53	0.0
.88	.0040	-14.8	464	254	51	0.0
.88	.0039	-11.2	442	258	48	0.0
.88	.0037	- 8.0	421	255	42	0.1
.89	.0036	- 5.4	402	253	35	0.2
	Ratio .88 .88 .88 .88	Ratio Coefficient .88 .0040 .88 .0039 .88 .0037	Ratio Coefficient Enthalpy (BTU/1bm) .88 .0040 -18.9 .88 .0039 -11.2 .88 .0037 - 8.0	Ratio Coefficient Enthalpy (BTU/lbm) Pressure (lbf/in.2) -18.9 488 .88 .0040 -14.8 464 .88 .0039 -11.2 442 .88 .0037 - 8.0 421	Ratio Coefficient Enthalpy (BTU/lbm) Pressure (lbf/in.2) Temperature (°R) .88 .0040 -14.8 464 254 .88 .0039 -11.2 442 258 .88 .0037 - 8.0 421 255	RatioCoefficientEnthalpy (BTU/lbm)Pressure (lbf/in.2)Temperature (OR)Density (lbm/ft3).88.0040-18.948824753.88.0039-14.846425451.88.0039-11.244225848.88.0037- 8.042125542

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VALIDATION OF THE SOLAR HEATING AND COOLING HIGH SPEED PERFORMANCS (HISPER) COMPUTER CODE

Prepared By:

Academic Rank:

University and Department

Donald B. Wallace, Ph.D., P.E.

Associate Professor

The University of Alabama in Huntsville Department of Mechanical Engineering

NASA/MSFC: (Laboratory) (Division) (Branch)

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VALIDATION OF THE SOLAR HEATING AND COOLING HIGH SPEED PERFORMANCE (HI S PER) COMPUTER CODE

By

Donald B. Wallace, Ph.D., P. E. Associate Professor of Mechanical Engineering The University of Alabama in Huntsville Huntsville, Alabama 35809

ABSTRACT

Computer simulations of solar energy systems are used to evaluate system performance, size system components and predict fossil fuel savings. The solar heating and cooling high speed performance (HISPER) computer code was previously developed to give a quick and accurate predictions. As a simplification of the TRNSYS program, it achieves its computational speed by not simulating detailed system operations or performing detailed load computations.

In order to validate the HISPER computer code for air systems the simulation was compared to the actual performance of an operational test site. The site selected was the Home Builders Association of Huntsville office building in Huntsville, Alabama. Solar insolation, ambient temperature, water usage rate, and water main temperatures from the data tapes for the site were used as input for the HISPER program.

The HISPER program was found to predict the heating loads and solar fraction of the loads with errors of less than ten percent. Good correlation was found on both a seasonal basis and a monthly basis. Several parameters (such as infiltration rate and the outside ambient temperature above which heating is not required) were found to require careful selection for accurate simulation.

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INTRODUCTION

Solar energy collection systems are installed on dwellings in order to reduce the occupant's dependence on conventional forms of energy for heating and cooling. Due to the diffuse nature of solar energy and the variability in the weather, it is impractical to install a system large enough to provide all of the energy needs. A back-up system, using conventional energy, must be provided for supplying energy during long periods of cloudy weather. The solar energy system is ideally sized to maximimize the benefit to cost ratio (the ratio of annual solar energy provided divided by the life cycle cost of the solar energy system). Because of the large number of design parameters and the weather variability, the design optimization problem is a difficult one.

Computer programs have been developed to help predict solar heating and cooling (SHAC) performance. One widely used program is TRNSYS which was developed at the University of Wisconsin Solar Energy Laboratory [1]*. This is a general purpose program that allows the user to simulate a system by defining the performance of and the interaction between various types of components (collectors, storage, pumps, control logic, etc.). TRNSYS gives an accurate indication of the model performance but it does require a significant amount of computer time to run. In a design situation it is more desirable to have an analysis tool which can quickly and economically give reasonable answers. This allows more design iterations to be analyzed.

In response to the need for a rapid solar simulation computer program, Northrup Services, Inc. (under contract to NASA) developed HISPER - a high speed performance program [2]. HISPER is basically a simplification of the TRNSYS program in which some detailed system operation and load calculations are not performed (such as room temperature oscillations and ductwork heat losses). When HISPER was compared to TRNSYS for identical cases, the predicted system performance differed by less than 7 percent while computer time required by TRNSYS was approximately 30 times greater than the time required by HISPER.

The usefulness of any simulation is dependent on how well it predicts the performance of the actual system. This study provides this validation of the HISPER program by comparing the predicted results using actual weather data

*numbers in brackets refer to references at the end of this paper.

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with the measured performance of the site during this same period. The Home Builders Association of Huntsville office building was selected as the site to be simulated.

This report first describes the Home Builders Association office building and its operation over a period from September, 1978 to April, 1979. Next, the basic organization of the HISPER program and the modifications necessary to run actual weather data is presented. Third, the simulation input data is discussed. The fourth part is a discussion of the results of the study including the effects of several critical design parameters. Finally the results of this study are discussed and recommendations made for the use of HISPER and for topics which need further study.

OBJECTIVES

The primary objective of this study are to verify the general computational philosophy of the HISPER program by demonstrating that the program adequately predicts the performance of actual solar energy systems. This verification will result in more confidence in HISPER for use as a design tool.

Secondary objectives include: identification of the effects of parameter values on the simulation performance, identification of any desirable changes in the HISPER logic, and the establishment of a post-design review procedure in which the actual performance of a site can be compared with the predicted design value.

OPERATIONAL TEST SITE

Site Description:

The operational test site selected for use in this study is the Home Builders Association of Huntsville office building located in Huntsville, Alabama (see Figure 1). This building has two stories with a total floor area of approximately 210 square meters (2270 square feet). The building was designed by IBM (under contract to NASA) to be a single family dwelling but is being used as an office building at this location.

The south facing roof is sloped at 45° and contains 30 solar collectors (manufactured by Solar Energy Products) for a gross area of 67 square meters (720 square feet). These collectors are single glazed and use air as the energy transport medium.

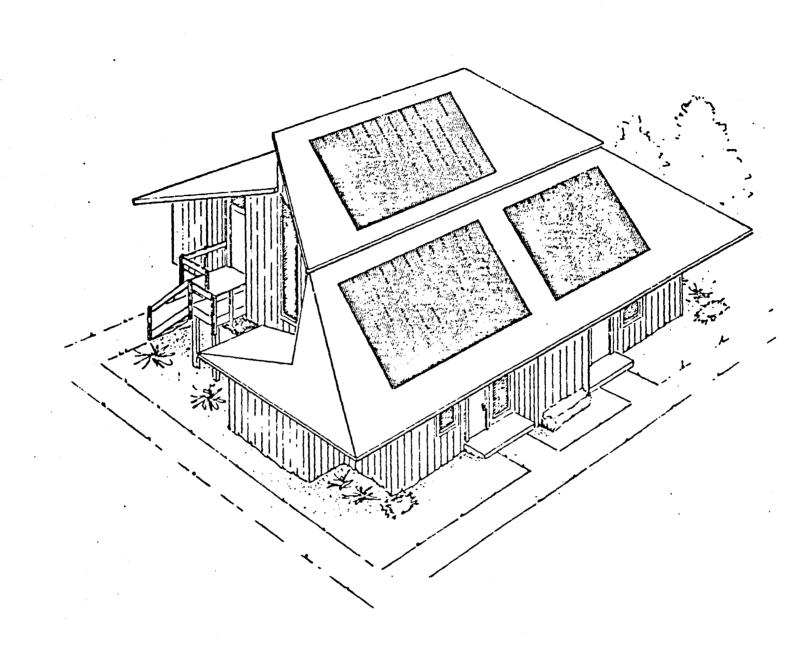


Figure 1. Office Building for the Home Builders Association of Huntsville, Alabama Figure 2 shows a schematic of the solar energy collection system. Air heated in the solar collectors first passes through an air/water heat exchanger so that the domestic water going to the hot water heater (not shown) may be preheated. Then the hot air is directed by the air handles either to the house or to the pebble bed storage, depending upon the energy needs of the house at that time. The pebble bed storage has a volume of 14 cubic meters (500 cubic feet) and contains 23 metric tons (25 tons) of river rock.

During periods when no solar energy is being collected, house air may be passed through the pebble bed thus providing energy to the house. If the temperature of the pebble bed storage is too low (below $32^{\circ}C$ or $90^{\circ}F$), a Westinghouse air-to-air heat pump provides the heat demanded by the house.

A complete description of the building and the solar energy system design and operation is given in various reports prepared by the designers - the IBM Corporation Federal Systems Division [3, 4, 5].

Operational Test Site Data:

The Home Builders Association office building was equipped with 45 sensors to obtain data for performance evaluation. These include 27 temperature sensors, 7 electric power recorders, 5 flow rate sensors, a solar insolation indicator, and 5 others. Each of these sensors is scanned at 320 second intervals (5 minutes and 20 seconds) and their average value is recorded by a data logger. All of this information is then put on to magnetic tapes.

This data is used by IBM to prepare a monthly performance summary for the site. The performance summary includes the climatic conditions, house heating and hot water loads, and solar collection information. The monthly summaries also include any known problems (such as malfunctioning equipment) which occur during the month.

The period chosen for the comparison of HISPER and the actual system was October 1978 through April, 1979. This was chosen since both the 5 minute recorder data (on magnetic tape) and the monthly reports were available. Summer months were not included since the real system is put in "Summer Mode" in which the pebble bed storage is not used. HISPER does not have the capabilities of the "Summer Mode".

A review of the information on the data tapes indicated that the data tapes could not be used directly with the HISPER program. First, periodic sensor malfunctions were discovered (such as ambient temperature going from about $29^{\circ}C$ ($84^{\circ}F$) to $-20^{\circ}C$ ($-4^{\circ}F$) for a period of about 45 minutes and then back again). Secondly, gaps in the data of up to 30 hours was observed on several occasions. These problems were overcome by correcting the obviously erroneous data and inserting reasonable values for the missing data. This data manipulation will result in a small discrepancy in the IBM Monthly Reports and the HISPER predictions. Since the total amount of bad or missing data is

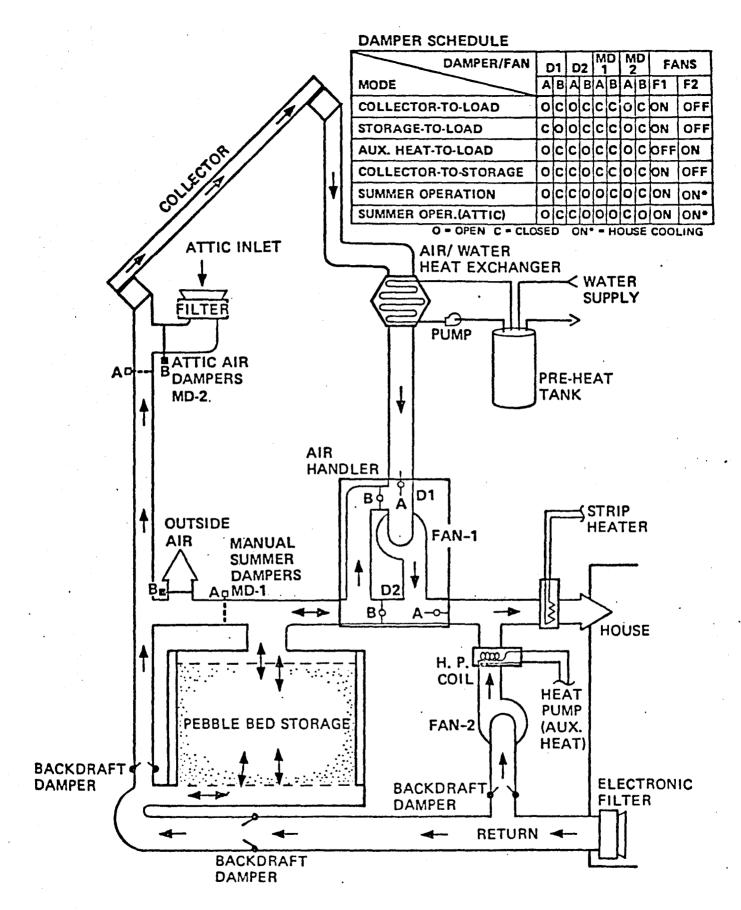


Figure 2. Actual System Configuration

very small, the discrepancy is insignificant.

Site Operation:

A review of the test site data revealed several areas which could lead to significant errors. These relate to hot water usage and temperature control of the house.

Since the site is being used as an office building rather than a residence, the demand for hot water was quite low. When the hot water tank was set at $63^{\circ}C$ (145°F), before November 22, 1978, an average of only 30kg (7.8 gallons) of hot water were used daily. After the hot water temperature was lowered to 51°C (123°F) on November 22, 1978, the hot water demand rose to an average of 36kg (9.5 gallons) a day. Very little water was used on the weekends.

The temperature of the water coming out of the water main fluctuates over the period of the year. Water main temperature varies from a high in August at about $25^{\circ}C$ ($77^{\circ}F$) to a low in February at about $9^{\circ}C$ ($48^{\circ}F$). This variation probably results from the water being transported in mains that are not buried deep enough to be unaffected by seasonal temperature variations.

Another problem is illustrated in Figure 3. March 24th and 25th were cold weekend days following a rather mild week (high ambient temperatures). Apparently someone had shut off the thermostat before leaving work on Friday so that when the workers returned on Monday, the building temperature was $8^{\circ}C$ (46°F). There is other evidence that the thermostat settings had been adjusted over the period of investigation.

Figure 3 also shows that a significant variation in room temperature does occur on a daily basis. This may be due to a direct solar gain either through the windows or through the walls. The resulting heating in the occupied area will reduce the amount of solar or auxiliary energy required.

HISPER MODIFICATIONS

The HISPER simulation program consists of a main program and three fundamental subroutines: MODEL, WETHR, and SOLAR. The purpose of each of these subprograms is:

1. MAIN - The main program reads the system parameters, performs the time integration, and controls the output of the program. The main program calls subroutine MODEL.

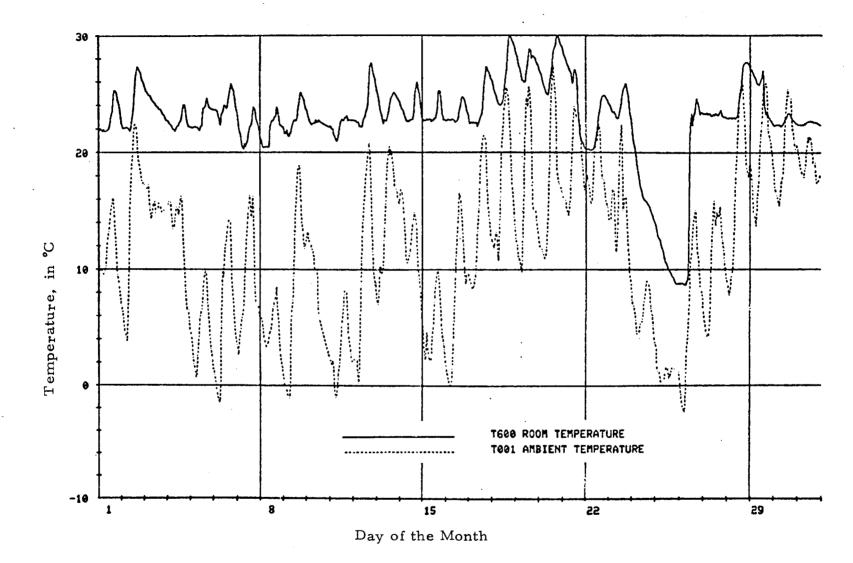


Figure 3. Ambient and Room Temperatures in March, 1979

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- 2. MODEL This subroutine calculates all of the energy flows in the system model in a given increment of time. It calculates collector performance, the dwelling heating load, the hot water load, and the flow of energy in the solar system (including in to and out of storage). It calls WETHR and SOLAR.
- 3. WETHR This subroutine provides ambient temperature and horizontal solar insolation to MODEL. These values are read from weather tapes containing typical weather data.
- 4. SOLAR This subroutine receives the total hourly horizontal insolation output by the WETHR subroutine and converts it to total insolation incident on the collector surface using one of four user determined methods.

The basic HISPER procedure, with the exception of WETHR, was not changed significantly. Some control logic was added to MODEL and different forms of output were provided for in MAIN. Subroutine SOLAR was not altered at all. Since the climatic data was from a new source, WETHR had to be changed completely.

HISPER was designed to read weather tapes which contain typical climatic data for a particular location. In order to validate HISPER predictions for an actual operational test site, it was necessary to modify HISPER to read the actual climatic conditions at that site. Sensors at the Home Builders Association office building include an ambient temperature sensor and a pyranometer to measure total solar insolation in the collector plane.

HISPER uses both insolation on a horizontal surface (for load calculations) and insolation on the collector plane (for solar collector performance) in its simulation procedure. Normally the horizontal insolation is provided by the weather tapes and HISPER uses one of four standard techniques to estimate the tilted insolation from the horizontal value (see references [1, 6, 7]). Since the site data includes only tilted insolation data, HISPER's WETHR subroutine had to be modified to reverse the standard procedures and predict the horizontal insolation value which will produce the observed tilted value when SOLAR is called.

HISPER is designed to simulate a typical domestic dwelling so the assumed hot water load represents what a normal family might use each day. Since the system being studied is an office building, the HISPER predicted daily usage (224kg -- 59 gal) does not correspond well to the actual usage of about 33kg (8.7 gallons) per day. Therefore HISPER was modified to read (in WETHR) actual water usage and true water main temperature from the site data. The control logic used in HISPER is designed to provide the maximum amount of energy to the loads as possible. The main control features include:

- 1. Turning on the collector when ever energy can be gained.
- 2. Energy is extracted by the hot water preheat tank whenever the collector is on and the outlet temperature is greater than the preheat tank temperature.
- 3. Energy is provided to the house whenever the house has a load and the air inlet temperature, either from the collector or storage, is greater than the house temperature (returns air temperature is assumed to be room temperature).
- 4. Energy is not allowed to go to storage if the temperature from the collector is less than the temperature at the top of the storage bed.

In order to study the effects of actual control settings on the HISPER Simulation performance, the following features were added to the MODEL subroutine control logic.

- 1. A minimum temperature difference between the collector plate temperature and the bottom of the pebble bed storage bin before the collector would collect energy (note: the plate temperature was based on steady state flow conditions) Actual site required a 15°C (28°F) temperature difference [8].
- 2. A minimum temperature difference between the collector outlet temperature and the preheat tank for domestic hot water before energy could be extracted from the air and be put in the preheat tank. Actual site required 11°C (20°F) to turn on and no less than 2°C (3°F) before turning off [8].
- 3. A minimum temperature at the top of the pebble bed storage was necessary for the storage to be allowed to provide energy to the house. Actual site required a temperature from the storage of at least 32°C (90°F). [8].
- 4. An air return temperature from the house that was lower than the nominal house temperature. Review of data tapes indicates that the return air is approximately 5°C (9°F) below room temperature. This is probably due to thermal stratification in the rooms.

The control logic as modified by these four features will be referred to as the "modified control logic."

SIMULATION INPUT DATA

A variety of sources was used to develop the input data for the HISPER simulation program. The attempt was to make the model for the Home Builders of Huntsville office building as realistic as possible.

The manufacturer's data for the sunworks solar collectors was used since other tests of the collectors by Wyle Laboratories showed good correlation with the predicted performance. Manufacturers data was also used for the air to water heat exchanger used for preheating water.

Flow rates in the air ducts were taken from actual tests performed by IBM after the construction of the test site.

Internally generated heat loads were assumed to be 2890 kJ/hr. This corresponds to the average heat generation at 10 people, lights, 2 refrigerators, and three coffee pots. The storage losses (since the storage bed is in the building, is considered to be part of the internally generated heat.

The parameters used for heat load calculations are primarily based on a design analysis done at NASA/MSFC in 1977. After the Home Builders Association building was constructed, tests were conducted to determine actual heat loss coefficients. The results of the study indicate that the building load demand is commensurate with the 720 BTUH/^oF design building loss coefficient-area product (UA).

The only significant difference between the NASA/MSFC design values and the values used in the HISPER simulation was for infiltration. The NASA/MSFC value was one half air change per hour which assumes very tight construction. The tests performed on the building were conducted at night with no employees present at the office building. ASHRAE [9] recommends design infiltration rates from one half to two air changes per hour, depending on the number of windows and doors in the building. Since the building is being used as an office, with the possibility of many people entering and leaving during the day, an air infiltration rate of 1.5 air changes per hour was used.

SIMULATION RESULTS

HISPER was designed to give a reasonable prediction of solar performance of a home. This means that it should be able to predict the loads and the amount of solar energy delivered to the loads. The accuracy of the predictions was investigated with respect to several of the input parameters: infiltration rate, the outside ambient temperature above which the house is assumed to require no heat (TLH), and the effect of control logic. The domestic hot water predictions were also investigated. Effect of TLH: (Temperature above which no energy is required by the house)

A house does not always require energy whenever the outside ambient temperature is less than the house temperature. This is because any house has some direct solar gain and a large amount of thermal capacitance. The direct solar gain statisfies some of the current house demand as well as being absorbed in the house for use at night. This effect can be seen in the room temperature fluctuations in Figure 3.

The data tapes for the Home Builders Association building were searched to determine what percentage of the time the heating/cooling system was on as a function of the outside temperature. Figure 4 shows this relationship. The increase in the percentages for the higher temperatures is due to cooling, not heating. From Figure 4, it is obvious that if the house is set at $22^{\circ}C$ $(72^{\circ}F)$, it actually does not require significant heating until well below that value, at say $10^{\circ}C$.

Initial simulations were run assuming heating was not required over an ambient temperature of 10° C, i.e. TLH 10° C (50° F). The HISPER results are shown on a monthly load basis in Figure 5 and on a daily load basis in Figure 6 (a). An infiltration rate of 1.1 was assumed so that the total house load for the period was correct. This value of TLH results in HISPER under predicting the actual loads on cold days (heavy loads) and over predicting on warm days (light loads).

A value of TLH of $7^{\circ}C$ (45°F) and an infiltration rate of 1.5 air changes an hour was then used to produce the results in Figure 5 and Figure 6 (b). These values result in a very good load prediction on both a monthly and daily basis.

Effects of Control Logic:

Once the actual loads were matched through the appropriate selection of TLH, the portion of the total load provided by solar energy was investigated. Figure 7 shows that a good correlation was achieved on a monthly basis.

A closer examination of component performance revealed that the temperature at the top of the storage was higher after periods of heavy load than HISPER predicted. (See Figures 8 and 9). This was due to the actual control logic which stopped energy withdrawal at temperatures of less than $32^{\circ}C$ ($90^{\circ}F0$. The storage bed model of HISPER is programmed to supply, on demand, energy to the building anytime the temperature at the top of the storage is greater than the thermostat setting.

The temperature at the bottom of the pebble bed storage at the actual site often reached temperatures $5^{\circ}C$ ($9^{\circ}F$) below the nominal room temperature setting. This is probably due to the return air from the house being at a lower temperature than the room.

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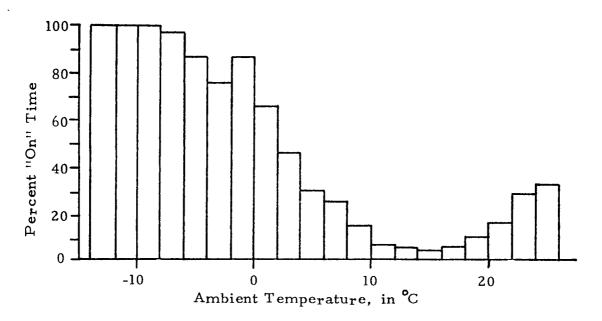


Figure 4. Heating/Cooling System "on" Time

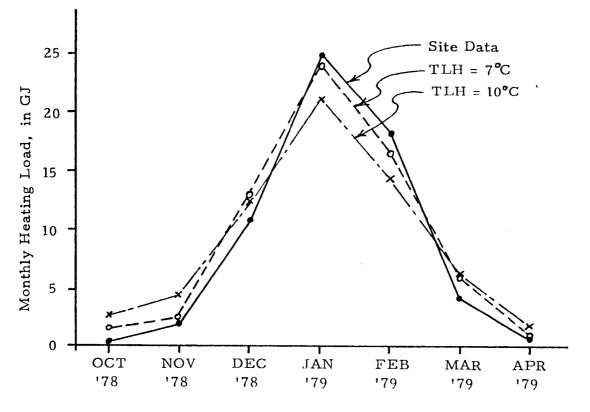


Figure 5. Predictions of Monthly Heating Load

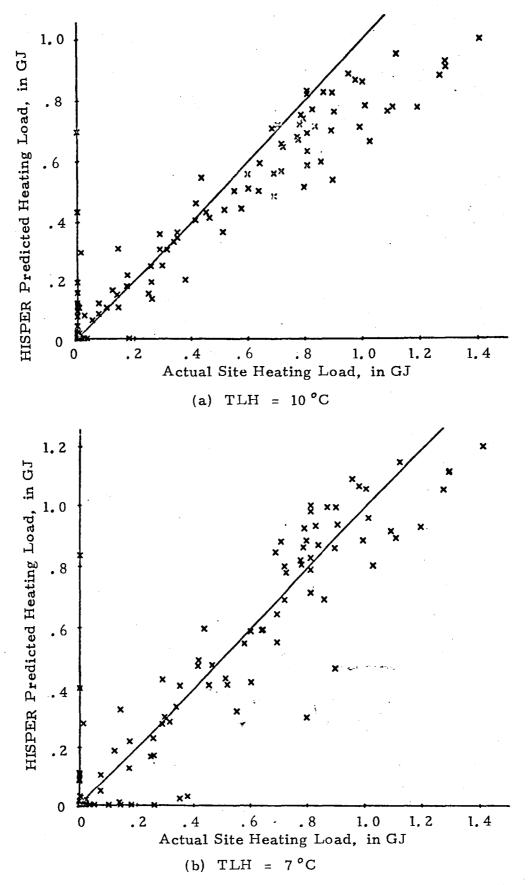


Figure 6. Actual Versus Predicted Daily Heating Loads

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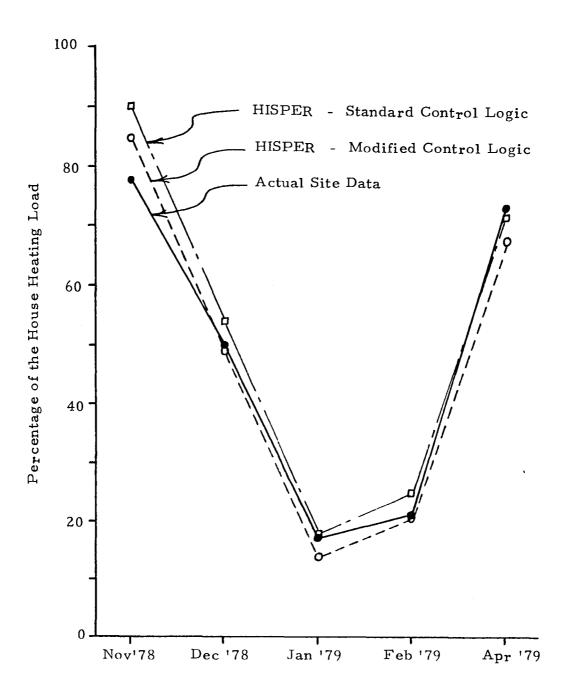
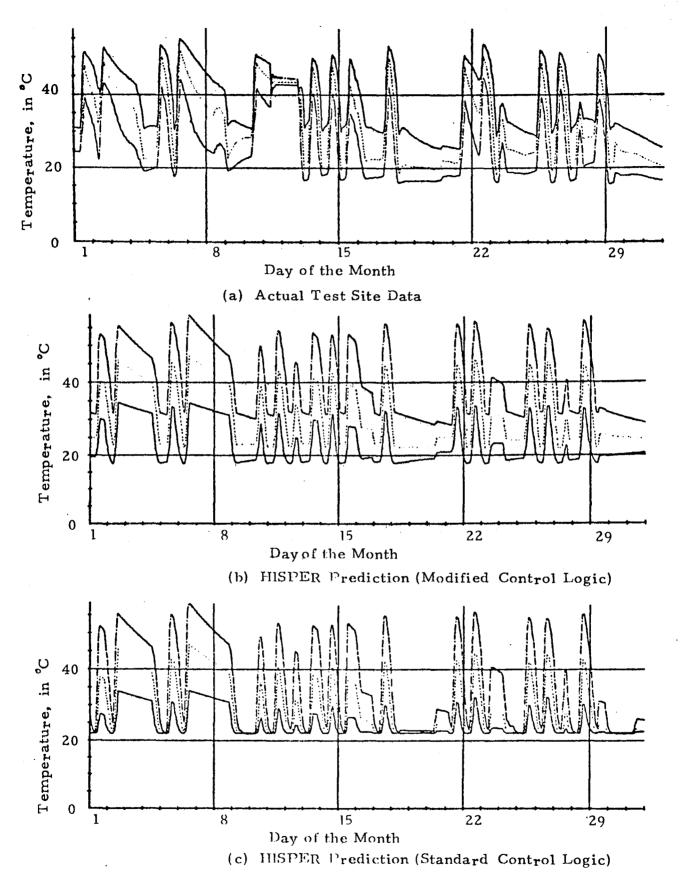
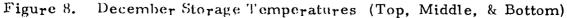
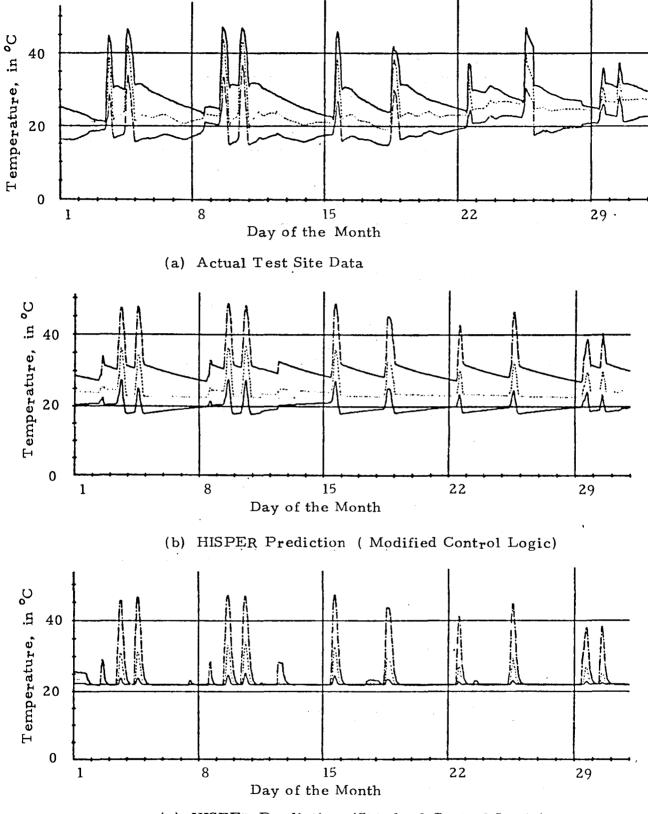


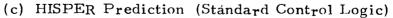
Figure 7. Solar Percent of House Heating Load

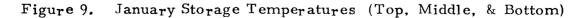




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The four changes to the "control logic detailed in the HISPER Modification section were made with the following values:

- 1. Minimum temperature difference between collector plate and storage bottom: 7°C (13°F)
- 2. Minimum temperature difference between collector output temperature and preheat tank temperature: $5^{\circ}C(9^{\circ}F)$
- 3. Minimum temperature from storage to house: $32^{\circ}C$ (90°F)
- 4. House return temperature: $5^{\circ}C(9^{\circ}F)$ below room temperature.

The effects of this modified control logic can be seen in Figures 7, 8 and 9. The temperature difference criteria for turning on the collector and preheat tank heat exchanger did not have a great effect on the results. The minimum temperature to the house requirement reduced the solar percentages since the storage could no longer be "drained." Lowering the return temperature from the house allowed the collectors to operate more efficiently, collect more energy and thus deliver more to the house. The last two effects tend to counteract one another thus resulting in still a very good prediction of solar percentages on a monthly basis (Figure 7). Figures 8 and 9 show that these changes do result in a significant improvement in the simulation of the pebble bed storage behavior.

Hot Water Percentages:

Hot water loads were accurately predicted since almost all of the information required for the load calculation was read from the data tapes. The percentage of the domestic hot water was not predicted well by HISPER. HISPER predicted a larger solar fraction than actually existed.

The inaccurate prediction is a result of several factors. First, the water usage demand is so low in this application that energy losses are of a much greater magnitude than energy used in heating the water actually used. Therefore any error in heat loss parameters will have a large effect on the solar percentage.

The second effect is that HISPER only simulates heat lost from the preheat tank, not the standard hot water tank. Due to the low usage rate, most of the solar provided heat will be lost before it is used.

The solar percentages of the hot water load for the actual site and the HISPER predictions are shown in Table 1.

Month	Actual Site Percentage	HISPER Predicted Percentage
October, 1978	52	99
November,1978	49	90
December, 1978	64	91
January, 1979	50	68
February, 1979	43	69
March, 1979	59	100
April, 1979	64	100

Table 1. Solar Percent of Hot Water Load

CONCLUSIONS AND RECOMENDATIONS

The major conclusion is that the HISPER solar simulation program seems to be a sufficiently accurate design tool. Errors of less than 10 percent were observed when comparing predicted results to the actual results. One caution, however, is that parameters such as infiltration and TLH (the temperature above which no heating is required) must be carefully selected to achieve good monthly predictions.

Although HISPER predictions have been validated against actual performance, several areas were discovered which need further attention:

- The value of 7°C for TLH (temperature above which heating is not required) seems lower than the site data indicates it should be (Figure 4). This should be investigated for this site and others.
- 2. The validation process should be carried out for several other operational test sites.
- 3. The four control options previously discussed should be added to HISPER. This will give HISPER greater flexibility but will not significantly increase computer time required for simulation.
- 4. The water main temperature was observed to vary from 8°C to 25°C over the period of a year. This variation could significantly affect the predicted hot water load since HISPER currently assumes a constant water main temperature.

REFERENCES

- "TRNSYS A Transient Simulation Program," Solar Energy Laboratory, University of Wisconsin, Engineering Experiment Station Report 38, March 31, 1975.
- Gray, J.C., "Solar Heating and Cooling High Speed Performance Program," Technical Note TN - 240 - 1661, Northrup Services, Inc., Huntsville, Alabama, July, 1976.
- 3. "System Design Package for IBM System One Solar Heating and Domestic Hot Water," <u>DOE/NASA CR - 150614</u>, prepared by IBM Corporation, February, 1977.
- 4. "SIMS Prototype System 1 Design Data Brochure," <u>DOE/NASA</u> <u>CR - 150534</u>, prepared by the IBM Corporation, January, 1978.
- 5. "Installation Package SIMS Prototype System 1A," <u>DOE/NASA</u> <u>CR - 150524</u>, prepared by the IBM Corporation, December, 1976.
- 6. Duffie, J.A. and Beckman, W.A., <u>Solar Energy Thermal Processes</u>, John Wiley and Sons, Inc., 1974.
- Liu, B.Y.H. and Jordan, R.C., "The Interrelationship and Characteristic Distribution of Direct, Diffuse and Total Solar Radiation," Solar Energy, IV (3), July, 1960.
- "SIMS Prototype System 1 Test Results Engineering Analysis," <u>DOE/NASA CR - 150522</u>, prepared by the IBM Corporation, January, 1978.
- 9. <u>ASHRAE Handbook 1977 Fundamentals</u>, American Society of Heating, Refrigerating, and Air-Conditioning Engineers, Inc., 1977.

1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

INCIPIENT FAILURE DETECTION (IFD) OF ANTI-FRICTION BALL BEARINGS

Prepared By:

David Y. Wang, Ph.D.

Academic Rank:

Associate Professor

University and Department:

University of Wisconsin-Plattville Civil Engineering Department

NASA/MSFC:

(Laboratory) (Division) (Branch)

MSFC Counterpart:

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Luke A. Schutzenhofer Jess H. Jones

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INCIPIENT FAILURE DETECTION (IFD) OF BALL BEARINGS

by

David Wang, Ph.D. Associate Professor of Civil Engineering University of Wisconsin - Platteville Platteville, Wisconsin

ABSTRACT

Because of the immense noise dissipated during the operation of a large engine such as the SSME, the unique ball bearing signals are often buried by the overall signature of the machine. As a result, the most commonly used technique, pattern recognition (1)*, to detect and predict ball bearing failures is rendered useless.

In order to retrieve the buried bearing defect signals from the background noises, a survey of two improved techniques was conducted.

Signal averaging method (2) averages n^{**} signals over a time duration equaling the known period of the ball bearing defect frequency. With this method, the desired ball bearing defect signal will be enhanced over the noise by a factor of \sqrt{n} .

A second method is to perform a further frequency analysis on the logarithmic spectrum. The purpose of this operation is to detect and separate different families of harmonics and sidebands, thereby giving information about modulating frequencies.

The above mentioned alternate IFD methods were carried out with a Norland 3001 waveform and data system. The obtained results are then compared with the power spectral density (PSD) of the overall machine signals and conclusions are made accordingly.

^{*} Number within the parentheses refers to the references at the back of the report.

^{**} Number of samples taken.

ACKNOWLEDGEMENTS

The author wishes to acknowledge contributions made to this project by his counterparts at NASA, Dr. L. A. Schutzenhofer and Mr. J. H. Jones. The author also wishes to express his appreciation to those of the Environmental Analysis Branch who provided help in preparing this report.

NOMENCLATURE

f(t)	- complex signal
S(t)	- periodic signal
n(t)	- noise
S/N	- signal to noise ratio
t _k	- beginning time of the kth repetitive signal
ts	- sample time interval
Т	- period
σ	- rms of noise n(t)
fo	- ball pass frequency over outer race
fi	- ball pass frequency over inner race
fs	- ball spin frequency
D _w	- ball diameter
dm	- bearing pitch diameter
α	- contact angle
N	- number of rolling elements
R	- shaft rotative speed in Hz

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INTRODUCTION

Every operating machine will generate a combination of sound and vibration in the form of 'signatures' which are unique to the devices being evaluated. The ability to monitor, analyze, and use this signature to locate vibratory sources and predict future failures is the goal of an IFD system. Pattern recognition is the most commonly used technique in evaluating ball bearing conditions. It is performed by spectrum analyzing the output of an accelerometer signal over a specified frequency range that includes the ball pass frequencies. An abnormal increase in amplitude of this 'signature' over some duration of time indicates deterioration of ball bearings.

During the operation of a large machine such as the SSME, the bearing 'signature' in the form of electronic signals is often buried by the large background noise. Therefore, it is necessary to extract the bearing defect signals from the combined signal before an evaluation of the bearings can be made. This report describes a signal averaging or summation technique to extract the repetitive bearing defect signal from a complex signal with a high level background noise.

A cepstrum (3) type of signal analysis was also carried out. The advantage of cepstrum over a spectrum analysis is that it can separate families of periodic sidebands around some fundamental frequencies. The purpose of this cepstrum analysis is to cross check its results with that obtained using the signature averaging method. The possibility of using cepstrum in evaluating ball bearing conditions was also explored.

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OBJECTIVE

The objective of this project is to seek a data reduction process in order to extract pertinent information from the time record of SSME during its tests. The extracted information is, then, used to predict condition of ball bearings installed in the high pressure oxygen turbopump (HPOTP) of the SSME. Of numerous available methods, signal averaging or summation analysis (3), (7), were chosen to be examined in this report.

SIGNATURE EXTRACTION

A signal with known periodicity, T, is buried deep in noise. To extract that signal by averaging or summation, the total signal plus noise is divided into m segments of duration T. Each segment is sampled at a number of points during every period. The samples are stored in a data storing device such as a magnetic tape which algebraically adds together all m samples.

A repetitive event of period T will always be added in the same sample location. The resulting sum, thus obtained, will be m times the average amplitude of that event. Events which do not repeat with the same period will result in a sum proportional to \sqrt{m} , therefore, the desired signal has been enhanced over the noise by a factor of \sqrt{m} .

The above described signal averaging process can be described as follows. Signal \neq (t) is

$$f(t) = S(t) + n(t) ----$$
(1)

the signal at kth repetition starts at time instant t_k , if taking $t_1 = 0$, is

$$f(t_{k} + it_{s}) = S(t_{k} + it_{s}) + n(t_{k} + it_{s})$$

= $S(it_{s}) + n(t_{k} + it_{s}) ----$ (2)

Since noise, n(t), is random and the m segments of samples are independent, summation of the m noises will yield a mean square value of $m\sigma^2$ and a rms value of $\sqrt{m\sigma}$. Hence, after adding the m segments of the signal, equation (2), which indicates values at the ith location of the signal f(t) within the period T becomes

$$\sum_{k=1}^{m} f(t_k + it_s) = mS(it_s) + \sum_{k=1}^{m} n(t_k + it_s)$$
$$= mS(it_s) + \sqrt{m\sigma}$$

Therefore, the signal to noise ratio, after m summations, is

$$\binom{S}{N}_{m} = \frac{mS}{\sqrt{m\sigma}} = \int m \binom{S}{N}_{m}$$

A Norland 3001 waveform analyzer was used to carry out the signal extraction procedures described above. There are two ways to accomplish signal extraction. One is to stack segments of signal of given period T by shifting the signal segment of duration T, without any gap, m times successively. Another way is to convolute the signal, f(t), with a train of m units impulses at a time delay of the given period T. The latter procedure is described by the following operation.

$$a(t) = \int_{\infty}^{\infty} f(t - \varepsilon) \sum_{k=1}^{m} (\varepsilon - kT) d\varepsilon$$

 $= \sum_{k=1}^{m} f(t - kT)$

CASE HISTORY AND DATA COLLECTION

Failures of bearings within the HPOTP were experienced during tests of the SSME. Spallings, damage cuts, pittings, and discoloration appear on both inner and outer races, rolling elements, and cages. The most serious failure seems to occur at bearings closest to the turbine of HPOTP. It was suspected from evidences of damages, that these failures were caused by excessive loading in the axial direction of the turbopump during the starting stage of the test. In order to be able to predict the oncoming bearing as well as other types of failures, data were collected from several strategically located accelerometers. These signals were usually low pass filtered at 10 KHz or under before storing on magnetic tapes for later analysis. Because of the low pass filtering of data, possibility of adopting various IFD techniques in the higher frequency range (30 KHz and higher) as diagnostic instruments is precluded.

The collected data in the form of time record is time averaged or summed and then transformed to the ferquency domain as power spectral density (PSD) diagrams. This allows one to relate specific periodic occurrence of the bearing to specific frequencies in the data.

In order to analyze the obtained results, ball pass frequencies over a defect were computed according to the following equations.

$$\mathcal{L}_{o} = \frac{N}{2}R(1 + \frac{D_{W}}{d_{m}}\cos\alpha)$$

$$\mathcal{L}_{i} = \frac{N}{2}R(1 - \frac{D_{W}}{d_{m}}\cos\alpha)$$

$$\mathcal{L}_{s} = R\frac{d_{m}}{D_{W}}(1 - \left(\frac{D_{W}}{d_{m}}\right)^{2}\cos^{2}\alpha)$$

For the given bearing configuration, various ball pass frequencies and other relevant data (at 100% SSME power) are listed as follows:

 $f_0 = 2570 \text{ Hz}, f_1 = 3530 \text{ Hz}, f_s = 2920 \text{ Hz}$ N = 13, R = 469 Hz

Static capacity of the bearing 🛎 3050 lbs.

DATA ANALYSIS

Results obtained from the test carried out in the Spring of 1980 are presented in figure 1 through figure 11. Figure 1 is a collection of three PSD's for the initial, middle, and final stage of the bearing life span. Narrow band spikes at shaft rotative speed and turbine blade frequencies appear in the diagram at 469 Hz, 1836 Hz and 3672 Hz, as expected. Note the growth of the narrow band spike at 2570 Hz region which indicates a possible fault on the outer race. Defects on the inner race and rolling elements are not obvious in this particular figure.

Figure 5 through figure 10 are the so-called zoom PSD's at the same time instants as those in figures 1 through figure 4. It is more desirable thatn the normal PSD's because it provides a considerably finer resolution over the portion of the spectrum that we are interested in. Again, possible defects on the outer race appear on these spectra in the neighborhood of 2600 Hz region. Finally, a useful technique in detecting the deterioration of bearings with respect to time is presented in the form of a three dimensional PSD with time as the third parameter.

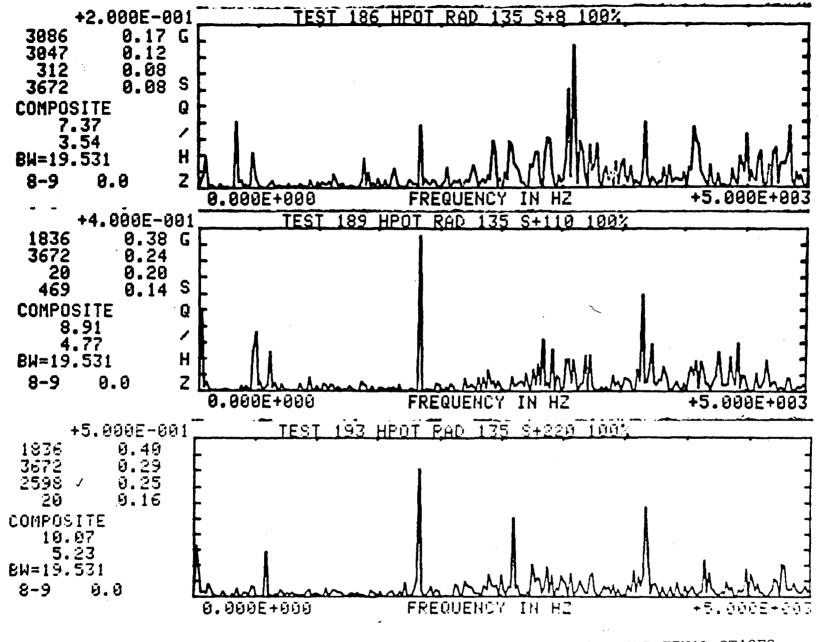
CONCLUSIONS AND RECOMMENDATIONS

A signal averaging or summation process has been described that is used to extract the bearing defect signals form the time record of SSME. It appears that this procedure can also be applied to the zoom and three-dimensional PSD for finer resolution and better presnetation of the deterioration history of ball bearings. An electronic equipment of black box type based on this principle could be constructed and adopted as part of an "on-line" IFD system. One drawback in applying this technique is that the period, T, of a repetitive signal cannot be absolutely time-locked in an analog to digital converter because of the limitation of the sampling time interval. The result of this deficiency is the tendency of the summed time signal being smeared gradually. Signal to noise gain will be decreased accordingly.

Even though other techniques such as "acoustic emission" (10) and "shock pulse" have been proposed, they have not been proven successful. One alternative method in ball bearing defects detection is to utilize the bearing high frequencies (13), (14), especially for large machines. Natural frequencies of ball bearings are usually far beyond that of the noise level, therefore, less noise interference to the bearing signal. In order to adopt this method of approach, real time signal should be collected without any low pass filtering.

REFERENCES

- 1. Taylor, J. I., "Determination of Anti-Friction Bearing Conditions by Spectral Analysis," Progress of Machinery Vibration Monitoring and Analysis Seminar, February 1977.
- 2. Randall, R. B., "Cepstrum Analysis and Gear Box Fault Diagnosis," B&D Application Note No. 13-150.
- 3. Weichbrodt, B., "Mechanical Signature Analysis, A New Tool for Product Assurance and Early Fault Detection," Annals of Reliability and Maintainability," Vol. 5, AIAA, 1966.
- 4. Hower, D. R., "Signal Analysis Techniques for Vibration Diagnostics," 22nd Meeting of Mechanical Failure Prevention Group, October 1976.
- 5. Harris, T. A., "Rolling Bearing Analysis," Wiley, New York, 1966.
- 6. Bickel, H. J., "Real Time Spectrum Analysis," Sound and Vibration, Sol. 5, No. 3, March 1971.
- 7. Trimble, C.R., "What is signal Averaging?" Hewlett_Packard Journal, April 1968.
- 8. Fieldhouse, K., "Techniques for identifying Sources of Noise," Sound and Vibration, Vol. 4, No. 12, 1970.
- 9. Bannister, R. L., and Donato, V., "Signature Analysis of Turbomachinery," Sound and Vibration, September 1971.
- 10. Balderston, H. L., "The Detection of Incipient Failure in Bearings," Materials Evaluation, June 1969.
- 11. Braun, S., "Signal Analysis for Rotating Machinery Vibrations," Pattern Recognition, Vol. 7, 1975.
- 12. Bendat, J. S., and Piersol, A. G., "Random Data Analysis and Measurement Procedures," Wiley-Interscience, New York, 1958.
- 13. Harting, D. R., "Incipient Failure Detection by Demodulated Resonance Analysis," Instrumentation Technology, September 1977.
- 14. Burchill, R. F., "Resonant Structure Techniques for Bearing Fault Analysis," 18th Meeting of Mechanical Failures Prevention Group, November 1971.
- 15. Randall, B., "Frequency Analysis," Bruel & Kjar, September 1977, 2nd edition.



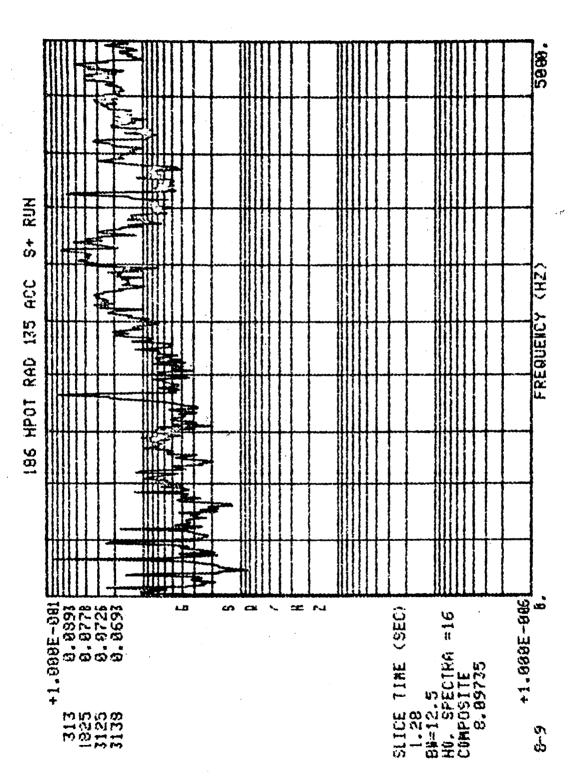
NARROW BAND FREQUENCY ANALYSIS - INITIAL, MIDDLE, AND FINAL STAGES

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Figure

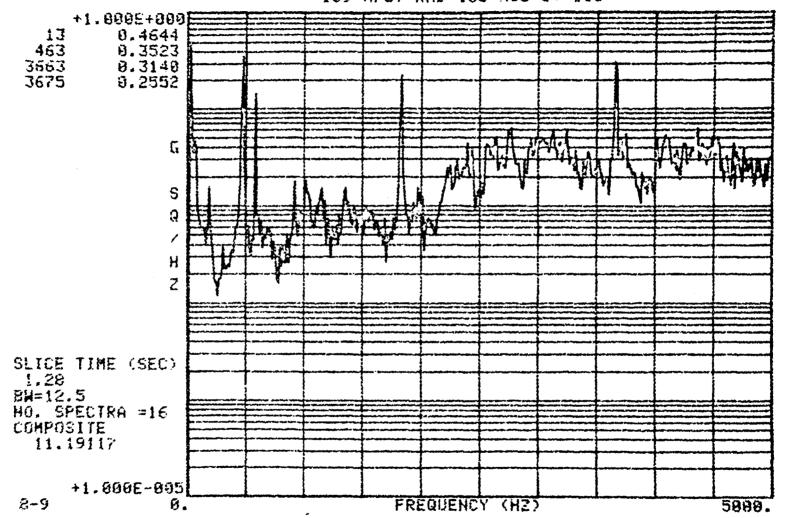
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PSD (INITIAL STAGE)



189 HPDT RAD 135 ACC S+ 110

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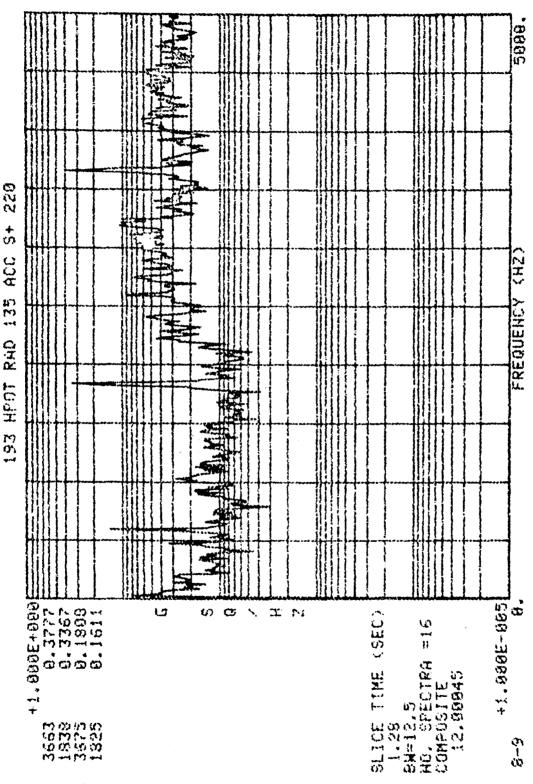
PSD (MIDDLE STAGE)

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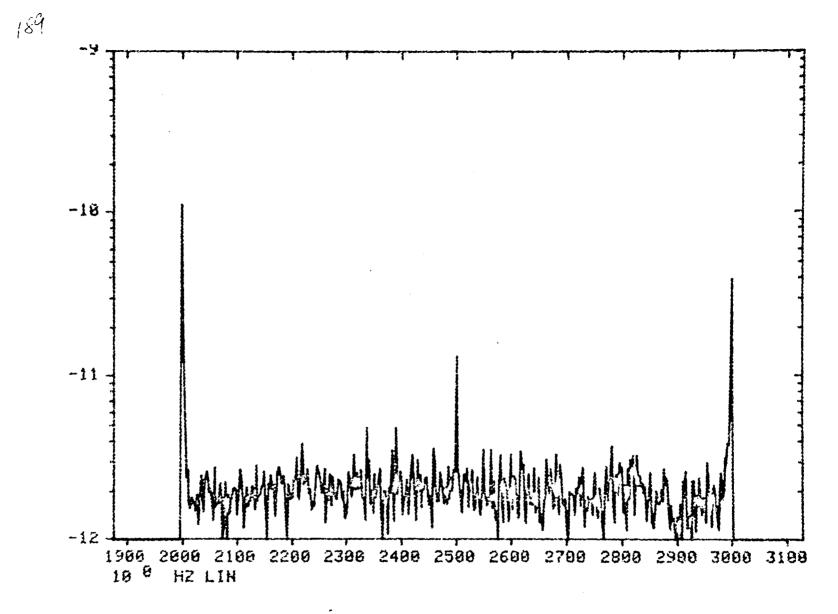
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(FINAL STAGE)

PSD

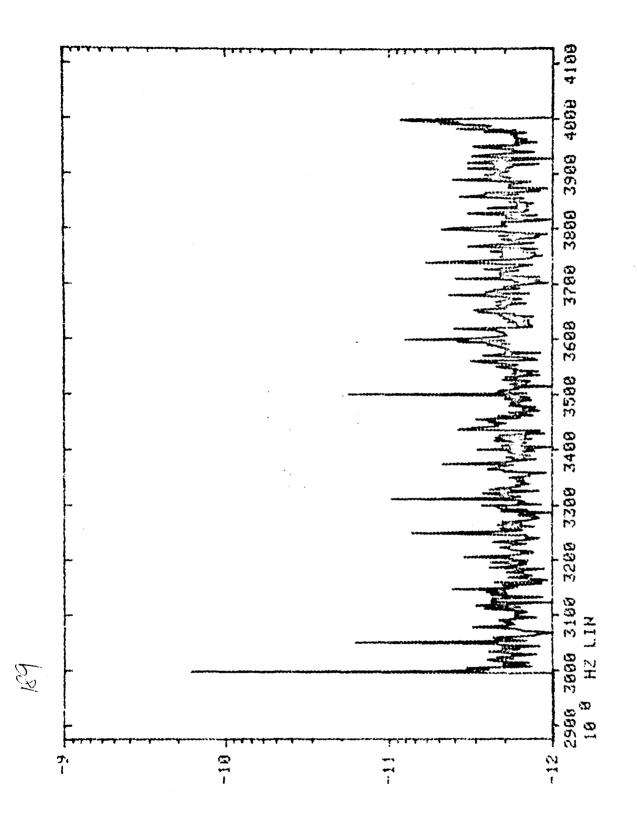


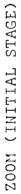
+ ഗ ACC 117 HPOT RHD



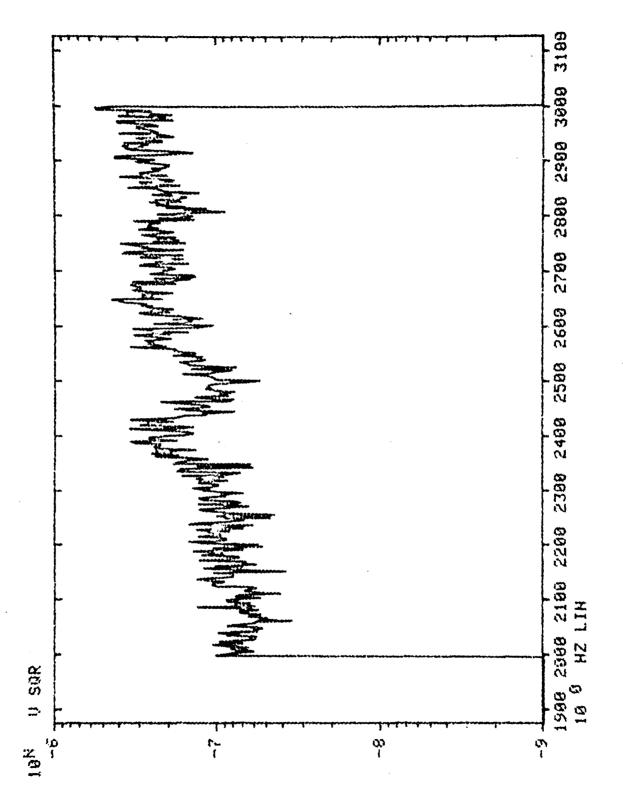
ZOOM (INITIAL STAGE)

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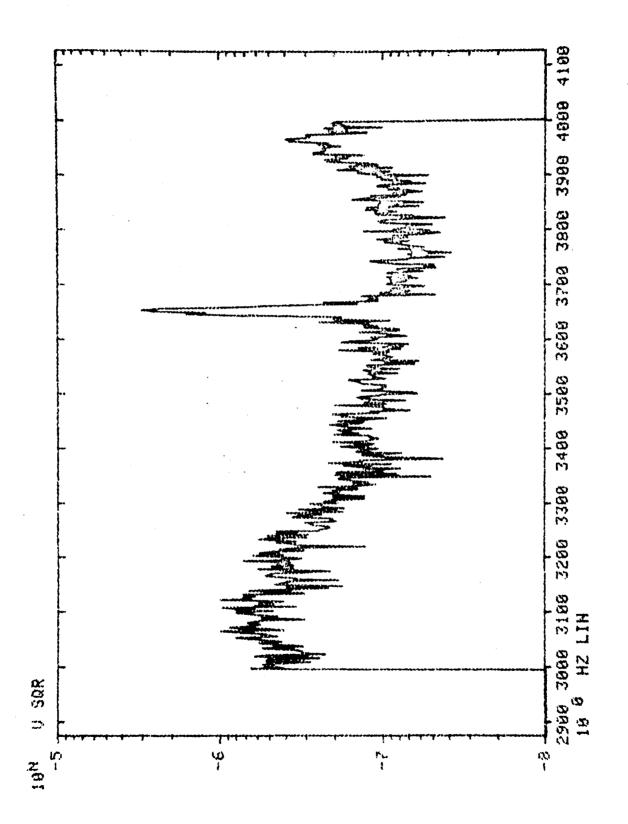


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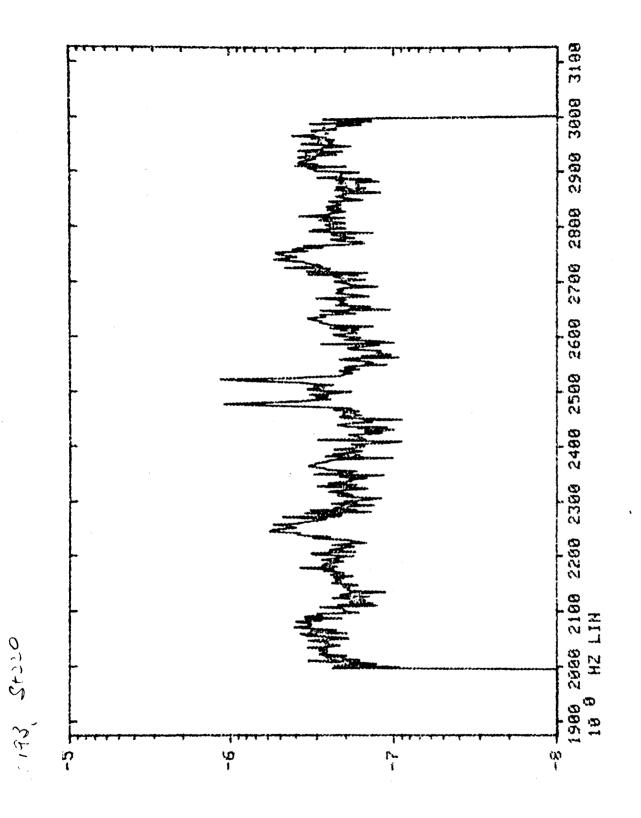




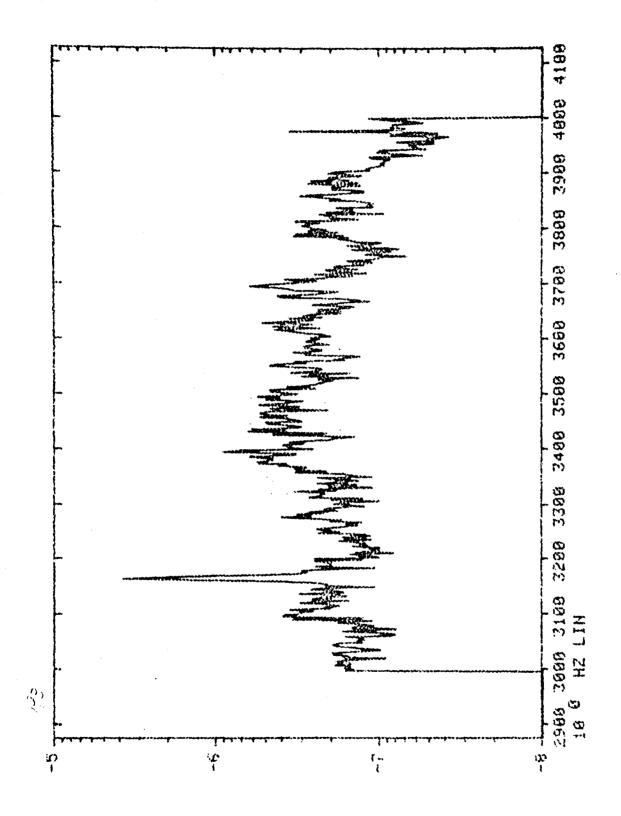
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ZOOM (MIDDLE STAGE)

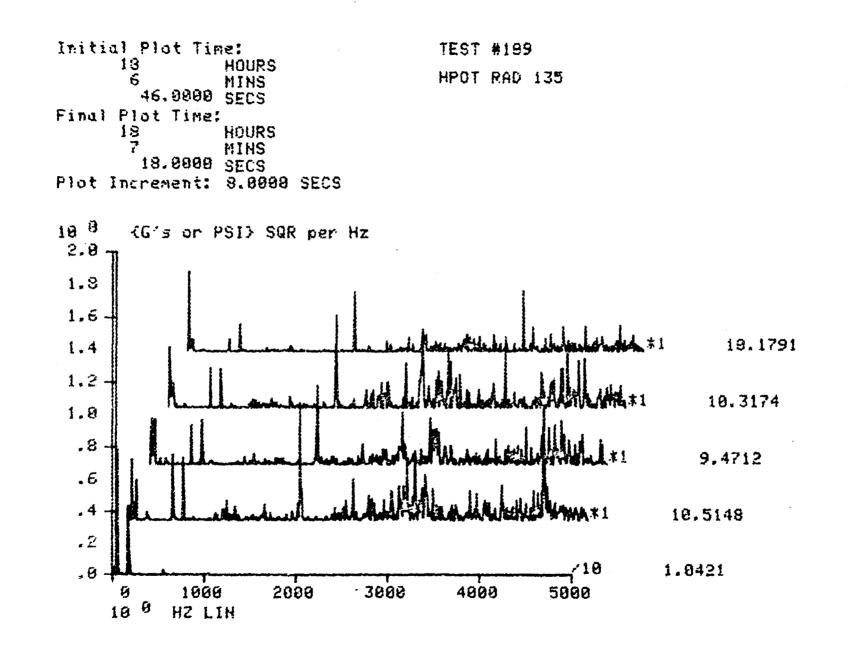


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L. St.

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THREE DIMENSIONAL PSD

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1980

NASA/ASEE FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

PULSED EDDY CURRENT TESTING

Prepared By:

Gary L. Workman, Ph.D.

Academic Rank:

Associate Professor

Athens State College

University and Department:

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Materials and Processes Engineering Physics Nondestructive Testing

Department of Scientific Technologies

J. M. Knadler, III

August 1, 1980

NGT-01-002-099 (University of Alabama)

PULSED EDDY CURRENT TESTING

ΒY

Gary L. Workman, PH.D. Associate Professor of Scientific Technologies Athens State College Athens, Alabama 35611

ABSTRACT

The traditional nondestructive testing technique used for inspecting welds in large space structures, such as the "External Tank", has been radiographic testing, in conjunction with surface inspection methods, such as dye penetrant testing. In order to improve upon the economics and reliability of these critical inspections, new approaches to the test methods are under consideration. One approach is to replace the very expensive radiographic inspection with an automated eddy current testing technique. A description of pulsed eddy current and its relationship to multifrequency techniques if presented here, as well as a some preliminary results obtained from observing pulsed waveforms with apparatus and algorithms currently in use for ultrasonic testing of welds. It can be shown the pulsed eddy current techniques can provide similar results.

Eddy current testing has primarily been used for sorting of materials due to differences in conductivities, detecting differences in surface due to corrosion on cracks, and in some instances to detect flaws below the surface. Manual operation of the test proves to be very unreliable due to such test parameters as lift off, non-uniform conductivity, depth of penetration, surface treatment, and complex geometries of the test object. The use of a pulsed eddy current techniques can eliminate some of the noncritical parameters affecting the eddy current signals and facilitate in the detection of critical parameter such as flaws, subsurface voids, and corrosion.

Acknowledgements

I wish to acknowledge Mr. John M. Knalder, III, for allowing me to conduct the research this summer and Mr. Ken Woodis for assisting me in setting up the computer interfaces. Also Mr. Ben Bankston and Mr. Claude Williamon, for implementing experimental fixtures. I also thank Miss Pearline Stone for typing the report.

Introduction

The nondestructive inspection of large space structures requires that a large part of the inspection procedure be automated in order to reduce human errors and to reduce the amount of time and money spent in this critical phase of production. Since a large number of the procedures used for inspecting the External Tank are concerned with determining flaws in welds, there is a need to develop an inspection technique, which can be automated, to determine flaws in welds and structures with complex geometries. The techniques; whereby one generates eddy current in a metallic material and observes the changes in the circuit parameters due to material differences, has been chosen as one possible approach. The principles of eddy current testing has been treated in a number of publications (1-3). Of particular interest to the researcher is the book by Libby¹ and the work of Dodd³.

Eddy currents are generated in metallic specimens by alternating currents in coils in close proximity to the material. Some of the properties associated with eddy currents are:

1

- They constitute a closed circuit; hence, in the material, they usually have a circular configuration
- 2. They penetrate into the material with decreasing

amplitude away from the surface.

- Variations in part distances introduces electrical phase differences.
- The magnitude of eddy current flow is proportional to the current in the test coil.
- Both the electrical conductivity and the magnetic permeasbility affect the eddy current flow in the material.
- Discontinuities which affect the eddy current flow are easily detected as the eddy current redistribute themselves around the discontinuity.

These general properties have enabled eddy current techniques to be useful for a number of applications. Some of the most notable are:

- 1. Measurements of electrical conductivity
- 2. Measurements of magnetic permeability
- 3. Sorting of metallic materials
- 4. Measurements of distances between components
- 5. Thickness determinations
- Detections of voids on discontinuities in metallic materials

A number of commerical eddy current instruments are available for these applications. The reliability and sensitivity of each test depends upon many parameters, for instance:

- 1. Coil-to-conductor spacing and geometry
- 2. Choice of frequency for a particular conductor
- 3. Orientation of defects to coil
- 4. Sample geometry
- 5. Coil characteristics

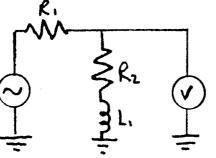
Such a large number of parameters affecting an eddy current measurement accounts for the range and flexibility of the technique, as well as its pitfalls. For instance, in the determination of flaws in welds, one has to keep all other parameters constant so that only the interaction of the eddy currents with a flaw would be observed as a signal. Even with a judicious laboratory arrangement, this is very difficult to accomplish. When testing complex geometries, in a production environment, it is practically impossible to maintain that many parameters constant. Also, most materials exhibit minute changes in composition which can affect the eddy current measurements and this can be misinterpreted as a flaw. Since it can be so difficult to make a single eddy current measurement and interpret the results correctly, many researchers in the field are developing methods which take a number of measurements and solve simultaneously for different parameters which can change. The use of computers in both hardware control and data acquisition and analysis allows these approaches to become easily automated.

Objectives

The primary objectives of this work is to determine the feasibility of using pulsed eddy current methods for nondestructive testing. As described in the report there are many ways of handling the data received from the eddy current test. Our approach has been to pulse an eddy current probe, use present FFT data processing and determine if the information contained in the frequency domain does allow one to solve for the properties of the material under inspection.

Multiparameter Methods

The electrical quanities of interest in an eddy current measurement are usually a voltage or current and its phase relationship to the driving voltage or current. Eddy current circuits are most easily represented as simple RLC circuits and their signal responses represented as impedance of the material being inspected. Hence impedance plane plots are most often used to illustrate and differentiate between the different parameters affecting the eddy current measurement. For example, if we consider a simple coil and driver arrangement such as



then if $V_0 = V_0 \cos w t$ (i.e. a sinusoidal waveform) then

$$V = V_0 \left(\frac{Z}{R+Z}\right) = V_0 \left|\frac{Z}{R+Z}\right|$$
$$= V_0 \sqrt{\frac{(R_2 + 2\pi f L_1)^2}{(R_1 + R_2 + 2\pi f L_1)^2}}$$

Notice that the measurement of a single voltage can determine only a single parameter or the combined effect of several parameters. In order to differentiate between the parameters in an eddy current measurement, one has to synthesize an excitation signal which then can be decomposed into frequencies affected by the desired parameters. One approach is to use multifrequency excitation from which we can extract a voltage and a phase value for each frequency. For example, if we use two frequencies, then we have four measurements for each test and

where A_1 , A_2 , A_3 , A_4 , represent the material properties as coefficients in response to the eddy current source and each Pij represents the magnitude of that effect in each measurement. V_1 through V_4 represents either a voltage or phase measurement. This procedure has been developed and successfully demonstrated by Dr. Dodd and his Union Carbide Corp. in Oak Ridge(4,5). Although the parameters A_1 through A_4 appear to be linearly independent, experimentally it can often be difficult to determine unique solutions and thus to identify such effects as lift-off, voids, or corrosion. However, their approach was to synthesize known types of defects, liftoff's, etc into the laboratory calibration. Then, during an inspection procedure, their system uses a pattern recognition approach to go back and fit the experimental data

to the calibration data in order to determine what responses are contained in each measurement. To date, this approach has worked very well in the inspection of tubing used in nuclear reactor heat exchangers.

The frequencies chosen for this approach use a high frequency signal for surface effects and a lower frequency (say 1/10) for multilayer identification, and even another lower frequency (say 1/100) for detection of voids farther away from the surface.

Pulsed Eddy Current Techniques

The multifrequency technique provides a workable solution to the eddy current testing problem. The major drawback to the method is that it is expensive. A multifrequency head using several coils will prove to be extremely costly. In addition to the data acquisition circuit, precise frequency filters are required to maintain a reliable signal-to-noise ratio for each phase and amplitude measurement. A promising technique, which should prove to be less expensive, is to use a pulsed excitation rather than the continuous alternating current as in the previous work. In reality pulsed eddy current and multifrequency eddy current provide similar information. For instance, a short duration pulse can be transformed into the frequency domain and the resulting wave form consists of a large number of frequencies. Hence, we can also produce multifrequency excitation using a pulse with a finite time period.

Pulsed eddy current data can be obtained in two different ways, although the two approaches are essentially equivalent. The pulsed eddy current experiment can be performed in a manner similar to the multifrequency experiment discussed above, the only deviation is using a finite pulse instead of several continuous wave frequencies. The interaction of the coil with the eddy currents is still dependent on the inherent frequency of interaction. Hence, one would still measure amplitudes and phases of the resulting signals and perform the data reduction in exactly the same way as before. The two methods are equivalent in that same sense.

The other approach one might take in using pulsed eddy currents is to actually transform the time domain data into the frequency domain and use the multifrequency analytical approach with these transformed values. Our approach this summer has been to obtain data in this mode in order to determine the feasibility of the approach. A number of algorithms currently in use for ultrasonic imaging applications one applicable, as well as a hardwired Fast Fourier Transform Analyzer (6). Hence we felt that this approach might proceed the quickest for a summer's work.

Experimental Results

All of the measurements taken in this work were obtained with a Microtek model 101 ultrasonic pulsing system driving a Kervonics reflection eddy current coil. The pick-up coil signals were digitized by a 8-bit A/D and stored in a Biomation Transient Analyzer. Model #8100. The data was then transferred to a PDP11/45 computer and analyzed via software generated by contractor personnel. An example of the data obtained here is shown in Figure 1-4. The pick-up signal itself is shown in Figure 1 and its FFT is shown in Figure 2. These observations were obtained statically with a Kervonics 40A probe (pick-up coil radius=0.7 cm) on aluminum.

Lift-off simulation is shown in Figure 3 and the resulting waveform when over a 20 mil surface defect in shown in Figure 4.

Note the changes which occur for the frequency domain data, particularly at 0.5, at 1.3, and at 3 MHZ. In all three cases the pick-up signal changes somewhat but not as quantitative as given in the frequency data. Obviously then the analysis of a three frequency probe at 0.5, 1.3, and 3 MHZ should provide the same type of information, but would require three frequencies with separate pick-up coils. Since the scanning mechanism, which was expected to be used in this work, never did arrive; all the measurements were made using a specially inprovised mount for the probe. Since no motion occurred during the measurement, all tests were static.

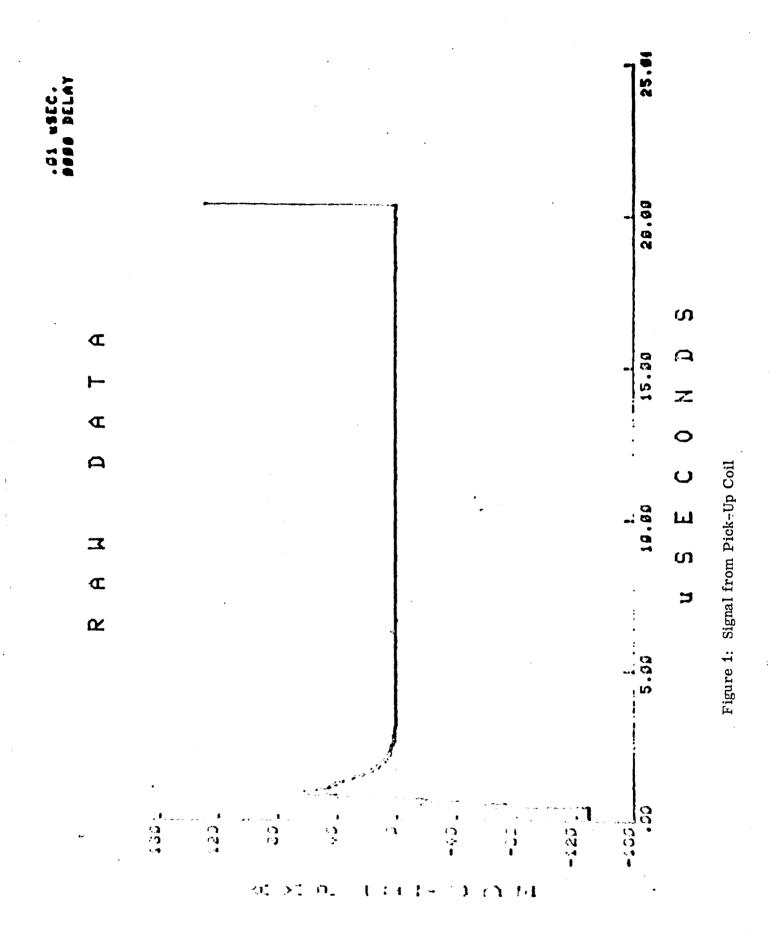
Conclusions and Recommendations

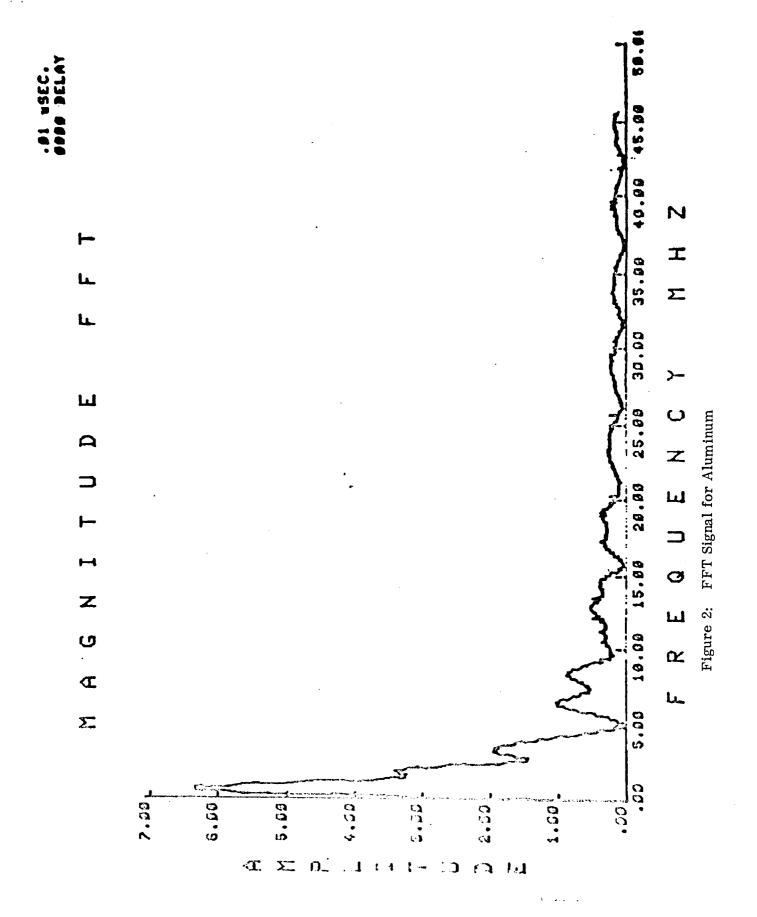
The pulsed eddy current technique can differentiate between various materials properties using FFT analysis. Since we were not able to obtain data while scanning over a specimin, that should be the next objective in this project. Also, in view of the information published by Dr. Dodd and his associates, the following recommendations can be made.

- Improve the resolution of the A/D converter presently being used by replacing it with at least a 12-bit A/D converter.
- Use waveform synthesis techniques to generate a pulse which contains those frequencies of interest for a particular inspection. A microprocessor generated pulse can do this.
- Implement the pattern recognition algorithms at Marshall Space Flight Center. Upon request most of his software is available.

If pulsed eddy current is to be developed at MSFC, a full time effort should be provided. The nondestructive testing requirements for large spacecraft will be more easily implemented using either a pulsed or multi-frequency method.

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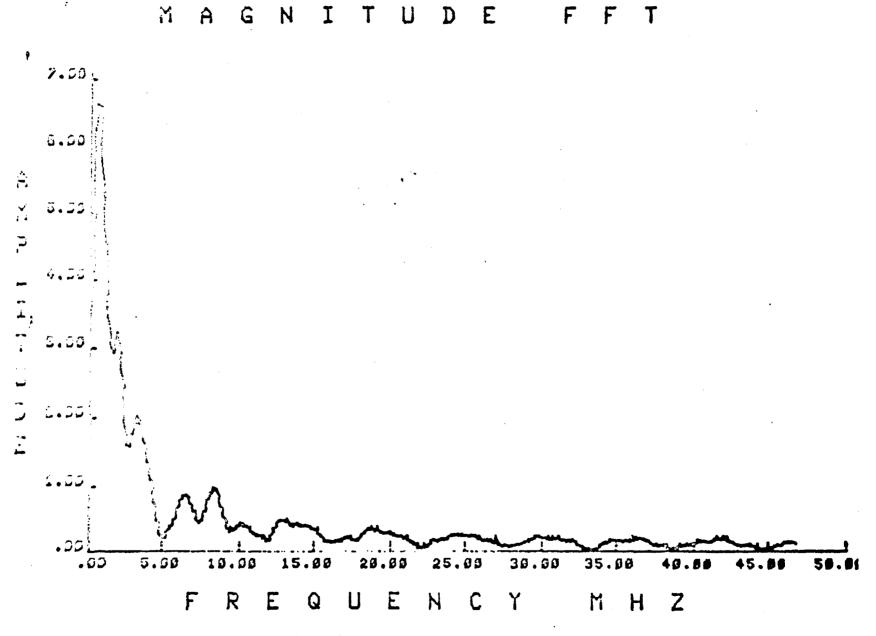


Figure 3: FFT Signal for Lift-Off

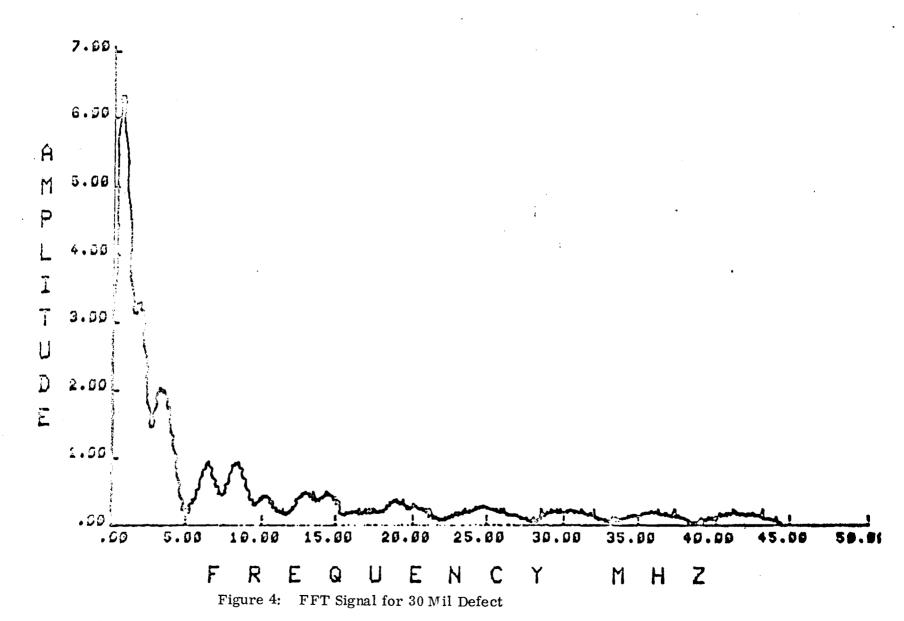
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MAGNITUDE FFT



REFERENCES

- 1. Libby, H. L., Introduction to Electromagnetic Nondestructive Test Methods, Krieger Pub., Huntington, N.Y., 1979.
- "Eddy Current Inspection", Nondestructive Inspection and Quality Control, vol. 11 of Metals Handbook, ASM. Metals Park, Ohio, 1976.
- 3. Dodd, C. V., <u>Basic Principles of Eddy Current Testing</u>, Internal document, Union Carbide Corp., Oak Ridee, TN, 1978.
- Deeds, W. E., Dodd, C. V., Scott, G. W.; "Computer-Aided Design of Multifrequency Eddy Current Tests for Layered Conductors with Multiple Property Variations", ORNL/TM-6858, 1979.
- 5. Dodd, C. V. and Deeds, W. E., "Multiple Property Variations in Coaxial Cylindrical Conductors Determined with Multiple-Frequency Eddy Currents".
- 6. Smith, G.R., "Interactive Signal Analysis and Ultrasonic Data Collection System"; Final Report, Contract MAS 8-32024.

1980

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALBAMA

STABILITY OF THE STRATIFIED CYLINDRICAL ANNULUS FLOW

Prepared By:

Academic Rank:

University and Department:

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Basil N. Antar, Ph.D.

Associate Professor

University of Tennessee Space Institute Department of Engineering Science and Mechanics

Space Sciences Atmospheric Sciences Fluid Dynamics

W. W. Fowlis

August 1, 1980

NAT-01-002-099 (University of Alabama)

STABILITY OF THE STRATIFIED CYLINDRICAL ANNULUS FLOW

by

Basil N. Antar Assistant Professor of Engineering Science and Mechanics The University of Tennessee Space Institute Tullahoma, Tennessee 37388

ABSTRACT

The linear stability analysis for the stratified flow between two rotating circular cylinders is formulated. Two approaches for the stability analysis are presented. The first approach results in an algebraic eigenvalue problem, while the second results in an initial value problem for the perturbation function. The advantages and disadvantages of both approaches are discussed and a preferable numerical solution technique is outlined.

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The linear stability analysis for the stratified flow between two rotating circular cylinders is formulated. Two approaches for the stability analysis are presented. The first approach results in an algebraic eigenvalue problem, while the second results in an initial value problem for the perturbation function. The advantages and disadvantages of both approaches are discussed and a preferable numerical solution technique is outlined.

INTRODUCTION

It is generally believed that the existence of flucuations in the circulation of the atmosphere can be attributed to the instability of the stationary dynamical state without fluctuations. Partial confirmation of this belief has recently come from detailed analysis of the numerical general circulation models. However, no exact laboratory experiments have been conducted of the atmospheric circulation in order to test this hypothesis. The absence of such a laboratory model is basically due to the difficulty arising from the large earth gravity field which is ever present in any terrestrial experiment. Recently hopes for the possibility of setting up an experiment model of the earth's global circulation has been rekindled again through the availability of the gravity free environment of the forthcoming missions of Spacelab. Indeed an experiment to model and study atmospheric fluctuations arising from stability considerations in a true spherical geometry is being considered by NASA for a future Spacelab Mission [1].

The existence of fluctuations in the flow inside a rotating cylindrical annulus in which the inner cylinder is cooled while the outer cylinder is heated have been observed experimentally by Fultz et al. [2] and Fowlis and Hide [3]. These fluctuations were found to resemble in a remarkable way those that are observed in the atmosphere although the geometries are different. Due to this striking resemblence between the waves in the annulus experiments and the waves in the atmosphere, the flow field in the rotating cylindrical annulus has been studied extensively (see Hide and Mason [4] for an up to date discussion on the annulus problem).

One of the major findings common to both the analytical and experimental studies of the annulus problem is that waves of different azimuthal wave lengths appear only for specific value of the governing parameters of the problem. This fact is usually shown by plotting a curve in the Taylor-Froude number plane which separates the wave regime from the no-wave (symmetric) regime. Such a curve has been obtained experimentally by Fowlis and Hide [3]. However, in the theoretical treatment of the exact annulus problem only few points on that curve have been obtained, originally by Williams [5] and subsequently by Quon [6]. These theoretical confirmations

have been obtained through the complete, three-dimensional, numerical solution of the governing equations for the annulus geometry. These numerical solutions of the exact problem consume excessive amounts of computer time to perform. For this reason the regime diagram which was obtained experimentally have not been reproduced in full through numerical solutions, since the amounts of computer time needed to perform these calculations are beyond the scope of present day computers.

In this report we will outline a different numerical method for obtaining the regime diagrams for the annulus Since the existence of the observed fluctuations problem. in the annulus is due to instability of the basic dynamical state, it then seems reasonable that a stability analysis of the basic state, symmetric annulus flow should predict under what conditions the fluctuations will appear. Now, it is well known that stability analysis requires far less calculations than the total solution of the governing system of equations. Hence, such an analysis may be efficiently used to generate the desired regime diagrams. In this report two methods of stability analysis will be presented and the advantages and disadvantages of each method will be discussed in detail. Furthermore, a recommendation will be given on the best numerical implementation of the techniques in question.

STABILITY ANALYSIS

As indicated in the introduction there are basically two analytical approaches that may be used to verify the experimental regime diagrams. The first approach is to integrate the initial-boundary value problem starting from a state of no motion and following the solution in time. This approach requires the soltuion of the full governing equation in three-dimensions and for long enough time to discern whether fluctuations will appear or not. However, the computation time that may be necessary to implement this approach could be prohibitive. Quon [6] has indicated that as much as 150 hours of computer time on the Univac 1108 is necessary to establish only one point in the regime diagram using this approach.

The alternative approach which is subsequently outlined is to use linear stability analysis to establish the regime diagrams. This approach may be implemented in one of two ways both of which will be discussed at length.

The linear stability analysis starts by first formulating the governing equation for the annulus flow in cylindrical geometry. These equations are the Navier-Stokes equations, the energy equation and the mass conservation equation. The Navier-Stokes equations in cylindrical geometry in a rotating reference frame, are the following:

$$\begin{aligned} u_{t} + uu_{r} + & \underbrace{\psi}_{u_{q}} + wu_{t} - \underbrace{\psi}_{r}^{2} - z \mathcal{N} \psi = -\frac{1}{2} P_{r} \\ + \sqrt{(\nabla^{2} u_{r} - \frac{2}{r} v_{q} - \frac{u_{r}}{r})} \end{aligned} \tag{1}$$

$$w_{1} + uw_{2} + \frac{\omega}{r}w_{q} + ww_{2} = -\frac{1}{5}p_{2} + \sqrt{\nu}w + \alpha gT \qquad (3)$$

where

$$\nabla^2 f = f_{rr} + \frac{1}{rc} f_{qq} + f_{zz} + \frac{1}{r} f_r$$

and

In Eq. (3) the Boussinesq approximation has been used. In the above equations u, v and w are the radial (r), aximuthal (ϕ) and axial (z) velocity components, respectively. p and ρ are the pressure and density; v is the kinematic viscosity; α is the volumetric expansion coefficient and T is the temperature. Ω is the angular velocity of both cylinders. In conjunction with above system the energy and mass conservation equations are needed which are given by

$$\overline{T_{t}} + U\overline{T_{t}} + \stackrel{\omega}{=} \overline{T_{q}} + \omega \overline{T_{t}} = K \overline{D^{2}T}$$

$$\tag{4}$$

$$u_{r} + \frac{1}{r} v_{\bar{q}} + w_{\bar{z}} + \frac{u}{r} = 0 \tag{5}$$

where κ is the thermal diffusivity of the fluid.

Next, all the field variables are separated into a stationary mean part and a perturbation part in the following way:

$$U = u_0 + u'$$

$$V = V_0 + V'$$

$$w = w_0 + w'$$

$$P = P_0 + P'$$

$$T = T_0 + T'$$

When the above variables are substituted into the governning system of equations (1)-(5) and using the fact that the mean variables (f_o) themselves satisfy equations (1)-(5)a system of equations governing the perturbation functions results. However, this latter system of equations is as complicated as the original system with the added complexity of having the basic state functions appearing as coefficients in the equations. This means that in order to solve for the perturbation functions one needs to solve for the basic state first.

If it is further assumed that the perturbation functions are very small, i.e.infinitesimal then it is possible to linearize the system of equation for the perturbations. These equations will be of the following form after dropping the primes:

$$u_{i} + u_{0}u_{r} + \frac{U_{0}}{r}u_{q} + w_{0}u_{z} + uu_{0r} + wu_{0z} - 2\frac{U_{0}}{r}u - 2NU = -\frac{1}{r}u_{r} + v(v_{1}u - \frac{2}{r}u_{q} - \frac{u}{r})$$
(6)

$$w_{t} + u_{0}w_{r} + \frac{v_{0}}{r}w_{p} + w_{0}w_{t} + uw_{0r} + ww_{0t} = -\frac{1}{8}P_{t} + vv_{w} + agT \qquad (8)$$

$$T_{t} + u_{o}T_{r} + \frac{U_{3}}{r}T_{q} + w_{o}T_{t} + uT_{or} + wT_{oz} = K \overline{V}^{2}T \qquad (9)$$

$$u_{\mu} + \frac{\mu}{r} + \frac{1}{r} \frac{\nu_{q}}{r} + \frac{\omega_{1}}{r} = 0$$
 (10)

In the above equations the stationary basic state variables are functions of r and z alone are assumed to be known everywhere.

EIGENVALUE PROBLEM

We continue with the stability analysis by looking for solutions to Eqs. (6)-(10) for the perturbation functions with the known stationary basic state for some known initial small values for the perturbations. The solution will proceed by first formulating the eigenvalue problem. Eqs. (6)-(10) can be simplified by noting that they are linear and the coefficients are independent of the azimuthal direction ϕ . Under these conditions the perturbation functions admit the following form of solution:

$$[u, \sigma, \omega, \tau, \rho] = [\hat{u}(r, t), \hat{\upsilon}(r, t), \hat{\omega}(r, t), \theta(r, t), \pi(r, t)] \times$$

$$e^{\gamma} p [ik(\varphi - ct)]$$

$$(11)$$

The solution form (11) represents a wave of wavelength $2\pi/k$ which is propagating in the azimuthal direction ϕ with a speed c. Upon substituting the functional form (11) into the perturbation equations (6)-(10), the following equations result

$$-ikc\dot{u} + u_{0}\dot{u}_{r} + \frac{v_{0}}{2}; k\dot{u} + w_{0}\dot{u}_{1} + \dot{u}u_{0r} + \dot{w}u_{0} - \frac{zv_{0}\dot{w}}{2} - \frac{zv_{0}\dot{w}}{2} - \frac{zv_{0}\dot{w}}{2} - \frac{zv_{0}\dot{w}}{2} - \frac{zv_{0}\dot{w}}{2} - \frac{zv_{0}\dot{w}}{2} - \frac{v_{0}\dot{w}}{2} - \frac{v_{$$

$$= -\frac{i}{2}kc\hat{\sigma} + u_{\sigma}\hat{\sigma}_{r} + \frac{i}{2}(v_{\sigma}\hat{\sigma} + w_{\sigma}\hat{\sigma}_{r} + \hat{w}\hat{\sigma}_{r} + \frac{i}{2}(u_{\sigma}\hat{\sigma} + \hat{u}\hat{\sigma}_{r}) + 2S\hat{u}$$
(13)
$$= -\frac{i}{2}k\hat{\rho} + i(\nabla^{2}\hat{\sigma} + \hat{v}_{r}\hat{u}\hat{\sigma}_{r})$$

$$-ikc\hat{w} + u_{0}\hat{w}_{1} + ikv_{0}\hat{w}_{2} + w_{0}\hat{w}_{2} + \hat{u}w_{0} + \hat{w}w_{0} = -\frac{1}{5}\hat{p}_{2} \qquad (14)$$

$$+\sqrt{p}\hat{w} + \alpha q \Theta$$

$$-ikc\theta + u_{\theta}\theta_{r} + ikto \theta + w_{\theta}\theta_{z} + \hat{u}T_{\theta_{r}} + \hat{w}T_{\theta_{z}} = K D^{*}\theta \qquad (15)$$

$$\hat{u}_{r} + \hat{\underline{u}}_{r} + \hat{\underline{v}}_{t} + \hat{w}_{t} = 0$$
(16)

In the present work the perturbation functions will be zero at the boundaries.

Now Eqs. (12)-(16) may be written in the following operator form:

$$\mathcal{A}_{\cdot} \begin{bmatrix} \hat{u} \\ \hat{v} \\ \hat{\mu} \\ \hat{\rho} \end{bmatrix} = \mathcal{L} \begin{bmatrix} \hat{u} \\ \hat{v} \\ \hat{w} \\ \hat{\rho} \\ \hat{\rho} \end{bmatrix}$$
(17)

where k, and k are linear partial differential operators which involve the basic state u, v, w, T and kSince Eq. (17) is homogeneous with Romogeneous boundary condition then it forms an eigenvalue problem for the eigenvalue c.

Normally the system (17) is solved through the method of separation of variables which will reduce the eigenvalue problem to an ordinary differential eigenvalue problem. Such a problem may be solved either by the shooting method or the matrix method (see Antar [7] for a discussion on the application of these techniques). However, in the present case we can not use the method of separation of variables since the coefficients in the operators are not separable. Thus, the partial differential eigenvalue problem (17) has to be solved as it stands.

A logical way to handle the system (17) is to use the finite difference method especially since the coefficients are given as a set of numbers at specific points in the domain. This is due to the fact that the basic state solution is obtained through a finite difference method. Without going too much into the details of the finite differencing procedure, if the derivatives in the operators of and for are approximated by finite differences, the equation (17) will be transformed to the following alegebraic problem

$$[A]\{\bar{x}\} + c[B]\{\bar{x}\} = 0 \qquad (18)$$

where [A] and [B] are banded matrices and is a vector representing the value of the field variables at the discrete points in the solution domain. Taking the inverse of [B] in (18), the equation may be rewritten as

$$\{ [b] - c [I] \} \{ x \} = 0$$
 (19)

where $[\mathbf{\delta}] = [\mathbf{B}]^{-1}$ [A]. Clearly (19) then is an algebraic eigenvalue problem which may be solved through matrix manipulations.

Although in principal the technique is straight forward, in reality the transformation of Eq. (17) into Eq. (18) is quite involved and will not be discussed here. The major difficulty in formulating Eq. (19) is that the vector \bar{x} is very large. In fact if one chooses to discretize Eq. (17) through 10 points in each of the z and r directions the vector x will have 400 elements. However, we know from experience that 10 points for the discretization is not enough and a reasonable choice will be 50 points. For this latter value the vector x will possess 10⁴ elements. This size of the vector will make the solution of the eigenvalue problem very lengthy. Another approach which might be more suitable for the present problem is discussed next.

INITIAL VALUE PROBLEM

The second alternative for obtaining a solution to the perturbation equations (6)-(10) is to assume a solution which possesses a sinusoidal form in the azimuthal direction, ϕ , i.e.

$$f(r, z, q, t) = \hat{f}(r, z, t; k) \sin kq$$
 (20)

This form of the solution is admissible due to the linearity of the equations and also the independence of the coefficient function of the azimuthal direction ϕ . Upon substituting the functional form (20) into eqs. (6)-(10), the following linear system of equations results:

$$\hat{u}_{1} + u_{0}\hat{u}_{r} + i k \underbrace{\sigma}_{p} \hat{u}_{1} + w_{0} \hat{u}_{2} + \hat{u} u_{0r} + \hat{w} u_{02} - \underbrace{z}_{r} \hat{v}_{0} \hat{v}_{1} - z \mathcal{R} \hat{v}_{1}$$

$$= - \frac{1}{2} \hat{P}_{r} + \mathcal{V} \left(\overline{v}^{2} \hat{u}_{1} - \underbrace{z}_{r} \underline{k} \underbrace{v}_{r} - \underbrace{u}_{r} \right)$$

$$(21)$$

$$\hat{v}_{1}^{2} + u_{0}\hat{v}_{1} + \dot{v}_{0}\hat{v}_{2} + \dot{w}\hat{v}_{0} + \dot{w}\hat{v}_{0} + \frac{1}{4}(u_{0}\hat{v} + \dot{u}v_{0}) + 2\Omega\hat{u}$$

$$= -\frac{ik}{r}\hat{b} + v\left(\nabla\hat{v} + \frac{i}{2}\hat{u}\hat{u} - \frac{\partial}{2}\right)^{(22)}$$

$$\hat{w}_{t} + u_{0}\hat{w}_{1} + \hat{w}_{0}\hat{w}_{2} + \hat{w}_{0}\hat{w}_{1} + \hat{w}_{0}\hat{w}_{3} = -\frac{1}{2}\hat{b}_{1} + \frac{1}{2}\hat{w}_{1}\hat{w}_{3}\hat{v}_{3} = -\frac{1}{2}\hat{b}_{1} + \frac{1}{2}\hat{w}_{1}\hat{v}_{3}\hat$$

$$O_{t} + u_{0} O_{r} + i k \sigma_{0} O_{t} + w_{0} O_{z} + \hat{u} T_{o_{t}} + \hat{w} T_{o_{t}} = \kappa \sigma^{2} O \qquad (24)$$

The set (21)-(24) constitute a linear partial differential set depicting an initial value problem. Thus a solution to the system may be obtained once the proper initial conditions are given. The initial conditions in this case are given again as a wave function with a specified wavenumber k and a specified amplitude function.

The stability or instability of the perturbations is resolved by examining the behavior of the amplitude functions with time. If for a given wavenumber k, the amplitudes appear to grow with time, then the perturbations are said to be unstable. If, on the other hand, the perturbation amplitudes seem to decay with time then the pertubations are said to be stable. As for the eigenvalue problem, however, the sign of the imaginary part of c determines stability or instability. In this case if the sign of c_i is negative, then the perturbations are said to be stable, otherwise it is unstable.

Now, due to the complexity of the coefficients of the system (21)-(24), no closed form solution is possible. Thus one has to resort to numerical solutions. In this case it seems logical to use finite difference techniques for both the time and the two spatial directions. Conceptually, a numerical solution for this system should be obtained without difficulty due to the linearity of the system.

REFERENCES

- Fowlis, W. W. and G. H. Fichtl, "Geophysical Fluid Flow Model Experiments in Spherical Geometry." Proc. Third NASA Weather and Climate Program Science Review, NASA Conf. Publ. 2029, Pap. No. 32, 1977.
- Fultz, D., R. R. Long, G. V. Owens, W. Bohan, R. Kaylor and J. Weil, "Studies of Thermal Convection in a Rotating Cylinder with Some Implications for Large-Scale Atmospheric Modeling," Meteor. Monogr., No. 21, 1959.
- 3. Fowlis, W. W. and R. Hide, "Thermal Convection in a Rotating Annulus of Liquid: Effects of Viscosity on the Transition Between Axisymmetric and Non-Axisymmetric Flow Regimes," J. Atmos. Sci., Vol. 22. 1965.
- 4. Hide, R. and P. J. Mason, "Sloping Convection in a Rotating Fluid," Advances in Physics, Vol. 24, 1975.
- 5. Williams, G., "Baroclinic Annulus Waves," J. Fluid Mech. Vol. 49, 1971.
- Quon, C., "A Mixed Spectral and Finite Difference Model to Study Baroclinic Annulus Waves," J. Comp. Phys. Vol 20, 1976.
- Antar, B. N., "On the Solution of Two-Point Linear Differential Eigenvalue Problems," J. Comp. Phys. Vol. 20, 1976.

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ON ROSEN'S THEORY OF GRAVITY AND COSMOLOGY

Prepared By:

Academic Rank:

Dr. Ronnie C. Barnes

Associate Professor

University and Department:

NASA/MSFC:

(Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Lambuth College

Physical Science Department

Space Sciences Cyrogenics Gravitation

Dr. Peter Eby

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ON ROSEN'S THEORY OF GRAVITY AND COSMOLOGY By

Dr. Ronnie C. Barnes Associate Professor of Astronomy Lambuth College Jackson, TN 38301

ABSTRACT

Formal similarities between general relativity and Rosen's bimetric theory of gravity have been used to analyze various bimetric cosmologies. The following results have been found:

- (1) Physically plausible model universes which have a flat static background metric, a Robertson-Walker fundamental metric and which allow co-moving coordinates do not exist in bimetric cosmology.
- (2) It is difficult to use the Robertson-Walker metric for both the background metric and $g_{\mu\nu}$ and require that $g_{\mu\nu}$ and $\gamma_{\mu\nu}$ have different time dependences because there are not enough bimetric field equations to simultaneously fix the time dependences.
- (3) A consistency relation for using co-moving coordinates in bimetric cosmology was derived.
- (4) Certain spatially flat bimetric cosmologies of Babala (1975) were tested for the presence of particle horizons.
- (5) An analytic solution for Rosen's (1978) $k = +1 \mod 1$ was found.
- (6) Rosen's singularity free k = +1 model arises from what appears to be an arbitary choice for the time dependent part of $\gamma_{\mu\nu}$.

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I. INTRODUCTION

In 1973 and 1974 Nathan Rosen proposed a bimetric theory of gravitation. The origin of its name arises from Rosen's assumption that two tensor fields $(g_{\mu\nu} \text{ and } \gamma_{\mu\nu})$ exist at each point (x^1, x^2, x^3, x^4) of spacetime. Rosen (1973,1974,1978) has given the following interpretations for $g_{\mu\nu}$ and $\gamma_{\mu\nu}$:

As in general relativity, $g_{\mu\nu}$ is the fundamental metric tensor of Riemannian geometry. The motion of material bodies and photons is determined by $g_{\mu\nu}$. Moreover, those aspects of spacetime which can be observed by using clocks and measuring rods are contained in $g_{\mu\nu}$.

On the other hand, $\gamma_{\mu\nu}$ may be thought of as representing the spacetime metric that one would obtain if all the mass-energy in the universe was removed. $\gamma_{\mu\nu}$ is called the "background" metric. According to Rosen (1978), " $\gamma_{\mu\nu}$ describes a geometry of spacetime not directly observable at present."

Bimetric gravitation has a number of interesting features:

- 1. It is a metric theory which, if suitable boundary conditions are chosen, agrees exactly with general relativity in the post-Newtonian limit (Lee, et. al. 1976).
- 2. It does not appear to predict the existence of black holes (Rosen 1973, 1974).
- 3. Considerably larger (5-6X) masses for neutron stars are allowed by bimetric gravitation than by general relativity (Rosen and Rosen 1975).
- 4. If the background metric is assumed to be Minkowskian $(\gamma_{\mu\nu} = \gamma_{\mu\nu})$, then it is possible to express gravitational energy and momentum density via a conserved tensor and the field equations and resulting calculations are simpler than in general relativity (Rosen 1973, 1974).
- 5. Homogeneous bimetric cosmological models which have acceptable values for the Hubble parameter H_o and the deceleration parameter q_o have been found and studied by Babala (k=0, 1975), Goldman and Rosen (k=-1, 1977), and Rosen (k=+1, 1978).
- 6. At the present time the only observation known to pose difficulties for Rosen's theory is the period change |P/P|= 1.2 x 10⁻⁷ yr⁻¹ determined by Taylor et.al. (1976) for the binary pulsar PSR 1913 + 16. Will and Eardley (1977) analyzed binary star motion using

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bimetric theory and, assuming that gravitational radiation from a source is described by a retarded potential solution of the wave equation, have shown that bimetric gravitation allows dipole radiation. (However, also see Rosen 1978b for a different view.) Will and Eardley concluded that Rosen's theory is consistent with observations only if PSR 1913 + 16 contains either (a) two very heavy neutron stars of nearly equal mass or (b) a very light neutron star whose companion is a rapidly rotating white dwarf or helium main sequence star.

II. MOTIVATION--COSMOLOGICAL HORIZONS

One of the most intriguing predictions of general relativity is its prediction of "event horizons" in spacetime. These are boundaries which separate observeable events from those not observeable. The best known example of an event horizon is that which surrounds a spherical black hole and which happens to arise at the same value of the radial coordinate, $r = 2GM/c^2$, for which the Schwarzchild metric

$$ds^{2} = (1 - \frac{26M}{Rc^{2}})dt^{2} - \frac{dr^{2}}{1 - 2GM/Rc^{2}} - r^{2}d\theta^{2} - r^{2}sin^{2}\theta d\phi^{2} \qquad (1)$$

has a singularity.

However, the existence of an horizon in spacetime is independent of the choice of a coordinate system and, in fact, it is possible to have event horizons even when spacetime is perfectly homogeneous. Rindler (1956), in his classic paper on cosmological horizons, showed that homogeneous cosmology permits universes with particle horizons, event horizons or both.

The presence of particle horizons (a surface in 3-space which divides all fundamental particles into two classes: those which have been observed by a given "fundamental observer" and those which have not been observed) has troubled cosmologists for the following reason. Particle horizons represent regions which have been and still are causally disconnected from other parts of the universe. However, according to the precepts of canonical cosmology, all fundamental observers must find identical physical conditions (of course disregard small scale local inhomogenities) at the same instant of cosmic time. The obvious and difficult question is: How do regions which are causally disconnected achieve such a remarkable degree of physical identity?

Of course, one can always postulate special initial conditions, but this approach seems more in the spirit of theology than of science. A more challenging approach is to assume that the universe began with anisotropies present and that through the actions of certain mechanisms homogeneity naturally arose (Misner, Thorne and Wheeler 1973, ch. 30). For this approach to be successful it is essential that the following two requirements be met:

- (1). A satisfactory mechanism for converting initial anisotropies into isotropy be found.
- (2). Whatever the mechanism, there must be sufficient time to create the degree of homogeneity and isotropy now observed.

It is possible to use the 3° K microwave background radiation to place some constraints on the time available for producing homogeneity (Misner, Thorne and Wheeler 1973, ch. 30). These constraints are so severe--the universe must be already homogeneous as early as the epoch z = 1000 (only 600,000 yr after creation in an Einstein-de Sitter model) that no entirely satisfactory chaotic cosmology has yet been found.

Therefore, because particle horizons pose an important problem for cosmology, it is important to analyze bimetric cosmology in order to discover if these horizons exist in bimetric models as they do in general relativity.

III. BIMETRIC COSMOLOGY WITH THE ROBERTSON-WALKER METRIC.

Nowadays it is popular to use Robertson-Walker spacetime $ds^{2} = dt^{2} - \frac{R^{2}(t)}{(1+kr^{2}t)^{2}} \cdot \left(\frac{\partial r^{2} + r^{2} d\theta^{2} + r^{2} \sin^{2} \theta}{\partial t^{2} + r^{2} \sin^{2} \theta} d\phi^{2} \right)$ (2)

for cosmological investigations. Here (r, θ, ϕ) form a set of "comoving" spherical coordinates, t is cosmic time, R(t) is the scale factor function, and k is the constant of space curvature.

Rindler (1956) derived equations which allow one to test for horizons in Robertson-Walker spacetime. In models which have R = 0 a particle horizon will exist at time t_0 if $\int_{0}^{t_0} \frac{\partial t}{\partial t}$ converges. (3)

Some cosmologies never pass through R = 0 and begin their motion as $t \rightarrow -\infty$. In these models a particle horizon will exist at t_0 if

 $\int_{-\infty}^{+\infty} dt'/R(t) \qquad \text{converges.} \qquad (4)$

(5)

For an event horizon there is only one test, viz.,

Although the above tests are, at least in principle, straightforward there do arise problems when one attempts to use them in bimetric cosmology. The principle difficulty arises from the

converges.

from the fact that none of the published bimetric cosmological models are consistent with the Robertson-Walker line element for $g_{\mu\nu}$ and, at the same time, consistent with the form used for $\gamma_{\mu\nu}$ in solving the bimetric field equations. For example, Babala (1975) initially assumes

$$ds_{I}^{2} = g_{\mu\nu} dx^{\mu} dx^{\nu} = e^{2\Phi} dt^{2} - e^{2\Psi} (dr^{2} + r^{2} d\theta^{2} + r^{2} sin^{2} \theta d\phi^{2})$$
(6)

$$ds^{2}_{II} = \chi_{\mu\nu} dx^{\mu} dx^{\nu} = dt^{2} - dr^{2} - r^{2} d\theta^{2} - r^{2} sin^{2} \theta d\phi^{2}$$
(7)

where Φ and Ψ are functions of t. By redefining the time coordinate it is possible to transform eq. (6) into Robertson-Walker form

$$\tau = \int e^{\frac{\pi}{2}} dt \qquad ; \qquad ds_{\pi}^2 = d\tau^2 - e^{2\Psi(\tau)} (dr^2 + r^2 d\theta^2 + r^2 sin^2 \theta d\phi^2). \quad (8);(9)$$

From equations given by Babala it is easy to show that his models require

$$\tau = e^{2\circ} \tanh t + k \tag{10}$$

where \mathbf{F}_{\bullet} and k are obtained from initial conditions. The simplest case to analyze is $k = \mathbf{F}_{\bullet} = 0$ and then one finds

(12)

 $\tau = \tanh t \qquad -|<\tau < | \qquad (11)$

$$t = \frac{1}{2} \ln \frac{1+\tau}{1-\tau}$$

$$e^{\psi(\tau)} = e^{\psi_0} \frac{1}{1-\tau^2} \left(\frac{l+\tau}{1-\tau}\right)^{(13)}$$

where λ' is a constant of integration related to the value of Ψ' . According to Babala, expanding models are given by $\lambda' > 1, \lambda' = 1$ $-1 < \lambda' < 1$. The simplest case to study is the one with $\lambda' = 1$ because eq. (13) simplifies to form easy to integrate.

If the Rindler tests are applied, then one finds that this model contains both a particle horizon and an event horizon at τ_o because both

$$\int_{-\infty}^{+0} \frac{dt}{R(t)} \rightarrow \int_{-1}^{\tau_0} e^{-\psi_0} (1-\tau)^2 d\tau = \frac{e^{-\psi_0}}{3} (8 - (1-\tau_0)^3)$$
(14)

$$\int_{t_0}^{\infty} \frac{dt}{R(t)} = \int_{T_0}^{t} e^{-\frac{t}{6}} (1-\tau) d\tau = \frac{e^{-\frac{1}{6}}}{3} (1-\tau_0)^3 \qquad (15)$$
converge.

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However, it is possible to find bimetric models without particle horizons. An example is afforded by choosing $\lambda = 2$. In this model the particle horizon integral logarithmically diverges as $\tau \rightarrow -1$. $\int_{-1}^{\infty} e^{-\frac{1}{2}} \left(\frac{1-\tau}{1+\tau}\right)^{3} d\tau = \left(8 \ln(1+\tau) - 8(1+\tau) + 2(1+\tau)^{2} - \frac{1}{3}(1+\tau)^{3}\right) e^{-\frac{1}{2}} \left(\frac{1+\tau}{1+\tau}\right)^{3} e^{-\frac{1}{2}}$ (1b)

The difficulty which results when the time coordinate is redefined lies in the fact that while eq. (8) succeeds in converting eq. (6) to Robertson-Walker form it also forces the background metric to be time dependent. However, in deriving eqs. (10)-(13), gotten by substituting eqs. (6) and (7) into the bimetric field euations, Babala assumed that $\gamma_{\mu\nu}$ was independent of time. Bimetric cosmologies published by Goldman and Rosen (1977) and Rosen (1978) are subject to the same criticism. However, in the latter paper Rosen assumes that $\gamma_{\mu\nu}$ is time dependent at the outset and that $q_{\mu\nu}$ is given by eq. (6). In this case the effect of making a new time coordinate via eq. (8) is to change the time dependence of $\gamma_{\mu\nu}$ away from the form used in setting up the field equations.

The preceding considerations suggest that the Robertson-Walker line element be introduced at the start rather than making an attempt to transform to it at a later stage. Unfortunately, this procedure does not work if one uses a flat, static background metric. On substituting eqs. (2) and (7) into the bimetric field equations (see Rosen 1973,1974 for more detail):

$$N_{\mu\nu} = \frac{1}{2} \chi^{\alpha\beta} g_{\mu\nu} |_{\alpha\beta} - \frac{1}{2} \chi^{\alpha\beta} g^{\alpha\beta} g_{\mu\lambda} |_{\alpha} g_{\nu\beta} |_{\beta} = -8\pi \pi \left(T_{\mu\nu} - \frac{1}{2} g_{\mu\nu} T \right)$$
(17)

and evaluating the (4,4) component one finds

(18)

for a matter dominated universe with neglible pressure. Apparently, only empty Robertson-Walker bimetric cosmologies are allowed by using eqs. (2) and (7).

The simplest way to produce non-empty bimetric Robertson-Walker models is to let $\gamma_{\mu\nu}$ be time dependent. A straightforward method would be to let $\gamma_{\mu\nu}$ have a Robertson-Walker form but with a different time dependence than $g_{\mu\nu}$, viz.,

$$ds_{I}^{2} = dt^{2} - \frac{R^{2}(t)}{(1+kr^{2}/4)^{2}} \left[\partial r^{2} + r^{2} \partial \theta^{2} + r^{2} \sin^{2} \theta d\phi^{2} \right]$$
(19)

$$ds_{II}^{2} = dt^{2} - \frac{S^{2}(t)}{(1+kr^{2}/4)} \left[dr^{2} + r^{2} d\theta^{2} + r^{2} \sin^{2}\theta d\theta^{2} \right].$$
(20)

However, this approach is not completely satisfactory because

there are not enough field equations to simultaneously determine both R(t) and S(t).

Rosen (1978) made an interesting attempt to fix the time dependence of $\gamma_{\mu\nu}$ by using the cosmological principle. It is well known, e.g., see Robertson and Noonan 1968, ch. 14, 347, that theonly non-static metric which satisfies the perfect cosmological principle is the spatially flat de Sitter line element

$$ds^{2} = dt^{2} - e^{\alpha t} (dr^{2} + r^{2} d\theta^{2} + r^{2} sin^{2} \theta d\theta^{2})$$
(21)

When $\gamma_{\mu\nu}$ is identified with de Sitter's metric it is possible to solve Rosen's field equations for R(t). Rosen, however, rejected the resulting model because it has a big-banglike singularity at $t \rightarrow \infty$. Assuming that the time dependence of $\gamma_{\mu\nu}$ is cosh t/a rather than exp 2t/a and that both $g_{\mu\nu}$ and $\gamma_{\mu\nu}$ are spacetimes of constant positive curvature, Rosen was able to derive an expanding, closed and singularity free universe.

However, it should be pointed out that these goals were reached by making what seems to be an arbitary choice for S(t). Also Rosen fails to point out that if $S(t) = \cosh t/a$, then $\gamma_{\mu\nu}$ no longer satisfies the perfect cosmological principle because the Hubble parameter of this metric

is not independent of time.

During the course of this fellowship the author attempted to develop a satisfactory mehod for fixing the time dependent part of $\gamma_{\mu\nu}$. One approach, which looked promising, was suggested by the fact that in general relativity one can start with a rather general spacetime

$$ds^{2} = e^{\nu(r_{1}t)}dt^{2} - e^{\mu(r_{1}t)}(dr^{2} + r^{2}d\theta^{2} + r^{2}\sin^{2}\theta d\theta^{2})$$
(23)

and derive the Robertson-Walker line element by making use of a "consistency relation" which stems from the use of co-moving coordinates (McVittie, 1965, ch.8).

Since no published work about bimetric consistency relations now exists, the author hoped to (1) derive a consistency relation for the use of co-moving coordinates and (2) use it, if possible, to pin down the time dependence of $\gamma_{\mu\nu}$.

(22)

The first step was accomplished for the following case. If one assumes that $\gamma_{\mu\nu}$ and $g_{\mu\nu}$ generate spherically symmetric spacetimes

$$ds_{I}^{2} = g_{\mu\nu} dx^{\mu} dx^{\nu} = e^{\nu(r,t)} dt^{2} = e^{\mu(r,t)} (dr^{2} + r^{2} d\theta^{2} + r^{2} sin^{2} \theta d\theta^{2})$$
 (24)

$$ds_{II}^{2} = \chi_{\mu\nu} dx^{\mu} dx^{\nu} = dt^{2} - e^{\lambda(r_{1}+1)} (dr^{2} + r^{2} d\theta^{2} + r^{2} sim^{2} \theta d\phi^{2})$$
 (25)

and then substitutes these into the bimetric field equations, then one finds the following non-vanishing components

$$N'' = \frac{2}{7} \left[H^{\mu} - y^{\mu} + (h^{\mu} - y^{\mu}) - \frac{8}{7} y_{f}^{f} e_{y+\eta} + \frac{7}{7} \left[y^{ff} - h^{ff} + \frac{3}{3} (y^{f} - h^{f}) \right] e_{h} \right]$$
(77)

$$N_{22} = r^2 N_{11}$$
 (27)
(38)

$$N_{33} = r^{4} N_{41} \sin^{2} \Theta$$

$$N_{41} = -\frac{1}{4} \left[v_{rr} + \left(\frac{2}{r} + \frac{1}{2}\lambda_{r}\right)v_{r} \right] e^{v_{r}\lambda} + \frac{1}{2} \left[v_{41} + \frac{3}{2}\lambda_{4}v_{4} \right] e^{v} + \frac{3}{8}\lambda_{4}^{2} \left[e^{H-\lambda} - e^{\lambda+3v-\mu} \right]$$
(29)

$$N_{H} = \frac{1}{4} \left[\lambda_{tr} + \lambda_{t} \left(\nu_{r} + \lambda_{r} - \mu_{r} \right) \right] e^{\nu} - \frac{1}{4} \left[\lambda_{tr} - \lambda_{t} \left(\nu_{r} + \lambda_{r} - \mu_{r} \right) \right] e^{\mu - \lambda}$$
(30)

The use of co-moving coordinates requires that, in spherically symmetric spacetime,

$$T'' = (p+p) \frac{dx'}{ds} \frac{dx'}{ds} - g''p$$
(31)

vanish since and g = 0. However, these conditions also require T^{14} -g¹⁴T/2 = 0 and that means N₁4 =0. Thus, the bimetric consistency relation is

$$\left[\lambda_{tr} + \lambda_{t} (\nu_{r} + \lambda_{r} - \mu_{r})\right] e^{\nu} = \left[\lambda_{tr} - \lambda_{t} (\nu_{r} + \lambda_{r} - \mu_{r})\right] e^{\mu - \lambda}$$
(3>)

For comparison, the analogous general relativistic equation is

$$\mu_{rt} = \frac{1}{2} \mu_{t} \nu_{r} \tag{33}$$

As outlined in McVittie's textbook, it is possible to integrate eq.(33) and, by requiring that the (1,1) and (2,2) components of the Einstein tensor yield the same cosmological fluid pressure, one arrives at the Robertson-Walker line element.

Unfortunately, when the same procedure is followed in bimetric theory, one discovers that N_{11} and N_{22} already give the samedifferential equation for pressure. Thus, unlike general

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relativity, the consistency relation for co-moving coordinates does not lead to appreciable progress and, in particular, no new information is gained about the time dependence of $\gamma_{\mu\nu}$.

IV. AN ANALYTIC SOLUTION FOR ROSEN'S 1978 k =+1 MODEL

As mentioned in the previous section, Rosen (1978) found a closed space expanding model which does not have a singularity. During the course of this summer's work the author found an analytic solution to the bimetric field equations of this model.

$$\ddot{\mp} + \frac{3}{a} (\tanh \frac{1}{a}) \dot{\mp} + \frac{3}{a^2} \tanh^2 \frac{1}{a} \sinh (2\Psi - 2\overline{\Phi}) = -4\pi e^{\frac{1}{2} + 3\Psi} (p + 3p) \quad (34)$$

$$\ddot{\psi} + \frac{3}{a} (\tanh \frac{1}{2}) \dot{\psi} - \frac{1}{a^2} \tanh^2 \frac{1}{a} \sinh((2\psi - 2\bar{\psi})) = 4\pi e^{\frac{1}{2} + 3\psi}(p-p)$$
 (35)

$$\dot{g} + 3(g+p)(\dot{\psi} + \frac{1}{4} \tanh \frac{1}{4}) = 0$$
 (36)

If one investigates a zero pressure model and sets $g_{44} = 1$, then these equations simplify and it is possible to use the integral of eq.(36)

$$ge^{3^{4}}cosh^{3} \frac{t}{a} = A$$
; A is a constant of integration, (37)

to set up a single second order differential equation which is linear in $\ \Psi$

$$\dot{\psi} + \frac{3}{2}\dot{\psi} \tanh t_{\alpha} = \frac{8\pi A}{3} \operatorname{sech}^{3} \frac{t}{\alpha}$$
 (38)

By using the classical variation of parameters method it is possible to arrive at the following general solution to eq.(38):

$$\Psi(t) = \left(\frac{8\pi A}{3} \pm \frac{1}{2} aB\right) \left[\operatorname{sech} \pm \tanh \pm \frac{1}{2} \tanh \frac{1}{2} + \tan^{-1}(\sinh \frac{1}{4})\right] + \frac{8\pi A}{3} \operatorname{sech} \pm \frac{1}{2} - \frac{8\pi A}{3} \int \tan^{-1}(\sinh \pm \frac{1}{2}) \frac{d\pm}{2} + (..., (39))$$

In this solution B and C are constants of integration and the integral can be carried out by expanding $\tan^{-1}(\sinh t/a)$ in a power series.

Although eq.(39) is not as simple as one would like, it is nevertheless, sufficiently simple to analyze for possible singularities in cosmic time. By letting $t \rightarrow -\infty$, $t \rightarrow 0$ and $t \rightarrow \infty$ and looking at each term in (39) one can verify that Rosen's conclusion about the absence of a singularity in this model is correct.

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V. SUMMARY

Because the formalism of Rosen's bimetric theory is so similiar to general relativity one would think that it would be a simple matter to set up bimetric cosmology by using the Robertson-Walker metric for g $\mu\nu$. However, if such is attempted, then one finds that there are too few bimetric field equations to simultaneously fix the time dependence of g $_{\mu\nu}$ and $\gamma_{\mu\nu}$.

Other results are:

- (1) Certain flat bimetric cosmologies of Babala have particle horizons; others do not. Apparently, the bimetric field equations (by themselves) are not strong enough to prevent the occurence of cosmological particle horizons, although strangely enough, they do exclude black holes.
- (2) An empty universe results if one tries to use a flat background metric along with a Robertson-Walker fundamental metric.
- (3) Rosen's singularity free universe of positive spatial curvature (Rosen 1978) arises from what seems to be a an arbitary choice for the time dependent part of $\gamma_{\mu\nu}$.
- (4) A consistency relation for using co-moving coordinates in bimetric cosmology was derived.
- (5) An analytic solution was found for the field equations which describe Rosen's (1978) universe of positive space curvature.

REFERENCES

Babala, D. 1975, J. Phys. A., 8, 1409.

- Goldman, I. and Rosen, N. 1977, Ap.J., <u>212</u>, 602.
- Lee, D.L., Caves, C.M., Ni, W-T, and Will, C.M. 1976, Ap.J., 206, 555.
- Mc Vittie, G.C. 1962, General Relativity and Cosmology (U. of Ill. Press: Urbana).
- Misner, C.W., Thorne, K.S. and Wheeler, J.A. 1973, Gravitation (Freeman Co.: San Francisco).
- Rindler, W. 1956, MNRAS, <u>116</u>, 662.

Robertson, H.P. and Noonan, T.W. 1968, Relativity and Cosmology, (W.B. Saunders Co.:Philadelphia).

Rosen, N. 1973, J. Gen. Rel. Grav., 4, 435.

_____ 1974, Ann. Phy. (N.Y.), <u>84</u>, 455.

_____ 1978, J. Gen. Rel. Grav., 2, 339.

____ 1978b, Ap.J., <u>221</u>, 284.

Rosen, N. and Rosen, J. 1975, Ap.J. 202, 782.

Taylor, J.H., Hulse, R.A., Fowler, L.A., Gullahorn, G.E. and

- Rankin, J.M. 1976 Ap.J. (Lett.), <u>206</u>, L53.
- Will, C.W. and Eardley, D.M. 1977 Ap. J. (Lett.), L91.

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NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

MEASUREMENTS OF THE SAMPLING VOLUME OF AN INFRARED LASER DOPPLER VELOCIMETER FOR THE STUDY OF FEASIBILITY OF SIMULTANEOUS THREE-DIMENSIONAL VELOCITY PROFILE MAPPING

Prepared By:	Robert L. Bond, Ph.D.
Academic Rank:	Associate Professor
University and Department:	University of Arkansas Department of Electronics and Instrumentation
NASA/MSFC: (Laboratory) (Division) (Branch)	Electronics and Control Optical and R.F. Systems Optics
MSFC Counterpart:	James W. Bilbro
Date:	August 1, 1980
Contractor No.:	NGT-01-002-099 (University of Alabama)

XXIV

MEASUREMENTS OF THE SAMPLING VOLUME OF AN INFRARED LASER DOPPLER VELOCIMETER FOR THE STUDY OF FEASIBILITY OF SIMULTANEOUS THREE-DIMENSIONAL VELOCITY PROFILE MAPPING

Ву

Robert L. Bond Associate Professor of Electronics & Instrumentation University of Arkansas Little Rock, Arkansas

ABSTRACT

This project is a continuation and configuration of the experimental work performed during the summer of 1979. Data obtained after the submission of last year's report gave strong indication that it would be possible by spatial masking to code information onto an outgoing laser beam which would yield three-dimensional velocity information when analyzed at an interferometer.

The work continued into this year has involved detailed, high resolution incoherent mapping of the sampling volume to determine the three-dimensional profile of the volume of the beam from which information will be obtained by scattering.

ACKNOWLEDGEMENTS

I wish to express my gratitude to Mr. James W. Bilbro for his continuous support and many suggestions offered throughout the performance of this work. I also appreciate Mr. Bilbro's assignment of Miranda James to assist me during the data-taking phase of this work. Miranda has substantially facilitated this normally time-consuming and tedious process. I also wish to thank Dr. Bob Barfield for his very competent management of the Summer Faculty Program.

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INTRODUCTION

Marshall Space Flight Center has been involved in the application of Laser Doppler Velocimeters (LDVs) since their inception by Yeh and Cummins¹ in 1964. Subsequent research has applied LDVs to the study of aircraft wake vortices,²,³ atmospheric turbulence measurements,⁴ remote intensity fluctuations,⁵ dust devil measurements,⁶ and presently to severe storms.⁷

Except for the earliest work in wind tunnels, most applications have used CO₂ lasers to increase the signal-to-noise ratio - an advantage resulting from the atmospheric window at 10.6 micrometers. Present work involves pulsed CO₂ laser to further increase the signal-to-noise ratio. Future work will utilize the advantages of off-axis telescopes and, possibly, lasers operating at wavelengths which more nearly match atmospheric windows.

Each LDV has certain limitations on the velocity vectors which it can measure. The early LDV of Yeh and Cummins, for example, could easily measure flow velocities perpendicular to the incident optical plane. However, it could not measure velocity components along the optical axis. Subsequent systems, such as those used to measure aircraft wake vortices, measure the components along the optical direction at the expense of the orthogonal components. A need exists for an LDV which is capable of reading all orthogonal velocity components. Such an LDV could be used to map the three-dimensional characteristics of a remote flow field.

Several developments in LDVs point to the possibility of coding sufficient information into a laser beam so that the returned radiation will contain three-dimensional velocity component data. These developments include the classical heterodyne, fringe, and the more recent time-of-flight LDVs. It is possible that a combination of these approaches would yield the three-dimensional information necessary to map a three-dimensional flow field. The most likely impediments appear to be instrumentation artifacts which may obscure the signal of one or more velocity components or produce an unacceptable signal-to-noise ratio.

OBJECTIVES

The purpose of this work has been to obtain a detailed understanding of the sampling volume of the infrared laser Doppler velocimeter assembled during the 1979 Summer Faculty Fellowship Program. Detailed, high resolution maps were to be made of both the coherent and incoherent sampling volumes. The response of the sampling volume to spatial masking within the telescope was also to be determined. The response of the system to various objects moving through the sampling volume was also to be determined.

EXPERIMENTS

The infrared laser Doppler velocimeter (Figure 1) assembled during the Summer Faculty Fellowship Program last summer was returned to its operational status. The system was improved by modifying the external mirrors to permit mechanical rather than electrical positioning. This step was necessary to maintain more precise alignment of the entire laser Doppler velocimeter (LDV).

The detailed study of the coherent sampling volume yielded essentially the same results as those obtained last summer. The double maximum present last summer in the Doppler Signal Power vs. Secondary Mirror Position curves was not as pronounced. The cause of this phenomenon was isolated to the z-scan mechanism of the secondary mirror. As the secondary mirror is moved along the optical (z) axis, the focal point projected onto the detector assembly (27-50 feet away) moves erratically about in the x-y plane of the detector. Of course, similar wandering of the beam occurs relative to a fixed position on the rotating sandpaper wheel.

Ironically, a similar phenomenon occurs when the rotating sandpaper wheel located on a carriage on the optical bench is moved along the z-axis of the beam. In this case the "wandering of the beam" is only apparent. A fixed point on the rotating wheel is actually moving in the x-y plane as the ztraverse is made. Obviously, data entailing movement of both the wheel and secondary mirror contains a compound error resulting from both phenomena.

Data taken by scanning the secondary mirror, first in one direction, then in the reverse direction, is reasonably repeatable but is not symmetric relative to the maximum power point in the scan. This asymmetry is in part due to the x-y motion introduced into the outgoing beam.

The same data taken by scanning the rotating wheel along the optical axis, first in one direction, then in the reverse direction, is approximately as repeatable but with much less asymmetry.

When the signal power (as measured by a specturm analyzer) was plotted against the voltage (as measured on an oscilloscope) there were discontinuities on the semi-log plot. The causes of these points which did not fall on the straight line plot were never determined. There is a remote possibility that they resulted from an ailing Honeywell specturm analyzer which failed to function after this data was taken.

The incoherent data verified and accentuated the conclusions arrived at from the coherent data - namely that the beam wanders about in the x-y plane as it rotates in an irregular arc about the optical axis. An example of the phenomenon is as follows. With the secondary mirror adjusted to focus at the near end of the optical bench let us assume the focal point is on the optical axis. As the secondary mirror is adjusted to move the focal point toward the far end of the optical bench, the beam first swings toward the left as it swings upward and finally swings back toward the right as it reaches its maximum height at the far end of the optical bench. This general pattern was repeated for many experiments. Occasionally there was some fine detail in the arc but the trend remained the same. It was determined that the relationship between the secondary mirror position and the focal point position on the optical axis was highly repeatable and agreed with theoretical predictions.

The incoherent sampling volume was measured by moving an aperture throughout the double cone defined by the focused beam from the laser. Figure 2 is a traverse in the x-axis of the laser beam without secondary mirror or wire filter. This x-axis profile is the standard for all additional work which was taken with the beam focused between 27 and 50 feet by the secondary mirror. A 500 μ m aperture was selected for the traverse of the unfocused beam to keep the power density within the limits of the pyroelectric detector and lock-in amplifier used for power detection. Apertures of suitable size were always used for taking data since neutral density filters operable at 10.6 µm were not available. Attempts to use existing flats of Ge, CaF2 or BaF2 were futile because of significant artifacts introduced into the data via diffraction and multiple reflections. Notice the asymmetry in the beam which is present to some extent in all data.

Initial data was taken with a 1.0 mm aperture as was done last year. It was determined that this aperture did not produce sufficient resolution for a laser beam that has a theoretical focal diameter of approximately 40 mil. The resolution would be particularily dismal if one were interested in detecting auxillary maxima about the center of the beam. As a result, the 1.0 mm aperture attached to a Coherent Radiation power head was replaced with pinholes of either 25 or 50 μ m attached to a pyroelectric detector. Depending upon the laser output power, the 50 μ m aperture was frequently so large that it produced slight saturation of the pyroelectric detector or lock-in amplifier. Therefore, much of the data was taken with the 25 μ m aperture when the laser power was too large for the other size aperture. The Coherent Radiation power head could not be used because it did not have sufficient sensitivity with small apertures and because its frequency response was below the 4 Hz minimum response of the lock-in amplifier and accompanying rotating aperture. Though data were taken at many z-axis positions, those which follow were taken at a nominal 47 feet.

Figure 3 is a trace along the optical axis obtained by moving the aperture along the optical bench with the secondary mirror set for a nominal distance of 47 feet. Notice the asymmetry relative to the nominal focal plane of the nominal focal plane is caused by variation in time of the x-y position of the focal point. The beam was originally positioned in x, y and z to provide maximum power at the nominal focal plane.

Figure 4 compares the trace along the z-axis taken with and without the wire filter in place. There is a natural reduction in total power when the wire is in place but otherwise the two traces are similar.

Additional information was obtained by x and y traverses of the sampling volume at the nominal focus and at distances before and after the nominal focal plane. Figure 5 compares the x-axis traverse taken at the nominal focal plane with and without the wire filter in position. Figure 6 presents the same information for the y-axis. Other than the normal reduction in power produced when the wire is introduced, the two traces are similar for both the x and y traverses.

Figure 7, in combination with Figure 8, represents several x traverses in the nominal focal plane and in several planes before and after the nominal focal plane. Figure 7 compares the x trace in the nominal focal plane to the x traces taken at nominal half (5 in.) and quarter (7 in.) power points before focus. Figure 8 compares the x trace in the focal plane to the x trace taken at nominal half power (3 in.) and quarter power (4 in.) points after focus.

Figure 9 and 10 present corresponding traces for the y axis.

The introduction of the wire filter just prior to the secondary mirror of the telescope splits the outgoing laser beam into two distinct nearly equal components - one left and one right. Each component is made up of a large lobe of radiation separated into several smaller diffraction lobes whose intensity decreases rapidly as one progresses from the z axis outward - to right for right hand lobe - to left for left hand lobe. Data indicates that the x and y tracings taken at the nominal focal plane for the whole beam is not the simple intensity sum of the two individual half beams. It is speculated that the system is sufficiently abberated that the individual lobes each form a non-detailed blur in the region of the focal plane. When the two lobes are simultaneously present at the focal plane, they coherently interfer to form an interferance pattern with its maximum on axis and usually with side lobes.

Figure 11 presents a trace along the x axis at the nominal focal plane for the total beam and for the left and right hand halves of the beam. Either the right or left half of the beam is removed by placing a mask near the telescope. The edge of the mask coincides with the center of the shadow of the wire filter. The introduction of only one half the beam at a time is guaranteed by viewing the pattern with liquid crystal paper.

Figure 12 presents a trace along the y axis. The x-y position of the focal point has moved in this plot.

Figure 13 and 14 present the same data for a nominal half power point 2.5 inches after the nominal focal plane.

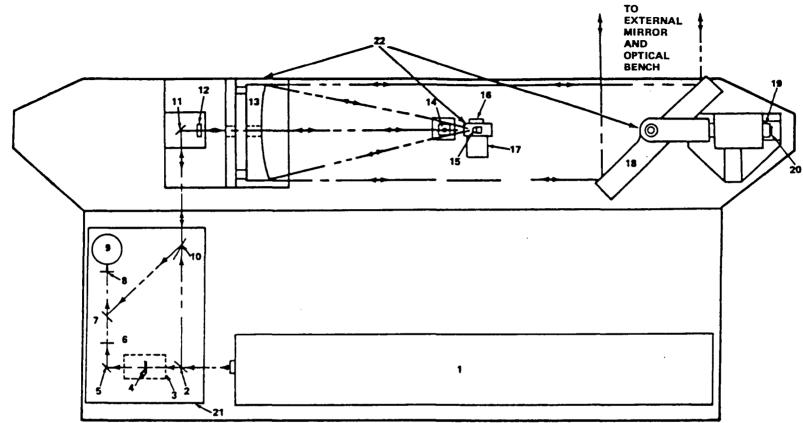
The normal sampling interval for the x, y traverses is 2.5 inches, which provides that there is no overlap of sample areas whether a 25 or 50 μ m aperture is used. For the data with the unfocused beam a 500 μ m (approximately 20 mil) aperture sampled every 25 mils.

CONCLUSIONS AND RECOMMENDATIONS

The incoherent sampling volume of the infrared laser Doppler velocimeter assembled in the 1979 Summer Faculty Fellowship Program was found to be heavily dependent upon the configuration of the outgoing laser beam. In general when only half the beam is allowed to focus at the focal plane, the x, y traverses indicate a broader beam with nominal side lobes. However, if the total beam is permitted to impinge on the focal plane there is substantial narrowing of the beam with increased maximum intensity and an increase in side lobes.

The double maximum indicated by the coherent sampling volume appears to be an artifact of the scanning process - not a fundamental entity of the sampling volume.

It is not known at this point how significant any of these effects are on the actual Doppler signal and its analysis. It is recommended that this information be obtained by first "calibrating" the laser output followed by automated data taking procedures to minimize the very apparent time factor which permeates all the data.



 1. LASER
 12.

 2. BEAM SPLITTER
 13.

 3. BRAGG CELL
 14.

 4. ATTENUATOR
 15.

 5. 45° MIRROR
 16.

 6. HALF WAVE PLATE
 17.

 7. BEAM SPLITTER
 18.

 8. FOCUSSING LENS
 18.

 9. DETECTOR
 20.

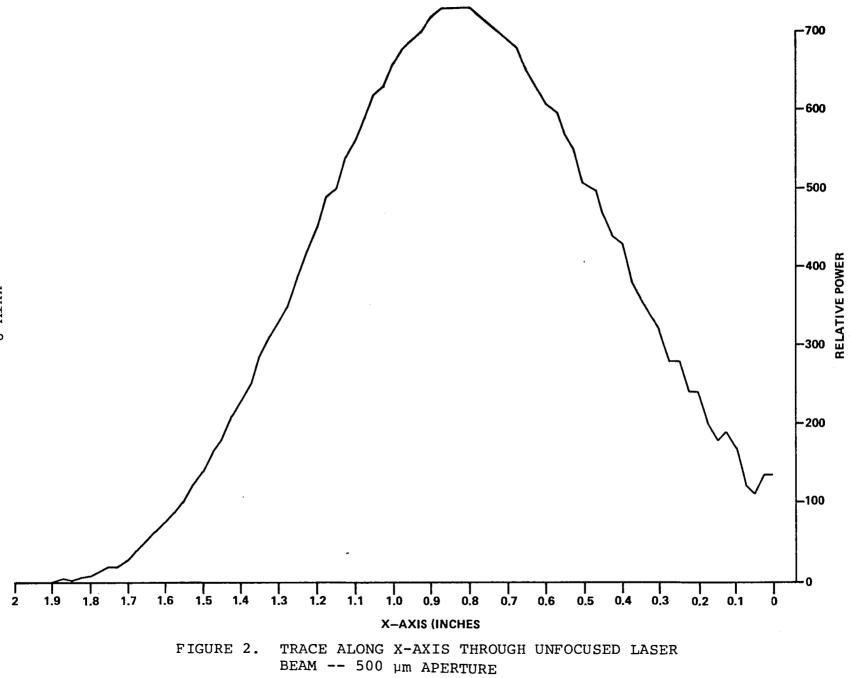
 10. BREWSTER WINDOW
 21.

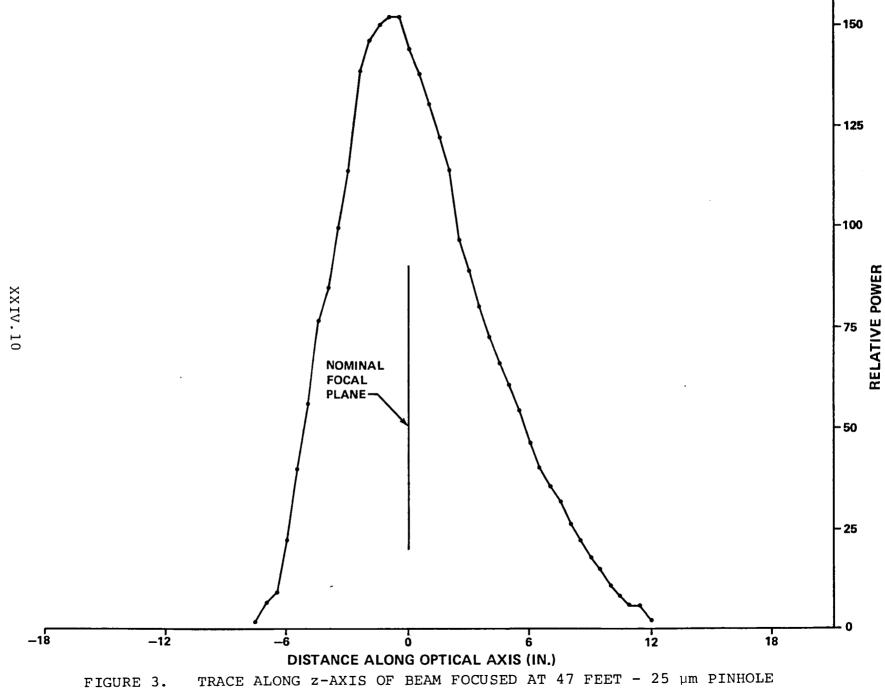
 11. 45° MIRROR
 22.

-

QUARTER WAVE PLATE
 PRIMARY MIRROR
 SECONDARY BACKSCATTER ATTENUATOR
 SECONDARY MIRROR
 SERVO MOTOR
 SCANNING MIRROR
 SERVO MOTOR
 SINE COSINE POT
 INTERFEROMETER
 TELESCOPE







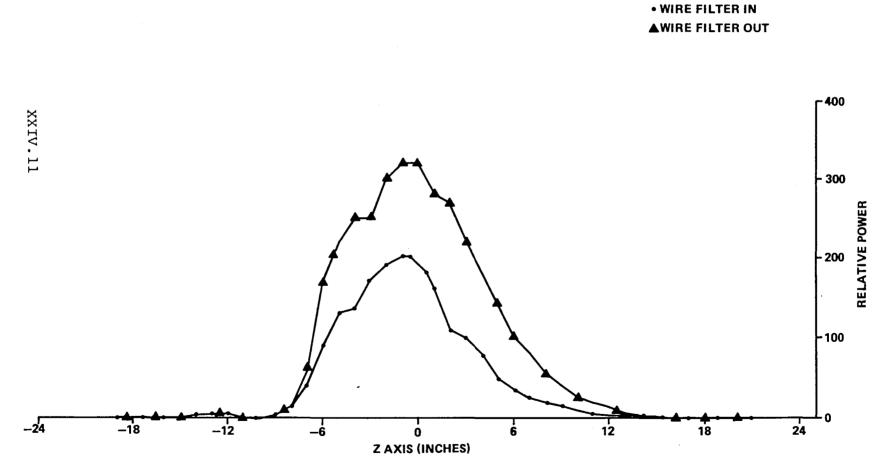
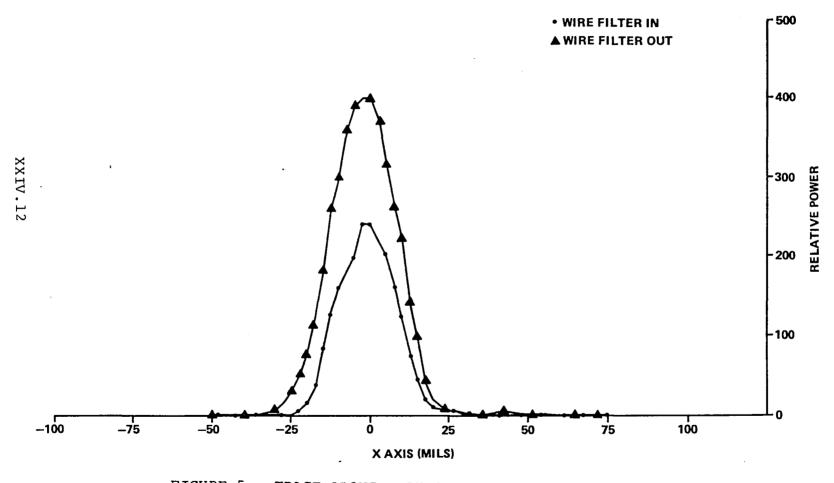
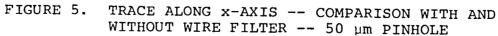
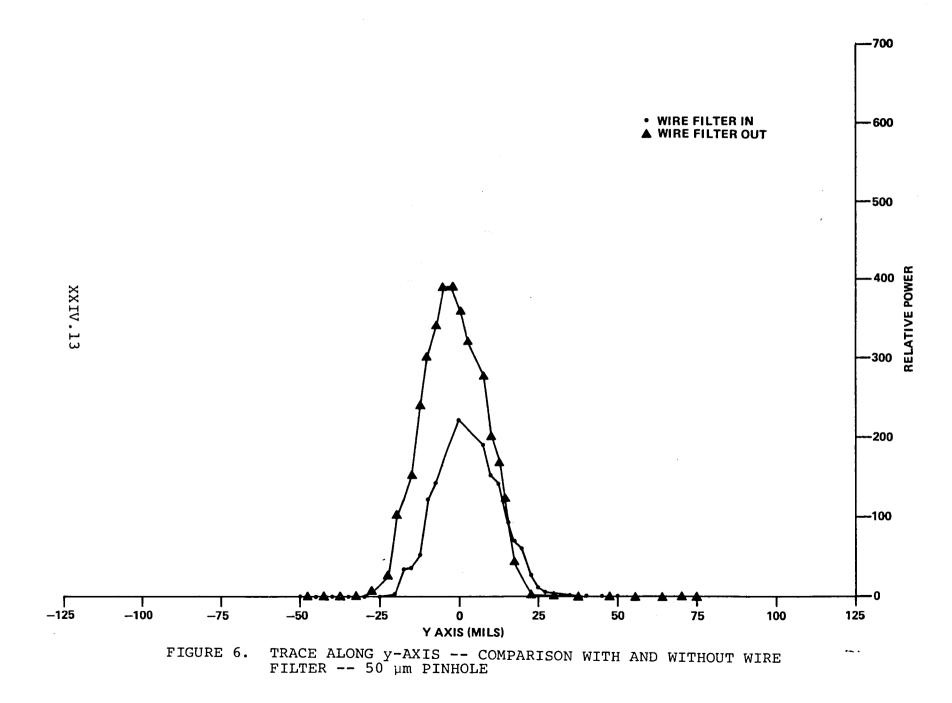


FIGURE 4. TRACE ALONG z-AXIS OF BEAM FOCUSED AT 47 FEET -- COMPARISON WITH AND WITHOUT WIRE FILTER -- 50 μm PINHOLE







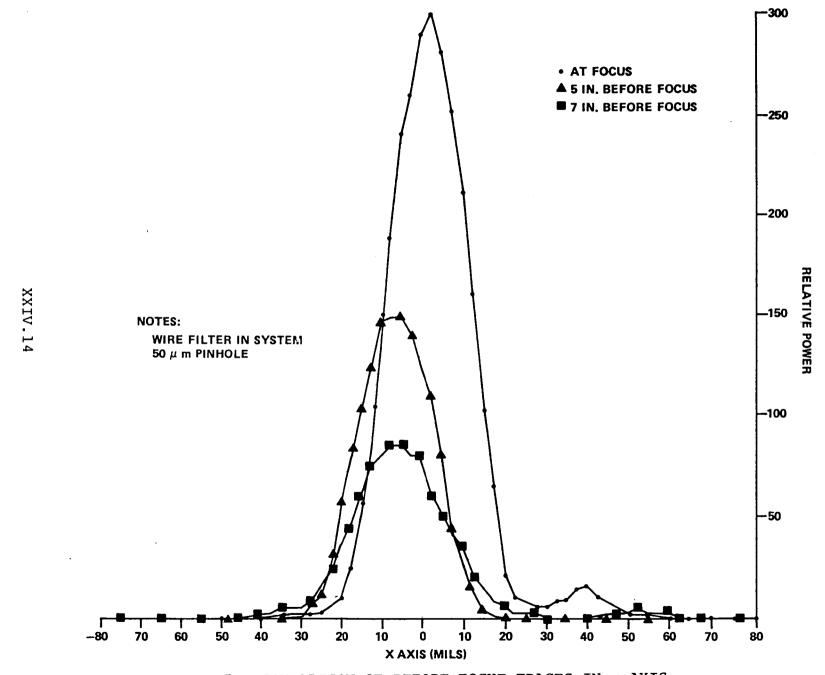
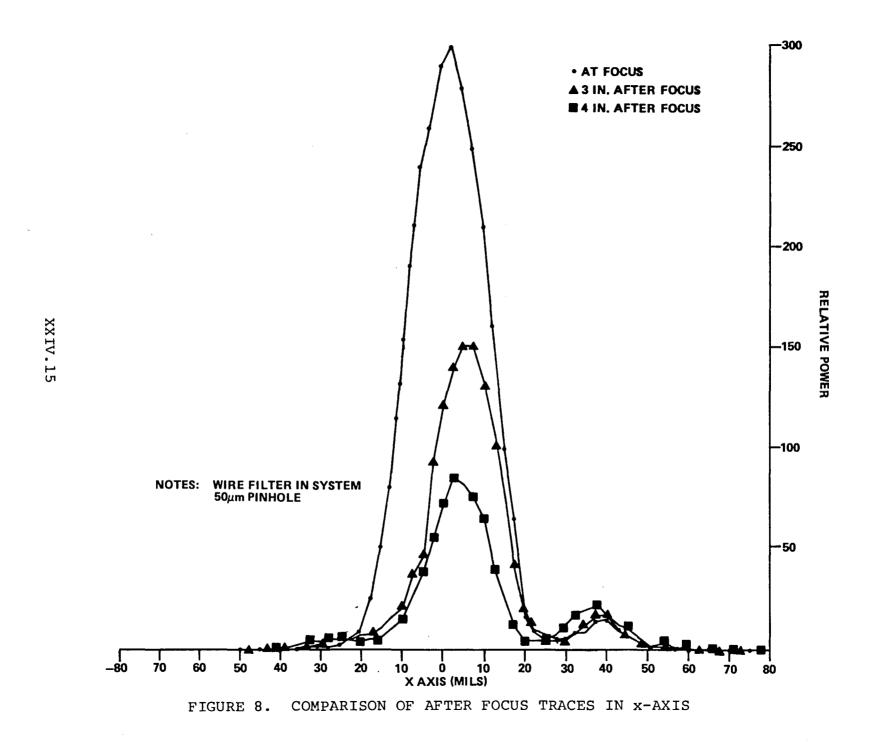


FIGURE 7. COMPARISON OF BEFORE FOCUE TRACES IN X-AXIS



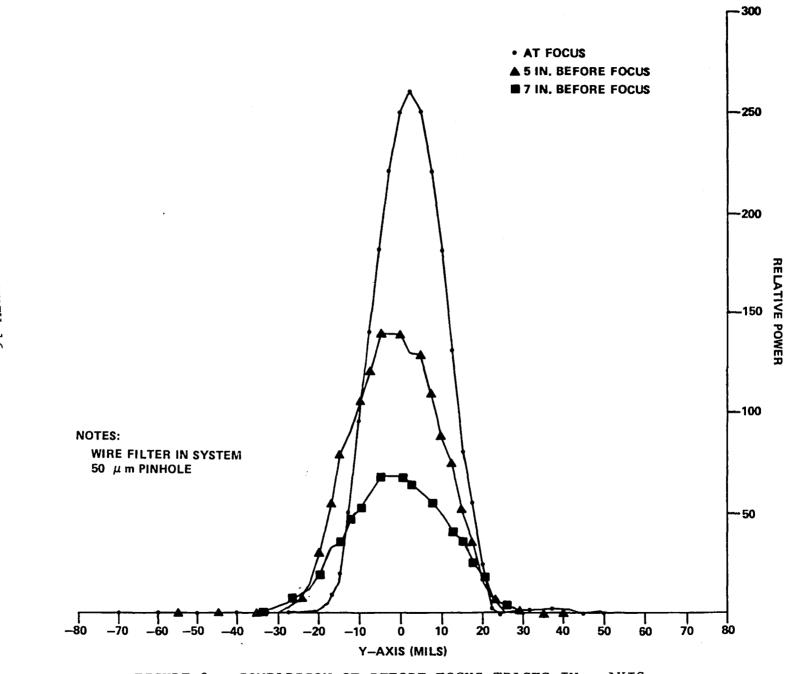


FIGURE 9. COMPARISON OF BEFORE FOCUS TRACES IN X-AXIS

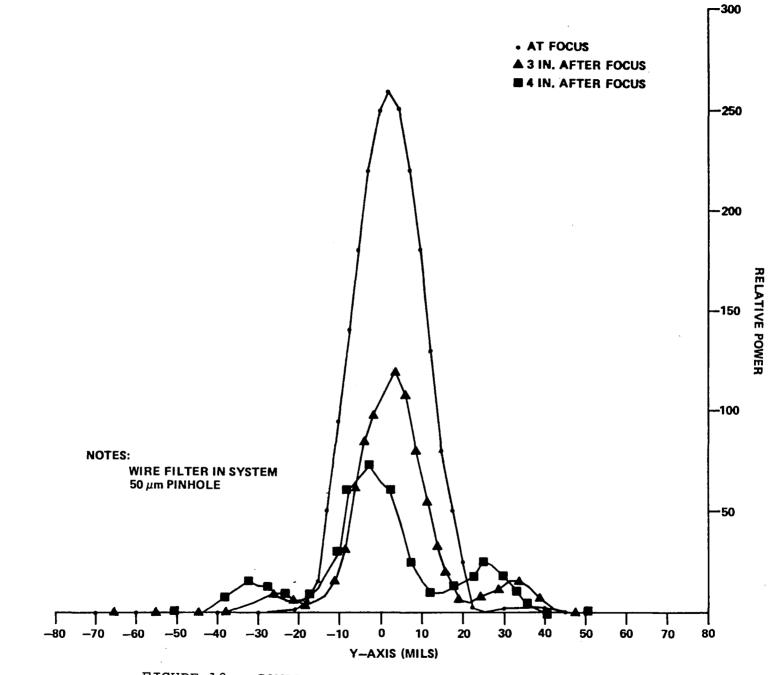
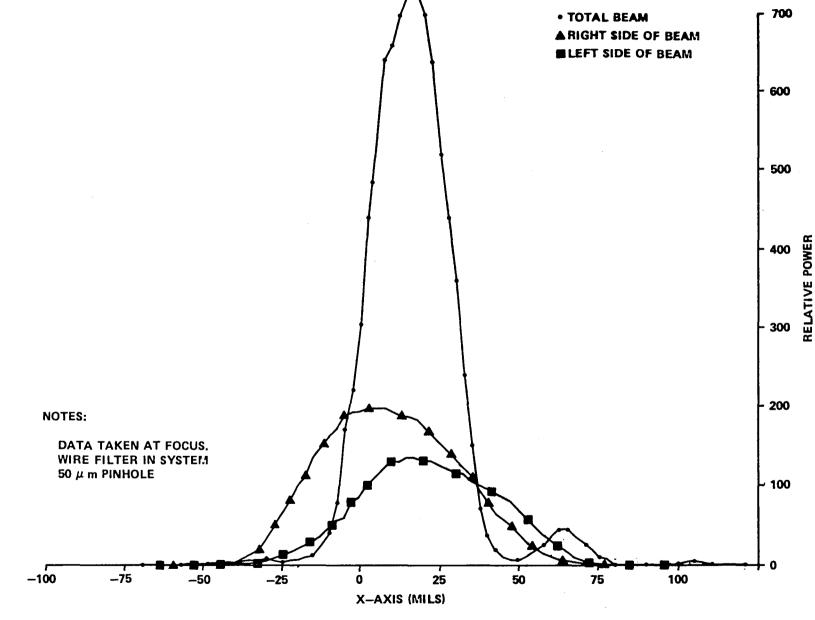
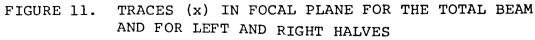


FIGURE 10. COMPARISON OF AFTER FOCUS TRACES IN y-AXIS





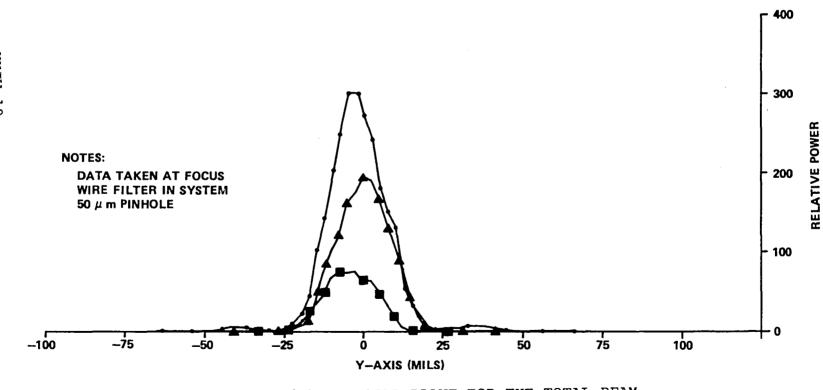
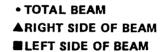


FIGURE 12. TRACES (y) IN FOCAL PLANE FOR THE TOTAL BEAM AND FOR LEFT AND RIGHT HALVES



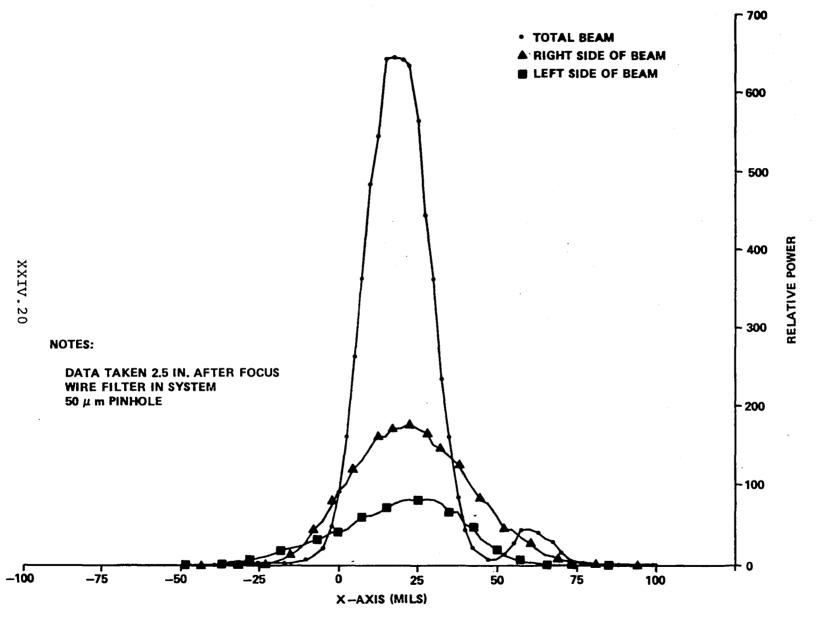


FIGURE 13. TRACES (x) 2.5 INCHES AFTER FOCUS FOR THE TOTAL BEAM AND FOR LEFT AND RIGHT HALVES

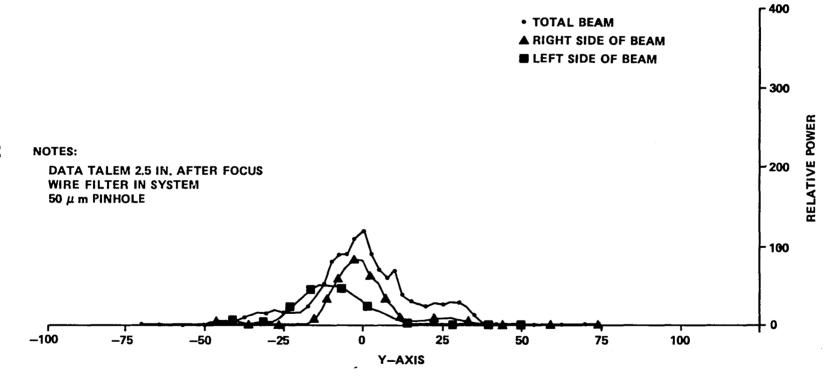


FIGURE 14. TRACES (y) 2.5 INCHES AFTER FOCUS FOR THE TOTAL BEAM AND FOR LEFT AND RIGHT HALVES

REFERENCES

1.	Yeh, Y. and H. Z. Cummins, Appl. Phys. Lett 4, 176 (1964).
2.	Huffaker, Robert M., et al. Development of a Laser Doppler System for the Detection, Tracking, and Measurement of Aircraft Wake Vortices, Final Report. FAA-RD-74-213, March 1975.
3.	Bilbro, J. W., et al. Laser Doppler Velocimeter Wake Vortex Tests, Final Report. NASA TM-X-64988, March 1976.
4.	Polge, R. J. and J. W. Bilbro, Proceedings of Southeastcon '78, IEEE, April 1978.
5.	Kennedy, L. Z., and J. W. Bilbro, Appl. Opt. 15, 2008 (1976).
6.	Bilbro, J. W., et al. Laser Dust Devil Measurements, NASA TN D-8429, February 1977.
7.	Bilbro, James W. and William W. Vaughn, Bull. Am. Meteorological Soc. 59, 1095 (1978).

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MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

A MONTE CARLO APPROACH TO TOLERANCE ANALYSIS OF A NEAR-DIFFRACTION LIMITED OPTICAL IMAGING SYSTEM

Prepared By:

Academic Rank:

W. Harold Davenport

Assistant Professor

University and Department:

Columbus College Department of Mathematics

NASA/MSFC: (Iaboratory) (Division) (Branch)

MSFC Counterpart:

Date:

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Electronics and Control Optical and R.F. Systems Optics

Donald B. Griner

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W. Harold Davenport, Ph.D. Assistant Professor of Mathematics Columbus College Columbus, Georgia

ABSTRACT

Using the theory of Don Knuth, a computer program has been developed to generate random numbers on the Sigma V computer. A subroutine gives a random real number from a uniform distribution on the interval (0,1). A normal distribution with mean zero and variance 1 is also obtained. In my NASA/ASEE report for 1979, a statistical description of 5 possible errors is given. These 5 errors are S₁ which is X-axis tilt; S₂ which is Y-axis tilt; S₃ which is X-decenter; S₄ = Y-decenter; and S₅ which is despace.

Combining the above results a quality function, $Q(S_1, S_2, \ldots, S_5)$, is obtained via a computer program. Values of Q^2 are given in fractional wavelengths. The (variables) errors S_1, \ldots, S_5 are assumed to be independent of each other with possibly different statistical descriptions. Consequently, Q is a linear function.

ACKNOWLEDGEMENTS

The author would like to express his gratitude to Dr. Bob Barfield for directing a very successful summer program. I would also like to express my appreciation to my MSFC counterpart, Don Griner, for his consultations and guidance during the summer; to Charlie Jones for the many technical discussions concerning this project; and to Julie Taylo for her programming assistance.

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I. INTRODUCTION

A Monte Carlo approach is used to determine wavefront error on a two mirror telescope. A computer program that traces N rays through an optical system and calculates the root mean square of the optical path difference (RMSOPD) is used for this analysis. An expanded version of this program will also give the Zernike coefficients. In this study only five errors are used, X-tilt, Y-tilt, X-decenter, Y-decenter, and despace of the secondary mirror. A statistical description is obtained for these five errors. Assuming a uniform probability of any one of these errors, we obtain a wavefront error.

The critical ingredient in a Monte Carlo method is a good random number generator. Knuth has developed a theory for an algorithm that generates a uniformly distributed set of pseudorandom numbers. This algorithm was programmed in FORTRAN IV and used to obtain a uniform distribution of random numbers on (0,1) and a normal distribution of mean zero and variance of 1.

II. OBJECTIVES

To develop a statistical simulation that uses the Monte Carlo techniques in the form of a computer program. The RSS method of tolerance analysis then is to be compared with the Monte Carlo method.

III. TOLERANCE ANALYSIS OF A NEAR-DIFFRACTION LIMITED OPTICAL SYSTEM

A near-diffraction limited optical imaging system is an optical system designed such that root mean square of the optical path difference (RMSOPD) is less than or equal .25 λ . In this analysis the wavelength (λ) used was 632.8 nm. A ray trace computer program written in FORTRAN IV for the Sigma V computer was used. This program can trace N (N = 101 to 1000) rays through an optical system and calculates RMSOPD. The optical path difference (wavefront errors) is obtained by subtracting the chief ray length from the remaining ray lengths. The chief ray is always traced by the program; ignoring the central obscuration it is used as our reference. A spherical wavefront on the image plane (focal plane) is used as the reference. The optical system is a modified Cassegrain or Ritchey-Chretian design. The exit pupil for this design is back of the secondary mirror (see Figure 1). The entrance pupil is the primary mirror.

The ray trace program has a graphics capability of plotting the spot. Figure 2 is a spot diagram of the program when no aberrations are present. Page XXV-10 gives the output that corresponds to this case. Recall that aberration describes the lack of homocentricity in a pencil of rays (input) or the deviation of a wave surface from a spherical form. In Figure 3 we have a spot diagram with 0.1 arc min of tilt about the X-axis. The output that corresponds to this case is on page XXV-9.

The primary lens (mirror) will remain fixed. The secondary has six degrees of freedom but is symmetric with respect to rotation; hence, we only consider:

- (i) X-axis tilt (S₁) -.95 to .95 arc min -1.7 to 1.7 arc min (centroid correction) -2.9 to 2.9 arc min (best chief ray)
- (ii) Y-axis tilt (S₂)
 Same range of numbers as X-axis tilt
- (iii) X-axis decenter (S₃) -165 to 165 μ m -300 to 300 μ m (centroid correction) -530 to 530 μ m (best chief ray)
- (iv) Y-axis decenter (S₄)
 Same range of values as X-axis decenter
- (v) Despace (S₅) -10.9 to 10.9

When the optical system is perturbed by one of the above five perturbations, it is the exit pupil wavefront that is operated upon.

IV. MONTE CARLO METHOD

A random number generator using the theory of Don Knuth (see reference 7) is used to give a uniform distribution on (0,1). Also see reference 2, page 240f, for a FORTRAN listing of this random number computer program. Let $Q = Q(S_1, S_2, \ldots, S_5)$ be the quality function where S_i , $i = 1, \ldots, 5$ are the errors given above. A Monte Carlo approach is used to obtain Figure 2. A random number is selected from the random number generator corresponding to a uniform distribution. This number is then scaled to the maximum range of the possible error for S_i . The five errors S_i are assumed to be independent of each other. Hence, one to five different uniformly distributed random numbers are obtained to give the histogram of Figure 2. The data for each of the histograms was formatted as follows:

Run No.	sl	s ₂	s ₃	s ₄	s ₅	RMSOPD	AVG.	RMSOPD
1								
2								

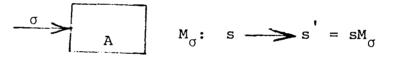
One thousand runs were made and the average RMSOPD for these 1000 runs was plotted on the Y-axis.

One thousand runs were made where X-Y tilt and X-Y decenter were given an equal probability of occurring. The tolerance range used for each of these four errors was the range that would give a .25 λ RMSOPD wavefront error. If we take the root sum square (RSS) of these four values we get 0.5 λ . From Figure 4 we see that the average RMSOPD is 0.302 λ . Hence, this would imply for this scenario that the RSS technique is a more rigid tolerance than a Monte Carlo technique. So an optical system designed with respect to the RSS method would be better than one designed relative to the tolerances given by a Monte Carlo method.

Taking the RSS of two errors with a value of .25 λ we get 0.354 λ . From Figure 4 we see that the Monte Carlo method used on any combination of errors X-tilt, Y-tilt, X-decenter, and Y-decenter taken two at a time will give a range of .158 λ to .226 λ .

V. AN OPTICAL SYSTEM VIEWED AS AN AUTOMATON

Let $A = (S, \Sigma, M)$ be an automaton



 $\Sigma = \{\sigma/\sigma = a \text{ unit of measure of X-axis tilt of the secondary mirror in arc min}\}.$

Let N be some positive integer such that $\sigma^N = 0$. To obtain $\sigma^2 = \sigma\sigma$ we define σ^2 to be two units of tilt, e.g., let $\sigma = 1$ arc min then σ^2 is two arc min, etc. With this binary operation Σ makes a semigroup with zero which acts like an identity. Note Σ is semigroup isomorphic to $(Z_N, +)$. $(Z_N, +)$ is the semigroup of integers modulo N with respect to addition.

S = states of the system which will be the Zernike coefficients A_1, \ldots, A_k in the polynomial approximation. $W(X,Y) = A_1P_1 + \ldots + A_kP_k$. Note: k is to be a fixed positive integer; P_1 , $i=1,2,\ldots,k$, are the Zernike polynomials.

Now for $M_{\sigma} \in M$, $\sigma \in \Sigma$, define

$$M_{\sigma}: S \rightarrow S$$

 $A_{i} \rightarrow A_{i}M_{\sigma} = A_{i}$

so
$$M_{\sigma} \sim \begin{pmatrix} A_1 & \cdots & A_k \\ & & & \\ A_1 & \cdots & A_k \end{pmatrix}$$

Choose P_1, \ldots, P_k such that they are orthogonal with respect to the inner product

$$(P_{i}, P_{j}) = f P_{i}P_{j} .$$

$$X^{2}+Y^{2} \leq 1$$
Thus, if $M_{\tau} = \begin{pmatrix} A_{1} \cdots A_{k} \\ A_{1} \cdots A_{k} \end{pmatrix}$ then
$$M_{\sigma}M_{\tau} = \begin{pmatrix} A_{1} \cdots A_{k} \\ A_{1} \cdots A_{k} \end{pmatrix}$$
where $A_{i} = A_{1} + A_{i}$

VI. RESULTS AND DISCUSSION

The usual method of combining wavefront errors is the root sum square (RSS). For example, using the five errors above and combining them as follows:

$$\left(s_{1}^{2} + s_{2}^{2} + s_{3}^{2} + s_{4}^{2} + s_{5}^{2}\right)^{\frac{1}{2}}$$

Also, another method is to combine those by a Monte Carlo technique. If S_5 , despace, is not considered then we obtain Figure 4. Hence, an optical system designed to operate within the tolerance criterion of the RSS method would also satisfy the tolerance conditions of the Monte Carlo method.

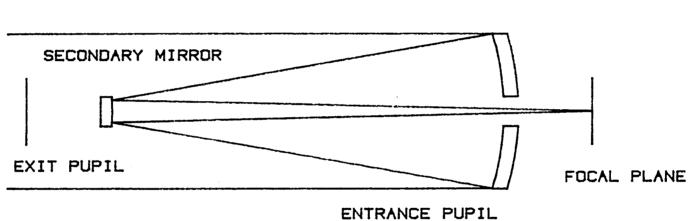
The maximum bounds of the error on focus were approximately -10 to $10 \ \mu\text{m}$. The focus proved to be a critical error in this analysis and further study is required.

Problems to consider:

1. Develop algorithms for an optical system considered as an automaton.

2. Extend this Monte Carlo method of tolerance analysis to more complicated optical systems.

3. Develop an algorithm and program it in FORTRAN for a Zernike coefficient program (see Davenport, NASA/ASEE summer report, 1979).



PRIMARY MIRROR



XXV-5

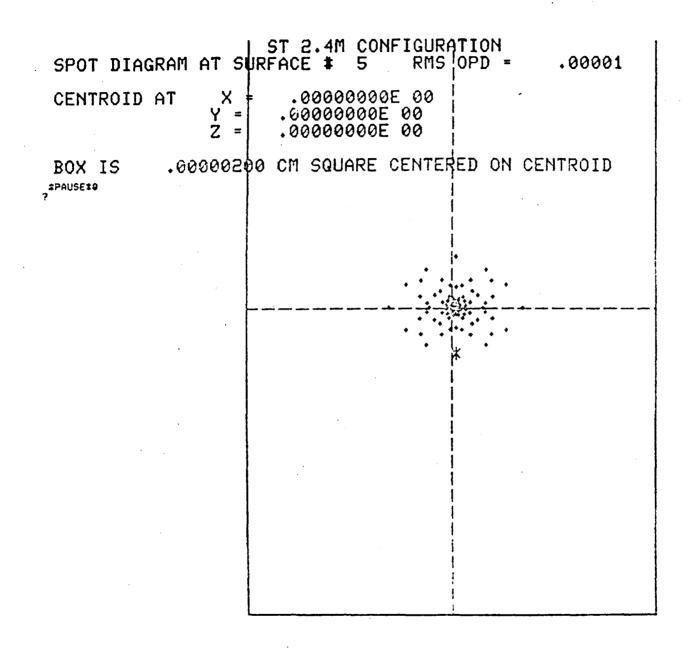


Figure 2. No Aberrations

8-VXX

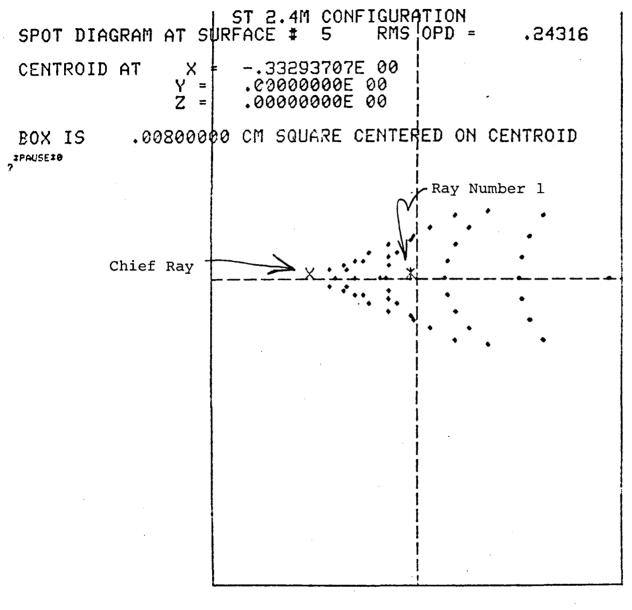


Figure 3. 0.1 Arc Min Tilt About X-Axis

XXV-7

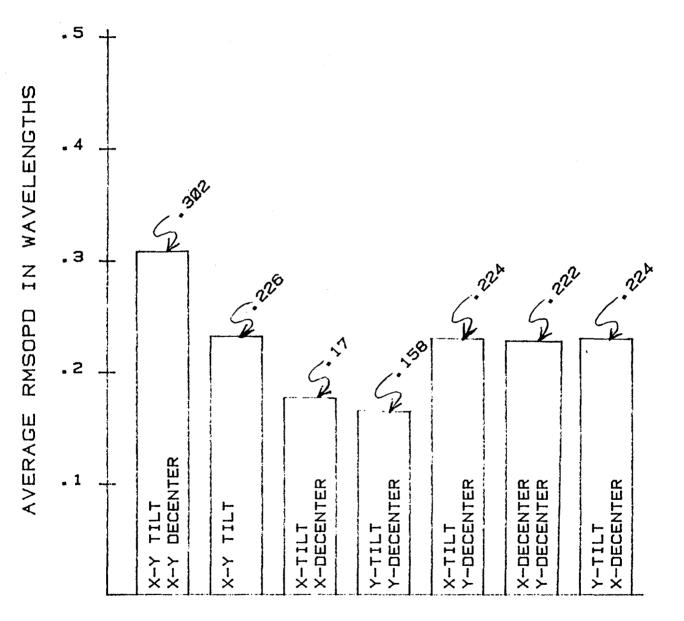


FIGURE 4 HISTOGRAM OF AVERAGE RMSOPD UNIFORM DISTRIBUTION OF 1000 POINTS.

ST 2.4M CONFIGURATION, FIELD PT. .100, .000 101 RAYS

RMS OPD	.2432
STREHL	-1.334
CENTROID:	
; X	3329
Y Y	.0000
Z	.0000
CHIEF RAY:	
Χ.	3351
Y	.0000

RANK 10

. 0

P1'= -.1079021E-01

ZERNIKE COEFFICIENTS **P**2 P3 P5 P6 P4 X TILT Y TILT FOCUS ASTIG(45) ASTIG(0) -1.46027 -.264879E-07 -.496520E-02 .163329E-18 -.642378E-04 P7 **P8** P11 P12 Y COMA X COMA 5TH-ORDER 3D-ORDER .182050E-07 -.495005 .141003E-04 .170063E-05 DESIGN PARAMETERS \$ D Ε QN · R -1.00228977203300 1 10.0000000000000 -1102.50000000000 -1.000002 DUMMY -490.0000000000000 3 -1.49599512243700 -135.464362850970 -1.000004 DUMMY 5 640.000000000000 -1.00000000000000 1000000000.00000 1.00000 í R1 =699.507000000000 120.000 R4 =5760.00 EE = .340000 FL = .632800E-04 NRAS = 13 QLO =\$STOP\$ 0

Table 1.

Zernike Coefficients with 0.1 Arc Min Tilt About the X-Axis <u> - ЛХХ</u>

ST 2.4M	CONFIGURATION,	FIELD	PT.	.000,	.000
	101 RAYS				

RMS OPD	.7330E-05
STREHL	1.000
CENTROID:	
×	.0000
Y	.0000
Z	.0000
CHIEF RAY:	
×	.0000
Y	.0000

RANK 10

P1 = .3218636E-04

÷ ZERNIKE COEFFICIENTS **P**2 P3 P4 **P**5 **P6** X TILT Y TILT FOCUS ASTIG(0) ASTIG(45) -.409722E-12 -.708786E-11 -.847167E-21 .298539E-12 .308430E-04 **P7 P8** P11 P12 3D-ORDER STH-ORDER Y COMA X COMA .251956E-12 .129803E-11 .163000E-05 .116445E-04 DESIGN PARAMETERS D R QN \$ E 1 10.000000000000 -1.00228977203300 -1102.50000000000 -1.00000 S DUWWA -490.000000000000 -135.464362850970 -1.49599512243700 3 -1.00000DUMMY 4 1000000000.00000 640.000000000000 5 -1.000000000000000 1.00000 699.507000000000 R1 = 120.000 R4 5760.00 .340000 FL • EE • .632800E-04 NRAS - 13 QL0 -

STOP 0

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REFERENCES

- Barakat, Richard, Optimum Balanced Wavefront Aberrations for Radially Symmetric Amplitude Distributions: Generalizations of Zernike Polynomials, Journal of the Optical Society of America, Vol. 70, No. 6, June 1980, pp. 739-742.
- Forsythe, George E., Michael A. Malcolm, and Cleve B. Moler, Computer Methods for Mathematical Computations, Prentice-Hall, 1977.
- 3. Freund, John E., Modern Elementary Statistics, 3rd ed., Prentice-Hall, 1967.
- 4. Hammersley, J. M., and D. C. Handscomb, Monte Carlo Methods, Wiley-Methuen and Co. LTD, 1964.
- 5. Ginzburg, A., Algebraic Theory of Automata, Academic Press, 1969.
- 6. Klein, Miles V., Optics, Wiley, 1970.
- 7. Knuth, Don, The Art of Computer Programming: Seminumerical Algorithms, Vol. 2, Addison-Wesley, 1969.
- Koch, Donald G., A Statistical Approach to Lens Tolerancing, Society of Photo-Optical Instrumentation Engineers, Vol. 147, Computer-Aided Optical Design (1978), pp. 71-82.
- 9. Kovach, Ladis D., Computer-Oriented Mathematics, Holden-Day, 1964.
- 10. Scalzo, Frank, and Rowland Hughes, Elementary Computer-Assisted Statistics, rev. ed., Van Nostrand, 1978.
- 11. Kuo, Shan S., Computer Applications of Numberical Methods, Addison-Wesley, 1972.

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

POWER DISTRIBUTION FOR ELECTRON BEAM WELDING

Prepared by:

Academic Rank:

University and Department:

NASA/MSFC: (Laboratory) (Division) (Branch)

NASA Counterpart:

Date:

Contract No.:

Eugene Edwards, M.S.E.E.

Assistant Professor

Alabama A.& M. University Department of Engineering and Engineering Technology,

Materials and Processes Process Engineering Metals Processes

Arthur C. Nunes, Ph.D.

August 1, 1980

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1980

POWER DISTRIBUTION FOR ELECTRON BEAM WELDING

By

Eugene Edwards, M.S.E.E. Assistant Professor of Electrical Engineering Alabama A&M University Normal, Alabama

ABSTRACT

Electron beams are being used to apply heat to metals for welding purposes.

In this report, the power distribution in the beam is analyzed. Experimental and digital computer techniques are used to evaluate the radial distribution of power detected by a wire probe circulating through the beam. Analytical models developed at MSFC of the Electron Beam-Workpiece interation contain two weld loss parameters in a function of the radial power distribution. The results from this research are intended to yield information on the mechanism by which these losses take place.

Additional information on the second loss mechanism is provided by measuring metal weight losses during welding. The power wasted on these weld losses can thus be estimated and distributed to obtain a more accurate weld penetration computation.

ACKNOWLEDGEMENTS

I am most grateful to ASEE, NASA-MSFC (Mr. Marion Kent), and the University of Alabama (Drs. Barfield and Karr) for making it possible for me to obtain this research fellowship.

I especially would like to recognize Dr. Arthur Nunes, my NASA Counterpart, for outstanding contributions toward the success of this work. Also at MSFC's Materials and Processes Laboratory, I acknowledge Paul H. Schuerer (Division Chief) and Melvin McIlwain (Branch Chief, EH42). The technical assistance of Pete Smith (EH15), Clifton Green (EH42) and Howard Novak (EH42) was appreciated, as well as the inspiration of Willibald Prasthofer (EH44) and Carl Wood (EH44). Many thanks to Doris Flowers and Beverly Robinson for typing and reviewing this paper, respectively.

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NOMENCLATURE

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AV	-	accelerating voltage
BC	-	beam current
BF	-	beam focus
D	-	diameter of the irradiated target
F	-	in welder's computer program, focus
FIL	-	filament
I	-	beam current
^I e	-	emission current
οK	-	degrees Kelvin
L	-	length of the wire probe
Р	-	perveance
Ρ _C	-	power lost to the crucible by conduction
Pd	-	power dissipated
Рe	-	power loss for electrons not reaching the target
Peff	-	effective power
Рg	-	power of the gun
Pi	-	power developed in the electron beam
P ₁ ,P'1	-	convergence points
PL	-	power loss due to latent heat
Pn	-	loss of power to the evaporant
Pr	-	power lost by heat radiation
Ρv	-	power lost by electrons not striking the target
Ρ _X	-	loss of power due to X-ray production
РОТ .	-	potentiometer
R	-	brightness of the beam
Т	-	cathode temperature
٧ _a	-	acceleration voltage
da	-	cathode diameter at point (a)
df		beam's spot diameter
ďk	-	cathode diameter at k
dm		smallest diameter of the beam
е	-	electron charge of 1.6 x 10^{-9} coul.
eVo	-	probable thermal energy

- j_F current density of the beam at the target
- j_k current density for beam current (J) at the cathode
- k constant, such as the k^{th} term, also K, as a constant
- kb Boltzman constant
- p power density at the target
- p'_F high power density
- α alpha, space-charge-limited current flow function
- $\boldsymbol{\alpha}_{1}$ aperture at the target
 - gamma, aperture angle of marginal rays between the anode
 - aperture of marginal rays at the position of the smallest beam cross section
- gamma, thickness of the wire probe
- omega, width of the wire probe
- ϕ phi, angle of current signal for computer input
- **π** pi, ≈ 3.14

Ż.

- ρ rho, beam target distance from motor-plate to some point on the probe
- $\boldsymbol{\Theta}$ theta, aperture angle of the cathode beam generator

OBJECTIVES

The objective of this study is to investigate the power distribution of the electron beam in order to understand the loss mechanisms associated with the beam-workpiece interation. Another objective of this research is to apply techniques that will yield distribution and profiles of the beam such that power losses may be carefully examined.

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INTRODUCTION

Electron Beam Welder (Sciaky Bros. Inc., Chicago, Illinois)

In an electron beam process, the part is welded by the Kinetic energy of a dense beam of accelerated electrons. As the electrons strike the workpiece, their Kinetic energy is converted to the thermal energy causing the metal to melt and fuse.

The device which emits, accelerates, and focuses the electrons on the workpiece is referred to as the electron beam gun. The gun can be mounted on a carriage to provide several axes of motion, or it can be held stationary while the work is driven, or both the gun and the work can be driven. See Figure 1 below for a simplified illustration.

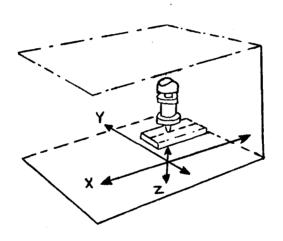
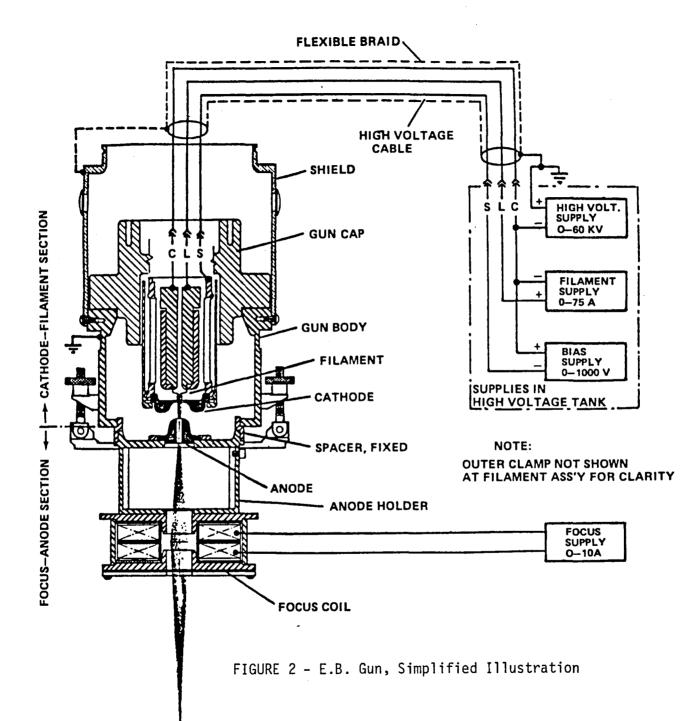


FIGURE 1 - Illustration of Travel Direction

The gun assembly is housed in a chamber which is pumped down to a fine vacuum level, usually to about 1×10^{-4} torr. or less. By operating in a vacuum the life of the filament (Figure 2) is protected and the electrons can be focused into a fine dense beam. The filament which is heated to a high temperature would be destroyed if exposed to air at the operating temperature. If there were a sufficient concentration of loose molecules in the chamber, the electrons would collide with these molecules. This would cause the beam to be diffused and the electrons to lose some of their Kinetic energy.

In Figure 2, the voltage of the bias allows the beam current to be changed independent of the accelerating potential (within the operating limits of the gun). Irrespective of the bias level, the maximum beam current cannot exceed the maximum current range of the gun. This range is determined by the mating elements installed in the gun. Maximum current is obtained at 0 bias. As the bias is increased, the current is reduced.



The spacing between the filament, cathode, and anode determines the internal resistance of the gun. This resistance is non-linear and varies directly with the distance between these elements. As the distance is increased, the beam current is decreased.

The model E-S9460 gun has 2 major ranges; 250 mA and 500 mA. A mating filament and cathode are provided for each range. In each of these 2 ranges, additional ranges can be obtained by inserting a spacer between the 2 main sections. The same anode is used in both major ranges.

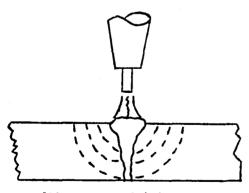
Ranges of Gun E-S9460

Range	:	500 mA		:	250 mA
Filament	:		Filament	::	C-G8615
Cathode	:	B-K2501	Cathode	:	B-H844
**Anode	:	B-M5740	**Anode	:	B-M5740
		onal ranges use: /0.050" spacer	150 mA 100 mA	١	tional ranges use: W/0.100" spacer W/0.200" spacer W/0.500" spacer

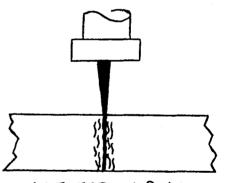
For application requiring shallow penetration with narrow welds at low currents, a small hole cathode is used.

Electron beam welds posses several unique characteristics:

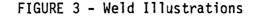
 Welds are produced in many cases 15 to 20 times narrower than comparable TIG/MIG welds. Notice the dotted isothermal lines of typical Tungsten Arc welding compared to Electron Beam welding (Figure 3).







(B) ELECTRON BEAM



**The same cathode is used in both major ranges.

- (2) Less thermal energy is transferred to the unmelted base metal and thus reducing residual stresses, warping, and increasing the strength of the weld.
- (3) Absorption of atmospheric gasses by the weld is eliminated by the vacuum system, thus highly oxidizable materials can be welded (Zirconium, Beryllium, etc.).

POWER RELATIONS

The output power from the electron beam gun follows the general power formula of the form

$$P_{g} = V_{a}I_{e}, \tag{1}$$

where, P_q is the gun power

 V_a is the acceleration voltage

and I is the emission current.

Power, P_d , is lost as the electrons are taken off of the anode. This dissipated power must be subtracted from the gun power.

Since all of the electrons will not reach the welding target the gun power becomes even smaller. The beam focus plays an important role in allowing more electrons to reach the target. Now, the effective power becomes

$$P_{eff} = P_g - P_d - P_e, \qquad (2)$$

where P_{ρ} is the power lost due to electrons not reaching the target.

For certain applications, a low precision gun with a diffused output beam may allow a little more than 50% of the emitted electrons to reach the target [v],

Energy is lost when the electrons collide with system components and target samples. Smith, reference [1], attribute some of this energy loss to:

- excitation and ionization of electrons, by electronelectron interaction, and displacement of atoms with the lattice by energy trnasferred to individual atoms. This occurs at energies of a few-hundred KEV.
- X-ray production by resonance absorption occurs at the high energy range. These losses are minor.
- o inadequate focusing method creating minor loss of electrons from the beam.

- o vapor between the welding sample and gun can cause minor losses due to collisions.
- back scatter of electrons reflected produces the main source of loss (25%).

It is generally safe to assume that the effective power that is converted to heat the sample is

$$P_{eff} = 0.60 P_{q}.$$
 (3)

Note that equation Eq. (3) is approximately the equation for average power. Information on evaporation losses will be presented later in this paper.

The important relationship between beam power and quality of the weld penetration can be obatained for various metals, metal sizes, and welding speeds. A computer routine is available that gives results similar to those shown in Figure 4 (NASA-MSFC).

A summary of the power-energy distribution is given by Hunt and Hughes 2 as the process heat balance:

$$P_{i} = P_{v} + P_{r} + P_{L} + P_{n} + P_{x} + P_{c}, \qquad (4)$$

- where, P_i is the power developed in the electron beam by acceleration to the anode voltage.
 - P_v is the power lost by electrons striking other than the target; e.g., filament-to-anode current (negligible).
 - P_r is the power lost by heat radiation from the liquid evaporant surface (hard to determine).
 - P_L is the power loss due to the latent heat of evaporation for the evaporant (available for most materials).
 - P_n is the loss of power to the evaporant by ionization and secondary electron generation (due to vapor density and electron back-scattering).
 - P_v is the power loss due to production of X-rays (negligible).

P_c is the power loss to the crucible by conduction (very high).

Table I (reference [2]) on page XXVI-15 shows data on power requirements of typical sources evaporating various materials from a 4-inch-diameter water cooled copper crucible, fed from the bottom with rods of evaporant. This data implies that there appears to be an optimum beam density above and below which evaporation rate decreases for a given level of electron beam power at the evaporant surface.

- 2219 Aluminum--Plate Thickness .75 Inches
- 50% Nonconductive Losses--10% to Nailhead
- .08 Inch Root Width

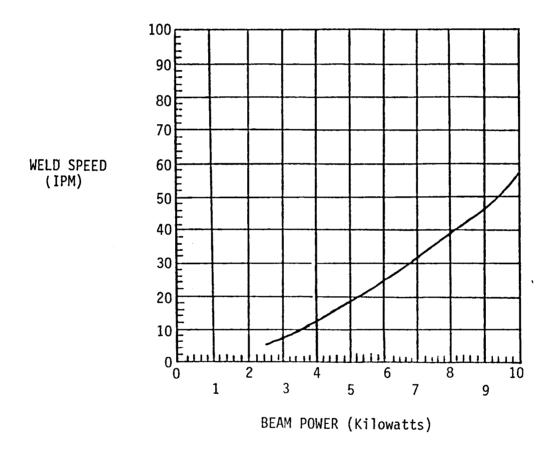


FIGURE 4 - Typical Minicomputer Output on Beam Power and Weld Speed

TABLE 1

•

<u>Material</u>	Power In	Evap. Rate	Specific Power Requirement
Zirconium	72kw	2.57 lb/hr	28.0 kwh/1b
20% Cr S.S.	70kw	9.38 lb/hr	7.5 kwh/1b
Ti-6A1-4V	70kw	5.17 lb/hr	13.5 kwh/1b
Aluminum-Bronze	80kw	8.96 lb/hr	13.1 kwh/1b
Inconel 600	70kw	5.06 lb/hr	13.8 kwh/1b

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TABLE 1 - Power Relations for Evaporation Loss

EQUATIONS OF THE BEAM CURRENT AND POWER DENSITY

When the E.B. gun is operating, its beam current follows the equation [3]

$$I = P V_a^{3/2}$$
 (5)

where, V_a is accelerating voltage and P is the perveance. Schiller, Heisig, and Lenk [3] used the equation below for perveance (P) as well as the following equations including a deviation of the power density (p_f) :

P = perveance =
$$15 \times 10^{-6} \frac{(1 - \cos \theta)}{\alpha^2}$$
, (6)

- where, \checkmark is a function of the space-charge-limited current flow depending on the relation,
 - \ll = f $\frac{d_k}{d_a}$, between the cathode diameter
 - d_k and the anode bore diameter d_a . The f implies function of.

 is the aperture angle of the cathode beam generator or the convergence angle of marginal rays in the beam generator (See Figure 5).

$$eV_0 = probable thermal energy = k_b T = \frac{1}{11,608} ev$$
, (7)

where,

- V_o is thermic electron potential
- e is the charge of 1.6×10^{-19} coul.
- k_b is the Boltgmann constant in $\frac{ev}{deg}$.
- T is the cathode temperature $^{\circ}K$.

A Maxwellian energy distribution is assumed with usual emission temperatures, perhaps 3000°K.

R = brightness of the beam =
$$\frac{J_k V_a}{\pi V_o} = \frac{J_k e V_a}{\pi k_b T_K}$$
, (8)

where,

 $\mathbf{j}_{\mathbf{k}}$ is the current density for beam current (J) at the cathode.

 $p = power density at target = j'_{F} V_{a}$ (9)

where,

 $\mathbf{j'}_{\mathsf{F}}$ is the current density of the beam at the target and

equals $R\Pi \alpha_1^2$. Here, α_1 is the aperture at the target.

An arrangement of Equations (7), (8), and (9) yields (a high power density p_F)

$$P_{\mathsf{F}} = \frac{\mathbf{j}_{\mathsf{k}}}{\mathbf{V}_{\mathsf{0}}} \, \boldsymbol{\alpha}_{\mathsf{L}}^{2} \cdot \, \mathbf{V}_{\mathsf{a}}^{2}. \tag{10}$$

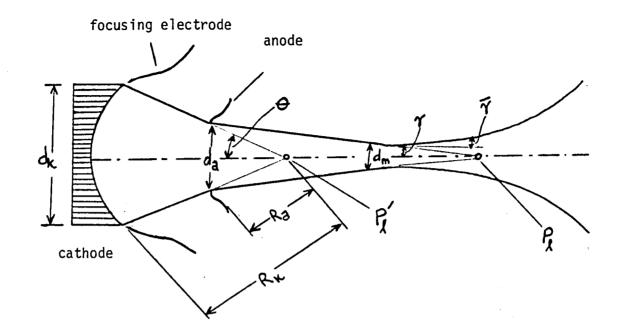


FIGURE 5 - Cross Section of a Spherical Beam Generation System of Pierce-type (Schematically)

d_k = diameter of cathode

da

Θ

d_m

Ÿ

- = diameter of anode bore
- = aperture angle of the beam generator convergence angle of marginal rays in the beam generator
- $P'_1 P_1 = convergence points$
 - = smallest diameter of the beam consider the intrinsic space charge
 - = aperture of marginal ray at the position of the smallest beam cross section

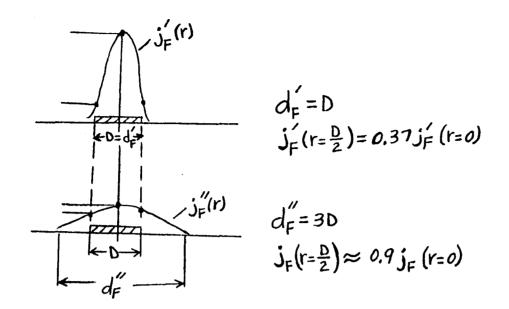


FIGURE 6 - Reduction of the Current Density Difference Between Center and Margin of an Irradiated Area of Diameter D by an Increased Electron Beam Diameter $d_{\rm E}^{\rm E}$

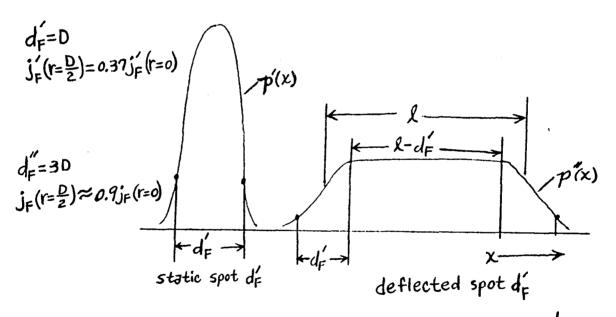


FIGURE 7 - Creating a Constant Power Density p''(x) on a Length $1 - d_F'$ by Time Linear Deflection of an Electron Beam d_F'' over the Length 1. (The range over which power density decreases to 0.63 p''(x=0) is = d_F')

Reference [3] also used Figures 6 and 7 to show that time-linear deflection is one coordinate results in a power density distribution (deflection $X = \pm .5$) of constant values in certain ranges. Power outside this range decreases and is a power loss. The term d_F is the beam's spot diameter and D is the target diameter (irradiated).

TECHNIQUES FOR MEASURING AND EVALUATING THE ELECTRON BEAM

This section explains the experimental and computer aided approaches used to evaluate the beam configuration.

The initial problem task for this section is that of determining what available method and apparatus to use to take measurements of the beam. The rotating probe (or wire) approach was chosen. This method utilizes a small heat protected motor to turn the wire probe through the beam. The signal obtained on the probe is fed into a storage oscilloscope. Pictures of the signal for various operating conditions of the system are taken from the scope and analyzed. The technique has the following advantages:

- (a) simple
- (b) apparatus available and not refined
- (c) beam current is measured
- (d) beam diameter is obtainable

The wire probe did not obtain appreciable damage from the beam and adequately allowed suitable measurements to take place.

Adjusting the accelerating voltage (thus power output) appears not to change the shape nor the general dimensions of the beam.

The experimental setup used to measure the beam is shown in Figure 8. The oscilloscope was used to monitor and record not only the beam signals but general signals (noise) in the vacuum. During pump down the oscilloscope was triggered by noise signals equivalent to those shown in Figure 9.

Since this electron beam welding machine receives its instructions from a computer controlled system, the sample computer program for a typical set of data inputs is shown on page XXVI-23.

When the beam comes into contact with the rotating wire probe, the electrons that are not backscattered or secondary-emitted are received as the current of the beam (perhaps 50%). The current through the wire has the common variation in magnitude at various angles of emergence. The surface contour of the current is obtained from the electron distribution. Care must be taken in determining the needed current signal configuration and associate information instead of bad data. The collecting wire probe has a symmetrically fixed amount of wire extending from two sides of the plate affixed to the

VACUUM CHAMBER

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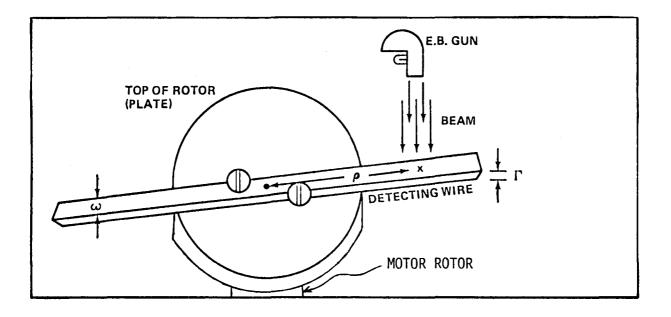
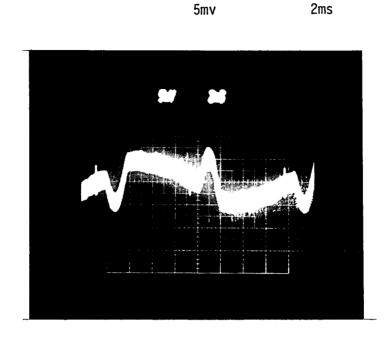
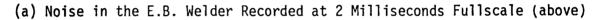
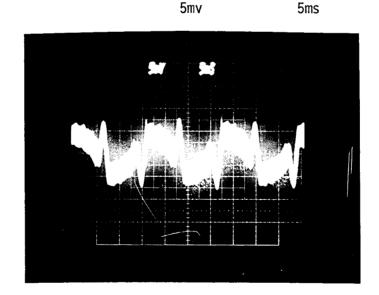


FIGURE 8 - EXPERIMENTAL SETUP







(b) Noise in the E.B. Welder Recorded at 5 Milliseconds Fullscale (above)

FIGURE 9 - Signals Triggering the Scope Due to Noise Present During Pumping of Vacuum Chamber

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TYPICAL DATA RESULTS

ρ = 2.833 INCHES

:

 ω = .048 INCHES

 Γ = .003 INCHES

BF (60 KV) 5.00 FIL 3.50

AV 30 KV

BC 50 MA

THE SAMPLE COMPUTER PROGRAM FOR THE WELDER IS SHOWN BELOW.

FILE RECORD NASA EB WELDER			
FILE NO. 110	BEAM CROSS SECTION STUDY PROGRAM	PAGE 1	
(BEAM CROSS SECT) OVE POT 0 G1APT X0 Y0 F500 FIL350 F20 BF500 F28 AV300 BC50 F20 AV300 BC50 F1000 AV0 BC0 F20	ON STUDY PROGRAM) M0		

A TABLE OF COMPUTER INPUT DATA (Ø,I,ETC...) IS DEVELOPED FROM FIGURE 11 FOR USE ON THE DIGITAL MINICOMPUTER.

rotor of the motor (as was shown in Figure 8). The motor was shielded from the beam and from heat by a thin sheet of metal. The current collected was sent through a load resistor and the output from this load to the oscillo-scope.

In order to adequately measure the beam current, the wire probe dimensions were carefully measured (Figure 10). The beam was also precisely applied to a target area of the probe at a chosen radius from the center of the plate of the motor.

Figure 11 shows the typical current signal and can be plotted versus angle position. With the given beam characteristics or welding operation status, the signal profile is always of a similar shape. To avoid possible aggravation of the signal, it should be saved on the scope as soon as a clear configuration is observed. The scope will automatically trigger to the signal prior to being available for saving the signal. Interferences during signal generation, such as surfaces disturbances not allowing adequate distribution of beam electrons, were considered minor.

Two cases representing the output current signal are in Figures 12 and 13. The first case (Fig. 12) is the peak beam current for two pulse signals. The second case (Fig. 13) is an expanded version of one of the pulse signals distributed with a wider profile. Several sweep time divisions of the scope were investigated for many conditions until satisfactory signals were realized.

The current vs. probe angle was plotted and used as inputs for the computer routine that calculates the current density per unit radius. Additional inputs to the computer included (a) the distance from the center of the plate of the motor to a target point on the wire, (b) the width of the wire probe (receptor), (c) the total beam current, and (d) the number of current-angle data points.

As can be seen from Table 2 (page XXVI-28) and Figures 14 and 15, a current density profile can be obtained. The radial distribution is modulated by a strong periodic error signal of a sort known to arise from small errors (about 0.1%) in data. Work is continuing to obtain current density profiles.

The measurements taken from this experimental setup show that much more beam power is lost than can be explained from measurements taken from the wire probe, unless very large amounts of energy are carried away per secondary emission or backscattered electrons. This justifies the need to continue investigating the evaporation power loss and other power losses accounting for this reduction of total beam power. This phenomenon occurs quite often when samples are being welded. Some information is available on secondary emission losses from a 1974 study conducted at NASA-MSFC.

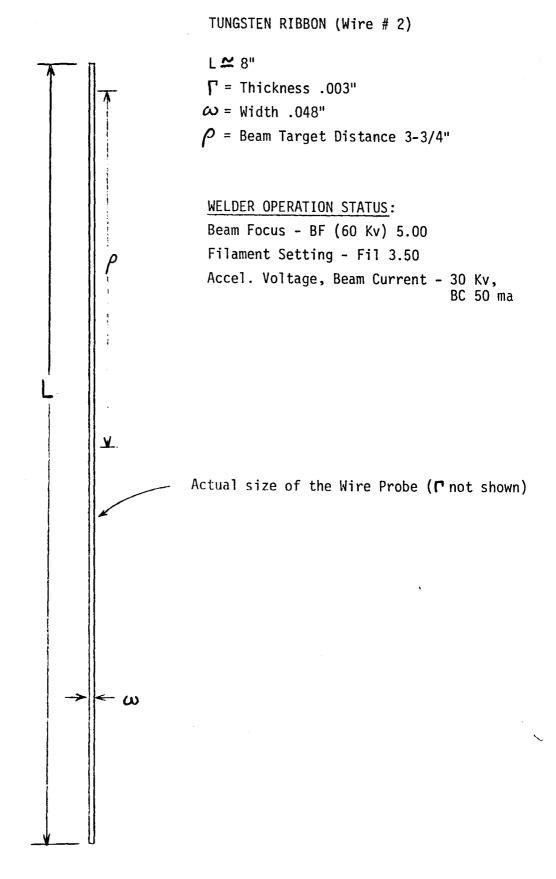


FIGURE 10 - Wire Probe Characteristics

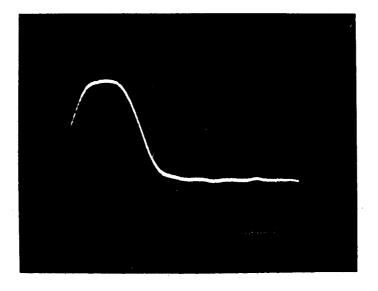


FIGURE 11 - Typical Beam Current Configuration (actual photo)

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EXPERIMENTAL RESULTS

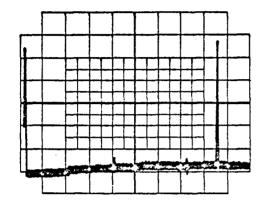


FIGURE 12 - Peak Beam Current for Two Pulse Signals

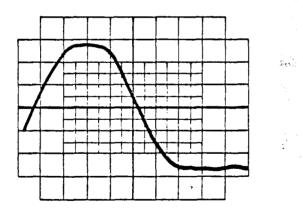


FIGURE 13 - Single Pulse Signal Distributed

THE CHARACTERISTICS OF THE TUNGSTEN WIRE ARE

.

- $\rho, \ \ \,$ THE DISTANCE FROM THE CENTER OF THE MOTOR PLATE TO A TARGET POINT ON THE WIRE
- ω , THE WIDTH OF THE WIRE
- Γ, THE THICKNESS OF THE WIRE

THE E.B. WELDER WAS COMPUTER CONTROL FOR AN OPERATION STATUS FOR VALUES OF

BF, BEAM FOCUS

- FIL, FILAMENT SETTING
- AV, ACCELERATING VOLTAGE BC, BEAM CURRENT

TYPICAL DATA RESULTS ARE SHOWN ON THE NEXT PAGE.

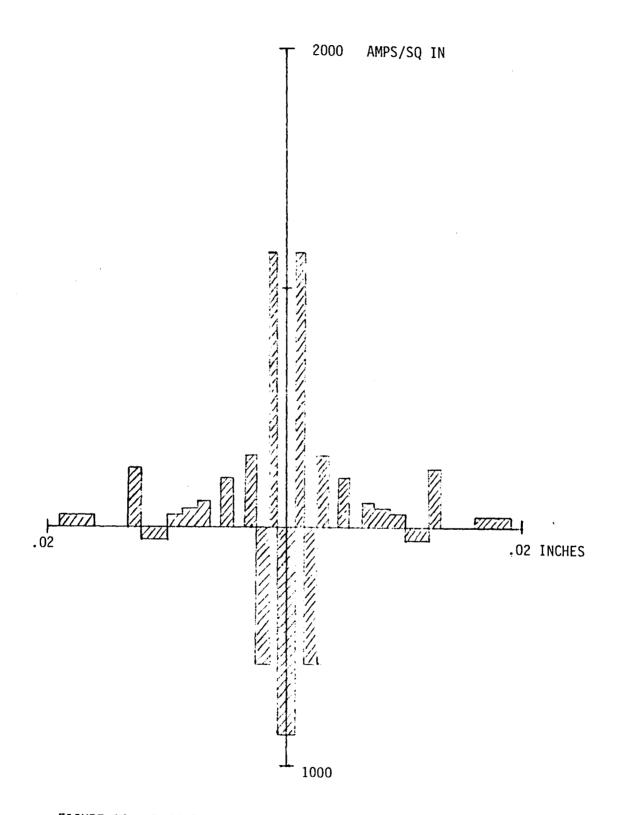
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TABLE 2

ANGLE	CURRENT	RADIUS	CURRENT DENSITY
(DEG)	(MA)	(IN)	(AMPS/SQ IN)
.017	26.32	8.40541E-04	-832.032
.031	28.15	1.53275E-03	1153.54
.053	28.97	2.62051E-03	-571.086
.074	31.79	3.65883E-03	301.177
.093	33.61	4.59827E-03	.0935703
.11	35.43	5.43882E-03	214.231
.134	37.25	6.62548E-03	-3.99142
.154	39.08	7.61436E-03	104.89
.178	40.89	8.80103E-03	75.7133
.211	42.72	.0104327	66.7768
.248	44.54	.0122622	-57.2598
.266	46.36	.0131522	246.266
.327	48.18	.0161683	6,38338
. 386	50	.0198856	45.3961

Ratio of Total Current Calculated to Total Beam Current.

TABLE 2 - Digital Computer Results on Current Density





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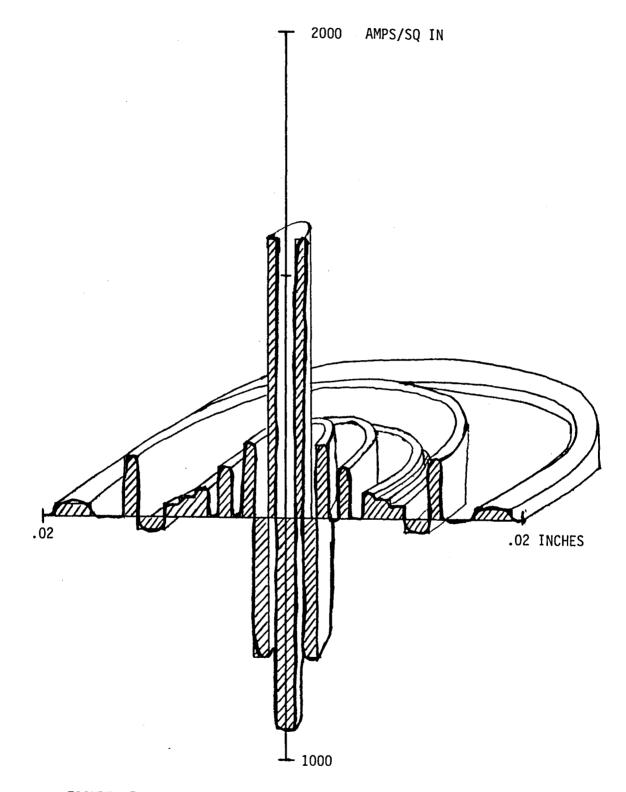


FIGURE 15: Radial Current Distribution (profile)

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CONCLUSIONS

At the present time we are able to capture by experimental techniques the current configuration for an electron beam. The computer routines allow us to theoretically determine the beam characteristics on the workpiece for a given power profile. Techniques developed and described in this report can be valuable for gathering information needed to assess the nature of weld power losses.

Future work should be done in order to relate accurately the total beam characteristics and power distribution to weld geometry. However, we have successfully completed the experimental task of obtaining the signal configuration which determines the cross sectional power distribution in the electron beam.

The results of Figures 11 and 13 compare with those obtained by Sayegh and Dumonte $\begin{bmatrix} 4 \end{bmatrix} \begin{bmatrix} 5 \end{bmatrix}$.

An analysis for obtaining the radial current density and the profiles in Figures 14 and 15 is available but small errors in data (or deviations from a radial distribution) apparently cause strong periodic error modulations in the radial distribution.

RECOMMENDATIONS

1. Techniques for obtaining more information on electron reflection should be investigated.

2. Controlled use of the welding system such that the time wasted in equipment venting and pump down is minimized.

3. Advanced studies should be done on the trajectories of the beam (both gaussian and nongaussian type).

4. Curvilinear distribution analyses should be considered for further study.

5. Relate the diameter of the beam and the power distribution on a welding sample in items 3 and 4 above.

6. Determine if the target surfaces cause any disturbance in the distribution of the beam's electrons on the wire probe.

7. Determine the optimum setting for the E.B. welder in order to obtain the optimum power density profile.

REFERENCES

- Smith, H. R., "Principles of Electron Beam Technology," 2nd Electron Beam Processing Seminar (R. M. Silva, ed.), Universal Technology, Dayton, Ohio, June 1972.
- Hunt, Charles d'A. and Hughes, J. L., "Principles of Electron Beam Evaporation," 2nd Electron Beam Processing Seminar (R. M. Silva, ed.), Universal Technology, Dayton, Ohio, June 1972.
- 3. Schiller, S., Heisig, U., and Lenk, P., "Theory and Development of Electron Beam Sources," 2nd Electron Beam Processing Seminar (R. M. Silva, ed.), Universal Technology, Dayton, Ohio, June 1972.
- Sayegh, G. and Dumonte, P., "100 KW EB Gun for Welding-Design, Realization and Applications," 4th International Electron Beam Processing Seminar (R. M. Silva, ed.), Universal Technology Corporation, Dayton, Ohio, April 1976.
- 5. Sayegh, G. and Dumonte, P., "Theoretical and Experimental Techniques for Design of Electron Beam Welding Guns, Effects of Welding Conditions on Electron Beam Characteristics, "3rd Electron Beam Processing Seminar Proceedings, Universal Technology Corporation, Dayton, Ohio, March 1974.

1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

MEASUREMENT OF VELOCITY FIELDS IN FLUID DYNAMICS EXPERIMENTS

Prepared by:

Academic Rank:

University and Department

1

NASA/MSFC: (Laboratory) (Division) (Branch)

MSFC Counterpart:

Date:

Contract No.:

Daniel E. Fitzjarrald

Research Associate

Geophysical Fluid Dynamics Institute Florida State University

Space Sciences Laboratory Atmospheric Sciences Fluid Dynamics

.

George H. Ficht1

August 6, 1980

NGT-01-002-099 (University of Alabama)

XXVII.1

MEASUREMENT OF VELOCITY FIELDS IN FLUID DYNAMICS EXPERIMENTS

by

Daniel E. Fitzjarrald Geophysical Fluid Dynamics Inst. Florida State University Tallahassee, FL

ABSTRACT

In many fluid dynamics experiments in which the motion has both spatial and temporal variations, it is necessary to measure the entire two-dimensional field of velocity at an instant. Two methods to accomplish this measurement have been investigated. In both cases the emphasis has been on using a microcomputer-video system in the analysis of data. The first method is the traditional one in which tracers are introduced into the fluid and their position compared at two closely spaced times. The second method involves scattering coherent light in the fluid and obtaining motion by analyzing the multipleexposed speckle pattern recorded on photographic film. Both of these velocity field methods can be used in the geophysical fluid flow and space processing experiments in progress at NASA.

1. Introduction

In many fluid dynamics experiments both spatial and temporal variations of the flow patterns are important. This is particularly true in experiments modelling geophysical flows and in space-processing experiments, both in progress at NASA. With conventional instrumentation the experimentalist must rely on the use of fine wire probes, laser-Doppler measurements, or other similar techniques that are all limited to a single point in space. It is possible to move the measuring point around in the fluid, but the technique is never satisfactory when there are significant time variations in the flow field.

Streak photography has often been used in such cases, i.e., taking photographs of the displacement of tracers in the velocity field. Data analysis is quite cumbersome, however, whenever there is much variation in the flow field, or when a dense coverage of data points is required.

It is the goal of this summer's research to investigate a promising alternative to conventional velocity field measurement. The technique to be investigated will be termed light-speckle techniques, and all are closely related to streak photography. The key in each case is that the tracers and time intervals are chosen so that optical correlation techniques can be used to recover velocities from the data.

2. Light-Speckle Velocity Field Measurement

Although there are several variations in details, these techniques are all basically the same. Light from scatterers that move with the fluid is recorded at closely-spaced time intervals on photographic film. The optical pattern from the first instant is slightly changed by the fluid motion at succeeding time intervals. Thus, when the multiple-exposed negative is interrogated by a narrow beam of coherent light, the correlation between patterns recorded at succeeding times is indicated by a pattern of regular fringes. The spacing of the fringes tells the amount of displacement, i.e., the time interval and the orientation of the fringes tells the direction of displacement. Thus, by illuminating the fluid with a thin sheet of light and viewing it at a right angle to the plane of illumination, an entire two-dimensional field of velocity can be obtained at an instant.

2.1 Generating the Optical Pattern

The optical pattern that is to be displaced by the fluid can be generated either by using the coherent nature of laser light with very small scatterers, or with white light and larger scatterers. The former method has been used by Simpkins and Dudderer (1978) and is based on the fact that laser light appears grainy The grainy when diffusely reflected off a surface. appearance is caused by constructive and destructive interference of coherent light scattered off roughness elements the same size as the wavelength. These "objective speckles" are a function of the size of the scatterers, while the "subjective speckles" seen by the observer or recorded on film are dependent on the aperture size of the lens through which they are viewed. When the displacement of the recorded subjective speckles due to fluid motion is between one and twenty speckle diameters the correlation will yield a good fringe pattern. The fringes will also be enhanced if more than two optical patterns are recorded on the same negative.

A useable optical pattern can also be obtained by using larger scatterers illuminated by white light (Chiang and Asundi, 1979), or by a laser (Meynart, 1980). In the latter case the spatial coherence of the laser light is used only to obtain a thin light sheet of high power density. Again, multiple exposures increase the visibility of the fringe.

2.2 Interrogating the Negative

Once the multiple-exposed negative has been obtained, there are two methods to recover the velocity field. Illuminating the negative point-by-point with a small coherent beam yields fringes that come from a small portion of the flow field where the velocity was constant. The negative is thus interrogated at each point in the field that the velocity is required.

A whole-field method of analysis can be obtained by illuminating the entire negative and using a spatial filtering setup (see Meynart, 1979) that consists of two lenses (Fourier transformers) and an off-axis pinhole

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(spatial filter). The result is the whole velocity field described in isolines of the component of velocity depending on the orientation of the spatial filtering hole.

There are advantages to both these interrogation methods. The first yields results that are more precise, with the limits of precision depending on the quality of the fringes that are obtained. The wholefield method is simpler and much quicker.

3. Use of Microcomputer in Analysis

In order to speed up analysis to make possible the evaluation of fast-changing geophysical and spaceprocessing fluid flows, it is desirable to automate the data analysis as much as possible. Using a microcomputer-TV system it should be possible to accomplish the point-by-point interrogation of the multipleexposed negatives automatically. The microcomputer will be programmed to count the fringes, determine orientation, and then move the film transport on to the next position. In addition, the TV system can determine velocity by simply subtracting successive positions of large scatterers to serve as a check on the velocities obtained using the speckle techniques.

4. Planned Work and Work Accomplished to Date

At this writing the experimental work is just getting into progress. Multiple exposures have been taken in a small water flume and thermal convection cells illuminated by a 1 watt CW argon laser. The pulse repetition rate and flow rate have been varied to determine optimum conditions for obtaining good fringes.

In addition considerable program development work has been accomplished on the TV-microcomputer system. Picture enhancement and pattern recognition programs have been written and debugged, so that they can be put together to measure fringes and track individual scatterers.

5. Conclusions

The light-speckle velocity field techniques described above offer tremendous possibilities toward the understanding of geophysical and space-processing fluid flow experiments. The ability to obtain an entire two-dimensional velocity field at an instant makes these techniques extremely powerful. The experiments and computer program development in progress should demonstrate the utility of these lightspeckle techniques.

References

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Chiang, F. P., and A. Asundi, 1979. White light speckle method of experimental strain analysis. Applied Optics, <u>18</u>, 409-411.

Meynart, R., 1980. Equal velocity fringes in a Rayleigh-Benard flow by a speckle method. Applied Optics 19, 1385-1386.

Simpkins, P. and T. Dudderer, 1978. Laser speckle measurements of transient Benard convection. J. Fluid Mech., 89, 665-671.

1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM MARSHALL FLIGHT CENTER THE UNIVERSITY OF ALABAMA

ON THE CONTROL AND STABILITY OF THE PINHOLE-OCCULTER FLEXIBLE BOOM FACILITY

PREPARED BY:	Michael Greene, PhD		
ACADEMIC RANK:	Assistant Professor		
COLLEGE AND DEPARTMENT:	University of Alabama Department of Electrical Engineering		
NASA/MSFC: (OFFICE) (DIVISION) (BRANCH)	Preliminary Design Subsystems Design Navigation and Control		
MSFC COUNTERPART:	J. R. Parker		

DATE:

CONTRACT NO:

July 15, 1980

NGT-01-002-099 (University of Alabama)

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ON THE CONTROL AND STABILITY OF THE PINHOLE-OCCULTER FLEXIBLE BOOM FACILITY

By

Michael Greene, PhD Assistant Professor of Electrical Engineering University of Alabama Tuscaloosa, AL

Abstract

The dynamics and control of the Shuttle based, gimbal mounted, boom pointing system for use in the x-ray pinhole experiments were studied. The overall system stability was studied, and root locus compensation was designed to achieve stability and pointing stability requirements. This single axis control study found the compensated system to be stable and capable of pointing to within 6.5 arc sec while the Shuttle rocks in its 0.1° deadband. The system can follow step (position) inputs with zero error and ramp (velocity) inputs with a steady state error coefficient of 4.5 sec^{-1} .

Introduction

The pinhole-occulter system is a Space Shuttle based experiment for the production of hard X-ray images taken primarily from the Sun. The system is basically a pinhole camera utilizing a deployable 50-m flexible boom for separating the pinhole from the recording devices located in the Shuttle as seen in Figure 1. At the distal end of the boom from the Shuttle is a 50 kg mask containing two sets of pinholes and two chronograph shields as seen in Figure 2. At the proximal end the detectors are located and mounted, along with the boom, to a gimbal pointing system (either IPS or AGS) mounted in the Shuttle Payload Bay.

The mask of Figure 2 must be pointed at the X-ray source with a pointing stability of less than 10 arc sec to align the axes of the detectors with the pinholes and shields. Failure to do so will result in a blurring of the images on the detectors and a loss of resolution. Being a Shuttle based experiment, the system will be subjected to the disturbances of the Shuttle. The worst of these is thruster firing for orbit correction; Shuttle uses a bang-bang thruster control system to maintain orbit to within $\pm 0.1^{\circ}$. Other disturbances include man motion, motion induced by other systems, and gravity gradient torques.

The control system of the pointing mount senses both position and velocity of the mask tip and uses these to accurately point the boom. The AGS employing perfect sensors was used in this study of the single axis (y-axis) control system. The extension of this study to a three-axis system will be the topic of a later study.

Background

The structural analysis of the boom [1,2,3] revealed the seven mode shapes and natural frequencies shown in Table 1. These natural frequencies varied with both boom length and tip mass as seen in Figure 3. For a 50-m boom with a 50 kg tip mass, the first two natural frequencies are 1 rad/sec for deflections about the x and y axes, respectively. The third natural frequency is 17.1 rad/sec for z axis motion. For x and y motions the next two natural frequencies are 17.4 rad/sec and 59.9 rad/sec, respectively.

For y-axis control (torques about y-axis), the equations of motion for the boom, assuming a damping factor of 10 percent, are [1]:

$$\dot{\eta}_1 + .2 \dot{\eta}_1 + \eta_1 = -3687.93 \dot{\theta}_2$$
 (1)

$$\dot{\eta}_2 + .2 \dot{\eta}_2 + \eta_2 = 0 \tag{2}$$

.

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$$\dot{\eta}_3 + 3.42 \dot{\eta}_3 + 292.4 \eta_3 = -.018396 \dot{\theta}_2$$
 (3)

$$\dot{\eta}_4 + 3.46 \dot{\eta}_4 + 300 \eta_4 = 1250.734 \theta_2$$
 (4)

$$\dot{\eta}_{5} + 3.46 \, \dot{\eta}_{5} + 300 \, \eta_{5} = 0 \tag{5}$$

$$\dot{\eta}_{6}^{*} + 11 \dot{\eta}_{6}^{*} + 3014 \eta_{6}^{*} = 52.3905 \theta_{2}^{*}$$
 (6)

$$\dot{\eta}_{7} + 11 \dot{\eta}_{7} + 3014 \eta_{7} = 0$$
 (7)

$$\theta_{\rm TIP} = 2.895 \times 10^{-4} \eta_1 - 8 \times 10^{-10} \eta_3 + 5.31 \times 10^{-5} \eta_4$$
$$- 3 \times 10^{-5} \eta_6 + \theta_2 \qquad (8)$$

Taking the Laplace transform of equations (1) through (7) and inserting into equation (7), we obtain:

$$\theta_{\text{TIP}}(S) = \left[\frac{-1.07S^2}{S^2 + .2S + 1} + \frac{1.5 \times 10^{-11} S^2}{S^2 + 3.42 S + 292.4} + \frac{.07S^2}{S^2 + 3.465 + 300} + \frac{1.57 \times 10^{-3} S^2}{5^2 + 11S + 3014} + 1\right] \theta_2(S) \quad .$$
(9)

Ignoring terms 2 and 4 and reducing, the ratio of input motion to output motion for the boom is obtained:

$$\frac{\theta_{\text{TIP}}(S)}{\theta_2(S)} = \frac{-.07 (S^2 - 2.865 - 14.28)}{S^2 + .2S + 1} + \frac{.07S^2}{S^2 + 3.48S + 300}$$
(10)
$$= \frac{-.0289 (S^3 + 688.17S^2 - 2194.12S^1 - 10380.62)}{S^4 + 3.67S^3 + 301.69S^2 + 63.47S^1 + 300}$$
(11)

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In its factored form, this transfer function for the boom is:

$$G_{\rm B}(S) = \frac{-.0289 (S + 2.63) (S - 5.88) (S + 671)}{(S^2 + .2S + 1) (S^2 + 3.46S + 300)} = \frac{N_{\rm B}(S)}{D_{\rm B}(S)} \quad .$$
(12)

The block diagram for the system is shown in Figure 4. The pointing mount uses a position plus integral control with a velocity feedforward [4]. For the sake of clarity, the velocity feedforward will be considered separately from the forward transfer function of the Gimbal Pointing System which is:

$$G_{G}(S) = \frac{K_{p}(S + 1.2)}{S}$$
 (13)

The velocity feedforward and the boom comprise a loop which can be reduced to a new forward transfer function, $G_{BS}(S)$, to treat the system as a unity feedback system as seen in Figure 5. The new forward transfer function is given by:

$$G_{BS}(S) = \frac{G_B(S)}{1 + K_R S G_B(S)} = \frac{N_B(S)/D_B(S)}{1 + K_R S N_B(S)/D_B(S)}$$
(14)

$$G_{BS}(S) = \frac{N_B(S)}{D_B(S) + K_R S N_B(S)}$$
(15)

From equation (15) we can see that the open loop eigenvalues can be arbitrarily set by adjustment of the rate gain, K_R . The closed loop eigenvalues will then be set by means of the compensator, $G_C(S)$, and the forward gain, K_P .

Results

The open loop eigenvalues were set by adjustment of the rate feedforward gain, K_R . The plot of these roots (eigenvalues) as K_R is increased from zero is shown in Figure 6. At a gain between 0.1 and 0.2, the system goes unstable. At a gain of 0.07 the open loop roots are:

XXXVIII-4

$$S_{1,2} = -.14 \pm j.98$$

änd,

$$S_{3,4} = -1.23 \pm j 17.38$$
.

Yielding the new transfer function:

$$G_{BS}(S) = \frac{-.02896 (S + 2.63) (S - 5.88) (S + 671)}{(S^2 + 28S + .99) (S^2 + 2.46S + 303)}$$
(16)

The major result of this rate compensation (which is built into the pointing mount) is to increase the damping ratio of the first quadratic root and decrease the damping of the second. This gives the basic system transfer function:

$$G_{SYS}(S) = \frac{K^{1} (S + 1.2) (S + 2.63) (S - 5.88) (S + 671)}{S (S^{2} + .28S + .99) (S^{2} + 2.46S + 303)}$$
(17)

Setting $G_c(S) = 1$, we can again plot the root locus for the system as K^1 is varied ($K^1 = -.02896 K_p$). This plot is shown in Figure 7. At a gain between -0.2 and -0.03 the system goes unstable and oscillates; the system is unstable for all values of positive gain. At a gain of -0.2 the dominate roots have a time constant of about 8 seconds and would be extremely underdamped with $\xi = 0.01$ ($S_{1,2} = -0.13 \pm j$ 12.3). This is not acceptable; compensation is required.

The compensation chosen was:

$$G_{c}(S) = \frac{(S^{2} + 2.46S + 303)(S + 2)}{(S + 671)(S + 10)^{2}}$$
(18)

This compensator cancels the open loop roots at $s = -1.23 \pm j$ 17.38 and the zero at S = -671. If the compensator does not cancel the complex roots exactly it is better to undersize the real part of the compensator roots than oversize them. The response will contain terms of Ae^{-1.23} COS (17.38t + ϕ) but A will be extremely small due to the proximity of the zeroes to these roots.

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The basic idea of the compensation is to cause the roots nearest the imaginary axis (first mode) to move to the left as gain is increased while higher order modes (roots) move to the right. This allows for a high value of the gain which produces low steady state ramp errors and a wide passband. Ideally a type 2 system is desired but due to the zero in the right half s-plane, this is impossible.

Using the compensator of equation (18), a new G_{total} is obtained:

$$G_{\text{total}}(S) = \frac{K^{1} (S + 1.2) (S + 2) (S + 2.6) (S - 5.88)}{S (S^{2} + .28S + .99) (S + 10)^{2}} = \frac{K^{1}N(S)}{D(S)}$$
(19)

With unity feedback, the closed loop control ratio is:

$$\frac{C}{R} = \frac{G}{1+G} = \frac{K^{1} N(S) / D(S)}{1 + K^{1} N(S) / D(S)}$$
(20)

which reduces to

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$$\frac{C(S)}{R(S)} = \frac{K^{1} N(S)}{D(S) + K^{1} N(S)}$$
(21)

The plot of the closed loop roots is seen in Figure 8. At a gain of approximately -16, system becomes unstable. At a gian of -12, the dominant root is S = -0.75 and the ramp error coefficients is $K_1 = 4.5 \text{ sec}^{-1}$. For a ramp input of 0.01°, the steady state error for the system would be 8 arc sec.

The error response of the system to the Shuttle's thruster firing and drift due to gravity gradient is seen in Figure 9. The peak error response is 3.6 arc sec as the Shuttle drifts through its limit cycle and 6.5 arc sec when one thruster is firing. The maximum gimbal torque required to oppose the thruster acceleration is 17.9 NT-m while the maximum torque during drift is 8.3 NT-m.

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The step response of the system to a 1° step is shown in Figure 10. Within 5 sec of the step, the tip angle is within 104 percent of steady state. The peak overshoot is 42 percent occurring at 7 sec. The response is within ± 1 percent of steady state within 16 sec. Steady state error to step commands is zero. The torque required at the gimbal motors for this response exceeds the rated 33 NT-m during the first 12 sec. The torque required to initially accelerate the boom is up to 182 NT-m/degree. This limits step commands to less than 0.18 degree.

The frequency responses for the boom and rate compensated boom are seen in Figure 11. The responses of the fixed boom is a series of single order zeroes and quadratic poles starting at 1 radian/sec (.16 Hz). The boom is itself stable having an 18 dB gain margin and a 35° phase margin. The effect of the gimbal system rate compensation is to increase the damping ratio of the first mode and decrease the damping of the second. At a $K_n = 0.07$, stability is unaffected.

The frequency response of the gimbal mounted boom system is shown in Figure 12 for a gain of 16. At this gain, the system is unstable having a gain margin of -50 dB. And a phase margin of nearly -90° . For a stable system, the gain would have to be reduced drastically and would yield a slow, low bandpass system with high ramp following errors.

The frequency response of the compensated system is shown in Figure 13 for a gain of 16. At this value of gain, the system is just unstable with both a phase margin and a gain margin of zero. Reducing the gain of 2.4 dB to a gain of 12, the system becomes stable and has the closed loop frequency response of Figure 14. The system bandpass is 2.5 Hz and it will follow up to 0.2 Hz signal without error or phase shift. This is the basic reason the system can tolerate the drift of the Shuttle in its deadband and the firing of one thruster. The drift of the Shuttle is very nearly a 0.01 Hz signal while the turnaround at the deadband is like a 1 Hz signal. The system response to thruster firing is not frequency (dynamics) limited but torque limited by the gimbal motors.

Conclusions and Recommendations

As compensated, the gimbal mounted boom system is stable with fast dynamics and low following errors. The system is relatively unaffected by the drift of the Shuttle in its deadband and by one thruster firing (maximum error 6.5 arc sec). The system is limited by the torque produced by the gimbal motors not by dynamics.

This study only considered motions about the y-axis, not x-axis or z. Motion about the x-axis should be very similar to that of the yaxis except modes shapes 2 and 5 would predominate. The control system for x-axis should be nearly identical to that of the y-axis. Motion about the z-axis is twisting motion of primary the third mode (see Table 1). This control system should be much simpler than either x or y control. This study assumed perfect motion detectors for the feedback loops. Good detectors [6] are available for x, y, and z motions but their sensitivities and accuracies and sensitivities need to be added to the control simulation as disturbance.

The dynamic as well as probabilistic characteristics of each of these disturbances as well as man motion need to be characterized and incorporated into the control model to determine the pointing accuracies and pointing stability of the gimbal mounted boom. Future work should include a stochastic analysis of the control system concerning probabilistic disturbances and the minimization of their effects on pointing stability. The results of such a study would be a three-axis model of the system compensated for stability as well as compensated for disturbance induced pointing variations.

References

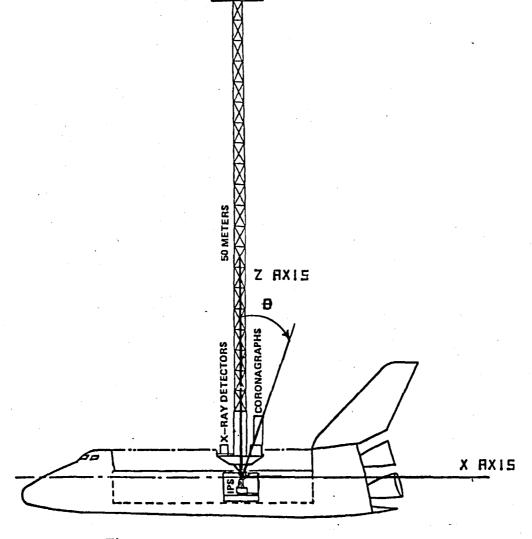
- 1. Sperry Support Services, Final Report Contract #NAS8-33588, "Sperry Interactive Graphics System (SIMS) and its Application to Control Simulation of the Pinhole X-Ray/Coronograph Experiment," 1980.
- 2. Snowdon, J. C. "Vibration and Shock in Damped Mechanical Systems," John Wiley and Sons, New York, N. Y., 1968.
- 3. Thomson, W. T. "Vibration Theory and Applications," Prentice-Hall, Englewood Cliffs, N. J., 1965.
- 4. Sperry Flight Systems, Report on Contract #NAS8-33413, "Annular Suspension Pointing System - Gimbal System Design Report - Volume I Data Package for Preliminary Design Review," Data Requirement 8, 1979.
- 5. D 'Azzo, J. J. and Houpis, C. H. "Linear Control System Analysis and Design, Conventional and Modern." McGraw-Hill, New York, N. Y., 1975.
- 6. Visidyne, Final Report Contract #NAS8-33697, "Pinhole X-Ray/ Coronograph Optical Systems Concept Definitions Study," 1980.

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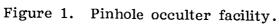
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TABLE 1. PINHOLE BOOM FREQUENCIES AND MODE SHAPES

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MASK



CONFIGURATION

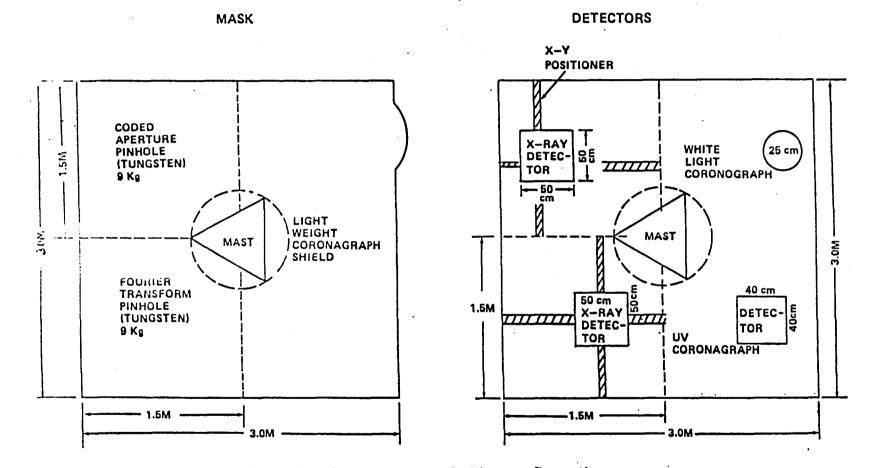
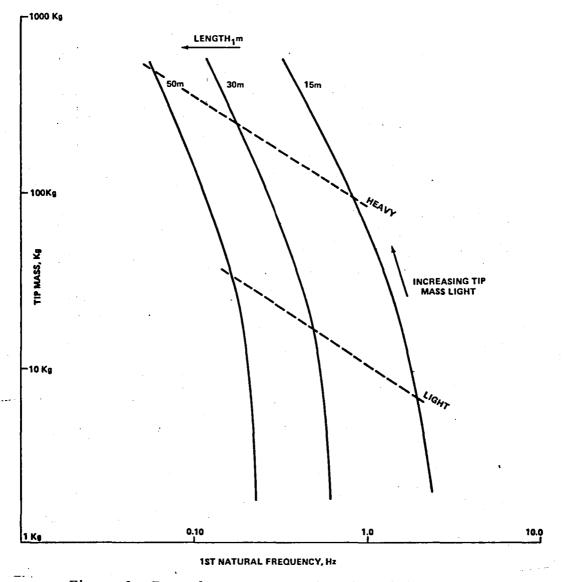
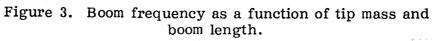


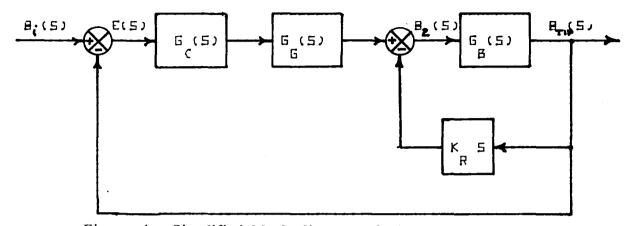
Figure 2. Pinhole occulter facility configuration.

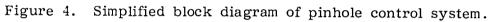
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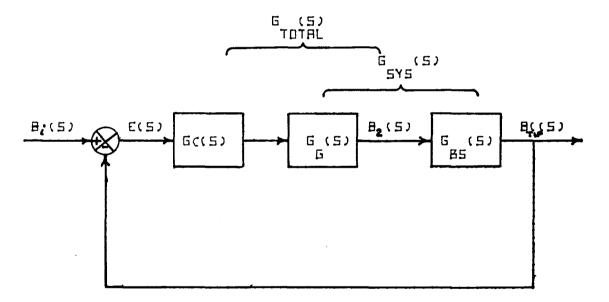
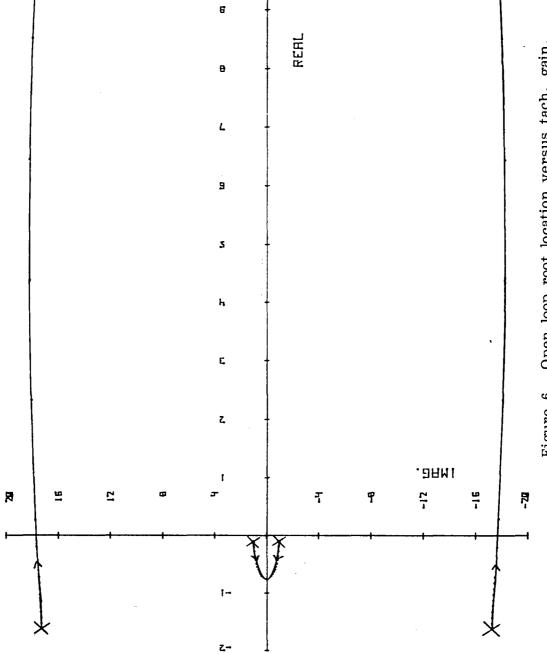


Figure 5. Reduced block diagram of pinhole control system.

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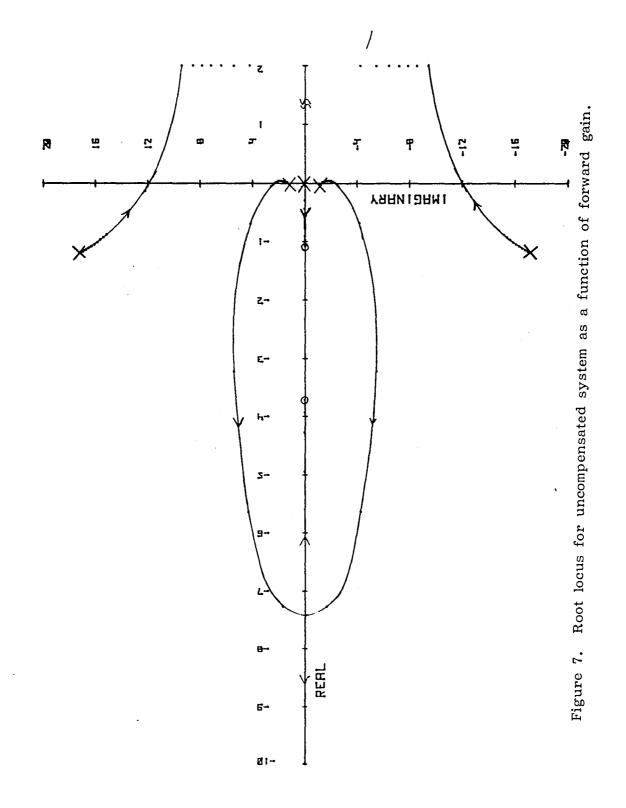
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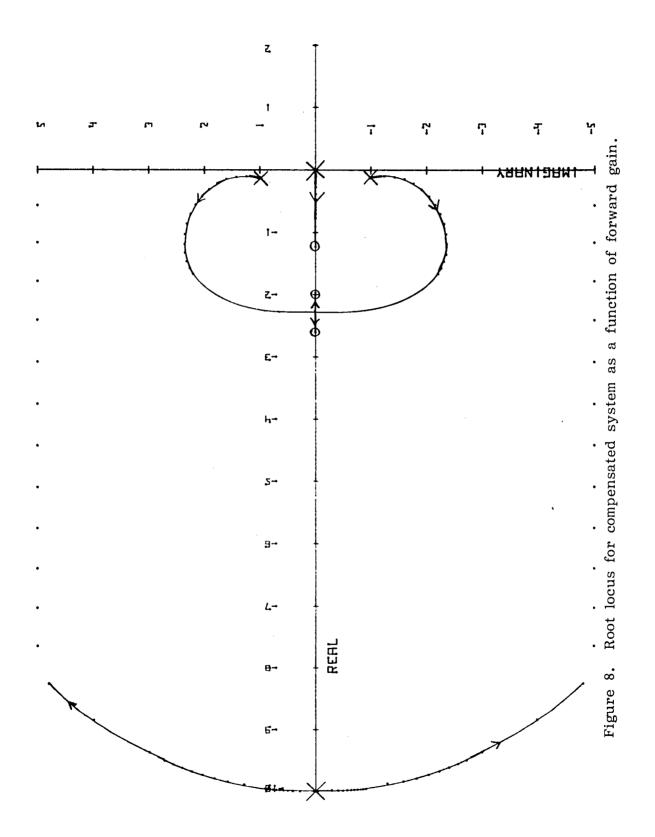


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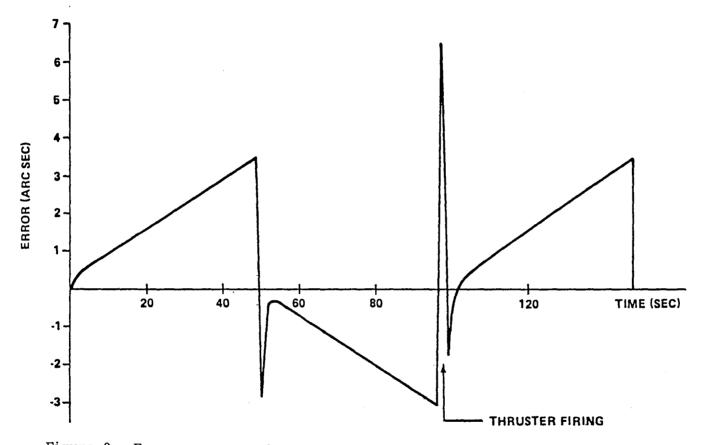
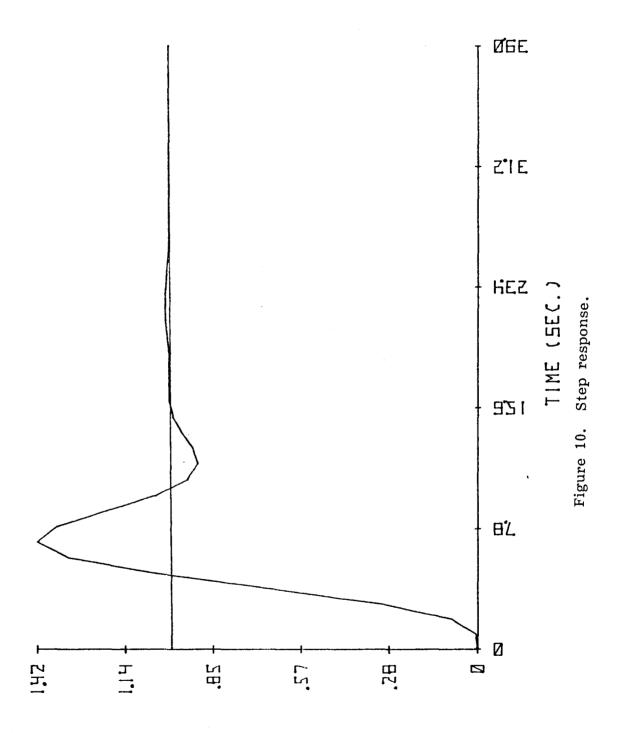
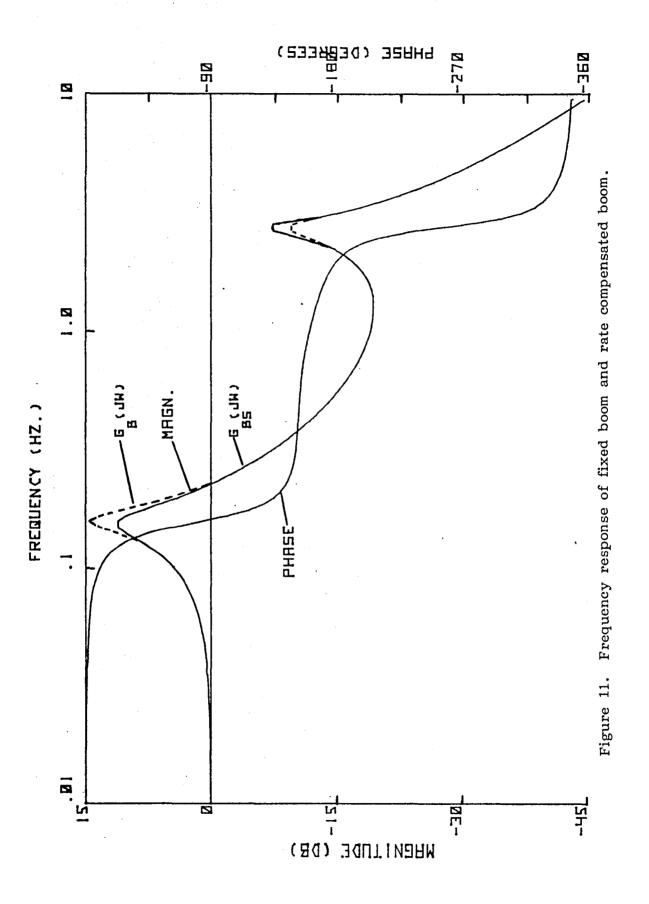


Figure 9. Error response of boom tip to thruster firing and gravity gradients.



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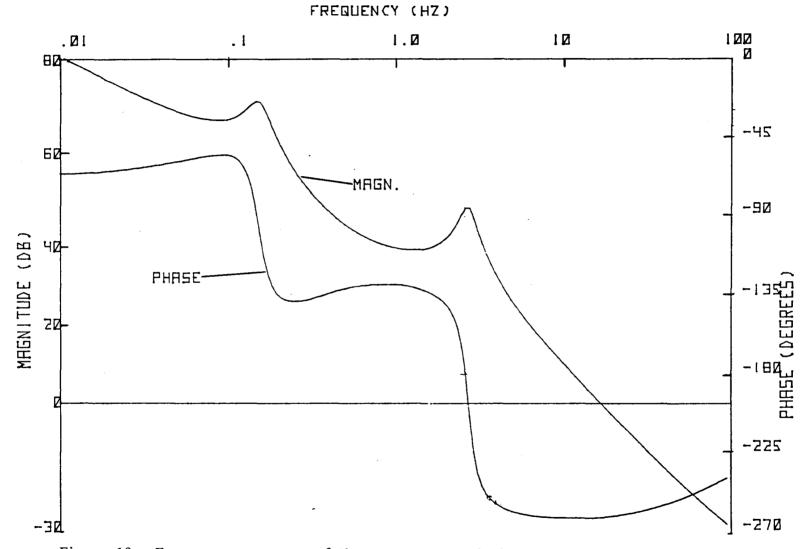
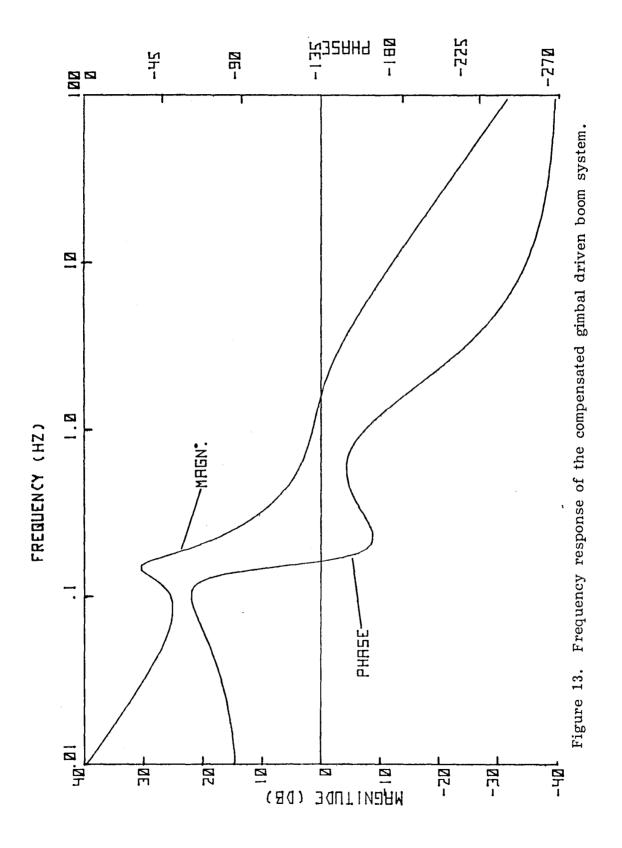
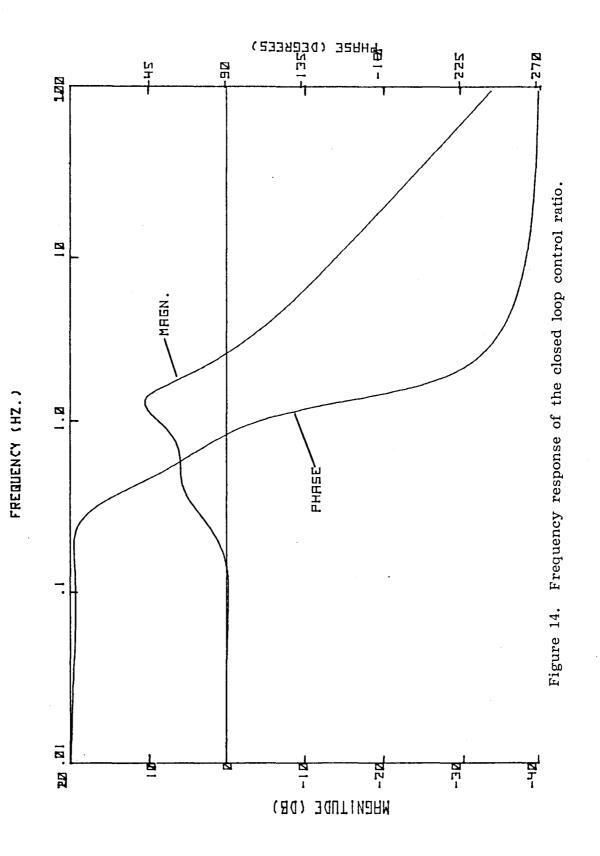


Figure 12. Frequency response of the uncompensated gimbal driven boom system.



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1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

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FFT METHODS IN SIGNAL PROCESSING OF THE COAL INTERFACE DETECTOR RADAR

Prepared By:Darko Kajfez, Ph. D.Academic Rank:ProfessorUniversity & DepartmentUniversity of Mississippi
Department of Electrical

NASA/MSFC:

(Laboratory) (Division) (Branch) MSFC Counterpart: Date: Contract No.: Communications & Tracking Billy R. Reed August 1, 1980 NGT-01-002-099 (University of Alabama)

Engineering

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FFT METHODS IN SIGNAL PROCESSING OF THE COAL INTERFACE DETECTOR RADAR

ΒY

Darko Kajfex, PH. D. Professor of Electrical Engineering University of Mississippi University, MS 38677

ABSTRACT

The FM radar for the coal interface detector, operating in the frequency band 2 to 4 GHz, is intended for the display of thicknesses between 2 cm and 20 cm. Because of such a short range, the thickness information is contained in the very few lowest spectral components of the output signal. To overcome this inconvenience, the Fourier series of the output signal has been augmented to approximate a Fourier integral. This modification in the signal processing resulted in a higher spectral density, which in turn enabled an easier identification of the interface position in the laboratory. The orientation and spacing of the receiving and transmitting antennas is found to have an important influence on the system performance. A field test in the coal mine is planned in the future.

ACKNOWLEDGEMENTS

I appreciate the opportunity of continuing the work on the challenging problem of the CID radar during my second year of NASA/ASEE Summer Faculty Research Fellowship Program. My NASA counterpart, Mr. Billy R. Reed provided an excellent support in all aspects of the project. Mr. Ivan A. Burroughs was very helpful in programming of the signal processing procedure, and Mr. Edmund H. Gleason provided the expertise in the selection and utilization of the various electronic instruments and subsystems. The typing of the report was performed by Joyce Gray.

INTRODUCTION

Coal interface detector (CID) radar is designed to measure the thickness of the coal layer in the longwall mining method. The frequency of the radar transmitter is modulated in a saw tooth fashion between 2 and 4 GHz. The thickness of the coal layer is measured by observing the frequency difference between the signals reflected from the front and rear surface of the coal. The radar output signal, which contains this difference frequency, is analyzed on a digital Fourier Analyzer, which then provides a display of the reflection amplitude vs. the coal thickness.

The basic principle of the CID radar has been under investigation at the MSFC for several years. The difficulty in implementing the radar lies in the presence of a multitude of erroneous signals, which can be misinterpreted as the rear surface return signals. Also, the frequency of the reflected signal is only several times (<10) higher than the modulation frequency. Therefore, it is difficult to determine which of the individual spectral components belongs to the position of the rear surface.

The presence of the multiple reflections may be somewhat reduced by using long cable sections between the individual components of the system, and by placing absorbing materials around the antennas. The second difficulty of distinguishing between the inidvidual components in the spectrum can be greatly reduced by processing the time-domain signal before taking the Fast Fourier Transform. Examples of measured wave-forms and their spectra are included in the report.

OBJECTIVES

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This report describes the work performed in the second period of the NASA/ASEE Summer Faculty Research Fellowship Program. The radar configuration is somewhat different from the one described in the last year's report (Reference (1)). The difference lies in the fact that a direct detector receiver has been utilized instead of the mixer-type receiver. The objective of the investigation described in the present report is to achieve the radar configuration which is suitable for application not only in the laboratory, but in the actual coal mine situation. Much of the effort is directed toward determining the most suitable signal processing procedure which produces a clear display of the coal thickness.

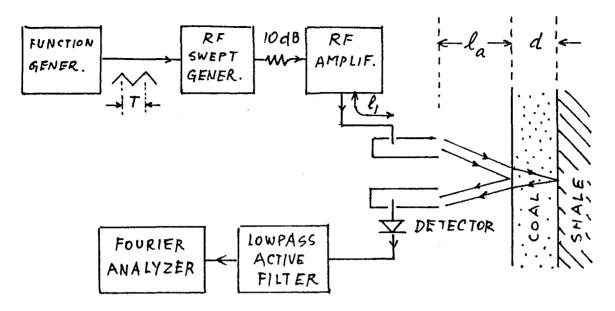


Fig. 1. Block diagram of the CID radar using the detector receiver

The transmitting part of the CID radar in Fig. 1 consists of the Hewlett Packard Function Generator Model 3312, the Wiltron Sweep Generator Model 610, the Avantek Transistor Amplifier Model 406-3M, and the Scientific Atlanta Ridged Horn Antenna Model 40M-2.0. The length of the coaxial cable between the amplifier and the antenna is $\mathcal{L}_{i} = |\mathbf{m}|$. The receiving part of the radar consists of an identical antenna, as in the transmitting part, and a Narda Detector Mount Model 4503. The detected signal is amplified and filtered in the Ithaco Dual Filter Model 4302. The filter is set for pulse operation (Bessel type response) in the lowpass mode.

The operating data are as follows:

Frequency range of the sweep oscillator -----2 to 4 GHz. Sweep period----- T = 25.6 ms RF power delivered to transmitting antenna-----138 mW Lowpass filter cutoff frequency ------3.15 KHz (two sections in cascade, each with 10 dB gain) Distance between the antenna and the coal -----1 = 60 cm

As indicated in Fig. 1, the received signal consists of two parts: One part comes from the wave which was reflected from the front surface of the coal, and the other part is a weaker signal from the wave which penetrated the coal and was then reflected from the rear interface. If the coal dielectric constant is denoted by \pounds r the sweep period by T, the coal layer thickness by d, and the RF deviation by B, the frequency difference between the front and rear reflection is

$$f_a = \frac{4Bd\sqrt{\mathcal{E}_r}}{cT}$$
(1)

c is the velocity of light in free space. For the data listed above, a coal layer of thickness d =6" (15.24 cm) and the dielectric constant of $\mathcal{E}r=5$ would produce a difference frequency approximately 355 Hz.

When such a mixture of the two microwave signals is applied to the square law detector, their difference frequency (here 355Hz) is obtained. This signal is interrupted by the repetition rate of the sweep generator. It is afterwards amplfied and filtered before being delivered to the Fourier Analyzer Hewlett Packard Model 5451B, which performs the signal processing and displays the result.

The radar system described above is similar in concept to the radar described in Reference (2). The main difference of the present system is the use of a direct detector system instead of the heterodyning receiver, and the use of a digital signal processing system instead of an analog one.

SIGNAL PROCESSING

The tests in the laboratory are performed on slabs of "artificial coal". These are slabs of different thicknesses, made from a mixture of dielectric and conducting materials, so that the dielectric constant and the loss tangent approximate the values of the actual coal from the mine. In the particular measurement to be described, the slab thickness 15.5 cm (6") and of the cross section 61x61 cm (2'x2') was backed by an aluminum foil for better rear surface reflection. The received signal at the input to the Fourier Analyzer is shown in Fig. 2.

The observed signal is periodic with the period T= 25.6 ms. The sawtooth waveform which modulates the transmitter is shown in Fig. 3. The characteristics of the transmitter are such that the modulation voltage 2.4 V produces an output of 2 GHz and 4.8 V produces an output frequency of 4 GHz. It is seen from Figures 2 and 3 that the detected signal displays an even symmetry within each half period. This is to be expected, because the transmitted frequency attains the same value twice within each period ; once on the upswing, and once on the downswing of the modulation signal.

The signal shown in Fig. 2 has been sampled with the time intervals 50 μ s, the total number of sample points being 1024. The waveform shown includes two full periods of the signal. If now only one period is retained (first 512 points) and the Fast Fourier Transform of this waveform is computed, the resulting spectrum is such as shown in Fig. 4. Each spectral component makes one point in the plot. Due to an automatic plotting procedure, the points are interconnected by straight lines, creating the appearance of a continuous spectrum. The spectrum actually consists of discrete frequencies, separated from each other. by the repetition frequency 39 Hz. According to (1), the beat frequency for the 6" artificial coal is 355 Hz. From Fig. 4 it can be seen that the two longest spectral components are components number one and number eight. Thus, by neglecting the first component which belongs to the repetition frequency, the frequency of the return signal is estimated to be eight times the repetition frequency, or 312 Hz (instead of 355 Hz).

There are several inconveniences associated with the display such as in Fig. 4. In the first place, the repetition frequency usually comes out to be the strongest component, which fact makes it difficult to read the thickness of the thin coal layers, where spectral components are very close to the repetition frequency. Furthermore, the components **immediately** next to the observed return frequency often have a small amplitude, so that the signal is not always readily identifyable as in Fig. 4, where the seventh, eighth, and ninth spectral components are all large. It was decided that a difficulty in locating the return signal is caused by the fact that the spectral components are too far apart

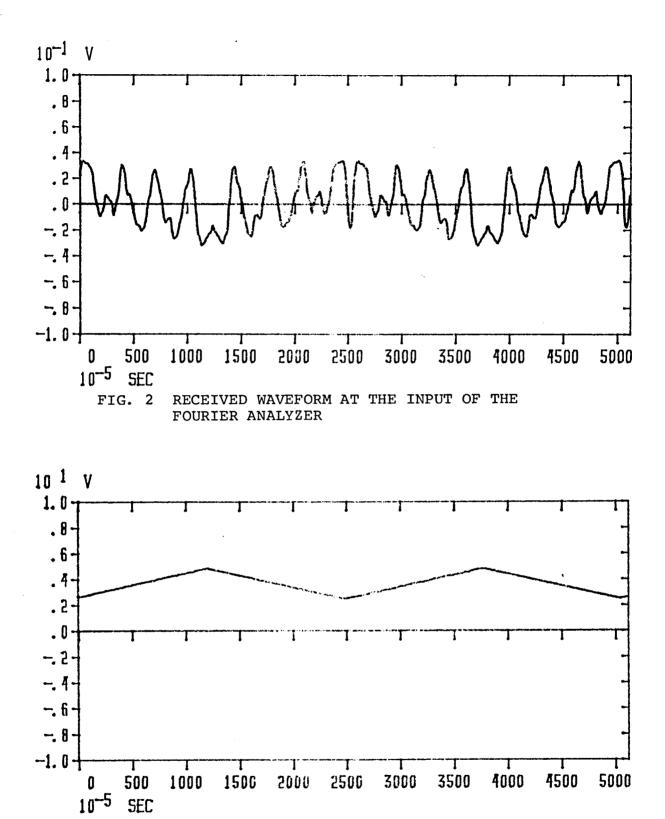
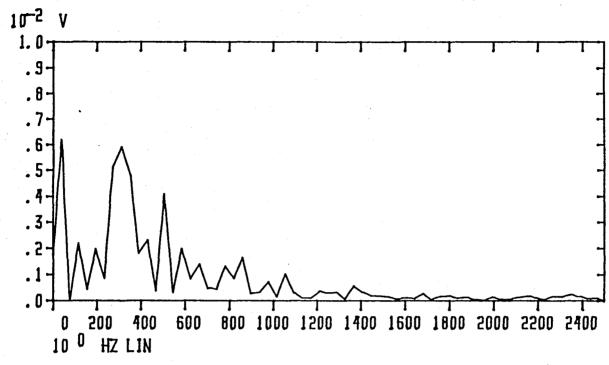
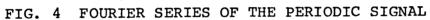
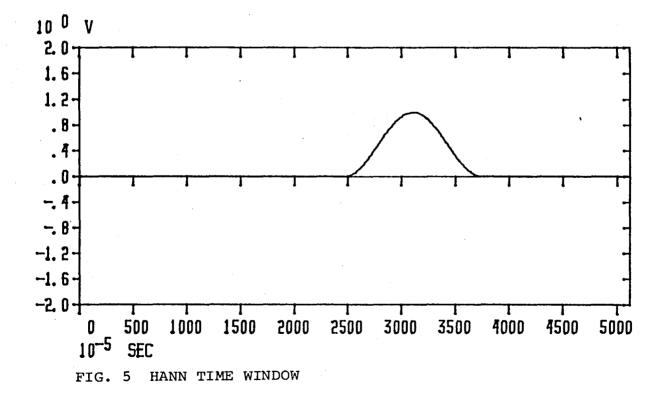


FIG. 3 MODULATION SIGNAL AT THE OUTPUT OF THE FUNCTION GENERATOR

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from each other. In order to create a more densely populated spectrum, advantage has been taken of the digital signal processing in the following manner.

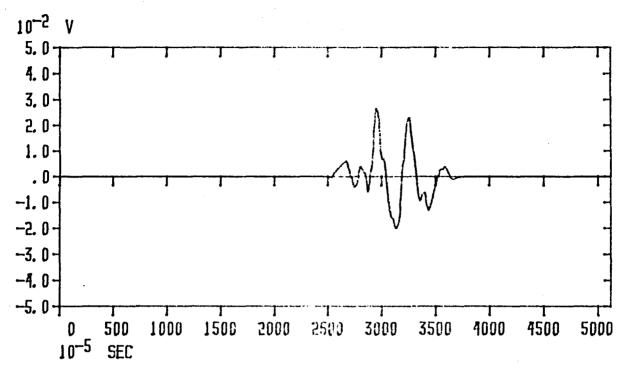
The original sampled signal from Fig. 2 consists of 1024 points. The sample extends over two periods, each period has a half-period symmetry. Therefore, the basic information is contained within 256 points. Thus, a time window is created, of the length 256 sample intervals, as shown in Fig. 5. This is the so-called Hann window (3) of the shape 1 -COS x. When the sampled signal is multiplied with the window, the result looks very much like a single burst of oscillations shown in Fig. 6. The total sample length is still 1024 points, but 75% of the sample points are forced to be zero. The Fast Fourier Transform performed on these 1024 sample points will have much larger spectral dnesity from the spectrum in Fig. 4, and thus will allow more accurate reading of the coal thickness.

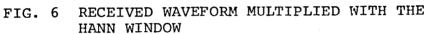
Before the spectrum is displayed, another operation is performed. The radar return signal is redundant in the sense that the half period which corresponds to the modulation upswing is a mirror image of the other half period which corresponds to the modulation downswing. Therefore, another Hann window is created, shifted for 256 points. Again the sampled signal is multiplied with this second window, and the Fast Fourier Transform is completed. The resulting spectrum is then multiplied with the previously evaluated spectrum. The procedure is equivalent to taking the convolution of the two half periods in the time domain. This insures that the information contained in each half period contributes equally to the final result. The amplitude of the resulting spectrum is then displayed in a logarithmic scale, as shown in Fig. 7.

The spectrum in Fig. 7 is dense enough for each strong spectral component to be easily identified. When comparing this spectrum with the one in Fig. 4, it is noticed that now the return signal has the largest amplitude, and is thus easily recognized. It appears that the amplitude of this return is about 20 dB larger than other spectral components on either side. However, it should be remembered that the spectrum in Fig. 7 has been obtained by multiplying the two spectra emanating from each of the two half periods, so that the resulting product is made of the squared amplitudes.

Because of this, the number of decibels in Fig. 7 must be divided by two, so that the actual difference between the largest lobe and the next largest lobe is only about 10 dB, and not 20 dB.

The frequency corresponding to the rear surface reflection in Fig. 7 is 310 Hz, instead of the predicted value 355 Hz. The discrepancy may have been caused by an inaccurate knowledge of the frequency deviation B and the dielectric constant \mathcal{E}_{r} which both enter into equation(1).





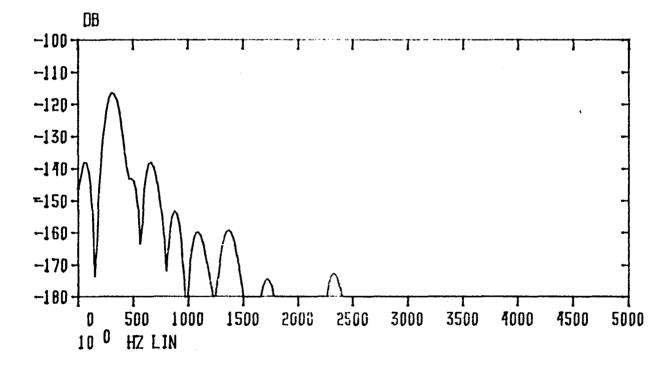


FIG. 7 FOURIER TRANSFORM DISPLAY

When a sinusoidal signal is interrupted by the square pulses, the specturm of such a modulated signal contains not only the carrier frequency, but also a number of sidelobes which are decreasing in amplitude according to sin x/x function. The first sidelobe of this function is 13.5 dB smaller than the maximum. The radar return signal is switched each time the modulation waveform changes its slope. Therefore, the observed spectrum will contain considerable sidelobes on each side of the carrier frequency (355 Hz in the example discussed above. These sidelobes may be misinterpreted as apparent radar returns. One way of reducing these sidelobes is to use the Hann window, such as shown in Fig. 5. The sidelobes of the sinusoidal signal modulated with the Hann window are 37 dB smaller than the maximum. This is the reason that the sampled signal in Fig. 6 has been multiplied by the Hann window.

One unavoidable consequence of reducing the sidelobes by using the Hann window is the doubling of the spectral width of the main lobe, when compared with the rectangular window (4) should two close surfaces be distinguished in the radar display, their difference in frequency should be at least half width of the main lobe. For the sinusoidal signal multiplied with the Hann window, the first zero appears at double the value of the sin x/x function:

$$\widehat{T} \Delta f \frac{T}{2} = 2 \widehat{I}$$

The frequency difference between the main lobe maximum and the first zero is

$$\Delta f = \frac{4}{T} \tag{2}$$

If two signals are reflected from two very closely located surfaces their frequencies may differ for less than Δf , and cannot be distinguished on the display. Substituting (2) into (1), we find the minimum resolution distance between the two reflections in the coal

$$d_{res} = \frac{2c}{B\sqrt{\epsilon_r}} \qquad (3)$$

The resolution distance is independent of the repetition frequency. The only way to obtain a finer resolution is to increase the frequency deviation For the system with B = 2 GHz, the resolution distance is $d_{res} = 6.7$ cm.

When two reflection surfaces are dloser to each other than d_{res} , their main spectral lobes will merge into a single lobe, and the two returns will not be distinguishable on the radar display. One way to reduce the resolution distance is to use the rectangular window instead of the Hann window. This would split the resolution distance in half at the expense of increasing the sidelobes. It follows from the previous discussion that large sidelobes are undesirable. Therefore, the fact that the frequency deviation of the radar is 2 GHz sets on a priori limit of 6.7 cm as the smallest recognizable coal thickness (3.3 cm with the rectangular window).

ANTENNA CONSIDERATIONS

Little is known about the design of microwave antennas suitable for sensing the material properties in close range. The majority of microwave antennas are designed to produce a given electromagnetic radiation in the far field region. In the early experiments with the CID radar, the transmitting and receiving antennas have sometimes been placed less than one wavelength apart from each other or from the coal interface. It has been empirically found afterwards that these distances must be increased to several wavelengths.

One important property of the antenna to be used for the CID radar is low dispersion. The time delay for the signal, to pass from the antenna terminal to an obstacle in front of the antenna and back to the terminal, should be independent of frequency. One of the popular wideband antennas, the log periodic array, is ill suited for the task, as pointed out in [2].

Another important property of the pair of antennas to be used in the radar is a high decoupling when they are placed side by side. By looking at Figurel, it becomes obvious that a direct propagation from the transmitting to the receiving antenna creates a third wave impinging upon the detector. The difference frequency of that wave and the front surface reflection may produce an erroneous "return" signal on the radar display.

Without making any theoretical investigation, and without even comparing the performance of various commercially available antennas, the pair of ridged horn antennas has been assigned to do the job. The orientation and the distance between them has been experimentally adjusted to give the best performance as follows.

Figure 8 depicts two possible relative positions of the receiving and transmitting antennas, denoted parallel and collinear. It has been found experimentally that the collinear orientation offers higher isolation between the antennas. The distance D has also been varied, and the value D=25 cm appears to be sufficient. The measurement has been conducted when the antenna-to-coal distance is $l_{a1} = l_{a2} = 60$ cm.

Another unexpected critical adjustment proved to be the angle 0 in figure 9. This angle measures the departure of the coal surface from making a right angle with the antenna system. It has been found experimentally that the return signal from the rear surface drops considerably when the difference $l_{al} - l_{a2}$ becomes larger than 2 cm. For D = 24 cm and b = 10 cm, this implies that the angle between the coal surface and the antenna boresight should not depart from the right angle for more than $\theta = 3.3$ degrees.

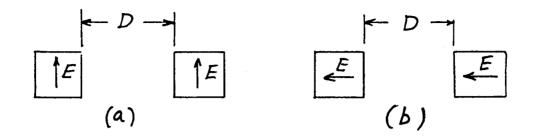


Fig. 8. POLARIZATION OF ANTENNAS: (a) PARA LLEL (b) COLLINEAR

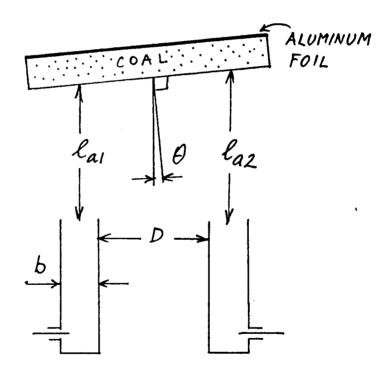


FIG. 9. INCLINATION OF THE COAL SURFACE WITH RESPECT TO THE ANTENNA SYSTEM

CONCLUSIONS

The CID radar system which has been developed has certain advantages over the previously utilized systems. The digital processing of the signal on a Fourier Analyzer System enables more convenient display which seems to improve the separation between the actual return and other unintentionally created signals. When also proper care is exercised in antenna positioning, the system seems to be capable of giving meaningful real-time results in a coal mine environment.

The improvement in the resolution distance may be achieved only by increasing the frequency deviation of the system. Further improvements on the antenna system may also bring an enhancement of the desired return signal and a reduction of the undesired signals.

REFERENCES

- 1. Kajfez, D., "Digital Processing of the Spectrum Emanating from the Coal Interface Detector (CID) Radar," University of Alabama in Huntsville, Report on NASA/ASEE Summer Faculty Research Fellowship Program, Huntsville, 1979.
- 2. Owen, T. E., Tranbarger, O., "Vol. II Investigation and Development of a High Resolution FM-CW Radar System for Residual Coal Thickness Measurements," Southwest Research Institute, San Antonio, Texas, June 1, 1977.
- 3. *******, "Fourier Analyzer Training Manual," Hewlett Packard Application Note 140-0.
- 4. Beauchamp, K. G., "Signal Processing, New York: John Wiley, 1973, p. 368.

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EXCHANGE OF STABILITY FOR A COLUMN OF FLUID WITH VARIABLE RAYLEIGH NUMBER UNDER FREE BOUNDARY CONDITIONS

Prepared By: Letitia Korbly, Ph.D. Academic Rank: University of Alabama in University and Department: Birmingham NASA/MSFC: (Laboratory) Space Sciences (Division) Atmospheric Sciences (Branch) Fluid Dynamics MSFC Counterpart: George H. Ficht1 August 15, 1980 Date: NGT-01-002-099 Contract No.: (University of Alabama)

EXCHANGE OF STABILITY FOR A COLUMN OF FLUID WITH VARIABLE RAYLEIGH NUMBER UNDER FREE BOUNDARY CONDITIONS

by

Dr. Letitita Korbly Department of Mathematics University of Alabama in Birmingham Birmingham, Alabama

ABSTRACT

Motion starts in a column of fluid subject to an adverse temperature gradient when the gravitational forces are large enough to overcome those of viscosity. This point is governed by the critical (smallest) Rayleigh number.

When the column is heated rapidly from below and not well mixed the Rayleigh number varies with height.

For such a system with harmonic boundary conditions the principle of exchange of stabilities is shown. Attempts to show this principle under rigid and semirigid boundary conditions failed.

The Rayleigh number was assumed to follow the nondimensional distribution $\Re(z) = \Re\left(\frac{2+2}{z_o}\right)^{-4}$, $\alpha = 4/3$ or $\alpha - 3/2$, $z_o = 10^{-5^-}$, $0 \le z \le 1$. The lowest values for at which there would be convection was estimated at 8.512 x 10⁸ for $\alpha = 4/3$ and 4.7023 x 10⁹ for $\alpha = 3/2$. The eigenfunction was approximated in both cases.

This work has application to both columns of air in the Earth's atmosphere and the GFFC experiment to be flown in Spacelab 3.

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INTRODUCTION

The Navier Stokes equations have classically been used to describe the motion of fluid in the presence of temperature gradients.

In the Benard problem, the fluid is assumed to be incompressible and at rest; it is assumed to be uniformly heated from below. The vertical temperature distribution is assumed to be linear. The equations are then linearized by looking at temperature and velocity perturbations.

The equations are then reduced to a single sixth order linear partial differential equation with constant coefficients in the vertical velocity perturbation ω . Fourier transform in the horizontal space variables has reduced the number of independent variables. Only time, t, and height, z, remain. The equation is then transformed in t and non-dimensionalized. It becomes:

$$\left[(D^{2} a^{2}) (D^{2} - a^{2} - \sigma) (D^{2} - a^{2} - \rho \sigma) + Ra^{2} \right] W(2) = 0 \qquad 0 \le 2 \le 1$$

Here the vertical velocity perturbation is

$$w = e^{-t + i(a_1 x + a_2 y)} \setminus \sqrt{(2)}$$

Here DW = W and $a^2 = a_1^2 + a_2^2$.

The principle of exchange of stabilities holds if convection modes which neither grow nor decay with time are time independent. If **R**, the Rayleigh number, and **p**, the Prandl number, are constant, exchange of stabilities can be shown to hold under both free and rigid boundary conditions.

If **R** is allowed to vary with height, what will happen? We were able to show the principle of exchange of stabilities holds under free, but not rigid or semirigid, boundary conditions.

Having this result, the onset of convection was investigated for certain distribution of $\mathbf{\hat{k}}$ which model what occurs in the summer atmosphere below the inversion. The values for $\mathbf{\hat{k}}$ used were

 $\mathcal{R}(z) = \mathcal{R}\left(\frac{z}{z_0}\right)^{-\alpha}$ Zo= 2 = d , x=4/302 d=3/2.

Here d is the height of the inversion, 2_0 is surface roughness length, and \propto is empirically chose **n**. This is nondimensionalized, and the critical (smallest) value of R for which convection takes place is estimated.

The results of the calculations were checked by determining the temperature gradient at the surface, $\frac{\partial \Theta}{\partial z}\Big|_{z=z_0} = \beta_0$. This determined $\frac{\partial \Theta}{\partial z} = \beta_0 z^{-4/3}$ which was then integrated to give a temperature drop of roughly 15° in the first meter above the surface. This figure checks reasonably well with the data.

EXCHANGE OF STABILITIES

These calculations show that the principle of exchange of stability holds for a column of fluid heated from below if the Rayleigh number is variable when the boundary conditions are harmonic. The Boussinesq approximations are used. The transverse velocities, u and v, the pressure, p, and the temperature perturbation, θ , are eliminated. The nondimensional equation is

$$\nabla^2 (\mathcal{O}_t - \nabla^2) (p \mathcal{O}_t - \nabla^2) \omega = -\nabla_h R \omega$$

Here $\nabla^2 \omega = \omega_{xx} + \omega_{yy} + \omega_{zz}$; $\nabla_h^{t} \omega = \omega_{xx} + \omega_{yy}$ $\rho = \gamma/R$ is the Prandtl number, and the Rayleigh number, $R = g \propto \beta d^{1/(rk)}$, depends on the height, z.

Proceeding with the analysis into normal modes, set

$$w = e^{\sigma t + i(a_1 \times a_2 y)} W(z)$$

The operator then becomes:

$$\approx (D^2 - a^2)(D^2 - a^2 - pc)W = -Ra^2W$$

Here $a^2 = a_1^2 + a_2^2$ and DW = W.

The harmonic boundary conditions are:

$$\begin{array}{l}
W(0) = 0 = V(1) \\
W''(0) = 0 = V''(1) \\
V'''(0) = 0 = V''(1)
\end{array}$$

Multiplying (*) by \sqrt{V} and integrating forms the following equation.

The following integral occurs in the analysis of (**):

$$I_1 = -S'(D^2 - a^2)WWdz = S'(DWl^2dz + a^2)Wl^2dz$$

The boundary conditions, W(0) = 0 = W(1), make the integrated terms vanish. Set $S_0' |W|^2 dz = ||W||^2$ and $S_0' |DW|^2 dz + \alpha^2 S_0' |W|^2 dz = ||\nabla V/||^2$

Both ||W|| and $||\nabla W||$ are strictly positive unless $W \equiv 0$ under the boundary conditions. So $I_1 = ||\nabla W||^2$ is strictly positive unless N = 0

Again the integrated terms vanish because W(o) = o = W(l)and W''(o) = 0 = W''(l). (The same result can be obtained for this integral under rigid boundary conditions.)

Finally

$$\begin{aligned}
\mathbf{I}_{3} &= -\int_{0}^{1} (D^{2} - a^{2})^{3} W \overline{W} dz \\
&= \int_{0}^{1} (D^{2} - a^{2})^{2} D W D \overline{W} + a^{2} (D^{2} - a^{2})^{3} W \overline{W} dz \\
&= -\int_{0}^{1} (D^{2} - a^{2})^{2} W (D^{2} - a^{2}) \overline{W} dz
\end{aligned}$$

Again the integrated terms vanish under both free and rigid boundary conditions.

$$I_{3} = S_{0}^{1} |(D^{2} - a^{2}) DW|^{2} dz + a^{2} S_{0}^{1} |(D^{2} - a^{2}) W|^{2} dz + (D^{2} - a^{2}) W|^$$

Under free boundary conditions the integrated terms vanish so

$$I_{3} = \int_{0}^{1} |(D^{2} - a^{2})DW|^{2} dz + a^{2} \int_{0}^{1} |(D^{2} - a^{2})W|^{2} dz = ||\nabla (D^{2} - a^{2})W||^{2}$$

When (**) is multiplied by -1, it becomes

$$-\int_{0}^{1} (D^{2} - a^{2})^{3}W \overline{W} dz + (p+1)\nabla \int_{0}^{1} (D^{2} - a^{2})^{2}W \overline{W} dz - p\sigma^{2} \int_{0}^{1} (D^{2} - a^{2})W \overline{W} dz$$

$$= a^{2} \int_{0}^{1} R |w|^{2} dz$$

Set $J = \int_{0}^{1} R |w|^{2} dz$. If I_{1}, I_{2}, I_{3} and J are
substituted in the equation above it becomes

$$p \underline{I}_{1} \sigma^{2} + (p+1) \underline{I}_{2} \sigma + \underline{I}_{3} - J = 0.$$

Since \mathcal{P} , the Prandtl number, is always positive, and \mathcal{I} , is positive unless $\omega=0$, the quadratic formula can be applied to solve for σ . This gives

$$\sigma = \left[-(p+1)I_2 = V(p+1)^2 I_2^2 - 4pI_1(I_3 - J) \right] / pI_1.$$

To show that exchange of stabilities holds, we must show that if $\mathcal{R}_{e} \circ = \mathcal{O}$, then $\mathcal{I}_{m} \circ = \mathcal{O}$.

There are two cases where $Re \sigma = O$

(i) $4 \neq J$, $(I_3 - J) = 0$ and (ii) $(p+1)I_3 = 0$ and $4 \neq J$, $(I_3 - J) < 0$. If (i) holds $\sigma = 0$, if (ii) holds, $\overline{I_3} = 0$ so $\int_0^1 |(D^2 - a^2) W|^2 dz = 0$. S implies

This implies

$$(0^2 - \alpha) W \equiv 0$$

The last implies that

$$\underline{T}_{1} = -\int_{0}^{1} (D^{2} - a^{2}) W \overline{W} dz = 0$$

In both cases $\mathcal{R}_{e} \sigma \circ o$ implies $\sigma \circ o$. This completes the proof of the principal of exchanges of stabilities for operators whose Rayleigh number is variable under free boundary conditions.

References

Chandrasekhar, S. "Hydrodynamic and Hydromagnetic Stability" Oxford University Press, 1961.

ANALYSIS INTO NORMAL MODES

Since the principle of exchange of stabilities holds for the variable Rayleigh number problme, we attempt to solve the steady state problem for certain functions, \mathbf{K}

 $\mathcal{R}(z) = \mathcal{R}\left(\frac{z}{z_0}\right)^{-\alpha}$ 20210-5, x=1/3 or x= 3/2, The equation is $\left(\left(\nabla^{2} \right)^{3} - R \left(\frac{z}{z_{o}} \right)^{-} \nabla_{\mu}^{t} \right] u = 0$ Z. = Z = d. with the harmonic boundary conditions 4(0) = 0 = 4(1) $\mu^{*}(0) = 0 = \mu^{*}(1)$ $\mu^{(m)}(n) = n = \mu^{(m)}(1)$ Here $\nabla^2 u = u_{xx} + u_{yy} + u_{zz}$ and $\nabla^2 u = u_{xx} + u_{yy}$ If we set $y = z - z_{\bullet}$, the equation becomes $(\nabla^2)^3 u = R(2 + z_{\bullet})^{-\alpha} \nabla_u^2 u_{\bullet}$ 0 = z = d - z .. with harmonic boundary conditions at 0 and d-z. The equation is normalized by setting x=x/(d-2.) y = y/(d - 20) 2 = 3/(d - 20) and becomes $(\nabla^2)^3 u = R[z_0^{\alpha}/(d-z_0)^4] (z+z_0)^{-\alpha} \nabla_h^2 u, \quad 0 \le z \le l_0$ with harmonic boundary conditions at 0 and 1. Let $R_{o} = R z_{o}^{4} / (d - z_{o})^{4}$ The equation reduces to $(\nabla^2)^3 u = R_0 (z + z_0)^{-2} \nabla_h^2 u$ with $a^2 = a_1^2 + a_2^2$, the equation reduces to the ordinary differential equation: $(D^{2}-a^{2})^{3}u = -R_{0}(z+z_{0})^{-a}a^{2}u$

If $A_n \cdot o$, $n \ge 2$, we can look for a first approximation to the critical (smallest) R_o for which the equation can be satisfied. This will be approximated by finding the smallest 'R' such that $0 = \int [(0^2 - a^2)^3 + (R_0(2 + 2_0)^{-\alpha} a^2] \sin \pi 2 \sin \pi 2 dz$ That is (12+a2) 3 So sin 12 dz = 'Roa' So (2+20) sin 2 112 dz I " = So' (2+20) - sin 2 12 dz Set S' sin 2 17 2 = 1/2. and note that This yields $R_{0} = (\pi^{2} + a^{2})^{3} / (2 a^{2} I_{1,1}^{*}) = (1/(2 I_{1,1}^{*})) (\pi^{2} + a^{2})^{3} / a^{2}$ It is easy to see that ' R_0 has a minimum when ' $R_0' = 0$, that is, when $\alpha^2 = \pi^2/2$. Then 'Ro = 27 17 1/(8 I 1.) Since numerical integration yields the values $\pm \frac{4/3}{1} = 1.760401$ $I_{1,1}^{3/2} = 2.157553.$ and The first estimate, 'R, of R is $'R_{\rm b} = 186.75045^{\circ}$, a = 4/3, if a = 3/2. and 'R = 152.3743 , if The first estimate of R, then is x = 4/3, 8.667841×108, if and $\alpha = 3/2$ 4. 81 3 3 0 65 × 109 if

The second approximation is made by assuming $A_n \circ D, n \ge 3$.

The problem is to find the smallest number, ${}^{z}R_{o}$, and a pair of numbers A_{1} and A_{2} , with

$$O = S_{0}^{1} [(D^{2} - a^{2})^{3} + {}^{*}R_{0} (z + z_{0})^{-1} a^{2}] (A_{1} \sin \pi z + A_{2} \sin 2\pi z) \sin \pi z dz$$

and

$$0 = S_{1} [(0^{2} - a^{2})^{3} + {}^{2}R_{0}(2+2_{0})^{-1}a^{2}] (A_{1} \sin \pi 2 + A_{2} \sin 2\pi 2) \sin 2\pi 2 dz.$$

.

That is

$$A_{1}(\pi^{3}+a^{2})^{3}\left[\int_{0}^{t}\sin^{3}\pi z\,dz\right] = a^{3}R_{0}\left[A_{1}\int_{0}^{t}(2+z_{0})^{-d}\sin^{3}\pi z\,dz + A_{2}\int_{0}^{t}(2+z_{0})^{-d}\sin^{2}\pi z\,dz\right]$$

and

$$A_{2}(\pi^{2}+a^{2})\left[S_{sin}^{2} a_{\pi 2}d_{z}\right] = a^{2} CR_{s}\left[A_{1}S_{s}(z+z_{s})^{-4}sin\pi z sina\pi z d_{z} + A_{2}S_{s}(z+z_{s})^{-4}sin^{2}a_{\pi 2}d_{z}\right]$$

Set

$$I_{j,K}^{d} = S_{o}(2+z_{o})^{-d} \sin j\pi z \sin k\pi z dz$$

Note that $I_{j,K}^{\alpha} = I_{K,j}^{\alpha}$ and that $S_0^{\prime} \sin^2 j \pi z dz = \sqrt{2}$. This means that

$$A_{1}(\pi^{2}+a^{2})^{3}/2 = a^{2} R_{0}[A_{1}] I_{1,1}^{4} + A_{2}I_{1,2}^{4}]$$

and

$$A_{2}(4\pi^{2}+a^{2})^{3}/2 = a^{2} R_{1} [A_{1} I_{1,2}^{*} + A_{2} I_{2,2}^{*}].$$

The numbers $\underline{I}_{\mathbf{j},\mathbf{k}}$ are found to any desired degree of accuracy by Simpson's rule.

The problem can be written now as an eigenvalue problem in \not/ R_{\bullet} with eigenvector (A_1, A_2) .

$$(1/2R_{o}) A_{1} = 2a^{2}/(1^{2}+a^{2})^{3} [I_{1,1}^{a}A_{1} + I_{1,2}^{a}A_{2}]$$

$$(1/2R_{o}) A_{2} = 2a^{2}/(4\pi^{2}+a^{2})^{3} [I_{1,2}^{a}A_{1} + I_{2,2}^{a}A_{2}]$$

At the value of a^2 corresponding to the smallest value of R₀, the function 1/2R₀ takes on its largest value. This will be the maximum of the largest eigenvalue of all the matrices,

$$M = \begin{pmatrix} m_{11} & m_{12} \\ m_{21} & m_{22} \end{pmatrix}$$
$$M_{j_1k} = (2a^2/(j^2\pi^2+a^2)^2) I_{j_1k}^{k}$$

where

M is not symetric.

It is convenient to set
$$b = \alpha^{3}/\pi^{2}$$
. Then
 $M_{j,k} = (2b/(j^{2}+b)^{3}) I_{j,k}^{2} / \pi^{4}$

Let $r = b/(1+b)^2$ and $s = b/(4+b)^3$. r reaches its maximum when $b = \sqrt{2}$; then $r = \frac{4}{27}$ and $s = \frac{4}{(27)^2}$. s reaches its maximum when b = 2; then $s = \frac{1}{108}$ and $r = \frac{2}{27}$.

It is clear that the entries of \mathcal{M} are positive since $\mathcal{I}_{\mathbf{j},\mathbf{k}}$ is the integral of a positive function and the entries of the second row are much smaller than those of the first row. $\alpha = \frac{4}{3}$

I "/3	;	1.760401	I'm	7	2.157553
L 1/3	=	1.177827	I 1,2	9	1.667493
I""	4	2.593355	工 3/2	P	3,446855

These numbers are accurate to the last place, though some of them are extrapolated to determine the last place.

The problem then is to find the largest eigenvalue, λ , of

$$2\left(\begin{array}{ccc} r \stackrel{\uparrow a}{\coprod} & r \stackrel{\uparrow a}{\coprod} \\ s \stackrel{\downarrow a}{\coprod} & s \stackrel{\downarrow a}{\coprod} \end{array}\right)$$

Then λ/π^4 will be the largest eigenvalue of \mathcal{M} and $\mathcal{R}_{\sigma} = \pi^{*}/4$

This yields $O = \lambda^2 - 2(r I_{i,i}^{\alpha} + s I_{i,2}^{\alpha}) \lambda + 4rs (I_{i,i}^{\alpha} I_{i,2}^{\alpha} - (I_{i,2}^{\alpha})^2)$

It is difficult to maximize λ as a function of a^2 analytically. Instead, using the values of $\Box_{j,\kappa}$ obtained by numerical intergration, shown in the tables, values of λ were calculated in a range between b = .5and $b = \lambda$.

When $a^* = \pi \frac{1}{2}$, $b = \frac{1}{2}$. This is the value of both the critical Rayleigh number in the constant coefficient case (a=0) and the value of b corresponding to the first estimate R, of R, we did earlier.

The calculation was done as follows. For $\propto =4/3$, λ and $\approx R_{o}$ were evaluated at numbers

b = 000001, 0.10001, 0.20001, ..., 1.90001

The largest value of λ occurred when 0.50001 . λ and ${}^{*}\!\!R_{\star}$ were then evaluated at

b = 0.50001, 0.5100 , 0.5200 , ..., 0.59001

The maximum value of λ this time occurred when b = 0.51 col. So λ and R, were evaluated at

b = 0.50101, 0.50201, ..., 0.51801, 0.51901

This process was continued until the machine produced no differences in the sixth decimal place for \mathcal{R}_o . This represents eight significant digits. The numbers $\mathcal{L}_{\mathbf{J},\mathbf{l}e}$ are known only to seven significant digits. The calculation is not very sensitive to changes in b (and therefore b^2) when is near its critical value, but it should be quite sensitive to changes in $\mathcal{L}_{\mathbf{i},\mathbf{k}}$.

When $\alpha = \frac{4}{3}$, the critical value was established at b = 0.50904, $a^2 = 5.024023$

2R. = 183.5693 Then 2R = 8.520189×108 and The eigenvector is $\begin{pmatrix} A_{1} \\ A_{2} \end{pmatrix} = \begin{pmatrix} 0.9996005 \\ 0.0260555 \end{pmatrix}$ b = 0.50904, $\lambda = 0.5306395$, At 5 = 0.0055526 r=0,1481322 and The corresponding critical values for a = 3/2 were b = 0.51208a² = 5.05 4 127 ${}^{2}R_{2} = 148.9276$ R = 4.709315 ×107 The eigenvector is $\begin{pmatrix} A_1 \\ A_2 \end{pmatrix} = \begin{pmatrix} 0.9995499 \\ 0.0301839 \end{pmatrix}$ At b= 6.51208, &= 0.6540702 r=0.1481199 and s= 0.0055 745 If we compare R_o , the first estimate of R_o with R_o , the second estimate of R_o , we have x = 4/3 x = 3/2. R. 186.7505 152.3742 $^{2}R_{n}$ 183.5693 198.9276 The numbers ${}^{j}R_{\circ}$ converge quickly to R_{\circ} . If a third estimate is made the entries $\mathcal{I}_{\mathbf{3},\mathbf{k}}$ are smaller than 5.0. $\mathcal{L}=4/3$ $\mathcal{L}_{\mathbf{3},\mathbf{k}}$ Lus (),832954 1,289302 2.791557 1.871715 4.442455 Ī"z, 3.182768

This problem is analogous to the last. We are looking \prime for a function

$$\begin{split} & \mathbf{W}_{(2)} = A_{1} \sin n\pi + A_{2} \sin 2\pi z + A_{3} \sin 3\pi z \\ & \text{with} \\ & \int_{0}^{1} \left[(D^{2} - a^{2})^{3} W + a^{2} \frac{3}{R_{0}} (z + z_{0})^{-d} W \right] \sin j\pi z \, dz = 0 \\ \text{Let } t = b/(9 + b)^{3} . \text{ Then we must find the largest eigenvalue} \\ & \text{of the matrix, } M^{3} \\ & M^{3} = 2 \begin{pmatrix} r \ I_{1,1}^{d} & r \ I_{1,2}^{d} & r \ I_{1,3}^{d} \\ s \ I_{1,2}^{d} & s \ I_{2,2}^{d} & s \ I_{2,3}^{d} \\ t \ I_{1,3}^{d} & t \ I_{2,3}^{d} & t \ I_{3,3}^{d} \end{pmatrix}$$

The factor t is much smaller than s, just as s is much smaller than r. The determinate of \mathcal{M}^{s} is then $\mathcal{O}(rst)$. The sum of the determinates of the principal submatrices is $\mathcal{O}(rs)$. The trace is $\mathcal{O}(r)$.

Since the determinate is the product of the eigenvalues, the trace is the sum of the eigenmodes, and the sum of the partial products of the eigenvalues is the sum of the determinates of the principal submatrices, it is easy to see that \mathcal{M}^3 has one relatively large eigenvalue, λ , and that its other eigenvalues are near \mathcal{O} . The large eigenvalue should be $\mathcal{O}(\mathbf{r})$.

Because the large eigenvector is so isolated and because we have a good guess $(2 \wedge L_{i,i})$ for it, the power method was used to calculate λ and the eigenvector. This converged, rapidly, giving a value of λ that depended on β .

 λ was maximized exactly as in estimating ${}^{\tau}R_{o}$.

For $\alpha = \frac{4}{3}$, the maximum occurred when b = 0.50970and $\alpha^{-5.5030532}$. Then, the third estimate, $R_0 = 183.3736$ and R = 8.5/237540? The eigenvector is $\begin{pmatrix} A_1 \\ A_2 \\ A_3 \end{pmatrix} = \begin{pmatrix} 0.999656 \\ 0.026135 \\ 0.001981 \end{pmatrix}$ For d = 3/2, the occurred when b = 0.5/32, and $a^{2} = 5.06508$. Then $^{3}R_{0} = 148.7001$ and $R = 4.702310 \times 109$

XXX-12

The eigenvector is

(A, \		$\left(\begin{array}{c} 0.99 \ 75 \ 37 \\ 0.0 \ 3 \ 0 \ 30 \ 8 \\ 0.0 \ 3 \ 5 \ 20 \end{array}\right)$
$ \begin{pmatrix} A_{i} \\ A_{z} \\ A_{3} \end{pmatrix} $	5	0.030308
$\left(A_{3}\right)$		0.0025201

The three estimates	of the eigenvalue	converged quickly.
'R.	186.75	152,37
2Ro	183.57	148.93
²R。 ³R。	183.39	148,70

It should be noted that none of these values, λ , R_{\bullet} , or the eigenvector are very sensitive to changes of **b** or **a**² near the critical value.

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REFERENCES

Chandraskhar, S. "Hydrodynamic and Hydromagnetic Stability", Oxford University Press, 1961.

Gantmacher, F. R. "The Theory of Matrices", Vol. II, Chelsea, New York, 1959.

Velarde, M. G. and Christian Normond "Convection", Vol. 243 #1, Scientific American, July 1980.

Conclusions and Recommendations

Work like this in the case of variable viscosity, thermal conductivity and gravity has atmospheric applications as well as applications to the space program. This would be applicable to the GFFC experiment to be flown in Spacelab 3, especially if rigid boundary conditions were handled. At least some of this is presently feasible.

1980

NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

IMPROVEMENTS TO THE NASAP CODE

Prepared By:

Academic Rank:

University and Department

NASA/MSFC: Division:

MSFC Counterpart:

Date:

Contract No.:

David Perel, Ph.D.

Assistant Professor

University of South Carolina College of Engineering

Program Development Office

Carl E. Colley

August 5, 1980

NGT 01-002-099 The University of Alabama

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IMPROVEMENTS TO THE NASAP CODE

ΒY

David Perel Assistant Professor of Engineering University of South Carolina Columbia, South Carolina

ABSTRACT

The object of this work is the modification and improvement of the FORTRAN code, NASAP, written by the author. The purpose of NASAP is the transformation of CAD-generated NASTRAN input data to input data for DESAP II and/or DESAP I. The latter programs are used for structural optimization, and in the initial design phase are often more useful than NASTRAN, which is used for structural analysis.

NASAP is designed to operate interactively, and to this end, modifications and refinements have been made which simplify the interactive participation of the engineer, thus reducing chances for error as well as time spent at the computer terminal. Redundancies have been eliminated, and a tabular form of input has been incorporated into NASAP.

Currently, NASAP is being expanded to include iterative design cycles involving DESAP II and/or DESAP I.

ACKNOWLEDGMENT

The author thanks Carl Colley for his help in this project.

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INTRODUCTION

The Preliminary Design Office currently has the capability of creating NASTRAN bulk data by use of an interactive graphics system implemented on a Computervision system. This method of creating NASTRAN input data is far more efficient in terms of time and accuracy than manual preparation of the data. The CADDS software associated with the Computervision creates NASTRAN input data. This data is then transmitted to a main computer for NASTRAN execution in the batch mode. The output of the program is essentially the stress and displacement fields of the structure being analyzed.

DESAP I and DESAP II are also structural programs, but differ from NASTRAN in that they optimize the structure with respect to some variable, usually weight (i.e., minimizes the weight) subject to user-supplied limits on stress and displacement (DESAP I), and stress and buckling (DESAP II).

The input required for DESAP I and II is rather lengthy, but much of it, including the geometric description of the structural model, may be derived from the CADDS-generated NASTRAN input data. This is the purpose of NASAP.

The current version of DESAP I allows constraints on stress and displacement, while DESAP II allows constraints on stress and buckling. In order to generate a design satisfying constraints on stress, displacement, and buckling, it is necessary to use both DESAP I and II in an iterative cycle. Development of this capability is currently underway.

OBJECTIVE

The objective of this project is the improvement and expansion of NASAP in the following ways: (1) elimination of redundancies; (2) simplification of user-supplied input; and, (3) incorporation of the capability of iterative design cycles involving DESAP I and DESAP II.

REPORT

The body of this report consists mainly of the NASAP program. Its major subroutines are self-explanatory. An overall flow chart, Figures la-ld, outlines the main features of the program. The program is based on the following structural equivalences between NASTRAN and DESAP I and II:

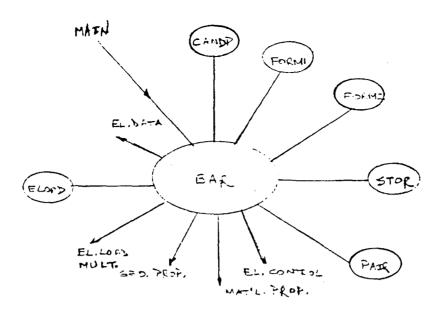
NASTRAN	DESAP I and II
ROD	BAR
BAR	BEAM
QDMEM	Plane Stress Quadrilateral
QUAD2	Plate Quadrilateral
TRIA2	or Triangle
SHEAR	Quadrilateral Shear Panel

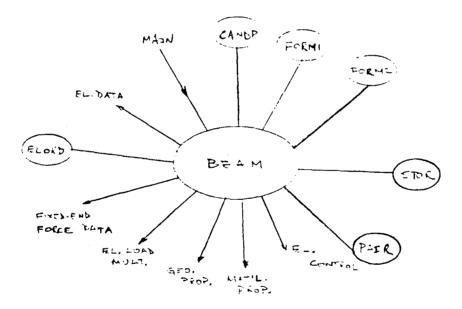
This is reflected in the five subroutines in NASAP called BAR, BEAM, QDMEM, QT, and SHEAR. These subroutines deal only with the structural elements their names imply, and perform ordering and manipulation of NASTRAN data via the subroutines CANDP, STOR, PAIR, VEC, and ID.

Figure 2 illustrates the design cycles that are possible using NASAP with DESAP I and DESAP II. The main feature of these cycles is the modification of file 13 (DESAP II) via calculated design variables and/or tolerances from DESAP I, and the corresponding route for file 14 (DESAP I).

When an acceptable design has been achieved using both DESAP I and DESAP II, there is no guarantee that this design is an optimum, i.e., that for the given structure it minimizes weight subject to the given constraints on stress, displacement, and buckling. There may exist other designs of lower weight also satisfying the constraints. This situation cannot be remedied without incorporating the three constraints in a single program.

MAIN HELDENG ORDER ŦĿ (VEC PROGRAM CONTROL SIGNA DESTEN CONTROL COUNT NODAT BAR BEAM QUMENI QUMEN QUMENZ SHEAR QT -STEUCTURAL LOAS DATA BUCKLING CONTROL NODAL LOAD DATA VARIABLE DATA restan FND LEGEND ' TO FILE IS (DESAP II) NASAP 53BROUTENE F=G. 12 XXXI-2





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FIG.11

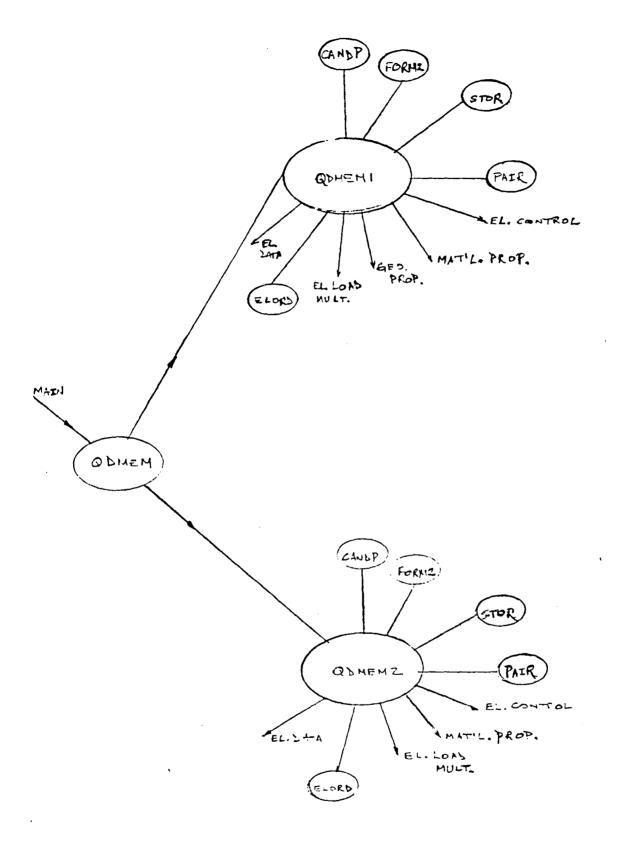
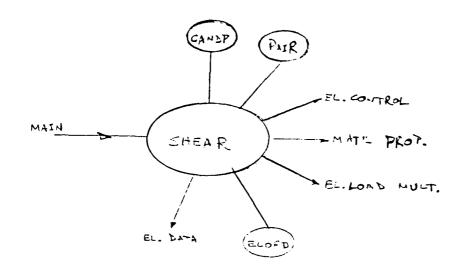


FIG. IC

XXXI-4



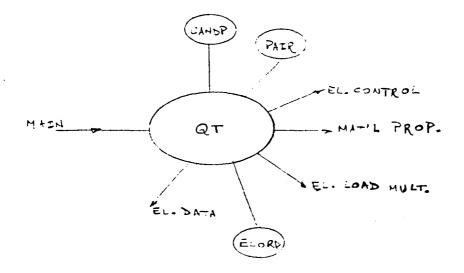
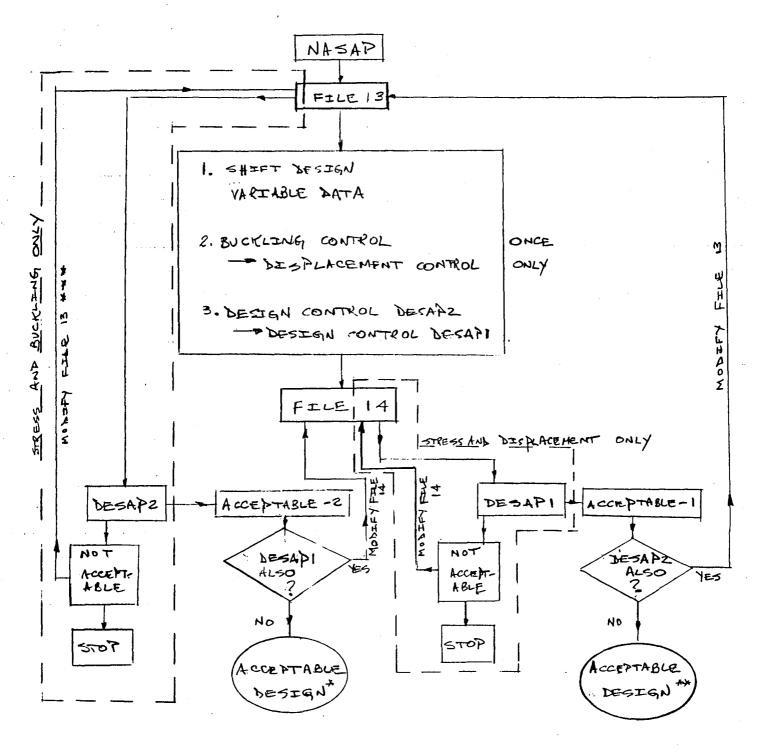


FIG. 1d



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- * SATISFIFS DEGAP2 OR DESAP2 AND DESAP1
- ** SATESFIES DESADI OR DESADZ AND DESADI

*** FILE MODIFICATIONS INVOLVE CHANGES IN DESIGN VARIABLES AND VARIOUS TOLERANCES XXXI-6

DESIGN CYCLES

FIG. 2

BAR	· ^			DATE 073080 PASE 2
uBan	1 007430	К1	6000 I 067444 K10 - 60cm I 007456 K12 - 0000 I 007451	K13 6000 1 007453 K15
	I 007455		COUD I DU7456 K17 COUD I DU7457 K16 0000 I DU7435	
0000	I 007431	K4	0000 I 007437 K5 0000 I 007440 K6 0000 I 007441	
	I 007443		5000 I 062114 MATCH 0004 00000 MAT11 0004 R 000024	MAT12 0000 1 006736 MTCON1
	R 067152		0000 R 007006 HTPROP 3000 I 007412 NEID 3000 I 007344	
ບນັ້ງເ	I 067423	NROW	0003 000022 NSTART 0003 I 000000 NUM 0000 T 007424	
0000	I 107401	NUM12	0000 I 007416 NUM18 0000 I 007400 NUM5 0000 I 006344	
1010 	1 032734	TOCRO	0000 I 007356 TYPE 0000 I 007352 VAR 0000 I 002424	VEID 0000 I 005214 VE1D2
L0101	1 #		SUBROUTINE BAR (CONFAC, NKODE, IDV1. NIDV1)	000002
60103	2*		COMMON/BACH/NUMI181,NSTART(18)/MOZART/HAT11(20),MAT12(20,10)	00002
0103	3*	<u>C **</u>	CHANGE GROUP. NO. OF DESAP BARINASTRAN RUDI ELEMENTS	<u> </u>
60104	4 *		INTEGER CROD(100,6), CRODID() J0,5), MATCH(100,2), VEID(100), INDEX(10	
	5*		J.TOCROD(190.6).LLOAT1(100.6).VE102(100)	000002
L0105	5 <i>*</i>		DIMENSION ELUAT2(100,4), DUHMY(100,1)	000002
60105		<u>C ++</u>	CHANGE GROUP. NO. OF PID'S	000002
60106	d≠		INTEGER PROD1(25,2)	000002
60107	9 *		UIMENSION PROD2(25,5),GEOPR2(25,3)	000002
60107	10*	C **	CHANGE GROUP. NO. OF HID'S	600002
<u> </u>	11*		INTEGER_MTCON1(26+2)	COUOO2
L0111	12*	• • • •	REAL HTPROP(20,5), HTCON2(20)	000002
<u></u>	13*	<u>C ##</u>	CHANGE GROUP. NO. OF DESIGN VARIABLES	600002
60112	14+		DIMENSION IDV1(100), IDV11(50), IDV12(50)	000002
	15*	<u>C **</u>	***************************************	000062
60113	16*		INTEGER KOAD(2), NPAR(5), VAR(4), TYPE(2)	000002
<u>60114</u>	<u> </u>		<u>REAL_MAT12,EMUL(4,4)</u> DATA_VAR/*A*,*B*,*C*,*D*/TYPE/*BAR*,**/	000002
00115				
<u> </u>	<u>19</u> ⊭ 2:J¢		NUM5=NUM(5) NUM12=NUM(12)	600004
ن121 10122	21*		CALL CANUP(5,NUM5,NUM12,CROD,DUMMY,PROU1, PP0D2)	600006
<u> </u>	22*	C	PRINT 14	000006
	23*	C14	FORMAT(" CROD IN BAR FROM CANOP")	000006
<u> </u>	<u> </u>	<u> </u>	UO 9 L=1.NUH5	000006
UD122	25*	C 9	PRINT 8. (CROD(L.J).J=1.4)	606006
L0122	26+	<u> </u>	FORMAT(8x,418)	000006
<u>cc122</u>	27*	C	PRINT 13	000006
60122	28+	<u> </u>	FORMAT (* PROD1 AND PROD2 IN BAR FROM CANDP*)	000006
J0122		(DO 16 L=1+NUM12	00006
00122	<u>3</u> U÷	<u>C16</u>	PRINI 12, (PROD1(L,J),J=1,2),(PROD2(L,J),J=1,4)	000006
	31*	C12	FORMAT(8%,218,4F6,4)	L00006
LC123	32*	7	FORMAT()	000017
UT124	33*		CALL FURHINNODE, KOAD, TYPEI	000017
L0125	34 \$		NPAR (1)=1	CCUN24
60120	35 *		NPAR (2) = NUH5	600026
60127	36*		NPAR (6)=1	UCG030
U0130	37*		Un 999 I=1, NUM5	000051
60133	38*		00 999 J=1,6	606051
Un136	39*	999	TOCROC(I,J)=CROD(I,J)	000051
L0141	4U#		00 1000 J1=1,NKOUE	000062
60144	41+		NPARISI=KOAD(J1)	<u>UDU162</u>
60144	42#	(**	IF ONLY ONE CONSTRUCTION CODE, SKIP TO 25	00062
<u>C0145</u>	43*		IF(NKOUE.EG.1) GU TO 25	C00064
ÜÜ147	44#		1F(J1+EQ+2) 60 TO 4PUU	COÜR66
06151	45÷		CALL FORM2(J1,KOAD, TYPE, NPAR(2), VEID, NUM5)	00071

на.	R		DATE 073080	PAGE	3
UG152	46+	ISTOP=NPAR(2)			
66153	47*	KOWNITEU	000163		
<u> </u>	48#	00 3100 IJK=1.NUH5	000114		
00157	49#	IFLAGIU	000114		
<u> </u>	50*	DO 3656 IUKL-1.ISTOP	<u>U0U117</u>		
CD163	51*	3050 IF(CROD(IJK, 1).EL.VEID(IJKL))IFLAG=1	000117		
<u> </u>	<u>52*</u>	IEIIELAG.EQ.11 GU TO 3100	<u>1001125</u>		
00170	53×	KOWNT=KOWNT+1	600130		
LC171	<u>54×</u>	VEIDZIKONTI=CROD(1JK.1)	000133		
60172	55*	3100 CONTINUE	000142		
<u> </u>	<u> </u>	<u>4600 CONTINUE</u> IF(J1.EC.1) GO TO 4100	000142		
60175	534	ISTOP=NUM5=ISTOP	<u> </u>		
		NPAR(2)=ISTOP	<u>000147</u>		
60201	59* 6n*	00-4650 IJK=1,ISTOP	600154		
	<u> </u>	4050 VELCIJK)=VELO2(IJK)	<u> </u>		
UG204	61# 62#	C PRINT 4051.ISTOP	<u> </u>		
	024	C DO 4L52 IJK=1,ISTOP	600154	· · · · · ·	
00254 66204	64*	C4052 PRINT 4053.VEID/IJK)	00154		
<u></u>	<u> </u>	C4052_PRINT_4055.VEID(10A)	<u>U20154</u>		
<u>00204</u>	<u>66</u> ≠	C4053 FORMAT(16) C4051 FORMAT(/* VEID IN BAR FOR J1=2*//* ISTOP=*+16)			
LC206	<u>67</u> ≠	4100 CONTINUE	<u> </u>		
LU200		UO 868 I=1.NUM5			
<u></u> 	69*	UU 068 J=1.6	LOUIST		
00215	7.2*	888 CROD(I.J)=TOCROD(I.J)	<u> </u>		
	71*	C PRINT 333	UD0164		
	724	L333 FORMATI' CROE IN BAR JUST BEFORE STOR ")			
	73=	C UO 777 I=1,10M5	ÜDG164		
60215	74#	C777 PRINT 222, (LROD(I.J),J=1.ISTOP)			_
6,1215	75*	C222 FORMAT(615)	U30164		
L0220	76*	CALL STORIISTOP, NUM5, CROD, VEID, 41			
uU22U	77+	C PRINT 67	606172		
<u></u>	7.4*	C67 FORMATI' CROL IN BAR JUST BACK FROM SIDE'/)			
<u>00220</u>	79×	C UO 77 J=1,ISTOP	C0U172		
<u>00220</u>	•Û*	<u>C77 PRINT 76, (CROU(J,K),K=1,6)</u>	L00172		
60220	61*	C PRINT 68	600172		
<u>10220</u>	<u>82</u> 2	C68 FORMAT(' VELL IN BAR JUST BACK FROM STUR')	<u>000172</u>		
65225	83≠	C PRINT 87, (VEÍU(I),I=1,ISTOP)	000172		
40220	04≠	CB7 FORMAT(1L14)			
LDZZJ	85¥	C76 FORMAT(615)	000172		
.UJ221	<u>⊽6</u> ≠	60 10 26	<u></u>		
C0222	⊳7 ≉	25 ISTOPENUM5	000203		
L0223		20 U0 7: J=1.ISTOP	L00205		
69226	69≄	NEID=CROL(J,1)	000211		
<u> </u>	<u>92*</u>	NPI0=CR0U(J, c)	006213		
LU227	91÷	C ** FINU PROUL CARD WITH SAME PIU	U00213		
ن23ن	<u> </u>	<u> 00.65.K=1+NUM12</u>	600220		
LU233	¥3≠	IFINPID .EQ.PRUDI(K,1))GU TO 61	000220		
LL1235	<u> </u>	<u>ьо IU 65</u>	LDJ223		
66230	95×	61 KK=K	600225		
<u>uu237</u>	96#	60 16 62	CU226		
00240	97*	65 CONTINUE	000232		
00242	98*	62 CRODIDIJ:11=NEIU	600232		
60243	93*	CROUID(J,2)=CRUD(J,3)	000234		
60244	<u></u>	CRUDID(J,3)=RUD(J,4)	<u>600236</u>		
60245	161*	CROUID(J,4)=PROD1(KK,2)	606240		
UC246	1.24	<u>CROUID(J,5)=4P1D</u>	600242		

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ВА	AR	. · · · · · · · · · · · · · · · · · · ·		DATE 073080	PAGE	4
60247	163*	70	CONTINUE	000246		
UD247	104+	С	PRINT 73	000246		
00247	105+	C73	FORMATE CRODIU FROM BAR*)	000246		
60247	106*	С	UO 74 J=1,ISTOP	000246		
00247	107*	C74	PRINT 75, (CRODID(J,K),K=1,5)	000246		
60247	168*	C75	FORMAT(515)	000246		
<u> 60247</u>	109#	C **	CRODID ARRAY IS DONE. IT CONTAINS.FOR THIS CONST. CODE.	000246		
60247	110*	C **	EID, G1, G2, MID, PID, IN ISTOPA = NPAR(2))ROWS.	000246		
00247	111*		FIND NUMMAT AND MATCH.	600246		
00247	112+	C **	IN CALL PAIR, PROD1 IS NOT USED	000246		
60251	113*		NUM18=NUM(18)	600246		
00252	114*		KKODE=U	000250		
60253	115+		CALL PAIR(KKODE, CRODIN, PRODI, TSTOP: NUM18, 4, NUMMAT, MATCH, INDEX)	000251		
Ú0254	116*		NPAR (3)=HUHHAT	000264		
<u>LI0254</u>	117*		SET UP HATERIAL CONTROL CARD FOR EACH OF THE NUMMAT	000264		
60254	118*	C **	MATI CARUS ASSOCIATED WITH CRODID,FOLLOWED BY THE	000264		
60254	119*	<u>C **</u>	MATERIAL PROPERTY CARD FOR THAT MATERIAL	000264		
60255	120*		DO 83 JIG =1,NUMMAT	000277		
00260	121*		MTCON1(J1C+1)= INDEX(11C)	600277		
00261	122*		MTCON1(J10,2)=1	000301		
60261	123*	C **	FOR MID NO. INDEX(J10) FIND CORRESPONDING ROW NO. IN MAT12	_000301		
F0262	124*		DO 84 J11=1,NUMMAT	000306		
<u>LU265</u>	125*		IF (HATCH(J11,1),EU.INDEX(J10)160 TO 85	000306		
60267	126*			000311		
00274	127*	85	NROW=HATCH(J11.2)	000313		
60271	128#			600314		
00272	129*	84	CONTINUE			
<u>0274</u>	130*	86	MTCON2(J10) = MAT12(NROW, 4)*CONFAC	000320		<u> </u>
60274	131*	c	PRINT 2222. MAT12(NROW.4)	000320		
60274	132 *		FORMAT(' HAT12 =',F10.0)	600320	<u> </u>	
60274	133*	C	PRINT 1111. MTCON2(J10)	<u>00320</u>		
00274	134*		FORMAT(2x, 'HITCON2 ='F2n.B)	600320		
00274	135×	c	PRINT 95. NROW.CONFAC			
LU274	136*	C 95	FORMAT(* NROW, CONFAC*, I3, 2X, F8.2)	000320		
60274	137 ±		SET UP MATERIAL PROPERTIES CARD FOR THIS MATERIAL	000320		
00275	138*	<u> </u>	MTPROP(J10,1):0.	600323		
00276	139*		HTP20P(J10,2)=MAT12(NROW,1)	000323		
60277	140*		MTPROP(j10,3)=NAT12(NROW,5)	000326		
60277	141*	С	PRINT_111, MIPROP(J10.2), MIPROP(J10.3)	600326		
60277	142*		FORMAT(' MTPROP+2 AND HTFROP-3 ='2F20.8)	000326		
<u>Un</u> 30u	143*		MTPROPIJIC.41=MAT12(NROW.6)	000328		
C0301	144+		MTPEOP(J10,5)=MAT12(NROW,9)	000332		
60302	145*	83	CONTINUE	000336		
60302	146*		MATERIAL PROPERTIES DATA IS LONE	<u> </u>		
00302	147*	C ##	IT IS NOT SENT TO FILE 12 UNTIL AFTER NPAR IS.	000336		
L0302	148*	C ##	SINCE NONE OF THE ELS OF PROUZ ARE USED BY DESAP	000336		
60302	149#		IT IS NOT NECESSARY TO MATCH THE PID'S OF THE CROD'S	000338 0nn336		
00302	150*	C ++	WITH THE ROWS OF PROD2	000336	· · · · ·	
60302	151*		SET UP GEOMETRIC PROPERTIES DATA. DEFAULT TO AREA =1Y=1Z=1.	000336		
<u> </u>	152*		WRITE (6,88) KUAH (J1), ISTOP	000336		
68310	153*	88	FORMAT 15X, "ENTER NUMBER OF UIFFERENT GEOMETRIC PROPERTIES FUR CON	000345		
6031g	154*		ISTRUCTION CODE , 14, '. MUST BE LESS THAN OR EQUAL TO ', 14)			
00311	155*	•	READ (5,7)NUMGEO	600345		
<u>UU311</u>	<u>155</u>		NPAR(4)=NUMGEO	000345		
	156*			660353		
<u></u>	<u>157*</u> 158*	7.05	WRITE(12,795)	000355		
00317		(75	FORMAT(* COMM ELEMENT CONTROL FROM BAR*)	686362		
U0320	159*		WRITE(12,79) (NPAR(IP), IP=1, b)	000362		

ja 🖌	<u>۲</u>		DATE 073060	PAGE	5
ບ _{ເປ} 32 _{ເມ}	16 ∂ ≄	C ++ NPAR TO FILE 12	000362		
60320	161*	C PRINT 799	UDU362		
60320	162#	C799 FORMAT(* NPAR TO FILE 12+)	C00362		
00320	163*	C PRINT 79, (NPAR(IP), IP=1,6)	600362		
00323	164*	79 FORMATI615)			
LU324	165#	wRITE(12,349)	608372		
LU326	166≠	349 FORMAT(COMM MATERIAL CONTROL AND PROPERTIES FROM BAR)			
60327	167 _{\$}	00 350 J12=1,NUMHAT	000404		-
66332	168*	BRITE(12,300)(MTCON1(J12,J13),J13=1,2),MTCON2(J12)			
00332	169*	C PRINT 8999	000404		
00332	17U#	CB999 FORMATI' MICONI AND HICONZ TO FILE 12')	606464		
UU 332	171*	C wRITE(6,300)(MTCON1(J12,J13),J13=1,2),MTCON2(J12)	000404		
LD341	172*		000416		
60342	173*	wRITE(12,301) (MTPROP(J12,J13),J13=1,5)	600416		
60342	174*	C PRINT 2999	000416		
60342	175 ≠	C9999 FORMAT(* MTPROP TO FILE 12*)	600416		
60342	<u>176</u> #	<u>C</u> <u>WRITE(6,3D1) (MTPROP(J'12,J13)+J13=1,5)</u>	000416		
6635	177*	301 FORMAT (5F10.0)	000437		
60351	178#	350 CONTINUE		• • • • • • • • • • • • • • • • • • • •	
60351	179#	C ** MATERIAL PROPERTIES DATA HAVE BEEN SENT TO FILE 12	000437		
_00353	<u>189*</u>	IF INUMBED. 6 T. 1360 TO 160	<u> </u>		
60355	161*	DO 101 K1=1,3	000446		
60360	<u>1d2#</u>	161 GEOFRZ(1,K1)=1.	000446		
60362	183*	00 1 ₀ 8 K4=1,1STOP	600453		
<u> </u>	164#	108_CRODID(84,5)=1	000453		
60367	185*	RITE (6,89)	000455		
60371	<u>166*</u> 167*	<u>E9 FORMAT (2X, FOR THIS CONST CUDE THERE IS ONLY ONE GEOMETRIC PROPER</u> 1TY. IT HAS BEEN ASSIGNED NU-1'/2X' AND THE FULLOWING PROPERTIES:	600462		
60371 60371	1074 188*	1AREA=1:IY=1:IZ=1. IF YOU WANT A DIFFERENT PROPERTY SET.*/	600462 600462		
L0371	189#	12X *ENTER IT IN THE SAME ORDER, OTHERWISE DEFAULT*)	080462		
LC372	1 9µ₹	READ (5.7) DUNI	000462		
00372	191#	C PRINT 666, DUM1	000462		
<u>.ca372</u>	192*	C666 FORMATI' DUMI LINE 173 MAIN' F10.5/1	BD462		
00375	193*	IF (DUM1 .EC, U.) GO TO 110	600470		
69377	194+	FEAD (5.7) DUM2.DUM3	000472		
00403	195#	GEOPR2(1,1)=DUM1	600501		
00404	156*	GEOPR2(1,2)-UUH2	000503		
L0405	1 - 7 +	GEOPE2(1,3)=0UM3	000505		
L.D.4116	198*	60 10 110	606507		
L04U7	199*	100 00 103 K2=1,NUMGE0	600511		
L0412_	2 <u>L0 +</u>	U0 1C3 K3=1,3	UOU522		<u>.</u>
JB415	≠1ر 2	1J3 GEOPF2 (K2,K3)=1.	600522		
<u>u042.</u>	2u2*	WRITE (6.9É)NUMGEO.	006530		
00423	263*	90 FORMAT (2X, FOR EACH OF THE +, I3, F GEOMETRIC PROPERTY NUMBERS, THE	000541		
<u>i0423</u>	21:4*	1FOLLOWING PROPERTIES HAVE BEEN ASSIGNED: 1/	COU\$41		
սն 423	265*	22X*AREA=1,IY=1,IZ=1。 IF YOU WANT A DIFFERENT PROPERTY SET,ENTER IT	000541		
00423	216±	3IN THE SAME GREER 1/2X FOLLOWING THE GEOMETRIC PROPERTY NO.TO BE G	L00541		
UU423	2.7*	41VEN OTHERWISE DEFAULT*)	000541		
40424	263*	<u> </u>	600541		<u></u>
00427	269*	WRITE (6,91)K4	000541		
<u>u0432</u>	21U*	91 FORMAT (5X, "GEUMETRIC PROPERTY NO.", 14)	000597		
60433	211*	READ(5,7)DUH1	600547		
00436	212*	1F (UUM1.EQ.4.)60 TO 144	000555		
60446	213*	READ(5,7) DUM2	000557		
10443	214	READ(5.7) DUH3	<u> </u>		
60446	215*	GEOPR2(K4,1)=DUM1	000573		
60447	216*	GEOFR21K4+21=DUM2			

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<u>.</u>	217*	GEOPR2(K4,3)=DUM3	000577		
LU451	218*	1U4 CONTINUE	000603	*	
00451	219*	C ** SET UP CRODIL FOR USE IN ELEMENT DATA	000663		
60453	220+	WRITE (6,92) NUMGEO	000603		
60450	221*	92 FORMATIZX, FOR EACH DESAP BAR ELEMENT NUMBER GIVEN, ENTER THE ASSOC	000614		
60456	222*	11ATED GEOMETRIC PROPERTY NUMBER 1 THROUGH ", 14)	600614		
00457	223*	UO 105 K5=1,1 _{ST} OP	000614		
LC462	224*	WRITE (6,7)CRODID(K5,1)	600614		
60465	225*	105 READ (5,7)CRODID (K5,5)	<u>600621</u> 000631		
60471	226÷ 227*	110 CONTINUE C ** GEOMETRIC PROPERTIES DATA CARDS ARE DONE	000631		
<u>00471</u> 00472	228*	wRITE(12,96)	000631		
60474	229*	96 FORMATE COMM GEOMETRIC PROPERTIES FROM BAR!			
60474	230*	DO 106 K6=1, NUMGEO	000635	_	
00500	231*	WRITE (12,93)K6, (GEOPR2(K6,K7),K7=1,3)	_000646_		
	232*		000646		
60500	233*	C5555 FORMATI' GEOPRE TO FILE 12'	000646		
00500	234#	C WRITE (6,93)x6, (GEOPR2(K6,K7), K7=1,3)	000646		
60507	235*	93 FORMAT (15.5x.3F10.3)	000665		<u></u>
60510	236*	106 CONTINUE	000665		
60516	237*	C ** GEOMETRIC PROPERTIES DATA HAS BEEN SENT TO DESAP	000665		
60510	238*	C ** GENERATE ELEMENT LOAD MULTIPLIERS	000665		
L0512	239*	WRITE (5:111) KOAD(J1)	000665		
60515	240*	111 FORMAT (15X, 'ELEMENT LOAD MULTIPLIERS FOR CONST. CODE +,14)	000677		
<u></u>	241*	DO 115 K6=1.4	000677		
66521	242×	DO 115 $K7=1,4$	000677		
L0524	<u>243</u> ≠	115 <u>EMULIK6,K7)=0.</u>	000677		
00527	244*	DO 120 K8=1,4	000710		
<u> </u>	245*	WRITE (6,112) VAR(K8) 112 FORMAT (2X,'FOR LOAD ',A1,' ENTER X,Y,ZELEMENT LOAD FRACTIONS	000710	·	
60535	246* 247*	I IN THAT ORDER. DEFAULT = ALL ZEROS*)	000715		
60536	248+	READ (5.7)DUM1	000715		
00541	249*	IF(DUM1.EQ.D.)60 TO 120	_00n723		
60543	250*	READ (5,7)DUM2	000725		
60546	251#	READ (5,7)CIIH3	000733		
00551	252*	EMUL (1, K8)=00M1	000741		
L0552	253*	EMUL (2,K8)=1,UM2	806743		
60553	254*	EMUL (3, k8)=00H3	000745		
L0554	255*	120 CONTINUE	000752		
00554	256×	C ** ELEMENT LOAD MULTIPLIERS ARE DONE	000752		
00554	257#	C PRINT 6666	000752		
60554	258*	C6666 FORMAT(* EMUL TO FILE 12+)	000752		
<u> </u>	259*	HRITE(12.117) 117 FORMAT(' COHM EL. LOAD MULTI. FROM BAR')	606752		-
0056U UU561	2604	UO 121 K9=1.4	000762 _600762		
60564	<u>261*</u> 262*	121 wRITE(12,116) (EMUL(K9,K10),K10=1,4)	<u>000762</u>		
E0584	262#	$\frac{121}{C121} = \frac{111}{WRITE(0, 116)} = \frac{100C(K9, K10), K10-1, 4}{(C121)} = 100C($	000762		
00573	264*	116 FORMAT (4F10.0)	001003		
60573	_265*	C ** GENERATE ELEMENT DATA	001003		
60574	266*	UO 150 KHOW=1,1STOP	U01003		
60577	267*	00 149 KCOL=1.5	L01003		
00602	2684	149 ELDAT1(KROW, KCOL)=CRODID(KROW, KCOL)	001003		
606.14	269*	IX_ELDAT1(KROW,1)			
00604	270*	C ELDAT1(KKOW,6)=IX	001005		
00604	<u>271*</u>	C LLUATZ(KROW,1)=2+IX	601005		
60605	272*	WRITE (6,130) 1X	601007		
00610	273*	130 FORMATIZX, 'ENTER THE DESIGN VARIABLE NO., IUV, FOR ELEMENT NO. ', 15)	001015		

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	274*	READ (5.7) ELDAT1(KROW.6)	001015		
60614	275*	ELDAT2(KROW,1)=1.	601023		
60614	276*	C WRITE(6,131) 1X	001023		
00615	277*	131 FORMATIZX'ENTER DESIGN VARIABLE FRACTION ,FRC, FOR EL.NO.+,14, DE	601025		
<u>00615</u>	278*	1FAULT=1.00')	001025		
LD615	279*	C READ(5,7)DUM	601025		
<u>Lü615</u>	28ila	C IF(DUM+E4+3+) 60 TO 151	L01025		
60615	281*	C ELDAT2(KROW,1)=DUM	CO1 ₀ 25		
00616	282*	151 CONTINUE	601025		
00617	263*	ELUAT2(KROW, 2)=D	001025		
60617	264*	C NRITE(6,132) IX	001025		•
00626 00626	265¥ 286≠	132 FORMAT(2X, "ENTER END-FIXITY COEFFICIENTS C FOR BUCKLING ABOUT THE 1Y AND 2 AXES, RESPECTIVELY, FOR EL.NO, ", 14)	001026		
- <u>110620</u> 	287#	C READ (5,7)ELDAT2(KROW,3),ELDAT2(KROW,4)	<u> </u>		
00020	268*	ELUATZIKROW, 3)=1.	U01026		
600022	269*	ELDATZ(KROW,4)=1.	001027		• •
60622	290+	C ELDATZ(KROW, 3)=IX	601027		
00622	291+	C ELDAT2(KROW,4)=IX*2.	001027		
LIN623	292×	150 CONTINUE	001034		
66623	293*	C PRINT 776	£01034		
00623	<u>294*</u>	C776 FORMATC ELDATA AND ELDATA IN BAR THAT IS SENT TO FLORD .	001034		
6623	295×	CC DO 1122 KROW=1,ISTOP	001034		
64623	296*	<u>C BRITE (6,7) IELDATI(KROH, 1), 1=1,6)</u>	001034		
LN623	297*	C1122 WRITE (6,7) (ELDAT2(KROW,J),J=1,4)	601034		
LOE23	<u> 298</u> #	<u>C ## URDER THE ELDAT ARRAYS IN ASCENDING ORDER OF FLEMENT</u>			
L0625	299×	CALL ELORD (ISTOP, ELDAT1, ELDAT2, 6, 4)	001034		
<u> 66525</u>	<u>3i</u>	C PRINT 4444	001034	· ,	
00626	361*	WRITE(12,135)	601043		
<u>ม่ถืองี่ม</u> ปฏุธงัญ	<u>3u2≉</u> 3u3≈	135 FORMATL' COMM. ELA. DATA FROM BAR') C4444 FORMATL' ELDATI AND ELDATZ FROM ELORD TO BAR TO FILE 12')	001050		
68631	3∟4≠	DO 160 K12=1.ISTOP	UC1050 001050		
60634	305 +	160 wRITE (12,133)(ELDAT1(K12,K13),K13=1,6),(ELDAT2(K12,K13),K13=1,4)	001056		
60634	3,,6 =	CI60 WRITE (6,133)(ELGATI(K12,K13),K13=1,6),(ELDAT2(K12,K13),K13=1,4)	<u> </u>		
6647	367*	133 FORMAT(615,4F10.U,5x)	001101		
60650	3 <u>Us</u> *	GO TO (450,460), J1	001101		
60651	369*	450 ISTGP1=ISTOP	U01111		
60652	<u> </u>	U0 451 K15=1,ISTOP	<u> </u>		
LD655	311*	451 IDV11(K15)=ĒLDAT1(K15,6)	001116		
<u>LU657</u>	312*		001120		
60666	313*	460 ISTOP2=ISTOP	601122		
00661	314#	00 461 K16=1+1510P	CD1123		
00664	315*	461 IDV12(K16)=ELDAT1(K16,6)	001127		
<u>00666</u> 00670	<u> </u>	1000 CONTINUE	001133	· · ·	
<u>CU676</u>	317# 318#	IF(NKODE.EQ.1) GO TO 1100 NIDV1=ISTOP1+ISTOP2	001133		
L0673	<u> </u>	UO 1L13 K17=1,1STOP1	<u> </u>		
00676	320*	1010 10v1(K17)=10v11(K17)	601157 001157		
60700	321*	UO 1020 K18=1,ISTOP2	001166		· · · · · · · · · · · · · · · · · · ·
00763	322#	1026 IDV1(ISTOP1+K18)=TDV12(K18)	001166		
03705	323≠	RETURN	U01170	·	
40706	324#	1100 NIDV1=ISTOP1	601174		
0707	325*	UO 111J K17=1,ISTOP1	001175		
00712	326≠	1116 IDV1(K17)=10V11(K17)	001103		
00714	327*	RETURN	601205		
Ln715	3284	£ND			
END FOR		-			

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	1 007330			UDOD I DIGZAY ENUFOR	0000 R 014544 FRC	0000 I 01477	Z GEOPRI	
	R 015251			0000 I 016322 I	0000 I 016352 IDUM	U000 I 01623		
0000	I 016057	10721		0000 I 016330 IFLAG	0000 I 016327 IJK	6000 I 01633		
-000	I 004064	INDEX	ŬŬŬO 017052 INJP\$	UDUO I 016325 ISTOP	0000 I 016407 ISTOP1	0000 I 01641		
	I 016344		LUJO I 016323 J	0000 I 016245 JOR	0000 I 016324 J1	0000 I 01634		
	I U16345		GOUD I 016350 J12	U000 I 016351 J13	0000 I 016353 J14	0000 I 01635		
0000	1 016335	K	0000 I 016336 KK	U000 I U16321 KKDE	0000 I 016340 KKODE	0000 I 01634	1 KKOL	
	I U16223		0000 I 016326 KOWNT	UDDD I 016370 KROW	0000 I 016364 K10	0000 I 01641		
	<u>I 016412</u>		<u>COOD I 016413 K17</u>	<u>СООЛ I 016414 К18</u> 0000 I 003554 матсн	0000 I 016355 K6 0004 000000 MAT11	0000 1 01635		
	1 016357		0000 I 016363 K9	0000 I 003554 MATCH 0000 I 016366 N1		0004 R 000024 0000 I 01640		
	I 015643 I 016367		L COUO R 015667 MTPRO2 COUO I U16372 M3	<u>- 0000 1 016366 H1</u> - 0000 Ι 016373 H4	<u>0000 I 016405 M12</u> 0000 I 016374 M4M1	UD ₀₀ I 01637		
	1 016400		GDDD I 016376 M7	0000 I 016401 M8	0000 I 010374 H4HI 0000 I 016402 M9	6000 I 01637		
	I 016332		0000 I 016403 NNCOL1	0000 I 016404 NNCOL2	0000 I 016265 NODE	GCUO I 01637		
	1 016225		0000 I 016334 NPID	0000 I 016346 NROW	0003 000022 NSTART	0003 I 000nn		
	I 016365		COCO I C16347 NUMGEO	0000 I 016342 NUMMAT	0000 I C16320 NUM13	0000 I 01633		
	1_016317		0000 I 014710 PBAR1	1000 R 015054 PBAR2	0000 R 016303 SFT	0000 I 001130		
	1 016255		0000 I 016233 VAR	U000 I 002114 VEID	6000 I 013724 VEID2			
				· · · ·				
00101	1*		SUBROUTINE BEAMICONFAC, NKO			600002		
00103	2*		COMMON/BACH/NUM(18),NSTART			000002		
<u>UL103</u>	3*	C ±+	CHANGE GROUP. NO. OF DESAP	BEAMINASTRAN BAR) ELEM	ENTS			
60104	4 *		INTEGER CBAR1(100,6), TCBAR	1(100,5), VEID(100), CBAR)	[D(100,7),MATCH(1	000002		
<u> </u>	5*	1	100.2).INDEX(100).ELDAT(100	161.ELDAT1,100,231.VEI	2(100)	000002		
60105	6* 7		DIMENSION CBAR2(100,3),FRC			000002		
60105	7*	C **	CHANGE GROUP, NO. OF PID'S			000002		
00106 00107	8 ¥ 9 #		INTEGER PBAR1(25,2), GEOPRI	163161 D2125 41 550002125 11		00002		
L0107	10*	C **	DIMENSION PBAR2125.51.GEOP CHANGE GROUP. NO. OF MID'S			<u> </u>		
60110	11+	• •	INTEGER MTPRO1(20)			<u>600002</u>		
C0111	12*		REAL MTPRO2(20,6)	·····	······································	000002		
L0111	13*	C **	CHANGE GROUP, NO. OF DESIG	N VARIABLES		000002		
60112	14*		DIMENSION IDV2(100), IDV21(50), IDV22(50)		000002		
00112	15*	C **	*********	****	****	<u>600002</u>		
60113	16*		INTEGER KOAD (2), NPAR (6), VA			000002		
00114	17*		LIMENSION DUMMIGI .NODE (2) .		_	00002		
60115	18*		REAL MAT12	-		000002		
60116	19*		DATA VAR/*A*.*6*.*C*.*D*/T	YPE/*BEAM*.* */JOR/*)	[*•*J*/	000002		
L0122	20×	7	FORMAT ()			600002		
00123	21*		NUM6=NUM(6)			000002		
00124	22*		NUH13=NUH(13)		•	000004		
<u> </u>	23*					400000		
UU126 UU126	2 ^{4 #} 25 *		CALL CANUP (KKUE, NUM6, NUH1 CBAR2 IS NEVER USED	3, CHARI, CHARZ, PBARI, PBAR	21	600010		
<u>00126</u> 00126		<u>c</u>	WRITE (6,5)	······································		000010		
00126	27*		FORMAT (25% BEAM ELEMENT	NATA CRAP IN MASTRANIES		000010 000010		
00126	28*	<u>c</u>	PRINT 14	VOIS TRAD AN HASTRANI		600010		
UU126	29*		FORMATI' CBAR1 AND CBARZ I	N HEAM FROM CANDD'S		000010		
00120	3.)*	C	UO 9 L=1,NUH6	HERE COLOR NOT A COMPLETE PROCESSION		000010		
L0126	31*		PRINT 88, (C6AR1(L,J),J=1,	5),(CB_R2(L,J),J=1.3)		600010		
00126	32*		FORMAT (2x, 516, 3F8.2)			000010		
60126	33*	С	PRINT 13		·	000010		
60126	34*		FORMATE' PBAN1 AND PBAR2 I	N BEAM FROM CANDP")		000010		
			00_16 L=1,NUM13					

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L0126	<u> </u>	C16	PRINT 12, (PEAR1(L,J), J-1, 2), (PBAR2(L,J), J=1,5)	006010		
00126	37*	C12	FORMAT(21,216,5F6.2)	000010		
L0127	38*		CALL FORHINKOUE, KOAD, TYPE)	000021		
00130	39+		NPAR(1)=2	000026		_
<u>Uc131</u>	<u>40</u> +		NPAR 121 ENUM6	000030	_	_
00132	41#		DO 999 I=1,200	000047		
60135	<u>42</u> #		<u>iu 999 1=1.5</u>	600047		
60146	43*	999	TCBAR1(I,J)≃CBAR1(I,J)	000047		
60143	44*		UO 1000 J1=1.NKODE	000060		
60146	45×		NPAR(5)=KOAD(J1)	600060		
66146	46±	<u> </u>	IF ONLY ONE CONSTRUCTION CODE.SKIP TO 25	000060		
L0147	47*		IF (NKUDE.EQ.1) GO TO 25	000062		
<u>00151</u>	<u>48</u>		1F(J1+E0+2) 60 TO 4000		• • • •	
60153	49#		CALL FORH2(J1,KOAD,TYPE,NPAR(2),VEID,NUH6)	000067		
Lc154	<u>50*</u>		LSTOP=NPAR(2)			
UD155	51*		KOWNTEU	000103		
60156	52*		00 3100 IJK=1,NUH6	000112	· · · · · · · · · · · · · · · · · · ·	
LD161	53×		IFLAG=0	000112		
60162	54*	•	00_3050_IJKL-1.ISTOP	000115	·	
UO165	55×	3050	IF{C&Akllijk,1).EQ.VEID{IJKL})IFLAG=1	000115		
<u> </u>	56*		IF(IFLAG.EQ.1) GO TO 3100	000123		
JU172	57*		KOWNT=KOWNT+1	600126		
L0173	58*		_YEID21K0HNT)=CHAR11IJK.1)			<u> </u>
LC174	59*	3100	CONTINUE	C00140		
<u> </u>	<u> </u>	4.300	CONTINUE			
60177	61*		IF(J1.EQ.1) 60 TO 4100	000140		
<u> </u>	<u>62¢</u>		_ISTOP=NUM6=ISTOP	000142		<u> </u>
63202	63*		NPAR(2)=ISTOP	000146		
00233	. 64*		UQ 4650 IJK=1.1STOP	600153		<u></u>
0206	65*	4050	VEIU(IJK)=VEID2(IJK)	000153		
<u>согаь</u>	<u>66*</u>	C	PRINT 4051, ISTOP	000153		
60206	67*	C	DO 4052 IJK=1,ISTOP	600153		
66200	68*		PRINT 4053, VEID(IJK)	000153		
00206	69#		FORMAT(16)	600153		
60266	<u> </u>		FORMATE VEID IN BEAM FOR JJ=2"//"ISTOP="+16)	000153	· · · ·	
0210	71 ±	4106	CONTINUE	600161		
60211	<u>. 72 ± </u>		<u>00 688 1=1.200</u>	<u> </u>		
60214	73#		UO 888 J=1,5	000161		
00217	74#		UBAK1(I.J)=TCBAR1(I,J)	<u> </u>		
60217	75 *	6	PRINT 333	000161		
<u> </u>	<u>76*</u> 77*		FORMATI' CBARL IN BEAM JUST HEFORE STOR')		· ·	
60217 60217	78*	C 777	UO 777 I=1,NUM6	600161 000161		
LD217	<u> </u>		PRINT 222. (CBAR1(I.J).J=1.51		·· · ·	
	/9≠ 80≠	6222	•	COU161 COU167		
<u>60222</u> 00223	<u>80</u> ≢ 51¥	· · ·	NNCOL=5	C00187 C00171		
	61¥ 	с	CALL STOR(ISTOP, NUM6, CBAR1, VEID, NNCOL)			
00223. L0223	<u>02</u> ∓ 83≠	<u> </u>	PRINT 67			
L0223	034 84#	C C C C C C C C C C C C C C C C C C C	FORMAT(* CEAR1 AND VEID IN BEAM JUST BACK FROM STOR*) NO 77 J=1.ISTOP	000171		
UU223	<u>04#</u> 65#	C77	PRINT 76. (CBAR1(J,K),K=1,5),VEID(J)			
00223	85*	C77		600171 600171		
<u> </u>	87*	<u> </u>	<u>FORMAT1615)</u> Go Tu 26	U00200		_
		75	ISTOP=NUM6			
<u>UD225</u> UD226	<u>88*</u> 89*	25	DO 70 J=1,ISTOP	<u> </u>		
60231	+€8 90+	26		00213		
60232	<u> </u>		NEID=CHAR1(J,1) NPID=CHAR1(J,2)	UUU213 UOU215		
60232	92×	C **	FIND PBAR1 CARD WITH SAME PIU			
	764	U ##	TINU FORNI CANU MITH SAME MIU	000215		

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00233	93*	D0 65 K=1,NUM13	000222		
60236	94.*	IFINPID.EQ.PHARI(K.1))GO TO 61	U00222		
U0240	95*	<u>60 TO 65</u>	000225		
60241	96*	61 KK=K	U00227		
U0242	97*	<u> </u>	00023n	· · · · · · · · · · · · · · · · · · ·	
LD243	\$8	65 CONTINUE	000234		
00245	99*	62 CBAR ID(J,1)=NEID	000234		<u> </u>
60246	100*	CBARID(J,2)=CBAR1(J,3)	600236		
00247	101*	CBARID(J,3)=CBAR1(J,4)	000240		
60250 10250	102* 103+	CBARID(J,4)_CBAR1(J,5)	000242		
60251	103+	CBARID(J,5)=PBAR1(KK,2) CBARID(J,6)=NPID	000244		
LD252 L0253	165*	7U CONTINUE	600246		
00253	106*	C PRINT 73	600252		
60253 60253	167*	C73 FORMAT(+ CBARID FROM BEAM*)	000252		
60253	168*	C DO 74 J=1,ISTOP	000252		
60253	109*	C74 PRINT 57, (CBARID(J.K)/K=1,7)	000252		
L0253	110+	C57 FORMAT(715)	000252		
L0253	111+	C ** CBARID ARRAY IS DONE EXCEPT FOR CBARID (.7)	600252		
L0253	112+	C ** FIND NUMMAT AND MATCH	000252		
L0255	113+	NUM18=NUM(18)	600252		
60250	114*	KKODE=0	600254		
00257	115+	KK0L=5	000255		
60257	116#	C PRINT 1011, ISTOP, NUM6, NUM18, NUM13	000255		
<u>60257</u>	117*	C1011 FORMAT(415)			
L026U	118*	CALL PAIRIKKODE, CBARID, PBARI, ISTOP, NUM18, KKOL, NUMMAT, MATCH, INDEX	000257		
L0261	119*	NPAR(3)=NUMMAT	000274		
UD262	12U#	WRITE(6,5D3)KOAD(J1)	600276		
60265	121#	503 FORMATIZX, "ENTER NO. OF FIXED END FORCE SETS FOR CONST. CODE . 14			
00266	122*	READ (5,7) NPAR(6)	000305		
<u> </u>	123*	C ** SET UP MATERIAL CONTROL CARD FOR EACH OF THE NUMMAT MAIL CARUS AS			
UU266 UU271	124* 125*	C ** WITH CHARID	- 606305		
00274	126*	UO 83 J1L=1.NUMHAT IX=INDEX(J10)	<u> </u>		
00275	127*	HTPRO1(J10)-IX	600310		
00275	128*	C ** FOR MIU # INDEX(J10),FIND CORRES. ROW # IN MAT12	000320		
60276	129*	UO 84 J11=1, NUMMAT			
60301	130*	IF(MATCH(J11,1).EQ. IX)GO TO 85	<u>000324</u>		
00303	131+	GO TO 84	000327		
L0304	132*	85 NROK = MATCH(J11,2)	000331		
00305	133*	GO TO 86	600332		
00306	134+	64 CONTINUE	000336		
00310	135*	85 MTPRG2(J10+1)=HAT12(NRQK,4)+CONFAC	000336		
60311	136*	HTPRU2(J10,2)=HAT12(NROW,1)	600341		
60312	137*	HTPP02(J10,3)=HAT12(NR0W,3)	<u>L0G343</u>		
60313	138*	MTPRU2(J10,4)=MAT12(NROW,8)	000345		
<u>CO314</u>	139*	MTPR021J10,51=HAT12(NROW,9)	<u>UOC347</u>		
60315	146*	MTPRO2(JIA,6)=HAT12(NROW,10)	Ung351		
60316	141*	B3 CONTINUE	<u> </u>		
60316 00316	142# 143#	C ** CALCULATE NPAR(4) C ** SET UP GEOMETRIC PROPERTIES DATA	600355		
60326	144#	KKOUE = 1	<u> </u>		
60321	145#	NKOL - I	U00355		
60322	146*	CALL PAIR(KKUDE, CRAPID, PBAR1, ISTOP, NUM13, KKOL, NUMGEO, MATCH, INDEX)	000357		
00323	147*	wRITE(12,795)	000361		
00325	148*	NPAR (4)=NUMGEO	<u>UUU374</u> LOB404		
	+ · · · ·	WRITE (12:504)(NPAR(J11),11=1.6)			

XXXI -15

81	_AM			DATE 07308n	PAGE	5
60331	150÷	795	FORMATL' COMM ELEMENT CONTROL FROM BEAM'I			
UD331	151*		NPAR TO FILE 12	000416		
60331	152*	c	PRINT 799	000416		
L0331	153*	6799	FORMAT(NPAR TO FILE 12.)	000416		
69331	154*	_ C	PRINT 504. (NPAR(IP),IP=1.6)	<u> </u>		
60332	155*	504	FORMAT (615)	000416		
60333	<u>156</u> *		wRITE(12.349)	600416		
60335	157*	349	FORMAT(* COMM MATERIAL PROPERTIES FROM BEAM*)	000427		
<u>L.0336</u>	<u> 158 </u>		UO 350 J12=1,NUMHAT	000427		
60341	159*	350		000427		
<u>L0351</u>	160*		<u>FORMATIIS-SX-F10-4-F10-0-F10-4-3F10-0</u>	000445		
6035 <u>1</u> 60351	161* 162*	C	PRINT 8999 Format(* MTPRO1 and MTPRO2 TO FILE 12*)	000445 000445		
<u>UU351</u>	<u> </u>	<u> </u>	PRINT 300, HTPRO1(J12), (HTPRO2(J12,J13), J13=1,6)	<u> </u>	· · · · · · · · · · · · · · · · · · ·	
60352	164#	C	<u>aRITE (6.270)KOAD(J1)</u>	000445		
U0355	105*	270	FORMAT (2X, ' FOR CONST. CODE', 14, + FOR THE FOLLOWING PROPERTY IDS,	000456		
	166*		1ENTER CROSS-SECTION CODE (KSEC):1=SYMMETRIC:2=ZEE:3=CIRCULAR . DEF	000456		
60355	167*		ZAULT=1')	000456		
L0350	<u>168*</u>		00 282 JINEINUMGEO	000456		
60361	169*		GEOPR1(J19,2)=1	C00456		
LU362	<u> 170* </u>		RITE_(6:271)INDEX(J10)	600457		
60365	171×	271	FORMAT(5X, 'PROPERTY ID=', 14)	600465		
<u> </u>	<u> 172 </u>		READ (5+7) IDUM	000465		
00371	173*		IF(IUUM.NE.0) GEOPRI(J10,2)=IDUM	000473		
<u> </u>	174*	282	CONTINUE	<u> </u>		
60375	1754		D0 283 J10=1,NUHGE0	000507		
<u> </u>	<u>176*</u> 177*	····	GEOFRI(J10,1)=INDEX(J10) For PID NO. INDEX(J10)FIND CORRES.ROW NO. IN PBAR2	<u>000507</u> 000507		<u> </u>
C1400	177 *	(#w		UUUSU7 UUUS14		
<u> </u>	179#		<u>00 264 J11=1+NUHGEO</u> IF(MATCH(J11+1)+EQ+ INDEX (,10)) GO TO 285	000514		
60404	183*		_60 TO 284	000517		
C0407	161*	285	NROW=MATCH(J11,2)	000521		
60410	182*		GO TO 286	000522		
60411	183*	284	CONTINUE	000526		
60413	184#	286	GEOPR2(J10,1)=PBAR2INROW.1)	000526		
60414	1a5*		GEOPR2(J10,2)=PBAR2(NRON,4)	600530		
60415	<u>196*</u>		GEOPK21J10.3)=PBAR2(NROW.3)			
CO41c	187*		GEOPR2(J13,4)=PBAR2(NROW,2)	600534		
0417	<u>188*</u>		<u> </u>	000543	<u> </u>	
60422	169*	290	GEOPK3(J10,J12)=0.	600543		
66424	190*		<u>WRITE(6.272)INDEx(J1D)</u>	000544		
00427 U0427	191* 192*		FORMAT(2X, "FOR PROPERTY 10", 14," ENTER THE SIX SECTION MODULI 2. Y	000554		
<u></u>	<u>1924</u> 193*		100 HAVE TO SPECIFY ALL OR NONE (DEEAULT).*/15x.*DEEAULT MEANS CORR	000554		
L0430	1934		<u>10 295 J13=1.6</u>	000554 000554		
<u> </u>	195*	295	GEOPR3(J1C, J13)=0.			
00,435	1 76 *	2,5	READ(5.7)DIMM(1)	000555		
<u>60440</u>	197*		IF(DUHH(1)+Eu+U+) 60 TO 297	000563		
UD442	198*		READ(5,7) (DUMM(J14),J14=2.6)	000565		
60445	199*		DO 298 J15=1,6	000602		
L0450	200*	298	<u>GEOPR3 (J10, J15)=DUPH(J15)</u>	006602		
C0452	2010	297		000615		-
60453	202#	283	CONTINUE	60615		
60453	2U3*	C **		000615		
60455	204*		WRITE(12,96)	000615		<u></u>
60457	265*	96	FORMAT(* COMN GEUMETRIC PRUPERTIES FROM BEAM *)	000627		
<u>60460</u>	206*		_D0_2cg_J10=1.NUMGE0	L00627		

BE	AH		DATE U73060	PAGE	6
60463	207+	299 WRITE (12,273)(GEOPR1(J10,J11),J11=1,2),(GEOPR2(J10,J12),J12=1,4),	000627		
U11463	218*	1(GEOPR3(J10,J13),J13=1,6)	000627		
៤០ី463	209*	C PRINT 327	000627		
60463	210+	C327 FORMAT(" GEOMETRIC PROPERTY CARDS TO FILE 12")	600627		
60463	211*	C DO 331 J10=1,NUMGEO	000627		
00463	212*	C331 WRITE (6,273)(GEOPRI(J10,J11=1,2), (GEOPR2(J10,J12),J12=1,4),	600627		
60463	213*	C 1(GEOPR3(J10,J13),J13=1,6)	000627		
00502	214#	273 FORMAT (215,4F10.3/6F10.3)	600661		
00502	215*	C ** GENERATE ELEMENT LOAD_MULTIPLIERS	000661		
60503	216*	WRITE(6,111)KOAD(J1)	000661		
<u>00506</u> 60507	<u>217*</u> 218*	111 FORMAT (//15x, *ELEMENT LOAD MULTIPLIERS FOR CONST. CODE*,14/) Do 115 K6=1,3	<u> </u>		
00512	<u>219*</u> 220*	DO 115 K7=1.4 115 EMUL(K6.K7)=0.	000673		
60515 60520	221*	115 EMUL(K6,K7)=0. D0 120 K8=1,4	ÚDD673 000704		
60523	222*	WRITE (6,112) VAR(K8)	<u> </u>		
<u>60525</u>	223*	112 FORMAT(2X, "FOR LOAD ".A1." ENTER X.Y.Z ELEMENT LOAD FRACTIONS, IN T	000711		
<u>00526</u>	224*	1HAT ORDERDEFAULT=ALL ZEROS*)	000711		
60520	_225*		000711		
60532	226*	IF(DUH1+EQ+D+)60 TO 120	000717		
60552	227#		000721		
<u>66540</u>	228*	EHUL(1,K8)=DUH1	000730		
60541	229*		400732		
LU542	230*	EMUL (3, K8)=DUM3	000734	-	
1.0543	231*	120 CONTINUE	606741		
60543	232*	C PRINT 6666	600741		
00543	233≠	C6666 FORMAT(" EMUL TO FILE 12")			
60545	234*	WRITE(12,1117)	600741		
60547	235*	1117 FORMATI' COMM EL. LOAD MULTIPLIERS FROM BEAM'	000751		
L055u	236*	DO 121 K9=1,4	000751		
60553	237#	121 WRITE (12,116) (EMUL(KQ,K10),K10=1,4)	000751		
GG553	238*	C121 PRINT 116, (EMUL(K9,K10),K10=1,4)	000751		
60562	239# 24û*	116 FORMAT (4F10.0)	000765		
50562 60563		C ** GENERATE FIXED-END FORCE DATA	000765		
<u> </u>	<u>241*</u> 242*	IF (NPAR(6)+EQ+0) GO TO 125 WRITE(6+117)KOAD(J1)	000767		
L0570	243+	117 FORMAT (20X. * FIXED END FORCE DATA FOR CONST. CODF*.14)	600775		
Ú0571	244*	NUMFX=NPAR(6)	600775		
00571	245*	C PRINT 1221	600775		
L0571	246*	C1221 FORMAT(' SFT TO FILE12')	000775		
00572	247*	wRITE(12,1116)	000777		
60574	248*	1118 FORMAT(COMM SFT FROM BEAN*)	001004		
<u>un\$75</u>	249*	00 123 M1=1.NUMEX			
00600	256*	WRITE (6,118)M1	001011		
60603	251*	118 FORMAT (2X. "FOR FIXED FORCE SET NO.".I3. ENTER RX.RY.RZ.MX.MY.HZ	<u> </u>		
00603	252*	1FOR NODE I .	001017		
60604	<u>253</u> *	READ (5,7)(SFT(M2),M2=1,6)	001017		
UU6U7	254 *	WRITE (6,119)	001027		
<u>U0611</u>	<u>255*</u> 256*	119 FORMAT (5X, 'SAME INFO FOR NOVE J')	<u></u>		
00612 UD612		READ (5,7)(SFT(M2), H2=7,12)	001034		
<u>60615</u> 60615	<u>257*</u> 258*	WRITE (12,122)M1,(SFT(M2),M2=1,12) C WRITE (6,122)M1,(SFT(M2),M2=1,12)	L01044 UC1044		
U0621	250+ 259+	122 FORMAT (15,6F10,2/5X,6F16,2)	601066		
LG622	263*	123 CONTINUE	601060		
00624	261*	125 CONTINUE	001060		
00624	262*	C ** UENERATE ELEMENT DATA	001060		
66625	263*	$W_{\rm R}$ ITE (6,127)KOAU(J1)	C01060		

B£	AM		DATE_073080	PAGE	7
6063ú	264*	127 FORMAT (7/15% BEAM ELEMENT DATA FOR CONST. CODE 15/)			
00631	265*	DO 150 KROW=1,ISTOP	C01101		
60634	266+	IX=CGARIU(KROW.1)	001101		
00635	267#	NODE(1)_CBARID(KROW,2)	001163		
00636	268*	NODE(2)=CBARID(KROL.3)	001105		
60637	269*	WRITE (6,130)IX	601107		
<u>UD642</u>	270*	130 FORMAT L2X. "ENTER DESIGN VARIABLE NO IDV.FOR EL. NO.". IS)			
60643	271*	READ (5,7)CBARID(KROW,7)	001115		
<u>U0646</u>	272*	WRITE (6,131)1X	£01123		
00651	273*	131 FORMAT (2x, 'ENTER DESIGN VARIABLE FRACTION, FRC, FOR EL. NO.", 14,	601131		
60651	2.7.4*	1º DEFAULT ELA 1	001131		
60652	275*	FRC (KROW,1)=1.	001131		
<u>_cc653</u>	276*	READ (5,7) OUM			_
00656	277*	IF (DUH.LQ.D.) GO TO 151	001141		
<u>00660</u>	278+	FRC (KRON, 1)=DUH	001143		
00661	279*	151 CONTINUE	001146		
LD662	2804	<u>DO 152 M3=1,4</u>	<u> </u>		
L0665	281*	152 ELDAT (KROW, H3)=+ +	601152		
<u> </u>	282*	IE(NPAR(6), ÉL.O.) GO. TO 153	601154		
60671	283¥	WRITE (6,132)IX	001156		
<u>60674</u> 60674	<u>284*</u> 285*	<u>132_FORMAT (2X, 'ENTER NUMBERS OF FIXED-END FORCE SETS FOR EL. LOADS A.</u> 18,C,D RESPECT.,FOR EL. NO.',14," OR DEFAULT TO ALL BLANKS')			
			U01164		
<u>0675</u>	286*	READ (5.7) IDUMH(1)	001164		-
00700 00702	287*	IF (IDUMM(1).EQ.0.) GO TO 153	001172		
<u>.00702</u> 00705	<u>288*</u> 289*	<u>READ(5,7) (IDUHH(H3),H3=2,4)</u> Do 154 M3=1,4	<u> </u>		
00705 00710	289# 29j]#	154 H3-1,4 154 LIDAT(KROW+M3)=IDUMM(M3)			
	<u></u> 291*	153 CONTINUE			
u0713_	2714	U0 160 M4=5,16	001217		
60716	293*	160 ELDAT(KROW.M4)=*	001217		
4072.	294#	DO 170 H4=1.2	001224		•
60723	295*	M4H1=M4+1	001224		
66724	296*	WRITE (6.135) IX. JOR(M4). NODE(M4)	001227		
ú0731	297*	135 FORMAT 12X "FOR EL. NO.", 13, " NODE ", A1, "(", I3, ") ENTER THE NO. OF	001237		_
00731	298×	1 END FORCES RX.RY.RZ.HX.HY.HZ KNOWN TO BE ZERO")	601237		
U11732	2494	READ (5.7) NONE	001237		
_60735	363.	1F (NONÉ, EG. 0) GO TO 170	001245		
60737	3 ₁₁ ‡	IF (NONE+EQ+6) GO TO 168	601247	· · · · · · · ·	
00741	362*	<u> </u>	001252		
60742	3113*	168 DO 165 M7=1,6	001254		
60745	304*	165 £1 DATIKROW,4+M4M1*6+M7)=*1*	001264		
LU747	345*	60 TO 175	U01266		
LÜ750	<u>306*</u>	169 WRITE (K.137)JUR(M4).NODE(M4)	001270		
CO754	3 u 7 ≠	137 FORMAT (2X, "FOR NOUE ", A1, "(", I3, ") ENTER THE END FORCES KNOWN TO	001305		
66754	<u>3u8*</u>	1 BE_ZEROUSE_THE_FOLLOWING_CODES: RX=1.RY=2MZ=6*)	601305		
U0755	369*	READ (5,7)(ENDFOR(M5),M5=1,NONE)	001305		
_L0763	310*	<u>ио 171 нь=1. NONE</u>	601327		
UU766	311*	LO 171 M7=1,6	001327		
<u> </u>	312*	LE [LNGFOR(H6)_FG_M7] ELGAT(KR04,4+H4H146+H7)=*1*	601327		
66773	313*	171 CONTINUE	001357		
66776	<u></u>	170 CONTINUE	001357		
u1600	315÷	150 CONTINUE	C01357		
<u>L1002</u>	316#	UO 175 Ho=1,1STOP	001357	· <u></u> ····	
01005	317*	U0 175 H9=1,7	001357		
61016	318*	175 ELDAT1(H8.M9)-CBAR1D(M6.M9)	<u> </u>		
01013	319*	00 176 M8=1, ISTOP	001371		
<u>01016</u>	320*	UO_176_M9=8,23	001371		

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	01021	321+	176 ELDAT1(M8,M9)=ELDAT(M8,M9-7)	001371		
	01024	322.	NNCOL1=23	001377		
	01025	323*	NNCOL2=1	601461		
	01026	324#	CALL ELORD(ISTOP,ELDAT1,FRC,NNCOL1,NNCOL2)	001403		
	<u>61026</u>	325+	C PRINT 4444	001403		
	U1026	326*	C4444 FORMAT(* ELEMENIT DATA TO FILE 12*)	601403		
•	01027	327*	wRITE(12,4445)	<u> </u>		
	01031	328*	4445 FORMAT(* COMH EL. DATA FROM BEAM*)	001417		
	<u>01632</u>	329*	UO 180 M12=1,ISTOP	001417		
	61635	330+	180 WRITE (12,136)(ELDAT1(H12,H13),H13=1,7),FRC(H12),(ELDAT1(H12,H13),	601426		
	01035	331*	1113=8,23)	<u>001426</u>		
	ü1035	332*	C180 WRITE (6,136)(ELDAT1(H12,H13),H13=1,7),FRC(M12),(ELDAT1(H12,H13),	601426		
	<u>u1035</u>	333*	<u>C 1H13=8+23)</u>	001426		
	61051	334 🛊	136 FORMAT (715,F10.0,4A5,12A1,3X)	001454		
	<u> </u>	335*	<u> </u>	<u>C01454</u>		
	61053	336*	450 ISTOP1=ISTOP	CO1464		
	<u>61654</u>	337*	00 451 K15=1, ISTOP	001465		
	ú1ú57	339+	451 IDV21(K15)=ELDAT1(K15,7)	001471		
	61661	339*	<u>60 TO 1000</u>	601473		
	<u>ί</u> 1ü6Ż	340*	46D ISTOP2=ISTOP	001475		
·	<u> </u>	341+	D0 461 K16=1.ISTOP	001476		-
	01066	342*	461 IDV22(K16)=ELDAT1(K16,7)	001502		
_	01070	343*	1000 CONTINUE	01506		
	61672	344*	IF(NKOUE.EQ.1) GO TO 1100	CD1506		
	01074	345*	NIDV2=ISTOP1+ISTOP2	601511		
	61675	346*	DO 1L10 K17=1,ISTOP1	601532		
XXXI-	61100	<u> 347* </u>	1010.1DV2(K17)=1DV21(K17)	601532		_
a l	01102	348*	D0 1020 K18=1,ISTOP2	001541		
n -	<u>01105</u>	349*	1020_I0V2(ISTOP1+K18)=IDV22(K18)	001541		
4	61107	350+	GO TO 1115	601543		
<u> </u>	01116	351*	1100 NIDV2=1STOP1	<u> </u>		
æ	i1111	352*	00 1110 K17=1,ISTOP1	001546		
	01114	353*	1110 IDv2(K17)=IDv21(K17)	001554		
	01116	354*	1115 CONTINUE	001557		
-	<u> </u>	355*	RETURN	001557		
	61126	356*	END	001645		
	ENU FOR	· · · · · · ·		·······		· · · · · · · · · · · · · · · · · · ·

.HDG,P CANDP

CANU	P					······································		<u> </u>		····-	······	DATE 073	080	PAGE	1
éFOR . S	CANDP.C.	ANDP													
HSA E3	07/30/80	-09:45	:12 (53,)												
SUBROU	TINE CAN	DP	ENTRY POIN	T CO0663				· <u> </u>							
STORAU	E USED: (CODE(1) 000710; D	ATA(0) OI	DO141; BI	ANK COM	1MON (2)	000000							
COMMO	N BLOCKS	:								········					
U D 3	BACH	00004													
6004	HANDEL				<u> </u>						·				
											<u> </u>				
EXTERN	AL REFER	ENCES	IBLOCK, NAM	ε,											
0005	ΙĹ														
<u></u>	<u>NRELS</u> NRDUg					······	<u></u>				<u> </u>	·······			
<u></u>	NI03														
0011	N1025														
<u>u012_</u>	NERR25														
0013	NI015														
1014	NERR35														
						<u> </u>									
SIOHAU	E ASSIGN	HENT	BLOCK, TYP	E, RELATI	IVE LOCAT	TION, NA	ME)								
0000	000033		0000	060054 1	OF	0000	000057		0061	200041	115G	0001	000064		
<u> </u>	<u> </u>			000150 1		<u> </u>	<u></u>			000156		00_01	000132		
0001 0000	000267 000034		0001 6001	000175 1		3001 3001	000242		0001	000230		0001	000261		
UUUU	000372		6061	000412	2466	0061	000427		0001	000332	2536	0001	000453		
6001				000561	276	4001	000522		0001	000530		0001	000577	3126	
	000605		0000	000062 3	55F	0000	000065	36F	0000	000070		0000	000042	5F	
<u> </u>	<u> </u>			060635 9			000044		0000	000047	.7 <u>F</u>	0000	000052		
	000000 9		0000	000100 1			000030	J		000031			I 000027		
	I 066632		<u>00i4</u>	<u>660000 k</u>	VCMID	<u>1004</u>		KVCPID_		060025			I <u>GUUU26</u>		· · · · · · · · · · · · · · · · · · ·
6004 0004	υίοι7ύ <u>Ι. Οι η172</u>			000171 H 000173 N		0000 1	000024	WE ND	0003 I	060022	NSIART	0003	1 000000	NUM	
00101	1*		SUBPOUTINE				,cout1,	COUT2,PO	UT1,POUT2)		000			
<u>60103</u> 60104	 3*		COMMON/BAC	H/NUMI181	INSIARTI	181					··	<u>000</u>	030		·····
	34 4±	C ±±	COMMON/HAN CHANGE GRO				U / PRAY	TOPHANDI	SANOUNIDAI	Unriv		000 1001			
C0105	5*	<u> </u>	INTEGER LO									000		·····	
00100	6*		DIMENSION									000			
60106		C **	CHANGE GRO									000			
60107	8¥		INTEGER PO									000			
00110	9#		DIMENSION									600			
00110	10*	<u>C_</u>	*********		******	******	****	*******	*			600			
00111 60111	11* 12*	C **	DIMENSION KDE. INPUT		-040 570							000			
	*	U 77	NUC INPUT	a 3-RUU+6	-DAKILIL	•							030		

CAN	UP.		· · · · · · · · · · · · · · · · · · ·	DATE 673080 PAGE 2
66111	13*	C **	NCCARD. INPUT NO. OF C CARDS TO BE READ	rinno 30
00111	14+	C **	NPCARD. INPUT NO. OF P CARDS ITO BE READ	000030
00111	15#	<u>C **</u>		000030
60111	16*	C **	POUT1,POUT2. OUTPUT	000030
60112	17*		REWIND 11	<u>000030</u>
00113	18*		NEND = NSTART (KUE)-1	000033
66114	<u>19*</u>		UO_5U_L=1,NEND	000036
0117	20*	50	READ (11,1)DUM	000041
00123	21*	1	_FORMAT (2044) LIM= NUM(KDE)	000051
60125	23*		LTM- NORTKDE/ KDEM4 =KDE+4	000053
LU125	24*	С	PRINT 100	000053
60125	25+	c100		00053
ju[126	26*		UO 51 L=1,LIM	000064
60131	27*		GO TG (15,16,17,18,17,20,21),KDEM4	600064
60132	∠8 ≠	15	READ (11,5) (COUT1(L,J),J=1,4)	600101
60132	29*	C	PRINT 5. (COUTI(L.,).J=1.4)	000101
66132	30×	C	PRINT 200,NUMPID	C00101
LC132	<u> </u>	C200		<u> </u>
60149	32*		CALL ID(2,NUMPID,COUT1(L,2),KC)	600116
60141	<u>33</u> *		COUT1(L.2)=KC	660126
60141	34*	С	PRINT 5, (COUT1(L,J),J=1,4)	000126
L0142	35*		<u>60 TO 51</u>	000130
00143 L0143	36≠ 37≠	16 C	READ (11,6) (COUT1(L,J),J=1,5); (COUT2(L,J),J=1,3)	000132
UD155	38*	<u> </u>	<u>WRITE (6.6) {COUT1(L.J).J=1.4}.(COUT2(L.J).J=1.3)</u> CALL ID(2,NUMPID,COUT1(L,2),KC)	000132
60155	30≠ 39≠		COUTI(L,2)=KC	000171
20157	40*		60 TO 51	600173
20160	41#	17	READ(11.7) (COUT1(L,J), J=1.6), COUT2(L,1)	006175
UU16U	42*	c	wRITE(6,7) (COUT1(L,J),J=1,6),COUT2(L,1)	000175
6c167	43¢		CALL ID(2.NUHPID.COUT1(L.2).KC)	000214
60175	44#		COUT1(L,2)=KC	606224
<u> </u>	45*		<u>60 T0 51</u>	U00226
60172	46*	18	READ(11,6) (COUT1(L,J),J=1,6)	600230
LG172	47#	C	WRITE(6,6) (COUT1(L,J),J=1,6)	000230
LC200	45≠ 49≠		CALL IU(2,NUHPID,COUT1(L,2),KC)	600245
<u> </u>	<u>49</u> 50*		COUT1(L_2)=KC GO TO 51	
00202	51×	20	_READ (11,10) (COUT1(L,J),J=1,5),COUT2(L,1)	UOD257 GCD261
<u> </u>	52*	<u> </u>	WRITE(6,10) (COUT1(L,J),J=1,5),COUT2(L,1)	000261
40212	53*	Ċ.	CALL IN(2.NUMPTD.COUT1(1.2].KC)	600300
LU213	54#		COUT1(L,2)=KC	000310
LC214	55*		60 TO 51	LOU312
66215	56*	21	READ (11,11) (COUT1(L,J),J=1,2),(COUT2(L,J),J=1,4)	000314
60215	57*	<u> </u>	RITE(6,11) (COUT1(L_J),J=1,2), (COUT2(L,J),J=1,4)	000314
LU227	58*		CALL ID(2,NUMPID,COUTI(1,2),KC)	000343
_L023,j	<u> </u>		COUT1(L,2)=KC	<u> </u>
60231	6Ü#	⁵ 1	CONTINUE	000361
<u> </u>	<u>61*</u>		REWIND 11	<u>UD0361</u>
6 ₀ 234 00235	62* 63*		"NEND = NSTART(KDE+7)-1 D0 52 L=1,NEND	UDQ364
UU235 UU24ui	<u> </u>	52	READ (11,1)0,1M	000367
00240	65 *	5 L	_LIM = NUM(kDE+7)	000372
00244	<u> </u>	Ċ	PRINT 101	LOU402
60244	67*	c101	FORMATILX. UTPUT FROM CANDP FOR POUTL AND POUT2. IN CANDP*/)	600402
60245	68*		DO 53 L=1,LIM	100404
0025u	69*		GO TO 125,26,27,27,27,27,27,KDEM4	000412

XXXI-20

CANDP			DATE 073080	PAGE	3
60251 70	* 25	READ (111-35) (pOUT1(1-1)-1=1-2) + (POUT2(1-1)-1=1-4)	000427		
60251 71		PRINT 35, (POUT1(L,J),J=1,2),(POUT2(L,J),J=1,4)	600427		
00263 72		CALL IU(2.NUMPID.POUT1(L.1).KP)	000456		
Lij264 73		POUT1(L,1)=KP	LO0466		
60264 74		PRINT 201.NUMMID	600466		
<u> </u>			000470		
		CALL IO(1.NUMMTD.ROUT1(1.21.KE)	000470		
L0267 77		POUT1(L,2)=KP	000500		
60267 78	* C	PRINT 35, (POUT1(L,J),J=1,2), (POUT2(L,J),J=1,4)	000500		
LO270 79		GO TO 53	000502		
60271 80			000504		
00271 81		PRINT 36, (POUT1(L, J), J=1,2), (POUT2(L,J), J=1,5)	000504		
60303 82		CALL ID (2. NUMPID. POUTILL. 1). KP:	000533		
60304 83		POUT1(L,1)=KP	U00543		
L0305 84	*	CALL ID(1,NUMMID.POUT1(L,2),KP)			
UQ306 85	\$	POUT1(L,2)=KP	000555		
6n3n6 86		PRINT 36, (POUT1(L,J), J=1,2), (POUT2(L,J), J=1,5)	000555		
66307 87		60 TO 53	000557		
60310 88		READ (11.37) (POUTL(L,J),J=1.2).(POUT2(L,J),J=1.2)	<u> </u>		
LO31ú 89		PRINT 37, (POUT1(L,J), J=1,2), (POUT2(L,J), J=1,2)	000561		
60322 90		CALL I0(2.NUMPTD.POUT1(1.1).KP)	<u></u>		
L0323 91		POUT1(L,1)=KP	000620		
L0324 92		CALL ID(1.NUMMID.PQUT'1(L.2),KP)			
60325 93		POUTI1L,21=KP	000632		
<u>un325</u> 94		PRINT 37. (POUT1(L.J).J=1.2). (POUT2(L.J).J=1.2)			
LO326 95			000640		
60330 96		FORMAT (8x,418)			
00331 97		FORMAT (8X,518,3F8.4)	600640		
Q8		FORMAT (8X.618.F5.4)	000640		
LC333 99		FORMAT (6X,618)	070640		
<u>u0334 1u0</u>	<u>* 10</u>	FORMAT (8X.518.58.4)	000640		
LO335 1L1		FORMAT (8X,218,4F8.4)	000640		
00336 102		FORMAT (8x,218,4F8,4)	000640	=	
60337 103		FORMAT (8X,218,5F8.4)	000640		
00340 104		FORMAT (8X,218,2F8,4)			
60341 165		RETURN	000640		
60342 106	*	END	C00707		
END FOR					
aHDG P	COUNT				
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		· ·			
·					

COUNT

<u>aFOR,S</u> <u>COUNT,COUNT</u> HSA E3 -07/30/80-09:45:18 (3,)

SUBROUTINE COUNT ENTRY POINT 600053

STORAGE USED: CODE(1) DODU63; DATA(U) DUDD12; BLANK COMMON(2) DODUDD

COMMON BLOCKS:

GUG3 HANDEL JC0172

EXTERNAL REFERENCES (BLOCK, NAME)

UUU4 NERR35

STURAGE ASSIGNMENT (BLOCK, TYPE, RELATIVE LOCATION, NAME)

CUU1 000011 1066	5001 000022 120L	U001 000046 121L	Ugu1 000027 1226	0001 000015 20L
0001 000CC33 21L	6000 I 050001 I	UDUO DODUOZ INJPS	0000 I DUNNED KOUNT	UDUJ I DUDUDU KVCHID
UCC3 I OUDU24 KVCPID	GOUS I OUUITO HAXMID	0003 I 000171 MAXPID		

DATE 073080

PAGE 1

			· · ·	
<u>cui01</u>	14	SUBRCUTINE COUNT(NUMMID_NUMPID)	<u></u>	
LU133	2 *	COMMON/HANDEL/KVCMID(20),KVCPIU(100),HAXHID,HAXPID	600005	
60103	<u>3</u> *	C ** NUMMID(NUMPID).OUTPUT.NO.OF NON ZERO ELS.	000005	
60103	4 ¢	C ★★ IN KVCMIU(KVCPID)	000005	
<u> </u>	5*	<u> </u>	000005	
60104	6*	KOUNT=0	600005	
LU105	7*	Uo_lco_l=1.MAXMID	<u>C00011</u>	
6011 0	8 *	1F(KVCMID(1)+NE+0) GO (TO 20	000011	
<u>L0112</u>	9*	<u>60 TO 126</u>	000013	
60113	≉ل1	20 KOUNT=KOUNT+1	600015	
60114		100 CONTINUE	000022	
60116	12*	120 CONTINUE	600022	
60117	13*	NUMMID=KOUNT	000022	
60120	14 =	KOUNT=ü	000023	
00120	15*	<u> </u>	006023	
U0121	16*	00 101 I=1,MAXPID	600027	
<u>i0124</u>	17#	IF(KVCPIU(I) NE.D) GO TO 21	000027	
60120	18 #	60 TO 121	600031	
66127	19*	21 KOUNT_KOUNT+1	0nnn 33	
60130	20×	101 CONTINUE	UC0040	
00132	21*	121 CONTINUE	000040	
60133	22*	NUMPID_KOUNT	606040	
60134	23#	RETURN	000041	
00135	24×	END	000062	
ENU FOR				

aHUG,P UIF

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Ð	1	F

_FOR.S__UIF.UIF_ HSA_E3_07/30/80-09:45:20 (3,)

SUBROUTINE DIF ENTRY POINT UDULUS

STORAGE USED: CODE(1) ODD123; DATA(U) ODD220; BLANK COMMON(2) ODD000

EXTERNAL REFERENCES (BLOCK . NAME)

LUD3 NERR35

 STORAGE	ASSIGNMENT	(BLOCK, TY	PE, RELA	IVE LO	DCATION, NA	ME)			_	_				
 6001	000016 106G	0001	000033	116G	0001	000036	1216	0001	000060	136G	0001	000050 1	191	
 	000042 30L	0000	I 000176	1 -	0000	000201	TNJPS	ODDO	000200	J	0000I	000177.8	COUNT	
3030 I	DUGUDO MAX	0000	I 060001	VIDV1	220									

DATE 073080

PAGE 1

	U0101	1*		SUBROUTINE DIF (NVIDV.VIDV.NDIF)	600011	
_	U0101	2*	(**	NVIUV.INPUT.NO. OF ELS. IN VIDV	000011	
X	601n1	3*	C ++	VIDV.INPUT.VECTOR OF DESIGN VARIABLE NOS.		
- X -	60101	4*	C ++	NDIF.OUTPUT.NO. OF DIFFERENT DESIGN VARIABLE NOS.	000011	
×	U0103	5*		INTEGER_VIDV(125).VIDV1(125)	800011	س
_ _	60104	6*		MAX=VTDV(1)	ū00011	
22 _	60105	7*		DO 10 1=2.NYIDY	6	×
	00110	8 *		IF(VIDV(I).GT.MAX) MAX=VIDV(I)	U00016	*
	60112	9 🖈	10		000025	•
	UC114	10+		KOUNTEG	000025	"
·	L0115	11+		DO 20 I=1.HAX	000026	
	60120	12+		DO 19 J=Ī,NVIDV	000036	*
	60123	13*		IF(VIDV(J) ELSI GO TO 3G	000036	
- 	ÚÜ125	14*		GO TO 19	000040	•'
		<u>15*</u>		KOUNT=KOUNT+1	000042	"
	60127	16*		VIDV1(KOUNT)=I	606045	43
	60130	17#	19	CONTINUE	U00053	"
	üD132	18 #	2 ü	CONTINUE	000053	**
	<u>00134</u>	<u>19*</u>	40	NDIF:1	600053	
	UD135	2U≠		UO 5L I=2,NVIDV	000060	**
	00140	21*		IF(v10v1(1).NE.VIDv1(I=1)) NUIE=NDIF+1	000060	
	00142	د2∗	50	CONTINUE	000067	-
·	60144	23*		RETURN	000067	
	L0145	24 *		END	600122	••
	END FOR					CT

GHDG,P ELORD FLORD

DATE 073080 PAGE 1

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.FOR, S ELURU, ELORD

HSA E3 -07/30/80-09:45:23 (20,)

SUBROUTINE ELORD ENTRY POINT DO0151

STORAGE USED: CODE(1) 000165; DATA(0) 000073; BLANK COMMON(2) 000000

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EXTERNAL REFERENCES (BLOCK, NAME)

UUU3 NERR35

STORAGE ASSIGNMENT (BLOCK, TYPE, RELATIVE LOCATION, NAME)

6001 060641 1166 0011 060047 1531	6801 6881	000105 1326	0001 000122 142G	0001 000023 151L 0000 I CU0001 H0LD1	6001 600053 152L 0000 R 000030 H0LD2
0000 I 000040 I	0000	OUDD44 INJPS	0CUO 1 000041 J	0000 I 000036 NEND	OGGO I OCOO37 NROWMX

.

	L0101	1#		SUBROUTINE ELORD(NUM.ELDAT1.ELDAT2.NCOL1.NCOL2)	COno17
	60101	2 #	(; **	CHANGE GROUP. NO. OF ELEMENTS	606017
5.	60103	34		INTEGER ELDATI(100,23)	000017
2	60104	4 #		DIMENSION ELUAT2(100,6)	L06017
	66104	<u>5¥</u>	<u>C ++</u>	*****	600017
ç	L0105	6*		INTEGER EIDMAX,HOLD1(23)	000017
<u> </u>	LN106	7 ¢		DIMENSION HOLD2(6)	600017
	ს 01ე6	8*	C	ORDERS ELDAT ARRAYS IN ASCENDING ORDER OF IEL (ELDAT1 (L,1))	000017
_	<u> </u>	9*	C	NUM.INPUT.NO. UF POWS IN ELDATI AND ELDAT2	000017
	60100	10÷	C	ELDATI, ELDAT2. ON INPUT, ELEMENT DATA ARRAYS. ON OUTPUT, EL. NUS. HAVE BEEN	000017
	<u> </u>	.1*	Ç	PUT IN ASCENCING ORDER	<u> </u>
	CO106	12*	C	NCOL1,NCOL2.INPUT.NO.OF COLS. IN ELDAT1 AND ELDAT2	600017
	UQIU7	13=		NEND=NUM+1	L00017
	L0113	140	151	NEND=NENU-1	000923
-	60110	15#	C	PRINT 111, NENU	600023
	60110	164	C111	FORMAT(* NEND=*, I3)	600023
	0111	17#		IF (NEND.EQ.1) GO TO 160	<u></u> C00025
	66113	16*		EIDMAX=ELDAT1(1,1)	00027
	60114	<u>19</u> #		NRQHMX=1	<u> </u>
	Ա <mark>Ը11</mark> 5	≠ل∡		UO 152 I=2,NLNU	U00041
_	L0120	21#		IF(ELDAT1(1,1).GT.EIDMAX) 60 TO 153	600041
	60122	22*		60 Τ ₀ 152	000045
	<u>(12)</u>	23*	153	EIUMAX=ELDAT1(1+1)	
	66124	24*		NROWMX=1	600050
	00125	25×			600055
	125يات	26 \$	С	PRINT 222, NENU,NROWMX,EIDMAX	LOU055
	<u> </u>	27*		FORMATI NENU, NROWMAX, ÉIDMAX= (33(2X, I3))	L00055
	LiC125	∠8 ≠	С	INTERCHANGE ROW WITH LARGEST EID WITH ROW NEND	C00055
	LU127	29*		IF (NKO_MX+EQ+NEND) GO TO 151	000055
	60131	≠⊎ک		UO 154 J=1,NCOL1	LOUO60
	<u>U0134</u>	31*		HOLD1(J)=ELDAT1(NEND,J)	000105
	60135	32*		LLDATI(NEND, J)=ELDAT1(NRGWMX, J)	600106
	L0136	33*		ELDAT1(NROWMX,J)=HOLD1(J)	600110

ELO	<u>RD</u>			DATE 073080	PAGE	
60137	34*	154	CONTINUE	000122		
L0141	35*		D0 155 J=1,NCOL2	600122	·····	
00144	36*		HOLD2(J)=ELDAT2(NEND,J)	ÜOn122		
00145	37*		ELUAT2(NEND,J)=ELDAT2(NROWHX,J)	600123		
6 <u>0145</u> 60147	3.3*		ELDAT2(NROWMX+J)=HOLD2(J)	600125	····	
60147	39*	155	CONTINUE	600130		
60147	40*	<u>(</u>	PRINT 533. ELDAT1(1.1).ELDAT1(251).ELDAT1(3.1) FORMAT(' ELDAT1-3 ROWS,FIRST COL.=',3(2X,I3))	<u> </u>		
UO147 UO151	41* 42*	6333	GO TO 151	000130 U00130		
L0152	43*	160	CONTINUE			
00153	44*	100	RETURN	000132		
00154	45*		END	000164		
END FOR						
aHDG,P	FORM	11				
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FURI	11			DATE 073080	PAGE	1
aFOR.S	FORM1,F	FORM1	•			
			5:27 (10,)			
				<u> </u>		
<u>SUBROL</u>	ITINE FOF	RH1	ENTRY POINT LONDS6			
STORAL	E USED:	CODE(LI COUNTE: DATAIGI DODUAE: BLANK COMMONIEI DUCODO			
EXTERN	AL REFER	RENCES	(BLOCK, NAME)			
<u></u>	NEDUS					
3004	NIO15					
<u>ບປີນ5</u> ມີມີນ6	NEDUS					<u> </u>
UUU7	NERR31	L.	·			
<u>aup</u>	0u6031	1 INJP	· · · · · · · · · · · · · · · · · · ·			
20101	1#		SUBROUTINE FORM1(NKODE,KOAD,TYPE)	000000		
<u> </u>	<u>2*</u> 3*		INTEGER KOAD(2).TYPE(2) NKODL.UUTPUT.NO. OF CONST. CODES	<u> </u>		
00103	_>≁ 4.¥	C C	KOAU+OUTPUT+NO+ OF CONST+ CODES	000000		
60103	5*	Č	TYPE.INPUT.NAME OF CURRENT ELEMENT	600000		
60104			WRITE (6,6)TYPE	600001		
E0112 E0112	7≠ 8≠	6	FORMAT (5X, 'ENTER THE NO. OF CUNST. CODES USED FOR ', A6, A2, ' ELEME INTS, 1 OR 2')	000013		
60113	<u>9</u> #		KOAU(1)=1	600013		
<u>CO114</u>	10+		KOAU(2)=2	666015		
	11*		READ(5,7)NKODE	600017		
00115	12*		FORMATI J IFINKOUE.EC.21 GU TO 20	000025		
00120			IFINKUUL+L+21 GU IV 2U	600025		
60120 60121	13 [#] 14±					
00120	13* 14* 15*	8	WRITE (6,8) FORMAT (5x. 'ENTER THE CONST. CODE NO. (1 OF 2)')	<u> </u>	· · · · - · - · - · - · - · - · - · - ·	
60120 60121 60123	14☆ 15≈ 16≉	8	WRITE (6,8) Format (5x, 'Enter the Const. Code No. (1 of 2)') Read(5,7) Koad(1)		· · · · · · · · · · · · · · · · · · ·	
60120 60121 60123 60125	14± 15≠	8 2ú	FORMAT 15X, 'ENTER THE CONST. CODE NO. (1 OF 2)')	<u>000030</u> 000035	· · · · · · · ·	

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600071

LD132 18*

HUG.P FORM2

END

	M2		DATE 073080 PAGE 1
SFOR,S	EORM2_FORM2		
	-07/30/80-09:		
	. <u> </u>		
SUBRO	UTINE FORM2	ENTRY POINT GOD141	
STORA	LE USED: CODE	(1) 000166; DATA(U) 700131; BLANK COMMON(2) 00000	· · · · · · · · · · · · · · · · · · ·
EXTER	NAL REFERENCE	S_(BLOCK . NAME)	
<u> </u>			······································
<u> </u>			
0007			
STORA	UE ASSIGNMENT	(BLOCK, TYPE, RELATIVE LOCATION, NAME)	
	0J0050 10F	0001 00006 1066 0001 000017 1146 0001 000046 13	UG 0001 000054 1346
บิกมา		<u>6 0001 000122 1616 0000 000014 2CF 0000 000047 7F</u>	
0000	I 000000 I	CUÚD DÚD113 INJP\$ GÖÜÜ I DODDO1 K	
	· · · · · · · · · · · · · · · · · · ·		
			······································
60101	1*	SUBROUTINE FORM2(J1.KOAD.TYPE.ISTOP.VEID.MAX)	600006
60101		* J1.INPUT.INDEX OF KOAD	000006
u0101	3* C*		000006
00101	4* C*	* TYPE.INPUT.TYPE OF CURRENT ELEMENT	000006
u0101		* ISTOP.OUTPUT.NO. OF ELS. FOR CURRENT CONST. CODE	60006
00101	6≑ C ≉	* VEID.OUTPUT.ID NUMBERS OF ELEMENTS	000006
60101		* MAX_INPUT_CONTAINS_MAXIMUM_NOOF_ELEMENTS	000006
	8.# C.#		000006
C0101			
UG101		* CHANGE GROUP. NO. OF ELEMENTS	000006
UC101	10+	INTEGER VEID(100)	000006
60101 60103 60103	10≠ 11≠ C ≠	INTEGER VEID(100) * **********************************	000006 000006 400006
UC101 UC103 UC103 UC104	10¢ 11* C * 12*	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006
UC101 UC103 UC103 UC104 UC104	10¢ 11¢ C¢ 12¢ 13‡	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006
0101 0103 00103 00104 00105 00110	10* 11* C* 12* 13* 14* 1	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006 000006
UC101 UC103 UC103 UC104 UC104	10¢ 11¢ C¢ 12¢ 13‡	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006
00101 00103 00103 00104 00105 00105 00110	10¢ 11¢ C¢ 12¢ 13¢ 14¢ 1 15¢	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006 000006 000006 000006
L0101 U0103 U0104 U0105 U0105 U0110 L0112 U0122 L0125	10¢ 11¢ C ¢ 12* 13* 14* 1 15* 16¢ 9 17* 18*	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006 000006 000006 000006 000024 000024 000024 000032
L0101 U0103 U0104 U0105 U0105 U0102 U0121 U0122 L0125 U0141	10¢ 11* C* 12* 13* 14* 1 15* 16¢ 9 17* 18* 19* 20	INTEGER VEID(100) * ***********************************	000006 000006 000006 000006 000006 000006 000006 000006 000024 600024 600024 6000224 000032 000032
L0101 U0103 U0104 U0105 U0102 U0102 U0102 U0121 U0122 U0125 U0124 U0124	10¢ 11* C* 12* 13* 14* 1 15* 16¢ 9 17* 18* 19* 20 20*	INTEGER VEID(100) * ***********************************	U00006 U00006 U00006 U00006 U00006 U00006 U00007 U0007 U0007 U00024 U00032 U00061 U00061 U00061
L0101 U0103 U0104 U0105 U0102 U0102 U0102 U012 U0122 U0125 U0141	$ \begin{array}{c} 10 \\ 11 \\ 12 \\ 12 \\ 13 \\ 14 \\ 15 \\ 16 \\ 7 \\ 16 \\ 16 \\ 9 \\ 17 \\ 18 \\ 19 \\ 20 \\ 21 \\ 21 \\ 21 \\ 21 \\ 20 \\ 21 \\ 21 \\ 21 \\ 21 \\ 21 \\ 21 \\ 21 \\ 21$	INTEGER VEID(100) * ***********************************	U00006 U00006 U00006 000006 000006 000006 000007 U0007 U00024 U00024 U00052 U00061 U00061 U00061
U0101 U0103 U0104 U0105 U0102 U0112 U0121 U0122 U0125 U0141 U0141 U0142	10¢ 11¢ C ¢ 12° 13° 14° 1 15* 16¢ 9 17* 18* 19° 20° 20° 21° 21* 22*	INTEGER VEID(100) * ***********************************	U00006 U00006 U00006 000006 000006 U00006 U00007 000024 000032 U00052 U00051 U00061 U00061
L0101 U0103 U0104 U0105 U01012 U0121 U0122 U0125 U0141 U0142 U0142	10¢ 11¢ C * 12* 13* 14¢ 1 15* 16¢ 9 17* 18* 19* 20 20* 21* 22* 23* 7	INTEGER VEID(100) * ***********************************	U00006 U00006 U00006 000006 000006 U00006 U000024 C00024 C00024 U00032 U00061 U00061 U00061 U00061 U00061 U00061 U00061
L0101 U0103 U0104 U0105 U01012 U0121 U0122 U0125 U0141 U0142 U0142 U0142	10¢ 11¢ C * 12* 13* 14* 1 15* 16¢ 9 17* 18* 19* 20 20* 21* 22¢ 23* 7 24*	INTEGER VEID(100) ***********************************	U00006 U00006 U00006 U00006 000006 U00006 U00024 C00024 C00024 U00032 U00061 U00061 U00061 U00061 U00073 U00073
L0101 U0103 U0104 U0105 U01012 U0121 U0122 L0125 L0141 L0141 L0142 L0143	10¢ 11¢ C * 12* 13* 14* 1 15* 16¢ 9 17* 18* 19* 20 20* 21* 22* 23* 7 24* 25* 10	INTEGER VEID(100) ***********************************	U00006 U00006 U00006 U00006 U00006 U00006 U00024 C00024 C00024 U00032 U00051 U00061 U00061 U00061 U00061 C00073 C00073 U00112
L0101 U0103 U0104 U0105 U01012 U0121 U0122 U0125 U0141 U0142 U0142 U0142	10¢ 11¢ C * 12* 13* 14* 1 15* 16¢ 9 17* 18* 19* 20 20* 21* 22¢ 23* 7 24*	INTEGER VEID(100) ***********************************	U00006 U00006 U00006 U00006 000006 U00006 U00024 C00024 C00024 U00032 U00061 U00061 U00061 U00061 U00073 U00073

000112 000112

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00157

29*

READ (5.7) (VEIU(K), K=1. ISTUP)

FURH2			DA	<u>TE 673080</u>	PAGE	2
00165 30* 10166 31* END FOR	RETURN END			000125		
àHDG,P ID					. <u></u>	
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1	4.	

éFOR,S 10,ID HSA E3 -07/30/80-09:45:34 (7.) ENTRY POINT UDU045 SUBROUTINE ID STORAGE USED: CODE(1) DUC057; DATA(g) DUC020; BLANK COMMON(2) DUC000 COMMON BLOCKS: UD113 HANDEL DC0172 EXTERNAL REFERENCES (BLOCK, NAME) NERR25 ដាំងដូ C0C5 NERR35 STORAGE ASSIGNMENT (BLOCK, TYPE, RELATIVE LOCATION, NAME) **UUU1** JL0015 1L 6061 000032 100L <u>0001 000005 1066</u> 0001 000021 21 0001 nnn025 50L 2006 000001 7F DUD007 INJPS U000 I 000000 K Unno Uno3 I DUDUDO KVCMID 0003 I 000024 KVCPID ចាព្រំ 3 DUD175 HAXMID 6663 000171 MAXPID UG101 1# SUBROUTINE ID(KIND, NDIFID, IDIN, IDOUT) 000005 60103 2* COMMON/HANDEL/KVCHID(20).KVCPID(100).MAXMID.MAXPID. 000005 00103 3 ¢ C ** KIND . INPUT . 1=MID . 2=PID 000005 60103 4 * C ** NDIFID.INPUT. THE NO OF ELS. IN KYCHID OR KYCPID 000005 IDIN. INPUT. THE ORIGINAL MTD OR PID CD103 5+ C ** 000005 L0103 C ** IDOUT.OUTPUT, MODIFIED MID OR PID 6* 000005 60103 7+ С 000005 00103 8≉ 000005 60103 9.0 С PRINT 7,KIND,NDIFID, IDIN 606005 60104 10+ 7 FORMATI * KIND, NUIFÍD, IDIN, INID * 318) 000005 11* DO 100 K=1,NDIFID 60105 000005 00116 12* 60 TO (1.21.KIND 000005 66111 13* 1 IF(IUIN_EQ_KVCHID(K)) GO TO 50 000015 L0113 14 = GO TO 100 006017 60114 15* 2 IF(IUIN.EQ.KVCPID(K)) GO TO 50 000021 00116 16# GO TO 100 600023 60117 17* 50 IDOUT=K 000025 00120 18+ RETURN 000026 100 CONTINUE 60121 19* 000056 -60123 23* ĔND 600056 END FOR aHDG P MAIN

0000 0000 0000 0000 0003 0003	I 002721 I 002676 I 002647 I 002646 I 002652 I 000022	L3 NELTYP	<u> пор I 062667 LBUCK 0000 I 002712 LIM поро I 002672 L1</u> 0030 I 062700 L4 0000 I 002704 L5 0000 I 002703 NCTOT	UCOO I 002674 L2 0000 I 002661 NCYCL
0000 0000 0000 0003 0003	<u>I 052647</u> I 052646 I 052652	NELTYP		0000 I 002661 NCYCL
0000 0000 0003 0003	I UL2646 I 002652			
0000 0003 0000	I 002652	ALCI TYC	<u>0000 I 002642 NELTY1 0000 I 002643 NELTY2 0000 I 002644 NELTY</u>	3 0000 I 002645 NELTY4
0003 0000	<u>I 002652</u> I 060022	-	ŬŨŪŊ I CU271Ū NĒND - ŪŪOO I UD2655 NIDVTO - ŪŪCŊ I DU2650 NIDV1	
0000	1 060022	NIDV3	<u> </u>	
			COUS I DECODO NUM 0000 I 002705 NUMDV DOGO I 002722 NUMNP	
- U(.)),	1 002671		0000 I 002673 NUM6 0000 I 002675 NUM7 0000 I 002677 NUM8	<u>0000 I 002701 NUM9</u>
	1 000000	R1	0060 R 000316 R2 0000 R 002506 STR	
	1*	C **	MAIN DRIVER	00000
0101	2*		COMMON/BACH/NUM(18),NSTART(18)	000000
.0101	3#	C ** f	FIRST DIMENSION OF R1 AND R2 MUST BE THE SAME AS	000000
_0101	4.*	C ** 1	THAT OF ELDATE AND ELDATE IN ELORD.	000000
0103	5.*		INTEGER R1(150,2)	000001
0104	6*	(DIMENSION R2(100,2)	000001
0104	7*		NO. OF DIFFERENT DESAP BAR DESIGN VARIABLES	600001
0105	8 *		UIMENSION IDV1(10D)	000001
0105	9*	<u>Ç ** </u>	NO, OF DIFFERENT DESAP BEAM DESIGN VARIABLES	000001
0106	10*		DIMENSION IDV2(100)	600001
<u>C106</u>	<u>11*</u> 12*		NO, <u>CF DIFFERENT DESAP QUAD DESIGN VARIABLES</u>	00001
0107	12*		DIMENSION IDV3(100)	
.011L	14#		NO. UF DIFFERENT DESAP SHEAR DESIGN VARIABLES DIMENSION I(V4(107)	
.0110	14+		NO, OF DIFFERENT DESAP OF DESIGN VARTABLES	690961 600991
,0111	16*		DIMENSION IDV5(100)	600001
0111	17*		TOTAL NO. OF DIFFERENT DESIGN VARIABLES	600001
0112	18*		DIMENSION IDV(150),A(150,2)	600001
.0.112	19*	<u> </u>	NO. OF STRUCTURAL LOAD MULTIPLIERS	
.0113	¥ن 2		DIMENSION STR(15,4)	000061
.0113	21*		***************************************	00001
0114	22		INTEGER DUMM(4),BUCK1(4)	00001
0115	23*		DIMENSION HEAD(20).BUCK2(3)	000001
0116	24#		DATA DUM/0./NELTY1,NELTY2,NELTY3,NELTY4,NELTY5,NELTYP,	600001
0116	25*		NIDV1.NIDV2.NIDV3.NIDV4.NIDV5.NIDVT0/12+D/DUHM/*_COM*/	<u> 400001</u>
0135	26*		WRITE (6,50)	000001
.0137 .014u	<u>27*</u> 23*		FORMAT (15X, 'ENTER JOB HEADING ON ONE LINE') READ(5,1) HEAD	<u>UODOD6</u> UOUOD6
.0145	20≠ <u>∠9</u> ≠		FORMAT(2UA4)	UCOD16
.0143	<u>+ ب</u> غن¢		PLACE CARD I ON FILE TO DESAP	000016
0144	31×		URITE_(12.1)HEAD	
0144	32*		CONVERT NASTRAN MASS DENSITY TO DESAP WEIGHT DENSITY	000016
<u>a147</u>	<u></u>		NRITE (6.10)	00026
0151	34*		FORMAT(15X, 'ENTER UNITS CODE: ENGLISH=1, HKS=2')	00033
0152	354		KEAD (5.7) IVCODE	600033
0155	36 \$		CONFAC=1.	600041
0156	37*		1F(1VCUDE.EC.2) CONFAC=9.8140 -	
0166	38*		FORMAT()	606050
0161	39#		CALL ORDER	<u></u>
0162	4U#		RITE(6,63D)	600052
0164	41*		EORMAT(15X, 'ENTER MAX NO.OF REDESIGN CYCLES, NCYCL 1C=INITIAL DESIG	L00057
.D164	42* 43*		V ONLY)'/10X'AND MAX• NO• SCALINGS,NSCALE (DEFAULT=3)• SEPARATE BY COMMA ONLY')	600057 606057
0165	444		READ (5,7) NUYCL,NSCALE	
	45 <u>*</u>		IF (NSCALE.EU.U) NSCALEIX	00066

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60173	46*	WRITE (6,601)	<u>uco</u> 072		
00175	47*	601 FORMAT (15x, "ENTER SCALING CODE, KSCALE")	000077		
00176	48*	READ 15,71 KSCALE	00077		
60201	49*	WRITE (6,602)	000105		
<u>CO2O3</u>	<u> </u>	602 FORMAT (15X, 'ENTER DELTAIDEFAULT=0.05).AND EPSILON (DEFAULT=0.1)'/	000112		
00203	51*	115X "SEPARATE BY COMMA ONLY")	000112		
60204	52*	REAU (5,7) DELTA,EPSIL	000112		
00210	53*	IF (UELTA-EQ. 0.) DELTA= .05	000121		
<u>60212</u> 80214	<u> </u>	IF (EPSIL-EQ. 0.) EPSIL = >1 WRITE (6,603)	000125		
	56*		000131		
<u>00216</u> CC216	57*	603 FORMAT (15X, *ENTER CODES FOR NODAL DISPL. PRINTOUT KPRINT (0=NO PR 1INT,1 = PRINT)*/10X*AND BUCKLING CONSTRAINT LBUCK. SEPARATE BY SPA	<u>000136</u>		
60216	58*	2CE AND/OR COMMA']			
60217	59*	READ (5,7)KPRINT,LBUCK	600136		
00223	60.	KPUNCH=0	000145		
60224	61*	WRITE(12,611)	600146		
60220	62*	611 FORMAT(* COMH DESIGN CONTROL FROM MAIN*)	000153		
00227	63*	WRITE (12,610)NCYCL,NSCALE,KSCALE,DELTA,EPSIL,KPUNCH,KPRINT,LBUCK	000153		
60227	64=	<u>C PRINT 114</u>	000153		
üG227	65*	C114 FORMAT(' IN MAINNCYCL+NSCALE+KSCALE,DELTA+EPSIL+KPUNCH+KPRINT+LB	000153		
00227	66#	<u>c 1uck•)</u>	000153		
L0227	67.≠	C WRITE (6,610)NGYCL,NSCALE,KSCALE,DELTA,EPSIL,KPUNCH,KPRINT,LBUCK	000153		
60241	<u>68</u> ≠	610 FORMAT (315.2F10.3,315)	<u>üQQ17c</u>		
60242	69*	CALL NODPT	000170		
60243	70+	NUH5=NUH153	000172		
60244	71*	IF(NUH5.NE.O) PRINT 15	000174		
<u>UU247</u>	72*	15 <u>FORMATI</u> ************************************	000202		
LO247 UO247	73*	1,/* *'/* DESAP BAR ELEMENT (NASTRAN ROD)*,/,	000202		
<u>00247</u> 00250	<u>74</u> #	2**/******************************	000202		
60252	76÷	NELTYP-NELTYPINELTYI	000202 00n212		
60253	77*	IF (NUM5.EQ.0) GO TO 622	000215		
00253	78*	C PRINT 333, NIDV1, NELTY1, NELTYP	000215		
00253	79#	C333 FORMAT(' NIDV1, NELTY1, NELTYP=', 3(2X, 14))	000215		
60253	80*	C PRINT 666	000215		
60253	81*	C666 FORMAT(+ LOOKOUT HERE COMES THE IDV ARRAY)	000215		
00255	82*	UO 821 L1=1,NIUV1	000224		
00250	63×	821 IDV(L1)=IDV1(L1)	000224		
<u></u>	84*	C821 PRINT 444, I(V(L)	<u></u>		
00263	65×	C444 FORMAT()	600224		
<u> </u>	86*	622 NIUVIO=NIDVIO+NIDVI	UD0227		
00262	87¥	C PRINT 555, NIDVTO	606227		
<u></u>	<u> </u>	<u>C555 FORMAT(* NIDVTO=*,14)</u>			<u></u>
LU263	894		000231		
<u>60263</u> 00263	<u> </u>	C MIKE=0 C 1F(MIKE+EQ+D)G0 TO 2000	000231		
00264	92*		000231		
<u>00264</u> 00267	<u> </u>	<u>IF(NUM6+NE+G) PRINT 20</u> 20 FORMAT(************************************	000233		
60267	93+ 94+	1./ * * / * DESAP BEAM ELEMENT (NASTRAN BAR) * / .	000241		
L0267	95*	2* **/* *******************************	U00241		
60270	96.	IF(NUMD.NE.D)CALL BEAM(CONFAC.NELTY2.IDV2.NInV2)	000241		
00272	97*		Ung254		
00273	98*	IF (NUM6.60.0) 60 TU 624	000257		
00273	¥9*	C PRINT 332, NIDV2, NELTY2, NELTYP	000257		
60273	160*	C332 FORMAT(* NIDV2, NEL TY2, NEL TYP=*, 3(2X, 14))	000257		
CU273	1J1#	C PRINT 667	000257		
00273	162*	C667 FORMATI LOOKOUT HERE COMES THE IDV ARRAY"	600257		

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60275	1:13*	D0 623 L2=1.NIDV2	000270	·	
UU 300	164+	623 IDV(L2+NIDVT0)=IDV2(L2)	Ú00270		
60300	105*	C623 PRINT 444. IUV(12*NIDVTO)	000270		
û03 ₀ 2	1ü6*	624 NIDVTO=NIDVTO+NIDV2	000273		
<u>6302</u>	<u>107*</u>	C PRINI 555 NIDVIO	000273		
60302	168+	C HIKE=0	000273		
<u>LD302</u>	149*	C IF (MIKE, EQ. 0) GO TO 2000	<u> </u>		
60303	110*	NUH7=NUH(7)	600275		
<u>10304</u> 00307	<u> 111# </u>	<u>IF(NUM7.NE.U) PRINT 25</u> 25 FORMAT(* ***********************************	000277		
60307	113*	1,/ * * / * DESAP QUAD ELEMENT (NASTRAN ODMEM) * . / .	000305		
60307	114#		000305		
60316	115=	IFINUM7.NE.DICALL GDMEN (CONFAC.NELTY3.IDy3.NIDy3)			
60312	116#	NELTYP=NELTYP + NELTY3	000320		
60313	117#	IF(NUM7, E0.n) 60 TO 626	000323		
L0315	118#	D0 625 L3=1,NIUV3	000334		
60320	<u>119*</u>	625 IDV(L3+NIDVTO)=IDV3(L3)	<u> </u>		
u0322	123*	626 NIDVTO=NIDVTO+NIDV3	000337		
<u> </u>		NUM8=NUM(8)	000341		·····
60324	122#	IF(NUM8.NE.D) PRINT 30	000343		
<u> </u>	123*	<u>30 </u>	000351		
LU327	124 +	1,/* **/* * DESAP SHEAR ELEMENT (NASTRAN SHEAR)*,/,	000351		
0327	125*	<u>2' *'/' ********************************</u>	000351		
U033u	126*	IF (NUMB.NE.D)CALL SHEAR (CONFAC, NELTY4, IDV4, NIDV4)	000351		
<u> </u>	1277	NELTYP=NELTYP+NELTY4	600364		
60335	128× 129×	1F(NUM8.20.0) GO TO 627	600367		
<u></u> LU346	<u>13</u> 0¥	628 L9=1+NIDV4 628 IDV(L4+NIDVT0)=IUV4(L4)	606400		
	131*	<u>627 NIDVIO=NIDVIO+NIDV4</u>	COG400 COU403		
60342	152+	C2DCU CONTINUE	000403		
En 343	133*				
66344	134×	NUH10=NUH(10)	000407		
UC 345	135*	NCTOT=NUM9+NUM10	000411		
ü()346	136 \$	IF(NCTOT.NE.U) PRINT 35	006413		
60351	137#	35 FORMATE ************************************	000421		
Ú0351	138*	1,/* ≠*/* ≠ DESAP QIT ELEMENTS (NASTRAN GUAD2 & TRIA₂)*,/,	000421		
60351	<u>139*</u>	<u>2' *'/' ********************************</u>	600421		
L0352	140*	IF (NCTOT.NE.U) CALL OT (CONFAC, NELTYS, IDV5, NIDV5)	UD0421		
60354	141*	NELTYP=NELTYP+NELTYS	600434		
លប355 ប្រ0357	142*	IF(NCTOT_EG.L) GO TO 629	000437		
<u>UU357</u>	<u>143</u> * 144*	630 IDv(L5+NIDV50)=10V5(L5)	000450		
LU362 LU364	145+	639 IDV(LS+NIDVIO)=IDV5(L5) 629 CONTINUE	000450 000453		
LD365	146*		000453	·····	
60366	147*	NIDVIG=NIDVIG+NIDV5	<u>600453</u>		
00367	148*	CALL DIF(NIDVTO,IDV,NUMDV)	<u> </u>		
Un367	149*	CPRINT_111	000455		
LU367	150*	C111 FORMAT(" IDV ARRAY AND NUMDY IN MAIN AFTER RETURN FROM DIF")	600455		
LG361_	151*	C PRINT 114. (IDV(I),I=1.NIDVTO)	000455		
LO367	152*	C112 FORMAT(2X,13)	000455		
60367	153*	C PRINT 113. NUMDY	000455		
60367	154*	C113 FORMAT(* NUMUY =*,13)	000455		
60370	155*	PRINT 40	000462		
UG 372	156≠	40 FORMAT(/// END OF ELEMENT PROCESSING SET OTHER CONTROL VARIABLE	000466		
60372	157+	<u>15*/)</u>			
00373	158#	WRITE (6,760)	600466		
	159*	700 FORMAT (20X, STRUCTURAL LOAD MULTIPLIERS 1/2CX, SENTER NO. OF LOAD	600473		

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0 ₀ 375	160*	1CONDITIONS •)	000473		
60376	161*	READ (5,7)NLOAD	U00473		
66401	162*	U0 750 L1=1,NLOAD	000501		
00404	163*	00 750 L2=1,4	000510		
60407	164*	750 STR(L],L2]=0.	000510		
60412	165*	DO 751 LI=1,NLOAD	000520		
60415	166 *	WRITE (6,701)L1	000523		
60420	167#	701 FORMAT (2X, "FOR LOAD CONDITION", 14, " ENTER STRUCTURAL LOAD MULTIPL	000531	· · · · · · · · · · · · · · · · · · ·	
60420	168*	IIERS FOR ELEMENT LOADS A, B, C, D, IN THAT ORDER. DEFAULT TO ALL ZEROS	000531		
00420	169*		C00531		
00421	170+	READ (5,7) DUHH(1)	000531		
L0424	171*	IF(DUMM(1).EU.D) GO TO 751	000537		
CD426	172*	READ(5,7) (DUMM(L2),L2=2,4)	000541		
00431	173#	UO 752 L2=1,4	400556		
60434	174*	752 STR(L1,L2)=DUMM(L2)	600556		
60436	175*	751 CONTINUE	000563		
00436	176*	C PRINT 115	000563		
60436	177*	C115 FORMAT(* IN HAINSTR ARRAY*)	000563	• • • • • • • • • • • • • • • • • • • •	
<u> </u>	178*	WRITE(12,759)	000563		
	<u> </u>	759 FORMAT(* COMM STRUCTURAL LOAD MULTIPLIERS FROM MAIN+)			
UU442 C0443	180*	DO 755 LI=1+NLQAD	600573 006573		
Li0445	181+	755 wRITE (12,702)(STR(L1,L2),L2=1,4)	C00573		
60446	181#				
		C755 WRITE (6,7G2)(STR(L1,L2),L2=1,4)			
68455	183*	7UZ FORMAT (4F1C.3)	000607		
00455	164#	C ** BUCKLING CONIROL CARD	600607		
üU456	185*	IF(LBUCK.EQ.O) GO TO 850	000607		
<u>C0460</u>	166#	WRITE (6.703)	600611		
00462	187*	703 FORMAT (10X, "BUCKLING CONTROL CARD")	000616		
<u></u>	158*	WRITE (6,704)	600616		
60465	189*	704 FORMAT (2x, "ENTER COEFFT. DEFAULT=1.")	000623		
<u> </u>	190*	<u>BUCK2(1)=1.</u>	000623		_
60467	191*	READ (5,7) DUH	000625		
60472	192*	IF [DUH=NE=0]BUCk211)=DUM	000633		
60474	193*	WRITE (61705)	000637		
00476	194*	705 FORMAT 12X, "HOUEIN HAS BEEN SET TO 0"/}	000644	- '	
00477	195*	BUCK1(1)=0	000644		
60500	196*	WRITE (6,706)	<u>CO0645</u>		
60502	197*	706 FORMAT (2X, "ENTER NMODE . DEFAULT=1")	600652		
<u> </u>	<u> 198</u> *		00652		<u>.</u>
60504	199*	READ (5,7)DUM	C00654		
60507	<u>200*</u>	IF (DUM.NE.D)BUCK1(2)=DUM	000662		
10511	261+	WRITE (6,707) 707 FORMAT (28 SENTED INDET AND NUMED IN THAT ODDEDIN	000673		
<u></u>	262*	707 FORMAT 12X. *ENTER INDET AND NVEC.IN THAT ORDER*)	000700		
60514	263*	READ (5,7) BUCK1(3), BUCK1(4)	000700		
<u>40520</u>	204*	WRITE (6,708)	<u> </u>		
00522	205*	708 FORMAT (2X, *ENTER ALPA*)	000714		
<u>U0523</u>	<u>206</u> * 207*	READ (517) BUCK2(2)	606714		<u> </u>
L0526	≉/⊔∠ ≉8ن2	wRITE $(6,709)$	600722		
60536	208* 209*	709 FORMAT (2X. "ENTER OMEGA . DEFAULT = 0.8")	600727		
00531		BUCK2(3)=.8	600727		
00532	210*	READ (5,7) DUM	000731		
00535	211*	IF (DUH+NE+0) BUCK2(3)=DUH	600737		
<u> </u>	212*	wRITE (6,710)	000743		
60541	213*	710 FORHAT (2X, "ENTER SHIFT. DEFAULT=J.")	000750		
<u> </u>	214	READ (5.7) BUCK2(4)	000750		
00542	215*	C PRINT 116	60 <u>c</u> 750		
U0542	216*	C116 FORMAT(' IN MAINBUCK2(1), BUCK1 ARRAY, BUCK2 ARRAY')	000750		-

MA	IN		DATE 073080	PAGE	6
L0545	217*	WRITE(12,713)	000756		
60547	218*	713 FORMATI' COMM BUCKLING CONTROL FROM MAIN')	000763		
00550	219*	WRITE (12,711) 8U-K2(1), (BUCK1(T), I=1,4), (BUCK2(I), I=2,4)	000763		
6055 u	220*	C mRITE (6,711)BUCK2(1),(BUCK1(I),I=1,4),(BUCK2(I),I=2,4)	000763		
_ <u>60555</u> _	221*	711 FORMAT (F10-3,415,3F10-3)	001000		
00556	222*	85n CONTINUE	001000		
60556	223*	C ++ NODAL LOAD DATA	001000		
60557	224*	IF(NUM(3)+EQ+D) GO TO 950	001000		
<u> </u>	225*	REWIND 11	001001	·	
60562	226*	NEND=NSTART(3)-1	001004		
<u> </u>	227.*	DO 925 K1=1+NEND	601067		
LU566	228+	925 READ(11,1)DUH	001012		
<u>u0572</u>	229*	901 FORMAT (/)	001020		
60573	230*	LIM=NUM(3)	601020		
<u> </u>	231*	00 930 K2=1,LIM		<u>. </u>	
60577	232*	930 READ(11,902) R1(K2,1), (R2(K2,K3),K3=1,3)	001030		
<u> </u>	233*	9U2 FORMAT (16X.18.16X.3Fg.1) C *+ SET ALL LOAD CONDITION NUMBERS=1. THIS MAY BE HODIFIED BY HAVING THE USE			
006 <u>0</u> 7 00607	234* 235*	C ## SET ALL LUAD CONDITION NUMBERS-1. THIS HAY BE MUDIFIED BY HAVING THE USE C ## ENTED THE LOAD CONDITION FOR EACH NODAL LOAD CARD	R 601051 01051		
CC610 L0613	236* 	D0 935 K4=1+LIH 935 R1(K4+2)=1	001051 601051		
<u>UU013</u>	238*	C ## SET ALL CONCENTRATED NODAL MOMENTS TO ZERO. THIS MAY BE MODIFIED BY HAVI	NG 001051		
	239*	C ++ SET ALL CONCENTRATED NODAL HOHENTS TO ZERO. THIS MAT BE HODIFIED BY HAVI	001051		
<u> </u>	<u> </u>	DO 936 K5=1,LIM	001057	·	
<u> </u>	240+	<u>10 936 K6=4.6</u>	001057		
L0623	242*	936 k2(k5,k6)=0.	U01057		
LL626	242+	CALL ELORD(LIM,R1,R2,2,6)	001064		
LC626	244=	C PRINT 117	601064		
LD626	245=		001064		
60627	246*	write(12,913)	001073		
60631	247≠	U0 940 K7=1.LIM	001100		
10634	248*	913 FORMATE COMM NGDAL LOAD DATA FROM MAIN+)	601106	·	
63635	249*	940 HRITE (12,903) (R1(K7,K8),K8=1,2),(R2(K7,K8),K8=1,6)			
66635	25J¥	C940 wRITE (6,903) (R1(K7,K8),K8=1,2),(R2(K7,K8),K8=1,6)	001106		
60650	251+	943 FORMAT (14.13.6E12.5)	001132		
u0651	252×	950 WRITE (12,904)	001132		
60653	<u>253</u> *	904 FORMAT(* *)	001136		_
60653	254*	C ** DESIGN VARIABLE DATA	601136		
60654	255*	#RITE (6.100U)	001136		
U l656	256*	1000 FORMAT (15%, DESIGN VARIABLE DATA"/)	001143		
<u>LU657</u>	257.*	60 1653 K1=1,NUMBY	601143		
66657	258×	C = A(K1,1) = 1	001143		
<u>UD657</u>	259*	C1050 A(M1+2)=+1.	001143		
6662	260*	WRITE (6,1001)K1	001151		
	261*	1050 READ (5.7) A(K1.1), A(K1.2)	601157		
LG672	262*	1001 FORMAT (2X, "FOR DESIGN VARIABLE NO", 13, " ENTER AOLD AND AMIN, IN T	001174		
<u> </u>	263*		601174		
LD672	264.ÿ 265.≠	C PRINT 118	601174		
<u> </u>	<u>265*</u> 266*	C118 FORMATIC' IN MAINK1.AND K ARRAYAOLD AND AMIN'J Do 1.51 k1=1.Numdy	601174		
	265≠ 267≠		001174		
<u> </u>		1051 WRITE (12,1 <u>602) K1,(A(K1,K2),K2=1,2)</u> C1051 WRITE (6,1002) K1,(A(K1,K2),K2=1,2)	001174		
60676 60706	268* 269*	$\frac{1_{102} - FORMAT(15, 2F10, 3)}{2}$	UO1174 UO1213		
<u>60708</u>	270+	NUHNP=NUH(1)			
<u> </u>	¥	NUMNF_NUM(I) 	001213		
<u></u>	272*		001215		
<u>u0712</u>		REWIND 12	001220		
1.0712	273*	REAU(12.1) HEAD	001223		

MAIN		DATE 07308n	PAGE 7
u0715 274 +	WRITE(13,1) HEAD	001233	
L0720 275*	WRITE (13,1003) NUMNP, NEL TYP, NLOAD, NUMDV	CO1243	
L072U 276+	C PRINT 1004	601243	•
60720 277*	CIUD4 FORMAT(* NUMNPNELTYPNEQADNUMDV NEXT PRINT*)	001243	
£6726 278≠	C WRITE (6,1003)NUMNP,NELTYP,NLOÅD,NUMDV	001243	
66726 279*	1005 READ(12,1,END=2) HEAD	001255	
<u>00731 280</u> #	IF(HEAD(1)+E4+DUHM)G0_T0_1005	001265	
00733 281*	WRITE(13,1) HEAD	C01270	
00736 262*	<u>60°TO 1005</u>	001300	
00737 263*	1003 FORMAT (415)	601302	
60740 284*	Z STOP	001302	
LO741 285+ ENU FOR	END	001305	
UHDG,P NAME			
APRT.S NAME			
FURPUR 28R1H1 E3	36 S74T11 07/30/80 09:45:45		
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NOUPT

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WFOR, S NOUPT, NOUPT

HSA E3 -07/30/80-09:45:46 (27,)

SUBROUTINE NOUPT ENTRY POINT GOU457

STORAGE USED: CODE(1) 000472; DATA(U) 015130; BLANK COMMON(2) 000000

COMMON BLOCKS:

UDU3 BACH UGOD44

EXTERNAL REFERENCES (BLOCK, NAME)

<u> </u>	NREWS	
5LÚJ	NEDUS	
<u></u> 6	NIO15	
0007	NIOZS	
6010	N1035	
0011	NEDUS	
	NERR35	

3	STORAGE ASSIGNMENT (BLOCK, TYPE, RELATIVE LOCA	TION, NAME)			
X —	0001 00017 116G	CCU1 00025 123G	0001 000031 127G	U001 00051 137G	U001 000053 142G	· · · · · ·
7	<u></u>	<u>CUU1 CG0102 1753</u>	<u>0051 000134 1776</u>	0001 000223 1853L	<u> </u>	
ట	LUJ1 - ÜJ0154-2136	0001 000171 223G	0001 UD0174 226G	0061 000216 241G	U061 060240 2566	
<u> </u>	<u> 2001 000242 2616</u>	60ú1 00n254 2676	0001 000312 314c	0001 000331 3246	0001 000343 3356	
	COO1 060351 340G	UDU1 000353 343G	0001 000365 3516	0001 000410 3676	UDU1 0C0415 373G	
	<u></u>	LUND 015062 5F	U001 000370 61L	0001 CH0360 62L	6000 015051 700F	
	0000 015055 701F	COUD 015/156 702F	U001 000057 751L	0001 0001n5 752L	0001 000107 753L	
	UUJ1 GUD115 765L	COUL 020200 851L	0001 nnu226_652L	C001 000230 853L	0001 nnn234 864L	
	0001 000265 881L	CG01 000247 882L	0000 015073 901F	0000 015064 903F	0000 015061 910F	
	<u></u>	CCU1 0004n3 959L	1000 I 000000 cR0n	0000 R 0150n1 DUM	0000 I 015030 I	
	0000 I 015631 II	LOUD 015110 INJPS	UCOD I 015032 J	0000 I 015036 JJ	0000 I n15033 JK	
	0000 I 015027 к		1000 I D15035 L	0000 I 015045 L1	0000 I 015047 12	
	1000 I 015050 L3	UNUD T U15046 L4	UDUD I U15040 M	UDUO I 015037 M1	UOLO I 015044 NCROD	
	UUUU_I 015143 NEND	<u>0000 I 015034 NEND3</u>	0000 I 005524 NODPT1	0003 I 000022 NSTART	6003 I 000nng NUM	
	LUCH I U15025 NUMNP	LUUD I U15026 NUM4	1000 I 000310 SHAR1	UCCO p 011610 SHAR2	0000 I 013414 SPC1	
		6030 I 014772 V				

LU101	1 *		SUBROUTINE NODPT	000000	
60105	2*		COMMON/BACH/NUM(18),NSTART(18)		
60103	3 ×	C ##	CHANGE GROUP. NO. OF DESAP BAR ELS. (NASTRAN ROD)	000000	•
UG104	4 *		INTEGER_CROD(110,2)	60000	''
60104	5 *	() * *	CHANGE GROUP. NO. OF NODAL POINTS	ենըընն	
00105	6\$		INTEGER SHARI(300,9)	000000	
u0106	7*		UIMENSION NOUPT1(300,7), SHAR2(300,3)	000000	
60106	8*	C **	CHANGE GROUP. NO. OF SPCI CARDS	000000	,
60107	9*		INTEGER SPC1(50,15)	000000	
00107	1J+	C **	**********	000000	

NOL	UP T		DATE UT3080	PAGE	2
60110	11*	INTEGER V(7)	000000		
C0111	12*	DIMENSION DUM(20)	000000		
00112	13*	REWIND 11	000000		
00113	14+	NUMNP=NUM(1)	000002		
60114	15*	NUH4=NUH(4)	000005	· <u> </u>	
L0114	16*	C ++ CREATES NODAL POINT DATA	000005		
60114	<u>17*</u>	<u>C.** WHEN NUH141=U (NO SPC1 CARDS).READS GRID CARDS FROM FILE 11.STORES IN</u>	000005		
60114	18#	C ** ARRAYS SHAR1 AND SHAR2 AND WRITES NODAL PT. DATA CARDS ON FILE 12	000005		
60114	19*	C ** SHAR1 CONTAINS NODE NO. AND NASTRAN CONSTRAINTS	<u> </u>		
60114 60115	∠0* 21*	C ** SHAR2 CONTAINS CORRESPONDING COORDINATES (SAME ROW NO.)	000005		
L0126	22*	D0 760 K=1.NUMNP READ (11,700) SHAR1(K,1),(SHAR2(K,I),I=1,3),(SHAR1(K,I),I=2,9)	600017		
<u>i0120</u>	23*	<u>C PRINT 700, SHARI(K,1),(SHAR2(K,1),1=1,3),(SHAR1(K,1),1=2,9)</u>			
L0133	24+	760 NOUFT1(K.1)=SHAR1(K.1)	000034	· · · ·	
60135	25≠	700 FORMAT(8X+18+8X, 3F8+4+8X, 811)	000053		
00135	26*	C ** NODE PT. NO. IS NOW NOOPTI(K,1)	000053		
L0135	27*	C ++ NODPTI CONTAINS DESAP NODE NO. AND 6 HOTION CODES	000053		
60135	28*	C ** INITIALIZE DESAP CONSTRAINT CODES TO ZERO	000053		
<u>u0136</u>	<u>29</u> *	00 765 K=1.NUMNP	00053		
60141	30¥	VO 766 I = 2,7	000053		
00144	<u> </u>	766_NODP11(K.I)-G	000053		
00144	32*	C ** II COUNTS THE 8 COLUMNS OF SHAR1 (K, 2 TO 9)	000053		
UC146	33*		600054		
00147 20150	34 * 35 ±	751 II=1I+1 IF (II.6T.9)60 T0 765	600057 600061		
L0150	<u> </u>	C ** CHECK IF COLUMN IS BLANK	00 ₀₀ 61		
00152	37*	IF(CHAR1(K,II),EQ,D) = GO TO 751	600064		
00152	38*	C ** J IS THE NASTRAN RESTRAINT CODE. J=1 MEANS UX=0 .J=6 MEANS DX=0	000064		
L0154	39*		600075		
00157	43#	IF(SHAR1(K,II),E0,J) G0 TO 1753	000075		
00161	41#	<u>60 TO 752</u>	000100		
69162	42×	1753 JK=J	600102		
00163	43*	<u>60 to 753</u>	000103		
ÚD164	44 +	752 CONTINUE	000107		
00166	45*	753 NODPT1(K.JK+1)=1	000107		
60167	46*	GO T _O 751	000113		
<u>UC170</u> U0170	<u>47*</u>	<u>C PRINT 1800</u>	000121	· · · · · · · · ·	
60170 60170	49.4	C1800 FORMAT(1X, 'NODPT: JUST BEFORE SPC1 CARDS')	000121		
60170	<u>+,+</u> 5(;+	C 00 1801 K:1.NUMNP	<u> </u>		
60170	51*	<u>C1801 pRINT 910.1N00PT1(K.I].I_1.7)</u>	000121		
00170	52*	C ** IF THERE ARE SPC1 CARDS, NOUIFY NODP11 ARRAY	000121		
<u>00172</u>	<u>53*</u>	IF (NUM(4).EL.0160 TO 956	000121		
L0172	54×	C ** READ SPC1 CARDS FIELUS 2 THROUGH 9 AND STORE IN SPC1	000121		
L0172	<u> </u>	C AN SPCI(K,2] THROUGH SPCI(K,9) CONTAIN CONSTRAINT CODES 1 THROUGH 6	000121		
60174	56*	REWIND 11	000123		
<u></u>	57*	NEND3 = NSTART (4)-1	000126		
60176 60201	58* 59*	UO 770 L=1,NEND3 770 READ(11.701)UUM	C00131		
00205	<u> </u>	701 FORMAT (20A4)	<u> </u>	······································	
00205 0 <u>020</u> 6	61×	00 771 L:1.NUM4	000147		
00211	62#	771 REAU (11,7C2) (SPC1(L,J),J=1,15)	006147		
60220	63*	742 FORMAT(8X,18,811,618)			
60221	64*	REWIND 11	000163		
<u> </u>	65 <i>+</i>	C ** TRANSFER CONSTRAINT CODES FROM SPC1 ARRAY TO NODP11 ARRAY	<u>inD0163</u>		
LU221	66 ₽	C ** V IS STORAGE VECTOR	600163		
00221	67*	C ++ K INDEXES SPC1 RUW(SPC1 CARD)	000163		

NU	UPT			DATE 073080	PAGE	.3
U0222	68×		UO 865 K=1.NUM4	600174		
60225	69*		UO 86(1 L=1,7	600174		
00230	7:0≠	860	v(_)=D	000174		
60232	71*		I I = 1	000175		
00233	72*	851	11=11+1	000200		
60234	73×		IF (II.6T.9) GO TO 864	600202		
60234	74*	<u> </u>	CHECK IF COLUHN IS BLANK	000202		
60236	75*	(° + +	IF (SPC1(K,II) .E0.D)GO TO 851	000205		
<u>60236</u> 60240	<u>76*</u> 77*	(**	J IS THE NASTRAN RESTRAINT CODE.ETC ** DO 852 J=1.6	000205		
60243	78 ≠		IF(SPC1(K, II).EQ.J)60 TO 1853	LIDG216		
00245	79#		GO TO 652	U00221		
60246	៦ដ≠	1853	JJ=J	000223		
60247	61*		60 TC 653	U00224		
00250	82×	852	CONTINUE	000230		
UG252	63×	853	V (J+1)=1	Ú00230		
<u>40253</u>	84*		<u>GO TO 851</u>	<u> </u>		
00253	85×	C *∗	KTH SPC1 IS DONE. V CONTAINS 6 DESAP CONSTRAINT CODES FOR THAT K	. 000232		
60254				000234		
60254	87*		PLACE NEW CONSTRAINT CODES IN NODPT1	000234		
60254	<u>68*</u>	C	PRINT 908, K.(V(I).I=1.7)	000234		
60254	89*	66938	FORMAT(1X, *V FOR SPC1*, 13/5X, 715)	000234		
<u> </u>	<u> </u>		DQ 88D M1=10,15 D0 681 M=1,NUMNP	<u>LDD234</u> DDU242		
00280	92*		IF (NOUPI1(M.1).EQ. SPC1 (K.M.)) GO TO 8g2	600242		
L0265	<u> </u>	· · · · · ·		000245		
00265	94*	(**	NODAL PT. NO. IN NODPIL (M.1) IS REFERENCED ON SPCI CARD INDEXED BY K			
60265	95 *	C **	PLACE NEW CONSTRAINTS ON THIS NODAL PT.	600245		
60266	96*		U0 863 L=2,7	406247		
L0271	97×		IF(V(L).NE.0.) NUDPT1 (M,L)=V(L)	000254		
L0273	<u>98*</u>		CONTINUE	000273		
60273	59#		THE MTH NODPT1 CARD IS NOW COMPLETE	000273		
60273	1*	Ç	PRINT 909,M	000273		•
60273	101*	6909	FORMAT(1x, "NODPT1 CARD", 13/)	000273		
<u> </u>	<u>162*</u> 103*	<u> </u>	PRINT 910.(ACDPT1(H.L).L=1.7) FORMAT(715)	<u> </u>		
L0276	+دن⊥ ¢4ن4		CONTINUE	000273		
L0300	1,5 *		CONTINUE	600273		
60306	1654		ALL NOUPT1 CARDS ARE NOW COMPLETE	C00273		
<u></u>	107*		CONTINUE	000273	· · · · ·	
<u> </u>	108*		CONTINUE	600273		
60365	109*		KN=0	COD273		
60306	<u> </u>		_ <u>I</u> Ξg	600273		
UU 3D6	111+		HOUIFY HUTION CODES IF THERE ARE ANY NASTRAN RODS	000273		
<u> </u>	_112*	<u>(</u> **	BRING CRODS FROM FILE 10 AND GET THE GRID PT NO'S.	000273		
60306	113×	(; * *	FIND CURRESPUNDING NODE PT NO'S IN NODPTI AND ENFORCE ZERO ROTATIONS	000273		
<u>13117</u>	114#		REWIND 11	<u> </u>		
60310	115*		IF (NUMIS).EQ. U)GO TO 959	000277		
<u>60312</u> 60 ³ 13	<u>116*</u> 117*		<u>NEND=NSTART(5)-1</u> DO 56 L=1,NEND	000301		
66316	118*	50	READ [11,701]DUM	LC0306 LC0312		
60322	119#		NCROD=NUM(5)	600322		
60323	120+		UO_51 L=1.NCROU	000324		
60326	121*	51	READ (11,5) CROD(L,1),CROD(L,2)	606331		
60333	122*	5	FORMAT (24X, 216)	606343		
60334	123*		DO EU LI=1,NCROD	<u> </u>		
60337	124#		UO + 4 L4=1,2	606353		

NOD	PT			DATE 073080	PAGE	4
60342	125*		00 61 L2=1.NUMNP	000353		
60345	126*		IF (NODPT1(L2,1).EQ.CROD (L1,L4)) GO TO 62	L00353		
00347	127*		60 10 61	<u> </u>		
60350	128*	62	UO 63 L3=5,7	C00360		
60353	129*	63	NOUp11(L2+L3)=1	000365		
60355	130*	61	CONTINUE	000375		
40357	131*	64	CONTINUE			
L0361	132*	60	CONTINUE	000375		
60361	133*	C	PRINT 200	000375		~
60361	134*	C200	FORMAT(+ FINAL OUTPUT OF NODPT.") Combines Nodpt1 and Shar2 to form desap Nodal PT. Data	000375		
<u>00361</u> 00363	<u>135*</u> 136*	<u> </u>	WRITE(12,903)	<u> </u>		
60365 60365	137*	0.03	FORMATE COMM NODAL POINT DATA FROM NODPT"	000403		
LO366	138+	959	DO 960 K=1,NUMNP	C00403		
60371	139#	960	WRITE (12.901) (NODP11(K.I).I=1.7). (SHAR2(K.I).T=1.3).KN.T	000410		
00371	143*	0960	<pre>HRITE (12.901)(NODPT1(K+1)+I=1+7)+(SHAR2(K+1)+I=1+3)+KH+T PRINT 9D1,(NODPT1(K+I)+I=1+7)+(SHAR2(K+I)+I=1+3)+KH+T</pre>	000410		
60400	141*	9:11	FORMAT (715.3F10.3.15.F10.3)	000436		
ŨŪ407	142*		RETURN	000436		
<u>60410</u>	143*		<u>END</u>			
END FOR						
aHDG.P	URDEI	R				
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OFUR, S URDER, ORDER

HSA E3 -07/30/80-09:45:50 (29,)

SUBROUTINE ORDER ENTRY POINT DOC552

STORAGE USED: CODE(1) DOD561; DATA(0) DOD570; BLANK COHMON(2) DODDDD

COMMON BLOCKS:

3ر (1)	васн	600044					
 <u>5004</u>	HUZART	000334		 			
UU U5	HANDEL	CCU174	·				

EXTERNAL REFERENCES (BLOCK, NAME)

_	<u>u006</u>	VËC	
	6067	COUNT	
	0100	SIGMA	· · · · · · · · · · · · · · · · · · ·
	uU11	10	
_	u012	NKDUS	
	13 لال	NI035	
~	014	NIO25	
	u ü 15	NEDUS	
XX	<u> </u>	NERTS	
- i 🗸 -	L017	NREw\$	
- <u>-</u> -	<u></u> <u></u>	NWEFS	
4	u021	NI015	
•		NERR35	

STORAGE ASSIGNMENT (BLOCK, TYPE, RELATIVE LOCATION, NAME)

0001 000002 1176	6001 C00062 15GG	GOU1 000075 1576	0001 r00141 2046	0001 000177 2226	
ULU1 060235 2406	0001 060270 2566	<u>0ar1 000327 2746</u>	0001 000336 3016	Unnt 000350 3100	
uJu1 30C4 <u>Cú</u> 327G	6061 060417 340G	0001 00044p 3476	0001 000466 3576	0001 000476 3716	
6866 066560 400F	0000 0L0501 402F	6060 000503 405F	0001 nu0002 5L	0001 000042 5C1L	· · · · · · · · · · · · · · · · · · ·
0001 000115 502L	6031 060111 503L	0001 000162 507L	0001 000262 550L	0001 000314 552L	
UD01 000617 6L	0000 000517 604F	0001 0nn525 655L	0001 000514 657L	0000 0cn523 658F	
JUCU DUNSD5 749F	0000 000515 750F	0000 R 000000 A	CHOC R RUDU24 B	0000 R 000046 DUM	
	<u> </u>	UNUD I GOU455 ID1	0060 1 000456 102	000555 INJPS	
4890 I 360475 J	1000 I 000453 K	L000 I 000476 KHOAD	0000 I 000477 KM	0000 I 000472 KOUNT	
UUUS I JUDUOU KVCMID	LOUS I NLOUZE KVCPID	JC00 I 000470 L	0000 I 000473 LL	0000 I 000474 LLL	
6004 I 300006 MAT11	GUG4 R GUGUZ4 MAT12	UDG5 I BOD17C MAXMID	0005 I 000171 HAXPID	0000 I 000466 MID	
UUUU I UUU462 NEND	0000 I 000467 NEND17	000 I 000465 NPCARD	0000 I 000457 NREC	0000 T 000461 NRECH1	
UUU3 I OUDU22 NSTART	603 1 360600 NUM	UCUS I GOU172 NUMMID	0000 I 000454 NUMNP	0005 I 000173 NUMPID	
6000 I 100463 NUMRE	LUUU I DUD460 NUMREC	GCCC I 000471 NUM18	0000 I 000426 PID	6000 I non427 XMAT11	· '
GUDU R GUD116 XMAT12					,

0101	<u> </u>	SUBROUTINE ORDER	u00002	· · · · · · · · · · · · · · · · · · ·
00103	2*	COMMON/BACH/NUH(18),NSTART(18)/MOZART/HAT11(20),HAT12(20,10)/HANUE	00002	
60103	3*	<u>1_/KVCHID(2_),KVCPID(100),MAXHID,MAXPID,NUMMID,NUMPID</u>	600002	·

08	DER		DATE 073080	PAGE	. 2
L0104	4*	DIMENSION A(20),8(18),DUM(20),GRDMAX(20)	000002		
L0105	5*	DATA NUM/18+U/B/'GRID', 'OMIT', 'FORC', 'SPC1', 'CROD', 'CBAR', 'CUDM',	600002		
cn105	6*	1*CSHE+_+CQUA+.+CTRI+.+CONM+.+PROD+.+PBAR+.+PUDH+.+PSHF+.+PQUA+.	000002		
60105	7#	2*PTRI*, *HAT1*/KVCHID/20+0/KVCPID/100+0/HAXHID,HAXPID/20,100/	000002		
00114	8*	REAL MATIZ, XHAT12(20,10)	L00002		
00115	9*	INTEGER PID, XMAT11(20)	000002		
60115	10+	C ** WRITE ON FILE 10 IN THIS ORDER: GRID.OMIT.FORCMATL	600002		
60115	11 <i>≠</i>	C ** CALCULATE NO. OF EACH TYPE.E.G.,NUM(1)=NO. OF GRID CARDS,NUM(18)=NO. OF	000002		
un115	12*	C ++ MAT1 CARDS	000002		
L0115	13*	C **	600062		
60115	14#	C ** HAXMID(MAXPID) IS THE MAX NO. OF DIFFERENT MID(PID)NOS.	<u>CONOD2</u>		
60115	15*	C ## THE PRUGRAM CAN HANDLE. IF IT IS INCREASED, ALSO	000002		
60115	16*	C ** INCREASE DIMENSIONS OF KVCHTDIKVCPID) ACCORDINGLY.	600002		
00116	17+	DO 560 K=1,18	600002		
60121	18*	5 READ (9,400,END=501) A			
60124	19*	IF (A(1).EQ. B(K)) GO TO 6	000012		
60126	20*	<u>60 TO 5</u>	660015		
UD127	21*	6 WRITE (10,400)A	000017		
00132	22≠	PRINT 402. A	000026		
UC135	23*	NUH(K)=NUH(K)+1	600735		
60136	24=	60 TO 5	C00040		
u0137	25*	501 REWIND 9	000042		
60140	26*	500 _ CONTINUE	000046		
60140	27*	C PRINT 801	000046		
00140	28*	C 801 FORMAT(' NUM(1) THRU NUM(18) //)	000046		
C0140	29*	C PRINT 22, (NUH(K),K=1,18)	000046		
<u>00140</u>	30*	C22 FORMAT(1814)	600046		
L0142	31*	4 ₀₀ FORMAT (20A4)	000046		
L0143	32≠	4JZ FORMAT(1X.1944)	000046		
60144	33×	ENDFILE 10	600046		
60144	34≠	C ** NUMBER OF NOUAL POINTS			
L0145	35*		000051		
00145	<u>36</u> ≠	C ** CALCULATE STARTING RECORD NUMBERS IN FILE 10	000051		
LD146	37*	NSTART(1)=1	600053		
.00147	38*	D0 2c K=2+18	600062		
00152	39*	20 NSTART (K)=NSTART (K_1)+NUH (K-1)	000062		
60154	40*	RELIND 11	<u></u>		
00154	41*	C ** CALCULATE LARGEST NODAL POINT NUMBER AND CARD ON WHICH IT OCCURS NRFC **	600065		
U0155	42*	101=1	600070		
00156	43≠	UO 502 K=1,NUMNP	000075		
60161	44*		600075		
60164	45*	4U5 FORMAT (8X,16)	600103		
<u>U0165</u>	<u>46¥</u>	IF (102 .61. 101) 60 10 503			
60167	47 \$	5g2 01 00	000107		
<u> </u>	48*	503 101-102			
U0171	49*	NREC=K	090112		
00172	<u>50*</u>	502_CONTINUE	000116		
CO174	51*	NUMREC=NSTART(18)+NUH(18)-1 -	000116		
<u> </u>	<u>52*</u>	C. IF(NUM(1R) EU.) NUMRECENSTART(1R)	600116		
UD174	53×	C PRINT 23, NUMREC	L00116		
<u> </u>	54#	C23 FORMAT(+ NUMREC = +, 14)	000116		
00175	55*	IF(NREC +EQ+ NUMNP) GO TO 550	600122		
<u>00175</u>	56*	C ** PLACE RECORD WITH LARGEST NOUAL POINT NUMBER LAST ON FILE 11	000122		
60177	57×	REWIND 1G	000125		
62200	<u>58*</u>		600130		
60200	59×	C ** READ FROM 10 ALL RECORDS BEFORE THE ONE WITH LARGEST NOUAL PT+ NO+ AND	00013 ₀		
60200	6ü#	C ≠≠ WRITE THEM ON 11	000130		

061	DER			DATE 07306n	PAGE	3
60202	61*		NRECH1=NREC-1	000133		
UC203	62*		D0 505 K=1,NRECH1	000136		
60206	63#		READ (10,400)DUM	000141		
00211	64*	505	WRITE (11,400)00M	000150		
.60211	65*	C505	PRINT 402, DUN	000150		
60211	66¥	C **	READ RECORD WITH LARGEST NODAL PT. NO. (GRID PT. NO.)	000150		
L0215	67*	507	READ 110,4CD1GRDMAX	000162		
u0215	68*	C¢≉		000162		
UD22U UD221	<u> </u>		NEND=NUMNP+NREC	600171		
00221	70# 71#		DO 506 K=1,NEND	600174		
<u>00224</u> 00227	72*	506	READ(10,400)00M	000177	·	·· · · · ·
uu227 uU227	724 734		PRINT 462, DUM	000206 600206		
UD227	74+		WRITE UN 11 RECORD WITH LARGEST NODAL PT. NO.	000206		
60233	75*	C ++	WRITE_(11.400)GRDMAX	600208		
<u>00233</u>	76*	C	PRINT 402, GRUMAX	000217		
CU233	77*	c	PRINT 399	000217		
60233	78*	C399		680217		
50235	79*	<i><i><i>v</i>₀,,,</i></i>	NUMRE=NUMREC-NUMNP	000227		
L0230	80*	C	PRINT 24, NUMRE, NUMNP	630227		
00236	81*	ČZ4	FORMATE NUMRE AND NUMNP = 214)	000227		
00237	62*		DO 508 K=1,NUMRE	600232		
ÜÜ242	63 *			600235		
60245	84*	548	HRITE (11,400)00H	C00244		
60245	85*	C508	PRINT 402. DUM	000244		
60251	66*		ENDFILE 11	000255		
60252	<u>87</u> ≉		60 T0 552	600260		
LU253	¤ 8≠		LONTINUE	600262		
00253_		<u>C **</u>	SINCE FILE 10 ALREADY IN CORRECT ORDER JUST WRITE TO FILE 11	000262		
C0254	90×		REWIND 16	000262		
60255	<u>91*</u>		<u>D0 551 1=1+NUPREC</u>	000264		·····-
U026U	92*		READ(10,40C) DUM	000270		
<u>L0263</u>	93*	551				
60263	94#	C551	· • -	000277		
<u>60267</u> 00267	<u>95*</u> 96*	C ++	ENDFILE 11 CREATE MATERIAL VECTOR MATII (NUM18)AND MATERIAL	000310		
60267	97*		_HATRIX_MATIZ(NUM18,10)_THE_CUNTENT_OF_MATIL(K)_IS_THE_HID_OF_MATIZ(K,J).	000310		
U027U	98*		REWIND 11	<u> </u>		
uu27u u027u	>8÷ 99≠		IT IS NECESSARY TO RENUMBER MID'S AND PID'S SINCE CAUDS	000314		
L027u	160*	C ##	MAY GENERATE NOS. THAT ARE TUO LARGE.	000314		
00270	1.1*		SET UP KVCMID AND KVCPID			
L0270	1u2#	C		U00314		
60271	1ù3+		NENCENSTART(12)-1	000316		
ü0272	164*		NPCAHDEO	000321		
. 11.273	105*			CD0327		
UU276	106*	7.00	NPCARD=NPCARD+NUM(1)	000327		
60300	107*		UQ 733 I=1.NENU	600332		
60303	145#	733	REAU(11,400)DUM	100336		
60307	<u>169*</u>		U0 710 1=1,NPCARU	600350		
60312	110*		REAC(11,75m) PID,MID	000350		
<u>u0312</u>	111*	C	PRINT 749,PIU,MID	000350		
00316	112*	749	FORMATE * PIU AND MID READ FROM 11 IN ORDER* ,2X, 218)	000356		
L0317	<u>113*</u>		FORMAT(8x,218)	600356		
60317	114*	C	PRINT 911	000356		
00317	115*	<u>C911</u>	FORMATI " IN URUER GOING TO VEC ")	L06356		
60326	116*	6	CALL VEC(MID,PID)	000356		
60320	117*	<u> </u>	PRINT 912	600356		

OR	DER			DATE 073080	PAGE	4
60.320	118+	6917	FORMAT(* IN ORDER JUST BACK FROM VEC*)	000356		
00321	119*	710	CONTINUE	000363		
60321	120+		THE NEXT NINE LINES ARE DEBUGS. TO BE COMMENTED OUT LATER.	600363		
00321	121*	C	PRINT 751	000363		
00321	122*	Č	00 760 I=1.HAXHIU	000363	_	
00321	123*	C760		£00363		
60321	124*	Č.	PRINT 752	600363		_
60321	125+	С	UO 770 I=1,MAXPID	000363		
J0321	126*	C770	PRINT 761.KVCPID(I)			
00321	127*	C761	FORMAT (18)	606363		
00321	128*	<u> </u>	FORMAT(* KYCHID IN ORDER.)	000363		
60321	129*	C752	FORMATI ' KUCPID IN ORDER')	000363		
60323	130*		CALL COUNT(NUMMID, NUMPID)	<u> </u>		
60323	131*	C	PRINT 771,NUMMID,NUMPID	000363		
60323	<u>132×</u>	C771	FORMAT(* NUMMID AND NUMPID IN ORDER * 218)	000363		
60324	133*		REWIND 11	000367		
00325	134*		NEND17=NSTAR1(18)-1	600372		
66326	135#		U0 654 L=1,NEND17	000375		
<u>40331</u>	1364	654	READ 111,400 IDUH	000460		
LU335	137#		NUM18=NUM(18)	606410		_
60335	138 +	C_**	CHANGE MID IN MAT11	606410		
60335	139#		HIU AND PID IN C AND P CARUS WILL BE CHANGED IN CANDP	030410		
60336	146*	-		UDn412		
00337	141*		U0 655 LL=1,NUM18	C00417		
L0342	142*			600417		
UD 34 3	143*			000424		
60344	144+		READ (11,604) XMAT11(L),(XMAT12(L,J),J=1,7)	L00426		
60353	145#	604	FORMAT(8X, 18, 268.3, 5F8.4)	000443		
LD354	146*		CALL SIGMA(XMAT11(L).LLL.XMAT12)	600443		
60355	147*			600452		
60356	148*		Do Boo IEL-NUMMID	606460	_	
L0361	149*		IF(XMAT11(L).EQ.KVCMID(I)) KHOAD=1	600460		
00363	150*	800				
00365	151*		IF(KH0AD.EQ.0) GO TO 657	<u>ÜDU466</u>		
60367	152#		MAT11(L)=XMAT11(L)	000470		
00370	153*		U0 8.5 J=1,10	000476		
00373	154*	865	HAT12(L.J)=XHAT12(L.J)	000476		
U0373	155*	C	PRINT 625,L,HAT11(L)	000476		
L0373	156*		FORMATE * BEFORE CHANGE . HATILE * .12. *)=* .13)	ong476		
00375	157*		CALL ID(1,NUMMID,MAT11(L),KM)	600500		
60376	158+		HAT11(L)=KH			
u(j 376	159*	C	PRINT o27,L,HAT11(L)	600510		
60376	_160*		FORMATI' AFTER CHANGE MATIL(*,12,*)=*,13)			
60377	161#		$GO T_{ij} = 55$	606212		
ԱոԿու	162*	657	_PRINT_658,XMATIL(_)	000514		
60463	163*		KOUNT=KOUNT+1	600521		
60404	164*	658	EQRHATE AN HIU NOT REFERENCED BY ANY PID: "+14+ "+ THE COPPESPONDI	000526		
00404	165*		ENG MATI CARD HAS NOT BEEN PUT IN MATIL OR MATIL'I	000526		
60405	166*	655	CONTINUE	U00526		
60405	167#	C	PRINT 420	000526		
60405	168#		FORMATI' MATIL AND MATIL IN URUER')	000526		
60407	169#	<u> </u>	NUM(18)=NUM(18)-KOUNT	000526		
60410	170*		NUM18=NUm(18)	<u> </u>		
60410	171*	<u> </u>	UO 660 M1=1,NUM18	00 ₀ 53 <u>1</u>		
60410	172*	C66D	PRINT 614.MAT11(M1).(MAT12(M1.J).J=1.10)			
		<u> </u>	FORMAT (/15,5E11.5/3X,5E11.5)	U00531		
u041j	173*					

08	DFB			- <u>.</u>					DATE 07	308C	PAGE	5
60412	175*	RETURN					<u> </u>		00	0535		
60412 60413 END FOR	176*	END							00	0560		
		<u>_</u>							4	····		
aHDG,P	PAIR	· · · · · · · · · · · · · · · · · · ·										·
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SFOR, S PAIR, PAIR HSA E3 -U7/30/80-09:46:12 (55,)

SUBROUTINE PAIR ENTRY POINT 000251

.

STORAGE USED: CODE(1) 000307; DATA(U) 700230; BLANK CONHON(2) 000000

COMMON BLOCKS:

MOZART 000334 BACH 000044 0003 <u>0034</u>

EXTERNAL REFERENCES (BLOCK, NAME)

UUU5 NERR2\$ LOG6 NERR35

STORAGE RESTORATION ADDRESS AND ADDRESS REPAILS REPAILS REPAILS ADDRESS	STORAGE ASSIGNMENT	(BLOCK, TYPE,	RELATIVE	LOCATION,	NAMEJ
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_	0000 0001	060160 0666.77		1001 1001	000202		0001	000052		0001	000053	-	6001 0001	000064 00n142	-	,
×	5061 anri 1	000206 000157		00 01 0001	000167		0001	000172	2046	~0001 0000	000132	78L	0001	000127		<u> </u>
XX	1 0000	000152	J3 C	1 000	000153	J4	1 0000 I	000155	J5	0000 I	000154	J6	0000	000156	J7	-
I-4		000157			000146 000150		<u>.0000 I</u> 0004	000145	** • • • • • •		000000			<u> </u>		— ,
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PAGE 1

DATE 073080

60101	T*.	SUBROUTINE PAIR (KGDE, CID, PRID, ISTOP, NCARDS, KOL, NOUT, MATCH, INDEX)	000053	
60101	2*	C ** KODE INPUT. D FOR MID PAIRING, 1 FOR PID PAIRING	000053	
CO101	3*	C ** CID. INPUT. ELEMENT ID ARRAY. E.G., CRODID	000053	
60101	4 #	C ** PRID. INPUT. PROPERTY ID ARRAY. E.G. PROD1		
L0101	5*	C ** ISTOP. INPUT. NO. OF ELS. WITH GIVEN CONST. CODE (=NPAR(2))	000053	
60101	6*	<u>C ** NCARDS, INPUT, NUMIA IF KODE=D, IF KODE=1, TOTAL NO. OF P. CARDS FOR GIVEN</u>		
00101	7*	C ★★ ELEMENT. E.G., IF NASTRAN RUD,NCARDS=NUM12	000053	
60101	<u>8*</u>	<u>C ** KUL. INPUT. IF KODE=D. COL. NO. OF MID IN ELEMENT ID ARRAY.</u>		<u>. </u>
60101	9 ‡	C ## E.G. IF GIVEN ELEMENT IS NASTRÅN ROD.KOL=4. IF KODE=1.COR.	000053	
60101	<u> 10* </u>	C ** NO. OF PID IN ELEMENT ID ARRAY. E.G., IF GIVEN EL. IS	600053	
60101	11*	C ## NASTRAN ROD, KUL=5	000053	
60101	12*	C ** NOUT, OUTPUT. IF KODE=D.NOUT=NUMMAT FUR GIVEN EL. AND CONST. CODE.	600053	
60101	13*	C ** IF KODE=1,NOUT+NUMGEO FOR GIVEN EL. AND CONST. CODE.	600053	
00101	14*	<u>C ** MATCH. OUTPUT. FIRST COL. CONTAINS MID(KODE=C) OR PID(KODE=1). SFCOND COL.</u>	00nn53	
C0101	15*	C ** CONTAINS ROW NO. OF THAT MIU IN MATI2(KODE=1),OR ROW NO. OF THAT PIU IN	600053	
60101	16*	C ** PROPERTY ARRAY OF GIVEN ELEMENT (E.G. PRODZ ARRAY)	000053	
00101	17*	C ** INDEX. OUTPUT. THE ELS. OF INDEX ARE HID OR PID NOS. OF THE CURRENT ELEMENT	000053	
00103	18*	COMMON/MOZART/MAT11(20), MAT12(20,10)	000053	
60164	19*	COMMON/BACH/NUM(18),NSTART(18)	600053	
60104	20+	C ** CHANGE GROUP. NO. OF ELEMENTS	600053	
60105	21*	INTEGER CID(100,7),VID(100),INDEX(100),MATCH(100,2)	000053	
00105	22*	C ** CHANGE GROUP . NO. OF PID'S	000053	

PA1	R		ATE 073080	PAGE	Z
60106	23×	INTEGER PRID(25,2)	000053		
60106	24#	(** **********************************	000053		
00107	25×	REAL MAT12	000053	_	
-υ ₀ 11.,	26*	1 FORMAT()	000053		
00110	27#	C PRINT 2	000053		
60110	28*	C2 FORMAT(5x,*KODE*)	000053		
00110	29*	C PRINT 1, KODE	000053		
60111 60114	30*	D0 666 KR=1,100	000053		
00117	<u> </u>	00 <u>666 KC=1,2</u> 666 MATCH(KR,KC)=0	000053		
UD122	33*	00 8 K=1,100	000055		
L0125	34*	B INDEX(K):0	<u> </u>		
60125	35*	C PRINT 3	000064		
60125	36#	C3 FORMATISX, "CID#S JUST INSIDE PAIR")	000064		
60125	37+	C DO 1C MI=1,ISTOP	000064		
00125	38*	C10 PRINT 11, (CID(M1,M2),M2=1,7)	000064		····
00125	39*	C11 FORMAT(716)	U00064		
UC127	≉(ن 4	NUM16=NUM(18)	000065		
<u>L0127</u>	41#	C PRINT 40	006065		
00127	42*	C40 FORMAT(" MAT11 AND MAT12 IN PAIR")	000065		
_60127	43*	L DO 41 I=1.NUM18	000065		····
60127	44*	C41 PRINT 45, MAT11(1), (MAT12(1,J), J=1,10)	600065		
<u>0127</u> 0130	45*	<u>C 45 FORMAT(/15.5E11.5/3X.5E11.5)</u> KODE = KODE+1	<u> </u>		
60130	40≠ 47≠	MAXID = CID (1.KOL)	600067 606072		
00132	48+	UO 75 J_2=2,ISTOP	000077		• • • • • • • • • • • • • • • • • • • •
66135	49.	IF (CID(J2,KOL),GT, MAXID HAXID = CID (J2,KCL)	000077		
GC137	≉ن ڏ	75 CONTINUE	000106	··	
60137	51*	C ** FOR ALL INTEGERS BETWEEN 1 AND MAXID, FIND OUT HOW MANY HID'S OR PIN'S THERE	000106		
60137	52*	C ** ARE OF EACH. ASSUME MAXID =100	000106		
60137	53*	C PRINT 4	001066		
60137	54 *	C4 FORMAT(5x, "MAXID IN PAIR")	UD0106		
L0137	55*	C PRINT 1, MAXID	000106		
UG141	56*	UO 76 J2=1, MAXID	000106		
<u>UD144</u> GC146	<u>57*</u> 58*	<u>76 yid (J2)=0</u> D0 77 J3=1, MAXID	000114		
60148 6 ₀ 151	50≠ 59≠	D0 77 J3-1, MAXID D0 76 J4=1, ISTOP	000123		
60154	<u> </u>	IF (CIU(J4,KOL).EQ. J3) GO TU 79	000123		
00156	61*	60 TO 78	600125		
UU157	62#	79 vid (J3) = vid (J3) + 1	000127		
00166	63*	78 CONTINUE	00n135		
CO162	64¥	77 CONTINUE	600135		
00162	65#	C PRINI 15	600135		
LG162	66 	C 15 FORMAT(* VID IN PAIR*)	006135		
40162	67*	C 00 16 J3=1.HAX1D	000135		
L016_	63*	C16 PRINT 1, VID(J3)	006135		
L0162	<u>69</u> ≉ 7J≄	C ** NOUT (=NUMMAT OR NUMGEO) IS THE NO. OF NONZERO ELS. OF VID.	000135		
0162	7J# 71#	C ** THE N'S ON THE MATERIAL CONTROL CARDS AND GEUMETRIC PROP. CARDS ARE THE C ** INDICES OF THESE NONZERO ELS. OF VID.	000135		
L0164	72+	L ** INDICES OF THESE NONZERO ELS. OF VID. J6=r,	000135		·
60165	73*	NOUT=C	000135		
L0166	74#	UO 80 J5=1,MAXID	000138		· · · · ·
L0171	75*	$IF(VI_{D}(J5)) \in (-6)GO TO BD$	U00142		
L0173	76#	NOUT=NOUT+1	UGU144		
60174	71*		000147		
0175	78*	INDEx(J6)=J5	000154		
10176	79¢	BU CONTINUE	000160		

PAI					PAGE	
L0176	83*	C **	INDEX(1)=J_MEANS THAT J IS AN MID OR PID NO.	000 160		
L0176	81*	C ++	FIND THE MATLL OR PID INDEX CORRESPONDING TO EACH OF THE ELS.OF INDEX	000160		
60200	6Z*	-	00 61 J7=1.NGUT	000160		
60203	83*		UO 82 J8=1,NCARDS	000172		_
UD206 _			60 TO (100-200)-KODE	600172		
60207	85¥	100	IF(MAT11(J8) ,EQ, INDEX(J7)) GO TO 83	000202		
00211	86*					
60212	87*	200	IF (PRID (J8,1, EQ. INDEX (J7)) GO TO 83	000206		
60214	88*			000210		
60215	89*	83	MATCH (J7+1)=INDEX(J7)	000212		
C0216	<u>91</u> *		HATCH (J7,2)=J8	000213		
00217	91×	d 2	CONTINUE	600224		
60221	92*	81		000224		
u0221	93 *	С	PRINT 5	000224		
<u>uD221</u>	94*	<u>C5</u>	FORMAT(5X, *MATCH AND INDEX IN PAIR*)	000224		
68221	95*	С	UO 150 M1=1,NOUT	000224		
00221	96*	<u>c150</u>	PRINT 1. (MATCH(M1.M2), M2=1, 2), INDEX(M1)	000224		
00221	97×	С	PRINT 6, NOUT	000224		
L0221	98*	<u>C6</u>	FORMAT(5X, *NOUT=*.13)	000224		
uU223	99*		RETUKN	000224		
L0224	100*		END	<u>CO0306</u>	·····	
END FOR						
<u>aHUG.P</u>	LDME	. <u>M</u>				
<u>aHUG.P</u>	LDME	<u>M</u>				
анµд.р	UDME	.m				
<u>анра, р</u>	UDME	.H				
<u>анра, р</u>		.н 				
<u>анра р</u>		.H				
<u>анра, р</u>		.H				
<u>анра, р</u>		.H				
<u>анис.</u>		.H				
<u>анис.</u>						
ани <u>с.</u>		.H	·			
ани <u>с.</u>		.H				
attibe . P		.H				
анис. 		.H				

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GFOR, S UDMEM, UDMEM HSA E3 -07/30/80-09:46:16 (6,)

SUBROUTINE ODMEN ENTRY POINT CO0145

STORAGE USED: CODE(1) DUC177; DATA(U) DDC176; BLANK COMMON(2) DOCODO

COHMON BLOCKS:

6023 BACH 600044

EXTERNAL REFERENCES (BLOCK, NAME)

JUJ4 FORM1 **UUU5 ODMEM1** UUn6_ UUMEM2

GUJ7 NERR35

STORAGE ASSIGNMENT (BLOCK, TYPE, RELATIVE LOCATION, NAME)

	<u>6031 0001nn 11L</u>	0001 000063 1216	0001 000072 1266	6001 000120 20L	0000 1 0000 <u>04 IDV31</u>	
\times	JUGO I JECU66 IDV32	C000 000161 INJP\$	0000 I 000153 ISTOP1	0000 I 000154 ISTOP2	0000 I 000002 KOAD	,
- X	<u></u>	<u>0000_1.000156_k18</u>	000022_NSTART	6063 I 000600 NUM	0060 I 000151 NUM14	•
2	UDUC I GUD152 NUM18	COUD I CCO150 NUM7	3000 I 000000 TYPE			
						<u> </u>

.

DATE 073080

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00101	1*	SUBROUTINE OUMEMICONFAC.NKODE.1DV3,NIDV3)	000002	
00103	2 \$	COMMON/BACH/NUM(18),NSTART(18)	000002	
<u>001:14</u>	3#	INTEGER TYPE(2)+KOAD(2)+IDV3(100)+IDV31(50)+IDV32(50)	600002	
60105	44	DATA TYPE/*GUADHE*,*M */	606002	
	<u>5,+</u>	CALL FORM1 (NKODE, KOAD . TYPE)	000002	
L011 _Ŭ	· 6*	NUH7=NUH(7)	000007	
60111	7*	<u>NUM14=NUM(14)</u>	000011	
60112	8*	NUM16=NUM(18)	000013	
LU113	<u> 7*</u>	IF(NKOGE+EQ-1) 50 TO 11	000015	
66115	lu*	CALL GUMEM1(NUM7, NUM14, NUM18, CONFAC, KOAD, IUV31, ISTOP1, NKODE)	000020	
L0116	11*	CALL QUMEMZINUM7, NUM14, NUM18, CUNFAC, KOAD, INV32, ISTOP2, NKODE)	660032	
60117	12*	NIDV3=1STOP1+ISTOP2	LCD_44	
LOIZU	13*	D_0 1,1,1, K17=1,1STOP2	000063	
60123	14≠	101U 10V3(K17)=I0V31(K17)	CODD63	
60123	15*	U0_1020_K18=1+ISTOP2	600072	
66139	16#	1026 IDV3(ISTOP1+K16)=I6V32(K18)	000072	
L0132	17#	RETURN	000074	
60133	18 🕈	11 IF (KOAD (1).E.v.1) GU TO 20	600160	
00135	19#	CALL ODMEMZ(NUM7,NUM14,NUM18,CONFAC,KOAD,IDV3,NIDV3,NKODE)	000102	
60130	∠U*	RETURN	60(114	
L0137	21*	20 CALL QUMEMI(NUM7, NUM14, NUM13, CONFAC, KOAD, IDV3, NIDV3, NKODE)	U00120	
00146	22*	RETURN	000131	
1.0141	<3×		CO0176	

QL	JM	É	M	1
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	60101	1*		SUBROUTINE QUMEHI(NUM7, NUM14, NUM18, CONFAC, KOAD, IDy31, ISTOP, NKODE)	U00000
	60103	2*		COMMON/BACH/NUM(18),NSTART(18)/HOZART/MAT11(20),MAT12(20,10)	000000
	00103	3*	C **	CHANGE GROUP. NO. OF DESAP PLANE STRESS QUADRILATERALS(NASTRAN	000000
	<u> 00103</u>	4≠	C **	QUADRILATERAL MEMBRANE, CONSTRUCTION CODE 1	00000
	00104	5*		INTEGER CODM1(100,6),VEID(100),CODMID(100,9),MATCH(100,2),INDEX(10	00000
·	60104	6*		14)	00000
	00135	7*		DIMENSION DUMMY(100,1)	00000
	60105	8*	C **	CHANGE GROUP, NO. OF PID'S	000000
	00106	9*		INTEGER PQDM1(25,2)	00000
۰	60107	<u>1ù*</u>		DIMENSION PAUMZ125,2).GEOPR2(25,6)	<u></u>
	C0107	11*	C **	CHANGE GROUP. NO. OF HID'S	00000
·	<u>- 40110</u> 60111	12*		DIMENSION MICRUIIZA, 21	000000
	00111	13* 14*	C **	REAL MTPRO2(20),MTPRO3(20,8) CHANGE GROUP. NO. OF DESIGN VARIABLES	0 ₀₀₀₀₀
·	60112	<u> </u>	U ¥¥	DIMENSION IDV31(5n)	000000
	00112	16*	c ++	<u>**^{**}********************************</u>	000000
	60112	17#	<u> </u>	INTEGER KOAD(2), TYPE(2), NPAR(7), VAR(4)	<u>000000</u>
	60114	18*		DIMENSION ENUI (4.4)-ELDAT2(100.4).DUMM(4)	
	L0115	19*	·····	REAL MAT12	L00000
	LU116	17÷ 2J≠		DATA_VAR/'A', 'B', 'C', 'D'/TYPE/'QUADHE', 'H '/	600000
	LU121	21#	7	FORMAT ()	00000
	60122	22*	,	CALL CANDPIT.NUMT.NUM14.CODH1.DUMMY.PODH1.PODH2)	u00000
	60123	23*		WRITE (6,5)	000010
	0125	24*	5	FORMAT (2X, *PLANE STRESS CHADRILATERAL LOUAD MEMBRANE, IN NASTRAN).C.	000015
	L0125	25¥		IONST. COLE 1')	600015
×	UC126	26*		NPAR(1)=3	00015
	UC127	27*		NPAR (2) = NUM (7)	600017
X	00130	28*		NPAR (4) = 1	000021
	60131	Z9*		NPAR(5)=1	000023
49	6.0132	30*		NUM7=NUM(7)	000024
•	66133	31*		ISTOP=NUH7	600025
	60134	32+			<u>u00026</u>
	60135	33+		IFINKODE.EQ.1) GO TO 12	000027
	00137	34*		CALL FORM2(JJ1,KOAU.TYPE,NPAR(2).VEID.NUM7)	<u>600031</u>
	UO140	35*		ISTOP=NPAR(2)	000041
	U0141			NNCOL=6	600043
	00142	37×		CALL STOR(ISTOP,NUH7,CQDH1,VEID,NNCOL)	000045
	L0143	38*	12	00 7C J=1,ISTOP	000055
	60146	39*		NEID=CyDH1(J,1)	000064
	00147	40*	C ++	NPID=CoDM1(J,2)	000066
	00147 00150	41* 42*	C ++	FIND PUDM1 CARD WITH SAME ID	CODO66 000073
	60153	43*		IF(NPIU_EQ,PuDM1(K,1))GO TO 61	<u> </u>
	ü0155	44 =			00075
	60156	45*	61		000100
•	60157	46*		<u>60 TU 62</u>	000101
	00160	47×	65	CONTINUE	000105
	60162	43*		CODMID(J,1)=NEID	000105
	60163	49#		CQDM1D(J,2)=CQUM1(J,3)	000107
		<u>50*</u>		(QDP1D(J,3)=(QUM1(J,4)	000111
	LU165	51*		CODMID(J,4)=COUM1(J,5)	000113
	L0166	<u>52 ×</u>		CQDMID(J,5)=CQDM1(J,6)	
	UD167	53×		CQDMID(J,6)=PQGM1(KK,2)	606117
·	60170	54*		<u>CQDMID(J,7)=NPID</u>	000121
	60171	55*	70	CONTINUE	600125
	66171	56*	<u>(</u> + +	FIND NUMMAT AND MATCH	000125

cu	HEM1		DATE 073080	PAGE	3
60173	57 *	**************************************	L00125		
UC174	58*	KKOL=6	000126		
60175	59 *	CALL PAIRIKKODE, CODMID, PUDM1, ISTOP, NUM18, KKOL, NUMMAT, MATCH, INDEX)	000130		
CO176	60+	NPAR (3) =NUMMAT	000145		
00177	61*	WRITE (6,501)	000147		
u0201	62*	501 FORMAT (2X, ENTER CODE FOR INCOMPATIBLE DISPL. MODES: O MEANS USE	000155		
60201	63*	1THEM.; 1 HEANS DON,T USE THEN"	600155	· <u> </u>	<u> </u>
60202	64#	READ (5,7) NPAR(6)	000155		
<u>60212</u>	65*	C ++ SET UP MATERIAL PROPERTIES DATA	000155		
60205	66¥	DO 83 J10=1, NUHMAT	006167		
60210	67+	$HPR_01(J1_0,1)=INDE_X(J10)$	000167		
UC211 00212	68* 69*	MTPR01(J10,2)=1	000171 000176		
C0212	70+	<u> </u>	<u> </u>		
00217	71¢	G0 TU 84	000178 0nn201		
00220	72*	85 NROW=MATCH(J11+2)	600203	· · ·	
00221	73×		006204		
60222	74*	84 CONTINUE	000210	· · · · · ·	
66224	75×	86 MTPR02(J10)=HAT12(NROW.4)*CONFAC	000210		
00225	76#	MTPR03(J10,1)=0.	000213		
60226	77*	HTPR03(J10.2)=MAT12(NROW.1)	000214		
00227	78*	H1PR03(J17,3)=MAT12(NROW_3)	000216		
00230	79#	MTPR03(J10,4)=n.	000220		
L0231	8:j+	MTPR03(J10,5)=MAT12(NRON,8)	600221		
60232	81*	MTPR03(J10.6)=MAT12(NR0H.9)	<u> </u>		
LO233	°2≠	HTPR03(J10,7)=HAT12(NROW,16)	LOU225		
60234	<u>63*</u>	<u>MTPR03(J10,6)=MTPR03(J10,6)</u>	600227		
00235	64 ♥	IX=INDEX(J10)	000231		
00236	<u>65</u> *	B3 CONTINUE	L00235	<u> </u>	
LO236	86¥	C ++ MATERIAL PROPERTIES DATA IS DONE.	000235		
<u>CC236</u>	87*	C ** SET_UP_GEOMETRIC_PROPERTIES_DATA	000235		
CO240 0n241	88* 89*	KKKODE=1	600235		
<u> </u>	9.7 *	KOL=7 CALL PAIR(KKKODE,CGDMID,PQDH1,ISTOP,NUM14,KKOL,NUMGEO,MATCH,INDEX)	000237		
00243	y]≠	NPAR(7)=NUMGEO	630254		
60244	92*	wRITE(12,795)	ü00256		
60246	93*	795 FORMATI' COMM ELEMENT CONTROL FROM QDMEM1')	000263		
UU247	94*	WRITE(12,504)(NPAP(J11),J11=1,7)	L00263	······	
1.0252	95#	504 FORMAT(715)			
L0252	¥6*	C ** WRITE MATERIAL PROPERTIES DATA	000277		
60253	974	WpItE(12,349)	000277		
00255	¥8*	349 FORMATE COMM MATERIAL PROPERTIES FROM QUMEMI+)	000313		
<u>60256</u>	99*	U0 350 J12=1+NUMMAT		·	
00261	160*	35n WRITE(12,300)(MTPR01(J12,J13),J13=1,2),MTpR02(J12),(MTPR03(J12,J13	606313		
00261	<u> 101# </u>	1), J13=1, 6)	600313		
60275	162*	300 FORMAT (215,F10.3/2F10.0,F10.3,5F10.0)	606341		
60276	163*	WRITE(0,301)	000341		
00300 00301	164*	301 FORMAT (10x, 'GLOMETRIC PROPERTIES FOR CONST. CODE 1*)	000355		
60304	<u>1ü5*</u> 1u6*	<u>00 283 J10=1,NUMGE0</u> D0 264 J11=1,NUMGE0			
60307	106# 1u7#	IF (MATCH(J11,1).EQ_INDEX(J13))GO_TU_285	000355		
60311	1074		606355		
60312	108+	285 NROW-MATCHIJ11,21	GDU360		
LO313	116*	60 Tu 286	<u> </u>		
UG314	111*	284 CONTINUE	COD363		
60316	112#	286 UEOPR2(J1G,1)=FQ(M2(NKOW,1)	U00367		
~~~~	113+	WRITE(0,362)1NUEX(U10)	LOU371		

<u> </u>	DHEM1		DATE 073080	PAGE	
ÚD 322	114+	302 FORMAT(2X, FOR PROPERTY ID + 14, * ENTER W.SA, SI.D. WE, IN THAT ORDER*	000377		
60322	115+	1)	000377		
60323	116*	READ(5,7) (GEOPR2(J12), J12=2,6)	U00377		
U0326	117*	283 CONTINUE	600411		
60336	118*	WRITE(12,96)	600411		
60332	119#	96 FORMAT( * COMM GEOMETRIC PROPERTIES FROM QDHEH1 * )	600422		
60333	120*	DO 299 J10=1.NUMGEO	600422		
00336	121*	299 WRITE(12,303)INDEX(J10),(GEOPR2(J10,J11),J11=1,6)	L00422		
60346	122=	303 FORMAT(15,6F10.0)	006440		
00346	123*	C ** GENERATE ELEMENT LOAD NULTIPLIERS	000440		
L0347	124*		600440		
60351	125*	111 FORMAT(10X, "ELEMENT LOAD MULTIPLIERS FOR CONST. CODE 1")	000451		
<u>LU352</u> LU355	<u> </u>	D0 115 K6=1,4 D0 115 K7=1,4	000451		
U036U	128*	115 ENUL (K6, K7)=1,4	COO451 000451		
<u> </u>	129*	115 ENO(1801877+0.000)	000451		
60363 61366	130*	WRITE(6.112)VAR(L8)			
60371	$\frac{130 \pm 1}{131 \pm 1}$	112 FORMATIZX, "FUR LOAD ",A1," ENTER X, Y, Z ELEMENT LOAD FRACTIONS, IN	000464		
	132#	ITHAT ORDER. DEFAULT = ALL ZEROS*)	_600472		
60372	133÷	READ(5,7)(DUMM(K11),K11=1,3)	Ung472		
00375	134*	IF(DUMM(1),E4.0.)G0 TO 120	000562		
60377	135*	00 113 K11=2,4	C00511		
60402	136*	113 EMUL(KA:K11)=DUMM(K11)	606511		
00404	137#	120 CONTINUE	600515		·· ···
00406	138+	WRITE(12-1117)	000515		
ú041u	137*	1117 FORMATI' COMM ELEMENT LOAD MULTIPLIERS FROM (DHEM1')	000525		
63411	143÷	UO 121 K9=1,4	600525		
60414	141#	121 WRITE(12,116)(EMUL(K9,K10),K10=1,4)	000525		
UD423	142*	116 FORMAT(4F10+U)	000541		
ŪŪ423	143*	C ** GENERATE ELEMENT DATA	600541		
00424	144+	WRITE(6.127)	000541	· · · · · · · · · · · · · · · · ·	
60426	145¥	127 FORMAT(15X, *ELEMENT DATA FOR CONST. CODE 1*)	000551		
60427	146*	DO 150 KROH=1.ISTOP	000551		
60432	147*	IX=CQDHID(KROW,1)	000551		
<u> </u>	<u>    148*                                </u>	write(6,130) ₁ x	000552		<u> </u>
00436	149*	13n FORMAT(2X, "ENTER DESIGN VARIABLE NO., IUV, FOR EL.NO.", IS)	000560		
<u>L0437</u>	<u>150*</u>	READIS.71CODHID(KROW.8)	000560		
00442	151*	WRITE(6,131)IX	000566		
<u> </u>	<u>152*</u> 153*	131 FORMAT(2X. *ENTER DESIGN VARIABLE FRACTION.FFC.FOR EL.NO.*.I4.*. DE 1FAULT=1.*)	<u> </u>		
00445	154*	ELDAT2(KROW.1)=1.	ú0n574		
L0447	155*	READ(5,7)DUM	<u> </u>		
60452	156 *	IF(DUM_NE_Q_)ELDAT2(KROW_1)=DUM	600604		
60454	157*	ELDATZ(KROW,2):0.	60061g		
00455	158*	WRITE(6,132)1v	000611		
ննկ6ս	159*	132 FORMATIZX, "ENTER ANGLE BETA FOR EL. NO.", 14,	000617		
LD461	160*	READ(5,7)ELDAT2(KROW,3)	000617		
60464	161*	WRITE(6,133)IX	600625		
60467	162*	133 FORMAT(2X, VENTER EFC FOR EL NO. V. 14. V. DEFAULT =1. V)	<u>LINN633</u>		
L0476	163*	ELDAT2(KROW, +)=1.	004633		
60471	164*	READ(5,7)DUM	000635		<u> </u>
60474	165*	IF (DUM+NE+0+)ELDAT2(KROW+4)=UUM	600643		
LU476	166*	WRITE(6,134)1X	600647		
00501	167*	134 FORHATIZX, "ENTER STRESS PRINTOUT CODE, KS, FOR EL.NO.", 14, ". DEFAULT	UDU655		
<u>in\$q1</u>	168*	1=3')	UDU655		···
60502	167*	CODMID(KROW, 9)=3	000655		
60503	170*	READ(5,7)IDUM	000657		

C500         71*         Ffluen.AC.0020000124004.001.0000         Openets           C6510         77*         155         C614.21         C00673           C6511         77*         C0073         C00673         C00673           C6511         77*         C         PRESS         C0070           C6511         77*         PRESS         C0070         C0070           C6511         77*         PRESS         C00711         C00711           C6520         160         PRESS         C00721         C00723           C6531         161         C0071100127001270012501         C00735         C00735           C6531         162	QDME	.н.		DATE 073080	PAGE	5
U0510       172*       150       CONTINUE       000673         U0512       173*       NNCOL129       000675         G0514       175*       CALL ELORD(ISTOP, CCDMID, ELDAT2, NNCOL1, NNCOL2)       000677         00514       176*       C       PRINT 179       000677         00515       177*       C170       FORMAT(//* ELEMENT RECORD ON FILE 12, IN QDHEM1*/)       C00677         00515       178*       WRITE(12,1135)       CDUTC6       C00713         C0515       178*       WRITE(12,1135)       CDUTC6         C0517       179*       1135       FORMAT(* COMELLEMENT DATA FROM QDMEM1*)       CD0713         C0523       181*       180       WRITE(12,165)(CQUMID(M12,M13),M13=1,8), (ELDAT2(M12,M13),M13=1,4), C       C00723         C0523       181*       180       WRITE(5,186) (CQUMID(M12,M13),M13=1,8), (ELDAT2(M12,M13),M13=1,4), C       C00723         C0523       182*       C       1COMID(M12,9)       C00723       C00755         C0537       185*       186       FORMAT(815,2710.0,15,5X)       C00755       C00755         C0543       187*       190       IOy31(A1)=CQUMID(K1,8)       C00755       C00755         C0543       187*       190       IOy31(A1)=CQUMID(K1,8)	60506	171*	IF (IDUM .NE . D) CODM t D (VROW. 9) = IDUM	Unn665		
U0512         173*         NNC0L1=9         000673           G0513         174*         NNC0L2=4         U00675           00514         175*         CALL ELORD(ISTOP,CGDMID,ELDAT2,NNC0L1,NNC0L2)         U00677           00514         176*         C         PRINT 179         U00677           00514         176*         C         PRINT 179         U00677           00514         176*         C         PRINT 179         U00677           00517         179*         GDEAT         CDEAT         CDEAT           U0518         178*         WRITE(12,1135)         CDUTGE         CDUTGE           U0520         183*         U0 180 MI2=1,155 UP         CDUT13         CDUT13           U0523         183*         180 (CQUMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C         C00723           U0523         184*         140 MIUM12,9)         C00723         C00723           U0523         184*         C         140 MIUM12,9)         U01723           U0523         184*         C         140 MIUM12,9)         U00723           U0543         184*         C         140 MIUM12,9)         U00723           U0543         186 FORMAT(815,2F1U-U,22X/2F1D+0,15,5X)         U00755			150 CONTINUE			
CUS13       174*       NNCOL2=4       UD0675         CUS14       175*       CALL ELORD(ISTOP,CCDHID,ELDAT2,NNCOL1,NNCOL2)       UD0677         UDS14       176*       C       PR[N1179       UD0677         UDS14       177*       C179       FORMAT(//* ELEMENT RECORD ON FILE 12, IN QDMEM1*/)       CD0677         UDS15       178*       WRITE(12,1135)       CD0713       CD0713         UDS20       180*       D0 180 H12=1,ISTOP       UD0713       UD0713         L0521       180*       D0 180 H12=1,ISTOP       U00713       U00713         L0523       181*       180       LRITE(12,186)(CQ0HID(M12,M13),H13=1,8),(ELDAT2(M12,M13),H13=1,4),C       U00723         L0523       182*       140HIU(M12,9)       U00723       U00723       U00723         L0523       183*       C180       WRITE(6,186) (CQ0HID(M12,M13),H13=1,8),(ELDAT2(M12,M13),H13=1,4),C       U00723         L0523       184*       C       140HIU(M12,9)       U00723       U00723         L0524       184*       C       140HIU(M12,9)       U00723       U00723         L0540       185*       186       FORMAT(815,2F10,0,15,5X)       U00755       U00755         L0544       186*       UD 19U K1=1,ISTOP       U00755	00512	173*				
G0514       175*       CALL ELORD(ISTOP, CGDHID, ELDAT2, NNCOL1, NNCOL2)       000677         00514       176*       C       PRINT 179       000677         00514       177*       C179       FORMAT(//* ELEMENT RECORD ON FILE 12, IN QDMEM1*/)       CQD677         00515       178*       WRITE(12,1135)       CQD72       CQD72         00517       179*       1135       FORMATI* COMM ELEMENT DATA FROM QDMEM1*)       CQ0713         00520       180*       U0 180 M12=1,ISTOP       C00713       C00723         00523       181*       180 WRITE(12,186) (CQDMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       C00723         00523       162*       140MIU(M12,9)       C00723       C00723         00523       184*       C       140MIU(M12,9)       000723         00523       184*       C       140MIU(M12,9)       C00755         00545       186       FORMAT(815,2F1U+G,20X/2F10-0,15,5X) <td></td> <td></td> <td>NNCOL2=4</td> <td></td> <td></td> <td></td>			NNCOL2=4			
00514       176*       C       PRINT 179       000677         10514       177*       C179       FORHAT(//* ELEMENT RECORD ON FILE 12, IN QDMEM1*/)       000677         00515       178*       WRITE(12,1135)       00076         00517       179*       1135       FORMATI' COMH LEMENT DATA FROM QDMEM1*)       000713         10520       180*       D0 180 M12=1,ISTOP       000713         00523       181*       180RITE(12,186)(COUMID(M12,M13),M13=1,81,(ELDAT2(M12,M13),M13=1,4),C       000723         00523       162*       1QDMID(M12,9)       000723       000723         00523       184*       C       1QDMID(M12,M13),M13=1,81,(ELDAT2(M12,M13),M13=1,4),C       000723         00523       184*       C       1QDMID(M12,9)       000723         00523       184*       C       1QDMID(M12,M13),M13=1,81,(ELDAT2(M12,M13),M13=1,4),C       000723         00523       184*       C       1QDMID(M12,9)       000723         00523       184*       C       1QDMID(M12,M13,M13=1,81,(ELDAT2(M12,M13),M13=1,4),C       000723         00523       184*       C       1QDMID(M12,M13,M13=1,81,0)       000755         00543       186*       00       190       190       190       190       190 </td <td>00514</td> <td>175*</td> <td>CALL ELORD(ISTOP,CGDMID,ELDAT2,NNCOL1,NNCOL2)</td> <td></td> <td></td> <td></td>	00514	175*	CALL ELORD(ISTOP,CGDMID,ELDAT2,NNCOL1,NNCOL2)			
177*       C179       FORMAT(//* ELEMENT RECORD ON FILE 12, IN QDHEM1*/1       C00677         C0515       178*       WRITE(12,1135)       C00706         C0517       179*       1135       FORMATI* COMM ELEMENT DATA FROM QDHEM1*)       C00713         C0520       180*       UO 180 M12=1,ISTOP       U00713       C00723         C0523       181*       180       WRITE(12,186)(CQUMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       C00723         C0523       162*       ICDMID(M12,9)       C00723       C00723         C0537       185*       186       FORMAT(815,2F1U,G,20X/2F10+0,15,5X)       C00755         C0540       186*       UO 190 K1=1,ISTOP       C00755         U0545       186*       00 190 K1=1,ISTOP       C00755         U0545       186*       00 190 K1=1,ISTOP       C00755         U0545       186*       00 190 K1=1,ISTOP       C00755         U0545       186*       C00755       C00755         U0546       180*       RETURN       C00755         U0545       180*       RETURN       C00757         U0546       189*       END       K01051         LNU FOK       KNU FOK       K01051       K01051	00514	176*	C PRINT 179	000677		
C0515       178*       WRITE(12,1135)       C0076         C0517       179*       1135       FORMATI* COMM ELEMENT DATA FROM QDMEM1*)       C00713         C0520       180*       D0 180 M12=1,ISTOP       C00723         C0523       181*       180 WRITE(12,186)(CQUMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       C00723         C0523       182*       C100 WRITE(6,186)       CQUMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       C00723         C0523       183*       C180       WRITE(6,186)       CQUMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       C00723         C0523       184*       C       1CDMID(M12,9)       C00723       C00723         C0523       184*       C       1CDMID(M12,9)       C00723       C00723         C0523       184*       C       1CDMID(M12,9)       C00755       C00755         C0537       185*       186       FORMAT(815,2F10.0,15,5X)       C00755       C00755         U0543       187*       190       IOY31(K1)=CQUMID(K1,8)       C00755       C00755         U0543       187*       190       IOY31(K1,8)       C00755       C00755         U0545       188*       RETURN       END       END       END       END <td< td=""><td><u>iin514</u></td><td>177*</td><td>C179 FORHAT(//* ELEMENT RECORD ON FILE 12, IN ODMEM1*/)</td><td>000677</td><td></td><td></td></td<>	<u>iin514</u>	177*	C179 FORHAT(//* ELEMENT RECORD ON FILE 12, IN ODMEM1*/)	000677		
L0517       179*       1135 FORMATI* COMH ELEMENT DATA FROM QDMEM1*)       L00713         L0520       180*       D0 180 M12=1,1550P       000713         L0523       181*       180 wRITE(12,186)(CQUMID(M12,H13),H13=1,8),(ELDAT2(M12,M13),M13=1,4),C       000723         L0523       181*       140 MIU(M12,9)       000723         L0523       183*       C180 WRITE(6,186) (CQUMID(M12,M13),H13=1,8),(ELDAT2(M12,M13),M13=1,4),C       000723         L0523       184*       C       140MIU(M12,9)       000755         L0537       185*       186 FORMAT(815,2F10.0,20X/2F10.0,15,5X)       000755         U0543       187*       190       10y31(k1)=CQUMID(k1,8)       000755         U0545       188*       RETURN       000755       000755         L0546       189*       FND       FND       001051         LNU FOK	00515	178#	WRITE(12,1135)			
LU>2U     180*     D0 180 M12=1,ISTOP     U00713       U0523     181*     180 wRTE(12,186)(CQDMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C     GD0723       LU523     162*     14DMID(M12,9)     000723       GD523     183*     C180 wRTE(6,186) (CQDMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C     000723       LU523     184*     C     14DMID(M12,9)     000723       LU537     185*     186     FORMAT(815,2F10.0,15,5X)     000725       LU540     186*     U0 190 K1=1,ISTOP     000755       LU543     187*     190     IDy31(K1)=CQDMID(K1,8)     000755       LU545     188*     RETURN     000757       LU545     189*     END     END     U01051	60517	179*	1135 FORMATLY COMM ELEMENT DATA FROM QDMEM1*)			
U0523       181*       180       WRITE(12,186)(CQDMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       000723         U0523       162*       14DMID(M12,9)       000723         U0523       183*       C180       WRITE(6,186) (CQDMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       000723         U0523       184*       C       14DMID(M12,9)       000723         U0540       186*       U0 190 K1=1,ISTOP       000755         U0543       187*       190       IDY31(K1)=CQDMID(K1,8)       000755         U0545       188*       RETURN       000757       000757         U0546       189*       END       001051       001051         ENU FOK       V       V       V       V       V	<u>เชริม</u>	180+	DO 180 H12=1,ISTOP	000713		
L0523       162*       14DHIU(H12,9)       000723         G0521       183*       C180       HRITE(6,186)       (CQUMID(M12,M13),H13=1,8),(ELDAT2(M12,M13),H13=1,4),C       000723         U0523       184*       C       14DHIU(H12,9)       000723         U0523       184*       C       14DHIU(H12,9)       000723         U0523       184*       C       14DHIU(H12,9)       000723         U0523       185*       186       FORMAT(815,2F1U.G,20X/2F10.0,15,5X)       000755         U0540       186*       U0 19U K1=1,ISTOP       000755         U0543       187*       19D       IDY31(K1)=CQUMID(K1,8)       000755         U0545       188*       RETURN       000757       000757         U0546       189*       END       END       001051         ENU FOK        001051       001051		181*	180RITE(12,186)(CQDHID(M12,H13),H13=1,8),(ELDAT2(M12,H13),H13=1,4),C			
GD523       183*       C180       WRITE(6,186)       (CQUMID(M12,M13),M13=1,8),(ELDAT2(M12,M13),M13=1,4),C       000723         U0523       184*       C       14DHID(M12,9)       000723         L0537       185*       186       FORMAT(815,2F10,0,15,5X)       000755         U0540       186*       U0 190 K1=1,ISTOP       000755         U0543       187*       190       IDy31(K1)=CQUMID(K1,8)       000755         U0545       188*       RETURN       000757       000757         U0546       189*       END       END       000751	66523	102*	14DHID(H12,9)	600723		
L0523       184*       C       140HID(H12,9)       U00723         L0537       185*       186       FORMAT(815,2F10.0,15,5X)       G00755         U0540       186*       U0 190 K1=1,ISTOP       G00755         U0543       187*       190       IDy31(K1)=CQUMID(K1,8)       000755         U0545       188*       RETURN       000757         U0546       189*       END       IN01051         L0546       189*       END       IN01051	60523	1 <u>8</u> 3*	C180 HRITE(6,186) (CQUHID(H12,H13),H13=1,8),(ELDAT2(H12,H13),H13=1,4),C			
L0537         185*         186         FORMAT(815,2F10.0,15,5X)         GOU755           U0540         186*         U0 19U K1=1,ISTOP         GOU755           U0543         187*         190         IOy31(K1)=CQUMID(K1,8)         OD0755           U0545         188*         RETURN         GOU757           U0546         1 ₈ 9*         END         U01051		184 *	C 14DHID(M12,9)			
UD54G     186*     UD 19U K1=1,ISTOP     UD0755       UD543     187*     19D     IDy31(K1)=CQUMID(K1,8)     000755       UD545     188*     RETURN     000757       UD546     159*     END     001051       ENU FOR		185*	186 FORMAT(815,2F10.0,20X/2F10.0,15,5X)	000755		
UD543         187*         190         IDy31(k1)=CQUMID(K1,8)         D00755           U0545         188*         RETURN         D00757           U0546         189*         END         U01051           ENU FOR         U01051         U01051	ប្រទ4ជ	186*	UO 190 K1=1,ISTOP	000755		
00545 188* RETURN 000757 10546 159* END ENU FOR	60543	187*	190 IDV31(K1)=COUMID(K1.8)			
_UD546 189* ΕΝΟ ΕΝΟ FOR		188*	RETURN			
END FOR	_LO546	189*				
>HDG,P         (DMEH2	ENU FOR					
	GHDG,P	4DM	IEM2			
				· · · · · · · · · · · · · · · · · · ·		
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@FOR,S	QDHEM2 , QDHEM2					
	-07/30/80-09:46:24	4 (28,)	·····	• • • • • • • • • • • • • • • • • • •	· · · · · · · · · · · · · · · · · · ·	
SUBROL	ITINE GDHEM2 EI	NTRY POINT GOUG64				
STORAC	E USED: CODE(1)	000732; DATA(U) 005575;	BLANK COMMONIZI DOCUDO	, , <u>, , , , , , , , , , , , , , , , , </u>	· · · · · · · · · · · · · · · · · · ·	· · · · · · · · · · · · · · · · · · ·
Сонн	N BLOCKS:	······································				
ម0 _{ប៉} 3	BACH 000044					
<u> </u>	MUZART 000334					
EXTER	AL REFERENCES (BI	LOCK, NAME)			·····	· · · · · · · · ·
ບປັນ5	CANDP					
	FURM2			····	<u> </u>	
0007 6010	STOR PAIR					
	ELORD			<u> </u>		
	NHOUS			<u> </u>		
υ013 	NIO25 NRDU5					
6015	NI035	· · · · · · · · · · · · · · · · · · ·			·····	• • • • • • • • • • • • • • • • • • • •
<u>J016</u>	NIO15					
6017	NERR 35		·······	<u></u>		
S LORAI	E ASSIGNMENT (BI	LOCK. TYPE. RELATIVE LOC.	ATION. NAME)	·····		
0000	005351 111F	COJO 065404 1117F	0000 005362 112F	0000 nri5521 1135F	0000 005414	116F
6001	000370 120L	6000 005416 127F	0000 005425 130F	0000 005436 131F	C000 C05453	
	005472 133F	0000 005505 134F	0000 005527 136F	<u>6001 00067 1456</u>	0001 000076	
0001 	000210 2156 000324 2756	0001 000217 2226 0001 000060 3L	0001 000266 252g 0000 005343 300F	6001 000273 2566 0001 000325 3006	0001 000301 000301 000335	
	000365 3226	0001 000401 3346	<u>0000 005343 300F</u> 0001 000406 340G	0000 005333 349F	0001 C00426	
<u> 6001</u>	000467 377G	0001 000607 4506	UOC1 000614 454G	0001 CU0620 460G	0001 000641	47UG
ັດຖາດ	JU5265 SF	0000 005303 501F	0000 005332 504F	GOO1 GUO103 61L	0001 000110	
<u></u>	<u>000231 86L</u>	<u>6000 065264 7F</u> 6000 I 001274 CQUMID	Unon 005323 795F	0001 CU0227 84L 0000 R 005254 DUM	0001 000224 CCUO R 005220 1	
	<u>R 003410 DUMMY</u>	_ 1.0.00 R 004403 ELDATZ	GOOD I <u>DOOGDO CCDM1</u> 	0000 R 005254 00H	003244	
- JUG J	005547 INJP\$	0000 I 005242 IX	0060 I 005227 J	6000 I 005225 JJ1	0000 1 005240	J10
	I 05237 J11	6000 I 305243 312	0000 I 005244 J13	0000 I 005232 K	0000 I nn5233	
	I UU5224 KKDE <u>I UU5252 K1U</u>	GUUD I 005234 KKODE GOUD I 005250 K11	<u>0000 I 005235 KKOL</u> <u>0000 I 005245 K6</u>	0000 I 005253 KROW	0000 I 005263 (	
	I 005252 KIU	UGUG I GUSZSU KII.		<u>0060 I 005246 K7</u> 0004 R 000024 HAT12	0000 I 005247   0000 I 004033	
นิยวบ	R DC4103 HTPR02	1000 R 064127 MTPR03	0000 I 005255 ml	0060 I 005261 H12	<u>0000 I 065262 I</u>	113
	I UUS230 NEID	OCOU I OUS226 NNCOL	UDUD I UD5257 NNCOL1	UNGO I DU5260 NNCOL2	U000 I 004345 M	
	<u>I 115231 NPID</u> I 003554 PQDM1	<u>UUUD I DU5241 NROW</u> DDUD R DU3636 PQDM2	<u>0003 000022 NSTART</u> 0000 I 004343 TYPE	UTUD I CUDUOD NUM UTUD I CU4354 VAR	0000 I 005236   0000 I 001130 \	
			0.000 I 004343 HTE	UIUU I UU4334 FAR	0000 1 001130	·····

QD	MEM2			DATE 073080	PAGE	2
60103	2*		COMMON/BACH/NUH(18),NSTART(18)/MOZART/MAT11(20),MAT12(20,10)	000000		
U0103	3=	( <b>*</b> *	CHANGE GROUP. NO. OF DESAP PLANE STRESS QUADRILATERALSINASTRAN	000000		
60103	4 *	C **	CUADRILATERAL MEMBRANE CONSTRUCTION CODE 2	000000		
60104	5*		INTEGER CODH1(100,6), VEID(100), CODHID(100,8), MATCH(100,2), INUEX(10	00000		
00104	6*		10)	00000		
60105	7*		DIMENSION DUMMY(100,1)	000000		
0105	8 *	C **	CHANGE GROUP. NO. OF PID'S	000000		
60106	9 🛊		INTEGER PODM1(25,2)	000000		
00107	<b>1</b> U*		DIMENSION POUMZ(25,5)	000000	•	
60107	11+	Ŭ **	CHANGE GROUP. NO. OF MID.S	000000		
<u> </u>	12*		DIMENSION MTPRO1(20,2)	000000		
00111	13*		REAL MTPRO2(20), MTPR03(20,7)	000000		
60111	14*	C **	CHANGE GROUP. NO. OF DESIGN VARIABLES	000000		
L0112	15+		DIMENSION IDV32(50)	000000		
UG112	16*	C **		000000		
66113	17*		INTEGER KOAD(2), TYPE(2), NPAR(7), VAR(4)	000000		
60114	18+		DIHENSION EHUL(4,4),ELDAT2(100,4),DUHH(4)	000000		
60115	19#		REAL MAT12	000000		
66116	20*		DATA VAR/*A*,*B*,*C*,+D+/TYPE/*QUADHE*,*H */	000000		
00121	21*	7	FORMAT( )	000000		
L0122	22*		KKDE=7	000000		
L0123	23*		CALL CANUP(KKDE+NUM7,NUM14,CQDM1,DUMMY,PQDH1,PQDM2)	000001		
<u>LC124</u>	24#		WRITE(6,5)	000012		
60126	25*	5	FORHAT(2x, PLANE STRESS QUADRILATERAL (QUAD. MEMBRANE IN NASTRAN),	000017		
<u> </u>	26*		1CONSTRUCTION CODE 2°)	000017		
00127	27*		HPAR (1)=3	606017		
<u>UC130</u>	28*	-	NPAR(2)=NUM(7)	000021		
60131	29*		NPAR (4)=1	U00023		
L0132	<u>\$</u> 0*		NPAR(5)=2	_000025		
60133	31+		NUH7=NUH(7)	000027		
<u>U0134</u>	32 +		ISTOP=NUM7	000030		
60135	33*		JJ1=2	600031		
60136	<u>34</u> +		IF(NHOUE • EQ • 1) GU TO 3			
60140	35*		CALL FORM2(JJ1,KOAD, TYPE, NPAR(2), VEID, NUM7)	000034		
00141	≉6ز		ISTOP=NPAR(2)			
60142	37≠		NNCOL=6	606046		
00143	38 <u>*</u>		CALL STORIISTOP, NUM7, CQDM1, VEID, NNCOL)			
60144	39÷	3	DO 70 J=1,ISTOP	00000		
60147	40≠		NEID=CQDM1(J,1)	000067		
60156	41*		NPID=CuDM1(J,2)	000071		
00150	42*	<u> </u>	FIND PODM1 CARD WITH SAME ID	000071		
LU151	43*		U0 65 K=1,NUH14	000076		
60154	44+		IF(NPID,E0,PGDH1(K,1))GO TO 61	000076		
LU156	45 =		60 10 65	000101		
60157	46*	61		000103		
ÚD16Ú	47*		60 10 62	000104		
00161	48*	65	CONTINUE	606110		
60163	49*	62	CQDMID(J,1);NEID	000110		
00164	50+		Canito of Ci-canito (2)	000112		
00165	51*		CCDHID(J,3)=CQUH1(J,4)	000114		
<u>CU166</u>	<u> </u>		CCDMID(J,4)=CQUM1(J,5)	000116		
C0167	53*		CODMID(J,5)=COUM1(J,6)	000120		
<u>L0170</u>	54+		CODMID(J,6)=FCUM1(kK,2)	L00122		
60171	55*	7U	CONTINUE	L00126		
00171	56*	<u>ر **</u>	FIND NUMMAT AND MATCH	L00126		
00173	57*		KKODE= J	000126		
LO174	58×		_KKOL=6	U00127		

90	MEM2		DATE 073n80	PAGE	3
u0175_	59*	CALL PAIRIKKODE.CODMID.PODM1.ISTOP.NUM18.KKGL.NUMMAT.MATCH.INDEX)	000131		
J0176	63*	NPAR(3)=NUMMAT	000144		
<u>00177</u>	61*	wRÎTE(6,501)	000146		
60201	62*	501 FORMATIZX, "ENTER CODE FOR INCOMPATIBLE DISPL. NODES: D MEANS USE T	000153		
<u> </u>	63*	IHEM: 1 MEANS DO NOT USE THEM .)	000153		
00202	64+	READ(5,7)NPAR(6)	600153		
00205	65*	<u>kRITE(12,795)</u>	<u> </u>		
00207 00210	66 <b>*</b> 67≠	795 FORMAT(' COM _m Element Control From QDMEM2') WRITE(12,504)(NPAR(J11),J11=1,6)	000166 600166		
60210	68*	504 FORMAT(615)	000210		
60213	69*	C ** SET UP MATERIAL PROPERITIES DATA	000210		
60214	70+	LO 83 JIU=1,NUMMAT	000210		
00217	71=	MTPR01(J10,1)=INDEx(J10)	000210		
Ü <u>n</u> 22ŭ	72*	MTPR01(J10,2)=1	000212		
60221	<u>73</u> *	DO 84 J11=1+NUMMAT	000217		
60224	74 <del>×</del>	IF(HATCH(J11,1),EQ.INDEX(J10))G0 T0 85	600217		
L0226		<u>60 TO 84</u>	000222		
00227	76*	85 NROW=MATCH(J11,2)	000224		
<u> </u>		<u>bo to 86</u>	000225		
00231	78*	84 CONTINUE	000231		
<u> </u>	<u>79≭</u> 80≭	86 HTPRO2(J1D)=HAT12(NROW,4)+CONFAC	000231		
60235	50≢ 51≠	MTPRU3(J10,1)=0. MTPRU3(J10,2)=MAT12(NROW,1)	000234		
<u> </u>	82*	MTPRO3(J10,3)=MAT12(NROW,3)	000237		
60230	83*	HTPRO3(,10,4)=0.	000241		
60240	84×	MTPR03(J10,5)=MAT12(NROW,8)	600242		
LJ241	85*	MTPP03(J10,6)=HAT12(NROW,9)	000244		
00242	66#	MTPR03(J10,7)=MAT12(NROW,1J)	606246		
00243	87*	IX_INDEX(JIG)	000250		
00244	88*	83 CONTINUE	600254		
00244	. 89*	C ** MATERIAL PROPERTIES DATA IS DONE	600254		
60246	90+	WRITE(12,349)	000254		
<u> </u>	<u> </u>	349 FORMAT('COMM MATERIAL PROPERTIES FROM ODMEM2')	<u> </u>		
60251	92*	DO 35D J12=1,NUMHAT	U00266		
<u>00254</u> 00254	<u>93</u> # 94#	350 WRITE(12,300)(HTPR01(J12,J13),J13=1,2),HTPR02(J12),(HTPR03(J12,J13 1),J13=1,7)	000266		
60276	95*	300 FORMAT(215,F10,3/2F10,0,F10,3,4F10,0)	600314		
60270	96*	C ** GENERATE ELEMENT LOAD MULTIPLIERS	600314		
	97*	wRITE(6,111)	000314		
<u> </u>	98÷	111 FORMATIIUX, LLEMENT LOAD MULTIPLIERS FOR CONST. CODE 2")	000325		
UC274		UO 115 K6=1.4	000325		
00277	100*	DO 115 K7=1,5	UC0325		
<u> </u>	101*	115 EMIL(K6.K7)-0.	600325		
u0305	162÷	UO 120 K8=1,4	L00335		
<u> </u>	103*	WRITE(6,112)VAR(k8)	600340		
00313	104≄ 105≉	112 FORMATIZX, "FUR LOAD ", A1," ENTER PRESSURE AND X, Y, Z ELEMENT LOAD F	000346		
<u>00313</u> 60314	<u>105#</u> 1,6*	IRACTIONS IN THAT ORDER, DEFAULT=ALL ZEROS*) REAU(5,7)(DUMH(K11),K'11=1,4)	<u>600346</u> tind346		
00317	167*	IF(DUMM(1).Eu.)GU TO 120	000356		
60321	108*	n0 113  k11=2.5	000365	<i></i>	
6 <u>0</u> 324	107*	113 EMUL(K6,K11)=DUMM(K11)	COU365		
L0326	11)*		600371		
60330	111*	WRITE(12,1117)	000371		
60332	112*	1117 FORMAT(" COMM ELÉMENT LOAD MULTIPLIERS FROM GDMEM2")	675461		
60333	113*	00 121 K9=1,4	000401		
60336	114#	121 WRITE(12,116)(EMUL(K9,K10),K10=1,5)	608401		
<u>L0345</u>	115*	116 FORMAT(5F10.0)	000415		

<u>QDH</u>	EH2		DATE 073080	PAGE	4
60345	116#	C ** GENERATE ELEMENT DATA	000415		
60346	117.	WRITE(6,127)	000415		
6035 ₁₀	118¥	127 FORMAT(15x, +ELEMENT DATA FOR CONST. CODE 2")	000426		
00351	119#	DO 150 KROW_1,ISTOP	000426		
00354	120*	IX=C_DHID(KROu,1)	000426		
60355	121+	WRITE(6,130) IX	000427		
ມ <u>ກ</u> 36 ມ	122#	130 FORMATIZX, "ENTER DESIGN VARIABLE NO., IDV, FOR EL. NO." 151	000435		
U0361	123*	READ(5,7)CUDHID(KRUW,7)	600435		
LD 364	124*	WRITE(6.13) IX	000443		
00367	125#	131 FORMAT(2X, ENTER DESIGN VARIABLE FRACTION , FRC, FOR EL. NO. *, 14, *.	000451		
<u>ú0367</u>	126*	1 DEFAULT=1.*J	000451		
60376	127*	ELDAT2(KROW,1)=1.	000451		
<u>60371</u>	128*	READ(5,7)DUM	660453		
60374	129*	IF(DUM.NE.D.)ELDAT2(KROW,1)=DUM	600461		
60376	<u>130*</u>	<u> </u>	000467		
UU401	131*	145 ELDAT2(KROW, M1)=0.	000467		
60403	132*	<u>urite(6,132)Ix</u>	000470		
00406	133*	132 FORMATIZX, 'ENTER COMP. FORCE/UNIT LENGTH APPLIED TO SIDE 1-2 OF EL	000476		
60466	134*	1.NO. *, I4, *. DEFAULTED. *)	000476		-
60407	135*	READ(5,7)DUM	000476		
00412	136*	IF (DUM.NE.D.)ELDAT2(KROW,3)=DUM	600504	· · · · · · · · · · · · · · · · · · ·	
60414	137+	WRITE(6,133)IX	000510		
66417	138*	133 FORMAT(2X. ENTER THE ANGLE BETA FOR EL. NO DEFAULT=0)	000516		
60420	139*	READ(5,7)DUH	000516		
<u>uCu23</u>	140*	IF (DUM+NE+G+)ELDAT2(KROH+4)=DUM	000524		
60425	141#	WRITE(6,134)IX	600530		
<u>L0430</u>	142*	134 FORMATIZX, "ENTER STRESS PRINTOUT CODE, KS, FOR EL. NO. "+14, " DEFAUL			
00430	143*	17:3*)	000536		
00431	144*	CQDHID(KROW, 6)_3	000536		
60432 6n435	145* 146*	READ(5,7)IDUH	000540		
LU437	147*	IF(ILUM.NE.O)CQDMID(KROW.8)=IDUM 150 CONTINUE	<u>000546</u> 000557		
60437	148#	NNCOL1=7	000557		
60442	149*	NNCOL2:4	000561		
60443	150+	CALL ELORD(ISTOP,CQDMID,ELUAT2.NNCOL1,NNCOL2)	000563		
00443	151*	C PRINT 179	000563		
60443	152#	C179 FORMATI// ELEMENT DATA DECORD ON FILE 12. IN QDMEM2'/)	000563		
<u>60444</u>	153*	WRITE(12,1135)	000572		
u0446	154¢	1135 FORMATI" COMM ELEMENT DATA FROM QDMEMZ")	000577		
60447	155*	DO 160 H12=1.ISTOP	000577		
00452	156*	180 WRITE(12,136)(COUMID(M12,M13),M13=1,7),(ELDAT2(M12,M13),M13=1,4),C	UDD607		
60452	157*		600607		
L0452	158*	C180 wRITE(6,136) (COUMID(M12,M13),M13=1.7).(ELUAT2(M12,M13).M13=1.4).C	000607		
LU452	159#	C 19DHID(M12.8)	000607	· <u> </u>	
00466	160*	136 FORMAT(715, F., U, 3F10, D, 15, 5X)	000641		
LC467	161=	UO 190 K1=1,ISTOP	006641		
00472	162*	190 10y32(K1)=CQDHID(K1,7)	000641		
60474	163*	RETURN	600643		
<u>611475</u>	164*	END	000731		
END FOR					
wHDG,P					<u> </u>
<u></u>	<u>₩'</u>				

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LU101	1+		SUBROUTINE QTICONFAC, NKODE, IUVS, NCTOT)	00000
L0103	2*		<u>COMMON/BACH/NUM(18),NSTART(18)/HOZART/MAT11(20),MAT12(20,1n)</u>	00000
66103	3*	C **		000000
00104	4.⊅		<u>INTEGER CQUAUIII0,6),CTRII(100,5),CQTID(100,7),HATCH(100,2),INDEX(1</u>	
66104	5*		1(100),CQT1(100,6),CQT2(100,1)	00000
60105	6≭		<u>DIMENSION_CQUAD2(10C+1)+CTRI2(100+1)+ELDAT2(100+6)</u>	00000
60105	7*	C **	CHANGE GROUP. NO. OF PID'S	00000
00106	8 🕈		INTEGER POUAU1(25,2), PITRI1(25,2), POIT1(25,2)	600000
60107	9*		DIMENSION PQUAU2(25,2), PTRI2(25,2), PQT2(25,2)	00000
60107	10×	C **		C00000
L0110	11*		INTEGER MTPRO1(20,2)	000000
ŭŭiii	12*		REAL HTPR02120,11,HTPR03120,7)	000000
00111	13+	C **	CHANGE GROUP. NO. OF DESIGN VARIABLES	000000
L0112	14 =		DIMENSION IDV5(50)	000000
60112	15*	C **	**********	00000
60113	16*	-	DIMENSION NPAR(5), EMUL(5,4), VAR(4)	000000
UD114	17*		RFAL MATI2	00000
00115	18*		UATA VAR/*A', 'B', 'C', 'D'/	000nnn
UD117	19#	7		60000
60120	20×	•	NUM9=NUM(9)	00000
L0121	21+		NUM1C=NUM(10)	C00001
60122	22*		NUH17:NUH(17)	606003
60123	23*		NUM167NUM(16)	
60123	23# 24#			
00124			NUH16=HUH(16)	60007
	25*		NCTOT=NUH9+NUH10	000011
60126	26*		NPTOT=NUH16+NUH17	000013
00127	21*		KKDE=9	000015
00130	28≠		IF (NUM9, NE, 0) CALL CANDP(KKDE, NUM9, NUM16, CQUAD1, CQUAD2, PQUAD1, PQUA	00017
60136	294		102)	00017
<u>U0132</u>	30*		KKDE=11	600032
66133	31*		IF{NUMID+NE+G}CALL CANDP(KKDE,NUM10,NUM17,CTRI1,CTRI2,PTRI1,PTRI2)	600034
<u> </u>			IF (NUM9+EQ+6)60.To 40	000052
60137	33*		D0 25 LR=1,NUM9	000054 .
00142	34*		U0 25 LC=1.6	000063
60145	35*	25	CGT1(LR,LC)=CQUAD1(LR,LC)	000063
60150	36*		UQ 26 LR=1+NUM9	<u>CODO74</u>
60153	37*	26	CQT2(LR,1)=CQUAD2(LR,1)	000074
60155	38*	<u>-</u>	UO 27 LR=1.NUA16	000107
<u>ČO160</u>	39×		00 27 LC=1,2	000107
UG163	40.4		POTI(LR.LC)=PQUAD1(LR.LC)	000107
00164	41*	27	PGT2(LR,LC)=PQUAU2(LR,LC)	C00110
60167	42*	<b>4</b> Q	IF (NUM10.EC.)60 TO 50	000120
C0171	43≠		ISTARC=NUM9+1	000121
UC172	44*		ISTARP=NUM16+1	000124
L0173	45*		DO 31 LR=ISTARC, NCTOT	L00136
60176	46×		00 31 LC=1.5	600156
<u>LOŻU1</u>	47*	31	CQT1(LK,LC)=CTRI1(LR_NUM9,LC)	600156
00204	48*		DO 32 LREISTARC. NCTOT	600173
60207	49#	32		600173
00207	50*	~ -	UO 33 LREISTARC, NCTOT	LOU202
60214	51*	33	CQT2(LR,1)=CTR12(LR-NUM9,1)	C00202
66216	52*	~ ~	UO 34 LREISTARP, NPTOT	000220
60221	<u> </u>	•••	UO 34 LC=1.2	J00220
60224	54≠		PQT1(LR,LC)=PTR11(LR-NUM16.LC)	600220
00224	<u> </u>	34		
00225	55+ 56+	54 C	PQT2(LR,LC)=PTHI2(LR-NUM16,LC)	600221
	707	L L	PRINT 30	00221

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60225	57¢	C 3 D	FORMATI' COTI AND POTI ARRAY IN QT")	000221		
UC225	58.	C	UO 36 I=1.NCTOT	000221		
60225	59*	C 36	PRINT 35. (COT1(1,J), J=1,6), (POT1(1,J), J=1,2)	000221		
60225	60+	C 3 5	FORMAT(615,215)	600221		
	61*	50	NKODE=1	006234		
<mark>ن 23 تا تا</mark>	62*	C **	THERE IS ONLY UNE CONST. CODE FOR THIS ELEMENT	600234		
E0231	63*		<u>NPAR(1)=6</u>	606235		
60232	64 #		NPAR (2)=NCTOT	600237		
	65*		NPAR14)=1	000241		
66234	66*		NPAR (5)=1	ÚDÚ242		
<u>µ0235</u>	<u> </u>		<u></u>			
É <u>0</u> 246	68*		NEID=cut1(J,1)	000252		
60241	69*		NPID=CQT1(J,2)	000254		
00242	73*		00 65 N=1,NPTOT	000261		
L0245	<u>71a</u>		<u>IFINPID-EQ-POTL(K-1)1GQ TO 61</u>	606261		
60247 60250	72*		GO TO 65	000264		
	<u>75*</u> 74*		<u>KK=K</u> Go To 62	000266		
60251	74# 			000267		
<u></u>	76#	65	CONTINUE CQTIU(J,1)=NEID	<u>000273</u>		
LU255	77*	62	COTID(J.2)=COTI(J.3)	<u>000275</u>		
6 ₆ 255	78*		CQTID(J.3)=CQT1(J,4)	<u>U00277</u>	· · · · · · · · · · · · · · · · · · ·	
00257			COTID(J.4)=CUTI/J.5)	600301		
00260			CQTID(J+5)=CUTI(J+6)	606303		
60260	50.÷ 81.≠		CQTIU(J,6)=PUTI(KK.2)			
LU262	62*	70	CONTINUE	000311		
0262	63¥	c	PRINT 71	000311		
Ur 262	84*	C71	FORMAT(* CUTID ARRAY IN OT JUST BEFORE GOING TO PAIR*/)	690311		
00262	65¥	C .	U0 72 1=1.ACTOT	600311		
60262	86*	C72	PRINT 73, (CutID(I,J),J=1,6)	600311		
L0262	6.7*	C73	FORMAT(615)	000311		
60262	68¥	( **	FIND NUMMAT AND MATCH	L00311		
00264	69*		KKOLE=ù			
L0265	¢0≉		KKOL=6	000312		
L0266	<u>.91#</u>		<u>CALL PAIRIKKODE.CQIID.POII.NCTOI.NUHIB.KKOL.NUHHAT.HATCH.INDEX)</u>	<u> </u>		
60267	92+		NPAR (3) = NUMMAT	000327		
<u>0027ú</u>	<u> </u>		<u>WpItt(12,795)</u>	600331	····	
00272	94 #	795	FORMAT(* COMM ELEMENT CONTROL FROM QT*)	000336		
60273	<u> </u>			<u>UDU336</u>		
60276	96×		FORMAT(515)	606360		
<u> </u>	<u> </u>	<u> </u>	SET UP MATERIAL PROPERTIES DATA	000360		
60277 60302	98* 		UO 63 JIUEI NUHMAT	600360		
<u> </u>	<u> </u>	- , <u></u>	MTPR01(J10.1)-INUEX(J10)	<u> </u>		
	163≠ 161≠		$HTPR_01(J1_0,2)=1$	000362		
<u> </u>	<u>    161*    </u> 162 •		<u>D0 64 J11=1,NUMHAT</u> IF(MATCH(J11,1).EQ.IN _D EX(J10))60 TO 85	006367		
UC311	1624		<u>60 TO 84</u>	UDU372		
<u> </u>	<u>1640</u>	85	ND04-MATCH(+)1 2)	ngu374		
60313	105*		<u>60 T0 86</u>	<u>600375</u>		
00314	164	84	CONTINUE	600401		
60316	107#	86	MTPRG2(J1C)=MAT12(NROW,4)+CONFAC	600401		
00317	168*	<u>×</u> •	MTPR03(J10,1)=0.	000404		
	149*		MTPR03(J10,2)=MAT12(NROW,1)	000404		
60321	110+	•	HTPR03(J10, 3)=MAT12(NROW, 3)	000407		
00322	111+		_MTPR03(J10,4)=0			
ÜC323	112*		MTPRU3(J10,5)=MAT12(NRUW,8)	000412		
60324	113+		_MTPR03(J10,6)=MAT12(NPON.9)			

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00325	114=	MTPR03(J10,7)=HA_12(NROy,10)	000416		
00326	115*	IX=INDEX(J10)	000420		
00327	<u>116</u> ¢	83 CONTINUE	000424		
60327	117+	C ++ MATERIAL PROPERTIES DATA IS DONE	000424		
60331	118*	WRITE(12,349)	000424		
60333	119+	349 FORMATI' COMM MATERIAL PROPERTIES FROM QT')	600436		
60334	123*	Do 350 J12=1, NUMMAT	000436		
00337	121*	350 WRITE(12,300)(MTPR01(J12,J13),J13=1,2),MTPR02(J12),(MTPR03(J12,J13	000436		
<u> </u>	<u> </u>	1),J13=1,7) 300 FORMAT(215,F10.3/2F10.0,F10.3,4F10.0)	<u> </u>	·	
00353	124*	C ** GENERATE ELEMENT LOAD HULTIPLIERS	000464		
i0354	125*	WRITE(6.111)	000464		<u> </u>
60356	126*	111 FORMAT(15X, 'ELEMENT LOAD MULTIPLIERS')	000475		
00357	127*	U0 115 K6=1,5	000475		
00362	128#	D0 115 K7=1,4	000475		
CD365	129*	115 EMUL(K6,K7)=0.	000475	_	
00370	130*	D0 120 K6=1,4	C00506		
60373	131*	WRITE(6,112)VAR(K8)	000506		
60376	<u>132#</u>	112 FORMATIZX. FOR LOAD .A1. ENTER LATERAL PRESSURE AND X.Y.Z FLEMEN	<u>u00513</u> _		
60376	133×	1T LOAD FRACTIONS,IN THAT ORDER. DEFAULT= ALL ZEROS*)	000513		
00377	134#	READ(5.7)DUH1.DUH2.DUH3.DUH4	000513		
00405	135*	IF(DUM1+EQ+D+)60 TO 120	000524		
<u> 40407</u>	136*	EMUL (1, K6) = DUM1	000526		
6841ŭ	137÷	EMUL(3,K8)=DUM2	080530		
1.1.411	138#	EMUL(4,KB)=DUM3	<u> </u>		
68412	139#	EHUL(5,K8)=DUM4	000534		
00413	140*	<u>120 CONTINUE</u> WRITE(12,1117)	000541		
UO415 UO417	141* 142*	1117 FORMATI' COMM ELEMENT LOAD MULTIPLIERS FRom (1")	600541		
60426	142#	DO 121 K9=1,5	<u> </u>		·
00420	144*	121 WRITE(12,116)(EMUL(K9,K10),K1n=1,4)			
00432	145+	116 FORMAT(4F10.0)	000565		
60432	146#	C ++ GENERATE ELEMENT DATA	000565		
LU433	147*	WRITE(6,127)	000565		
60435	148 *	127 FORMATIISX, "ELEMENT DATA"	000575		
00436	149#	DO 150 KROW=1,NCTOT	600575		
60441	<u>150</u> *	Iv=COTID(KROW, 1)	000575		
ÚC442	151#	wRITE(6,130)IX	000576		
<u>UC445</u>	152*	130 FORMAT(2X, "ENTER DESIGN VARIABLE NO IDV, FOR EL. NO. ", 15)	000604		
L0446	153*	READ(5,7)CGTID(KROW,7)	000604		
L0451	154+	WRITEIO, 1311IX	600612		
60454	155*	131 FORMAT(2X, "ENTER LATERAL PRESSURE ON EL. NO.", 14, ". DEFAULT=0.")	000620		
00455	156*	ELDAT2(KROW+1)=0+	600620		
00456	157*	READ(S,7)DUH Telaum Ng B Ng DAT2(kRou 1)=(km	600621		
<u>611461</u> 00463	<u>158*</u> 159*	IF(DUM.NE.G.)ELDAT2(KROW.1)=DUM ELUAT2(KROW.2)=D.	<u>000627</u> 000633		
60464	163*	write(6.132)1x	U00633		
<u>00467</u>	161.	132 FORMAT(2X, "ENTER DESIGN VARIABLE FRACTION, FRC, FOR EL. NO.", 14, * D	000642		·
60467	162*	LEFAULT=1,*)			
U047u	163#	ELDAT2(KROW, 3)=1	000642		
40471	164*		LOQ644		
L0474	165*	IF (DUM.NE.O.)ELDAT2(KROW,3)=DUM	600652		
60476	166*	WRITE(6,133)1X	000656		
605 ₀ 1	167*	133 FORMAT(2X, . ENTER THE ANGLE BETA FOR EL. NO. ", 14, ". DEFAULT="")	000664		
<u> </u>	168*	£LDAT2(KKÔW+4)=0	<u>UDU664</u>		
66503	167*	REAU(5,7)DUM	000665		
5լլ6	170*	IF(LUM.NE.D.)ELDAT2(KRON.4)=DUM	000673		

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00510171+	150 CONTINUE	006761	
60512 172*	NNCOL1=7	000701	
L0513 173*	NNCOL 2 = 4		
60514 174*	CALL ELORDINCTOT, CQTID, ELDATZ, NNCOL1, NNCOL2)	600705	
<u>un515 175+</u>	WRITE112,1135) 1135 Format(° Comm el. Data from ut°)	<u> </u>	
UO517 176* UC520 177*	1135 FORMAT(* COMH EL. DATA FROM UT*) DO 180 M12=1.NCTUT	000721 600721	
60523 178*	180 WRITE(12,136)(CQTID(H12,H13),H13=1,7),(ELDAT2(H12,H13),	H13=1,4) 000730	
<u>20523 179*</u>		113=1,41 000730	
00536 180*	136 FORMAT(615,5X,15,4F10.0)	C00760	
LD537 181#	DO 190 K1=1.NCTOT		
UU542 182≠	190 IDV5(K1)=CQTID(K1,7)	000760	
60544 183*	<u>PETURN</u>	000762	<u>.                                    </u>
00545 184* END FOR	END	G01033	
äHDG,P REMAR	P		
OPRIS KENAP			
FURPUR 20R1H1 E36	6 S74T11 07/30/80 09:46:32		
		· · · · · · · · · · · · · · · · · · ·	

*)* 

SHEAR	
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DATE 073080 PAGE 1

aFOR, S SHEAR, SHEAR

HSA E3 -07/30/80-09:46:49 (30,)

SUBROUTINE SHEAR ENTRY POINT DOU527

STORAGE USED: CODE(1) DDD560; DATA(0) D05060; BLANK COMMON(2) DDDUDO

COMMON BLOCKS:

UUU3 BACH UUU044 UUU4 MUZART 000334

EXTERNAL REFERENCES (BLOCK, NAME)

	······		··· · - · - · · · · · · · · ·			
ŰŰ	OS CANDP.					
	U6PAIR					
UD UD	G7 ELORD					
06	116 NEDUS					
ίιU	11 NI025					
60	12 NI035					
00	13 NIO15					
	114 NRDUS					
	115 NERR35					
	<u> </u>					
S To	RAGE ASSTGNMENT (	BLOCK . TYPE . RELATIVE L	OCATION. NAMEL			
μΩ	00 064677 111F	0000 aa473a 1117F	6000 004710 112F		SE 0000 004740 116E	
	U1 060317 120L	0000 064742 127F	0000 004751 130F	0000 004762 131	•	
	UU UU5017 136F	6001 000044 1376				
	01 000221 2326	0001 0L0226 236G	6001 000234 2436			
	11 <u>000271 2666</u>	<u> </u>	· · · · · · · · · · · · · · · · · · ·			
¥¥	US		<u> </u>		<u>6 uon1 non355 3326</u>	

-IXXX

000064677_111F	COOD	6000 004710 112F	0000 005011 1135E	0000 004740 116F
UDU1 000317 120L	0000 004742 127F	0000 004751 130F	0000 004762 131F	0000 004774 132F
0000 005017 136F	<u>0001 000044 1376</u>	0001 000053 144G	0001 000152 201G	0001 000161 2066
JOO1 000221 2326	0001 060226 236G	6001 000234 243g	0001 no0257 255G	0001 000260 Ž60G
1	<u>6000 064672 300F</u>	JC01 000331 3146	0001 000336 3206	001000355_332G
ບົບບັບ _{ບໍ່ບົ} 4663 349F	6001 060461 3766	0001 000466 4026	0001 000504 412g	0000 004647 5F
0000 004662 504F		0001 000065 62L	0001 0U0063 65L	0000 004646 7F
0000 004653 795F	6031 060171 84L	U001 000166 85L	0001 000173 86L	0060 I 001130 CSHRID
<u>0000 1 060000 CSHR1</u>	0030 R 004640 DUM	<u>0000 p.003244.00MMy</u>	0000 R 004631 DUM1	000n R 004632 00H2
UUUU R 004633 DUM3	0000 R 003410 ELDAT2	0000 R 004572 EMUL	0000 I 003100 INDEX	0000 005033 INJPs
6000 I 004637 IX	<u>0000 I 004611 J</u>	000I_004621_J10	0000 T 004622 J11	0000 I 004624 J12
UUUU I 625 J13	00JO I 004614 K	ύ600 Ι ₀₀ 4615 KK	0000 I 004610 KKDE	6000 I 004617 KKODE
COUD I 004616 KKOL	0000 I 004636 KRON	U000 I 004645 K1	0000 I 004635 K10	0000 I 004626 K6
6830 I 884627 K7	6000 I 064636 KB	UCOD I 004634 K9	0000 I 002570 MATCH	0004 000000 HAT11
004 R 000024 MAT12	0020 I 004343 HTPR01	0000 R 004413 MTPR02	0000 R 004437 MTPR03	000n I 004643 M12
0000 I 004644 M13	COUD I 004612 NEID	0000 I 004641 NNCOL1	0000 I 004642 NNCOL2	U000 I 004561 NPAR
0000 I 004613 NPID	0060 I 004623 NROW	0003 000022 NSTART	0003 T 000000 NUM	0000 I 004620 NUMMAT
0000 1 304606 NUM15	LOUD I 004607 NUM18	0000 I 004064 PSHR1	0000 R 004146 PSHR2	0000 I 004557 TYPE
GOOD I DG4566 VAR				

60101	1*	SUBROUTINE SHEAR (CONFAC, NKODE, IDV4, NUMB)	00000	•
<u>C0103</u>	2*	<u>COMMON/BACH/NUM(18).NSTART(16)/MOZART/MAT11(20).MAT12(20,10)</u>		
60103	3*	C ** CHANGE GROUP. NO. OF SHEAR PANELS	000000	, .
60104	4*	INTEGER_CSHR1(100,6).CSHRID(100,8).MATCH(100,2).INDEX(100)		

5н	LAR			DATE 073080	PAGE	Z
00105	<u>5</u> *		DIMENSION	000000		
C0105	6*	C ++	CHANGE GROUP. NO. OF PID'S	000000		
<u>UU106</u>	7*		INTEGER PSHR1(25,2)	<u>600000</u>		
60107	8 🕈		DIMENSION PSHR2(25,5)	U000 ₀₀		
00107	9*	<u>C</u> \$*	CHANGE GROUP. NO. OF MID.S	000000		
00110	10+		INTEGER MTPRO1(20,2)	00000		
<u> </u>	<u> </u>	C **	REAL HTPR02(20), HTPp03(20,4)	000000		
60112	124	C ##	CHANGE GROUP. NO. OF DESIGN VARIABLES	000000		
00112	14#	C **	DIMENSION IDV4(50)	000000		<del>-</del>
60113	15+	C ++	INTEGER TYPE(2),NPAR(5),VAR(4)	000000		
<u>u0114</u>	16*		REAL EMUL(3,4), HAT12	000000	······································	
60115	17*		UATA VAR/*A*,*B*,*C*.*D*/TYPE/*SHEAR*.* */			
L012U	18*	7	FORMAT ( )	000000		
UD121	19*	•	NUM8=NUM(8)			
UU122	20+		NUM15=NUM(15)	000001		
L0123	21#		NUM16=NUM(18)	600003		
00124	22*		KKDE_8	000005		
00125	23*		CALL CANDP(KKDE, NUH8.NUH15.CSHR1.DUHMY.PSHR1.PSHR2)	600007		
U0125	24*	с	PRINT 13	U00007		
<u>пп125</u>	25*	C13	FORMATI/ PSHR1 IN SHEAR JUST BACK ERON CANDP 1/1	000007		
60125	26=	С	00 15 J=1,NUH15	000007		
60125	27+	C15	PRINT 14. (PSHR1(J.K),K=1,2)	00007	_	
60125	26*	C14	FORMAT(216)	600007		
00126	29*		LRITE (6,5)	000020		
E013U	30≠	5	FORMAT (20X, "SHEAP PANEL")	600025		
00131	31*		NKODE=1	000025		·
60131	32*	C **	THERE IS ONLY ONE CONST. CODE FOR SHEAR PANELS	608025		
<u>LU132</u>	<u>33</u> *		NPAR (1) = 4	00027		
60133	34*		NPAR (2)=NUMB	000031		
00134	<u> </u>		NPAR141=1	666633		
60135	36*		NPAR(5)=1	606034		
<u></u>	37*		0076 JE1.NUMB			
60141	38*		NEID=CSHR1(J,1)	600044		
<u>60142</u> 60143	<u>39</u> ≠ 4ù≉		NPIDECSHR1(J+2)			
C0146	41*		D0 65 K=1,NUM15 IF(NPID.EQ.PSHR1(K.1))G0 T0 61	600053		
U0150	42+		GO TO 65	600053		
60151	43#	61	50 10 85 KK=K	000 ₀ 56		
L0152	<u> </u>		GQ TQ 62	<u> </u>		
00153	45*	65		UD0065		
	46*	62		U00065		
60156	47#		CSHRID(J.2)=CSHR1(J.3)	000067		
60157	48*		CSHRID(J,3)=CSHR1(J,4)	000071	<u> </u>	
00160	49#		CSHRID(J,4)_CSHR1(J,5)	000073		
00161	50*		CSHR1D(J,5)=CSHR1(J,6)	000075		
60162	51*		CSHR1D(J,6)=PSHR1(KK,2)			
L0163	52*	70	CONTINUE	606163		
60163	53*	<u> </u>	PRINT 71	000103		
LB163	54 ±	C71	FORMATE CSHRID ARRAY JUST BEFORE LEAVING SHEAR FOR PAIR ")	200103		
00163	55*	Ç	<u>00 72 J=1,HUMB</u>	e001n3		
60163	56*	C72	PRINT 73, (CSHRID(J,K),K=1,7)	000103		
60163	57*	<u>C73</u>	FORMAT(716)	000103	·	·
60163	58×	C **	FIND NUMMAT AND MATCH	000103		
LD165	<u>59*</u>			000103		
60166	` ຽວີ≄		KKODE=U	000105		
60167	01*		CALL PAIRIKKUDE, CSHRID, PSHRI, NUMB, NUM16, KKOL, NUMMAT, MATCH, INDEX)	000106		

<u>S</u> H	EAR		DATE 073080	PAGE	3
0017 ₁₁	62*	NPAR (3) = NUMMAT	000121		
00171	63*	WgITE(12,795)	000123		
60173	64*	795 FORMAT( COMM ELEMENT CONTROL FROM SHEAR )	000130	····	
60174	65*	WRITE(12,504)(NPAR(J),J=1,5)	000130		
60177	66*	504 FORMAT (515)	600152		
60177	67 <b>+</b>	C ** SET UP HATERIAL PROPERTIES DATA	600152		
. 0200	68*	00 83 J1G=1,NUHHAT			
00203	69*	MTPRO1(J10,1)=INDEX(J10)	600152		
<u>60204</u>	70*	MTPR01(J10,2)=1	000154		
CO2U5	71*	DO 84 J11=1, NUMMAT	000161		
<u>60216</u> 60212	<u> </u>		<u> </u>		
L0212	74≠	85 NROW=MATCH(J11.2)	000166		
00214	75*	GO TO 86	000167		
_60215	76*	84 CONTINUE	000173		
00217	77#	B6 HTPRO2(J10)_HAT12(NROW,4)+CONFAC	000173		
60220	78*	MTPRo3(J10,1)=0.	600176		
6 ₀ 221	79+	HTPRO3(J10,2)=HAT12(NR ₀ W,1)	600177		
<u> hũ222</u>	81) 🕈	MTPR03(J10,3)=HAT12(NR09,3)	000201		
UC223	81*	MTPR03(J10,4)=MAT12(NROW,10)	000203		
00224	62*	83 CONTINUE			
60226	ь3 <b>*</b>	WRITE(12,349)	600207		
00230	84 *	349 FORMAT(* COMM MATERIAL PROPERTIES FROM SHEAR*)	<u>000221</u>		
L0231	85*	DO 350 J12=1,NUMMAT	600221		
	<u> </u>	350. WRITE(12,300)(HTPR01(J12,J13),J13=1,2),WTpR02(J12),(HTPR03(J12,J13	000221		
LU234	87*	1),J13=1,4)	600221		
<u>00250</u>	68*	300 FORMAT (215,F10,3/2F10,0,F10,3,F10,0)	000247		
00250	89*	C ** GENERATE ELEMENT LOAD HULTIPLIERS	600247		
00251	90*	WPITE (6.111)	000247		
UO253 LO254	91×	111 FORMAT (1DX, "ELEMENT LOAD HULTIPLIERS FOR CONST. CODE 1")	000260		
60257	<u>92*</u> 93*	DO 115 K6=1,3 DO 115 K7=1,4	<u> </u>		
LG26Z	94±	115 EMULIK6.K7)=G.	000280		
00265	95#	D0 120 K8=1,4	000271	· · · · · ·	<u> </u>
UQ276	96#	WRITE (6.112) VAR(K8)	600271	,	
60273	97*	112 FORMAT 12X, 'FOR LOAD ', A1, ' ENTER X, Y, Z ELEMENT LOAD FRACTIONS, IN	000276		
	<u> 98</u> ≠	1 THAT ORDER. DEFAULT =ALL ZEROS 1	600276		
60274	99*	REAU (5,7)DUH1,DUH2,DUH3	000276		
L0301	100*	IF 10UML + EQ + U + )GO TO 120	000306		
00303	101*	EMUL(1,K8)=DUM1	C00310		
<u> </u>	1u2*	EMUL (2+K8)=DUM2	000312		
00305	163*	EMUL (3, K6)=DUM3	000314		
<u> </u>	<u>    104*                                </u>		000321		
00310	105*	WRITE(12,1117)	600321		
<u> </u>	<u>106*</u> 107*		000331		
00315	107≠ 1⊔8≠	121 K9=1,3 121 WRITE(12,116)(EMUL(K9,K10),K10-1,4)	000331		
<u>uu</u> 325	109*	116 FORMAT (4F10+0)	<u>00331</u> C00345		
60325	110+	C ** GENERATE ELEMENT DATA	600345		
L0326	111+	WRITE (6,127)	ü00345		
	.112#	127 FORMAT (15x, ELEMENT DATA FOR CONST.CODE 2")	000345		
00331	113*	LO 15U KROW=1.NUM8	000355		· · ·
00334	114*	IX=CSHpIU(KRUW,1)	600355		
60335	115*	WRITE (6,130)TX	000356		
00340	116#	130 FORMAT 12X. "ENTER DESIGN VARIABLE NO IDV.FOR EL. NO.". 15)	000364		
L0341	117+	READ(5,7) CSHRID(KROw,7)	000364		
60344	118#	WRITE (6,131)IX	<u> </u>		

	EAR			DATE 073080	PAGE 4
00347	119*	131	FORMAT (2X, "ENIER EDGE SUPPORT CONDITION, ISU, FOR EL. NO.", ISI	600400	
u ₀ 35u	120#		READ (5,7)CSHRID(KROW,8)	600400	
60353	121*		WRITE (6,132)1X	000406	
LU356	122#	132	FORMAT (2X, "ENTER DESIGN VARIABLE FRACTION, FRC, FOR EL. NO.", 14, "	000414	
UC 356	<u>123*</u>		1DEFAULT=1)	000414	
ÚÖ357	124*		ELDAT2(KROW,1)=1.	006414	
ن 26 0 م	_125*		READ (5,7)DUH	000416	· · · · · · · · · · · · · · · · · · ·
L0363	126*		IF (UUN+NE+D+JELDAT2(KROH+1)=DUM	000424	
<u>C0365</u>	127*	<u>    15    </u>	CONTINUE	000432	
L0367	120*		NNCOL1=8	600432	
<u>16370</u>	129#		NNCOLZ=3	<u> </u>	····
60371	130*		CALL ELORD (NUH8, CSHRID, ELDAT2, NNCOL1, NNCOL2)	000436	
00371	131#	<u> </u>	PRINT 179	000436	
60371	132*	C179	FORMAT(//* ELEMENT DATA RECORD ON FILE 12, IN SHEAR*/)	000436	
L0372	$\frac{133 *}{134 *}$	1170	WRITE(12,1135) Format(" Comm el. data from Shear")	000445	
	134# 135*	1132		000452	
<u>10375</u> 1040u	<u>135</u> * 136*	190	D0 160 H12=1,NUH8 WRITE (12,136)(CSHRID(H12,H13),H13=1,8),ELDAT2(H12,1)	<u>000452</u> 000461	
00400	138#	100		<u> </u>	
0041u	138+	174	FORMAT (815,F10.0,25x)	000504	
00410	139#	¥ 30	D0 190 K1=1.NUM8	688504	
00414	140*	190	IDV4(K1):CSHRID(K1,7)	000504	
UC416	141*	170	RETURN	000506	
00417	142#		ENU	000557	
END FOR					····
⊙HÚG+P	SI6	MA			
•HÚG+P	516	HA			
•HUG+P	SI 6	Η A			
•нџб,Р	SIU!	HA			
•нџб,Р	SIG	H A			
•нџб,Р	SI 6	H A			·
•нџб,Р	SI 6!	M A			
•нџс,р	SIG	H A			
•HUG,P		M A			

SIGNA			DATE U73080	PAGE	_1
GFOR, S SIGMA, SIGMA		<u> </u>			
HSA E3 -07/30/86-09:46:53 (6,)					
SUBROUTINE SIGMA ENTRY POIN	1_600036				
STORAGE USED: CODE(1) DUD044; D	ATA(D) 000052; BLANK COMMON(2) 0000	20		<u> </u>	
EXTERNAL REFERENCES (BLOCK . NAM	F)				
UDUJ NPRTS					
0004 NIO25 0005 NRDUS					
UUUG NERR3S			<u> </u>		
STORAGE ASSIGNMENT (BLOCK, TYP	E. RELATIVE LOCATION, NAME)				
4030 940001 10F COUL	000016 1116 0000 000036 7F	0000 00042 INJPs		L	
		- <u> </u>			
U0191 1≠ SUBROUTINE <u>10103 2≠ DIMENSION</u>	SIGMA(MATID,LLL,XMAT12) XMAT12(20.10)		000005		
	ATID		000005		
10 FOPMAT(//*	FOR MATERIAL NO 15 ENTER ALLOW	BLE TENSILE.COMPRESS	000016		
L0107 5* 11VE, AND SH	EAR STRENGTHS, IN THAT ORDER ON THREE	LINES. ",/ * BE SURE	600016		
	FORMAT WITH DECIMAL POINT !/ )		000016		
C0110 7≠ U0 15 J=6, Cn113 8≠ 15 READ(5,7)X			000016		
<u>UI113 8* 15 READ(5,7)X</u> UU113 9* C 15 MAT12(LLL	MAT12(LLL,J) ,J)=MATID*1000.+J*100.		600016		
40117 10* 7 FORMATIF15	.2)		000024		
UD117 11* C PRINT 50			000024		
60117 12* C 00 20 J=8.	10		000024		
00117 13* C 20 PRINT 60,L	LL,J,XMAT12(LLL,J)		000024		
<u>CU117 14★ C 50 FORMAT(* X</u> CD117 15* C 69 FORMAT(* X	MAIL2 IN SIGMA") MATL2(+ ,I3,',' ,I3,')=',E12.5)		<u>000024</u> 000024		
GO120 16* RETURN	HAITSIA 1231.1. 1231.1-12151		000024		
LC121 17* E ND			000043		
END FOR					
WHDG,P STOR					
·			<u></u>		<u> </u>
		· · · · · · · · · · · · · · · · · · ·			
		**************************************		. <u> </u>	

STO	٠			DATE UT3080 PAGE	1
aFUR,S	STUR	108			
HSA E3	-07/30/80	J-09:47	9 (18,)		
SUBROI	UTINE STO	)R	NTRY POINT LOUISO		
STORAL	E USED:	CODE(1	DG0152; DATA(G) CG1174; BLANK COMMON(2) D00000		
EXTER	AL REFER	RENCES	LOCK, NAME)	·····	
<u> </u>	NERR 3	L			
STORAL	E ASSIGN	MENT	LOCK, TYPE, RELATIVE LOCATION, NAME)	·	
ບບິບ 1 ປມີປີ 1	000031	105G	COC1 000046 1106 0001 000055 1176 0001 00072 130 6001 000045 211 0000 1 001131 1 0000 1 C01132 II	G 0001 000073 133G 0000 001135 INJP\$	
	I 001130	L C	COUD I DEDUDE STORE		
L0101	1*		UBROUTINE STORIISTOP.NUM.C.VEID.NCOL)	000031	
CC101 CC101	2* 3≠		STOP=NPAR(2)= NO. OF ELEMENTS WITH SPECIFIED CONST. CODE. INPUT UM=Total NO. OF C cards for this element. E.G., for Rods NUM=NUM5. If	000031 NPUT 000031	
60101	<u> </u>		IS THE INTEGER ARRAY FORMED FROM THE NUM C CARDS * E.G., CROD.	000031	
<u> </u>	5 *	C **	NPUT/QUTPUT.	000031	
- E0101 - E0101	6* 7*	C ** C **	VEID CONTAINS USER INPUT ELEMENT ID NOS FOR THIS CONST. CODE. INPUT Col= No. OF Cols of integer c array in calling program f.g., for Nasi	U00031	
CO101	8*		<u>COL- NOA OF COLS OF INTEGER C ARRAY IN CALLING PROGRAM FAGAE FOR NAS</u> D ,NCOL=4. INPUT	<u>1RAN000031</u> _000031	
60101	9 <b>≠</b>	( **		u00031	
00101	10×	C **	FOR THE GIVEN EL. AND CONST.CODE THIS SUBROUTINE	000031	
<u>    00101                             </u>	<u>11#</u> 12#	<u>C **</u>	FINDS THE C CARDS AND VEID CARDS WITH MATCHING EID'S, PUTS THE C'S N STORE, THEN PUTS THEM BACK IN C IN THE FIRST ISTOP ROWS.	000031	
	134	++			;
60101	14+	C **	HANGE GROUP. NO. OF ELEMENTS	600031	
	<u>15 ÷</u>		NTEGER C(10C,6),VEID(10C),STORE(100,6)	000031	
60103 60103	16≠ 17≠	C (200	RINT 200 ORMAT(1x,*C JUST INSIDE STOR*)	COUO31 UOO031	
66103	13*	<u> </u>	0 2(1 I=1,NUM	00031	
60103	19*	CZU1	RINT 202. (C(I.J).J=1.4)		<u> </u>
60103	∠ ا⊭		ORMAT(8x,418)	000031	
60103	21*	<u>C</u>	RINT 100, (VEID(I),I=1,NUH)	00031	
60103 601 ₁₃ 4	22* 23*	C100	ORMAT(1X, *VEIU AT INPUT*/2014)	00031	
60107	24*		0 21 J=1, ISTOP 0 19 I=1, NUM	UDD031 UDC040	
66112	25*		F (C(1.1).EC.VEID(J)) GO TO 21	000040	
00114	26*		0 TO 19	606043	
00115	27*	21	ONTINUE	000045	
60116	28≠ 29≠	~ ,	0 22 II=1,NCOL	600045	
LD121 60123	<u>294</u> ວິນ*	<u>22</u> 19	TORE (J.II)= C(I.II) ONTINUE	<u> </u>	
<u>00125</u>	31+	20	ONTINUE	60073	
60127	32*		0 23 1 = 1,15104	000073	
60132	33×		0.23 J = 1 MCOL	L00073	

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510	R			DATE_U73080	PAGE	2
00135	34 *	23	C(I,J)=STORE(I,J) PRINT 102 FORMAT(1X, °C JUST BEFORE LEAVING STOR °) UO 25 I=1,ISTOP PRINT 103, (C(I,J),J=1,6) FORMAT(615) RETURN FND			
00135 00135	35*	C	PRINT 102	U00073		
00135 00135 00135 00135	36*	C102	FORMATIIX, "C JUST BEFORE LEAVING STOR")	000073		
00135	37=	С	UO 25 I=1,ISTOP	000073		
<u>00135</u>	<u> </u>	C25	PRINT 103, (C(1,J),J=1,6)	<u> </u>		
60135	39*	C1U3	FORMAT(615)	000073		
<u>60146</u> 60141	<u>41*</u>		RETURN	000104		
END FOR	41#		END	000151		
aHDG,P	VEC					
		<u> </u>				
<u> </u>						
				· · · · · · · · · · · · · · · · · · ·		
				······································		
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					<u></u>	

_	VEC				DATE 073080	PAGE	1
	FOR,S	VEC.VEC	5				
	HSA ES -			:12 (14,)			
	SUBROU	TINE VEC		ENTRY POINT 000053			
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#### CONCLUSIONS AND RECOMMENDATIONS

Currently NASAP transforms CAD-generated NASTRAN data to data for DESAP I or DESAP II with minimum input from the engineer. The program has been thoroughly checked.

An effort is underway to implement the design cycles shown in Figure 2. Such a capability will greatly enhance the effectiveness of DESAP I and DESAP II by allowing the simultaneous satisfaction of stress, displacement, and buckling constraints.

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#### REFERENCES

- Kiusalaas, J. and Reddy, G. B.; DESAP II: "A Structural Design Program with Stress and Buckling Constraints" Department of Engineering Science and Mechanics, The Pennsylvania State University, University Park, PA, April 1976.
- Kiusalass, J. and Reddy, G. B.; DESAP I: "A Structural Design Program with Stress and Displacement Constraints" Department of Engineering Science and Mechanics, The Pennsylvania State University, University Park, PA, April 1976.

#### 1980

#### NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

DEPLOYABLE STRUCTURE DESIGN FOR THE SCIENCE AND APPLICATIONS SPACE PLATFORM

Prepared by: Henry W. Stoll, Ph.D.

Academic Rank:

University and Department:

University of Wisconsin -Platteville Department of Industrial and Mechanical Engineering

Associate Professor

**H** 

NASA/MSFC: Division: Branch:

MSFC Counterpart:

James K. Harrison

Program Development Special Projects

Date:

Contract No.:

August 1, 1980

NGT-01-002-099 (University of Alabama)

#### DEPLOYABLE STRUCTURE DESIGN FOR THE SCIENCE AND APPLICATIONS SPACE PLATFORM

By

Henry W. Stoll Associate Professor of Mechanical Engineering Univeristy of Wisconsin - Platteville Platteville, Wisconsin

#### ABSTRACT

NASA has long been interested in deployable structure technology as a means for achieving efficient spacecraft packaging for launch. An important application currently under study is the Science and Applications Space Platform (SASP). In this study, basic concepts regarding deployable structures design, including systematic design/classification schemes and a "deployability" criterion, are proposed for use in synthesis, analysis and evaluation of alternative deployable structure designs. Using design guidelines based on SASP requirements and the basic concepts developed, a variety of new designs are synthesized, and these along with previously proposed designs are analyzed and evaluated. Recommendations and conclusions regarding optimal deployable structure design are made based on insights gained in the study.

#### I. INTRODUCTION

NASA has long been interested in deployable structure technology as a means for achieving efficient spacecraft packaging for launch. An important application currently under study by NASA involves the Science and Applications Space Platform (SASP). A representative configuration (Fig. 1) has been evolved which satisfies the SASP objectives of providing accommodation for the simultaneous operation of several payloads of various science and application disciplines. The configuration incorporates three independently oriented platform arms, dedicated to celestrial, solar, and Earth viewing, respectively. Each platform arm provides interface provisions for payload carriers at two locations, and each location has provisions for two interfaces on opposite sides of the platform arm. The platform arms utilize deployable structures which allow compaction for launch in the Orbiter cargo bay.

A variety of deployable structure concepts have been proposed over the years (see Ref. [1]* for example) to meet different spacecraft packaging needs. In addition, several specific design concepts have been proposed for use in SASP, some of which are presently under development by NASA/MSFC. There does not exist, however, any clear-cut agreement regarding a preferred design approach for SASP, nor does it appear that all viable design alternatives have been identified. This report documents a 10-week, one-man effort, conducted between late May 1980 and August 1980, to take a fresh look at the deployable structure design problem in light of SASP design goals and requirements. Emphasis in this study has been placed on the mechanism aspects of the deployable structures problem. Structural considerations have been addressed only in so far as they impact the mechanism problem.

#### II. OBJECTIVES

The principal objectives of this study are (1) to completely define the deployable structure design problem in light of SASP design goals and requirements; (2) develop basic concepts relating to design morphology, systematic synthesis, deployment method, and performance criteria such as deployability, compactability, etc. which can be used to assist in the synthesis, analysis, and evaluation of deployable structure designs; (3) identify alternative deployable structure designs which appear to hold promise for the SASP application;

*Numbers in brackets refer to references at end of paper.

(4) analyze and evaluate new and previously proposed designs; (5) investigate various detail design considerations relating to structural dimensions, control of structural stiffness and damping, etc.; and (6) summarize findings in the form of recommendations for optimal design.

#### III. DESIGN GUIDELINES

For the purposes of this study, it is assumed that, to be acceptable, a deployable structure design should conform to the SASP design goals and requirements as currently envisioned in reference [2]. These requirements are summarized as follows:

(1) <u>State-of-the-Art Constraints</u>. The design should utilize current design techniques and materials in order to minimize the need for development of new technologies and the use of unproven space operations.

(2) Astronaut Participation. Deployment utilizing an astronaut in EVA mode should be avoided. The astronaut may be utilized in a back-up role.

(3) <u>Retract Capability</u>. In order to provide maximum flexibility and contengencies to conduct mission operations, it is desirable, but not necessary, that a retract capability of the deployed structure on orbit be provided.

(4) Payload Utilities Accommodation. The platform structure should facilitate means for distribution of payload utilities in the form of a separate electrical harness to each payload carrier interface (Fig. 1). Each harness will consist of 4-1/0 power cables plus other electrical and communication leads.

(5) Launch Packaging. The envelope of the platform arm in its compacted form must be such that the three platform arms can be accommodated within the Orbiter cargo bay with each of the arms connected to the SASP platform support module.

(6) <u>Misalignment and Distortion</u>. The maximum misalignment between the payload carrier interface and the platform arm mechanical interface at the rotary joint should not exceed <u>+1.0</u> degree. This alignment requirement is to include effects of fabrication tolerances, joint mechanical dead band, and interface misalignment. Dynamic stability of the platform arm alignment will be held to <u>+0.1</u> degree, including effects such as cyclic thermal distortion, structural distortion due to environmental and induced loads, and mechanical dead band.

- (5) The deployment drive should be small, mechanically simple, and straight forward in its operation.
- (6) A retract capability is desirable.
- C. Deployed Structure Considerations
  - (1) The deployed and rigidized structure should have a minimum of dead-band or other unwanted relative motion.
  - (2) Structural properties should not degrade or otherwise vary due to fluctuating loads or other adverse conditions inherent in the space environment.
  - (3) Mounting points of the deployed structure should lie in a plane normal to the longitudinal axis of the structure in order to permit interfacing of the structure with the platform support module and the payload carrier interface.
- D. General Considerations
  - (1) The overall deployable structure concept should be mechanically simple and free from unnecessary design complexities.
  - (2) Structural links and connections should be easily manufacturable from materials suitable for space applications.
  - (3) The design should be insensitive to manufacturing error and tolerance build-up; close tolerances should be avoided.

#### IV. Basic Concepts

The purpose of this section is to develop some basic concepts which can be used to assist in the synthesis, analysis, and evaluation of deployable structure design. As a starting point, it is useful to note that deployable structures operate both as mechanisms during the deployment stage and as structures after they are fully deployed. Conversion of the deployable structure from a mechanism into a structure is facilitated by hardening or freezing one or more of the kinematic pairs which connect the structural links together. It should also be noted that, depending on the deployable structure design, deployment

m = 0		m =	= 1	m	= 2	m = 3		
n	³ 1	n	^j 1	n	j 1	n	j 1	
3	3	4	4	3	2	4	3	
5	6	6	7	5	5	6	6	
7	9	8	10	7	8	8	9	
9	12	10	13	9	11	10	12	
11	15	12	16	11	14	12	15	

TABLE 1. Possible Link-Joint Combinations to Give a Specified Mobility Based on Equation (1)

Examination of the deployable, three-sided planar truss structure of Figure 2a shows it to be unsatisfactory with regard to interfacing and mounting needs. Altering link proportions and mounting point locations gives the configuration of This design is unsatisfactory structurally, but its Figure 2b. square ends meet interfacing and mounting requirements. From Table 1, the next possible mechanism with a mobility of 1 has n = 6,  $j_1 = 7$ . Using this combination results in the foursided planar truss structure depicted in Figure 2c. This configuration is acceptable structurally and meets interfacing and mounting needs. Also, like the three-sided truss (Figure 2a), it requires only one imput motion for constrained deployment and can be converted to a structure by freezing only one joint.

Efforts to synthesize deployable structure designs utilizing more complex arrangements and numbers of links and joints resulted generally in designs which, if they were acceptable structurally, turned out to be various combinations of the three and four-sided planar structures characterized by Figure 2a and 2c. This insight suggests that the simplest deployable structure designs consist of a series of repeating cells, each of which are based on very simple planar structures. This conclusion is substantiated by the fact that most practical designs proposed to date are applications of this approach.

B. <u>Morphological Analysis</u>. Having decided that the simplest and therefore most practical deployable structures are made up of a series of repeating cells, the next step is to carefully examine the morphology of the simple planar structure or cell. Figure 2 shows that each simple planar structure is composed of three different structural links, namely the longeron, the lateral, and the diagonal. Consider now that each of these links can be constructed to have one of five different properties. A link can be rigid in which case its length is fixed; it can be telescoping in which case its length is changeable; it can fold at its midspan such as the longeron of Figure 2a or the diagonal member of Figure 2c; it can be flexible in which case it is rigid in tension but buckles elastically in compression; or it can be separable in which case parts of the link are not contiguous in the compacted mode.

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Different joint types can also be used to connect the various structural links. Observing the single-degree-of-freedom joint restriction limits joint type selection to pinned joints (revolute pairs) and sliding type joints (prismatic pairs). A screw pair can be substituted for a slider to facilitate actuation. In the case of flexible diagonal links, ball and socket joints and wrapping pairs (e.g., cable and drum, belt and pulley, etc.) can be utilized as long as the structure remains properly constrained. Also, properly designed flexural pivots or plastic hinges can be used in place of pinned or ball and socket joints. Finally, similar to the separable link, joints can be made detachable to facilitate more efficient compaction of the structure.

It should be noted that, kinematically speaking, the telescoping link and sliding pair are identical. However, from the standpoint of visualizing deployable structure design, the ability of the telescoping link to change length is an important property and hence it is treated separately from other sliding pairs. In mobility determinations, the telescoping link should be considered as two rigid links joined by a sliding pair. Similarly, the folding link is acutally two rigid links connected by a revolute pair. But, for the purposes of visualizing deployable structure design, it is treated as a particular link type. As will be shown subsequently, this deviation from conventional kinematic thinking leads to a powerful systematic synthesis procedure.

C. Deployment Methods. Deployment of the structure can be facilitated (1) manually using astronauts in EVA, (2) mechanically using remote manipulators, teleoperators, etc.; or (3) automatically using a self-contained drive system. All deployment methods other than the automatic methods may be classified as "erectable" methods. Automated deployment methods may be further subdivided as follows:

<u>Case 1.</u> The entire, multi-cell deployable structure is designed as a linkage having a mobility of 1. Hence, only one link need be driven to produce constrained motion during deployment.

Case 2. Actuating links in each cell are interconnected such that actuation of one link produces quasi-simultaneous deployment of all cells. angles, and other parameters of mechanisms which give insight as to whether a mechanism is a good one or a poor one are commonly used. Examples of these include mechanical advantage, transmission angle, and pressure angle. Many of these have a number of features in common, including the fact that most can be related to the velocity ratios of the mechanism and therefore can be determined solely from the geometry of the mechanism. However, most also depend upon some knowledge of the application of the mechanism, especially of which are the input and output links.

In the case of deployable structures, a more general "index of merit" is required since there are no clearly defined output links. A survey of the literature indicates that the index proposed by Denavit [4] may be useful as a measure of deployability. This index of merit is the determinant ( $\Delta$ ) of the coefficients of the simultaneous equations relating the dependent velocities of a mechanism. When this determinant becomes small, the mechanical advantage also becomes small and the deployability of the structure is reduced. This same determinant also appears in the denominator of the dependent accelerations and all other quantities which require taking derivatives of the loop-closure equation. Consequently, it is true in general that, if this determinant is small, the deployment mechanism will function poorly in all respects - force transmistransformation, sensitivity to manufacturing sion. motion errors, and so on.

The deployability index ( $\Delta$ ) proposed above depends on the deployable structure geometry, link dimensions, and on which link is driven to effect deployment (i.e., rotation of a link, contraction of a telescoping member, etc.). Consequently, the deployability index can be used to determine the most suitable actuation method for a particular design and to compare deployability potential of alternative designs.

Design Complexity. Design complexity is a nebu-F. lous concept which is difficult to quantify and yet it has an important bearing on the acceptability of alternative deployable structure designs. In general, it is safe to say the the probability of achieving high deployment reliability as well as meeting deployed structure alignment and stiffness design goals will decrease with increasing design complexity. When examined carefully, design complexity is found to involve a variety of factors such as manufacturing methods, materials, space, and the case of deployable In structures, these economics. considerations correlate well with the number of kinematic links, n, the number of joints,  $j_1$ , and the mobility, m, of the structure. Mobility is involved because deployment system complexity increases with increasing degrees-of-freedom of the

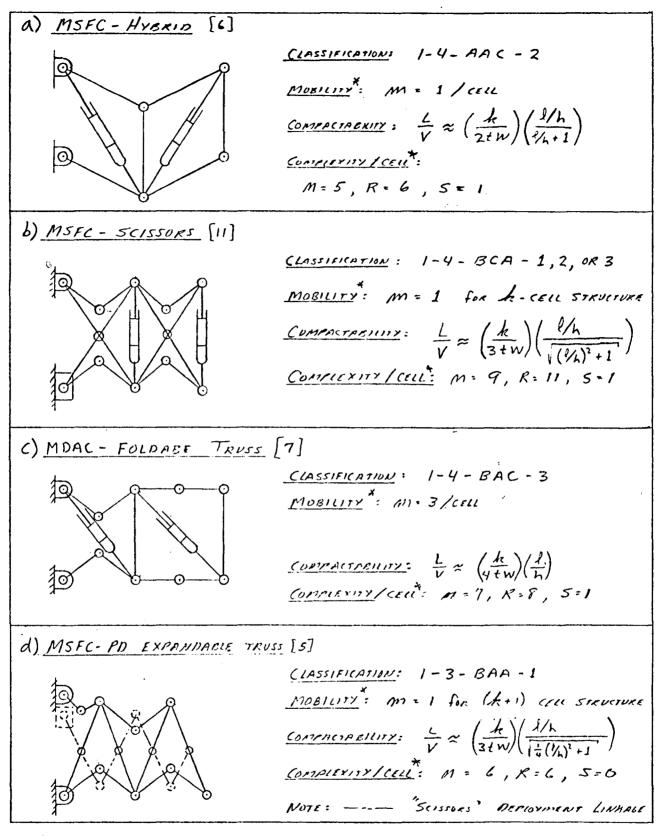
Examination of this approach shows that, although the four considerations proposed are valid, they are not independent of each other in every case. Hence, the real driver for systematic synthesis in this approach is found to be the number of different link type combinations which are available. That is, the five link types (rigid, fold, telescoping, flexible, and separable) taken in groups of three (longeron, lateral/diagonal, and diagonal) results in a total of  $5^3 = 125$  different possible combinations. Further variatons are possible by using different combinations of pinned and sliding joints as well as other joint types.

Possible link type combinations were organized by grouping the three-sided planar truss links (Fig. 2a) in the order of longeron-diagonal-diagonal and grouping the four-sided planar truss in the order of longeron-lateral-diagonal. Each link type was given the letter designation shown in Table 2. This notational convention gives the following 25 column by 5 row array of possible link type combinations for both the three and foursided planar truss cell.

AAA,...,AAE,ABA,...,ABE,ACA,...,ACE,ADA,...,ADE,AEA,...,AEE BAA,...,BAE,BBA,...,BBE,BCA,...,BCE,BDA,...,BDE,BEA,...,BEE CAA, DAA, EAA,....,EEE

In addition to systematic synthesis, the organizational scheme of Table 2 suggests a convenient method for classifying or coding various deployable structure concepts. Employing the notation of Table 2 gives the following alpha-numeric code,

X - X - XXX/XXX - X Deployment Method (Case 1 thru 4 and Erectable) Additional Link Combination groups used as needed. Diagonal Link Type (A,B,C,D, or E) Lateral or Diagonal Link Type (A,B,C,D, or E) Longeron Link Type (A,B,C,D, or E) Basic Repeating Cell Structure Type (3 or 4-sided) Deployment Direction (1,2, or 3- dimensional)



* MOBILINY AND COMPLEYINY ARE CALCULATED FOR PLANAR STRUCTURE

Figure 3. Previously Proposed Deployable Structure Concepts for SASP

XXXII-15

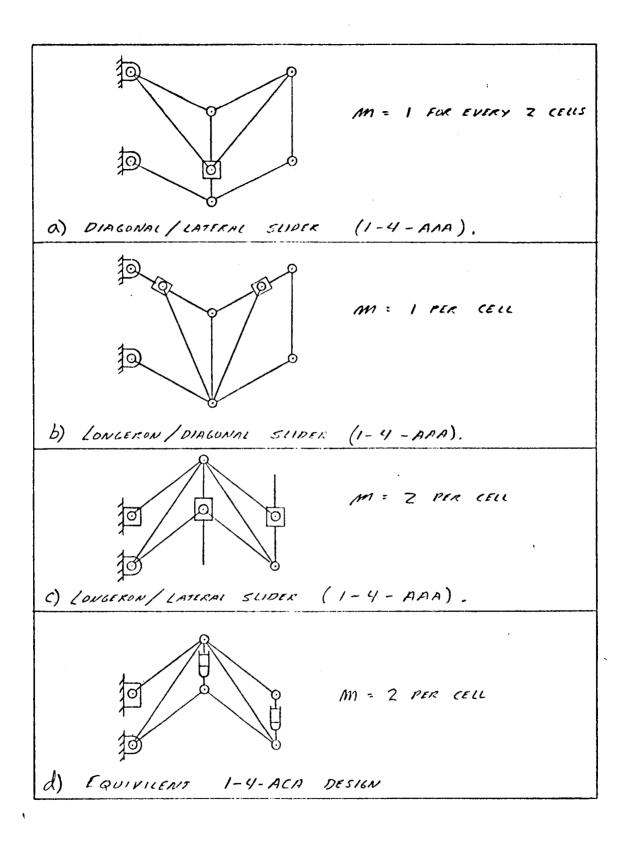
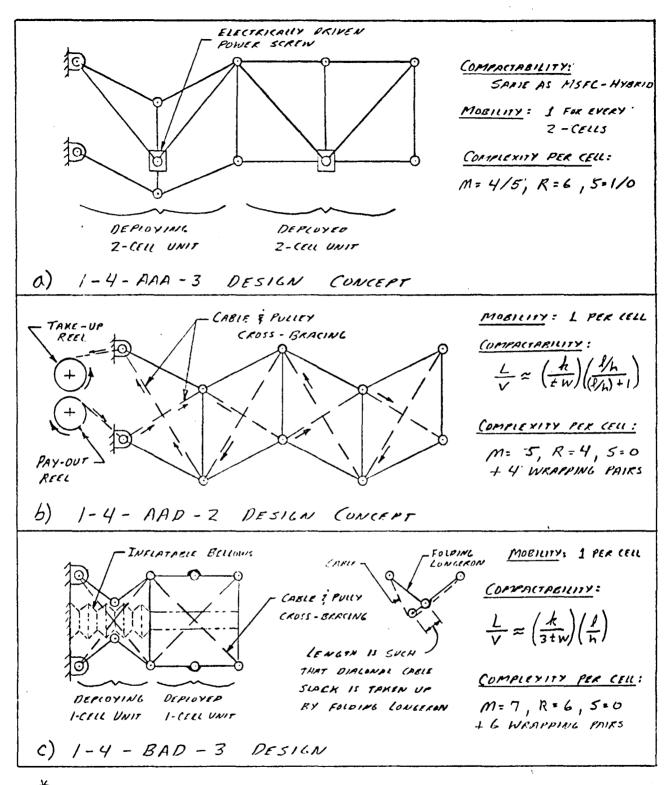


Figure 4. Alternative 1-4-AAA Designs



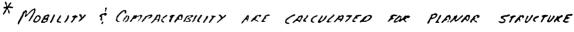
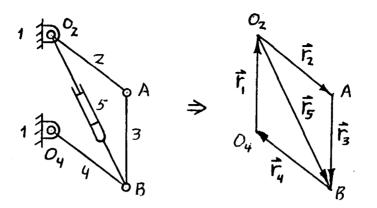


Figure 5. New Deployable Strucutre Concepts for SASP Application

#### XXXII-19



Loop Closure Equations:

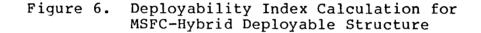
$$\vec{r}_{1} + \vec{r}_{2} + \vec{r}_{3} + \vec{r}_{4} = 0$$

$$\vec{r}_{1} + \vec{r}_{5} + \vec{r}_{4} = 0$$

$$\Delta_{1} = f (\vec{r}_{1}, \vec{r}_{2}, \vec{r}_{3}, \vec{r}_{4}, \vec{r}_{2}, \vec{r}_{4})$$

$$\Delta_{2} = f (\vec{r}_{1}, \vec{r}_{4}, \vec{r}_{5}, \vec{r}_{4}, \vec{r}_{5})$$

$$\Delta_{1} = \int_{1}^{2} \Delta_{1} = \Delta_{1} + \Delta_{2}$$



Use of equation (3) is illustrated in Figure 6. Here it is seen that one cell of the MSFC-Hybrid design (Fig. 3a) contains two loop-closures and hence, its deployability index is the product of the "index of merit" or  $\Delta$  calculated for each loop-closure. Since this design utilizes Case 2 deployment, the deployability index for the cell is also the deployability index `for the structure as a whole.

Examination of the large variety of deployable structure design concepts synthesized in this study shows that each design is made up of different combinations of basic mechanisms or linkages, where each linkage defines a particular loopclosure. The basic linkages comprising the selected designs of Figures 3 and 5 are presented in Figure 7. The deployability index ( $\Delta$ ) for each possible input motion to these linkages has been derived as a function of the link dimensions 1, h, and d and the degree of deployment,  $x^* = x/1$ , where  $0 \le x^* \le 1$ . In addition, each deployability index has been plotted as a function of percent deployment.

The results depicted in Figure 7 can be used in conjunction with equation (3) to gain considerable insight regarding detail design of deployable structures. Remembering It should be noted that a large 1/h ratio (or for a given h, a large 1 dimension) not only improves deployability, but also decreases the number of cells required to produce a required deployed length, L. Hence, design complexity of the overall structure will be reduced by using large 1/h ratios.

B. <u>Compactability Analysis</u>. Based on the design guidelines, a high degree of compactability is not needed in the SASP application. Nor is high compactability particularly desirable in this application because of complications arising due to handling of the utility cables. It is informative, however, to examine the sensitivity of compactability to changes in structural dimensions. Using the ratio of deployed length to compacted volume (L/V) as a measure of compactability, it is found that, for all designs, the magnitude of L/V and therefore, compactability, is increased as the l/h ratio is made larger. This is at least partly due to the reduction in the number of cells produced by a large l dimension. It should be noted that, for designs where L/V ratio involve complicated functions of l/h, compactability is improved by making h small as well as by making l large.

C. <u>Rigidization Considerations</u>. It has been stated previously that the deployable structure can be converted into a structure by hardening or freezing joints until the mobility is less than or equal to zero. The question is where to place the locking device, i.e., which joints should be hardened or frozen.

In a system where energy is not dissipated, the products of deflection and forces are constant (equal to the energy)

energy in = energy out

$$F_{a} \bullet \delta_{a} = F_{b} \bullet \delta_{b} = T_{c} \bullet \Theta_{c} = T_{d} \bullet \Theta_{d}$$
(4)

where a, b, c, and d are arbitrary points on the deployable structure, F is the force in the direction of the deflection  $\delta$ , and T is the torque in the direction of the angular deflection  $\theta$ . The product of velocity and force is also constant (equal to the power)

power in = power out  

$$F_a \bullet V_a = F_b \bullet V_b = T_c \bullet \omega_c = T_d \bullet \omega_d$$
(5)

where V is the linear velocity and  $\omega$  is the angular velocity.

These new concepts, as well as other previously proposed designs, were analyzed and evaluated.

Although no clear-cut, "best design" emerged from this effort, several insights into deployable structure design for the SASP application did become apparent. These insights are summarized by the following conclusions and recommendations:

(1) It is important to keep the deployable structure design as simple as possible in order to achieve required reliability goals. Poor deployability associated with Case 1 deployment, randomness and other "uncertainties" associated with Case 2 deployment, and extra weight and complexity associated with Case 4 deployment make Case 3 deployment appear most appropriate for the SASP application. Designs which minimize the number of deployment drives and are, at the same time simple, should be preferred.

(2) Both deployability and compactability are enhanced by large 1/h ratios. In addition, the number of cells which must be deployed to achieve a desired deployed length is reduced by large 1/h. It is therefore recommended that the largest 1/h ratio consistent with structural and spacecraft requirements be used.

(3) Because of the large number and wide distribution of joints in a deployable structure, it is likely that the dynamic behavior of the structure can be tailored through selective hardening of various joints and use of joints designed to have specific damping properties. It is recommended that the feasibility of such an approach be further investigated and verified.

Although much of the work in this study was directed (4) toward the SASP application, many of the results are applicable to a wide variety of deployable structure needs. The systematic design procedure developed led to a number of promising deployable structure concepts which were rejected, in this study, because they did not fit the SASP application. Since man first ventured into space, deployable structure technology has been an integral part of the spacecraft "packaging" problem. It appears that there will be many future needs for deployable structures in space. In view of this, and to avoid future duplication of effort, it is recommended that the material developed in this study, as well as that developed in other studies concerned with deployable structures, be consolidated into a Deployable Structures Handbook. This handbook should be designed in such a way that it can be used to help identify the best and most appropriate deployable structure approach for a particular application.

#### NASA/ASEE SUMMER FACULTY RESEARCH FELLOWSHIP PROGRAM

## MARSHALL SPACE FLIGHT CENTER THE UNIVERSITY OF ALABAMA

#### PAYLOAD MAINTENANCE COST MODEL

#### FOR THE SPACE TELESCOPE

Prepared By:

William L. White, Ph.D.

Academic Rank:

Jackson State University

Associate Professor

University and Department:

Department of Mathematics

NASA/MSFC: Laboratory: Division: Branch:

MSFC Counterpart:

Systems Analysis and Integration Systems Engineering Flight Systems Analysis

Frank Pizzano

August 1, 1980

Contract No.:

Date:

NGT-01-002-099 (University of Alabama)

XXXIII

#### PAYLOAD MAINTENANCE COST MODEL

#### FOR THE SPACE TELESCOPE

ВY

#### WILLIAM L. WHITE

#### ASSOCIATE PROFESSOR OF MATHEMATICS

#### JACKSON STATE UNIVERSITY

JACKSON, MISSISSIPPI 39217

#### ABSTRACT

The research project focuses on the development of an optimum maintenance cost model for the Space Telescope for a fifteen (15) year mission cycle. A review of various documents and subsequent updates of failure rates and configurations were made. The reliability of the Space Telescope for one year, two and one-half years, and five years were determined using the failure rates and configurations. The failure rates and configurations were also used in the maintenance simulation computer model which simulates the failure patterns for the fifteen (15) year mission life of the Space Telescope. Cost algorithms associated with the maintenance options as indicated by the failure patterns were developed and integrated into the model.

#### INTRODUCTION

This report presents the summer's activities directed toward determining the most feasible maintenance policy for the Space Telescope for a fifteen (15) year mission cycle. These include cost algorithms associated with part replacements and/or complete refurbishment. Participating in the summer's activities with me were Frank Pizzano, my counterpart, Rodney Stewart, Allen Forney, Molly Anderson, and Carol Cleveland.

Two basic tasks were adopted which were as follows:

- (1) To update a maintenance simulation computer model which simulates the failure patterns for the fifteen (15) year mission cycle of the Space Telescope.
- (2) To develop cost algorithms for the maintenance options according to the failure patterns to ascertain the optimum maintenance policy.

In order to accomplish the tasks the activities performed were as follows:

- (1.1) A detailed study of the maintenance simulation computer model.
- (1.2) Updates of failure rates were made for various parts whenever necessary.
- (1.3) Model configuration updates were made where revisions in the specifications called for a reconfiguration.
- (1.4) The reliability of the Space Telescope was determined for one year, two and one-half years, and five years.
- (1.5) Updates were made in the computer program where model configuration changed.
- (2.1) Cost Data were collected and organized for parts whose prices were known or available information would allow a price estimate.
- (2.2) Cost Algorithms were written involving a base-line algorithm, i.e. calculates cost if time schedule is standard and extended algorithm, i.e. takes base-line figure and adds a certain percentage according to the relationship between the time left in the new schedule and the time left in the standard schedule.

(b) Update the maintenance simulation computer model with respect to component/systems configurations, failure rates, and orbit replaceable unit cost estimates.

(c) Run the program and obtain the following cost comparisons against the specification plan of having three (3) in-orbit maintenance actions during the fifteen (15) year mission operating life of the Space Telescope:

(1) Assume in-orbit corrective maintenance actions only for the fifteen (15) year mission operating life with frequencies as dictated by the maintenance simulation computer model.

(2) Assume in-orbit corrective (Plus Preventive) maintenance actions only for the fifteen (15) year mission. (The preventive actions would involve components subject to degradation with age such as batteries.)

(3) Assume in-orbit corrective actions plus ground return actions with frequencies as dictated by the maintenance simulation computer model.

Reliability versus cost will be compared for each case using the Shuttle cost constants plus orbit replaceable unit cost estimates to obtain an indication of a preferable maintenance policy.

#### RELIABILITY OF THE SPACE TELESCOPE

The Space Telescope consist of three major subsystems: An Optical Telescope Assembly, A Support System Module, and The Scientific Instruments. The support system module encloses the optical telescope assembly and scientific instruments and also provide all interfaces with the shuttle orbiter.

The maintenance of the space telescope will involve either in-orbit maintenance or ground return maintenance or both. The items which will be maintained in orbit will be called orbit replaceable units. Each item is given a part number, a string number, and an orbit replaceable unit number for identification. The orbit replaceable units are configured on trays (i.e., A component in which all parts have the same orbit replaceable unit number), in such a way that they are easily accessible to the maintenance crew. The items which will be replaced on ground return will be called ground replaceable units.

The components of the optical telescope assembly were identified and failure rates associated with each component were obtained from references [1,4] to determine the reliability of each item for one year, two and one-half years, and five years.

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However, in the case of the DF-224 computer which consist of three memory units required and six provided with three active and three inactive, the reliability according to reference [3] is given By:

$$R(t) = \tilde{e}^{3\lambda t} + \frac{3\lambda}{\lambda'} \tilde{e}^{3\lambda t} (1 - \tilde{e}^{\lambda' t}) + \frac{3\lambda(3\lambda + \lambda')}{2\lambda'} \tilde{e}^{3\lambda t} (1 - \tilde{e}^{\lambda' t})^{2} + \frac{3\lambda}{6\lambda'^{3}} (3\lambda + \lambda') (3\lambda + 2\lambda') \tilde{e}^{3\lambda t} (1 - \tilde{e}^{\lambda' t})^{3}$$
(5)

Where  $\lambda$  is the failure rate and  $\lambda'$  is the standby failure rate, i.e.,  $\lambda' = 0.1 \lambda$ . The central processor unit and the in-put-output unit are the other components of the DF-224. In these cases there are three provided with one required so the reliability formula is given by:

$$Rd = \bar{e}^{\lambda t} + \frac{\lambda}{\lambda'} \bar{e}^{\lambda t} (1 - \bar{e}^{\lambda' t}) + \frac{\lambda(\lambda + \lambda')}{2\lambda'^2} \bar{e}^{\lambda t} (1 - \bar{e}^{\lambda' t})^2 \quad (6)$$

The program for the computation of reliability of each component is given below.

The reliability of the support systems module for one year, two and one-half years, and five years is given in Table [2].

A breakdown of the base-line algorithm includes cost associated with maintenance analysis and planning, management of the orbit replaceable units, ground maintenance facility, and orbiter launch. The total expenditures being the sum of the cost for the maintenance analysis and planning, orbit replaceable unit management, ground maintenance facility and orbiter launch. The extended algorithm takes the base-line figure and adds a certain percentage according to the relationship between the time left in the new schedule and the time left in the standard schedule. The cost per part is used to determine the total cost for parts being replaced at a downstate.

#### MAINTENANCE SIMULATION MODEL

The maintenance simulation computer model is programmed to test each item once every 72 hours for failure. The failure rate of each item is used to determine the probability that the item has failed. With a random number generator the item is either failed or not failed by comparing a random number with the probability (computed) of failure. Each item is given a part number, a string number, and an orbit replaceable unit number for identification, as items are failed a log is kept of the failures by specifically identifying the failed item. This is done so the computer will know when the number of failed items result in a downstate. The Shuttle is then called for maintenance action. The time of each failure is also logged. The tests can be continued every 72 hours for a 15 year mission cycle. Each 15 year mission cycle test is called a pass. The model has the capabilities of making a number of passes sufficient to predict the mean time to failure with reasonable confidence. Plots of the failures versus time can also be made. This information together with the cost algorithm which is integrated in the model will give the total cost relative to maintenance of the Space Telescope according to the standard schedule as well as cost relative to the simulated schedule.

#### CONCLUSION

A maintenance policy for the Space Telescope is a must if the Space Telescope is to serve the needs of the scientific community effectively. The policy must optimize both cost for maintenance and availability for use by the scientific community. Therefore, the policy must be based on all available resources and information. Recognizing that many parts of the Telescope are of new design, very little is known other than simulated testing concerning the reliability thus the task of defining a policy becomes even more difficult. To this end, generic failure rates are used to determine the reliability of each item. Various configurations of the Space Telescope are continuously being updated in an effort to simulate a model which will more closely parallel the actual performance of the Space Telescope. The reliability of the Space Telescope with updated configurations was determined for one year, two and one-half years, and five years. The hope is that reliability as indicated by the model will provide a basis for a maintenance policy according to a particular time schedule. Due to the lack of acquired data relative to the actual item performance of the Space Telescope the model simulation approach seems appropriate. This approach has been used with much success in the past.

#### REFERENCES

- 1. Contractor's Reliability Allocations
- 2. IBM Nas 5-25630, "SI C&DH Reliability Prediction Report" November 20, 1980
- 3. Memorandum EF01, "Reliability Analysis and Evaluation of DF-224/NSSC-II Computers for Space Telescope," October 1978.
- 4. MIL-HDBK-217 B/C, "Reliability Prediction of Electronic Equipment."

## TABLE 2

#### SUMMARY OF SUPPORT SYSTEMS MODULE ORBIT

## REPLACEABLE UNIT (ORU) RELIABILITY STUDY

## Reliability (Probability of Success)

	R (1 year)	R (21/2 years)	R (5 years)
Support Systems Module with present ORU capability	0.828343	0.214951	0.0 ₅ 3446
Support Systems Module with degraded mission capability	0.949195*	0.540912*	0.0 ₃ 3066*
Reliability enhance- ment with limited LMSC suggested replacement ORU's	0.909337	0.709793	0.378308
(20 ORU's Total)	0.994114*	0.966236*	0.883826*
Reliability enhance- ment with MSFC suggested replacement	0.890368	0.628765	0.246092
ORU's (12 ORU's Total)	0.973377*	0.855933*	0.574936*

NOTE: Present ST Design includes approximately 37 ORU types with 69 total ORU's.

## TABLE 4

## SUMMARY OF TOTAL SPACE TELESCOPE SYSTEMS

## RELIABILITY STUDY

# Reliability (Probability of Success)

		R (1 year)	R (2 1/2 years)	R (5 years)
(1)	Probability of the Space Telescope functioning successfully in accordance with SPEC and including at least 3 of 6 science functions.	0.713088	0.108935	0.06 473
(2)	Item (1) above except including mission degradation allowances.	0.794267	0.290403	0.0 ₄ 54

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