

Fig.1 The modulation of 11 years variation of cosmic ray intensity on the ozone

distribution of various Forbush decrease events as table 1 shown.

Table 1 The distribution of Forbush decrease events

importance season	>1%	>2%	>3%	>5%
winter	119	49	22	4
summer	115	48	28	4
total	234	97	50	8

The method of the statistical study has used the superposed epoch method. The day of the minimum of Forbush decrease is considered as the zero day. The abscissa denotes the time in days for all of the figures in this section. The ordinate represents the ozone content. The horizontal line refers to the average value of the total assembly of  $[O_3]$ . The dashed line represents confidence level with probability 0.05 or the degree of confidence 95%.  $n$  is the number of Forbush decrease events.

2. Analysis results. The statistical study for all Forbush decrease events shows that the disturbance of Forbush decrease on  $[O_3]$  is of statistical significance, and the amplitude of  $[O_3]$  disturbance increases with the increase of Forbush decrease amplitude. The disturbance of Forbush decrease on  $[O_3]$  increases with the increase of the latitudes. But at

tensity to minimum year  $[O_3]$  increased by 7.8% and 5.8% at Kodaikanal and Kagoshima respectively. The increase of  $[O_3]$  decreased with the increase of the latitudes,  $[O_3]$  only increased by 4.8% at Bismarck. But 11 years variation of  $[O_3]$  become indistinct at Resolute in the polar region.

III. The disturbance of Forbush decrease on  $[O_3]$ . 1. Data analysis. In order to study the disturbance stage of Forbush decrease for variant amplitudes on  $[O_3]$ , Forbush decrease events are divided >1%, >2%, >3% and >5% four importances based on their amplitudes. The disturbance analysis of the solar flare and Forbush decrease on vorticity arear index (VAI) showed that the disturbances of VAI have evident seasonal property [9]. So the time occurred Forbush decrease is also divided the winter (10-3 months) and the summer (4-9 months).

According to the seasons and importances occurred Forbush decrease, their variant combination is analysed. There were 234 Forbush decrease events during 1965-1976. The distribution of various Forbush decrease events as table 1 shown.

high latitudes, the effect of the solar cosmic ray and precipitating particles on  $[O_3]$  will be over Forbush decrease, so that the ozone disturbance produced by Forbush decrease all is covered.

The disturbance of Forbush decrease occurred in the winter on  $[O_3]$  is larger and more evident. Usually on the day started Forbush decrease, i.e. about at -1 day,  $[O_3]$  begins increase, and there is a larger increase after the fourth day. The amplitudes of  $[O_3]$  disturbance increase with the increase of the latitudes of the observational stations. Fig. 2 shows the disturbance of  $[O_3]$  for variant latitude regions in the winter. Because the data of  $[O_3]$  observed in the winter at Resolute is a few numbers, the change of  $[O_3]$  can not be given.

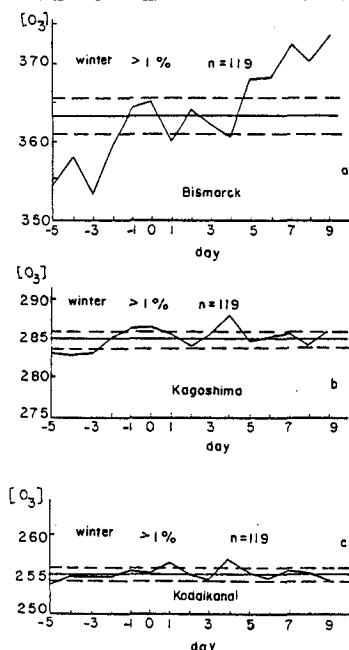


Fig.2 The disturbance of Forbush decrease in the winter on  $[O_3]$   
IV. Conclusions and discussion. According to the results described above, we obtain the following conclusions:

$[O_3]$  is modulated by 11 years variation of cosmic ray intensity. The relationship between the ozone variation and cosmic ray intensity variation is the negative correlation. The 11 years vibration of  $[O_3]$  is most evident at the middle and the low la-

The disturbance Forbush decrease occurred in the winter on  $[O_3]$  at same latitude increases with the increase of Forbush decrease amplitudes. If the average value of  $[O_3]$  from the third day to the fifth day before the zero day is considered as the undisturbance value of  $[O_3]$ , that the largest disturbance of  $>1\%$  Forbush decrease on 0 was about 5.3%. The largest disturbance of  $>2\%$ ,  $>3\%$  and  $>5\%$  Forbush decrease events were 5.9%, 6.2% and 10.3% respectively.

The disturbance of Forbush decrease occurred in the summer on  $[O_3]$  in low latitudes can also be seen, but its disturbance is very less than in the winter. The disturbance of Forbush decrease on  $[O_3]$  all is covered at high latitudes. It always is the decrease tendency after the solar flare, and it lasts several days (see Fig.3).

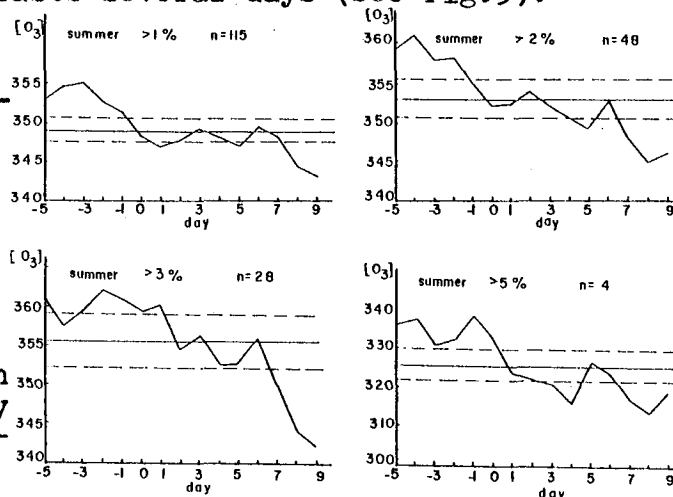


Fig.3 The change of  $[O_3]$  during Forbush decrease in the summer at Bismarck.

titudes. The law of this variation is more indistinct at the high latitudes. Because the influences of the solar cosmic ray and precipitating particles on  $[O_3]$  are so large that the effect caused by the long period variation of cosmic ray intensity is covered.

The disturbance of Forbush decrease occurred in the winter on  $[O_3]$  is very remarkable. The disturbance of  $[O_3]$  evidently increases with the rising of the latitudes. Usually  $[O_3]$  increases after Forbush decrease started, and lasts several days. The disturbance occurs the maximum after the fourth or the fifth day. The amplitude of  $[O_3]$  disturbance also increases with the increase of Forbush decrease amplitude.

The disturbance of Forbush decrease occurred in the summer on  $[O_3]$  is very interesting. The effect of Forbush decrease on  $[O_3]$  is also evident at the low latitudes, but the disturbance of  $[O_3]$  is just opposite to the effect caused by Forbush decrease at the high latitudes. Evidently, this is not produced by Forbush decrease. It is produced by the solar cosmic ray and precipitating particle events. As was shown by Shah [10], the disturbance of the solar cosmic ray events occurred in the summer on VAI is the most remarkable. We discover that the larger solar cosmic ray events occur in the summer. So the disturbance of the solar cosmic ray events occurred in the summer on  $[O_3]$  can also be the most remarkable.

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