Computations of Boiling in Microgravity

G. Tryggvason University of Michigan Ann Arbor, MI 48109 Dave Jacqmin NASA Glenn Research Center Cleveland, OH 44135

ABSTRACT

The absence (or reduction) of gravity, can lead to major changes in boiling heat transfer. On Earth, convection has a major effect on the heat distribution ahead of an evaporation front, and buoyancy determines the motion of the growing bubbles. In microgravity, convection and buoyancy are absent or greatly reduced and the dynamics of the growing vapor bubbles can change in a fundamental way. In particular, the lack of redistribution of heat can lead to a large superheat and explosive growth of bubbles once they form. While considerable efforts have been devoted to examining boiling experimentally, including the effect of microgravity, theoretical and computational work have been limited.

Here, the growth of boiling bubbles is studied by direct numerical simulations where the flow field is fully resolved and the effects of inertia, viscosity, surface deformation, heat conduction and convection, as well as the phase change, are fully accounted for. Boiling involves both fluid flow and heat transfer and thus requires the solution of the Navier-Stokes and the energy equations. The numerical method is based on writing one set of governing transport equations which is valid in both the liquid and vapor phases. This local, single-field formulation incorporates the effect of the interface in the governing equations as source terms acting only at the interface. These sources account for surface tension and latent heat in the equations for conservation of momentum and energy as well as mass transfer across the interface due to phase change. The single-field formulation naturally incorporates the correct mass, momentum and energy balances across the interface. Integration of the conservation equations across the interface directly yields the jump conditions derived in the local instant formulation for twophase systems. In the numerical implementation, the conservation equations for the whole computational domain (both vapor and liquid) are solved using a stationary grid and the phase boundary is followed by a moving unstructured two-dimensional grid. While two-dimensional simulations have been used for preliminary studies and to examine the resolution requirement, the focus is on fully three-dimensional simulations.

The numerical methodology, including the parallelization and grid refinement strategy is discussed, and preliminary results shown. For buoyancy driven flow, the heat transfer is in good agreement with experimental correleations. The changes when gravity is turned off and/or fluid shear is added are discussed, as well as the difference between simulations of a layer freely releasing bubbles versus simulations using only one wavelength initial perturbation. Figure 1 shows the early stages of the formation of a three-dimensional bubble from a thin vapor layer. The boundary conditions are periodic in the x and y direction, the bottom is a hot and the top allows a free outflow. The jagged edge of the surface close to the bottom of the computational domain is due to some of the surface elements being on the other side of the domain and some elements not plotted by our plotting routine. In the second figure, we show the temperature distribution through two perpendicular planes.



Figure 1. The formation of a bubble from a thin vapor layer. The grid representing the phase boundary is shown after the bubble has formed. The jagged edges of the front are an artifact of plotting of the periodic boundaries.



Figure 2. The temperature distribution in a bubble rising from a thin vapor layer. The high temperature in the bubble is represented by a light shades of gray.

Computations of Boiling Flows

1 51 C 14 ---

NASA Glen Research Center Dave Jacqmin Worcester Polytechnic Institute Gretar Tryggvason

Ozgur Ince University of Michigan Asghar Esmaeeli

Motivation

Phase change is the most efficient way of transporting results of design or operator error can be catastrophic. heat (as in boiling and refrigeration). However, the

Effect of gravity, flow, binary solute, thermocapillary, surfactants, electric fields, acoustic, and so on are important but still poorly understood. Motivation

Direct numerical simulations of bubbly flows now allow relatively routine predictions of bubbly flows.





Motivation





The formation of bubbles is the critical **•** part of boiling

\mathbf{O}
Ċ
<u> </u>
~
\mathbf{O}
<u> </u>
$\overline{\Delta}$
Ū.
m
()
5
$\overline{\mathbf{A}}$
U
Ň

2D simulations by:

D. Juric and G. Tryggvason. Computations of Boiling Flows. Int'l. J. Multiphase Flow. 24 (1998), 387-410.

boiling near critical pressures with a level set method G. Son and V.K. Dhir. Numerical simulation of film J. Heat Trans. 120 (1998), 183-192.

S.W.J. Welch and J. Wilson. A volume of fluid based method for fluid flows with phase change. J. Comp. Phys. 160 (2000), 662-682.

Governing Equations	Momentum (conservative form, variable density and viscosity)	$\mathbf{j} + \nabla \cdot \nabla \mathbf{u} = \mathbf{u} + \nabla \cdot \nabla \mathbf{h} + \frac{\partial \phi}{\partial t}$	$+\nabla \cdot \mu (\nabla \mathbf{u} + \nabla^T \mathbf{u}) + \int_F \sigma \kappa \mathbf{n} \delta (\mathbf{x} - \mathbf{x}_f) dA$	Singular Interface Term	Mass conservation (incompressible flows) $\nabla \cdot \mathbf{u} = 0$	Equations of state (properties of each fluid remain constant)	$\frac{D\rho}{Dt} = 0; \frac{D\mu}{Dt} = 0$
---------------------	--------------------------------------------------------------	--------------------------------------------------------------------------------------------------------------------------------	-----------------------------------------------------------------------------------------------------------------------------------------	-------------------------	------------------------------------------------------------------------	---------------------------------------------------------------	----------------------------------------------

I

Numerical Method



101 1





887





Mass conservation

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \, \mathbf{n} = 0$$

The Rankine-Hugoniot conditions give:

$$-V_n(\rho_f - \rho_v) + (u_f \rho_f) - (u_v \rho_v) = 0$$

where

 V_n :Velocity of phase boundary

Rewrite:

 $\rho_f(u_f - V_n) = \rho_v(u_v - V_n) = \dot{m}$

Phase boundary

$$\frac{\dot{m}}{\rho_f} = u_f - V_n$$
 and $\frac{\dot{m}}{\rho_v} = u_v - V_n$

Volume Source (subtract)

$$Q = u_f - u_v = \dot{m} \left(\frac{1}{\rho_f} - \frac{1}{\rho_v} \right)$$

Velocity of phase boundary (add)

Q

 $\left(\begin{array}{c}1\\\rho_{f}\end{array}\right)$

 $V_n = \frac{1}{2} \left(u_f + u_v \right) - \frac{\dot{m}}{2} \left(\frac{1}{2} \left(\frac{1}{2} + \frac{1}{2} \right) \right)$

Phase boundary

The mass transfer at the phase boundary is proportional to the heat source

 $\dot{m} = \frac{\dot{q}}{L}$

where

 $\dot{q} = k \frac{\partial T}{\partial n} \Big|_{1} - k \frac{\partial T}{\partial n} \Big|_{2}$



Computing the heat source at the phase boundary

$$\dot{q} = k \frac{\partial T}{\partial m}\Big|_{1} - k \frac{\partial T}{\partial m}\Big|_{2}$$

Interpolate temperature here



Governing equations

Momentum equation

$$\frac{\partial \boldsymbol{\omega} \mathbf{u}}{\partial t} + \nabla \cdot \boldsymbol{\omega} \mathbf{u} = -\nabla p + \mathbf{f}$$
$$+ \nabla \cdot \mu (\nabla \mathbf{u} + \nabla^T \mathbf{u}) + \int_F \sigma \kappa \, \mathbf{n} \delta (\mathbf{x} - \mathbf{x}_f) dA$$

Energy equation

 $\frac{\partial c_p \rho T}{\partial t} + \nabla \cdot c_p \rho \mathbf{u} T = \nabla \cdot k \nabla T$

 $+ \int \dot{q} \delta(\mathbf{x} - \mathbf{x}_f) dA$

Contraction of the second s

-

Governing equations

Mass conservation

$$\nabla \cdot \mathbf{u} = \int \dot{m} \left(\frac{1}{\rho_f} - \frac{1}{\rho_v} \right) \delta(\mathbf{x} - \mathbf{x}_f) dA$$

Thermodynamic

$$T_b = T_v(P):$$

Q

 ρ_f

 $\dot{q} = L\dot{m}$

Heat source

Normal velocity of boundary
$$\frac{d\mathbf{x}_b}{dt} = V_n \mathbf{n}$$

NASA/CP-2000-210470

Algorithm



Find corrected velocity











Four frames from a simulation of the boiling of hydrogen is shown to the right. The color represents temperature, with red hot and blue cold. The bottom wall has a constant heat flux.











Film Boiling—Velocity

.



Film Boiling—Temperature















Film Boiling—Heat Flux



The average heat flux versus time. Predictions by various empirical correlations are shown by horizontal lines. Nucleate Boiling





Nucleate Boiling



Solidification and the effect of flow on microstructure formation

Dendrite growing in a uniform flow

401 by 401grid

Anisotropy=0.4

 $\mathrm{St} = \frac{c(T_{\infty} - T_m)}{2} = -0.3$ $\text{Re} = \frac{\rho UZ}{\rho 00} = 600$ $\sigma = \frac{cT_m\gamma}{L^2Z} =$







Next Step

- method into our parallel code for fluid flow Implementation of the phase change
- Add adaptive grid refinement capability
- Examination of the effect of gravity and flow on heat transfer
- Adding other effects (electric field, etc)