

Engineering the Future

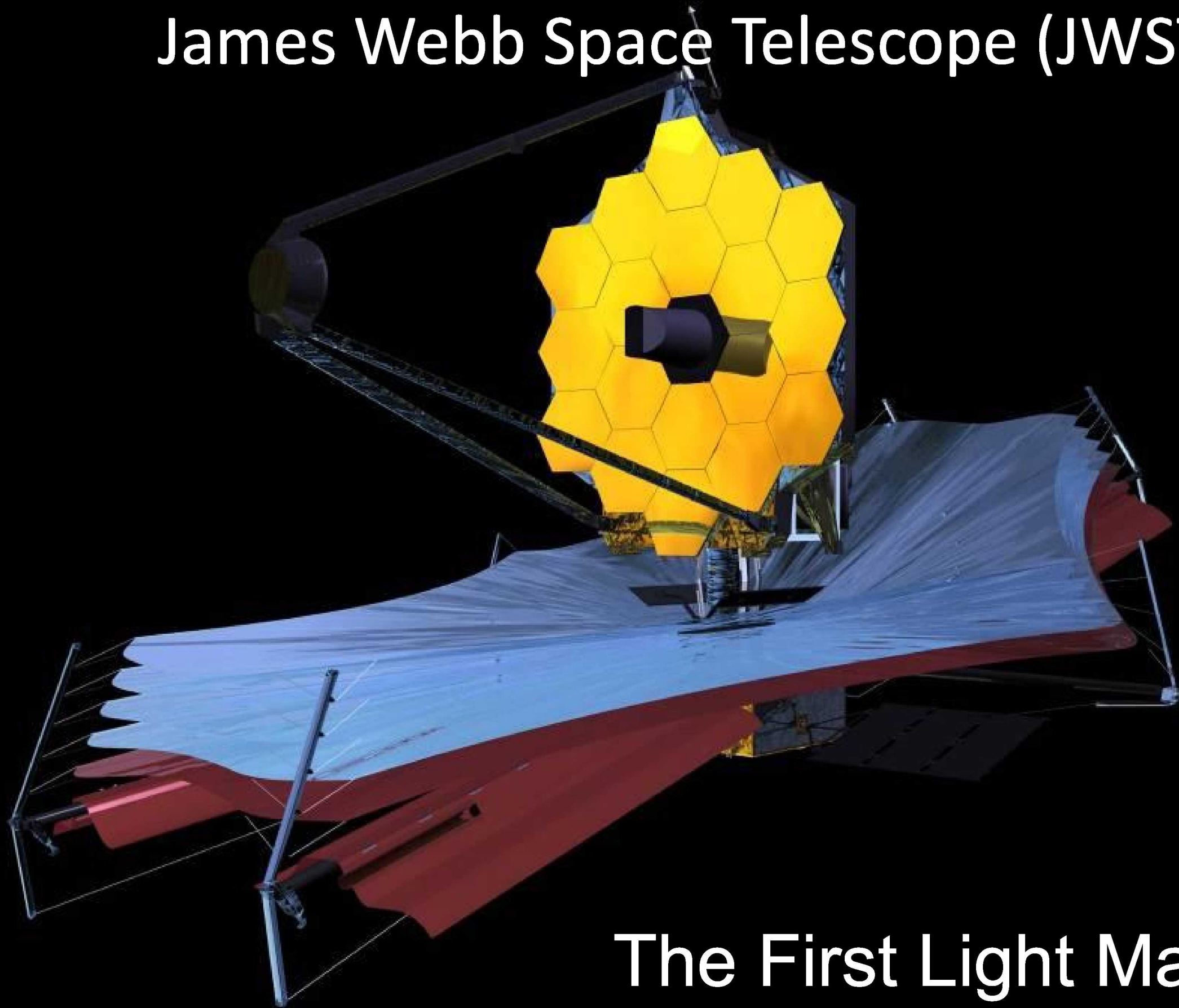


Cell 6

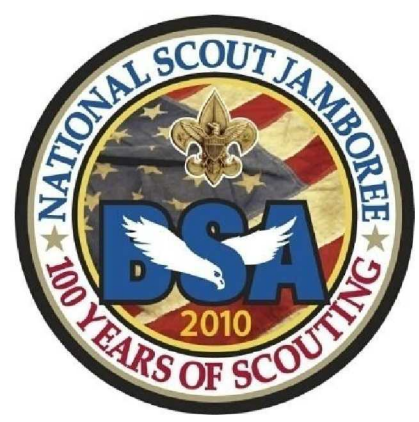
Requirement 4) Visit with an engineer:

- a. Discuss the engineer's work and the tools used.
- b. Discuss a current project and the engineer's role.
- c. Find out how the engineers work is done and how results are achieved.
- d. Look at written reports.

James Webb Space Telescope (JWST)



The First Light Machine



Engineering the Future



How to Engineer a Space Telescope

Use a Systems Engineering Methodology

Organize the Project

Engineering Specifications flow from Science Requirements

Plan Activities and Schedule

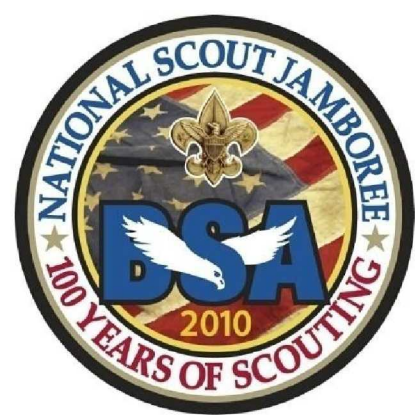
Preliminary Design & Analysis

Technology Development

Competition of Designs and Down Select

Engineering Development Units

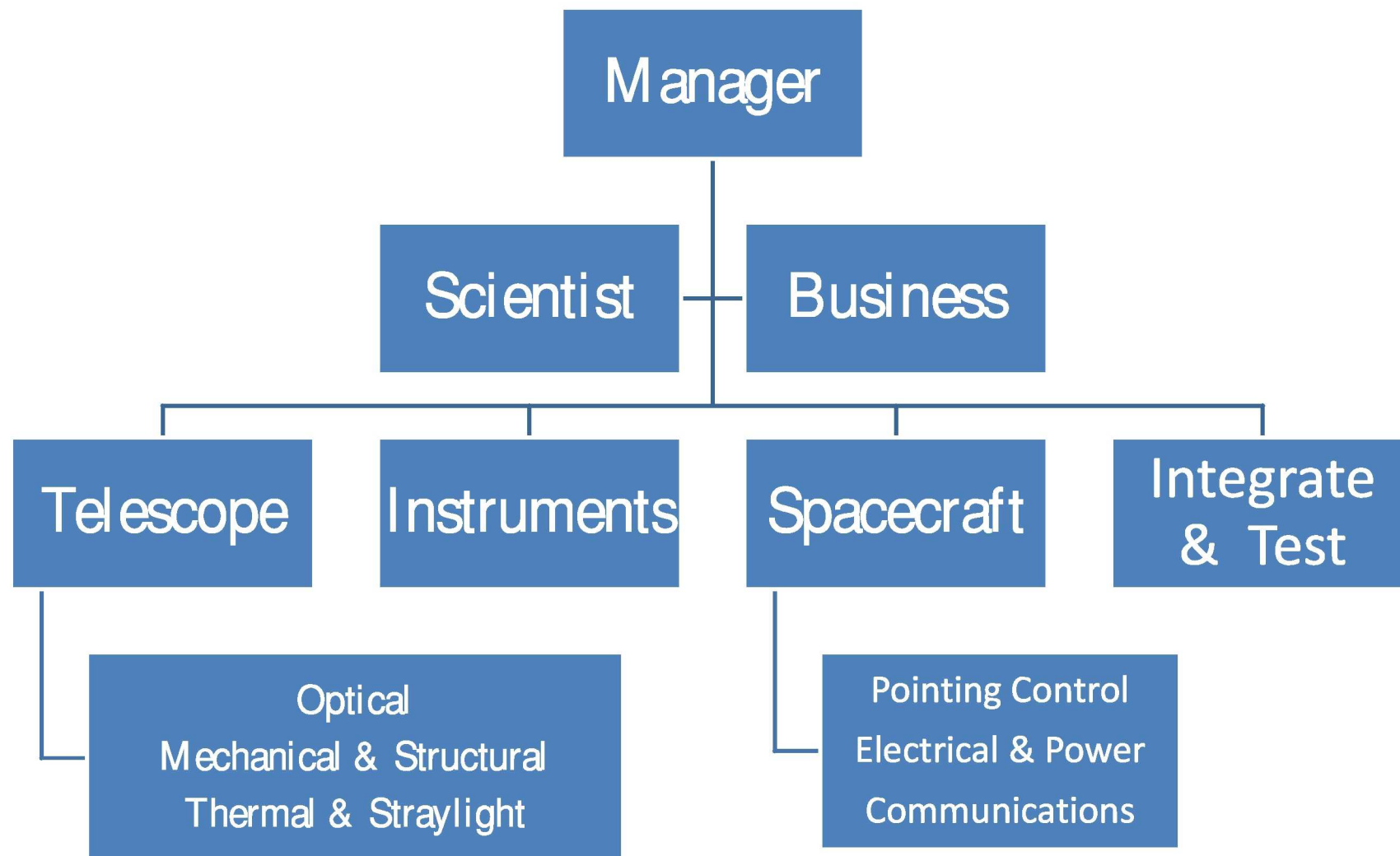
Verification and Validation by Test and Analysis



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Organize the Project

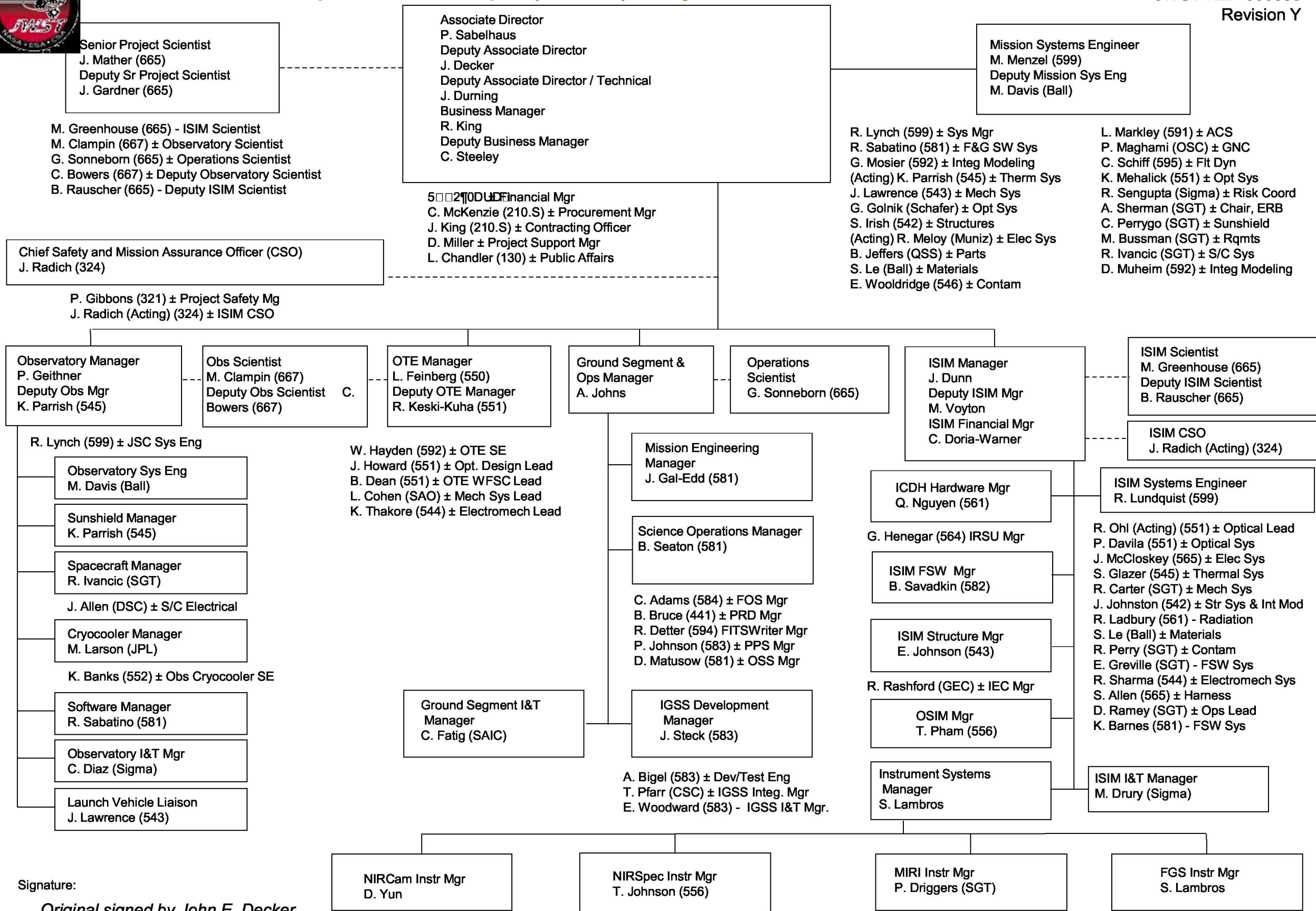




James Webb Space Telescope (JWST) Project ± Code 443

JWST-REF-000838

Revision Y



Signature:

Original signed by John E. Decker

3/20/2009



Engineering the Future



Specifications flow from Requirements

Science Requirements

Science Goals

Sensitivity

Exposure Times

Engineering Specifications

Operating Wavelength

Aperture Diameter

Pointing Stability

Operating Wavelength drives:

Mirror Quality

Optical Coatings

Operating Temperature

Mechanical Stability

Straylight

Orbit & Launch Vehicle drives:

Packaging

6.5 m telescope in 4.5 m fairing

Total Mass

6600 kg to L2

Structural Launch Survival

Vibe & acoustic

Hubble Science Goals

- a) Definitively establish the cosmic distance scale (i.e., measure the Hubble constant to 10%)
- b) Study galaxy evolution out to $z = 1$ (this was considered very high redshift back then)
- c) Study the intergalactic distribution of gas from quasar absorption lines (required UV coverage)

Required a 2.4 meter diameter, UV/Visible (0.5 micrometer diffraction limited) observatory with milli-arc-second pointing stability and wide field of view.



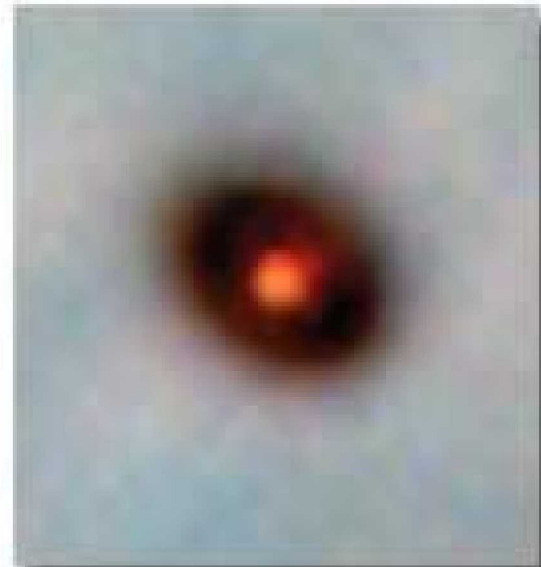
JWST Science Goals



End of the dark ages: First light and reionization



The assembly of galaxies



Birth of stars and proto-planetary systems



Planetary systems and the origin of life

Requires a 6.5 meter diameter, near-infrared (2.0 micrometer diffraction limited) observatory with milli-arc-second pointing stability and wide field of view.

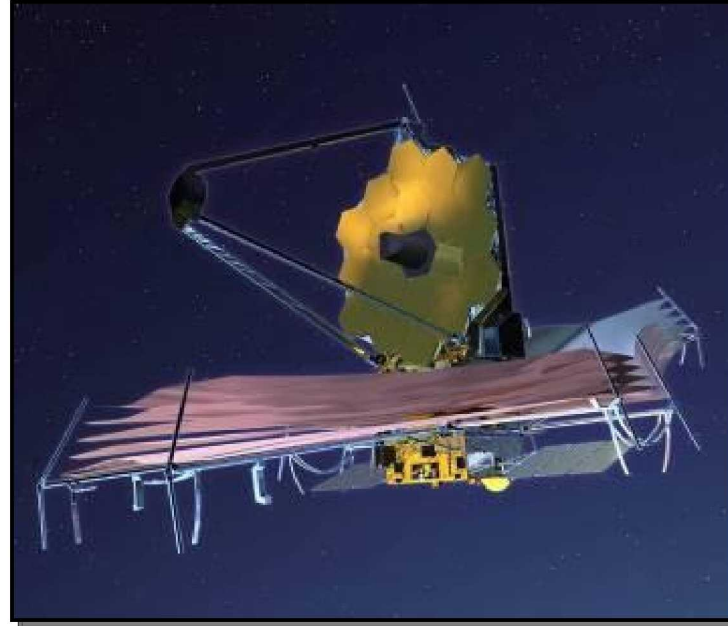
JWST Expands on HST & Spitzer Capabilities

HST



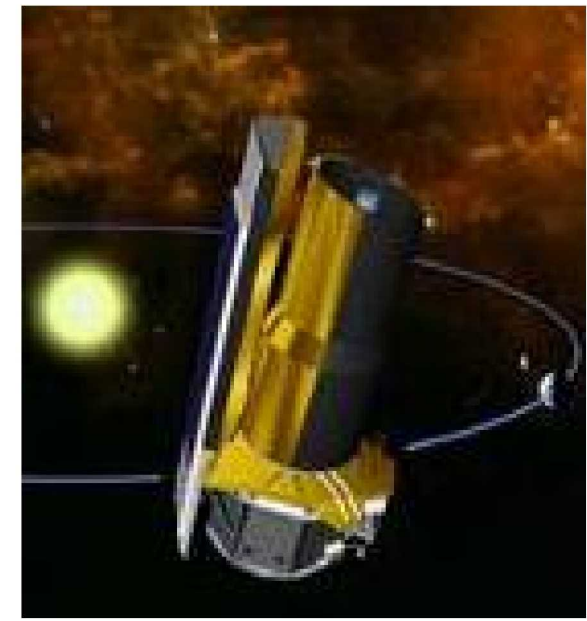
2.4 m Primary
0.5 micrometer
Room Temperature

JWST



6.5 m Primary
2 micrometer
< 50 K (-223C, -370F)

Spitzer



0.85 m Primary
5 micrometer
4K

JWST has

6x the light gathering capability of the Hubble Space Telescope

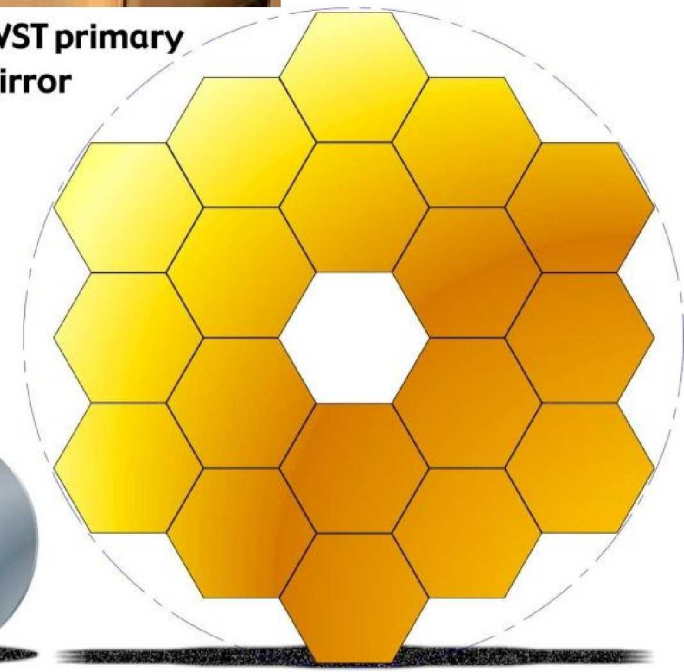
44x the light gathering capability of the Spitzer Space Telescope

JWST has same angular resolution in the near-IR as HST in visible

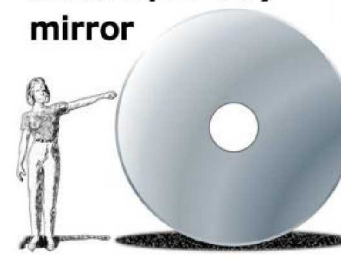
How big is JWST?



JWST primary mirror



Hubble primary mirror



Full Scale JWST Mockup

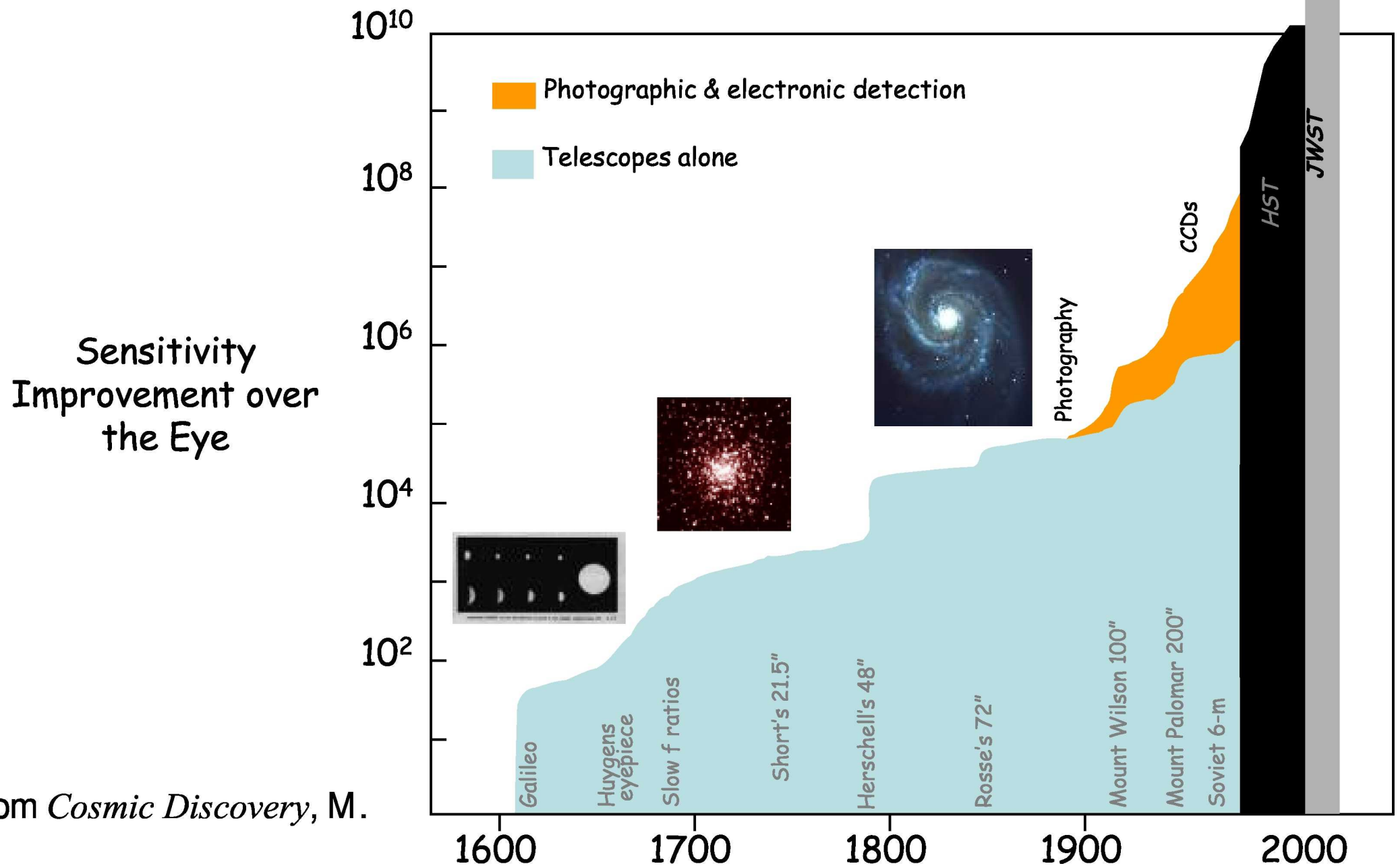


21st National Space Symposium, Colorado Springs, The Space Foundation

How to win at Astronomy

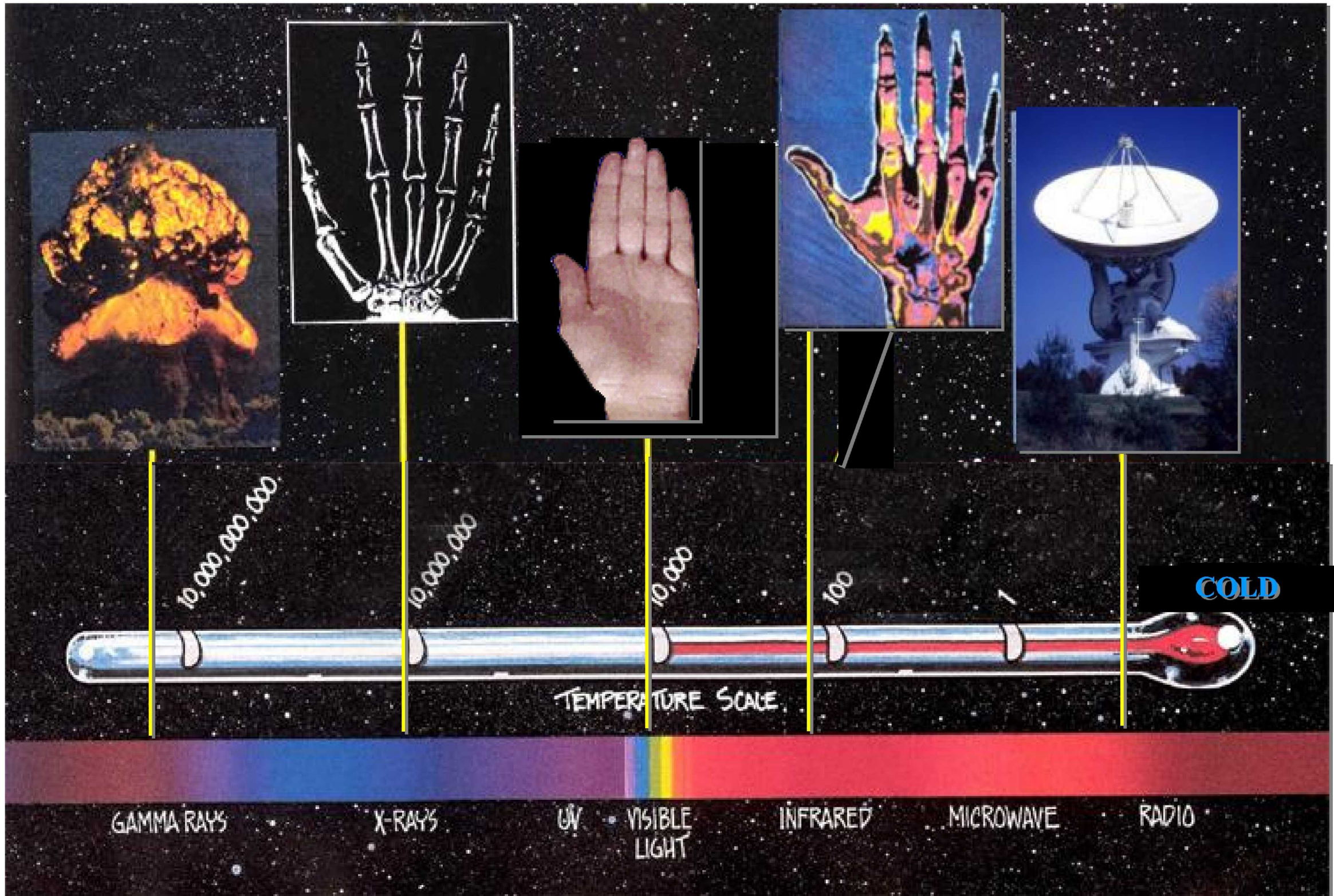
Aperture = Sensitivity

Big Telescopes with Sensitive Detectors In Space



Adapted from *Cosmic Discovery*, M. Harwit

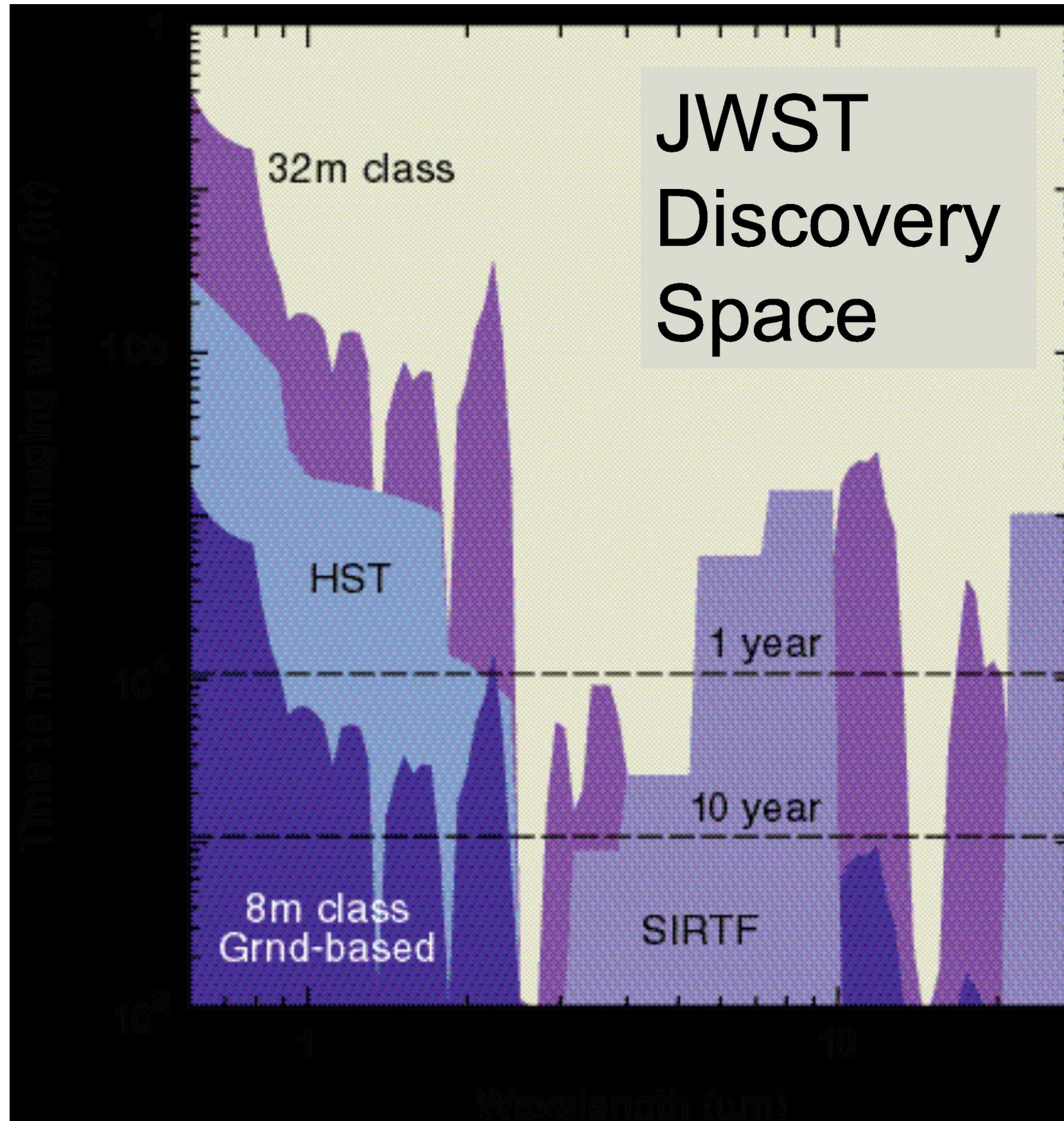
Infrared Light

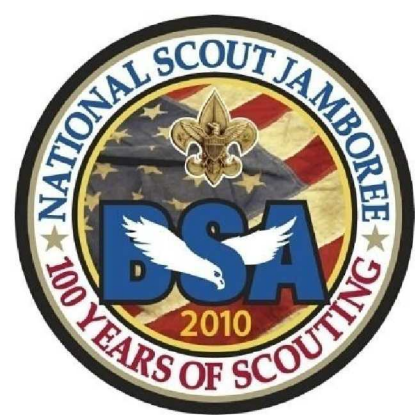


Why Infrared ?



Why go to Space ± Wavelength Coverage





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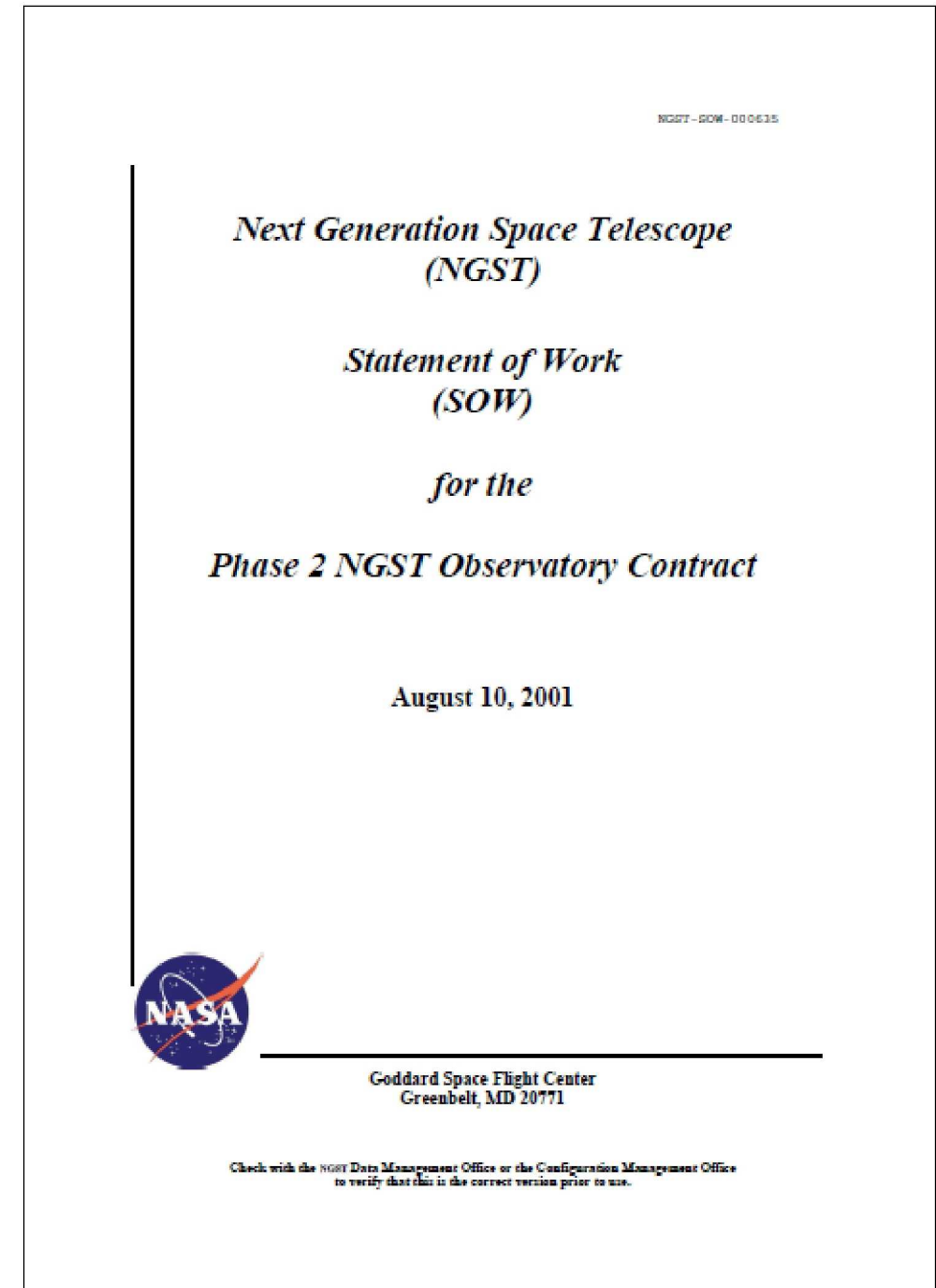


Planning

Statement of Work (SOW)

Work Breakdown Schedule (WBS)

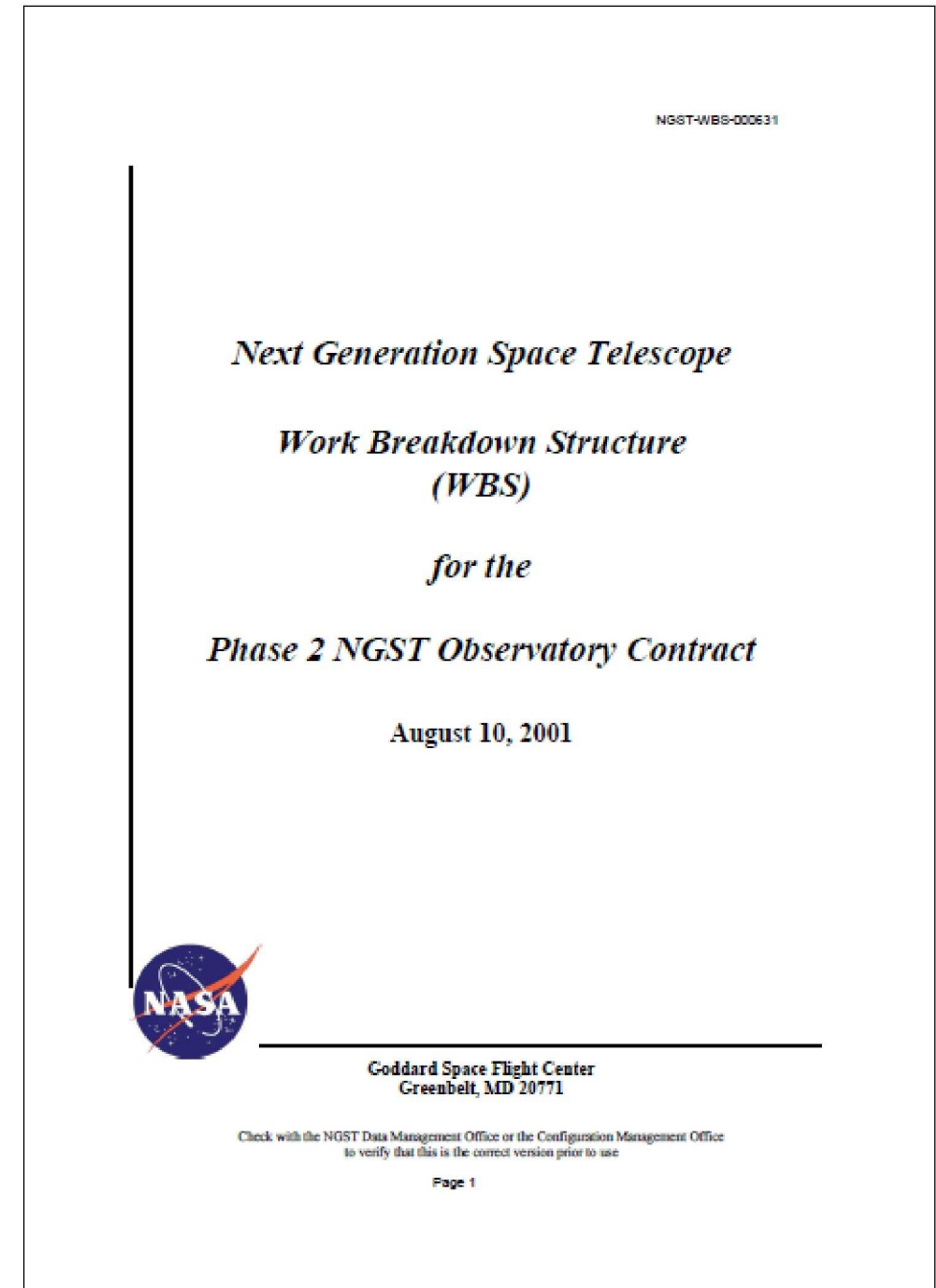
Schedule (Gant Chart)





JWST Optical Telescope Element WBS

1. Project Management
2. Reserved
3. Observatory Systems Engineering
 - 3.1 Systems Engineering Management
 - 3.2 Requirement Analysis and Verification
 - 3.3 Interface Management and Configuration Control
 - 3.4 Trade Studies and Life Cycle Cost (LCC) Analysis
 - 3.5 Technical Risk Management
 - 3.6 Technology Development and Validation
 - 3.7 Integrated Modeling and Analysis
4. Observatory Integration and Test
5. Optical Telescope Element
 - 5.1 OTE Management
 - 5.2 OTE Systems Engineering
 - 5.3 OTE I&T
 - 5.4 OTE Simulators
 - 5.5 OTE Subsystem Design, Manufacture, Assembly and Test
 - 5.5.1 Primary Mirror
 - 5.5.2 Other Optics and Structure
 - 5.6 Wavefront Sensing and Control (WFS&C)
6. Sunshield
7. Spacecraft
8. Integrated Science Instrument Module (ISIM)
9. Flight Software (FSW) Systems Development
10. Reserved
11. Ground Segment and Operations/Science





JWST Optical Telescope Requirements

Requirements

Primary goal is to observe early universe, at an age between 1 million and a few billion years during when the first stars and galaxies began to form. To accomplish this, JWST will be an infrared observatory instrumented for imaging and spectroscopy, diffraction limited at $2\mu\text{m}$, with approximately a 0.1 arc-second resolution, nano-Jansky sensitivity, and a large field of view.

Science Mission Lifetime

Science mission lifetime shall be a minimum of five years, with a total mission lifetime goal of ten years.

Wavelength Range

Spectral coverage shall extend from $0.6\mu\text{m}$ to $>10\mu\text{m}$.

Optical Telescope Element

Primary Mirror Area

The unobscured primary mirror area shall be greater than or equal to 25 square meters.

Wavefront Error Allocation


The OTE shall be allocated 131 nanometers rms of wavefront error.

NGST-RQMT-000634

*Next Generation Space Telescope
(NGST)*

Level 2 Requirements

August 10, 2001



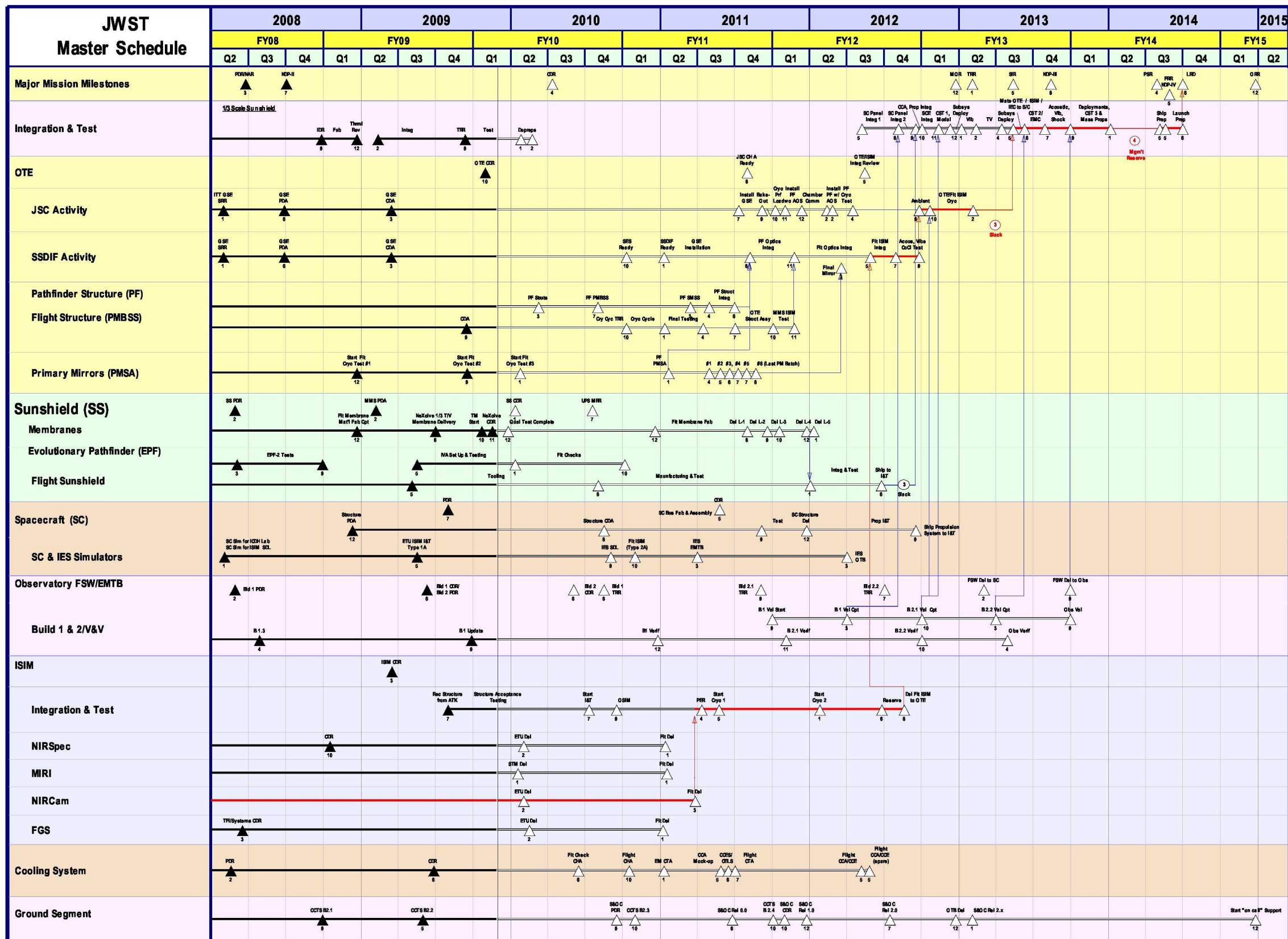
Goddard Space Flight Center
Greenbelt, MD 20771

Check with the NGST Data Management Office or the Configuration Management Office to verify that this is the correct version prior to use.

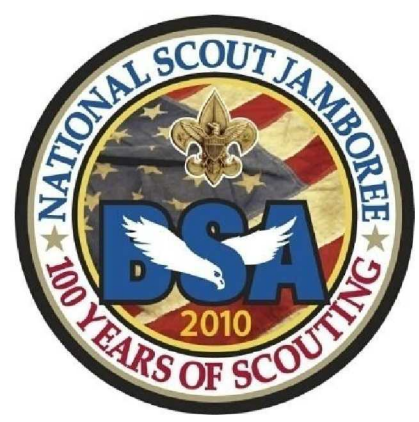


JWST Master Schedule

11/30/09



Rev-I DRAFT



Engineering the Future



Optical Engineering Design Tools

Specialized Software Tools

Optical Design

Computer Aided Design (CAD)

Dynamic Response

Thermal Design

Straylight

Generic Software Tools

Word Processing

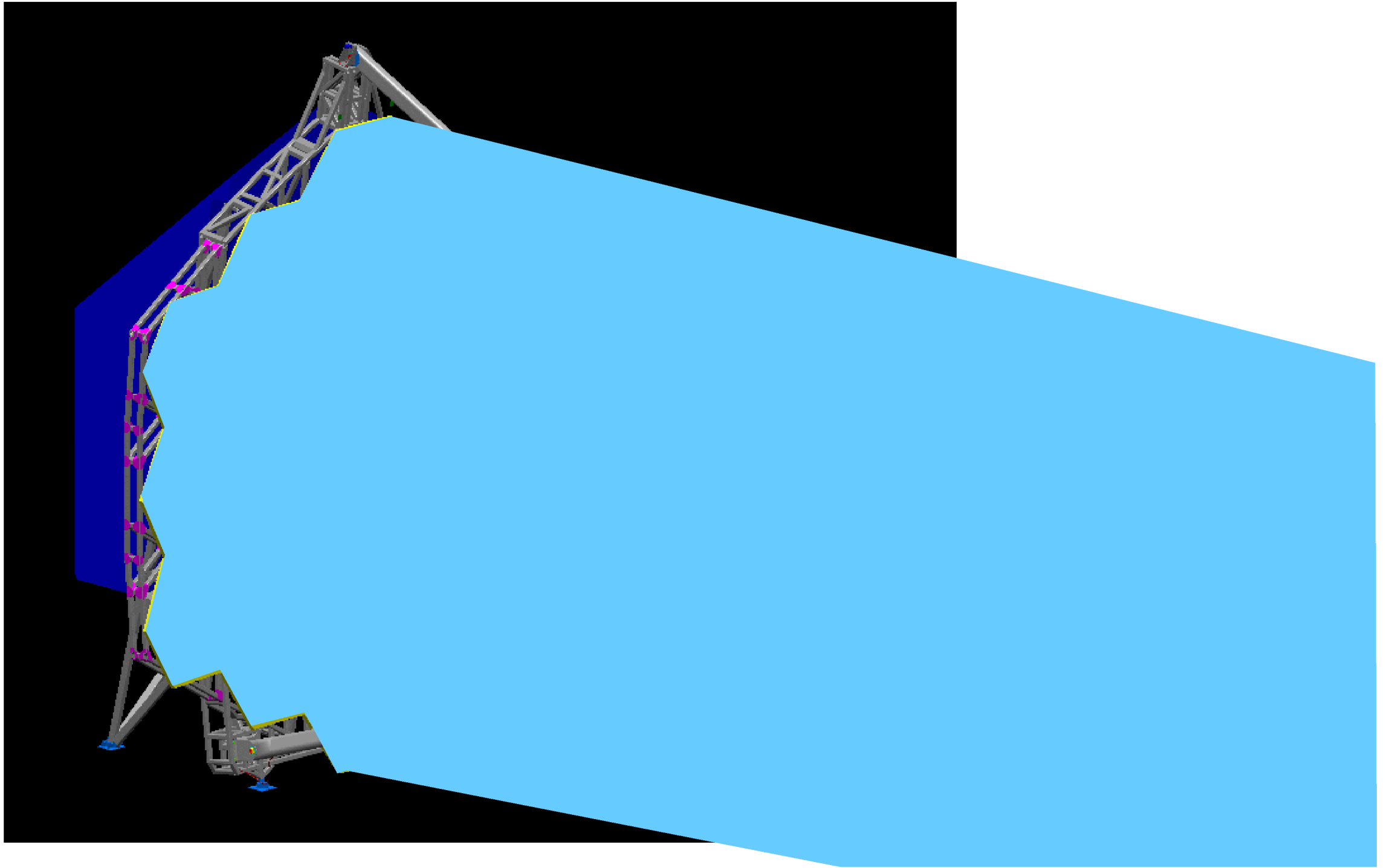
Spread Sheets

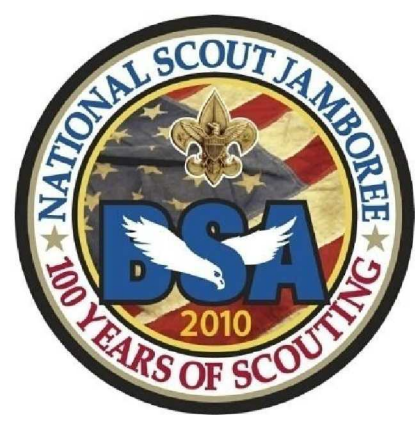
Power Point

Scheduling

Engineering Development Units

JWST Optical Path





Engineering the Future

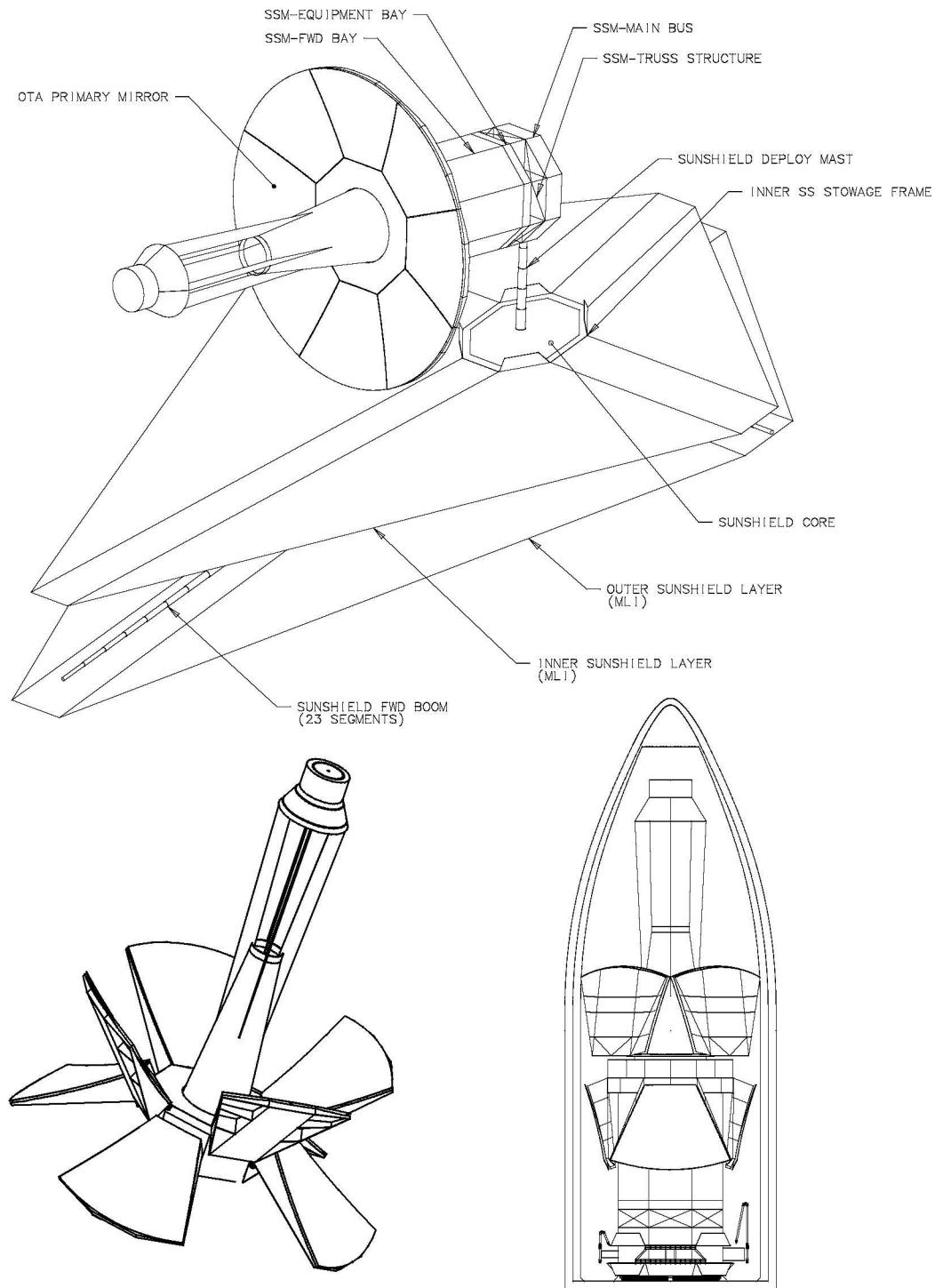
Preliminary Design & Analysis



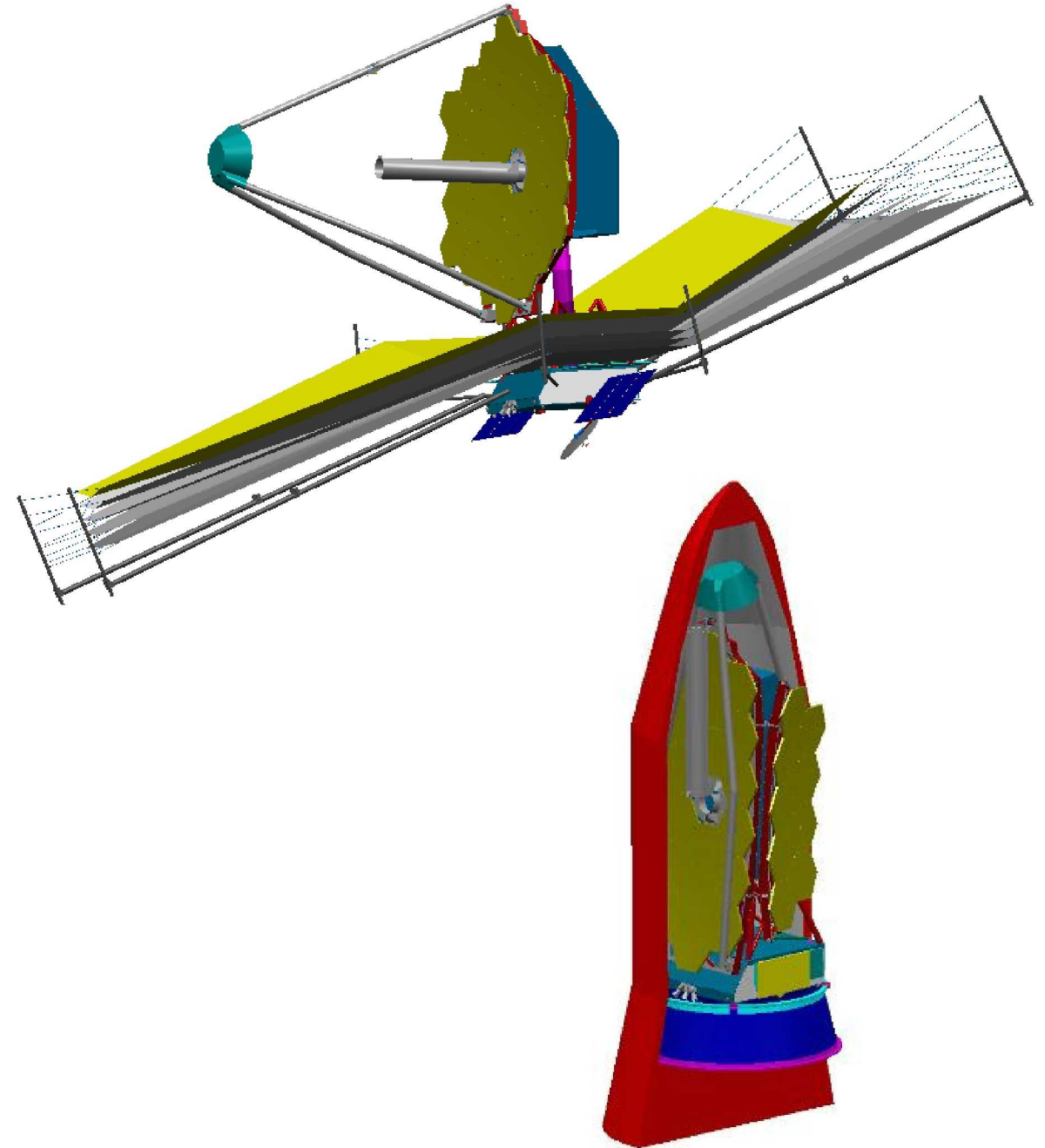
- Space Telescope Missions Take a Long Time to Plan
 - ± Science Planning for JWST began in 1989
 - ± Preliminary Design began in 1996
 - ± Launch is scheduled for 2014
 - 25 years after initial idea
 - 18 years after preliminary design began.
- Because of Launch Vehicle Fairing Size Constraints, JWST must be segmented.
 - ± JWST Collecting Aperture is 6.5 meters diameter
 - ± Largest Launch Vehicle Fairing Size is 4.5 meters diameter
- Preliminary Design Contracts resulted in two design concepts
 - ± These two concepts competed against each other until in 2003 when NASA selected the TRW/Ball concept

Two Competing Design Concepts

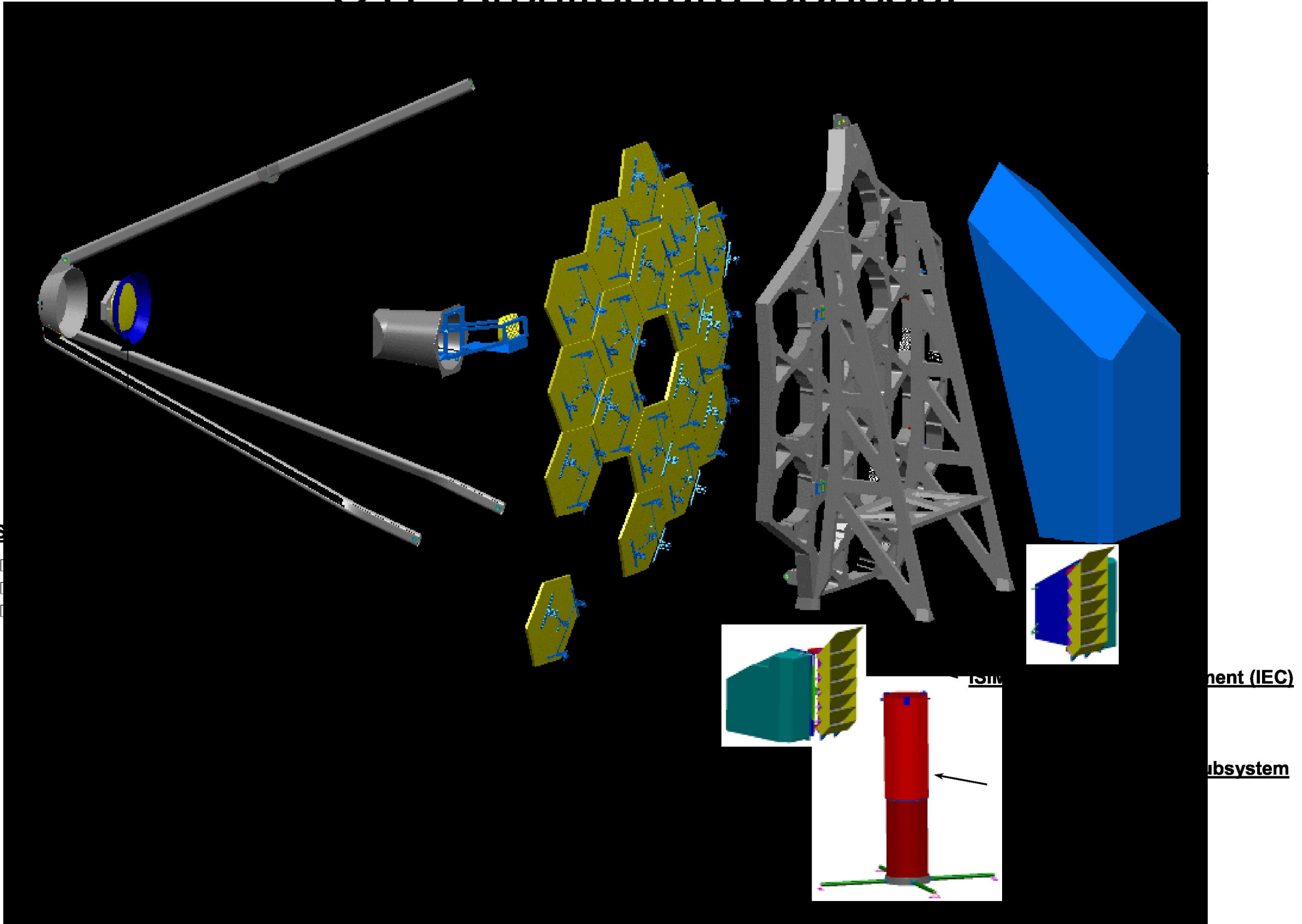
Lockheed/Raytheon

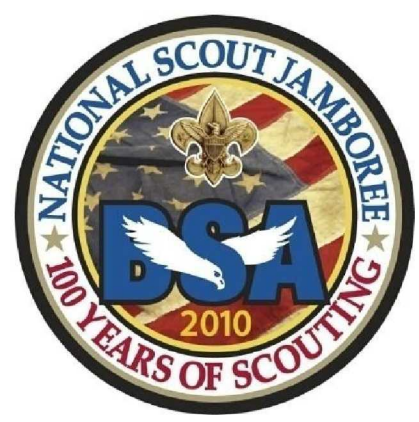


TRW/Ball



OTF Architecture Concept





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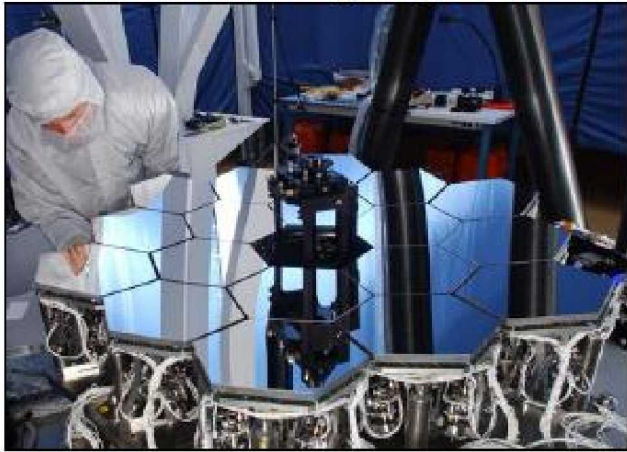


Technology Development

- In 1996, the technology did not exist to build JWST.
- NASA invested over \$300M in technology development.
- Starting in 1999, it was my job to lead the Mirror Technology Development effort.
- Specific mirror technology that had to be developed:
 - ± New kind of Beryllium alloy
 - ± New ways to manufacture mirrors
 - ± New ways to test mirrors
- Same as for the observatory concepts, there was a competition between mirror technologies and eventual selection.

JWST Technology Demonstrations for T-NAR

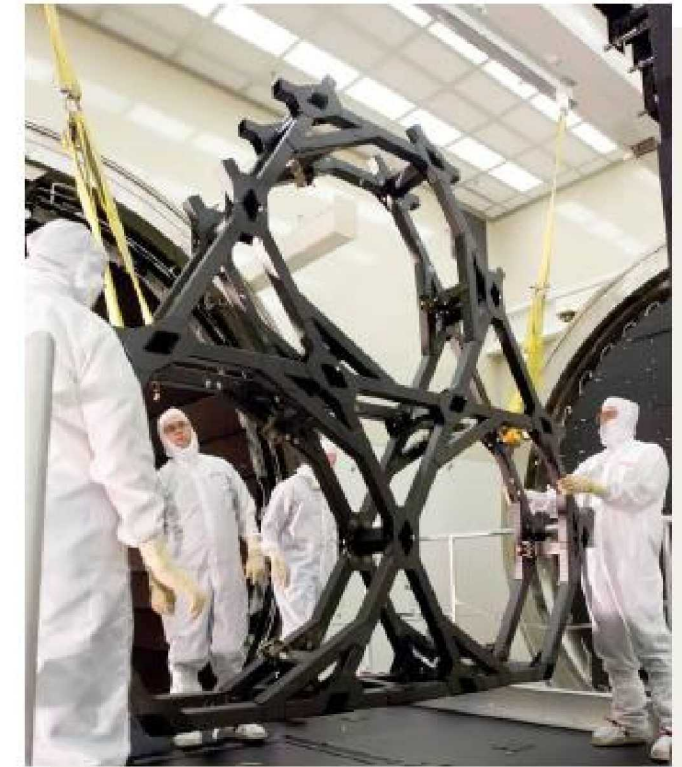
Mirror Phasing Algorithms



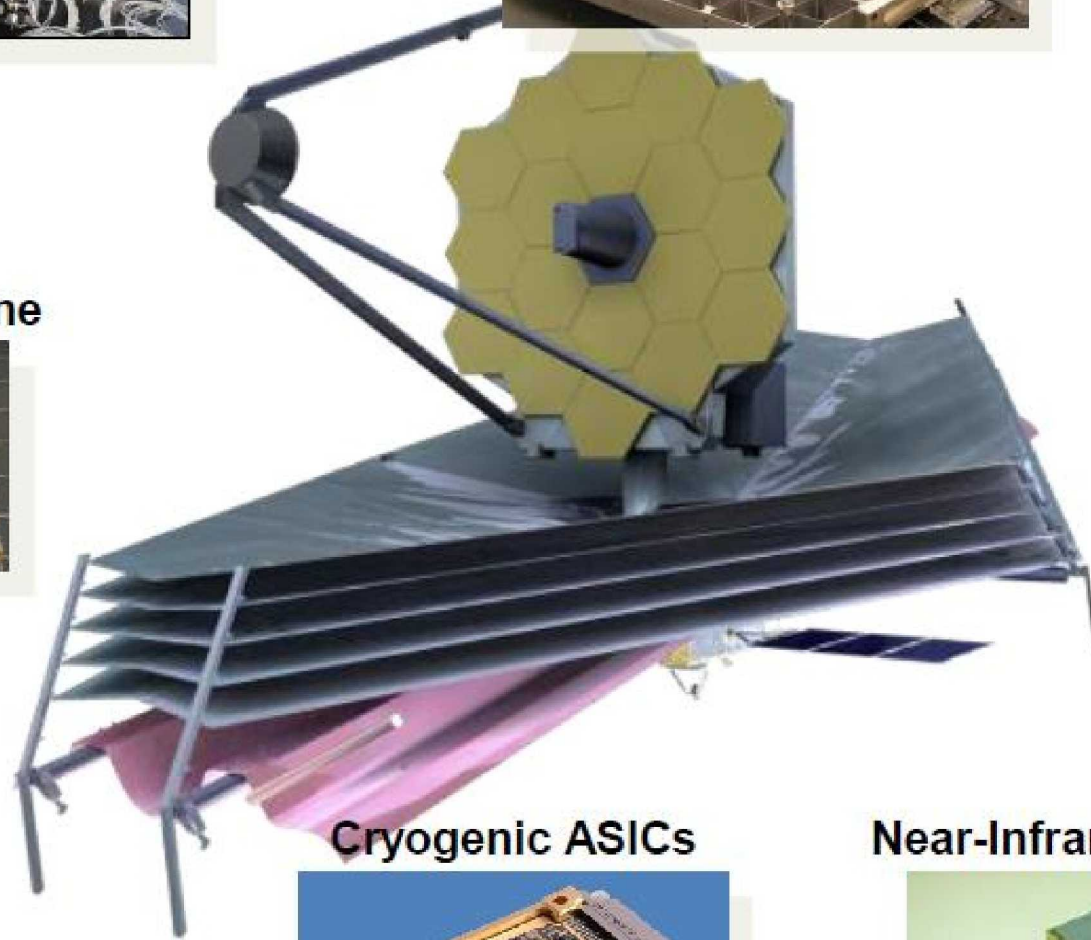
Beryllium Primary Mirror Segment



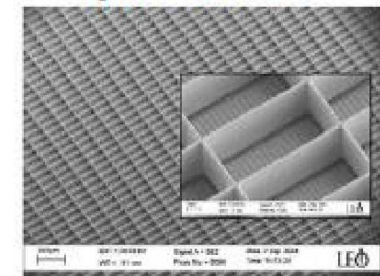
Backplane



Sunshield Membrane



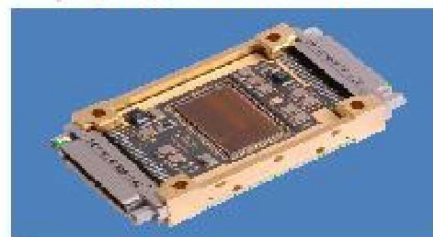
μ Shutters



Cryocooler



Cryogenic ASICs



Near-Infrared Detector



Mid-Infrared Detector



Investments Have Reduced Risk

Mirror Actuators



Mirrors

AMSD



SBMD



Mirror System



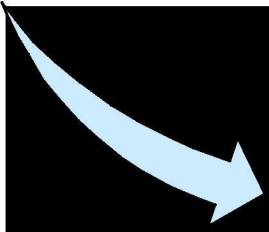
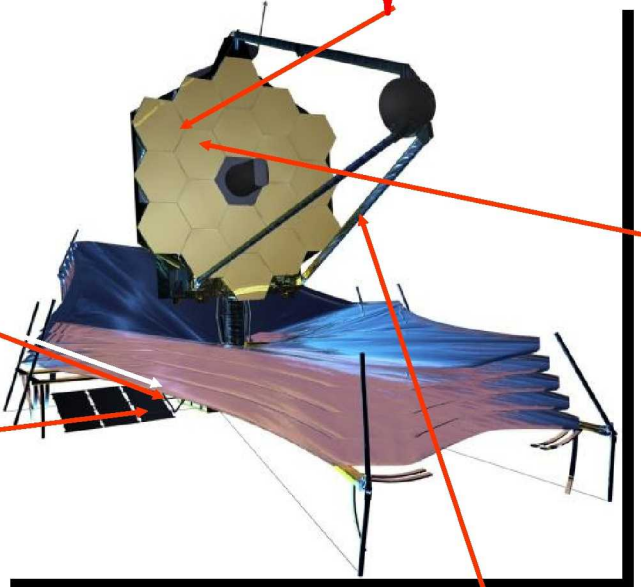
Wavefront Sensing and Control, Mirror Phasing



1 Hz OTE Isolators

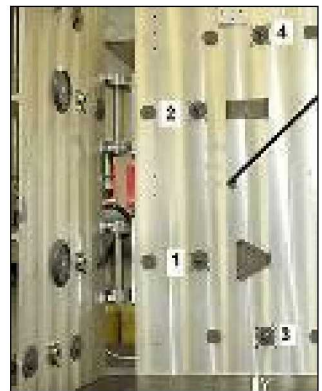


Reaction Wheel Isolators



Primary Deployable Optical Telescope Assembly (DOTA)

Primary Mirror Structure Hinges and Latches



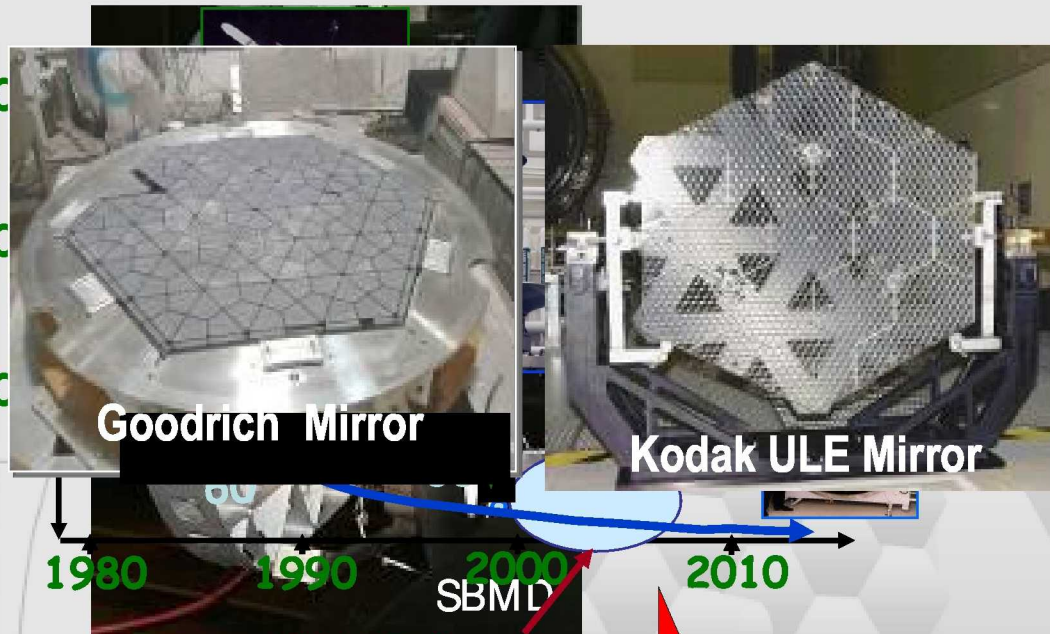
Half-Scale Sunshield Model



Secondary Mirror Structure Hinges

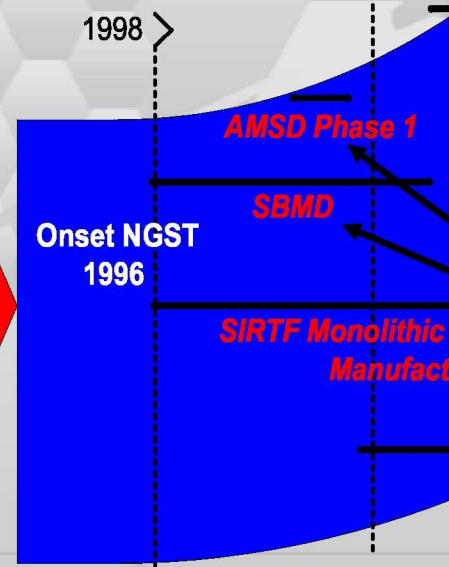
JWST Mirror Technology History

Areal Density (Kg/m²)



JWST Requirement

*** NASA HST, Chandra, SIRTf Lessons Learned**
 - TRL 6 by NAR
 - Implement an active risk management process early in the program (Early investment)

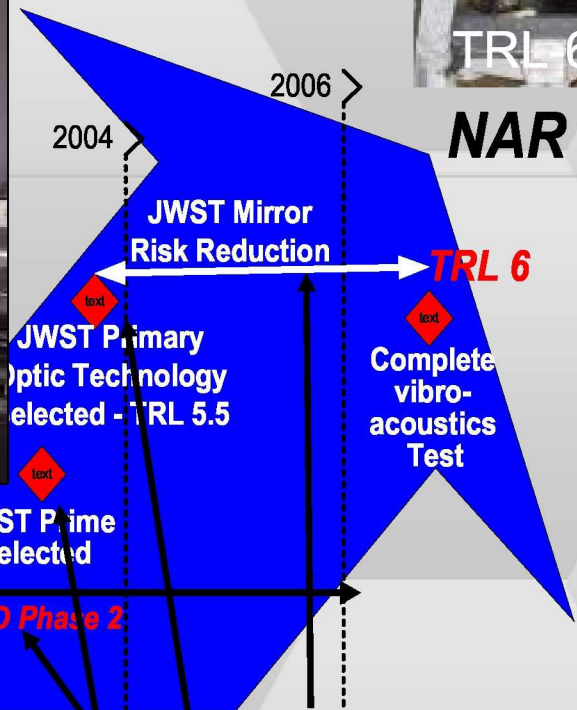


Mirror Material/Technology Selection, September, 2003

- Beryllium chosen for technical reasons (cryogenic CTE, thermal conductance, issues with glass stress issues with Be noted)

Prime Contractor Selection

- Ball (Beryllium) and ITT/Kodak (ULE) proposed as options, Goodrich dropped from AMSD (meter demonstrations)



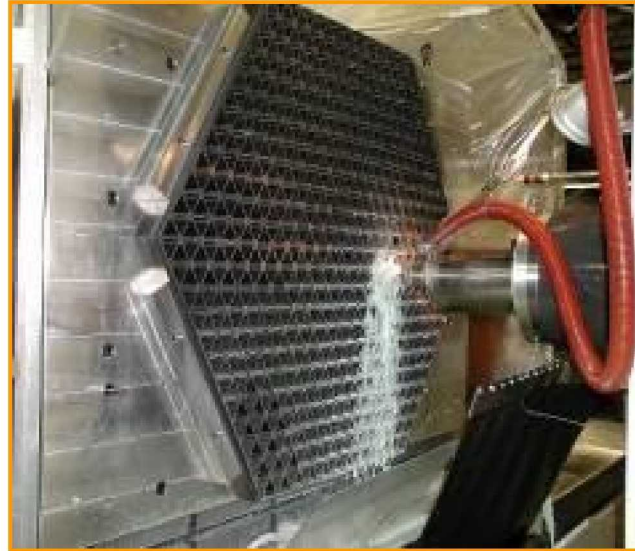
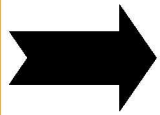
Based on lessons learned, JWST invested early in mirror technology to address lower areal densities and cryogenic operations

Mirror Manufacturing Process

Blank Fabrication

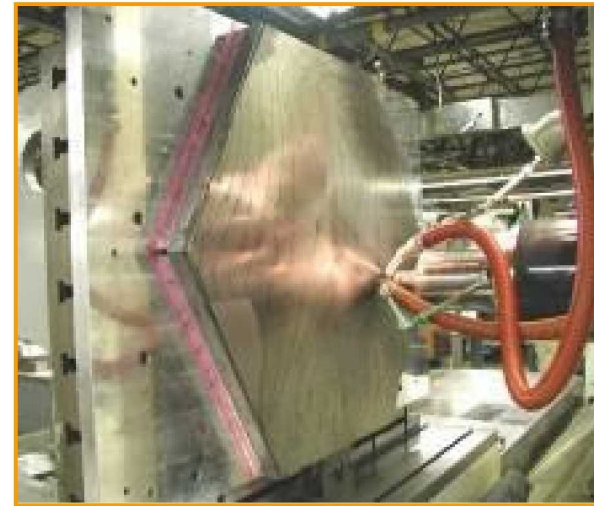


HIP Vessel being loading into chamber



Machining of Web Structure

Machining



Machining of Optical Surface



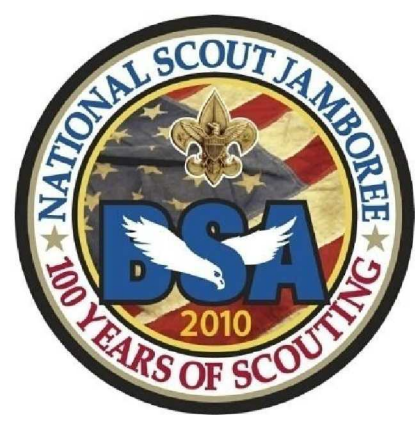
Completed Mirror Blank

Polishing



Mirror System Integration





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Engineering Development Units (EDUs)

EDUs are extremely important.

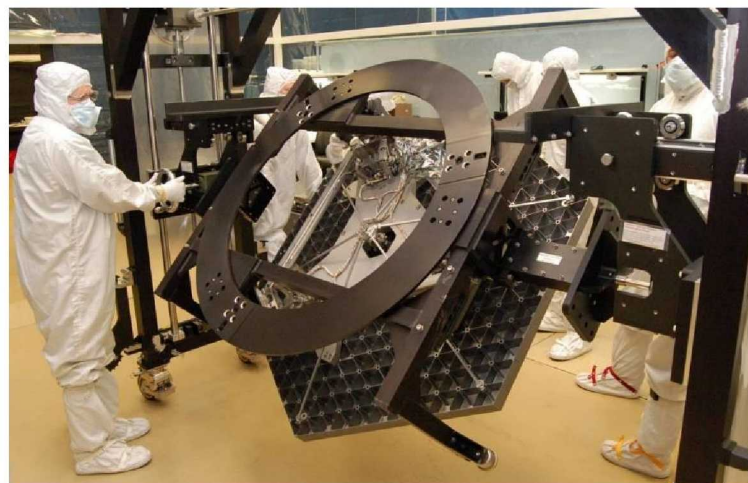
Flight Hardware is VERY EXPENSIVE

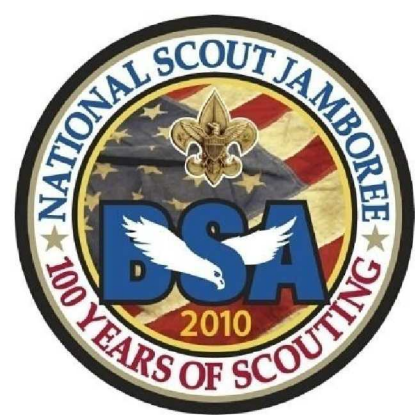
EDUs allow us to practice how to manufacture, handle, test, and assemble the flight hardware.

EDUs allow us to make mistakes without the risk of damaging the flight hardware.



JWST EDUs and Simulators





Engineering the Future



Verification & Validation by Test & Analysis

Testing is performed at every level of the program:

Component

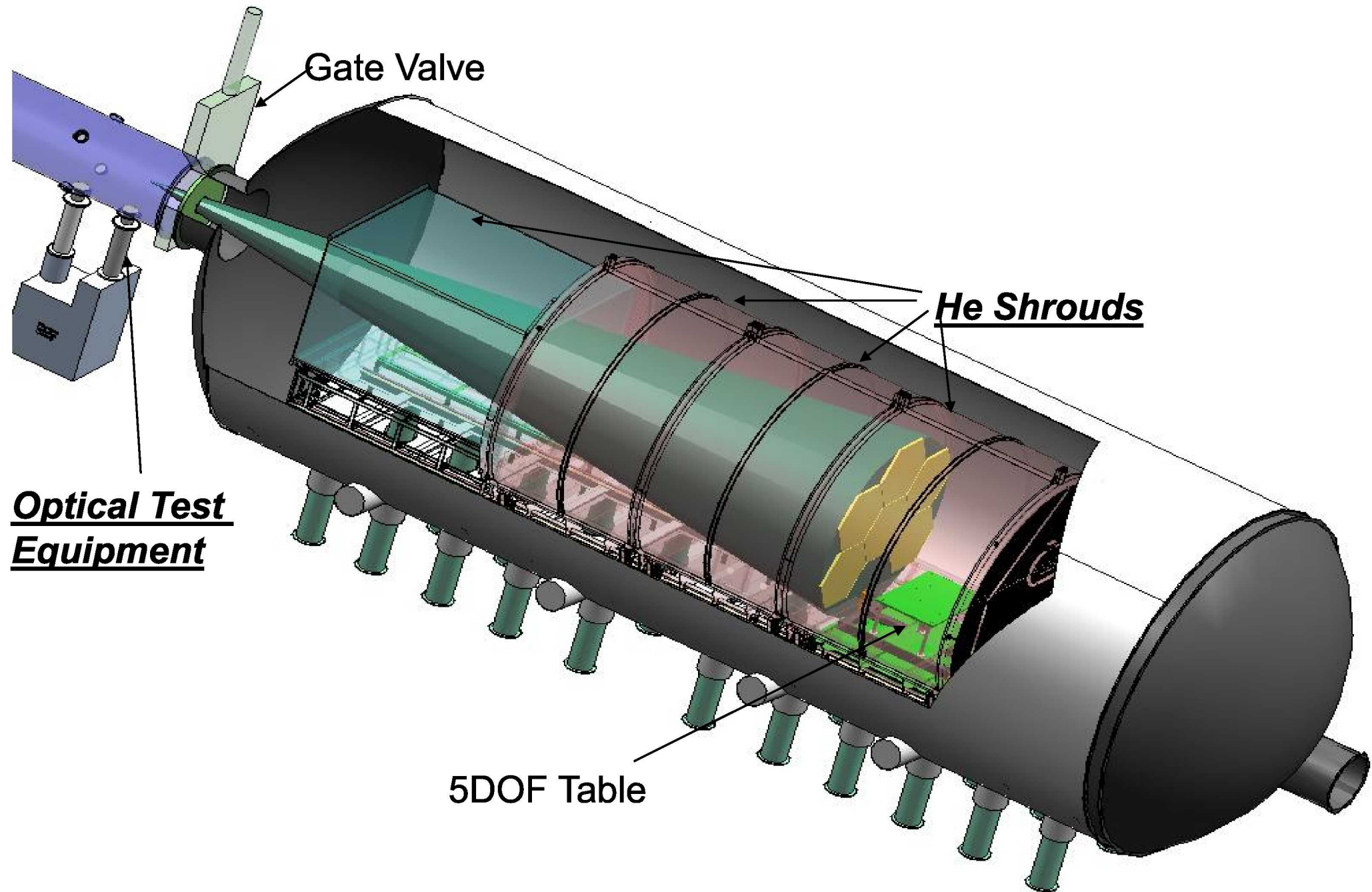
Sub-System/Sub-Assembly

Assembly

System

Observatory

MSFC Cryogenic Test Facility



MSFC Cryogenic Test Stand

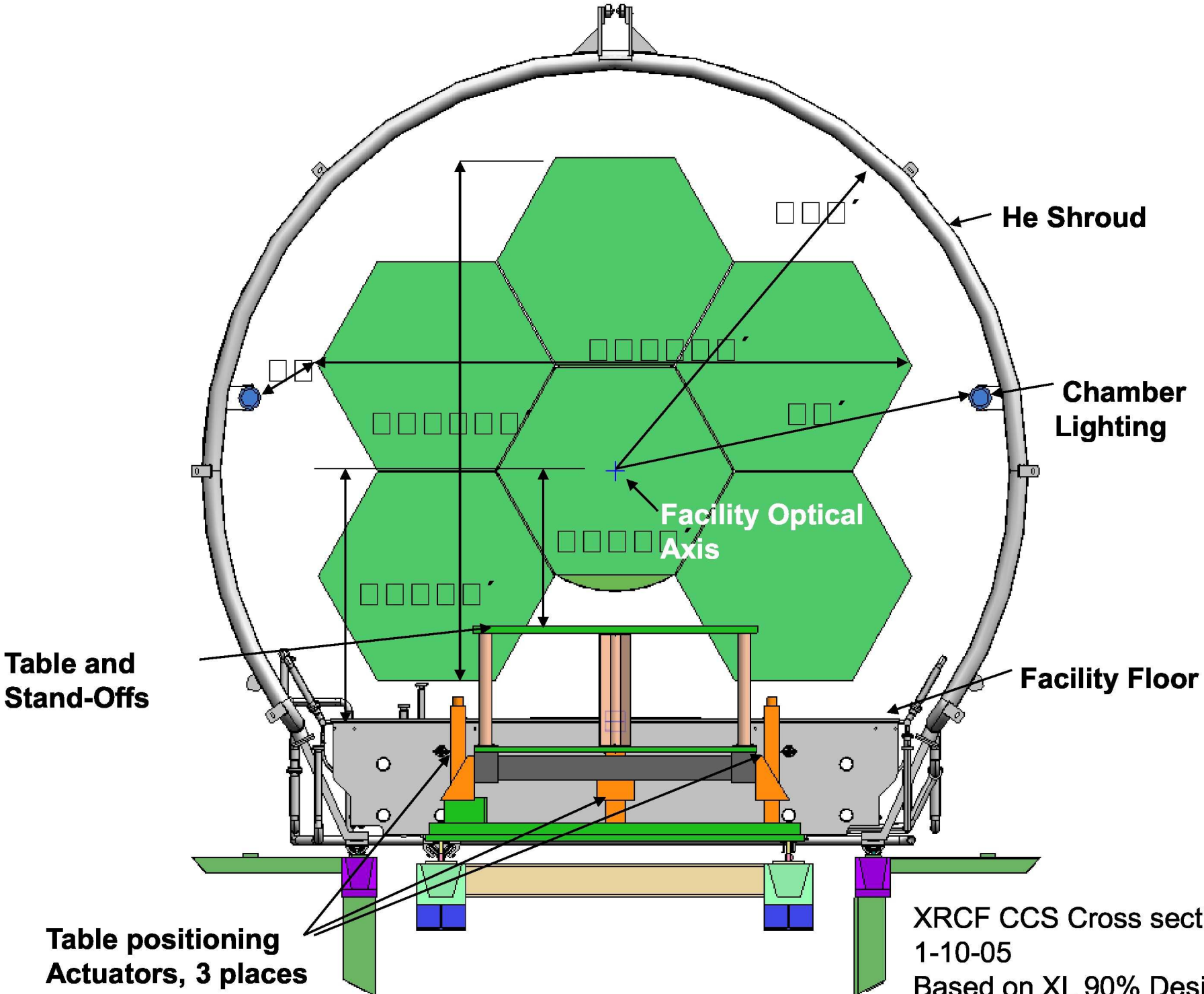
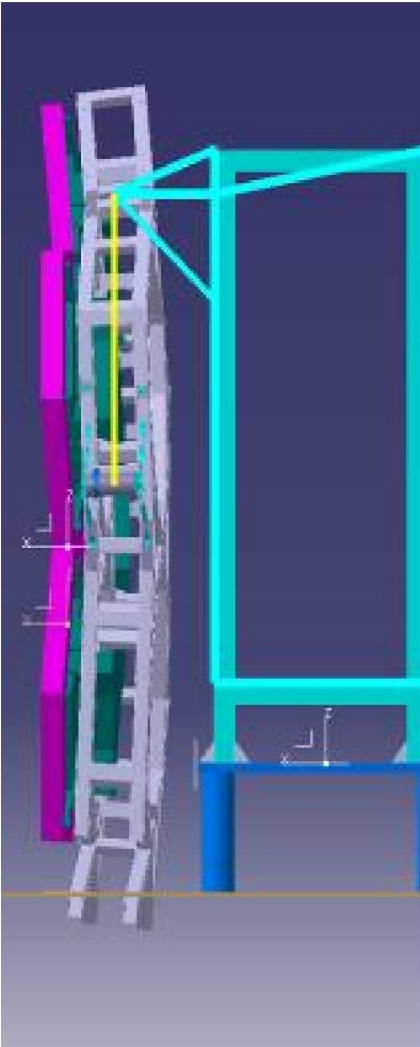
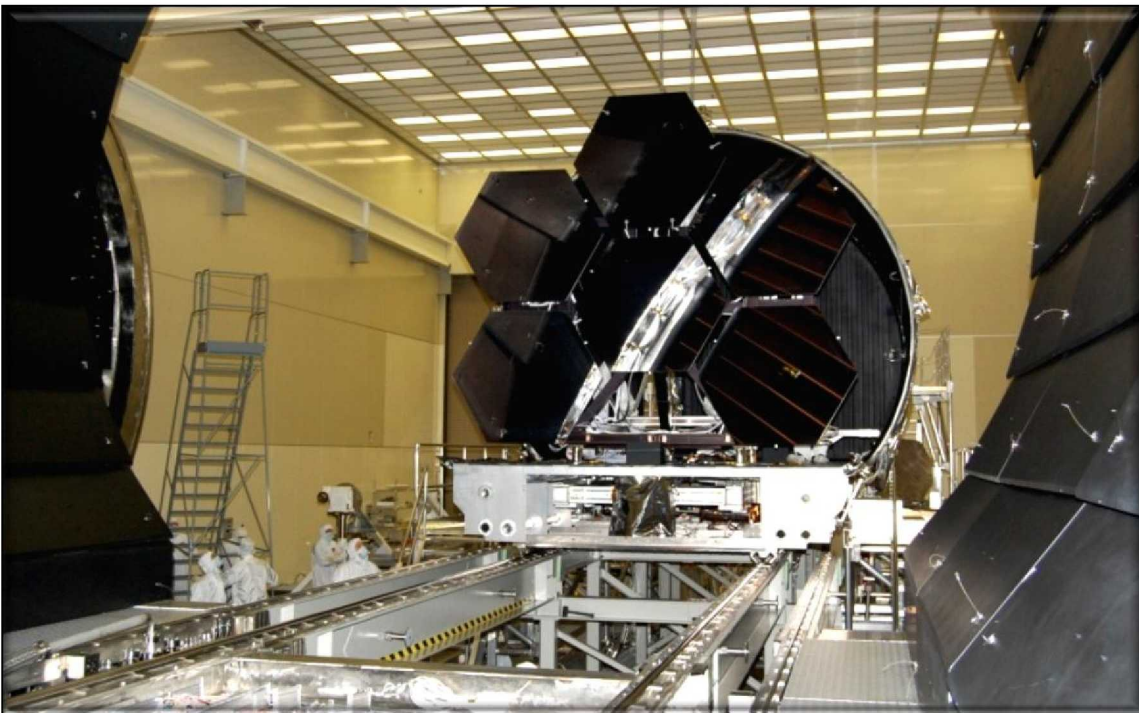
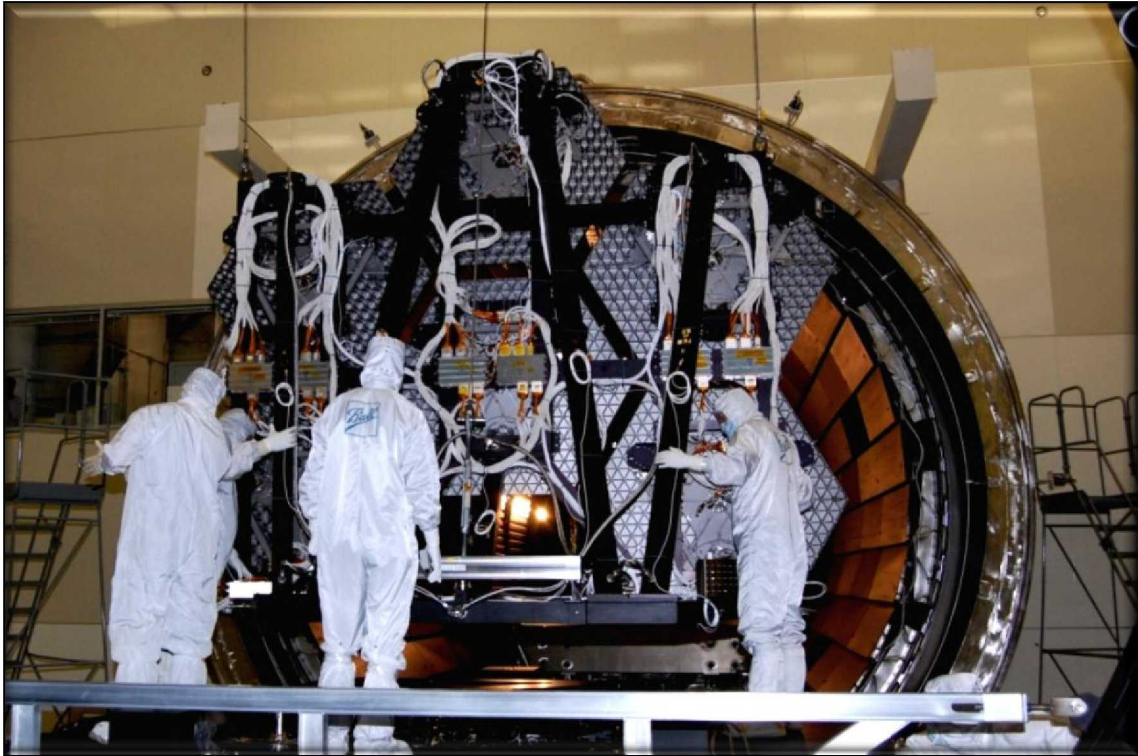


Table positioning Actuators, 3 places

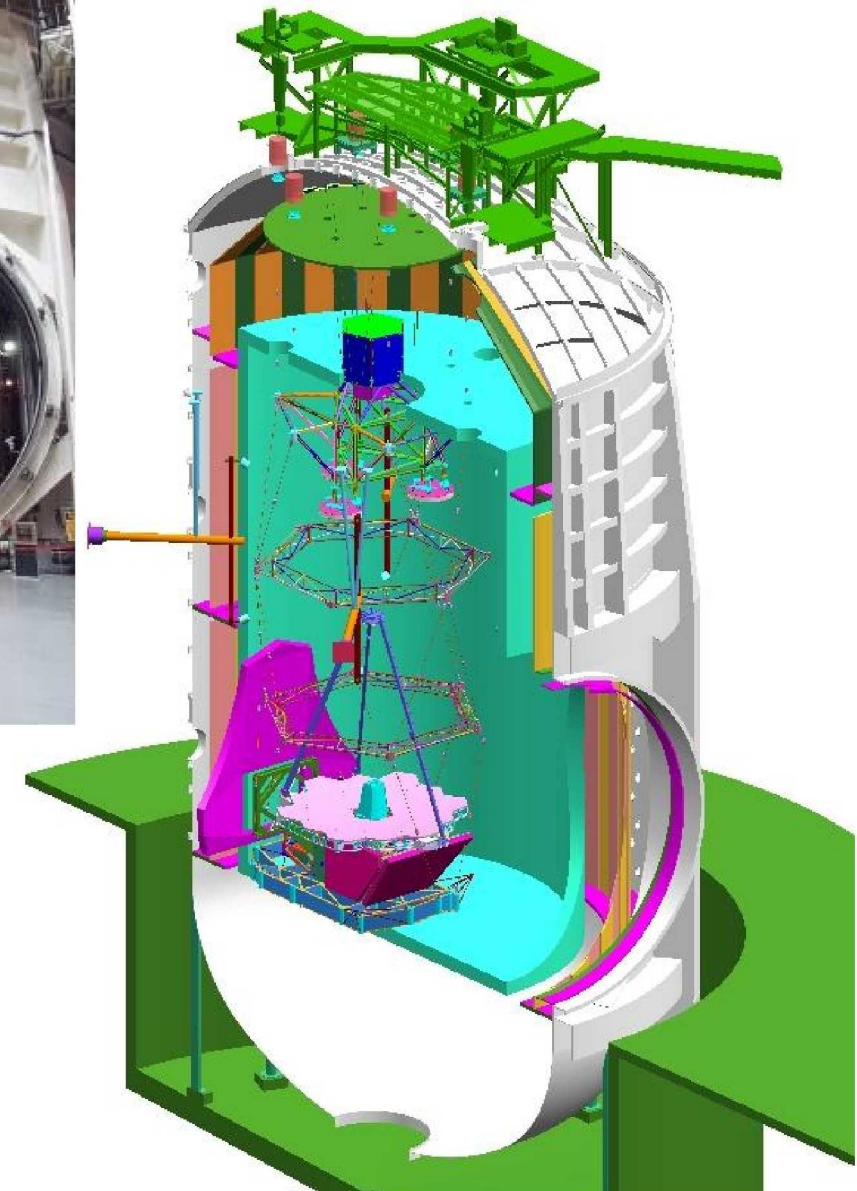
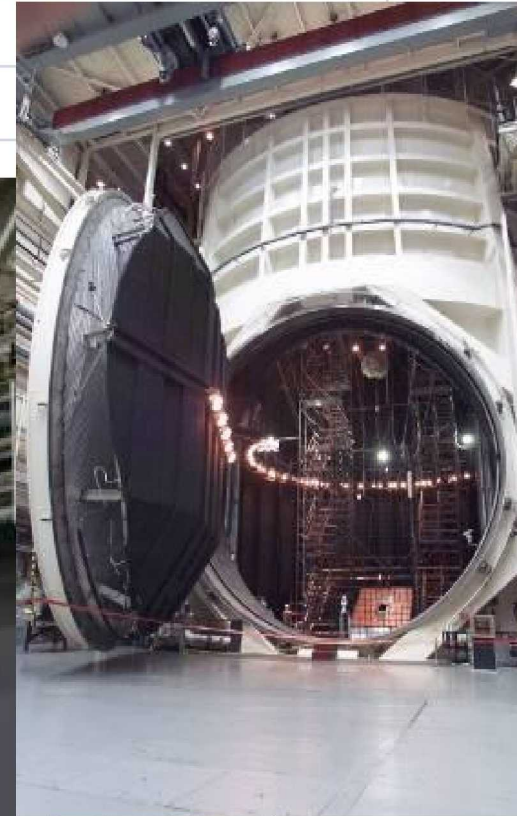
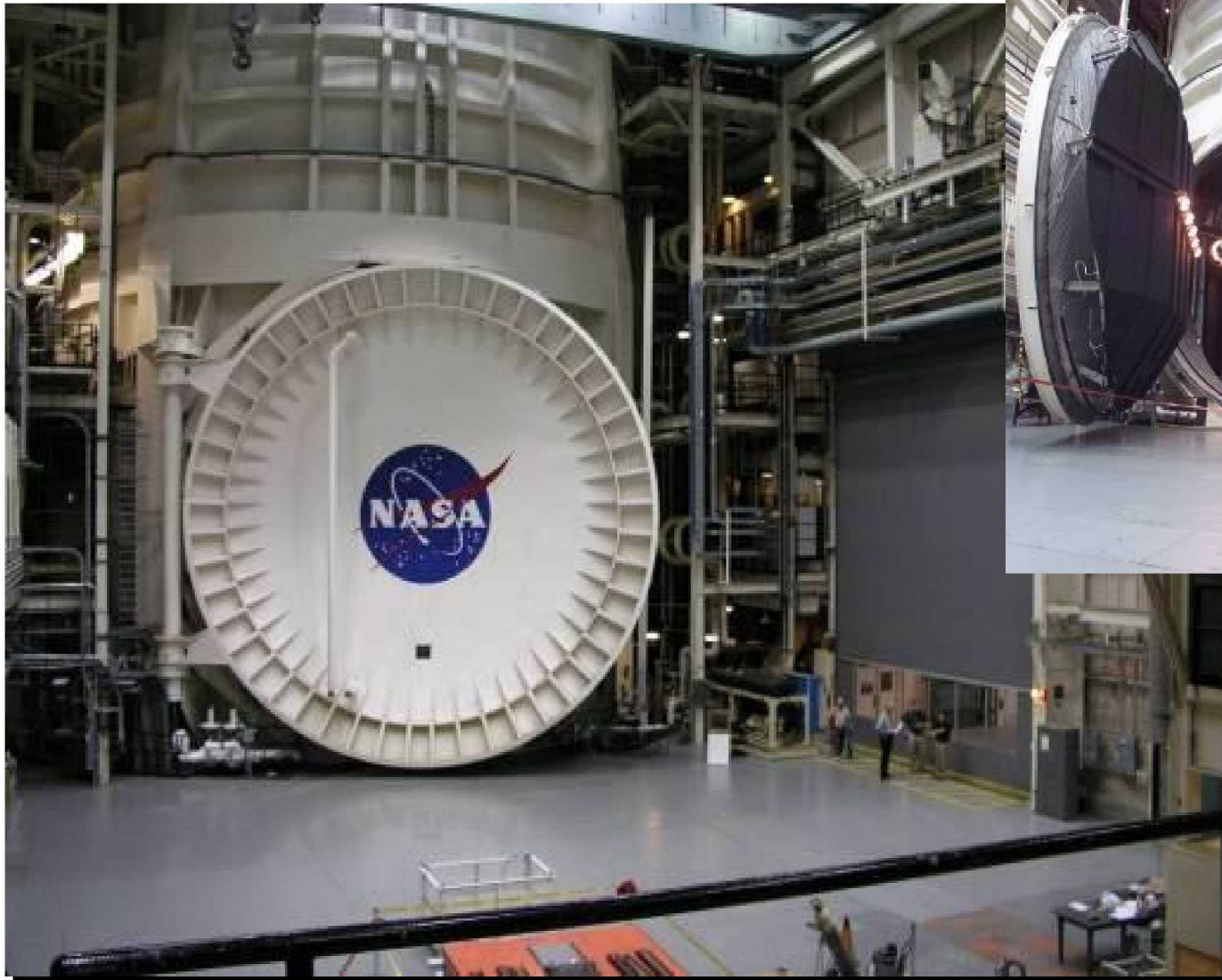
XRCF CCS Cross section
1-10-05
Based on XL 90% Design Review Data

Primary Mirror Cryogenic Tests



JWST Observatory Level Integration & Test

Johnson Space Center Chamber A	
Chamber size	16.7 meter diameter, 35.6 meter tall
Existing Shrouds	LN2 shroud, GHe panels
Chamber Cranes	4 x 7.6 meter fixed, removable
Chamber Door	12 meter diameter
High bay space	~31 m L x 21.6 m W





JWST vs. HST - orbit

Sun

Earth

Moon



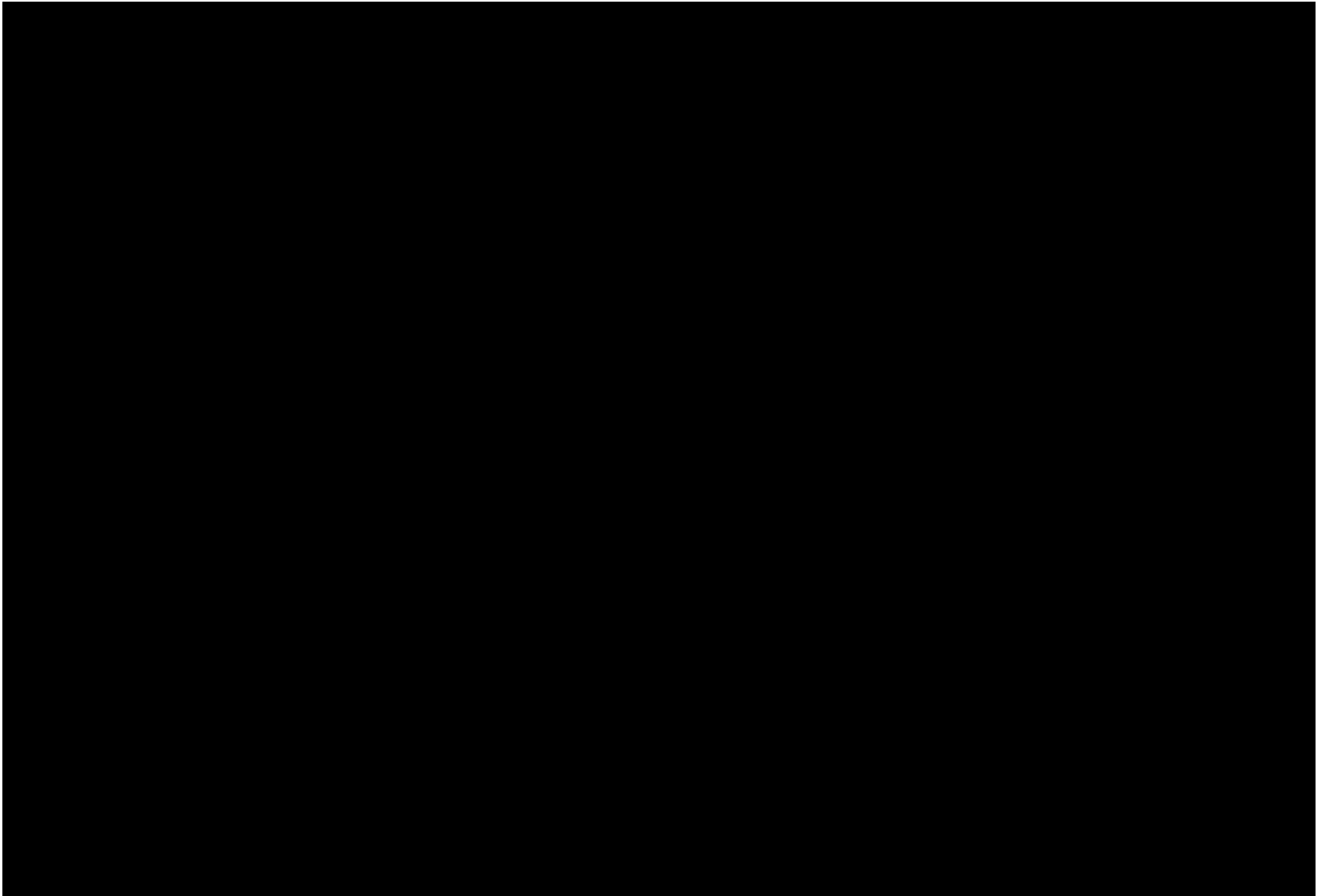
HST in Low Earth Orbit, ~500 km up.
Imaging affected by proximity to Earth

L2

JWST will operate at the 2nd Lagrange Point (L2) which is 1.5 Million km away from the earth



JWST Deployment Movie





Any Questions

