

Millimeter- and Submillimeter-Wave Remote Sensing Using Small Satellites

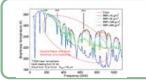
NASA Goddard Space Flight Center

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Outline





Motivation and science drive



GSFC mm- and sub-mm-wave instruments

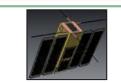




Candidate instrument architecture



Small rocket technology (SMART)



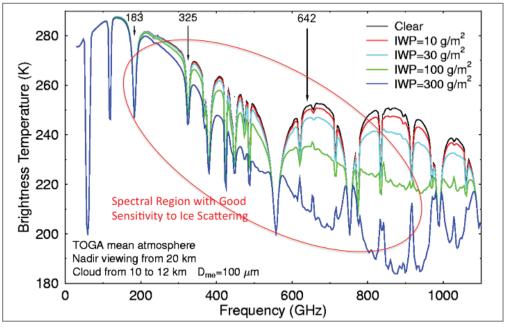
IceCube

Conclusion

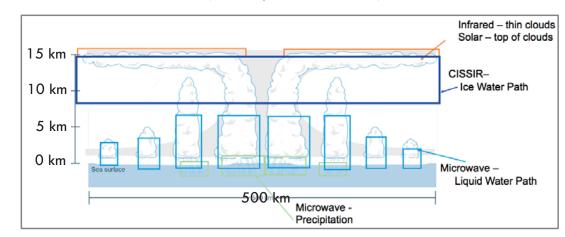
Motivation and science drive



- Cloud ice properties are fundamental controlling variables of atmospheric radiation and precipitation
- Large discrepancies in Ice Water Path (IWP) exist in Global Circulation Models
- Limited availability of data and poor assumptions about the cloud micro- and macro-physical properties of clouds are principle contributors to the discrepancy
- No ice cloud measurements currently exist for the intermediate altitudes
- mm- and submm-wave radiometry offers great potential to fill the measurement gap in the middle and upper troposphere

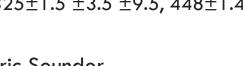


The spectral region with good sensitivity to ice cloud scattering (courtesy of Frank Evans).

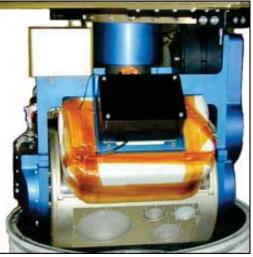


GSFC mm- and sub-mm-wave instruments

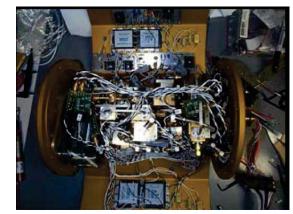
- CoSMIR: Conical scanning mm-wave imaging radiometer
 - Azimuth over Elevation dual axes gimbals for cross-track and concial scanning
 - 6 receivers, 9 channels at 50.3, 52.6, 89 (H&V), 165.5 (H&V), 183.3 \pm 1, 183.3±3, ±7
 - FOV: 4° beam width
 - Scan head: cylinder with 21.5 cm diameter and 28 cm length
- CoSSIR: Compact scanning sub-mm-wave imaging radiometer
 - 6 receivers, 12 channels at 183.3 \pm 1 \pm 3 \pm 7, 325 \pm 1.5 \pm 3.5 \pm 9.5, 448 \pm 1.4 \pm 3 ±7.2, 642 (H&V), 874
- HyMAS: Hyperspectral Microwave Atmospheric Sounder
 - Hyperspectral with 52 channels (in partnership with MIT-LL)
 - 6 receivers at 172-183 GHz (H&V),
 - 2x108-119 (H&V)
- GMI: GPM Microwave Imager
 - **Receiver Components:**
 - Noise sources X-band—Ka-band
 - Mixers up to G-band





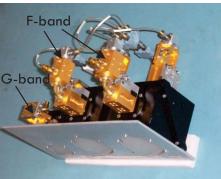


CoSSIR interior









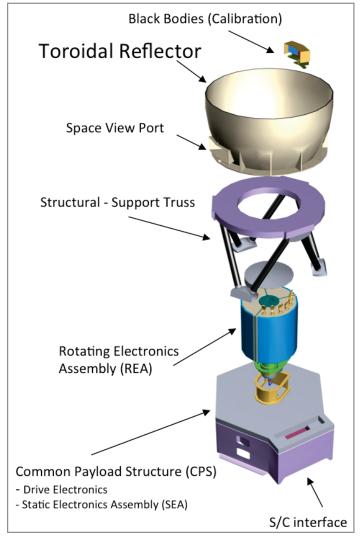
System Study

NASA

- A small, compact sub-mm instrument has many opportunities in future NASA, NOAA and international missions
- GSFC developed design and analysis tools to evaluate size, weight, and power (SWaP) of sub-mm-wave instruments
 - Cross-track vs. Conical scan
 - Spatial resolution, aperture size, and spatial resolution
 - Receiver selection and number of channels
- Scientifically-valuable sub-mm-instrument configuration fit within 30–50 kg mass, suitable for Earth-venture class opportunities

Projected Aperture		Receivers						# of Ch	Spatial Altit 400 km		Power (W)	Mass (kg)
36.0 cm	Conical	183 V	325 V	448 V	640 V			11	4.5	9.8	61.2	107.5
50.0 CIII						640 H		8			61.1	107.3
19.8 cm		1021/			640 V			7	8.2	17.7	52.7	38.0
		183 V				640 H		8			61.1	41.3
	Conical		325 V	325 V 448 V				8	4.6	10.0	54.0	37.7
				448 V		C 4 0 1 1		9			62.4	41.0
						640 H	874 V	3	2.3	5.1	54.3	37.7
	Conical		325 V	448 V				8	6.2	13.5	54.0	28.9
					640 V		2	3.2	6.8	45.9	25.6	
14.7 cm						640 H	874 V	3	3.2	6.8	54.3	28.9
	Cross-		225.14					8	3.5	6.9	49.5	
	Track		325 V	448 V		640 H		9	3.5	6.9	57.9	
36.0 cm	Conical	cal 183 V	V 325 V	448 V	640 V	640 H	074.14	12	4.5	9.8	165	137.8
							874 V	12			78.0	114.0

Conical Scan Instrument concept



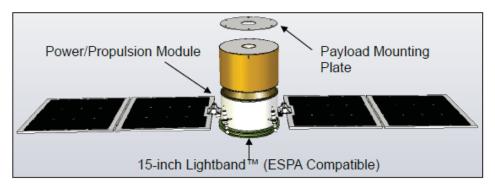
Small rocket/spacecraft technology (SMART)

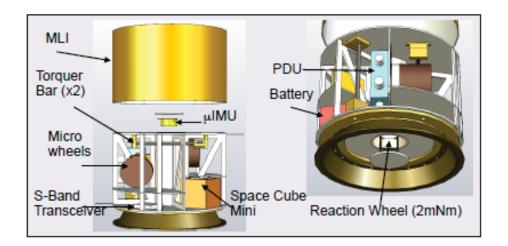


□ SMART is a

- Microsatellite prototype that enables focused science and technology missions with limited resources
- Demonstration of a point design instantiation of a modular, open systems architecture micro spacecraft prototype
- Miniaturized high-performance, power-efficient processing avionics for small expendable launch vehicles, and small orbiting spacecraft
- Modular, reconfigurable, and rapid architecture
- SMART core components are mostly COTS, and replaceable
- Payload mass capability: 30–50 kg
- Payload available power: 284–314 W
- □ Bus dissipated power: 45–75 W

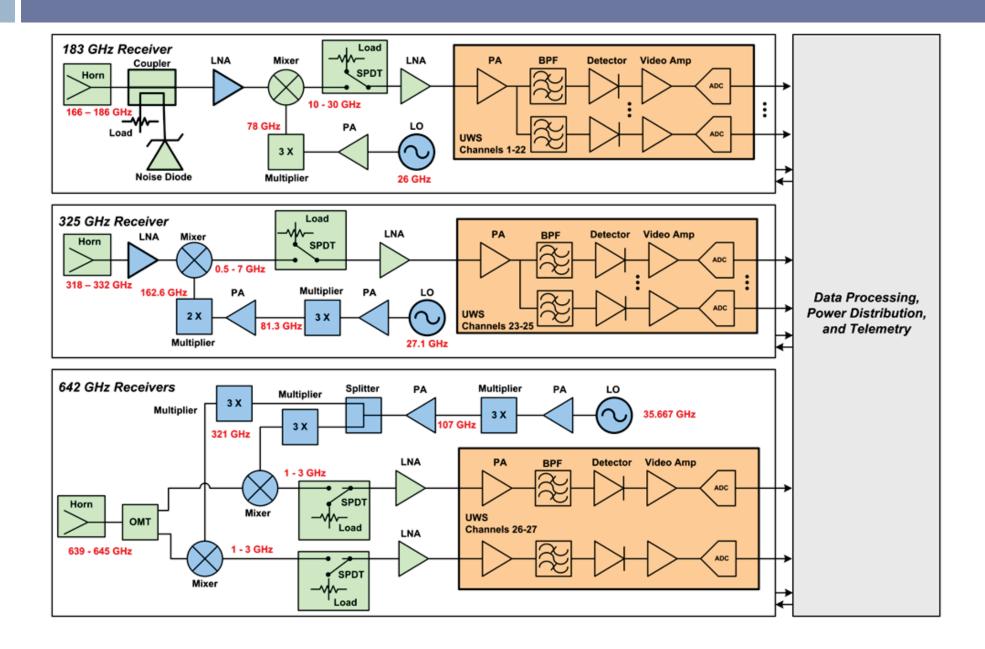
Simplified illustration of SMART







Candidate instrument architecture (SCAMPER)



SCAMPER performance parameters



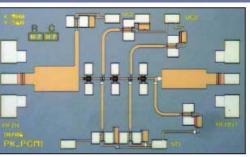
Center frequency & polarization	183 GHz V-pol	325 GHz V-Pol	642 GHz H-Pol	642 GHz V-Pol	
RF pass-band [GHz]	166-186	318-332	639-645	639-645	
IF pass-band [GHz]	10-30	0.5-7	1-3	1-3	
Number of channels	22	3	1	1	
Polarization	Vertical	Vertical	Horizontal	Vertical	
Noise temperature [K]	1060 (SSB)	1540 (DSB)	4120 (DSB)	4120 (DSB)	
NETD [K] @ $\tau_{int} = 7 \text{ ms}$	0.64 (650 MHz BW)	0.49 (2 GHz BW)	1.18 (2 GHz BW)	1.18 (2 GHz BW)	

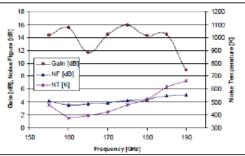
New instrument technologies



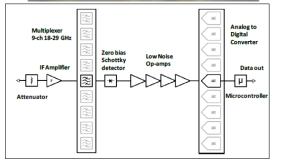
LNAs at 183 GHz and 325 GHz

- Advantages
 - Improved noise temperature
 - Reduced power (mixer with higher conversion loss)
 - Single side-band detection (hyperspectral receivers)
- Available technologies
 - 35 nm InP technology NGAS/JPL [Kangaslati et al., 2008] and BAE
 - SiGe technology Georgia Tech [Coen et al.]
 - Teledyne HBT InP
- Coupled noise source at 183 GHz
 - 9 dB ENR @ 200 GHz noise source developed at GSFC
 - Improving receiver calibration
 - Eliminates requiring view of clear sky
- Ultra Wide Band Spectrometer (UWS) developed by MIT-LL
 - 52-channel, 100 cm³ UWS developed for HyMAS (ESTO/ACT project)









IceCube: 874 GHz radiometer on a CubeSat

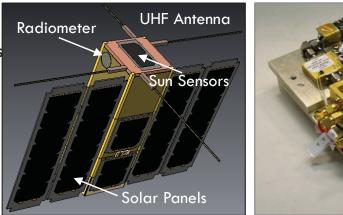


Objective

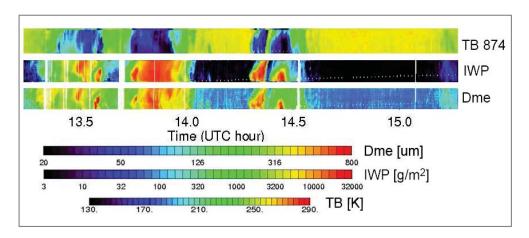
 Develop and validate a flight-qualified 874 GHz receiver for future use in ice cloud radiometer missions

Operation Concept

- Orbit: LEO @ 350 km
- Duration: 28 days
- 15 orbits per day (31 mins pointed at Earth, 9 mins view of clear sky)
- Technologies
 - Sub-mm-wave 874 GHz receiver
 - 2nd Harmonic mixer
 - Frequency doubler & triplers
 - V-band multiplied LO with power modulation
 - IF with noise injection
- □ Receiver performance
 - Noise temperature < 6000 K</p>
 - NEDT < 0.1 K for 1 s integration time
 - 2 K calibration error



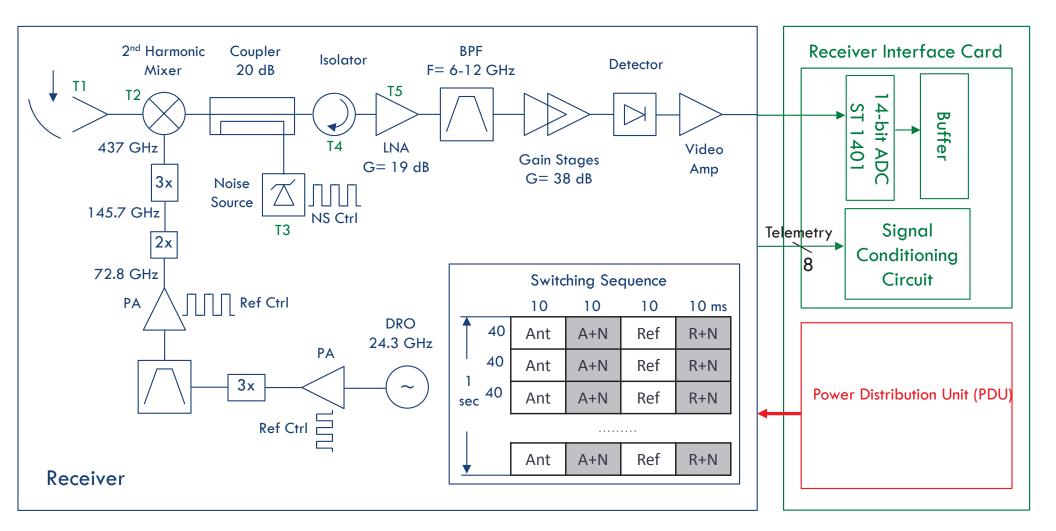




(a) Conceptual drawing of IceCube (radiometer at top). (b) The 874 GHz CoSSIR. (c) First ever 874 GHz cloud measurements acquired by CoSSIR in 2008. CoSSIR measurements of ice clouds were used to successfully demonstrate retrieval of ice water path (IWP) and ice particle median mass-weighted ice particle size (D_{me}) .

IceCube: receiver





Conclusion



- Distributed satellite systems are required to retrieve IWP with 25 % accuracy, to enable improved ice cloud modeling
 - Highly sensitive
 - Cost effective
 - Compact mm- and sub-mm-wave instrument with multiple channels

		SCAMPER	SSMIS	MHS	GMI	
	Cloud ice	Yes	Yes	Yes	Yes	
Science capabilities	lce particle size	Yes	Limited	Limited	Limited	
• • • • • • • • • • • • • • • • • • •	lce particle shape	Yes	No	No	Limited	
	Frequencies [GHz]	183, 325, 642, 874	19, 22, 37, 91, 50- 63,183	89,1 <i>5</i> 7, 183, 190	10 18, 23, 36, 89, 165, 183	
Instrument	IF channels	28	24	5	13	
properties	Dual-pol [GHz]	642	19, 37, 91	None	10.7, 18.7, 36.5, 89,165.5	
	RF cal [GHz]	183	None	None	10.7, 18.7, 23.8, 36.5	

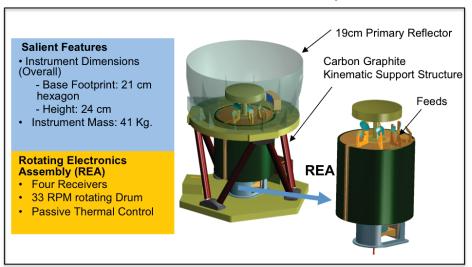
Questions?



Backup: system study

Distributed satellites

- Observations of ice clouds requires frequently updated measurements by a distributed satellite system, due to a strong diurnal variation
- Conically scanning/ Cross track
 - Polarization preservation
- Radiometric stability
 - Antenna is not exposed to open air
- mm- and sub-mm-wave receivers housed in the compact scan head
- Mass budget
 - 41 kg
- Power budget:?



Small Satellite Mission Concept



Backup: concept study table

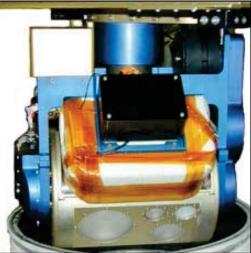


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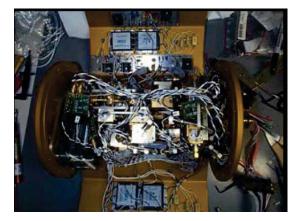
Backup: previous mm- and sub-mm-wave work



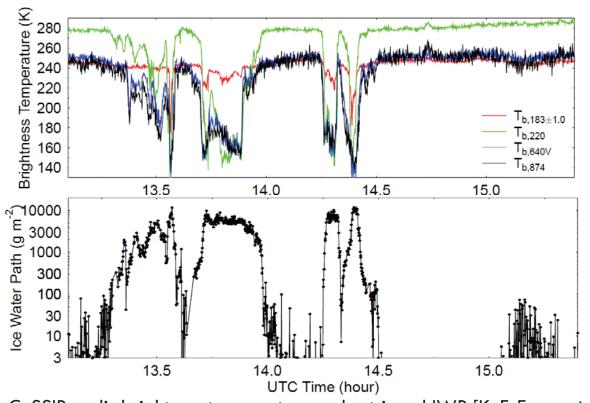
CoSMIR



CoSSIR interior



- CoSMIR: Conical scanning mm-wave imaging radiometer
 - Four receivers
 - □ 9 channels at 50.3, 52.6, 89 (H&V), 165.5 (H&V), 183.3±1, 183.3±3, ±7
- CoSSIR: Compact scanning sub-mm-wave imaging radiometer
- □ HyMASS:

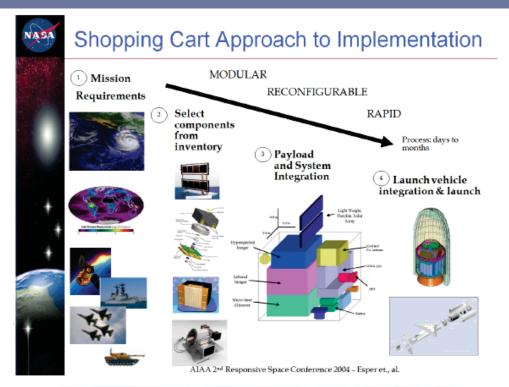


CoSSIR nadir brightness temperature and retrieved IWP [K. F. Evans et al.]

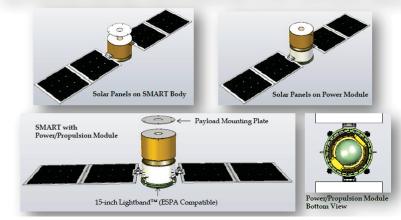
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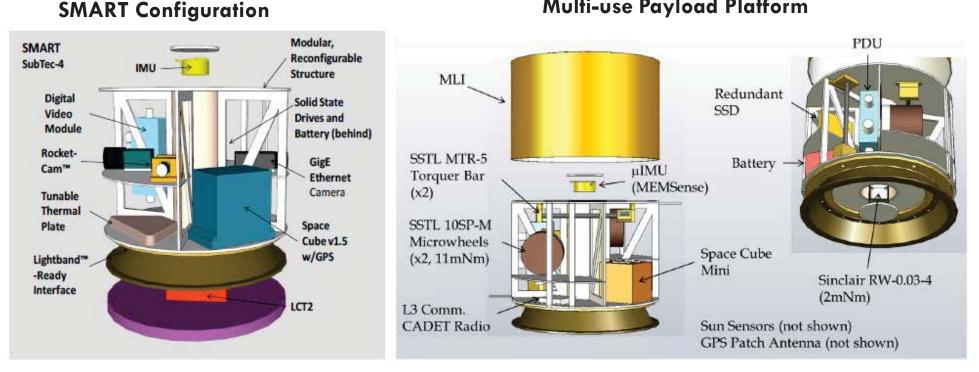


Shown 359W BOL Solar Arrays with TJ Cells (Si Cells were also used in mission analyses)





- SMART core components are mostly COTS, and replaceable
- Payload mass capability: 30-50 kg
- Payload available power: 284–314 W
- Bus dissipated power: 45–75 W



Multi-use Payload Platform

Backup: atmospheric sensing spectral regions

