

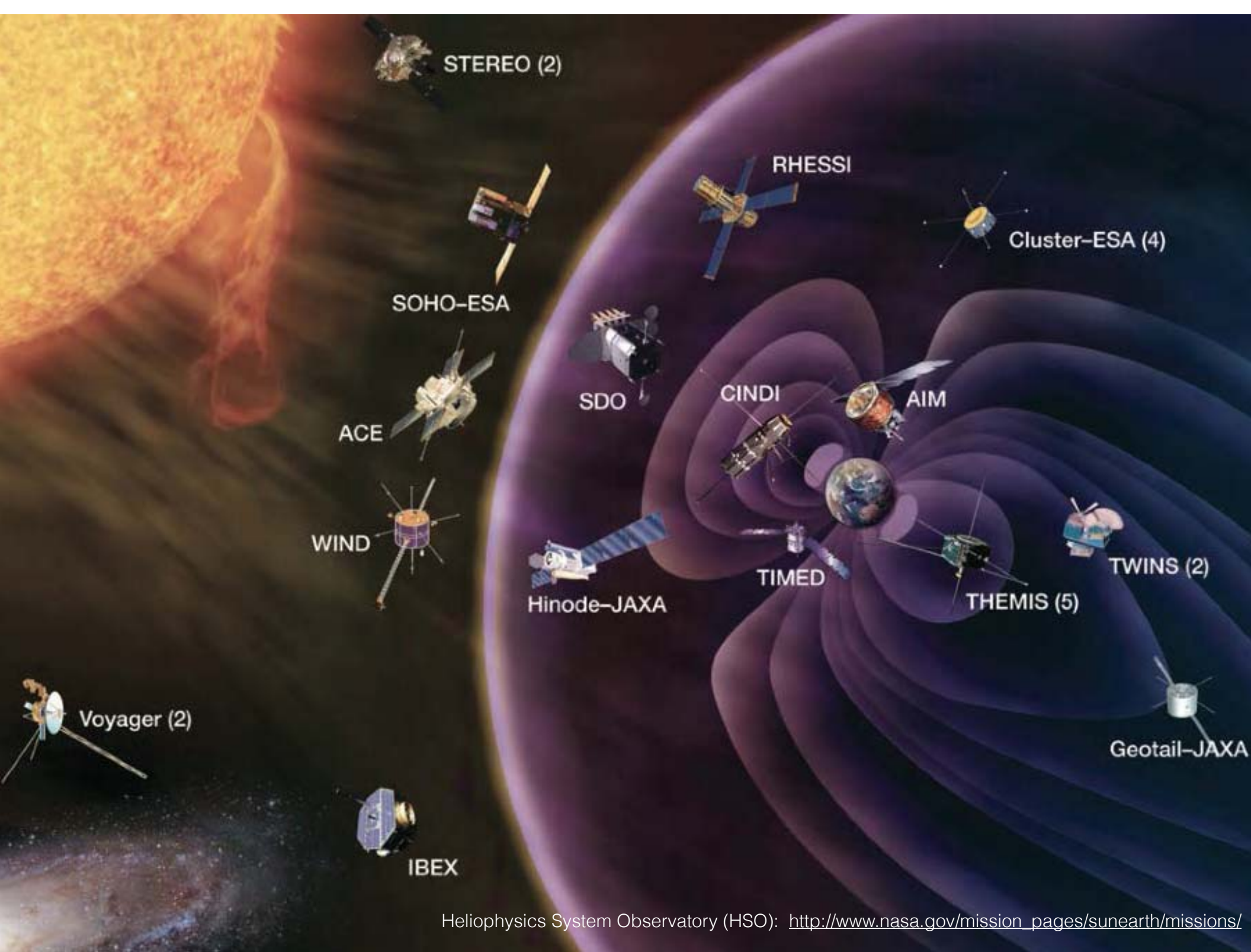
Solar Flares



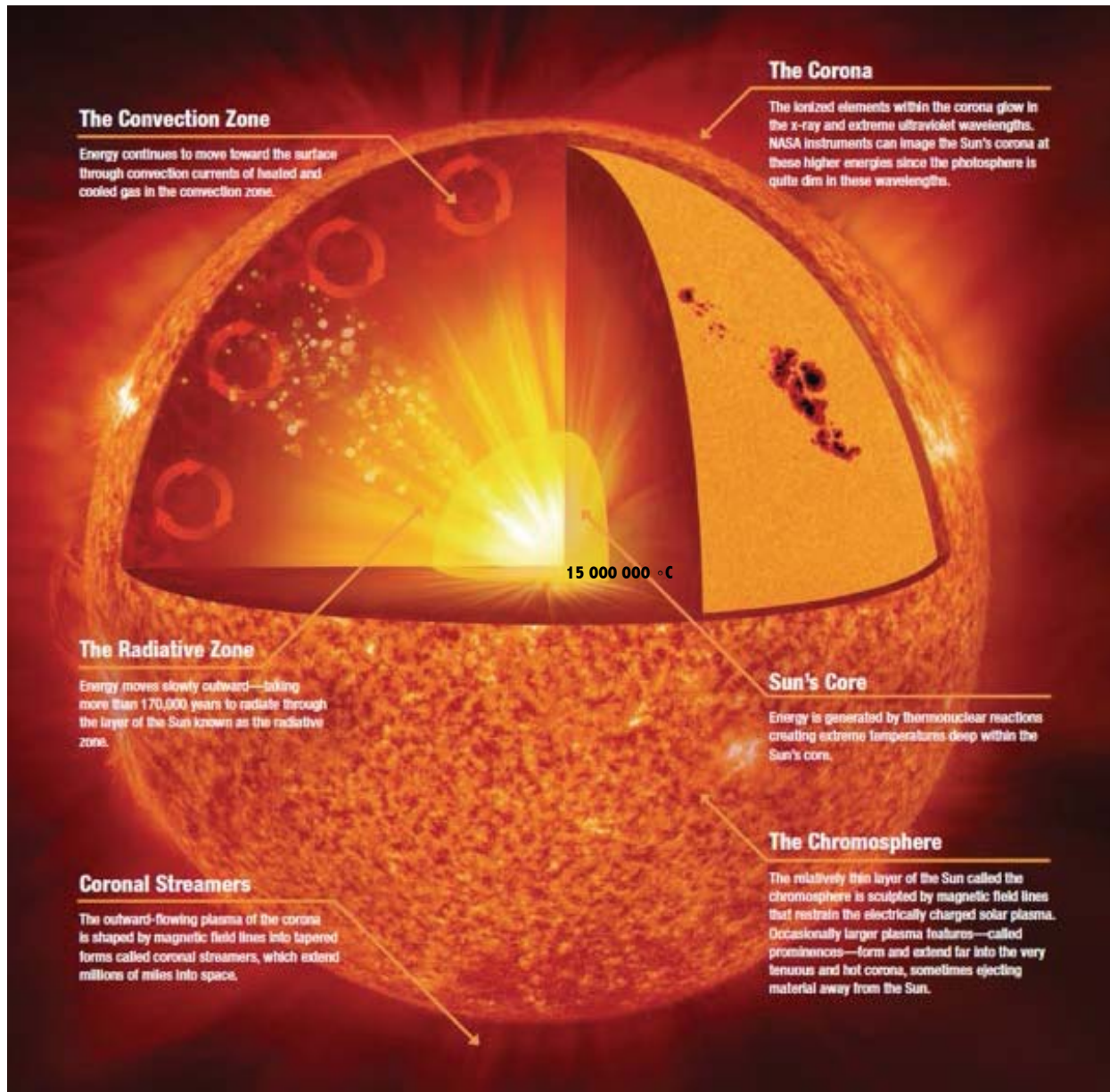
Sabrina Savage (NASA/MSFC)

Heliophysics System Observatory (HSO)

- Fleet of solar, heliospheric, geospace, and planetary satellites designed to work independently while enabling large-scale collaborative investigations.



The Sun in Layers



Converts 4 million tons of matter into energy every second.

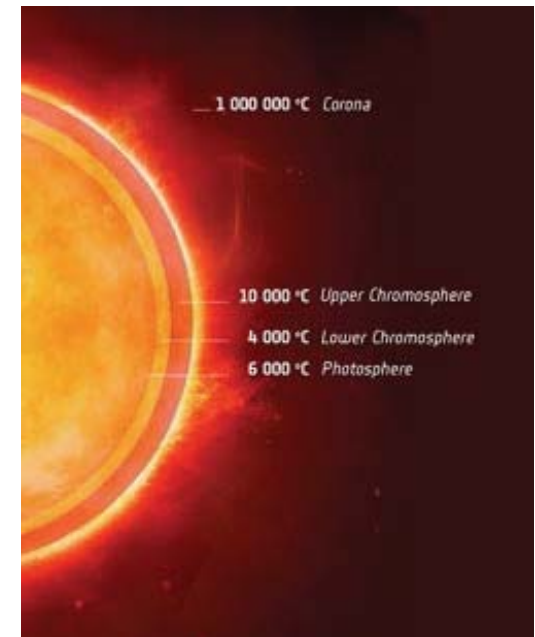
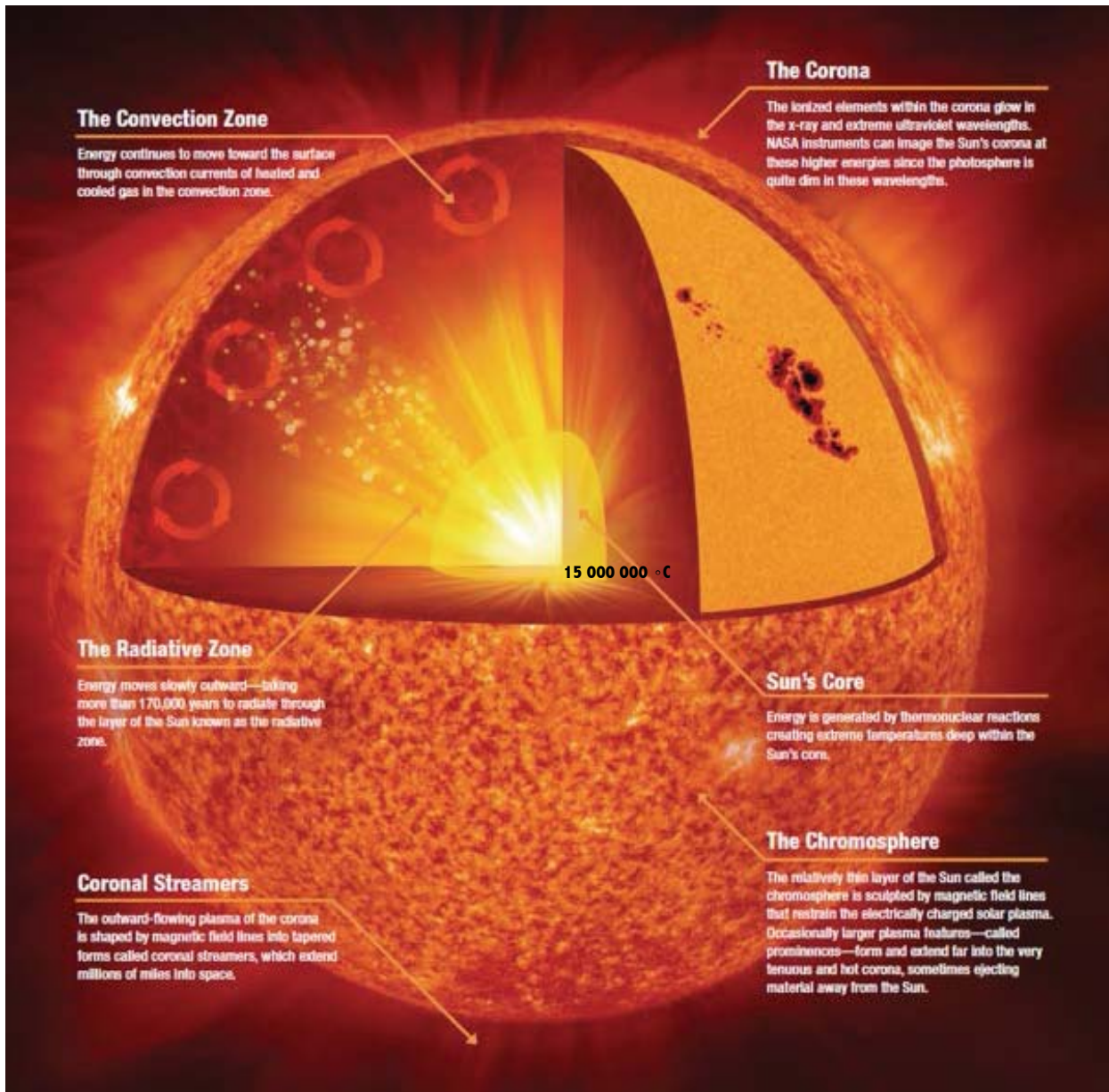
Core is as dense as lead.

Interplay between magnetic pressure and gas (plasma) pressure.

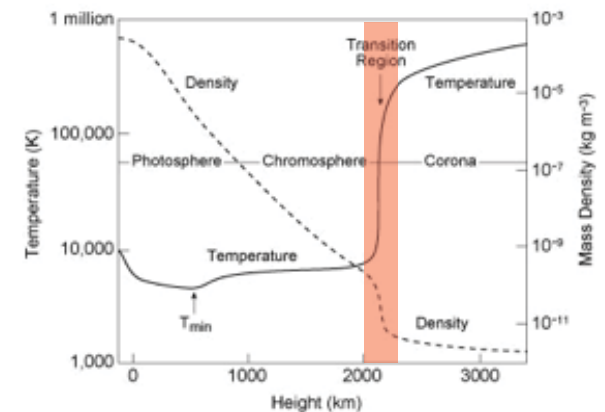
"Mysteries of the Sun": NASA / Jenny Mottar

Sun Facts: <http://solarscience.msfc.nasa.gov/>

The Sun in Layers



European Space Agency (ESA)

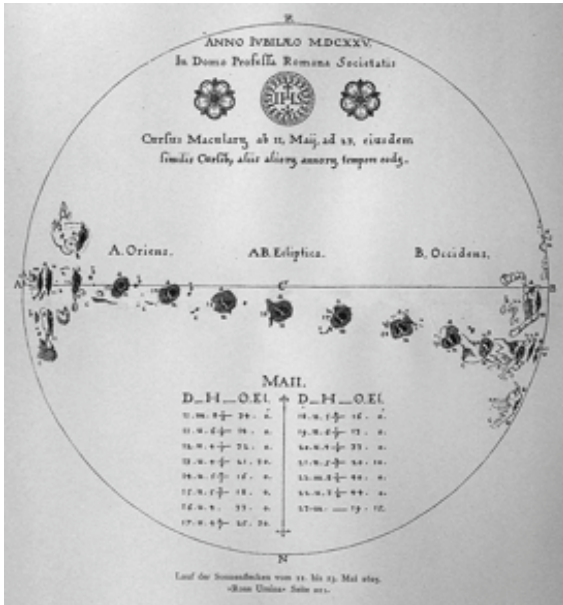


Smithsonian Astrophysical Observatory (SAO)

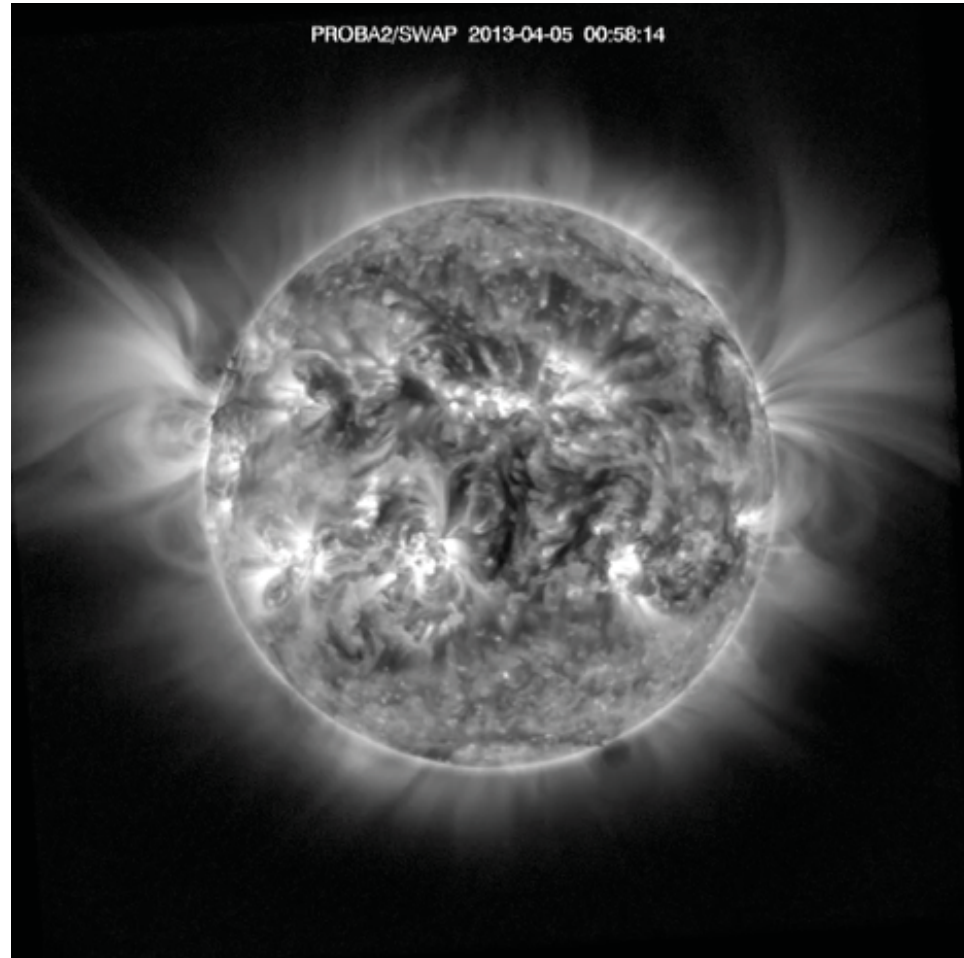
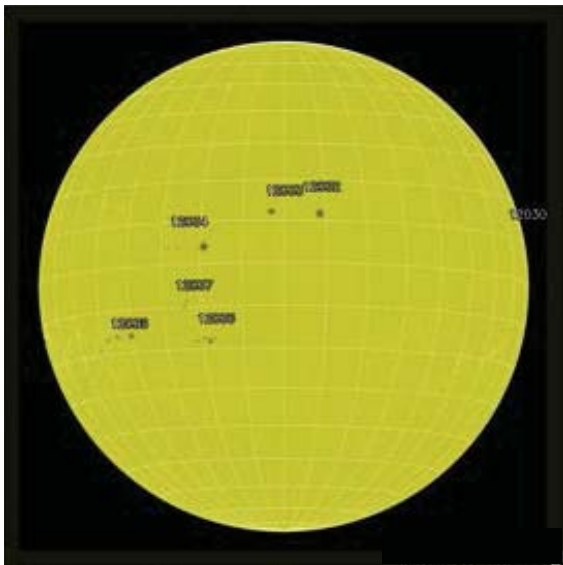
"Mysteries of the Sun": NASA / Jenny Mottar

Sunspots & Active Regions

1625 May: Christoph Scheiner



2014 April 14: SDO HMI 6173 A

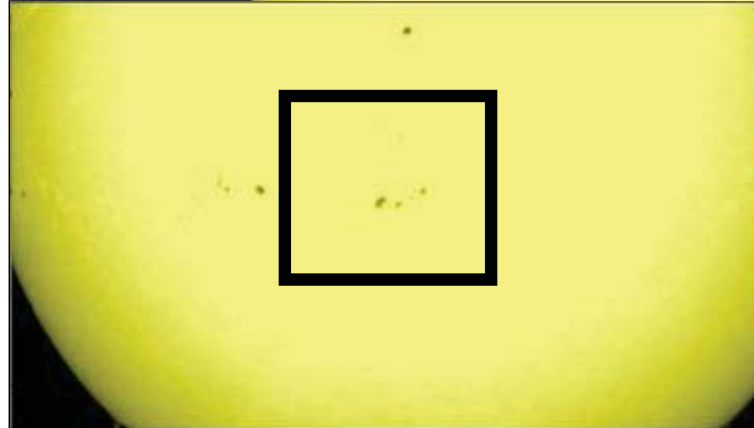


European Space Agency (ESA) / Royal Observatory Belgium (ROB)

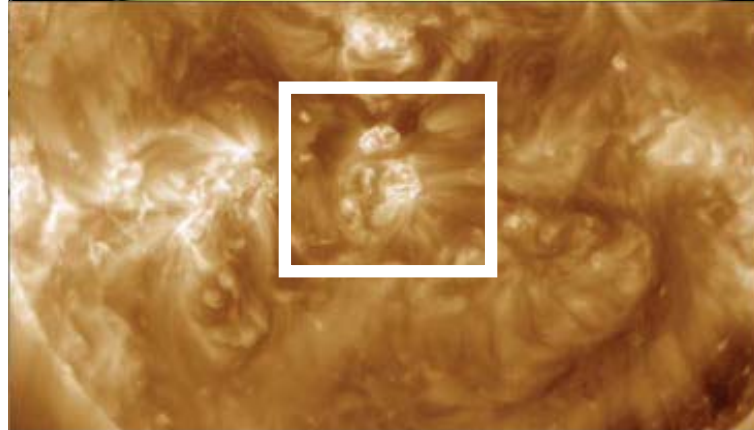
Sunspots & Active Regions

Formation

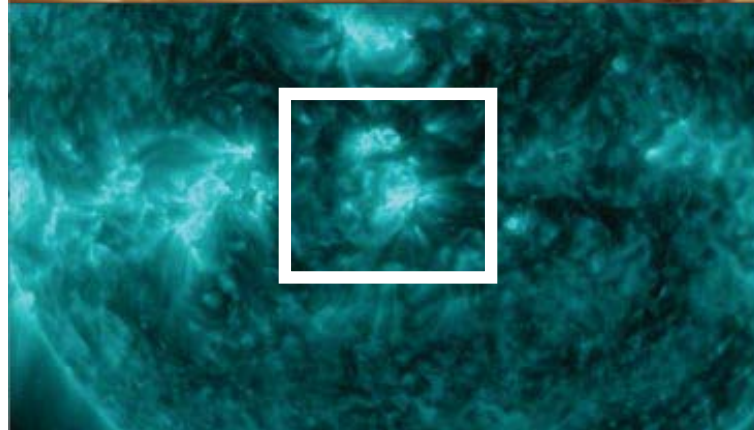
4500 A



193 A

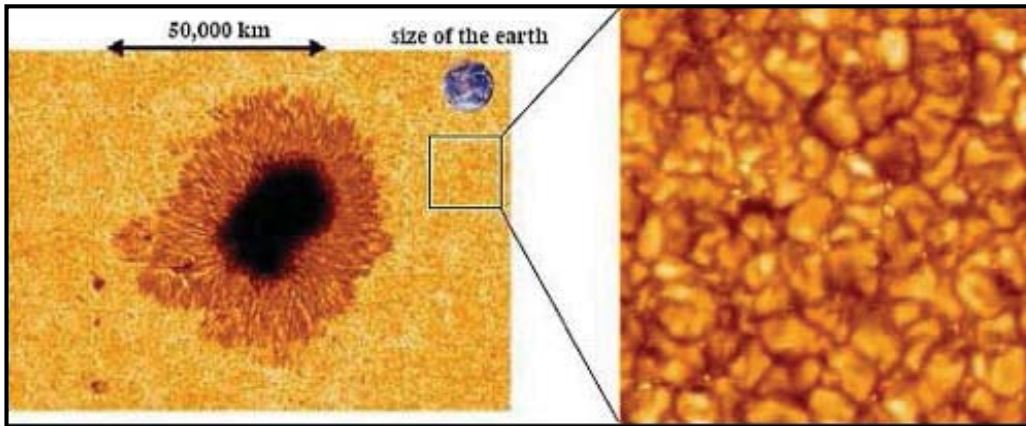


131 A



SDO / AIA
2014 Apr 13 - 15

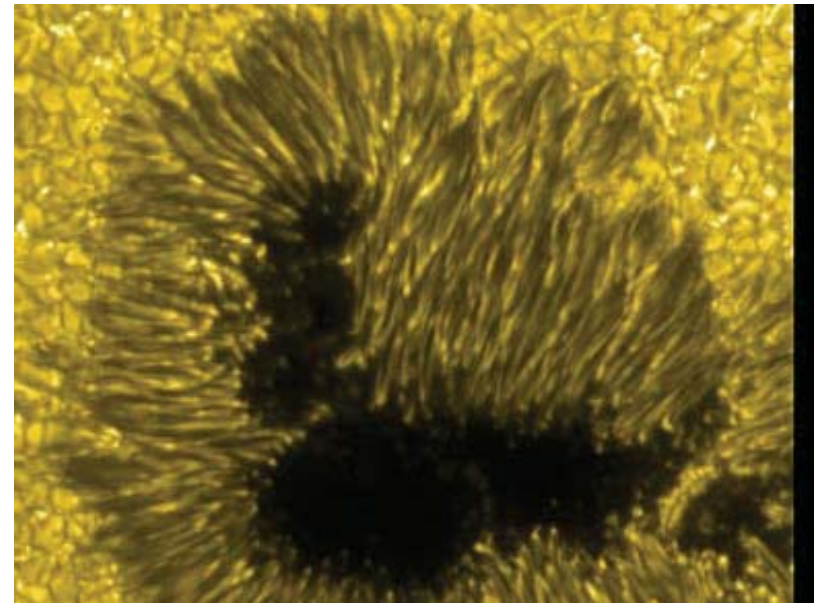
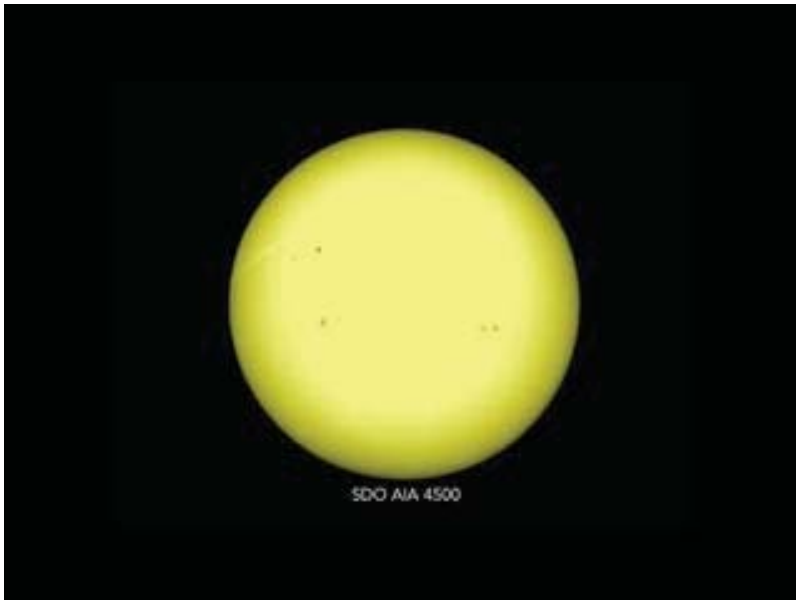
Sunspots & Active Regions



Hinode SOT: NASA / JAXA / NAOJ

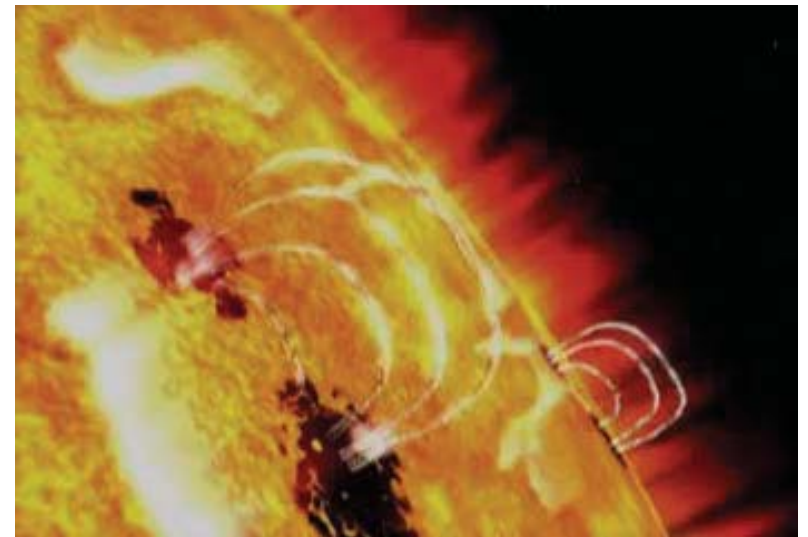
Sunspot Magnetic fields ~ 3000-6000 times stronger than Earth's field.
Magnetic pressure dominates gas pressure in spot, thus inhibiting convective flow of heat.

JHelioviewer *SDO* / AIA 2014 Apr 04

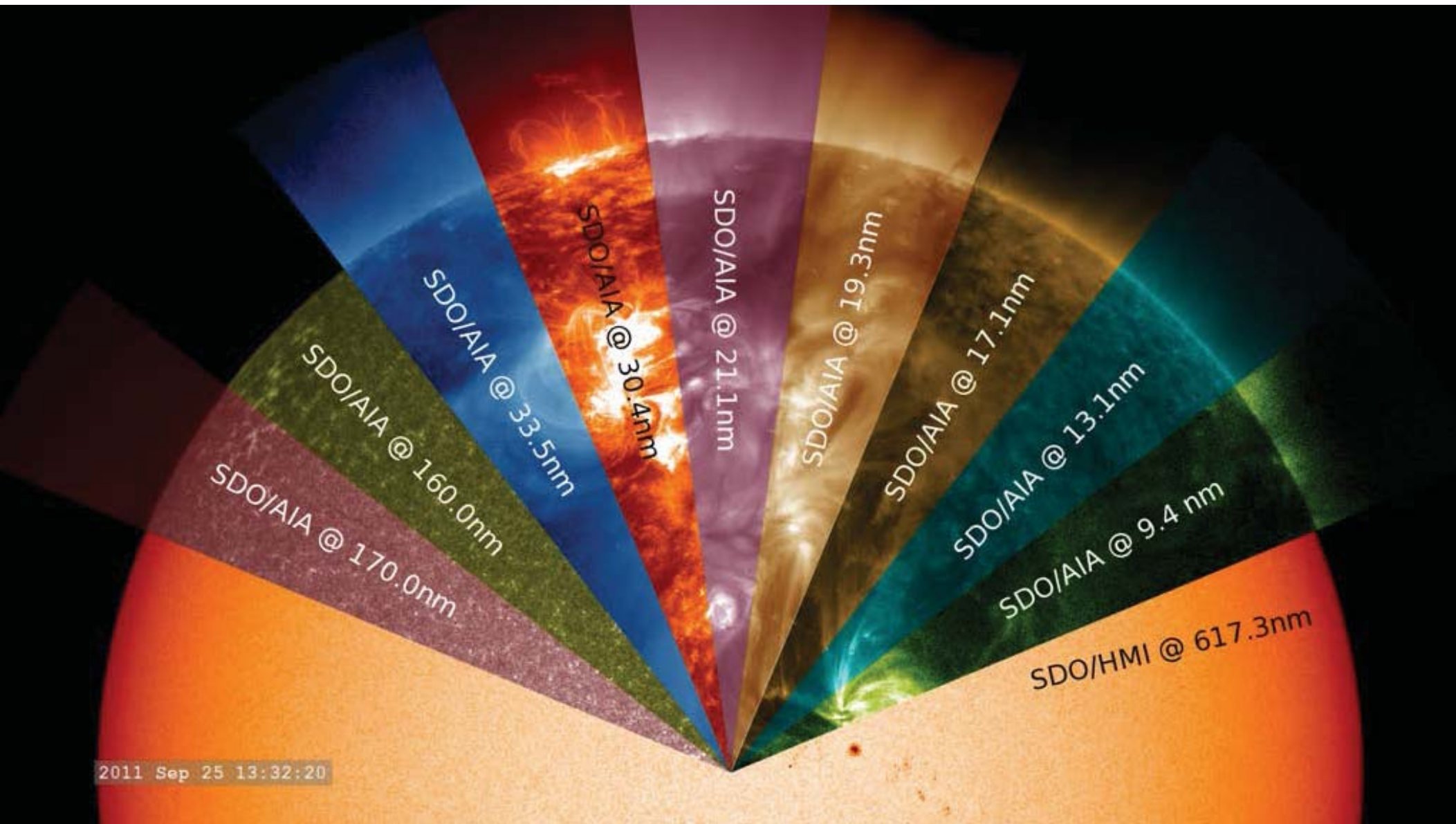


SOT (CN line 3883 Å); 2007 May 2

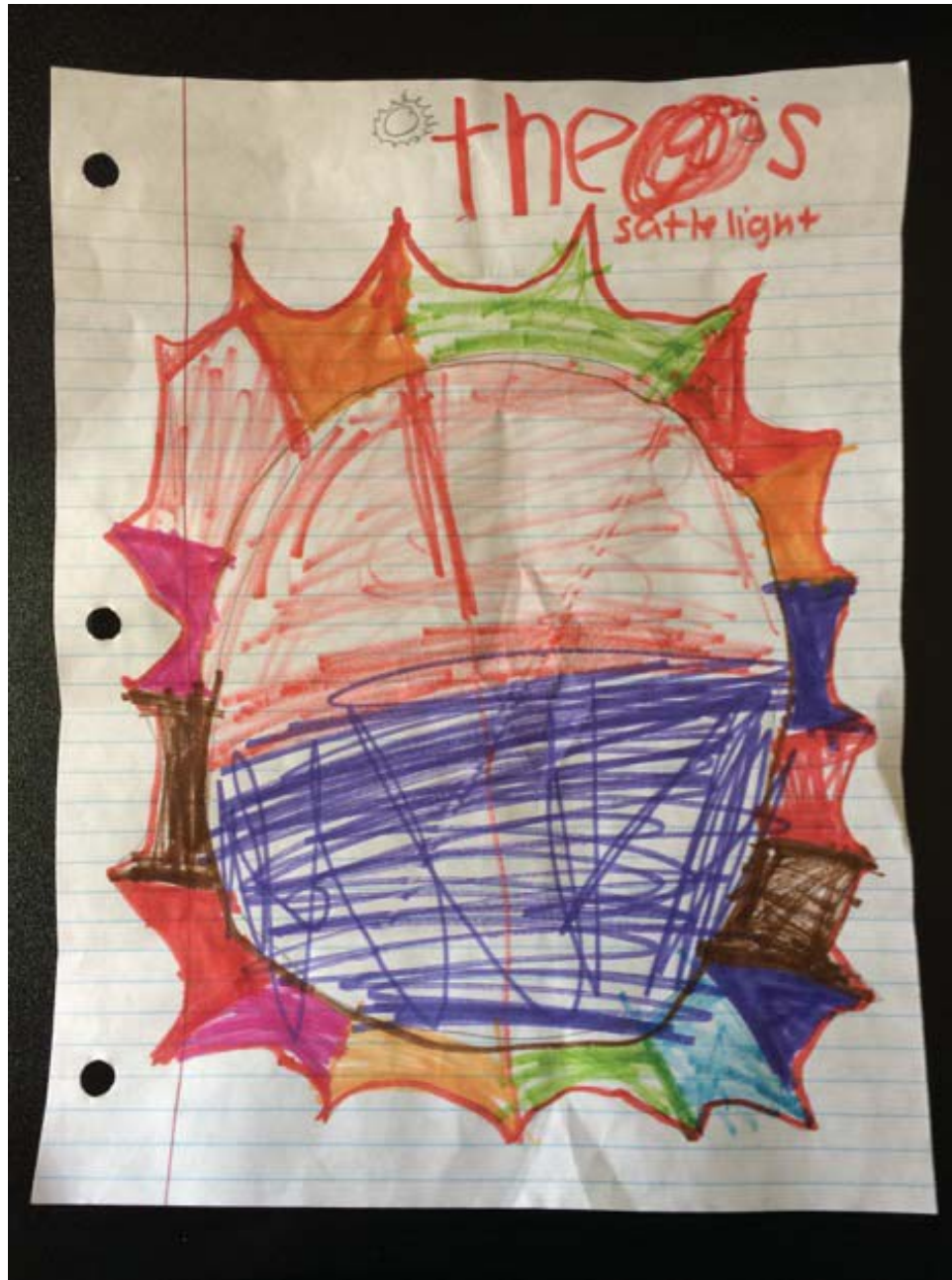
SOHO animation gallery



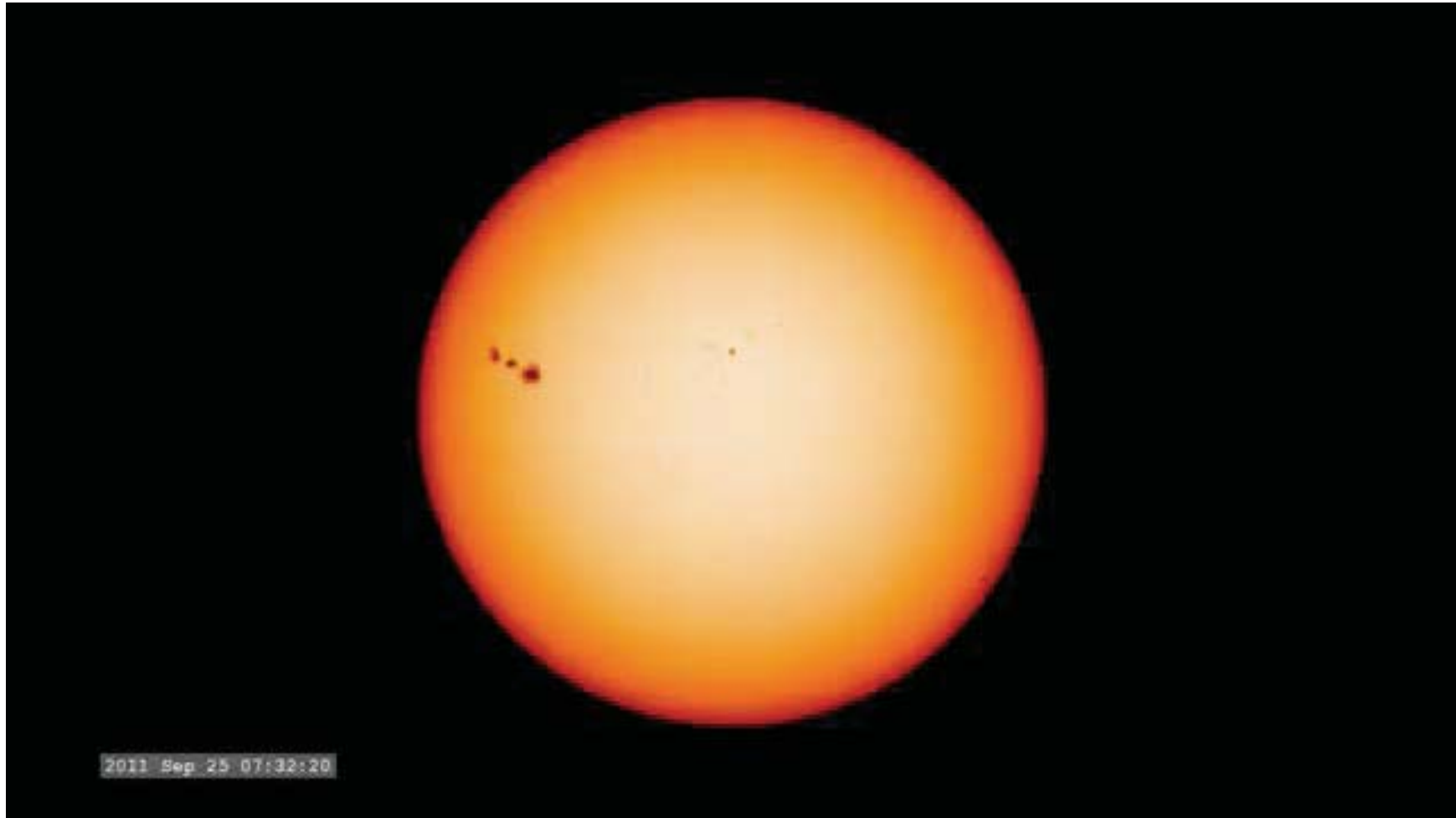
Sunspots & Active Regions



Sunspots & Active Regions



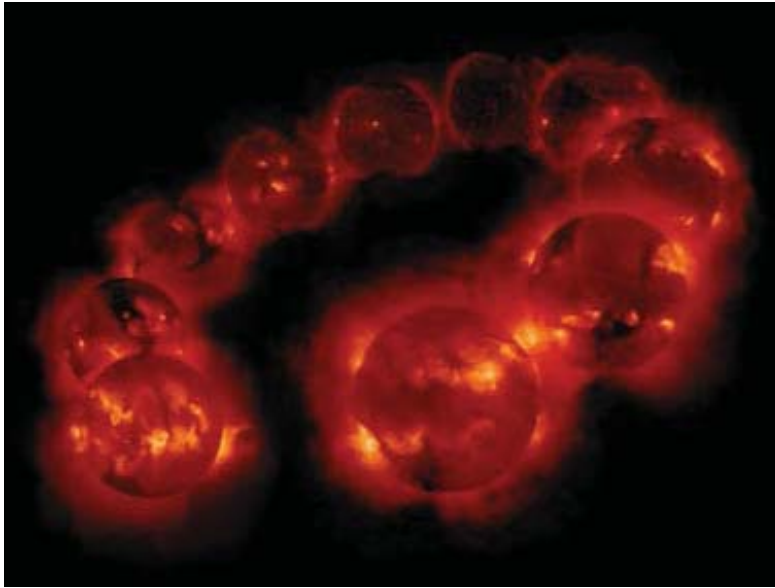
Sunspots & Active Regions



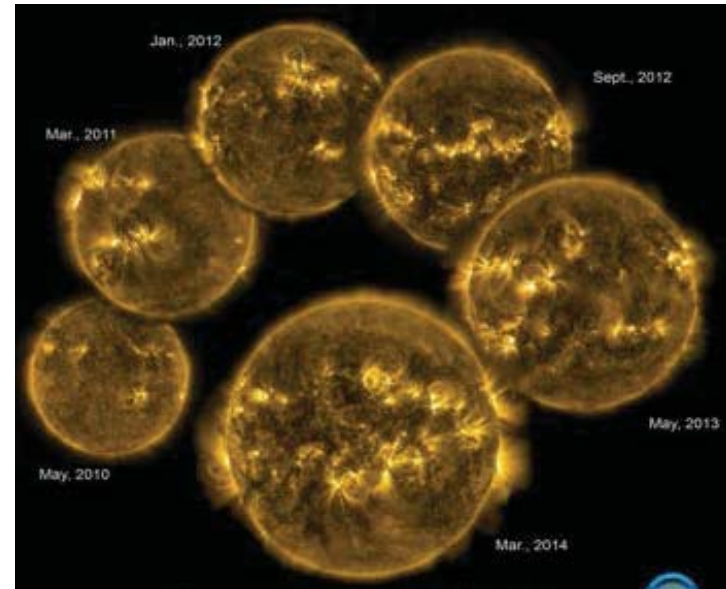
“SDO Jewel Box”

Solar features as seen with 10 different filters (i.e., plasma at different temperatures).

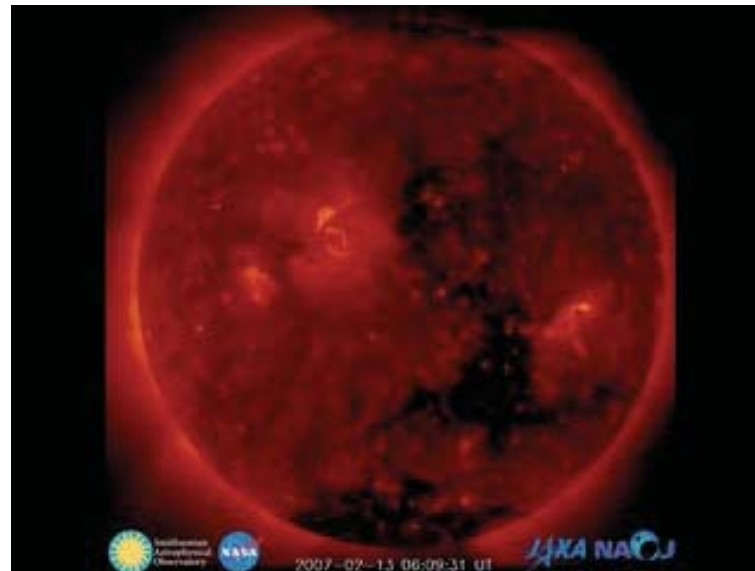
Solar Cycle



Yohkoh / SXT
1991 - 1999



SDO / AIA 171 A
2010 - 2014

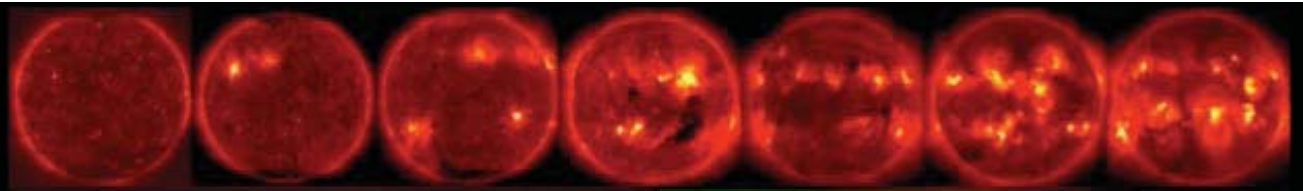


Hinode / XRT 2007 - 2012

9 - 14 year cycle

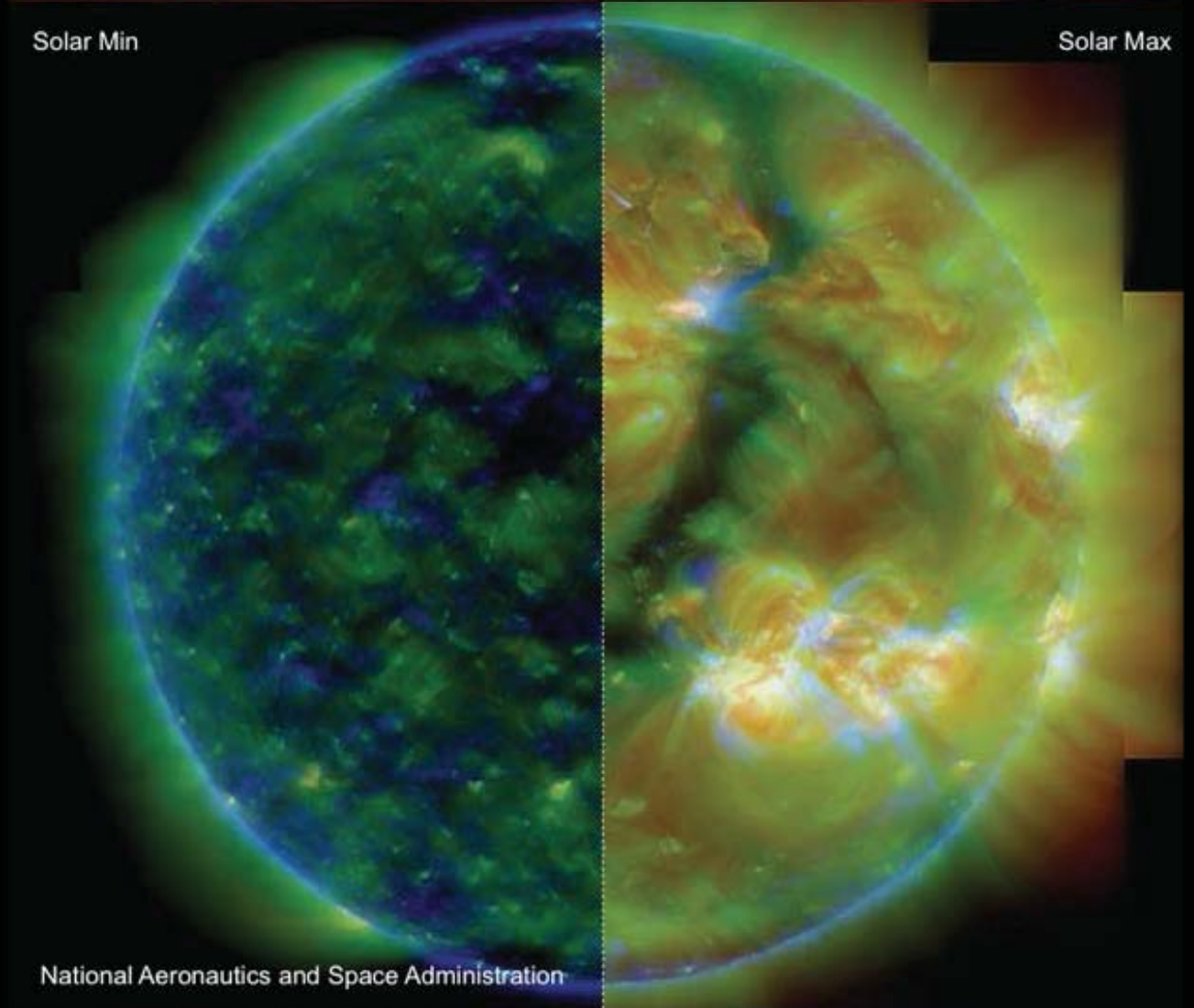
Hinode:

A Comprehensive
Mission to Study
the Variable Sun

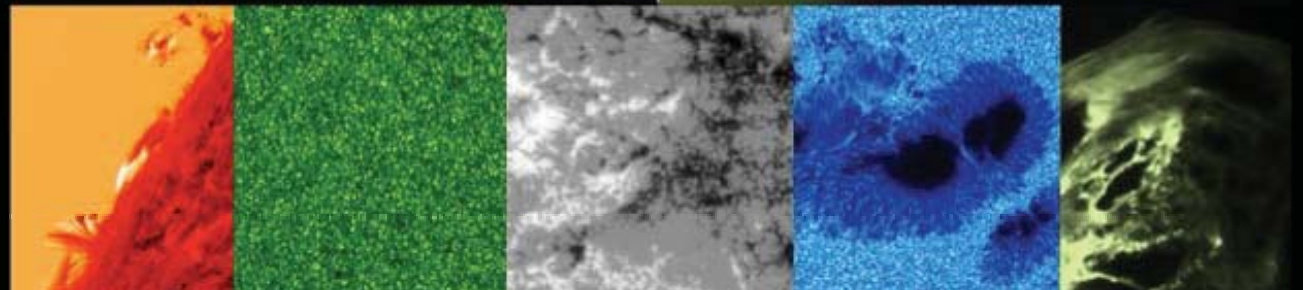


Solar Min

Solar Max

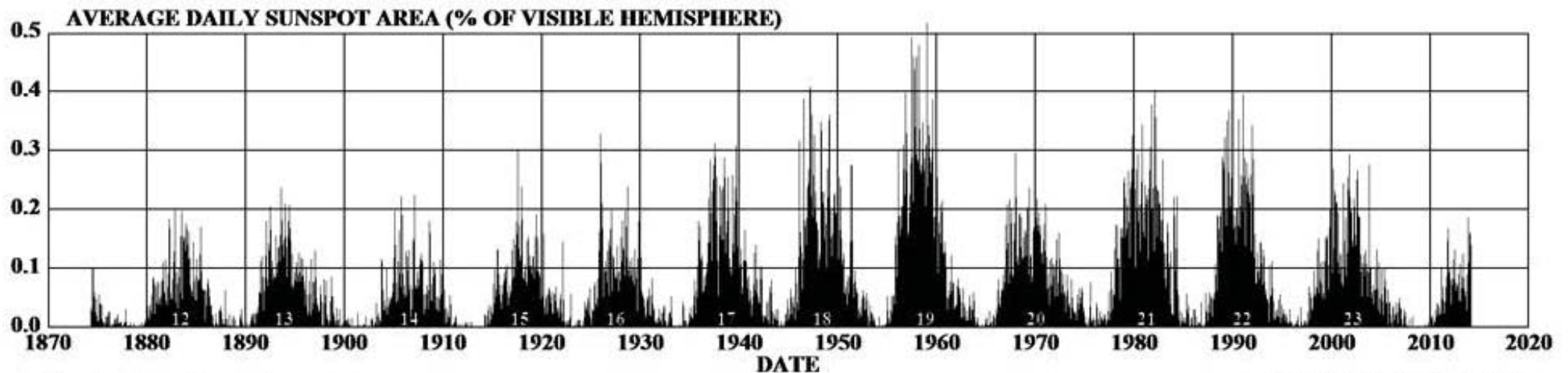
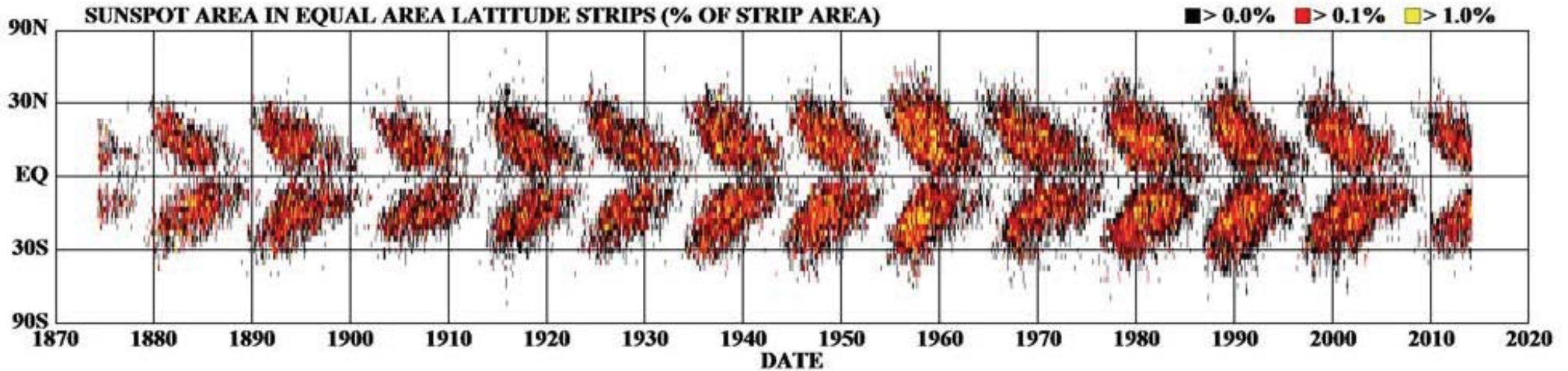


National Aeronautics and Space Administration



Solar Cycle

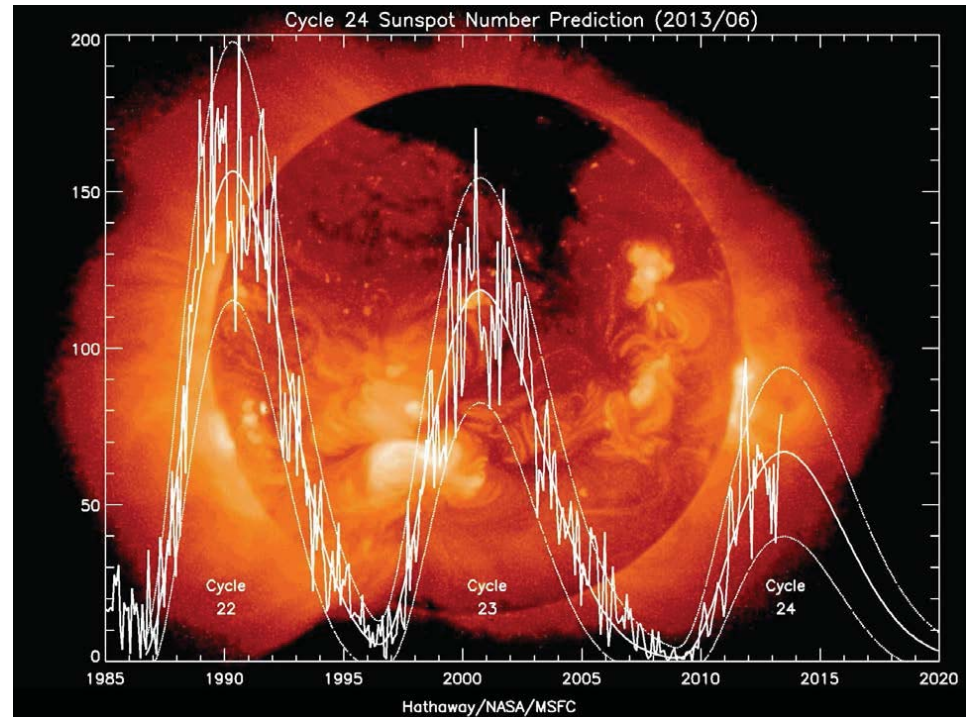
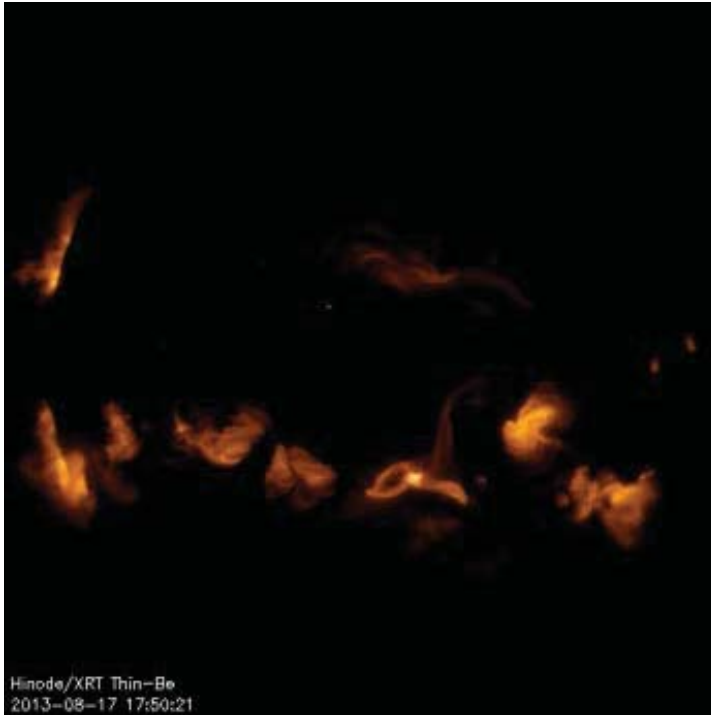
DAILY SUNSPOT AREA AVERAGED OVER INDIVIDUAL SOLAR ROTATIONS



<http://solarscience.msfc.nasa.gov/>

HATHAWAY/NASA/MSFC 2014/04

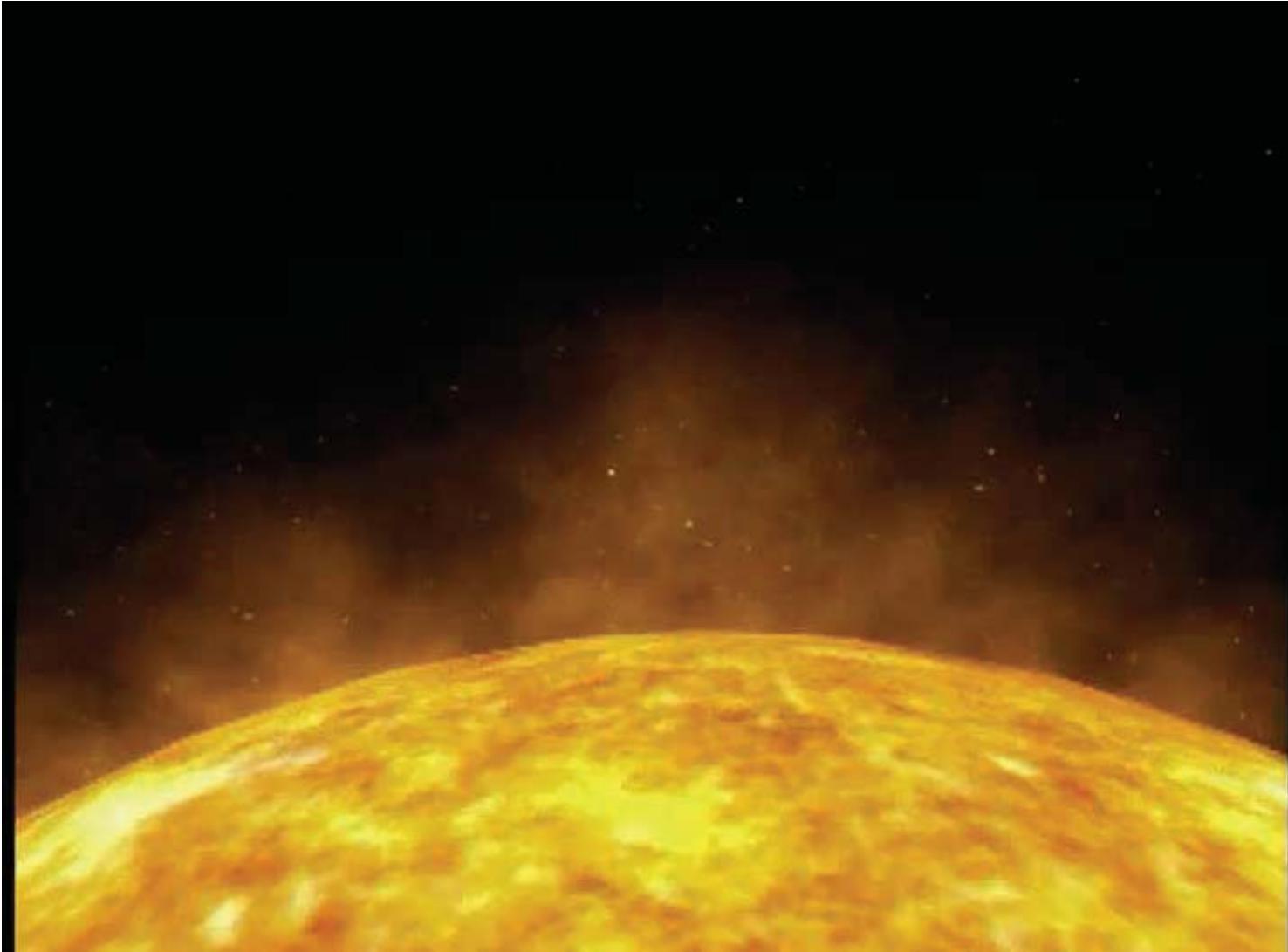
Current Cycle



#24 — Smallest cycle in ~100 years

<http://solarscience.msfc.nasa.gov/SunspotCycle.shtml>

Sun-Earth Interaction

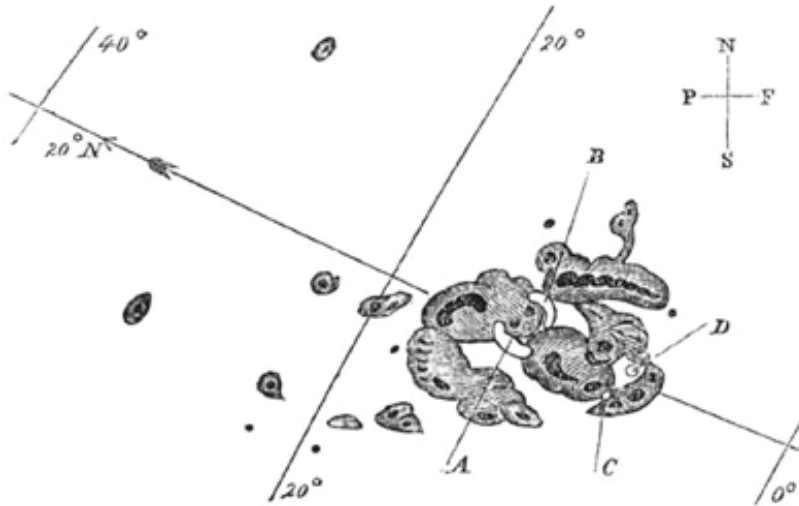


Solar storms cause the *Earth* to lose up to 100 tons of atmosphere into space.

Aurora mostly caused by ionospheric particles disrupted by currents induced from the coronal mass ejection — not the solar wind directly.

Aurora can generate up to 100 trillion watts of power.

Impacts of Space Weather



1859 Carrington Event
Largest Geomagnetic storm recorded



M. A. Shea, Geophysics Directorate, Phillips Laboratory
1989 Superstorm Blackout, \$6 Billion loss to economy



J. Kappenman
2008

NOAA / Space Weather Prediction Center
Space Weather Now
2014 Apr 11 16:43 UTC (Apr 11 10:43 MDT)

Latest GOES Solar X-ray Image

NOAA Scales Activity
Range 1 (minor) to 5 (extreme)

NOAA Scale	Past 24 hrs	Current
Geomagnetic Storms	none	none
Solar Radiation Storms	none	none
Radio Blackouts	none	none

Alerts
Latest Alert: Apr 10 0052 UTC ALERT: Type II Radio Emission

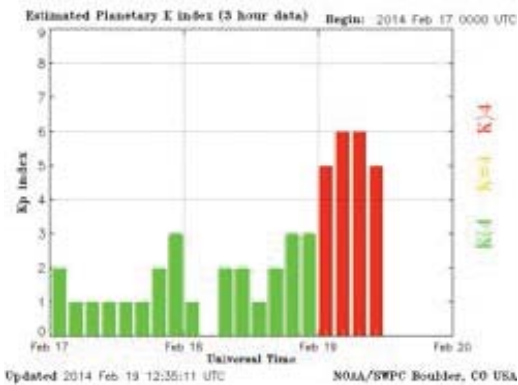
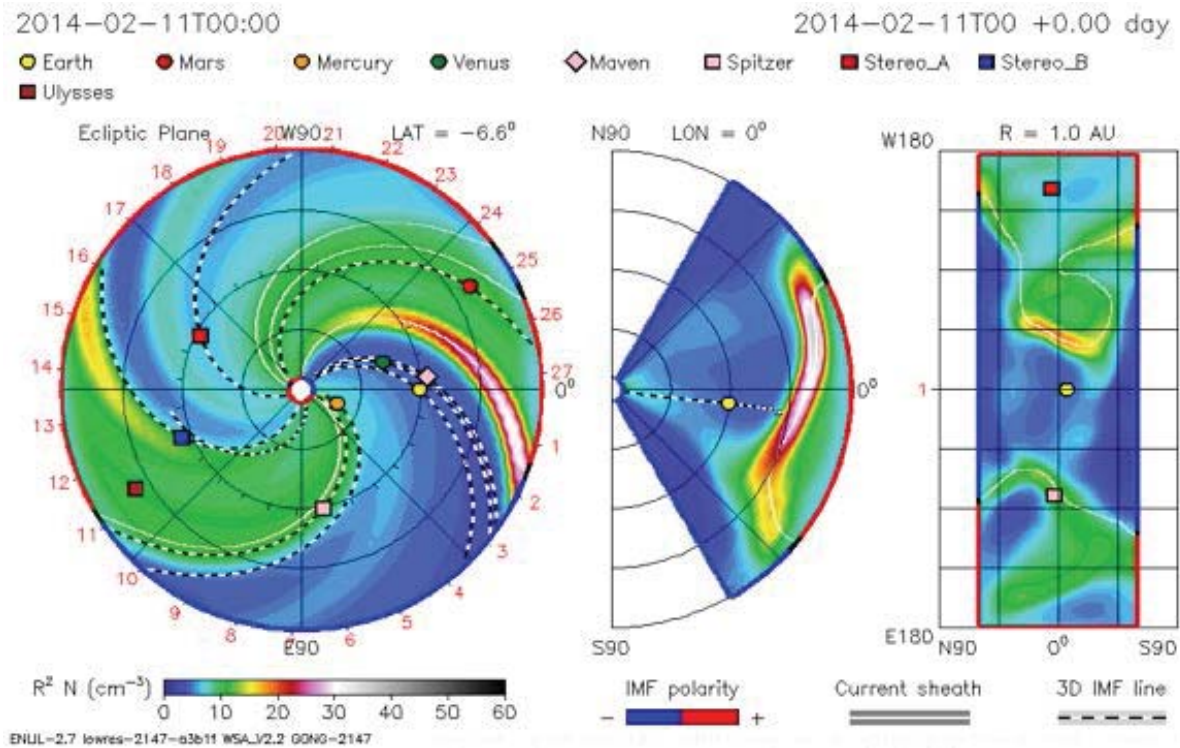
ACE Real-Time Solar Wind Pages
2014 Apr 11 1643 UTC

Auroral Map
2014 Apr 11 16:30 UTC

Solar Cycle Progression

<http://www.swpc.noaa.gov/SWN/>

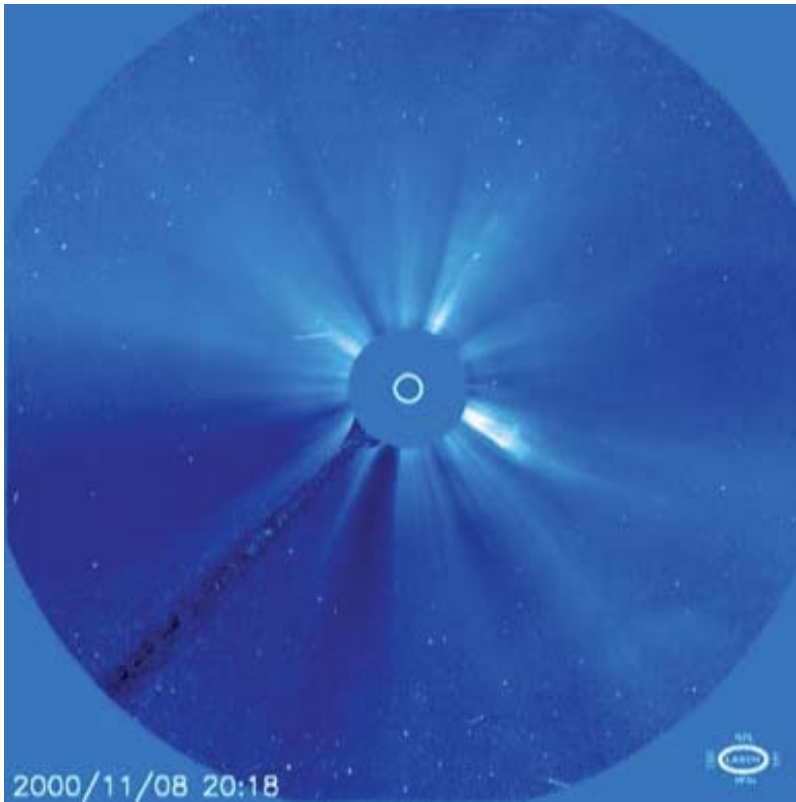
Impacts of Space Weather



NOAA / SWPC

<http://www.spaceweather.com>

Impacts of Space Weather



SOHO Large Angle and Spectrometric Coronagraph Experiment (LASCO)

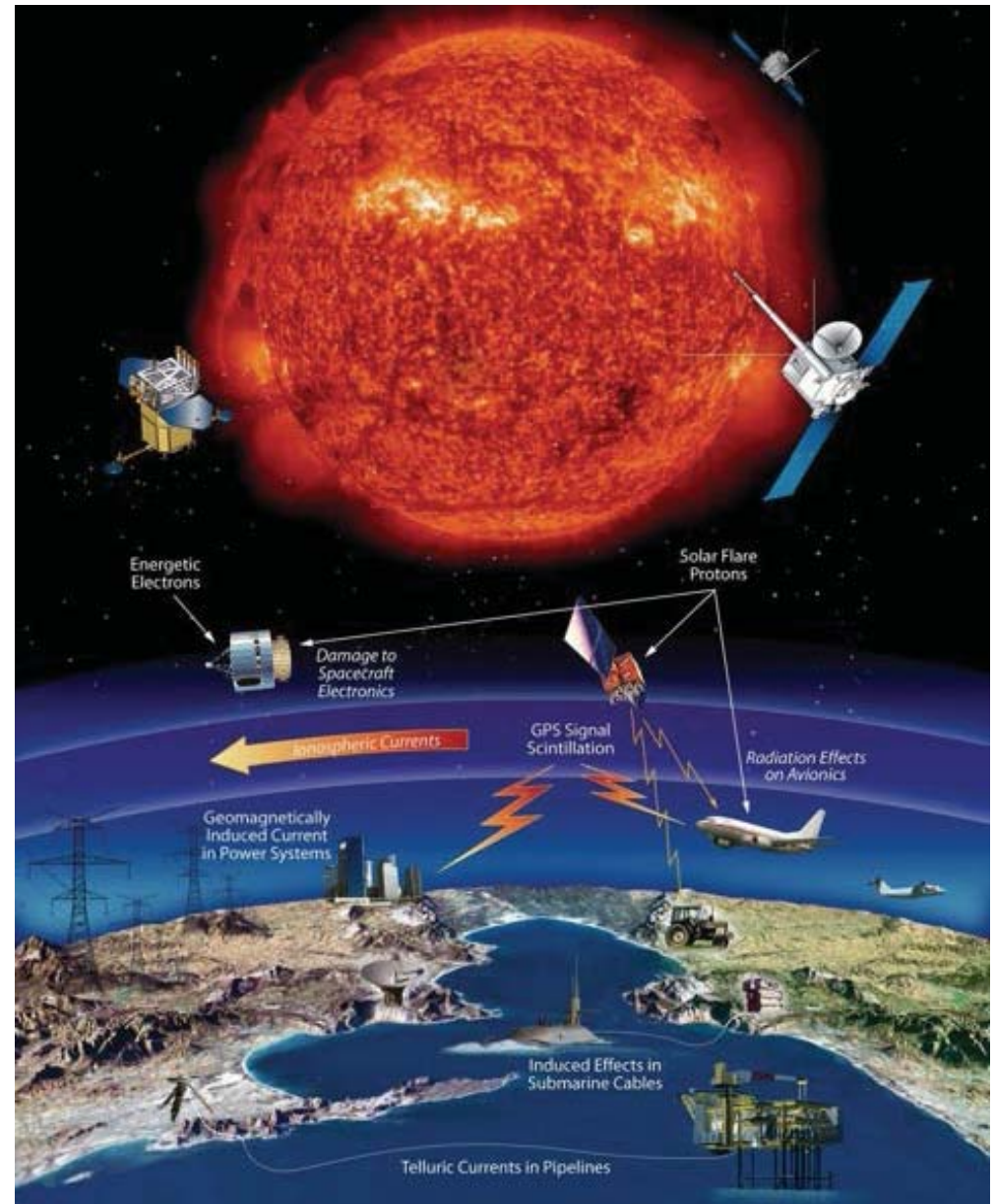
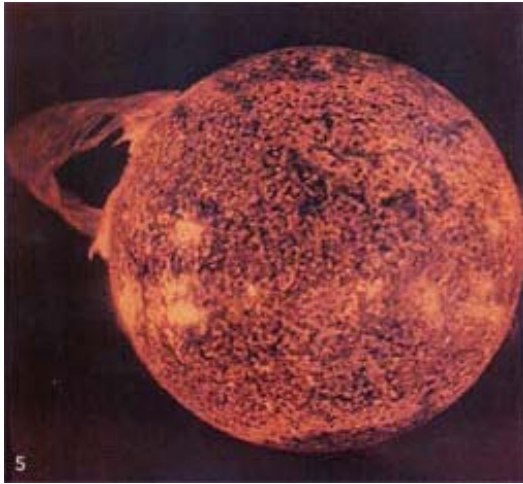
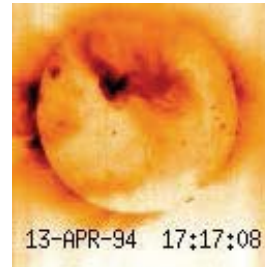


Image credit: NASA & L. Lanzerotti (NJIT)

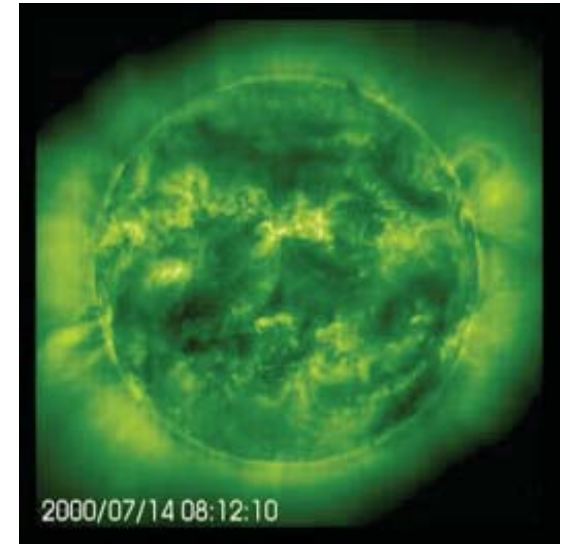
Solar Flares (A Space-Based Tour)



Skylab

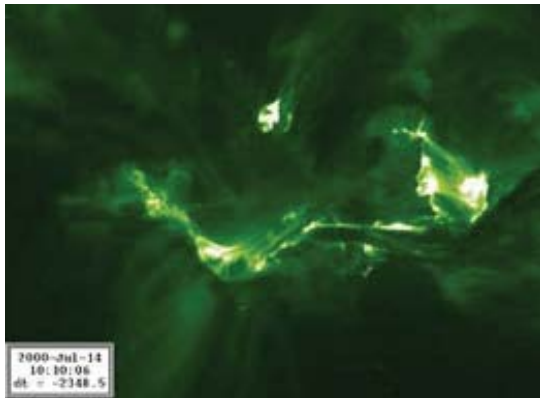


Yohkoh / SXT

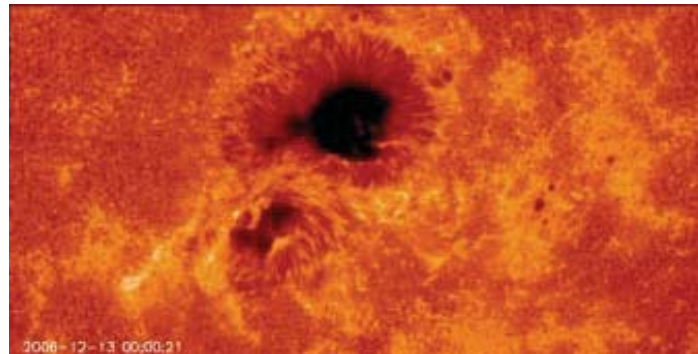


SOHO / EIT+LASCO

TRACE



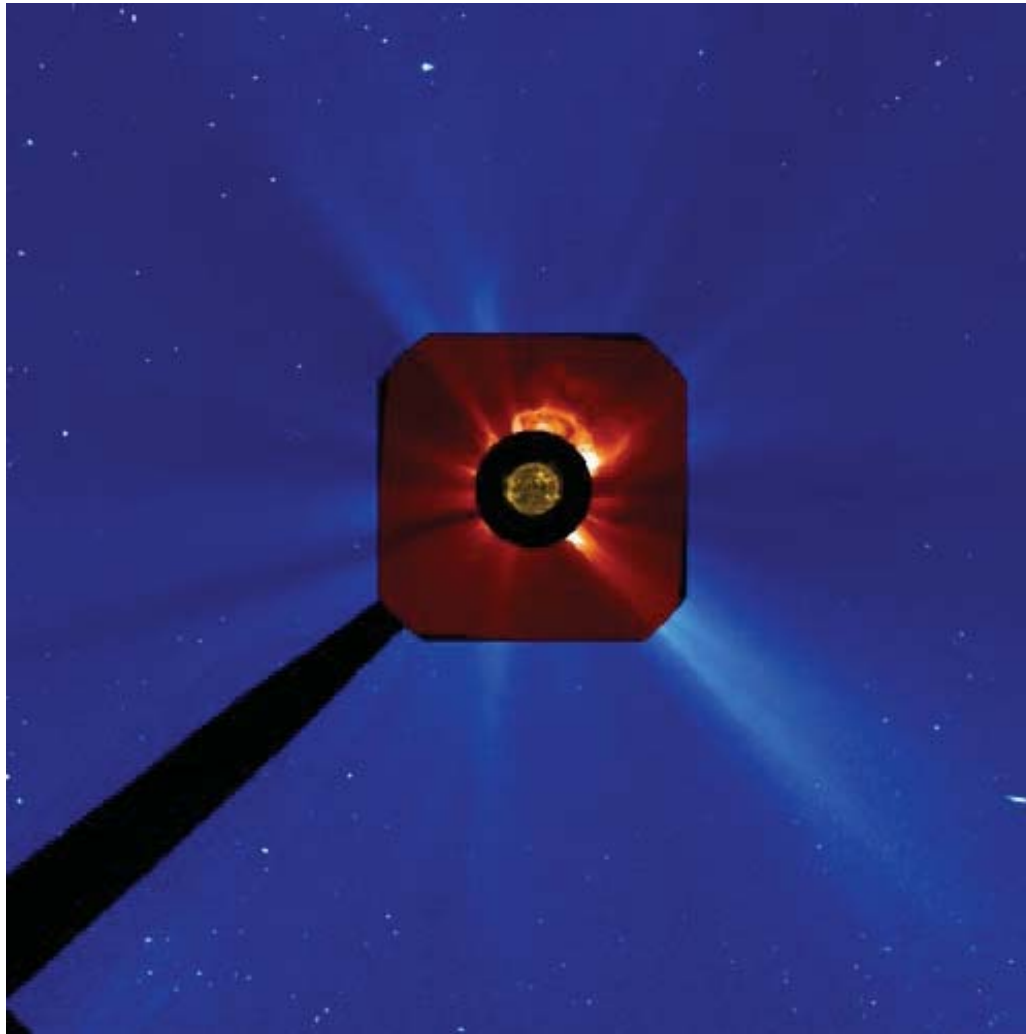
Hinode / SOT



Hinode / XRT

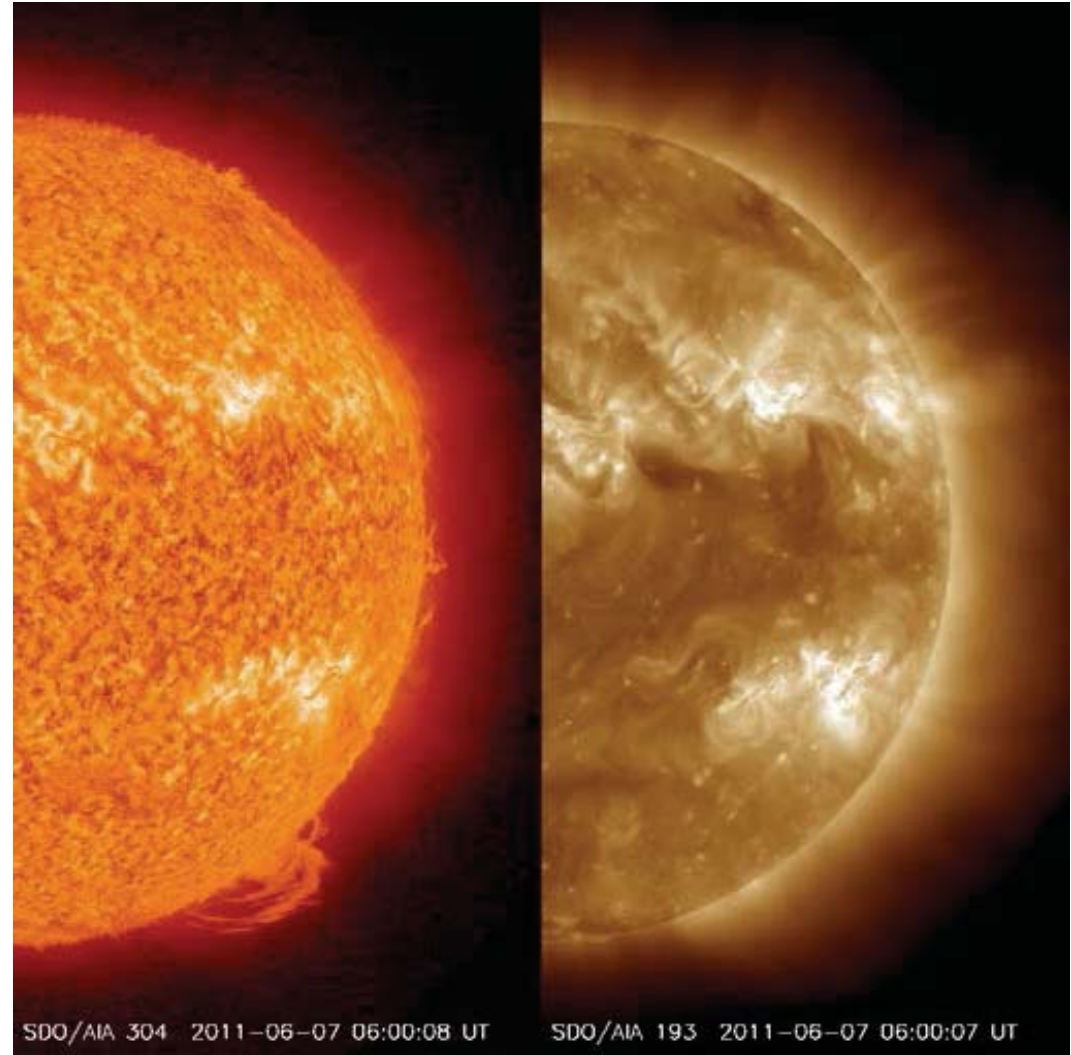
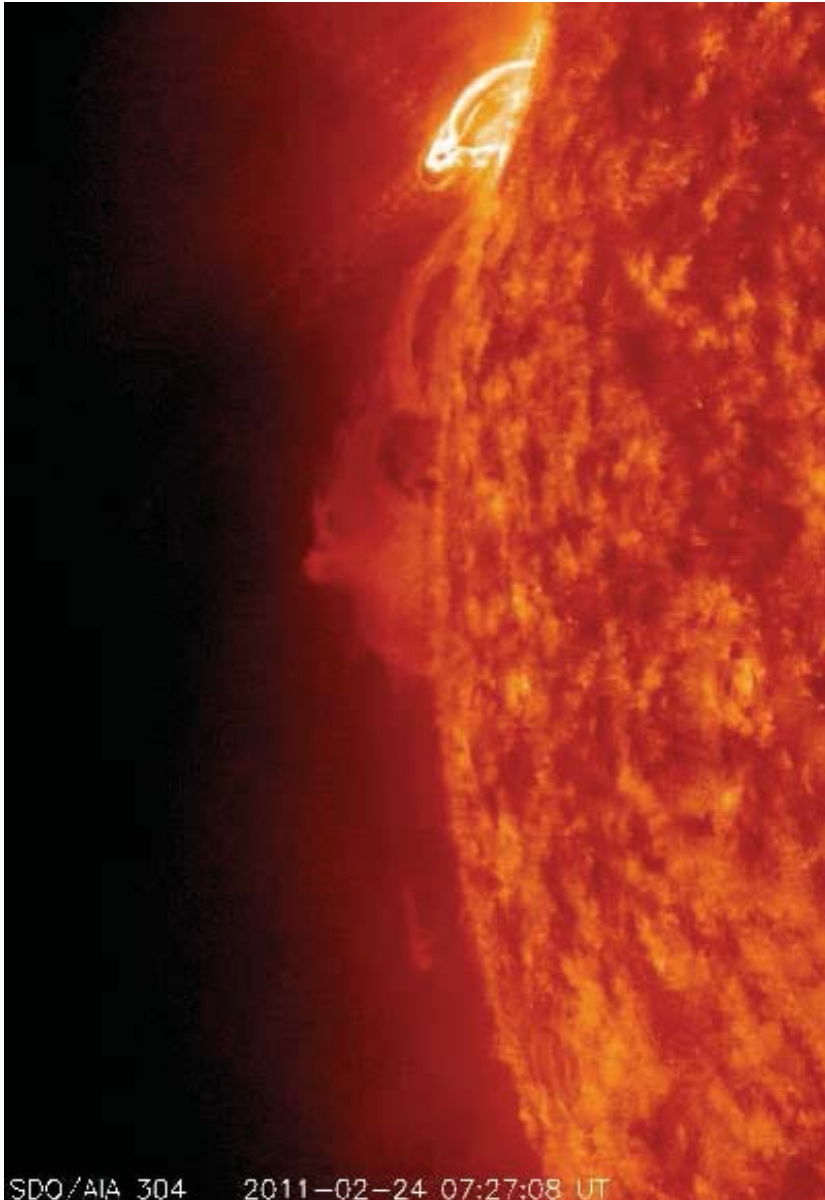


Solar Flares (A Space-Based Tour)



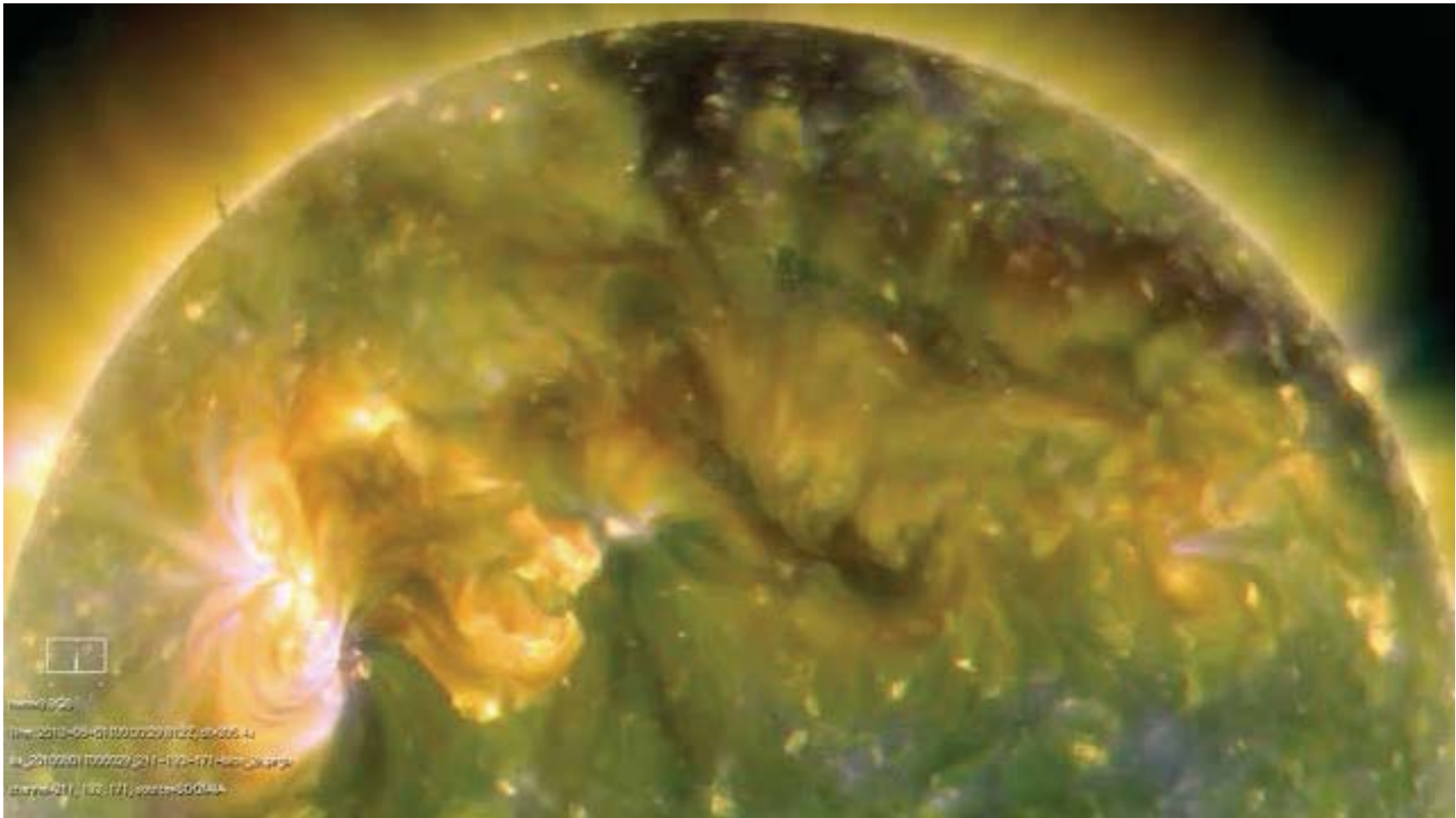
SDO / AIA + SOHO / LASCO

Solar Flares (A Space-Based Tour)



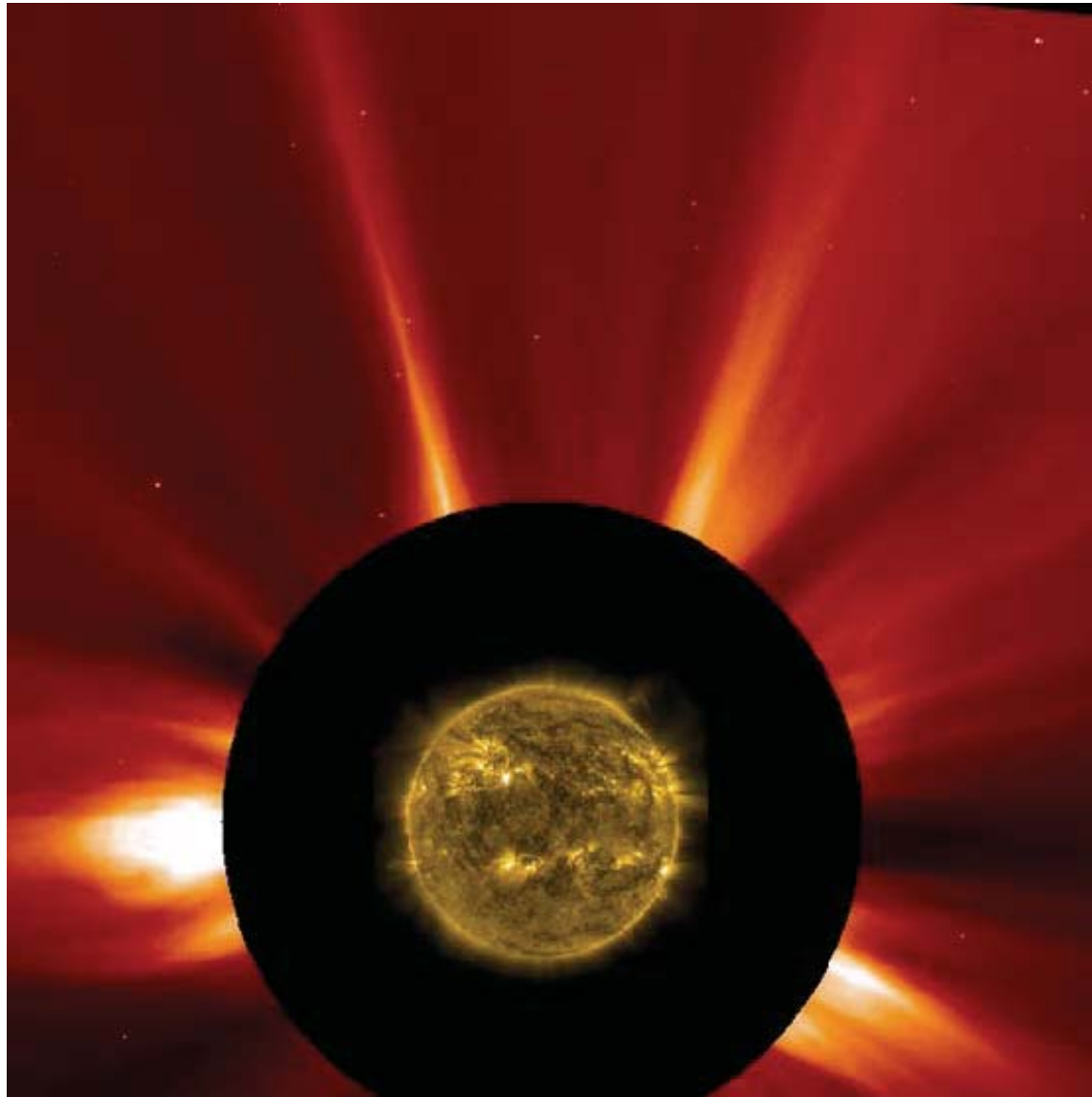
SDO / AIA

Solar Flares (A Space-Based Tour)



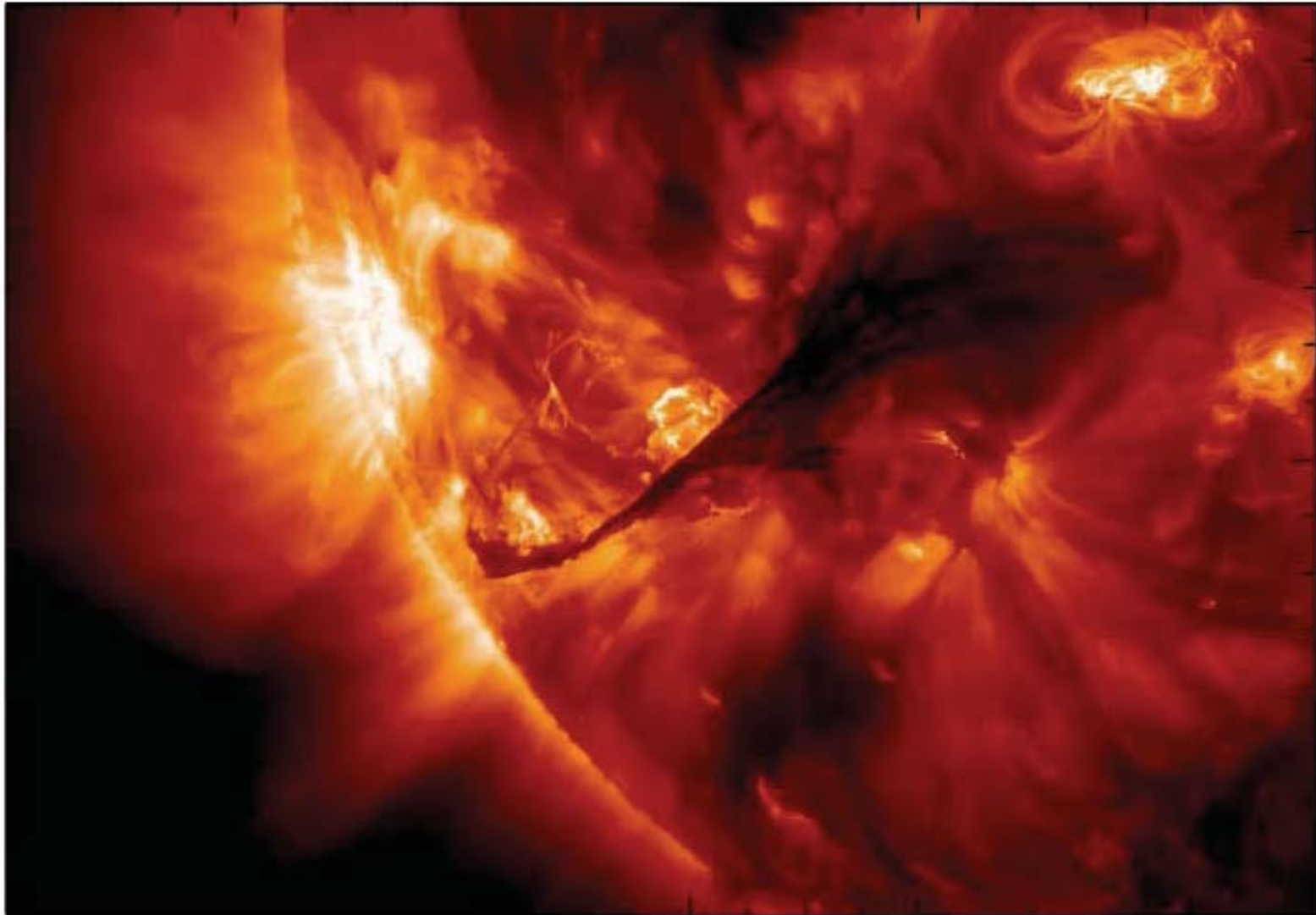
SDO / AIA

Solar Flares (A Space-Based Tour)



SDO / AIA + SOHO / LASCO

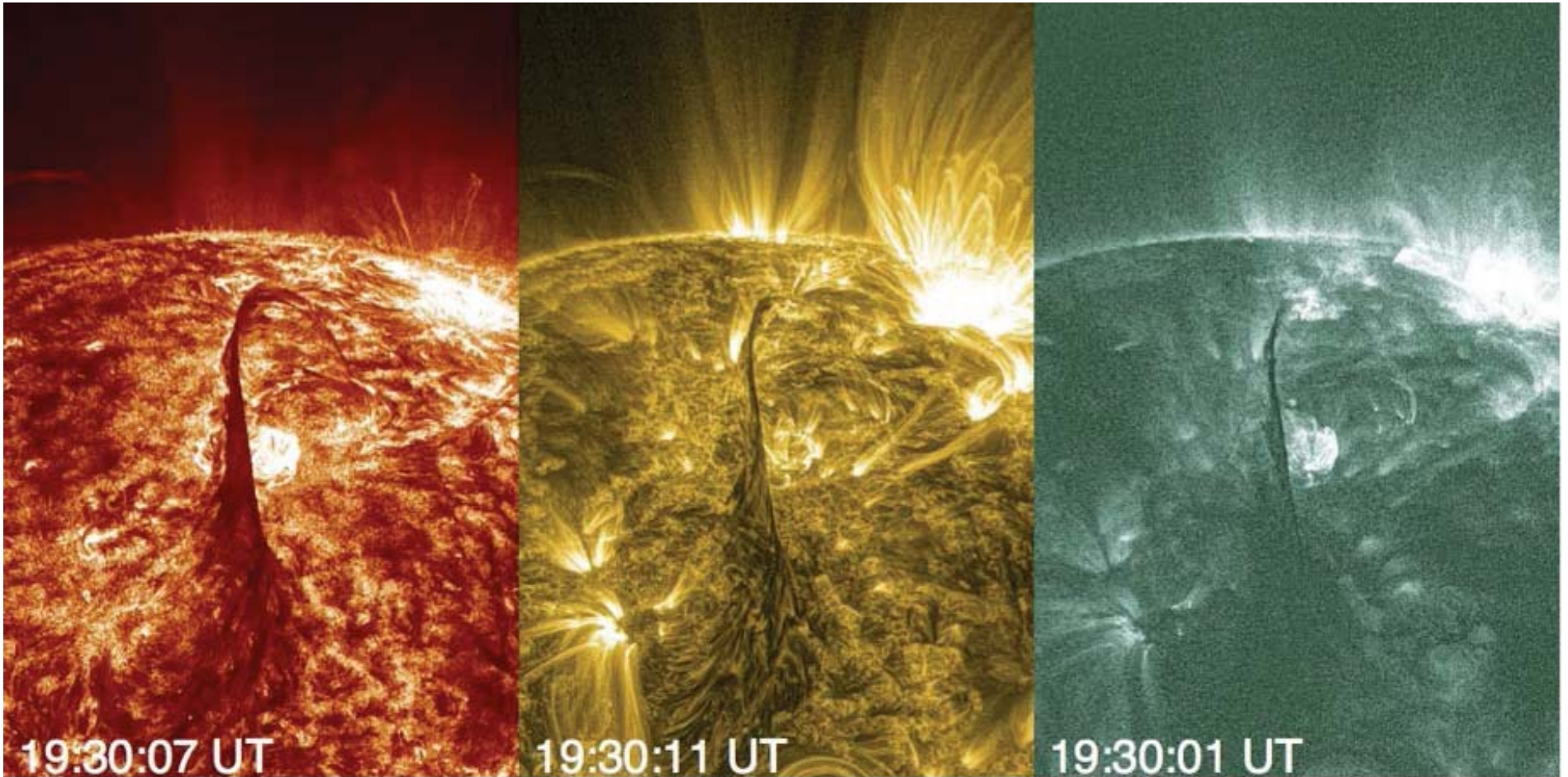
Solar Flares (A Space-Based Tour)



SDO / AIA + Hinode / EIS

Solar Flares (A Space-Based Tour)

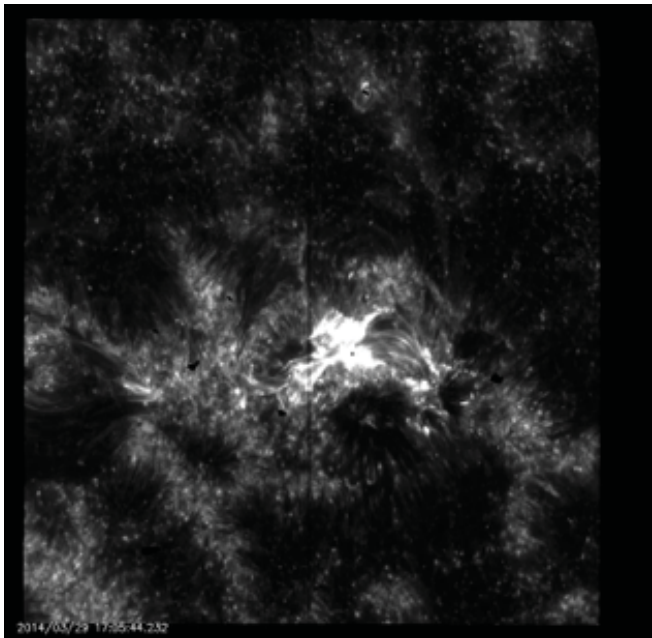
Same flare as previous slide but in 3 different AIA channels and enhanced for contrast.



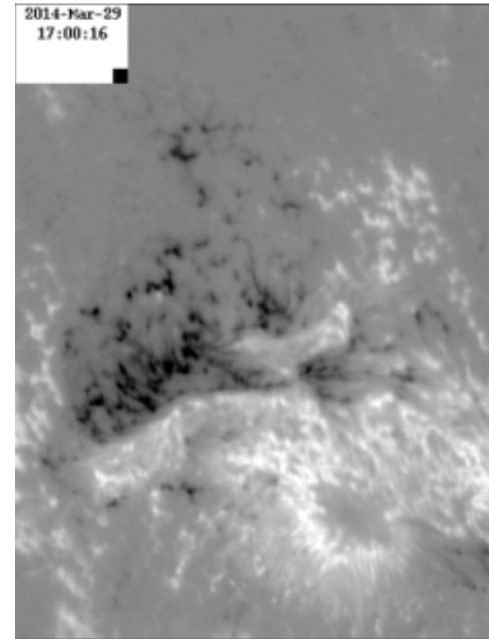
SDO / AIA

Solar Flares (A Space-Based Tour)

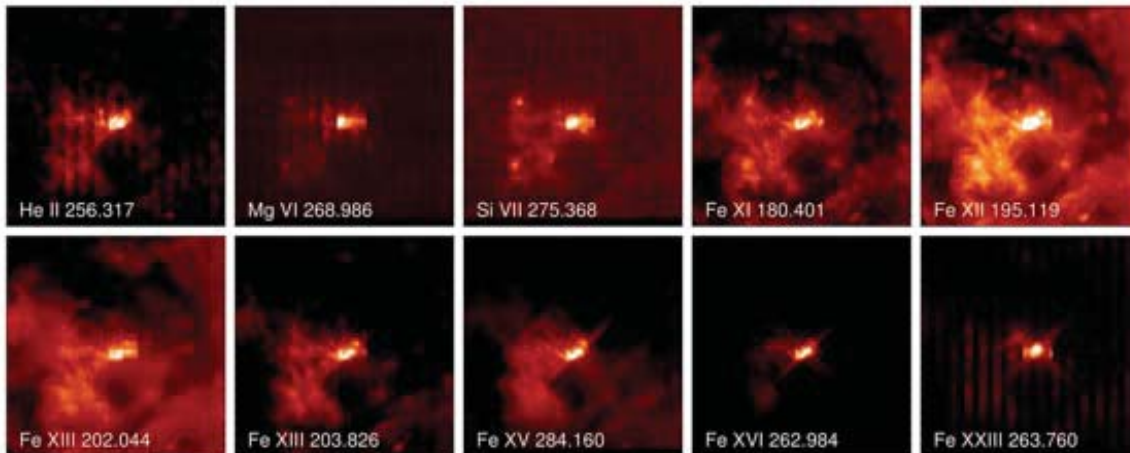
IRIS



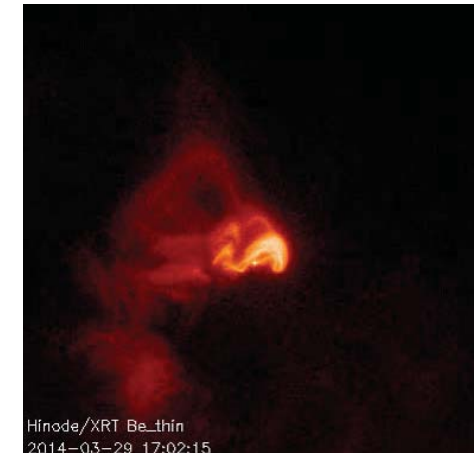
Hinode / SOT [Magnetogram]



Hinode / EIS



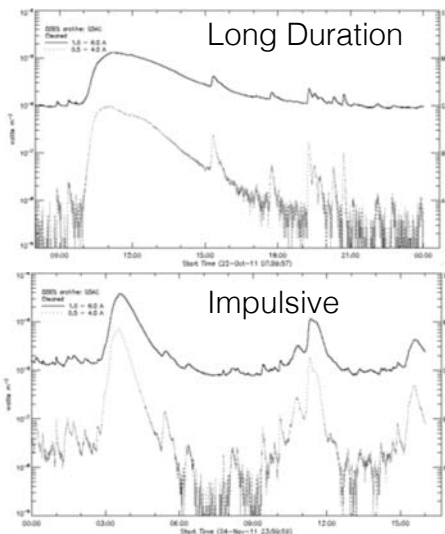
Hinode / XRT



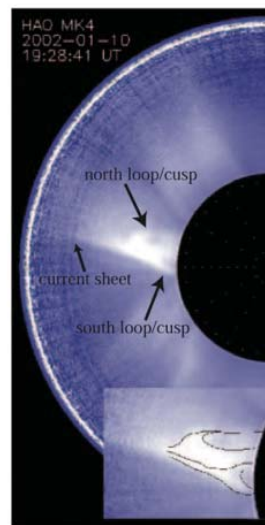
Investigating Energy Release

Focus on Long Duration Events

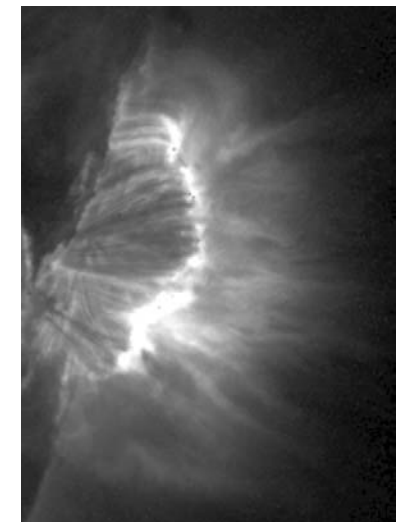
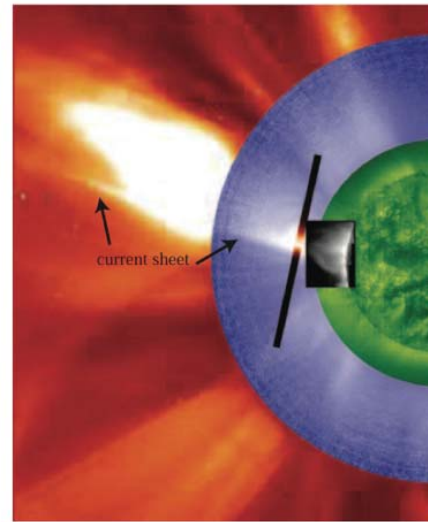
- Energy released for many hours
- Associated with Coronal Mass Ejections (CMEs)
- Development of current sheets and supra-arcade fans



Example GOES lightcurves



Ko et al. 2003



Savage & McKenzie 2011

Investigating Energy Release

Standard **2-D** Flare Model

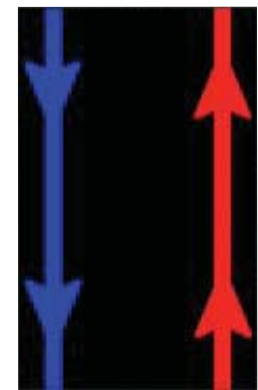
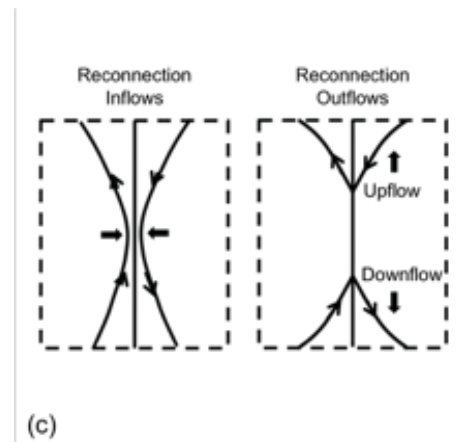
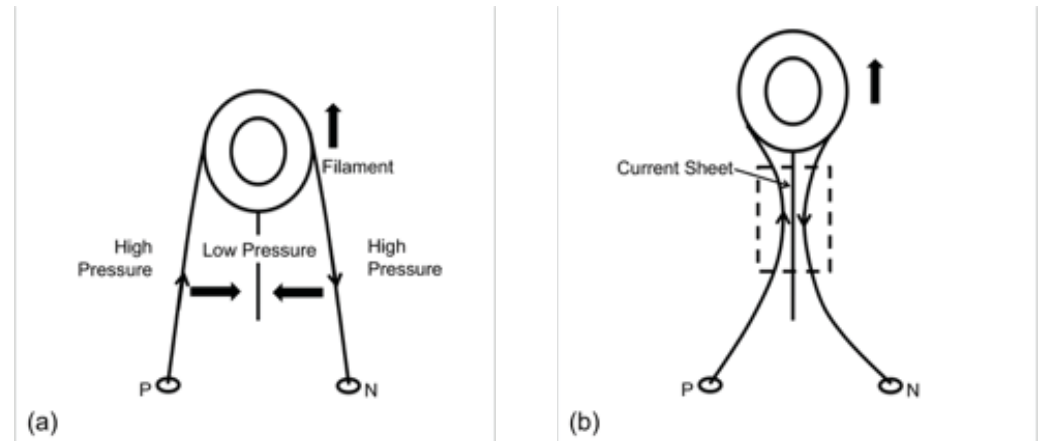
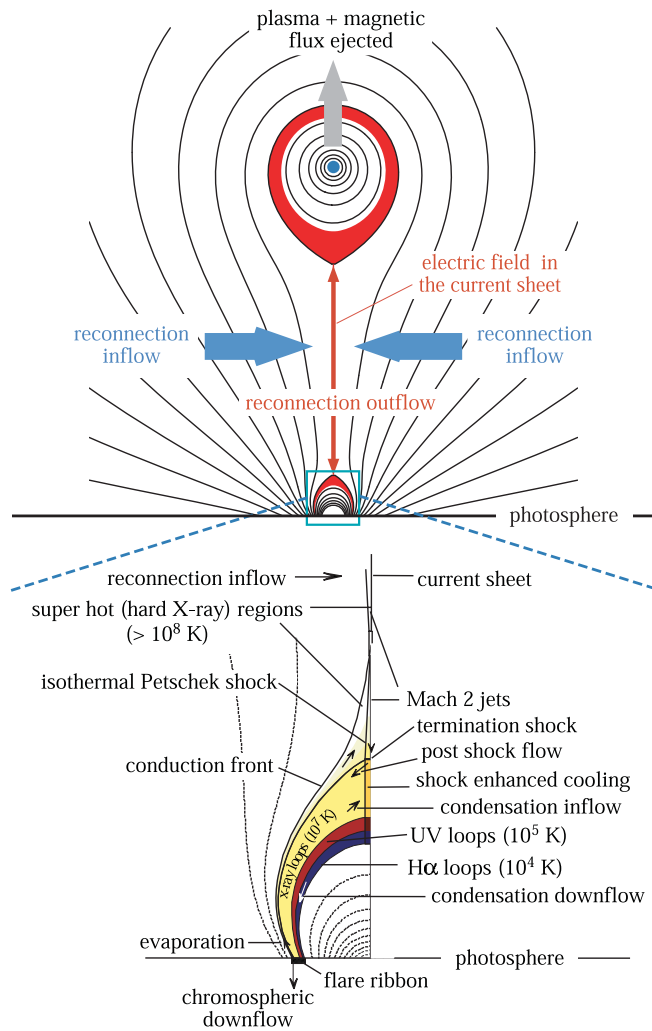


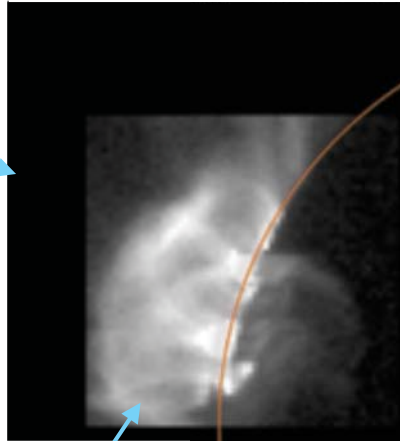
Image Credit: ESA

Investigating Energy Release

Early observations of Supra-Arcade Downflows (SADs) & Downflowing Loops (SADLs)

Yohkoh / SXT 1999 Jan 20

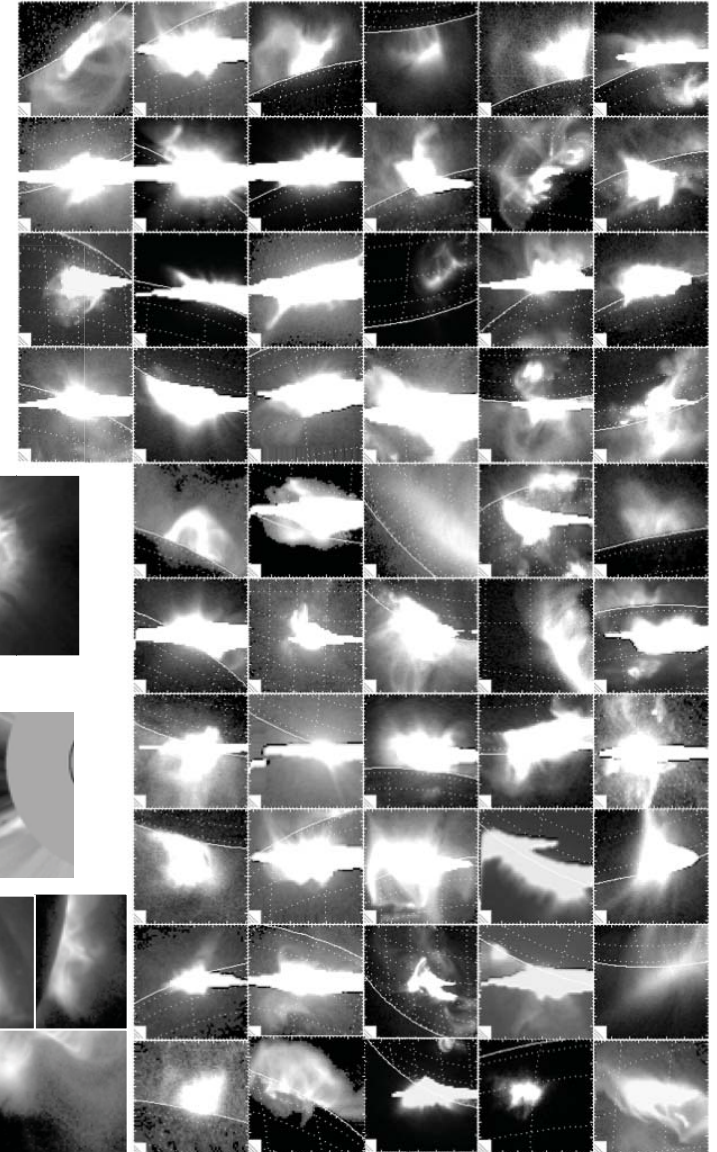
Downflowing
Voids Above
Arcade



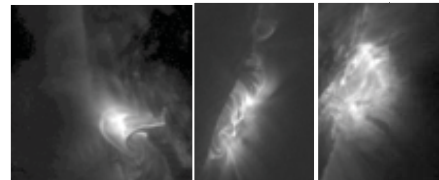
Post-eruption
Arcade
(Saturated)

Solar Limb

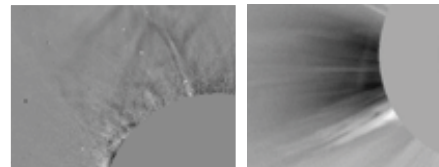
Yohkoh / SXT



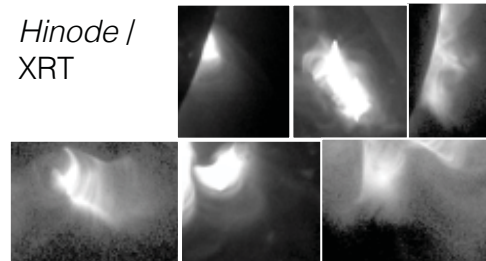
TRACE



SOHO / LASCO



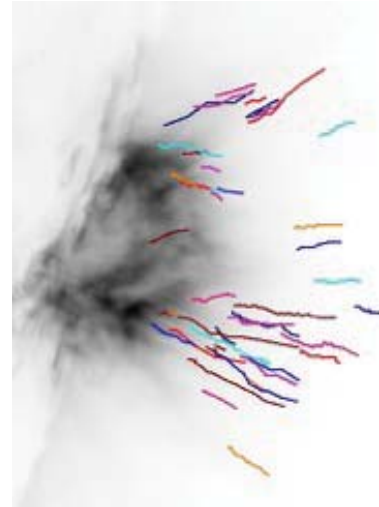
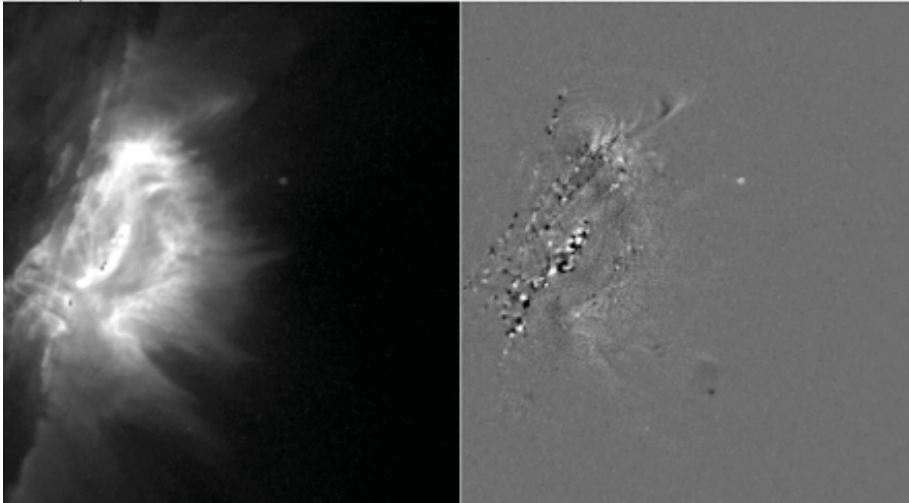
Hinode / XRT



Investigating Energy Release

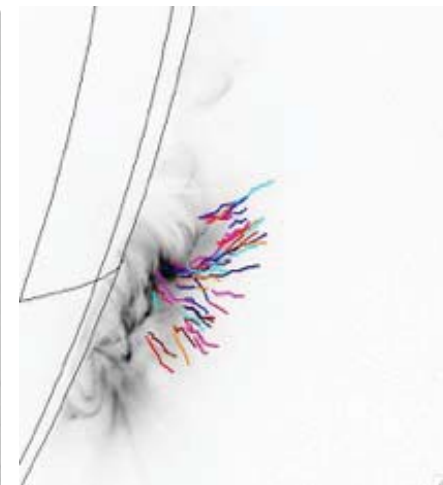
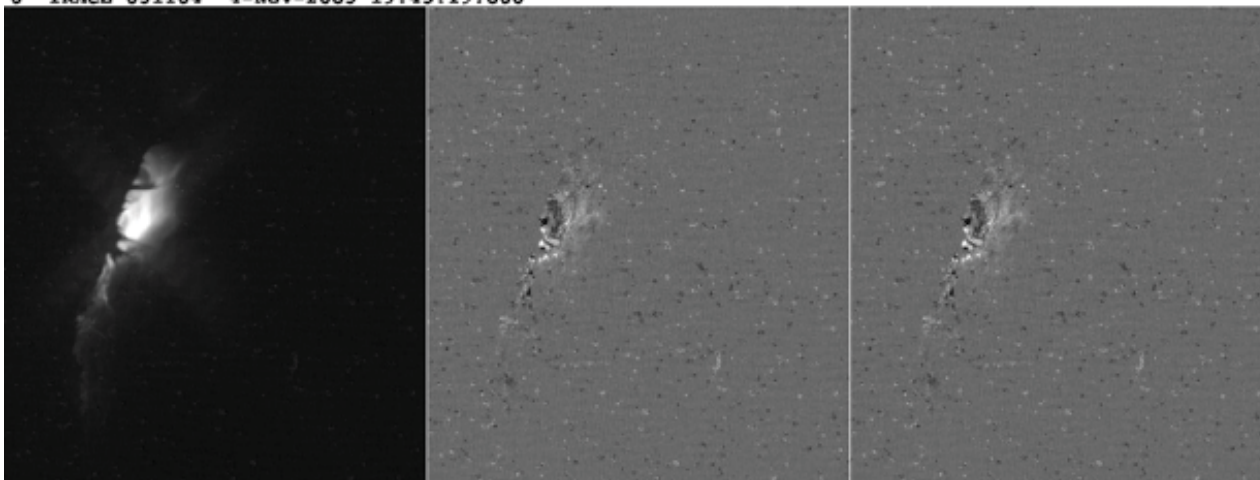
TRACE 193 A, X-flare, 2002 Apr 21

B4 21-Apr-2002 01:33:32.000



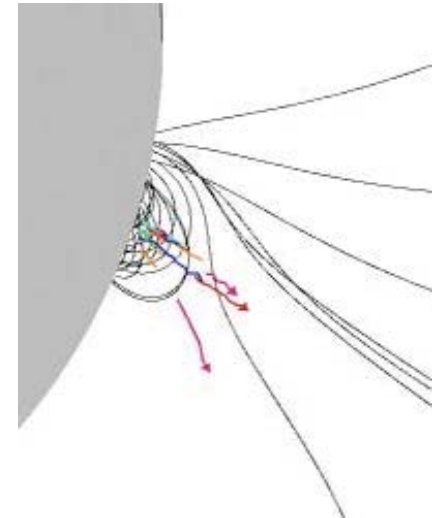
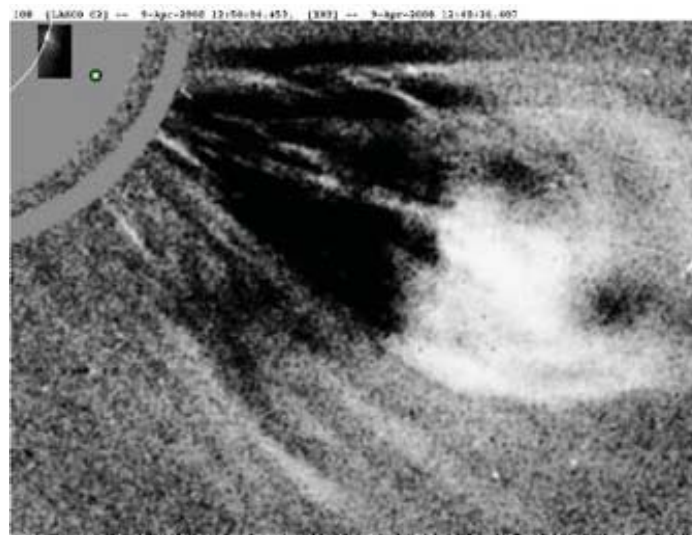
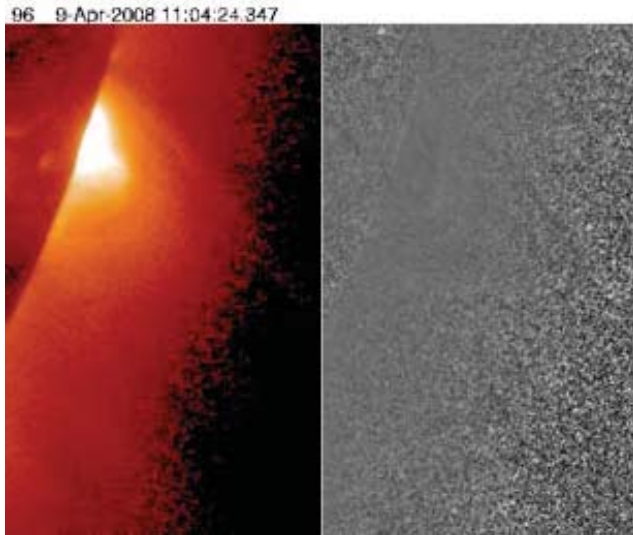
TRACE 193 A, X-flare, 2003 Nov 4

0 TRACE 031104 4-Nov-2003 19:45:19.000



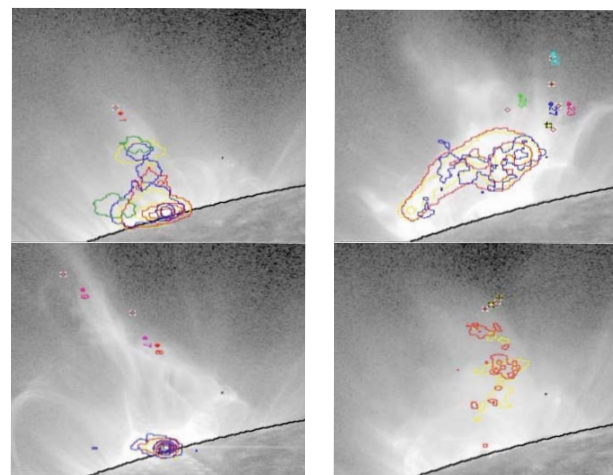
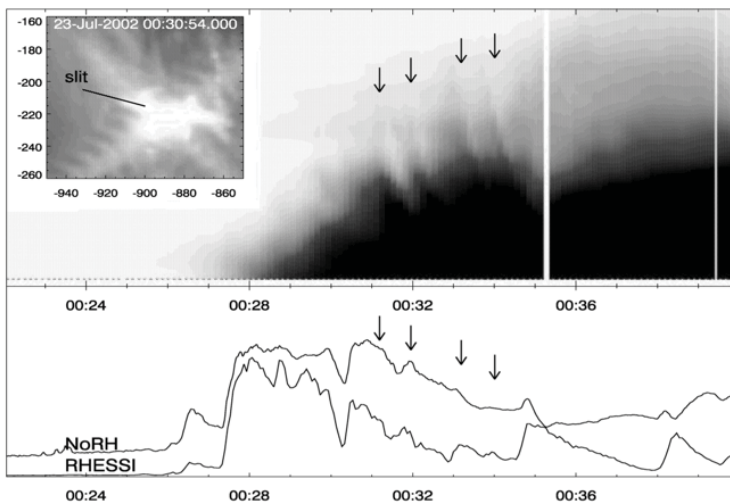
Investigating Energy Release

Hinode / XRT, 2008 Apr 9



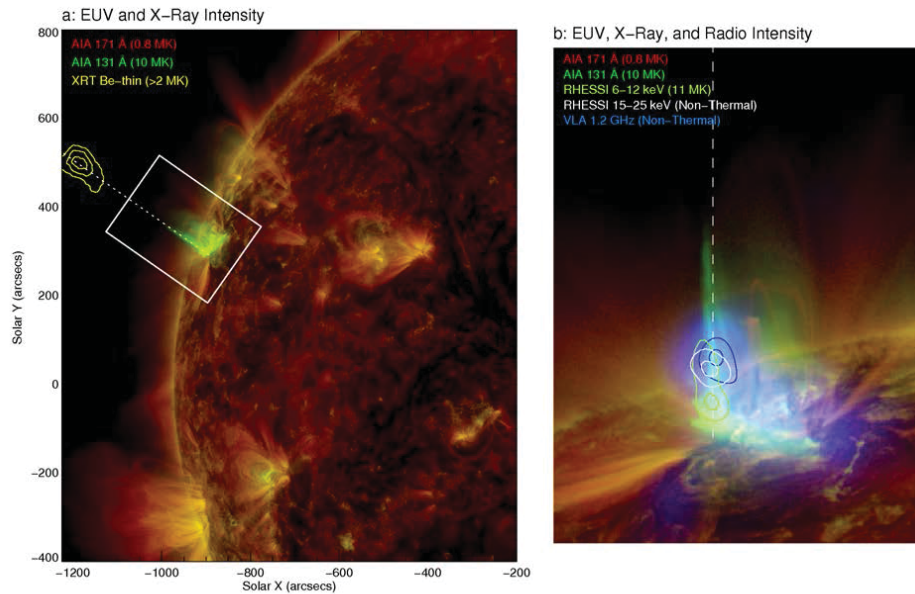
TRACE + RHESSI + NoRH radio (lightcurve),
2002 Jul 23

SDO / AIA + RHESSI (contours), 2010 Nov 3



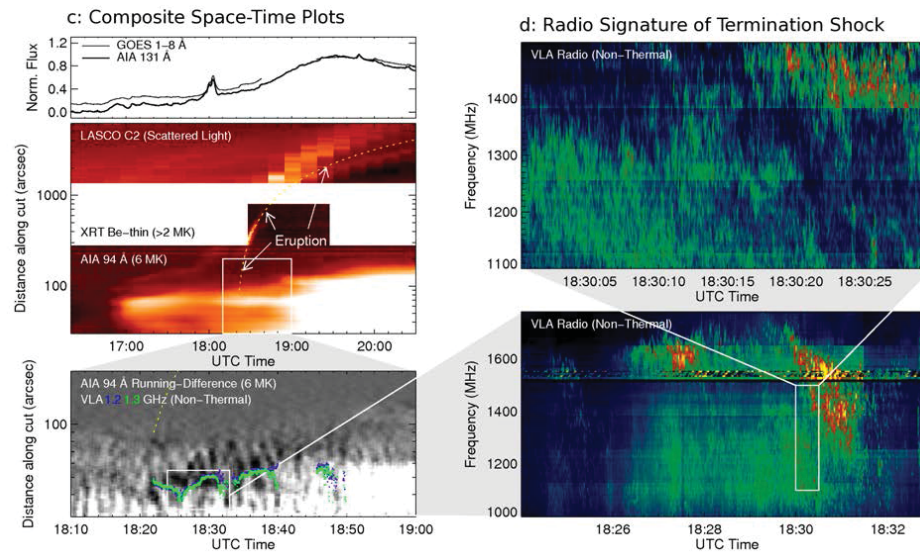
And now RADIO!...

Investigating Energy Release



Jansky Very Large Array (VLA) observations:

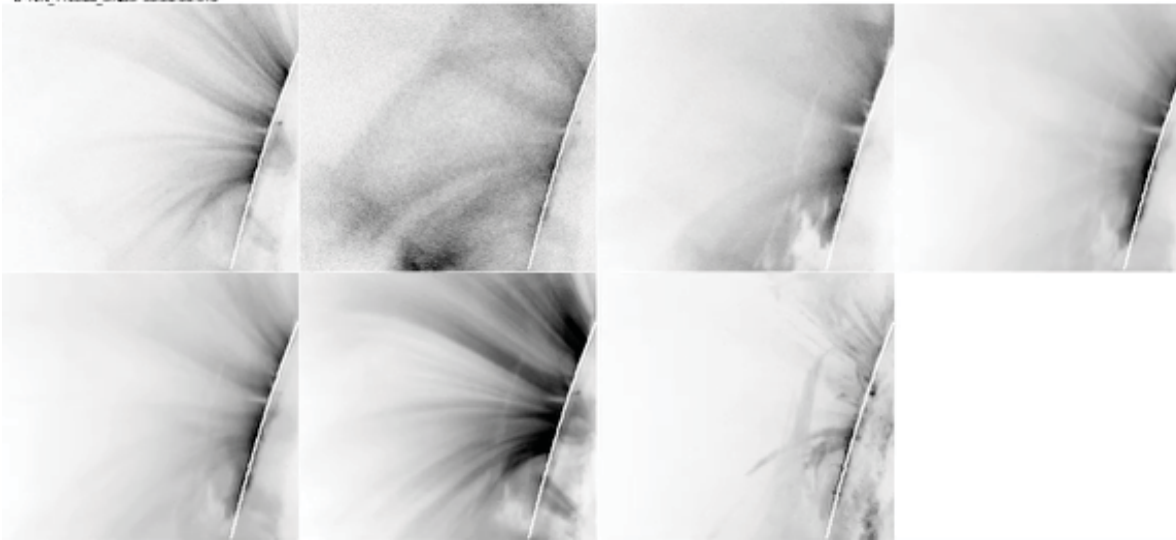
D) Possible termination shocks at the arcade looptops.



Investigating Energy Release

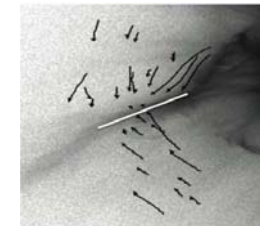
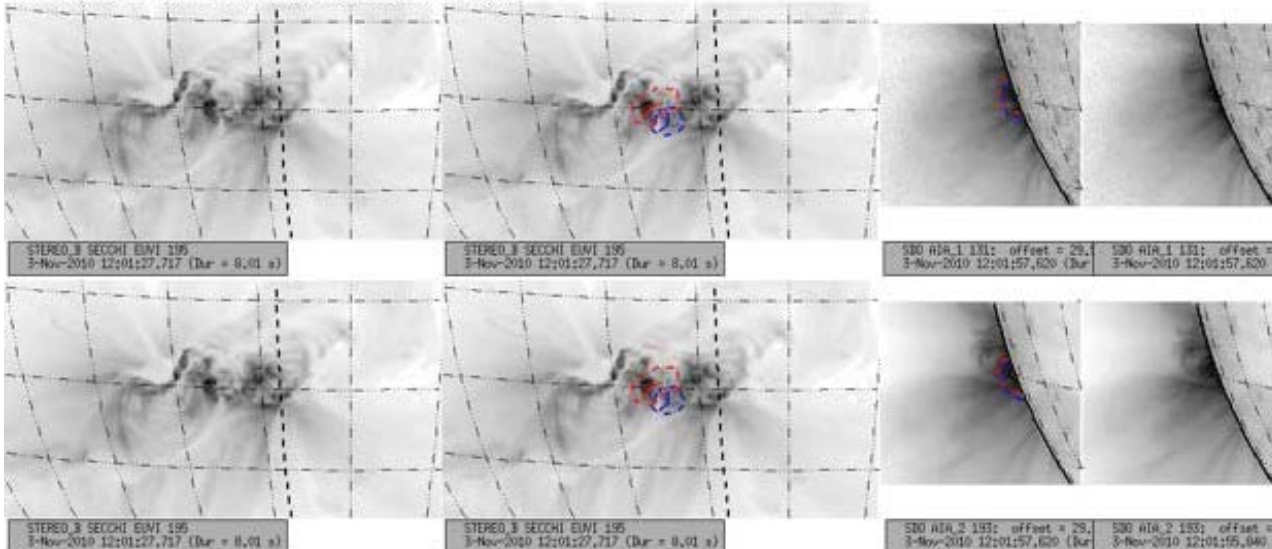
SDO / AIA, 2011 May 9

AIA_110009_SADS_20200940

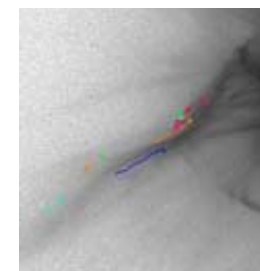


SDO / AIA + STEREO / SECCHI, 2010 Nov 3

STEREO-B 195 | AIA 131 | AIA 131 20101103 Reference time: 12:01:27,71



Inflows
Composite

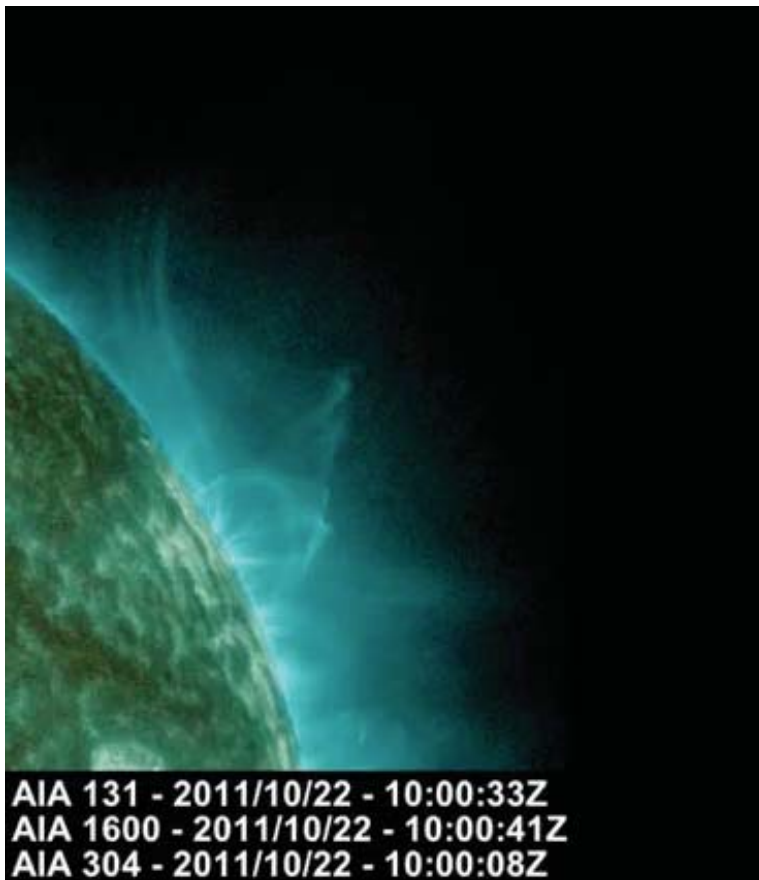


Outflows
131 A

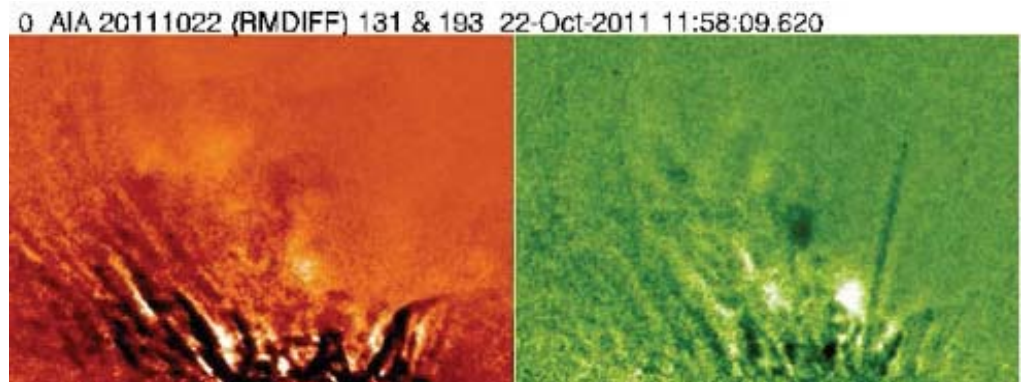
Investigating Energy Release

Explanation for SADs & SADLs converging ...

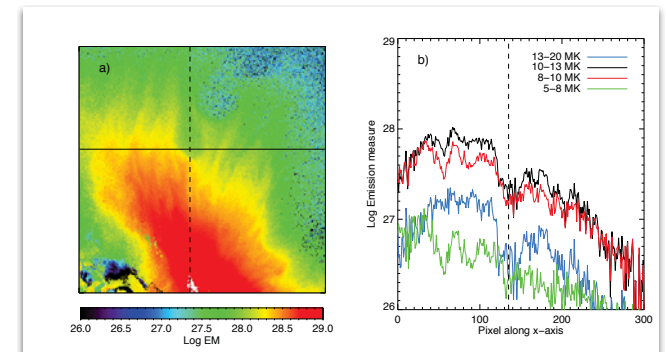
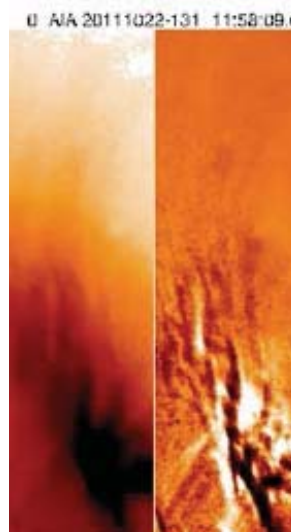
SDO / AIA, 2011 Oct 22



Movie Credit: D. E. McKenzie, Mont. State Univ



Bright thin loops retracting below voids.

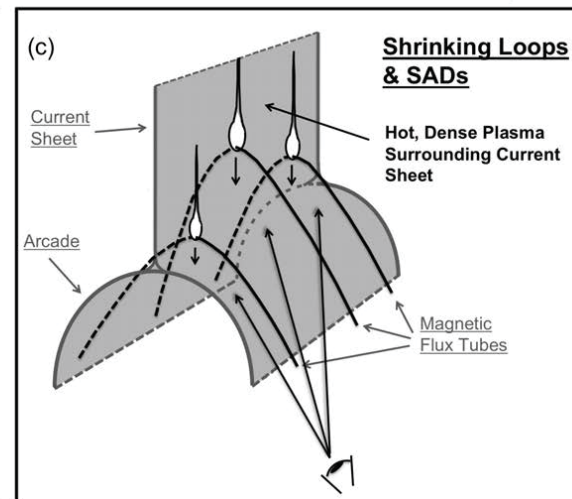
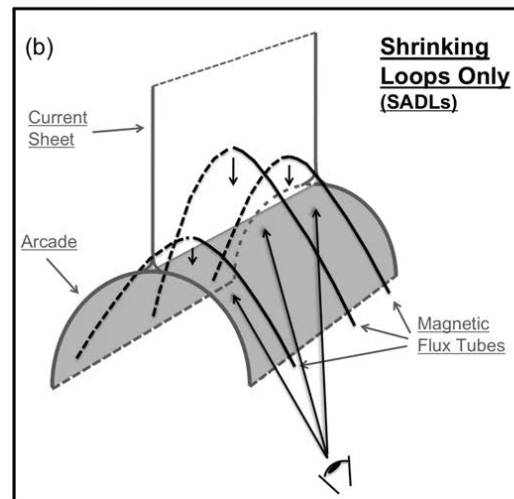
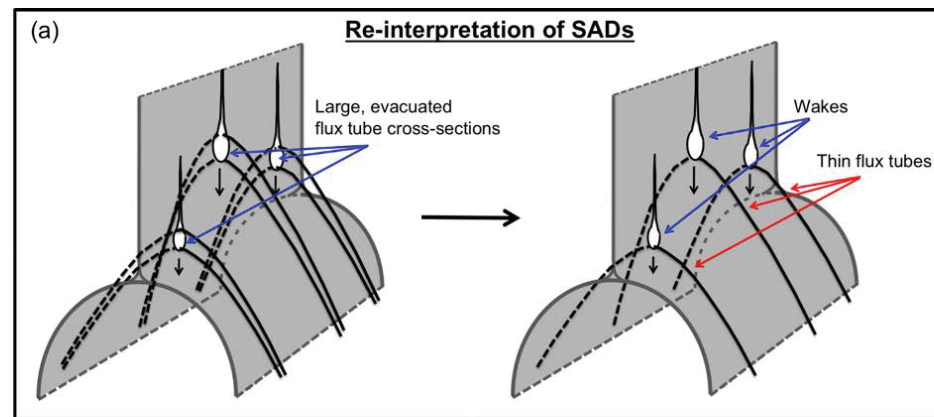


SADs cooler than fan (and much less dense)

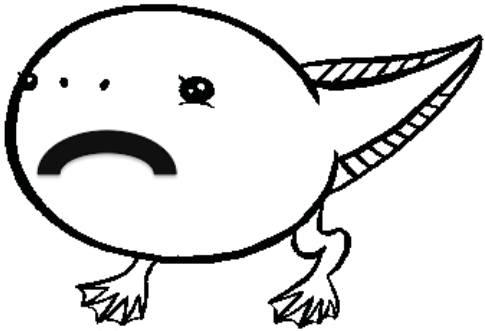
Investigating Energy Release

Explanation for SADs & SADLs converging ...

- > Loops outflows of patchy, bursty magnetic reconnection?!
- > Voids rarefaction regions behind retracting loops?

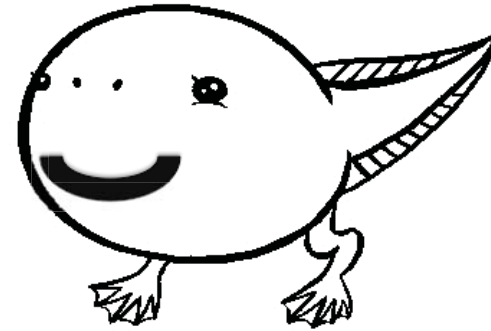


Sadpoles



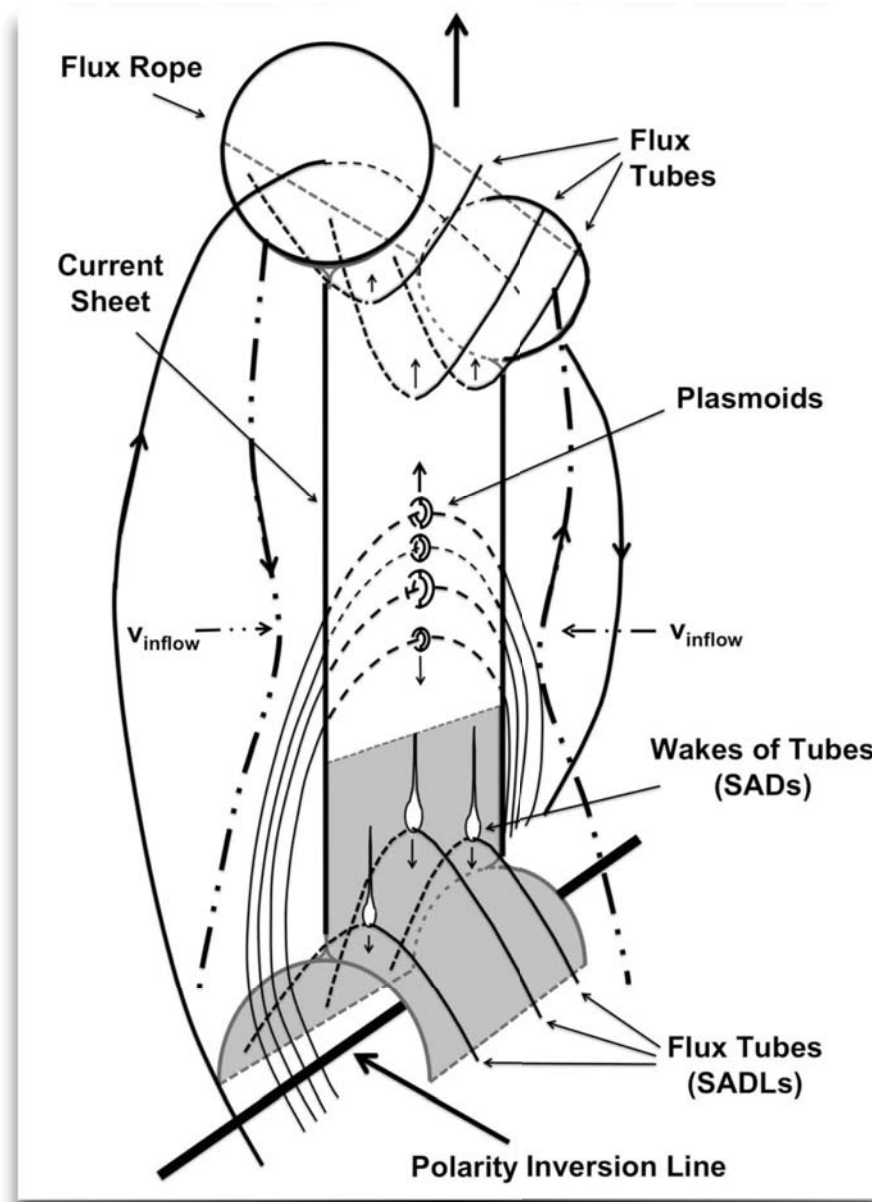
?

Happis



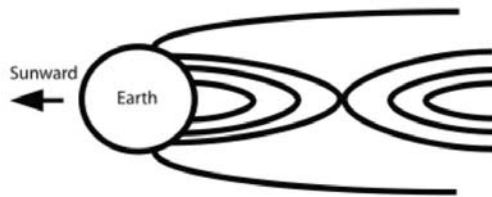
HAPPIs: High-Altitude Propagating Pressure Imbalances?

A Simplified **3-D** Solar Flare Model

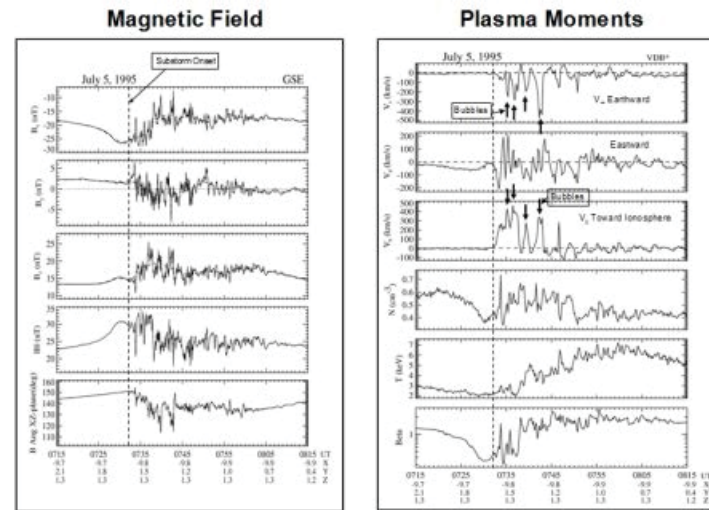


Observing Magnetic Reconnection

Solar flares comparable to Magnetotail substorms

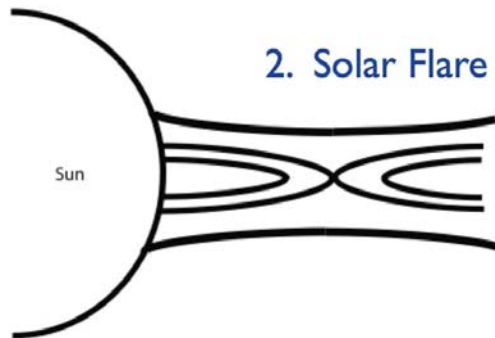


1. Magnetotail Substorm

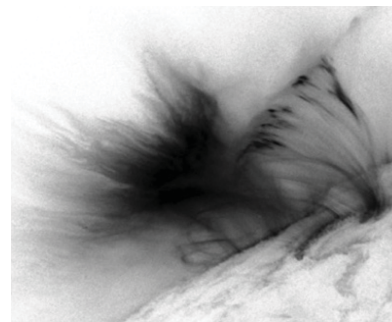


Magnetotail:

In Situ Measurements



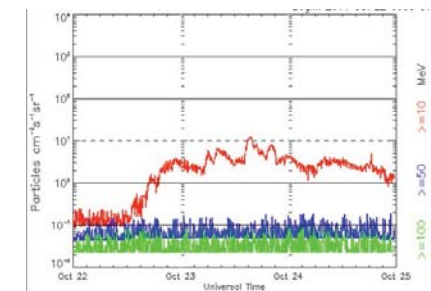
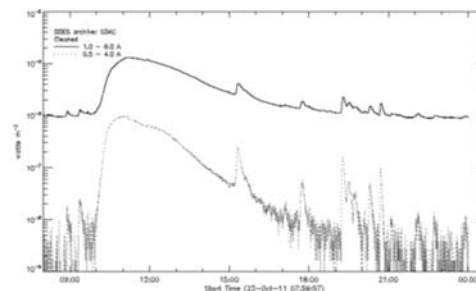
2. Solar Flare



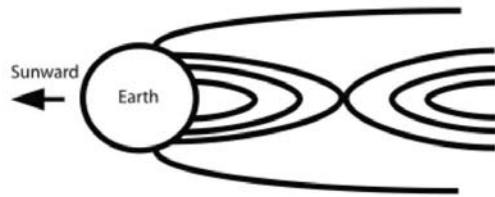
Solar:

Global Context

Note: Very different scales and plasma regimes. **BUT COMPARABLE ALFVEN SPEEDS!**



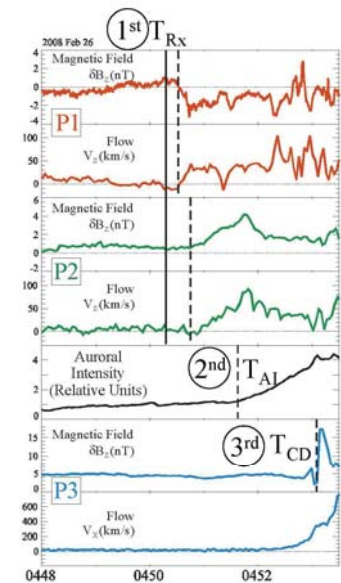
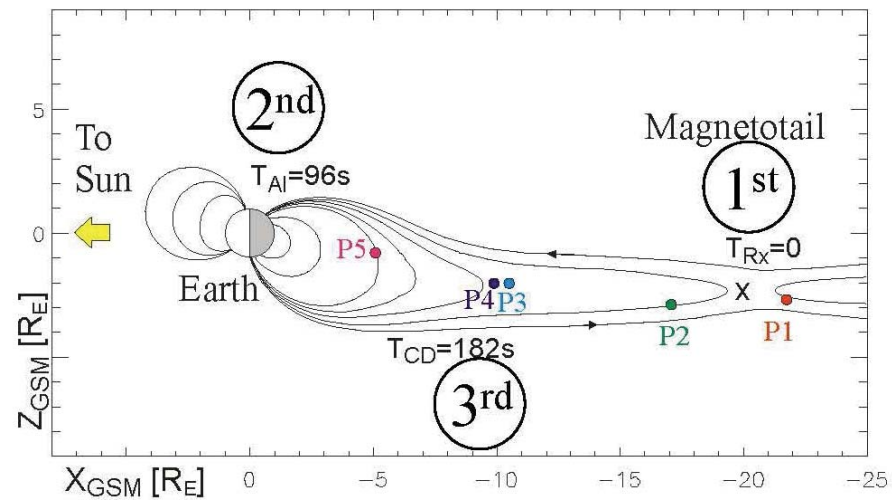
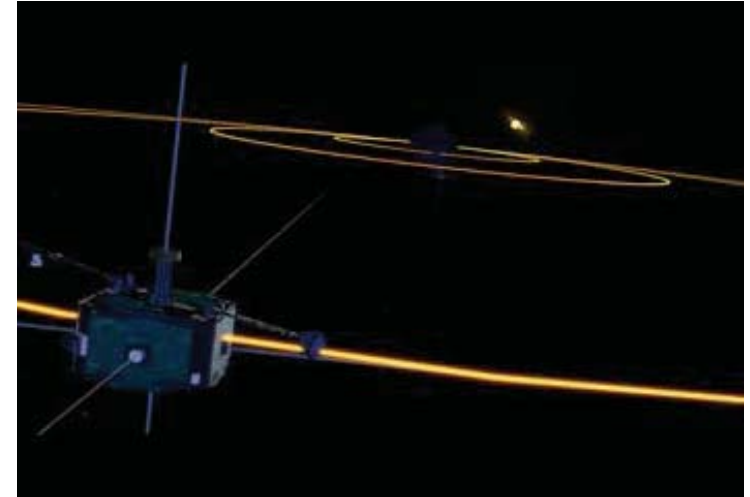
Observing Magnetic Reconnection



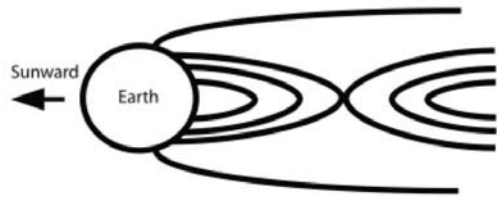
I. Magnetotail Substorm

THEMIS:

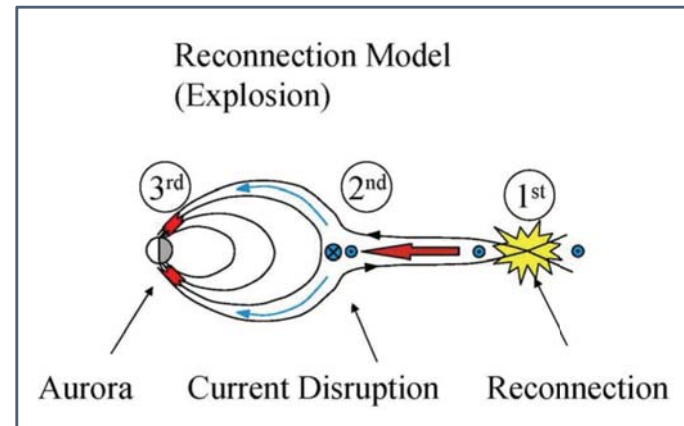
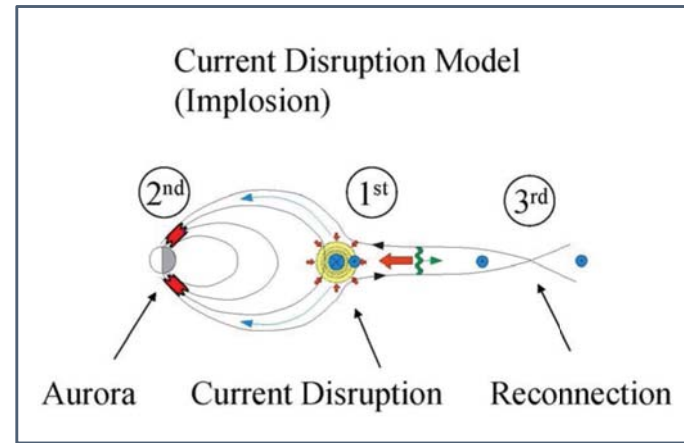
Constellation of satellites
Measure B, E, density, etc.



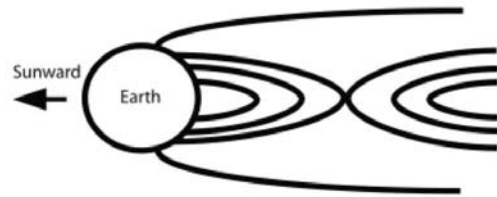
Observing Magnetic Reconnection



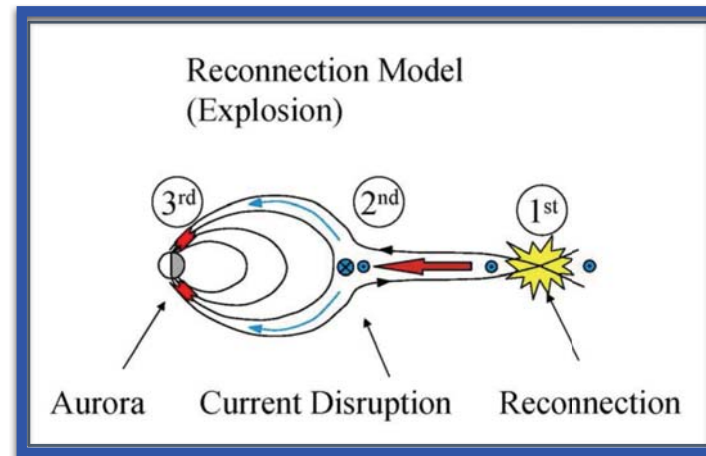
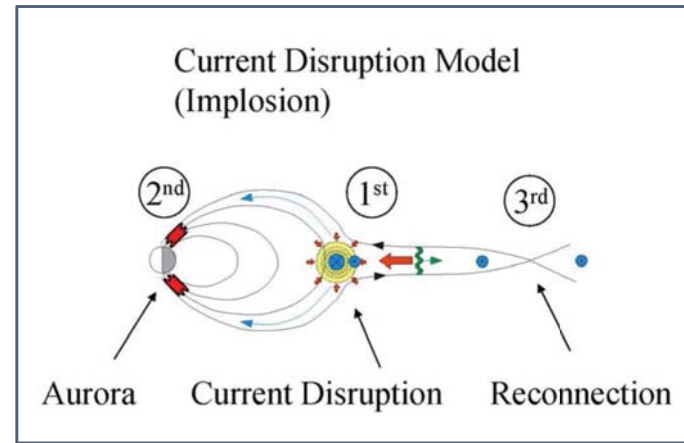
I. Magnetotail Substorm



Observing Magnetic Reconnection



I. Magnetotail Substorm



Akin to Solar Flares!!

SDO 2nd Year Highlights

