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Produced in High Energy Proton-Proton Collisions by
The Two-Temperature Statistical Model

by

J. R. Wayland

Technical Report No. 844

June 1968

GPO PRICE \$ _____

CSFTI PRICE(S) \$ _____

Hard copy (HC) _____

Microfiche (MF) _____

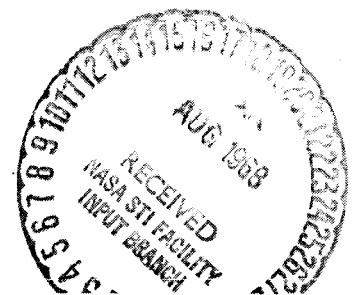
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N 68-33548	
(ACCESSION NUMBER)	(THRU)
19	1
(PAGES)	(CODE)
CK-96317	24
(NASA CR OR TMX OR AD NUMBER)	(CATEGORY)



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**A Program to Calculate Particle Spectra
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*This research was supported by the National Aeronautics and Space
Administration under Grant (NsG-695.)

NGL-21-002-033

I. INTRODUCTION

In the past few years, the lack of success of older statistical model led to new models that rely more on phenomenological observations. One of the most successful models is the two-temperature model of Wayland and Bowen [1]. The analytic form of this model allows easy extension of the predictions to cosmic ray energies.

We will give a brief outline of the two-temperature model in Section II. In Section III, we consider the application of the model. The parameters for the π^\pm , K^\pm , p and \bar{p} spectra are given. Copies of the programs can be obtained from the author.

II. THE TWO-TEMPERATURE MODEL

The main assumption of the two-temperature model is that there are two characteristic temperatures associated with each type of particle produced; one is associated with the transverse momentum distribution (T_0), the other with the longitudinal momentum distribution (T). The temperature T_0 (a constant) describes the natural "thermal" motion of a hot interaction region. Thus, the transverse temperature will give a transverse momentum distribution that is independent of the incident energy. This will be reflected in a constant value of $\langle p_\perp \rangle$, the average transverse momentum. However, T_0 will be different for different mass particles. The temperature, T , can be considered as the quasi-equilibrium condition at which the various particles "boil off" as the interaction volume cools. T will be greater than T_0 because it will include the effect of the collective motion of the interacting gas along the direction of the incoming nucleon. We start with a quantum-statistical mechanical average occupation number density

$$v = \frac{1}{\exp[\sqrt{p^2 + m^2}/T] \pm 1}, \quad \begin{array}{l} + \rightarrow \text{fermions} \\ - \rightarrow \text{bosons} \end{array} \quad (1)$$

where p and m are the momentum and mass of the particles. The normalized momentum distributions are:

$$\omega_{\ell}^N(p_{\ell}^*) dp_{\ell}^* = \frac{T}{m^2 c^3} \frac{\sum_{k=1}^{\infty} (\pm)^{k+1} [\exp(-\frac{k\mu_1}{T})/k^{3/2}] [1 + \frac{k\mu_1}{T}]}{\sum_{k=1}^{\infty} (\pm)^{k+1} K_2(kmc^2/T)/k} dp_{\ell}^* \quad (2)$$

$$\omega_t^N(p_t^*) dp_t^* = \frac{p_t^* \mu_2}{T_o m^2 c^2} \frac{\sum_{k=1}^{\infty} (\pm)^{k+1} K_1(k\mu_2/T_o)}{\sum_{k=1}^{\infty} (\pm)^{k+1} K_2(kmc^2/T_o)/k} dp_t^* \quad (3)$$

where

$$\mu_1^2 = p_{\ell}^{*2} + m^2,$$

$$\mu_2^2 = p_t^{*2} + m^2,$$

$$p_{\ell}^* = \text{longitudinal momentum,}$$

$$p_t^* = \text{transverse momentum,}$$

and the K 's are modified Bessel functions. Starred quantities are in C.M. system. The flux of particles is given by

$$\frac{d^2 N}{dp^* d\Omega^*} = \frac{2V_o T^2}{T_o m^2 c^4 h^3} p^{*2} \mu_2 \frac{\sum_{k=1}^{\infty} (\pm)^{k+1} K_1 \left(\frac{k\mu_2}{T_o} \right) \sum_{k=1}^{\infty} (\pm)^{k+1} \exp\left(\frac{-k\mu_2}{T}\right) \left(1 + \frac{k\mu_1}{T}\right) / k^{3/2}}{\sum_{k=1}^{\infty} (\pm)^{k+1} \frac{K_2(kmc^2/T_o)}{k}} \quad (4)$$

We can write

$$T = \text{constant } (E_o^*)^{1/4} \quad (5)$$

which allows us to make calculations as a function of the incident energy, E_o .

For pions the interaction volume varies as

$$V_o = V_o' \left\{ 1 + \frac{k'}{\sigma\sqrt{2\pi}} \exp\left\{ \frac{-\langle p^* \rangle^2}{2\sigma^2} [1 - \cos(2\langle \theta^* \rangle - \theta^*)]^2 \right\} \right\} \quad (6)$$

where

$$\tan \langle \theta^* \rangle = \frac{\langle p_t^* \rangle}{\langle p_\ell^* \rangle} = \frac{\pi}{2} \sqrt{\frac{T_o}{T}} \frac{K_2(m/T) K_{5/2}(m/T_o)}{K_2(m/T_o) K_{5/2}(m/T)}, \quad (7)$$

and

$$\langle p^* \rangle^2 = \langle p_t^* \rangle^2 + \langle p_\ell^* \rangle^2. \quad (8)$$

III. USE OF PROGRAM

One must supply the following input parameters for the program:

- W1 = the rest mass of the proton in GeV/c^2 ,
- E0 = the incident energy of the proton in GeV,
- W = the rest mass of the produced particle in GeV/c^2 ,
- AL = the angle in the lab system at which the particles are produced (in radians),
- T0 = the transverse temperature in GeV,
- A5 = the conversion factor for the longitudinal temperature,
- C = the normalization constant,
- A = index for passes through program,
- SG = the first constant for meson volume,
- AK = the second constant for meson volume,
- AM = +1 for fermions, -1 for bosons,
- M1 = the number of terms kept in the sums.

The values of the fixed parameters in the correct format are shown in Table I.

For energies $E_0 < 50$ GeV, one should take $M1 = 3$ to insure accurate results for low momentum secondaries. However, above 50 GeV, $M1 = 1$ is normally acceptable. The A term should be zero (blank) except on the cards for the last spectra wanted. Then one should use any positive floating number greater than zero.

The required Bessel functions are generated by function subprograms. These are generated by a series approximation method [2]. Of course, one can easily use existing programs if it is more expedient.

If the program is to be used for incident energies greater than 1000 GeV the version given in appendix C should be used.

If one does not want the volume variation for the pions set
 $AK = SG = 0.$

Table I

Particle	To	A5	C	SG	AK
π^+	1.4E-01	2.89E-01	1.6E+03	5.0E-02	1.7E-01
π^-	1.4E-01	2.41E-01	1.5E+03	5.0E-02	1.7E-01
K^+	1.15E-01	1.92E-01	1.4E+04	-	-
K^-	1.15E-01	1.28E-01	3.3E+04	-	-
P	1.2E-01	3.01E+00	1.8E+03	-	-
\bar{P}	1.4E-01	1.02E-01	6.0E+03	-	-

Appendix A

A Listing of the Fortran II Program

THE FOLLOWING PROGRAM COMPUTES THE SPECTRA OF PARTICLES
PRODUCED IN PROTON-PROTON COLLISIONS AT HIGH ENERGIES BY THE
TWO TEMPERATURE STATISTICAL MODEL. THE INPUT PARAMETERS REQUIRED
ARE AS FOLLOWS

W1 = THE REST MASS OF THE PROTON IN GEV/C*C
EO = THE INCIDENT ENERGY OF THE PROTON IN GEV
W = THE REST MASS OF THE PRODUCED PARTICLE IN GEV/C*C
AL = THE ANGLE IN THE LAB SYSTEM AT WHICH THE PARTICLES ARE
PRODUCED (IN RADIANS)
TO = THE TRANSVERSE TEMPERATURE IN GEV
A5 = THE CONVERSION FACTOR FOR THE LONGITUDINAL TEMPERATURE
C = THE NORMALIZATION CONSTANT
A = INDEX FOR PASSES THROUGH PROGRAM
SG = THE FIRST CONSTANT FOR MESON VOLUME
AK = THE SECOND CONSTANT FOR THE MESON VOLUME
AM = +1 FOR FERMIONS, -1 FOR BOSONS
M1 = THE NUMBER OF TERMS KEPT IN THE SUMS

10 READ INPUT TAPE 5, 900, W1, EO, W, AL, TO, A5, C, A
900 FORMAT(8E10.6)
READ INPUT TAPE 5, 903, SG, AK, AM, M1
903 FORMAT(3E10.6, I10)

COMPUTE THE KINEMATIC FACTORS TO CONVERT FROM LAB TO CMS (TO LAB)

T1 = EO - W1
ETCM = SQRT((W1 + W1)**2 + 2.0*T1*W1)
GAMMA = (T1 + W1 + W1)/ETCM
ETA = SQRT(T1*(T1 + 2.0*W1))/ETCM
BETA = ETA/GAMMA

COMPUTE THE PARAMETERS FOR COMPUTATION OF CROSS-SECTION

T = A5*ETCM**0.25
9 B2 = (AL*180.0)/3.1415926
WRITE OUTPUT TAPE 6, 901, B2, TO, T, C, EO, W
901 FORMAT(///20X, 14HANGLE IN LAB =, F7.3, 5X, 4HTO =, F7.3, 5X, 3HT =,
\$ F7.3, 5X, 3HC =, F1.3, 5X, 4HEO =, F7.3, 5X, 3HW =, F7.4//
\$ 20X, 12HANGLE IN CM, 5X, 15HMOMENTUM IN LAB, 5X,
\$ 14HMOMENTUM IN CM, 5X, 16HCM CROSS SECTION, 5X,
\$ 17HLAB CROSS SECTION/)
PL = 1.00
57 FL = SQRT(PL*PL + W*W)
B1 = SIN(AL)/(GAMMA*(COS(AL) - (BETA*EL/PL)))
A2 = ATAN(ABS(B1))
IF(B1) 14, 15, 15
14 ACM = 3.1415926 - A2
GO TO 16
15 ACM = A2
16 B3 = ACM*(180.0/3.1415926)
P = SQRT((PL*SIN(AL))**2 + (GAMMA**2)*(PL*COS(AL) -
1 BETA*EL)**2)
IF(AK) 21, 21, 22
21 C1 = C
GO TO 56
22 AP1 = (SQRT(W*T/1.5707963)*BK52(W/T))/BK2(W/T)
AP2 = (SQRT(1.5707963*W*TO)*BK52(W/TO))/BK2(W/TO)
AP = SQRT(AP1*AP1 + AP2*AP2)
A1 = SG*SQRT(6.2831852)
ATH = ATAN(AP2/AP1)

```

      C1 = C*(1. + AK*EXPE(-(AP*AP*(1. - COSF(2.*ATH - ACM))**2)
1 / (2.*SG*SG))/A1)
56 P1 = P*COSE(ACM)
      P2 = P*SINE(ACM)
      U1 = SQRTF(P1*P1 + W*W)
      U2 = SQRTF(P2*P2 + W*W)

C
C
C      COMPUTE THE CROSS SECTION IN THE CMS

      X = U2/T0
      Y = U1/T
      B4 = 0.0
      B5 = 0.0
      DO 67 I=1, M1
      AI = I
      X1 = AI*X
      B4 = B4 + BK1(X1)*AM**(I+1)
      Y1 = AI*Y
67 B5 = B5 + (EXPE(-Y1)*(1.0 + Y1)*AM**(I+1))/(AI**1.5)
      DCS = C1*P*P*T*T*U2*B4*B5

C
C
C      COMPUTE THE CROSS SECTION IN THE LAB

      ECM = SQRTF(P*P + W*W)
      CONV = (ECM*PL*PL)/(EL*P*P)
      DCSL = CONV*DCS
      WRITE OUTPUT TAPE 6, 902, B3, PL, P, DCS, DCSL
902 FORMAT(20X, F7.3, 9X, F10.6, 10X, F10.6, 10X, F10.5, 10X, F10.5)
      PL = PL + 1.00
      IF(PL - F0) 57, 57, 5
      5 IF(A) 10, 10, 4
      4 CALL EXIT
      END
      FUNCTION BK1(X)
      IF(X - 2.0) 20, 21, 21
20 S = X/2.
      T = X/3.75
      BK1 = (1.0/X)*(X* LOGF(X/2.0)*X*(0.5 + .87890504*T*T +
1 .51498869*T**4 + .15084934*T**6 + .02658733*T**8 +
2 .00301532*T**10 + .00032411*T**12) + 1. +
3 .15443144*S*S - .67278579*S**4 - .18156879*S**6 -
4 .01919402*S**8 - .00110404*S**10 - .00004686*S**12)
      RETURN
21 S = 2./X
      BK1 = EXPE(-X)*(1.0/SQRTF(X))*(1.25331414 + .23498619*S -
1 .03655620*S*S + .01504268*S**3 - .00780353*S**4 +
2 .00325614*S**5 - .00068245*S**6)
      RETURN
      END
      FUNCTION B10(X)
      T = X/3.75
      B10 = 1.0 + 3.5156229*T*T + 3.0899424*T**4 +
1 1.2067492*T**6 + .2659732*T**8 + .0360768*T**10 +
2 .0045813*T**12
      RETURN
      END
      FUNCTION BK0(X)
      IF(X - 2.0) 100, 101, 101
100 R = X/2.0
      BK0 = -LOGF(R)*B10(X) - 0.57721566 + 0.42278420*R*R +
1 .23069756*R**4 + .0348859*R**6 + .00262698*R**8 +

```

```

2 .0001075*R**10 + .0000074*R**12
RETURN
101 S = 2.0/X
BK0 = (EXP(-X)/SQRT(X))*(1.25331414 - 0.07832358*S +
1 .02189568*S**2 - .01062446*S**3 + .0058782*S**4 -
2 .0025154*S**5 + .00053208*S**6)
RETURN
END
FUNCTION BK2(X)
BK2 = (2.0/X)*BK1(X) + BK0(X)
RETURN
END
FUNCTION BK52(X)
BK52 = SQRT(1.707963/X)*EXP(-X)*(1.+3./X+3./(X*X))
RETURN
END
* DATA

```

Appendix B

A Listing of the Fortran IV Program

SIPETC MAIN

```
C
C THE FOLLOWING PROGRAM COMPUTES THE SPECTRA OF PARTICLES
C PRODUCED IN PROTON-PROTON COLLISIONS AT HIGH ENERGIES BY THE
C TWO TEMPERATURE STATISTICAL MODEL. THE INPUT PARAMETERS REQUIRED
C ARE AS FOLLOWS
C W1 = THE REST MASS OF THE PROTON IN GEV/C*C
C E0 = THE INCIDENT ENERGY OF THE PROTON IN GEV
C W = THE REST MASS OF THE PRODUCED PARTICLE IN GEV/C*C
C AL = THE ANGLE IN THE LAB SYSTEM AT WHICH THE PARTICLES ARE
C PRODUCED (IN RADIANS)
C TO = THE TRANSVERSE TEMPERATURE IN GEV
C A5 = THE CONVERSION FACTOR FOR THE LONGITUDINAL TEMPERATURE
C C = THE NORMALIZATION CONSTANT
C A = INDEX FOR PASSES THROUGH PROGRAM
C SG = THE FIRST CONSTANT FOR MESON VOLUME
C AK = THE SECOND CONSTANT FOR THE MESON VOLUME
C AM = +1 FOR FERMIONS, -1 FOR BOSONS
C M1 = THE NUMBER OF TERMS KEPT IN THE SUMS
C
10 READ(5,900) W1, E0, W, AL, TO, A5, C, A
900 FORMAT(8F10.6)
READ(5,903) SG, AK, AM, M1
903 FORMAT(3E10.6, I10)
C
C COMPUTE THE KINEMATIC FACTORS TO CONVERT FROM LAB TO CMS (TO LAB)
C
T1 = E0 - W1
ETCM = SQRT((2.*W1)**2 + 2.*T1*W1)
GAMMA = (T1 + W1 + W1)/ETCM
ETA = SQRT(T1*(T1 + 2.*W1))/ETCM
BETA = ETA/GAMMA
C
C COMPUTE THE PARAMETERS FOR COMPUTATION OF CROSS-SECTION
C
T = A5*ETCM**0.25
9 B2 = (AL*180.0)/3.1415926
WRITE(6,901) B2, TO, T, C, E0, W
901 FORMAT(///20X,14HANGLE IN LAB =, F7.3,5X, 4HTO =, F7.3,5X,3HT =,
$ F7.3, 5X, 3HC =, F8.3 , 5X, 4HEO =, F7.3, 5X, 3HW =, F7.4//
$ 20X, 12HANGLE IN CM , 5X, 15HMOMENTUM IN LAB, 5X,
$ 14HMOMENTUM IN CL, 5X, 16HCM CROSS SECTION, 5X,
$ 17HLAB CROSS SECTION/)
PL = 1.00
57 EL = SQRT(PL*PL + W*W)
B1 = SIN(AL)/(GAMMA*(COS(AL) - (BETA*EL/PL)))
A2 = ATAN(ABS(B1))
IF(B1) 14, 15, 15
14 ACM = 3.1415926 -A2
GO TO 16
15 ACM = A2
16 B3 = ACM*(180.0/3.1415926)
P = SQRT((PL*SIN(AL))**2 + (GAMMA**2)*(PL*COS(AL) - BETA*EL)**2)
IF(AK) 21, 21, 22
21 C1 = C
GO TO 56
22 AP1 = (SQRT(W*T/1.5707963)*BK52(W/T))/BK2(W/T)
AP2 = (SQRT(1.5707963*W*TO)*BK52(W/TO))/BK2(W/TO)
AP = SQRT(AP1*AP1 + AP2*AP2)
ATH = ATAN(AP2/AP1)
C1 = C*(1. + AK*EXP(-(AP*AP*(1. - COS(2.*ATH - ACM))**2))
```

```

1 / (2.*SG*SG)) / (SG*.3993899))
56 P1 = P*COS(ACM)
   P2 = P*SIN(ACM)
   U1 = SORT(P1*P1 + W*W)
   U2 = SORT(P2*P2 + W*W)
C
C   COMPUTE THE CROSS SECTION IN THE CMG
C
   X = U2/T0
   Y = U1/T
   R4 = 0.0
   R5 = 0.0
   DO 67 I=1, M1
   AI = I
   X1 = AI*X
   R4 = R4 + BK1(X1)*AM**(I+1)
   Y1 = AI*Y
67 R5 = R5 + (EXP(-Y1)*(1. + Y1)*AM**(I+1))/(AI**1.5)
   DCS = C1*P*P*T*T*U2*R4*R5
C
C   COMPUTE THE CROSS SECTION IN THE LAB
C
   ECM = SORT(P*P + W*W)
   CONV = (ECM*PL*PL)/(EL*P*P)
   DCSL = CONV*DCS
   WRITE(6,902) R3, PL, P, DCS, DCSL
902 FORMAT(20X, F7.3, 9X, F10.6, 10X, F10.6, 10X, F10.5, 10X, F12.8)
   PL = PL + 1.00
   IF(PL - FO) 57, 57, 5
5 IF(A) 10, 10, 4
4 RETURN
   END
SIBETC IO
   FUNCTION RIO(X)
   IF(X- 3.75) 3, 3, 4
3 T = X/3.75
   RIO = 1.0 + 3.5156229*T*T + 3.0899424*T**4 +
1 1.2067492*T**6 + .2659732*T**8 + .0360768*T**10 +
2 .0045813*T**12
   RETURN
4 T = 3.75/X
   RIO = (EXP(X)/SQRT(X))*(0.39894228 + 0.01328592*T +
5 0.00225319*T*T - 0.00157565*T**3 + 0.00916281*T**4
6 - 0.02057706*T**5 + 0.02635537*T**6
7 - 0.01647633*T**7 + 0.00392377*T**8)
   RETURN
   END
SIBETC KO
   FUNCTION RK0(X)
   IF(X - 2.0) 100, 101, 101
100 R = X/2.0
   RK0 = -ALOG(R)*RIO(X) - 0.57721566 + 0.42278420*R*R +
1 .23069756*R**4 + .0348859*R**6 + .00262698*R**8 +
2 .0001075*R**10 + .0000074*R**12
   RETURN
101 S = 1.0/X
   RK0 = EXP(-X)*SQRT(S)*(1.2533141373 - 0.1566641916*S
5 + 0.0881112782*S*S - 0.0913909546*S**3
6 + 0.1344596228*S**4 - 0.2299850328*S**5
7 + 0.3792409730*S**6 - 0.5247277331*S**7)
   RETURN

```



```

      END
$IRFIC K1
      FUNCTION BK1(X)
      IF (X = 2.0) 20, 21, 21
      20 S = X/2.
      T = X/3.75
      BK1 = (1./X)*(X*ALOG(S)*X*(.5 + .87890594*T*T +
      1 .51498869*T**4 + .15084234*T**6 + .02658733*T**8 +
      2 .00301532*T**10 + .00032411*T**12) + 1. +
      3 .15443144*S*S - .67278579*S**4 - .18156879*S**6 -
      4 .01919402*S**8 - .00110404*S**10 - .00004686*S**12)
      RETURN
      21 S = 2./X
      BK1 = EXP(-X)*(1./SQRT(X))*(1.25331414 + .23498619*S -
      1 .03655620*S*S + .01504268*S**3 - .00780353*S**4 +
      2 .00325614*S**5 - .00068245*S**6)
      RETURN
      END
$IRFIC K2
      FUNCTION BK2(X)
      BK2 = (2.0/X)*BK1(X) + BK0(X)
      RETURN
      END
$IRFIC K52
      FUNCTION BK52(X)
      BK52 = (1.2533141*EXP(-X)*(1.+3./X+3./(X*X)))/SQRT(X)
      RETURN
      END
$DATA

```

Appendix C

A Listing of the Program for Energies Greater than 1000 GeV

*IFETC MAIN

THE FOLLOWING PROGRAM COMPUTES THE SPECTRA OF PARTICLES
PRODUCED IN PROTON-PROTON COLLISIONS AT HIGH ENERGIES BY THE
TWO TEMPERATURE STATISTICAL MODEL. THE INPUT PARAMETERS REQUIRED
ARE AS FOLLOWS

W1 = THE REST MASS OF THE PROTON IN GEV/C*C
EO = THE INCIDENT ENERGY OF THE PROTON IN GEV
W = THE REST MASS OF THE PRODUCED PARTICLE IN GEV/C*C
AL = THE ANGLE IN THE LAB SYSTEM AT WHICH THE PARTICLES ARE
PRODUCED (IN RADIANS)
TO = THE TRANSVERSE TEMPERATURE IN GEV
A5 = THE CONVERSION FACTOR FOR THE LONGITUDINAL TEMPERATURE
C = THE NORMALIZATION CONSTANT
A = INDEX FOR PASSES THROUGH PROGRAM
SG = THE FIRST CONSTANT FOR MESON VOLUME
AK = THE SECOND CONSTANT FOR THE MESON VOLUME
AM = +1 FOR FERMIONS, -1 FOR BOSONS
PC = THE LOWEST MOMENTUM OF PRODUCED PARTICLE IN GEV/C
M1 = THE NUMBER OF TERMS KEPT IN THE SUMS

THIS VERSION IS FOR INCIDENT ENERGIES GREATER THAN 1000 GEV

10 READ(5,900) W1, EO, W, AL, TO, A5, C, A
900 FORMAT(8F10.6)
READ(5,903) SG, AK, AM, PC, M1
903 FORMAT(4F10.6, I10)

COMPUTE THE KINEMATIC FACTORS TO CONVERT FROM LAB TO CMS (TO LAB)

T1 = EO - W1
ETCM = SQRT((2.*W1)**2 + 2.*T1*W1)
GAMMA = (T1 + W1 + W1)/ETCM
ETA = SQRT(T1*(T1 + 2.*W1))/ETCM
BETA = ETA/GAMMA

COMPUTE THE PARAMETERS FOR COMPUTATION OF CROSS-SECTION

T = A5*ETCM**0.25
9 B2 = (AL*180.0)/3.1415926
WRITE(6,901) TO, T, C, EO, W
901 FORMAT(///20X, 4HTO =, F10.4, 5X, 3HT =, F10.4, 5X, 3HC =, F10.4,
\$ 5X, 4HEO =, F14.7, 5X, 3HW =, F10.4//
\$ 20X, 15HMOMENTUM IN LAB, 10X,
\$ 14HMOMENTUM IN CM, 10X, 16HCM CROSS SECTION, 10X,
\$ 17HLAB CROSS SECTION//
PL = PC
57 FL = SQRT(PL*PL + W*W)
IF(AK) 21, 21, 22
21 C1 = C
GO TO 56
22 AP1 = (SQRT(W*T/1.5707963)*BK52(W/T))/BK2(W/T)
AP2 = (SQRT(1.5707963*W*TO)*BK52(W/TO))/BK2(W/TO)
AP = SQRT(AP1*AP1 + AP2*AP2)
ATH = ATAN(AP2/AP1)
C1 = C*(1. + AK*EXP(-(AP*AP*(1. - COS(2.*ATH - ACM))**2)
1 / (2.*SG*SG))/(SG*.3993899))
56 P1 = (PL*COS(AL))/(2.*GAMMA)
P2 = PL*SIN(AL)
P = SQRT(P1*P1 + P2*P2)
U1 = SQRT(P1*P1 + W*W)

```

      U2 = SQRT(P2*P2 + W*W)
C
C      COMPUTE THE CROSS SECTION IN THE CMS
C
      X = U2/T0
      Y = U1/T
      R4 = 0.0
      R5 = 0.0
      DO 67 I=1, M1
      AI = I
      X1 = AI*X
      R4 = R4 + BK1(X1)*AM**(I+1)
      Y1 = AI*Y
67  R5 = R5 + (EXP(-Y1)*(1. + Y1)*AM**(I+1))/(AI**1.5)
      DCS = C1*P*P*T*T*U2*R4*R5
C
C      COMPUTE THE CROSS SECTION IN THE LAB
C
      ECM = SQRT(P*P + W*W)
      CONV = (ECM*PL*PL)/(EL*P*P)
      DCSL = CONV*DCS
      WRITE(6,902) PL, P, DCS, DCSL
902  FORMAT(20X, 1PE14.7, 10X, 1PE14.7, 10X, 1PE14.7, 10X, 1PE14.7)
      PL = PL + PC
      IF(PL = 10.*PC) 57, 30, 30
30  PC = 10.*PC
      PL = PC
      IF(PC = E0) 57, 5, 5
5  IF(A) 10, 10, 4
4  RETURN
      END
SIBFIC IO
      FUNCTION BIO(X)
      IF(X= 3.75) 3, 3, 4
3  T = X/3.75
      BIO = 1.0 + 3.5156229*T*T + 3.0899424*T**4 +
1  1.2067492*T**6 + .2659732*T**8 + .0360768*T**10 +
2  .0045813*T**12
      RETURN
4  T = 3.75/X
      BIO = (EXP(X)/SQRT(X))*(0.39894228 + 0.01328592*T +
5  0.00225319*T*T - 0.00157565*T**3 + 0.00916281*T**4
5  - 0.02057706*T**5 + 0.02635537*T**6
5  - 0.01647633*T**7 + 0.00392377*T**8)
      RETURN
      END
SIBFIC KO
      FUNCTION BKO(X)
      IF(X = 2.0) 100, 101, 101
100 R = X/2.0
      BKO = -ALOG(R)*BIO(X) - 0.57721566 + 0.42278420*R*R +
1  .23069756*R**4 + .0348859*R**6 + .00262698*R**8 +
2  .0001075*R**10 + .0000074*R**12
      RETURN
101 S = 1.0/X
      BKO = EXP(-X)*SQRT(S)*(1.2533141373 - 0.1566641816*S
5  + 0.0881112782*S**2 - 0.0913909546*S**3
5  + 0.1344596228*S**4 - 0.2299850328*S**5
5  + 0.3792409730*S**6 - 0.5247277331*S**7)
      RETURN
      END

```

\$IRFIC K1

FUNCTION BK1(X)

IF(X = 2.0) 20, 21, 21

20 S = X/2.

T = X/3.75

BK1 = (1./X)*(X*ALOG(S)*X*(.5 + .87890594*T*T +
1 .51498869*T**4 + .15084934*T**6 + .02658733*T**8 +
2 .00301532*T**10 + .00032411*T**12) + 1. +
3 .15443144*S*S - .67278579*S**4 - .18156879*S**6 -
4 .01919402*S**8 - .00110404*S**10 - .00004686*S**12)

RETURN

21 S = 2./X

BK1 = EXP(-X)*(1./SQRT(X))*(1.25331414 + .23498619*S -
1 .03655620*S*S + .01504268*S**3 - .00780353*S**4 +
2 .00325614*S**5 - .00068245*S**6)

RETURN

END

\$IRFIC K2

FUNCTION BK2(X)

BK2 = (2.0/X)*BK1(X) + BK0(X)

RETURN

END

\$IRFIC K52

FUNCTION BK52(X)

BK52 = (1.2533141*EXP(-X)*(1.+3./X+3./(X*X)))/SQRT(X)

RETURN

END

\$DATA