



STANFORD RSL Technical Report 69-2

QUANTITATIVE GEOLOGIC ANALYSIS OF
MULTIBAND PHOTOGRAPHY FROM THE
MONO CRATERS AREA, CALIFORNIA

by

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Geology Department
School of Earth Sciences
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Stanford, California

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February 1969

*Now at Bendix Corporation, Ann Arbor, Michigan

REMOTE SENSING LABORATORY
SCHOOL OF EARTH SCIENCES

STANFORD UNIVERSITY • STANFORD, CALIFORNIA

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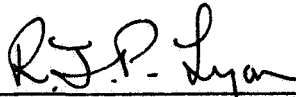
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This research report was submitted to the Department of Geology of Stanford University in partial fulfillment of the requirements for the Degree of Master of Science, March 1968. Partial support was provided under NASA Grant NGR-05-020-115 and NASA Contract NAS9-7313 in the Stanford Remote Sensing Laboratory.



R.J.P. Lyon
Principal Investigator

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I. INTRODUCTION

The primary purpose of this report is to determine the effectiveness of remote sensing techniques in geologic mapping applications. This was accomplished by selecting an area of previously mapped geology and analyzing it with all remote sensing data available. The area selected for study was the northern portion of the Mono Craters, Mono County, California, primarily because data of all the major wavelength regions were available when research for this report began in September, 1966. Specific areas of interest were field checked and samples collected in April, August and October of 1967 in attempts to correlate local geology with aircraft data, using quantitative analytical approaches.

I would like to acknowledge the assistance provided by Dr. R. J. P. Lyon, who was responsible for obtaining the basic National Aeronautics and Space Administration remote sensing data and advising in its interpretation. Only the multiband photographic data from the area was analyzed in great detail. Radar and infrared scanner data were also studied but not in the depth required for quantitative geologic evaluation.

The portion of the Mono Craters studied in this report is located in the northeastern quadrant of the Mono Craters 15-minute quadrangle and includes the area from the south shore of Mono Lake to about Crater Mountain, near the center of the main arc of the Mono Craters. This area is from three to five miles east of the frontal scarp of the Sierra Nevada and is on the bor-

der between the Sierra Nevada and Basin and Range physiographic provinces. Topographic relief of the craters above the surrounding pumice plains varies from 200 feet at North Crater, to 2000 feet at Crater Mountain (Figure 1). All major streams in the area originate in the Sierras and flow into Mono Lake, a highly saline body of water which has no external drainage. Water from rain or snowfall in the craters drain into the sparsely vegetated and porous areas of pumice and does not produce significant surface drainage. There are also areas of springs along the shore of Mono Lake, some of which are hot springs.

The climate of this region is arid to semi-arid and supports little vegetation in areas of rock outcrop, other than lichen and a few widely scattered pine trees. Regions of pumice surround these outcrops, and may vary from essentially unvegetated on steep slopes, to dense growths of sagebrush in areas of gently-sloping topography. Other locations of little or no vegetation include deposits of beach sand and formations of calcareous tufa along the shore of the lake. This combination of sparse vegetation, distinctive terrain types, and available data has provided an excellent area for geologic remote sensing investigations.

II. GEOLOGY

The regional geology of the Mono Craters area consists of a basement complex of metamorphosed Paleozoic and Mesozoic sedimentary and volcanic rocks and Mesozoic granitic rocks. This was demonstrated by Kistler (1966), who also states that the granitic rock is composed chiefly of granodiorite and quartz monzonite and occurs in plutons of moderate size. This complex is covered by Tertiary and Quaternary volcanic rocks and Quaternary lake sediments, moraines, pumice deposits and alluvium.

The geologic units studied are all of Recent age, including the flows of the main crater complex. This was indicated in work done by Kistler (1966) and Friedman (1966), and emphasizes the extremely youthful age of these units. The arcuate form of the main crater complex has suggested that this is a portion of a ring-fracture zone that formed after the Sherwin Glaciation and before the eruption of the Bishop Tuff (Kistler, 1966).

The geologic maps used in this report are all derived from Friedman's compilation of mapping done in the area, as they provide the most detailed information available (Fig. 2 and 3).

III. ANALYSIS OF VISUAL AND NEAR INFRARED FIELD SPECTRA*

Although reflectance spectras are available in the visible and near infrared region from 0.4 to 1.5 microns, there is very little data of a geological nature which can be used for field analysis. This situation is especially evident when analyzing the remote sensing data presently available in the form of multiband photography from Mono Craters, California. This is currently being studied in an attempt to perfect a method of multi-spectral geologic mapping.

The field spectra presented here were obtained at North Crater (sometimes known as Panum Crater), a rhyolite, obsidian and pumice explosion crater approximately one-half mile in diameter and rising 200 to 300 feet above a plain of moderately vegetated pumice. It is located about a mile south of Mono Lake and is a northern extension of the main arc of the Mono Craters (See Fig. 1 and 2).

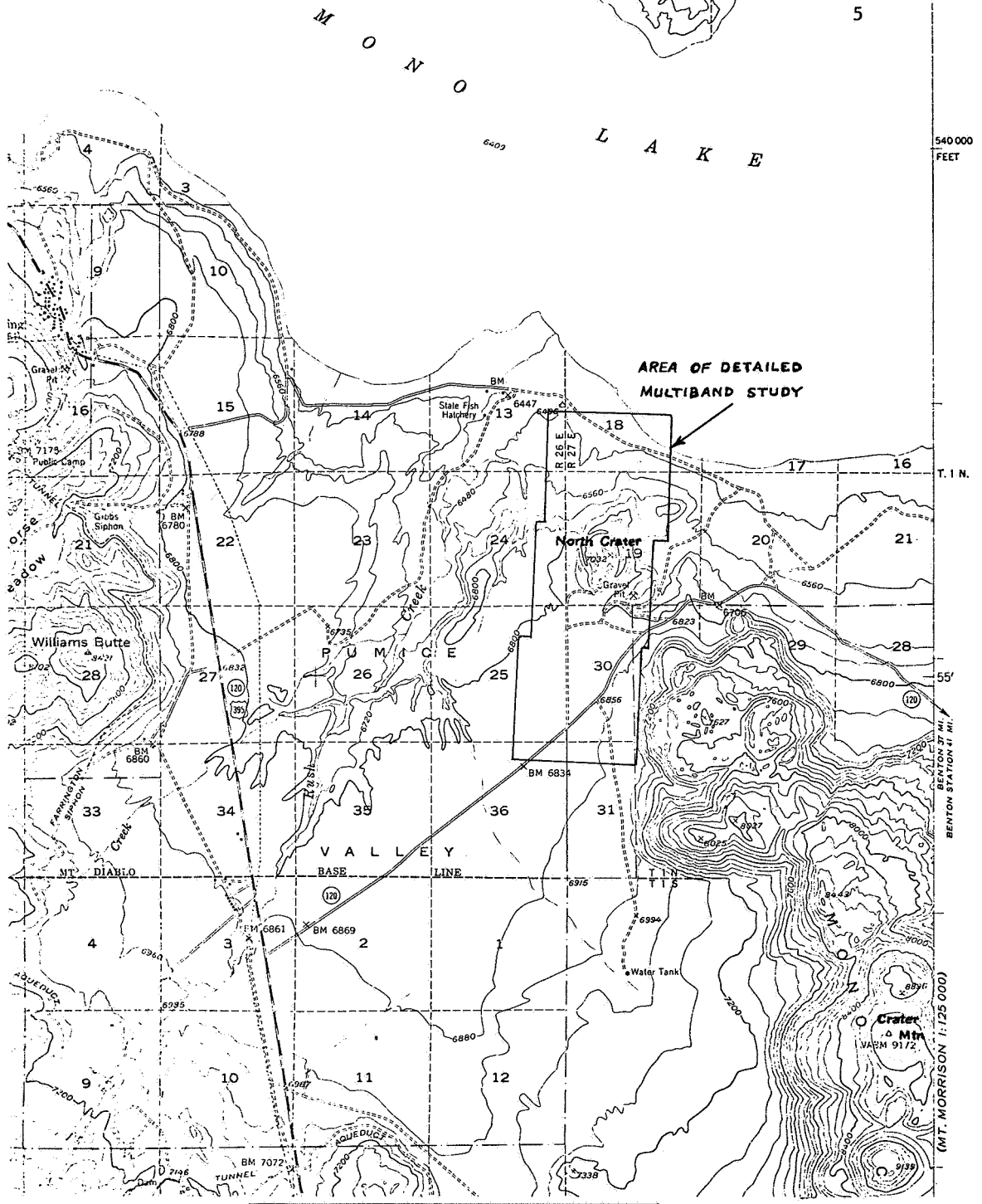
Although the conditions for obtaining on-site reflectance spectra on April 1, 1967 were not ideal, the methods employed were as consistent as possible and agree with spectra collected during August, 1967.

An ISCO spectro-radiometer fitted with a fiber-optics, remote probe was used to obtain these spectra, which ranged in wavelength from 0.425 to 1.55 μ . The reflectance at each wavelength was calculated by taking the ratio of reflected to total intensity, both intensities being measured over a 180^o (2 π steradian) field of view.

*From: Semi-Annual Report dated 15 May 1967, "Field Infrared Analysis of Terrain", NASA Grant NGR-05-020-115.

ORT 1:125 000) R. 26 E. 5' 12420.000 FEET R. 27 E. 119°00' 38"00'

MONO CRATERS QUADRANGLE
CALIFORNIA
15 MINUTE SERIES (TOPOGRAPHIC)



SCALE 1:62,500

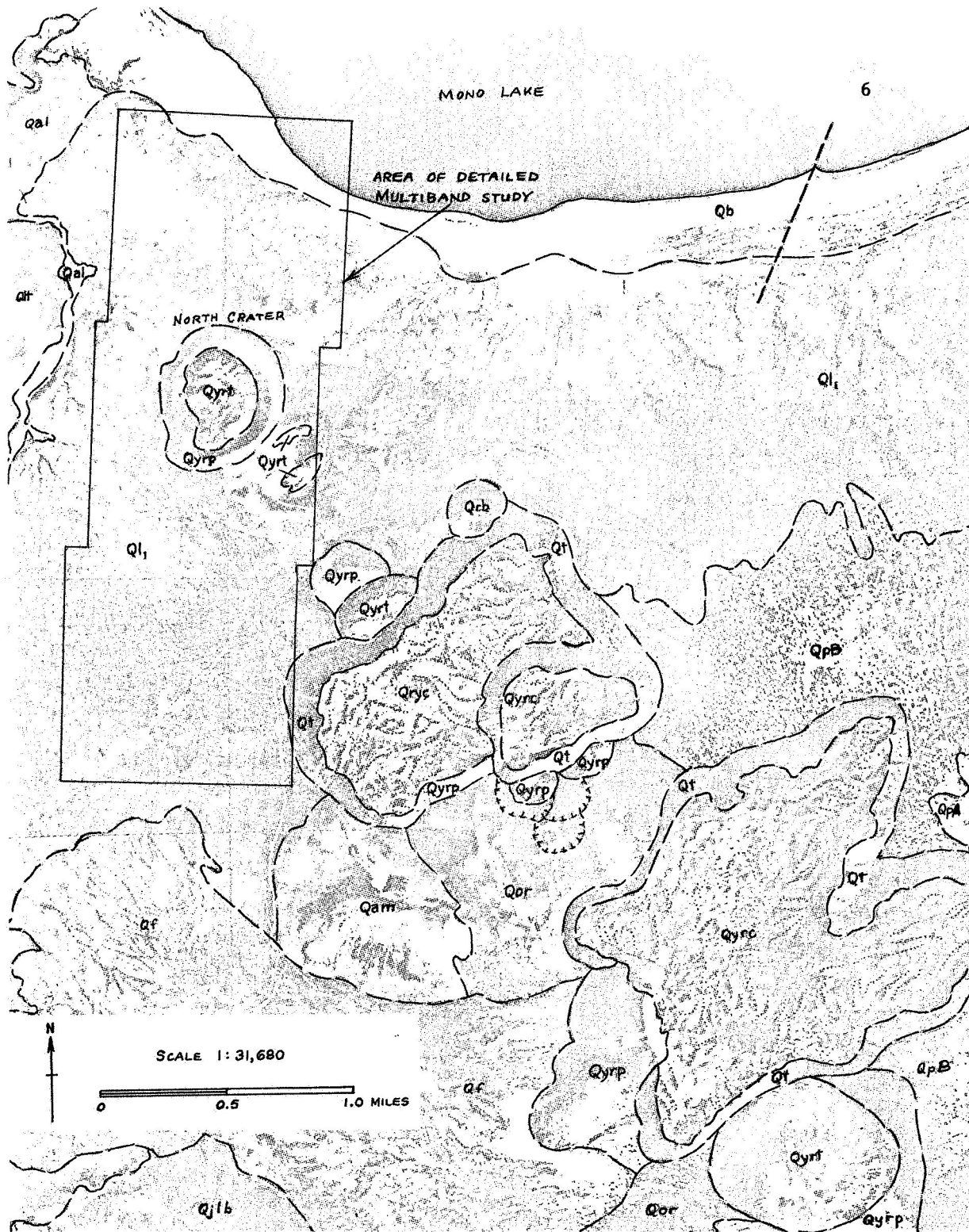
(CONTOUR INTERVAL 80 FT.) 1000 5000 10000 15000 FEET

1 .5 0 1 2 3 4 KILOMETERS

FIGURE 1



QUADRANGLE LOCATION



PHOTOGEOLOGIC MAP OF THE NORTHERN MONO CRATERS, CALIFORNIA

After J. D. Friedman, 1966

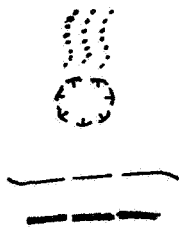
Figure 2

7

**GEOLOGIC COLUMN TO ACCOMPANY PHOTOGEOLOGIC
MAP OF NORTHERN MONO CRATERS AREA, CALIFORNIA**

AGE	SEDIMENTARY UNITS	IGNEOUS UNITS
RECENT	Qb BEACH DEPOSITS OF MONO LAKE INCLUDING CALCAREOUS TUFA AND DELTAIC DEPOSITS.	
	Qt SLIDEROCK AND TALUS OF COULEE FRONT.	
	Qf FAN AND SLOPEWASH DEPOSITS	
	Qa1 STREAM AND OTHER ALLUVIAL DEPOSITS INCLUDING REWORKED ASHFALL DEPOSITS.	
	Qit DISSECTED STREAM-CUT TERRACES IN LACUSTRINE DEPOSITS WITH SOME Qa1.	
		QpA PUMICE TEPHRA (ASH AND LAPILLI) ASHFALL VENEER AND BARREN PUMICE SANDFLATS.
		QpB SAME AS QpA BUT WITH VEGETATION AND MINOR RELIEF.
		Qyrc YOUNGER PUMICEOUS RHYOLITE OBSIDIAN OF COULEES.
		Qyrt YOUNGER RHYOLITE AND OBSIDIAN OF VOLCANIC DOMES IN PLACES WITH OVERLYING Qyrc.
		Qrb YOUNGER OBSIDIAN AND RHYOLITE OF BLOCK CRATERS.
		Qyrp YOUNGER TEPHRA (LAPILLI, ASH) RINGS, RAMPARTS AND CONES.
	Ql1 LACUSTRINE DEPOSITS INCLUDING OLDER DELTAIC GRAVELS OVERLYING Qyrc EAST AND SOUTHEAST OF NORTH CRATER.	
		Qor OLDER RHYOLITE DOMES AND FLOWS.
		Qam ANDESITE OF MONO CRATERS.
	Qjlb BASALT OF JUNE LAKE JUNCTION	

MAP SYMBOLS:



FLOW RIDGES ON COULEE SURFACES, TREND GENERALLY TRANSVERSE TO DIRECTION OF MOVEMENT.

EXPLOSION CRATERS OR VENTS IN PYROCLASTIC DEPOSITS, RHYOLITE OR ANDESITE-BASALT EXTRUSIVES.

CONTACTS

INFERRED FAULTS

FIGURE 3

As this is a completely portable instrument, spectra of samples were taken in place and the samples collected for further study. Of the twenty wavelengths used, nine correspond to the filtered bands of the multiband camera.

When these reflectance values (Table I) are plotted against wavelength (Fig. 4) we notice that most of the variation occurs in the visible region from 0.425 to 0.750 μ and that, with the exception of black obsidian, the highest reflectance occurs in the near infrared. The two samples of gray obsidian are both from the Mono Craters area, but the one showing highest spectral reflectance is from a sample provided for field analysis by the University of Nevada's NASA Project. Both of these spectra are similar below 0.575 μ (violet to yellow light), but become less similar at wavelengths greater than this. Note particularly the higher overall intensity of the rough gray obsidian relative to the smooth one. This may be caused by the rough surface of the first obsidian, apparently is more reflectant than the smooth surface of the other sample.

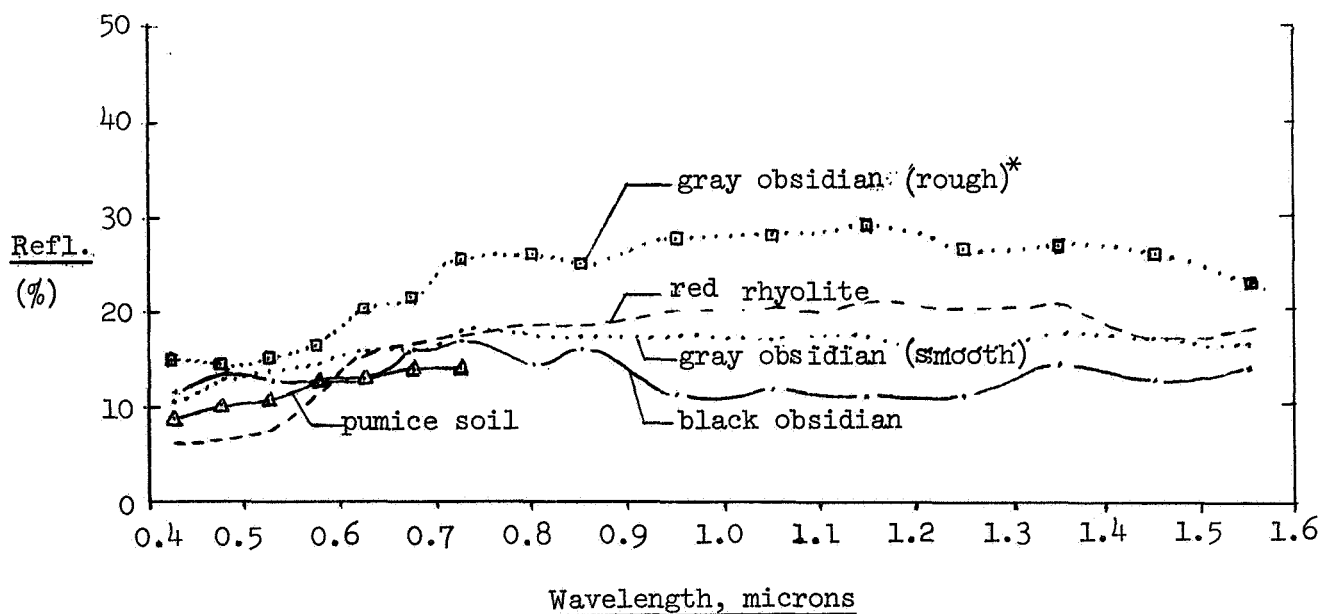
The spectrum of black obsidian is uniform and unexpectedly high over much of the visual region with reflectance reaching a minimum in the infrared. The smooth surface of this obsidian block helps to explain its high reflectance, as the sun was at a very low angle and the outlet was on a steep slope facing the sun.

The spectrum of red rhyolite shows very low reflectance in the shorter wavelength visible region, with an expected maximum in the 0.75 μ (red) region. This high reflectance persists into

TABLE I

VISUAL AND NEAR INFRARED FIELD DATAFROM MONO LAKE AREA, CALIFORNIA

<u>Reflectance, %;</u> wavelength, microns	gray rough obsidian*	gray smooth obsidian	black smooth obsidian	red smooth rhyolite	pumice soil
0.425	14.9	10.3	11.4	5.7	8.8
0.475	14.5	13.0	13.3	6.7	10.2
0.525	15.2	14.0	12.9	7.6	10.7
0.575	16.7	14.6	12.9	11.4	12.2
0.625	20.3	16.0	12.5	15.6	12.9
0.675	21.4	15.8	15.6	16.7	14.2
0.725	25.6	18.3	17.0	17.7	14.6
0.800	25.9	17.5	14.3	18.7	--
0.850	25.0	17.5	16.0	18.6	--
0.950	27.5	17.7	11.1	20.0	--
1.050	27.8	16.7	11.8	20.0	--
1.150	29.1	17.0	11.0	20.8	--
1.250	26.3	15.9	11.1	20.0	--
1.350	26.7	17.4	14.3	20.6	--
1.450	25.9	17.0	12.7	17.0	--
1.550	23.0	16.3	13.7	17.9	--

REFLECTANCE SPECTRA

* Sample collected by Univ. of Nevada NASA Project.

Figure 4

the infrared and shows little variation except at very long wavelengths. In contrast, the pumice soil spectrum shows higher values in the violet and lower values in the red regions of the visual, with a crossover near 0.60μ (orange). The pumice spectrum does not extend into the infrared, as time did not permit its completion.

A characteristic of all the longer wavelength infrared spectra is their tendency to converge toward the same reflectance in the region near 1.55μ , with no crossovers occurring in the entire near infrared region. Combining this information with that obtained in the visual portion, we note that there are a number of optimum* wavelengths which could be used for rock spectra classification. In the visible these are 0.425μ (violet), 0.475μ (blue), and 0.725μ (red), with 0.80μ in the photographic infrared and 0.90 to 1.35μ and 1.5μ in the region beyond photographic detection. Thus, by using these bands, the non-vegetated areas near Mono Craters could be classified as red rhyolite, black obsidian, gray obsidian or pumice. Hopefully, other areas of similar lithology and vegetation cover could be mapped using presently available data in the area around Mono Lake. This assumes, of course, that repeatable reflectance spectra could be obtained under other field conditions. Clearly this would form the foundation of any future research.

*Optimum: Defined as those wavelengths at which relative maximum or minimum reflectance values occur.

IV. MULTIBAND GEOLOGIC ANALYSIS

A. System Description:

This system is designed to gather information in the visible and near or short wavelength infrared region of the electromagnetic spectrum by means of nine combinations of films and filters. The nine-lens camera (Fig. 5) accomplishes this by simultaneously exposing three frames of each of three rolls of film as the aircraft flies over the terrain to be "sensed". Two of these films are conventional panchromatic (black-and-white) film, while the third is infrared black-and-white film (Fig. 6).

Unfortunately, the roll of infrared film for this flight was spoiled in the developing process and only the six visual filter combinations, or "bands" were able to be studied.* The wavelengths of visual light recorded by these bands correspond roughly to the colors violet, blue, green, yellow, orange, and red--in order of band-number from one to six (Fig. 7). Even with the drawback of not having the infrared bands, it is still clearly possible to obtain meaningful data from visual film and will serve as a demonstration of a general method of multi-spectral analysis.

Of the approximately sixty sets (each set consists of six spectral bands) of exposures which were taken over the region near Mono Lake, three sets of six spectral bands were chosen because they showed a diversity of features. As this flight was

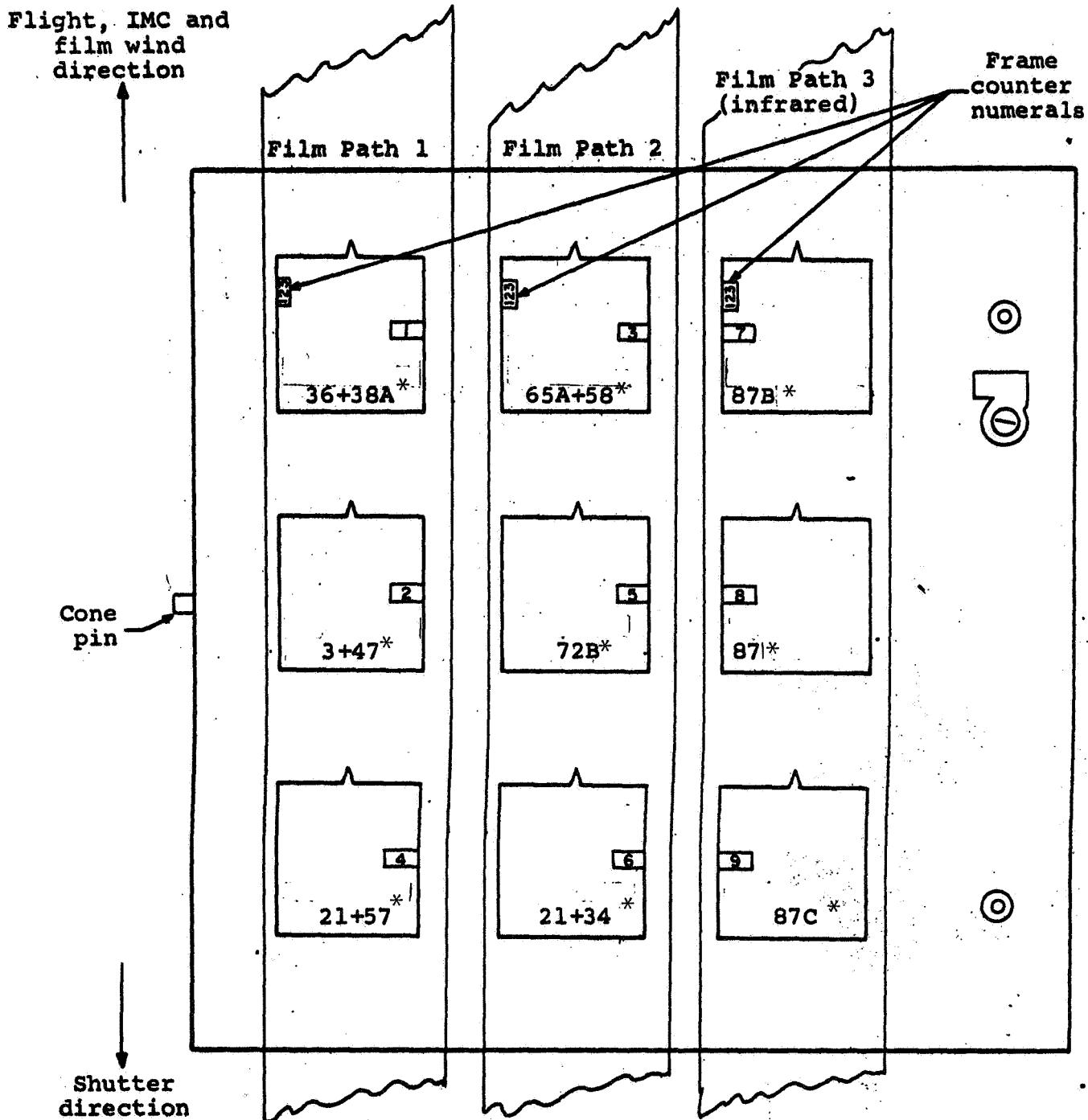
*The Mono Craters area had only been flown once with the NASA 9-lens camera prior to January 1, 1968.



Itek Nine-Lens 70-mm Camera, Model 2.

(From NASA Anthology, 1966)

Figure 5



* Wratten filter combinations

Schematic view of format area nine-lens camera model 2.

(From NASA Anthology, 1966)

Figure 6

IMPORTANT PARAMETERS FOR MULTIBAND PHOTOGRAPHY

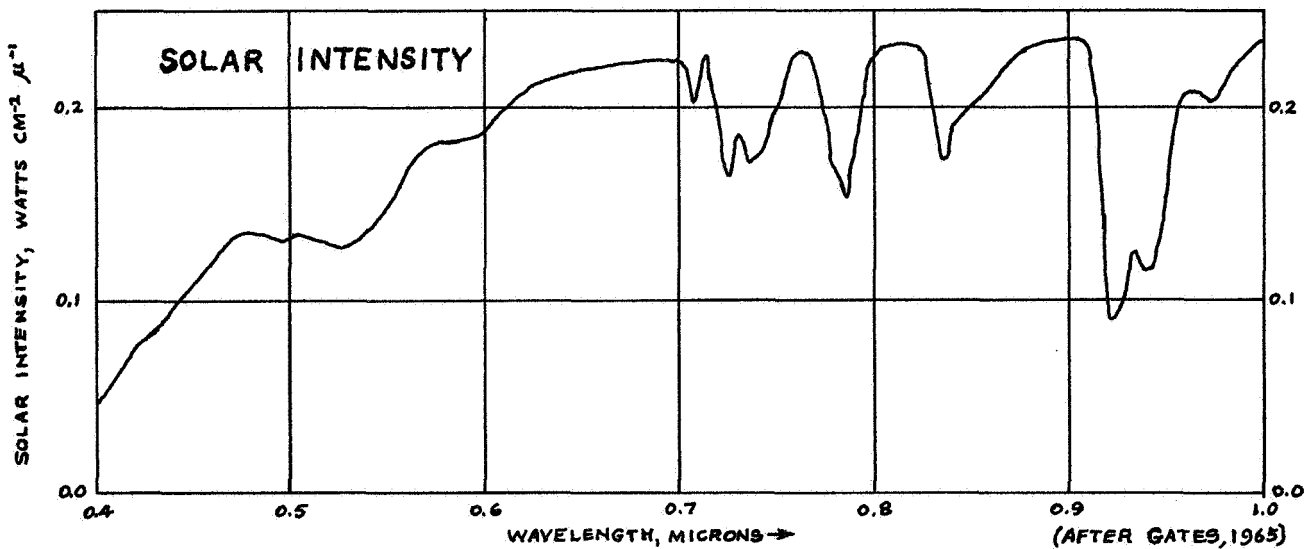
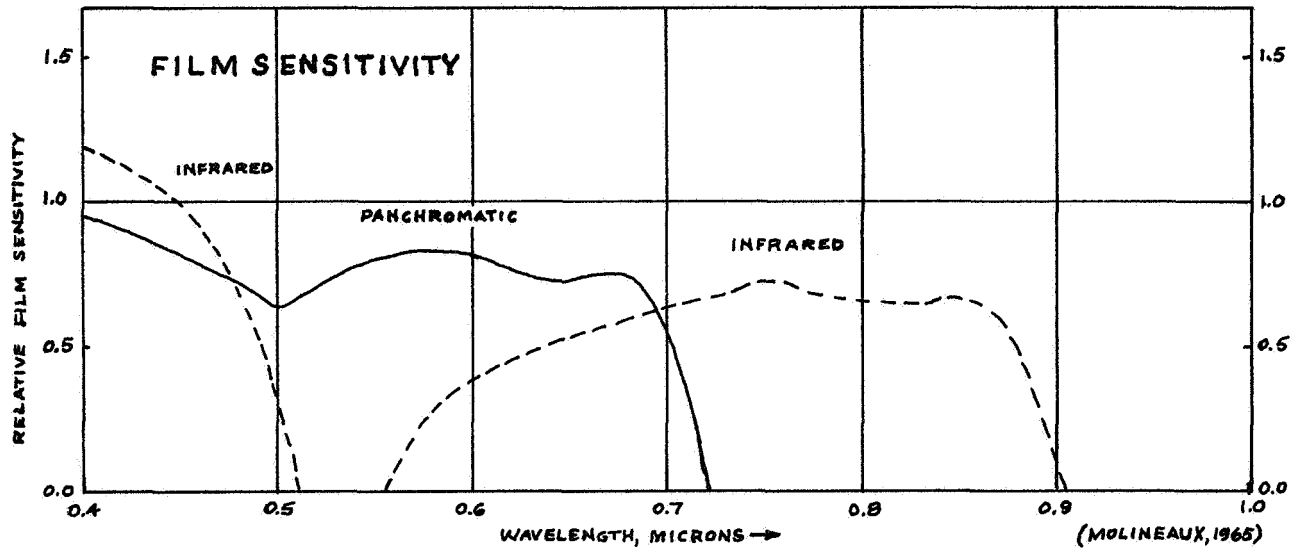
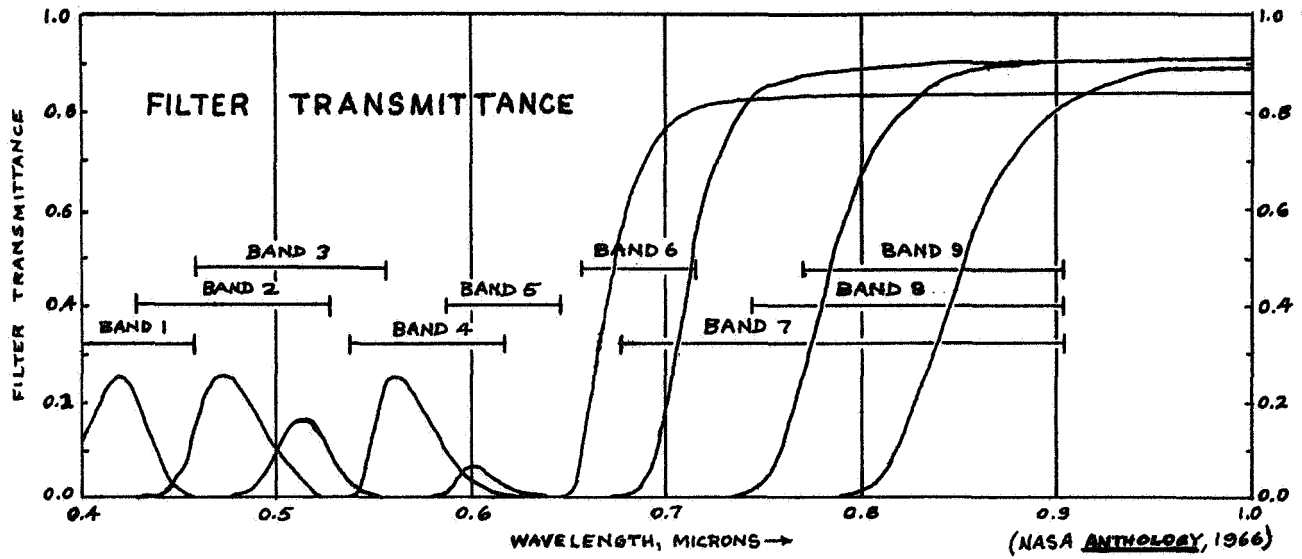


Figure 7

flown at about 9:00 A.M. on September 30, 1965, there were several areas of shadow which also served as another potential type of feature for classification. The three sets of photographs show adjacent areas along a north-south distance of about three miles, with the average width of each photograph being about one mile and overlap amounting to less than 10 percent on each. The northern-most photograph shows a portion of the south shore of Mono Lake, a strip of beach deposits, and an area of varying amounts of pumice sand and sagebrush, while the middle photo contains a small rhyolite and obsidian dome surrounded by a plain of pumice vegetated with sagebrush and crossed by roads. The southern-most photo also shows an area of pumice sand covered with sagebrush and crossed by California Highway 120.

B. Data Acquisition:

Having selected the areas to be studied, the next step was to find a satisfactory method of obtaining "gray scale" or, as these were transparencies, the transmittance values from each of the 18 exposures. After experimenting with several densitometers, the Jarrel-Ash Recording Spectrographic Densitometer was chosen for its sensitivity, film positioning accuracy, and the permanent record of transmittance values which it records on chart paper. With this instrument five profiles of transmittance were obtained along north-south lines spaced 10 mm apart and centered about the middle of each exposure. The graphical output of the densitometer was then digitized at intervals equivalent to 5 mm on the film, with this being done for each

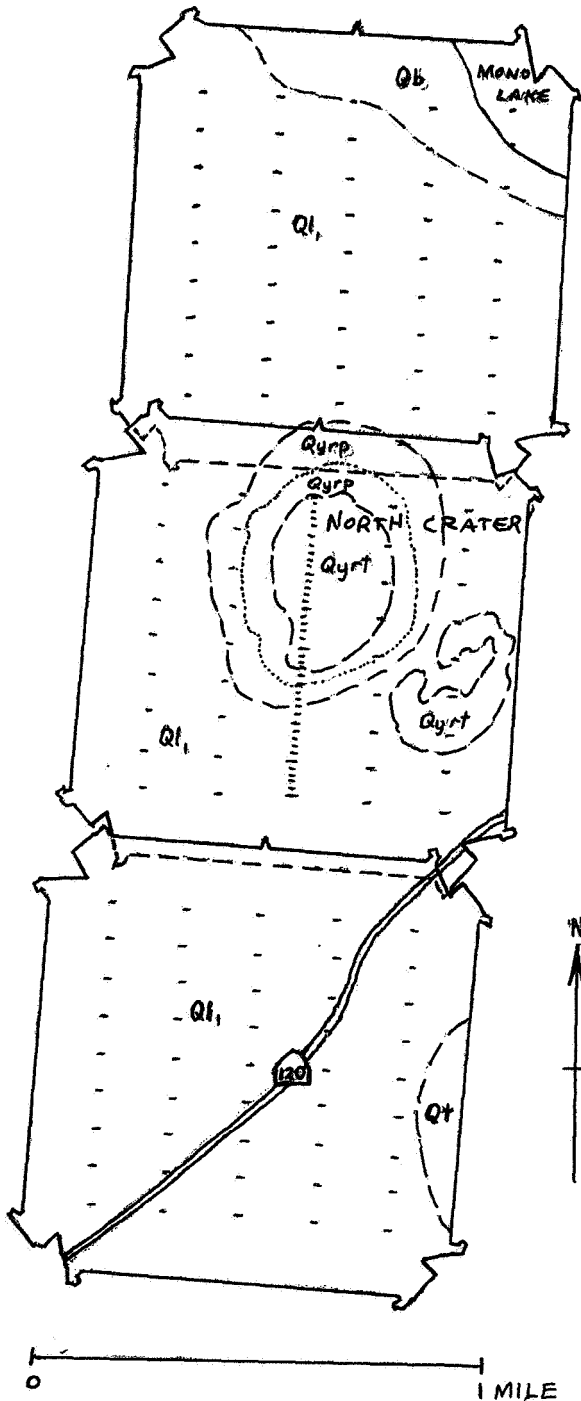
of the five profiles obtained from each of the 18 photos (Fig. 8). There were nine of these data points on each of the profiles, so that no measurements were taken closer than 10 mm from the edges of the positives--in the region of abnormally high film transmittance. The photographs used for Figures 9 and 10 are rejected duplicates and are used for illustration purposes only. The original transparencies are to be used in further experiments and were not cut apart for illustration purposes.

Having obtained five profiles of nine points each, for a total of 45 points per transparency, additional data was obtained from the central north-south profile over North Crater from forty points spaced at 1 mm intervals for use in one of the later methods of analysis. The first method of analysis required the use of only the three sets of 45 points for a total of 135 data points, each covering a field of view of 0.3 by 0.5 mm and representing six filtered bands. Each of these transmittance values were recorded on computer data cards in groups of six bands per location preceded by a coded description of the geology and vegetation as decided by ground inspection and photographic interpretation.

C. Methods of Analysis:

There were three primary methods of analysis used on this data, with each method being an attempt to discriminate the various terrain, geology, vegetation and lighting types present in the area covered by each of the photos. These three methods

PHOTOGRAPH GRID SYSTEM



DIMENSIONS:

	MM. ON PHOTO	FEET ON GROUND
DENSITOMETER SAMPLE SIZE	0.3 x 0.5	21 x 45
<u>SAMPLE SPACING:</u>		
OVERALL GRID	5.0 x 10.0	450 x 900
CENTERLINE	1.0	90

GEOLOGIC UNITS:

- Qb - BEACH DEPOSITS.
- Qt - SLIDEROCK AND TALUS.
- Qyrt - YOUNGER RHYOLITE AND
OBSIDIAN DOME.
- Qyfp - YOUNGER TEPHRA
- Ql₁ - LACUSTRINE DEPOSITS
INCLUDES OLDER DELTAIC
GRAVELS AND OVERLYING
Qyfp NEAR NORTH CRATER.

- CALIFORNIA HIGHWAY 120

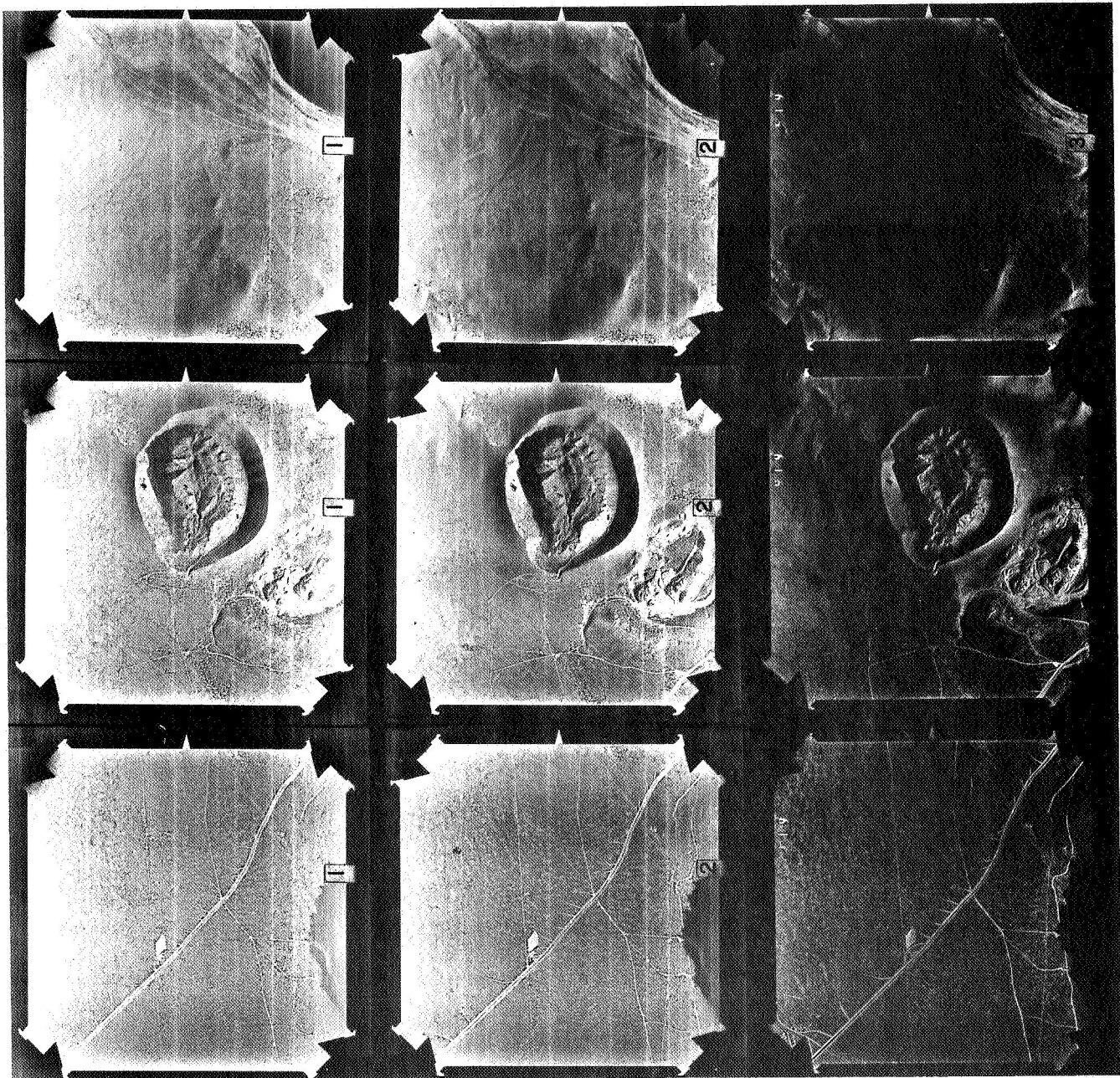
CONTACTS

CRATER RIM

SAMPLE AREAS

SCALE 1:36,600

FIGURE B



BAND 1
0.355 - 0.460 μ

BAND 2
0.430 - 0.530 μ

BAND 3
0.460 - 0.560 μ

Figure 9



BAND 4
0.540 - 0.630 μ

BAND 5
0.590 - 0.650 μ

BAND 6
0.640 - 0.720 μ

Figure 10

may be classed as investigations of trend surfaces, six-dimensional distances, and statistical models. The trend surfaces were the first to be studied and in general they were least valuable in discriminating terrain types, perhaps because only the trend surfaces themselves were studied and not the residual values of the transmittance data.

1. Trend Surfaces

In order to apply the method of trend surfaces to the multi-band data, each wavelength of light was assumed to have its relative reflectance from the ground, proportional to the transmittance values obtained from the photography. This is not necessarily correct, as solar intensity varies with wavelength and even this reflected energy is approximately proportional to the film density of the negative, where density = $\log_{10} (1/\text{transmittance})$. This also assumes that the terrain being investigated is of uniform nature and does not consider variations in slope, surface irregularities, or minor vegetation changes. Variations from one site to another may be influenced by changes in sun angle, aircraft altitude, atmospheric transmission and cloud cover, as well as variations in film condition and processing.

Having made the assumption that reflectance is related to film transmittance, the 45 transmittance values on each of the three sets of six bands were fitted to a second degree trend surface (Harbaugh, 1964) of the form $Z = AX^2 + BXY + CY^2 + D$; where Z is the value of the transmittance in percent, X and Y

are the coordinates of the point in millimeters from the upper left-hand corner of the photo, and A, B, C and D are the coefficients. When the resulting computer print-out trend contours were examined (Figures 11 and 12), most geologic discrimination was shown by band six, or red light. Even this did not bring out the exact location of the crater in the middle photo,* but did delineate the shore of Mono Lake fairly well. Another drawback of these trend surfaces is that they are not in alignment at photograph boundaries, even assuming 10 percent overlap.

A second method of trend surface analysis was used when the value of each of the spectral bands at each geographic point was compared to the mean of all six band transmittances. The formula for this was $DEV = \frac{TRANS}{MEAN} - 1$; where DEV is the positive or negative value of the transmittance of one of the six bands when compared to the average of all the bands at that point, TRANS is the transmittance of the band under study, and MEAN is the mean of all six bands at the point of interest. An intermediate program was written to make this computation and punch the six values of DEV on one card per "spectrum". Trend surfaces of these photographs were then contoured using the second-degree function in the above example, with much the same results (Fig. 13, 14). Once again the red wavelengths showed most discrimination.

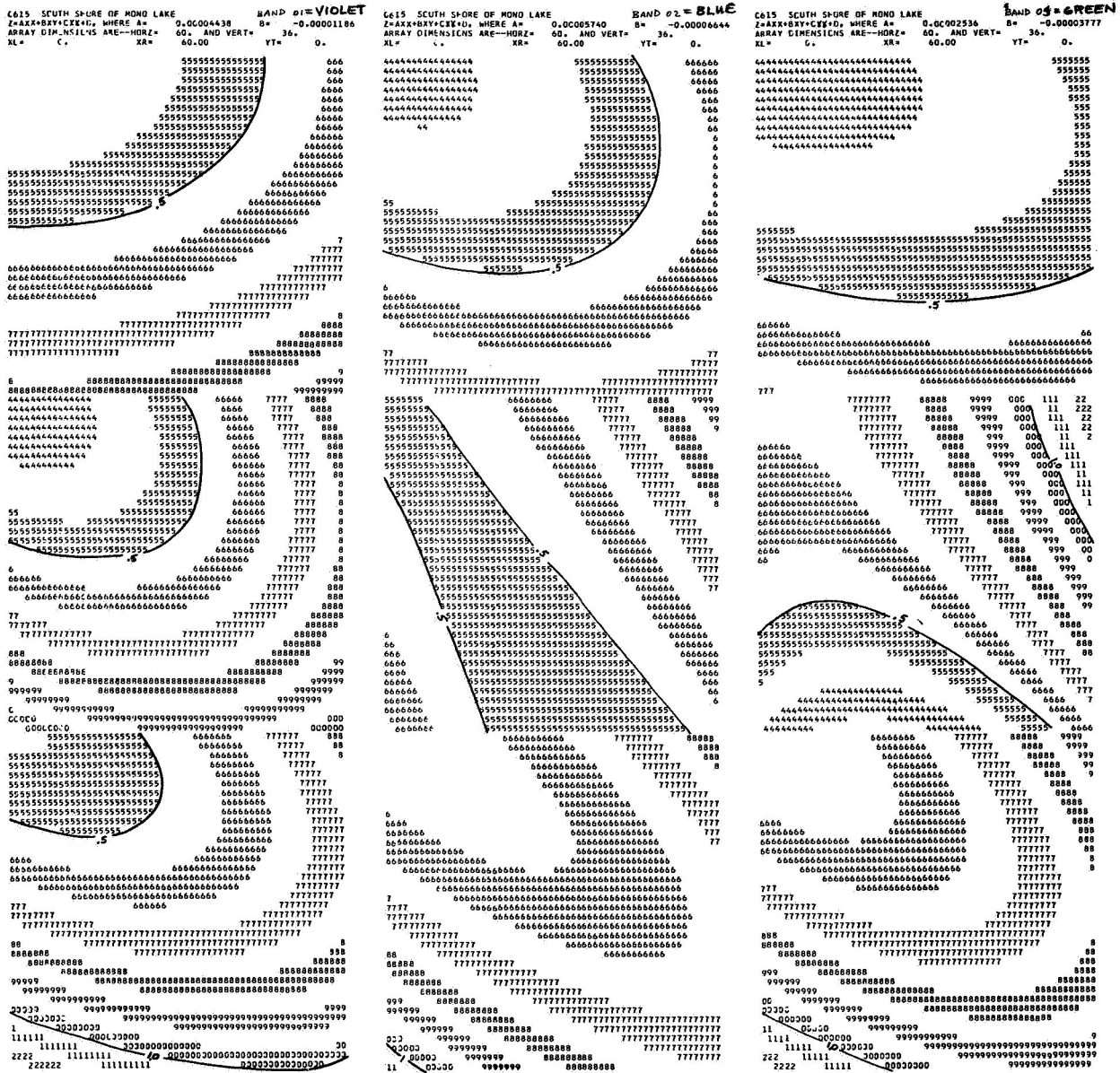
2. Six-dimensional distances

The second major method of analysis assumed that if an area could be divided into several general terrain and geology

* Compare Fig. 11 with Fig. 9, and Fig 12 with Fig. 10 etc.

SECOND DEGREE TREND SURFACES FITTED TO 45

MULTIBAND FILM TRANSMITTANCE VALUES PER PHOTO



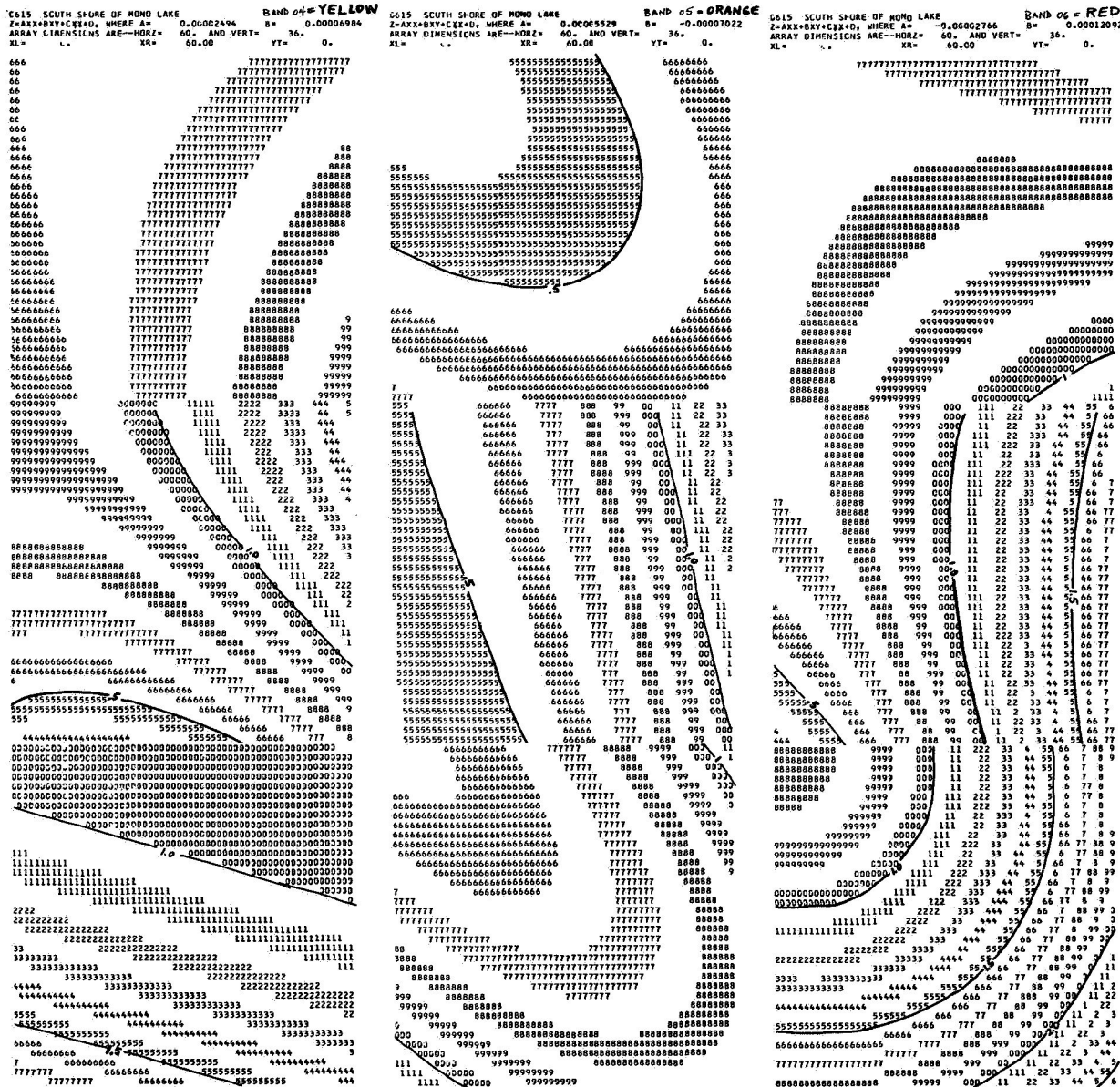
BAND 1

BAND 2

BAND 3

Figure 11

SECOND DEGREE TREND SURFACES FITTED TO 45
MULTIBAND FILM TRANSMITTANCE VALUES PER PHOTO



BAND 4

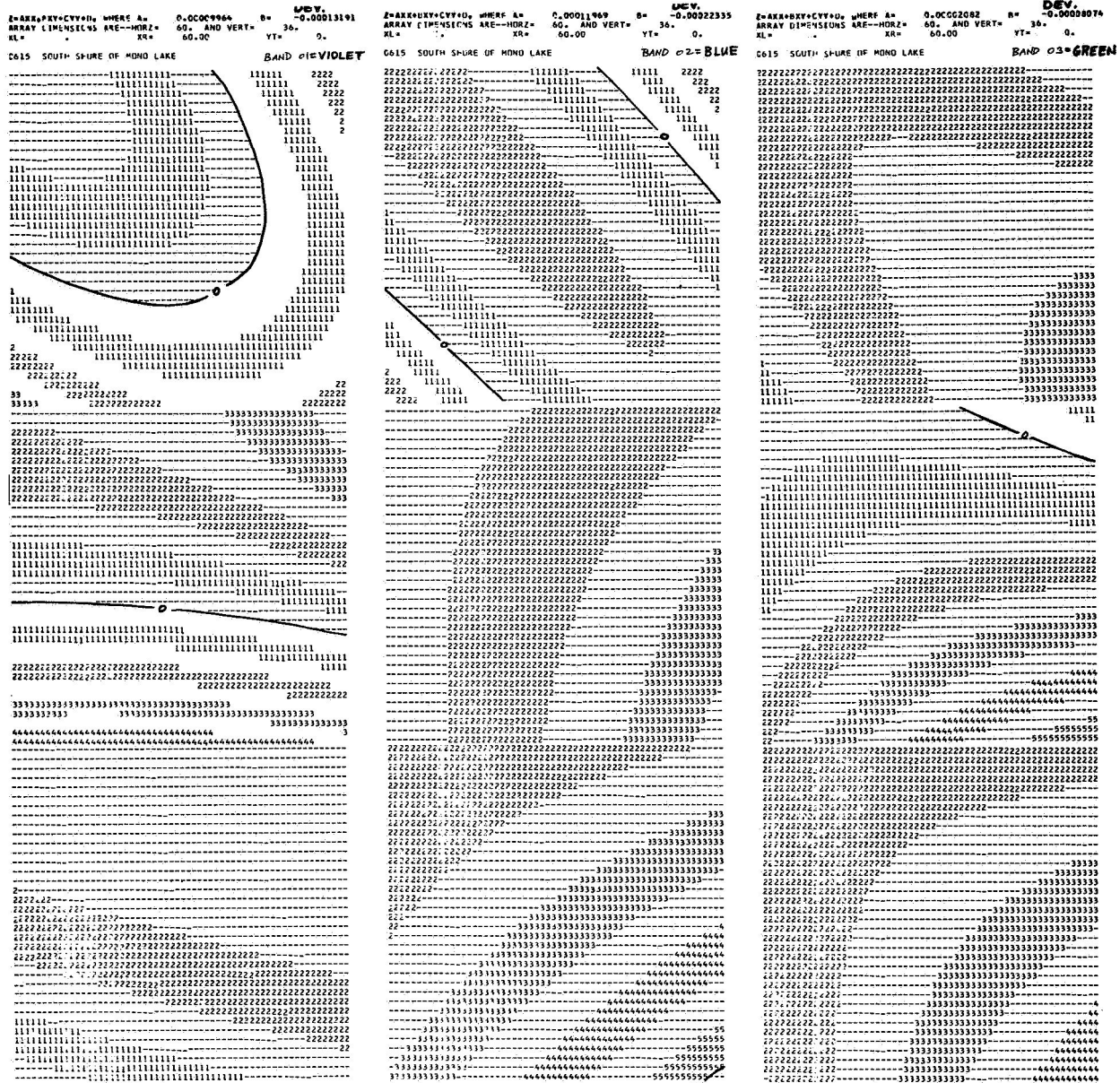
BAND 5

BAND 6

Figure 12

SECOND DEGREE TREND SURFACES FITTED TO DEVIATION
FROM AVERAGE MULTIBAND FILM TRANSMITTANCE

(45 VALUES PER PHOTO)



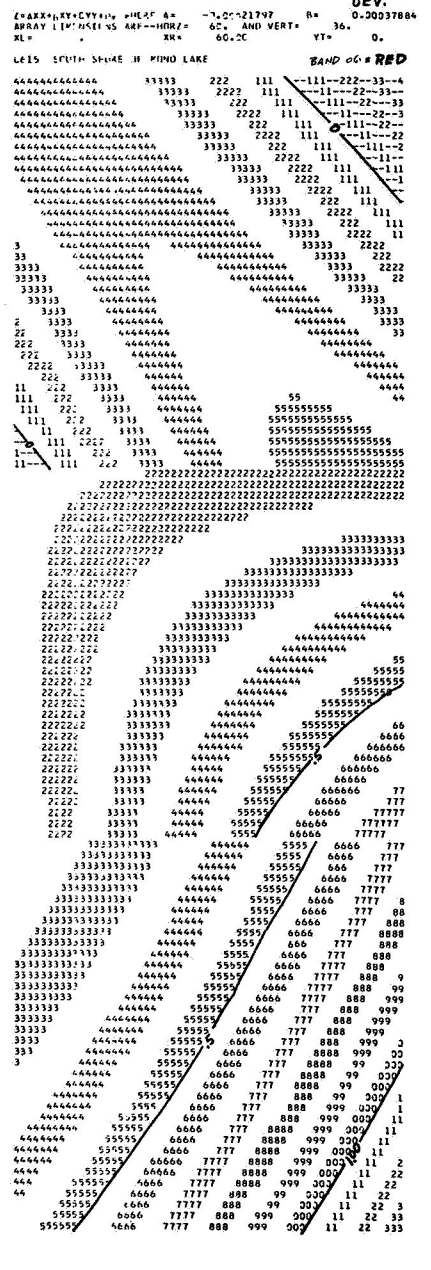
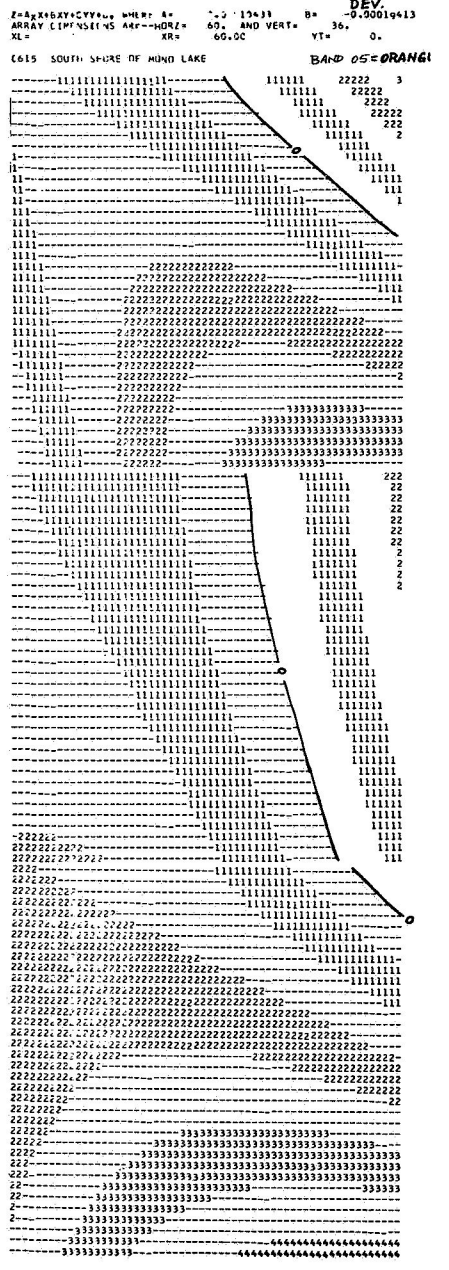
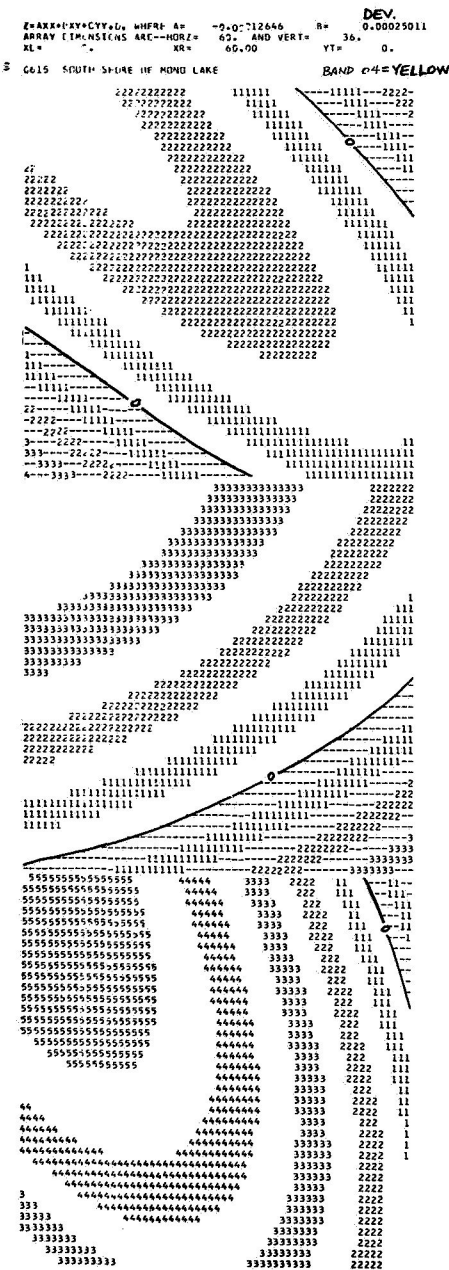
BAND 1

BAND 2

BAND 3

Figure 13

SECOND DEGREE TREND SURFACES FITTED TO DEVIATION
FROM AVERAGE MULTIBAND FILM TRANSMITTANCE
(45 VALUES PER PHOTO)



BAND 4

BAND 5

BAND 6

Figure 14

types, then the mean of the value of each of the spectral bands within each type would provide an "ideal" set of numbers for every terrain type. Having made this assumption, we might also assume that the group of values which are "closest" to an unknown spectrum would become the most nearly correct terrain type for this unknown spectrum. One method of comparing the spectrum to be classified with the several possible spectrum is by means of an extension of the Pythagorean theorem into the sixth dimension.* In this case, each dimension is the value of the transmittance of each band. For an example in the second dimension, the distance from an unknown spectrum with transmittance values for each of two bands B1 and B2 can be compared with a mean spectrum for rhyolite, R1 and R2, by computing the distance D between the two. Here, $D^2 = (B1 - R1)^2 + (B2 - R2)^2$ with D being the square-root of the right-hand side of the equation.

The distance thus obtained can be used with a value in a third dimension, and this in turn may be used to find distances in higher dimensions. (This calculation can be done very simply by means of a DO loop in FORTRAN IV, and the minimum distance among several terrain types chosen either by means of an IF statement or the MIN function). The effectiveness of this method in classifying terrain types was not as good as expected, largely because of the extreme variability of many of the band transmission values. If the amount of variation in the spectra being classified

* See page 53 of the Appendix for additional information.

is great, then this may cause many of them to be incorrectly identified. A large variation of one band relative to another is the greatest weakness of this system. In order that this variability could be accounted for, statistical methods were resorted to in the last method to be investigated.

3. Statistical Models

The use of statistical methods was first suggested by a computer program which was used in connection with a statistics course (Geology 205) given by Dr. Paul Switzer in the Spring quarter, 1967, at Stanford. This was a program for Stepwise Discriminant Analysis (BMD07M) and was written by the UCLA Health Science Computing Facility. The following method requires means and standard deviations of film transmittances for each of the terrain types to be calculated as they are used in classification. The results of this general program indicated a maximum accuracy, from a choice of three groups, of 80% correct identification for an unknown spectrum obtained from six multiband transparencies. To be more specific a new FORTRAN IV program to produce a classification map was written for this project.

As each of the six bands may be considered a variable and each of the three possible terrain types may be considered a group, the method of multivariate analysis can be used and the resulting classification checked against actual conditions. First, a "perfect" spectrum for each terrain group is decided upon--the means of each group--and is calculated thus:

$$H^R = (h_1^R, h_2^R, \dots, h_6^R); h_i^R = \frac{1}{n_R} \sum_{j=1}^n (b_{ij}^R); \text{ e.g. } h_1^R = \frac{1}{18} \sum_{j=1}^{18} (b_{1j}^R) = 0.65167$$

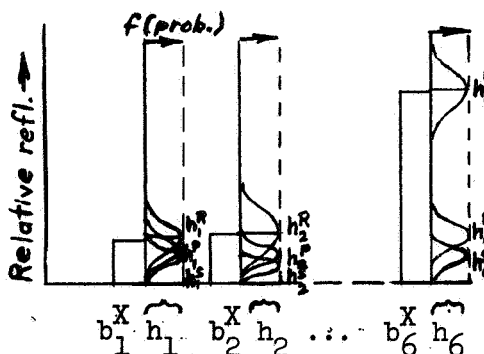
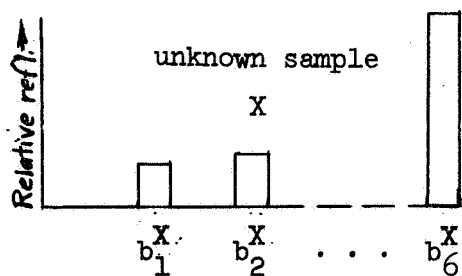
$$H^S = (h_1^S, h_2^S, \dots, h_6^S); h_i^S = \frac{1}{n_S} \sum_{j=1}^n (b_{ij}^R); \text{ e.g. } h_1^S = \frac{1}{15} \sum_{j=1}^{15} (b_{1j}^S) = 0.4000$$

$$H^P = (h_1^P, h_2^P, \dots, h_6^P); h_i^P = \frac{1}{n_P} \sum_{j=1}^n (b_{ij}^R); \text{ e.g. } h_1^P = \frac{1}{46} \sum_{j=1}^{46} (b_{1j}^P) = 0.54696$$

where H^R, H^S, H^P are each a set of 6 band means for rhyolite, shadow, or pumice; h_i^R, h_i^S, h_i^P are the means of an individual band i ; n_R, n_S, n_P are the number of spectra in each of the three groups; and $b_{ij}^R, b_{ij}^S, b_{ij}^P$ are the relative reflectance values of an individual band in an individual spectrum j .

By using these means as the most probable estimates we may also assume that any variation within an individual band is not completely random, but may be approximated by a Gaussian function of some form. This implies that the three distributions in a given band are each represented by a characteristically shaped normal curve with its peak at the mean of that band. Unknown spectra may then be compared to these ideal spectra and the most probable group decided upon (Fig. 15).

Comparison illustration:



Spectrum means:

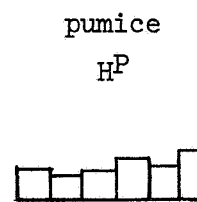
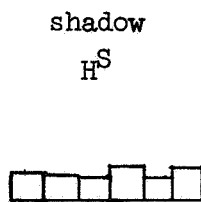
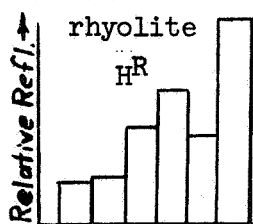


Figure 15

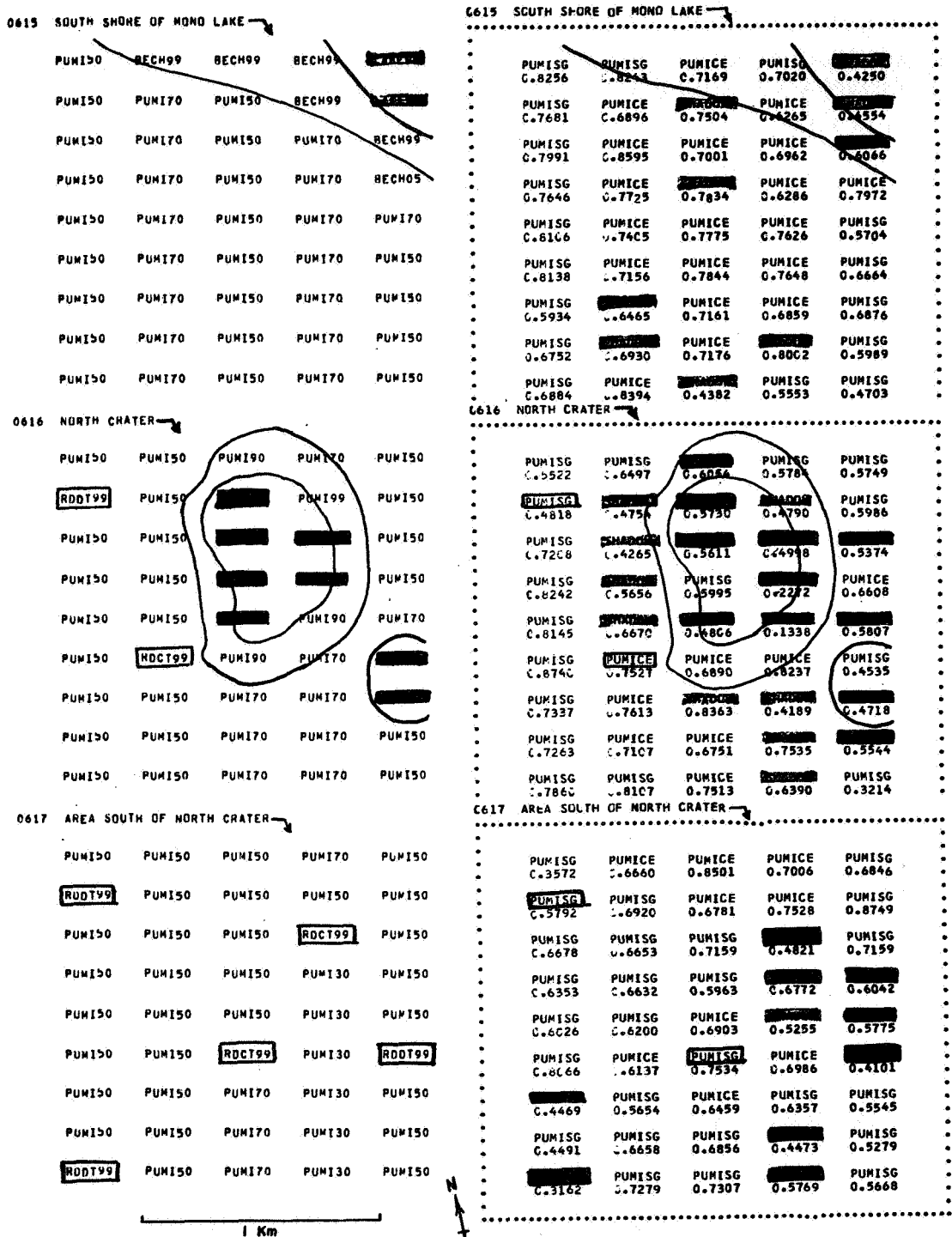
This was done for each known group and for each of the six bands in sequence. Next, the probabilities were totalled and the group with the largest probability was assigned to the unknown spectrum. The value of this probability was computed by dividing the sum of the total probability of the six bands by six and was printed beneath the code name of the group in its proper location relative to the photo.

The three terrain types used were areas of rhyolite obsidian (RHYOLT), rhyolite pumice with little or no sage (PUMICE), and areas of shadow falling on any terrain (SHADOW). The results of this method were about 50% correct using these three groups, however, a fourth group was added which consisted of areas of moderately vegetated pumice (PUMISG), and the original program was modified to consider this group with the others (Fig. 16). The results were more encouraging, with five of the six unknown terrain types in the Crater correctly classified. There were, however, few probability values greater than 0.80, but this was acceptable. Even dark lake water was "correctly" classified as shadow though only having a probability of 0.425.

Other possible terrain types were included in subsequent versions of this new program, but did not show promising results, because those groups had standard deviations (i.e. broad "vague" distribution curves) and were always selected. Another approach involved substituting the values of DEV obtained from a previous program for the values of transmittance. The results of this DEV program were also unsatisfactory, as groups with little variation between bands were confused with each other. This was the case with lake water, which was

MULTIVARIATE CLASSIFICATION OF TERRAIN TYPES NEAR MONO LAKE, CALIFORNIA

DATA FROM SIX VISUAL BANDS OF MULTIBAND PHOTOGRAPHY



TERRAIN TYPES FOUND IN AREA:

- RHY099 = RHYOLITE WITH LITTLE OR NO VEGETATION
- PUM199 = PUMICE WITH LITTLE OR NO VEGETATION
- PUM170 = PUMICE WITH SMALL PERCENTAGE OF VEGETATION
- PUM150 = PUMICE AND VEGETATION (SAME) ABOUT EQUAL
- PUM130 = HEAVILY VEGETATED PUMICE
- BECH099 = BEACH DEPOSITS WITH LITTLE OR NO VEGETATION
- LAKE099 = LAKE WATER
- RDDT99 = DIRT ROAD OF PUMICE MATERIAL

TERRAIN TYPES CONSIDERED IN ANALYSIS:

- RHYOLT = RHYOLITE OBSIDIAN
- PUMICE = PUMICE WITH LITTLE OR NO VEGETATION
- PUM15G = PUMICE MODERATELY VEGETATED WITH SAGE
- SHADOW = SHADED AREAS OF EITHER RHYOLITE OR PUMICE

Figure 16

classified as being pumice. This occurs because the flat, but high reflectance of pumice is compared to the flat, but low reflectance of lake water.

V. CONCLUSIONS

In any attempt at remote sensing data interpretation it must be born in mind that the final classification scheme can only be as reliable as the basic data itself. Any future experiments must be designed so that as many variables as possible can be evaluated. These variables include field reflectance data, major and minor topographic features, soil and vegetation distribution, atmospheric attenuation, aircraft altitude, camera film and filter spectral responses, solar illumination changes with sun angle, and as many other parameters as possible which might influence data obtained during the aircraft overflight.

Assuming all these conditions to be constant for a given time and area, then quantitative multiband data can be obtained by use of recording densitometers, with future data reduction by means of flying-spot densitometers being the next logical step. Secondly, computerized analysis of data provides a rapid means of testing and perfecting methods of classification and discrimination of terrain types, particularly overall geology, from multiband regions under study. It appears that of the three methods of analysis investigated, that statistical multivariate analysis holds the most promise of being an ideal solution for pattern recognition purposes, especially if attempts to classify other areas of similar geology by multiband photography are attempted.

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APPENDIXPROGRAM DESCRIPTIONS*Data Cards:

Each data card used in the following programs contained a coded description of the terrain in the area sampled, the six transmittance values for each of the "spectra", and roll, frame and photograph coordinate numbers. The following examples illustrate this format.

<u>Card column</u>	<u>Description</u>	<u>Example</u>
1-6	Geologic formation	QR = Quaternary rhyolite
7-12	Rock type and percent	RHYØ70 = 70% rhyolite
13-18	Vegetation percent	SAGE25 = 25% sagebrush
19-24	Secondary veg. percent	LICH05 = 5% lichen
25-30	Water type and percent	LAKE99 = entirely lake
31-36	Miscellaneous terrain	RDDT90 = 90% dirt road
37-60	Transmittance values for six visual bands (6F4.2)	Format 6F4.2
61-72	Reserved for IR data	Format 3F4.2
73-76	Exposure number	0615
77-80	X- and Y-coordinates	X = 30, Y = 30

These data cards were then arranged into five groups each having the same X-value and increasing Y-values and a title card placed before each group of 45 spectra. This title card allows a maximum of 72 characters and includes the photograph number and a brief description of the photo area. This is illustrated by the following listing from three digitized photos and one photo profile.

* Written for Stanford System IBM 7090 Computer.

0617 AREA SOUTH OF NORTH CRATER									
QP	PUMI50SAGE50		110	185	90	150	70	110	0617 1 1
QP		RDDT	70	80	105	110	60	80	0617 1 2
QP	PUMI50SAGE50		70	80	55	120	75	110	0617 1 3
QP	PUMI50SAGE50		55	80	60	120	80	110	0617 1 4
QP	PUMI50SAGE50		65	65	70	150	70	170	0617 1 5
QP	PUMI50SAGE50		70	70	70	90	70	100	0617 1 6
QP	PUMI50SAGE50		90	60	100	20	110	175	0617 1 7
QP	PUMI50SAGE50		80	80	90	120	75	140	0617 1 8
QP		RDDT99	140	150	190	280	170	250	0617 1 9
QP	PUMI50SAGE50		50	55	50	110	45	80	0617 2 1
QP	PUMI50SAGE50		60	60	70	145	50	95	0617 2 2
QP	PUMI50SAGE50		60	60	70	160	70	80	0617 2 3
QP	PUMI50SAGE50		50	50	55	130	70	110	0617 2 4
QP	PUMI50SAGE50		55	60	50	130	65	170	0617 2 5
QP	PUMI50SAGE50		60	40	45	130	45	100	0617 2 6
QP	PUMI50SAGE50		60	50	40	130	35	115	0617 2 7
QP	PUMI50SAGE50		65	50	45	130	50	90	0617 2 8
QP	PUMI50SAGE50		65	45	50	100	45	90	0617 2 9
QP	PUMI50SAGE50		55	45	50	80	40	70	0617 3 1
QP	PUMI50SAGE50		50	50	40	95	40	70	0617 3 2
QP	PUMI50SAGE50		50	50	55	110	55	110	0617 3 3
QP	PUMI50SAGE50		50	50	35	115	55	120	0617 3 4
QP	PUMI50SAGE50		55	40	50	90	45	125	0617 3 5
QP		RDSH99	60	50	70	100	55	125	0617 3 6
QP	PUMI70SAGE30		65	40	35	80	40	85	0617 3 7
QP	PUMI70SAGE30		70	70	60	135	55	125	0617 3 8
QP	PUMI70SAGE30		70	50	50	120	55	100	0617 3 9
QP	PUMI70SAGE30		40	40	40	65	40	80	0617 4 1
QP	PUMI50SAGE50		50	50	60	80	50	70	0617 4 2
QP		RDCT99	60	100	135	110	160	355	0617 4 3
QP	PUMI30SAGE70		60	70	70	125	115	125	0617 4 4
QP	PUMI30SAGE70		40	35	30	70	70	100	0617 4 5
QP	PUMI30SAGE70		50	45	40	80	50	115	0617 4 6
QP	PUMI30SAGE70		70	45	60	80	40	150	0617 4 7
QP	PUMI30SAGE70		50	80	110	80	80	140	0617 4 8
QP	PUMI30SAGE70		60	60	55	160	90	225	0617 4 9
QP	PUMI50SAGE50		75	75	55	100	75	125	0617 5 1
QP	PUMI50SAGE50		65	60	60	70	60	95	0617 5 2
QP	PUMI50SAGE50		50	55	50	75	70	95	0617 5 3
QP	PUMI50SAGE50		75	70	70	140	110	190	0617 5 4
QP	PUMI50SAGE50		70	60	70	155	95	240	0617 5 5
		RDDT99	100	105	235	195	140	345	0617 5 6
QP	PUMI50SAGE50		90	65	70	105	75	170	0617 5 7
QP	PUMI50SAGE50		95	70	60	115	55	185	0617 5 8
QP	PUMI50SAGE50		100	65	55	110	45	135	0617 5 9

0616 PROFILE ON N-S CENTER LINE

QP	PUMI	SAGE	62	65	95	105	88	90	06163010
QP	PUMI	SAGE	55	57	75	95	61	67	06163011
QR	RHYO		27	18	10	08	10	00	06163012
QR	RHYO		30	19	08	15	15	05	06163013
QR	RHYO		55	58	70	40	55	105	06163014
QR	RHYO		68	83	187	180	110	380	06163015
QR	RHYO		42	40	58	95	55	140	06163016
QR	RHYO		85	102	210	300	172	402	06163017
QR	RHYO		65	78	265	210	120	290	06163018
QR	RHYO		78	85	345	325	224	570	06163019
QR	RHYO		55	60	135	165	158	157	06163020
QR	RHYO		35	30	80	55	61	40	06163021
QR	RHYO		60	66	121	150	107	211	06163022
QR	RHYO		60	57	102	175	90	125	06163023
QR	RHYO		55	42	65	175	100	210	06163024
QR	RHYO		45	38	40	100	63	140	06163025
QR	RHYO		35	23	07	35	30	55	06163026
QR	RHYO		65	60	145	205	183	502	06163027
QR	RHYO		25	20	06	22	25	22	06163028
QR	RHYO		37	21	07	22	22	30	06163029
QR	RHYO		57	77	42	127	60	240	06163030
QR	RHYO		25	10	20	20	23	17	06163031
QP	PUMI	SAGE	50	48	48	95	46	73	06163032
QP	PUMI	SAGE	40	21	37	60	35	47	06163033
QP	PUMI	SAGE	35	16	25	40	25	40	06163034
QP	PUMI	SAGE	65	45	67	87	60	81	06163035
QP	PUMI	SAGE	70	50	60	107	60	88	06163036
QP	PUMI	SAGE	57	34	57	75	50	87	06163037
QP	PUMI	SAGE	60	30	30	60	40	48	06163038
QP	PUMI	SAGE	55	28	35	45	35	72	06163039
QP	PUMI	SAGE	55	31	33	45	35	50	06163040
QP	PUMI	SAGE	55	32	32	50	36	42	06163041
QP	PUMI	SAGE	53	36	45	60	41	63	06163042
QP	PUMI	SAGE	54	36	41	55	38	50	06163043
QP	PUMI	SAGE	55	38	45	55	40	50	06163044
QP	PUMI	SAGE	57	40	43	55	40	50	06163045
QP	PUMI	SAGE	57	35	48	55	48	62	06163046
ROADDT			RD	RD	RD	RD	RD	RD	06163047
ROADDT			RD	RD	RD	RD	RD	RD	06163048
QP	PUMI	SAGE	60	43	56	65	43	64	06163049
QP	PUMI	SAGE	60	40	40	50	40	70	06163050

Conversion of Data to "Map" Form

In order for some of the following trend surface programs to work in a less complex manner, the original data was "rearranged" and punched on cards. This was accomplished by the following program, which inputs the title card and profile-type data on 45 cards and outputs six sets of band transmittances and the appropriate title card. In essence, this provides six sets of numbers for the individual wavelengths and arranges them from Band 1 to Band 6. Also the format for each card changes from 36X, 6F4.2 to 5F10.2; or from a spectrum to a line of five equal Y-values.

This data rearrangement program is shown in the following listing, along with a listing of the output.

SJOB S060 IBJOB 2 800 BALLEW GARY 233 BANDPASS MAPS

\$IBJOB

\$IBFTC MAIN NODECK

C ACTUAL BAND-PASS MAPS

DIMENSION TITLE(12), B(5,9,6)

99 WRITE(6,98)

98 FORMAT(1H1,1H)

READ(5,1) (TITLE(K),K=1,12)

1 FORMAT (12A6)

READ(5,3)((B(L,J,I), I=1,6), J=1,9),L=1,5)

3 FORMAT(36X,6F4,2)

DO 100 I=1,6

WRITE(6,10) (TITLE(K),K=1,12),I

10 FORMAT(1HP,12A6, 4HBAND,13/)

DO 100 J=1,9

WRITE(6,4)(B(L,J,I),L=1,5)

4 FORMAT(1HP,5F10,2/)

100 CONTINUE

GO TO 99

RETURN

END

732

PRECEDES B5000 AND IBJOB DATA CARDS

0615	SOUTH SHORE OF MOND LAKE					BAND 1
	0.65	0.55	0.45	0.60	0.30	
	0.55	0.45	0.30	0.35	0.30	
	0.60	0.50	0.40	0.60	0.70	
	0.65	0.50	0.40	0.45	0.55	
	0.70	0.55	0.50	0.50	0.90	
	0.70	0.50	0.50	0.55	0.70	
	0.80	0.60	0.50	0.60	0.85	
	0.80	0.50	0.50	0.40	0.90	
	0.75	0.50	0.25	0.70	0.90	
0615	SOUTH SHORE OF MOND LAKE					BAND 2
	0.65	0.65	0.55	0.85	0.30	
	0.65	0.40	0.35	0.50	0.25	
	0.60	0.45	0.40	0.60	0.70	
	0.45	0.35	0.35	0.40	0.45	
	0.55	0.35	0.40	0.35	0.70	
	0.50	0.45	0.35	0.35	0.50	
	0.55	0.40	0.45	0.40	0.55	
	0.75	0.40	0.50	0.30	0.75	
	0.75	0.50	0.30	0.95	0.85	
0615	SOUTH SHORE OF MOND LAKE					BAND 3
	0.80	0.60	0.45	0.65	0.05	
	0.55	0.30	0.25	0.35	0.45	
	0.45	0.45	0.40	0.30	0.75	
	0.75	0.40	0.35	0.35	0.25	
	0.55	0.40	0.40	0.35	0.65	
	0.55	0.40	0.50	0.40	0.75	
	0.80	0.40	0.40	0.45	0.55	
	0.60	0.35	0.40	0.25	0.50	
	0.70	0.45	0.10	0.70	0.85	
0615	SOUTH SHORE OF MOND LAKE					BAND 4
	1.00	0.90	0.65	1.00	0.07	
	0.80	0.80	0.50	0.70	0.05	
	0.75	0.70	0.55	0.70	1.50	
	0.75	0.65	0.55	0.70	0.80	
	0.80	0.60	0.70	0.70	1.30	
	0.85	0.60	0.70	0.60	0.70	
	0.50	0.50	0.85	0.40	0.70	
	0.65	0.45	0.60	0.35	0.60	
	0.85	0.65	0.20	1.20	1.15	
0615	SOUTH SHORE OF MOND LAKE					BAND 5
	0.65	0.60	0.50	0.80	0.27	
	0.65	0.45	0.35	0.60	0.25	
	0.60	0.50	0.45	0.50	0.90	
	0.60	0.45	0.35	0.45	0.45	
	0.50	0.40	0.45	0.50	0.60	
	0.60	0.35	0.40	0.50	0.55	
	0.60	0.35	0.35	0.40	0.60	
	0.70	0.30	0.40	0.25	0.60	
	0.80	0.50	0.20	0.90	0.80	
0615	SOUTH SHORE OF MOND LAKE					BAND 6
	1.10	0.85	0.60	1.00	0.07	
	0.85	0.70	0.50	0.70	0.05	
	1.10	0.85	0.70	0.70	1.30	
	1.10	0.80	0.75	1.20	0.75	
	1.05	0.85	0.90	0.80	1.40	
	0.90	0.60	0.70	0.90	0.75	
	0.75	0.55	0.80	0.85	0.80	
	1.10	0.60	0.85	0.55	1.10	
	1.00	0.70	0.25	1.10	1.10	

0616	NORTH CRATER				BAND 1
	0.85	0.60	0.60	0.65	0.70
	0.70	0.40	0.60	0.25	0.60
	0.60	0.25	0.50	0.50	0.65
	0.60	0.30	0.40	0.75	0.45
	0.60	0.40	0.50	0.15	0.80
	0.70	0.50	0.50	0.45	1.00
	0.90	0.55	0.50	0.30	1.00
	0.85	0.55	0.50	0.40	1.40
	0.85	0.75	0.55	0.55	0.90
0616	NORTH CRATER				BAND 2
	0.70	0.50	0.70	0.70	0.70
	0.95	0.40	0.85	0.25	0.75
	0.70	0.20	0.60	0.55	0.85
	0.65	0.20	0.40	1.20	0.40
	0.60	0.40	0.80	0.10	0.70
	0.55	0.35	0.40	0.40	0.60
	0.60	0.35	0.30	0.20	0.70
	0.65	0.40	0.40	0.30	1.00
	0.60	0.60	0.40	0.40	0.60
0616	NORTH CRATER				BAND 3
	0.75	0.60	1.20	1.00	0.90
	1.10	1.00	1.90	0.20	0.85
	0.55	0.15	0.90	1.00	0.85
	0.50	0.17	0.35	1.55	0.35
	0.55	0.35	0.30	0.	0.95
	0.60	0.35	0.70	0.50	0.75
	0.50	0.40	0.35	0.05	1.05
	0.55	0.40	0.45	0.20	1.20
	0.60	0.65	0.40	0.35	0.95
0616	NORTH CRATER				BAND 4
	0.80	0.60	1.00	0.70	1.00
	1.40	1.10	1.75	0.20	1.00
	1.10	0.20	2.05	1.10	1.50
	0.95	0.30	0.95	3.10	0.60
	0.80	0.60	1.20	0.	1.40
	0.85	0.70	0.65	0.70	1.40
	0.80	0.60	0.45	0.15	1.10
	1.00	0.50	0.50	0.35	1.90
	0.80	0.70	0.45	0.35	0.50
0616	NORTH CRATER				BAND 5
	0.50	0.55	0.90	0.80	0.95
	0.70	0.45	1.10	0.25	0.95
	0.75	0.25	1.20	1.05	1.20
	0.60	0.30	0.60	2.20	0.55
	0.55	0.50	0.60	0.05	1.05
	0.55	0.45	0.60	0.50	0.70
	0.60	0.45	0.35	0.25	0.85
	0.60	0.45	0.40	0.30	0.85
	0.65	0.60	0.40	0.40	2.00
0616	NORTH CRATER				BAND 6
	0.60	0.50	0.90	0.80	0.80
	1.10	0.50	1.70	0.10	0.95
	0.90	0.20	1.50	1.50	1.25
	0.90	0.30	1.20	5.30	0.60
	0.90	0.60	2.40	0.05	1.80
	1.05	0.80	0.80	0.70	1.95
	0.95	0.60	0.55	0.25	1.25
	0.80	0.45	0.50	0.30	2.45
	0.90	1.00	0.70	0.50	1.60

0617	AREA SOUTH OF NORTH CRATER					BAND 1
	1.10	0.50	0.55	0.40	0.75	
	0.70	0.60	0.50	0.50	0.65	
	0.70	0.60	0.50	0.60	0.50	
	0.55	0.50	0.50	0.60	0.75	
	0.65	0.55	0.55	0.40	0.70	
	0.70	0.60	0.60	0.50	1.00	
	0.90	0.60	0.65	0.70	0.90	
	0.80	0.65	0.70	0.50	0.95	
	1.40	0.65	0.70	0.60	1.00	
0617	AREA SOUTH OF NORTH CRATER					BAND 2
	1.85	0.55	0.45	0.40	0.75	
	0.80	0.60	0.50	0.50	0.60	
	0.80	0.60	0.50	1.00	0.55	
	0.80	0.50	0.50	0.70	0.70	
	0.65	0.60	0.40	0.35	0.60	
	0.70	0.40	0.50	0.45	1.05	
	0.60	0.50	0.40	0.45	0.65	
	0.80	0.50	0.70	0.80	0.70	
	1.50	0.45	0.50	0.60	0.65	
0617	AREA SOUTH OF NORTH CRATER					BAND 3
	0.90	0.50	0.50	0.40	0.55	
	1.05	0.70	0.40	0.60	0.60	
	0.55	0.70	0.55	1.35	0.50	
	0.60	0.55	0.35	0.70	0.70	
	0.70	0.50	0.50	0.30	0.70	
	0.70	0.45	0.70	0.40	2.35	
	1.00	0.40	0.35	0.60	0.70	
	0.90	0.45	0.60	1.10	0.60	
	1.90	0.50	0.50	0.55	0.55	
0617	AREA SOUTH OF NORTH CRATER					BAND 4
	1.50	1.10	0.80	0.65	1.00	
	1.10	1.45	0.95	0.80	0.70	
	1.20	1.60	1.10	1.10	0.75	
	1.20	1.30	1.15	1.25	1.40	
	1.50	1.30	0.90	0.70	1.55	
	0.90	1.30	1.00	0.80	1.95	
	0.20	1.30	0.80	0.80	1.05	
	1.20	1.30	1.35	0.80	1.15	
	2.80	1.00	1.20	1.60	1.10	
0617	AREA SOUTH OF NORTH CRATER					BAND 5
	0.70	0.45	0.40	0.40	0.75	
	0.60	0.50	0.40	0.50	0.60	
	0.75	0.70	0.55	1.60	0.70	
	0.80	0.70	0.55	1.15	1.10	
	0.70	0.65	0.45	0.70	0.95	
	0.70	0.45	0.55	0.50	1.40	
	1.10	0.35	0.40	0.40	0.75	
	0.75	0.50	0.55	0.80	0.55	
	1.70	0.45	0.55	0.90	0.45	
0617	AREA SOUTH OF NORTH CRATER					BAND 6
	1.10	0.80	0.70	0.80	1.25	
	0.80	0.95	0.70	0.70	0.95	
	1.10	0.80	1.10	3.55	0.95	
	1.10	1.10	1.20	1.25	1.90	
	1.70	1.70	1.25	1.00	2.40	
	1.00	1.00	1.25	1.15	3.45	
	1.75	1.15	0.85	1.50	1.70	
	1.40	0.90	1.25	1.40	1.85	
	2.50	0.90	1.00	2.25	1.35	

This program provides a method of obtaining a trend surface of each of the colors or bands used in this analysis. This is accomplished by means of a trend surface contour plotting program which has been spliced into another program which solves for the coefficients of the second-order equation. The input for this consists of the six sets of the 45 values of transmittance obtained from each photo and the corresponding title cards. The format of each card is the 5F10.2 format previously described. In addition, the first data card must contain the parameters of the plotted trend surface dimensions.

These output parameters include the horizontal dimension HORZ, vertical dimension VERT, both in terms of lines and spaces; the left, right, top, and bottom dimensions XL, XR, YT, and YB, respectively; the reference contour REF; and the contour interval CON. The output consists of a pattern of alternating blanks and numbers from zero to one for positive numbers, and alternating minus signs and numbers for negative numbers. For best results, the contour interval should be half that of the expected contour interval if whole numbers are used.

The equation used is $Z = AX^2 + BXY + CY^2 + D$, where Z is the transmittance of the trend surface at coordinates X and Y in millimeters from the upper left-hand corner of the photograph, and A, B, C and D are the coefficients which are determined by the SOLVEX subroutine of the program. This subroutine essentially solves the 4 by 4 matrix generated by the least-squared method employed here. This sub-routine may be obtained from the computation center's library of programs under the name SOLVE, program number one, which can solve up to 20 by 20 matrices.

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$IBJOB          MAP
$IBFTC MAIN
C CONTOUR FROM CARDS IN COORDINATE FORMAT USING SOLVEX
  INTEGER PLUS,MINU,CV
  DIMENSION WC( 60),CV( 60),PLUS( 36),MINU( 36),
  1          AA(4,4),BB(4),COEF(4),X(45,3),Z(9,5),TITLE(12)
  DATA(PLUS(I),I=1,20)/1H,1H1,1H,1H2,1H,1H3,1H,1H4,1H,1H5,1H,1
  1H6,1H,1H7,1H,1H8,1H,1H9,1H,1H0/, (MINU(I),I=1,20)/1H-,1H1,1H-,1
  2H2,1H-,1H3,1H-,1H4,1H-,1H5,1H-,1H6,1H-,1H7,1H-,1H8,1H-,1H9,1H-,1H
  30/
101 READ(5,99)          HORZ,VERT,XL,XR,YT,YB,REF,CON
  99 FORMAT (28X,F3.0,F5.0,6F7.2)
200 READ(5,7) (TITLE(K),K=1,12)
  7 FORMAT(12A6)
  READ(5,8) ((Z(I,J),J=1,5),I=1,9)
  8 FORMAT(5F10.3)
  X4=0.0
  X3Y=0.0
  X2Y2=0.0
  XY3=0.0
  Y4=0.0
  XY=0.0
  Y2=0.0
  X2=0.0
  ZX2=0.0
  ZXY=0.0
  ZY2=0.0
  Z=0.0
  N=0
  X(I,2)=5.0
  DO 100 L=1,9
  X(I,2)=X(I,2)+5.0
  X(I,1)=0.0
  DO 100 M=1,5
  X(I,1)=X(I,1)+10.0
  X(I,3)=Z(L,M)
  X4=X4+X(I,1)*X(I,1)*X(I,1)*X(I,1)
  X3Y=X3Y+X(I,1)*X(I,1)*X(I,1)*X(I,2)
  X2Y2=X2Y2+X(I,1)*X(I,1)*X(I,2)*X(I,2)
  XY3=XY3+X(I,1)*X(I,2)*X(I,2)*X(I,2)
  Y4=Y4+X(I,2)*X(I,2)*X(I,2)*X(I,2)
  XY=XY+X(I,1)*X(I,2)
  Y2=Y2+X(I,2)*X(I,2)
  X2=X2+X(I,1)*X(I,1)
  ZX2=ZX2+X(I,3)*X(I,1)*X(I,1)
  ZXY=ZXY+X(I,3)*X(I,1)*X(I,2)
  ZY2=ZY2+X(I,3)*X(I,2)*X(I,2)
  Z=Z+X(I,3)
  N=N+1
100 CONTINUE
  RN=N
  AA(1,1)=X4
  AA(1,2)=XY3
  AA(1,3)=X2Y2
  AA(1,4)=X2

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AA(2,1)=X3Y
AA(2,2)=X2Y2
AA(2,3)=XY3
AA(2,4)=XY
AA(3,1)=X2Y2
AA(3,2)=XY3
AA(3,3)=Y4
AA(3,4)=Y2
AA(4,1)=X2
AA(4,2)=XY
AA(4,3)=Y2
AA(4,4)=RN
BB(1)=ZX2
BB(2)=ZXY
BB(3)=ZY2
BB(4)=Z
CALL SOLVE (4, AA, BB, 1, 0, 0, 10, COEF, IT)
A=COEF(1)
B=COEF(2)
C=COEF(3)
D=COEF(4)
WRITE(6,98)A,B,C,D
98 FORMAT(1H1,25HZ=AXX+BXY+CY+D, WHERE A=, F15.8,5X,2H8=,F15.8,5X,
12HC=,F15.8, 5X, 2HD=,F15.8 )
WRITE(6,97)HORZ,VERT
97 FORMAT(1H ,27HARRAY DIMENSIONS ARE--HORZ=, F6.0,11H AND VERT=,F8.
10)
WRITE(6,96)XL,XR,YT,YB,REF,CON
96 FORMAT(1H ,3HXL=,F10.2,10X,3HXR=,F10.2,10X,3HYT=,F10.2,10X,3HYB=,
1F10.2,10X,4HREF=,F10.2,10X,4HCON=,F10.2//)
WRITE(6,5) (TITLE(K),K=1,12)
5 FORMAT(1H ,12A6//)
DX=(XR-XL)/HORZ
DY=(YB-YT)/VERT
C1=(A*XL*XL)+(B*XL*YT)+(C*YT*YT)+D
C2=(2.*A*XL*DX)+(B*YT*DX)
C3=A*DX*DX
C4=(B*XL*DY)+(2.*C*YT*DY)
C5=C*DY*DY
C6=B*DX*DY
IHORZ=HORZ
IVERI=VERT
DO 1 I=1,IVERT
RI=I
C7=C1+(RI*C4)+(RI*RI*C5)
C8=(C6*RI)+C2
DO 2 J=1,IHORZ
RJ=J
W(J)=C7+(C8*RJ)+(C3*RJ*RJ)
IF(W(J).LT.REF) GO TO 3
XX=(W(J)-REF)/CON
IX=XX
IA=MUD(IX,20)
CV(J)=PLUS(IA+1)
GO TO 2

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3 YY=(HEF-W(J))/CON
  IY=YY
  IB=MUD(IY,20)
  CV(J)=MINU(IB+1)
2 CONTINUE
  WRITE(6,95)(CV(K),K=1,IHORZ)
95 FORMAT(1HT,132A1)
1 CONTINUE
  GO TO 200
102 RETURN
  END
$IBFTC SOLVEX

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	SLV4000		
CSOLVE	LINEAR EQUATION SOLVER WITH ITERATIVE IMPROVEMENT	VERSION IV	SLV4001
	SUBROUTINE SOLVE(NN,A,B,IN,EPS,ITMAX,X,IT)		SLV4002
C	SOLVES AX=B WHERE A IS NXN MATRIX AND B IS NX1 VECTOR		SLV4003
C	IN=		SLV4004
C	1 FOR FIRST ENTRY		SLV4005
C	2 FOR SUBSEQUENT ENTRIES WITH NEW B		SLV4006
C	3 TO RESTORE A AND B		SLV4007
C	EPS AND ITMAX ARE PARAMETERS IN THE ITERATION		SLV4008
C	IT=		SLV4009
C	-1 IF A IS SINGULAR		SLV4010
C	0 IF NOT CONVERGENT		SLV4011
C	NUMBER OF ITERATIONS IF CONVERGENT		SLV4012
C	CALLS MAP SUBROUTINES ILOG2,DOT,SDOT AND DAD		SLV4013
C			SLV4014
C	TO MODIFY DIMENSIONS, CHANGE THE NEXT 3 (NOT 2 BUT 3) CARDS.		SLV4015
	DIMENSION A(4, 4),B(4),X(4),AA(4, 4),DX(4),R(4),		SLV4016
	* Z(4),RM(4),IRP(4)		SLV4017
	MA= 4		SLV4018
C	MA MUST = DECLARED DIMENSION OF SYSTEM		SLV4019
	EQUIVALENCE(R,DX)		SLV4020
	GO TO (1000,2000,3000),IN		SLV4021
1000	N=NN		SLV4022
	NM1=N-1		SLV4023
	NP1=N+1		SLV4024
C			SLV4025
C	EQUILIBRATION		SLV4026
C			SLV4027
	DO 510 I=1,N		SLV4028
	KTOP=ILOG2(A(I,1))		SLV4029
	DO 503 J=2,N		SLV4030
503	KTOP=MAX0(KTOP,ILOG2(A(I,J)))		SLV4031
	RM(I)=2.0**(-KTOP)		SLV4032
	DO 509 J=1,N		SLV4033
509	A(I,J)=A(I,J)*RM(I)		SLV4034
510	CONTINUE		SLV4035
C			SLV4036
C	SAVE EQUILIBRATED DATA		SLV4037
C			SLV4038
	DO 548 I=1,N		SLV4039
	DO 548 J=1,N		SLV4040
548	AA(I,J)=A(I,J)		SLV4041
C			SLV4042

C	GAUSSIAN ELIMINATION WITH PARTIAL PIVOTING	SLV4043
C		SLV4044
	DO 99 M=1,NM1	SLV4045
	TOP=ABS (A(M,M))	SLV4046
	IMAX=M	SLV4047
	DO 12 I=M,N	SLV4048
	IF(TOP=ABS (A(I,M)))10,12,12	SLV4049
10	TOP=ABS (A(I,M))	SLV4050
	IMAX=I	SLV4051
12	CONTINUE	SLV4052
	IF(TOP)14,13,14	SLV4053
13	IT=-1	SLV4054
C	*SINGULAR*	SLV4055
	RETURN	SLV4056
14	IRP(M)=IMAX	SLV4057
23	IF(IMAX=M)29,29,24	SLV4058
24	DO 25 J=1,N	SLV4059
	TEMP=A(M,J)	SLV4060
	A(M,J)=A(IMAX,J)	SLV4061
25	A(IMAX,J)=TEMP	SLV4062
29	MP1=M+1	SLV4063
	DO 33 I=MP1,N	SLV4064
	EM=A(I,M)/A(M,M)	SLV4065
	A(I,M)=EM	SLV4066
	IF(EM)31,33,31	SLV4067
31	DO 32 J=MP1,N	SLV4068
32	A(I,J)=A(I,J)-A(M,J)*EM	SLV4069
33	CONTINUE	SLV4070
99	CONTINUE	SLV4071
	IRP(N)=N	SLV4072
	IF (A(N,N))120,113,120	SLV4073
113	IT=-1	SLV4074
	RETURN	SLV4075
120	CONTINUE	SLV4076
C	STORAGE FOR A NOW CONTAINS TRIANGULAR L AND U SO THAT (L+I)*U=A	SLV4077
C		SLV4078
C	DUPLICATE INTERCHANGES IN DATA	SLV4079
C		SLV4080
	DO 229 I=1,N	SLV4081
	IP=IRP(I)	SLV4082
	IF(I=IP)221,229,221	SLV4083
221	DO 222 J=1,N	SLV4084
	TEMP=AA(I,J)	SLV4085
	AA(I,J)=AA(IP,J)	SLV4086
222	AA(IP,J)=TEMP	SLV4087
229	CONTINUE	SLV4088
C		SLV4089
C	PROCESS RIGHT HAND SIDE	SLV4090
C		SLV4091
2000	CONTINUE	SLV4092
	DO 601 I=1,N	SLV4093
601	B(I)=B(I)+RM(I)	SLV4094
	DO 609 I=1,NM1	SLV4095
	IP=IRP(I)	SLV4096
	TEMP=B(I)	SLV4097

	B(I)=B(IP)	SLV4098
	B(IP)=TEMP	SLV4099
609	CONTINUE	SLV4100
C		SLV4101
C	SOLVE FOR FIRST APPROXIMATION TO X	SLV4102
C		SLV4103
199	DO 200 I=1,N	SLV4104
200	Z(I)=-SDOT(I-1,A(I,1),MA,Z(1),1,-B(I))	SLV4105
	DO 201 K=1,N	SLV4106
	I=NP1-K	SLV4107
201	X(I)=-SDOT(N-I,A(I,I+1),MA,X(I+1),1,-Z(I))/A(I,I)	SLV4108
C		SLV4109
C	ITERATIVE IMPROVEMENT	SLV4110
C		SLV4111
	IF(ITMAX)370,370,300	SLV4112
300	TOP=0.0	SLV4113
	DO 303 I=1,N	SLV4114
303	TOP=AMAX1(TOP,ABS(X(I)))	SLV4115
	EPSX=EPS*TOP	SLV4116
	DO 369 IT=1,ITMAX	SLV4117
C	FIND RESIDUALS	SLV4118
	DO 319 I=1,N	SLV4119
319	R(I)=-DOT(N,AA(I,1),MA,X(1),1,-B(I))	SLV4120
C	FIND INCREMENT	SLV4121
	DO 329 I=1,N	SLV4122
329	Z(I)=-SDOT(I-1,A(I,1),MA,Z(1),1,-R(I))	SLV4123
	DO 339 K=1,N	SLV4124
	I=NP1-K	SLV4125
339	DX(I)=-SDOT(N-I,A(I,I+1),MA,DX(I+1),1,-Z(I))/A(I,I)	SLV4126
C	INCREMENT AND TEST CONVERGENCE	SLV4127
	TOP=0.0	SLV4128
	DO 342 I=1,N	SLV4129
	TEMP=X(I)	SLV4130
	X(I)=DAD(X(I),DX(I))	SLV4131
	DELX=ABS(X(I)-TEMP)	SLV4132
	TOP=AMAX1(TOP,DELX)	SLV4133
342	CONTINUE	SLV4134
	IF(TOP=EPSX)381,381,369	SLV4135
369	CONTINUE	SLV4136
370	IT=0	SLV4137
381	RETURN	SLV4138
C		SLV4139
C	RESTURE A AND B	SLV4140
C		SLV4141
3000	CONTINUE	SLV4142
	DO 709 K=1,N	SLV4143
	I=NP1-K	SLV4144
	IP=IRP(I)	SLV4145
	IF(I=IP)701,709,701	SLV4146
701	TEMP=B(I)	SLV4147
	B(I)=B(IP)	SLV4148
	B(IP)=TEMP	SLV4149
	DO 702 J=1,N	SLV4150
	TEMP=AA(I,J)	SLV4151
	AA(I,J)=AA(IP,J)	SLV4152

702		AA(IP,J)=TEMP		SLV4153
709	CONTINUE			SLV4154
	DO 729	I=1,N		SLV4155
		B(I)=B(I)/RM(I)		SLV4156
	DO 729	J=1,N		SLV4157
		A(I,J)=AA(I,J)/RM(I)		SLV4158
729	CONTINUE			SLV4159
	RETURN			SLV4160
	END			SLV4161
\$IBMAP	DOT	84		DOT4000
*	DOT AND FRIENDS		ROUTINES FOR USE WITH SOLVE	DOT4001
	ENTRY	DOT (N,A(1),MA,B(1),MB,C)	DOUBLE INNER PRODUCT	DOT4002
	ENTRY	SDOT (N,A(1),MA,B(1),MB,C)	INNER PRODUCT	DOT4003
	ENTRY	ILOG2 (A)	FLOATING POINT EXPONENT	DOT4004
	ENTRY	DAD (A,B)	ADD WITH ROUND	DOT4005
*	SNAD	MACRO	M STORE NEGATIVE OF ADDRESS IN DECREMENT	DOT4006
		SUB	=0100000 COMPLEMENT IF POSITIVE	DOT4007
		ALS	18	DOT4008
		STD	M	DOT4009
		ENDM	SNAD	DOT4010
*	DOT	SAVE	1,2,4	DOT4011
		STZ	S	DOT4012
		STZ	S+1	DOT4013
		CLA*	8,4 C	DOT4014
		LDQ	C+1	DOT4015
		STO	C	DOT4016
		CLA*	3,4 N	DOT4017
		TZE	NONE SKIP LOOP IF N = 0	DOT4018
		STQ	N	DOT4019
		CLA	4,4 BASE ADDRESS OF A	DOT4020
		PAC	,1 X1=-(BASE OF A)	DOT4021
		CLA*	5,4 MA	DOT4022
		SNAU	MA	DOT4023
		CLA	6,4 BASE ADDRESS OF B	DOT4024
		PAC	,2 X2=-(BASE OF B)	DOT4025
		CLA*	7,4 MB	DOT4026
		SNAU	MB	DOT4027
		LXA	N,4 X4=N	DOT4028
LOOP		LDQ	0,1 A(I)	DOT4029
		FMP	0,2 B(I)	DOT4030
		DFAU	S	DOT4031
		DST	S	DOT4032
MA		TXI	**1,1,** (X1)=(X1)+MA	DOT4033
MB		TXI	**1,2,** (X2)=(X2)+MB	DOT4034
		TIX	LOOP,4,1 END OF MAIN LOOP	DOT4035
NONE		DFAU	C	DOT4036
		FRN		DOT4037
		RETURN	DOT	DOT4038
*	SDOT	SAVE	1,2,4	DOT4039
		STZ	S	DOT4040
		CLA*	8,4	DOT4041
		STO	C	DOT4042
				DOT4043
				DOT4044
				DOT4045

	CLA*	3,4	DOT40460
	TZE	SNONE	DOT40470
	STO	N	DOT40480
	CLA	4,4	DOT40490
	PAC	,1	DOT40500
	CLA*	5,4	DOT40510
	SNAU	SMA	DOT40520
	CLA	6,4	DOT40530
	PAC	,2	DOT40540
	CLA*	7,4	DOT40550
	SNAU	SMB	DOT40560
	LXA	N,4	DOT40570
SLOOP	LDQ	0,1	DOT40580
	FMP	0,2	DOT40590
	FAD	S	DOT40600
	STO	S	DOT40610
SMA	TXI	**+1,1,**	DOT40620
SMB	TXI	**+1,2,**	DOT40630
	TIX	SLOOP,4,1	DOT40640
SNONE	FAD	C	DOT40650
	RETURN	SDOT	DOT40660
*			DOT40670
ILOG2	CAL*	3,4	DOT40680
	ANA	=03770000000000	DOT40690
	SUB	=02000000000000	DOT40700
	ARS	27	DOT40710
	TRA	1,4	DOT40720
*			DOT40730
DAD	CLA*	3,4	DOT40740
	FAD*	4,4	DOT40750
	FRN		DOT40760
	TRA	1,4	DOT40770
*			DOT40780
	EVEN		DOT40790
C	PZE		DOT40800
	PZE		DOT40810
S	PZE		DOT40820
	PZE		DOT40830
N	PZE		DOT40840
	END		DOT40850

This program essentially computes the distance from an unknown spectrum of six bands to the means of three general terrain types-- all this in theoretically six dimensions of transmittance. The group which is closest to the unknown spectrum is then assumed to be the most likely choice and is printed in the proper location. This distance is known as the Mahalanobis or M-distance and is an extension of the Pythagorean theorem. In this program the distance in two dimensions is computed and this distance used to compute the third dimension, and higher dimensions in sequence up to six.

The means of the bands for each group are input using a DATA statement, as are the names of these groups. The spectra cards are used in the format first presented for the remaining input, while the output consists of the six-letter code name of the group selected for each spectrum and its distance in percent transmittance to the unknown spectrum. If the amount of variation in the spectra being classified is great, then this may cause many of them to be incorrectly identified. A large variation of one band relative to another is the greatest weakness of this system.

\$JOB R563 IBJOB 2 800 BALLEW, GARY 233 6=D GEOMAP
 \$IBJOB
 \$IBFTC MAIN

C MULTIDIMENSIONAL CLASSIFICATION OF MULTIBAND SPECTRA
 DIMENSION TITLE(12), SYMBOL(60), XNAME(5), BAND(5,9,6),
 1 AMEAN(3,6), ANAME(3), RMIN(5), D2(3)
 DATA (ANAME(NG), NG=1,3) / 6HRHYOLT, 6HPUMICE, 6HSHADOW/
 DATA ((AMEAN(NG, I), I=1,6), NG=1,3) / 0.65167, 0.71111, 1.33167,
 1 1.82056, 1.19667, 2.64833,
 2 0.54696, 0.43913, 0.47870, 0.68239, 0.48196, 0.74174,
 3 0.40000, 0.27467, 0.31400, 0.45133, 0.35533, 0.52067/
 DATA (SYMBOL(L) ,L=1,60) / 60*1H./
 1000 READ(5,1)(TITLE(K),K=1,12)
 1 FORMAT(12A6)
 WRITE(6,2) (TITLE(K), K=1,12)
 2 FORMAT(1H1,12A6)
 WRITE(6,401)(SYMBOL(L),L=1,60)
 WRITE(6,303)
 WRITE(6,303)
 READ(5,3) (((BAND(NX,NY,I),I= 1,6), NY=1,9) ,NX=1,5)
 3 FORMAT (36X,6F4.2)
 DO 400 NY=1,9
 DO 300 NX=1,5
 RMIN(NX)=0.0
 DO 200 NG=1,3
 D2(NG) = 0.0
 DO 100 I=1,6
 D2(NG) = D2(NG) + (AMEAN(NG,I) -BAND(NX,NY,I)) *
 1 (AMEAN(NG,I) - BAND(NX,NY,I))
 100 CONTINUE
 IF(D2(NG) .LE.RMIN(NX)) GO TO 201
 GO TU 200
 201 RMIN(NX) = SQRT(D2(NG))
 XNAME(NX)= ANAME(NG)
 200 CONTINUE
 300 CONTINUE
 WRITE(6,301)(XNAME(NX),NX=1,5)
 301 FORMAT(1HT, 7H. , 5(A6,4X), 3H .)
 WRITE(6,302) (RMIN(NX),NX=1,5)
 302 FORMAT(3H. , 5F10.4 , 7H .)
 WRITE(6,303)
 303 FORMAT(1HT, 1H., 58X, 1H.)
 400 CONTINUE
 WRITE(6,401)(SYMBOL(L),L=1,60)
 401 FORMAT(1HT, 60A1)
 GO TU 1000
 RETURN
 END

?32

PRECEDES B5000 AND IBJOB DATA CARDS

Sum of Probabilities Method of Classification

In order to account for the variation between deviations within six bands, a statistical method of multivariate analysis appears to give the best results. By using this program, the probability that an unknown spectrum may belong to a group is computed for each band by assuming that this is a function of a normal error curve. The position and shape of this curve is determined by the mean and standard deviation of the ideal groups in each of the six bands, and the probabilities are summed for each of the major groups being considered. For this purpose the error function ERF is called from the library of mathematical functions and suitably modified.

The input for the program consists of data cards of 45 spectra arranged as first described in this writeup. The output is a map of most probable spectra and their probabilities of being correct. A six-letter code word is printed along with the probability of its being the group selected as most likely, as the code words are input with the means and standard deviations by use of DATA statements.

Groups with extremely large standard deviations or an exceptionally large number of ideal spectra should be avoided, as the groups with the largest deviations will be chosen a disproportionate percent of the time.


```

$JOB R563 IBJOB 2 800 BALLEW, GARY 233 SUM PROBS 4 POSS
$IBJOB
$IBFTC MAIN
C SUM OF PROBABILITIES MAP
  DIMENSION TITLE(12), SYMBOL(60), XNAME(5), BAND(5,9,6),
  1 SUM(3), AMEAN(4,6), SIGMA(4,6), ANAME(4), RMAX(5)
  DATA (ANAME(NG), NG=1,4) / 6HRHYOLT, 6HPUMICE, 6HSHADOW,
  1 6HPUMISG /
  DATA ((AMEAN(NG, I), I=1,6), NG=1,4) / 0.65167, 0.71111, 1.33167,
  1 1.82056, 1.19667, 2.64833,
  2 0.54696, 0.43913, 0.47870, 0.68239, 0.48196, 0.74174,
  3 0.40000, 0.27467, 0.31400, 0.45133, 0.35533, 0.52067,
  4 0.65270, 0.58919, 0.60635, 0.92770, 0.62703, 1.03851/
  DATA ((SIGMA(NG, I), I=1,6), NG=1,4) / 0.16745, 0.20041, 0.82748,
  1 0.68870, 0.51649, 1.47826,
  2 0.12811, 0.15430, 0.22833, 0.30602, 0.18508, 0.30488,
  3 0.11582, 0.12141, 0.23561, 0.27126, 0.16591, 0.43310,
  4 0.20135, 0.22041, 0.28852, 0.36734, 0.26782, 0.50920/
  DATA (SYMBOL(L), L=1,60) / 60*1H./
1000 READ(5,1)(TITLE(K), K=1,12)
  1 FORMAT(12A6)
  WRITE(6,2) (TITLE(K), K=1,12)
  2 FORMAT(1H1,12A6)
  WRITE(6,401)(SYMBOL(L), L=1,60)
  WRITE(6,303)
  WRITE(6,303)
  READ(5,3) ((BAND(NX,NY,I), I=1,6), NY=1,9), NX=1,5)
  3 FORMAT (36X,6F4.2)
  DO 400 NY=1,9
  DO 300 NX=1,5
  RMAX(NX)=0.0
  DO 200 NG=1,4
  SUM(NG) = 0.0
  DO 100 I=1,6
  ABSDIF = ABS(BAND(NX,NY,I) - AMEAN(NG,I))
  SUM(NG) = SUM(NG) + (1.0 - ERF(ABSDIF / (SIGMA(NG,I) * 1.414214)))
  1 /6.0
100 CONTINUE
  IF(SUM(NG).GE.RMAX(NX)) GO TO 201
  GO TO 200
201 RMAX(NX) = SUM(NG)
  XNAME(NX) = ANAME(NG)
200 CONTINUE
300 CONTINUE
  WRITE(6,301)(XNAME(NX), NX=1,5)
301 FORMAT(1HT, 7H, 5(A6,4X), 3H .)
  WRITE(6,302) (RMAX(NX), NX=1,5)
302 FORMAT( 3H, 5F10.4, 7H .)
  WRITE(6,303)
303 FORMAT(1HT, 1H., 58X, 1H.)
400 CONTINUE
  WRITE(6,401)(SYMBOL(L), L=1,60)
401 FORMAT(1HT, 60A1)
  GO TO 1000
  RETURN

```

232 END

PRECEDES B5000 AND IBJOB DATA CARDS