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Code CUPED

A Code to Unfold
Scintillation Spectrometer
Polyenergetic Gamma Photon
Experimental Distributions

CASE FILE
COPY

September 1969

NASA Contract:
NAS 5-10337

NUS CORPORATION

Prepared by

J. J. Steyn
U. Born
J. T. Strahl
R. Huang

For

Goddard Space Flight Center
Greenbelt, Maryland

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SUMMARY

A FORTRAN code has been developed for the IBM-360 digital computer to unfold sodium-iodide (thallium-activated) scintillation spectrometer polyenergetic gamma photon experimental distributions. It was specifically designed to analyze the combination bremsstrahlung and monoenergetic gamma radiation field of cylindrical radioisotope power generators. The code generates the detector system response matrix function and applies it to the monoenergetic spectral components discretely and to the bremsstrahlung iteratively. It corrects for iodine K X-ray escape, detector non-linearity, system drift, source decay, background, and detection efficiency. Results are presented in digital form for differential and integrated photon number and energy distributions, and exposure dose.

1. INTRODUCTION

This report presents a description of and the user requirements for code CUPED --- a digital computer code to unfold scintillation spectrometer polyenergetic gamma photon experimental distributions developed under NASA-GSFC contract NAS5-10337. Code CUPED, written in the FORTRAN IV language for the GSFC IBM-360/91 digital computer, is a much modified version of code CUBED which was developed under contracts NAS5-10133 and NAS5-10337 and previously reported in NUS-315, -316, and -395^(1,2,3). The experimental distributions which the code has been specifically developed to analyze, are those recorded by the right-cylindrical sodium-iodide (thallium-activated) ---- NaI(Tl) ---- scintillation crystal, coupled to a multi-channel pulse-height analyzer and exposed to the bremsstrahlung and line energy photons emitted by right-cylindrical radioisotope power generators fuelled with such as either plutonium, thallium or strontium oxides.

The code can either read or generate the scintillation detector system response function matrix and apply it to unfold the pulse-height analyzer distributions to determine differential and integrated photon number and energy distributions, and exposure dose. The response matrix generation procedure relies on the spectra of standard radioisotopes such as Cd¹⁰⁹, Hg²⁰³, Sr⁸⁵, Cs¹³⁷, Nb⁹⁵, Mn⁵⁴, Zn⁶⁵, Co⁶⁰ and Na²⁴. The standard spectra are normalized with respect to photopeak pulse-height and area, and their photopeaks subtracted to obtain normalized Compton continua. The response matrix vectors are determined at each energy by interpolating the normalized continua and computing the associated Gaussian photopeaks. The thus interpolated vectors are redistributed in pulse-height to correspond to the detector system energy response and to satisfy the requirements of the spectra to be unfolded. Quadratic interpolation of the normalized continua is carried out either

directly for gamma photon energies ≤ 0.6616 MeV or by a method of parts , described in this report , for energies > 0.6616 MeV.

Code CUPED applies the response function matrix according to two distinct procedures to unfold either bremsstrahlung spectra or complex spectra consisting of a limited number of photopeaks or spectra consisting of bremsstrahlung-plus-photopeaks. The code determines the detector incident photon lines in the unknown spectra through the fitting of a Gaussian-plus-straight-line function to each photopeak. The corresponding photopeak associated Compton continua are determined by an interpolation of the standard Compton continua. The thus determined photopeak-plus-continuum spectra are then subtracted to leave, ideally, either a continuous or zero residual spectrum. Continuous and residual spectra are iteratively unfolded according to the matrix inversion technique of Scofield⁽⁴⁻⁷⁾, to determine detector-incident continuous photon spectra. The total incident gamma photon spectrum is determined as the sum of the line and continuous components.

The code corrects for partial photon energy deposition in the NaI(Tl) detector through the application of the response matrix. It corrects for the number of photon interactions in the detector crystal and for absorptions by the crystal cladding materials, interposed absorbers such as Lucite and the air medium between the source and the crystal. In addition, it corrects for primary source decay, iodine K X-ray escape, pulse-height drift and natural background. It corrects for detection system nonlinear energy response either inherently during unfolding or directly before unfolding according to option.

Since the present report is considered primarily as a code user's manual, a detailed description of the mathematics and the logic of the code is referred to report NUS-316⁽²⁾ and NUS-395⁽³⁾.

2. CODE DESCRIPTION

2.1 INTRODUCTION

Code CUPED is written in the FORTRAN IV compiler language for the GSFC IBM-360/91 digital computer. It was designed to run under the System 360 monitor system at NASA Goddard Space Flight Center. Input data are read from card-to-tape ---- TAPE 5, digital output is written on tape for print out ---- TAPE 6. The code calls only standard library subroutines, such as transcendental functions.

The code consists of a main control program and thirty-seven subprograms, numbered from 1 to 38. A subprogram glossary is given in Appendix I, in alphabetic name order, and a code FORTRAN punch card deck listing in Appendix II. Appendix III consists of a sample input card deck listing and Appendix IV of a program output listing corresponding to the input given in Appendix III: the execution time for the sample data was approximately five seconds (360/91) with compile and loading time being approximately 28 seconds.

Although CUPED is written for the -360/91 it may be run on smaller 360 machines, down to a -360/50. It has total byte length of less $\sim 2.05 \times 10^5$ (the -360 is 32 bit word third generation computer with 8 bits per byte) and thus a 'word' size of about 62k in the vernacular of second generation computers.

The logic and function of the main program, referred to as MAIN, and its subprograms are discussed in Section 2.2, in some detail. Those subprograms not discussed are considered as being adequately described in either Appendix I or references (1), (2) and (3). The user is also referred to those

same references for the theoretical bases of code CUPED. The constants required by the code are explained in Section 2.3. Reference to this Section will allow the user to make changes as necessary, in for example, the relationship of such as the detector system photopeak resolution and pulse-height with photon energy. The code operation is discussed in Section 3.1 with special reference to the various options available. The code input is detailed in Section 3.2, and the output is defined in Section 3.3. FORTRAN names and variables are shown capitalized in what follows, with 'zero' and 'oh' thus: 0, Ø.

2.2 CODE LOGIC

2.2.1. MAIN PROGRAM

The main program was designed to execute data input and output operations, many of them under initially input option signals, and provide the control connectivity for the hierarchy of thirty-eight subprograms presented in Figure 1. Figure 2 shows a simplified flow diagram of the main program. Program MAIN calls subprograms:

SHAPE	ØMITS	DEC	GANE	SINGLE
SØLN	DECAY	GEØMTR	ENLIN	XTAL

Subprogram SHAPE is called by MAIN to generate and return the detector system response matrix and the associated vectors relating pulse-height to photon energy. It also returns a vector of photopeak-area-to-total-spectrum-area ratios, i.e. experimental photofractions, for code check purposes. Under control of an input option signal, the matrix and its associated vectors may

be read as a card deck, instead of generated. For the analysis of many sets of unknown spectra one response matrix may be applicable, and thus the main program includes an option to bypass both the calling of SHAPE and the input of a matrix card deck. For similar reasons an option is provided to call SHAPE to generate a response matrix based on previously input standard spectra.

Input options allow either the execution of MAIN to continue, return to start or call EXIT, after the calling of SHAPE. Thus, for example, the code may be run only for the purpose of generating a response matrix.

The code takes advantage of the fact that a series of input data cards for a set of unknown source spectra, may vary in only one or two variables, and so the remainder need not be repeated, the unchanging variables being supplied through the automatic calling of subprogram ØMITS by MAIN.

Unknown spectra may be input to the code in uninterrupted blocks of up to twenty spectra through the calling of subprogram DEC by MAIN. Since pulse-height analyzer background subtracted counts are normally recorded in a complement mode, i. e. as positive numbers, the code converts them to true negative numbers in subprogram DEC. A count greater than 9×10^5 is assumed to be in the complement mode.

MAIN is coded to subtract background spectra from source-plus-background spectra under an input option control signal. This option allows the subtraction or addition, of a fraction or multiple of a background spectrum. It further allows the continued reuse, as desired, of a previously stored background spectrum, and of course the addition of similar spectra if this is the requirement of the user.

The energy correspondence of an unknown spectrum to the response function matrix is matched through MAIN calling subprogram GANE^(1,2,8). Subprogram GANE returns the unknown spectrum to the main program after normalization to pulse-height analyzer true zero pulse-height and gain changing such that channel width corresponds to that of the response matrix.

Additional automated modification of an unknown spectrum may be optionally carried out by the code, namely:

- (a) the counts in the first n 'dead' channels of the spectrum may be replaced with either the count stored in channel n + 1 or counts determined by a straight line function of specified slope; and/or
- (b) the counts between specified limiting channels may be replaced by counts determined by a straight line function of specified slope.

Item (a) allows the user to make judgments with respect to the first few percent of the spectrum which is often either suspect or actually electronically distorted. Item (b) allows the removal of a known spurious peak, etc. in order to allow more meaningful subsequent analysis.

At this point in the main program on both the first and subsequent loops, the response matrix is stored, and an unknown pulse-height analyzer spectrum is ready for analysis and conversion to a photon number spectrum. According to an input option signal the code is instructed that the unknown spectrum is either a pure continuum or a specified number of monoenergetic spectra "super-imposed" on a continuous spectrum. If the input option indicates the presence of superimposed monoenergetic spectra and if either their line energies and/or photopeak tail channel limits are input, MAIN calls subprogram SINGLE prior to calling subprogram SOLN to carry out iterative unfolding of the continuum.

Subprogram SINGLE is called by MAIN to analyze the photopeaks indicated by the input options for the monoenergetic components of the unknown spectrum. Gaussian distributions are fitted to each photopeak and the associated Compton continua are determined by SINGLE calling subprogram SHAPE and the corresponding line photon numbers calculated. In this manner the monoenergetic spectral components are established in turn and subtracted, i.e. stripped, from the unknown spectrum to leave a residual continuum spectrum. The residual spectrum is returned to MAIN for subsequent analysis and unfolding. The monoenergetic photon numbers thus obtained are added to the residual photon number spectrum determined later by unfolding, to give the total photon number spectrum corresponding to the input pulse-height analyzer spectrum.

Before calling subprogram SØLN the code tests certain input options (M(10), M(11), M(12), to determine the users special requirements, three of these are, namely:

- (a) is unknown spectrum to be corrected for non-linear energy response by calling subprogram ENLIN prior to or during unfolding?;
- (b) is unknown spectrum to be corrected for detection efficiency by calling subprogram XTAL prior to gain changing, (eg. from 200 to 30 channels) and before unfolding or after?; or
- (c) is unknown spectrum to be actually unfolded?

Item (a) allows the user to make decisions with respect to the effects of the linearity correction. Item (b) allows the user to make decisions with respect to efficiency correction effects and accuracies. Item (c) allows the user to make decisions with respect to such as the significance of actually carrying out the unfolding procedure on a low energy spectrum such as that of Pm^{147} . The optional linearizing of unknown spectra causes the code (subprogram SHAPE) to automatically linearize standard spectra and generate a linear response matrix. The above options and all others, are defined in more detail in Section 3.2. The called subprograms SHAPE, ENLIN, SINGLE, XTAL, SØLN

and GEØMTR are discussed further in Subsections 2.2.2 through 2.2.7. The code corrects unknown spectra for the non-linear energy response of NaI(Tl) either during the unfolding process or if according to an input option this process is by-passed, by the calling of subprogram ENLIN prior to unfolding. Subprogram ENLIN returns an energy linearized spectrum to MAIN which may be subsequently corrected for other phenomena.

At this point in the main program on both the first and subsequent loops, unknown spectra are considered as prepared for unfolding, and thus, they and the response matrix are communicated to subprogram SØLN. This subprogram is called by MAIN to control the iterative unfolding process according to the Scofield method^(2,4,5) and to apply efficiency corrections. It returns the corrected photon number spectrum to the main program.

The main program calls subprogram DECAY to determine the primary source decay factor. Subprogram DECAY returns a correction factor by which the number spectrum is later multiplied. The calling of subprogram DECAY may be optionally bypassed, in which case the multiplying factor is assumed as unity.

Subprogram GEØMTR is called by the main program to apply the decay correction factor, carry out geometrical corrections, compute the final differential and integral photon number and energy distributions, and exposure dose.

Subprogram GEØMTR returns the final results to the main program for output, after which the code loops back along either paths 1, 2, or 3, as shown in Figure 2. The code loops back along path 1 primarily to read new data pertaining either to the response matrix or to the control options or both. The code loops back along path 2 to read new data pertaining either to the unknown source or if the maximum number of allowable passes (twenty) along path 3 have been equaled; the second reason is dictated by either code DIMENSION or computer finite capacity. The code loops back along path 3 to read a new unknown spectrum.

2.2.2 RESPONSE FUNCTION MATRIX

2.2.2.1 Response Matrix Generation

The detector system response function matrix is generated under the control of subprogram SHAPE. The subprograms called by SHAPE are those shown in Figure 3, namely:

GANP	RESGEN	TA	TE	PØLATE	VETCMX
PULSE	RAXEL	PEAKS	PEEK	CØMPLX	ENLIN

The main program supplies SHAPE with a number of control parameters, and three variables. SHAPE begins execution by input of a card deck of spectra of standard radioisotopes. The number of such spectra is equal to NSTAND, where $NSTAND \leq 9$. This card deck is preceded by one parameter card containing information regarding the number of spectra in the deck (NSTAND); the number of "dead" or unused channels at the low energy end of each spectrum (NPHA); the number of channels in the input standards (NXLIM); and the pulse-height analyzer reference coarse gain at which they were measured (UNGAIN). The parameter card is followed by a set of NSTAND cards, each pertaining to and in the same order as the spectra to which they refer. These cards contain the source identity, data regarding peak approximate locations and the deviation of the spectrum from true zero pulse-height, i.e. the $\pm$ normalizing spectrum shift required. A typical card deck is shown in Figure 4. (Further details are referred to Section 3)

The standard spectra allowed by the code must have been measured from the following radioisotope sources:

Cd ¹⁰⁹	Sc ⁴⁷	Hg ²⁰³	Cr ⁵¹	Sr ⁸⁵	Cs ¹³⁷
Mn ⁵⁴	Nb ⁹⁵	Na ²²	Zn ⁶⁵	Co ⁶⁰	Na ²⁴

The order in which they are input to the code is immaterial excepting that Na²², Zn⁶⁵, Co⁶⁰ and Na²⁴ must be input in the above order. The user may employ

sources not shown above by an obvious modification of the DATA statements in subprogram SHAPE (located at subprogram statement numbers 9000 + 7 and +8), so long as such sources are monoenergetic and do not contain an 0.51 MeV energy photon line. For example, suitable alternate spectra might be either those originating from Au¹⁹⁸ and F¹⁸ sources, or as hand prepared. Multipeak potential standard spectra may be prepared by treating them first as an unknown and using CUPED to determine components.

According to an input option and after the first call, calling of SHAPE allows the by-passing of input of standard spectra. This allows the code to generate a response matrix based on already stored and normalized standard spectra. Similarly subprogram SINGLE calls SHAPE to determine Compton continua based on already stored current standard spectra.

After spectral data is input to SHAPE, counts in the complement mode are converted to their true negative value. The spectra are shifted to true-zero pulse-height by the calling of subprogram GANE. Their order of input is established prior to the calling of subprogram RESGEN for spectral normalization. The standard spectra may also be corrected for non-linear energy response in accord with an input-option (M(14)) if the unknown is to be similarly treated; this is as described in Sections 2.2.1 and 3.2.

Subprogram RESGEN is called by subprogram SHAPE to normalize the single-photopeak and Na²² and Zn⁶⁵ standard spectra with respect to photopeak area and pulse-height; photopeaks and source-characteristic X-ray peaks are subtracted. The multipeak spectra of Co⁶⁰ and Na²⁴ are normalized later by the calling of subprogram CØMPLX. The residual spectra thus determined by RESGEN consist of Compton continua characteristic of the primary photon energy. The X-ray peaks are subtracted since they are not representative of the primary photon energy but rather of the source. The 0.51 MeV component photopeaks and their Compton continua are subtracted for the same reason. Figure 5 shows a typical set of spectral continua as normalized by subprogram RESGEN for SHAPE; the 1.17 MeV continuum determined by CØMPLX is included in this figure. The logic of subprogram RESGEN is described in section 2.2.2.2.

Subprogram RESGEN returns the normalized differential standard Compton distributions to subprogram SHAPE. These distributions are re-ordered with respect to the ascending order of primary photon energy, prior to the normalization of the multipeak standard spectra by CØMPLX. Subprogram SHAPE calls subprogram CØMPLX to determine normalized standard continua at either 1.33 and/or 2.76 MeV using the Co⁶⁰ and Na²⁴ standards, as described in section 2.2.2; the already described normalized Compton continua are required by CØMPLX for this analysis. The energy ordered normalized continua are interpolated quadratically with respect to the energy axis of the desired response function matrix through subprogram SHAPE calling subprogram PØLATE. The result of this interpolation consists of N (corresponding to matrix size) Compton continuum vectors normalized to unit photopeak area and pulse-height. Subprogram SINGLE calls SHAPE to obtain interpolated normalized Compton continua at specific energies as described in this paragraph.

The Gaussian photopeaks plus iodine K X-ray escape peaks of unit (total) area and at unit pulse-height (photopeak) are added to the differential Compton continua to give N (matrix size) complete spectra. The peaks are computed through SHAPE calling subprogram PEAKS which in turn calls GAUSS, which calls PEEK. The Gaussian photopeak photon energy dependent standard deviation $\sigma(E)$, is computed here from an expression of the type

$$\sigma(E) = k \cdot E^n \quad (1)$$

where

k and n are (presently) user-determined either by a regression analysis or by a plot of $\sigma(E)$ on log-log graph paper to obtain the slope n and the intercept k (at $E = 1.0$ keV).

The energy dependent K X-ray escape fraction is computed by calling function RAXEL, which interpolates a stored table of escape fractions described in Section 2.3.

The N determined unit length differential spectra are redistributed in pulse-height linearly or non-linearly in accord with the option chosen. The redistribution is obtained through SHAPE calling subprogram GANE. The non-linear pulse-heights are obtained by the calling of function subprogram PULSE. Linear pulse-heights are determined if the standards were linearized by subprogram ENLIN.

At this point in the subprogram SHAPE execution, a response function matrix has been determined. This matrix and its corresponding vectors for pulse-height, photon energy and photofraction are returned to the main program. Figure 6 compares the photofractions of the response matrix vectors determined by SHAPE with actual experimental values separately determined for the standard source spectra.

2.2.2.2 Spectrum Normalization

Subprogram RESGEN normalizes the standard spectra input to SHAPE. This subprogram begins execution by carrying out necessary initializations. In the event that an 0.51 MeV spectrum and either Zn^{65} and/or Na^{22} have been included in the input standard spectra, they are also stored in dummy vectors for later 0.51 MeV component subtraction. SHAPE calls control subprogram CØBALT, which in turn calls subprogram STDFIT, to carry out a Gaussian fit to the primary and 0.51 MeV (of Na^{22} and Zn^{65}) spectral photopeaks. Subprogram STDFIT calls subprogram GUESS to estimate the necessary initial values of the five function parameters: straight-line slope and intercept, Gaussian photopeak standard deviation, area and mean pulse-height. Subprogram STDFIT returns the parameters of the fitted photopeaks to subprogram RESGEN.

Subprogram RESGEN uses the determined photopeak parameters to carry out normalization and to subtract the photopeaks from the standard spectra. In

addition to photopeak subtraction, a subtraction of the characteristic X-ray peaks is carried out in the case of those spectra where they occur. The 0.51 MeV photopeak and continuum of Sr^{85} , if it has been input, is employed to subtract the 0.51 MeV spectrum contribution of Na^{22} and/or Zn^{65} . This operation requires both count and pulse-height gain normalization, the gain normalization being carried out through the calling of subprogram GANE.

The residual Compton continua are gain normalized to a photopeak pulse-height of 100 channels by the calling of subprogram GANE and count normalized to unit photopeak area by a division operation. The iodine K X-ray escape peak is subtracted for the case of primary photon energies less than 300 keV by the calling of RAXEL. The resulting residual normalized continua are checked for negative count values, which are replaced by zero, and returned to the calling subprogram SHAPE. The fitted photopeak parameters are also returned to subprogram SHAPE.

2.2.2.3 Multipeak Standard Spectrum Analysis

Subprogram CØMPLX, called by SHAPE, separates the components of the multipeak spectra of Co^{60} and Na^{24} . It calls subprogram CØBALT to determine the photopeak parameters and to subtract the photopeaks. It calls subprogram PØLATE to either interpolate or extrapolate a normalized Compton continuum from the already prepared monoenergetic standards for the 1.17 MeV energy component in the case of Co^{60} and the 1.368 MeV component in the case of Na^{24} . If the standards included Na^{22} then an extrapolation is only required for the Na^{24} analysis. The interpolated or extrapolated continuum is modified using the photopeak fitted parameters to determine a continuum which is subtracted from the Co^{60} or Na^{24} continua, as the case may be, to leave a residual 1.33 MeV or 2.754 MeV continuum. The residual continua so found are then normalized and included as additional members of the previously determined set of standards. Figure 6 shows an example 1.33 MeV photofaction

obtained by the code. Na^{24} standard spectra measured in the geometry of the other standards were not available for similar comparison.

2.2.2.4 Compton Continuum Interpolation

Subprogram PØLATE quadratically interpolates or extrapolates normalized Compton continua. It is called by SHAPE for response matrix vector generation purposes and by CØMPLX for multipeak standard spectrum analysis. Actual interpolation is performed in function subprogram TE.

For energies less than 0.6616 MeV, the interpolation is carried out directly on the differential standard continua. The interpolation assumes reasonably that at zero energy, the channel counts are also zero. For energies greater than 0.6616 MeV, a 'method of parts' was developed. This method divides the continuum into three characteristic regions: A) zero energy to backscatter peak, B) backscatter peak to Compton edge, and C) Compton edge to photo-peak mean pulse-height. The mutual boundary of each region is overlapped for continuity reasons. The dashed curves in Figure 5 indicates the regional boundaries.

Prior to interpolation (or extrapolation) the three regions are aligned such that the regional boundaries on which the backscatter peaks and Compton edges lie would be vertical and straight if shown in Figure 5, e.g. backscatter peaks normalized with respect to pulse-height. Alignment is carried out with respect to the continuum of highest energy in the set of three standards in the quadratic, and by the calling of subprogram GANE. The three interpolated components, A, B, and C, are unaligned by gain changing to obtain the desired continuum. The pulse-height axis direction of Region C is reversed for convenience during the whole operation, i.e. zero pulse-height is taken at the photo-peak mean pulse-height. The central portion of Region B is determined by a direct interpolation of the standard continua. An empirically modified form of

the Compton angular-energy equation is used to aid in automatically locating the Compton edges and backscatter peaks.

POLATE returns the interpolated Compton continuum to the calling program.

2.2.3 Energy Response Correction

Subprogram ENLIN is called by MAIN and SHAPE according to an input to correct pulse-height analyzer spectra for the non-linear response of the NaI(Tl) scintillation spectrometer system to gamma photon energy. The logic of subprogram ENLIN is described in this section.

The main program supplies ENLIN with a number of control parameters and four variables. The variables are, namely:

- a. The number of channels in the spectrum to be linearized (NX),
- b. The counts in the spectrum to be linearized (FM(I)),
- c. The energy at which the spectrum was calibrated (EG), and
- d. The measured pulse-height corresponding to the calibration energy (VG).

For the sake of discussion in this section the above input to ENLIN will be referred to as n , C_1' , E_m and V_m' , respectively.

Subprogram ENLIN begins execution by defining the linear response slope, or channel energy worth Δ as

$$\Delta = E_m / V \tag{2}$$

where

$$V = V'_m / (1 + \partial V_m (E_m)) \quad (3)$$

and

$\partial V_m (E_m)$ is the fractional deviation of pulse-height from a linear response for energy E_m , as determined by ENLIN calling subprogram PULSE; ∂V is further discussed in Section 2.3. The linear response is normalized at energies $E = 0$ and 1.3325 MeV.

The non-linear channel energy worth of channel i , is $V'_i - V'_{i-1}$

where

$$V' = V (1 + \partial V (E)) \quad (4)$$

and

$$V = E/\Delta \quad (5)$$

With the above relationships established the linear response count in channel i , C_i , may be determined as

$$C_i = C'_i \cdot (V'_i - V'_{i-1}) / (V_i - V_{i-1}) \quad (6)$$

where the denominator is equal to one channel.

2.2.4 Analysis of Monoenergetic Spectral Contributions

Subprogram SINGLE is called by MAIN according to an input option to analyze photopeaks and their associated Compton continua, in complex spectral distributions. Subprogram SINGLE will fit a 'single or double Gaussian plus

straight line' function to the photopeaks of a multipeak PHA spectrum. It will subtract the fitted photopeaks and their associated continua to leave a residual continuous spectrum. In the case of a 'no-bremsstrahlung' complex PHA spectrum the residual will, ideally, have zero intensity.

The code CUPED user may input either the energy of the photopeaks to be fitted in SINGLE or alternately the photopeak fitting limits (channel numbers). In the event that the energies of only certain peaks are known they may be input, while the remainder may be defined by fitting limits. The choice of energy order of monoenergetic peak analysis and spectral stripping is left to the user; the code permits a mixed order to be chosen.

In addition to fitting and subtracting the monoenergetic components of an unknown spectrum, SINGLE determines their corresponding efficiency-corrected detector-incident photon number. The thus determined photon number may be optionally, either added to the bremsstrahlung photon number determined by later unfolding or diverted for separate output. In this way the separated radiations can be studied, eg. bremsstrahlung analyses may be carried out even though the subject source emits monoenergetic photons.

Although SINGLE is designed primarily for single or double (merged) photopeaks, it may be applied to x-ray peaks with a reasonable degree of accuracy. For x-ray peak analysis a single Gaussian will give an approximate result. The fitting of a double Gaussian requires that the peak have good counting statistics and thus be 'well' formed. Alternately, an x-ray peak may be analyzed approximately by the iterative unfolding procedure under the control of subprogram SØLN. It is noted that the iterative unfolding convergence rate is generally slower in spectra containing prominent residual peaks.

Subprogram SINGLE begins execution by determining whether the photopeak fitting limits have been input to MAIN by the user or whether they are to be determined. In the event that they are to be determined, SINGLE begins execution by estimating their channel locations based on the peak input energies. It first establishes the approximate channel region of the photopeak. It then ascertains it more accurately for the approximated channel region by calling subprogram VECTMX to establish the channel of maximum count. The fitting limits are determined as a function of the photon energy dependent standard deviation. A check is made to ensure that the limits are within the spectrum and that their domains do not overlap each other.

With the fitting limits established, a single or double Gaussian distribution is fitted to each photopeak in turn. Actual photopeak function fitting is carried out by the calling of subprogram STDFIT for single peaks and STDFT2 for double peaks. Although subprogram STDFT2 is capable of carrying out the fitting of a single peak and so leaving STDFIT redundant, it is not required to do so because of the other code CUPED changes that would be necessary. It is proposed that those modifications be considered as future work. Subprograms STDFIT and STDFT2 return fitted photopeak parameters to SINGLE.

The Compton continuum associated with each photopeak is determined by calling subprogram SHAPE, which returns an interpolated continuum normalized with respect to a photopeak of unit area and pulse-height (100 channels). The continuum is then scaled and gain changed according to the peak area and pulse-height determined by STDFIT or (STDFT2).

Gain changing is carried out through SINGLE calling subprogram GANE. The photopeak and Compton continuum are then subtracted from the unknown spectrum for each monoenergetic spectral component in turn to finally leave

a bremsstrahlung continuum residual spectrum. If no bremsstrahlung or other continuous contribution was present in the unknown, then ideally a zero spectrum will result.

Prior to returning the residual continua to MAIN for iterative unfolding, SINGLE determines the photon number corresponding to each monoenergetic spectral component. This is done by computing the photofraction, $P(E)$, the detector interaction efficiency, $\epsilon(E)$, and the attenuation term for detector cladding, air and lucite material interposed between the source and NaI(Tl) crystal, $\eta(E)$, all as outlined in reference (2). The photon number is then determined from the relationship:

$$N(E) = \frac{\text{Photopeak Area (or Counts)}}{P(E) \cdot \epsilon(E) \cdot \eta(E)} \quad , \quad (7)$$

The corrections noted are carried out by subprogram SINGLE calling subprogram XTAL. $N(E)$ is returned to MAIN either to be added to the iteratively unfolded continuum number spectrum or to be output separately.

Figure 7 shows a PHA spectrum of PuO_2 before and after analysis by subprogram SINGLE.

2.2.5 Spectral Unfolding

The reduction of pulse-height analyzer continuous spectra to photon number spectra and the application of efficiency corrections are carried out under the control of subprogram SØLN called by the main program. Subprogram SØLN begins execution by carrying out certain initializations after which it calls subprogram RESMAT to unfold the pulse-height analyzer spectra according to the Scofield method ^(4,5). The number spectra returned by subprogram RESMAT are corrected for efficiency by SØLN calling XTAL. The thus corrected number spectra are returned to MAIN. The remainder of this

section describes the logic of the unfolding subprogram RESMAT and of the efficiency vector subprogram XTAL, EFFIC, AIRABS, PERSPX and CLAD.

Subprogram RESMAT unfolds the pulse-height analyzer spectra by solving the matrix equation (in matrix notation)

$$\vec{P} = \bar{R} \vec{N}' \quad (8)$$

where $\vec{P}$ and $\vec{N}'$ are the m-dimensional vectors of the PHA spectrum and the efficiency uncorrected photon number spectrum, respectively, and $\bar{R}$ is the m x m square response function matrix. Equation (8) is formally solved as

$$\vec{N}' = \bar{R}^{-1} \vec{P} \quad (9)$$

where $\bar{R}$ is non-singular and $\bar{R}^{-1}$ is its inverse. Subprogram RESMAT executes equation (9) iteratively according to the Scofield method^(4,5). Figure 8 shows a flow diagram of the iterative algorithm coded in subprogram RESMAT. Further details are referred to references (1-3).

The efficiency corrected photon number spectrum $\vec{N}$, is determined from equation (9), as

$$\vec{N} = \bar{\eta}^{-1} \vec{N}' \quad (10)$$

where $\bar{\eta}$ is a diagonal efficiency matrix accounting for interaction efficiency and photon attenuation by detector cladding, air and lucite materials interposed between the source and the crystal. Subprogram SØLN calls subprogram XTAL which in turn calls function subprograms EFFIC, CLAD, AIRABS and PERSPX, to determine $\bar{\eta}$. SØLN then executes equation (10) and returns the determined photon number spectrum to MAIN. Figure 9 shows a typical spectrum before ($\vec{P}$) and after ($\vec{N}$) unfolding.

2.2.6 Analysis of Unfolded Spectra

Subprogram GEØMTR is called by MAIN to carry out a final analysis on the unfolded photon number spectra. The spectra are corrected for primary source decay and converted by GEØMTR to differential photon number flux at the detector per unit time, $N_x(E)$, (coded as FNXTAL), as

$$N_x(E) = \frac{N(E)}{\pi R_x^2}, \quad \gamma/\text{cm}^2 \text{ sec} \quad (11)$$

where

R_x = NaI(Tl) crystal radius, cm.

The differential energy flux incident on the crystal per unit time, $I_x(E)$, (codes as ENXTAL), is determined as

$$I_x(E) = N_x(E) \cdot E, \quad \text{MeV}/\text{cm}^2 \text{ sec} \quad (12)$$

The energy integrated exposure dose rate at the crystal, D , (coded as DOSDET), is determined as

$$D = \int_{\text{energy}} N_x(E) E \mu_{\text{air}}(E) K dE, \quad \text{roentgens/hours} \quad (13)$$

where

$\mu_{\text{air}}(E)$ = energy mass absorption coefficient of air, cm^2/gm

K = conversion constant

= $3600/5.24 \times 10^7$, (roentgens-second-gm air)/
MeV-hour

The integration in equation (13) is carried out numerically by GEØMTR, as

$$D = \sum_{i=1}^m N(E_i) E_i \mu_{\text{air}}(E_i) K \Delta E_i \quad , \quad (14)$$

The energy integrated photon number and photon energy flux at the crystal is determined by integrating $N_x(E)$ and $I_x(E)$ over E , (coded as SUMNUM and SUMENY); the units are $\gamma/\text{cm}^2\text{-sec}$ and $\text{MeV}/\text{cm}^2\text{-sec}$. The following tabulated data are also determined by subprogram GEØMTR for output by the calling main program:

<u>FORTTRAN NAME</u>	<u>EQUAL TO</u>	<u>DEFINITION & UNITS</u>
(AT THE CRYSTAL)		
AVENGY	$\frac{\sum N(E) \cdot E \Delta E}{\sum N(E) \cdot \Delta E}$	average energy, MeV
PHNUBE	$\frac{\sum N(E) \cdot \Delta E}{N_{\beta}}$	integrated photon number flux per beta source strength, $(\gamma/\text{cm}^2\text{-sec})/(\beta/\text{sec})/\text{MeV}$; (N_{β} defined below) *
ENBENY	$\frac{\sum N(E) \cdot E \cdot \Delta E}{E_{\beta \text{max}}}$	integrated energy flux per beta maximum energy, $(\text{MeV}/\text{cm}^2\text{-sec})/\text{MeV}$; ($E_{\beta \text{max}}$ defined below)
PHENBE	$\frac{\sum N(E) \cdot \Delta E}{N_{\beta}}$	integrated energy flux per beta source strength, $(\text{MeV}/\text{cm}^2\text{-sec})/(\beta/\text{sec})$
DOXBEX	D/N_{β}	dose rate per beta source strength, $(\text{r/hr})/(\beta/\text{sec})$; (D defined in equation (13)).
(AT THE BETA SOURCE CYLINDER)		
DOSCYL	$\frac{\text{DOXBEX}}{G}$	dose rate per beta source strength, $(\text{r/hr})/(\beta/\text{sec})$; (G defined below)

*Beta, $N_{\beta A}$ and $E_{\beta \text{max}}$ are used in this section because of earlier bremsstrahlung analyses; gamma (γ) may be conceptually substituted.

DCYVOL

DOSCYL
Source Volume

dose rate per beta source
strength per cm^3 of source
volume, = $\text{DOSCYL}/\text{cm}^3$

where

$E_{\beta_{\text{max}}}$ = maximum beta (or chosen γ) energy in MeV

= EBMAX of card ⑥ of report section 3.2*

N_{β} = number of source emitted betas or γ 's per unit time

= (SBETA of card ⑥ of report section 3.2) $\div$ 3.7×10^{10} .**

and

$$G = \Omega_x / \Omega = 1/2(1 - r/(r^2 + R_x^2)^{1/2})$$

where

Ω_x = solid-angle subtended by the crystal at the source
geometric center.

Ω = total solid-angle at the source = 4π steradian

r = source to crystal distance

Subprogram GEØMTR returns all of the above data to the main program for
output.

* if not meaningful to code user, then input as, EBMAX = 1.0

** if not meaningful to code user, then input as, SBETA = $1.0/(3.7 \times 10^{10})$.

2.3 Code Constants

In this section the origin and meaning of certain constants coded into CUPED, are discussed. All attenuation coefficients used were those given in references (9-13). The discussion is carried through in alphabetic order of subprograms, except for General Discussion and MAIN which are presented first. Certain subprograms require no discussion.

General Discussion: Certain constant values appear periodically throughout CUPED. The values 2.35482, 2354.82, 0.3989423 and 1.065 are Gaussian or normal distribution constants⁽¹⁴⁾. The value 2.54 is for conversion from cm to inches. The values 60 (minutes/hr and seconds/min) and 1440 (minutes/day) are clock time conversion constants. The values 0.51097, 0.514, 1.173, 1.332, 1.368, and 2.754 are radioisotope peak energy values⁽¹⁵⁾ in MeV, for β^+ , Sr^{85} , Co^{60} , Co^{60} , Na^{24} , Na^{24} . Other values are either obvious or constants unique to code logic except as explained below.

MAIN: Certain constants required by CUPED subprograms are coded in MAIN and communicated by CØMMØN/CNSTNT/. The constant T 30 = 30.48 is source-to-crystal distance in cm for use by the code in instances where it (DOST(I)) is omitted from input. The constant T 75 = 0.75 is the coded thickness of Lucite absorber in cm^2/gm ; presently not coded for input, though it may easily be. The constants T 20 = 20.0, T 50 = 50.0, T 01 = .0001, T 38 = 3.81 and T 76 = 7.62 are used by the code if input (ØN, HITMAX, EPS, RX, H) left blank; this is discussed in Section 3.2. The constant T 1293 = 0.001293 is the density of air coded for subprogram AIRABS. The constants T 90 = 900000. and T 1000000. are used by subprograms DEC and SHAPE for checking for spectral counts in the complement mode. The constant, T 366 = 3.671 is the density of NaI(Tl) in gm/cc . The constant, T 3316 = 0.03316, is the energy of the K electron shell absorption edge for

iodine in NaI(Tl) , in MeV. The constant T 06 = 10^{-6} is an integration criterion for subprogram SIMPSN. The constants T 321 = .321 and T 7677 = .7677 are k and n of Equation (1). The constant T 285 = 0.0285 is the energy of the NaI(Tl) iodine k x-ray in MeV. The constants associated with UT at statement numbers 40 to 47 are explained in Section 3.2. The constant 3.7 E + 10 at statement number 128 + 1 is the conversion factor for Curies to disintegrations/second.

AIRABS: The mass absorption coefficients of air are given in the DATA statement in cm^2/gm . The coefficients include coherent scattering. They are multiplied by the density of air (T 1293) in gm/cc to give output units in cm^{-1} .

CLAD: The attenuation factors (a fraction) of the detector cladding material are given in the DATA statement. They were determined for the following material composition and thickness

Material	Density X Thickness gm/cm^2	Density gm/cm^3
Aluminum	130	2.70
Neoprene	43	1.30
Polythene	13	0.90
Aluminum Oxide	67	4.0

and the expression

$$\text{"Factor"} = e^{-\mu(E) \cdot \text{Thickness}}$$

where

$\mu(E)$ is the material weighted mass absorption coefficient.

DØSE: the energy mass absorption coefficients for air are given in the DATA statement in cm^2/gm . They are based at 20°C and a fractional weight composition of

Nitrogen	0.755
Oxygen	0.232
Argon	0.013

EFFIC: The total mass absorption coefficients are given in the DATA statement in cm^2/gm ; they do not include coherent scattering.

GEØMTR: The constant defined as CØNST has been already discussed for Equation (13).

PERSPX: The mass absorption coefficients for perspex (ie. Lucite) are given in the DATA statement in cm^2/gm .

SHAPE: The isotopic names and photopeak energies in MeV of the allowed standard source spectra are given in the DATA statement located at sub-program statement number 9095 + 5.

PULSE: The fractional deviations of pulse height at specific energies are given in the DATA statement. They are based on an analysis of experimentally measured standard source spectra whose photopeak energies are well known. A step-by-step description of such an analysis can be found in reference (16). The deviation values coded in PULSE are typical of the spectrometer system and not just characteristic of NaI(Tl).

RAXEL: The iodine K X-ray escape fractions derived from the equations of Axel⁽¹⁷⁾ for parallel beam geometry and as modified in reference (18) are given in the DATA statement for energies up to 0.150 MeV. For incident

photon energies greater than 0.150 MeV, the fraction is computed by RAXEL as

$$F_k(E) = 5.0233 \times 10^{-5} \times E^{-2.787}$$

and for $E > 0.5$ MeV, as

$$F_k(E) = 0.0$$

SIMPSN: The constant $TMAX = 2048.0 = 2^{11}$ (first execution statement) is a criterion to stop integration in the event of a non-convergence.

STDFIT: The constant $EPS = .00001$ (at statement 100 - 3), is the fitting criterion for the single peak non-linear regression. The constant $NI = 10$ is a stopping criterion for non-linear regression in the event of a non-convergence.

STDFT2: Constants EPS and NI (first and second statements) are as for STDFIT, for double peak fitting.

3. CODE OPERATING INFORMATION

3.1 GENERAL

Code CUPED is written in FORTRAN-IV for the IBM-360/91. It may be run on any IBM-360 with sufficient core size, ie. the present version requires 205,000 bytes, (4 bytes/word). There are no Sense Switch or special tape requirements. Input formats are standard FORTRAN-IV, as given in any IBM or CDC Fortran manual; the code has been designed with a view to ease of translation for use on other than IBM computers. Input/output tapes are presently coded as LI and LØ equal to 5 and 6, respectively, at the beginning of MAIN (statement 19181 + 2 and + 3). A code listing is given in Appendix II .

Figure 10 shows a general arrangement for the data input cards. Input card details, order, formats, restrictions and location are given in Section 2.3. Card numbers are encircled and defined in the order in which they are read by the code. A sample input listing is presented in Appendix III.

The code CUPED input data card deck consists of thirteen (13) types of cards, referred to as Card ①, Card ②, etc. If the type requires more than a single card the reference is made to Card Set ○. Card ① is a single card, input only once. Cards ②③ and ④ are single cards input at least once. Card set ⑤ (⑤A or ⑤B) is input at least once in order to define the response matrix. If the user only wishes to generate a response matrix but not to apply it to any data, then no further input is required. If the matrix is to be applied to analyze unknown spectra, further input is required to define the spectral data and the required analysis.

Card set ⑥ consists of two cards which must be input to define parameters which are common to all PHA spectra in the PHA spectral data set (card set ⑬), eg. source size, counting time, source strength, etc. Card set ⑦ must be input to define parameters unique to each PHA spectrum in the PHA spectral data set, eg. zero shift required, smoothing required or not, source-to-detector distance, energy calibration data, etc. Card ⑧ and card sets ⑨ through ⑫ are optional. Card set ⑬ consists of the unknown PHA-spectra to be analyzed and their PHA-background-spectra. The code will analyze spectra of up to 256 channels, although up to 512 channels may be input and code gain-changed to 256 for analysis.

The optional cards are as follows: Card ⑧ is a single card to allow the user to study the iterative unfolding convergence, ie. intermediate unfolding data is output. Card set ⑨ allows the user to replace undesirable peaks, prominences or spurious spectral counts with a straight-line shape. Card set ⑩ allows the user to load initial spectral channels with a straight-line shape. Card set ⑪ allows the user to input the energy of photopeaks to be analyzed. Card set ⑫ allows the user to input the channel region of photopeaks to be analyzed. The input of both card set ⑪ and ⑫ allows the user to give the energy of certain peaks and the channel region of others. Card set ⑫ allows the user to input energy as well as channel domain for peaks, in which case the code will use the input energy as opposed to the code determined energy for such as efficiency calculations. This last is useful in applications where a priori knowledge indicates that the energy which the code would determine would be too approximate.

The input of channel number values must be as recorded by the pulse-height analyzer. The code will change the input values in accord with requested shifts or gain changes. In Section 3.2, 'rounded-up' refers to the 'next highest integer value', i.e., 3.7 rounded-up is 4. The input order of card

sets ⑦ , ⑨ , ⑩ , ⑪ and ⑫ must correspond to the spectra of card set ⑬ .

The code output is reviewed in Section 3.3. Appendix IV is a sample output listing. It corresponds to sample input of Appendix III. Debug type output may be obtained by input of $M(8) = 8$ on card ③ . The user is cautioned with respect to profusion of output under this option - - - a trial using sample data is recommended first.

3.2 Card Input Details

Card ① (one card; once only)

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
SET	1-10	F10.5	Total number of spectra to be unfolded by a code run (= RUNS * ØJSØ * number of times card ② to ④ input)

Card ② (single card)

CASE	2-72 (column 1 for printer control)	A	User's problem description (alphanumeric)
------	--	---	--

Card ③ (single card)

M (1)	1-3	I3	Signal for routing after response matrix generation > 0 CALL EXIT = 0 Continue < 0 Return to READ card ②
M (2)	4-6	I3	= 0, use existing response matrix ≠ 0, generate new response matrix using existing standard spectra
M (3)	7-9	I3	If ≠ 0, read card ⑧ (iterative unfolding output signal)
M (4)	10-12	I3	If ≠ 0, read card ⑨ (replace peak with straight line)
M (5)	13-15	I3	If ≠ 0, read card ⑩ (dead channel fill-in)
M (6)	16-18	I3	If ≠ 0, add photopeak contributions to iteratively unfolded continuous photon number, else separate
M (7)	19-21	I3	If ≠ 0, SINGLE called for peak analysis

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
M (8)	22-24	I3	Debugging output if $\neq 0$
M (9)	25-27	I3	If $\neq 0$, the peaks analyzed by SINGLE are corrected for source decay
M (10)	28-30	I3	See Table I (routing option)
M (11)	31-33	I3	See Table I (routing option)
M (12)	34-36	I3	See Table I (routing option)
M (13)	37-39	I3	If $= 0$, output the GANE reduced PHA spectrum prior to unfolding
M (14)	40-42	I3	If $\neq 0$, ENLIN called before entry to SINGLE (subsequent ENLIN call by-passed)
M (15)	43-45	I3	If $\neq 0$, output the iterative unfolding convergence differences
M (16)	46-48	I3	If $\neq 0$, do <u>not</u> correct for source decay, i.e., assume decay factor $= 1.0$
M (17)	49-51	I3	If $\neq 0$, output the PHA spectrum
M (18)	52-54	I3	If $\neq 0$, by-pass final result computations, i.e., by-pass GEØMTR

NOTE: The choice of non-zero values required for M (1) is arbitrary, however, positive subscript index values will aid in identity, e.g., if M (7) $\neq 0$ then input as $= 7$; the user is cautioned that arbitrary negative values are allowed for M (11) and M (12) in accord with Table I.

Card ④ (single card; last 5 variables usually blank)

ELIMIT	1-10	F10.5	The energy of the upper edge of the response matrix highest channel, (MeV) (CHANLS * EG(J)/VG(J); see cards ⑥ and card set ⑦)
--------	------	-------	---

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
ØJSØ	11-20	F10.5	Loop limit; number of sets of source data (card sets ⑦ through ⑫, except ⑧, before loopback to READ card ②).
ØMM	21-30	F10.5	If < 0 READ a response matrix, (card set ⑤B), = 0 use already computed matrix, > 0 generate new matrix; CALL SHAPE. Choice of values are arbitrary, eg. -1. and +1.
ØN	31-40	F10.5	The size of the response matrix, ie. number of channels; also the size of final flux spectra, ≤ 40.0.
HITMAX	41-50	F10.5	The maximum number of unfolding iterations; an even number such as 50.0 unless iterating input per M (3) required. ≤ 100.0.
EPS	51-60	F10.5	Convergence tolerance at which iteration will cease. eg. .0002.
RX	61-66	F6.4	Radius of NaI(Tl) crystal, cm.
H	67-72	F6.4	Cylindrical length of NaI(Tl) crystal, cm.

NOTE: if ØN zero (or blank) code sets = 20.
 if HITMAX " (" ") " " = 50.
 if EPS " (" ") " " .0001
 if RX " (" ") " " 3.81 cm. (=1.5")
 if H " (" ") " " 7.62 cm. (=3")

Card ⑤

Card ⑤ refers to a deck of cards of which two kinds are allowable, namely: ⑤A or ⑤B.

Card set ⑤A will be input, if ØMM > 0; read by subprogram SHAPE and a response matrix generated.

Card set ⑤B will be input if ØMM < 0; response matrix input to program MAIN.

Neither set input if ØMM = 0, ie. a correct response matrix assumed as existing in storage.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
Card Set (5A) (NPHA, NSJ, NFNJ, NSXJ, NFXJ and SHIFT are defined in Figure (11).			
Card (5A) - 1 (single card) (See Figure 4)			
NSTAND	1-5	I5	The number of standard source spectra.
NPHA	6-10	I5	The count in the first NPHA channels of each standard spectrum are replaced by the count in channel NPHA + 1.
NXLIM	11-15	I5	Number of channels in a standard spectrum (≤ 260).
UNGAIN	16-25	F10.5	The reference coarse gain of the pulse-height analyzer. Use 1.0, 2.0, 4.0, 8.0, 16.0 or 32.0 (for future code extension).
Card Set (5A) - 2 (I= 1 to NSTAND cards) (See Figure 4)			
ALABEL (I), BLABEL (I)	2-6	2A3	Standard source identity. Must be one of CD109, SC47, HG203, CR51, SR85, CS137, MN54, NB95, ZN65, NA22, CØ60 or NA24 in card field; see Section 2.2.2.1
NSJ (I)	11-15	I5	A Gaussian-plus-straight-line function is fitted to standard spectra from channel NSJ to NFNJ.
NFNJ (I)	16-20	I5	A Gaussian-plus-straight-line function is fitted to standard spectra from channel NSJ to NFNJ.
NSXJ (I)	21-25	I5	X-ray peaks between channel NSXJ and NFXJ are subtracted from standard spectra; if NSXJ is negative the 0.51 MeV spectra of Na ²² and Zn ⁶⁵ are subtracted if they are present in the standard deck providing Sr ⁸⁵ (or a substitute 0.51 spectrum) also exists (-NSXJ and +NFXJ are .51 MeV peak channels defining this fitting range).

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
NFXJ (I)	26-30	I5	X-ray peaks between channel NSXJ and NFXJ are subtracted from standard spectra; if NSXJ is negative the 0.51 MeV spectra of Na ²² and Zn ⁶⁵ are subtracted if they are present in the standard deck providing Sr ⁸⁵ (or a substitute 0.51 spectrum) also exists (-NSXJ and +NFXJ are .51 MeV peak channels defining this fitting range).
SHIFT (I)	31-40	F10.5	The channel location ($\pm$) of the standard spectrum true zero pulse-height. The code carries out a shift correction.
Card Set <u>5A</u> - 3 (NSTAND times NXLIM/10 (rounded-up) cards); ((R (I,J), I = 1, NXLIM), J=1, NSTAND); (See Figure 4)			
R (1, 1)	1-7	F7.1	The count in the first channel of the first input standard spectrum.
R (2, 1)	8-14	F7.1	The count in the second channel of the first input standard spectrum.
.			
.			
.			
R (10, 1)	63-70	F7.1	The count in the tenth channel of the first input standard spectrum.
R (11, 1)	1-7	F7.1	The count in the eleventh channel of the first input standard spectrum.
.			
.			
.			
R (NXLIM, NSTAND)	63-70	F7.1	The count in the NXLIMth channel of the NSTANDth input standard spectrum.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
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The above Card Set (5A) - 3 may be summarized as:

10 channels of information/card per 10F7.1 format,
 NXLIM/10 (rounded-up) cards per spectrum; if NXLIM = 256
 then $256/10 = 25.6$ taken as 26, NSTAND spectra.

The spectrum input order must correspond with the order of
 card set (5A) - 2; Na^{22} precedes Zn^{65} precedes Co^{60} precedes
 Na^{24} . The order of the remaining spectra is immaterial. If
 either Zn^{65} and/or Na^{22} input then Sr^{85} must have been input.

Card Set (5B)

Card Set (5B) - 1 (($\emptyset N \times \emptyset N/5$) Cards); ((R(J,I), I=1, $\emptyset N$), J=1, $\emptyset N$)

R (1, 1)	1-11	E11.4	Response Matrix Element 1,1
R (1, 2)	12-22	E11.4	" " " 1,2
R (1, 3)	23-33	E11.4	" " " 1,3
R (1, 4)	34-44	E11.4	" " " 1,4
R (1, 5)	45-55	E11.4	" " " 1,5
R (1, 6)	1-11 (second card)	E11.4	" " " 1,6
.			
.			
.			
.			
R ($\emptyset N$, $\emptyset N$)	45-55	E11.4	" " " $\emptyset N, \emptyset N$

NOTE: The first $\emptyset N$ elements input represent the lowest energy matrix
 vector spectrum (analogous to a PHA spectrum); similarly, the
 second $\emptyset N$ elements, etc. The sum over each vector must =
 unity.

Card Set (5B) - 2 (($\emptyset N/2$) cards)

Q (1)	1-7	F7.4	Mid-increment energy of first energy interval of response of matrix.
Q (2)	8-14	F7.4	Mid-increment energy of second energy interval of response matrix.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
Q (10)	64-70	F7.4	Mid-increment energy of tenth energy interval of response matrix.
Q (11)	1-7	F7.4	Mid-increment energy of eleventh energy interval of response matrix.
Q ($\emptyset$ N)	64-70	F7.4	Mid-increment energy of $\emptyset$ Nth energy interval of response matrix.

Card Set (5B) - 3 (($\emptyset$ N/2) cards)

PV (1)	1-7	F7.4	Mid-channel value of first channel of response matrix.
PV (2)	8-14	F7.4	Mid-channel value of second channel of response matrix.
.			
.			
PV (10)	64-70	F7.4	Mid-channel value of tenth channel of response matrix.
PV (11)	1-7	F7.4	Mid-channel value of eleventh channel of response matrix.
.			
.			
PV ($\emptyset$ N)	64-70	F7.4	Mid-channel value of $\emptyset$ Nth channel of response matrix.

Card (6) (two cards)

First Card:

BTAG,BTAGA	2-6	A4,A2	Unknown source identity (alphanumeric)
SBETA	11-20	F10.5	Unknown source strength, curies; see page 23 footnote.
EBMAX	21-30	F10.5	Unknown source maximum or reference energy, MeV; see page 23 footnote.
CYLDIA	31-40	F10.5	Unknown source cylindrical diameter, cm.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
TH	41-50	F10.5	Unknown source half-life (optional units; see UT this card).
RUNS	51-60	F10.5	Number of spectra per unknown source data set, ≤ 20.0 .
CHANLS	61-66	F6.0	See footnote; ≤ 260
UT	67-72	F6.0	Multiplier for TH: UT = 0.0; TH in years = 1.0; TH in seconds = 60.0; TH in minutes = 24.0; TH in hours = 365.0; TH in days (Values other than these will cause output of error flag followed by CALL EXIT)

Second Card:

M222	1-5	I5	The number of peaks to be fitted with a Gaussian for which channel limits will be input (per card set (12)) ≤ 20 -M66; See card set (12) note.
M66	6-10	I5	The number of peaks to be fitted with a Gaussian for which only the energy is to be input (per card set (11)); ≤ 20 -M22.
MZ	11-15	I5	The number of channels, in each spectrum, to be loaded with zero counts (beginning at channel 1).
MNX	16-20	I5	See footnote; ≤ 512

Footnote: The code will expect unknown spectra of MNX channels (MBX for backgrounds) to be input; it will analyze as if CHANLS channels were input; if $257 \leq \text{MNX} \leq 512$ then unknown spectra are reduced to 256 channels by DEC calling GANE and analysis carried out on CHANLS channels; if $\text{MNX} = 0$, then CHANLS channels are input and analyzed.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
MBX	21-25	I5	Same as MNX, except that it refers to background spectra; see footnote; ≤ 512
MSM	26-30	I5	If $\neq 0$, smooth spectra before analysis; use 1.0 or 2.0 for single or double smoothing pass by sub-program GANE.

NOTE: BTAG/BTAGA, SBETA and EBMAX may be 'blank', 1.0 and 1.0 if not known prior to analysis. Actual values are used only for normalizing in sub-program GEOMTR prior to output of analysis results; see page 23 footnote.

Card Set ⑦

"Number of cards in set ⑦" = RUNS. Input of card I = 1 detailed below, cards 2 to RUNS similar. Card order must correspond to related pulse-height analyzer unknown (beta) spectra, $I \leq 20.0$.

DOST (I)	1-7	F7.3	Distance from geometric center of unknown source to front face of NaI(Tl) crystal, cm.
TETA (I)	8-14	F7.3	Polar angle of NaI(Tl) crystal axis with respect to the unknown source, degrees (a dummy variable for future use).
FIE (I)	15-21	F7.3	Azimuth angle of NaI(Tl) crystal axis respect to the unknown source, degrees (a dummy variable for future use).
DELT (I)	22-28	F7.3	Live time counting duration of unknown spectrum, minutes.
TM1 (I)	29-35	F7.3	Time duration from reference time to start of counting, days.
COGAIN (I)	36-42	F7.3	Pulse-height analyzer coarse gain setting for unknown spectrum, ≥ 1.0 .
VG (I)	43-49	F7.3	Monitor pulse-height corresponding to EG(I), channels.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
EG (I)	50-56	F7.3	Monitor energy corresponding to VG(I), MeV.
TTZ (I)	57-63	F7.3	Pulse-height analyzer channel location of true zero pulse-height, channels.
BK (I)	64-70	F7.3	BK(I) times a background spectrum may be subtracted from the unknown spectrum (if BK(I) negative, then is added).
BNBK (I)	71-72	F2.0	Background spectrum signal: < 0 subtract previously stored background = 0 no background > 0 read and subtract background Background spectra are read following unknown spectra with which they are associated.

Card (8) (one card input of M (3) $\pm$ 0)

This card may contain up to 18 integer numbers, which are the iterating or unfolding loop at which intermediate output is desired. $\leq$ eighteen indices may be input. The card format is 18I4

Example:

MN (1)	1-4	I4	Iterating loop index eg. 3
MN (2)	5-8	I4	" " " eg. 5
MN (3)	9-12	I4	" " " eg. 9

will cause subprogram RESMAT to output on iterating loops 3, 5 and 9.

Card Set (9) (1, 2, 3 or 4 cards input if M (4) $\pm$ 0) (See Figure 12)

N1X (1)	1-5	I5	Replace a peak (or other prominence in spectrum 1 of card set (13) with a straight line from channel N1X(1) to N2X(1); count rate in N1X(1) and N2X(1), used to determine slope and intercept of straight line.
---------	-----	----	---

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
N2X (1)	6-10	I5	See Above
N1X (2)	11-15	I5	See Above
N2X (2)	16-20	I5	See Above
.			
.			
.			
N2X (5)	46-50	I5	See Above
N1X (6)	1-5	I5	See Above
.			
.			
.			
N1X (RUNS)	—	I5	Replace a peak (or other prominence in spectrum RUNS, card set ⑬, with a straight line from channel N1X(J) to N2X(J); count rate in N1X(J) and N2X(J), used to determine slope and intercept of straight line; J = RUNS.
N2X (RUNS)	—	I5	See Above

NOTE: N2X (5) would be the tenth and last field of first card; the number of cards(≤ 4) depends on RUNS.

Card Set ⑩ (1 to 5 cards input if M (5) $\neq$ 0) (See Figure 12)

NFILL (1)	1-5	I5	Replace in spectrum 1 of card set ⑬, channels 1 to NFILL (1) with a straight line of slope VAL (1) based on count in channel NFILL (1).
VAL (1)	6-15	F10.5	See Above
NFILL (2)	16-20	I5	See Above
VAL (2)	21-30	F10.5	See Above
.			
.			
.			
VAL (4)	51-60	F10.5	See Above
NFILL (5)	1-5	I5	See Above
.			
.			
.			
.			

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
NFILL (RUNS)	—	I5	See Above
VAL (RUNS)	—	F10.5	See Above

NOTE: VAL (4) would be the eighth and last field of first card; the number of cards (≤ 5) depends on RUNS.

Card Set (11) ($1 * \text{RUNS}$ to $4 * \text{RUNS}$ cards input if $M66 \neq 0$) (See Figure 12)

This card set is input only when $M66 \neq 0$. It must consist of $\text{RUNS} * (M66/5 \text{ rounded up})$ cards corresponding to the number of unknown energy-specified spectral peaks in card set (13). Each card contains ≤ 5 photopeak photon energies corresponding to the unknown complex-plus-continuous spectrum to be analyzed. From one to 20(= M222) photopeaks may be energy specified per spectrum ie. a maximum of 4 cards per spectrum.

EUK (1, 1)	1-10	F10.5	Energy (MeV) of the first energy specified photopeak in card set (13) first spectrum.
EUK (2, 1)	11-20	F10.5	Energy (MeV) of the second energy-specified photopeak in card set (13) first spectrum.
EUK (3, 1)	21-30	F10.5	Similar to above.
EUK (M66, 1)	—	F10.5	Similar to above.
EUK (1, 2)	1-10	F10.5	Similar to above.
.			
.			
.			
EUK (M66, RUNS)	—	F10.5	Similar to above.

NOTE: The first photopeak in each spectrum is that of lowest energy with the remainder being in energy ascending order. The user is cautioned that peak channels may also be input instead or in addition, per card set (12) if $M222 \neq 0$; $M222 + M66 \leq 20$.

Card Set (12) ($1 * \text{RUNS}$ to $7 * \text{RUNS}$ cards input if $M222 \neq 0$) (See Figure 12)

This card set is input only when $M222 \neq 0$. It must consist of ' $\text{RUNS} * (M222/3 \text{ rounded-up})$ ' cards corresponding to the number of unknown channel-limit specified spectral photopeaks in card

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
set (13) . Each card contains information for ≤ 3 photopeaks. The information advises code of whether peak is a single peak or is instead one of a pair, of the upper and lower fitting limits (channel) and of the energy if it is not known.			
NJJ (1, 1)	1-4	I4	Signal for first channel-limit- specified peak of first card set (13) spectrum: = 1, if a single peak = 2, if one of two peaks in a pair.
NSSS (1, 1)	5-8	I4	Channel number defining fitting limit on low energy side of first channel-limit-specified peak, ie. fit peak from channel NSSS(1, 1) to NFNNN(1, 1).
NFNNN (1, 1)	9-12	I4	Channel number defining fitting limit on high energy side of first channel-limit-specified peak, ie. fit peak from channel NSSS(1, 1) to NFNNN(1, 1).
EXRAY (1, 1)	13-20	F8.4	Energy of first channel-limit- specified peak, MeV .
NJJ (2, 1)	21-24	I4	Similar to NJJ (1, 1).
.			
.			
EXRAY (3, 1)	53-60	F8.4	Similar to EXRAY (1, 1).
NJJ (4, 1)	1-4	I4	Similar to NJJ (1, 1).
.			
.			
EXRAY (M222, RUNS) —		F8.4	Similar to EXRAY (1, 1).

NOTE: The user is cautioned that peak energy data may also be input instead or in addition, per card set (11) if $M66 \neq 0$; $M222 + M66 \leq 20$. If $M66 = 0$ and $M222 \neq 0$ then photopeaks data may be input in any energy order; if $M66 \neq 0$ and $M222 \neq 0$ then input is expected in energy ascending order. Where data is input for double peaks $NJJ(I, J)$ and $NJJ(I, J + 1) =$

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
-------------	---------------	---------------	------------------------------------

2 and 0, respectively, then energy ascending order is expected if EXRAY (I,J) specified. An example of input data for double peak (I and I + 1), is given (indices omitted):

<u>NJJ</u>	<u>NSSS</u>	<u>NFNNN</u>	<u>EXRAY</u>	<u>NJJ</u>	<u>NSSS</u>	<u>NFNNN</u>	<u>EXRAY</u>
2	46	—	0.501	0	7	66	0.575

This specifies that the code shall carry out a double peak analysis between channels 46 and 66, that peak energies are 0.501 and 0.575 MeV and that the peaks are approximately 7 channels apart. The code analysis will determine the actual separation distance and thus 7 is given only as an estimate.

Card Set (13)

The number of cards in this set = (CHANLS (or MNX) * RUNS/10.0 + the number of background cards if any)*. The cards will contain the unknown source spectra to be unfolded. The number of spectra which may be input is limited by the DIMENSION (20) = RUNS. The spectra, corresponding to card sets (6) and (7), may be stacked together. A background spectrum, if input, must directly follow the unknown spectrum from which it is to be subtracted. Twenty unknown spectra, each followed by a background spectrum, are regarded as twenty spectra from the standpoint of 20 being the maximum number. Each spectrum contains CHANLS (or MNX) channels and background spectra must correspond*. Each card contains 10 channels of information. Thus, the following is typical of card set (13) as read by subprogram DEC:

S (1)	1-7	F6.0	Pulse-height analyzer count in channel 1.
S (2)	8-13	F6.0	Pulse-height analyzer count in channel 2.
S (3)	15-20	F6.0	Pulse-height analyzer count in channel 3.
.			
.			
S (10)	64-70	F7.1	Pulse-height analyzer count in channel 10.

<u>NAME</u>	<u>COLUMN</u>	<u>FORMAT</u>	<u>DESCRIPTION, PURPOSE OR USE</u>
S (11)	1-7	F7.1	Pulse-height analyzer count in channel 11.
.			
.			
.			
S (CHANLS)*	—	F7.1	Pulse-height analyzer count in channel CHANLS.*

(last card in spectrum)

* For input purposes, if $MNX > CHANLS$ then replace CHANLS above with MNX for unknown and MBX for backgrounds, ie. CUPED will read spectra up to 512 channels in size (per MNX and MBX) but will immediately reduce them to $CHANLS = 256$.

3.3 CODE OUTPUT

Throughout the discussion in this section, reference to Appendix V, Sample Code Output Listing, is necessary and understood. Those outputs which are clearly defined by format headings are either not discussed or are mentioned only briefly. Output pages are referred to through the encircled letters A, B, C, etc.

A. The values on this page are output by MAIN, and are as input on card sets (1) to (4), with the exception of EM(=EN/ELIMIT) and the obvious modifications to NaI(Tl) crystal dimensions (note the units are output in inches).

B. The values on this page are output by SHAPE, and correspond to those standard source spectral parameters input on card sets (5A) - 1 and - 2. Indicated channel numbers are those values after shifting with respect to true zero channel has been carried out.

C. The values on this and following similar pages are the standard source spectral counts corrected for input in the complement mode and true zero channel. This output by SHAPE corresponds to card set (5A) - 3 input.

D. The results of the Gaussian function regression analysis by STDFIT for the standard spectra photopeaks are output on this page by RESGEN. The output is self-explanatory.

E. The output on this and the following similar pages, by SHAPE, consists of the Compton continua of the standard source spectra normalized with respect to unit photopeak area and a pulse-height of 100 channels.

F. This page presents the response matrix generated by SHAPE and output by MAIN. It corresponds to that input which would be required for card set (5B) - 1.

G. This page presents the energies (MeV), pulse-heights (channels) and photofractions corresponding to the generated response matrix, at increment midpoints, as determined by SHAPE and output by MAIN. The energies and pulse-heights correspond to that input which would be required for card sets (5B) - 2 and - 3. The photofractions correspond to the solid curve in Figure 6.

H. The output on this page, by MAIN, corresponds to the input specified for card sets (6) and (7), excepting that the units in some cases are modified before output.

I. Optional output by MAIN giving the indices for which unfolding iteration output has been requested by input of card (8).

J. Output by MAIN of the options requested through input of card sets (9), (10), (11) and/or (12).

K. The output on this page by MAIN corresponds to the (first) spectrum to be analyzed and as input on card set (13). Background spectrum subtraction and complement mode correction is carried out before output.

L. Output of spectrum before entry to SINGLE; in the example the result of smoothing is demonstrated.

M. The output on this page by SINGLE is self-explanatory and refers to the fitting of an input specified monoenergetic spectral component of the unknown spectrum.

N. The output gives the PHA spectrum after stripping of photopeaks and associated continua by SINGLE; gain parameters for subsequent unfolding are also output.

O. The optional output on this page by MAIN corresponds to the unknown spectrum after gain changing and before unfolding analysis.

P. The output on this page by RESMAT is that requested by input of card (8). It consists of the gain changed unknown spectrum normalized to unit integral count; output at loop IT, corresponding to that requested (per MN), of the determined photon number spectrum (PHI) and the iterated input spectrum (PP); the iterated spectrum and the iteration convergence loop (IT), the normalizing integral count (SU) and the final value of the iteration arresting criterion term ($TERM = \chi^2$, Pearson's Chi Square).

Q. The output on this page by XTAL is self-explanatory and consists of the components of the diagonal efficiency matrix, η , defined by equation (10) of section 2.2.5.

R. The optional output on this page by MAIN, consists of the efficiency corrected and unfolded spectrum after post-normalization and the rate of convergence or fitting differences, during unfolding ($\Delta\chi^2$).

S. The optional output on this page, by MAIN, is self-explanatory and consists of $N_x(E)$ and $I_x(E)$, as already discussed in section 2.2.6.

T. The output on this page, by MAIN, is self-explanatory and consists of SUMNUM, SUMENY, D (Equation 14), AVENGY, etc. in the order of, and as already discussed in, section 2.2.6.

4. SUMMARY AND CONCLUSIONS

A FORTRAN-IV IBM-360 package code has been developed for the rapid analysis of bremsstrahlung spectra. The code is readily applicable to the analysis of sodium-iodide scintillation detector complex continuous spectra. The response matrix generating portion of the code is suitable for use as a separate entity for problems in spectral analysis such as are encountered in the various fields of gamma spectrometry.

The code employs an iterative unfolding method which has been used successfully by its authors, N. E. Scofield^(4,5) and R. Gold⁽⁶⁾, by the present author^(1-3,18), and others⁽⁷⁾. While this method is approximate because of the iterative technique used, it generally is the most suitable where continuous spectra are involved. It is suggested that degree of accuracy be the subject of future work, wherein the iterative method results would be compared with results obtained by other methods. The best value of the matrix size consistent with non-oscillatory good results and computer efficiency would be of interest here⁽¹⁸⁾. Contract circumstances did not allow for detail studies during the development of code CUPED.

It is proposed that the response matrix generating portion of the code be made more versatile by studying the use of additional standard sources. Contract circumstances did not allow a detailed debugging of the code in the energy range 1.4 to 3.0 MeV. It is mentioned that during the code development, considerable difficulty was experienced in using Na²² because of the relatively high intensity of the 1.28 MeV photopeak relative to that at 0.51 MeV, thus Zn⁶⁵ is recommended instead. Further work is proposed with a respect to use of the Na²² spectrum above 0.6 MeV.

The 200 channel PuO_2 spectrum of Figure 7 represents only half of the experimentally measured spectrum. The complete 400 channel distribution is shown in Figure 13. This spectrum is repeated in Figure 14 with a non-linear pulse-height scale proportional to the detection system resolution ⁽¹⁹⁾; the peaks in both figures are numbered for comparison purposes. The advantages of this form of representation are self-evident. Photopeaks are given equal importance throughout the energy range of measurement. It is proposed that the further incorporation of an energy non-linear analysis unfolding method be considered for code CUPED.

It is concluded that the developed code ,CUPED , is an operable and useful addition to the field of scintillation spectrometry. It allows the semi-automatic generation of detector system response function matrices , spectral unfolding process and final analysis of unknown complex-continuous spectra to be carried out in a single computer run , i.e. , without human interfacing. It is very suited to on-line applications.

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FIGURES

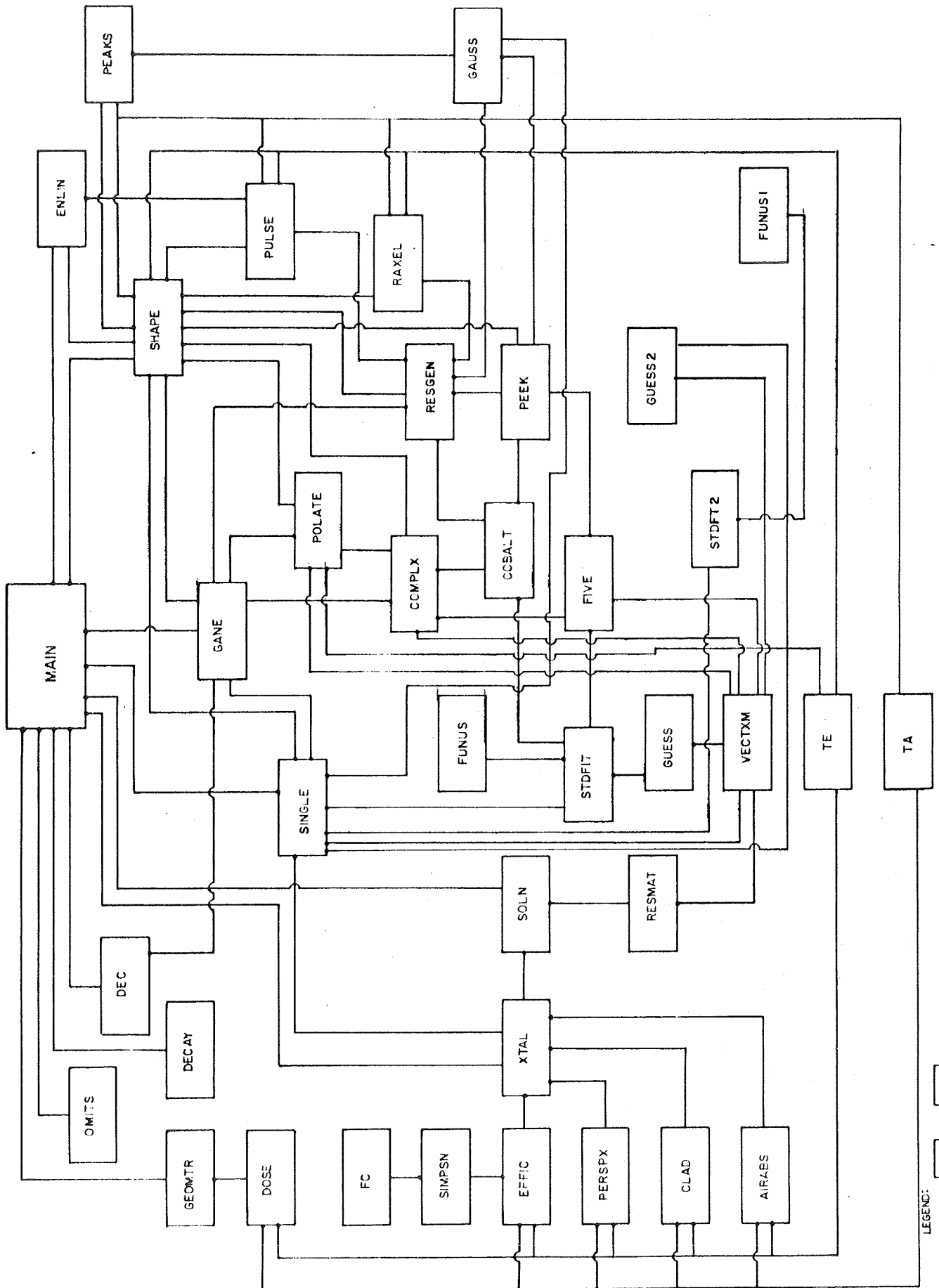


FIGURE 1
CODE CUPED SUBPROGRAM CONNECTIVITY

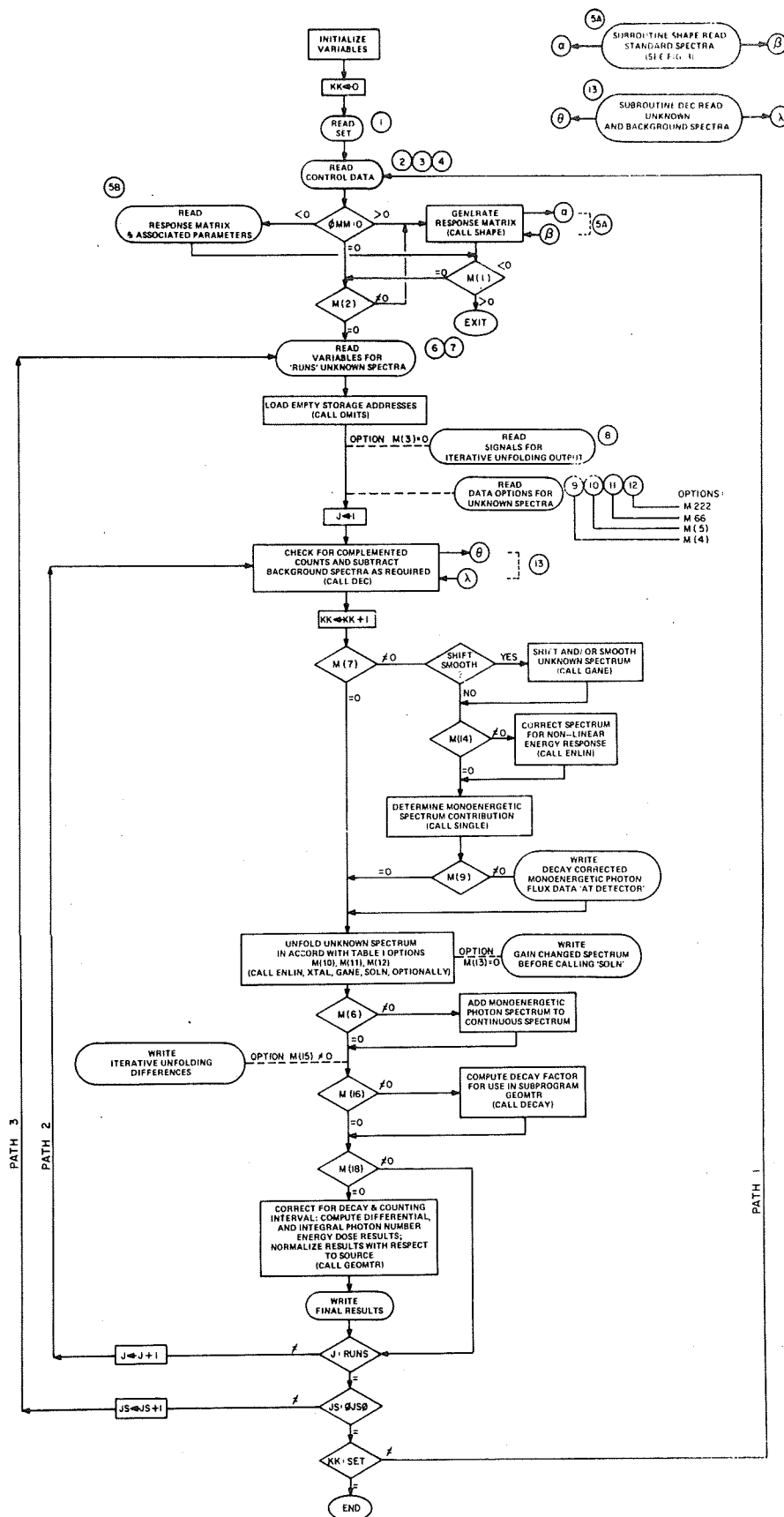
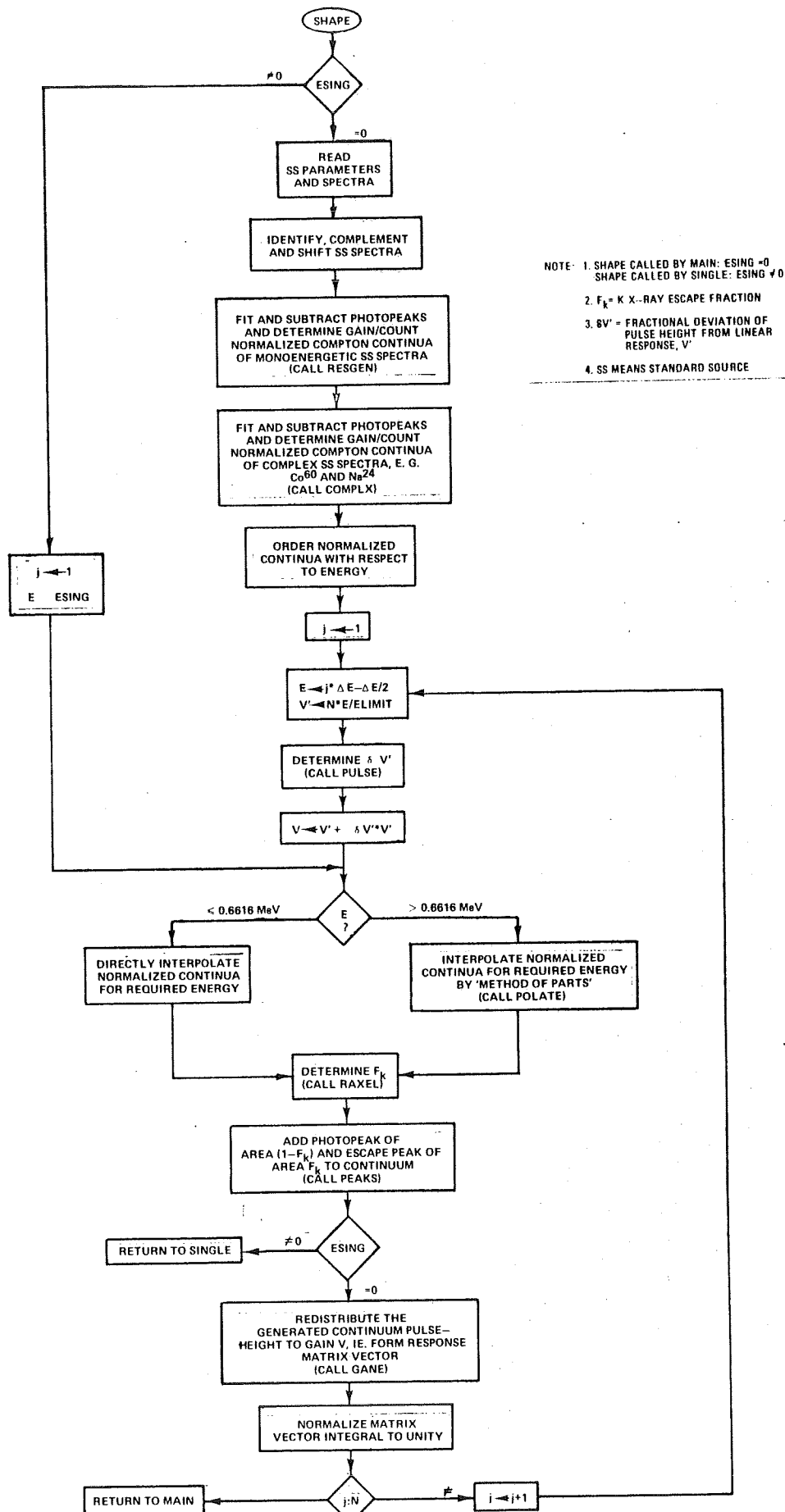


FIGURE 2
CODE CUPED SIMPLIFIED
FLOW DIAGRAM



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FIGURE 3
FLOW DIAGRAM OF RESPONSE MATRIX GENERATING
PROGRAM LOGIC;
SUBPROGRAM SHAPE

NSTAND = NO. OF SPECTRA
 NXLIM = NO. OF CHANNELS PER SPECTRUM
 (10 CHANNELS PER CARD)
 R(I,J) I = NXLIM
 J = 1, NSTAND
 E.G. FOR 8 SPECTRA J=1,8

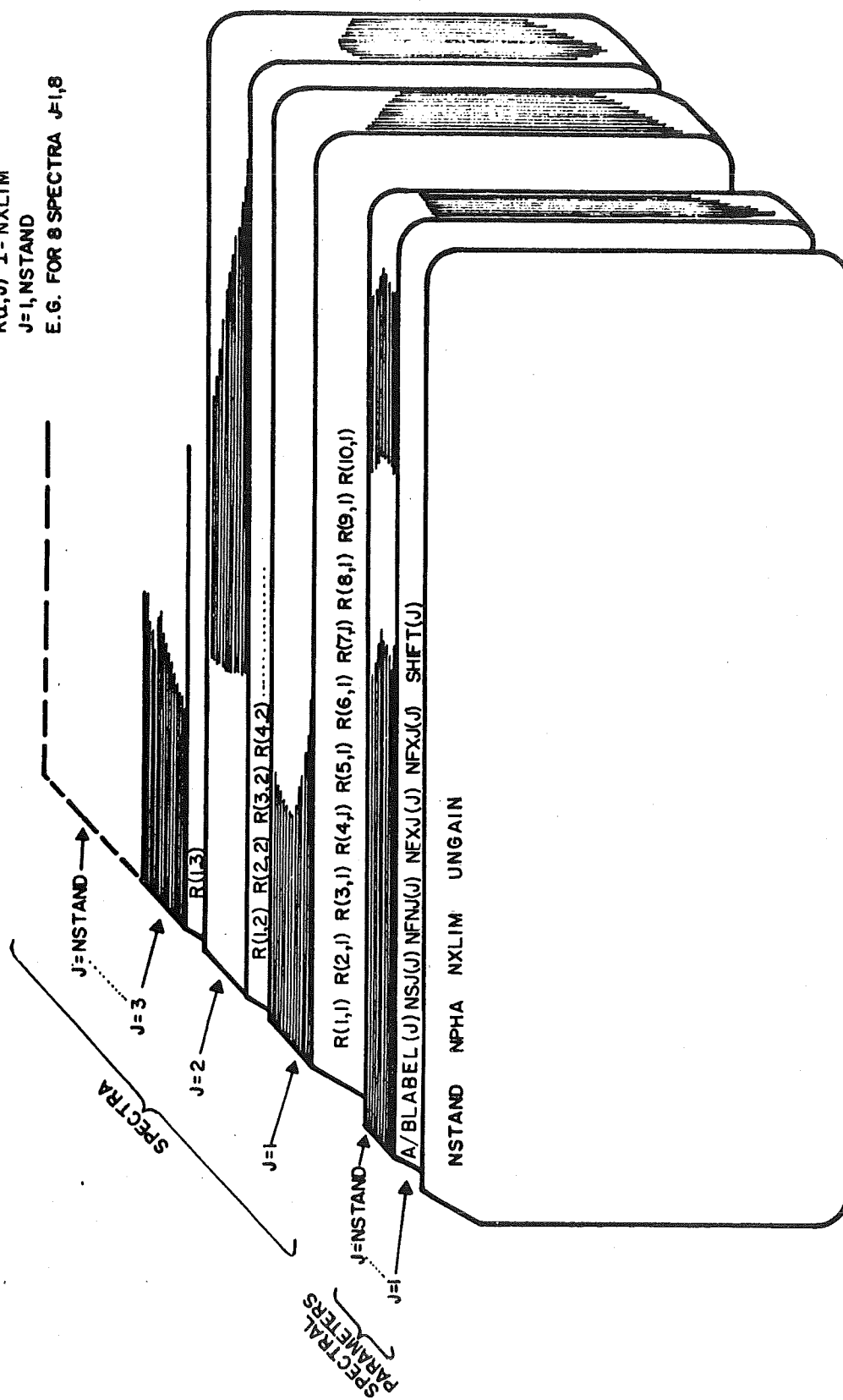


FIGURE 4
STANDARD SOURCE SPECTRA LIBRARY DECK ARRANGEMENT

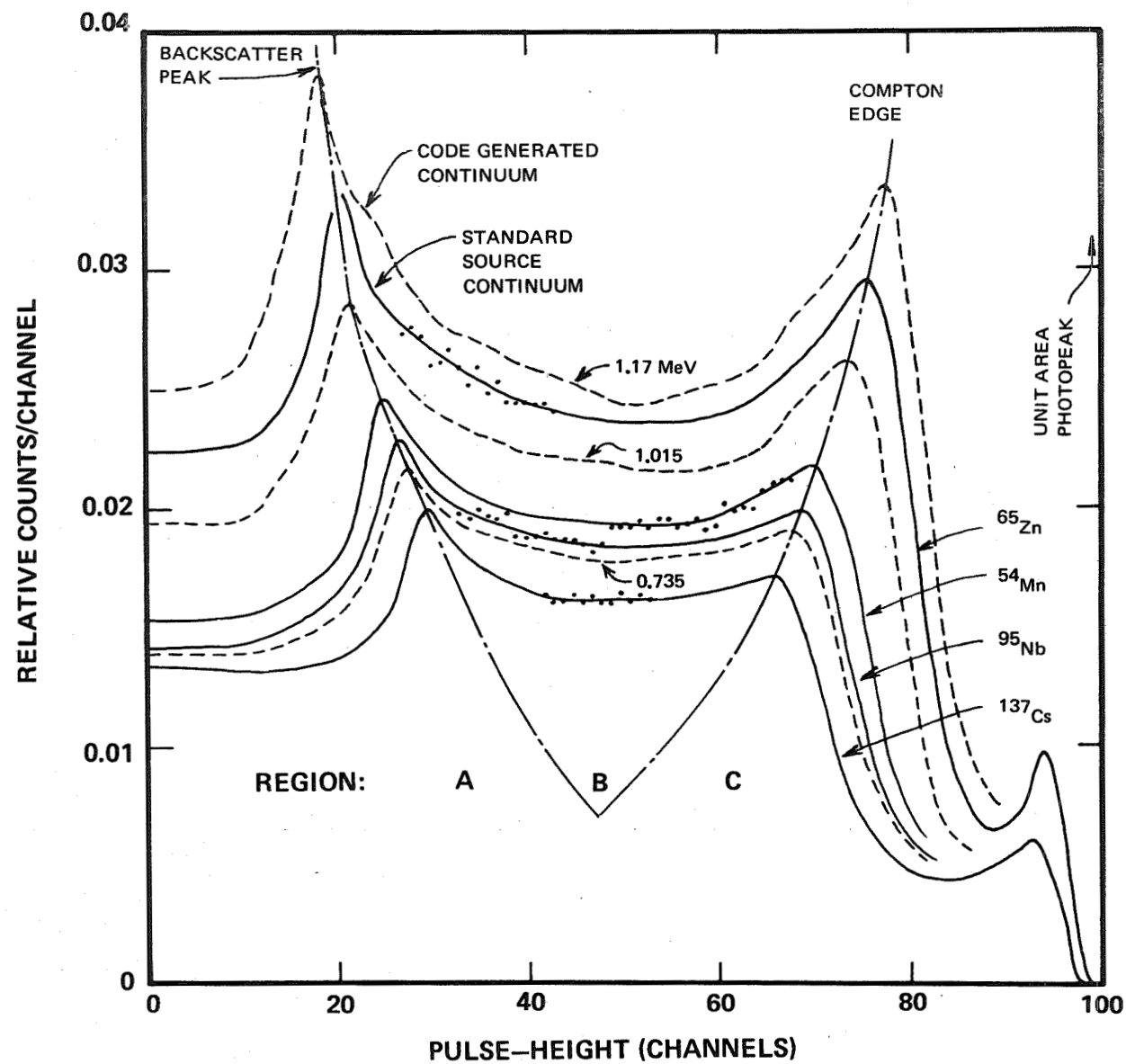


FIGURE 5
COMPARISON OF STANDARD SOURCE AND CODE GENERATED COMPTON
CONTINUA PULSE-HEIGHT NORMALIZED TO UNIT AREA PHOTOPEAK

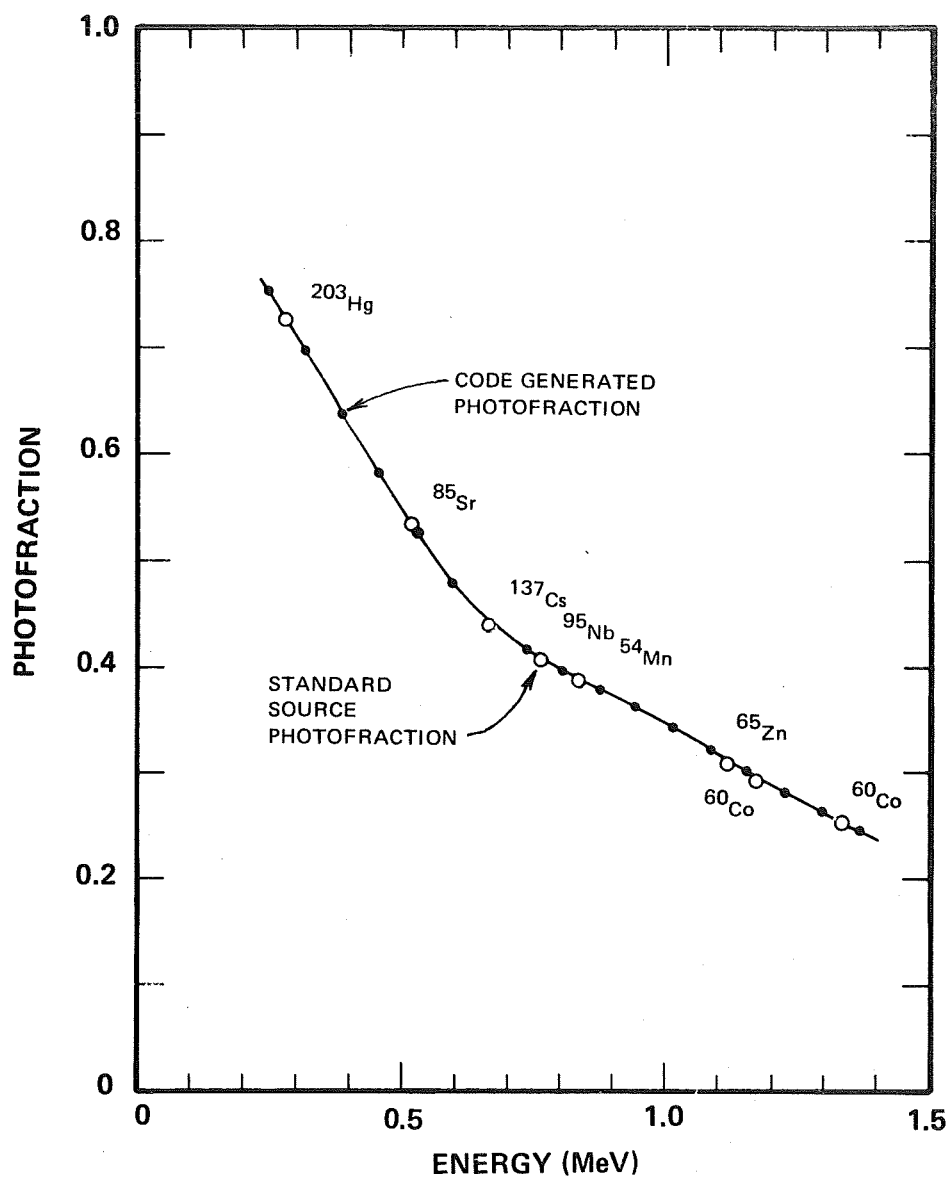


FIGURE 6

COMPARISON OF STANDARD SOURCE AND CODE
GENERATED RESPONSE MATRIX PHOTOFRACTIONS

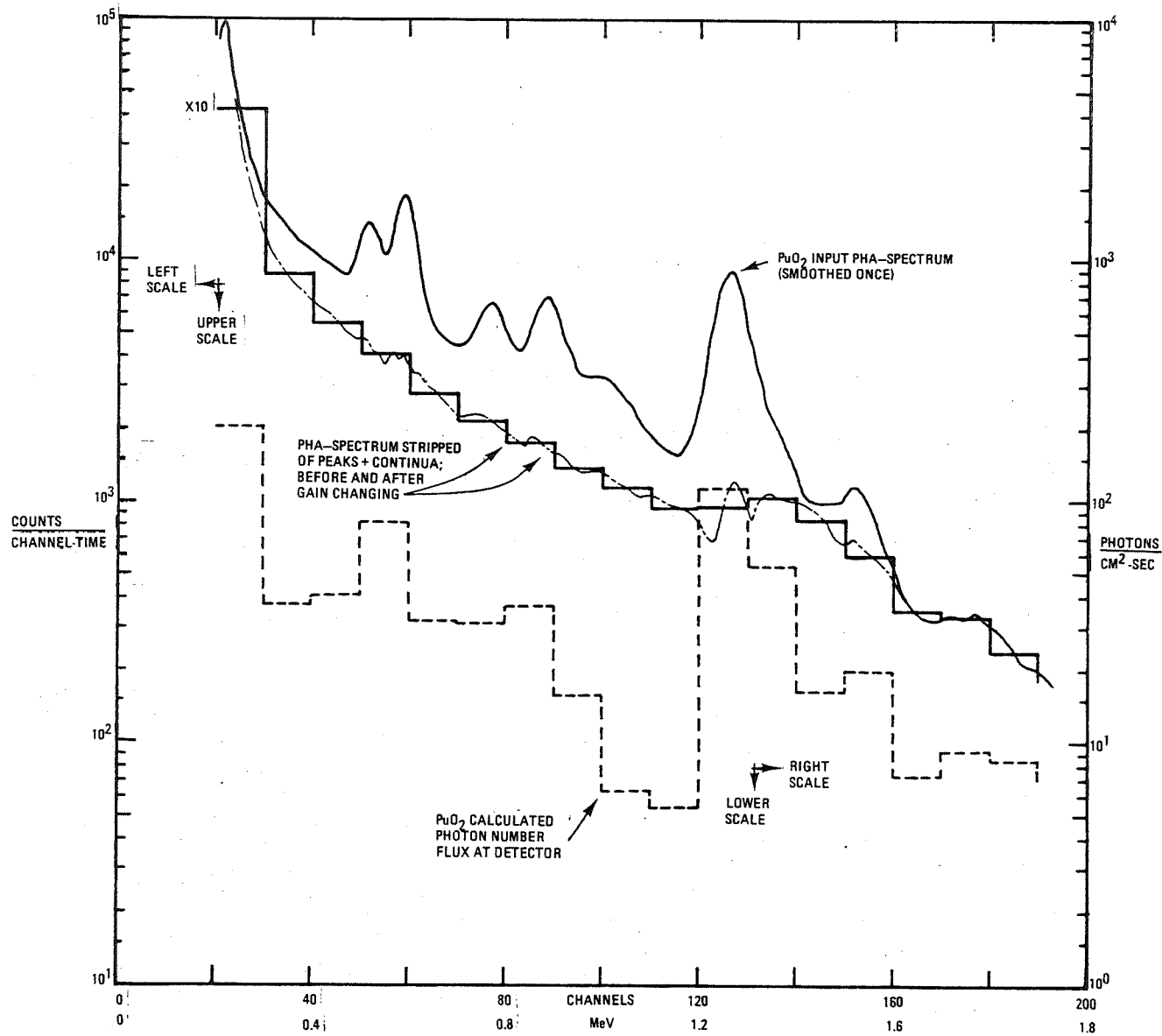


FIGURE 7
PLUTONIUM OXIDE SPECTRUM ANALYSIS

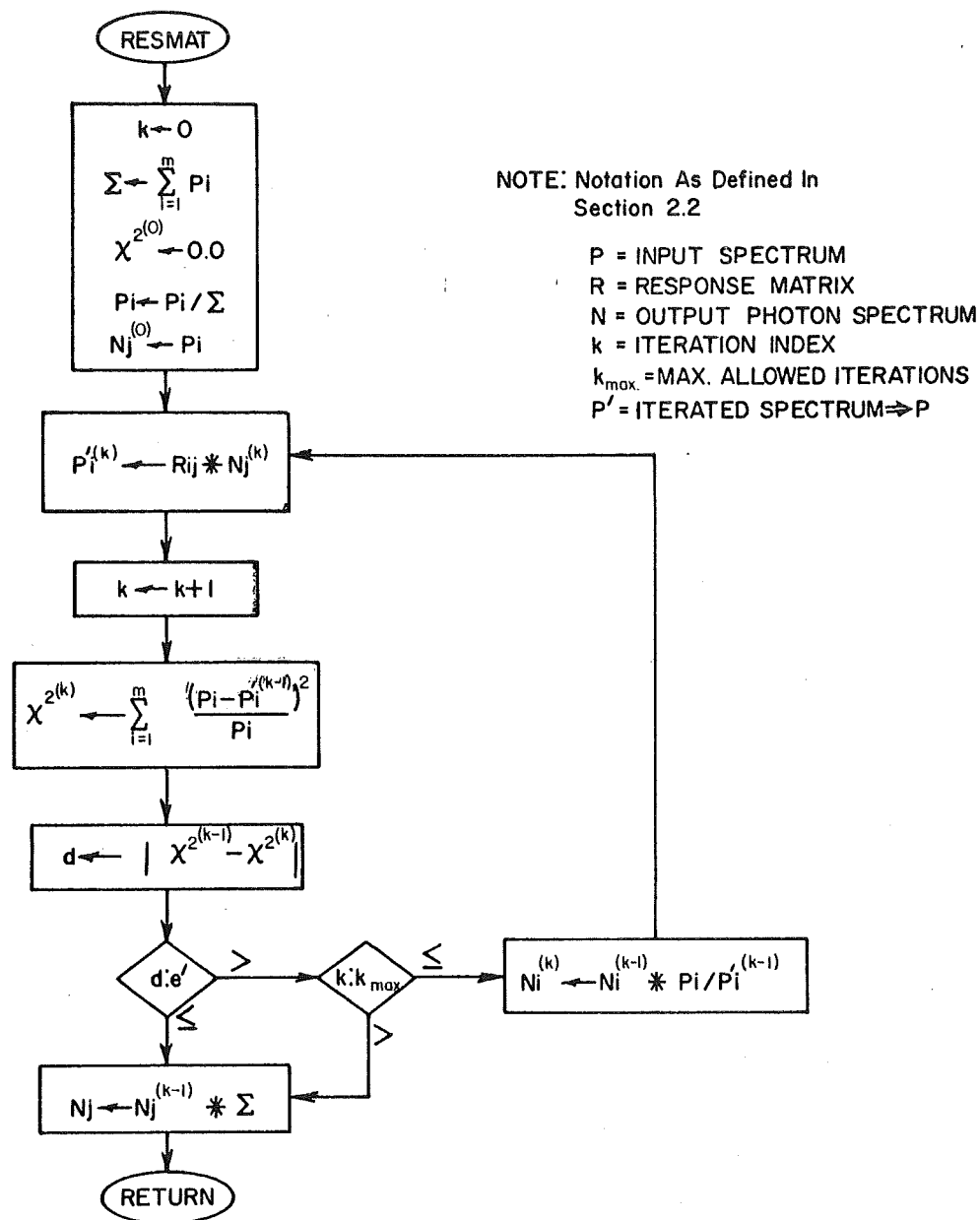
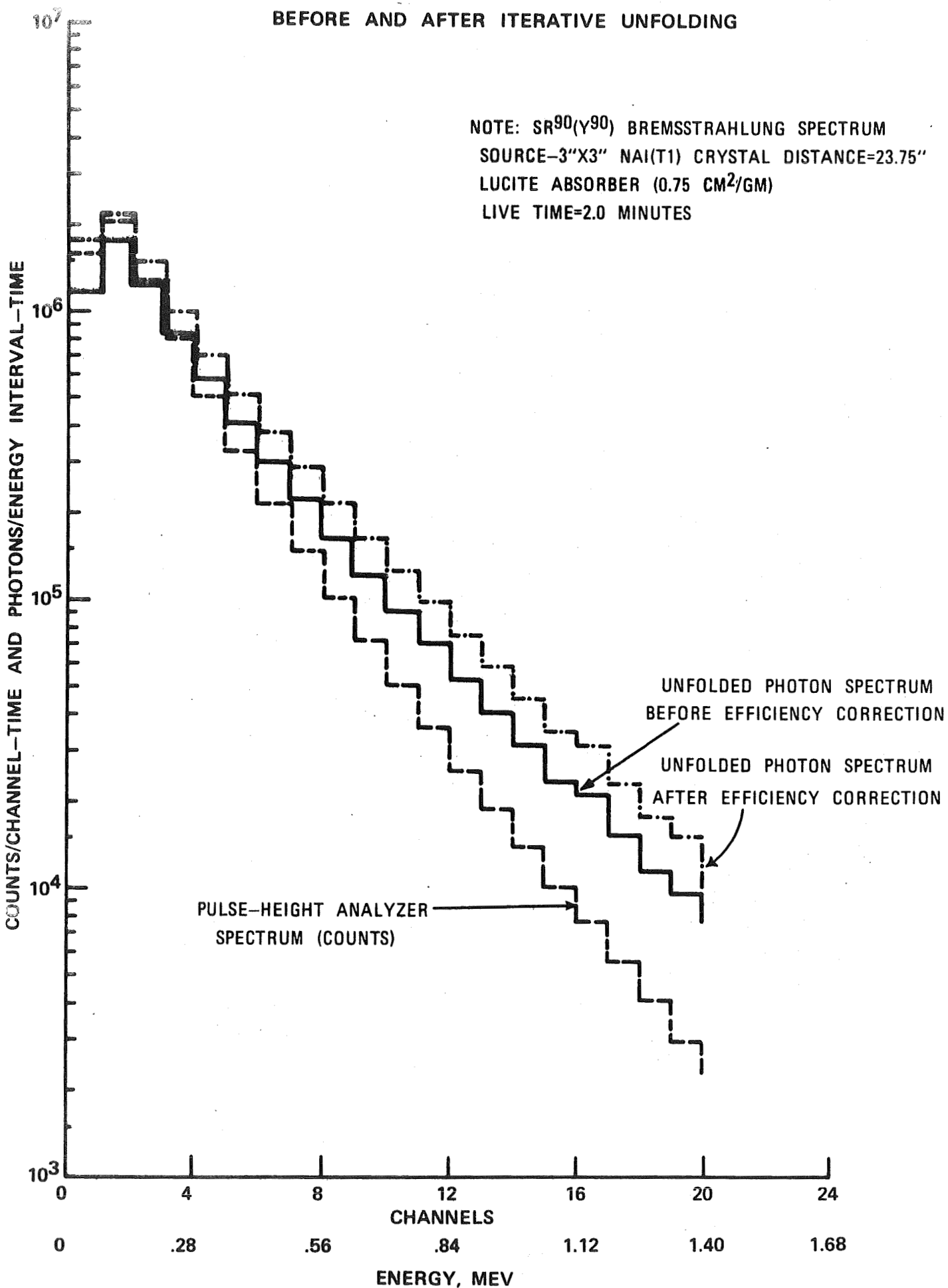


FIGURE 8

FLOW DIAGRAM SHOWING
THE GENERAL LOGIC OF
SUBPROGRAM RESMAT

FIGURE 9 BREMSSTRAHLUNG SPECTRUM
BEFORE AND AFTER ITERATIVE UNFOLDING



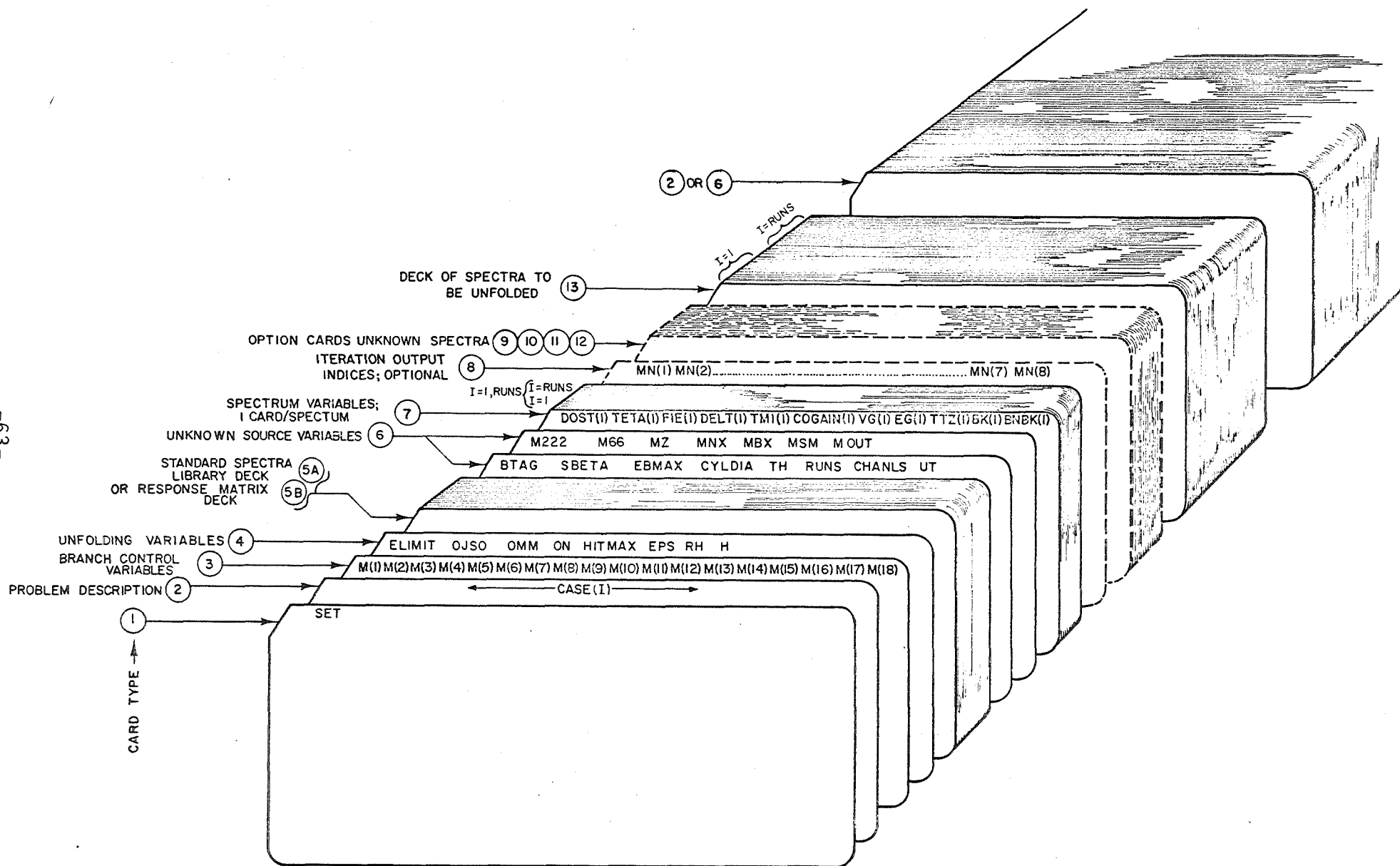


FIGURE 10
GENERAL ARRANGEMENT FOR INPUT CARD DATA DECK

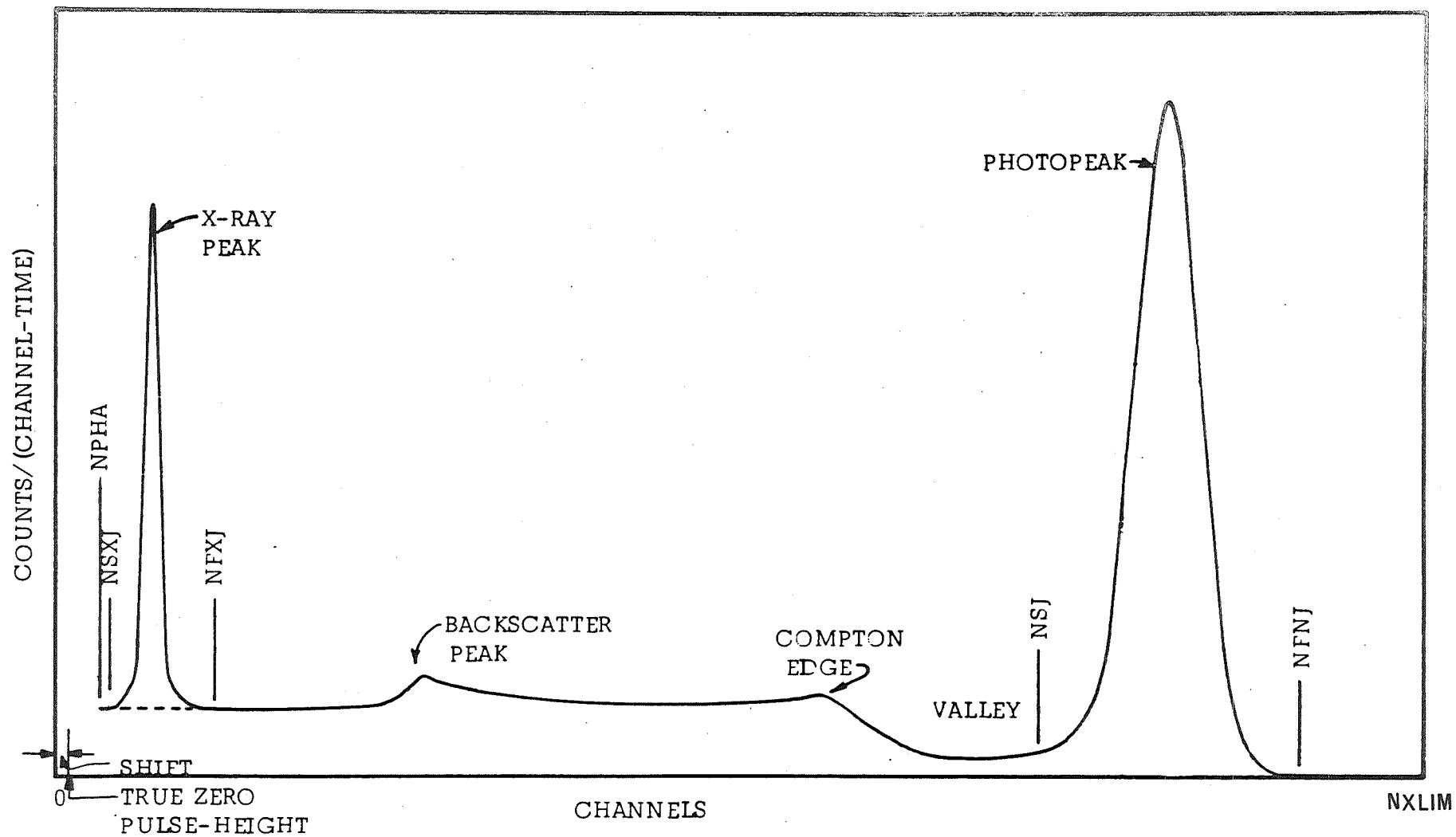
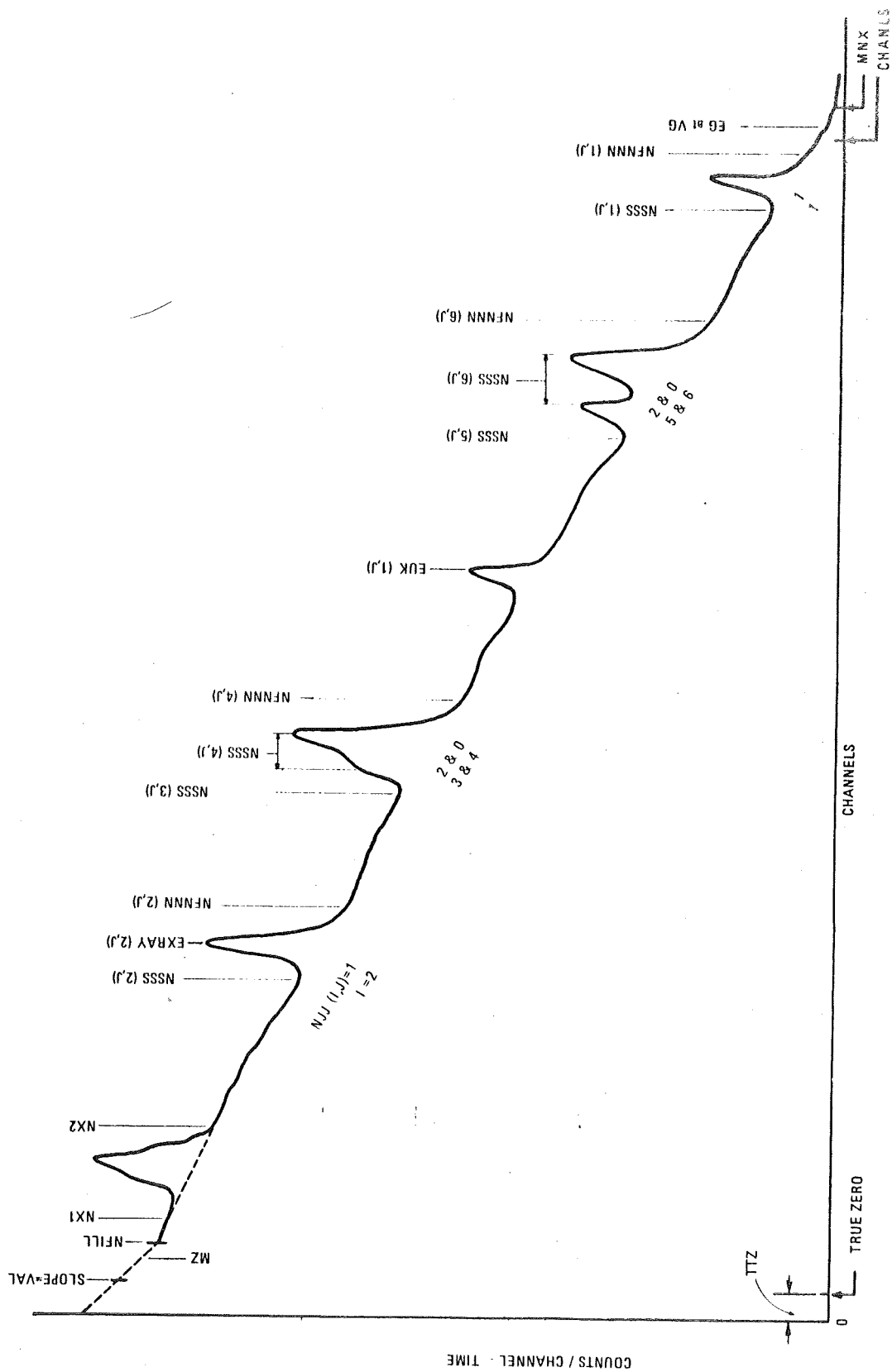


FIGURE 11
DEFINITION OF PARAMETERS FOR STANDARD
SOURCE SPECTRA DATA INPUT



NOTE:

M222 = 6

M66 = 1

EXRAY (1,J) = 0, EXCEPT EXRAY (2,J)

FIGURE 12

DEFINITION OF INPUT DATA PARAMETERS FOR UNKNOWN SPECTRA

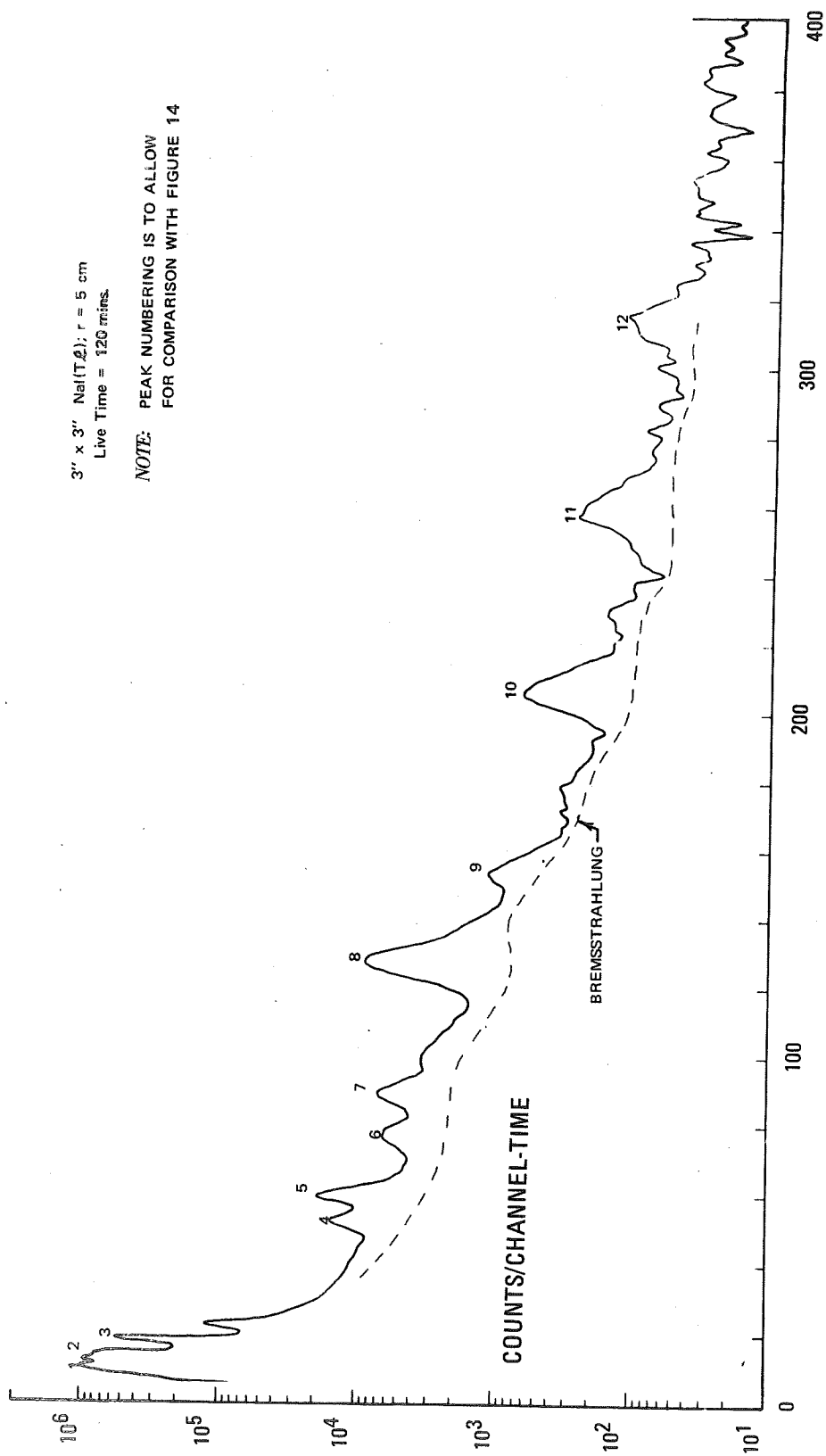


FIGURE 13
PLUTONIUM-OXIDE PHA SPECTRUM

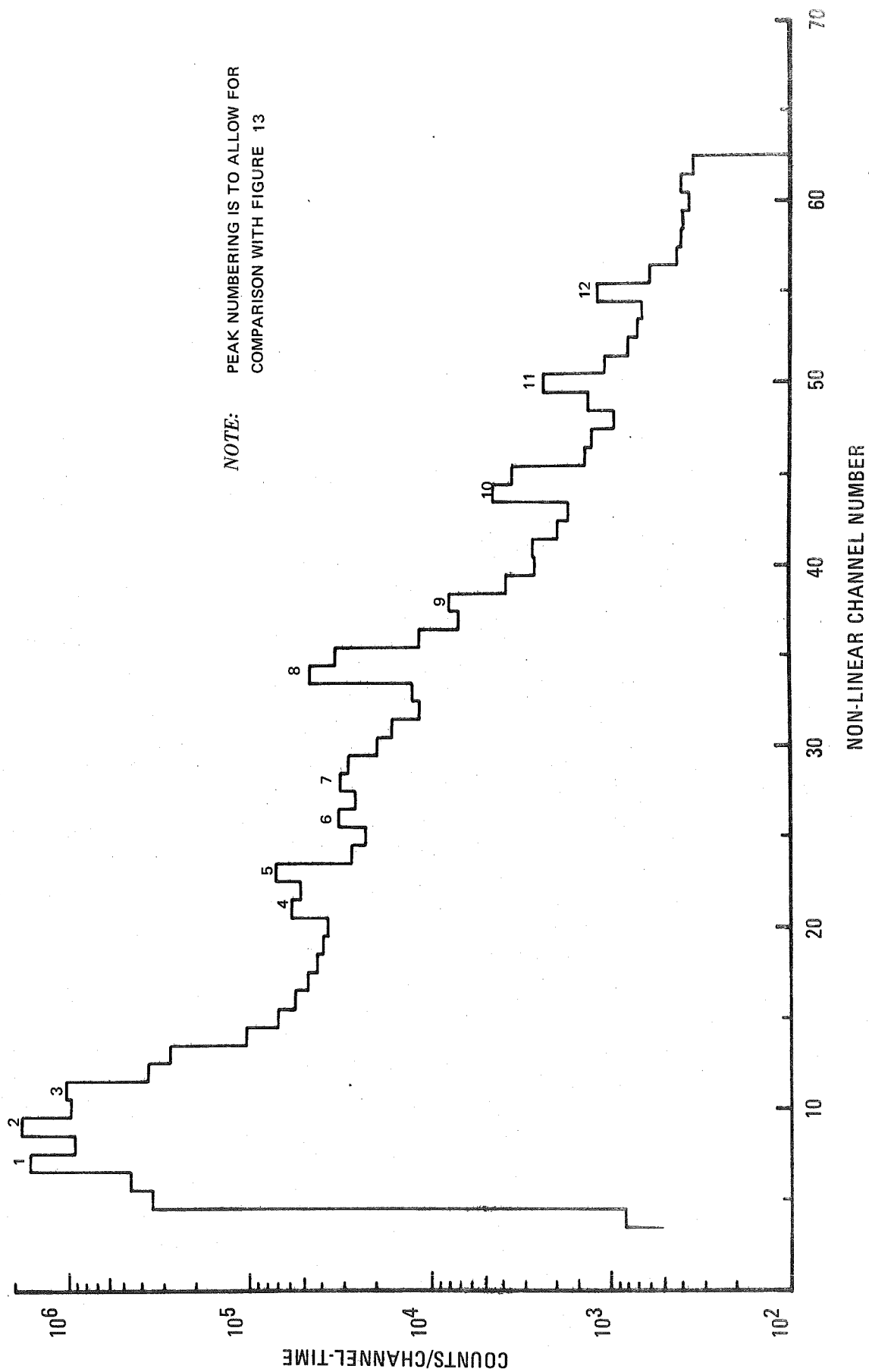


FIGURE 14

NON-LINEAR ENERGY INTERVAL REPRESENTATION OF PuO_2 PHA SPECTRUM⁽¹⁹⁾

Table I
Allowable Values of Input
Options M(10) , M(11) and M(12)

M(I)	Called Subprogram Operation Identifier (Defined Below)					
	1	2	3	4*	5*	6*
M(10)	0	0	0	10	0	10
M(11)	0	11	-11	11	11	11
M(12)	0	12	12	0	-12	12
Subprograms Called (See Appendix I for meaning)	GANE RESMAT XTAL	ENLIN XTAL GANE RESMAT	XTAL GANE RESMAT	ENLIN SØLN XTAL	ENLIN GANE RESMAT	ENLIN XTAL

Example: M(10) , M(11) , M(12) = 0 , 11 , 12 causes the code to correct an unknown spectrum for non-linear energy response , detection efficiency , gain change (e.g. , from 200 to 30 channels) and then unfold . N. B. M(14) dictates whether ENLIN is to be called before or after entry to SINGLE .

* Not normally useful , i.e. , for study or debugging purposes only .

APPENDIX I
GLOSSARY OF SUBPROGRAMS

APPENDIX 1

GLOSSARY OF PROGRAMS FOR CODE CUPED

(In alphabetical order, except for MAIN)

<u>NUMBER</u>	<u>NAME</u>	<u>FUNCTION or USE</u>
1	MAIN	Input, output and linking of subprograms.
2	AIRABS	Computes air attenuation factor. (F)
3	CLAD	Computes detector cladding attenuation factor. (F)
4	COBALT	Control program for photopeak fitting and subtraction.
5	CØMPLX	Analyzes Co ⁶⁰ and Na ²² standard spectra for SHAPE.
6	DEC	Reads pulse-height analyzer spectra; checks for PHA-complemented counts.
7	DECAY	Computes source decay correction factor. (F)
8	DØSE	Converts gamma photon flux to dose. (F)
9	EFFIC	Computes elements of detector interaction efficiency vector. (F)
10	ENLIN	Applies non-linear energy response correction factor.
11	FC	Crystal interaction efficiency function. (F)
12	FIVE	Controls fitting and subtraction of Na ²² escape and 0.51 MeV peaks.
13	FUNUS	Photopeak fitting function, partial derivatives and Chi-square term for STDFIT (one Gaussian plus straight line).

14	FUNUS II	Photopeak fitting function, partial derivatives and Chi-square term for STDFT 2 (two Gaussians plus straight line).
15	GANF	Gain changing program; also spectral shifting and smoothing.
16	GAUSS	Computes Gaussian photopeak for given parameters.
17	GEOMTR	Computes geometry factors, integrates number and energy spectra and calculates normalized dose data for final code results.
18	GUESS	Provides initial estimates of the photopeak function parameters for non-linear regression analysis in subprogram STDFIT.
19	GUESS 2	Provides initial estimates of the photopeak function parameter for non-linear regression analysis in subprogram STDFT 2.
20	OMITS	Code for insertion of repetitive variables omitted on all but first-card of set.
21	PEAKS	Adds photopeaks and escape peaks; computes photofractions.
22	PEEK	Computes Gaussian function ordinate. (F)
23	PERSPX	Computes perspex absorption factor. (F)
24	POLATE	Interpolates normalized Compton continua by the method of parts.
25	PULSE	Computes the detector system pulse-height at a given energy. (F)
26	RAXEL	Computes NaI(Tl) iodine K X-ray escape fraction. (F)
27	REGEN	Orders and normalizes standard spectra for response matrix interpolation.
28	RESMAT	Pulse-height analyzer spectrum unfolding according to the Scofield algorithm.

29	SHAPE	Control program for response matrix generation.	
30	SIMPSN	Simpson's rule integrating program for function FC.	(F)
31	SINGLE	Determines monoenergetic spectral contribution.	
32	SØLN	Control program for unfolding and detector efficiency correction.	
33	STDFIT	Non-linear regression analysis of standard spectra single photopeaks.	
34	STDFT 2	Non-linear regression analysis of standard spectra photopeak pairs.	
35	TA	Binary table searching program.	
36	TE	n-degree Lagrangian interpolation program.	(F)
37	VECTMX	Determines the index and value of the maximum valued element in a vector of elements.	
38	XTAL	Controls computation of detector total efficiency.	

APPENDIX II
FORTRAN CARD DECK LISTING

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```

WRITE(L0,701) (EUK(I,J),I=1,M66)
701 FORMAT(35H1 ENERGIES IN MEV OF UNKNOWN PEAKS // (1X,4F10.5))
1771 CONTINUE
771 CONTINUE
IF(M222) 1496,1495,1494
C READ SINGLE OR DOUBLE SIGNAL AND PEAK LIMITS (3 PXS/CARD) FOR M222
C PEAKS FOR *SINGLE* FITTING
1494 DO 772 J=1,NRIN
READ(L1,1496) (NJ(J,JP,J),N5SS(JP,J),N5NN(JP,J),EXPAY(JP,J),JP=1,
1M222)
1496 FORMAT(3(314,F8.4))
IF(MNX,LT,256) GO TO 772
DO 702 JP=1,M222
N5SS(JP,J)=N5SS(JP,J)*EACTOR
702 N5NN(JP,J)=N5NN(JP,J)*EACTOR
772 WRITE(L0,1490) (NJ(J,JP,J),N5SS(JP,J),N5NN(JP,J),JP=1,
1M222)
1490 FORMAT(//14H CARD SET 12 // (2X,3(314,F10.4)))
1495 CONTINUE
C
C MAIN EXECUTION LOOP FROM HERE
C
DO 500 J=1,NRIN
KK=KK+1
MENIN=0
GAIN=VG(J)=ELIMIT/EG(J)
DIST=DIST(J)
DT=DELTA(J)
TM=TM1(J)
RKG=RG(J)
NRK=NRK(J)
N=NX
JJ=J
CLAMDA=ALOG(2.0)/T4
C READ SPECTRA
C NOTE THAT SPECTRA AND RKGROS OF SIZE 257 TO 512 CHANNELS MAY BE READ
C BY SDECK, BUT ARE REDUCED TO SIZE NX WHERE NX IS LESS THAN 257.
C RKG=MULTIPLIER FOR BACKGROUND ADDITION/SUBTRACTION OPERATIONS
C NRK=SIGNAL SUBT/HOT READ/AND A BACKGROUND DEP -/0/4
C IF NRK =0 NEITHER READ NOR SUBTRACT A BACKGROUND.
C IF NRK NEG. THEN SUBTRACT THE CURRENT EXISTING BACKGROUND.
C IF NRK POS. THEN READ AND SUBTRACT RKG*BACKGROUND.
C SPECTRA READ AND COMPLEMENTED BY WDEC*
C
CALL DEC (NX,EM,MNX)
IF(NRK)6,2222,11
C
11 CALL DEC (NX,4,MNX)
C
4 DO 3 I=1,NX
3 EM(I)=EM(I)-RKG*EM(I)
2222 CONTINUE
WRITE(L0,25123) J1,RTAG,PTAG6,(E(I),I=1,NX)
25123 FORMAT(1H1,17= SPECTRUM NUMBER ,13,5H FOR ,A6,12,4H SMOOPE
25123133H (AFTER BACKGROUND SUBTRACTION) // (1X,10F8.0))
9 T2=TT7(J)
NT7=17
S0=0.0
C CALL *PEAKS TO SUBT AND/OR SUM WITH FIVE SPECIFIC ANALYSES IN *SAMPLE*.
44=RTAG
A0=0.53

```

```

IF(I7+40)1402,27,1402
1402 CALL GAIN (I7,0,1,0,1,0,AN,FM)
C
32 IF(M66) 602,601,602
602 DO 34 I=1,M66
34 FUC(I)=PHI(I)*GAIN/FLIMIT
601 CONTINUE
WRITE(LO,22333) (FUC(I),I=1,M)
22333 FORMAT(40H1 UNKNOWN SPECTRUM BEFORE PEAK FITTING //14X,10F10,0)
NSIN=N
36 IF(M14) 60,129,60
C
60 CALL ENLIN(N,FM,FC,VC,NGC,JD,NGC,0,FLIMIT)
MEMINI=1
129 CALL SINGLE(NSIN,GC)
C
WRITE(LO,22332)I7,N,GAIN,ROGAIN,SC,(FM(I),I=1,N)
22332 FORMAT(1H1,43P GAIN PARAMETERS AFTER CALL SINGLE IN MAIN
1/1X,4P,3,15,3,10,4/(1X,10F12,4))
1403 IF(M61) 300,306,300
300 WRITE(LO,302)
302 FORMAT(1H0,38X,43SUBTRACTED DISCRETE PEAK PHOTON NUMBER FLUX //
13X,5HINDEX,15X,4HENERGY,17X,11HNUMBER FLUX /
25X,5HMEV),14X,10H(PHOTONS/CM**2-SEC) //)
DE=DECAY(OT,TH,TW,CLAMBDA)
DO 301 I=1,N
PHOT(I)=(DE*PHOT(I))/(OT*60.83,141592653589793X)
WRITE(LO,303)I,FUC(I),PHOT(I)
303 FORMAT(1H,34X,11,2F25,7)
301 CONTINUE
DO 305 I=1,N
305 PHOT(I)=0.0
306 CONTINUE
9 DO 205 I=1,NLIMIT
205 PHI(I)=0(I)
124 IF(M10) 124,10,124
124 IF(M11)110,21,110
110 ME=1
MSKIP=1
GO TO 116
10 IF(M12)111,113,111
116 IF(M12)111,112,111
111 M14=M(12)
IF(M11)1209,113,112
209 M17=M(11)
GO TO 201
112 M17=M(11)
IF(MENI,N) 600,888R,600
C
888 CALL ENLIN (N,FM,FC,VC,NGC,JD,NGC,0,FLIMIT)
600 MEMINI=0
C
WRITE(LO,100) N,(FM(I),I=1,N)
100 FORMAT (24H1 MEASURIZED SPECTRUM OF ,16,2X,10H CHANNELS //11Y,10F
10019,0)
201 IF(M(12)1206,113,202
202 NGFI=1
C
CALL XTAL(NS,M,0,M)
C

```



```

1V(260),PARAV(12),R(5),PARAS(12),PARAF(12),NOSUHT(12)
NRA=0
WRITE(10,90)N1,N2,NST,NST02
C 90 FORMAT(10H N1,N2,ETC.,4I10)
IF(INR)122,120,121
122 ENCO=3,0
GO TO 130
120 ENCO=1,173
GO TO 130
121 ENCO=1,368
130 DO 5000 J=N1,N2
NRA3=1
INDIC=NLIMIT
IF(NST-2)233,235,233
235 JZ=J
IF(JZ-NST02)233,5001,233
233 NS=NSJ(J)
NEN=NENJ(J)
ENV=STDENY(J)
DO 2005 I=1,NLIMIT
2005 Y(I)=0
DO 5002 I=NS,NEN
5002 Y(I)=R(I,J)
JJ=J
R(3)=0
C CALL REGRESSION SURPROGRAM.
C CALL STFIT (Y ,NEN , R , NS , ENV)
C
PARAV(J)=R(1)
PARAS(J)=R(2)
PARAF(J)=R(3)
PAV=PARAV(J)
PAK=PARAS(J)
C SURTRACT FITTED PEAK.
C
NRI=R(1)
DO 5010 I=NS ,NRI
II=I
R(I,J)=Y(I)-PEFK(II,PAV,PAK)
NRA3=R(1)-6,0#R(2)
NRA=NS-1
IF(NRA-NRA4)30592,30592,30593
30592 DO 30591 I=NRA3,NRA4
II=I
30591 R(I,J)=R(I,J)-PEFK(II,PAV,PAK)
30593 IF(ENV-ENCO)100,101,100
101 NRA2=NRA
DO 104 I=NRA,NRA
II=I
104 R(I,J)=R(I,J)-PEFK(II,PAV,PAK)
NRA2=NRA2+1
GO TO 103
100 NRA2=NRA1+1
103 DO 5011 I=NRA2,NLIMIT
5011 R(I,J)=0
IF(NSX(J))5020,5001,5020
C SURTRACT X-RAY PEAKS IF ANY.
C
5020 IF(NOSUHT(J))5020,5001,5020
5020 DEL=NEFX(J)-NSX(J)
NSX=NSX(J)
NSX=NSX(J)+1
NEFY=NEFY(J)
NEFY=NEFY(J)+1
DO 5021 I=NSX,NEFY
Y1=NSX
5021 R(I,J)=R(NXX,J)+X*(R(NEXX,J)-R(NXX,J))/DEL
5001 CONTINUE
C ZERO FROM THE FIRST NEGATIVE CHANNEL TO 260
DO 9010 I=NRA3,NLIMIT
IF(R(I,J))9011,9012,9012
9011 INDIC=1
GO TO 9014
9012 CONTINUE
9010 CONTINUE
GO TO 9015
9014 DO 9013 I=INDIC,NLIMIT
9013 R(I,J)=0
9015 CONTINUE
5000 CONTINUE
NRA=NRI
WRITE(10,819)NRA,INDIC,NRI,NR2,NRA3,NRA4,NTR,NRA,ENCO
C 819 FORMAT(9H COB-1 ,F15,F14.7/)
END
SUBROUTINE COMPLEX (NST,NST02)
C ***** PROGRAM NUMBER - 5 *****
C
C CALLED BY *SHAPE*
C CALLS *SIVE* - *GANE* - *POLATE* - *CHALIT*
C
COMMON /LID/ LI,LI0,IP,NLIMIT
COMMON/PLX/R,STDENY,P,PAV,PARAS,PARAF,ALABEL,VX,Z,NSJ,
1 NENI,NSXJ,NEFY,NISUHT,NX,NSTAND,NTR
COMMON/POL/PI,F,CEA,MIN
COMMON/ODRM/ODRM,NORM
DIMENSION NSJ(12),NENJ(12),STDENY(12),R(5),NSXJ(12),
1 NFXJ(12),NOSUHT(12),Z(6),YS(6),R(260,12),ALABEL(12),
2 RI(50,3),RIABEL(12),PARAS(12),PARAF(12)
DIMENSION SMO(520),F(520),PV(520)
WRITE(10,120)STDENY(NSTAND),NSJ(NSTAND),NENJ(NSTAND),
1 NOSUHT(NSTAND),NST,NST02,NTR,NX,NSTAND
C 120 FORMAT (1H1,F9.4,G15)
IAZ=0,0
SM=0,0
NVS=0
NVS2=0
NSTA1=NSTAND
NSTA2=NSTAND
IF(NTR)10,10,12
12 NS=NSJ(NSTAND)
NEN=NENJ(NSTAND)
DO 26 I=1,NLIMIT
26 R(I)=R(I,NSTAND)
CALL FIVE (FV,NS,NEN,PV,PAV)

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R763 SMO(I)=R(I,NSTAND)
SMX=2.0
NXSM=100
CALL GANE(TAZ,NXSM,1.0,1.0,SMX,SMO)
DO R762 I=30,78
R762 R(I,NSTAND)=SMO(I)
C WRITE(LO,120)ANRM,GAIN,ROGAIN,RRI,RR3
C WRITE(LO,120)YMAX,IMAX,NR,NX,NCE,NL,N2,NXII,II,NDEL
C 201 WRITE(LO,130)ALREL(NSTAND),ALARFL(NSTAND),STDEWY(NSTAND),IP(TINST)
C IAND),I=1,110
C WRITE(LO,120)STDEWY(NSTAND),MSJ(NSTAND),MFNJ(NSTAND),NSX(NSTAND),
C INDSURT(NSTAND),NST,NSTD2,NIR,NX,NSTAND
C RETURN
END
SURROUNTIME DEC (NX,S,MX)
C***** PROGRAM NUMBER = 6 CIPED *****
C
C CALLED BY *MAIN*
C CALLS *GANE*
C READS PULSE-HEIGHT ANALYZER SPECTRA, CHECKS FOR PHA-COMPLEMENTED COUNTS.
C
C DIMENSION S(260),FM(520)
C COMMON /LIO/ LI,LO,LP,NLIMIT
C COMMON/CONST/T38,T01,T20,T50,T76,T1293,T90,T100,T366,T3316,T06,
C 1 T321,T7677,T285
C 1 IF(MX-NX) 1,1.2
C 2 READ(LO,3000)FM(I),I=1,MX
C MX=MX
C GO TO 7
C 1 READ (LI,3000)FM(I),I=1,NX
C MX=MX
C 3000 FORMAT (10(F6.0,1X))
C 7 DO 4 I=1,MX
C IF(FM(I)-T90) 5,5,6
C 6 FM(I)=FM(I)-T100
C 5 CONTINUE
C 4 CONTINUE
C 20 IF(MX-NLIMIT)11,11,21
C 21 ROGAIN=NX
C GAIN=MX
C MX=MX
C CALL GANE(LO,0,MXX,GAIN,ROGAIN,0.0,FM)
C 11 DO 3 I=1,NX
C 3 S(I)=FM(I)
C RETURN
END
FUNCTION DECAY ( OT,TH,T1, CLAMDA )
C***** PROGRAM NUMBER = 7 CIPED *****
C
C CALLED BY *MAIN*
C COMPUTES DECAY CORRECTION FACTOR.
C
C T1=T1*1440.0
C IF ( OT/TH ) - 0.00112,2.1
C T2=T1 + OT
C TE=(-1.0/CLAMDA) *ALOG ((EXP(-CLAMDA *T1)-EXP(-CLAMDA *T2))/
C 1 (CLAMDA *OT))
C GO TO 3
C 2 TE=T1 + OT/2.0

```


[illegible]

```

CORR=0.389423AR(3)
COPD=0.389423AR(NM2)
C COMPUTE STRAIGHT LINE BASE FOR PHOTOPeAK
DO 400 I=NS,NEN
  X=I-NS
  X=X-0.5
  P(I)=R(NM1)*X+R(NR)
  A(I,NM1)=X
  400 A(I,NR)=1.0
C COMPUTE GAUSSIAN FUNCTIONS AND PARTIAL DERIVATIVES
DO 700 I=NS,NEN
  X=I
  X=X-0.5
  PNM1=(X-R(1))/R(2)
  IF(ABS(PNM1)-7.0)99.999.999
  ARG1=EXP(-0.5*PNM1*PNM1)
  GO TO 299
  999 ARG1=0.0
  299 PP1=CORR*ARG1/R(2)
  A(I,1)=PP1*PNM1/R(2)
  A(I,2)=PP1*(PNM1*PNM1-1.0)/R(2)
  A(I,3)=PP1/R(3)
  IF(NR=5)4.5*4
  4 IF(R(4))500.500.510
  500 PP2=0.0
  GO TO 490
  510 IF(NR=8)512.514.700
  512 IF(NR=6)600.514.600
  600 RZ=R(2)
  GO TO 515
  514 RZ=R(5)
  515 PNM2=(X-R(4))/RZ
  IF(ABS(PNM2)-7.0)199.1999.1999
  ARG2=EXP(-0.5*PNM2*PNM2)
  GO TO 399
  1999 ARG2=0.0
  399 PP2=CORR*ARG2/RZ
  IF(NR=8)520.522.700
  520 IF(NR=6)620.522.620
  620 A(I,2)=A(I,2)+PP2*(PNM2*PNM2-1.0)/R(2)
  GO TO 550
  522 A(I,5)=PP2*(PNM2*PNM2-1.0)/R(5)
  550 A(I,4)=PP2*PNM2/RZ
  A(I,NM2)=PP2/R(NM2)
  GO TO 690
  690 PP2=0.0
  690 P(I)=P(I)+PP1+PP2
  FC(I)=Y(I)-P(I)
  700 FIT=FIT+FC(I)*FC(I)/Y(I)
  IF(MOD(I),15)=100.15
  100 FORMAT(15H FUMIS R-VALUES /1X,FF14.4)
  100 WRITE(LO,10) (A(I),I=1,8)
  WRITE(LO,18) (FC(I),I=NS,NEN)
  WRITE(LO,19) (Y(I),I=NS,NEN)
  WRITE(LO,14) (FIT,KLM,NM1,NM2,NR)
  14 FORMAT(13H *F11.4,414)
  18 FORMAT(16H FUMIS FC-VALUES /1X,10F10.2)
  19 FORMAT(16H FUMIS Y-VALUES /1X,10F10.2)
  15 RETURN
END

SUBROUTINE GAIN (IT,NY,GAIN,DOGAIN,SMOOTH,C)
C ***** PROGRAM NUMBER = 15 (FIXED) *****
C REF. JO. NIC. SCI. AND ENG. VOL 35, MAR. 1960 STEP 603 ANDREWS
C MARCH 1968 VERSION. 1.5. DIMENSION CAPACITY CHECK
C CALLED BY MAIN - *SINGLES - *DESCRIPTORS - *SHAPE
C GAIN CHARGING PROGRAM. ALSO SPECTRAL SHIFTING.
C
C DIMENSION C(520),FM(520)
C NLIMIT=520
C WDMX=NX
C NMNY=NMX
C IF(NMXY-NLIMIT)620.620.621
  621 NX=NX-(NLIMIT-NMXY)-1
  620 IF(ITZ)600.601.601
  600 ITZ=ABS(ITZ)
  NX=NX-ITZ
  IF(NX-NLIMIT)601.601.603
  603 NX=NLIMIT
  601 IF(ITZ)1000.275.1000
  1000 NZC=ITZ
  NXD=NX-NZC
C
C INTEGER SHIFT IF NZC# NOT EQUAL TO ZERO.
C
  IF(ITZ)913.910.911
  913 NZC=NZC-1
  NS=NZC*(-1) + 1
  NSX=NS-1
  NSXD=1
  NXD=NSX+NX
  945 DO 956 I=NS,NXD
    K=I+NZC
  956 FM(I)=C(K)
  957 C(I)=FM(I)
  GO TO 93
  911 NS=1
  NSXD=NXD+1
  NSX=NX
  945 DO 91 I=NS,NXD
    K=I+NZC
  91 C(I)=C(K)
  93 DO 92 I=NSXD,NSX
    92 C(I)=0.0
  910 NX=NXD
C
C DECIMAL SHIFT.
C
  274 TMZC=NZC
  916 I=I-TMZC
  916 I=1.0-DIF
  IF(NIT)271.272.271
  271 NXD=NX-2
  C(I)=C(I) + C(2)*DIF
  DO 270 I=2,NXD
  270 C(I)=C(I)*DIF + C(1+1)*DIF
  C(NXD+1)=C(NXD+1)*DIF + C(NX)
  GO TO 273

```

```

272 IF(NZC1273+275+275
273 NX=NX-1
275 IZ=0,
I=1,
G=0.5
DO 50 I=1,NLIMIT
50 FM(I)=0.0
FMULT=GAIN/ROGAIN
C DEL=0.5 WHEN GAIN/ROGAIN=1.72. I.F. DOUBLING REQUIRED.
C DEL=2.0 WHEN GAIN/ROGAIN=2.71. I.F. HALVING REQUIRED.
C
1 DEL=GAIN/ROGAIN
IF(DEL-2.0)402, 204, 402
402 IF(DEL-0.5)3, 3, 4
3 L=L+1
GAIN=GAIN*2.0
GO TO 1
C INITIALIZE FOR REDUCING ALGORITHM.
C
4 I=1
K=1
X=0.
NX=1.0
DEL=0.
IF(DEL-1.0)5,499,105
499 IF(L-1)497,500,497
497 DELT=ROGAIN/2.0
L=L-1
496 DO 498 I=1,NX
498 FM(I)=C(I)
GO TO 201
C INCREASE. .... LOAD EVEN CHANNELS WITH QUADRATICALLY
C INTERPOLATED COUNT. SHIFT ENTIRE SPECTRUM UPWARD ONE-HALF CHAN.
C
5 DELT=ROGAIN/2.0
DELT=GAIN/DELT
GO TO 40
303 NX=J-4
DO 305 J=1,NX,2
J=J+1
C(I1)= (3.0*C(I1) + 6.0*C(J+2) - C(J+4))*125
C(I1+2)= (3.0*C(I1+3) + 6.0*C(I1+1) - C(I1))*125
NX=I+3
DO 936 I=2,NX
936 FM(I)=(C(I)+C(I-1))*0.5
FM(I)=C(I)*0.5
DO 935 I=1,NX
935 C(I)=FM(I)
GO TO 304
C REDUCTION ALGORITHM.
C
105 XW=DEL
DEL=XN-X
XW=XW
106 FM(K)=FM(K)+(DEL*C(I)

```

```

IF(1-NX)112,202,202
112 X=X+DEL
IF(X-XK+1,0E-9)104,107,107
107 K=K+1
XN=I
DEL=XN-X
XK=XK+XKX
GO TO 106
108 XN=XK
DEL=XN-X
109 IF(DEL-1.0)111,111,110
110 DIF=DEL-1.0
DEL=DEL-DIF
GO TO 109
111 I=I+1
GO TO 106
202 DO 1949 INX=1,NLIMIT
1999 C(INX)=0.0
XNX=NX
XNX=XNX/DF
NX=XNX
C LOAD ODD CHANNELS IF *DELT* EQUAL TO ZERO.
C
201 IF(DEL) 300,301,300
300 NC=2
DO 302
GO TO 302
301 NC=1
DO 302
GO TO 302
302 DO 200 I=1,NX
J=NC*I - NO
C(J)=FM(I)
IF(DEL) 303,304,303
304 IF(L-1)404, 500,404
404 L=L-1
GO TO 496
C HALVING.
C
204 I=1
NCHCK=NX
NX=NX/2
NCK=NCHCK -(NX+NX)
IF(NCK)555, 504, 555
555 NX=NX+1
504 K=2*I-1
C(I)=C(K)+C(K+1)
IF(1-NX)502,503,503
502 I=I+1
GO TO 504
503 IF(1-1)525,525,205
C SMOOTHING IF *SMOOTH* NOT EQUAL TO ZERO.
C
525 IF(SMOOTH)526,500,526
526 DEL=DEL
SMOOTH=SMOOTH - 1.0
FM(I)=0.5
GO TO 496

```

```

205 L=L-1
GO TO 204

C COUNTS SCALED IN ACCORD WITH *FMULT* FOR INCREASED SPECTRA.
C
500 KK=0
915 CONTINUE
IF(KK)1499,1500,1499
1500 FAC=1.0
GO TO 1501
1499 FAC=1.0/2.0**KK
1501 IF(FMULT-FAC)916,917,917
916 KK=KK+1
GO TO 915
917 DO 920 I=1,NX
920 C(I)=C(I)*FAC
IF(SMOOTH)204,527,204
527 GAIN=C
RETURN
END

SUBROUTINE GAUSS (FM,V,S,PKAREA,SUM,NMAX)
C***** PROGRAM NUMBER - 16 CIPED *****
C
C CALLED BY *PEAKS* - *RFGCEN*
C CALLS *PEAK*
C COMPUTES GAUSSIAN PHOTOPeAK FOR GIVEN PARAMETERS.
C
DIMENSION FM(260)
COMMON /LIN/ L1,L0,LP,MLIMIT
COMMON /THIN/ THIN
SUMEO=0
NMIN=V-6.0*S
IF(NMIN)7,7,A
7 NMIN=1
A NMAX=V+6.0*S+1.0
IF(NMAX-MLIMIT-1)9,9,10
10 NMAX=MLIMIT-1
9 DO I=NMIN,NMAX
11 I=I
6PEFK(I),V,S,PKAREA)
SUMESUM=SUMESUM+FM(I)
1 PHIT(I)=FM(I)*G
20 IF(I)7,7,20
30 IF(I)7,7,20
5 FORMAT(10H GAUSS /1,215,5X,4F14.7/1X,5F14.7)
20 RETURN
END

SUBROUTINE GOMTR
C***** PROGRAM NUMBER - 17 CIPED *****
C
C CALLED BY *MATH*
C CALLS *DOSE*
C COMPUTES GEOMETRY FACTORS. INTEGRATES NUMBER AND ENERGY SPECTRA.
C CALCULATES NORMALIZED DOSE DATA FOR FINAL CODE RESULTS.
C
COMMON /C/ SUMMIN,SUMEN,DOSDET,AVENGY,PNUMRE,FMRENY,PNUMRE
1 DOSREX=DOSCYL*DOXVOL*DCYVOL
COMMON/SHA/SHMIN(1405),M,F,DH42(104)
COMMON/GEN/RETNIM,DE,PHI,MS,BETENY,NGD,DT,VOL,CYLINDIA,EMYTAL
COMMON/STN/STN(560),NX,NDIMA(3),TX,HC,PX,IN,PHIT(20),NGC,DHMS

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C FM(1) NOW ADDS UP TO UNIT AREA.
C
END
FUNCTION PERK(I,PAV,PA5,PAK)
C***** PROGRAM NUMBER - 22 CUPED *****
C
C CALLED BY *SHAPE* - *GAUSS* - *REFGEN* - *FINE* - *SINGLE* -
C *CONJ*
C COMPUTES GAUSSIAN FUNCTION AT X. (F)
C
X=I
X=X-0.5
POM=(X-PAV)/PA5
ARG=0.5*POM*POM
IF(ARG-20.0)2,2,3
2 PERK=0.3949423*PAK*EXP(-ARG)/PA5
RETURN
3 PERK=0.0
RETURN
END
FUNCTION PERSPX (E,NGX,N,TKLUC)
C***** PROGRAM NUMBER - 23 CUPED *****
C
C CALLED BY *XTAL*
C CALLS *TA* - *TF*
C COMPUTES PERSPEX ABSORPTION FACTOR. (F)
C
DIMENSION X(23),R(23),Z(6),Y(6)
GO TO (1,2),NGX
1 NGX=2
X(1)=0.01
X(2)=0.015
X(3)=0.020
DO 7 I=4,11
7 X(I)=X(I-1)+.010
DO 8 I=12,20
8 X(I)=X(I-1)+.100
X(21)=1.5
X(22)=2.0
X(23)=3.0
DATA R
1.1707,1.660,1.603,1.320,1.148,1.028,1.0869,1.0763,1.0720,
2.0687,0.559,0.478,0.384/
LOW=1
M=23
C
2 CALL TA (E,X,M,LOW,MOM,MIN,Z,Y,R,N,1.0)
C
NN=N+1
PATE (NN,Z,Y,F)
PERSPX=EXP(-P*TKLUC)
RETURN
END
SUBROUTINE PLOTATE(FOUT)
C***** PROGRAM NUMBER - 24 CUPED *****
C
C CALLED BY *SHAPE* - *CONJ*
C CALLS *VECTMX* - *GAMF* - *TF*
C
COMMON /10/ LT,LD,LP,NLIMIT
COMMON /PLX/ STORV(12),DIM(7),C,MINIM(6),MSTAND,MTA
COMMON /POLY/ RUI,F,CEA,IMIN
COMMON /CONSP/ CON,CONP
DIMENSION Z(240,12),O(6),FOUT(520),Y(6),FM(520),ERS(3),VRS(3),
1 FCF(3),VCF(3),NRS(3),VRS(3),PVI(240,3),VNRS(6),VNGF(6),FDM(520)
2,NMX(3),NRX(3),FAM(520),RUI(50,3),OUI(6)
MTREF=3
SM=0.0
TAZ=0.0
G1=1.0
DO 3 I=1,NLIMIT
FOUT(I)=0.0
FDM(I)=0.0
FAM(I)=0.0
3 FM(I)=0.0
DO 230 I=1,100
Y(1)=R(1,MIN)
Y(2)=R(1,MIN+1)
Y(3)=R(1,MIN+2)
FOUT(I)=TF(MTREF,O,Y,F)
WRITE (10,30) (FOUT(I),I=1,100)
1 FCF=FM(6)240,100,100
C
C DETERMINE COMPTON FINE AND AS PEAK LOCATIONS
C
100 J=3
JNE=J+JMIN-1
DO 2 I=1,100
2 FM(I)=R(1,JN)
VRS(J)=O(J)/(1.0+O(J)/O.511)*1.84
FCF(J)=ERS(J)*CONP/O(J)
VCF(J)=ECF(J)*CONP/O(J)
N1=VRS(J)-5.0
N2=N1+10
YMAX=0.0
CALL VECTMX (FM,N1,N2,IMAX,YMAX)
NRS(J)=JMAX
N1=VCF(J)-15.0
N2=N1+30
YMAX=0.0
CALL VECTMX (FM,N1,N2,IMAX,YMAX)
NCE(J)=JMAX
FNRS=NRS(3)
FNCE=ENRS/VRS(3)
ENGE=ENCF/VCF(3)
DO 70 J=1,2
ERS(J)=O(J)/(1.0+O(J)/O.511)*1.84
VRS(J)=ERS(J)*CONP/O(J)
FCF(J)=O(J)-ERS(J)
VCF(J)=ECF(J)*CONP/O(J)
VNRS(J)=VRS(J)*FNRS
VNGF(J)=VCF(J)*FNCE
VNGF(3)=ENCF(3)
VNRS(3)=ENRS(3)
NRS(1)=VRS(1)+0.5
NRS(2)=VNRS(2)+0.5

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X(1)=T3316
X(2)=0.0350
X(3)=0.040
DO 20 I=4,14
  R(1)=X(I)-0.010
  R(2)=25500
  R(3)=21001
  R(4)=14533
  R(5)=10157
  R(6)=07200
  R(7)=05301
  R(8)=03950
  R(9)=03001
  R(10)=02330
  R(11)=01830
  R(12)=01480
  R(13)=01205
  R(14)=00994
M=14
LOW=1
2 CALL TA (E,X,M,LOW,MOX,MIN,Z,Y,R,N,I,NZERO)
C WRITE(LO,100) M,LOW,MOX,MIN,L,N,NGR,Z(1),Z(2),Z(3),Y(1),Y(2),Y(3),
C 1 (1,X(1),R(1),I=1,14) ,715/1X,6E14,7/(1X,15,5X,2E14,7)
C 100 FORMAT (10H RAXEL
NN=N+1
RAXEL= TE (NN,7,Y,E)
10 RETURN
9 IF(E=0.5)11,11,12
11 RAXEL=(5.0233E-05)*E**(-2.7872)
GO TO 10
12 RAXEL=0.0
GO TO 10
END
SUBROUTINE RESGEN
***** PROGRAM NUMMER - 27 CUPEN *****
C
C CALLED BY *SHAPE*
C CALLS *STOFIT* - *PULSE#* - *RAXEL#* - *PEEK#* - *GAUSS#* - *GAMF#*
C *CORALT*
C ORDERS AND NORMALIZES STANDARD SPECTRA FOR RESPONSE MATRIX INTERPOLATION.
C
DIMENSION R(260,12),ALABEL(12),TAG(24), STDENY(12),NENH(12),
2 PAREA(12),NUSURT(12),RLABEL(12)
COMMON /LIO/ LI,LO,LP,NLIMIT
COMMON /PLX/R,STDENY,R,PARAV,PARAS,PAREA,ALABEL,RLABEL,NIMH(12),
1 NSJ,NENJ,NSXJ,NFXJ,NUSURT,NX,NSTAND,NIR
COMMON /GEN/ TAG,G,SCONST,SN,PM,IMGB,IN,NTAG,NA,N7,N5,NPHA
COMMON /CNSTNT/ T3A,T01,T20,T50,T76,T1293,T90,T1100,T366,T3316,T04,
1 T321,T7677,T285
COMMON /SHA/SDIM(1969),MEW
COMMON /DORRW/DRM,DORR
NAT=0
C INITIALIZE
NST=0
NSTD=0
NSTD2=0
NGAIN=0
NS2=NS#2
NA2=NA#2
NZ2=NZ#2
NENJ=NENJ#1
NENH=NENH#1
DO 2 I=1,NSTAND
  IF(9,ALABEL(1),NE,TAG(NS2))GO TO 4
  3 I5=1
  4 CONTINUE
  2 CONTINUE
  NAT=0
  EG=1.114
  IF(STDENY(NSTAND-1)-EG)50,51,50
  51 NAT=NAT+1
  50 EG=1.2R
  NSTNAT=NSTAND-NAT-1
  IF(STDENY( NSTNAT )-EG)52,53,52
  53 NAT=NAT+1
  52 NSTAND=NSTAND-NAT
  NSTD3=NSTAND
  GO TO 200
C SEARCH FOR NA22.
10 DO 20 I=1,NSTAND
  IF(9,ALABEL(1),NE,TAG(NA2))GO TO 40
  30 INA=1
  NST=1
  GO TO 201
  40 CONTINUE
  20 CONTINUE
C SEARCH FOR ZNA5.
201 DO 2000 I=1,NSTAND
  IF(9,ALABEL(1),NE,TAG(INZ))GO TO 400
  300 INZ=1
  NST=NST+2
  GO TO 1111
  400 CONTINUE
  2000 CONTINUE
C INITIAL LIMITS FOR START AND STOP FITTING.
1111 NSJ=1
  NENJ=NENJ(15)
  NSTD1=NSTAND+1
  NSTD2=NSTD1+1
  NSTD3=NSTD2+1
  NENJ=NENJ(15)
C LOAD 3 EMPTY VECTORS FOR TEMPORARY USE.
111 NSJ(NSTD1)=NSJ(15)
  NENJ(NSTD1)=NENJ(15)
  STDENY(NSTD1)=STDENY(15)
  DO 150 I=NSJ5,NENJ5
  150 R(1,NSTD1)=R(1,15)
  GO TO (159,170,1701,NST
  170 NSJ(NSTD3)=NSJ(INZ)
  NENJ(NSTD3)=NENJ(INZ)
  STDENY(NSTD3)=STDENY(INZ)
  NSJ7=NSJ(NSTD3)
  NENJ7=NENJ(NSTD3)
  DO 180 J= 1 ,NENJ7
  180 R(1,NSTD3)=R(1,INZ)

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7010 CONTINUE
WRITE(LO,220)(J,ALABEL(J),PARAV(J),PARAS(J),PARFA(J),
1 J=1,INSTAND)
8201 ENDOF(11),45X,2PHASES OF PHOTOPeAK FITTING. ///
8202 MAXIMUM X=12X,15STANDARD SOURCE=11X,12PULSE-HEIGHT=10X,
8203 1STANDARD DEVIATION=14X,44AREA/ 52X,INTEGRALS=14X,
8204 10H(CHANNELS),14X,15H(COUNTS/TIME)/(9X,15,18X,20X,5X,3F5.7))
RETURN
END
SUBROUTINE RESMAT
***** PROGRAM NUMBER - 28 (CIPED) *****
C
C CALLED BY *SOLN*
C CALLS *VECTMX*
C PULSE-HEIGHT ANALYZER SPECTRUM UNFOLDING ACCORDING TO THE SCOFIELD ALGORITHM.
C
COMMON/SHA/R,N,DIM1(365),K,DIM2(3)
COMMON/SOLRES/EPS,IT,ITMAX,DIF,MN,DIM3(2)
COMMON/SDN/P,DIM4(329)
COMMON/GE0/DIM5(2),PHI,DIM6(266)
COMMON/LIO/LI,LO,LP,NLIMIT
DIMENSION PP( 50),P(520),R( 40, 40),PHI(260),FIT(100),DIF(100)
1 ,MN(18)
C
C INITIALIZE
C
NCHECK=ITMAX/2
NCHECK=ITMAX-2*NCHECK
SU=LO
DO 1999 I=K,N
SU=SU + P(I)
1999 SU=SU + P(I)
DO 1800 I=K,N
1800 P(I)=P(I)/SU
YMAX=0.0
C
CALL VECTMX (P,K,N,JMAX,YMAX)
C
YLOW=YMAX*(1.0E-15)
DO 1 I=1,N
PP(I)=0.0
1 PHI(I)=P(I)
C
WRITE IF ITMAX IS AN ODD-NUMBER.
8999 WRITE (LO,8999) (P(I),I=1,N)
8999 FORMAT (LO,8999) / (1X,10F11.4))
INDEX=1
8000 CONTINUE
IT=1
FIT(1)=0.
C
C MATRIX MULTIPLICATION P=R*PHI
C
DO 2 I=K,N
DO 2 J=K,N
IF(R(I,J))2001,2000,2001
2001 PP(I)=PP(I) + R(I,J)*PHI(J)
2000 CONTINUE
2 CONTINUE
TERM = 0.0

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C
C ARRESTING CHECK WHEN DIF(SUM(P-PP)*2/P) FOR LOOP(IT) AND LOOP(IT-1)
C ARE LESS THAN EPS.... OR IT=ITMAX.
C
DO 5 I=K,N
IF(P(I))4,4,4
4 TERM = (P(I) - PP(I))*2/P(I) + TERM
5 CONTINUE
6 CONTINUE
IF(NCHECK)16000,15000,16000
16000 IF(MN(INDEX) - IT)15000,17000,15000
17000 WRITE(LO,19000) IT,MN(INDEX),(PHI(I),I=K,N)
19000 WRITE(LO,19000) IT,MN(INDEX),(PP(I),I=K,N)
INDEX=INDEX+1
15000 CONTINUE
19000 FORMAT (51H INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 215
19001/ (1X,10F11.4))
19005 FORMAT (51H INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PP(I)) 215
19005/ (1X,10F11.4))
IT=IT+1
FIT(IT)=TERM
DIF(IT)=ARS (FIT(IT-1) - FIT(IT) )
IF( DIF(IT) ) -EPS)7,7,29
29 IF(IT-ITMAX)229,7,7
C
C SCOFIELD CORRECTION FACTOR METHOD.
C
229 DO 25 I=K,N
IF(PP(I))125,52,125
125 PHI(I)=PHI(I)*P(I)/PP(I)
52 CONTINUE
25 CONTINUE
DO 1000 I=K,N
IF(ARS(PHI(I)) - YLOW )1001,1002,1002
1001 PHI(I)=0.0
1002 CONTINUE
1000 CONTINUE
GO TO 10
C
C RESULTS.
C
7 DO 600 I=K,N
P(I)=PP(I)*SU
600 PHI(I)=PHI(I)*SU
C
WRITE IF ITMAX IS AN ODD-NUMBER.
9999 WRITE (LO,9999) (PP(I),I=1,N)
9999 FORMAT (LO,9999) / (1X,10F11.4))
10000 CONTINUE
60 FORMAT (20H0 IT,SU,TERM
/110,2F14.7)
WRITE (LO,60) IT,SU,TERM
RETURN
END
SUBROUTINE SHAPE
***** PROGRAM NUMBER - 29 (CIPED) *****
C
C CALLS *SPEAK* - SPEAK* = SPECFGEN* - SPEULSE* - STAS - STE*
C
C *SPEULSE* = *COMPLEX* - *VECTMX* - *PHI*IN*
C
C CONTROL PROGRAM FOR RESPONSE MATRIX GENERATION.
C
DIMENSION PM(40,40),P(260,12),FM(520),X(12),PV(50),PEFACT(50),

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      DO 3456 J=1,NTAG
      JK=J*2
      IF(RLAREL(KK),FO,TAG(JK))GO TO 4321
      6543 CONTINUE
      3456 CONTINUE
      WRITE (LO,9) (ALAREL(JJ),RLAREL(JJ),JJ=1,NSTAND), (TAG(JJ),
      1 JJ=1,24)
      9 FORMAT (1H1,24H ERROR FLAG FOR PHOTRA /11X,1H*,2A33)
      CALL EXIT
      4321 STDENY(KK)=STDEN(J)
      1234 CONTINUE
      IF(MEM)2720,2720,2721
      2721 NGC=1
      DO 2722 J=1,NSTAND
      NXL=NX LIM
      JJ=J
      N1=NSJ(J)
      N2=NFNJ(J)
      DO 2723 I=1,NX LIM
      2723 FM(I)=R(I,J)
      YW=0.0
      CALL VECTMX(FM,N1,N2,JM,YM)
      V6(J)=JM
      FLIM=STDENY(J)*1.5
      CALL ENLIN(NXL,FM,STDENY,V6,NGC,JJ,NGC,DM,FLIM)
      DO 2724 I=1,NX LIM
      2724 R(I,J)=FM(I)
      2722 CONTINUE
      WRITE(LO,13456)JK,NSTAND,NTAG,NTZ,STDENY(1),STDENY(2),STDENY(R)
      C
      C CALL RESGEN TO CALC GAUSSIAN PARAMETERS AND UNIT CONTINUA FOR STANDS.
      C
      2720 CONTINUE
      WRITE(LO,7) NSTAND,NPHA,UMGAIN, (ALAREL(J),
      1 BLAREL(J),NSJ(J),NFNJ(J),NSX(J),NEX(J),SHIFT(J),STDENY(J),
      2 J=1,NSTAND)
      7 FORMAT(11H1,15X,37H STANDARD SOURCE SPECTRAL PARAMETERS ///
      71 14X,15,33H SPECTRA IN STANDARD SOURCE DECK //
      72 13X,16H CHANNELS ONE TO 14,22H ASSUMED AS REMINANT //
      73 17X,25H REFERENCE COARSE GAIN = *F10.5///
      741X,72HSTANDARD PHOTOPeAK X-RAY OR .5 PEAK SHIFT
      75 PHOTOPeAK /71H SOURCE FROM FROM
      76 SPECTRUM ENERGY //10X,60H CHANNEL CHANNEL
      77 CHANNEL CHANNELS MEV //12X,2A3,6X,14,5X,14,7X,13,6X,13,
      784X,E9,4,1X,F10.5)
      DO 27123 J=1,NSTAND
      27123 J=1,NSTAND
      WRITE (LO,8999) ALAREL(J),RLAREL(J),R(I,J),I=1,NX LIM
      8999 FORMAT (1H1,32X,25H STANDARD SOURCE SPECTRA //11X,10F9.0)
      8991 RH SOURCE
      27123 CONTINUE
      WRITE (LO,3)
      3 FORMAT (11H1)
      NSTOOD=NSTAND
      IF(NPHA)7000,7000,9005
      9005 DO 7002 J=1,NSTAND
      DO 7002 I=1,NPHA
      7002 P(I,J)=R(NPHA+1,J)
      7000 NCOR=0
      IF(RLAREL(NSTAND),FO,TAG(24))GO TO 4322

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      R21 KJ=1
      IF(E=-5)R23,R24,R24
      R23 IF(E-T3316) R24,R25,R25
      R25 EK=E-T285
      SCALE=1.0
      IF(EK-.01)R24,5.5
C
C SET ENERGY AND PULSE-HEIGHT SCALES
R22 Q(KJ)=E
      V=E*EM
      IF(E=0-.010)60,60,59
      60 PV(KJ)=V
      K=K+1
      NMAX=N
      DO 30R80 I=1,NLIMIT
      30R80 FM(I)=0.0
      P=0.0
      GO TO 61
      59 EK=E-T285
      SCALE=GAIN/V
      SIG=(SCNST*(E*1000.0)**SN)*SCALE*EM/2354.R2
      R NMA=GAIN+1.0+6.0*SIG
      C WRITE(LO,952) NMA,NO,SCALE,GAIN,V,SIG
      C 952 FORMAT(10H SHAPE-1A ,215.4E14.7)
      124 GN=GAIN
      GAIN=GAIN/2.0
      SIG=SIG/2.0
      SCALE=GAIN/V
C REDUCE VECTOR LENGTH TO ALLOW COMPLETE PEAK TO BE FORMED (1.F,LOW E)
      CALL GANE (TAZ,NMAX,GN,GAIN,SM,FM)
      GO TO 8
      123 V=E*EM
      IF(EM)1521,1521,1522
      1521 V=V+V*PULSE (F,NGN,NDEGRE)
      1522 PV(KJ)=V
      4 VV=0.0
      OR=0.0
      SI=0.0
      GO TO 14
      5 VK=EK*GAIN/F
      IF(EM)1523,1523,1524
      1524 VV=VK
      GO TO 1525
      1523 VV=VK+VK*PULSE (EK,NGN,NDEGRE)
      1525 SI=(SCNST*(EK*1000.0)**SN)*SCALE*EM/2354.R2
      OR=RXEL(F,NGR,NDEGRE)
      IF(ESING)R26,14,R26
      R26 NREGIN=VV-6.0*SI
      IF(NREGIN)R28,R28,R29
      R28 NREGIN=1
      R29 DO R30 I=NREGIN,100
      II=I
      R30 FM(II)=FM(II)+PEEK(II,VV,SI,OR)
      C WRITE(LO,12321)NMA,NO,NREGIN,NMAX,F,SN,SCNST,SCALE,GAIN,GN,VV,OR,
      C 151,FSING
      GO TO R24
      14 CONTINUE
      CALL PEAKS (FM,GAIN,F,OR,VV,SI,SIG,NO,P,NMAX)

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      C
      NMAX=N/7
      CALL GANE (TAZ,NMAX,GAIN,V,SM,FM)
      IF(MOUT-NE.0) GO TO 61
      C WRITE(LO,12321) NMA,NO,K,NMAX,F,V,SIG,SCALE,GAIN,GN,VV,OR,SI,FSING
      C 12321 FORMAT (10H PHOTON ? ,415/(1X,5F14.7))
      SUM=0.0
      DO 700 I=1,NMAX
      700 SUM=SUM+FM(I)
      WRITE(LO,500) SUM,NMAX
      500 FORMAT(10H SUM,NMAX F15.8,I10)
      WRITE(LO,R6) KJ,P,F,FM(I),I=1,40)
      8A FORMAT(//////////)
      8A1 36H INTERPOLATED RESPONSE MATRIX VECIOR,I3,4H OF ,I3,
      86214H FOR ENERGY = ,F10.5,4H MEV /((1X,5F14.7))
      61 H=NMAX+2
      IF(M-40)90,90,91
      91 M=40
      90 PERACT(KJ)=P
      DO 7 I=NS,M
      2 FM(I,KJ)=FM(I)
      IF(KJ=MM)918,919,919
      918 KJ=KJ+1
      GO TO 920
      919 CONTINUE
      R24 RETURN
      END
      FUNCTION SIMPSM (A,R,EPS)
      C***** PROGRAM NUMBER - 30 CIPED *****
      C
      C CALLED BY #FFFC#
      C CALLS #FC#
      C SIMPSM'S RULE INTEGRATING PROGRAM FOR FUNCTION -FC-.
      C
      COMMON /R/ KL,IFGAM,FX,DIST,H
      IMAX=2048.
      AM=2.0
      FA=FC(A)
      FB=FC(B)
      FC=FC(C)
      SI=0.
      DELT=B-A
      FDIR=0.0
      405 AK=1.0
      AKN=AM/2.0
      399 FFY=FC(A+((12.0*AK-1.0)*DELT)/AM))
      FDIR=FDIR+FFY
      IF (AKN-AK)400,401,400
      400 GO TO 399
      401 SIM=(DELT/(AM*2.0))*((FA+FB+4.0*FFYR+2.0*FFD)
      IF (IMAX-AM)503,403,409
      409 DIFX=ARS(SIM-SI)
      IF (EPS-DIFX/SIM)402,403,403
      402 AN=AM*AN
      FDIR=FDIR+DIFX
      SI=SIM
      GO TO 405
      503 CONTINUE

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5005 CONTINUE
R3=R(3)
CALL GAUSS(FM,R(1),R(2),R3,SAM,NMAX)
R0014 DO 46 I=NS,NMAX
IF(FM(I))147,48,48
47 FM(I)=0.0
48 CONTINUE
49 CONTINUE
C
C DETERMINE THE CONTINUUM ASSOCIATED WITH A PEAK OF AREA = UNITY FOR
C ENERGY = ENY, AND OF PULSE-HT. = 100 CHAN. GAIN CHANGE PER R(1)
C AND MULTIPLY BY R(3) AND FINALLY SUBTRACT IT FROM FM(I). ALSO DET-EM.
C THE PHOTOINTERACTION TO CALC PHOTON FLUX.
C
R0020 DO 99 I=1,NLIMIT
99 PP(I)=0.0
MRPE=MRPEAT
MENY=REM
NNPSN
NENP
C
C CALL SHAPE
C
NENP
MRPEAT=MRPE
MENY=REM
DO 850 I=80,NLIMIT
IF(PP(I))851,851,852
851 INDEX=I
GO TO 853
852 CONTINUE
850 CONTINUE
853 DO 854 I=INDEX,NLIMIT
854 PP(I)=0.0
SUM=0.0
DO 999 I=1,100
SUM=SUM+PP(I)
PHOT(I)=1.0/(1.0+SUM)
N=100
GAINR(I)
PHOTPHIT(I)=PHOT(I)*PP(I)*SUM*PHOT(I).*(PP(I)-1.0/NLIMIT)
C9777 FORMAT (10H SINGLE 3 /IX,110,4F14.7/(1X,5F14.7))
C
CALL GANE(0.0,NX,NDRM, GAINR1,0.0,PP)
C
IF(NY-NLIMIT)1800,1801,1801
1801 NY=NLIMIT-1
1800 DO 1100 I=1,NX
FM(I)=FM(I)-PP(I)*R(3)
IF(FM(I))1873,1874,1874
1873 FM(I)=0.0
1874 CONTINUE
1100 CONTINUE
C 800 WRITE(10,800)NY,GAINR1,(FM(I),I=1,NLIMIT).*(PP(I)-1.0/NLIMIT)
C 800 FORMAT (10H SINGLE 3A /110,4F14.7/(1X,5F14.7))
C
FM(I) NOW WITHOUT PEAK J AND ASSOC CONTINUUM. PHOT ABOVE IS PHOTPHIT
C NOW IS LINE SPECTRUM AREA AT CHANNEL CORRESPONDING TO ENY (NO FILL)
C AND MENY (NO FILL).

```

```

C
PHOT(I)=R(3)/PHOT(I)
C
SLINE(I)=ENY
C
CALL XTAL(1,1,SLINE,44)
C
PHOT(I)=PHOT(I)/ETA(I)
WRITE(10,555)PHOT(I),PHO
555 FORMAT(10H//4X,25PHOTON NUMBER
555 189. /1X,1000 /25PHOTOINTERACTION
555 189. /1X,1000 /25PHOTOINTERACTION
1500 FOR (140/43X 31401SCRETE PHOTODEAK FITTING DATA ///
1500 1X,4MLIMIT,15X,4ENERGY,14X,12PHOTON SE-HEIGHT,10X, 18USTANDARD
15002 DEVIATIONS,1X,4AREA, 19H CHANNEL TO CHANNEL,10X,5TIMEV,18X,
15003 104CHANNELS,15X,104CHANNELS,14X,134CHANNELS/TIMEV//)
1500 WRITE(10,1400)NS,NENP,ENY,R(1),R(2),R(3)
1400 FORMAT (1H,2X,13,4X,13,F22.7,3F25.7)
100 IF(J-1)180200,8030,80200
80200 J=J+1
8040 IF(NR)8040,8015,8040
8040 NR=0
DO 18041 I=1,NLIMIT
8041 SLINE(I)=R3*NFN
DO 18042 I=1,NLIMIT
8042 SLINE(I)=FM(I)
MENY=FM(I)
R(1)=R(1)+R(3)
R(2)=R(2)+R(3)
R(3)=R(3)+R(3)
IF(FR(J))532,532,534
534 ENY=EXR(I)
GO TO 533
532 ENY=RR(4)*ELIMIT/66
533 GO TO 8020
8030 NCF=1
NY=NX
ETA(I)=ETA
RETURN
END
SUBROUTINE SPIN
***** PROGRAM NUMBER = 32 CLIPED *****
C
CALLED BY SPIN
C
CALLS XTAL = SPIN
C
DETERMINES AND APPLIES DETECTOR EFFICIENCY VECTOR.
C
COMMON /A/ DI,H,WF
COMMON/SHA/R,N,NIM,ELIMIT,NIM1(2),NDEGRF,F,NIM2(100),K,NIM3(3)
COMMON/SPH/P,NIM4(4),MCA,NX,NCG,TCLUC,IX,NIM5(21),MGO,FTA,NIST
COMMON/GEN/DIMA(2),PHI,NS,DIM7(265)
COMMON/STUBS/EPS,IT,ITMAX,DIF,N,NSKIP,M14
COMMON/10/1,10,10,NLIMIT
DIMENSION P(520),P(40,40),PHI(260),ETA(260),DIF(100),M(18)
IF(K)50,50,50
30 K=1
50 KK=K-1
IF(KK-1)601,602,602
602 DO 600 I=1,KK

```

```

600 P(1)=0.0
601 CONTINUE
IF(MF)12,11,12
C
C 11 CALL RESMAT
C
12 IF(MDIST-0)13,40,3
40 IF(MSKIP)3,4,3
3 NG0=1
IF(MI4)12,13,2
C
C 13 CALL XTAL(NS,N,F,H)
C
2 NG0=1
DI=DIST
IF(MF)19,4,9
4 IF(MI4)54,44,54
44 DO 22 I=NS,N
22 PHI(I)=PHI(I)/ETA(I)
54 MF=0
C WRITE DIF(I) IF (IT=ITMAX)
100 WRITE(10,105) (DIF(I),I=1,IT)
105 FORMAT (1H1,36H FLUX FITTING DIFFERENCES (IT=ITMAX)/(1X,5F14.7))
89 RETURN
9 IF(MI4)98,99,98
98 DO 96 I=NS,N
96 ETA(I)=1.0
99 DO 9999 I=NS,N
9999 PHI(I)=P(I)/ETA(I)
RETURN
END
SUBROUTINE STDFIT (Y,MEN,R
*NS,FNV)
*****
C ***** PROGRAM NUMBER - 33 CIPED *****
C
C CALLED BY *CORALT* - *SINGLE* - *FIVE*
C CALLS *GUESS* - *FINISH* - *FIVE*
C NON-LINEAR REGRESSION ANALYSIS OF STANDARD SPECTRA SINGLE
C PHOTONPEAKS
C
DIMENSION FC(260),Y(260),R(5),A(260,5),G(6,6)
COMMON /L10/ L1,L0,LP,NLIMIT
COMMON/TMOUT/MOUT
EPS=0.00001
L=8(3)
DO 100 I=1,5
100 R(I)=0.0
R(3)=L
NRC=5
NF=NFEN
NI=10
L=0
FITLM=0.0
LL=0
C
C CALL GUESS TO INITIALIZE PARAMETERS.
C CALL FUNCTION TO BE FITTED.
CALL GUESS (NS,MEN,Y,R,FNV)
FIT=0.0

```

[illegible]

APPENDIX III
SAMPLE INPUT CARD DECK LISTING

1.0
TEST RUN PH02 400 CH. TO 256. 6 PFAKS

0	2	3	0	0	6	7	0	0	0	0	0	0	0	15	16	0	0	
2.0	7	1.0	1.0	0.0	51.0													
7	7	200	1.0															
HG203	100	150	15	47	0.0													
SR25	110	145	0	0	0.0													
CS137	110	147	0	0	0.0													
NR95	110	144	0	0	0.0													
MR54	110	144	0	0	0.0													
7MA5	115	143	-54	67	0.0													
C060	122	142	0	0	-1.2													
000000	000000	000670	000902	000760	000860	000860	000917	000966	000953									
001094	001047	001061	001110	001051	001223	001357	001532	001751	001874									
001734	001544	001324	001333	001421	001568	001576	001784	002047	002622									
003782	005941	008361	010705	011311	010229	007962	005928	004732	004017									
003306	002680	001962	001423	001101	000974	000991	000996	001005	001080									
001023	001074	001134	001123	001195	001317	001330	001399	001562	001654									
001829	001911	002048	001983	001963	001887	001860	001684	001578	001509									
001425	001269	001271	001175	001127	001095	001037	000993	000989	000948									
000429	000929	000924	000921	000953	000918	000950	001002	001035	001047									
001066	001071	001063	001107	001081	001168	001088	001182	001225	001201									
001237	001251	001260	001271	001262	001259	001324	001409	001415	001495									
001543	001513	001489	001833	001937	002295	002796	003525	004557	006194									
000818	010850	013586	017036	020230	023422	025834	028083	029117	029829									
028868	027857	025426	022621	019213	016047	012746	009898	007310	005393									
003635	002413	001524	000889	000551	000310	000175	000122	000082	000078									
000028	000044	000016	000006	000058	000027	000031	000033	000030	000032									
000035	000049	000041	000041	000044	000034	000031	000041	000046	000026									
000056	000031	000046	000031	000058	000025	000032	000034	000048	000045									
000041	000037	000037	000026	000031	000037	000047	000040	000043	000030									
000035	000032	000042	000027	000025	000037	000025	000025	000037	000000									
000000	000000	005224	006569	001895	001767	001724	001823	001898	001878									
001883	001920	001848	001959	001920	002024	002019	002127	002176	002129									
002062	002040	002057	002062	002055	002000	002098	002068	002098	002091									
002176	002119	002241	002225	002223	002306	002286	002353	002358	002509									
002686	002920	003039	003038	003208	003105	003044	002938	002938	002865									
002748	002842	002716	002762	002680	002655	002630	002616	002587	002642									
002568	002560	002540	002515	002439	002544	002676	002633	002775	002721									
002757	002896	002758	002823	002858	002774	002925	002688	002796	002611									
002616	002453	002258	002098	001959	001795	001564	001416	001343	001252									
001234	001206	001126	001096	001039	001080	001042	001041	000974	000941									
000927	000946	000887	000947	000917	000868	000859	000925	000971	000974									
000971	001016	001130	001235	001339	001477	001726	002310	002873	003963									
000546	007331	009907	012520	015779	018960	021819	024258	025505	026004									
024786	022676	019649	016234	012537	008971	006186	004058	002538	001402									
000813	000388	000177	000084	000025	000030	000003	000020	999985	000025									
000014	000001	000007	000013	000006	000011	000017	000015	000025	000004									
000010	000002	000030	000025	000001	000015	000004	000014	000018	000009									
999993	000008	000022	000020	000013	000013	000009	000011	000003	000013									
000023	000015	000021	000010	000028	000016	999999	000023	000013	000024									
000005	000017	000009	000018	000011	000005	000011	000009	000024	000000									
000000	000000	000391	002460	005458	018591	011156	003228	002239	002131									
002095	002090	002222	002348	002280	002286	002169	002125	002195	002189									
002165	002247	002251	002348	002320	002220	002275	002399	002362	002416									
002510	002501	002643	002764	003020	003204	003216	003312	003354	003103									
003190	002973	002687	002977	002892	002955	002881	002800	002866	002771									
002760	002691	002649	002735	002694	002645	002697	002740	002711	002620									
002745	002652	002655	002717	002757	002650	002764	002685	002648	002729									
002703	002839	002698	002754	002706	002745	002654	002885	002789	002761									
002807	002751	002857	002843	002865	002777	002744	002590	002476	002344									
002098	001902	001633	001487	001341	001231	001118	001086	000988	000906									
000923	000807	000809	000772	000746	000741	000752	000740	000730	000670									
000775	000761	000835	000873	001004	001102	001180	001388	001705	002272									
002062	004099	005760	007998	010583	013552	016386	019038	020885	022004									
021289	018441	016849	013395	010019	007096	004634	002866	001504	000843									
000615	000174	000102	000039	000040	000021	000015	000019	000007	000015									
000030	000050	000040	000066	000053	000053	000045	000044	000035	000032									
000043	000017	000021	000017	000008	000035	000013	000008	000021	000014									
000022	000012	000005	000007	000000	000017	000019	999993	000004	000011									
000025	000023	000003	000020	999993	000008	000022	000011	000007	000007									
000015	000014	000020	000018	000006	000004	000012	000011	000013	000000									
000000	000000	002668	003593	004041	003031	002654	002603	002671	002640									
002628	002865	002977	002937	002779	002771	002785	002791	002892	002917									
002862	002963	002942	002962	003174	003165	003140	003338	003366	003438									
003419	004031	004182	004364	004312	004197	004080	003867	003927	003858									
003830	003802	003733	003688	003796	003759	003705	003730	003813										

000000 000000 002100 002363 001965 002048 001976 002036 002016 002083 80
002144 002218 002144 002132 002117 002067 002146 002203 002213 002165 81
002148 002255 002308 002247 002291 002381 002408 002590 002727 003057 82
003153 003230 003123 003163 003022 003054 002913 002931 002794 002878 83
002709 002743 002679 002662 002569 002618 002586 002718 002588 002568 84
002661 002587 002554 002487 002599 002576 002527 002574 002633 002585 85
002639 002531 002497 002583 002488 002545 002563 002480 002628 86
002638 002523 002497 002573 002552 002493 002445 002605 002668 002589 87
002648 002588 002629 002735 002762 002800 002762 002787 002850 002870 88
002734 002678 002508 002535 002400 002257 001918 001742 001468 001229 89
001109 001061 000892 000863 000789 000717 000704 000632 000578 000599 90
000606 000596 000593 000676 000740 000784 000870 001018 001218 001528 91
002842 002842 004116 005784 008124 010700 013782 016575 018092 018626 92
018116 015873 013377 009948 006871 004508 002510 001388 000731 000327 93
000138 000054 000012 000017 000016 999999 000016 000003 999997 000006 94
000002 999993 999991 000011 000003 000006 000003 999997 000010 000010 95
000016 999991 000018 000009 000003 000005 000017 000018 000016 000006 96
000003 000001 000016 999999 000005 000034 000001 000008 000012 000000 97
000010 000006 999997 999996 000010 000001 000007 999994 000011 000004 98
000015 000010 999999 000007 000001 999991 000009 000020 000012 000000 99
000000 000000 002845 003407 002638 002760 002868 002787 002915 002907 80
002768 002879 002787 002863 003015 003074 003178 003027 003365 003365 81
003502 003544 003710 004056 004128 004307 004297 003968 003932 003837 82
003816 003670 003737 003552 003611 003608 003675 003606 003364 003322 83
003397 003265 003196 003085 003222 003109 003014 003160 002995 002992 84
003003 003018 002997 003078 003115 003340 003365 004377 005240 006726 85
005888 005526 004689 003404 003085 002916 002856 002709 002795 002801 86
002912 002737 002637 002780 002851 002909 002821 002806 002862 002844 87
002084 002778 002963 002937 003039 002962 003066 003163 003070 003160 88
003074 003329 003228 003319 003341 003498 003507 003427 003383 003438 89
003133 002903 002565 002260 001985 001703 001326 001240 001073 001007 90
000906 000830 000802 000769 000718 000849 000852 000845 000945 001219 91
001512 001984 002831 004206 006172 009050 012199 015770 018207 019904 92
018275 016786 012846 009182 005920 003343 001815 000786 000404 000144 93
000073 000007 000003 999992 999994 000023 000003 999982 000019 000025 94
000019 000022 999986 000005 000012 000015 000022 000012 000007 999995 95
000025 999996 000030 000015 000020 000052 999995 999971 999996 999970 96
000062 999974 999983 999995 999986 999999 000017 000009 000017 000000 97
999991 000010 000003 000006 999990 000006 000002 000010 000008 999997 98
000007 000010 999997 000009 999995 000007 000009 999997 000008 000000 99
000000 000008 000611 007029 007604 007866 008379 008611 008250 007996
008025 008414 008303 008515 008611 008729 009112 009505 010104 011167
012003 012483 011557 011543 010804 010737 010571 010423 010384 010211
009843 009641 009548 009353 009317 009121 009074 009125 008960 008970
008775 008908 008836 008753 008651 008562 008457 008458 008466 008488
008459 008482 008259 008219 008205 008171 008145 008249 008167 008136
007945 007977 008105 008164 008206 008265 008314 008312 008393 008356
008408 008236 008354 008500 008724 008722 008888 008726 008987 009183
009469 009606 009579 009625 009737 009762 009948 009782 009715 009312
008941 008555 008047 007599 007090 006856 006624 006585 006423 006415
006470 006320 006399 006341 006186 006054 006057 006670 008057 010847
015340 020879 027077 031373 032347 029421 023838 016855 010792 006544
004215 003636 003986 005727 008628 012771 017774 022343 025820 026541
024011 019373 013710 008721 004843 002534 001041 000432 000264 000164
000116 000092 000109 000107 000084 000086 000105 000098 000091 000102
000095 000102 000104 000104 000103 000102 000099 000095 000084 000100
000076 000075 000092 000081 000076 000088 000069 000067 000057 000076
000068 000066 000074 000075 000066 000059 000069 000073 000062 000080
000061 000059 000052 000069 000062 000063 000071 000055 000057 000063
000067 000034 000049 000048 000059 000048 000058 000068 000042 000000
p002 1.0 1.0 1.0 1.0 20.0 1.0 200.0 0.0
7 0 20 0 0 1 1.0 1.0 127.0 1.27 0. 0. 0.
100. 0. 0. 0. 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0
1 2 3 4 5 7 10 15 20 30
2 115 0 1.2700 0 10 144 1.36 2 46 0 0.50
0 7 66 0.5750 2 69 0 0.7600 0 12 96 0.8750
1 143 167 1.5400

100513 124001 81387 43872 33958 31632 26716 22787 20076 18482
16846 16063 14977 14320 13427 12768 12231 11751 11418 11230
10681 9918 9722 9510 9269 8743 8626 8950 9931 11764
13968 15364 14202 11787 9944 10473 13311 17066 19334 18016
14363 10043 7109 5760 5198 5078 4830 4574 4404 4342
4327 4723 4915 5502 6167 6422 6663 6361 5726 4988
4457 4222 4061 4420 4869 5758 6432 7025 7106 6655
5809 5046 4196 3705 3371 3347 3318 3407 3304 3336
3182 3124 2915 2730 2549 2501 2162 2095 1958 1937
1744 1735 1676 1576 1508 1589 1638 1808 2116 2575
3324 4313 5555 6977 8397 8902 9095 8633 7318 5924
4451 3475 2854 2476 2244 2082 1960 1634 1485 1300
1144 1050 1001 906 930 915 915 886 967 1051
1124 1160 1168 1020 965 886 797 660 597 530
433 425 375 345 333 334 333 306 315 327
364 308 334 325 322 325 345 352 364 288
293 277 282 253 221 221 203 206 190 207
168 208 179 166 224 198 217 235 290 311
293 277 282 253 221 221 203 206 199 207
164 208 179 166 224 198 217 235 290 311

10
10

4281 CAPS

APPENDIX IV
SAMPLE OUTPUT LISTING

~~BRIEF DESCRIPTION OF PHA RUNS~~

TEST RUN PU02 400 CH. TO 256. 6 PEAKS

CONTROL NUMBERS

M(1) = 0 M(2) = 2 M(3) = 3 M(4) = 0 M(5) = 0 M(6) = 0
M(7) = 7 M(8) = 0 M(9) = 0 M(10) = 0 M(11) = 0 M(12) = 0
M(13) = 0 M(14) = 0 M(15) = 15 M(16) = 16 M(17) = 0 M(18) = 0
M(19) = 0 M(20) = 0 M(21) = 0 M(22) = 0 M(23) = 0 M(24) = 0

EM = 10.00000 CHANNELS/MEV

ELIMIT= 2.00000

ITERATIVE ERROR TOLERANCE, EPS = 0.00010

NUMBER OF BETA SOURCE SETS, OJSD = 1 MM = 1

MAX NUMBER OF ITERATIONS, ITMAX= 51

NUMBER OF CHANNELS INPUT, N = 20

NAI(TL) CRYSTAL SIZE = 3.00 X 3.00 INCHES.

(A)

STANDARD SOURCE SPECTRAL PARAMETERS

7 SPECTRA IN STANDARD SOURCE DECK

CHANNELS ONE TO 8 ASSUMED AS REDUNDANT

REFERENCE COARSE GAIN = 1.00000

STANDARD SOURCE	PHOTOPEAK		X-RAY OR .5 PEAK		SHIFT SPECTRUM	PHOTOPEAK ENERGY
	FROM	TO	FROM	TO		
	CHANNEL	CHANNEL	CHANNEL	CHANNEL	CHANNELS	MEV
HG203	110	150	15	47	0.0	0.27900
SR85	110	145	0	0	0.0	0.51500
CS137	110	147	0	0	0.0	0.66162
NR95	110	144	0	0	0.0	0.76400
MN54	110	144	0	0	0.0	0.83500
ZN65	115	143	54	67	0.0	1.11400
CO60	123	143	0	0	-1.2000	1.33200

(B)

STANDARD SOURCE SPECTRA

H6203 SOURCE

0.	0.	670.	902.	760.	860.	860.	917.	966.	983.
1094.	1047.	1061.	1110.	1051.	1223.	1357.	1532.	1751.	1874.
1734.	1544.	1324.	1333.	1421.	1568.	1576.	1784.	2047.	2622.
3782.	5941.	8361.	10705.	11311.	10229.	7962.	5928.	4732.	4017.
3306.	2680.	1962.	1423.	1141.	974.	891.	686.	1205.	1380.
1023.	1074.	1134.	1123.	1195.	1317.	1330.	1399.	1562.	1654.
1829.	1911.	2048.	1983.	1983.	1887.	1860.	1684.	1578.	1500.
1425.	1269.	1271.	1175.	1127.	1095.	1037.	993.	989.	948.
929.	929.	924.	921.	953.	918.	950.	1002.	1035.	1047.
1066.	1071.	1063.	1107.	1081.	1168.	1088.	1182.	1225.	1201.
1237.	1251.	1260.	1271.	1262.	1250.	1324.	1400.	1415.	1405.
1543.	1513.	1689.	1833.	1937.	2295.	2796.	3525.	4557.	6194.
8118.	10850.	13586.	17036.	20230.	23422.	25834.	28083.	29117.	29829.
28968.	27557.	25426.	22621.	19213.	16047.	12746.	9898.	7310.	5393.
3635.	2413.	1624.	889.	561.	310.	175.	122.	82.	78.
28.	44.	16.	56.	58.	27.	31.	33.	30.	32.
35.	49.	41.	41.	44.	34.	31.	41.	46.	26.
56.	31.	46.	31.	59.	25.	32.	34.	48.	45.
41.	37.	37.	26.	31.	37.	47.	40.	43.	30.
35.	32.	42.	27.	25.	37.	25.	25.	37.	0.

C

STANDARD SOURCE SPECTRA

SP85 SOURCE

0.	0.	5224.	6569.	1855.	1767.	1724.	1823.	1598.	1874.
1883.	1920.	1848.	1959.	1921.	2024.	2019.	2127.	2176.	2129.
2062.	2040.	2057.	2062.	2055.	2000.	2098.	2068.	2098.	2091.
2176.	2119.	2241.	2225.	2223.	2306.	2286.	2353.	2358.	2500.
2686.	2920.	3039.	3038.	3208.	3105.	3044.	2938.	2938.	2865.
2748.	2842.	2715.	2742.	2661.	2655.	2630.	2516.	2587.	2642.
2858.	2560.	2540.	2515.	2439.	2544.	2676.	2633.	2775.	2721.
2757.	2896.	2759.	2823.	2858.	2774.	2925.	2688.	2796.	2611.
2616.	2453.	2258.	2098.	1959.	1795.	1564.	1416.	1343.	1252.
1274.	1206.	1126.	1096.	1039.	1080.	1042.	1041.	974.	941.
927.	946.	887.	947.	917.	868.	859.	925.	971.	976.
971.	1016.	1130.	1235.	1339.	1477.	1726.	2310.	2873.	3943.
5346.	7331.	9917.	12520.	15779.	18960.	21819.	24258.	25505.	26034.
26784.	22676.	19649.	16234.	12537.	8971.	6186.	4058.	2538.	1402.
813.	388.	177.	84.	25.	30.	3.	20.	-15.	25.
14.	1.	7.	13.	6.	11.	17.	15.	25.	0.
10.	2.	30.	25.	1.	15.	4.	14.	18.	0.
-7.	8.	22.	20.	13.	13.	9.	11.	3.	13.
23.	15.	21.	10.	28.	16.	-1.	23.	13.	24.
5.	17.	9.	18.	11.	5.	11.	9.	24.	0.

STANDARD SOURCE SPECTRA

CS137 SOURCE

0.	0.	391.	2460.	6458.	10591.	11156.	3228.	2239.	2131.
2095.	2090.	2222.	2348.	2280.	2286.	2169.	2125.	2195.	2140.
2166.	2247.	2251.	2348.	2340.	2220.	2276.	2309.	2362.	2416.
2510.	2501.	2643.	2764.	3020.	3204.	3216.	3312.	3354.	3103.
3190.	2973.	2987.	2977.	2892.	2955.	2881.	2800.	2866.	2771.
2760.	2691.	2665.	2735.	2694.	2665.	2697.	2740.	2711.	2620.
2745.	2652.	2665.	2717.	2757.	2650.	2746.	2685.	2668.	2729.
2703.	2839.	2698.	2754.	2706.	2745.	2654.	2885.	2788.	2761.
2807.	2751.	2857.	2843.	2865.	2777.	2744.	2580.	2476.	2344.
2098.	1972.	1633.	1487.	1341.	1231.	1118.	1086.	988.	916.
921.	817.	819.	773.	746.	741.	752.	740.	739.	670.
775.	761.	835.	873.	1004.	1102.	1180.	1388.	1745.	2272.
3062.	4099.	5740.	7098.	10543.	13552.	16386.	19138.	20885.	22004.
21289.	19441.	16849.	13395.	10019.	7096.	4634.	2866.	1594.	843.
415.	174.	112.	39.	40.	21.	15.	19.	7.	15.
30.	50.	40.	46.	53.	53.	45.	44.	35.	32.
43.	17.	21.	17.	8.	35.	13.	8.	21.	14.
22.	12.	5.	7.	0.	17.	19.	-7.	4.	11.
5.	23.	3.	20.	-7.	8.	22.	11.	7.	7.
15.	14.	20.	18.	6.	4.	12.	11.	13.	0.

STANDARD SOURCE SPECTRA

NR95 SOURCE

0.	0.	2668.	3593.	4041.	3031.	2654.	2603.	2671.	2640.
2629.	2845.	2577.	2937.	2779.	2771.	2785.	2791.	2892.	2917.
2862.	2963.	2932.	2963.	3174.	3165.	3140.	3338.	3366.	3438.
3819.	4031.	4132.	4364.	4312.	4197.	4180.	3867.	3927.	3858.
3830.	3992.	3733.	3688.	3796.	3759.	3705.	3739.	3813.	3598.
3505.	3546.	3579.	3637.	3510.	3618.	3507.	3535.	3491.	3526.
3356.	3551.	3432.	3420.	3499.	3494.	3379.	3506.	3541.	3678.
3490.	3518.	3495.	3529.	3553.	3542.	3494.	3462.	3597.	3519.
3666.	3588.	3649.	3674.	3588.	3725.	3784.	3734.	3776.	3669.
3655.	3375.	3185.	2884.	2516.	2245.	2016.	1728.	1565.	1518.
1335.	1272.	1110.	1147.	1034.	965.	925.	884.	871.	869.
895.	821.	924.	1004.	1053.	1184.	1278.	1445.	1934.	2381.
1354.	4731.	6401.	9042.	12312.	16121.	19855.	23267.	25337.	26389.
25220.	22348.	18716.	14126.	9950.	6597.	3934.	2183.	1084.	475.
246.	139.	30.	19.	23.	25.	2.	38.	14.	7.
21.	-17.	7.	3.	8.	28.	17.	16.	17.	13.
17.	24.	4.	12.	6.	11.	12.	-11.	20.	-1.
8.	15.	6.	-4.	48.	17.	5.	6.	17.	4.
-2.	17.	10.	0.	13.	10.	2.	24.	6.	7.
19.	6.	1.	8.	14.	11.	7.	0.	13.	0.

STANDARD SOURCE SPECTRA

MNS4 SOURCE

0.	0.	2100.	2363.	1905.	2048.	1976.	2036.	2016.	2083.
2144.	2218.	2144.	2132.	2117.	2067.	2146.	2203.	2213.	2165.
2148.	2255.	2398.	2247.	2291.	2341.	2408.	2590.	2727.	3057.
3153.	3230.	3123.	3163.	3022.	3054.	2913.	2931.	2794.	2878.
2709.	2743.	2679.	2662.	2569.	2618.	2586.	2718.	2588.	2568.
2661.	2587.	2554.	2487.	2599.	2576.	2527.	2574.	2633.	2585.
2593.	2639.	2531.	2497.	2583.	2489.	2546.	2563.	2480.	2628.
2538.	2523.	2497.	2573.	2552.	2493.	2445.	2505.	2658.	2589.
2648.	2588.	2629.	2725.	2762.	2800.	2762.	2787.	2850.	2870.
2764.	2678.	2598.	2535.	2400.	2257.	1918.	1742.	1468.	1229.
1109.	1061.	892.	863.	789.	717.	704.	632.	578.	599.
606.	596.	593.	676.	740.	784.	870.	1018.	1218.	1528.
2012.	2442.	4116.	5784.	8124.	10703.	13782.	16576.	18092.	18626.
18116.	15873.	13377.	9948.	5871.	4508.	2510.	1388.	731.	327.
138.	54.	12.	17.	6.	-1.	16.	3.	-3.	6.
2.	-7.	-9.	11.	3.	6.	3.	-3.	10.	10.
16.	-9.	18.	9.	3.	5.	17.	18.	16.	6.
3.	1.	16.	-11.	5.	34.	1.	8.	12.	0.
10.	6.	-3.	-4.	4.	1.	7.	-6.	11.	4.
15.	19.	-1.	7.	1.	-9.	9.	20.	12.	0.

STANDARD SOURCE SPECTRA

ZM65 SOURCE

0.	0.	2845.	3497.	2638.	2760.	2868.	2787.	2915.	2907.
2768.	2879.	2787.	2863.	3015.	3074.	3178.	3027.	3345.	3365.
3502.	3544.	3719.	4056.	4125.	4307.	4297.	3948.	3932.	3837.
3814.	3670.	3737.	3552.	3611.	3608.	3675.	3406.	3354.	3322.
3397.	3245.	3196.	3085.	3222.	3109.	3014.	3160.	2995.	2992.
3033.	3018.	2997.	3078.	3115.	3343.	3465.	4377.	5240.	5726.
5888.	5526.	4689.	3804.	3085.	2916.	2856.	2709.	2705.	2801.
2912.	2737.	2637.	2780.	2801.	2909.	2921.	2806.	2862.	2844.
2844.	2778.	2963.	2937.	3039.	2962.	3166.	3163.	3070.	3160.
3074.	3329.	3228.	3312.	3311.	3498.	3507.	3427.	3383.	3438.
3133.	2903.	2565.	2260.	1965.	1703.	1326.	1240.	1077.	1007.
904.	839.	602.	749.	718.	849.	552.	845.	945.	1210.
1412.	1084.	2834.	4206.	6172.	9050.	12109.	15770.	18207.	19904.
14373.	16786.	12846.	9182.	5920.	3343.	1815.	786.	404.	144.
73.	7.	3.	-8.	-6.	23.	3.	-18.	19.	25.
10.	22.	-14.	5.	12.	15.	22.	12.	7.	-5.
25.	-4.	30.	19.	20.	52.	-5.	-29.	-4.	-30.
-18.	-25.	-17.	-5.	-14.	-1.	17.	9.	17.	0.
-19.	13.	3.	6.	-10.	6.	2.	10.	8.	-3.
7.	10.	-3.	9.	-5.	7.	9.	-3.	8.	0.

STANDARD SOURCE SPECTRA

CO60 SOURCE

0.	0.	6.	400.	5745.	7480.	7814.	8276.	8565.	8322.
8047.	8010.	8336.	8325.	9473.	8592.	8775.	9035.	9426.	9984.
10954.	11836.	12387.	12062.	11626.	10952.	10750.	10604.	10453.	10302.
10246.	9917.	9681.	9567.	9302.	9324.	9150.	9083.	9115.	8993.
8968.	8834.	8911.	8850.	8770.	8671.	8580.	8478.	8438.	8454.
8484.	8465.	8477.	8374.	8227.	8208.	8178.	8150.	8228.	8187.
8142.	7983.	7971.	8079.	8152.	8198.	8253.	8304.	8312.	8377.
8363.	8398.	8270.	8330.	8471.	8679.	8722.	8855.	8758.	8935.
9144.	9412.	9579.	9584.	9616.	9715.	9757.	9911.	9815.	9728.
9303.	9015.	8632.	8149.	7689.	7192.	6903.	6570.	6593.	6455.
6417.	6459.	6353.	6383.	6353.	6217.	6080.	6056.	6547.	7780.
10289.	14441.	19771.	25837.	37514.	32152.	30076.	24955.	18252.	12005.
7394.	4681.	3592.	3876.	5379.	8048.	11942.	16773.	21429.	25125.
26397.	24517.	20311.	14843.	9719.	5619.	2996.	1340.	554.	298.
184.	126.	97.	106.	107.	89.	86.	101.	99.	92.
110.	96.	101.	104.	104.	103.	102.	100.	96.	86.
97.	81.	75.	89.	83.	77.	86.	73.	67.	59.
72.	71.	66.	72.	75.	68.	60.	67.	72.	64.
76.	65.	59.	53.	66.	63.	63.	69.	58.	57.
62.	66.	41.	46.	48.	57.	50.	56.	66.	56.

RESULTS OF PHOTOPEAK FITTING

INDEX	STANDARD SOURCE	PULSE-HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
1	HG203	0.1293640E 03	0.5359105E 01	0.1047150E 06
2	SR85	0.1299184E 03	0.4466077E 01	0.2855184E 06
3	CS137	0.1293185E 03	0.4018573E 01	0.2157564E 06
4	NB95	0.1291098E 03	0.3796879E 01	0.2447797E 06
5	MN54	0.1292172E 03	0.3673843E 01	0.1693172E 06
6	ZN65	0.1294717E 03	0.3192249E 01	0.1631872E 06
7	CO60	0.1303343E 03	0.2804309E 01	0.1764731E 06
7	CO60	0.1154571E 03	0.2711362E 01	0.1925015E 06



NORMALIZED CONTINUUM OF STANDARD SOURCE SPECTRA

CO60 SOURCE

ENERGY= 1.1731 MEV

0.2646846E-01	0.2646846E-01	0.2547525E-01	0.2646874E-01	0.2503630E-01
0.2531907E-01	0.2512569E-01	0.2463023E-01	0.2434987E-01	0.2528646E-01
0.2717772E-01	0.2806072E-01	0.2855738E-01	0.3077833E-01	0.3115567E-01
0.3316884E-01	0.3587442E-01	0.3923366E-01	0.3713191E-01	0.3380536E-01
0.3354563E-01	0.3283470E-01	0.3208455E-01	0.3227669E-01	0.3342755E-01
0.3198130E-01	0.2949777E-01	0.3050390E-01	0.2772917E-01	0.2766851E-01
0.2787067E-01	0.2761759E-01	0.2722292E-01	0.2732985E-01	0.2750296E-01
0.2741931E-01	0.2645379E-01	0.2631678E-01	0.2572940E-01	0.2568012E-01
0.2590969E-01	0.2605839E-01	0.2606782E-01	0.2573343E-01	0.2528319E-01
0.2522488E-01	0.2526300E-01	0.2482699E-01	0.2423742E-01	0.2401540E-01
0.2408047E-01	0.2443208E-01	0.2482979E-01	0.2474272E-01	0.2450509E-01
0.2474865E-01	0.2520159E-01	0.2644264E-01	0.2649899E-01	0.2631078E-01
0.2507915E-01	0.2519117E-01	0.2647607E-01	0.2558392E-01	0.2570218E-01
0.2718165E-01	0.2884625E-01	0.2894346E-01	0.2844362E-01	0.2914467E-01
0.2962512E-01	0.2887173E-01	0.3033938E-01	0.3050105E-01	0.3124147E-01
0.3302990E-01	0.3357356E-01	0.3327994E-01	0.3156444E-01	0.2820840E-01
0.2482439E-01	0.2102492E-01	0.1684629E-01	0.1350956E-01	0.1125828E-01
0.1002641E-01	0.8908115E-02	0.8158208E-02	0.7496569E-02	0.8136358E-02
0.7983018E-02	0.8376796E-02	0.9859957E-02	0.1090555E-01	0.1092027E-01
0.8007824E-02	0.3478980E-02	0.4815639E-03	0.0	0.0
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0



NORMALIZED CONTINUUM OF STANDARD SOURCE SPECTRA

CO60 SOURCE

ENERGY= 1.3320 MEV

0.4341295E-01	0.3994905E-01	0.3101853E-01	0.3134372E-01	0.3194692E-01
0.3264210E-01	0.3255892E-01	0.3041316E-01	0.2796849E-01	0.2813519E-01
0.2718278E-01	0.2592969E-01	0.2646169E-01	0.2691513E-01	0.2834446E-01
0.3362410E-01	0.4400415E-01	0.4849734E-01	0.4649885E-01	0.4134053E-01
0.3941939E-01	0.3774909E-01	0.3776591E-01	0.3795122E-01	0.3678231E-01
0.3625144E-01	0.3527296E-01	0.3460072E-01	0.3376563E-01	0.3345424E-01
0.3307296E-01	0.3302681E-01	0.3308183E-01	0.3283355E-01	0.3238788E-01
0.3181666E-01	0.3129408E-01	0.3119950E-01	0.3129676E-01	0.3116640E-01
0.3093897E-01	0.3084192E-01	0.3031213E-01	0.3081194E-01	0.3072731E-01
0.3023370E-01	0.2963825E-01	0.2947183E-01	0.2969426E-01	0.2981513E-01
0.3009455E-01	0.3041202E-01	0.3069021E-01	0.3076787E-01	0.3073591E-01
0.3075473E-01	0.3077631E-01	0.3062458E-01	0.3050767E-01	0.3074014E-01
0.3131723E-01	0.3249704E-01	0.3374201E-01	0.3418598E-01	0.3404662E-01
0.3363177E-01	0.3299099E-01	0.3211954E-01	0.3122766E-01	0.3070489E-01
0.3058577E-01	0.3093320E-01	0.3167510E-01	0.3280288E-01	0.3411749E-01
0.3541201E-01	0.3650334E-01	0.3723139E-01	0.3755042E-01	0.3749723E-01
0.3566836E-01	0.3210024E-01	0.2939012E-01	0.2627845E-01	0.2164732E-01
0.1768945E-01	0.1368147E-01	0.1119418E-01	0.0771071E-02	0.8544035E-02
0.7739123E-02	0.7466595E-02	0.7829677E-02	0.7647868E-02	0.8917959E-02
0.1752238E-01	0.1566444E-01	0.1152772E-01	0.8399166E-02	0.2011044E-02
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

NORMALIZED CONTINUUM OF STANDARD SPECTRA

HG203 SOURCE

ENERGY= 0.2790 MEV

0.3187532E-02	0.3187532E-02	0.3187532E-02	0.3187532E-02	0.3187532E-02
0.3187532E-02	0.3184200E-02	0.3288179E-02	0.3526246E-02	0.3490566E-02
0.3617539E-02	0.3466145E-02	0.3437530E-02	0.3448054E-02	0.3440871E-02
0.343256E-02	0.3425153E-02	0.3416597E-02	0.3408982E-02	0.3401353E-02
0.3392323E-02	0.3384708E-02	0.3377003E-02	0.3368510E-02	0.3360434E-02
0.3352819E-02	0.3344709E-02	0.3336160E-02	0.3328545E-02	0.3320979E-02
0.3311886E-02	0.3304270E-02	0.3296655E-02	0.3288065E-02	0.3279996E-02
0.3272391E-02	0.3286453E-02	0.3293376E-02	0.3461643E-02	0.3506635E-02
0.32731350E-02	0.32816306E-02	0.4223704E-02	0.4421636E-02	0.4825767E-02
0.5342465E-02	0.5971413E-02	0.6447189E-02	0.6641217E-02	0.6490875E-02
0.6205522E-02	0.5866211E-02	0.5312603E-02	0.4951231E-02	0.4501006E-02
0.4191447E-02	0.3910828E-02	0.3685640E-02	0.3508935E-02	0.3309116E-02
0.3231318E-02	0.3058784E-02	0.3065442E-02	0.3048909E-02	0.3080393E-02
0.3072749E-02	0.3133685E-02	0.3340623E-02	0.3436827E-02	0.3503629E-02
0.3527709E-02	0.3575736E-02	0.3592726E-02	0.3791634E-02	0.3711868E-02
0.3988657E-02	0.3984392E-02	0.4095737E-02	0.4142653E-02	0.4181981E-02
0.4153151E-02	0.4235953E-02	0.4524719E-02	0.4619278E-02	0.4891210E-02
0.4640743E-02	0.4734606E-02	0.4335008E-02	0.4071482E-02	0.3692162E-02
0.2941125E-02	0.2464349E-02	0.2133672E-02	0.2546392E-02	0.2432301E-02
0.2713073E-02	0.2489286E-02	0.1379860E-02	0.9845577E-03	0.1097073E-03
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

NORMALIZED CONTINUUM OF STANDARD SPECTRA

SP85 SOURCE

ENERGY= 0.5150 MEV

0.8613080E-02	0.8613080E-02	0.8613080E-02	0.8613080E-02	0.8613080E-02
0.8613080E-02	0.8608215E-02	0.8528713E-02	0.8630430E-02	0.8471753E-02
0.8836497E-02	0.8912381E-02	0.9170048E-02	0.9616729E-02	0.9803064E-02
0.9489969E-02	0.9279750E-02	0.9340190E-02	0.9342127E-02	0.9149693E-02
0.9468738E-02	0.9437423E-02	0.9501018E-02	0.9795393E-02	0.9782858E-02
0.1013109E-01	0.1009006E-01	0.1043982E-01	0.1050327E-01	0.1069301E-01
0.1141827E-01	0.1256599E-01	0.1356578E-01	0.1381770E-01	0.1443259E-01
0.1395306E-01	0.1346521E-01	0.1527272E-01	0.1278428E-01	0.1274103E-01
0.1241328E-01	0.1235266E-01	0.1209937E-01	0.1195703E-01	0.1149287E-01
0.1185503E-01	0.1175744E-01	0.1160680E-01	0.1148387E-01	0.1120498E-01
0.1149947E-01	0.1208745E-01	0.1228074E-01	0.1241053E-01	0.1253285E-01
0.1287364E-01	0.1270933E-01	0.1290812E-01	0.1292199E-01	0.1266135E-01
0.1256735E-01	0.1185435E-01	0.1151481E-01	0.1050323E-01	0.9460454E-02
0.8596171E-02	0.7492784E-02	0.6499633E-02	0.5967560E-02	0.5637776E-02
0.6502811E-02	0.5082760E-02	0.4857678E-02	0.4840765E-02	0.4744411E-02
0.4614001E-02	0.4331760E-02	0.4222203E-02	0.4219570E-02	0.4161753E-02
0.4198148E-02	0.3940500E-02	0.4021488E-02	0.4329155E-02	0.4461407E-02
0.4438423E-02	0.4767235E-02	0.5238943E-02	0.5440801E-02	0.5418684E-02
0.5867004E-02	0.5450126E-02	0.5084891E-02	0.4318912E-02	0.3620144E-02
0.1118185E-02	0.2926134E-03	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

NORMALIZED CONTINUUM OF STANDARD SPECTRA

CS137 SOURCE

ENERGY= 0.6615 MEV

0.1348738E-01	0.1348738E-01	0.1348738E-01	0.1348738E-01	0.1348737E-01
0.1348738E-01	0.1348738E-01	0.1277048E-01	0.1260379E-01	0.1319968E-01
0.1404967E-01	0.1375093E-01	0.1328460E-01	0.1288334E-01	0.1320855E-01
0.1309776E-01	0.1345938E-01	0.1373693E-01	0.1405391E-01	0.1351476E-01
0.1387120E-01	0.1434949E-01	0.1445173E-01	0.1506506E-01	0.1538948E-01
0.1636503E-01	0.1302287E-01	0.1932214E-01	0.1958003E-01	0.2014377E-01
0.1882005E-01	0.1862280E-01	0.1796592E-01	0.1786355E-01	0.1756370E-01
0.1752998E-01	0.1697312E-01	0.1709351E-01	0.1645710E-01	0.1631015E-01
0.1616671E-01	0.1636372E-01	0.1610909E-01	0.1626462E-01	0.1644033E-01
0.1599937E-01	0.1637444E-01	0.1598057E-01	0.1618997E-01	0.1654923E-01
0.1609733E-01	0.1637650E-01	0.1610442E-01	0.1638399E-01	0.1659661E-01
0.1659748E-01	0.1651491E-01	0.1636949E-01	0.1624810E-01	0.1703753E-01
0.1676790E-01	0.1675555E-01	0.1668051E-01	0.1713950E-01	0.1717465E-01
0.1694055E-01	0.1650336E-01	0.1635910E-01	0.1650016E-01	0.1670513E-01
0.1113158E-01	0.0446379E-02	0.4364104E-02	0.7385693E-02	0.6664019E-02
0.6189469E-02	0.5548727E-02	0.5758253E-02	0.4867404E-02	0.4710667E-02
0.4490603E-02	0.4492920E-02	0.4481297E-02	0.4396524E-02	0.4264932E-02
0.4613373E-02	0.4944768E-02	0.5413757E-02	0.6103054E-02	0.6305855E-02
0.6681945E-02	0.6975420E-02	0.6940488E-02	0.5243298E-02	0.3612872E-02
0.1719738E-02	0.1551420E-03	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

NORMALIZED CONTINUUM OF STANDARD SPECTRA

NR95 SOURCE

ENERGY= 0.7645 MEV

0.1415910E-01	0.1415910E-01	0.1415910E-01	0.1415910E-01	0.1415910E-01
0.1415910E-01	0.1414458E-01	0.1397611E-01	0.1453245E-01	0.1560773E-01
0.1539663E-01	0.1471295E-01	0.1473884E-01	0.1486048E-01	0.1537790E-01
0.1629199E-01	0.1560097E-01	0.1558772E-01	0.1627072E-01	0.1679043E-01
0.1685172E-01	0.1775734E-01	0.1805158E-01	0.2112529E-01	0.2163953E-01
0.2271691E-01	0.2289753E-01	0.2218954E-01	0.2108008E-01	0.2072535E-01
0.2044604E-01	0.2024311E-01	0.1992327E-01	0.1963515E-01	0.2005988E-01
0.1976925E-01	0.1978377E-01	0.1994222E-01	0.1976910E-01	0.1887726E-01
0.1900685E-01	0.1902123E-01	0.1895773E-01	0.1869390E-01	0.1866879E-01
0.1860431E-01	0.1800544E-01	0.1877378E-01	0.1811987E-01	0.1841314E-01
0.1879971E-01	0.1815475E-01	0.1869570E-01	0.1915274E-01	0.1854251E-01
0.1858554E-01	0.1305366E-01	0.1883464E-01	0.1866694E-01	0.1841001E-01
0.1893909E-01	0.1392444E-01	0.1019534E-01	0.1928506E-01	0.1949552E-01
0.1964601E-01	0.1907231E-01	0.1984799E-01	0.1978168E-01	0.1940193E-01
0.1798542E-01	0.1636032E-01	0.1420084E-01	0.1218775E-01	0.1031225E-01
0.8749984E-02	0.8120578E-02	0.7087164E-02	0.6406233E-02	0.6059905E-02
0.5541414E-02	0.5049977E-02	0.4786350E-02	0.4633125E-02	0.4637405E-02
0.4563745E-02	0.4718992E-02	0.5288213E-02	0.5706612E-02	0.5158177E-02
0.6429482E-02	0.7381929E-02	0.7223569E-02	0.6926879E-02	0.3655544E-02
0.1226191E-02	0.8027162E-03	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

NORMALIZED CONTINUUM OF STANDARD SPECTRA

MN54 SOURCE

ENERGY= 0.8359 MEV

0.1546274E-01	0.1546274E-01	0.1546274E-01	0.1546274E-01	0.1546273E-01
0.1546274E-01	0.1549361E-01	0.1611589E-01	0.1674518E-01	0.1658096E-01
0.1632813E-01	0.1606495E-01	0.1626572E-01	0.1686817E-01	0.1683760E-01
0.1652733E-01	0.1718110E-01	0.1756697E-01	0.1740975E-01	0.1808926E-01
0.1875734E-01	0.2032693E-01	0.2261020E-01	0.2418567E-01	0.2447919E-01
0.2413413E-01	0.2338626E-01	0.2312220E-01	0.2241294E-01	0.2170475E-01
0.2179625E-01	0.2089286E-01	0.2071525E-01	0.2034710E-01	0.1983532E-01
0.1993791E-01	0.2061477E-01	0.1980856E-01	0.2015233E-01	0.1999616E-01
0.1956236E-01	0.1943695E-01	0.1981984E-01	0.1946681E-01	0.1989577E-01
0.1998572E-01	0.1988596E-01	0.2002691E-01	0.1928532E-01	0.1962597E-01
0.1919847E-01	0.1957549E-01	0.1925175E-01	0.1993940E-01	0.1942864E-01
0.1924902E-01	0.1967858E-01	0.1946146E-01	0.1894521E-01	0.1962217E-01
0.2033175E-01	0.2003773E-01	0.2002107E-01	0.2018455E-01	0.2009079E-01
0.2134420E-01	0.2125172E-01	0.2149119E-01	0.2103122E-01	0.2160615E-01
0.2056758E-01	0.1973923E-01	0.1880374E-01	0.1727609E-01	0.1420547E-01
0.1223116E-01	0.9852346E-02	0.8424409E-02	0.7544734E-02	0.6688692E-02
0.6016973E-02	0.5461734E-02	0.5064208E-02	0.4511420E-02	0.4619132E-02
0.4666742E-02	0.4550171E-02	0.5261857E-02	0.5741943E-02	0.6178394E-02
0.6034442E-02	0.7381525E-02	0.7173019E-02	0.6324206E-02	0.4283320E-02
0.1376360E-02	0.1186633E-03	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

NORMALIZED CONTINUUM OF STANDARD SPECTRA

MN55 SOURCE

ENERGY= 1.1143 MEV

0.2270554E-01	0.2270554E-01	0.2270554E-01	0.2271953E-01	0.2269836E-01
0.2260320E-01	0.2245085E-01	0.2207566E-01	0.2181071E-01	0.2161495E-01
0.2234658E-01	0.2347403E-01	0.2428272E-01	0.2368174E-01	0.2556257E-01
0.2600495E-01	0.2686243E-01	0.2933428E-01	0.3176779E-01	0.3333212E-01
0.3255036E-01	0.3166966E-01	0.3006403E-01	0.2942329E-01	0.2847259E-01
0.2770959E-01	0.2753382E-01	0.2796156E-01	0.2705504E-01	0.2604904E-01
0.2639725E-01	0.2663009E-01	0.2592496E-01	0.2528101E-01	0.2594583E-01
0.2484092E-01	0.2554516E-01	0.2445731E-01	0.2448476E-01	0.2434481E-01
0.2431791E-01	0.2435781E-01	0.2423321E-01	0.2408299E-01	0.2448367E-01
0.2420381E-01	0.2424288E-01	0.2485366E-01	0.2476824E-01	0.2253740E-01
0.2419280E-01	0.2334866E-01	0.2351841E-01	0.2399490E-01	0.2410341E-01
0.2269776E-01	0.2358418E-01	0.2441110E-01	0.2428246E-01	0.2389934E-01
0.2428243E-01	0.2434974E-01	0.2383744E-01	0.2510896E-01	0.2534937E-01
0.2538831E-01	0.2609337E-01	0.2659136E-01	0.2655491E-01	0.2639998E-01
0.2804121E-01	0.2784570E-01	0.2835280E-01	0.2973722E-01	0.2950066E-01
0.2887240E-01	0.2379969E-01	0.2586716E-01	0.2291945E-01	0.1944315E-01
0.1613903E-01	0.127533E-01	0.1070836E-01	0.8990251E-02	0.8153722E-02
0.7246811E-02	0.6763161E-02	0.6262356E-02	0.6785888E-02	0.7186978E-02
0.7212043E-02	0.8427605E-02	0.9583656E-02	0.9327102E-02	0.8332899E-02
0.5082455E-02	0.1722422E-02	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0

SYSTEM RESPONSE FUNCTION MATRIX

0.1000E-01	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.7850E-01	0.9214E-00	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.8247E-01	0.1086E-00	0.4190E-01	0.2367E-00	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.9137E-01	0.1294E-00	0.8230E-01	0.6969E-00	0.0	0.3039E-04	0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.9340E-01	0.1233E-00	0.1249E-00	0.5471E-01	0.6033E-00	0.3223E-03	0.0	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.9418E-01	0.1156E-00	0.1217E-00	0.1131E-00	0.4401E-01	0.6290E-00	0.1234E-02	0.0	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.9015E-01	0.1059E-00	0.1149E-00	0.1110E-00	0.8372E-01	0.3772E-01	0.4537E-00	0.2712E-02	0.0	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.7772E-01	0.9474E-01	0.1099E-00	0.1101E-00	0.1920E-00	0.7023E-01	0.3430E-01	0.4074E-00	0.5052E-02	0.0	0.0
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.7217E-01	0.8471E-01	0.0997E-01	0.9013E-01	0.8809E-01	0.9303E-01	0.5723E-01	0.3270E-01	0.3725E-00	0.8344E-02	0.0
0.4642E-10	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.6768E-01	0.8091E-01	0.9247E-01	0.8148E-01	0.7902E-01	0.7801E-01	0.8788E-01	0.4774E-01	0.3393E-01	0.3404E-00	0.0
0.1183E-01	0.2631E-03	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.6422E-01	0.7416E-01	0.3721E-01	0.7623E-01	0.7274E-01	0.6992E-01	0.7265E-01	0.8257E-01	0.4267E-01	0.3523E-01	0.0
0.3791E-00	0.1411E-01	0.3579E-07	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.6280E-01	0.6723E-01	0.3339E-01	0.7264E-01	0.6753E-01	0.6534E-01	0.6523E-01	0.6934E-01	0.7692E-01	0.4059E-01	0.0
0.3576E-01	0.2769E-00	0.1631E-01	0.1810E-06	0.0	0.0	0.0	0.0	0.0	0.0	0.0
0.6233E-01	0.5952E-01	0.8032E-01	0.6956E-01	0.6396E-01	0.6151E-01	0.6023E-01	0.6253E-01	0.6527E-01	0.7157E-01	0.0
0.4770E-01	0.3634E-01	0.2378E-00	0.1923E-01	0.7976E-06	0.0	0.0	0.0	0.0	0.0	0.0
0.5997E-01	0.6136E-01	0.7802E-01	0.6758E-01	0.6137E-01	0.5839E-01	0.5642E-01	0.5787E-01	0.6219E-01	0.5886E-01	0.0
0.6770E-01	0.4127E-01	0.3579E-01	0.2244E-00	0.1864E-01	0.2163E-05	0.0	0.0	0.0	0.0	0.0
0.5884E-01	0.4398E-01	0.7480E-01	0.6511E-01	0.5925E-01	0.5576E-01	0.5330E-01	0.5477E-01	0.5600E-01	0.6019E-01	0.0
0.5290E-01	0.6444E-01	0.4235E-01	0.3458E-01	0.2936E-00	0.1914E-01	0.5179E-05	0.0	0.0	0.0	0.0
0.5751E-01	0.3782E-01	0.7022E-01	0.6501E-01	0.5711E-01	0.5406E-01	0.5208E-01	0.5217E-01	0.5365E-01	0.5760E-01	0.0
0.5334E-01	0.4781E-01	0.6241E-01	0.4247E-01	0.3425E-01	0.1843E-00	0.2011E-01	0.0323E-05	0.0	0.0	0.0
0.5596E-01	0.3329E-01	0.6310E-01	0.6439E-01	0.5614E-01	0.5247E-01	0.5020E-01	0.4710E-01	0.5113E-01	0.5135E-01	0.0
0.5958E-01	0.4613E-01	0.4365E-01	0.6190E-01	0.4772E-01	0.3445E-01	0.1680E-00	0.2735E-01	0.2164E-04	0.6471E-10	0.0
0.5345E-01	0.3167E-01	0.5242E-01	0.6435E-01	0.6511E-01	0.5148E-01	0.4854E-01	0.4698E-01	0.4580E-01	0.5108E-01	0.0
0.5212E-01	0.5557E-01	0.3984E-01	0.4153E-01	0.6359E-01	0.3655E-01	0.3540E-01	0.1544E-00	0.2016E-01	0.2846E-04	0.0
0.5913E-01	0.3291E-01	0.3627E-01	0.6508E-01	0.5474E-01	0.5062E-01	0.4768E-01	0.4642E-01	0.4260E-01	0.8861E-01	0.0
0.4873E-01	0.5729E-01	0.4561E-01	0.3437E-01	0.4116E-01	0.6711E-01	0.2913E-01	0.3809E-01	0.1427E-00	0.1968E-01	0.0
0.4611E-01	0.3590E-01	0.1726E-01	0.7117E-01	0.5422E-01	0.4924E-01	0.4729E-01	0.4532E-01	0.4339E-01	0.4289E-01	0.0
0.4954E-01	0.4746E-01	0.5919E-01	0.3765E-01	0.2744E-01	0.4331E-01	0.6853E-01	0.2074E-01	0.4208E-01	0.1325E-00	0.0

(F)

RESPONSE MATRIX ENERGY INTERVAL MIDPOINTS IN MEV

0.0500 0.1500 0.2500 0.3500 0.4500 0.5500 0.6500 0.7500 0.8500 0.9500
1.0500 1.1500 1.2500 1.3500 1.4500 1.5500 1.6500 1.7500 1.8500 1.9500

PULSE HEIGHT IN CHANNELS (MIDPOINTS)

0.54 1.54 2.53 3.52 4.51 5.51 6.51 7.51 8.51 9.51
10.51 11.51 12.50 13.50 14.50 15.50 16.50 17.50 18.50 19.50

RESPONSE MATRIX PHOTOPEAKS

0.9541 0.8567 0.7593 0.6659 0.5846 0.5094 0.4503 0.4114 0.3846 0.3599
0.3329 0.3046 0.2765 0.2504 0.2269 0.2044 0.1849 0.1679 0.1532 0.1405

(G)

GAMMA SOURCE	SOURCE STRENGTH (CURIES)	REFERENCE ENERGY (MEV)	SOURCE DIAMETER (CM.)	SOURCE HALF-LIFE (MINUTES)	NUMBER OF PHA RUNS THIS SET	NUMBER OF CHANNELS PER SPECTRUM	HALF-LIFE MULTIPLIER
RU02	1.0000	1.0000	1.0000	0.1053E-00	1	200	0.5260E-00

M222	M66	MZ	MX	MPX	MSM
7	0	20	0	0	1

SOURCE CRYSTAL DISTANCE (CM.)	CRYSTAL AT-SOURCE (DEGREES) POLAR	ANGLE (DEGREES) AZIMUTH	COUNTING DURATION (MINUTES)	REFERENCE TIME (DAYS)	PHA COARSE GAIN	MONITOR PULSE HEIGHT (CHANNELS)	MONITOR ENERGY (MEV)	SPECTRA ZERO SHIFT (CHANNELS)	BACKGROUND SIGNAL	BACKGROUND MULTIPLIER
100.000	0.0	0.0	1.000	1.00	1.00	127.0000	1.2700	0.0	0.0	0.0

(H)

ITERATING OUTPUT ON ITERATION LOOPS NUMBERED BELOW

1	2	3	4	5
7	10	15	20	30
0	0	0	0	0
0	0	0	0	0

(I)

CARD SET 12

2	115	0	1.2700	0	10	144	1.3600	2	45	0	0.5000
0	7	66	0.5750	2	69	0	0.7600	0	12	96	0.8750
1	143	167	1.5400								

(J)

SPECTRUM NUMBER 1 FOR PU02 SOURCE

(AFTER BACKGROUND SUBTRACTION)

0.	0.	0.	0.	0.	0.	0.	0.	0.	0.
0.	0.	0.	0.	0.	0.	0.	0.	0.	0.
100513.	124001.	81387.	43872.	33958.	31632.	26716.	22787.	20076.	18482.
16846.	16763.	14977.	14320.	13427.	12768.	12231.	11751.	11418.	11237.
10481.	9918.	9722.	9510.	9269.	8743.	8626.	8950.	9931.	11764.
13964.	15364.	14272.	11787.	9944.	10473.	13311.	17066.	19334.	18716.
14363.	10043.	7109.	5746.	8198.	5078.	4830.	4574.	4404.	4742.
4327.	4721.	4915.	5502.	6167.	6482.	6663.	6361.	5726.	4988.
4457.	4222.	4051.	4420.	4869.	5758.	6432.	7025.	7106.	6655.
5800.	5046.	4196.	3705.	3371.	3347.	3318.	3407.	3304.	3339.
3102.	3124.	2915.	2730.	2549.	2501.	2162.	2195.	1958.	1937.
1744.	1735.	1676.	1576.	1508.	1589.	1678.	1838.	2116.	2575.
3324.	4313.	5555.	6977.	8397.	8042.	9195.	8633.	7318.	5224.
4551.	3475.	2854.	2476.	2244.	2182.	1960.	1634.	1485.	1300.
1144.	1050.	1001.	906.	930.	915.	915.	886.	967.	1051.
1124.	1160.	1140.	1020.	965.	886.	797.	660.	597.	530.
433.	425.	375.	345.	333.	334.	333.	306.	316.	327.
354.	308.	334.	325.	322.	325.	345.	352.	344.	299.
293.	277.	292.	253.	221.	221.	211.	206.	199.	207.
159.	208.	179.	166.	224.	198.	217.	235.	230.	311.

(K)

UNKNOWN SPECTRUM BEFORE PEAK FITTING

0.	0.	0.	0.	0.	0.	0.	0.	0.	0.
0.	0.	0.	0.	0.	0.	0.	0.	0.	0.
94310.	98613.	73792.	53809.	38802.	30621.	26602.	23218.	20482.	18500.
17447.	15987.	15194.	14273.	13457.	12793.	13240.	11853.	11519.	11100.
10447.	10055.	9764.	9439.	9134.	8800.	8700.	9193.	10223.	12145.
14054.	14318.	13482.	11813.	11420.	11547.	14737.	16697.	18426.	17014.
13778.	13478.	7588.	5950.	5404.	5022.	4804.	4590.	4435.	4378.
4454.	4650.	5111.	5586.	6104.	6453.	6549.	6215.	5677.	5065.
4836.	4241.	4192.	4487.	5024.	5746.	6454.	6457.	6043.	6518.
3752.	5003.	4244.	3755.	3470.	3363.	3364.	3360.	3740.	7280.
3255.	3064.	2903.	2754.	2606.	2412.	2214.	2085.	1994.	1800.
1736.	1709.	1652.	1591.	1552.	1564.	1651.	1625.	2177.	2657.
3389.	4432.	5655.	7014.	8229.	8888.	8972.	8326.	7205.	5800.
4596.	3623.	2949.	2531.	2280.	2074.	1891.	1680.	1487.	1100.
1151.	1154.	982.	948.	930.	909.	898.	926.	930.	1054.
1121.	1134.	1105.	1047.	968.	874.	776.	695.	603.	524.
459.	400.	375.	353.	339.	329.	322.	310.	320.	306.
329.	331.	330.	326.	323.	313.	346.	340.	324.	307.
202.	283.	274.	254.	231.	216.	208.	205.	204.	199.
192.	181.	173.	194.	203.	211.	219.	241.	278.	

(L)

EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.12700E 01	0.99271E 00	0.98589E 00	0.95513E 00	0.71771E 00	0.67090E 00
				PHOTON-NUMBER	=	0.3554105E 06
				PHOTOFRACTION	=	0.2700069E 00

DISCRETE PHOTOPEAK FITTING DATA

LOCATION CHANNEL TO CHANNEL	ENERGY (MEV)	PULSE-HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
115 144	0.1270000E 01	0.1260716E 03	0.3371937E 01	0.5459844E 05



EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.13610E 01	0.99297E 00	0.98630E 00	0.95671E 00	0.70603E 00	0.66065E 00
				PHOTON-NUMBER	=	0.3619769E 05
				PHOTOFRACTION	=	0.2477741E 00

DISCRETE PHOTOPEAK FITTING DATA

LOCATION CHANNEL TO CHANNEL	ENERGY (MEV)	PULSE-HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
115 144	0.1360000E 01	0.1352512E 03	0.2853853E 01	0.5925289E 04

EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.50000E 00	0.98884E 01	0.97840E 00	0.93207E 00	0.87879E 00	0.79245E 00
					PHOTON-NUMBER PHOTOFRACTION	0.7474144E 00 0.5475301E 00

DISCRETE PHOTOPEAK FITTING DATA

LOCATION CHANNEL-TO-CHANNEL	ENERGY (MEV)	PULSE-HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
46 65	0.5000000E 00	0.5142572E 02	0.1918112E 01	0.3320754E 05

EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.57500E 00	0.98947E 01	0.97964E 00	0.93574E 00	0.85553E 00	0.77600E 00
					PHOTON-NUMBER PHOTOFRACTION	0.1643307E 06 0.4920496E 00

DISCRETE PHOTOPEAK FITTING DATA

LOCATION CHANNEL-TO-CHANNEL	ENERGY (MEV)	PULSE-HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
46 65	0.5750000E 00	0.5862791E 02	0.2084412E 01	0.6274670E 05

EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.76100E 00	0.99070E 00	0.99196E 00	0.94309E 00	0.80624E 00	0.71970E 00
					PHOTON NUMBER PHOTOFRACTION	= 0.5311842E 00 0.4080024E 00

DISCRETE PHOTOPEAK FITTING DATA

LOCATION CHANNEL TO CHANNEL	ENERGY (MEV)	PULSE HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
69 95	0.7600000E 00	0.7622591E 02	0.2505323E 01	0.1603270E 05

EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.87500E 00	0.99129E 00	0.98308E 00	0.94674E 00	0.78161E 00	0.72112E 00
					PHOTON NUMBER PHOTOFRACTION	= 0.4638912E 00 0.3784605E 00

DISCRETE PHOTOPEAK FITTING DATA

LOCATION CHANNEL TO CHANNEL	ENERGY (MEV)	PULSE HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
69 96	0.8750000E 00	0.8821091E 02	0.2643516E 01	0.2357678E 05

EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.15406E 01	0.00342E 02	0.00725E 00	0.05067E 00	0.00312E 00	0.06281E 00
				PHOTON NUMBER		0.077653E 05
				PHOTOFRACTION		0.006362E 00

DISCRETE-PHOTOPEAK FITTING-DATA

LOCATION CHANNEL TO CHANNEL	ENERGY (MEV)	PULSE-HEIGHT (CHANNELS)	STANDARD DEVIATION (CHANNELS)	AREA (COUNTS/TIME)
143 167	0.1540002E 01	0.1824641E 03	0.3287717E 01	0.3583153E 04

GAIN PARAMETERS- AFTER CALL SINGLE- IN MAIN

[illegible]

INPUT SPECTRUM GAIN CHANGED TO 19 CHANNELS
GAIN CHANGE RATIO = 200.100000/ 20.0000

0.0 0.327934E 05 0.4236717E 06 0.9729356E 05 0.5499171E 05
0.0014493E 05 0.2318827E 05 0.2177730E 05 0.1760173E 05 0.1388685E 05
0.1141115E 05 0.0435738E 04 0.9538375E 04 0.1037918E 05 0.9399789E 04
0.5958785E 04 0.3509482E 04 0.3290112E 04 0.2364875E 04

①

NORMALIZED INPUT SPECTRUM

0.0 0.4180E-01 0.5400E-01 0.1113E-01 0.7009E-01 0.5119E-01 0.3593E-01 0.2776E-01 0.2243E-01 0.1779E-01
0.1454E-01 0.1203E-01 0.1216E-01 0.1323E-01 0.1059E-01 0.7595E-02 0.4473E-02 0.4194E-02 0.3014E-02
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 1 1
0.0 0.4180E-01 0.5400E-01 0.1113E-01 0.7009E-01 0.5119E-01 0.3593E-01 0.2776E-01 0.2243E-01 0.1779E-01
0.1454E-01 0.1203E-01 0.1216E-01 0.1323E-01 0.1059E-01 0.7595E-02 0.4473E-02 0.4194E-02 0.3014E-02
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 1 1
0.9248E-01 0.1403E-01 0.4782E-01 0.1026E-01 0.5893E-01 0.3936E-01 0.2501E-01 0.1806E-01 0.1381E-01 0.1053E-01
0.8111E-02 0.6182E-02 0.5104E-02 0.4434E-02 0.3236E-02 0.2112E-02 0.1141E-02 0.8533E-03 0.6148E-03
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 2 2
0.0 0.1246E-01 0.5098E-01 0.1207E-01 0.8337E-01 0.6832E-01 0.5162E-01 0.4267E-01 0.3645E-01 0.2976E-01
0.2609E-01 0.2340E-01 0.2896E-01 0.3987E-01 0.3466E-01 0.2731E-01 0.1754E-01 0.2061E-01 0.1765E-01
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 2 2
0.1730E-01 0.1381E-01 0.5559E-01 0.1275E-01 0.8270E-01 0.5996E-01 0.4305E-01 0.3389E-01 0.2806E-01 0.2317E-01
0.1935E-01 0.1604E-01 0.1453E-01 0.1439E-01 0.1140E-01 0.8239E-02 0.4741E-02 0.4211E-02 0.2935E-02
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 3 3
0.0 0.3772E-02 0.5924E-01 0.1053E-01 0.7126E-01 0.5833E-01 0.4308E-01 0.3404E-01 0.2914E-01 0.2274E-01
0.1969E-01 0.1756E-01 0.2423E-01 0.3654E-01 0.3198E-01 0.2518E-01 0.1654E-01 0.2053E-01 0.1813E-01
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 3 3
0.9347E-01 0.1192E-01 0.5325E-01 0.1103E-01 0.6999E-01 0.5107E-01 0.3628E-01 0.2846E-01 0.2354E-01 0.1942E-01
0.1638E-01 0.1373E-01 0.1289E-01 0.1331E-01 0.1078E-01 0.7789E-02 0.4563E-02 0.4196E-02 0.3001E-02
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 4 4
0.0 0.1723E-02 0.6707E-01 0.1062E-01 0.7147E-01 0.5847E-01 0.4267E-01 0.3408E-01 0.2777E-01 0.2272E-01
0.1741E-01 0.1537E-01 0.2286E-01 0.3630E-01 0.3142E-01 0.2455E-01 0.1629E-01 0.2051E-01 0.1821E-01
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 4 4
0.9328E-01 0.1171E-01 0.5383E-01 0.1100E-01 0.6913E-01 0.5031E-01 0.3533E-01 0.2746E-01 0.2252E-01 0.1837E-01
0.1545E-01 0.1297E-01 0.1243E-01 0.1318E-01 0.1063E-01 0.7659E-02 0.4489E-02 0.4192E-02 0.3012E-02
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 5 5
0.0 0.4720E-03 0.6026E-01 0.1074E-01 0.7247E-01 0.5953E-01 0.4338E-01 0.3443E-01 0.2767E-01 0.1997E-01
0.1630E-01 0.1426E-01 0.2235E-01 0.3644E-01 0.3130E-01 0.2435E-01 0.1623E-01 0.2052E-01 0.1822E-01
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 5 5
0.9353E-01 0.1168E-01 0.5400E-01 0.1109E-01 0.6961E-01 0.5064E-01 0.3541E-01 0.2737E-01 0.2229E-01 0.1798E-01
0.1506E-01 0.1261E-01 0.1227E-01 0.1318E-01 0.1060E-01 0.7618E-02 0.4476E-02 0.4193E-02 0.3014E-02
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 7 7
0.0 0.6749E-04 0.6024E-01 0.1078E-01 0.7313E-01 0.6048E-01 0.4441E-01 0.3531E-01 0.2807E-01 0.1953E-01
0.1550E-01 0.1317E-01 0.2204E-01 0.3664E-01 0.3126E-01 0.2425E-01 0.1622E-01 0.2053E-01 0.1822E-01
INTERMEDIATE ITERATING OUTPUT (IT, MN, AND PHI(I)) 7 7
0.9369E-01 0.1167E-01 0.5401E-01 0.1113E-01 0.7007E-01 0.5109E-01 0.3575E-01 0.2756E-01 0.2229E-01 0.1774E-01
0.1473E-01 0.1229E-01 0.1218E-01 0.1322E-01 0.1059E-01 0.7598E-02 0.4473E-02 0.4194E-02 0.3014E-02
ITERATED SPECTRUM
0.9370E-01 0.1167E-01 0.5401E-01 0.1113E-01 0.7011E-01 0.5116E-01 0.3583E-01 0.2763E-01 0.2232E-01 0.1771E-01
0.1466E-01 0.1221E-01 0.1217E-01 0.1322E-01 0.1059E-01 0.7595E-02 0.4473E-02 0.4194E-02 0.3014E-02

IT, SU, TERM

9 0.7845656E 06 0.1341255E 00

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EFFICIENCY FACTORS

INDEX	ENERGY (MEV)	AIR ATTENUATION	CLADDING ATTENUATION	LUCITE ATTENUATION	CRYSTAL EFFICIENCY	TOTAL EFFICIENCY
1	0.90000E-01	0.97449E 00	0.92534E 00	0.86291E 00	0.99959E 00	0.77744E 00
2	0.15100E 00	0.98243E 00	0.95827E 00	0.89711E 00	0.99069E 00	0.84282E 00
3	0.25100E 00	0.98547E 00	0.97149E 00	0.91268E 00	0.97227E 00	0.84907E 00
4	0.35100E 00	0.98711E 00	0.97502E 00	0.92190E 00	0.93679E 00	0.83120E 00
5	0.45000E 00	0.98833E 00	0.97741E 00	0.92911E 00	0.89770E 00	0.80570E 00
6	0.55000E 00	0.98926E 00	0.97925E 00	0.93457E 00	0.86285E 00	0.78110E 00
7	0.65100E 00	0.99002E 00	0.98060E 00	0.93908E 00	0.83258E 00	0.75904E 00
8	0.75100E 00	0.99064E 00	0.98184E 00	0.94275E 00	0.80820E 00	0.74109E 00
9	0.85100E 00	0.99116E 00	0.98286E 00	0.94609E 00	0.78497E 00	0.72525E 00
10	0.95100E 00	0.99162E 00	0.98371E 00	0.94863E 00	0.76746E 00	0.71017E 00
11	0.11500E 01	0.99200E 00	0.98446E 00	0.95084E 00	0.75080E 00	0.69715E 00
12	0.11500E 01	0.99234E 00	0.98514E 00	0.95287E 00	0.73548E 00	0.68511E 00
13	0.12500E 01	0.99265E 00	0.98576E 00	0.95477E 00	0.72061E 00	0.67324E 00
14	0.13500E 01	0.99294E 00	0.98634E 00	0.95654E 00	0.70641E 00	0.66177E 00
15	0.14500E 01	0.99321E 00	0.98686E 00	0.95817E 00	0.69300E 00	0.65003E 00
16	0.15500E 01	0.99344E 00	0.98729E 00	0.95960E 00	0.68010E 00	0.64207E 00
17	0.16500E 01	0.99364E 00	0.98765E 00	0.96084E 00	0.66731E 00	0.63476E 00
18	0.17500E 01	0.99383E 00	0.98800E 00	0.96206E 00	0.65447E 00	0.62770E 00
19	0.18500E 01	0.99401E 00	0.98834E 00	0.96320E 00	0.64621E 00	0.62004E 00

Q

DIFFERENTIAL FLUX AT ITERATION NUMBER = 0

0.0	0.20170E 02	0.55051E 06	0.11173E 06	0.11960E 06
0.22245E 06	0.87224E 05	0.85465E 05	0.10067E 06	0.43120E 05
0.17220E 05	0.14755E 05	0.30998E 06	0.14714E 05	0.44062E 05
0.54603E 05	0.20053E 05	0.25661E 05	0.23020E 05	

FITTING DIFFERENCES

0.0	0.2994000E 00	0.5198365E-01	0.9333937E-01	0.8179724E-02
0.1331687E-02	0.3445745E-03	0.1285672E-03	0.5682978E-04	

R

NUMBER AND ENERGY SPECTRUM AT THE CRYSTAL

INCREMENT	ENERGY (MEV)	NUMBER FLUX (PHOTONS/CM**2-SEC)	ENERGY FLUX (MEV/CM**2-SEC)
1	0.05000	0.0	0.0
2	0.15000	0.7371482E-02	0.1105722E-02
3	0.25000	0.2439800E-03	0.5664695E-02
4	0.35000	0.3717918E-02	0.1301271E-02
5	0.45000	0.4006633E-02	0.1803440E-02
6	0.55000	0.8129715E-02	0.4471341E-02
7	0.65000	0.3167755E-02	0.2072848E-02
8	0.75000	0.3123466E-02	0.2342598E-02
9	0.85000	0.3679332E-02	0.3127431E-02
10	0.95000	0.1575904E-02	0.1497108E-02
11	1.05000	0.6293324E-01	0.6607986E-01
12	1.15000	0.5396131E-01	0.6205543E-01
13	1.25000	0.1130977E-03	0.1416398E-03
14	1.35000	0.5377463E-02	0.7259566E-02
15	1.45000	0.1410333E-02	0.2334981E-02
16	1.55000	0.1995564E-02	0.3093123E-02
17	1.65000	0.7238831E-01	0.1209266E-01
18	1.75000	0.9378304E-01	0.1641202E-01
19	1.85000	0.8412992E-01	0.1564402E-01

(S)

INTEGRATED RESULTS AT SOURCE AND CRYSTAL

ENERGY INTEGRATED PHOTON (BREMS.) VALUES AT THE CRYSTAL

PHOTON NUMBER (PHOTONS/CM**2-SEC) = 0.7175212E-03
 PHOTON ENERGY (MEV/CM**2-SEC) = 0.5421594E-03
 PHOTON DOSE (ROENTGENS/HOUR) = 0.1714621E-02

AVERAGE ENERGY (MEV) = 0.7558792E-01

PHOTON NUMBER / SOURCE EMITTED BETA NUMBER (PHOTONS/CM**2-SEC)/(BETA/SEC) = 0.1939246E-07
 PHOTON ENERGY / SOURCE EMITTED BETA ENERGY ((MEV/CM**2-SEC)/(MEV)) = 0.6423594E-03
 PHOTON ENERGY / SOURCE EMITTED BETA NUMBER (MEV/CM**2-SEC)/(BETA/SEC) = 0.1465936E-07
 PHOTON DOSE / SOURCE EMITTED BETA NUMBER (R/HR)/(BETA/SEC) = 0.2742219E-13
 PHOTON DOSE / SOURCE EMITTED BETA NUMBER PER SOURCE VOLUME (((R/HR)/(BETA/SEC))/CM**3) = 0.3491502E-13

AT THE SOURCE CYLINDER

PHOTON DOSE / SOURCE EMITTED BETA NUMBER (R/HR)/(BETA/SEC) = 0.7565362E-10
 PHOTON DOSE / SOURCE EMITTED BETA NUMBER PER SOURCE VOLUME (((R/HR)/(BETA/SEC))/CM**3) = 0.0632901E-10

(T)