

N71-20622

SERIES BASE NO. / VOL. / REISSUE

TM-(L)-HU-033 / 003/00

NASA CR-103076

TECHNICAL MEMORANDUM

(TM Series)

CASE FILE COPY

This document was produced by SDC in performance of contract NAS8-25471

EXPERIMENT SIMULATION
SOFTWARE (FPE 5.1)

March 15, 1971

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March 15, 1971

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ABSTRACT

This document describes the software required to automate the acquisition, analysis and display of the prime data for the X-ray Polarimeter Experiment (FPE 5.1). In addition, software to generate simulated input to test the experiment software is described. This work was performed under contract number NAS8-25471 for the Computation Laboratory of the George C. Marshall Space Flight Center, Huntsville, Alabama.

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SECTION 1. INTRODUCTION

This report is one of two produced during the extension of NASA study contract NAS8-25471, "Analyses of the Requirements for Computer Control and Data Processing Experiment Subsystems." A second report, SDC document TM-(L)-HU-033/004/00 entitled "Experiment Support Software Techniques Analysis (FPE 5.3A)," describes the software techniques applicable to Space Station experiment FPE 5.3A - Solar Imaging X-ray Telescope. Both reports were prepared by the System Development Corporation's Huntsville Space Projects staff.

This report describes and documents the software required to analyze the primary data of the X-ray Polarimeter Experiment (FPE 5.1). Software specifications were presented in SDC document TM-(L)-HU-033/001/00, dated October 15, 1970. A description of the experiment procedure and hardware is contained in SDC document TM-(L)-HU-033/000/00, dated May 15, 1970.

This report is presented in two parts:

- 1) X-ray Polarimeter Experiment Program
- 2) Data Generation Program

1.1 Use of Existing Software

All programs were coded in FORTRAN IV and were tested and run on the MSFC IBM 7094 computer. Where possible, use was made of existing library routines including:

- QUIK3V - SC-4020 plot routine
- BIT - Logical bit testing function
- CTOBCD - Convert to binary routine
- BIN - AND/OR routine
- LAGRNG - Lagrangian interpolation routine
- EVAL - Power spectrum routine
- HARM - Harmonic analysis (FFT) routine

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LSF - Bit manipulation routine
FLD - Bit transfer routine
RANDOM - Random number generator
PDUMP - Partial core dump

Subroutines EVAL, HARM and LSF are not standard library routines but were made available by the MSFC Computation Laboratory personnel. Brief descriptions of these three routines are included herein. Descriptions of other library routines may be found in the MSFC IBM 7094 Preliminary Reference Manual.

1.2 Description of Experiment

The X-ray polarimeter experiment hardware consists of the optics and instrumentation necessary to determine the degree of polarization of X-ray celestial sources. The polarimeter is located at the focal point of a large grazing incidence X-ray telescope. Data is collected in electronic format and is processed in real time on board the Space Station. Two types of data are considered--polarimeter and pulsar. Polarimeter data is analyzed to determine the degree of polarization and the intensity of the X-ray source. Pulsar data is analyzed to determine if an X-ray source is a pulsar, and if so, the frequency of pulsation is calculated. Analyzed data is presented in both graphic and tabular format. (For a more detailed explanation of the experiment, see SDC document TM-(L)-HU-033/000/00.)

The software described in Section 2 is limited to that required for acquisition, analysis and display of primary data. Section 3 describes the software to simulate the output of the polarimeter experiment. This simulated data serves as input to the experiment program.

SECTION 2. X-RAY POLARIMETER EXPERIMENT PROGRAM

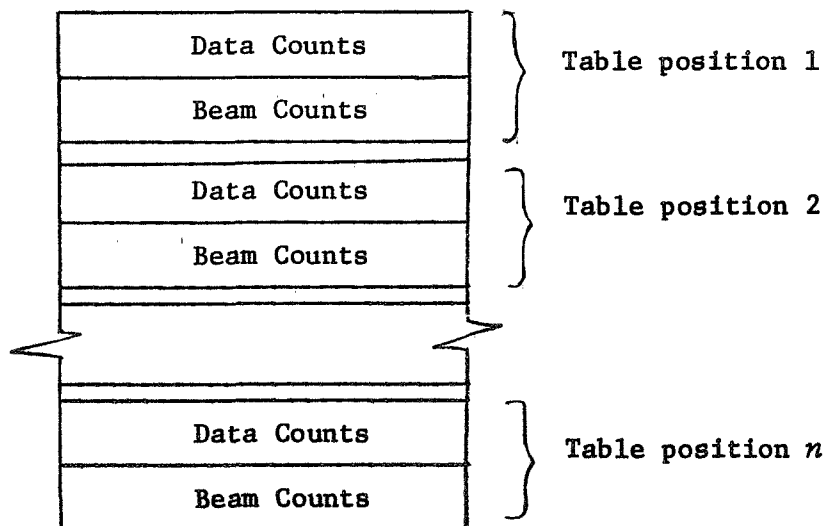
The X-ray Polarimeter Experiment Program acquires, analyzes and displays the primary data from the X-ray polarimeter and pulsar mode counter.

2.1 Input Parameters

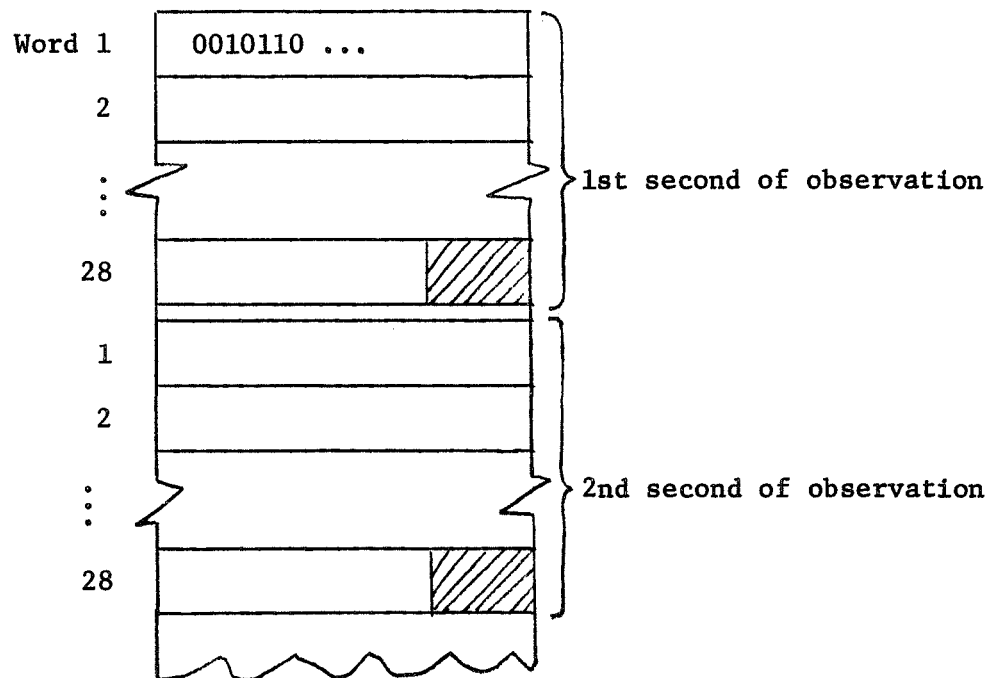
Experiment control parameters input by cards in NAMELIST format are:

NS - Number of Sources
NP - Plot Flag (1 = Plot all Levels of Intensity
 0 = Plot Cumulative Only)
ID - Source Identification
RAHR - Right Ascension (Hours)
RAMIN - Right Ascension (Minutes)
DDEG - Declination (Degrees)
DMIN - Declination (Minutes)
OBSTIM - Observation Time
BTP - Beginning Table Position (Degrees)
ETP - Ending Table Position (Degrees)
TSS - Table Step Size (Degrees)

Simulated polarimeter data is input from tape (channel A5) in the following format:



Simulated pulsar data is input from tape (channel A6) in bit string format as follows:



Pulsar data is in blocks of 1000 bits--where 1 bit represents 1 millisecond of observation. The data is packed so that the entire block is stored in 28 thirty-six bit words. The least significant 8 bits of the 28th word are not used.

2.2 Program Flow

Figure 2-1 presents the general flow of the X-ray polarimeter experiment program.

2.3 Subroutines

CAMRAV - Open SC-4020 file
INIT - Initialize experiment parameters
PULDAQ - Pulsar data acquisition
PULDAN - Pulsar data analysis
PULDDI - Pulsar data display
POLDAQ - Polarimeter data acquisition
POLDDP - Polarimeter data display
SUMUP - Experiment summary
CLEAN - Close SC-4020 file

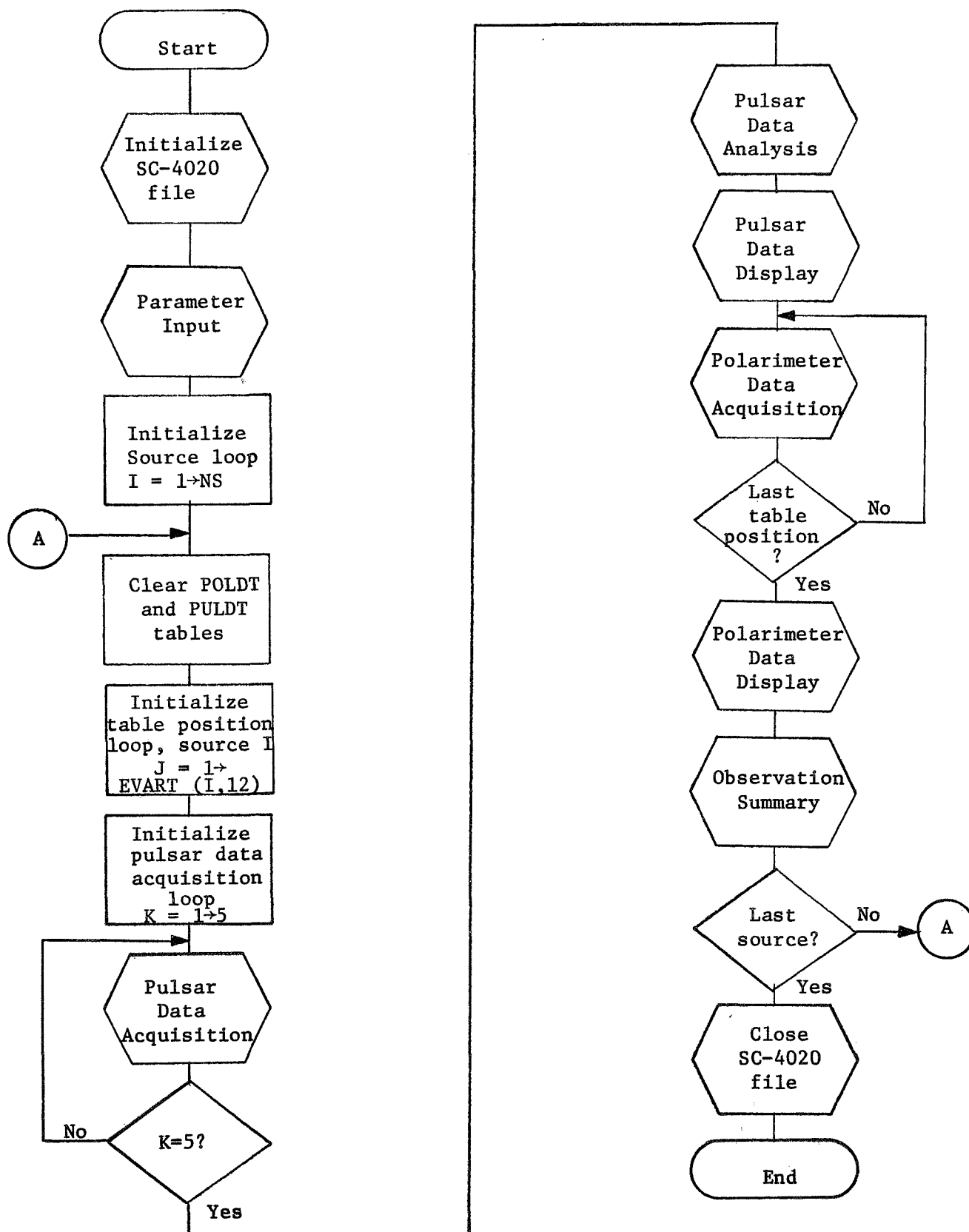


Figure 2-1. General Flow for Simulated Experiment

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CAMRAV and CLEAN are part of the SC-4020 QUIK package. All other subroutines were developed as part of this contract and are documented in this section.

Subroutine INIT

Purpose: Subroutine INIT inputs all experiment control variables and initializes the experiment variable table, EVART.

Usage: CALL INIT (NS, EVART, NP)

NS - Number of sources

EVART - Experiment variable table

NP - Plot flag

Input: Input to subroutine INIT is by card in NAMELIST format.

NS - Number of Sources

NP - Plot Flag (1 = Plot all Levels of Intensity
0 = Plot Cumulative Only)

ID - Source Identification

RAHR - Right Ascension (Hours)

RAMIN - Right Ascension (Minutes)

DDEG - Declination (Degrees)

DMIN - Declination (Minutes)

OBSTIM - Observation Time

BTP - Beginning Table Position (Degrees)

ETP - Ending Table Position (Degrees)

TSS - Table Step Size (Degrees)

All of the above input items are stored in table EVART except NS and NP.

Since NAMELIST does not provide for BCD input, the source identification is input in code format as follows:

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1 - Tau X-1	16 - Lup XR-1	31 - Oph XR-2
2 - Vir XR-1	17 - Nor XR-2	32 - Sco XR-3
3 - Cen XR-2	18 - Sco X-2	33 - GX 9+9
4 - Sco X-1	19 - Ara XR-1	34 - Oph XR-1
5 - Sco XR-2	20 - GX-5.6	35 - Sco XR-5
6 - GX 3+1	21 - Sgr XR-1	36 - Sco XR-6
7 - GX 5-1	22 - Ser XR-1	37 - Lyr XR-1
8 - GX 9+1	23 - Cep XR-1	38 - Sgr XR-5
9 - Sgr XR-2	24 - Leo XR-1	39 - Aql XR-1
10 - Ser XR-2	25 - Cen XR-3	40 - Cyg XR-3
11 - Cyg XR-1	26 - Cen XR-1	41 - Cyg X-3
12 - Cyg X-4	27 - Nor XR-1	42 - Vul XR-1
13 - Cyg XR-2	28 - Sco XR-4	43 - Lac XR-1
14 - Cas XR-1	29 - L7	44 - Cep XR-2
15 - Vel XR-1	30 - L8	45 - Cep XR-3

Storage: 370₍₈₎

Subroutine PULDAQ

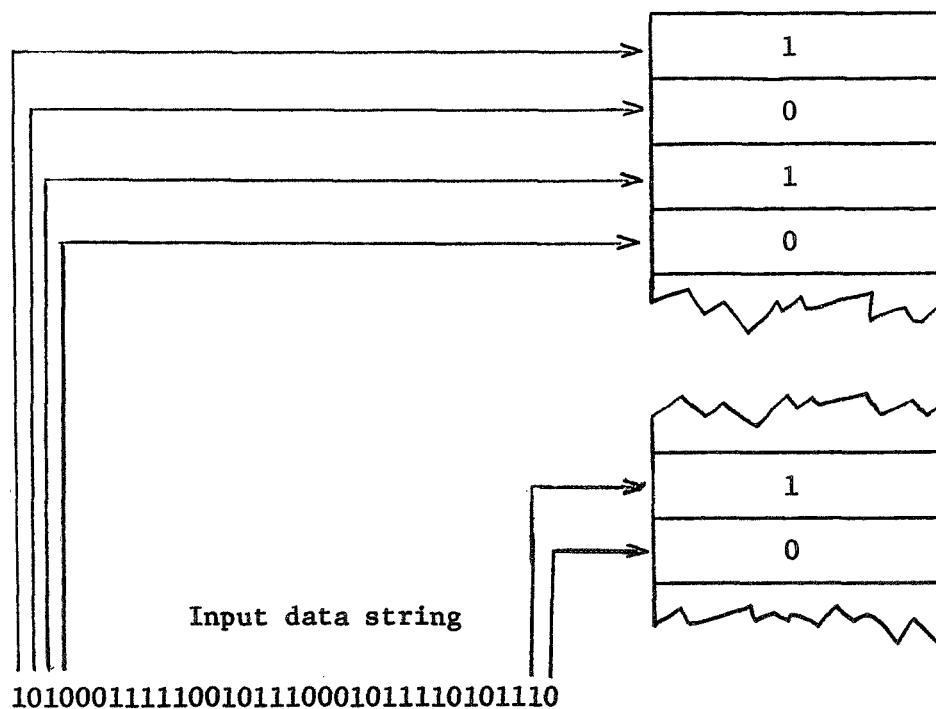
Purpose: Subroutine PULDAQ acquires pulsar data from tape and converts it from bit to word format.

Usage: CALL PULDAQ (I, PULDT)

I - Source index

PULDT - Pulsar data table

Method: Pulsar data is read from tape in string format where 1 bit represents 1 millisecond of observation. The data is unpacked to 1 bit per word to facilitate later analysis.



Subroutines: BIT (Logical bit testing routine)

Storage: 213₍₈₎

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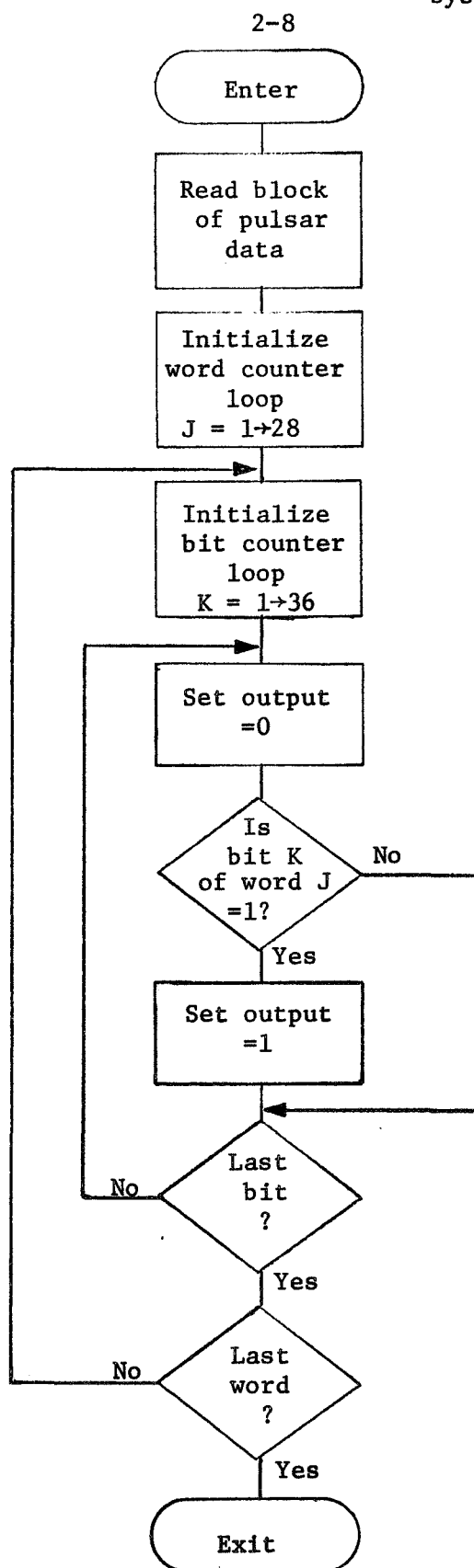


Figure 2-2. Subroutine PULDAQ

Subroutine PULDAN

Purpose: Subroutine PULDAN detects X-ray pulsations and determines the approximate pulsation period.

Usage: CALL PULDAN (I, PULDT, DUMMY, KFLAG, POWSP, TOP10, KTOP10)

I - Source index
 PULDT - Pulsar data table
 DUMMY - Duplicate of pulsar data table required by FFT
 KFLAG - Pulsar indicator (1 = pulsar, 0 = no pulsar)
 POWSP - Power spectrum of pulsar data
 TOP10 - Ten highest amplitudes of power spectrum
 KTOP10 - Index of 10 highest amplitudes

Method: The method used to detect pulsars and to estimate their fundamental frequency is an application of a Cooley-Tukey type Fast Fourier Transform (FFT) to obtain a power spectrum of the detected X-radiation. The pulse train of unity amplitude serves as the input to the FFT which produces cosine and sine amplitude coefficients:

$$A_r = \frac{2}{N} \sum_{k=0}^{N-1} X_k (\cos 2\pi rk/N) \quad r = 0, 1, \dots, N-1$$

$$B_r = \frac{2}{N} \sum_{k=0}^{N-1} X_k (\sin 2\pi rk/N) \quad r = 0, 1, \dots, N-1$$

where A_r and B_r are the r^{th} coefficients, X_k denotes the k^{th} sample of the time series of N samples.

From the cosine and sine amplitude coefficients, a set of power amplitudes is computed as

$$PA_r = \frac{(A_r + jB_r)(A_r - jB_r)}{T}$$

where T is the number of time increments in the sample and

$$j = \sqrt{-1}$$

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The spectrum of power amplitudes is sorted to isolate the frequencies at which the amplitudes are the highest. The 10 highest amplitudes are summed and if the resulting value is greater than 5% of the sum of all amplitudes, the source is considered a pulsar. The frequency at which the power amplitude is the greatest is the fundamental frequency of the X-ray source.

Subroutines: EVAL (Power spectrum density)

Storage: 225₍₈₎

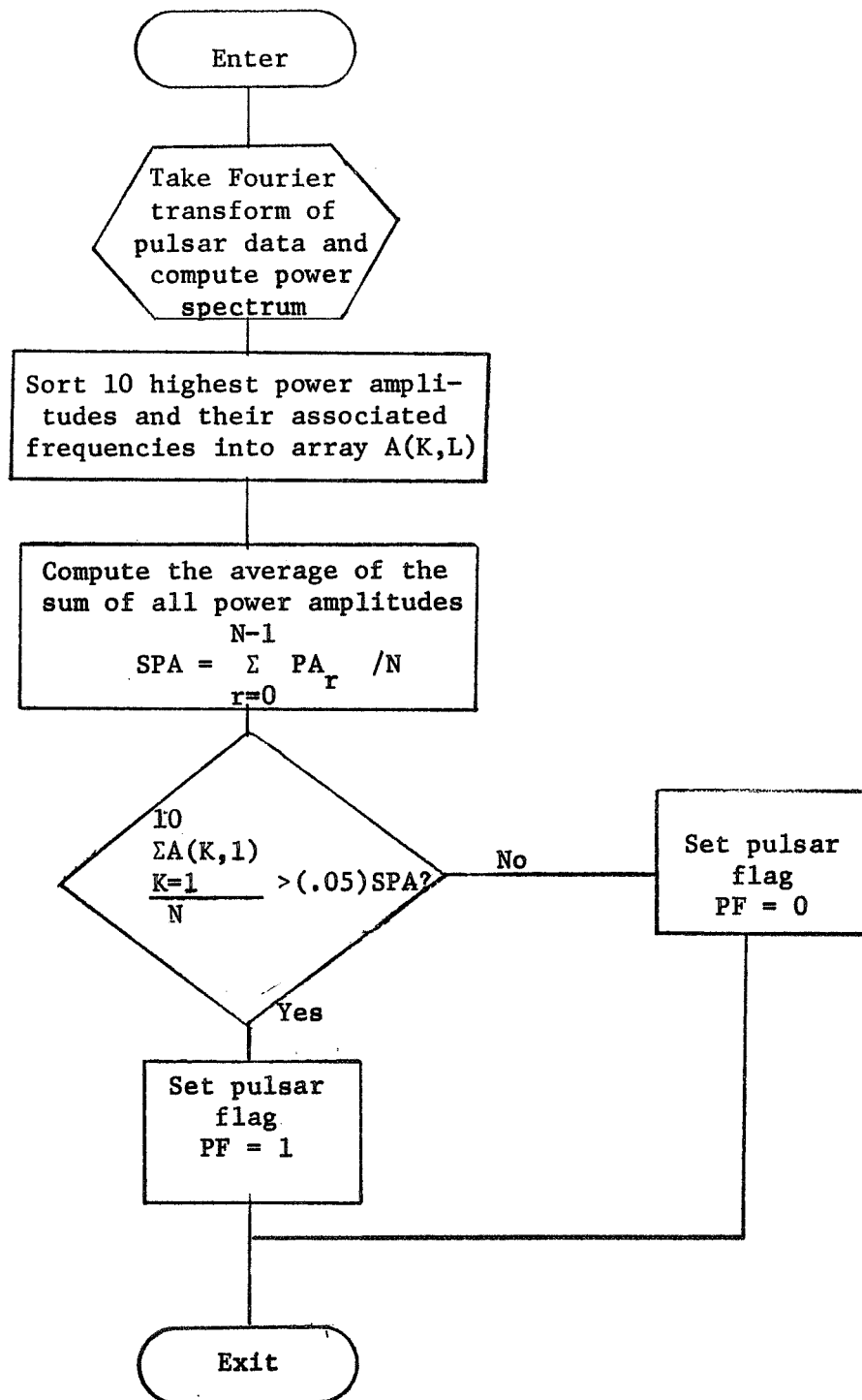


Figure 2-3. Subroutine PULDAN

Subroutine EVAL

Purpose: Subroutine EVAL calculates the power spectrum using the real and imaginary arrays.

Usage: CALL EVAL (X1, X2, N, SR, Y1)

X1 - Input data for the first data array.

X2 - Input data for the second data array.

N - The number of points in the X1 and X2 arrays where $N \leq 2^{12}$. If N is not a power of 2, the value of N will be set to a negative value on error return and the routine will not calculate the power spectrum.

SR - The sample rate for the data.

Y1 - The cross power spectrum array containing N spectral estimates.

Subroutines: HARM (Subroutine HARM is an IBM Share routine which has been converted to the IBM-7094.)

LSF (The subroutine LSF is a bit manipulation routine with entry points IRS and INVERT.)

Restrictions:

1. Subroutine EVAL removes the mean from the X1 and X2 data arrays before the cross power spectrum is calculated.
2. After a successful call to EVAL, the data is destroyed in the X1 and X2 data arrays.
3. A flag can be set in the routine to aid in program checkout by printing the power spectrum for each data array.
4. If the contents of X1 and X2 data arrays contain the same data, the Y1 array will be a single power spectrum.

Storage: 501(8)

Subroutine LSF

Purpose: Subroutine LSF moves a string of bits left, right, or produces a mirror image for N bits where $N \leq 32$.

Usage: CALL LSF (IWORD, NBITS)

IWORD - The routine will move a string of bits in IWORD to the left.

NBITS - The number of places to move the string of bits in IWORD.

Entry Point II

CALL IRS (IWORD, NBITS)

IWORD - The entry IRS will move a string of bits in IWORD to the right.

NBITS - The number of places to move the string of bits in IWORD.

Entry Point III

CALL INVERT (IWORD, NBITS)

IWORD - The entry will invert a string of bits to produce a mirror image.

NBITS - The number of bits in IWORD to invert.

Storage: 216₍₈₎

Subroutine PULDDI

Purpose: Subroutine PULDDI presents the analyzed pulsar data in graphic form.

Usage: CALL PULDDI (I, EVART, POWSP, TOP10, KTOP10, X, Y, DOMFRQ)

I - Source index
EVART - Experiment variable table
POWSP - Power spectrum of pulsar data
TOP10 - Amplitude of 10 highest frequencies
KTOP10 - Index of 10 highest amplitudes
X - Working storage array
Y - Working storage array
DOMFRQ - Dominant frequency

Output: The SC-4020 plotter is used as the output device for plots of the power spectrum of the pulsar data. See Figure 2-7.

Subroutines: CTOBCD (Convert to BCD)
 QUIK3V (SC-4020 plot routine)

Storage: 411₍₈₎

Subroutine POLDAQ

Purpose: Subroutine POLDAQ acquires from tape, edits and maintains cumulative totals of the polarimeter data.

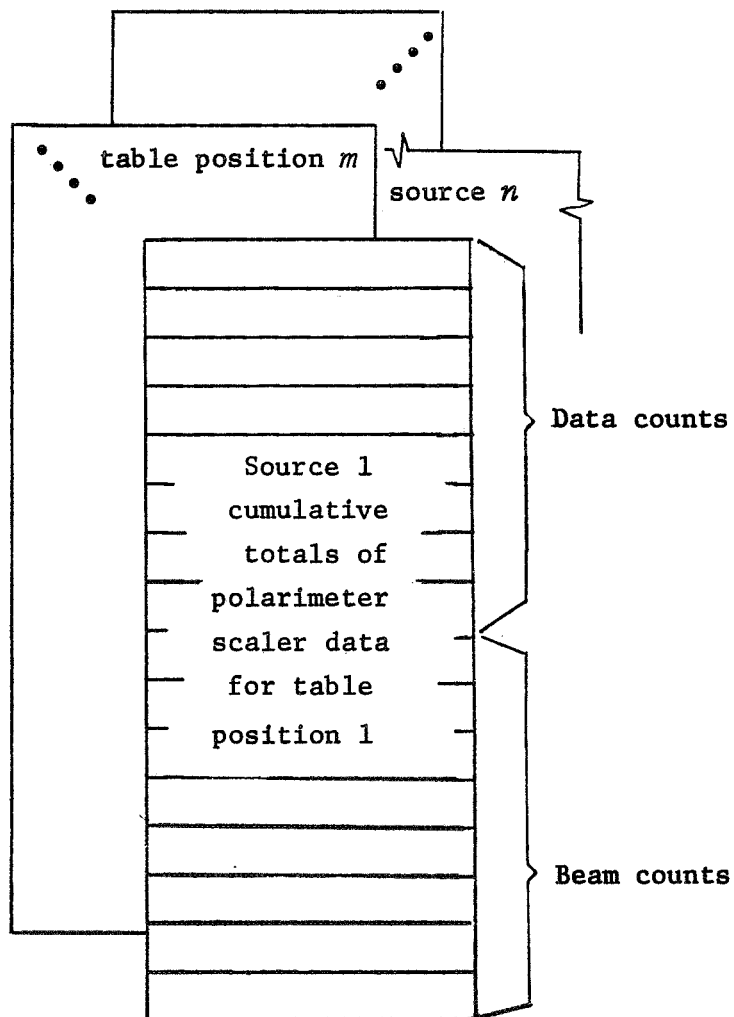
Usage: CALL POLDAQ (I, J, POLDT)

I - Source index

J - Polarimeter table position index

POLDT - Polarimeter data table

Method: Polarimeter data is read from tape, extraneous bits stripped off and the resulting values stored in table POLDT as follows:



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Subroutines: BIN (AND operator)

Storage: 174₍₈₎

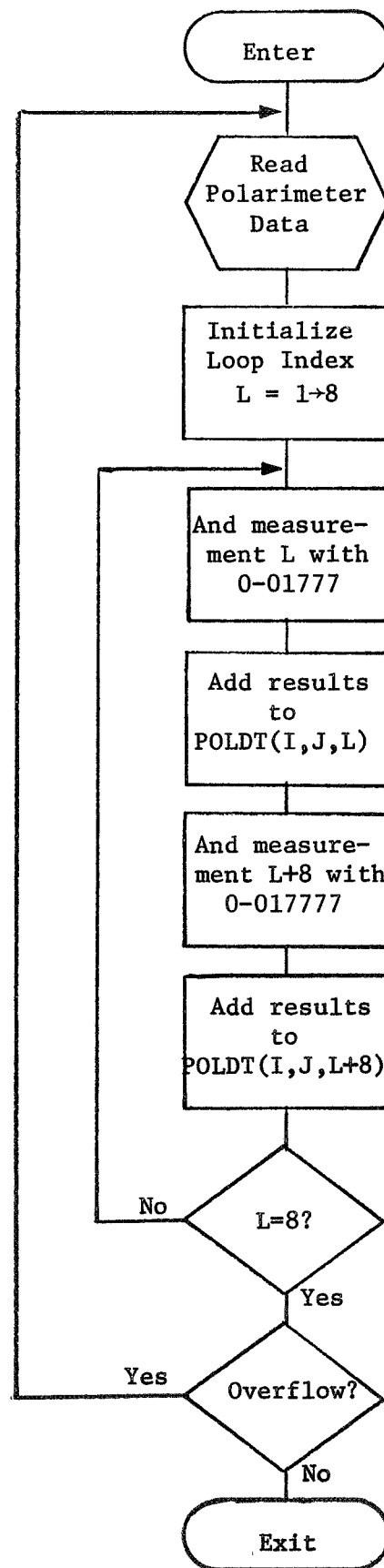


Figure 2-4. Subroutine POLDAQ

Subroutine POLDDP

Purpose: Subroutine POLDDP formats and presents the polarimeter data in graphic form.

Usage: CALL POLDDP (I, POLDT, EVART, NP, DEGPOL, PM)

I - Source index
POLDT - Polarimeter data table
EVART - Experiment variable table
NP - Plot flag
DEGPOL - Degree of polarization
PM - Angle of maximum polarization

Method: Degree of polarization is calculated as

$$D = \frac{\text{max} - \text{min}}{\text{max} + \text{min}} \times 100.$$

Angle of maximum polarization is determined by interpolation.

Output: The SC-4020 is used as the output device. See Figure 2-8.

Subroutines: CTOBCD (Convert to BCD)
LAGRNG (Lagrangian interpolation)
QUIK3V (SC-4020 plot routine)

Storage: 1075₍₈₎

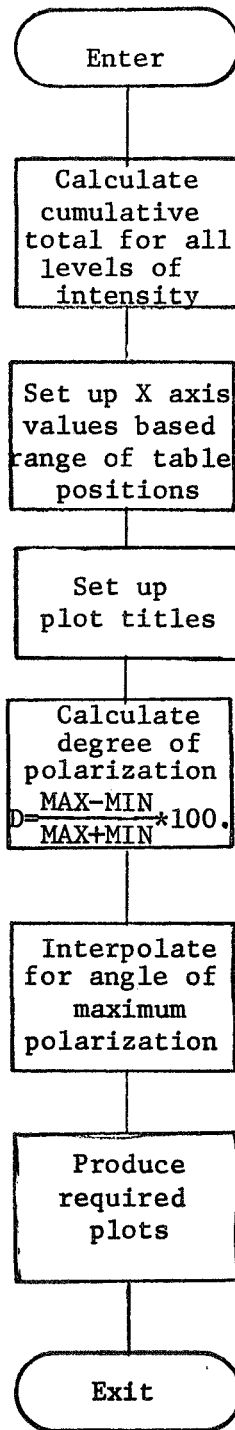


Figure 2-5. Subroutine POLDDP

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Subroutine SUMUP

Purpose: Subroutine SUMUP outputs a summary of the total observation of each source.

Usage: CALL SUMUP (I, EVART, POLDT, TOP10, KTOP10, DEGPOL, PM, KFLAG, DOMFRQ)

I - Source index
EVART - Experiment variable table
POLDT - Polarimeter data table
TOP10 - Ten highest amplitudes of power spectrum
KTOP10 - Index of ten highest amplitudes
DEGPOL - Degree of polarization
PM - Angle of maximum polarization
KFLAG - Pulsar indicator
DOMFRQ - Dominant frequency

Output: See Figures 2-20 through 2-23.

Storage: 624₍₈₎

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2.4 Sample Input

Card input for 4 sample X-ray sources is presented in Figure .

\$DATA		
\$NAMI		
NS = 4	}	Number of Sources and Plot Flag
NP = 1\$		
\$NAM2		
ID = 2	}	Source 1
RAHR = 12.		
RAMIN = 30.7		
DDEG = 12.		
DMIN = 30.		
OBSTIM = 20.		
BTP = 0.		
ETP = 350.		
TSS = 10.\$		
\$NAM2		
ID = 18	}	Source 2
RAHR = 16.		
RAMIN = 53.		
DDEG = -40.		
DMIN = 0.		
OBSTIM = 10.		
TSS = 10.\$		
\$NAM2		
ID = 21	}	Source 3
RAHR = 17.		
RAMIN = 48.		
DDEG = -30.		
DMIN = 0.		
OBSTIM = 20.		
TSS = 10.\$		
\$NAM2		
ID = 13	}	Source 4
RAHR = 21.		
RAMIN = 42.6		
DDEG = 38.		
DMIN = 5.		
OBSTIM = 15.		
BTP = 0.		
ETP = 175.		
TSS = 5.\$		
\$END		

Figure 2-6. Sample Experiment Program Input

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2.5 Sample Output

Output for 4 sample X-ray sources is presented in Figures 2-7 through 2-19.

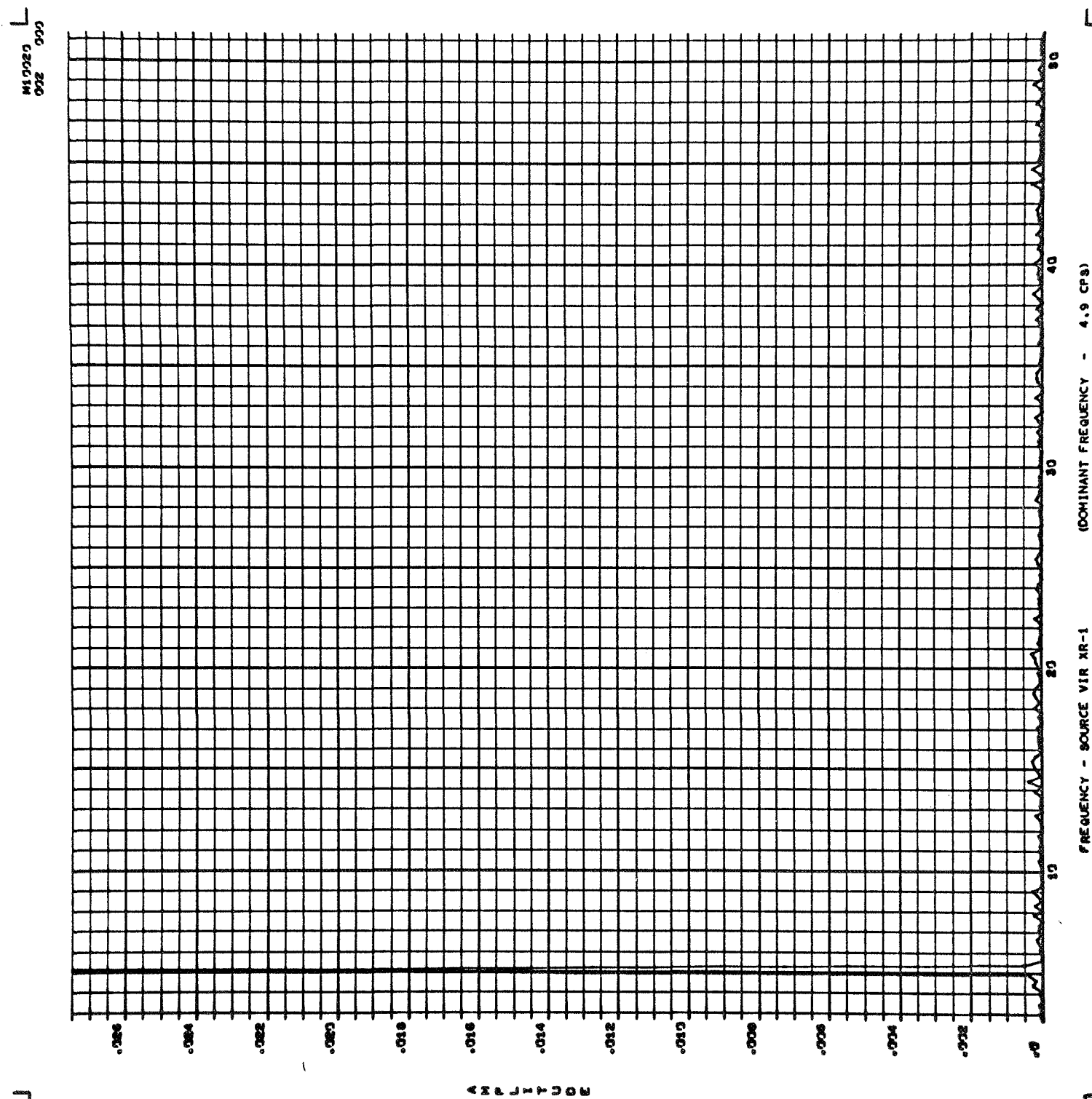


Figure 2-7. Power Spectrum of Pulsar Data - Source 1

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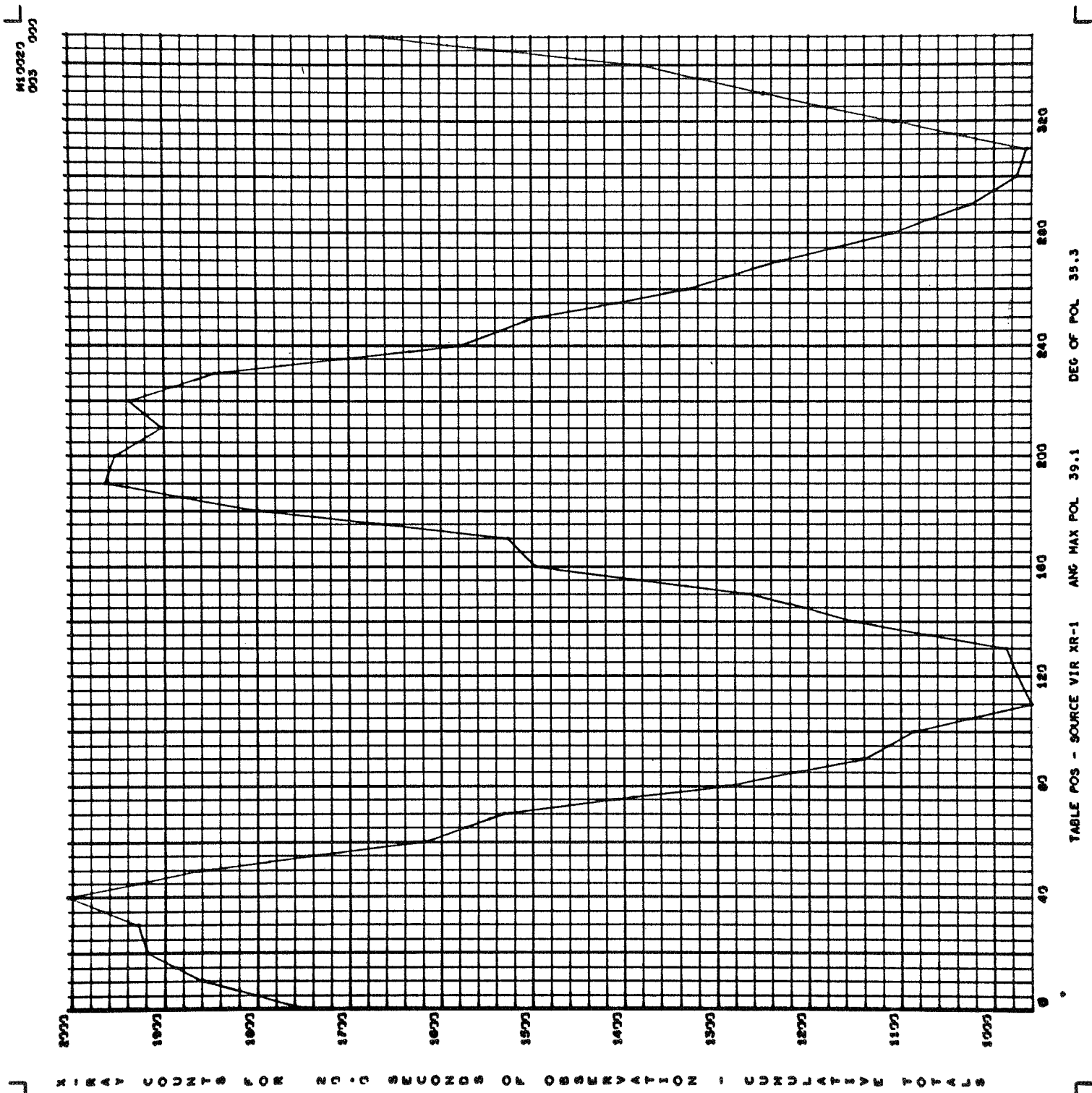


Figure 2-8. Cumulative Totals of Polarimeter Data Counts - Source 1

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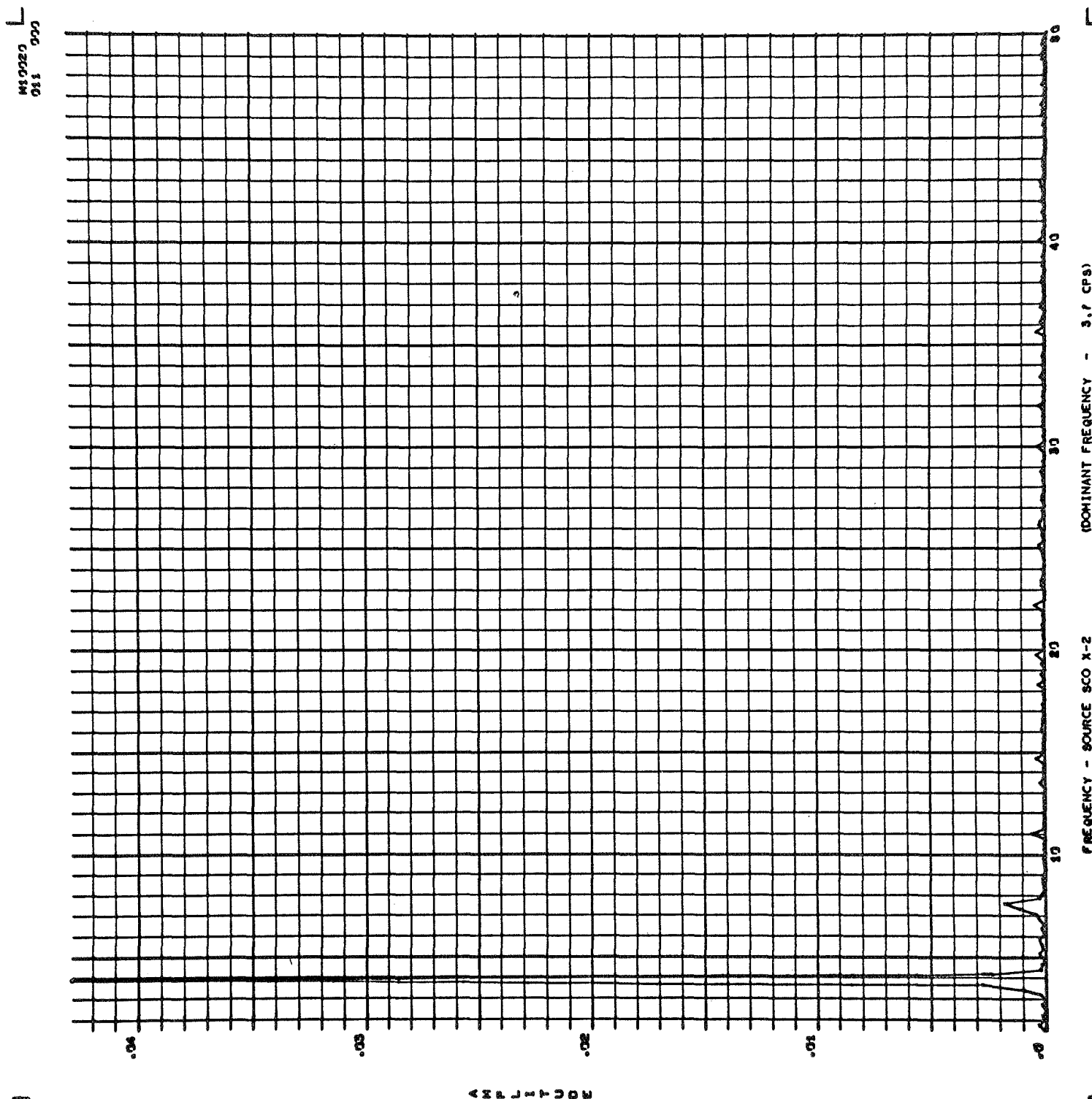


Figure 2-9. Power Spectrum of Pulsar Data - Source 2

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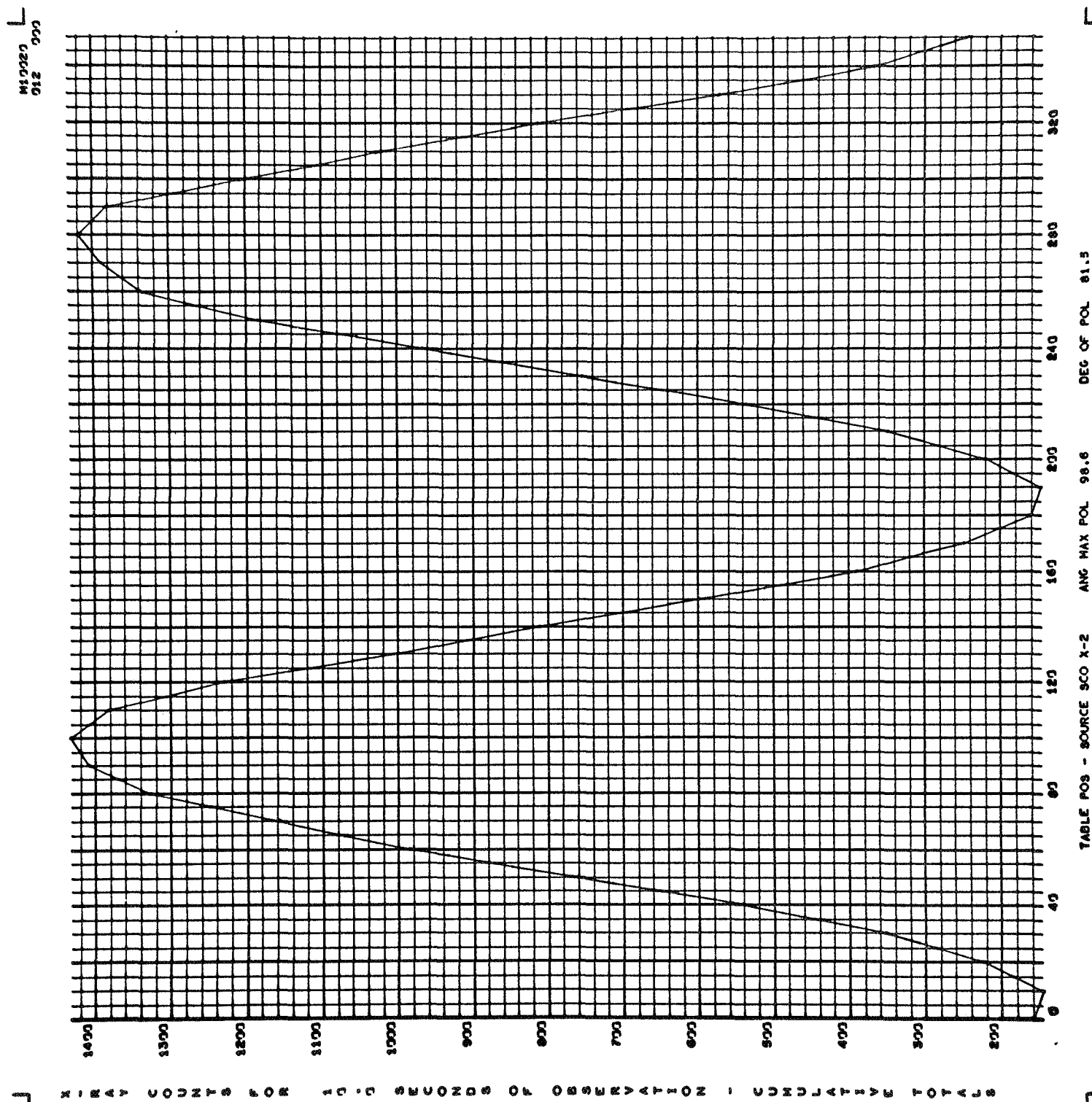


Figure 2-10. Cumulative Totals of Polarimeter Data Counts - Source 2

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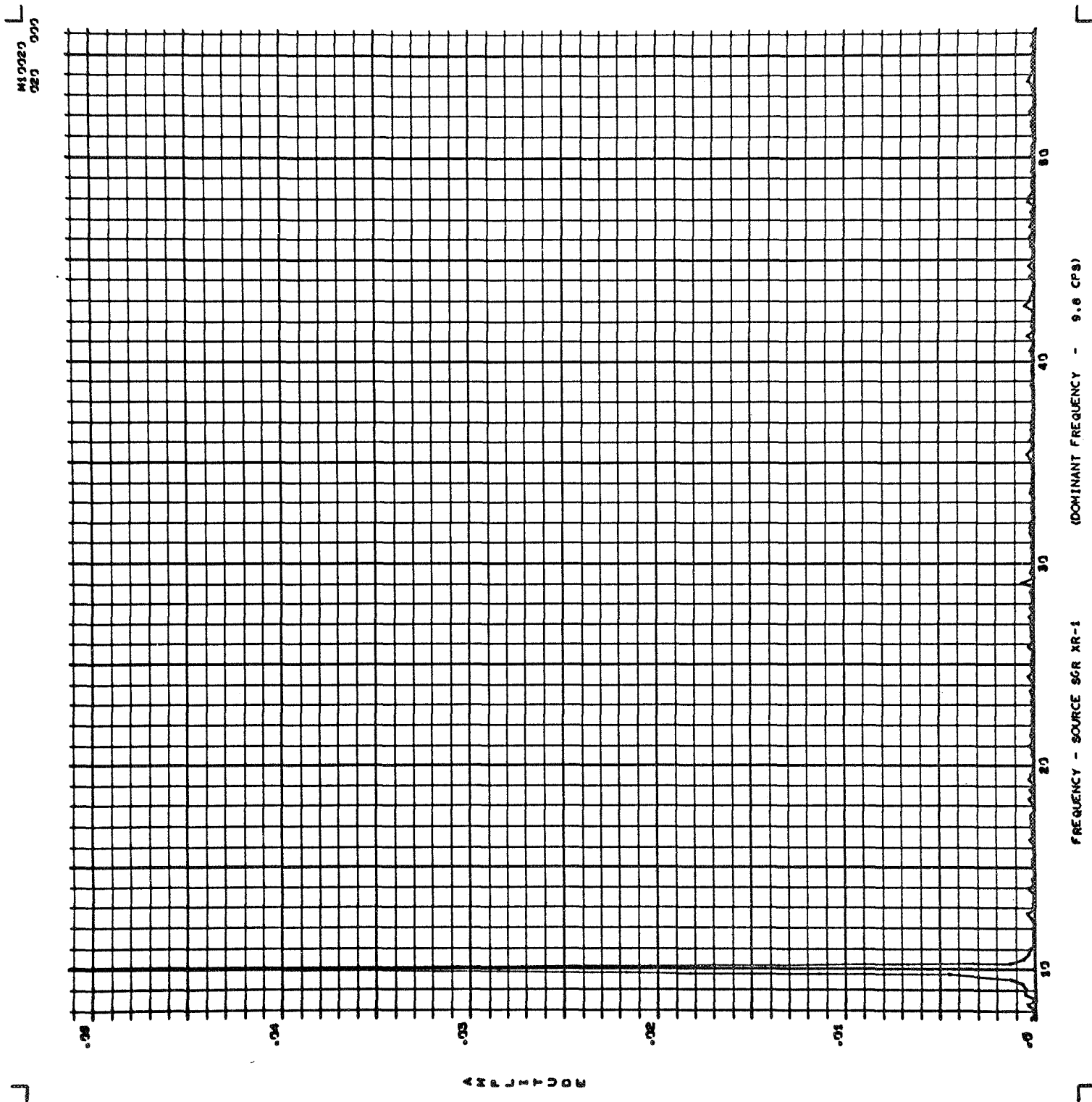


Figure 2-11. Power Spectrum of Pulsar Data - Source 3

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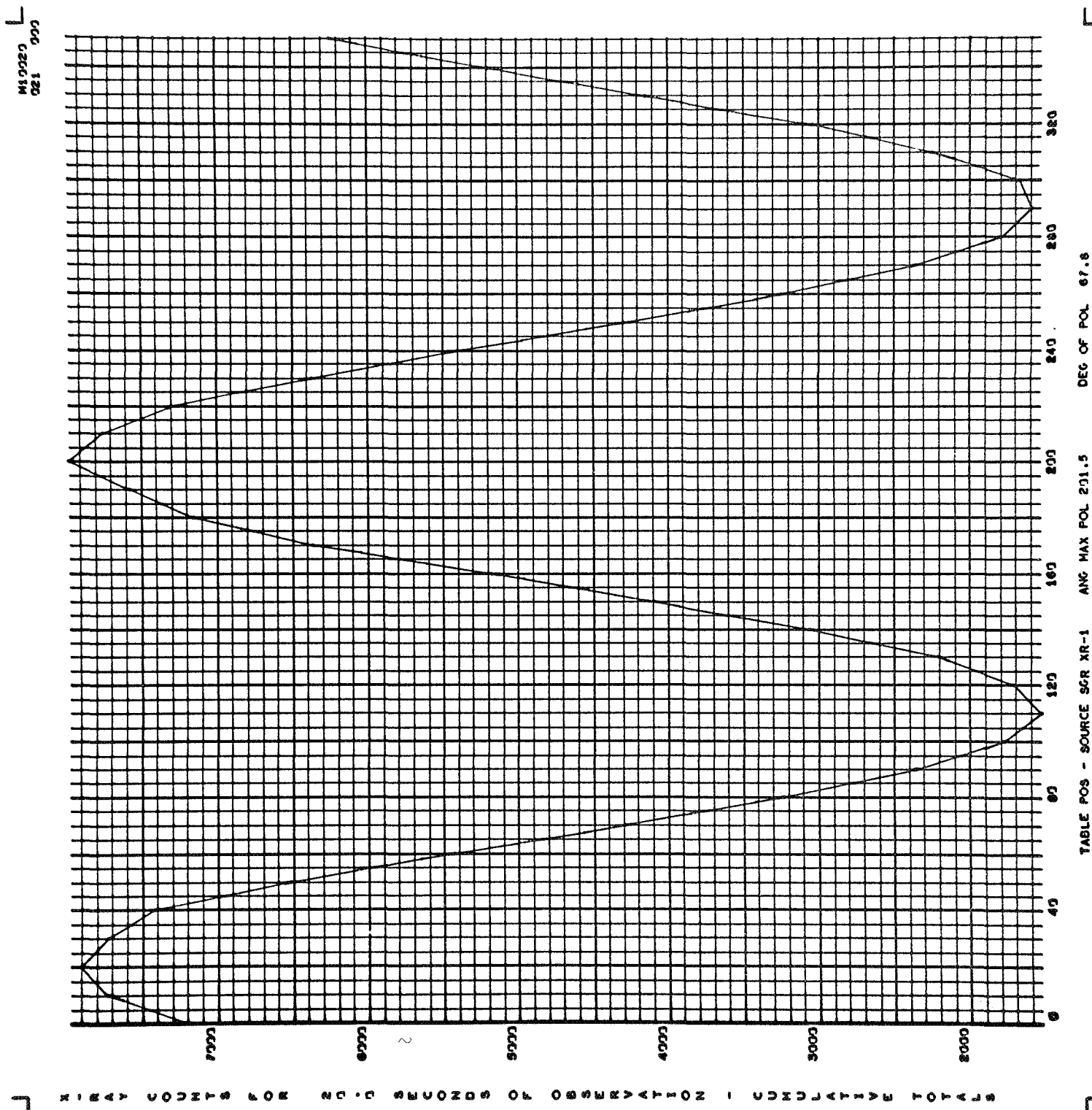


Figure 2-12. Cumulative Totals of Polarimeter Data Counts - Source 3

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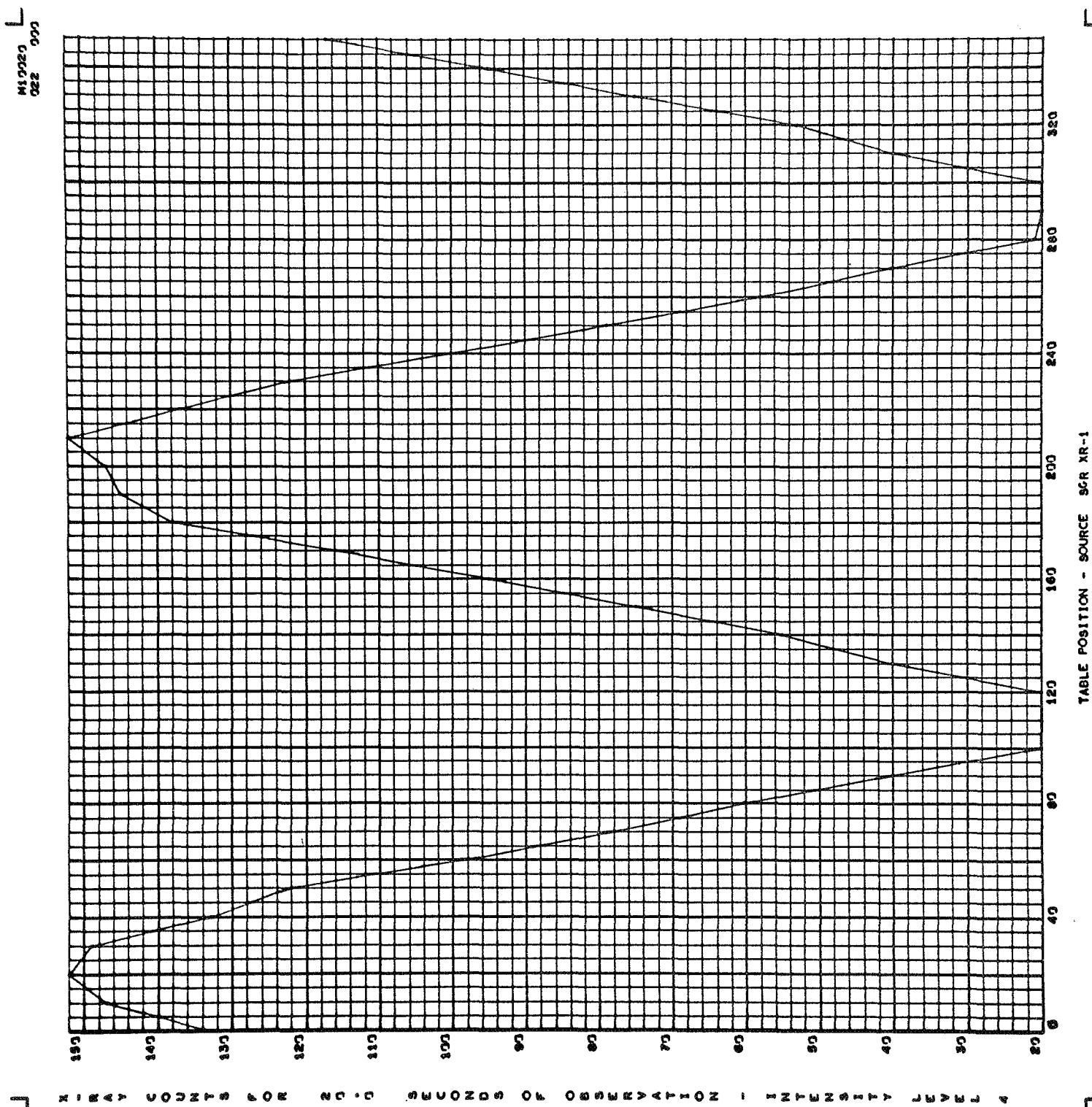


Figure 2-13. Intensity Level 4, Polarimeter Data Counts - Source 3

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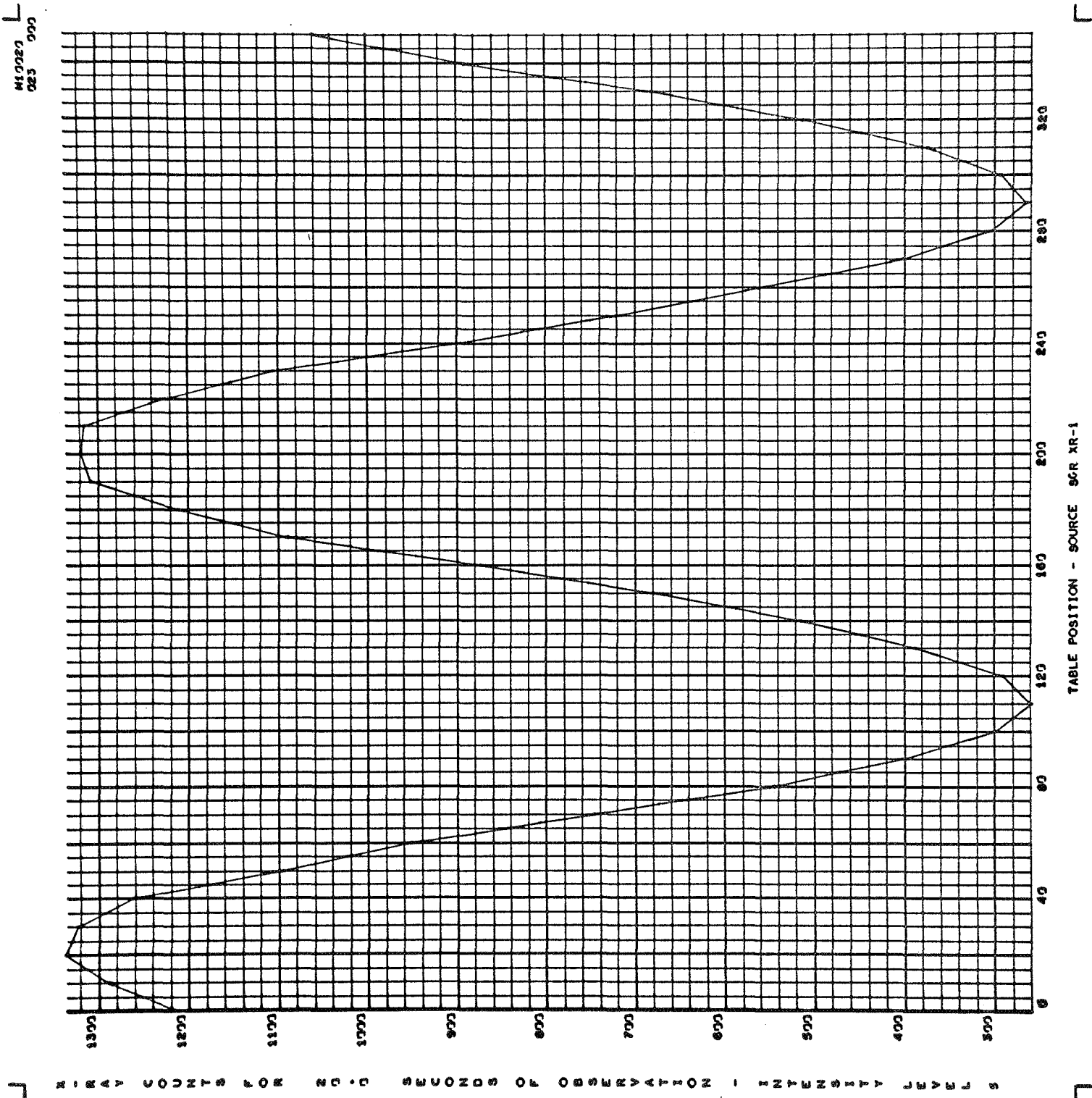


Figure 2-14. Intensity Level 5, Polarimeter Data Counts - Source 3

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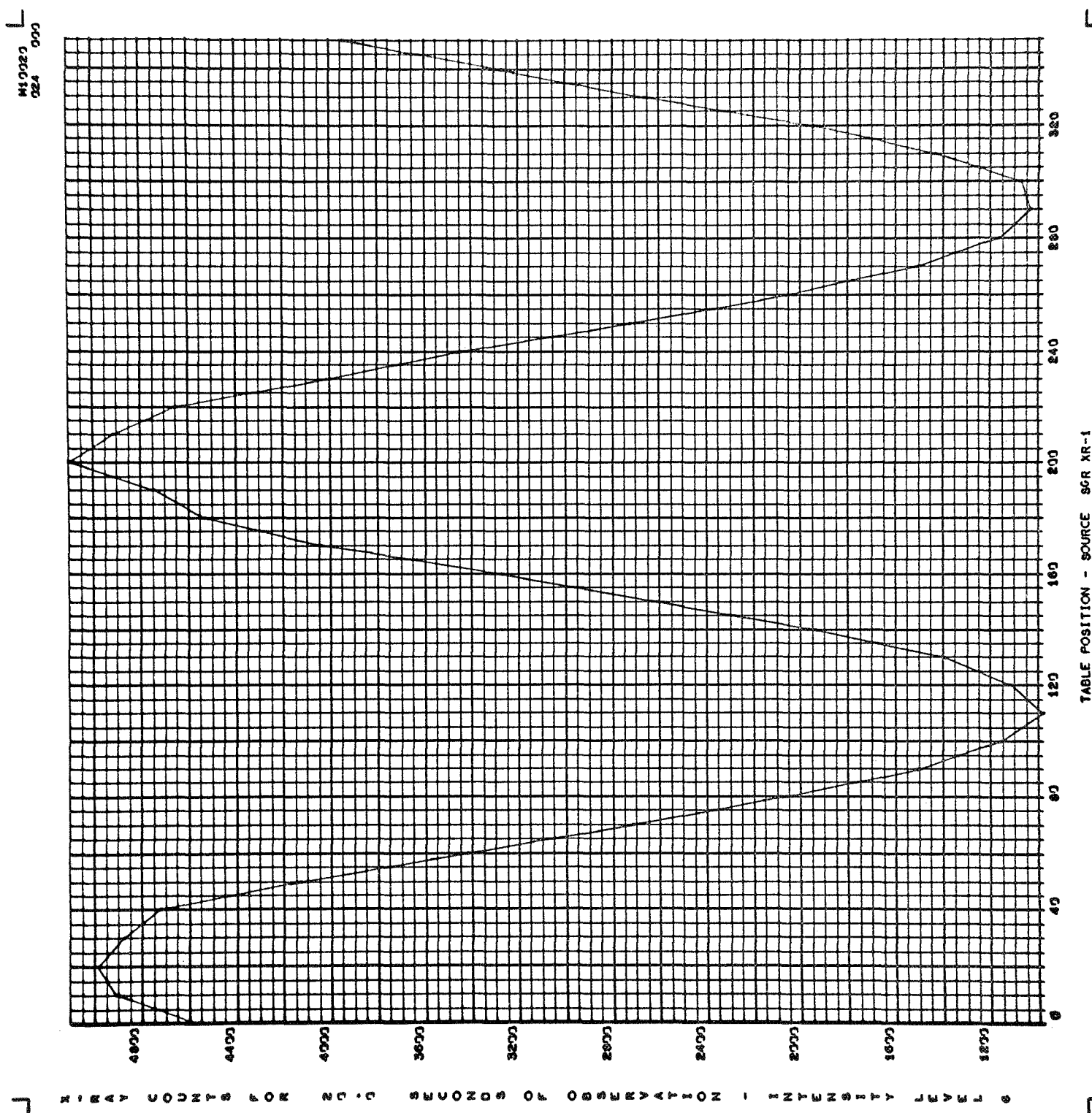


Figure 2-15. Intensity Level 6, Polarimeter Data Counts - Source 3

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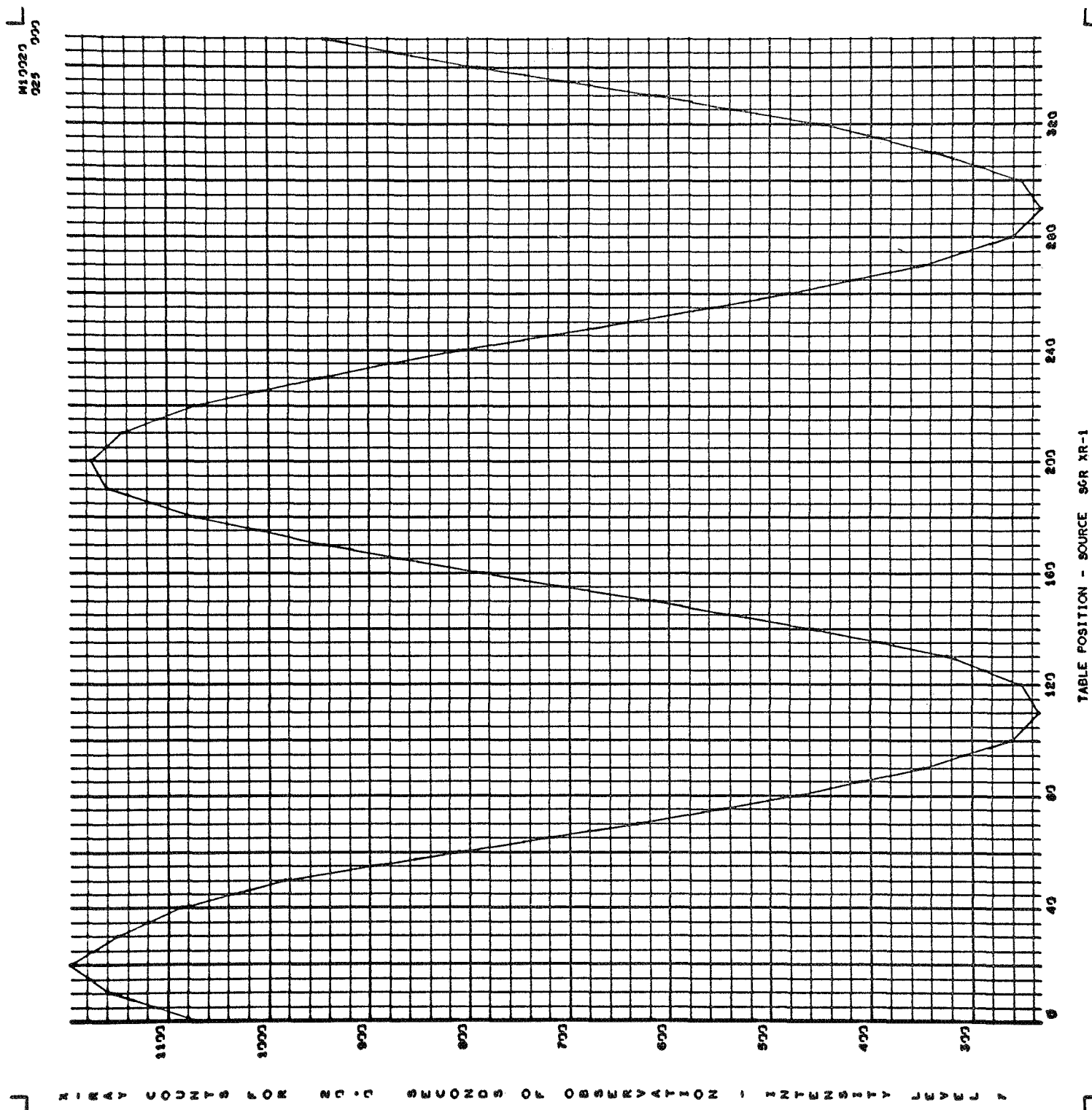


Figure 2-16. Intensity Level 7, Polarimeter Data Counts - Source 3

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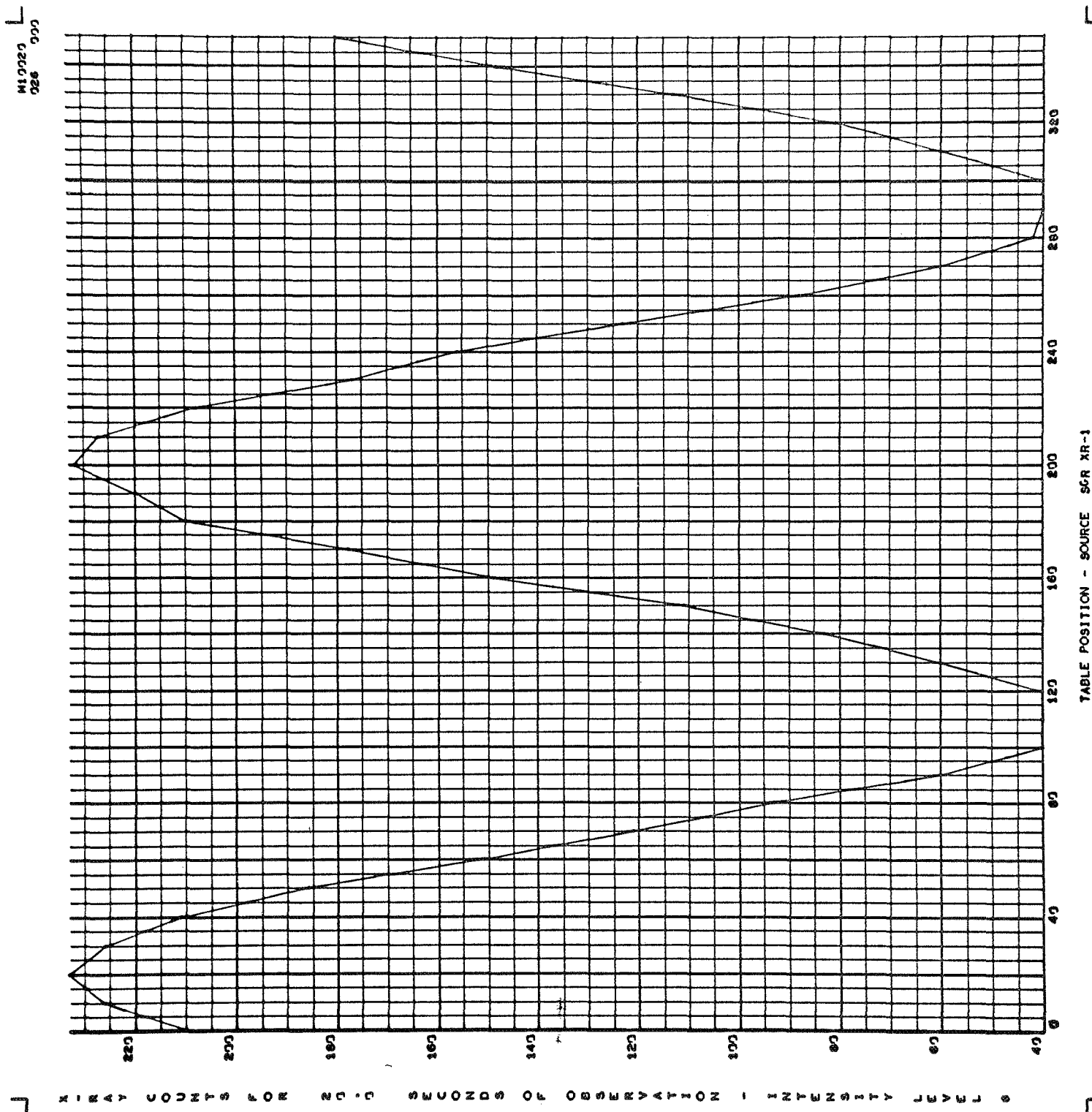


Figure 2-17. Intensity Level 8, Polarimeter Data Counts - Source 3

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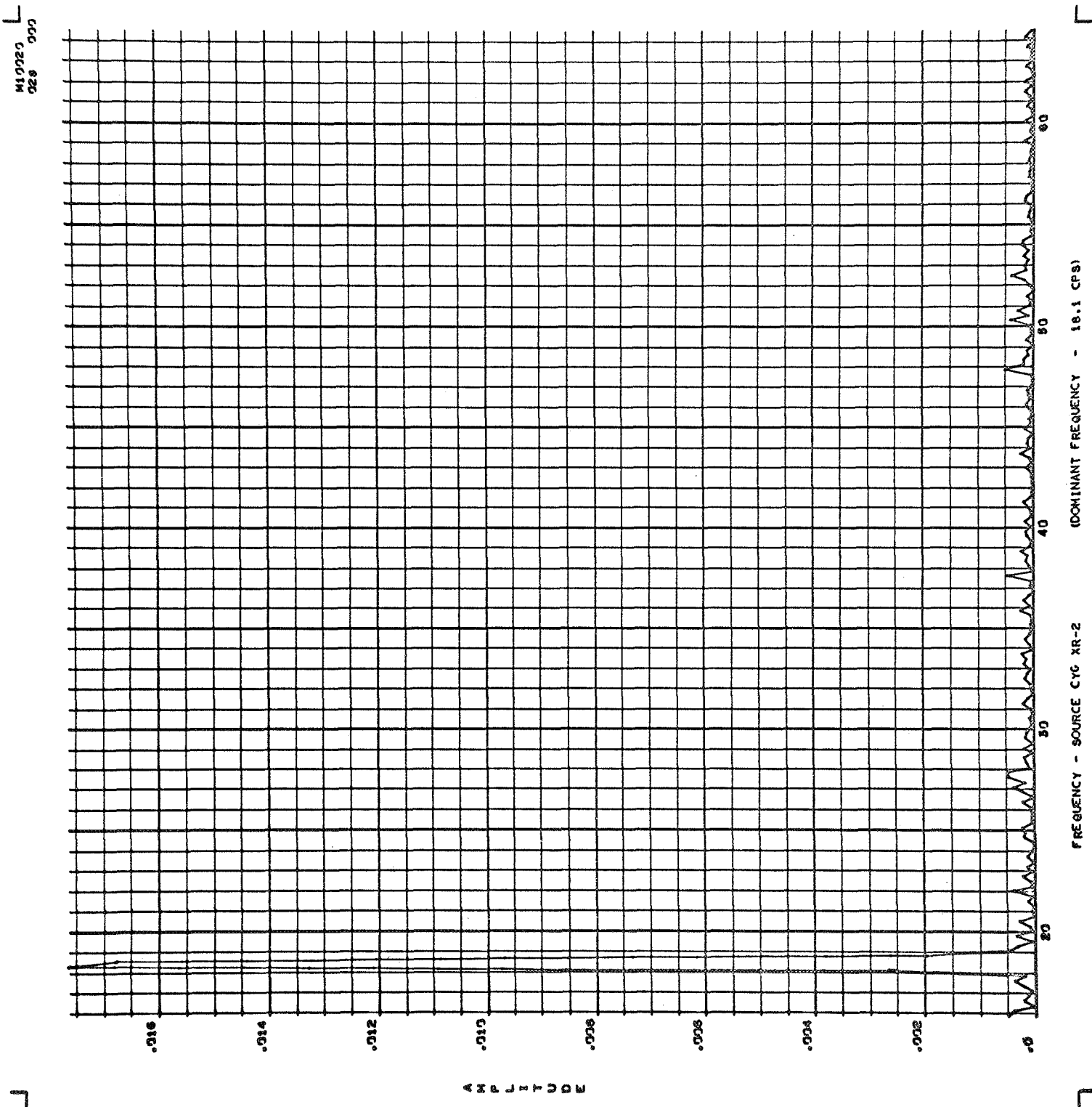


Figure 2-18. Power Spectrum of Pulsar Data - Source 4

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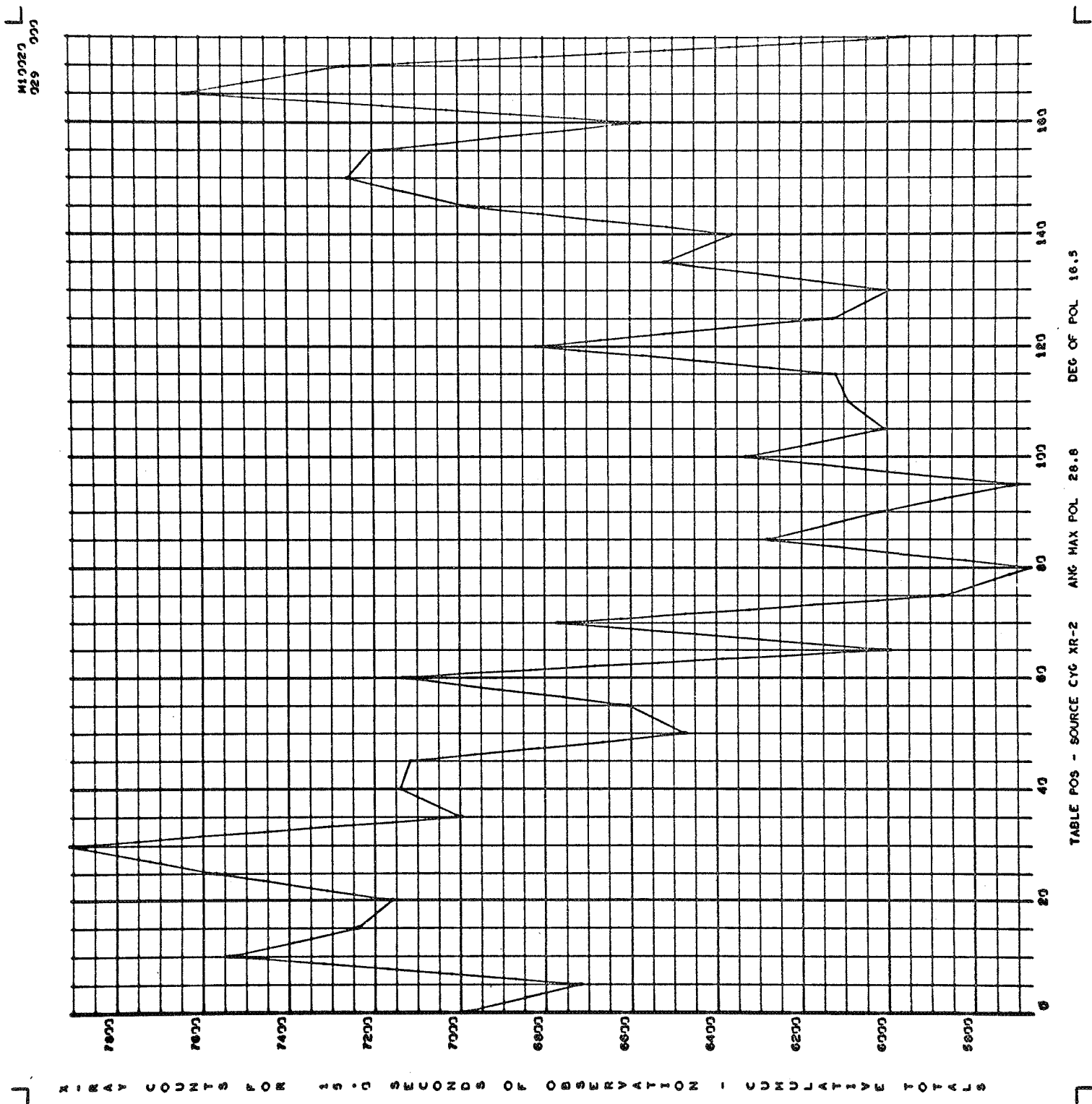


Figure 2-19. Cumulative Totals of Polarimeter Data Counts - Source 4

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SUMMARY OF OBSERVATION

SOURCE - VIR XR-1

LOCATION - RIGHT ASCENSION 12. HOURS, 31. MINUTES

DECLINATION 12. DEGREES, 30. MINUTES

OBSERVATION TIME - 20. SECONDS AT EACH TABLE POSITION

X-RAY POLARIZATION

BEGINNING TABLE POSITION 0. DEGREES

ENDING TABLE POSITION 350. DEGREES

TABLE STEP SIZE 10. DEGREES

ANGLE OF MAXIMUM POLARIZATION - 39.1 DEGREES

DEGREE OF POLARIZATION - 35.3 PERCENT

X-RAY INTENSITY

LEVEL	DATA COUNTS		BEAM COUNTS	
	NUMBER	PERCENT	NUMBER	PERCENT
1	5215	9.9	3460	9.9
2	8999	17.0	6080	17.4
3	15399	29.1	10155	29.1
4	21981	41.6	14531	41.6
5	895	1.7	544	1.6
6	374	0.7	165	0.5
7	0	0.0	0	0.0
8	0	0.0	0	0.0

X-RAY PULSATION

THIS SOURCE APPEARS TO BE A PULSAR WITH A
FREQUENCY OF 4.86 CYCLES PER SECOND

Figure 2-20. Observation Summary - Source 1

SUMMARY OF OBSERVATION

SOURCE - SCO X-2

LOCATION - RIGHT ASCENSION 16. HOURS, 53. MINUTES

DECLINATION -40. DEGREES, 0. MINUTES

OBSERVATION TIME - 10. SECONDS AT EACH TABLE POSITION

X-RAY POLARIZATION

BEGINNING TABLE POSITION 0. DEGREES

ENDING TABLE POSITION 350. DEGREES

TABLE STEP SIZE 10. DEGREES

ANGLE OF MAXIMUM POLARIZATION - 98.6 DEGREES

DEGREE OF POLARIZATION - 81.5 PERCENT

X-RAY INTENSITY

LEVEL	DATA COUNTS		BEAM COUNTS	
	NUMBER	PERCENT	NUMBER	PERCENT
1	0	0.0	24	0.1
2	1778	6.3	2513	6.5
3	11064	39.2	14851	38.3
4	7222	25.6	10030	25.9
5	5166	18.3	7131	18.4
6	2685	9.5	3814	9.8
7	302	1.1	432	1.1
8	0	0.0	0	0.0

X-RAY PULSATION

THIS SOURCE APPEARS TO BE A PULSAR WITH A
FREQUENCY OF 3.66 CYCLES PER SECOND

Figure 2-21. Observation Summary - Source 2

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SUMMARY OF OBSERVATION

SOURCE - SGR XR-1

LOCATION - RIGHT ASCENSION 17. HOURS, 48. MINUTES

DECLINATION -30. DEGREES, 0. MINUTES

OBSERVATION TIME - 20. SECONDS AT EACH TABLE POSITION

X-RAY POLARIZATION

BEGINNING TABLE POSITION 0. DEGREES

ENDING TABLE POSITION 350. DEGREES

TABLE STEP SIZE 10. DEGREES

ANGLE OF MAXIMUM POLARIZATION - 201.5 DEGREES

DEGREE OF POLARIZATION - 67.8 PERCENT

X-RAY INTENSITY

LEVEL	DATA COUNTS		BEAM COUNTS	
	NUMBER	PERCENT	NUMBER	PERCENT
1	0	0.0	0	0.0
2	0	0.0	0	0.0
3	0	0.0	0	0.0
4	3112	1.8	2641	1.8
5	29012	17.0	25286	16.9
6	108626	63.5	95273	63.7
7	25519	14.9	22275	14.9
8	4822	2.8	4145	2.8

X-RAY PULSATION

THIS SOURCE APPEARS TO BE A PULSAR WITH A
FREQUENCY OF 9.77 CYCLES PER SECOND

Figure 2-22. Observation Summary - Source 3

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SUMMARY OF OBSERVATION

SOURCE - CYG XR-2

LOCATION - RIGHT ASCENSION 21. HOURS, 43. MINUTES

DECLINATION 38. DEGREES, 5. MINUTES

OBSERVATION TIME - 15. SECONDS AT EACH TABLE POSITION

X-RAY POLARIZATION

BEGINNING TABLE POSITION 0. DEGREES

ENDING TABLE POSITION 175. DEGREES

TABLE STEP SIZE 5. DEGREES

ANGLE OF MAXIMUM POLARIZATION - 28.8 DEGREES

DEGREE OF POLARIZATION - 16.5 PERCENT

X-RAY INTENSITY

LEVEL	DATA COUNTS		BEAM COUNTS	
	NUMBER	PERCENT	NUMBER	PERCENT
1	208854	87.0	47137	87.8
2	29123	12.1	6288	11.7
3	2182	0.9	245	0.5
4	0	0.0	0	0.0
5	0	0.0	0	0.0
6	0	0.0	0	0.0
7	0	0.0	0	0.0
8	0	0.0	0	0.0

X-RAY PULSATION

THIS SOURCE APPEARS TO BE A PULSAR WITH A
FREQUENCY OF 18.07 CYCLES PER SECOND

Figure 2-23. Observation Summary - Source 4

2.6 Program Listing

```
C      SIMULATED X-RAY POLARIMETER EXPERIMENT
C      KEN BAUER - SYSTEM DEVELOPMENT CORPORATION, HUNTSVILLE, ALABAMA
C      PHONE 539-7711
C
C      THIS PROGRAM SIMULATES THE PROCESSING OF PRIME DATA FROM THE
C      X-RAY POLARIMETER ASSOCIATED WITH A PROPOSED SPACE STATION
C      EXPERIMENT - FPE 5.1.
C
C      DIMENSION EVART(4,12)
C      DIMENSION POLDT(4,36,16)
C      DIMENSION PULDT(4097)
C      DIMENSION DUMMY(4097)
C      DIMENSION POWSP(4097)
C      DIMENSION TOP10(10)
C      DIMENSION KTOP10(10)
C
C      INTEGER POLDT
C
C.....INITIALIZE SC 4020
C      CALL CAMRAV(9)
C
C.....INPUT EXPERIMENT CONTROL PARAMETERS
C      CALL INIT( NS,EVART,NP)
C
C.....INITIALIZE SOURCE LOOP
C      DO 300 I=1,NS
C
C.....CLEAR POLDT AND PULDT TABLES
C      DO 30 K=1,4
C      DO 30 L=1,36
C      DO 30 M=1,16
C      30 POLDT(K,L,M) = 0
C      DO 40 K=1,4096
C      40 PULDT(K) = 0.
C
C.....INITIALIZE TABLE POSITION LOOP
C      NTP = EVART(I,12)
C      J = 1
C
C.....INITIALIZE PULSAR DATA ACQUISITION LOOP AND ACQUIRE DATA
C      DO 100 K=1,5
C      100 CALL PULDAQ (K,PULDT)
C
C.....PULSAR DATA ANALYSIS
C      CALL PULDAN (I,PULDT,DUMMY,KFLAG,POWSP,TOPI0,KTOP10)
C
C.....IF NO PULSAR, DO NOT PLOT
C      IF (KFLAG.EQ.0) GO TO 200
C
C.....PULSAR DATA DISPLAY
C      CALL PULDDI (I,EVART,POWSP,TOPI0,KTOP10,PULDT,DUMMY,DOMFRQ)
C
C.....ACQUIRE POLARIMETER DATA FOR SOURCE I, TABLE POSITION J
C      200 CALL POLDAN (I,J,PULDT)
C
```

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C.....INCREMENT TABLE POSITION

J = J+1

IF (J.LE.NTP) GO TO 200

C

C.....POLARIMETER DATA DISPLAY

CALL POLDDP (I,POLDT,EVART,NP,DEGPOL,PM)

C

C.....SUMMATION FOR SOURCE I

CALL SUMUP (I,EVART,POLDT,TOPIO,KTOPIO,DEGPOL,PM,KFLAG,DOMFRQ)

300 CONTINUE

C

C.....CLOSE SC-4020 FILE

CALL CLEAN

STOP

END

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C.....PARAMETER INITIALIZATION SUBROUTINE
SUBROUTINE INIT (NS,EVART,NP)

C

DIMENSION EVART(4,12)
DIMENSION SNAME(90)

C

INTEGER POLDT

C

DATA BLANKS/6H /
DATA SNAME/

16H	TAU	X,6H-1	,6H	VIR	X,6HR-1	,6H	CEN	X,6HR-2	,
26H	SCD	X,6H-1	,6H	SCD	X,6HR-2	,6H	GX	3+,6H1	,
36H	GX	5-,6H1	,6H	GX	9+,6H1	,6H	SGR	X,6HR-2	,
46H	SER	X,6HR-2	,6H	CYG	X,6HR-1	,6H	CYG	X,6H-4	,
56H	CYG	X,6HR-2	,6H	CAS	X,6HR-1	,6H	VEL	X,6HR-1	,
66H	LUP	X,6HR-1	,6H	NOR	X,6HR-2	,6H	SCD	X,6H-2	,
76H	ARA	X,6HR-1	,6H	GX	-5,6H.6	,6H	SGR	X,6HR-1	,
86H	SER	X,6HR-1	,6H	CEP	X,6HR-1	,6H	LEO	X,6HR-1	,
96H	CEN	X,6HR-3	,6H	CEN	X,6HR-1	,6H	NOR	X,6HR-1	,
16H	SCD	X,6HR-4	,6H	L7	,6H	,6H	L8	,6H	,
26H	UPH	X,6HR-2	,6H	SCD	X,6HR-3	,6H	GX	9+,6H9	,
36H	UPH	X,6HR-1	,6H	SCD	X,6HR-5	,6H	SCD	X,6HR-6	,
46H	LYR	X,6HR-1	,6H	SGR	X,6HR-5	,6H	AQL	X,6HR-1	,
56H	CYG	X,6HR-3	,6H	CYG	X,6H-3	,6H	VUL	X,6HR-1	,
66H	LAC	X,6HR-1	,6H	CEP	X,6HR-2	,6H	CEP	X,6HR-3	/

C

NAMelist /NAM1/NS,NP
NAMelist /NAM2/ID,RAHR,RAMIN,DDEG,DMIN,OBSTIM,BTP,ETP,TSS

C

INPUT VARIABLES

C

NS - NUMBER OF SOURCES

C

NP - PLOT FLAG (1 = PLOT ALL LEVELS OF INTENSITY
0 = PLOT CUMULATIVE ONLY)

C

ID - SOURCE IDENTIFICATION

C

RAHR - RIGHT ASCENSION (HOURS)

C

RAMIN - RIGHT ASCENSION (MINUTES)

C

DDEG - DECLINATION (DEGREES)

C

DMIN - DECLINATION (MINUTES)

C

OBSTIM - OBSERVATION TIME

C

BTP - BEGINNING TABLE POSITION (DEGREES)

C

ETP - ENDING TABLE POSITION (DEGREES)

C

TSS - TABLE STEP SIZE (DEGREES)

C

ALL INPUT VARIABLES ARE STORED IN TABLE EVART EXCEPT NS AND NP

C

C.....READ NUMBER OF SOURCES AND PLOT FLAG

READ (5,NAM1)

C

C.....INITIALIZE SOURCE INPUT LOOP

IF (NS.GT.4) NS = 4

DO 100 I=1,NS

C

C.....READ PARAMETERS FOR SOURCE 1

READ (5,NAM2)

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```
K1 = ID*2-1
K2 = K1+1
EVART(1,1) = SNAME(K1)
EVART(1,2) = SNAME(K2)
EVART(1,3) = BLANKS
EVART(1,4) = RAHR
EVART(1,5) = RMIN
EVART(1,6) = DDEG
EVART(1,7) = DMIN
EVART(1,8) = GBSTIM
EVART(1,9) = BTP
EVART(1,10) = ETP
EVART(1,11) = TSS
EVART(1,12) = (ETP - BTP) / TSS + 1.0001
100 CONTINUE
RETURN
END
```


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```
C.....PULSAR DATA ACQUISITION SUBROUTINE
      SUBROUTINE PULDAQ (I,PULDT)
C
      LOGICAL BIT
      LOGICAL X
C
      DIMENSION NIN(28)
      DIMENSION PULDT(4096)
C
      READ (10) (NIN(M),M=1,28)
      DO 100 J=1,28
      DO 100 K=1,36
      L = K-1
      X = BIT(NIN(J),L)
      N = (J-1)*1000 + (K-1)*36 + K
      PULDT(N) = 0.
      IF (X) PULDT(N) = 1.
      IF (J.EQ.28.AND.K.EQ.28) RETURN
      IF (N.EQ.4096) RETURN
100  CONTINUE
      RETURN
      END
```

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```
C.....PULSAR DATA ANALYSIS SUBROUTINE
      SUBROUTINE PULDAN (I,PULDT,DUMMY,KFLAG,POWSP,TOPI0,KTOP10)
C
      DIMENSION PULDT(4097)
      DIMENSION DUMMY(4097)
      DIMENSION POWSP(4097)
      DIMENSION TOPI0(10)
      DIMENSION KTOP10(10)
C
C.....SET ARRAY DUMMY = PULDT
      DO 100 K=1,4096
        DUMMY(K) = PULDT(K)
      100 CONTINUE
C
C.....TAKE FOURIER TRANSFORM OF PULSAR DATA AND COMPUTE POWER
C.....SPECTRUM DENSITY
      CALL EVAL (PULDT,DUMMY,4096,1000.,POWSP)
C
C.....CALCULATE TOTAL POWER SPECTRUM DENSITY
      TOT = 0.
      DO 200 K=2,2048
        TOT = TOT + POWSP(K)
      200 CONTINUE
C
C.....DETERMINE TOP 10 VALUES OF POWER SPECTRUM DENSITIES
      DO 400 K=1,10
        TOPI0(K) = 0.
        DO 300 J=2,2048
          IF (POWSP(J).LT.TOPI0(K)) GO TO 300
          TOPI0(K) = POWSP(J)
          KTOP10(K) = J
        300 CONTINUE
        L = KTOP10(K)
        POWSP(L) = -POWSP(L)
      400 CONTINUE
C
C.....RESET TOP 10 VALUES IN POWER SPECTRUM
      DO 500 K=1,10
        L = KTOP10(K)
        POWSP(L) = TOPI0(K)
      500 CONTINUE
C
C.....SUM TOP 10 POWER SPECTRUM DENSITIES
      SUM = 0.
      DO 600 K=1,10
        SUM = SUM + TOPI0(K)
      600 CONTINUE
C
C.....CHECK FOR SUM OF TOP 10 GREATER THAN SOME PERCENTAGE OF TOTAL
      PCT = .05
      KFLAG = 0
      IF (SUM.GT.TOT*PCT) KFLAG = 1
      RETURN
      END
```

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```
      SUBROUTINE EVAL(X1,X2,N,SR,Y1)
C
C      IBM 7094 FAST FOURIER CONVERSION
C
C      SUBROUTINE EVAL REMOVES THE MEAN FROM THE X1 AND X2 ARRAYS
C      AND CALLS SUBROUTINE HARM. EVAL TAKES THE TWO ARRAYS
C      RETURNED FROM HARM AND CALCULATES THE POWER SPECTRUM.
C
C      INPUT
C
C      X1      DATA FOR CHNL NO. 1
C
C      X2      DATA FOR CHNL NO. 2
C
C      N      NUMBER OF POINTS TO TRANSFORM IN FFT
C      IF N IS NOT A POWER OF 2, N WILL BE SET TO A NEGATIVE
C      VALUE AND RETURN TO THE CALLING PROGRAM.
C
C      SR      SAMPLE RATE
C
C      Y1      THE POWER SPECTRUM COMPUTED FROM THE FFT
C
C      IDEBUG  IS SET TO 1 FOR DEBUG PRINT
C
C      OUTPUT
C
C      Y1      CROSS PSD
C
C      DIMENSION X1(1),X2(1)
C      DIMENSION M(3)
C      DIMENSION Y1(1)
C      IDEBUG=0
C      DO 210 I=1,12
C      ITEMP=2**I
C      IF(N.LE.ITEMP) GO TO 212
210 CONTINUE
212 M(1)=I
C      IF(2**M(1).EQ.N) GO TO 213
C      WRITE(6,214)
214 FORMAT(34H ***** N IS NOT A POWER OF 2 ***** )
C      N=-N
C      GO TO 25
213 M(2)=0
C      M(3)=0
C      NP=N
C      ANP=NP
C
C      REMOVE MEAN FROM DATA
C
C      T1=0.0
C      T2=0.0
C      DO 10 I=1,N
C      T1=T1+X1(I)
10 T2=T2+X2(I)
C      T1=T1/ANP
C      T2=T2/ANP
```

```

      DO 15 I=1,4
      X1(I)=X1(I)-T1
15   X2(I)=X2(I)-T2
217  CALL HARM(X1,X2,M)
      NDX=NP+1
      X1(NDX)=X1(1)
      X2(NDX)=X2(1)
      AN=1.0/(ANP*ANP)
      AB=AN*2.0
      T=C.0
220  DO 20 I=1,NP
      NDX=NP-I+2
      T1=(X1(I)+X1(NDX))*0.5
      T2=(X2(I)-X2(NDX))*(-0.5)
      Z1=(X2(I)+X2(NDX))*0.5
      Z2=(X1(NDX)-X1(I))*(-0.5)
      H1=T1*Z1+T2*Z2
      H2=T1*Z2-T2*Z1
      Y1(I)=SQRT(H1*H1+H2*H2)*AB
      IF(IDEBUG.EQ.0) GO TO 20

C
C      CALCULATIONS FOR DEBUG PRINT
C
C      F      FREQUENCY
C      X1      REAL ARRAY
C      X2      IMAGINARY ARRAY
C      TEMP6    PSD CHNL 1
C      TEMP7    PSD CHNL 2
C      Y1      CROSS PSD
      TEMP6=(T1*T1+T2*T2)*AB
      TEMP7=(Z1*Z1+Z2*Z2)*AB
      F=1-1
      WRITE(6,8348) F,X1(1), X2(1),TEMP6,TEMP7,Y1(1)
8348  FORMAT(1X,F15.2,5E16.7)
      20 CONTINUE
      25 RETURN
      END

```

```

SUBROUTINE HARM(AR,AI,N)
C
C   HARMONIC ANALYSIS SUBROUTINE. EVALUATES COMPLEX FOURIER SERIES
C   ND=NO. OF DIMENSIONS
C   M(1)=LOG2N1,N1=NO. OF PTS. IN 1-TH DIRECTION, ORDERED SO M(1) IS M
C   AR(1)=REAL PART OF A(I1,I2,I3),I=I1+I2*N1+I3*N1*N2
C   AI(1)=IMAG. PART OF A(I1,I2,I3)
C
C   ALL ARGUMENTS ARE TO BE SET BEFORE ENTRY.  THE AR,AI ARRAYS ARE
C   REPLACED BY THE REAL AND IMAGINARY PARTS, RESP. OF THE FOURIER
C   SUM-
C           I1*J1  I2*J2  I3*J3
C   X(J1,J2,J3)=A(I1,I2,I3)*W1*W2*W3
C   SUMMED OVER I1=0,N1-1,I2=0,N2-1,I3=0,N3-1.
C
C   WHERE W1=N1-TH ROOT OF UNITY
C   X(J1,J2,J3) IS STORED IN LOCATION K=K1+K2*N1+K3*N1*N2 WHERE
C   K1,K2,K3 ARE OBTAINED FROM J1,J2,J3 BY INVERTING THEIR BITS.
C
C   DIMENSION AR(1),AI(1),M(1)
C   DIMENSION W(2),T(2),N(3),NP(3)
C
C   EQUIVALENCE (N1,N),(N2,N(2)),(N3,N(3))
C
C   N1=2**M(1)
C   N2=2**M(2)
C   N3=2**M(3)
C   NP(1)=N(1)
C   NP(2)=N(1)*N(2)
C   NP(3)=N(3)*NP(2)
C   DO 120 ID=1,3
C   MML=M(ID)
C   IF(N(1)-N(ID))10,20,20
10  WRITE(6,12)M
12  FORMAT(27HOM(1) IS NOT LARGEST, M(1)= 315)
C   CALL EXIT
20  IF(MML)130,130,30
30  IL=NP(3)-NP(ID)+1
C   NBIT=MML
C   JBIT=NP(ID)
C   IDI=JBIT
C   FN=JBIT
C   LZ=6.2831853/FN
60  LLAST=M(ID)
C   KLAST=0
C   KKDI=NP(ID)
C   DO 110 IL6=1,LLAST
C   MML=MML-1
C   JBIT=JBIT/2
C   JL=JBIT-1
C   DO FOR K=0
C   DO 70 I=1,IL,IDI
C   JLAST=JL+1
C   DO 70 J=1,JLAST
C   JD=J+JBIT

```

```

      T(1)=AR(JD)
      T(2)=AI(JD)
      AR(JD)=AR(J)-T(1)
      AI(JD)=AI(J)-T(2)
      AR(J)=AR(J)+T(1)
70  AI(J)=AI(J)+T(2)
      IF(KLAST)100,100,72
72  KK=KKDI+1
      DO 90 K=1,KLAST
      IBIT=KK-1
      CALL IRS(IBIT,MML)
      CALL INVERT(IBIT,NBIT)
      ITDP=IBIT
      FN=ITDP
      ARG=ZZ*FN
8050 W(1)=COS(ARG)
      W(2)=SIN(ARG)
      ILAST=IL+KK
      DO 80 I=KK,ILAST,IDI
      JLAST=JL+I
      DO 60 J=1,JLAST
      JD=J+JBIT
      T(1)=AR(JD)*W(1)-AI(JD)*W(2)
      T(2)=AR(JD)*W(2)+AI(JD)*W(1)
      AR(JD)=AR(J)-T(1)
      AI(JD)=AI(J)-T(2)
      AR(J)=AR(J)+T(1)
      AI(J)=AI(J)+T(2)
80  KK=KK+KKDI
90  KK=KK+KKDI
C    END OF K LOOP
100 KKDI=KKDI/2
110 KLAST=2*KLAST+1
C    END OF L LOOP
120 CONTINUE
C    END OF ID LOOP
130 CONTINUE
C    PUT IN CORRECT ORDER
      DO 40 J3=1,N3
      NBIT=M(3)
      I3=J3-1
      I2=I3
      CALL INVERT(I3,NBIT)
      JJJ=I2
      CALL LSF(JJJ,M(2))
      CALL LSF(I3,M(2))
      III=I3
      DO 40 J2=1,N2
      NBIT=M(2)
      I2=J2-1
      CALL INVERT(I2,NBIT)
      JJ=JJJ+J2-1
      CALL LSF(JJ,M(1))
      II=III+I2
      CALL LSF(II,M(1))
      NBIT=M(1)
      DO 40 J1=1,N1

```

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```
      I1=J1-1
      CALL INVERT(I1,NBIT)
      J=J0+J1
      I=I1+I1+1
      IF(I-J)40,40,133
133  T(1)=AR(1)
      T(2)=AI(1)
      AR(1)=AR(J)
      AI(1)=AI(J)
      AR(J)=T(1)
      AI(J)=T(2)
40  CONTINUE
      RETURN
      END
```

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```
      SUBROUTINE LSF(IWORD,NBITS)
C
C
C      LSF      LSF TAKES A STRING OF BITS FROM IWORD AND RESTORES
C               IN DIFFERENT BIT POSITIONS IN IWORD LEFT SHIFTED
C
C      IWORD    CONTAINS THE BIT STRING TO BE MOVED
C      NBITS    THE NUMBER OF BITS IN IWORD TO BE MOVED
C
      IA=0
      NAB=36-NBITS
      CALL FLD(IA,0,NAB,NBITS,IWORD)
      IWORD=IA
      RETURN
C
      ENTRY IRS(IWORD,NBITS)
C
C      IRS      IRS TAKES A STRING OF BITS FROM IWORD AND RESTORES
C               IN DIFFERENT BIT POSITIONS IN IWORD RIGHT SHIFTED
C
      IA=0
      NAB=36-NBITS
      CALL FLD(IA,NBITS,NAB,0,IWORD)
      IWORD=IA
      RETURN
C
      ENTRY INVERT(IWORD,NBITS)
C
C      INVERT    INVERT GIVES A MIRROR IMAGE FOR A STRING OF NBITS
C               AND RESTORES THEM IN IWORD
C
      IF(NBITS.EQ.0) GO TO 20
      I1=35
      I2=36-NBITS
      IA=0
      DO 10 I=1,NBITS
      CALL FLD(IA,I2,I,I1,IWORD)
      I1=I1-1
10  I2=I2+1
      IWORD=IA
20  RETURN
      END
```


C.....PULSAR DATA DISPLAY SUBROUTINE

SUBROUTINE POLDD1 (I,EVART,POWSP,TOPIQ,KTOPIQ,X,Y,DOMFRQ)

C

```

      DIMENSION EVART(4,12)
      DIMENSION POWSP(4097)
      DIMENSION TOPIQ(10)
      DIMENSION KTOPIQ(10)
      DIMENSION X(4097)
      DIMENSION Y(4097)
      DIMENSION XLAB(12),YLAB(12)
      DIMENSION DFBCD(5)

```

C

```

      DATA XLAB/72HFREQUENCY - SOURCE
      IENCI = CPS /
      DATA YLAB/72HAMPLITUDE
      DATA FMT/6H(F6.1)/

```

(DOMINANT FREQU...

C

```

      C = 1000. / 4096.
      NPP = 682
      NP2 = 200
      DO 200 K=1,682
      Y(K) = C.
      DO 100 L=2,4
      M = (K-1)*3+L
      Y(K) = Y(K) + POWSP(M)
100  CONTINUE
      FK = K
      X(K) = ((FK-1.) * C * 3.) + (2. * C)
200  CONTINUE

```

C.....SET UP PLOT TITLE

```

      DO 300 K=1,3
      XLAB(K+3) = EVART(1,K)
300  CONTINUE
      TOP = KTOPIQ(1) - 1
      DOMFRQ = TOP * C
      CALL CTBCD (NBCDWS,DFBCD(1),FMT,1,DOMFRQ)
      XLAB(11) = DFBCD(1)
      CALL QUIK3V (-1,42,XLAB,YLAB,-NPP,X,Y)
      K1 = KTOPIQ(1) - 10
      IF (K1.LT.2) K1 = 2
      FK1 = K1
      DO 400 K = 1,200
      FK1C = FK1*C
      FK = K
      X(K) = FK1C + (FK-1.)*C
400  CONTINUE
      CALL QUIK3V(-1,42,XLAB,YLAB,-NP2,X(1),POWSP(K1))
      RETURN
      END

```

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```
C.....POLARIMETER DATA ACQUISITION SUBROUTINE
      SUBROUTINE POLDAQ(I,J,POLDT)
C
      DIMENSION KDATA(8)
      DIMENSION KBEAM(8)
      DIMENSION POLDT(4,36,16)
C
      INTEGER POLDT
C
      DATA KD/017777/
      DATA KB/017777/
C
C.....INPUT POLARIMETER DATA FOR SOURCE I, TABLE POSITION J
100  READ(8)KPOS,KOFF,(KDATA(M),M=1,8),(KBEAM(M),M=1,8)
C
C.....UNPACK POLARIMETER DATA
      DO 200 L=1,8
      CALL BIN (KDATA(L),0,36,0,KD,1)
      POLDT(I,J,L) = POLDT(I,J,L) + KDATA(L)
      CALL BIN (KBEAM(L),0,36,0,KB,1)
      POLDT(I,J,L+8) = POLDT(I,J,L+8) + KBEAM(L)
200  CONTINUE
      IF (KOFF.EQ.1) GO TO 100
      RETURN
      END
```

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```
C.....POLARIMETER DATA DISPLAY SUBROUTINE
      SUBROUTINE POLDDP (I,POLDT,EVART,NP,DEGPOL,PM)
C
      DIMENSION POLDT(4,36,16)
      DIMENSION EVART(4,12)
      DIMENSION TPOS(36)
      DIMENSION LPF(8)
      DIMENSION TPLAB(12)
      DIMENSION CTPLAB(12)
      DIMENSION CCNTLB(12)
      DIMENSION CNTLAB(12)
      DIMENSION CNT(36)
      DIMENSION TEMP(5)
      DIMENSION LEVEL(8)
      DIMENSION X(5)
      DIMENSION Y(5)
C
      REAL LEVEL
C
      INTEGER POLDT
      INTEGER CCNT
C
      DATA CNTLAB/72H  X-RAY COUNTS FOR          SECONDS OF OBSERVATION =
1 INTENSITY LEVEL      /
      DATA CCNTLB/72H  X-RAY COUNTS FOR          SECONDS OF OBSERVATION = C
1 CUMULATIVE TOTALS    /
      DATA TPLAB /72H TABLE POS = SOURCE          ANG MAX POL
1 DEG OF POL           /
      DATA CTPLAB/72H TABLE POSITION = SOURCE
1                      /
      DATA FMT2/6H(F6.1)/
      DATA LEVEL /6H 1      ,6H 2      ,6H 3      ,6H 4      ,6H 5      ,
16H 6      ,6H 7      ,6H 8      /
      DATA KONE/1/
      DATA KFIVE/5/
C
C.....CLEAR FLAGS FOR MULTILEVEL PLOTS
      DO 50 K=1,8
      LPF(K) = 0
50  CONTINUE
C
C.....SET NTP = NUMBER OF TABLE POSITIONS FOR SOURCE 1
      NTP = EVART(1,12)
C
C.....CALCULATE CUMULATIVE TOTALS FOR Y AXIS
      DO 100 K=1,NTP
      CCNT = 0
      DO 100 L=1,8
      CCNT = CCNT + POLDT(I,K,L)
      CNT(K) = CCNT
C
C.....SET FLAG FOR MULTILEVEL PLOTS IF THERE IS DATA FOR THAT LEVEL
      IF (POLDT(I,K,L).NE.0) LPF(L) = 1
100  CONTINUE
C
```

```
C.....SET UP X AXIS VALUES
      TS = EVART(1,9)
      DO 200 K=1,NTP
        TPOS(K) = TS
        TS = TS + EVART(1,11)
      200 CONTINUE
C
C.....SET UP PLOT TITLES
      TIME = EVART(1,8)
      CALL CTOBCD (NBCDWS,TEMP(1),FMT2,1,TIME)
      CNTLAB(4) = TEMP(1)
      CCNTLB(4) = TEMP(1)
C
C.....CALCULATE DEGREE OF POLARIZATION
      CMIN = CNT(1)
      CMAX = CNT(1)
      KMAX = 1
      DO 250 K = 2,NTP
        IF (CNT(K).LE.CMAX) GO TO 240
        CMAX = CNT(K)
        KMAX = K
      240 CONTINUE
      IF (CNT(K).LT.CMIN) CMIN = CNT(K)
      250 CONTINUE
      DEGPOL = ((CMAX - CMIN) / (CMAX + CMIN)) * 100.
      CALL CTOBCD (NBCDWS,TEMP(1),FMT2,1,DEGPOL)
      TPLAB(12) = TEMP(1)
C
C.....CALCULATE ANGLE OF MAXIMUM POLARIZATION
      KS = KMAX-2
      IF (KMAX.LT.3) KS = 1
      IF (KMAX.GT.(NTP-2)) KS = NTP-4
      DO 260 K=1,5
        X(K) = TPOS(KS)
        Y(K) = CNT(KS)
        KS = KS+1
      260 CONTINUE
      PM = TPOS(KMAX)
      CM = CNT(KMAX)
      ARG = TPOS(KMAX)
      STEP = .01
      IF (Y(2).GT.Y(4)) STEP = -.01
      IDUM = 0
      DO 270 K=1,100
        ARG = ARG + EVART(1,11) * STEP
        CALL LAGNG (IERR,IDUM,KONE,ARG,KFIVE,X,KFIVE,Y,ANS)
        IF (IERR.NE.0) WRITE (6,1000) ARG
      1000 FORMAT (1X,33HERROR IN INTERPOLATION - ARG = ,E15,7)
        IF (ANS.GT.CM) PM = ARG
        IF (ANS.GT.CM) CM = ANS
      270 CONTINUE
      CALL CTOBCD (NBCDWS,TEMP(1),FMT2,1,PM)
      TPLAB(8) = TEMP(1)
C
C.....PLOT CUMULATIVE
      TPLAB(4) = EVART(1,1)
```

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```
      TPLAB(5) = EVART(I,2)
      CALL QUIK3V (-1,42,TPLAB,CCNTLB,-NTP,TPOS,CNT)
C
C.....CHECK FOR MULTIPLE PLOTS OF ALL INTENSITIES
      IF (NP.EQ.0) RETURN
C
C.....PLOT MULTIPLE PLOTS OF ALL INTENSITIES
      DO 400 K=1,8
      IF (LPF(K).EQ.0) GO TO 400
      DO 300 L=1,NTP
      CNT(L) = POLDT(I,L,K)
300  CONTINUE
      CTPLAB(5) = EVART(I,1)
      CTPLAB(6) = EVART(I,2)
      CNTLAB(12) = LEVEL(K)
      CALL QUIK3V (-1,42,CTPLAB,CNTLAB,-NTP,TPOS,CNT)
400  CONTINUE
      RETURN
      END
```

```

SUBROUTINE SUBUP (I,EVART,POLDT,TOPI0,KTOP10,DEGPOL,PM,KFLAG,
1 DOMFRQ)
  DIMENSION EVART(4,12)
  DIMENSION POLDT(4,36,16)
  DIMENSION TOP10(10)
  DIMENSION KTOP10(10)
  DIMENSION N(16)
  INTEGER POLDT
  WRITE (6,1000) (EVART(I,H),H=1,11)
1000 FORMAT (1H1,20X,22HSUMMARY OF OBSERVATION///1X,10HSOURCE = ,
1 3A6//1X,28HLOCATION = RIGHT ASCENSION,F5.0,7H HOURS,,F7.0,
2 8H MINUTES//14X,11HDECLINATION,F9.0,9H DEGREES,,F5.0,8H MINUTES//
3 1X,19HOBSERVATION TIME = ,F6.0,31H SECONDS AT EACH TABLE POSITION
4 ///19H X-RAY POLARIZATION//6X,24HBEGINNING TABLE POSITION,F6.0,
5 8H DEGREES//6X,21HENDING TABLE POSITION,F9.0,8H DEGREES//
6 6X,15HTABLE STEP SIZE,9X,F6.0,8H DEGREES///)
  WRITE (6,1001) PM,DEGPOL
1001 FORMAT (6X,32HANGLE OF MAXIMUM POLARIZATION = ,F7.1,8H DEGREES///
1 6X,25HDEGREE OF POLARIZATION = ,F7.1,8H PERCENT///
2 1X,15HX-RAY INTENSITY//
3 22X,11HDATA COUNTS,14X,11HBEAM COUNTS/
4 6X,5HLEVEL,8X,6HNUMBER,4X,7HPERCENT,8X,6HNUMBER,4X,7HPERCENT//)
  DO 100 K=1,16
    N(K) = 0
    DO 100 L=1,36
      N(K) = N(K) + POLDT(I,L,K)
100 CONTINUE
    NDT = 0
    NBT = 0
    DO 200 K=1,8
      NDT = NDT + N(K)
      K1 = K+8
      NBT = NBT + N(K1)
200 CONTINUE
    DT = NDT
    FN = N(K)
    PD = (FN/DT)*100.
    BT = NBT
    FN = N(K+8)
    PB = (FN/BT)*100.
    K1 = K+8
    WRITE (6,1002) K,N(K),PD,N(K1),PB
1002 FORMAT (8X,11,9X,16,5X,F6.1,8X,16,5X,F6.1)
300 CONTINUE
    WRITE (6,1003)
1003 FORMAT (///1X,15HX-RAY PULSATION//)
    IF (KFLAG.EQ.0) WRITE (6,1004)
1004 FORMAT (6X,42HTHIS SOURCE DOES NOT APPEAR TO BE A PULSAR)
    IF (KFLAG.NE.0) WRITE (6,1005) DOMFRQ
1005 FORMAT (6X,41HTHIS SOURCE APPEARS TO BE A PULSAR WITH A/
1 5X,13H FREQUENCY OF,F7.2,18H CYCLES PER SECOND)
  RETURN
END

```

SECTION 3. DATA GENERATION PROGRAM

The Data Generation Program simulates the output of the X-ray polarimeter and pulsar mode counter. The output of this program serves as the input to the simulated experiment program.

3.1 Input

The following parameters are input in NAMELIST format for each X-ray source to be simulated:

NAME - Source Identification Code
ITIME - Observation Time at Each Table Position (Seconds)
STEP - Table Step Size (Degrees)
BEGIN - Beginning Table Position (Degrees)
END - Ending Table Position (Degrees)
AMAX - Angle of Maximum Polarization (Degrees)
DATAM - Data Count in Plane of Maximum Polarization (Average Data Counts per Second)
DATAP - Data Count in Plane Perpendicular to Plane of Maximum Polarization (Average Data Counts per Second)
TOTI - Total Intensity (Average Data and Beam Counts per Second)
DEV - Desired Random Deviation of Polarimetry Data (Percentage)
FREQ - Pulsation Frequency (Pulses per Second)
AMP - Pulsation Wave Form Amplitude
PHS - Pulse Height Spread (Percentages)

3.2 Program Flow

Figure 3-1 presents the flow of the Data Generation Program.

3.3 Subroutines

The following subroutines are called by the Data Generation Program:

RANDOM - Random Number Generator
PULSAR - Pulsar Data Generator
OVRFLO - Polarimeter Buffer Overflow

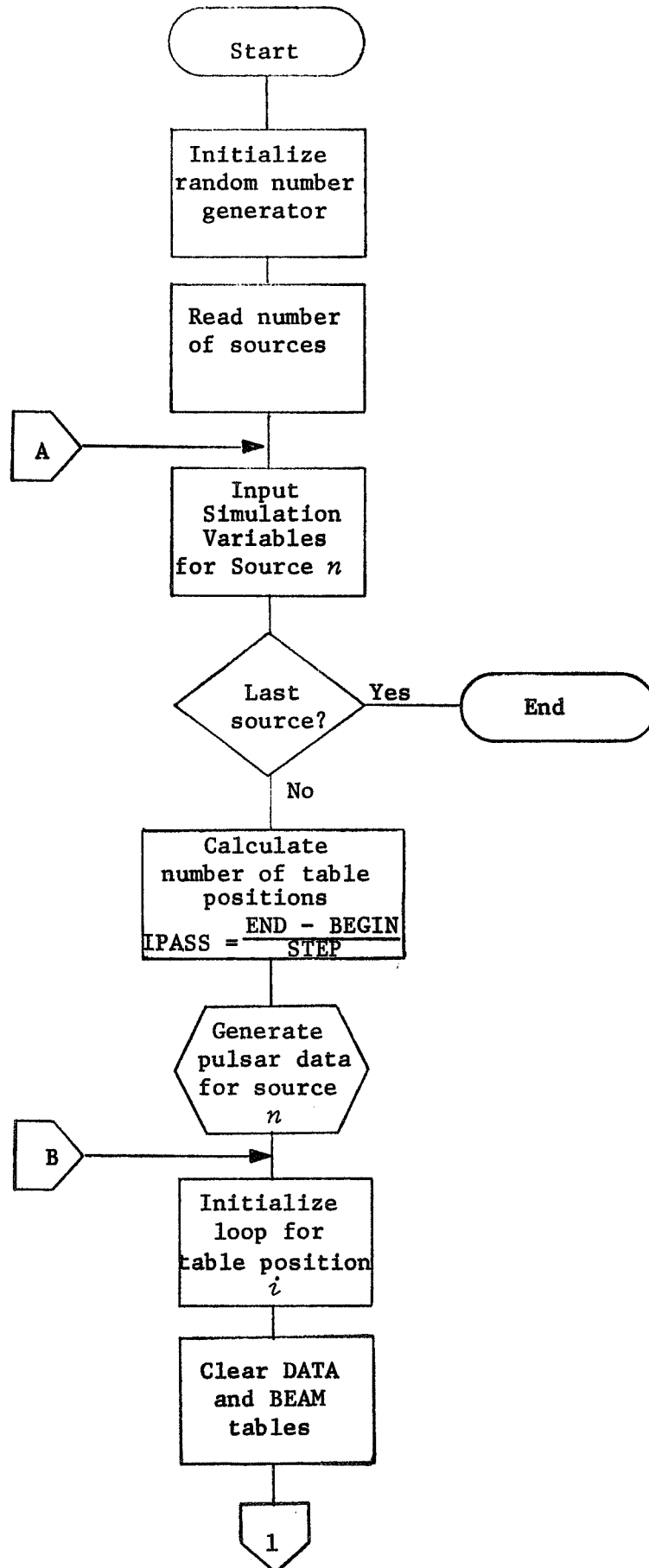


Figure 3-1. Routine DRIVER

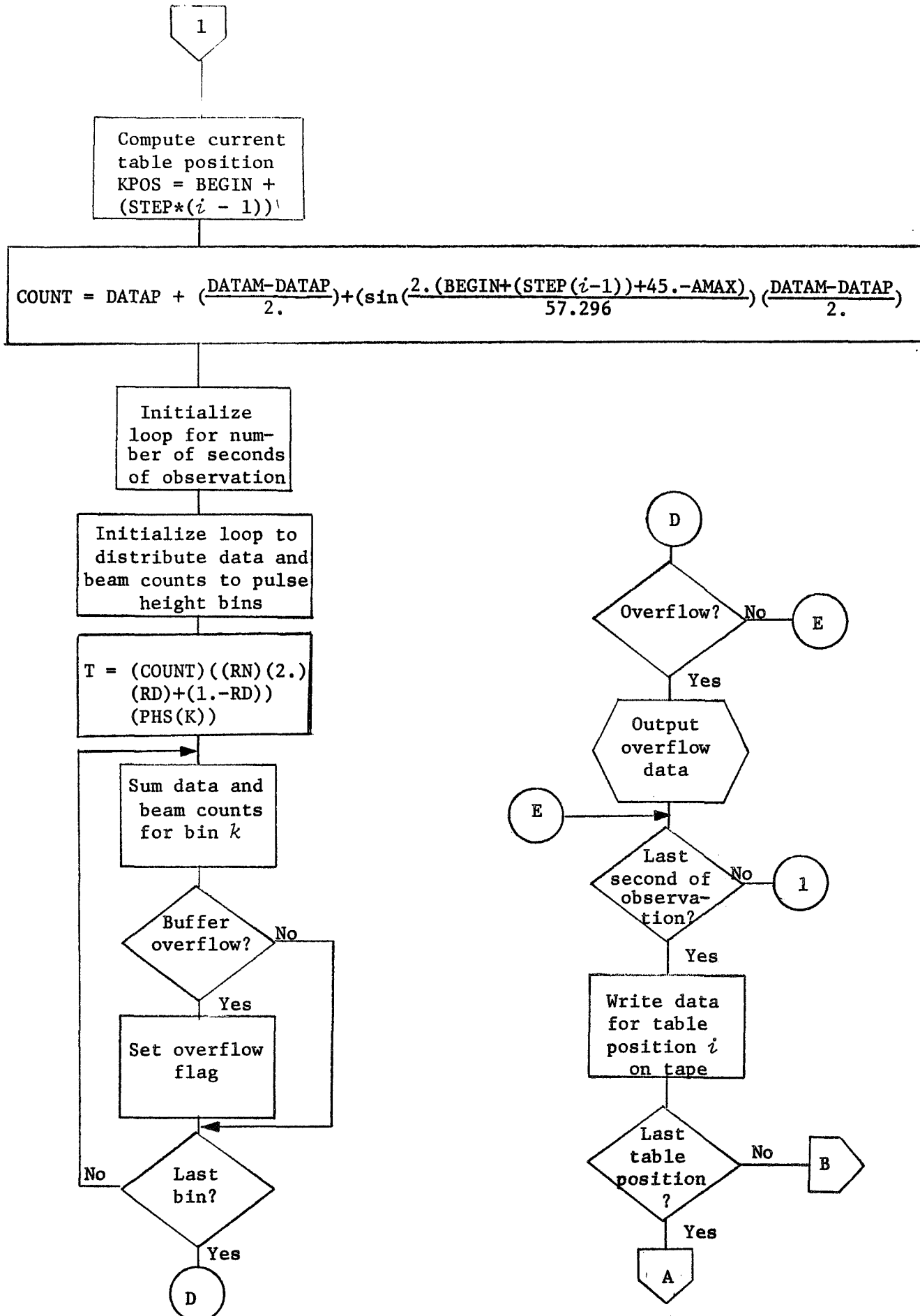


Figure 3-1. (Cont.) Routine DRIVER

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Subroutine OVRFLO

Purpose: Subroutine OVRFLO outputs overflow polarimeter data and beam counts and adjust the internal counters accordingly.

Usage: CALL OVRFLO (DATC, BEAM, KPOS)

DATC - Data count array

BEAM - Beam count array

KPOS - Table position

Storage: 207₍₈₎ locations

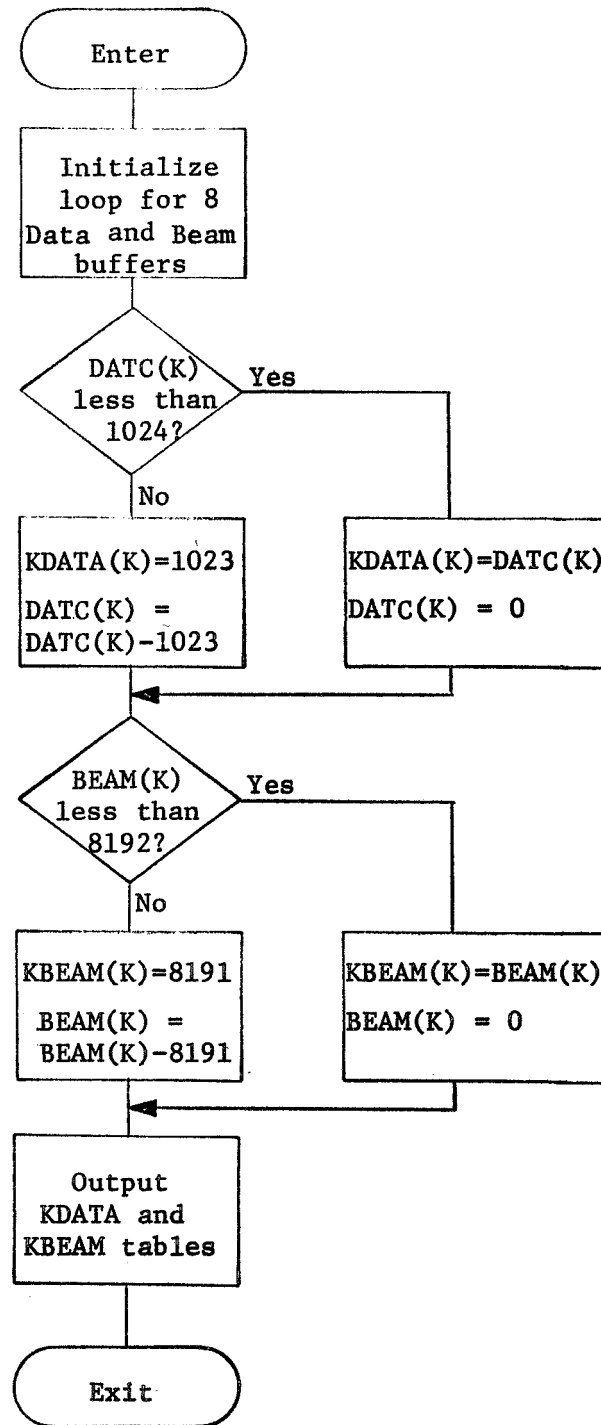


Figure 3-2. Subroutine OVRFLO

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Subroutine PULSAR

Purpose: Subroutine PULSAR computes simulated pulsar data as a sinusoidal function of time modified by a normally distributed probability function.

Usage: CALL PULSAR (TOTI, FREQ, AMP)

TOTI - Total intensity of simulated source
FREQ - Frequency of simulated source
AMP - Peak amplitude of sine wave

Method: Pulsar data is generated as a sinusoidal function of time, i.e.,

$$F = \frac{AMP}{2} \sin \left(\frac{2 \pi (i-1)}{P-1} \right) + RN$$

where i indicates the i^{th} millisecond of observation during period P where

$$P = (1./FREQ)1000.$$

and RN is a normally distributed random number between 0. and 1.0. F is compared against a threshold,

$$THRESH = 1. - (TOTI/1000.)$$

If F is greater than or equal to $THRESH$, N_i is set to 1 indicating an X-ray photon hit during millisecond i ; otherwise, N_i is set to 0 indicating no hit.

Subroutines: PULOUT (Pulsar data output)
RANDOM (Random member generator)

Storage: 2160₍₈₎

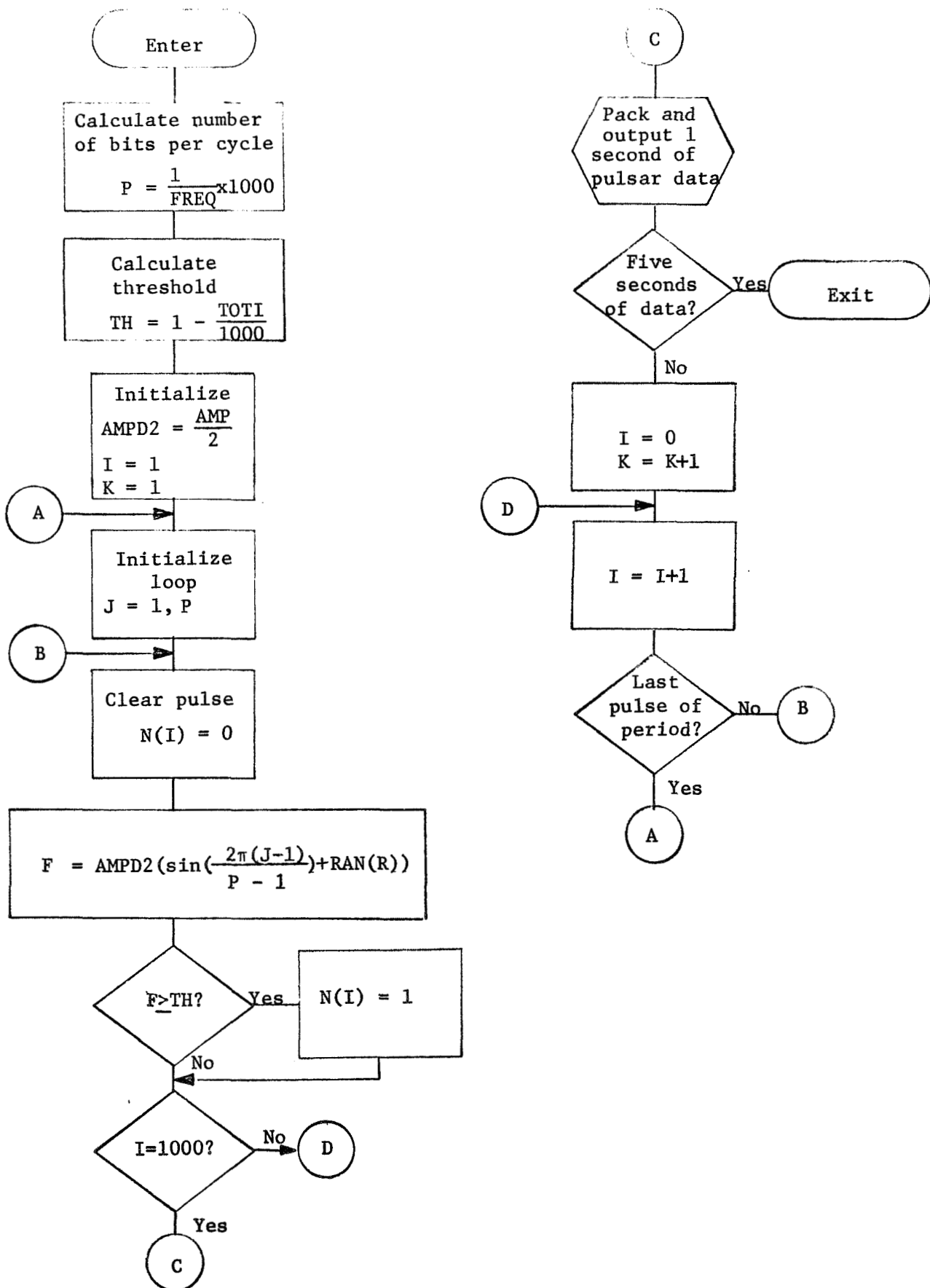


Figure 3-3. Subroutine PULSAR

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Subroutine PULOUT

Purpose: Subroutine PULOUT packs and outputs the pulsar data generated by subroutine PULSAR.

Usage: CALL PULOUT (N)

N - Pulsar data array

Method: Pulsar data is packed in string format where 1 bit represents 1 millisecond of observation. All 36 bits of the computer word, which are read from left to right, are used. One second of observation is output on tape per call to subroutine PULOUT.

Subroutines: BIN (or operator)

Storage: 433₍₈₎

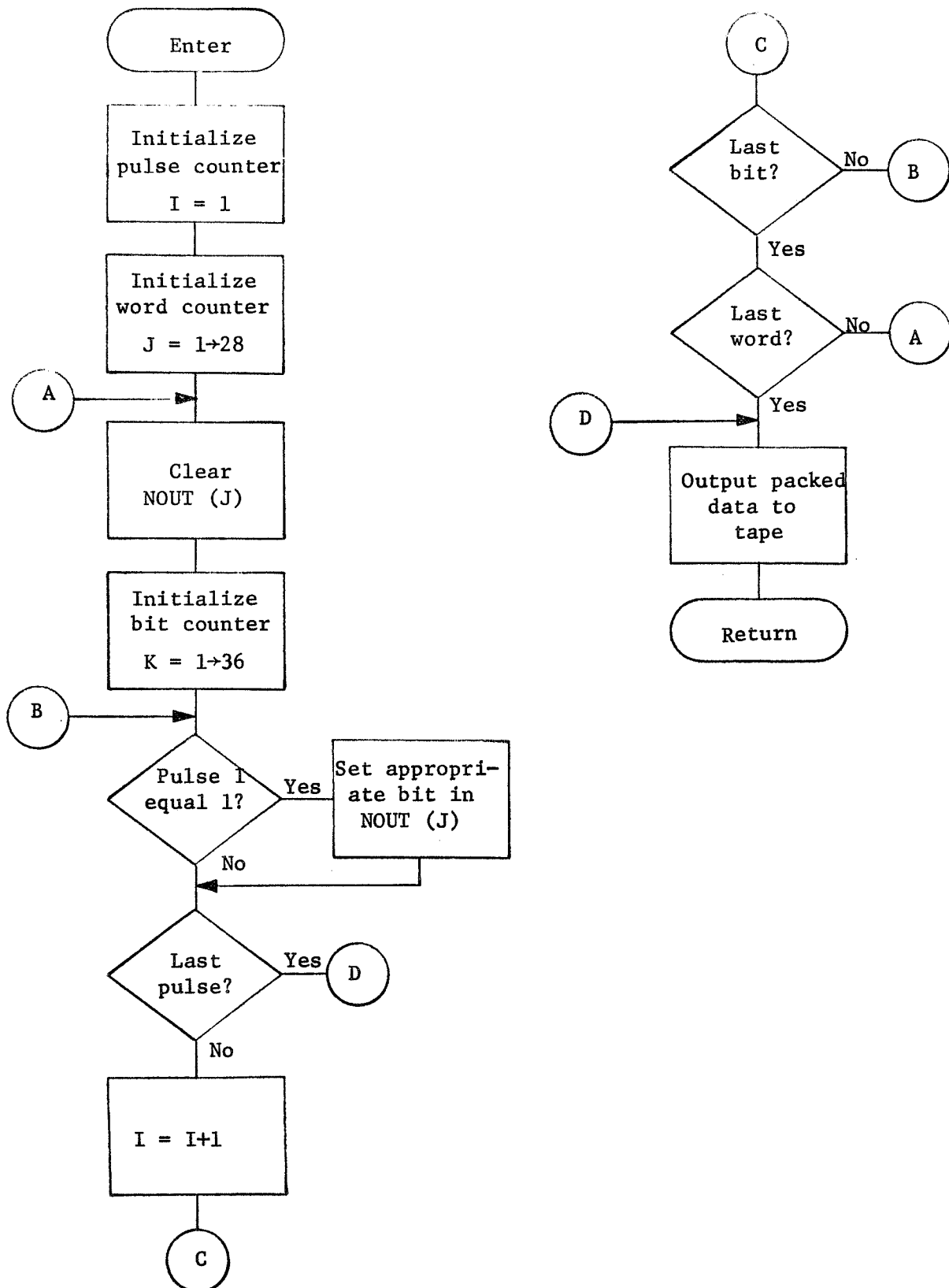


Figure 3-4. Subroutine PULOUT

3.4 Sample Input

Figure 3-5 presents the input data for 4 X-ray sources. These sample cases correspond to the sample cases presented in Section 2.

\$DATA	
\$INPUTS	
NS = 4\$	} Number of Sources
\$INPUT	
NAME = 2	
ITIME = 10	
STEP = 10.	
BEGIN = 0.	
END = 350.	
AMAX = 27.1	
DATAM = 200.	
DATAP = 100.	
DEV = .3	
TOTI = 250.	
FREQ = 4.9	
PHS = .0,.17,.29,.41,.02,.01,0.,0.	
AMP = .5\$	
\$INPUT	
NAME = 18	
ITIME = 20	
STEP = 10.	
AMAX = 96.	
DATAM = 75.	
DATAP = 10.	
DEV = .2	
TOTI = 100.	
FREQ = 3.6	
PHS = .01,.07,.37,.25,.18,.10,.02,.0	
AMP = 1.\$	
\$INPUT	
NAME = 21	
ITIME = 20	
AMAX = 21.	
DATAM = 400.	
DATAP = 80.	
DEV = .1	
TOTI = 450.	
FREQ = 9.7	
PHS = 0.,0.,0.,.02,.17,.63,.15,.03	
AMP = .7\$	
\$INPUT	
NAME = 13	
ITIME = 15	
STEP = 5.	

Figure 3-5. Sample Data Generator Program Input

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BEGIN = 0.
END = 175.
AMAX = 11.7
DATAM = 500.
DATAP = 400.
DEV = .5
TOTI = 550.
FREQ = 18.1
PHS = .87,.12,.01,0.,0.,0.,0.,0.
AMP = .6\$
\$END

Source 4

Figure 3-5 (cont). Sample Data Generator Program Input

3.5 Sample Output

Sample output data for two X-ray sources are presented in Figures 3-6 through 3-9.

LOC	06630	AC		QP	HQ		SI	KEYS		
		000000000000 -00000000+00			000000000014 -00000002-38			000000000000 000000000000		
		XR1		XR2		XR3		XR4		
		77634 -00124		77774 -00004		01474 -76304		71150 -06630		
		XR5		XR6		XR7		XR8		
		00000 -00000		00005 -77773		00000 -00000		63072 -14706		
		SENSE LIGHT		SENSE SWITCH		SENSE SWITCH		SENSE SWITCH		
		1 2 3 4		1 2 3 4		1 2 3 4		1 2 3 4		
		OFF OFF		OFF OFF		OFF OFF		OFF OFF		
		ON ON		ON ON		ON ON		ON ON		
06332	001010100100	000100000101	001001110000	00100110110	00001100001	001010001111	000010101101	010010010010		
06342	010000000101	000000000100	010000000000	000000000000	000000000000	000000000000	000000000000	000000000000		
06352	000010001000	101010011100	010100000111	001100001001	101100010001	101011011111	000011010101	110101110111		
06362	110000000000	000000000000	100010000000	001000010100	000000000000	000000000000	000000000000	000000000000		
06372	000000000100	000010100000	000111100010	1000000001100	001001011111	011001100101	000111110101	100100100101		
06402	110000101101	01101100010	000001001100	000000000100	000010001000	100000000000	000000000000	000000000000		
06412	000000000000	0000000001001	0010000000010	0110000000011	001101010000	001010011010	001000011100	010110010101		
06422	110010000010	000001000000	010010000000	011000110010	000011000000	011000000000	000000000000	000000000000		
06432	000000001000	000000000000	000000000000	000001100000	000010000101	010010001011	00101100100	111000000110		
06442	011111000111	010000111110	110000111110	000000100000	001001010011	010001011010	000010000000	000000010000		
06452	000000000000	000000000000	100001000100	000000000000	000000000000	000000000000	000000000000	000000000000		

Figure 3-6. Pulsar Data - Source 1

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Figure 3-7. Polarimeter Data - Source 1

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LOC	04630	AC	QP	MQ	SI	KEYS				
		000000000000 -00000000+00	00	000000000014 -00000002-38	3000000000000	00000000000000				
		XR1	XR2	XR3	XR5	XR6				
		77654	77774	77767	00000	00005				
		-00124	-00004	-00011	-00000	-77773				
						XR7				
						63072				
						-14706				
TRAP	DCT	10T	OFF	SENSE LIGHT	1	2	3	4	5	6
OFF	OFF	OFF	ON	OFF	OFF	OFF	OFF	OFF	OFF	OFF
06332	011111011011	011110110111	111011001111	001011001011	110110100011	000000010010	000000000000	000001000000		
06342	100111001101	111111111111	001111011111	010010001000	110110111110	011000100000	000000000000	000010000000		
06352	100000001011	000100001101	101111011111	111011110110	011111101010	100101100100	000000001000	010000000000		
06362	0100000000001	101010110110	000101111111	111111101011	110111111111	111001101111	110001000000	000000000000		
06372	100011000100	000000100001	010011110100	111001110110	110111111111	111010011111	100100110001	011000100110		
06402	000000000000	000001010000	010000100110	101000011111	111111111111	010111111011	110111011011	000101110001		
06412	010001000000	000100010110	010100000000	100010101111	100110111001	111111111101	010011100000	011100010111		
06422	101010010001	000001000000	101000000000	000101011100	010011001011	111110111110	111001111011	101010000111		
06432	110010010010	100001110000	000000001001	000100000011	000000000111	010011101111	110111110111	011110110100		
06442	110110010111	010101101010	100000010000	000000001000	001100010010	000110000000	011111100111	110111110101		
06452	100001010101	001000010001	001101010000	000000000000						

Figure 3-8. Pulsar Data - Source 3

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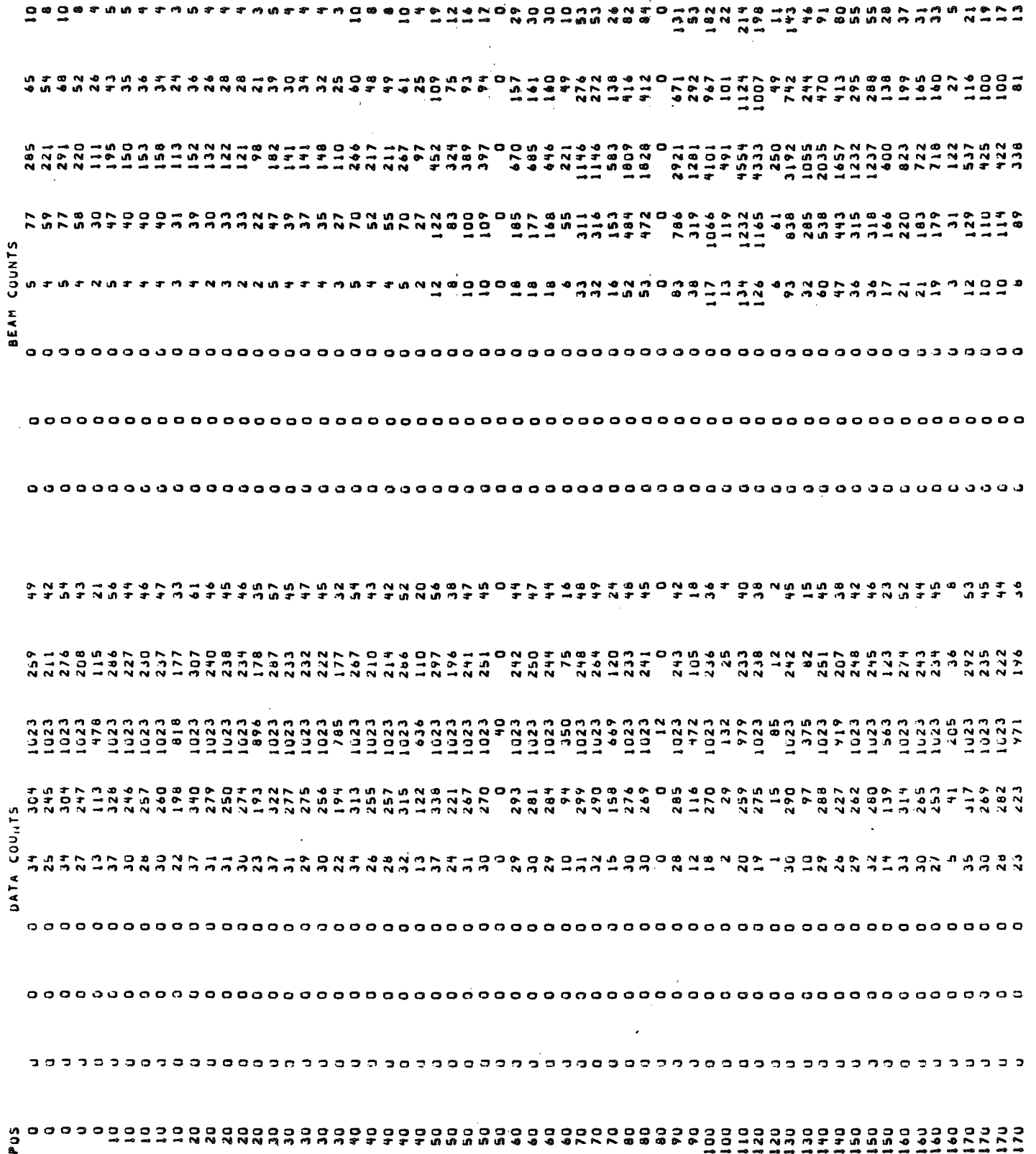


Figure 3-9. Polarimeter Data - Source 3

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Figure 3-9 (cont.). Polarimeter Data - Source 3

3.6 Program Listing

```

C      POLARIMETER AND PULSAR MODE COUNTER DATA SIMULATION ROUTINE
C      KEN CAUFF - SYSTEM DEVELOPMENT CORPORATION, HUNTSVILLE, ALABAMA
C      PHONE 559-7711
C
C      THIS ROUTINE SIMULATES THE OUTPUT OF AN X-RAY POLARIMETER AND
C      PULSAR MODE COUNTER. SEE SDC DOCUMENT TM-(L)-HU-033/000/00 FOR A
C      DESCRIPTION OF THIS SPACE BORNE EXPERIMENT.
C
C      INPUT VARIABLES
C
C      NAME      - SOURCE IDENTIFICATION CODE
C      ITIME     - OBSERVATION TIME AT EACH TABLE POSITION (SECONDS)
C      STEP      - TABLE STEP SIZE (DEGREES)
C      BEGIN     - BEGINNING TABLE POSITION (DEGREES)
C      END       - ENDING TABLE POSITION (DEGREES)
C      AMAX      - ANGLE OF MAXIMUM POLARIZATION (DEGREES)
C      DATAM     - DATA COUNT IN PLANE OF MAXIMUM POLARIZATION (AVERAGE DATA
C      COUNTS PER SECOND)
C      DATAP     - DATA COUNT IN PLANE PERPENDICULAR TO PLANE OF MAXIMUM
C      POLARIZATION (AVERAGE DATA COUNTS PER SECOND)
C      TOTI      - TOTAL INTENSITY (AVERAGE DATA AND BEAM COUNTS PER SECOND)
C      DEV       - DESIRED RANDOM DEVIATION OF POLARIMETRY DATA (PERCENTAGE)
C      FREQ      - PULSATION FREQUENCY (PULSES PER SECOND)
C      AMP       - PULSATION WAVE FORM AMPLITUDE
C      PHS       - PULSE HEIGHT SPREAD (PERCENTAGES)
C
C      DIMENSION PHS(8)
C      DIMENSION DATC(8)
C      DIMENSION BEAM(8)
C
C      INTEGER DATC
C      INTEGER BEAM
C
C      DATA PHS/.05,.10,.15,.20,.20,.15,.10,.05/
C      DATA KZERO/L/
C
C      NAMLIST /INPUTS/ NS
C      NAMLIST /INPUT/NAME,ITIME,STEP,BEGIN,END,AMAX,DATAM,DATAP,
C      1          TOTI,DEV,FREQ,AMP,PHS
C
C      C.....INITIALIZE RANDOM NUMBER GENERATOR
C      NN = 245
C      CALL RNDN(NN)
C      C.....INPUT NUMBER OF SOURCES
C      REAL (5,INPUTS)
C      NC = 0
C      C.....CHECK FOR LAST SOURCE
C      10  CONTINUE
C      NC = NC+1
C      IF (NC.GT.NS) GO TO 400
C      C.....REAL SIMULATION VARIABLES FROM CARDS FOR 1 X-RAY SOURCE
C      REAL (5,INPUT)
C      C.....FOR EACH TABLE POSITION
C      IPASS = 1

```

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```
      IF (STEP.EQ.C.) GO TO 25
      IPASS = (END - BEGIN) / STEP + 1.0001
25    CONTINUE
      CALL PULSAR (TOTI,FREQ,AMP)
      WRITE (6,1000)
1000  FORMAT (1H1,3HPUS,22X,11HDATA COUNTS,47X,11HBEAM COUNTS/)
      DO 300 I=1,IPASS
      FI = I
      DO 50 M=1,8
      DATC (M) = 0
50    BEAM(M) = 0
      KPOS = BEGIN + (STEP * (FI - 1.))
      COUNT = DATAP + (DATAM - DATAP) / 2. + SIN(2. * (BEGIN + (STEP *
      (FI - 1.)) + 45. - AMAX) / 57.296) * (DATAM - DATAP) / 2.
C.....FOR EACH SECOND AT TABLE POSITION I
      DO 200 J=1,ITIME
C.....DISTRIBUTE DATA COUNTS BETWEEN THE 8 PULSE HEIGHT ANALYZER BINS
      KOF = 0
      DO 100 K=1,8
      CALL RANDOM (R)
      INT = COUNT * (R * 2. * DEV + 1. - DEV) * PHS(K)
      DATC(K) = DATC(K) + INT
      IF (DATC(K).GT.1023) KOF = 1
      CALL RANDOM (R)
      INT = (TOTI - COUNT) * (R * 2. * DEV + 1. - DEV) * PHS(K)
      BEAM(K) = BEAM(K) + INT
      IF (BEAM(K).GT.8191) KOF = 1
100  CONTINUE
C.....IF BUFFER OVERFLOW, OUTPUT DATA
      IF (KOF.EQ.1) CALL OVRFLD (DATC,BEAM,KPOS)
200  CONTINUE
C.....OUTPUT DATA AND BEAM COUNTS FOR TABLE POSITION I
      WRITE (8)KPOS,KZERO,(DATC(M),M=1,8),(BEAM(M),M=1,8)
      WRITE (6,2000) KPOS,(DATC(M),M=1,8),(BEAM(M),M=1,8)
2000  FORMAT (1X,I3,8I7,3X,8I7/)
300  CONTINUE
      GO TO 10
400  CONTINUE
      END FILE 8
      REWIND 8
      END FILE 10
      REWIND 10
      STOP
      END
```


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```
      SUBROUTINE OVERFLO (DATC,BEAM,KPOS)
C
C.....OUTPUT OVERFLOW DATA AND BEAM COUNTS
C
C      DIMENSION DATC(8),BEAM(8),KDATA(8),KBEAM(8)
C
C      INTEGER DATC,BEAM
C
C      DATA KONE/1/
C
C      DO 400 K=1,8
C      IF (DATC(K).LT.1024) GO TO 100
C      KDATA(K) = 1023
C      DATC(K) = DATC(K) - 1023
C      GO TO 200
100  KDATA(K) = DATC(K)
C      DATC(K) = 0
200  IF (BEAM(K).LT.8192) GO TO 300
C      KBEAM(K) = 8191
C      BEAM(K) = BEAM(K) - 8191
C      GO TO 400
300  KBEAM(K) = BEAM(K)
C      BEAM(K) = 0
400  CONTINUE
C      WRITE (8,KPOS,KONE,(KDATA(M),M=1,8),(KBEAM(M),M=1,8)
C      WRITE (6,2000) KPOS,(KDATA(M),M=1,8),(KBEAM(M),M=1,8)
2000 FORMAT (1X,I3,8I7,3X,8I7/)
C      RETURN
C      END
```

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```
      SUBROUTINE PULSAR (TOTI,FREQ,AMP)
C
C.....COMPUTE PULSAR STRING DATA
C
      DIMENSION N(1000)
C
      INTEGER SLOTS
C
      DATA TWOPI/6.2832/
C
      SLOTS = (1./FREQ) * 1000.
      THRESH = 1. - (TOTI/1000.)
      AMPD2 = AMP/2.
      I = 1
      K = 1
50    DO 100 J=1,SLOTS
      N(I) = 0
      X = J - 1
      S = SLOTS - 1
      CALL RANDOM(R)
      TEMP = AMPD2 * SIN((TWOPI * X) / S) + R
      IF (TEMP.GE.THRESH) N(I) = 1
      IF (I.NE.1000) GO TO 100
      CALL PULOUT (N)
      IF (K.EQ.5) RETURN
      I = 0
      K = K+1
100   I = I+1
      GO TO 50
      END
```

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```
      SUBROUTINE PULOUT (N)
C
C.....PACK PULSAR DATA AND OUTPUT TO TAPE
C
      DIMENSION N(1000),NOUT(28),NBIT(36)
      DIMENSION NPRNT(84)
C
      DATA NBIT /04000000000000,02000000000000,01000000000000,
1 040000000000,020000000000,010000000000,040000000000,020000000000,
2 010000000000,040000000000,020000000000,010000000000,040000000000,
3 0200000000,0100000000,04000000,02000000,01000000,04000000,02000000,
4 01000000,04000000,02000000,01000000,04000000,02000000,01000000,04000000,02000000,01000000,
5 040,020,010,04,02,01/
C
      I = 1
      DO 100 J=1,28
        NOUT(J) = 0
        DO 100 K=1,36
          IF (N(I).EQ.1) CALL BIN(NOUT(J),0,36,0,NBIT(K),2)
          IF (I.EQ.1000) GO TO 200
          I = I+1
100    CONTINUE
200    CONTINUE
      WRITE (10) (NOUT(M),M=1,28)
      I = 1
      DO 1000 J=1,84
        NPRNT(J) = 0
        DO 1000 K= 1,12
          KT3 = K*3
          IF (N(I).EQ.1) NPRNT(J) = NPRNT(J) + NBIT(KT3)
          IF (I.EQ.1000) GO TO 2000
          I = I + 1
1000   CONTINUE
2000   CONTINUE
      CALL PDUMP(NPRNT(1),NPRNT(84),0)
      RETURN
      END
```

SECTION 4. SUMMARY

The X-ray polarimeter experiment of FPE 5.1 was analyzed and developed to the point where on-board operational software could be designed and implemented for the acquisition, analysis and display of primary data. The purpose of this effort was to establish some estimates of computer speed and capacity required for on-board experiment processing. Since the software was coded in FORTRAN IV and implemented on the MSFC Computation Laboratory's IBM 7094 II, estimates of on-board timing and load factors must be adjusted to take into consideration the difference in speed and capacity of future on-board computers and the IBM 7094.

Processing time required on the IBM 7094 to acquire, analyze and display the primary data for a typical observation sequence of one X-ray source is approximately 33 seconds. Since the acquisition time for simulated data is considerably faster than for actual real time data acquisition, the total processing time for data analysis and display can be spread out over the entire observation sequence. For example, a typical observation sequence of 20 seconds at each of eighteen table positions represents six minutes of elapsed time as compared to 33 seconds of processing time in the simulated mode of operation. This would indicate that the processing of the primary data can be time shared with other experiment operations.

Computer storage was utilized as follows:

Experiment modules	2,600 words
System and library routines	15,500 words
Data and working storage	<u>14,600</u> words
Total	32,700 words

The Fourier analysis of the pulsar data was limited by the available core storage of the IBM 7094 to 4096 words of input---just over four seconds of observation. It would be desirable to increase the data input to 16,384 words or 16+ seconds of observation to improve the accuracy and reliability of the analysis process. This is especially desirable for weak sources or sources with a relatively long pulsation period.

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This study has indicated that real time, on-board processing of primary experiment data can be beneficial not only in reducing the amount of data that must be transmitted, but also in reducing the reaction time to modify or update the experiment sequence due to unexpected events or experiment results.

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