# LUNAR PHOTOGRAPHS <br> FROM APOLLOS 8, 10, and II 





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## Acknowledgments

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## Introduction

The purpose of this photographic atlas is to make available to scientists and to the interested public a comprehensive sample of photographs from the first three Apollo lunar missions. This atlas contains photographs of almost every type of feature and terrain that exists on the lunar surface.

Selected photographs, though not as many as are in this atlas, are available to the public through

The Manned Spacecraft Center<br>Public Affairs Office<br>Houston, Tex. 77058

Scientific users such as universities, institutions, or schools may contact

National Space Science Data Center<br>Goddard Space Flight Center<br>Code 601<br>Greenbelt, Md. 20771

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## Apollo Lunar Photography

Much of the photography from Apollos 8 and 10 was oriented operationally; prime photographic targets were the landing sites and landmarks used for guidance to landing sites. Landmarks of particular significance are noted in captions accompanying the photographs.

The front side, or Earth-facing hemisphere, of the Moon has been studied and mapped by astronomers for several centuries, and most of the prominent features in this Earth-side region have been named, usually after famous scientists or astronomers. Prominent named features shown in the photographs are identified by name in the captions.
All of the names by which lunar features are identified must be approved by the International Astronomical Union (IAU). This scientific body governs the nomenclature of lunar features; any new or changed designations must be submitted to the IAU before the new designations can become officially recognized.

However, the Apollo program injected an entirely new, though unofficial, vocabulary into the listings of lunar features. Names such as U.S. 1, Diamondback Rille, and Boot Hill gained prominence in publications and in the news media. The reason for the development and use of this Apollo nomenclature is quite simple. The lunar-surface photographs from the mapping missions of Lunar Orbiter produced a level of detail that had never before been discernible -even by the best telescopes on Earth. From Lunar Orbiter photographs, features as small as 5 m were detectable; before the Lunar Orbiter missions, only the most powerful telescopes operating under the best atmospheric conditions could resolve objects with dimensions of 2 km . Among the thousands of lunar features and formations that had been unseen or unidentifiable from Earth, there were distinct or unique features that were especially useful for landmark tracking to the landing sites. The IAU could not assign names to these features because of the operational time constraints imposed by the Apollo flight schedule. Therefore, the Apollo crewmembers and the mission planners selected names arbitrarily
for those features that were used as identification points and landmarks for navigation. These designations are not intended to be submitted to the IAU for consideration.

A few of the features of the lunar far side have been named on the basis of photographs returned from the 1959 U.S.S.R. Lunik 3 mission, the first mission to photograph the back side of the Moon. Later, the U.S. Lunar Orbiter program took extensive photographs of the lunar far side. From these Lunar Orbiter photographs, detailed lunar maps have been made. Most of the major features of the lunar far side have been temporarily assigned numerical designations by the IAU, and the IAU is currently reviewing names that have been suggested for these features. At present, however, the majority of lunar far-side features are unnamed.

Most of the designations of named lunar features are of classical extraction, and on reference charts and maps, these named features are referred to in the classical context. Thus, the Sea of Tranquility is referred to as Mare Tranquillitatis, and the Hyginus Rille is designated Rima Hyginus. In this atlas, the classical denotation will be adhered to, although the translated or current nomenclature is provided when ambiguities might occur.

Strict geologic or other interpretive terms describing the features or terrain shown in the photographs have been avoided because it is thought that the photography in this volume is so significant that the atlas should reach and appeal to as wide an audience as possible. Therefore, captions generally have been restricted to the identification of known features and unique or peculiar formations or conditions.
Because the terminator, the line of demarcation between darkness and sunlight, was at most a few degrees west of the zero meridian during the Apollo 8,10 , and 11 missions, the photography from these missions is primarily of the eastern hemisphere of the Moon. As a result, much of this atlas is comprised of lunar far-side photography. Therefore, many of the photographs in this atlas are valuable for updating existing maps and charts of the lunar far side. In ap-


APOLLO 10
AS10-34-5014
The Earth as it appears from a distance of $36000 \mathrm{n} . \mathrm{mi}$. is shown here. North America is the dominant landmass, with the southwestern regions of the United States and the northern areas of Mexico clearly visible. At the lower right, the outline of the uppermost portion of South America can be distinguished.
pendix $A$, indexes of the photography have been prepared and are presented according to mission. For each frame in the atlas, the index includes a description of the area of coverage; coordinates of the principal point; focal length of camera lens; notations on the Sun elevation and the general type of view, whether vertical or oblique; and the page on which the photograph appears in this book. The maps used in this index were printed by the USAF Aeronautical Chart and Information Center, St. Louis Mo.

Because the orbital photography taken by all three missions was similar, integrating the photography from the missions into a coherent order seemed more appropriate than attempting to separate the photographs by mission. However, the mission from which a photographic frame was selected is easily determined by the frame number. The prefixes (AS8, AS10, or AS11) to the frame numbers identify a photograph as having been taken during Apollo 8, 10, or 11. For example, AS8-2606 was taken during Apollo 8, AS10-4433 during Apollo 10, and AS11-5903 during Apollo 11.

For the most part, the photography in this atlas has been oriented so that north is at the top of each page. Sometimes this orientation presents a view that is difficult for the eye to perceive. This situation occurs especially in high-oblique photographs. In such situations, the orientation most comfortable to the eye has been selected.

The onboard cameras for all three missions were modified Hasselblad 500 EL cameras, with $80-\mathrm{mm}$ and $250-\mathrm{mm}$ Zeiss panacolor lenses. For certain photographs of the lunar surface, a $60-\mathrm{mm}$ lens with a reseau was used. Use of this lens and reseau is apparent in the views that show crosslike fiducial marks. Although several emulsions (including Kodacolor) have been used experimentally on these missions, the photographs in this atlas were exposed on SO-368 and SO-168 (Ektachrome-type emulsions) for color and SO-168 (Plus X) and 3400 (Panatomic-X emulsions) for black and white.

For analytical purposes, black-and-white emulsions have been determined to provide a higher degree of resolution and image clarity than the color emulsions provide; therefore, much of the orbital photography is in black and white. However, most of the photographs taken on the lunar surface by astronauts Neil A. Armstrong and Edwin E. Aldrin, Jr., were in color.


AS8-14-2485
Top: Here the Moon appears in an orientation that is not seen by terrestrial observers. The eastern portion contains features of the lunar far side that until recently were unknown and, for the most part, remain unnamed. In the western portion, Mare Crisium (Sea of Crises) and just below it Mare Fecunditatis (Sea of Fertility) predominate. Near the center, the smaller Mare Marginus (Border Sea) and Mare Smythii (Smyth's Sea, named for the English admiral and astronomer, 1788-1865) are in sharp contrast to the surrounding terrain.

Right: This view of the Moon contains a level of detail similar to that obtainable from small terrestrial telescopes and provides an appearance not unlike that seen by the naked eye. The manner in which the blackness of space is emphasized when contrasted with an illuminated celestial object is of interest.



APOLLO 10
This high-oblique view of the lunar far side includes Mare Smythii in the lower left and Mare Moscoviense (Sea of Moscow) near the horizon at the center top. Note how the rugged terrain of the lunar landscape near Mare Moscoviense stands out in relief against the black background.


This near-vertical view of Mare Smythii was photographed at a high Sun elevation. The illumination conditions tend to flatten out the appearance of the landscape and make smaller craters appear as bright points of reflection.


None of the features in this region has been named; however, the keyhole-shaped crater was used for training the Apollo 8 crewmen in landmark tracking. The large crater is approximately 20 statute miles in diameter. Note the range in sizes of craters visible at the low $\left(7^{\circ}\right)$ Sun elevation.


Top left: This vertical photograph of the northern floor of a far-side basin was exposed near the subsolar point (point of maximum Sun elevation). The unnamed bright crater at the center is the focal point of the scene; thousands of tiny craters stand out as pinpoints of bright light.

Left: The southeastern edge of Mare Smythii is exposed at a high Sun elevation. emphasizing the large number of small craters existing in this region. The lineations running diagonally across the upper portion of the area also are of interest.


APOLLO 8
AS8-12-2192


APOLLO 8
AS8-12-2189
Top: This is a southerly looking high-oblique photograph of the Iunar far side. The dark-floored crater is approximately 40 miles in diameter and lies on the southeastern edge of Mare Australe (Southern Sea). Numerous other large craters can be seen in this region, including a very bright ray crater near the horizon at left.

Above: This is a high-oblique photograph of the lunar surface with the large crater Humboldt (German statesman, 1767-1835) at the lower left center and the crater Langrenus (Belgian selenographer, 1600-1675) at the far upper right. Numerous light areas and small bright-rayed craters are visible.


APOLLO 8


APOLLO 8
Top: A closer view of the Humboldt crater shows in detail the interesting linear patterns emanating from the center of the crater. The bright central peaks are characteristic of many lunar craters.

Above: This is an oblique view of the far-side crater Tsiolkovsky (Russian aerodynamics and rocketry scientist, 1857-1935). The crater Tsiolkovsky, approximately 94 miles in diameter, was first discovered by the Soviet Moon probe Lunik 3 in 1959. This feature particularly is distinguished by the bright central peak rising from the very dark crater floor.


APOLLO 8
AS8-12-2203
The crater Langrenus is located on the eastern edge of Mare Fecunditatis.
Langrenus, which is approximately 85 miles in diameter, shows the
classic features commonly associated with large craters, such as a
relatively smooth floor, terraced walls, and a central peak.


The crater Joliot-Curie (French physicist, 1900-1958), partially shown within this high-oblique photograph at left center, is approximately 110 miles in diameter. The dark-bottomed crater Lomonosov (a Russian founder of present-day natural science, 1711-65) is approximately 50 miles in diameter. Long narrow rays have been reported in the polar region of the Earth-facing hemisphere of the Moon. The very bright-rayed crater near the horizon is now thought to be the source of these rays.


APOLLO 8
AS8-13-2220
This is an oblique photograph (looking southwest) of Mare Fecunditatis. The large crater at the top is Bellot (French explorer, 1823-56), approximately 13 miles in diameter. The interesting double, or concentric, crater at the upper left is Bellot B.


APOLLO 8
In this oblique view, a portion of the Pyrenees Mountains can be seen in the center background at the top edge of the photograph. The large crater Goclenius (German scholar, 1572-1621) in the foreground lies on the southern edge of Mare Fecunditatis and measures approximately 45 miles in diameter. The numerous rilles scarring the floor of Goclenius can be seen; one rille extends across the entire crater floor, over the central peak, and across the rim into the smooth mare.


APOLLO 8
This view was taken looking southwest across the highlands just north of the Pyrences Mountains. The low Sun elevation emphasizes the relief. The large crater in the foreground is Lubbock D (English mathematician, 1803-85). In the middleground, the Gutenberg Rilles (German printer, 1398-1468) can be seen crossing both hills and craters. The craters in the background belong to the Capella (Carthaginian lawyer, ca. A.D. 450) system of craters.


APOLLO 8
AS8-13-2243
The large, heavily shadowed crater on the horizon is Fracastorius (Italian astronomer, 1483-1553). The shallow crater in the near foreground is Daguerre (French pioneer in photography known for his daguerreotypes, 1789-1851). The bright peak on the horizon at left center is approximately 270 miles from Daguerre.


APOLLO 8
AS8-13-2269
Top: The large crater in the center is Behaim (German navigator, 1436-1506). Of special interest is the very even appearance of this crater, especially in the center where the central peak resembles a smooth dome.

Right: This photograph is a westward-looking view into Mare Tranquillitatis (Sea of Tranquility). The two small craters in the lower right corner are Secchi A (Italian astronomer, 1818-78) and Secchi B. The large sharp-rimmed crater just beyond Secchi $A$ and Secchi $B$ is Taruntius $F$ (Roman philosopher, ca. 88 B.C.). The large crater remnant at the center background is Maskelyne F (English astronomer, 17321811).

APOLLO 8
AS8-13-2271



Top left: The prominent feature in the lower center is the crater Maskelyne $F$, an old ring crater which is 13 miles in diameter. In the center background near the terminator is the welldefined crater Maskelyne $H$. The distance between these two features is approximately 56 miles. The sharp protuberance near the center was selected as a control point because of its unique structure.

Left: This south-looking oblique view of the far-side terminator was exposed at a Sun elevation of $1^{\circ}$. The extreme contrasts of black and white lend a curious, almost artistic, quality to the scene. In the far background, a ridge apparently protrudes high enough to reffect sunlight, causing the bright, narrow, diagonal streak.


This oblique view of the northeastern portion of Mare Tranquillitatis depicts a rather wide variety of lunar topographic features. To the right center is the crater Cauchy (French mathematician, 1789-1856). The accompanying escarpment has been designated Rima Cauchy (Cauchy's Rille). Numerous small craters, peaks, and domes also prevail in this region.


This oblique view is looking northwest across the northern portion of Mare Tranquillitatis toward the near-side terminator. The Sun elevation was $1^{\circ}$. The large crater at the center of the picture is Vitruvius (Roman architect, 100 B.C.). Note the presence of a "ghost" crater just to the right of Vitruvius.


APOLLO 8
AS8-14-2399


Top left: This photograph is one in a sequence of vertical views of the central region of a far-side basin. The deep shadows near the terminator emphasize the relief of the area.

Left: Another in the sequence of vertical views, this heavily cratered area appears forbidding because of the accentuation of detail at the low Sun elevation.


APOLLO 8
The shadow cast by the large protuberance in this far-side view indicates
that this is a feature of considerable size. This area, as well as most of the other areas in this region, is heavily cratered.

Right: The large ridges running diagonally across the area are the predominant features. Numerous small gouges can be detected running at general right angles to the ridges. The heavy cratering seems typical of this far-side region.

Below: The lunar surface in this area shows a definite fibrouslike texture not generally seen in the other views of this far-side region. Also, this area does not seem to have been subjected to the heavy cratering evident in other portions of this region.


APOLLO 8
AS8-14-2410


APOLLO 8
AS8-14-2420
Above: The smooth appearance of this lunar far-side crater suggests that it may possibly be an older crater. Nevertheless, it exhibits a central peak and considerable terracing of its walls.


APOLLO 8
The large crater near the center of the photograph is approximately 10 miles in diameter. The apparent softness of the scene is belied by the presence of large numbers of small sharp-walled craters that are visible upon close inspection.


Above: The central highland area and a 5 -mile-diameter crater are evident in this low-oblique view of the central region of a large far-side crater. A recently formed bright crater and its associated ray system are located in the upper left region.

Right: Many of the smaller craters in this lunar far-side lowoblique photograph respond to the higher Sun elevation by assuming a bright appearance. Several such craters are located near the large crater at the center of the picture.
APOLLO 8
AS8-14-2439



APOLLO 8
Two craters apparently share an adjacent wall in this far-side view. The rim of the larger crater is more than 50 miles across and has been subjected to a steplike terracing. At this relatively high Sun elevation, numerous small bright craters can be seen in the area.


Top left: This oblique view of the crater Tsiolkovsky depicts the prominent central peak of this unusual crater. The contrast between the peak and the surrounding crater floor is especially striking.

Left: In this lunar far-side oblique photograph, exposed with a telephoto lens, the foreshortening effects provide an interesting portrayal of lunar far-side relief. APOLLO 8

AS8-14-2453


APOLLO 8
This is a photograph of earthrise as seen from lunar orbit. The visible
land area is the western portion of Africa, with heavy cloud patterns
concealing much of the Atlantic Ocean.


APOLLO 8
An oblique view of the crater Langrenus shows in detail the steep terracing of the inner walls and the smooth crater floor broken by the central peak. The large shallow crater to the upper left is Vendelinus (Belgian astronomer, 1580-1667). The fact that Vendelinus is smooth and worn as compared with Langrenus would suggest that Vendelinus is
a much older crater.

Top right: These unnamed far-side craters are located near the terminator. The large crater near the center of the photograph is approximately 18 miles in diameter. Because of the low Sun elevation, much of the detail in this region is obscured by shadow.

Right: This is the floor of a far-side crater that was described by the Apollo 8 lunar module pilot as being an area of possible flows. The heavy shadows on the basin floor are from the crater rim. A number of small, rolling, hummocklike structures can be seen to the lower right of this crater.


APOLLO 8
AS8-17-2664



This well-defined crater chain is located on the lunar far side. Much
of the surlace surrounding this crater chain seems heavily scarred and pockmarked, and linear gouges run diagonally across the area.

Topright: An interesting feature located near the center of this lunar far-side view is the large keyhole-shaped crater. This is a particularly rugged region with considerable relief and heavy cratering.

Right: The sharp line of demarcation just to the left of center in this view is the rim of a large, unnamed, far-side crater. The sloping crater wall appears to be very bright because the wall reffects more sunlight than the surrounding terrain.



The crater near the center of this view stands out sharply against the surrounding terrain. This is a crater within a much larger far-side crater. The central peak of the larger crater is seen near the small bright crater at the left, while a portion of the terraced walls that form the larger crater's boundary can be seen at the lower right.


APOLLO 8


Top: This is a high-oblique westward view looking toward what would become Tranquility Base in Mare Tranquillitatis. The sea at the bottom of the picture is Mare Fecunditatis, and the sea extending to the horizon is Mare Tranquillitatis. The highland separating the two seas is the Secchi Peninsula.

Above: The large, unnamed, far-side crater at the lower center is approximately 30 miles in diameter. Terracing and a central peak, features common to other large craters of this type, are shown in the photograph.
The mottled appearance of the surrounding area is produced by numerous small, bright, halo craters that stand out against the darker upland surface.


APOLLO 8
AS8-17-2748
The large crater in the left-central foreground of this lunar far-side oblique photograph is approximately 60 miles in diameter. The relatively high Sun angle causes a large number of craters near the horizon to stand out as bright streaks or spots. The sharp relief of this region can be seen outlined against the black horizon.


APOLLO 8
AS8-17-2776
Above: A number of large craters are evident in this far-side view. The crater near the lower center is approximately 15 miles in diameter. A portion of a much larger, dark-appearing crater can be seen at the lower left.

Right: This view is directly south of Mare Smythii. The crater at the top center is approximately 20 miles in diameter. The high Sun angle lends a luminescent quality to the outer walls of the larger craters. The lip on the side of the crater at the lower left is also of interest.



This photograph was taken from the Apollo 10 lunar module, which was being inspected by the command module pilot. The lunar background in this view is a portion of the limb region east of Mare Smythii.

Top right: In this photograph of Sinus Medii (Central Bay), the two prominent craters are Bruce (American patron of the arts, 1816-1900) and Blagg (English selenographer, 1858-1944). Bruce (the crater at the top) is approximately 4 miles in diameter, and Blagg (just below Bruce) has a diameter of approximately 3 miles. The topography on the surface of Sinus Medii is accentuated by the low Sun elevation.

Right: Bruce is the prominent crater near the bottom. The low Sun elevation (ranging from nearly $6^{\circ}$ in the east to less than $1^{\circ}$ in the west) emphasizes the undulations occurring in an apparently smooth mare area.


APOLLO 10
AS10-27-3905


APOLLO 10
AS10-27-3907



Top left: This unnamed far-side crater, approximately 26 miles in diameter, shows terracing on its inner walls. It also has an unusual break (at the right rear) in the outer rim. The peculiar shade variations and black spots are, at present, unexplained.
Left: The diameter of this far-side crater is approximately 15 miles. While the crater appears like many others of this size and age, the surrounding terrain appears quite wrinkled. Under the high Sun elevation, numerous small craters appear as pinpoints of light.


Left: This view of the south-central portion of Mare Tranquilitatis illustrates the differences in the character of the mare and the highland regions. (Compared with the highland areas, the mare appears quite smooth.) The small, clearly defined crater at left center is Maskelyne $T$, which is approximately 4 miles in diameter.

Below: This is a high-oblique view. looking east across Mare Tranquillitatis. The landing site for Apollo 11 is near the bottom center of the picture. The crater in the lower right is Moltke (Prussian general, 1800-1891). The large crater at the upper lef is Maskelyne.

APOLLO 10


APOLLO 10
AS10-28-4040
Above: This near-vertical view was taken over the southern edge of Mare Tranquillitatis. The small bright-rayed crater at the bottom right corner is Censorinus (Roman grammarian and mathematician, ca. A.D. 238). The larger crater on the right side of Censorinus is Censorinus A. The diameter of Censorinus is approximately 5 statute miles. Even at a medium Sun elevation, the ray structure associated with Censorinus is quite apparent.


APOLLO 10
AS10-28-4052


APOLLO 10
AS10-28-4067
This large far-side crater is approximately 40 miles in diameter. The
three smaller craters symmetrically situated on the lower third of the rim are of interest. The texture of the crater walls is readily seen because
of the low Sun angle.

Right: The large unnamed crater at the center of this far-side oblique photograph is approximately 100 miles in diameter. The rough terrain in this region results from heavy cratering. A bright-rayed crater can be seen on the south-central rim of the large crater.

Below: The unusual alinement of these three far-side craters made them especially useful as landmark identification points. The large crater at the right, which is approximately 25 miles in diameter, has a rough floor and considerable terracing, but the crater has no welldefined central peak.


AS10-29-4180



APOLLO 10
AS10-29-4183


APOLLO 10

AS10-29-4189

Above: The rectilinear appearance presented by the large crater in this far-side oblique view is unusual. From the smooth appearance, this crater appears to be much older than the bright, sharply defined crater near its center. A series of hummocky protuberances, located to the right of the large crater at rear center, can also be seen.
Left: This crater (visible at a distance in the preceding photograph) is approximately 20 miles in diameter. Close inspection of the crater shows a distorted lip on the left side. Terracing is also quite evident in this feature. The crater is located on the edge of a far-side unnamed basin, and the rugged terrain (generally associated with highland regions) can be seen at the left and top of this view.


APOLLO 10

Left: The unnamed far-side craters in this rather unusual grouping are from approximately 20 to 35 miles in diameter. A small distinctive crater is shown at the lower right edge of the crater at the top.
Below: A large, unnamed, far-side crater with a smaller crater on its outer rim is shown in the foreground of this oblique photograph. A small bright crater (left center) and a recencly formed crater with a distinctive ray pattem (right center) can be seen.


APOLLO 10


APOLLO 10
AS10-29-4226
This far-side oblique view shows the eastern edge of Mare Smythii
(the dark area at the central horizon). A large bright crater is predominant
in the foreground, and a mountainous ridge isolates a basin from the surrounding terrain at right center. Numerous small, bright-rayed craters also can be seen in this region.

Right: This low-oblique view is of a small area located within Mare Smythii. The area shows several unusual features. The large trenchlike feature at the left extends from a multiple-ringed structure, which lies out of view, into a basinlike area at center. Numerous bright-rayed craters and craterlets can be seen in both the basin and in the highland areas. A large crater is visible at top center.

Below: In this view of Mare Smythii, the high Sun elevation provides a sharp contrast between prominent features and their surrounding background. A sinuous rille, winding across the center of the region, can be seen clearly. At right center, a large bright crater can be seen, and small ray craters are scattered throughout the area.


APOLLO 10



AS10-29-4253
Above: This oblique photograph provides a closeup view of Messier B (French astronomer, 1730-1817), a small lunar crater shown at center. The shadowed portion of the crater shows that the steep interior walls are subject to considerable tonal variations, because of differences in reflected light. It can also be seen that the crater has outer walls that slope off gently into the surrounding mare.

Left: In this low-altitude, oblique photograph, the three predominant craters are Messier at left center, Messier A just to the right of Messier, and Messier $B$ in the foreground. These relatively small craters range from 4 to 8 miles in diameter. Topographically higher features on the horizon are contrasted against the black sky.
APOLLO 10
AS10-29-4256

Top right: This unique double crater is Messier A. There is an apparent break in the common wall between the older crater toward the top and the more recently formed crater in the foreground.

Right: This is a view of Secchi K, which is approximately 5 miles in diameter and is located in Mare Fecunditatis. A small bright-rayed crater can be seen at the bottom edge of Secchi K . The highland regions that mark the boundary of the mare are seen on the horizon. At the lower left is the shadow of a thrustor nozzle of the lunar module.

APOLLO 10
AS10-29-4261


APOLLO 11
AS11-38-5602



APOLLO 10
AS10-29-4265
The bright crater at the top right in this oblique photograph is Secchi UA which is located in the western portion of Mare Fecunditatis. The bright highlands can be readily distinguished from the darker mare. A broad rille to the right of center runs linearly across the area.


APOLLO 10
This low-oblique photograph was taken from the Apollo 10 lunar module during the descent approach to the Apollo 11 lunar landing site in Mare Tranquillitatis. The clongated hill in the center of the photograph is near Secchi B and is approximately 780 m above the surrounding mare floor. The shadowed hill at lower left provides an indication of the texture of these mare hills.


APOLLO 10
AS10-29-4324


APOLLO 10
This low-altitude oblique photograph of the central portion of a 50 -milediameter far-side crater provides a closeup view of the interior structure of such craters. The prominent mountainous formation at right center is the central peak of the crater. The crater floor is heavily pockmarked and contains many hummocky protuberances near the top center.
To the left, terracing of the inner crater walls is evident.


APOLLO 10
AS10-30-4371
In this view of a far-side double crater, the smaller crater gives the appearance of having been scalloped out of the lunar surface. The larger. older crater appears worn in comparison and does not display the sharp terracing of the smaller crater. The terrain in the background appears wrinkled under the relatively high Sun elevation.


APOLLO 10
AS10-30-4372
The large, brilliant ray structure of this relatively small far-side crater implies that the crater is of fairly recent origin. It is thought that rays such as these are formed by material ejected when a meteorite impact forms a crater and that the material ejected has not yet been subjected to the effects of solar erosion and darkening.


Taruntius $A$, the crater in the center of the photograph, is approximately 10 miles in diameter and is located in the northern portion of Mare Fecunditatis. Near the edge of the mare, a linear rille and a small bright-rayed crater (lower right) can be seen.


The large crater in the foreground is Manners (English naval officer, 1800-1870), and the smaller one at bottom center is Arago B (French astronomer, 1786-1853). The double crater Ariadaeus (Macedonian king, died 317 B.C.) and Ariadaeus A are at the terminus of the sinuous Rima Ariadaeus (Ariadaeus Rille) in the center background. Note how the rille in the vicinity of the double crater Ariadaeus appears to be partially filled in.


Above: This is a near-vertical view of a northern portion of Mare Fecunditatis. The highlands at the upper left are part of the rugged terrain that separates Mare Fecunditatis from Mare Crisium.

Right: The crater at right center, Taruntius G, is approximately 5 miles in diameter and is located in the northern portion of Mare Fecunditatis. The roughness of the mare is evident by the presence of numerous wrinkle ridges and small rilles rumaing diagonally across the area.



Top left: The bright crater at the tip of the highland peninsula is Secchi $\theta$, located in the eastern portion of Mare Tranquillitatis. The highest point on this peninsula is approximately 1300 m above the mare floor. A pronounced gouge in the mare surface is evident at the left, and numerous craters are grouped together in a variety of formations.

Left: This is a near-vertical view of the southern edge of Mare Tranquillitatis The semicircular feature in the upper left corner is Maskelyne $D$, a significant landmark on the approach to Tranquil ity Base. Maskelyne D is sometimes re ferred to as "Bob's Bend"; similarly, the peninsular-shaped landmass has been referred to as "Barbara Mesa." These are operational nicknames given to distinc tive or unusual landmarks by the Apollo 10 astronauts.


APOLLO 10
AS10-31-4546
The sharply defined crater at right center is Theon Senior (Greek astronomer, ca. A.D. 100), approximately 6 miles in diameter. At the upper left, a portion of the crater Delambre (French astronomer, 1749-1822) can be seen. This highland area is located west of Mare Tranquillitatis.


Above: This 40 -mile-diameter crater is Taruntius, located in Mare Fecunditatis. The smaller, bright crater on the rim is Taruntius C. Taruntius has a central peak, terracing, and some interesting arcuate rilles near the center.

Left: The hook-shaped feature in this view is the ring crater Maskelyne F. From the appearance, Maskelyne $\mathbf{F}$ is probably a very old crater that was formed before Mare Tranquillitatis was fully developed and had achieved its present level.


This view provides an indication of the size of Rima Hypatia I. Just beyond the crater Moltke (center) the rille forks; one branch continues through the mare and the other branch crosses the highlands that mark the southern boundary of Mare Tranquillitatis.


APOLLO 10
AS10-31-4621
Above: A parallel rille pattern is the outstanding feature of this oblique view of the southwestern portion of Mare Tranquillitatis. The broad rille at left is Rima Hypatia I, and the rille passing through the center of the area is Rima Hypatia II. At the top left is the crater Hypatia E.

Right: The large crater in the center of the photograph is Arago. Arago is located in the western portion of Mare Tranquillitatis and is approximately 18 miles in diameter. Terracing, characteristic of many lunar craters, is evident in Arago.


APOLLO 10
AS10-31-4630


APOLLO 10
The broad linear rille in this unusual oblique view is Rima Ariadacus,
which is approximately 3 miles wide. The large crater below the rille is
Silberschlag (German astronomer, 1721-91), which is approximately 9 miles in diameter. As can be seen from the photograph, the rille crosses a variety of terrain (ranging from mare to highland surfaces).


## APOLLO 10

AS10-31-4647
The large crater at center is Godin (French explorer and mathematician,
1704-60). This crater, which is approximately 27 miles in diameter,
is located in the highland region that separates Mare Tranquillitatis
from Sinus Medii.


APOLLO 10
AS10-31-4654
The large crater in this far-side oblique photograph is approximately
60 miles in diameter. Of particular interest are the crater chains occurring in the floor of this crater and in the rough area at top center. Two other large craters can be seen near the horizon at top left and top right.


APOLLO 10
AS10-31-4665


APOLLO 10


Above: The frying-pan-shaped crater in this oblique photograph is located on the outer rim of a large, unnamed, far-side crater. This frying-pan-shaped crater has a relatively flat floor that has apparently been subjected to fracturing.

Right: A distinct crater chain can be seen running from the top to the bottom of this far-side oblique photograph. Crater chains of the size shown in this photograph are rare on the lunar surface and interpretations as to their cause (i.e., impact or volcanism) vary. At the center and to the right of the crater chain is a large, recently formed crater with a smaller crater located on its outer rim.

AS10-33-4975


APOLLO 10
AS10-33-4914


APOLLO 10
The two craters seen in this low-altitude oblique photograph are Theophilus (Saint and Bishop of Alexandria, ca. A.D. 412) in the center foreground and Cyrillus (Saint, A.D. 444) on the horizon behind Theophilus. Both craters are quite large (approximately 65 miles in diameter), and both have extensive terracing of their side walls and prominent central peaks.


APOLLO 10
AS10-32-4734


Above: The crater in the center of this oblique view is Chladni (German physicist, 1756-1827), located at the southern tip of the highlands extending into Sinus Medii. The crater that is partially visible at the lower right is Triesnecker (Austrian astronomer, 1745-1817). A highland mountain range is illuminated under the low Sun elevation and stands out against the blackness of the terminator.

Left: This near-vertical photograph of the highland region between Sinus Medii and Mare Tranquillitatis shows, at bottom, a portion of the large crater Lade (German selenographer, 1817-1904). The smaller crater, inside the western rim of crater Lade, is Lade M. Godin B is at the top center. A concentration of small craters can be seen on the upper portion of the rim of Lade.


This near-vertical view shows a highland area that lies between the crater Lade and Sinus Medii. The low Sun angle, which emphasizes the topographic features of the area, clearly reveals a surface pockmarked by thousands of small craters. This view lends credibility to the hypothesis that the lunar surface is continually bombarded by meteorites.


APOLLO 10
AS10-32-4819


Top left: The crater Triesnecker (Austrian astronomer, 1745-1817), approximately 17 miles in diameter, and its associated network of crisscrossing rilles (the Triesnecker rilles) are located in the northeastern portion of Sinus Medii. This extensive series of rilles extends beyond Sinus Medii into the smooth floor of Mare Vaporum at upper right. The large crater at the edge of the highlands above Triesnecker is Ukert (German historian, 1780-1851).

Left: This oblique view is centered on Sinus Medii, with the crater Hyginus (Spanish astronomer, ca. A.D. 100) and the sharply defined Rima Hyginus (Hyginus Rille) to the right. The crater Hyginus, located where the rille makes a bend, lies near the northeast margin of Sinus Medii and is approximately 6 miles in diameter. From the crater, Rima Hyginus extends east-southeast toward Mare Tranquillitatis and northwest toward Mare Vaporum (Sea of Vapors). The rille is approximately 2 miles wide and more than 130 miles long. The rilles visible at center left are the Triesnecker rilles. The crater Triesnecker is just out of view at center left.


APOLLO 10
Above: 'The diameter of the large crater (at upper right) in this farside view is approximately 65 miles. This crater is typical of the more recently formed, large, far-side craters. A bright-rayed crater is located above and to the left of the central peak of the large crater. In the center are two adjacent craters with peculiar fracture patterns in their floors. These peculiar patterns imply that these craters have been subjected to volcanic activity.

Right: The heavy shadows, caused by the low Sun elevation, give an indication of the roughness of this far-side region. The terraces of the large crater at the top of the photograph are distinctly steplike in appearance.
APOLLO 10
AS10-32-4823




APOLLO 10
AS10-34-5073
The area shown in this view is on the southern edge of Mare Tranquillitatis.
The crater at top left is Moltke. Rima Hypatia I is the rille that extends
diagonally across the photograph. The landing site for Apollo 11 is
approximately $15 \mathrm{n} . \mathrm{mi}$. northwest of this area.


This photograph of the eastern limb is a high-oblique view of the crater Neper (Scottish mathematician, 1550-1617) looking west-northwest. Mare Smythii is in the lower left corner. Neper is at upper center, and Mare Marginus is at upper right.


APOLLO 10
This is a high-oblique view of a portion of Mare Crisium and the adjoining
highland. Mare Crisium is the dark mare in the upper right of the photograph, and the crater Picard (French astronomer, 1620-82) is the most prominent feature visible in that mare.


APOLLO 10
AS10-34-5099
Top: This is a view of the approach to Tranquility Base, which is just out of view in upper right center. The Apollo 10 crew gave nicknames (or code names) to most of the prominent features shown in this photograph. The features shown were given nicknames such as Thud Ridge, The Gashes, Faye Ridge, Diamondback Rille, Sidewinder Rille, and Last Ridge.

Right: This low-oblique photograph was taken on the approach to Tranquility Base, approximately 55 n . mi. east of the landing site. It shows Diamondback Rille connecting with an elongated crater. The circular crater at right center is Maskelyne $\mathbf{X}$, which is approximately 2 miles in diameter.


Apollo 10
This is a view of a northern segment of Diamondback Rille, which shows the rille branching off into a fork. The crater at upper left center is
Maskelyne G. A footprint-shaped crater is located to the left of Maskelyne G.



Above: This crater, located in Mare Spumans (Foaming Sea), has a very bright ray structure. It is generally believed that such brilliant ray patterns are associated with craters that have been formed relatively recently.

Left: The largest crater in this view of a part of Mare Tranquillitatis, at left center, measures approximately $11 / 4$ miles in diameter and is 13 miles southeast of a proposed Apollo program touchdown point. The small crater on its rim is used as a navigation landmark.


Top: The large bright crater in the center of this view of Mare Fecunditatis is Taruntius H. This crater, approximately 8 miles in diameter, appears very bright at high Sun elevations. Tonal variations can be detected on the crater slopes.

Above: The large crater at upper right is Taruntius $\mathbf{F}$, approximately 9 miles in diameter and located in the eastern portion of Mare Tranquillitatis. The texture of the crater walls is clearly visible in this view. At lower center are several hummocky features.


APOLLO 10
AS10-34-5160
This is the crater Ritter (German geographer, 1779-1859), located on the western edge of Mare Tranquillitatis and approximately 18 miles in diameter. The interior floor of Ritter shows sharp fracturelike lineaments, which are emphasized by the heavy shadows cast at low Sun elevations.


Apollo 10
This is the crater Schmidt (German selenographer, 1825-84), located at the western edge of Mare Tranquillitatis, south of crater Ritter and west of Sabine, and approximately 10 miles in diameter. The most notable features of Schmidt are the sharp rim, ray pattern, relatively rough floor, and surrounding hummocky terrain. Numerous boulders, easily resolvable on a photograph of this scale and quality, can be seen in the area surrounding the crater and on the crater floor. Most of these boulders range in size from 68 to 122 m .


APOLLO 10
AS10-34-5167
This view shows a portion of the highlands between Mare Tranquillitatis and Sinus Medii. The crater at right center is Godin D. The rough terrain of the highlands area is evident under the low Sun elevation.


APOLLO 10
The large crater in this lunar far-side view is approximately 100 miles in diameter. Terracing and a prominent central peak are evident. The relief, contrasted against the black sky, shows the ruggedness of the surrounding terrain.


APOLLO 10
AS10-34-5172
Above: The sharply defined, steep-walled crater in the center of this far-side oblique view is approximately 12 miles in diameter. Rubble, probably ejecta material, can be seen on the outer slopes surrounding the crater. A larger, more subdued crater is located above the steep-walled crater.

Right: This far-side oblique view is dotted with several distinct, bright craters. The photograph provides a representative view of a rugged area of the lunar surface.


APOLLO 10
AS10-34-5173


APOLLO 11
AS11-44-6581
The Apollo 11 lunar module Eagle is shown as it appeared following separation from the command module. Separation was performed in preparation for the descent to the lunar surface. The rods protruding from the footpads are touchdown sensors that shut off the descent engine automatically when contact is made with the lunar surface.


Top left: In this view, the command module, located just right of center, nearly blends into the background that shows the surface of Mare Tranquillitatis. Moltke is the large crater at lower right, and Rima Hypatia I crosses the area in the lower left corner of this photograph.

Left: In this photograph, the command module is at the southeast rim of the crater Schmidt (top center). Portions of craters Sabine (top right) and Ritter (above and to the left of Sabine) are also visible. Almost the entire spectrum of lunar features can be seen in this view: smooth mare, rugged highlands, large and small craters, and well-defined rilles.


This westward-looking low-oblique view is centered approximately on Tranquility Base, which is located at the terminator. The major crater in the lower right foreground is Maskelyne, with Maskelyne B just above. Torricelli C (Italian physicist, 1608-47) is the crater above the shadow of the lunar module thrustor nozzle on the left side of the photograph; and Moltke lies near the center of the photograph, near the terminator and just to the right of Rima Hypatia.

Top right: This view was taken before the Apollo 11 astronauts began their extravehicular activity. The area nearby is pockmarked with many small craters, and numerous boulders are strewn about the surface. Note the coarse, granular texture of the lunar surface.

Right: Looking southwest from Tranquility Base, the lunar surface appears relatively smooth, even under a low Sun elevation. The large crater has smooth undulations and the gentle rim lines that are associated with older craters. However, some of the minute craters seen are steep-walled and covered with small pebbles, features which perhaps denote a more recent origin.


APOLLO 11
AS11-37-5456


APOLLO 11


APOLLO 11
AS11-37-5517
In the foreground is the U.S. flag. The brace on which the flag was mounted was bent slightly, giving the flag the appearance of being unfurled in a breeze. In the center background is the television camera mounted on its stand; the connecting cable can be seen at center foreground. The surface around the flag and the television camera has been disturbed by the footprints of the first men on the Moon, and the imprints of the cleated soles of the astronaut's boots are clearly evident.


APOLLO 11
AS11-37-5549


Right: Astronaut Edwin E. (Buzz) Aldrin, Jr., descends the ladder of the lunar module to the lunar surface for the first time.

Below: Astronaut Aldrin stands near the American flag. Footprints and the television camera cable are visible in the foreground.

APOLLO 11
AS11-40-5875



APOLLO 11
AS11-40-5877
Top: This almost-perfect outline of the traction cleats on the astronauts' boots implies that the lunar soil has a very fine texture. A more coarse material would not have permitted the appearance of such a fine detail.

Above: This overhead view shows a detailed section of the lunar surface and the amount of compaction caused by the weight of a man.



APOLLO 11
Top: To the right of the center, the television camera keeps a lonely vigil on man's first activities on the lunar surface. The paths made by the astronauts have uncovered a portion of the lunar surface that is darker than the undisturbed surrounding region.

Above: This view looks south across a portion of the rim of a $120-\mathrm{ft}$-diameter crater. The interior slopes of this crater are strewn with rocks and boulders. The intense blackness of shadows on the lunar surface can be seen from the shadows cast by the crater rim in the foreground of this photograph.


APOLLO 11
AS11-40-5902
Astronaut Aldrin stands looking at the lunar module at Tranquility Base.
The lack of a pronounced footpad imprint demonstrates the overall firmness of the surface. The scuffed footprints show the extremely fine granular texture of the lunar soil.


APOLLO 11

The image of astronaut Neil Armstrong photographing astronaut Aldrin is clearly seen in the reflections on the helmet visor of astronaut Aldrin. Astronaut Armstrong, the lunar module, and the shadow cast by Aldrin are mirrored in the reflection.


APOLLO 11
AS11-40-5921
Above: In this view, the lunar surface beneath the descent engine nozzle (top center) is shown. The area has been swept clean of loose, unconsolidated material by the engine exhaust. Small radial striations emanate outward from the exhaust center, indicating the path taken by the expelled surface material.

Right: A small cluster of rocks, some smooth and others sharply angular, can be seen near the center of the landscape. The television camera mounted on its stand and the power cable leading from the camera to the lunar module are seen in the background. A crater is at top right on the horizon.


APOLLO 11


APOLLO 11
AS11-40-5926


APOLLO 11

Right: Astronaut Aldrin is shown in this view in the process of deploying the seismometer and the laser-ranging retroreflector experiment. This area is strewn with boulders and is the site of several large craters. At the right are the footprints Aldrin made as he carried the experiment equipment. The footprints are deeper than those seen in other photographs, probably because of the added weight of the equipment.

Below: The rather uniform cleavage between the two large rocks in the foreground indicates that they may have previously been a single large rock. In the background is a small crater with footprints around it, and at the extreme left center is the stereo camera that was used to photograph lunar sample locations.


APOLLO 11
AS11-40-5932


APOLLO 11
AS11-40-5945
Astronaut Aldrin is shown in this view deploying the seismometer and the laser-ranging retroreflector.


Left: This view shows Tranquility Base and the arrangement of the scientific experiments. From left to right are the television camera, the American flag, the laser-ranging retroreflector, and the stereoscopic camera. The lunar seismic experiment equipment is visible in the foreground.

Below: Astronaut Aldrin is deploying the solar panels of the passive seismic experiment package. The laser-ranging retroreflector has already been deployed in the background. To the right of the laserranging retroreflector and in front of the lunar module is the stereo camera. On the horizon, just left of center, the television camera is visible.
APOLLO 11
AS11-40-5950



APOLLO 11
AS11-40-5952
This is a closeup view of the laser-ranging retroreflector experiment
The purpose of this experiment is to reflect back to Earth the laser beams directed at the reflector from Earth. From data obtained from this experiment, more precise Earth-to-Moon distances can be determined.

Top right: In this view of the central portion of a $120-\mathrm{ft}$-diameter crater, a concentration of rubble and large debris can be seen. This is probably material that was not ejected during the crater-formation process.

Right: This view of the western portion of the 120 -ft-diameter crater shows a fairly large crater at left center on the edge of the rim. Two other craters of considerable size are also visible; one crater is just inside the rim at top center, and the other crater is at top left outside the rim. In the foreground is the stereo camera.


APOLLO 11
AS11-40-5956


APOLLO 11


APOLLO 11
AS11-40-5963
Here, a core sample tube is being pounded into the lunar surface by
astronaut Aldrin. Next to the core sample tube is the mast that is holding the solar-wind foil panel. The television camera is at the upper left.


This is the plaque that was left on the Moon proclaiming that the Apollo 11 astronauts "came in peace for all mankind."


APOLLO 11
Panorama A
This panorama was taken from the lunar module before extravehicular activity was begun.


APOLLO 11
Panorama A
This panorama was taken from the lunar module after the historic "Moon walk."


APOLLO 11
Panorama B
This panorama was taken looking east from the lunar surface.


APOLLO 11
Panorama B
This panorama, taken from the lunar surface, shows the large crater that was approximately 20 ft east of the lunar module.


APOLLO 11
Panorama C
This panorama was taken looking northwest from the lunar surface.


APOLLO 11
Panorama C
This panorama was taken looking north from the lunar surface.


Right: The ascent stage of the lunar module Eagle is shown preparing to rendezvous with the command module Columbia.

Below: Eagle prepares to dock with Columbia after man's first successful landing on the Moon.


AS11-44-6623


APOLLO 11

## APPENDIX A PHOTOGRAPHIC INDEXES

## APOLLO 8



| Frame no. | Description | Principal point |  | Focal length, mm | Sun angle |  |  | View |  | $\begin{gathered} \text { Page } \\ \text { no. } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Longitude | Latitude |  | High | Medium | Low | Vertical | Oblique |  |
| AS8-12-2052 | Keyhole-shaped crater. | 157 W | 5 S | 80 |  |  | X | X |  | 6 |
| 2148 | Far-side basin.---- | 106 E | 10 S | 80 | X |  |  | X |  | 7 |
| 2169 | Mare Smythii | 90 E | 8 S | 80 | X |  |  | X |  | 7 |
| 2189 | Humboldt and Langrenus. | 72 E | 26 S | 80 |  | X |  |  | X | 8 |
| 2192 | Mare Australe.----- | 100 E | 40 S | 250 |  | X |  |  | X | 8 |
| 2193 | Humboldt. | 90 E | 27 S | 250 | X |  |  |  | X | 9 |
| 2196 | Tsiolkovsky | 130 E | 21 S | 250 | X |  |  |  | X | 9 |
| 2203 | Langrenus. | 68 E | 9 S | 250 |  | X |  | X |  | 10 |
| 2209 | Joliot-Curie | 100 E | 24 N | 250 | X |  | X |  | X | 11 |
| AS8-13-2220 | Mare Fecunditatis and Bellot. | 49 E | 12 S | 80 |  |  | X |  | X | 12 |
| 2225 | Goclenius | 45 E | 11 S | 80 |  |  | X |  | X | 13 |
| 2228 | Lubbock D and Gutenberg Rilles | 38 E | 6 S | 80 |  |  | X |  | X | 14 |
| 2243 | Fracastorius and Daguerre.---- | 35 E | 15 S | 80 |  |  | x |  | X | 15 |
| 2269 | Behaim.- | 79 E | 18 S | 80 |  | X |  |  | X | 16 |
| 2271 | Taruntius F- | 40 E | 4 N | 80 |  |  | X |  | X | 16 |
| 2279 | Maskelyne F | 35 E | 4 N | 80 |  |  | X |  | X | 17 |
| 2314 | Far-side terminator | 150 W | 12 S | 80 |  |  | X |  | X | 17 |
| 2344 | Rima Cauchy | On ho |  | 80 |  |  | X |  | X | 18 |
| 2347 | Vitruvius. | In sp | pace | 80 |  |  | X |  | X | 19 |
| AS8-14-2383 | Earthrise |  |  | 250 | X |  |  |  | X | 27 |
| 2399 | Far-side basin_ | 155 W | 3 S | 250 |  |  | X | X |  | 20 |
| 2401 | Far-side basin | 156 W | 3 S | 250 |  |  | x | X |  | 20 |
| 2409 | Far-side basin | 162 W | 4 S | 250 |  |  | X | X |  | 21 |
| 2410 | Far-side basin | 163 W | 6 S | 250 |  |  | X | X |  | 22 |
| 2412 | Far-side basin | 165 W | 7 S | 250 |  |  | X |  | X | 22 |
| AS8-14-2420 | Far-side crater | 175 W | 10 S | 250 |  | X |  | X |  | 22 |
| 2423 | Far-side crater | 180 W | 8 S | 250 |  | X |  | X |  | 23 |
| 2433 | Far side, central highland. | 161 E | 10 S | 250 |  | X |  |  | X | 24 |
| 2439 | Far-side craters--------- | 150 E | 12 S | 250 | X |  |  |  | X | 24 |
| 2442 | Far side, adjacent-walled craters | 137 E | 12 S | 250 | X |  |  |  | X | 25 |
| 2451 | Tsiolkovsky- | 128 E | 21 S | 250 | X |  |  |  | X | 26 |
| 2453 | Far side-- | 113 E | 12 S | 250 | X |  |  |  | X | 26 |
| 2485 | Lunar disk |  |  | 250 |  |  |  |  |  | 3 |
| AS8-16-2616 | Langrenus. | 61 E | 9 S | 250 |  | X |  |  | X | 28 |


| Frame no. | Description | Principal point |  | Focal length, mm | Sun angle |  |  | View |  | $\begin{aligned} & \text { Page } \\ & \text { no. } \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Longitude | Latitude |  | High | Medium | Low | Vertical | Oblique |  |
| AS8-17-2664 | Far-side craters. | 157 W | 4 S | 80 |  |  | X | X |  | 29 |
| 2670 | Far-side crater. | 162 W | 7 S | 80 |  |  | X |  | X | 29 |
| 2673 | Far-side crater chain. | 166 W | 6 S | 80 |  |  | X |  | X | 30 |
| 2676 | Keyhole-shaped crater | 170 W | 0 | 80 |  |  | X | X |  | 31 |
| 2697 | Far-side crater | 170 E | 11 S | 80 |  | X |  |  | X | 31 |
| 2704 | Far-side crater | 164 E | 10 S | 80 |  | X |  | X |  | 32 |
| 2736 | Far-side crater | 127 E | 12 S | 80 | X |  |  |  | X | 33 |
| 2748 | Far-side crater | 110 E | 10 S | 80 | X |  |  |  | X | 34 |
| 2776 | Far-side craters | 93 E | 9 S | 80 | X |  |  |  | X | 35 |
| 2785 | South of Mare Smythii. | 81 E | 9 S | 80 | X |  |  |  | X | 35 |
| 2814 | Mare Tranquillitatis and Mare Fecunditatis |  |  | 80 |  |  | X |  | X | 33 |



| Frame no. | Description | Principal point |  | Focal length, mm | Sun angle |  |  | View |  | $\begin{gathered} \text { Page } \\ \text { no. } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Longitude | Latitude |  | High | Medium | Low | Vertical | Oblique |  |
| AS10-27-3873 | Command module | 85 E | 0 | 250 | X |  |  |  | X | 36 |
| 3905 | Sinus Medii | 4 E | 1 N | 250 |  |  | X |  | X | 37 |
| 3907 | Sinus Medii. | 1 E | 1 N | 250 |  |  | X |  | X | 37 |
| 3915 | Mare Smythii | 88 E | 3 S | 250 | X |  |  |  |  | 5 |
| 3929 | Mare Smythii and Mare Moscoviense |  |  | 250 |  |  |  |  |  | 4 |
| 3955 | Moon |  |  | 250 |  |  |  |  |  | 3 |
| AS10-28-4012 | Far-side crater | 123 E | 5 S | 250 | X |  |  |  | X | 38 |
| 4013 | Far-side crater | 124 E | 3 S | 250 | X |  |  |  | X | 38 |
| 4035 | Maskelyne T. | 37 E | 0 | 80 |  | X |  | X |  | 39 |
| 4040 | Censorinus. | 32 E | 0 | 80 |  | X |  | X |  | 39 |
| 4052 | Tranquility Base_ | 26 E | 1 N | 80 |  | X |  |  | X | 39 |
| 4067 | Far-side crater- | 172 E | 0 | 80 | X |  |  | X |  | 40 |
| 4106 | Far-side craters | 133 W | 1 N | 80 | X |  |  | X |  | 41 |
| AS10-29-4180 | Far-side craters | 149 E | 7 S | 80 |  | X |  |  | X | 41 |
| 4183 | Far-side crater | 142 E | 2 N | 80 | X |  |  |  | X | 42 |
| 4189 | Far-side crater | 139 E | 2 N | 80 | X |  |  |  | X | 42 |
| 4205 | Far-side craters | 119 E | 0 | 80 | X |  |  |  | X | 43 |
| 4224 | Far-side crater. | 100 E | 3 N | 80 | X |  |  |  | X | 43 |
| 4226 | Eastern limb | 97 E | 1 N | 80 | X |  |  |  | X | 44 |
| 4230 | Mare Smythii | 81 E | 1 S | 80 | X |  |  |  | X | 45 |
| 4253 | Messier B_. | 48 E | 1 S | 80 |  | x |  |  | X | 46 |
| 4256 | Messier craters | 47 E | 3 S | 80 |  | X |  |  | X | 46 |
| 4261 | Secchi K. | 45 E | 0 | 80 |  | X |  |  | X | 47 |
| 4265 | Secchi UA. | 42 E | 0 | 80 |  | X |  |  | X | 48 |
| 4276 | Mare Tranquillitatis_ | 38 E | 0 | 80 |  | X |  |  | X | 49 |
| 4312 | Mare Tranquillitatis. | 25 E | 0 | 80 |  | X |  |  | X | 50 |
| AS10-29-4324 | Moltke-------- | 24 E | 0 | 80 |  | X |  |  | X | 50 |
| AS10-30-4356 | Far-side crater. | 119 E | 4 N | 250 |  | X |  |  | X | 51 |
| 4371 | Far-side craters. | 107 E | 0 | 250 |  | X |  |  | X | 52 |
| 4372 | Far-side ray crater | 100 E | 4 N | 250 |  | X |  |  | X | 53 |
| 4426 | Taruntius A. | 50 E | 7 N | 250 | X |  |  |  | X | 54 |
| 4450 | Rima Ariadaeus. | 17 E | 5 N | 250 |  | X |  |  | X | 55 |


| Frame $n$ o. | Description | Principal point |  | $\begin{gathered} \text { Focal } \\ \text { length, }, \\ \mathrm{mm}, \end{gathered}$ | Sun angle |  |  | View |  | $\begin{gathered} \text { Pagg } \\ \text { no } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Longitude | Latitude |  | High | Medium | Low | Vertical | Oblique |  |
| AS10-31-4506 | Mare Fecunditatis. | 56 E | 2 N | 80 | X |  |  | X |  | 56 |
| 4512 | Taruntius G. | 50 E | 1 N | 80 | X |  |  | X |  | 56 |
| 4521 | Secchi $\theta_{-}$ | 40 E | 1 N | 80 |  | X |  | X |  | 57 |
| 4528 | Maskelyne D. | 33 E | 2 N | 80 | X |  |  |  | X | 57 |
| 4546 | Theon Senior | 16 E | 0 | 80 |  |  | X | X |  | 58 |
| 4566 | Taruntius | 46 E | 6 N | 250 | X |  |  |  | X | 59 |
| 4580 | Maskelyne F | 35 E | 4 N | 250 | X |  |  |  | X | 59 |
| 4601 | Moltke..- | 24 E | 1 S | 250 | X |  |  |  | X | 60 |
| 4621 | Rima Hypatia I and II | 22 E | 0 | 250 | X |  |  |  | X | 61 |
| 4630 | Arago-- | 22 E | 6 N | 250 | X |  |  |  | X | 61 |
| 4646 | Rima Ariadaeus. | 13 E | 7 N | 250 |  | X |  |  | X | 62 |
| 4647 | Godin.- | 11 E | 2 N | 250 |  | X |  |  | X | 63 |
| 4654 | Far-side crater | 164 E | 10 N | 250 |  | X |  |  | X | 64 |
| 4665 | Far-side basin | 143 E | 7 N | 250 | X |  |  |  | X | 65 |
| 4673 | Far-side basin. | 140 E | 7 N | 250 | X |  |  |  | X | 65 |
| AS10-32-4716 | Theophilus. | 25 E | 12 S | 250 |  | X |  |  | 8 | 67 |
| 4734 | Chladni. | 1 E | 4 N | 250 |  | X |  |  | X | 68 |
| 4771 | Lade. | 10 E | 0 | 80 | X |  |  | X |  | 68 |
| 4774 | Highlands | 8 E | 0 | 80 |  |  | X | X |  | 69 |
| 4813 | Hyginus | 5 E | 8 N | 80 |  |  | X |  | X | 70 |
| 4819 | Triesnecker | 4 E | 5 N | 80 |  |  | X |  | X | 70 |
| AS10-32-4823 | Far-side craters | 162 E | 10 S | 80 |  |  | X |  | X | 71 |
| 4828 | Far-side crater. | 146 E | 4 S | 80 | X |  |  |  | X | 71 |
| 4856 | Rhaeticus | 2 E | 0 | 80 |  |  | X |  | X | 72 |
| AS10-33-4914 | Far-side crater chain. | 139 E | 7 N | 250 | X |  |  |  | X | 66 |
| 4947 | Rhaeticus A. | 6 E | 1 N | 80 |  |  | X | X |  | 72 |
| 4975 | Far-side crater | 139 E | 6 S | 250 | X |  |  |  | X | 66 |
| 4999 | Mare Smythii | 82 E | 1 S | 250 | X |  |  |  | X | 45 |
| AS10-34-5014 | Earth. | In sp | pace | 80 |  |  |  | T |  | 2 |
| 5073 | Rima Hypatia I | 24 E | 1 S | 80 |  |  | X |  |  | 73 |
| 5081 | Neper--..-- | 85 E | 4 N | 80 | X |  |  |  | X | 74 |
| 5096 | Picard. | 50 E | 11 N | 80 |  | X |  |  | X | 75 |
| 5099 | Mare Tranquillitatis | 27 E | 1 N | 80 | X |  |  |  | X | 76 |
| 5129 | Mare Spumans. | 64 E | 1 N | 250 | X |  |  |  | X | 78 |
| 5136 | Taruntius H.. | 50 E | 0 | 250 | X |  |  |  | X | 79 |
| 5145 | Taruntius F- | 40 E | 3 N | 250 | X |  |  |  | X | 79 |
| 5150 | Mare Tranquillitatis | 35 E | 2 N | 250 |  | x |  | X |  | 78 |
| 5152 | Maskelyne X . | 27 E | 1 N | 250 |  | X |  |  | X | 76 |
| 5153 | Maskelyne G- | 27 E | 1 N | 250 |  | X |  |  | X | 77 |
| 5160 | Ritter | 19 E | 2 N | 250 |  |  | X |  | X | 80 |
| 5162 | Schmidt. | 20 E | 1 N | 250 |  |  | X |  | X | 81 |
| 5167 | Godin D. | 8 E | 2 N | 250 |  |  | X | X |  | 82 |
| 5171 | Far-side crater | 161 E | 5 S | 80 |  | X |  |  | X | 83 |
| 5172 | Far-side crater. | 158 E | 6 S | 80 |  | X |  |  | X | 84 |
| 5173 | Far-side craters | 157 E | 9 S | 80 |  | X |  |  | X | 84 |




Apollo 11-Concluded

| Frame no. | Description | Principal point |  | $\begin{gathered} \text { Focal } \\ \text { length, } \\ \mathrm{mm} \end{gathered}$ | Sun angle |  |  | View |  | $\begin{gathered} \text { Page } \\ \text { no. } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Longitude | Latitude |  | High | Medium | Low | Vertical | Oblique |  |
| 5956 | Central portion of $120-\mathrm{ft}-$ diameter crater |  |  | 60 |  |  |  |  |  | 102 |
| 5958 | Western portion of $120-\mathrm{ft}-$ diameter crater |  |  | 60 |  |  |  |  |  | 102 |
| 5963 | Aldrin driving core sampler. |  |  | 60 |  |  |  |  |  | 103 |
| AS11-44-6581 | LM preparing descent... |  |  | 80 |  |  |  |  |  | 85 |
| 6623 | LM preparing to rendezvous |  |  | 80 |  |  |  |  |  | 107 |
| 6643 | LM preparing for docking. |  |  | 80 |  |  |  |  |  | 107 |
| Panorama A... | Panoramas from the LM... |  |  |  |  |  |  |  |  | 104 |
| Panorama B | _ Panoramas from the surface_ |  |  |  |  |  |  |  |  | 105 |
| Panorama $\mathrm{C}_{\text {. }}$ | _Panoramas from the surface. |  |  |  |  |  |  |  |  | 106 |

## APPENDIX B <br> CREW AND LAUNCH PHOTOGRAPHS



Apollo 8 crew (from left to right): William A. Anders, James A. Lovell, and Frank Borman.


Apollo 10 crew (from left to right): Eugene A. Cernan, John W. Young, and Thomas P. Stafford.


Apollo 11 crew (from left to right): Neil A. Armstrong, Michael Collins, and Edwin E. Aldrin, Jr.

F


To the Moon.



