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Natural convection in an enclosed cavity

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## NATURAL CONVECTION IN AN ENCLOSED CAVITY

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## ABSTRACT

We are concerned with the problem of buoyancy driven flow in a vertical, rectangular cavity whose vertical sides are at different temperatures and whose horizontal sides are insulated. An application of the dynamic A.D.I. method to obtain numerical solutions to this problem is described. For large non-dimensional temperature differences characterized by the Rayleigh number the flow patterns develop strong boundary layers. These boundary layers are resolved by applying the D.A.D.I. method to the discretization of this problem on a non-uniform grid.

[^0]$$
1183-25433 x
$$

## Introduction

The problem to be considered was proposed by Jones [4] as a test bed for numerical methods for solving a variety of practical thermal problems. It is known as the "double glazing" or "window cavity" problem and has many applications, particularly in the field of thermal insulation. The most wellknown application is double glazing where a stagnant layer of air acts as an insulant between a warm room and a cold outside. The formulation of the problem is in terms of stream function, vorticity and temperature. The vorticity is eliminated to obtain a coupled pair of partial differential equations, one of which is fourth order. A dynamic A.D.I. (D.A.D.I.) method for solving these equations numerically is described. Grid stretching techniques originally due to Kalnáy de Rivas [8] are used to resolve the boundary layers which develop for large values of the Rayleigh number.

## 2. Formulation of the Problem

We are concerned with the problem of fluid flow in an upright, rectangular cavity described in non-dimensional terms by $0 \leqslant x \leqslant 1$, $0 \leqslant z \leqslant 1$, with $z$ vertically upwards. The cavity has different constant temperatures on the vertical walls, $T_{1}$ on the hot wall and $T_{2}$ on the cooler wall, and has insulated horizontal walls. We shall consider the twodimensional flow of a Boussinesq fluid of Prandt1 number 0.71 in which the flow takes place perpendicular to the walls. The Boussinesq approximation (see Mallinson and De Vahl Davis [9]) assumes that the physical properties and the density are constant except in the buoyancy term in the equations where the density is taken into account. This approximation is quite realistic and can give rise to predictions that are in good agreement with experiment (Jones
[5]) for small termperature differences. The governing equations are considerably simplified by use of this approximation. A full discussion and detailed description of this problem is given by Jones [6].


## Figure 1.

A non-dimensional temperature, $T$, is defined by

$$
T=\frac{\left(T^{*}-T_{2}\right)}{\left(T_{1}-T_{2}\right)}
$$

where $T^{*}$ is the temperature. The equations representing the conservation of mass, momentum and energy may be written as

$$
\begin{gather*}
\nabla \cdot \underline{v}=0  \tag{1}\\
(\nabla \times \underline{v}) \times \underline{v}=-\nabla p-\operatorname{Ra} \operatorname{Pr} T \underline{k}-\operatorname{Pr} \nabla^{2} \underline{v} \tag{2}
\end{gather*}
$$

$$
\begin{equation*}
\nabla \cdot(\underline{\mathrm{v}} \mathrm{~T})=\nabla^{2} \mathrm{~T} \tag{3}
\end{equation*}
$$

where $\underline{v}=(u, 0, w), \underline{k}=(0,0,1)$ and $p$ is the perturbation pressure. The Rayleigh number is given by

$$
\mathrm{Ra}=\mathrm{g} \beta\left(\mathrm{~T}_{1}-\mathrm{T}_{2}\right) / k \nu
$$

and the Prandtl number by

$$
\operatorname{Pr}=v / k,
$$

where $g$ is the acceleration due to gravity, $B$ the coefficient of volumetric expansion, $v$ the coefficient of kinematic viscosity and $k$ the coefficient of thermal conductivity.

Mallinson and De Vah1 Davis [10] show that the governing equations may be recast in the form

$$
\begin{align*}
& \operatorname{Pr} \nabla^{2} \zeta+\operatorname{Pr} \operatorname{Ra} \frac{\partial T}{\partial x}=u \frac{\partial \zeta}{\partial x}+w \frac{\partial \zeta}{\partial z}  \tag{4}\\
& \nabla^{2} \psi=-\zeta  \tag{5}\\
& -\frac{\partial}{\partial x}(u T)+\frac{\partial}{\partial z}(w T)=\nabla^{2} T  \tag{6}\\
& u=-\frac{\partial \psi}{\partial z}, w=\frac{\partial \psi}{\partial x} \tag{7}
\end{align*}
$$

where $\zeta$ is known as the vorticity and $\psi$ the stream function. Eqs. (4) to (7) represent what is known as the vorticity - stream function formulation of the problem.

The boundary conditions are

$$
\psi=\frac{\partial \psi}{\partial n}=0, \quad \text { on } \Gamma
$$

where $\Gamma$ is the boundary of $\Omega=\{(x, z): 0 \leqslant x \leqslant 1,0 \leqslant z \leqslant 1\}$,

$$
\begin{aligned}
& T=1, \\
& T=0, \\
& \frac{\partial T}{\partial n}=0, \\
& \text { on } x=0 \\
& \text { on } z=0 \quad \text { and } z=1
\end{aligned}
$$

Eliminating the velocities $u, w$ and the vorticity $\zeta$ we obtain the following system of equations:

$$
\begin{align*}
& \nabla^{4} \psi-\operatorname{Ra} \frac{\partial T}{\partial x}+\frac{1}{\operatorname{Pr}} \frac{\partial \psi}{\partial z} \frac{\partial}{\partial x}\left(\nabla^{2} \psi\right)-\frac{1}{\operatorname{Pr}} \frac{\partial \psi}{\partial x} \frac{\partial}{\partial z}\left(\nabla^{2} \psi\right)=0  \tag{8}\\
& \nabla^{2} T-\frac{\partial}{\partial z}(w T)-\frac{\partial}{\partial x}(u T)=0 \tag{9}
\end{align*}
$$

An advantage of eliminating the vorticity is that the need for a vorticity boundary condition is avoided. The system of equations (8) and (9) represents a fourth order equation for $\psi$ and a second order equation for $T$. The major Interest is in heat transfer so there are two important quantities namely the temperature fields and the overall heat transfer defined by a Nusselt number

$$
\mathrm{Nu}=\left.\int_{0}^{1} \frac{\partial \mathrm{~T}}{\partial \mathrm{x}} \mathrm{~d} z\right|_{\mathrm{x}=0}
$$

In order to solve this problem using an A.D.I. method we must first convert Eqs. (8) and (9) to the parabolic equations

$$
\begin{gather*}
\frac{1}{\lambda} \frac{\partial \psi}{\partial t}+\nabla^{4} \psi-\operatorname{Ra} \frac{\partial T}{\partial x}+\frac{1}{\operatorname{Pr}} \frac{\partial \psi}{\partial z} \frac{\partial}{\partial x}\left(\nabla^{2} \psi\right)-\frac{1}{\operatorname{Pr}} \frac{\partial \psi}{\partial x} \frac{\partial}{\partial z}\left(\nabla^{2} \psi\right)=0  \tag{10}\\
\frac{\partial T}{\partial t}=\nabla^{2} T-\frac{\partial}{\partial x}(u T)-\frac{\partial}{\partial z}(w T) \tag{11}
\end{gather*}
$$

whose steady state solution, if one exists, solves Eqs. (8) and (9). The parameter $\lambda$ controls the interaction between the equations. It means that for $\lambda \neq 1$ we are effectively using different time scales for the two equations.

## 3. Solution of a Nonlinear Equation

In a numerical process to find the solution of Eqs. (8) and (9) we use Eq. (8) to solve for $\psi$ and Eq. (9) to solve for $T$. Eq. (8) is nonlinear with respect to $\psi$ and can be written in the form

$$
\begin{equation*}
L(\psi)=\mathrm{f}(\mathrm{~T}), \tag{12}
\end{equation*}
$$

where $L$ is a nonlinear operator.
A Newton-type method is used to solve Eq. (12). Suppose that $\psi^{*}$ is some approximation to the solution of Eq. (12). We replace $L$ by its linearization about $\psi^{*}$ and then attempt to partially solve the linearized problem

$$
\begin{equation*}
L^{\prime}\left(\psi^{*}\right) \cdot\left(\psi-\psi^{*}\right)+L\left(\psi^{*}\right)-f(T) \simeq 0 \tag{13}
\end{equation*}
$$

where $L^{\prime}\left(\psi^{*}\right)$ is the Fréchet derivative of $L$ at $\psi^{*}$. We use the D.A.D.I. method to solve Eq. (13), the linearization being updated after each D.A.D.I. step. This will be explained more fully later.

The Fréchet derivative of $L$ at $\psi$ is given by

$$
\begin{align*}
L^{\prime}(\psi) \cdot \phi=\nabla^{4} \phi & +\frac{1}{\operatorname{Pr}} \frac{\partial \psi}{\partial z} \frac{\partial}{\partial x}\left(\nabla^{2} \phi\right)+\frac{1}{\operatorname{Pr}}\left\{\frac{\partial}{\partial x}\left(\nabla^{2} \psi\right)\right\} \frac{\partial \phi}{\partial z} \\
& -\frac{1}{\operatorname{Pr}} \frac{\partial \psi}{\partial x} \frac{\partial}{\partial z}\left(\nabla^{2} \phi\right)-\frac{1}{\operatorname{Pr}}\left\{\frac{\partial}{\partial z}\left(\nabla^{2} \psi\right)\right\} \frac{\partial \phi}{\partial x} \tag{14}
\end{align*}
$$

## 4. Finite Difference Equations

We cover $\Omega$ with a square grid of mesh length $h=1 / N$ where $N$ is a positive integer. Let $\psi_{i, j}$ and $T_{1, j}$ be the values of $\psi$ and $T$ at the point $\left(x_{i}, z_{j}\right)$ respectively where $x_{i}=i h$ and $z_{j}=j$. We use standard second-order centered difference approximations. At this point it is convenient to introduce the central difference operators $\delta_{x}^{2}$ and $\bar{\delta}_{x}$ as follows:

$$
\delta_{x}^{2} v_{i, j}=v_{i+1, j}-2 v_{i, j}+v_{i-1, j}, \quad \delta_{x} v_{i, j}=v_{i+1, j}-v_{i-1, j}
$$

The central difference operators $\delta_{z}^{2}$ and $\delta_{z}$ are defined in a similar manner. Eqs. (8) and (9) are discretized using these approxmiations to give

$$
\frac{1}{h^{4}}\left(\delta_{x}^{4}+2 \delta_{z}^{2} \delta_{x}^{2}+\delta_{z}^{4}\right) \psi_{i, j}-\frac{R a}{2 h} \bar{\delta}_{x} T_{i, j}-\frac{1}{2 h P r}\left(u_{i, j} \bar{\delta}_{x} p_{i, j}+w_{i, j} \bar{\delta}_{z} p_{i, j}\right)=0,(15)
$$

and

$$
\begin{equation*}
\frac{1}{h^{2}}\left(\delta_{x}^{2}+\delta_{z}^{2}\right)_{i, j}-\frac{1}{2 h} \bar{\delta}_{x}(u T)_{i, j}-\frac{1}{2 h} \bar{\delta}_{z}(w T)_{i, j}=0 \tag{16}
\end{equation*}
$$

where

$$
u_{i, j}=-\frac{1}{2 h} \bar{\delta}_{z} \psi_{i, j}, \quad w_{i, j}=\frac{1}{2 h} \bar{\delta}_{x} \psi_{i, j}, \quad p_{i, j}=\frac{1}{h^{2}}\left(\delta_{x}^{2}+\delta_{z}^{2}\right) \psi_{i, j} .
$$

The normal derivative boundary conditions are also discretized using central differences and these are used to remove imaginary exterior points from Eqs. (15) and (16) near the boundary. We have $2 \mathrm{~N}(\mathrm{~N}-1)$ equations for the unknowns $\Psi_{i, j}, 1=1, \ldots, N-1, j=1, \ldots, N-1$, and $T_{i, j}, i=1, \ldots, N-1$, $j=0, \ldots, N$.

Eq. (15) can be written in the form

$$
L_{h}\left(\psi_{i, j}\right)=f_{h}\left(T_{i, j}\right)
$$

which is the discrete form of Eq. (12), where $L_{h}$ is a nonlinear discrete operator.

## 5. Method of Solution

Here we describe how the D.A.D.I. method of Doss and Miller [2] can be applied to this problem. Since Eqs. (10) and (11) are parabolic they may be advanced in time by a direct method and the complete solution procedure may be regarded as a single iterative scheme. Our interest is not in solving the parabolic equations (10) and (11) accurately for finite times but to reach the steady state solution as soon as possible. We therefore use the D.A.D.I. method which uses a strategy that attempts to keep the time step $\Delta t$ within a region of fast convergence. An advantage of using an automatic step size changer is that it avoids the necessity of choosing a priori iteration parameters. The strategy of Doss and Miller [2] attempts to recognize

Instabilites as they start to occur and to bypass them by decreasing $\Delta t$. We start with initial approximations $T^{(0)}$ and $\psi^{(0)}$ to $T$ and $\psi$ respectively at time $t=0$. Suppose that we have reached the nth time step, where $n=2 m$ and $m$ is even, and that the current approximations of $T$ and $\psi$ at the grid points are $T^{(n)}$ and $\psi^{(n)}$ respectively. Starting from these approximations we describe a step of the D.A.D.I. method with time step $\Delta t$.

We begin by putting

$$
\phi_{i, j}^{(n)}=0,
$$

for $1, j=0,1, \ldots, N$.
The A.D.I. scheme due to Peaceman and Rachford [11] is used to advance Eq. (11) in time. The following systems are solved along lines in the $x$ direction:

$$
\begin{align*}
& \left(r_{1} I-\delta_{x}^{2}\right) T_{i, 0}^{(n+1)}=\left(r_{1} I+\delta_{z}^{2}\right) T_{i, 0}^{(n)},  \tag{17}\\
& \begin{array}{ll}
\left(r_{1} I-\delta_{x}^{2}\right) T_{i, j}^{(n+1)}+1 / 2 h \bar{\delta}_{x}\left(u^{(n)} T_{T}^{(n+1)}\right)_{i, j} \\
& =\left(r_{1} I+\delta_{z}^{2}\right) T_{i, j}^{(n)}-1 / 2 h_{z}\left(\bar{w}^{(n)} T_{T}^{(n)}\right)_{i, j} \\
& 1=1, \ldots, N-1, j=1, \ldots, N-1, \\
\left(r_{1} I-\delta_{x}^{2}\right) T_{i, N}^{(n+1)}=\left(r_{1} I+\delta_{z}^{2}\right) T_{i, N}^{(n)}, & i=1, \ldots, N-1,
\end{array}
\end{align*}
$$

with

$$
T_{0, j}^{(n+1)}=1, \quad T_{N, j}^{(n+1)}=0, \quad j=0,1, \ldots, N,
$$

homogeneous Neumann boundary conditions along $z=0$ and $z=1$ and where $r_{1}=h^{2} / \Delta t$.

The following systems are solved along lines in the $z$ - direction:

$$
\begin{gather*}
\left(r_{1} I-\delta_{z}^{2}\right) T_{i, j}^{(n+2)}+1 / 2 h \bar{\delta}_{z}\left(w^{(n)} T^{(n+2)}\right)_{i, j} \\
=\left(r_{1} I+\delta_{x}^{2}\right) T_{i, j}^{(n+1)}-1 / 2 h \bar{\delta}_{x}\left(u^{(n)} T_{T}^{(n+1)}\right)_{i, j}  \tag{20}\\
j=0, \ldots, N, i=1, \ldots, N-1
\end{gather*}
$$

with homogeneous Neumann boundary conditions along $z=0$ and $z=1$.
We advance $\psi$ in time using the A.D.I. scheme of Douglas and Rachford [3] to solve the fourth order linear equation (13). The extension of this scheme to solve the biharmonic equation is due to Conte and Dames [1].

We solve the following systems, described by $j=1, \ldots, N-1$, along lines in the $x$ - direction:

$$
\begin{align*}
& \left\{r_{2} I+\delta_{x}^{4}+\frac{1}{4 \operatorname{Pr}}\left(\bar{\delta}_{z} \psi_{i, j}^{(n)}\right) \bar{\delta}_{x} \delta_{x}^{2}-\frac{1}{4 \operatorname{Pr}}\left[\delta_{z}\left(\delta_{x}^{2}+\delta_{z}^{2}\right) \psi_{i, j}^{(n)}\right] \delta_{x}\right\} \phi_{i, j}^{(n+1)} \\
& =\left\{r_{2} I-\delta_{z}^{4}-2 \delta_{z}^{2} \delta_{x}^{2}-\frac{1}{4 \operatorname{Pr}}\left[\delta_{x}\left(\delta_{x}^{2}+\delta_{z}^{2}\right) \psi_{i, j}^{(n)}\right] \bar{\delta}_{z}+\frac{1}{4 \operatorname{Pr}}\left(\delta_{x} \psi_{i, j}^{(n)}\right) \bar{\delta}_{z}\left(\delta_{x}^{2}+\delta_{z}^{2}\right)\right. \\
& \left.-\frac{1}{4 \operatorname{Pr}}\left(\bar{\delta}_{z} \psi_{i, j}^{(n)}\right) \bar{\delta}_{x} \delta_{z}^{2}\right\} \phi_{i, j}^{(n)}+L_{h}\left(\psi_{i, j}^{(n)}\right)-f_{h}\left(T T_{i, j}^{(n)}\right) \\
& 1=1, \ldots, N-1, \tag{21}
\end{align*}
$$

with homogeneous Dirichlet and Neumann boundary conditions, and where $r_{2}=h^{4} /(\lambda \Delta t)$.

The following equations, described by $i=1, \ldots, N-1$, are solved along lines in the $z$ - direction:

$$
\begin{align*}
& \left\{r_{2} I+\delta_{z}^{4}-\frac{1}{4 \operatorname{Pr}}\left(\bar{\delta}_{x} \psi_{i, j}^{(n)}\right) \bar{\delta}_{z} \delta_{z}^{2}+\frac{1}{4 \operatorname{Pr}}\left[\bar{\delta}_{x}\left(\delta_{x}^{2}+\delta_{z}^{2}\right) \psi_{i, j}^{(n)}\right] \bar{\delta}_{z}\right\} \phi_{i, j}^{(n+2)} \\
& \quad=\left\{\delta_{z}^{4}-\frac{1}{4 \operatorname{Pr}}\left(\bar{\delta}_{x} \psi_{i, j}^{(n)}\right) \bar{\delta}_{z} \delta_{z}^{2}+\frac{1}{4 \operatorname{Pr}}\left[\bar{\delta}_{x}\left(\delta_{x}^{2}+\delta_{z}^{2}\right) \psi_{i, j}^{(n)}\right] \bar{\delta}_{z}\right\} \phi_{i, j}^{(n)}+r_{2} \phi_{i, j}^{(n+1)}, \tag{22}
\end{align*}
$$

$$
j=1, \ldots, N-1
$$

We define an A.D.I. step to be the process by which we obtain $T^{(n+2)}$ and $\phi^{(n+2)}$ from $T^{(n)}$ and $\phi^{(n)}$ i.e. the solution of equations (17) to (22). Starting with the approximations $T^{(n+2)}$ and $\phi^{(n+2)}$ we perform a second A.D.I. step using the same time step $\Delta t$ to obtain $T^{(n+4)}$ and $\phi^{(n+4)}$. The next part of the process is the bookkeeping stage of the D.A.D.I. step. Here we start with the approximations $T^{(n)}$ and $\phi^{(n)}$ and perform an A.D.I. step with time step $2 \Delta t$ to obtain $\widetilde{T}^{(n+4)}$ and $\tilde{\phi}^{(n+4)}$. We compute the test parameter, $T P$, which is given by

$$
\mathrm{TP}=\sqrt{[\mathrm{SUM} / \mathrm{ASUM}]},
$$

where

$$
\operatorname{SUM}=\left\|T{ }_{T}^{(n+4)}-\tilde{T}^{(n+4)}\right\|_{2}^{2}+\| \phi_{-}^{(n+4)}-\tilde{\phi}^{(n+4)} \eta_{2}^{2}
$$

and

$$
\text { ASUM }=\left\|T^{(n+4)}-T^{(n)}\right\|_{2}^{2}+\left\|\phi_{-\phi}^{(n+4)}{ }_{2}^{(n)}\right\|_{2}^{2}
$$

The test parameter is an estimate of the relative local truncation error. If we are interested in solving the parabolic equations (10) and (11) accurately then $\Delta t$ will be small, and so will TP. Our main concern, however, is to
accelerate convergence and attempt to push TP into an interval where convergence for A.D.I. is rapid. The strategy for changing $\Delta t$ is described in Doss and Miller [2]. Although their analysis of the step size strategy rests on rigid assumptions, they obtain good results in situations beyond the realm of the assumptions. If $T P>0.6$ then we reject the present D.A.D.I. step and start the step again with $\Delta t$ reduced by a factor of $1 / 16$. If TP $\leqslant 0.6$ then we accept the present D.A.D.I. step and change $\Delta t$ according to the strategy: if $T P$ falls in the intervals $[0,0.05]$, $0.05,0.1]$, ( $0.1,0.3$ ], $(0.3,0.4],(0.4,0.6]$ change $\Delta t$ by the factors $4,2, \sqrt{3}, 1 / 2,1 / 4$ respectively for the next D.A.D.I. step.

If the present step is accepted then we use one step of Newton's method to update the approximation to $\psi$ by setting

$$
\psi^{(n+4)}=\psi^{(n)}-\phi^{(n+4)} .
$$

We are now in a position to start the next D.A.D.I. step.

## 6. Numerical Results

The initial approximations $\mathrm{T}^{(0)}$ and $\psi^{(0)}$ are chosen to be the values at the grid points of the functions $T_{I}$ and $\psi_{I}$ respectively, where

$$
T_{I}(x, y)=1-x, \quad \psi_{I}(x, y)=0
$$

The algorithm described in the previous section was run for different grid sizes and for Rayleigh numbers of $10^{3}, 10^{4}, 10^{5}$, and $10^{6}$. The algorithm is terminated when the maximum modulus of the difference between successive
iterates on even time steps is less than $10^{-6}$ for $h=1 / 16,1 / 32$ and $5 \times 10^{-5}$ for $h=1 / 64$.

We experimented with different values of the parameter $\lambda$. For $\lambda=1$ we found that for large values of the Rayleigh number the iterates oscillated due to the interaction between the equations. For suitable choices of this parameter the instabilities were controlled and the method converged. The value of $\lambda$ was chosen to be $0.05,0.01,0.002$ and 0.002 respectively for Rayleigh numbers of $10^{3}, 10^{4}, 10^{5}$ and $10^{6}$. The choice of $\lambda$ was not critical in the sense that similar convergence behavior was observed for a wide range of values in the neighborhood of the chosen value.

Suppose that for a particular value of the Rayleigh number the problem has been solved numerically on a grid of mesh size $h$. To find the solution on a grid of mesh size $1 / 2 h$ we use the best available information to begin the new calculation i.e. we use as our initial approxmiation values interpolated from those obtained on the coarser grid. Cubic interpolation is used for values of $\psi$ and inear interpolation for values of $T$.

The average value of the Nusselt number is calculated using the trapezoidal rule, where we use the following approximation to the normal derivative of $T$ on the hot wall:

$$
\frac{\partial T}{\partial n}(0, z) \simeq \frac{\{-3 T(0, z)+4 T(h, z)-T(2 h, z)\}}{2 h}
$$

Table $I$ contains the average Nusselt number on the hot wall and the maximum and minimum local Nusselt numbers on the hot wall, and their location. Table $I I$ contains the number of D.A.D.I. steps and the run time, in seconds, required to reach the convergence criterion. The results were obtained on the 0xford University ICL 2980 computer. Contours for the
temperature and stream function are shown in Figures 2 to 9 for different values of the Rayleigh number.

The results we have obtained mainly show a good agreement and the convergence is approximately quadratic. Also it is clear that some form of non-uniform grid is needed to resolve the boundary layers accurately when Ra is large. In the next section we show that this technique has the effect of improving the rate of D.A.D.I. convergence.

TABLE I
Heat Transfer Results

| Ra | Mesh | Average Nusselt Number | Maximum <br> Nusselt Number |  | Minimum <br> Nusselt Number |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | $\overline{\mathrm{Nu}}$ | $\mathrm{Nu}_{\text {max }}$ | $@ z=$ | $N \mathrm{~N}_{\mathrm{min}}$ | $@_{z}=$ |
| $10^{3}$ | $17 \times 17$ | 1.120 | 1.515 | 0.86 | 0.697 | 0.00 |
|  | $33 \times 33$ | 1.118 | 1.508 | 0.91 | 0.693 | 0.00 |
|  | $65 \times 65$ | 1.118 | 1.507 | 0.91 | 0.692 | 0.00 |
| $10^{4}$ | $17 \times 17$ | 2.357 | 3.838 | 0.81 | 0.615 | 0.00 |
|  | $33 \times 33$ | 2.270 | 3.628 | 0.84 | 0.592 | 0.00 |
|  | $65 \times 65$ | 2.250 | 3.554 | 0.86 | 0.587 | 0.00 |
| $10^{5}$ | $17 \times 17$ | 5.146 | 8.423 | 0.88 | 0.854 | 0.00 |
|  | $33 \times 33$ | 4.758 | 8.508 | 0.91 | 0.755 | 0.00 |
|  | $65 \times 65$ | 4.573 | 7.972 | 0.92 | 0.735 | 0.00 |
| $10^{6}$ | $33 \times 33$ | 10.062 | 19.085 | 0.91 | 0.926 | 0.03 |
|  | $65 \times 65$ | 9.272 | 19.590 | 0.95 | 0.999 | 0.00 |

TABLE II
Computational Details

| Ra | Mesh | Number of D.A.D.I steps | Time |
| :---: | :---: | :---: | :---: |
| $10^{3}$ | $17 \times 17$ | 24 | 4.9 |
|  | $33 \times 33$ | 19 | 14.9 |
|  | $65 \times 65$ | 9 | 30.3 |
| $10^{4}$ | $17 \times 17$ | 36 | 6.7 |
|  | $33 \times 33$ | 36 | 26.9 |
|  | $65 \times 65$ | 21 | 66.0 |
| $10^{5}$ | $17 \times 17$ | 111 | 21.1 |
|  | $33 \times 33$ | 92 | 68.9 |
|  | $65 \times 65$ | 49 | 147.5 |
| $10^{6}$ | $33 \times 33$ | 362 | 268.5 |
|  | $65 \times 65$ | 335 | 1012.0 |



Figure 2. Temperature contours for $\operatorname{Ra}=10^{3}$. Contour levels are 0.1(0.1)0.9.


Figure 3. Stream function contours for $\mathrm{Ra}=10^{3}$.
Contour levels are $-0.12(-0.12)-1.08$.


Figure 4. Temperature contours for $\mathrm{Ra}=10^{4}$ Contour levels are 0.1(0.1)0.9.


Figure 5. Stream function contours for $\mathrm{Ra}=10^{4}$. Contour levels are $\mathbf{- 0 . 5 5 2 ( - 0 . 5 5 2 ) - 4 . 9 6 8 .}$


Figure 6. Temperature contours for $\mathrm{Ra}=10^{5}$. Contour levels are 0.1(0.1)0.9.


Figure 7. Stream function contours for $\mathrm{Ra}=10^{5}$. Contour levels are -1.06(-1.06)-9.54.


Figure 8. Temperature contours for $\mathrm{Ra}=10^{6}$. Contour levels are 0.1(0.1)0.9.


Figure 9. Stream function contours for $R a=10^{6}$. Contour levels are $-2.1(-2.1)-18.9,-19.3,-19.8$.

## 7. Non-Uniform Grids

For large values of the Rayleigh number the flows develop strong boundary layers. This effect can be seen in Figure 9, for example. When $\mathrm{Ra}=10^{6}$ we have had to use an extremely fine mesh, $65 \times 65$, in order to resolve these boundary layers. This is rather wasteful since the grid points are also densely distributed away from these layers where they are not needed. Here we apply the technique of Kálnay de Rivas [8] to resolve the boundary layers using a non-uniform grid.

Basically, the idea is to make a change of independent variable so that the domain is mapped into a new co-ordinate system where the variations of the solution are not so rapid. The grid intervals are varied by defining stretched co-ordinates $\zeta$ and $\eta$ such that $x=x(\zeta)$ and $z=z(\eta)$ where the grid intervals $\Delta \zeta$ and $\Delta \eta$ are constant and $x$ and $z$ are the old physical co-ordinates. The mapping is chosen so that the solution, when regarded as a function of the new variables, has no boundary layers.

Kálnay de Rivas [8] and Jones and Thompson [7] show how to express derivatives in terms of the stretched co-ordianates. For example, we can express the first derivative in terms of $\zeta$ in the following manner

$$
\begin{equation*}
\frac{\partial v}{\partial x}=\frac{\partial v}{\partial \zeta} \cdot \frac{d \zeta}{d x} . \tag{23}
\end{equation*}
$$

Eq. (23) can be discretized using central differences to give the following approximation

$$
\begin{equation*}
\left.\frac{\partial v}{\partial x} \simeq \frac{v_{i+1, j^{-v}}}{2 \Delta \zeta}{ }_{i-1, j} \frac{d \zeta}{d x}\right|_{x_{i}}, \tag{24}
\end{equation*}
$$

where $v_{i, j}$ is the value of $v$ at the grid point $\left(x_{i}, z_{j}\right)$ with
$x_{i}=x(1 \Delta \zeta)$ and $z_{j}=z(j \Delta \eta)$.
The transformation can be differentiated using central differences to obtain the following approximation to the first derivative:

$$
\begin{equation*}
\frac{\partial v}{\partial x} \simeq \frac{\left(v_{i+1, j^{-v}}^{i-1, i}\right.}{2 \Delta \zeta} \cdot \frac{2 \Delta \zeta}{\left(x_{i+1}-x_{i-1}\right)}=\frac{\left(v_{i+1, j^{-v_{i-1, i}}}\right)}{\left(x_{i+1}-x_{i-1}\right)} \tag{25}
\end{equation*}
$$

Finite difference approximations for higher order derivatives are obtained in a similar way.

Let $x(\zeta)$ and $z(n)$ be two grid stretching functions with constant grid intervals $\Delta \zeta$ and $\Delta \eta$ respectively. The region $\Omega$ is covered with a variable grid defined by the above mappings. We define $T_{i, j}$ and $\psi_{i, j}$ to be the values of $T(x, z)$ and $\psi(x, z)$ respectively at the grid point $\left(x_{i}, z_{j}\right)$. The finite difference equations are constructed using approximations like Eq. (25) for the derivatives. The system of finite difference equations are solved using the D.A.D.I. method described earlier.

The problem is solved on the following non-unfform grids:
(a)

$$
\begin{align*}
& x(\zeta)=\sin ^{2}(1 / 2 \pi \zeta) \\
& z(\eta)=\sin ^{2}(1 / 2 \pi \eta) \tag{26}
\end{align*}
$$

(b)

$$
\begin{align*}
& x(\zeta)=6 \zeta^{5}-15 \zeta^{4}+10 \zeta^{3} \\
& z(n)=6 n^{5}-15 n^{4}+10 n^{3} \tag{27}
\end{align*}
$$

These grid stretching functions give a smaller spacing of the grid points near the boundary.

The problem was solved for Rayletgh numbers of $10^{5}$ and $10^{6}$ on the nonuniform grids defined by Eqs. (26) and (27) with $\Delta \zeta=\Delta \eta=0.04$. In Table III we compare the average value of the Nusselt number along the hot wall obtained on these non-uniform grids, a uniform $26 \times 26$ mesh and a uniform $65 \times 65$ mesh. In Tables IV and $V$ we compare the number of D.A.D.I. steps and run time respectively to reach the convergence criterion, which is the same as that used for the results in Table II.

In Table III we see that we have obtained a good estimate of the Nusselt number by using a stretched grid even though we have only a $26 \times 26$ mesh. For small values of the Rayleigh number the use of a stretched grid has little effect. It is only when the Rayleigh number becomes large that we obtain an improvement by using non-uniform grids. The use of non-uniform grids to resolve the boundary layers has had the effect of improving the rate of D.A.D.I. convergence. Approximately twice as many steps of D.A.D.I. were required for convergence of the method using the uniform $26 \times 26$ mesh than for stretched $26 \times 26$ meshes.

The use of non-uniform grids has allowed us to resolve the boundary layers using comparatively few mesh points. Reasonable accuracy has also been obtained on the stretched grids compared with an extremely fine mesh.

The iterative solution of finite difference equations constructed on a non-uniform grid usually presents great difficulties. This is due to the problem of finding suitable parameters for the acceleration of convergence of any selected iterative method. Hence, an advantage of the D.A.D.I. method over standard iterative methods for solving problems of this type is that we do not require an a priori choice of parameters to accelerate convergence.

TABLE III
Average value of the Nusselt number

| Ra | Uniform Mesh | Stretched Mesh <br> (a) | Stretched Mesh <br> (b) | Uniform Mesh |
| :---: | :---: | :---: | :---: | :---: |
| $10^{5}$ | $46 \times 26$ | $26 \times 26$ | $26 \times 26$ | $65 \times 65$ |
| $10^{6}$ | 10.163 | 9.596 | 4.595 | 4.573 |

TABLE IV
Number of D.A.D.I. steps

|  | Uniform Mesh | Stretched Mesh <br> Ra | (a) | Stretched Mesh <br> (b) |
| :---: | :---: | :---: | :---: | :---: |
| $10^{5}$ | 151 | $26 \times 26$ | $26 \times 26$ | Uniform Mesh |
| $10^{6}$ | 676 | 85 | 84 | $65 \times 65$ |

TABLE V
Computational time

|  | Uniform Mesh | Stretched Mesh <br> (a) | Stretched Mesh <br> (b) | Uniform Mesh |
| :---: | :---: | :---: | :---: | :---: |

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