FINITE DIFFERENCE ANALYSIS OF ROTORDYNAMIC SEAL<br>COEFFICIENTS FOR AN ECCENTRIC SHAFT POSITION<br>R. Nordmann and F.J. Dietzen<br>Department of Mechanical Engineering<br>University of Kaiserslautern<br>Kaiserslautern, Federal Republic of Germany

The dynamic coefficients of seals are calculated for shaft movements around an eccentric position. The turbulent flow is described by the Navier-Stokes equations in connection with a turbulence model. The equations are solved by a finite-difference procedure.

## INTRODUCTION

To model the dynamic behaviour of turbopumps properly it is very important to consider the fluid forces which are developed in the seals. This has been clearly demonstrated by some authors, like for example Diewald (1987). The fluid forces are normally described by the following equation.

$$
-\left[\begin{array}{l}
F_{z}  \tag{1}\\
F_{y}
\end{array}\right]=\left[\begin{array}{cc}
K & k \\
-k & K
\end{array}\right]\left[\begin{array}{l}
z \\
y
\end{array}\right]+\left[\begin{array}{cc}
D & d \\
-d & D
\end{array}\right]\left[\begin{array}{l}
\mathbf{z} \\
\dot{y}
\end{array}\right]+\left[\begin{array}{ll}
M & 0 \\
0 & M
\end{array}\right]\left[\begin{array}{c}
\dot{z} \\
\hdashline \mathbf{y}
\end{array}\right]
$$

But this equation is only valid for a shaft moving around the center of the seal, which very seldom occurs in reality. In most machines the shaft will orbit around an eccentric position, so that the fluid forces must be described by

$$
-\left[\begin{array}{l}
F_{z}  \tag{2}\\
F_{y}
\end{array}\right]=\left[\begin{array}{cc}
K_{z z} & k_{z y} \\
-k_{y z} & K_{y y}
\end{array}\right]\left[\begin{array}{l}
Z \\
Y
\end{array}\right]+\left[\begin{array}{cc}
D_{z z} & d_{z y} \\
-d_{y z} & D_{y y}
\end{array}\right]\left[\begin{array}{l}
\dot{Z} \\
\dot{Y}
\end{array}\right]+\left[\begin{array}{cc}
M_{z z} & 0 \\
0 & M_{y y}
\end{array}\right]\left[\begin{array}{l}
\dot{Z} \\
\ddot{Y}
\end{array}\right]
$$

The dynamic coefficients in such a case have been investigated by Jenssen (1970). Allaire et al. (1976) and recently in an excellent paper by Nelson and Nguyen (1987). To model the turbulent flow, all these have used so called "Bulk-Flow Theories" in which the shear stress at the wall is described as a function of the average fluid velocity relative to the wall. Because some authors affirm that in the case of great eccentricities recirculation in circumferential direction occurs (which can't be described by a bulk-flow model) and strongly effects the dynamic coefficients, we extended the theory of Dietzen and Nordmann (1987) and Nordmann (1987) to investigate the flow and the coefficients in the case of an eccentric shaft. In that theory the Navier Stokes equations in connection with the $k-\epsilon$ turbulence model were used to calculate the dynamic coefficients of incompressible and compressible seals for a shaft motion around the centric position.

## GOVERNING EQUATIONS

In a turbulent flow, turbulent stresses occur which are of ten modeled like laminar stresses by introducing a turbulent viscosity. The turbulent and the laminar viscosity are then added to an effective viscosity.

$$
\begin{equation*}
\mu_{e}=\mu_{l}+\mu_{t} \tag{3}
\end{equation*}
$$

The turbulent viscosity must be described by a turbulence model. We use the $k-\epsilon$ model, but also much simpler mixing-length models will be appropriate for a straight seal. The turbulent viscosity is given by

$$
\begin{equation*}
\mu_{t}=C_{\mu} \rho \frac{k^{2}}{\epsilon} \tag{4}
\end{equation*}
$$

So we have the Navier-Stokes equations, the continuity equation and the equations of the $k-\epsilon$ model to describe the turbulent flow in a seal. These equations have the following form.
$\frac{\partial(\rho \phi)}{\partial \mathrm{t}}+\frac{\partial}{\partial \mathrm{x}}(\rho \mathrm{u} \phi)-\frac{\partial}{\partial \mathrm{x}}\left(\Gamma \frac{\partial \phi}{\partial \mathrm{x}}\right)+\frac{1}{\mathrm{r}} \frac{\partial}{\partial r}(\mathrm{r} \rho \mathrm{v} \phi)-\frac{1}{\mathrm{r}} \frac{\partial}{\partial \mathrm{r}}\left(\mathrm{r} \Gamma \frac{\partial \phi}{\partial \mathrm{r}}\right)+\frac{1}{\mathrm{r}} \frac{\partial}{\partial \theta}(\rho w \phi)-\frac{1}{\mathrm{r}} \frac{\partial}{\partial \theta}\left(\Gamma \frac{1}{\mathrm{r}} \frac{\partial \phi}{\partial \theta}\right)=\mathrm{S}_{\phi}$

| $\phi$ | $\Gamma$ | $\mathbf{S}_{\phi}$ |
| :---: | :---: | :---: |
| u | $\mu_{\text {e }}$ | $-\frac{\partial p}{\partial x}+\frac{\partial}{\partial x}\left(\mu \frac{\partial u}{\partial x}\right)+\frac{1}{r} \frac{\partial}{\partial r}\left(r \mu e \frac{\partial v}{\partial x}\right)+\frac{1}{r} \frac{\partial}{\partial \theta}\left(\mu \frac{\partial w}{\partial x}\right)$ |
| v | $\mu_{\text {e }}$ | $\begin{aligned} & -\frac{\partial \mathrm{p}}{\partial \mathrm{r}}+\frac{\partial}{\partial \mathrm{x}}\left(\mu_{\mathrm{e}} \frac{\partial \mathrm{u}}{\partial \mathrm{r}}\right)+\frac{1}{\mathrm{r}} \frac{\partial}{\partial \mathrm{r}}\left(\mathrm{r} \mu_{\mathrm{e}} \frac{\partial \mathrm{v}}{\partial \mathrm{r}}\right)+\frac{1}{\mathrm{r}} \frac{\partial}{\partial \theta}\left(\mathrm{r} \mu_{\mathrm{e}} \frac{\partial}{\partial \mathrm{r}}\left(\frac{\mathrm{w}}{\mathrm{r}}\right)\right) \\ & -\frac{2}{\mathrm{r}^{2} \mu_{\mathrm{e}}} \frac{\frac{\partial \mathrm{w}}{\partial \theta}}{}-\frac{2}{\mathrm{r}^{2}} \mu_{\mathrm{e}} \mathrm{v}+\frac{\rho}{\mathrm{r}} w^{2} \end{aligned}$ |
| w | $\mu_{\text {e }}$ | $\begin{aligned} & -\frac{1}{r} \frac{\partial p}{\partial \theta}+\frac{\partial}{\partial x}\left(\frac{1}{r} \mu e \frac{\partial u}{\partial \theta}\right)+\frac{1}{r} \frac{\partial}{\partial r}\left(\mu \frac{\partial v}{\partial \theta}\right)+\frac{1}{r} \frac{\partial}{\partial \theta}\left(\frac{1}{r} \mu e \frac{\partial w}{\partial \theta}\right) \\ & +\frac{1}{r^{2}} \mu_{e} \frac{\partial v}{\partial \theta}-\frac{w}{r^{2}} \frac{\partial}{\partial r}\left(r \mu_{e}\right)+\frac{1}{r} \frac{\partial}{\partial \theta}\left(\frac{2}{r} \mu_{e} v\right)-\frac{e^{\prime}}{r} v w \end{aligned}$ |
| 1 | 0 | 0 |
| k | $\mu_{\mathrm{e}} / \sigma_{\mathrm{k}}$ | $\mathrm{G}-\mathrm{\rho} \mathrm{\epsilon}$ |
| $\epsilon$ | $\mu_{e} / \sigma_{\epsilon}$ |  |

Table 1: The governing equations of the turbulent seal flow.
(The constants of the $k-\epsilon$ model are given in appendix A)

## PERTURBATION ANALYSIS



Fig. 1 Geometry of the eccentric shaft

To describe the flow if the shaft is moving around an eccentric position we follow the procedure of Dietzen and Nordmann (1987) and use a similar transformation. (see Fig. 1)

$$
\begin{equation*}
\eta=r_{a}(\theta)-\frac{r_{a}(\theta)-r}{\delta(\theta, t)} C_{o}(\theta) \tag{6}
\end{equation*}
$$

But now $r_{a}$ and $C_{o}$ are functions of $\theta$. This is not so if the shaft moves around the center. $\delta$ is the seal clearance, varying with angle $\theta$ and time $t$.

The radius $r_{a}$ can be described by the following equation:

$$
\begin{equation*}
r_{a}(\theta)=\sqrt{\mathrm{R}_{\mathrm{aN}}^{2}-\mathrm{ex}^{2} \sin ^{2}(\theta)}-\mathrm{ex} \cos (\theta) \tag{7}
\end{equation*}
$$

As result of this transformation the shaft orbiting around an eccentric position is transformed to the stationary eccentric position.
If we introduce this transformation into our equations, we must obey the following relations.

$$
\begin{equation*}
\left(\frac{\partial}{\partial \theta}\right)_{r}=\left(\frac{\partial}{\partial \theta}\right)_{\eta}+\left(\frac{\partial}{\partial \eta}\right)_{\theta}\left(\frac{\partial \eta}{\partial \theta}\right)_{r} \quad\left(\frac{\partial}{\partial t}\right)_{r}=\left(\frac{\partial}{\partial t}\right)_{\eta}+\left(\frac{\partial}{\partial \eta}\right)_{t}\left(\frac{\partial \eta}{\partial t}\right)_{r} \tag{8}
\end{equation*}
$$

To calculate the rotordynamic coefficients we assume that the shaft moves around the eccentric position on small orbits, so that we are allowed to introduce a perturbation analysis.

$$
\begin{array}{lll}
\delta=C_{0}-e h_{1} & u=u_{0}+e u_{1} & v=v_{0}+e v_{1}  \tag{9}\\
& w=w_{0}+e w_{1} & p=p_{0}+e p_{1}
\end{array}
$$

If we introduce these expressions and the coordinate transformation in our governing equation, neglecting terms with power of e greater than 1 and separating the parts with and without e we will get a set of zeroth order and first order equations. The zeroth order equations describe the stationary flow for the eccentric shaft, the first order equations the perturbation of this flow, if the shaft moves around the stationary position.

We assume that the shaft moves with the precession frequency $\Omega$ on a circular orbit with radius $r_{0}$ around the eccentric position. So the change of the clearance is given by

$$
\begin{equation*}
e h_{1}=r_{0}(\cos \Omega t \cos \theta+\sin \Omega t \sin \theta) \tag{10}
\end{equation*}
$$

Because this change is periodic in time we introduce also periodic functions for the flow variables.

$$
\begin{align*}
& u_{1}=u_{1} c \cos \Omega t+u_{1} s \sin \Omega t \quad v_{1}=v_{1} c \cos \Omega t+v_{1} s \sin \Omega t  \tag{11}\\
& w_{1}=w_{1} c^{\cos \Omega t+w_{1}} \sin \Omega t \quad p_{1}=p_{1} c^{\cos \Omega t+p_{1}} \sin \Omega t
\end{align*}
$$

By separating now in the first order equations the terms with cos $\Omega \mathrm{t}$ and $\sin \Omega \mathrm{t}$ we obtain two real equations for every first order equation. These equations are then arranged in a new form by introducing complex variables.

$$
\begin{array}{ll}
\hat{u}_{1}=u_{1} c \\
\hat{w}_{1} & i u_{1} s  \tag{12}\\
\mathbf{w}_{1} c \\
& \hat{v}_{1} w_{1} s
\end{array}
$$

Finally supplementary to our real zeroth order equations we have a set of complex first order equations. These equations have the following form.

$$
\begin{equation*}
\frac{\partial}{\partial \mathrm{x}}\left(\rho \mathrm{u}_{0} \phi\right)-\frac{\partial}{\partial \mathrm{x}}\left(\Gamma \frac{\partial \phi}{\partial \mathrm{x}}\right)+\frac{1 \partial}{\eta \partial \eta}\left(\eta \rho \mathrm{v}_{0} \phi\right)-\frac{1 \partial}{\eta \partial \eta}\left(\eta \Gamma \frac{\partial \phi}{\partial \eta}\right)+\frac{1 \partial}{\eta \partial \theta}\left(\rho w_{0} \phi\right)-\frac{1 \partial}{\eta \partial \theta}\left(\Gamma \frac{1 \partial \phi}{\eta \partial \theta}\right)=\mathbf{s}_{\phi} \tag{13}
\end{equation*}
$$

\begin{tabular}{|c|c|c|}
\hline $\phi$ \& $r_{\phi}$ \& $\mathbf{S}_{\phi}$ <br>
\hline $$
\begin{aligned}
& u_{0} \\
& v_{0} \\
& \mathbf{w}_{0} \\
& 1 \\
& k_{0} \\
& \epsilon_{0}
\end{aligned}
$$ \& $$
\begin{gathered}
\mu_{\mathrm{e}} \\
\mu_{\mathrm{e}} \\
\mu_{\mathrm{e}} \\
0 \\
\mu_{\mathrm{e}} / \sigma_{\mathrm{k}} \\
\mu_{\mathrm{e}} / \sigma_{\mathrm{t}}
\end{gathered}
$$ \& $$
\begin{aligned}
& S_{u_{0}} \\
& S_{v_{0}} \\
& \mathbf{s}_{w_{0}} \\
& 0 \\
& S_{k_{0}} \\
& S_{\epsilon_{0}}
\end{aligned}
$$ <br>
\hline $$
\hat{\mathrm{u}}_{1}
$$
$$
\hat{\mathbf{v}}_{1}
$$
$$
\hat{w}_{1}
$$ \& $\mu_{\mathbf{e}}$
$\mu_{\mathbf{e}}$

$\mu_{\mathbf{e}}$ \&  <br>
\hline
\end{tabular}

Table 2: Source terms of zeroth and first order equations.

Only the first order continuity equation to calculate $\hat{\mathrm{p}}_{1}$ has a slightly different form.

$$
\begin{equation*}
\frac{\partial}{\partial \mathrm{x}}\left(\rho \hat{u}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \eta}\left(\eta \rho \hat{\mathrm{v}}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \theta}\left(\rho \hat{w}_{1}\right)=\mathrm{D}_{\mathrm{p}_{0}} \tag{14}
\end{equation*}
$$

You get $S_{u_{0}}, S_{v_{0}}, S_{w_{0}}, S_{k_{0}}, S_{\epsilon_{0}}$, if you replace in table 1 in the corresponding terms $r$ by $\eta$ and $u, v, w, p, k, \epsilon$ by $u_{o}, v_{0}, w_{0}, P_{0}, k_{o}, \epsilon_{0}$. The terms $C_{u_{1}}, C_{v_{1}}, C_{w_{1}}$ result from the perturbation of the convective terms in the Navier Stokes equations. $D_{u_{0}}{ } D_{v_{0}} \cdot D_{w_{0}} . D_{p_{0}}$ are constants resulting from the coordinate transformation, which are not functions of $\hat{u}_{1}, \hat{\mathbf{v}}_{1}, \hat{w}_{1}, \hat{p}_{1}$. The terms with $\Omega$ represent the time dependent parts.

Because we assume that the viscosity $\mu_{e}$ remains constant for the small motions, we do not need a $\hat{k}_{1}$ and $\hat{\epsilon}_{1}$ equation.
$\left(D_{u_{0}}, D_{v_{0}}, D_{w_{0}}, D_{p_{0}}, C_{u_{i}}, C_{v_{1}}, C_{w_{1}}\right.$ are given in appendix B)

## BOUNDARY CONDITIONS

The zeroth order boundary conditions are

$$
\begin{aligned}
& \text { stator : } u_{\mathbf{o}_{\mathbf{s}}}=0 \quad \mathbf{v}_{\mathbf{o}_{\mathbf{s}}}=0 \quad \mathbf{w}_{\mathbf{o}_{\mathbf{s}}}=0 \\
& \text { rotor : } \quad \mathbf{u}_{\mathbf{o}_{\mathbf{r}}}=0 \quad \mathbf{v}_{\mathbf{o}_{\mathbf{r}}}=0 \quad \mathbf{w}_{\mathbf{o}_{\mathbf{r}}}=\mathrm{R}_{\mathbf{i} N^{\omega}} \\
& \text { entrance : } p_{o_{E n}}=p_{\text {Res }}-\frac{1}{2} \rho u_{o_{E n}^{2}}^{2}(1+\xi) \\
& \text { exit : } p_{o_{E x}}=0
\end{aligned}
$$

$u_{\text {En }}$ is the average axial entrance velocity for every plane with $\theta$ constant. $p_{\text {Res }}$ is the reservoir pressure and $\omega$ the rotational frequency of the shaft.

The first order boundary conditions are

$$
\begin{aligned}
& \text { stator : } \hat{u}_{1_{s}}=(0.0) \quad \hat{\mathbf{v}}_{\mathbf{1}_{\mathbf{s}}}=(0.0) \quad \hat{w}_{\mathbf{w}_{\mathbf{s}}}=(0.0) \\
& \text { rotor : } \hat{\mathbf{u}}_{1_{r}}=(0,0) \\
& \hat{\mathbf{v}}_{\mathbf{1}_{\mathbf{r}}}=\left[\mathrm{C}_{\mathbf{o}_{\mathrm{N}}}(\Omega-\omega) \sin \theta,-\mathrm{C}_{\mathbf{o}_{\mathrm{N}}}(\Omega-\omega) \cos \theta\right] \\
& {\hat{\mathbf{w}_{1}}}_{\mathbf{r}}=\left[\mathrm{C}_{\mathbf{o}_{\mathrm{N}} \Omega} \cos \theta, \mathrm{C}_{\mathrm{o}_{N} \Omega \sin \theta}\right] \\
& \text { entrance : } \hat{\mathrm{p}}_{1_{\mathrm{En}}}=-\rho{u_{\mathrm{o}_{\mathrm{En}}}(1+\xi) \hat{\mathrm{u}}_{1_{\mathrm{En}}}, ~}_{\text {ent }} \\
& \text { exit: } \hat{\mathbf{p}}_{1 \mathrm{Ex}}=(0,0)
\end{aligned}
$$

THE FINITE DIFFERENCE RETHOD
For solving these equations a finite-difference procedure is used which is based on the method published by Gosman and Pun (1974). The seal is discretized by a grid (Fig.2) and the variables are calculated at the nodes. The velocities $u_{0}, v_{0}, w_{0}$ $\left(\hat{u}_{1}, \hat{v}_{1}, \hat{w}_{1}\right)$ are determined at points which lie between the nodes where the variables $\mathrm{p}_{\mathrm{o}}, \mathrm{k}_{\mathrm{o}}, \epsilon_{\mathrm{o}}\left(\hat{\mathrm{p}}_{1}\right)$ are calculated (Fig.3)


Fig. 2 3-dimensional grid in the seal.


Fig. 3 Velocities and pressure in a staggered grid.

To calculate the flow we proceed as follows:

1. We start our procedure with guessed values for all variables.
2. First the velocities $u_{0}, v_{0}, w_{0}\left(\hat{u}_{1}, \hat{v}_{1}, \hat{w}_{1}\right)$ are calculated.
3. Then the velocities and the pressure are corrected to satisfy the continuity equation.
To do that we use a modified version of the 'PISO' procedure of Benodekar et al. (1985).
4. After this $k$ and $\epsilon$ are calculated (only for the zeroth order solution).

We repeat step 2 to 4 until we reach a convergent solution. First we solve the zeroth order equations and then the first order equations.

Of course we need a 3 -dimensional finite-difference method to calculate the flow in the case of an eccentric shaft,while a two-dimensional method is sufficient for movements around the centric position.

## DETERMINATION OF THE DYNAHIC COEFFICIENTS

By integration the pressure $\hat{\mathrm{p}}_{1}$ around the shaft we get forces in $z$ and $y$ directions.
Then we introduce $z, y, z, y, \underset{z}{ }, y$ from our circular orbit into equation (2). This gives us the following equations:

$$
\begin{aligned}
& \int_{0}^{L} \int_{0}^{2 \pi} p_{1_{1}} \cos \theta R_{i N} d \theta d x=C_{o_{N}}\left(K_{z z}+d_{z y} \Omega-M_{z z} \Omega^{2}\right) \\
& \int_{0}^{L} \int_{0}^{2 \pi} p_{1_{1}} \cos \theta R_{i N} d \theta d x=C_{o_{N}}\left(k_{z y}-D_{z z} \Omega\right) \\
& L_{0}^{2 \pi} \int_{0}^{2 \pi} p_{1_{1}} \sin \theta R_{i N} d \theta d x=C_{o_{N}}\left(-k_{y z}+D_{y y} \Omega\right) \\
& \int_{0}^{L} \int_{0}^{2 \pi} p_{1} \sin \theta R_{i N} d \theta d x=C_{o_{N}}\left(K_{y y}+d_{y z} \Omega-M_{y y} \Omega^{2}\right)
\end{aligned}
$$

If we calculate the forces for several precession frequencies $\Omega$ of the shaft, we can obtain the coefficients by a 'Least-Square-Fit'.

RESULTS
We compare our theory with the model of Nelson and Nguyen (1987) and some experimental results of Falco et al. (1986) which also have been published in the paper of Nelson and Nguyen (1987).

In Fig. 4-10 dynamic coefficients are calculated as a function of the eccentricity. We compare our results with Nelson and Nguyen's theory and the stiffness coefficients also with experimental and theoretical results of Falco et al.

The seal data are

$$
\begin{aligned}
& \text { length }: L=40.0 \mathrm{~mm} \quad \text { pressure drop }: 1.0 \mathrm{Mpa} \\
& \text { shaft radius : } R_{i N}=80.0 \mathrm{~mm} \quad \text { shaft speed : } 4000 \mathrm{RPM} \\
& \text { nominal clearance } \quad: \mathrm{C}_{0_{\mathrm{N}}}=0.36 \mathrm{~mm} \\
& \text { density } \quad: \rho=1000 \mathrm{~kg} / \mathrm{m}^{3} \\
& \text { preswirl ratio } \quad: \bar{w}(0, \theta) / R_{i N} \omega=0.3 \\
& \text { viscosity } \quad: \mu_{1}=1.0 * 10^{-3} \mathrm{Ns} / \mathrm{m}^{2} \\
& \text { entrance lost-coefficient : } \xi=0.5
\end{aligned}
$$

## CONCLUSIONS

We have shown that it is possible to calculate the dynamic coefficients of eccentric seals by a finite difference method based on the Navier Stokes equations. This method can also be extended to calculate the coefficients of eccentric gas seals by following the procedure of Nordmann (1987) and to calculate the dynamic coefficients of bearings by neglecting the turbulence model.



Fig. 6 Cross coupled stiffness $k_{z y}$ versus eccentricity.




Fig. 9 Cross coupled damping versus eccentricity.


## NOIENCLATURE

| $\mathrm{F}_{\mathrm{z}}, \mathrm{F}_{\mathrm{y}}$ | Forces on the shaft in $z$ and $y$ direction |
| :---: | :---: |
| K, k | direct and cross-coupling stiffness |
| D, d | direct and cross-coupled damping |
| M, m | direct and cross-coupling inertia |
| u,v,w | axial, radial and circumferential velocity |
| p | pressure |
| k | turbulence energy |
| $\epsilon$ | energy dissipation |
| $\mu_{\mathrm{e},}, \mu_{1}, \mu_{\mathrm{t}}$ | effective, laminar and turbulent viscosity |
| $\rho$ | density |
| t | time |
| x,r, $\theta$ | axial, radial and circumferential coordinate |
| x,y,z | rectangular coordinate directions |
| Y, Z | rotor displacements from its steady-state position |
| $\eta$ | radial coordinate after transformation |
| G | production term in $k$ - $\epsilon$-model |
| $\sigma_{k}, \sigma_{\epsilon}$ | constants of the $k$ - $\epsilon$-model |
| $\mathrm{C}_{\mu}, \mathrm{C}_{1}, \mathrm{C}_{2}$ | constants of the $k-\epsilon$-model |
| ${ }_{\text {¢ }}^{\text {S }}$ ¢ | general variable standing for $u, v, w, p, k$ general source term |
| ${ }^{\text {c }}$ | nominal seal clearance |
|  |  |
| $\mathrm{C}_{\mathrm{O}}^{\mathrm{N}}$ | steady-state clearance for an eccentric shaft |
| $\delta$ | varying seal clearance for orbiting shaft |
| $r_{0}$ | radius of the precession motion of the shaft |
| $e=\frac{r_{0}}{C_{o_{N}}}$ | perturbation parameter |
| $\omega$ | rotational frequency of the shaft $=R P M * \pi / 30$ |
| RPM | revolutions per minute |
| $\Omega$ | precession frequency of the shaft |
| $\xi$ | entrance lost-coefficient |
| L | Length of the seal |
| $\mathrm{r}_{\mathrm{a}}$ | stator radius (Fig. 1) |
| $\mathrm{R}_{\mathrm{a}_{\mathrm{N}}}$ | nominal stator radius (Fig. 1) |
| $\mathrm{R}_{\mathbf{i}}$ | shaft radius (Fig. 1) |
| $\mathrm{h}_{1}$ | change in the clearance |
| ex | eccentricity |
|  |  |
| w | average circumferential velocity |
| ${ }_{\text {u }}^{\text {En }}$ | average axial velocity at the entrance |
| $\mathrm{p}_{\text {Res }}$ | sump pressure |

APPENDIX A: Constants of the $k-\epsilon$ model.
$C_{\mu}=0.09 \quad C_{1}=1.44 \quad C_{2}=1.92 \quad \sigma_{k}=1.0 \quad \sigma_{\epsilon}=1.3$
and G is a given by
$G=\mu_{e}\left[2\left(\left(\frac{\partial v}{\partial r}\right)^{2}+\left(\frac{\partial u}{\partial x}\right)^{2}+\left(\frac{1 \partial w}{r \partial \theta}+\frac{v}{r}\right)^{2}\right)+\left(\frac{\partial v}{\partial x}+\frac{\partial u}{\partial r}\right)+\left(\frac{1 \partial v}{r \partial \theta}+\frac{\partial w}{\partial r}-\frac{w}{r}\right)^{2}+\left(\frac{\partial w}{\partial x}+\frac{1 \partial u}{r \partial \theta}\right)^{2}\right]$

APPENDIX B: Terms of first order source term.

$$
\begin{aligned}
& \mathrm{C}_{\mathrm{u}_{1}}=\frac{\partial}{\partial \mathrm{x}}\left(\rho \mathrm{u}_{0} \hat{\mathrm{u}}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \eta}\left(\eta \rho \mathrm{u}_{0} \hat{\mathrm{v}}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \theta}\left(\rho \mathrm{u}_{0} \hat{w}_{1}\right) \\
& \mathrm{C}_{\mathrm{v}_{1}}=\frac{\partial}{\partial \mathrm{x}}\left(\rho \mathrm{v}_{0} \hat{u}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \eta}\left(\eta \rho \mathrm{v}_{0} \hat{\mathrm{v}}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \theta}\left(\rho \mathrm{v}_{0} \hat{w}_{1}\right)-\rho_{\eta} \hat{w}_{0} \hat{w}_{1} \\
& \mathrm{C}_{\mathrm{w}_{1}}=\frac{\partial}{\partial \mathrm{x}}\left(\rho \mathrm{w}_{0} \hat{u}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \eta}\left(\eta \rho \mathrm{w}_{0} \hat{v}_{1}\right)+\frac{1}{\eta} \frac{\partial}{\partial \theta}\left(\rho \mathrm{w}_{0} \hat{w}_{1}\right)+{\frac{\rho}{w_{0}}}^{\eta} \hat{\mathrm{v}}_{1}
\end{aligned}
$$

$$
D_{u_{0}}=\left(\left(\left(r_{a}-\eta\right) D F C+R C S\right) U 1+O 0 c * U 2+0 C S * U 3\right)+i\left(\left(\left(r_{a}-\eta\right) D F S+R S S\right) U 1+O C S * U 2+000 * U 3\right)
$$

$$
D_{v_{0}}=\left(\left(\left(r_{a}-\eta\right) D F C+R C S\right) V 1+\infty 00 * V 2+0 C S * V 3\right)+i\left(\left(\left(r_{a}-\eta\right) D F S+R S S\right) V 1+O C S * V 2+000 * V 3\right)
$$

$$
D_{W_{0}}=\left(\left(\left(r_{a}-\eta\right) D F C+R C S\right) W 1+000 * W 2+0 C S * W 3\right)+i\left(\left(\left(r_{a}-\eta\right) D F S+R S S\right) W 1+\infty C S * W 2+\infty 00 * W 3\right)
$$

$$
D_{P_{o}}=\left(\left(\left(r_{a}-\eta\right) D F C+R C S\right) P 1-000 * P 2\right)+i\left(\left(\left(r_{a}-\eta\right) D F S+R S S\right) P 1-O C S * P 2\right)
$$

$$
\mathrm{DFS}=\frac{\mathrm{C}_{\mathbf{o}_{\mathrm{N}}}}{\overline{\mathrm{C}}_{\mathbf{o}}^{-}}\left[\cos (\theta)-\frac{\mathrm{ex}}{\bar{C}_{\mathbf{o}}}{ }_{*} \sin (\theta) \sin (\theta) *\left(1-\frac{\mathrm{ex} \cos (\theta)}{\sqrt{\overline{\mathrm{R}}_{\mathrm{aN}}^{2}-\operatorname{ex}^{2}-\sin ^{2}(\theta)}}\right)\right]
$$

$$
\mathrm{DFC}=-\frac{\mathrm{C}_{\mathrm{O}}}{\bar{C}_{\mathrm{o}}}\left[\sin (\theta)+\frac{\mathrm{ex}}{\mathrm{C}_{\mathrm{o}}} * \cos (\theta) \sin (\theta) *\left(1-\frac{\mathrm{ex} \cos (\theta)}{\sqrt{\mathrm{R}_{\mathrm{aN}}^{2}-\operatorname{ex}^{2}} \sin ^{2} \overline{(\theta)}}\right)\right]
$$

$$
\mathrm{RCS}=\operatorname{ex}_{\frac{\mathrm{C}_{0}}{} \overline{\mathrm{C}}_{\mathrm{o}}}^{-}\left[\operatorname { c o s } ( \theta ) \operatorname { s i n } ( \theta ) * \left(1-\frac{\operatorname{ex} \cos (\theta)}{\left.\left.{\sqrt{\mathrm{R}_{\mathrm{aN}}^{2}}-\operatorname{ex}^{2} \sin ^{2}(\theta)}_{(\theta)}\right)\right]}\right.\right.
$$

$$
\operatorname{RSS}=\operatorname{ex} \frac{C_{0} N}{C_{0}}\left[\sin (\theta) \sin (\theta) *\left(1-\frac{e x \cos (\theta)}{\sqrt{R_{a N}^{2}-\operatorname{ex}^{2} \sin ^{2}(\theta)}}\right)\right]
$$

$$
\begin{aligned}
& \alpha C=\frac{\mathrm{C}_{\mathrm{o}_{\mathrm{N}}}}{\mathrm{C}_{\mathrm{o}}} \cos (\theta) \quad \quad \alpha \operatorname{S}=\frac{\mathrm{C}_{\mathrm{o}_{\mathrm{N}}}}{\mathrm{C}_{\mathrm{o}}} \sin (\theta) \\
& \alpha C=-\Omega \frac{C_{o_{N}}}{C_{o}} \cos (\theta) \\
& O C S=\Omega \frac{C_{0_{N}}}{C_{0}} \sin (\theta) \\
& P 1=\frac{1 \partial}{\eta \partial \eta}\left(\rho w_{0}\right) \\
& \mathrm{P} 2=\frac{1 \partial}{\eta \partial \eta}\left(\eta \rho \mathrm{v}_{\mathrm{o}}\right)+\left(1-\frac{\mathrm{r}_{\mathrm{a}}}{\eta}\right) \frac{1 \partial}{\eta \partial \theta}\left(\rho \mathrm{w}_{\mathrm{o}}\right)-\frac{\mathrm{r}_{\mathrm{a}}}{\eta^{2}} \rho \mathrm{v}_{0} \\
& \mathrm{U} 1=\frac{1}{\eta}\left[\frac{\partial}{\partial \eta}\left(\rho \mathrm{u}_{\mathrm{o}} \mathrm{w}_{\mathrm{o}}\right)-\frac{\partial}{\partial \eta}\left(\frac{\mu}{\eta} \frac{\mathrm{e}{u_{0}}^{\partial \theta}}{\partial \theta}\right)-\frac{\partial}{\partial \eta}\left(\mu \frac{\partial w_{o}}{\partial \mathrm{x}}\right)\right]
\end{aligned}
$$

$$
\begin{aligned}
& -\frac{1 \partial}{\eta \partial \eta}\left(\eta \rho v_{o} u_{0}\right)+\frac{\mathrm{r}_{\mathrm{a}}}{\eta^{2}} \rho \mathrm{v}_{\mathrm{o}} \mathrm{u}_{\mathrm{o}}-\left(1-\frac{\mathrm{r}_{\mathrm{a}}}{\eta}\right) \frac{1 \partial}{\eta \partial \theta}\left(\rho \mathrm{u}_{\mathrm{o}} \mathrm{w}_{\mathrm{o}}\right) \\
& \mathrm{U} 3=\left(\mathrm{r}_{\mathrm{a}}-\eta\right) \frac{\partial}{\partial \eta}\left(\rho \mathrm{u}_{\mathrm{o}}\right) \\
& \mathrm{V}_{1}=\frac{1}{\eta}\left[\frac{\partial}{\partial \eta}\left(\rho \mathrm{v}_{\mathrm{o}} \mathrm{w}_{0}\right)-\frac{\partial}{\partial \eta}\left(\frac{\mu}{\eta} \frac{\partial \mathrm{v}_{0}}{\partial \theta}\right)-\frac{\partial}{\partial \eta}\left(\mu \mathrm{e} \frac{\partial}{\eta \eta}\left(\frac{w_{0}}{\eta}\right)\right)\right] \\
& \mathrm{V} 2=\left(1-\frac{\mathrm{r}_{\mathrm{a}}}{\eta}\right)\left[\frac{1 \partial}{\eta \partial \theta}\left(\frac{\mu}{\eta} \frac{\partial \mathrm{v}_{0}}{\partial \theta}+\mu_{\mathrm{e}} \eta^{\frac{\partial}{\partial \eta}}\left(\frac{w_{0}}{\eta}\right)\right)-\frac{2}{\eta}\left(\frac{\mu}{\eta} \frac{\mathrm{e}^{\partial w_{0}}}{\partial \theta}+\frac{\mu_{\mathrm{e}}}{\eta} \mathrm{v}_{\mathrm{o}}\right)-\frac{1 \partial}{\eta \partial \theta}\left(\rho \mathrm{v}_{\mathrm{o}} \mathrm{w}_{\mathrm{o}}\right)\right] \\
& -2 \frac{r_{a}}{\eta^{2}} \mu \frac{\partial v_{0}}{\partial \eta}+\frac{1 \partial}{\eta \partial \eta}\left(2 \pi \mu e \frac{\partial v_{0}}{\partial \eta}\right)-\frac{1 \partial}{\eta \partial \eta}\left(\eta \rho v_{o} v_{0}\right)+\frac{r_{a}}{\eta^{2}} \rho v_{o} v_{0}+\frac{\eta-r_{a}}{\eta} \frac{\rho_{w_{0}} w_{0}}{\eta}-\frac{\partial p_{0}}{\partial \eta} \\
& \mathrm{v} 3=\left(\mathrm{r}_{\mathrm{a}}-\eta\right) \frac{\partial}{\partial \eta}\left(\rho \mathrm{v}_{\mathrm{o}}\right) \\
& W_{1}=\frac{1}{\eta}\left[\frac{\partial}{\partial \eta}\left(\rho w_{o} w_{0}\right)-\frac{\partial}{\partial \eta}\left(2 \frac{\mu_{e}}{\eta} \mathbf{v}_{0}\right)-\frac{\partial}{\partial \eta}\left(2 \frac{\mu}{\eta} \frac{\partial w_{0}}{\partial \theta}\right)+\frac{\partial p_{0}}{\partial \eta}\right]
\end{aligned}
$$

$$
\begin{aligned}
& \mathrm{w} 3=\left(\mathrm{r}_{\mathrm{a}}-\eta\right) \frac{\partial}{\partial \eta}\left(\rho \mathrm{w}_{\mathrm{o}}\right)
\end{aligned}
$$

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