Quantum Well Infrared Photodetectors (QWIP)

B. F. Levine AT&T Bell Laboratories Murray Hill, NJ 07974

There has been a lot of interest in III-V long wavelength detectors in the $\lambda = 8$ -12 µm spectral range as alternatives to HgCdTe.¹⁻⁶ Recently high performance quantum well infrared photodetectors (QWIP) have been demonstrated. They have a responsivity of R = 1.2 A/W, and a detectivity $D_{\lambda}^{\star} = 2x10^{10}$ cm Hz^{1/2} / W at 68 K for a QWIP with a cutoff wavelength of $\lambda_c = 10.7$ µm and a R = 1.0 A/W, and $D_{\lambda}^{\star} = 2x10^{10}$ cm Hz^{1/2} / W at T = 77 K for $\lambda_c = 8.4$ µm. These detectors consist of 50 periods of MBE grown layers doped n = 1x10¹⁸ cm⁻³ having GaAs quantum well widths of 40 Å and barrier widths of 500 Å of Al_xGa_{1-x}As.

Due to the well-established GaAs growth and processing techniques these detectors have the potential for large, highly uniform, low cost, high performance arrays as well as monolithic integration with GaAs electronics, high speed and radiation hardness.

Our latest results on the transport physics, device performance and arrays will be discussed.

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Quantum Well Infrared Photodetectors QWIP

Research

B. F. LevineC. G. BetheaS. D. GunapalaR. J. MalikG. Hasnain

Government Systems

C. L. Allyn V. O. Shen

Development

P. J. Anthony W. A. Gault J. W. Stayt K. G. Glogovsky R. A. Morgan Y. M. Wong M. T. Asom S. J. Hsieh R. M. Braun

LWIR GaAs Quantum Well Detectors

Esaki, Sakaki

Smith, Chiu, Margalit, Yariv, Cho

Coon, Karunasiri

Goosen, Lyon

Capasso, Mohammed, Cho

Kastalsky, Duffield, Allen, Harbison

Janousek, Daugherty, Bloss, Rosenbluth, O'Loughlin, Kauter, DeLuccia, Perry

Woodall

Wu, Sato, Wen

Maserjian

Döhler

Mii, Karunasiri, Wang, Bai

Abstreiter et al.

MATERIAL FOR 10 μ m DETECTORS HgCdTe DETECTORS

- DIFFICULT GROWTH AND PROCESSING TECHNOLOGY
- POOR UNIFORMITY OF ARRAYS
- LOW QUALITY CdTe SUBSTRATES

GaAs DOPED QUANTUM WELL DETECTORS

- PERFORMANCE COMPARABLE TO HgCdTe
- MATURE GROWTH AND PROCESSING TECHNOLOGY
- EXCELLENT 3" GaAs SUBSTRATES
- MONOLITHIC INTEGRATION WITH GaAs ELECTRONICS

MONOLITHICALLY INTEGRATED GaAs QUANTUM WELL DETECTOR ARRAY AND IMAGE PROCESSING ELECTRONICS



VERTICALLY INTEGRATED GaAs QUANTUM WELL INFRARED SPECTROMETER



INFRARED RADIATION







DARK CURRENT CALCULATION

$$n(V) = \frac{m^*}{\pi \hbar^2 Lp} \int_{E_0}^{\infty} f(E)T(E)dE$$

- $\mathrm{E} > \mathrm{E}_{\mathrm{b}}$ Thermionic
- $E < E_b$ Tunneling

$$I_D = nevA$$







$$\lambda_{\rm c} = 10 \ \mu {\rm m}$$

D*(theory) > $10^{10} \ {\rm cm} \sqrt{{\rm Hz}} / {\rm W}$

 $T=\,77~{\rm K}$

• *

















Optical Gain

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$$g = \frac{\tau_{\rm L}}{\tau_{\rm T}} = \frac{L}{\ell}$$







NOISE EQUIVALENT TEMPERATURE CHANGE

NE $\Delta T = \frac{(A\Delta f)^{1/2}}{D_B^* (dP_B/dT) \sin^2(\theta/2)}$

 $A = (50 \ \mu m)^2$

 $\Delta f = 60 \text{ Hz}$

$$f/2 \text{ optics } (\theta/2 = 14^{\circ})$$

 $D^* = 1 \times 10^{10} cm \sqrt{Hz} / W$

 $NE\Delta T = 0.01 K$

ARRAY NONUNIFORMITY

To Obtain Background Limited Array Performance

$$U < \frac{1}{\sqrt{N}}$$

U = uniformity
N = number of photoelectrons
N = 10⁶ \Rightarrow U < 0.1%
(NE Δ T)_U = $\frac{T_B^2 \lambda U}{1.44}$
T_B = 295 K, λ = 10 μ m, U = 0.1%
(NE Δ T)_U = 0.06 K







Conclusions

- Demonstrated detectors having $\lambda_c = 4 13.5 \ \mu m$
- Spectral width $\Delta \nu / \nu = 13\% 36\%$
- $D_{BB}^* = 1 \times 10^{10} \text{ cm} \sqrt{\text{Hz}} / \text{W}$ T = 68 K $\lambda_c = 10.7 \ \mu\text{m}$
- $D_{BB}^* = 3 \times 10^{10} \text{ cm} \sqrt{\text{Hz}} / \text{W}$ T = 77 K $\lambda_c = 8.4 \ \mu \text{m}$
- $D_{BB}^* = 1 \times 10^{13} \text{cm} \sqrt{\text{Hz}} / \text{W}$ T < 40 K $\lambda_c = 10.7 \ \mu\text{m}$
- D^{*} sufficiently large (arrays uniformity limited)
- Calculated dark current (thermionic, tunneling)
- Hot electron continuum transport resonances
- High speed $\tau < 200$ psec
- Optical gain
- Graded barrier tunable spectral response
- Demonstrated grating detectors
- High uniformity
- Large arrays
- Camera demonstration