NUCLEAR THERMAL ROCKET WORKSHOP REFERENCE SYSTEM -ROVER/NERVA-

Dr. Stanley K. Borowski NASA Lewis Research Center

INTRODUCTION

The Rover/NERVA engine system is to be used as a "reference," against which each of the other concepts to be presented in this workshop will be compared. In this presentation I'll review the operational characteristics of the nuclear thermal rocket (NTR), the accomplishments of the Rover/NERVA programs, and performance characteristics of the NERVA-type systems for both Mars and lunar mission applications. I'll also briefly touch on the issues of ground testing, NTR safety, NASA's nuclear propulsion project plans, and NTR development cost estimates before concluding my presentation.

NERVA REFERENCE ENGINE

The NTR is basically a monopropellant liquid rocket system which utilizes a nuclear reactor core for power generation and propellant heating (Figure 1). High pressure hydrogen from a turbopump assembly passes through a high power reactor core where it is heated to high temperatures and then exhausted through a convergent-divergent nozzle at high speeds to produce thrust. Before entering the reactor core, hydrogen flowing from the pumps is first "preheated" by cooling the nozzle, reflector, control rods, peripheral shield, and core support structure.

In the "hot bleed cycle" (see Figure 2), this preheated hydrogen is routed down though the reactor core for heating to design temperatures and subsequent nozzle expansion. Approximately 3% of the heated hydrogen is diverted from the nozzle plenum chamber, cooled, and then used to drive the turbopumps with the exhaust being utilized either for roll control or readmitted into the diverging portion of the nozzle for additional thrust generation. In the "full flow topping" or "expander cycle" engine, the preheated hydrogen is routed to the turbopumps and then through the reactor core with the entire propellant flow being heated to design temperatures (Figure 2) providing more optimum performance in terms of higher engine specific impulse (I_{sp}).

The accomplishments of the Rover/NERVA program are summarized in Figures 3, 4, and 5. As Figure 3 indicates, the achievements were quite impressive with a total of 20 rocket reactors designed, built, and tested between 1955 and 1973 at a cost of \$1.4 Billion. From program start in 1955 to testing of the first KIWI-A reactor was only 4

years which is pretty impressive in itself. Major performance accomplishments were demonstrated in the areas of power and thrust levels, peak and fuel exit temperatures and equivalent specific impulse, and full power burn duration. Most notable was the NERVA program's NRX-A6 test in which the system operated for 62 minutes at a thrust level of about 55,000 pounds-force (55klbf) and a thermal power level of about 1125 megawatts (MWt).

The NERVA program's NRX series of reactors culminated in the downward test firing of the Experimental Engine Prototype (the XE-P) in 1969. The NRX-XE underwent 28 startup/shutdown cycles and demonstrated rather convincingly the practicality of the NTR systems. In addition to these "full scale" integrated engine tests, electric and nuclear furnace (NF-1) tests were also conducted in an effort to develop higher temperature/longer life reactor fuels. Anticipated performance for the "composite" and "carbide" fuel forms, which you will be hearing about at this workshop, is about 10 hours at I_{sp} values of about 925 seconds and 1020 seconds for the composite and carbide fuel forms, respectively.

Again, 20 reactors were tested in the Rover/NERVA programs and the chronology of system tests for both programs is shown in Figure 4. After demonstrating feasibility of the basic KIWI-B series concept, the Los Alamos Rover program concentrated its efforts on fuel research and higher power density systems. The Phoebus-1B system, tested in 1967, was approximately the same physical size as KIWI-B (see Figure 5) but was operated at 1500 MWt. Phoebus-2A (shown in Figures 5 and 6), was designed for 5000 MWt and 250 klbf. It was operated at about 80% of its rated design conditions for about 12.5 minutes in July 1968 and was the most powerful nuclear rocket reactor ever built. It was to be the prototype for the 200-250 klbf-class NERVA II engine contemplated by NASA at that time. Figure 6 is a picture of Phoebus-2A being transported to "Test Cell C" (Figure 7) on the Jackass & Western Railroad for full power testing.

A final noteworthy reactor system was the Nuclear Furnace (NF-1). It was operated in 1972 at about 44 MWt and was utilized primarily as a inexpensive "test bed" system for screening advanced fuels and reactor structural materials. A special feature of the NF-1 reactor was its "effluent cleanup system" which effectively removed radioactive contaminants from effluent reactor gas. The database provided by the Nuclear Furnace is of particular interest today because of environmental restrictions which would prevent open-air testing.

Figures 8, 9, and 10 show three of the six NRX-series reactor systems developed by Aerojet and Westinghouse for NASA and the AEC during the Nuclear Engine for Rocket Vehicle Application (NERVA) program. Figure 8 shows the NRX-A3 being prepared for test firing at Test Cell C at the Nuclear Rocket Development Station (NRDS) at Jackass Flats, Nevada. Figure 9 shows the 62 minute "continuous full-power burn" of the NRX-A6 system in December 1967 with its two large 500,000 gallon liquid hydrogen tanks off to the right. Last, Figure 10 shows the XE prototype engine installed for downward test firing at the ETS-1 test facility also at the NRDS.

The very large database accumulated in both the Rover/NERVA programs was integrated into a reference NERVA engine design in 1972. A mockup of the 1972 NERVA is shown in Figure 11. The fuel form was coated UC₂ particles in a graphite matrix, the chamber pressure was 450 psia, and hydrogen exhaust temperatures from the reactor ranged from 2,350 to 2,500 K. Both hot bleed and expander cycle versions of the 1972 NERVA were examined with I_{sp} values ranging from 825 to 870 seconds. The engine shown in Figure 11 had an overall length of about 10.5 meters with a 100-to-1 nozzle expansion ratio; it weighed a little over 11 metric tons, resulting in an engine thrust-to-weight ratio of 3. In terms of NASA's technology maturity ranking, the XE engine was rated at an overall system technology readiness level of about 6 (TRL=6 is the prelude to the next development step, which is the "flight engine"). Some of the NRX components were rated at about the TRL=5 level and required some further development (see Figure 12).

On the "non-nuclear" subsystem side, there have been major advances in chemical rocket technology in the 17 years since termination of the NERVA program. Of particular note are the significant performance improvements and accompanying weight reductions in the turbopump and nozzle areas. Figure 13 compares the Space Shuttle Main Engine (SSME) and the 1972 NERVA. You can see that the SSME nozzle is lighter and is capable of handling exhaust gas temperatures in excess of 3,100 K (equivalent to those anticipated from the advanced carbide fuels). It also operates with heat fluxes four times greater than those encountered in the NERVA program. Pump discharge pressures from the SSME hydrogen turbopump are also a factor of 5 greater than those of the 1972 NERVA. Chemical propulsion system development has therefore provided us with a significant database for use in the design of current day NERVA-type engine systems. Performance projections for "state-of-the-art" NERVA derivative reactor systems are shown in Table 1. Assuming a full-flow expander cycle engine operating at about 1000 psia, the I_{sp} values for a 500-to-1 nozzle expansion ratio vary from about 850 to 885 seconds for graphite fuel, about 925 seconds for the composite fuel, and about 1020 seconds for the pure carbide fuel form. Higher performance/lower weight non-nuclear components also result in a 2 to 3 metric ton savings in overall engine mass and the improved engine thrust-to-weight ratios shown.

REFERENCE MARS MISSION ANALYSIS

I would now like to review with you the results of trajectory and mission analysis work performed at the Lewis Research Center for the reference Mars mission. Both 1972 vintage and "state-of-the-art" NERVA-type systems were examined. But first I'd like to briefly show you some previous NASA work in this area from the 1960-1970 time frame to set the stage for the current results I will be showing you shortly. I'll also point out

the many similarities that exist between these earlier studies and our current day results. In August of 1969, just one month after the Apollo 11 moon landing, Werner von Braun described NASA's proposal for a piloted mission to Mars (around 1981) at a hearing of the Senate Committee on Aeronautics and Space Science. The mission would be accomplished using two spacecraft, each carrying a 6-person crew and having an initial mass in low Earth orbit (IMLEO) of about 727 tons. Each spacecraft would carry three 445 kilonewton (about 100klbf) NERVA-class engines (with an Isn of 850 seconds) of which two would be used only for departing Earth orbit for the 270-day journey out to Mars. After this trans-Mars insertion (TMI) burn, the two strap-on NERVA-powered booster stages would separate, retrofire, and return to Earth for liquid hydrogen refueling and reuse (see Figure 14). Subsequent mission maneuvers would be accomplished by the remaining NERVA engine on the core spacecraft. Later mission studies assumed a single 75klbf-class NERVA engine for spacecraft propulsion (see Figure 15), and a multiple perigee burn Earth departure scenario was adopted. Two large tanks attached to the core spacecraft would carry the TMI propellant and would be jettisoned after completion of the TMI maneuver. The remaining propellant would be accommodated in the central core tank(s).

The mission profile proposed by von Braun was a 640-day opposition class mission with an 80-day stay at Mars and inbound Venus swingby. Twenty-one years later, NASA's reference Mars mission scenario is a 2016 opposition class mission with 30-day surface stay and an inbound Venus swingby (see Figure 16). For this particular opportunity, the overall mission duration is attractive--on the order of 434 days. Most opposition class missions have mission durations somewhere in the 420- to 650-day ballpark.

The 2016 reference NTR mission profile originally assumed for the workshop is shown in Figure 17. The "all propulsive" NTR vehicle features expendable TMI and Mars orbital capture (MOC) tanks attached to an optional central truss structure. Trans-Earth injection and Earth orbital capture (EOC) propellant would be contained in a common core propellant tank in the vehicle "reuse" mode. In the "expendable" vehicle mode, the return of the crew to Earth could be accomplished utilizing an Earth Crew Capture Vehicle (ECCV).

The mission assumption and ground rules are shown in Table 2 and the propulsion system, boil off, and tankage assumptions are summarized in Table 3. Because our principle "figure-of-merit" for this analysis is IMLEO, a single 75klbf NERVA-class engine has been assumed as the baseline engine thrust level, along with perigee propulsion. By utilizing a multi-perigee burn departure scenario, we can more effectively impart propulsive energy to our spacecraft while reducing gravity losses associated with the finite burn durations accompanying lower thrust-to-weight ratio vehicle designs.

The motivation for going to multiple perigee burns with lower thrust engine systems is illustrated quite dramatically in Figure 18. If we tried a "one burn" Earth departure maneuver using a single 75klbf engine with a vehicle thrust-to-weight ratio of about 0.05,

gravity losses ("g-losses") would add 1500 meters per second (m/s) to the ideal TMI Delta-V requirement. By going to the "3 perigee burn" approach, g-losses are reduced to about 350 m/s. The actual g-loss value will vary, of course, depending on the mission C_3 requirement, the I_{sp} of the NTR, the orbital departure altitude, and the vehicle thrust-to-weight ratio. By using a single higher thrust engine or by clustering several lower thrust engines, the vehicle thrust-to-weight ratio can be increased, and single burn departure scenarios are possible with acceptable g-loss. As will be shown later in this talk, a single 250klbf Phoebus-2A class NTR can perform the 2016 Mars mission opportunity for an IMLEO of about 750 tons using a single burn Earth departure. With a thrust-to-weight ratio of about 0.15, the g-losses incurred during TMI are on the order of 400 m/s.

The "reference trajectory" assumed for this workshop (and shown in Figure 16) was originally established during the "90-Day Study" for the aerobrake chemical vehicle that was baselined at that time. The trajectory was subsequently adjusted somewhat for the NTR analysis purposes, although it was by no means optimum. An aerobrake-optimized trajectory weights both the arrival velocities at Mars and Earth more heavily since it assumes that a lightweight, high, heat-flux-resistant aerobrake will be developed in the future. By weighting the MOC and EOC velocities more heavily, the TMI and TEI Delta-V requirements can be reduced, thereby compensating for the limited capability of the chemical propulsion system. Table 4 summarizes trajectory data and associated IMLEO estimates for both the "doctored-up" NTR reference trajectory and a new "all propulsive optimized" NTR trajectory recently developed by Lewis Research Center's Advanced Space Analysis Office. The NTR optimized trajectory weights the departure maneuvers from Earth and Mars more heavily than the capture maneuvers thereby exploiting more fully the high I_{sp} capability of the NTR system.

Estimates of IMLEO from Marshall Space Flight Center's contractor, Boeing, and from the Lewis Research Center (LeRC) are shown for the reference trajectory and a "state-of-the-art" composite fuel NERVA derivative system operating at an I_{sp} of about 925 seconds. The Boeing estimate for IMLEO is about 735 tons and is based on the assumption of a fixed 200 m/s g-loss value and use of advanced composite cryogenic tanks. The LeRC IMLEO estimate is somewhat higher because of a more accurate g-loss estimate and different tankage assumptions. What is most impressive, however, is the impact on IMLEO of using the "all propulsive optimized" trajectory that results in a 150-ton mass savings!

A comparison of vehicle size for the 2016 Mars mission using the optimized and nonoptimized trajectories of Table 4 are shown in Figure 19. The two TMI drop tanks are limited in size to the payload shroud dimensions of anticipated heavy lift launch vehicles currently under study and are approximately 10 meters in diameter by about 30 meters in length.

The performance potential of different 75klbf-class NERVA engines of the type shown in Table 1 were examined and compared in terms of IMLEO and total engine burn time

requirements for the "all propulsive optimized" 2016 Mars trajectory described in Table 4. The results for "state-of-the-art" NERVA derivative reactor (NDR) systems using an expander engine cycle and a variety of fuel forms (graphite, composite, and carbide) are shown in Figure 20. At a 1000 psia chamber pressure and a 500-to-1 nozzle expansion ratio, a "current day" graphite NERVA system operating at 2,350 K (a temperature routinely demonstrated in the NERVA program) would deliver an I_{sp} of 850 seconds. The associated IMLEO and engine burn time for this system is 725 tons and 3.38 hours, respectively. Going to the higher performance composite and carbide fuel forms, the IMLEO and burn time requirements decrease to 613 tons/2.99 hours and 518 tons/2.64 hours, respectively. These values are to be compared to the reference aerobrake chemical vehicle from NASA's "90-Day Study" which had an IMLEO of about 752 tons for the expendable ECCV Earth return option, and about 830 tons for the reusable propulsive return option. The aerobrake mass fraction assumed for the MOC aerobrake was about 13 percent, which is also somewhat optimistic.

A "state-of-the-art," graphite fuel NDR engine propulsively returning the basic core spacecraft to LEO can therefore outperform the best aerobraked chemical vehicle design currently on the "drawing boards" by 27 tons when the chemical/aerobrake vehicle is operated in the expendable ECCV recovery mode, and by 105 tons in the vehicle reuse mode. Even the 1972 graphite fuel NERVA design outperforms the aerobraked chemical vehicle in the reuse mode with an IMLEO and engine burn time of about 755 tons and about 3.75 hours, respectively.

The relative vehicle size comparison for the graphite, composite, and carbide fuel NDR systems is shown in Figure 21. The individual burn duration for both 75klbf and 250klbfclass NTR systems are summarized in Table 5, and the relative vehicle sizes for the "3 perigee burn" 75klbf and "one burn" 250klbf-class NTR systems are shown in Figure 22. The 75klbf and 250klbf engines both assume a 1000 psia chamber pressure and a 500-to-1 nozzle expansion ratio, and utilize a composite fuel capable for delivering 925 seconds of I_{so} .

In contrast to the approximately 3-hour total engine burn duration for the composite fuel 75klbf NDR system, the 250klbf engine burn time totals a little over one hour at 65.3 minutes. The IMLEO requirement of 749 tons is comparable to that of the expendable aerobrake chemical vehicle due to the higher g-loss accompanying the "one burn" departure scenario and the heavier weight (about 21.8 tons) of this higher thrust engine. Perigee propulsion can reduce the IMLEO requirements further, at the expense of the more complex "3 burn" departure scenario.

Other Mars mission opportunities have been examined besides the 2016 opportunity in order to assess the magnitude of IMLEO variation across a synodic period. Figure 23 shows the sensitivity of IMLEO to mission roundtrip time (for a 925-second NTR system with multiple perigee burns) for a variety of mission modes and two different opportunities--an easy one (2018) and a tough one (in 2014). The mission modes

examined include a reusable, all propulsive mode, one with an ECCV for Earth return, and a split mission in which cargo is carried on a "minimum energy" conjunction-class trajectory while the piloted portion of the mission travels a faster, higher energy opposition-class trajectory. Stay times at Mars are in all cases assumed to be 30 days. This split-type mission is often referred to as the "split-sprint." A more advanced (but potentially greater risk) variation of the split mission involves having the cargo vehicle also carry the "return propellant" for the piloted vehicle. This variation was referred to during the 1960's as the "Hohmann tanker/dual vehicle" mission mode.

As we push from 434 days to round trip times on the order of one year, the IMLEO for the all-propulsive single vehicle case in 2018 almost doubles increasing from about 700 tons to about 1350 tons. By utilizing an ECCV for Earth return, one can shave off about 300 tons from the IMLEO requirement for the one-year mission. In the split-sprint mission mode the piloted vehicle IMLEO is on the order of 375 tons for the one-year mission although the total IMLEO requirement including the cargo vehicle is on the order of 750 tons. Even in the most difficult mission year of 2024, trip times from 400 to 500 days are possible with the various mission modes available. This is an important operational advantage of the NTR system over NEP systems--the ability to shorten trip times across the entire spectrum of Mars mission opportunities using a technology with a proven experimental database.

LUNAR MISSION ANALYSIS

Lewis Research Center has also been conducting "in-house" and contracted study efforts aimed at assessing the benefits of using NTR technology for lunar mission applications. During the "90-Day Study" the establishment of a lunar outpost was considered a prelude to undertaking missions to Mars. The flight schedule for the proposed lunar outpost scenario covered a 15-year period and required 30 separate flights involving either cargo, piloted, or combination missions (see Figure 24). The base line piloted Lunar Transportation Vehicle (LTV) in the 90-Day Study utilized chemical propulsion and required an aerobrake for Earth return to keep the IMLEO within a reasonable range (see Figure 25). The IMLEO for the first piloted lunar missions, which was used to size the system, was about 194 tons.

In the next several vugraphs you'll see some of the findings resulting from our contracted effort with SAIC. The specific mission and NTR system definition assumptions used in the SAIC study are shown in Figure 26 and 27, respectively, and a comparison of the IMLEO requirements for the first piloted mission using aerobraked chemical and NTR technologies is summarized in Figure 28. Figure 28 shows a mass savings of about 32 tons using an NTR-powered LTV in a "4 burn" all-propulsive lunar mission profile. By "4 burn" we refer to the four major propulsive maneuvers of trans-lunar injection, lunar orbit capture, trans-Earth injection, and Earth orbit capture.

In the SAIC study, the mass penalty associated with disposing of "end-of-life" NTR systems was also assessed and included in the IMLEO comparisons. A number of disposal modes were examined using 1-, 2-, 3-, and 4-burn lunar NTR scenarios, and the results are shown in Figure 28. One can see that disposing of the spent NTR propulsion module (consisting of a small propellant capacity run tank, an avionics package, and the NTR) into a 1,000 kilometer parking orbit (following Earth orbit capture of the NTR vehicle back into LEO) results in a modest 2-ton penalty. The mass penalty increases for the more demanding disposal modes into heliocentric and super-geo orbits. The overall impact on IMLEO is modest, however, compared to the chemical/aerobrake baseline system.

The overall mass savings resulting from using NTR technology in the lunar outpost scenario is summarized in Figure 29. Over a 15-year flight schedule, the total computed mass delivered to LEO for the reference aerobraked chemical LTV system was in excess of 5,000 metric tons. Using a conservative NTR growth assumption (I_{sp} of 900 seconds and nozzle expansion ratio of 200-to-1), a "4 burn", all-propulsive NTR LTV system would reduce the delivered mass to LEO to about 4040 tons--a savings of approximately 20 percent.

Since it's probably going to be tough to have the NTR system ready for the proposed first piloted mission in the early 2000's, without a major commitment of resources, the SAIC study also looked at "phasing in" the NTR system into the reference 90-Day Study scenario. This approach would still provide an IMLEO savings and would also provide valuable operational experience in the use of NTR systems in a "nearby" space environment prior to undertaking the more demanding Mars mission. Even with the phased NTR approach, a 15 percent IMLEO savings is indicated with disposal penalties again taken into consideration.

TESTING

In my last few vugraphs I would like to touch briefly on a number of peripheral issues that are very important. The first deals with the ground testing of full scale integrated reactor and flight engine systems. It is obvious that we cannot operate as we did in the past at NRDS with "open air" testing. The Nuclear Propulsion Project will therefore have to address a number of programmatic and development issues associated with NTR ground testing (see Figures 30 and 31). Concepts for "fully contained" test facilities have been proposed based on the earlier Nuclear Furnace experience. A schematic for one such facility, proposed by the Idaho National Engineering Laboratory, is shown in Figure 32. The facility would contain a number of debris traps, water sprays, cooler/scrubbers, filters and charcoal beds for removing particulates, soluble fission products, and noble gases from the engine exhaust prior to the hydrogen being released to the burn stack. Another option for confining engine exhaust gases might be to use some of the weapons test tunnels at the Nevada Test Site. Tunnel testing could have a number of advantages (Figure 33), and its usefulness for NTR testing will have to be assessed more fully in the future. A number of NASA, DOE and industry people visited the Nevada Test Site about a month ago and toured a weapons test tunnel and portions of the NRDS at Jackass Flats. There are a lot of site assets that still exist at the NRDS (see Figures 34 and 35) that could be put to good use in a future NTR development program.

With regard to NTR safety, the Rover/NERVA programs had an exemplary safety record handling large quantities of liquid hydrogen (on the order of a million gallons or more during some engine tests) and large radioactive systems remotely in its E-MAD facility during the post irradiation disassembly and fuel examination periods. The 1972 NERVA reference engine was also designed to be a "man-rated" system and included redundant turbopumps and valve sets (see Figure 36). Probablistic design and failure mode effects analyses were also done. The NERVA system that resulted from this analysis approach (see Figure 37) had good component redundancy to eliminate a number of identified failure modes that could develop during various phases of a typical lunar mission that was selected by NASA for its Design Reference Mission. A good database and starting point for a "man-rated" NTR system can therefore be found in the NERVA program.

Another issue that has surfaced recently deals with the diffusion of fission product gases from the NTR system during powered operation and the overall dose rates experienced by the crew of an NTR-powered spacecraft during a typical Mars mission. Although work is just being restarted in this area, Figure 38 provides us with some rough numbers. Shown is the temporal variation of dose rate for the "non-optimized" 2016 Mars reference mission that was originally assumed for this workshop. The burn duration for the major maneuvers and the approximate elapsed time between burns is shown at the top of the figure; the variation of dose rate experienced by a crew member standing 100 feet away from the unshielded reactor core center-line (a rather pessimistic assumption) is shown at the bottom. It is quite evident that during the full power TMI burn, the dose rate is lethal. One day after TMI, however, the dose rate has dropped by a factor of 6500, and after the 156-day coast period to Mars it is down to 0.23 Rem/hour. Following the MOC burn, the crew would depart the Mars spacecraft staying within the protected cone area provided by the NTR engine's external disk shield. After a 30-day surface stay, the returning Mars excursion vehicle could fly past the unshielded NTR and receive less than 2 Rem/hour at the 100-foot separation distance. Following the TEI burn-and-coast phase, the dose rate at our reference location is on the order of 75 millirem per hour prior to EOC. Up in the front of the vehicle where the crew will actually be located, the benefits of the external disk shield, core propellant tank, truss structure distance, and solar flare storm shelter will reduce overall accumulated crew dose to the required 5 Rem per year.

Because the NTR system is a high-thrust system, it provides all of its impulse to the

spacecraft quickly, unlike the NEP systems that must operate for a major portion of the total mission time--on the order of 10,000 to 20,000 hours. As a result of the NTR system's short burn duration, the radioactive inventory has a significant period of time to decay, thereby reducing the system's overall radiological hazard.

PROJECT PLANNING

We are working and reworking the Project Plan, taking into account inputs from industry sources, NASA sources, and DOE inputs. Our earlier speaker, Gary Bennett, outlined a three-phase program in which the important project elements are system development, nonnuclear component development and nuclear component development.

Obviously, a number of critical tests have to be done right up front. Facilities requirements must be defined in the first couple of years. We need to identify not only the components to be tested on the ground, but also the big ticket items, such as the ground test facility for doing the integrated and full scale engine tests.

Also we will include innovative technology (aimed at 2nd and 3rd generation systems) throughout a good part of the first two phases; we will also be conducting mission studies for a good portion of the early phases, identifying system concepts, and going through preliminary, critical and final design reviews. Potentially there will be a design freeze in which we could be really focusing in on the component and subsystem tests that will be tested in the latter years. Then ultimately, we get into reactor tests.

The NERVA program cost \$1.4 billion; escalating that to today's dollars would be almost \$10 billion. However, it is important to remember that the NERVA program was a gold-plated program; whole integrated reactors were put together just to test improvements in coating. We think there are better ways to do that with smaller subscale electric furnace, and nuclear furnace tests. Plus, there is now an established database, so while we have to reverify it, I don't know that it's necessary for us to go through the same number of tests. Obviously we must develop a Project Plan in the course of the next couple of months and over the course of the first few years. Also, a number of critical nonnuclear and nuclear component tests have to be done.

DEVELOPMENT COST

My first estimate on the cost of this program is close to \$3 billion to take it to technology level readiness 6. Somebody might get up and say they think it's more like 5 billion and I wouldn't argue very strongly. I think the results of this workshop will pull in a lot more information for us to make a more informed judgment on what the program will realistically cost.

Again, I think a critical thing in the program is the facility cost for the full scale engine test. We are certainly going to need a study by an unbiased major contractor who has experience in doing the large scale nuclear facilities.

CONCLUSION

My last vugraph (Figure 39) summarizes my conclusions and observations. The Rover/NERVA programs definitely established an impressive database that demonstrated convincingly the feasibility of the graphite core NTR concept. This database was used in putting together the 1972 NERVA reference engine design. Based on our analysis a "state-of-the-art," graphite core NDR system would have and IMLEO of 725 tons which is 105 tons lighter than the best aerobrake chemical system that NASA can envision today. Even 1972 NERVA can outperform it.

The ground test experience gained during the Rover/NERVA programs was substantial even though most of it was done in the open air. The Nuclear Furnace experiment with its effluent control system provides us with an important database for designing a "contained" test facility meeting today's environmental standards.

With the continued advances in chemical propulsion technology over the last 17 years, higher performance/lighter weight turbopumps, nozzles, and valves should help to improve the engine thrust-to-weight ratio for today's NERVA derivative engine. One should not overlook the impact of a radiation environment on component performance that could present some unforseen problems in a future development effort.

The NTR is an *enabling* technology for future piloted missions to Mars. It can shorten roundtrip mission times substantially allowing one-year missions to be contemplated. We also think that the NTR is *enhancing* for lunar mission applications, providing not only IMLEO savings but valuable operational experience with this impressive new propulsion technology.

A Nuclear Propulsion Program will certainly require a lot of work and a significant infusion of resources to become a reality. For the NTR I think test facilities are the key item with high-temperature fuel development being very important also.

Lastly, I'd like to point out that the projected performance parameters for NTR that we have been using in our analyses thus far are within a factor of 2 or less of those already demonstrated in the Rover/NERVA programs. This provides real confidence that piloted missions to the Moon and Mars will someday be a reality with the NTR system!

BIBLIOGRAPHY

Stanley K. Borowski NTR Workshop Reference System

- 1. P. G. Johnson, "Beyond Apollo with Nuclear Propulsion," <u>Astronautics and Aeronautics</u>, Dec. 1964, p. 22.
- 2. H. B. Finger, "Nuclear Rocketry-Confidence Substantiated," <u>Astronautics and Aeronautics</u>, June 1965, p. 34.
- 3. N. Y. Jordan, Jr., R. J. Harris, and D. R. Saxton, "Toward Modular Nuclear-Rocket Systems," <u>Astronautics and Aeronautics</u>, June 1965, p. 48.
- 4. R. W. Schroeder, "NERVA-Entering a New Phase," Astronautics and Aeronautics, May 1968, p. 42.
- 5. L. C. Corrington, "The Nuclear Rocket Program Its Status and Plans," J. Spacecraft, Vol. 6, p. 465, July 1969.
- 6. D. Buden, "Operational Characteristics of Nuclear Rockets," J. Spacecraft Vol. 7, p. 832, July 1970.
- 7. J. H. Altseimer, et al, "Operating Characteristics and Requirements for the NERVA Flight Engine," J. Spacecraft, Vol. 8, p. 768, July 1971.
- D. R. Koenig, "Experience Gained from The Space Nuclear Rocket Program (Rover)," LA-10062-H, Los Alamos National Laboratory, Los Alamos, New Mexico, May 1986.
- 9. R. R. Holman and B. L. Pierce, "Development of NERVA Reactor for Space Nuclear Propulsion," AIAA 86-1582, June 1986.
- 10. D. Buden, "Nuclear Rocket Safety," 38th Congress of the International Astronautical Federation, Brighton, England, IAF paper 87-297.
- 11. "Evaluation of NTR for Lunar Missions," NASA Contract NAS3-25809, July 9, 1990.







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Phoebus 2A in Transit to Test Cell C



NRX-A3 BEING PREPARED FOR TEST FIRING AT THE NRDS JACKASS FLATS, NEVADA



NRX-A6 TEST FIRING (DEC. 13, 1967): APPROXIMATELY 62 MINS. AT 1124MWt



ORIGINAL PAGE IS OF POOR QUALITY

PROTOTYPE NERVA ENGINE - THE NRX/XE -



ORIGENAL PACE IS OF POOR QUALITY 71



		LEWIS RESEARCH CENTER	
NUCLEAR	SUBSYSTEM CON	MPONENT MATURITY AND READINESS	
	LEVEL OF MATURITY	READINESS	
• FUEL - MATRIX - COMPOSITE - CARBIDE	6 5 4	MATERIALS AND DESIGN READY FOR FLIGHT TESTS REQUIRES SOME R&D REQUIRES SOME R&D	
• FUEL CLUSTER - HARDWARE	5	HOT END SUPPORT REQUIRES ADDITIONAL DESIGN AND ANALYSIS	
AXIAL/LATERAL SUPPORT SYSTEMS	6	MATERIALS AND DESIGN READY FOR FLIGHT TESTS	
• CORE PERIPHERY	6	MATERIALS AND DESIGN READY FOR FLIGHT TESTS	
REFLECTOR	5	ADDITIONAL DESIGN AND ANALYSIS REQUIRED	
	6	MATERIALS AND DESIGN READY FOR FLIGHT TESTS	
• CORE SUPPORT PLATE	6	MATERIALS AND DESIGN READY FOR FLIGHT TESTS	
INTERNAL DOME SHIEI	LD 6	MATERIALS AND DESIGN READY FOR FLIGHT TESTS	
		STEM COMPONENT MATURITY	Figure 12
	CLEAR SUBSY	STEM COMPONENT MATURITY READINESS	Figure 1
HYDROGEN T SHOULD ALL AND REDUCE SSME: NERVA:	URBOPUMPS: AN EX OW SIGNIFICANT RE D DEVELOPMENT TIL 72.6 KG/S @ 704 ~ 40 KG/S @ 130	LEWIS RESEARCH CENTER STEM COMPONENT MATURITY READINESS XTENSIVE DATABASE DEVELOPED SINCE NERVA EDUCTIONS IN WEIGHT, INCREASES IN RELIABILITY ME FOR NTR APPLICATIONS 10 PSI, 350 KG TOTAL MASS 60 PSI, 243 KG TOTAL MASS	Figure 1
NASA NON-NU HYDROGEN T SHOULD ALL AND REDUCE SSME: NERVA: REACTOR PR SST, SPACE TECHNOLOGY HIGH TEMPER	CLEAR SUBSY: AND URBOPUMPS: AN EX OW SIGNIFICANT RE D DEVELOPMENT TH 72.6 KG/S @ 704 ~ 40 KG/S @ 130 ESSURE VESSEL: AI SHUTTLE) HAVE AD 7 TO THE POINT THA TATURE TITANIUM PR	LEWIS RESEARCH CENTER STEM COMPONENT MATURITY READINESS XTENSIVE DATABASE DEVELOPED SINCE NERVA EDUCTIONS IN WEIGHT, INCREASES IN RELIABILITY ME FOR NTR APPLICATIONS 10 PSI, 350 KG TOTAL MASS 60 PSI, 243 KG TOTAL MASS EROSPACE DEVELOPMENT PROGRAMS (BOEING'S VANCED TITANIUM FORMING AND WELDING NT FABRICATION OF A HIGH STRENGTH, LOW MASS, RESSURE VESSEL SHOULD BE POSSIBLE	Figure 1
NON-NUC HYDROGEN T SHOULD ALL AND REDUCE SSME: NERVA: REACTOR PR SST, SPACE TECHNOLOGY HIGH TEMPER NOZZLE DESI THEORETICAL THAT USED C	CLEAR SUBSY AND URBOPUMPS: AN EX OW SIGNIFICANT RE D DEVELOPMENT TH 72.6 KG/S @ 704 ~ 40 KG/S @ 130 ESSURE VESSEL: AN SHUTTLE) HAVE AD Y TO THE POINT THA RATURE TITANIUM PR GN AND COOLING: L EFFICIENCY WITH ON NERVA	LEWIS RESEARCH CENTER STEM COMPONENT MATURITY READINESS XTENSIVE DATABASE DEVELOPED SINCE NERVA EDUCTIONS IN WEIGHT, INCREASES IN RELIABILITY ME FOR NTR APPLICATIONS 10 PSI, 350 KG TOTAL MASS 60 PSI, 243 KG TOTAL MASS EROSPACE DEVELOPMENT PROGRAMS (BOEING'S VANCED TITANIUM FORMING AND WELDING AT FABRICATION OF A HIGH STRENGTH, LOW MASS, RESSURE VESSEL SHOULD BE POSSIBLE TYPICAL NOZZLE DESIGNS NOW CAPABLE OF ~ 98% PERFORMANCE SIGNIFICANTLY GREATER THAN	Figure 1
NON-NU • HYDROGEN T SHOULD ALL AND REDUCE - SSME: - NERVA: • REACTOR PR SST, SPACE TECHNOLOGY HIGH TEMPER • NOZZLE DESI THEORETICAL THAT USED C	CLEAR SUBSY AND URBOPUMPS: AN EX OW SIGNIFICANT RE D DEVELOPMENT TH 72.6 KG/S @ 704 ~ 40 KG/S @ 130 ESSURE VESSEL: AN SHUTTLE) HAVE ADV Y TO THE POINT THA RATURE TITANIUM PH GN AND COOLING: L EFFICIENCY WITH DN NERVA ex ~ 3116°K, Pc ~ IEAT FLUX CAPABILI 200LING)	LEWIS RESEARCH CENTER STEM COMPONENT MATURITY READINESS XTENSIVE DATABASE DEVELOPED SINCE NERVA EDUCTIONS IN WEIGHT, INCREASES IN RELIABILITY ME FOR NTR APPLICATIONS 40 PSI, 350 KG TOTAL MASS 60 PSI, 243 KG TOTAL MASS EROSPACE DEVELOPMENT PROGRAMS (BOEING'S VANCED TITANIUM FORMING AND WELDING AT FABRICATION OF A HIGH STRENGTH, LOW MASS, RESSURE VESSEL SHOULD BE POSSIBLE TYPICAL NOZZLE DESIGNS NOW CAPABLE OF ~ 98% PERFORMANCE SIGNIFICANTLY GREATER THAN 3150 PSI, NOZZLE ASSEMBLY MASS ~ 600 kg, hTY ~ 16.4 KW/CM ² (HYDROGEN REGENERATIVE	Figure 1
NON-NU • HYDROGEN T SHOULD ALL AND REDUCE - SSME: - NERVA: • REACTOR PR SST, SPACE TECHNOLOGY HIGH TEMPER • NOZZLE DESI THEORETICAL THAT USED C SSME: T H C NERVA: T	CLEAR SUBSY AND URBOPUMPS: AN EX OW SIGNIFICANT RE D DEVELOPMENT TH 72.6 KG/S @ 704 ~ 40 KG/S @ 130 ESSURE VESSEL: AN SHUTTLE) HAVE ADA Y TO THE POINT THA RATURE TITANIUM PH GN AND COOLING: L EFFICIENCY WITH DN NERVA ex ~ 3116°K, Pc ~ HEAT FLUX CAPABILI COOLING) ex ~ 2500-3000°K, P ~ 1050 kg, HEAT FL	LEWIS RESEARCH CENTER STEM COMPONENT MATURITY READINESS XTENSIVE DATABASE DEVELOPED SINCE NERVA EDUCTIONS IN WEIGHT, INCREASES IN RELIABILITY ME FOR NTR APPLICATIONS 40 PSI, 350 KG TOTAL MASS 50 PSI, 243 KG TOTAL MASS 50 PSI, 243 KG TOTAL MASS EROSPACE DEVELOPMENT PROGRAMS (BOEING'S VANCED TITANIUM FORMING AND WELDING AT FABRICATION OF A HIGH STRENGTH, LOW MASS, RESSURE VESSEL SHOULD BE POSSIBLE TYPICAL NOZZLE DESIGNS NOW CAPABLE OF ~ 98% PERFORMANCE SIGNIFICANTLY GREATER THAN 3150 PSI, NOZZLE ASSEMBLY MASS ~ 600 kg, htty ~ 16.4 KW/CM ² (HYDROGEN REGENERATIVE PC ~ 450 psi, NOZZLE ASSEMBLY MASS LUX CAPABILITY ~ 4.1 KW/CM ²	Figure 1

CARBIDE "STATE-OF-THE-ART" NERVA DERIVATIVES** 1,020 9,313 500:1 OFFICE 100 3100 3.7 **LEWIS RESEARCH CENTER** 8,816 INITIATIVE 500:1 1000 925 ++THRUST-TO-WEIGHT RATIOS FOR NERVA/NDR SYSTEMS ARE ~5-6 AT THE 250 klbf LEVEL COMPOSITE 3.9 INFORMATION PROVIDED BY LERC PROPULSION TOC WITH SAIC AND WESTINGHOUSE ENGINE WEIGHTS CONTAIN DUAL TURBOPUMP CAPABILITY FOR REDUNDANCY 75 klbf NERVA-TYPE ENGINE CHARACTERISTICS* **TOPPING (EXPANDER)** 2700 8,483 200:1 915 500 4.0 Exploration 2350-2500 850-885 8,000 500:1 **GRAPHITE** 1000 4.3 200:1 7,721 2500 500 875 4.4 SPACE **'72 NERVA**** HOT BLEED/ GRAPHITE TOPPING 2350-2500 845-870 825-850/ 11,250 100:1 450 3.0 + W/O EXTERNAL DISK SHIELD ENGINE THRUST/WEIGHT CHAMBER PRESS. (psia) **ENGINE FLOW CYCLE** SPECIFIC IMPULSE(s) ENGINE WEIGHT+(kg) CHAMBER TEMP. (K) **NOZZLE EXP. RATIO** (W/INT. SHIELD)++ PARAMETERS FUEL FORM 1







76

2010 117					
			2	SENERAL	
PAYLOAD OUTBOU	IND:	73.12 t 34.94 t 7.00 t	MARS EXC MARS TRAI EARTH CRE	URSION MODULE (MI NSFER VEHICLE (MT EW CAPTURE VEHICI	EV) V) LE (ECCV)
PAYLOAD RETURN	:	34.94 t 7.00 t 0.50 t	MTV ECCV (USE MARS RET(D ONLY W/"EXPEND URN SAMPLES	ABLE MODE")
PLANETARY PARK	ING ORBITS:	407 kr 250 kr 500 kr	m CIRCULAR m x 1 SOL*(1 m x 24 hr+(E	(EARTH DEPARTURE MARS ARRIVAL/DEP EARTH ARRIVAL)	E) ARTURE)
g-LOSSES MODELI	ED FOR EARTH D	EPARTUR	EONLY		
EARTH DEPARTUR - 340 m/s - 100 m/s	IE PLANE CHANG (dia > 28.5°) (dia < 28.5°)	E AV PENA	ALTIES:		
MARS APSIDAL AL	IGNMENT AV PEN	NALTIES: 5	560 m/s		
PLANETARY TRAJ		IZED FOR	"ALL PROPUL M 120 TO 434 D	SIVE" MISSION SCEN	ARIO. FOR
SINGLE BURN AND) "3-BURN" PERK	GEE DEPAI	TURES FROM	EARTH EXAMINED	
50 km x 33 852 km =	: 1 SUL URDI I = 4	4.66 HUUr	15		
50 km x 33,852 km = 300 km x 77,604 km :	= 24 HOUR ORBIT	4.66 NOUP	15		
250 km x 33,852 km = 500 km x 77,604 km =	= 1 SOL ORBIT = 2 = 24 HOUR ORBIT	:4.66 HOUF	ns Nge expli	oration initia Lewis Resear	NTIVE OFFICE
250 km x 33,852 km = 500 km x 77,604 km = NASA : <u>PROPULSIO</u>	N SYSTEM/F	24.55 HOUH	LANT/TA	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u>	NTIVE OFFICE CH CENTER == MPTIONS
250 km x 33,852 km = 500 km x 77,604 km = NIR - PRIMARY	N SYSTEM/F	24.55 HOUH 2899 2809El 18 8	NGE EXPLA	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU	NTIVE OFFICE CH CENTER == MPTIONS JLSE
50 km x 33,852 km = 500 km x 77,604 km = PROPULSIOI NTB - PRIMARY - AUXILIARY	N SYSTEM/F PROPELLANT LH2 LH2	2.000 HOUH	AGE EXPLA	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU MID-COUR	NTIVE OFFICE CH CENTER == MPTIONS JLSE ISE CORRECTION
SO km x 33,852 km = 500 km x 77,604 km = PROPULSIOI NIB - PRIMARY - AUXILIARY - AUXILIARY	N SYSTEM/F PROPELLANT LH2 LH2 STOR. BIPROP	2.000 HOUH	AGE EXPLA LLANT/TA SD(S) 150-1020 100 (NERVA 10LE MODE'') 120	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU MID-COUR ATTITUDE	ATIVE OFFICE CH CENTER == MPTIONS JLSE ISE CORRECTION MID-COURSE
250 km x 33,852 km = 500 km x 77,604 km = 500 km x 77,604 km = PROPULSIOI NTR - PRIMARY - AUXILIARY - AUXILIARY ENGINE DESIGN	N SYSTEM/F PROPELLANT LH2 LH2 STOR. BIPROP	PROPEI	LANT/TA Sp(s) 50-1020 500 (NERVA 10LE MODE") 520 ENGINE+ MASS (1)	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU MID-COUR ATTITUDE EXT. SHIELD (I)* <u>MASS (I)</u>	ATIVE OFFICE CH CENTER == MPTIONS JLSE ISE CORRECTION /MID-COURSE TOTAL** MASS (1)
250 km x 33,852 km = 500 km x 77,604 km = 500 km x 77,604 km = PROPULSIOI NIR - PRIMARY - AUXILIARY - AUXILIARY ENGINE DESIGN 0 GRAPHITE NERVA	N SYSTEM/F PROPELLANT LH2 LH2 STOR. BIPROP ISD(S) 850	2ROPEI	AGE EXPLA LLANT/TA Sp(s) S00 (NERVA DDLE MODE'') S20 ENGINE+ MASS (1) 8.00 8.00	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU MID-COUR ATTITUDE EXT. SHIELD (I)* <u>MASS (I)</u> 4.5	ATIVE OPPICE CH CENTER = MPTIONS JLSE ISE CORRECTION MID-COURSE TOTAL** MASS (1) 19.4 20.2
250 km x 33,852 km = 500 km x 77,604 km = 500 km x 77,604 km = PROPULSIOI NIB - PRIMARY - AUXILIARY - AUXILIARY	N SYSTEM/F PROPELLANT LH2 LH2 STOR. BIPROP ISD(S) 850 74 925 1020	2ROPEI	AGE EXPLO AGE EXPLO LLANT/TA SD(S) 500-1020 500 (NERVA 10LE MODE") 320 ENGINE+ MASS (1) 8.00 8.82 9.31 1.26	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU MID-COUR ATTITUDE EXT. SHIELD (I)* <u>MASS (I)</u> 4.5 4.5 4.5	ATIVE OFFICE CH CENTER = MPTIONS JLSE ISE CORRECTION (MID-COURSE TOTAL** MASS (1) 19.4 20.2 20.7 37 65
SO km x 33,852 km = SOO km x 77,604 km = ENCORENT CONTRACTOR PROPULSION NIB PRIMARY - AUXILIARY - AUXILIARY - AUXILIARY ENGINE DESIGN O GRAPHITE NERVA COMPOSITE NERVA COMPOSITE NERVA COMPOSITE PHOEB	N SYSTEM/F 24 HOUR ORBIT 24 HOUR ORBIT PROPELLANT LH2 LH2 STOR. BIPROP ISD(S) 850 74 925 1020 US 925	2ROPEI B PROPEI B S THRUST (kN/klbf) 334/75 334/75 334/75 1112/250	AGE EXPLO AGE EXPLO LLANT/TA SD(S) 150-1020 100 (NERVA 10LE MODE'') 120 ENGINE+ MASS (1) 8.00 8.82 9.31 21.76 ENGINE2: 00/07	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPL MID-COUR ATTITUDE EXT. SHIELD (I)* <u>MASS (I)</u> 4.5 4.5 4.5 9.0	CH CENTER = MPTIONS ULSE ISE CORRECTION MID-COURSE TOTAL** MASS (1) 19.4 20.2 20.7 37.65
SO km x 33,852 km = 500 km x 77,604 km = 500 km x 77,604 km = PROPULSIOI NIB - PRIMARY - AUXILIARY - AUXILIARY - AUXILIARY ENGINE DESIGN 0 GRAPHITE NERVA 0 COMPOSITE NERVA 0 COMPOSITE NERVA 0 COMPOSITE PHOEB RESERVE/COOLDO	N SYSTEM/F PROPELLANT LH2 LH2 STOR. BIPROP ISD(S) 4 925 1020 US 925 OWN PROPELLAN	2ROPEI	LANT/TA LLANT/TA LLANT/TA SD(S) 150-1020 100 (NERVA 10LE MODE") 120 ENGINE+ MASS (1) 8.00 8.82 9.31 21.76 FRATES: 2%/3	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPL MID-COUR ATTITUDE EXT. SHIELD (I)* <u>4.5</u> 4.5 4.5 4.5 9.0 3%/.0.65 kg/m ² /mth	CH CENTER = MPTIONS ULSE ISE CORRECTION MID-COURSE TOTAL** MASS (1) 19.4 20.2 20.7 37.65
SO km x 33,852 km = SOO km x 77,604 km = ENTR PROPULSION NTR PRIMARY AUXILIARY AUXILIARY ENGINE DESIGN CARBIDE NERVA COMPOSITE NERVA COMPOSITE PHOEB RESERVE/COOLDO PROPELLANT TAN	N SYSTEM/F PROPELLANT LH2 LH2 STOR. BIPROP ISD(S) 7A 925 1020 US 925 DWN PROPELLAN IKS JETTISONED	2ROPEI 2ROPEI 2ROPEI 4.55 2ROPEI 4.5 2 2 2 2 2 2 2 2 2 2 2 2 2	LANT/TA LLANT/TA LLANT/TA Sp(s) So-1020 So0 (NERVA 'IDLE MODE'') S20 ENGINE+ MASS (1) 8.00 8.82 9.31 21.76 F RATES: 2%/3 II AND MOC BU	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPU MID-COUR ATTITUDE EXT. SHIELD (I)* <u>MASS (I)</u> 4.5 4.5 4.5 9.0 3%/.0.65 kg/m ² /mth JRNS	ATIVE OFFICE CH CENTER = MPTIONS JLSE ISE CORRECTION /MID-COURSE TOTAL** MASS (1) 19.4 20.2 20.7 37.65
250 km x 33,852 km = 500 km x 77,604 km = PROPULSIOI NIB - PRIMARY - AUXILIARY -	N SYSTEM/F 24 HOUR ORBIT 24 HOUR ORBIT 25 HOUR ORBIT 25 HOUR ORBIT 26 HOUR ORBIT 26 HOUR ORBIT 27 HOUR ORBIT 26 HOUR ORBIT 27 HOUR ORBIT 27 HOUR ORBIT 26 HOUR ORBIT 27 HOUR ORBIT 27 HOUR ORBIT 28 HOUR O	2ROPEI 2ROPEI 2ROPEI 4 5 5 5 5 5 5 5 5 5 5 5 5 5	LLANT/TA LLANT/TA Sp(S) 150-1020 100 (NERVA 10LE MODE'') 120 ENGINE+ MASS (1) 8.00 8.82 9.31 21.76 F RATES: 2%/3 II AND MOC BU AL PROPELLA C(- 15%), COM	DRATION INITIA LEWIS RESEAR <u>NKAGE ASSU</u> <u>USAGE</u> MAIN IMPL MID-COUR ATTITUDE EXT. SHIELD (I)* <u>4.5</u> 4.5 4.5 9.0 3%/.0.65 kg/m²/mth JRNS NT REQUIRED PER M	ATIVE OFFICE CH CENTER = MPTIONS JLSE ISE CORRECTION /MID-COURSE TOTAL** MASS (1) 19.4 20.2 20.7 37.65



2016 NTR MARS VEHICLE SIZE COMPARISON (OPTIMIZED VS. NON-OPTIMIZED TRAJECTORIES-COMPOSITE FUEL/Isp=925s)



			WIS NEGLA		
NERVA-DERIVAT	IVE ENGINI	E*/ISP TR	ADE RES	ULTS	
ALL PROPULSIVE OPT	IMIZED 2016	6 MARS	MISSION -	434 DAYS)+	
IML	<u>.EO (1)/TOTAL BU</u>	<u>irn time (Hr</u>	<u>5)</u>		
SINGLE CORE STAGE VEHIC	LE				
W/"CUSTOMIZED" DROP TAN 75 kibt Engine	IKS **		EHICLE REUSE	E MODE"	
W/"3 PERIGEE BURN"		(AL	L PROPULSIVE	MISSION	
EARTH DEPARTORE					
1. GRAPHITE CORE NDR (2350 K/lsp = 850 s)			725/3.38		
2. COMPOSITE CORE NDR			613/2 .9 9		
(2100 Kisp = 323 3)					
3. CARBIDE CORE NDR			518/2.64		
(3100 K/Isp = 1020 s)					
REFERENCE MTV (90 DAY STUDY	 D: CHEM/AB IMLE	EO=752t FOR	ECCV RETURN	/=830t FOR	
PROPULSIVE EARTH CAPTURE					
CHAMBER PRESSURE = 1000 psi DROP TANKS ASSUMED TO BE C	ia, ε = 500:1) Cylindrical W/R	ROOT2 ELLIPS	OIDAL DOMES	; DIA.=10M, LENGTI	н
CONSTRAINED TO BE ≤35 M					
CONSTRAINED TO BE ≤35 M		e exolor	A 11 11 CO 10 10 10 10 10 10 10 10 10 10 10 10 10	NATINE OFFICE	J
CONSTRAINED TO BE ≤35 M	SPACE	e explor	ation init	iative office	リ Fi
CONSTRAINED TO BE ≤35 M	SPACE	e explor	ation init	iative office RCH CENTER -	リ _{Fi}
CONSTRAINED TO BE ≤35 M	SPACE	e explor	ation init WIS RESEA	iative office RCH CENTER =	J _{Fi}
CONSTRAINED TO BE ≤35 M	SPACE	e explor	ation init Wis Resea	IATIVE OFFICE RCH CENTER =	J _{Fi}
CONSTRAINED TO BE ≤35 M NOSA	ATION FOR	e explor le <u>''All PR</u> ON - 434	ation init wis resea <u>opulsive</u> days	IATIVE OFFICE RCH CENTER = <u>" OPTIMIZED</u>	J _{Fi}
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR 2016 N	ATION FOR	e explor Le <u>"All PR</u> ON - 434	ATION INIT WIS RESEA OPULSIVE DAYS	iative office RCH CENTER = <u>" OPTIMIZED</u>	J _{Fi}
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR 2016 N	ATION FOR	e explor Le <u>"All PR</u> <u>ON - 434</u>	ATION INIT WIS RESEA OPULSIVE DAYS	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u>	J
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR 2016 N DURATION (mins)	ATION FOR MARS MISSIO	E EXPLOR LE <u>"ALL PR</u> ON - 434 75 kibf COMPOSITE	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 kibf</u> <u>COMPOSITE</u>	Fi
CONSTRAINED TO BE ≤35 M NDIVIDUAL BURN DURA 2016 N DURATION (mins)	ATION FOR MARS MISSIO	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 kibf</u> COMPOSITE	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 kibf</u> <u>COMPOSITE</u> 28 2/1	J _{Fi}
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR 2016 N DURATION (mins) TMI (TOTAL /# PEBIGEE BURNS)	ATION FOR ARS MISSI GRAPHITE ~122.1/3	E EXPLOR LE <u>"ALL PR</u> <u>75 kibf</u> COMPOSITE ~104/3	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 klbf</u> <u>COMPOSITE</u> 38.2/1	Fi
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS)	ATION FOR ARS MISSI GRAPHITE	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 kibf</u> COMPOSITE ~104/3	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 kibf COMPOSITE</u> 38.2/1	J _{Fi}
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC	ATION FOR MARS MISSIO GRAPHITE ~122.1/3 40.0	E EXPLOR 	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8	IATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 klbf</u> <u>COMPOSITE</u> 38.2/1 13.4	J _{Fi}
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC	ATION FOR MARS MISSI GRAPHITE ~122.1/3 40.0	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 kibf</u> COMPOSITE ~104/3 36.8	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 kibf</u> <u>COMPOSITE</u> 38.2/1 13.4	J _{Fi}
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC TEI	SPACE <u>ATION FOR</u> <u>ARS MISSIC</u> <u>GRAPHITE</u> -122.1/3 40.0 30.0	E EXPLOR 	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1	IATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 klbf</u> <u>COMPOSITE</u> 38.2/1 13.4 11.0	Fi
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC TEI	ATION FOR MARS MISSI GRAPHITE ~122.1/3 40.0 30.0	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 kibf</u> COMPOSITE ~104/3 36.8 28.0	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1	ATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 kibf</u> <u>COMPOSITE</u> 38.2/1 13.4 11.0	Fi
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC TEI EOC	SPACE <u>ATION FOR</u> <u>ARS MISSIC</u> <u>GRAPHITE</u> -122.1/3 40.0 30.0 7.1	E EXPLOR 	АТІЮН ІНІТ WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1 6.7	DATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 klbf</u> <u>COMPOSITE</u> 38.2/1 13.4 11.0 2.7	Fi
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) (TOTAL/# PERIGEE BURNS) MOC TEI EOC	SPACE ATION FOR ARS MISSIG GRAPHITE ~122.1/3 40.0 30.0 7.1	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 kibf</u> COMPOSITE ~104/3 36.8 28.0 6.9	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1 6.7	DATIVE OFFICE RCH CENTER = <u>"OPTIMIZED</u> <u>250 klbf COMPOSITE</u> 38.2/1 13.4 11.0 2.7	Fi
CONSTRAINED TO BE ≤35 M CONSTRAINED TO BE ≤35 M INDIVIDUAL BURN DUR. 2016 N 2016 N DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC TEI EOC	SPACE <u>ATION FOR</u> <u>ARS MISSIC</u> <u>GRAPHITE</u> -122.1/3 40.0 30.0 7.1 IOUSLY FOR 63	E EXPLOR 	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1 6.7 AT 1125 MW	DATIVE OFFICE RCH CENTER =	Fi
CONSTRAINED TO BE ≤35 M NOIVIDUAL BURN DUR. 2016 N DURATION (mins) (TOTAL/# PERIGEE BURNS) MOC TEI EOC NOTE: NRX-A6 RAN CONTINU HYDROGEN FU	SPACE ATION FOR ARS MISSIG GRAPHITE ~122.1/3 40.0 30.0 7.1 JOUSLY FOR 62 UEL EXIT TEME	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 kibf</u> COMPOSITE ~104/3 36.8 28.0 6.9 2 MINUTES PERATURE	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1 6.7 AT 1125 MWI ≥ 2550 K (DE	DATIVE OFFICE RCH CENTER = <u>250 klbf</u>	Fi
CONSTRAINED TO BE ≤35 M CONSTRAINED TO BE ≤35 M INDIVIDUAL BURN DUR. 2016 M DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC TEI EOC NOTE: NRX-A6 RAN CONTINU HYDROGEN FI NRX-XE ACCUMULATE ODEBATION D	ATION FOR ARS MISSI GRAPHITE -122.1/3 40.0 30.0 7.1 JOUSLY FOR 62 UEL EXIT TEMI D APPROXIMA DUBING 28 ENG	E EXPLOR LE <u>"ALL PR</u> ON - 434 <u>75 klbf</u> COMPOSITE ~104/3 36.8 28.0 6.9 2 MINUTES PERATURE TELY 115 M SINF RESTA	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1 6.7 AT 1125 MWI ≥ 2550 K (DE INUTES OF F BT TESTS OF	DATIVE OFFICE RCH CENTER	J
CONSTRAINED TO BE ≤35 M = NASA NDIVIDUAL BURN DUR. 2016 M DURATION (mins) TMI (TOTAL/# PERIGEE BURNS) MOC TEI EOC NOTE: NRX-A6 RAN CONTINU HYDROGEN FI NRX-XE ACCUMULATE OPERATION D BETWEEN MA	ATION FOR ARS MISSI GRAPHITE ~122.1/3 40.0 30.0 7.1 JOUSLY FOR 62 UEL EXIT TEMI D APPROXIMA DURING 28 ENG ARCH AND AUG	E EXPLOR LE "ALL PR ON - 434 <u>75 kibf</u> COMPOSITE ~104/3 36.8 28.0 6.9 2 MINUTES PERATURE TELY 115 M SINE RESTA SUST 1969	ATION INIT WIS RESEA OPULSIVE DAYS CARBIDE ~87.8/3 33.8 26.1 6.7 AT 1125 MWI ≥ 2550 K (DE INUTES OF F RT TESTS OG	RCH CENTER = <u>"OPTIMIZED</u> <u>250 klbf</u> <u>COMPOSITE</u> 38.2/1 13.4 11.0 2.7 a, 55 klbf AND A CEMBER 1967) CURRING	Fi





2016 NTR MARS VEHICLE SIZE COMPARISON (OPTIMIZED TRAJECTORIES - COMPOSITE FUEL/75 klbf & 250 klbf)





LUNAR OUTPOST FLIGHT SCHEDULE CHEM/AERO REFERENCE







ASE DESIGN IS 75,000	LBF THRUS	T NERVA-DERIVATIV	E ENGINE WITH
- (U,Zr)C-COMPOSIT - 2700 K CHAMBER - ISP = 900 SECON	TE FUEL ELE TEMP; 500 P DS HOUR LIEFTI	MENTS (NUCLEAR ISI CHAMBER PRESS	FURNACE TESTED) SURE ISSIONS INCL. DISPOSAL)
	MASS (KG)	SOURCE	COMMENTS
REACTOR	5,662	WESTINGHOUSE	NERVA-derivative
NTERNAL SHIELD	1,527	WESTINGHOUSE	•
NOZZLE	867	MMAG [*]	200:1 expansion 7.4 m length
NON-NUCLEAR HARDWARE	1,194	MMAG'	Incl. pumps, valves, lines, thrust structure, etc., 2% contingency
Subtotal: Engine	9,250		F/W = 3.69
EXTERNAL SHIELD	4,545	NERVA DESIGN	To be resized based on final design
Total NTR System	13,795		

Figure 27

as Applies



Figure 28











TEMPORAL VARIATION OF DOSE RATE FOR MSFC-BOEING "NON-OPTIMIZED" REFERENCE 2016 NTR MISSION

	1575 MW_	
	Engine	Mission
	Operating	Elapsed
	Time	Time
Maneuver	(minutes)	(days)
Trans Mars Injection	123.5	0
Mars Orbital Capture	62.3	156
Trans Earth Injection	24.1	187.
Earth Orbital Capture	10.7	435

Fvent	* Dose Rate
Full Power Operation	7.2×10^5
Trans Mars Injection Plus 1 Day	1.1×10^2
Prior to Mars Orbital Capture	2.3×10^{-1}
Prior to Trans Earth Injection	1.9 x 10 ⁰
Prior to Earth Orbital Capture	7.5×10^{-2}

*Dose point on axial midplane 100 feet from core centerline

REF. B. SCHNITZLER (INEL)

	OBSERVATIONS/CONCLUDING REMARKS	 ROVER/NERVA PROGRAMS ESTABLISHED A SIGNIFICANT DATA BASE ON WHICH THE '72 REFERENCE NERVA ENGINE WAS BASED 	 EXPERIENCE ALSO OBTAINED IN OPERATING "FULL-SCALE" ENGINE FACILITIES (ALBEIT IN "OPEN CYCLE" MODE), HANDLING LARGE QUANTITIES OF LH₂ AND RADIOACTIVE SYSTEMS (E-MAD FACILITY), SAFETY, AND IN THE BEGINNINGS OF "EFFLUENT CLEAN-UP" (WITH THE NUCLEAR FURNACE) 	 CONTINUED DEVELOPMENT OF CHEMICAL PROPULSION SYSTEMS HAVE ADVANCED SUBSTANTIALLY THE STATE-OF-THE-ART OF NON-NUCLEAR ENGINE COMPONENT (E.G., NOZZLES, TURBOPUMPS, ETC.) 	 NTR PROPULSION IS <u>ENABLING</u> FOR MARS MISSIONS AND CAN BE <u>ENHANCING</u> FOR LUNAR MISSIONS PROVIDING BOTH IMLEO BENEFITS AND OPERATIONAL EXPERIENCE IN A RELATIVELY "NEARBY" SPACE ENVIRONMENT 	• AN NTR PROGRAM WILL REQUIRE A LOT OF WORK - FACILITY REQUIREMENTS KEY FOLLOWED BY HIGH TEMPERATURE FUEL DEVELOPMENT	 PERFORMANCE PARAMETERS ACHIEVED IN ROVER/NERVA PROGRAM ARE WITHIN A FACTOR OF TWO OR LESS OF THOSE CURRENTLY BEING EXAMINED FOR SEI'S LUNAR AND MARS MISSIONS 		SPACE EXPLORATION INITIATIVE OFFICE
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