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Prepared for the
32nd Aerospace Sciences Meeting and Exhibit
sponsored by the American Institute of Aeronautics and Astronautics
Reno, Nevada, January 10-13, 1994



National Aeronautics and
Space Administration

(NASA-TM-106766) OBSERVATIONS OF
METHANE AND ETHYLENE DIFFUSION
FLAMES STABILIZED AROUND A BLOWING
POROUS SPHERE UNDER MICROGRAVITY
CONDITIONS (NASA. Lewis Research
Center) 11 p

N95-18211

Unclass

G3/29 0031424

OBSERVATIONS OF METHANE AND ETHYLENE DIFFUSION FLAMES STABILIZED AROUND A BLOWING POROUS SPHERE UNDER MICROGRAVITY CONDITIONS

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Abstract

This paper presents the experimental and theoretical results for expanding methane and ethylene diffusion flames in microgravity. A small porous sphere made from a low-density and low-heat-capacity insulating material was used to uniformly supply fuel at a constant rate to the expanding diffusion flame. A theoretical model which includes soot and gas radiation is formulated but only the problem pertaining to the transient expansion of the flame is solved by assuming constant pressure infinitely fast one-step ideal gas reaction and unity Lewis number. This is a first step toward quantifying the effect of soot and gas radiation on these flames. The theoretically calculated expansion rate is in good agreement with the experimental results. Both experimental and theoretical results show that as the flame radius increases, the flame expansion process becomes diffusion controlled and the flame radius grows as \sqrt{t} . Theoretical calculations also show that for a constant fuel mass injection rate a quasi-steady state is developed in the region surrounded by the flame and the mass flow rate at any location inside this region equals the mass injection rate.

I. Introduction

The absence of buoyancy-induced flows in a microgravity environment and the resulting increase in the reactant residence time significantly alters the fundamentals of many combustion processes. Substantial differences between normal gravity and microgravity flames have been reported during droplet combustion[1], flame spread over solids[2,3], candle flames[4] and others. These differences are more basic than just in the visible flame shape. Longer residence time and higher concentration of combustion products create a

thermochemical environment which changes the flame chemistry. Processes such as soot formation and oxidation and ensuing flame radiation, which are often ignored under normal gravity, become very important and sometimes controlling. As an example, consider the droplet burning problem. The visible flame shape is spherical under microgravity versus a teardrop shape under normal gravity. Since most models of droplet combustion utilize spherical symmetry, excellent agreement with experiments is anticipated. However, microgravity experiments show that a soot shell is formed between the flame and the evaporating droplet of a sooty fuel[1]. This soot shell alters the heat and mass transfer between the droplet and its flame resulting in significant changes in the burning rate and the propensity for flame extinction. This change in the nature of the process seems to have occurred because of two reasons: (i) The soot formed could not be swept out of the flame due to the absence of buoyant flows. Instead, it was forced to go through the high temperature reaction zone increasing the radiative heat losses, and (ii) soot formation was enhanced due to an increase in the reactant residence time.

Recently, some very interesting observations of candle flames under various atmospheres in microgravity have been reported[4]. It was found that for the same atmosphere, the burning rate per unit wick surface area and the flame temperature were considerably reduced in microgravity as compared with normal gravity. Also, the flame (spherical in microgravity) was much thicker and further removed from the wick. It thus appears that the flame becomes "weaker" in microgravity due to the absence of buoyancy generated flow which serves to transport the oxidizer to the combustion zone and remove the hot combustion products from it. The buoyant flow, which may be characterized by the strain rate, assists the diffusion process to execute these essential functions for

the survival of the flame. Thus, the diffusion flame is "weak" at very low strain rates and as the strain rate increases the flame is initially "strengthened" and eventually it may be "blown-out." The computed flammability boundaries[5] show that such a reversal in material flammability occurs at strain rates around 5 sec^{-1} . Model calculations for a zero strain rate *I-D* diffusion flame show that even gas radiation is sufficient to extinguish the flame[6].

The above observations suggest that flame radiation will substantially influence diffusion flames under microgravity conditions, particularly the conditions at extinction. This is because, flame radiation at very low or zero strain rates is enhanced due to: (i) high concentration of combustion products in the flame zone which increases the gas radiation, and (ii) low strain rates provide sufficient residence time for substantial amounts of soot to form which is usually responsible for most of the radiative heat loss. It is anticipated that this radiative heat loss may extinguish the already "weak" diffusion flame.

To investigate the possibility of radiation-induced extinction limits under microgravity conditions, spherical geometry is chosen. This is convenient for both experiments and theoretical modeling. In this work, a porous spherical burner is used to produce spherical diffusion flames in μg . Experiments conducted with this burner on methane (less sooty) and ethylene (sooty) diffusion flames are described in the next section. A general theoretical model for transient radiative diffusion flames is then formulated and calculations are presented for the transient expansion of the spherical diffusion flame. These calculations are compared with the experimental measurements in the discussion section. This work is the first necessary step toward investigating radiative-extinction of spherical diffusion flames.

II. Experimental Apparatus and Results

The μg experiments were conducted in the 2.2 sec drop tower at the NASA Lewis Research Center. The experimental drop-rig used is schematically shown in Color Plate 1. It consists of a test chamber, burner, igniter, gas cylinder, solenoid valve, camera, computer and batteries to power the computer and the solenoid valves. The spherical burner (1.9 cm in diameter) is constructed from a low density and low heat capacity porous ceramic material. A 150 cc gas cylinder at approximately 46.5 psig is used to supply the fuel to the porous spherical burner. Typical gas flow rates used were in the range of $3\text{-}15 \text{ cm}^3/\text{s}$. Flow rates to the burner are controlled by a needle valve and a gas solenoid valve is used to open and close the gas line to the burner upon

computer command. An igniter is used to establish a diffusion flame. After ignition the igniter is quickly retracted from the burner and secured in a catching mechanism by a computer-controlled rotary solenoid. This was necessary for two reasons (i) The igniter provides a heat sink and will quench the flame (ii) Upon impact with the ground (after 2.2 sec) the vibrating igniter may damage the porous burner.

As shown in the Color Plate 1, the test chamber has a 5" diameter Lexan window which enables the camera to photograph the spherical diffusion flame. The flame growth can be recorded either by a 16mm color movie camera or by a color CCD camera which is connected to a video recorder by a fiber-optic cable during the drop. Since the fuel flow may change with time, it had to be calibrated for various settings of the needle valve for both methane and ethylene. A soap bubble flow meter was used to calibrate the flow for various constant gas cylinder pressures. Constant pressures were obtained by connecting the cylinder to the main 200 lb gas cylinder using a quick-disconnect. An in-line pressure transducer was used to obtain the transient flow rates. Changes in the cylinder pressure during the experiment along with the pressure-flow rate calibration, provides the transient volumetric flow rates. These are shown plotted in Figure 1.

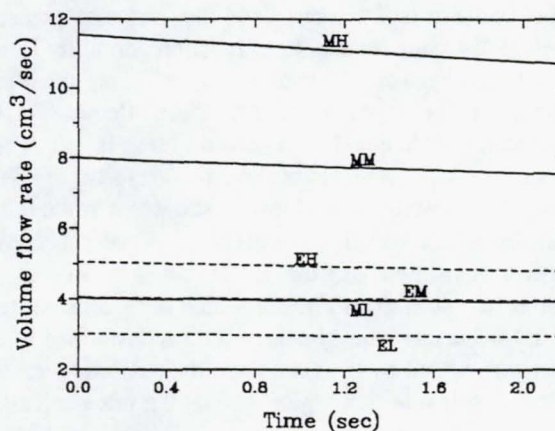
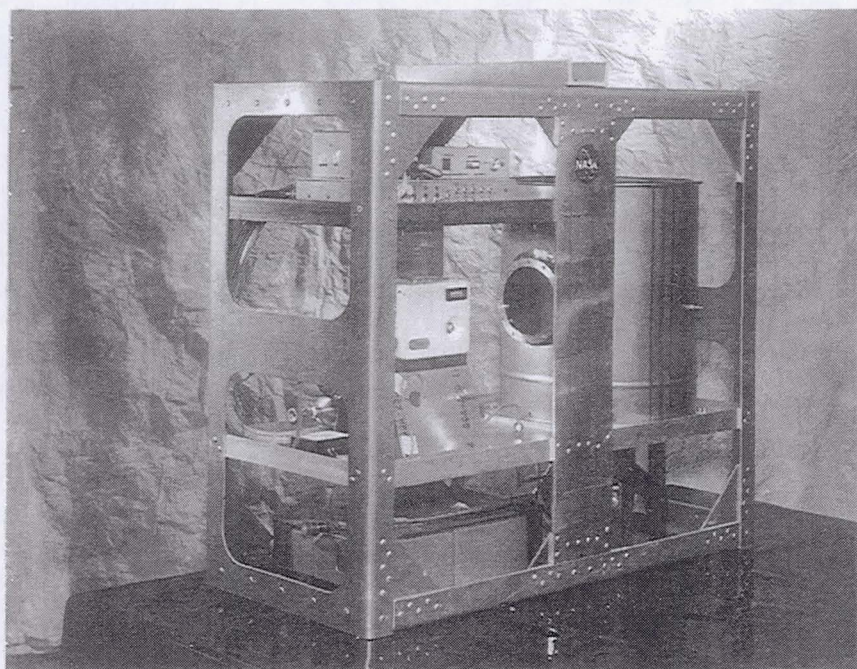


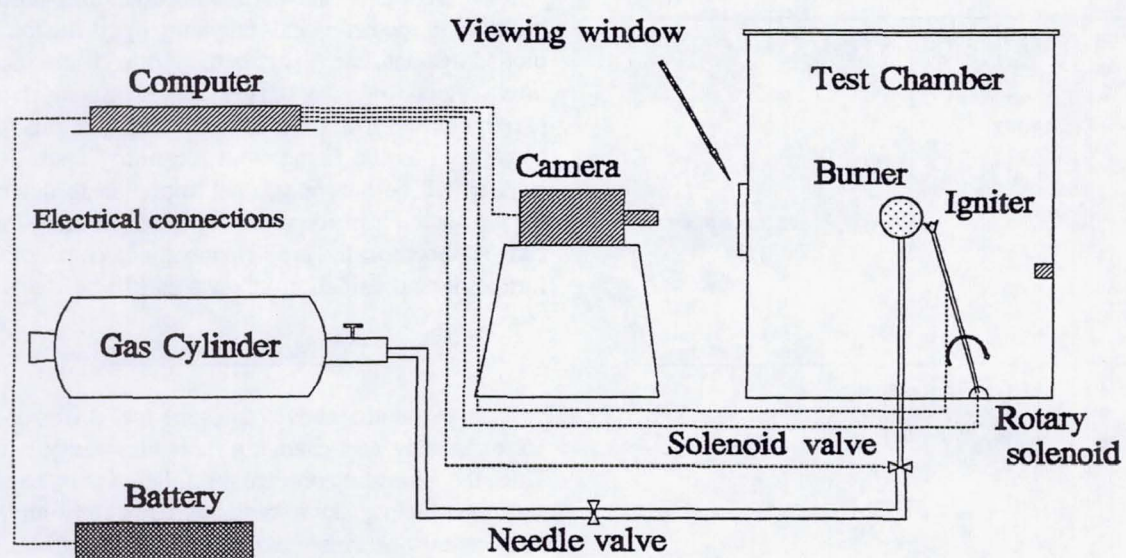
Figure 1: Volume flow rates versus time.

In Figure 1, the letters "M" and "E" represent methane and ethylene respectively and the letters "L", "M" and "H" represent low, medium and high flow rates. Thus, MM implies medium flow rate of methane. Note that low flow rate for methane is nearly equal to the medium flow rate of ethylene. For these experiments, the gas velocity at the burner wall was between 0.25-1 cm/sec.

The porous spherical burner produced a nearly spherical diffusion flame in microgravity. Some observed



Picture of the microgravity spherical diffusion flame apparatus



Schematic of the microgravity spherical diffusion flame apparatus

disturbances are attributed to slow large-scale air motion inside the test chamber. Several microgravity experiments were performed under ambient pressure and oxygen concentration conditions for different flow rates of methane and ethylene (as shown in Fig. 1). Methane was chosen to represent a non-sooty fuel and ethylene was chosen to represent a moderately sooty fuel. In these experiments, ignition was always initiated in 1-g just prior to the drop. The package was typically dropped within one second after ignition. The primary reason for not igniting in μg was the loss of time in heating the igniter wire and in stabilizing the flame after the initial ignition disturbances. Some photographs from these experiments are shown in the Color Plate 2.

The flame radius measured from such photographs along with the model predictions (to be discussed later) are shown in Fig. 2. As expected, for the same flow rates it was found that ethylene flames were much sootier and smaller. Immediately after dropping the package, the flame shape changed from a teardrop shape (see Color Plate 2) to a spherical shape (although it was not always completely spherical, probably because of slow large-scale air motion persisting inside the test chamber). The photographs shown in the Color Plate 2 are for medium flow rates of methane and low flow rates of ethylene. For the data presented in Fig. 2, an average flame radius determined from the photographs was used.

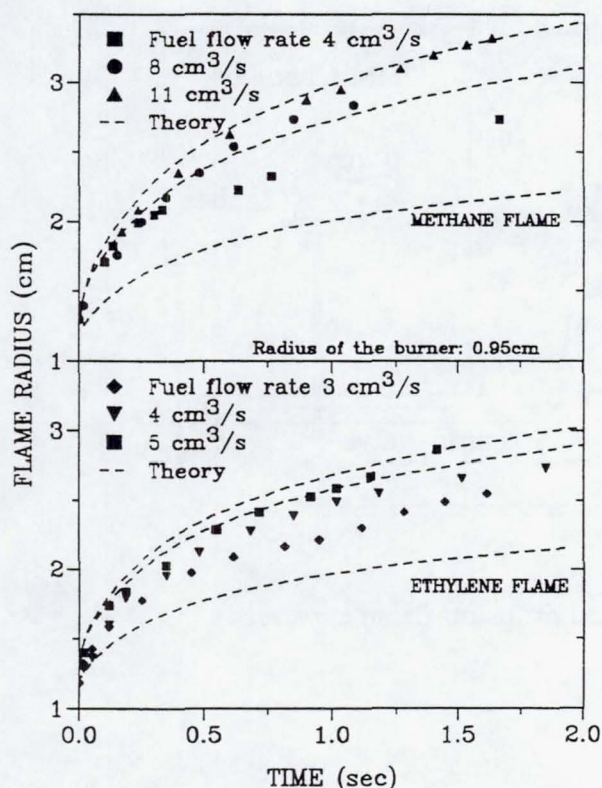


Figure 2: Flame radius versus time

It is interesting to note that for both methane and ethylene (see the progressive flame growth in Color Plate 2), initially and in 1-g (e.g. photographs 'e' & 'f') the flame is nearly blue (non-sooty) but becomes bright yellow (sooty) immediately after the onset of μg conditions. Later, as the μg time progresses, the flame grows in size and becomes orange and less luminous and the soot seems to disappear. A possible explanation for this observed behavior is suggested by the theoretical calculations of Ref. 6. The soot volume fraction first quickly increases and later decreases as the local concentration of combustion products increases. Essentially, further soot formation is inhibited by the increase in the local concentration of the combustion products [Ref.7,8] and soot oxidation is enhanced. Thus, at the onset of μg conditions, initially a lot of soot is formed in the vicinity of the flame front (the outer faint blue envelope) resulting in bright yellow emission. As the flame grows, several events reduce the flame luminosity: (i) The soot is pushed toward cooler regions by thermophoresis. In fact, for sootier fuels this leads to the formation of a soot shell. (ii) The high concentration of combustion products left behind by the flame front inhibits soot formation and promotes soot oxidation. (iii) The dilution and radiative heat losses caused by the increase in the concentration of combustion products reduces the flame temperature which in turn reduces the soot formation rate and the flame luminosity.

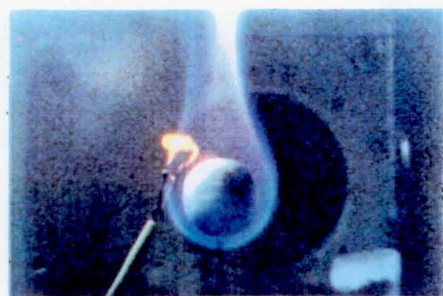
Figure 2 shows the average measured flame radius for methane and ethylene μg diffusion flames plotted against time. This is the radius of the outer faint blue region of the flame as measured from the photographs. To a good approximation this may be considered as the flame front location. Thus, as a first step, it will be interesting and important to determine if the transient expansion of the μg spherical diffusion flame can be theoretically predicted without considering soot formation and oxidation kinetics and flame radiation.

III. Model Formulation

As noted above, the spherical diffusion flames are expanding and changing their luminosity with time. Thus, the general theoretical formulation must be transient and must include flame radiation. For the simplest case of *constant pressure ideal gas reactions with $Le=1$* , we may write the following governing equations for any geometrical configuration (spherical or counterflow geometry):

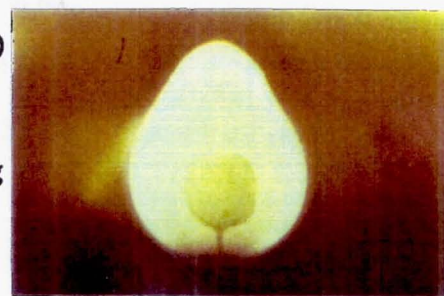
Mass Conservation:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{v}) = 0 \quad (1)$$



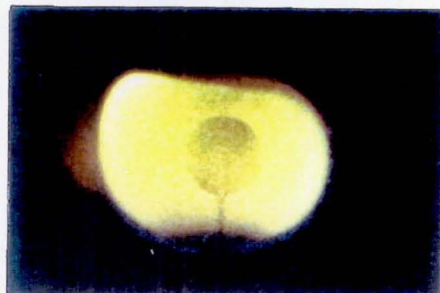
(a)

1-g flame



(b)

0.0667sec
into μ g



(c)

0.3333sec
into μ g



(d)

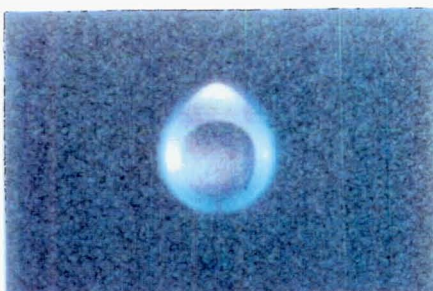
0.6667sec
into μ g

METHANE FLAME, FUEL FLOW RATE $8 \text{ cm}^3/\text{sec}$



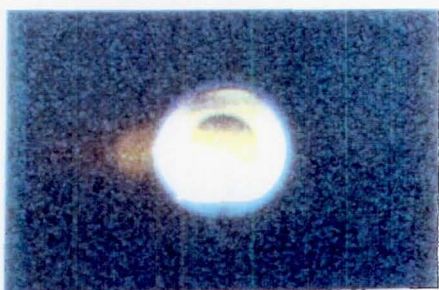
(e)

1-g flame



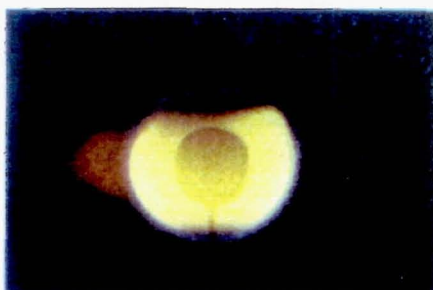
(f)

0.0667sec
into μ g



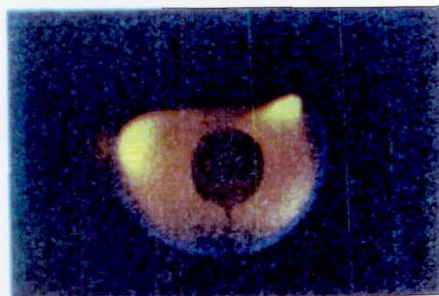
(g)

0.1667sec
into μ g



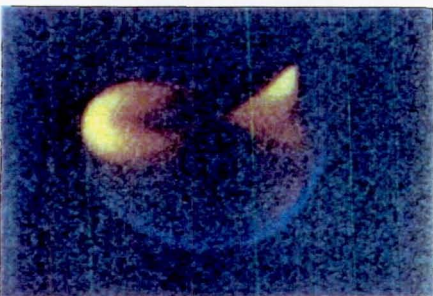
(h)

0.5667sec
into μ g



(i)

1.233sec
into μ g



(j)

1.667sec
into μ g

ETHYLENE FLAME, FUEL FLOW RATE $3 \text{ cm}^3/\text{sec}$

COLOR PLATE 2

Energy Conservation:

$$\rho \frac{\partial h^s}{\partial t} + \rho \vec{v} \cdot \nabla h^s - \nabla \cdot (\rho D \nabla h^s) = - \sum_i h_i^o W_i - Q_s \dot{m}_{s_s}'' - \nabla \cdot \vec{q}_r \quad (2)$$

Constant Pressure Ideal Gas:

$$\rho T = \rho_\infty T_\infty \quad \text{or} \quad \rho h^s = \text{const.} \quad (3)$$

Here, the symbols have their usual definitions with ρ = density, T = temperature, \vec{v} = velocity, Y_i = mass fraction of species i , h^s = sensible enthalpy, w_i = mass production or destruction rate per unit volume of species i and D = diffusion coefficient. The last three terms in Equ (2) respectively are: the chemical heat release rate due to gas phase combustion, chemical heat released due to soot oxidation and the radiative heat loss rate per unit volume. The above equations, however, are insufficient for our problem because the soot volume fraction must be known as a function of space and time to determine the radiative heat loss. To enable describing soot volume fraction in a simple manner, we define the mass fraction of atomic constituents as follows:

$$\xi_j = \sum_i \left(M_j v_i^j / M_i \right) Y_i, \quad \text{where } M_i \text{ is the molecular weight}$$

of species i , M_j is the atomic weight of atom j and v_i^j is the number of atoms of kind j in specie i . Assuming that the only atomic constituents present in the hydrocarbon flame are C, H, O & Inert and with $Y_{\text{soot}} \equiv \Phi \equiv \rho_s f_v / \rho$ (where: ρ_s = soot density & f_v = soot volume fraction), we obtain: $\xi_C + \xi_H + \xi_O + \xi_I + \rho_s f_v / \rho = 1$. Defining $\xi_C + \xi_H = \xi_F$ and $Z_F = \xi_F / Y_{F\infty}$ and $Z_o = \xi_o / Y_{o\infty}$, we obtain $Z = [(\xi_F / Y_{F\infty}) Z_F + \rho_s f_v / \rho]$ as the conserved scalar for a sooty flame. This yields the following soot, fuel and oxidizer conservation equations in terms of their scalar variables:

Soot Conservation:

$$\rho \frac{\partial \Phi}{\partial t} + \rho \vec{v} \cdot \nabla (\Phi) - \nabla \cdot (\rho D_s \nabla (\Phi)) = \dot{m}_{s_s}''' - \dot{m}_{s_s}'' = \dot{m}_{\text{net}}''' \quad (4)$$

Fuel Conservation:

$$\rho \frac{\partial Z_F}{\partial t} + \rho \vec{v} \cdot \nabla (Z_F) - \nabla \cdot (\rho D \nabla (Z_F)) = - \frac{1}{Y_{F\infty}} (\dot{m}_{s_s}''' - \dot{m}_{s_s}'') = - \frac{\dot{m}_{\text{net}}'''}{Y_{F\infty}} \quad (5)$$

Oxygen Conservation:

$$\rho \frac{\partial Z_o}{\partial t} + \rho \vec{v} \cdot \nabla (Z_o) - \nabla \cdot (\rho D \nabla (Z_o)) = 0 \quad (6)$$

Under conditions of small soot loading, the soot terms in the fuel and energy conservation equations can be ignored except when studying radiative extinction. Thus, Equ (5) may be considered homogeneous to a good approximation. Also, as a first crude approximation, the heat lost by flame radiation may be subtracted from the heat of combustion in the form of a radiative fraction. Thus, the energy equation (Equ(2)) can also be made homogeneous if written in terms of the total enthalpy [$h = \sum_i Y_i (h_i^o + h_i^s)$]. This approach may be adequate for calculating the observed expansion rate of the spherical diffusion flames, but it is completely inadequate for predicting radiative extinction. However, the great mathematical advantage of this approach is that it makes Eqs. (2, 4, 5, & 6) identical and only one conserved scalar equation need be considered. As a first step, it is of interest to see how well the transient expansion of the μ g spherical diffusion flames be predicted without rigorously considering soot and gas radiation. This will also help in quantifying the effect of soot and gas radiation by comparison with more detailed calculations. Re-writing the above equations in spherical coordinates, we get:

Mass Conservation:

$$\frac{\partial \rho}{\partial t} + \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \rho v) = 0 \quad (7)$$

Fuel Conservation:

$$\rho \frac{\partial Z}{\partial t} + \rho v \frac{\partial Z}{\partial r} - \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \rho D \frac{\partial Z}{\partial r} \right) = 0 \quad (8)$$

These two equations along with the ideal gas law at constant pressure, Equ.(3), are sufficient to describe the transient growth of non-radiative spherical diffusion flames and are expected to approximate this growth in the presence of flame radiation. It is also assumed that a fast one-step overall reaction occurs at the flame surface. This

is represented by: $v_F F + v_O O \rightarrow \sum_{i=1}^{N-2} v_i P_i$; with

q^o as the standard heat of reaction and $Q = q^o / M_F v_F$ the heat released per unit mass of fuel. Clearly,

$$q^o = h_F^o M_F v_F + h_O^o M_O v_O - \sum_{i=1}^{N-2} h_i^o M_i v_i. \quad \text{The}$$

corresponding initial and boundary conditions for a sphere of radius 'R' blowing fuel gases at a rate $\dot{M}(t)$ are discussed below and illustrated in Figure 3.

the domain $r > r_f$, we only need to find expressions for ρ and v in terms of Z .

For $r_f(t) < r < \infty$:

$$\rho = \rho_o \left(1 + \frac{z Q Y_{F\infty}}{h_{\infty}^s} \right)^{-1} \quad (14)$$

$$v = \frac{r_f^2(1-Z_c)}{r^2 Z_c} \left[\frac{v_f}{(1-Z_c)} - \frac{\dot{M}(t)}{4\pi r_f^2 \rho_o} - \frac{\dot{M}(t) Q Z_c}{4\pi r_f^2 \rho_o h_{\infty}^s} \left(Y_{FR} - Y_{F\infty} \left(\frac{Z_R - Z_c}{1 - Z_c} \right) \right) \right] - \frac{D}{\rho} \frac{\partial \rho}{\partial r} \quad (15)$$

In the derivation of Equ. (15), gas velocity and density at the flame front are made continuous i.e. $v(r_f^+, t) = v(r_f^-, t) = v_f$ and $\rho(r_f^+, t) = \rho(r_f^-, t) = \rho_f$. Thus, v_f in Equ. (15) can be obtained from Equ. (11). Once again, the third term inside the bracket of Equ (15) becomes zero for reasons discussed above. Eqs. (14), (15) and (8) along with the boundary conditions are sufficient to determine $Z(r,t)$ for $r > r_f$.

IV Solution

Before discussing the solution procedure, let us examine the porous sphere used in the experiments. This sphere is quite small (19 mm dia.) and is constructed from a high porosity, low density and low heat capacity insulating material. Thus, its capacity to store heat and mass is negligible compared to the fuel injection rate which is injected inside the sphere (see Fig. 3). Hence, conditions inside the material of the sphere equilibrate on a time scale much shorter than the flame expansion time i.e. convection balances diffusion for any variable under consideration that can be described by an equation similar to Equ. (8). Neglecting radiation from the surface of the sphere, conservation conditions yield equations identical to Eqs. 9(a) and 9(b) where ambient conditions are assumed to exist near the center of the sphere. Physically, the only purpose the porous sphere serves is to provide a radially uniform flow and it does not participate in energy and species balances because of its low storage capacity. Thus, the boundary conditions at the source can be applied at an arbitrarily small radius 'R' (chosen for numerical convenience) such that $Y_{F\infty} \approx Y_{FR}$; $h_{\infty}^s \approx h_R^s$ and $Z_R \approx 1$. This considerably simplifies Eqs. (11) & (15).

Equation (8), with Eqs. (10) & (11) for the fuel side and Eqs. (14) & (15) for the air side were numerically solved using the method of lines. A computer package entitled DSS2 was employed for this purpose. The calculated results for the flame location are shown plotted by dotted lines in Fig. 2. Property values used were those for air ($\rho_o = 1.16 \times 10^{-3}$ gm/cm³, $D_o = 0.226$ cm²/s, $T_{\infty} = 298$ K, $C_p = 1.35$ J/kgK) and the diffusion coefficient was assumed to vary as $T^{3/2}$ as predicted by kinetic theory of gases. Heat of combustion (Q) and mass based stoichiometric coefficient (v) used for methane and ethylene were $Q = 47465$ J/gm and $v = 4$ and $Q = 47465$ J/gm and $v = 3.429$ respectively. No assumptions other than those stated above were made to match the experimental data. Initial spatial distribution of $Z(r,0)$ required for the flame at the start of μg time (i.e. at $t=0$) was taken as:

$$Z(r,0) = \operatorname{erfc} \left(\frac{(r-R) \operatorname{erfc}^{-1}(Z_c)}{(r_f-R)} \right) \quad (16)$$

V Results and Discussion

Figure 2 shows the average radius of the outer faint blue regions for both methane and ethylene μg diffusion flames plotted against time. This radius was measured from the photographs. As stated above, the corresponding calculated results for the flame location are shown plotted with dotted lines. Given the approximations made in the model and the experimental errors, the comparison between the experimental and predicted flame radius is quite encouraging.

Numerical calculations also yield the instantaneous velocity and density profiles around the porous sphere during the flame expansion. These are shown plotted in Figures 4 & 5. Starting from the porous

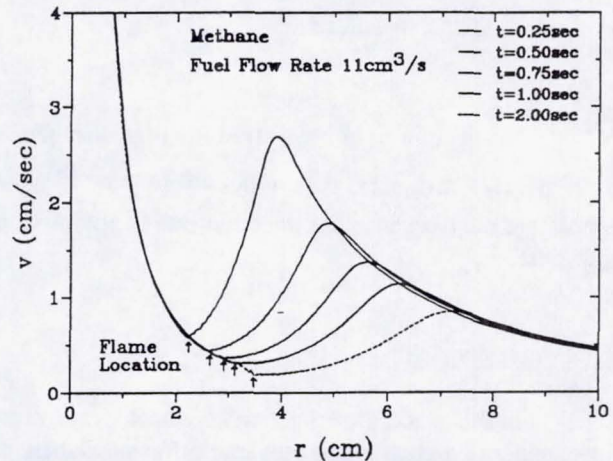


Figure 4: Radial velocity distribution at various instants

sphere ($r=0.95$), the gas velocity drops sharply and becomes a minimum at the flame location ($r=r_f$). Surprisingly, the mass flow rate at any location $r < r_f$ is found to be equal to the mass injection rate (i.e. $4\pi\rho v r^2 = \dot{M}(t)$). This implies that a similarity exists in the normalized coordinate $r/r_f(t)$ in the region $r < r_f(t)$. The density profiles in this region (Fig. 5) also show a similarity. Further reflection shows that this is to be

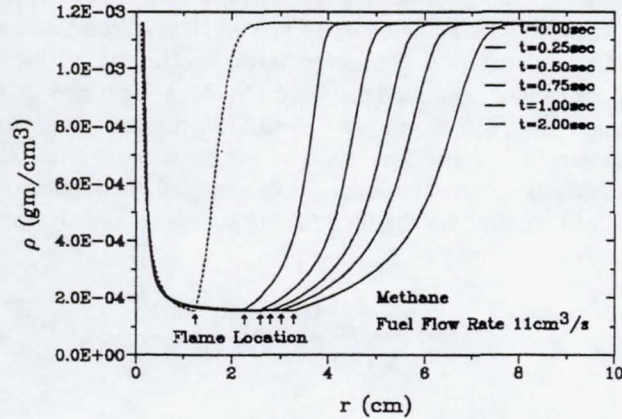


Figure 5: Radial density distribution at various instants

expected. In this problem, a constant temperature (adiabatic flame temperature) spherical flame is propagating outward starting from a small radius. In the spherical geometry, heat loss from the region surrounded by the flame is not possible. Thus, the only heat required by this region from the flame (in the absence of radiation) is to heat the injected mass $\dot{M}(t)$ to the flame temperature. Since, the injected mass is taken to be constant with time, a quasi steady state is developed. This is also observed in the density gradients at the flame on the fuel side (which are constant and are proportional to the temperature gradients). Applying a simple energy balance over the region $r \leq r_f$ we obtain:

$$\frac{\dot{M}(t)}{4\pi r_f^2 \rho_f} = \frac{\dot{M}(t)}{4\pi r_f^2 \rho_o} - \frac{D}{\rho_f} \left(\frac{\partial \rho}{\partial r} \right)_{r_f} = v_f \quad (17)$$

Using Equ. (11) we find that at the flame $v_f = \dot{M}(t)/4\pi\rho_f r_f^2$. It is important to note that this is possible only because the injection rate is not varying with time.

VI Conclusions

In this work, experimental and theoretical results for expanding methane and ethylene diffusion flames in microgravity are presented. A small porous sphere made from a low-density and low-heat-capacity insulating material was used to uniformly supply fuel at a constant

rate to the expanding diffusion flame. A theoretical model which includes soot and gas radiation is formulated but only the problem pertaining to the transient expansion of the flame is solved by assuming constant pressure infinitely fast one-step ideal gas reaction and unity Lewis number. This is a first step toward quantifying the effect of soot and gas radiation on these flames. The theoretically calculated expansion rate is in good agreement with the experimental results. Both experimental and theoretical results show that as the flame radius increases, the flame expansion process becomes diffusion controlled and the flame radius grows as \sqrt{t} . Theoretical calculations also show that for a constant fuel mass injection rate a quasi-steady state is developed in the region surrounded by the flame and the mass flow rate at any location inside this region equals the mass injection rate.

Acknowledgements: We would like to thank Mr. Mark Guether for his initial work on the Drop Rig. This project is supported by NASA under contract no. NAG3-1460.

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REPORT DOCUMENTATION PAGE			Form Approved OMB No. 0704-0188	
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1. AGENCY USE ONLY (Leave blank)	2. REPORT DATE November 1994	3. REPORT TYPE AND DATES COVERED Technical Memorandum		
4. TITLE AND SUBTITLE Observations of Methane and Ethylene Diffusion Flames Stabilized Around a Blowing Porous Sphere Under Microgravity Conditions		5. FUNDING NUMBERS WU-656-61-23		
6. AUTHOR(S) Arvind Atreya, Sanjay Agrawal, Kurt Sacksteder, and Howard R. Baum				
7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS(ES) National Aeronautics and Space Administration Lewis Research Center Cleveland, Ohio 44135-3191		8. PERFORMING ORGANIZATION REPORT NUMBER E-9209		
9. SPONSORING/MONITORING AGENCY NAME(S) AND ADDRESS(ES) National Aeronautics and Space Administration Washington, D.C. 20546-0001		10. SPONSORING/MONITORING AGENCY REPORT NUMBER NASA TM-106766 AIAA-94-0572		
11. SUPPLEMENTARY NOTES Prepared for the 32nd Aerospace Sciences Meeting and Exhibit sponsored by the American Institute of Aeronautics and Astronautics, Reno, Nevada, January 10-13, 1994. Arvind Atreya and Sanjay Agrawal, University of Michigan, Department of Mechanical Engineering and Applied Mechanics, Ann Arbor, Michigan 48109-2125; Kurt Sacksteder, NASA Lewis Research Center; Howard R. Baum, National Institute of Standards and Technology, Gaithersburg, Maryland 20899. Responsible person, Kurt Sacksteder, organization code 6711, (216) 433-2857.				
12a. DISTRIBUTION/AVAILABILITY STATEMENT Unclassified - Unlimited Subject Category 29		12b. DISTRIBUTION CODE		
13. ABSTRACT (Maximum 200 words) This paper presents the experimental and theoretical results for expanding methane and ethylene diffusion flames in microgravity. A small porous sphere made from a low-density and low-heat-capacity insulating material was used to uniformly supply fuel at a constant rate to the expanding diffusion flame. A theoretical model which includes soot and gas radiation is formulated but only the problem pertaining to the transient expansion of the flame is solved by assuming constant pressure infinitely fast one-step ideal gas reaction and unity Lewis number. This is a first step toward quantifying the effect of soot and gas radiation on these flames. The theoretically calculated expansion rate is in good agreement with the experimental results. Both experimental and theoretical results show that as the flame radius increases, the flame expansion process becomes diffusion controlled and the flame radius grows as \sqrt{t} . Theoretical calculations also show that for a constant fuel mass injection rate a quasi-steady state is developed in the region surrounded by the flame and the mass flow rate at any location inside this region equals the mass injection rate.				
14. SUBJECT TERMS Microgravity combustion		15. NUMBER OF PAGES 12		
		16. PRICE CODE A03		
17. SECURITY CLASSIFICATION OF REPORT Unclassified	18. SECURITY CLASSIFICATION OF THIS PAGE Unclassified	19. SECURITY CLASSIFICATION OF ABSTRACT Unclassified	20. LIMITATION OF ABSTRACT	