

SiC-based photo-detectors for UV, VUV, EUV and soft X-ray detection

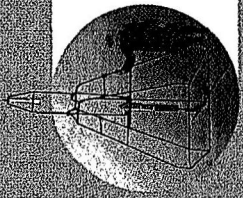
Feng Yan

NASA-Goddard Space Flight Center / MEI

In collaboration with

SiCLab at Rutgers Univ. & United Silicon Carbide, Inc.

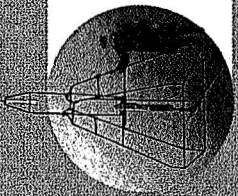
ECSCRM2006, Sept. 7th, 2006, Newcastle upon Tyne, UK



Acknowledgement

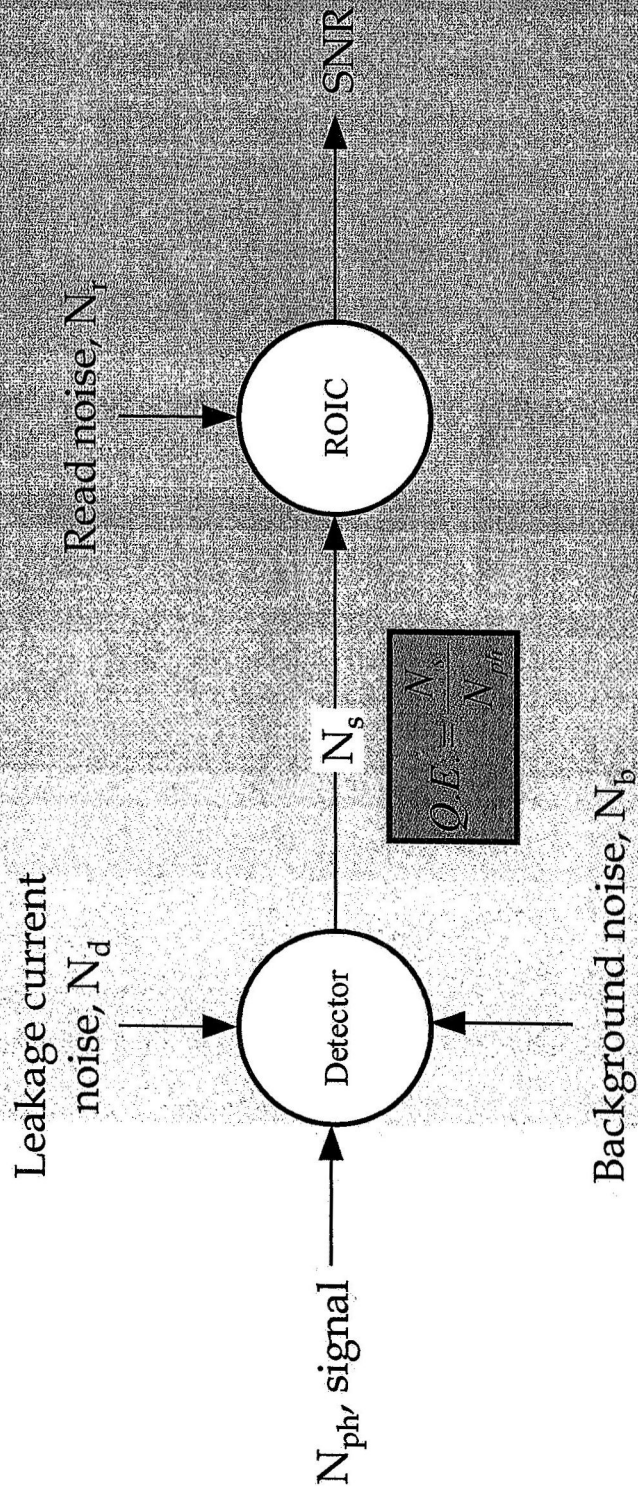
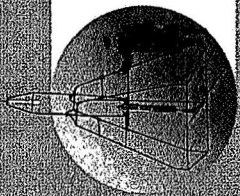
- **Carl. M. Stahle, Fred Herrero, Shahid Aslam, Sachi Babu at NASA Goddard Space Flight Center**
- **Xiaobin Xin, Jun Hu, Jian H. Zhao at Rutgers SiCLab**
- **Maurice Weiner at United Silicon Carbide, Inc.**



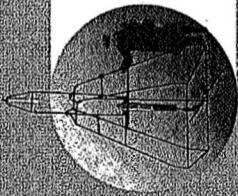


Outline

- An idea detector
 - SNR discussion
 - Features for an ideal detector
- SiC detectors
 - Dark current
 - Read noise
 - Quantum efficiency
- Applications
- Conclusion



N_s , N_d , N_r & N_b : the number of electrons generated by signal, leakage current, read noise and background during the observing period



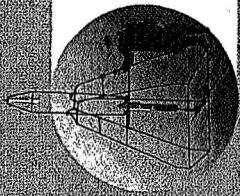
An ideal photo-detector

$$SNR = \frac{Q.E. * N_{ph}}{\sqrt{Q.E. * N_{ph} + N_b + N_d + N_r^2}}$$

An ideal detector

- $Q.E. = 1$ for low energy photons or E_{ph}/E_{pair} for high energy photons
- N_d negligible (detector related)
- N_b negligible (instrument related)
- N_r negligible (ROIC related)





An ideal photo-detector

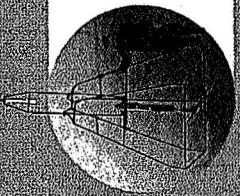
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An ideal detector

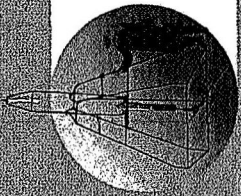
- $Q.E. = 1$ for low energy photons or E_{ph}/E_{pair} for high energy photons (detector related)
- N_d negligible (detector related)
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Ultimate limit: $QE * N_{ph} = 1 \rightarrow SNR = 1$

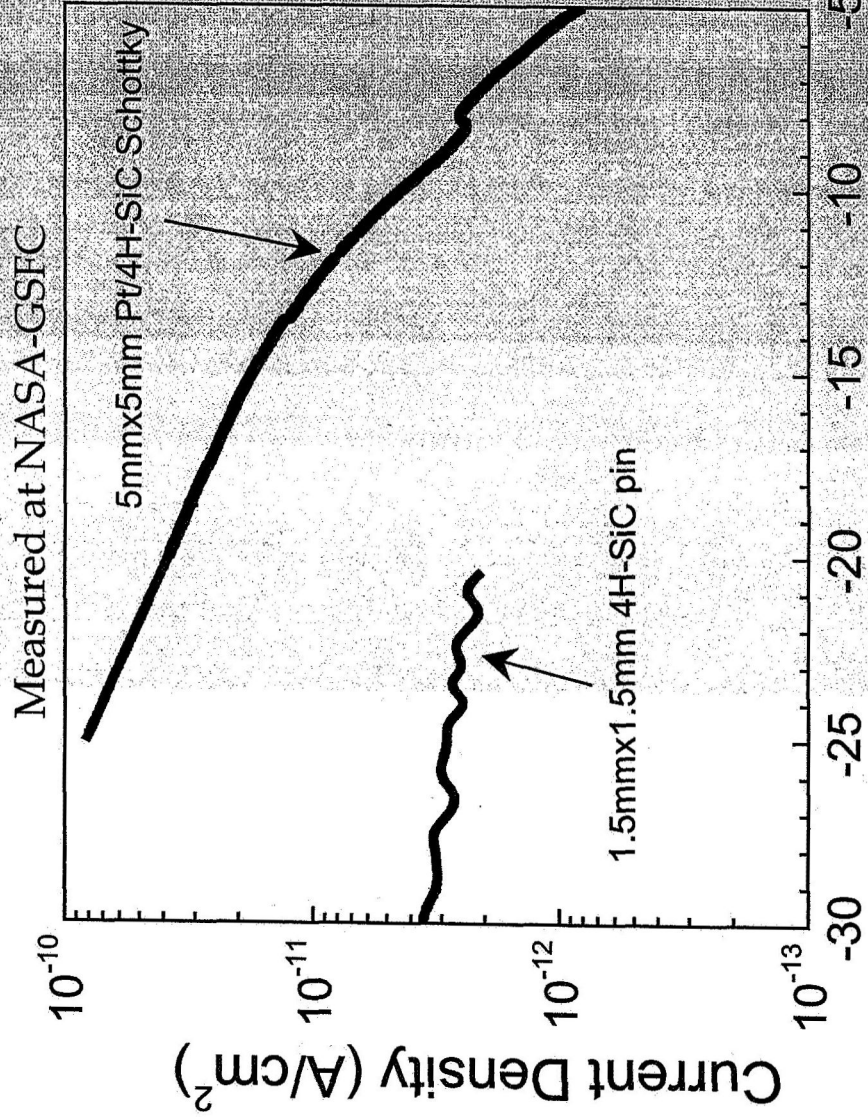




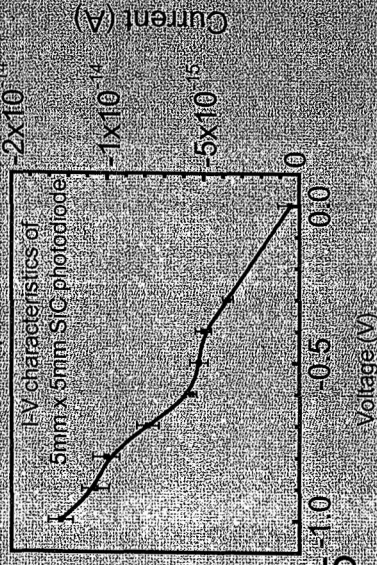
Dark current



Dark current density of SiC photodiodes



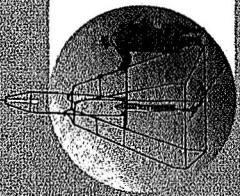
Measured at Keithley, Inc



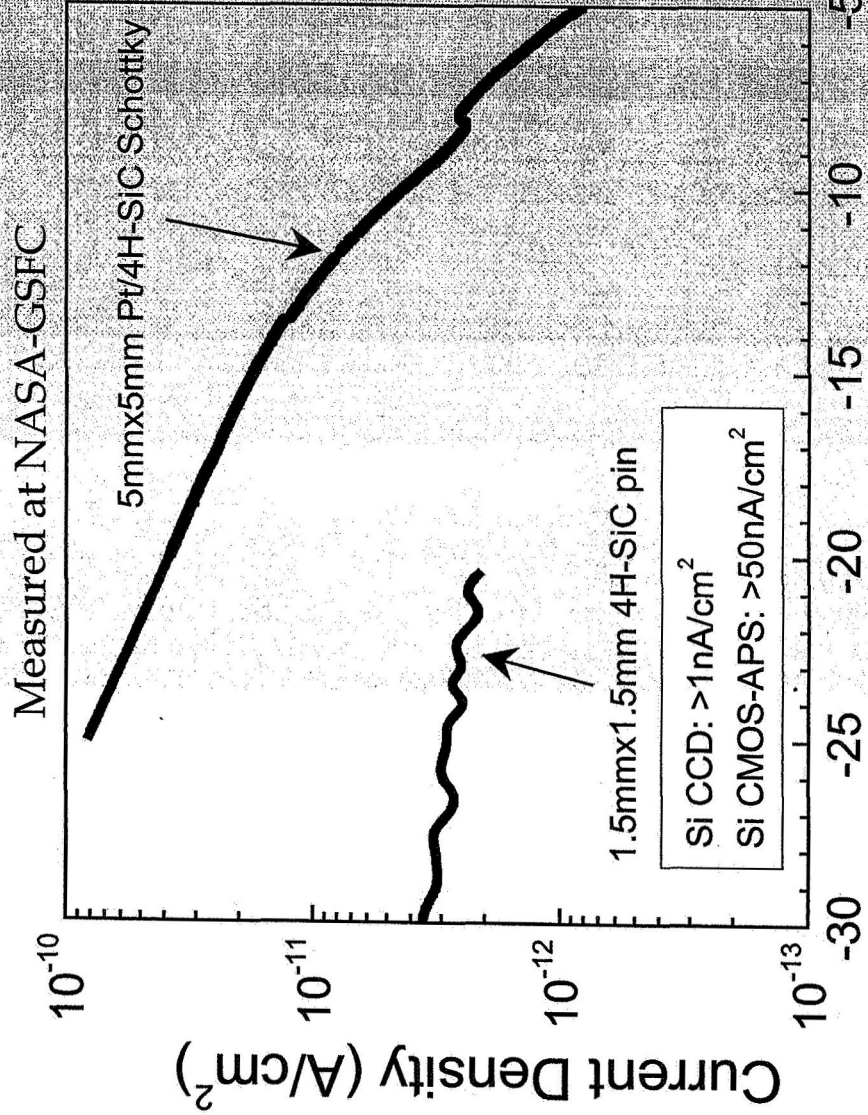
Reverse Bias (V)



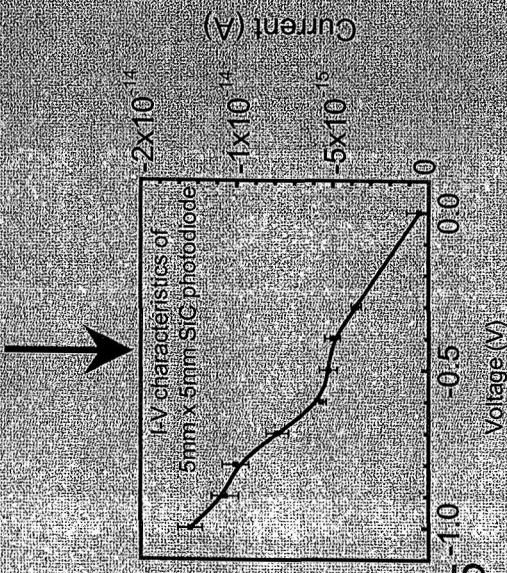
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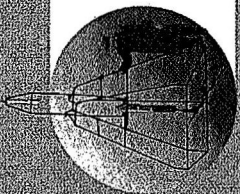
Dark current density of SiC photodiodes at room temperature



Measured at Keithley, Inc.



Reverse Bias (V)



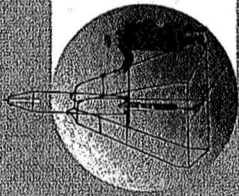
Dark current in SiC detectors

	I_d^{SiC}/I_d^{Si} (present experimental)	I_d^{SiC}/I_d^{Si} (theoretical)
Schottky	10^{-6} (surface defects)	$10^{-17} \dagger$
PIN	$\sim 10^{-7-8}$ (bulk defects)	$10^{-18} \ddagger$

$$\dagger I_d = A \cdot A^{**} \cdot T \cdot \exp\left(-\frac{q\Phi_b}{kT}\right) \cdot \exp\left(\frac{q\sqrt{qE/4\pi\epsilon_s}}{kT}\right)$$

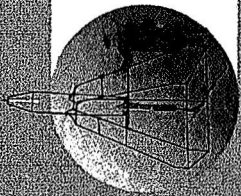
$$\ddagger I_d \sim q \sqrt{\frac{D_p}{\tau_p} \cdot \frac{n_i^2}{N_D} + \frac{qn_i W}{\tau_e}}$$



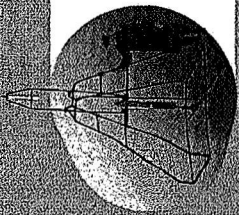


- $N_{Si} > 10^{10} / \text{cm}^2 \cdot \text{sec}$
- $N_{SiC} \sim 10^4 / \text{cm}^2 \cdot \text{sec}$ (can be lower)
- $dN_{Si}/dT \sim 0.2/K$
- N_{SiC} at RT $< N_{Si}$ at 77K

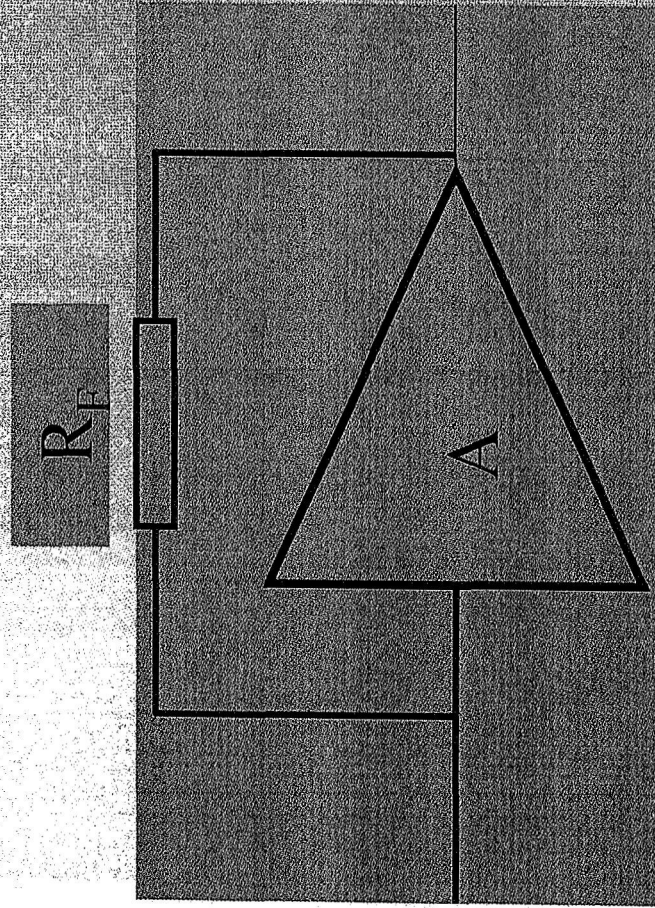




Read noise



Resistive feedback trans-impedance amplifier



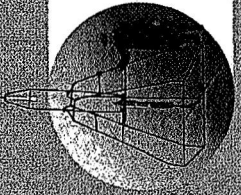
Noise

$$(4kTB/R_F)^{1/2}$$

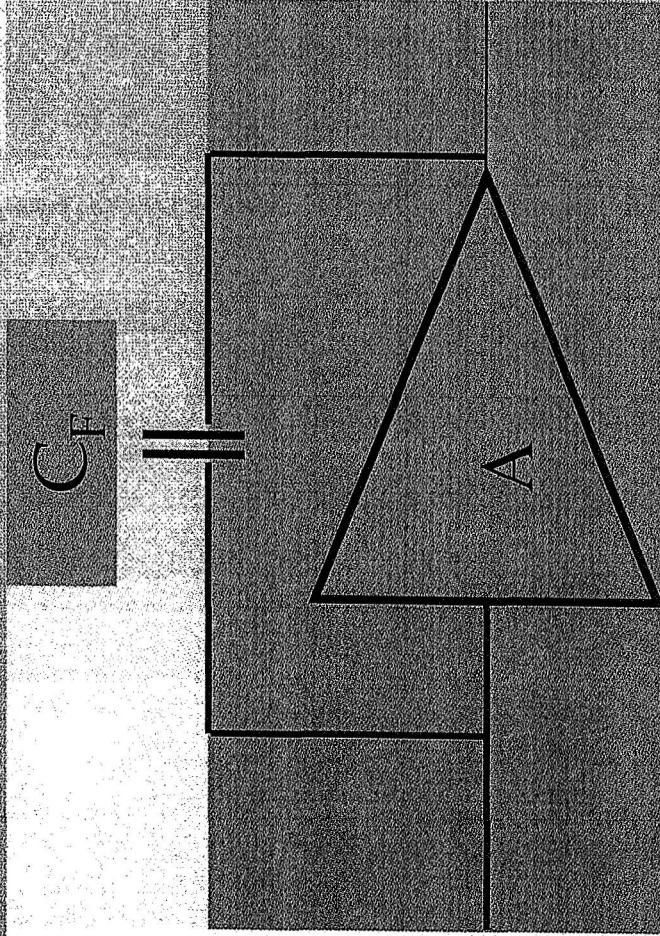
Bandwidth

$$R_F C$$

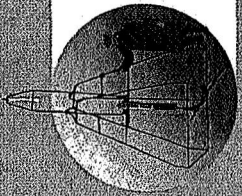




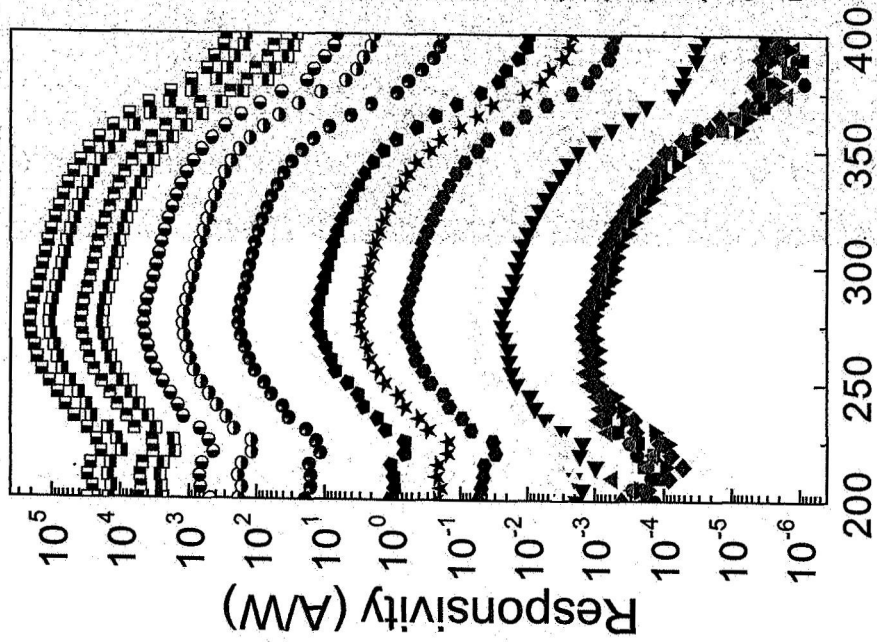
Capacitive feedback trans-impedance amplifier



$$\text{Noise} \sim (kTC_F B)^{1/2}$$



Avalanche gain

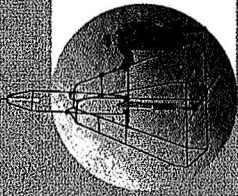


- Very large gain possible
- The functional form of the responsivity does not change over several orders of magnitude
- Linear gain $>10^6$
- Linear gain of SiC = 10^6
- Linear gain of Si = 10^3
- Linear gain of InGaAs/InP $<10^2$
- Best APD
- $k=0.1$

Wavelength (nm)



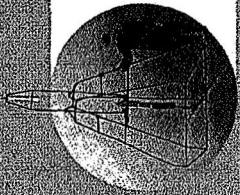
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Excess noise

$$SNR = \frac{Q.E. * N_{ph} * M}{\sqrt{(Q.E. * N_{ph} + N_b + N_d) * M^2 * (kM) + N_f^2}}$$

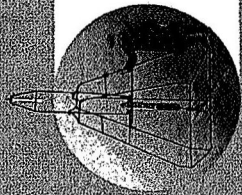




Excess noise

$$SNR = \frac{Q.E. * N_{ph} * M}{\sqrt{(Q.E. * N_{ph} + N_b + N_d) * M^2 * (kM) + N_x^2}}$$

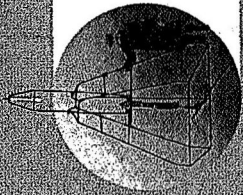




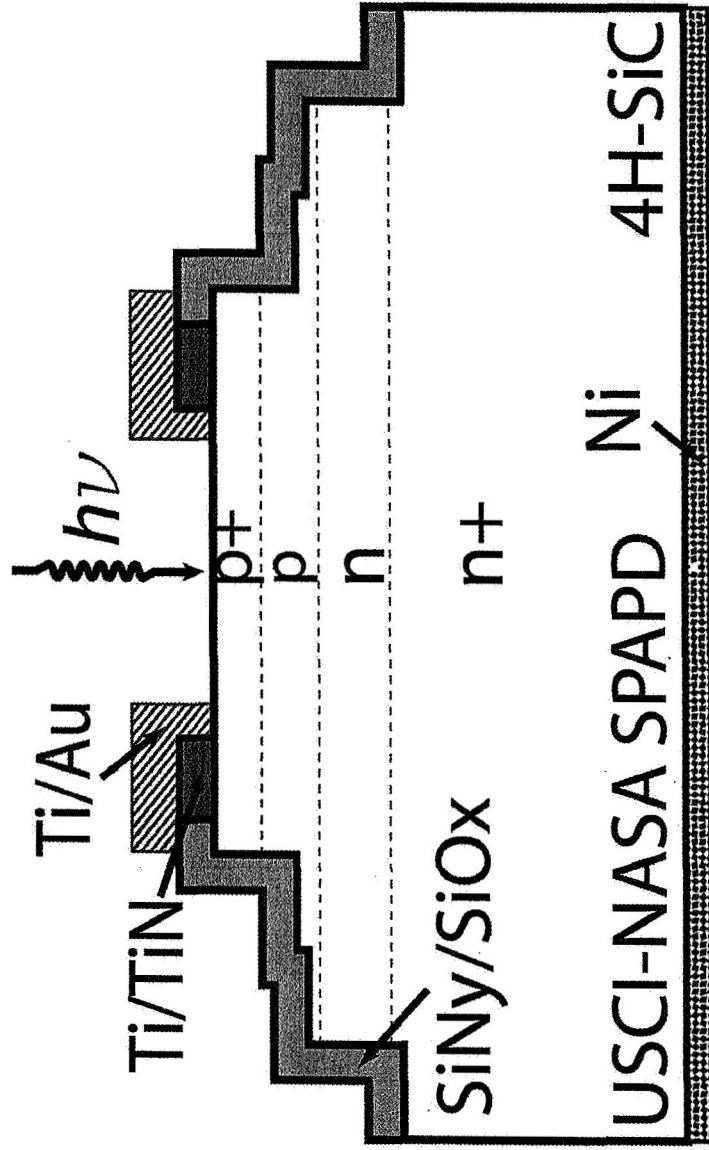
SNR in single photon counting mode

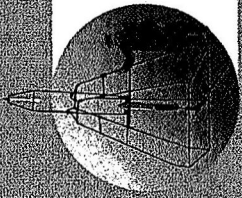
$$SNR = \frac{Q.E. * N_{ph}}{\sqrt{Q.E. * N_{ph} + N_b + N_d + N_r^2}}$$



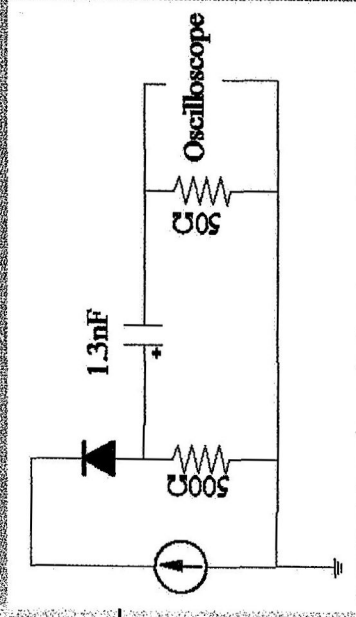
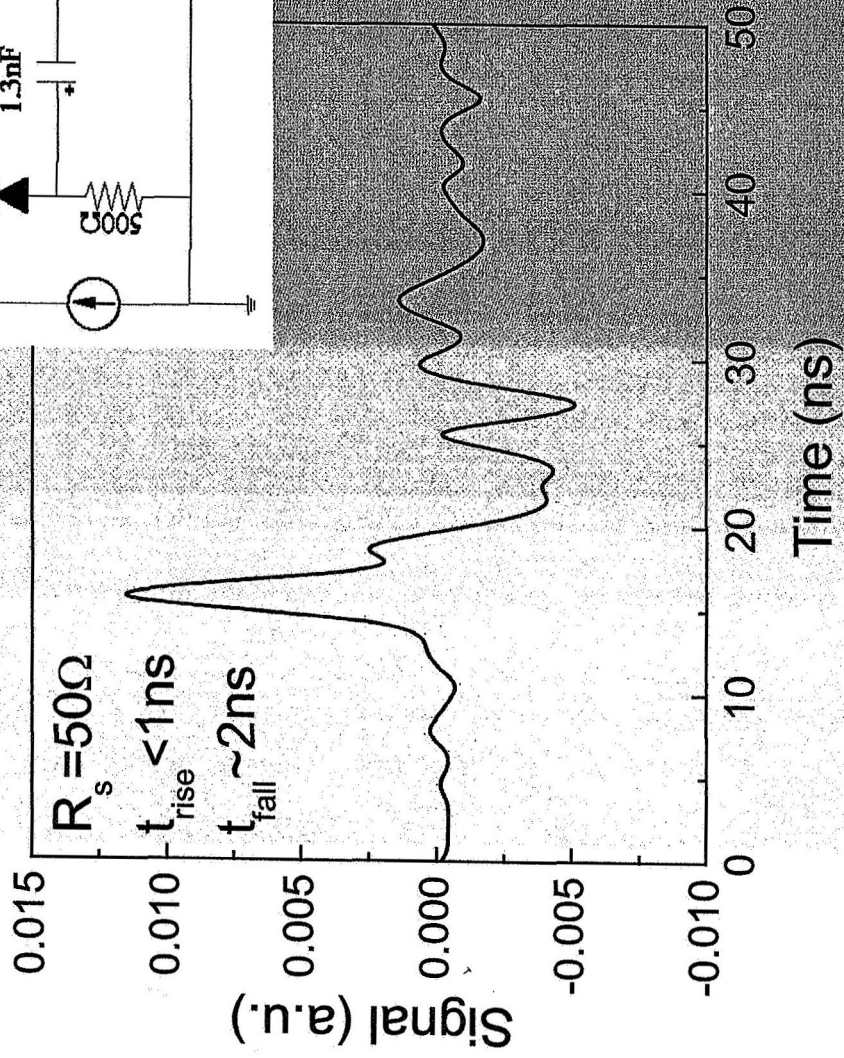


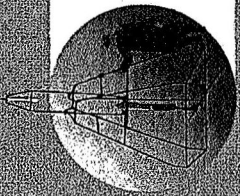
Structure of SiC single photon counting APD and testing structure



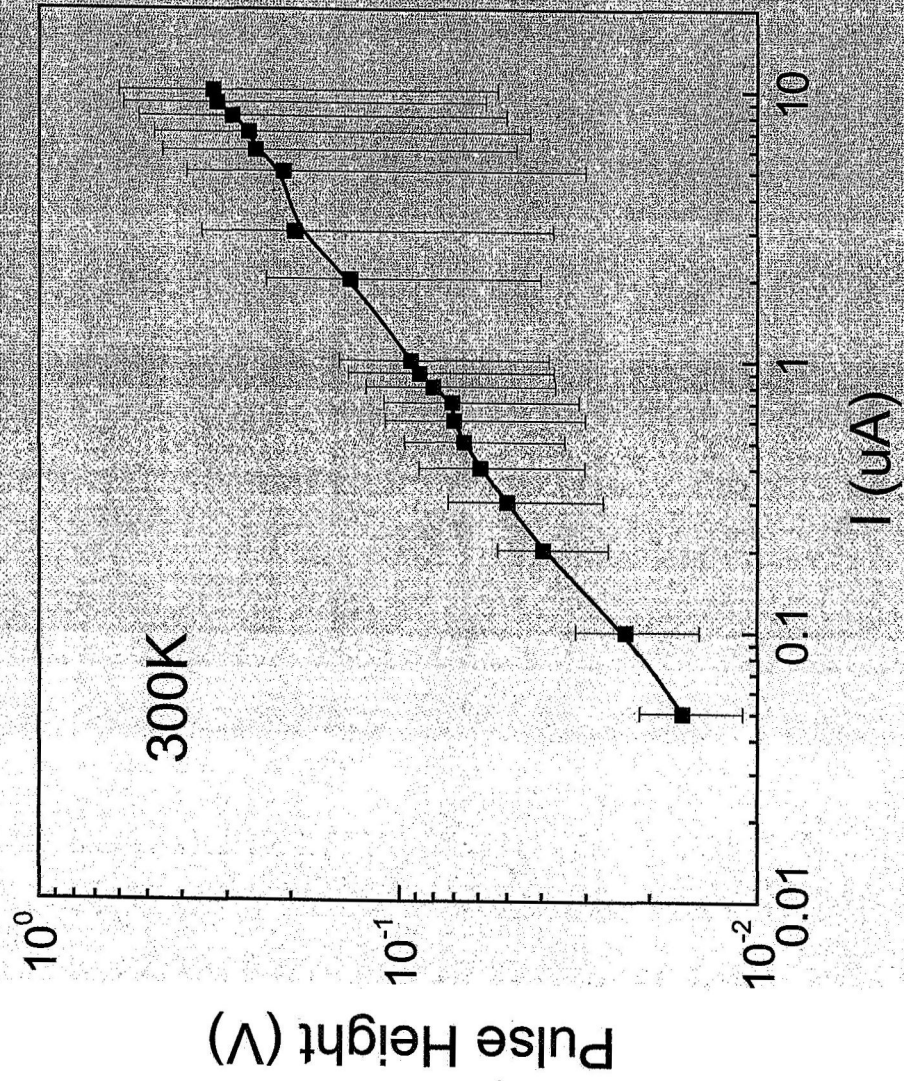


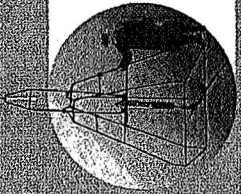
Single photon counting waveform and testing circuit



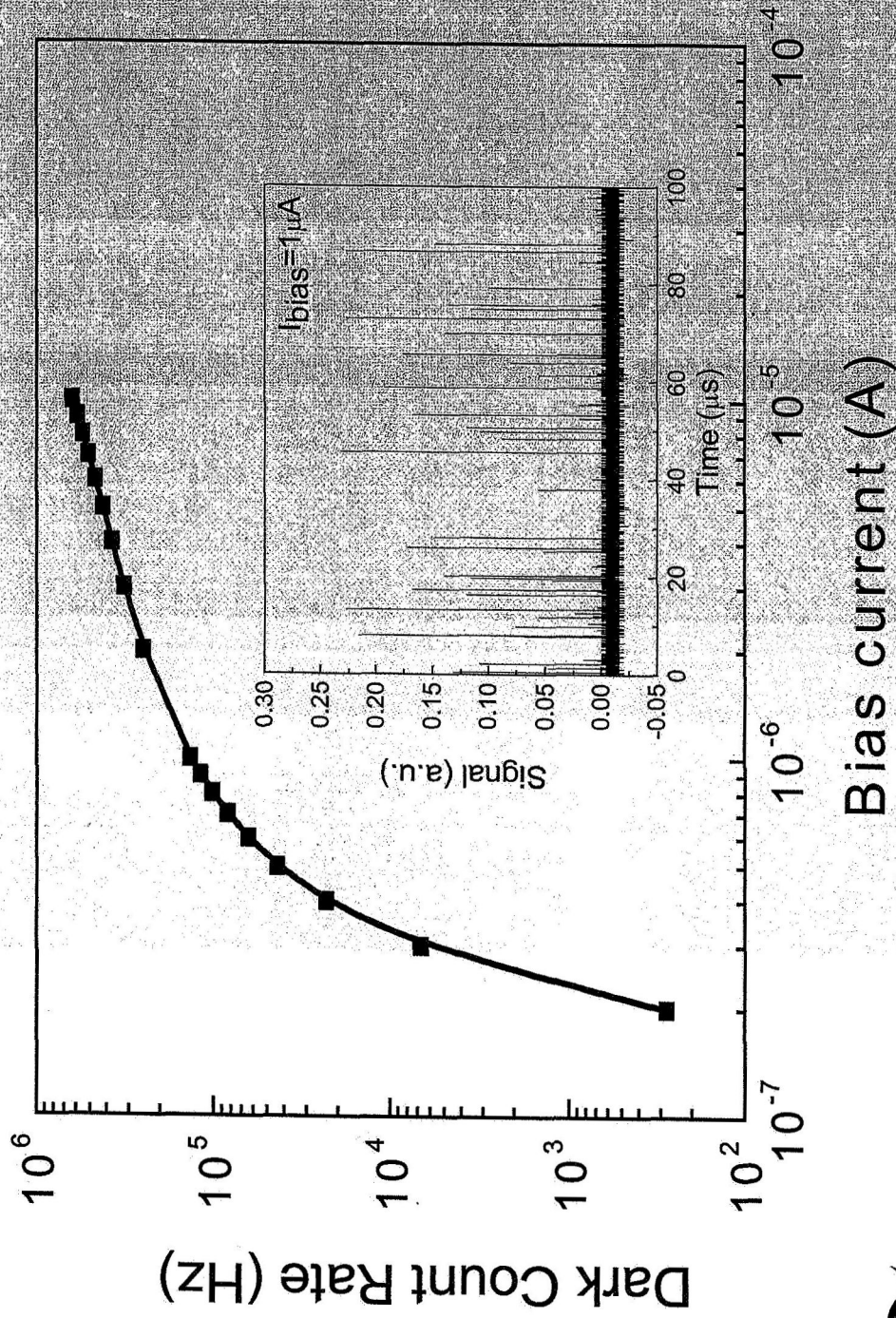


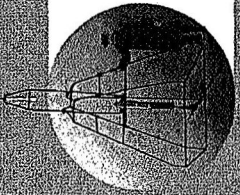
Amplitude of SiC single photon counter



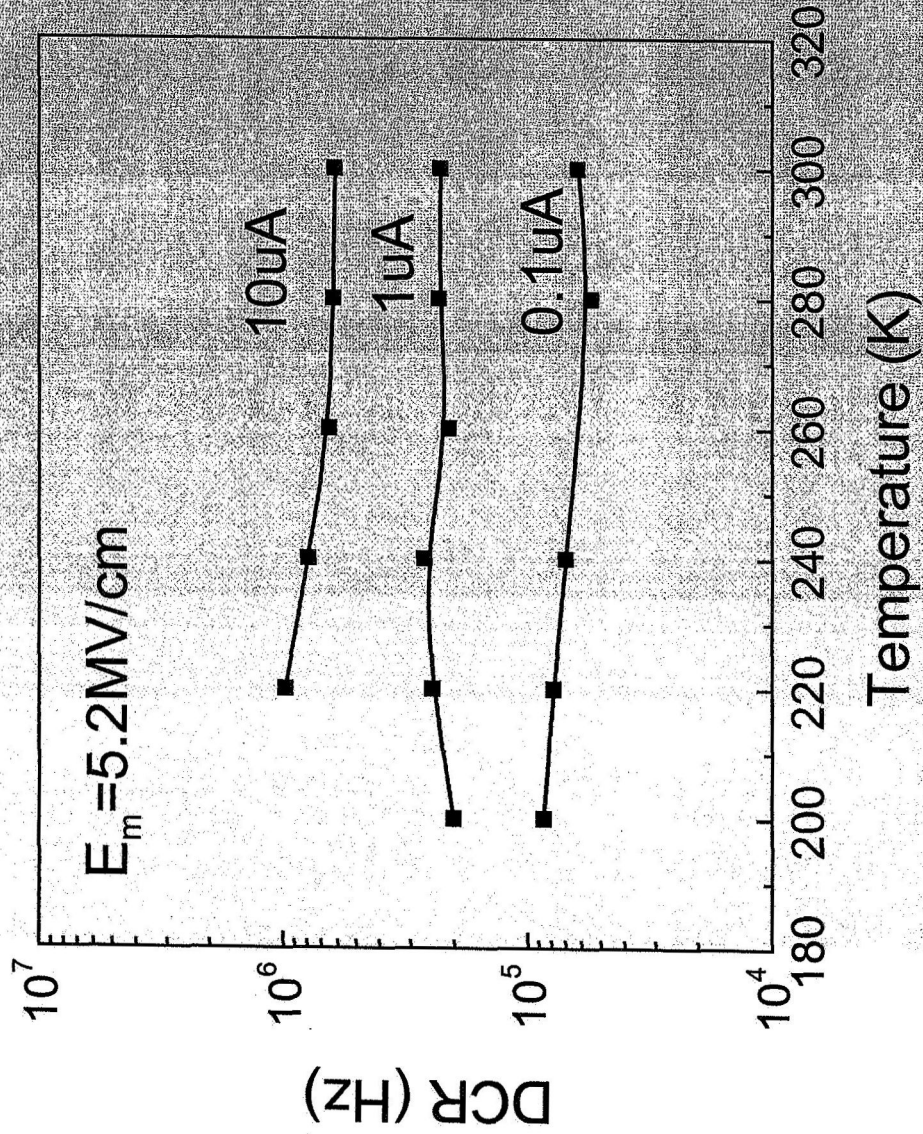


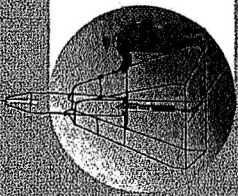
Dark count of SiC APD photon-counters



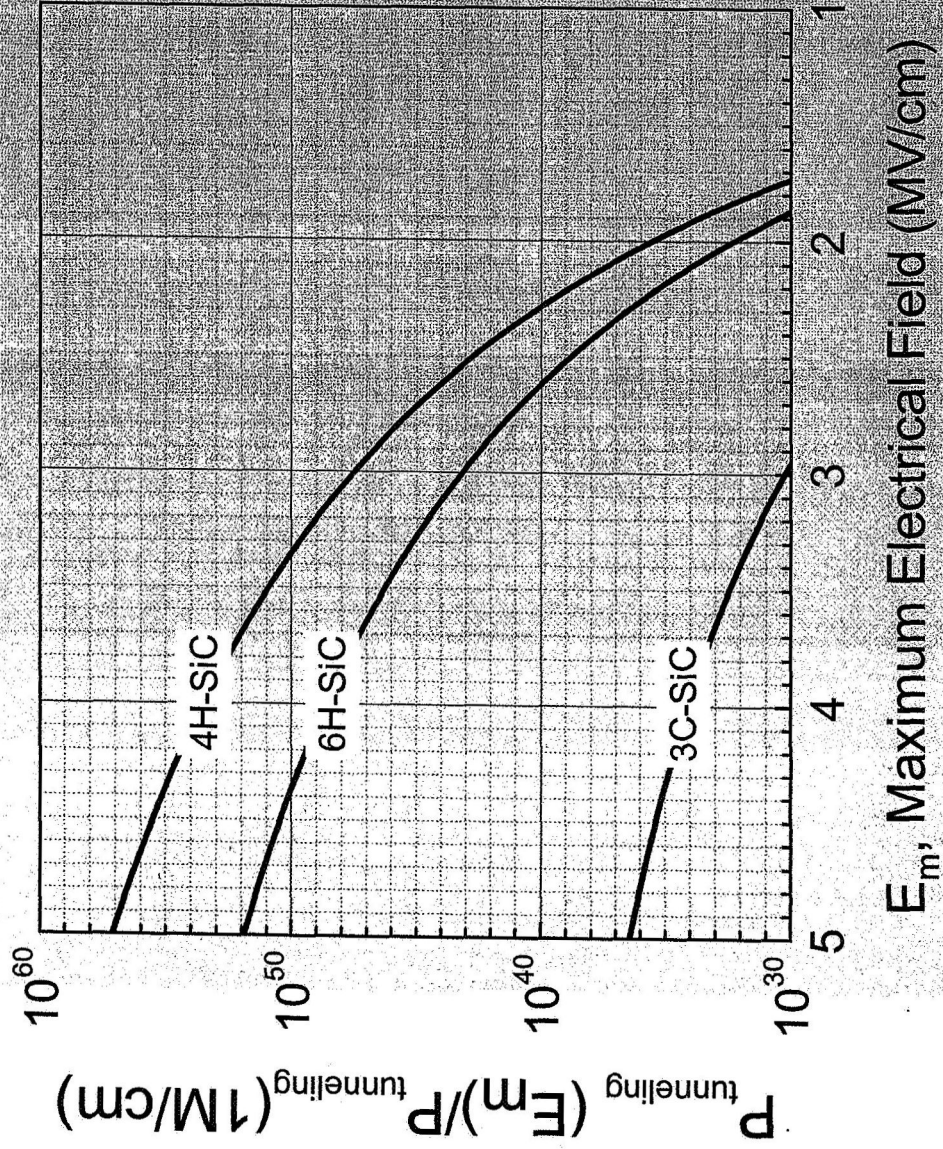


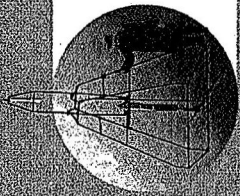
Temperature-dependence of dark count rate





Reduce the dark count rate by reducing the breakdown electric field

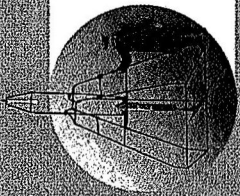




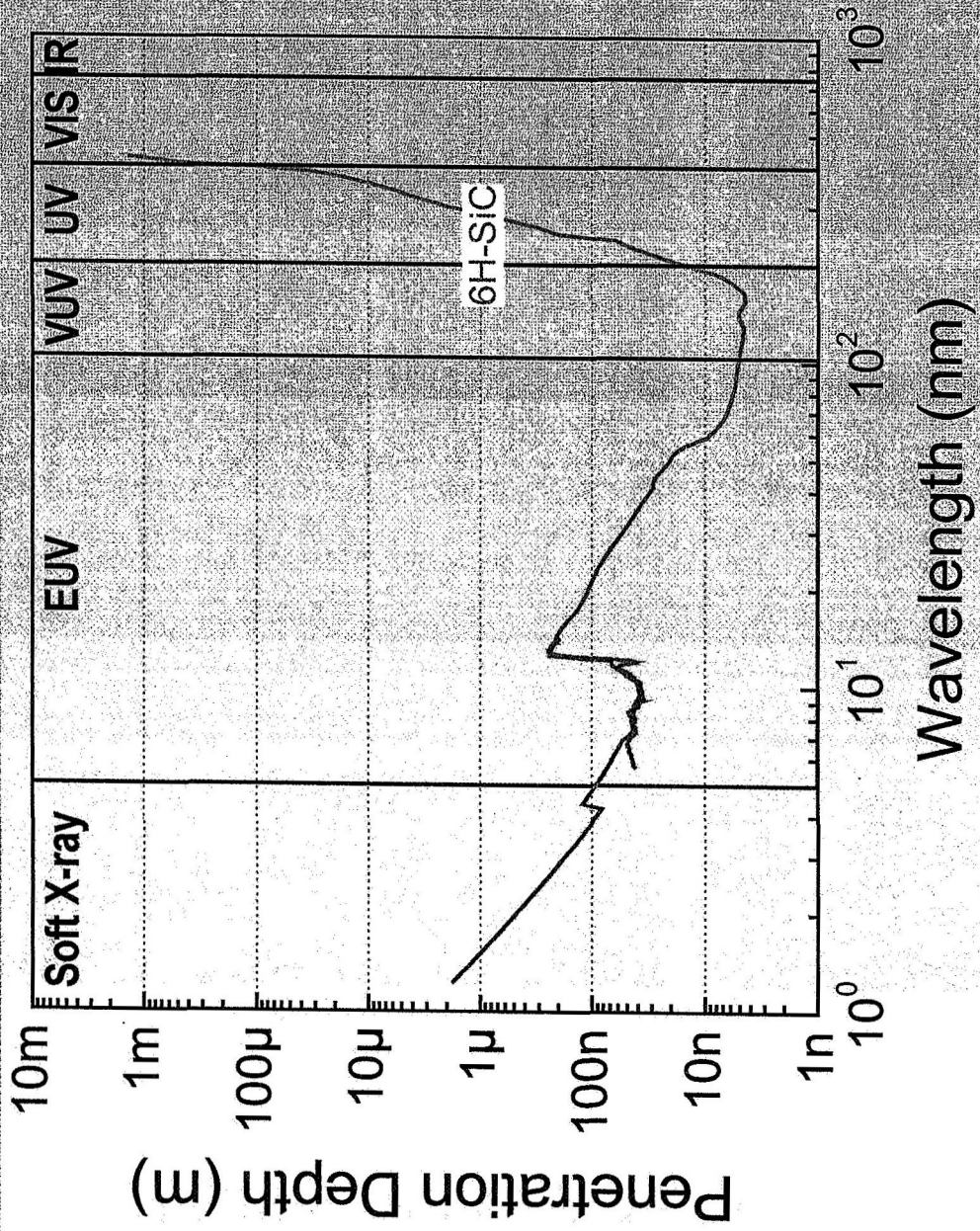
Quantum Efficiency

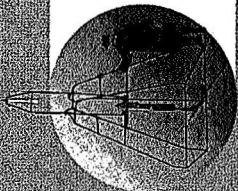


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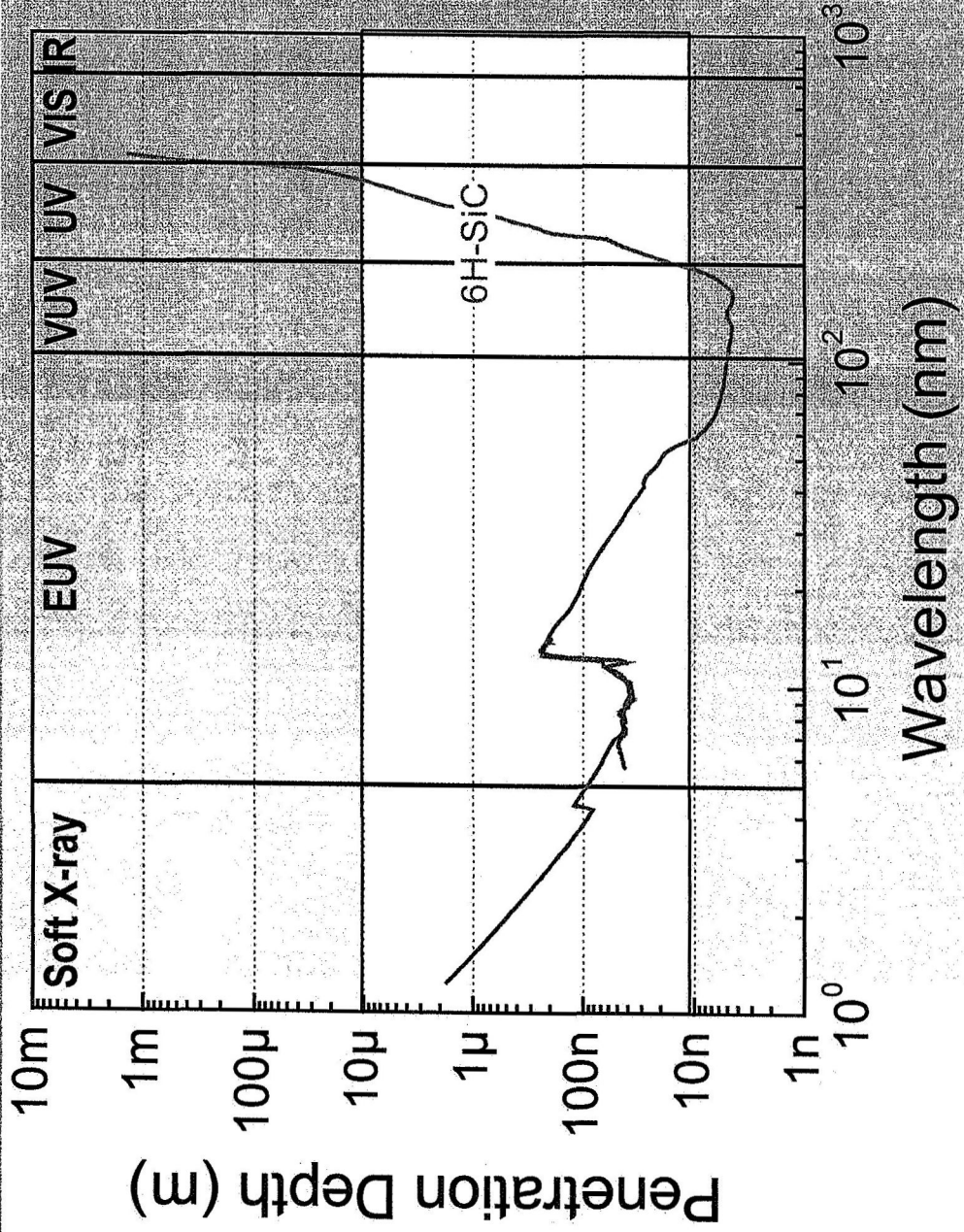


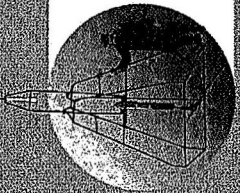
Spectrum range for SiC detectors



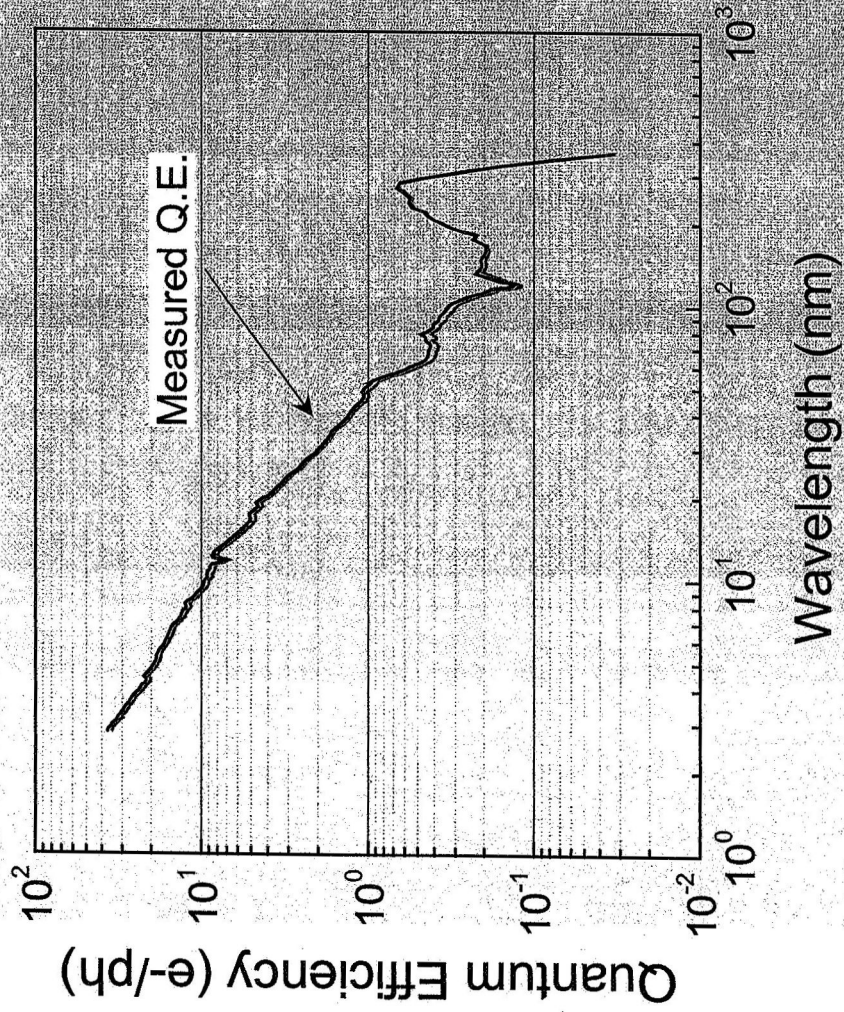


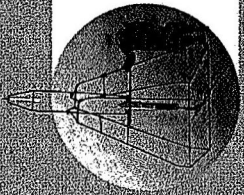
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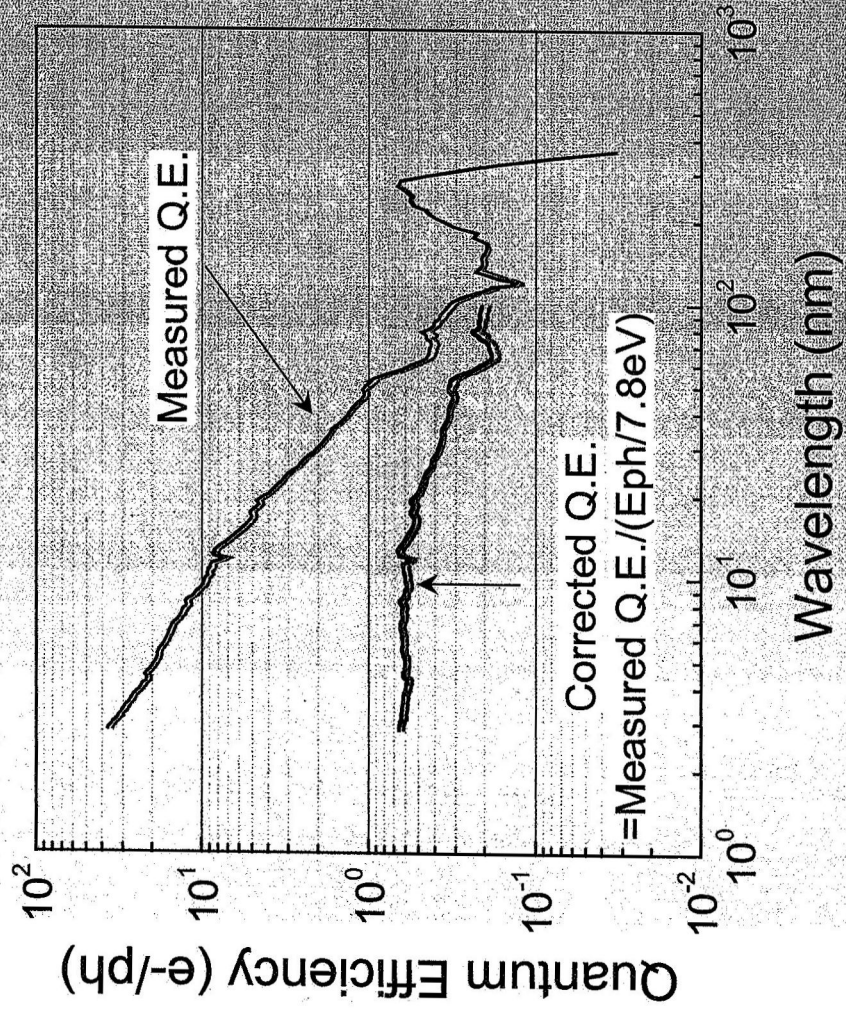


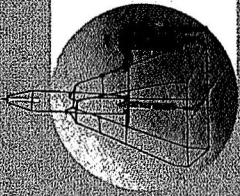
QE curves of Pt/4H-SiC photodiodes



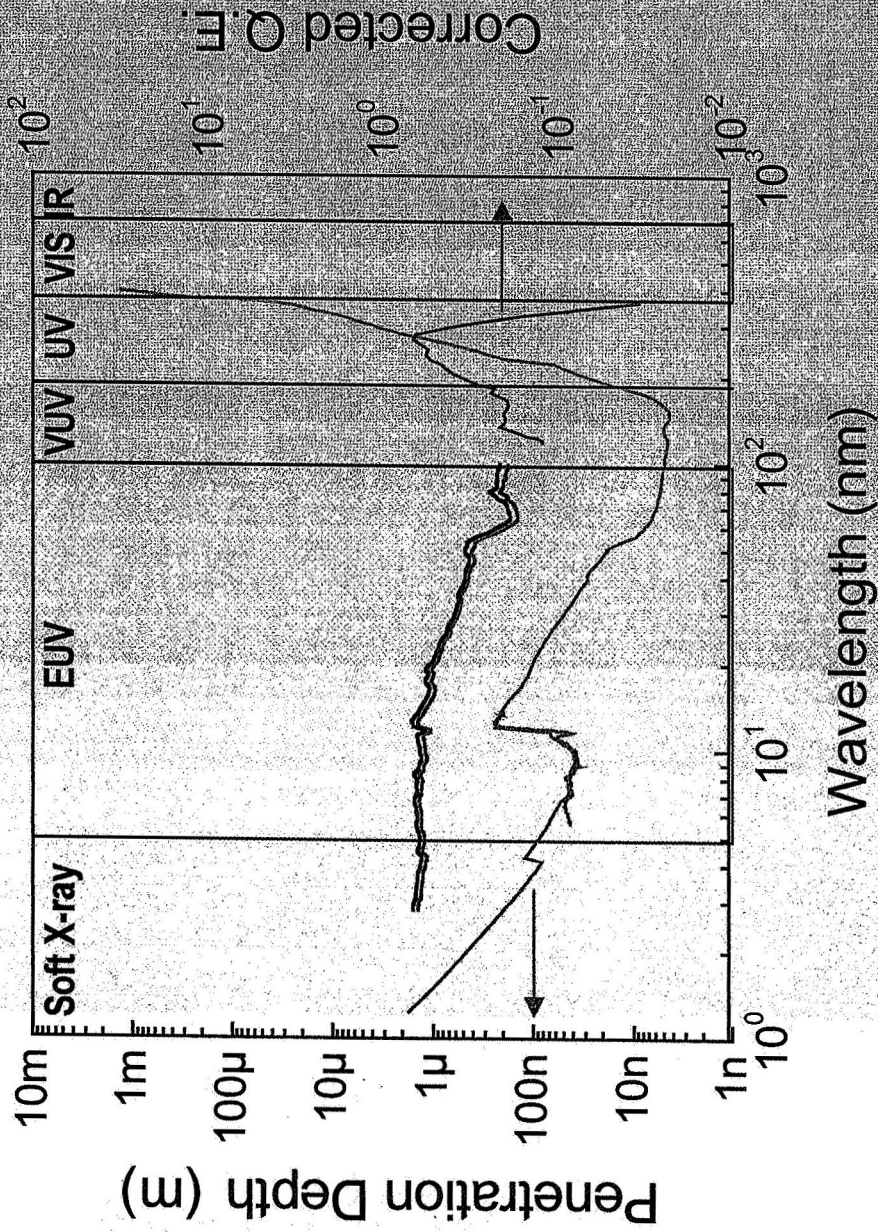


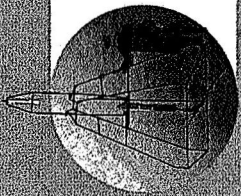
QE curve of SiC



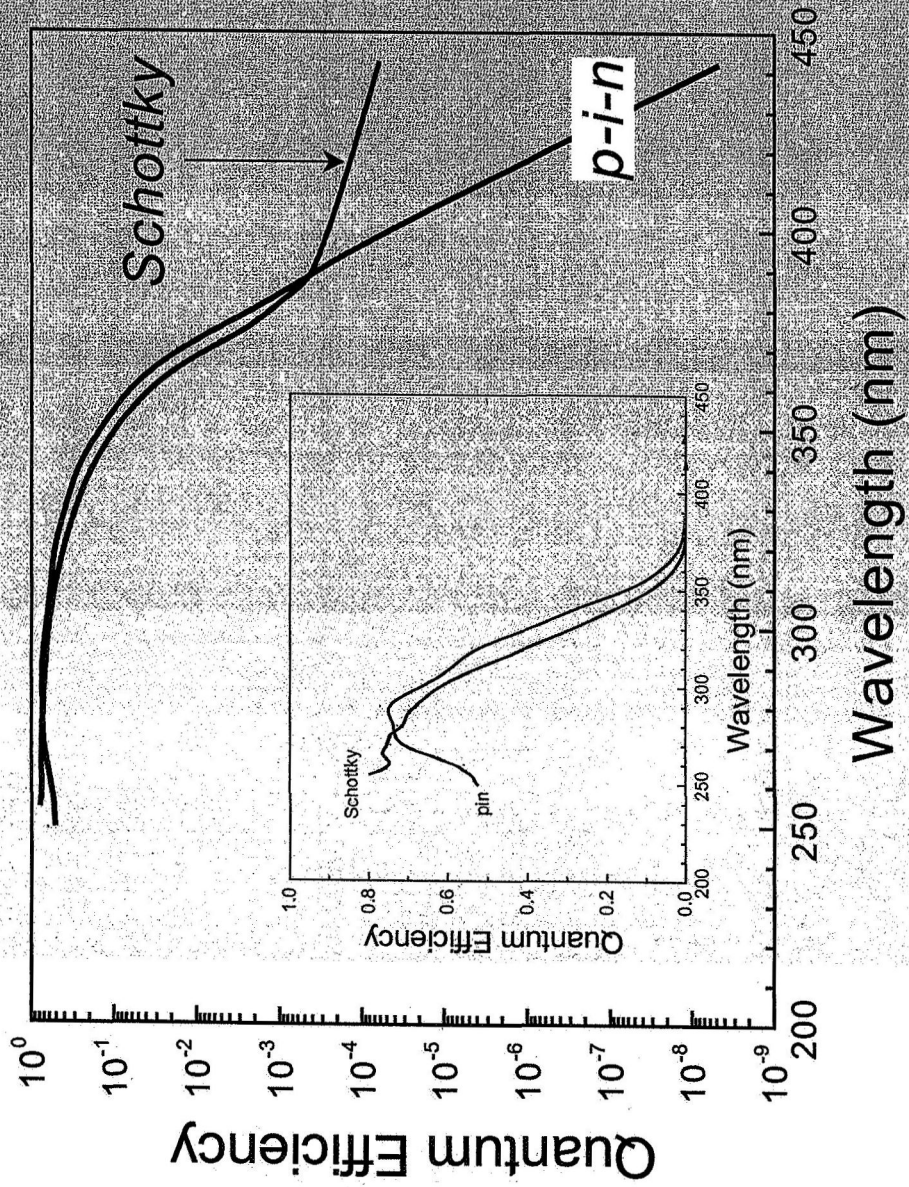


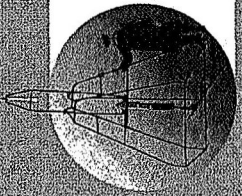
QE curves of SiC photodiode vs. penetration depth





Visible rejection of SiC photodiodes

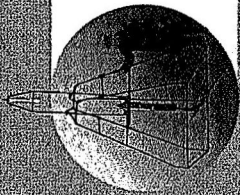




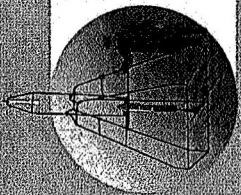
Advantages of SiC photodiodes

- Ideal for detector fabrication
 - Negligible dark current
 - Negligible read noise (SPADs)
 - Good Q.E. for 1nm~300nm
 - Blind to visible photons
- Additionally,
 - Good MOS interface, which allows to fabricate monolithic (!!!)
- SiC CCD (SiC EMCCD?)
- SiC CMOS-APS (active pixel sensor)
- Radiation sensors (for x-ray and particles)
- Wider operating temperature
- Excellent radiation hardness

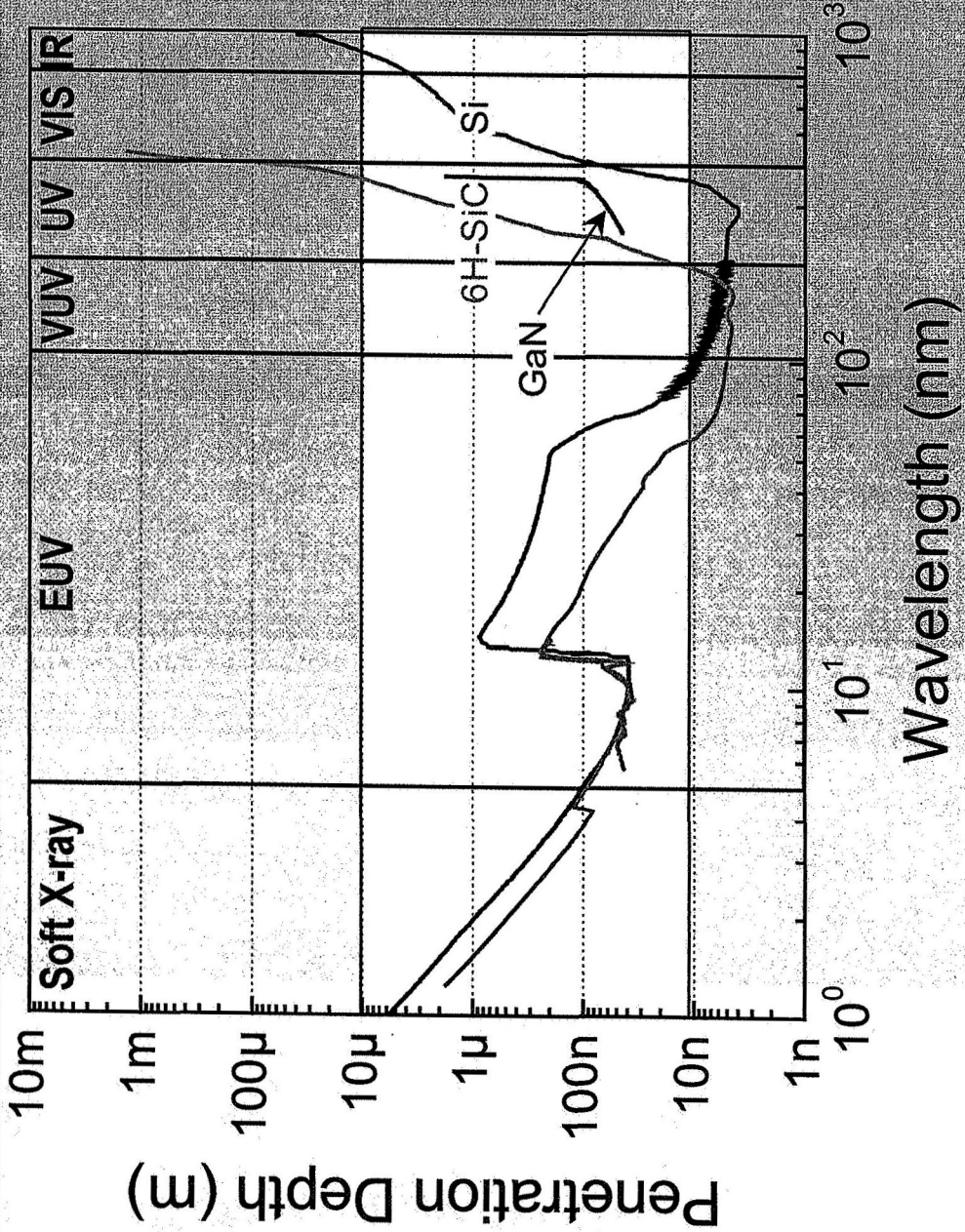


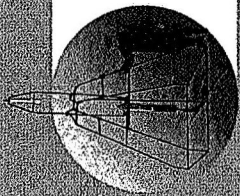


Applications

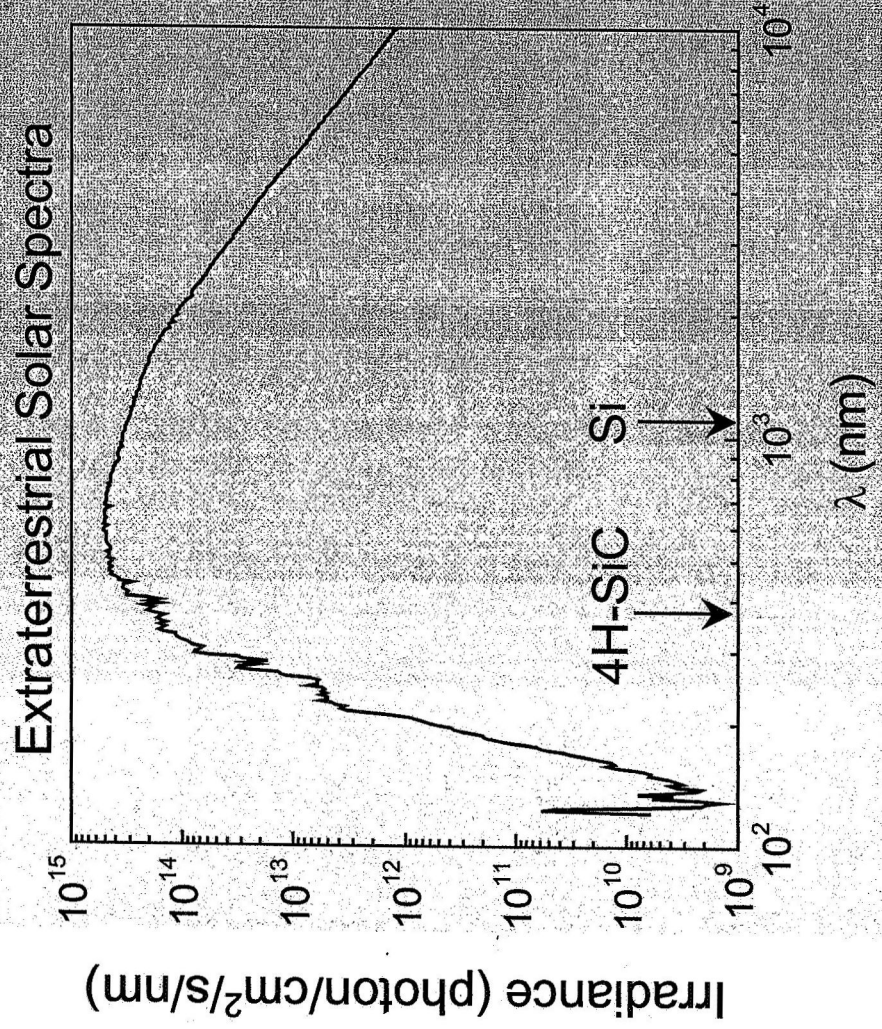


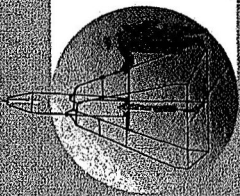
Competitors of SiC detectors



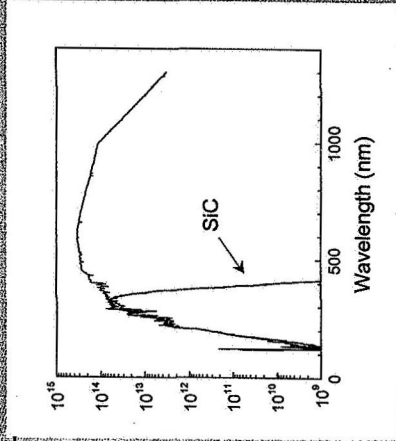
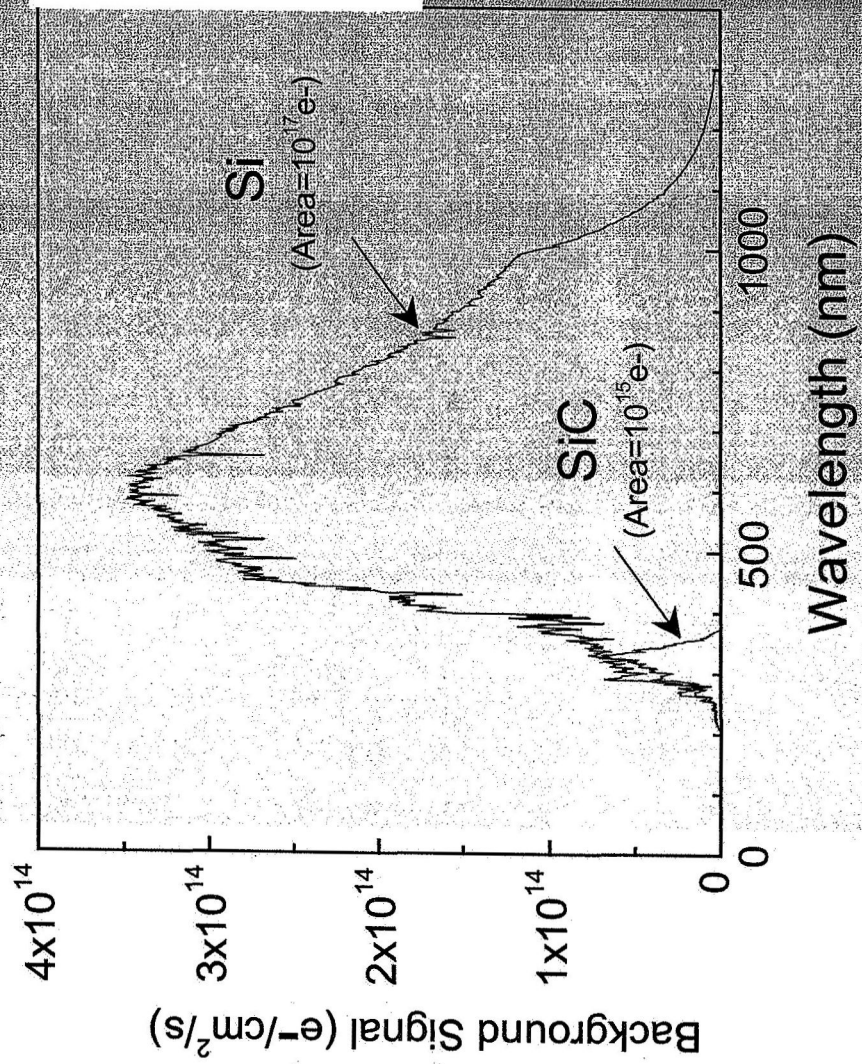


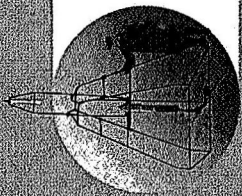
Extraterrestrial solar spectra



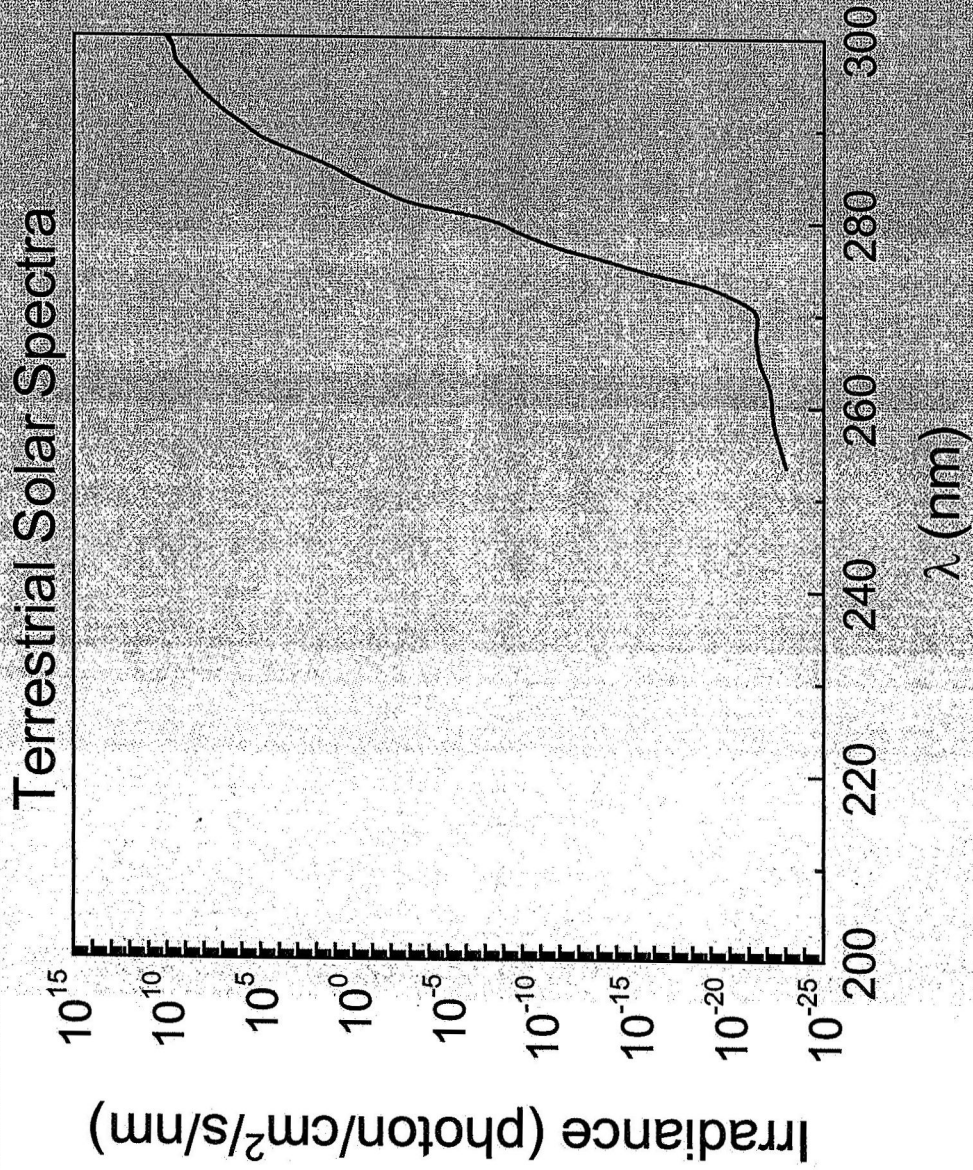


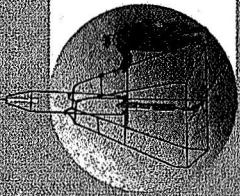
Visible-blind EUV detection





Terrestrial solar spectra

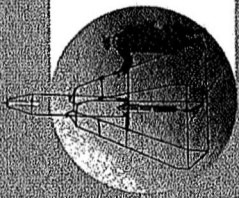




Less than 1keV soft x-ray detection

- Critical emission lines
 - Carbon $K\alpha_1$ (277 eV) → ~35 e⁻
 - Nitrogen $K\alpha_1$ (392.4 eV) → ~50 e⁻
 - Oxygen $K\alpha_1$ (524.9 eV) → ~67 e⁻
- No radiation detector module can cover this spectrum
- Radiation detector module: providing energy spectrum





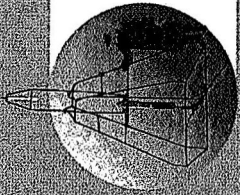
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Could SiC detectors or detector modules help ?

- *single photon-counting or*
- *avalanche gain*





Conclusion

- SiC has the elements to build nearly "ideal" detectors
 - Ultra low dark current
 - High avalanche gain with low k value
 - Single photon-counting at room temperature
- SiC is suitable for the visible blind detection of
 - UV
 - VUV and EUV
 - Soft X-ray for less than 1keV
- SiC can be a good candidate for solar blind UV, EUV and Soft X-ray (<1keV) detection

