

**<sup>39</sup>Ar-<sup>40</sup>Ar ‘AGE’ OF BASALTIC SHERGOTTITE NWA-3171.**  
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North-West-Africa 3171 is a 506 g, relatively fresh appearing, basaltic shergottite with similarities to Zagami and Shergotty, but not obviously paired with any of the other known African basaltic shergottites [1,2]. Its exposure age has the range of 2.5-3.1 Myr [3], similar to those of Zagami and Shergotty [4]. We made <sup>39</sup>Ar-<sup>40</sup>Ar analyses of a “plagioclase” (now shock-converted to maskelynite) separate and of a glass hand-picked from a vein connected to shock melt pockets.. Plagioclase was separated using its low magnetic susceptibility and then heavy liquid with density of <2.85 g/cm<sup>3</sup>.

The <sup>39</sup>Ar-<sup>40</sup>Ar age spectrum of NWA-3171 plag displays a rise in age over 20-100% of the <sup>39</sup>Ar release, from 0.24 Gyr to 0.27 Gyr. The first 20% of the <sup>39</sup>Ar release involves terrestrial weathering products characterized by adsorbed terrestrial Ar and likely terrestrial K contamination. Over the last 80% of the <sup>39</sup>Ar release, constant values of the <sup>36</sup>Ar/<sup>37</sup>Ar ratio indicate that essentially all <sup>36</sup>Ar released is cosmogenic. An isochron plot (<sup>40</sup>Ar/<sup>36</sup>Ar vs. <sup>39</sup>Ar/<sup>36</sup>Ar) of these data (R<sup>2</sup>=0.996) has a slope corresponding to an age of 225 ±4 Myr. Essentially the same age is obtained whether we use total <sup>36</sup>Ar or correct for trapped <sup>36</sup>Ar. A radiometric formation age for NWA-3171 has not yet been reported. However, the Ar-Ar age spectrum of NWA-3171 closely resembles that of Zagami, and the Arrhenius diffusion plots of <sup>39</sup>Ar for the two shergottites also are similar. Thus, the “true” age of NWA-3171 may be similar to the Zagami age (177 ±3 Myr; [5]). This implies NWA contains an extra component of <sup>40</sup>Ar, not accompanied by significant trapped <sup>36</sup>Ar, an inference that we have made for Zagami as well (Bogard & Park, this volume). We suggest this excess <sup>40</sup>Ar was inherited from the basaltic melt.

The <sup>39</sup>Ar-<sup>40</sup>Ar age spectrum for the glass inclusion is very different and shows apparent Ar-Ar ages ranging between 0.3 and 1.9 Gyr. Variations in the <sup>36</sup>Ar/<sup>37</sup>Ar ratios indicate release of trapped <sup>36</sup>Ar throughout most of the extraction. An isochron plot of <sup>36</sup>Ar/<sup>40</sup>Ar vs <sup>39</sup>Ar/<sup>40</sup>Ar suggests the release of terrestrial Ar in the first ~30% of the <sup>39</sup>Ar release, and high K/Ca ratios in these extractions also suggest terrestrial weathering. We used trapped <sup>36</sup>Ar in the isochron by subtracting a cosmogenic <sup>36</sup>Ar<sub>cos</sub> component obtained from average data reported for Zagami and Shergotty whole rock [4, 6-8]. Measured <sup>36</sup>Ar/<sup>37</sup>Ar ratios were used to apportion this <sup>36</sup>Ar<sub>cos</sub> over individual releases. The temperature extraction data of 780-1160 °C (corresponding to ~37%-93% of the <sup>39</sup>Ar release) define a mixing line between a radiogenic component and Martian atmospheric Ar with <sup>40</sup>Ar/<sup>36</sup>Ar ≅1800, consistent with previously reported values for Mars atmospheric Ar [9].. Like other impact glasses in shergottites, NWA-3171 glass contains martian atmosphere incorporated at the time of impact formation, and does not directly yield a formation age.

References: [1] Irving A. J. et al. 2005, **MaPS** 39, A49. [2] Meyer C. 2004, The Mars Meteorite Compendium. [http://www.curator.jsc.nasa.gov/antmet/mmc/XXXII\\_NWA3171.pdf](http://www.curator.jsc.nasa.gov/antmet/mmc/XXXII_NWA3171.pdf) [3] Nishiizumi & Caffee 2006, **MaPS** 69, A133. [4] Eugster et al. 1997, **GCA** 61:2749-2757. [5] Nyquist et al. 2001, **Space Sci. Rev.** 96,:105-164. [6] Park J. 2005, Ph.D. Thesis, Univ. of Tokyo. [7] Schwenzer et al. 2007, **MaPS** 42, 387-412. [8] Terribilini et al. 1998, **MaPS** 33, 677-684. [9] Bogard & Garrison, 1999. **MaPS** 34, 451.