

# Matlab based toolkits used to interface with optical design software for NASA's James Webb Space Telescope

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*2007 SPIE Optics and Photonics, San Diego, August 28*





## Acknowledgments

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Special thanks to  
Blair Unger (University of Rochester)  
and  
Mark Wilson (NASA)  
for their assistance in creating the  
Matlab-CodeV Toolkit



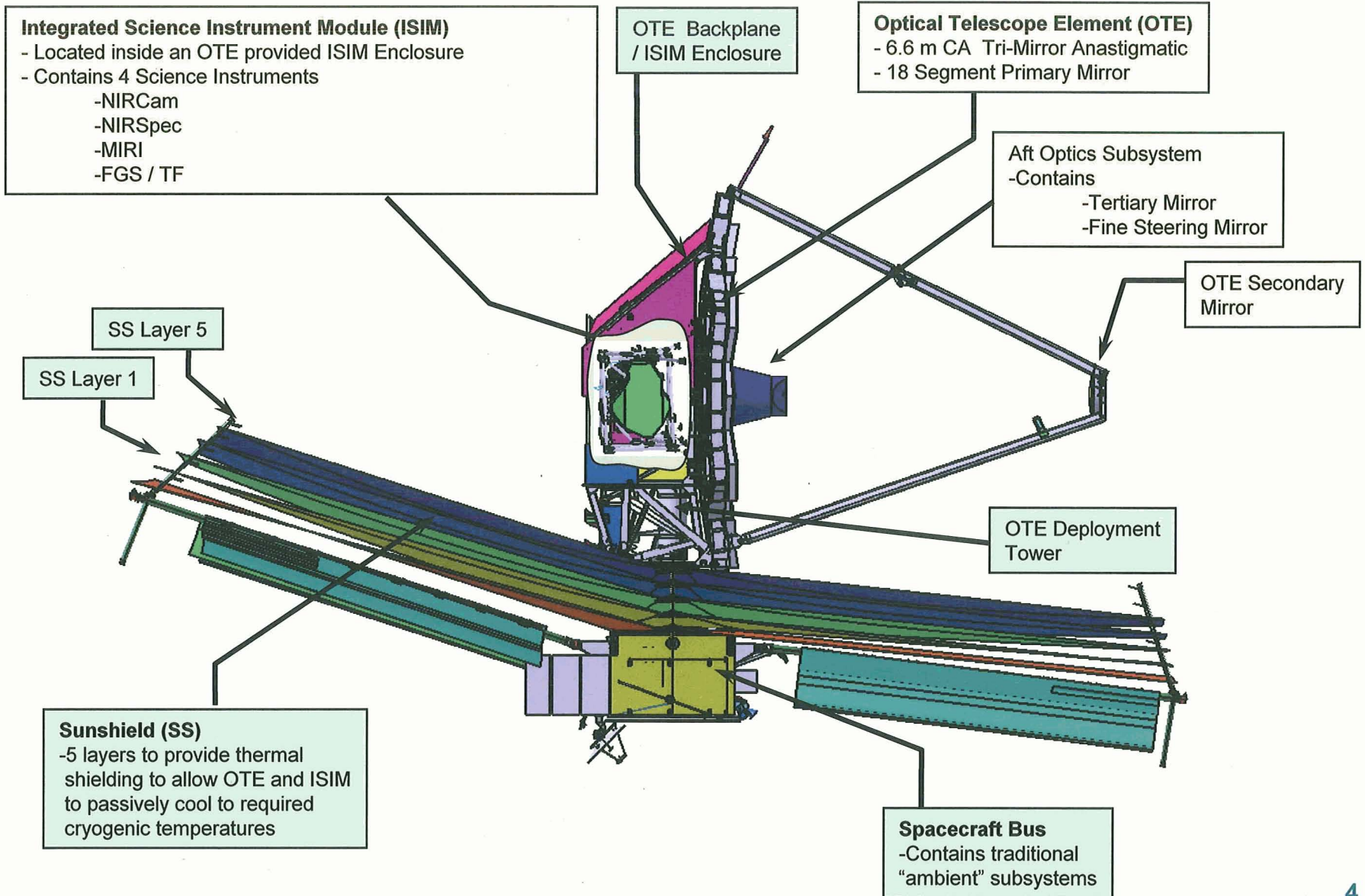
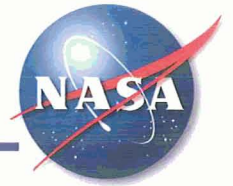
## Outline

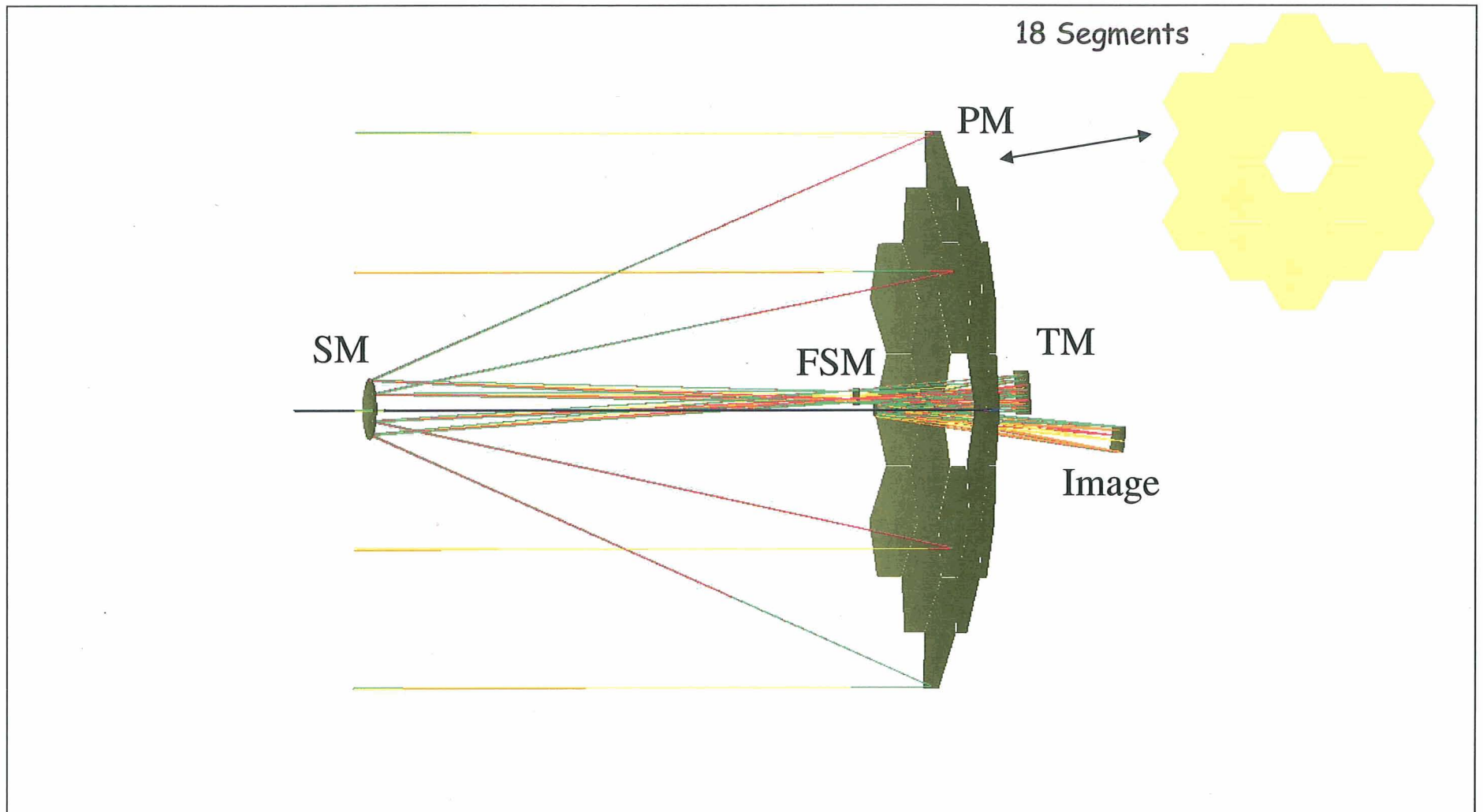
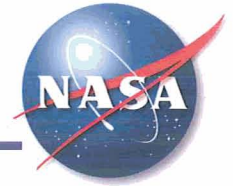


1. Introduction to JWST
2. Brief overview of Matlab toolkits
  - CodeV Toolkit
  - OSLO Toolkit
  - Zemax Toolkit
3. Examples of use with JWST
  - Wavefront sensitivities
  - Alignment simulations
4. Where to get them
5. Concluding remarks



# James Webb Space Telescope (JWST)





Three-Mirror-Anastigmat (TMA) wide-field telescope design



## Outline



### Part 1: Toolkit Overview:

- ◆ Purpose of the Toolkit
  - Intended audience
  - Value added over CodeV alone
- ◆ Toolkit Layout
  - System Functions / Utilities
  - Lens Info / Manipulation Functions
  - Analysis Tools / Ray Tracing Functions
- ◆ How Matlab gets data from CodeV
  - COM+
  - The CodeV Buffer
- ◆ Function Layout
  - Inputs: Use and syntax of individual functions
  - Outputs: Displaying data and using raw matrix data
- ◆ Using cvHELP

### Part 2: A Brief Tutorial

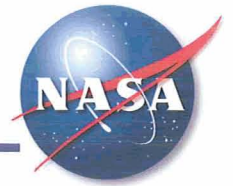
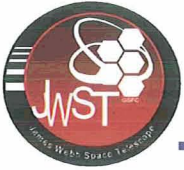
- ◆ Starting CodeV from Matlab, loading a lens file, and getting system data
- ◆ OPD Maps - cvPMA
- ◆ PSFs - cvPSF
- ◆ Ray Tracing
  - Single rays - cvr, cvRSI
  - Grids of rays - cvRayGrid, cvRayTra
- ◆ Encircled Energy - cvENC
- ◆ Sensitivity Analysis
  - Decentering / Tilting Lens Elements
  - Linear Optical Model: cvLOM
- ◆ Displaying a lens from CodeV in Matlab

### Part 3: Live demonstration (hopefully)

- ◆ OPD output using cvPMA
- ◆ PSF output using cvPSF



- ◆ **Intended audience**
  - **Non-CodeV Users:** Those with optical design and analysis experience wanting to perform analyses using CodeV.
  - **CodeV Users:** Those wanting additional analysis functions and an easy method to run myriad sensitivity analyses.
- ◆ **Value added over CodeV alone**
  - Commands are standardized and easy to pick up for anyone with Matlab experience.
  - Data output is Matlab MAT files. Graphical functions are easily modified.
  - Sensitivities can be analyzed with simple "for" loops.



## ◆ System Functions / Utilities

- cvon** - establishes the COM link between Matlab and CodeV
- cvoff** - kills the COM link between CodeV and Matlab
- cvin** - inputs .seq file into CodeV COM
- cvopen** - inputs .len file into CodeV COM
- cvsave** - saves current lens file under pathfilename

## ◆ Lens Info / Manipulation Functions

- cvgetf** - gets the desired fieldpoints in CodeV
- cvsetf** - sets fieldpoints in CodeV
- cvgetw** - gets desired wavelengths and weights in CodeV
- cvsetw** - sets wavelengths and weights in CodeV
- cvsetz** - sets active zoom positions in CodeV
- cvgetap** - gets the aperture data from the current lens
- cvims** - gets image surface number for lens in CodeV
- cvshift** - perturbs the CodeV file by rigid body motion
- cvrshift** - "shifts" the decenters for single surface 'surfnum'
- cvgc** - get the global coordinates for a lens
- cvlensdata** - CVLENSDATA gets the lens data for the current lens
- cvdraw** - draws the current lens in Matlab





### ◆ Ray Tracing Functions

- cvr** - gets CodeV ray trace data for reference rays from database
- cvray** - gets CodeV ray trace data from RAYRSI macro function
- cvRSI** - gets CodeV ray trace data from RSI command
- cvraygrid** - returns a grid of ray data for a particular surface and datatype
- cvraytra** - uses CodeV RAYTRA engine to calculate a grid of rays.

### ◆ Analysis Tools

- cvENC** - gets CodeV PSF based encircled energy for fieldpoint 1
- cvPMA** - gets CodeV exit pupil wavefront, mask, RefRad, f/#, and focal length data
- cvPSF** - gets CodeV PSF for fieldpoint 1, normalized to perfect lens (Strehl)
- cvsens** - gets CodeV sensitivity data at image based on
- cvWAV** - gets wavefront analysis data from CodeV
- cvspot** - graphs a spot diagram for a given field and ray density
- cvlom** - calculates the linear optical model for the given parameters



### ◆ Utilities

- cvbufgetarray** - gets continuous data from rows and cols in CodeV buffer
- cvbufgetrow** - gets continuous data from rows in CodeV buffer
- cvbufnum** - gets a single number from the applicable CodeV buffer
- cvbufstr** - gets a single string from the applicable CodeV buffer
- cvcmd** - sends command to CodeV command line over existing COM link
- cvEVA** - Sends command to CodeV command line over existing COM link
- cvgetINT** - Reads in an int file and plots it

### ◆ Project Specific Scripts

- cvnircam** - this script loads JWST NIRCAM Short wavelength channel into CodeV
- cvjwst** - This script loads JWST segmented OTE file into CodeV, and puts fieldpoints into the workspace.



### ◆ CodeV and COM

- CodeV API (Application Programming Interface) uses the Windows COM (Component Object Model) interface for passing commands and data between programs.
- Any VB, C++, program can send commands to CodeV and get data back from CodeV (Refer to the *CODE V API Reference Guide* for details.)
- Many functions are built into the COM for easy calls to ray trace data, OPD data, and PSFs. Other data can be sent to the CodeV buffer.

### ◆ The CodeV Buffer

- The CodeV Worksheet Buffer allows for fast access to any data output from CodeV.
- Once CodeV writes data to the buffer, COM can immediately read that data Matlab



## Function Layout



- ◆ **Inputs:** Use and syntax of individual functions  
`[output1, output2, ...] = function(input1, input2, ...);`
  - "function" is any of the functions starting with CV
  - "input" is the particular input requested by the function
  - "output" is the variable name where the output is going

- ◆ **Outputs:** Displaying data and using raw matrix data (example)

- To display a PSF to screen, omit the outputs

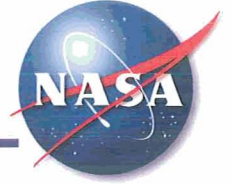
`cvpsf(128,64);` will create a PSF plot with a transform grid of 128 and 64 rays across the diameter.

- To output the PSF data to a matrix, enter the outputs preceding the function.

`[psf,grid_spacing] = cvpsf(128,64);` will create 2 arrays: psf- containing the psf data (128 by 128), and grid\_spacing- containing the grid spacing for the psf (single value).



# Toolkit output- data or graphical



Input: `[psf, grid_spacing] = cvpsf(256,64);`

```

psf(124:134,124:134) =
3.4408    2.3350    0.9989    0.2010    0.0471    0.0909    0.0471    0.2010    0.9989    2.3350    3.4408
3.0428    0.8452    0.0551    1.8657    4.9077    6.3941    4.9077    1.8657    0.0551    0.8452    3.0428
1.6787    0.0032    3.1947    12.2560    22.6098    27.1931    22.6098    12.2560    3.1947    0.0032    1.6787
0.5103    1.2104    11.2099    30.6998    50.9006    59.5509    50.9006    30.6998    11.2099    1.2104    0.5103
0.0719    3.5264    19.9491    48.6187    77.2257    89.3051    77.2257    48.6187    19.9491    3.5264    0.0719
0.0199    4.5292    23.2390    55.0783    86.5443    99.7826    86.5443    55.0783    23.2390    4.5292    0.0199
0.0939    3.1336    18.5847    45.8443    73.1451    84.6886    73.1451    45.8443    18.5847    3.1336    0.0939
0.6374    0.8681    9.5687    27.0564    45.3521    53.2123    45.3521    27.0564    9.5687    0.8681    0.6374
1.8470    0.0102    2.3169    9.7872    18.5329    22.4334    18.5329    9.7872    2.3169    0.0102    1.8470
3.0027    0.9877    0.0012    1.1263    3.3522    4.4724    3.3522    1.1263    0.0012    0.9877    3.0027
3.0544    2.2104    1.0918    0.3145    0.0374    0.0052    0.0374    0.3145    1.0918    2.2104    3.0544

```

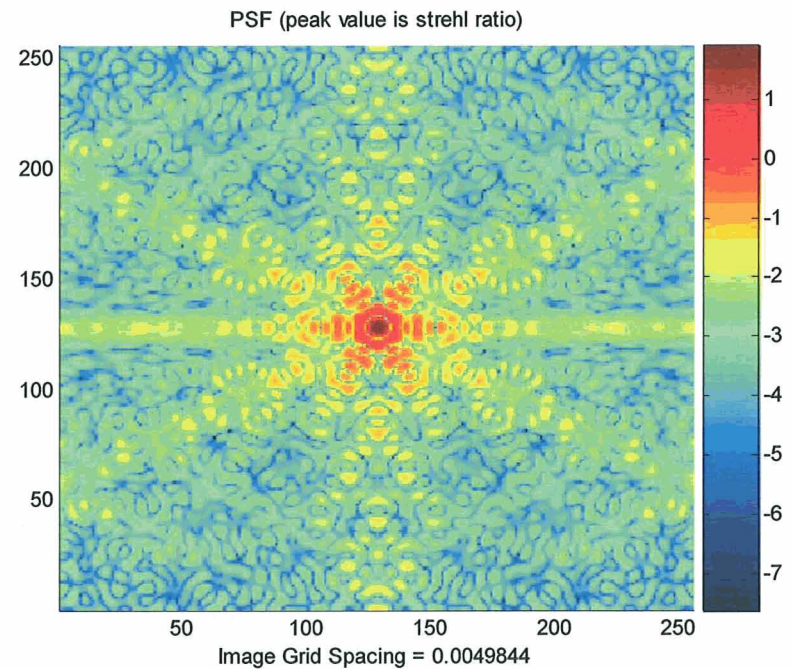
```

grid_spacing =
0.0049843814 mm

```

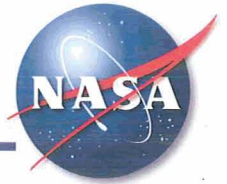
Numeric Output

Graphical Output





## Using cvHELP



- ◆ To look up the syntax for any function, simply type "*help function*" where "function" is the CodeV function you want.
- ◆ To see a list of all the Matlab-CodeV toolkit functions, simply type "*cvhelp*" and a list with descriptions of all the functions will appear. Then just click on one.
- ◆ Each function has help on command syntax and usage. Also most functions have defaults set so you can omit all or most inputs for quick analyses.



## Tutorial: Starting CodeV from Matlab, loading a lens file, and getting system data



### 1. Start a CodeV server session:

- ◆ Type `cvon;`
- ◆ A variable called CVhandle will appear. This shows that CodeV has been started.

### 2. Load a lens file:

- ◆ For .seq files, use: `cvin('filename');` (you need the '' part)  
or just `cvin;` (Matlab will graphically prompt you for a filename.)
- ◆ For .len files, use `cvopen('filename');`  
or just `cvopen;` (Matlab will graphically prompt you for a filename.)

### 3. Get fields and wavelengths:

- ◆ To get field info. and put it in the workspace, type: `[fx,fy] = cvgetf;`
- ◆ To get system wavelengths and weights, type: `[wl,wt] = cvgetw;`

### 4. Set fields and wavelengths:

- ◆ To set fields or wavelengths, either enter a vector containing all of the parameters or enter the field number. If a field number is entered, all others will be removed and that field will now be field 1.
- ◆ Set field angle to field 10: `cvsetf(10);` Set wavelength to wl 5: `cvsetw(5);`
- ◆ Set field to x, y angles = 0: `cvsetf(0,0);` Set wl to 2000nm with weight 1: `cvsetw(2000,1);`



# Tutorial: OPD Maps - cvPMA



```
[opd,mask,pin,Xfno,Yfno,RefRad,Xf1,Yf1]=  
cvPMA(TGR,fit_option,f,w,z,NRD_GRI);
```

## INPUTS:

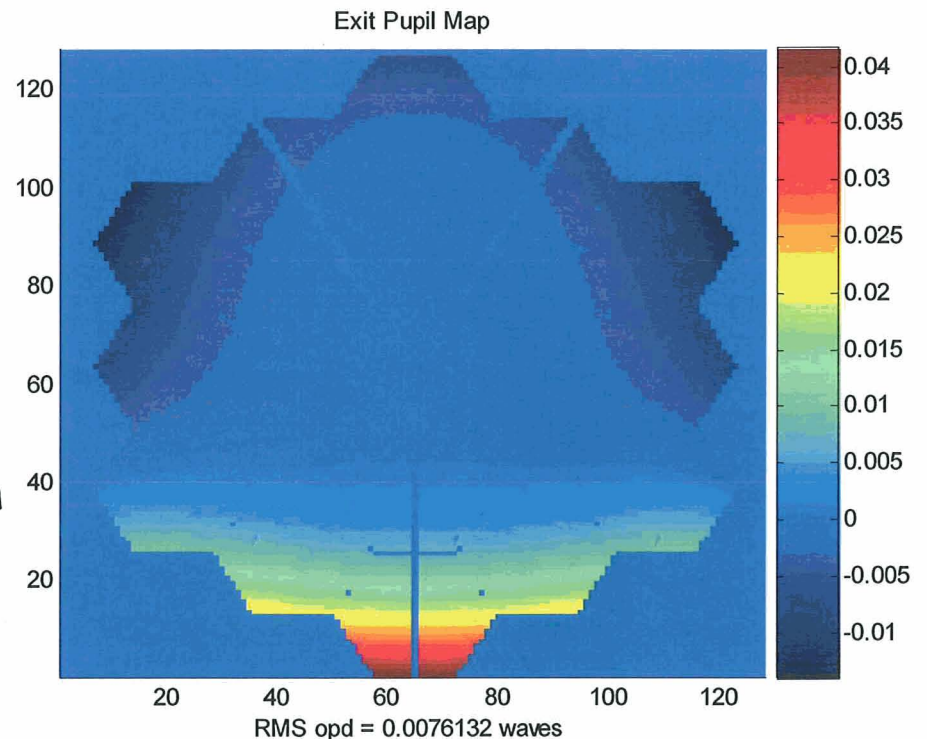
- ◆ TGR of square array, power of 2
- ◆ fit\_option, default = 0 (no fit), 1 = best fit tilt removed, 2 = best fit focus
- ◆ f,w,z = field number, wavelength number, zoom position (all default=1)

## OUTPUT:

- ◆ opd, mask, pin (pupil intensity), X Y real f/#, reference radius, X Y real focal length

## NOTES:

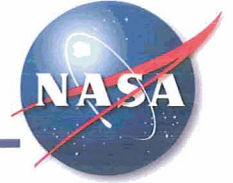
- ◆ Only gets data from first output in CodeV lens file. Field 1 and wave 1 should be redefined for data of interest!
- ◆ Since defaults are defined for most input parameters, use only as many output as you need- no need to output everything.
- ◆ If all outputs are omitted, **cvPMA** will return a figure of the OPD







# Tutorial: Point Spread Functions - cvPSF



```
[psf, grid_spacing] =  
  cvPSF (TGR, NRD_GRI, PGR, PRO);
```

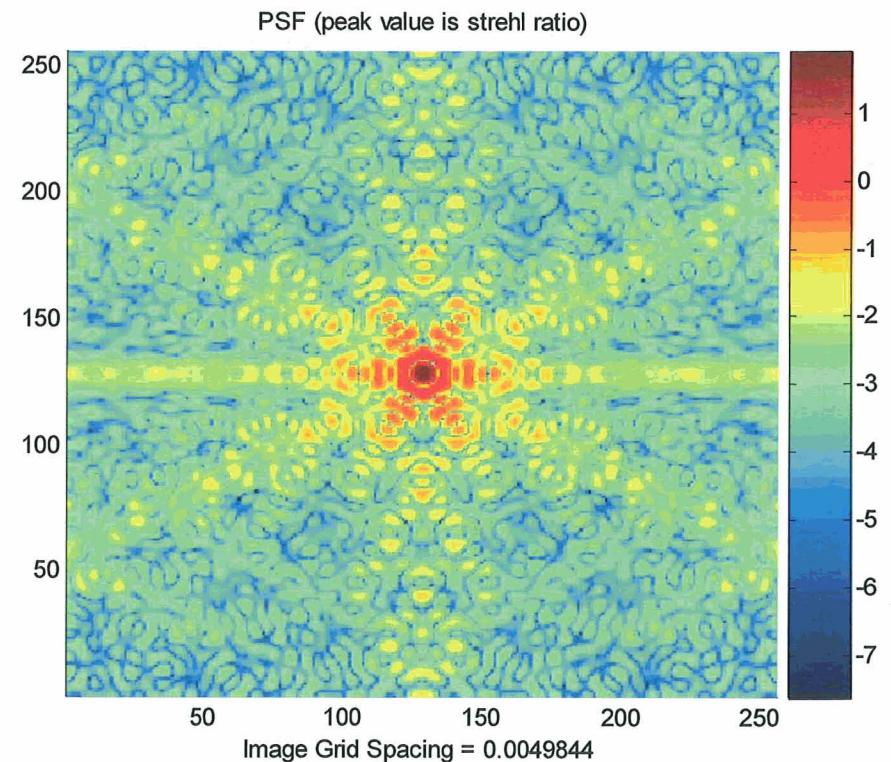
## INPUTS:

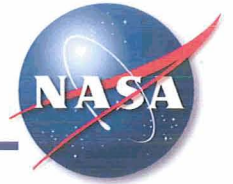
- TGR = Transform Grid Size
- NRD\_GRI = # of rays across pupil diameter
- PGR = Size of output array, default = TGR
- PRO = Use propagate equations for defocused image

## OUTPUT:

- psf: A matrix of the PSF data (size = PGR x PGR)
- grid\_spacing: Image plane pixel size in lens units

Command used to create plot: `cvPSF(256, 64);`





## Single rays

- ◆ **cvr** - allows the user to quickly look at reference ray parameters (position and direction cosines)

Typing `cvr('y', 1, 1, cvims, 1, 1)`

returns `-2.1356e-005`, the chief ray height for a field of 0,0 at the image plane.

- ◆ **cvRSI** - returns ray data for a single ray at every surface

## ◆ Grids of rays

- **cvRayGrid** - produces a grid of ray parameters for a specific surface. Any database object can be retrieved (i.e. position, direction, AOI, etc...)
- **cvRayTra** - uses a fast ray trace engine to calculate ray properties at the image plane. (Used in **cvspot**)



# Tutorial: cvSpot



```
[raygridx,raygridy] =  
  cvspot (fnum,wavnum,znum,gridsize,center);
```

## INPUTS:

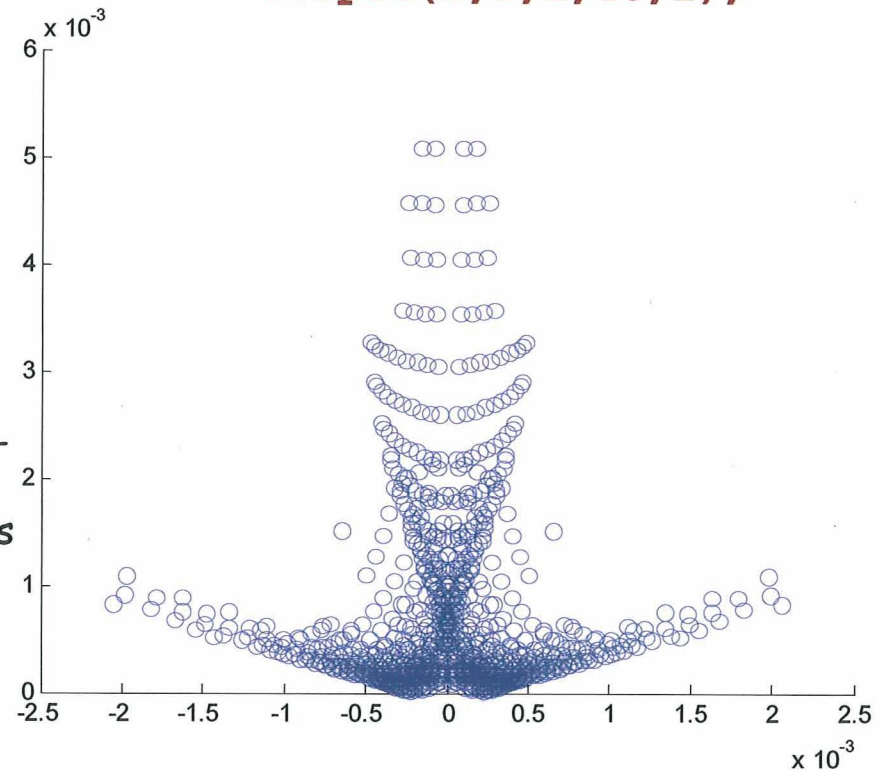
- fnum : can be a vector of field numbers (def = 1)
- wavnum, znum: (defaults =1)
- gridsize: number of rays across diameter to be traced.
- center: no centering, centered on zero, multiple fields

## OUTPUT:

- radgridx,y: the grid of data at the image plane (x- and y- coordinates).
- A scatter plot showing the spot diagram (output is in lens units)

Command used to create plot:

```
cvspot (1,1,1,40,1);
```





# Tutorial: Encircled Energy - cvENC



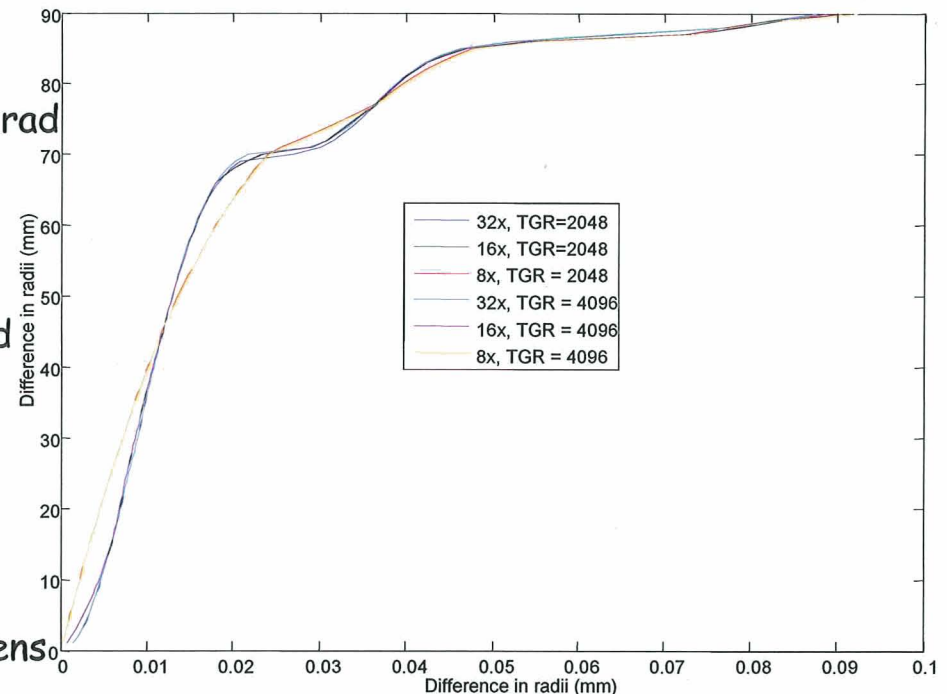
```
[percent, radii] =
  cvENC (PER, TGR, NRD_GRI, PGR, PRO, CEN) ;
```

## INPUTS

- PER = percent values to get encircled energy rad
- TGR = Transform Grid Size
- NRD\_GRI = # of rays across pupil diameter
- PGR = Size of output array, default = TGR
- PRO = Use propagate equations for defocused image
- CEN = Centered on Chief Ray or Centroid

## OUTPUT

- Percent = percentages at which radii were calculated
- radii = circle size encompassing % energy in lens units

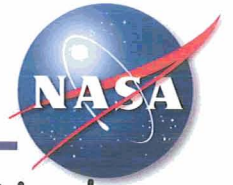


Data created with a simple "for" loop:

```
for ii=1:6
  [radii(:,ii),percent] =
  cvENC(percent,TGR(ii),NRD(ii),TGR(ii));
end
```

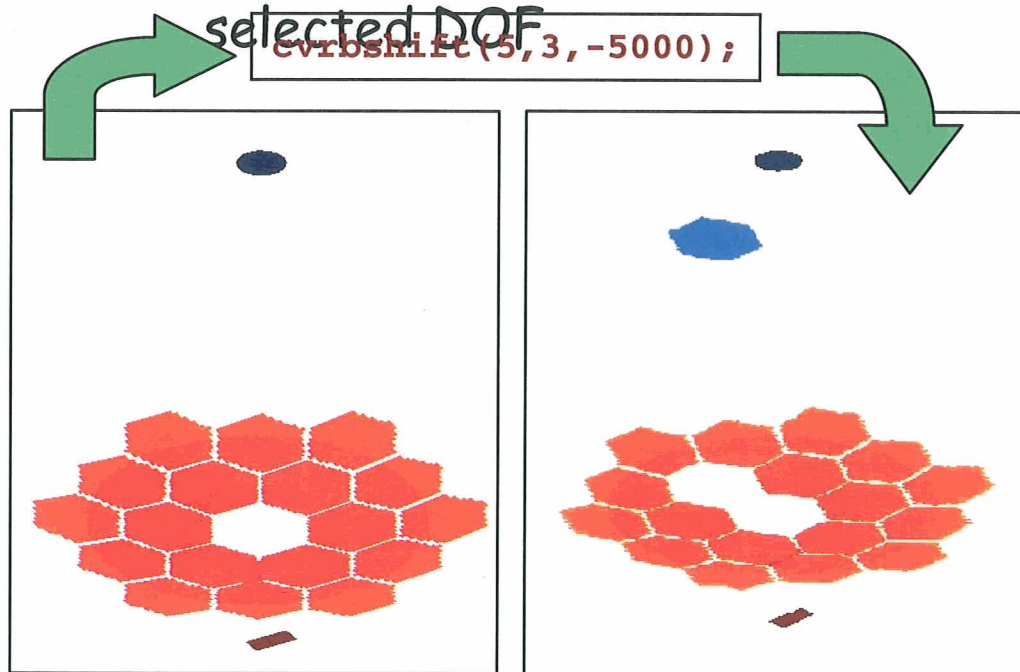


## Tutorial: Sensitivity Analysis



### ◆ Decentering / Tilting Lens Elements:

- **cvshift** - shifts the selected surface in 6 DOF at once.
- **cvrbsft** - shifts the selected surface in 1 selected DOF

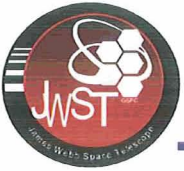


Shift surface 5 by -5000mm along the z-axis

### ◆ Sensitivities / Linear Optical Model:

- **cvLOM** - This function outputs the L, W, and C matrices needed for integrated modeling activities. cvlom is a specific form of cvsens, with the additional ability to take data with respect to a stationary exit pupil.

**cvsens** - computes the sensitivity of the selected surfaces to 6DOF. The output is a stack of matrices. Using a **zern\_fit** routine, one can easily access a simple numeric sensitivity output.



## Getting lens data from CodeV



```
[lens] = cvlensdata(refsurf);
```

- ◆ `cvlensdata` returns all of the construction parameters for the current lens file as a Matlab structure array

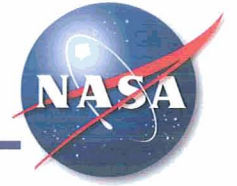
```
lens =  
    counts: [1x1 struct]  
 stopsurf: 2  
    dim: 2  
    fx: 0  
    fy: 0.001750000000000  
    wvl: [1x1 struct]  
    wl: [1x1 struct]  
    surf: [1x35 struct]  
 g_coords: [35x6 double]
```

```
lens.counts =  
    numz: 2  
 surface: 35  
    numf: 1  
    numw: 1
```

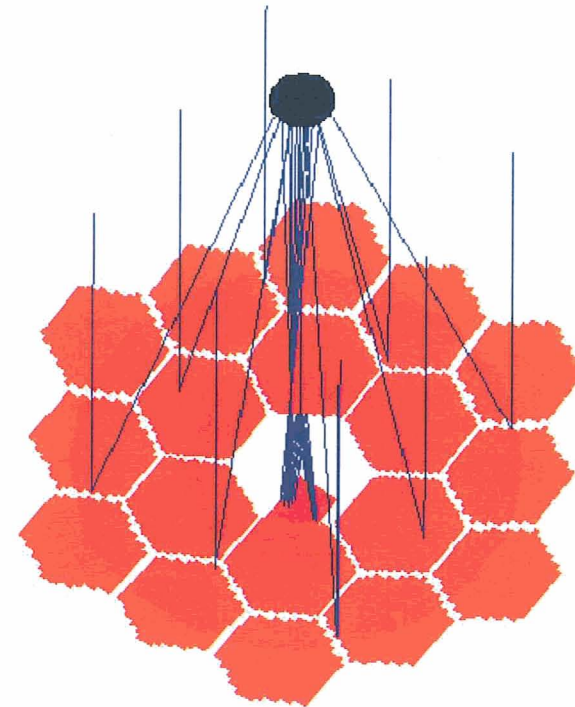
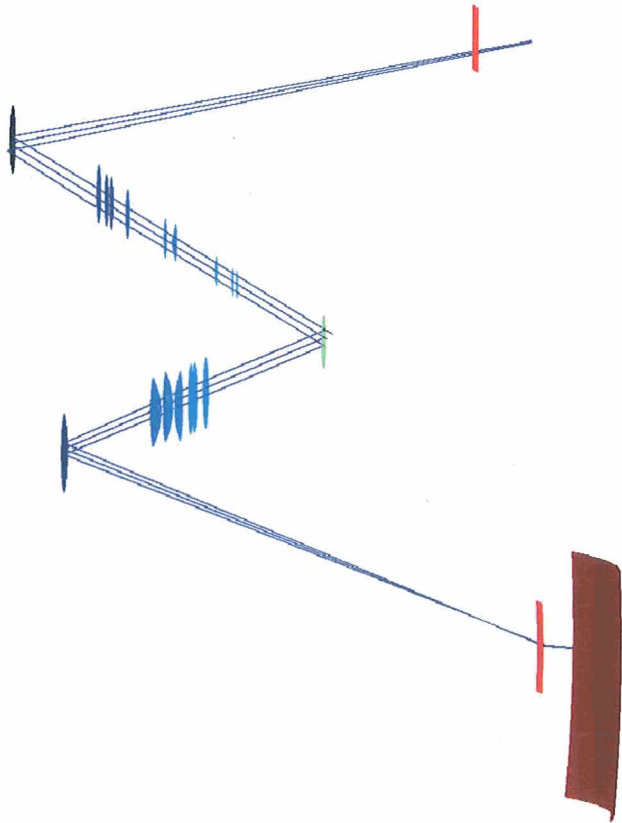
```
lens.surf =  
1x35 struct array with fields:  
    label  
    number  
    shape  
    rdx  
    rdy  
    k  
    ap
```



## Tutorial: Displaying a lens from CodeV in Matlab



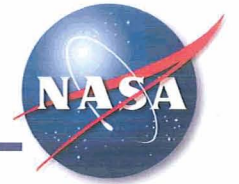
**`cvdraw(surfaces)`** ; Draws the selected surfaces in 3D and plot the reference rays for the system.





GO NAVY!!! BEAT ARMY!!!

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Any Questions?