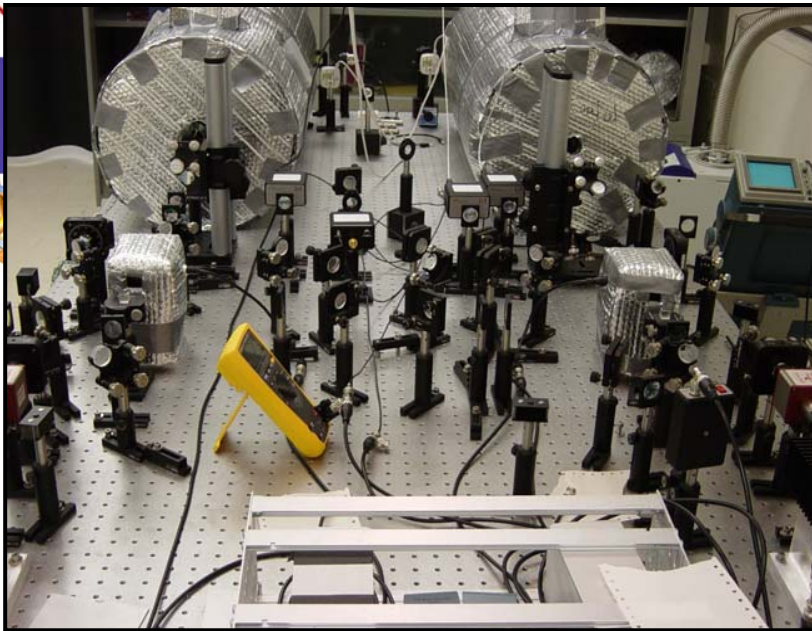


The "Lab"



From the Big Bang to Black Holes

*Information using Numerical
Waveforms*

J.I. Thorpe, S. McWilliams, B. Kelly, R. Fahey, K. Arnaud, J. Baker

7th International LISA Symposium

July 19th, 2008

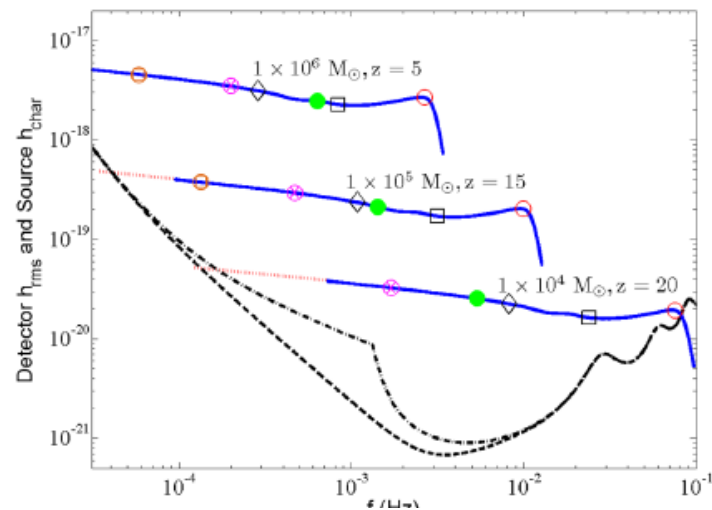
Barcelona, Spain



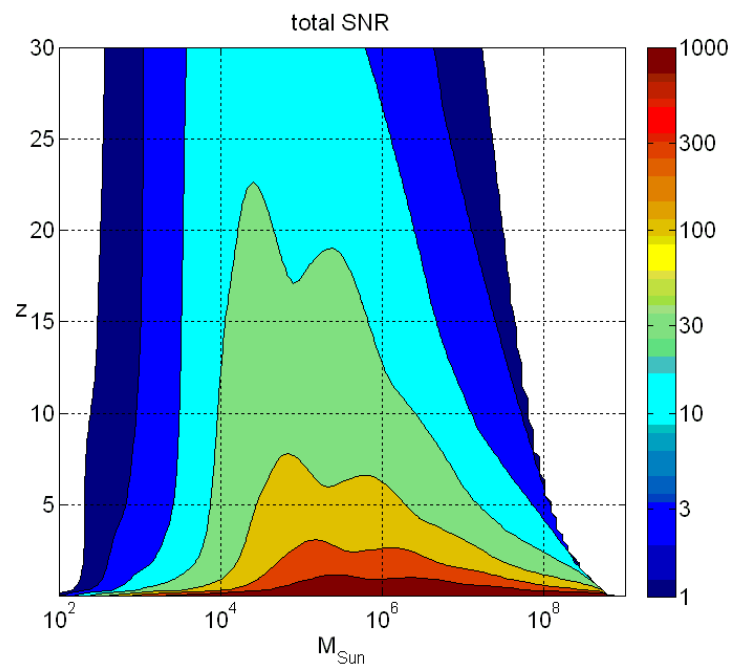
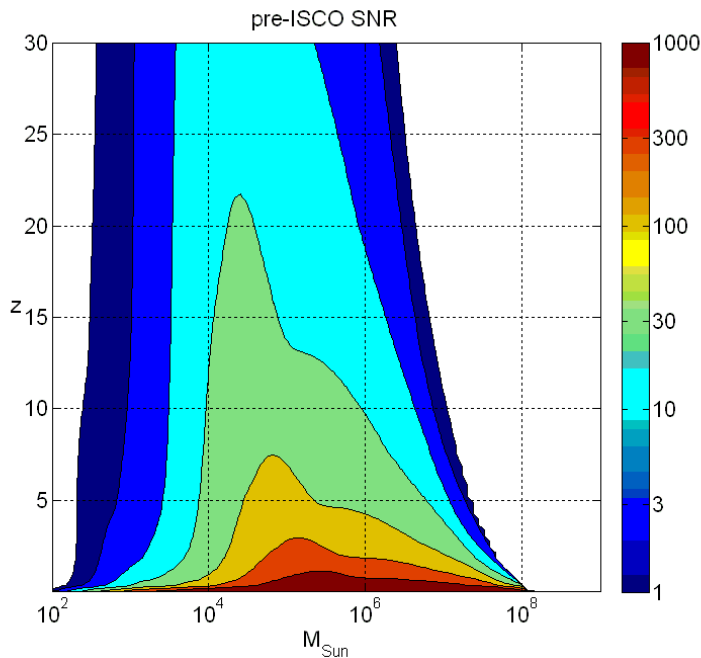
GSFC - JPL



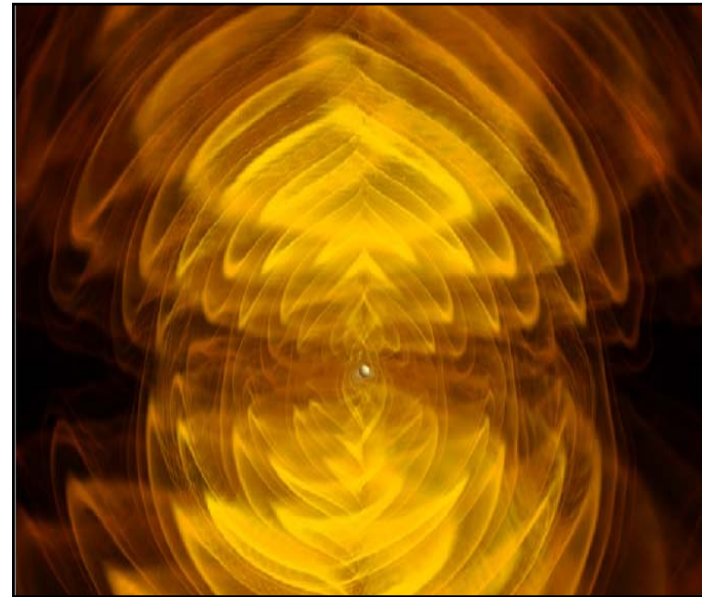
- Merger makes significant contribution to SNR
 - Depends on M_{tot}
 - Depends on distance/redshift



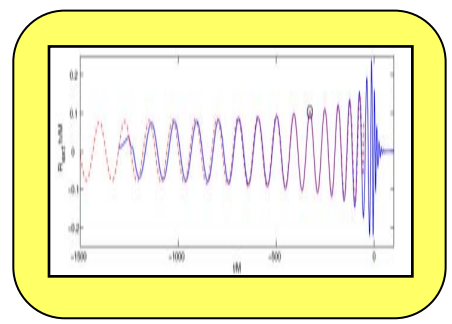
SNR contours for equal-mass, non-spinning case



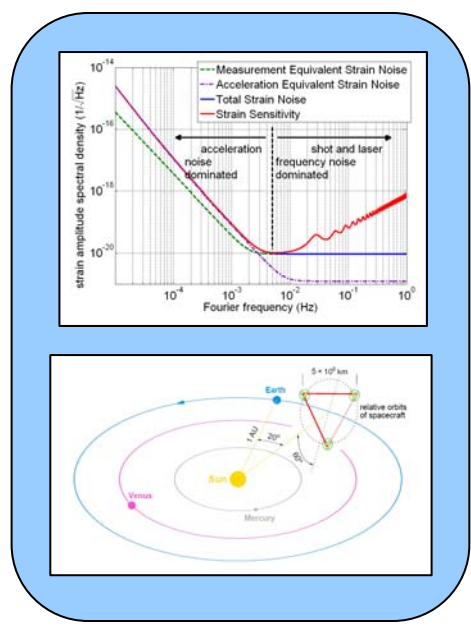
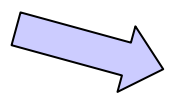
- 🌍 Traditional thought
 - Despite high SNR, merger is short-lived
 - Little orbital modulation
 - Small impact on parameter estimation
- 🌍 Possible paths for impact
 - Direct
 - Covariance



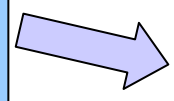
- Develop Parameter Estimation Code for Full Waveform
 - Complete inspiral, merger, and ringdown waveform
 - High-frequency instrumental response & noise
 - Fisher matrix approach



Waveform Model



Instrument Model



$$\Gamma_{ab} = \left(\frac{\partial h}{\partial \theta^a} \middle| \frac{\partial h}{\partial \theta^b} \right)$$

$$(a|b) = 4 \operatorname{Re} \int_0^\infty df \frac{\tilde{a}^*(f) \tilde{b}(f)}{S_n(f)}$$

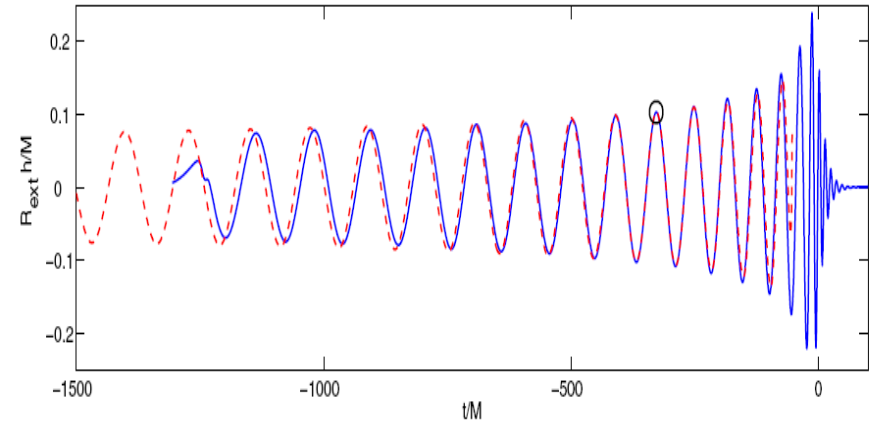
$$\Delta \theta^a \equiv \sqrt{\langle (\delta \theta^a)^2 \rangle} = \sqrt{\Sigma^{aa}}$$

Fisher Matrix

Waveform Model

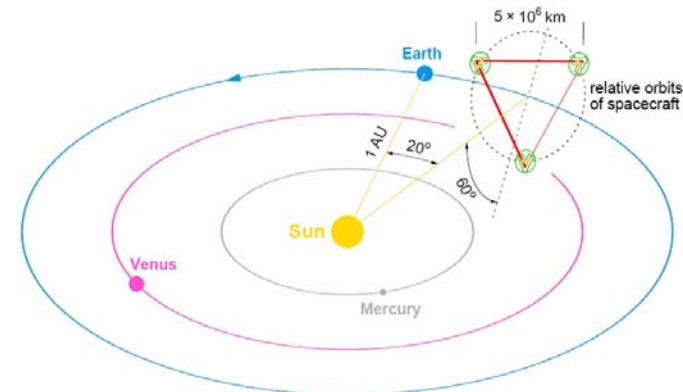
- Combined numerical & PN
- Equal mass, non-spinning
- Can be scaled to generate set of templates

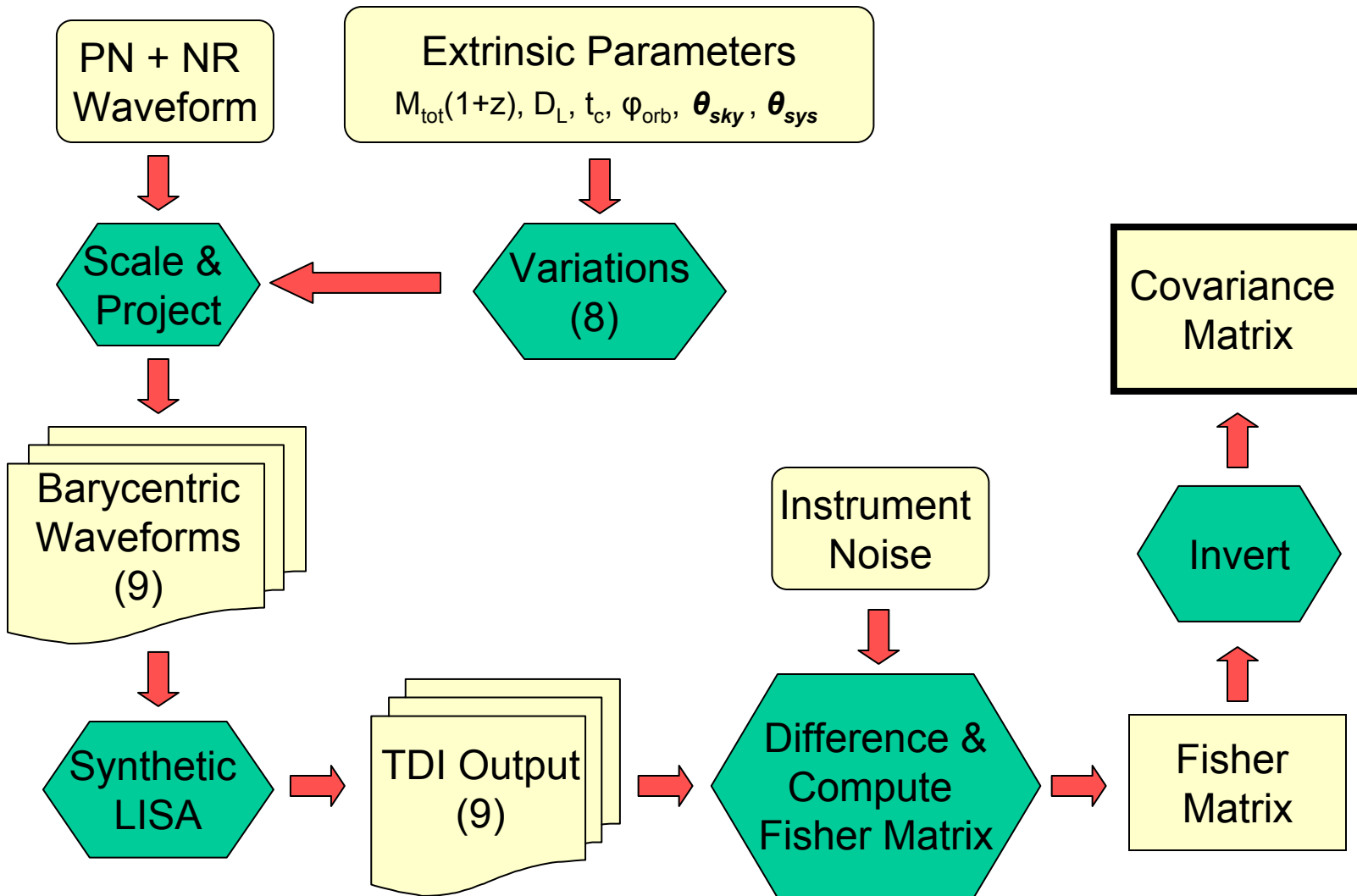
$$M_{\text{tot}}(1+z), D_L, t_c, \varphi_{\text{orb}}, \theta_{\text{sky}}, \theta_{\text{sys}}$$



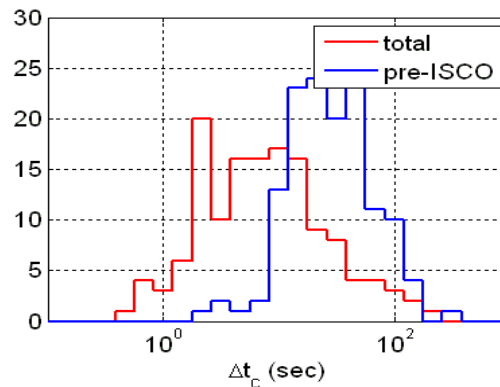
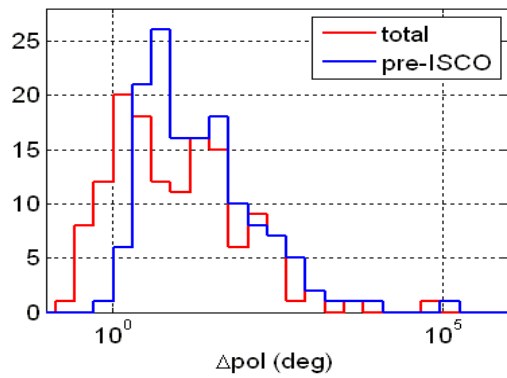
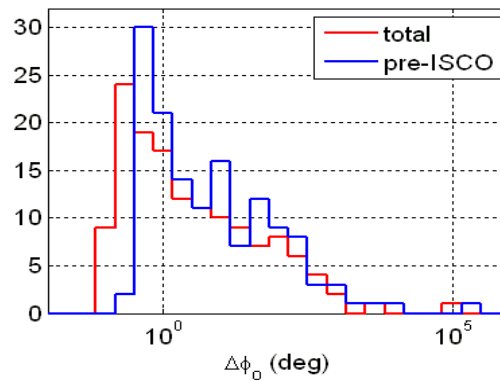
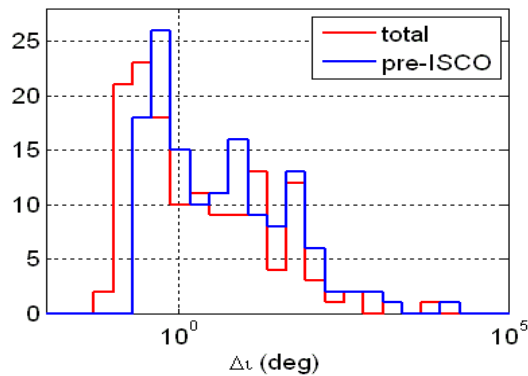
Instrument Modeling

- Synthetic LISA (Vallisneri)
- Realistic orbital model
- TDI outputs (pseudo A,E,T)
- Instrumental noise





- Two $2.5 \times 10^5 M_{\odot}$ non-spinning BH at $z = 1$ (~ 6 Gpc), 140 cases, pseudo A & E channels only





Results: $M_{tot}(1+z) \sim 10^6 M_{\odot}$

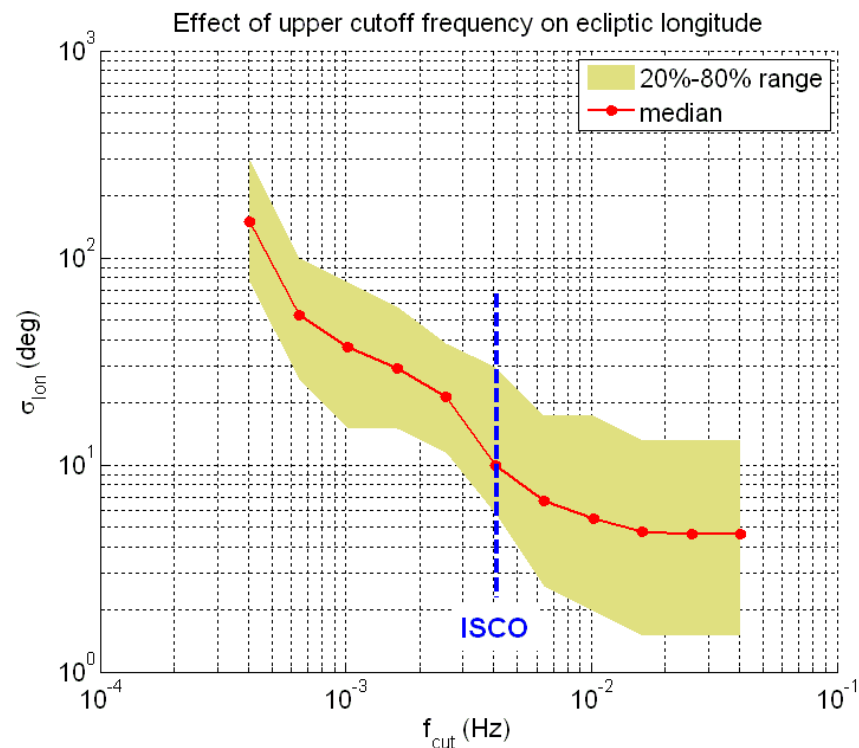
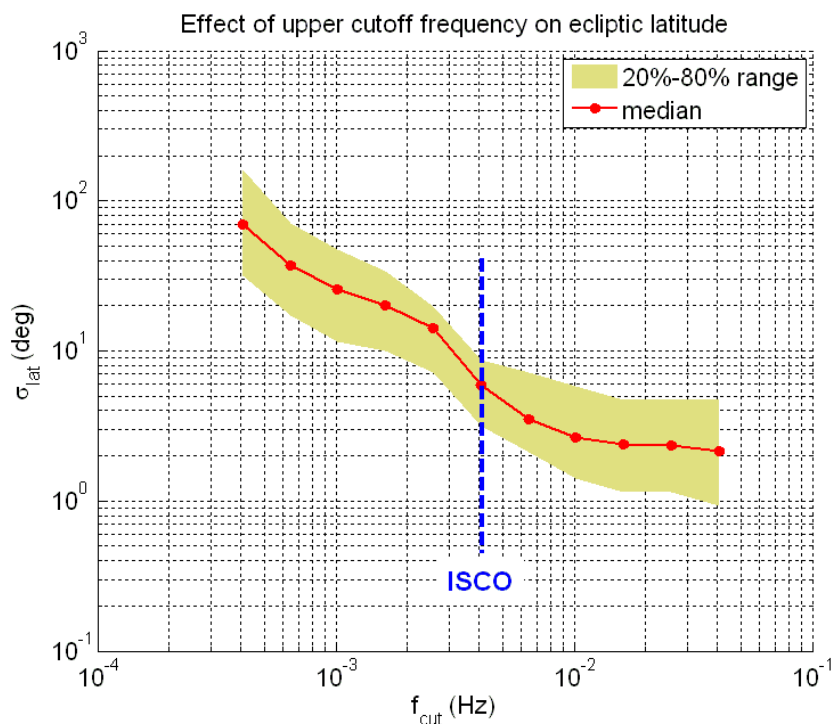


Beyond Einstein: From the Big Bang to Black Holes

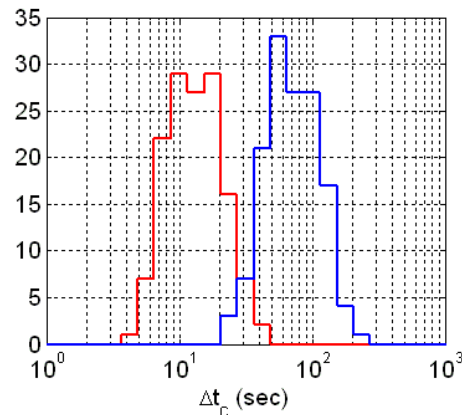
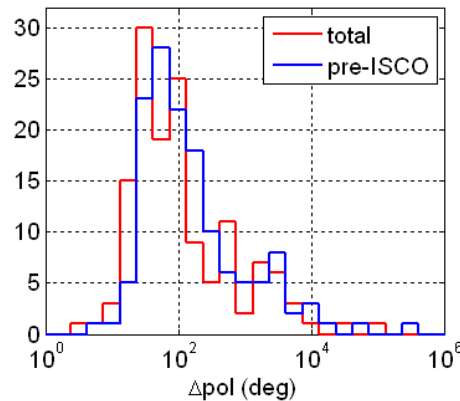
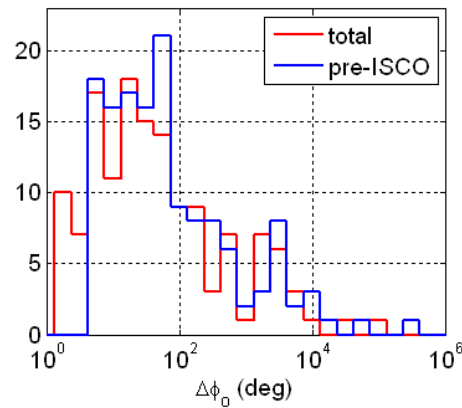
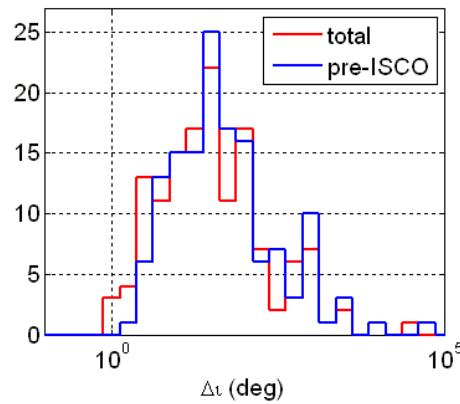
- Two $2.5 \times 10^5 M_{\odot}$ non-spinning BH at $z = 1$ (~ 6 Gpc), 140 cases, pseudo A & E channels only

	Total Mean	pre-ISCO Mean	Total Median	pre-ISCO Median
SNR	1.12E+03	8.35E+02	9.28E+02	6.98E+02
$\Delta \ln M$	2.35E-06	4.00E-06	1.66E-06	3.26E-06
$\Delta D/D$	3.87E-01	6.07E-01	4.48E-02	1.18E-01
$\Delta \text{lat (deg)}$	3.84E+00	7.19E+00	2.15E+00	5.28E+00
$\Delta \text{lon (deg)}$	1.77E+01	2.69E+01	4.64E+00	9.19E+00
$\Delta \text{iota (deg)}$	8.19E+01	1.37E+02	1.08E+00	3.12E+00
$\Delta \text{phio (deg)}$	6.07E+02	1.27E+03	1.59E+00	4.26E+00
$\Delta \text{pol (deg)}$	6.20E+02	1.29E+03	7.46E+00	1.59E+01
$\Delta \text{tc (sec)}$	1.71E+01	3.95E+01	6.87E+00	2.73E+01

- Two $2.5 \times 10^5 M_{\odot}$ non-spinning BH at $z = 1$ (~ 6 Gpc), 140 cases, pseudo A & E channels only



- Two $2.5 \times 10^6 M_{\odot}$ non-spinning BH at $z = 1$ (~ 6 Gpc), 140 cases, pseudo A & E channels only





Results: $M_{tot}(1+z) \sim 10^7 M_{\odot}$

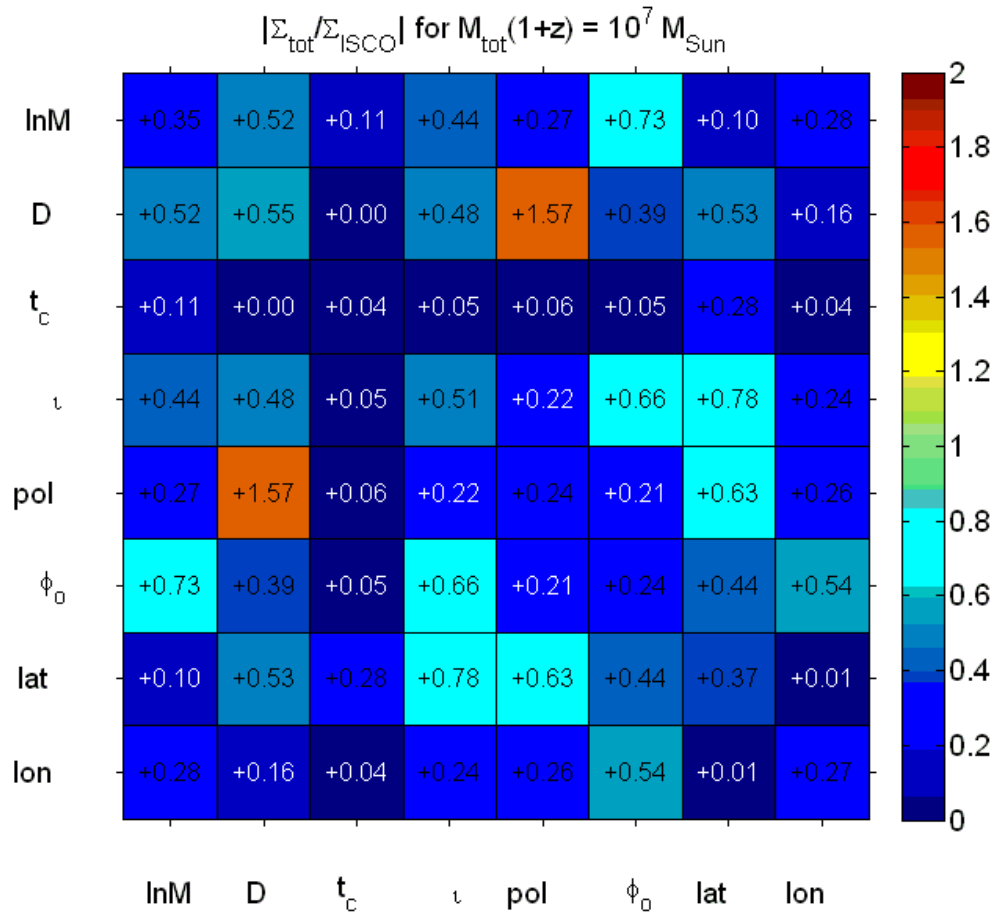


Beyond Einstein: From the Big Bang to Black Holes

Two $2.5 \times 10^6 M_{\odot}$ non-spinning BH at $z = 1$ (~ 6 Gpc), 140 cases, pseudo A & E channels only

	Total Mean	pre-ISCO Mean	Total Median	pre-ISCO Median
SNR	9.85E+01	2.68E+01	8.52E+01	2.29E+01
$\Delta \ln M$	1.55E-06	2.66E-06	1.01E-06	1.89E-06
$\Delta D/D$	2.17E+00	2.83E+00	8.04E-01	1.07E+00
$\Delta \text{lat (deg)}$	2.24E+01	3.35E+01	1.98E+01	2.76E+01
$\Delta \text{lon (deg)}$	5.84E+01	1.03E+02	2.20E+01	3.51E+01
$\Delta \text{iota (deg)}$	9.65E+00	2.52E+01	5.88E-01	1.40E+00
$\Delta \text{phio (deg)}$	1.58E+03	2.80E+03	3.33E+01	4.39E+01
$\Delta \text{pol (deg)}$	1.62E+03	2.87E+03	7.99E+01	1.11E+02
$\Delta \text{tc (sec)}$	1.47E+01	7.72E+01	1.31E+01	6.85E+01

- Changes in Covariance as merger is added
 - Most covariances decrease, merger is breaking degeneracies
 - Some covariances increase
 - Strong impact on coalescence time



- 🪐 Developed parameter estimation model integrating complete waveforms and improved instrumental models
- 🪐 Initial results for equal-mass non-spinning systems indicate moderate improvement in most parameters, significant improvement in some
- 🪐 Near-term improvement
 - Improved statistics
 - T-channel
 - Larger parameter space coverage
- 🪐 Combination with other results
 - Higher harmonics
 - Spin precession
 - Instrumental effects



Beyond Einstein: From the Big Bang to Black Holes

Backup Slides



GSFC - JPL



- Changes in Covariance as merger is added
 - Most covariances decrease, merger is breaking degeneracies
 - Some covariances increase

