

# Satellite-Scale Snow Water Equivalent Assimilation into a High-Resolution Land Surface Model

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## ***Abstract***

An ensemble Kalman filter (EnKF) is used in a suite of synthetic experiments to assimilate coarse-scale (25 km) snow water equivalent (SWE) observations (typical of satellite retrievals) into fine-scale (1 km) model simulations. Coarse-scale observations are assimilated directly using an observation operator for mapping between the coarse and fine scales or, alternatively, after disaggregation (re-gridding) to the fine-scale model resolution prior to data assimilation. In either case observations are assimilated either simultaneously or independently for each location. Results indicate that assimilating disaggregated fine-scale observations independently (method 1D-F1) is less efficient than assimilating a collection of neighboring disaggregated observations (method 3D-Fm). Direct assimilation of coarse-scale observations is superior to a priori disaggregation. Independent assimilation of individual coarse-scale observations (method 3D-C1) can bring the overall mean analyzed field close to the truth, but does not necessarily improve estimates of the fine-scale structure. There is a clear benefit to simultaneously assimilating multiple coarse-scale observations (method 3D-Cm) even as the entire domain is observed, indicating that underlying spatial error correlations can be exploited to improve SWE estimates. Method 3D-Cm avoids artificial transitions at the coarse observation pixel boundaries and can reduce the RMSE by 60% when compared to the open loop in this study.

*Key words:* assimilation, Kalman filter, downscaling, snow water equivalent, spatial correlation, multiscale

# 1. Introduction

Land surface data assimilation is mainly focused on surface temperature, soil moisture and snow. These variables interact with the atmosphere, which explains their direct impact on weather and climate predictions (Koster et al. 2004; Dirmeyer 2000). Land surface variables have a large spatial variability that cannot be captured by the existing operational observing systems alone. Land surface models could help in either downscaling the available coarse satellite observations or interpolating the scattered point observations. Snow differs from the other land surface states in its discontinuous local presence or absence, its cumulative character and often long temporal autocorrelation length (Slater and Clark 2006).

For snow, a variety of observations can be assimilated. Firstly, there are numerous snow water equivalent (SWE) or snowdepth point measurements, such as long term records from the U.S. Department of Agriculture's (USDA) Natural Resources Conservation Service (NRCS) SNOwpack TELEmetry (SNOTEL) network of snow pillow sites, the National Weather Service (NWS) Cooperative (COOP) weather stations network and short term, but spatially dense, ground measurements from intensive field campaigns like the Cold Land Processes Experiments (CLPX). Direct point snow data assimilation has been demonstrated by, e.g., Huang and Cressie (1996), Slater and Clark (2006) and Liston and Hiemstra (2008). Point data have been widely used as validation for satellite data assimilation.

Satellite observations of snow cover area or fraction and SWE-related quantities provide a second type of (partial) information about the snow state. As examples, the Advanced Very High Resolution Radiometer (AVHRR), Landsat Thematic Mapper (TM) and Moderate Resolution Imaging Spectroradiometer (MODIS) provide snow cover and albedo products

1 at fine resolutions ( $\leq 1$  km), but these visible or near infrared data cannot be collected  
2 at night or in cloudy conditions. Passive microwave sensor data do not suffer from these  
3 shortcomings, but have a much coarser resolution and come with large errors (Kelly et al.  
4 2003; Pulliainen et al. 1999). The Advanced Microwave Scanning Radiometer for the Earth  
5 Observing System (AMSR-E), the Scanning Multichannel Microwave Radiometer (SMMR)  
6 and the Spatial Sensor Microwave Imager (SSM/I) provide data at a 25 km resolution. Use  
7 of remote sensing data to update the SWE has focused on direct assimilation of MODIS snow  
8 cover area data (Rodell and Houser (2004); Andreadis and Lettenmaier (2006); Zaitchik and  
9 Rodell (2009)) or directly inverted (real or synthetic) SWE data from passive microwave  
10 sensors (Sun et al. 2004; Andreadis and Lettenmaier 2006; Dong et al. 2007).

11 Thirdly, different approaches have been developed to merge complementary sources of  
12 remotely sensed snow data (Durand et al. 2008), to combine remote sensing products with  
13 ground measurements (Pulliainen 2006; Tait et al. 2000), and to assimilate a priori merged  
14 snow products into models (Liston et al. 1999). The merger of different observational sources  
15 at multiple scales can also be achieved within the data assimilation framework, for exam-  
16 ple through the direct assimilation of multi-channel passive microwave brightness and near  
17 infrared data (Durand and Margulis 2006, 2007, 2008).

18 With few exceptions, notably Durand and Margulis (2007, 2008), snow data assimila-  
19 tion efforts have been limited to 1D filtering, i.e. the observations and model units were at  
20 or brought to the scale of the model grid and assimilated independently for each grid cell.  
21 The integration of coarse-scale observations for land surface state estimation in a dynami-  
22 cal framework is often complex, because satellite data and computational model units (grid  
23 cells, catchments units, etc.) typically differ in scale of support. With more advanced as-

1 simulation and scaling techniques, however, there is potential to extract information through  
2 downscaling satellite-scale observations into higher-resolution land model integrations.

3 Coarse observations can be re-gridded to the finer modeling units prior to data assimilation. Examples include retrieving SMMR SWE over catchment units (Dong et al. 2007),  
4 and reconstructing MODIS snow cover and AMSR-E SWE to match the modeling scale  
5 (Andreadis and Lettenmaier 2006). Additionally, many synthetic studies take this approach,  
6 such as Sun et al. (2004) for SWE assimilation. More sophisticated a priori disaggregation  
7 approaches based on statistical relationships have been used for soil moisture (Dubayah et al.  
8 1997; Kumar 1999; Crow and Wood 2002; Parada and Liang 2003b; Merlin et al. 2006). In  
9 any case, the disaggregated products could then be useful for dynamical assimilation into a  
10 finer-scale land model.  
11

12 Alternatively, the scale discrepancy can be dealt with more effectively and dynamically  
13 within the data assimilation framework by using a scaling observation operator. The dis-  
14 aggregation of coarse-scale information is then based on spatial error correlations that are  
15 modeled within the assimilation system. For example in land surface modeling, Reichle  
16 et al. (2001), Caparrini et al. (2004), Durand and Margulis (2007) and Zaitchik et al. (2008)  
17 used an averaging of a number of fine-scale forecasts to generate observation predictions at  
18 the coarse observation scale, respectively for assimilation of soil moisture (synthetic L-band  
19 passive microwave), land surface temperature (AVHRR, SSM/I, Geostationary Operations  
20 Environmental Satellite GOES), SWE (synthetical AMSR-E) and water storage (Gravity  
21 Recovery and Climate Experiment, GRACE).

22 In this paper, techniques for downscaling coarse-scale SWE observations to the under-  
23 lying fine-scale state variables within data assimilation is described and tested for the first

time within the National Aeronautics and Space Administration (NASA) Land Information System (Kumar et al. 2006; Peters-Lidard et al. 2007; Kumar et al. 2008a,b), version 5.0 (LIS5.0). One of the challenges in assimilating coarse observations over a large domain of fine-scale model units is in the ensemble estimation of the forecast error covariance matrix, where there is often a problem with spurious long range correlations. Therefore, ensemble filter techniques mostly include some type of localization (Keppenne and Rienecker 2002; Ott et al. 2004; Hunt et al. 2006; Reichle and Koster 2003), which will be discussed in this paper as well. As an alternative, multiscale Kalman filters could provide a way to replace the forecast error covariance matrix with a multiscale tree (Zhou et al. 2008; Parada and Liang 2004, 2008; Pan et al. 2009).

A number of ensemble Kalman filter (EnKF) approaches are explored that combine fine-scale land surface model simulations and synthetic satellite-scale observations. The focus will be on (i) the difference between 1D (point) and 3D (spatial) filtering and (ii) the study of different 3D filter approaches in view of subpixel snow variability estimation. Reichle and Koster (2003) demonstrated that the 3D EnKF was superior to 1D EnKF in a synthetic soil moisture estimation study. In that study, observations and modeling units corresponded in scale of support (catchments), and the advantage in the 3D filtering was primarily in data-sparse regions, because of the horizontal information propagation through the spatial forecast error structure (Hamill and Snyder 2000). This advantage of 3D filtering is also illustrated by Houser et al. (1998) and De Lannoy et al. (2009). In the synthetic experiment discussed here, however, observations are available for the entire domain and the focus is on the spatial disaggregation of coarse-scale observational information.

## 2. Experiment Setup

Different filters are tested in an identical twin experiment. Snow pack evolution is simulated at the fine-scale (1 km) as a reference “truth” for validation. A second integration with degraded forcing data mimics forecast modeling errors. Synthetic observations are generated from the truth integration at the coarse satellite-scale (25 km) by averaging the fine-scale snow pack and adding observation error. The coarse observation assimilation into the finer-scale degraded model simulations is then validated against the fine-scale truth.

The study domain includes the North Park area and part of the Rabbit Ears area in Colorado, USA, which have been intensively monitored during CLPX-I. The rectangular study area (left bottom corner: 40.255 N, -106.745 W; upper right corner: 40.995 N, -105.755 W) includes 75×100 fine-scale grid cells (0.01°~1 km) covering a central valley surrounded by mountains (Figure 1).

SWE is simulated with the land surface model of the National Centers for Environmental Prediction (NCEP)/Oregon State University/Air Force/Hydrologic Research Lab (Noah, version 2.7.1, Ek et al. (2003)) at a 0.01° (~1 km) resolution for the winter of 2002-2003, after a spin up starting in 2000. The NASA LIS5.0 (Kumar et al. 2008b) is used and adapted for this study. The forcings for the truth run are obtained from the North American Land Data Assimilation System (NLDAS, original resolution 1/8°, Cosgrove et al. (1999)). For all simulations in this paper a fixed temperature lapse rate of -6.5 K.km<sup>-1</sup> is used to disaggregate temperature data (Mitchell et al. 2004). The LSM parameters include a MODIS-based landmask and land cover parameters (1 km), the United States Geological Survey’s (USGS) GTOPO30 elevation map (~0.01°) to correct NLDAS forcings at the 1 km

scale, the Penn State University-USDA State Soil Geographic Database (STATSGO, 1/120°), a snow-free albedo product and a Noah-specific maximum snow albedo product from NCEP at 1°, National Oceanic and Atmospheric Administration (NOAA)/AVHRR-based greenness fraction (0.144°), and an NLDAS-based climatological bottom temperature (1/8°). The default Noah soil, vegetation and general parameter tables are used.

Synthetic SWE observations that mimic AMSR-E SWE retrievals are derived by aggregating the 1 km reference truth SWE simulation into twelve coarse-scale (25 km × 25 km) grid cells (3 rows, 4 columns, see Figure 1). Random noise is added to mimic observation errors and drawn from a normal distribution with a SWE-dependent standard deviation. Specifically, we use  $\sigma_{obs} = 5 + 0.095 \cdot \text{SWE} \text{ [mm]}$ . The perturbation standard deviation is minimally 5 mm, when no SWE is observed, and linearly increases to 100 mm for 1000 mm SWE. The average observation error magnitude is in the range reported for real passive microwave data, i.e. 5 to 45 mm in non-forested areas, increased by 5-10 mm over forests (Pulliainen et al. 1999; Derksen et al. 2003). Obviously, the assumed error function is only a crude approximation of actual errors. In practice, the type of snow emission model (Kelly et al. 2003; Pulliainen et al. 1999), local conditions and a number of physical details (including grain size, presence of liquid water, vegetation, Foster et al. (2005); Dong et al. (2005)) will strongly influence the observation error in satellite-based passive microwave SWE products. Furthermore, the AMSR-E products are known to have limited sensitivity to SWE for both very thin and deep snow packs (Dong et al. 2005). Nevertheless, we keep the whole range of synthetic observations, to illustrate the potential of the different assimilation techniques.

A second set of Noah model forecasts (open loop) are generated using forcings from the



1 Global Data Assimilation System (GDAS), which is originally at a Gaussian grid ( $T170 \sim 0.7^\circ$   
2 until 29 October 2002,  $T254 \sim 0.5^\circ$  until 31 May 2005). The forcings are of major impor-  
3 tance for the correct characterization of the SWE evolution (Slater et al. 2001; Mote et al.  
4 2003), and their errors will probably contribute most to SWE forecast errors in real experi-  
5 ments. Because the GDAS forcings result in heavily underestimated SWE simulations over  
6 the selected study domain, the precipitation is multiplied by a factor 3 to correct for the  
7 otherwise large forecast bias during the assimilation. This correction factor was determined  
8 through trial and error. More sophisticated bias correction methods such as matching of the  
9 cumulative distribution functions (Reichle and Koster 2004) or dynamic bias estimation (De  
10 Lannoy et al. 2007) could be used, but are beyond the scope of this paper, which focuses on  
11 the estimation of fine-scale spatial variability. The precipitation, air temperature, shortwave  
12 and longwave radiation, as well as the state variables are perturbed to generate ensembles  
13 for the assimilation, as will be discussed below.

14 The satellite-scale (25 km) synthetic observations are assimilated into the fine-scale  
15 (1 km) degraded model simulations from 30 September 2002 to 30 June 2003 (winter 2002-  
16 2003), either twice per month (every 15th and 30/28th of the month), or daily, to study the  
17 impact of the assimilation frequency on the analyses. To mimic the AMSR-E overpasses,  
18 the assimilation time is set to 8:00 am UTC (1:00 am local). The night overpass is selected  
19 (Durand and Margulis 2007), because mid-day SWE retrievals are typically affected by liquid  
20 (melted) water in the snowpack.

### 3. Data Assimilation

#### *a. Ensemble Kalman Filter*

The EnKF is a Monte-Carlo variant of the Kalman filter (Evensen 2003). The idea behind the EnKF is that a small ensemble of model trajectories captures the relevant parts of the forecast error structure. Each member of the ensemble experiences perturbed instances of the observed forcing fields (representing errors in the forcing data) and is also subject to randomly generated noise that is added to the model parameters and prognostic variables (representing errors in model physics and parameters). The error covariance matrices that are required for the filter update can then be diagnosed from the spread of the ensemble at the update time. The EnKF is flexible in its treatment of errors in model dynamics and parameters. It is also very suitable for modestly nonlinear problems and has become a popular choice for land data assimilation (Andreadis and Lettenmaier 2006; Durand and Margulis 2008; Kumar et al. 2008b; Pan and Wood 2006; Pauwels and De Lannoy 2006; Reichle et al. 2002a,b; Zhou et al. 2008).

The EnKF works sequentially by performing in turn a model forecast and a filter update. To reflect the uncertainty in the state forecasts, perturbations of the model forcings and the initial state estimate are applied at each time step to generate  $N$  ensemble forecasts  $\hat{\mathbf{x}}_i^{j-}$  ( $j = 1, \dots, N$ ) with the model  $\mathbf{f}$ :

$$\hat{\mathbf{x}}_i^{j-} = \mathbf{f}(\hat{\mathbf{x}}_{i-1}^{j+}, \mathbf{u}_i, \mathbf{w}_{i-1}^j), \quad (1)$$

where  $i$  denotes time,  $\hat{\mathbf{x}}_{i-1}^{j+}$  is the analysis from the previous time step (see below),  $\mathbf{u}_i$  represents the forcings, and  $\mathbf{w}_{i-1}^j$  denotes the perturbations to the  $j$ -th ensemble member. The

1 model forecast  $\hat{\mathbf{x}}_i^-$  is given by the ensemble mean. As a convention, we use bold upper case  
 2 symbols to refer to two-dimensional matrices, bold lower case symbols for one-dimensional  
 3 vectors and non-bold italic symbols for scalars.

4 When observations  $\mathbf{y}_i$  are available at time  $i$ , each ensemble member  $j$  is updated indi-  
 5 vidually to obtain the a posteriori state estimate (or analysis):

$$\hat{\mathbf{x}}_i^{j+} = \hat{\mathbf{x}}_i^{j-} + \mathbf{K}_i[\mathbf{y}_i^j - \mathbf{h}_i(\hat{\mathbf{x}}_i^{j-})]. \quad (2)$$

6 where  $\mathbf{y}_i^j$  is a suitably perturbed observation vector. The analysis  $\hat{\mathbf{x}}_i^+$  is again given by  
 7 the ensemble mean. The function  $\mathbf{h}_i$  is the observation operator that maps the state to  
 8 the observation space, and  $\mathbf{h}_i(\hat{\mathbf{x}}_i^{j-})$  denotes the model's observation prediction for ensemble  
 9 member  $j$ . The Kalman gain  $\mathbf{K}_i$  is identical for all ensemble members and determined  
 10 by the (sample) error covariance  $\text{Cov}[\hat{\mathbf{x}}_i^-, \mathbf{h}_i(\hat{\mathbf{x}}_i^-)]$  between the forecast and the observation  
 11 predictions, the (sample) error covariance  $\text{Cov}[\mathbf{h}_i(\hat{\mathbf{x}}_i^-), \mathbf{h}_i(\hat{\mathbf{x}}_i^-)]$  of the observation predictions,  
 12 and the observation error covariance  $\mathbf{R}_i$ :

$$\mathbf{K} = \text{Cov}[\hat{\mathbf{x}}_i^-, \mathbf{h}_i(\hat{\mathbf{x}}_i^-)] [\text{Cov}[\mathbf{h}_i(\hat{\mathbf{x}}_i^-), \mathbf{h}_i(\hat{\mathbf{x}}_i^-)] + \mathbf{R}_i]^{-1}. \quad (3)$$

13 For a linear (or linearized) observation operator  $\mathbf{H}_i$ , i.e.  $\mathbf{h}_i(\hat{\mathbf{x}}_i^-) \approx \mathbf{H}_i \hat{\mathbf{x}}_i^-$ , the sample error  
 14 covariances can be written as:

$$\text{Cov}[\hat{\mathbf{x}}_i^-, \mathbf{h}_i(\hat{\mathbf{x}}_i^-)] = \mathbf{P}_i^- \mathbf{H}_i^T \quad \text{and} \quad \text{Cov}[\mathbf{h}_i(\hat{\mathbf{x}}_i^-), \mathbf{h}_i(\hat{\mathbf{x}}_i^-)] = \mathbf{H}_i \mathbf{P}_i^- \mathbf{H}_i^T, \quad (4)$$

15 where  $\mathbf{P}_i^-$  is the forecast error covariance and superscript  $T$  denotes the matrix transpose.

16 In this paper we use the EnKF modules of the NASA Global Modeling and Assimilation  
 17 Office (GMAO, Reichle et al. (2009)) within LIS5.0 (after modifying the LIS5.0 implemen-

1 tation to enable the GMAO EnKF capability for horizontal error correlations, covariance  
 2 localization, and distributed filtering). In each fine-scale (1 km) grid cell, the Noah model  
 3 simulates a single snow layer with two prognostic variables, one for total column snow water  
 4 equivalent ( $swe$ ) and one for snow depth ( $snd$ ). The EnKF is used to update these two  
 5 model prognostic variables at the 1 km scale. Dropping the time index  $i$ , equation (2) and  
 6 (3) can then be re-written for a given fine-scale grid cell  $k$  as

$$\begin{pmatrix} swe_k^{j+} \\ snd_k^{j+} \end{pmatrix} = \begin{pmatrix} swe_k^{j-} \\ snd_k^{j-} \end{pmatrix} + \mathbf{K}[\mathbf{y}^j - \hat{\mathbf{y}}^{j-}] \quad (5)$$

$$\mathbf{K} = \text{Cov} \left[ \begin{pmatrix} swe_k^- \\ snd_k^- \end{pmatrix}, \hat{\mathbf{y}}^- \right] [\text{Cov}[\hat{\mathbf{y}}^-, \hat{\mathbf{y}}^-] + \mathbf{R}]^{-1} \quad (6)$$

7 with  $\mathbf{y}^j$  perturbed SWE observations that are used to update  $swe_k^{j-}$  and  $snd_k^{j-}$ ,  $\hat{\mathbf{y}}^{j-} \equiv \mathbf{h}(\hat{\mathbf{x}}^{j-})$   
 8 is short-hand for the corresponding model predictions of these observations, and  $\mathbf{R}$  denotes  
 9 the corresponding observation error covariance.

10 The first term in the expression for the Kalman gain includes the error correlation between  
 11 forecasted snow depth ( $snd^-$ ) and SWE ( $\hat{\mathbf{y}}^-$ ). This correlation information is the basis for  
 12 updating snow depth ( $snd^-$ ) in response to SWE observations ( $\mathbf{y}$ ). If  $m$  observations are  
 13 used to update SWE and snow depth for a given fine-scale grid cell, the Kalman gain  $\mathbf{K}$  in  
 14 equation (5) is a  $2 \times m$  matrix, and the observations  $\mathbf{y}$  and observation predictions  $\hat{\mathbf{y}}^-$  are  
 15 column vectors with  $m$  elements. In sections c and d, we will specify how exactly the coarse-  
 16 scale SWE observations are used (to update the fine-scale model prognostic variables  $swe_k^{j-}$   
 17 and  $snd_k^{j-}$ ) by relating  $\mathbf{y}^j$  and  $\hat{\mathbf{y}}^{j-}$  to observation and model variables. In any case, successful  
 18 downscaling of coarse-scale observational information relies on ancillary information at the  
 19 smaller scale, for example from micrometeorological or terrain information, and on horizontal

error correlations that enable horizontally distributed updates.

## *b. Ensemble Perturbation*

Random perturbation fields (normally distributed, zero mean, spatially correlated, temporally uncorrelated) are added hourly to the longwave radiation ( $LW$ , standard deviation,  $stdv = 10 \text{ W.m}^{-2}$ ) and near surface air temperature ( $AT$ ,  $stdv = 1 \text{ K}$ ) forcings, as well as to the forecasted SWE ( $stdv = 0.0025 \text{ m}$ ) and snowdepth ( $stdv = 0.01 \text{ m}$ ). The precipitation ( $P$ ) and shortwave radiation ( $SW$ ) are perturbed through multiplication with a random lognormally distributed field with  $stdv = 0.2$  and  $stdv = 0.5$ , respectively, for the standard deviation of the multiplication factor. Cross-correlation between the forcing perturbations is included ( $LW-SW$ : -0.3,  $SW-P$ : -0.1,  $SW-AT$ : 0.3,  $LW-P$ : 0.5,  $LW-AT$ : 0.6,  $P-AT$ : -0.1), as well as cross-correlation between  $swe^-$  and  $snd^-$  forecast state (0.9) perturbations. The magnitude of the perturbations is chosen in a realistic range and checked to allow near-optimal filter performance (see later).

Experiments are conducted with several different spatial correlation lengths ( $l_{corr}$ ) of the perturbation fields. Values of  $l_{corr}$  range between approximately 5 km and 100 km (i.e.,  $0.05^\circ$ ,  $0.1^\circ$ ,  $0.2^\circ$ ,  $0.3^\circ$ ,  $0.5^\circ$ ,  $0.7^\circ$ ,  $1^\circ$ ) and are always chosen to be identical for all different perturbation fields in a given experiment. When the spatial correlations are not included in the Kalman gain calculation, then the choice of  $l_{corr}$  does not matter and the spatial correlations are effectively forced to zero in the Kalman gain calculation. The number of ensembles is set to  $N = 12$ .

*c. Observation Pre-processing: A Priori Observation Disaggregation (1D-F1, 3D-Fm)*

Perhaps the simplest approach for the assimilation of coarse-scale observations into finer-scale model simulations is to create a surrogate observational grid with the same resolution as the fine-scale model grid. Here, the simplest possible disaggregation operation is used by assigning to each fine-scale (1 km) grid cell the value of the observation in the coarse (25 km) grid cell that contains the fine-scale grid cell.

An easy assimilation approach is then to apply the EnKF independently at each fine-scale grid cell  $k$ , using a single disaggregated and collocated observation  $swe_k^{obs}$  as illustrated in Figure 2. In this approach, spatial forecast error correlations are entirely disregarded. This filter is referred to as 1D-F1, a one-dimensional (1D) filter that uses exactly one fine-scale observation (F1) for the analysis increment computation at a given fine-scale grid cell. Formally, we compute the analysis increments  $\mathbf{K}[\mathbf{y}^j - \hat{\mathbf{y}}^{j-}]$  for the entire domain by looping through all fine-scale grid cells  $k$  and using (for each  $k$ )

$$\mathbf{y}^j = swe_k^{obs,j} \quad (7)$$

$$\hat{\mathbf{y}}^{j-} = swe_k^{j-} \quad (8)$$

as inputs to equation (5).

A straightforward extension of the 1D-F1 approach is to use several ( $m$ ) neighboring observations (after disaggregation to the fine scale) in the analysis increment computation for a given fine-scale grid cell  $k$  (Figure 2). This approach limits the edge-effect at the transition line between neighboring coarse observation pixels through the consideration of horizontal error correlations in the Kalman gain computation. We refer to this approach as

1 3D-Fm, because horizontal error correlations are taken into account in a “three-dimensional”  
 2 (3D) filtering approach (following the convention of Reichle and Koster (2003)) and because  
 3  $m$  (disaggregated) fine-scale (Fm) observations are used to update a given fine-scale grid  
 4 cell. Formally, we compute the analysis increments for the entire domain by looping through  
 5 all fine-scale grid cells  $k$  and using (for each  $k$ )

$$\mathbf{y}^j = \begin{pmatrix} swe_{k_1}^{obs,j} \\ swe_{k_2}^{obs,j} \\ \dots \\ swe_{k_m}^{obs,j} \end{pmatrix} \quad (9)$$

$$\hat{\mathbf{y}}^{j-} = \begin{pmatrix} swe_{k_1}^{j-} \\ swe_{k_2}^{j-} \\ \dots \\ swe_{k_m}^{j-} \end{pmatrix} \quad (10)$$

6 as inputs to equation (5), where  $k_1, k_2, \dots, k_m$  indicate the  $m$  (fine-scale) observations and  
 7 predictions that are included in the update of fine-scale grid cell  $k$ . The choice of the  
 8 observation selection area (influence radius around the analysis grid cell  $k$ ) determines the  
 9 observation dimension  $m$ . Because a larger influence radius also means a more complicated  
 10 inversion for the Kalman gain calculation, we limit the influence radius to 10 km when  
 11 1 km resolution observations are used (maximum  $m = 317$ ,  $m$  is smaller near the domain  
 12 boundaries). The observation error covariance matrix  $\mathbf{R}_i$  is assumed diagonal. Obviously,  
 13 as the influence radius approaches zero, the 3D-Fm approach reduces to the 1D-F1 method.

14 In our synthetic experiment, the dynamic observation error variance is known for the  
 15 coarse-scale observations, but not necessarily for the disaggregated observations. The ob-  
 16 servation error includes both the instrument error and the representativeness error of an  
 17 observation operator. While the first error term is typically assumed to be white Gaussian,

the second one can affect the actual observation error magnitude and correlations for the disaggregated observations. For simplicity, we do not consider spatially correlated observation errors and the same observation error magnitude is used for both fine- and coarse-scale observation assimilation.

*d. Observation Operator: Forecast Upscaling (3D-C1, 3D-Cm)*

It is more appealing to avoid the observation disaggregation prior to assimilation and to perform the downscaling within the filtering algorithm instead (Reichle et al. 2001; Durand and Margulis 2007; Zaitchik et al. 2008). Two disparate grids are then present in the multiscale filtering algorithm: the coarse observation grid and the fine-scale simulation grid, indicated by Greek and Roman letters, respectively. For example, a coarse-scale observation is denoted with  $swe_{\kappa}^{obs}$ .

A first option is to use only the single coarse-scale observation  $swe_{\kappa}^{obs}$  that covers the fine-scale grid cell  $k$  to update  $swe_k^{j-}$  and  $snd_k^{j-}$  with equation (5) as illustrated in Figure 2. In this case, the coarse-scale observation predictions are computed as the appropriate average over the corresponding fine-scale model predictions. Formally, we have

$$\mathbf{y}^j = swe_{\kappa}^{obs,j} \quad (11)$$

$$\hat{\mathbf{y}}^{j-} = 1/625 \sum_{l=1}^{625} swe_{k_l}^{j-} \quad (12)$$

where  $k_l$ ,  $l = 1, 2, \dots, 625$  indexes the  $25 \times 25$  fine-scale (1 km) grid cells contained within the coarse-scale (25 km) grid cell  $\kappa$ . This approach is referred to as the 3D-C1 EnKF here, because it uses one coarse-scale observation pixel to update each fine-scale state. The



1 error correlation between the fine-scale model state (to be updated) and the coarse-scale  
 2 observation prediction (first term of the Kalman gain in equation (5)) then downscales the  
 3 observational information into fine-scale updates.

4 The 3D-C1 technique can be expanded to include surrounding coarse observations in  
 5 the analysis, thereby using spatial forecast error correlations that overarch the boundaries  
 6 between the coarse observation pixel areas. Now, multiple ( $m$ ) coarse-scale observations are  
 7 used. Formally, equation (5) will be used with

$$\mathbf{y}^j = \begin{pmatrix} swe_{\kappa_1}^{obs,j} \\ swe_{\kappa_2}^{obs,j} \\ \dots \\ swe_{\kappa_m}^{obs,j} \end{pmatrix} \quad (13)$$

$$\hat{\mathbf{y}}^{j-} = \begin{pmatrix} 1/625 \sum_{l=1}^{625} swe_{k_{l,1}}^{j-} \\ 1/625 \sum_{l=1}^{625} swe_{k_{l,2}}^{j-} \\ \dots \\ 1/625 \sum_{l=1}^{625} swe_{k_{l,m}}^{j-} \end{pmatrix} \quad (14)$$

8 where  $\kappa_m$ ,  $m = 1, 2, \dots, 12$  denotes a coarse-scale grid cell in our study domain and  $k_{l,m}$ ,  
 9  $l = 1, 2, \dots, 625$  indexes the fine-scale grid cells contained within  $\kappa_m$ . This technique is  
 10 referred to as 3D-Cm, involving multiple coarse-scale observations and in a 3D filtering  
 11 approach. As in Reichle and Koster (2003) we suppress spurious long range correlations  
 12 in the forecast ensemble through (element-wise) Hadamard multiplication of the sample  
 13 error covariance terms in equation (5) with a distance-dependent and compactly-supported  
 14 function. Specifically, we use the 5th-order polynomial equation (4.10) of Gaspari and Cohn  
 15 (1999) and set the compact support scale equal to 2.5 times the spatial correlation length  
 16  $l_{corr}$  that is used for the ensemble spatial perturbation fields (i.e.  $2.5^\circ$ ,  $1.75^\circ$ ,  $1.25^\circ$ ,  $0.75^\circ$ ,  
 17  $0.5^\circ$  and  $0.25^\circ$ ). In practice, the localization means that observations beyond a given distance

1 from a specific fine-scale grid cell cannot affect the analysis at that grid cell.

2 Equation (5) is written for a single fine-scale grid cell  $k$  but involves knowing a potentially  
3 large number of fine-scale model states for the observation prediction ( $\hat{\mathbf{y}}^j$ ) computation.  
4 For implementation on a parallel architecture, parallelizing the loop through  $k$  is thus not  
5 necessarily the best strategy. Here, we implemented the 3D-C1 and 3D-Cm update methods  
6 by looping through coarse-scale grid cells and simultaneously updating all 625 fine-scale grid  
7 cells within a given coarse grid cell.

## 8 4. Results

9 The assimilation results are validated against the reference truth at the fine scale. Unless  
10 otherwise specified, the results are analyzed for assimilation twice a month (19 events be-  
11 tween 30 september 2002 and 30 june 2003) over the entire domain of 7500 fine-scale 1 km<sup>2</sup>  
12 simulation grid cells (covered by 12 coarse-scale observation pixels).

### 13 a. *SWE Open Loop Integration and Coarse Observations*

14 We first examine the synthetic coarse-scale observations and the (ensemble mean) open  
15 loop (no assimilation) integration. The top three rows of Figure 3 show five snapshots of  
16 the SWE field for the truth integration, the observations, and the open loop integration.  
17 Figure 4 shows the root-mean-square error (RMSE) of SWE (in time and in space) along  
18 with a time series of the domain average SWE. By construction the domain average SWE  
19 for the observations very closely approximates the true domain average SWE. Note also that

1 the open loop fields for the various choices of perturbation correlation lengths ( $l_{corr}$ ) are  
2 nearly identical, thus we do not distinguish between them for plotting. The observations  
3 show a smaller RMSE than the open loop integration. Because the observations are based  
4 on the truth simulation, there is no bias in the observations. For the open loop integrations,  
5 by contrast, some local or temporal bias (in addition to random errors) cannot be avoided.

6 Table 1 summarizes the space-time RMSE over the 19 analysis time steps. The RMSE  
7 for the observations is about 44 mm. As mentioned above, the ensemble mean open loop  
8 simulations for different  $l_{corr}$  perform similarly, with an RMSE of about 77 mm.

9 The difference between the truth and the open loop is caused by the difference in NLDAS  
10 and (scaled) GDAS forcings. Even though the GDAS forcings are at a very coarse resolu-  
11 tion, the modeled SWE distribution is quite reasonable. This is because the temperature  
12 correction (lapse rate of  $-6.5 \text{ K.km}^{-1}$ ) downscales the temperature allows a fine-scale dis-  
13 crimination between snowfall and rainfall. (But recall that a multiplication factor of 3 was  
14 used for the precipitation to bring the GDAS snowfall input to a more realistic level.) The  
15 peak in the spatial RMSE for both the open loop estimates and observations occurs during  
16 the melt period. This is caused by an increasing patchiness and variability in the actual  
17 snow amounts (as evidenced in the last column of Figure 3), which is impossible to reflect in  
18 the coarse-scale observation and also not fully captured by the open loop integration. The  
19 open loop shows almost a month delay in melting off all snow during early spring (Figure 4),  
20 because of an accumulated snowpack error.

21 Figure 5 shows a detail of the temporal SWE evolution for two individual fine-scale  
22 pixels. At the fine-scale, the coarse observations can differ significantly from the truth (e.g.  
23 during the snow accumulation in pixel (49,49)), because they do not represent the local

1 fine-scale SWE evolution. This fine-scale observation bias is important to keep in mind for  
2 1D filtering with disaggregated observations. Note also that the SWE forecast uncertainty  
3 ( $P_i^-$  as measured by the ensemble spread; not shown) increases from zero when no snow is  
4 present to a maximum value during deep snow packs, and then drops back to zero again  
5 after snowmelt.

## 6 *b. SWE Analyses*

### 7 1) SPACE-TIME AVERAGED STATISTICS

8 For most filter scenarios the space-time average RMSE (Table 1) is lower than for either  
9 the open loop (77 mm) or the observations (44 mm). This indicates that both the spatial  
10 mean field and the fine-scale variability are improved and supports the premise that assim-  
11 ilation products are better than either forecasts or observations alone (Reichle and Koster  
12 2005).

13 The 1D-F1 and 3D-Fm results show a similar RMSE (40-44 mm). Larger reductions  
14 in RMSE can be observed for the 3D-C1 and 3D-Cm with  $l_{corr} > 20$  km. For just 19  
15 assimilation events, the 3D-Cm approach achieves an RMSE decrease up to 60% when com-  
16 pared to the open loop and up to 25% when compared to the observations. For 3D-C1,  
17 the RMSE values are slightly higher (around 38 mm) than for 3D-Cm (around 34 mm) for  
18  $l_{corr} > 20$  km, but for shorter  $l_{corr}$  the 3D-Cm shows less assimilation impact. With daily  
19 assimilation, the RMSE over the same 19 time steps could not be significantly reduced. In  
20 the following subsections we analyze these findings in more detail.

## 2) TEMPORAL AND SPATIAL SWE PATTERNS

Figure 3 shows the spatial analysis SWE pattern for the different filter approaches. The time series of spatial RMSE for the assimilation analyses show a relatively smooth evolution (Figure 4). This means that the updates are accepted without significant drift in the spatial mean and that the snowpack has a considerable memory. During early snow accumulation, none of the filters, except the 3D-Fm (see dip in the spatial RMSE evolution in Figure 4 and spatial field in Figure 3 on 15 October 2003), is able to remove the excessive snow to match the true absence of snow. A very thin layer is kept, which is generated through the ensemble perturbation. During this initial period with extremely low snow amounts the update is negligible, because the model spread is still negligible, while the observation uncertainty has a set minimum value of 5 mm. During the ablation period, the open loop snow melt delay can be largely reduced (Figure 4) as a result of earlier snow pack corrections and the few instantaneous updates during the melt. The melt delay cannot be reduced for the filter scenarios relying excessively on the model predictions and hence resulting in analyses close to the open loop, like 3D-C1 with  $l_{corr} = 0$  km and 3D-Cm with very low  $l_{corr}$  and a small localization scale. With a limited spatial error correlation, the error covariance between a single fine-scale grid state and the observation prediction (based on a number of fine-scale grid elements) is low and the Kalman gain is limited.

In the spatial dimension, a clear coarse block structure in the analyses is evident for the 1D-F1, 3D-Fm and 3D-C1 filters (Figure 3). The 1D-F1 strongly forces all fine-scale state values to the observed coarse-scale SWE. Towards the melt season, the forecasts have a reduced spread and the spatial structure of the model simulations becomes apparent. For the

3D-Fm with an influence radius of 10 km, the coarse pixel boundaries are slightly smoothed, but they are still very obvious. By design, the 3D-C1 will always show the coarse pixel boundaries, because the state variables within a coarse area are all updated using only the single overlying coarse observation. However, when  $l\_corr=0$  km, the boundaries are less noticeable, because of a negligible observation impact (see above). By construction, the 3D-Cm does not show coarse observation transitions in the analysis result, since it uses a collection of coarse observations centered on each individual analysis point.

The spatial pattern of the RMSE (Figure 4) again shows the block patterns for some filters. Here, it is also clear that the spatial variability in the RMSE reduces with a longer  $l\_corr$  for 3D-C1 and 3D-Cm, and fine-scale locations with particular high RMSE values in the open loop integration are generally much improved through assimilation.

### 3) SPATIAL FORECAST ERROR CORRELATION AND LOCALIZATION

The results for the 3D-Cm and the 3D-C1 perform poorly when  $l\_corr < 20$  km (Table 1). No drastic change in performance is found with  $l\_corr > 20-30$  km. This threshold distance is roughly equal to the dimension of the coarse-scale observation, and also approximates the mean of the observed precipitation error correlation length (as measured by analyzing the spatial correlation length in the NLDAS minus GDAS precipitation difference fields), although the latter was found to vary significantly in time (not shown).

For 3D-C1, updates near the borders in the coarse observation pixel may be limited, when  $l\_corr$  is smaller than the pixel dimension. In the border areas, the average distance to all other fine-scale pixels within the coarse observation area is longer than for center pixels,

1 and the cross-correlation with fine-scale variables at longer distances is smaller, especially  
 2 when the distance reaches beyond  $l\_corr$ . This causes a limited covariance  $\text{Cov}[\hat{\mathbf{x}}_i^-, \mathbf{h}_i(\hat{\mathbf{x}}_i^-)]$   
 3 between the forecast error in a single fine-grid location and an observation prediction that  
 4 is based on a number of fine-scale grid forecasts. The sample observation prediction error  
 5 covariance  $\text{Cov}[\mathbf{h}_i(\hat{\mathbf{x}}_i^-), \mathbf{h}_i(\hat{\mathbf{x}}_i^-)]$  (which is independent of the fine-scale update location) is  
 6 also smaller for shorter  $l\_corr$ . Therefore, a smaller  $l\_corr$  results in a smaller gain factor,  
 7 mainly for the fine-scale analysis closer to the borders. For  $l\_corr > 20 - 30$  km, the forecast  
 8 error correlations are near-uniformly high over the each individual coarse satellite pixel area.  
 9 Increasing the  $l\_corr$  beyond the coarse pixel dimension adds little to the covariances in  
 10 the Kalman gain for each fine-scale update within the coarse pixel. This may explain why  
 11 Durand and Margulis (2008) also found a maximum assimilation efficiency for an exponential  
 12 variogram correlation length of 25 km when assimilating 25 km resolution SWE-related  
 13 observations, without significant degradation for longer  $l\_corr$ .

14 For the 3D-Cm approach, multiple coarse observations are included in each fine-scale  
 15 update and a covariance localization scale equal to  $2.5 \cdot l\_corr$  is applied to limit the impact  
 16 of more distant observations. For small  $l\_corr$  (relative to the domain size), this avoids  
 17 adverse effects caused by spurious long range correlations. However, the smaller  $l\_corr$  plus  
 18 the localization limit the magnitude of the covariances in the calculation of the Kalman gain.  
 19 This results in analyses close to the open loop simulation (compare the spatial and temporal  
 20 RMSE in Figure 4 and see the analysis RMSE in Table 1 for  $l\_corr = 5$  and 10 km). For  
 21 longer  $l\_corr$ , the corresponding longer covariance localization scale only marginally reduces  
 22 the sampled spatial correlations in this bounded study area. With longer  $l\_corr$ , the update  
 23 is increased locally and spread to other locations as well (smoothing), because of larger

1 forecast ensemble-based error covariances in the calculation of the Kalman gain.

2 In our study, a spatially isotropic error correlation and localization is imposed on the  
3 random fields, but the error field may not exactly represent the actual forecast error structure.  
4 Given the large elevation gradients and forecast errors introduced by different forcings, a  
5 complex error field can be expected. A good characterization of the error field and a region-  
6 dependent localization may further improve the results.

#### 7 4) FINE-SCALE SUBPIXEL SWE VARIABILITY

8 Figure 6 shows scatter plots of the fine-scale assimilation estimates (forecasts and anal-  
9 yses) within a single coarse grid cell (2,2) versus the true values just before and just after  
10 one particular assimilation update. The assimilated coarse-scale observation is also shown.  
11 These plots explain how each assimilation algorithm has a different effect on the spatial  
12 mean value and spread in the fine-scale SWE analyses. With the 1D-F1 filter, all fine-scale  
13 SWE values are drawn towards the same observed coarse-scale value. The spatial variability  
14 in the analysis result is much smaller than in the forecast. The same holds for the 3D-Fm  
15 (not shown), but the analysis has slightly more spatial variability, because of the interaction  
16 between neighboring innovations at each point. With the 3D-C1, the cloud of fine-scale pixel  
17 values is moved towards the observation, based on the coarse-scale difference between the  
18 observation and the forecast, but the forecasted spatial variability is generally maintained.  
19 The perhaps most interesting finding is that for the 3D-Cm, the spatial variability of the  
20 analysis is strongly improved and better mimics the truth, indicated by the analysis scatter  
21 points lined up around the 1-1 line. Apparently, including neighboring observations and a



reasonable forecast error structure can help to improve the fine-scale SWE estimation over using only the observation at the analysis point of interest.

We also examined the spatial correlation coefficient of SWE fields from assimilation versus the reference truth. The correlation coefficient measures the skill of the assimilation estimates in terms of fine-scale structure. Figure 7 shows the time series mean of the correlation coefficient for each coarse-scale pixel area and for a variety of filter scenarios. By construction, the spatial correlation coefficient between the synthetic observations and the truth is negligible, because the observations are purely randomly perturbed. The filters with disaggregated observations (1D-F1, 3D-Fm) degrade the spatial patterns with respect to the open loop. Overall, the 3D-C1 has only a limited positive effect and the spatial structure is only better than the open loop when  $l_{corr} > 20 - 30$  km. However, the 3D-Cm scenarios with larger  $l_{corr}$  show a great correlation improvement, mainly in the center coarse areas that benefit most from surrounding information and a relatively gentle topography (and hence relatively isotropic precipitation error correlation field). For coarse grid cell (2,2), for example, the correlation coefficient increases from a mean of 0.15 for the open loop to 0.64 for the 3D-Cm with  $l_{corr} = 20$  km. Only in two grid cells ((3,2) and (3,3)) the 3D-Cm does not improve the spatial variability, which is most likely because of the simple forecast error correlation structure in combination with a highly dynamic and spatially variable snow pack distribution in that area.

Figure 5 shows the temporal SWE evolutions at two individual fine-scale points within the coarse region (2,2). Most of the analyses better approach the truth than either the open loop or the observations, but none of the 1D and 3D filters are necessarily optimal at all locations within the coarse-scale pixel. At fine-scale grid cells where the disaggregated

coarse observations are consistently biased when compared to the fine-scale truth, 1D-F1 assimilation will degrade the analysis results compared to the open loop (e.g. during the snow accumulation period in pixel (49,49)). The 1D approach with disaggregated observation values has an observation bias problem at some individual locations. This problem can be overcome by assimilating the observations at the coarse-scale with a well defined error correlation structure and the proper scaling observation operator (as in 3D-C1 or 3D-Cm). For each update using multiple observations (3D-Fm, 3D-Cm), the SWE analysis for a given fine-scale pixel can exceed (or be less than) the forecast and the local observation. This is mainly obvious during the melt phase, when some patches still have a high SWE value with a high forecast uncertainty, while other patches have a very shallow snowpack with a very limited uncertainty. An innovation at the deeper snowpack in the influence area around an analysis point strongly impacts a location with a limited SWE (limited uncertainty and limited update by the collocated observation). In general, the 3D-Cm analyses are closest to the true reference simulation. For 3D-C1, most analyses are moved towards the same direction with respect to the forecast, but that is not necessarily the best solution at each fine-scale location. The 3D-C1 filter with small  $l_{corr}$  results in an analysis close to the open loop, because of the small Kalman gain, as discussed earlier.

## 5) FILTER DIAGNOSTICS

Figure 8 shows for different assimilation cases the frequency distribution of the normalized ensemble mean innovations, i.e. each innovation value  $[\mathbf{y}_i - \mathbf{H}_i \hat{\mathbf{x}}_i^-]_k$  (at the fine scale) or  $[\mathbf{y}_i - \mathbf{H}_i \hat{\mathbf{x}}_i^-]_\kappa$  (at the coarse scale) is normalized by the square root of its filter-estimated

1 standard deviation  $[\mathbf{H}_i \mathbf{P}_i^- \mathbf{H}_i^T + \mathbf{R}_i]_{kk}^{1/2}$  or  $[\mathbf{H}_i \mathbf{P}_i^- \mathbf{H}_i^T + \mathbf{R}_i]_{\kappa\kappa}^{1/2}$  (fine or coarse scale, resp.).  
 2 The normalized ensemble mean innovations should approximately obey a standard-normal  
 3 distribution (Gaussian with mean zero and variance one), if the model is linear and the filter  
 4 operates in accordance with its underlying assumptions. The innovation histograms provide  
 5 a rough indication of whether the model and observation error parameters are appropriately  
 6 chosen or not. The histograms show all innovations over space and time (for the 3D-C1 and  
 7 3D-Cm, there are only 12·19 innovations, while for the 1D-F1 and 3D-Fm there are 7500·19).  
 8 The figure shows a small bias in the innovation distributions. This is caused by local biases  
 9 in the forcings. When a daily assimilation is performed during the period from 30 September  
 10 2002 through 30 June 2003 (274 assimilation time steps instead of 19), the histograms (not  
 11 shown) exhibit similar features, but they show no bias.

12 For most filter scenarios, the standard deviation of the histograms is close to 1, indicating  
 13 that the filter parameters are near optimal. During early snow accumulation and late abla-  
 14 tion, the forecast uncertainty is bounded and it increases during deep snow packs. Similarly,  
 15 the observation error was increased with deeper snowpacks. This similar trend in observa-  
 16 tion and forecast errors keeps the filter close to its optimal operation during the entire snow  
 17 season. However, with real observations, it remains to be seen how the observation error  
 18 evolves in time and how it compares to the forecast uncertainty in each time period. Of all  
 19 filter configurations, 3D-C1 with  $l_{corr} = 0$  km suggests the least optimal filter operation,  
 20 with an excessive innovation spread. This is because the update is too limited to reduce the  
 21 forecast uncertainty during filtering. Observation error variance sensitivity tests (not shown)  
 22 indicate that scaling of the chosen dynamic observation error variance for the disaggregated  
 23 observations could slightly reduce the RMSE for the 3D-Fm filter (for the specified forecast

error). However, the corresponding innovation statistics show a suboptimal filter operation. This suggests that the better RMSE estimates may have resulted from compensation for imperfectly defined model errors.

Note that the innovation diagnostics are always available, including when satellite observations are assimilated, and can be used to identify obviously inadequate assimilation parameters (such as  $l_{corr} = 0$  km for 3D-C1) in the absence of reliable validation data.

## 5. Conclusions

The assimilation of coarse-scale (25 km) water equivalent (SWE) observations into fine-scale (1 km) model simulations is studied through a number of synthetic experiments with GMAO EnKF variants implemented in LIS5.0. Both a 1-dimensional (1D) or point filter and a variety of 3-dimensional (3D) or spatial filters are explored in a synthetical study. The truth is generated at the fine scale by forcing the Noah model with NLDAS data. Synthetic observations are generated through aggregation of these fine-scale simulations and addition of observation error. To assure a reasonable coefficient of variation in the observation error and in favor of the filter performance, the observation error variance is a function of the spatial mean SWE amounts. A degraded open loop model integration is simulated by forcing the Noah model with coarse GDAS data and imposing different types of spatially correlated random perturbations.

It is shown that coarse satellite products can improve both the SWE spatial mean and variability estimation over the open loop simulations, or in other words, fine-scale model simulations can downscale coarse satellite products to extract useful information for the

1 fine-scale SWE estimation. Because of the temporal evolution of the SWE amounts over the  
2 winter, the forecast uncertainty is small during early accumulation and later during the melt-  
3 ing season. Therefore, during the first days of simulated snow accumulation, the very small  
4 simulation spread and the fixed minimum observation error limits the assimilation impact.  
5 The predicted month delay in snow melt can be mostly removed through assimilation.

6 When the coarse observations are disaggregated prior to assimilation, a number of prob-  
7 lems arise. The 1D filter pushes the analyses to the coarse observation value, which improves  
8 the spatial mean SWE estimation but removes most of the subpixel spatial variability. Fur-  
9 thermore, significant fine-scale observation bias can be observed. If a number of fine-scale  
10 observations are used to update each analysis point in the 3D-Fm, the spatial variability is  
11 slightly better, but the computational cost becomes increasingly expensive in the inversion  
12 part of the Kalman gain calculation. For the actual assimilation of passive microwave prod-  
13 ucts (e.g. AMSR-E), disaggregating the coarse observations to finer scale observations may  
14 render the observations nearly useless, because the coarse products already have an unfav-  
15 orable signal-to-noise ratio. The correct scaling of the observation error for disaggregated  
16 observations needs further research to optimize the filter at the fine scale.

17 The best approach is to assimilate the coarse-scale observations directly with a 3D filter  
18 and a properly defined forecast error correlation structure. For all 3D filters using the coarse  
19 observations, it is found that the analyses are best when the spatial forecast error correlation  
20 length is equal to or larger than 20-30 km, which corresponds to the dimension of the coarse  
21 observation pixels, and also to the approximate correlation length of the precipitation error  
22 field. The results degrade significantly for shorter correlation lengths. When each fine-scale  
23 grid point is updated using the overlying coarse observation only (3D-C1), the spatial mean

1 SWE field can be improved over that of the observations or open loop simulations alone,  
2 and the spatial subpixel variability can be enhanced slightly. With additional inclusion of  
3 surrounding coarse observations (3D-Cm), there is a significant improvement in estimating  
4 the subpixel variability. Furthermore, artificial boundaries in the analysis field, caused by  
5 the boundaries in the coarse observations, are completely removed.

6 Even though the entire study domain is observed, there is a substantial value in including  
7 coarse observations from neighboring areas to enhance the fine-scale SWE structure estima-  
8 tion within the 3D-Cm filter. The SWE estimation can probably be further improved after  
9 optimization of the spatial error structure and covariance localization for the forecast error  
10 covariances in spatially complex terrain areas. In summary, the 3D-Cm filter is the method  
11 of choice to assimilate coarse observations, because it (i) improves the spatial mean SWE  
12 analysis, (ii) substantially enhances the subpixel SWE variability estimation, (iii) avoids ar-  
13 tificial transitions at the coarse observation boundaries and (iv) is computationally no more  
14 expensive than any of the other filter approaches.

## 15 6. Acknowledgments

16 The first author is a postdoctoral research fellow of the Research Foundation Flanders  
17 (FWO). Rolf Reichle was supported by NASA grant NNX08AH36G. We thank Sujay Ku-  
18 mar and the LIS development team at NASA/GSFC for providing the LIS5.0 beta version,  
19 Hiroko Kato from NASA/GSFC for providing the GDAS forcing data and Megan Larko  
20 from CREW/IGES for the computational support. This study is partially supported by the  
21 NOAA CPPA grant NA07OAR4310221.

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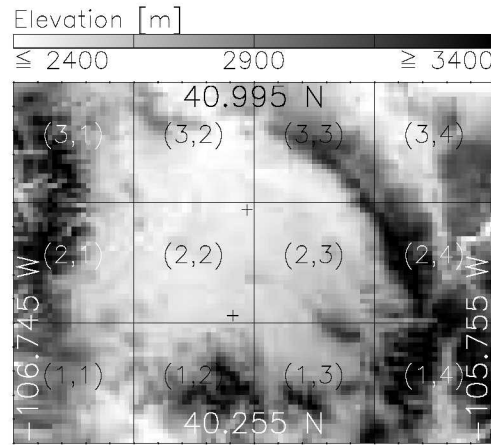


FIG. 1. Digital elevation model of the studied domain. The two plus-symbols mark locations for which individual time series are plotted in figure 5, with the upper one at fine-scale row and column (49,49) and the lower one at (27,46).

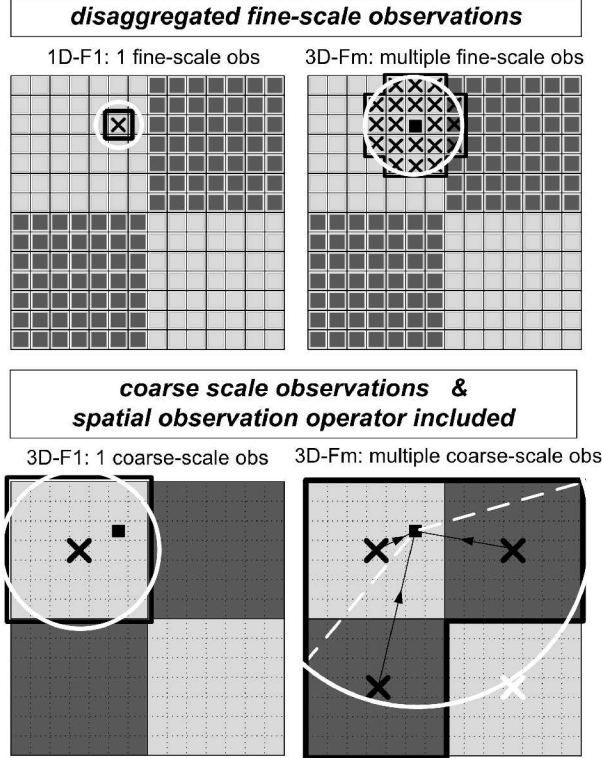


FIG. 2. Schematic of EnKF approaches illustrated for four coarse-scale pixels (gray shading), each containing  $7 \times 7$  fine-scale pixels. For the state update of a given fine-scale pixel (black filled square), we use observations within an influence area (white circle) and a corresponding set of fine-scale model forecasts (thick black line). For 1D-F1 and 3D-Fm fine-scale (disaggregated) observations are used (small crosses). For 3D-C1 and 3D-Cm, coarse-scale observations are used (center location is indicated by a larger crosses). In the 3D-Cm example, the observation at the white cross location is outside the influence area and not used to update the marked fine-scale state variable.

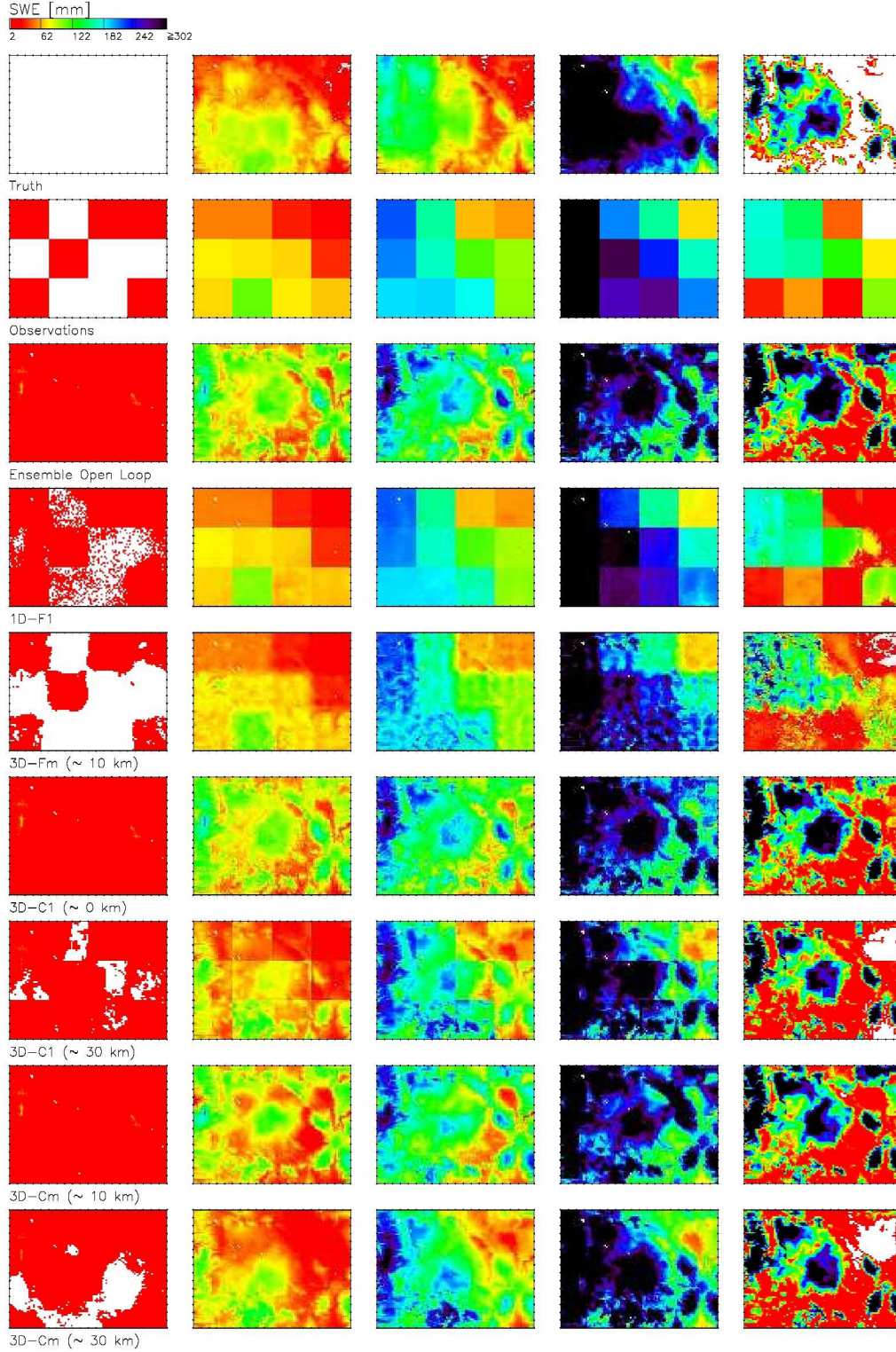


FIG. 3. Snapshots of SWE fields for (columns) 15 Oct 02, 30 Nov 02, 15 Jan 03, 28 Feb 03 and 15 Apr 03. (Top row) truth, (second row) synthetic observations, (third row) ensemble mean open loop forecasts and (rows 4-9) analyses obtained with several filter approaches. The spatial correlation length  $l_{corr}$  in the forecast perturbations is indicated in parentheses.

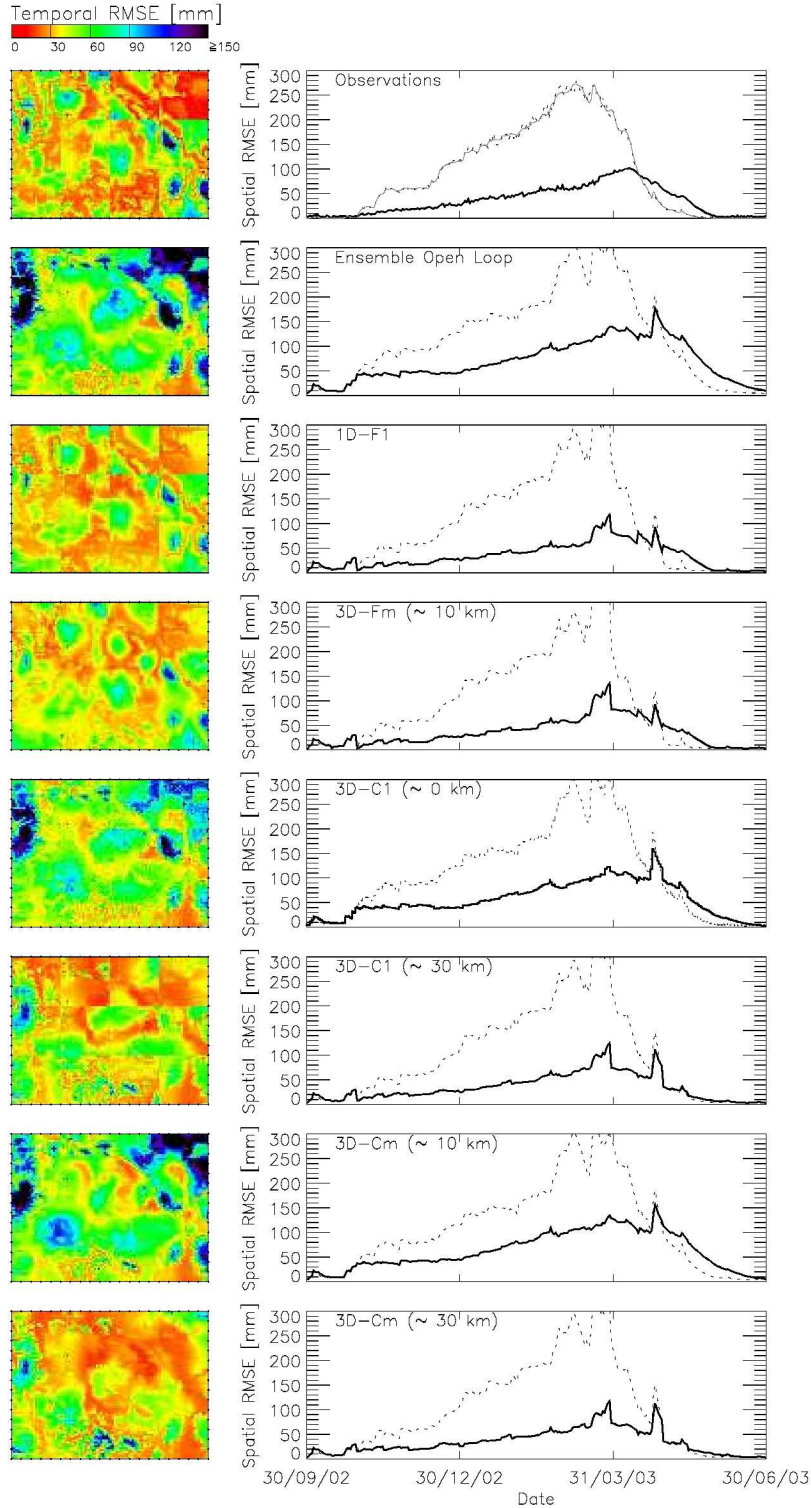


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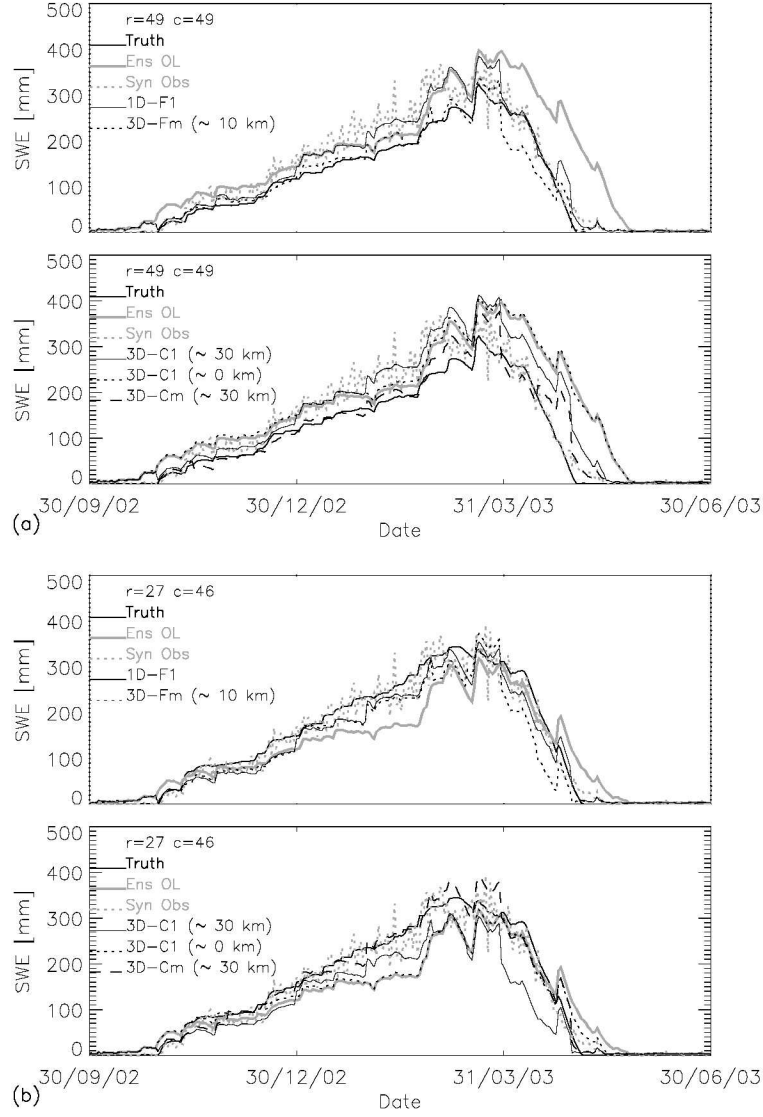


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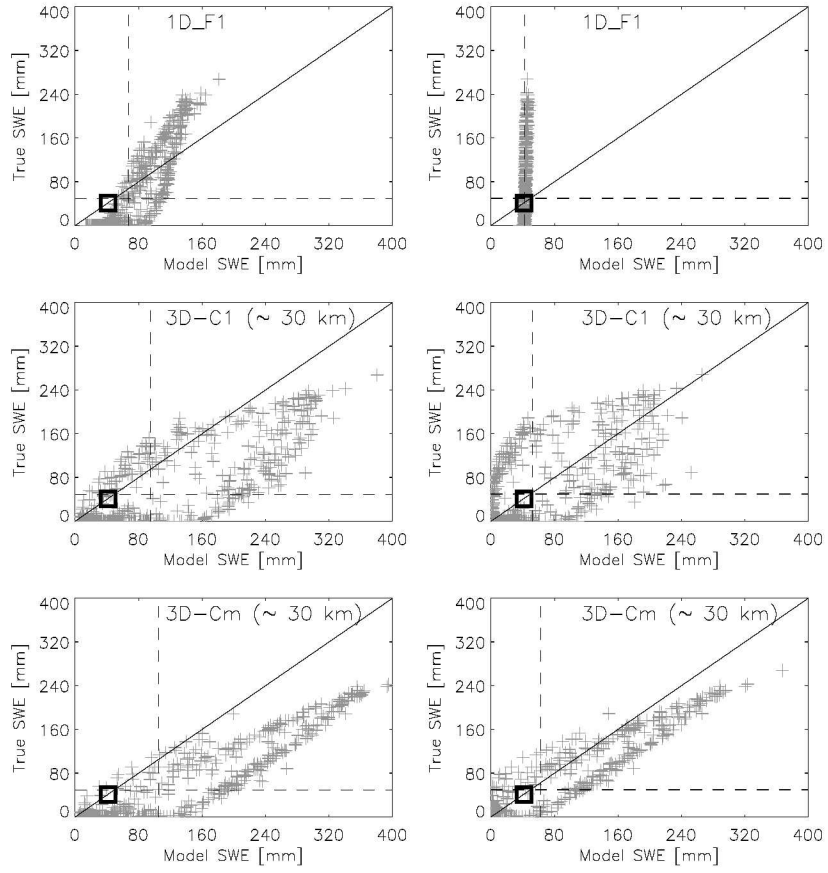


FIG. 6. Comparison of the fine-scale (left) forecasts and (right) analyses with the true SWE distributions within a single  $25 \times 25 \text{ km}^2$  coarse pixel (second row, second column of coarse grid) at 1 particular time step (30 Apr 2003; 8:00 am UTC). The dashed lines indicate the spatial means and the box shows the assimilated coarse-scale observation value.

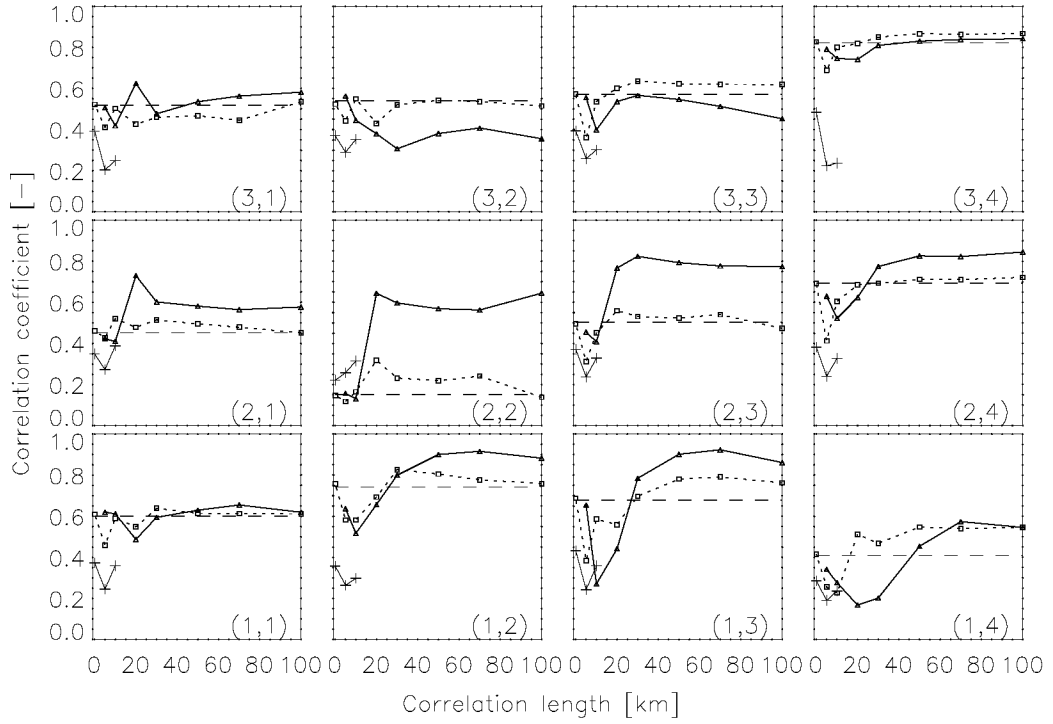


FIG. 7. Temporal mean of the spatial correlation coefficient between analyzed and true SWE as function of perturbations correlation length for (dashed) open loop, (black dotted) 3D-C1, and (black solid) 3D-Cm. Methods 1D-F1 ( $l_{corr} = 0$  km) and 3D-Fm ( $l_{corr} = 5$  and 10 km) are shown with + symbols. Each subplot corresponds to a coarse grid cell (Figure 1).

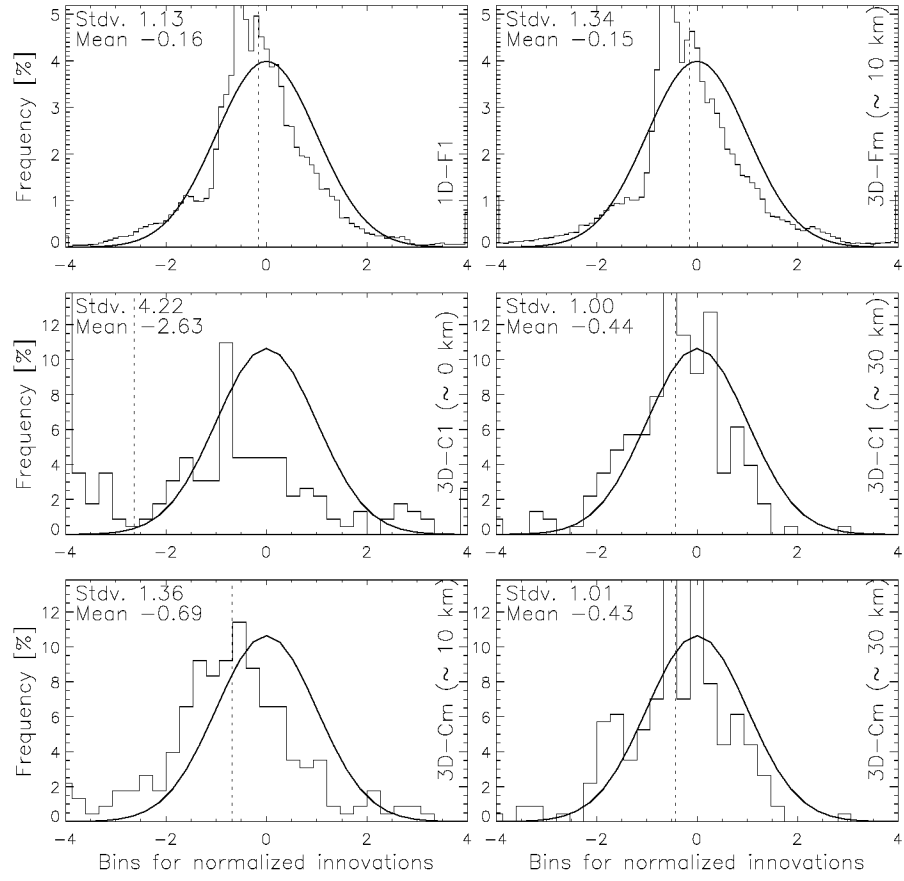


FIG. 8. Histograms of normalized innovations for different assimilation approaches.

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TABLE 1. RMSE of SWE [mm] over 19 update times and 7500 fine scale locations.  $l_{corr}$  is the spatial correlation length in the forecast perturbation fields.

| $l_{corr}$ | 0 km | 5 km | 10 km | 20 km | 30 km | 50 km | 70 km | 100 km |
|------------|------|------|-------|-------|-------|-------|-------|--------|
| Obs        | 44.4 | n/a  | n/a   | n/a   | n/a   | n/a   | n/a   | n/a    |
| Open loop  | 77.5 | 77.2 | 77.1  | 77.0  | 75.4  | 76.4  | 78.2  | 78.2   |
| 1D-F1      | 40.9 | 41.0 | 41.0  | 40.9  | 40.8  | 40.8  | 40.2  | 40.2   |
| 3D-Fm      | n/a  | 43.8 | 42.2  | n/a   | n/a   | n/a   | n/a   | n/a    |
| 3D-C1      | 61.7 | 53.1 | 42.0  | 39.0  | 37.1  | 38.0  | 39.2  | 37.7   |
| 3D-Cm      | n/a  | 77.2 | 70.4  | 41.7  | 35.4  | 33.7  | 34.5  | 32.3   |

***Title: Satellite-Scale Snow Water Equivalent Assimilation into a High-Resolution Land Surface Model***

**Journal of Hydrometeorology (American Meteorological Society)**

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*Abstract:*

An ensemble Kalman filter (EnKF) is used in a suite of synthetic experiments to assimilate coarse-scale (25 km) snow water equivalent (SWE) observations (typical of satellite retrievals) into fine-scale (1 km) model simulations. Coarse-scale observations are assimilated directly using an observation operator for mapping between the coarse and fine scales or, alternatively, after disaggregation (re-gridding) to the fine-scale model resolution prior to data assimilation. In either case observations are assimilated either simultaneously or independently for each location. Results indicate that assimilating disaggregated fine-scale observations independently (method 1D-F1) is less efficient than assimilating a collection of neighboring disaggregated observations (method 3D-Fm). Direct assimilation of coarse-scale observations is superior to a priori disaggregation. Independent assimilation of individual coarse-scale observations (method 3D-C1) can bring the overall mean analyzed field close to the truth, but does not necessarily improve estimates of the fine-scale structure. There is a clear benefit to simultaneously assimilating multiple coarse-scale observations (method 3D-Cm) even as the entire domain is observed, indicating that underlying spatial error correlations can be exploited to improve SWE estimates. Method 3D-Cm avoids artificial transitions at the coarse observation pixel boundaries and can reduce the RMSE by 60% when compared to the open loop in this study.

*Popular Summary:*

Satellites can indirectly measure the amount of snow that is present on the land surface. The measurement principle involves an analysis of the electromagnetic radiation in the microwave frequency range (centimeter wavelengths) that is naturally emitted by the land surface. Due to the characteristics of the satellite instruments and the measurement technique, the satellite observations can only provide information about the average snow amount within areas of about 25 km by 25 km. Such coarse-scale information is of limited value for water resources applications, particularly in mountainous areas. Complementary information on snow amounts can be obtained by using estimates of land surface characteristics (such as topography, vegetation, and soil information) along with precipitation (and other surface meteorological information) in a numerical model of land

surface processes. The model keeps track of the water and energy balance at the land surface and thereby also estimates snow amounts. Model estimates of snow can be obtained at relatively fine scales by setting up the numerical model on a 1 km by 1 km grid.

Both the satellite observations and the model estimates are affected by necessary simplifications in the respective computational algorithms and by errors in the corresponding input data. An optimization technique known as “data assimilation” can be used to merge the information from the satellite observations and the land surface model. The resulting estimates are superior to the estimates from the satellite or the land surface model alone. By design, the data assimilation system distributes the coarse-scale satellite information onto the finer-scale model grid, thereby making the satellite observations more useful for water resources applications.

In this paper, we investigate several options for implementing such a data assimilation system. One alternative is to disaggregate the coarse-scale satellite observations to the fine scale of the model prior to data assimilation or within the data assimilation system. Another alternative is for each model location to use only satellite observations that are local to that location or to also use observations from neighboring locations. We test these options for a 75 km by 100 km mountainous area in western Colorado, USA. We find that it is indeed possible to improve fine scale estimates of snow amounts through the assimilation of coarse-scale satellite observations. Our results indicate that it is best (i) to avoid disaggregating the satellite observations prior to data assimilation and (ii) to have satellite observations from the entire domain influence the assimilation estimate at a given location. The most sophisticated assimilation method avoids artificial boundaries around coarse-scale satellite observations and, for our study domain, reduces errors by 60% when compared to the estimates provided by the land surface model alone.