

# Development of a Multivariable Parametric Cost Analysis for Space-Based Telescopes



Although cost models must evolve as new technology becomes available and as program objectives are modified, the research and analysis done in this project have improved the foundation upon which cost models for future space telescopes will be based. Major results from my research include confirmation of the relationship` between OTA cost and diameter, and OTA Cost and OTA mass. As diameter increases, so does OTA Cost and OTA Mass, both increasing in similar ways. This shows that mass is still the most significant factor in cost and provides graphical evidence that we are taking the rights steps towards a cost estimate by separating the variables that influence mass and cost. Research on the variables that affect cost will continue in the future in an effort to create the most updated and accurate cost model for space telescopes possible.

Better Balance of Science and Budget

More

formation

and

Research

Over the past 400 years, the telescope has proven to be a valuable tool in helping humankind understand the Universe around us. The images and data produced by telescopes have revolutionized planetary, solar, stellar, and galactic astronomy and have inspired a wide range of people, from the child who dreams about the images seen on NASA websites to the most highly trained scientist. Like all scientific endeavors, astronomical research must operate within the constraints imposed by budget limitations. Hence the importance of understanding cost: to find the balance between the dreams of scientists and the restrictions of the available budget. By logically analyzing the data we have collected for over thirty different telescopes from more than 200 different sources, statistical methods, such as plotting regressions and residuals, can be used to determine what drives the cost of telescopes to build and use a cost model for space-based telescopes Previous cost models have focused their attention on ground-based

telescopes due to limited data for space telescopes and the larger number and longer history of ground-based astronomy. Due to the increased availability of cost data from recent space-telescope construction, we have been able to produce and begin testing a comprehensive cost model for space telescopes, with guidance from the cost models for ground-based telescopes. By separating the variables that effect cost such as diameter, mass, wavelength, density, data rate, and number of instruments, we advance the goal to better understand the cost drivers of space telescopes.. The use of sophisticated mathematical techniques to improve the accuracy of cost models has the potential to help society make informed decisions about proposed scientific projects. An improved knowledge of cost will allow scientists to get the maximum value returned for the money given and create a harmony between the visions of scientists and the reality of a budget.



# **Data Collection:**

 Interviews with project managers •NASA websites. Data falls into four categories: •Programmatic Primary Mirror Mission

•OTA Data

# **Methodology:**

•Using statistical methods and selected information from the database, regressions and residuals for variable analysis were created.

•Plots were analyzed for significance in affecting cost. •Over 20 variables were tested. •Variables tested thus far were chosen based upon how important they were thought to be, using logic to determine if their engineering significance would affect cost.

WBS CADRe

## **Conclusions:**

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		General Info
Confidence Level	90%	How likely th
		Measure of "n
S <sub>log\$</sub>	0.62	an unwanted
n	17	Num
r <sup>2</sup>	75%	Measu
		Measure of Go
r <sup>2</sup> <sub>adj</sub>	73%	account the i
		number of
SPE	80%	Measure
p-value	0%	P-value fro
OTA		Optical

- Data on space telescopes and cost was gathered using: •Work Breakdown Structures (WBS) Database •Research Data Storage and Retrieval(Redstar) • Redstone Arsenal Information Center(RSIC) Apply Regression •Cost Acquisition Data Report(CADRe) on Excel

Redstar

Interviews

Collected data was then added to our Excel database.

	А	1.00	J	К	L	М	N	0	Р	Q	R	S	T
1	Telescope	Aperture Diameter	PM F Len.	PM F/N	OTA Volume	FOV	Pointing Accuracy	Total Mass	OTA Mass	Total Areal Density	OTA Areal Density	Spectral Range minimum	Wavel h Diffrac
2		(m)	(m)	unitless	(m³)	(°)	(Arc-Sec)	(kg)	(kg)	(kg/m <sup>2</sup> )	(kg/m²)	<b>(μ)</b>	<mark>(μ</mark>
3	FUSE	0.83	2.24536	2.692989	1.22596656	0.333	0.5	1361	780	2492.674	1428.571	0.0905	
4	GALEX	0.50	1	2	0.19634954	1.2	0.15	280	133	1426.028	677.3634	0.135	
5	HST	2.40	5.28	2.2	23.8861573		0.01	11110	2400	2455.849	530.5165	0.09	0.2
6	IUE	0.45	1.1684	2.63	0.18171947	0.2667	0.1	644	117	4140.721	752.2738	0.115	0.3
7	Kepler	0.95	2.785	3.48125	1.97406883	13.88	0.009	1071	322	1510.958	454.2749	0.43	
8	OAO-3	0.80	2.72	3.4	1.36722112		0.1	2203	428	4382.729	851.4789	0.071	
9	SOHO/ EIT	0.12				0.75	1	1850		163575.9		0.0171	
10	TRACE	0.30				0.14167	5	250	59	3536.777	834.6793	0.0171	
11	COM_0.7	0.70	0.84	1.2	0.32326988	0.93	90		114		296.2231	0.45	0.2
12	CALIPSO	1.00	0.7	0.7	0.54977871	0.0007448	0.05	587	261	747.3916	332.3155	0.532	6.(
13	ICESat	1.00					1.5	970	300	1235.042	381.9719	0.532	6.(
14	HIRISE	0.50	0.865	1.73	0.16984235	1.15	32.17731	953	34	4853.589	173.1606	0.4	0.6
15	IRAS	0.57	0.854	1.5	0.21715623	0.5	30	809	228	3181.516	896.6448	8.5	8
16	ISO	0.60	1.2	2	0.33929201	0.333	2.7	2498		8834.868	_	2.5	2.5
17	JWST	<u>6.</u> 50	16	2.461538	530.929158		0.007	6530	2500	196.7871	75 <mark>.33962</mark>	0.6	2
18	Spitzer	0.85	1.02	1.2	0.57879918	0.5333	0.1	950	851.5	1674.156	15 <mark>00.572</mark>	3	6.5

RSIC



0.10

CER and

Statistics



References om Introduction to Cost Analysis gital image. Web. < http://imgsrc.hubblesite.org/hu/db/images/hs-2008-12-a-full\_jpg.jpg>. al image. Web. <http://light.wittenberg.edu/blog/marywilson/files/2008/10/200px-witt\_seal.gif: gital image. Web. < http://www.jwst.nasa.gov/images\_mirror20.html>.



### ation

- e interval is to contain the parameter
- oise" in log-space. Noise is perturbation to a wanted signal
- ber of data points. re of Goodness of fit odness of fit that takes into number of data points and estimated coefficients. of "noise" in real space. m the coefficient t-test.
- elescope Assembly





Aperture Diameter (m)

1.00

l image. Web. <http://www.nasa.gov/images/content/399983main\_enceladus\_full\_full\_full.jpg>