

Mars In Situ Resource Utilization Technology Evaluation

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We have examined the technologies required to enable Mars In-Situ Resource Utilization (ISRU) because our understanding of Mars resources has changed significantly in the last five years as a result of recent robotic missions to the red planet. Two major developments, (1) confirmation of the presence of near-surface water in the form of ice in very large amounts at high latitudes by the Phoenix Lander and (2) the likely existence of water at lower latitudes in the form of hydrates or ice in the top one meter of the regolith, have the potential to change ISRU technology selection. A brief technology assessment was performed for the most promising Mars atmospheric gas processing techniques: Reverse Water Gas Shift (RWGS) and Methanation (aka Sabatier), as well as an overview of soil processing technology to extract water from Martian soil.

Nomenclature

°C	=	degrees Celsius
ISRU	=	In Situ Resource Utilization
I_{sp}	=	Specific Impulse
K	=	degrees Kelvin
OCT	=	NASA Office of the Chief Technologist
Psia	=	pounds per square inch absolute
TRL	=	Technical Readiness Level
v.p.	=	vapor pressure
W	=	Watts

I. Introduction

Recent national policy statements have established Mars as the ultimate destination of NASA's human exploration program. In Situ Resource Utilization (ISRU) is a key technology required to enable such missions and it is appropriate to review progress in this area and continue to advance the systems required to produce rocket propellant and other consumables on Mars using the carbon dioxide atmosphere and other potential resources. Our understanding of Mars resources have changed significantly in the last five years as a result of recent robotic missions to the red planet^{1,2,3,4}. The indication of the presence of water in significant amounts at high latitudes, and possibly at lower latitudes, has a significant impact on ISRU technology and human exploration architectures. Prior to these robotic findings, the ISRU program has focused on the production of (1) methane from atmospheric CO₂ and hydrogen brought from Earth via the Sabatier reaction and (2) oxygen from a secondary atmospheric CO₂ processing unit in order to produce oxygen at the oxidizer-to-fuel mass ratio (3.4:1) required for propellant usage. The secondary atmospheric processing unit was focused on processes that do not consume Earth-supplied resources (i.e. hydrogen).

Smith and coworkers⁴ reported a water ice table at the Phoenix landing site at depths of 5 to 18 centimeters. Also, a water vapor pressure in the atmosphere of 2 pascals, corresponding to the saturation vapor pressure of water at 210 K, was measured by Phoenix instruments. The utilization of water ice on Mars will require excavation techniques and soil processing technology to extract water. Although the lunar ISRU program was not investigating water ice mining technology, lunar regolith mining and processing are applicable to Mars ISRU. The technology needs for mining water from Mars will be further discussed in this paper.

II. Mars Atmospheric Gas Processing

The production of propellant and life support consumables in-situ on Mars is an enabling technology required for sustainable and affordable human exploration of Mars. Mars offers useful resources in the form of carbon dioxide

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masses, power, volume, and advantages of packed-bed reactors and microchannel reactors as well as ways to prevent catalyst deactivation and regenerate catalysts are needed.

Methanation catalysts are available commercially from BASF, Sud-Chemie, Haldor Topsoe, and others. The catalyst is available in the nickel and ruthenium form for high and low temperature application, respectively. Inquiries have been made to industry with respect to catalyst performance and lifecycle at an industrial level, but no response has been received so far.

Catalyst materials are typically sensitive to impurities and contaminants that would poison and deactivate the catalytic material. In the case of methanation catalyst, the catalytic materials (nickel or ruthenium) are sensitive to sulfur compounds, and slightly sensitive to water and oxygen. Sulfur compounds tend to poison and passivate the catalyst; whereas oxygen and water tend to oxidize the catalyst, deactivating it. As seen from Viking data summarized in Table 1, sulfur is not present in the Martian atmosphere, and oxygen and water are present in small amounts that do not present any immediate concern on the life of the catalyst, since typical methanation reactions are carried out

in a fuel rich environment. However, sulfates and perchlorates do occur in the Martian regolith and dust, requiring high efficiency removal of fines from the gas stream fed to the methanation reactor.

Pioneer Astronautics successfully ran a combined Sabatier and RWGS process continuously for approximately 120 hours using a simulated Mars atmosphere gas mixture (95% CO₂, 2% Ar, and 3% N₂) with no detectable degradation of the process.¹⁰ Combining the exothermic Sabatier reaction and the endothermic RWGS reaction reduces the amount of heat generated by 37%, decreasing concerns about catalyst sintering noted above, but at the cost of a methane-carbon monoxide cryodistillation step. However, the combined reaction also produces oxygen in the proper ratio for combusting all the methane generated, eliminating a separate process to make supplemental oxygen. This advantage is eliminated if Martian water sources are available.

These Martian atmosphere trace gases do not affect the processes' efficiency since they are not playing any role in the CO₂ reaction (i.e. not changing enthalpy or entropy); hence separation of CO₂ from Martian atmosphere may not be critical. These gases slightly decrease the partial pressure of CO₂ in the stream, but this small dilution, and 5% reduction in efficiency, would be off-set by the extra hardware, complexity, and reliability associated with an additional unit operation to perform the CO₂/trace gases separation. However, the power and mass of a pump required to pressurize the Martian atmosphere for processing relative to that required by a freezing process (see below) should be taken into consideration.

Martian trace gases are a useful resource for use as a buffer gas for life support systems, and would eventually require separation and capture. However, separation and capture of the buffer gases (i.e. N₂ and Ar) from CO₂ reduction product gas stream may prove to be simpler than the initial CO₂ separation.

The Sabatier process offers the most straightforward buffer gas separation scheme. The state-of-the-art Sabatier process has a high conversion of reactants into product, and could be considered to be 100% efficient. As shown previously, the CO₂ reduction via Sabatier process yields methane and water. In this process, water and methane can be easily condensed (below 373 K for water and below 112 K for methane) leaving the buffer gases in the gas phase. This separation technique would have to be evaluated to estimate the methane purity required for propulsion, and the solubility of the buffer gases in methane at the condensation temperature and pressure. However, liquid methane could be cryodistilled to remove any impurities that could be present as a result of the process.

C. RWGS Process

The RWGS/Electrolysis process has been developed for Mars in situ propellant production (ISPP) purposes for several years. The process is based on the conversion of Martian carbon dioxide into oxygen and carbon monoxide

Table 1. Composition of the Atmosphere of Mars

Viking atmospheric measurements		
Composition	95.32%	carbon dioxide
	2.7%	nitrogen
	1.6%	argon
	0.13%	oxygen
	0.07%	carbon monoxide
	0.03%	water vapor
	trace	neon, krypton, xenon, ozone, methane
Surface pressure	1-9 millibars, depending on altitude; average 7 mb	

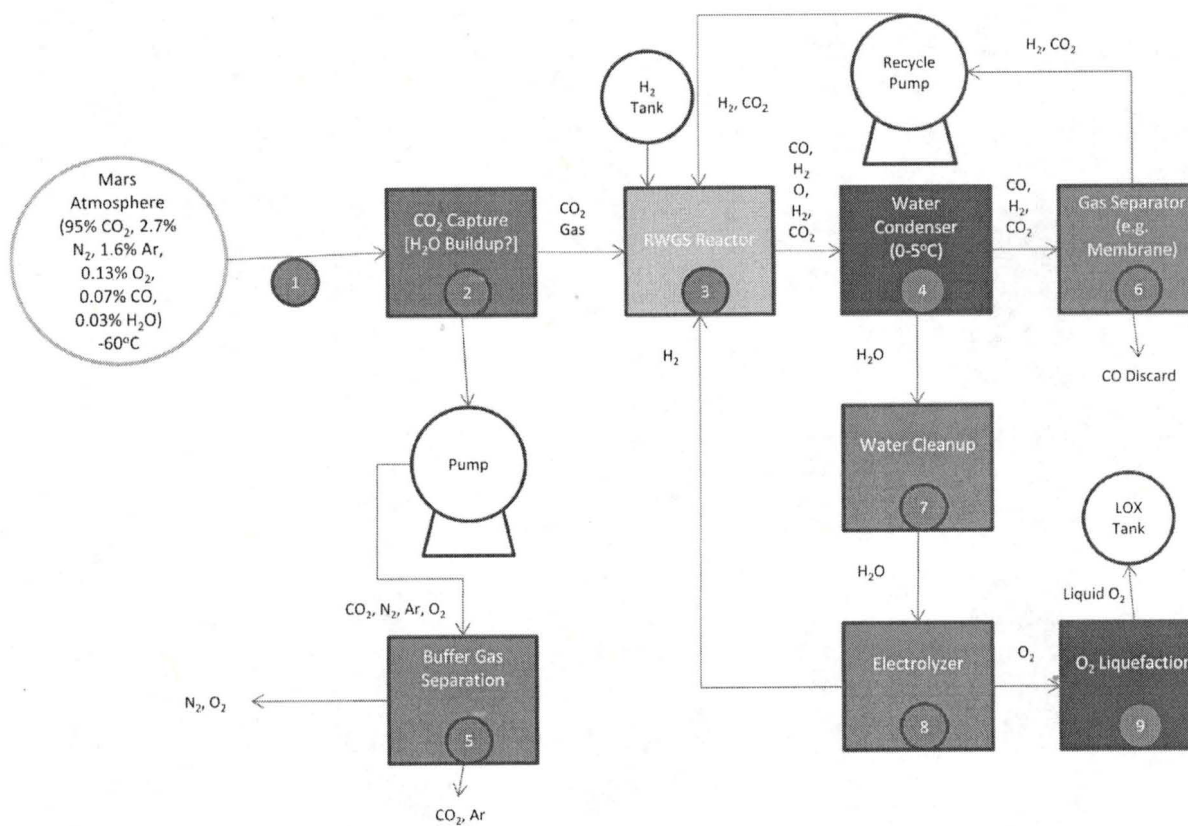
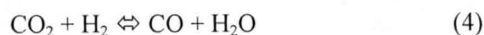


Figure 2. Schematic representation of the RWGS/Electrolysis process for oxygen production.

via hydrogen reduction of carbon dioxide into water and CO, with electrolysis of water to produce oxygen and hydrogen for recycling. Theoretically, the process yields infinite leverage (i.e. the mass of oxygen product divided by the mass of hydrogen brought from Earth), but some loss of hydrogen is inevitable. Nevertheless, the process can give very high leverage in the production of oxygen for both propellant and for life support functions on Mars. Figure 2 shows a simplified schematic of the RWGS process to enable discussion of available and potential technologies for process optimization.

The RWGS process starts (Steps 1, 2) with the collection and concentration of carbon dioxide from the ambient atmosphere of Mars. If desired, potentially useful buffer gases such as nitrogen and argon may be collected as well and possibly separated (Step 5). Depending on the collection process, a pump may be needed to bring the atmosphere into the collection device. Carbon dioxide is sent to the RWGS reactor (Step 3) where it is combined with hydrogen (either from Earth or from indigenous water electrolysis). Normally, the RWGS reaction is incomplete:



Reaction (4) is slightly endothermic with a ΔH of +9 kcal/mol. To improve the yield, the gases are separated and recycled, incorporating more carbon dioxide feed and make up hydrogen, if needed. After the RWGS reactor, the reaction gases are fed to a water condenser (Step 4) where liquid water is collected, deionized (Step 7), and electrolyzed (Step 8) into hydrogen, which is recycled into the RWGS reactor, and oxygen, which is liquefied and stored (Step 9). Alternatively, hot gases could be scrubbed of contaminants before water condensation (not shown). After removal of water vapor, the remaining gas, which consists of carbon monoxide and unreacted hydrogen and carbon dioxide, is passed through a separator, usually a membrane module that removes the carbon monoxide, which is normally discarded, but could be used to produce other hydrocarbons when combined with more hydrogen over the proper catalyst. The carbon dioxide and hydrogen are pumped back into the RWGS reactor to improve the conversion, as noted above.

Previous Mars ISRU efforts (before Phoenix Lander) have included RWGS and SOE as secondary CO₂ processing process in order to produce oxygen in a ratio, with respect to methane, that is appropriate for the propulsion used. The oxygen-to-methane mass ratio produced after electrolysis of the water produced in reaction 2 is 2:1, which is short of the typical 3.4:1 oxidizer:fuel mass ratio used in methane-oxygen engines. Hence, SOE and RWGS provide an oxygen production process option without the consumption of hydrogen brought from Earth. In the new era of Mars exploration and resource utilization, oxygen would be generated from water mined from the soil or subsurface water ice.

D. Electrolysis of Carbon Dioxide in Ionic Liquids

Another method for the production of oxygen from the Martian atmosphere is the electrolysis of carbon dioxide in ionic liquids at much lower temperatures than SOE, i.e. <100°C. The solubility of CO₂ in ionic liquids makes this a potentially promising technology in terms of combining CO₂ capture and oxygen production as well as energy efficiency. However, the Mars Design Reference Architecture 5.0¹¹ notes that this process is at too low a TRL for consideration:

It should be noted that while other technologies and methods for CO₂ processing are possible and have been evaluated, these processes were considered to be too low of a TRL to be evaluated at a system level for mission applicability. These alternative, low-TRL technologies include: molten carbonate electrolysis, non-aqueous electrolysis of CO₂, ionic liquid electrolysis, liquid CO₂ electrolysis, and lower-temperature mobile oxide ceramics. Finding H₂O in sufficient amounts, and of sufficient accessibility, provides an alternative to these low-TRL technologies, via use of electrolysis.

A thorough literature search found very few publications focused on oxygen production using this process beyond a few scouting experiments at KSC and some work by Rosen at the University of Illinois at Urbana-Champaign,¹² The KSC tests had little success in making oxygen because the two ionic liquids tested started decomposing during electrolysis. It was felt that other ionic liquids would be more successful. Rosen concluded that “1-ethyl-3-methylimidazolium tetrafluoroborate (EMIM BF₄) has many favorable characteristics that can enable the reduction of CO₂ at less negative potentials compared to previous work, as well as at low temperatures.” Unlike the KSC work, Rosen used the ionic liquid with 1 M water added, which participated in the electrolysis. He observed the production of hydrogen, oxygen, and carbon monoxide in the same gas stream, attributing some of the oxygen to reduction of CO₂. Use of Martian water would make this approach feasible, but the purpose is unclear if the point is to make oxygen, which could be obtained from water electrolysis alone, unless the power required could be reduced. Rosen’s lab scale apparatus also suffered a short working life as the ionic liquid diffused through carbon paper and coated the cathode catalyst. During a Phase II NASA SBIR project, Kimble et al.¹³ investigated the process with two ionic liquids ([emim-BF₄] and [emim-tFMS]), each absorbed into Nafion membranes. Oxygen was produced, but some problems were encountered with the long-term operation of the electrochemical cells and poisoning of the Pt electrodes by carbon monoxide.

Lueck¹⁴ has summarized the issues encountered for the electrolytic reduction of CO₂ to oxygen:

“Hurdles for Ionic Liquids

- Find one that carbonate is soluble in, or carbonate is the anion.
- (Working with Prof. R. Rogers @Univ. of Alabama)
- Confirm CO₂ reduction, preferably to CO.
- Confirm O₂ production at anode (2:1, CO₂/O₂).
- Confirm long term stability and balanced cell reactions.
- Minimize cell voltage.
- Electrode materials
- Minimize I*R drop → thin electrolyte film, highly conductive.
- Construct porous support for electrolyte (similar to carbonate).
- Construct cell manifolds and multi-cell assemblies.”

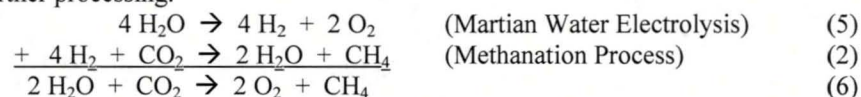
The lack of significant progress in this area over the past decade is likely due to the difficulty in fulfilling the requirements listed above and the fact that there is little need for this process on Earth. More work should be encouraged by the ISRU program, especially in academia.

E. Electrochemical Reduction of CO₂ to Hydrocarbons

During the literature review for the above topic, it was realized that there is an alternative approach that becomes feasible with the known presence of large amounts of water on Mars. This technology is the combined electrolysis of CO₂ and water to produce hydrocarbons. This process has been the subject of intense research for several years and viable methods are emerging. The most relevant study by Sammels¹⁵ was done for NASA JSC to improve the Sabatier process for use on the ISS. The authors state "A typical cell used to implement the process includes a polymer solid-electrolyte membrane, onto which are deposited cathode and anode films. The cathode film is catalytic for electrolytic reduction of CO₂ at low overpotential. The anode film is typically made of platinum. When CO₂ is circulated past the cathode, water is circulated past the anode, and a suitable potential is applied, the anode half-cell reaction is $4\text{H}_2\text{O} \rightarrow 2\text{O}_2 + 8\text{H}^+ + 8\text{e}^-$. The H^+ ions travel through the membrane to the cathode, where they participate in the half-cell reaction $\text{CO}_2 + 8\text{H}^+ + 8\text{e}^- \rightarrow \text{CH}_4 + 2\text{H}_2\text{O}$." The net reaction is $\text{CO}_2 + 2\text{H}_2\text{O} \rightarrow \text{CH}_4 + 2\text{O}_2$. The reaction occurs at room temperature and appears to be efficient, based on the limited technical information available. The Tech Brief refers to a future Ph.2 study, but confirmation that it is underway was not found. This work and similar work also based on copper catalysts for this intriguing reaction should be pursued as a potentially high-payoff technique for fuel and oxygen production on Mars in a single chemical step, avoiding high temperatures, catalysts prone to deactivation, heat exchangers, water condensers, demisters, gas separation membranes, and a separate water electrolysis step. The CO₂/H₂O electrolysis process would provide enough oxygen to burn all the methane plus extra for life support systems (see discussion below).

III. Mining Martian Water

Recent robotic missions to Mars have reported the presence of water ice in subsurface deposits. The water ice deposit can be mined to extract hydrogen and oxygen that can be used for the production of propellants, as well as drinking water and oxygen for life support systems. From a mission architecture perspective, a combination of water electrolysis and methanation process provides a mass ratio of oxidizer:fuel just above the 3.4 oxidizer:fuel typically used for liquid oxygen-methane engines. As shown below, the Mars water-methanation combination provides an oxygen:methane mole ratio of 2 or a mass ratio of 4, which has the potential to provide sufficient oxygen for life support without any further processing.



Clearly, Martian water mining will require excavation systems to obtain the water ice. The ISRU program has begun collaborating with Human Robotic Systems (HRS) to develop excavation implements for roving platforms for lunar applications, which are likely to be applicable to Mars surface systems. However, a different excavation implement will be required for water ice, as the excavation forces will be completely different and the sublimation of water on Mars may be significant. Further development of excavation techniques for Mars application should be pursued given the latest data from Mars.

Extracting water from Martian soil will require the development of soil processing reactors. Although the Mars water extraction process may not be at as high a temperature as those for lunar ISRU, typical terrestrial components have temperature limitations around, and not much higher than, the boiling point of water. Mars soil processing process would require processing temperatures higher than 100°C; hence, technology development on high temperature seals, fluid components (valves, sensors, etc), soil feed systems, and energy efficient thermal design (to minimize the dependency on electrical power) will be required to enable the capability to extract water from Martian soil.

IV. Mars ISRU Supporting Unit Operations

A. Mars Atmospheric Gas Capture

There are several ways to capture and pressurize carbon dioxide from the Martian atmosphere, which is required to purify it from the minor components and to pressurize it sufficiently to obtain a significant feed rate. The literature review we recently performed¹⁶ evaluated several technologies available for this step, which included freezing at cryogenic temperatures, mechanical compression, and adsorption pumping. That study concluded that freezing was the most advantageous process because it required less power than compression pumping, gave high collection rates, purified the carbon dioxide, yielded high pressure feed upon warming, and required less mass than

other options, especially adsorption pumping. Care must be taken that any material used to increase the condensation surface area does not adsorb other Martian atmospheric components, as occurred in one study that used activated carbon.¹⁷

Polymeric membranes have been considered for Mars atmospheric capture and gas separations, but their utility appears to be better for a feed that is already compressed to operating conditions. Thus, membranes will be evaluated below for buffer gas preparation because argon creates problems when used at high concentrations as a component in breathing gas. Membranes are also important devices for process gas separation.

We have also surveyed ionic liquids were also surveyed for carbon dioxide capture.¹⁶ This is an intense area of study for terrestrial carbon dioxide capture and many candidate materials provide good separations from nitrogen, methane, and hydrogen. However, ionic liquids are expensive or require synthesis to prepare and processes using them for Mars applications are at low TRL. Given that there are no clear advantages for them relative to freezing, their pursuit for only this purpose is not recommended at this time, but they could be candidates for proposals to the NASA Office of the Chief Technologist (OCT). Developments should be monitored, however, and OCT has recently funded project formulation activities at KSC for combining carbon dioxide capture and electrolysis of the CO₂ to oxygen and CO or oxygen and methane. The latter process has the potential to greatly simplify the conversion of Martian CO₂ into propellant by combining capture and conversion in a single process vessel at room temperatures as compared to the high temperatures and many unit processes and vessels required for the Sabatier/Electrolysis process.

Similarly there are several other carbon dioxide capture technologies being studied for use on Earth, but some work is aimed at Mars. Amines, amino acid salts, cold dimethyl ether (Selexol), and cold methanol (Rectisol) are commercial CO₂ capture reagents that can be used selectively on flue gases. Although they work on Earth, they do not provide the capability to work at very low Martian pressures and do not provide high pressure CO₂ as a product as freezing does. Brooks and coworkers at the Pacific Northwest National Laboratory (PNNL) have investigated microchannel 13X zeolite absorbers¹⁸ to overcome some of the disadvantages of sorption beds, i.e. being heavy and slow. PNNL calculated a productivity of >2 g CO₂ per kg of hardware for multiple parallel units for a compression ratio of 100, but they did not compare their results to other options. A rough estimate using several assumptions show the PNNL system may have a mass comparable to a CO₂ freezer of the same capacity, but this is a low confidence estimate. On the other hand, Precision Combustion, Inc. has a novel reactor concept that combines CO₂ separation from the Martian atmosphere and conversion to methane using the same lightweight, compact catalyst bed.^{19,20} Such combinations of two or more steps in propellant production technologies should be pursued if they can be shown to save mass, power, volume, etc.

However, if trace gases are not required to be separated from CO₂, mechanical compression of atmospheric gases would have to be considered as an option. This option would require dust filtration technology and reliable and low power design, providing long life-cycles (possibly longer than 5 yrs of operation).

1. Development of a Carbon Dioxide Freezer

Following our own advice that the best current technology for CO₂ capture is cryofreezing,¹⁶ we have been designing a CO₂ freezer to provide feed to a Sabatier reactor as part of the MARCO POLO (Mars Atmosphere and Regolith Collector/Processor for Lander Operations) project being led by JSC. The purpose of the MARCO POLO project is to demonstrate the conversion of Martian CO₂ and water from regolith into methane/oxygen bipropellant on the scale needed for a Mars Sample Return mission. The requirements are to provide 88 g of >99% pure CO₂/hr at 50 psia for a 14-hr day with an initial power limit of 500 W (later increased to 900 W). We selected a Sunpower Cryotel GT cryocooler with 34 W of lift at 150 K, the amount of cooling estimated to freeze the CO₂ and offset heat leaks to the environment. Figure 3 shows a notional drawing of the current design of the module with the Sabatier reactor module in the background. Unlabeled parts of the module include the fluid system, valves, regulators, pressure transducers, a CO₂ pump, and the freezer vessels. The design is still underway and fluid connections for the CO₂ storage tanks remain to be added. Two cryocoolers are required for one to collect dry ice while the other supplies it to the Sabatier reactor. The CO₂ storage tanks are needed because the flange on the cryocooler cannot withstand elevated pressures and the CO₂ pump is used to transfer gaseous CO₂ to the tanks.

We have initiated testing of the freezing process and the cold head size and shape required to collect sufficient dry ice. An initial finned copper cold head based on prior work turned out to be too heavy (1.3 kg), resulting a cool down time too long (43 min) to meet the requirements and minimal dry ice collection. A second test with only the cryocooler tip (1" o.d. x 1.33" long) with a small solid copper cylinder (0.75" o.d. x 1.5" long) attached was much more successful, with a cool down time of only 4.5 min. After 1 hr, 15 min (including cool down time) at 150 K or less and a CO₂ pressure of 11 mbar, a layer of dry ice ~0.2-0.3" thick accumulated on the copper parts (Figure 4). Much of the dry ice was chipped off and it weighed 40 g. Some was lost by sublimation and uncaptured chips. Based on measured thickness, ~78 g of CO₂ was collected in 1.25 hr or 62.4 g/hr, which is 71% of the goal.

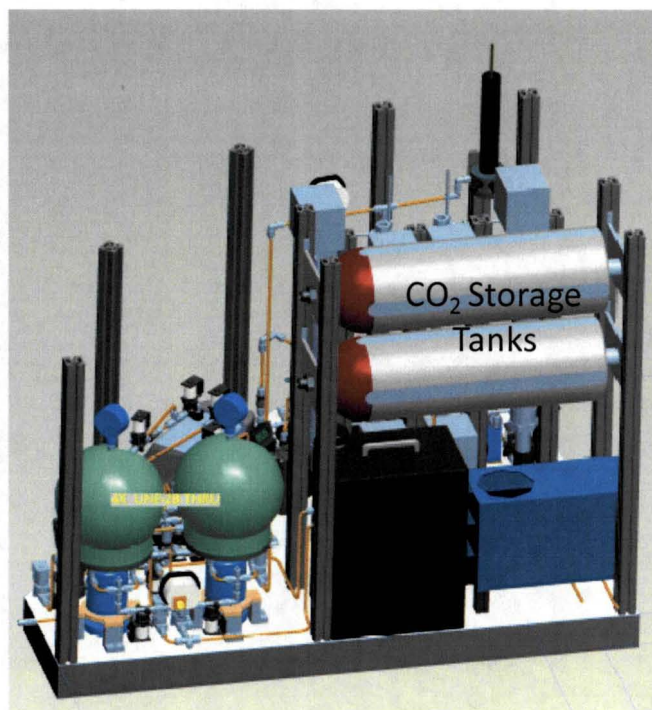
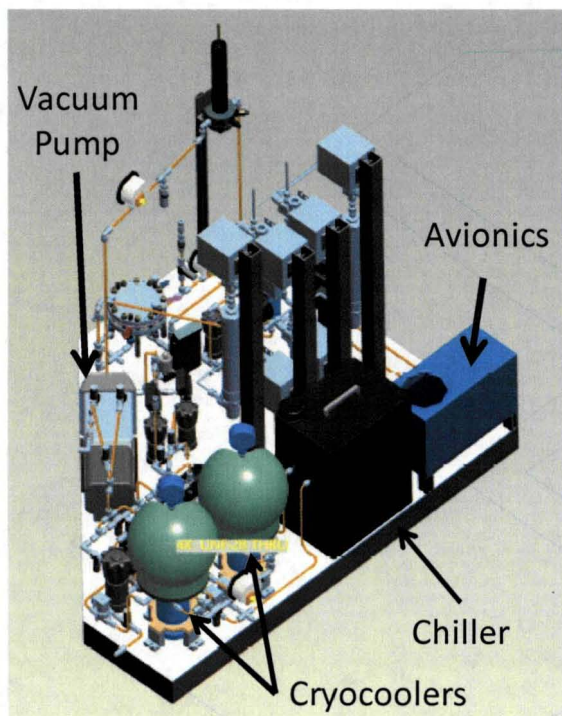


Figure 3. Drawing of the CO₂ Freezer module. The CO₂ Storage Tanks are omitted on the left for clarity.

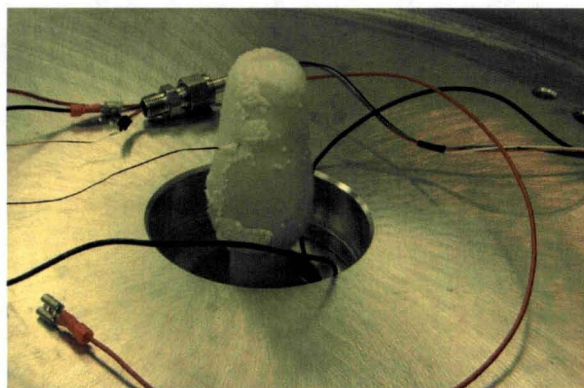


Figure 4. Photo of dry ice collected at 10 mbar CO₂ pressure and 150 K on the cryocooler cold tip plus a 1"x 3/4" cold head.

Based on these results, we designed a new cold head that minimizes copper mass to minimize cool down time and maximizes surface area for dry ice collection. Figure 5 shows the new design that also takes into account the ability of our machine shop to make the item from solid copper. However, testing of the new design also failed to meet requirements. Testing and redesign of the cold head will continue until the desired freezing rate and mass is obtained.

B. Water Condenser

Water condensation is a straightforward cooling process to convert the water vapor into liquid water that has a very low vapor pressure just above 0°C. Previous testing in this process has not identified any issues. Effort must be taken to prevent misting loss of the water downstream for the sake of efficiency. No new development activities need to be undertaken in this area except implementation of a new light-weight material with high thermal conductivity (such as carbon foam).

C. Gas Separation

The gas separation process is very dependent on the ISRU process under consideration. For example: (1) carbon monoxide, hydrogen, and carbon dioxide separation after water condensation

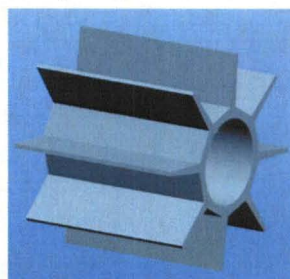
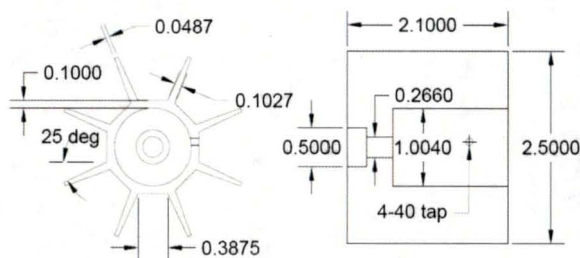


Figure 5. Drawings of the new copper cold head. All units are in inches.

in the RWGS process, (2) unreacted hydrogen and methane after water condensation in the Sabatier process, and (3) volatile contaminants from the soil processing reactor. For the RWGS process, work by Whitlow and Parrish²¹ showed the commercial Permea Prism membrane module was highly efficient. If the mass and volume of the membrane module could be reduced, it should be pursued, but not at a high priority because the existing Permea technology is adequate. For the Sabatier process, membranes separation and cryogenic condensation also offer a good solution to recover hydrogen from methane because the methane would have to be liquefied for ISRU purposes.

Trace gases, nitrogen and argon, are also a useful resource and would require purification and storage. Depending on the Mars atmospheric gas capture technique, buffer gas purification can take place before CO₂ reduction or as a final step in the process. As previously discussed, these trace gases are inert and should not affect the reduction of CO₂ in reactions 1-3. However, if the Sabatier reaction is carried slightly fuel rich (which may extend the life of the catalyst) and trace gases remain in the gas stream, a hydrogen-trace gas separation step would be required.

D. Water Cleanup

NASA's Phoenix lander mission to the northern hemisphere of Mars (68° N) found significant amounts of water-soluble constituents in the Martian soil.^{1,3} These constituents represent potential contaminants to water produced from soil processing. These constituents are potassium, calcium, sodium, magnesium, perchlorate, chloride, and carbonate. The ion exchangers used in the past for this process require large amounts of consumables for regeneration and are therefore not acceptable for Mars ISRU. Regenerable (non-consumable) contaminant removal systems to provide electrolysis-grade water will be required as part of the water extraction and electrolysis system. Various water purification techniques could be applicable to this application: membranes, distillation, reverse-osmosis, electrochemical filtration, among others. Technologies being developed for water cleanup from hydrogen or carbothermal reduction of regolith could be useful for this process as well. Development of all these technologies is needed.

E. Electrolyzer

Water electrolysis will be a critical step required to support Mars ISRU. Water electrolysis is a common need for multiple surface systems, and the ISRU project should continue collaboration in technology development with Power and Exploration Life Support System projects. Currently available water electrolyzers are already highly efficient, although electrolyzers that are not as sensitive to impurities would be a welcome improvement.

F. Cryogenic Liquefaction and Storage

Cryogenics liquefaction and storage has been outside the scope of ISRU for the last several years. This technology development is led by the Cryogenic Fluid Management project, and ISRU should continue to follow the effort as it is required for a complete ISRU architecture.

V. Technology Development Recommendations

The following recommendations are set forward by the authors with the best knowledge of the data presented in this point in time.

1. Mars Atmospheric Gas Processing: Select Sabatier (Methanation) Process for Fuel Production Research and Technology:

- Lifetime of commercial catalyst
- Efficient thermal control
- Effect of pure CO₂ versus raw atmospheric gas on catalyst
- Alternatives (for dissimilar redundancy): microchannel reactors (poison resistant/regenerable), ionic liquid electrolysis of CO₂, combined electrolysis of CO₂ and H₂O, improved SOE subsystems (ruggedized)

Rationale:

- Given the recent robotic mission findings, significant water is available on the surface of Mars
- Water mining-Sabatier Process provides oxygen:methane in a ratio appropriate for propulsion

- This approach will reduce Earth departure mass (i.e. no hydrogen logistics)
2. **Mining Water:** Continue soil excavation system and soil processing reactor development

Research and Technology:

 - High temperature seals, valves, sensors/instrumentation, gas pumps, auxiliary equipment
 - Efficient thermal design
 - Excavation systems: reliable, long life
 - Dust tolerant mechanisms

Rationale:

 - Martian water provides a better architecture for ISRU (i.e. eliminates Earth-supplied hydrogen logistics)
 3. **Mars Atmospheric Gas Capture:** Parallel developmental effort for CO₂ freezer and mechanical compression:

Research and Technology:

 - High efficiency (freezer/compressor)
 - Efficient thermal design (freezer)
 - Long life (cryocooler compressor)
 - Improvements in microchannel absorbers and new approaches

Rationale:

 - The gas capture technique downstream of CO₂ reduction unit depends on the gas capture technique employed to capture atmospheric CO₂
 4. **Water Condenser:** Engineering design to take place during integrated system development effort

Research and Technology:

 - None

Rationale:

 - This is an engineering effort, not a technology development task
 5. **Gas Separation:** Parallel effort of cryogenic condensation and membrane separation

Research and Technology:

 - Polymer Electrolyte Membranes
 - Efficient cryogenic system (combine with cryocooler development above)
 - Solubility of gases in cryogenic liquids

Rationale:

 - Methane purity and atmospheric gas capture selection will drive the final selection of the gas separation (separation of CH₄-H₂ or N₂-Ar-H₂-CH₄)
 6. **Water Cleanup:** Continue regenerable water cleanup

Research and Technology:

 - Regenerable water cleanup to meet electrolysis grade

Rationale:

 - Clean water is required for electrolysis
 7. **Electrolyzer:** Continue Small Business Innovation Research (SBIR) and collaboration with the NASA Environmental Life Support project and the Power project; encourage industry development

Research and Technology:

 - Contaminant resistant electrolysis

Rationale:

- Electrolysis is a common sub-system for various surface system elements.
- Continue team collaboration and SBIR hardware.

8. **Cryogenic Liquefaction and Storage:** Continue collaboration with the Cryogenic Fluid Management project.

Research and Technology:

- Zero boil-off storage
- High efficiency and reliable cryocooler

Rationale:

- Cryogenic storage is needed to provide methane and oxygen propellant
- Cryogenic storage provides high density storage of consumables.

VI. Conclusion

This paper has examined the technologies required to enable Mars In-Situ Resource Utilization (ISRU). Our understanding of Mars resources has changed significantly in the last five years as a result of recent robotic missions to the red planet. The indication of the presence of water in significant amounts at high latitudes, and possibly at lower latitudes, has the potential to change ISRU technology selection. A brief technology assessment is described for the most promising Mars atmospheric gas processing techniques: Reverse Water Gas Shift (RWGS) and Methanation (aka Sabatier), as well as an overview of soil processing technology to extract water from Martian soil. The authors conclude that the technologies needed to (1) concentrate carbon dioxide, nitrogen, and argon on Mars, (2) separate them, (3) process them into oxygen, methane, and buffer gases, and (4) store them for use are already available at a sufficiently high level of development at the component level that further efforts should be focused on engineering development and field demonstrations of integrated systems of these processes. A few technologies that promise improvements in the performance of these subsystems, such as ionic liquids for carbon dioxide collection and electrolysis to oxygen, low temperature electrolysis of carbon dioxide and water to methane and oxygen, and microchannel collectors and reactors, deserve closer scrutiny and possibly more development through the proper funding channels. Furthermore, current lunar soil processing technologies (i.e. excavation, soil reactor, and fluid processing) have applicability to Mars soil processing to extract water and these efforts should continue further development. It is expected that these integrated system engineering units will reveal remaining vulnerabilities caused by interactions of the various subsystems that are currently unknown, leading to new technological needs. Once satisfactory performance is demonstrated in the laboratory, oxygen, propellant, and buffer gas engineering units should be demonstrated at analog Mars sites along with other surface systems to demonstrate a full operational cycle (resource acquisition, processing, purification, storage, and utilization) to raise the TRLs to the 5-6 range. The time is right to push the development of these technologies to convince mission planners that Mars ISRU is a powerful tool for reducing costs and expanding exploration capabilities for Mars missions.

Furthermore, this paper recommends that the ISRU position on whether hydrogen for oxygen and propellant production on Mars should be imported from Earth or obtained from Martian water sources should be the use of Mars water resources. Years of debate on whether it is feasible to transport hydrogen in liquid or slush form to Mars is not likely to be resolved in the near future. The best strategy, which also enables future Mars ISRU, is to use the water resources that are already known by direct measurement to exist on Mars. The problem then is simplified to something we can address: the development and demonstration of efficient Mars water mining and refining technologies, which would also provide pure water, another extremely useful supply item for human missions for both drinking water and radiation shielding. One other extremely important advantage of this strategy is that the amount of oxygen and propellant that could be produced would no longer be limited to the amount of hydrogen brought from the Earth; a virtually unlimited supply would become available, restricted by only the lifetime of the hardware needed and the energy supply. Looking to the more distant future, this strategy is precisely what needed for a Mars outpost and a future Martian civilization. The availability of local Martian water resources will limit landing sites somewhat to known locations, but this is a minor disadvantage because Mars is a large planet with the same surface area as the continents of Earth. In addition, sites with water would enhance the possibilities of finding signs of past or current Martian lifeforms, a significant scientific goal for Mars exploration.

VII. Acknowledgments

This work was sponsored by the NASA In Situ Resource Utilization program. William E. Larson is the ISRU Program Manager. We wish to thank James Captain and Robert Devor of KSC for CO₂ Freezer testing and new cold head designs.

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50th AIAA Aerospace Sciences Meeting
January 9-12, 2012
Nashville, Tennessee

Mars In Situ Resource Utilization Technology Evaluation

**Anthony Muscatello and
Edgardo Santiago-Maldonado**
**NASA – Kennedy Space
Center**



Outline

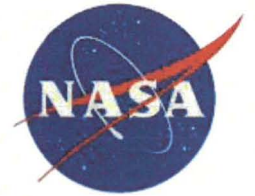
- Background
- Purpose
- Methodology
- CO₂ Capture and Buffer Gases
- Chemical Processes
- Recommendations
- Conclusions

Background – Our Changing View of Mars

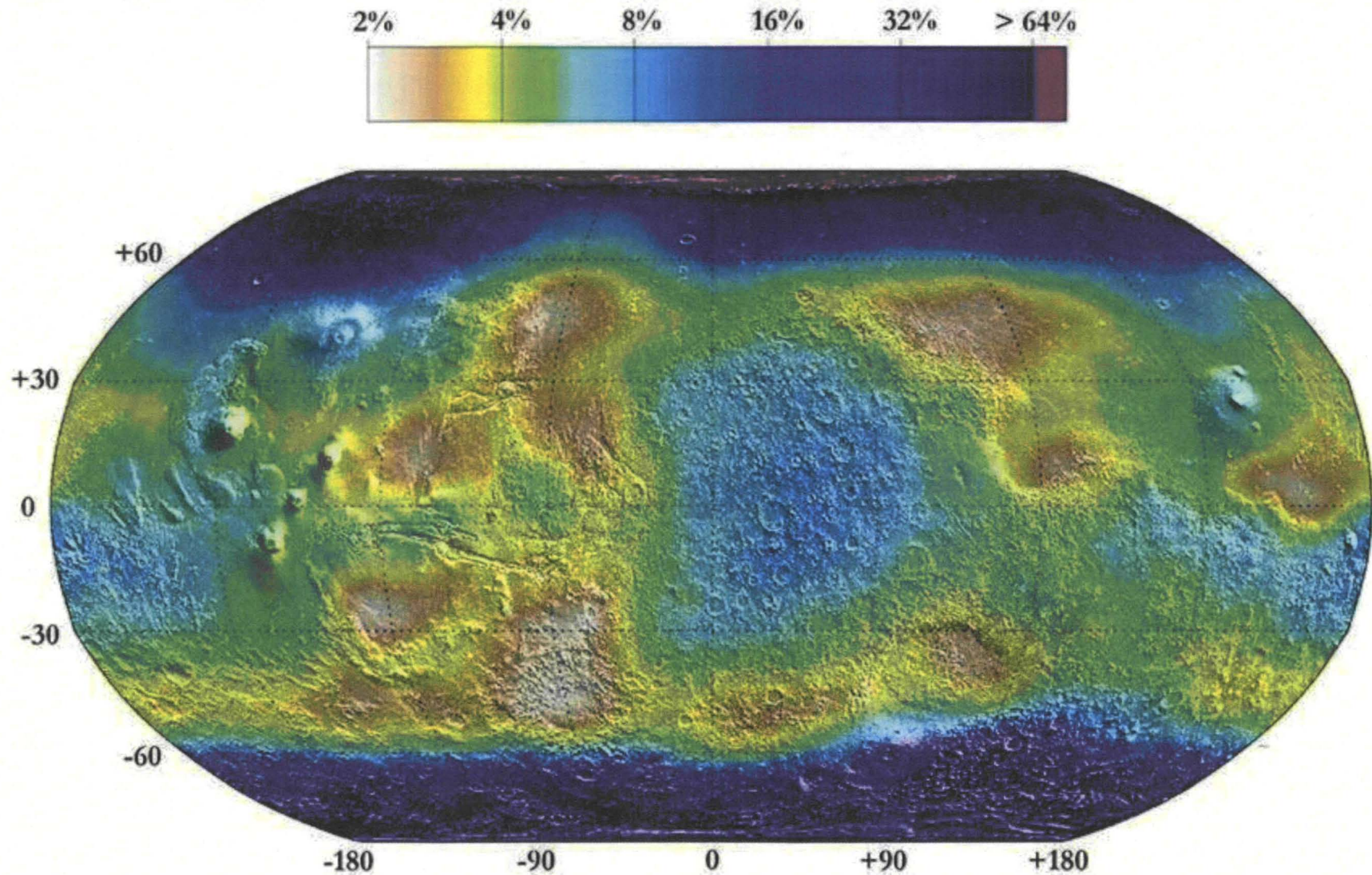


- Mariners – dry, dusty, cold, Moon-like
 - Evidence for ancient water flow
- Vikings – dry, dusty, cold, but intermediate between Moon and Earth
 - RH ~100% (?)
 - Oxidizers in soil, no life (?)
- Pathfinder/Mars Global Surveyor – dry, dusty, cold, but interesting
 - Possible current liquid water flows
- Mars Odyssey – evidence for extensive water ice
- Mars Rovers – in situ proof of ancient, long duration surface water
- Mars Phoenix – in situ proof of near-surface water ice
 - Near neutral pH, perchlorate oxidizer

Water, water everywhere!



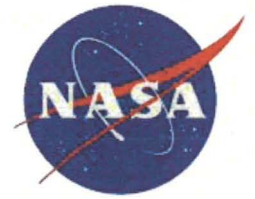
Lower-Limit of Water Mass Fraction on Mars





Purpose

- Update status of Mars ISRU technologies
- Provide guidance for future investments
- Provide basis for Mars ISRU planning



Methodology

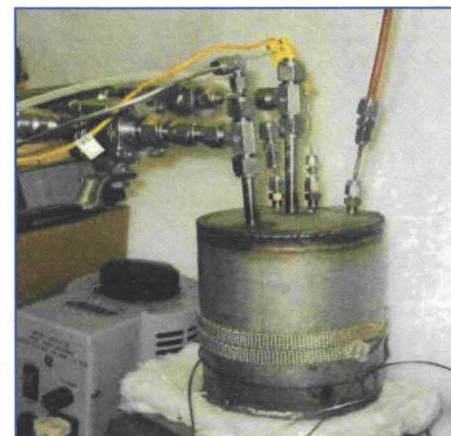
- Evaluated CO₂ capture and gas separation technologies
 - Included new options: ionic liquids, amine-based solvents, preparative GC
- Selected RWGS and Sabatier systems
 - Examined unit processes for TRL status
- Summarized results in a report to ISRU Program Manager Bill Larson
- Full paper in 50th AIAA ASM Proceedings



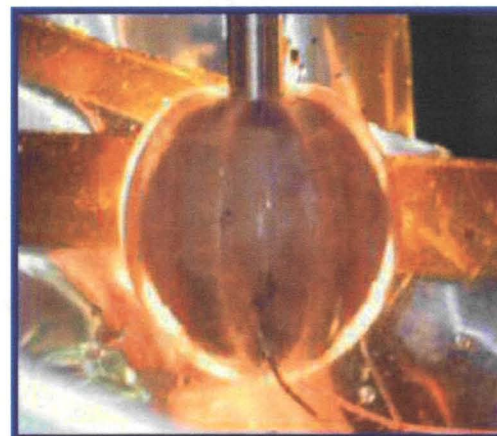
CO₂ Capture Technologies

- CO₂ Freezers Look Promising
- CO₂ freezers have been tested by Pioneer Astronautics and Lockheed-Martin
- Results show accumulation rates of ~20, 13, and 80 g/hr using lab-scale systems (equiv. 5-30 g/hr CH₄)
- N₂/Ar was not measured or purified
- Rapp estimated a CO₂ freezer for 0.5 kg/hr needs ~1/3 the power and 11% the mass of a compression pump/membrane CO₂ purifier
- JPL investigated liquefaction of the Martian atmosphere, but power requirements are high
- Adsorption beds also rejected because of high mass, volume, and power

TRL 3-4



Pioneer MACDOF



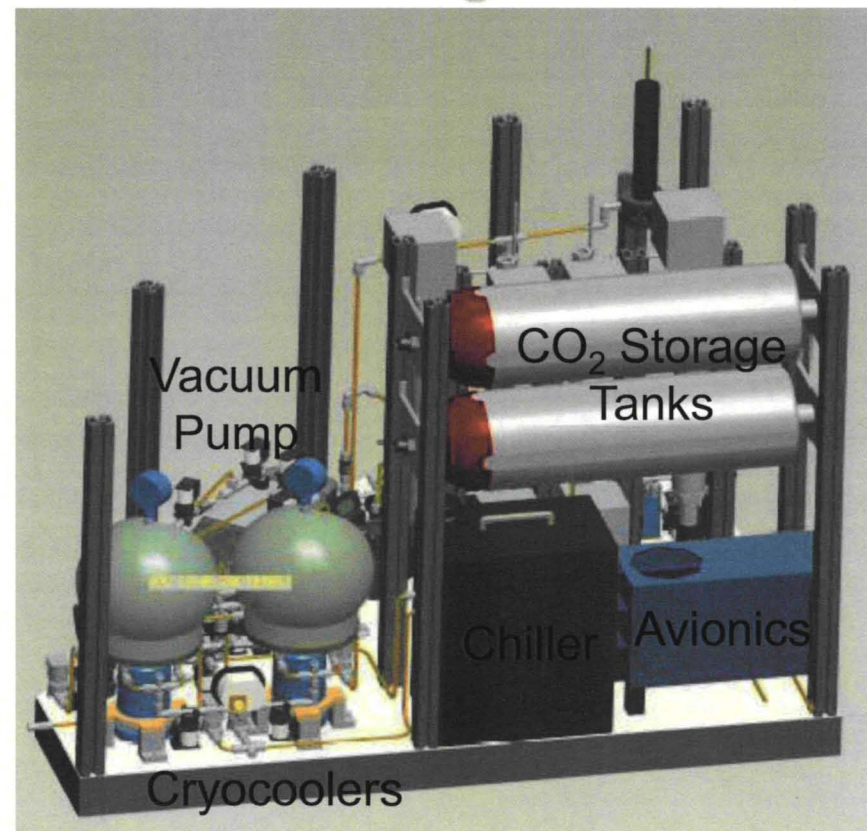
Lockheed Cryocooler Freezer 7

CO₂ Freezer Development

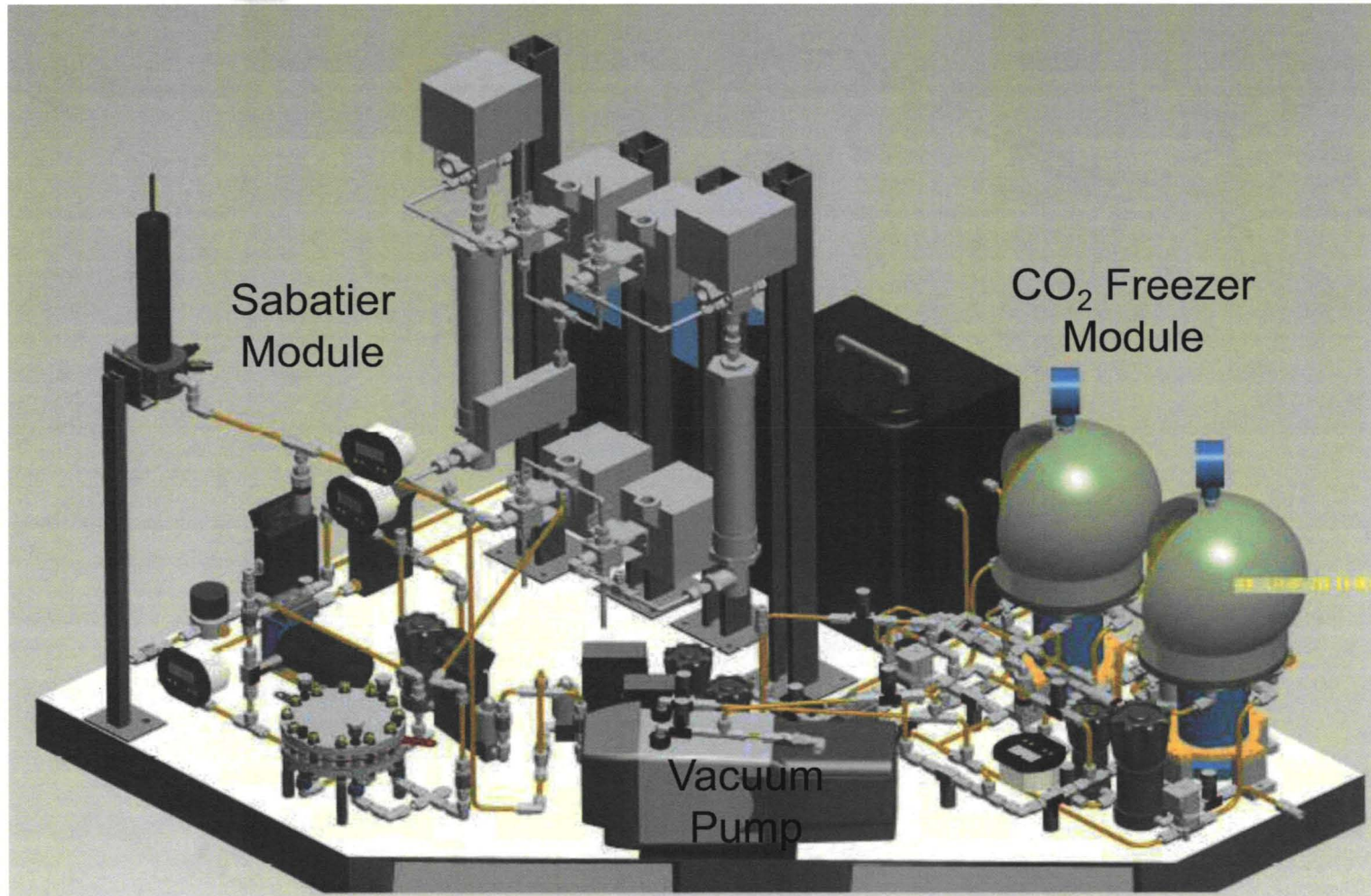
Followed own advice

- Designing/building/testing cryofreezer to test process and separate buffer gases
- Joined forces with MARCO POLO project
- Currently procuring equipment and developing freezer cold head

3-D Model of Atmospheric Processing Module



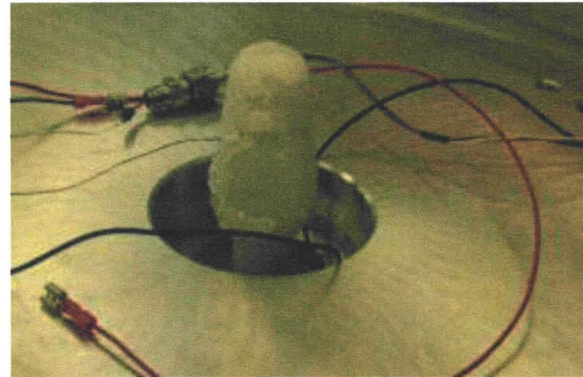
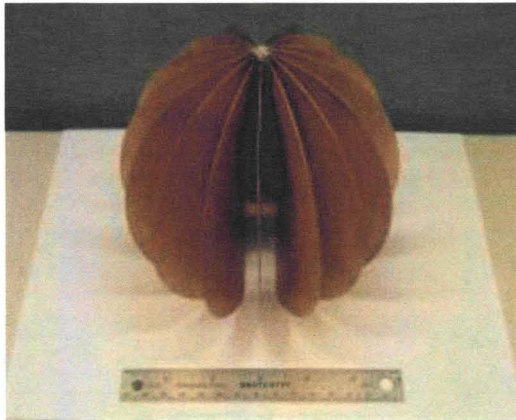
CO₂ Freezer Development



**3-D Model of Atmospheric Processing
Module (CO₂ Tanks Omitted)**

CO₂ Freezer Development

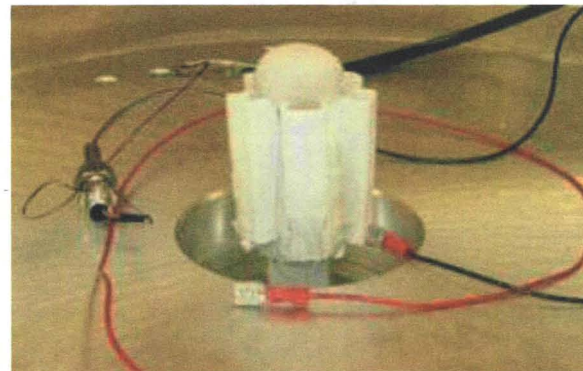
Requirement: 88 g CO₂/hr @ 50 psia



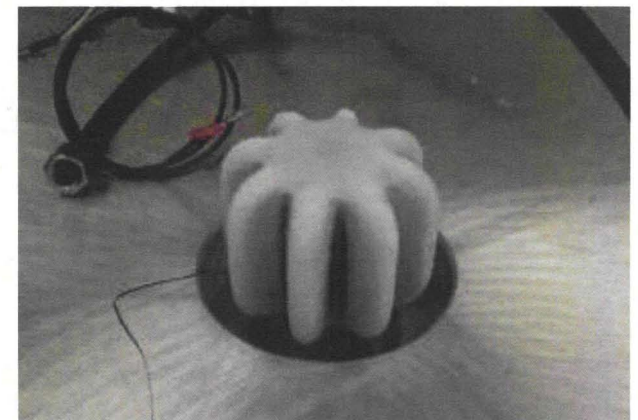
Cold tip + 1x3/4" rod
~60 g/hr



Based on Lockheed
6" fins
~5 g/hr



Cold tip + 1x3/4" rod + Al fins
~20 g/hr

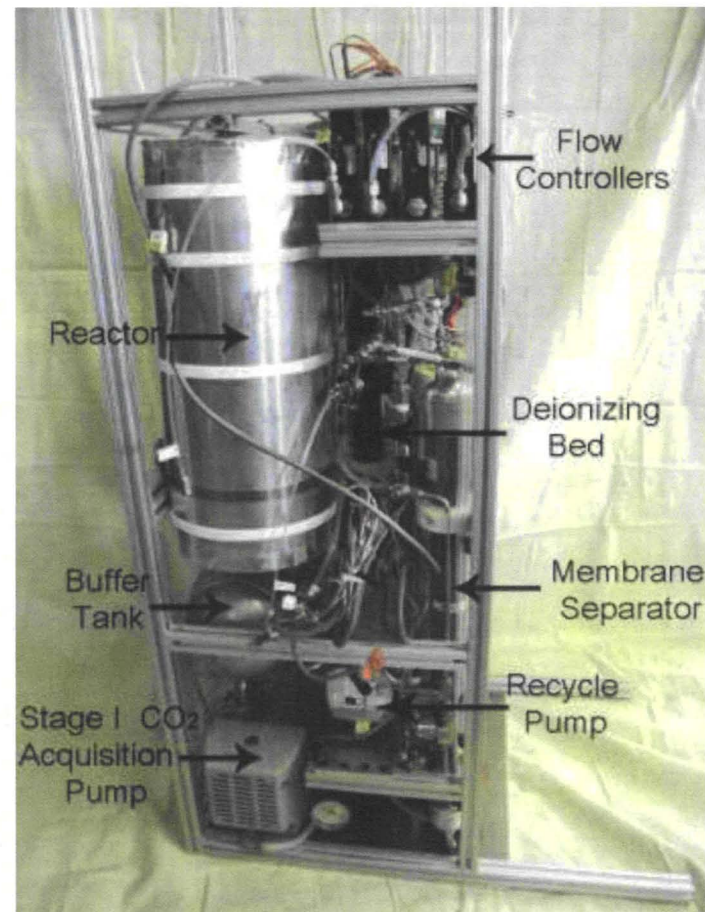


2x2.5" machined fins
~35 g/hr

Alternative Approach- Direct Mars Atmospheric Gas Processing

- ISRU processes (SOE, RWGS, Sabatier) may not require high purity CO₂.
- Pioneer Astronautics ran a combined RWGS-Sabatier process with CO₂/N₂/Ar for 5-continuous days without degradation of catalyst.
- N₂ and Ar were not separated from feed, but were removed during condensation or cryodistillation of products
- Gas separation downstream from CO₂ reduction process may be easier and still provide useful buffer gases
- Mechanical compression is required, and may require more power but was claimed to be less complex.
 - 240 W for 0.6 lpm (71 g/hr) @ 5 bar
- A mass comparison needs to be done, as well.

TRL 3-4



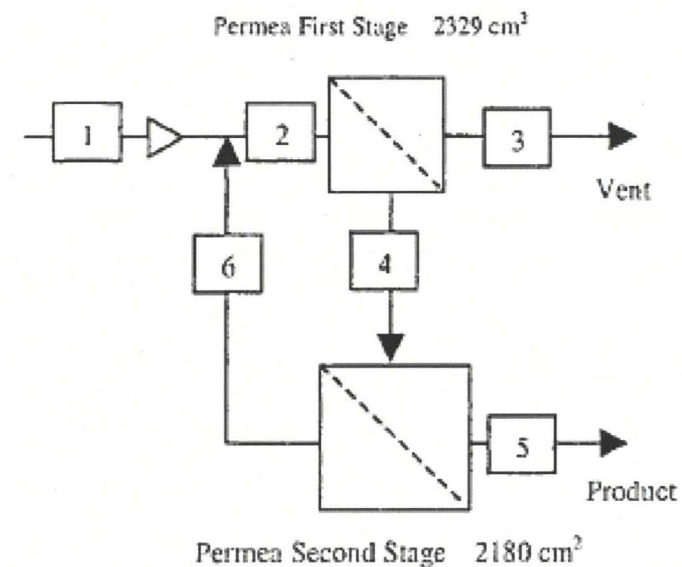
Pioneer IMISPPS



Buffer Gas Separation

- **COTS Membrane Modules Are Adequate**
- **Parrish (KSC, 2002) performed a study of several commercial membranes:**
 - Permea Prism® Alpha Separators PPA-20.
 - Neomecs GT #020101 .
 - Enerfex SS.
 - Enerfex SSP-M100C Membrane sheet.
- **Temperatures = -45°C to +30°C.**
- **Variety of pressures.**
- **Designed a system that would operate at -44°C and 780 mm Hg (1.03 atm)**
- **Feed = 30% CO₂, 26% Ar, and 40% N₂.**
- **Predicted product = 6 lpm, 600 ppm CO₂, 38% Ar and 62% N₂.**
- **47% recovery of the feed.**
- **Work is needed on Ar/N₂ separation.**
 - Ar leads to potential bends issue.

TRL 3-4



**Membrane
purification of feed
from the capture of
CO₂ (Parrish, 2002)**



Chemical Processes

- **Solid Oxide Electrolysis:**



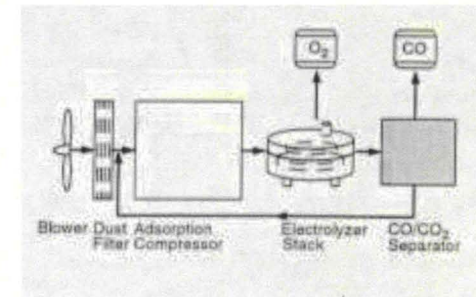
- **Sabatier:**



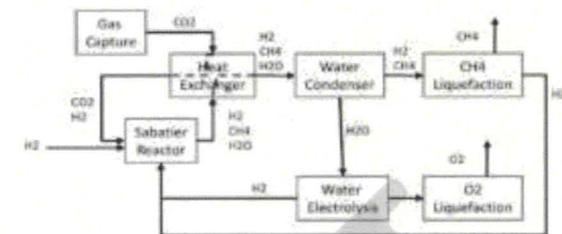
- **RWGS:**



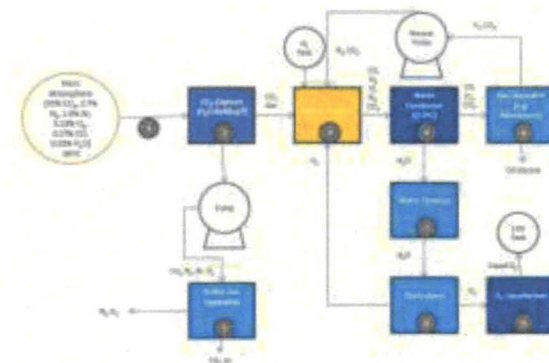
- **SOE:** fragile ceramic membrane, but new developments at Bloom Energy
- **RWGS:** produces only oxygen, but recycles hydrogen many times
- **Sabatier:** well known, but need to take H_2 to Mars and makes only half the oxygen needed
- **So?**



SOE Process



Sabatier Process



RWGS Process

TRL 3-4

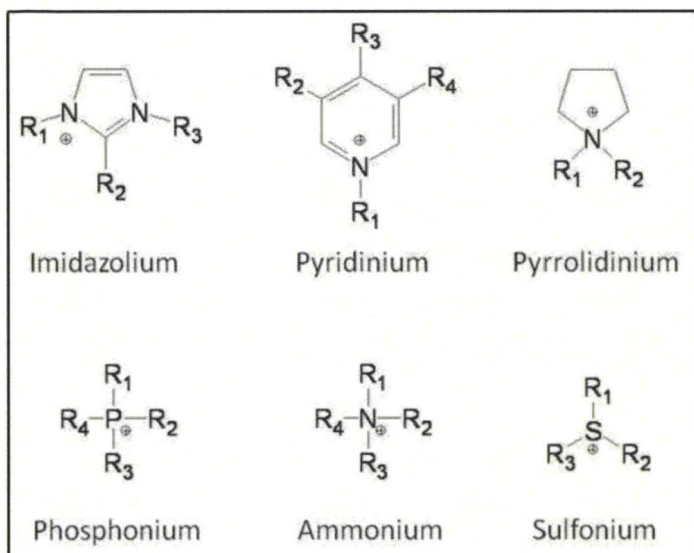
Major Recommendation – Use Mars

Surface Water Resources

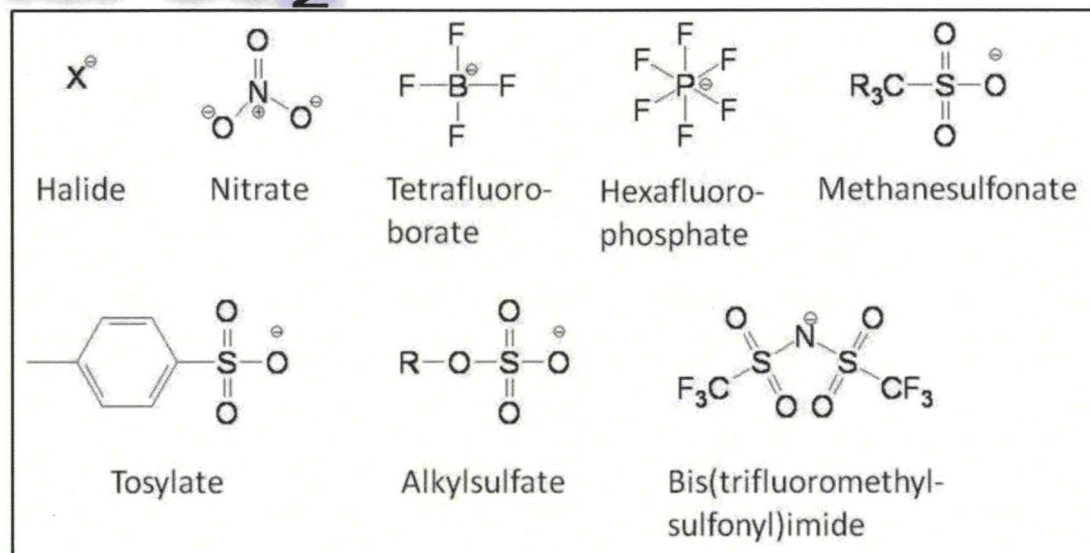


- Considered by DRA 5.0
- We know where the water (ice) is and how much (Phoenix and Mars Odyssey)
- Combined with Sabatier, we can make H_2 on Mars and get the right ratio of CH_4/O_2 with surplus O_2 for life support
- Avoids difficulty of LH_2 storage
- Requires surface mining
 - Base on lunar mining technologies
- Provides a path forward for human exploration and settlement
 - i.e. water for life support and shielding

Chemical Processes - Ionic Liquids for CO₂



Cations



Anions

Typical cations and anions for ILs

- **Ionic Liquids are ionic compounds that are liquid at or near room temperatures**
- **Absorb up to 7% CO₂**



Chemical Processes - Ionic Liquids for CO₂

- Electrolysis of CO₂ in ionic liquids
 - Low TRL, but simplifies processing
 - Requires compression of Martian atmosphere
 - OCT project formulation at KSC
- Electrocatalytic reduction of CO₂ and H₂O to CH₄ and O₂
 - Room temperature, low voltage
 - Also low TRL, but very promising
 - Eltron Research for JSC for ISS application

Chemical Processes - Membrane Separations



- Permea modules have been tested by LMA, Pioneer Astronautics and KSC
 - KSC results are good with minor H₂ losses (0.26% average)
- 28 other candidate membrane materials evaluated
 - Top 10 identified
- Selectivity and permeability are inversely related
- Pressurization is required (1-10 atm)
- Synthesis required in some cases

TRL 3-4

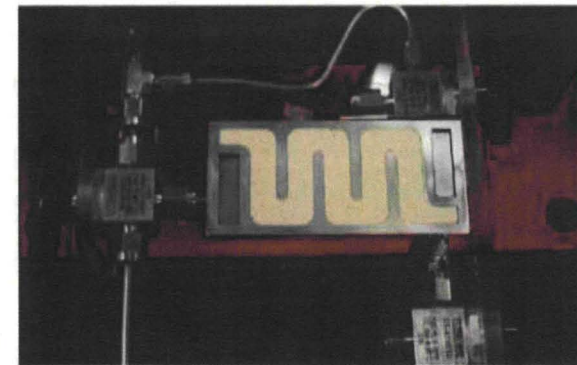
	Calculated	Measured	Calculated	Measured	Calculated	Measured
Stream	H ₂ (slpm)	H ₂ (slpm)	CO ₂ (slpm)	CO ₂ (slpm)	CO (slpm)	CO (slpm)
Base Case Feed	16.825	16.861	7.56	7.522	2.114	2.116
Base Case Permeate	16.812	16.812	7.547	7.547	1.126	1.126
Base Case Reject	0.013	0.013	0.013	0.013	0.988	0.988
Sim-1 Feed	26.265	25.333	7.265	8.126	1.793	1.864
Sim-1 Permeate	26.058	26.205	7.162	7.189	0.64	0.65
Sim-1 Reject	0.207	0.059	0.103	0.076	1.153	1.144
Sim-2 Feed	30.748	30.555	7.984	8.07	1.537	1.644
Sim-2 Permeate	30.011	30.48	7.681	7.801	0.3831	0.541
Sim-2 Reject	0.737	0.267	0.303	0.184	1.154	0.996
Sim-3 Feed	29.552	29.428	8.605	8.696	2.028	2.047
Sim-3 Permeate	29.324	29.463	8.486	8.523	0.723	0.834
Sim-3 Reject	0.228	0.09	0.119	0.082	1.305	1.194
Sim-4 Feed	25.603	25.099	7.035	7.332	1.779	1.986
Sim-4 Permeate	25.515	25.563	6.987	6.96	0.753	0.747
Sim-4 Reject	0.088	0.04	0.048	0.076	1.026	1.032
Sim-5 Feed	16.507	17.248	8.921	8.323	2.47	2.327
Sim-5 Permeate	16.43	16.5	8.843	8.909	1.032	1.384
Sim-5 Reject	0.077	0.007	0.078	0.012	1.438	1.086
Sim-6 Feed	20.591	20.864	7.036	6.786	1.983	1.96
Sim-6 Permeate	20.552	20.564	7.008	7.007	0.952	1.024
Sim-6 Reject	0.039	0.026	0.028	0.029	1.031	0.959

KSC RWGS Membrane Results
(Whitlow and Parrish, 2003)

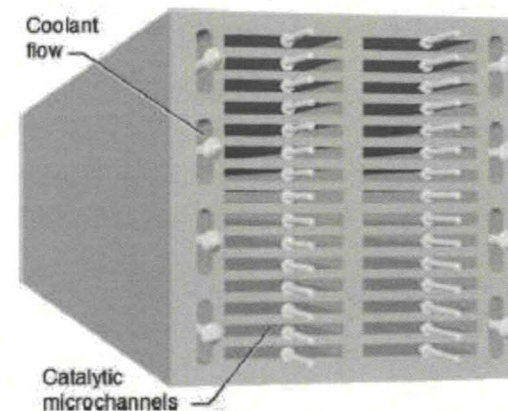


Microchannel Technologies

- **Microchannel reactors offer:**
 - Better temperature control of the catalyst bed
 - Reduce temperature gradients and localized “hot spots”
 - Prevent sintering of a packed bed catalyst
 - Large mass savings over the traditional packed bed reactor design,
 - Penalty of increased pressure drop and increased probability of complete catalyst deactivation.
- **Potentially improved CO₂ absorption for concentration**
 - Lower mass, volume, and power
- **Further development is justified**



PNNL Microchannel Zeolite CO₂ Absorber



TRL 3 PNNL illustration of a section of microchannel reactor.



Recommendations

1. Mars Atmospheric Gas Processing: Select Sabatier (Methanation) Process for Fuel Production

Research and Technology:

- Lifetime of COTS catalyst
- Efficient thermal control
- Effect of pure CO₂ versus raw atmospheric gas on catalyst
- Alternatives (for dissimilar redundancy): microchannel reactors (poison resistant/regenerable), ionic liquid electrolysis of CO₂, combined electrolysis of CO₂ and H₂O, improved SOE subsystems (ruggedized)

Rationale:

- Given the recent robotic mission findings, significant water is available on the surface of Mars
- Water mining-Sabatier Process provides oxygen:methane in a ratio appropriate for propulsion
- This approach will reduce Earth departure mass (no hydrogen logistics)

2. Mining Water: Continue soil excavation system and soil processing reactor development

Research and Technology:

- High temperature seals, valves, sensors/instrumentation, gas pumps, auxiliary equipment
- Efficient thermal design
- Excavation systems: reliable, long life
- Dust tolerant mechanisms

Rationale:

- Martian water provides a better architecture for ISRU (i.e. eliminates Earth-supplied hydrogen logistics)



Recommendations

3. Mars Atmospheric Gas Capture: Parallel developmental effort for CO₂ freezer and mechanical compression:

Research and Technology:

- High efficiency (freezer/compressor)
- Efficient thermal design (freezer)
- Long life (cryocooler compressor)
- Improvements in microchannel absorbers and new approaches

Rationale:

- The gas capture technique downstream of CO₂ reduction unit depends on the gas capture technique employed to capture atmospheric CO₂

4. Water Condenser: Engineering design to take place during integrated system development effort

Research and Technology:

- None

Rationale:

- This is an engineering effort, not a technology development task



Recommendations

5. Gas Separation: Parallel effort of cryogenic condensation and membrane separation

Research and Technology:

- PEM-based membranes
- Efficient cryogenic system (combine with cryocooler development above)
- Solubility of gases in cryogenic liquids

Rationale:

- Methane purity and atmospheric gas capture selection will drive the final selection of the gas separation (separation of $\text{CH}_4\text{-H}_2$ or $\text{N}_2\text{-Ar-H}_2\text{-CH}_4$)

6. Water Cleanup: Continue regenerable water cleanup

Research and Technology:

- Regenerable water cleanup to meet electrolysis grade

Rationale:

- Clean water is required for electrolysis



Recommendations

7. Electrolyzer: Continue SBIR and collaboration with ELS and Power; encourage industry development

Research and Technology:

- Contaminant resistant electrolysis

Rationale:

- Electrolysis is a common sub-system for various surface system elements.
- Continue team collaboration and SBIR hardware.

8. Cryogenic Liquefaction and Storage: Continue collaboration with CFM.

Research and Technology:

- Zero boil-off storage
- High efficiency and reliable cryocooler

Rationale:

- Cryogenic storage is needed to provide methane and oxygen from propellant
- Cryogenic storage provides high density storage of consumables.



Conclusions

- Most of the technology needed for integrated Mars ISRU demonstrations is at a high enough TRL
- Further development is required on more efficient and high-reliability options for a large-scale, fully operational system for human missions
 - Microchannel devices
 - Catalysts
 - CO₂ Freezers