

Conference: First Stars IV  
Location: Kyoto, Japan  
Dates: May 21-25, 2012

Title: "Properties of Massive Stars in Primitive Galaxies"

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Abstract:

According to R. Dave, the phases of galaxy formation are distinguished by their halo mass and governing feedback mechanism. Galaxies in the birth phase (our 'primitive galaxies') have a low halo mass ( $M < 10^9 M_{\text{sun}}$ ); and star formation is affected by photo-ionizing radiation of massive stars. In contrast, galaxies in the growth phase (e.g. Lyman Break galaxies) are more massive ( $M = 10^9 - 10^{12} M_{\text{sun}}$ ); star formation is fueled by cold accretion but modulated by strong outflows from massive stars.

I Zw 18 is a local blue, compact dwarf galaxy that meets the requirements for a birth-phase galaxy: halo mass  $< 10^9 M_{\text{sun}}$ , strong photoionizing radiation, no galactic outflow, and very low metallicity,  $\log(\text{O}/\text{H}) = 7.2$ . We will describe the properties of massive stars in I Zw 18 based on analysis of ultraviolet spectra obtained with HST.