# Mars Global Reference Atmospheric Model 2010 Version: Users Guide 

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The first release of Mars-GRAM occurred in May 1988. A complete history of Mars-GRAM program versions is contained in the marshist.txt file that can be found in the documentation folder within the Mars-GRAM 2010 zip file.
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## PREFACE

The 2010 version of the Mars Global Reference Atmospheric Model (Mars-GRAM 2010) was developed by the Natural Environments Branch, Spacecraft and Vehicle Systems Department, Engineering Directorate, of the NASA Marshall Space Flight Center.

For those unfamiliar with earlier versions of Mars-GRAM, NASA TM-108509, "Mars Global Reference Atmospheric Model (Mars-GRAM 3.34) Programmer's Guide," NASA TM-108513, "A Revised Thermosphere for the Mars Global Reference Atmospheric Model (Mars-GRAM Version 3.4)," NASA/ TM-1999-209629, "Mars Global Reference Atmospheric Model (Mars-GRAM) Version 3.8: Users Guide," NASA/TM-2000-210279, "Mars Global Reference Atmospheric Model 2000 Version (MarsGRAM 2000): Users Guide," and NASA/TM-2001-210961, "Mars Global Reference Atmospheric Model 2001 Version (Mars-GRAM 2001): Users Guide" are recommended. These Technical Memorandums are available electronically from the NASA Technical Report Server at [http://ntrs.nasa.gov/](http://ntrs.nasa.gov/) or in hardcopy from

National Technical Information Service 5285 Port Royal Road
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For information on obtaining Mars-GRAM 2010 (or earlier) code and data, as well as additional copies of this Technical Memorandum, contact

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Examples of output from the University of Michigan Mars Thermospheric General Circulation Model (MTGCM) are available for browsing by interested readers at the following Web site: <http://data. engin.umich.edu/tgem planets archive/thermo.html>. This Web site has a constantly changing archive of available MTGCM case runs for use by the scientific community at large.

Examples of output from the NASA Ames Mars General Circulation Model (MGCM) are available for browsing by interested readers at the following Web site: [http://humbabe.arc.nasa.gov/](http://humbabe.arc.nasa.gov/). This Web site includes example MGCM output based on Mars Orbiter Laser Altimeter (MOLA) topography and pre-MOLA topography.

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## LIST OF ACRONYMS AND ABBREVIATIONS

| ASCII | American Standard Code for Information Interchange |
| :--- | :--- |
| AU | astronomical unit |
| COSPAR | Committee on Space Research |
| ERT | Earth-receive time |
| FTP | General Circulation Model |
| GCM | International Astronomical Union protocol |
| IAU | latitude |
| Lat | longitude-dependent waves |
| LDW | local true solar time |
| Lon | Mars Global Reference Atmospheric Model |
| LTST | Mars Orbiting Laser Altimeter |
| Mars-GRAM | Mars General Circulation Model |
| MET | Mars Global Surveyor |
| MGCM | Mars |
| MGS | Mars |

# LIST OF ACRONYMS AND ABBREVIATIONS (Continued) 

| NPOS | desired number of positions |
| :--- | :--- |
| ODY | Mars Odyssey |
| PC-DOS | personal computer Disc Operating System |
| PDS | Planetary Data System |
| TES | thermal emission spectrometer |
| TT | coorrestrial (dynamical) time |
| UTC |  |

## NOMENCLATURE

| A | coefficient |
| :---: | :---: |
| $A_{0}$ | diurnal mean value of the given parameter (temperature, pressure, density and wind components) |
| $A_{1}$ | amplitude of the diurnal tide component |
| $A_{2}$ | amplitude of the semi-diurnal tide component |
| B | coefficient |
| $B_{0}$ | diurnal mean value of longitude-dependent wave |
| $B_{1}, B_{2}, B_{3}$ | amplitude of the LDW peak for wave-1, wave-2, and wave-3 components |
| $d$ | Julian day at which LDW is evaluated |
| $d_{0}$ | Julian day for the primary peak(s) of the LDW traveling component |
| $F$ | adjustment factor |
| F10 | solar activity parameter |
| F10.7 | solar flux at 10.7 cm wavelength ( $10^{-22} \mathrm{~W} / \mathrm{cm}^{2}$ at 1 AU ) |
| $F_{h}$ | height factor |
| $F(z)$ | logarithmic height factor |
| $g$ | gravity ( $3.712 \mathrm{~m} / \mathrm{s}^{2}$ at the surface of Mars) |
| H | local pressure scale height (km) |
| Lat | latitude (degrees) |
| $L_{S}$ | areocentric longitude of the Sun from Mars (degrees) |
| $m_{d}$ | areal dust density ( $\mathrm{kg} / \mathrm{m}^{2}$ ) |
| $n$ | number of peaks and troughs the wave component has through 360 degrees of longitude |

## NOMENCLATURE (Continued)

| $p$ | pressure ( $\mathrm{N} / \mathrm{m}^{2}$ ) |
| :---: | :---: |
| $p_{s f c}$ | surface pressure ( $\mathrm{N} / \mathrm{m}^{2}$ ) |
| $q_{0}$ | dust mixing ratio, mass of dust per unit mass of air, at the surface |
| $R$ | gas law 'constant' |
| $R_{i}$ | Richardson number |
| $S$ | wave scale parameter |
| $T$ | temperature (K) |
| $T_{5}$ | temperature at $5 \mathrm{~m}(\mathrm{~K})$ |
| $T_{g}$ | temperature at ground surface (K) |
| $t$ | local solar time (hours) |
| $u$ | wind component (m/s) |
| $v$ | wind component ( $\mathrm{m} / \mathrm{s}$ ) |
| ZF | height of the 1.26 nbar pressure level (km) |
| $z$ | height (km) |
| $z_{0}$ | surface roughness parameter |
| $\lambda$ | longitude (degrees) |
| $\Delta z_{0}$ | input value of constant height offset (km) |
| $\varsigma$ | height parameter |
| $\rho$ | density ( $\mathrm{kg} / \mathrm{m}^{3}$ ) |
| $\tau$ | optical depth |
| $\Phi_{1}, \Phi_{2}, \Phi_{3}$ | phases (longitude) of the LDW peak for wave-1, wave-2, and wave-3 components |
| $\Phi_{1}^{\cdot}, \Phi_{2}^{\cdot}, \Phi_{3}^{\cdot}$ | rate of movement of the LDW peak for wave-1, wave-2, and wave-3 components |

## NOMENCLATURE (Continued)

 phase (local time in hours) of diurnal tide componentphase (local time in hours) of semi-diurnal tide component

# MARS GLOBAL REFERENCE ATMOSPHERIC MODEL 2010 VERSION: USERS GUIDE 

## 1. INTRODUCTION

### 1.1 Background and Overview

The Mars Global Reference Atmospheric Model (Mars-GRAM) is an engineering-level atmospheric model widely used for diverse mission applications. Applications include systems design, performance analysis, and operations planning for aerobraking, entry, descent and landing, and aerocapture. Mars-GRAM has been utilized during the aerobraking operations of Mars Global Surveyor (MGS), ${ }^{1}$ Mars Odyssey (ODY), and Mars Reconnaissance Orbiter (MRO). MarsGRAM has also been used in the prediction and validation of Mars Pathfinder hypersonic aerodynamics, ${ }^{2}$ the aerothermodynamic and entry dynamics studies for Mars Polar Lander, ${ }^{3}$ the Mars Aerocapture System Study, as well as the Aerocapture Technology Assessment Group.

Mars-GRAM versions ${ }^{4-8}$ prior to Mars-GRAM 2000 were based on ad hoc parameterizations to data observed by the Mariner and Viking missions. Mars-GRAM 2000, ${ }^{9}$ Mars-GRAM 2001, Mars-GRAM 2005, and the current version, Mars-GRAM 2010, are based on input data tables derived from output results from the NASA Ames Mars General Circulation Model (MGCM) ${ }^{10,11}$ and the University of Michigan Mars Thermospheric General Circulation Model (MTGCM). ${ }^{12,13}$

Section 1 provides an overview of Mars-GRAM 2010. Section 2 of this Technical Memorandum describes the MGCM and MTGCM data and how they are applied in Mars-GRAM. The new features of Mars-GRAM 2010 are described in section 3. Section 4 explains how to obtain the Mars-GRAM code and data files and how to set up and run the program. Sample results are presented in section 5. Appendices A-F provide additional details of Mars-GRAM 2010 input and output files and how to interpret program results. Appendix G describes several auxiliary programs that are provided with Mars-GRAM 2010.

### 1.2 Basic Description of Mars-GRAM 2010

Mars-GRAM's perturbation modeling capability is commonly used, in a Monte Carlo mode, to perform high-fidelity engineering end-to-end simulations for entry, descent, and landing. ${ }^{14}$ Mars-GRAM has been validated ${ }^{15}$ against Radio Science data, and both nadir and limb data from thermal emission spectrometer (TES). ${ }^{16}$

From the surface to 80 km altitude, Mars-GRAM is based on the NASA Ames MGCM. Above 80 km , Mars-GRAM is based on the MTGCM. Mars-GRAM and MGCM use surface topography from the MGS Mars Orbiter Laser Altimeter (MOLA), with altitudes referenced to the MOLA constant potential surface (areoid).

There are several traditional Mars-GRAM options for representing the mean atmosphere along entry corridors. The first option is mapping year 0 , with user-controlled dust optical depth and Mars-GRAM data interpolated from MGCM results driven by selected values of globallyuniform dust optical depth. The second is the auxiliary profile option in which the user can read and use any auxiliary profile of temperature and density versus altitude in the mapping year 0 option. In exercising the auxiliary profile Mars-GRAM option, the values from the auxiliary profile replace data from the original MGCM databases. Examples of auxiliary profiles include data from TES nadir or limb observations or Mars mesoscale model output at a particular location and time. The final option is mapping years 1 and 2, with Mars-GRAM data coming from MGCM results driven by the observed TES dust optical depth during TES year 1 and 2.

Mars-GRAM standard inputs are geographic position and time. The user can also adjust the optical depth of the uniformly mixed background dust level, add a seasonal dust optical depth, set the dust particle diameter and density, and provide the starting areocentric longitude of the Sun from Mars $\left(L_{s}\right)$, position, duration, intensity, and radius of a dust storm. Mars-GRAM outputs include density, temperature, pressure, winds, and selected atmospheric constituents. Three Mars-GRAM parameters allow standard deviations of Mars-GRAM perturbations to be adjusted: rpscale can be used to scale density perturbations up or down, rwscale can be used to scale wind perturbations, and wlscale can be used to adjust wavelengths (spectral range) of the perturbations.

### 1.3 Significant Changes in Mars-GRAM 2010

Mars-GRAM 2010 has been updated to FORTRAN 90/95. Mars-GRAM 2010 now includes adjustment factors that are used to alter the input data from MGCM and MTGCM for the mapping year 0 user-controlled dust case. ${ }^{17}$ The greatest adjustments are made at large optical depths such as $\tau>1$. The addition of the adjustment factors has led to better correspondence to TES limb data from zero to 60 km altitude as well as better agreement with MGS, ODY, and MRO data at approximately 90 to 130 km altitude.

# 2. PREEXISTING FEATURES OF MARS-GRAM 2010 

### 2.1 Mars Orbiter Laser Altimeter Topography Data

### 2.1.1 MOLA Areoid

Flying on MGS, the MOLA produced topographic data ${ }^{18-20}$ at a variety of high resolutions from 1 by 1 degree to $1 / 16$ by $1 / 16$ degree latitude-longitude grids. MOLA topography is measured with respect to a zero elevation surface level known as the MOLA areoid, which is defined as the gravitational equipotential whose average value at the equator is equal to the mean radius determined by MOLA. Mars-GRAM 2010 uses half-degree latitude-longitude resolution data for both MOLA areoid and topography.

Prior to Mars-GRAM 2001, a simple ellipsoid of revolution as zero elevation level was used. Previous resolution for Mars-GRAM topography was 7.5 by 9 degrees, a resolution that is consistent with the evaluation grid of the Ames MGCM. Although Mars-GRAM 2010 works internally with MOLA areoid and topography, and uses these as defaults, program input options also allow users to input and output heights relative to the old Mars-GRAM ellipsoid.

### 2.1.2 MOLA Topography

MOLA topography and areoid radius in Mars-GRAM 2010 are specified in a textformat input file, MOLATOPH.TXT. For Mars-GRAM use, this file must be converted to a binary file, molatoph.bin, by running a conversion program, makebin.f90, discussed in appendices D and F . Each line of the text file is for a given latitude-longitude grid, and contains grid-averaged values of longitude ( ${ }^{\circ} \mathrm{E}$.), latitude ( ${ }^{\circ} \mathrm{N}$.), planetary radius ( m ), areoid radius ( m ), topographic altitude (m), and number of data points in the grid. Planetary radius (radius to the local topographic surface) and areoid radius (radius to the zero elevation surface) are measured along a planetocentric radius direction from the center of mass of the planet. MOLA latitude data are planetocentric, hence, differ slightly from planetographic latitudes. Topographic altitude is the difference between planetary radius and areoid radius. Because of nonlinearities and interpolation methods used in processing MOLA data, grid-averaged topography that is provided in the MOLATOPH.TXT file and used in Mars-GRAM 2010 is slightly different from the difference between the grid-averaged planetary radius and grid-averaged areoid radius.

Relative to MOLA areoid, the highest point on Mars, the peak of Olympus Mons, is 21.2 km , and lowest, a point in the Hellas Basin, is -7.8 km . For the half-degree resolution MOLA data that is used in Mars-GRAM 2010, the highest and lowest elevations are 21 and -7.6 km , respectively.

### 2.2 Mars General Circulation Model Input Data

### 2.2.1 Introduction to MGCM and MTGCM Data

Mars-GRAM 2010 utilizes input data tables from the NASA Ames MGCM ${ }^{10,11}$ and the University of Michigan MTGCM. ${ }^{12,13}$ These tables give the variation of temperature, density, pressure, and wind components with height, latitude, time of day, and $L_{s}$. The tables also provide boundary layer data at the topographic surface, as well as 5 and 30 m above the surface as a function of longitude, latitude, time of day, and $L_{s}$. MGCM data tables cover altitudes from the surface to 80 km . MTGCM data tables cover altitudes of 80 to 170 km . A modified latitude-longitude dependent Stewart-type thermospheric model ${ }^{6}$ is used for altitudes above 170 km , and for dependence on solar activity at higher levels. The Stewart-type thermosphere model starts at a lower boundary condition height of the 1.26 nbar pressure level referred to as height $Z F$. Between 80 km and height $Z F$ (typically at about 125 km ), MTGCM data are used directly and also for dependence on solar activity. MTGCM values are interpolated/extrapolated to any desired solar activity value from MTGCM input data for $\mathrm{F} 10.7=70$ and 130 . F 10.7 is the solar flux at 10.7 cm wavelength in units of $10^{-22} \mathrm{~W} /$ $\mathrm{cm}^{2}$ at 1 astronomical unit (AU). Above 170 km , modified Stewart-type thermosphere model data are used directly. Between height $Z F$ and 170 km , a fairing process is used that smoothly transitions from MTGCM values to Stewart-type model values.

Details and formats of MOLA, MGCM, and MTGCM data files are given in appendices D and F. To facilitate transfer, these files are provided to the user in American Standard Code for Information Interchange (ASCII format). For a shorter Mars-GRAM run time it is best to read the files in binary form. A program, makebin.f, which is discussed in appendix F is provided to convert the ASCII format MGCM and MTGCM data files to binary files on the user's machine.

### 2.2.2 Evaluation of MGCM and MTGCM Tidal Components

For each atmospheric parameter of temperature, pressure, density, and wind components, MGCM and MTGCM data tables provide a diurnal (daily) mean value, and the amplitudes and phases of the diurnal and semi-diurnal tidal components. Tidal values for each parameter are computed from the relation:

$$
\begin{equation*}
\text { Tide }=A_{0}+A_{1} \cdot \cos \left[\frac{\pi\left(t-\phi_{1}\right)}{12}\right]+A_{2} \cdot \cos \left[\frac{\pi\left(t-\phi_{2}\right)}{6}\right], \tag{1}
\end{equation*}
$$

where $A_{0}$ is diurnal mean value of the given parameter, $A_{1}$ is amplitude of the diurnal tide component, $t$ is local solar time in hours, $\phi_{1}$ is phase (local time in hours) of the diurnal component, $A_{2}$ is amplitude of the semi-diurnal tide component, and $\phi_{2}$ is phase (local time in hours) of the semidiurnal tide component.

MGCM and MTGCM tidal coefficients are provided at 5 km height increments starting at 0 km relative to datum level, and ending at 80 km (MGCM) or 170 km (MTGCM) relative to datum level. MGCM coefficient data are provided at 7.5 degrees latitude spacing, while MTGCM data have 5 degrees latitude spacing. Both MGCM and MTGCM data are available at every 30 degrees of $L_{S}$ angle, and include three levels of dust optical depth ( $\tau=0.3,1$, and 3 ). MGCM
tidal coefficients are also provided at the topographic surface and heights 5 and 30 m above local topography. Surface layer MGCM data are at 9 degree longitude spacing (for the same latitudes, $L_{s}$ values, and dust optical depths as MGCM data above the surface layer).

### 2.2.3 Interpolation Methods

Equation (1) is used to evaluate each atmospheric parameter at the desired local solar time $(t)$, at 'corners' of a multidimensional 'box' of grid points. This box contains the desired interpolation location, $L_{s}$, and dust optical depth $(\tau)$. Multidimensional interpolation routines are used to evaluate all atmospheric parameters at locations between the MGCM or MTGCM grid points. For data above the surface layer, interpolation is three-dimensional in latitude, $L_{s}$, and $\tau$. For surface layer data (topographic surface, and 5 or 30 m above the surface), interpolation is fourdimensional in longitude, latitude, $L_{s}$, and $\tau$. Interpolation is logarithmic for $\tau$ and linear for all other dimensions.

Interpolation to a desired height in $\mathrm{km}(z)$ is done by interpolating between two height levels ( $z 1$ and $z 2$ ) from grid point altitudes just above and below $z$. Above the surface layer, $z 1$ and $z 2$ are at the 5 km vertical grid spacing of the MGCM or MTGCM data. Near the surface layer (topographic surface or 5 and 30 m above surface height), altitudes $z 1$ and $z 2$ are adjusted as appropriate. Temperature in $\mathrm{K}, T(z)$, and wind components in meters per second, $u(z)$, and $v(z)$, are found by linear interpolation on height. Pressure, $p(z)$, is found by first computing pressure scale height $(H)$ in kilometers:

$$
\begin{equation*}
H=\frac{(z 2-z 1)}{\ln \left[\frac{p(z 1)}{p(z 2)}\right]} \tag{2}
\end{equation*}
$$

and evaluating pressure in $\mathrm{N} / \mathrm{m}^{2}, p(z)$, from the hydrostatic relation:

$$
\begin{equation*}
p(z)=p(z 1) \exp \left[\frac{z 1-z}{H}\right] . \tag{3}
\end{equation*}
$$

Gas law 'constant' $R$ is evaluated from pressure ( $p$ ), density ( $\rho$ ), and temperature $(T)$ at heights $z 1$ and $z 2$ by:

$$
\begin{equation*}
R(z 1)=\frac{p(z 1)}{[\rho(z 1) T(z 1)]} \tag{4}
\end{equation*}
$$

and

$$
\begin{equation*}
R(z 2)=\frac{p(z 2)}{[\rho(z 2) T(z 2)]} . \tag{5}
\end{equation*}
$$

Density in $\mathrm{kg} / \mathrm{m}^{3}, \rho(z)$, at height $(z)$ is then determined by the gas law relation and a linearly interpolated $R$ value, $R(z)$ :

$$
\begin{equation*}
\rho(z)=\frac{p(z)}{[R(z) T(z)]} . \tag{6}
\end{equation*}
$$

### 2.2.4 Interpolation in the Boundary Layer

MGCM data tables used by Mars-GRAM include ground surface temperature. Between the surface and 5 m height, large temperature gradients can exist. There can also be a difference between ground surface temperature and air temperature 'immediately' above ground. These features must be represented by a boundary layer model. Following the approach used in the Ames MGCM, ${ }^{21}$ Mars-GRAM assumes temperature varies from $T_{g}$ at ground surface to $T_{5}$ at the 5 m level according to the relation:

$$
\begin{equation*}
T(z)=T_{g}+\frac{\left(T_{5}-T_{g}\right)\left[1+F_{h}^{1 / 2} F(z)\right]}{\left[1+F_{h}^{1 / 2}\right]} \tag{7}
\end{equation*}
$$

where the factor $F_{h}$ is given by:

$$
\begin{equation*}
F_{h}=\left(1-16 R_{i}\right)^{1 / 2} \text { if } R_{i}<0 ; F_{h}=\left[1+\frac{15 R_{i}}{\left(1+5 R_{i}\right)^{1 / 2}}\right]^{-1} \text { if } R_{i} \geq 0 \tag{8}
\end{equation*}
$$

as a function of Richardson number $\left(R_{i}\right)$ determined from wind and temperature gradients between the ground and 5 m level. Logarithmic height factor, $F(z)$, is given by

$$
\begin{equation*}
F(z)=\frac{\ln \left(\frac{z}{z_{0}}\right)}{\ln \left(\frac{5}{z_{0}}\right)} \tag{9}
\end{equation*}
$$

where $z_{0}$ is the surface roughness parameter assumed to be 0.01 m , except over surface ice, where $z_{0}=0.0001 \mathrm{~m}$ is used. ${ }^{22}$

Wind components at heights $<5 \mathrm{~m}$ above the surface are evaluated from a logarithmic boundary layer profile relation:

$$
\begin{equation*}
u(z)=u(5) F(z) \tag{10}
\end{equation*}
$$

and

$$
\begin{equation*}
v(z)=v(5) F(z) . \tag{11}
\end{equation*}
$$

### 2.3 Longitude-Dependent (Terrain-Fixed) Wave Model

Tide components evaluated by equation (1) depend only on local solar time. Implicitly, this equation also depends on longitude. At any given instant, solar time varies at a rate of 1 hour for
every 15 degrees of longitude. During aerobraking operations, measurements by MGS, ${ }^{23-26}$ ODY and MRO revealed substantial longitude-dependent wave patterns for atmospheric density. Being in Sun synchronous orbits, MGS, ODY and MRO passed through each periapsis at essentially the same latitude and local solar time. Nevertheless, they found substantial variations that tended to repeat as a function of the periapsis longitude. The density variations were of the form of longi-tude-dependent (i.e., terrain fixed) wave patterns. Mars-GRAM 2010 includes an optional model for these longitude-dependent waves (LDW) of the form:

$$
\begin{align*}
\mathrm{LDW}= & B_{0}+B_{1} \cos \left[\pi\left(\lambda-\Phi_{1}-\Phi_{1}^{\cdot}\left(d-d_{0}\right)\right) / 180\right]+B_{2} \cos \left[\pi\left(2 \lambda-\Phi_{2}-\Phi_{2}^{\cdot}\left(d-d_{0}\right)\right) / 180\right] \\
& +B_{3} \cos \left[\pi\left(3 \lambda-\Phi_{3}-\Phi_{3}^{\cdot}\left(d-d_{0}\right)\right) / 180\right], \tag{12}
\end{align*}
$$

where $\lambda$ is longitude (in degrees), $B_{0}$ is the diurnal mean value of longitude-dependent wave, $B_{1}$, $B_{2}$, and $B_{3}$ are amplitude, $\Phi_{1}, \Phi_{2}$, and $\Phi_{3}$ are phases (longitudes), $\Phi_{1}{ }^{\circ}, \Phi_{2}{ }^{\circ}$, and $\Phi_{3}{ }^{\circ}$ are the rate of movement of the LDW peak for wave-1, wave-2, and wave-3 components, $d_{0}$ is the Julian day for the primary peak(s) of the LDW traveling component, and $d$ is the Julian day at which LDW is evaluated. The term wave- $n$ means the wave component has $n$ peaks and troughs through 360 degrees of longitude. LDW perturbations computed by equation (12) are applied as a multiplier to the mean density and pressure computed from MGCM and MTGCM data, as interpolated by methods described in section 2.2.3. Wave model coefficients for equation (12) can be input from the NAMELIST format input file (see app. B), or from an auxiliary file of time-dependent wave model coefficients named by the input parameter WaveFile (secs. 4.2 and 4.3). Values of LDW coefficients may be determined empirically by accelerometer observations, ${ }^{23,25}$ or theoretically, from wave characteristics of Mars GCMs. ${ }^{24,26}$

For altitudes above 100 km , LDW perturbations from equation (12) are assumed to be altitude independent. For altitudes below 100 km , LDW perturbations are assumed to diminish in magnitude at an exponential rate, namely,

$$
\begin{equation*}
\operatorname{LDW}(z)=1+(\operatorname{LDW}(100)-1) \exp [(z-100) / S] \tag{13}
\end{equation*}
$$

where $S$ is the wave scale parameter Wscale, from the NAMELIST format input file.
One way of adjusting Mars-GRAM 2010 density values up or down (at altitudes below 100 km ) is by changing the LDW mean term, $B_{0}$. For example, if the user wants to adjust Mars-GRAM 2010 density values by a factor $W 1=B_{0}(z 1)$ at height $z 1$ and $W 2=B_{0}(z 2)$ at height $z 2$ (where $z 1$ and $z 2$ are both $<100 \mathrm{~km}$ ), then use the scale parameter value:

$$
\begin{equation*}
S=(z 2-z 1) / \ln [(W 2-1) /(W 1-1)], \tag{14}
\end{equation*}
$$

which yields an LDW multiplier value at $100 \mathrm{~km}, B_{0}(100)$, given by

$$
\begin{equation*}
B_{0}(100)=1+(W 2-1) \exp [(100-\mathrm{z} 2) / \mathrm{S}] . \tag{15}
\end{equation*}
$$

Once values of $S$ and $B_{0}(100)$ are input to the program, density at any height ( $z$ ) (below 100 km ) is adjusted by the factor,

$$
\begin{equation*}
B_{0}(z)=1+\left(B_{0}(100)-1\right) \exp [(z-100) / S] . \tag{16}
\end{equation*}
$$

Note that multipliers may be larger or smaller than 1 (yielding density increase or decrease, respectively).

### 2.4 Mars-GRAM Climate Factors and Height Adjustment

Below 170 km, Mars-GRAM 2010 is based directly on MGCM and MTGCM output and does not need any of the climate factors that were included prior to Mars-GRAM 2001. The only climate adjustment factor that is still included in Mars-GRAM 2010 is deltaTEX, which adjusts the exospheric temperature. Mars-GRAM 2010 model output can be affected as a result of the choice of dust optical depth through the input parameter Dusttau and by the LDW parameters that were discussed in section 2.3.

Mars-GRAM 2010 density output can also be adjusted through the use of the input parameter zoffset, the direct height offset of the MTGCM data. The height offset adjustment only affects density above 80 km . The function of the height offset is very similar to that of the LDW wave parameter $B_{0}$ that was discussed in section 2.3. $B_{0}$ shifts a height versus density plot to the right or left as it increases or decreases density at a given height. Height offset shifts such a height-versusdensity curve up or down as it increases or decreases the height at which a given density applies. The net result of a positive or negative height offset is to increase or decrease the density at a given height.

Zoffset can be specified in several ways. A specific offset value in kilometers may be specified in the NAMELIST format input file. Alternatively, the offset value required to match MTGCM data to MGCM data at 80 km can be computed and applied by Mars-GRAM. The MTGCM-MGCM matchup can be specified as applicable locally or globally, based on data shown in table 1 . A local matchup can be based on either density at a given location and time of day or the daily average density at a given location.

Another option is to have the program compute and use a global height offset that depends on time of year given by the solar longitude $\left(L_{S}\right)$. Based on comparisons of MTGCM with density observed during MGS aerobraking, ${ }^{25}$ time-of-year dependence of height offset is given as:

$$
\begin{equation*}
\text { Height Offset }(\mathrm{km})=\Delta z_{0}-2.5 \sin \left(\pi L_{s} / 180\right) \tag{17}
\end{equation*}
$$

where $\Delta z_{0}$ is an input value of constant height offset given by input parameter zoffset. With a default value of $\Delta z_{0}=5 \mathrm{~km}$, height offset values from equation (17) vary seasonally from 2.5 km at $L_{s}=90$ degrees to 7.5 km at $L_{S}=270$ degrees.

Table 1. Global average height offset (km) required for MTGCM-MGCM matchup, as a function of solar longitude and dust optical depth.

| Solar <br> Longitude <br> (deg) | Dust Optical Depth |  |  |
| :---: | ---: | ---: | ---: |
|  | 0.3 | 1 | $\mathbf{1}$ |
| 0 | -0.002 | -0.001 | -0.002 |
| 30 | -0.001 | 0.000 | 0.003 |
| 60 | 0.001 | -0.003 | -0.002 |
| 90 | -0.002 | -0.002 | -0.000 |
| 120 | 0.001 | -0.002 | -0.001 |
| 150 | 0.004 | -0.002 | 0.011 |
| 180 | -0.003 | -0.004 | -0.002 |
| 210 | -0.002 | 0.001 | -0.001 |
| 240 | -0.002 | -0.001 | 0.018 |
| 270 | -0.002 | 0.004 | -0.003 |
| 300 | -0.001 | 0.008 | -0.002 |
| 330 | 0.033 | -0.001 | -0.003 |
| 360 | -0.002 | -0.001 | -0.003 |

Mars-GRAM allows optional regional or global scale dust storms to be activated at any desired $L_{s}$. Dust storm simulations are discussed in section 2.5. Based on comparisons between Mars-GRAM and density observed by the MGS accelerometer ${ }^{23}$ during the regional Noachis dust storm, an additional height offset of MTGCM data is applied during simulated dust storms. This additional offset amount in kilometers is seven times the dust storm optical depth.

### 2.5 Quantitative Dust Concentration Model

Background dust optical depth $(\tau)$ in a non dust storm case is specified by the input parameter Dusttau. Interpolation routines given in section 2.2.3 interpolate logarithmically between $\tau$ values for both MGCM and MTGCM input data. If Dusttau $=0$ is input, a prescribed Viking-like seasonal variation of dust optical depth is used, in which case variation of $\tau$ with $L_{S}$ (in degrees) is specified by:

$$
\begin{equation*}
\tau=0.65-0.35 \sin \left(\pi L_{s} / 180\right) \tag{18}
\end{equation*}
$$

A model for global or local scale dust storms ${ }^{5}$ is retained for mapping year 0 . In MarsGRAM 2010, the input value for dust storm intensity given by input parameter, INTENS, is equivalent to the peak dust optical depth for the storm. Allowable values for INTENS range from zero for no dust storm to three for the maximum intensity dust storm. Dust storm intensity is added to background dust optical depth to give total dust optical depth. Mars-GRAM does all necessary interpolations on dust optical depth as it varies with time, $L_{s}$, and space for local storms. DUST$L A T$ and DUSTLON input parameters give the location of the dust storm. The input option,
$A L S D U R$, allows users to control the duration of simulated dust storms. Input parameter $R A D M A X$ is the maximum radius (kilometers) a dust storm can attain. The radius develops according to the parameterized space and time profile of buildup and decay in the program. If a value of zero or more than $10,000 \mathrm{~km}$ is used for $R A D M A X$, the storm is taken to be of global dimensions and uniformly covers Mars, but is still assumed to build up and decay in intensity according to the same temporal profile.

Mars-GRAM 2010 computes several dust concentration parameters from dust optical depth. Methods used by Haberle et al. in the MGCM are employed. ${ }^{21,27}$ Areal dust density $\left(m_{d}\right)$, total column mass of dust per unit ground surface area, is 0.005 times $\tau$. Dust mixing ratio, the mass of dust per unit mass of air, at the surface $\left(q_{0}\right)$ is computed by:

$$
\begin{equation*}
q_{0}=m_{d} g /\left(0.994 \exp ^{-v} p_{s f c}\right), \tag{19}
\end{equation*}
$$

where $g$ is gravity, $3.712 \mathrm{~m} / \mathrm{s}^{2}$ at the surface of Mars; $v$ is input parameter Dustnu, which controls the vertical dust distribution; and $p_{s f c}$ is surface pressure. Dust mixing ratio at height $z$ is determined by:

$$
\begin{equation*}
q(z)=q_{0} \exp \left\{v\left[1-p(z) / p_{s f c}\right]\right\} \tag{20}
\end{equation*}
$$

where $p(z)$ is pressure at height $z$. Dust mass density, the mass of dust per unit volume of air, is the product of dust mixing ratio and atmospheric density. From the dust mixing ratio, assuming that dust particles are spheres of a given diameter and mass, Mars-GRAM 2010 also computes dust number density, number of dust particles per unit volume of air. Users can also input values of dust particle diameter through the use of the input parameter Dustdiam and dust particle density using the input parameter Dustdens. Consistent with the MGCM, ${ }^{27}$ Mars-GRAM 2010 assumes a default particle diameter, $5 \mu \mathrm{~m}$, and a default particle density, $3,000 \mathrm{~kg} / \mathrm{m}^{3}$.

Dust model output values are written to the file MarsRad.txt details of which are discussed in appendix A.

### 2.6 Solar and Thermal Radiation From Mars-GRAM Output

Several auxiliary programs used with Mars-GRAM are described in appendix G. One, marsrad.f90, computes the upwelling and downwelling components of the solar (shortwave) and thermal (longwave) radiation at the surface and the top of the atmosphere. This auxiliary program uses the Mars-GRAM output file MarsRad.txt, which includes the dust concentration information discussed in section 2.5. A full discussion of MarsRad.txt is contained in appendix A .

To compute upwelling shortwave radiation at the surface, Mars-GRAM uses surface albedo values from the data file albedol.txt that contains the surface albedo at 1 degree latitudelongitude resolution. ${ }^{28,29}$

The auxiliary program marsrad.f90 produces two output files: radlist.txt and radout. txt. The computation methods that are used within marsrad.f90 to determine the radiation
components contained in the output files are discussed in appendix G. Mars-GRAM output that is used as input to marsrad.f90 consists of one or more vertical profiles of temperature versus pressure (altitude). To facilitate the production of latitude-longitude, latitude-time, or longitudetime arrays of such profiles, program radtraj.f90 is also provided as an auxiliary program and is discussed in appendix $G$.

### 2.7 Slope Wind Model

The slope wind model contained in the Mars-GRAM 2010 subroutine slopewind is based on Ye et al. ${ }^{30}$ In this model, the slope winds depend on terrain slopes that are determined from MOLA 0.5 degree resolution topography for the MOLA grid containing the latitude-longitude of interest. The method of Ye et al. ${ }^{30}$ provides an analytical relationship to compute winds as a function of altitude and terrain slope for daytime, thermally-driven, upslope winds. The slopewind subroutine assumes a diurnal pattern of slope wind variation with time of day. The slope winds apply up to 4.5 km above the surface during the day and up to 2.5 km above the surface at night and are added to background MGCM winds at $7.5 \times 9$ degree resolution. Peak daytime winds are assumed to occur at 15 hours local solar time. Peak nighttime winds, due to downslope drainage flows, are assumed to occur at 3 hours local solar time. Lighter, cross-slope flows occur between these two times. Vertical component slope winds are also computed. These are proportional to the terrain slope, to the horizontal slope winds, and have an assumed variation with altitude, which gives maximum vertical winds near the middle of the slope wind altitude region, and zero vertical winds near the surface and near the top of the slope wind altitude region. The slope winds can also be scaled with the input factor, blwinfac, with a value of zero suppressing the slope wind output.

### 2.8 Traditional Mars-GRAM Options for Representing the Mean Atmosphere Along Entry Corridors

Mars-GRAM has several options for representing the mean atmosphere along entry corridors and these can be set using the input parameter MapYear. The MapYear $=0$ option allows the user to control the dust optical depth and utilizes data interpolated from MGCM model results driven by selected values of globally-uniform dust optical depth. If MapYear $=1$ is chosen, data are from MGCM model results driven by observed TES dust optical depth during TES mapping year 1 which was from April 1999 through January 2001. If MapYear $=2$ is chosen, data are from MGCM model results driven by observed TES dust optical depth during TES mapping year 2 which was from February 2001 through December 2002. TES mapping year 1 had no global dust storm while TES mapping year 2 had a major, global-scale dust storm, peaking at $L_{S}=210$. The Map Year 1 and 2 data sets have a 1 km vertical resolution up to 10 km .

### 2.9 Auxiliary Profile Option

The auxiliary profile option provides the user with the choice to read and use any auxiliary profile of temperature and density versus altitude. In exercising the auxiliary profile MarsGRAM option, the values from the auxiliary profile replaces data from the original MGCM and MTGCM databases. This option is controlled by setting parameters profile, profnear, and proffar in the NAMELIST input file. Each line of the auxiliary profile input file consists of: (1) height, in
km , (2) latitude, in degrees, (3) longitude, in degrees, (4) temperature, in K , (5) pressure, in $\mathrm{N} / \mathrm{m}^{2}$, (6) density, in $\mathrm{kg} / \mathrm{m}^{3}$, (7) eastward wind, in $\mathrm{m} / \mathrm{s}$, and (8) northward wind, in $\mathrm{m} / \mathrm{s}$. Heights are relative to the MOLA areoid or reference ellipsoid, as set by input parameter MOLAhgts. Latitudes are planetocentric or planetographic, as set by input parameter ipclat. Longitudes are east or west, as set by input parameter LonEW. MGCM/MTGCM temperature, pressure, and density data are used if any of the profile inputs for temperature, pressure, or density are zero. MGCM/MTGCM winds are used if both wind components are zero on the profile file. A sample auxiliary profile file named profiledata.txt is provided with the Mars-GRAM 2010 distribution. Additional examples of auxiliary profiles include data from TES nadir or limb observations or Mars mesoscale model output at a particular location and time.

## 3. NEW FEATURES OF MARS-GRAM 2010

### 3.1 Mars-GRAM 2010 Adjustment Factors

### 3.1.1 Adjustment Factor Requirements

The adjustment factors generated by this process had to satisfy the gas law: $p=\rho \mathrm{RT}$, as well as the hydrostatic relation: $d p / d z=-\rho g$. If $T$ is assumed to be unchanged and both $p$ and $\rho$ are adjusted by a common factor, $F$, both relations are preserved. The adjustment factors $\left[F\left(z, L a t, L_{s}\right)\right]$ were expressed as a function of height $(z)$, latitude (Lat), and areocentric solar longitude $\left(L_{s}\right)$. This adjustment factor $(F)$ is applied to the daily mean density and pressure values from MGCM $(0-80 \mathrm{~km})$ and MTGCM (above 80 km ). The pressure scale height $(R T / g)$ is unchanged by this process. However, since the pressure has been changed by the adjustment factor, the height of the 1.26 nbar pressure level, referred to as $Z F$ in Mars-GRAM, has also been changed.

The daily mean MGCM or MTGCM density, DTA0, and the daily mean MGCM or MTGCM pressure, PTA0, depend on height (z), latitude (Lat), solar longitude ( $L_{s}$ ), dust amount $(\tau)$, and solar activity parameter (F10). The adjusted values of DTA0' and PTA0' are computed from the adjustment factors $(F)$ using the following equations:

$$
\begin{equation*}
\mathrm{DTA} 0^{\prime}=\mathrm{DTA} 0 * F\left(z, L a t, L_{s}\right) \tag{21}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{PTA}^{\prime}=\operatorname{PTA} 0 * F\left(z, L a t, L_{s}\right), \tag{22}
\end{equation*}
$$

where the adjustment factor $(F)$ has been determined as described above. Adjustment factors $(F)$ are also used to adjust $Z F$ by the relation:

$$
\begin{equation*}
Z F^{\prime}=Z F+H \ln (F) \tag{23}
\end{equation*}
$$

where $H$ is local pressure scale height.

### 3.1.2 Development of MTGCM Factors

Mars-GRAM density and pressure need to be consistent at 80 km , where the transition from MGCM to MTGCM data occurs. Thus, the assumption was made that $F\left(80, L a t, L_{s}\right)$ for the MTGCM data had to be the same as the adjustment factor at 80 km for the MGCM data. After adjustment factors $F\left(80, L a t, L_{s}\right)$ were determined from the MGCM analysis, they were used to determine MTGCM adjustment factors by use of the following equation:

$$
\begin{equation*}
F\left(z, L a t, L_{s}\right)=F\left(80, L a t, L_{s}\right) *\left(1+A \varsigma+B \varsigma^{2}\right), \tag{24}
\end{equation*}
$$

where the height parameter $\varsigma=(z-80)$ and the coefficients $A$ and $B$ depend on Lat and $L_{s}$.
Final adjustment factors $F\left(z, L a t, L_{s}\right)$ for MTGCM data were implemented into Mars-GRAM and a validation run comparing Mars-GRAM 2010 versus MGS, ODY, and MRO aerobraking data from the Planetary Data System (PDS) was completed. Any residual variation of aerobraking density about mean values that became apparent during this process was used to update the height dependence of Mars-GRAM perturbation standard deviations.

## 4. HOW TO RUN MARS-GRAM 2010

### 4.1 How to Obtain the Program

All source code and required data files are available as a downloadable zip file issued by the Marshall Space Flight Center Natural Environments Branch (EV44). The software is offered free of charge. To obtain the program source code and data files, refer to the contact information in the preface. See appendices $D$ through $G$ for summaries of the program and data files available in the downloaded file.

### 4.2 Running the Program

There are two ways to run Mars-GRAM: (1) as a subroutine in a user-provided main driver program (such as a trajectory program) and (2) as a stand-alone program, using a NAMELIST format input file, in which values for all input options are provided. To use Mars-GRAM as a subroutine, see discussion in appendix E and use example file dumytraj.f (available in the zip file distribution) as a guide. File README2.txt (available in the zip file distribution) also discusses use of dumytraj_M10.f90 as an example for using Mars-GRAM as a subroutine.

The steps involved in setting up and running Mars-GRAM in stand-alone mode are the following:

- To compile marsgram_M10.f90, dumytraj_M10.f90, and multtraj_M10.f90 under UNIX, to produce executable files marsgram_M10.x, dumytraj_M10.x, and multtraj_M10.exe, you can use the commands:
f90 Cfiles_M10_C.f90 Ifiles_M10_I.f90 marsgram_M10.x marsgram_M10.f90 marssubs_M10. f90/
setup_M10.f90 TESsubs_M10.f90
mv a.out marsgram_M10.x
erase *.o
erase *.mod
f90 Cfiles_M10_C.f90 Ifiles_M10_I.f90 dumytraj_M10.f90 marssubs_M10.f90 setup_M10.f90 /
TESsubs_M10.f90 wrapper_M10.f90
mv a.out dumytraj_M10.x
erase *.o
erase *.mod
f90 Cfiles_M10_C.f90 Ifiles_M10_I.f90 multtraj_M10.f90 marssubs_M10.f90 setup_M10.f90 / TESsubs_M10.f90 wrapper_M10.f90
mv a.out multtraj_M10.x
erase *.o
erase *.mod
- To compile marsgram_M10.f90, dumytraj_M10.f90, and multtraj_M10.f90 under personal computer Disk Operating System (PC-DOS) (for example, with gfortran), to produce executable files marsgram_M10.exe, dumytraj_M10.exe, and multtraj_M10.exe, you can use the example compile-and-link macros in the Code directory.
- To compile the auxiliary programs bldtraj.f90, finddate.f90, marsrad.f90, julday.f90, or radtraj. f90, or the binary conversion programs makebin.f90, readalb.f90, or READTOPO.f90 under UNIX, just use the FORTRAN compile statement for the specific auxiliary program source code file, i.e.
f90 -o auxiliary.x auxiliary.f90
or for PC-DOS, use the compile-and-link macros in the Utilities directory.
- Make sure that necessary data files albedol.txt (surface albedo data), COSPAR2.DAT (Committee on Space Research (COSPAR) model atmosphere data), MOLATOPH.TXT (MOLA topographic height information), and hgtoffst.dat (global average height offset values) are in an appropriate directory whose pathname is specified by parameter DATADIR in the NAMELIST format input file.
- Compile and run programs READTOPO.f90 to convert the MOLA data from ASCII to binary and readalb.f90 to convert the albedo data from ASCII to binary.
- Compile and run program makebin.f90 (see app. F) and convert the ASCII format MGCM and MTGCM data files provided to binary form (see app. D); this conversion process needs to be done only once on each user's machine.
- Make sure that the binary format MGCM and MTGCM data files (see apps. D and F) are in an appropriate directory whose pathname is specified by parameter $G C M D I R$ in the NAMELIST format input file.
- Prepare a NAMELIST format input file whose name is specified at run time with the desired values of all input options (example in app. B).
- If trajectory input mode rather than automatic profile mode is desired, prepare a trajectory input file whose name is set by parameter $T R A J F L$ in the NAMELIST input file containing time, height, latitude, and longitude further discussion of this is included below.
- If time-dependent coefficients for longitude-dependent wave model are to be used, prepare a file whose name is specified by parameter WaveFile in the NAMELIST format input file. This file contains one set of coefficients per line: time (seconds from start time) and wave model coefficients ( $B_{0}$ through $\Phi_{3}$, defined in sec. 2.3; further discussion below).
- Run the program by entering its executable name (e.g., marsgram); the program automatically opens and reads the NAMELIST input file, the TRAJFL file (if the trajectory mode is used), the data files albedo1.txt, COSPAR2.DAT, MOLATOPH.TXT and hgtoffst.dat, all MGCM and MTGCM binary data files, and the WaveFile file (if time-dependent coefficients are used).

If the program is run in profile mode, the user provides in the NAMELIST format input file fixed values for increments of time, height, latitude, and longitude. In this mode, the program automatically increments position until the desired number of positions (NPOS) are evaluated. In trajectory mode, selected by using $N P O S=0$, Mars-GRAM reads time and position information from the $T R A J F L$ file.

If constant values of longitude-dependent wave model coefficients are used, values for these are read in as part of the NAMELIST input file (sec. 3.3). For time-dependent coefficients, values are read from the WaveFile file. Each set of coefficients applies from the time given with the coefficient data, until a new time and set of coefficients are given on the next line of WaveFile. The last set of coefficients in WaveFile applies indefinitely, beginning with its given time.

### 4.3 Program Input

Appendix B gives a sample of a NAMELIST format input file for Mars-GRAM 2010. Whether the subroutine or stand-alone version is used, input variables whose values are supplied in the INPUT file are as follows:

LSTFL Name of LIST file; example of a LIST file is given in appendix C; for a listing to the console in the stand-alone version enter filename CON

OUTFL Name of OUTPUT file; a complete description of this file is contained in appendix A
$T R A J F L \quad$ (Optional) trajectory input file name; file contains time (seconds) relative to start time, height (km), latitude (degrees), longitude (degrees West if LonEW $=0$ or degrees East if $L o n E W=1$; see below)
profile (Optional) auxiliary profile input file name
WaveFile (Optional) input file for time-dependent wave coefficient data; see file description under parameter iuwave, below

DATADIR Pathname to directory for COSPAR data, topographic height data, surface albedo data, and global height offset data files

| GCMDIR | Pathname to directory for MGCM and MTGCM binary data files |
| :---: | :---: |
| IERT | 1 for time input as Earth-receive time (ERT) or 0 Mars-event time (MET) |
| IUTC | 1 for time input as Coordinated Universal Time (UTC), or 0 for Terrestrial (Dynamical) Time (TT) |
| MONTH | Integer month (1 through 12) for initial time |
| MDAY | Integer day of month for initial time |
| MYEAR | Integer year for starting time, a four-digit number; alternatively years 1970-2069 can be input as a two-digit number |
| NPOS | Maximum number of positions to evaluate, if an automatically-generated profile is to be produced; use 0 if trajectory positions are to be read in from a $T R A J F L$ file |
| IHR | Integer initial time, hour of day (ERT or MET, controlled by IERT value and UTC or TT, controlled by IUTC value) |
| IMIN | Integer initial time, minute of hour (meaning controlled by IERT and IUTC values) |
| SEC | Initial time, seconds of minute (meaning controlled by IERT and IUTC values) |
| LonEW | Longitude switch, 0 for input and output with West longitude positive (default) or 1 for East longitude positive |
| Dusttau | Optical depth of background dust level (no time-developing dust storm, just uniformly mixed dust), 0.1 to 3 (if 0 is input, a Viking-like annual variation of background dust is assumed) |
| Dustmin | Minimum seasonal dust $\tau$ if input Dusttau $=0(>=0.1)$ |
| Dustmax | Maximum seasonal dust $\tau$ if input Dusttau $=0(<=3)$ |
| Dustnu | Parameter for vertical distribution of dust density (Haberle et al. ${ }^{21}$ ) |
| Dustdiam | Dust particle diameter (micrometers, assumed monodisperse) |
| Dustdens | Dust particle density ( $\mathrm{kg} / \mathrm{m}^{3}$ ) |
| ALS0 | Value of areocentric longitude of the $\operatorname{Sun}\left(L_{s}\right.$, in degrees) at which a dust storm is to start; use a value of 0 if no dust storm is to be simulated; dust storm can be |

simulated only during the season of the Mars year for which $L_{s}$ is between $180^{\circ}$ and $320^{\circ}$
$A L S D U R \quad$ Duration in $L_{S}$ degrees for dust storm $($ default $=48)$
INTENS Dust storm intensity, measured as peak dust optical depth of the storm, with allowable values ranging from 0 (no dust storm) to 3 (maximum intensity dust storm). Dust storm intensity is added to background dust optical depth to give total dust optical depth

RADMAX Maximum radius (km) a dust storm can attain, developing according to the parameterized space and time profile of buildup and decay in the program; if a value of 0 or more than $10,000 \mathrm{~km}$ is used, the storm is taken to be of global dimensions and therefore uniformly covering the planet, but still assumed to build up and decay in intensity according to the same temporal profile

DUSTLAT Latitude (degrees, North positive) for center of dust storm

DUSTLON Longitude (degrees, West positive if LonEW $=0$, or East positive otherwise) for center of dust storm

MapYear $\quad 1$ or 2 for TES mapping year 1 or 2 GCM input data, or 0 for Mars-GRAM 2001 GCM input data sets
10.7 cm solar flux in its usual units of $10^{-22} \mathrm{~W} / \mathrm{cm}^{2}$ at average Earth orbit position ( 1 AU ); solar flux is automatically converted by the program to its value at the position of Mars in its orbit

NR1 Seed value (integer) for random number generator; allowable range is 1 to 29999; to do Monte Carlo simulations with a variety of perturbations, use a different random number seed on each model run; to repeat a given random number sequence on a later model run, use the same random number seed value
$N V A R X \quad x$-code for the plotable output ( $x-y$ pairs for line graphs or $x-y-z$ triplets for contour plots); appendix A lists the variables associated with the $x$-code (e.g., if NVAR $X=1, x$ output for plotting is height above the MOLA areoid)

NVARY $\quad y$-code for contour plot output ( $x-y-z$ triplets); use a $y$-code value of 0 for line graph ( $x-y$ pair) plots; appendix A lists $y$-code values and parameters represented

LOGSCALE Parameter to control units of output values of density and pressure on output plot files; a value of 0 means use regular density and pressure units $\left(\mathrm{kg} / \mathrm{m}^{3}\right.$ and $\mathrm{N} / \mathrm{m}^{2}$ ); 1 means to output logarithm (base-10) of the regular units; 2 means to output percentage deviation from COSPAR values of density and pressure; 3 means use SI units, with density in $\mathrm{kg} / \mathrm{km}^{3}$ (suitable for high altitudes)

| FLAT | Latitude of initial point to simulate (degrees, North positive) |
| :--- | :--- |
| $F L O N$ | Longitude of initial point to simulate (degrees, West positive if Lon $E W=0 ;$ |
|  | East positive otherwise) |

FHGT Height (km) of initial point to simulate above the reference ellipsoid; use FHGT $\leq-10$. km to specify that surface altitude should be used

MOLAhgts 1 for input heights relative to the MOLA areoid, otherwise input heights are relative to the old reference ellipsoid
hgtasfcm $\quad$ Height above surface ( $0-1,000 \mathrm{~m}$ ); use if $F H G T \leq-10 \mathrm{~km}$
zoffset Constant height offset ( km ) for MTGCM data or constant part of $L_{s}$-dependent (Bougher) height offset ( 0.0 means no constant offset). Positive offset increases density, negative offset decreases density
ibougher $\quad 0$ for no $L_{s}$-dependent (Bougher) height offset term; 1 means add $L_{s}$-dependent (Bougher) term, $-A^{*} \sin \left(L_{s}\right)(\mathrm{km})$, to constant term (zoffset), offset amplitude $A=2.5$ for MapYear $=0$ or 0.5 for MapYear $>0 ; 2$ means use global mean height offset from data file hgtoffst.dat; 3 means use daily average height offset at local position; 4 means use height offset at current time and local position. Value of zoffset is ignored if ibougher $=2$, 3, or 4

DELHGT Height increment (km) between successive steps in an automatically generated profile (positive upward)

DELLAT Latitude increment (degrees, Northward positive) between successive steps in an automatically generated profile

DELLON Longitude increment (degrees, Westward positive if LonEW = 0; Eastward positive if $L o n E W=1$ ) between successive steps in an automatically generated profile

DELTIME Time increment (s) between steps in an automatically generated profile
deltaTEX Additive adjustment to modify temperature (K) of the exosphere (asymptotic temperature approached at very high altitudes), nominal $=0$
profnear Latitude-longitude radius (degrees) within which weight for auxiliary profile is 1 (use profnear $=0.0$ for no profile input)
proffar Latitude-longitude radius (degrees) beyond which weight for auxiliary profile is 0.0
rpscale $\quad$ Random density perturbation scale factor ( $0-2,1=$ nominal $)$

| rwscale | Random wind perturbation scale factor ( $>=0$ ) |
| :---: | :---: |
| wlscale | Scale factor for perturbation wavelengths (0.1-10) |
| wmscale | Scale factor for mean winds |
| blwinfac | Scale factor for boundary layer slope winds ( $0=$ none $)$ |
| NMONTE | Number of Monte Carlo runs during one execution of the program; new/ different starting random numbers are automatically generated for each of the Monte Carlo profiles or trajectories |
| iup | Option controlling output of LIST file and graphics output files ( $0=$ none, other than 0 (default) indicates generate these files) |
| WaveA0 | Mean term of longitude-dependent wave multiplier for density |
| WaveDate | Julian date for (primary) peak(s) of wave (0 for no traveling component) |
| WaveA1 | Amplitude of wave- 1 component of longitude-dependent wave multiplier for density |
| Wavephil | Phase of wave-1 component of longitude-dependent wave multiplier (longitude, with West positive if $\operatorname{Lon} E W=0$, East positive otherwise) |
| phildot | Rate of longitude movement (degrees per day) for wave-1 component (Westward positive if $\operatorname{Lon} E W=0$, Eastward positive if $\operatorname{Lon} E W=1$ ) |
| WaveA2 | Amplitude of wave-2 component of longitude-dependent wave multiplier for density |
| Wavephi2 | Phase of wave-2 component of longitude-dependent wave multiplier (longitude, with West positive if Lon $E W=0$, East positive otherwise) |
| phi2dot | Rate of longitude movement (degrees per day) for wave-2 component (Westward positive if $\operatorname{Lon} E W=0$, Eastward positive if $\operatorname{Lon} E W=1$ ) |
| WaveA3 | Amplitude of wave-3 component of longitude-dependent wave multiplier for density |
| Wavephi3 | Phase of wave-3 component of longitude-dependent wave multiplier (longitude, with West positive if $\operatorname{Lon} E W=0$, East positive otherwise) |
| phi3dot | Rate of longitude movement (degrees per day) for wave- 3 component (Westward positive if $\operatorname{Lon} E W=0$, Eastward positive if $\operatorname{Lon} E W=1$ ) |


| iuwave | Unit number for (Optional) time-dependent wave coefficient data file named <br> by input parameter WaveFile (or 0 for none); WaveFile contains time (sec) rela- <br> tive to start time, and wave model coefficients (WaveA0 through Wavephi3) from <br> given time to next time in the data file |
| :--- | :--- |
| Wscale | Vertical scale (km) of longitude-dependent wave damping at altitudes below <br> $100 \mathrm{~km}(10 \leq W$ scale $\leq 10,000 \mathrm{~km})$ |
| corlmin | Minimum relative step size for perturbation updates (0.0-1.0); 0.0 means always <br> update perturbations, x.x means only update perturbations when corlim > x.x |
| ipclat | 1 for planetocentric latitude and height input, 0 for planetographic latitude <br> and height input |
| requa | Equatorial radius (km) for reference ellipsoid |
| rpole | Polar radius (km) for reference ellipsoid |
| idaydata | 1 for daily max/min data output; 0 for none |

Four auxiliary input files are required in addition to the MGCM and MTGCM input data files. The file MOLATOPH.TXT contains the MOLA areoid and topography data at $1 / 2$ by $1 / 2$ degree latitude-longitude resolution. Each line of MOLATOPH.TXT contains East longitude and latitude at the center of the $1 / 2$ by $1 / 2$ degree grid box, grid-box-average radius ( m ) to topographic surface, areoid radius ( m ) which is the radius to reference constant potential surface, evaluated at the center of the grid box, topography which is the grid-box-average difference ( m ) between local planetary radius and areoid, analogous to local terrain height above sea level for Earth, and Num, the number of MOLA laser measurements averaged over the grid box. MOLA latitudes are planetocentric. Longitudes in the MOLA input file are with respect to the IAU 1991 prime meridian. A shift (of about $0.24^{\circ}$ ) is made automatically within Mars-GRAM 2010, in order to convert to longitudes relative to the International Astronomical Union (IAU) 2000 prime meridian.

File COSPAR2.DAT contains COSPAR reference values of temperature, density, and pressure as a function of height. File hgtoffst.dat contains global values of MTGCM height offset versus $L_{s}$ and dust optical depth. File albedol.txt contains 1-degree resolution surface albedo data. Each line of albedol.txt contains latitude, and West longitude at the center of the 1 by 1 degree grid box, and grid-box-average surface albedo, the ratio of surface-reflected solar flux to that incident on the surface. Before Mars-GRAM is first used, the ASCII format files MOLATOPH.TXT and albedol.txt must be converted to binary with the auxiliary programs READTOPO.f90 and readalb.f90, that are supplied with Mars-GRAM 2010.

If the number of positions to be calculated, $N P O S$ is set to 0 , an optional trajectory input file is read from a file whose name is given by the input parameter TRAJFILE. Each line of the trajectory file consists of: time, in seconds past the start time specified in the NAMELIST input, height in km, latitude in degrees, and longitude in degrees. Heights are relative to the MOLA
areoid or reference ellipsoid, as set by the input parameter MOLAhgts. Latitudes are planetocentric or planetographic, as set by the input parameter ipclat. Longitudes are East or West, as set by the input parameter LonEW. Any additional information included on each line (e.g. orbit number, measured density, etc.) is ignored. Trajectory positions in these files do not have to be at small time or space steps. For example, a trajectory file may consist of successive periapsis times and positions for a simulated or observed aerobraking operation. Trajectory files may also contain arrays of locations used for computing height-latitude cross sections or latitude-longitude cross sections. Such trajectory input files can be as built by the program BLDTRAJ.f90.

If time-dependent wave parameters WaveA0 through Wavephi3 are desired, these are input from the file whose pathname is specified by the parameter Wavefile in the NAMELIST format input file. Parameter iuwave determines whether time-dependent WaveFile values are read or not. Iuwave $=0$ means no WaveFile data; otherwise iuwave is the WaveFile unit number. Each data line in the WaveFile file contains the time in seconds relative to start time, and wave model coefficients Wave $A 0$ through Wavephi3. Wave parameter values apply from the given time on each data line until the time given on the subsequent data line. Time-dependent wave parameters read in from WaveFile supercede any values given in the NAMELIST format input file.

### 4.4 Program Output

There are three general types of program output provided in Mars-GRAM. The first is a LIST file whose name is specified by the input parameter LSTFL and contains header and descriptor information that is suitable for printing or viewing by an analyst. An example LIST file is given in appendix C. Second is an OUTPUT file whose name is specified by the input parameter OUTFL which contains one header line and one line per output position and is suitable for reading into another program for additional analysis. A description of this file can be found in appendix A. Finally, there is a set of plotable output files, or graphics output files, that are text files suitable for input to a graphics program. Descriptions of these files can be found in appendix A.

The graphics output files contain either $x-y$ data pairs or $x-y-z$ data triplets, determined by the selected values for the input parameters $N V A R X$ and $N V A R Y$. If line-graph ( $x-y$ pair) data is the selected plot output option, then $N V A R Y=0$ is input. If contour plot ( $x-y-z$ triplet) data are the selected plot output option, then a non-zero value of $N V A R Y$ is input. Appendix A lists codes for $N V A R X$ and NVARY.

If the user desires to suppress the LIST, OUTPUT and graphics output files so that output can be handled in a user-provided program, they must set the LIST file unit number, iup, to 0 in the NAMELIST format input file. The unit number associated with the 'screen' output, iu0 or iustdout, normally 6 in the stand-alone version, can be set to any other value, by changing the assigned value of iustdout at program code line MGRM 21, and then recompiling the program.

## 5. SAMPLE RESULTS

### 5.1 Improvement of Mars-GRAM 2010 at Lower Altitudes

Application of adjustment factors to the Ames MGCM data yields improved comparisons between Mars-GRAM and TES limb data, as shown by the density ratios (Mars-GRAM/TES Limb) given in figure 1 . Prior to adjustment these density ratios were as low as 0.65 near 60 km .


Figure 1. Latitude-height contours of density ratio (Mars-GRAM/TES limb) after application of MGCM adjustment factors.

Mars-GRAM 2005 and Mars-GRAM 2010 MapYear $=0$ results have also been compared for three locations at local true solar time (LTST) 2 and 14:

- Location $1(\mathrm{~L} 1)=22.5^{\circ}$ S., $180^{\circ}$ E., $L_{S}=90 \pm 5, \tau=0.11$
- Location 2 (L2) $=22.5^{\circ}$ S., $180^{\circ}$ E., $L_{s}=75 \pm 5, \tau=0.12$
- Location $3(\mathrm{~L} 3)=2.5^{\circ} \mathrm{N} ., 180^{\circ} \mathrm{E}$., $L_{S}=210 \pm 5, \tau=2.65$ *dust storm case*

Figure 2 provides the density ratios of Mars-GRAM to TES for Mars-GRAM 2005.
As figure 3 shows, the application of the adjustment factor in Mars-GRAM 2010 results in ratios of $\approx 1$ at lower altitudes.


Figure 2. Density ratio (Mars-GRAM/TES) for Mars-GRAM 2005.


Figure 3. Density ratio (Mars-GRAM/TES) for Mars-GRAM 2010.

At the higher altitudes, Mars-GRAM 2010 results have corrected the effect of the underestimated dust aloft in the MGCM. At location 3, the Mars-GRAM 2010 density ratio has shifted closer to 1 . This demonstrates that the addition of adjustment factors to Mars-GRAM 2010 has improved the results for the MapYear $=0$ cases for large $\tau$ values.

### 5.2 Improvement of Mars-GRAM 2010 at Aerobraking Altitudes

Mars-GRAM modeled data output has improved at aerobraking altitudes by adding MTGCM adjustment factors which included height parameters and thermosphere coefficients. Improvement has been quantified by examining profile data density ratios for MGS, MRO, and ODY.

Taking the 99th percentile of all the density profiles illustrates the significant change the updated Mars-GRAM 2010 has on the profile density ratios. As shown in figure 4, the least amount of change was observed in the MGS data over the 99th percentile profile data, with an overall change of 2 units across the altitude range. The MRO data showed a significant improvement from the previous version of Mars-GRAM, reducing the higher altitude ratios from 6 to close to the optimal value of 1 on the updated data. However, the greatest change in ratio values occurred with the ODY data where the older data reached values close to 20 but the newer data brought the ratios down to a range between $<1$ to over 4 at the higher altitudes. All of the ratio values of the datasets improved from the old Mars-GRAM data output to the updated MarsGRAM 2010 version.


Figure 4. The 99th percentile density ratios of the profile data from MGS, MRO, and ODY to Mars-GRAM 2010 output versus height.

Figures 5 and 6 show the ratio of the observed density values to the Mars-GRAM output values for the old version and the updated Mars-GRAM 2010 version versus height and Mars latitude. Before the MTGCM adjustment factors, including thermosphere coefficients, were added to the Mars-GRAM code (fig. 5), the ratio values were higher than the optimal value of 1, especially at locations towards the poles. The contour lines are very tight near the poles, meaning lots of variability exists with the comparisons. In the updated plot shown in figure 6, a large area of the map is covered with the 1 ratio value, especially between $-30^{\circ} \mathrm{S}$. and $15^{\circ} \mathrm{N}$. Although a large discrepancy of ratio values still exists towards the poles, the variability has decreased with the inclusion of the adjustment factors.


Figure 5. Contour plots of the ratio of observed PDS density values to MarsGRAM output values (before adjustment) versus height and latitude.


Figure 6. Contour plots of the ratio of observed PDS density values to MarsGRAM 2010 output values (after adjustment) versus height and latitude.

## APPENDIX A—HEADERS FOR MARS-GRAM 2010 OUTPUT FILES

Mars-GRAM 2010 produces several output files suitable for passing to a graphics program for plotting and further analysis. Several of these files allow run-time selection from among several plotable parameters as the X parameter in an X-Y graph, or the X and Y parameters in an $\mathrm{X}-\mathrm{Y}-\mathrm{Z}$ graph. See the list of parameter selection codes at the end of this appendix. The graphics output file names and their descriptive headers are as follows:

## OUTPUT.txt (or other name, as prescribed in the NAMELIST INPUT file)

| Time | Time after initial input time (seconds) |
| :---: | :---: |
| Height | Planetocentric height $(\mathrm{km})$ above MOLA areoid $($ Height $=$ HgtMOLA $)$ OR planetocentric height $(\mathrm{km})$ above ellipsoid $($ Height $=$ HgtELPS) OR planetocentric height $(\mathrm{km})$ above local MOLA topographic surface (Height $=$ HgtSFCM) OR planetographic height ( km ) above ellipsoid (Height $=$ HgtGRPH), as determined by input parameters MOLAhgts, NVARX, NVARY, and ipclat |
| Lat | Planetocentric latitude $($ Lat $=$ LatPC $)$ or planetographic latitude $($ Lat $=$ LatPG $)$ in degrees (North positive) |
| LonW/LonE | Longitude (degrees, West positive or East positive) |
| Denkgm3 | Average (mean plus wave perturbed) density $\left(\mathrm{kg} / \mathrm{m}^{3}\right)$ or "Logkgm3" for $\log 10\left(\mathrm{~kg} / \mathrm{m}^{3}\right)$ or "Den $\%$ Avg" for percent deviation from COSPAR average, or "Denkgkm3" for kg / $\mathrm{km}^{3}$, depending on input value of $L O G S C A L E$ |
| Temp | Average temperature (K) |
| EWind | Eastward wind component (m/s, positive toward East) |
| NWind | Northward wind component ( $\mathrm{m} / \mathrm{s}$, positive toward North) |
| sigD | Standard deviation for density perturbations (\% of unperturbed mean) |
| Ls | Areocentric longitude of Sun from Mars (degrees) |
| Dust | Dust optical depth |
| LTST | Local true solar time (Mars hours) |
| CO2\%m | Carbon Dioxide mass concentration (\% by mass) |
| N2\%m | Nitrogen mass concentration (\% by mass) |


| $\mathrm{Ar} \% \mathrm{~m}$ | Argon mass concentration (\% by mass) |
| :--- | :--- |
| $\mathrm{O} 2 \% \mathrm{~m}$ | Molecular Oxygen mass concentration (\% by mass) |
| $\mathrm{CO} \% \mathrm{~m}$ | Carbon Monoxide mass concentration (\% by mass) |
| $\mathrm{O} \% \mathrm{~m}$ | Atomic Oxygen mass concentration (\% by mass) |
| $\mathrm{He} \% \mathrm{~m}$ | Helium mass concentration (\% by mass) |
| $\mathrm{H} 2 \% \mathrm{~m}$ | Molecular Hydrogen mass concentration (\% by mass) |
| $\mathrm{H} \% \mathrm{~m}$ | Atomic Hydrogen mass concentration (\% by mass) |
| $\mathrm{H} 2 \mathrm{O} \% \mathrm{~m}$ | Water vapor mass concentration (\% by mass) |
| DensP | Ratio of perturbed density to mean density |

## DayData.txt (Daily averages for heights below 1.26 nbar level)

Var_X User-selected plot variable determined by $N V A R X$ value

Var_Y
TempDay Local daily average temperature (K)
PresDay Local daily average pressure ( $\mathrm{N} / \mathrm{m}^{2}$ or as prescribed by LOGSCALE)
DensDay Local daily average density ( $\mathrm{kg} / \mathrm{m}^{3}$ or as prescribed by LOGSCALE)
EWwnDay Local daily average Eastward wind (m/s)
NSwnDay Local daily average Northward wind (m/s)
Tempmin Local daily minimum temperature (K)
Tempmax Local daily maximum temperature (K)
Densmin Local daily minimum density ( $\mathrm{kg} / \mathrm{m}^{3}$ or as prescribed by LOGSCALE)
Densmax Local daily maximum density ( $\mathrm{kg} / \mathrm{m}^{3}$ or as prescribed by LOGSCALE)
LOGSCALE Option controlling units of pressure and density output
DensAV Local density ( $\mathrm{kg} / \mathrm{m}^{3}$ or as prescribed by LOGSCALE)

## Density.txt

Var_X User-selected plot variable determined by NVARX value

Var_Y (Optional) user-selected plot variable from $N V A R Y$ value
DENSLO Low ( $\sim$ average - 1 standard deviation) density ( $\mathrm{kg} / \mathrm{m}^{3}$ or Log-10 or $\%$ from COSPAR or $\mathrm{kg} / \mathrm{km}^{3}$, as controlled by LOGSCALE)

DENSAV Average (mean plus wave-perturbed) density ( $\mathrm{kg} / \mathrm{m}^{3}$ or $\log -10$ or $\%$ from COSPAR, or $\mathrm{kg} / \mathrm{km}^{3}$, as controlled by LOGSCALE)

DENSHI High ( average +1 standard deviation) density ( $\mathrm{kg} / \mathrm{m}^{3}$ or Log -10 or $\%$ from COSPAR, or $\mathrm{kg} / \mathrm{km}^{3}$, as controlled by LOGSCALE)

DENSTOT Total (mean plus perturbed) density ( $\mathrm{kg} / \mathrm{m}^{3}$ or Log-10 or $\%$ from COSPAR, or $\mathrm{kg} / \mathrm{km}{ }^{3}$, as controlled by LOGSCALE)

DustOD Dust optical depth
Radius $\quad$ Radial distance from planetary center of mass to spacecraft position (areoid radius plus altitude)

Grav Local acceleration of gravity $\left(\mathrm{m} / \mathrm{s}^{2}\right)$
RadAU Mars orbital radius (Astronomical Units)
LOGSCALE Option controlling units of density output
Hgtoffset Local height offset (km) for MTGCM and MGCM data
Ibougher Input parameter controlling height offset option
MapYear TES mapping year (0 for Mars-GRAM 2001 data)
Profwgt Weight factor for auxiliary input profile data

## MarsRad.txt

Var_X User-selected plot variable (determined by NVARX value)
Var_Y (Optional) user-selected plot variable (from NVARY value)
alb Surface albedo (ratio upward/downward SW radiation at surface)
mu0 Cosine of solar zenith angle
Dareaden Dust column areal density $\left(\mathrm{kg} / \mathrm{m}^{2}\right)$
Dmixrat Dust mixing ratio (kg dust / kg air)
Dmasden Dust mass density (micrograms dust / m ${ }^{3}$ )
Dnumden Dust number density (number dust particles / $\mathrm{m}^{3}$ )

Ice $\quad$ Surface polar ice indicator $(0=$ no, $1=$ yes $)$

## File $=$ Perturb.txt

| Ice | Surface polar ice indicator $(0=$ no, $1=$ yes $)$ |
| :--- | :--- |
| Var_X | User-selected plot variable (determined by NVARX value) |
| Var_Y | (Optional) user-selected plot variable (from NVARY value) |
| SigD | Standard deviation of density perturbations (\% of unperturbed mean) |
| DensRand | Density perturbation from random model (\% of unperturbed mean) |
| DensWave | Density perturbation from wave model (\% of unperturbed mean) |
| DensP | Total density perturbation value (\% of unperturbed mean) |
| corlim | Fraction of minimum step size for accuracy of perturbations (should be $>1$ for insured <br> accuracy of perturbations) |
| SigU | Standard deviation of horizontal wind perturbations (m/s) |
| SigW | Standard deviation of vertical wind perturbations (m/s) |
| iupdate | 1 if perturbations updated, 0 if perturbations not updated but perturbation step <br> updated, -1 if neither perturbations nor step updated |

## ThrmData.txt (Thermospheric parameters for heights above 80 km )

| Var_X | User-selected plot variable (determined by NVARX value) |
| :--- | :--- |
| Var_Y | (Optional) user-selected plot variable (from NVARY value) |
| Tbase | Temperature at 1.26 nbar level (K) |
| Zbase | Altitude of 1.26 nbar level (km) |
| Flpeak | Altitude of F1 ionization peak (km) |
| MolWgt | Mean molecular weight (kg / kg.mole) |
| Texos | Exospheric temperature (K) |
| hgtoffset | Height offset for thermospheric (MTGCM) data (km) |
| ibougher | Input parameter controlling height offset option |

## TPresHgt.txt

| ibougher | Input parameter controlling height offset option |
| :---: | :---: |
| Var_X | User-selected plot variable (determined by $N V A R X$ value) |
| Var_Y | (Optional) user-selected plot variable (from NVARY value) |
| Temp | Mean temperature (K) |
| Pres | Mean (plus wave-perturbed) pressure ( $\mathrm{N} / \mathrm{m}^{2}$, or as controlled by LOGSCALE) |
| TdegC | Mean temperature ( ${ }^{\circ} \mathrm{C}$ ) |
| Pres_mb | Mean (plus wave-perturbed) pressure (mb) |
| Hrho | Density scale height (km) |
| Hpres | Pressure scale height (km) |
| MolWt | Molecular weight (kg/kg-mole) |
| TerHgt | Altitude of local surface above MOLA 1/2-degree areoid (km) |
| Tgrnd | Ground surface temperature (K) |
| Areoid | Local radius (km) of MOLA 1/2-degree areoid |
| dAreoid | MOLA areoid minus radius of old reference ellipsoid (km); equal to height from old ellipsoid minus height from MOLA areoid |
| CO2\% ${ }^{\text {v }}$ | Mole fraction (\%) Carbon Dioxide concentration (\% by volume) |
| N2\% ${ }^{\text {v }}$ | Mole fraction (\%) Nitrogen concentration (\% by volume) |
| Ar\%v | Mole fraction (\%) Argon concentration (\% by volume) |
| O2\%v | Mole fraction (\%) Molecular Oxygen concentration (\% by volume) |
| CO\%v | Mole fraction (\%) Carbon Monoxide concentration (\% by volume) |
| O\%v | Mole fraction (\%) Atomic Oxygen concentration (\% by volume) |
| He\%\% | Mole fraction (\%) Helium concentration (\% by volume) |
| H2\% ${ }^{\text {v }}$ | Mole fraction (\%) Molecular Hydrogen concentration (\% by volume) |
| H\%v | Mole fraction (\%) Atomic Hydrogen concentration (\% by volume) |
| H2O\%v | Mole fraction (\%) Water vapor concentration (\% by volume) |
| LOGSCALE | Option controlling units of pressure output |

## Winds.txt

| Var_X | User-selected plot variable (determined by NVARX value) |
| :--- | :--- |
| Var_Y | (Optional) user-selected plot variable (from NVARY value) |
| EWmean | Mean eastward wind component (m/s, positive eastward) |
| EWpert | Eastward wind perturbation (m/s) |
| EWtot | Total (mean plus perturbed) eastward wind (m/s) |
| NSmean | Mean northward wind component (m/s, positive northward) |
| NSpert | Northward wind perturbation (m/s) |
| NStot | Total (mean plus perturbed) northward wind (m/s) |
| VWpert | Vertical wind perturbation (m/s) |
| iupdate | 1 if perturbations updated, 0 if perturbations not updated but perturbation step <br> updated, -1 if neither perturbations nor step updated |

Model input codes used to select the plotable x and y parameters (Var_X and Var_Y) are as follows:

```
Code Parameter
```

1 Planetocentric height above local MOLA areoid (km)
2 Planetocentric height above local MOLA topographic surface (km)
3 Planetocentric latitude (degree (deg.))
4 Longitude (deg.) West+ if Lon $E W=0$, East + if LonEW $=1$
5 Time from start (Earth seconds)
6 Time from start (Martian Sols)
7 Areocentric Longitude of Sun, $\mathrm{L}_{\mathrm{s}}$ (deg.)
8 Local Solar Time (Mars hours)
9 Pressure (mb)
10 Pressure Height (km) $\left[-\mathrm{H}^{*} \log (\right.$ Pres $/$ PresSurf $)=-\mathrm{H}^{*} \log ($ sigma $\left.)\right]$
11 Sigma coordinate [sigma $=$ Pressure/(Pressure at Surface)]
12 Planetocentric Height (km) above reference ellipsoid (km)
13 Planetographic Height (km) above reference ellipsoid (km)
14 Planetographic Latitude (deg.)
15 Longitude in range -180 to +180 (East or West, controlled by LonEW)
Run-time selection of these plotable parameters is made by the input variables NVARX and NVARY contained in the NAMELIST format input file as described in section 4.3 and appendix B.

## APPENDIX B—EXAMPLE NAMELIST FORMAT INPUT FILE

The following is an example of the NAMELIST format input file required by Mars-GRAM 2010. Input data given here are provided as file inputstd0.txt. Values given are the default values assigned by the program. Only values that differ from the defaults actually have to be included in the NAMELIST file:

```
$INPUT_M10
LSTFL = 'LIST.txt'
OUTFL = 'OUTPUT.txt'
TRAJFL = 'TRAJDATA.txt'
profile = 'null'
WaveFile = 'null'
DATADIR = 'D:\Mars\Mars2010\Release1.0_Nov10\binFiles\'
GCMDIR = 'D:\Mars\Mars2010\Release1.0_Nov10\binFiles\'
IERT = 1
IUTC = 1
MONTH = 7
MDAY = 20
MYEAR = 76
NPOS = 41
IHR = 12
IMIN = 30
SEC = 0.0
LonEW = 0
Dusttau = 0.3
Dustmin = 0.3
Dustmax = 1.0
Dustnu = 0.003
Dustdiam = 5.0
Dustdens = 3000.
ALS0 = 0.0
ALSDUR = 48.
INTENS = 0.0
RADMAX = 0.0
DUSTLAT = 0.0
DUSTLON = 0.0
MapYear = 0
F107 = 68.0
NR1 = 1234
NVARX = 1
NVARY = 0
LOGSCALE = 0
FLAT = 22.48
FLON = 47.97
FHGT = -5.
MOLAhgts = 1
hgtasfcm = 0.
zoffset = 3.25
ibougher = 1
DELHGT = 5.0
DELLAT = 0.5
```

DELLON $=0.5$
DELTIME $=500.0$
deltaTEX $=0.0$
profnear $=0.0$
proffar $=0.0$
rpscale $=1.0$
rwscale $=1.0$
wlscale $=1.0$
wmscale $=1.0$
blwinfac $=1.0$
NMONTE $=1$
iup $=13$
WaveA0 $=1.0$
WaveDate $=0.0$
WaveA1 $=0.0$
Wavephil $=0.0$
phildot $=0.0$
WaveA2 $=0.0$
Wavephi2 $=0.0$
phi2dot $=0.0$
WaveA3 $=0.0$
Wavephi3 $=0.0$
phi3dot $=0.0$
iuwave $=0$
Wscale $=20$.
corlmin $=0.0$
ipclat $=1$
requa $=3396.19$
rpole $=3376.20$
idaydata $=1$
\$END

Explanation of variables:
LSTFL = List file name (CON for console listing)
OUTFL = Output file name
TRAJFL $=$ (Optional) Trajectory input file. File contains time (sec) relative to start time, height ( km ), latitude (deg), longitude (deg W if LonEW=0, deg E if LonEW=1, see below)
profile $=($ Optional $)$ auxiliary profile input file name
WaveFile $=($ Optional $)$ file for time-dependent wave coefficient data. See file description under parameter iuwave, below.
DATADIR $=$ Directory for COSPAR data and topographic height data
GCMDIR = Directory for GCM binary data files
IERT $=1$ for time input as Earth-Receive time (ERT) or 0 Mars-event time (MET)
IUTC $=1$ for time input as Coordinated Universal Time (UTC), or 0 for Terrestrial (Dynamical) Time (TT)
MONTH $=($ Integer $)$ month of year
MDAY $=($ Integer $)$ day of month
MYEAR $=$ (Integer) year (4-digit; 1970-2069 can be 2-digit)
NPOS $=$ max \# positions to evaluate $(0=$ read data from trajectory input file)
IHR = Hour of day (ERT or MET, controlled by IERT and UTC or TT, controlled by IUTC)
IMIN = minute of hour (meaning controlled by IERT and IUTC)
SEC $=$ seconds of minute (meaning controlled by IERT and IUTC). IHR:IMIN:SEC is time for initial position to be evaluated
LonEW $=0$ for input and output West longitudes positive; 1 for East longitudes positive

Dusttau $=$ Optical depth of background dust level (no time-developing dust storm, just uniformly mixed dust), 0.1 to 3.0 , or use 0 for assumed seasonal variation of background dust
Dustmin $=$ Minimum seasonal dust tau if input Dusttau $=0(>=0.1)$
Dustmax $=$ Maximum seasonal dust tau if input Dusttau=0 $(<=1.0)$
Dustnu = Parameter for vertical distribution of dust density (Haberle et al., J. Geophys. Res., 104, 8957, 1999)
Dustdiam = Dust particle diameter (micrometers, assumed monodisperse)
Dustdens $=$ Dust particle density ( $\mathrm{kg} / \mathrm{m}^{* *} 3$ )
ALS0 $=$ starting Ls value (degrees) for dust storm ( $0=$ none)
ALSDUR $=$ duration (in Ls degrees) for dust storm $($ default $=48)$
INTENS $=$ dust storm intensity ( $0.0-3.0$ )
RADMAX $=$ max. radius $(\mathrm{km})$ of dust storm $(0$ or $>10000=$ global $)$
DUSTLAT $=$ Latitude (degrees) for center of dust storm
DUSTLON $=$ Longitude (degrees) (West positive if LonEW=0, or East positive if LonEW = 1) for center of dust storm
MapYear $=1$ or 2 for TES mapping year 1 or 2 GCM input data, or 0 for Mars-GRAM 2001 GCM input data sets
F107 $=10.7 \mathrm{~cm}$ solar flux $\left(10^{* *}-22 \mathrm{~W} / \mathrm{cm}^{* * 2}\right.$ at 1 AU$)$
NR1 = starting random number $(0<\mathrm{NR} 1<30000)$
NVARX = x-code for plotable output ( $1=$ hgt above MOLA areoid). See file xycodes.txt
NVARY = y-code for 3-D plotable output ( 0 for 2-D plots)
LOGSCALE $=0=$ regular SI units, $1=$ log-base-10 scale, $2=$ percentage deviations from COSPAR model, $3=$ SI units, with density in $\mathrm{kg} / \mathrm{km}^{* * 3}$ (suitable for high altitudes)
FLAT = initial latitude ( N positive), degrees
FLON = initial longitude (West positive if LowEW $=0$ or East positive if LonEW =1), degrees
FHGT = initial height (km); <=-10 means evaluate at surface height; $>3000 \mathrm{~km}$ means planeto-centric radius
MOLAhgts $=1$ for input heights relative to MOLA areoid, otherwise input heights are relative to reference ellipsoid
hgtasfcm $=$ height above surface $(0-4500 \mathrm{~m})$; use if FHGT $<=-10 . \mathrm{km}$
zoffset $=$ constant height offset $(\mathrm{km})$ for MTGCM data or constant part of Ls-dependent (Bougher) height offset ( 0.0 means no constant offset). Positive offset increases density, negative offset decreases density.
ibougher $=0$ for no Ls-dependent (Bougher) height offset term; 1 means add Ls-dependent (Bougher) term, -A*Sin(Ls) (km), to constant term (zoffset) [offset amplitude A $=2.5$ for MapYear $=0$ or 0.5 for MapYear > 0]; 2 means use global mean height offset from data file hgtoffst.dat; 3 means use daily average height offset at local position; 4 means use height offset at current time and local position. Value of zoffset is ignored if ibougher $=2,3$, or 4 .
DELHGT $=$ height increment $(\mathrm{km})$ between steps
DELLAT $=$ Latitude increment (deg) between steps (Northward positive)
DELLON = Longitude increment (deg) between steps (Westward positive if LonEW = 0, Eastward positive if LonEW = 1)
DELTIME $=$ time increment (sec) between steps
deltaTEX = adjustment for exospheric temperature (K)
profnear $=$ Lat-lon radius (degrees) within which weight for auxiliary profile is 1.0 (Use profnear $=0.0$ for no profile input)
proffar $=$ Lat-lon radius (degrees) beyond which weight for auxiliary profile is 0.0
rpscale $=$ random density perturbation scale factor (0-2)
rwscale $=$ random wind perturbation scale factor ( $>=0$ )
wlscale $=$ scale factor for perturbation wavelengths (0.1-10)
$\mathrm{wmscale}=$ scale factor for mean winds
blwinfac $=$ scale factor for boundary layer slope winds $(0=$ none $)$

NMONTE = number of Monte Carlo runs
iup $=0$ for no LIST and graphics output, or unit number for output
WaveA0 = Mean term of longitude-dependent wave multiplier for density
WaveDate = Julian date for (primary) peak(s) of wave (0 for no traveling component)
WaveA1 = Amplitude of wave-1 component of longitude-dependent wave multiplier for density
Wavephil $=$ Phase of wave-1 component of longitude-dependent wave multiplier (longitude, with West positive if LonEW $=0$, East positive if LonEW = 1)
phildot $=$ Rate of longitude movement (degrees per day) for wave-1 component (Westward positive if LonEW $=0$, Eastward positive if LonEW = 1)
WaveA2 = Amplitude of wave-2 component of longitude-dependent wave multiplier for density
Wavephi2 $=$ Phase of wave-2 component of longitude-dependent wave multiplier (longitude, with West positive if LonEW $=0$, East positive if LonEW = 1)
phi2dot $=$ Rate of longitude movement (degrees per day) for wave-2 component (Westward positive if LonEW $=0$, Eastward positive if LonEW = 1)
WaveA3 = Amplitude of wave-3 component of longitude-dependent wave multiplier for density
Wavephi3 = Phase of wave-3 component of longitude-dependent wave multiplier (longitude, with West positive if LonEW $=0$, East positive if LonEW = 1)
phi3dot $=$ Rate of longitude movement (degrees per day) for wave-3 component (Westward positive if LonEW $=0$, Eastward positive if LonEW = 1)
iuwave $=$ Unit number for (Optional) time-dependent wave coefficient data file "WaveFile" (or 0 for none). WaveFile contains time (sec) relative to start time, and wave model coefficients (WaveA0 thru Wavephi3) from the given time to the next time in the data file.
Wscale $=$ Vertical scale (km) of longitude-dependent wave damping at altitudes below $100 \mathrm{~km}(10<=$ Wscale $<=10,000 \mathrm{~km})$
corlmin $=$ minimum relative step size for perturbation updates ( $0.0-1.0$ ); 0.0 means always update perturbations, x.x means only update perturbations when corlim > x.x
ipclat $=1$ for Planeto-centric latitude and height input, 0 for Planeto-graphic latitude and height input
requa $=$ Equatorial radius $(\mathrm{km})$ for reference ellipsoid
rpole $=$ Polar radius ( km ) for reference ellipsoid
idaydata $=1$ for daily max $/ \mathrm{min}$ data output; 0 for none

## APPENDIX C—SAMPLE OUTPUT LIST FILE

Following is the LIST file output produced by the standard input parameters given in appendix B. Standard input files for MapYear $=0,1$, and 2 are provided to users as the files inputstd0.txt, inputstd1.txt, and inputstd2.txt. The output data given below are provided in the file ListMapYr0.txt. Output files that result from the use of inputstd1.txt and inputstd2.txt are provided to users in the files ListMapYrl.txt and ListMapYr2.txt. Availability of these files allows users to complete a test run after compiling Mars-GRAM 2010 on their own computer and to electronically check their output by a file-compare process (e.g. the 'diff' command in UNIX or the 'fc' command in DOS). Please note that, due to machine-dependent or compiler-dependent rounding differences, some output values may differ slightly from those shown here. These differences are usually no more than one unit in the last significant digit displayed:

```
Mars-GRAM 2010 (Version 1.0) - Nov 2010
LIST file= LIST.txt
OUTPUT file= OUTPUT.txt
Data directory= D:\Mars\Mars2010\Release1.0_Nov10\binFiles\
GCM directory= D:\Mars\Mars2010\Release1.0_Nov10\binFiles\
Input time is Earth-Receive Time (ERT)
Input time is Coordinated Universal Time (UTC)
Date = 7/20/1976 Julian Day =2442980.02083 Time = 12:30: 0.0
Input heights are planeto-centric, relative to MOLA areoid
Reference ellipsoid radii (km): Equator = 3396.19 Pole =3376.20
Output heights are planeto-centric, except as noted.
Longitude & ephemeris use IAU 2000 rotational system.
F10.7 flux = 68.0 (1 AU) 25.0 (Mars)
Dust optical depth from NAMELIST input
Dustnu = 0.0030 Dustdiam = 5.00 E-6 meters Dustdens = 3000.0 kg/m**3
Random seed = 1234 Dens.Pert.Scale Factor = 1.00 corlmin = 0.000
Wind.Pert.Scale Factor = 1.00 Wavelength Scale Factor = 1.00
Mean Wind Scale Factor = 1.00 Slope Wind Scale Factor = 1.00
A0,A1,phi1,A2,phi2,A3,phi3=1.000 0.000 0.0 0.000 0.0 0.000 0.0
Wave Scale = 20.0 km. Wave phases are in degrees of West Longitude
Time (rel. to T0) = 0.0 sec. ( 0.000 sols) Ls =97.0 Dust =0.30
Height Above MOLA (or Surface) = -5.000 km (-1.372 km) OWLT = 19.01 Min
Topographic Height = -3.628 km Radius (Areoid) = 3387.954 (3392.954) km
Hgt Above Ellipsoid = -5.292 km Scale Hgt H(p)=10.94 H(rho)=10.94 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 0.000 km
Planeto-Centric Lat = 22.48 deg Longitude = 47.97 W (312.03 E) deg.
Planeto-Graphic Lat = 22.72 deg Planeto-Graphic Hgt (Ellps)= -5.292 km
Planeto-Cent Sun Lat = 25.02 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 108.69 deg.W Local True Solar Time = 16.05 Mars hrs
Temperature = 241.8 K Pressure = 8.953E +02 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 1.922E-02 1.960E-02 1.999E-02 kg/m**3
Departure, COSPAR NH Mean = -21.6 % -20.0 % - -18.4 % iupdate = 1
Tot.Dens. =2.006E-02 kg/m**3 Dens.Pert. = 2.34% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = 0.0-1.1 -1.1 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = 0.0-0.4-0.4 m/s -0.2 m/s
CO2=2.577E+23 N2=7.809E+21 Ar=4.685E+21 O2=3.692E+20 CO=2.840E+20 #/m**3
    96.084
    94.445 2.862 1.717 0.135 0.104 % by volume
```

```
H2O=2.025E+21 O=0.000E +00 He=0.000E+00 H2=0.000E+00 Total=2.729E+23 #/m**3
    0.308 0.000 0.000 0.000 % by mass MolWgt=43.259
    0.742 0.000 0.000 0.000 % volume (mole) fraction
```

Time (rel. to T0) $=500.0 \mathrm{sec} .(0.006$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=0.000 \mathrm{~km}(3.585 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.585 \mathrm{~km}$ Radius $($ Areoid $)=3392.852(3392.852) \mathrm{km}$
Hgt Above Ellipsoid $=-0.268 \mathrm{~km}$ Scale Hgt H(p) $=10.85 \mathrm{H}(\mathrm{rho})=12.96 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=22.98$ deg Longitude $=48.47 \mathrm{~W}(311.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=23.22$ deg Planeto-Graphic Hgt $($ Ellps $)=-0.268 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=110.72$ deg. W Local True Solar Time $=16.15$ Mars hrs
Temperature $=227.5 \mathrm{~K}$ Pressure $=5.670 \mathrm{E}+02 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=1.294 \mathrm{E}-021.320 \mathrm{E}-02 \quad 1.346 \mathrm{E}-02 \mathrm{~kg} / \mathrm{m}^{*}{ }^{*} 3$
Departure, COSPAR NH Mean $=-16.5 \%-14.9 \% \quad-13.2 \%$ iupdate $=1$
Tot.Dens. $=1.338 \mathrm{E}-02 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=1.41 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=2.82 .95 .7 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-7.9-2.5-10.4 \mathrm{~m} / \mathrm{s}-0.7 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.739 \mathrm{E}+23 \mathrm{~N} 2=5.267 \mathrm{E}+21 \mathrm{Ar}=3.160 \mathrm{E}+21 \mathrm{O} 2=2.490 \mathrm{E}+20 \mathrm{CO}=1.915 \mathrm{E}+20 \mathrm{\#} / \mathrm{m}^{* *} 3$

| 96.290 | 1.857 | 1.589 | 0.100 | 0.068 |
| :--- | :--- | :--- | :--- | :--- |$\%$ by mass

$94.928 \quad 2.875 \quad 1.725 \quad 0.136 \quad 0.105 \%$ by volume
$\mathrm{H} 2 \mathrm{O}=4.244 \mathrm{E}+20 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=1.832 \mathrm{E}+23$ \#/m**3
$\begin{array}{llll}0.096 & 0.000 & 0.000 & 0.000\end{array}$ \% by mass MolWgt=43.386
$\begin{array}{llll}0.232 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}$

```
Time (rel. to T0) = 1000.0 sec. ( 0.011 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 5.000 km ( 8.545 km) OWLT = 19.01 Min
Topographic Height = -3.545 km Radius (Areoid) = 3397.749 (3392.749) km
Hgt Above Ellipsoid = 4.756 km Scale Hgt H(p)= 9.92 H(rho)= 11.27 km
Height Offset Parameters: ibougher =1 Local Height Offset = 0.000 km
Planeto-Centric Lat = 23.48 deg Longitude = 48.97 W (311.03 E) deg.
Planeto-Graphic Lat = 23.73 deg Planeto-Graphic Hgt (Ellps)= 4.756 km
Planeto-Cent Sun Lat =25.02 deg Mars Orbital Radius =1.649 AU
Sun Longitude = 112.74 deg.W Local True Solar Time = 16.25 Mars hrs
Temperature = 211.0 K Pressure = 3.577E+02 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 8.793E-03 8.969E-03 9.148E-03 kg/m**3
Departure, COSPAR NH Mean = -11.2 % -9.4 % -7.6 % iupdate = 1
Tot.Dens. = 8.701E-03 kg/m**3 Dens.Pert. = -2.99% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = 2.5 -1.9 0.6 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = 3.0 1.3 4.3 m/s -0.3 m/s
CO2=1.183E+23 N2=3.581E+21 Ar=2.149E+21 O2=1.693E+20 CO=1.302E+20 #/m**3
    96.366
    95.107 2.880 1.728 0.136 0.105 % by volume
H2O=5.525E+19 O=0.000E+00 He=0.000E E 00 H2=0.000E+00 Total=1.244E+23 #/m**3
    0.018 0.000 0.000 0.000 % by mass MolWgt=43.434
    0.044 0.000 0.000 0.000% volume (mole) fraction
```

Time (rel. to T0) $=1500.0 \mathrm{sec} .(0.017 \mathrm{sols}) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=10.000 \mathrm{~km}(13.513 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.513 \mathrm{~km}$ Radius $($ Areoid $)=3402.642(3392.642) \mathrm{km}$
Hgt Above Ellipsoid $=9.778 \mathrm{~km}$ Scale Hgt H(p) $=9.48 \mathrm{H}(\mathrm{rho})=10.60 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=23.98$ deg Longitude $=49.47 \mathrm{~W}(310.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=24.23$ deg Planeto-Graphic Hgt $($ Ellps $)=9.778 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=114.77$ deg.W Local True Solar Time $=16.35$ Mars hrs
Temperature $=198.7 \mathrm{~K} \quad$ Pressure $=2.161 \mathrm{E}+02 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=5.624 \mathrm{E}-035.753 \mathrm{E}-035.885 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-13.1 \%-11.1 \% \quad-9.0 \%$ iupdate $=1$

Tot.Dens. $=5.749 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=-0.07 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total) $=-16.4-0.9-17.3 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=9.52 .011 .4 \mathrm{~m} / \mathrm{s}-0.8 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=7.588 \mathrm{E}+22 \mathrm{~N} 2=2.297 \mathrm{E}+21 \mathrm{Ar}=1.378 \mathrm{E}+21 \mathrm{O} 2=1.086 \mathrm{E}+20 \mathrm{CO}=8.352 \mathrm{E}+19 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llllll}96.380 & 1.857 & 1.589 & 0.100 & 0.068 & \%\end{array}$
$95.139 \quad 2.880 \quad 1.728 \quad 0.136 \quad 0.105 \%$ by volume
$\mathrm{H} 2 \mathrm{O}=9.639 \mathrm{E}+18 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=7.975 \mathrm{E}+22$ \#/m**3
$\begin{array}{llll}0.005 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt $=43.442$
$0.012 \quad 0.000 \quad 0.000 \quad 0.000 \%$ volume (mole) fraction
Time (rel. to T0) $=2000.0 \mathrm{sec} .(0.023$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=15.000 \mathrm{~km}(18.545 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.545 \mathrm{~km}$ Radius $($ Areoid $)=3407.533(3392.533) \mathrm{km}$
Hgt Above Ellipsoid $=14.800 \mathrm{~km}$ Scale Hgt H(p) $=9.08$ H(rho) $=10.13 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=24.48 \mathrm{deg}$ Longitude $=49.97 \mathrm{~W}(310.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=24.73$ deg Planeto-Graphic Hgt $($ Ellps $)=14.800 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=116.80$ deg.W Local True Solar Time $=16.46$ Mars hrs
Temperature $=188.3 \mathrm{~K}$ Pressure $=1.276 \mathrm{E}+02 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) = 3.494E-03 3.587E-03 3.682E-03 kg/m**3
Departure, COSPAR NH Mean $=-16.2 \%-14.0 \%-11.7 \%$ iupdate $=1$
Tot.Dens. $=3.597 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=0.28 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-19.7-6.0-25.7 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=10.80 .211 .0 \mathrm{~m} / \mathrm{s} \quad 1.4 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=4.731 \mathrm{E}+22 \mathrm{~N} 2=1.431 \mathrm{E}+21 \mathrm{Ar}=8.589 \mathrm{E}+20 \mathrm{O} 2=6.767 \mathrm{E}+19 \mathrm{CO}=5.205 \mathrm{E}+19 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llllll}96.385 & 1.857 & 1.589 & 0.100 & 0.067 & \%\end{array}$
$95.149 \quad 2.879 \quad 1.727 \quad 0.136 \quad 0.105 \%$ by volume
$\mathrm{H} 2 \mathrm{O}=1.827 \mathrm{E}+18 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total=4.972E+22 \#/m**3
$\begin{array}{llll}0.002 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt $=43.445$
$0.004 \quad 0.000 \quad 0.000 \quad 0.000 \%$ volume (mole) fraction
Time (rel. to T0) $=2500.0 \mathrm{sec} .(0.028$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=20.000 \mathrm{~km}(23.571 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.571 \mathrm{~km}$ Radius $($ Areoid $)=3412.420(3392.420) \mathrm{km}$
Hgt Above Ellipsoid $=19.821 \mathrm{~km}$ Scale Hgt H(p) $=8.69 \mathrm{H}(\mathrm{rho})=9.71 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=24.98$ deg Longitude $=50.47 \mathrm{~W}(309.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=25.24$ deg Planeto-Graphic Hgt $($ Ellps $)=19.821 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=118.83$ deg. W Local True Solar Time $=16.56$ Mars hrs
Temperature $=178.1 \mathrm{~K} \quad$ Pressure $=7.374 \mathrm{E}+01 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=2.125 \mathrm{E}-032.190 \mathrm{E}-032.257 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{*} * 3$
Departure, COSPAR NH Mean = -19.2 \% -16.7 \% -14.2 \% iupdate = 1
Tot.Dens. $=2.229 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=1.78 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-30.6-0.7-31.3 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=8.2-6.12 .1 \mathrm{~m} / \mathrm{s}-0.3 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=2.888 \mathrm{E}+22 \mathrm{~N} 2=8.737 \mathrm{E}+20 \mathrm{Ar}=5.242 \mathrm{E}+20 \mathrm{O} 2=4.130 \mathrm{E}+19 \mathrm{CO}=3.177 \mathrm{E}+19 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llllll}96.388 & 1.856 & 1.588 & 0.100 & 0.067 \% \text { by mass }\end{array}$
$95.153 \quad 2.878 \quad 1.727 \quad 0.136 \quad 0.105 \%$ by volume
$\mathrm{H} 2 \mathrm{O}=3.014 \mathrm{E}+17 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=3.036 \mathrm{E}+22$ \#/m**3
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt=43.446
$\begin{array}{llll}0.001 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}$
Time $($ rel. to T 0$)=3000.0 \mathrm{sec} .(0.034$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=25.000 \mathrm{~km}(28.359 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.359 \mathrm{~km}$ Radius $($ Areoid $)=3417.305(3392.305) \mathrm{km}$
Hgt Above Ellipsoid $=24.841 \mathrm{~km}$ Scale Hgt H(p) $=8.34 \mathrm{H}(\mathrm{rho})=9.43 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=25.48 \mathrm{deg}$ Longitude $=50.97 \mathrm{~W}(309.03 \mathrm{E}) \mathrm{deg}$.

Planeto-Graphic Lat $=25.74$ deg Planeto-Graphic Hgt $($ Ellps $)=24.841 \mathrm{~km}$ Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=120.85$ deg.W Local True Solar Time $=16.66$ Mars hrs
Temperature $=168.0 \mathrm{~K} \quad$ Pressure $=4.160 \mathrm{E}+01 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=1.265 \mathrm{E}-031.310 \mathrm{E}-031.356 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{*} * 3$
Departure, COSPAR NH Mean = -21.9 \% -19.1 \% -16.3 \% iupdate = 1
Tot.Dens. $=1.302 \mathrm{E}-03 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=-0.63 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-42.13 .7-38.5 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=9.55 .414 .9 \mathrm{~m} / \mathrm{s} \quad 1.3 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.728 \mathrm{E}+22 \mathrm{~N} 2=5.224 \mathrm{E}+20 \mathrm{Ar}=3.134 \mathrm{E}+20 \mathrm{O} 2=2.469 \mathrm{E}+19 \mathrm{CO}=1.900 \mathrm{E}+19 \mathrm{\#} / \mathrm{m}^{* *} 3$
$\begin{array}{llllll}96.389 & 1.855 & 1.588 & 0.100 & 0.067 \% \text { by mass }\end{array}$
$95.156 \quad 2.877 \quad 1.726 \quad 0.136 \quad 0.105 \%$ by volume
$\mathrm{H} 2 \mathrm{O}=4.024 \mathrm{E}+16 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=1.816 \mathrm{E}+22$ \#/m**3
$\begin{array}{lllll}0.000 & 0.000 & 0.000 & 0.000 & \%\end{array}$ by mass MolWgt=43.446
$0.000 \quad 0.000 \quad 0.000 \quad 0.000 \%$ volume (mole) fraction
Time $($ rel. to T0) $=3500.0 \mathrm{sec} .(0.039 \mathrm{sols}) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=30.000 \mathrm{~km}(33.384 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.384 \mathrm{~km}$ Radius $($ Areoid $)=3422.188(3392.188) \mathrm{km}$
Hgt Above Ellipsoid $=29.861 \mathrm{~km}$ Scale Hgt H(p) $=8.01 \mathrm{H}(\mathrm{rho})=8.72 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=25.98$ deg Longitude $=51.47 \mathrm{~W}(308.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=26.25$ deg Planeto-Graphic Hgt $($ Ellps $)=29.861 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02 \mathrm{deg}$ Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=122.88$ deg. W Local True Solar Time $=16.76$ Mars hrs
Temperature $=157.1 \mathrm{~K}$ Pressure $=2.293 \mathrm{E}+01 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=7.418 \mathrm{E}-047.719 \mathrm{E}-048.032 \mathrm{E}-04 \mathrm{~kg} / \mathrm{m}^{*} * 3$
Departure, COSPAR NH Mean $=-24.3 \%-21.2 \%-18.0 \%$ iupdate $=1$
Tot.Dens. $=7.296 \mathrm{E}-04 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=-5.47 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-60.3-4.5-64.8 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind (Mean,Perturbed,Total) $=-4.5-4.2-8.8 \mathrm{~m} / \mathrm{s}-0.0 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.018 \mathrm{E}+22 \mathrm{~N} 2=3.077 \mathrm{E}+20 \mathrm{Ar}=1.846 \mathrm{E}+20 \mathrm{O} 2=1.455 \mathrm{E}+19 \mathrm{CO}=1.119 \mathrm{E}+19 \mathrm{\#} / \mathrm{m}^{* *} 3$
$\begin{array}{llllll}96.391 & 1.855 & 1.587 & 0.100 & 0.067 \%\end{array}$
$\begin{array}{lllll}95.158 & 2.876 & 1.726 & 0.136 & 0.105\end{array} \%$ by volume
$\mathrm{H} 2 \mathrm{O}=3.427 \mathrm{E}+15 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=1.070 \mathrm{E}+22$ \#/m**3
$\begin{array}{lllll}0.000 & 0.000 & 0.000 & 0.000 & \%\end{array}$ by mass MolWgt=43.446
$0.000 \quad 0.000 \quad 0.000 \quad 0.000 \%$ volume (mole) fraction

```
Time (rel. to T0) = 4000.0 sec. ( 0.045 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 35.000 km ( }38.545\textrm{km})\mathrm{ OWLT = 19.01 Min
Topographic Height = -3.545 km Radius (Areoid)=3427.069 (3392.069) km
Hgt Above Ellipsoid = 34.882 km Scale Hgt H(p)= 7.76 H(rho)= 8.72 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 0.000 km
Planeto-Centric Lat = 26.48 deg Longitude = 51.97 W (308.03 E) deg.
Planeto-Graphic Lat = 26.75 deg Planeto-Graphic Hgt (Ellps)= 34.881 km
Planeto-Cent Sun Lat = 25.02 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 124.91 deg.W Local True Solar Time = 16.86 Mars hrs
Temperature = 149.4 K Pressure = 1.234E+01 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 4.175E-04 4.371E-04 4.576E-04 kg/m**3
Departure, COSPAR NH Mean = -28.3% -24.9% -21.4% iupdate = 1
Tot.Dens. = 4.215E-04 kg/m**3 Dens.Pert. = -3.58% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) =-84.4 4.2-80.2 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -14.5 4.7 -9.7 m/s -0.5 m/s
CO2=5.765E+21 N2=1.742E+20 Ar=1.045E+20 O2=8.234E+18 CO=6.334E+18 #/m**3
    96.392
    95.160
H2O=4.724E+14 O=0.000E +00 He=0.000E+00 H2=0.000E+00 Total=6.059E+21 #/m**3
    0.000 0.000 0.000 0.000 % by mass MolWgt=43.446
    0.000 0.000 0.000 0.000% volume (mole) fraction
```

```
Time (rel. to T0) = 4500.0 sec. ( 0.051 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 40.000 km (43.648 km) OWLT = 19.01 Min
Topographic Height =-3.648 km Radius (Areoid) = 3431.950 (3391.950) km
Hgt Above Ellipsoid = 39.903 km Scale Hgt H(p)= 7.56 H(rho)= 8.12 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 0.000 km
Planeto-Centric Lat = 26.98 deg Longitude = 52.47 W (307.53 E) deg.
Planeto-Graphic Lat = 27.25 deg Planeto-Graphic Hgt (Ellps)= 39.903 km
Planeto-Cent Sun Lat =25.02 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 126.94 deg.W Local True Solar Time = 16.96 Mars hrs
Temperature = 139.3 K Pressure = 6.522E+00 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 2.350E-04 2.477E-04 2.610E-04 kg/m**3
Departure, COSPAR NH Mean = -30.9% -27.2 % -23.2 % iupdate = 1
Tot.Dens. = 2.232E-04 kg/m**3 Dens.Pert. = -9.88% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -89.0 -4.8-93.7 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -9.0-4.4-13.4 m/s -0.0 m/s
CO2=3.267E+21 N2=9.865E+19 Ar=5.919E+19 O2=4.663E+18 CO=3.587E+18 #/m**3
    96.393 1.853 1.586 0.100 0.067 % by mass
    95.162 2.874 1.724 0.136 0.105 % by volume
H2O=2.599E+13 O=0.000E +00 He=0.000E+00 H2=0.000E+00 Total=3.433E+21 #/m**3
    0.000 0.000 0.000 0.000 % by mass MolWgt=43.447
    0.000 0.000 0.000 0.000 % volume (mole) fraction
```

```
Time (rel. to T0) \(=5000.0 \mathrm{sec} .(0.056\) sols \() \mathrm{Ls}=97.0\) Dust \(=0.30\)
Height Above MOLA (or Surface) \(=45.000 \mathrm{~km}(48.628 \mathrm{~km})\) OWLT \(=19.01 \mathrm{Min}\)
Topographic Height \(=-3.628 \mathrm{~km}\) Radius \((\) Areoid \()=3436.830(3391.830) \mathrm{km}\)
Hgt Above Ellipsoid \(=44.926 \mathrm{~km}\) Scale \(\mathrm{Hgt} \mathrm{H}(\mathrm{p})=7.33 \mathrm{H}(\mathrm{rho})=7.41 \mathrm{~km}\)
Height Offset Parameters: ibougher \(=1\) Local Height Offset \(=0.000 \mathrm{~km}\)
Planeto-Centric Lat \(=27.48\) deg Longitude \(=52.97 \mathrm{~W}(307.03 \mathrm{E})\) deg.
Planeto-Graphic Lat \(=27.75\) deg Planeto-Graphic Hgt \((\) Ellps \()=44.926 \mathrm{~km}\)
Planeto-Cent Sun Lat \(=25.02 \mathrm{deg}\) Mars Orbital Radius \(=1.649 \mathrm{AU}\)
Sun Longitude \(=128.97\) deg.W Local True Solar Time \(=17.07\) Mars hrs
Temperature \(=133.2 \mathrm{~K}\) Pressure \(=3.389 \mathrm{E}+00 \mathrm{~N} / \mathrm{m}^{*} * 2\) profwgt \(=0.000\)
Density (Low, Avg., High) \(=1.267 \mathrm{E}-041.346 \mathrm{E}-041.430 \mathrm{E}-04 \mathrm{~kg} / \mathrm{m}^{*}{ }^{*} 3\)
Departure, COSPAR NH Mean \(=-34.7 \%-30.6 \% \quad-26.3 \%\) iupdate \(=1\)
Tot.Dens. \(=1.225 \mathrm{E}-04 \mathrm{~kg} / \mathrm{m}^{* *} 3\) Dens.Pert. \(=-8.99 \%\) Wave \(=0.00 \%\) of mean
Eastward Wind (Mean,Perturbed,Total) \(=-86.71 .6-85.1 \mathrm{~m} / \mathrm{s}\) VertWind
Northward Wind(Mean,Perturbed,Total) \(=-12.8 \quad 0.3-12.4 \mathrm{~m} / \mathrm{s}-1.0 \mathrm{~m} / \mathrm{s}\)
\(\mathrm{CO} 2=1.776 \mathrm{E}+21 \mathrm{~N} 2=5.361 \mathrm{E}+19 \mathrm{Ar}=3.216 \mathrm{E}+19 \mathrm{O} 2=2.534 \mathrm{E}+18 \mathrm{CO}=1.949 \mathrm{E}+18\) \#/m**3
    \(\begin{array}{llllll}96.395 & 1.853 & 1.585 & 0.100 & 0.067 \%\end{array}\)
    \(95.163 \quad 2.873 \quad 1.724 \quad 0.136 \quad 0.104 \%\) by volume
\(\mathrm{H} 2 \mathrm{O}=3.620 \mathrm{E}+12 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00\) Total \(=1.866 \mathrm{E}+21 \# / \mathrm{m}^{* *} 3\)
    \(\begin{array}{lllll}0.000 & 0.000 & 0.000 & 0.000 & \% \\ \text { by mass MolWgt }=43.447\end{array}\)
    \(0.000 \quad 0.000 \quad 0.000 \quad 0.000 \%\) volume (mole) fraction
```

Time (rel. to T0) $=5500.0 \mathrm{sec} .(0.062$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=50.000 \mathrm{~km}(53.523 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.523 \mathrm{~km}$ Radius $($ Areoid $)=3441.711(3391.711) \mathrm{km}$
Hgt Above Ellipsoid $=49.951 \mathrm{~km}$ Scale Hgt H(p) $=7.86 \mathrm{H}(\mathrm{rho})=7.21 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$ Planeto-Centric Lat $=27.98$ deg Longitude $=53.47 \mathrm{~W}(306.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=28.26$ deg Planeto-Graphic Hgt $($ Ellps $)=49.951 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02 \mathrm{deg}$ Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=130.99$ deg.W Local True Solar Time $=17.17$ Mars hrs
Temperature $=132.4 \mathrm{~K} \quad$ Pressure $=1.726 \mathrm{E}+00 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=6.440 \mathrm{E}-056.903 \mathrm{E}-057.399 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-40.4 \% \quad-36.1 \% \quad-31.5 \%$ iupdate $=1$
Tot.Dens. $=6.888 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=-0.21 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-86.40 .5-86.0 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=12.27 .820 .0 \mathrm{~m} / \mathrm{s}-0.5 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=9.105 \mathrm{E}+20 \mathrm{~N} 2=2.747 \mathrm{E}+19 \mathrm{Ar}=1.648 \mathrm{E}+19 \mathrm{O} 2=1.299 \mathrm{E}+18 \mathrm{CO}=9.990 \mathrm{E}+17 \# / \mathrm{m} * * 3$

| 96.396 | 1.852 | 1.585 | 0.100 | $0.067 \%$ by mass |
| :--- | :--- | :--- | :--- | :--- |
| 95.165 | 2.872 | 1.723 | 0.136 | $0.104 \%$ by volume |
| $\mathrm{H} 2 \mathrm{O}=2.765 \mathrm{E}+12$ | $\mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00 \quad$ Total $=9.568 \mathrm{E}+20$ \#/m**3 |  |  |  |
| 0.000 | 0.000 | 0.000 | $0.000 \%$ by mass MolWgt $=43.447$ |  |
| 0.000 | 0.000 | 0.000 | $0.000 \%$ volume (mole) fraction |  |

Time (rel. to T0) $=6000.0 \mathrm{sec} .(0.068$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=55.000 \mathrm{~km}(58.286 \mathrm{~km})$ OWLT $=19.01 \mathrm{Min}$
Topographic Height $=-3.286 \mathrm{~km}$ Radius $($ Areoid $)=3446.592(3391.592) \mathrm{km}$
Hgt Above Ellipsoid $=54.978 \mathrm{~km}$ Scale Hgt H(p) $=9.17$ H(rho) $=8.50 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.000 \mathrm{~km}$
Planeto-Centric Lat $=28.48 \mathrm{deg}$ Longitude $=53.97 \mathrm{~W}(306.03 \mathrm{E}) \mathrm{deg}$.
Planeto-Graphic Lat $=28.76$ deg Planeto-Graphic Hgt $($ Ellps $)=54.978 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=133.02$ deg. W Local True Solar Time $=17.27$ Mars hrs
Temperature $=140.4 \mathrm{~K}$ Pressure $=9.204 \mathrm{E}-01 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=3.209 \mathrm{E}-053.475 \mathrm{E}-053.764 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m} * * 3$
Departure, COSPAR NH Mean $=-45.8 \%-41.3 \%-36.4 \%$ iupdate $=1$
Tot.Dens. $=2.639 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=-24.06 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-68.2-6.4-74.6 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-1.9 \quad 1.7-0.1 \mathrm{~m} / \mathrm{s}-4.6 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=4.584 \mathrm{E}+20 \mathrm{~N} 2=1.383 \mathrm{E}+19 \mathrm{Ar}=8.296 \mathrm{E}+18 \mathrm{O} 2=6.536 \mathrm{E}+17 \mathrm{CO}=5.028 \mathrm{E}+17 \# / \mathrm{m}^{* *} 3$ $\begin{array}{llllll}96.398 & 1.851 & 1.584 & 0.100 & 0.067 & \%\end{array}$ $\begin{array}{lllll}95.167 & 2.870 & 1.722 & 0.136 & 0.104\end{array} \%$ by volume
$\mathrm{H} 2 \mathrm{O}=3.607 \mathrm{E}+13 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00 \mathrm{Total}=4.817 \mathrm{E}+20 \mathrm{\#} / \mathrm{m}^{* *} 3$
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt=43.447
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}$

```
Time (rel. to T0) \(=6500.0 \mathrm{sec} .(0.073\) sols) Ls \(=97.0\) Dust \(=0.30\)
Height Above MOLA \((\) or Surface \()=60.000 \mathrm{~km}(63.123 \mathrm{~km})\) OWLT \(=19.01 \mathrm{Min}\)
Topographic Height \(=-3.123 \mathrm{~km}\) Radius \((\) Areoid \()=3451.473(3391.473) \mathrm{km}\)
Hgt Above Ellipsoid \(=60.007 \mathrm{~km}\) Scale Hgt H(p) \(=12.00 \mathrm{H}(\mathrm{rho})=11.87 \mathrm{~km}\)
Height Offset Parameters: ibougher \(=1\) Local Height Offset \(=0.000 \mathrm{~km}\)
Planeto-Centric Lat \(=28.98 \mathrm{deg}\) Longitude \(=54.47 \mathrm{~W}(305.53 \mathrm{E}) \mathrm{deg}\).
Planeto-Graphic Lat \(=29.26\) deg Planeto-Graphic Hgt \((\) Ellps \()=60.006 \mathrm{~km}\)
Planeto-Cent Sun Lat \(=25.02\) deg Mars Orbital Radius \(=1.649 \mathrm{AU}\)
Sun Longitude \(=135.05\) deg. W Local True Solar Time \(=17.37\) Mars hrs
Temperature \(=146.6 \mathrm{~K}\) Pressure \(=5.372 \mathrm{E}-01 \mathrm{~N} / \mathrm{m}^{* *} 2\) profwgt \(=0.000\)
Density (Low, Avg., High) \(=1.774 \mathrm{E}-051.943 \mathrm{E}-052.129 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m} * * 3\)
Departure, COSPAR NH Mean = -44.2 \% \(-38.9 \% \quad-33.0 \%\) iupdate \(=1\)
Tot.Dens. \(=1.943 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m}^{*} * 3\) Dens.Pert. \(=0.01 \%\) Wave \(=0.00 \%\) of mean
Eastward Wind \((\) Mean,Perturbed,Total \()=-45.72 .1-43.6 \mathrm{~m} / \mathrm{s}\) VertWind
Northward Wind(Mean,Perturbed,Total) \(=-12.5-2.2-14.7 \mathrm{~m} / \mathrm{s}-2.0 \mathrm{~m} / \mathrm{s}\)
\(\mathrm{CO} 2=2.563 \mathrm{E}+20 \mathrm{~N} 2=7.728 \mathrm{E}+18 \mathrm{Ar}=4.637 \mathrm{E}+18 \mathrm{O} 2=3.653 \mathrm{E}+17 \mathrm{CO}=2.810 \mathrm{E}+17 \# / \mathrm{m}^{*} * 3\)
    \(\begin{array}{llllll}96.399 & 1.850 & 1.583 & 0.100 & 0.067 & \%\end{array}\)
    \(\begin{array}{lllll}95.169 & 2.869 & 1.721 & 0.136 & 0.104 \% \text { by volume }\end{array}\)
\(\mathrm{H} 2 \mathrm{O}=2.233 \mathrm{E}+14 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00 \mathrm{Total}=2.693 \mathrm{E}+20 \mathrm{\#} / \mathrm{m}^{* *} 3\)
\(\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000\end{array} \%\) by mass \(\mathrm{MolWgt}=43.448\)
\(\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}\)
```

Time (rel. to T0) $=7000.0 \mathrm{sec} .(0.079$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=65.000 \mathrm{~km}(68.349 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-3.349 \mathrm{~km}$ Radius $($ Areoid $)=3456.354(3391.354) \mathrm{km}$
Hgt Above Ellipsoid $=65.038 \mathrm{~km}$ Scale Hgt H(p) $=9.35 \mathrm{H}(\mathrm{rho})=9.59 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=0.498 \mathrm{~km}$ Planeto-Centric Lat $=29.48$ deg Longitude $=54.97 \mathrm{~W}(305.03 \mathrm{E}) \mathrm{deg}$.
Planeto-Graphic Lat $=29.77$ deg Planeto-Graphic Hgt $($ Ellps $)=65.037 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=137.08$ deg.W Local True Solar Time $=17.47$ Mars hrs

Temperature $=147.0 \mathrm{~K} \quad$ Pressure $=3.605 \mathrm{E}-01 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$ Density (Low, Avg., High) $=1.170 \mathrm{E}-051.299 \mathrm{E}-051.443 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Departure, COSPAR NH Mean $=-30.3 \%-22.7 \%-14.1 \%$ iupdate $=1$
Tot.Dens. $=1.228 \mathrm{E}-05 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=-5.46 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-29.47 .5-21.9 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-23.0-9.5-32.6 \mathrm{~m} / \mathrm{s}-2.1 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.714 \mathrm{E}+20 \mathrm{~N} 2=5.165 \mathrm{E}+18 \mathrm{Ar}=3.099 \mathrm{E}+18 \mathrm{O} 2=2.442 \mathrm{E}+17 \mathrm{CO}=1.878 \mathrm{E}+17 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llllll}96.401 & 1.849 & 1.583 & 0.100 & 0.067 & \%\end{array}$
$95.171 \quad 2.868 \quad 1.721 \quad 0.136 \quad 0.104 \%$ by volume
$\mathrm{H} 2 \mathrm{O}=2.476 \mathrm{E}+14 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=1.801 \mathrm{E}+20 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt=43.448
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000\end{array} \%$ volume (mole) fraction


Time $($ rel. to T 0$)=8000.0 \mathrm{sec} .(0.090$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=75.000 \mathrm{~km}(77.908 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-2.908 \mathrm{~km}$ Radius $($ Areoid $)=3466.114(3391.114) \mathrm{km}$
Hgt Above Ellipsoid $=75.101 \mathrm{~km}$ Scale Hgt H(p) $=6.87 \mathrm{H}(\mathrm{rho})=7.12 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=1.481 \mathrm{~km}$
Planeto-Centric Lat $=30.48$ deg Longitude $=55.97 \mathrm{~W}(304.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=30.77$ deg Planeto-Graphic Hgt $($ Ellps $)=75.100 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=141.13$ deg. W Local True Solar Time $=17.68$ Mars hrs
Temperature $=141.2 \mathrm{~K} \quad$ Pressure $=1.160 \mathrm{E}-01 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) = 3.779E-06 4.334E-06 4.971E-06 kg/m**3
Departure, COSPAR NH Mean = -15.5 \% $-3.0 \% \quad 11.2 \%$ iupdate $=1$
Tot.Dens. $=4.403 \mathrm{E}-06 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=1.59 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-47.9-10.4-58.2 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-19.0-13.1-32.1 \mathrm{~m} / \mathrm{s} \quad 0.2 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=5.718 \mathrm{E}+19 \mathrm{~N} 2=1.721 \mathrm{E}+18 \mathrm{Ar}=1.033 \mathrm{E}+18 \mathrm{O} 2=8.138 \mathrm{E}+16 \mathrm{CO}=6.260 \mathrm{E}+16 \# / \mathrm{m}^{* *} 3$
$\begin{array}{lllll}96.404 & 1.848 & 1.581 & 0.100 & 0.067\end{array} \%$ by mass
$\begin{array}{lllll}95.176 & 2.865 & 1.719 & 0.135 & 0.104\end{array} \%$ by volume
$\mathrm{H} 2 \mathrm{O}=4.693 \mathrm{E}+13 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=6.008 \mathrm{E}+19 \mathrm{\#} / \mathrm{m}^{*} * 3$
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt=43.448
$\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}$
Time (rel. to T0) $=8500.0 \mathrm{sec} .(0.096$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=80.000 \mathrm{~km}(82.699 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-2.699 \mathrm{~km}$ Radius $($ Areoid $)=3470.992(3390.992) \mathrm{km}$

Hgt Above Ellipsoid $=80.133 \mathrm{~km}$ Scale Hgt H(p) $=6.87 \mathrm{H}(\mathrm{rho})=7.11 \mathrm{~km}$ Height Offset Parameters: ibougher = 1 Local Height Offset $=1.961 \mathrm{~km}$ Planeto-Centric Lat $=30.98 \mathrm{deg}$ Longitude $=56.47 \mathrm{~W}(303.53 \mathrm{E})$ deg. Planeto-Graphic Lat $=31.27$ deg Planeto-Graphic Hgt $($ Ellps $)=80.132 \mathrm{~km}$ Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$ Sun Longitude $=143.16$ deg.W Local True Solar Time $=17.78$ Mars hrs Temperature $=133.8 \mathrm{~K} \quad$ Pressure $=5.632 \mathrm{E}-02 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$ Density (Low, Avg., High) $=1.843 \mathrm{E}-062.155 \mathrm{E}-062.521 \mathrm{E}-06 \mathrm{~kg} / \mathrm{m} * * 3$ Departure, COSPAR NH Mean $=-19.5 \%-5.9 \% 10.1 \%$ iupdate $=1$ Tot.Dens. $=2.363 \mathrm{E}-06 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=9.61 \%$ Wave $=0.00 \%$ of mean Eastward Wind $($ Mean,Perturbed,Total $)=-10.73 .3-7.5 \mathrm{~m} / \mathrm{s}$ VertWind Northward Wind(Mean,Perturbed,Total) $=-10.3-13.9-24.2 \mathrm{~m} / \mathrm{s}-0.1 \mathrm{~m} / \mathrm{s}$ $\mathrm{CO} 2=2.843 \mathrm{E}+19 \mathrm{~N} 2=8.557 \mathrm{E}+17 \mathrm{Ar}=5.134 \mathrm{E}+17 \mathrm{O} 2=4.045 \mathrm{E}+16 \mathrm{CO}=3.111 \mathrm{E}+16 \mathrm{\#} / \mathrm{m}^{* *} 3$ $\begin{array}{llllll}96.406 & 1.847 & 1.581 & 0.100 & 0.067 \% \text { by mass }\end{array}$ $\begin{array}{lllll}95.178 & 2.864 & 1.718 & 0.135 & 0.104\end{array} \%$ by volume $\mathrm{H} 2 \mathrm{O}=4.214 \mathrm{E}+12 \mathrm{O}=0.000 \mathrm{E}+00 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00$ Total $=2.987 \mathrm{E}+19 \# / \mathrm{m}^{* *} 3$ $\begin{array}{lllll}0.000 & 0.000 & 0.000 & 0.000 & \% \\ \text { by mass MolWgt=43.449 }\end{array}$ $\begin{array}{llll}0.000 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}$

```
Time (rel. to T0) = 9000.0 sec. ( 0.101 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 85.000 km ( 87.133 km) OWLT = 19.02 Min
Topographic Height = -2.133 km Radius (Areoid) = 3475.868 (3390.868) km
Hgt Above Ellipsoid = 85.164 km Scale Hgt H(p)= 6.18 H(rho)= 6.31 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 2.280 km
Planeto-Centric Lat = 31.48 deg Longitude = 56.97 W (303.03 E) deg.
Planeto-Graphic Lat = 31.77 deg Planeto-Graphic Hgt (Ellps)= 85.163 km
Planeto-Cent Sun Lat =25.02 deg Mars Orbital Radius =1.649 AU
Sun Longitude = 145.19 deg.W Local True Solar Time = 17.88 Mars hrs
Exospheric Temp. = 186.9 K Tbase = 149.1 K Zbase = 123.1 km
Solar Zenith Angle = 75.8 deg F1 peak = 134.1 km
Temperature = 128.6 K Pressure =2.617E-02 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 8.547E-07 1.022E-06 1.221E-06 kg/m**3
Departure, COSPAR NH Mean = -27.0 % -12.7% 4.4 % iupdate = 1
Tot.Dens. = 8.968E-07 kg/m**3 Dens.Pert. = -12.23% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = 7.9-13.0 -5.1 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -7.9-16.0-23.9 m/s -0.3 m/s
CO2=1.313E+19 N2=4.190E+17 Ar=2.510E+17 O2=2.104E+16 CO=3.575E+16 #/m**3
    93.902 1.908 1.630
    89.097 2.843 1.704 0.143 0.243 % by volume
O}=8.797\textrm{E}+17\textrm{He}=0.000\textrm{E}+00\textrm{H}2=0.000\textrm{E}+00\textrm{H}=0.000\textrm{E}+00\mathrm{ Total=1.474E+19 #/m**3
    2.287 0.000 0.000 0.000 % by mass MolWgt=41.757
    5.970}00.000\quad0.000\quad0.000% volume (mole) fractio
```

Time (rel. to T0) $=9500.0 \mathrm{sec} .(0.107 \mathrm{sols}) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=90.000 \mathrm{~km}(90.974 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-0.974 \mathrm{~km}$ Radius $($ Areoid $)=3480.742(3390.742) \mathrm{km}$
Hgt Above Ellipsoid $=90.195 \mathrm{~km}$ Scale Hgt H(p) $=6.04 \mathrm{H}(\mathrm{rho})=6.19 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=31.98$ deg Longitude $=57.47 \mathrm{~W}(302.53 \mathrm{E}) \mathrm{deg}$.
Planeto-Graphic Lat $=32.28$ deg Planeto-Graphic Hgt $($ Ellps $)=90.194 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=147.22$ deg. W Local True Solar Time $=17.98$ Mars hrs
Exospheric Temp. $=186.4 \mathrm{~K} \quad$ Tbase $=148.8 \mathrm{~K}$ Zbase $=123.2 \mathrm{~km}$
Solar Zenith Angle $=76.9 \mathrm{deg}$ F1 peak $=134.7 \mathrm{~km}$
Temperature $=126.3 \mathrm{~K}$ Pressure $=1.163 \mathrm{E}-02 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=3.765 \mathrm{E}-074.613 \mathrm{E}-075.654 \mathrm{E}-07 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-37.4 \%-23.2 \%-5.9 \%$ iupdate $=1$
Tot.Dens. $=4.314 \mathrm{E}-07 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=-6.48 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=8.64 .913 .5 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind (Mean,Perturbed,Total) $=-12.5-12.2-24.7 \mathrm{~m} / \mathrm{s} \quad 0.3 \mathrm{~m} / \mathrm{s}$

```
CO2=5.917E+18 N2=1.882E+17 Ar=1.126E+17 O2=1.001E+16 CO=2.537E+16 #/m**3
    93.730
    88.755 2.823 1.689 0.150 0.381 % by volume
O=4.135E+17 He=0.000E+00 H2=0.000E+00 H=0.000E+00 Total=6.667E+18 #/m**3
    2.381 0.000 0.000 0.000 % by mass MolWgt=41.673
    6.202 0.000 0.000 0.000 % volume (mole) fraction
```

Time (rel. to T0 $)=10000.0 \mathrm{sec} .(0.113$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=95.000 \mathrm{~km}(95.839 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-0.839 \mathrm{~km}$ Radius $($ Areoid $)=3485.614(3390.614) \mathrm{km}$
Hgt Above Ellipsoid $=95.225 \mathrm{~km}$ Scale $\mathrm{Hgt} \mathrm{H}(\mathrm{p})=6.14 \mathrm{H}(\mathrm{rho})=6.11 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=32.48$ deg Longitude $=57.97 \mathrm{~W}(302.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=32.78$ deg Planeto-Graphic Hgt $($ Ellps $)=95.224 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=149.24$ deg. W Local True Solar Time $=18.08$ Mars hrs
Exospheric Temp. $=185.9 \mathrm{~K} \quad$ Tbase $=148.6 \mathrm{~K}$ Zbase $=123.3 \mathrm{~km}$
Solar Zenith Angle $=77.9 \mathrm{deg}$ F1 peak $=135.3 \mathrm{~km}$
Temperature $=125.7 \mathrm{~K} \quad$ Pressure $=5.170 \mathrm{E}-03 \mathrm{~N} / \mathrm{m}^{* * 2}$ profwgt $=0.000$
Density (Low, Avg., High) $=1.636 \mathrm{E}-072.061 \mathrm{E}-072.597 \mathrm{E}-07 \mathrm{~kg} / \mathrm{m}^{*}{ }^{*} 3$
Departure, COSPAR NH Mean = -47.1 \% $-33.3 \%-16.0 \%$ iupdate $=1$
Tot.Dens. $=2.028 \mathrm{E}-07 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=-1.62 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=6.7-1.05 .7 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-19.0-3.5-22.5 \mathrm{~m} / \mathrm{s} \quad 0.0 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=2.641 \mathrm{E}+18 \mathrm{~N} 2=8.355 \mathrm{E}+16 \mathrm{Ar}=4.992 \mathrm{E}+16 \mathrm{O} 2=4.696 \mathrm{E}+15 \mathrm{CO}=1.543 \mathrm{E}+16 \mathrm{\#} / \mathrm{m} * * 3$
$\begin{array}{lllll}93.638 & 1.886 & 1.607 & 0.121 & 0.348\end{array} \%$ by mass
$88.602 \quad 2.803 \quad 1.675 \quad 0.158 \quad 0.518 \%$ by volume
$\mathrm{O}=1.861 \mathrm{E}+17 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00 \mathrm{H}=0.000 \mathrm{E}+00 \mathrm{Total}=2.980 \mathrm{E}+18 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llll}2.399 & 0.000 & 0.000 & 0.000\end{array} \%$ by mass MolWgt=41.642
$6.245 \quad 0.000 \quad 0.000 \quad 0.000 \%$ volume (mole) fraction

```
Time (rel. to T0) = 10500.0 sec. ( 0.118 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 100.000 km (100.979 km) OWLT = 19.02 Min
Topographic Height = -0.979 km Radius (Areoid) = 3490.485 (3390.485) km
Hgt Above Ellipsoid = 100.255 km Scale Hgt H(p)= 6.10 H(rho)= 6.17 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 2.280 km
Planeto-Centric Lat = 32.98 deg Longitude = 58.47 W (301.53 E) deg.
Planeto-Graphic Lat = 33.28 deg Planeto-Graphic Hgt (Ellps)= 100.253 km
Planeto-Cent Sun Lat = 25.02 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 151.27 deg.W Local True Solar Time = 18.19 Mars hrs
Exospheric Temp. = 185.3 K Tbase =148.3 K Zbase = 123.4 km
Solar Zenith Angle = 78.9 deg F1 peak = 136.0 km
Temperature = 126.1 K Pressure =2.316E-03 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 7.110E-08 9.243E-08 1.202E-07 kg/m**3
Departure, COSPAR NH Mean = -55.3% -41.9% -24.4 % iupdate = 1
Tot.Dens. = 6.524E-08 kg/m**3 Dens.Pert. = -29.42% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = 1.9 1.5 3.4 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -26.4 1.5-24.9 m/s 2.4 m/s
CO2=1.188E+18 N2=3.703E+16 Ar=2.210E+16 O2=2.194E+15 CO=8.707E+15 #/m**3
    93.892 1.864 1.586 0.126 0.438% by mass
    89.262 2.784 1.661 0.165 0.654 % by volume
O=7.283E+16 He=0.000E+00 H2=0.000E+00 H=0.000E+00 Total=1.330E+18 #/m**3
    2.093 0.000 0.000 0.000 % by mass MolWgt=41.839
    5.474 0.000 0.000 0.000 % volume (mole) fraction
```

Time (rel. to T0 $)=11000.0 \mathrm{sec} .(0.124$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=105.000 \mathrm{~km}(106.163 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.163 \mathrm{~km}$ Radius $($ Areoid $)=3495.353(3390.353) \mathrm{km}$
Hgt Above Ellipsoid $=105.284 \mathrm{~km}$ Scale Hgt H(p) $=6.61 \mathrm{H}(\mathrm{rho})=6.41 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$

Planeto-Centric Lat $=33.48$ deg Longitude $=58.97 \mathrm{~W}(301.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=33.78$ deg Planeto-Graphic Hgt $($ Ellps $)=105.282 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=153.30$ deg.W Local True Solar Time $=18.29$ Mars hrs
Exospheric Temp. $=184.8 \mathrm{~K} \quad$ Tbase $=148.0 \mathrm{~K}$ Zbase $=123.5 \mathrm{~km}$
Solar Zenith Angle $=79.8 \mathrm{deg}$ F1 peak $=136.7 \mathrm{~km}$
Temperature $=127.2 \mathrm{~K}$ Pressure $=1.072 \mathrm{E}-03 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=3.196 \mathrm{E}-084.235 \mathrm{E}-085.612 \mathrm{E}-08 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean = -60.1 \% -47.1 \% -29.9 \% iupdate = 1
Tot.Dens. $=2.125 \mathrm{E}-08 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=-49.83 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-5.4-3.1-8.5 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-32.915 .8-17.0 \mathrm{~m} / \mathrm{s} \quad 1.4 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=5.435 \mathrm{E}+17 \mathrm{~N} 2=1.687 \mathrm{E}+16 \mathrm{Ar}=1.005 \mathrm{E}+16 \mathrm{O} 2=1.051 \mathrm{E}+15 \mathrm{CO}=4.824 \mathrm{E}+15$ \#/m**3
$\begin{array}{lllll}93.787 & 1.853 & 1.575 & 0.132 & 0.530 \% \text { by mass }\end{array}$
$89.077 \quad 2.765 \quad 1.647 \quad 0.172 \quad 0.791 \%$ by volume
$\mathrm{O}=3.386 \mathrm{E}+16 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00 \mathrm{H}=0.000 \mathrm{E}+00$ Total $=6.102 \mathrm{E}+17 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llll}2.124 & 0.000 & 0.000 & 0.000 \% \text { by mass MolWgt=41.799 }\end{array}$
$\begin{array}{llll}5.548 & 0.000 & 0.000 & 0.000 \% \text { volume (mole) fraction }\end{array}$

```
Time (rel. to T0) = 11500.0 sec. ( 0.130 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 110.000 km (111.266 km) OWLT = 19.02 Min
Topographic Height = -1.266 km Radius (Areoid) = 3500.219 (3390.219) km
Hgt Above Ellipsoid = 110.312 km Scale Hgt H(p)= 6.48 H(rho)= 6.43 km
Height Offset Parameters: ibougher =1 Local Height Offset =2.280 km
Planeto-Centric Lat = 33.98 deg Longitude = 59.47 W (300.53 E) deg.
Planeto-Graphic Lat = 34.28 deg Planeto-Graphic Hgt (Ellps)=110.310 km
Planeto-Cent Sun Lat =25.02 deg Mars Orbital Radius =1.649 AU
Sun Longitude = 155.33 deg.W Local True Solar Time = 18.39 Mars hrs
Exospheric Temp. = 184.3 K Tbase =147.8 K Zbase = 123.6 km
Solar Zenith Angle = 80.8 deg F1 peak = 137.5 km
Temperature = 129.4 K Pressure = 5.056E-04 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 1.459E-08 1.970E-08 2.660E-08 kg/m**3
Departure, COSPAR NH Mean = -64.8 % -52.4 % -35.8 % iupdate = 1
Tot.Dens. = 1.592E-08 kg/m**3 Dens.Pert. = -19.19% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -13.7 -8.5-22.2 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -36.3 7.7-28.6 m/s 0.4 m/s
CO2=2.532E+17 N2=7.770E+15 Ar=4.623E+15 O2=5.081E+14 CO=2.621E+15 #/m**3
    93.931 1.835 1.557 0.137 0.619 % by mass
    89.481 2.746 1.634 0.180 0.926 % by volume
O}=1.424\textrm{E}+16\textrm{He}=0.000\textrm{E}+00\textrm{H}2=0.000\textrm{E}+00 H=0.000\textrm{E}+00 Total=2.830E+17 #/m**
    1.921 0.000 0.000 0.000% by mass MolWgt=41.924
    5.034 0.000 0.000 0.000% volume (mole) fraction
```

Time $($ rel. to T 0$)=12000.0 \mathrm{sec} .(0.135$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA ( or Surface ) $=115.000 \mathrm{~km}(116.271 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.271 \mathrm{~km}$ Radius $($ Areoid $)=3505.084(3390.084) \mathrm{km}$
Hgt Above Ellipsoid $=115.339 \mathrm{~km}$ Scale Hgt H(p) $=6.94 \mathrm{H}(\mathrm{rho})=6.41 \mathrm{~km}$
Height Offset Parameters: ibougher = 1 Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=34.48$ deg Longitude $=59.97 \mathrm{~W}(300.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=34.79$ deg Planeto-Graphic Hgt $($ Ellps $)=115.338 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=157.35$ deg. W Local True Solar Time $=18.49$ Mars hrs
Exospheric Temp. $=183.7 \mathrm{~K} \quad$ Tbase $=147.5 \mathrm{~K}$ Zbase $=123.7 \mathrm{~km}$
Solar Zenith Angle $=81.8 \mathrm{deg}$ F1 peak $=138.2 \mathrm{~km}$
Temperature $=133.9 \mathrm{~K}$ Pressure $=2.434 \mathrm{E}-04 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=6.659 \mathrm{E}-09$ 9.156E-09 1.259E-08 kg/m**3
Departure, COSPAR NH Mean $=-69.7 \%-58.4 \% \quad-42.8 \%$ iupdate $=1$
Tot.Dens. $=8.091 \mathrm{E}-09 \mathrm{~kg} / \mathrm{m}^{* * 3}$ Dens.Pert. $=-11.64 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total) $=-22.0-4.4-26.5 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind (Mean,Perturbed,Total) $=-36.9-6.4-43.3 \mathrm{~m} / \mathrm{s} \quad 0.4 \mathrm{~m} / \mathrm{s}$

```
CO2=1.176E+17 N2=3.591E+15 Ar=2.133E+15 O2=2.460E+14 CO=1.397E+15 #/m**3
    93.823 1.825 1.546 0.143 0.710% by mass
    89.289 2.727 1.620 0.187 1.061 % by volume
O=6.734E+15 He=0.000E+00 H2=0.000E+00 H=0.000E+00 Total=1.317E+17 #/m**3
    1.954 0.000 0.000 0.000% by mass MolWgt=41.882
    5.115}00.000 0.000 0.000 % volume (mole) fractio
```

Time (rel. to T 0$)=12500.0 \mathrm{sec} .(0.141 \mathrm{sols}) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=120.000 \mathrm{~km}(121.331 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.331 \mathrm{~km}$ Radius $($ Areoid $)=3509.947$ ( 3389.947 ) km
Hgt Above Ellipsoid $=120.366 \mathrm{~km}$ Scale Hgt H(p) $=7.21 \mathrm{H}(\mathrm{rho})=6.65 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=34.98$ deg Longitude $=60.47 \mathrm{~W}(299.53 \mathrm{E})$ deg.
Planeto-Graphic Lat = 35.29 deg Planeto-Graphic Hgt (Ellps)= 120.364 km
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=159.38$ deg.W Local True Solar Time $=18.59$ Mars hrs
Exospheric Temp. $=183.2 \mathrm{~K} \quad$ Tbase $=147.2 \mathrm{~K}$ Zbase $=123.7 \mathrm{~km}$
Solar Zenith Angle $=82.7 \mathrm{deg}$ F1 peak $=139.1 \mathrm{~km}$
Temperature $=140.8 \mathrm{~K} \quad$ Pressure $=1.213 \mathrm{E}-04 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=3.078 \mathrm{E}-094.309 \mathrm{E}-096.033 \mathrm{E}-09 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean = -74.1 \% $-63.8 \%-49.3 \%$ iupdate $=1$
Tot.Dens. $=8.499 \mathrm{E}-09 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=97.23 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-27.92 .7-25.2 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-35.7-17.8-53.5 \mathrm{~m} / \mathrm{s}-1.5 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=5.504 \mathrm{E}+16 \mathrm{~N} 2=1.690 \mathrm{E}+15 \mathrm{Ar}=1.003 \mathrm{E}+15 \mathrm{O} 2=1.211 \mathrm{E}+14 \mathrm{CO}=7.461 \mathrm{E}+14$ \#/m**3
$\begin{array}{lllll}93.343 & 1.825 & 1.544 & 0.149 & 0.805\end{array} \%$ by mass
$88.228 \quad 2.709 \quad 1.607 \quad 0.194 \quad 1.196$ \% by volume
$\mathrm{O}=3.784 \mathrm{E}+15 \mathrm{He}=0.000 \mathrm{E}+00 \mathrm{H} 2=0.000 \mathrm{E}+00 \mathrm{H}=0.000 \mathrm{E}+00$ Total $=6.239 \mathrm{E}+16$ \#/m**3 $\begin{array}{lllll}2.333 & 0.000 & 0.000 & 0.000 & \%\end{array}$ by mass MolWgt=41.597 $6.066 \quad 0.000 \quad 0.000 \quad 0.000 \%$ volume (mole) fraction

```
Time (rel. to T0) = 13000.0 sec. ( 0.146 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 125.000 km ( 126.297 km) OWLT = 19.02 Min
Topographic Height =-1.297 km Radius (Areoid) = 3514.808 (3389.808) km
Hgt Above Ellipsoid = 125.392 km Scale Hgt H(p)= 7.62 H(rho)= 7.00 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 2.280 km
Planeto-Centric Lat = 35.48 deg Longitude = 60.97 W (299.03 E) deg.
Planeto-Graphic Lat = 35.79 deg Planeto-Graphic Hgt (Ellps)= 125.390 km
Planeto-Cent Sun Lat = 25.02 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 161.41 deg.W Local True Solar Time = 18.70 Mars hrs
Exospheric Temp. = 182.7 K Tbase =146.9 K Zbase = 123.8 km
Solar Zenith Angle = 83.6 deg F1 peak = 140.0 km
Temperature = 149.0 K Pressure =6.236E-05 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 1.468E-09 2.091E-09 2.980E-09 kg/m**3
Departure, COSPAR NH Mean = -77.9% -68.5% -55.1 % iupdate = 1
Tot.Dens. = 3.509E-09 kg/m**3 Dens.Pert. = 67.79% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -30.2 2.7-27.4 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -35.8-21.3-57.1 m/s -2.4 m/s
CO2=2.665E+16 N2=8.580E+14 Ar=4.889E+14 O2=6.273E+13 CO=4.131E+14 #/m**3
    93.136}1.909 1.551 0.159 0.919 % by mass
    87.940
O=1.830E+15 He=1.645E+12 H2=6.592E+10 H=1.925E+11 Total=3.031E+16 #/m**3
    2.325 0.001 0.000 0.000 % by mass MolWgt=41.554
    6.039 0.005 0.000 0.001 % volume (mole) fraction
```

Time (rel. to T0) $=13500.0 \mathrm{sec} .(0.152$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA ( or Surface ) $=130.000 \mathrm{~km}(131.276 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.276 \mathrm{~km}$ Radius $($ Areoid $)=3519.668(3389.668) \mathrm{km}$
Hgt Above Ellipsoid $=130.418 \mathrm{~km}$ Scale Hgt H(p) $=8.18$ H(rho) $=7.42 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$

Planeto-Centric Lat $=35.98 \mathrm{deg}$ Longitude $=61.47 \mathrm{~W}(298.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=36.29$ deg Planeto-Graphic Hgt $($ Ellps $)=130.416 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.02$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=163.44$ deg.W Local True Solar Time $=18.80$ Mars hrs
Exospheric Temp. $=182.2 \mathrm{~K} \quad$ Tbase $=146.7 \mathrm{~K}$ Zbase $=123.9 \mathrm{~km}$
Solar Zenith Angle $=84.5 \mathrm{deg}$ F1 peak $=141.0 \mathrm{~km}$
Temperature $=157.1 \mathrm{~K} \quad$ Pressure $=3.331 \mathrm{E}-05 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=7.256 \mathrm{E}-101.052 \mathrm{E}-091.526 \mathrm{E}-09 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-80.7 \%-72.0 \%-59.4 \%$ iupdate $=1$
Tot.Dens. $=2.051 \mathrm{E}-09 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=94.97 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-29.531 .21 .7 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-38.0-18.7-56.7 \mathrm{~m} / \mathrm{s}-1.5 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.327 \mathrm{E}+16 \mathrm{~N} 2=5.269 \mathrm{E}+14 \mathrm{Ar}=2.569 \mathrm{E}+14 \mathrm{O} 2=3.657 \mathrm{E}+13 \mathrm{CO}=2.537 \mathrm{E}+14 \mathrm{\#} / \mathrm{m} * * 3$
$\begin{array}{llllll}92.205 & 2.330 & 1.620 & 0.185 & 1.122 \%\end{array}$
$86.451 \quad 3.431 \quad 1.673 \quad 0.238 \quad 1.652 \%$ by volume
$\mathrm{O}=1.005 \mathrm{E}+15 \mathrm{He}=1.378 \mathrm{E}+12 \mathrm{H} 2=5.683 \mathrm{E}+10 \mathrm{H}=1.689 \mathrm{E}+11$ Total $=1.535 \mathrm{E}+16$ \#/m**3
$2.537 \quad 0.001 \quad 0.000 \quad 0.000 \%$ by mass MolWgt=41.263
$\begin{array}{lllll}6.543 & 0.009 & 0.000 & 0.001 \% \text { volume (mole) fraction }\end{array}$

```
Time (rel. to T0) = 14000.0 sec. ( 0.158 sols) Ls = 97.0 Dust = 0.30
Height Above MOLA (or Surface) = 135.000 km ( 136.272 km) OWLT = 19.02 Min
Topographic Height =-1.272 km Radius (Areoid) = 3524.526 (3389.526) km
Hgt Above Ellipsoid = 135.443 km Scale Hgt H(p)= 8.69 H(rho)= 8.00 km
Height Offset Parameters: ibougher =1 Local Height Offset =2.280 km
Planeto-Centric Lat = 36.48 deg Longitude = 61.97 W (298.03 E) deg.
Planeto-Graphic Lat = 36.79 deg Planeto-Graphic Hgt (Ellps)= 135.440 km
Planeto-Cent Sun Lat = 25.01 deg Mars Orbital Radius =1.649 AU
Sun Longitude = 165.46 deg.W Local True Solar Time = 18.90 Mars hrs
Exospheric Temp. = 181.7 K Tbase = 146.4 K Zbase = 124.0 km
Solar Zenith Angle = 85.3 deg F1 peak = 142.1 km
Temperature = 163.9 K Pressure = 1.843E-05 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 3.803E-10 5.515E-10 7.996E-10 kg/m**3
Departure, COSPAR NH Mean = -80.1 % -71.2% -58.2 % iupdate = 1
Tot.Dens. = 9.922E-10 kg/m**3 Dens.Pert. =79.92% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -26.7 -1.7-28.4 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -40.7-25.3-66.0 m/s -2.0 m/s
CO2=6.861E+15 N2=3.338E+14 Ar=1.405E+14 O2=2.206E+13 CO=1.608E+14 #/m**3
    90.923 2.816 1.690
    84.264 4.099 1.725 0.271 1.974 % by volume
O=6.229E+14 He=1.171E+12 H2=4.963E+10 H=1.501E+11 Total=8.143E+15 #/m**3
    3.001 0.001 0.000 0.000 % by mass MolWgt=40.786
    7.650}00.014\quad0.001 0.002% volume (mole) fraction
```

Time $($ rel. to T 0$)=14500.0 \mathrm{sec} .(0.163$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=140.000 \mathrm{~km}(141.237 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.237 \mathrm{~km}$ Radius $($ Areoid $)=3529.383(3389.383) \mathrm{km}$
Hgt Above Ellipsoid $=140.467 \mathrm{~km}$ Scale Hgt H(p) $=9.75 \mathrm{H}(\mathrm{rho})=8.91 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=36.98 \mathrm{deg}$ Longitude $=62.47 \mathrm{~W}(297.53 \mathrm{E})$ deg.
Planeto-Graphic Lat $=37.29$ deg Planeto-Graphic Hgt $($ Ellps $)=140.465 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=167.49$ deg. W Local True Solar Time $=19.00$ Mars hrs
Exospheric Temp. $=181.2 \mathrm{~K} \quad$ Tbase $=146.2 \mathrm{~K} \quad$ Zbase $=124.1 \mathrm{~km}$
Solar Zenith Angle $=86.2 \mathrm{deg}$ F1 peak $=143.2 \mathrm{~km}$
Temperature $=169.0 \mathrm{~K} \quad$ Pressure $=1.069 \mathrm{E}-05 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=2.104 \mathrm{E}-103.051 \mathrm{E}-104.423 \mathrm{E}-10 \mathrm{~kg} / \mathrm{m} * * 3$
Departure, COSPAR NH Mean $=-80.7 \%-72.0 \% \quad-59.4 \%$ iupdate $=1$
Tot.Dens. $=4.091 \mathrm{E}-10 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=34.10 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-22.8-6.1-28.9 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-44.5-17.2-61.7 \mathrm{~m} / \mathrm{s}-1.6 \mathrm{~m} / \mathrm{s}$


Time (rel. to T 0 ) $=15000.0 \mathrm{sec} .(0.169$ sols $) \mathrm{Ls}=97.0$ Dust $=0.30$
Height Above MOLA (or Surface) $=145.000 \mathrm{~km}(146.335 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.335 \mathrm{~km}$ Radius $($ Areoid $)=3534.239$ (3389.239) km
Hgt Above Ellipsoid $=145.492 \mathrm{~km}$ Scale Hgt H(p) $=10.14$ H(rho) $=9.20 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=37.48$ deg Longitude $=62.97 \mathrm{~W}(297.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=37.79$ deg Planeto-Graphic Hgt $($ Ellps $)=145.490 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=169.52$ deg. W Local True Solar Time $=19.10$ Mars hrs
Exospheric Temp. $=180.7 \mathrm{~K} \quad$ Tbase $=145.9 \mathrm{~K} \quad$ Zbase $=124.1 \mathrm{~km}$
Solar Zenith Angle $=87.0$ deg F1 peak $=144.5 \mathrm{~km}$
Temperature $=172.7 \mathrm{~K} \quad$ Pressure $=6.444 \mathrm{E}-06 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=1.208 \mathrm{E}-101.752 \mathrm{E}-102.540 \mathrm{E}-10 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-82.9 \% \quad-75.2 \% \quad-64.0 \%$ iupdate $=1$
Tot.Dens. $=2.445 \mathrm{E}-10 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=39.54 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-19.2-7.0-26.2 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-49.3-5.6-54.9 \mathrm{~m} / \mathrm{s}-0.6 \mathrm{~m} / \mathrm{s}$ $\mathrm{CO} 2=2.079 \mathrm{E}+15 \mathrm{~N} 2=1.535 \mathrm{E}+14 \mathrm{Ar}=4.868 \mathrm{E}+13 \mathrm{O} 2=9.234 \mathrm{E}+12 \mathrm{CO}=7.396 \mathrm{E}+13$ \#/m**3
$\begin{array}{lllll}86.732 & 4.077 & 1.844 & 0.280 & 1.963\end{array} \%$ by mass
$\begin{array}{lllll}76.952 & 5.682 & 1.802 & 0.342 \quad 2.737 \% \text { by volume }\end{array}$
$\mathrm{O}=3.362 \mathrm{E}+14 \mathrm{He}=9.462 \mathrm{E}+11 \mathrm{H} 2=4.226 \mathrm{E}+10 \mathrm{H}=1.320 \mathrm{E}+11$ Total $=2.702 \mathrm{E}+15$ \#/m**3
$5.099 \quad 0.004 \quad 0.000 \quad 0.000 \%$ by mass MolWgt=39.046
$12.444 \quad 0.035 \quad 0.002 \quad 0.005 \%$ volume (mole) fraction

```
Time (rel. to T0) = 15500.0 sec. ( 0.175 sols) Ls = 97.1 Dust = 0.30
Height Above MOLA (or Surface) = 150.000 km ( 151.272 km) OWLT = 19.02 Min
Topographic Height =-1.272 km Radius (Areoid) = 3539.095 (3389.095) km
Hgt Above Ellipsoid = 150.517 km Scale Hgt H(p)= 10.31 H(rho)= 9.58 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 2.280 km
Planeto-Centric Lat = 37.98 deg Longitude = 63.47 W (296.53 E) deg.
Planeto-Graphic Lat = 38.29 deg Planeto-Graphic Hgt (Ellps)= 150.514 km
Planeto-Cent Sun Lat = 25.01 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 171.55 deg.W Local True Solar Time = 19.21 Mars hrs
Exospheric Temp. = 180.1 K Tbase =145.7 K Zbase = 124.2 km
Solar Zenith Angle = 87.8 deg F1 peak = 146.0 km
Temperature = 175.0 K Pressure = 3.962E-06 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 7.125E-11 1.033E-10 1.498E-10 kg/m**3
Departure, COSPAR NH Mean = -84.9 % -78.2 % -68.3 % iupdate = 1
Tot.Dens. = 1.792E-10 kg/m**3 Dens.Pert. = 73.48% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -15.5 -8.1-23.6 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -53.6-14.6-68.2 m/s 2.9 m/s
CO2=1.187E+15 N2=1.083E+14 Ar=2.989E+13 O2=6.219E+12 CO=5.218E+13 #/m**3
    83.980
    72.413 6.607 1.823 0.379 3.183 % by volume
O=2.546E+14 He=8.800E +11 H2=4.035E+10 H=1.282E+11 Total=1.640E+15 #/m**3
    6.548 0.006 0.000 0.000 % by mass MolWgt=37.948
    15.531 0.054 0.002 0.008 % volume (mole) fraction
```

Time (rel. to T0) $=16000.0 \mathrm{sec} .(0.180 \mathrm{sols}) \mathrm{Ls}=97.1$ Dust $=0.30$
Height Above MOLA ( or Surface) $=155.000 \mathrm{~km}(156.136 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-1.136 \mathrm{~km}$ Radius $($ Areoid $)=3543.950(3388.950) \mathrm{km}$
Hgt Above Ellipsoid $=155.542 \mathrm{~km}$ Scale $\mathrm{Hgt} \mathrm{H}(\mathrm{p})=10.81 \mathrm{H}(\mathrm{rho})=9.89 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$

Planeto-Centric Lat $=38.48 \mathrm{deg}$ Longitude $=63.97 \mathrm{~W}(296.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=38.80$ deg Planeto-Graphic Hgt $($ Ellps $)=155.540 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=173.58$ deg. W Local True Solar Time $=19.31$ Mars hrs
Exospheric Temp. $=179.5 \mathrm{~K} \quad$ Tbase $=145.4 \mathrm{~K}$ Zbase $=124.3 \mathrm{~km}$
Solar Zenith Angle $=88.6$ deg F1 peak $=147.6 \mathrm{~km}$
Temperature $=176.2 \mathrm{~K}$ Pressure $=2.474 \mathrm{E}-06 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=4.275 \mathrm{E}-116.198 \mathrm{E}-118.987 \mathrm{E}-11 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-87.3 \%-81.6 \%-73.3 \%$ iupdate $=1$
Tot.Dens. $=1.247 \mathrm{E}-10 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=101.12 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-11.7-9.2-20.9 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind $($ Mean,Perturbed,Total $)=-57.6-18.4-75.9 \mathrm{~m} / \mathrm{s} 3.1 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=6.844 \mathrm{E}+14 \mathrm{~N} 2=7.744 \mathrm{E}+13 \mathrm{Ar}=1.861 \mathrm{E}+13 \mathrm{O} 2=4.246 \mathrm{E}+12 \mathrm{CO}=3.731 \mathrm{E}+13 \mathrm{\#} / \mathrm{m}^{*} * 3$
$\begin{array}{lllll}80.695 & 5.813 & 1.992 & 0.364 & 2.799\end{array} \%$ by mass
$\begin{array}{lllll}67.278 & 7.612 & 1.829 & 0.417 & 3.667\end{array} \%$ by volume
$\mathrm{O}=1.943 \mathrm{E}+14 \mathrm{He}=8.284 \mathrm{E}+11 \mathrm{H} 2=3.899 \mathrm{E}+10 \mathrm{H}=1.259 \mathrm{E}+11$ Total $=1.017 \mathrm{E}+15$ \#/m**3
$8.328 \quad 0.009 \quad 0.000 \quad 0.000 \%$ by mass MolWgt=36.692
$19.099 \quad 0.081 \quad 0.004 \quad 0.012 \%$ volume (mole) fraction

```
Time (rel. to T0) = 16500.0 sec. ( 0.186 sols) Ls = 97.1 Dust = 0.30
Height Above MOLA (or Surface) = 160.000 km ( 160.950 km) OWLT = 19.02 Min
Topographic Height = -0.950 km Radius (Areoid) = 3548.806 (3388.806) km
Hgt Above Ellipsoid = 160.568 km Scale Hgt H(p)=11.25 H(rho)= 10.06 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 2.280 km
Planeto-Centric Lat = 38.98 deg Longitude = 64.47 W (295.53 E) deg.
Planeto-Graphic Lat = 39.30 deg Planeto-Graphic Hgt (Ellps)= 160.566 km
Planeto-Cent Sun Lat =25.01 deg Mars Orbital Radius = 1.649 AU
Sun Longitude = 175.60 deg.W Local True Solar Time = 19.41 Mars hrs
Exospheric Temp. = 179.0 K Tbase = 145.2 K Zbase = 124.3 km
Solar Zenith Angle = 89.3 deg F1 peak = 149.3 km
Temperature = 176.7 K Pressure = 1.577E-06 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 2.591E-11 3.757E-11 5.447E-11 kg/m**3
Departure, COSPAR NH Mean = -89.3 % -84.5 % -77.6 % iupdate = 1
Tot.Dens. = 6.128E-11 kg/m**3 Dens.Pert. = 63.11% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -8.5 6.0 -2.5 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -61.5-36.3-97.8 m/s -6.1 m/s
CO2=3.917E+14 N2=5.610E+13 Ar=1.174E+13 O2=2.937E+12 CO=2.703E+13 #/m**3
    76.187 6.947 2.073 0.415 3.346 % by mass
    60.610}80.681 1.817 0.455 4.182 % by volum
O}=1.558\textrm{E}+14\textrm{He}=7.899\textrm{E}+11\textrm{H}2=3.815\textrm{E}+10\textrm{H}=1.253\textrm{E}+11 Total=6.462\textrm{E}+14 #/m**
    11.017 0.014 0.000 0.001 % by mass MolWgt=35.011
    24.108 0.122 0.006 0.019 % volume (mole) fraction
```

Time (rel. to T0) $=17000.0 \mathrm{sec} .(0.191$ sols $) \mathrm{Ls}=97.1$ Dust $=0.30$
Height Above MOLA (or Surface) $=165.000 \mathrm{~km}(165.854 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-0.854 \mathrm{~km}$ Radius $($ Areoid $)=3553.661(3388.661) \mathrm{km}$
Hgt Above Ellipsoid $=165.595 \mathrm{~km}$ Scale Hgt H(p) $=11.63 \mathrm{H}(\mathrm{rho})=10.40 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=39.48$ deg Longitude $=64.97 \mathrm{~W}(295.03 \mathrm{E})$ deg.
Planeto-Graphic Lat $=39.80$ deg Planeto-Graphic Hgt $($ Ellps $)=165.592 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01$ deg Mars Orbital Radius $=1.649 \mathrm{AU}$
Sun Longitude $=177.63$ deg. W Local True Solar Time $=19.51$ Mars hrs
Exospheric Temp. $=178.5 \mathrm{~K} \quad$ Tbase $=144.9 \mathrm{~K}$ Zbase $=124.4 \mathrm{~km}$
Solar Zenith Angle $=90.0$ deg F1 peak $=999.9 \mathrm{~km}$
Temperature $=177.0 \mathrm{~K}$ Pressure $=1.020 \mathrm{E}-06 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=1.590 \mathrm{E}-112.305 \mathrm{E}-113.342 \mathrm{E}-11 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean = -91.2 \% -87.2 \% $-81.5 \%$ iupdate = 1
Tot.Dens. $=3.656 \mathrm{E}-11 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=58.62 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-6.10 .7-5.4 \mathrm{~m} / \mathrm{s}$ VertWind


Time (rel. to T0 $)=17500.0 \mathrm{sec} .(0.197$ sols $) \mathrm{Ls}=97.1$ Dust $=0.30$
Height Above MOLA (or Surface) $=170.000 \mathrm{~km}(170.917 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-0.917 \mathrm{~km}$ Radius $($ Areoid $)=3558.517(3388.517) \mathrm{km}$
Hgt Above Ellipsoid $=170.623 \mathrm{~km}$ Scale Hgt H(p) $=12.23 \mathrm{H}(\mathrm{rho})=10.72 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=39.98$ deg Longitude $=65.47 \mathrm{~W}(294.53 \mathrm{E}) \mathrm{deg}$.
Planeto-Graphic Lat $=40.30$ deg Planeto-Graphic $\operatorname{Hgt}($ Ellps $)=170.620 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01$ deg Mars Orbital Radius $=1.648 \mathrm{AU}$
Sun Longitude $=179.66$ deg.W Local True Solar Time $=19.61$ Mars hrs
Exospheric Temp. $=177.9 \mathrm{~K} \quad$ Tbase $=144.7 \mathrm{~K}$ Zbase $=124.4 \mathrm{~km}$
Solar Zenith Angle $=90.7$ deg F 1 peak $=999.9 \mathrm{~km}$
Temperature $=177.1 \mathrm{~K} \quad$ Pressure $=6.716 \mathrm{E}-07 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) = 9.886E-12 1.434E-11 $2.079 \mathrm{E}-11 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-92.7 \%-89.4 \% \quad-84.6 \%$ iupdate $=1$
Tot.Dens. $=1.551 \mathrm{E}-11 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=8.21 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-3.99 .65 .7 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-69.0-41.7-110.7 \mathrm{~m} / \mathrm{s} \quad 0.9 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.279 \mathrm{E}+14 \mathrm{~N} 2=2.985 \mathrm{E}+13 \mathrm{Ar}=4.737 \mathrm{E}+12 \mathrm{O} 2=1.425 \mathrm{E}+12 \mathrm{CO}=1.438 \mathrm{E}+13 \mathrm{\#} / \mathrm{m}^{* *} 3$
$\begin{array}{lllll}65.184 & 9.688 & 2.193 & 0.528 & 4.667 \%\end{array}$
$46.551 \quad 10.867 \quad 1.724 \quad 0.519 \quad 5.237 \%$ by volume
$\mathrm{O}=9.553 \mathrm{E}+13 \mathrm{He}=7.288 \mathrm{E}+11 \mathrm{H} 2=3.708 \mathrm{E}+10 \mathrm{H}=1.258 \mathrm{E}+11 \mathrm{Total}=2.747 \mathrm{E}+14 \# / \mathrm{m}^{* *} 3$
$\begin{array}{llll}17.704 & 0.034 & 0.001 & 0.001 \% \text { by mass MolWgt=31.429 }\end{array}$
$34.778 \quad 0.265 \quad 0.013 \quad 0.046 \%$ volume (mole) fraction
Time (rel. to T 0$)=18000.0 \mathrm{sec} .(0.203$ sols $) \mathrm{Ls}=97.1$ Dust $=0.30$
Height Above MOLA (or Surface) $=175.000 \mathrm{~km}(175.556 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-0.556 \mathrm{~km}$ Radius $($ Areoid $)=3563.373(3388.373) \mathrm{km}$
Hgt Above Ellipsoid $=175.651 \mathrm{~km}$ Scale Hgt H(p) $=12.56 \mathrm{H}(\mathrm{rho})=11.25 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=40.48 \mathrm{deg}$ Longitude $=65.97 \mathrm{~W}(294.03 \mathrm{E}) \mathrm{deg}$.
Planeto-Graphic Lat $=40.80$ deg Planeto-Graphic Hgt $($ Ellps $)=175.648 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01$ deg Mars Orbital Radius $=1.648 \mathrm{AU}$
Sun Longitude $=181.69$ deg. W Local True Solar Time $=19.71$ Mars hrs
Exospheric Temp. $=177.4 \mathrm{~K} \quad$ Tbase $=144.5 \mathrm{~K}$ Zbase $=124.5 \mathrm{~km}$
Solar Zenith Angle = 91.4 deg F1 peak $=999.9 \mathrm{~km}$
Temperature $=177.1 \mathrm{~K} \quad$ Pressure $=4.470 \mathrm{E}-07 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=6.284 \mathrm{E}-129.113 \mathrm{E}-121.321 \mathrm{E}-11 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-93.9 \%-91.2 \% \quad-87.2 \%$ iupdate $=1$
Tot.Dens. $=9.602 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=5.37 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=-2.62 .60 .0 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-71.4-23.2-94.6 \mathrm{~m} / \mathrm{s} \quad 2.8 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=7.495 \mathrm{E}+13 \mathrm{~N} 2=2.172 \mathrm{E}+13 \mathrm{Ar}=3.000 \mathrm{E}+12 \mathrm{O} 2=9.901 \mathrm{E}+11 \mathrm{CO}=1.047 \mathrm{E}+13 \mathrm{\#} / \mathrm{m}^{* *} 3$
$\begin{array}{lllll}60.104 & 11.087 & 2.184 & 0.577 & 5.342 \% \text { by mass }\end{array}$
$40.997 \quad 11.879 \quad 1.641 \quad 0.542 \quad 5.725 \%$ by volume
$\mathrm{O}=7.083 \mathrm{E}+13 \mathrm{He}=6.988 \mathrm{E}+11 \mathrm{H} 2=3.649 \mathrm{E}+10 \mathrm{H}=1.258 \mathrm{E}+11$ Total $=1.828 \mathrm{E}+14 \# / \mathrm{m}^{* *} 3$
$20.651 \quad 0.051 \quad 0.001 \quad 0.002 \%$ by mass MolWgt=30.018
$38.746 \quad 0.382 \quad 0.020 \quad 0.069 \%$ volume (mole) fraction
Time $($ rel. to T 0$)=18500.0 \mathrm{sec} .(0.208$ sols $) \mathrm{Ls}=97.1$ Dust $=0.30$
Height Above MOLA (or Surface) $=180.000 \mathrm{~km}(180.274 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=-0.274 \mathrm{~km}$ Radius $($ Areoid $)=3568.229(3388.229) \mathrm{km}$
Hgt Above Ellipsoid $=180.679 \mathrm{~km}$ Scale $\mathrm{Hgt} \mathrm{H}(\mathrm{p})=13.20 \mathrm{H}(\mathrm{rho})=11.60 \mathrm{~km}$

Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$ Planeto-Centric Lat $=40.98$ deg Longitude $=66.47 \mathrm{~W}(293.53 \mathrm{E})$ deg. Planeto-Graphic Lat $=41.30$ deg Planeto-Graphic Hgt $($ Ellps $)=180.676 \mathrm{~km}$ Planeto-Cent Sun Lat $=25.01 \mathrm{deg}$ Mars Orbital Radius $=1.648 \mathrm{AU}$
Sun Longitude $=183.71$ deg.W Local True Solar Time $=19.82$ Mars hrs
Exospheric Temp. $=176.9 \mathrm{~K} \quad$ Tbase $=144.3 \mathrm{~K} \mathrm{Zbase}=124.5 \mathrm{~km}$
Solar Zenith Angle $=92.1 \mathrm{deg}$ F1 peak $=999.9 \mathrm{~km}$
Temperature $=176.7 \mathrm{~K} \quad$ Pressure $=3.027 \mathrm{E}-07 \mathrm{~N} / \mathrm{m}^{*} * 2$ profwgt $=0.000$
Density (Low, Avg., High) $=4.040 \mathrm{E}-125.858 \mathrm{E}-128.495 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-94.9 \%-92.6 \%-89.2 \%$ iupdate $=1$
Tot.Dens. $=9.702 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=65.60 \%$ Wave $=0.00 \%$ of mean
Eastward Wind $($ Mean,Perturbed,Total $)=-1.412 .911 .5 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-73.9 \quad 0.8-73.1 \mathrm{~m} / \mathrm{s}-2.8 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=4.338 \mathrm{E}+13 \mathrm{~N} 2=1.582 \mathrm{E}+13 \mathrm{Ar}=1.902 \mathrm{E}+12 \mathrm{O} 2=6.887 \mathrm{E}+11 \mathrm{CO}=7.626 \mathrm{E}+12 \mathrm{\#} / \mathrm{m} * * 3$
$\begin{array}{llllll}54.107 & 12.564 & 2.154 & 0.625 & 6.054 \% \text { by mass }\end{array}$
$34.957 \quad 12.750 \quad 1.533 \quad 0.555 \quad 6.146 \%$ by volume
$\mathrm{O}=5.384 \mathrm{E}+13 \mathrm{He}=6.714 \mathrm{E}+11 \mathrm{H} 2=3.598 \mathrm{E}+10 \mathrm{H}=1.261 \mathrm{E}+11 \mathrm{Total}=1.241 \mathrm{E}+14$ \#/m**3
$24.414 \quad 0.076 \quad 0.002 \quad 0.004 \%$ by mass MolWgt=28.433
$43.388 \quad 0.541 \quad 0.029 \quad 0.102 \%$ volume (mole) fraction

```
Time (rel. to T0) = 19000.0 sec. ( 0.214 sols) Ls = 97.1 Dust = 0.30
Height Above MOLA (or Surface) = 185.000 km (185.015 km) OWLT = 19.02 Min
Topographic Height =-0.015 km Radius (Areoid) = 3573.084 (3388.084) km
Hgt Above Ellipsoid = 185.708 km Scale Hgt H(p)= 13.96 H(rho)= 12.02 km
Height Offset Parameters: ibougher = 1 Local Height Offset = 2.280 km
Planeto-Centric Lat = 41.48 deg Longitude = 66.97 W (293.03 E) deg.
Planeto-Graphic Lat = 41.80 deg Planeto-Graphic Hgt (Ellps)= 185.705 km
Planeto-Cent Sun Lat = 25.01 deg Mars Orbital Radius = 1.648 AU
Sun Longitude = 185.74 deg.W Local True Solar Time = 19.92 Mars hrs
Exospheric Temp. = 176.4 K Tbase =144.1 K Zbase = 124.6 km
Solar Zenith Angle = 92.7 deg F1 peak = 999.9 km
Temperature = 176.3 K Pressure =2.090E-07 N/m**2 profwgt = 0.000
Density (Low, Avg., High) = 2.632E-12 3.816E-12 5.534E-12 kg/m**3
Departure, COSPAR NH Mean = -95.7 % --93.8% -90.9 % iupdate = 1
Tot.Dens. = 5.493E-12 kg/m**3 Dens.Pert. = 43.92% Wave = 0.00% of mean
Eastward Wind (Mean,Perturbed,Total) = -0.2-13.4-13.7 m/s VertWind
Northward Wind(Mean,Perturbed,Total) = -76.4 24.1 -52.3 m/s 0.4 m/s
CO2=2.474E+13 N2=1.152E+13 Ar=1.205E+12 O2=4.787E+11 CO=5.553E+12 #/m**3
    47.366 14.044 2.094 0.666 6.768% by mass
    28.805 13.415 1.403 0.557 6.466 % by volume
O=4.158E+13 He=6.452E+11 H2=3.549E+10 H=1.263E+11 Total=8.588E+13 #/m**3
    28.941 0.112 0.003 0.006 % by mass MolWgt=26.764
    48.414 0.751 0.041 0.147 % volume (mole) fraction
Time \((\) rel. to T 0\()=19500.0 \mathrm{sec} .(0.220\) sols \() \mathrm{Ls}=97.1\) Dust \(=0.30\)
Height Above MOLA (or Surface) \(=190.000 \mathrm{~km}(189.754 \mathrm{~km})\) OWLT \(=19.02 \mathrm{Min}\)
Topographic Height \(=0.246 \mathrm{~km}\) Radius \((\) Areoid \()=3577.940(3387.940) \mathrm{km}\)
Hgt Above Ellipsoid \(=190.737 \mathrm{~km}\) Scale Hgt H(p) \(=14.85 \mathrm{H}(\mathrm{rho})=12.53 \mathrm{~km}\)
Height Offset Parameters: ibougher \(=1\) Local Height Offset \(=2.280 \mathrm{~km}\)
Planeto-Centric Lat \(=41.98\) deg Longitude \(=67.47 \mathrm{~W}(292.53 \mathrm{E})\) deg.
Planeto-Graphic Lat \(=42.30\) deg Planeto-Graphic Hgt \((\) Ellps \()=190.733 \mathrm{~km}\)
Planeto-Cent Sun Lat \(=25.01 \mathrm{deg}\) Mars Orbital Radius \(=1.648 \mathrm{AU}\)
Sun Longitude \(=187.77\) deg.W Local True Solar Time \(=20.02\) Mars hrs
Exospheric Temp. \(=175.9 \mathrm{~K} \quad\) Tbase \(=143.9 \mathrm{~K}\) Zbase \(=124.6 \mathrm{~km}\)
Solar Zenith Angle = 93.3 deg F1 peak \(=999.9 \mathrm{~km}\)
Temperature \(=175.8 \mathrm{~K} \quad\) Pressure \(=1.473 \mathrm{E}-07 \mathrm{~N} / \mathrm{m}^{*} * 2\) profwgt \(=0.000\)
Density (Low, Avg., High) \(=1.741 \mathrm{E}-122.524 \mathrm{E}-123.660 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{*} * 3\)
Departure, COSPAR NH Mean = -96.3 \% \(-94.7 \% \quad-92.3 \%\) iupdate \(=1\)
Tot.Dens. \(=3.176 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{* *} 3\) Dens.Pert. \(=25.83 \%\) Wave \(=0.00 \%\) of mean
```

Eastward Wind $($ Mean,Perturbed,Total $)=0.7-9.5-8.8 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind(Mean,Perturbed,Total) $=-79.0 \quad 13.9-65.1 \mathrm{~m} / \mathrm{s}-0.5 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=1.380 \mathrm{E}+13 \mathrm{~N} 2=8.383 \mathrm{E}+12 \mathrm{Ar}=7.622 \mathrm{E}+11 \mathrm{O} 2=3.325 \mathrm{E}+11 \mathrm{CO}=4.041 \mathrm{E}+12 \mathrm{\#} / \mathrm{m}^{* *} 3$

| 39.955 | 15.450 | 2.004 | 0.700 | $7.446 \%$ by mass |
| :--- | :--- | :--- | :--- | :--- |

$\begin{array}{lllll}22.751 & 13.819 & 1.256 & 0.548 & 6.662\end{array} \%$ by volume
$\mathrm{O}=3.256 \mathrm{E}+13 \mathrm{He}=6.201 \mathrm{E}+11 \mathrm{H} 2=3.500 \mathrm{E}+10 \mathrm{H}=1.266 \mathrm{E}+11$ Total $=6.066 \mathrm{E}+13$ \#/m**3
$34.269 \quad 0.163 \quad 0.005 \quad 0.008 \%$ by mass MolWgt=25.059
$53.676 \quad 1.022 \quad 0.058 \quad 0.209 \%$ volume (mole) fraction
Time (rel. to T0) $=20000.0 \mathrm{sec} .(0.225$ sols) Ls $=97.1$ Dust $=0.30$
Height Above MOLA $($ or Surface $)=195.000 \mathrm{~km}(194.639 \mathrm{~km})$ OWLT $=19.02 \mathrm{Min}$
Topographic Height $=0.361 \mathrm{~km}$ Radius $($ Areoid $)=3582.794(3387.794) \mathrm{km}$
Hgt Above Ellipsoid $=195.765 \mathrm{~km}$ Scale Hgt H(p) $=15.85 \mathrm{H}(\mathrm{rho})=13.15 \mathrm{~km}$
Height Offset Parameters: ibougher $=1$ Local Height Offset $=2.280 \mathrm{~km}$
Planeto-Centric Lat $=42.48 \mathrm{deg}$ Longitude $=67.97 \mathrm{~W}(292.03 \mathrm{E}) \mathrm{deg}$.
Planeto-Graphic Lat $=42.80$ deg Planeto-Graphic Hgt $($ Ellps $)=195.762 \mathrm{~km}$
Planeto-Cent Sun Lat $=25.01 \mathrm{deg}$ Mars Orbital Radius $=1.648 \mathrm{AU}$
Sun Longitude $=189.80$ deg.W Local True Solar Time $=20.12$ Mars hrs
Exospheric Temp. $=175.5 \mathrm{~K} \quad$ Tbase $=143.7 \mathrm{~K} \quad$ Zbase $=124.6 \mathrm{~km}$
Solar Zenith Angle $=93.8 \mathrm{deg}$ F1 peak $=999.9 \mathrm{~km}$
Temperature $=175.4 \mathrm{~K} \quad$ Pressure $=1.060 \mathrm{E}-07 \mathrm{~N} / \mathrm{m}^{* *} 2$ profwgt $=0.000$
Density (Low, Avg., High) $=1.171 \mathrm{E}-121.699 \mathrm{E}-122.463 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{* *} 3$
Departure, COSPAR NH Mean $=-96.8 \%-95.4 \% \quad-93.4 \%$ iupdate $=1$
Tot.Dens. $=1.686 \mathrm{E}-12 \mathrm{~kg} / \mathrm{m}^{* *} 3$ Dens.Pert. $=-0.73 \%$ Wave $=0.00 \%$ of mean
Eastward Wind (Mean,Perturbed,Total) $=1.64 .25 .8 \mathrm{~m} / \mathrm{s}$ VertWind
Northward Wind $($ Mean,Perturbed,Total $)=-81.6 \quad 1.2-80.4 \mathrm{~m} / \mathrm{s} \quad 0.0 \mathrm{~m} / \mathrm{s}$
$\mathrm{CO} 2=7.443 \mathrm{E}+12 \mathrm{~N} 2=6.096 \mathrm{E}+12 \mathrm{Ar}=4.819 \mathrm{E}+11 \mathrm{O} 2=2.308 \mathrm{E}+11 \mathrm{CO}=2.939 \mathrm{E}+12 \mathrm{\#} / \mathrm{m}^{* *} 3$

| 32.021 | 16.697 | 1.882 | 0.722 | $8.048 \%$ by mass |
| :--- | :--- | :--- | :--- | :--- |

$\begin{array}{lllll}17.007 & 13.930 & 1.101 & 0.527 & 6.716\end{array} \%$ by volume
$\mathrm{O}=2.581 \mathrm{E}+13 \mathrm{He}=5.960 \mathrm{E}+11 \mathrm{H} 2=3.453 \mathrm{E}+10 \mathrm{H}=1.269 \mathrm{E}+11$ Total=4.376E+13 \#/m**3
$40.377 \quad 0.233 \quad 0.007 \quad 0.013 \%$ by mass MolWgt=23.374
$\begin{array}{llll}58.989 & 1.362 & 0.079 & 0.290 \% \text { volume (mole) fraction }\end{array}$

## APPENDIX D—SUMMARY OF FILES PROVIDED WITH MARS-GRAM 2010

The Mars-GRAM 2010 zip file contains the following directories:

| binFiles | PC binary version of data files required |
| :--- | :--- |
| Code | Source code for Mars-GRAM stand-alone and example trajectory programs |
| Documentation | README and other .txt and .pdf documentation files |
| Executables | PC executables for Mars-GRAM and the various Utilities programs executables <br> for other platforms must be compiled from the source code |
| IOfiles | Reference input and output files [sample output files ListMapYrx.txt (x $=0,1,2)$ ] |
| txtFiles | Text version of data files required. Source code makebin.f90 in Utilities directory <br> is a program to read these text files and create the binary files on non-PC systems <br> (see Conversion of ASCII Data to Binary in appendix F). |
| Utilities | Source code for utilities programs to perform various tasks related to Mars- <br> GRAM. |

NOTE: Text files provided have end-of-line marks in PC format. These may need to be converted to end-of-line marks on your non-PC platform. File transfer protocol (FTP) file transfer to your target platform from the zip files, read on a PC, can make the transform, if the FTP file transfer is done in ASCII mode. On-line or GNU utilities are also available that can perform this end-of-line transform for you.

The directories contain the following files:

## Directory Code:

Cfiles_M10_C.f90 Mars-GRAM common block modules
Ifiles_M10_I.f90 Mars-GRAM interface modules
marsgram_M10.f90 Source code for the "stand-alone" version main program
marssubs_M10.f90 Subroutines used by marsgram_M10, dumytraj_M10, and multtraj_M10
setup_M10.f90 Setup routines used by marsgram_M10, dumytraj_M10, and multtraj_M10
TESsubs_M10.f90 Subroutines to read and interpolate TES mapping year 1 and 2 data; used by marsgram_M10, dumytraj_M10, and multtraj_M10
multtraj_M10.f90
wrapper_M10.f90
ldMarsGRAM.bat
lddumytraj.bat
ldmulttraj.bat

## Directory Utilities:

makebin.f90
readalb.f90 Program to read ASCII albedos and convert to binary
READTOPO.f90 Program to read ASCII MOLA topography and convert to binary
BLDTRAJ.f90 Program to build pseudo-trajectory file for use in Mars-GRAM to compute output for maps or cross-sections
finddate.f90 Utility to find Earth date/time for desired $L_{s}$ or Mars time
julday.f90 Utility to find Julian day from year, month, day input
marsrad.f90 Uses Mars-GRAM output to compute various solar (shortwave) and thermal (longwave) fluxes at the surface and the top of the atmosphere
radtraj.f90

| LdBldTraj.bat | Example macro to compile and link program BLDTRAJ |
| :--- | :--- |
| LdFindDate.bat | Example macro to compile and link program finddate |
| LdJulDay.bat | Example macro to compile and link program julday |
| LdMakebin.bat | Example macro to compile and link program makebin |
| LdMarsRad.bat | Example macro to compile and link program marsrad |
| LdRadTraj.bat | Example macro to compile and link program radtraj |
| LdReadalb.bat | Example macro to compile and link program readalb |
| LdReadtopo.bat | Example macro to compile and link program READTOPO |

## Directory txtFiles:

sfc00xxy.txt MGCM boundary layer data at ground surface for Mars-GRAM 2001 data at dust optical depths $0.3,1.0$, and $3.0(x x=03,10,30)$ and for TES mapping years 1 and 2 $(x x=y 1, y 2)$, and version number $y$

MGCM boundary layer data at 5.0 m height for Mars-GRAM 2001 data at dust optical depths $0.3,1.0$, and $3.0(x x=03,10,30)$ and for TES mapping years 1 and 2 ( $x x=y 1, y 2$ ), and version number $y$

MGCM boundary layer data at 30.0 m height for Mars-GRAM 2001 data at dust optical depths $0.3,1.0$, and $3.0(x x=03,10,30)$ and for TES mapping years 1 and 2 ( $\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2$ ), and version number y

## MGCM data files (ASCII format) for 0 to 80 km or $\mathbf{- 5}$ to 80 km

Mars-GRAM 2001 MGCM 0 to 80 km temperature, pressure, and density data for 3 dust optical depths $(x x=03,10,30)$, version number $y$, or TES mapping year 1 and 2 data $(x x=y 1, y 2)$ for -5 to 80 km , version number y
uvloxxy.txt
Special "trajectory" building program to compute vertical profiles at latitudelongitude, latitude-time, or longitude-time cross sections, for input to Mars-GRAM runs to produce output for input to marsrad radiation calculations

Example macro to compile and link program BLDTRAJ

Example macro to compile and link program finddate

Example macro to compile and link program julday

Example macro to compile and link program READTOPO

## Data files (ASCII format) for MGCM near-surface data

sfc05xxy.tx
sfc30xxy.txt
tpdloxxy.txt
depths $(x x=03,10,30)$, version number $y$, or TES mapping year 1 and 2 data ( $x x=$ $\mathrm{y} 1, \mathrm{y} 2$ ) for -5 to 80 km , version number y

## Data files (ASCII format) for MTGCM 80 to 170 (or 240) km data

| tpdlsxxy.txt | Mars-GRAM 2001 MTGCM 80 to 170 km temperature, pressure, and density data for 3 dust optical depths ( $\mathrm{xx}=03,10,30$ ), version number y , for solar activity F10.7 = 70; or MTGCM data for 80 to 240 km for TES mapping years 1 and 2 ( $\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2$ ) |
| :---: | :---: |
| tpdmsxxy.txt | Mars-GRAM 2001 MTGCM 80 to 170 km temperature, pressure, and density data for 3 dust optical depths ( $\mathrm{xx}=03,10,30$ ), version number y , for solar activity F10.7 = 130; or MTGCM data for 80 to 240 km for TES mapping years 1 and 2 ( $\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2$ ) |
| tpdhsxxy.txt | MTGCM 80 to 240 km temperature, pressure, and density data for TES mapping years 1 and $2(x x=y 1, y 2)$, version number y , for solar activity F10.7 $=200$ |
| uvlsxxy.txt | MTGCM 80 to 170 km EW wind and NS wind data for 3 dust optical depths xx , version number y, for solar activity F10.7 $=70$; or 80 to 240 km for TES mapping years 1 and $2(x x=y 1, y 2)$ |
| uvmsxxy.txt | MTGCM 80 to 170 km EW wind and NS wind data for 3 dust optical depths xx , version number y , for solar activity F10.7 $=130$; or 80 to 240 km for TES mapping years 1 and $2(x x=y 1, y 2)$ |
| uvhsxxy.txt | MTGCM 80 to 240 km EW wind and NS wind data for TES mapping years 1 and 2 $(\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2)$ version number y , for solar activity F10.7 $=200$ |
|  | Other text files, which must be converted to binary format |
| albedol.txt | Global surface albedo at 1 by 1 degree latitude-longitude resolution; must be converted to binary with program readalb.f90 |
| MOLATOPH.TXT | MOLA areoid and surface topography at $1 / 2$ by $1 / 2$ degree latitude-longitude resolution; must be converted to binary with program READTOPO.f90 |
| TESdust1.txt | TES dust optical depth data versus $\mathrm{L}_{\mathrm{s}}$, latitude, longitude for TES mapping years 1 and 2 ; must be converted to binary with program makebin.f90 |

## Directory binFiles:

| sfc*.bin | Binary format MGCM surface data files |
| :--- | :--- |
| tpdlo*.bin | Binary format MGCM thermodynamic files |
| uvlo*.bin | Binary format MGCM wind files |
| tpdxs*.bin | Binary format MTGCM thermodynamic files ( $\mathrm{x}=1, \mathrm{~m}$, or h for low, medium, <br> or high solar activity) |


| uvxs*.bin | Binary format MTGCM wind files ( $\mathrm{x}=1, \mathrm{~m}$, or h for low, medium, or high solar <br> activity) |
| :--- | :--- |
| albedol.bin | Binary format albedo file |
| molatoph.bin | Binary format topography file |
| TESdust1.bin | Binary format dust file |
| COSPAR2.DAT | Text format data file for COSPAR reference model atmosphere |
| hgtoffst.dat | Text format height offset file |
| zfhtlsy.txt | Height ZF of 1.26 nbar level for all dust optical depths, version number y, <br> for solar activity F10.7 = 70 for Mars-GRAM 2001 MTGCM data |
| zfhtmsy.txt | Height ZF of 1.26 nbar level for all dust optical depths, version number y, <br> for solar activity F10.7 = 130 for Mars-GRAM 2001 MTGCM data |
| zfTESlsy.txt | Height ZF of 1.26 nbar level for TES mapping years 1 and 2, MTGCM data <br> version y, for solar activity F10.7 = 70 |
| zfTESmsy.txt | Height ZF of 1.26 nbar level for TES mapping years 1 and 2, MTGCM data <br> version y, for solar activity F10.7 = 130 |
| zfTEShsy.txt | Height ZF of 1.26 nbar level for TES mapping years 1 and 2, MTGCM data <br> version y, for solar activity F10.7 = 200 |

Note: For PC users, binary version MGCM files, MTGCM files, and albedol.bin, molatoph.bin, and TESdustl.bin are supplied and the readalb, readtopo, and makebin programs do not have to be run. For more details, see appendix F.

## Directory IOfiles:

| inputstd0.txt | Commented test input file for reference case using mapping year 0 MGCM <br> and MTGCM input data |
| :--- | :--- |
| inputstd1.txt | Commented test input file for reference case using TES mapping year 1 MGCM <br> and MTGCM input data |
| inputstd2.txt | Commented test input file for reference case using TES mapping year 2 MGCM <br> and MTGCM input data |
| ListMapYr0.txt | List output file for inputstd0.txt reference case |
| ListMapYr1.txt | List output file for inputstd1.txt reference case |
| ListMapYr2.txt | List output file for inputstd2.txt reference case |


| profiledata.txt | Sample auxiliary profile |
| :--- | :--- |
| dumytraj.bat | Example execute macro for dummy trajectory program dumytraj_M10 |
| marsgram0.bat | Example execute macro for mapping year 0 reference input |
| marsgram1.bat | Example execute macro for mapping year 1 reference input |
| marsgram2.bat | Example execute macro for mapping year 2 reference input |
| multtraj.bat | Example execute macro for dummy trajectory program multtraj_M10 |

## Directory Documentation:

| headers.txt | List of output files and file header definitions |
| :--- | :--- |
| marsfix.txt | List of code changes since Mars-GRAM 2010 Version 1.0 (Nov 2010) |
| marshist.txt | History file summarizing various Mars-GRAM versions and changes |
| Mars2000.pdf | A portable document format (PDF) version of the Mars-GRAM 2000 user guide |
| Mars2001.pdf | A PDF version of the Mars-GRAM 2001 user guide. |
| Mars2010.pdf | A PDF version of the Mars-GRAM 2010 user guide |
| MarsEnvironment.pdf | A PDF version of the Mars Transportation Environment Definition Document <br> (NASA/TM-2001-210935) |
| README1.txt | General program introduction file |
| README2.txt | Discussion of dumytraj_M10.f90 dummy trajectory program <br> README3.txt |
| grams to convert ASCII files into binary files for faster running of Mars-GRAM |  |

Plotable output files can be generated with data given versus several selected parameters. Generation of LIST file output and plotable output files is controlled by the value of iup on input.

For Mars-GRAM 2010, a number of plotable output files are generated, each containing several parameters suitable for plotting. These plotable files have headers to help identify parameters in the files.

File names and definitions of headers are given in the file headers.txt.
Plotable x and y parameters and their code values (set by input variables $N V A R X$ and $N V A R Y$ ) are as follows:

## Code Parameter

1 Planetocentric Height (above local MOLA areoid, km)
2 Planetocentric Height (above local MOLA topographic surface, km )
3 Planetocentric Latitude (deg.)
4 Longitude (deg.) West+ if Lon $E W=0$, East+ if LonEW $=1$
5 Time from start (Earth seconds)
6 Time from start (Martian Sols)
7 Areocentric Longitude of Sun, $\mathrm{L}_{\mathrm{s}}$ (deg.)
8 Local True Solar Time (Mars hours)
9 Pressure (mb)
10 Pressure Height $(\mathrm{km})\left[-\mathrm{H}^{*} \log (\right.$ Pres $/$ PresSurf $)=-\mathrm{H}^{*} \log ($ sigma $\left.)\right]$
11 Sigma coordinate [sigma $=$ Pressure/(Pressure at Surface)]
12 Planetocentric Height (km) above reference ellipsoid
13 Planetographic Height (km) above reference ellipsoid
14 Planetographic Latitude (deg.)
15 Longitude in range -180 to +180 (East or West, controlled by LonEW)

## APPENDIX E—EXAMPLE APPLICATION OF MARS-GRAM IN A TRAJECTORY CODE

With earlier versions of Mars-GRAM a dummy trajectory program, marstraj.f, was supplied. Beginning with Mars-GRAM version 3.8, an alternate version of a double precision dummy trajectory calculating program, dumytraj.f, was included. Although similar in general function to the original marstraj.f code, some details of dumytraj.f are different:
(1) In the original marstraj.f, interaction with Mars-GRAM was via calls to three subroutines:

$$
\begin{aligned}
& \text { Call Setup(...) } \\
& \text { Call Randinit(...) } \\
& \text { Call Datastep(...) }
\end{aligned}
$$

These three subroutines are part of the Mars-GRAM 2010 code and are automatically available to be called whenever the Mars-GRAM 2010 code (marssubs_M10.f90 and setup_M10.f90) is linked to the user's main trajectory driver program. If you already have a trajectory program built like this, with calls to Setup, Randinit, and Datastep it might be easily modified to incorporate Mars-GRAM 2010 subroutines without using the approach taken in dumytraj_M10.f90. Note, however, that the number of arguments in these subroutines has changed, so appropriate modifications in your trajectory programs must be made.
(2) In dumytraj_M10.f90, interaction with Mars-GRAM 2010 is via three calls to one wrapper subroutine named Marstraj_M10 and provided as file wrapper_ M10.f90, but with different values of three control parameters called isetup, jmonte, and istep:

$$
\begin{aligned}
& \text { Call Marstraj_M10(...) } \quad \text { with isetup }=1 \\
& \text { Call Marstraj_M10(...) } \quad \text { with isetup }=0, \text { jmonte }>0 \text {, istep }=0 \\
& \text { Call Marstraj_M10(...) } \quad \text { with isetup }=0, \text { jmonte }=0 \text {, istep }>0
\end{aligned}
$$

where isetup $=1$ triggers the call to the Setup subroutine, jmonte $>0$ triggers the call to the reinitialization process including the call to the Randinit subroutine, and istep $=1$ to MAXNUM is a counter for steps along the trajectory with a call to the Datastep subroutine at each step. Marstraj_M10 is a subroutine in the dumytraj_M10.f90 code, and must be included along with the basic Mars-GRAM code setup_M10.f90 and marssubs_M10.f90 as a subroutine in the user's calling trajectory program.
(3) In the original marstraj.f dummy trajectory main code, the transfer of double precision trajectory variables to and from single precision Mars-GRAM variables was assumed to be done within the user's main trajectory code. In the dumytraj_M10.f90 code this transfer is handled within the Marstraj_ M10 subroutine which must be included as a subroutine in the user's trajectory program.
(4) In the original marstraj.f dummy trajectory main code, single precision values of position increments $D E L H G T, D E L L A T$, and $D E L L O N$ were presumed to be calculated within the user's main trajectory program. In the dumytraj_M10.f90 code, input variables to the Marstraj_M10 subroutine are current and next double precision position values of height, latitude, and longitude and the position increments to be passed to the Datastep subroutine including increments of height, latitude, and longitude are computed within the Marstraj_M10 subroutine.

Regardless of which dummy trajectory code you decide to use as your starting model from which to build the interface to Mars-GRAM 2010 for your own trajectory code, it is worthwhile to read the comments embedded in the dumytraj_M10.f90 code. These comments give more explicit descriptions of the functions that are being performed. They also provide valuable hints about what to do if you are using predictor-corrector or other trajectory approaches that require mid-point corrections along trajectory steps and/or the use of density variations that occur within each trajectory step.

Another feature of dumytraj_M10.f90 is that it allows high precision Mars ephemeris values for sun latitude, longitude, and $L_{s}$ angle to be passed from the trajectory program for use by Mars-GRAM 2010 subroutines.

The multiple perturbation version, multtraj_M10.f90, allows different realizations of perturbation profiles to be generated on a single trajectory run. This can be used, for example, to test sensitivity of a guidance algorithm to different gain levels.

## APPENDIX F—DETAILS OF MGCM, MTGCM, AND MOLA DATA FILES

## Time-of-Day Variation of MGCM and MTGCM Data

ASCII format MGCM and MTGCM data files are provided, each having values for amplitudes and phases of diurnal and semi-diurnal components. The diurnal period is 24 Mars hours and the semidiurnal period is 12 Mars hours. Generically, the amplitudes and phases are as follows:

$$
\begin{aligned}
& \text { A0 }=\text { Diurnal mean value of the given parameter } \\
& \text { A1 }=\text { Amplitude of the diurnal tide component } \\
& \text { phi1 }=\text { Phase (local time in Mars hours) of the diurnal component } \\
& \text { A2 = Amplitude of the semi-diurnal tide component } \\
& \text { phi2 }=\text { Phase (local time in Mars hours) of the semi-diurnal component }
\end{aligned}
$$

For temperature, wind components, and height of 1.26 nbar level (ZF), data files give amplitudes in the same units as those of the parameter, K for temperature, $\mathrm{m} / \mathrm{s}$ for wind, or km for ZF . For pressure and density, data files give amplitudes in units of percent of the mean value A0. A0 units for pressure are $\mathrm{N} / \mathrm{m}^{2}$, while density A0 units are $\mathrm{kg} / \mathrm{m}^{3}$.

Three slightly different functions for tidal variation versus time of day are used. For MGCM and MTGCM temperatures and winds and MTGCM ZF heights, function TideX is used, where

$$
\operatorname{TideX}=\mathrm{A} 0+\mathrm{A} 1 * \cos ((\pi / 12) *(\text { time }- \text { phi1 }))+\mathrm{A} 2 * \cos ((\pi / 6) *(\text { time }- \text { phi2 }))
$$

and time is the local solar time in Mars hours. Note that units for A1 and A2 are same as those for A0 in this function form. For Mars-GRAM 2001, MGCM and MTGCM pressure and density data which have A1 and A2 in units of percent of A0, function TideY is used, where

$$
\operatorname{TideY}=\mathrm{A} 0 *(1+0.01 * \mathrm{~A} 1 * \cos ((\pi / 12) *(\mathrm{t}-\mathrm{phi1}))+0.01 * \mathrm{~A} 2 * \cos ((\pi / 6) *(\mathrm{t}-\mathrm{phi} 2)))
$$

TES MTGCM data extend to higher altitude, 240 km versus 170 km for Mars-GRAM 2001 MTGCM data. Consequently, tidal amplitudes for density and pressure for the new data grow to larger values than for the Mars-GRAM 2001 data. To accommodate this situation, an alternate model is adopted for pressure and density variation with time of day, whereby it is the log of pressure and density that are assumed to vary as cosine of time of day. Namely TES mapping year 1 and 2 MTGCM pressure and density tides are computed by

$$
\mathrm{tTideY}=\mathrm{A} 0 *((1.0 \mathrm{~d} 0+0.01 \mathrm{~d} 0 * \mathrm{~A} 1) * * \mathrm{c} 1) *((1.0 \mathrm{~d} 0+0.01 \mathrm{~d} 0 * \mathrm{~A} 2) * * \mathrm{c} 2)
$$

where exponents cl and c 2 are given by

$$
\begin{gathered}
\mathrm{c} 1=\cos ((\pi / 12) *(\text { time }- \text { phi1 })) \\
\mathrm{c} 2=\cos ((\pi / 6) *(\mathrm{t}-\mathrm{phi} 2)
\end{gathered}
$$

## Near-Surface MGCM Data

Near-surface MGCM data files are provided as follows in ASCII format:
sfc00xxy.txt MGCM temperature data at topographic surface
sfc05xxy.txt MGCM temperature and wind data at 5 m height above the surface
sfc30xxy.txt MGCM temperature and wind data at 30 m height above the surface
Naming convention for these files is:
$\mathrm{xx}=03,10,30$ for Mars-GRAM 2001 data at globally-uniform dust optical depths of 0.3, 1.0, and 3.0
$\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2$ for data at time-and space-variable dust optical depth as observed during TES mapping years 1 and 2
$\mathrm{y}=$ version number
Each record of these files contains $\mathrm{L}_{\mathrm{s}}$ value, latitude, longitude, and tidal coefficients (A0, A1, phi1, A2, phi2) for temperature, and for Eastward and Northward wind components, except for ground surface data files, which contain only temperature information.

## MGCM Data Up to 80.0 km Altitude

Files of zonally-averaged MGCM data up to 80.0 km altitude above the MOLA areoid are provided as follows in ASCII format:
tpdloxxy.txt MGCM temperature, pressure, and density data
uvloxxy.txt MGCM Eastward and Northward wind data
Naming convention for these files is:
$\mathrm{xx}=03,10,30$ for Mars-GRAM 2001 data at globally-uniform dust optical depths of 0.3, 1.0, and 3.0. These data are at 5.0 km intervals from 0.0 km to 80.0 km above MOLA
$\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2$ for data at time-and space-variable dust optical depths as observed during TES mapping years 1 and 2. These data are at 1.0 km interval from -5.0 km to +10.0 km above MOLA, and 5.0 km interval from 10.0 km to 80.0 km above MOLA
$\mathrm{y}=$ version number

For the MGCM data sets, each record of the tpdloxx.txt files contains $L_{s}$ value, height, latitude, and tidal coefficients for temperature and pressure. Only the A 0 coefficient is given for density. Tidal variations in density are computed from those for pressure and temperature by the perfect gas law relation. For the TES mapping year 1 and 2 data, each record of the tpdloxx.txt files contains $L_{s}$ value, height, latitude, and tidal coefficients for temperature and density. Only the A0 coefficient is given for pressure. Tidal variations in pressure are computed from those for density and temperature by the perfect gas law relation.

For both MGCM and TES mapping year 1 and 2 data, each record of the uvloxx.txt files contains $L_{s}$ value, height, latitude, and tidal coefficients for the Eastward and Northward wind components.

## MTGCM Data Up to $\mathbf{1 7 0 . 0}$ or $\mathbf{2 4 0 . 0} \mathbf{~ k m}$ Altitude

Zonally-averaged MTGCM data are provided in ASCII format at 5.0 km height resolution for the altitude range of 80.0 to 170.0 km for Mars-GRAM 2001 data and 80.0 to 240.0 km for TES mapping year 1 and 2 data. Data sets are provided for different levels of solar activity as characterized by 10.7 cm solar flux, F10.7, as measured at 1.0 AU .
tpdlsxxy.txt MTGCM temperature, pressure, and density data for low solar activity (F10.7 = 70)
tpdmsxxy.txt MTGCM temperature, pressure, and density data for moderate solar activity ( $\mathrm{F} 10.7=130$ )
tpdhsxxy.txt MTGCM temperature, pressure, and density data for high solar activity (F10.7 = 200; available for TES mapping years 1 and 2 only)
uvlsxxy.txt MTGCM Eastward and Northward wind data for F10.7 $=70$
uvmsxxy.txt MTGCM Eastward and Northward wind data for F10.7 $=130$
uvhsxxy.txt MTGCM Eastward and Northward wind data for F10.7 = 200 (available for TES mapping years 1 and 2 only)

Naming convention for these files is:
$\mathrm{xx}=03,10,30$ for Mars-GRAM 2001 data at globally-uniform dust optical depths of $0.3,1.0$, and 3.0.
$\mathrm{xx}=\mathrm{y} 1, \mathrm{y} 2$ for data at time-and space-variable dust optical depths as observed during TES mapping years 1 and 2.
$y=$ version number
Each record of the tpdlsxxy.txt, tpdmsxxy.txt and tpdhsxxy.txt files contains $L_{s}$ value, height, latitude, and tidal coefficients for temperature, pressure, and density. Because of height variations in molecular weight, tidal coefficients are retained for all three of these thermodynamic components. Each record of the uvlsxxy.txt, uvmsxxy.txt and uvhsxxy.txt files contains $L_{s}$ value, height, latitude, and tidal coefficients for the Eastward and Northward wind components.

Files zfhtlsy.txt, and zfhtmsy for Mars-GRAM 2001 data, F10.7 $=70$ or 130 and zfTESlsy.txt, zfTESmsy.txt, and zfTEShsy.txt for TES mapping years 1 and 2, F10.7 $=70$, 130, or 200 provide tidal coefficient information for altitude ZF, the height of the 1.26 nbar level. Each record of the ZF files contains dust optical depth or TES mapping year, $\mathrm{L}_{\mathrm{s}}$, latitude, and tidal coefficient values. The mean ZF value and tidal amplitudes are given in km .

## MOLA Areoid and Topography Data

MOLA areoid and topography data at $1 / 2$ by $1 / 2$ degree latitude-longitude resolution is provided in ASCII file MOLATOPH.TXT. Each line of this file contains East longitude and latitude at the center of the $1 / 2$ by $1 / 2$ degree grid box, grid-box-average radius in meters to the topographic surface, areoid radius in meters which is the radius to reference constant potential surface, evaluated at the center of the grid box, topography which is the grid-box-average difference in meters between the local planetary radius and areoid; this is analogous to the local terrain height above sea level for Earth, and Num which equals the number of laser shots averaged over the grid box. MOLA latitudes are planetocentric. Longitudes in the MOLA input file are with respect to the IAU 1991 prime meridian. A shift of about 0.24 degrees is made automatically within Mars-GRAM 2010, in order to convert to longitudes relative to the IAU 2000 prime meridian.

## Surface Albedo Data

Global surface albedo at 1 by 1 degree latitude-longitude resolution is given in file albedol.txt. Each line of this file contains latitude, and West longitude at the center of the 1 by 1 degree grid box, and grid-box-average surface albedo which is the ratio of the surface-reflected solar flux to that incident on the surface.

## Global Mean MGCM-MTGCM Height Offset Data

Height offsets can be used to control the smoothness of the transition at 80 km altitude between MGCM data and MTGCM data. Height offset options are controlled by input parameters zoffset and ibougher. The file README6.txt contains more details about height offsets and these input parameters. Option ibougher $=2$ causes the height offsets to be evaluated from global average height offset data given in the file hgtoffst.dat. In the first part of this file, global-average offsets in km are given for Mars-GRAM 2001 data as a function of $\mathrm{L}_{\mathrm{s}}$ and dust optical depth of $0.3,1.0$, and 3.0. In the second part of this file, global-average offsets are given versus $\mathrm{L}_{\mathrm{s}}$ for TES mapping years 1 and $2, \mathrm{y} 1$ or y 2 .

## TES Dust Optical Depth Data

Observed average dust optical depth from TES mapping years 1 and 2 are provided in the file TESdust1.txt. TES mapping year 1 was from April 1999 through January 2001. TES mapping year 2 was from February 2001 through December 2002. In terms of $\mathrm{L}_{\mathrm{s}}$, TES mapping years 1 and 2 cover from $\mathrm{L}_{\mathrm{S}}=$ 115 , through $\mathrm{L}_{\mathrm{s}}=360 / 0$, and back to $\mathrm{L}_{\mathrm{s}}=115$ the following Mars year. A conventional Mars year runs from $\mathrm{L}_{\mathrm{s}}=0$ to $\mathrm{L}_{\mathrm{s}}=360$. There were no global-scale dust storms during TES mapping year 1 . However, a very intense, global-scale dust storm began near the end of June 2001, during TES mapping year 2. Each line of file TESdust1.txt contains TES mapping year ( 1 or 2 ), $\mathrm{L}_{\mathrm{s}}$, latitude, West longitude, and the average TES optical depth expressed as visible-wavelength optical depth, approximately twice the optical depth values measured by TES at its 9 micron observing wavelength. Data in this file are at a resolution of 5.0 degrees in Ls, 7.5 degrees in latitude, and 9.0 degrees in longitude.

## COSPAR Reference Data

COSPAR Northern hemisphere mean reference data, as given in Pitts et $\mathrm{al}^{31}$, are provided in file COSPAR2.DAT. Each line of this file contains height (km), temperature (K), pressure (mbar), and density $\left(\mathrm{g} / \mathrm{cm}^{3}\right)$.

## Conversion of ASCII Data to Binary

Source code is provided for a program called makebin.f90 that reads the ASCII format MGCM and MTGCM data files and the TES dust optical depth data file and writes them out in binary format. After this ASCII-to-binary conversion is completed once, subsequent reading of the binary format files significantly shortens the time required to initialize Mars-GRAM on each run. Source code is also provided for the ASCII-to-binary conversion programs called READTOPO.f90 and readalb.f90 that covert the MOLA topography data and the albedo data files. ZF data files, height offset file, and COSPAR reference data file are sufficiently small that they can be read in ASCII format, so no conversion to binary is required for these files.

Note: For PC users, all necessary binary version files are supplied and the ASCII-to-binary conversion programs (readalb, readtopo, and makebin) do not have to be run.

To run Mars-GRAM, the binary-version MGCM and MTGCM files, together with the ASCIIformat ZF data files, must be in the directory whose pathname is given by the input parameter GCMDIR in the NAMELIST input file. The binary version albedo, MOLA topography, and TES dust optical depth data files, together with ASCII-format files COSPAR2.DAT and hgtoffst.dat must be in the directory whose pathname is given by the input parameter DATADIR in the NAMELIST input file. In this distribution of Mars-GRAM, the default pathname for DATADIR and GCMDIR are the same, with all necessary files together in the binFiles directory.

## Optional Trajectory Input Files

If the number of positions to be calculated given by the input parameter NPOS is set to 0 , an optional trajectory input file is read from a file with the name given by the input parameter TRAJFL. Each line of the trajectory file consists of: time, in seconds past the start time specified in the NAMELIST input, height, in km, latitude in degrees, and longitude in degrees. Heights are relative to the MOLA areoid or reference ellipsoid, as set by the input parameter MOLAhgts. Latitudes are planetocentric or planetographic, as set by the input parameter ipclat. Longitudes are East or West, as set by the input parameter $L o n E W$. Any additional reference information included on each line (e.g. orbit number, measured density, etc.) is ignored. Trajectory positions in these files do not have to be at small time or space steps. For example, a trajectory file may consist of successive periapsis times and positions for a simulated or observed aerobraking operation. Trajectory files may also contain arrays of locations used for computing heightlatitude cross sections or latitude-longitude cross sections. Such trajectory input files can be built by the program BLDTRAJ.f90 (see appendix G).

## Optional Auxiliary Profile Input File

As an option, data read from an auxiliary profile may be used to replace data from the MGCM and MTGCM data files. This option is controlled by setting the input parameters profile, profnear, and proffar in the NAMELIST input file. Input parameter profile gives the file name of the file containing the profile data values. Input parameter profnear is the latitude-longitude radius in degrees within which the
weight for the auxiliary profile is 1.0 . Input parameter proffar is the latitude-longitude radius in degrees beyond which the weight for the auxiliary profile is 0.0 . A weighting factor for the profile data, profwgt, having values between 0 and 1 , is applied between the proffar and profnear radii. Mean conditions are as given in the profile file if the desired point is within a latitude-longitude radius of profnear from the profile latitude-longitude at the given altitude; mean conditions are as given by the original MGCM or MTGCM data if the desired point is beyond a latitude-longitude radius of proffar from the latitude-longitude of the profile at the given altitude. If profnear $=0$, then profile data are NOT used. The profile weight factor, profwgt, for the auxiliary profile also varies between 0 at the first profile altitude level and 1 at the second profile altitude level and between 1 at the next-to-last profile altitude level and 0 at the last profile altitude level. First and second profile points and the next-to-last and last profile points should therefore be selected widely enough apart in altitude that a smooth transition can occur as profwgt changes from 0 to 1 near these profile end points.

Each line of the auxiliary profile input file consists of: height, in km, latitude, in degrees, longitude, in degrees, temperature, in K, pressure, in $\mathrm{N} / \mathrm{m}^{2}$, density, in $\mathrm{kg} / \mathrm{m}^{3}$, Eastward wind, in $\mathrm{m} / \mathrm{s}$, and Northward wind, in $\mathrm{m} / \mathrm{s}$. Heights are relative to the MOLA areoid or reference ellipsoid, as set by the input parameter MOLAhgts. Latitudes are planetocentric or planetographic, as set by the input parameter ipclat. Longitudes are East or West, as set by the input parameter LonEW. MGCM/MTGCM temperature, pressure, and density data are used if any of the profile inputs for temperature, pressure, or density are 0 . MGCM/MTGCM winds are used if BOTH wind components are 0 in the profile file. A sample auxiliary profile file named profiledata.txt is provided in the IOfiles directory.

## APPENDIX G—AUXILIARY PROGRAMS FOR USE WITH MARS-GRAM

Four auxiliary programs are provided for use with Mars-GRAM: BLDTRAJ.f90, finddate.f90, marsrad.f90, and radtraj.f90. The stand-alone version Mars-GRAM program files (marsgram_M10.f90, marssubs_M10.f90, and setup_M10.f90) are discussed in appendix D. Appendixes D and E discuss dummy trajectory programs dumytraj_M10.f90 and multtraj_M10.f90, which provide examples of how to use Mars-GRAM as a subroutine in trajectory programs or orbit propagator programs. Programs makebin. f90, readalb.f90, and READTOPO.f90, which convert data provided in ASCII format into binary files, are discussed in appendixes D and F.

## Program BLDTRAJ

It is frequently desirable to produce Mars-GRAM output for graphing as a map (i.e. latitudelongitude cross section at a given height) or other cross-section (e.g. height- latitude cross section at a given longitude). Program BLDTRAJ.f90 generates a trajectory file with input lines containing time, height, latitude, and longitude that can be used as Mars-GRAM input for generation of such maps or cross-sections. Program BLDTRAJ is interactive and prompts the user to input starting values, ending values, and step sizes for height ( $\mathrm{z} 1, \mathrm{z} 2, \mathrm{dz}$ ), latitude (lat1, lat2, dlat), and longitude (lon1, lon2, dlon). The program also prompts for a value of time increment which is applied between each trajectory step. This time increment may be 0 , if all trajectory points are at the same time. Time values in the trajectory file are time in seconds from the start time that is specified by the date and time information provided in the Mars-GRAM NAMELIST-format input file.

NOTE: If heights $>3,000.0 \mathrm{~km}$ are used, they are interpreted as planetocentric radius values $(\mathrm{km})$.
Example:
For a latitude-longitude map at height 10.0 km above the MOLA areoid, between latitudes -30 and 30 degrees in steps of 5 degrees, and longitudes 0 to 180 degrees in steps of 20 degrees, enter 10100 for zl, z2, dz; enter -30 305 for lat1, lat2, dlat; and enter 018020 for lon1, lon2, dlon. All of these input quantities are of type real, and can be entered to one or more significant digits beyond the decimal.

## Program finddate

Program finddate.f90 allows calculation of the areocentric longitude of the Sun, $\mathrm{L}_{\mathrm{s}}$, and Mars local true solar time, LTST, for a given Earth date and time. It also computes the next occurrence beyond the initial input date and time of the Earth date and time for which $\mathrm{L}_{\mathrm{s}}$ and LTST are any desired values. Accuracy information and other documentary comments are given within the source code. The program is interactive and prompts for all necessary inputs.

## Program julday

Program julday.f90 calculates the Julian day from year, month, and day input. Traveling wave components require the Julian day for initialization of wave phases. The program is interactive and prompts for all necessary inputs.

## Program marsrad

Program marsrad.f90 uses Mars-GRAM output files containing height profile information to compute various solar (shortwave) and thermal (longwave) fluxes at the surface and the top of the atmosphere. These profiles must start at the surface, should usually extend upward to a height of from 10.0 to 30.0 km , and may be at any desired height resolution with a limit of 1,000 points per profile. Mars-GRAM output files used by marsad.f90 are TPresHgt.txt, Density.txt, and MarsRad.txt. See the file headers.txt that is provided in the Documentation Directory of the Mars-GRAM 2010 zip file for a description of the parameters contained in these files.

Program marsrad.f90 runs interactively, with the only user input required being the number 1 or 2 of plot variables Var_X and, optionally, Var_Y used in the Mars-GRAM output files. The marsrad program computes various solar (shortwave) and thermal (longwave) fluxes at the surface and the top of the atmosphere. Two marsrad output files are produced. Output file Radlist.txt contains an annotated set of radiation fluxes, equivalent black-body temperatures, and albedos. Output file Radout.txt, suitable for input to a plot program, contains fluxes and other information in one line for each set of output. With an input trajectory file that is generated by program radtraj.f90 - see below, output file Radout.txt can be used to plot solar and thermal radiation data as a map lat-lon cross section, or as latitude-time or longitudetime cross sections.

Longwave radiative fluxes are computed by a broad-band (emissivity) method, patterned after Savijarvi ${ }^{32}$. Dust optical depth tau is for the shortwave (solar) spectrum. For longwave calculations, infrared emissivity versus shortwave solar optical depth curves are used, adapted from Haberle et al. ${ }^{27}$. Infrared emissivities for $\mathrm{CO}_{2}$ and water vapor are functions of pressure-scaled optical path lengths with emissivities from Staley and Jurica ${ }^{33}$. Shortwave fluxes are computed from total dust optical depth adjusted for small amount of clear-sky optical depth by a delta-Eddington method ${ }^{34}$. Both longwave and shortwave effects of water vapor are included, with relative humidity assumed constant at $20 \%{ }^{32}$. Dust optical properties assumed are 0.7 for asymmetry parameter, and 0.9 for single-scatter albedo. Other reasonable values may be found in Table 1 of Murphy et al. ${ }^{35}$. Values of asymmetry parameter and single-scatter albedo as well as assumed relative humidity can be changed in data statements near the beginning of the marsrad program.

Output parameters given in Radlist.txt output file in addition to plot variables, Var_X and, optionally, Var_Y are:

## tau Total vertical dust optical depth

MarsAU Mars orbital radius (AU)
mu0 Cosine of solar zenith angle
ice $\quad 0$ for no ice on the surface, 1 for ice on the surface (affects surface albedo)
Longwave (LW) fluxes F (W/m²):
Fdown(sfc) Downwelling LW flux at surface
Fup(sfc) Upwelling LW flux at surface (related to Tsfc)
Fup(toa) Upwelling LW flux at top-of-atmosphere

Femit(atmos) LW flux emitted by atmosphere [Fup(toa) - net LW at sfc]
Radiative (equivalent black-body) temperatures (K):
Tsky(sfc) Equivalent sky temperature [related to Fdown(sfc)]
Tsfc Ground surface temperature [related to Fup(sfc)]
Teff(toa) Effective black-body temperature at top-of-atmosphere
Shortwave (SW) fluxes E (W/m²):

| Edown(sfc) | Downwelling SW flux at surface |
| :--- | :--- |
| Eup(sfc) | Upwelling SW flux at surface [albedo times Edown(sfc)] |
| Eup(toa) | Upwelling SW flux at top-of-atmosphere |
| Eabsorb(atmos) | Net SW flux absorbed by atmosphere |
| Edown(toa) | Solar flux at toa = mu0*(solar constant)/MarsAU ${ }^{2}$ |
| rface albedo | Surface albedo interpolated from file albedol.txt |
| netary albedo | Ratio Eup(toa)/Edown(toa) |

SW+LW Fluxes (W/m²):

| Absorbed(sfc) | SW+LW flux absorbed at the surface |
| :--- | :--- |
| Emitted(toa) | Upwelling SW+LW flux at top-of-atmosphere |

Controlled by parameter "heatrate" set via user input, marsrad.f90 also outputs optical path lengths for water vapor $\left(\mathrm{H}_{2} \mathrm{O}\right), \mathrm{CO}_{2}$, and dust. $\mathrm{H}_{2} \mathrm{O}$ and $\mathrm{CO}_{2}$ optical path lengths are scaled by pressure to the 0.75 power. With heatrate set to 1 , the program also outputs various fluxes $\left(\mathrm{W} / \mathrm{m}^{2}\right)$ and heating rates ( $\mathrm{K} /$ day ) as a function of pressure level. These optional outputs are:

| Pres | Pressure (mb) |
| :--- | :--- |
| uH 2 O | Pressure-scaled $\mathrm{H}_{2} \mathrm{O}$ optical path (precipitable micrometers) |
| uCO 2 | Pressure-scaled $\mathrm{CO}_{2}$ optical path (atmosphere-centimeters) |
| udust | Dust optical depth from surface to given pressure level |
| LWFup | Upwelling LW flux at pressure level |
| LWFdn | Downwelling LW flux at pressure level |


| LWFnet | Net (up - down) LW flux at pressure level |
| :--- | :--- |
| LWdTdt | LW heating rate at pressure level |
| SWdTdt | SW heating rate at pressure level |
| TotdTdt | Total (LW+SW) heating rate at pressure level |

Parameters given in Radout.txt output file in addition to plot variables, Var_X and, optionally, Var_Y are:

| albsfc | Surface albedo (interpolated from file albedo1.txt) |
| :---: | :---: |
| tau | Total vertical dust optical depth (for solar wavelengths) |
| RadAU | Mars orbital radius (AU) |
| mu0 | Cosine of solar zenith angle |
| ice | 0 for no ice on the surface, 1 for ice on the surface |
| Tsfc | Ground surface temperature (K) |
| Fusfc | Upwelling LW flux at surface ( $\mathrm{W} / \mathrm{m}^{2}$ ) |
| Tsky | Equivalent sky temperature (K) |
| Fdsfc | Downwelling LW flux at surface ( $\mathrm{W} / \mathrm{m}^{2}$ ) |
| Teff | Effective black-body temperature at top-of atmosphere (K) |
| Futoa | Upwelling LW flux at top-of-atmosphere ( $\mathrm{W} / \mathrm{m}^{2}$ ) |
| Edsfc | Downwelling SW flux at surface ( $\mathrm{W} / \mathrm{m}^{2}$ ) |
| Eusfc | Upwelling SW flux at surface ( $\mathrm{W} / \mathrm{m}^{2}$ ) |
| Edtoa | Solar flux at toa $=\mathrm{mu} 0 *($ solar constant $) / \mathrm{RadAU}^{2}\left(\mathrm{~W} / \mathrm{m}^{2}\right)$ |
| Eutoa | Upwelling SW flux at top-of-atmosphere ( $\mathrm{W} / \mathrm{m}^{2}$ ) |
| Planalb | Planetary albedo $=$ ratio Eutoa/Edtoa |
| thet | Solar zenith angle (degrees) |
| Tdif | Diffuse transmittance for diffuse irradiance |
| Tdir | Diffuse transmittance for beam irradiance |
| Tbeam | Beam transmittance |

## Program radtraj

Radtraj. 990 is a special trajectory building program to compute vertical profiles at latitude-longitude, latitude-time, or longitude-time cross sections, for input to Mars-GRAM runs to produce output for input to marsrad radiation calculations. The radtraj.f90 program is similar in function and use to the BLDTRAJ.f90 trajectory building program, except that radtraj.f90 is especially designed for use in conjunction with Mars-GRAM runs for which radiation calculations are to be done by the marsrad.f90 program.

Program radtraj.f90 produces a trajectory file consisting of sets of vertical profiles that are constrained to start at the surface by automatically setting $\mathrm{zl}=-10$. Any heights that are below the surface are automatically ignored by the marsrad.f90 program. Any upper height z2 and height step dz may be used for the height profiles. However, it is not necessary to use z2 higher than about 30 km . The input time step if other than zero applies only as the trajectory goes from one vertical profile to the next.

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