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Improvements in Modeling Thruster Plume Erosion Damage to Spacecraft Surfaces



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Introduction



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- Spacecraft bipropellant thrusters impact spacecraft surfaces with high speed droplets of unburned and partially burned propellant.
 - These impacts can produce erosion damage to optically sensitive hardware and systems (e.g., windows, camera lenses, solar cells and protective coatings)
- On the International Space Station (ISS), operational constraints are levied on the position and orientation of the solar arrays to mitigate erosion effects during thruster operations
- In 2007, the ISS Program requested evaluation of erosion constraint relief to alleviate operational impacts due to an impaired Solar Alpha Rotary Joint (SARJ)
- Boeing Space Environments initiated an activity to identify and remove sources of conservatism in the plume induced erosion model to support an expanded range of acceptable solar array positions
- The original plume erosion model over-predicted plume erosion and was adjusted to better correlate with flight experiment results
- This paper discusses findings from flight experiments and the methodology employed in modifying the original plume erosion model for better correlation of predictions with flight experiment data
- The updated model has been successful employed in reducing conservatism and allowing for enhanced flexibility in ISS solar array operations

Thruster Operations & Plume Effects



Impacts to Optically Sensitive Surfaces

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Thruster plume induced erosion of solar arrays

- Solar cell coverglass is damaged by droplet impacts. Damage impacts the performance of UVE filter coatings and increases optical scatter on the coverglass.
- The laminates used on solar array thermal side can also be damaged. Thin silicon oxide (SiOx) coatings protect Kapton from Atomic Oxygen erosion.

Representative Solar Array Construction





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Plume Induced Erosion Concerns

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Erosion events of concern for the ISS solar arrays

- Soyuz/Progress approach and separation to Russian Segment docking ports
- Soyuz relocations
- Russian Segment reboost and attitude control
- Soyuz thruster tests
- Progress and Service Module thruster tests
- Commercial cargo transportation vehicles approach and separation
- Commercial crew transportation vehicles approach and separation
- Orbiter approach and separation to PMA2
- Orbiter reboost and attitude control



ISS Solar Array Constraints

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Example of Current Erosion Constraint Table

Service Module Roll/Pitch/Yaw Attitude Control - Inboard SAW Example

 Results for each alpha/beta combination represent maximum erosion on the entire solar array surface if that particular solar array alpha/beta combination is selected for every event within a year

																	2/	4																		
2A SARJ	0	10	20	30	40	50	60	70	80	90	100	110	120	130	140	150	160	170	180	190	200	210	220	230	240	250	260	270	280	290	300	310	320	330	340	350
0				8131		2261	432	0.00	0.00	0.00	68.2	359	1050	1920		6111								2461	555	0.00	0.00	0.00	51.3	381	1104	1998		6313	8019	
15		***	1221	506	119	8.75	0.00	0.00	0.00	64.6	581	1525	***	***	7414	***	***			2410	1363	585	160	8.08	0.00	0.00	0.00	36.6	447	1276	3418		7410	7141		5651
30		***	1812	837	100	0.00	0.00	0.00	586	3126	***				***		***			***	1683	784	134	0.00	0.00	0.00	446		***							
45			1641	265	0.00	0.00	0.00	26.4	638	1352		3312			***			6921			1712	334	0.00	0.00	0.00	7.22	645	1466	2177	3512			7021	7617		
60	855	492	129	0.00	0.00	0.00	0.00	21.7	65.9	167	314	654	1144	1646		2241	2134	1598	893	532	148	0.00	0.00	0.00	0.00	21.5	72.1	182	343	702	1229	1761	2173		2187	1587
75	53.3	17.2	2.28	0.00	0.00	0.00	0.00	0.98	3.43	12.4	26.4	52.4	92.9	136	167	171	142	91.2	58.2	20.4	3.30	0.00	0.00	0.00	0.00	1.03	3.84	13.8	29.8	57.7	101	146	176	175	141	86.5
90	1.46	0.44	0.02	0.00	0.00	0.00	0.00	0.02	0.12	0.40	1.09	1.98	3.60	4.95	5.97	5.85	4.51	3.27	1.67	0.52	0.03	0.00	0.00	0.00	0.00	0.03	0.14	0.44	1.23	2.17	3.84	5.28	6,19	5.86	4.33	3.08
105	0.19	0.05	0.00	0.00	0.00	0.00	0.00	0.01	0.05	0.21	0.43	0.90	1.50	2.08	2.25	1.86	1.16	0.59	0.23	0.07	0.01	0.00	0.00	0.00	0.00	0.01	0.06	0.23	0.47	0.96	1.55	2.09	2.17	1.72	1.02	0.50
120	0.08	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.07	0.20	0.33	0.55	0.71	0.75	0.63	0.43	0.22	0.10	0.03	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.08	0.21	0.34	0.55	0.69	0.70	0.57	0.37	0.18
135	0.03	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.05	0.13	0.18	0.22	0.24	0.21	0.15	0.08	0.04	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.05	0.12	0.18	0.20	0.22	0.18	0.12	0.06
150	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.05	0.08	0.09	0.07	0.06	0.04	0.02	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.04	0.07	0.08	0.06	0.05	0.03
165	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.02	0.02	0.02	0.02	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.02	0.02	0.01
180	0.01	0.01	0.02	0.02	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.02	0.02	0.02	0.02	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01
195	0.02	0.04	0.06	0.08	0.09	0.07	0.04	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.03	0.05	0.08	0.10	0.10	0.08	0.04	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01
210	0.06	0.11	0.19	0.26	0.29	0.27	0.24	0.14	0.05	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.03	0.08	0.13	0.22	0.30	0.32	0.29	0.26	0.14	0.05	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02
225	0.15	0.34	0.65	0.95	1.12	1.06	0.80	0.48	0.29	0.09	0.02	0.00	0.00	0.00	0.00	0.00	0.01	0.06	0.19	0.41	0.74	1.05	1.19	1.09	0.80	0.46	0.27	0.08	0.02	0.00	0.00	0.00	0.00	0.00	0.01	0.05
240	0.39	1.03	2.01	3,19	3.83	3.54	2.56	1.57	0.72	0.35	0.08	0.00	0.00	0.00	0.00	0.00	0.03	0.16	0.48	1.22	2.25	3.42	3.94	3.52	2.45	1.46	0.66	0.32	0.07	0.00	0.00	0.00	0.00	0.00	0.02	0.12
255	0.97	2.77	5.69	9.06	9.66	9.84	6.88	3.85	1.84	0.76	0.23	0.04	0.00	0.00	0.00	0.00	0.06	0.40	1.20	3.20	6.24	9,45	9.60	9.42	6.43	3.49	1.64	0.65	0.20	0.04	0.00	0.00	0.00	0.00	0.04	0.30
270	2.19	6.01	12.2	16.4	13.6	15.3	13.2	7.21	3.34	1.28	0.42	0.08	0.00	0.00	0.00	0.00	0.15	0.91	2.67	6.86	13.2	16.7	13.4	14.5	12.2	6.48	2.94	1.09	0.36	0.07	0.00	0.00	0.00	0.00	0.10	0.70
285	55.0	105	146	163	148	110	68.8	36.0	16.4	6.91	2.03	0.61	0.00	0.00	0.00	0.00	8.59	32.5	58.7	107	143	155	137	101	61.9	32.3	14.4	6.21	1.79	0.61	0.00	0.00	0.00	0.00	7.31	29,1
300	906	1267	1422	1373	1129	808	516	298	155	103	43.1	0.09	0.00	0.00	0.00	55.1	261	483	893	1214	1344	1287	1054	754	480	275	142	92.5	39.8	0.33	0.00	0.00	0.00	43.3	241	483
315	1340	1752	2167	***	2361	2413	2125	1605	882	472	67.0	0.00	0.00	0.00	64.4	310	491	845	1255	1646	***	2101	***		1997	1528	841	486	88.7	0.00	0.00	0.00	52.0	319	538	911
330		5816	5501	***	***	***	1192	496	157	28,4	0.00	0.00	0.00	7.02	426	1273	***				***	4431			1288	531	200	26.4	0.00	0.00	0.00	6.77	332	1140		
345			8918	***		2811	1691	747	54.6	0.00	0.00	0.00	751	2681			***	9781			***	7144			1603	709	98.0	0.00	0.00	0.00	660					
Interpolati	on_N	/leth	o d		Maxi	imum		Gre	en_	Min		0.00		Yel	ow	Min		0.10		Re	ed_M	lin		1.00												

NASA JSC 29181 Plume Model

Original ISS Plume Erosion Model



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 Original liquid droplet distribution/flux model used by Space Environments was developed by NASA (JSC 29181)

Reference: Larin, M.E., "Model For Predicting Liquid Droplet Distribution and Velocities In The Plumes Of Small Bipropellant Thrusters," JSC 29181, September 2000

Model was designed to determine feathering angles for ISS sensitive surfaces

- Quantifies liquid droplet distribution and velocities in the thruster plume through the analysis of the two-phase droplet-gas flow
- Model was based on ground-based vacuum chamber data, and not correlated against light experiment data [SPIFEX, PIC]
- Space Environments' assessment of model predictions with SPIFEX flight experiment measurements demonstrated overestimation of droplet flux when compared to flight experiment data
- Plume erosion model was reassessed and updated to correlate with flight and high-speed impact test data



Plume Erosion Model



- A value of 4.57 was determined for 'p' through an iterative procedure requiring a close match with droplet mass angular distribution results from a MBB 10 N thruster test by H. Trinks.
- Implementation by Space Environments of the droplet distribution model follows AIAA 2001-2816:



Plume Contamination Model

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The plume erosion model uses the total liquid-phase contaminant flux from the plume contamination model (AIAA 2002-2016):



- r is range (e.g. in cm)
- \succ θ is angle off thruster plume centerline (degrees)
- T is thrust (N)
- \blacktriangleright \dot{M}_{T} is the total propellant mass flow rate (e.g. in g/s)
- R_{exit} is the nozzle exit radius (e.g. in cm; note this is just a reference distance)
- I_{noz} is nozzle length (e.g. in cm)
- \succ θ_0 is the "dispersion coefficient" (5 degrees)

Reference: Soares et al., "International Space Station Bipropellant Plume Contamination Model," AIAA 2002-3016, June 2002



Flight Experiment Data Pitsweeper Image Analysis

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- Individual images were loaded into Pitsweeper image analysis program
- Pits were identified by user and separated into two groups
 - Pit (red bar)
 - Potential Pit (green bar)
- Pitsweeper outputed pit size distribution



Observed Erosion of SPIFEX Samples

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Green bars represent potential pits, or features in images of the samples suggestive of pits, but not well-resolved.

 The Kapton and aluminum samples were adjacent in the experiment. Significant differences between the plots above indicate that surface material properties (i.e., hardness) has an important effect.



Comparison of SPIFEX Kapton Pitting Fluences with Fluences Calculated with JSC 29181 Model



JSC 29181 model assumptions



Update to Plume Erosion Model

- JSC 29181plume model assumes a maximum droplet diameter of 100 µm
- Flight experiment measurements support a maximum droplet size of 12 to 24 µm (based on calculations showing that droplets produce pits one to two times their diameter):
 - Top five pit diameters (µm) on SPIFEX samples:
 - Kapton: 20.3, 19.7, 18.4, 17.9, 17.4
 - Aluminum: 19.8, 15.5, 15.5, 15.5, 15.3
 - Largest pit diameter (µm) reported for PIC experiment: 24
- Boeing Space Environments conducted a parametric study and arrived at an updated plume model producing good correlation with the SPIFEX Kapton data. This was achieved by lowering the maximum droplet size and adjusting the droplet size distribution coefficient p



Initial Contaminant Flux

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For the thrust level of interest (100-3870N), the contamination ratio (KT^{-β}) was revised to fit the on-orbit data from the PIC flight experiment.



Comparison of SPIFEX Flight Experiment Data with Boeing Space Environments' Updated Plume Model

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Updated Boeing Plume Model

Model assumptions/notes

- Droplet size distribution coefficient: p = 1.70
- The maximum droplet diameter is 12 µm <u>Requires</u> 1:2 droplet-to-pit size assumption (allows pits up to 24 µm, as reported for the PIC flight experiment)
 - 1:1 droplet-to-pit ratio did not correlate



Summary of Boeing Space Environments' Plume Model Update

- Simulations of SPIFEX Kapton pitting levels with the JSC 29181 plume erosion model demonstrated that model needed to be updated
- Boeing Space Environments' plume model update correlates well with observed SPIFEX Kapton pitting results when:
 - Maximum droplet diameter changed from 100 µm to 12 µm based on observed maximum pit diameters (and based on 1:2 relationship of droplet diameter to pit diameter)
 - Droplet limiting angle calculation updated per input from NASA JSC Aerosciences
 - Droplet size distribution coefficient "p" (which controls particle size distribution) changed from 4.57 to 1.70



Returned Node 1 Nadir Window Hyzod Cover

- Node 1 nadir window Hyzod cover was deployed on ISS from Flight UF2 (June 2002) to Flight LF1 (July 2005) during which it was exposed to thruster firings for various Soyuz and Progress proximity operation events (approaches and separations of vehicles)
- NASA JSC Materials Evaluation Laboratory performed a microscopy imaging survey on the returned Hyzod material. Image survey results were delivered to the Space Environments Team for analysis
- Boeing Space Environments team developed an image analysis technique for this study to measure the damage crater sizes and performed the analysis
 - Top five impact feature (pit) diameters (μm): 16.6, 13.5, 7.5, 7.4, 7.3



Comparison Between Observed Hyzod Pitting Levels with Boeing Updated Plume Model Calculations



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Incident Droplet Fluence





 Hyzod, as well as aluminum, which are harder materials than Kapton, demonstrate lower pit damage than Kapton which is a softer material



Ground Testing & Analyses



Hypervelocity Impact (HVI) tests supported of reduction of solar array constraints from thruster plume induced erosion

- HVI test program was conducted by the NASA JSC Hypervelocity Impact Technology Facility (HITF)
 - Light gas gun testing conducted at the NASA JSC White Sands Test Facility (WSTF)
- Detailed test objectives:
 - Address impact craters as a function of particle size, particle velocity and impacted material
 - Define calibration data for SPHINX
 - Assess solar cell power collection degradation as a function of surface damage
 - Assess scrim cloth mechanical damage due to particle impacts



Summary of Ground-Based Light Gas Gun Test Results

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- High-speed impact testing (using a light gas gun system) was performed at the NASA WSTF on Hyzod, solar cell, and scrim cloth samples
- Optical microscopic imaging was performed on the samples at the HITF prior to and after the samples were shot with the light gas gun
- The core damage diameter was estimated from the test samples images
- Results show that the ratio of the core damage diameter/projectile diameter was less than 2.0 for the samples that were tested (Hyzod, solar cell, calibration samples)
 - SPHINX analysis results confirm core damage diameter / projectile diameter less than 2.0 for plume droplet impacts



Example Solar Cell Ground Test Image 5 micron impacting at 1.82 km/s HITF08212 Post Test image (5000X)



Application of Methodology

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2A 2A 0 10 20 30 40 50 60 70 80 90 100 110 120 130 140 150 60 70 80 90 100 110 120 130 140 150 160 170 180 190 200 210 220 230 240 250 260 270 280 290 300 310 320 330 340 350 2.13 25.3 104 264 451 729 961 1157 1339 1298 1198 994 636 329 66.8 0.00 0.00 0.00 .07 0.00 0.00 0.00 0.09 11.1 29.2 76.9 109 124 147 118 103 92.6 78.6 42.3 16.5 4.24 0.07 0.00 0.00 0.00 0.52 8.05 212 106 61.2 23.0 2.89 <mark>0.00 0.00 0.00</mark> 90.7 175 317 540 616 584 590 480 383 316 198 98.6 56.0 20.7 3.59 <mark>0.00 0.00 0.00</mark> 22.2 158 309 551 212 138 77,8 12.3 **0.00 0.00 0.00 0.74** 20.0 37.2 54.1 88.2 128 144 160 180 203 231 199 131 80.0 156 <mark>0.00 0.00 0.00 0.13</mark> 21.5 41.7 59.5 95.8 60 5.14 11.1 20.2 27.7 33.9 39.7 43.9 34.1 19.6 0.03 0.11 0.23 0.48 0.88 1.34 1.69 1.77 1.50 0.97 0.62 0.24 105 120 135 150 165 180 195 210 225 240 255 270 285 300 13.6 17.2 15.5 13.4 11.2 9.01 6.39 3.51 1.79 1.17 0.58 0.00 0.00 0.00 0.86 4.56 7.53 13.3 16.2 14.4 12.3 10.2 8.20 5.85 3.19 1.60 1.02 0.52 0.00 0.00 0.00 0.00 9.2 20.5 21.8 25.5 26.3 26.2 28.6 25.9 15.3 9.46 1.24 0.00 0.00 0.00 1.06 5.27 7.16 11.8 17.5 19.2 20.3 23.8 24.6 24.3 26.4 24.2 14.4 9.58 1.64 0.00 0.00 0.00 0.88 5.62 7.1 24 387 362 280 207 118 50.4 22.7 5.34 0.48 0.00 0.00 0.00 0.00 0.19 18.1 65.2 159 272 336 383 344 257 185 103 43.5 19.3 4.67 0.44 0.00 0.04 0.00 0.05 12.7 52.3 330 613 366 193 56.6 3.56 0.00 0.00 0.00 108 371 561 840 1075 1220 133 1260 1065 861 590 354 185 53.9 5.81 <mark>0.00 0.00 0.00</mark> 95.6 Interpolation Method Maximum Green Min Yellow_Min 0.10 Red Min

JSC 29181 Plume Model

Service Module Roll/Pitch/Yaw Attitude Control Port Inboard SAW

Boeing Space Environments Updated Plume Model



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Application of Methodology (concluded)

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Conclusions



- This paper presents the methodology employed in updating the ISS plume induced erosion model for better correlation with flight experiment data and for increased accuracy
- The plume induced erosion model originally developed to support the Program significantly over-predicted erosion damage
- Boeing Space Environments succeeded in adjusting the model for better correlation with flight experiment results
- The updated plume model was successful applied in the definition of updated constraints for ISS solar array operations while mitigating against excessive erosion to the arrays. Erosion keep-out zones for ISS solar arrays were reduced by 30% to 60% with the updated model
- The authors hope that future efforts to improve the characterization of plume induced erosion will draw upon the expertise developed for the ISS Program in the development of space environments effects modeling









Hyzod Window Cover Background

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Hyzod Exposures to Thruster Plumes

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Hyzod window cover exposed to thruster firings during Orbiter, Soyuz, and Progress Proximity Operations

Thruster	Docking Port	Number of Approaches	Number of Separations
Orbiter	PMA2	3	3
Soyuz*	FGB	4	4
Soyuz	DC1	4	4
Progress	DC1	1	1
Progress	SM Aft	10	10

* Expected to be the dominant source of plume effects



Example Hyzod Image (1000x Magnification)







Summary of Hyzod Image Analysis Results

- Most common pit size is 2 3 µm for both magnification levels
- Largest pit diameter is approximately 17 µm
- 99.8% of pits are below 8 µm
- Pits with diameters less than 1 µm are visible in 3000x images but not in 1000x images
- Number densities inferred from the 3000x images are up to two times larger than the 1000x images





Note: Histograms show normalized pit size distributions

Note: 3000x magnification images provided only for comparison purposes. Image samples insufficient for quantitative assessment





Hyzod Image Analysis Results

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Magnification	Approximate Pit Number Density Excluding Potential Pits (pits / mm²)	Approximate Pit Number Density Including Potential Pits (pits / mm ²)
1000x	230	240
3000x	410	410

Note: Pit density rounded to the nearest multiple of ten pits / mm²

Magnification	Approximate % Total Area Pitted Excluding Potential Pits	Approximate % Total Area Pitted Including Potential Pits
1000x	0.16 %	0.16 %
3000x	0.24%	0.24%

Note: 3000x magnification images provided only for comparison purposes. Image samples insufficient for quantitative assessment



SPIFEX Background

- Experiment conducted by the U.S. in September 1994 (STS-64)
- Exposed samples of Kapton tape and aluminum foil to:
 - 84 PRCS firings
 - 17 VRCS firings
 - Average pulse of 248.5 ms
 - Average distance of 46 feet
 - Angles off centerline varied from 0° to 90°
- Of interest to compare results for Kapton tape and aluminum foil to evaluate differences in material susceptibility to impact damage



Example SPIFEX Kapton Tape Image: 1500x Magnification (20 µm Scale)

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Example SPIFEX Kapton Tape Image: 100x Magnification (500 µm Scale)





Example SPIFEX Aluminum Foil Image: 1000x Magnification (50 µm Scale)

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Example SPIFEX Aluminum Foil Image: 100x Magnification (500 µm Scale)





SPIFEX Image Analysis Results

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Sample	Approximate Pit Number Density Excluding Potential Pits (pits / mm ²)	Approximate Pit Number Density Including Potential Pits (pits / mm ²)
Kapton Tape	6710	9300
Aluminum Foil	2160	2160

Note: Pit density rounded to the nearest multiple of ten pits / mm²

Sample	Approximate % Total Area Pitted Excluding Potential Pits	Approximate % Total Area Pitted Including Potential Pits
Kapton Tape	6.4%	7.5%
Aluminum Foil	2.8%	2.8%





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Note: Histograms scaled to show similarities and differences in both distributions



Summary of SPIFEX Image Analysis Results

- Most common pit size is 1 2 μm
- Largest Kapton tape pit diameter is approximately 20 µm
 - Largest potential pit diameter is approximately 28 µm
- Largest aluminum foil pit diameter is approximately 20 µm
- Both samples were exposed to the same thruster firings
 - Can compare results to evaluate differences in material susceptibility to impact damage



Comparison with Original Aluminum Foil Image Analysis Results (By KSC)

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Comparison of number of pits identified per mm²

Pit Diameter	Original Count / mm ²	Approximate Current Count / mm ²
≤ 4 µm	449	1600
5 – 10 µm	231	320
11 – 20 µm	60	30

Reference: Soares, C., Barsamian, H., Rauer, S.: "*Thruster Plume Induced Contamination Measurements From the PIC and SPIFEX Flight Experiments*," The Boeing Company, NASA JSC



PIC Background

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- Conducted in 1996 (STS-74)
- Measured initial and permanent plume induced molecular contamination using Quartz Crystal Microbalances (QCM)
- Exposed QCM to:
 - 20 PRCS firings
 - 100 Russian 130 N thruster firings
- Impact features were observed on the camera lens of the Orbiter RMS
 - Consistent with observations from the SPIFEX flight experiment

Soares, C., Barsamian, H., Rauer, S.: "*Thruster Plume Induced Contamination Measurements From the PIC and SPIFEX Flight Experiments*," The Boeing Company, NASA JSC



PIC Statistics

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- 61 pits / mm²
- Pitted area represents 1.8% of the camera lens surface area¹
- PIC statistics given below
 - 6 13 µm bin is the most common PIC pit size
 - $2 3 \ \mu m$ bin is the most common Hyzod pit size
 - $1 2 \ \mu m$ bin is the most common SPIFEX pit size

Pit Diameter	Pit Density / mm ²
2 - 5 µm	21
6 - 13 µm	30
14 - 24 μm	10

PIC Camera Lens Pit Density

1. Soares, C., Barsamian, H., Rauer, S.: "Thruster Plume Induced Contamination Measurements From the PIC and SPIFEX Flight Experiments," The Boeing Company, NASA JSC

