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Venus Science Priorities for Laboratory Measurements and Instrument Definition Workshop April 7–8, 2015 Hampton, Virginia



Coherent Doppler Wind Lidar Technology Development at NASA Langley

**Research Center for Space-based Observations** 

- 1. Hardware Development
- 2. Instrument Demonstration
- 3. Measurement Technique Validation
- 4. NASA GRIP Hurricane Study Field Campaign on DC-8

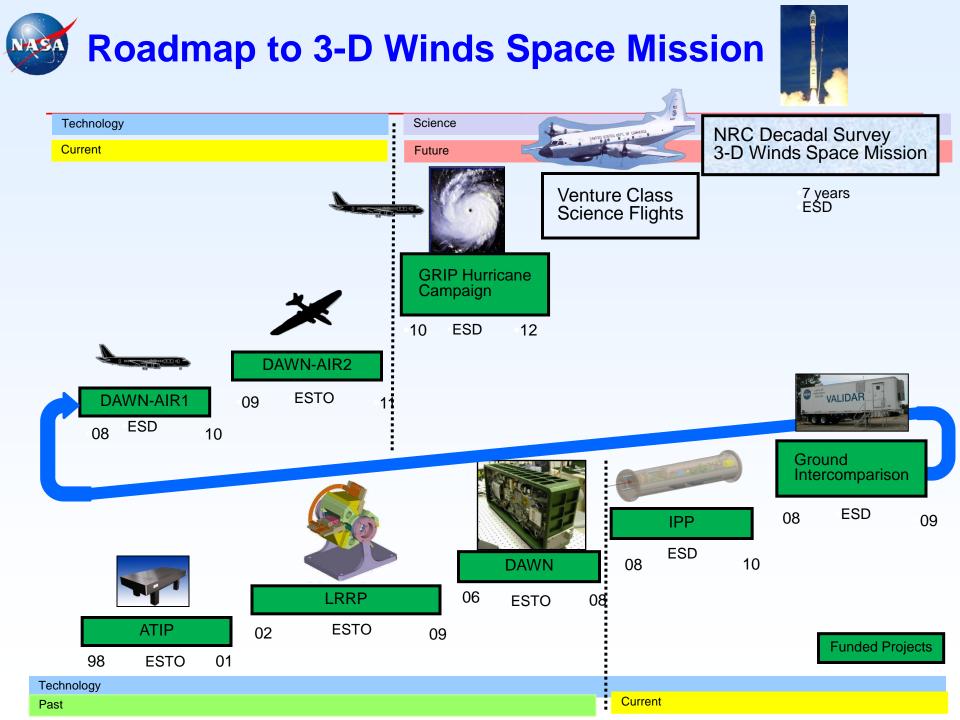
**Coherent Doppler Wind Lidar for Venus** 

- 1. Objectives and Approach
- 2. Venus Proposed Lidar Instrument
- 3. Lidar Performance Simulation for Venus



## **Decadal Survey 3D-Winds Mission**

FARTH SCIENCE		Decadal Survey Mission 2010-2013 CLARRED (NASA portion) SMAP ICESat-II	Launch, Orbit, and Instrume Mission Description Solar and Earth radiation; spect forcing and response of the clin Soil moisture and freeze-thaw f water cycle processes Ice sheet height changes for clin diagnosis	Orbit <sup>a</sup> rally resolved nate system Precessing or weather and LEO, SSO	Instruments Absolute, spectrally resolved interferometer L-band radar L-band radar Laser altimeter	Rough Cost Estimate (FY 06 \$million	
APPLICATIONS FROM SPACE		DESDyni 2013-2016 HyspiRi	Surface and ice shee understanding natu vegetation structure Land surface compo mineral characteriza	Survey's efits	ey's		
N FO App Ass Div NATI	EARTH SCIENCE AND APPLICATIONS FROM SPACE	ASCENDS SWOT	ecosystem health Day/night, all-latitud integrals for climate Ocean, lake, and rive and inland water dy	Extreme Weather Warnings Human Health			✓ ✓
	NATIONAL IMPERATIVES FOR THE NEXT DECADE AND BEYOND	GEO-CAPE	Atmospheric gas col forecasts; ocean colo health and climate e	the second second	e Early Warning		
	Committee on Earth Science and Applications from Space: A Community Assessment and Strategy for the Future	ACE 2016-2020 LIST	Aerosol and cloud p water cycle; ocean o biogeochemistry Land surface topogr and water runoff	mproved We Sea-	Level Rise	ction	√#1
	Space Studies Board Division on Engineering and Physical Sciences	PATH GRACE-II	High-frequency, all-v humidity soundings and sea-surface tem High-temporal-resol tracking large-scale		e Prediction		
	NATIONAL RESEARCH COUNCIL OF THE NATIONAL ACADEMIES THE NATIONAL ACADEMIES PRESS	SCLP GACM	Snow accumulation Ozone and related g air quality and strate prediction		em Services		
3D-Winds (Demo)	2007 Washington, D.C. www.nap.edu Tropospheric winds for weather forec and pollution transport	3D-Winds (Demo)	Tropospheric winds and pollution transp	Air Doppler I	Quality idar	Yalisi	650



# NASA

## **Technology Maturation Example**



Modeling

A fully conductively cooled 2-micron solid-state pulsed laser has been demonstrated to enable space-borne 3-D Wind

measurements





**System Integration** 



Space Qualifiable Design



Testing and Model Verification

## Mobile Ground based High Energy Wind Lidar Transceiver – LRRP/DAWN Funded

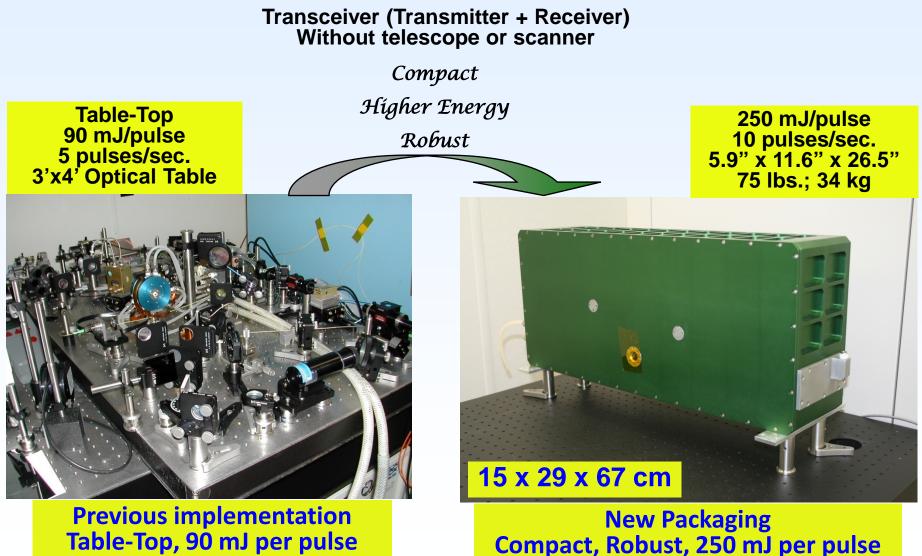
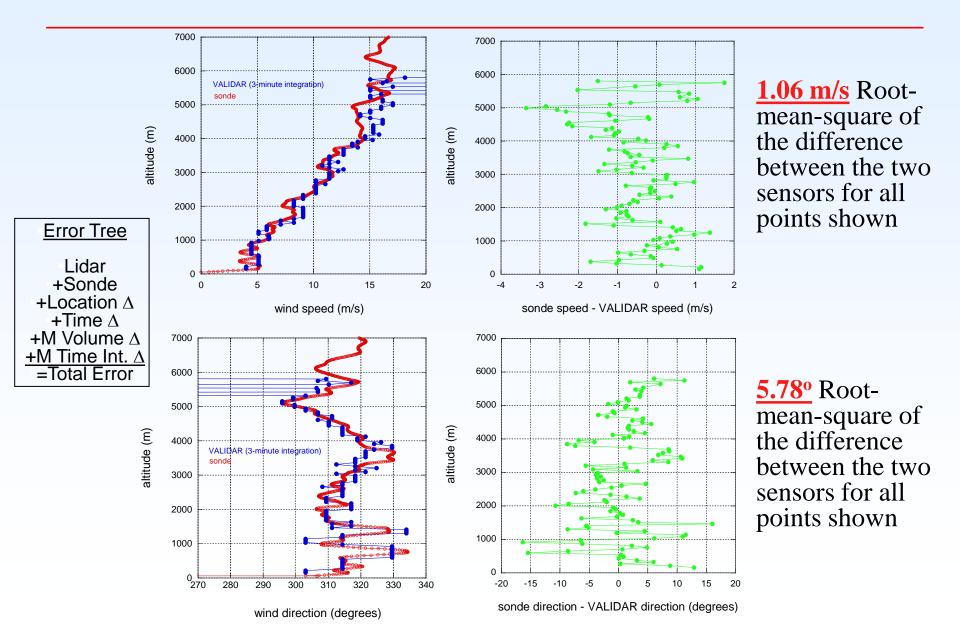


Table-Top, 90 mJ per pulse

## Comparison of Coherent Lidar and Sonde





## **GRIP DC-8 Payload**

#### Dropsondes

(Vertical Profiles of Temp, Press, Humidity and Winds)

#### CAPS, CVI, PIP

(Cloud Particle Size distributions, Precip Rate, Rain & Ice water content)

#### DAWN

Doppler Aerosol Wind Lidar (Vertical Profiles of Vectored Horizontal Winds)

#### LASE

Lidar Atmospheric Sensing Experiment (H2Ov, Aerosol profiles and Cloud distributions)

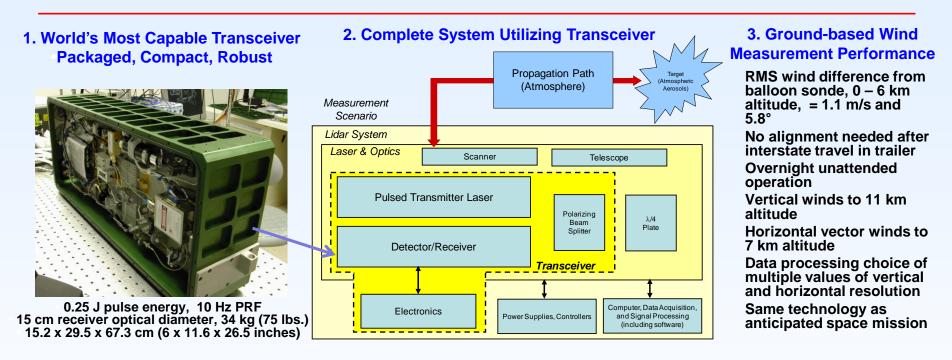
#### APR-2

Airborne Precipitation Radar Dual Frequency (Vertical Structure Rain Reflectivity and Cross Winds)

#### MMS

Meteorological Measurement System (Insitu Press, Temp, 3D Winds and Turbulence)

### Genesis and Rapid Intensification Process (GRIP) Campaign Coherent Pulsed Doppler Wind Profiling Lidar System



#### 4. Enclosure for All Lidar Optics Robust Aircraft Design

5. Optics in DC-8

6. Lidar System in DC-8

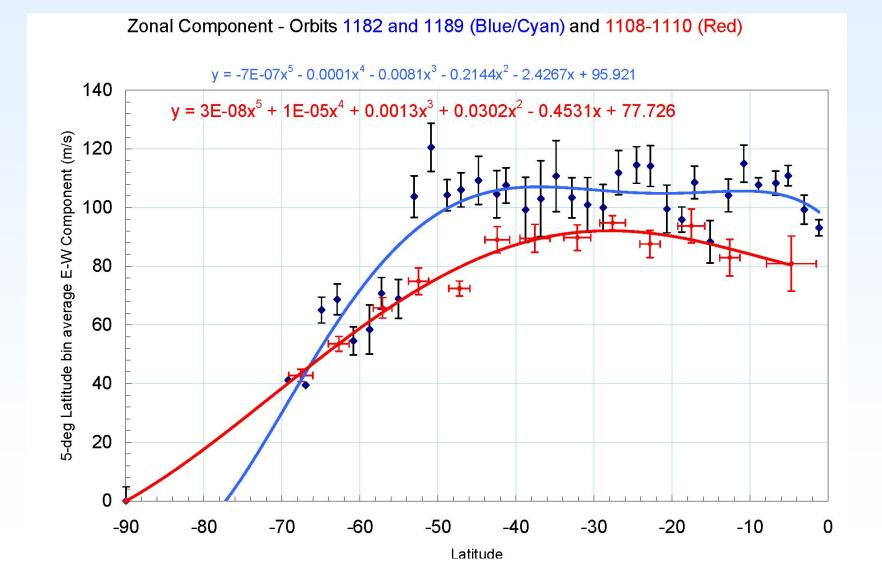




## **Need for Doppler Wind Measurements on Venus**

- Global super-rotation is a puzzle
- Keys to understanding the circulation are spatially distributed measurements at the same vertical level to enable estimates of eddy and mean contributions to transport of angular momentum meridionally
- Cloud motions provide only day side winds with high spatial resolution
- Night side cloud tracking measurements refer to a different level
- Hence the true zonal average circulation is not well known







- Obtain direct measurement of Venus winds within the upper cloud layer and the overlying aerosol (haze) layer
- Obtain directly measured heights of cloud tops and their optical depths
- Obtain aerosol concentration and distribution within the upper haze layer

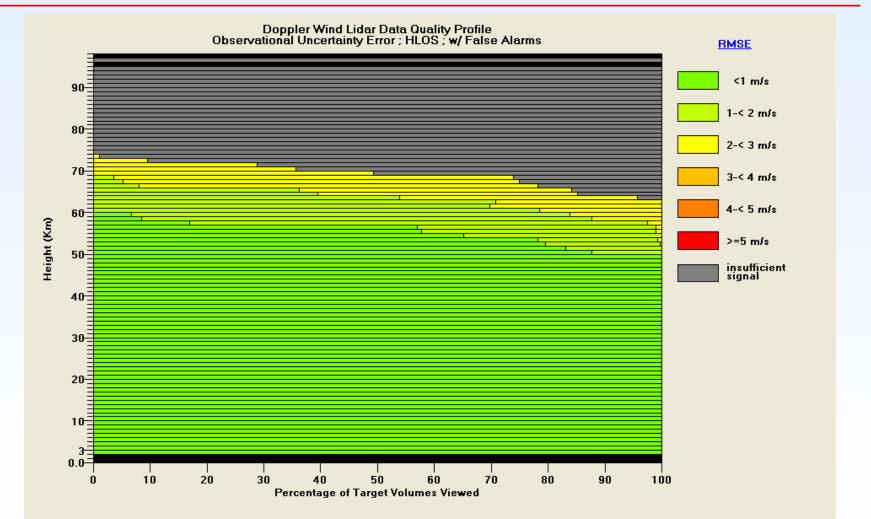


- The thick uniform cloud cover (1 micron radius and sub-micron sized haze particles in polar latitudes) should enable good Doppler measurements and provide height resolved results of atmospheric motions
- Use an orbiting Doppler lidar to obtain u, v, w components of the winds (clouds and aerosols permitting).
- Design instrument for 1km vertical resolution and 200 km horizontal spacing of profiles; enable onorbit changes to integration and sampling strategies.
- Operate full or partial orbits depending upon platform power availability



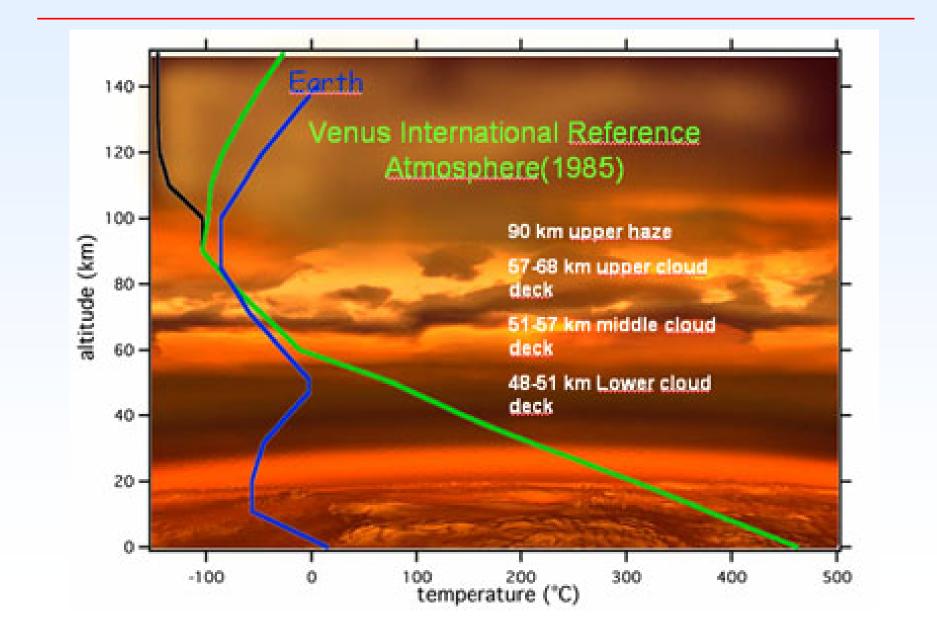
- Build upon the NASA funded Mars Lidar Simulation Model (MLSM) developed by Simpson Weather Associates.
- Modify as needed to represent Venus conditions; preference is to use a community Venus weather model.
- Assume similarities in cloud reflectance at 2  $\mu m$  and spherical particles in the haze layer.

# Doppler Wind Lidar Profile Simulation on Mars (Example)



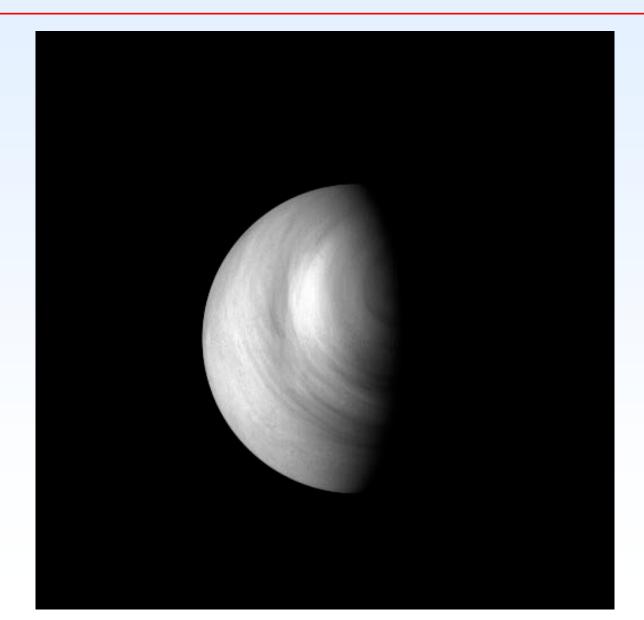
RUN: MCD Ls=0 deg 2001Z (2 hr sim.) Global ALTITUDE: 400 Km ORBIT INC: 98 deg TARGET ATM.: median Backscatter, R= 0.30, constant Q ELEVATED DUST: user defined R = 0.00 um DETECTION: DWL SCAN: 4 pt. Step/Stare WAVELENGTH: 2.053472 um NADIR: 45. deg ENERGY: 0.25 J APERTURE: 0.5 m PLT PWR: 63 W PRF: 10. Hz ACCUM. TIME: 25 s BETA THR.: 50.%

## Venus International Reference Atmosphere



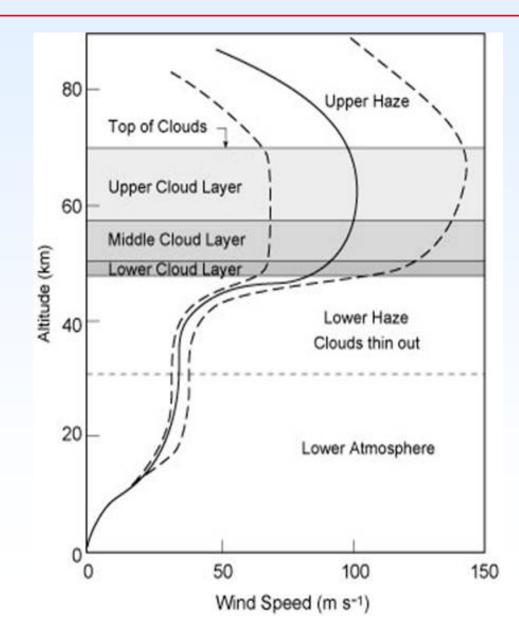


## **Venus Cloud Layers**





## **Target Region for VenWinds**





### **VenWinds Instrument**

Wavelength (µm)	2.0	
Energy per pulse (J)	0.001	
Pulse length (nsec)	180	
PRF (Hz)	500	
Optical output (watts)	0.5	
Wallplug power (watts)	10	Includes cooling but not data collection/transmission
Telescope diameter (m)	0.1	
Scanner (step stare 4 azimuth angles plus 1 nadir)		Other beam director options can be considered
Nadir scanning angle (degrees)	20	Can be varied
Dwell time (seconds)	10	This would result in 5000 samples being integrated for one LOS wind component.
Sample volume diameter (m)	0.1	
Sample volume length; range gate (m)	1100	The number of pulse lengths in this volume
Beta50 (msr-1)	6.4×10 <sup>-7</sup>	Backscatter for Venus clouds ~ 1.0×10 <sup>-5</sup>
Weight (kg)	15 -50	Depends on effort to reduce weight of current technology
Dimensions (I x w x h)	TBD	



### **VenWinds Data**

Height resolution(m)	1000	
(average wind in layer)		
Horizontal spacing between wind	200	This can be changed to be as small as 25km with
profiles (u, v, w) (km)		no sacrifice on accuracy, just sensitivity.
LOS velocity precision (m/s)	< 1	
Maximum horizontal wind (m/s)	200	
Aerosol profiles (m)	1000	Each LOS observation yields a wind speed,
(layer average during scan)		turbulence estimate (TKE) and signal intensity
Number of profiles per orbit	100 - 1000	Varies with integration time
Vertical coverage	TBD	Minimum of cloud top speeds and heights
Cloud top range resolution (m)	~ 100 meters	



- NASA Langley Research Center is the world leader in developing pulsed 2-micron coherent Doppler/DIAL/backscatter lidar for space remote sensing of Earth's atmosphere and have successfully developed and matured the DWL technologies and techniques
- These technologies can be customized and matured for Venus through leveraging the knowledge and knowhow acquired by the LaRC team in last two decades
- Simpson Weather Associates and NASA LaRC have jointly developed a Mars Lidar Simulation Model to test DWL concepts for a Mars mission and it can be used for Venus wind simulation
- Based upon general available information regarding the atmosphere of Venus, a small Doppler Wind Lidar (DWL) could provide wind, cloud and aerosol information from an orbit of several 100 kms above the surface of Venus.
- Issues of power, weight, volume need to be addressed to identify the tall poles in this proposed instrument.