Erosion Data from the MISSE 8 Polymers Experiment after 2 Years of Space Exposure on the International Space Station

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Abstract

The Polymers Experiment was exposed to the low Earth orbit (LEO) space environment for 2.14 and 2.0 years as part of the Materials International Space Station Experiment 8 (MISSE 8) and the Optical Reflector Materials Experiment-III (ORMatE-III), respectively. The experiment contained 42 samples, which were flown in either ram, wake, or zenith orientations. The primary objective was to determine the effect of solar exposure on the atomic oxygen erosion yield (Ey) of fluoropolymers. This paper provides an overview of the experiment with details on the polymers flown, the characterization techniques used, the atomic oxygen fluence for each exposure orientation, and the LEO Ey results. The Ey values for the fluoropolymers range from $1.45 \times 10^{-25} \text{ cm}^3/\text{atom for white Tedlar}^{\otimes}$ (polyvinyl fluoride with white titanium dioxide pigment) flown in the ram orientation to $6.32 \times 10^{-24} \text{ cm}^3/\text{atom for aluminized-Teflon}$ fluorinated ethylene propylene (Al-FEP) flown in the zenith orientation. Erosion yield data for FEP flown in ram, wake and zenith orientations are compared, and the Ey was found to be highly dependent on orientation, hence environmental exposure. Teflon FEP had an order of magnitude higher Ey when flown in the zenith direction (6.32 $\times 10^{-24}$ cm³/atom) as compared to the ram direction (2.37 x 10^{-25} cm³/atom). The Ey of FEP was found to increase with a direct correlation to the solar exposure/AO fluence ratio showing the effect of solar radiation and/or heating due to solar exposure on FEP erosion. In addition, back-surface carbon painted FEP (C-FEP) flown in the zenith orientation had a significantly higher Ey than clear FEP or Al-FEP further indicating that heating has a significant impact on the erosion of FEP, particularly in the zenith orientation.



Erosion Data from the MISSE 8 Polymers Experiment after 2 Years of Space Exposure on the International Space Station

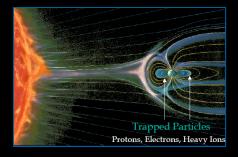


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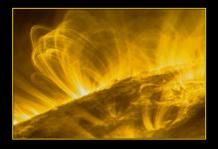
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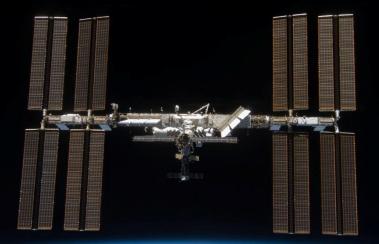


The Space Environment

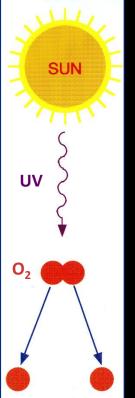


Materials on the exterior of spacecraft are exposed to many harmful environmental threats In low Earth orbit (LEO) environmental threats include:

- Solar radiation (ultraviolet (UV), x-rays)
- Charged particle radiation (electrons, protons)
- Cosmic rays (energetic nuclei)
- Temperature extremes & thermal cycling
- Micrometeoroids & orbital debris (space particles)
- Atomic oxygen (reactive oxygen atoms)



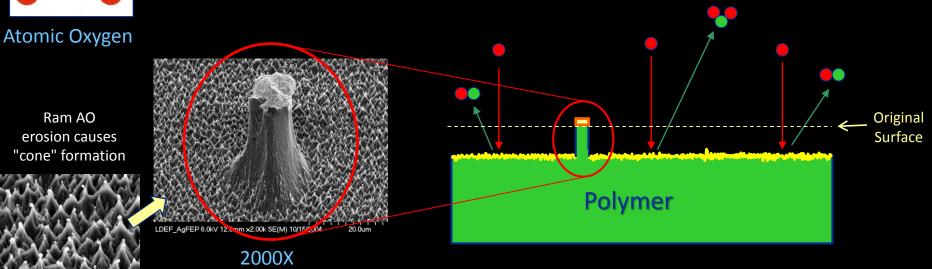




Atomic Oxygen (AO)



- AO is the predominant species in LEO (\approx 200-650 km)
- It is formed by photodissociation of molecular oxygen (O₂) by short wavelength energetic UV radiation
- At ram impact velocities (17,000 mph) the average impact energy is 4.5 eV
- AO oxidizes certain materials (such as polymers) with resulting gas formation so *the material erodes away...*
 - \Rightarrow AO is a serious threat to spacecraft survivability





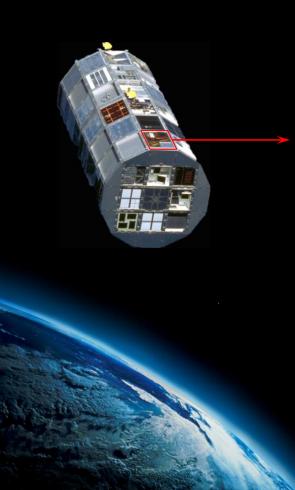
Atomic Oxygen Erosion of a Kapton Insulation Blanket on the Long Duration Exposure Facility (LDEF)

Pre-flight



Tray F-9

Post-flight







After 5.8 years in space

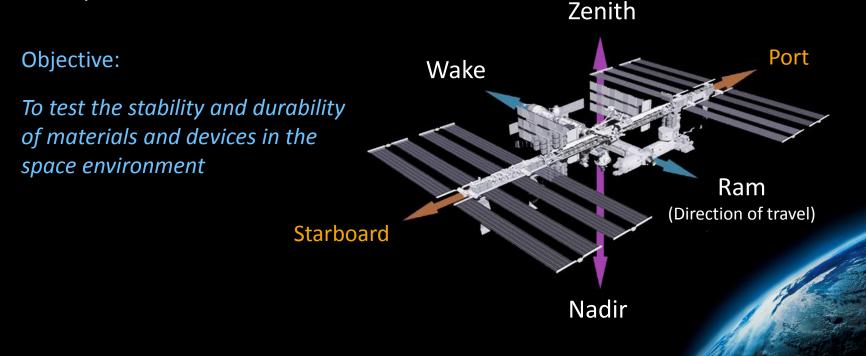


Materials International Space Station Experiment (MISSE)



MISSE is a series of materials flight experiments consisting of trays called Passive Experiment Containers (PECs), that were exposed to the space environment on the exterior of the International Space Station (ISS).

The PECs were positioned in either a *ram/wake orientation* or a *zenith/nadir orientation*.



Flight Orientations & Environmental Exposures

Ram:

- Facing the direction of travel (i.e. forward pointing or leading edge)
- <u>Highest AO</u> & moderate solar exposure

Wake:

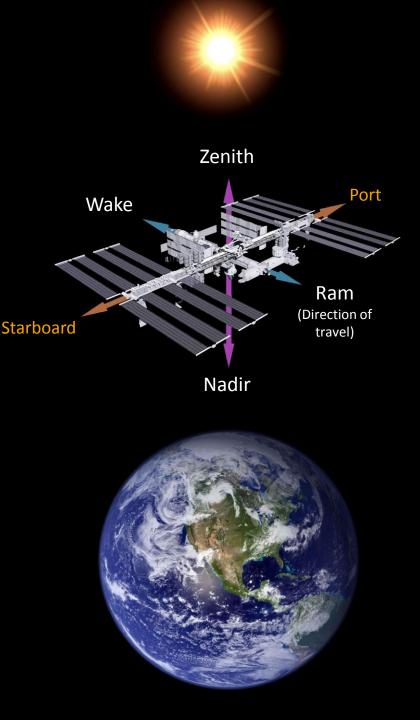
- Facing away from the direction of travel (i.e. aft pointing or trailing edge)
- <u>Very little AO</u> & moderate solar exposure

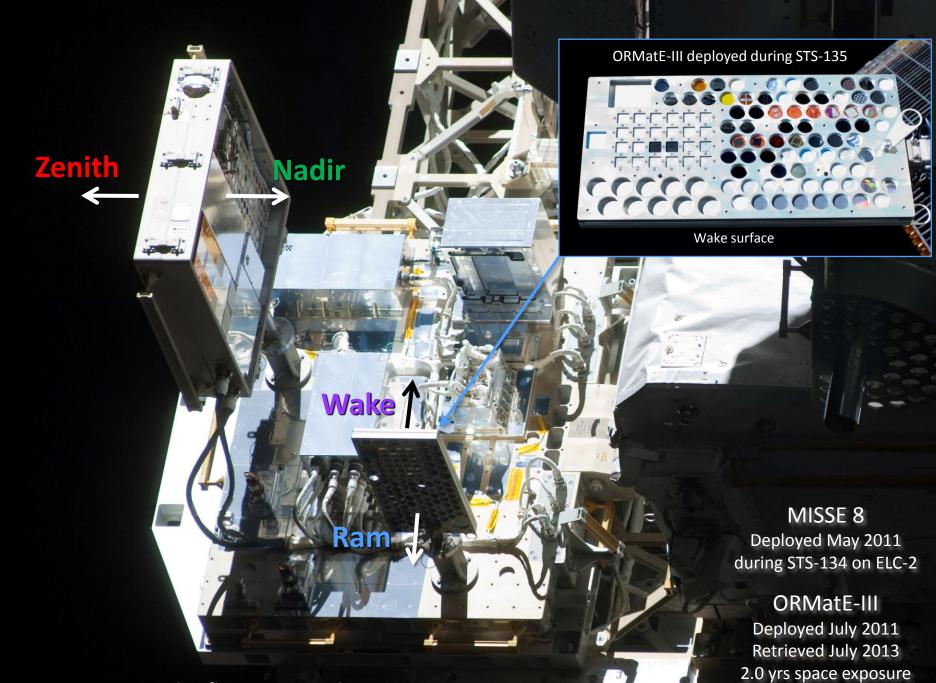
Zenith:

- Direction facing away from Earth (i.e. directly above)
- Grazing AO & <u>highest solar exposure</u>

Nadir:

- Direction facing towards Earth (i.e. straight down)
- Grazing AO & <u>lowest solar exposure</u>





ORMatE-III: Optical Reflector Materials Experiment III

MISSE 8 Deployed: May 2011 Retrieved: July 2013

2.14 years space exposure

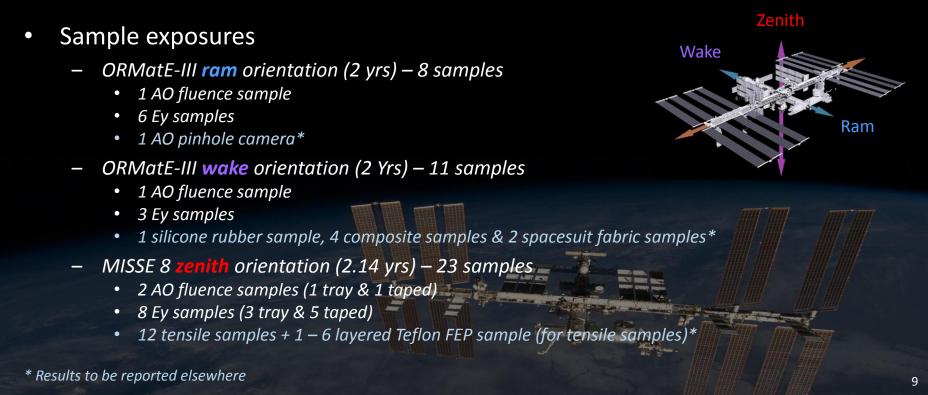
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MISSE 8 Polymers Experiment



- Passive experiment w/ 42 samples exposed to space for 2 or 2.14 years
 - Primary objective:
 - Determine the effect of solar exposure on the atomic oxygen erosion yield (Ey, cm³/atom) of fluoropolymers
 - Additional samples were included for environmental durability assessment, and will be reported elsewhere
 - Tensile samples, composites, spacesuit fabrics, pinhole camera, docking seal sample, etc.*





MISSE 8 Zenith Samples

Zenith Electronics and Polymer Exposure Experiment (ZEPEE)

Polymer Experiment Samples: 4 tray & 11 taped samples

MISSE 8 Zenith Tray (pre-flight)

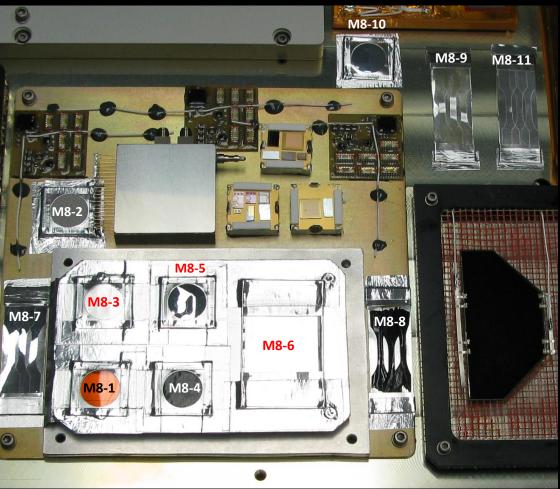


Photo credit: Naval Research Laboratory

MISSE 8 Zenith Tray (pre-flight)



M8- M8- M8-Z2B Z3B Z4B Photo credit: Naval Research Laboratory

GRC ID	MISSE 8 Sample							
M8-Z1B	Kapton H							
M8-Z2B	FEP							
M8-Z3B	AI-FEP*							
M8-Z4B	HST Al-FEP* (SM3A R1 - 9.7 yrs)							
11 Taped Samples								
M8-1	Kapton H							
M8-2	Pyrolytic Graphite (PG)							
M8-3	FEP							
M8-4	Carbon back-surface painted FEP (C-FEP)*							
M8-5	AI-FEP*							
M8-6	6 layers FEP (3 tensile samples x 6)							
M8-7	AI-FEP tensile samples* (3)							
M8-8	C-FEP tensile samples* (3)							
M8-9	CP1 tensile samples (3)							
M8-10	Ag-FEP*							
M8-11	FEP tensile samples (3)							

* FEP is space-facing



ORMatE-III Ram Samples



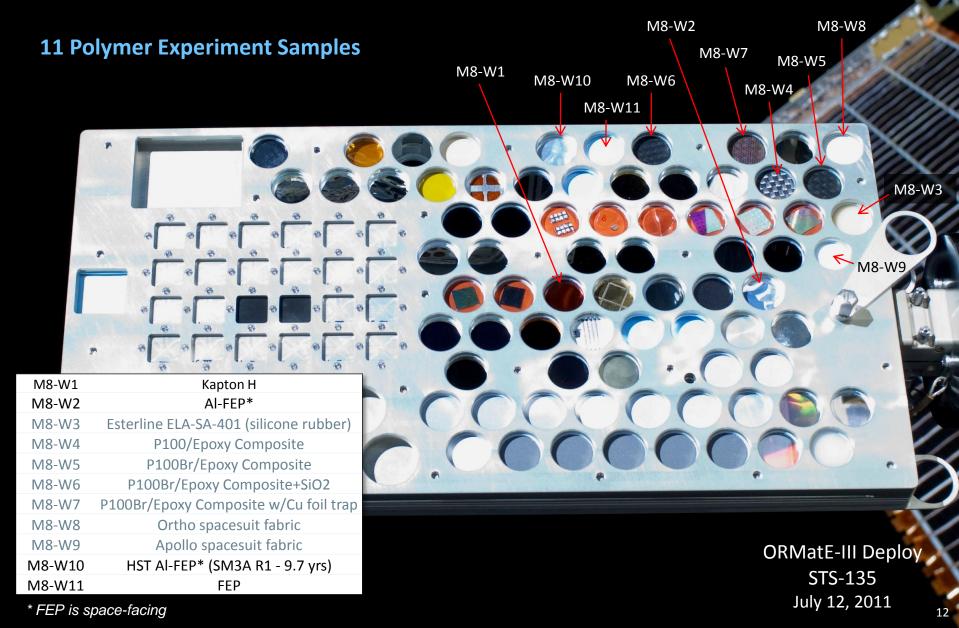
8 Polymer Experiment Samples

			M8-R7				
		M8-R8					
	M8-R5				-		
					1		
				00	0		
		V					
			0			-	
M8-R10						21	
						-	0 0 0 0
M8-R6	0						
		The Co	E.C.				0 0
M8-R9 —		2310					
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			()		0		0 0 0 0 0
		N (.				0	
						100	
M8-R2						M8-R1	Kapton H
							Aluminized-fluorinated ethylene propylene
To U						M8-R2	(AI-FEP)*
- Pris	He I I I I	TIE.				M8-R5	White Polyvinyl fluoride (PVF),
1	A THE						White Tedlar
	1 1 mar			0	(M8-R6	Hubble Space Telescope (HST)
		18-R1	66	0			Al-FEP* (SM3A R1 - 9.7 yrs)
	14					M8-R7	Polytetrafluoroethylene (PTFE)
F 22	20			(()) (M8-R8	AO Pinhole Camera
Eva 22					00	M8-R9 M8-R10	Fluorinated ethylene propylene (FEP) DC 93-500 silicone/fused silica
July 9,	2013		1	ACAC			space-facing ¹¹
						- FEF IS S	space-racing 11



ORMatE-III Wake Samples





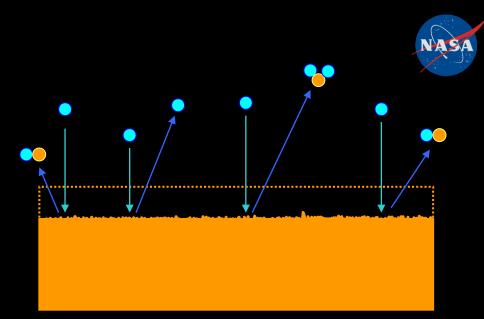


Atomic Oxygen

Erosion Yield (E_y)

(Also called Reaction Efficiency or Recession Rate)

Ey is the volume loss per incident oxygen atom (cm³/atom)



Ey based on Mass Loss Measurements

Erosion Yield (E_y) of Sample

$$E_{y} = \frac{\Delta M_{s}}{A_{s}\rho_{s}F_{k}}$$

Atomic Oxygen Fluence

$$F_k = \frac{\Delta M_k}{A_k \rho_k E_k}$$

where: $\Delta M_s =$ Mass loss of polymer sample (g) $A_s =$ Area of polymer sample (cm²) $\rho_s =$ Density of sample (g/cm³) $F_k =$ AO fluence measured by Kapton H witness samples (atom/cm²)

where:
$$\Delta M_k =$$
 Mass loss of Kapton H witness (g)

 $A_k =$ Area of Kapton H witness (cm²)

- $\rho_k = Density of Kapton H sample$ (1.427 g/cm³)
- E_k = Erosion yield of Kapton H (3.0 x 10⁻²⁴ cm³/atom)



Experimental Procedures

Mass Measurements:

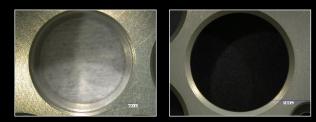
Hygroscopic materials can have large weight variations due to moisture absorption (i.e. Kapton can absorb up to 2% of its weight in moisture)

Samples were vacuum dehydrated (\approx 60-100 mtorr) for a min. of 72 hours prior to measuring mass before, & after, flight using a Sartorius ME 5 Microbalance (0.000001 g sensitivity)

Exposed Surface Area (SA) Measurements:

Tray Samples:

4 diameter measurements were taken of each tray opening with a Bore Gauge (± 0.001 mm) and averaged



Density Measurements: Density Gradient Columns*

Solvents for low density

Solvents for high density

*Ref: de Groh et al. HPP 20: 388-409, 2008

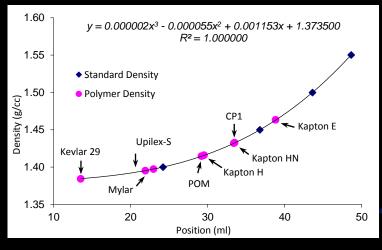
- Cesium Chloride (CsCl, $\rho = 2 \text{ g/cm}^3$) & Water (H₂O, $\rho = 1 \text{ g/cm}^3$)
- Carbon Tetrachloride (CCl₄, ρ = 1.594 g/cm³) & Bromoform (CHBr₃, ρ = 2.899 g/cm³

Taped & irregularly shaped samples:

AutoCAD computer design software was used to trace the border of the sample on a photograph & the SA was computed with the AutoCAD software



AutoCAD photo-trace











MISSE 8 Polymers Experiment Flight Data

Atomic Oxygen Fluences Solar Exposures

Erosion Yield Values

&

MISSE 8 Deployed May 2011 on ELC-2



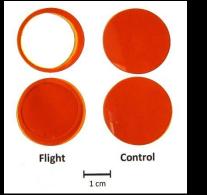
MISSE 8 AO Fluence



GRC AO Fluence Data

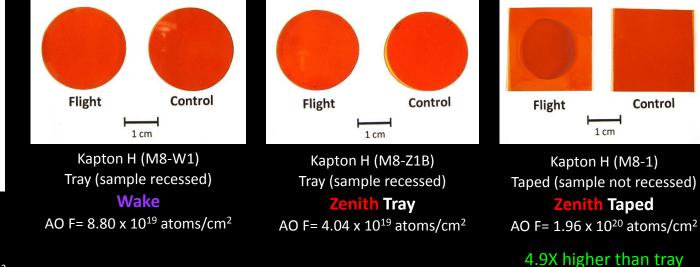
Orientation	Experiment	Sample ID	Material	Thickness	Holder	Technique	AO Fluence (atoms/cm ²)
Ram	Polymers Experiment	M8-R1	Kapton H*	5 mil	Beveled Tray	Mass loss	4.62 E21
Wake	Polymers Experiment	M8-W1	Kapton H*	5 mil	Beveled Tray	Mass loss	8.80 E19
Zenith	Polymers Experiment	M8-Z1B	Kapton H*	5 mil	Beveled Tray	Mass loss	4.04 E19
Zenith	Polymers Experiment	M8-1	Kapton H*	5 mil	Thin Al foil (taped)	Mass loss	1.96 E20

* Kapton H E_v based on prior LEO flight experiments



Kapton H (M8-R1) Tray (sample recessed) Ram



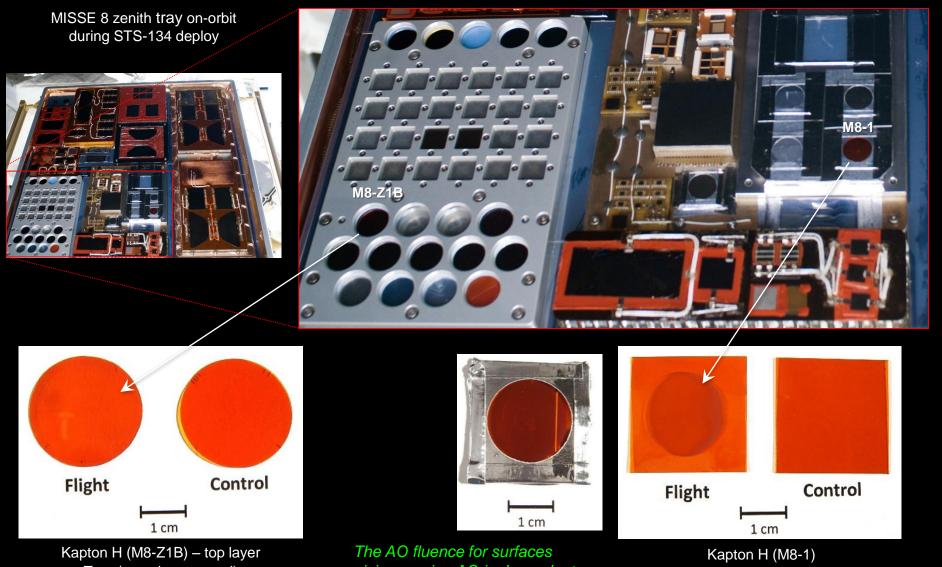




MISSE 8 Zenith Samples

Kapton H AO Fluence Witness Samples





Tray (sample recessed) Very little texture visible The AO fluence for surfaces receiving grazing AO is dependent on the sample holder geometry

Kapton H (M8-1) Taped Al Holder (sample not recessed) Erosion texture visible





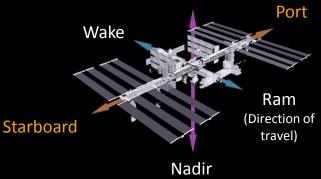


Ram & Wake Solar Exposure Approximations

Flight Orientation	MISSE 7 Exposure (Yrs)	MISSE 7 Equivalent Sun Hours (ESH)	MISSE 7 ESH Relative to Zenith	MISSE 8 ESH Relative to MISSE 7 ESH	MISSE 8 (ESH)
Zenith	2.14	4,300	1	6,100*	$6,100 \pm 1,000^{*}$
Ram	2.00	2,400	0.56	3,182**	3,200 ⁺
Wake	2.00	2,000	0.47	2,652**	2,700+
Nadir	2.14	<<2,000	-	-	$800\pm300^{+}$

* Reference: Phil Jenkins, Naval Research Laboratory, Calculated ESH
** Ram & wake also multiplied by 2.00/2.14 (0.93)

Approximations
 (Nadir approximation provided by M. Finckenor, NASA MSFC)



Zenith





MISSE 8 Ram Samples



GRC ID	Material	Thickness (mil)	Number of layers	Mass Loss (g)	Bore Gauge Area (cm ²)	Density (g/cm³)	AO Fluence (atoms/cm ²)	MISSE 8 Ey (cm ³ /atom)
M8-R1	Kapton H	5	2	0.079104	3.9953	1.4273	4.62E+21	3.00E-24*
M8-R2	AI-FEP**	5	1	0.009489	3.9969	2.1443	4.62E+21	2.39E-25
M8-R5	White Tedlar	1	7	0.004358	3.9945	1.6241	4.62E+21	1.45E-25
M8-R6	Hubble Space Telescope (HST) 9.7 yrs, Al-FEP**	5	1	0.009870	3.9925	2.137	4.62E+21	2.50E-25
M8-R7	PTFE	2	1	0.007712	3.9956	2.1503	4.62E+21	1.94E-25
M8-R9	FEP	2	1	0.009399	3.9924	2.1443	4.62E+21	2.37E-25
M8-R10	DC 93-500 silicone on fused silica	10	1	0.000076	3.9940	1.08	4.62E+21	3.81E-27

* Kapton H Ey based on prior LEO flight experiments

** FEP is space facing



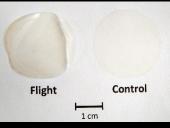
AI-FEP (M8-R2)



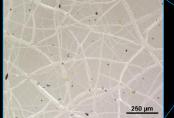
White Tedlar (M8-R5)



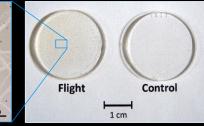
HST AI-FEP (M8-R6)



PTFE (M8-R7)







AO induced surface cracks

19



MISSE 8 Wake Samples



GRC ID	Material	Thickness (mil)	Number of layers	Mass Loss (g)	Bore Gauge Area (cm²)	Density (g/cm³)	AO Fluence (atoms/cm ²)	MISSE 8 Ey (cm ³ /atom)
M8-W1	Kapton H	5	1	0.001523	4.0415	1.4273	8.80E+19	3.00E-24*
M8-W2	AI-FEP**	5	1	0.000839	4.0440	2.1443	8.80E+19	1.10E-24
M8-W10	HST Al-FEP** (9.7 yrs)	5	1	0.000790	4.0437	2.137	8.80E+19	1.04E-24
M8-W11	FEP	2	1	0.000827	4.0410	2.1443	8.80E+19	1.08E-24

* Kapton H Ey based on prior LEO flight experiments

** FEP is space facing





MISSE 8 Zenith Tray Samples



GRC ID	Material	Thickness (mil)	Number of layers	Mass Loss (g)	Bore Gauge Area (cm²)	Density (g/cm³)	AO Fluence (atoms/cm ²)	MISSE 8 Ey (cm ³ /atom)
M8-Z1B	Kapton H	5	1	0.000670	3.8771	1.4273	4.04E+19	3.00E-24*
M8-Z2B	FEP	2	1	0.001916	3.8772	2.1443	4.04E+19	5.71E-24
M8-Z3B	AI-FEP**	5	1	0.002120	3.8742	2.1443	4.04E+19	6.32E-24
M8-Z4B	HST Al-FEP** (9.7 yrs)	5	1	0.002113	3.8774	2.137	4.04E+19	6.32E-24

* Kapton H Ey based on prior LEO flight experiments

** FEP is space facing





MISSE 8 Zenith Taped Samples



GRC ID	Material	Thickness (mil)	Number of layers	Mass Loss (g)	Bore Gauge Area (cm²)	Density (g/cm³)	AO Fluence (atoms/cm ²)	MISSE 8 Ey (cm ³ /atom)
M8-1	Kapton H	5	1	0.002330	2.7819	1.4273	1.96E+20	3.00E-24*
M8-2	Pyrolytic Graphite	80	1	0.000513	2.9277	2.22	1.96E+20	4.04E-25
M8-3	FEP	2	1	0.001573	2.7310	2.1443	1.96E+20	1.37E-24
M8-4	C-FEP**	2	1	0.004134	2.7264	2.1443	1.96E+20	3.61E-24
M8-5	AI-FEP**	2	1	0.002172	2.7213	2.1443	1.96E+20	1.90E-24
M8-10	Ag-FEP**	5	1	0.001565	2.7606	2.1443	1.96E+20	1.35E-24

* Kapton H Ey based on prior LEO flight experiments

** FEP is space facing



Flight Control

Ag-FEP (M8-10)



The Ey of C-FEP was 1.9X greater than AI-FEP:

 \Rightarrow Heating has a significant impact on the erosion of FEP



MISSE 8 Teflon FEP



Ram vs Wake vs Zenith Ey

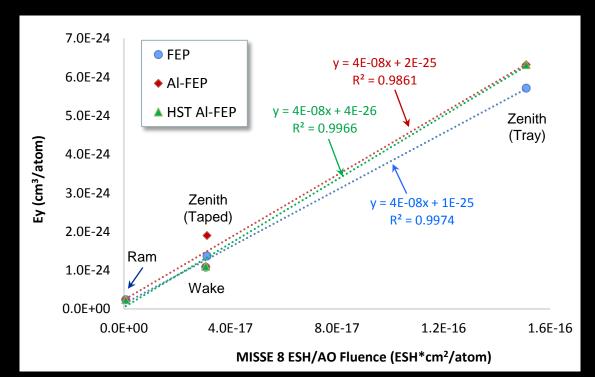
Orientation	MISSE 8 ESH	MISSE 8 AO Fluence (atoms/cm ²)	ESH/AO Fluence (ESH*cm ² /atom)	FEP Ey (cm ³ /atom)	Al-FEP Ey (cm ³ /atom)	HST Al-FEP Ey (cm ³ /atom)
Ram	3,200	4.62E+21	6.92E-19	2.37E-25	2.39E-25	2.39E-25
Wake	2,700	8.80E+19	3.07E-17	1.08E-24	1.10E-24	1.10E-24
Zenith (Taped)	6,100	1.96E+20	3.12E-17	1.37E-24	1.90E-24	-
Zenith (Tray)	6,100	4.04E+19	1.51E-16	5.71E-24	6.32E-24	6.32E-24

Al-FEP consistently slightly higher than FEP



Excellent correlation of Ey to ESH/AO F ratio:

⇒ Shows the effect of solar radiation and/or heating due to solar exposure on erosion of FEP





MISSE 8 Teflon FEP Ram vs Wake vs Zenith Ey



AI-FEP HST AI-FEP MISSE 8 **FEP C-FEP ESH/AO Fluence** MISSE 8 Orientation **AO Fluence** Ey Ey Ey Ey (ESH*cm²/atom) ESH (cm³/atom) (cm³/atom) (cm³/atom) (atoms/cm²) (cm3/atom) 4.62E+21 2.37E-25 2.39E-25 Ram 3,200 6.92E-19 2.39E-25 Wake 2,700 8.80E+19 3.07E-17 1.08E-24 1.10E-24 1.10E-24 Zenith (Taped) 1.96E+20 3.12E-17 1.37E-24 1.90E-24 6.100 3.61E-24 4.04E+19 Zenith (Tray) 6,100 1.51E-16 5.71E-24 6.32E-24 6.32E-24

Al-FEP consistently slightly higher than FEP

Ey vs. ESH/AO F

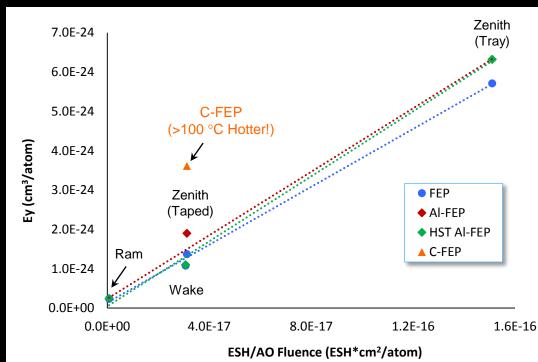
Temperature Estimates:

Assumptions:

- Same sample configuration and orientation
- Solar absorptance:
 - 0.179 for 2 mil Al- FEP
 - 0.96 for 2 mil C-FEP
- Thermal emittance:
 - 0.66 for both
- Constraint:
 - Radiated heat is ejected only in the space direction (due to Al on the back or Al in the sample holder)

Maximum temperature:

- *AI-FEP:* **2°C**
- C-FEP: 170 °C



Heating has a major impact on the Ey of FEP in the zenith orientation



Summary & Conclusions



- The MISSE 8 Polymers Experiment was successfully flown & exposed to the LEO space environment for 2-2.14 years
 - Samples flown in ram, wake & zenith orientations \Rightarrow received different exposures
 - Kapton H samples were used to determine the AO fluence in each orientation
 - > The AO fluence on grazing surfaces is dependent on the sample holder geometry
- LEO Ey of 17 samples were determined based on:
 - Mass loss, density, surface area & AO fluence
- Ey of FEP was dependent on orientation, hence environmental exposure:
 - The zenith & wake exposures (high ESH/AO fluence) provided an <u>order of magnitude</u> greater Ey than ram exposure
 - FEP Ey increased with a direct correlation to the **ESH/AO fluence** ratio showing the effect of solar radiation and/or heating due to solar exposure on FEP erosion
- C-FEP had a significantly higher Ey than FEP or Al-FEP in zenith orientation
 - Heating has a significant impact on the erosion of FEP
- MISSE flight data is available in MISSE MAPTIS Database *http://maptis.nasa.gov/*
 - MAPTIS: Materials and Processing Technical Information System



Acknowledgements



We would like to express our sincere appreciation to Phil Jenkins (NRL, MISSE 8 Mission PI) for providing the opportunity to fly this experiment as part of the MISSE 8 mission

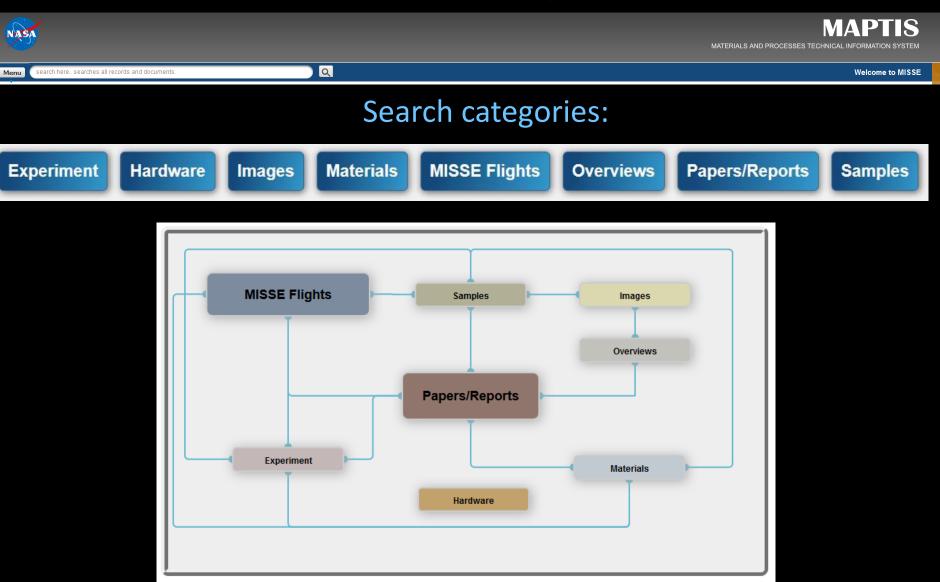
Glenn's MISSE research has been supported by various projects over the past 15 years including the ISS Research Program, the MISSE-X Project and currently, the MISSE Informatics Project



Extra Slides



Materials and Processing Technical Information System (MAPTIS) MISSE Database http://maptis.nasa.gov/



NASA