The use of D-criteria to assess meteor shower significance

Introduction

- The anisotropy of the sporadic meteor background complicates meteor shower extraction.
- Using static orbital similarity criteria to identify shower members can produce too many false positives near sporadic sources.
- Concept: We use shower "analogs" to characterize the density of meteor orbits in a region of parameter space when the shower is *not* active.

Orbital Similarity Criteria

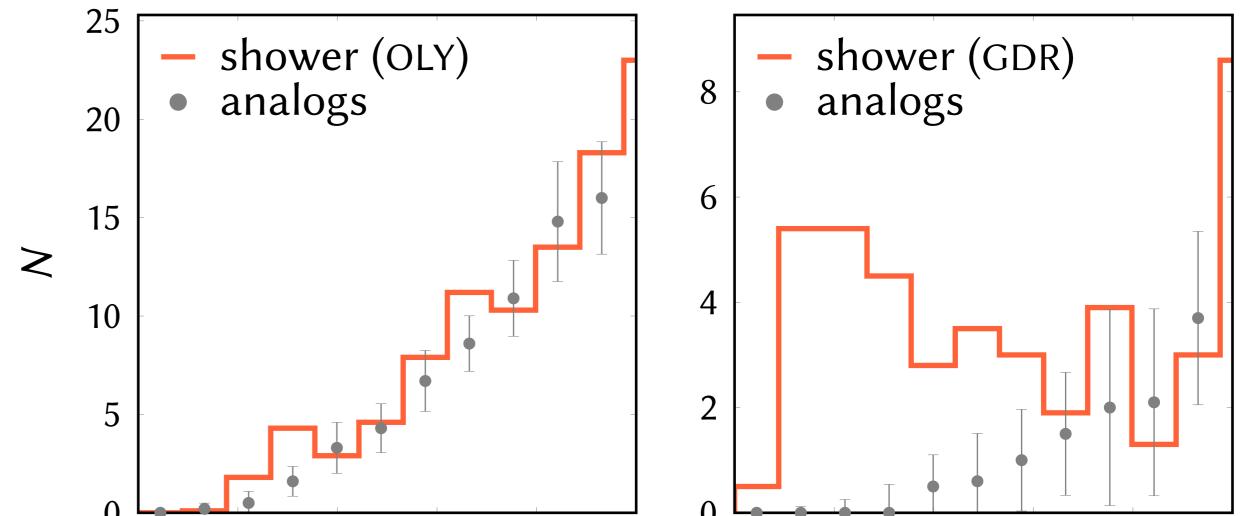
- D-parameters quantify orbital similarity.
- We obtained the best results using D_N [1]:

$$\Delta \phi_a = 2 \sin \frac{1}{2} (\phi_2 - \phi_1)$$

$$\Delta \phi_1 = 2 \sin \frac{1}{2} (\pi + \phi_2 - \phi_1)$$

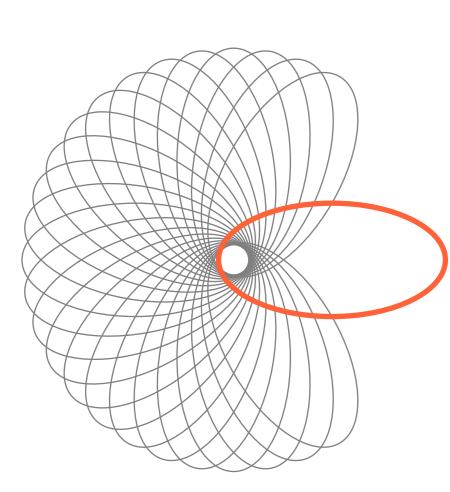
Application #1: Testing Shower Significance

- Example 1: The October Lyncids (OLY) were not detected: meteor density around the shower and its analogs is similar.
- **Example 2:** The July γ Draconids (GDR) were detected: meteor density around the shower exceeded the false positive rate.



 $\Delta \phi_b = 2 \sin \frac{1}{2} (\pi + \phi_2 - \phi_1)$ $\Delta \lambda_a = 2 \sin \frac{1}{2} (\lambda_{\odot,2} - \lambda_{\odot,1})$ $\Delta \lambda_b = 2 \sin \frac{1}{2} (\pi + \lambda_{\odot,2} - \lambda_{\odot,1})$ $\Delta \xi^2 = \min(\Delta \phi_a^2 + \Delta \lambda_a^2, \ \Delta \phi_b^2 + \Delta \lambda_b^2)$ $D_N^2 = (u_2 - u_1)^2 + w_1 (\cos \theta_2 - \cos \theta_1)^2 + \Delta \xi^2$ where $u = v_g / v_{\oplus}, \ \theta = \cos^{-1} (u_y / u)$, and $\phi = \tan^{-1} (u_x / u_z)$. D_N is based on geocentric speed and radiant (\vec{v}_g) and solar longitude (λ_{\odot}) instead of orbital elements.

Construction of shower analogs



- We construct a set of analogs for each shower.
- Analogs have the same geocentric speed and sun-centered ecliptic radiant, but are offset from the shower by at least 60° in λ_{\odot} .
- This effectively defines a shower as an enhancement lasting < 4 months.</p>
- We calculate D_N of *all* meteors relative to each analog and to the shower and compare.

Application #2: Shower membership probability

The ratio of meteors that lie close to the shower orbit vs. its shower analogs provides an estimate of shower membership probability as a function of *D*:

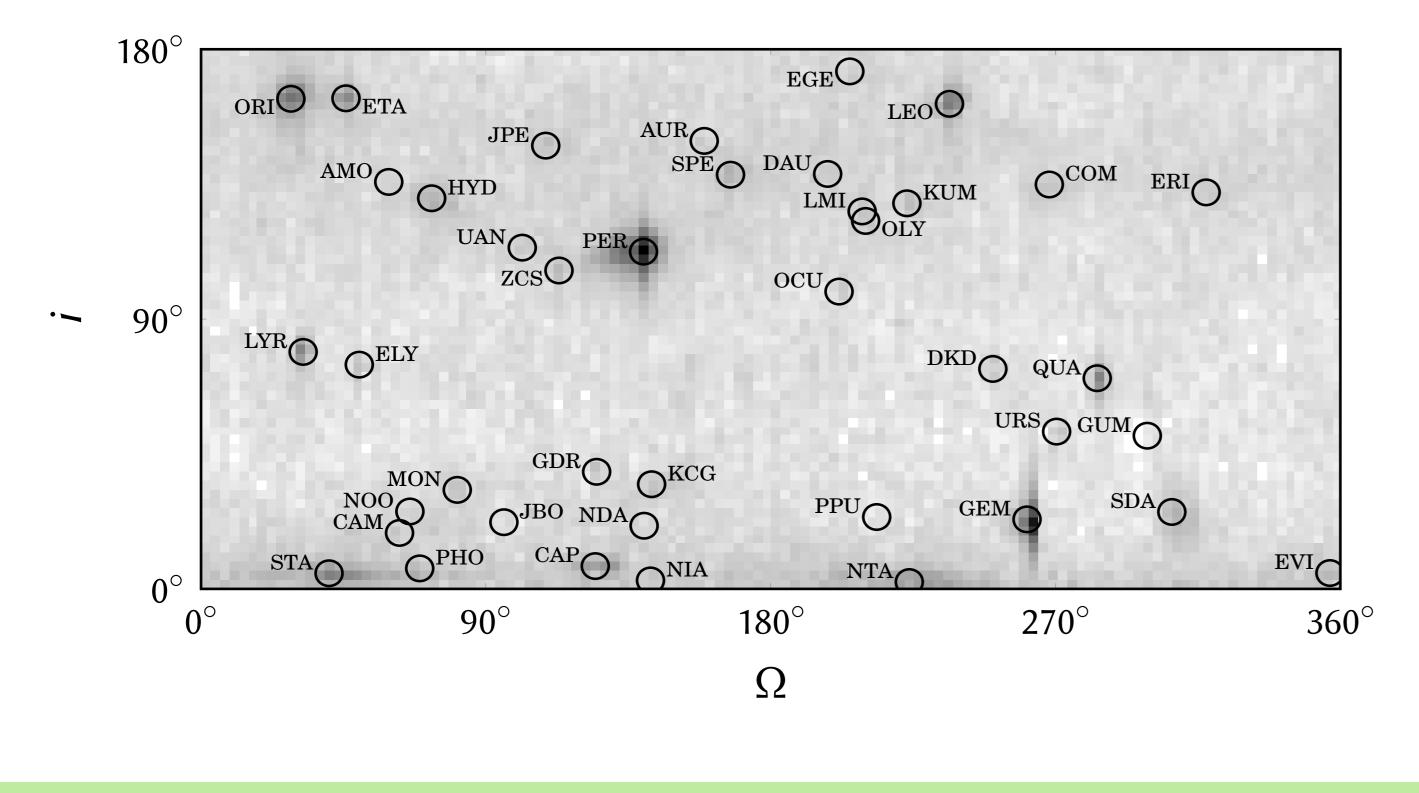
$$P(shower|D) \simeq rac{N_D - N_{spor,D}}{N_D}$$

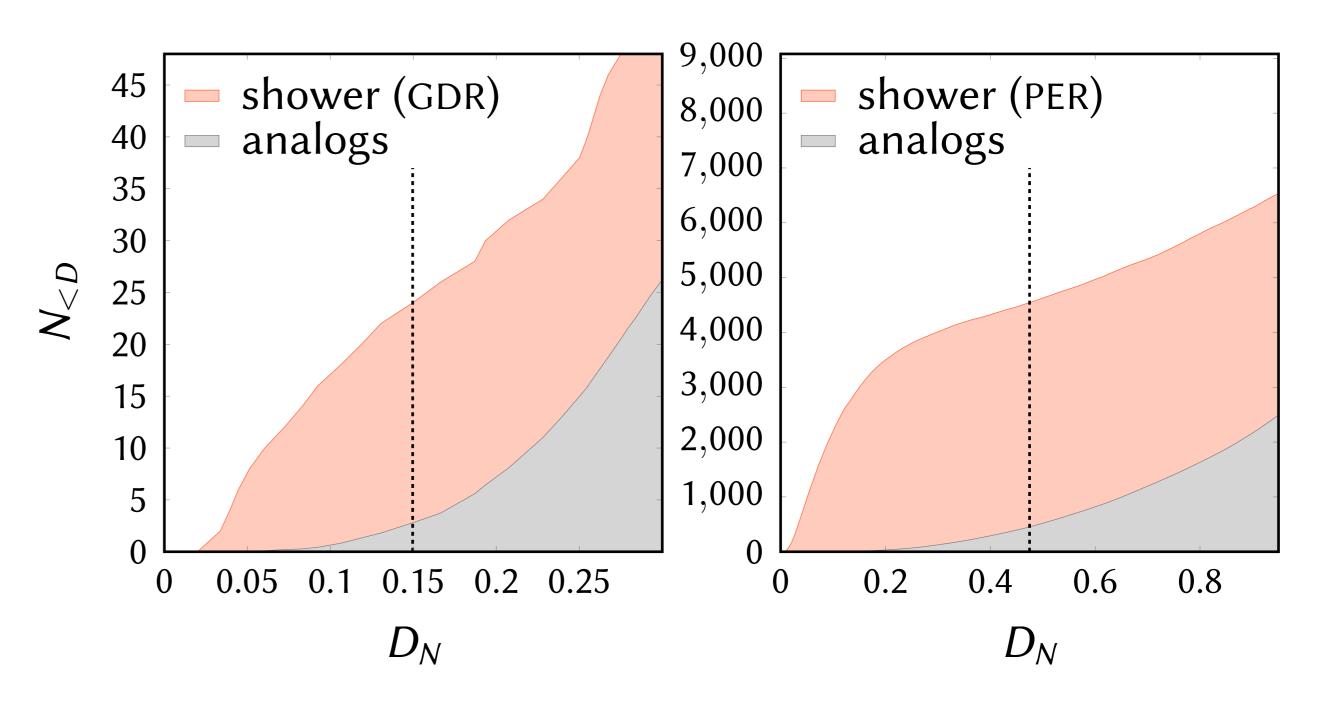
Application #3: Limiting Sporadic Contamination

- Shower analogs yield a false positive rate
- \triangleright D_{max} can be chosen to limit this to a desired percentage
- Example 1: $D_{max} = 0.15$ limits sporadic contamination to less than 10% for the July γ Draconids (GDR).
- **Example 2:** $D_{max} = 0.475$ limits sporadic contamination to less than 10% for the Perseids (PER).
- This provides us with an estimate of the false positive rate for shower association as a function of D_N.

Data

- We apply our method to 36,617 all-sky meteors from NASA All Sky Fireball Network [2] and the Southern Ontario Meteor Network (SOMN) [3].
- Possible showers were identified using orbital element heat maps (see below) or as short-lived clusters of meteors [4]



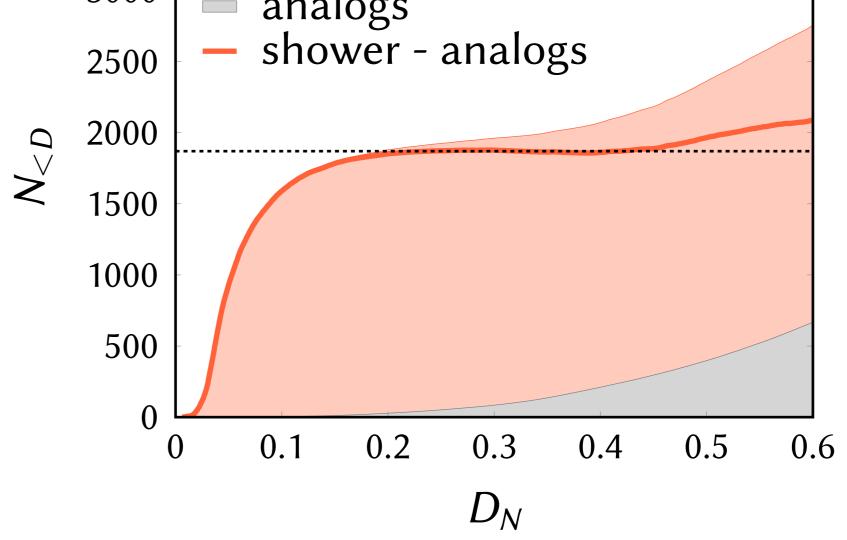


Application #4: Shower strength estimates

- It is not necessary to identify each shower member in order to estimate the strength of a meteor shower.
- **Example:** This CDF indicates that we have \approx 1870 Geminids (GEM) in our data set.

References

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Weryk R. J. et al., 2008, *EM&P*, 102, 241-246.
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