

Tutorial on Atomic Oxygen Effects and Contamination

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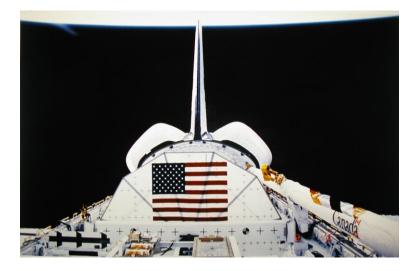
Applied Space Environments Conference Huntsville, Alabama May 15-19, 2017

www.nasa.gov

National Aeronautics and Space Administration

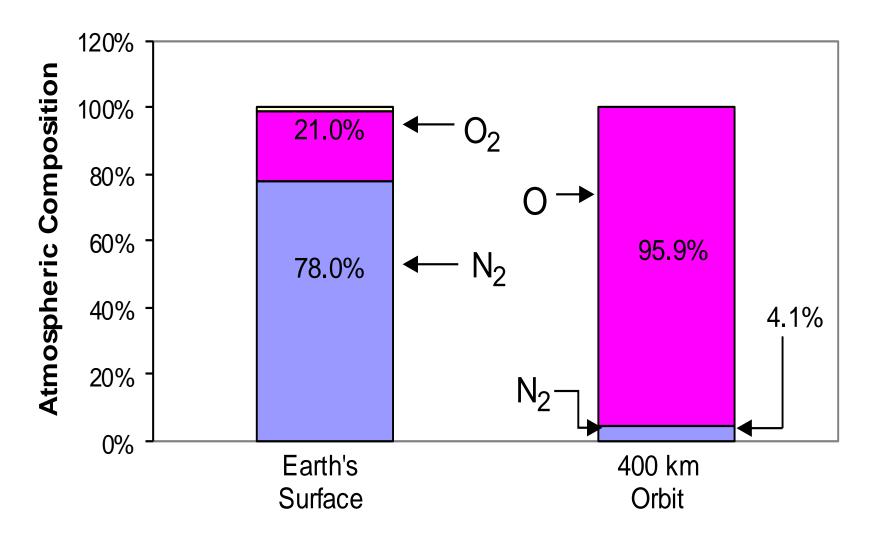
ATOMIC OXYGEN

Environment Interaction Visible on Space Shuttle Tail Section



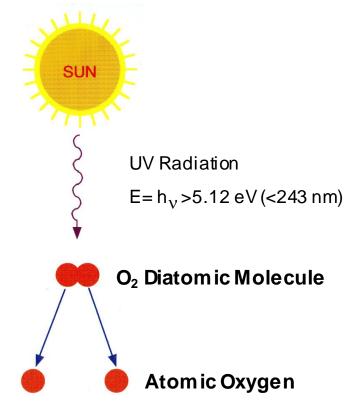


Atmospheric Composition



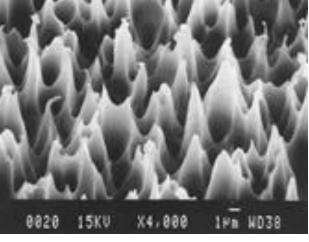
Atomic Oxygen in Low Earth Orbit





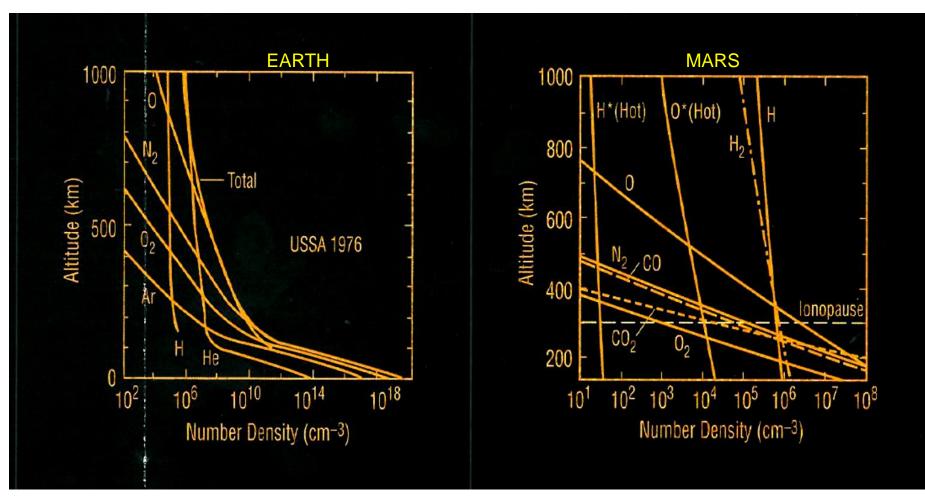
- AO is the predominant species from 180-650 km
- Average ram energy \approx 4.5 eV

LDEF Spacecraft CTFE after 8.99 x 10²¹ atoms/cm²



Polychlorotrifluoroethylene (CTFE)

Atmospheric Composition Comparison Between Earth and Mars



Graphs Courtesy of NASA JPL

Composition of Mars Atmosphere

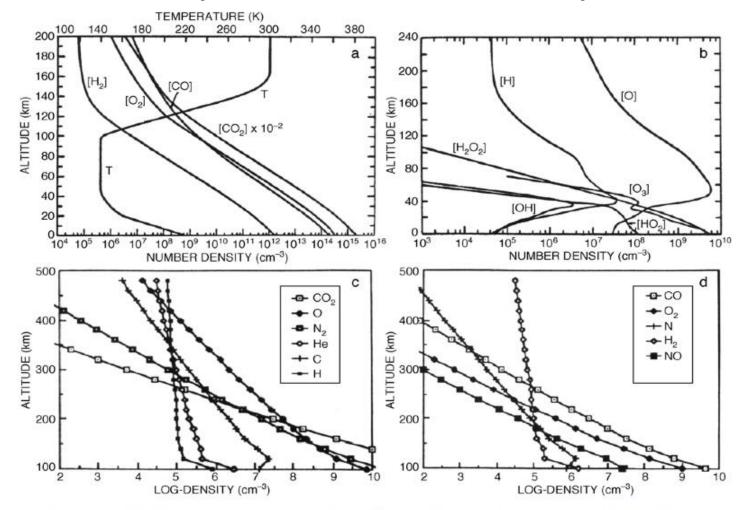
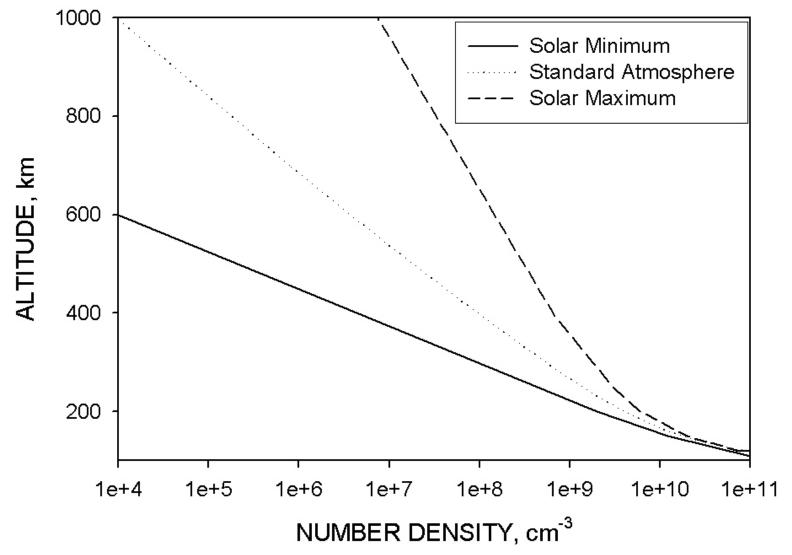


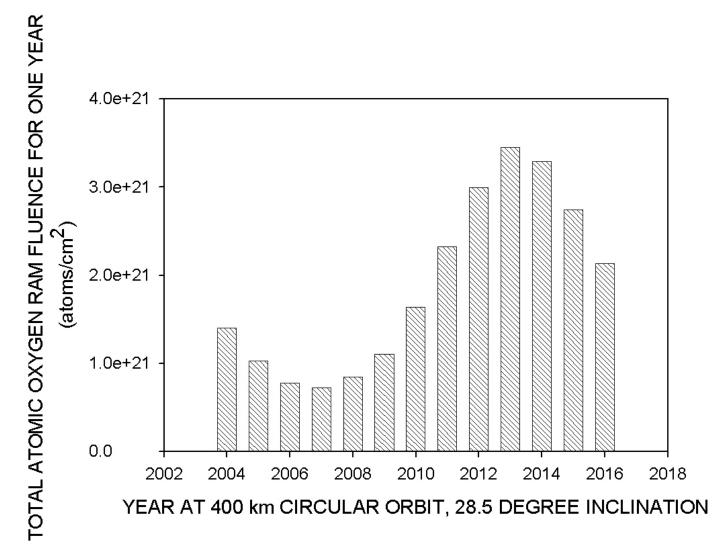
Figure 4-2. Martian Atmospheric Density Profiles for Various Constituents. a) CO₂, CO, O₂, and H₂ in an altitude range between 0 and 200 km, b) O, H, OH, H₂O₂, and O₃ from 0 to 240 km, c) CO₂, O, N₂, He, C, and H between 100 and 500 km and d) CO, O₂, N, H₂, and NO between 100 and 500 km.

Graphs Courtesy of Hank Garrett at NASA JPL

Atomic Oxygen Earth Atmosphere Number Density Dependence Upon Solar Activity

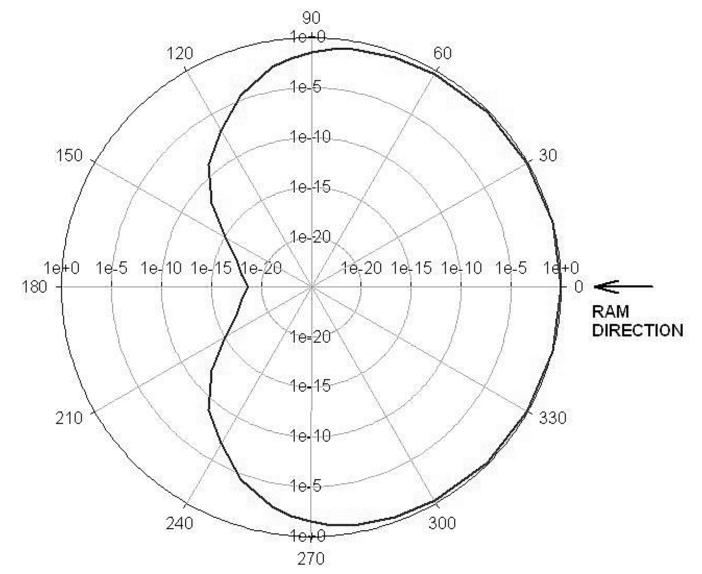


Solar Cycle Caused Variation in Level of Atomic Oxygen in Low Earth Orbit at 400 km

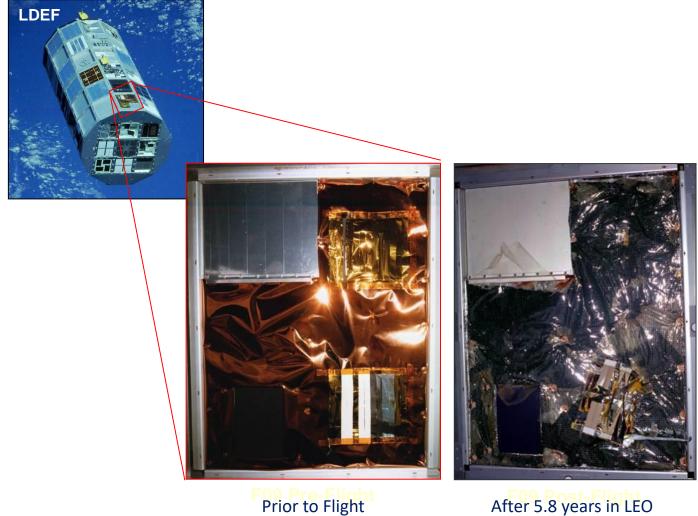


Logarithmic Polar Plot of Atomic Oxygen Arrival Flux

(400 km Earth orbit at 28.5° inclination and 1000 K thermosphere)

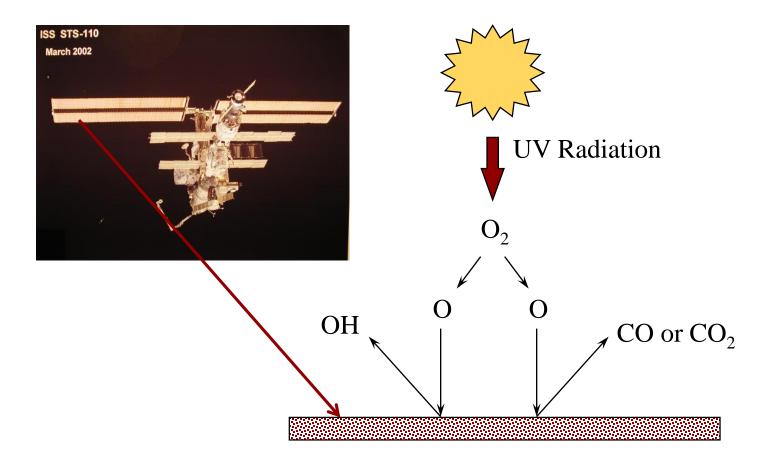


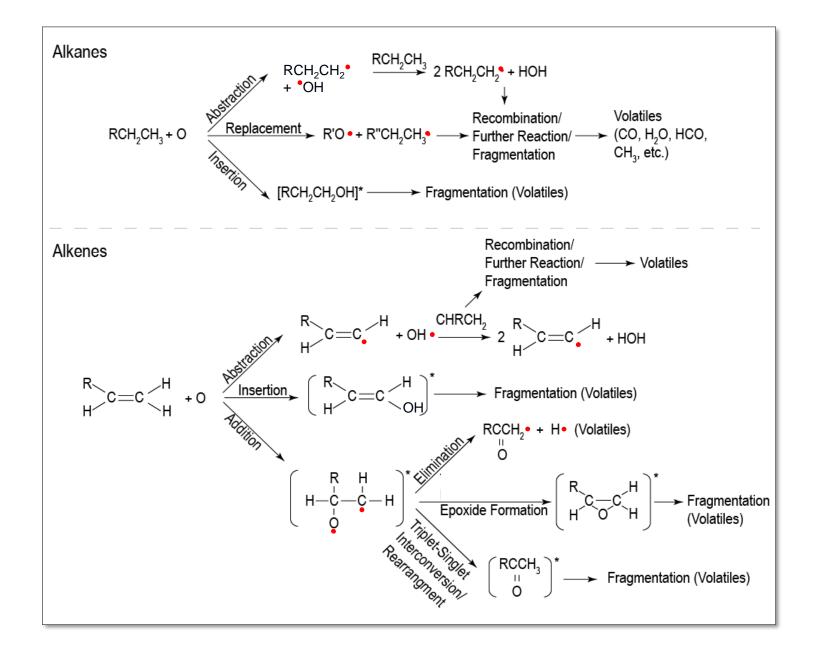
So What Can Atomic Oxygen Do to Spacecraft?



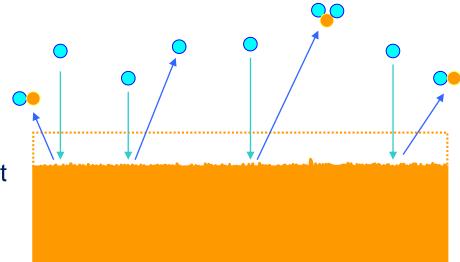
After 5.8 years in LEO

Basic Atomic Oxygen Interaction with Organic Surfaces





Atomic Oxygen Erosion Yield (E_y)



E_y is the **volume loss** per incident **oxygen atom** *(cm³/atom)*

Ey based on Mass Loss Measurements

Erosion Yield (E_y) of Sample where:

$$E_{y} = \frac{\Delta M_{s}}{A_{s}\rho_{s}F_{k}}$$

Atomic Oxygen Fluence

where:

$$F_k = \frac{\Delta M_k}{A_k \rho_k E_k}$$

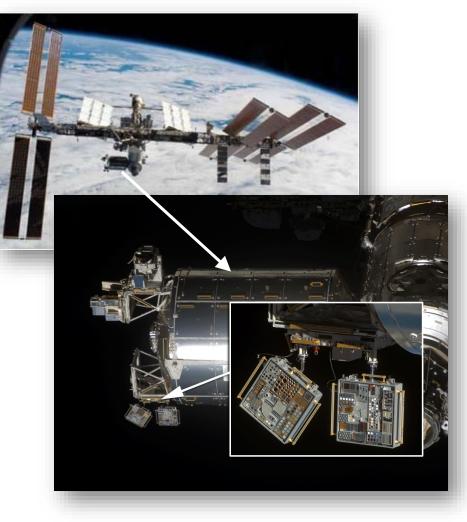
 $\Delta M_s =$ mass loss of polymer sample $A_s =$ area of polymer sample density of sample ρ_s = **F**_k = AO fluence measured by Kapton H witness samples ∆**M**⊾ = mass loss of Kapton H witness A_k = area of Kapton H witness density of Kapton H sample $\rho_k =$ (1.427 grams/cm³) $E_{k} =$ erosion yield of Kapton H

(3.0 x 10⁻²⁴ cm³/atom)

Material Tests in Low Earth Orbit (LEO) for Environment Interactions

Materials International Space Station Experiment (MISSE)





Atomic Oxygen Erosion Yields of Polymers Flown on MISSE-2 (PEACE)

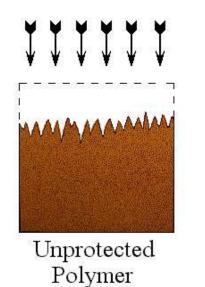
Material	Abbrev.	Ey (cm³/atom)	Ey Uncertainty (%)	Material	Abbrev.	Ey (cm³/atom)	Ey Uncertainty (%)
Acrylonitrile butadiene styrene	ABS	1.09E-24	2.7	Polyamide 6 or nylon 6	PA6	3.51E-24	2.7
Cellulose acetate	CA	5.05E-24	2.7	Polyamide 66 or nylon 66	PA 66	1.80E-24	12.6
Poly-(p-phenylene terephthalamide)	PPD-T (Kevlar)	6.28E-25	2.6	Polyimide	PI (CP1)	1.91E-24	2.8
Polyethylene	PE	3.74E-24	2.6	Polyimide (PMDA)	PI (Kapton H)	3.00E-24	2.7
Polyvinyl fluoride	PVF (Tedlar)	3.19E-24	2.6	Polyimide (PMDA)	PI (Kapton HN)	2.81E-24	2.6
Crystalline polyvinylfluoride w/white pigment	PVF (White Tedlar)	1.01E-25	4.1	Polyimide (BPDA)	PI (Upilex-S)	9.22E-25	3.0
Polyoxymethylene; acetal; polyformaldehyde	POM (Delrin)	9.14E-24	3.1	Polyimide (PMDA)	PI (Kapton H)	3.00E-24	2.6
Polyacrylonitrile	PAN	1.41E-24	3.3	High temperature polyimide resin	PI (PMR-15)	3.02E-24	2.6
Allyl diglycol carbonate	ADC (CR-39)	6.80E-24	2.6	Polybenzimidazole	PBI	2.21E-24	2.6
Polystyrene	PS	3.74E-24	2.7	Polycarbonate	PC	4.29E-24	2.7
Polymethyl methacrylate	PMMA	5.60E-24	2.6	Polyetheretherkeytone	PEEK	2.99E-24	4.5
Polyethylene oxide	PEO	1.93E-24	2.6	Polyethylene terephthalate	PET (Mylar)	3.01E-24	2.6
Poly(p-phenylene-2 6- benzobis oxazole)	PBO (Zylon)	1.36E-24	6.0	Chlorotrifluoroethylene	CTFE (Kel-f)	8.31E-25	2.6
Epoxide or epoxy	EP	4.21E-24	2.7	Halar ethylene- chlorotrifluoroethylene	ECTFE (Halar)	1.79E-24	2.6
Polypropylene	PP	2.68E-24	2.6	Tetrafluorethylene-ethylene copolymer	ETFE (Tefzel)	9.61E-25	2.6
Polybutylene terephthalate	PBT	9.11E-25	2.6	Fluorinated ethylene propylene	FEP	2.00E-25	2.7
Polysulphone	PSU	2.94E-24	3.2	Polytetrafluoroethylene	PTFE	1.42E-25	2.6
Polyeurethane	PU	1.56E-24	2.9	Perfluoroalkoxy copolymer resin	PFA	1.73E-25	2.7
Polyphenylene isophthalate	PPPA (Nomex)	1.41E-24	2.9	Amorphous Fluoropolymer	AF	1.98E-25	2.6
Graphite	PG	4.15E-25	10.7	Polyvinylidene fluoride	PVDF (Kynar)	1.29E-24	2.7
Polyetherimide	PEI	3.31E-24	2.6	*Ey > this value because sam	ple stack was par	tially or fully ero	ded through

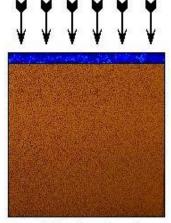
Material Testing in an Atomic Oxygen Environment Using Ground-Based Systems



Issues With Protective Coatings

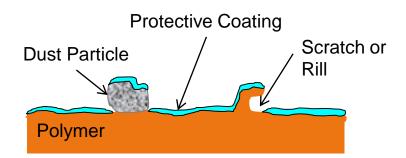




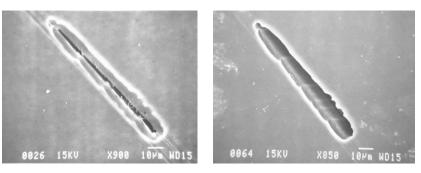


Protected Polymer

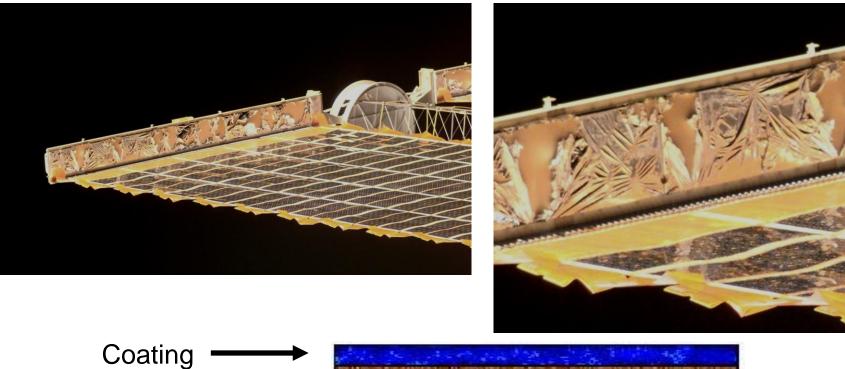
Imperfections in Thin Film Coatings

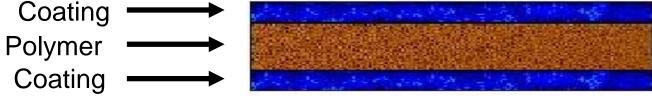


Aluminized Kapton Flown on LDEF



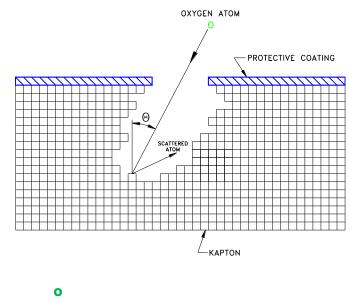
Blanket Box Cover Failure of Aluminized Kapton Observed on ISS





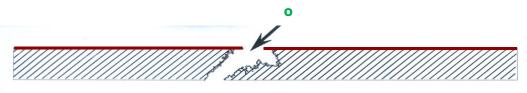
Monte Carlo Computational Model Predictions

- 2-D Computational modeling of atomic oxygen erosion of polymers based on observed in-space results
- Takes into account:
 - Energy dependence of reaction probability
 - Angle of impact dependence on reaction probability
 - Thermalization of scattered oxygen atoms
 - Partial recombination at surfaces
 - Atomic oxygen scattering distribution functions
- Modeling parameters tuned to replicate in-space erosion



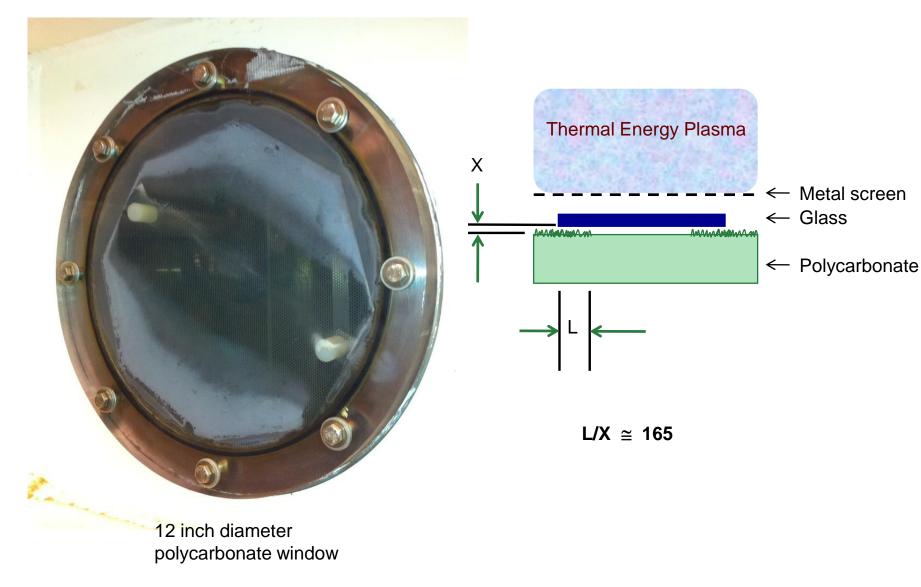


Aluminized on both sides

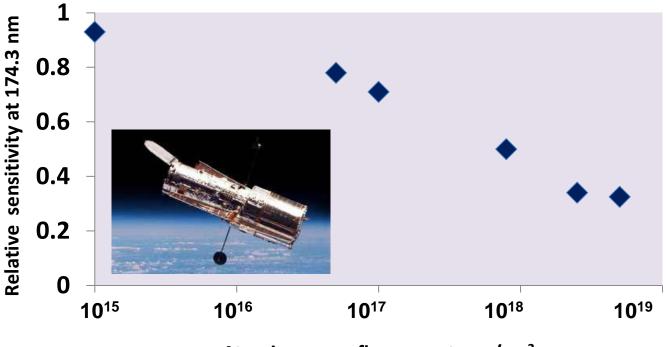


Aluminized on exposed side only

Atomic Oxygen Scattering



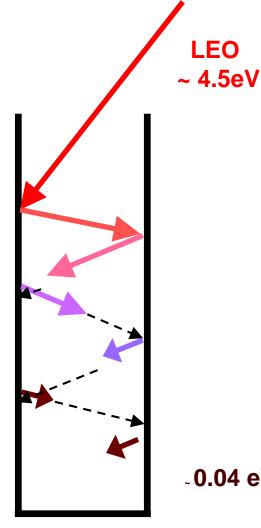
Change in Sensitivity of Cosmic Origins Spectrograph on Hubble Space Telescope



Atomic oxygen fluence, atoms/cm²

Experienced a far UV sensitivity decline ranging from 3-15%/year (based on data from June 2009 through mid-February 2010)

Scattering and Thermal Accommodation of Low Earth Orbital Atomic Oxygen



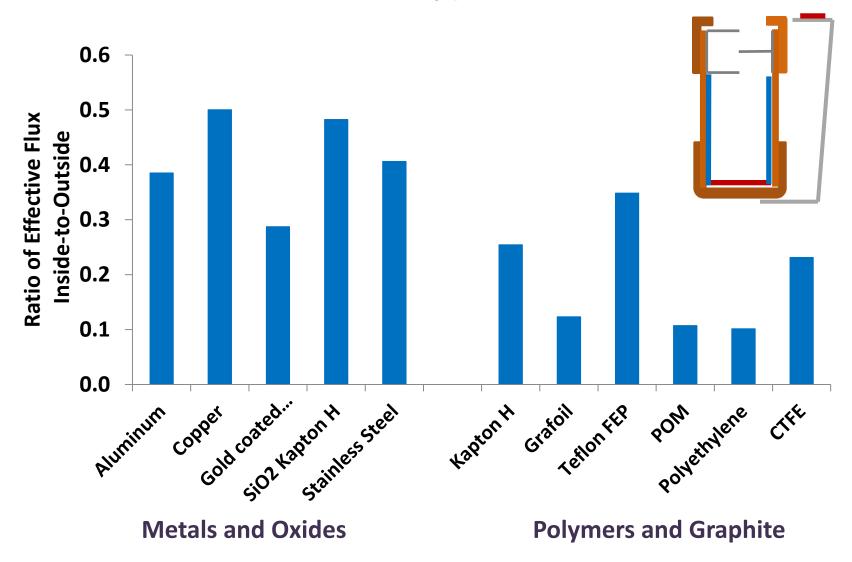
~0.04 eV

LEO

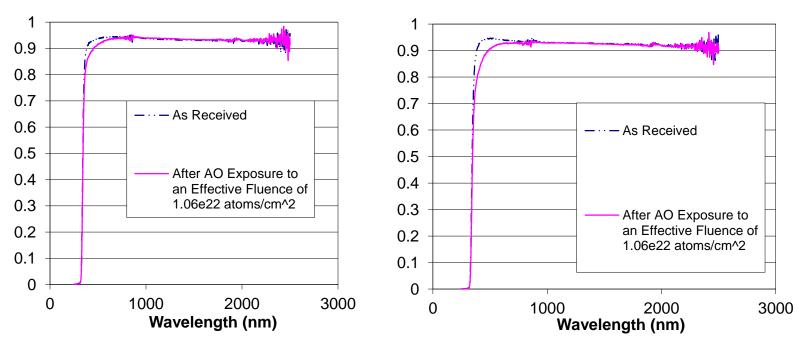
Possible Events Upon Impact:

- Reaction
- Recombination
- Scattering
- Partial thermal accommodation
- Ejection out the entrance

Test of Mock Aperture with Various Types of Liners



Total Transmittance as a Function of Wavelength for Coverglass Prior to and After Exposure to Atomic Oxygen



AR Coated

Conductive AR Coated

Mirrored Silver Back of Solar Cell Prior to and After Exposure to Atomic Oxygen

As Received

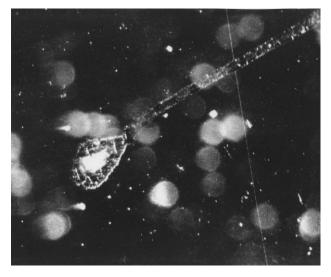
After Exposure to an AO Effective Fluence of 2x10²¹ atoms/cm²





Oxidative Cracking of Silicone

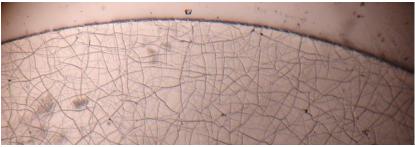
DC 93-500 Silicone Exposed to LEO Atomic Oxygen on STS-46 Fluence = 2.3×10^{20} atoms/cm²





Pre-flight

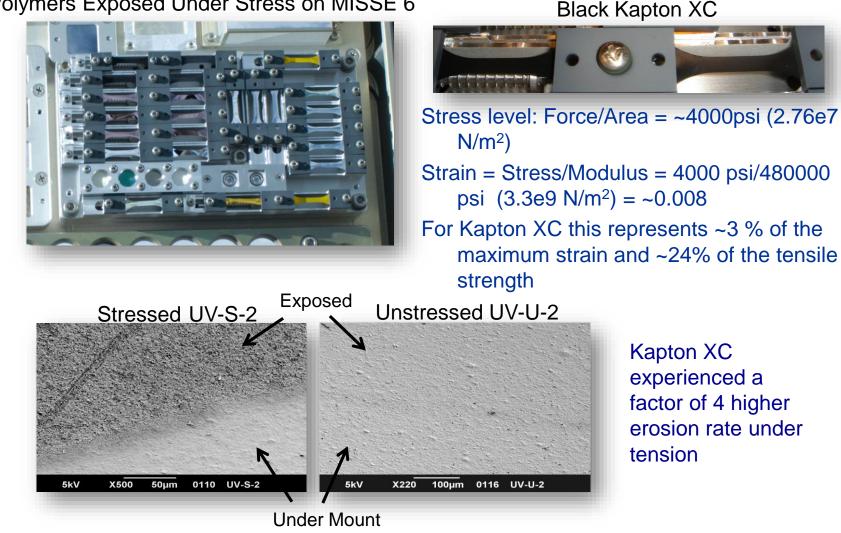
Post-flight



Stress Dependent Atomic Oxygen Erosion of Black Kapton XC

Stressed (left) and Unstressed (right)

Polymers Exposed Under Stress on MISSE 6

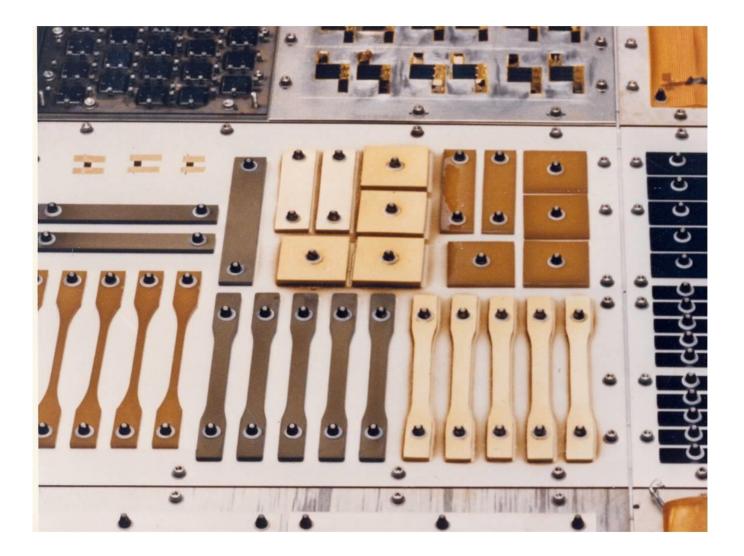


CONTAMINATION

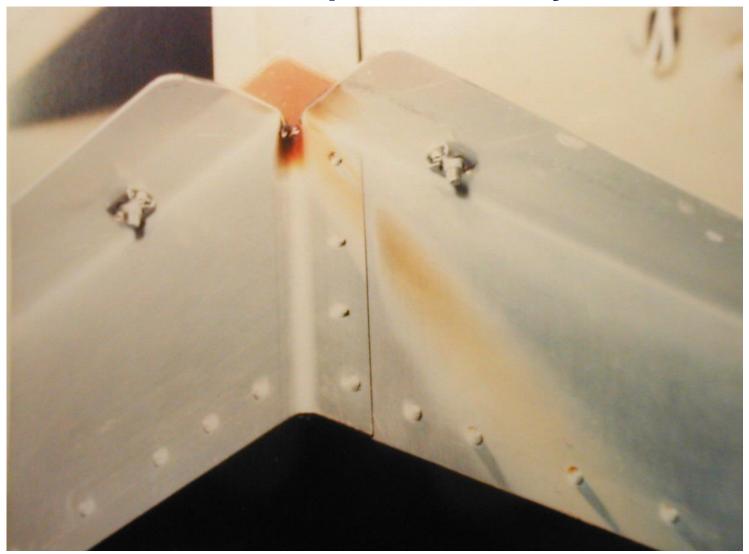
Waste Water Dump Residue on Polished Plate Micrometeoroid and Debris Collector Experiment on MIR



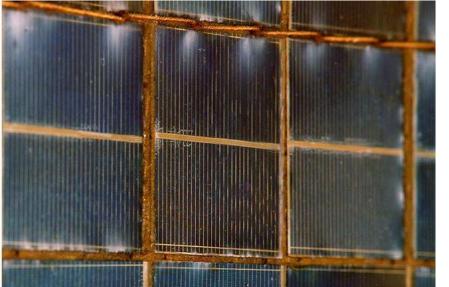
Silicone Contamination on LDEF Tray

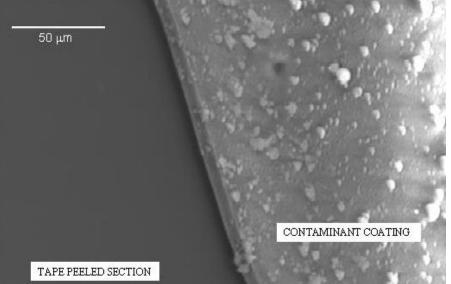


LDEF Experiment Tray



Oxidized Silicone Contamination on the MIR Solar Array After 10 Years in LEO





Frosty deposits on solar cell cover glasses

~ 4.6 micron thick silica deposits

Silicone Contamination on Mir PV Array Optical Solar Reflector (OSR)

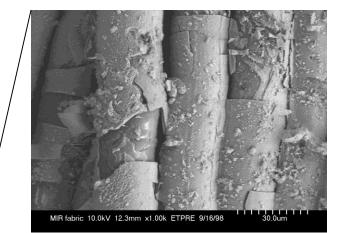


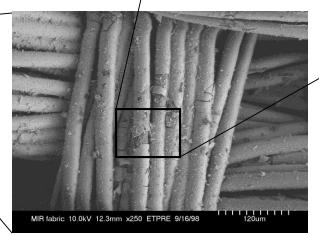
Suture

<image>

Silicone Contamination on MIR Flexible Handhold Fabric Tape







Summary

Atomic Oxygen:

- Atomic oxygen is the most predominant specie in LEO
- Atomic oxygen is reactive and energetic enough to break chemical bonds in materials
- Reaction products with polymers and carbon containing materials are volatile (typically CO and CO₂)
- Metals and inorganics experience surface oxidation in some cases leading to shrinkage and cracking or spalling
- Atomic oxygen can thermalize on contact and scatter from surfaces leading to further reaction, which is dependent on the materials it contacts and geometry
- The effect that atomic oxygen has on a particular material on a spacecraft is dependent upon how much atomic oxygen arrives at the surface, atom energy, and can be affected by mechanical loading, temperature, and other components in the environment (UV radiation, charged particles...)

Summary

Contamination:

- Contamination can be as a result of actions of a nearby neighboring spacecraft or experiment, or by self contamination
- Contaminants can be fixed on a surface by UV radiation or atomic oxygen
- Contamination can cause detrimental changes in optical and thermal properties

Overall:

 Each situation is unique and for accurate prediction of degradation of a material or component, it should be tested or modeled in a configuration representative of how it will be used

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Publications:

http:/ntrs.nasa.gov

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