

Toward large FOV high-resolution X-ray imaging spectrometer: microwave multiplexed readout of 32 TES microcalorimeters





NASA/GSFC: Wonsik Yoon, Joseph S. Adams, Simon R. Bandler, James A. Chervenak, Aaron M. Datesman, Megan E. Eckart, Fred M. Finkbeiner, Richard L. Kelley, Caroline A. Kilbourne, Antoine R. Miniussi, Samuel H. Moseley, Frederick S. Porter, John E. Sadleir, Kazuhiro Sakai, Stephen J. Smith, Thomas R. Stevenson, Nicholas A. Wakeham, Edward J. Wassell, Edward J. Wollack National Radio Astronomy Observatory (NRAO), University of Virginia (UVA), NASA/GSFC: Omid Noroozian

NIST/Boulder: Dan Becker, Douglas A. Bennett, Joseph W. Fowler, Johnathon D. Gard, Gene C. Hilton, John A.B. Mates, Carl D. Reintsema, Daniel S. Swetz, Joel N. Ullom, Leila R. Vale

Flux ramp much faster than input signal

and proportional to TES current

 $\phi = \frac{2\pi M}{\Phi_0} I_{TE}$

For demodulation

 \rightarrow phase shift ϕ is almost constant during ramp

Email: wonsik.yoon@nasa.gov

Abstract

We performed a small-scale demonstration at GSFC of highresolution x-ray TES microcalorimeters read out using a microwave SQUID multiplexer. This work is part of our effort to develop detector and readout technologies for future space based x-ray instruments such as the microcalorimeter spectrometer envisaged for Lynx, a large mission concept under development for the Astro 2020 Decadal Survey. In this paper we describe our experiment, including details of a recently designed, microwave-optimized lowtemperature setup that is thermally anchored to the 50 mK stage of our laboratory ADR. Using a ROACH2 FPGA at room temperature, we simultaneously read out 32 pixels of a GSFC-built detector array via a NIST-built multiplexer chip with Nb coplanar waveguide resonators coupled to RF SQUIDs. The resonators are spaced 6 MHz apart (at ~5.9 GHz) and have quality factors of ~15,000. Using fluxramp modulation frequencies of 160 kHz we have achieved spectral resolutions of 3-5 eV FWHM on each pixel at 6 keV. We will present the measured system-level noise and maximum slew rates, and briefly describe the implications for future detector and readout

Microwave SQUID multiplexing



• A few GHz of bandwidth per amplifier channel • TESs couple to unique microwave resonator • RF SQUIDs built into microwave resonator • Inductance modulates resonance frequency • Linearize output without feedback (common modulation)

• Single microwave feed-line can read out hundreds of pixels

Lynx

A large mission concept under development by NASA for the Astro 2020 Decadal Survey

Microcalorimeter detector array for Lynx

- Energy resolution: better than 3 eV FWHM at 0.2 10 keV
- Number of readout channel > 56,000
- (number of pixel > 150,000 with hydra)
- Various angular resolution, energy resolution and count rates
- Sensor: Transition-edge sensor(TES) or magnetically coupled calorimeter (MCC)
- Initial approach: Use position-sensitive TES microcalorimeter, "Hydras". These have multiple absorbers attached to each sensor
- See also

5 resonators with 5 TESs

- PE-46, "The Design of the Lynx X-ray Microcalorimeter," Simon Bandler et al.

- PE-59, "Design and optimization of multi-pixel transition-edge sensors for X-ray









100 Hz 1 kHz NEP with X-ray: 2.60 – 2.81 eV @ ~ 5.9 keV No NEP degradation compared to five resonator multiplexing !

16 resonator multiplexing

(5 TESs, 11 without TES)

Low temperature uMUX setup



Room temperature electronics

• Coupling constant : 8.73

(SQUID input – flux ramp circuit)



Ongoing work: microwave multiplexing of 32 TES microcalorimeters

Changed TES detector chip with the goals of:

1. Improved energy resolution – expect non-multiplexed ΔE_{FWHM} =1.6 eV based on measurements of similar chip 2. 32-channel multiplexing – bond pad layout compatible with μ MUX chip layout





5860	5880	5900 Energy (eV)	5920	5940	5860	5880	5900 Energy (eV)	5920	5940
• <∆E> = 3	.53 eV l	FWHM @) 5.9 ke	V includi	ng all me	asured	pixels		
- TES 18: Re	sonator	intentiona	lly turne	ed off to av	oid interfe	erence wi	th other r	esonator	^
- TES 20 / 2	5: Relati	vely poor	energy r	esolution o	due to exc	ess detec	tor noise		
In progress	: fixed v	viring, rep	placed d	amaged N	lyquist ch	ip (4/20	17), requ	ired nev	ν μΜՍΧ
chip due to	subseq	uent dam	lage dur	ing handl	ing, μMU	X screen	ing unde	rway (4-	7/2017)
Future wor	·k : micro	owave rea	ndout of	'Hydra' p	ixels for L	.vnx (see	Bandler.	Smith)	