

# Simulation-Based Height of Burst Map for Asteroid Airburst Damage Prediction



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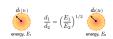
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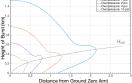
#### Overview

Entry and breakup models predict that airburst in the Earth's atmosphere is likely for asteroids up to approximately 200 meters in diameter. Objects of this size can deposed over 250 megations of energy into the atmosphere. Fast-unning ground diamage prediction costs for such events with heavity our methods developed from moder warpon treasarch to estimate method and the second second and the second se

#### Improving the Height of Burst (HOB) Map

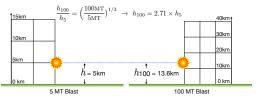
- With no buoyancy or other length scales, point sources blasts are self-similar
- · Yield scaling allows us to relate overpressures between blasts of different strengths





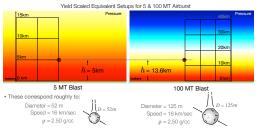
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- · Buoyancy introduces an additional length-scale breaking the self-similarity of the blasts -- can be neglected for "small" blasts ( < 5-10MT) since
- · Propagation times are short so acceleration due to gravity doesn't have time to act
- Distances are small as compared to the scale-height of the earth's atmosphere (~7-8km).
- e.g. Yield scaling predicts that a 100 MT blast at a burst height of 13.6 km will have the same scaled ure as a 5 MT blast detonated at an altitude of 5 km around over



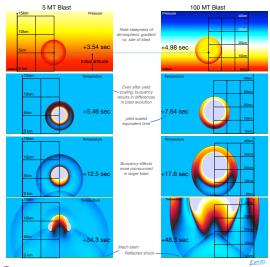
#### Introducing Buoyancy

- Buovancy enters through the pressure gradient due to gravity. <a href="https://doi.org/action.org/action.org/light-action.org/lig characteristic length: the scale height of the atmosphere
- . With this second length scale, the two blasts are no longer strictly self-similar.
- . When the scale height (7-8 km for Earth) is large compared to the blast footprint, buoyancy effects are small
- . However, at the high energy levels associated with asteroids whose diameter is greater than about 100m, these effects can be significant.
- . Here is the yield scaled equivalent setup for the 5 & 100 MT airburst in (1) with buoyancy. Note that while the blasts themselves are similar, the background pressure gradient in the atmosphere is significantly steeper at the larger yield



### Blast Evolution

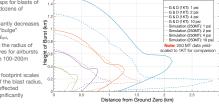
- The similarity parameter = Const establishes the time scale equivalence between two blasts.
- $\frac{t_2}{t_1} = \frac{t_{100MT}}{t_1} = 1.4079$ • For 5 and 100MT, this implies  $\left(\frac{t_2}{t_1}\right)^2 = \frac{\rho_{\circ 2}}{\rho_{\circ 1}} \frac{E_1}{E_2} \left(\frac{d_2}{d_1}\right)^5 \text{giving}$
- . The increase in time scale compounds the effects of buoyancy. Not only is the pressure gradient steeper. but it acts over a longer period of time.
- . As an example, compare 5 & 100 MT blasts. In the image sequence below, burst height, length and time scales have been sized to provide yield scaled equivalent snapshots of the blast evolution
- · Buovancy modifies the ground overpressure footprint, resulting in different damage predictions



#### Updated Height of Burst Map

- · Used simulation to numerically generate HOB maps for blasts of various yields at dozens of altitudes
- Buovancy significantly decreases the height of the "bulge associated with Heat
- This can increase the radius of lethal overpressures for airbursts of asteroids in the 100-200m range.
- Since the ground footprint sca with the square of the blast radius, estimates of the effected population are significantly improved.

#### Overpressure for 1 KT Blast vs. Burst Height Ground Overpress



#### Outcome

The probabilistic risk tool (Mathias, 2017) now interpolates between appropriate numerically generated HOB maps to give improved prediction of ground footprint for larger entry energies. Since large asteroids typically have lower burst heights, these predictions can be significantly effected by the steep gradients seen in the HOB map above. This modified risk tool was used for analysis supporting NASA's Science Definition Team's work to quantify the threat posed by Near Earth Objects.

## Acknowledgements

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