Differential deposition for the figure correction of X-ray optics

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Full-shell electroformed optics @MSFC





State of art - full-shell optics

X-ray Optics - State of art

- Full-shell: <5 arc sec FWHM; 10 to 15 arc sec HPD
- A key factor that limits the angular resolution is <u>Axial Figure Imperfections</u>





Imaging quality of X-ray optics can be significantly improved

if the RMS height variations can be reduced

Concept of differential





Use of physical vapor deposition to selectively deposit material on the mirror surface to smooth out figure imperfections





Differential Deposition – proof of concept

•Proof of concept on smaller scale NIH optics





•Use of existing sputter deposition chambers

•Demonstrated improvement through metrology data





Custom built vacuum chambers



- Horizontal chamber for full shell optics can accommodate upto 0.25m diameter, 0.6m length
- Vertical chamber for segmented and very large full-shell optics (0.5 m diameter)
- Computer controlled translation and rotation stages with encoders
- Matlab GUI interface to control the stages



Metrology results













X-ray test results -1





X-ray test results -2





Working towards 2nd stage improvement



Future work



In-situ metrology - VLTP approach



Schematic of in-situ metrology. The path from the optical board to the test surface passes into the vacuum chamber through an optical feed-through flange to a penta-prism which directs the laser light to and from the test surface.

Detailed stress analysis

Active slit approach



Design concept of active slit approach



Spin-off applications

NIST – Neutron microscope

- Prototype optics demonstrated 70 microns spatial resolution
- Goal 10 microns spatial resolution (~1 arc sec)

National Ignition Facility

- X-ray imaging is critical to the physical understanding of ICF implosions
- Need for high-resolution 5microns (FWHM) spatial resolution (few arc secs) imaging optics for hard 10-25 keV x-rays





(Newly-Funded) Direct fabrication of full-shell X-ray optics

Replication



Direct Fabrication





Material	Density (g/cm³)	СТЕ (10 ⁻⁶ / К ⁻¹)	Elastic Modulus GPa	Yield Strength MPa
Fused Silica	2.2	0.5	72	48*
Beryllium	1.8	12	318	240
BeAL-162MET	2.1	24	69	276
AlSi	2.8	13.9	193	314
Duralcan F3S.30S AlSi+SiC(30% by vol)	2.8	14.6	120	210

Mechanical Properties of Potential Mirror Substrate Materials

*Maximal achievable value. The 'working' value is typically much less and depends on the surface/subsurface condition.

Ideally, the mirror shell has low density, low coefficient of expansion (CTE), high modulus of elasticity and high yield strength. It should also be a material that is not too difficult to figure and polish.

- Be + NiP (CATS-ISS telescope)
- BeAl +NiP
- AlSi + NiP

Zeeko polishing machine





Wear function characterization:

Wear rate is proportional to

- Velocity of bonnet depends on
 - Spindle rotation
 - Head attack angles
- Bonnet pressure depends on
 - Internal pressure of bonnet
- Bonnet structural and mechanical prope
 Parameter optimization

Bonnet pressure Spindle speed Tool Offset

Test polishing runs

- 100mm dia NIP electroplated Al flat samples.
- Initially diamond turned and then polished over small linear regions "trenches".
- The polishing parameters were varied from one trench to another and the wear function dependence on these parameters was determined.





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NASA

ZEEKO polishing – demonstration on existing mandrel





Surface height error profile of the hyperbolic side of a NiP plated mandrel before (upper curve) and after (lower curve) polishing using a 40 mm bonnet.

	before	after
Figure error (St. Dev.)	500 nm	10.7 nm
Slope error (> 2 cm) (RMS)	6.32 arcsec	0.30 arcsec
Low frequency (> 7 cm) slope error (RMS)	2.66 arcsec	0.09 arcsec
Mid frequency (2-7 cm) slope error (RMS)	5.73 arcsec	0.29 arcsec



Mandrel > 5x better than any made with conventional polishing