

Heavy Ion Testing at the Galactic Cosmic Ray Energy Peak



Jonathan A. Pellish¹, M. A. Xapsos¹, K. A. LaBel¹, P. W. Marshall², D. F. Heidel³, K. P. Rodbell³, M. C. Hakey⁴, P. E. Dodd⁵, M. R. Shaneyfelt⁵, J. R. Schwank⁵, R. C. Baumann⁶, X. Deng⁶, A. Marshall⁶, B. D. Sierawski⁷, J. D. Black⁷, R. A. Reed⁷, R. D. Schrimpf⁷, H. S. Kim⁸, M. D. Berg⁸, M. J. Campola⁸, M. R. Friendlich⁸, C. E. Perez⁸, A. M. Phan⁸, and C. M. Seidleck⁸

¹: Radiation Effects and Analysis Group, NASA/GSFC Code 561.4, 8800 Greenbelt RD, Greenbelt, MD 20771 USA;
²: NASA Consultant, Brookneal, VA 24528 USA; ³: IBM T. J. Watson Research Center, Yorktown Heights, NY 10598 USA
⁴: Formerly with IBM System and Technology Group, Essex Junction, VT 05452 USA; ⁵: Sandia National Laboratories, Albuquerque, NM 87185 USA
⁶: Texas Instruments, Dallas, TX 75243 USA; ⁷: Department of Electrical Engineering and Computer Science, Vanderbilt University, Nashville, TN 37235 USA
⁸: MEI Technologies (NASA/GSFC), Greenbelt, MD 20771 USA.



Jonathan Pellish

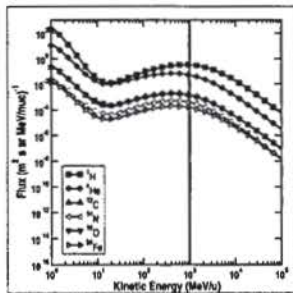
Abstract

A 1 GeV/u ⁵⁶Fe ion beam allows for true 90° tilt irradiations of various microelectronic components and reveals relevant upset trends for an abundant element at the galactic cosmic ray (GCR) flux-energy peak.



Space Environment

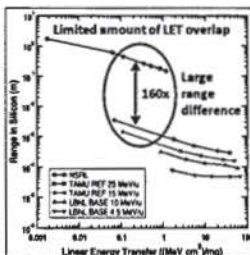
- The galactic cosmic ray (GCR) environment has a flux-energy peak at ≈1 GeV/u
- At this level of kinetic energy, LETs > 1 ((MeV·cm²)/mg) are dominated by iron
- 1 GeV/u ⁵⁶Fe has an LET of 1.2 ((MeV·cm²)/mg) and a range in silicon of 15 cm



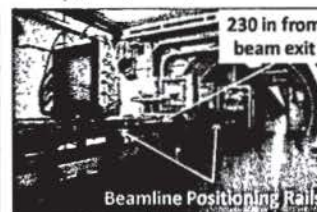
GCR Fluxes for geostationary orbit at solar minimum (CREME96)

Facility

The NASA Space Radiation Effects Laboratory
 Brookhaven National Laboratory, Long Island, NY USA



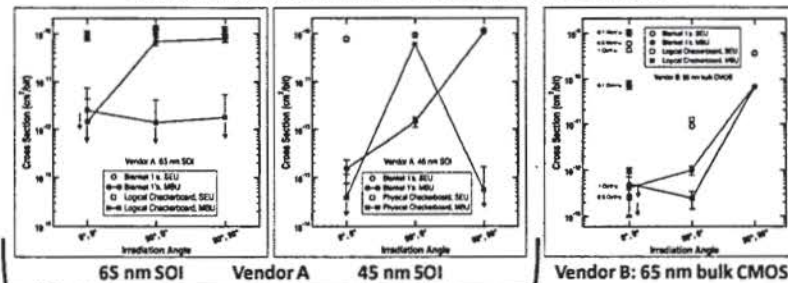
- The NSRL compared to other heavy ion facilities
- Note that the JYFL REF is similar to LBNL BASE facility



- Target Room
- Facility has rotation and translation stages
 - Can stack multiple experiments
 - Experiments conducted ≈230 in from beam exit

Static Random Access Memories

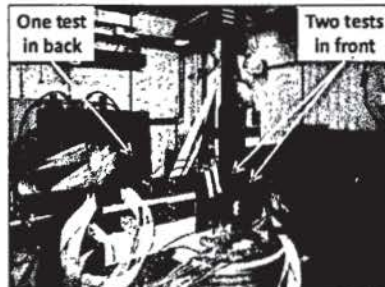
Note: Different scales for Vendor A's SOI and Vendor B's bulk CMOS



→ Solid lines connect 1 GeV/u MBU cross sections ←

- Vendor A's SOI devices have both a data pattern and orientation dependence
- Vendor B's bulk CMOS device only has an orientation dependence

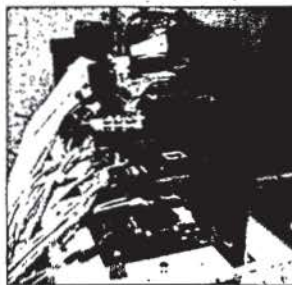
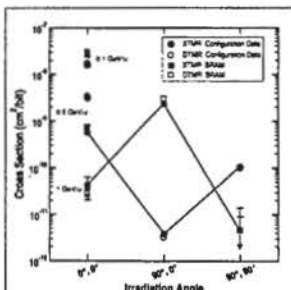
Pattern and orientation sensitivities arise from SOI's inter-device isolation and bit cell layout. These features are muted in bulk CMOS due to charge transport.



- Test setup for normal incidence irradiations
- Taking advantage of 8 in × 8 in beam spot
 - Two setups in front and one in back
- Roughly ± 2% uniformity over this area

Field Programmable Gate Array

→ Solid lines connect 1 GeV/u configuration and BRAM cross sections ←



- Extreme variation between configuration and BRAM cross section based on orientation
- Suspect physical layout is responsible

- Test setup for 90° tilt, 0° roll irradiations
- Taking advantage of 8 in × 8 in beam spot
- All four DUTs are stacked (FPGA on bottom)
- Roughly ± 2% uniformity over this area

Conclusions

- First time the NSRL facility has been used to characterize highly-scaled commercial technologies
- 1 GeV/u ⁵⁶Fe beam allowed for true 90° tilt irradiations
- Extreme upset cross section variation observed as a function of roll angle, data pattern, and storage elements
- Data provides impetus to study these limiting irradiation conditions with radiation transport modeling since studies like these are not feasible on a regular basis