The *Lynx* Mission Concept 2017 Accomplishments and 2018 Goals

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-Presented On behalf of the *Lynx* Team





Meet Lynx!

One of 4 large missions under study for the 2020 Astrophysics Decadal, Lynx is an X-ray observatory that will directly observe the dawn of supermassive black holes, reveal the invisible drivers of galaxy and structure formation, and trace the energetic side of stellar evolution and stellar ecosystems.



- 50–100× gain in sensitivity via high throughput with high angular resolution
- 16× field of view for arcsecond or better imaging
- 10–20× higher spectral resolution for point-like and extended sources



RAY OBSERVAT

Lynx will contribute to nearly every area of astrophysics and provide synergistic observations with future-generation ground-based and space-based observatories, including gravitational wave detectors.

Decadal Deliverables Updated Schedule



ROPH	Study Deliverables	NASA
M1	 Comments on Study Requirements and Deliverables Accept the study requirements/deliverables and submit plan or Provide rationale for modifying requirements/deliverables 	April 29 2016
01	Optional: Initial Technology Gap Assessment – To impact PCOS/COR/ExEP 2016 technology cycle	June 30 2016
02	Optional: Update Technology Gap Assessments	June 2017
M4	 Interim Report Provide science case and mission concept (use CML 3 as a guide) Deliver initial technology roadmaps; estimate technology development cost/schedule CML 4 tailored approach (optional) 	March 2018
03	Update Technology Gap Assessments	June 2018
M6	 Draft Final Report at Concept Maturity Level 4 Audit / Freeze Point Design Provide science case and mission concept (use CML4 as a guide) Support independent cost estimation/validation process Submit to HQ for CATE 	Jan 2019
M6'	CATE report returned by HQ to STDTs for incorporation into M7	May 2019
M7	Final Report / incorporate CATE report + final changes – As described in study success criteria chart 15	June 2019
		4

*Note: Schedule relaxed from original by ~4 months due to decadal committee schedule delay

New Members!

STDT Members





Zoltan Haiman, Columbia



Terri Brandt, **PCOS Program Office Acting Chief Scientist**

- 22 STDT Members
- 8 Science Working Groups
- OpticsWorking Group
- Calibration Working Group
- CommunicationsWorkingGroup
- Instrument Working Group
- Ex-officio International members

Over 275 total members!

Ex-Officio



Peter Jonker, **SRON-Appointed**



Giovanni Pareschi **INAF-Appointed**





Andrey Kravtsov, Chicago

Science of Lynx



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The Dawn of Black Holes

The Invisible Drivers of Galaxy and Structure Formation

The Energetic Side of Stellar Evolution and Stellar Ecosystems



Endpoints of stellar evolution

Stellar birth, coronal physics, feedback

Impact of stellar activity on habitability of planets

AAS Lynx Science Talks



- Monday, 9:45AM-10:05AM, Grav. Wave SIG, Lynx and LISA, R. Petre
- Tuesday, 10:00AM-11:30AM, 103.04: Lynx Mission Concept Study, A. Vikhlinin
- Wednesday, 2:00PM-3:30PM, 223.08, Future prospects with the Chandra and XMM source catalogs: Setting the stage for Lynx, R. Hickox
- Thursday, 2:00PM-3:30PM, 332.01: Implications from XMM and Chandra Source Catalogs for Future Studies with Lynx, A. Ptak
- Thursday, 5:30PM-6:00PM, 350.01: Looking for Dust Scattering Light Echoes, B. Mills

Hyperwall

Tuesday & Thursday, 9:10AM-9:35AM NASA's Decadal Mission Concept Studies: HabEx, LUVOIR, Lynx, OST, D. Pooley & A. Vikhlinin (Lynx)

Wednesday, 9:10AM-9:20AM Revealing the Dawn of Black Holes with the Lynx X-ray Observatory, R. Hickox

Friday, 9:20AM-9:30AM Revealing the Invisible Drivers of Galaxy and Structure Formation and Evolution, J. Kollmeier

Science Driven Instrument Requirements



High-Definition X-ray Imager

Optimized for deep survey science

- Silicon sensors with ~ 0.3" pixels
- FOV ≥ 20'×20'
- Δ E ~ 100 eV over 0.1–10 keV band
- High frame rates to minimize pile-up.



Monolithic CMOS



Hybrid CMOS



Digital CCD with CMOS readout



Off-Plane Grating Array



Critical Angle Transmission – Grating Array

- Resolving power $\lambda/\Delta\lambda > 5000$
- Effective area > 4000 cm² covering X-ray emission and absorption lines of C, O, Mg, Ne, and Fe-L.

X-ray Grating Spectrometer

Detail outflow velocities and mass loss rates to provide information on matter and energy feedback in accreting galaxies.

Map the unobserved, large fraction of baryons that likely exists in the hot phase of the intergalactic medium.

Key Instrument Requirements

Lynx X-ray Microcalorimeter

- Main Array provides non-dispersive spectroscopy with Δ E < 3 eV over the 0.2–7 keV band and imaging with 1" pixels over a 5'×5' FOV
- Several subarrays are optimized for sub-arcsec imaging, 0.3 eV energy resolution, and coverage of 20'x20' FOV.

*Enhancement Main Array: O*ptimized to allow for higher count-rates, such as from AGN

High-Res Inner Array: Optimized to allow for higher count-rates, such as from AGN

Ultra-High-Res Array: Enables the study of turbulent line broadening around individual galaxies through the study of the highly ionized oxygen lines

Extended Array: Surveys over large regions of the sky for observations of the soft diffuse emission from extended galaxies, the outer regions of galaxy groups and clusters and also cosmic filaments



- RAY OBSERVATOR

Instrument Design Studies

MSFC Advanced Concept Office performed a comprehensive Design Study for HDXI and XGS (for both Off-Plane and Critical Angle Transmission Grating readouts)

Configuration Structures Mechanisms Thermal Power Electronics Cost



- RAY OBSERVATORY

GSFC contributed an Instrument Design Lab for LXM, including baseline and updated cost modeling



Optical Assembly Requirements

Angular resolution (on-axis)	0.5 arcsec HPD (or better)	
Effective area @ 1 keV	$\sim 2 \text{ m}^2$ (met with 3-m OD)	
Grasp, A*(FOV for HPD < 1 arcsec)	~600 m ² arcmin ²	
Wide FOV sub-arcsec Imaging	10 arcmin radius	

Simulated 2'×2' deep fields: JWST (Illustris-TNG light cone)

> Lynx (purple = AGNs, green=galaxies)

- RAY OBSERVATOR

Science Traceability:

Lynx will find the first supermassive black holes in the first galaxies detected by JWST, trace their growth from the seed phase, and shed light on how they subsequently co-evolve with the host galaxies. Needed sensitivities, 10^{-19} erg/s/cm², are ~ 200× below *ATHENA* confusion limit.

- Angular resolution < 1" (50% power diameter) will avoid source confusion and limit background
- An Effective area > 2m² and FOV > 10' in radius with arcsecond or better imaging will survey sufficient volume at z=10 in less than 25 Msec.

Athena (5" PSF, same area as Lynx)

Lynx Optics Technology Study

Lynx will use the Kepner-Tregoe trade process to select an optics technology for the Lynx Design Reference Mission and to establish feasibility for alternate viable technologies. – Supported by G. Blackwood, NASA Exoplanet Exploration Program

Study has been initiated for 3 Viable Optics Technologies



Adjustable Optics (Study Lead: P. Reid/SAO)



Si Meta-Shell Optics (Study Lead: W. Zhang/GSFC)



Full Shell Optics (Study Leads: K. Kilaru/USRA/MSFC, G. Pareschi/INAF/OAB)

Selection will be based on Science, Technical and Programmatic criteria (TBD)

The Lynx Optics Working Group will make a formal recommendation to STDT in Summer 2018!





Schattenburg – NASA PCOS SIG, 04/2016 - Modified

Mission Design Study



Launch Vehicle Most likely will need a Heavyclass launch vehicle (TBC)



Mission Lifetime: Baseline mission is 5 years, extendable for an additional 20 years based on consumables (still need to complete analysis on L2 radiation environment)

<u>Mission Operations</u>: Chandra-like. Lynx will have a primary science program combined with a general observer program

Lynx Observatory







2017 Additional Accomplishments

- RAY OBSERVATORY

- Awarded 5 Cooperative Agreement Notices (CANs) to industry partners to support payload design and programmatics
- Science Traceability Matrix was developed
- Interim Report first draft completed and reviewed by Red Team
- Major improvements to the Lynx simulation package (<u>http://hea-www.cfa.harvard.edu/~jzuhone/soxs/</u>)
 - Astrophysical backgrounds now include resolved point sources
 - A module to generate an X-ray light cone from a cosmological situation was added
 - Instrument specifications were added for imaging observations of ACIS-I and ACIS-S, Cycles 0 and 19, Hitomi/SXS, and AXIS.
 - The ability to generate gratings spectra for Lynx and Chandra ACIS-S/HETG was added.



Updated Website





LYNX SCIENCE



2018 Key Tasks



- Submit Interim Report to HQ (due 03/2018)
- Continue to strengthen science case and traceability to observatory architecture
- Improve fidelity of instruments, observatory, and mission concept design (MSFC and GSFC)
- Complete Optics Technology Study
- Complete Technology Roadmap for Optics and Instruments
- Complete Risk Assessment and Independent Costing for Lynx
- Carry out Informal CATE with Aerospace
- Initiate Final Report

*Next STDT F2F is 01/25/18-01/26/18 in Houston, TX



Thank you!

- Please visit the Lynx Display next to the Chandra Table and the Decadal Studies Table for more information
- Participation is open and welcome at any level. For more information and to signup to our News Distribution, visit our website at: <u>https://wwwastro.msfc.nasa.gov/lynx/</u>