J. Edmunson<sup>1</sup>, J. A. Gaskin<sup>2</sup>, Z. E. Gallegos<sup>3</sup> and the Miniaturized Variable Pressure Scanning Electron Microscope Science and Instrument Development Teams gineering Services and Science Capability Augmentation Contract, NASA Marshall Space Flight Center (MSFC), ST23/4201/228A, Huntsville AL 35812; <sup>3</sup>University of New Mexico, Department of Earth and Planetary Sciences, Albuquerque NM 87131

# Miniaturized Variable Pressure Scanning **Electron Microscope (MVP-SEM)**

- Funded by the NASA Planetary Instrument Concepts for the Advancement of Solar System Observations (PICASSO) Opportunity.
- Designed to utilize the CO<sub>2</sub>-rich atmosphere of Mars as an imaging medium, preventing charging of a sample.
- Developed under a partnership between NASA's Marshall Space Flight Center, the Jet Propulsion Laboratory (JPL), Applied Physics Technologies (AP-Tech), and Creare LLC.
- Capabilities will include a resolution of 100nm or better in both backscattered electron and secondary electron imaging, micron-scale calibrated energy dispersive spectroscopy (EDS) for geochemical analysis.

## **Instrument Development Status**

- LaB<sub>6</sub> and Schottky emitters tested for suitability and lifespan in a  $CO_2$ -rich environment; LaB<sub>6</sub> selected.
- Emitter mounts vibration tested to gauge resistance to launch and landing loads; both LaB<sub>6</sub> and Schottky emitters withstood 2 minutes at a maximum 14.11 G-RMS (~20s ramp) on x, y, and z axes with minimal resulting offsets that do not exceed the ability of the electron optics to accommodate.
- Optics modeling resulted in a working design for a field of view of 1mm square at an environmental distance of 2mm.
- •Electron gun and column assembly has been built and tested in a vacuum; testing of electronics and software is ongoing (Figures 1-3).
- Sample chamber design uses 3D printed components (Figure 4) and utilizes a custom sample rod to insert samples; pressure inside the chamber is controlled to allow for imaging and EDS under optimum conditions.

#### **Future Work**

- Testing of the prototype instrument, with the integrated sample chamber, will be completed in a Mars Environment Chamber at JPL during the late spring/early summer of 2018.
- The MVP-SEM team will respond to the 2018 Maturation of Instruments for Solar System Exploration Call in which the status of the prototype PICASSO instrument will be presented, as well as a plan to increase the MVP-SEM's technology readiness level.
- The Science Team will publish details of the Concept of Operations, Science Requirements, and other information relevant to MVP-SEM science.

# The Science Case for a Scanning Electron Microscope on Mars



Figure 1: MVP-SEM prototype column housing. Further miniaturization is possible with custom flanges and fittings. Image provided by AP-Tech

Figure 2: Test image of tin spheres without noise reduction and **Pressure Limiting Aperture 1**. Image provided by AP-Tech

> Figure 3: Prototype instrument image of a gold and carbon grid. mage provided by AP-Tech

# **Understanding of Mars**



Figure 4: 3D printed base of sample chamber Image provided by JPL

### Acknowledgements

- The Science and Instrument Development Teams would like to thank the NASA Science Mission Directorate and the PICASSO Program for funding the instrument development.
- The authors of this poster would like to thank the numerous participants and contributions of the Instrument Development and Science Teams.

**Origin and Evolution** 

Petrology

Human Interests

Habitability



**Geochronology:** Determine the relative petrologic sequence of events on a microscopic scale using imaging and EDS.

**Igneous Rocks:** Analyze existing phases to calculate magmatic conditions and evaluate microscopic texture, zoning, and resorbed minerals.

**Metamorphism:** Observe fabrics, textures, and specific minerals indicative of metamorphic processes on a microscopic scale.

**Geochemistry:** Analyze minerals and amorphous phases using EDS (EDS analysis does not rely on crystal structure), as well as resolve flow banding compositional differences.

**Mineralogy:** Obtain modal mineralogy via backscattered electron imaging and element mapping via EDS; the mineralogy of small phases such as martian dust will also be determined.

**Resource Prospecting and Utilization:** Identify water-ice and hydrated minerals, ores, metal/silicate phases, etc. to provide feedstock for applications such as in-space manufacturing, agriculture, life support consumables, propellant, and other in-situ resource utilization needs.

**Alteration:** Investigate evidence of weathering and metasomatism, and characterize alteration (e.g., surface rinds on individual grains).

Sediments: Detrital mineralogical analysis, characterization of cementitious materials (e.g., salts), and visualization of oxidation/reduction reaction products and depositional fabrics.

Environments Favorable for Life: Positively identify clays and other weathering products that are indicative of conditions favorable for life in the past or present, and assess the habitability of Mars for human life.

**Biosignatures:** Study small-scale structures and geochemical signatures indicative of life, including fossil morphologies.

**Toxicity of the Environment:** Identify fine particles that could potentially enter a human habitat, recognize materials that are toxic to humans, and characterize any organisms that could impact human operation on the martian surface.

**Potential Existing Life:** Resolve micron-scale bacteria or other organisms that could be toxic to humans should they exist on Mars, even in the presence of larger grains as an SEM enjoys a greater depth of field than optical microscopes.