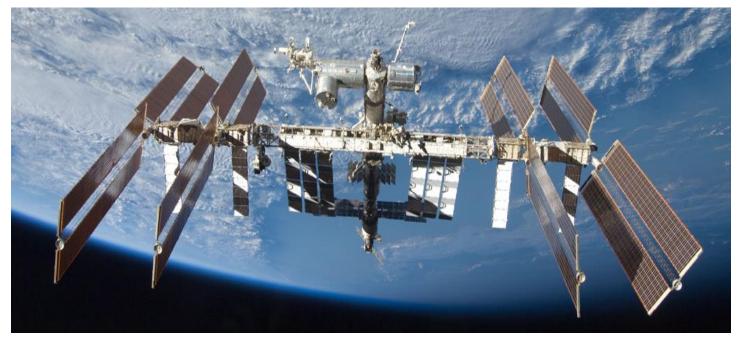
#### **Increment 57 / 58 Science Symposium**



## Advanced Colloids Experiment (Temperature controlled) – ACE-T2 PI: Prof. P. Schall– University of Amsterdam







Dr. William V. Meyer (a.k.a. Bill Meyer) ACE NASA Project Scientist

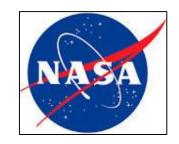
USRA at NASA GRC, Tel: (216) 433-5011, Email: William.V.Meyer@NASA.Gov ACE NASA Project Manager: Ron Sicker, Tel.: (216) 433-6498 ZIN-Technologies Project Lead: Kelly Bailey, Tel: (440) 625-2251 ZIN-Technologies Science Lead: John Eustace, Tel: (440) 625-2244



May 24, 2018

## ASSEMBLY OF COLLOIDAL SUPERSTRUCTURES BY CRITICAL CASIMIR FORCES IN MICROGRAVITY

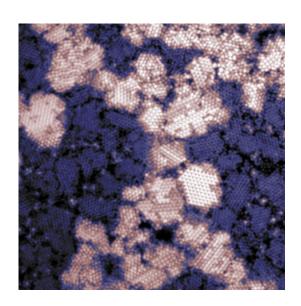




#### Piet Swinkels, Gerard Wegdam, Peter Schall

University of Amsterdam –





## Science Team

### **Amsterdam**



Piet Swinkels



Simon Stuij



Prof. G. Wegdam



Prof. P. Schall

## <u>Milan</u>



Dr. M. Potenza

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https://iop.fnwi.uva.nl/scm/

## ISS Increments 57 and 58 Science Symposium

Advanced Colloids Experiment (Temperature controlled) – ACE-T2 Prof. P. Schall– University of Amsterdam

- Science Background and Hypothesis
- Investigation goals and objectives
- Measurement approach
- Importance and reason for ISS
- Expected results and how they will advance the field
- Earth benefits/spin-off applications

## Science Background and Hypothesis – 1/5

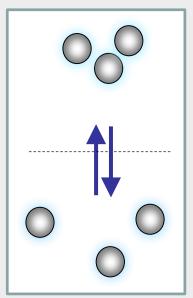
#### **The Critical Casimir effect**

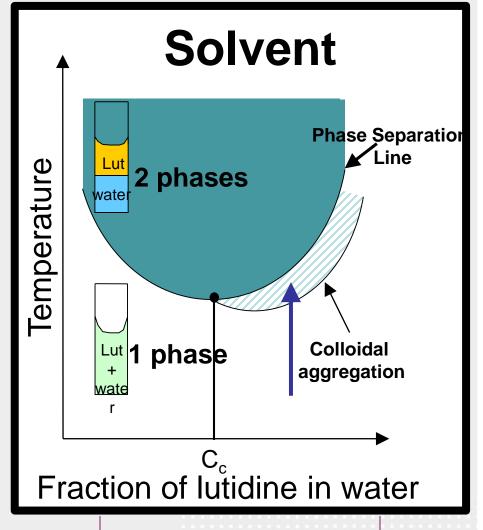


## Science Background and Hypothesis – 2/5

#### The Critical Casimir effect

Concentration fluctuations will become larger as you increase the temperature towards the phase separation line. As the fluctuations grow bigger, the attraction will become stronger and longer ranged and colloids will start to aggregate. See "Notes page" – below.





## Science Background and Hypothesis – 3/5

#### **Particle interactions**

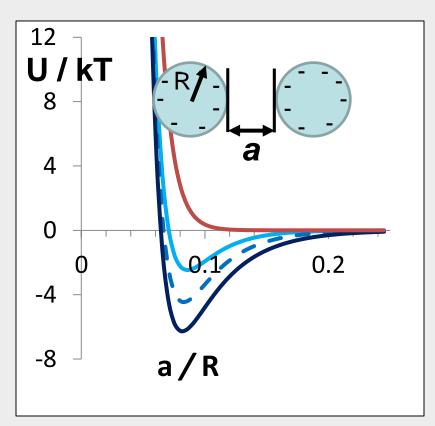
#### Electrostatic repulsion

$$V_{el}(a) \propto \lambda_d^2 e^{-a/\lambda_d}$$



#### Critical Casimir attraction

$$V_{casimir}(a) \propto -\frac{1}{\xi} e^{-a/\xi}$$



where  $\xi$  is the bulk correlation length of the solvent

Schall et al., PRL **103**, 156101 (2009)

## Science Background and Hypothesis – 4/5

### **Tuned Interactions: A/B particle mixtures**

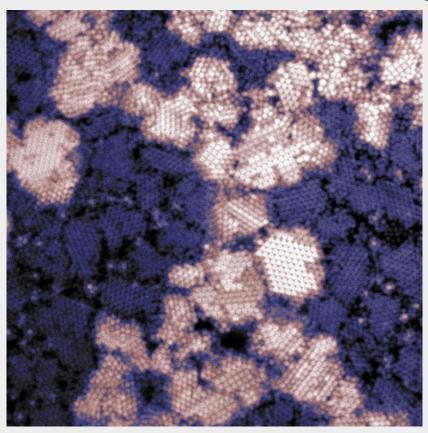




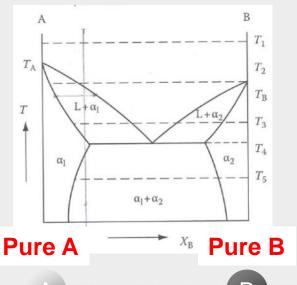
Interactions  $u_{AA} > u_{AB} > u_{BB}$ 

## Science Background and Hypothesis – 5/5

### **Tuned Interactions: A/B particle mixtures**



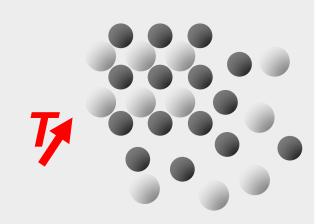
Model intermetallic phases, e.g. Eutectic mixtures



## Investigation goals and objectives – 1/3

### **Study Critical Casimir assembly** Use temperature control

- → Tune attractions
- → Follow structure formation
- → "Complex crystallisation kinetics"



**Vary Temperature** → Vary attraction strength

Vary Rate of change → Eqilibrium vs. out-of-equilibrium

**Reverse Temperature** → Repeat Experiment

## Investigation goals and objectives – 2/3

Investigation of assembly of complex structures from micron-scale colloidal particles interacting via tunable attractive interactions.

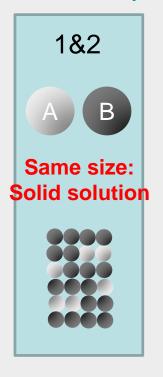
Samples contain suspensions of particle type A and B (trifluoroethyl methacrylate, FEMA colloidal particles) in binary solvents of water ( $H_2O$ , 68%mass) and lutidine (32%mass), that upon nearing the critical solvent temperature ( $T_c\sim32^\circ\text{C}$ ) give rise to critical Casimir interactions between the particles.

Regulating temperature enables control of particle A, B interactions.

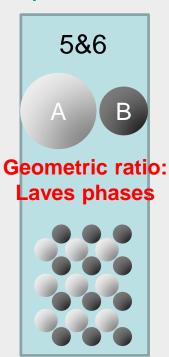
The goal is to understand how complex interactions lead to complex structures, and to understand the dynamics of growth of these structures.

## Investigation goals and objectives – 3/3

#### Proposed Experiment - Samples







- Particles in solvents of Lutidine and water + sugar for index match
- Particle volume fraction ~ 10%
- Small amount of salt (NaCl, 2mM)
- Sample details provided in back-up slides

## Measurement approach – 1/5

Particle A, B mixtures have been recently investigated by us on ground. Suspended in binary solvent, at temperatures around 28-30°C, they show complex liquid and crystal phases, reminding of complex intermetallic phases.

Structures depend on the particle number and size ratio, as well as on temperature setting their respective attractive strength  $u_{AA}$ ,  $u_{AB}$ , and  $u_{BB}$ .

These structures and their formation kinetics will be directly followed with the LMM microscope on the ISS.



National Aeronautics and Space Administration John H. Glenn Research Center at Lewis Field



Using the Light Microscopy Module (microgravity microscope) aboard the International Space Station (ISS).

## Measurement approach – 2/5

#### **Operational Requirements:**

Overall, one experiment (capillary cell) is conducted every two days, depending on the number of *z*-layers and speed of growth of the structures. The experiment is repeated until all capillaries are tested. Sample modules are switched if air bubbles are too big. Use the rest of the week to analyze data, re-write scripts, and adjust parameters. The number of capillaries per experiment is limited by data bottlenecks on IPSU (internet protocol setup utilities) and IOP (input/output operations per second). We will need to check *XYZ* position repeatability and oil availability. More specifically, we will need to return to the same Region of Interest (ROI), e.g., single particle crystallite – or to maintain *XYZ* coordinates during an experiment. This requirement implies maintaining one capillary posiStion. If this requirement takes an excessive amount of time, we will need to find a solution.

## Operational Protocols: Descriptive overview of the investigation on orbit procedures. Experiment Steps:

The general experimental protocol includes imaging the dynamics of colloidal assemblies in each capillary. We will inspect all samples before mixing the colloid using the *in situ* mixer. Next, define XY offsets. Locate and image particles, adjusting the camera parameters using the 2.5x objective and Texas red (or FITC) filter (fluorophore = Cy3MM, Harvard Fluorophore, Cy3-like). Survey capillary at 2.5x, scanning in the X direction over a range of at least 10 millimeters. Determine bubble locations and possible primary / secondary Regions of Interest (ROI). If the 2.5x objective is difficult to switch in and out with the 63x air objective, then find ROI capillary cells before using 100x oil objective. If both ends of the capillary can be used (no big air bubble), then apply immersion oil to half of the capillary to allow imaging with 100x oil *and* 63x air objective without astronaut intervention. Select primary locations away from stir bar or bubble.

## Measurement approach – 3/5

#### **Operational Protocols [Continued]:**

#### A. Determine aggregation and solvent phase separation temperature

Use low magnification lens to determine particle aggregation temperature,  $T_{AA}$ , and solvent phase transition temperature  $T_c$ . To do so, survey and record best Z-depth at each primary test location. Ramp temperature slowly from ~27°C to ~31°C during 4h (ramp rate 1°C/hour) while recording microscope video. Note aggregation temperature  $T_a = T_{AA}$  when particle aggregates start to appear, and phase separation temperature  $T_c$ , where the solvent separates into two components.

#### B. Confocal video, time series

Go back to room temperature, and stir for 30min. After that, switch in 63x air objective. Increase temperature to  $T_a$ -0.3°C. Record microscope video (confocal imaging), while slowly ramping temperature to  $T_a$ +0.3°C during 6 hours (ramp rate 0.1°C/hour). These direct particle-scale observations should allow determining the colloidal liquid condensation and crystallization temperatures  $T_{liq}$  and  $T_{cryst}$ . If possible, record confocal z-stacks of the characteristic liquid phase (at T> $T_{liq}$ ) and crystal phase (at T> $T_{cryst}$ ). Each z-stack should consist of ~200-300 images, 0.1 $\mu$ m apart, starting from ~10-20 $\mu$ m into the capillary (above the glass wall). The Z-scanning rate should be set as fast as possible, potentially 10 frames per second. No pixel binning, 8 bits per pixel (highest supported), full frame images. After z-stack is taken, continue recording microscope video (at same location and depth as before)

[Continued]

## Measurement approach – 4/5

#### **Operational Protocols [Continued]:**

#### C. Confocal stacks

Go back to room temperature, and stir for 30min. Use 63x air objective.

After aggregation temperatures have been refined in (B.), go to  $T_{\rm liq}$  and keep the temperature constant. Record confocal stack time series (each z-stack consisting of ~200-300 images, 0.1 µm apart, starting from ~10-20 µm into the capillary, taken with high scan rate). Record 1 z-stack per 2 min, in total 50 z-stacks over 100 minutes. After this time series, when colloidal liquid has formed, go to  $T_{\rm cryst}$  and repeat the same z-stack time series, again taking one z-stack every ~2min, 50 –stacks total over 100min.

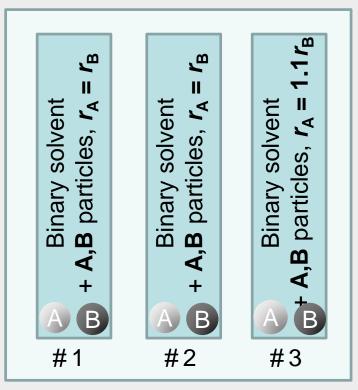
#### D. Imaging with 100x oil

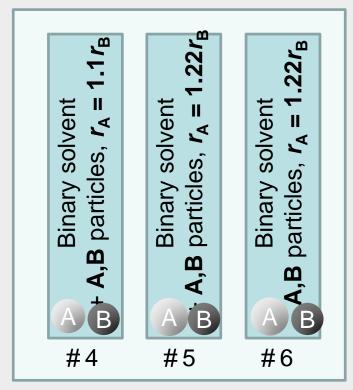
For better image quality, imaging with 100x oil objective is preferred. However, this will introduce a heat gradient as heat is sucked away from the imaging location into the objective. The heat gradient may also induce particle flows, which are detrimental for the study of particle assembly. To find  $T_{\text{liq}}$  and  $T_{\text{cryst}}$  with the 100x air objective, start from  $T_{\text{a}}$  and repeat ramping procedure in (A), ramping to ~31°C during 4h (ramp rate 1°C/hour). After aggregation temperatures have been found, refine with a slower ramp rate 0.1°C/hour. Cool down and stir for 30min. Then, go to refined  $T_{\text{liq}}$  and record confocal z-stack time series as described in (C). After that, go to  $T_{\text{cryst}}$ , and again record confocal z-stack time series.

Repeat A. – D. for all sample capillaries.

## Measurement approach – 5/5

#### **Pictorial summary**





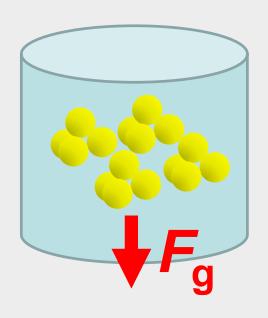
Module 1

Module 2

FEMA particles (trifluoroethyl methacrylate) in water (68%wt) and lutidine (32%wt)

## Importance and reason for ISS – 1/2

#### 1. Growth of large structures



$$F_{g} = \Delta \rho g V$$

where:

 $F_{\alpha}$  [force of gravity]

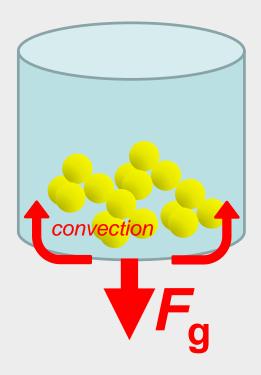
 $\Delta \rho = \rho_{\text{Colloid}} - \rho_{\text{Solvent}} \text{ [density]}$ 

g [gravity]

V [Volume]

## Importance and reason for ISS – 2/2

#### 2. Temperature is the control parameter

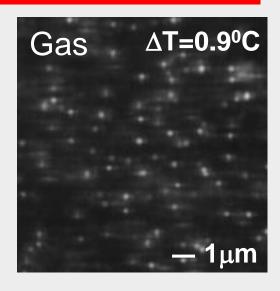


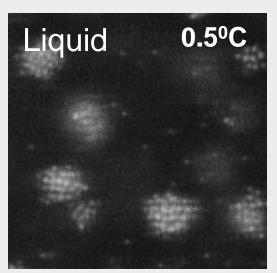
- Change Temperature
   Δρ changes
- Temperature → Convection disrupts growth

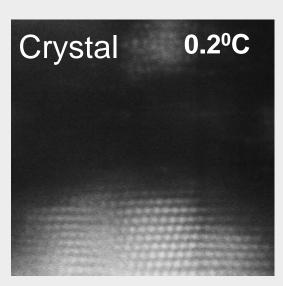
These structures and the growth process are strongly affected by sedimentation and convection. The goal of the ACE-T-2 experiment is to study the diffusion-limited growth of these complex structures without the disturbing effects of gravity.

# Expected results and how they will advance the field

#### 1-Component system







Microscopy: Ground-based work

# Expected results and how they will advance the field

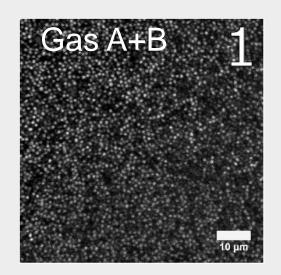
2-Component system

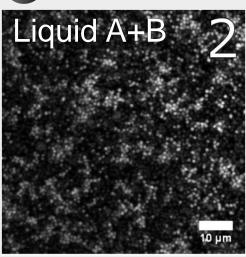


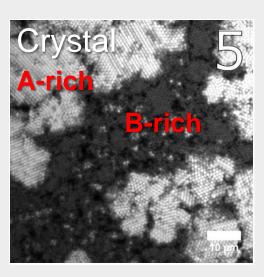
Particle A



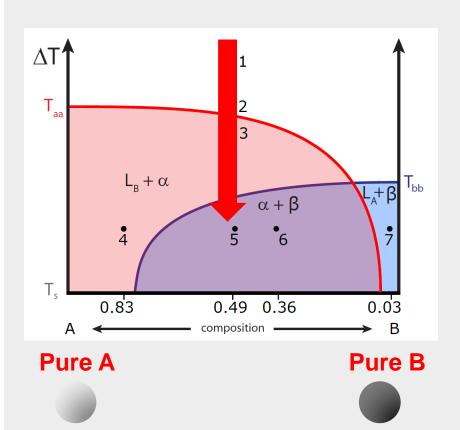
Particle B







Microscopy: Ground-based work



#### **Liquid phases**

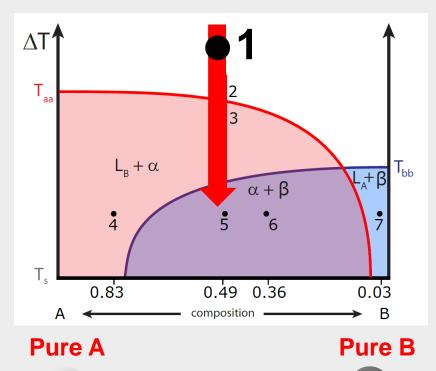
L<sub>A</sub>: liquid, A-rich L<sub>B</sub>: liquid B-rich

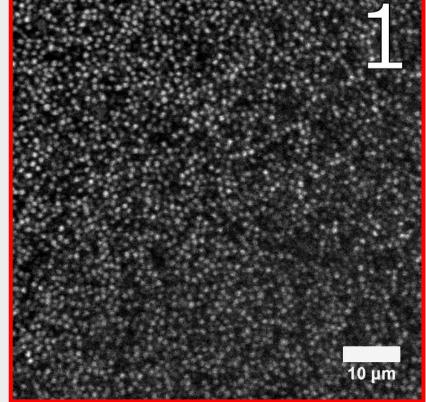
#### **Crystal phases**

α: crystal, A-rich β: crystal, B-rich

#### **Crystallization temperatures**

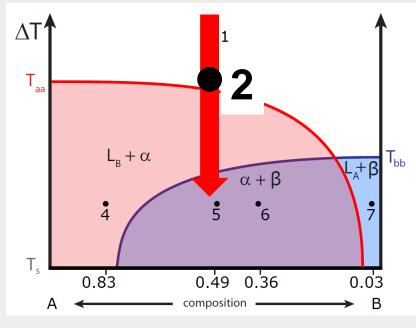
T<sub>aa</sub>: crystallizaton of A T<sub>bb</sub>: crystallizaton of B







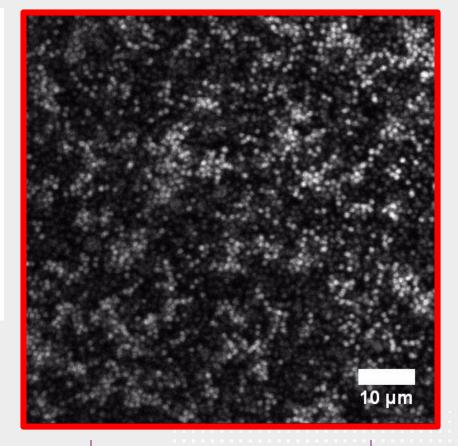




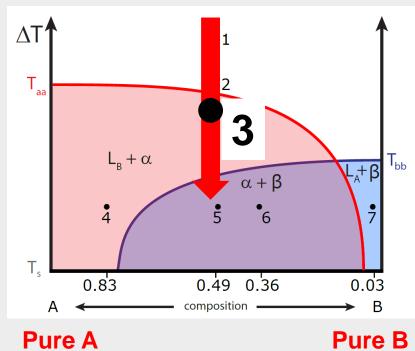
**Pure A** 

**Pure B** 



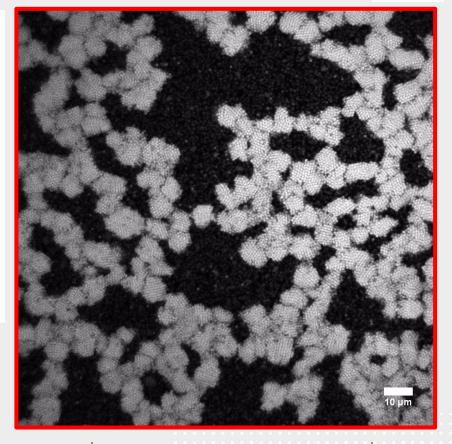






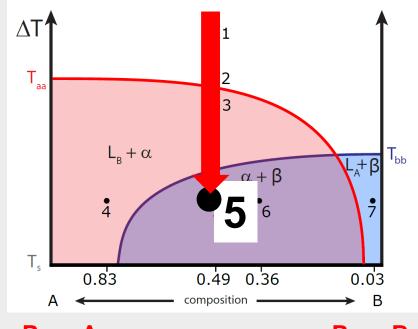








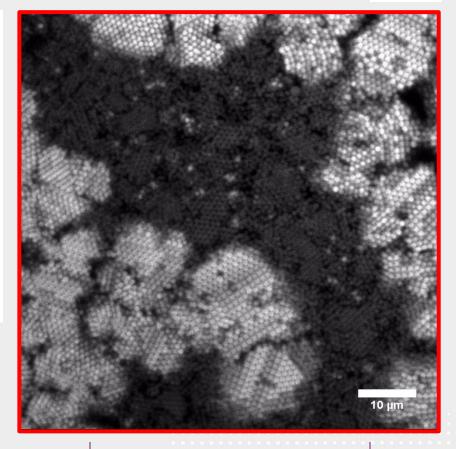






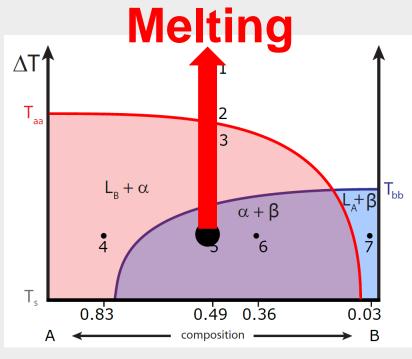








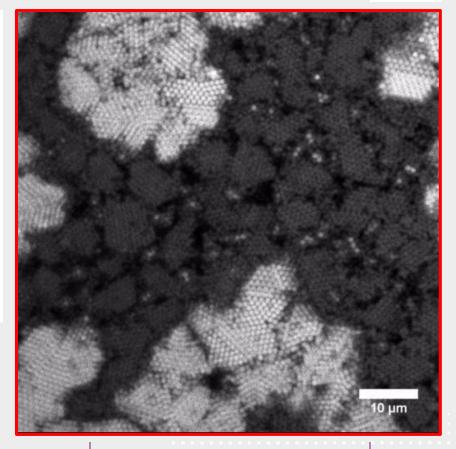




**Pure A** 

**Pure B** 





## Expected results and how they will advance the field

Space-based (microgravity) work – 3D

- What do we expect to see?
  - Observation of temperature controlled assembly
  - Structures as function of A/B mix and quench rate
  - Dynamics of "eutectic" crystal formation

## Earth benefits / spin-off applications

Understanding self-assembly processes will teach ways to grow complex nanostructured materials. Furthermore, these experiments will provide fundamental insight into the nature of self-assembly, the relation between complex interactions and complex equilibrium and non-equilibrium structures, and the dynamics of both equilibrium and non-equilibrium growth processes.

#### **Collaborators**

- Tom Kodger (Wageningen)
- Peter Bolhuis (Amsterdam, Simulations)
- S. Dietrich (MPI Stuttgart)
- Ania Maciolek (MPI Stuttgart)

## The International Space Station (ISS) Light Microscopy Module (LMM)

**ACE-T2 BACKUP SLIDES** 

### **Mission Success Criteria for ACE-T2**

Complete success is the achievement of all of the science requirements. This means that there will be sufficient information to provide a crosscheck of all data and calculated factors. Processing, manipulation and characterization of the samples in micro-gravity are as important as the measurements during the experiments themselves. *e.g.*, sample homogenization is essential to conduct of any of the flight experiments. This allows for the homogenization of the crystallites or any structures formed from phase separation or gelation that have occurred in 1g before launch, and provides a proper starting point in micro-g.

Success Level		Accomplishment
•	Minimum Success	<ul> <li>Homogenize and observe at least 4 of the 6 capillary samples. This includes observing the time evolution with confocal microscopy (using bright field and fluorescent microscopy as needed) imaging for several days to weeks, depending on rates of change determined in real-time as data is downlinked to Earth (these cannot be predicted accurately ahead of time in the 1g environment).</li> <li>Have sufficient data (both in terms of frequency and duration) from microscopy of sufficient quality to observe, characterize and quantify the rates of growth of structures formed as a result of the physical process of interest in microgravity, including but not limited to self-assembly, the relation between complex interactions and complex equilibrium and non-equilibrium structures, and the dynamics of both equilibrium and non-equilibrium growth processes.</li> <li>We hope that these processes will generate new structures formed in microgravity, that may direct further earthbound studies and inspire new directions for materials synthesis and fundamental physics understanding.</li> </ul>
•	Significant Success	Accomplish the above for 5 of the 6 different types fluid samples launched.
•	Complete Success	<ul> <li>Accomplish the above for all launched samples, with multiple runs to repeat the experiment and assess reproducibility.</li> </ul>

## **ACE-T2** samples

Description: UvA ACE-T2 Flight Sample #1

P/N: UvA\_T2\_Flight\_Sample1

Total volume: 1 ml

Suspension of 10%vol FEMA particles dyed with Cy3MM in a binary solvent of 32% Lutidine and 68%  $H_2O$ , with 40%wt sucrose in the  $H_2O$  fraction + 2 mM salt (KCl). (binary particle mixture A, A' (same radius) with number ratio 20/80%)

Description: UvA ACE-T2 Flight Sample #2

P/N: UvA\_T2\_Flight\_Sample2

Total volume: 1 ml

Suspension of 10%vol FEMA particles dyed with Cy3MM in a binary solvent of 32% Lutidine and 68%  $H_2O$ , with 40%wt sucrose in the  $H_2O$  fraction + 2 mM salt (KCl). (binary particle mixture A, A' (same radius) with number ratio 50/50%)

Description: UvA ACE-T2 Flight Sample #3

P/N: UvA\_T2\_Flight\_Sample3

Total volume: 1 ml

Suspension of 10%vol FEMA particles dyed with Cy3MM in a binary solvent of 32% Lutidine and 68% H2O, with 40%wt sucrose in the H2O fraction + 2 mM salt (KCl). (binary particle mixture A, A' (same radius) with number ratio 80/20%)

Description: UvA ACE-T2 Flight Sample #4

P/N: UvA\_T2\_Flight\_Sample4

Total volume: 1 ml

Suspension of 10%vol FEMA particles dyed with Cy3MM in a binary solvent of 32% Lutidine and 68% H2O, with 40%wt sucrose in the H2O fraction + 2 mM salt (KCl). (binary particle mixture A, B ( $R_B$ =1.2  $R_A$ ) with number ratio 20/80%)

Description: UvA ACE-T2 Flight Sample #5

P/N: UvA\_T2\_Flight\_Sample5

Total volume: 1 ml

Suspension of 10%vol FEMA particles dyed with Cy3MM in a binary solvent of 32% Lutidine and 68% H2O, with 40%wt sucrose in the H2O fraction + 2 mM salt (KCl). (binary particle mixture A, B ( $R_R$ =1.2  $R_A$ ) with number ratio 50/50%)

Description: UvA ACE-T2 Flight Sample #6

P/N: UvA\_T2\_Flight\_Sample6

Total volume: 1 ml

Suspension of 10%vol FEMA particles dyed with Cy3MM in a binary solvent of 32% Lutidine and 68% H2O, with 40%wt sucrose in the H2O fraction + 2 mM salt (KCl). (binary particle mixture A, B ( $R_a$ =1.2  $R_A$ ) with number ratio 80/20%)