

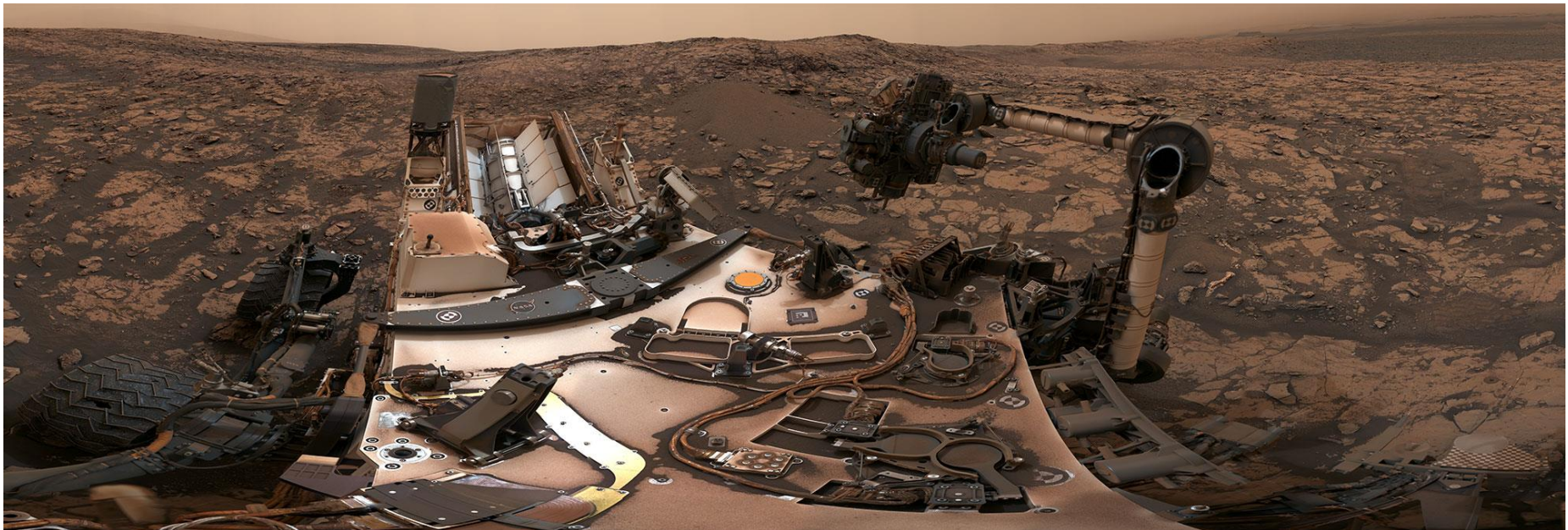
2018 Conference on Advanced Power Systems for Deep Space
October 21-24
Pasadena, California

Designing Spacecraft Power Systems to Solve Planetary Protection Challenges

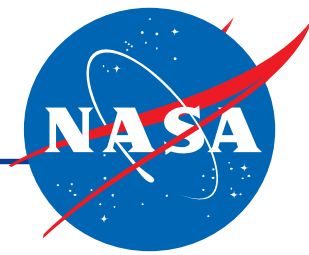
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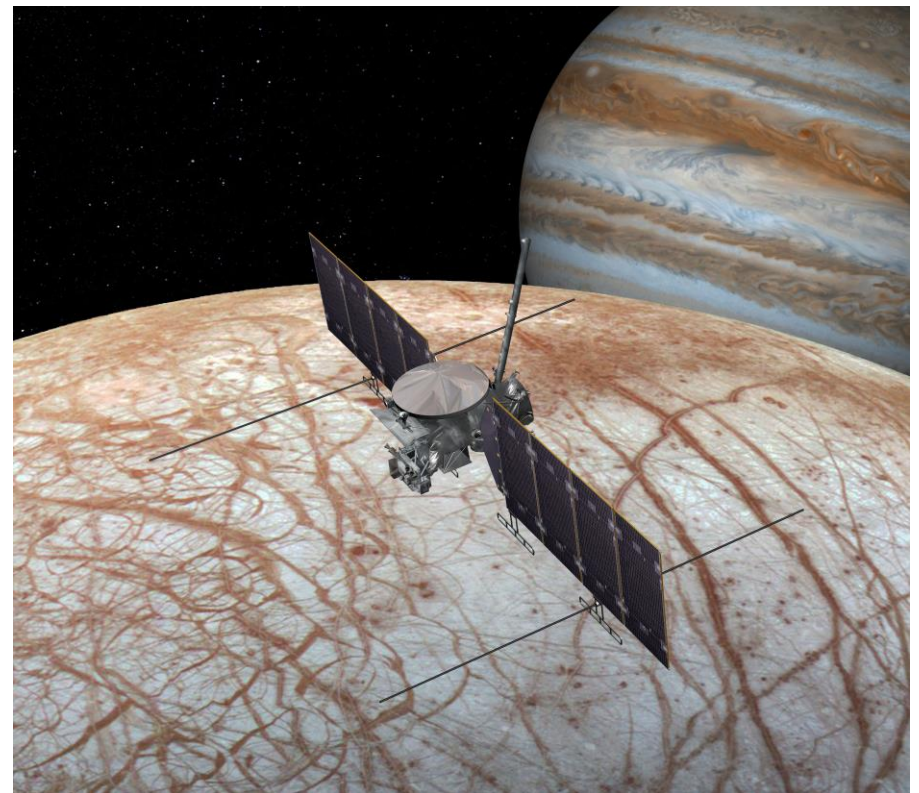
August 9, 2018 NASA's Curiosity rover on Mars, a 360-degree panorama of its current location on Vera Rubin Ridge



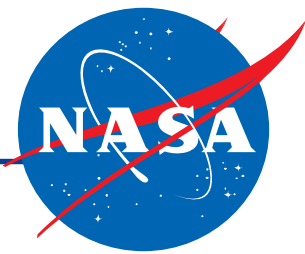
Planetary protection requirements ensure *valid* and *safe* scientific exploration for extraterrestrial life

NASA Objectives during robotic missions beyond Earth:

- (*valid*) Avoid forward contamination of other worlds by terrestrial organisms carried on spacecraft.
- (*safe*) Prevent backward contamination of Earth by extraterrestrial life or bioactive molecules in samples returned for scientific study.



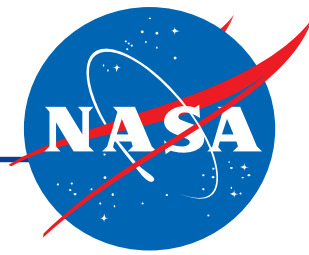
Artist vision of the Clipper spacecraft orbiting Europa which is an icy moon of Jupiter.



Enduring Authority of the Outer Space Treaty

The Outer Space Treaty of 1967:

- Proposed to the UN in 1966
- Signed at Washington, London, Moscow on January 27, 1967
- Entered into force October 10, 1967
- **Article IX.** States Parties to the Treaty shall pursue studies of outer space, including the Moon and other celestial bodies, and conduct exploration of them so as **to avoid their harmful contamination and also adverse changes in the environment of the Earth resulting from the introduction of extraterrestrial matter and, where necessary, shall adopt appropriate measures for this purpose.**
- **Article VI.** States Parties to the Treaty shall bear international responsibility for national activities in outer space, including the moon and other celestial bodies, **whether such activities are carried on by governmental agencies or by non-governmental entities**, and for assuring that national activities are carried out in conformity with the provisions set forth in the present Treaty.

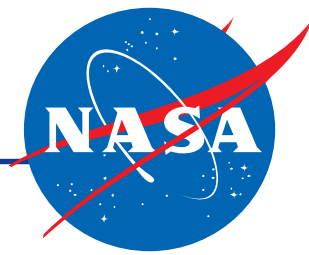


Relationship to COSPAR

NASA recognizes that only the 1967 Outer Space Treaty (OST) sets forth legal planetary protection requirements. It is however NASA policy to follow the Committee on Space Research (COSPAR) Planetary Protection Policy of 2017 (“COSPAR Planetary Protection Policy”), as amended. In carrying out this policy, NASA is guided by advice and recommendations by the National Academies of Sciences, Engineering, and Medicine as well as other independent advisory panels.



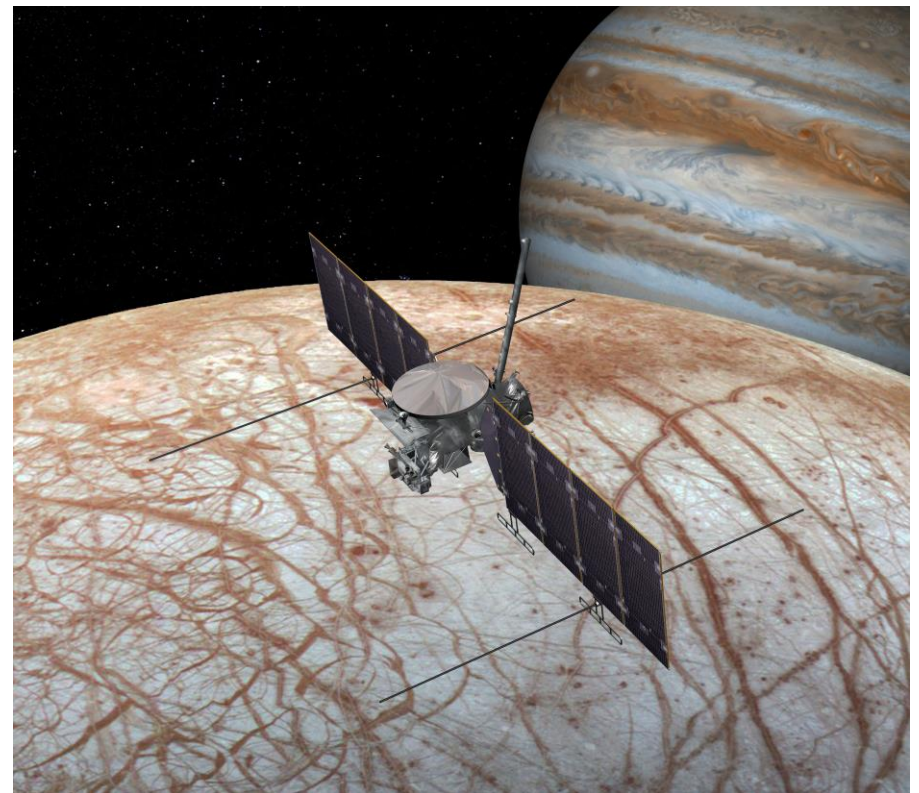
*Expanding the knowledge frontier of space
for the benefit of humankind*



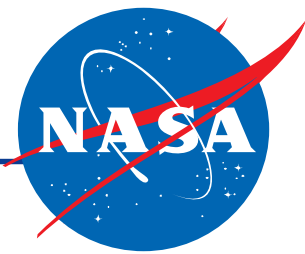
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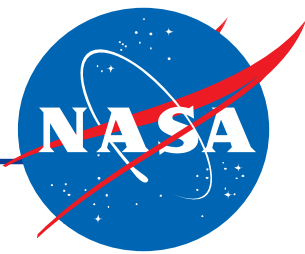
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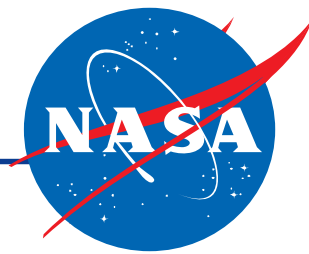


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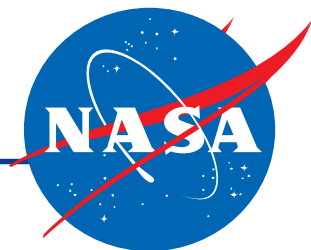


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Planetary Protection Mission Categories (NASA/ESA/COSPAR Policy)

Types of Planetary Bodies	Mission Type	Mission Category
Not of direct interest for understanding the process of chemical evolution. No protection of such planets is warranted.	Any	I
Of significant interest relative to the process of chemical evolution, but only a remote chance that contamination by spacecraft could jeopardize future exploration. Documentation is required.	Any	II
Of significant interest relative to the process of chemical evolution, and/or the origin of life or for which scientific opinion provides a significant chance of contamination which could jeopardize a future biological experiment. Substantial documentation and mitigation is required.	Flyby, Orbiter <i>Mars, Europa, Enceladus</i>	III
Same as Cat III	Lander, Probe <i>Mars, Europa, Enceladus</i>	IV <i>IVa, IVb, IVc</i> Mars
Any solar system body. Unrestricted applies only to bodies deemed by scientific opinion to have no indigenous life forms.	Earth Return <i>Restricted or Unrestricted</i>	V



Categorization for Current New Frontiers Missions

New Horizons

Pluto-Kuiper Belt

Category II



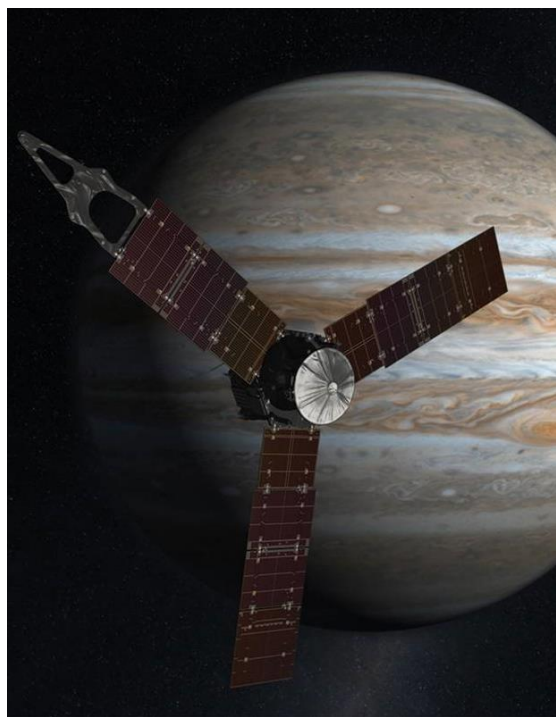
*January 2006 Launch
Flyby Pluto July 2015
Kuiper image August 2018*

JUNO

JUNO

Jupiter Polar Orbiter

Category III



*August 2011 Launch
Arrival Jupiter July 2016
13th flyby July 2018*

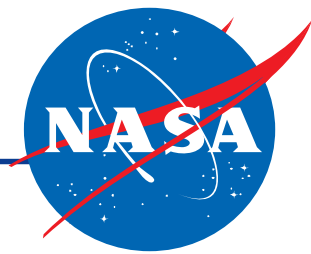
OSIRIS-REx

Asteroid Sample Return

Category V Unrestricted



*September 2016 Launch
Final approach Bennu 2018*



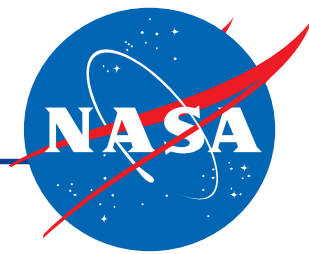
Planetary Protection Provisions for Robotic Extraterrestrial Missions

NPR 8020.12D

Effective Date: April 20, 2011

Expiration Date: December 20, 2018

5.1		Numerical Implementation Limits for Forward Contamination Calculations not Otherwise Specified
	5.1.1	To the degree that numerical limits are required to support the overall policy objectives of this document, and except where numerical requirements are otherwise specified, the limit to be used is that the probability that a planetary body will be contaminated during the period of biological exploration shall be no more than 1×10^{-3} . No specific format for probability of contamination calculations is specified.
	5.1.2	The period of biological exploration shall extend at least 50 years after a PP Category III or IV mission arrives at its protected target and no longer than the time point after which no organisms remain viable on the spacecraft.
	5.1.3	For all launch vehicle elements leaving Earth's orbit, the probability of impacting Mars shall be less than 1×10^{-4} for a period of 50 years. The probability of impact assessment should be provided in the Planetary Protection Plan.
	5.1.4	For all spacecraft crossing Mars orbit en route to other targets, the probability of impacting Mars shall be less than 1×10^{-2} for a period of 50 years. The probability of impact assessment should be provided in the Planetary Protection Plan.
	5.1.5	In the context of missions to icy satellites, "contamination" is defined as the introduction of a single viable terrestrial microorganism into a liquid-water environment.



NPR 8020.12D

Effective Date: April 20, 2011

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Planetary Protection Provisions for Robotic Extraterrestrial Missions

5.3		PP Category-specific Requirements for Mars	
5.3.2			PP Category IV for Mars is subdivided into IVa, IVb, and IVc. Missions shall comply with requirements appropriate to the subcategory they have been assigned. Requirements for missions carrying life detection instruments that access special regions will include a combination of those listed under IVb and IVc, as determined on a mission-by-mission basis.
	5.3.2.1		PP Category IVa. Lander systems not carrying instruments for the investigations of extant Martian life shall:
		a	Be restricted to a surface biological burden level of 3×10^5 spores in total and an average of 300 spores per square meter of exposed external and internal spacecraft surfaces.
		b	Provide an assessment of Entry, Descent, and Landing (EDL) expected performance against environmental and other design cases, identifying included and excluded factors, and, to the extent available, quantitative assessments of confidence levels.
	5.3.2.2		PP Category IVb. Lander systems designed to investigate extant Martian life shall comply with all of the requirements of PP Category IVa and also with one of the following requirements:
	either	a	The entire landed system is restricted to a surface biological burden level of 30 spores (see 5.3.2.4) or to levels of biological burden reduction driven by the nature and sensitivity of the particular life-detection experiments, whichever are more stringent, and protected from recontamination.
	or	b	The entire landed system is restricted to a surface biological burden level of 30 spores (see 5.3.2.4) or to levels of biological burden reduction driven by the nature and sensitivity of the particular life-detection experiments, whichever are more stringent, and protected from recontamination.

Cleanroom Garments and Sample Acquisition

Garment Zones



Booties, hairnet and mask followed by hood, bunny, suit, boots, gloves and wrist tape.

Polyester Wipes



Spacecraft hardware and transfer surfaces are sampled by wiping or swabbing.

Extraction of Microbes



Polyester wipe inserted into sterile glass bottle with deionized water for extraction.

NASA Standard Spore Assay



Sterilization of glassware and growth media in autoclave



Heat shock of the extracted samples at 80°C for 15 minutes



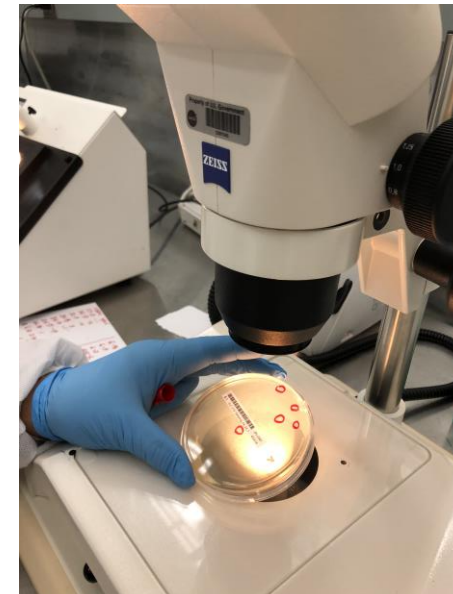
Quenching shocked samples in ice bath



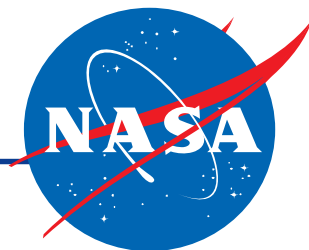
Quantitative aliquots plated on Trypticase Soy Agar



Visual enumeration of colony forming units after incubation at 32°C for 24, 48, and 72 hours



Bar coding enables reliable tracking of bioburden data

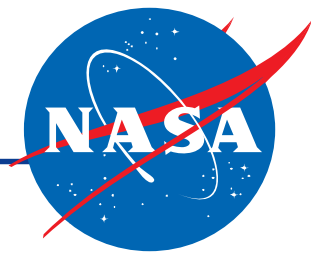


Lethality is determined by treatment and growth of ordinary terrestrial organisms found on spacecraft

Table 1. Experimental D-values of *Bacillus atrophaeus* spores under controlled and ambient humidity

Bacillus Atrophaeus						
	Ambient humidity			Controlled humidity		
Temp. (C)	D-value (min.)	St. Dev. (min.)	3 sigma D-value (min)	D-value (min.)	St. Dev. (min.)	3 sigma D-value (min)
115	92.6	5.6	115	27.2	0.64	29.65
125	33.9	0.8	37	19.7	0.56	21.9
135	9.56	0.30	10.8	9.09	0.42	10.74
150	1.37	0.068	1.65	1.66	0.057	1.89
170	0.35	0.02	0.41	0.40	0.031	0.531

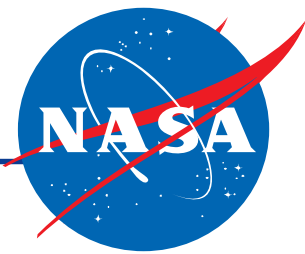
D-value refers to a decadal (factor of 10) decrease in viable spores



Lethality is determined by treatment and growth of very heat resistant terrestrial organisms

Table 3. Experimental D-values of *Bacillus* ATCC 29669 spores

Bacillus ATCC 29669 Spores							
		Ambient Humidity			Controlled Humidity		
Temp.(°C)	Units	D-value	St. dev.	3 sigma D-value	D-value	St. dev.	3 sigma D-value
115^a	Days	3.35	NA	NA	1.92	NA	NA
125	hr.	18.8	3.79	22.6	10.10	0.26	11.0
150	Min	66.4	3.96	80.8	40.0	1.14	44.1
170	Min	9.76	0.423	11.3	4.90	4.90	5.39
200	Sec	20.5	0.660	22.9	17.8	0.46	19.5



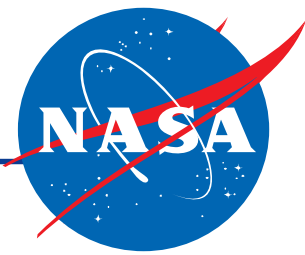
Dryness and Lethality

At 125°C, the ambient humidity and controlled humidity lethality rate constants are statistically different at the 99.00% confidence limit:

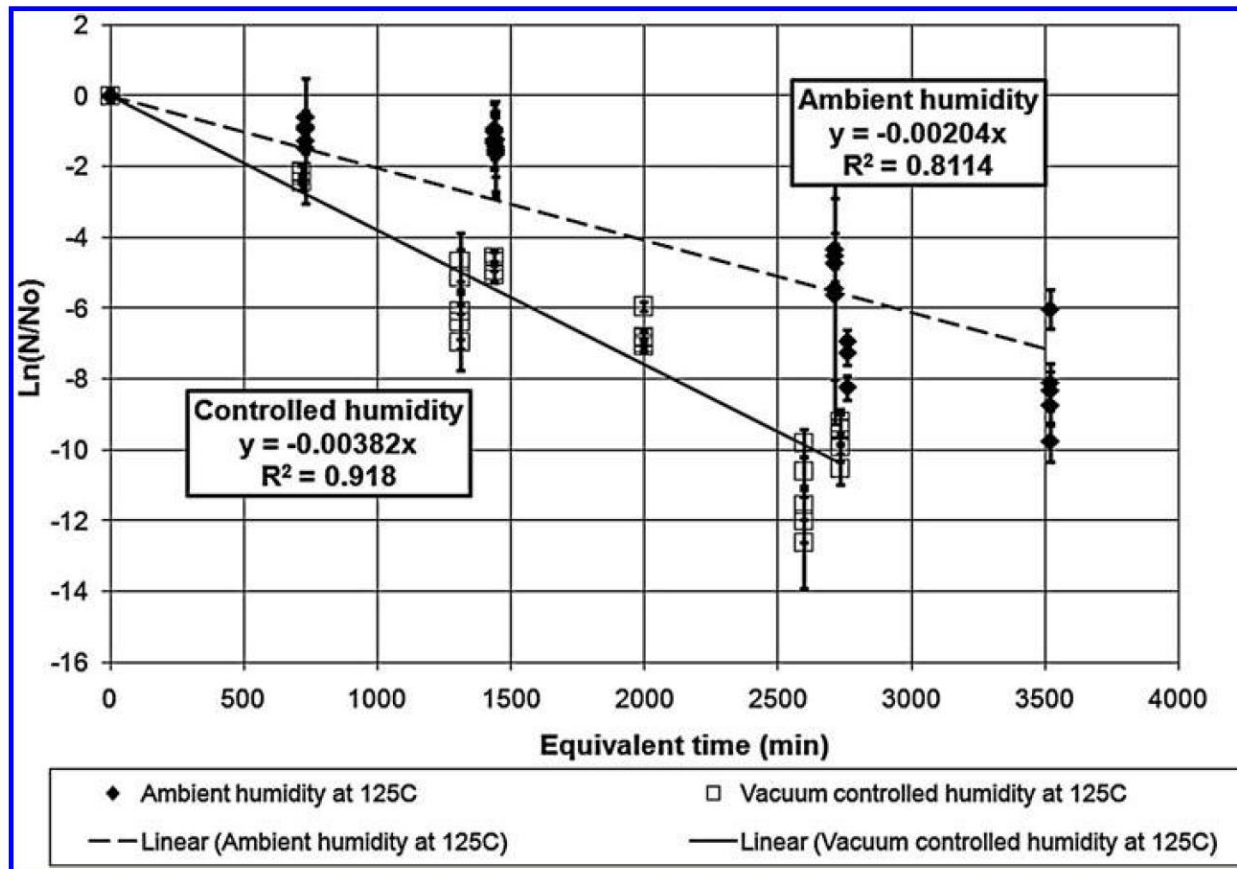
Dryness increases lethality of exposure to heat at 125°C or below

At 135°C, 150°C and 170°C, the ambient humidity and controlled humidity lethality rate constants are statistically the same at the 95% confidence limit:

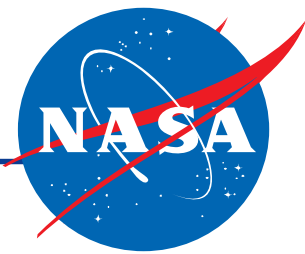
Dryness does not increase lethality at or about 130°C



Graphic representation of lethality data for a very dry-heat resistant organism



N_0 = number of visible colonies growing on agar before heat treatment
 N = number of visible colonies growing on agar after heat treatment



ECSS-Q-ST-70-57C
30 August 2013

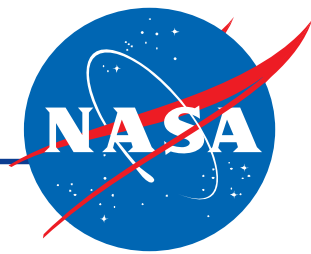
Dry Heat Bioburden Reduction for Flight Hardware exposed, mated and encapsulated surfaces are distinct cases

This standard defines procedures for the reduction of microbiological contamination of flight hardware using heat.

The procedures described in this standard cover:

- Reduction of microbiological contamination on exposed surfaces, mated surfaces and encapsulated in materials.
- Reduction of microbiological contamination in dry, ambient and uncontrolled humidity environments.

This standard also sets requirements for the conditioning of the flight hardware, bioburden reduction cycle development, and equipment to be used for applying a bioburden reduction procedure.



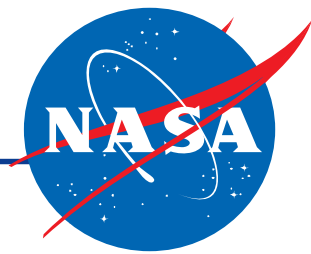
ECSS-Q-ST-70-57C
30 August 2013

Temperature dependent D-values in minutes for 2 to 3 order of magnitude surface bioburden reduction under dry humidity conditions shall be calculated using equation [5-1] for temperatures ≤ 140 °C and equation [5-2] for temperatures > 140 °C, with the temperature (T) in °C:

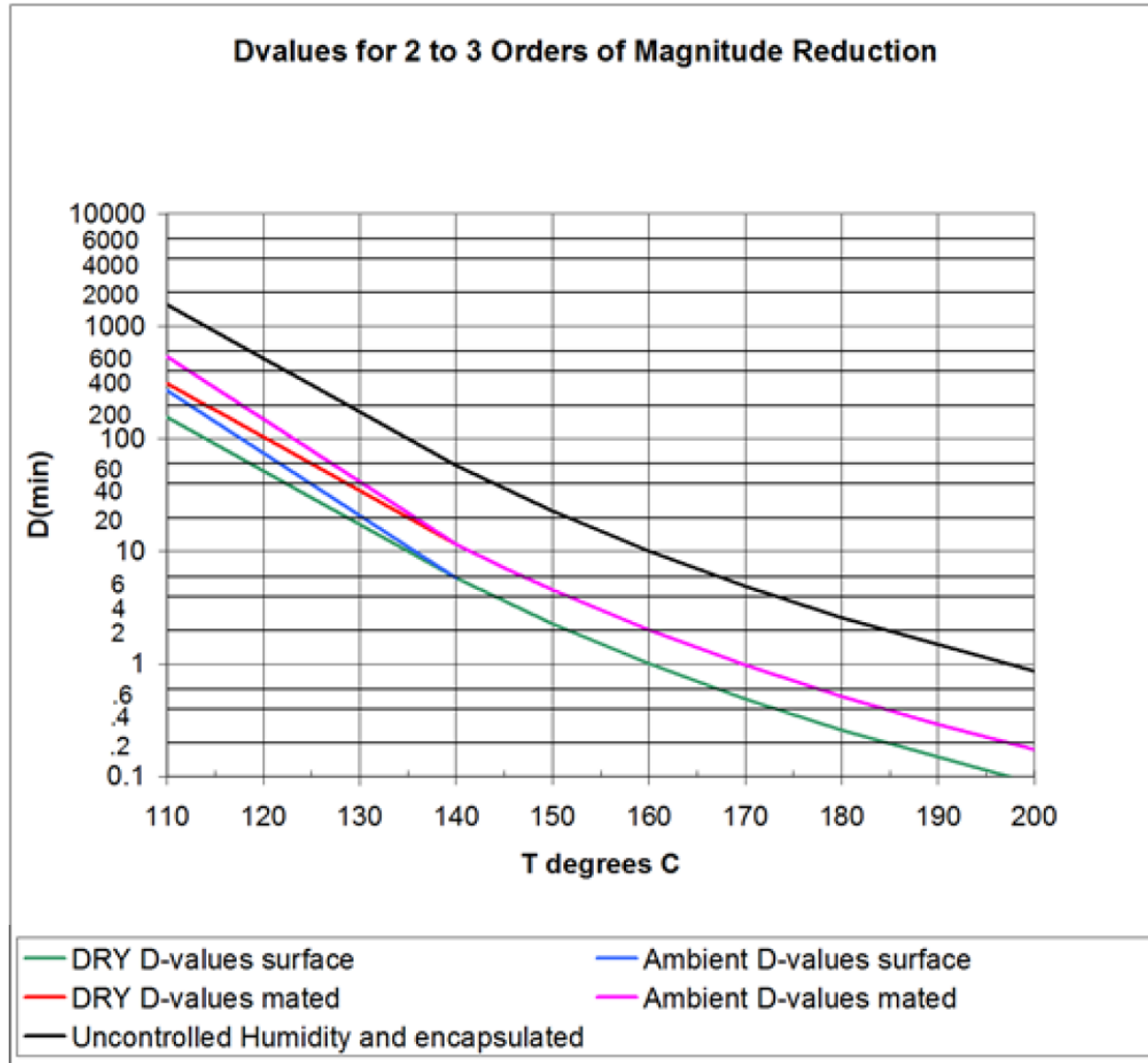
$$D(\text{min}) = 30 \times 10^{\frac{(125-T)}{21}}$$

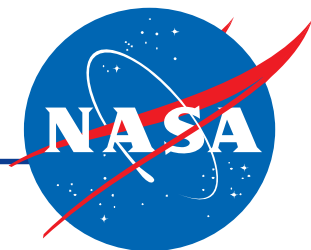
$$D(\text{min}) = 5,79 \times 10^{\frac{(140-T)}{(23 \times T)/140}}$$

NOTE 1 A 2 or 3 order of magnitude reduction is achieved by multiplying the respective D-values in equations [5-1], [5-2], and [5-3] by a factor of 2 or 3, respectively.



ECSS-Q-ST-70-57C
30 August 2013



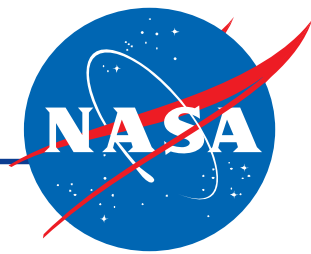


So Why Talk About Nuclear Power Sources?

- Heat
 - Radioisotope Thermal Generators convert heat from decaying ^{238}Pu with an efficiency of $\sim 6\%$
 - Space nuclear reactor concepts vary:
 $6\% < \epsilon < 35\%$
- Radiation
 - RTGs: Fairly benign
 - Reactors: How much do you need?

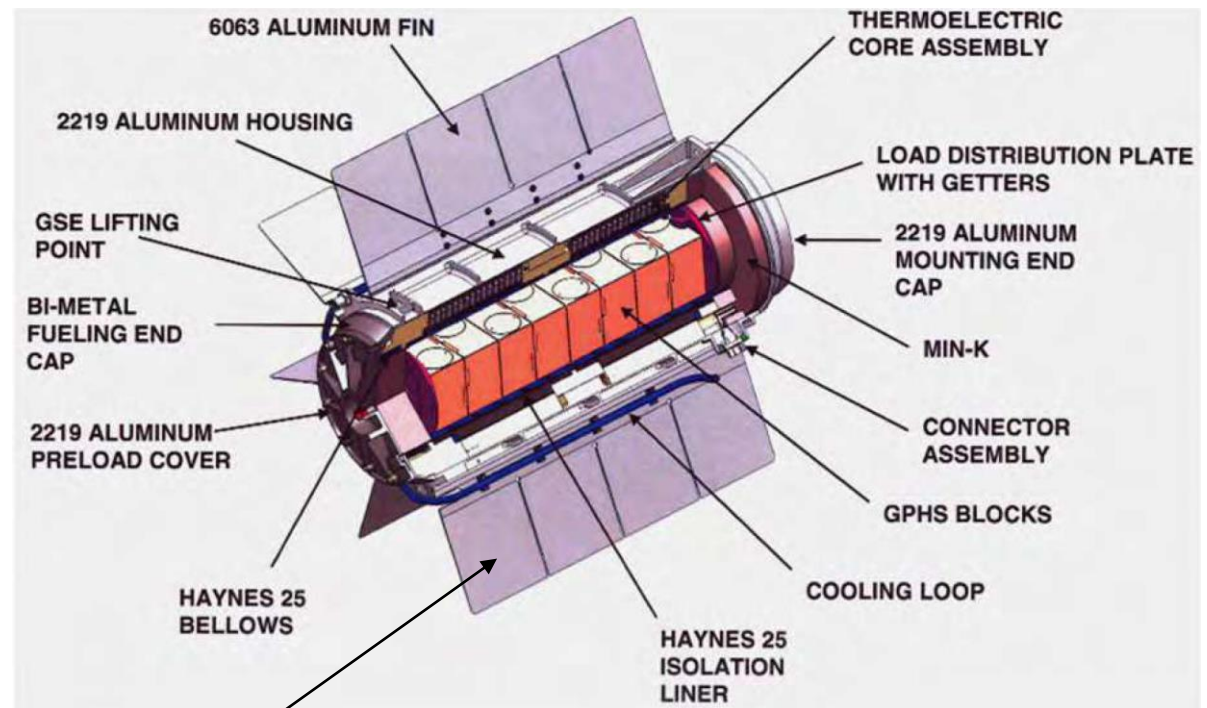
“One man’s trash
is another man’s
treasure”

Olde English
Proverb

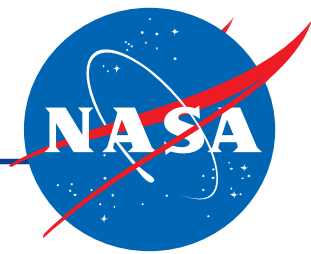


MMRTG Overview

- The MMRTG generator is about 25 inches (64 centimeters) in diameter (fin-tip to fin-tip) by 26 inches (66 centimeters) tall and weighs about 94 pounds (45 kilograms)
- 10.6 pounds (4.8 kilograms) of plutonium dioxide (including Pu-238) that
- Initially provides approximately 2,000 watts of thermal power and 110 watts of electrical power when exposed to deep space environments.



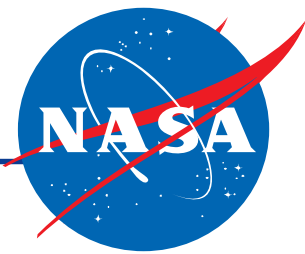
Nominal 200° C Surface Temp, Radiative Flux 2000 W/m²



Getting the heat where you need it

- Use electric power for Ohmic heating → wasteful...
- Smarter ways to do it
 - Forced Fluid Loops
 - Flexible means of transferring heat from a hot surface to heat another colder surface. Still needs power to drive pumps
 - Louvers and Insulation
 - Typically spacecraft elements are protected from radiant heat
 - Variable emissivity insulation layers can be used to create hot spots in areas for long term heating
 - Louvers with bimetallic actuators could act as “windows” to a radiator to hold an area in a desired temperature range
 - Heat Pipes
 - Multiple technology options: constant conductance, variable conductance, loop, and diode heat pipes
 - Attachment of heat pipes between hot and cold surfaces is accomplished by use of thermally conductive adhesive, press fit, or clamping

Where you go matters: Moon, planets with atmospheres, deep space all have special considerations

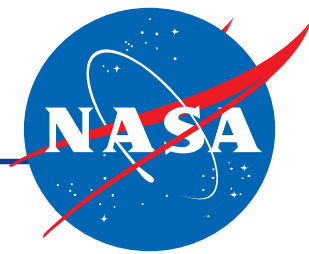


Radiation Sterilization

- Industrial Standard: ISO 11137-2:2013 *Sterilization of health care products -- Radiation -- Part 2: Establishing the sterilization dose*
 - Specifies methods for determining the minimum dose needed to achieve a specified requirement for sterility and methods to substantiate the use of **25 kGy or 15 kGy as the sterilization dose to achieve a sterility assurance level, SAL, of 10^{-6} .**
 - ISO 11137-2:2013 also specifies methods of sterilization dose audit used to demonstrate the continued effectiveness of the sterilization dose.

1 gray = 1 joule/kilogram

2.5 kGy = 2.5 Mrad → For gamma rays this equates to 2.5Mrem



Nuclear Power System Source Terms

RPS

- Primary
 - Alpha decay of ^{238}Pu
 - Associated gamma yield

Gamma Radiation From Plutonium-238 Product

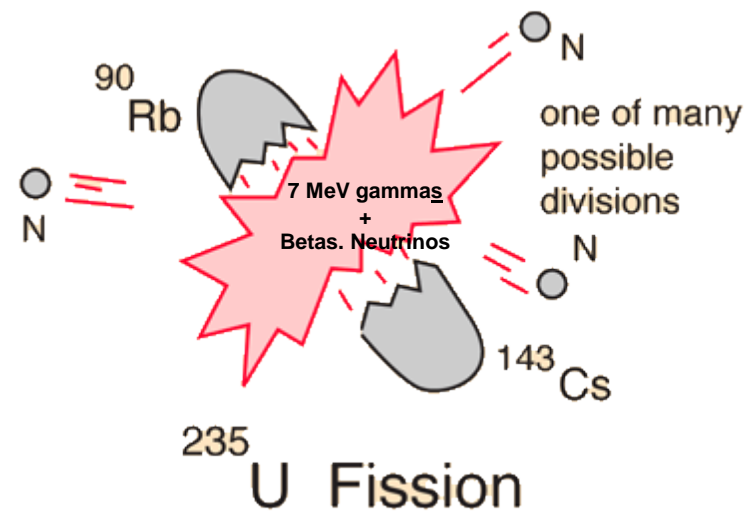
Isotope	Energy, Mev.	Photons/gram-sec.
^{238}Pu	0.017 (L x-ray)	6.7×10^{10}
	0.043	2.0×10^8
	0.099	4.6×10^7
	0.150	5.2×10^6
	0.203	2.0×10^4
	0.760	2.6×10^5
^{239}Pu	0.017 (L x-ray)	3.4×10^7
	0.052	2.4×10^4
^{240}Pu	Less abundant than ^{239}Pu	
^{241}Pu	0.145	$\leq 4 \times 10^4$
^{236}Pu	0.048	$\leq 7 \times 10^7$

- Secondary
 - Neutrons and prompt gammas from spontaneous fission of ^{238}Pu
 - Depends of ^{18}O “enrichment”
 - 11,300 n/sec → 113 n/sec

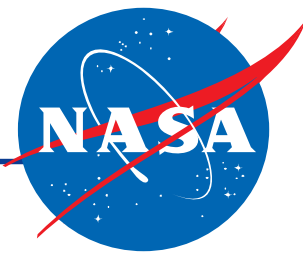
RPS radiation source diminishes over time, Nuclear Reactor scales with power and duration of operation

Nuclear Reactor

- Primary
 - Neutron induced fission of ^{235}U
 - ~200 MeV/fission



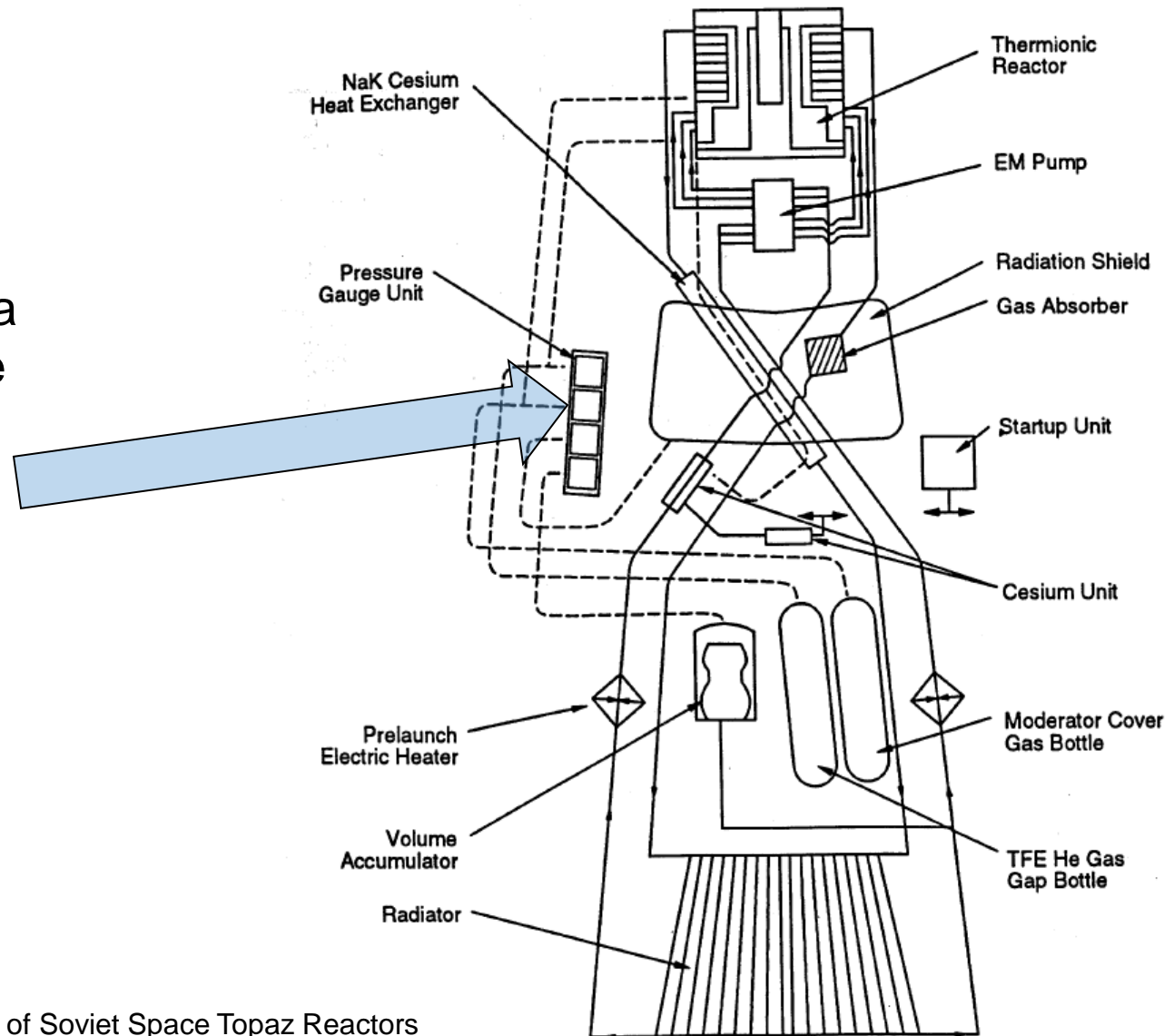
- Secondary
 - Radioactive decay of fission products
 - Depends on burnup (i.e., how long the reactor has operated at power)

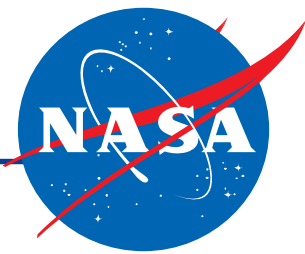


Schematic Reactor with Shield

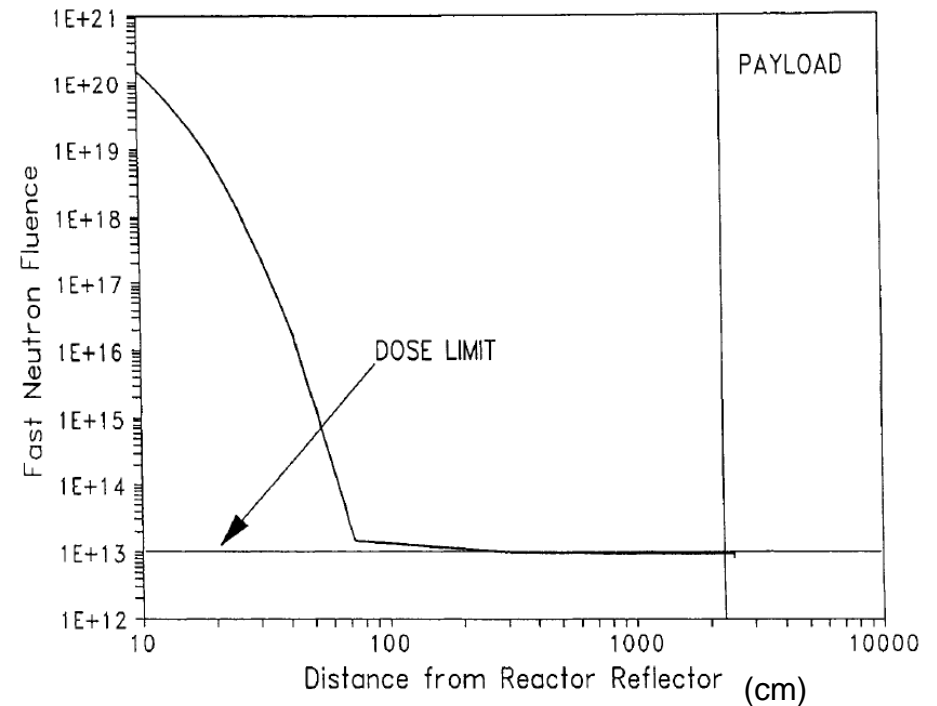
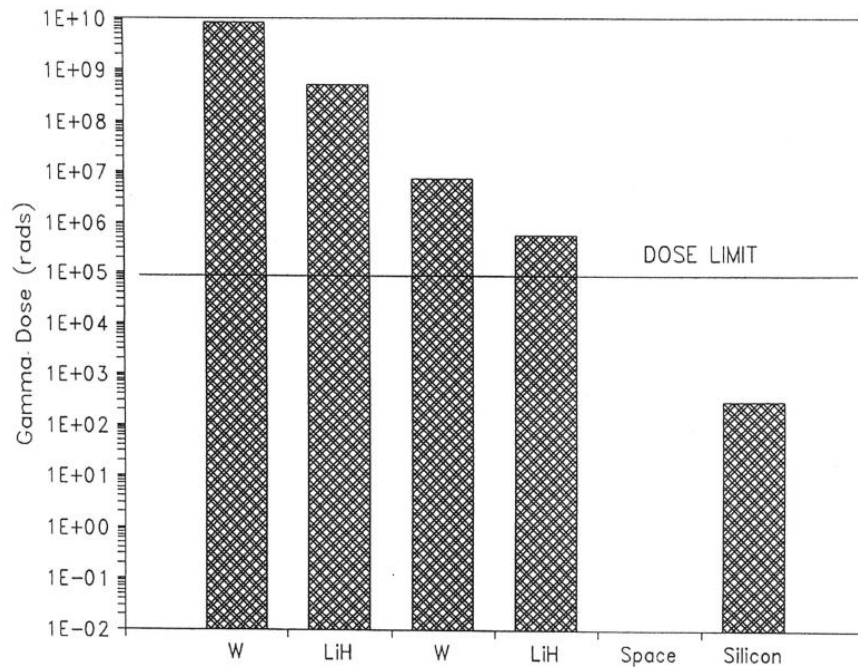
Radiation shield is a composite structure

- High Z materials shield intense gamma flux
- Low Z materials shield neutrons





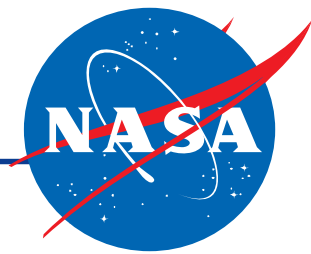
Representative Radiation Doses from 10 Years at Nominal 500 kW



System Modeling and Reactor Design Studies of the Advanced Thermionic Initiative Space Nuclear Reactor

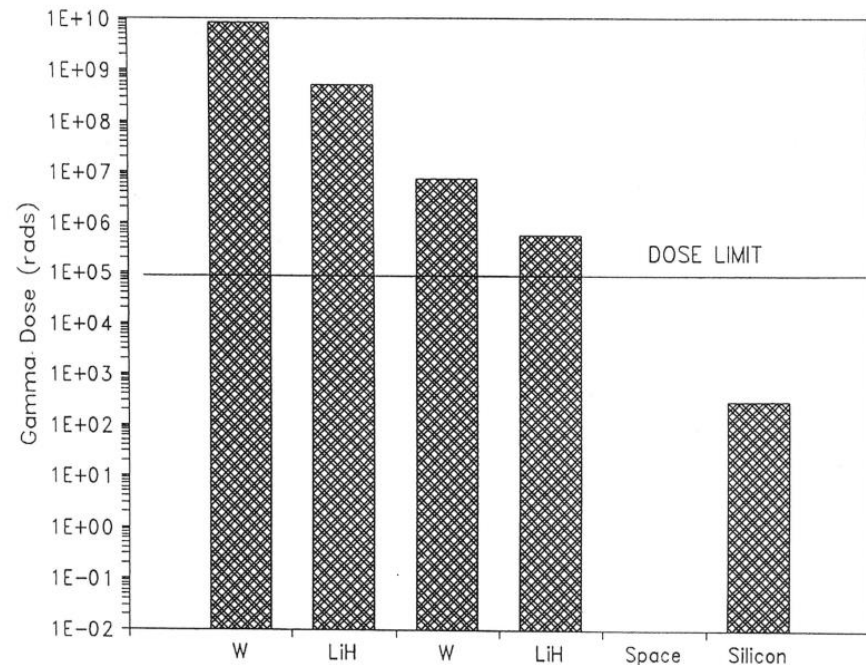
Hsing H. Lee, Shahab Abdul-Hamid & Andrew C. Klein

To cite this article: Hsing H. Lee, Shahab Abdul-Hamid & Andrew C. Klein (1996) System Modeling and Reactor Design Studies of the Advanced Thermionic Initiative Space Nuclear Reactor, Nuclear Technology, 115:1, 1-21, DOI: [10.13182/NT96-A35271](https://doi.org/10.13182/NT96-A35271)



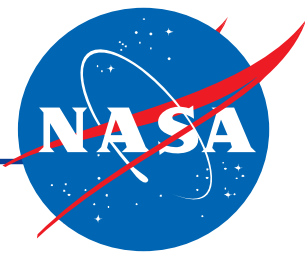
Design Considerations

- 1 Mrad/year dose environment in front of second tungsten shield @ 500 kW thermal power – scales linearly with time and power
- Fast neutron dose in this region will increase the dose; however, neutron activation of the sterilized article will be to be taken into account



Key takeaways

1. Reactors can provide ample amounts of heat, ohmic or passive methods of heat transfer are equally viable
2. Using radiation sterilization methods will require complex engineering of radiation shields, but may be worth the effort depending on planetary protection objectives and goals



Conclusions

- Heat and radiation from RTG's and reactors are potentially useful to increase lethality for terrestrial organisms during multi-year cruises to bioburden-protected destinations in the outer solar system.
- Research is needed on lethality in the temperate range of 80° to 120° C for time periods of years.
- Research is needed on lethality for specific types of ionizing radiation and for secondary products (chemical cascade) from radiolysis of water molecules such as peroxides and radicals.