



### Change of Inertia Tensor Due to a Severed Radial Boom for Spinning Spacecraft

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### **Spin-Stabilized Spacecraft**



- Gyroscopic action
- Easy to tension long wire booms due to centrifugal force
- Magnetic and electric fields, plasma, etc.
- Thin boom structure can be a risk





## **History of Radial Boom Breakage**

- Fast Auroral SnapshoT (FAST)
  - 1996: Deployment failure
- Imager for Magnetopause-to-Aurora Global Exploration (IMAGE)
  - 2000: -X antenna damage
  - 2001: +Y antenna damage
  - 2004: <sup>+</sup>Y antenna damage (again)
- ARTEMIS P1 (formerly THEMIS B)
   2010





Image from https://www.nasa.gov/vision/universe/solarsystem/fast\_10yr.html



#### **Boom Contribution to Moment of Inertia**

- Change in mass may be negligible (100 g lost versus 1,400 kg spacecraft)
- Mass moment increases with square of distance

$$J = mr^2$$

• At a radius of 100 m, a 100 g mass contributes 1,000  $kg \cdot m^2$  (typical total moment of inertia could be 5,000  $kg \cdot m^2$ )





### **Motivation**

- 1. Impact of radial boom anomaly to mass moment of inertia tensor is significant
- While inertia tensor is not directly observable, direction of Major Principal Axis (MPA) is observable for some missions
- 3. Location of break along boom should be related to some change in MPA





### Assumptions

- 1. Motion is steady-state, all vibrations damped
- 2. With no internal motion, inertia tensor is same as a rigid body
- 3. Torque-free motion
- 4. Given (1), (2), and (3): *MPA*, angular velocity  $\vec{\omega}$ , and angular momentum  $\vec{L}$  all coincide





#### Magnetospheric MultiScale (MMS) Mission



- Spin-plane Double Probe (SDP)
- Axial Double Probe (ADP)
- Analog Flux Gate (AFG)
- Search Coil Magnetometer (SCM)
- Digital Flux Gate (DFG)
- Electron Drift Instrument (EDI)

The main coordinate system considered is the Observatory Coordinate System (OCS)



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#### Existing MMS Attitude Ground System (AGS)

- Based on a software suite that has been used for many missions
- Center of Mass (CM) and inertia tensor models developed specifically for MMS
  - Use pre-launch determined values
  - Account for deployment status of booms
  - Assume nominal spin axis (OCS Z-axis)
- Inertia tensor calibration (fuel asymmetry)





## **Proposed Improvements**

- Account for directions of booms at steady-state
  - Net torques and forces are zero
  - Radial to spin vector, not Z-axis
  - "Radial" = intersecting + perpendicular
- Account for mutual dependence of boom directions and MPA
- Account for fully or partially severed boom(s)





### **Big Picture of Improved Model**







#### "Inner" Iteration







# Inertia Tensor (1 of 2)

- $\hat{b}$  and *CM* must be given (and  $f_b$  and  $f_f$ )
- "Build" inertia tensor of system from constituents that have known inertia tensors
  - Spacecraft body
  - Basic 3D solids (thin rod, sphere, cylinder)
- Parallel axis theorem translates inertia tensor to/from center of mass and arbitrary point



# Inertia Tensor (2 of 2)

- Boom Direction Coordinate System (BDCS)  $\begin{bmatrix} J_{axial} & 0 & 0 \\ 0 & J_{transverse} & 0 \\ 0 & 0 & J_{transverse} \end{bmatrix}$
- Change tensor orientation via similarity transformation

$$A_{BDCS\leftarrow OCS} = f(\hat{b})$$
  

$$A = A_{BDCS\leftarrow OCS}^{T}$$
  

$$J_{OCS} = A J_{BDCS} A^{T}$$

- Overall process:
  - 1. Build each boom tensor (parallel axis theorem)
  - 2. Transform each boom tensor from *BDCS*<sub>boom</sub> to OCS
  - 3. Build total spacecraft tensor (parallel axis theorem)





### "Outer" Iteration (1 of 2)







## "Outer" Iteration (2 of 2)

- Accelerated method converges in approximately  $^{1}\!/_{10}$  to  $^{1}\!/_{3}$  the number of iterations
- Based on assumption that error decreases roughly as a geometric progression ( $\epsilon = ar^n$ ; |r| < 1)
- Related to Aitken's  $\delta^2$ -process (a.k.a. Aitken extrapolation)





#### **Big Picture of Improved Model (Review)**







## Results (1 of 5)

Suppose a coordinate system has its X-axis parallel to the nominal direction of the severed boom. Let  $\varphi_1$  and  $\varphi_2$  define the "tilt" of the MPA from the OCS Z-axis.







### Results (2 of 5)







## Results (3 of 5)

- Uncertainty in *MPA* is approx.  $0.003^{\circ}$  (3 $\sigma$ )
- Let  $\Phi$  denote the change in *MPA* due to break
- Uncertainty in  $\Phi$  is approx. 0.006° (3 $\sigma$ )
- Let *X* denote the break location, measured in meters from the attachment point of the boom
- Want to know uncertainty  $\Delta X$  given uncertainty  $\Delta \Phi$
- First order approximation:

$$\Delta X \approx \left| \frac{d\Phi}{dX} \right|^{-1} \Delta \Phi$$





### **Results (4 of 5)**







### **Results (5 of 5)**

Approximate values for  $3\sigma$  uncertainty in break location for various regions of the boom.

Break Region	$3\sigma$ Uncertainty
Near Boom Attachment	6 m
Near Boom Midpoint	50 cm
Near Boom Tip	20-30 cm





## **MMS Application**

- Predictive products:
  - Rigid body inertia tensor is used to calculate gravity gradient torque
- Definitive products:
  - Extended Kalman Filter (EKF) uses inertia tensor in propagation step
- Mass properties
  - CM and rigid body inertia tensor are reported for onboard use





### **Future Work**

- Investigate whether choice of independent variable(s) ( $\varphi_1$ ,  $\varphi_2$ ,  $\Delta MPA$ ) affects accuracy of boom fraction mapping
- Implementation is already generalized for multiple breaks ( $f_b$  is a row vector)
  - May result in ambiguous solutions
  - Requires analysis
- Model boom deployment failure (requires little modification)
- Incorporate new model into inertia tensor calibration tool





## Summary

- The effects of a radial boom break were shown to be observable and quantifiable
- An improved model for CM and inertia tensor was developed for the MMS mission
- Based purely on attitude observations, location of boom break can be estimated to within a small uncertainty





### Questions





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#### Backup Slide: Why MPA, $\vec{\omega}$ , and $\vec{L}$ Coincide

- 1.  $\vec{L}$  is fixed relative to space
- 2. *MPA* is fixed relative to the body
- 3.  $\vec{\omega}$  nutates, tracing out "body cone" and "space cone"
- 4. Nutation of  $\vec{\omega}$  induces internal motion
- 5. If all motion is damped,  $\vec{\omega}$  is no longer nutating (angle between vectors is zero)



