

Comparisons between Supercell Kinematics and Lightning Optical Energy Output from GLM

Dustin Conrad ¹, Christopher Schultz ², Mason Quick ³

1 - University of Alabama - Huntsville

2 - NASA Marshall Space Flight Center

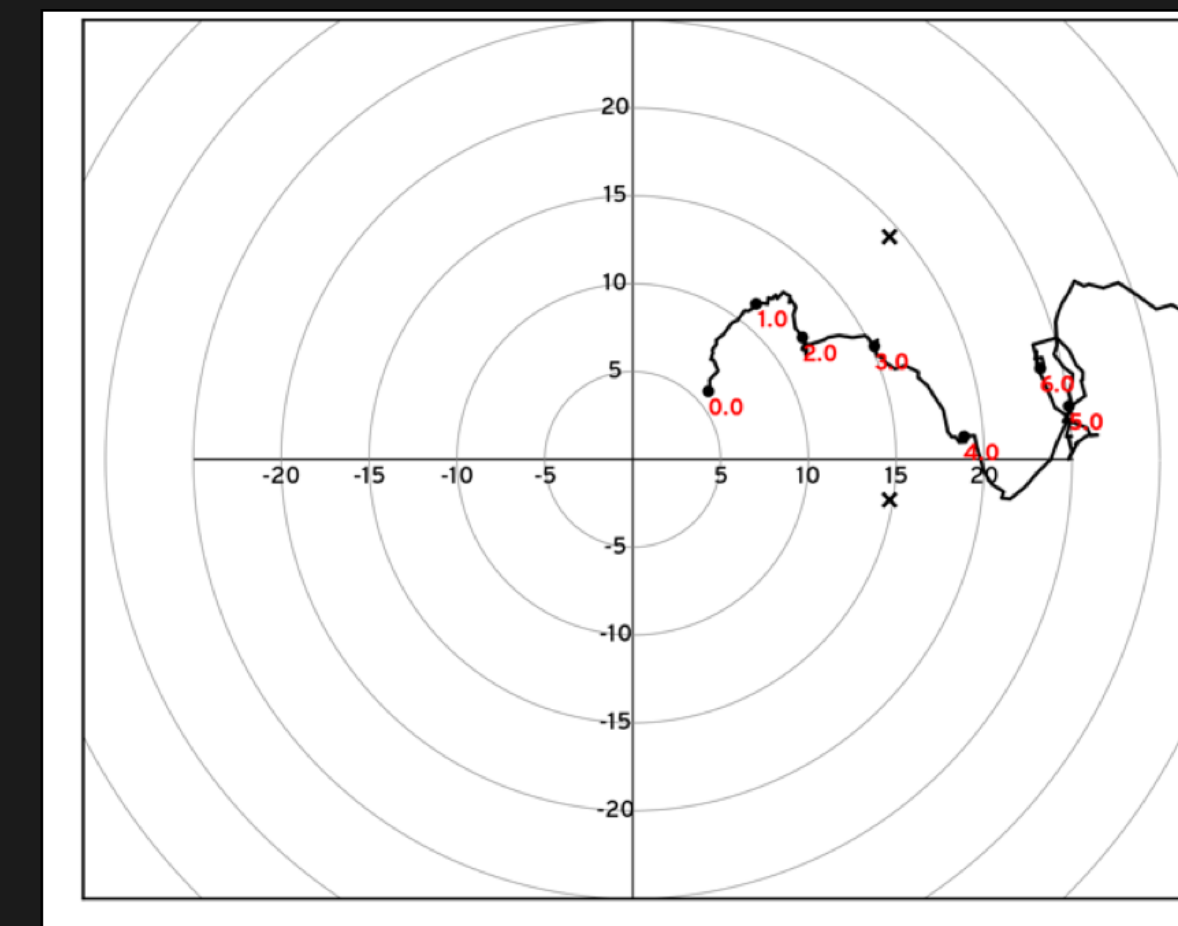
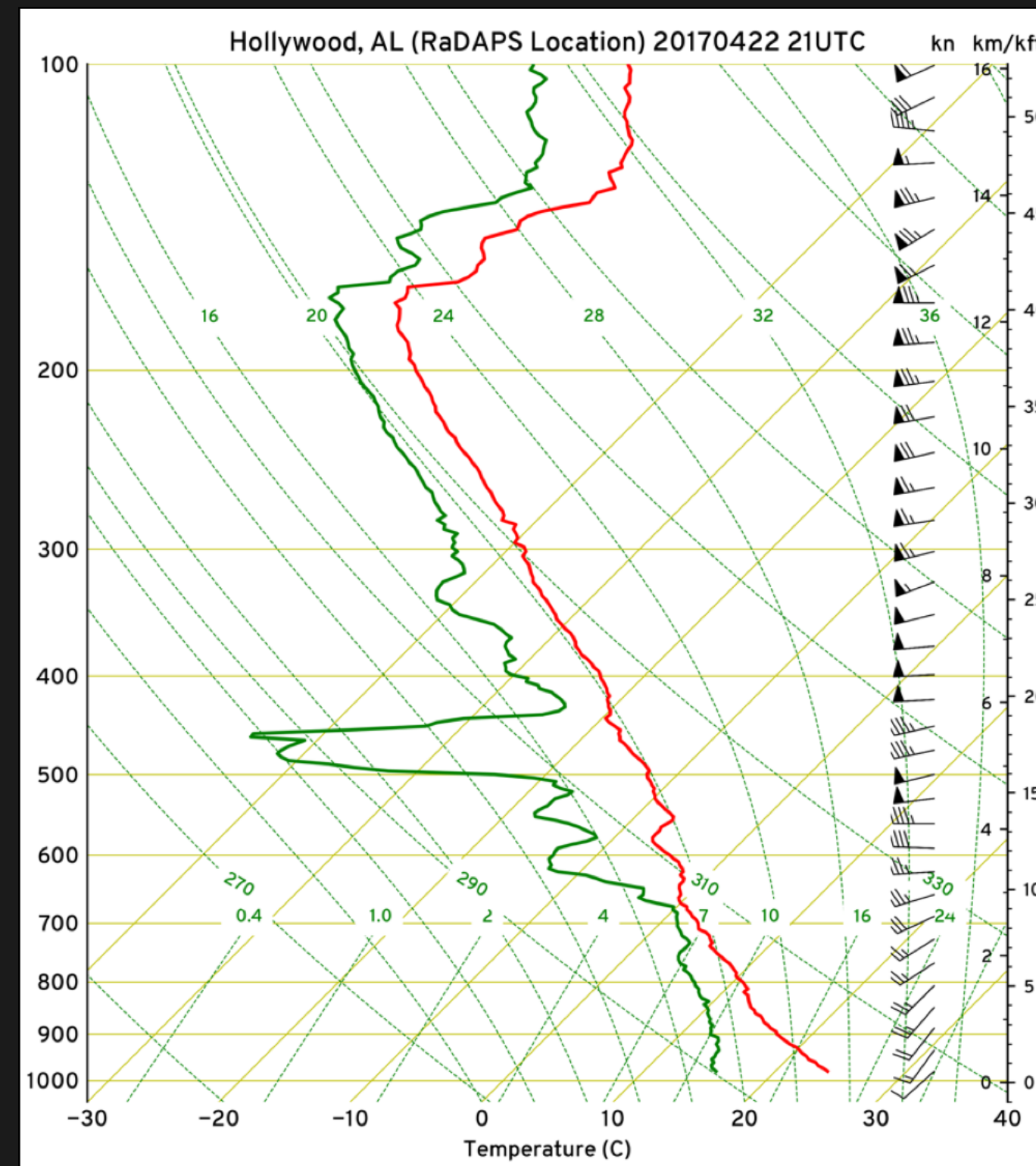
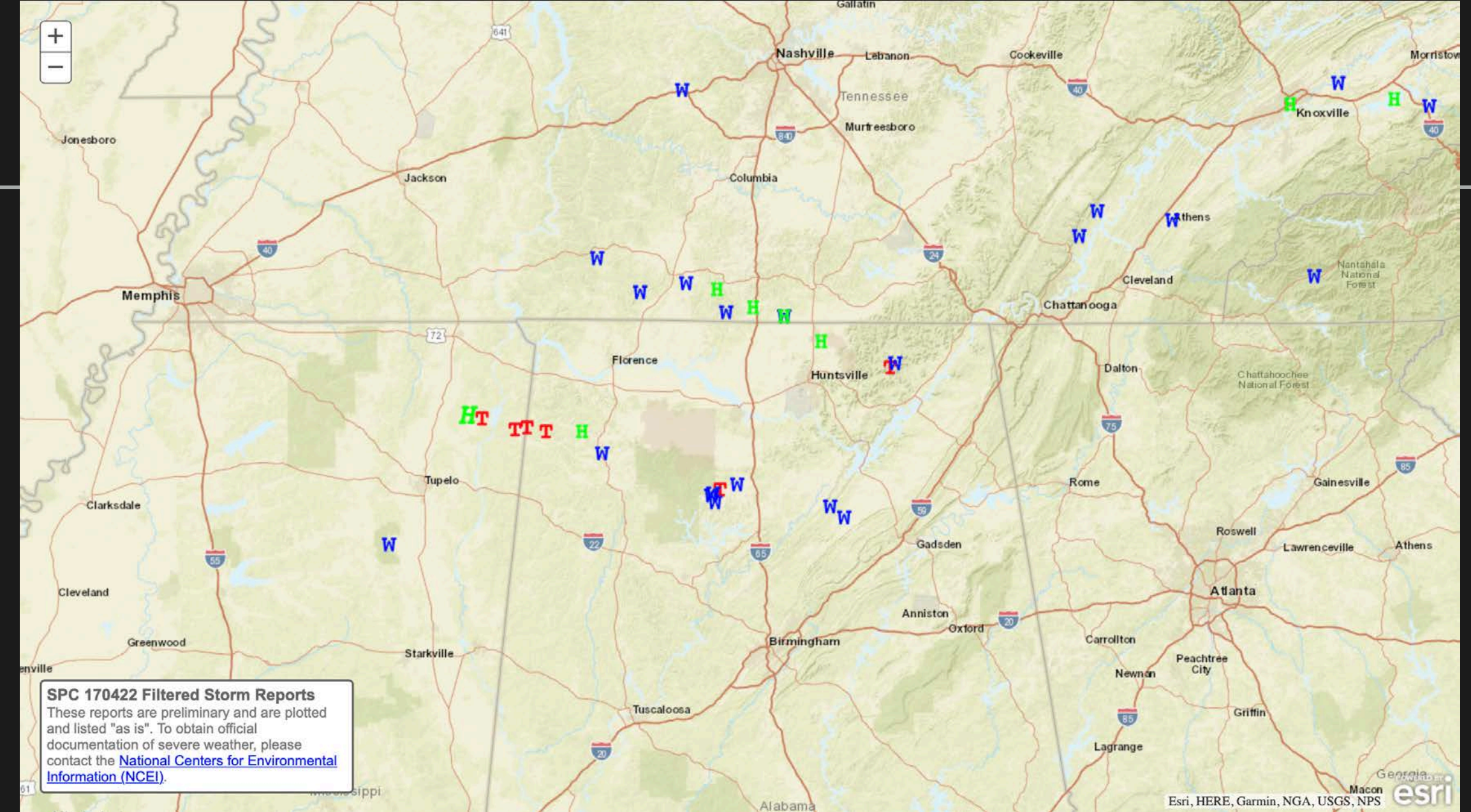
3 - Earth System Science Center University of Alabama - Huntsville

Motivation

- ▶ With the launch of the Geostationary Lightning Mapper (GLM) in Nov 2016, continuous lightning observations of the Western Hemisphere from space are now available
- ▶ Launches of additional and future lightning mappers on GOES-17 (Mar. 2018), Meteosat (2020s), and other geostationary satellites will allow continuous lightning measurements from $\pm 55^\circ$ latitude
- ▶ In addition to climatology, lightning activity can be used as an aid in nowcasting severe weather in regions with poor radar or LMA coverage
- ▶ Before being applied on a global scale, need to understand lightning trends with severe weather using space based measurements coincident with ground based measurements

22 Apr 2017

- ▶ Supercells across the North Alabama mesoscale domain
- ▶ EF-1 and EF-0 tornadoes, along with multiple hail and wind reports, associated with two supercells
- ▶ 552 J/kg of mixed layer CAPE with 37 kts of 0-6 km shear
- ▶ Occurred during the VORTEX-SE and GOES-R Validation field campaigns



500 m Mixed Layer Parcel

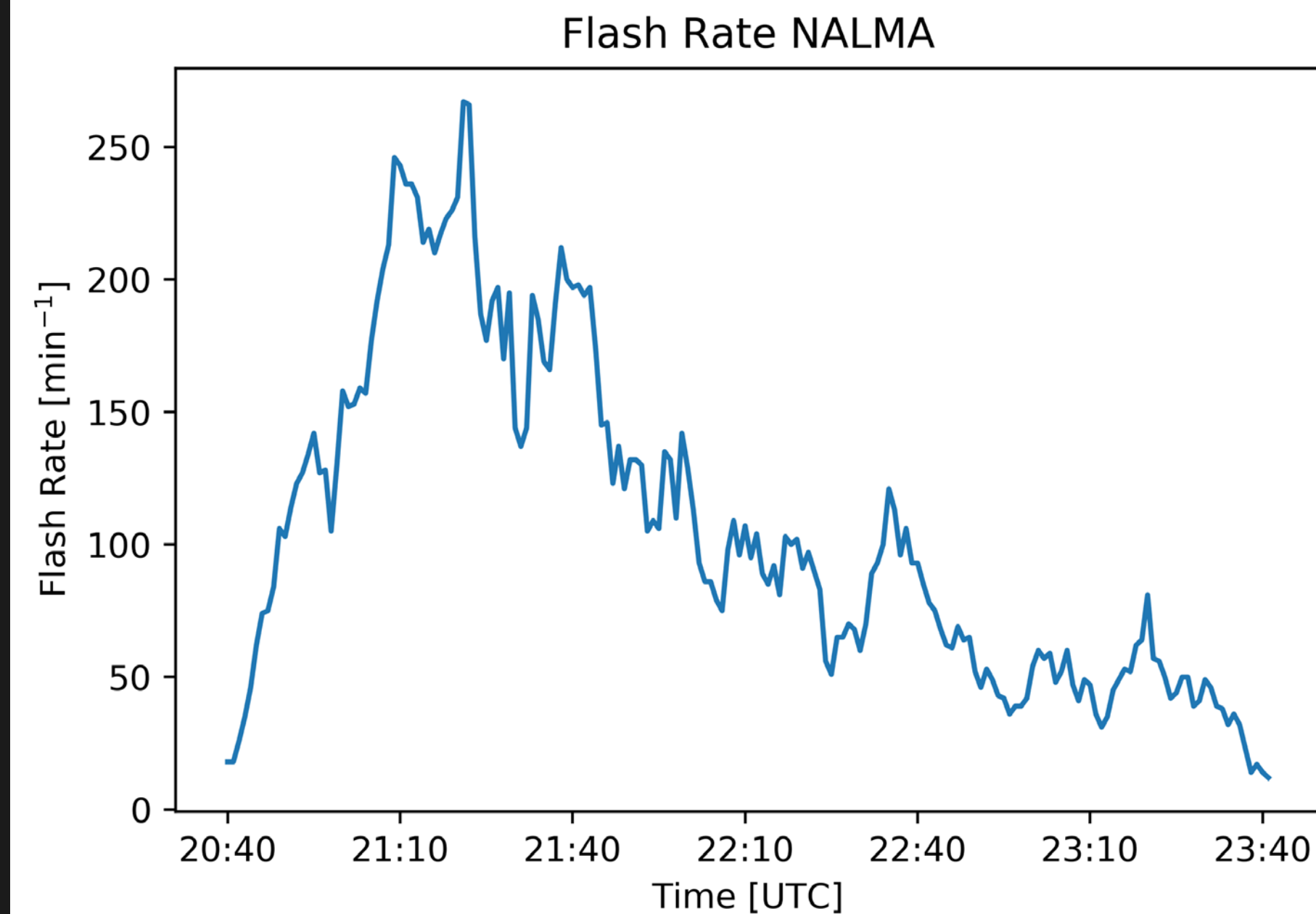
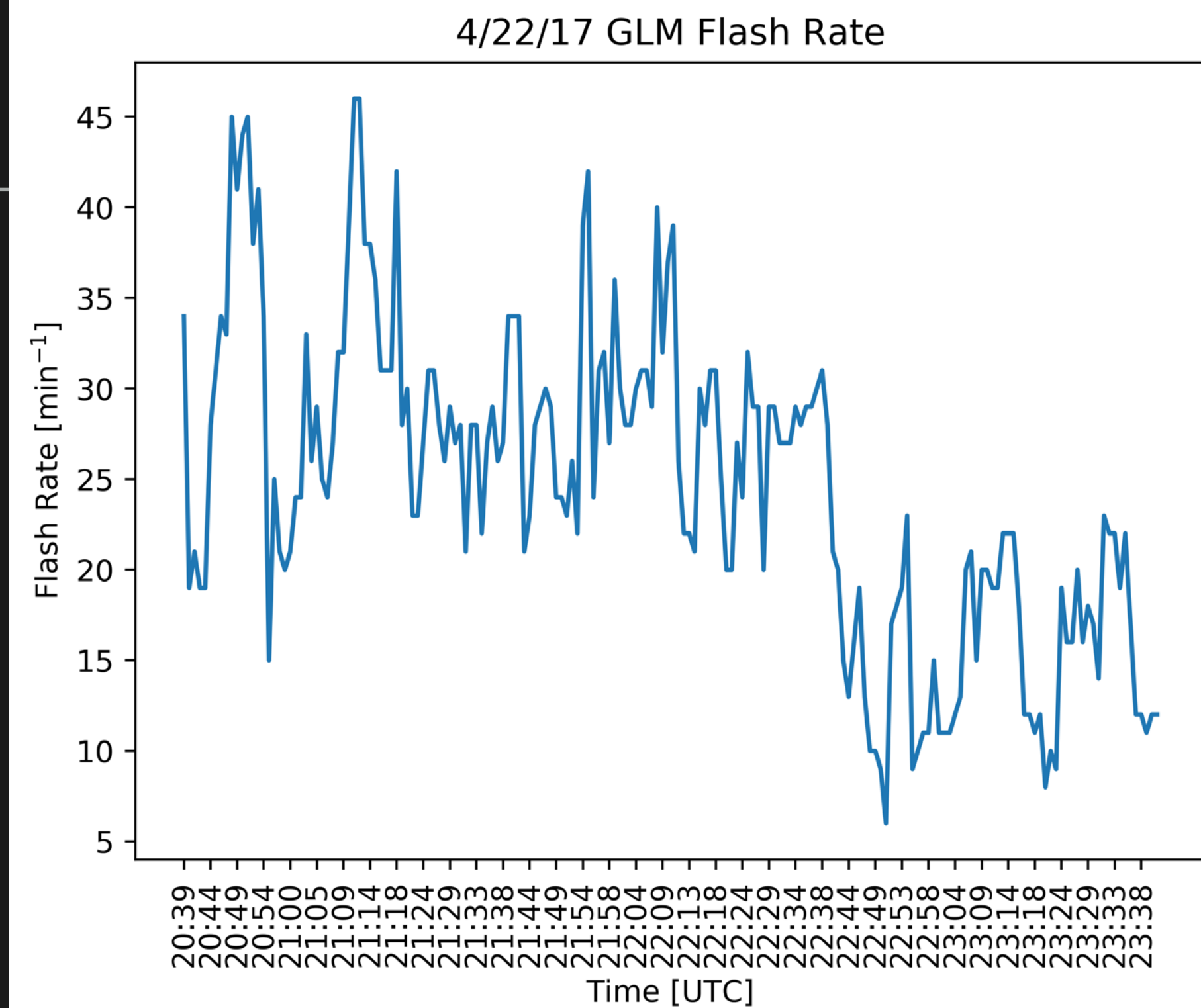
CAPE	552	J kg^{-1}
CIN	15	J kg^{-1}
TOPS	1651.73	km
θ_e	330.1	K
L_{MAX}	-4.6	C
L_{500}	-2.2	C
L_{300}	0.3	C
Parcel	979	mb

Hodograph

0-1 km shear	28° 11.0 kts
0-3 km shear	74° 19.0 kts
0-6 km shear	86° 36.8 kts
SRH 0-1	$73 \text{ m}^2/\text{s}^2$
SRH 0-3	$129 \text{ m}^2/\text{s}^2$
BRN	10
0-1 km shear	54° 5.7 kts
1-2 km shear	244° 3.3 kts
2-3 km shear	188° 4.1 kts
3-4 km shear	262° 7.3 kts
4-5 km shear	143° 6.2 kts
5-6 km shear	-72° 2.7 kts

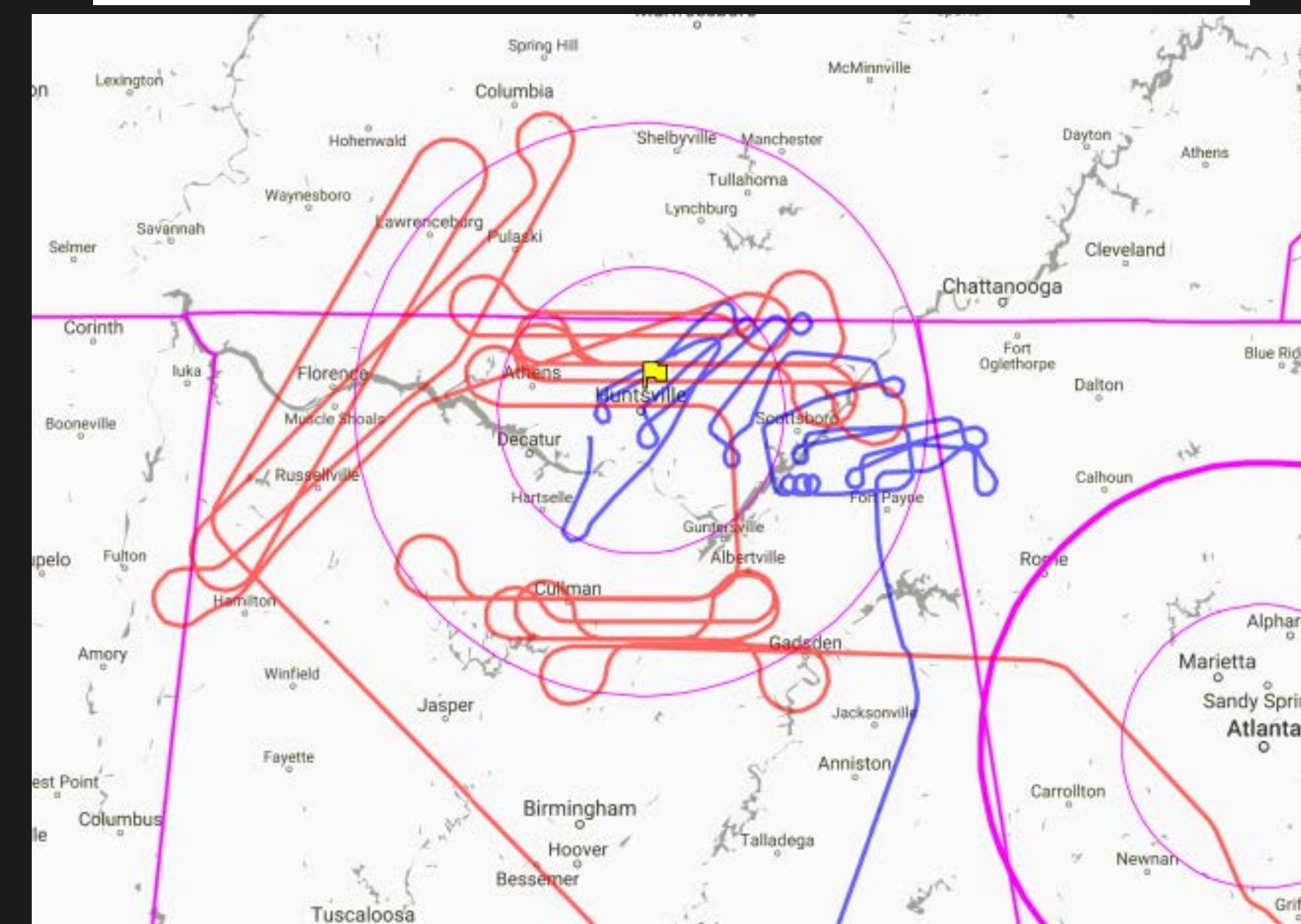
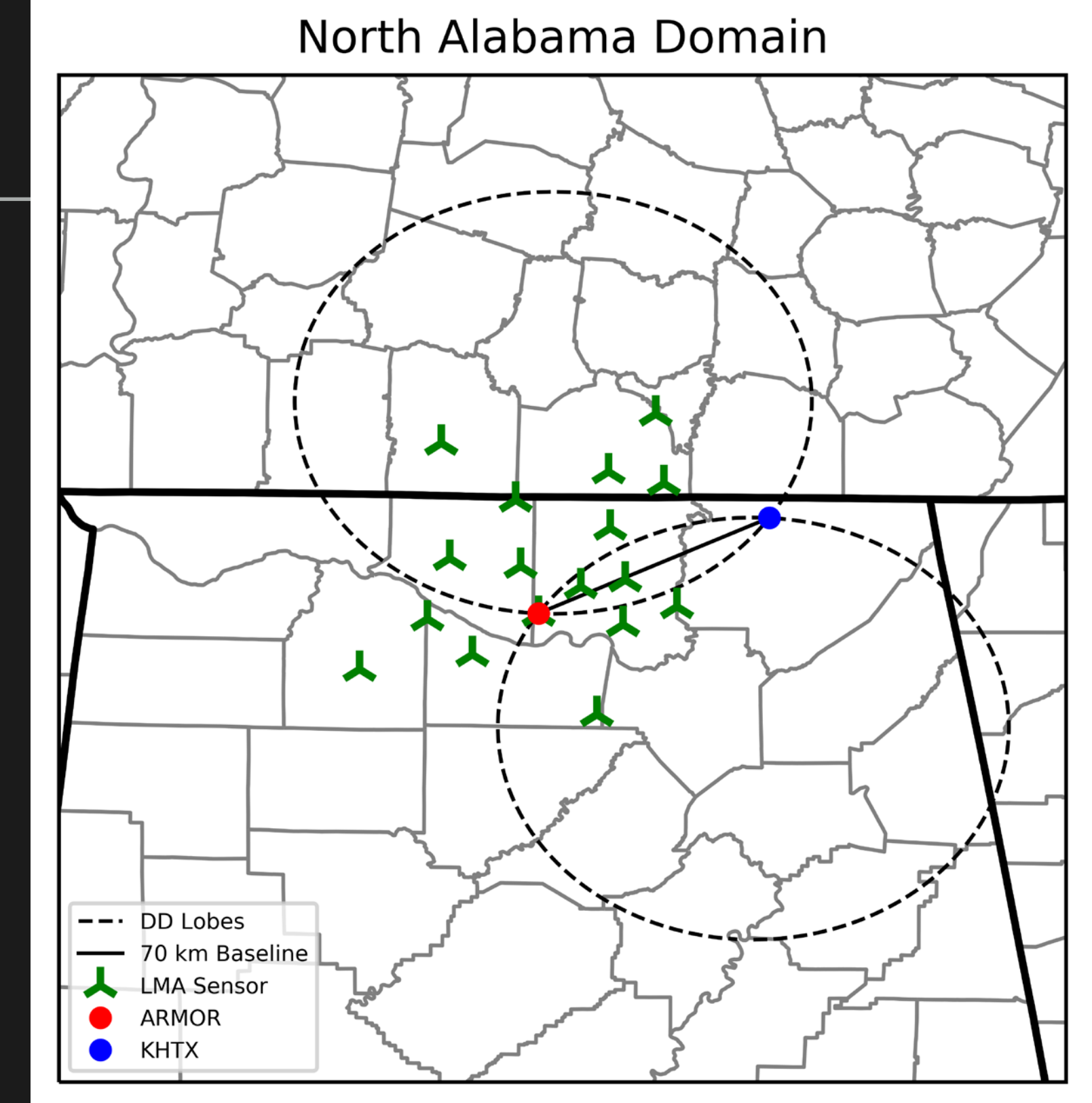
LMA vs GLM Flash Rates

- ▶ GLM flash rates much lower than flash rates from LMA
 - ▶ GLM max flash rate around 45 flashes per minute, LMA near 275 flashes per minute
 - ▶ This trend seen with different LMA clustering thresholds and algorithms
 - ▶ Implications with lightning jump (poster 1018 Wed. 4-6pm)
- ▶ Cause of this difference?
 - ▶ Resolution?
 - ▶ Optical extinction?
 - ▶ Flash clustering?
- ▶ Goal of using optical energy is to remove the flash clustering process



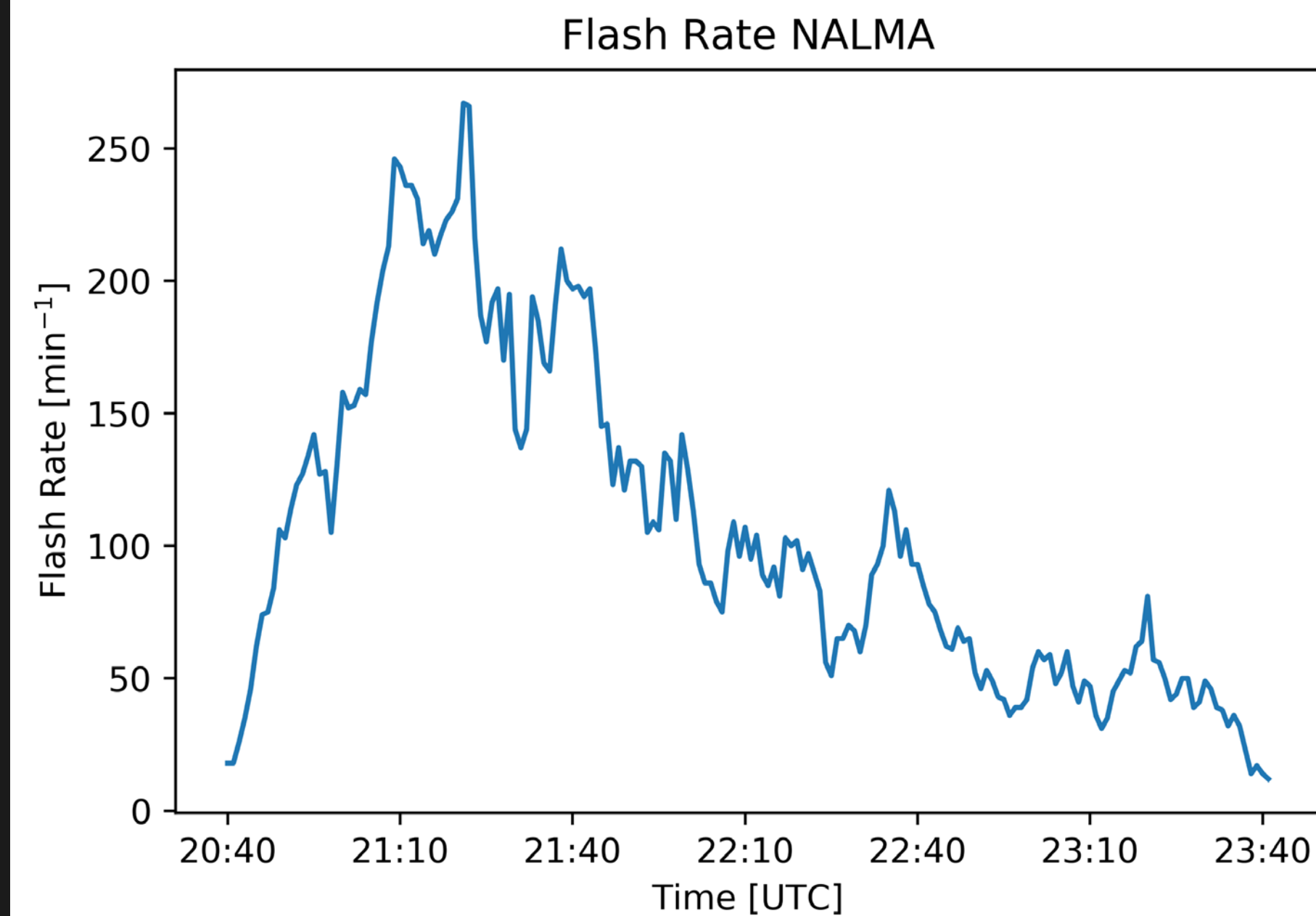
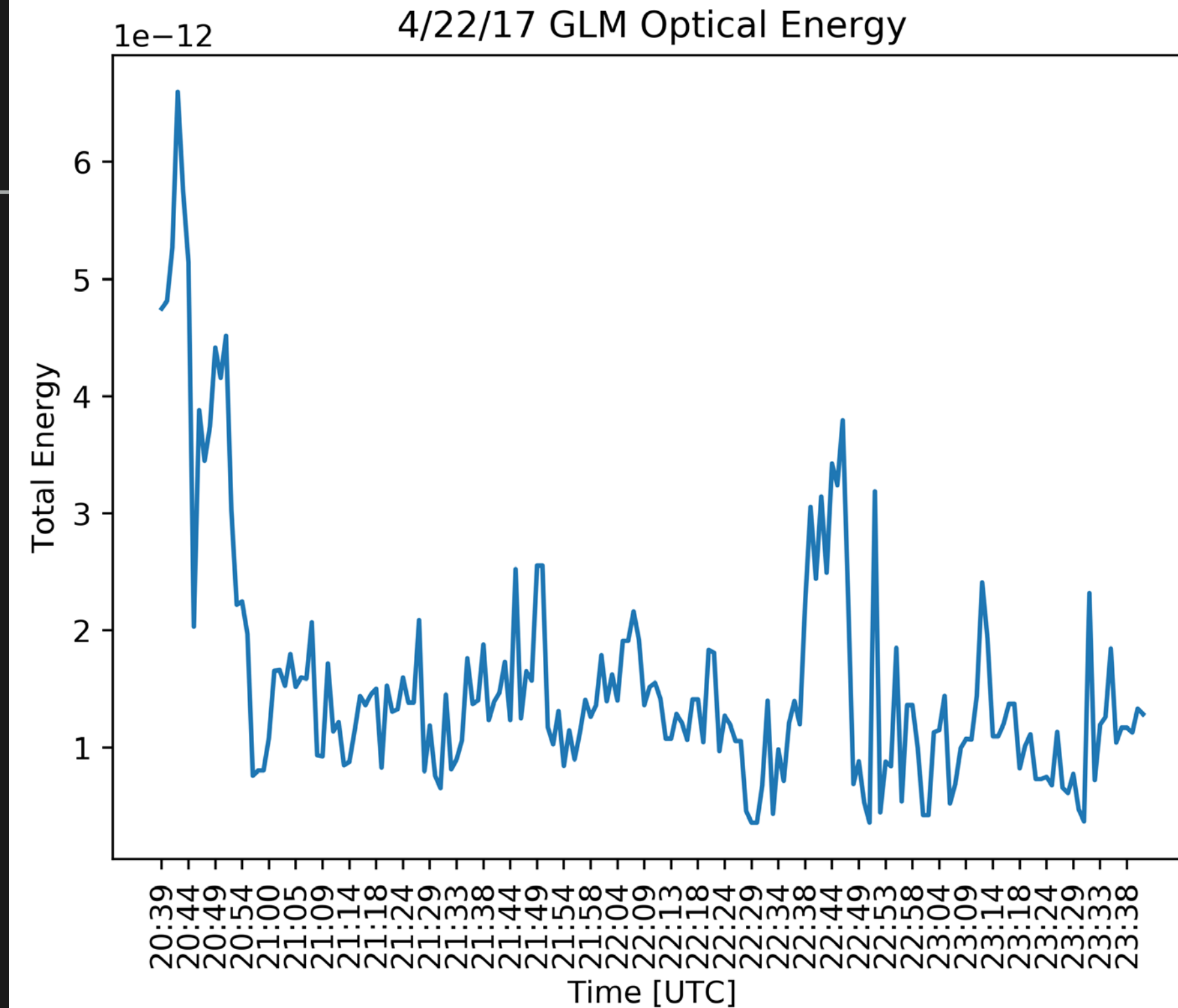
Data and Methodology

- ▶ Lockheed-Martin reprocessed GLM data
- ▶ Undecimated LMA data from the North Alabama LMA (NALMA)
- ▶ LMATools flash clustering
 - ▶ 0.3 s time threshold
 - ▶ 3 km distance threshold
 - ▶ Chi square value of 1
 - ▶ 6 station minimum
- ▶ Radar data from KHTX and ARMOR manually unfolded and quality controlled
- ▶ Data gridded to a 1 km x 1 km x 0.5 km grid
- ▶ Multi-Doppler winds found using PyART Multidop package which uses a 3DVAR approach (Shapiro et al. 2009, Potvin et al. 2012)
- ▶ HID using CSURadartools (Dolan et al. 09, Dolan et al. 13)



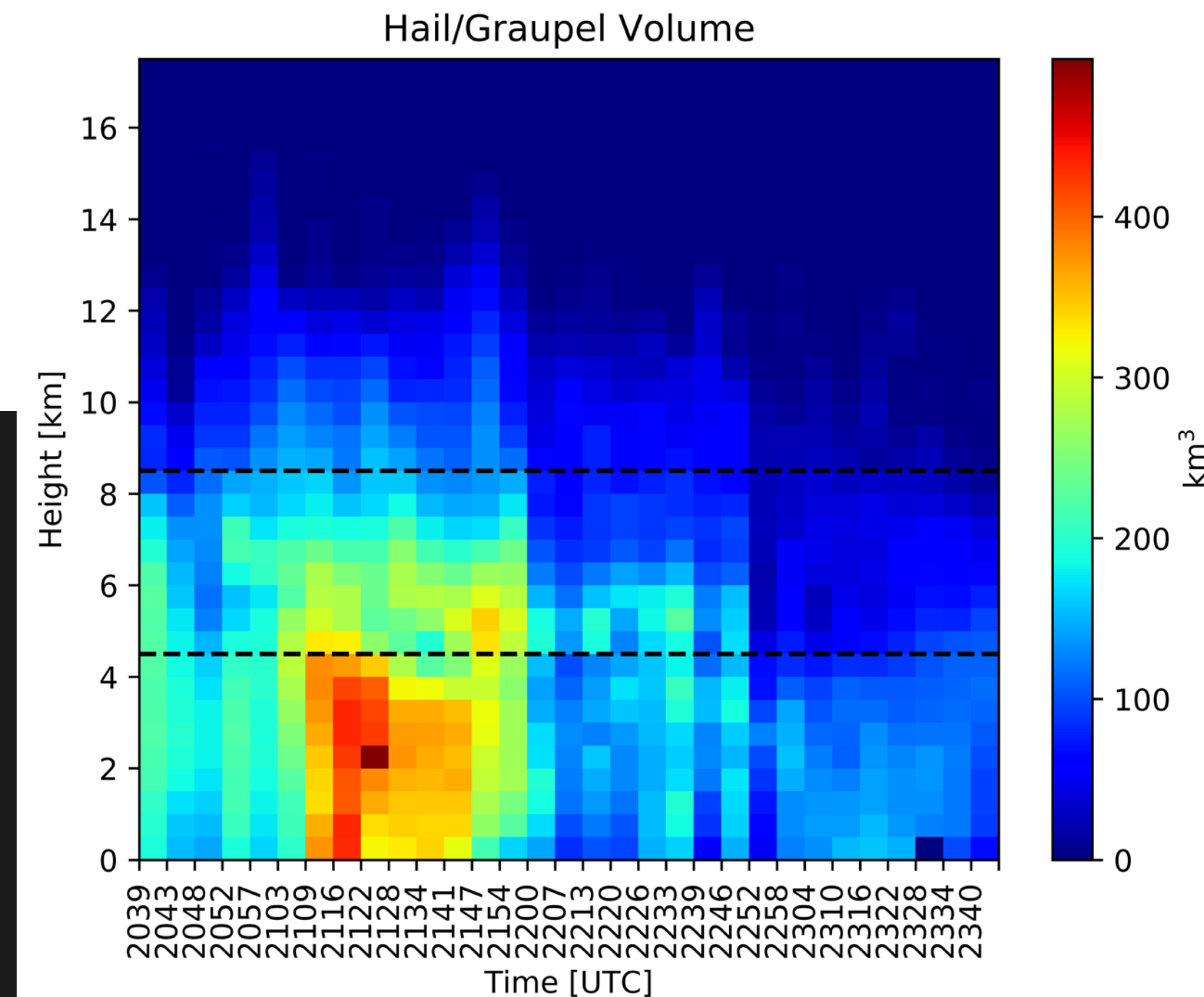
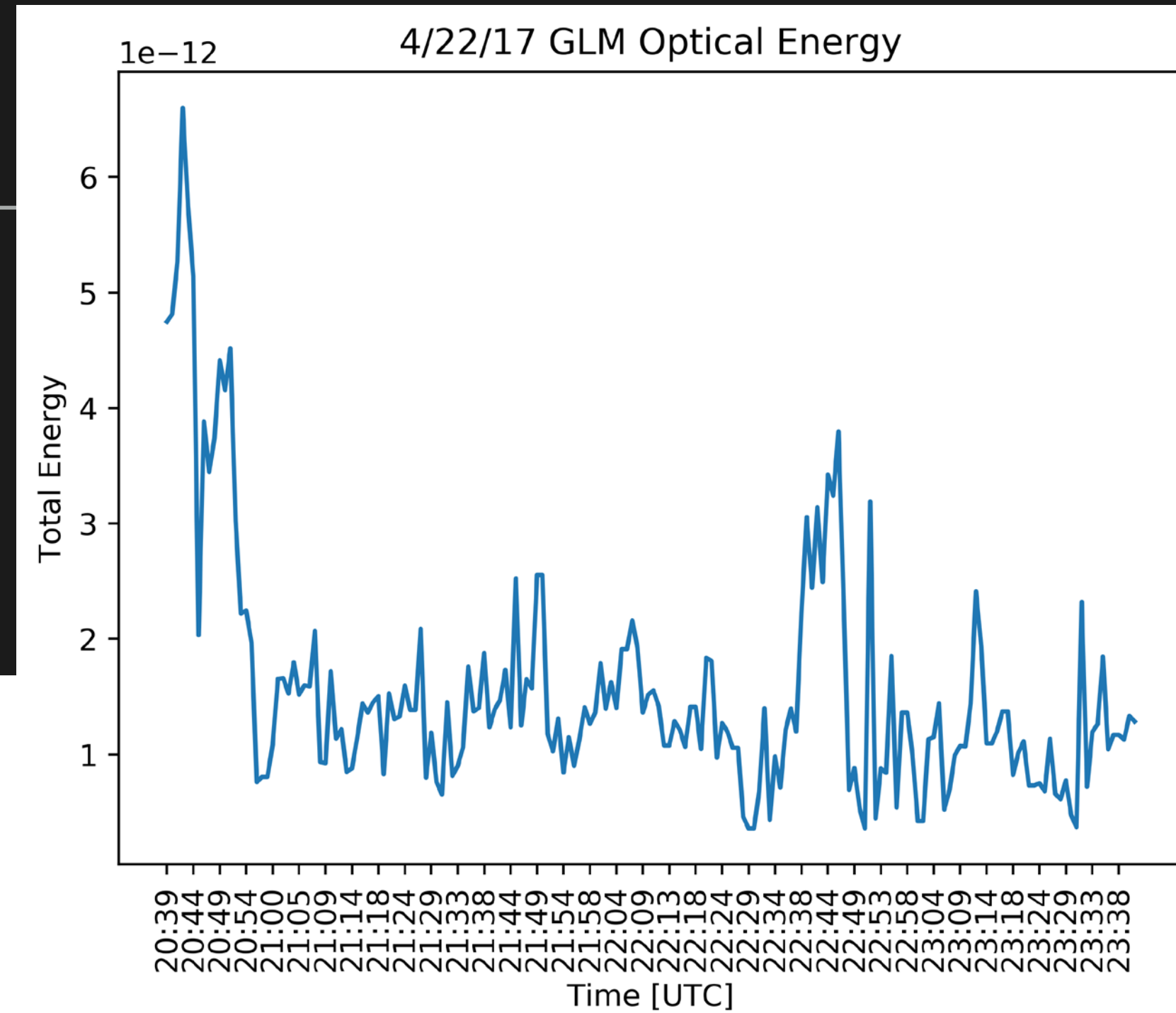
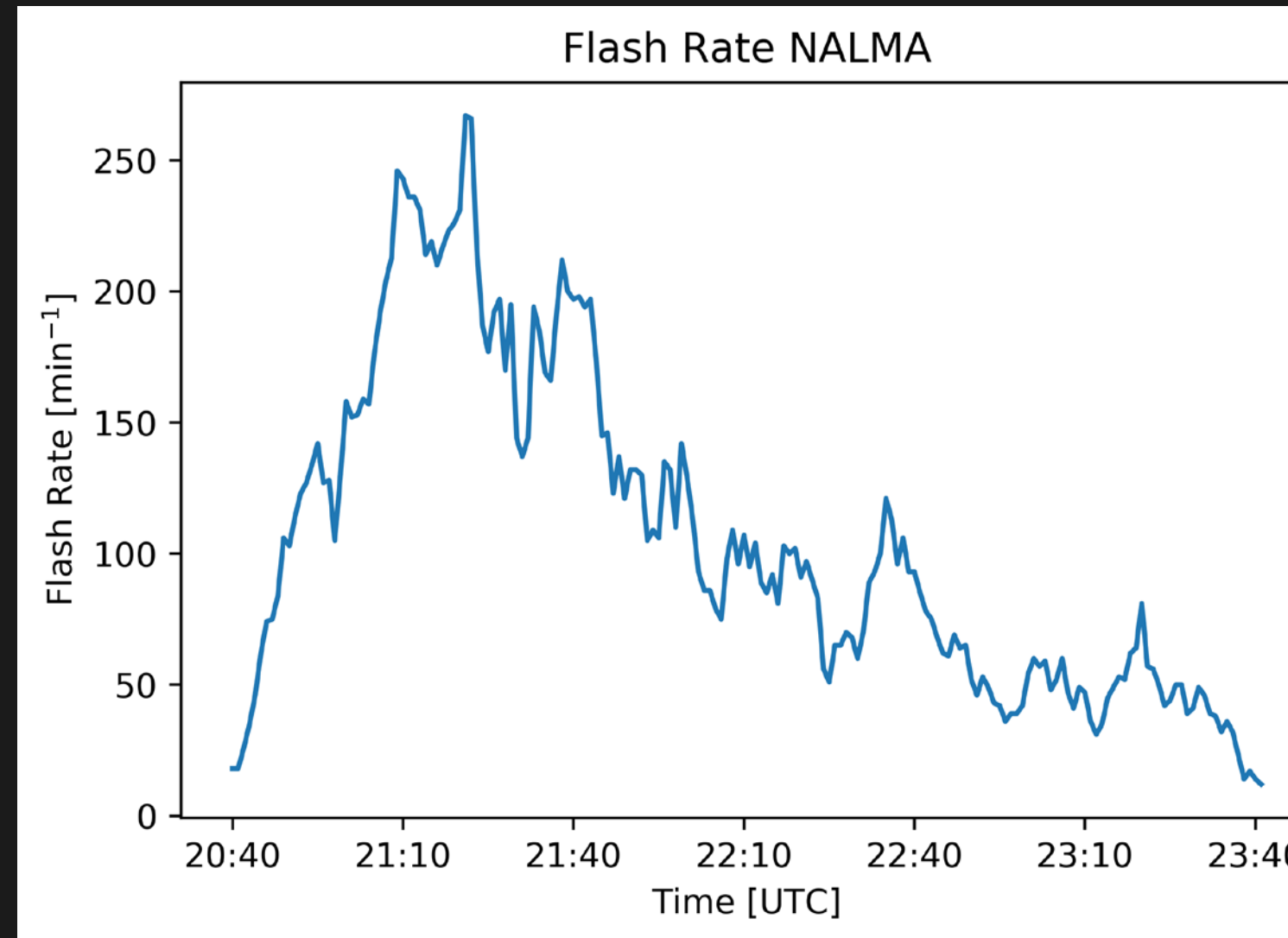
Optical Energy

- ▶ Optical energy high as storm enters range of NALMA, then dives down to a minimum for an extended period of time
- ▶ Optical energy remains low even during times of high flash rates in NALMA during period of severe hail
- ▶ Increase in optical energy associated with increase in flash rate around 2240 UTC, coincident with time of the EF-0 tornado



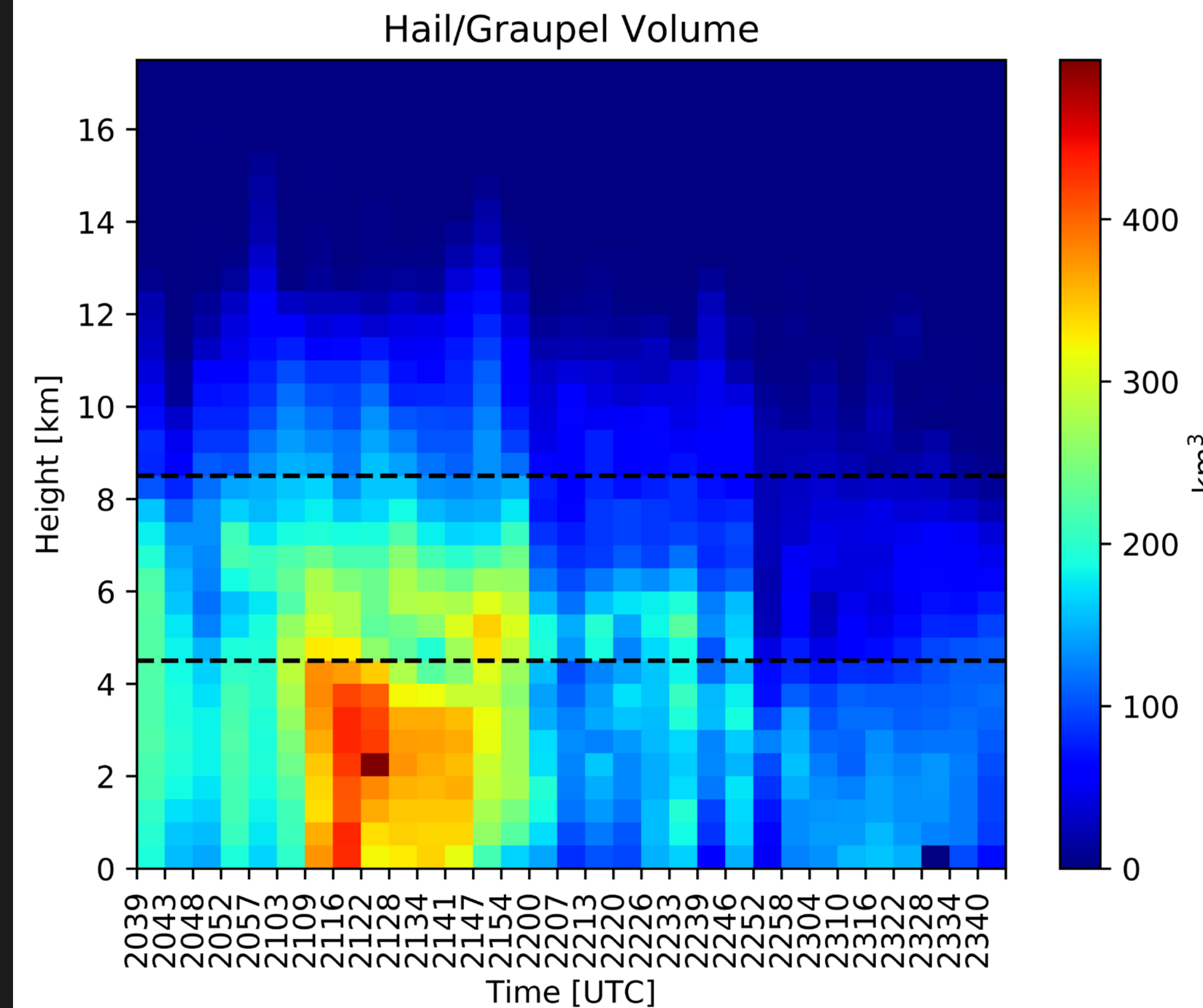
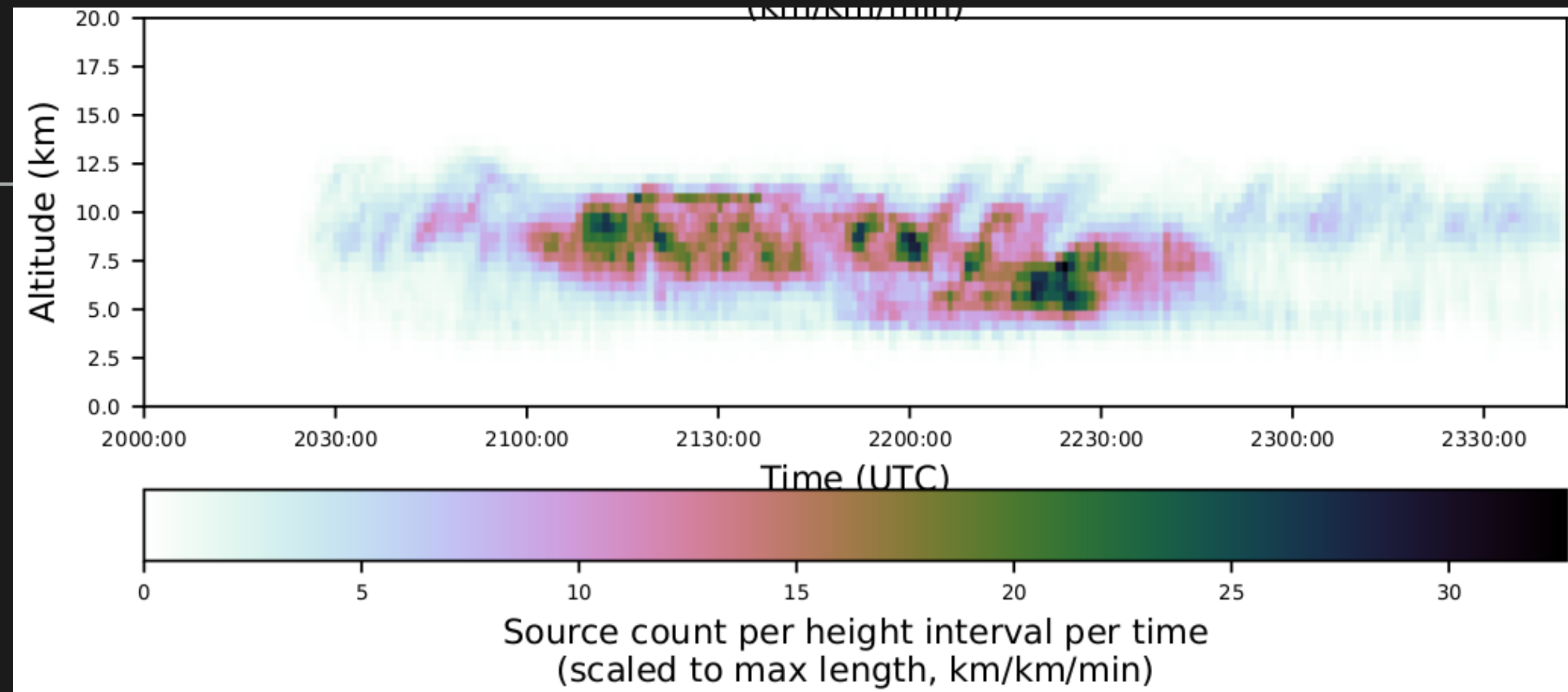
Graupel Volume vs Optical Energy

- ▶ Period of lowest optical energy from 2054-2234 UTC associated with highest hail/graupel volumes
- ▶ Peak of optical energy at 2238 UTC occurred with less hail/graupel
- ▶ However, optical energy does not increase as the hail/graupel volume decreases starting at 2200 UTC



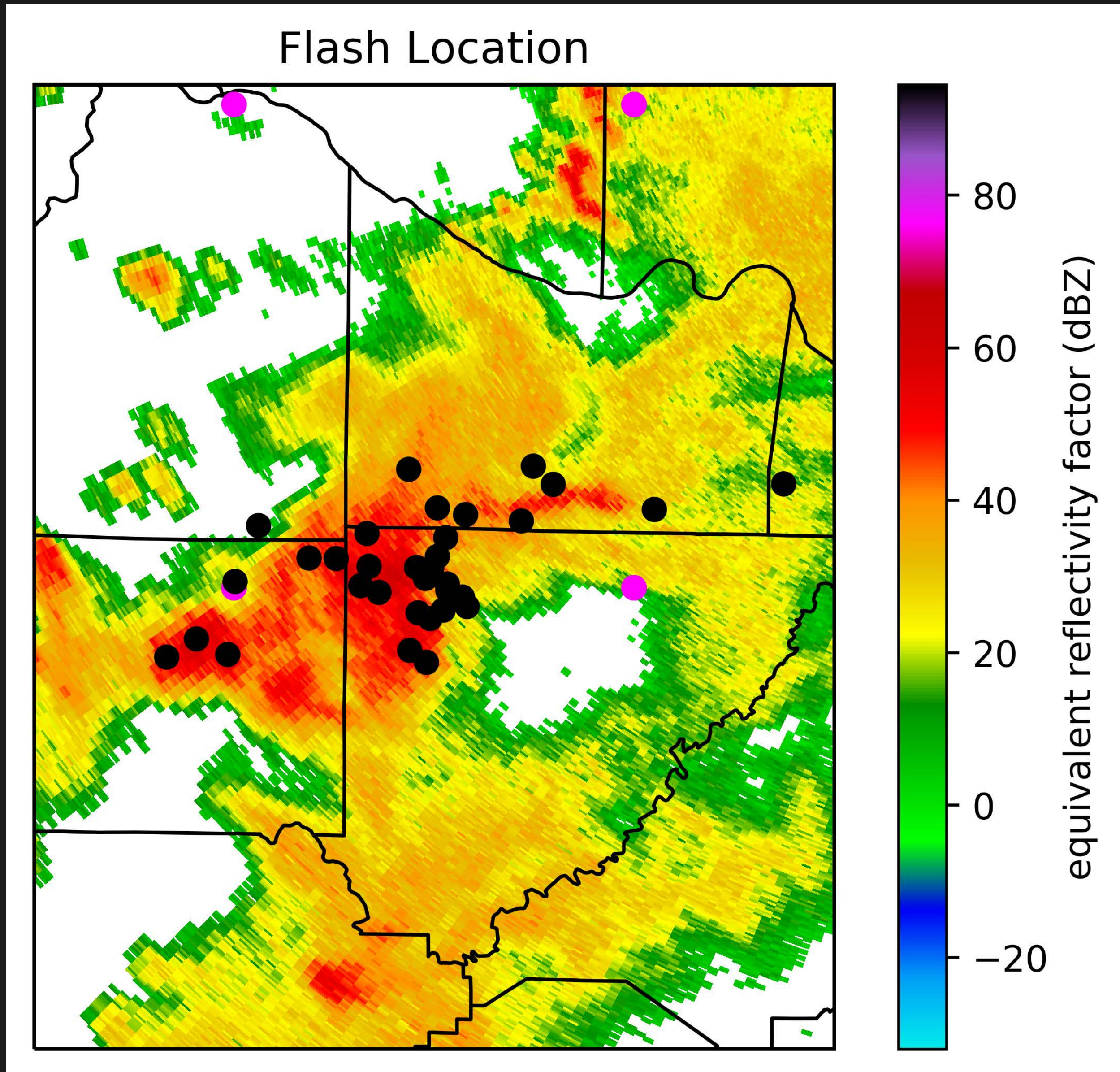
Lightning Height

- ▶ Majority of sources are located above higher regions of hail/graupel
- ▶ LMA sources descend prior to tornadogenesis, when optical energy and flash rate both increase
- ▶ Hail and graupel are not the only hydrometeors that may be contributing to optical extinction -> cloud water/ice have an impact



ISS LIS - Cullman Supercell

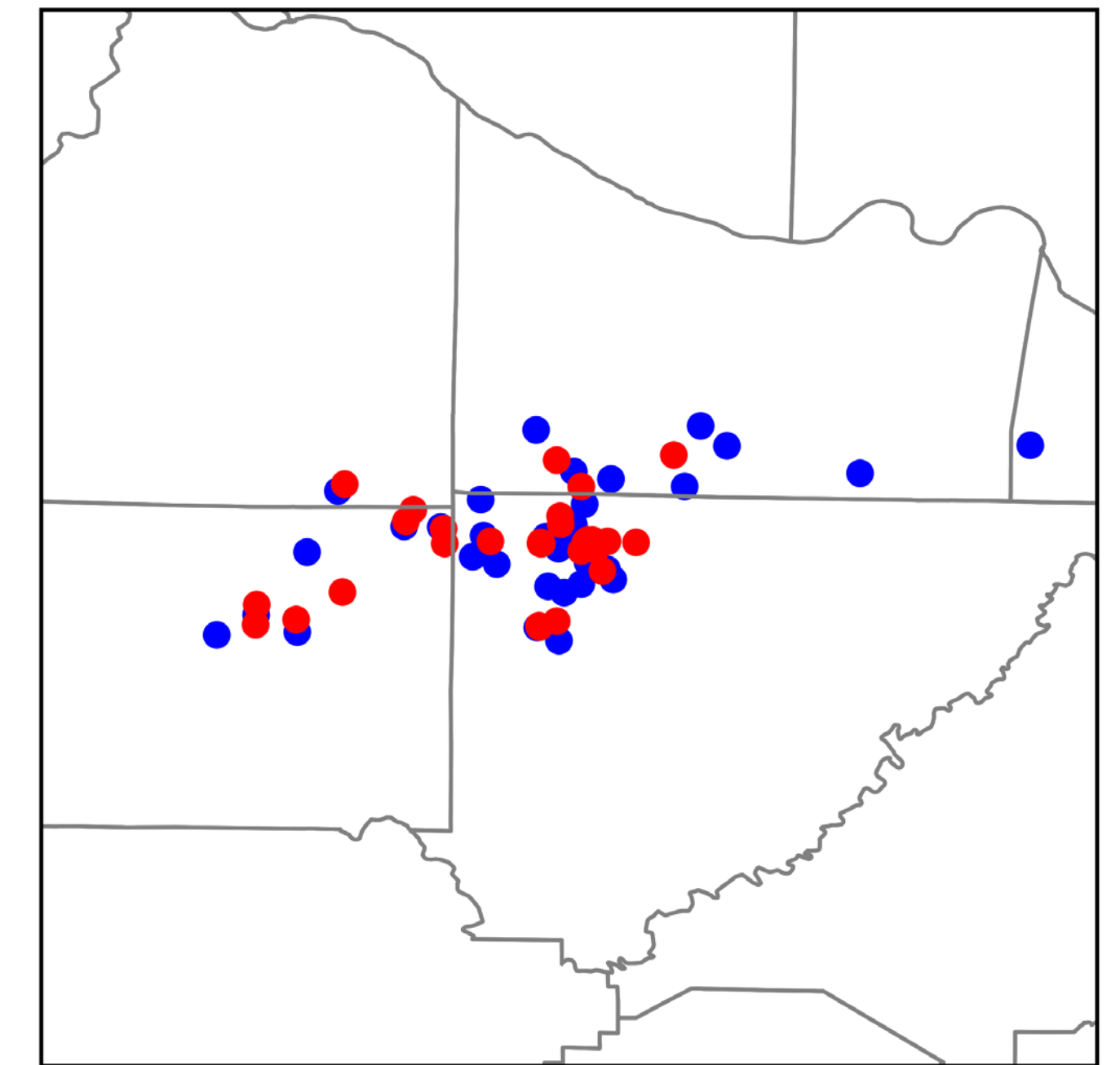
- ▶ ISS LIS overpass around 2245 UTC over supercell near Cullman County
- ▶ ISS LIS had 33 flashes during the 1 minute overpass, while GLM had 31 flashes; LMA showed 86 flashes
- ▶ Low Earth Orbit and Geostationary orbit showing similar flash rates, eliminating distance to the satellite as a possible means of error in underestimating flash rates for this case



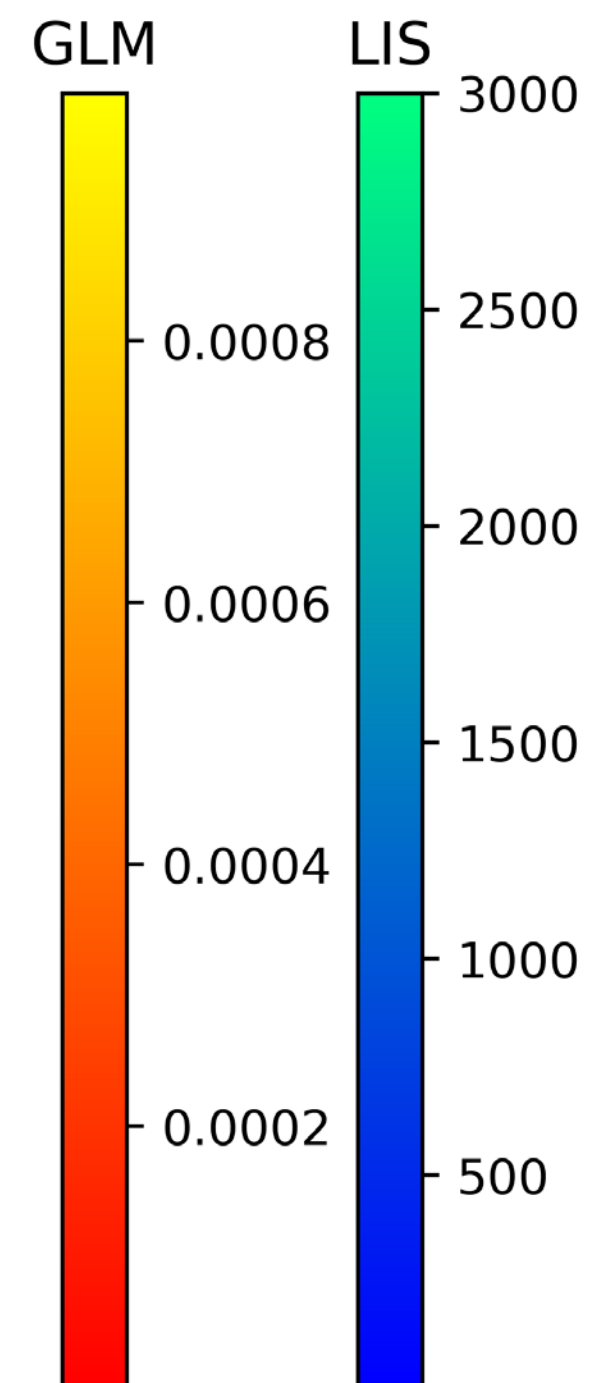
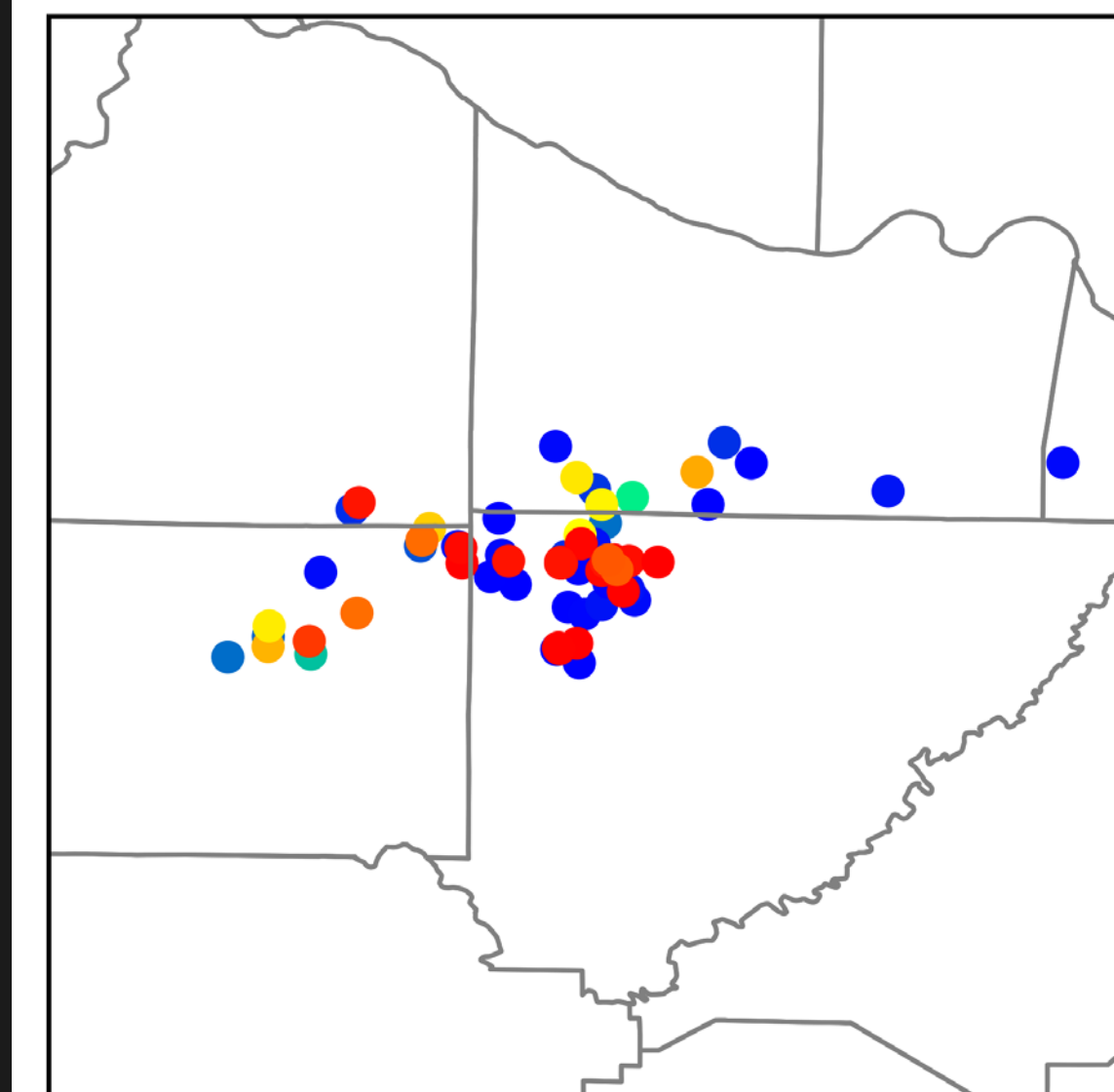
ISS LIS - Cullman Supercell

- ▶ GLM and LIS flashes located in similar regions of the storm, with similar flash rates
- ▶ LIS did show 2 flashes east of the storm, where GLM did not
- ▶ LIS and GLM energies for each flash show similar trends, with higher optical energies located in similar areas

LIS (blue) and GLM (red) Flash Centroid Location



Radiance for GLM and LIS



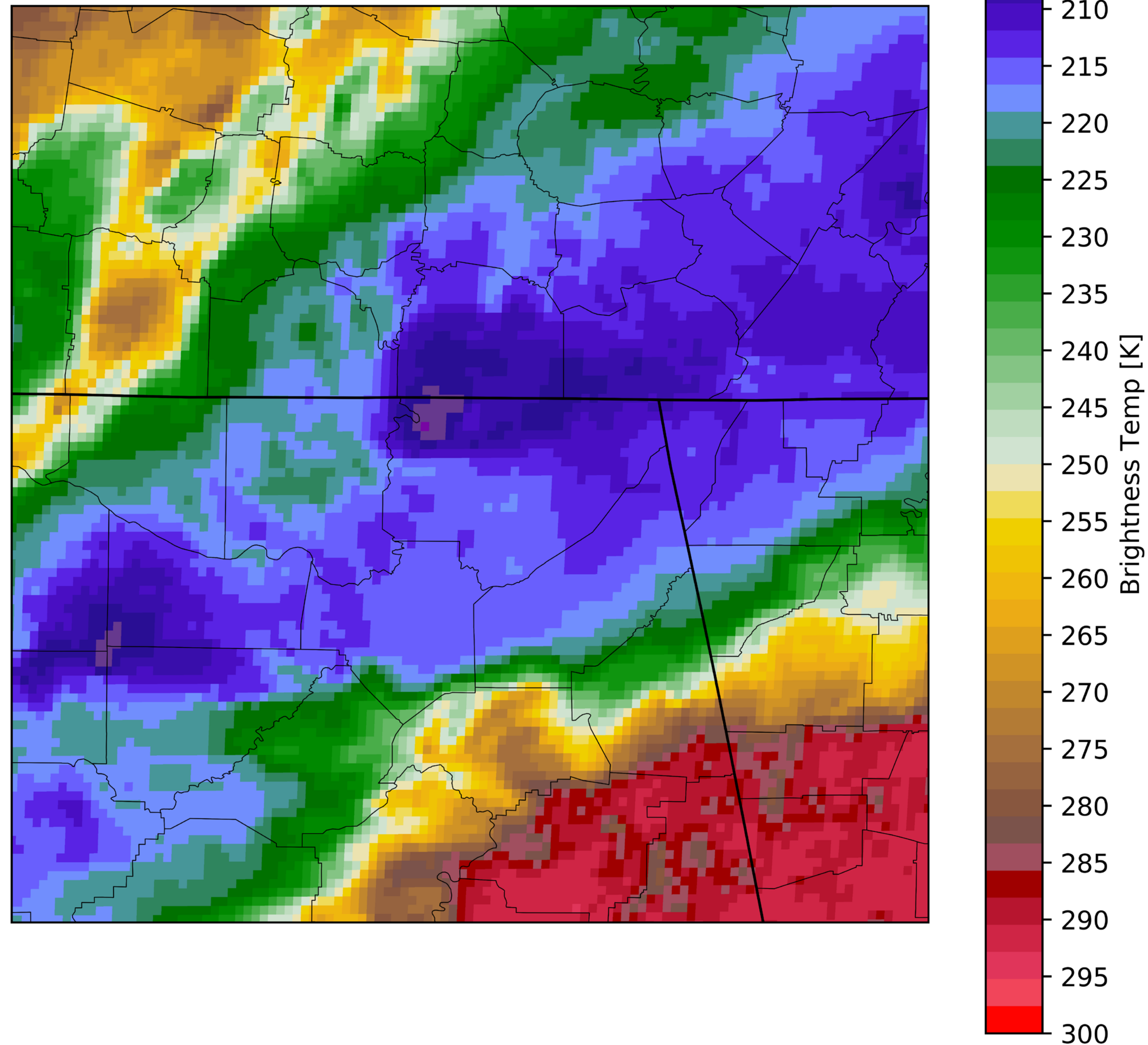
Conclusions

- ▶ Flash rates between LMA and space based observations differ, with satellite observations showing lower flash rates than LMA
- ▶ Updraft volume and optical energy may be inversely related, however it can not be confidently stated due to storm passing through dual-Doppler baseline
- ▶ Optical energy from GLM does not appear to be correlated with flash rates, from either LMA or GLM
- ▶ Optical energy remains low after hail/graupel volume decreases, however LMA sources are located above most of the hail/graupel
- ▶ ISS LIS and GLM both show reduced radiances within heavy precipitation, however further investigation into the flash properties of that storm is needed

Future Work

- ▶ Flash properties
- ▶ Cloud properties from GOES-16 ABI
- ▶ ER-2 Data
 - ▶ Electric field changes
 - ▶ FEGS
- ▶ Further investigation with ISS LIS
- ▶ P3 tail radar

2231 Z GOES-16 IR Brightness Temp



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- ▶ Contact: dconrad@nsstc.uah.edu