

NASA Small Business Innovative Research (SBIR) Subtopic S2.04

“X-Ray Mirror Systems Technology, Coating Technology: X-Ray, Ultraviolet (UV), Optical and Infrared (IR), and Free-Form Optics”

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Outline

- **Overview of NASA / GSFC Optics Branch recent missions**
- **SBIR Topic: S2. Advanced Telescope Systems**
- **SBIR Subtopic: S2.04**
 - **X-Ray Mirror Systems Technology**
 - **Optical Coatings from X-Ray, Extreme UV (EUV) to Optical and IR**
 - **Free-Form Optics Design, Manufacturing and Metrology**

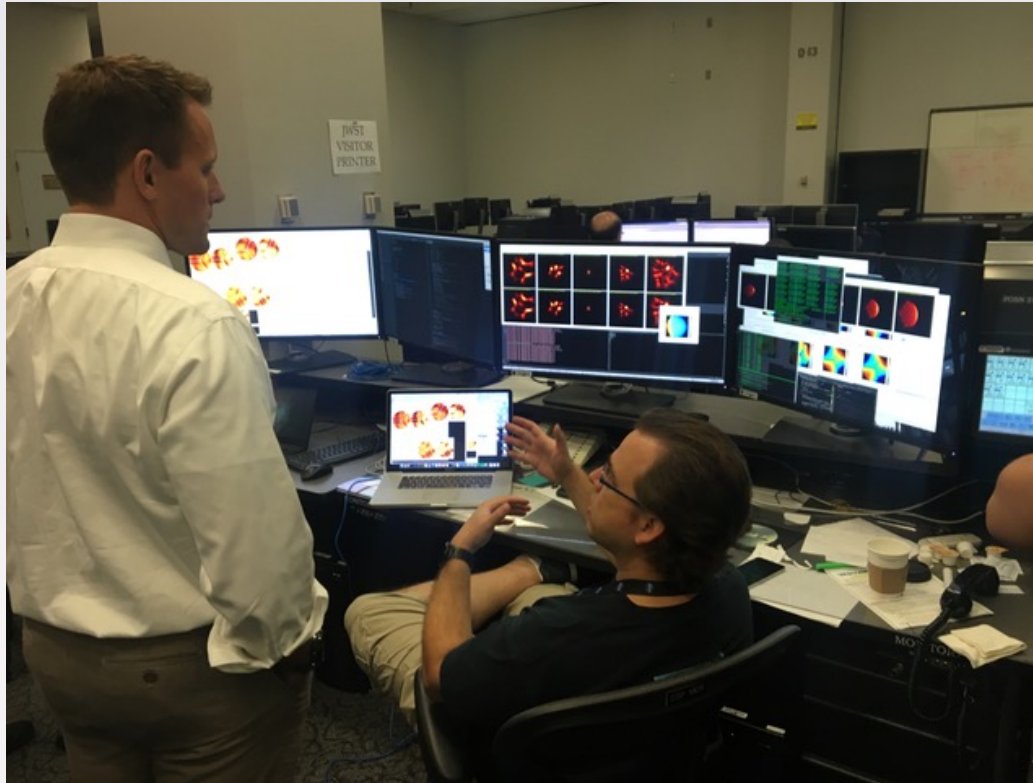


Goddard Optics Branch, Code 551

- **JWST Wavefront Sensing Commissioning**
- **WFIRST**
- **GEDI**
- **OCI**
- **Lucy-Ralph**
- **OSIRIS-REx**
- **FUV Coating**
- **Black-Silicon Coating**
- **Free-form Optics**



James Web Space Telescope Wavefront Sensing and Control Commissioning



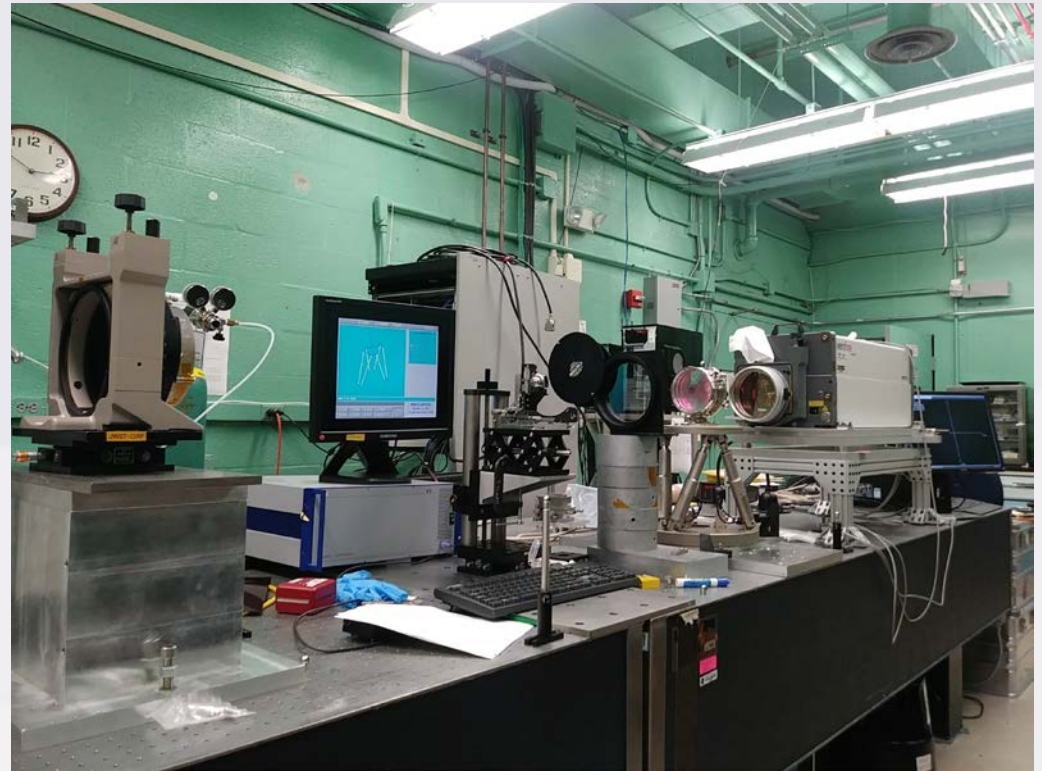
Phase Retrieval Metrology System (PRMS) software to be used during JWST Wavefront Sensing Commission



Wide Field Infrared Space Telescope (WFIRST) Grism Testing



Chamber to perform interferometric testing of grism prototype under 150K

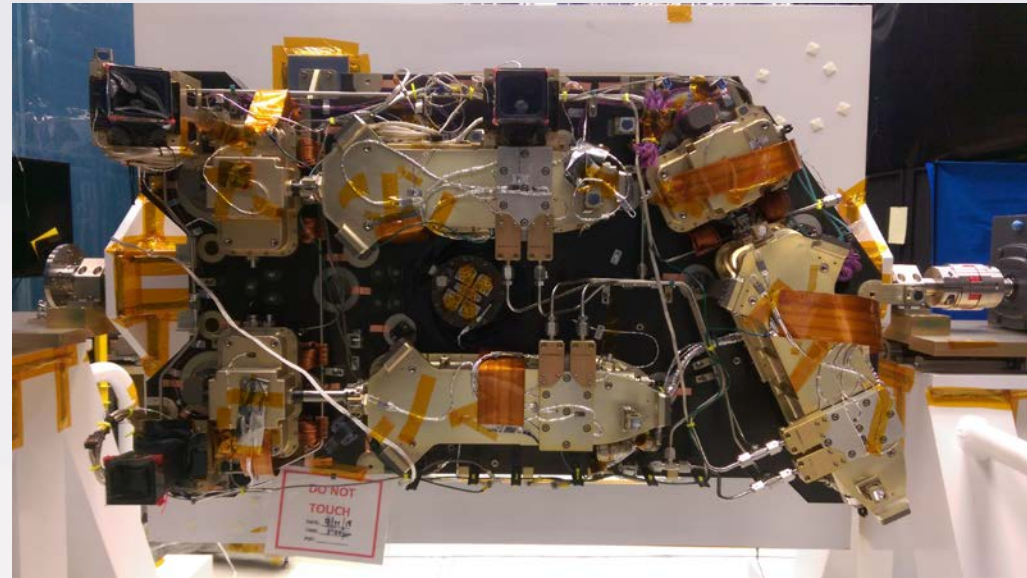


Interferometric ambient testing of the grism prototype for Wide Field Infrared Space Telescope (WFIRST)

Global Ecosystem Dynamic Investigation (GEDI)



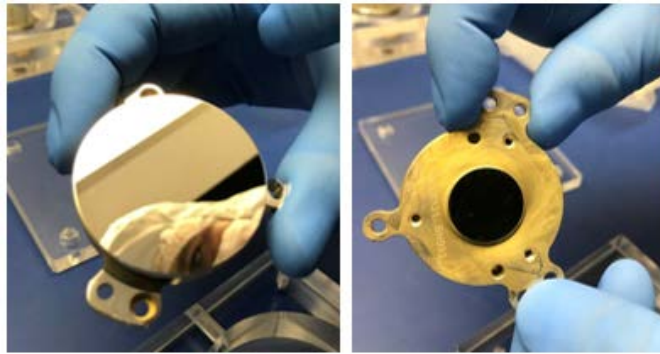
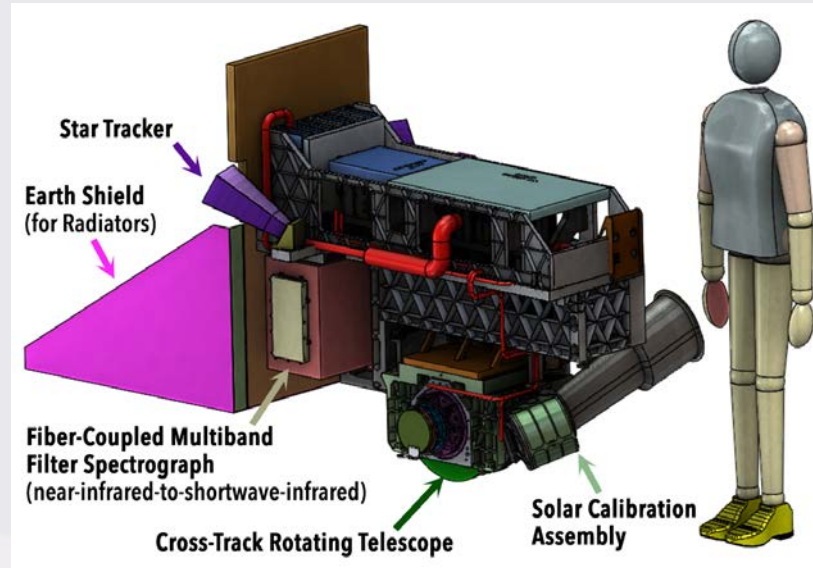
Global Ecosystem Dynamic Investigation (GEDI) optical bench installed in the GEDI Turnover Dolly. RTA (Receiver Telescope Assembly) and transmitter beam expanders seen



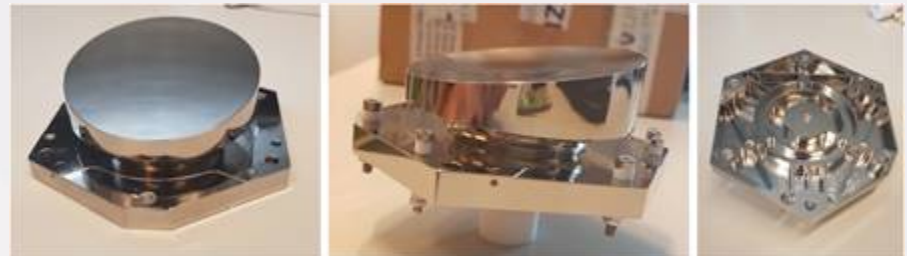
GEDI Optical bench seen from minus Z-side populated with lasers, beam dithering units and fold mirrors

Ocean Color Instrument (OCI)

Plankton, Aerosol, Cloud, ocean Ecosystem (PACE) primary sensor



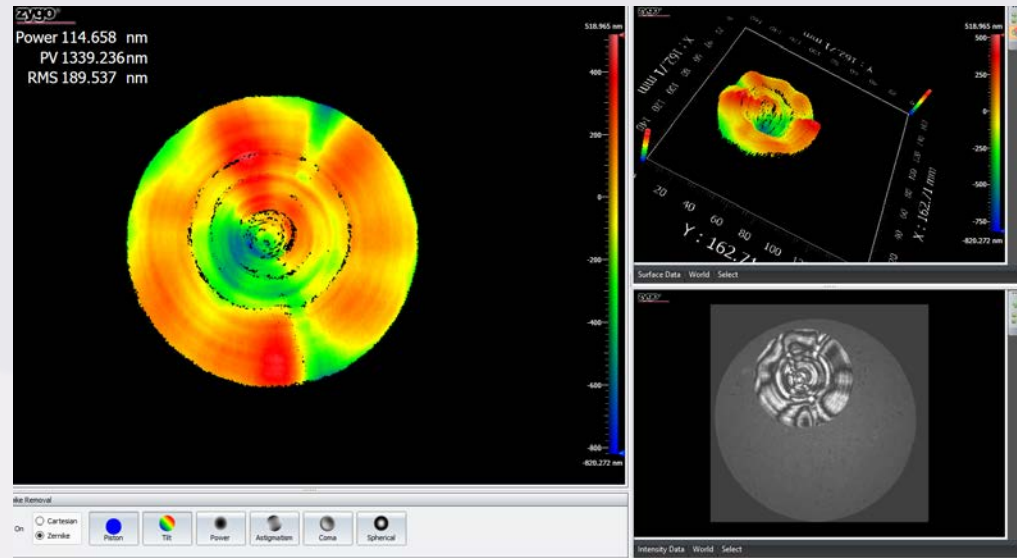
A couple of the fold flats



Collimator mirror images



Lucy-Ralph Three Mirror Anastigmat Alignment

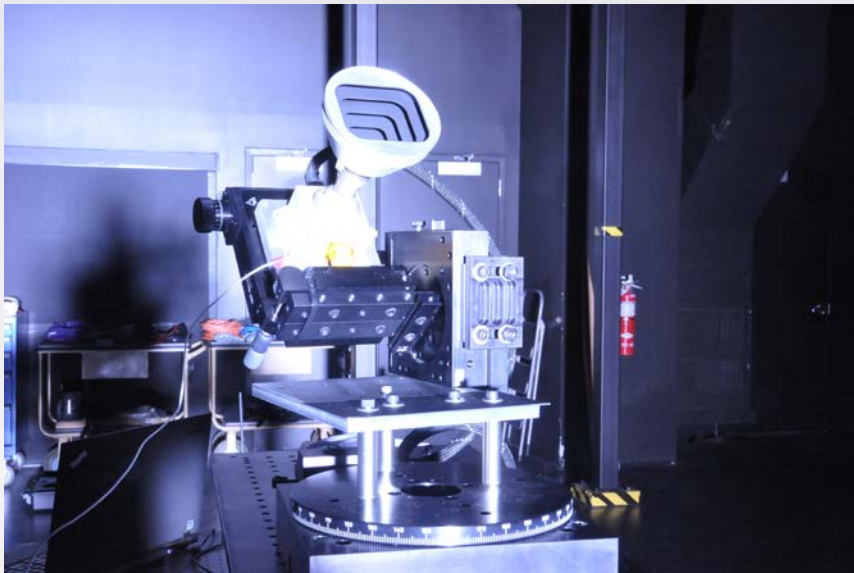


Lucy Ralph breadboard with fixed M2 at front of Zygo interferometer; non-contact Coordinate Measurement Machine (CMM) combined as built prescription and tooling ball fiducials

Wavefront after 2nd iteration Rx, Ry, and tooling ball focus adjustments made



Origins-Spectral Interpretation-Resource Identification Security--Regolith Explorer (OSIRIS-Rex)

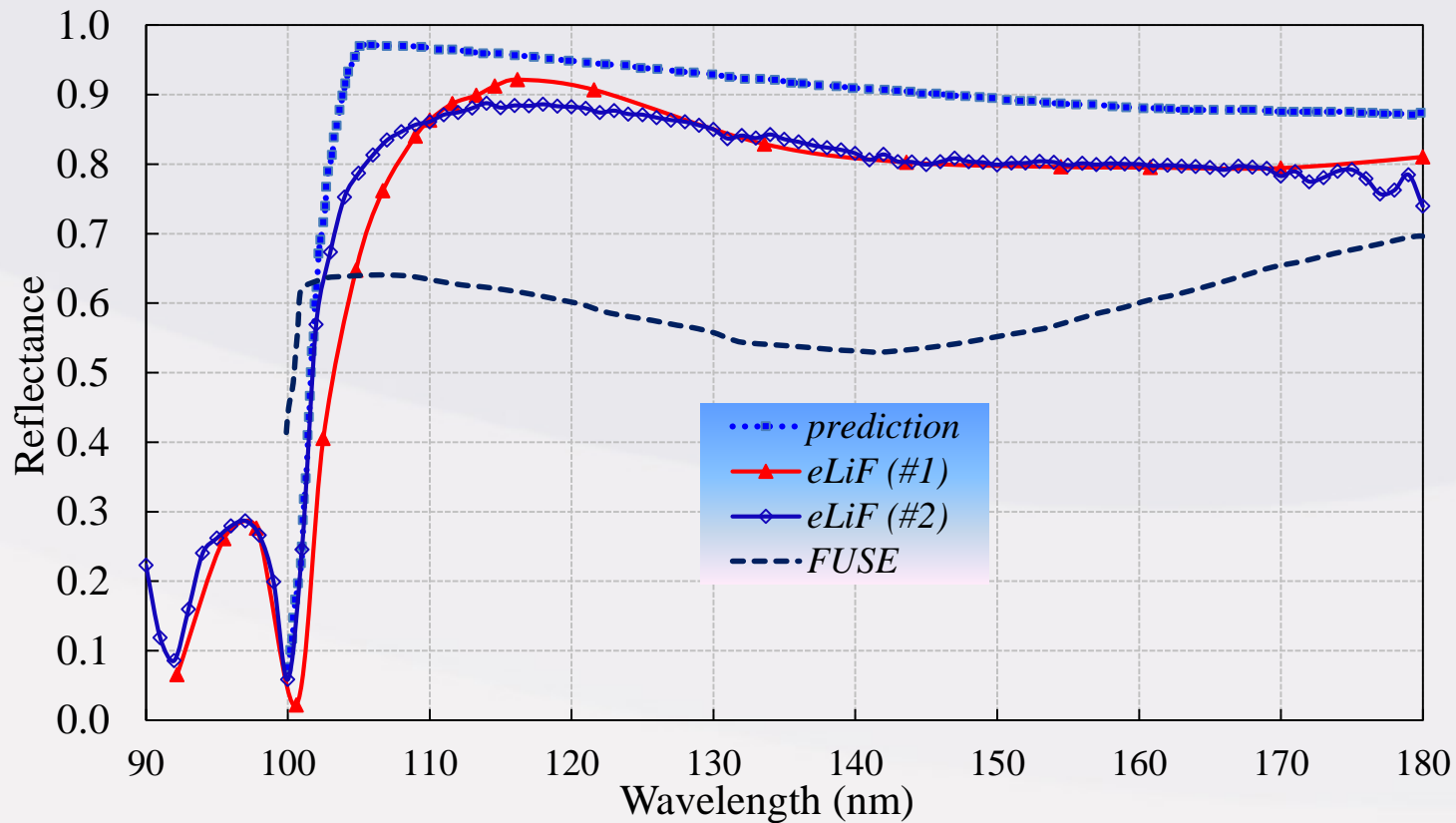


OSIRIS-Rex NavCam-1, a panchromatic imager, is part of the TAGCAMS (Touch-And-Go Camera System) navigation camera suite.



The OSIRIS-REx spacecraft captured this image using its PolyCam imager. The close up image of Bennu was taken on October 29th from 205 miles away.

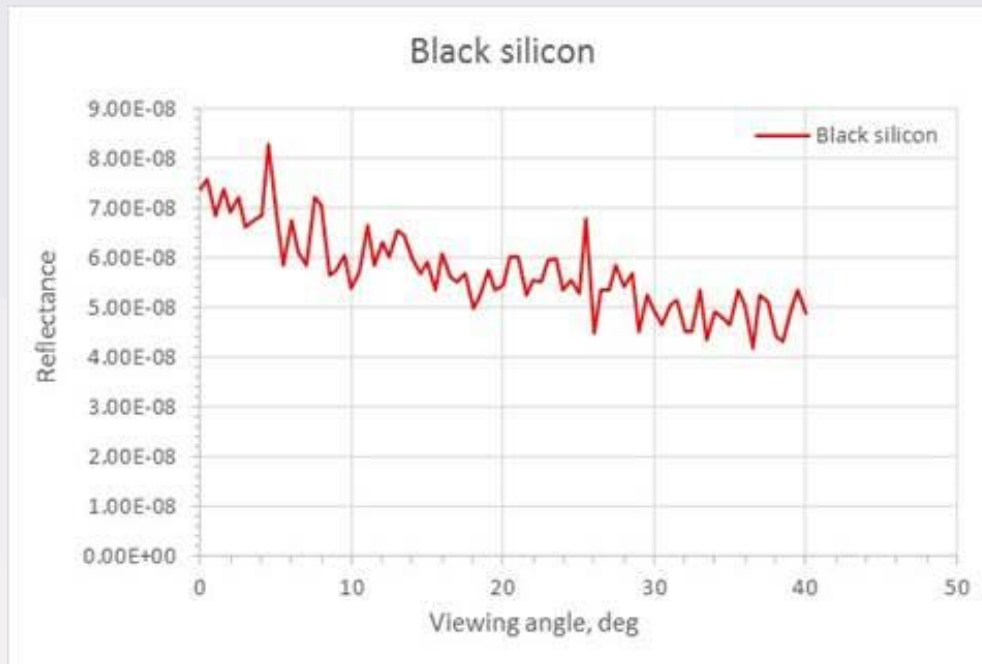
Optimization Al+LiF (eLiF) Hot Coatings



The 2-meter chamber was used to successfully the 0.5-meter SISTINE primary (Sounder Rocket project led by PI Kevin France from U. of Colorado).



Black Silicon Coating



Reflectance of uniform black-silicon coating on silicon wafer etched cryogenically. The specular reflectance is 6E-8 on incident beam of 632.8 nm

Freeform optics

Large UV/Optical/Infrared Surveyor (LUVOIR) Instruments

1. LUVOIR Ultraviolet Multi Object Spectrograph (LUMOS)

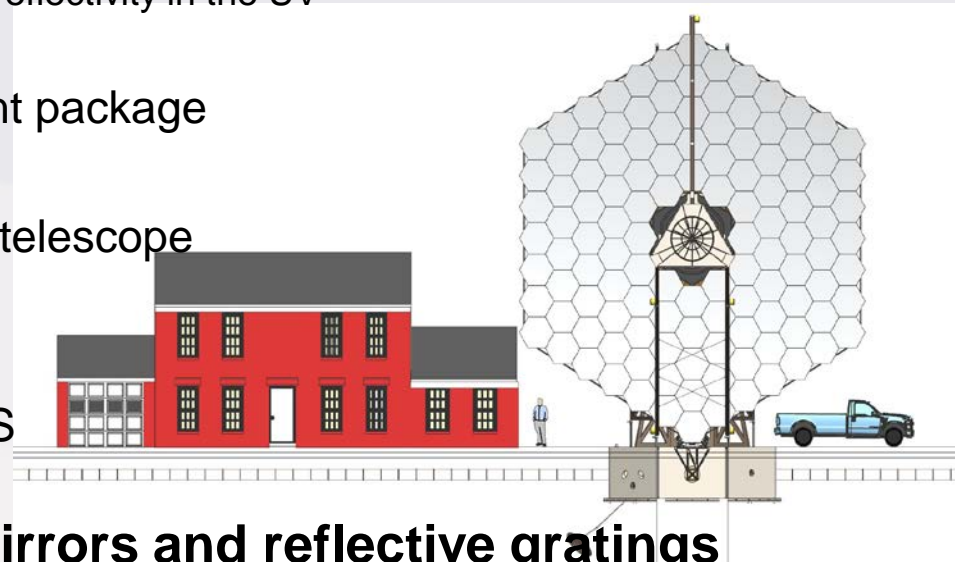
- Fewer mirrors
 - Throughput is limited by the Al coating reflectivity in the UV
- Improve image quality in UV
- Reduce volume for limited instrument package

2. Coronagraph

- Correct aberrated off-axis field from telescope

3. High Definition imager (HDI)

- Correct wide field of view
- Improve image quality in UV and VIS
- Reduce volume and mass



NASA Needs: UV grade freeform mirrors and reflective gratings



SBIR Subtopic S2.04

- **X-Ray Mirror Systems and Components Technology**
- **Optical Coatings from X-Ray, EUV to Optical and IR**
- **Free-Form Optics Manufacturing and Metrology**



X-Ray Mirror Systems Technology

- Optical Components, systems, and instrument for X-Ray missions
- Light-weight, low-cost, ultra-stable mirrors for large X-Ray observatories
- Stray-light suppression systems (baffles) for large advanced X-Ray observatories
- **Horizon:** 1 to 3 years, mature the technology in advance of decadal 2020 proposal call
- **State of Art:** costly and time consuming to produce X-Ray mirrors. Require improvement to about 0.5-1.0 arc-seconds of angular resolution
- The current stray light suppression is bulky and ineffective for wide-field of view telescopes. We seek significant reduction in both expense and time
- **Importance:** Very-high value, critical need where no feasible competitor and only government is the major player in this technology



Mirror Technology Days 2018

Subtopic: (S2.04, X-Ray Mirror Systems Technology, Coating Technology for X-Ray-UV-OIR, and Free-Form Optics)
Manager: (Ron Shiri / GSFC, Kunjithapatham Balasubramanian / JPL, Philip Stahl / MSFC, David Broadway/ MSFC)
Center(s): (GSFC, JPL, MSFC)

Optical Components, Systems, and Stray Light Suppression for X-Ray Missions

- Light-weight, low-cost, ultra-stable mirrors for large X-Ray observatory
- Epoxies to bond silicon mirrors to absorb IR radiation between 1.5 to 6 microns that traverses silicon with little or no absorption and cured quickly
- Stray light suppression systems (baffles) for large advanced X-Ray observatories
- Ultra-stable low-cost light-weight X-Ray telescope using grazing-incidence optics for high altitude balloon-borne and rocket-borne mission

Science Traceability

The 2010 National Academy Decadal Report specifically identifies optical components and the ability to manufacture and perform precise metrology on them needed to enable several different future missions (LYNX)

The NRC NASA Technology Roadmap Assessment ranked advanced mirror technology for new x-ray telescopes as the #1 Object C technology requiring NASA investment.

Need Horizon

1 to 3 years, Need to mature technology in advance of proposal Decadal 2020

State of Art

It's very costly and time consuming to produce X-Ray mirrors. Most of SOA requiring improvement is 0.5-1.0 arc-seconds angular resolution. SOA stray light suppression is bulky and ineffective for wide-field of view telescopes. We seek significant reduction in both expense and time. Reduce the areal cost of telescope by 2X such that the larger collecting area can be produced for the same cost or half the cost.

Importance

Very high – Critical need, no feasible competitors. X-Ray mirror technology is inherently in government. There is no commercial application.



Coating Technology: X-Ray, Extreme UV to Visible and IR

- Metrics for X-Ray:
 - Multilayer high-reflectance coatings for hard X-Ray mirrors
 - Multilayer depth gradient coatings for 5 to 80 keV with high broadband reflectivity
 - Zero-net stress coating of iridium or other high reflectance elements on thin substrates (< 0.5 mm)
- Metric for EUV:
 - Reflectivity greater than 90% from 6 nm to 200 nm and depositable onto < 2 meter mirror substrate
- Metric for UVOIR:
 - Broadband reflectivity greater than 60% and uniform polarization from 90 nm to 2500 nm and depositable onto 2, 4, and 8 meter mirror substrate
- Non-Stationary Optical Coating:
 - Used in both reflection transmission that vary with location on the optical surface. The variation refers to ratio of reflectivity transmissivity, optical field amplitude , phase, and polarization change.
 - The optical surface range of diameter is 0.5 cm to 6 cm that could either be flat, conic or free-form



Coating Technology: X-Ray, Extreme UV to Visible and IR (Continued)

- **Horizon:** 1 to 3 years, mature the technology in advance of decadal 2020 proposal call
- **State of Art:** costly and time consuming to produce X-Ray mirrors. Require improvement to about 10 arc-seconds of angular resolution
- The current stray light suppression is balky and ineffective for wide-field of view telescopes. We seek significant reduction in both expense and time
- **Importance:** Very-high value, critical need where no feasible competitor and only government is the major player in this technology



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Optical Coatings for X-Ray, EUV, LUV, UV, Visible, and IR Telescopes

- Meet low temperature operation requirement

Metrics for X-Ray:

- Multilayer high-reflectance coatings for hard X-Ray mirrors

- Multilayer Depth Gradient Coatings with a reflective spectral response from 1 to 150 KeV or higher

- Near-zero-residual stress single & multilayer thin-film coatings on thin substrates (< 0.5 mm)

Metrics for EUV:

- Meet temperature requirement, 35 Kelvin
- Reflectivity > 90% from 6 nm to 90 nm onto a < 2 meter mirror substrate

Metrics for LUV/OIR:

- Meet temperature requirement, 35 Kelvin
- Broadband Reflectivity > 70% from 90nm-120nm (LUV) and > 90% from 120nm-2.5um (VUV/Visible/IR). Reflectivity Non-uniformity < 1% 90nm-2.5um
- Induced polarization aberration < 1% 400nm-2.5um depositable onto a 1-8m substrate

Non-stationary Optical Coatings:

- Used in reflection & transmission that vary with location on the optical surface

Carbon Nanotube (CNT) Coatings:

- Broadband Visible to NIR, Reflectivity of 0.1% or less, adhere to the multi-layer dielectric or protected metal coating

Science Traceability

Astrophysics Decadal specifically calls for optical coating technology investment for: Future UV/Optical and Exoplanet missions (HabEX or LUVOIR)
 Heliophysics 2009 Roadmap identifies optical coating technology investments for: Origins of Near-Earth Plasma (ONEP); Ion-Neutral Coupling in the Atmosphere (INCA); Dynamic Geospace Coupling (DGC); Fine-scale Advanced Coronal Transition-Region Spectrograph (FACTS); Reconnection and Micro-scale (RAM); & Solar-C
 Nulling polarimetry/coronagraph for exoplanet imaging and characterization, dust and debris disks, extra-galactic studies and relativistic and non-relativistic jet studies

Need Horizon

1 to 3 years
 Affordable high-performance optical component system technology needs to achieve TRL-4-5 by approximately 2018 to support the 2020 Astrophysics Decadal process. Heliophysics missions need mirror technology sooner. Historically, it takes 10 years to mature mirror technology from TRL-3 to 6. To achieve these objectives requires sustained systematic investment. 1 to 5 years for CNT coating applications.

State of Art

Current X-Ray is defined by NuSTAR and Chandra
 Current EUV is defined by Heliophysics (80% reflectivity from 60 to 200 nm)
 Current UVOIR is defined by Hubble. MgF2 over-coated Aluminum on a 2.4 meter mirror. This coating has birefringence concerns and a marginally acceptable reflectivity between 100 and 200 nm.

Importance

Very High – optical technology is mission enabling. The technical capabilities of the optical systems will determine performance and science return.



Free-Form Optics: Design, Manufacturing, Metrology

- Freeform Optical Surfaces
 - 0.5 cm to 6 cm diameter optical surfaces (mirrors) with free form optical prescriptions with surface tolerances about 1-2 nm rms
 - Freeform refers to either 2nd order conic prescription with higher order surface polished onto it or without underlying conic prescription but such that is no steps in the surface.
 - The optics with underlying conic prescription would need to be in $F/\#$ range of $F/2$ or $F/20$
- Metrology of Freeform Optics
 - Component metrology is difficult because of very large departure from the planar or spherical shapes that can be accommodated by conventional interferometric testing
 - New Methods such as multibeam low-coherence optical probe and slope sensitive optical probe are highly desirable
- **Horizon:** 3 to 5 years
- **State of Art:** Never been done before
- **Importance:** Highly desirable, allows efficient, small package, and lower cost that expands operational temperature range in un-observed system. It allows coronagraphic nulling without shearing and increases the useful science field-of-view



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<p>Free-Form Optical Surfaces</p> <ul style="list-style-type: none"> -0.5 cm to 6 cm diameter optical surfaces (mirrors) with free form optical prescriptions with surface tolerances are 1-2 nm rms. -The Optics with large field of view and fast F/#s to provide non-rotationally symmetry. Optical freeform surfaces enabling additional degrees of freedom to reduce volume and eliminate traditional design constraints on the surface. -Metrology of 'freeform' optical components is difficult. New methods such as multibeam low-coherence optical probe and slope sensitive optical probe are highly desirable -Optics design, fabrication, and metrology for package constrained imaging systems -Integrated SWIR imaging spectrometer 	<p>Science Traceability</p> <p>NASA missions with alternative low-cost science and small size payload are increasing. However, the traditional interferometric testing as a means of metrology are unsuited to freeform optical surfaces due to changing curvature and lack of symmetry. Metrology techniques for large fields of view and fast F/#s in small size instruments is highly desirable specifically if they could enable cost-effective manufacturing of these surfaces for constrained imaging systems. (CubeSat, SmallSat, NanoSat, various coronagraphic instruments)</p>	
	<p>Need Horizon</p> <p>3 to 5 years</p>	
	<p>State of Art</p> <p>Early stages of development. Improve optical surfaces with large field of view and fast F/#s. Reduce design and fabrication cost of freeform surfaces.</p>	
	<p>Importance</p> <p>High – Highly desirable, allows efficient, small package, and lower cost that expands operational temperature range in un-observed system. Can allow new coronagraphic instruments that adhere to high-contrast imaging while maintains high throughput.</p>	



Conclusion

- GSFC has a robust and productive SBIR program in the Optics, with high quality proposals being submitted every year, leading to advances in key Optics Technologies. Companies with successful SBIR efforts have submitted high quality New Technology Reports (NTRs)
- Focus areas,
 - X-Ray Optical Systems, Mirrors, Coating, and Components
 - Optical Coating from X-Ray to UV + Optical + IR
 - Freeform Optics Design, Development, and Metrology

