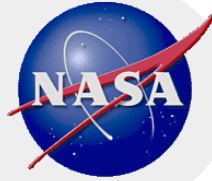


Tutorial :

An Overview of the Orbital Debris

Environment

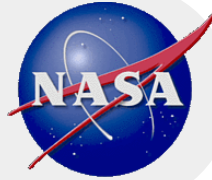
Mark Matney, Ph.D.
Orbital Debris Program Office
NASA Johnson Space Center



Space Debris

- **Space debris = any man-made object in space that no longer serves a useful purpose**
 - Note there are also natural debris – meteoroids
- **Intact objects, > 1 m**
 - Old rocket bodies and spacecraft
 - “Operational” debris – shrouds, mounts, lens caps, etc
- **Fragmentation debris, 1 mm – 1 m**
 - **Deliberate or accidental explosions from on-board energy sources**
 - Unvented rocket fuel
 - Active batteries
 - Self-destruct mechanisms
 - **Deliberate or accidental collisions**
 - Weapons tests
 - Random collisions
 - **Solid rocket motor slag**
- **Small debris, < 1 mm**
 - **Deterioration of satellite surfaces in space environment**
 - Small debris impact ejecta
 - Deterioration of paint and other materials





Brewster Rokit on Debris Sources



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www.gocomics.com/brewsterrokit brewrockit@yahoo.com



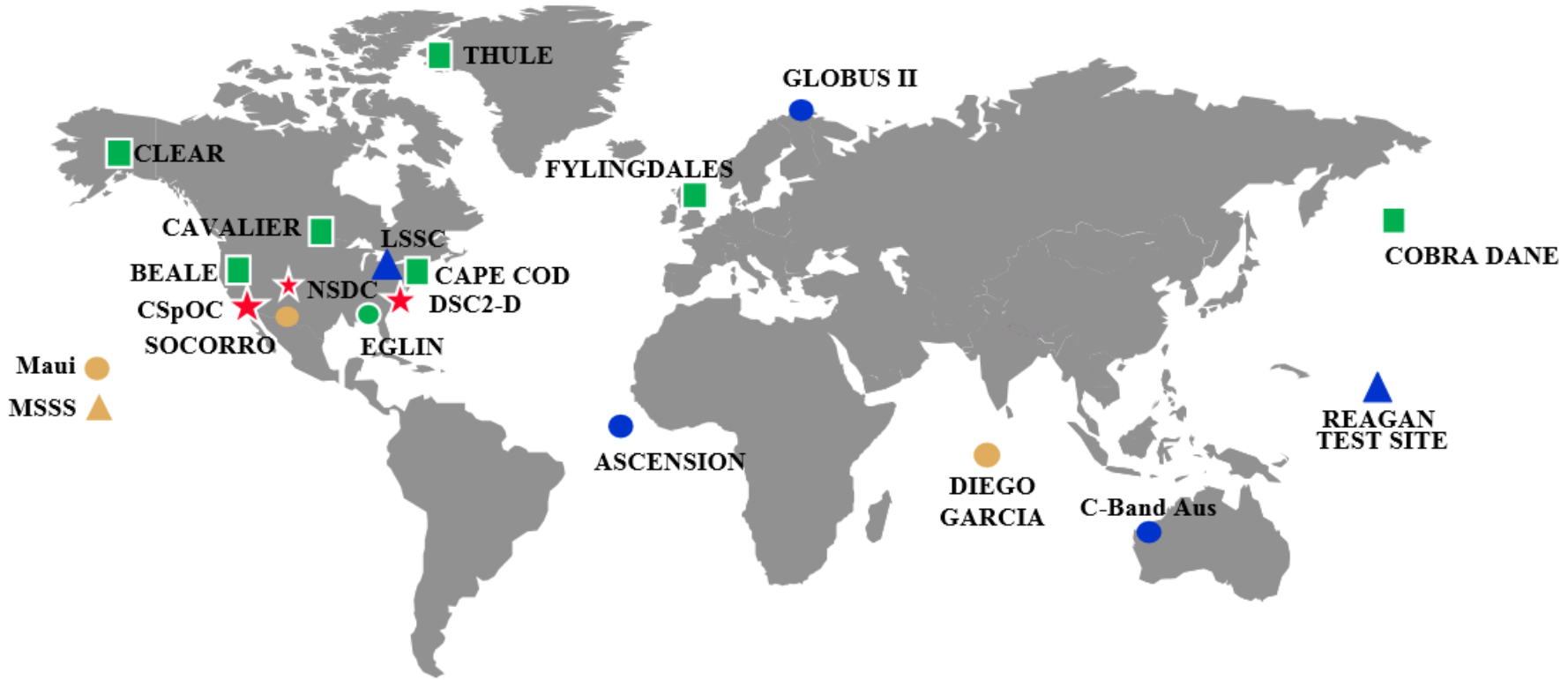
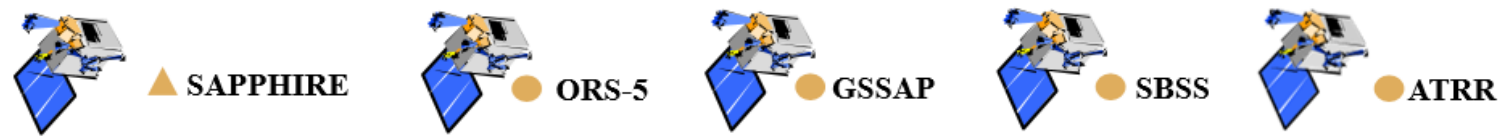


Space Surveillance Network (SSN)

- **Almost all of our operational knowledge of the space environment is from the U.S. Department of Defense's (DoD) Space Surveillance Network (SSN) and its parallels in other countries**
 - **New launches**
 - Payloads
 - Rocket Bodies
 - Operational Debris (brackets, shrouds, etc.)
 - **Breakup events**
 - Anomalous breakups
 - Explosions – both accidental and deliberate
 - Collisions between tracked objects
- **It is through the SSN Catalog that we know the various orbits where humans have launched their satellites and how they have evolved over time**



Today's SSN

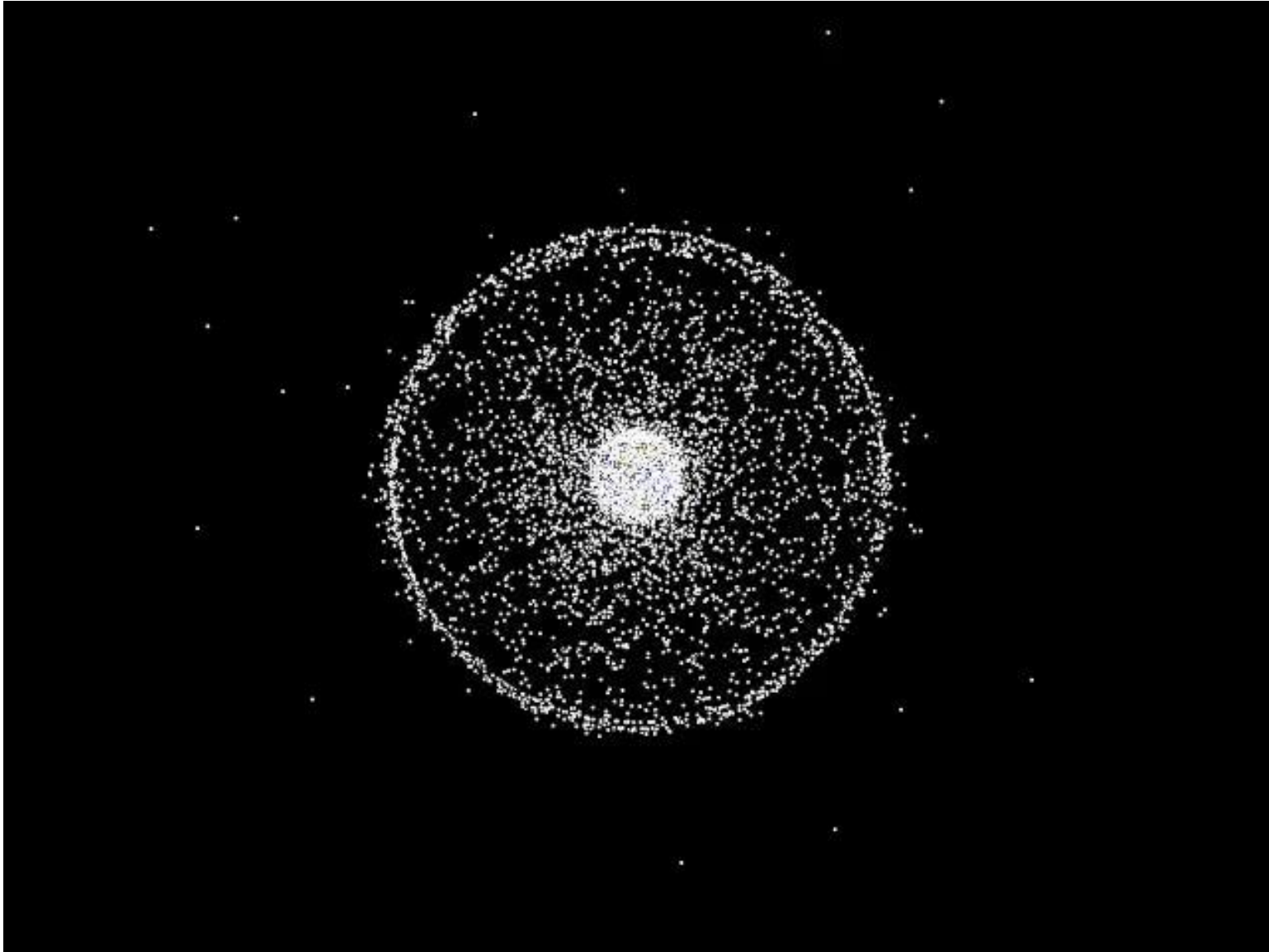


- Phased Array Radar
- Mechanical Radar
- Optical Systems
- SSN C2

- Dedicated - subordinate to USSTRATCOM, primary mission of SSA
- Collateral - subordinate to USSTRATCOM, primary mission other than SSA
- ▲ Contributing – under contract or agreement to support SSN
- ★ SSN C2



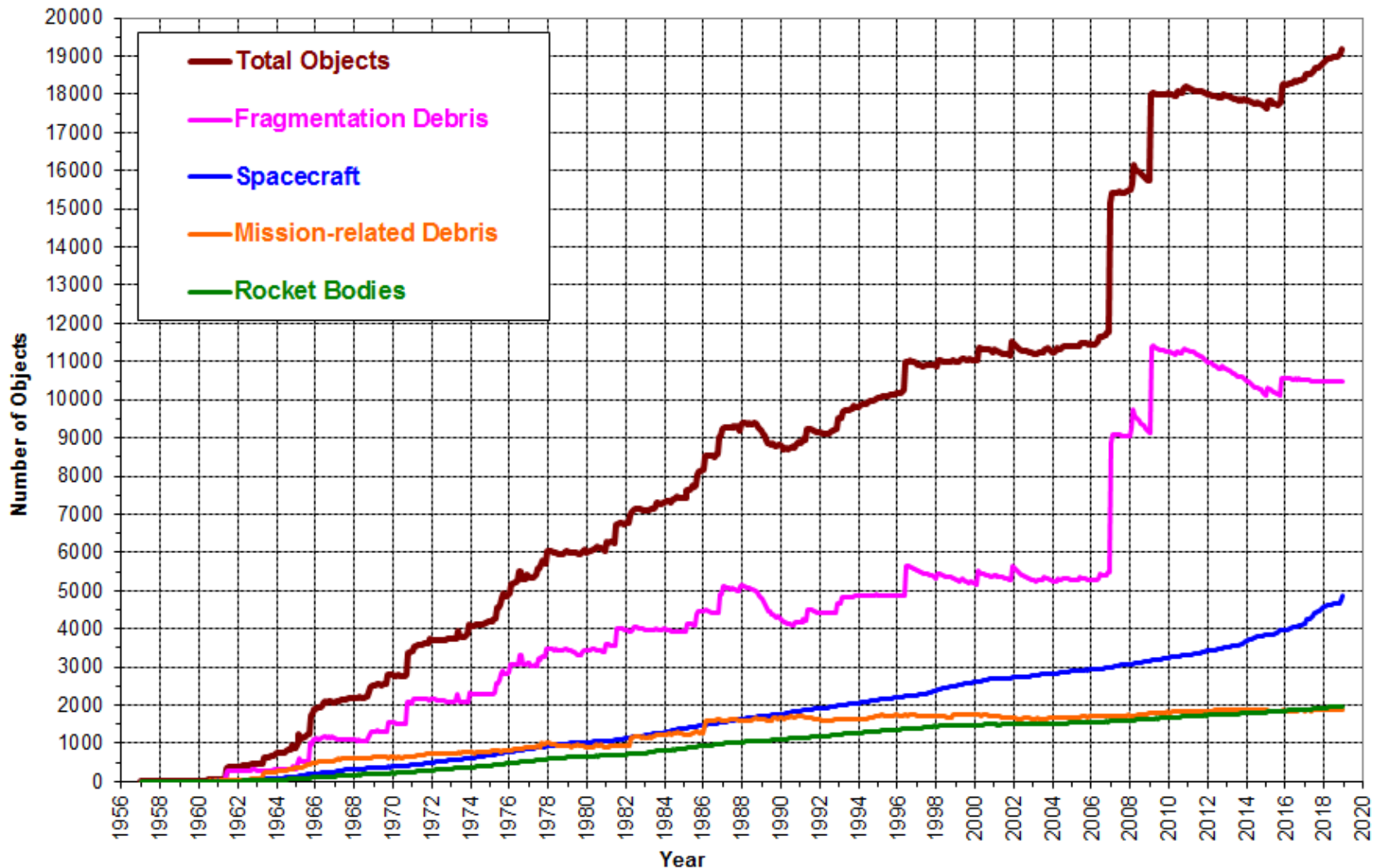
SSN Catalogue Orbital Environment





Evolution of the Catalogued (>10 cm) Satellite Population by Number

Monthly Number of Objects in Earth Orbit by Object Type



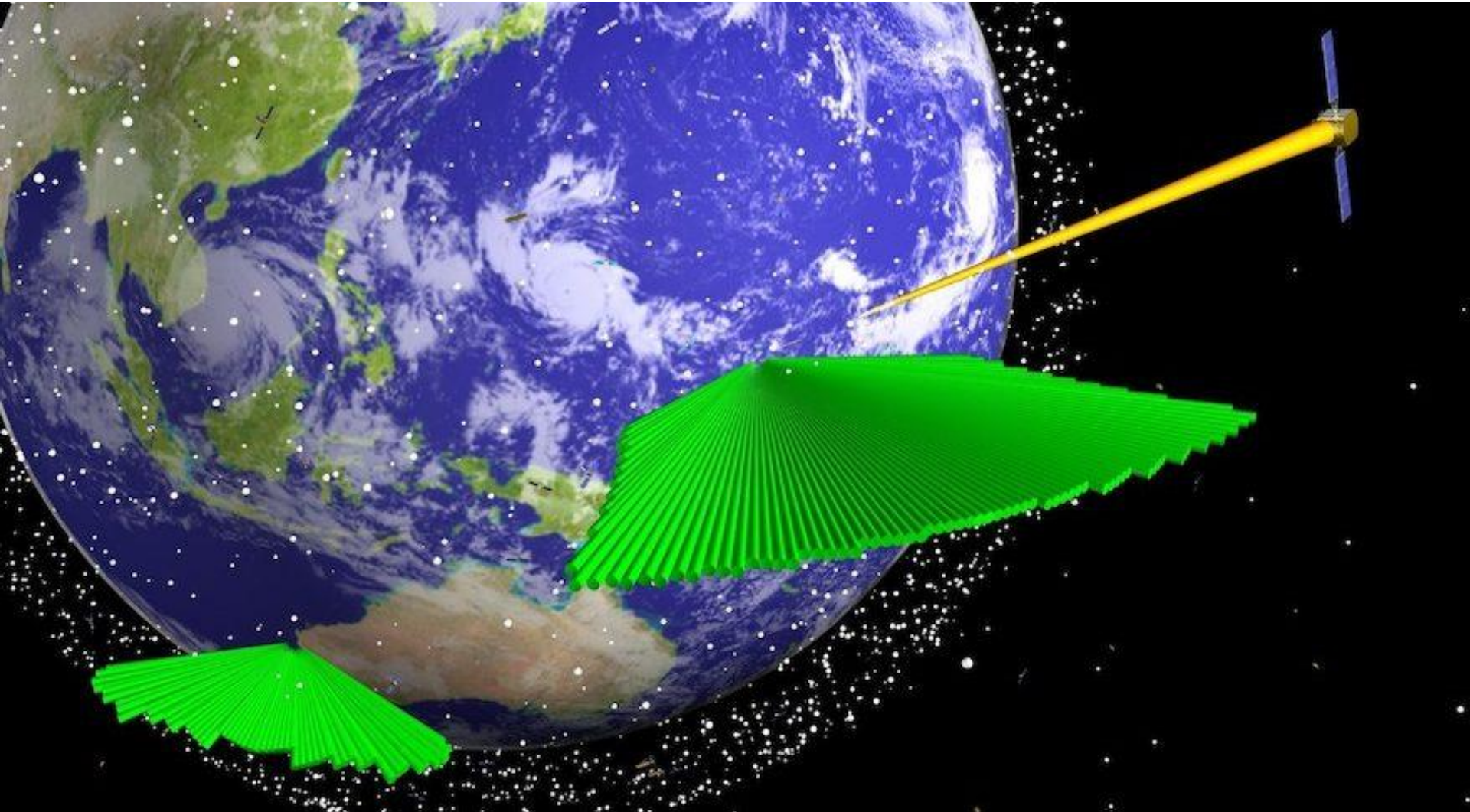


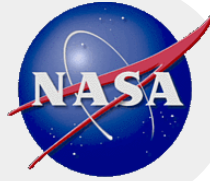
CSpOC

- **The Combined Space Operations Center (CSpOC) is tasked with using the measurement data of the SSN to maintain a Catalog of space objects**
 - **Catalog consists of objects large enough to be detected by sensors of the SSN and observed often enough to determine their orbits with sufficient accuracy to recover the object on a future pass over an SSN sensor**
 - **This tracking capability allows the CSpOC to perform conjunction assessment calculations for satellite users**
 - **There is a sensitivity limit for the SSN sensors, generally given as >10 cm in low-Earth orbit (LEO), and losing sensitivity for deep space objects**
 - **However, we know there are many debris smaller than 10 cm in size that cannot be tracked**



Future Space Fence





Collision Avoidance

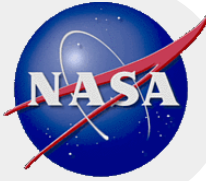
- **The current statistical technique used by CSpOC was developed as a joint project by the DoD and NASA originally to ensure the safety of Shuttle and ISS astronauts**
- **Service now provided to any space user**
 - **Possible conjunction warning given to registered user**
 - **Contains the covariance matrix and encounter geometry for each object**
 - Covariance matrix gives uncertainty ellipsoid of the position of each object
 - Information can be used to compute a probability of collision
- **Conjunction assessment for NASA**
 - **Human spaceflight handled by Mission Control in Houston and by their counterparts in Moscow**
 - **Robotic spacecraft handled by NASA's Conjunction Assessment Risk Analysis (CARA) team at GSFC**



Collision Avoidance

- **While Collision Avoidance is a prudent thing for a spacecraft operator to do, it is not a cure-all for space debris issues**
- **While a collision-avoidance maneuver reduces the collision risk, depending on the maneuver threshold chosen and the geometry of a conjunction, it does not mitigate 100% of the collision probability – a (sometimes substantial) residual risk remains**
- **For every object tracked, there are tens to hundreds of objects we cannot track that can still cause serious damage to a spacecraft**
- **Vast majority of objects tracked (~95%) are inert and cannot maneuver**
 - **By itself not a solution for problem of long-term collisional growth of the debris environment**

Complementary NASA and DoD MMOD Environment Efforts



Activity

Lead Agency

Environment Definition (>10 cm)

DoD

Environment Definition (<10 cm)

NASA (ODPO)

Environment Definition (Meteoroids)

NASA (MEO)

Risk Assessments (>10 cm)

DoD

Risk Assessments (<10 cm)

NASA

OD Mitigation Measures

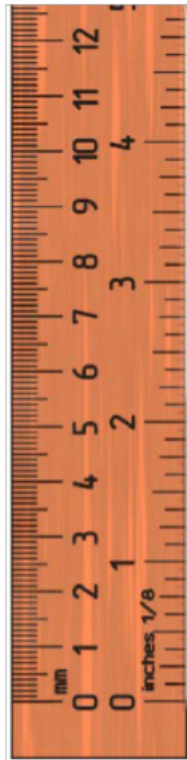
NASA

OD Environment Projection

NASA



Debris Sizes



**Softball size or larger (≥ 10 cm): ~23,000
(tracked by U.S. Combined Space Operations Center, CSpOC)**



Marble size or larger (≥ 1 cm): ~500,000



**Dot or larger (≥ 1 mm): >100,000,000
(a grain of salt)**

- Due to high impact speed in space (~ 10 km/sec in LEO), even **sub-millimeter** debris pose a realistic threat to human spaceflight and robotic missions
 - 10 km/sec = 22,000 miles per hour (the speed of a bullet $\sim 1,500$ miles per hour)
- **Mission-ending threat is dominated by small (mm-to-cm sized) debris impacts**
- Total mass: >7600 tons LEO-to-GEO (~ 3000 tons in LEO)



NASA Orbital Debris Program

- **NASA uses a number of assets to monitor the orbital debris environment <10 cm in order to characterize:**
 - **Size distribution**
 - **Orbit distributions (inclination, altitude, eccentricity)**
 - **Possible sources**
 - **Material types**
 - **Shape**
- **NASA uses a statistical sampling technique - a sensor samples the environment over time in order to make statistical conclusions about the debris populations**
 - **Determine how the debris are distributed in orbit**
 - Allows the ability to calculate the collision/damage risk to spacecraft
 - Allows the spacecraft designers to build their spacecraft with better shielding or other techniques to minimize failure risk
 - **Identify new sources and prevent future debris-creating events**
 - **Accurately assess the danger from known sources**
 - **Assess how space activities might be degrading the debris environment**
 - **Monitor for unforeseen new events invisible to the SSN**

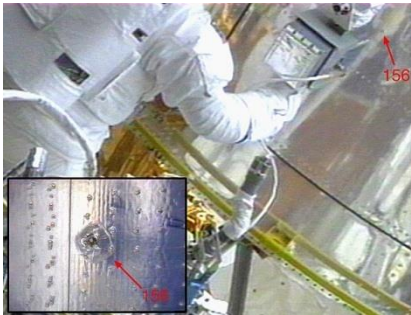
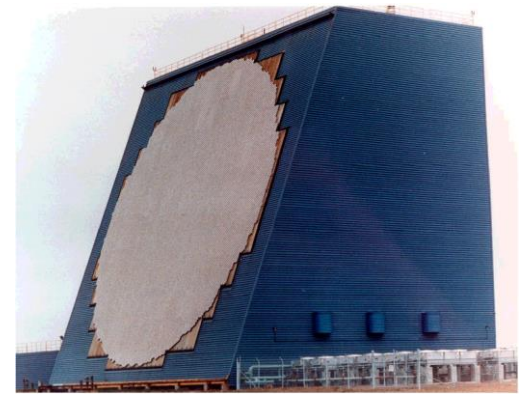


Damage Potential

EVA Suit Penetration

TPS Penetration for Crew Return Vehicles

Damage to ISS Shields

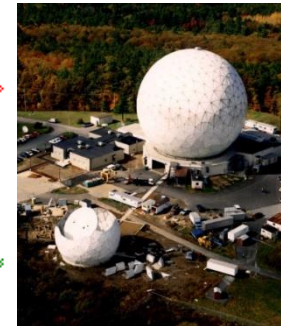


Goldstone Radars

Space Surveillance Network



Haystack Auxiliary Radar

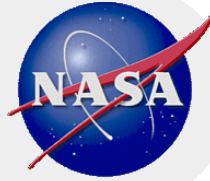


**Spacecraft
Surface
Inspections**

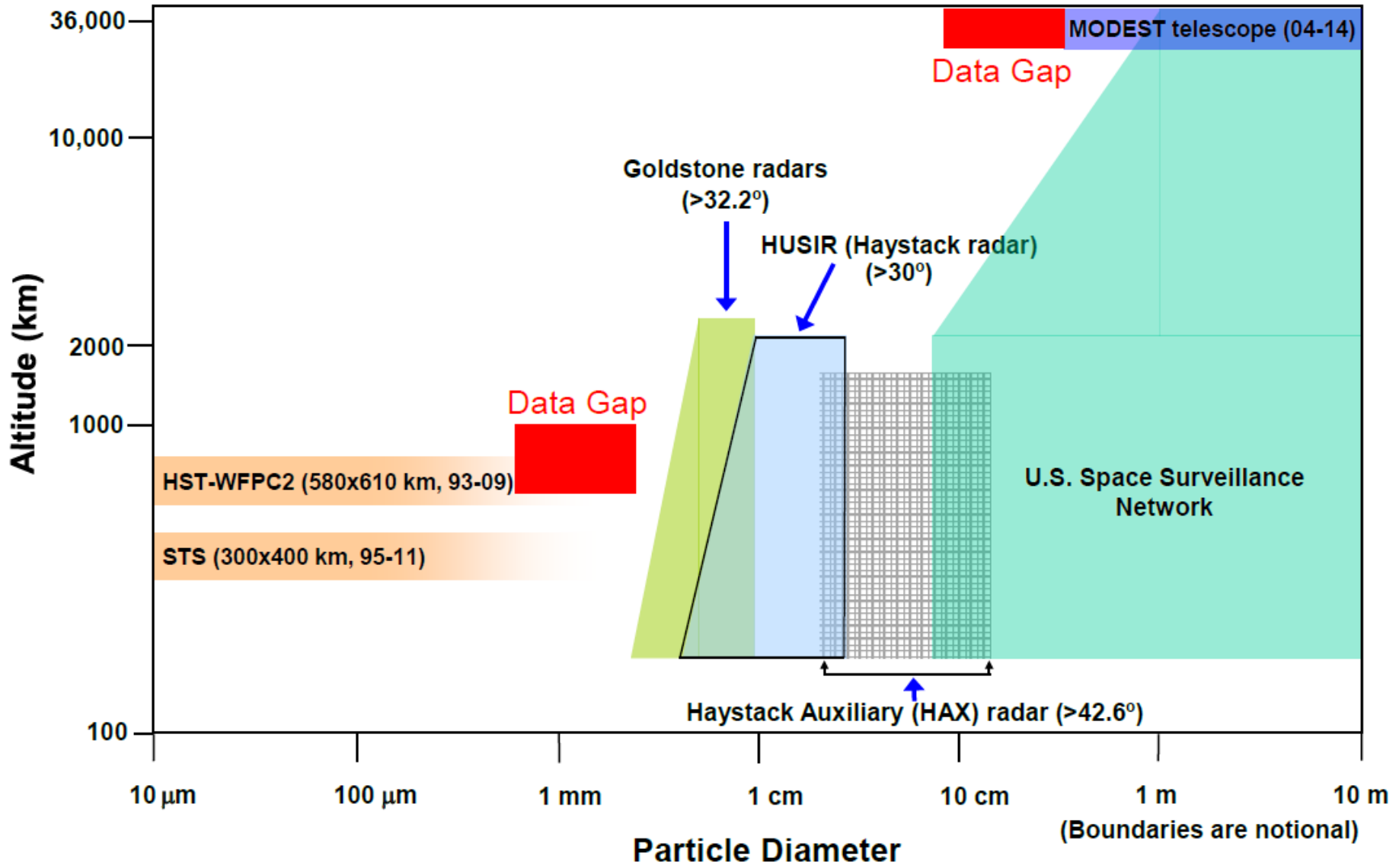
Haystack Radar

0.001 0.01 0.1 1 10 100 1000

Debris Diameter in Centimeters



NASA Measurements

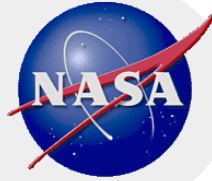




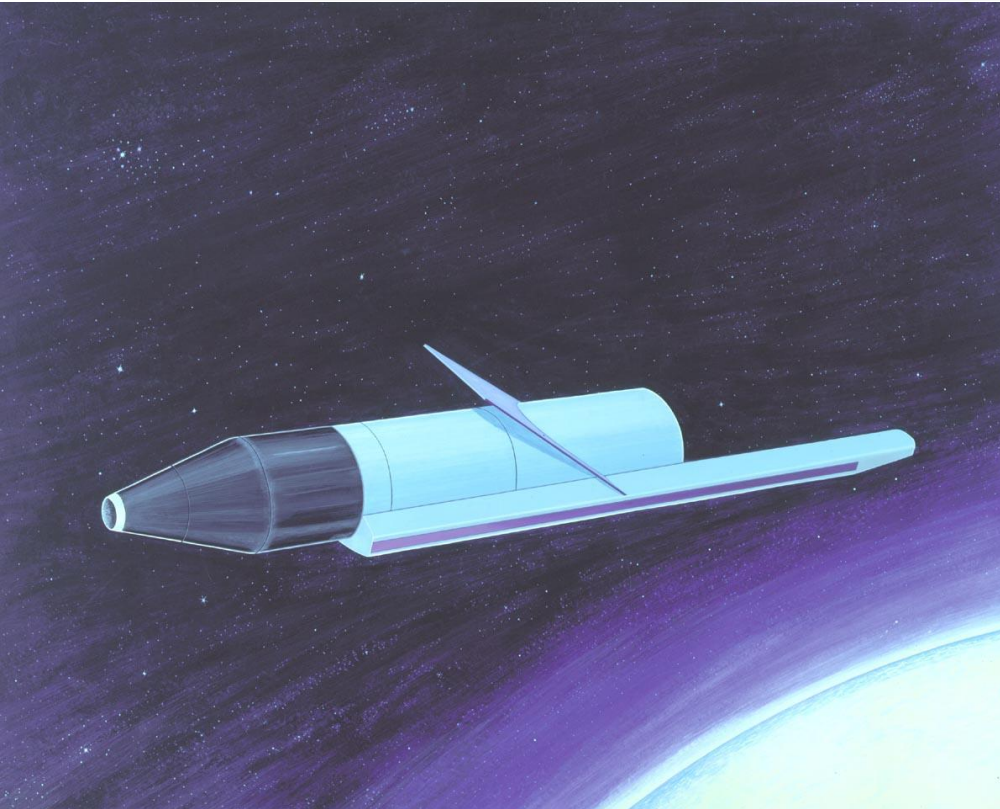
HUSIR/HAX Radars

- **Located in Massachusetts – 42.6° latitude, operated by MIT Lincoln Laboratories**
- **Haystack Ultrawide Satellite Imaging Radar (HUSIR – previously known as Haystack)**
 - 36 m diameter
 - 3 cm wavelength (X-band)
 - Can detect debris > 5 mm in LEO
- **Haystack Auxiliary Radar (HAX)**
 - 15 m diameter
 - 1.8 cm wavelength (Ku-band)
 - Can detect debris > 2 cm in LEO
- **These radars accurately measure RCS, range, and Doppler velocity along line of sight, but have trouble with other velocity components, so they usually operate in an off-vertical mode (75° East), where Doppler velocity can be used to infer orbit inclination**

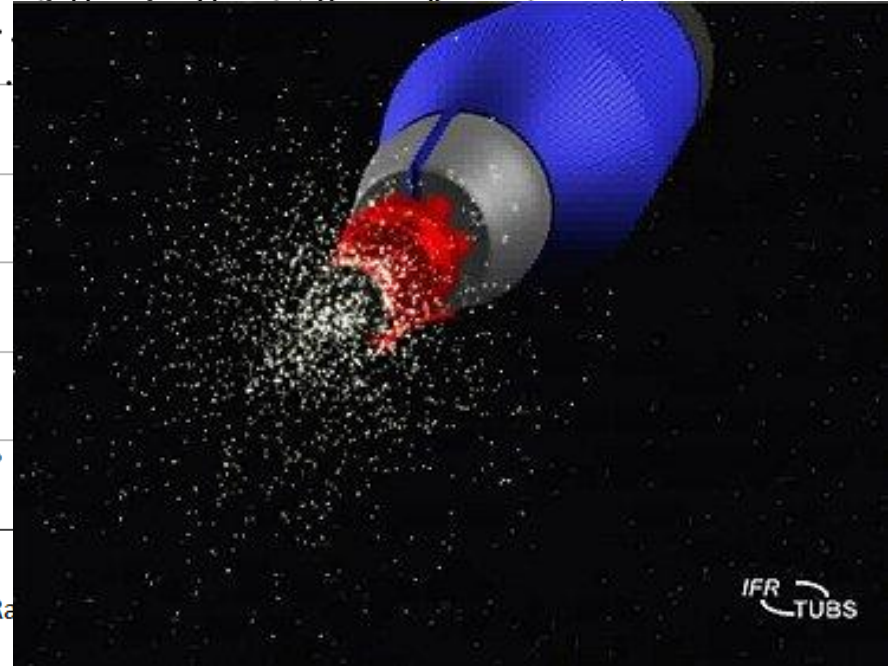
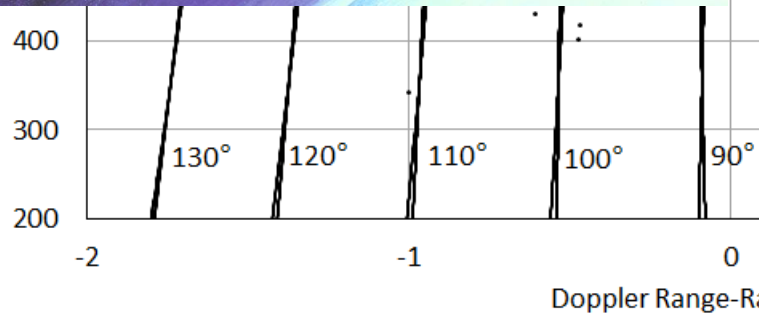
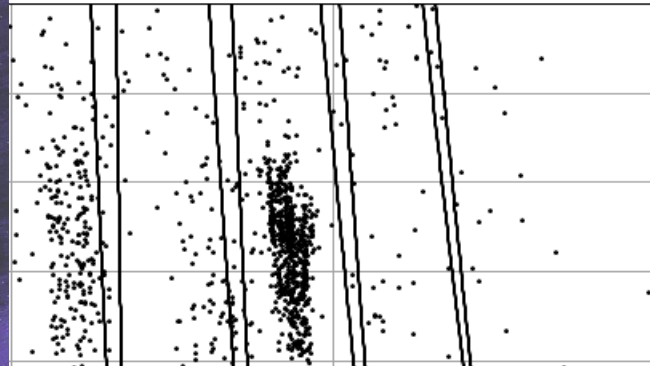




Discovery of RORSAT NaK



ta, 75 East





Goldstone Radar

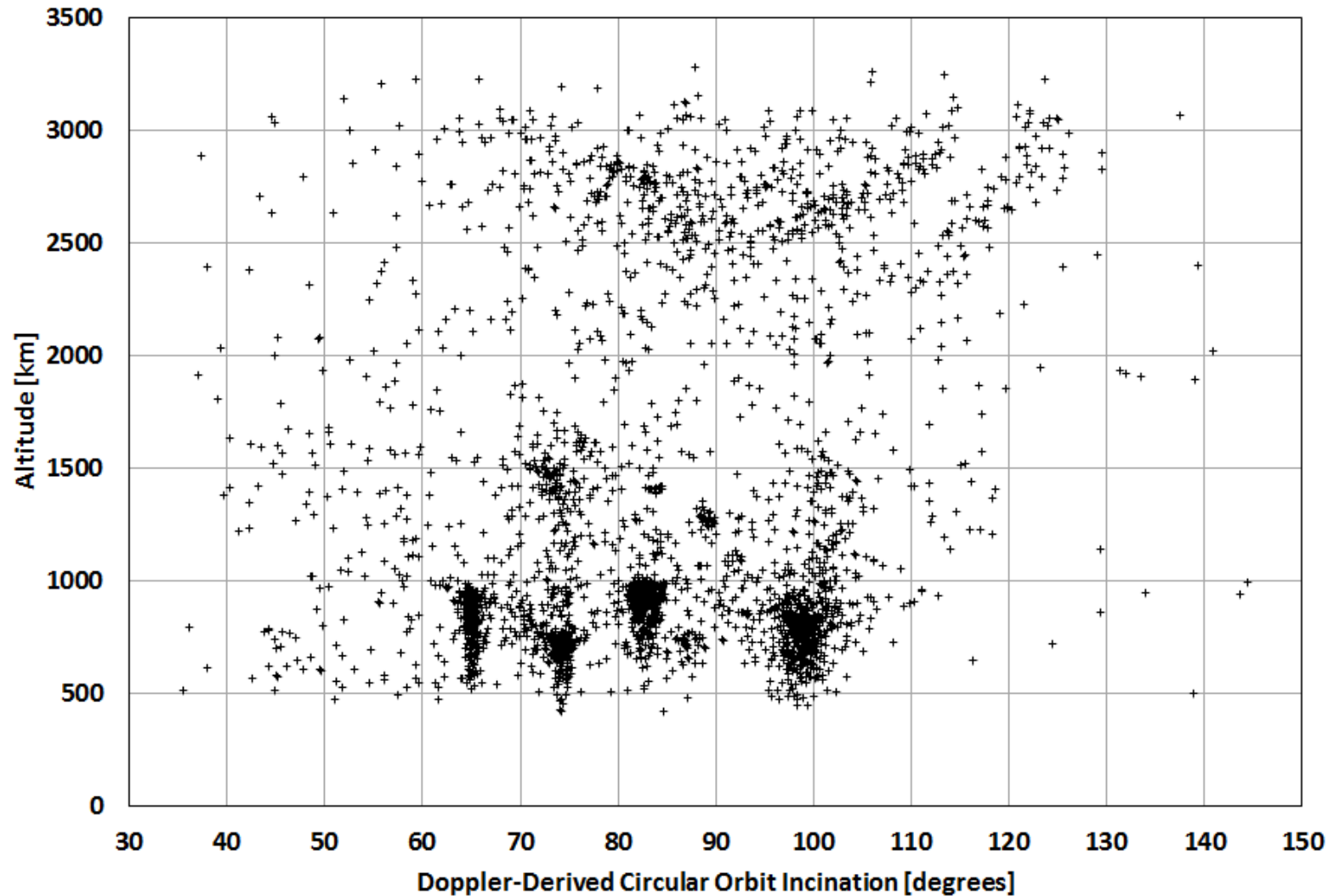
- **Located in southern California – 35.4° latitude**
- **Part of NASA's Deep Space Network**
- **Bistatic system**
 - 70 m dish + 34 m dish
 - 3.5 cm wavelength (X-band)
 - Can detect 2 mm – 5 mm debris in LEO
- **Limited capability and time available**
- **Due to upgrade of sensors, we lost the 34 m dish close to the 70 m dish (in background)**
- **New longer-baseline configurations have much reduced altitude overlap**





Goldstone Data

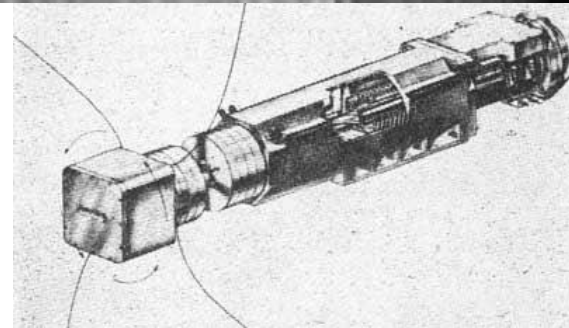
2016 Goldstone Data





West Ford Needles

- **The West Ford Needle project was a series of experiments from 1961-1963 to launch hundreds of millions of tiny copper needles (1.78 cm long, thickness of a human hair) into short-lived polar orbits to test ability to bounce signals off the resulting “ring” around the Earth**
- **Solar radiation pressure should have removed individual needles from orbit in a matter of weeks, but many stuck together in large mats that continue to orbit the Earth and are tracked by the SSN**
- **Goldstone data indicates there are also many tiny clumps of needles still in orbit**





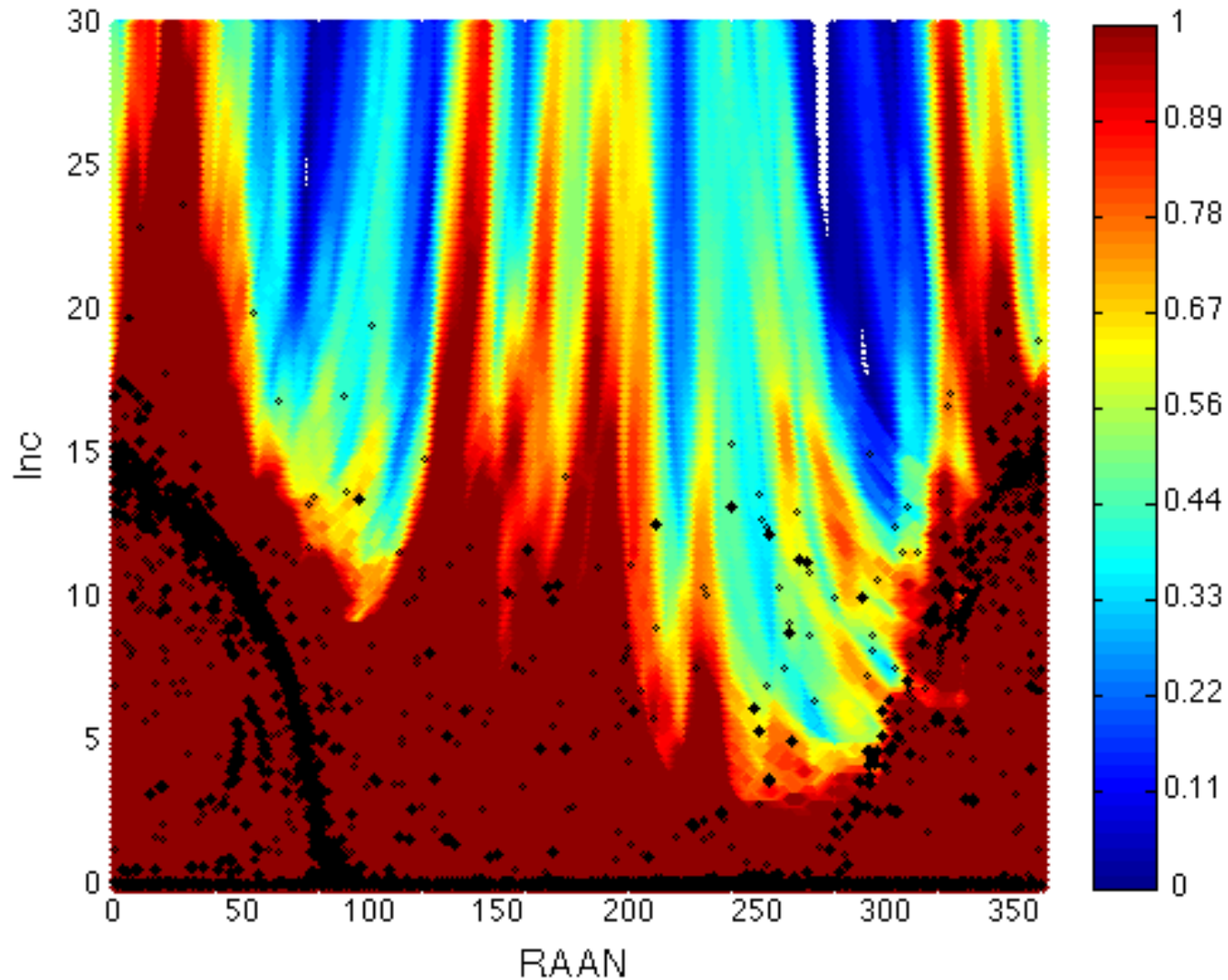
Optical Telescopes

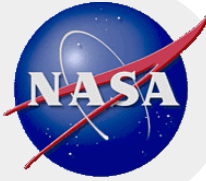
- **Telescopes are the preferred method to observe small debris at Geosynchronous orbit (GEO) altitudes**
- **For more than a decade, NASA has used the Michigan Orbital Debris Survey Telescope (MODEST) to statistically monitor the GEO environment**
 - 0.61 m aperture Curtis Schmidt optical telescope
 - Located in Chile, operated by University of Michigan
- **Observations are conducted near the Earth's shadow to maximize the reflected sunlight from debris**
 - Can detect objects down to about 30 cm in size
- **Statistical survey is corrected for probability of detecting an object in a particular orbit**



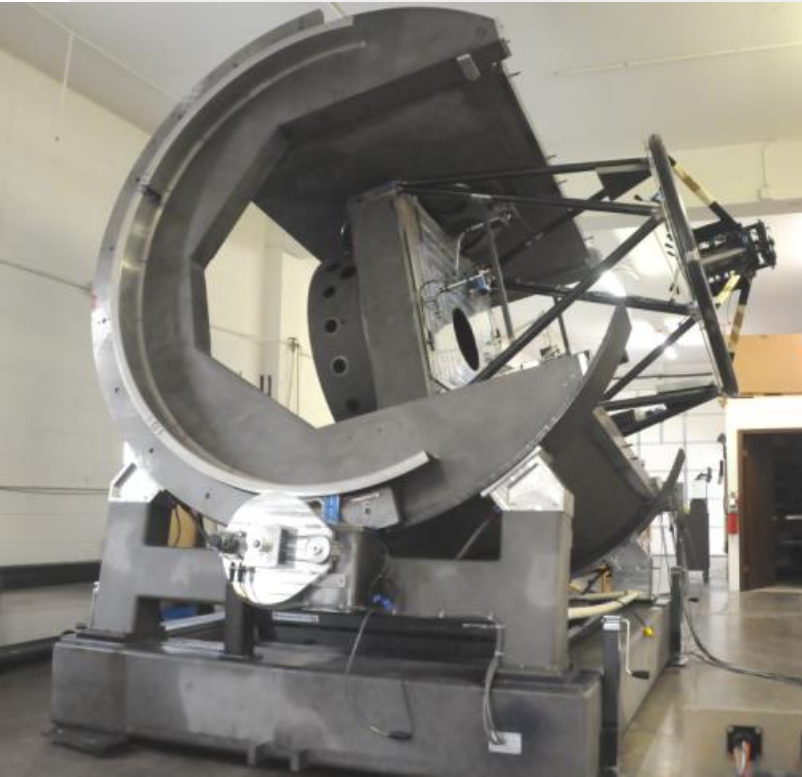


Statistical Survey





MCAT – Meter Class Autonomous Telescope



- **NASA has recently deployed the Eugene Stansbery Meter-Class Autonomous Telescope (MCAT), a 1.3 m aperture Ritchey-Chretien reflecting telescope to Ascension Island (8.0° S), in the Atlantic Ocean near the Equator**
- **Has the ability to extend statistical surveys in GEO to smaller debris (~ 20 cm)**
- **Also has the capability to look for low-altitude, low-inclination debris in LEO**



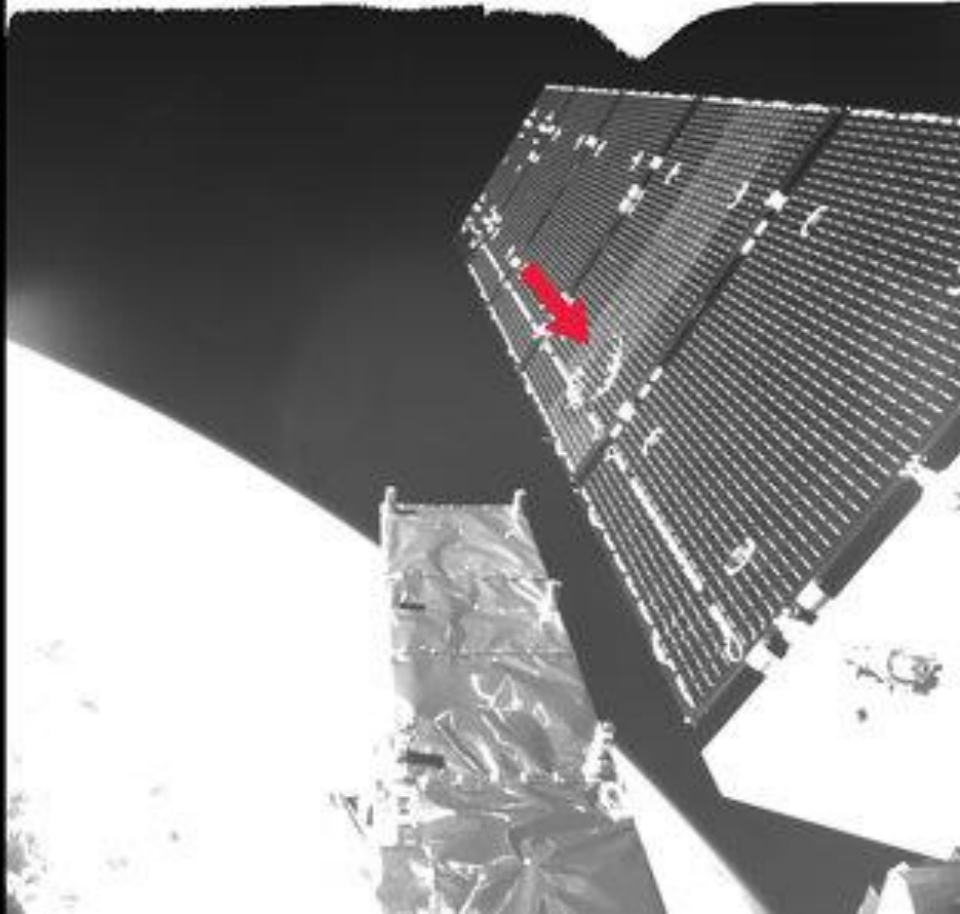
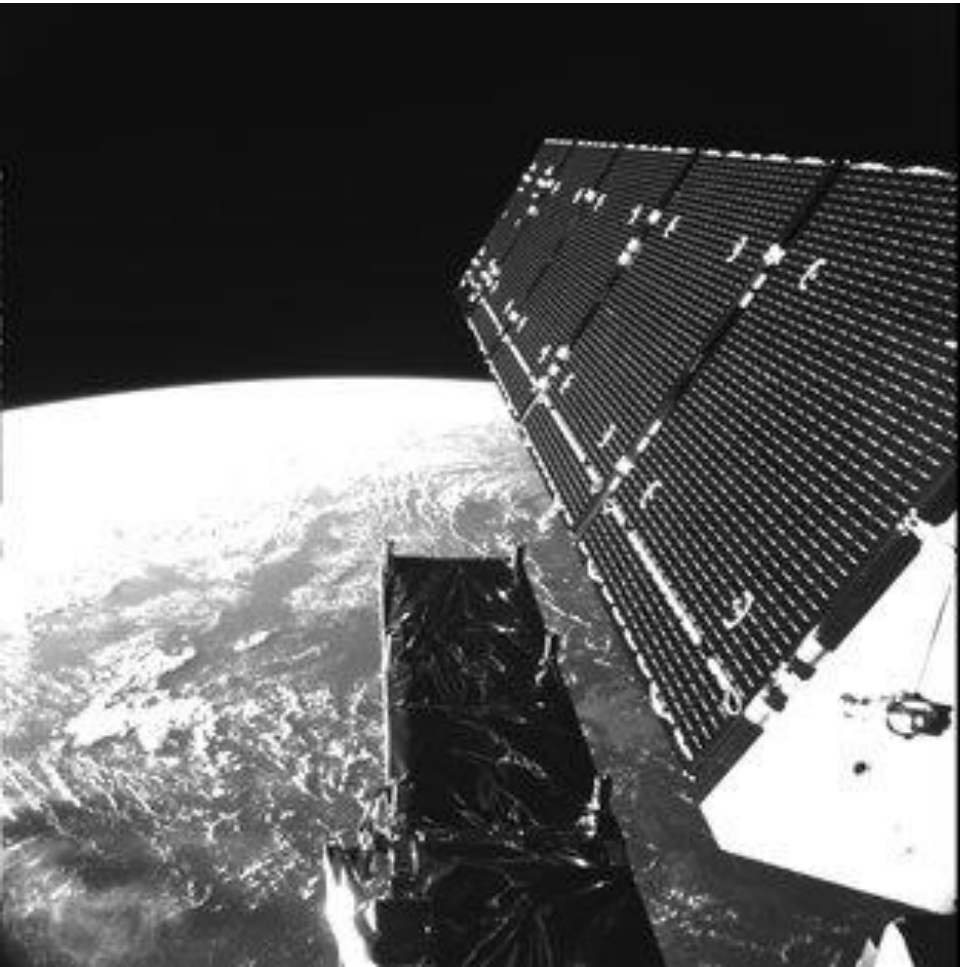


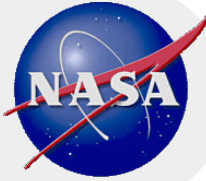
In Situ

- **For sizes smaller than ~2-3 mm, we rely on returned spacecraft surfaces**
- **Small impactors leave a damage feature – a hole or crater**
- **Feature size is a function of**
 - Particle size
 - Particle mass
 - Particle shape
 - Particle density
 - Particle speed and angle of impact
 - Characteristics of impacted surface
- **The chief problem is that we do not typically know these things for each particle, all we have is the feature size and position**
 - Sometimes, electron microscope analysis of feature yields melted residue of impactor, letting us know the material of the particle (e.g., aluminum, steel, meteoroid).
- **Use statistical techniques to “back out” debris characteristics**

National Aeronautics and Space Administration

Sentinel-1A Impact 2016/08/23 – Onboard Camera

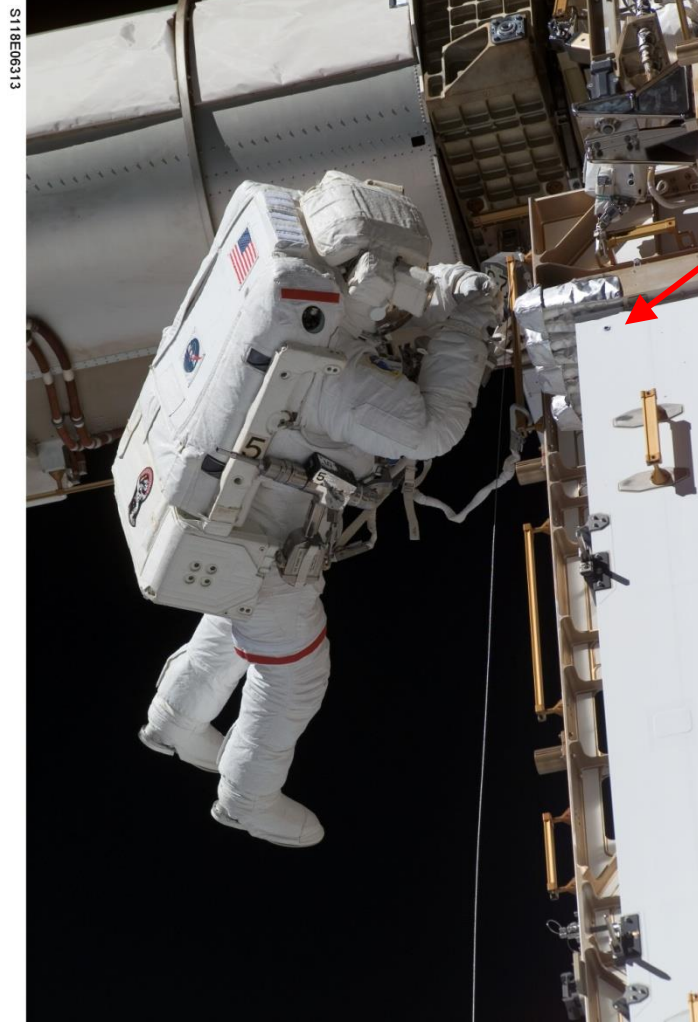




National Aeronautics and Space Administration

Debris Impacts Observed during EVA's

- Also in 2007, a crew member on EVA noticed a hypervelocity impact crater while working near a large aluminum panel.



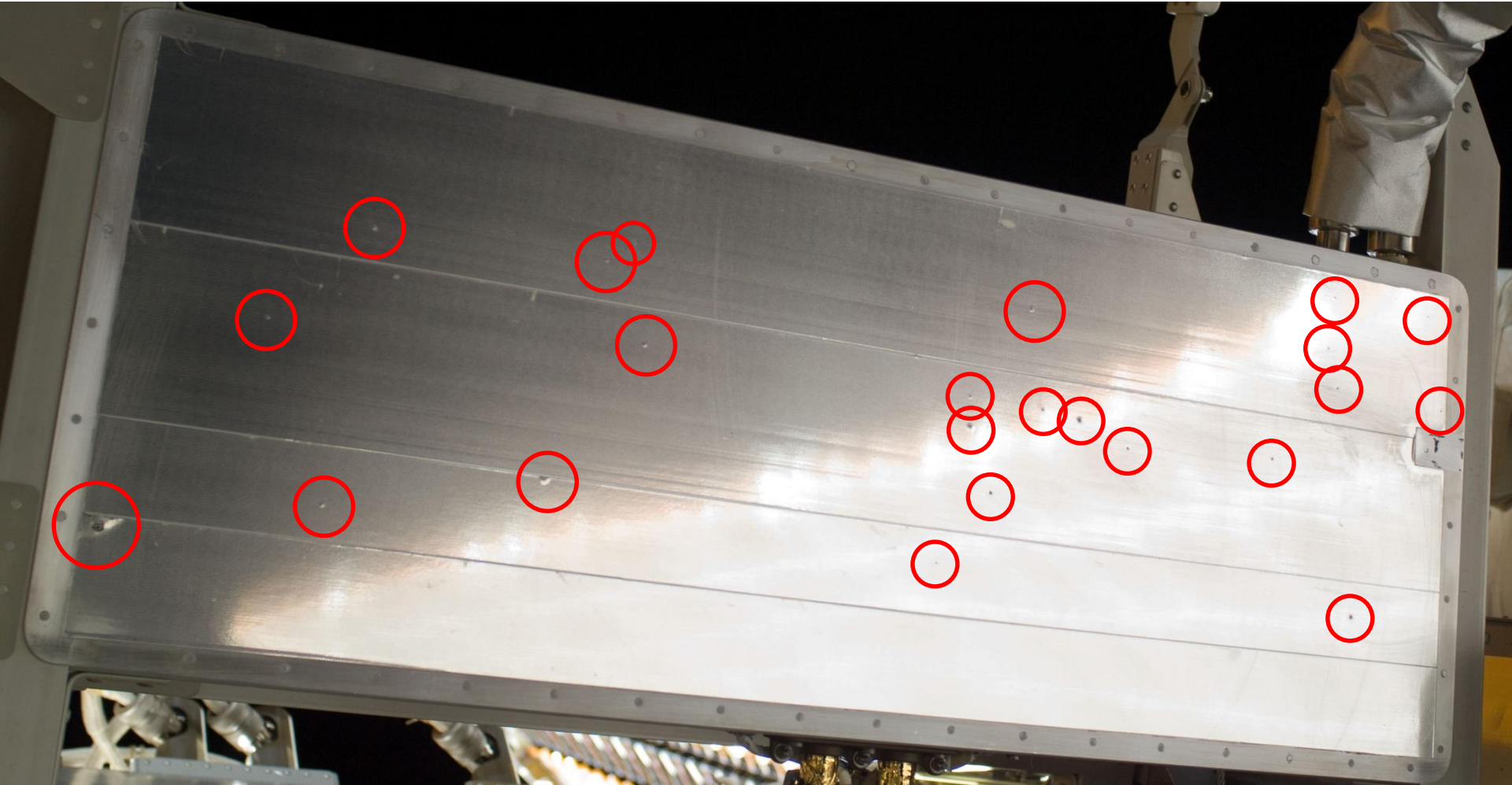
S118E06313

Space debris
impact site



MMOD Damage to ISS

- **MMOD impact damages observed to radiator panel during EVA-20 (Nov. 2012)**



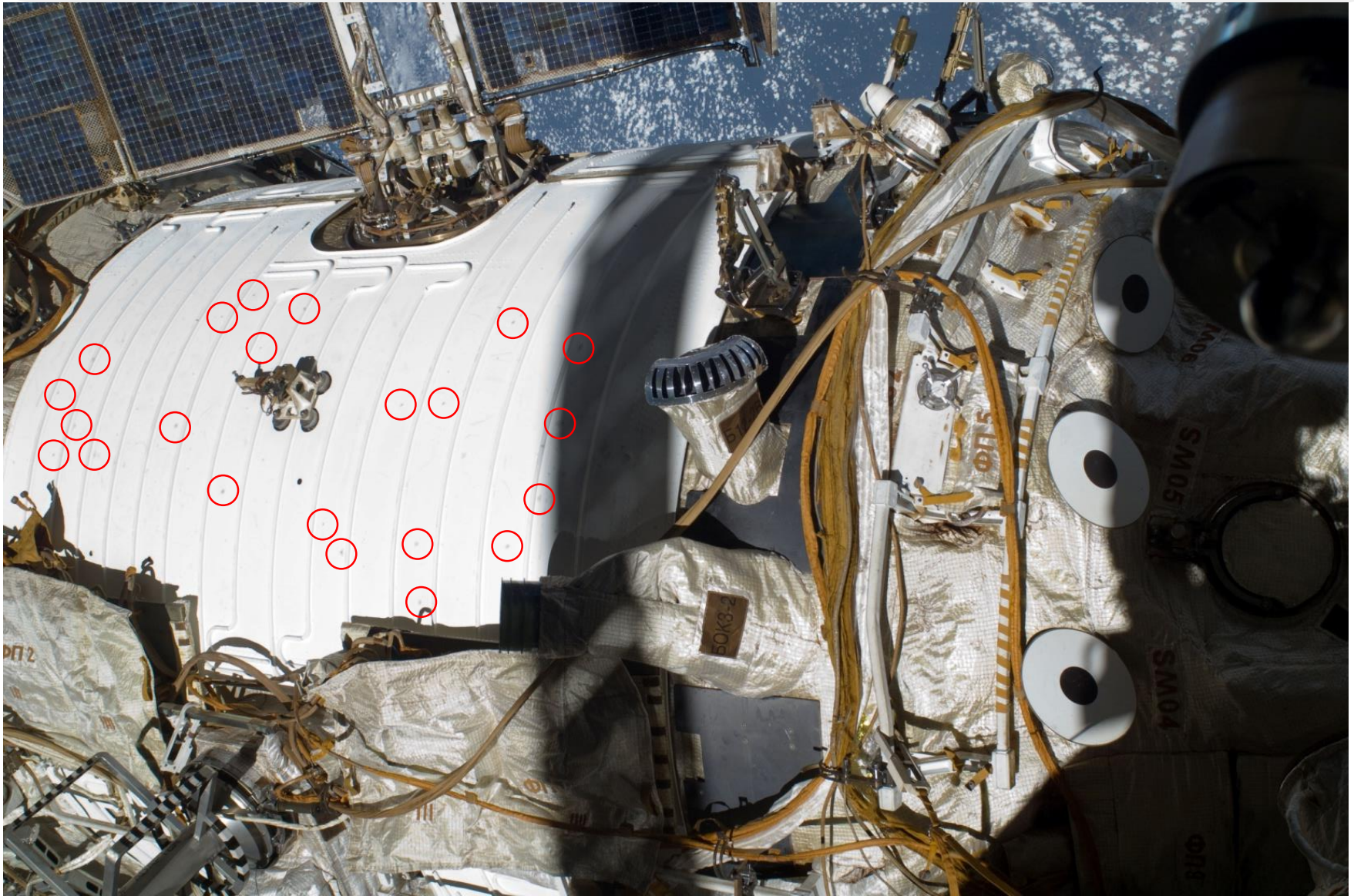


National Aeronautics and Space Administration

MMOD Damage to ISS

observed to Service Module during Russian EVA-31 (Aug. 2012)

ISS032e020579





ISS032e020579

Close-up of SM radiator damage (1/4)





ISS032e020579

Close-up of SM radiator damage (2/4)





ISS032e020579

Close-up of SM radiator damage (3/4)

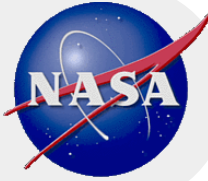




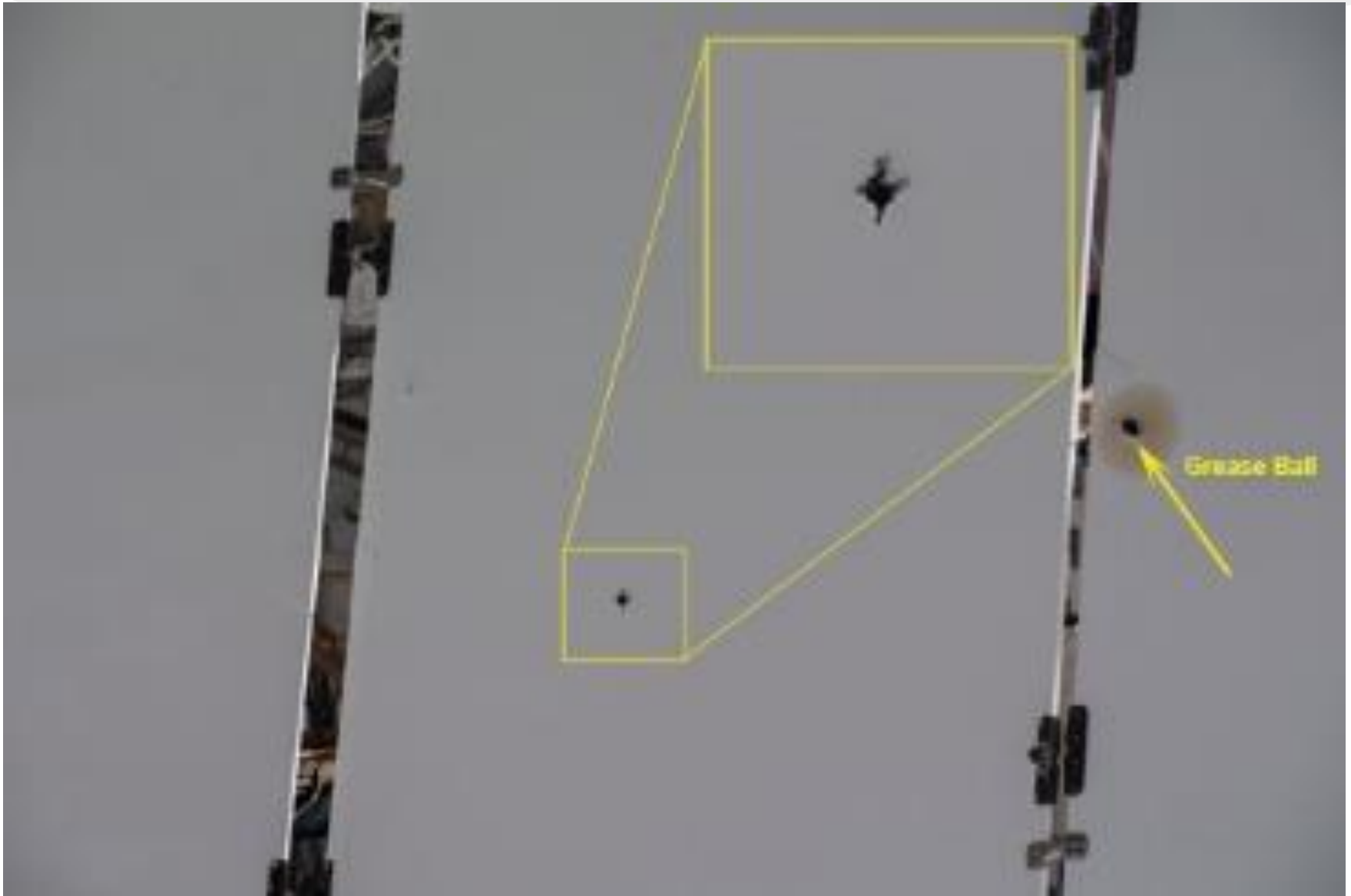
ISS032e020579

Close-up of SM radiator damage (4/4)





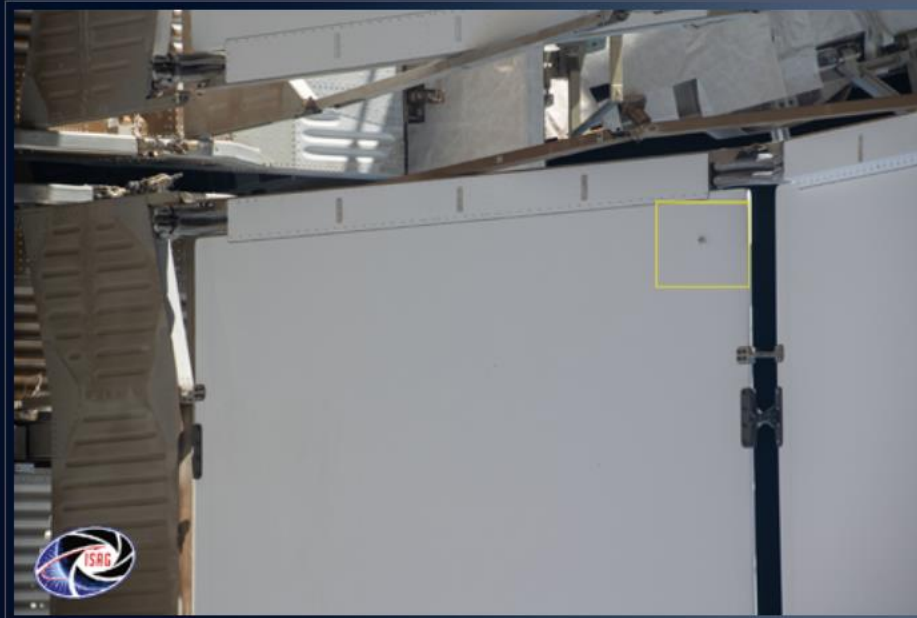
Recent ISS Radiator Imagery

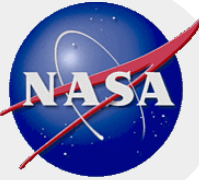




Recent ISS Radiator Imagery

MMOD Strikes Observed on S1-2 HRS Center Radiators; Panel 1





National A



Tim Peake ✓
@astro_timpeake

Follow

Often asked if [@Space_Station](#) is hit by space debris. Yes – this chip is in a Cupola window
[esa.int/spaceinimages/ ...](http://esa.int/spaceinimages/)



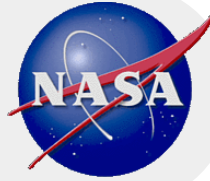
RETWEETS
1,735

LIKES
2,840

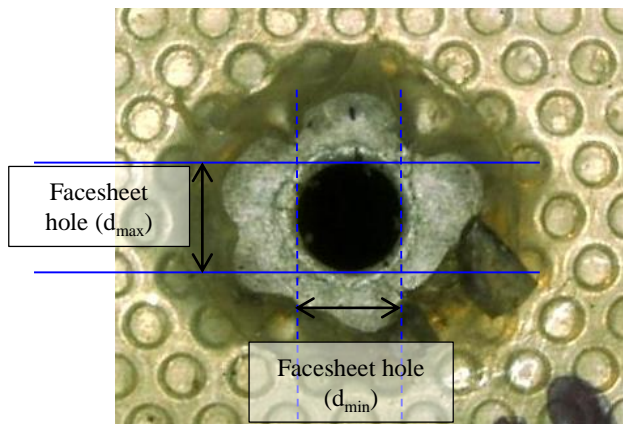
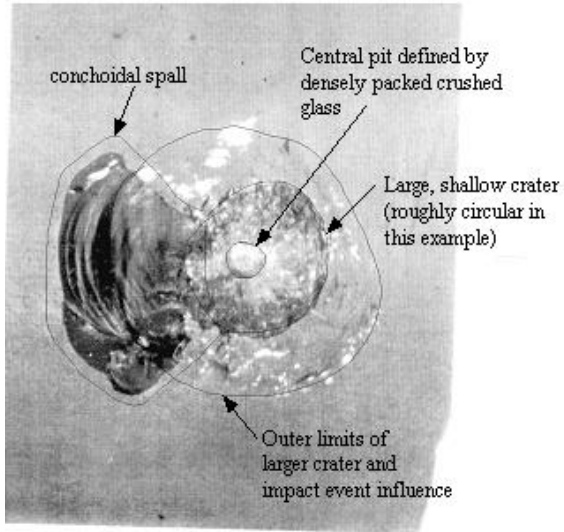


6:07 AM - 12 May 2016





Shuttle *In Situ* Data

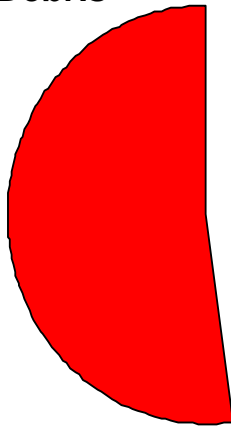




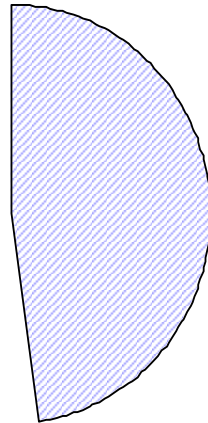
Material Types of Shuttle Window Impacts

- It is possible to put craters from space-exposed hardware into an electron microscope and identify the material of the impactor

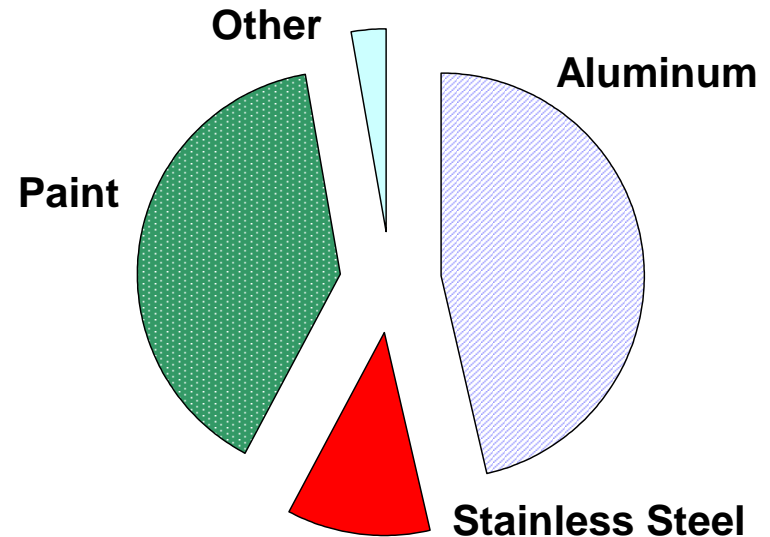
Orbital Debris



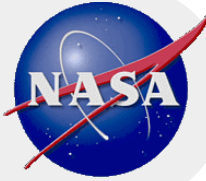
Meteoroids



Identified Impactors



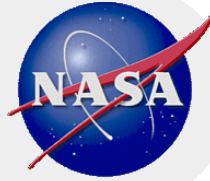
Types of Orbital Debris Impactors



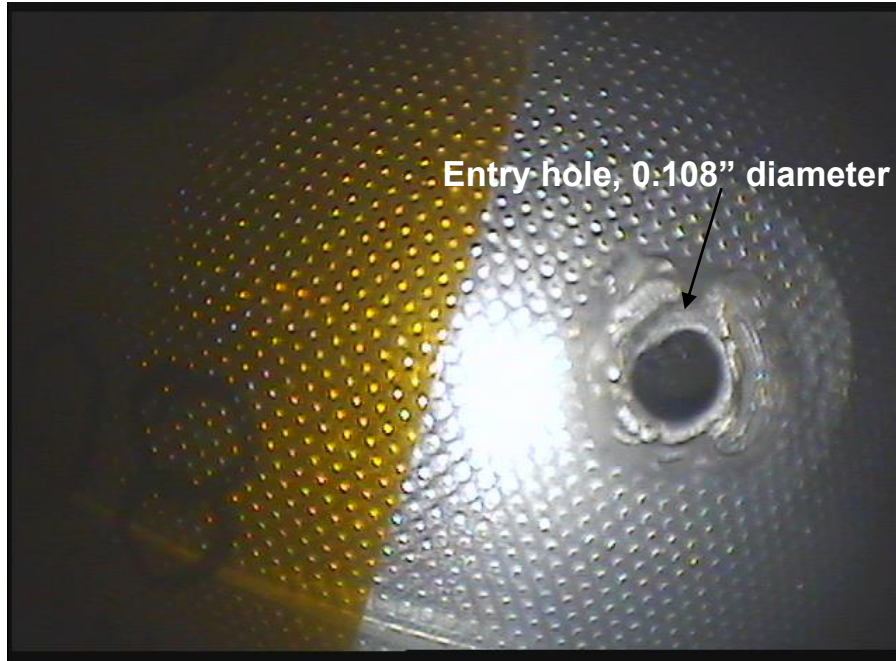
STS-115 MMOD Impact Damage

- The debris punched all the way through the radiator.
- The face sheet hole was 2.8 mm in diameter.
- The core inside the panel was completely destroyed for at least a 2.5 cm diameter below the face sheet damage.
- This is the most significant MMOD damage recorded on the Orbiter radiators up to that time

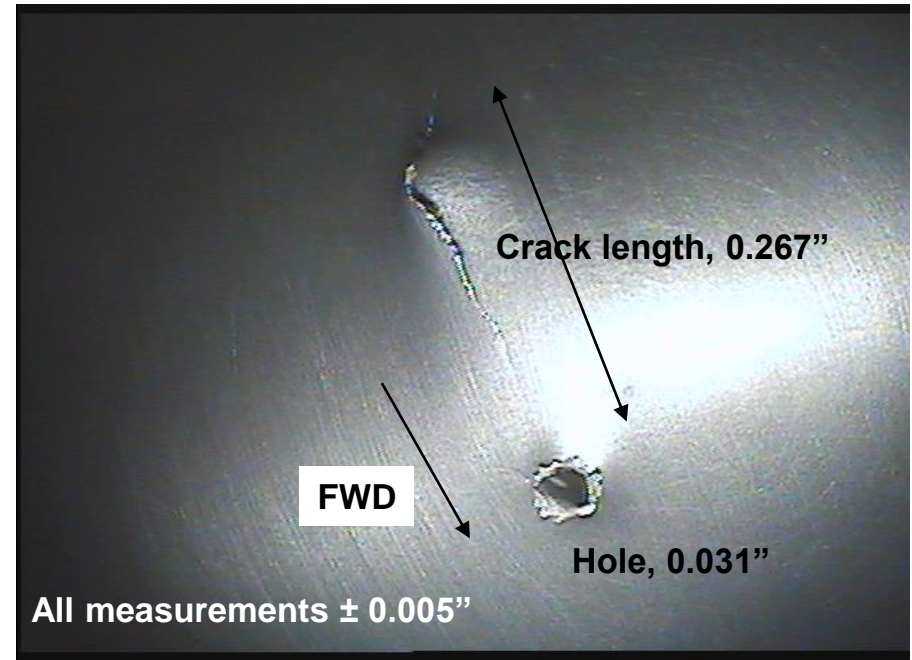




STS-115 MMOD Impact Damage



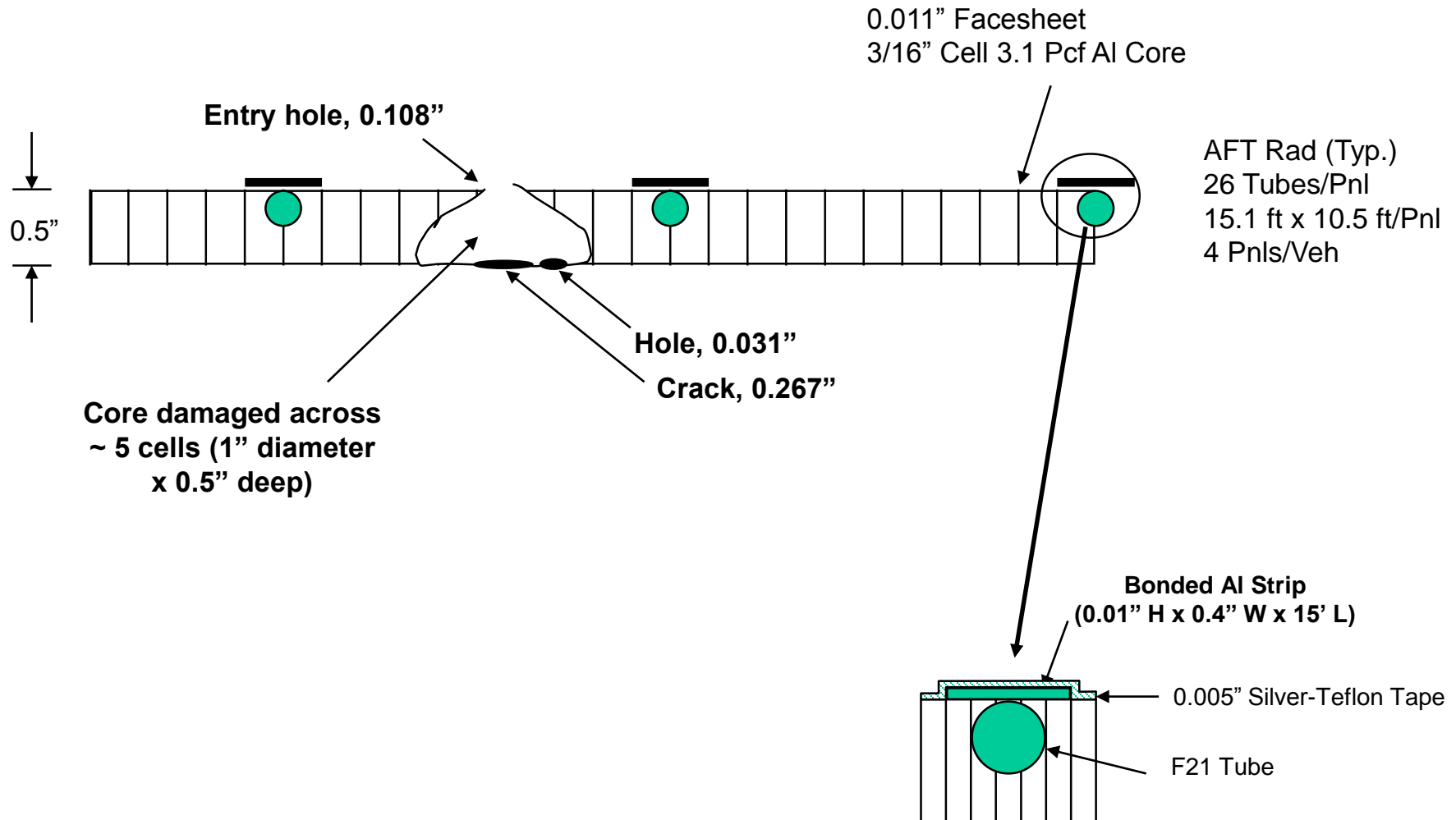
Outer face sheet damage



Inner face sheet damage

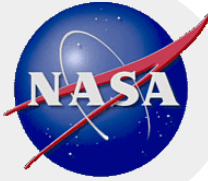


Schematic of Radiator and Sketch of Damage

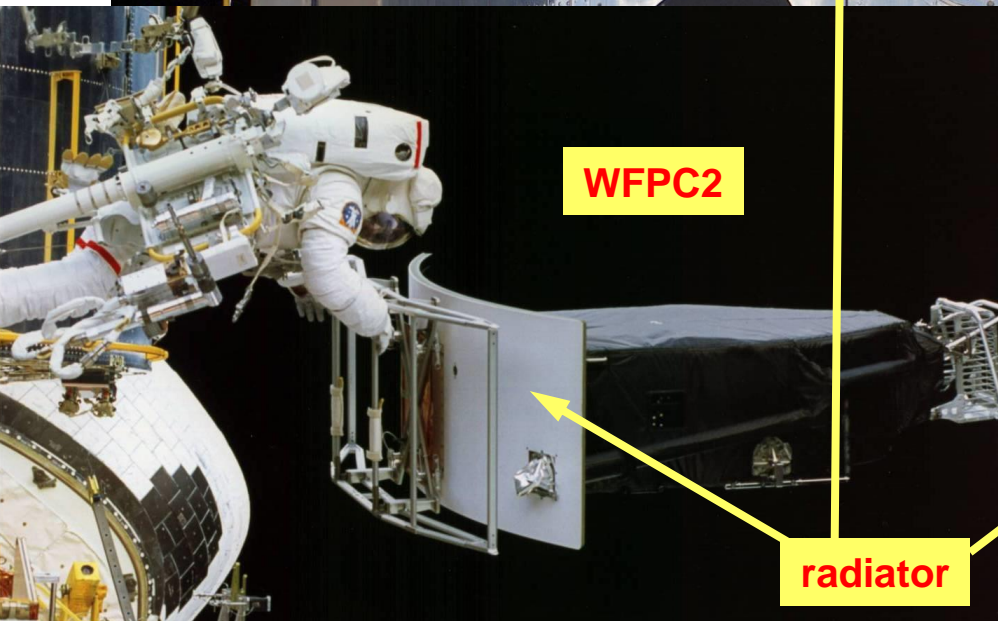
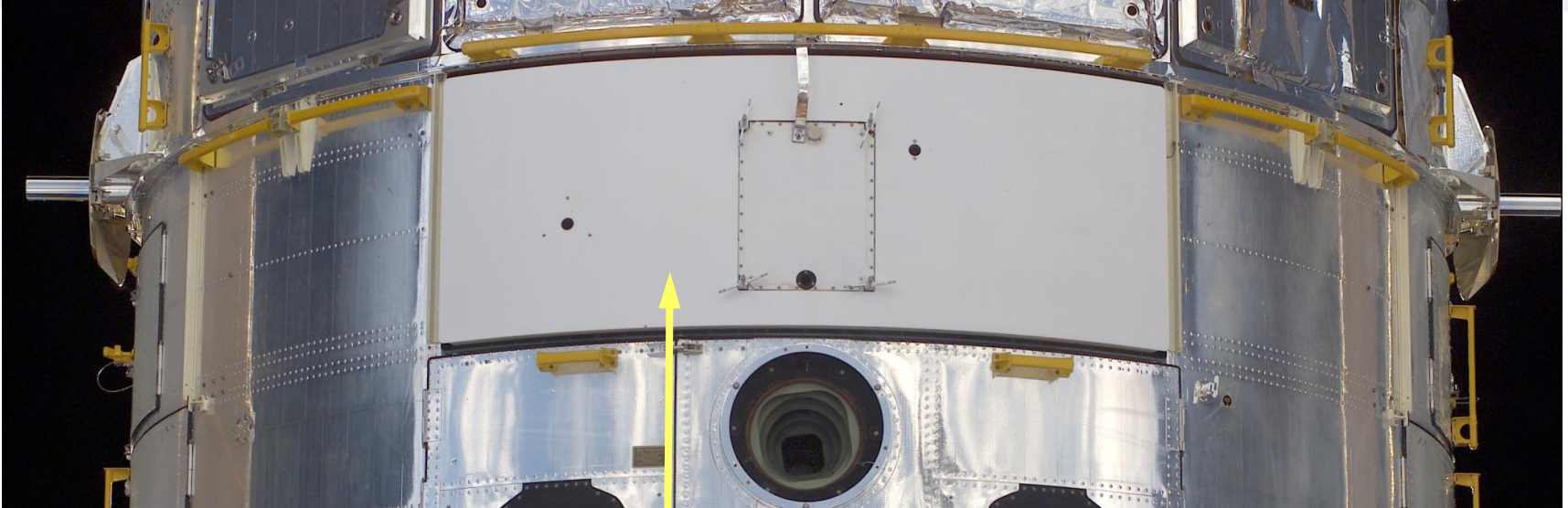




What day is trash pickup around here?

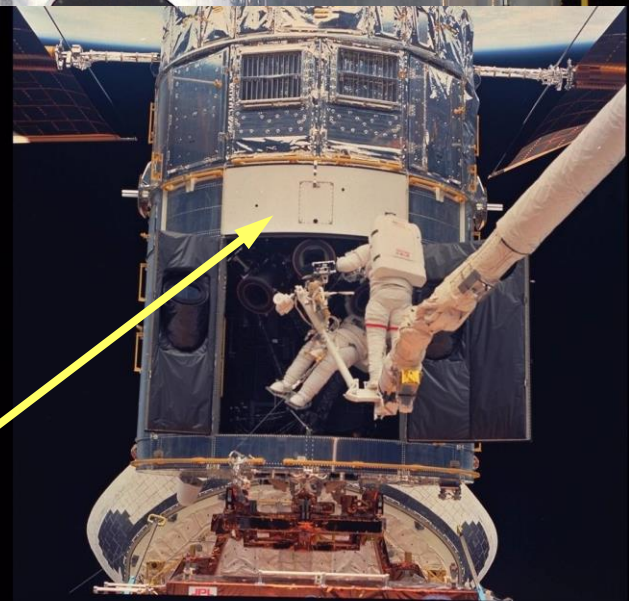


HST SM1 (STS-61, 1993)



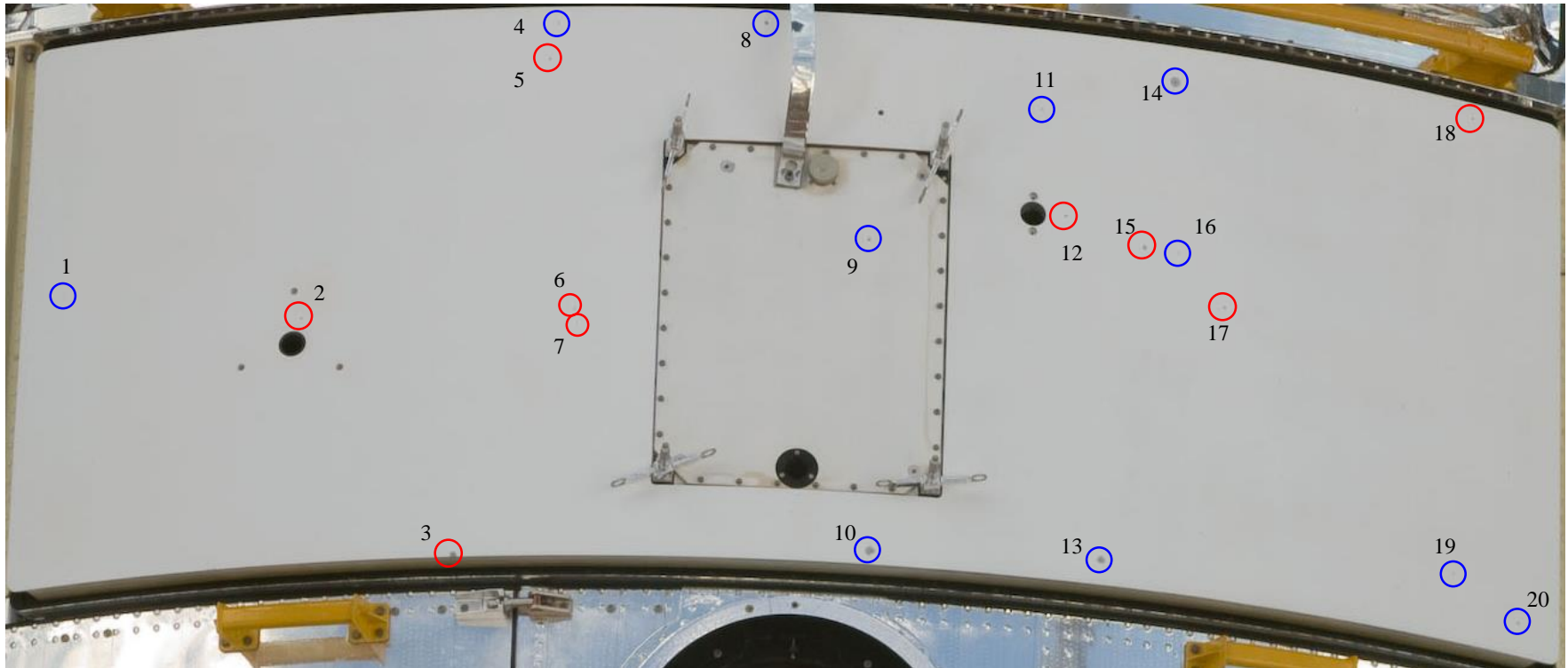
WFPC2

radiator



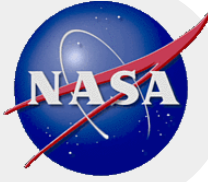


Visible MMOD Impact Damage on WFPC2 Radiator from the On-orbit Imagery Survey

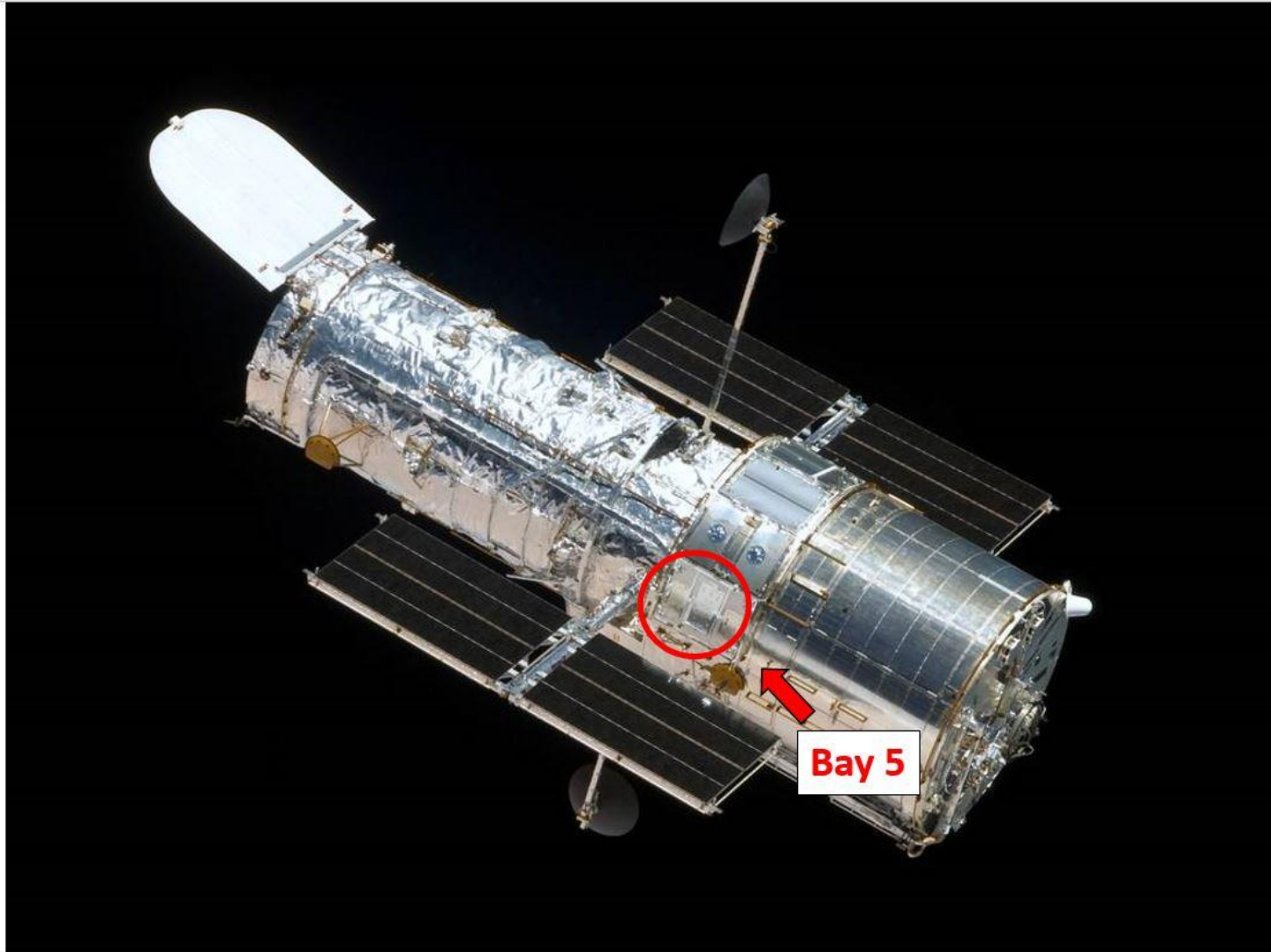


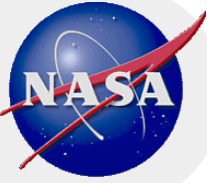
S125e006995.jpg (edited)

- Red circles: Impacts identified from SM3B images (2002)
- Blue circles: Additional impacts identified from SM4 images (2009)

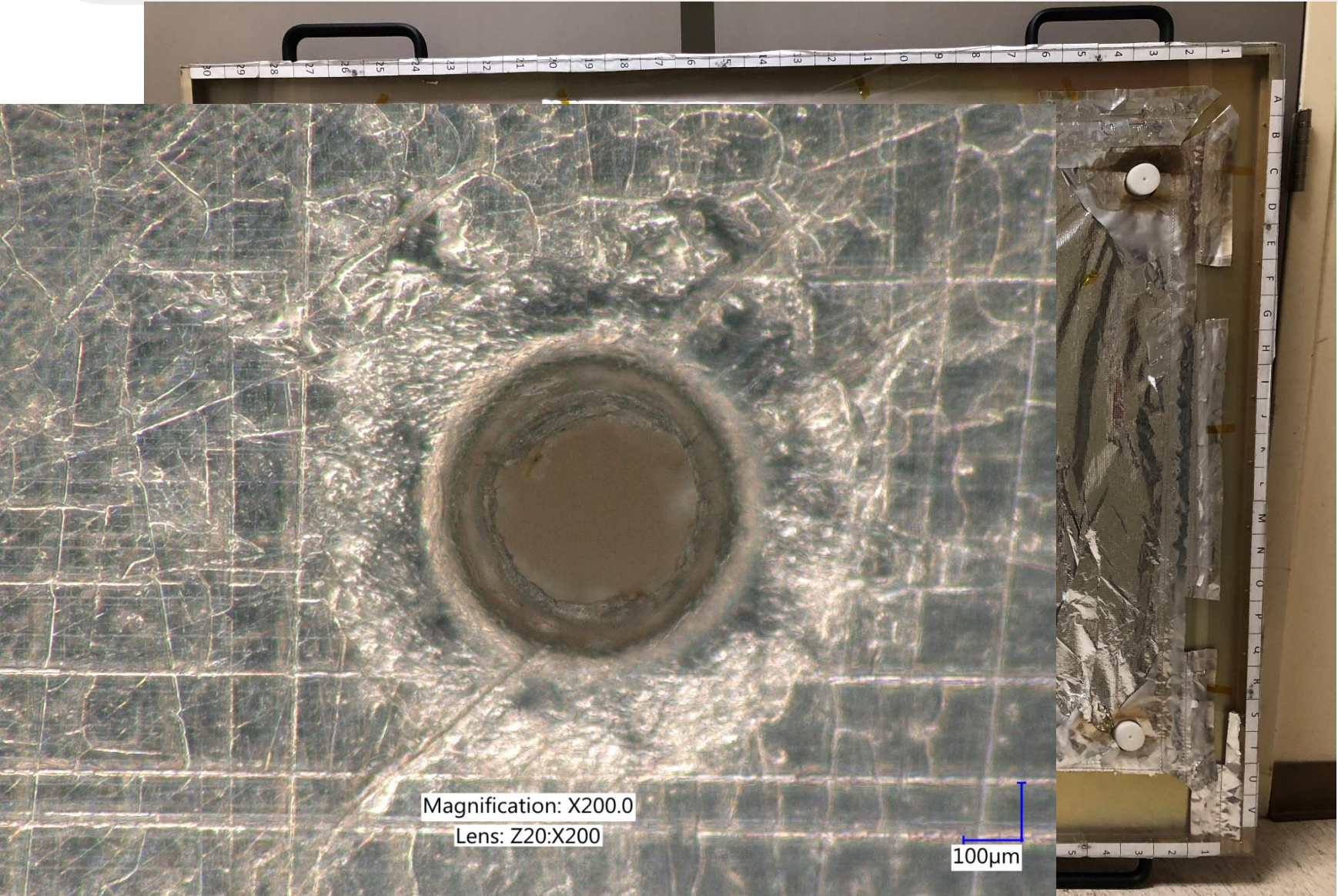


Bay 5 MLI





Bay 5 MLI



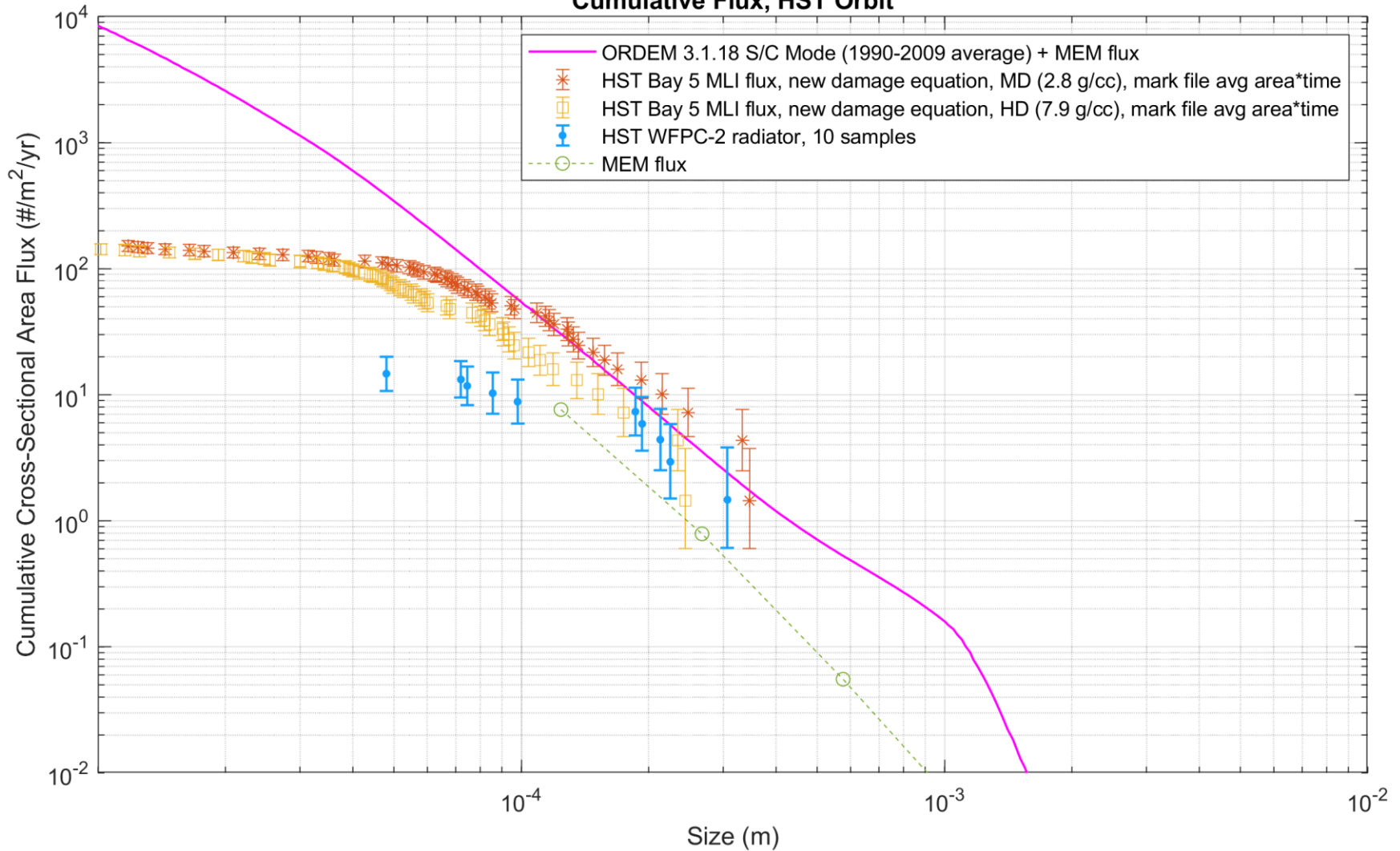
Magnification: X200.0
Lens: Z20:X200

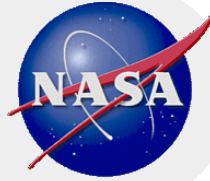
100µm



HST Crater Data

Cumulative Flux, HST Orbit





Future of *In Situ*

- **The Shuttle no longer flies, so NASA currently has no dedicated sensor to monitor the small particle environment**

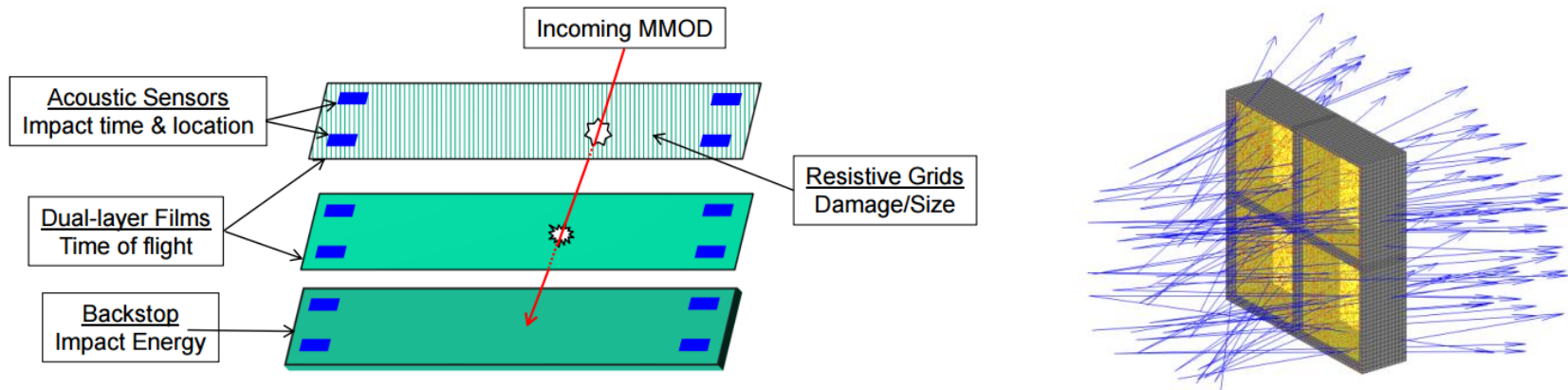


- **The best way to measure small particles is by using a dedicated, calibrated sensor, designed to measure the impactor properties of most interest**
 - Size
 - Shape
 - Material Density
 - Speed and Direction
 - Time of Impact (combined with position of sensor, can be used to determine particle orbit)

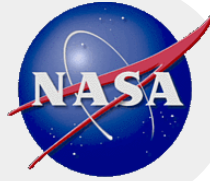


DRAGONS

- The Debris Resistive Acoustic Grid Orbital Navy-NASA Sensor (DRAGONS) is a new technology initiative to measure *in situ* debris

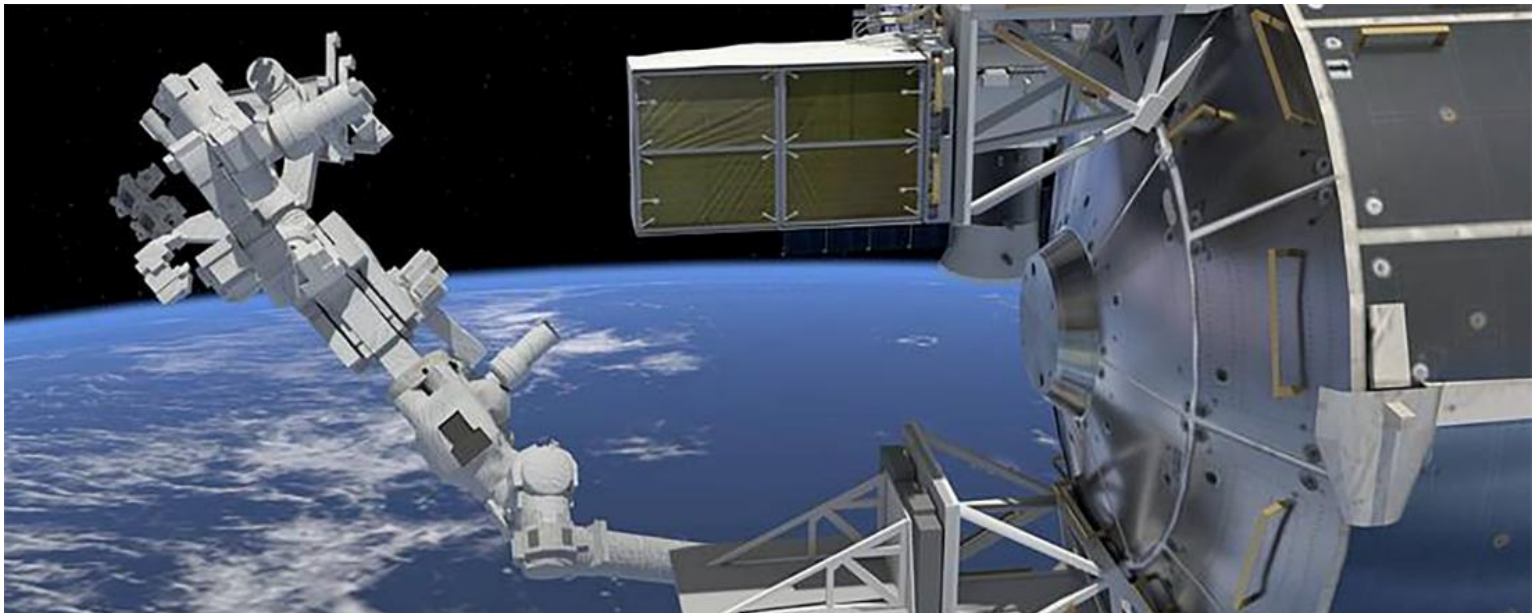


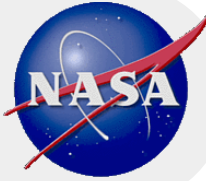
- The resistive grid on the first layer used to estimate the particle size
- Acoustic sensors at each layer to measure path and time-of-flight
- Backstop to record total energy
- Using velocity, energy, and size, should be able to estimate mass and material density
- Impact time to compute debris orbit



Space Debris Sensor

- **The technology had a flight technology demonstration on the Space Debris Sensor (SDS) aboard the ISS**
- **While some engineering data was obtained, the instrument suffered a fatal failure and is no longer operational**





Future DRAGONS

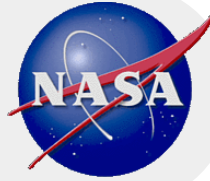
- **We are currently awaiting flight opportunities of the DRAGONS technology, especially at higher orbit altitudes**
 - **Debris environment is predicted to be worse at altitudes between 700 and 1000 km altitude**
 - **NASA and other spacecraft are spending money and resources to mitigate the predicted risk**
 - **We have little to no data on these small particles – we are relying on models to extrapolate the risk from lower altitudes**



Ground Experiments

- **Sometimes it is not enough to measure events in space, they need to be studied in the laboratory under controlled conditions**
- **There is a long history of studying collision or explosion debris on the ground by picking up the pieces afterwards**
 - Number of debris
 - Size distribution
 - Shapes
 - Delta-velocities
- **The primary source of data has been the Satellite Orbital debris Characterization Impact Test (SOCIT), which used an intact Transit satellite built in the 1960's for the target of a hypervelocity impact test**





Ground Experiments

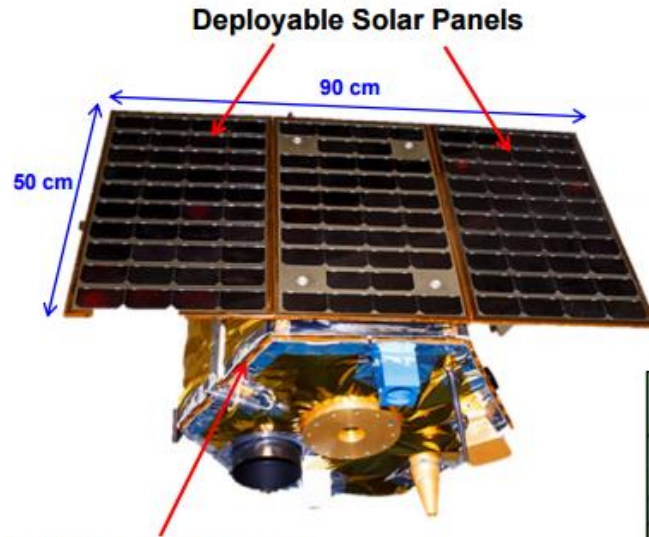
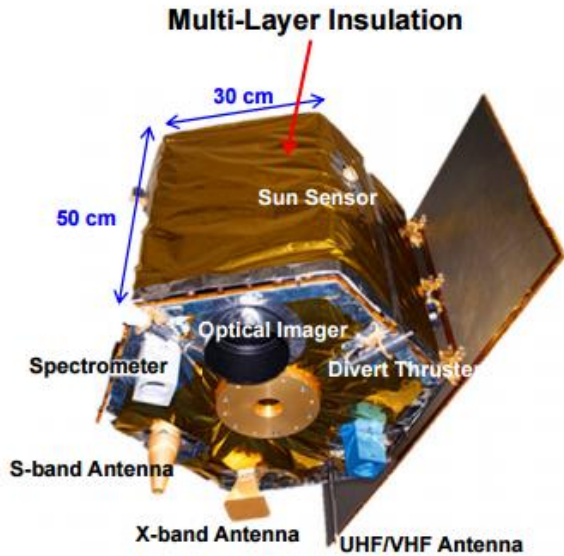
- However, there have been major changes in spacecraft construction materials over the years, so a need to test the breakup models using more modern spacecraft materials
- NASA, in conjunction with US DoD and the Aerospace Corporation, conducted the DebrisSat impact experiment, using a mock satellite made of modern materials
 - Included a test of a mock tank, designated DebrisLV

A comparison between Transit and DebrisSat

	Transit (SOCIT)	DebrisSat
Target body dimensions	46 cm (dia) × 30 cm (ht)	60 cm (dia) × 50 cm (ht)
Target mass	34.5 kg	56 kg
MLI and solar panel	No	Yes
Projectile	Al sphere	Hollow Al cylinder
Projectile dimension, mass	4.7 cm (dia), 150 g	8.6 cm × 9 cm, 570 g
Impact speed	6.1 km/sec	6.8 km/sec
Impact energy to target mass ratio	78 J/g (2.7 MJ total impact energy)	235 J/g (13.2 MJ total impact energy)



DebrisSat



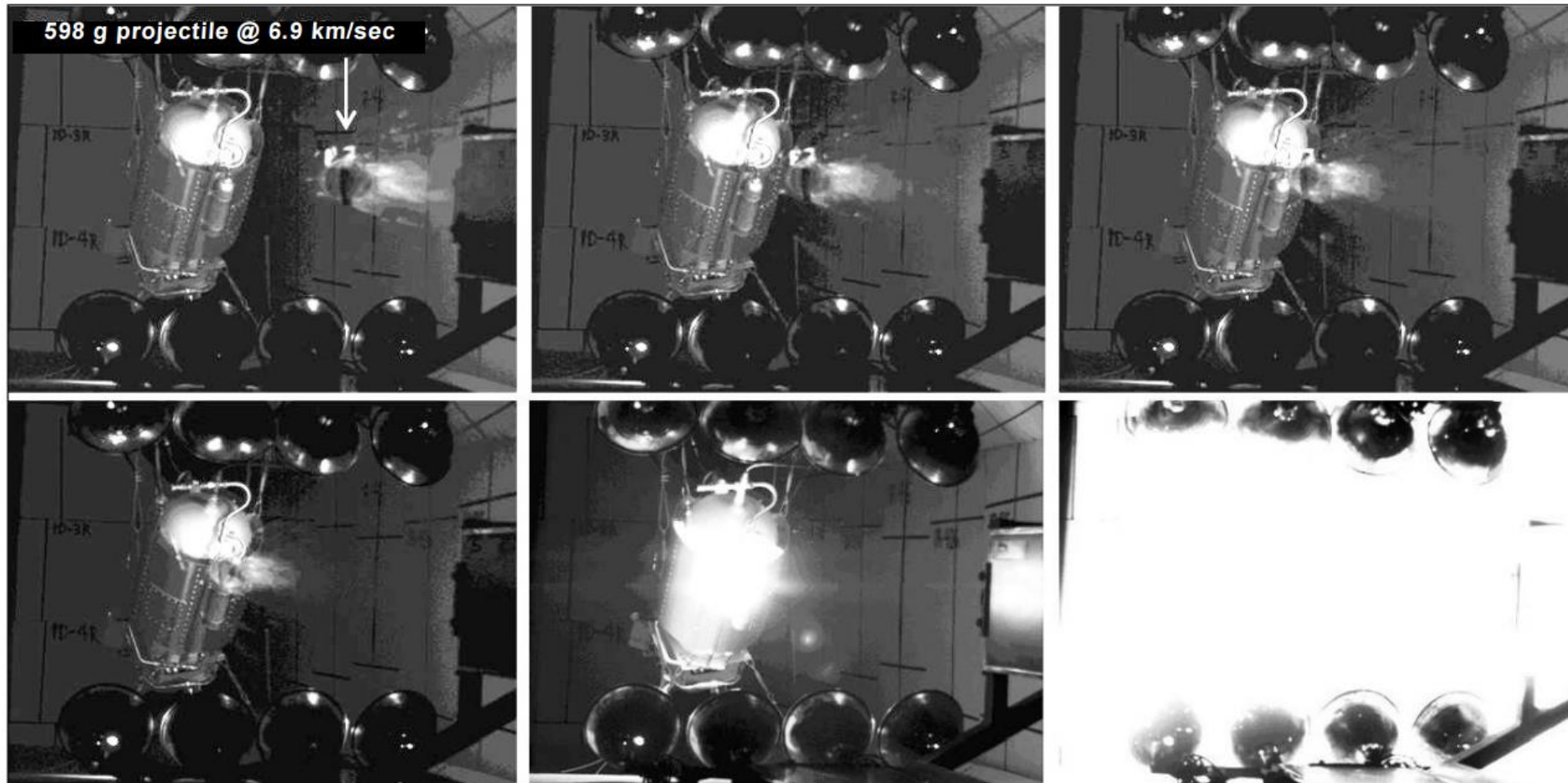
DebrisSat



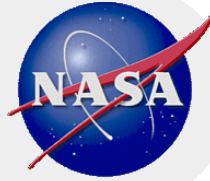
DebrisLV



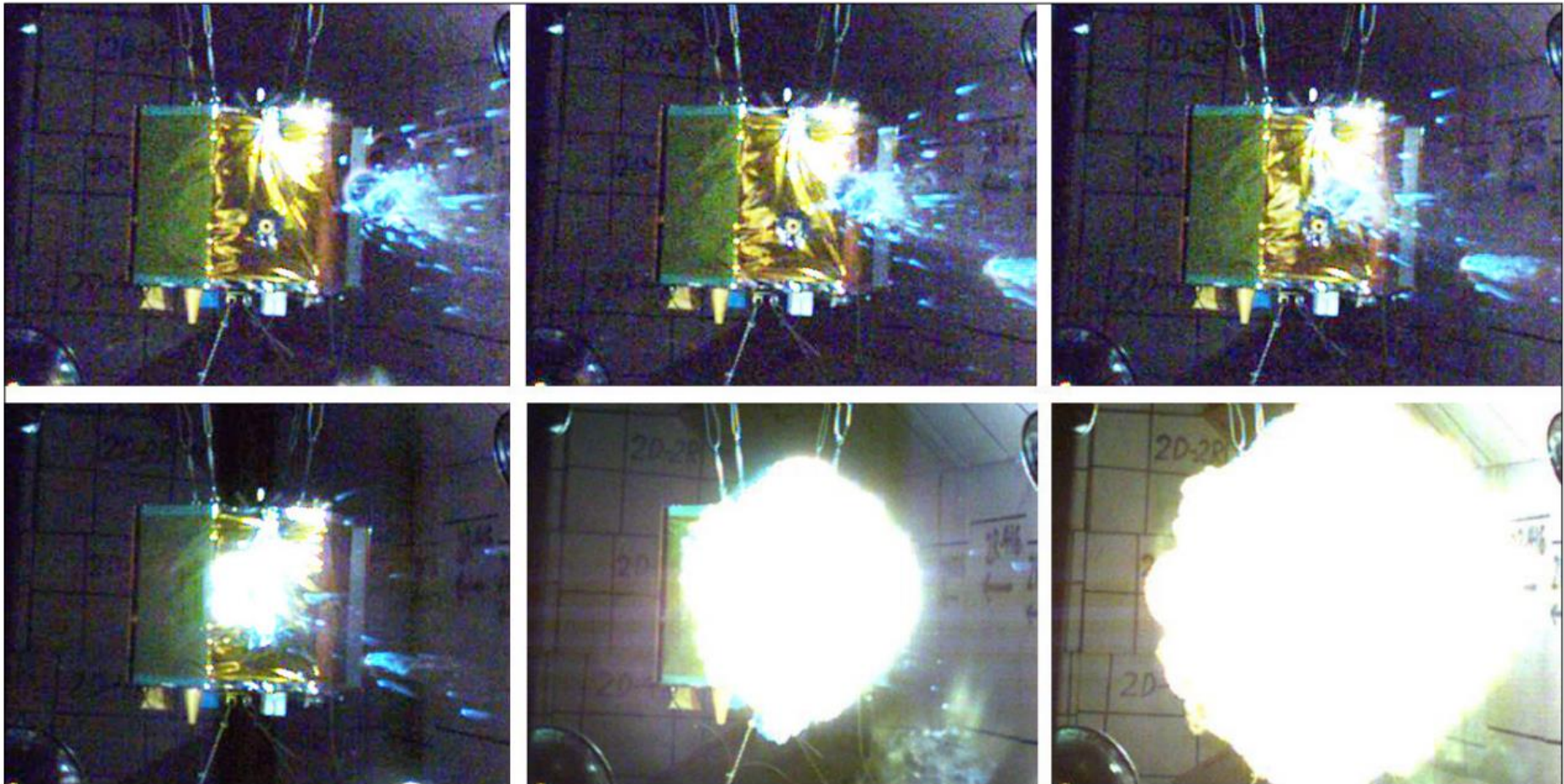
DebrisLV



Impact sequences of DebrisLV



DebrisSat



Impact sequences of DebrisSat

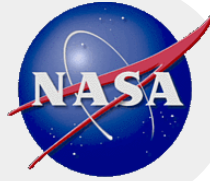


DebrisSat

- **The debris from the impacts have all been collected, and are being carefully analyzed by a team from the University of Florida**
 - Digital photos of each object
 - Mass, 3D dimensions
 - Material components identified
 - Soft-catch material being x-rayed to ascertain particle velocities and particle shapes

- **More debris were recovered than we anticipated based on previous models**

- **Final dataset will be a detailed resource**
 - Shape studies
 - RCS studies
 - Material distributions
 - Size distributions



ORDEM 3.1

- **An Engineering Model is a tool (primarily) for spacecraft designers and users to understand the long-term risks of debris collisions with their spacecraft**
- **NASA's Orbital Debris Engineering Model ORDEM represents NASA's best estimate of the current and near future orbital debris environment**
 - The environment is dynamic and must be updated periodically
 - Populations based on empirical data as much as possible
- **The ORDEM 3 series of models have significant new capabilities over previous ORDEM models**
 - Uncertainties
 - Material density categories
 - Model extended to GEO
 - Can easily calculate flux for satellites in highly elliptical orbit
- **ORDEM 3.1 is an update of the environment based on the latest data, but with minimal changes to the model structure**
 - Model completed and is undergoing review
 - Should be available later in 2019

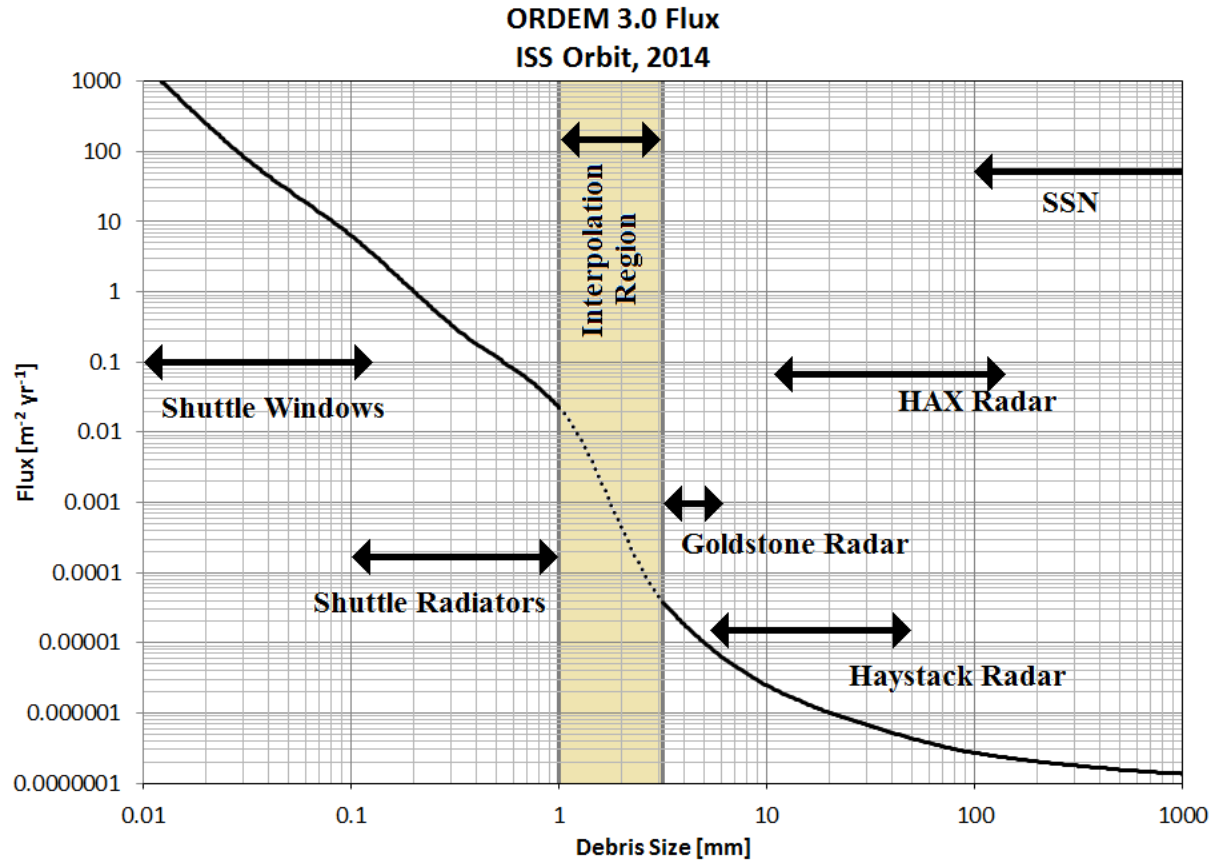


Data Sources for ORDEM Models

Data Source	Data Type	Size Limits	ORDEM 3.0	ORDEM 3.1
STS Windows and Radiators	<i>In situ</i>	10 μm - 1 mm	1995-2011	1995-2011
HST WFPC-2 Radiator	<i>In situ</i>	50 μm - 300 μm		1990-2009
HST Bay 5 MLI	<i>In situ</i>	10 μm - 300 μm		1990-2009
HUSIR/Haystack	Radar	>5.5 mm	1999-2003, 2007-2009	2013-2017
HAX	Radar	>1 cm	1999-2003, 2007-2009	
Goldstone	Radar	2 mm - 8 mm	2001, 2005- 2007, 2009	2016-2017
SSN Catalog	Radar	>10 cm	1957-2007	1957-2017
MODEST (GEO)	Optical	>30 cm	2004-2006	2004-2009, 2013-2014



Data and Size Regimes



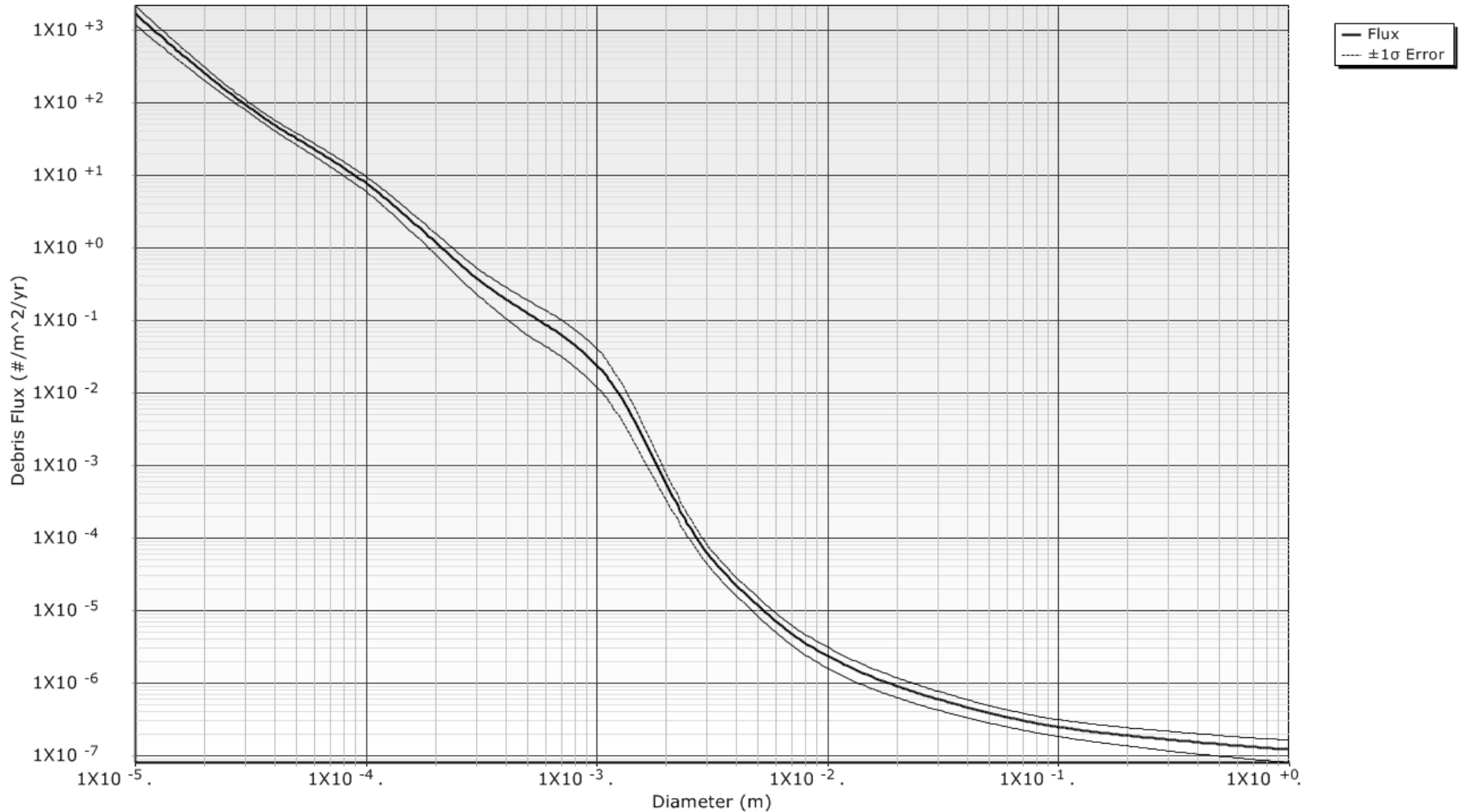
- Small particle populations are fit separately from large particle populations



ORDEM 3.0 Flux for ISS 400km

Average Cross-Sectional Flux vs. Size

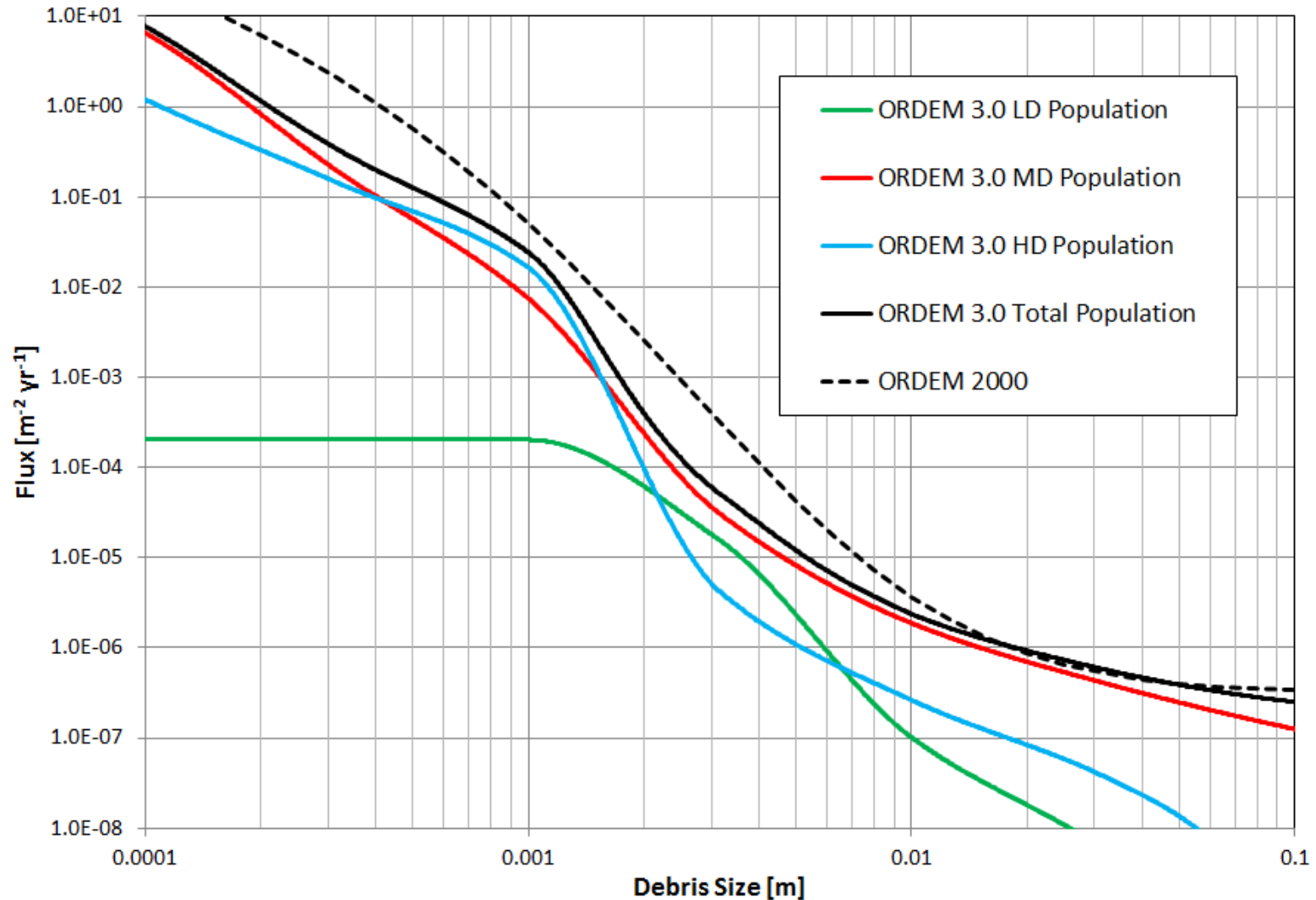
Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60





Material Distributions - ISS

ORDEM Populations for 2013 ISS Flux as a Function of Debris Size

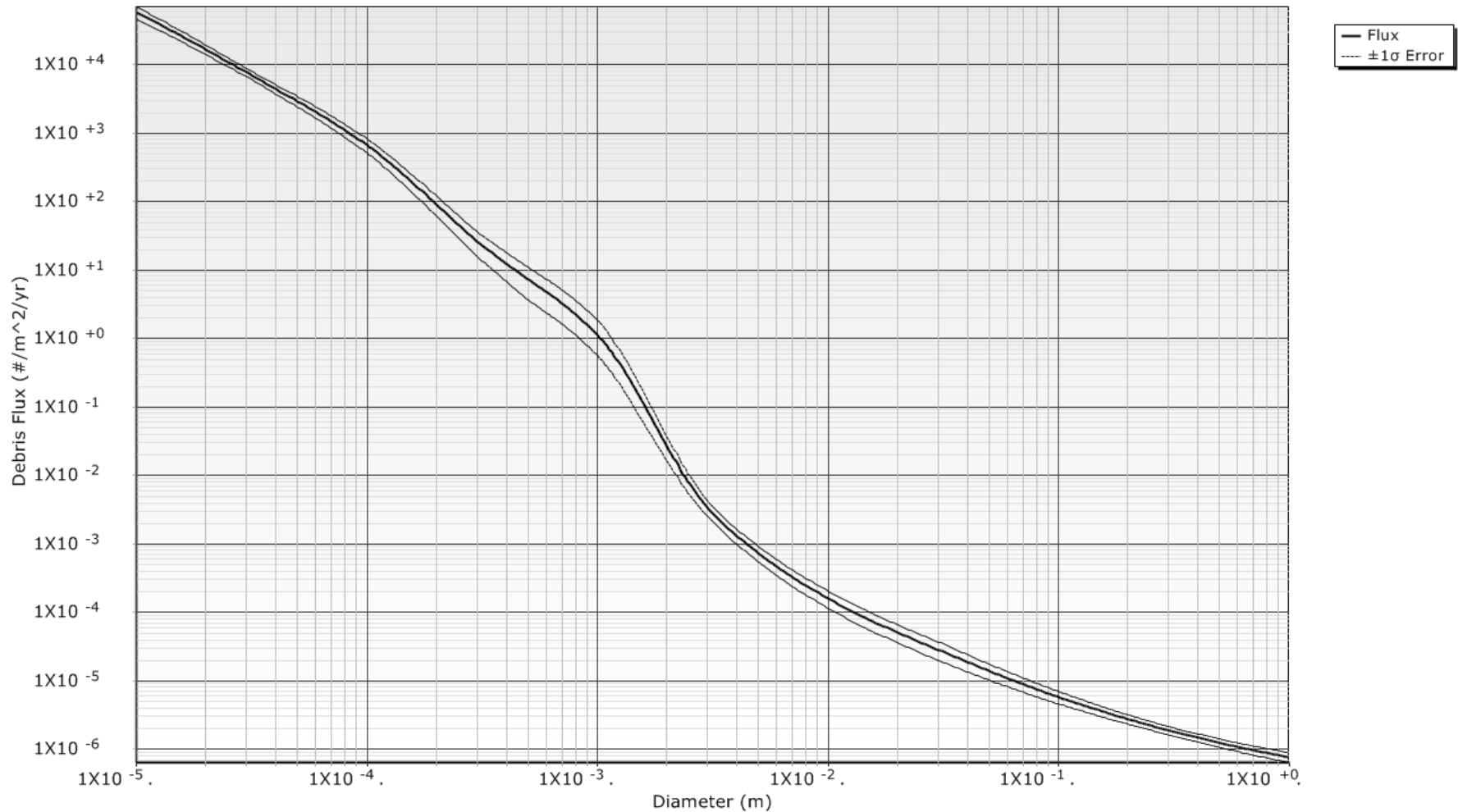


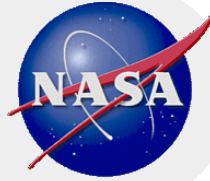


ORDEM 3.0 Flux for A-Train 705km

Average Cross-Sectional Flux vs. Size

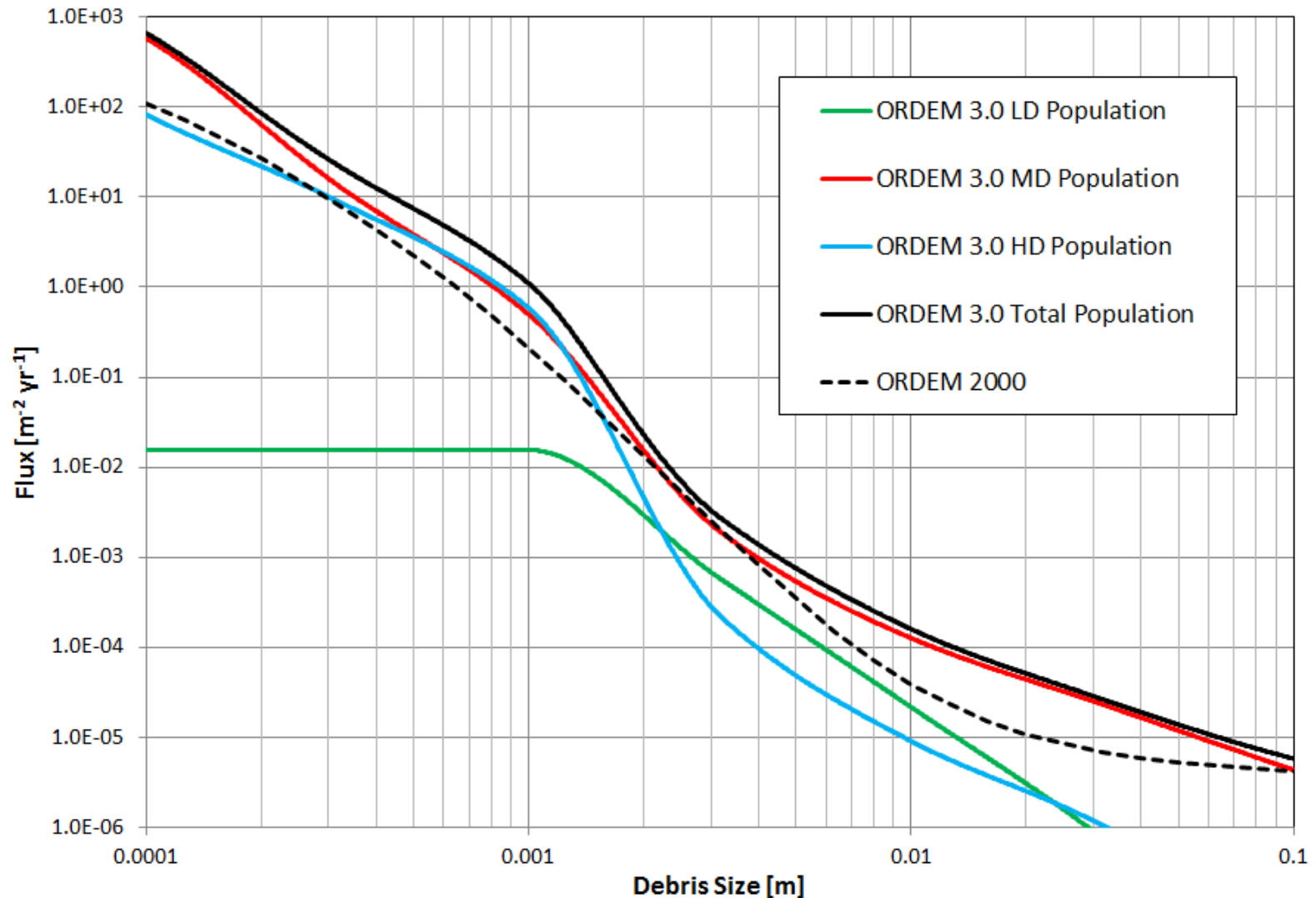
Year: 2013 Perigee Altitude = 705.000 Apogee Altitude = 705.000 inc = 98.00





Material Distribution – A-Train

ORDEM Populations for 2013 98° 705 km Orbit Flux as a Function of Debris Size

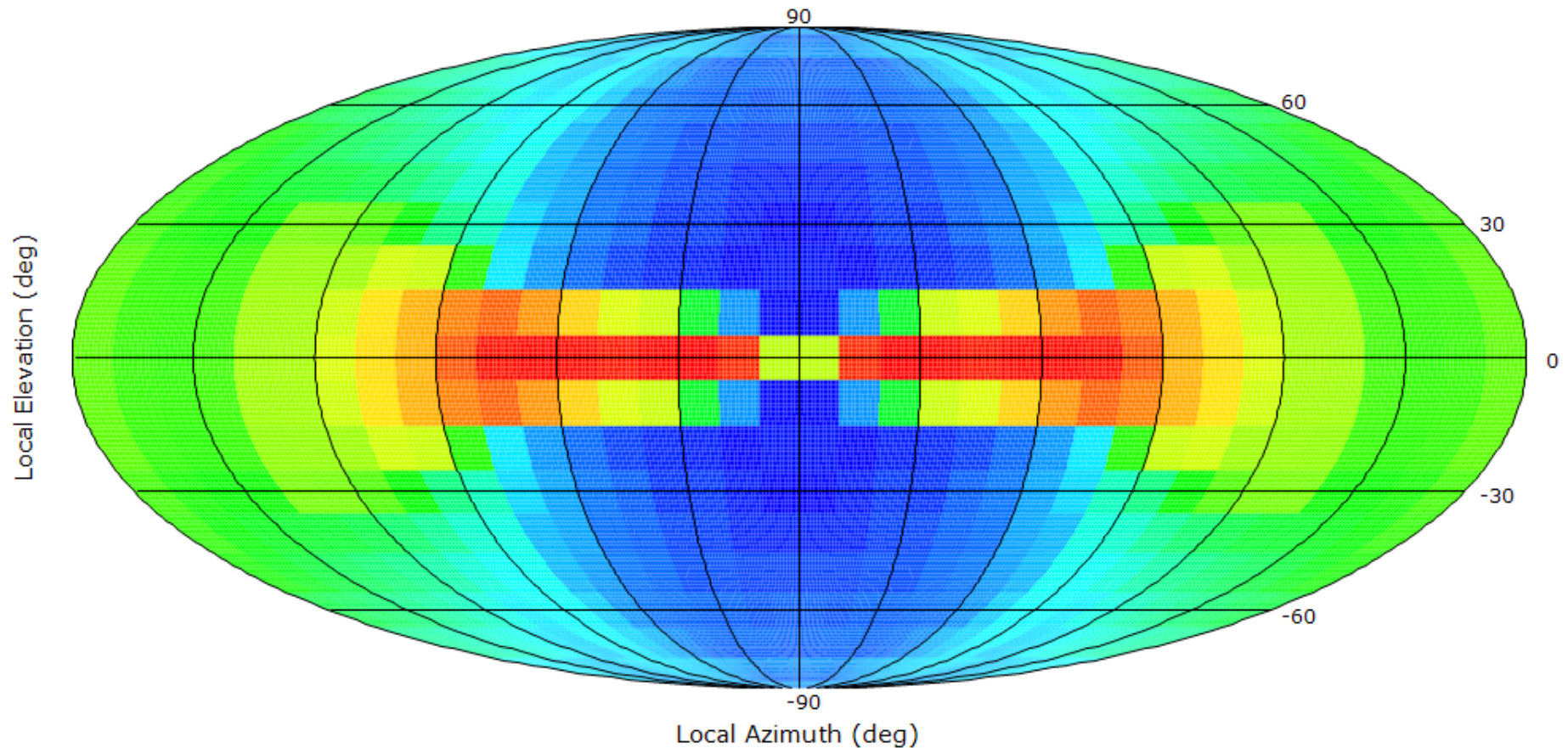




2D Flux Distribution

2-D Directional Flux

Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60 particle size = >1mm



7/23/2013 5:06:05 PM

10^{-12}

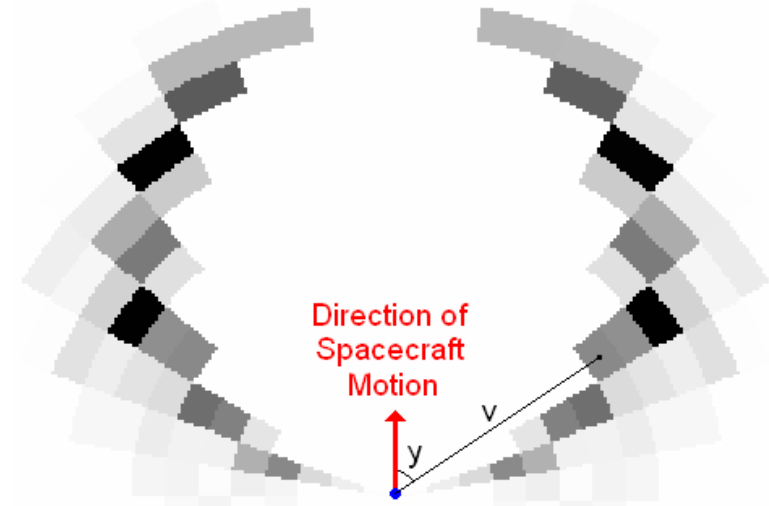
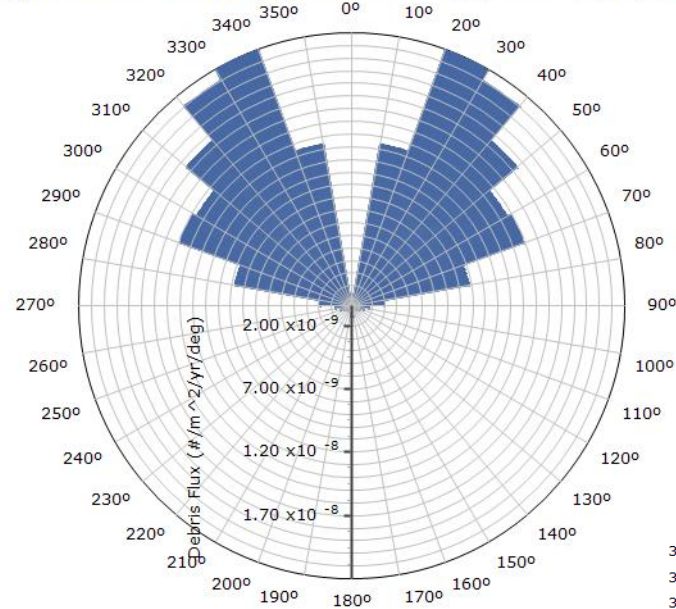
10^{-4}



Flux Dependence on Velocity and Direction

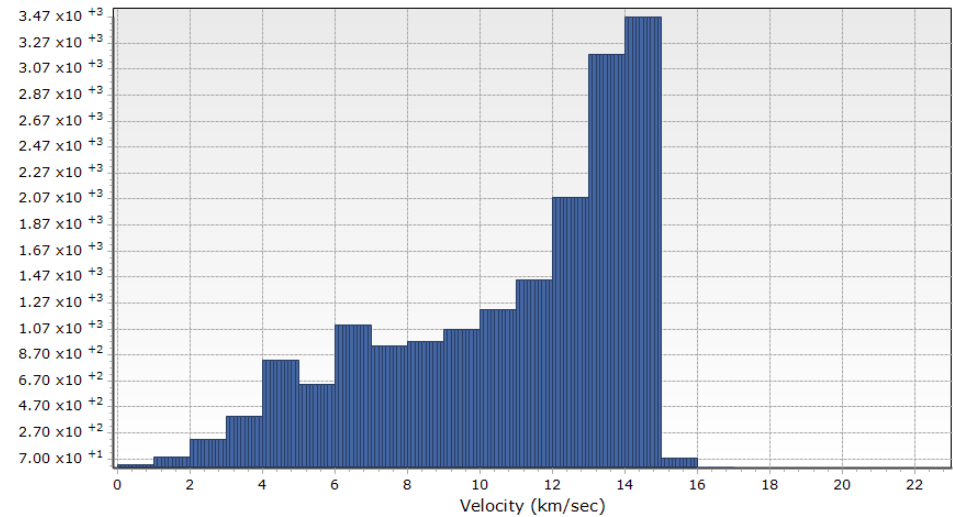
Flux vs. Local Azimuth

Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60 particle size = >1cm



Velocity Distribution

Year: 2019 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60 particle size = >10um



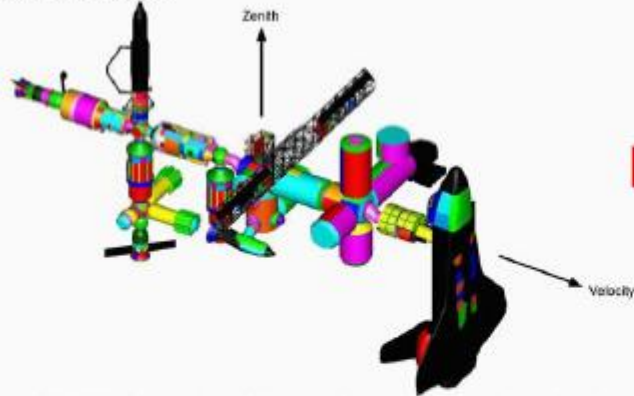


BUMPER

NASA/JSC BUMPER-II Meteoroid/Debris Threat Assessment Code

Spacecraft Configuration (I-DEAS Finite Element Model)

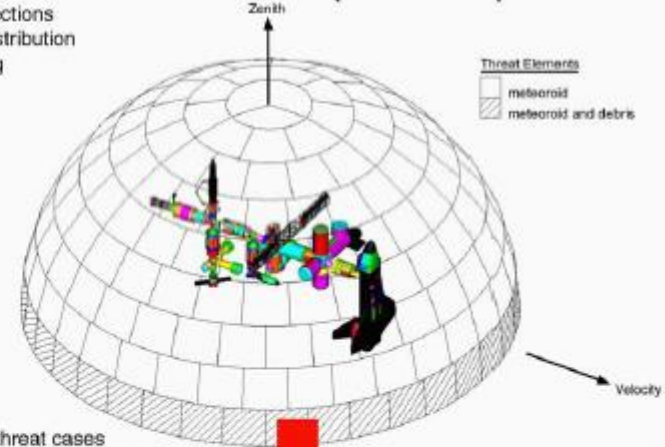
- Describes spatial relationships of spacecraft components
- Defines spacecraft orientation (velocity and zenith directions)
- Defines M/OD shield regions



- Approximately 120,000 elements in ISS assembly complete mated configuration FEM

Meteoroid & Debris Environments (GEOMETRY)

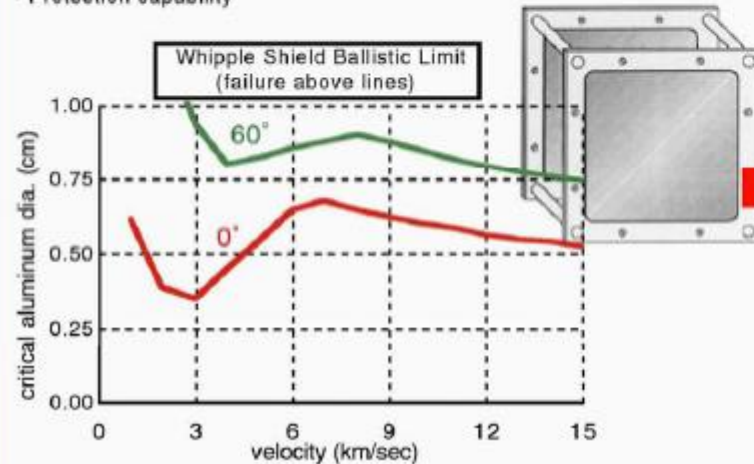
- Threat directions
- Velocity distribution
- Shadowing



- 90 debris threat cases and 149 meteoroid threat cases assessed for each element in the FEM

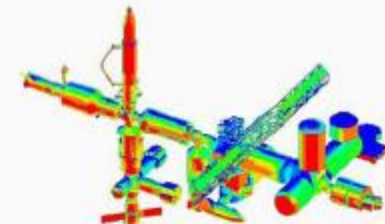
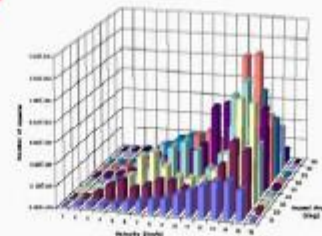
Critical Particle Diameter Calculation (RESPONSE)

- Protection capability



Computation of Penetrating Flux and PNP (SHIELD) Graphical Interpretation of Results (EXCEL & I-DEAS)

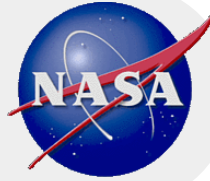
Station Region	Impact Risk From 1mm Q Debris		Debris Penetration Risk	
	Probability No Impact	Odds of Impact	Probability No Penetration	Odds of Penetration
FCB	0.996538	1/214	0.996541	1/224
Service Module	0.999336	1/1506	0.999796	1/4912
Node 2	0.990466	1/106	0.999966	1/625000
Hab Module	0.996074	1/29	0.998623	1/928
Lab Module	0.986522	1/69	0.999022	1/1025
CPV	0.997446	1/61	0.999630	1/6220
TOTALS	0.934622	1/15	0.996132	1/146



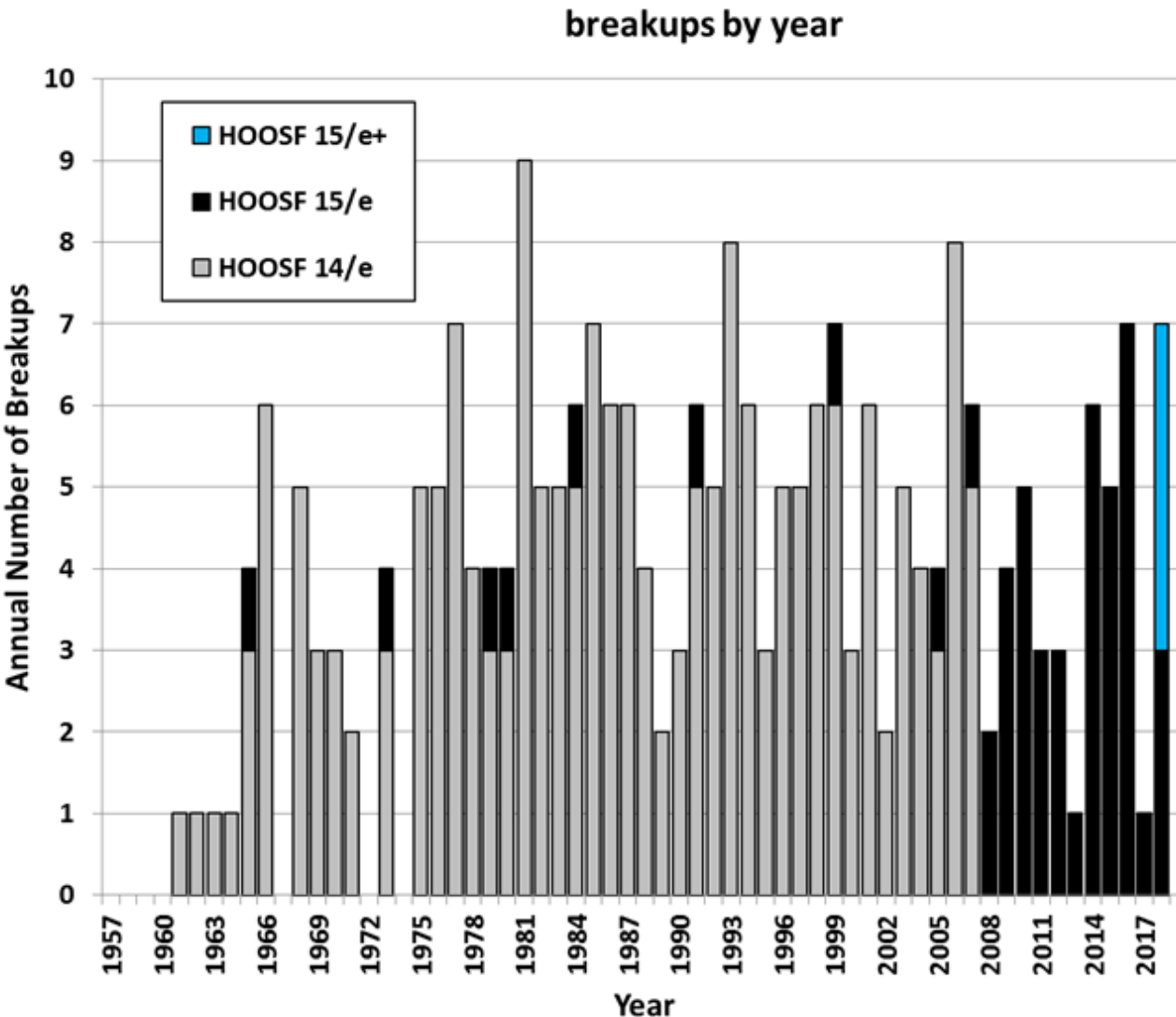


Spacecraft Environment Considerations

- **Orbital debris fluxes are a function of spacecraft orbit inclination and altitude**
- **Debris flux and velocity are direction-dependent**
 - **Custom multi-layer shields work best when optimized for particular velocities and directions**
- **Debris flux from different material types can skew the risk**
 - **High-density (e.g., steel) debris have a disproportionate effect on risk**
- **Spacecraft must design for the long-haul**
 - **A spacecraft will hopefully operate for many years**
 - **Risk is primarily a function of exposed area and exposed time**



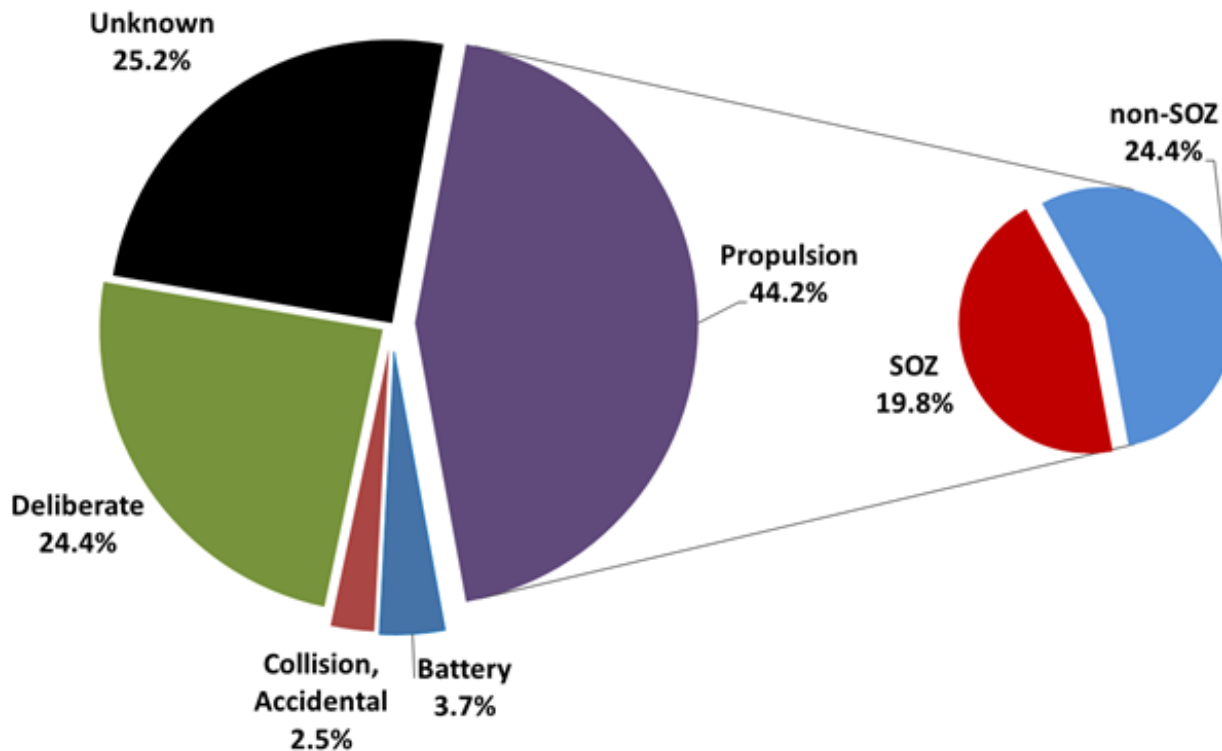
Breakups by year: **246** 1961-date



- The primary source of larger debris (> 1 cm) is from explosive breakups of spacecraft and rocket bodies
- HOOSF: the NASA ODPO History of On-orbit Satellite Fragmentations
 - 14th ed. published 2008
 - 15th ed. published 2018
 - Four events occurred in 2018 *after* information cut-off for 15th ed.
 - More have already occurred in 2019



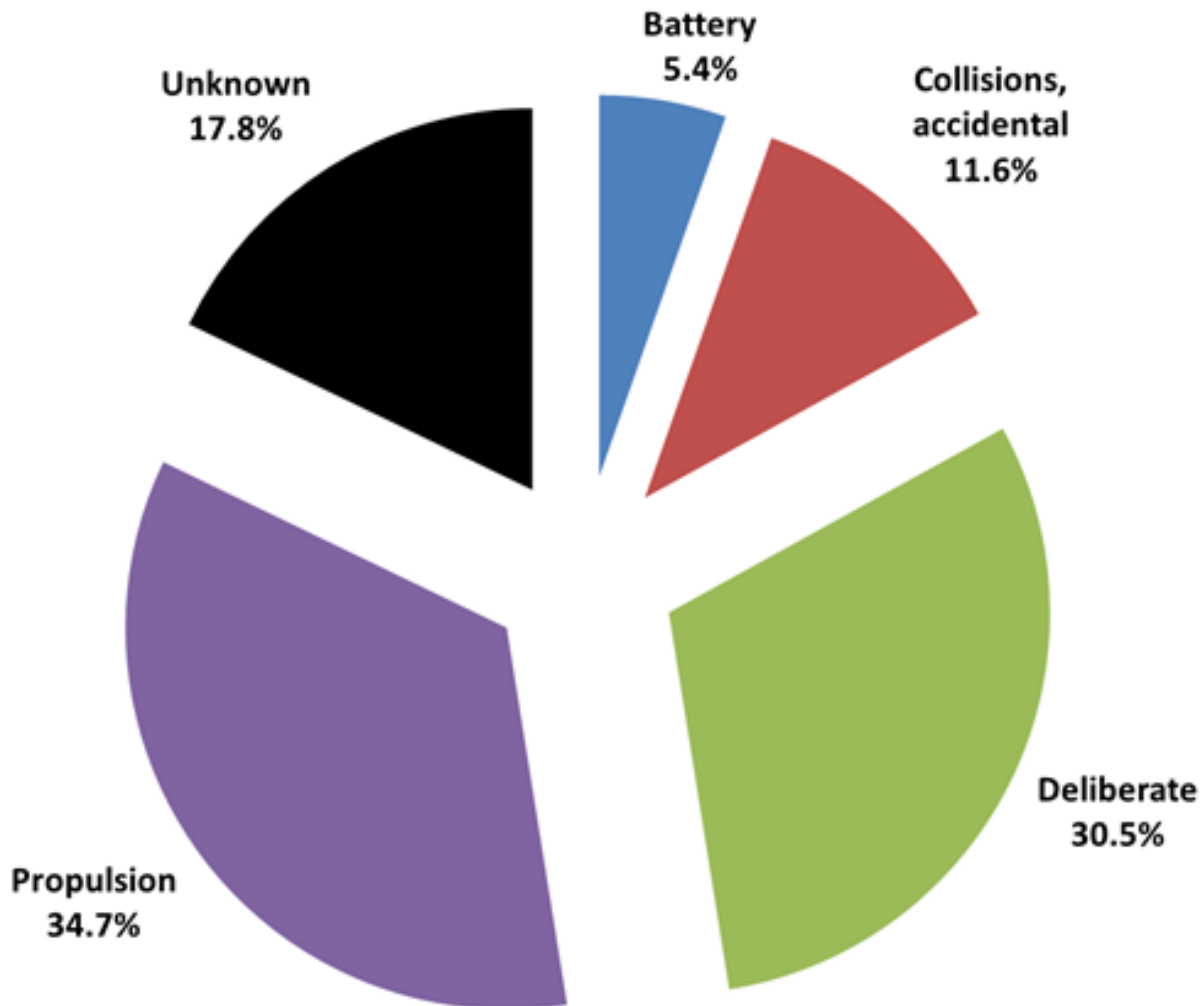
Breakups by cause: 246 1961-date



- Propulsion category accounts for **majority** of breakup events
- *SOZ units* are Proton 4th stage ullage motors
- When **SOZ** breakups are segregated, % of propulsion breakups equals historical deliberate breakup events
- **Unknown** category includes events whose root cause has not been uniquely identified or the breakup mechanism is unknown

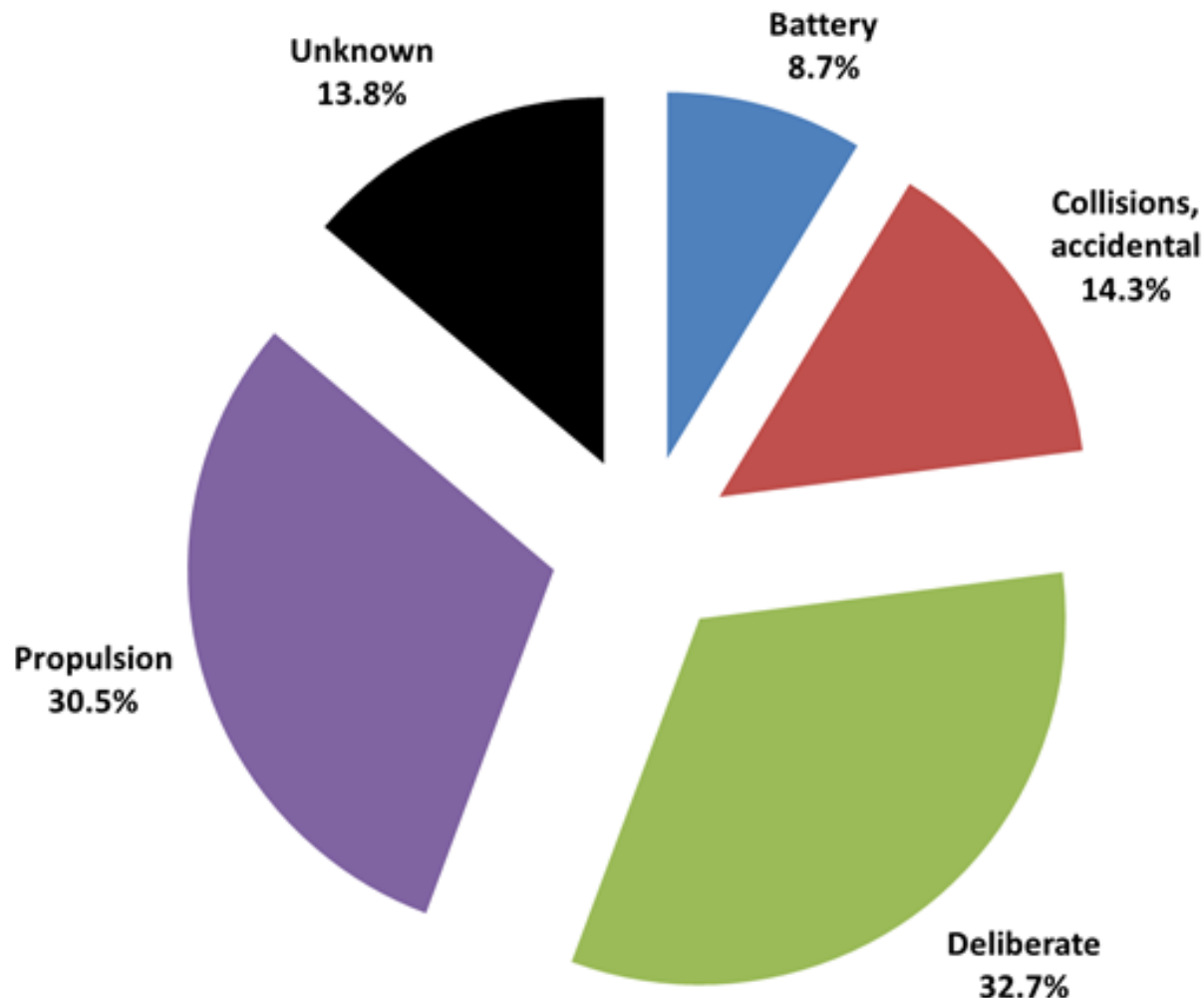


Cataloged breakup debris, 1961-date



- Multiply by **20044**, the total number of breakup debris cataloged, to get absolute number in any category
- Propulsion category accounts for **majority** of breakup debris cataloged
- While SOZ breakups typically result in few cataloged fragments, their eccentric parent orbits pose challenges to cataloging

On-orbit Cataloged breakup debris, 1961-date



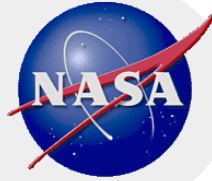
- Multiply by **9953**, the total number of cataloged breakup debris *remaining on orbit*, to get absolute number in any category
- Deliberate category accounts for **majority** of breakup debris on orbit due to intentional FY-1C Anti-Satellite (ASAT) weapon test in 2007
- While SOZ breakups typically result in few cataloged fragments, their eccentric parent orbits pose challenges to cataloging



Example Breakup - BRIZ-M

- **On August 6, 2012, the Russians attempted to launch two communications satellites using a Proton rocket**
- **The BRIZ-M upper stage failed to burn properly, and was left stranded in an elliptical orbit with about 5 metric tons of its propellant still aboard**
- **On October 16, the rocket body exploded, creating at least 700 trackable pieces of debris (and probably many more too small to be tracked) in orbits that cross ISS altitude**
- **Observed by astronomers at the Siding Springs Observatory**





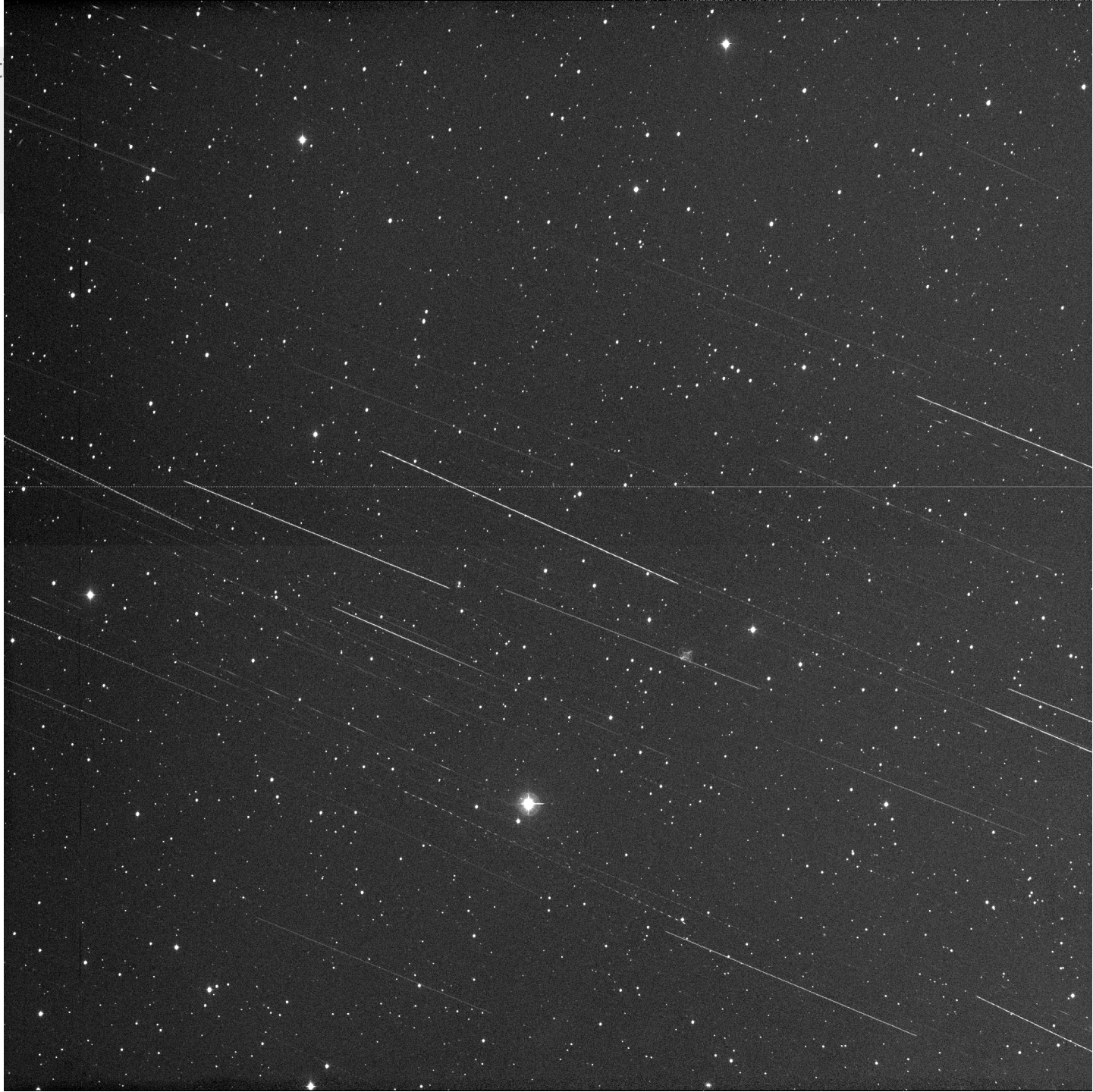
BRIZ-M Breakup

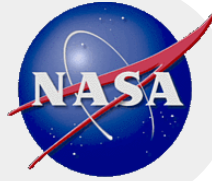


2012/10/16 14:05:00 UT



Nat



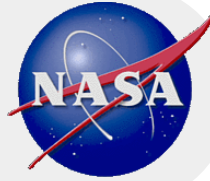


Previous Briz-M explosion – Feb 19, 2007

Rob McNaught,
Siding Springs Observatory

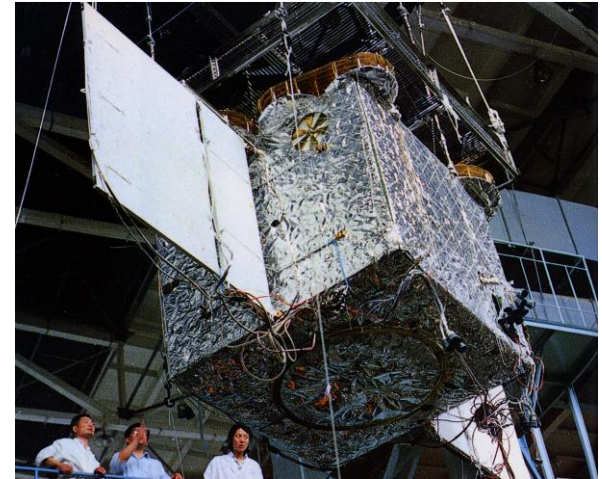


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www.NaturesPeak.com.au

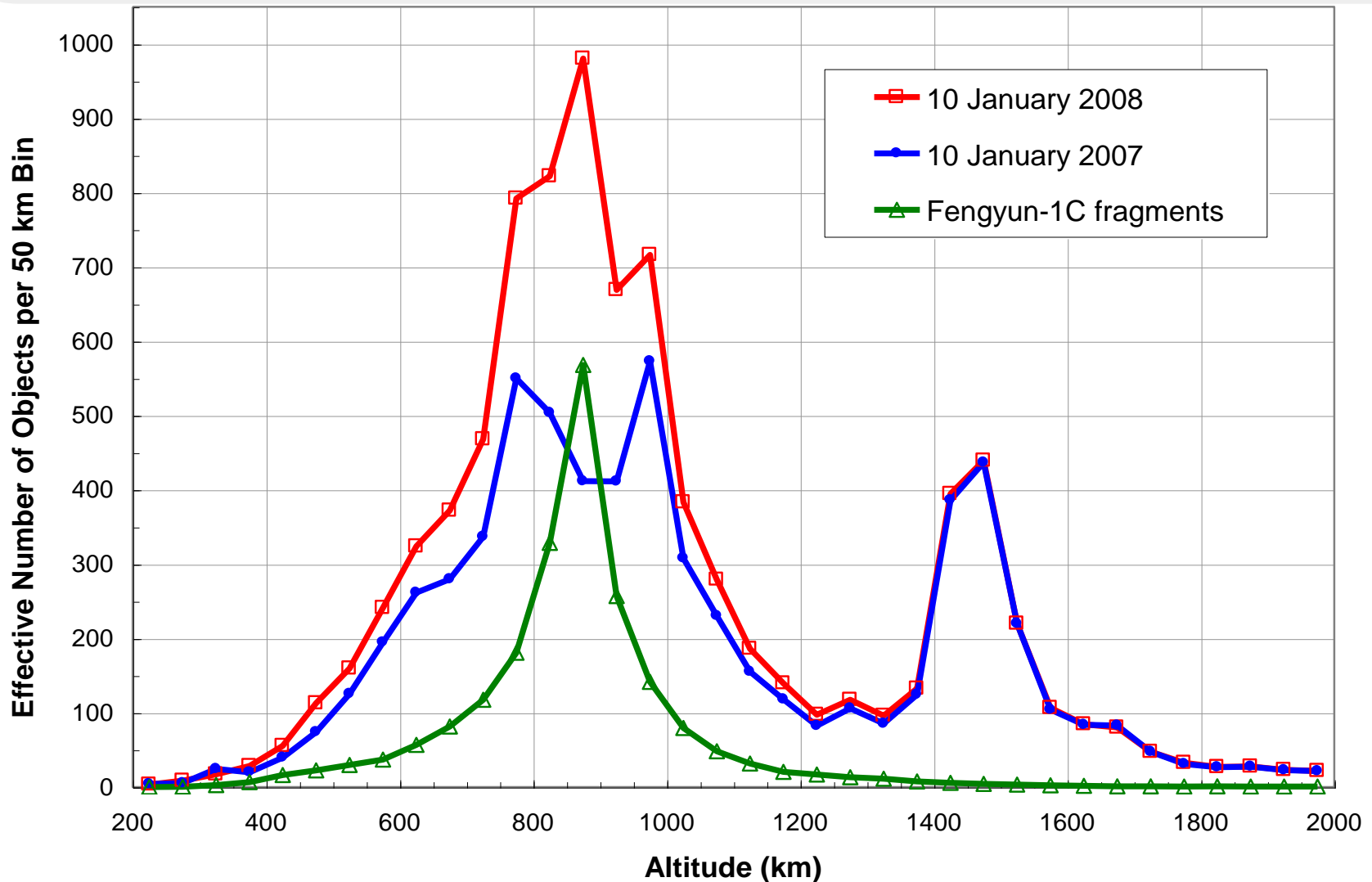


Chinese ASAT - Fengyun-1C

- **950 kg Chinese weather satellite**
- **865 km x 845 km, 98.6° orbit**
- **Destroyed by Chinese military using a ground-based anti-satellite (ASAT) missile on January 11, 2007**
- **Created an unprecedented number of tracked debris**



Effect of a Single Event (Catalog Populations in LEO)

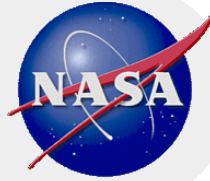




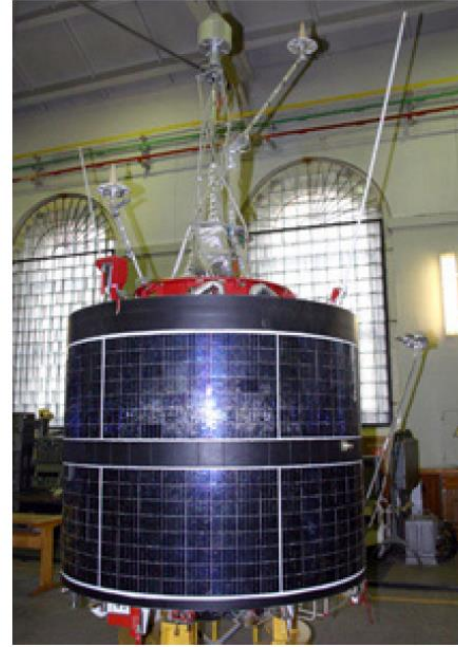
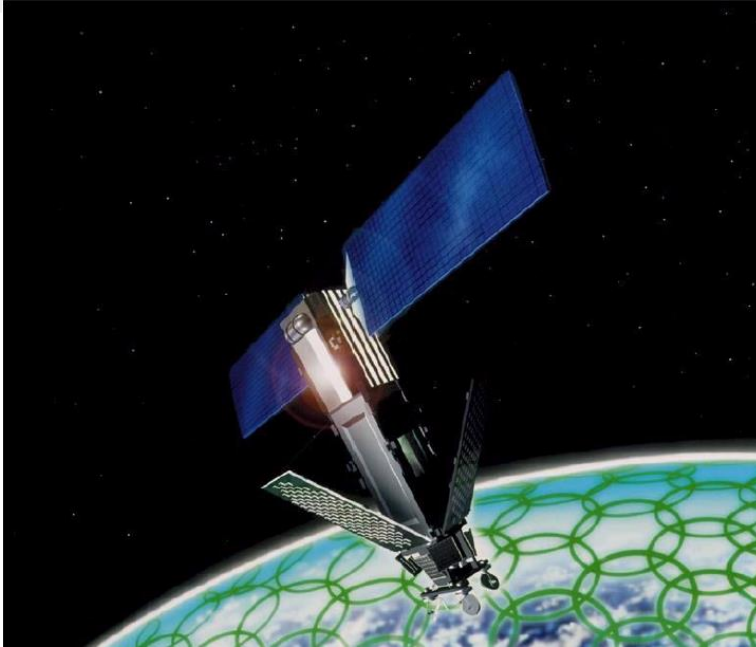
Indian ASAT

- On March 27, 2019, Indian announced it had successfully destroyed one of its own satellites with an ASAT weapon
- The target was destroyed at an altitude where most of the debris would likely reentry in a few weeks to months
- 90 debris catalogued so far
- **Microsat-R**
 - 96.6° inclination
 - 291x252 km
 - 740 kg





2009 Collision



February 10, 16:56 GMT two satellites collided near 789 km altitude

Iridium 33 (24946, 97051C)

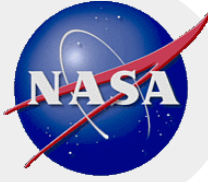
779 x 808 km, 86.4° orbit, 556 kg

Operational US Commercial Communication Satellite

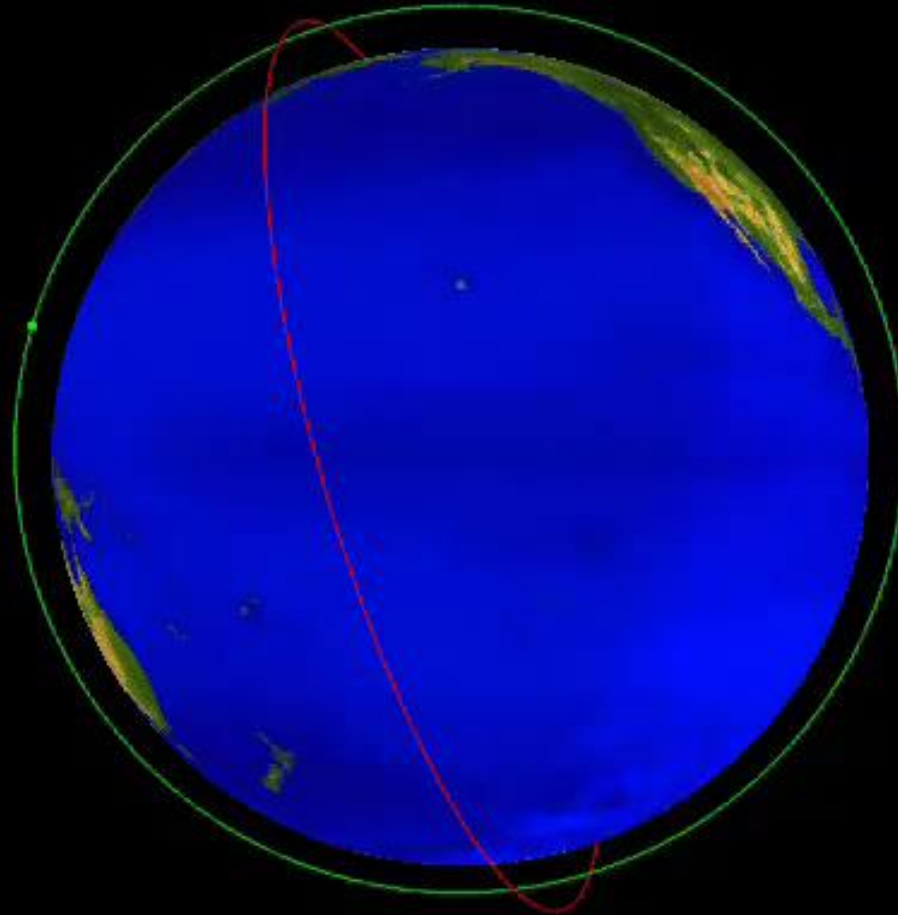
Kosmos 2251 (22675, 93036A)

786 x 826 km, 74.0° orbit, 900 kg

Non-operational Russian Communication Satellite



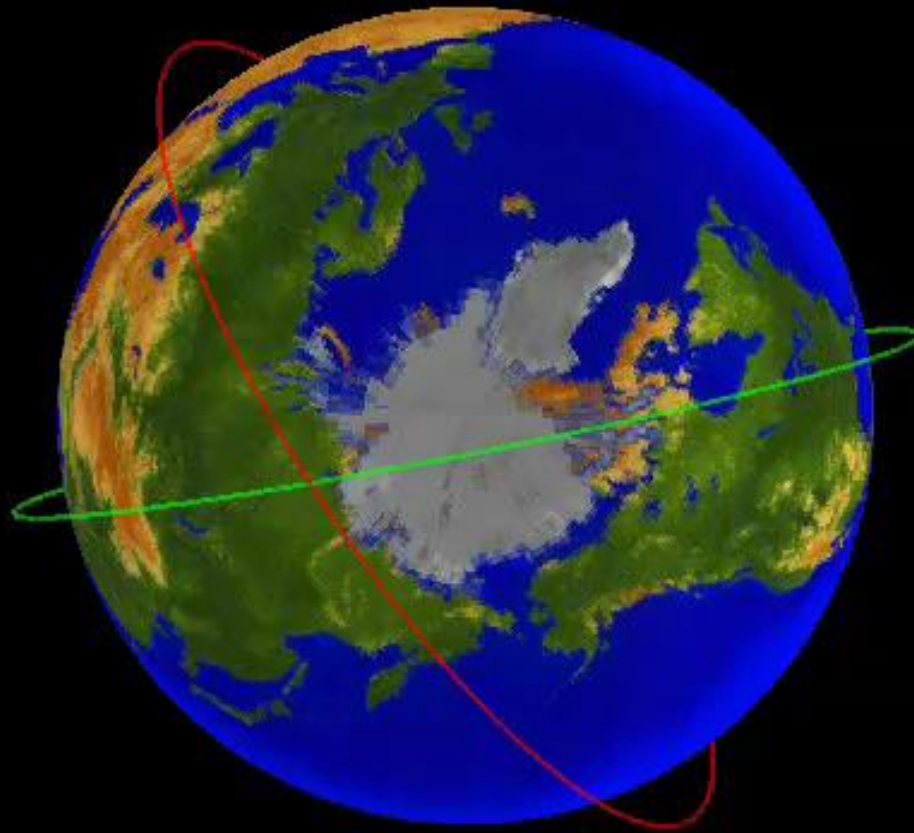
Iridium Collision



2009/02/10 15:00:00 UT



Iridium Collision

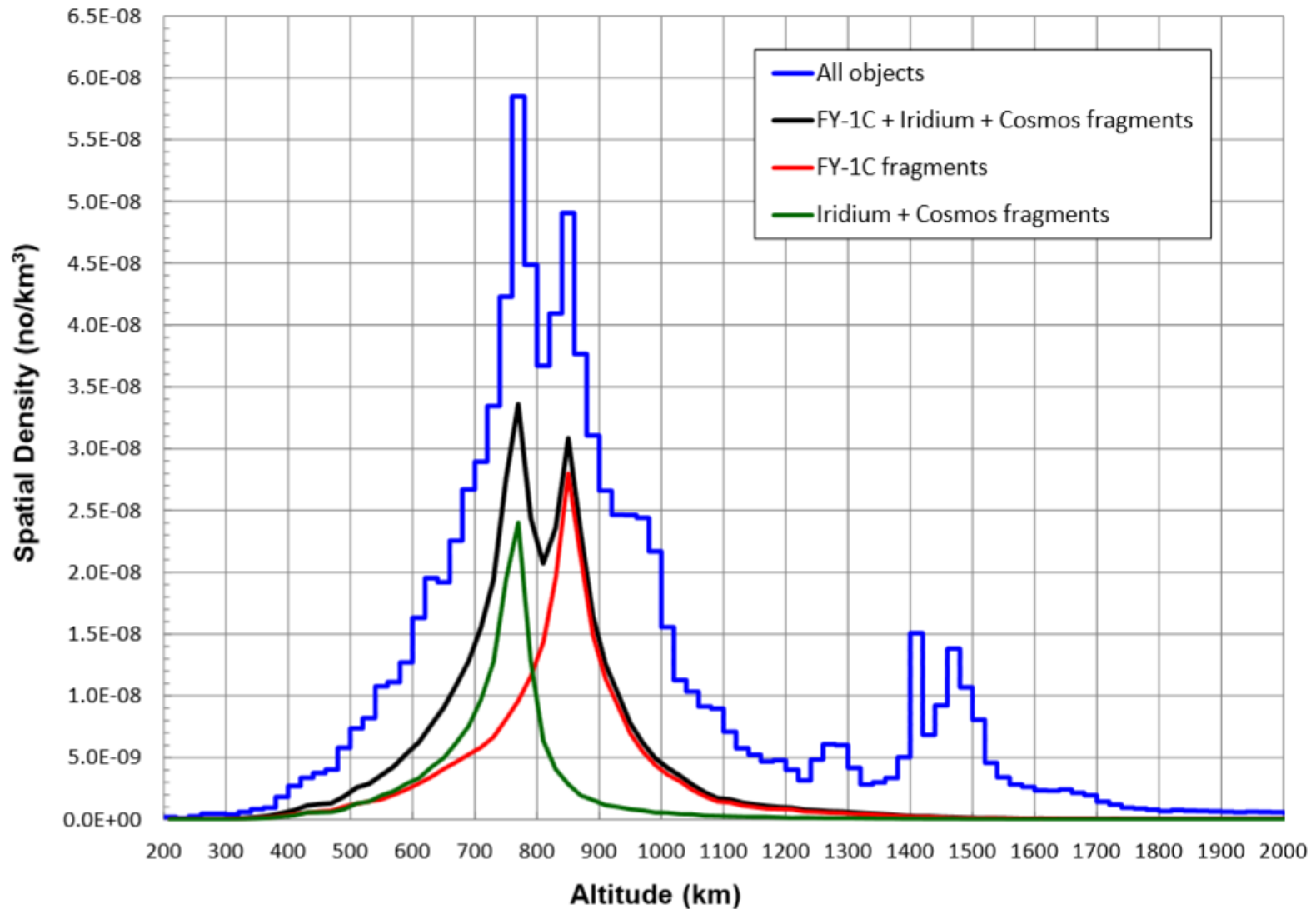


2009/01/15 18:38:16 UT



Effect of Collision on Catalog

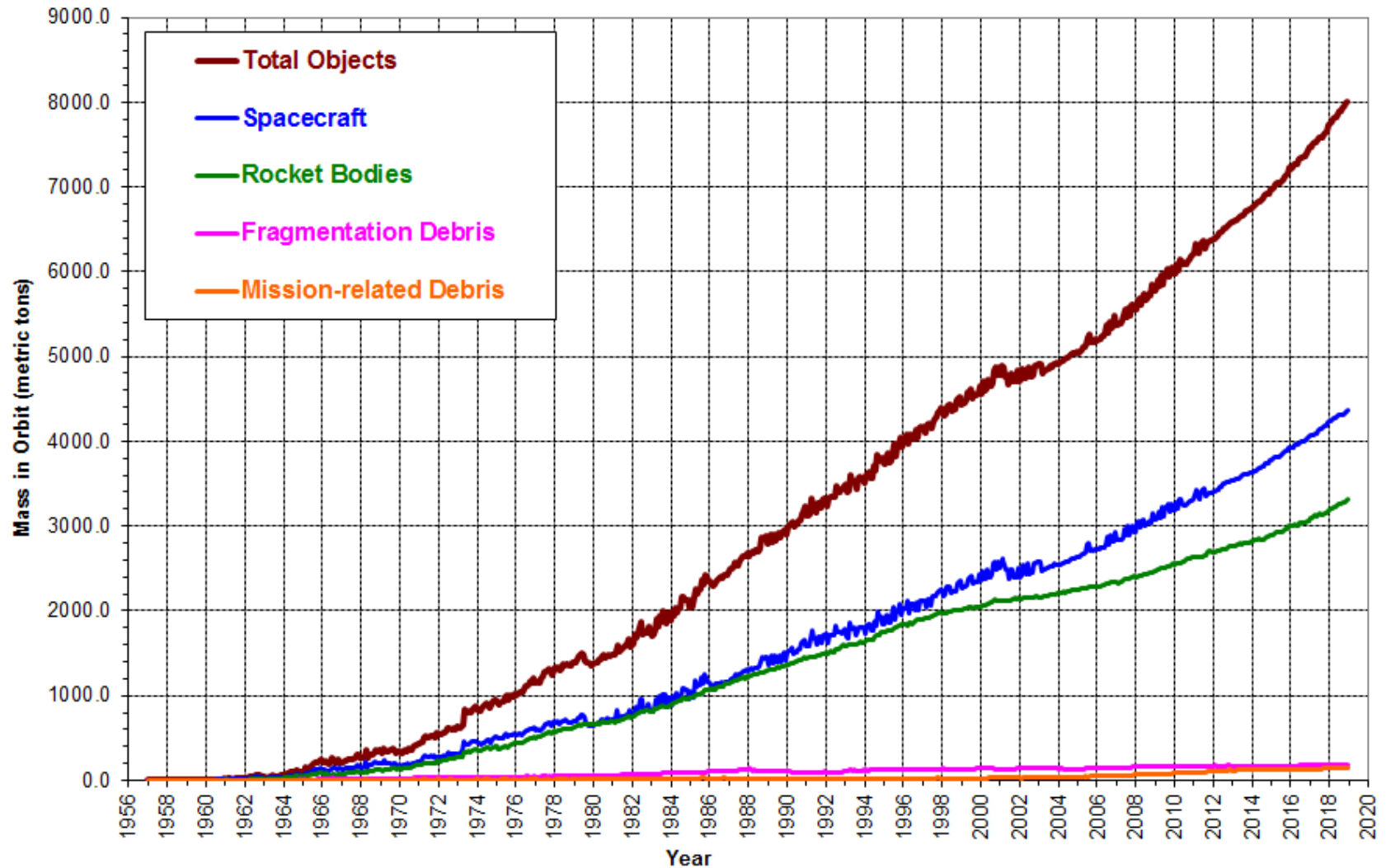
January 2013 Catalog

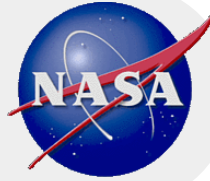




Evolution of the Catalogued (>10 cm) Satellite Population by Mass

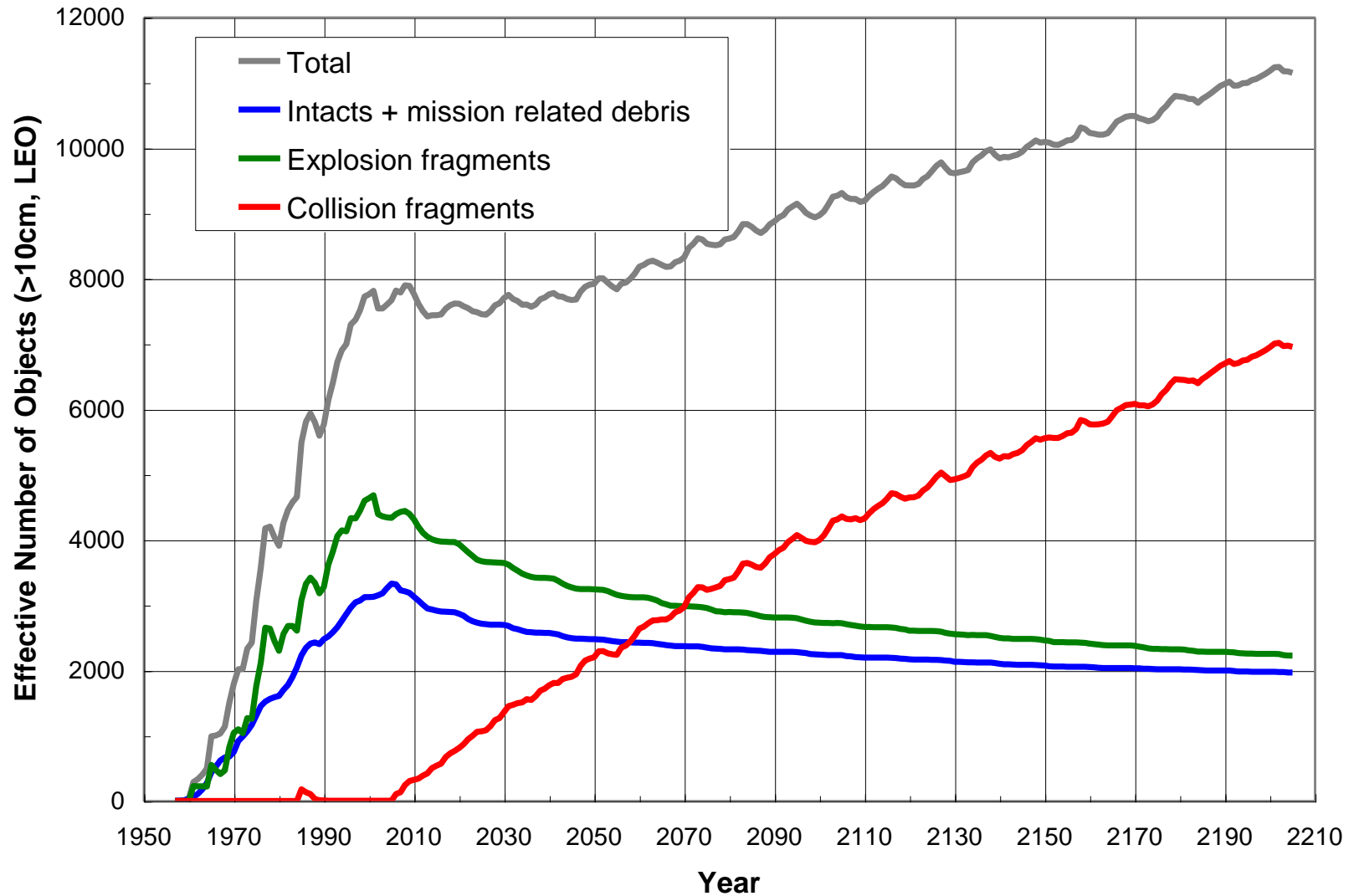
Monthly Mass of Objects in Earth Orbit by Object Type





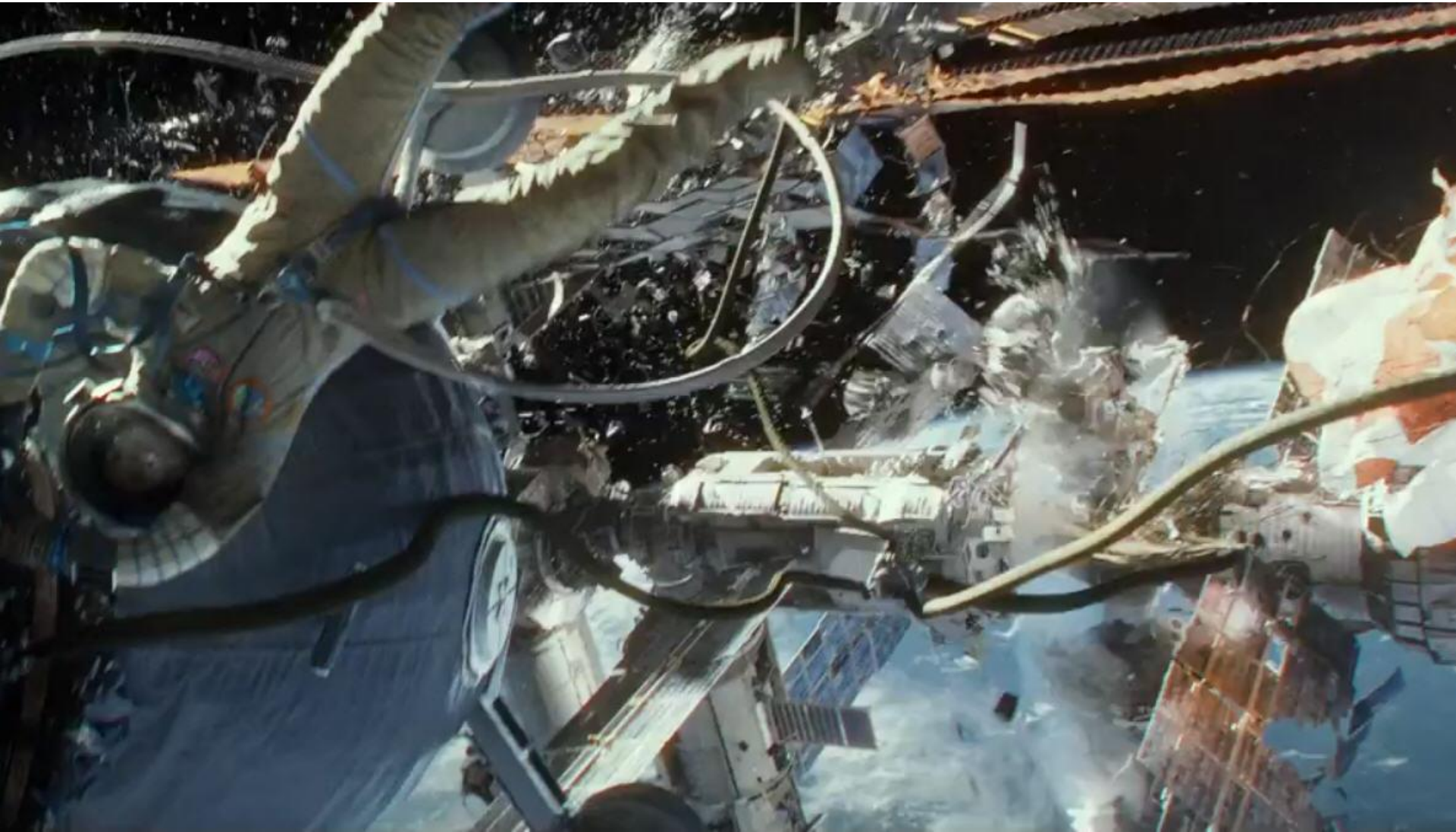
Growth with no future launches

Kessler Syndrome





Gravity





Gravity





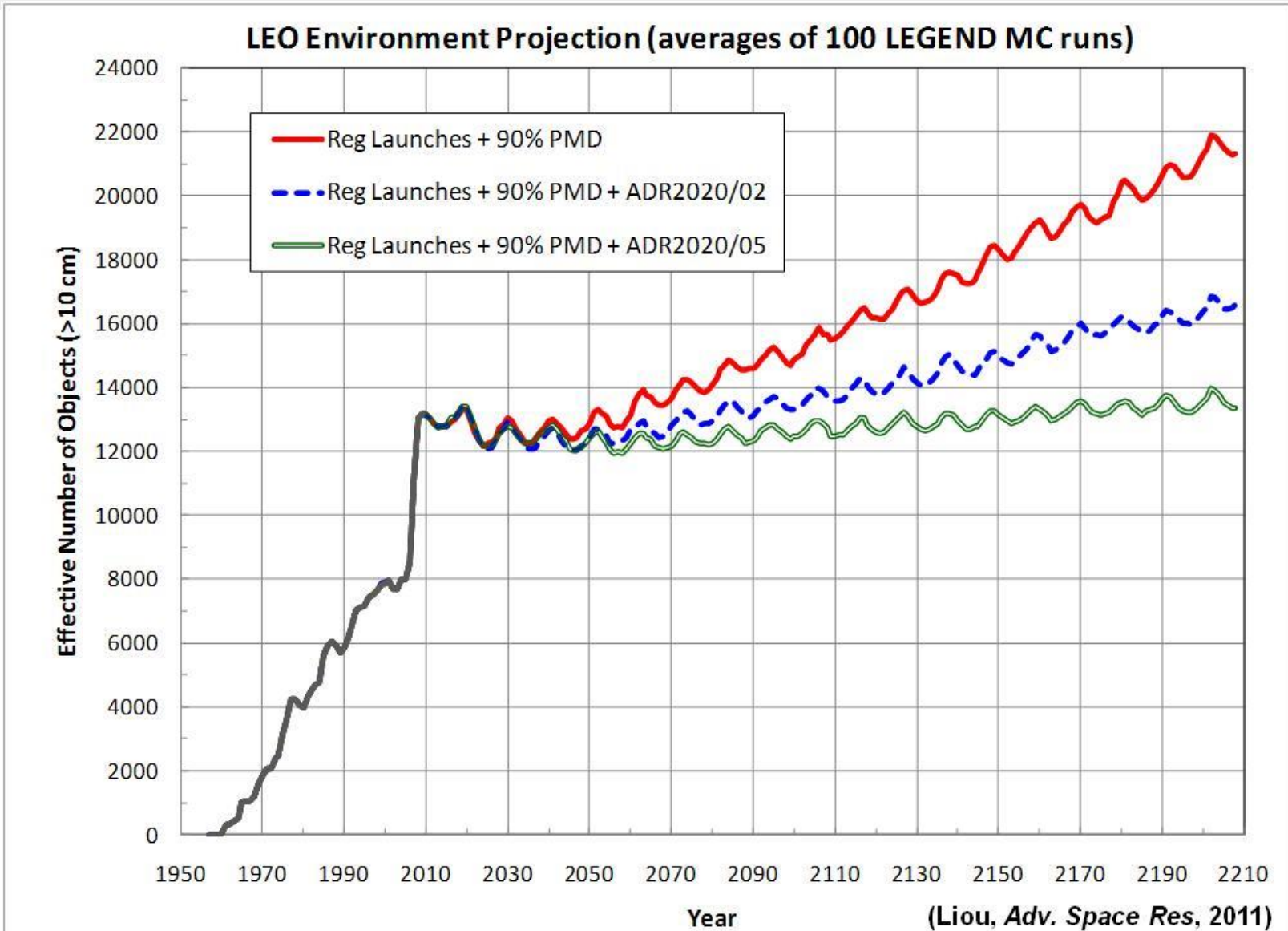
Rubes

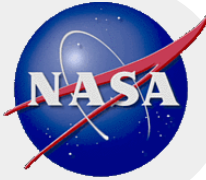


"Well, I'll be ... I guess the little chicken was right."

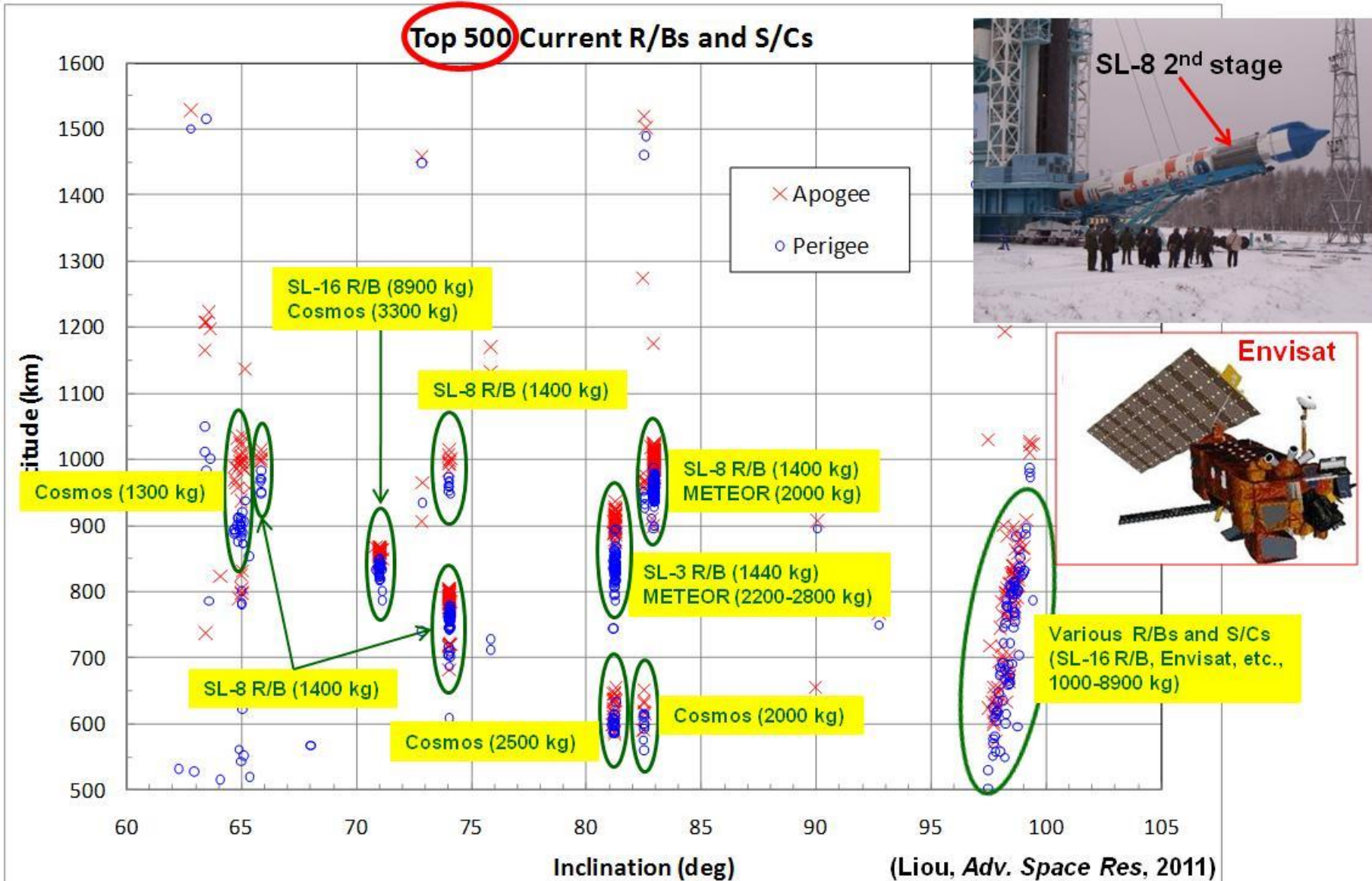


Fix the Problem? – Remove Mass





Highest Mass Objects





Active Debris Removal

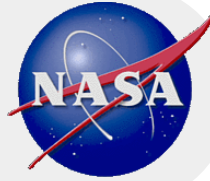
- **ESA has begun experimenting with technologies that might be used for active debris removal**



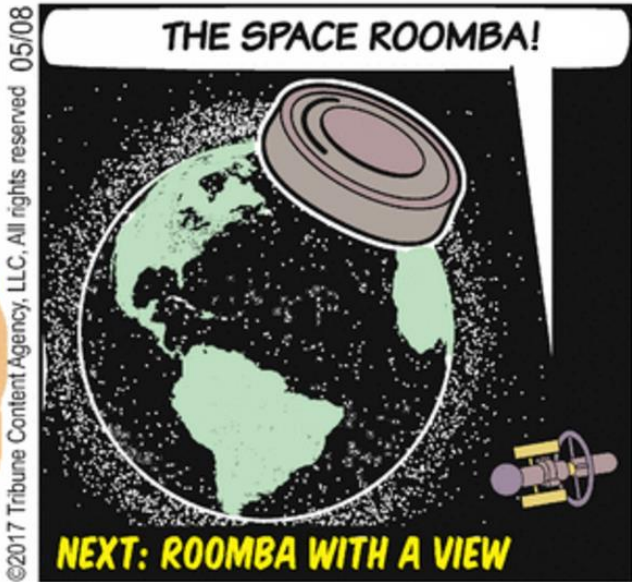


Active Debris Removal - 1965 (!)





Active Debris Removal – 2019?



www.gocomics.com/brewsterrockit brewrockit@yahoo.com

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Inter–Agency Space Debris Coordination Committee

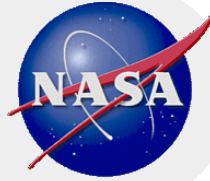


- **The Interagency Space Debris Coordination Committee (IADC) is composed of subject matter experts from 13 spacefaring nations, who meet together annually and address technical and policy issues**
- **Space Debris is a regular topic at the UN’s Committee on the Peaceful Uses of Outer Space (COPUOS)**
- **NASA has worked closely with the US government, IADC, and UN to come up with non-binding (but taken seriously nevertheless) “guidelines” for what a “good citizen” does in space:**
 - **Don’t make any messes you can’t clean up – do not create lots of long-lived debris**
 - **Clean up after yourself – make sure to remove satellites and rocket bodies from busy regions of space within 25 years after end of use and passivate them so they don’t explode later**
 - **Don’t hurt anyone – design your spacecraft and/or mission profile to minimize risk to other missions and people on the ground**



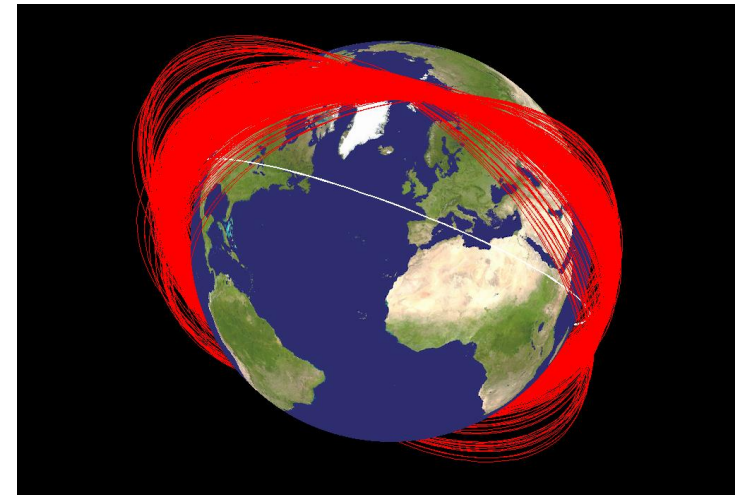
Conclusions

- **Monitoring the Earth space environment is critical**
 - **SSN catalog insufficient to characterize all debris**
 - **Environment is dynamic – even if we get it right today, it will change tomorrow**
- **With the loss of the Space Shuttle, new *in situ* data sources are needed to understand the small particle environment**
- **Models provide spacecraft designers and operators with tools to be able to make informed decisions about the safety of their space activities**
- **Models provide policy makers with tools to be able to make informed decisions about guidelines and regulations concerning space activities**
- **However, models are only as good as the assumptions made and the quality of the data behind them**



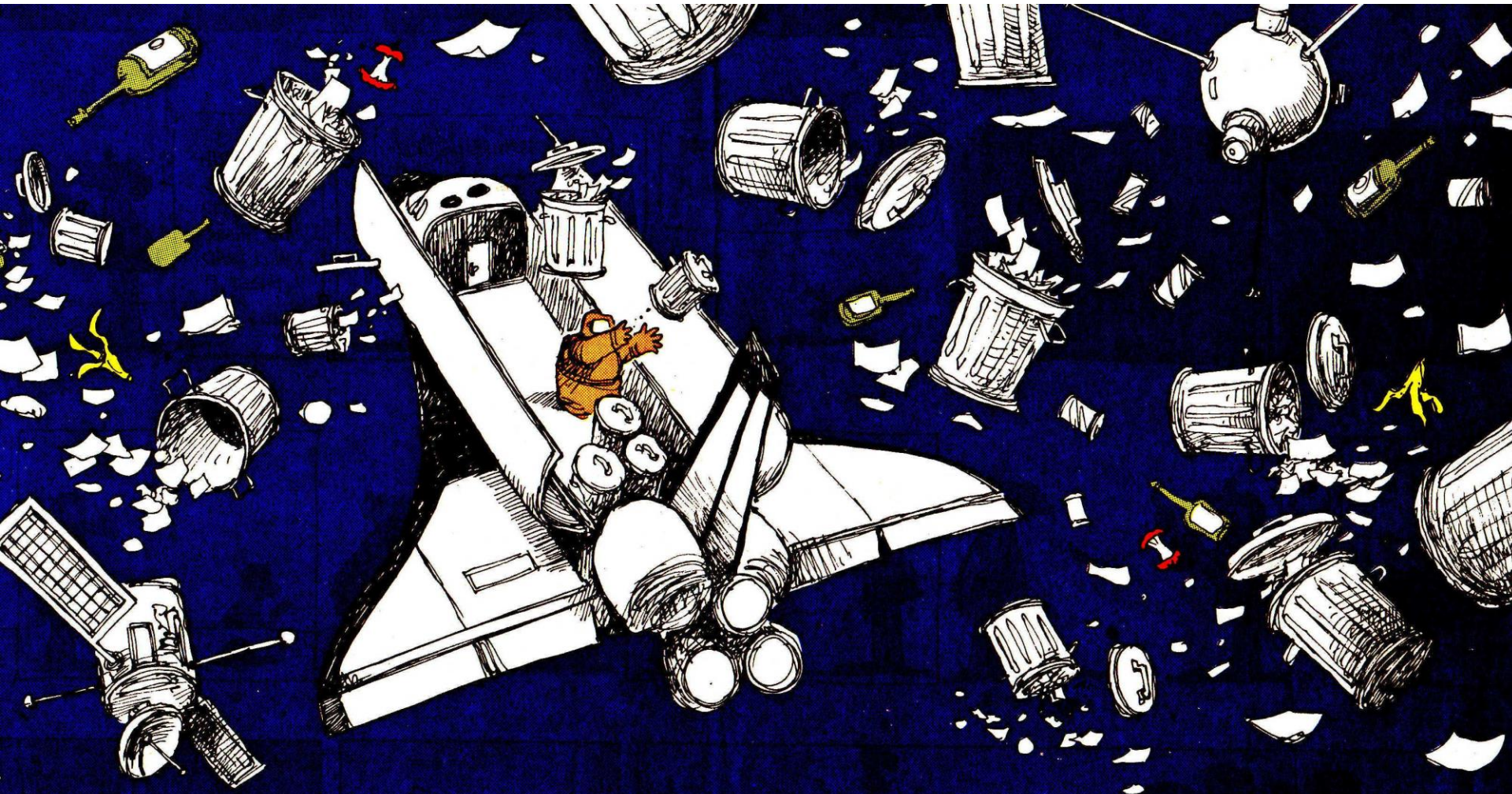
Challenges Remain

- **Adherence to national and international orbital debris mitigation guidelines is essential if the debris population is to be controlled**
- **Despite efforts to reduce accidental explosions of spacecraft and rocket bodies, such events continue to have dramatic effects in near-Earth space**
- **The deliberate testing of an anti-satellite weapon at high altitude by China in January 2007 created the worst orbital debris cloud in history**
 - The majority of the debris will remain in Earth orbit for decades to come
- **The accidental 2009 collision is only the harbinger – collisions are expected to become more common in the future**
 - Growing consensus that we may have to be more proactive in removing large debris





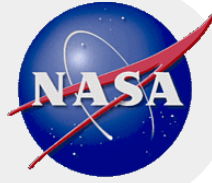
Questions?



HOW TO TELL WHEN MAN HAS OFFICIALLY CONQUERED SPACE

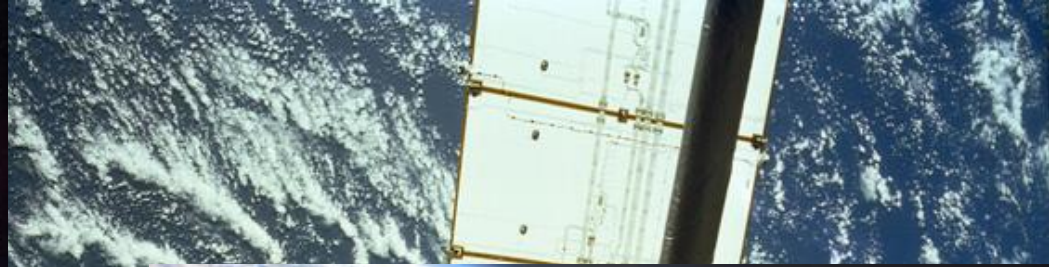
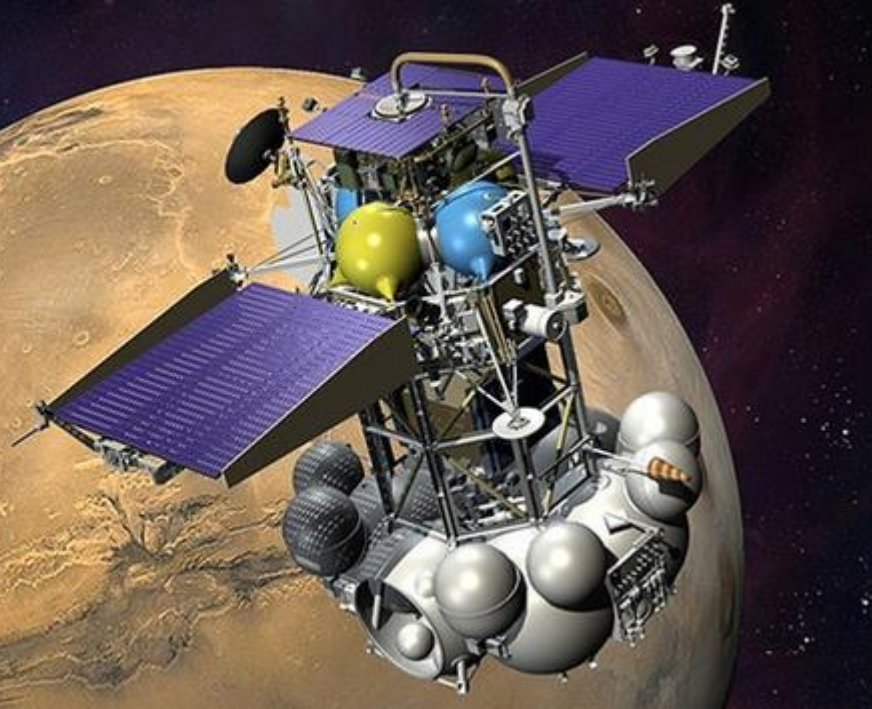


Backups



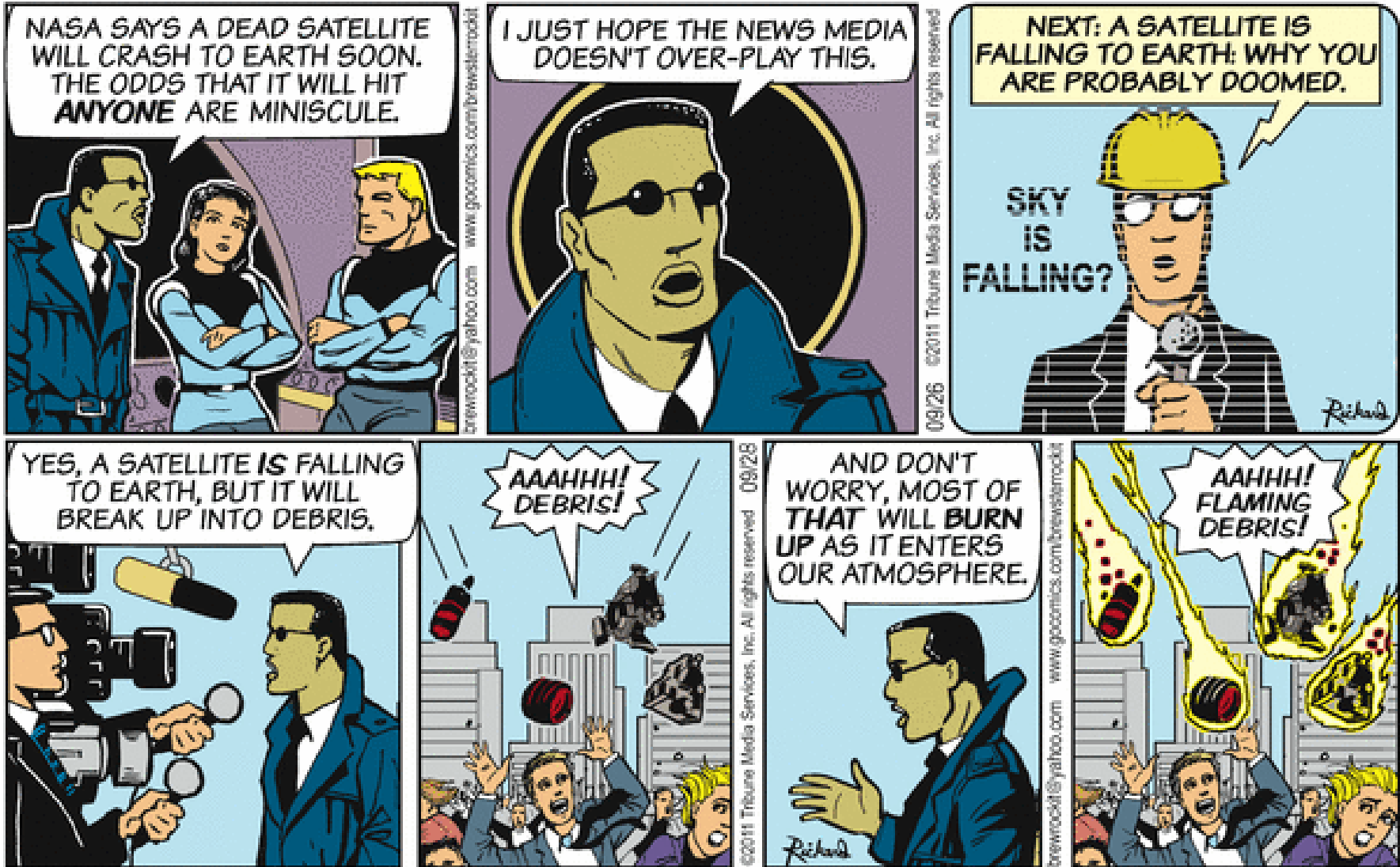
Recent Reentries

UARS, ROSAT, Phobos-Grunt, TRMM



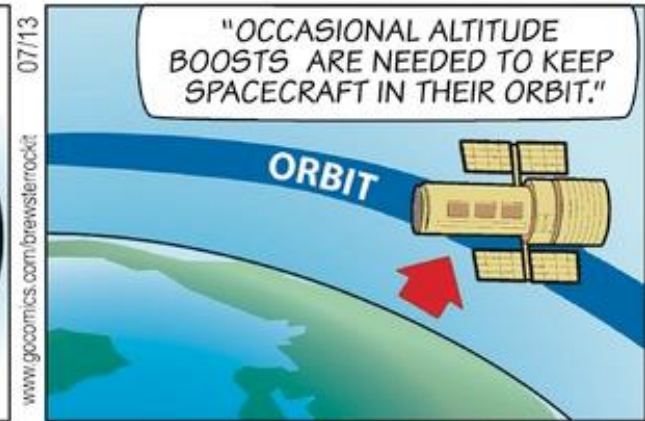
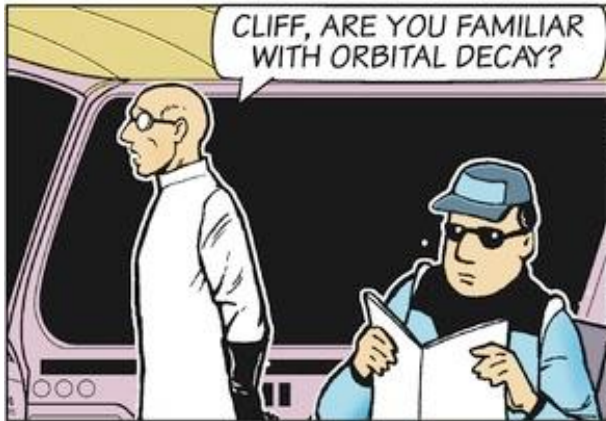


UARS Reentry in the Popular Imagination



BREWSTER ROCKIT: SPACE GUY!

BY TIM RICKARD





That Which Survives...



Texas, 1997



South Africa, 2000



Zimbabwe, 2013



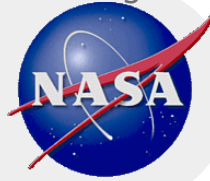
Guatemala, 2003



Argentina, 2004



Saudi Arabia, 2001



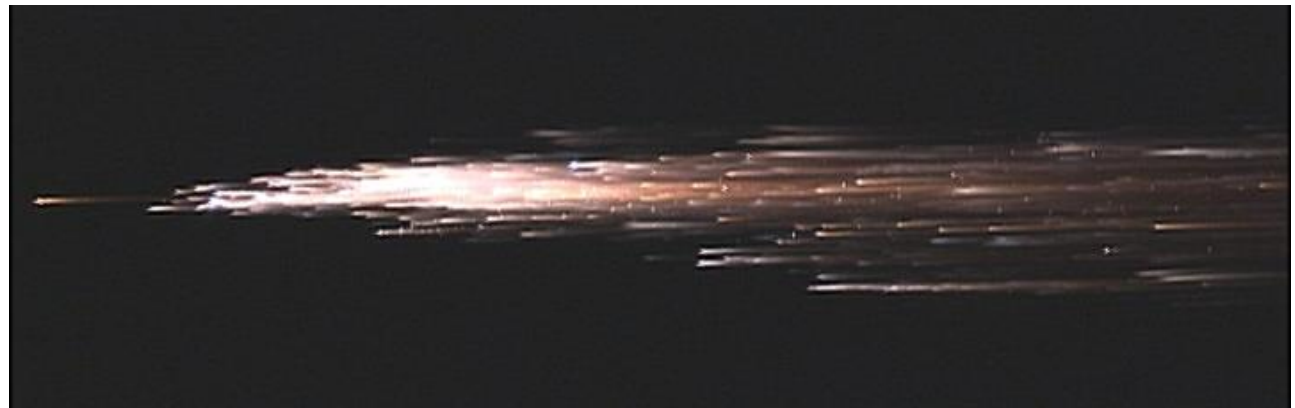
Reentry of the Jules Verne ATV

- **NASA and ESA conducted a joint observation campaign of the reentry of the Jules Verne ATV on 29 September 2008.**
 - Two aircraft collected a wide variety of data from vantage points over the Pacific Ocean near the reentry path of the Jules Verne.



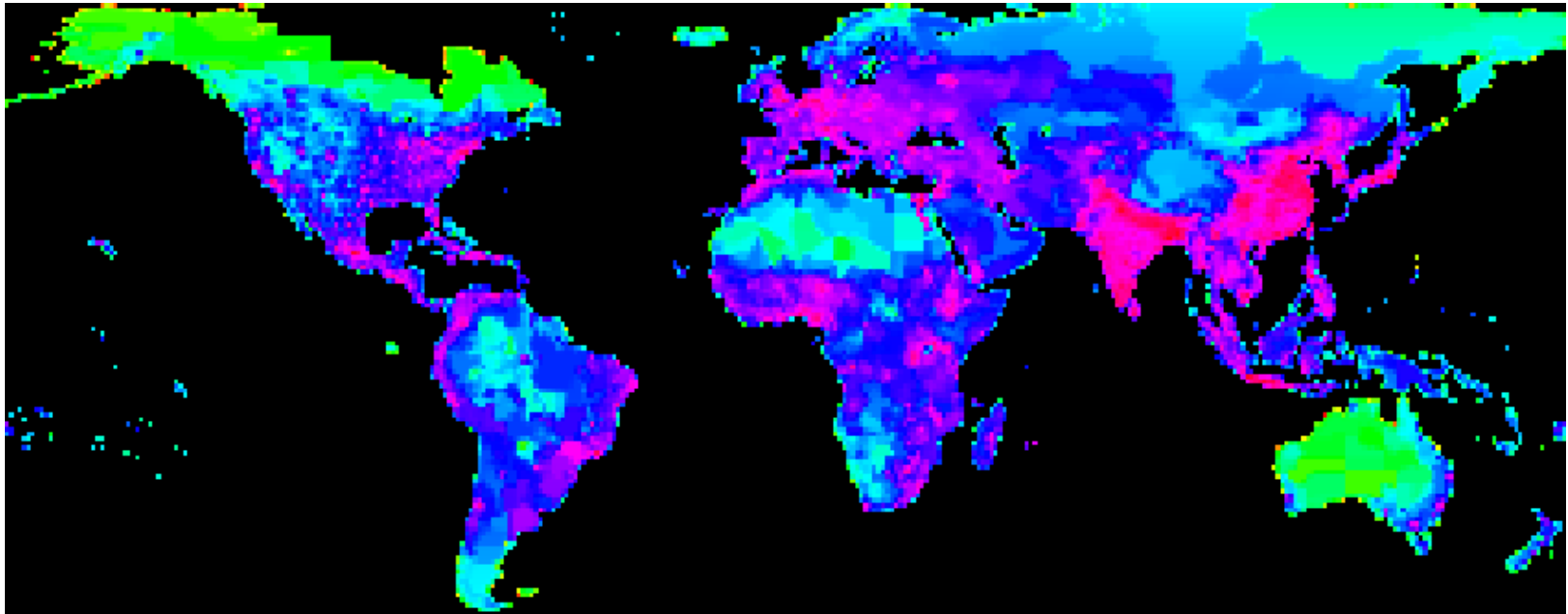
**Jules Verne undocking on
5 September 2008**

**Reentry over
Pacific Ocean**

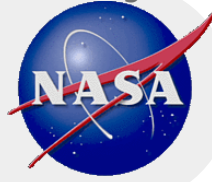




Population Distribution on the Earth

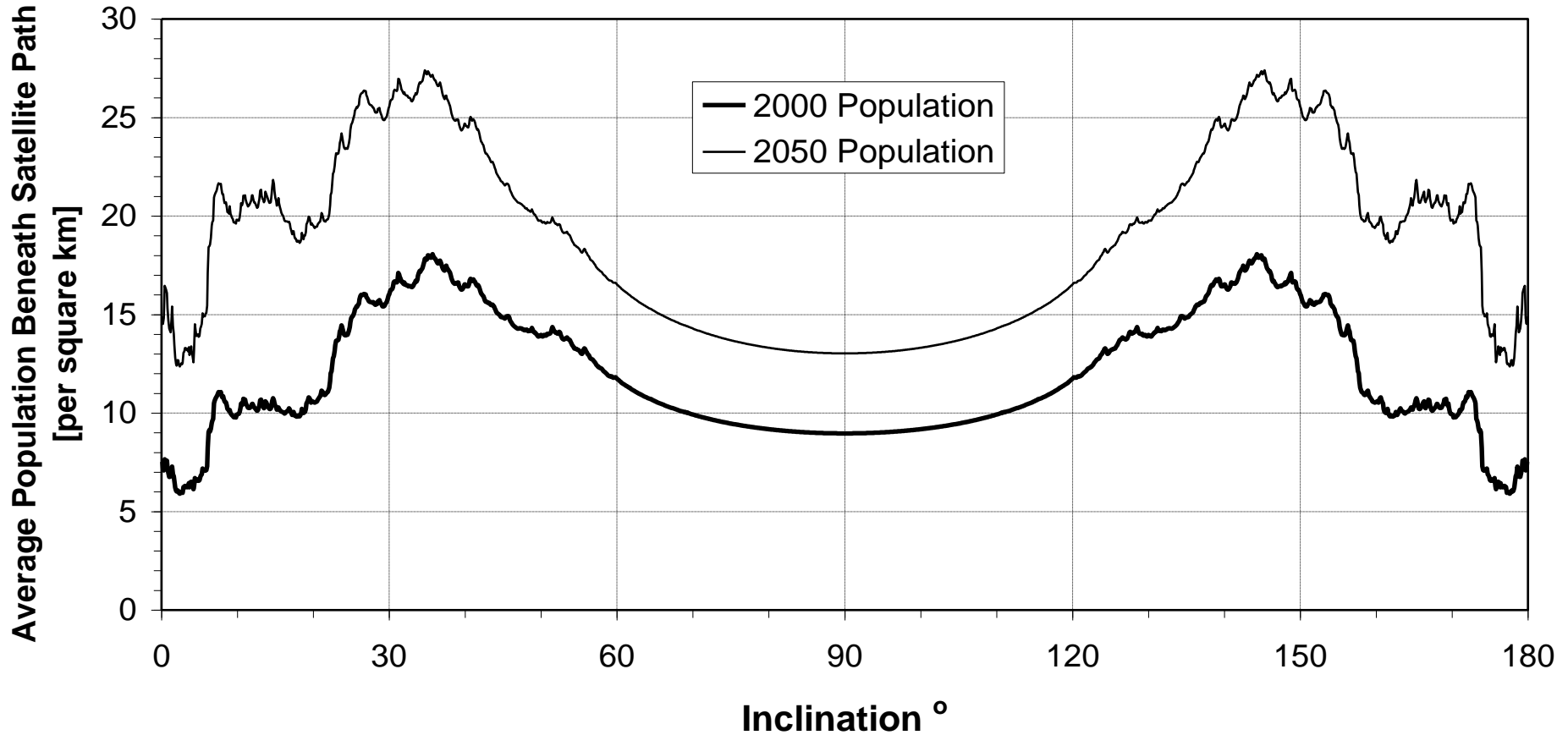


- Gridded Population of the World, version 3 (GPWv3)
- Socioeconomic Data and Applications Center (SEDAC) at Columbia University
- 2.5×2.5 arc minute cells = 4.6 km×4.6 km cells at the Equator
- Reference years 1990-2015 in 5-year intervals



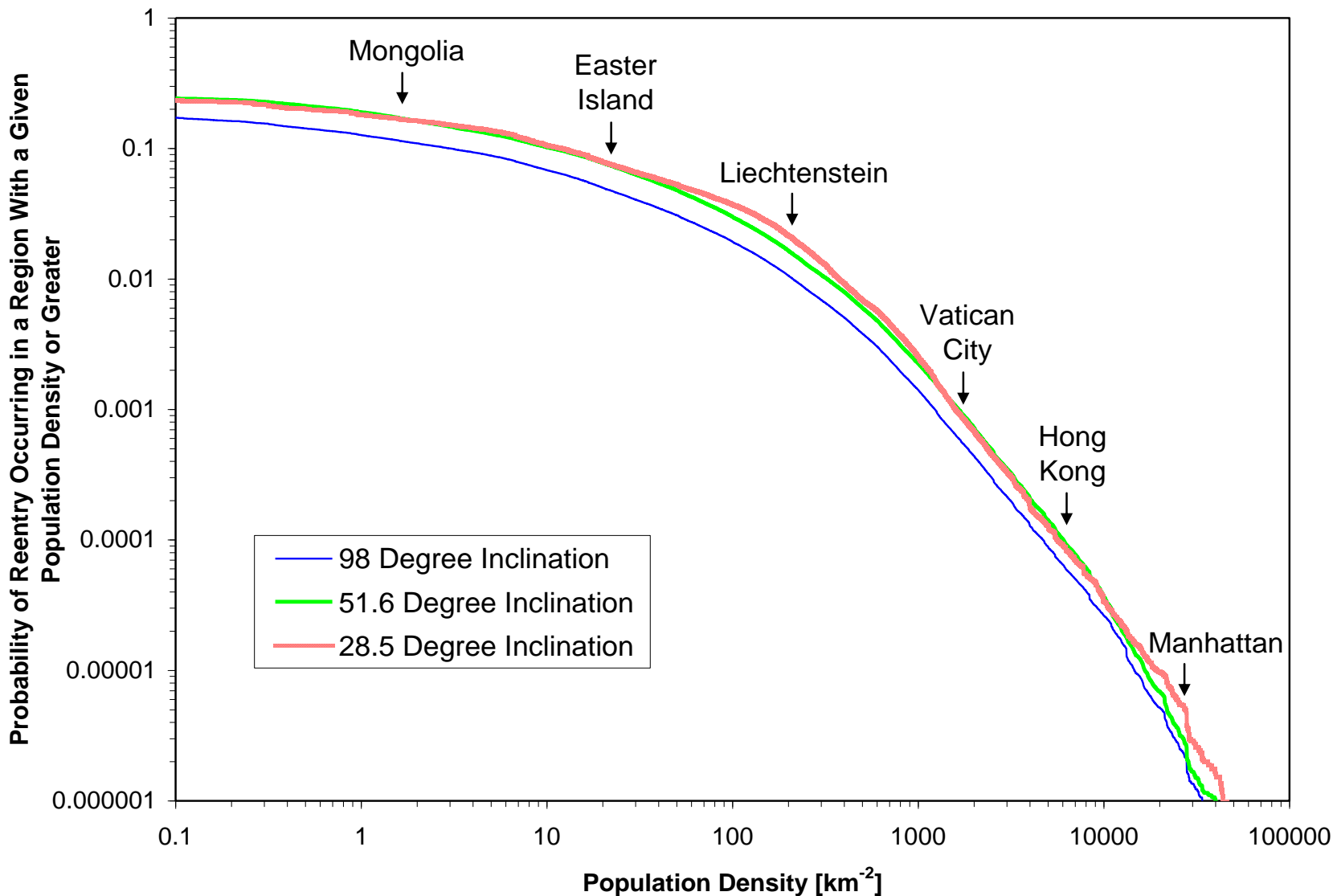
Average Density of People Below Satellite Path

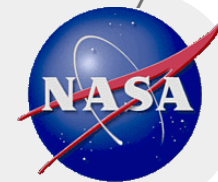
Inclination-Dependent Latitude-Averaged Population Density





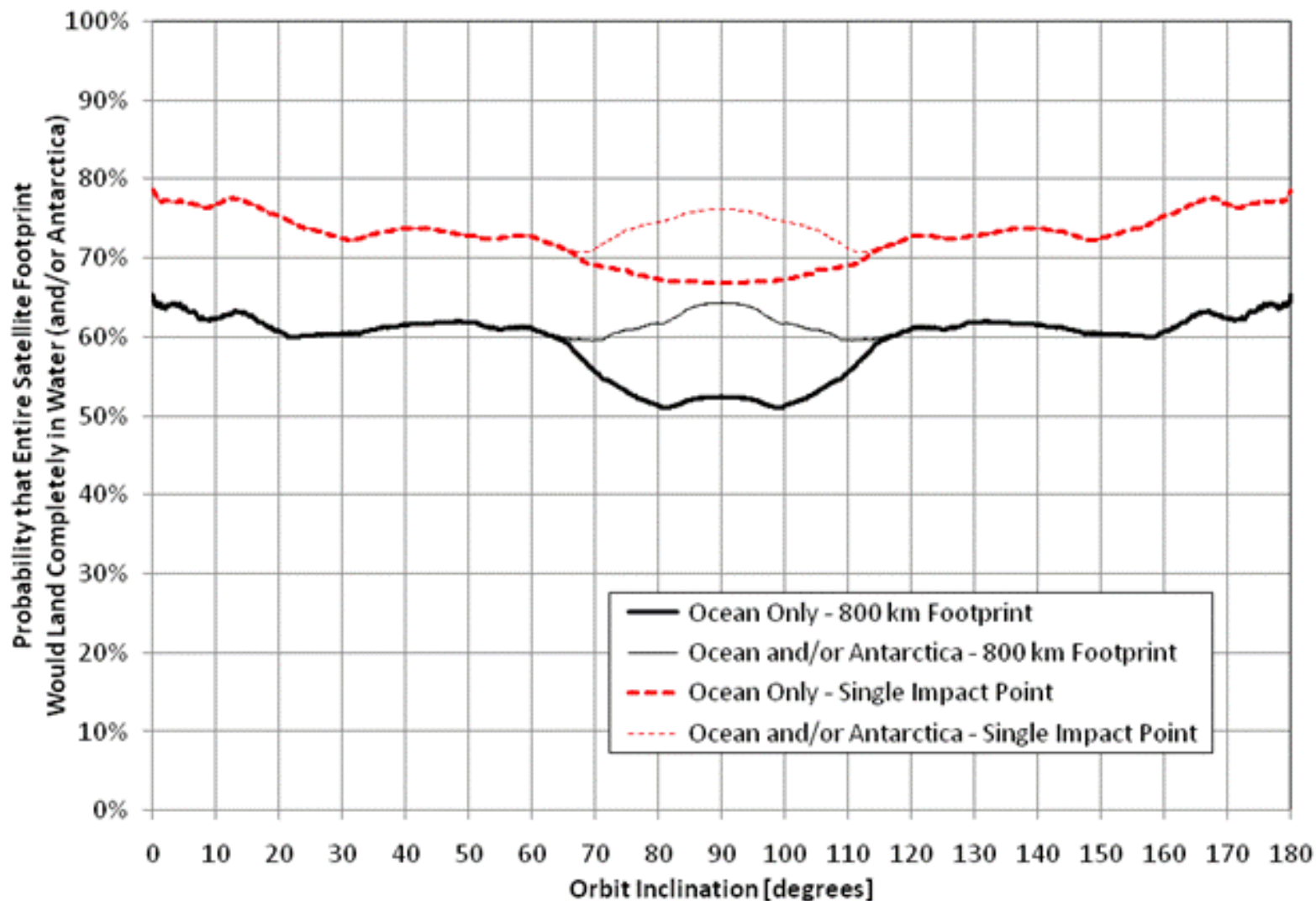
Probability of Falling in Populated Areas

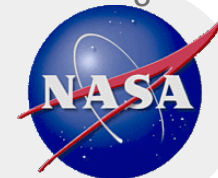




Probability of Ocean Reentry

Probabilities of Satellite Reentry Avoiding Land





Brewster Rokit on Reentry Risks

