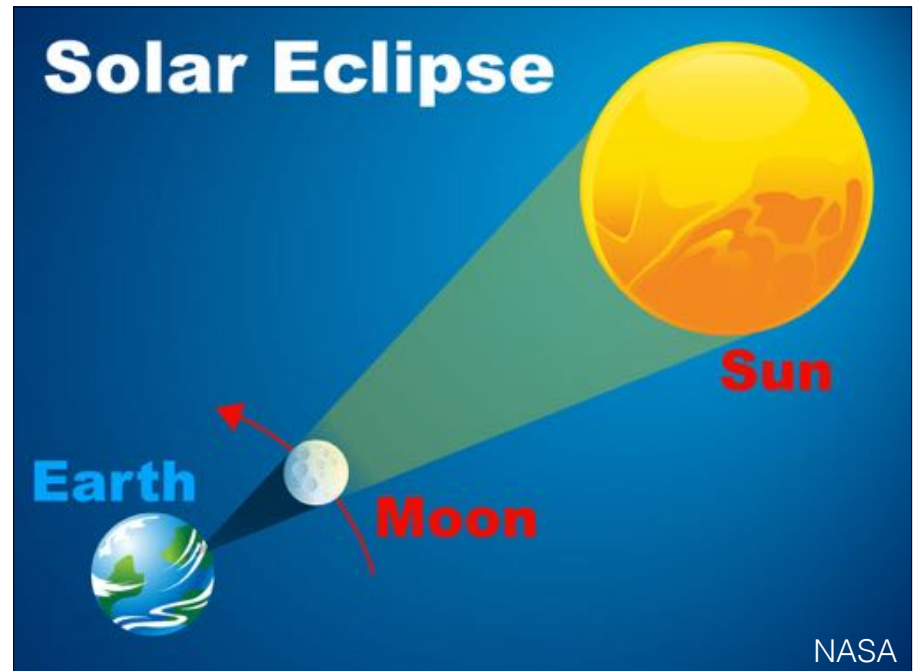
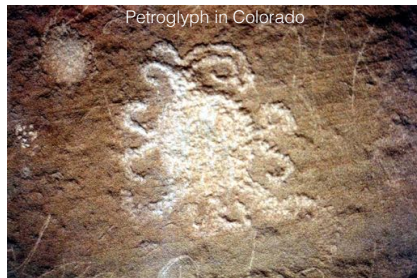
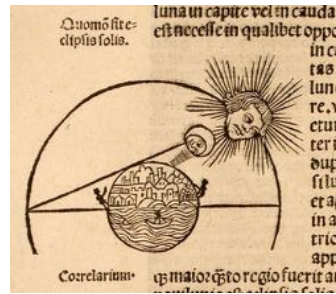


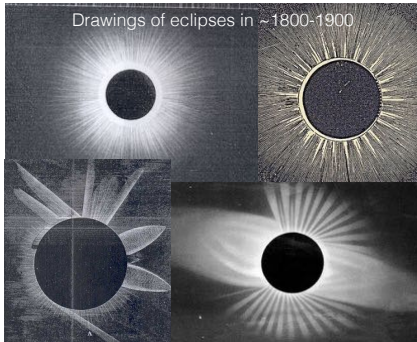




Petroglyph in Colorado



~1000 AD, De temporibus anni, Aelfric



Drawings of eclipses in ~1800-1900



2008 eclipse, ApJ, 10.1088/0004-637X/719/2/1362

Activity

Get in a group (~2-4 people) with your nearest neighbors

Record your answers.

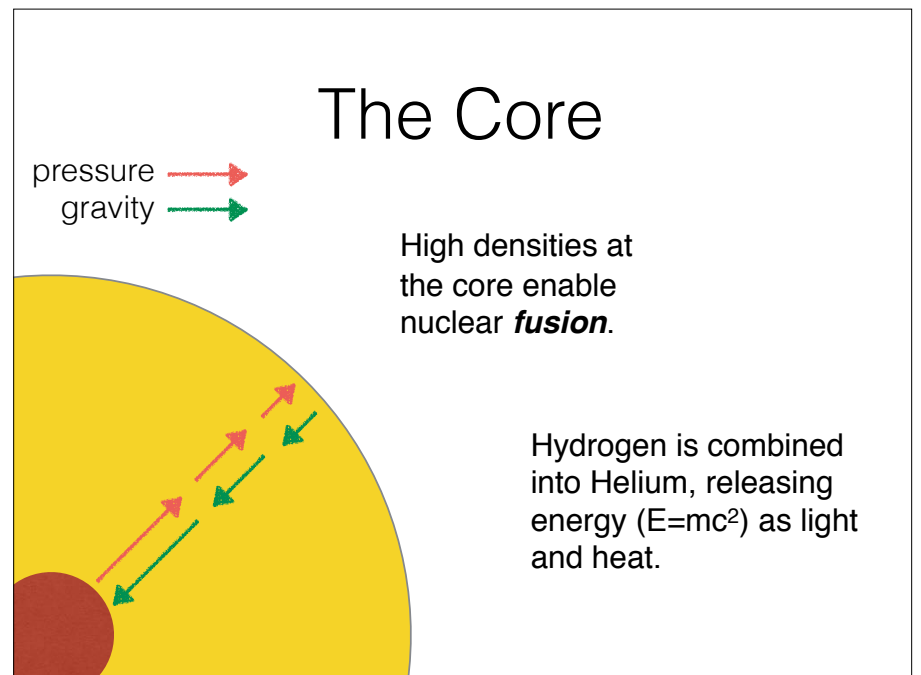
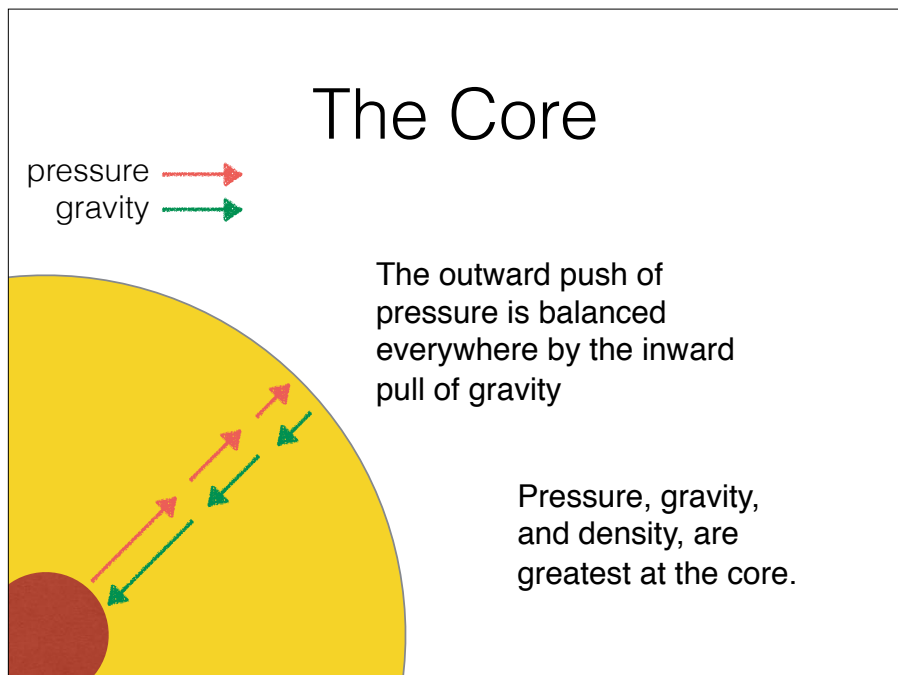
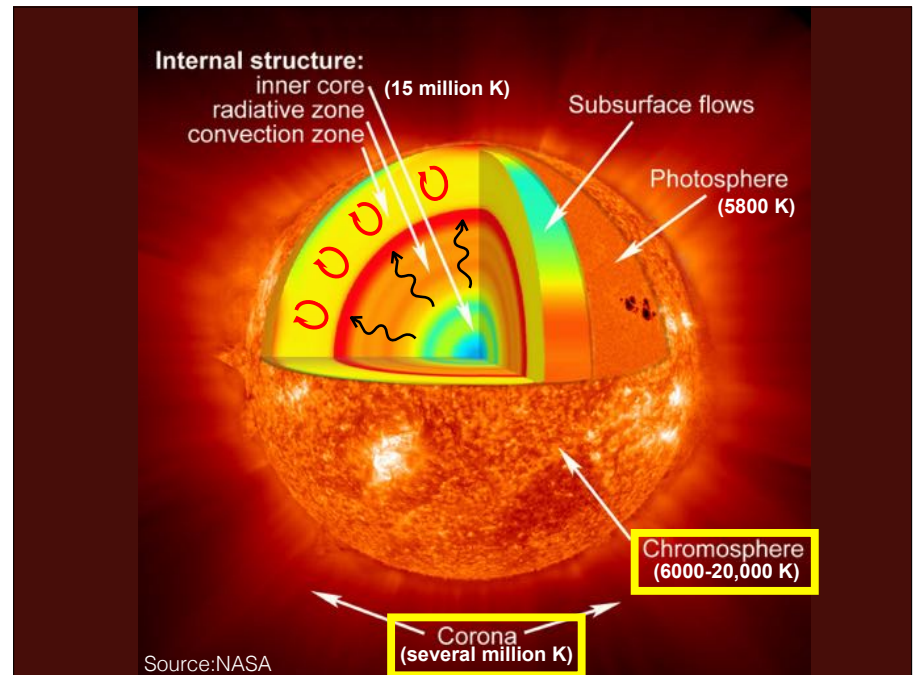
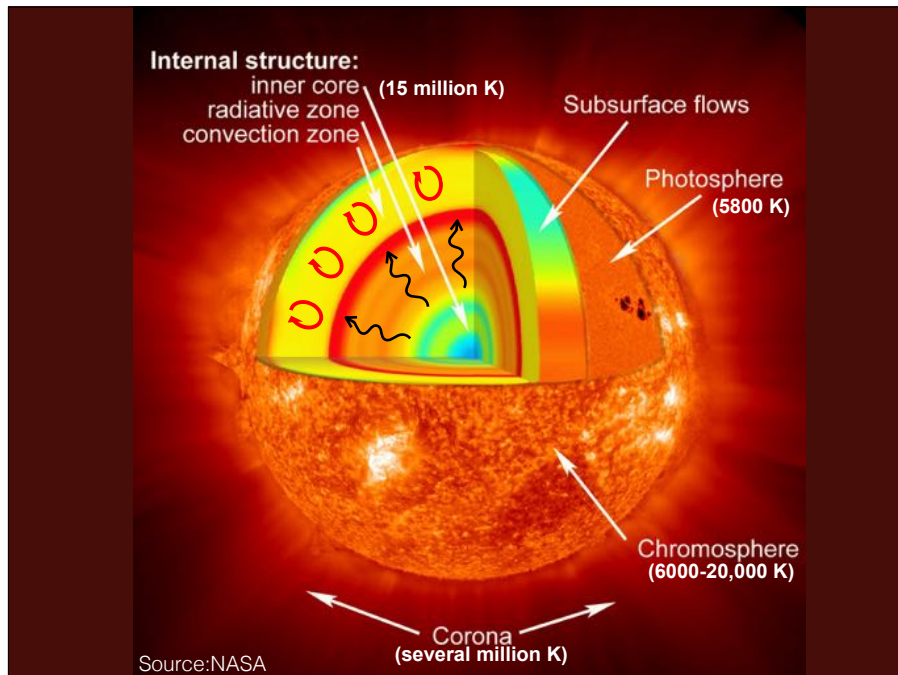
Order from coolest to hottest

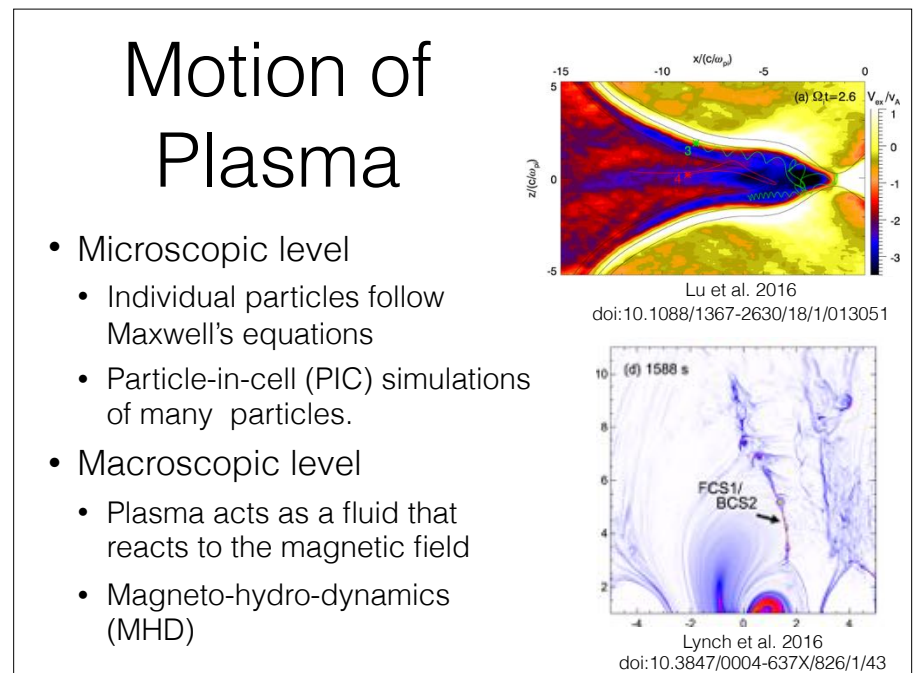
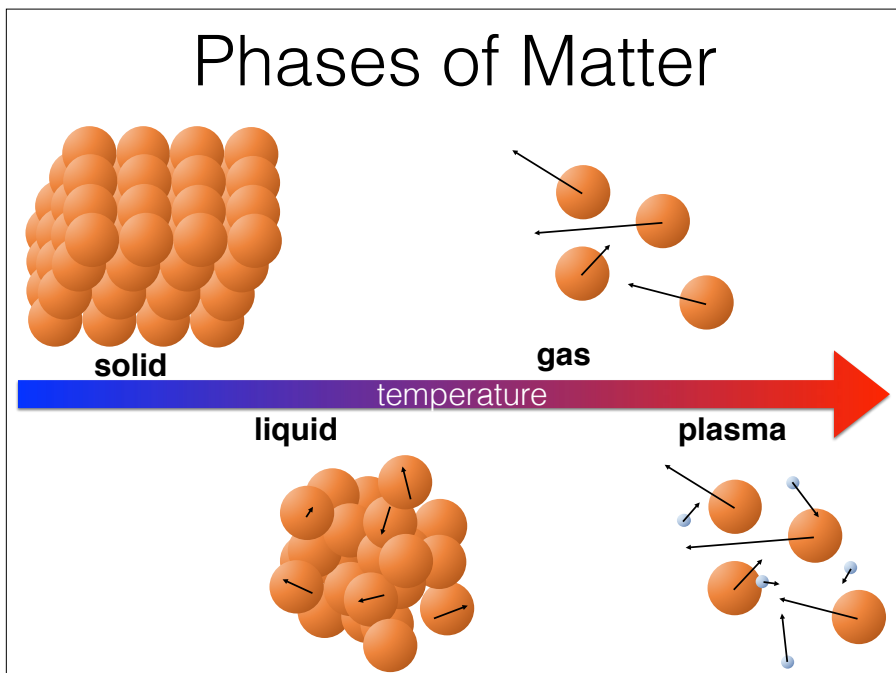
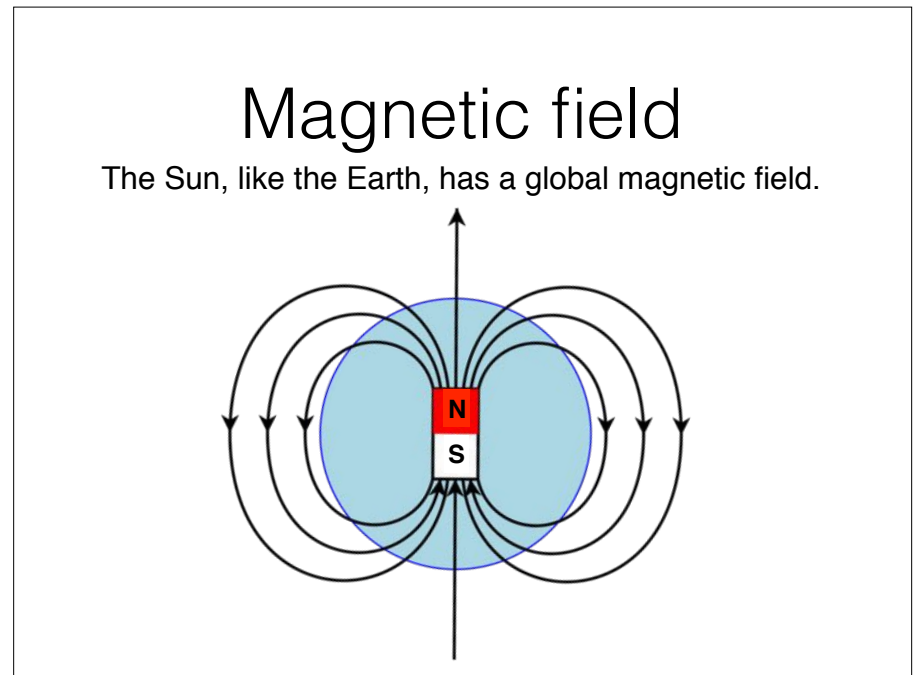
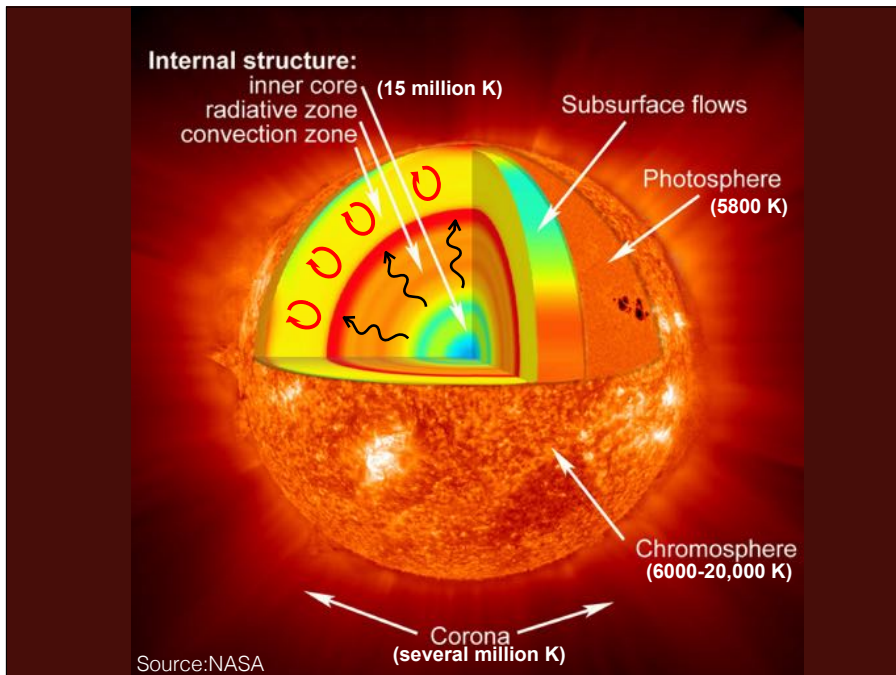
A Sunspot	B The Sun's Corona ("atmosphere")	C Earth's Core
D Meteor	E The Sun's Core	F Surface of the Sun
G Volcanic lava	H Comet	I Lightning

Answer

H Comet -450°F to 200°F	C Earth's core 6200°K
G Lava 1450°F to 2000°F	I Lightning 30,000°K
A Sunspot 6300°F	B Sun's corona 5 million °K
D Meteor 10,000°F or 5800°K	E Sun's core 15 million °K
F Sun's surface 6000°K	

Source: <http://solar-center.stanford.edu/activities/HowBig/How-Big-Far-Hot-Old.pdf>





Frozen-in-flux

- Induction equation (Ampere's law, Faraday's law, Ohm's law):

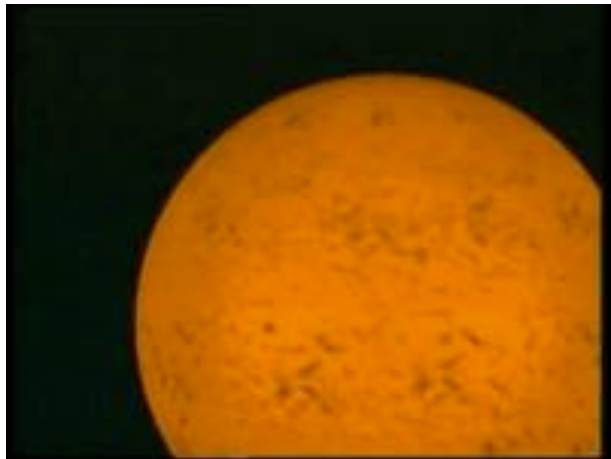
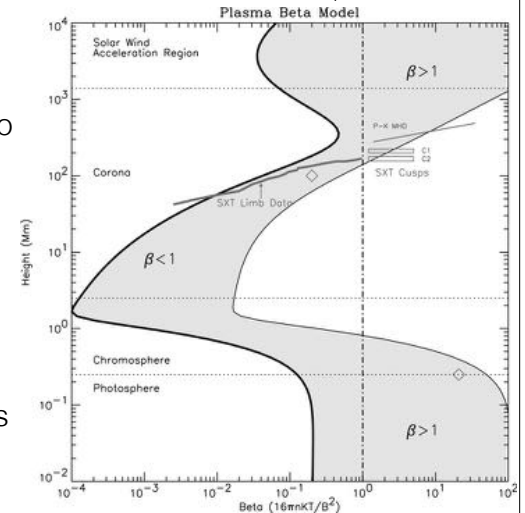
$$\frac{\partial \mathbf{B}}{\partial t} = \underbrace{\nabla \times (\mathbf{v} \times \mathbf{B})}_{I.1} + \underbrace{\eta \nabla^2 \mathbf{B}}_{I.2}$$

- Almost all astrophysical plasmas have very small *magnetic diffusivity*, η . (hotter plasmas have lower diffusivity)
- $I.2 \gg I.1$
- The fluid motion is tied to or 'frozen into' the magnetic field. (figure 2.17 on p 89 of your book)

Plasma- β

G.A. Gary 2001
doi:10.1023/A:1012722021820

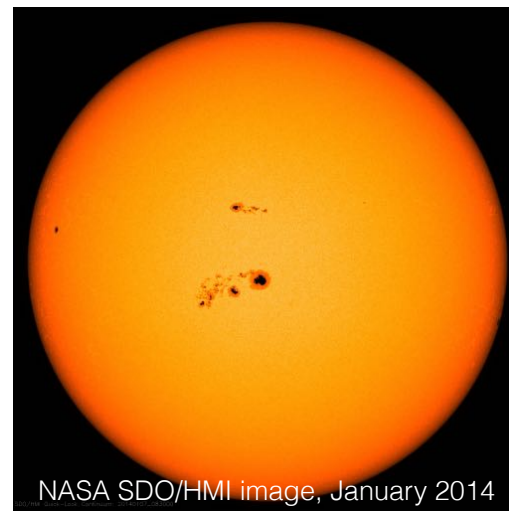
- $\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B})$
- Plasma- β is the ratio of plasma pressure to magnetic pressure.
- In the photosphere the plasma moves the magnetic field.
- In the corona the magnetic field moves the plasma.



Solar Dynamo

Spinning plasma drags the magnetic field.

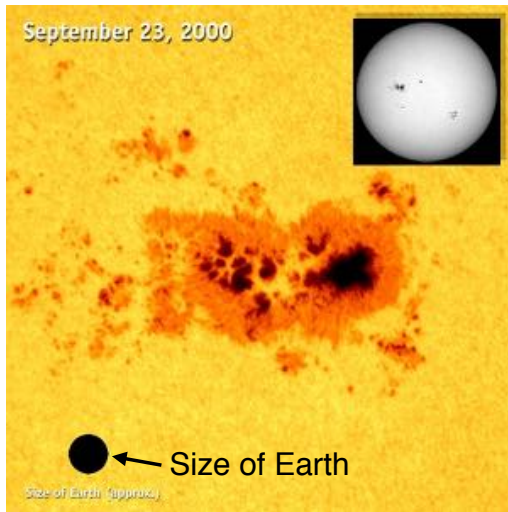
Source:SOHO (ESA/NASA)



Sunspots

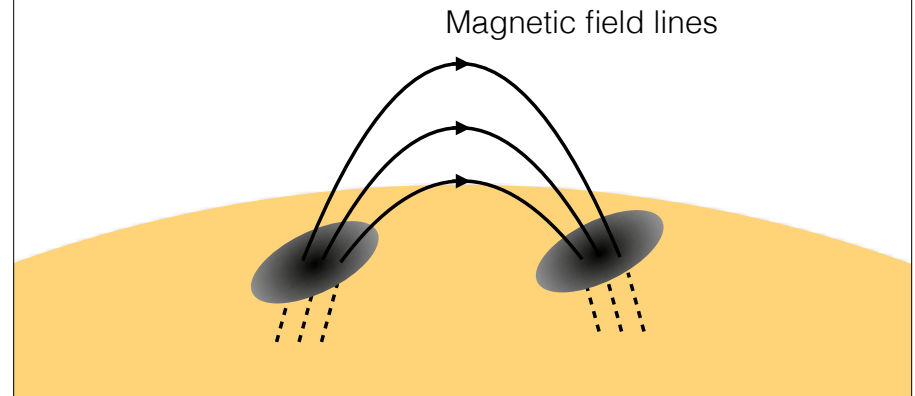
Sunspots form where concentrated magnetic field emerges through the photosphere.

source: SOHO (NASA/ESA)



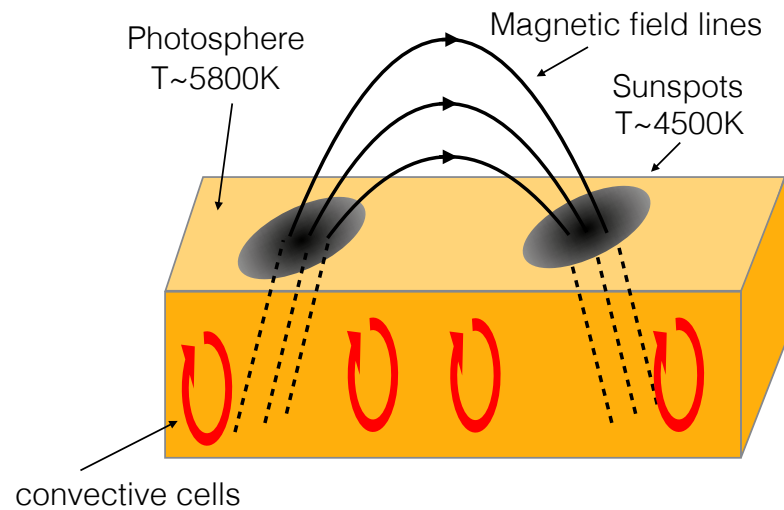
Sunspots

Smaller sunspots are about the size of the Earth.



Sunspots

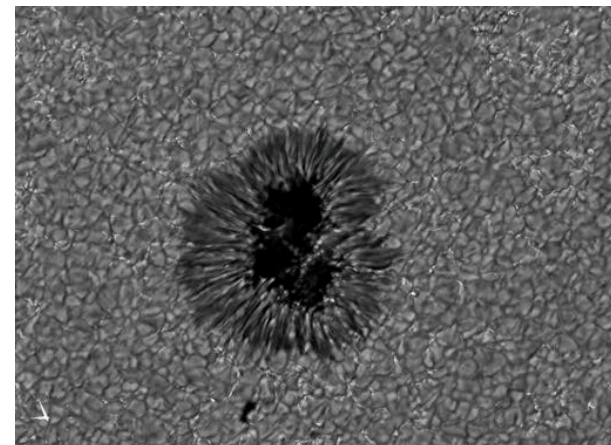
Strong magnetic field threads through sunspots.



Sunspots

Strong vertical field inhibits convection, making sunspots cooler than the surrounding photosphere.

Sunspot

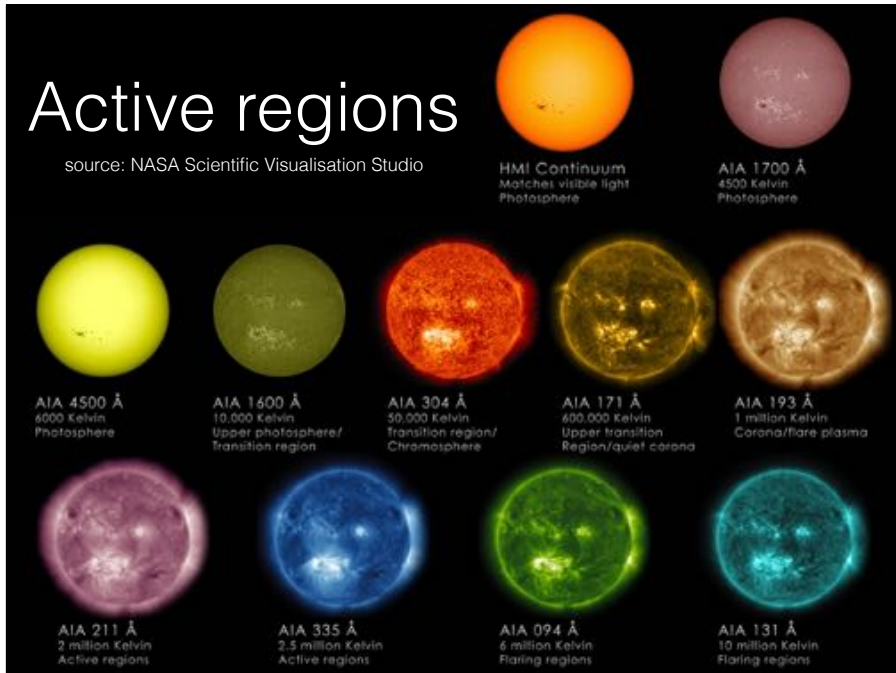


<http://www.staff.science.uu.nl/~rutte101/dot/albums/movies/album.html>

Dutch Open Telescope - 1 April 2001

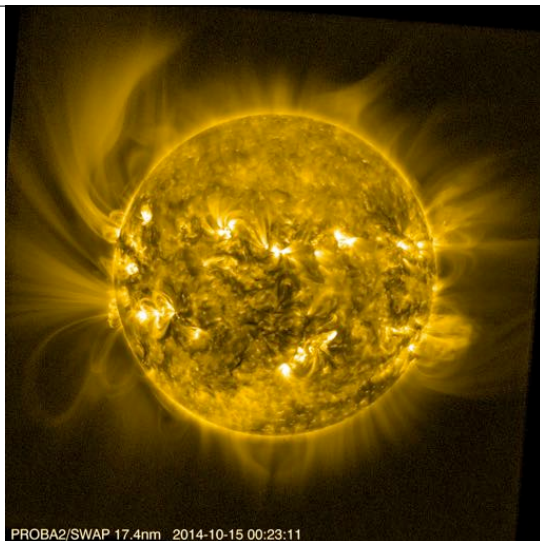
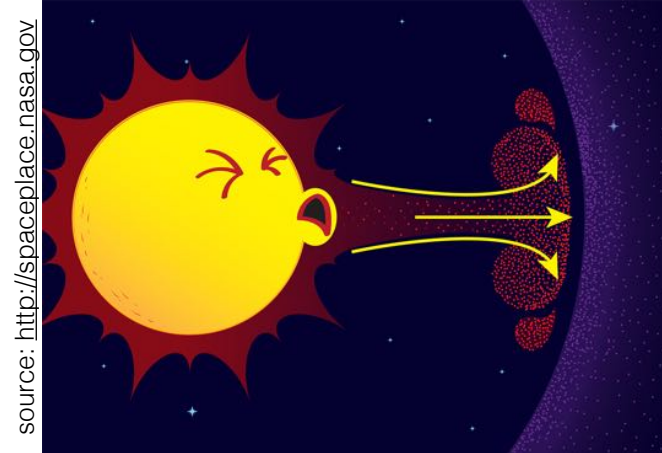
Active regions

source: NASA Scientific Visualisation Studio



Heliosphere

The bubble-like volume surrounding solar system caused by the *solar wind*. Outside the heliosphere is *interstellar space*.

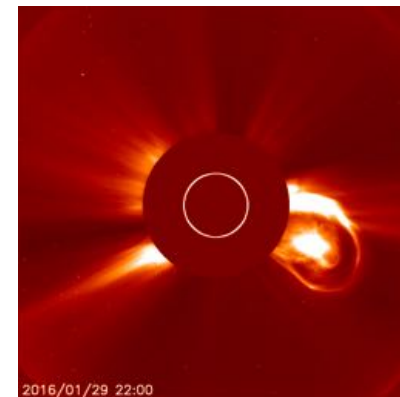


The dynamic corona

PROBA2/SWAP movie of 3 solar rotations

Eruptions

Major disturbances in the heliosphere are caused by massive explosions in the Sun's atmosphere: **coronal mass ejections**.



NASA LASCO C2 movie

Quiz

How fast are these eruptions?

		
5 m/s	5 km/s	500 km/s

Quiz

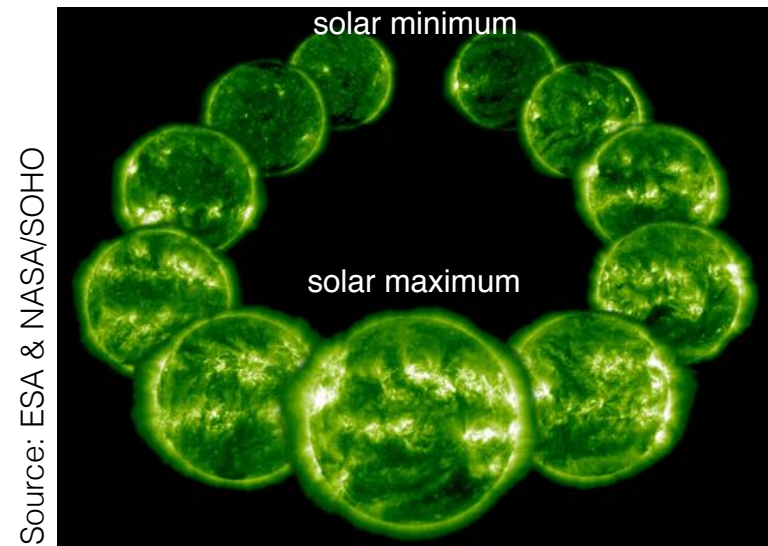
How massive are these eruptions?
(1 m³ of water = 1000 kg = 1 metric ton)

		
1 km ³ 10 ⁹ m ³	1000 km ³ 10 ¹² m ³	100,000 km ³ 10 ¹⁵ m ³

Eruption statistics

- How massive? 1 cubic km³ of water
- How fast? About 500 km/s (1.1 million mph)
- How much energy?

Activity cycle



CLASP launch 3 September 2015, White Sands Missile Range
MSFC Sounding Rocket



LASP Sounding Rocket



Thank
you!

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National Aeronautics and
Space Administration



MARSHALL
SPACE FLIGHT CENTER