

Proposed LISA Telescope Design

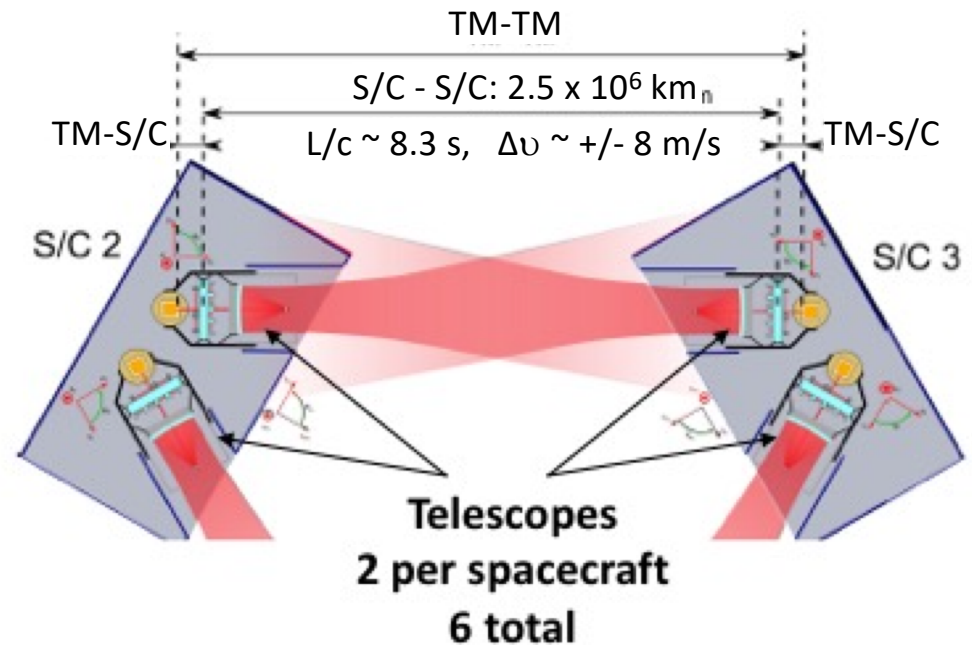
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Telescope Functional Description

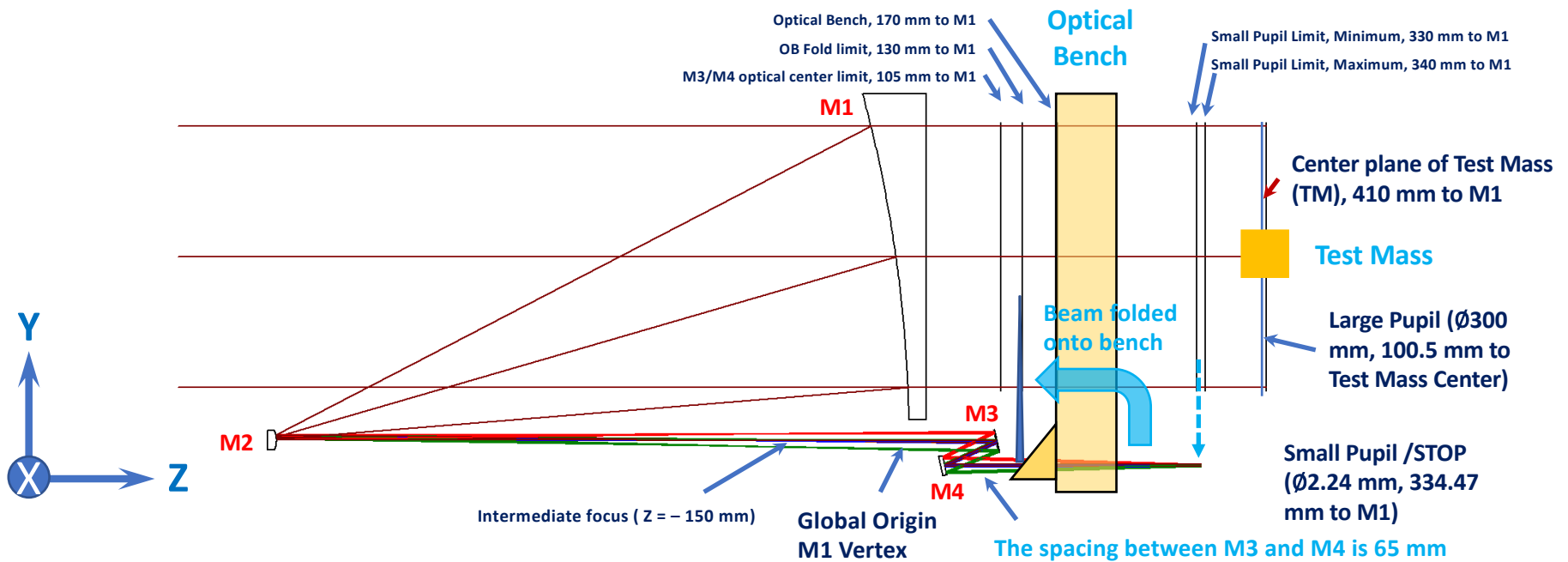
- Efficiently deliver power on-axis to the far spacecraft
- Simultaneous transmit and receive (TX/RX)
- Afocal beam expander
 - 300 mm dia. large beam
 - 2.24 mm dia. on bench
 - 134X magnification
- **Application is PRECISION LENGTH MEASUREMENT, not image formation**
 - Keep optical pathlength stable to ~ 1 pm/ $\sqrt{\text{Hz}}$ over the measurement BW
 - Minimize phase noise from coherent transmitter backscattered light
 - Minimize tilt to length (TTL) coupling



Key Challenges/design drivers

- Dimensional Stability
 - Low CTE material and thermal stability at the 1×10^{-5} K/VHz level, using baffles fixed to the spacecraft that act as a passive thermal filter
 - Spacecraft is stable at the 1×10^{-3} K/VHz level per LISA Pathfinder experience.
- Coherent backscattered light
 - Adopt unobstructed design
 - Low scatter coatings and dimensionally stable structure so scattered light phase is stable
 - Keep spacing between the telescope and optical bench stable
- Tilt-to-length (TTL) coupling
 - Careful design to minimize pathlength differences with field angle
 - Careful alignment of measurement axes (bench to bench and bench to GRS) such that spacecraft tilts do not couple to the length measurement (movable aperture stops help during integration)
 - Wavefront errors projected into the far-field will also couple TX angular jitter into an apparent length signal

725-V7a-M4-Flat Design Layout



Global Coordinate System

- Origin at M1 Vertex
- Z-axis: parallel to optical axis (object and image space)
- Y-axis: perpendicular to optical axis, within screen
- X-axis: perpendicular to optical axis, into screen

	Type	Conic	Radius (mm)	Clear Aperture (mm)	Edge Diameter (mm)
M1	Parabola	-1	-1475.246	345.3	365
M2	Hyperbola	-1.0928	-25.813	6.4	25.4
M3	XY Polynomial	0	Infinity	24.6	27.4
M4	Flat	0	Infinity	21.0	27.4

"freeform" optic

M1 Clear Aperture Analysis

Stop Size Calculation

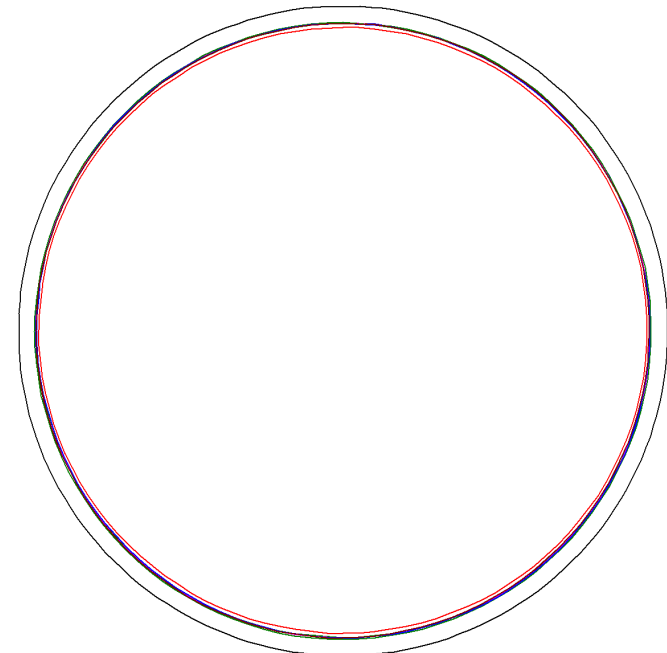
Nominal STOP Diameter (mm)	2.24	
Design Tolerance (mm)	+/- 0.0075	
Alignment/Interface Margin (mm)	+/- 0.170	+/- 0.140 Rx clip
Oversized STOP (Worst Case, mm)	2.5875	worst case expected

M1 Footprint Analysis

STOP SIZE (mm)	M1 Footprint (Y, X, mm) Science FOV (20 μ R)	M1 Footprint (Y, X, mm) Acquisition FOV (225 μ R)
2.2400	299.4, 299.4	300.2, 300.3
2.5875	345.9, 345.9	346.7 346.8

M1 Clear Aperture Analysis

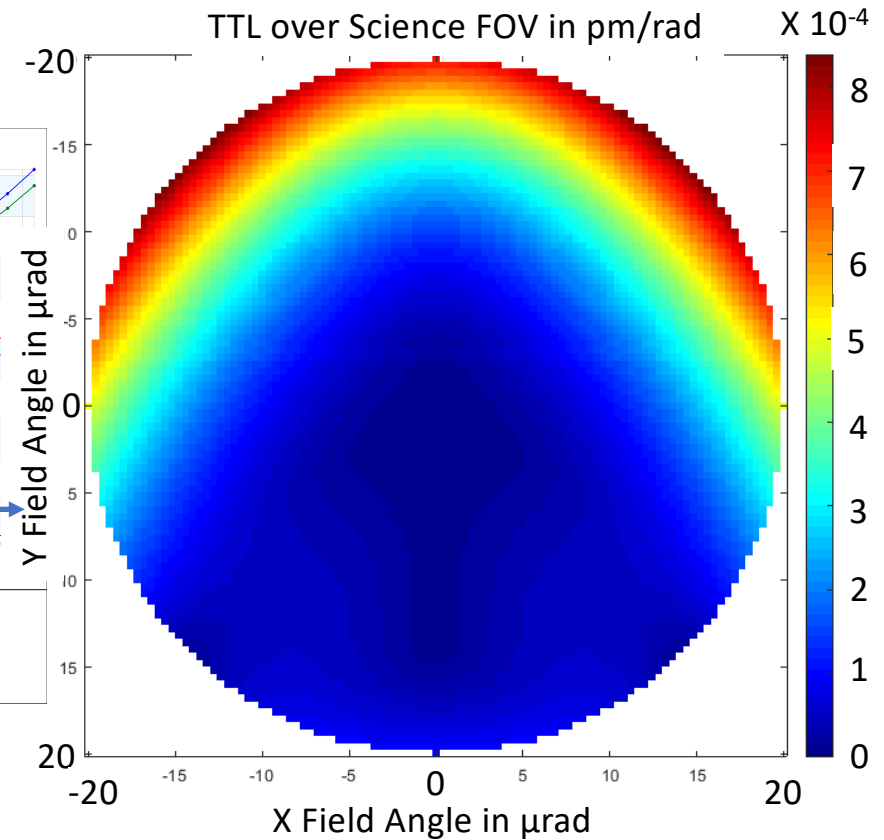
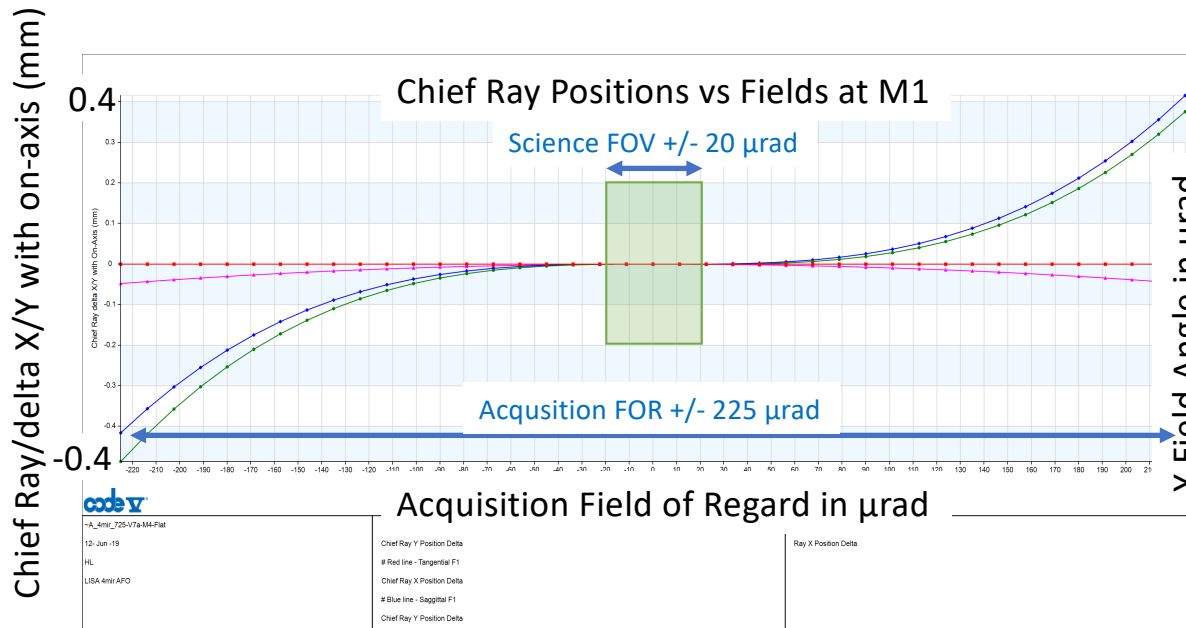
M1 Footprint (diameter, mm)	346.8
Clear Aperture Radius Margin (Alignment/Tolerance, mm)	3
Coating Radius Margin (included above, mm)	2
Mechanical Edge Radius Margin (mm)	4.1
M1 Physical Diameter (mm)	365



M1 Footprint Plot

- Physical Diameter 365 mm
- Footprint 346.8 mm
- Mirror center 210 mm to M1 vertex
- On-axis chief ray Y position 210 mm to M1 vertex

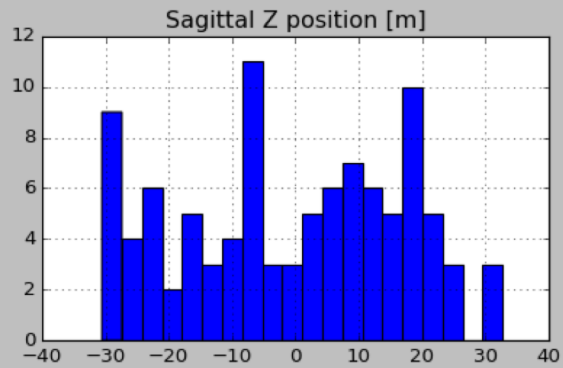
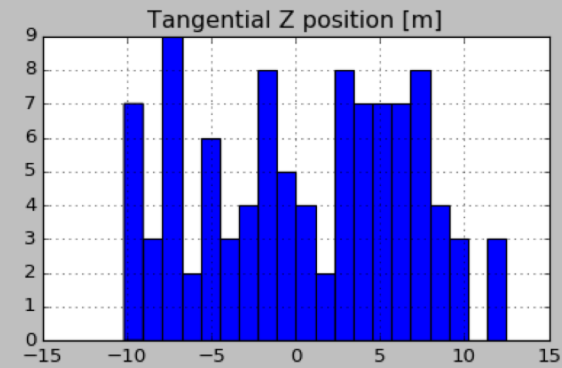
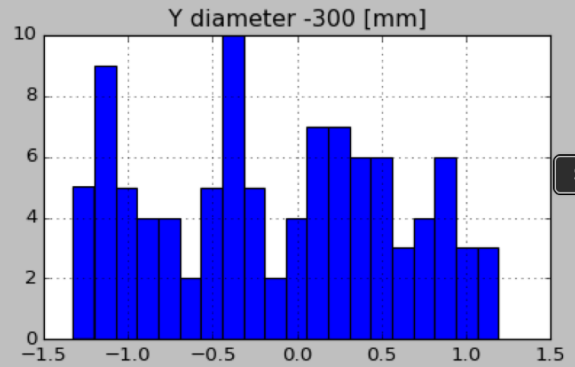
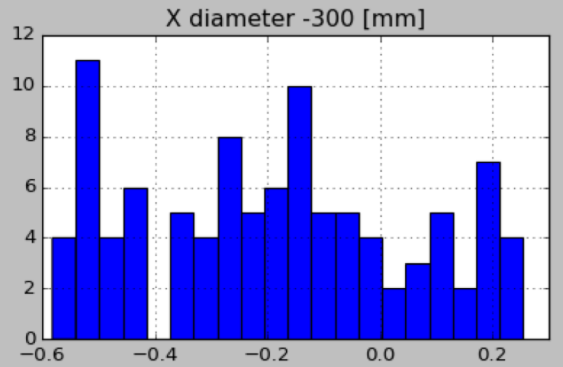
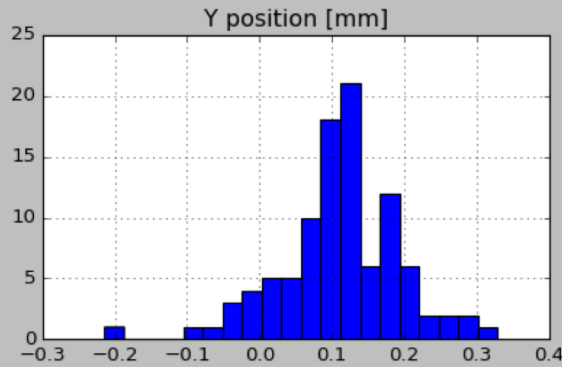
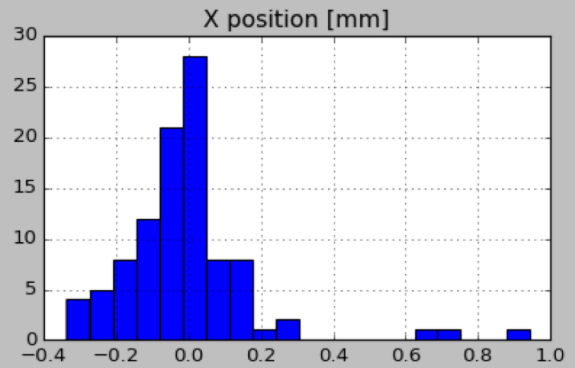
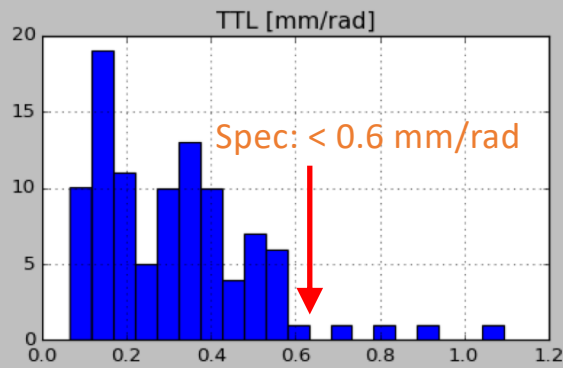
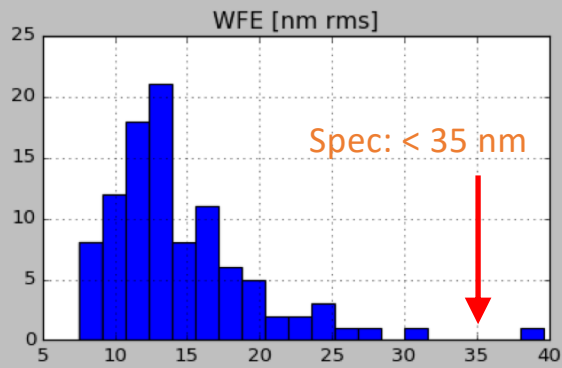
Pupil wander/dOPL vs. Field



“Intrinsic” optical path length variation nearly eliminated by design.
 Max over science field of view is 10^{-3} pm/rad = 1 fm/rad.

Alignment Tolerances

- Cassegrain tolerancing inputs are at the edge of metrologists' advice
 - few microns in ROC, 0.01% in conic, few microns in position
 - Driven by angles of incidence, need a longer telescope or a smaller primary to significantly alter
- M3/M4 tolerancing inputs are much more relaxed, angular alignment requirements are difficult but within the realm of "routine" (2 arcsec = 10 μ rad requirement)
- If these inputs are satisfied, then the image of the stop places well and the boresighting is tightly controlled (keeps TTL of integrated system to an acceptable level)
 - Magnification variation produces +/- 1.5 mm spread in beam sizes
 - WFE within requirements

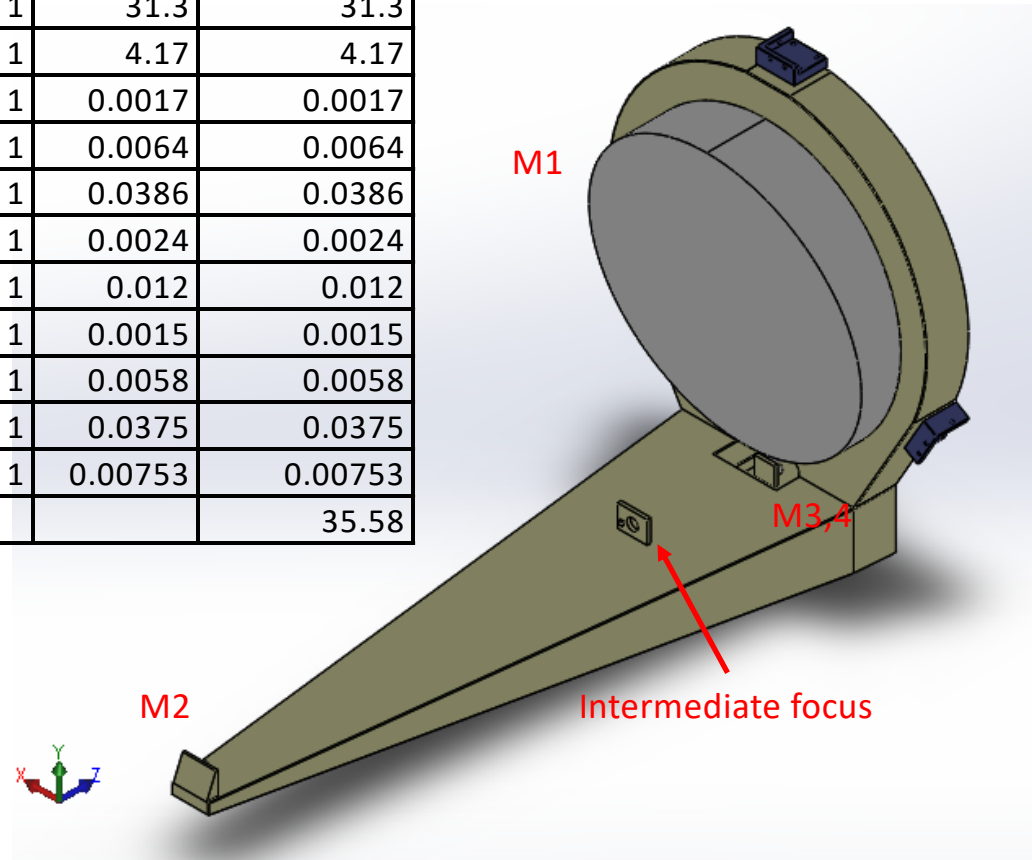


Monte Carlo Alignment
Tolerance Simulation
Results

Mechanical Design Overview

Component	Material	QTY	Mass (kg)	Total (kg)
One Arm Structure	Zerodur	1	31.3	31.3
M1	Zerodur	1	4.17	4.17
M2	Zerodur	1	0.0017	0.0017
M2 Adapter	Zerodur	1	0.0064	0.0064
M2 Mount	Zerodur	1	0.0386	0.0386
M3	Zerodur	1	0.0024	0.0024
M3 Mount	Zerodur	1	0.012	0.012
M4	Zerodur	1	0.0015	0.0015
M4 mount	Zerodur	1	0.0058	0.0058
M3,M4 Stage	Zerodur	1	0.0375	0.0375
Int Image Aperture	Zerodur	1	0.00753	0.00753
				35.58

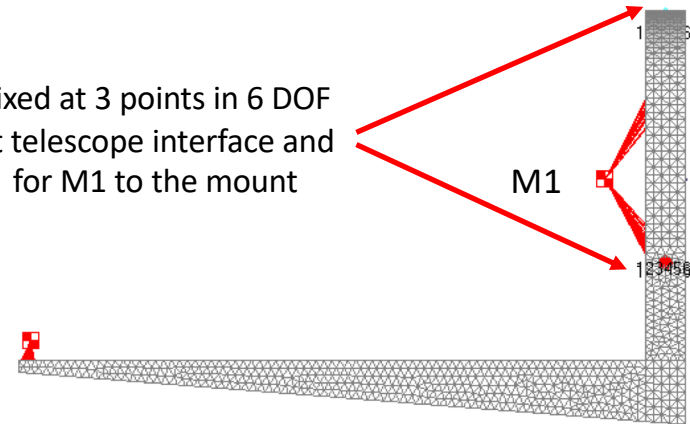
Further reduction in mass appears possible by light-weighting M1 mounting backplane (see next slide)



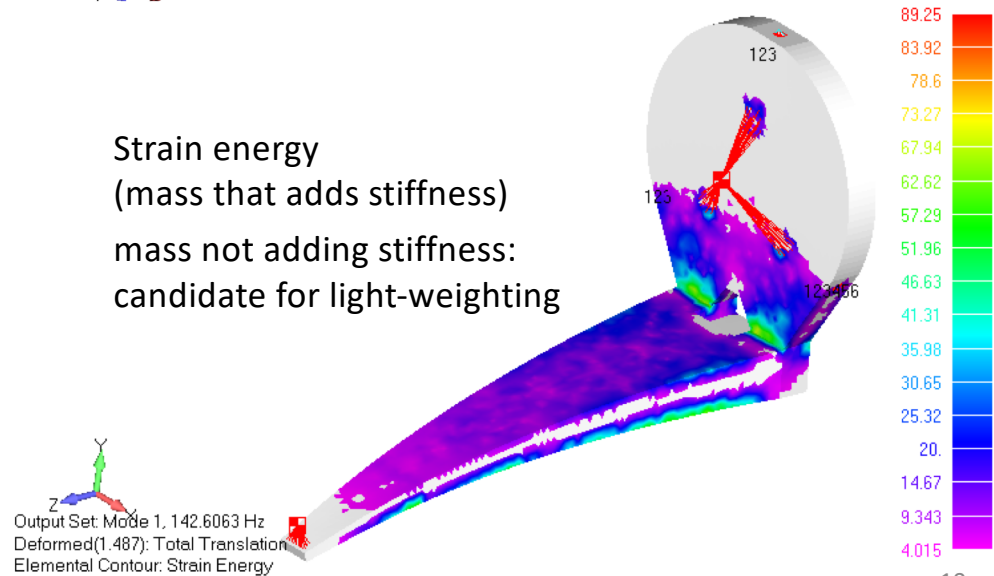
Preliminary FEM Analysis

- Single arm design
- Mass 30.3 Kg (Zerodur)
- First mode 142 Hz (diving board)
 - Goal is 100 Hz + 15% margin = 115 Hz
 - Second mode is 333 Hz
- M2 mass
 - Nominal ~.05Kg
 - 3x (account for alignment mechanism) ~.15Kg
- Output in global coordinate system
- M3/M4 mass/distortions not included
 - Fluid optical design
 - Sag should be negligible wrt to M2
 - Will update as optical design matures

Fixed at 3 points in 6 DOF at telescope interface and for M1 to the mount

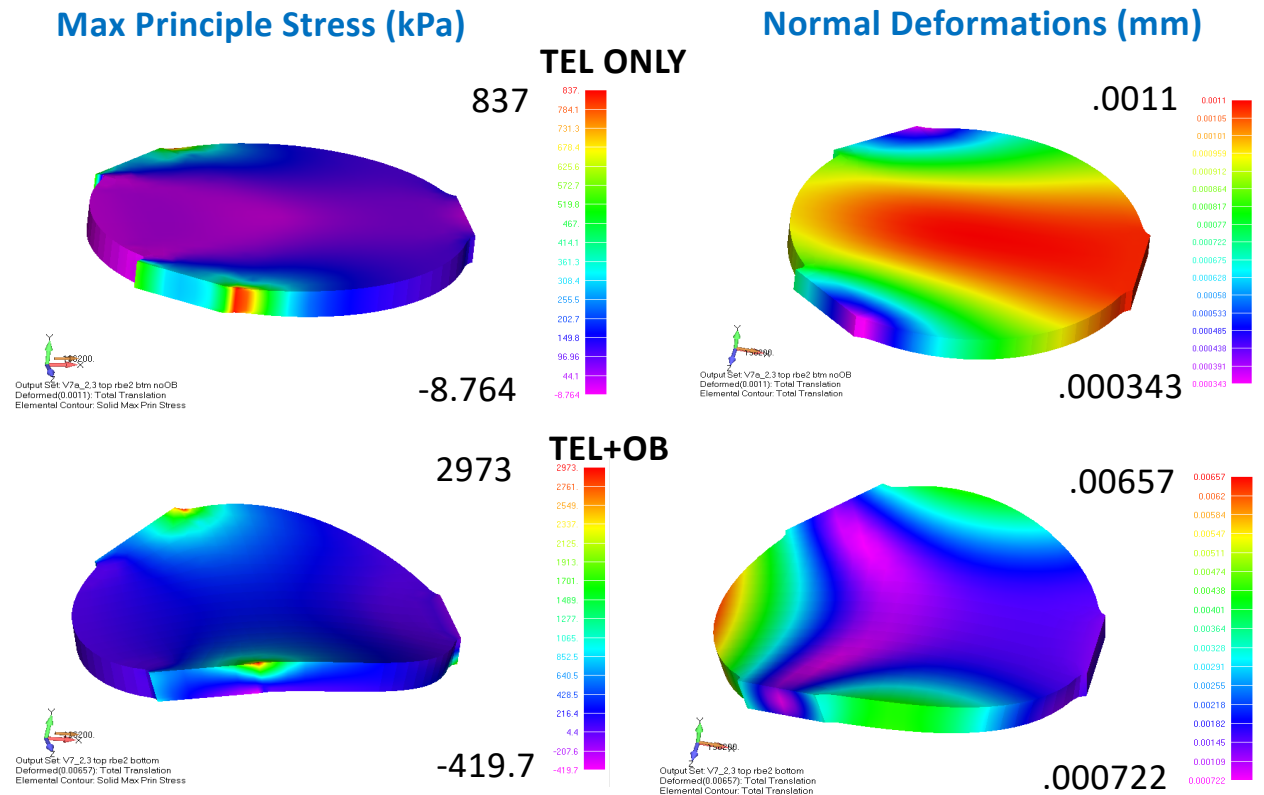


Strain energy (mass that adds stiffness)
mass not adding stiffness:
candidate for light-weighting



Preliminary FEM Analysis

- ~3x increase in telescope plate stress due to OB. Similar trends when looking at out of plane distortions across telescope face.
- Magnitude depends on (amongst other things) stiffness at supporting structure interface. The stiffer that interface the less loads will be reacted by the telescope.



Summary

- Proposed design meets key requirements
 - Dimensional Stability
 - Coherent backscattered light
 - “Intrinsic” tilt-to-length coupling
 - Can accommodate dynamic stop placement during alignment
- Tertiary mirror design is a freeform optic, but with small (< 2 microns RMS) deviations from a sphere
- Alignment requirements are tight, but possible
- Mechanical design with all-glass construction meets stiffness requirements
- Tolerance and mechanical analysis is ongoing
- Next step is to procure a structural thermal model (STM) and 2 engineering development units (EDU) for testing