



Dynamic Analyses of the Proposed Habitable Exoplanet Astrophysics Facility

SPIE August 2019

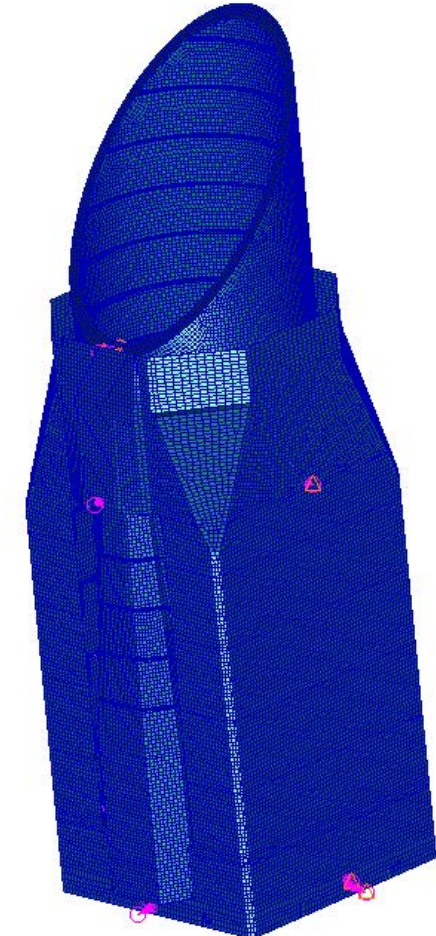
J. Brent Knight

NASA/MSFC/ES63

HabEx

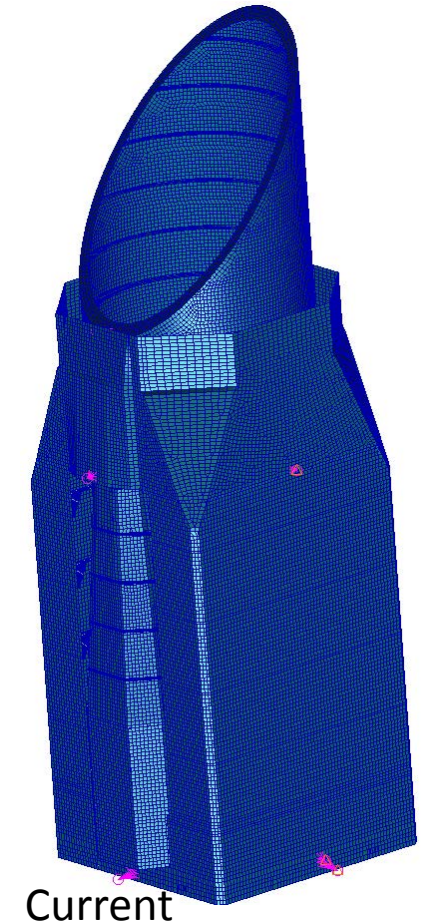
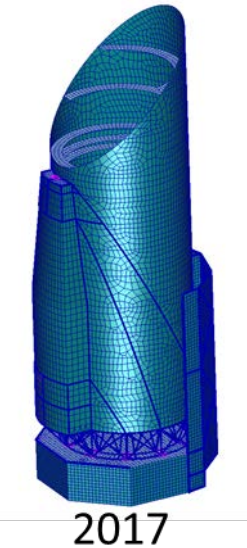
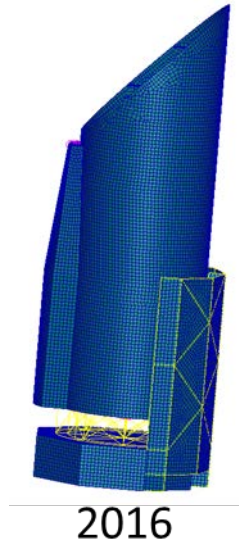
- HabEx is one of 4 large astrophysics facilities being considered by the 2020 decadal
 - It is intended to, among other things, directly image planetary systems around sun-like stars
 - Its main goal is to directly image earth like exoplanets and characterize their atmospheric content
- The current design is on the order of 17.2X5.25 m and the Primary Mirror (PM) is 4 m in diameter
- Performance requirements include an extremely stable system

Line of Sight Stability (Jitter)	< 0.5 milli-arc-seconds per axis
Wavefront Error Stability	1 to 250 pm depending on coronagraph and spatial frequency



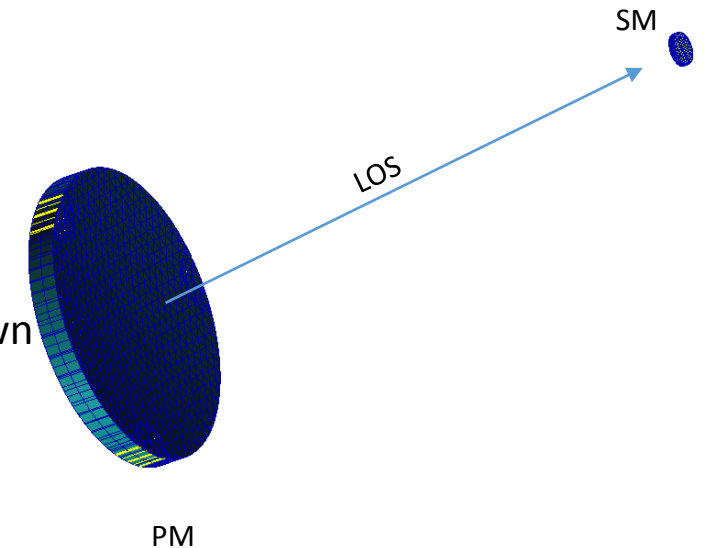
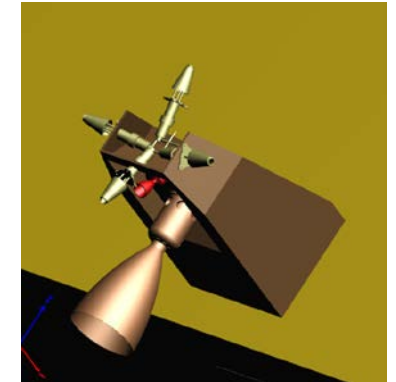
Objective

- As part of the HabEx Pre-Phase A (feasibility) study, structural dynamic analyses have been performed to provide Systems Engineers (SE) with order of magnitude estimates of dynamic responses
 - Results are consistent with feasibility studies and are based on first cut requirements, assumptions, and simplified inputs
 - These results are rolled into system level performance predictions
 - The objective of this presentation is to describe the models used, simulations performed, and results



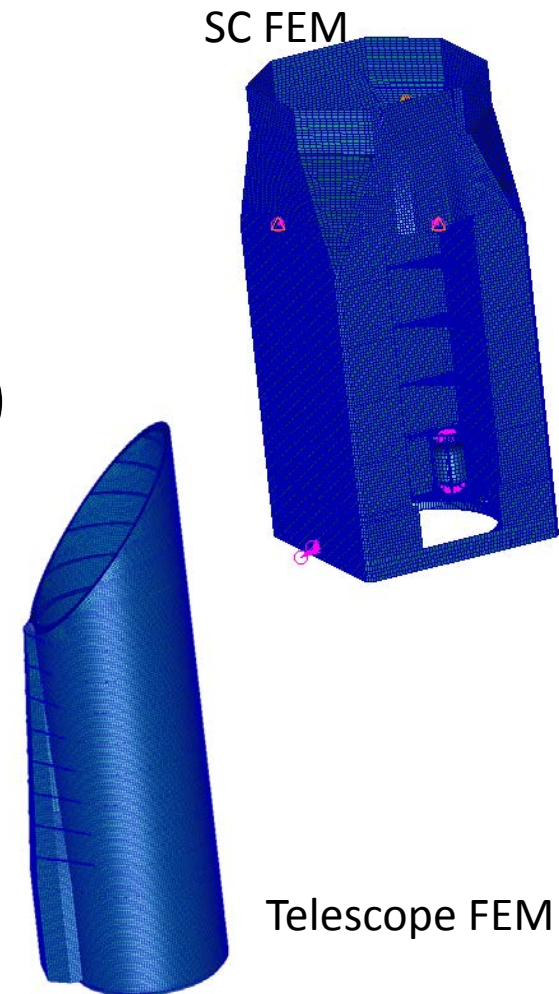
Jitter

- Line Of Sight (LOS) jitter < 0.5 mas per axis
 - Equates to low Nanolevel allowable linear and rotational jitter motion
- Dynamic analyses were performed to predict jitter due to Attitude Control System (ACS) dynamic inputs
 - **Ring Down** - Transient ring down associated with turning the system
 - ACS thrusters will start turning the system then fire again to stop that motion
 - How long does it take for transient jitter to subside?
 - **Jitter** while pointing at a target and collecting data
 - Micro-Thrusters (MT) will be used to maintain orientation
 - Thrust level is expected to change, ramp up/down, very slowly
 - Considered to have no frequency content
 - The MT have a continuous noise level up to 10 Hz and then it ramps down
 - The MT noise is the only identified dynamic disturbance during science windows



Model

- The analyses was performed via the Finite Element (FE) method
- MSC/NASTRAN was used as the solver and MSC/PATRAN was the pre/post processor
- The two primary systems are the Spacecraft (SC) and the telescope
 - SC
 - Modeled by JPL personnel
 - Telescope
 - Modeled by MSFC personnel



Model

- The SC and Telescope FEMs were integrated by JPL personnel
- FEM details

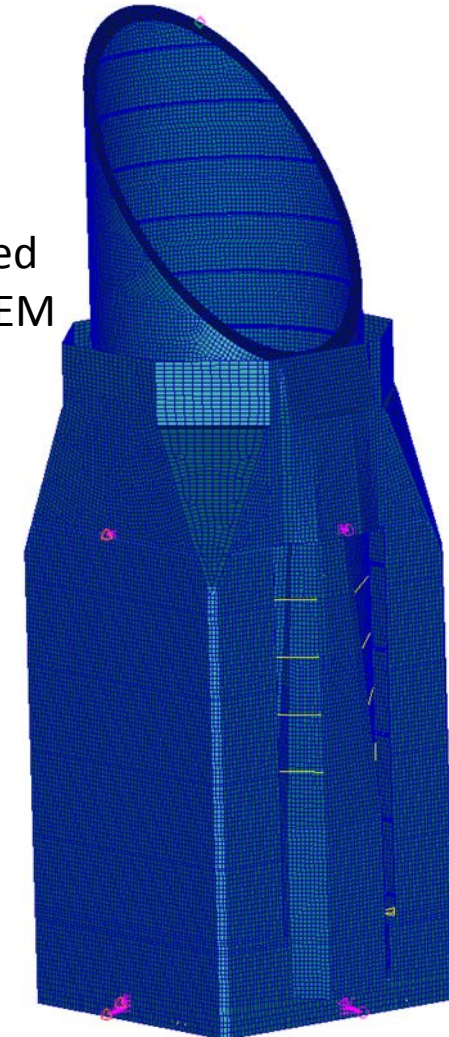
Grid	106381
Linear Elements	7382
Planar Elements	109023
Solid Elements	128
Point Elements	9596
RBE2 Elements	302
RBE3 Elements	53
MPC's	12

Approximately 636K DOF

- FEM mass properties

Mass, Kg	10,687
CGx, m	0.00
CGy, m	-.25
CGz, m	2.04

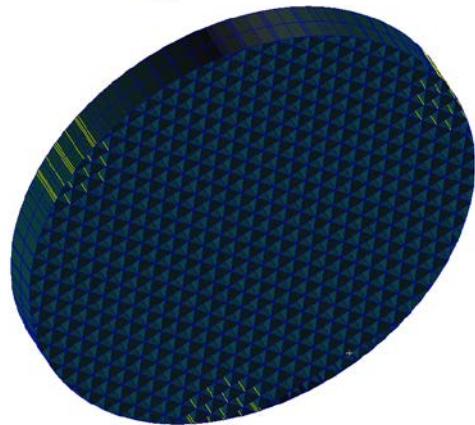
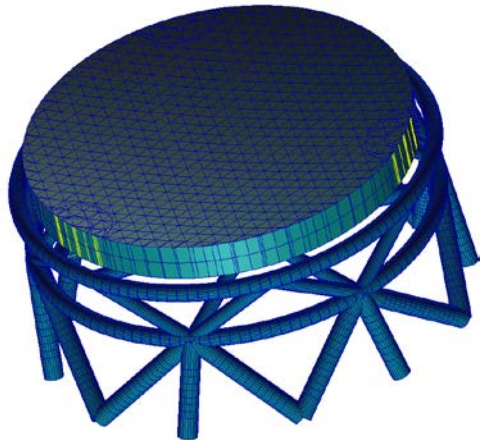
Integrated
HabEx FEM



Model

Mirror Models

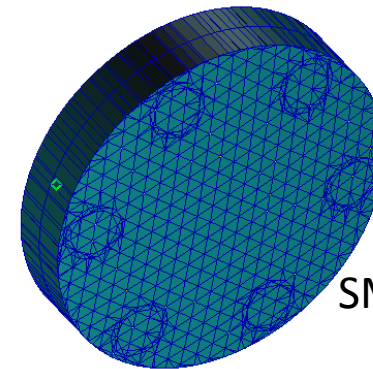
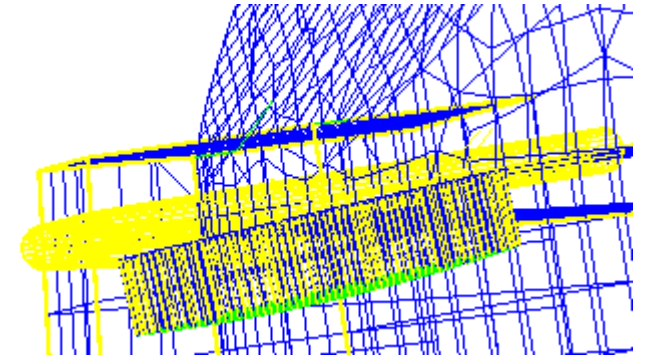
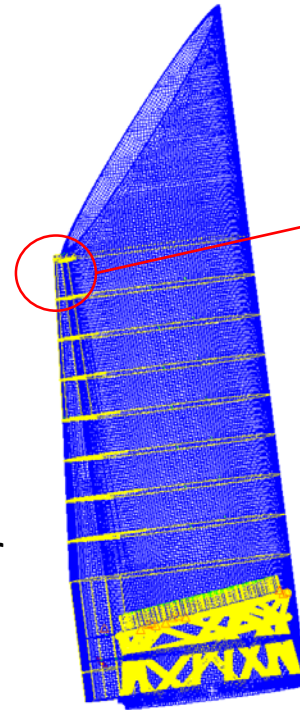
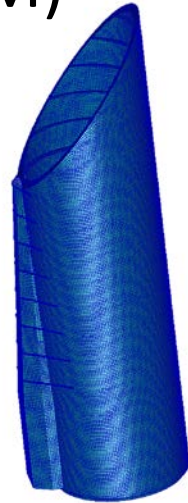
- Primary Mirror (PM)



PM FEM created using
The Arnold Mirror Modeler

PM has \approx 14K DOF

- Secondary Mirror (SM)

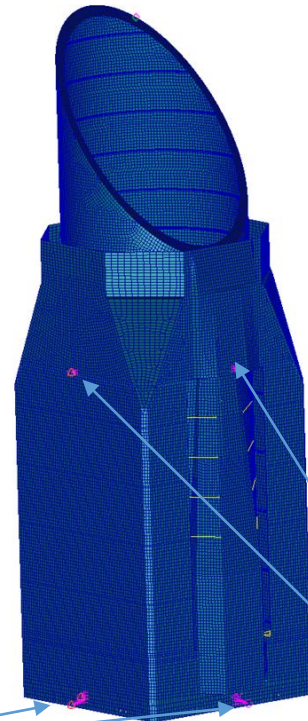
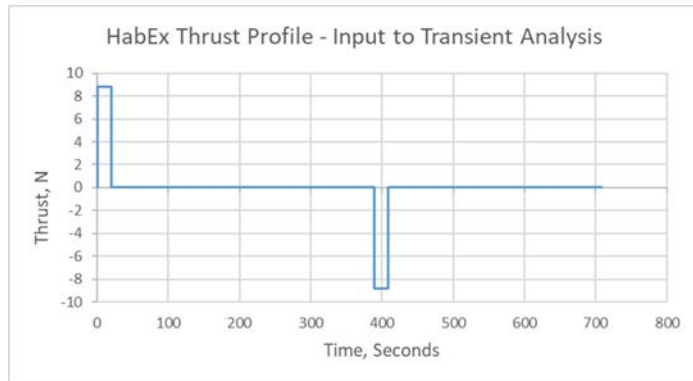


SM has \approx 10K DOF

Analysis Input Loads and Locations

- Ring Down

- JPL provided thrust loads associated with a simple maneuver
- Step function applied at 2 of the 4 ACS thruster locations – Y axis selected
- 8.8N for 20.5 s, drift 368 s, -8.8N for 20.5 s



4 ACS thruster locations and 4 large MT pod locations 90° apart

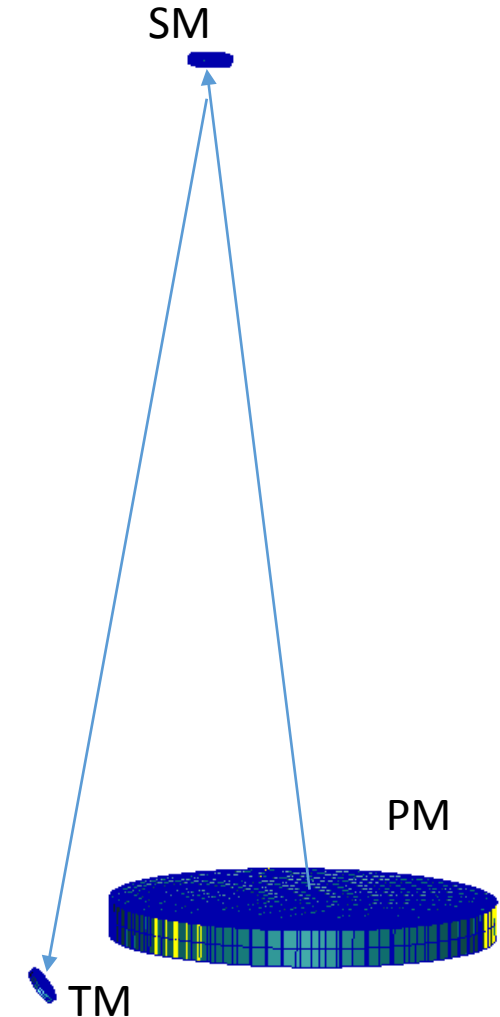
- Jitter

- MT noise is the disturbance during science windows
- Noise input at all 8 MT locations
- Large MT pods
 - $F=0.8 \mu\text{N}$, $.1 < f < 10 \text{ Hz}$ (spec)
 - $F=0.8 \mu\text{N}$, $.1 < f < 20 \text{ Hz}$ (applied)
- Small MT pods
 - $F=0.4 \mu\text{N}$, $.1 < f < 10 \text{ Hz}$ (spec)
 - $F=0.4 \mu\text{N}$, $.1 < f < 20 \text{ Hz}$ (applied)

4 small MT pod locations 90° apart

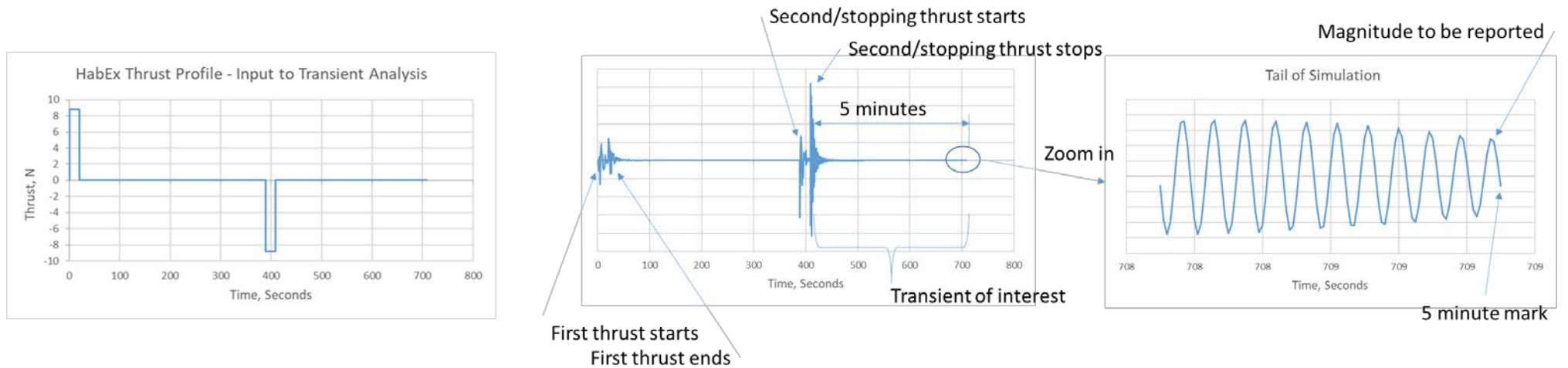
Analysis Output/Results

- NASTRAN Multi-Point Constrain (MPC) equations were incorporated to calculate Relative Motions (RM) within the solution sequence
- Per HabEx SE request, MPCs were written to calculate RM's between the PM/SM, the PM and Tertiary Mirror (TM), and the SM/TM



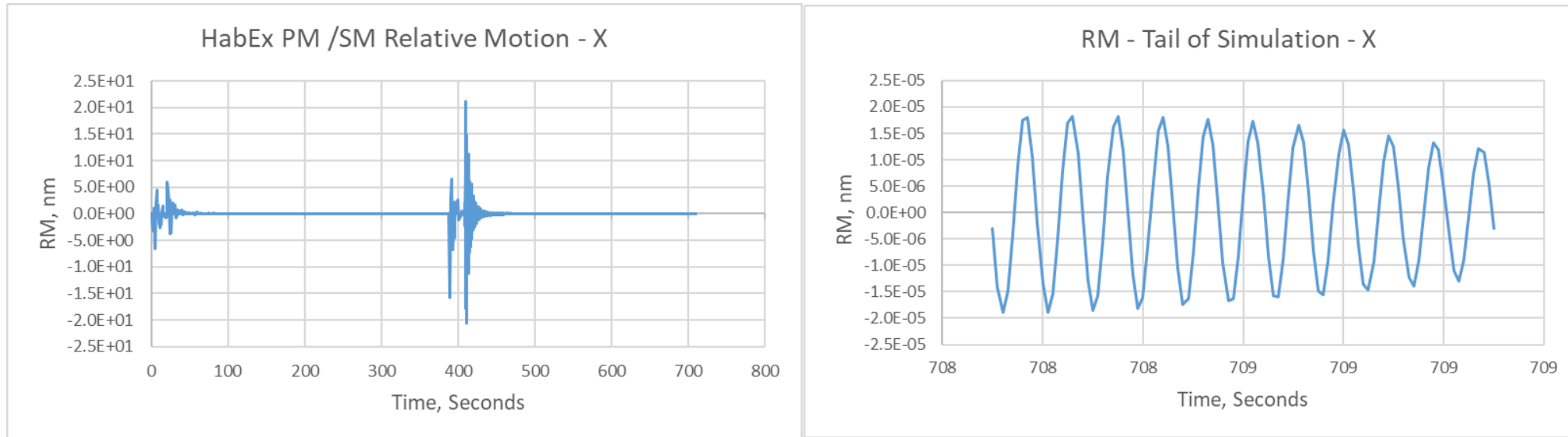
Ring Down Analysis

- Physically, ring down is a transient event
- To assess the design to estimate the ring down time a NASTRAN transient (time domain) dynamic analysis (Solution 112) was performed
 - Required damping of .05% was used
 - No ring down time requirement has been determined in this feasibility study
 - In the absence of a requirement, the simulation was run 5 minutes past the “stopping thrust”
 - PM/SM RM for each DOF was output
 - The max RM 5 minutes after the last thrust was reported



Ring Down Results

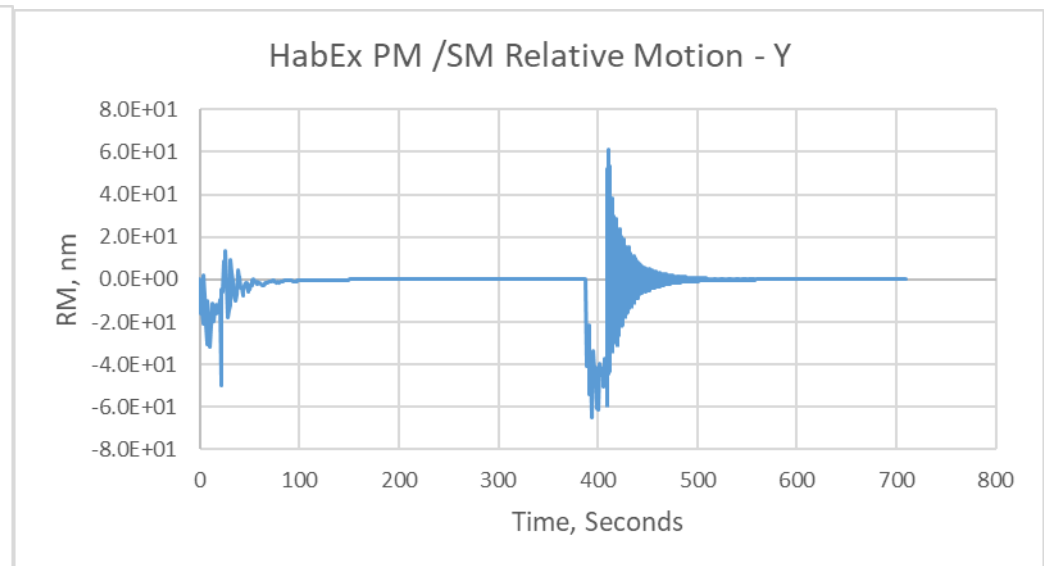
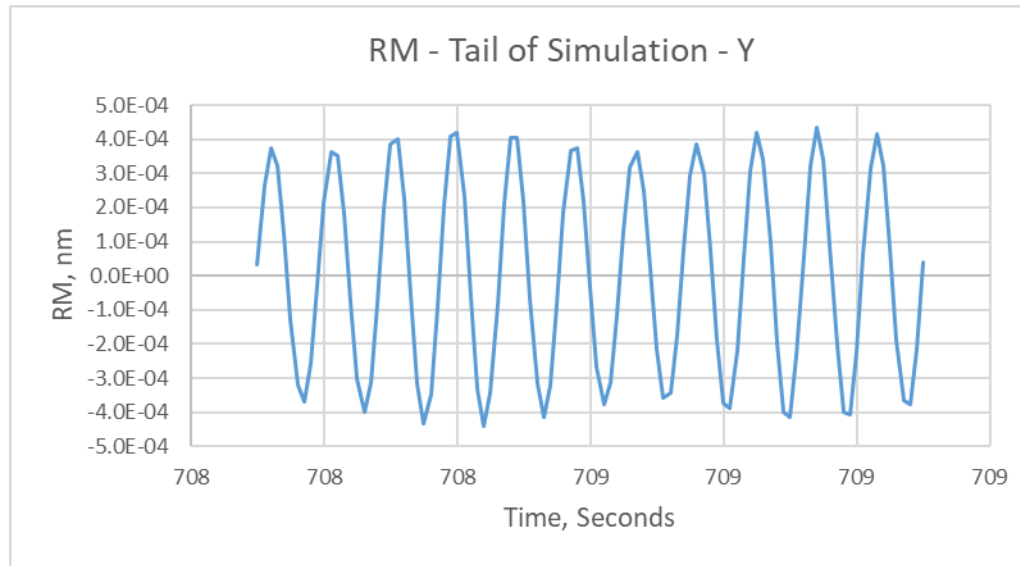
- PM to SM Relative Motion was the selected metric
- X Direction



Peak after 5 minutes = $1.2E-4$ pm

Ring Down Results

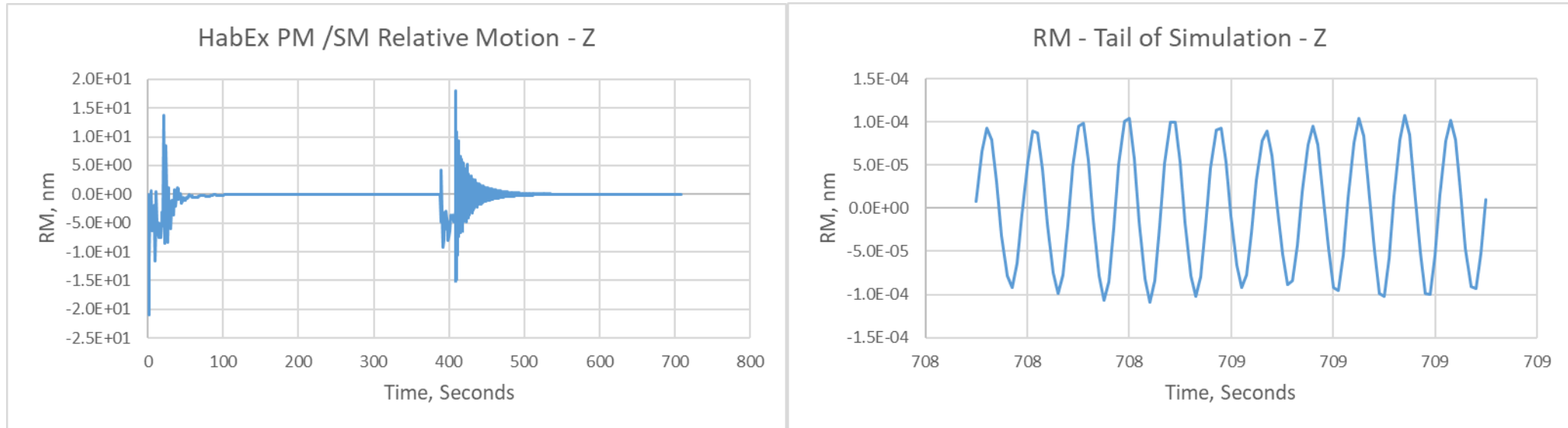
- PM to SM Relative Motion
- Y Direction



Peak after 5 minutes = $4.1E-4$ pm

Ring Down Results

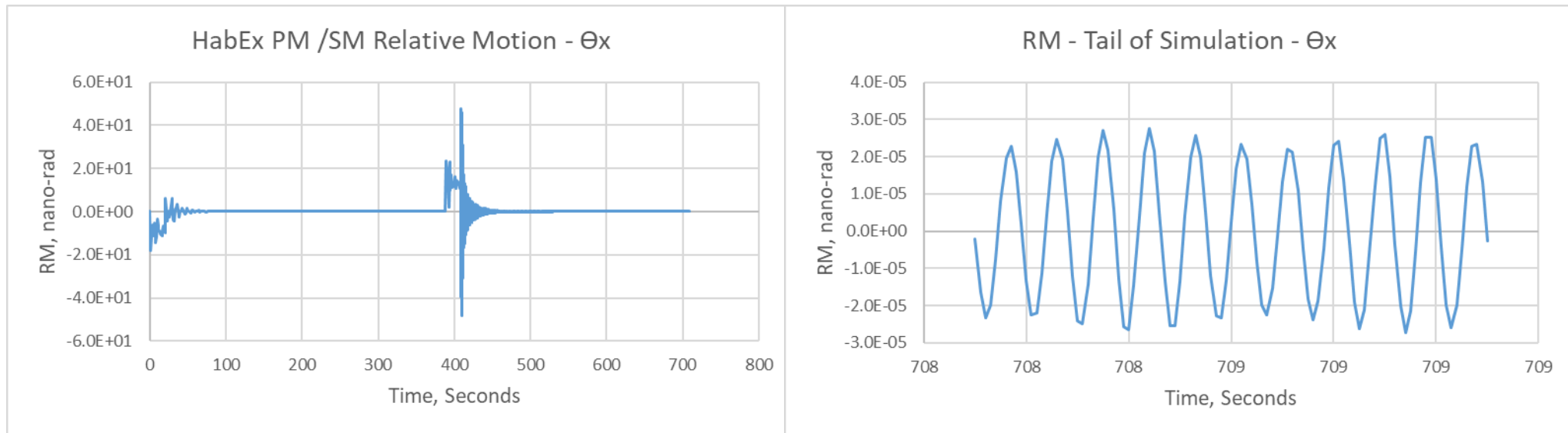
- PM to SM Relative Motion
- Z Direction



Peak after 5 minutes = $1.0E-4$ pm

Ring Down Results

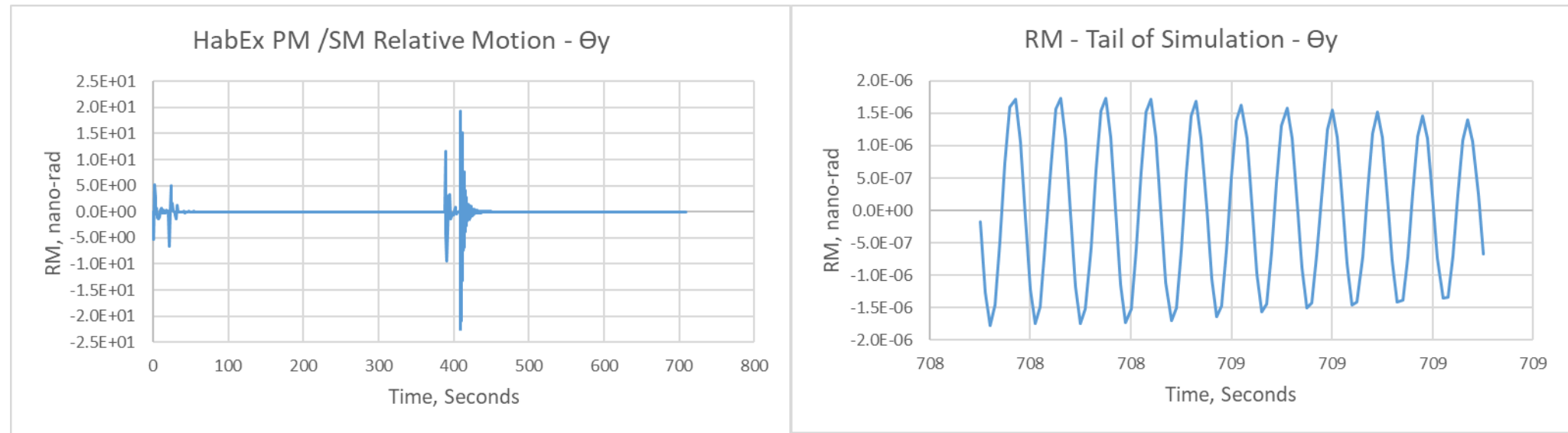
- PM to SM Relative Motion
- Θ_x Direction



Peak after 5 minutes = $2.3E-5$ pico-radians

Ring Down Results

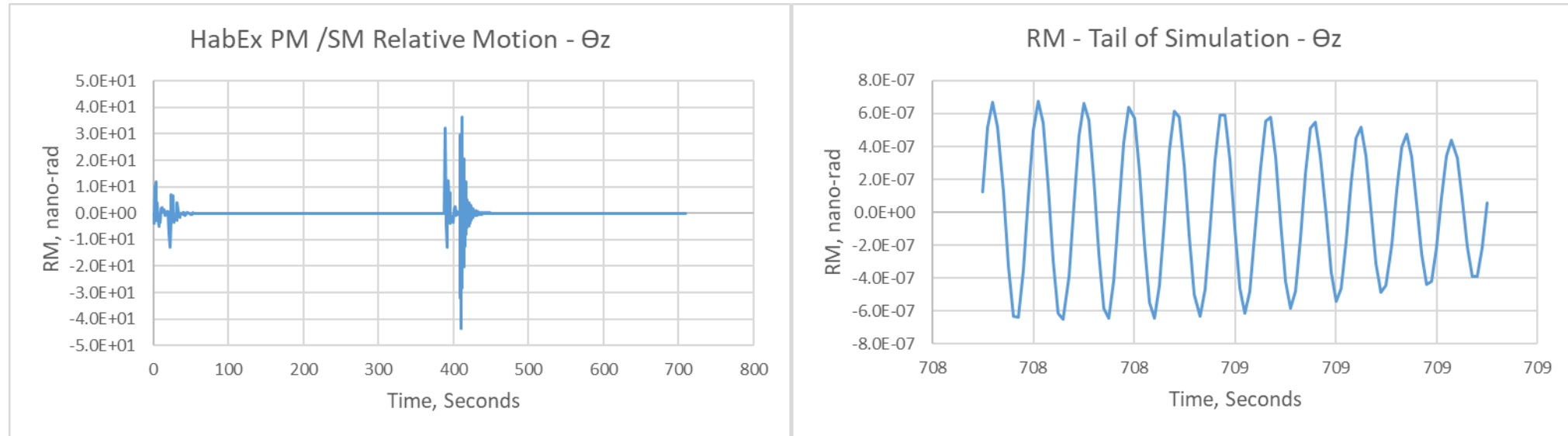
- PM to SM Relative Motion
- Θ_y Direction



Peak after 5 minutes = $1.4E-6$ pico-radians

Ring Down Results

- PM to SM Relative Motion
- Θ_z Direction



Peak after 5 minutes = $4.4E-7$ pico-radians



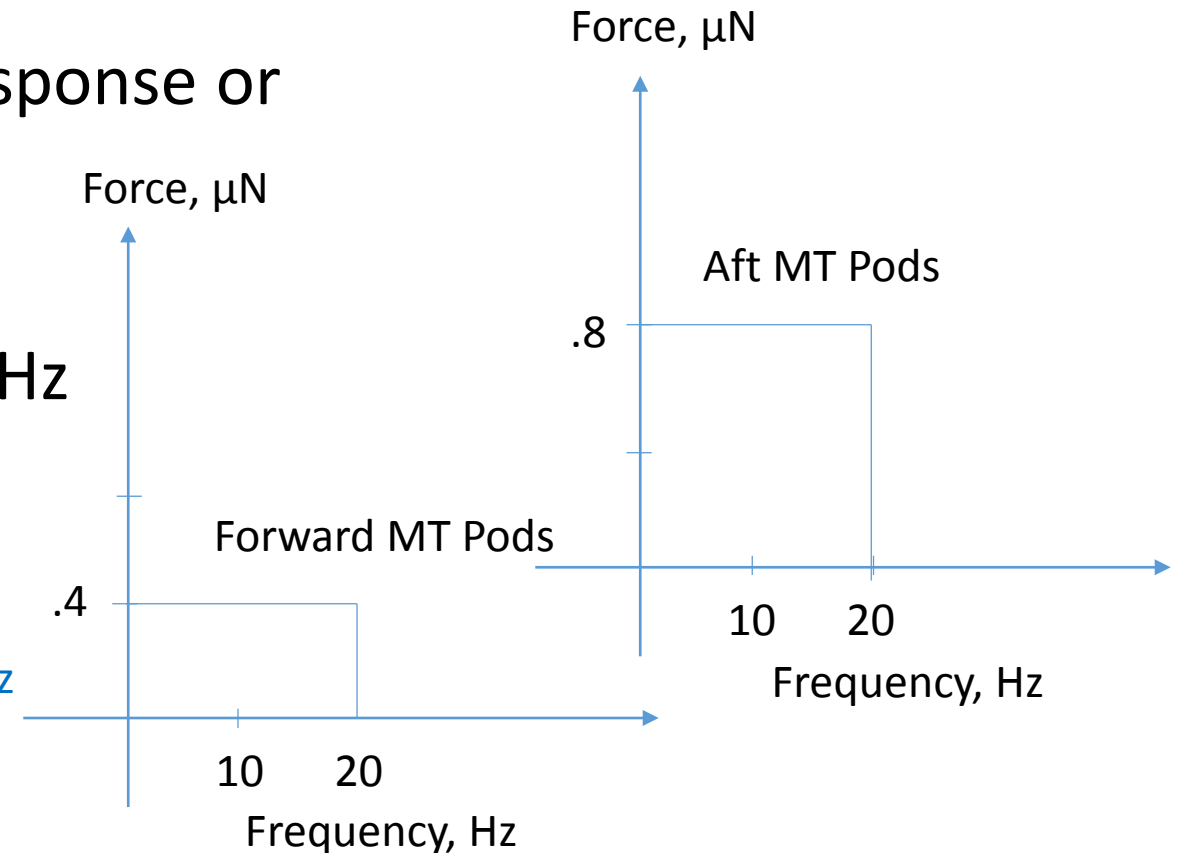
Ring down Conclusion

- No required time to regain stability after thruster induced transients has been determined
- The **maximum linear RM** is on the order of **10^{-4} pm** after 5 minutes of settling time
- The **maximum rotational RM** is on the order of **10^{-5} p-rad** after 5 minutes of settling time
- The LOS stability requirements are in the nm range
- Therefore, with small fractions of pm range motion predicted, 5 minutes of settling time is a conservative number to use in this early study

Jitter Analysis

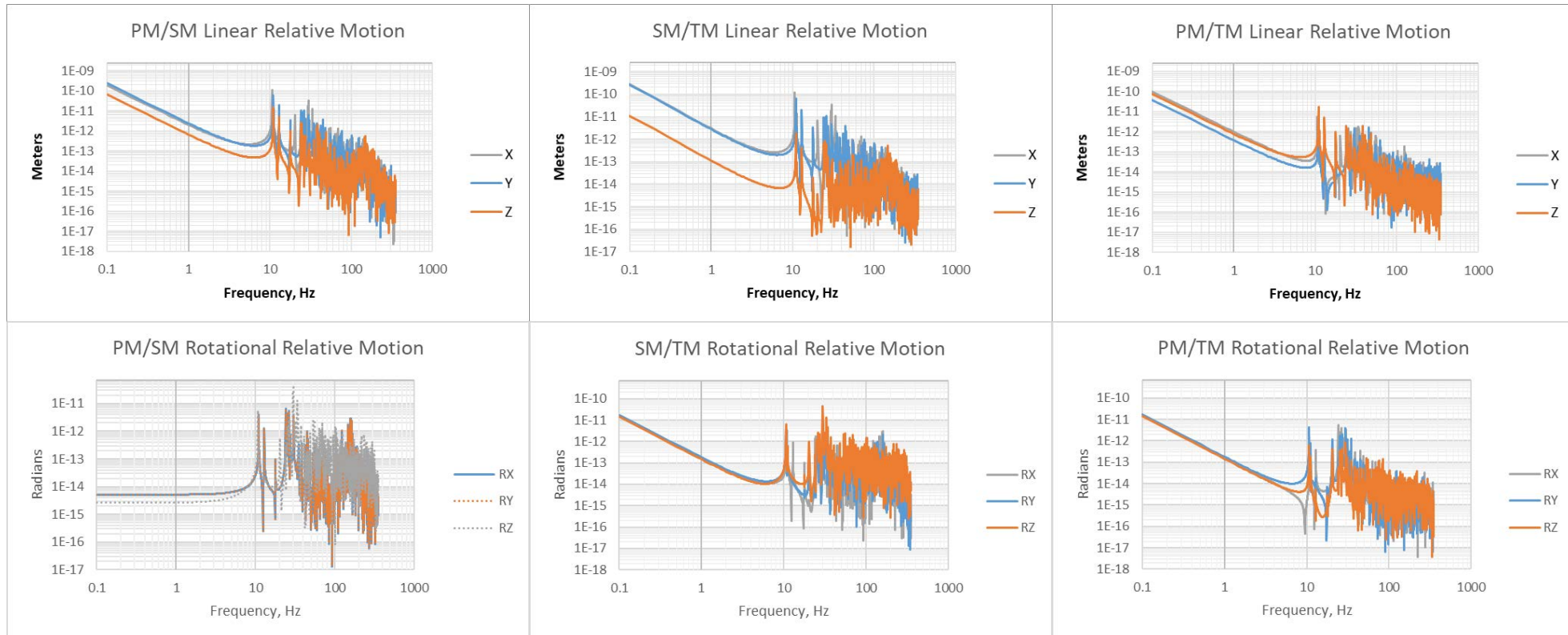
- Jitter is a continuous event that occurs as long as the source vibration is in operation
- It is well suited for a frequency response or harmonic dynamic analysis
- Damping was required to be .05%
- Results were predicted up to 350 Hz

Noise specified to 10 Hz
Conservatively applied out to 20 Hz



Jitter Results

- Results up to 300 Hz were provided to the HabEx SE team and were rolled into system level performance assessments





Final Comments

- Feasibility fidelity dynamic analyses have been performed for the HabEx Pre-Phase A engineering effort
- Based on the provided FEM and required inputs, no show stoppers were identified