Radiation Hardness Assurance: Evolving for *NewSpace*

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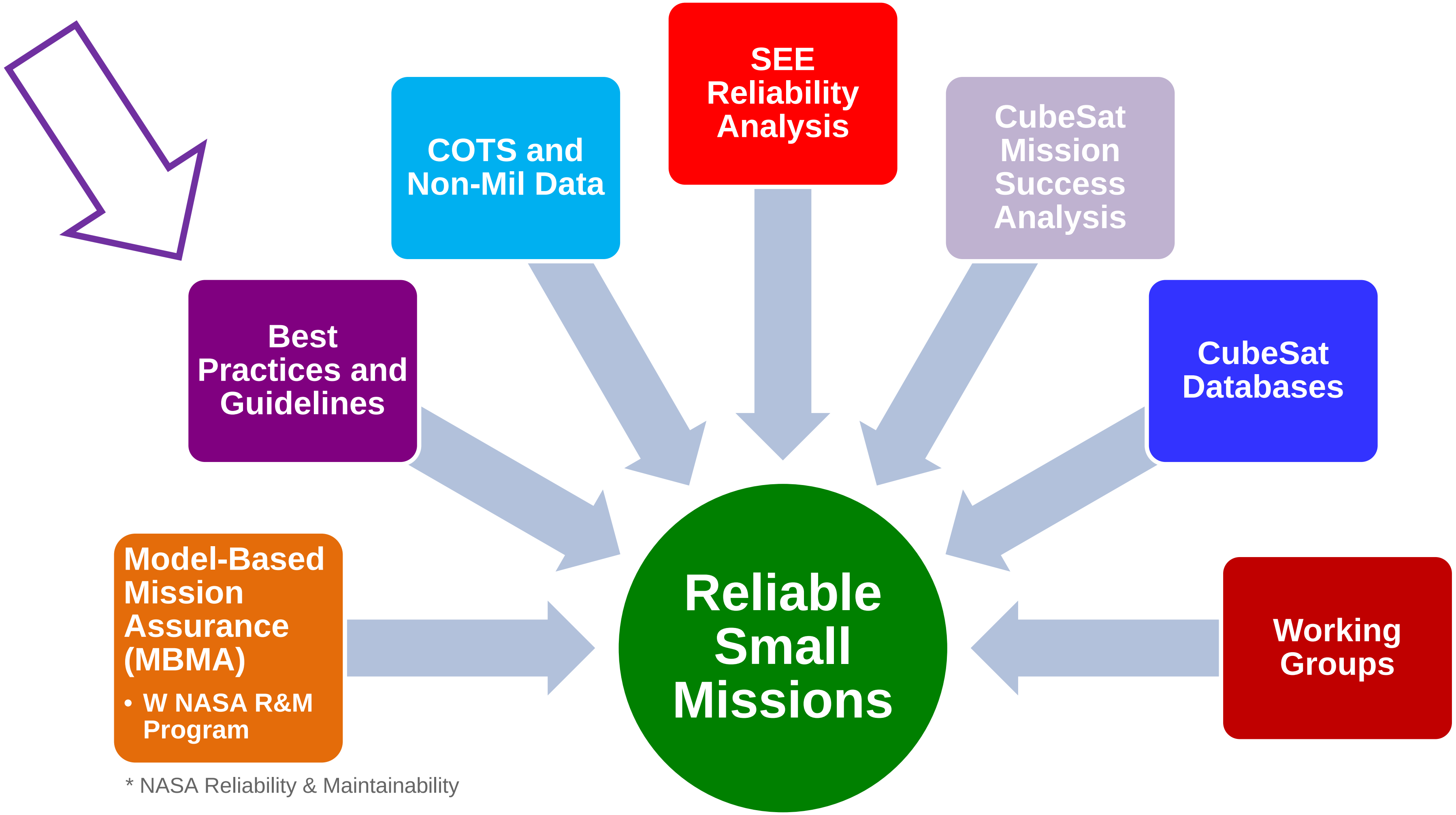
Acronyms

CME	Coronal Mass Ejection
COTS	Commercial Off The Shelf
DDD	Displacement Damage Dose
EEE	Electrical, Electronic, and Electromechanical
ELDRS	Enhanced Low Dose Rate Sensitivity
EP	Enhanced Performance
ESA	European Space Agency
GCR	Galactic Cosmic Ray
GOMAC	Government Microcircuits Applications and Critical Technologies Conference
GSFC	Goddard Space Flight Center
GSN	Goal Structuring Notation
HEART	Hardened Electronics and Radiation Technology
LEO	low earth orbit
LET	Linear Energy Transfer
MBMA	model based mission assurance
MRQW	Microelectronics Reliability and Qualification Workshop
NAND	Negated AND or NOT AND
NASA	National Aeronautics and Space Administration
NEPP	NASA Electronic Parts and Packaging
NEPP ETW	NASA Electronic Parts and Packaging (NEPP) Program Electronics Technology Workshop
NSREC	Nuclear and Space Radiation Effects Conference

RADECS	Radiation Effects on Components and Systems
RHA	Radiation Hardeness Assurance
SAA	South Atlantic Anomaly
SEE	Single Event Effects
SEE/MAPLD	SEE-MAPLD Single Event Effects (SEE) Symposium/
	Military and Aerospace Programmable Logic Devices (MAPLD) Workshop
SEGR	Single Event Gate Rupture
SEL	Single Event Latchup
SEP	Single Event Effects Phenomena (includes SEU, SEL, SEGR and SET)
SERESSA	School on the Effects of Radiation on Embedded Systems for Space Applications
SET	Single Event Transient
SEU	Single Event Upset
SLU	Saint Louis University
SwaP	Size, weight, and power
TID	Total Ionizing Dose
TID	Total Ionizing Dose
TMR	triple-modular redundancy
TNID	Total Non-Ionizing Dose
UV	Ultra-Violet

NEPP Program- Small Mission Efforts



Best Practices and Guidelines

COTS and Non-Mil Data

SEE Reliability Analysis

CubeSat Mission Success Analysis

CubeSat Databases

Model-Based Mission Assurance (MBMA)
- **W NASA R&M Program**

Reliable Small Missions

Working Groups

* NASA Reliability & Maintainability

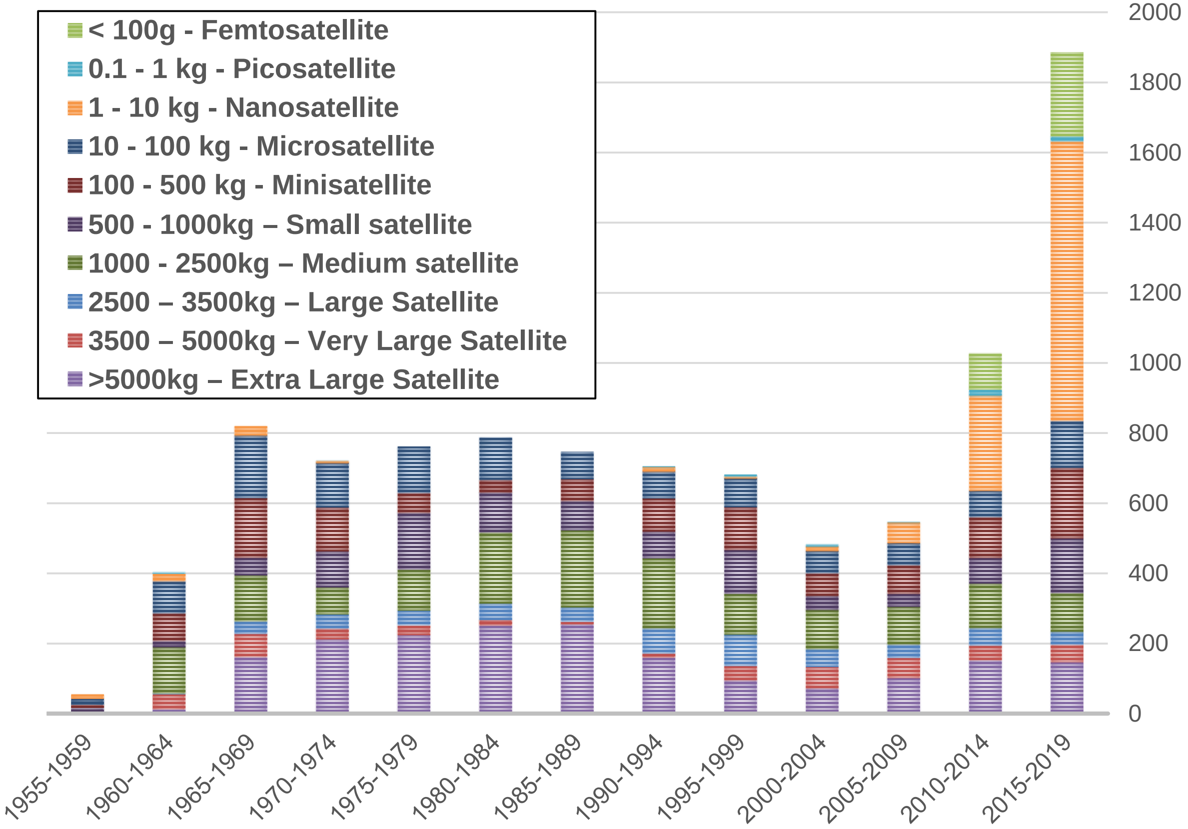
To be presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16th, 2019

Outline

- **New Space and SmallSat Considerations**
- **The Natural Space Radiation Environment Hazard**
- **Radiation Effects on Micro-Electronics**
- **Hardness Assurance, as a Discipline, with its Challenges**
  - **New Technologies**
  - **New Architectures**
  - **Unbound Risks**
- **Building Smart Requirements**
- **Risk Acceptance and Guidance**

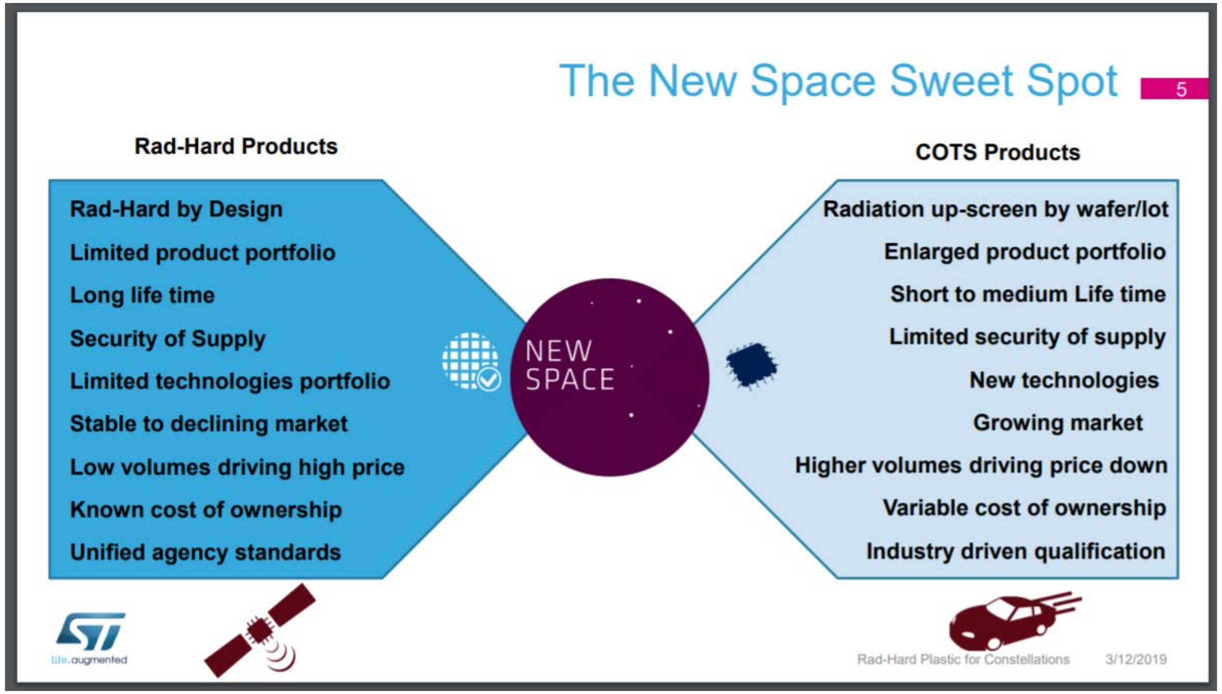
New Space – New Point of View

SmallSats Come in Many Sizes



Seradata SpaceTrak Data

Component Grades are Merging



The New Space Sweet Spot 5

Rad-Hard Products

- Rad-Hard by Design
- Limited product portfolio
- Long life time
- Security of Supply
- Limited technologies portfolio
- Stable to declining market
- Low volumes driving high price
- Known cost of ownership
- Unified agency standards

NEW SPACE

COTS Products

- Radiation up-screen by wafer/lot
- Enlarged product portfolio
- Short to medium Life time
- Limited security of supply
- New technologies
- Growing market
- Higher volumes driving price down
- Variable cost of ownership
- Industry driven qualification

life.augmented

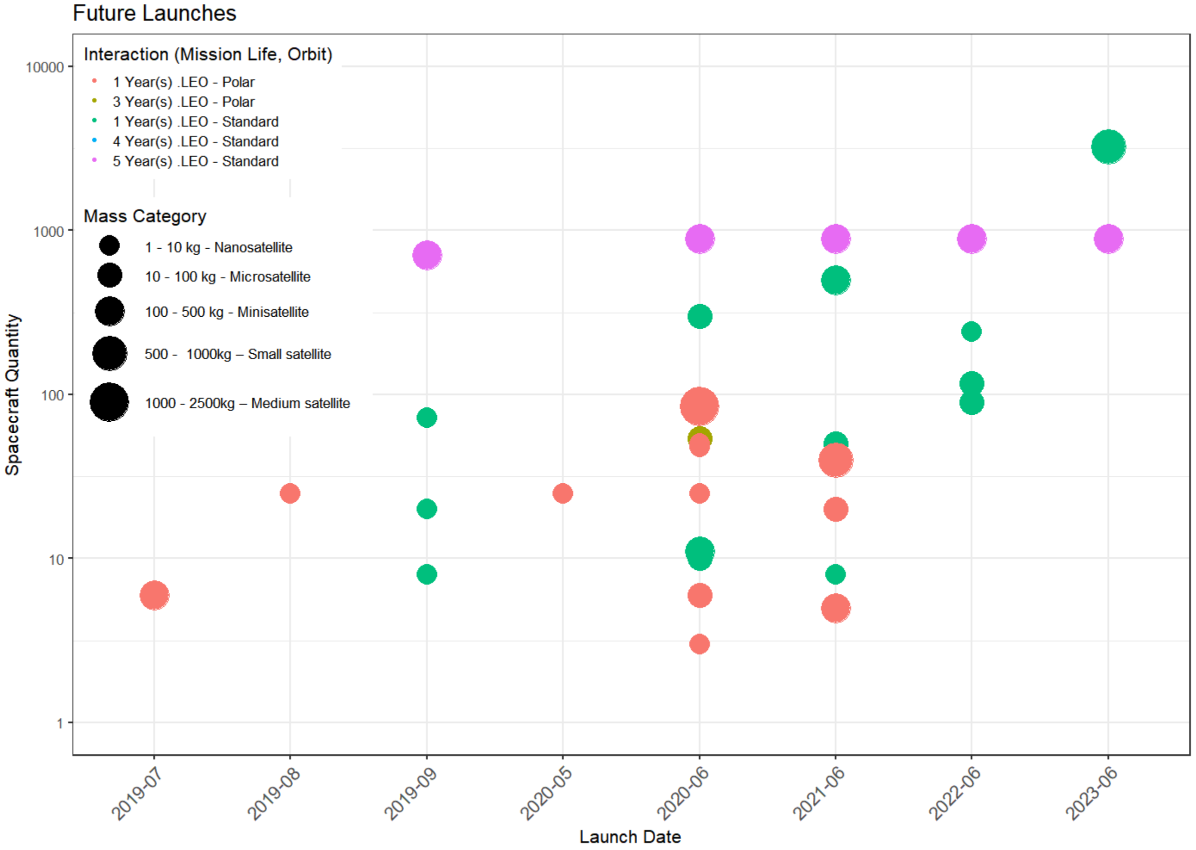
Rad-Hard Plastic for Constellations 3/12/2019

ESSCON : Eccofet

Risk acceptance is being used as a means to enable innovation

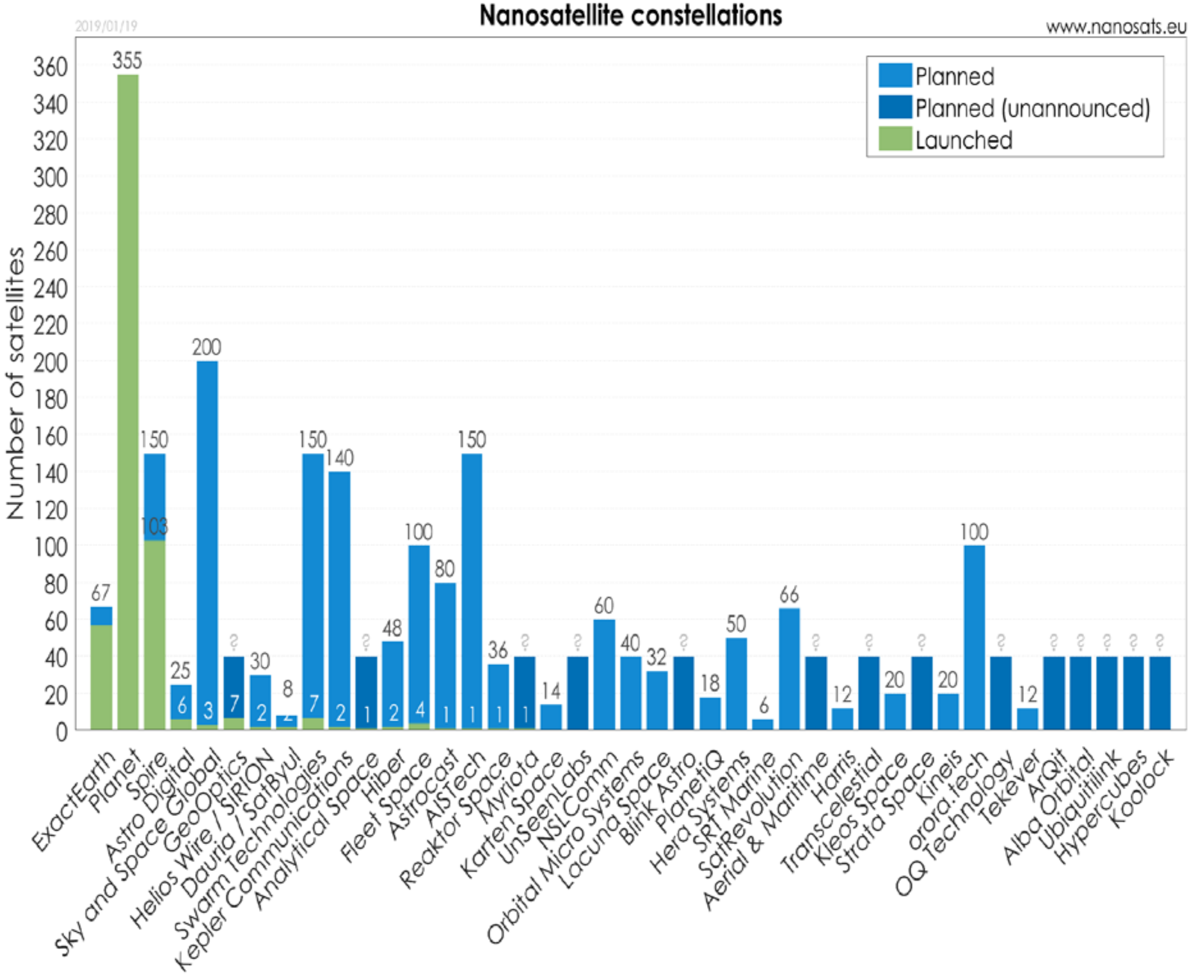
New Space – Looking Ahead

Constellations and Swarms



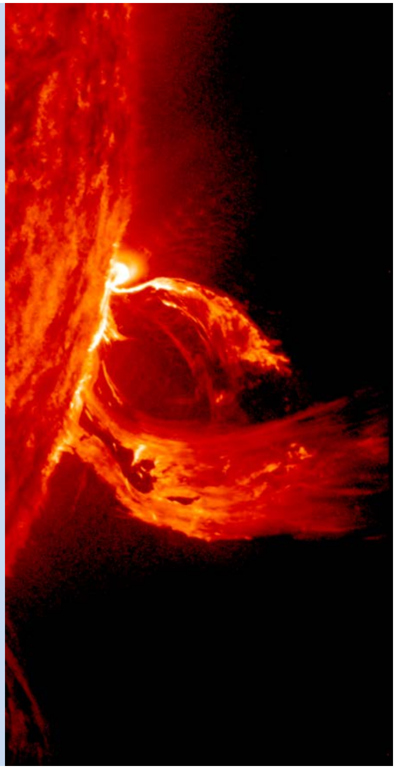
Seradata SpaceTrak Data (Notional Launches)

New Space = New Companies



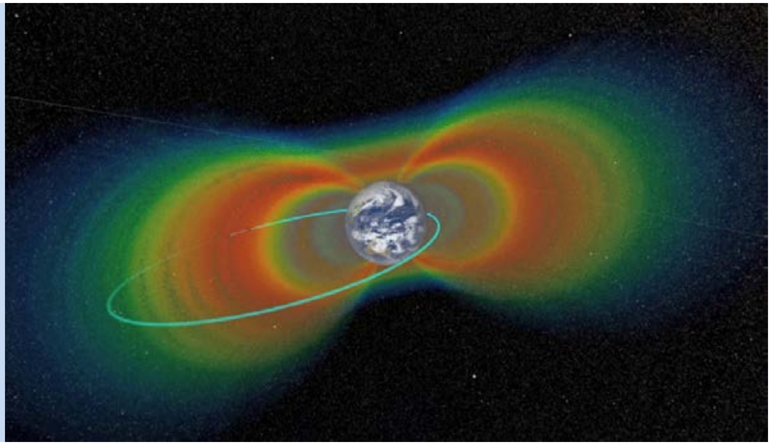
New Space – Same Old Radiation

- **New mission concepts and SmallSat paradigm**
  - Radiation challenges identified in the past are here to stay; adoption of new technologies are often the risk driver
  - Commercial Space, Constellations, Small missions, etc. will benefit from detailed hazard definition and mission specific requirements
- **The need for Radiation Hardness Assurance (RHA)**
  - Radiation effects are a mix of disciplines, evolve with technologies and techniques
  - Misinterpretation of failure modes / misuse of available data can lead to over/under design
  - RHA flow doesn't change, risk acceptance needs to be tailored
- **Some Top Level Resources**
  - NPR-7120.5 – NASA Agency Program Management
  - GPR-8705.4 – NASA Goddard Risk Classification Guidelines
  - NASA-STD-8739.10 – NASA Parts Assurance Standard



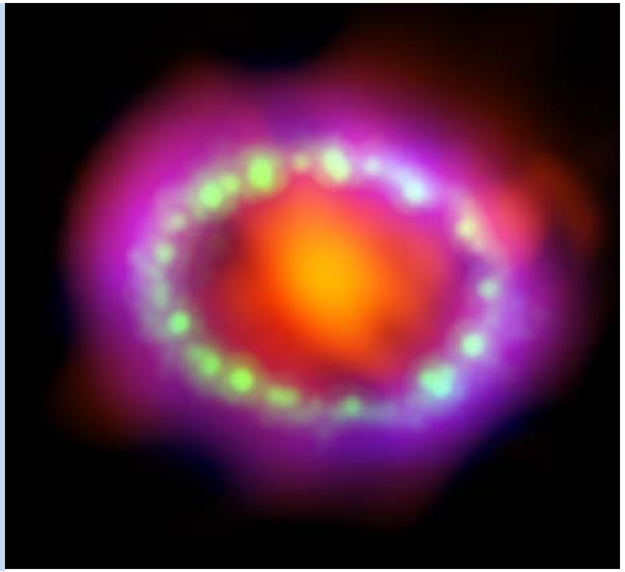
Solar Events & Activity

https://sdo.gsfc.nasa.gov



Trapped Radiation

https://www.nasa.gov/van-allen-probes



Supernovae

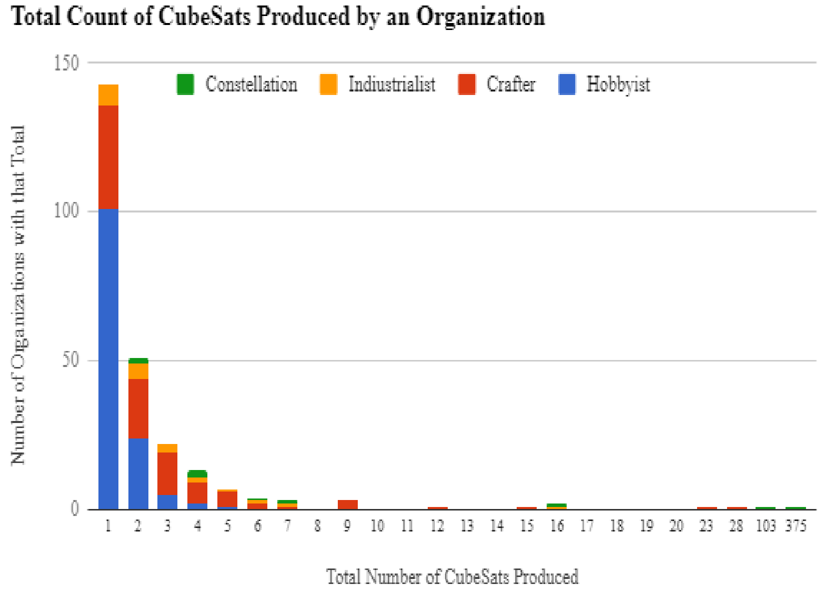
NASA, ESA, and L. Hustak (STScI)

To be presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16th, 2019

Who Needs This Guidance?

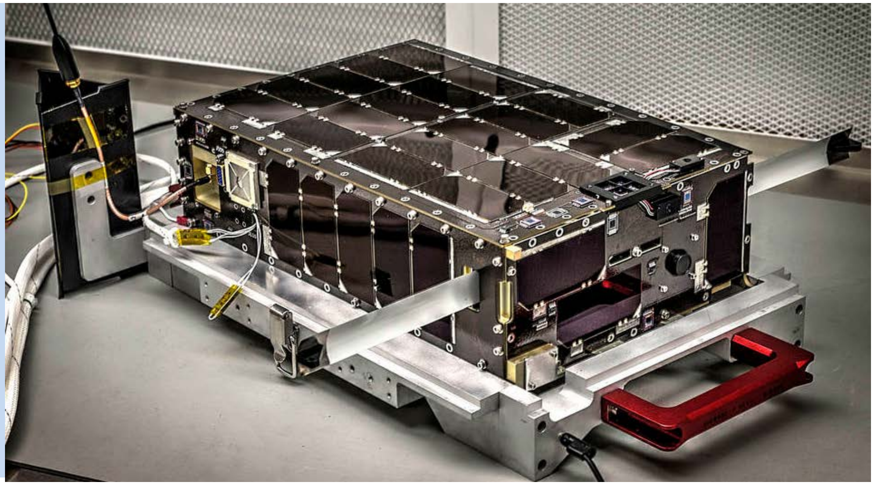
- **Universities / CubeSats**
  - May be first-time designers, or previous missions did not have requirements
  - Schedule driven, limited time for development
  - Rideshares – could end up in multiple environments
- **Space Agencies / Government**
  - More compact designs in new destinations
  - Cost savings of SmallSat platform, with more reliable outcome
  - More willing to trade risk for capability
- **Device / Subsystem Manufacturers**
  - Product / Device offerings: Space Plastic, EP, LeanRel, radiation tolerant, modified HiRel, etc.
  - Fault tolerance in designs

CubeSat Metrics



Michael Swartwout, SLU CubeSat Database

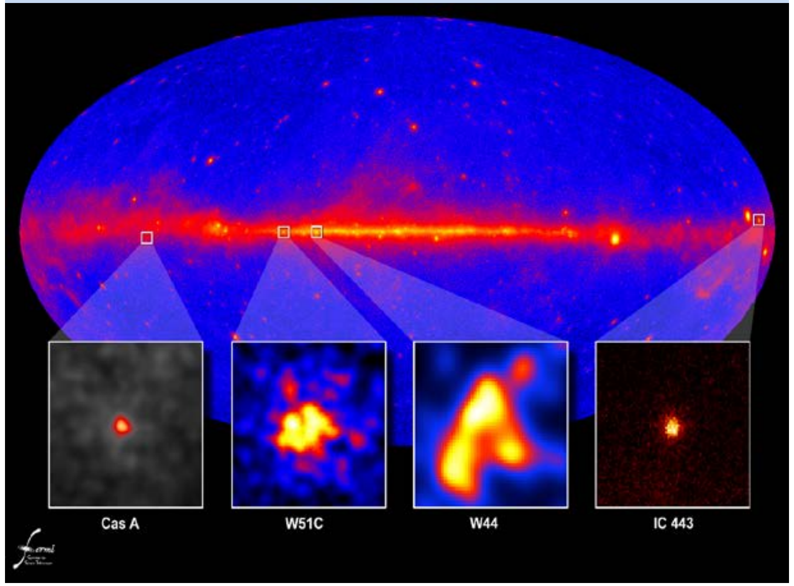
Dellingr



NASA's Goddard Space Flight Center/Bill Hrybyk

Natural Space Radiation Environment

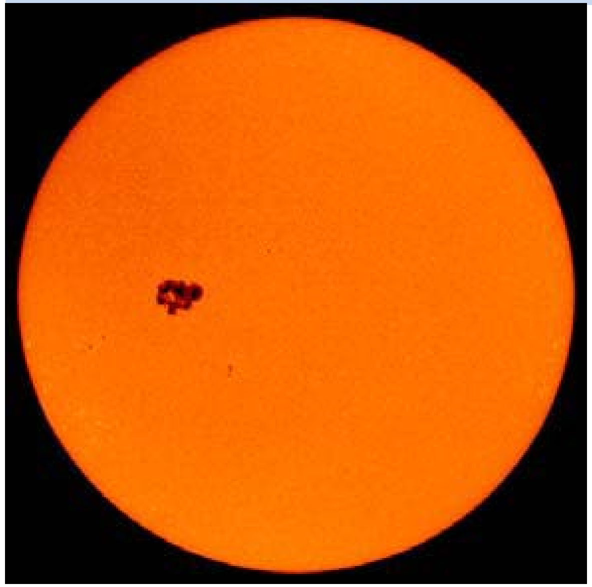
Galactic Cosmic Rays



Energetic supernovae remnants (~GeV, Z=1-92)

Originate outside of our solar system

Solar Activity

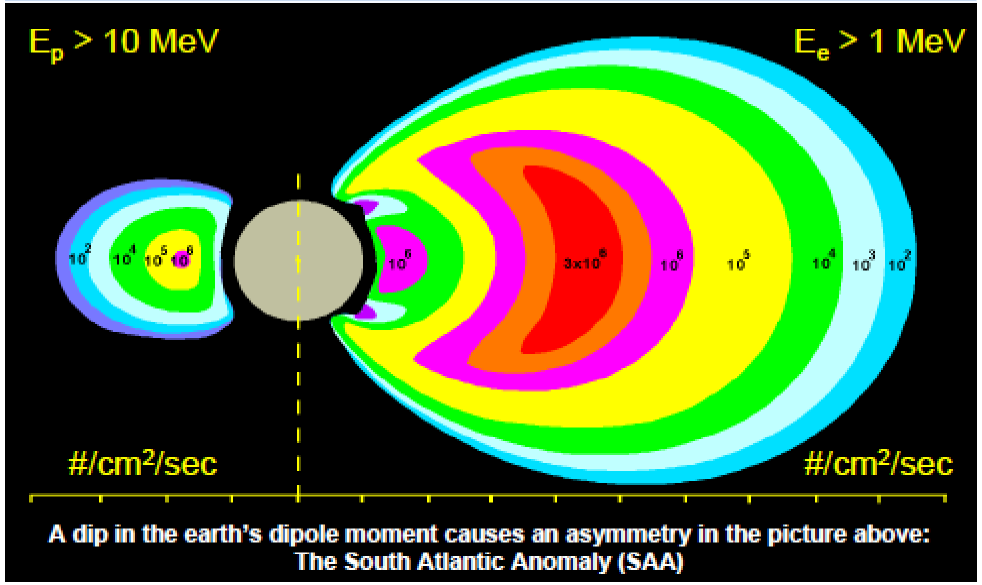


Solar Wind, Solar Cycle

CMEs (proton rich)

Flares (heavy ion rich)

Trapped Particles in Planetary Magnetic Fields



A dip in the earth's dipole moment causes an asymmetry in the picture above: The South Atlantic Anomaly (SAA)

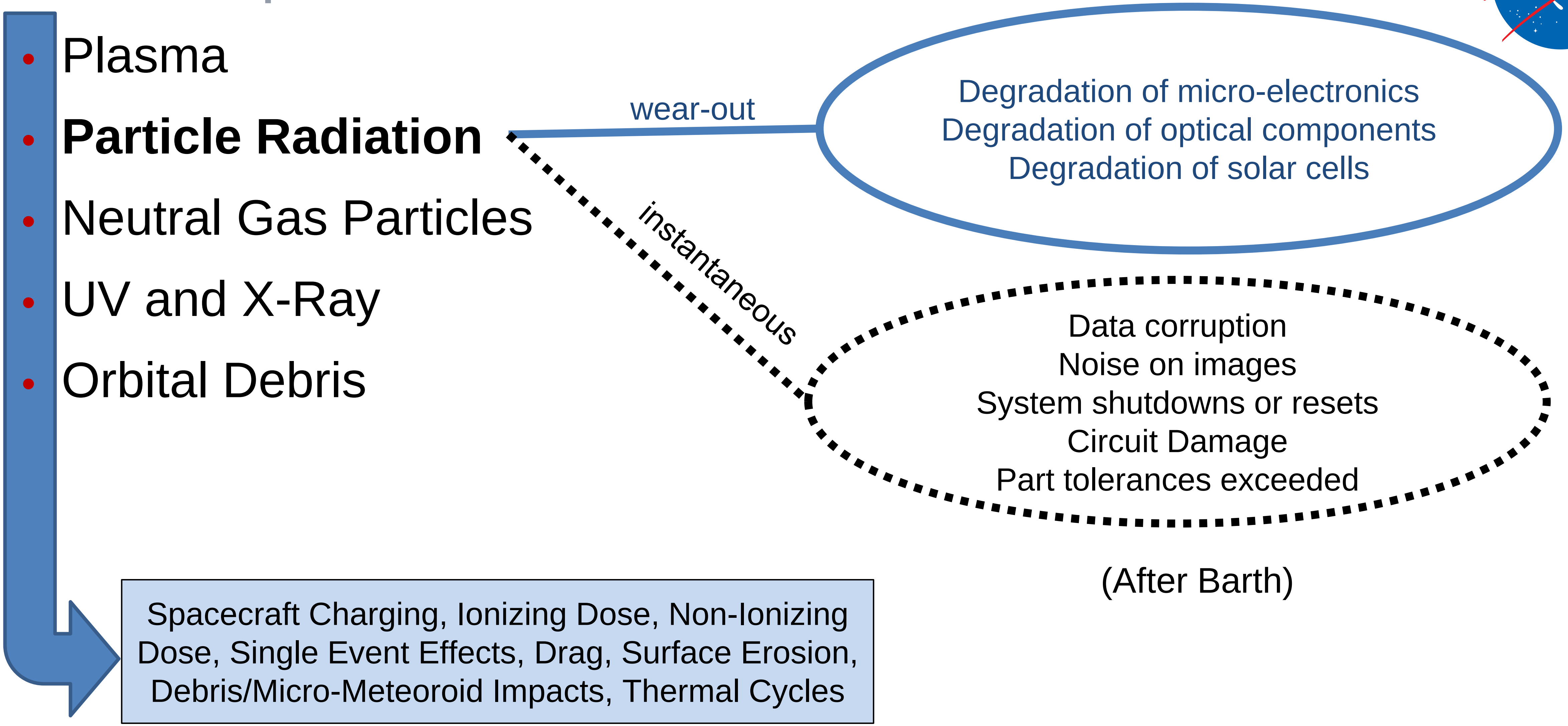
Fluctuate with Solar Activity and Events

Not a perfect dipole

Protons and Electrons trapped at different L-shell values and energies

Images from left to right – NASA FERMI X-ray telescope, Solar Dynamics Observatory, Janet Barth (radhome.gsfc.nasa.gov)

Natural Space Radiation Environment



- Plasma
- **Particle Radiation**
- Neutral Gas Particles
- UV and X-Ray
- Orbital Debris

wear-out

Degradation of micro-electronics
Degradation of optical components
Degradation of solar cells

instantaneous

Data corruption
Noise on images
System shutdowns or resets
Circuit Damage
Part tolerances exceeded

Spacecraft Charging, Ionizing Dose, Non-Ionizing Dose, Single Event Effects, Drag, Surface Erosion, Debris/Micro-Meteoroid Impacts, Thermal Cycles

(After Barth)

Natural Space Radiation Environment

- **Particle Radiation**

wear-out

Degradation of micro-electronics
Degradation of optical components
Degradation of solar cells

instantaneous

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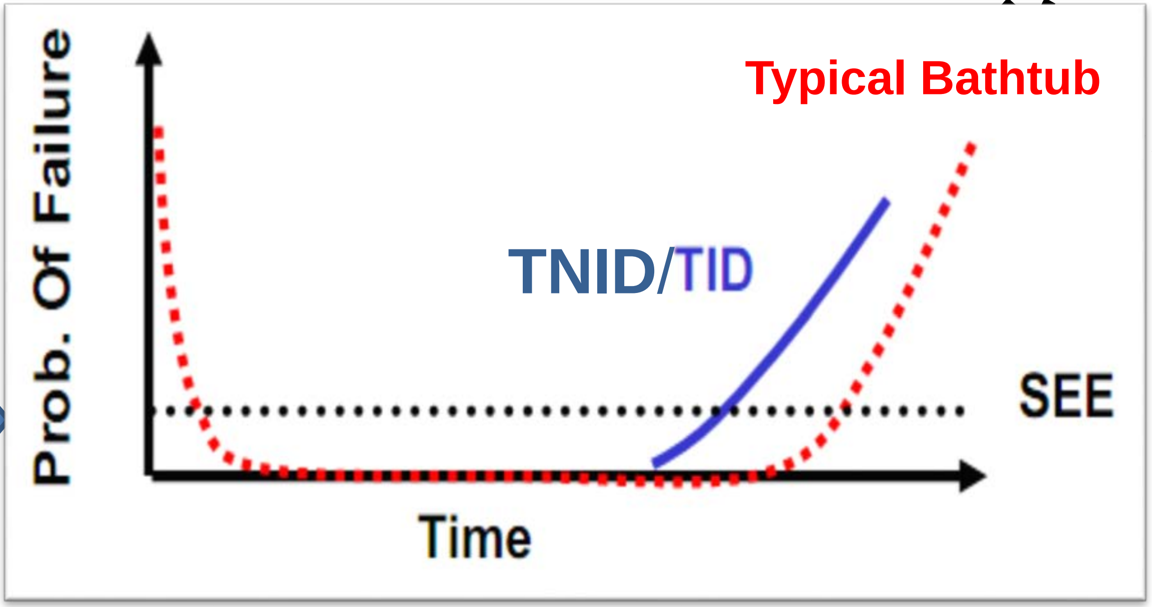


(After Buchner)

Conventional Units Explanation

Degradation

- **Total Ionizing Dose (TID)**
  - Absorbed dose (rad(Si))

    1 rad = 100 erg/g = 0.01 J/kg; 100 rad = 1 Gy
  - Always specified for a particular material

    1 rad(SiO_2), 10 krad(Si), 100 Gy(H_2O)
  - This is not exposure (R), or dose equivalent (Sv)
- **Total Non-Ionizing Dose (TNID)**
  - Fluence (particles/cm^2)

    Number of particles per unit area
  - Displacement Damage Dose (DDD)

    Specified at a given incident particle energy - e.g., 10 MeV p+, 50 MeV p+, 1 MeV eq. neutrons, etc.

Single Event

- **Linear Energy Transfer (LET)**
  - Stopping power normalized to target material

$$S = -\frac{dE}{dx} \Rightarrow \mathrm{LET} = -\frac{1}{\rho}\frac{dE}{dx}$$

  - Units are MeV·cm^2/mg
- **Cross Section (σ)**
  - Device particle interaction (cm^2)
  - Used in calculation of rate

    Can be /device or /bit per time interval

Degradation Contributors vs. Single Event

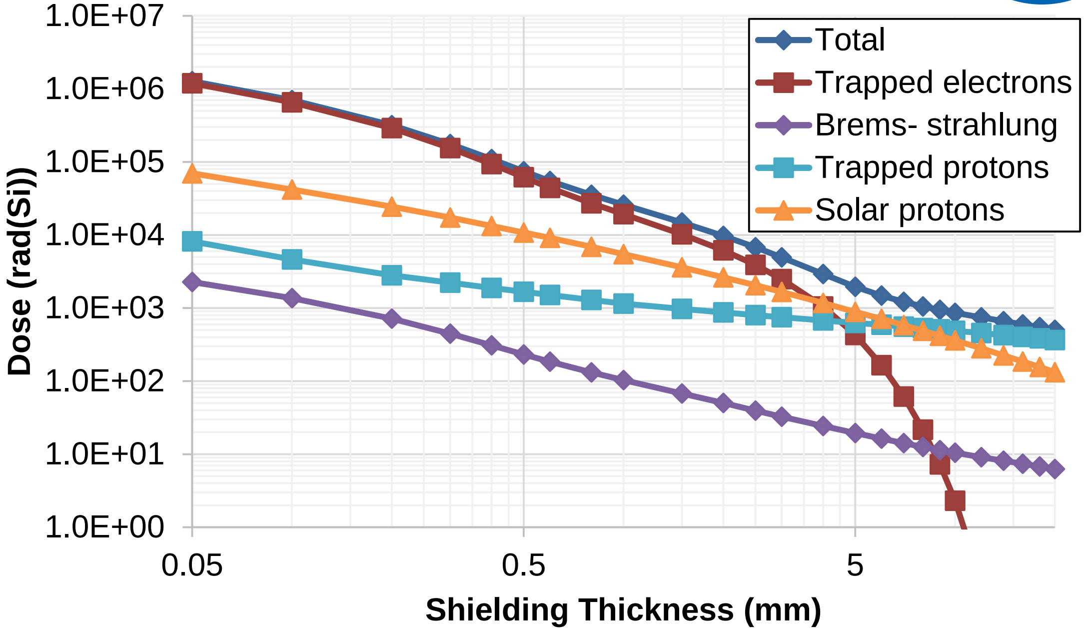
- **Cumulative effects**
  - Depend highly on which contributors and duration in their presence
  - Mimic wear-out/aging
  - TNID and TID must be accounted for

- **Typical destinations (LEO, GEO)**
  - LEO at low altitude/inclination is more protected by the Geomagnetic field
  - Proximity to the poles & SAA show a large variability in dose despite short mission durations
  - Electrons and their braking radiation are the big offender in Geostationary orbits (don't forget about spacecraft charging…)

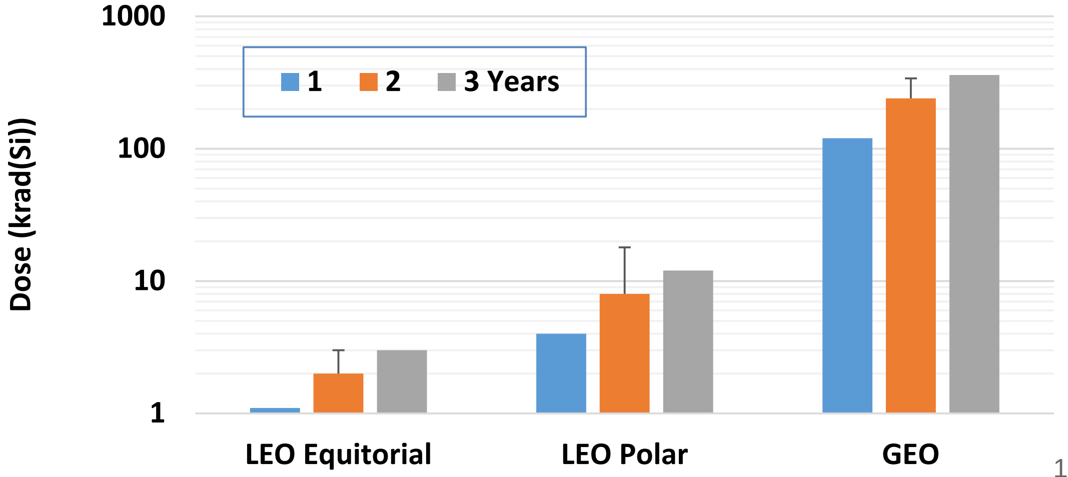
- **Note that**
  - A little bit of shielding goes a long way
  - Altitude plays a huge role when in/near the radiation belts (even transiting)
  - Beyond Geomagnetic field, highly variable solar environment contributions (Solar cycle)

Degradation has a strong dependence on where you go, not just how long you are on orbit

Total Ionizing Dose vs. Shielding



Approximate Dose Behind ~2.5mm Al



Degradation vs. **Single Event Contributors**

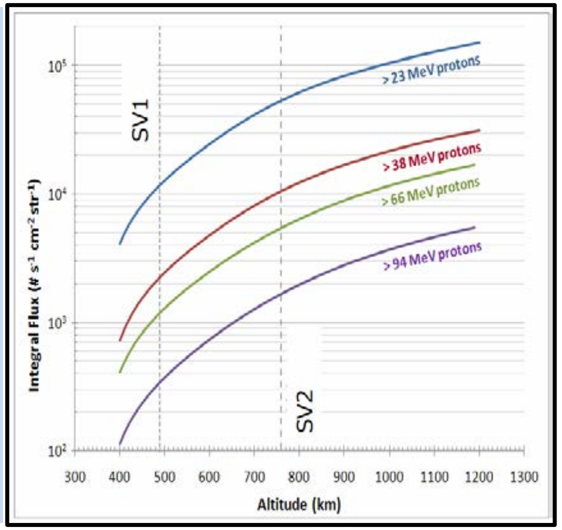
- **One particle causes the effect**
  - Random in nature, particle must traverse sensitive structure within device and have sufficient charge creation along its path
  - Shielding doesn't do so much for highly energetic particles
  - Device technology can be dependent on particle species

- **Typical Destinations (LEO, GEO)**
  - Again altitude plays a role; for some devices that is a direct threat
  - You are exposed to more GCR + Solar contribution as geomagnetic protection is reduced
  - Natural phenomena like the South Atlantic Anomaly (SAA), magnetic poles, are temporal drivers

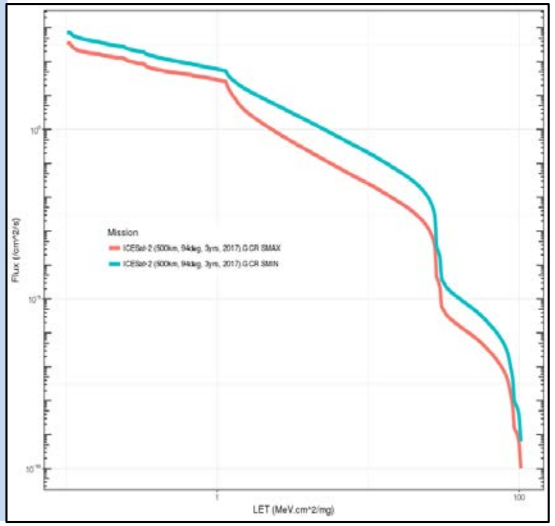
- **Note that**
  - There will be a background rate, solar cycle dependence, solar event rate, increased rate for poles or SAA – **not just one rate to consider**

Single event contributors benefit very little from shielding, have dependence on where you are

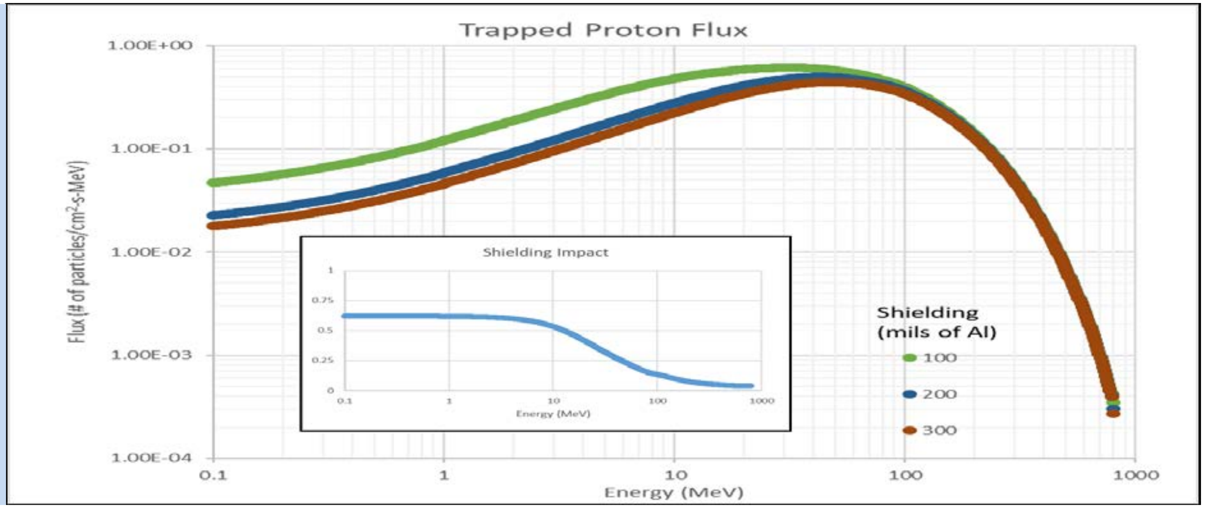
Altitude and Orbit



GCR – Solar Cycle



Shielding Effectiveness



Summary of Environmental Hazards

	Plasma (charging)	Trapped Protons	Trapped Electrons	Solar Particles	Cosmic Rays	Human Presence	Long Lifetime (>10 years)	Nuclear Exposure	Repeated Launch	Extreme Temperature	Planetary Contaminates (Dust, etc)
GEO	Yes	No	Severe	Yes	Yes	No	Yes	No	No	No	No
LEO (low-incl)	No	Yes	Moderate	No	No	No	Not usual	No	No	No	No
LEO Polar	No	Yes	Moderate	Yes	Yes	No	Not usual	No	No	No	No
International Space Station	No	Yes	Moderate	Yes - partial	Minimal	Yes	Yes	No	Yes	No	No
Interplanetary	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	Yes	Yes	No	Yes	Maybe	No	Yes	Maybe
Exploration – Lunar, Mars, Jupiter	Phasing orbits	During phasing orbits	During phasing orbits	Yes	Yes	Possibly	Yes	Maybe	No	Yes	Yes

https://radhome.gsfc.nasa.gov/radhome/papers/SSPVSE05_LaBel.pdf

Radiation Hazard Contributors for Dose and SEE

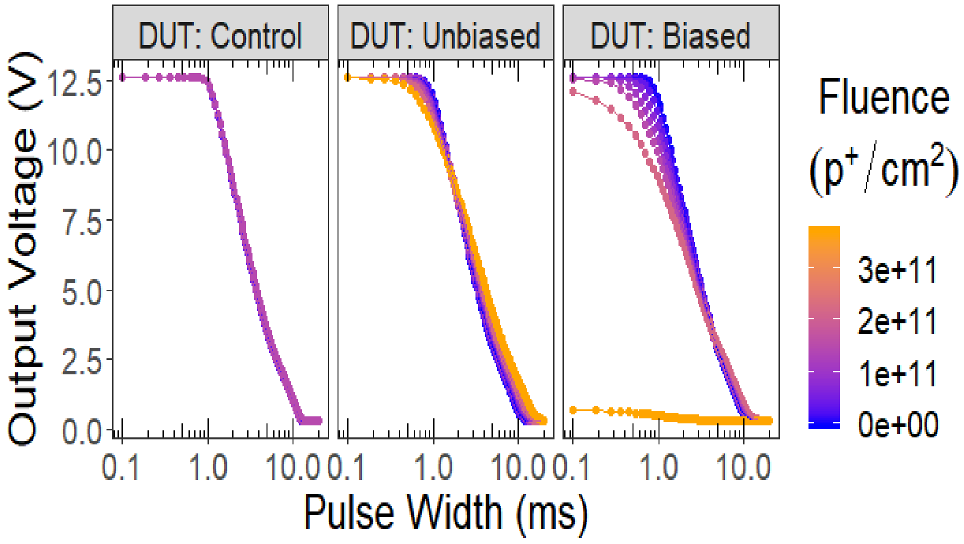
Environment

		LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
Mission Lifetime	> 3 Years	Moderate Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	High Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
	1- 3 Years	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
	< 1 Year	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	Moderate Dose / High GCR, High Solar Proton Variability

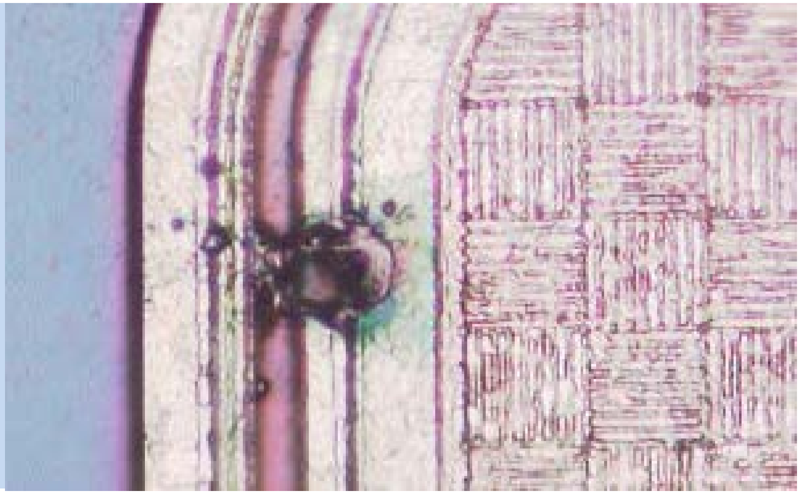
Radiation Effects on Active Microelectronic Devices

- **Cumulative effects and single event effects can both be permanently damaging**
  - TID/DDD lead to wear-out of device operation and degrade devices beyond acceptable operations internally and externally
  - Single Event Effects can be catastrophic instantaneously by turning on parasitic devices within the semiconductor or inducing electric field across dielectrics that eventually break down
  - Synergistic effects can make ground based testing very difficult
- **Destructive Single Event Effects (SEEs)**
  - Irreversible processes
  - Terms: Latchup, Burnout, Gate Rupture
- **Non-Destructive SEEs**
  - Lead to interruptions in operation and/or errors leading to unknown state spaces or loss of science / mission if not accounted for
  - Terms: Functional Interrupt, Transients, Upsets
- **IEEE / Papers / Short Courses / Presentations**
  - GOMAC, HEART, MRQW, NEPP ETW, NSREC, RADECS, SEE/MAPLD, SERESSA, SPWG

Degradation

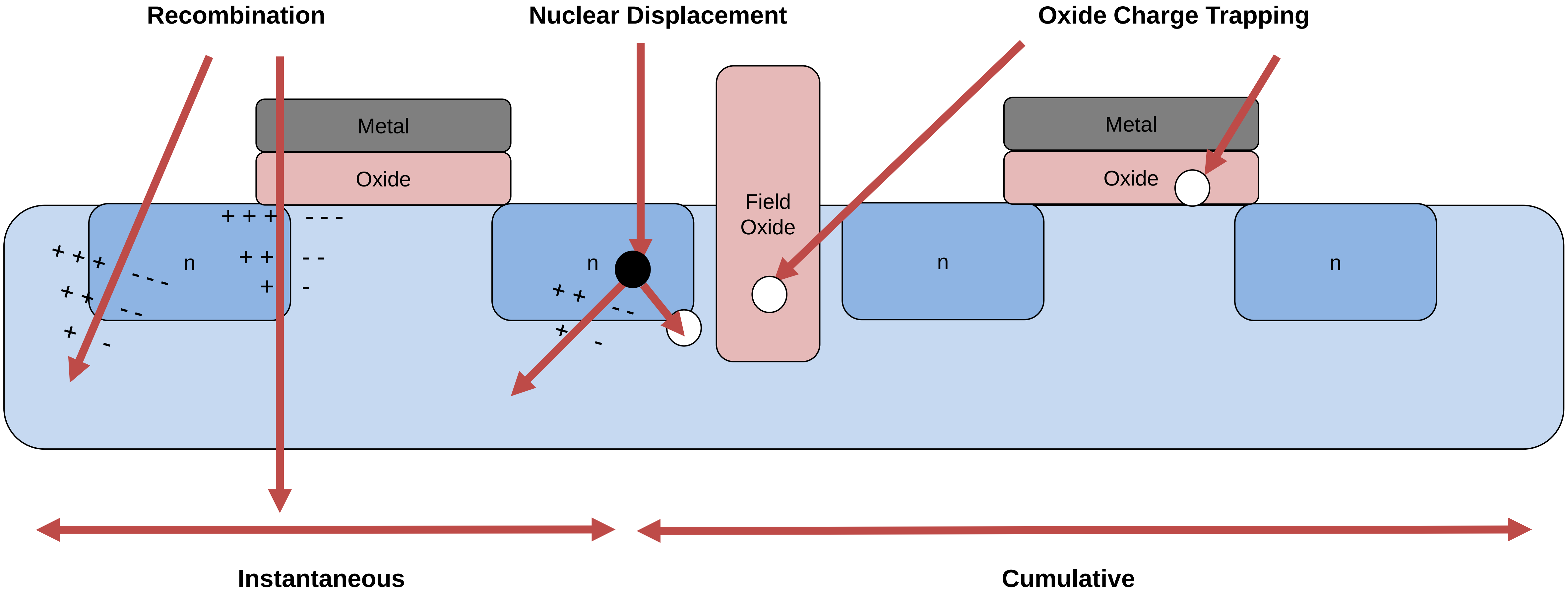


Single Event



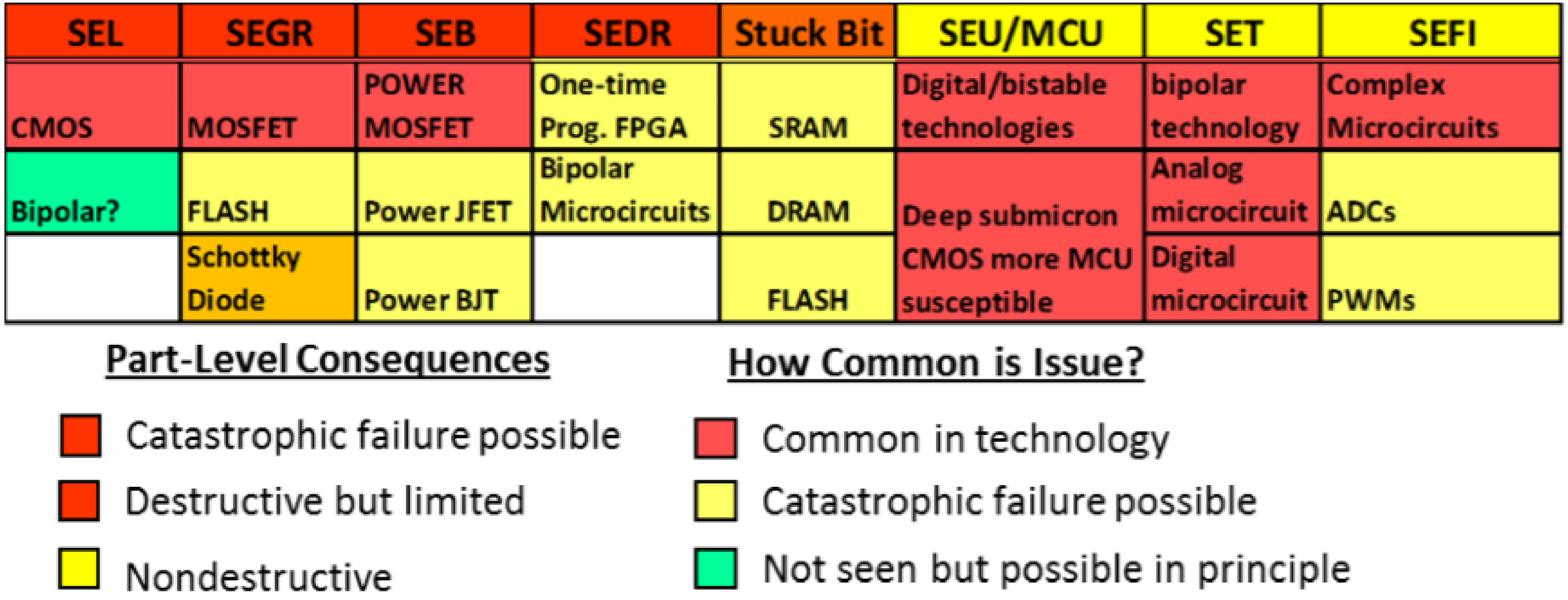
Megan Casey - https://nepp.nasa.gov/files/26196/2014-561-Casey-Final-Web-Pres-ETW-Diodes-TN16278_v2.pdf

Device and Particle Interaction



Brock J. LaMeres, Colin Delaney, Matt Johnson, Connor Julien, Kevin Zack, Ben Cunningham Todd Kaiser, Larry Springer, David Klumpar, "Next on the Pad: RadSat – A Radiation Tolerant Computer System," Proceedings of the 31st Annual AIAA/USU Conference on Small Satellites, Logan UT, USA, Aug. 5-10, 2017, paper: SSC17-III-11,
URL: http://digitalcommons.usu.edu/cgi/viewcontent.cgi?article=3618&context=smallsat

Table of SEE Susceptibility



SEL	SEGR	SEB	SEDR	Stuck Bit	SEU/MCU	SET	SEFI
CMOS	MOSFET	POWER MOSFET	One-time Prog. FPGA	SRAM	Digital/bistable technologies	bipolar technology	Complex Microcircuits
Bipolar?	FLASH	Power JFET	Bipolar Microcircuits	DRAM	Deep submicron CMOS more MCU susceptible	Analog microcircuit	ADCs
	Schottky Diode	Power BJT		FLASH		Digital microcircuit	PWMs

Part-Level Consequences

- 🟥 Catastrophic failure possible
- 🟥 Destructive but limited
- 🟨 Nondestructive

How Common is Issue?

- 🟥 Common in technology
- 🟨 Catastrophic failure possible
- 🟩 Not seen but possible in principle

Ray Ladbury, *https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf*

List is not exhaustive, but new failure modes are found in new devices, so it would not be possible to capture all

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  - **New Technologies**
  - **New Architectures**
  - **Unbound Risks**
- **Building Smart Requirements**
- **Risk Acceptance and Guidance**

The Job: Watch For the 'ilities

Survivability

- Must survive until needed
- Entire mission?
- Screening for early failures in components

Availability

- Must perform when necessary
- Subset of time on orbit
- Operational modes
- Environmental response

Criticality

- Impact to the system
- Part or subsystem function
- Mission objectives

Reliability

- Resultant of all
- Many aspects and disciplines
- Known unknowns

The People: Radiation Effects Engineers

Materials

- Material Property degradations with radiation
- Energy loss in materials

Device Physics

- Charge transport
- Device Process Dependencies
- Charge dependency of device operation

Electrical Engineering

- Part to part interconnections
- Understanding circuit response
- Device functions and taxonomy

Systems Engineering

- Requirements
- System Level Impacts
- Understanding interconnections
- Understanding functionality

Space Physics

- Space weather
- Environment models/modeling
- Radiation Sources and variability

Paths to Space Radiation

Space Radiation Ecosystem

Systems Engineering Background

- Radiation Reqs. Definition
- SPENVIS, OMERE, Fastrad, etc.
- Radiation Testing Management

Device Physics / Electrical Engineering Background

- Radiation Testing + Qualification
- EEE Parts Programs

Space Weather Physics Background

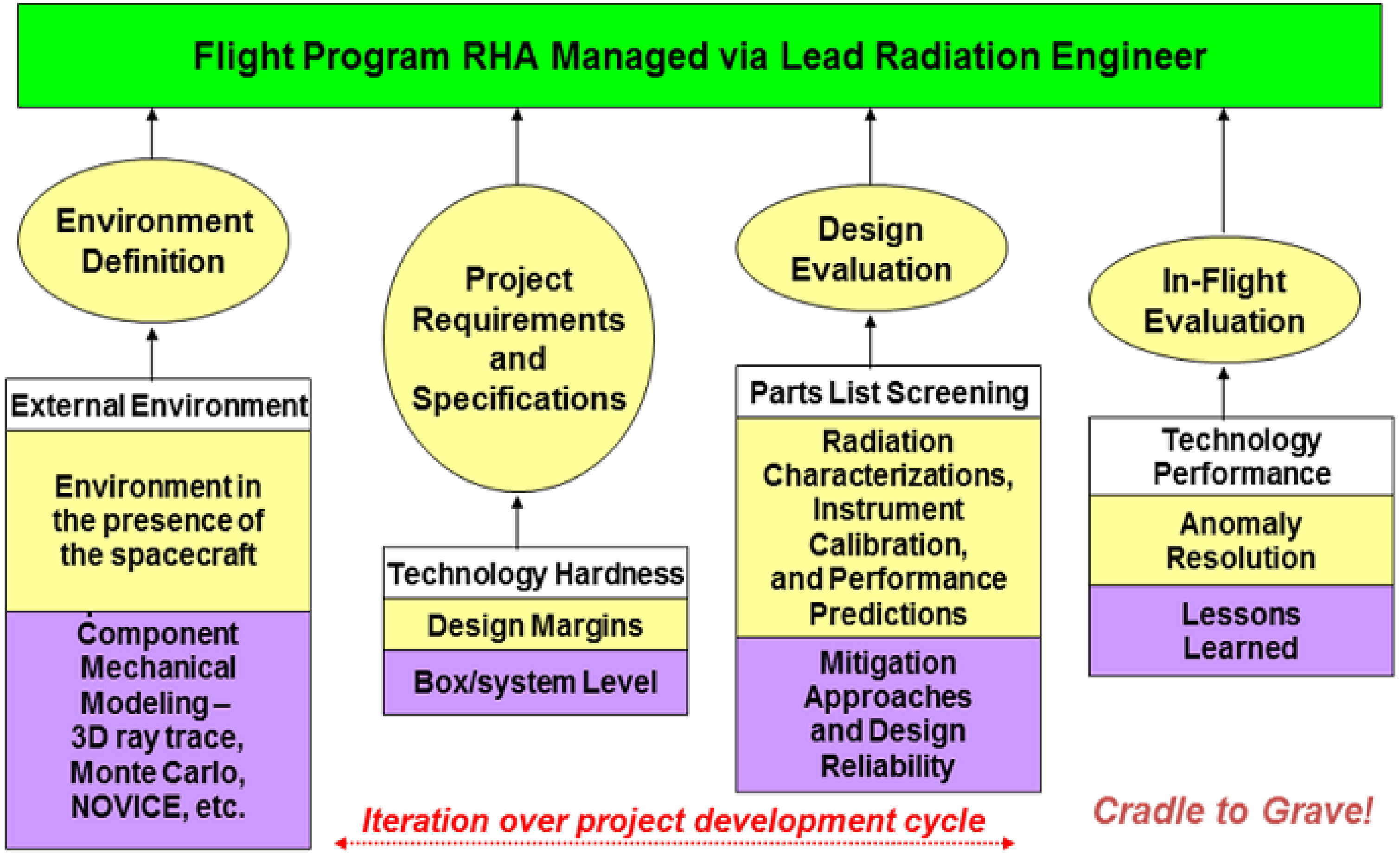
- Mission Scientists / PIs
- Model Developers (e.g. AP9/AE9)
- Often University + Research Lab based

After Whitney Lohmeyer, presented at JPL meeting 2019

Radiation Hardness Assurance (RHA) Overview

RHA consists of all activities undertaken to ensure that the electronics and materials of a space system perform to their *design* specifications throughout exposure to the mission space environment

(After Poivey 2007)



Flight Program RHA Managed via Lead Radiation Engineer

Environment Definition

External Environment

Environment in the presence of the spacecraft

Component Mechanical Modeling – 3D ray trace, Monte Carlo, NOVICE, etc.

Project Requirements and Specifications

Technology Hardness

Design Margins

Box/system Level

Design Evaluation

Parts List Screening

Radiation Characterizations, Instrument Calibration, and Performance Predictions

Mitigation Approaches and Design Reliability

In-Flight Evaluation

Technology Performance

Anomaly Resolution

Lessons Learned

Iteration over project development cycle

Cradle to Grave!

(After LaBel 2004)

To be presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16th, 2019

RHA Challenges…
Not So Small

- **Always in a *dynamic* environment**
- **New Technologies**
  - Device Topology / Speed / Power
  - Increased COTS parts / subsystem usage
- **New Mission Architectures**
  - Profiles of mission life, objective, and cost are evolving
  - Oversight gives way to insight in some mission classifications
  - Ground systems, do no harm, hosted payloads
  - Similarity and heritage data requirements widening
- **Quantifying Risk**
  - Translation of system requirements to radiation trades can be problematic
  - Determining appropriate mitigation level (operational, system, circuit/software, device, material, etc.)

Unbound radiation risks are likely

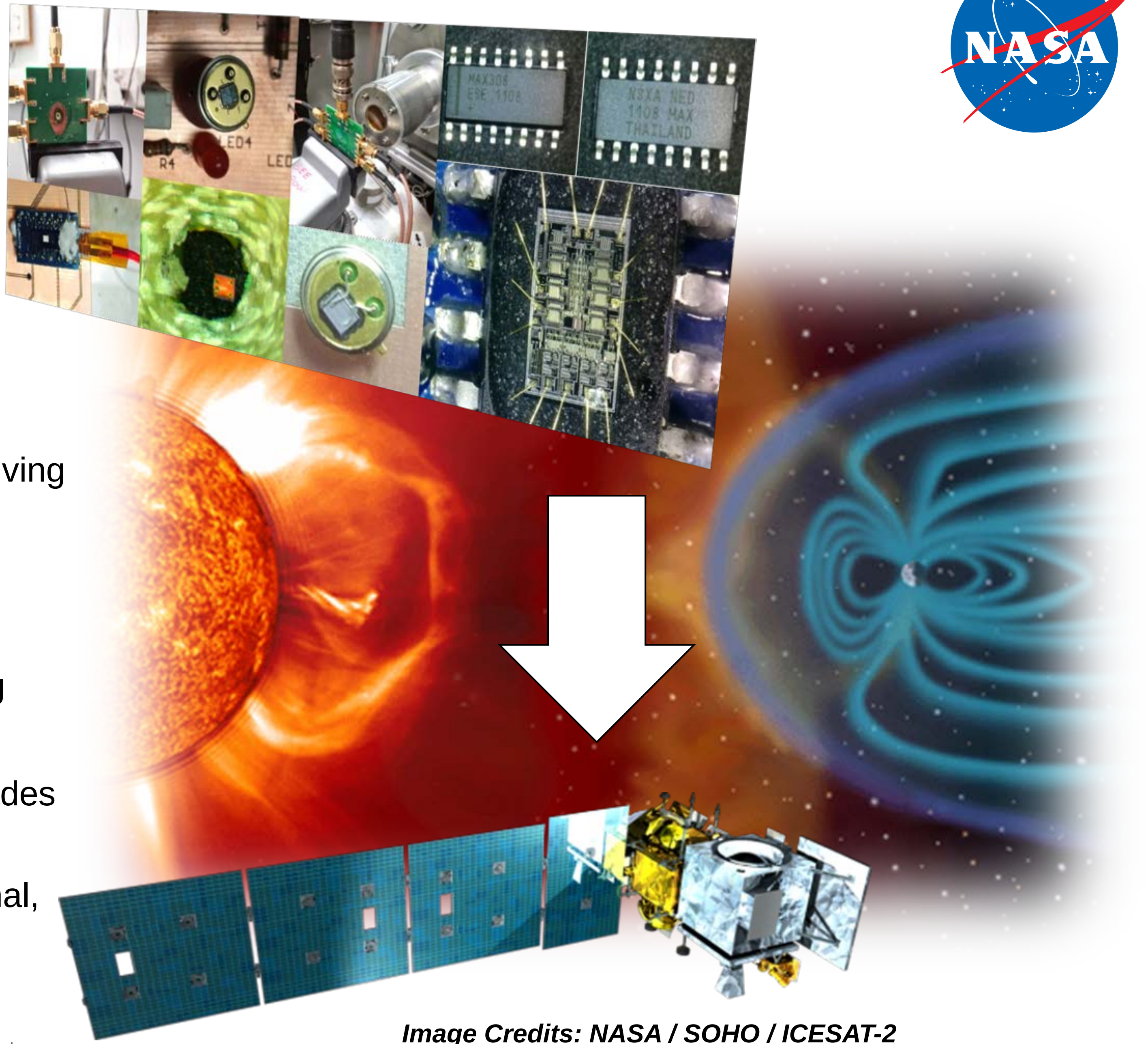


Image Credits: NASA / SOHO / ICESAT-2

New Technologies - New Susceptibilities

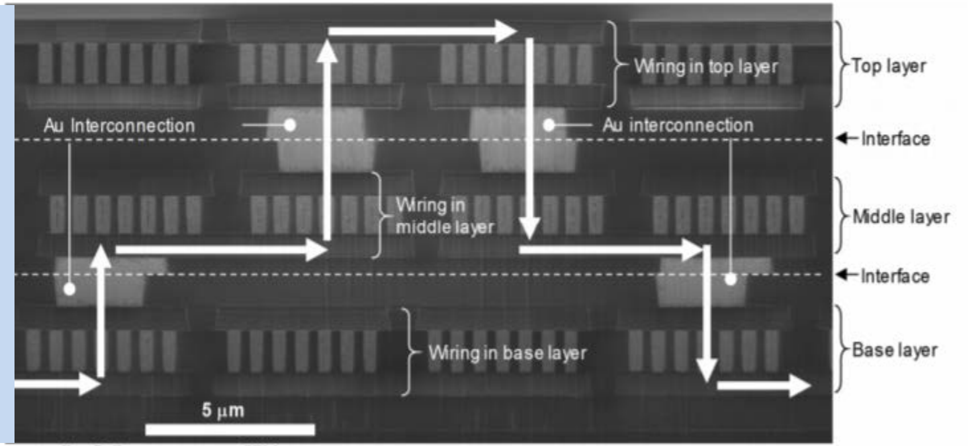
- **Feature Size / Critical Charge**
  - Sensitivity to muons? Low energy protons?
- **3D Stacking / Structures**
  - Deep sensitive volumes
  - New materials within structure
- **Testing Challenges**
  - Complexity (e.g., Systems-on-a-Chip)
    - Speed of interfaces
    - Obfuscation of state-space
  - Flux / range of beam at facilities
- **Function**
  - Integrated Photonics, MEMS, Hybrids

Without detailed part information you do not have certainty of the radiation threats

3D Memory Structures

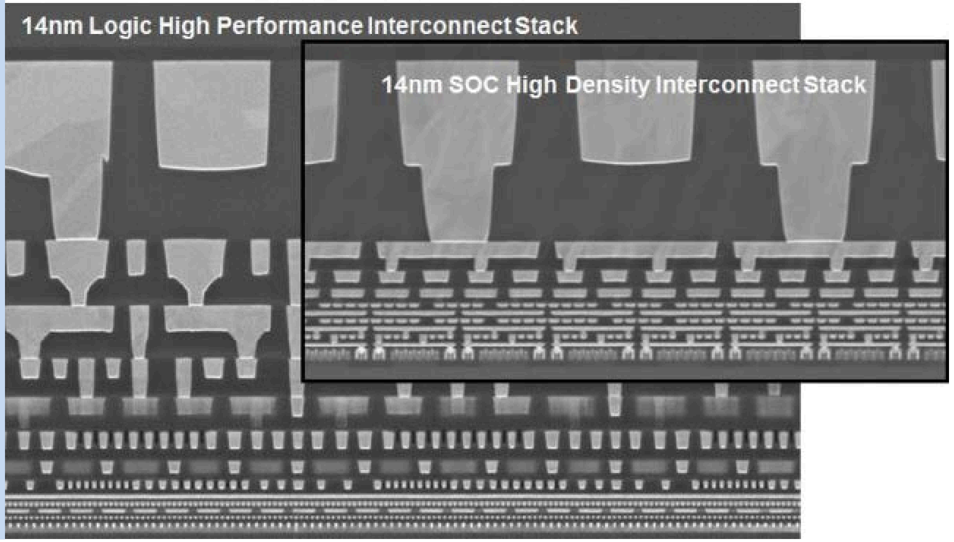
www.micron.com

Gold Interconnects



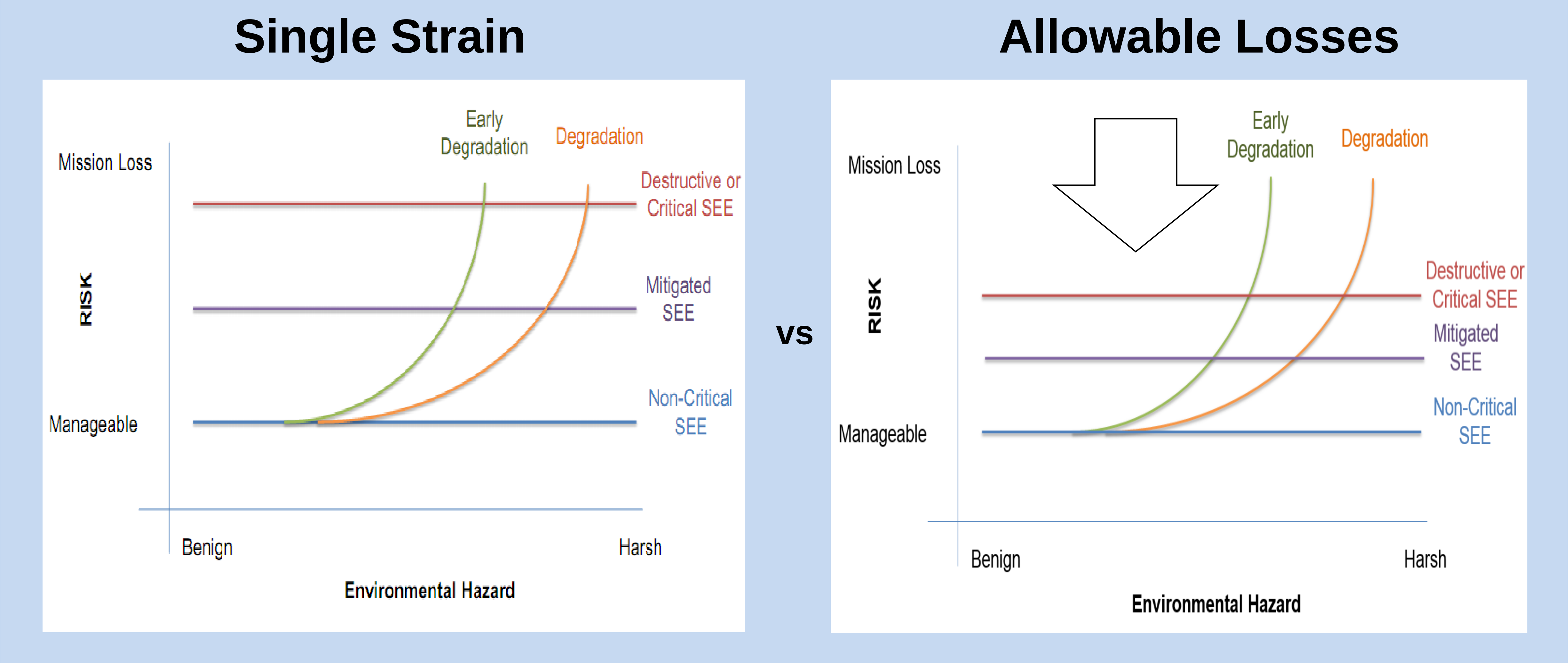
IEEE/DOI: 10.1109/TCPMT.2019.2910863

High Density Stacks



IEEE/DOI:10.1109/IITC-AMC.2016.7507637

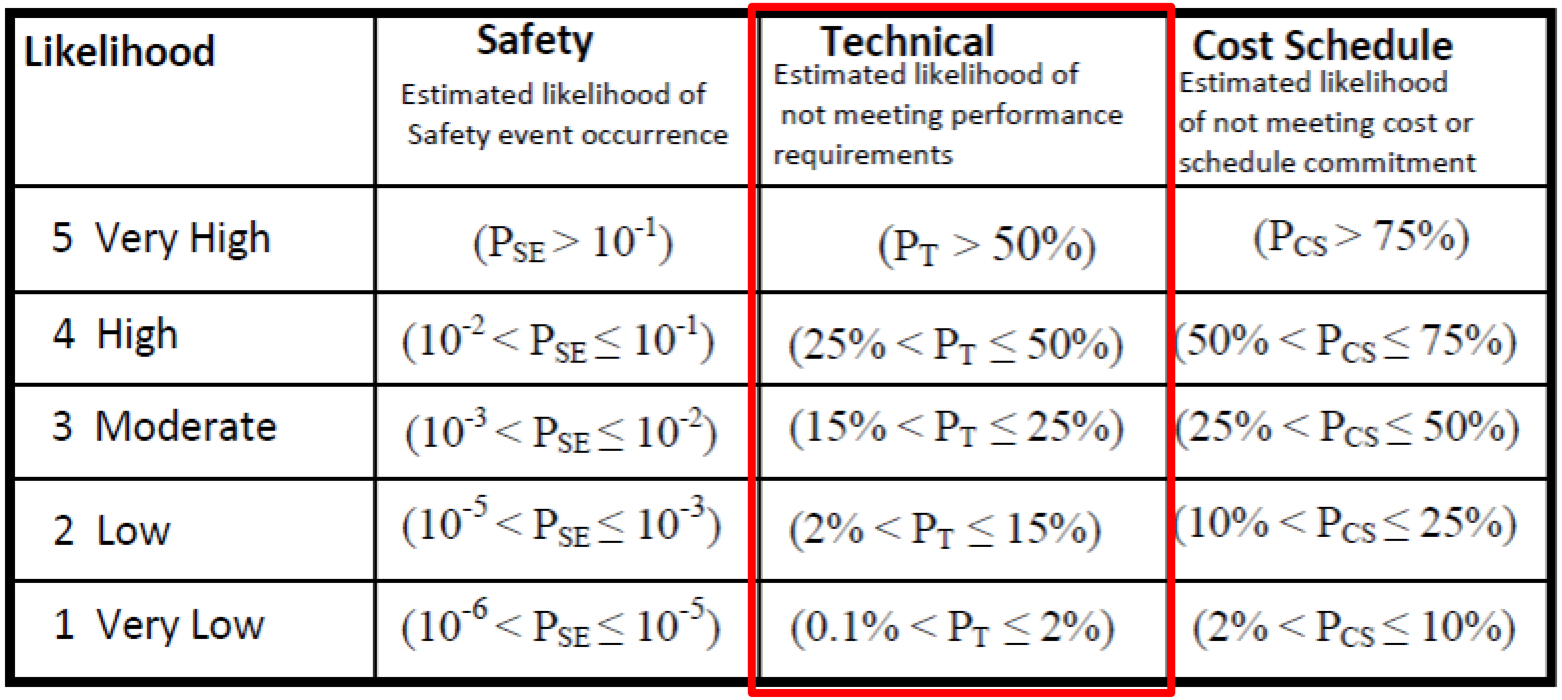
New Mission Architectures - How Many to Succeed?



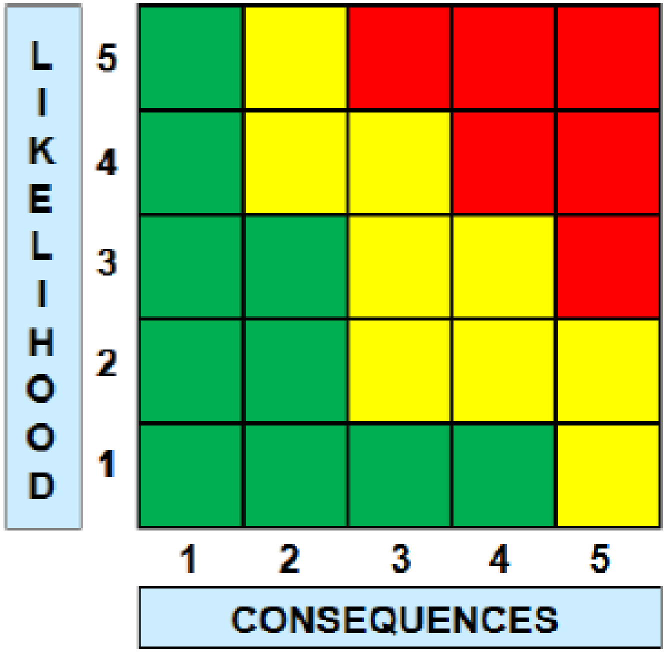
Redundancy alone does not remove the threat, adds complexity

New Challenges in Quantifying Risk

From Risk Assessment section of NASA Program Management 7120.5



Likelihood	Safety Estimated likelihood of Safety event occurrence	Technical Estimated likelihood of not meeting performance requirements	Cost Schedule Estimated likelihood of not meeting cost or schedule commitment
5 Very High	$(P_{SE} > 10^{-1})$	$(P_T > 50\%)$	$(P_{CS} > 75\%)$
4 High	$(10^{-2} < P_{SE} \leq 10^{-1})$	$(25\% < P_T \leq 50\%)$	$(50\% < P_{CS} \leq 75\%)$
3 Moderate	$(10^{-3} < P_{SE} \leq 10^{-2})$	$(15\% < P_T \leq 25\%)$	$(25\% < P_{CS} \leq 50\%)$
2 Low	$(10^{-5} < P_{SE} \leq 10^{-3})$	$(2\% < P_T \leq 15\%)$	$(10\% < P_{CS} \leq 25\%)$
1 Very Low	$(10^{-6} < P_{SE} \leq 10^{-5})$	$(0.1\% < P_T \leq 2\%)$	$(2\% < P_{CS} \leq 10\%)$

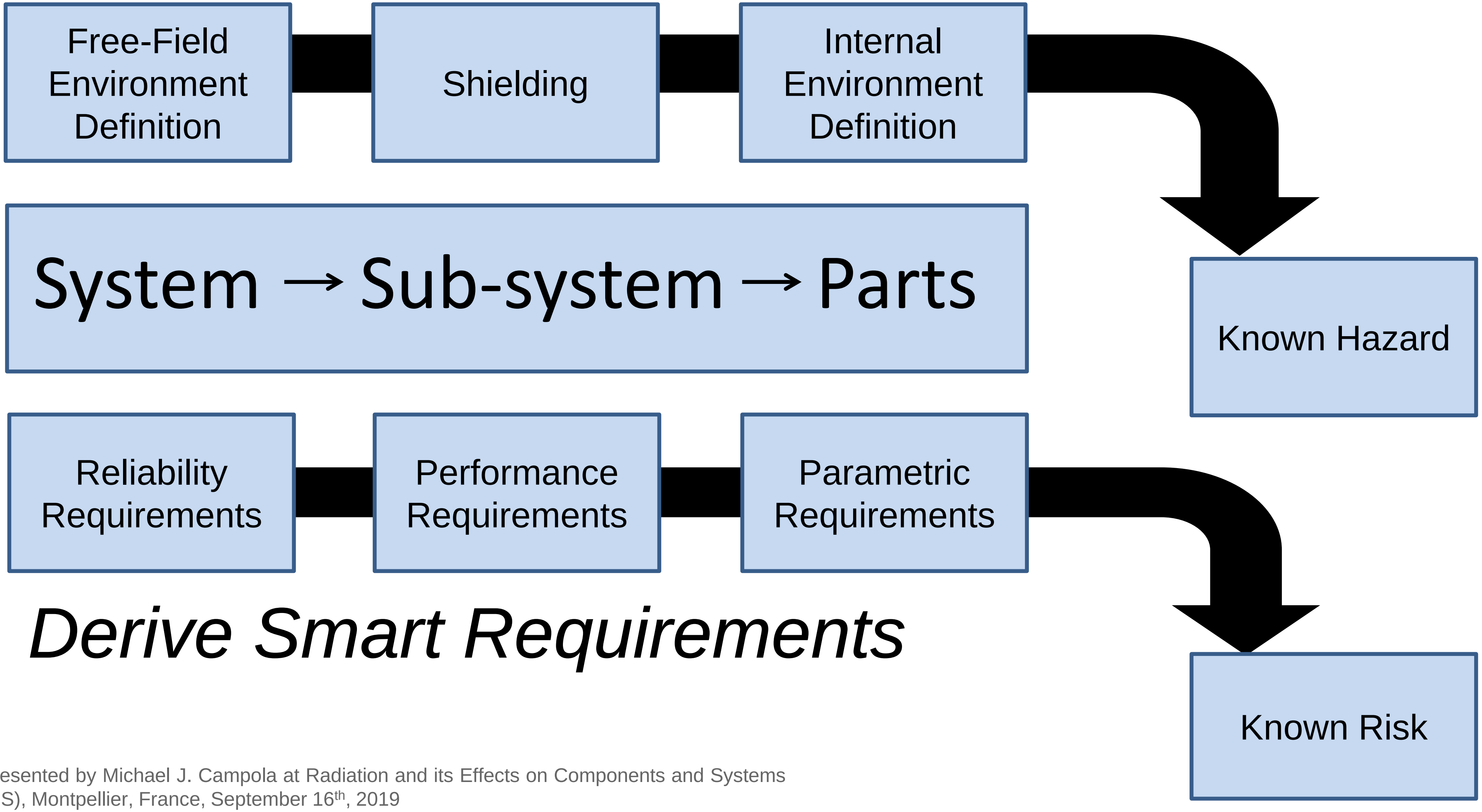


LIKELIHOOD	1	2	3	4	5
5	🟩	🟨	🟥	🟥	🟥
4	🟩	🟨	🟨	🟥	🟥
3	🟩	🟩	🟨	🟨	🟥
2	🟩	🟩	🟨	🟨	🟨
1	🟩	🟩	🟩	🟩	🟨

Can only get there with enough information about the system or the chosen device, need to have a known hazard and a known response

RHA Building Blocks

Define and Evaluate the Hazard



Free-Field Environment Definition

Shielding

Internal Environment Definition

System → Sub-system → Parts

Known Hazard

Reliability Requirements

Performance Requirements

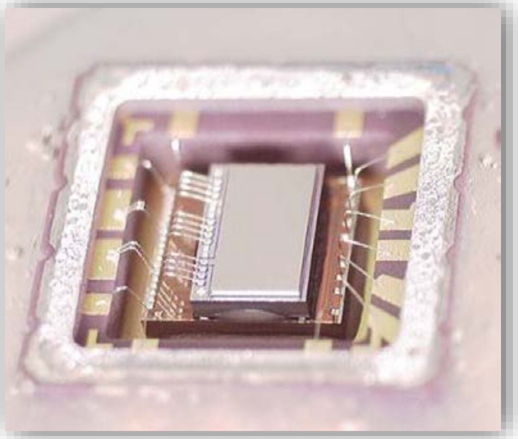
Parametric Requirements

Known Risk

Derive Smart Requirements

Risks Abound, What is Critical?

- **Parts**
    - Parametric degradation and leakage currents allowable in application?
    - Downstream/peripheral circuits considered?
    - Reset/refresh capability?
    - Mitigation within too complex?
    - Predicted radiation response unknown– loss of part functionality critical?
- **Subsystem**
    - Functionally required to mission that the subsystem work?
    - Interfaces allow you to get to a known state if all goes wrong?
- **System**
    - Increased power dissipation a mission ender?
    - Availability outweighed by error circumvention?
    - Data retention through reboots? What if there is science data loss?
    - Communications interruptions overwhelm?
    - Navigation or Attitude determination unable to deal with faults?

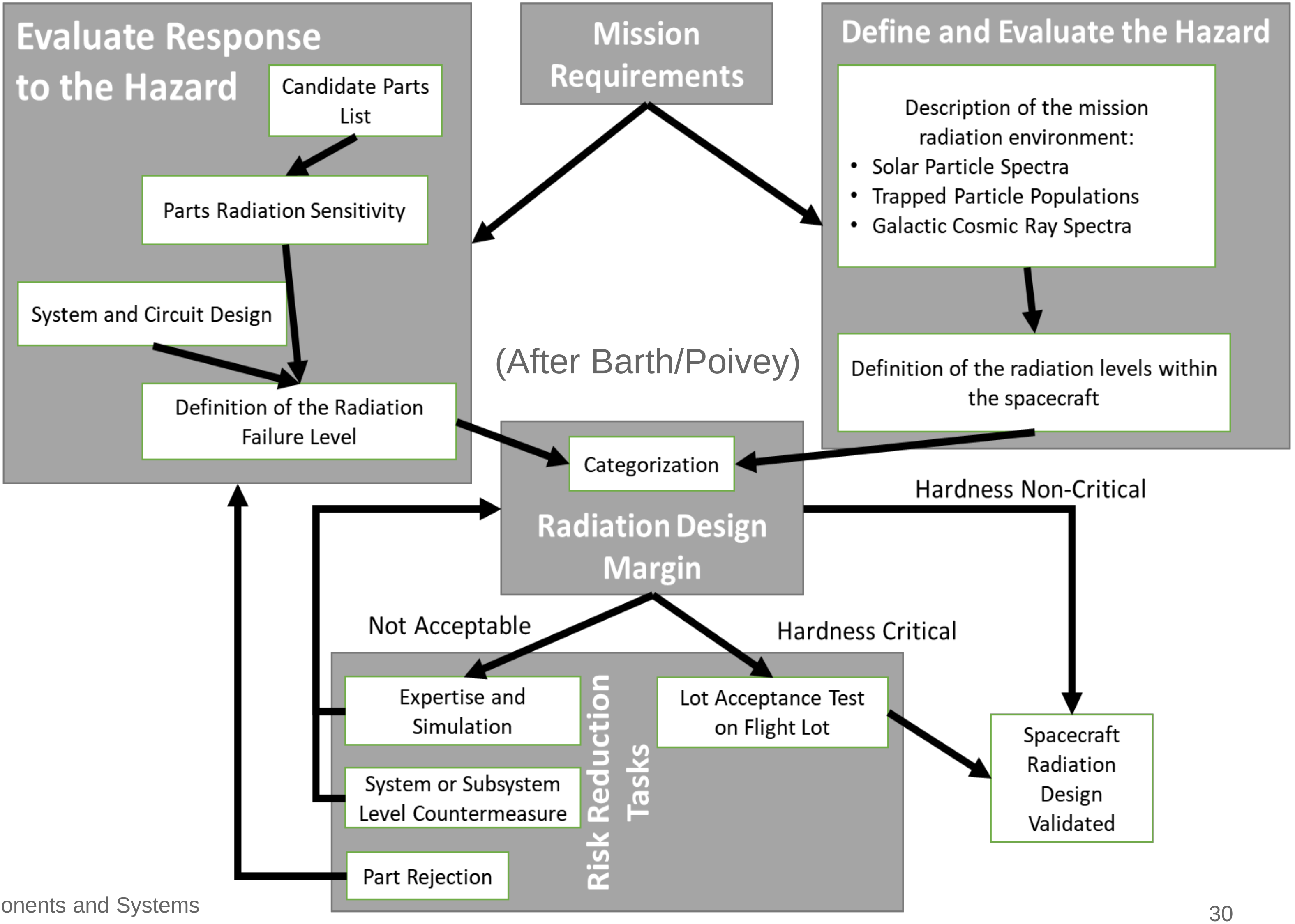


vs.

To be presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16th, 2019

RHA Flow Doesn't Change With Accepted Risk

- **Hardness Assurance is the practice of designing for radiation effects**
- **What it takes to overcome the radiation challenges**
- **Competing failure modes**



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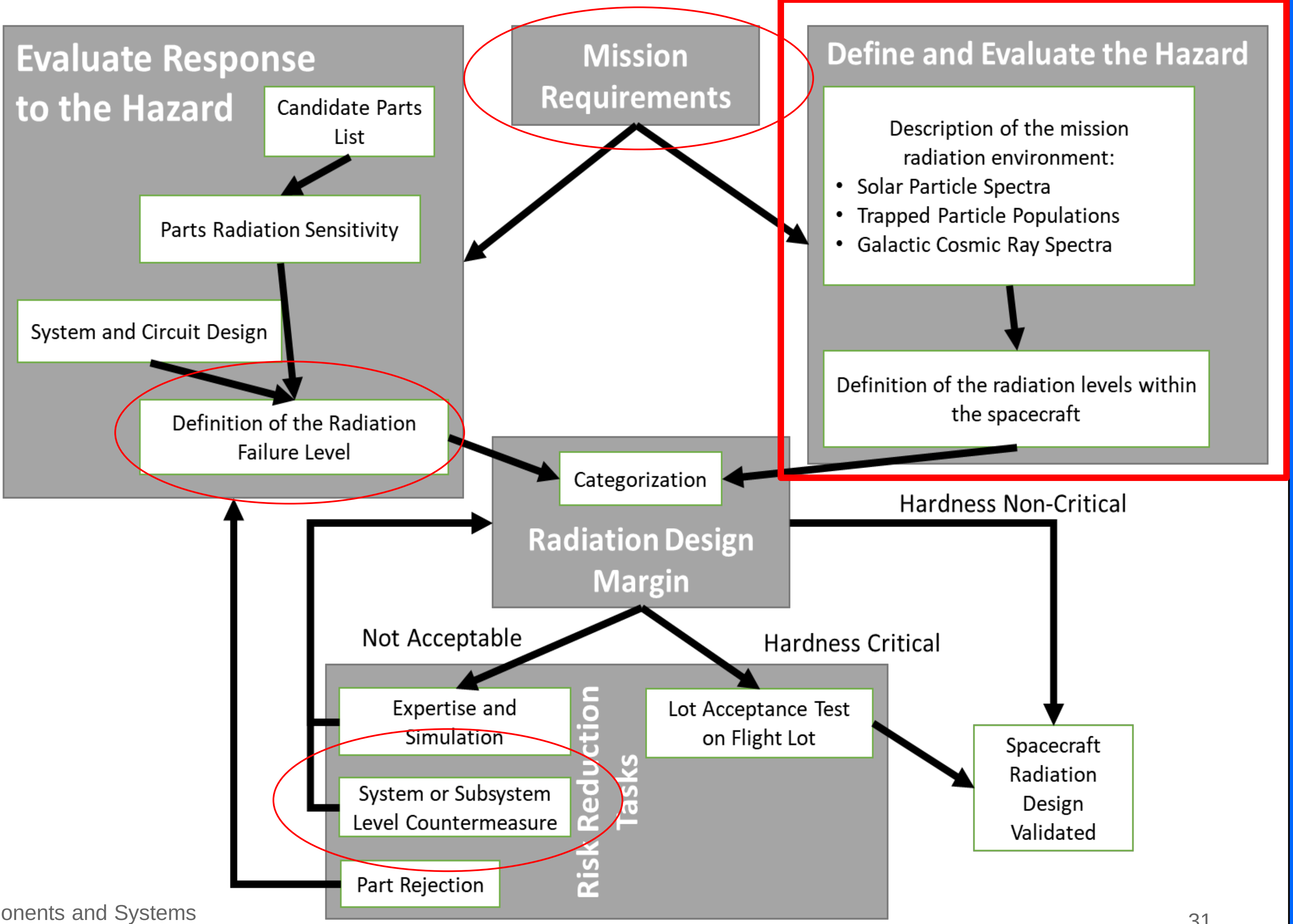
RHA Flow Doesn’t Change With Accepted Risk

- **Hardness Assurance is the practice of designing for radiation effects**
- **What it takes to overcome the radiation challenges**
- **Competing failure modes**
- **Focus for impact on risk acceptance:**
  - Failure Awareness
  - Countermeasures/Mitigation
  - Mission Requirements

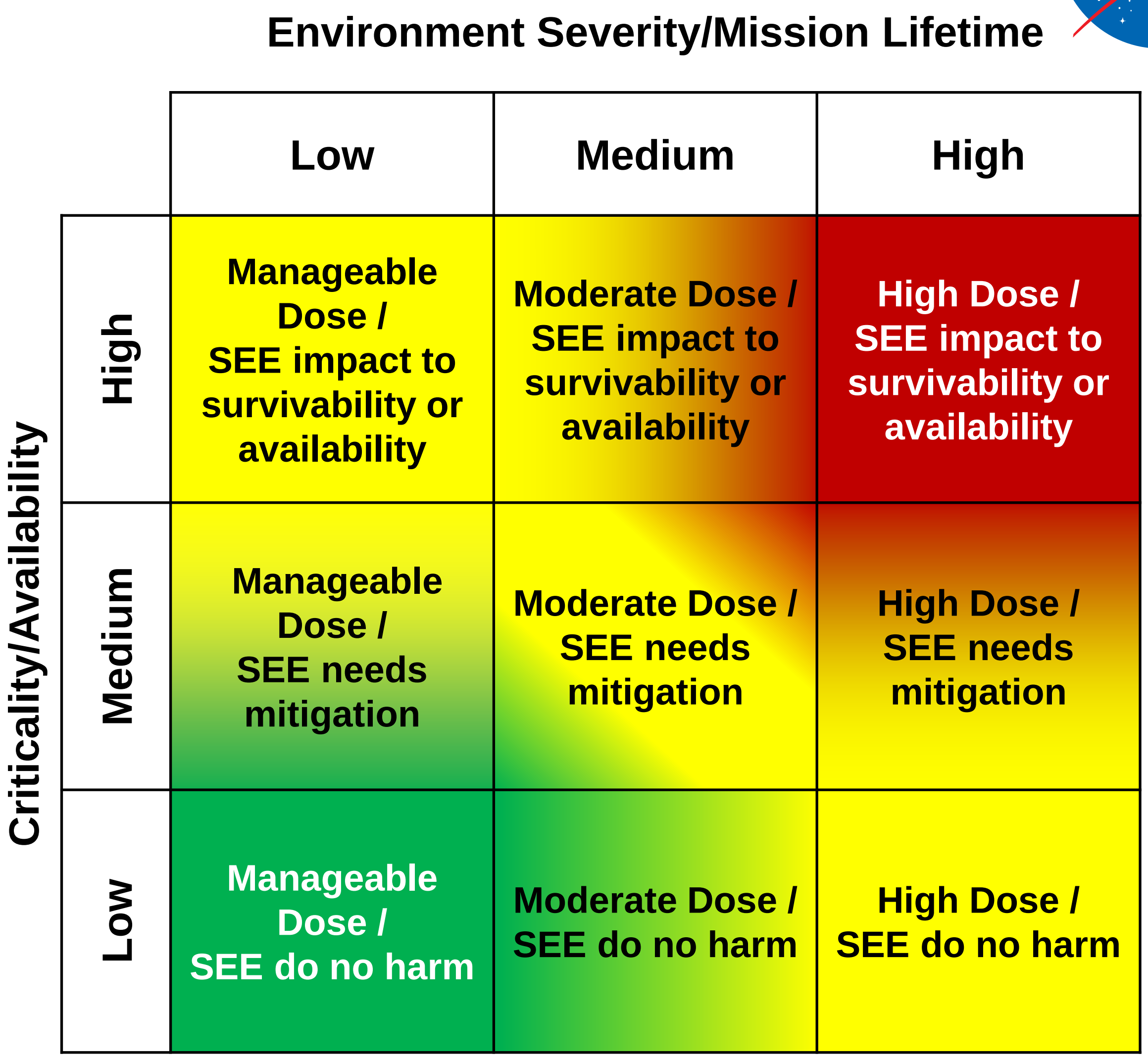


Focus For Risk Acceptance

- **Failure Awareness**
  - Know your hazard from the natural environment
  - Know your devices potential failure mechanisms or response (data)
- **Countermeasures and Mitigation**
  - Where are they necessary?
  - At what level (part, card, box, mission)
- **Smart Requirements – and Eventually Smart Trades**

Define and Evaluate the Hazard

- **Define the Environment**
  - External to the spacecraft
- **Evaluate the Environment**
  - Internal to the spacecraft
- **Define the Requirements**
  - Define criticality factors
- **Evaluate Design/Components**
  - Existing data/Testing
  - Performance characteristics
- **“Engineer” with Designers**
  - Parts replacement/Mitigation schemes
- **Iterate Process**
  - Review parts list based on updated knowledge



Environment Severity/Mission Lifetime

Criticality/Availability		Low	Medium	High
	High	**Manageable Dose / SEE impact to survivability or availability**	**Moderate Dose / SEE impact to survivability or availability**	**High Dose / SEE impact to survivability or availability**
	Medium	**Manageable Dose / SEE needs mitigation**	**Moderate Dose / SEE needs mitigation**	**High Dose / SEE needs mitigation**
	Low	**Manageable Dose / SEE do no harm**	**Moderate Dose / SEE do no harm**	**High Dose / SEE do no harm**

Derive Smart Requirements

- **Define the Environment**
  - External to the spacecraft
- **Evaluate the Environment**
  - Internal to the spacecraft
- **Define the Requirements**
  - Define criticality factors
- **Evaluate Design/Components**
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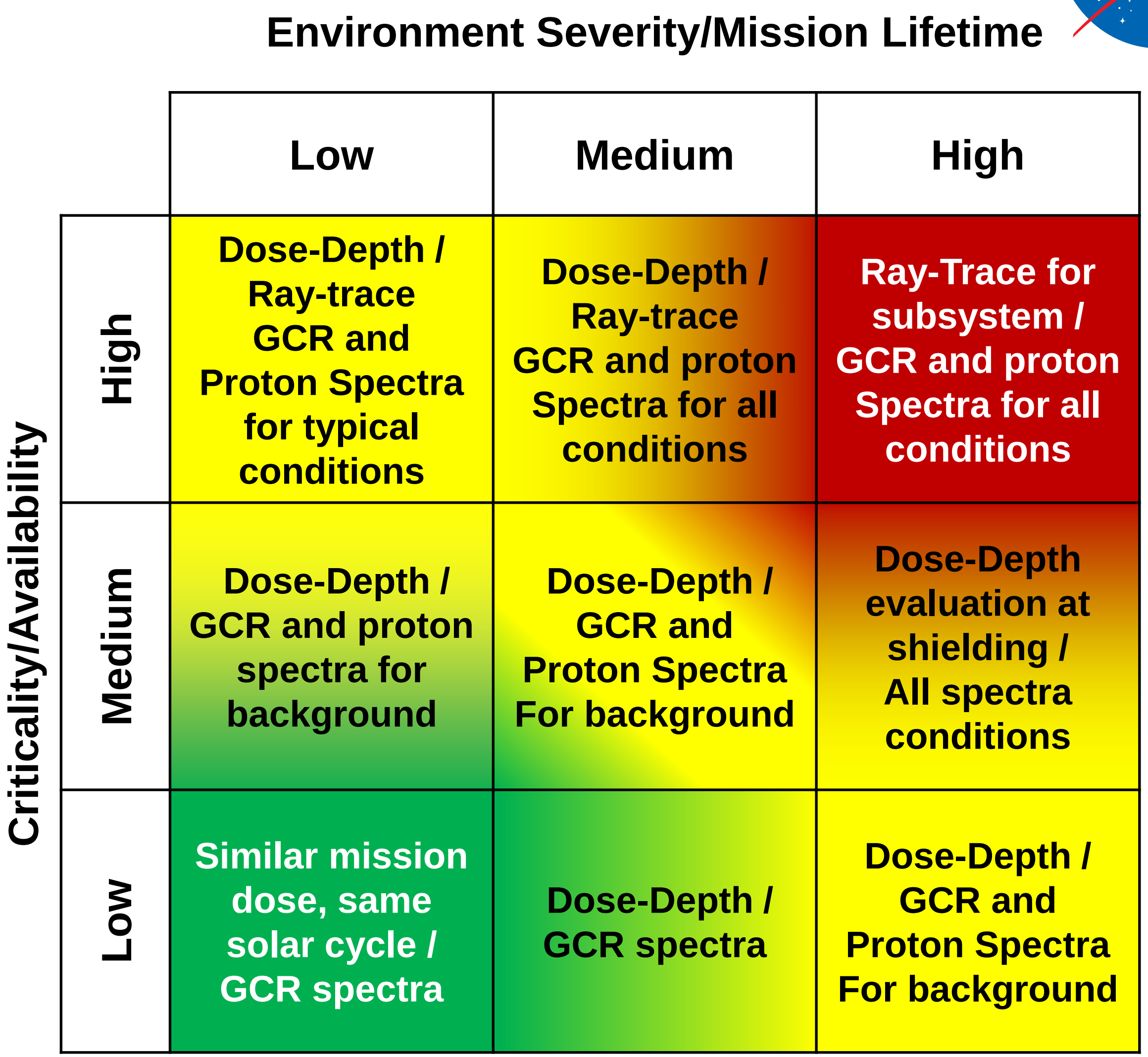


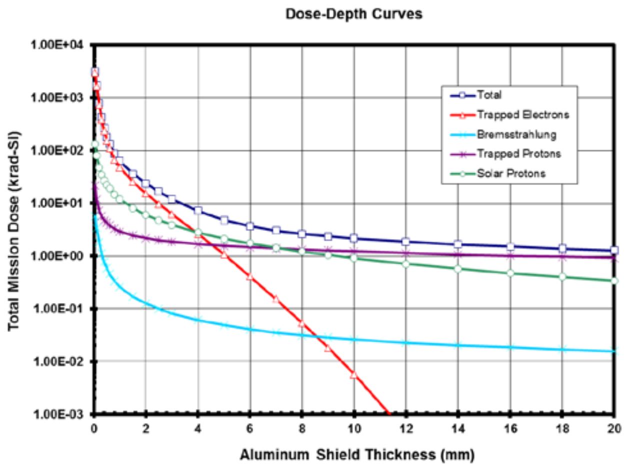
Environment Severity/Mission Lifetime

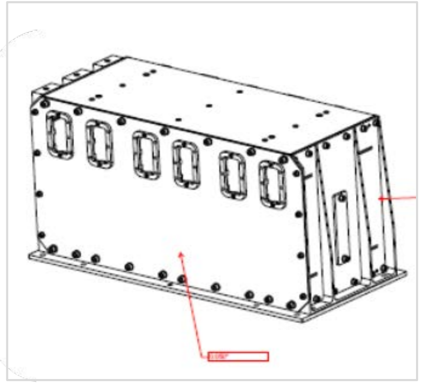
Criticality/Availability	**Low**	**Medium**	**High**
High	**Dose-Depth / Ray-trace GCR and Proton Spectra for typical conditions**	**Dose-Depth / Ray-trace GCR and proton Spectra for all conditions**	**Ray-Trace for subsystem / GCR and proton Spectra for all conditions**
Medium	**Dose-Depth / GCR and proton spectra for background**	**Dose-Depth / GCR and Proton Spectra For background**	**Dose-Depth evaluation at shielding / All spectra conditions**
Low	**Similar mission dose, same solar cycle / GCR spectra**	**Dose-Depth / GCR spectra**	**Dose-Depth / GCR and Proton Spectra For background**

Mitigation and Countermeasure Optimization

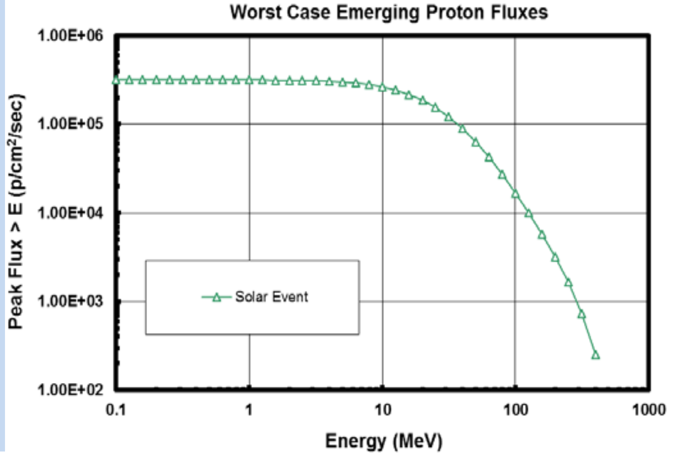
- **Define the Environment**
  - External to the spacecraft
- **Evaluate the Environment**
  - Internal to the spacecraft
- **Define the Requirements**
  - Define criticality factors
- **Evaluate Design/Components**
  - Existing data/Testing
  - Performance characteristics
- **“Engineer” with Designers**
  - Parts replacement/Mitigation schemes
- **Iterate Process**
  - Review parts list based on updated knowledge

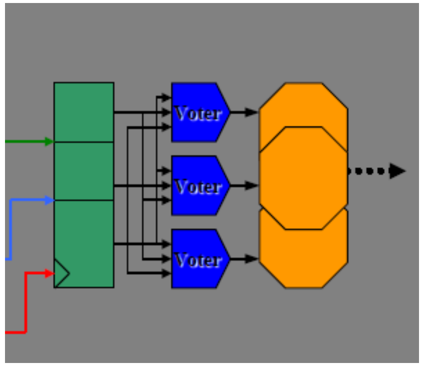
Degradation





Single Event





K.A. LaBel, A.H. Johnston, J.L. Barth, R.A. Reed, C.E. Barnes, “Emerging Radiation Hardness Assurance (RHA) issues: A NASA approach for space flight programs,” IEEE Trans. Nucl. Sci., pp. 2727-2736, Dec. 1998.

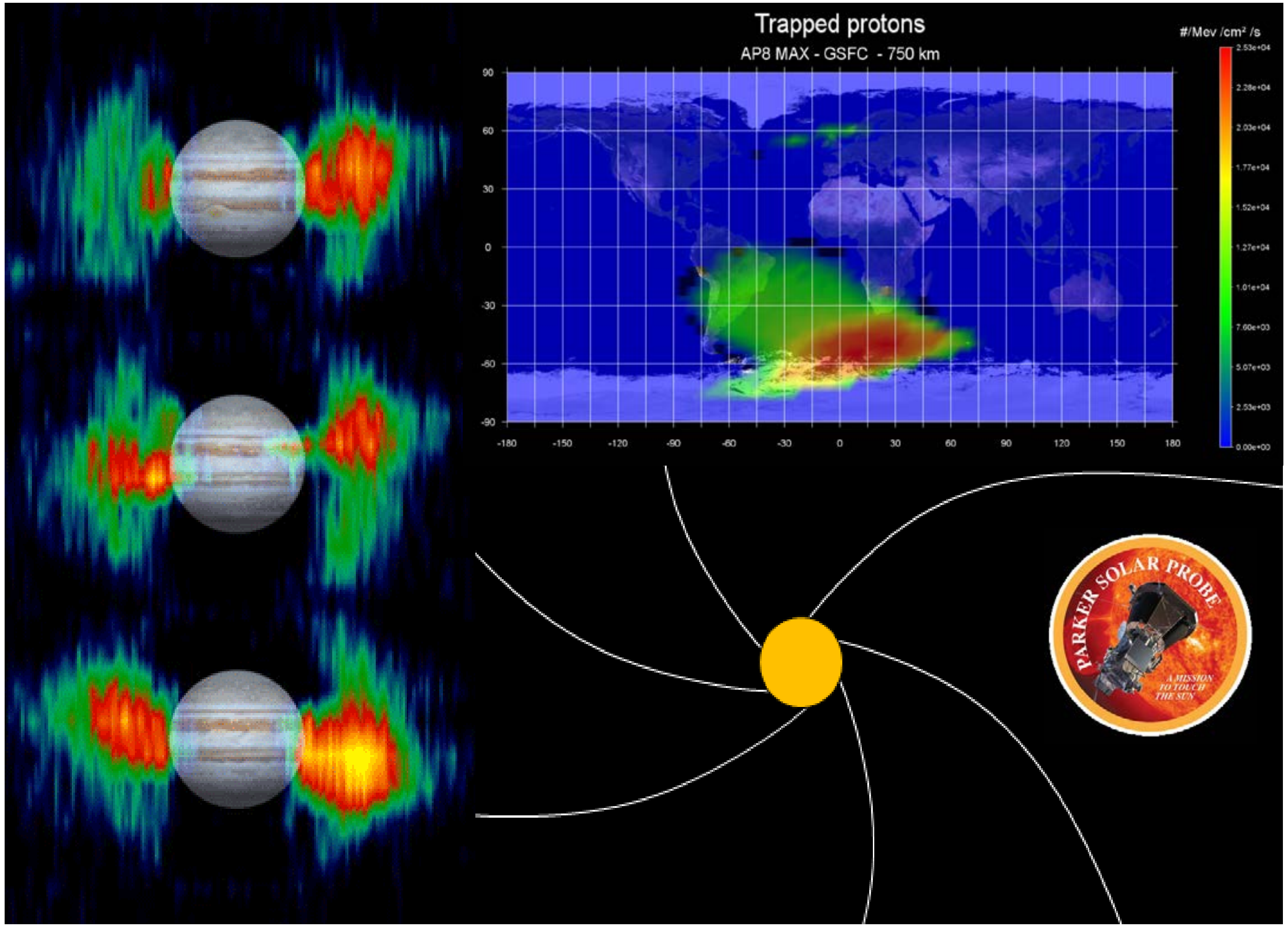
Building Requirements

- **Requirements by Environment**
- **Requirements by Technology**
- **Additional Considerations**
    - **LET Requirements for SEE**
    - **Dose Calculation**
    - **Operation During Flare Conditions**
    - **Radiation Data**

Requirements by Environment

- **Trapped Radiation Belts**
  - Can lead to high doses in a short mission: Jovian
  - Can lead to spatially dependent SEE responses: South Atlantic Anomaly (SAA)
- **Heliocentric Orbits**
  - Solar Events, highly dynamic, energetic, directional
  - Solar Wind, will depend on the solar cycle
  - No planetary magnetic field attenuation

In essence the requirements are always driven by the environment, some more than others create a unique challenge



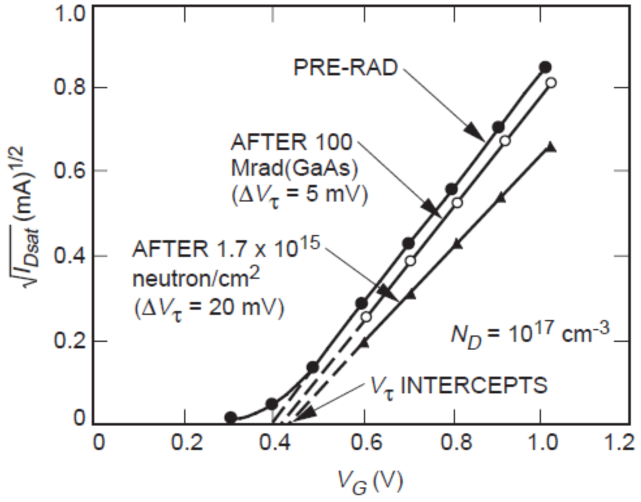
NASA JPL Cassini, http://saturn.jpl.nasa.gov,
Output from OMERE freeware http://www.trad.fr/en/space/omere-software/

Requirements by Technology

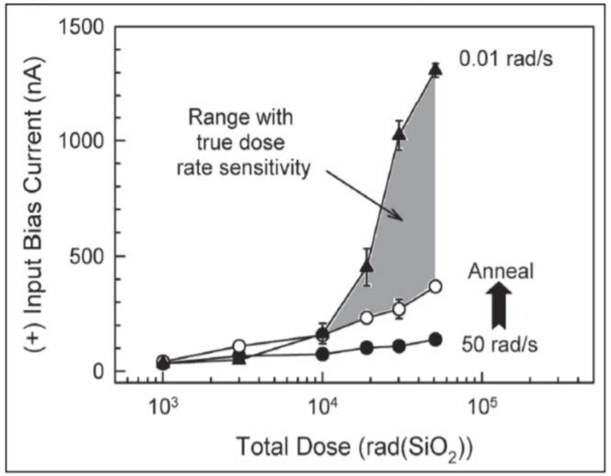
- **Technologies exhibit specific physics of failure**
  - *Not easy to group them all*
  - Opto-electronics - Displacement in the material
  - Bipolar - Enhanced Low Dose Rate Sensitivity
  - Digital CMOS - Latchup or SEFI
  - Power devices - SEGR/SEB
  - Analog/Mixed-Signal – Interruptions on PLLs, SERDES, clock dividers, etc.
- **Test Data requirements**
  - Failure distributions, often not enough parts
  - Destructive effects are one data point, variability from part to part
  - Statistics of the fit for rate calculations

Requirements should only be made applicable to the technologies that need to meet mission objectives and can benefit

R. Zuleeg, “Radiation Effects In GaAs FET Devices,” *Proc. IEEE*, Vol. 77, p.389, 1989.

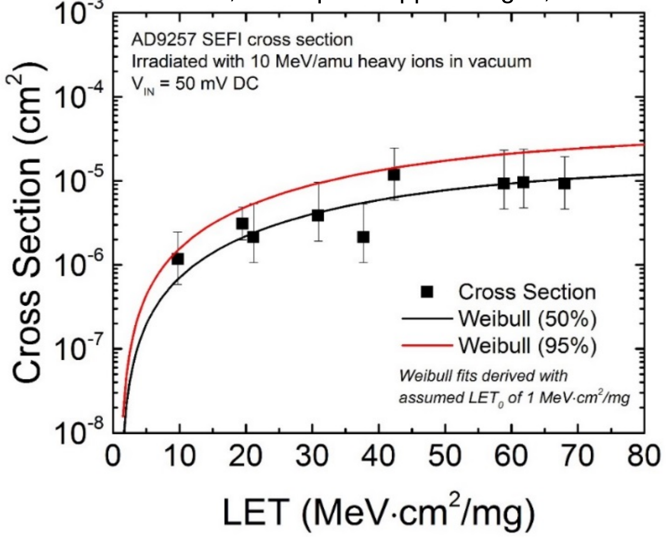


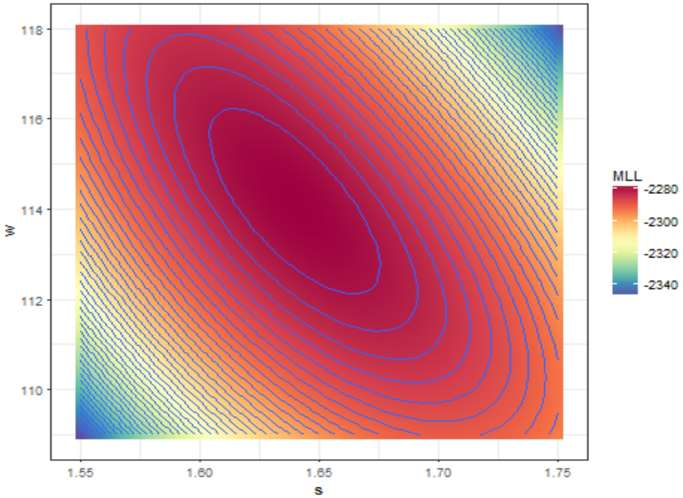
I_{B+} *vs.* Total Dose for LM111 Voltage Comparators



M. R. Shaneyfelt, et al., *IEEE TNS*, 2000.

D. Chen, test report nepp.nasa.gov, 2016

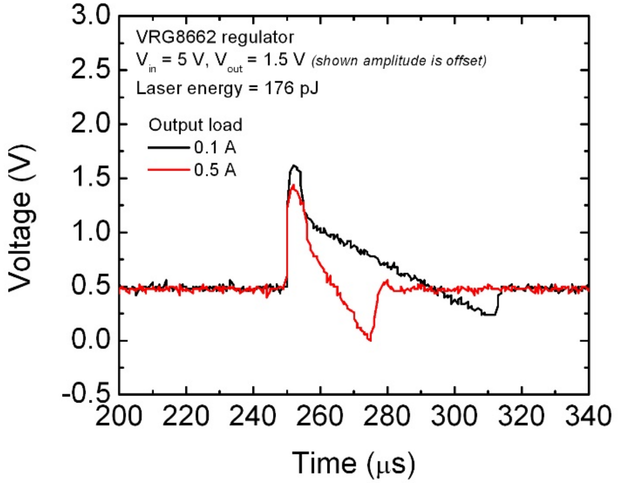


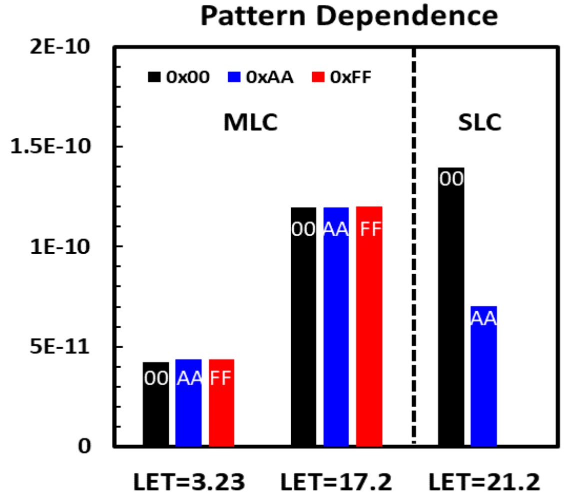


Considerations for SEE Requirements

- **SEL**
  - Environment and technology driven, risk avoidance
  - Protection circuitry / diode deratings
- **SEGR, SEB**
  - Effect driven, normally incident is usually the worst case
  - Testing to establish Safe Operating Area (SOA)
- **SET**
  - Don’t harm downstream parts via overvoltage/overstress on I/O, or accumulate over integrations
  - Can be internal - hybrids
- **SEU**
  - Tailored Filtering, EDAC, or Scrubbing
- **MBU, MCU, SEFI, Locked States**
  - Application Voltage or Pattern dependence
  - Watchdogs / reset capability
- **Proton SEE susceptible parts need evaluated in detail:**
  - Low-energy proton effects:
  - May have direct ionization
  - RHA for proton sensitivity update coming:
  - https://nepp.nasa.gov/files/25401/Proton_RHAGuide_NASAAug09.pdf

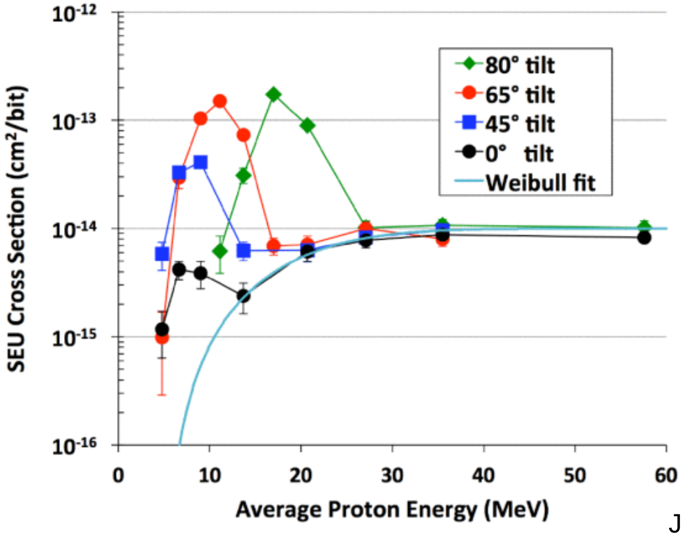
D. Chen, test report radhome.nasa.gov, 2015

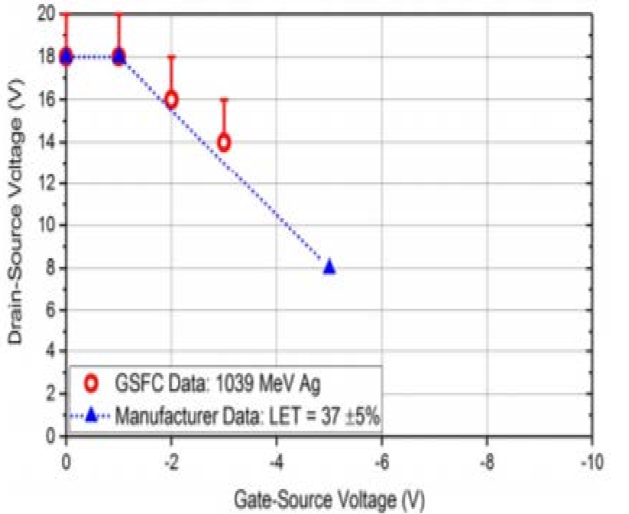




T. Wilcox, NSREC Poster DW, 2019

N. A. Dodds *et al.*, doi: 10.1109/TNS.2015.2486763



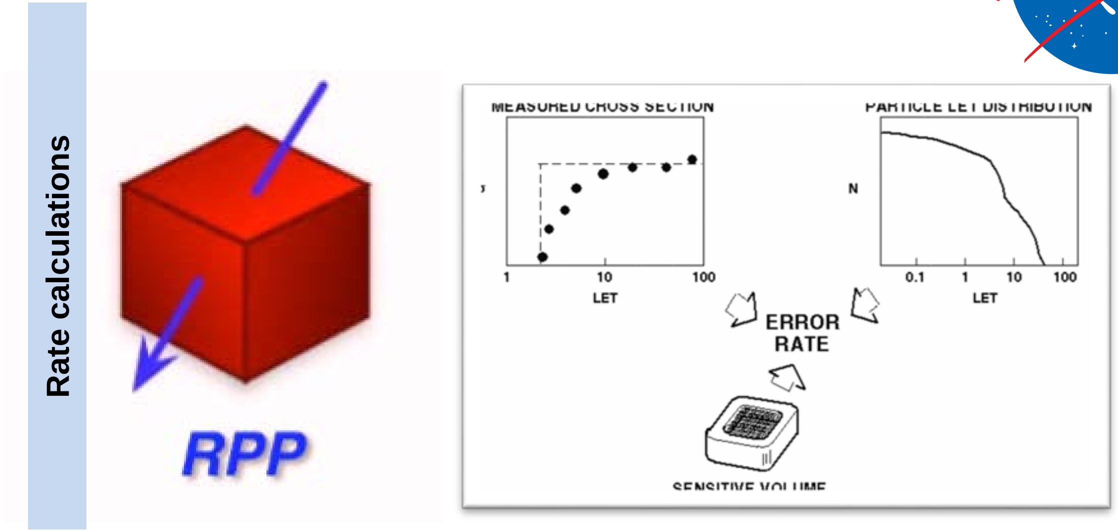


J. Lauenstein *et al.*, doi:10.1109/NSREC.2017.8115473

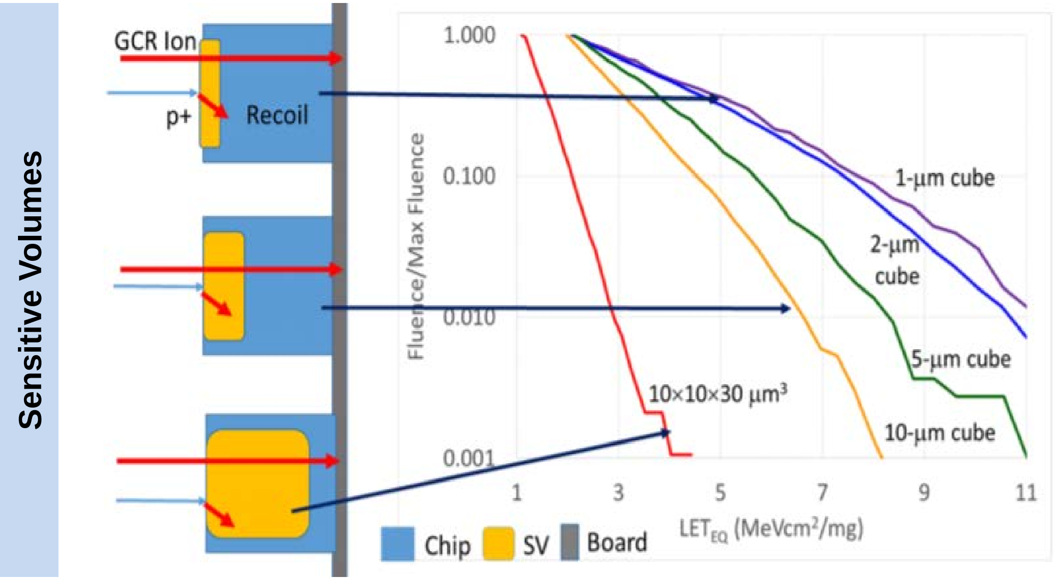
To be presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16th, 2019

Why You Can’t Relax an LET Requirement

- **Rate calculations are not the same for Destructive vs. Non-destructive**
  - Data are a limiting factor, one part = one data point
  - For SEE types that exist in a given technology, they present a constant risk in time domain
- **When you require by LET:**
  - Spectrum from environment is imparted on sensitive volumes, where we get LET thresholds (>75 vs. 60 vs. 37 MeV·cm²/mg)
  - Effective LET increases with angle – critical charge is what we are trying to determine
  - CRÈME calculation integrates the two
  - Deep sensitive volumes won’t necessarily get same LET each time with mono-energetic beams



"*Space Radiation Effects on Microelectronics*," *NASA Jet Propulsion Laboratory*



Ray Ladbury, NSREC2017 SC, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

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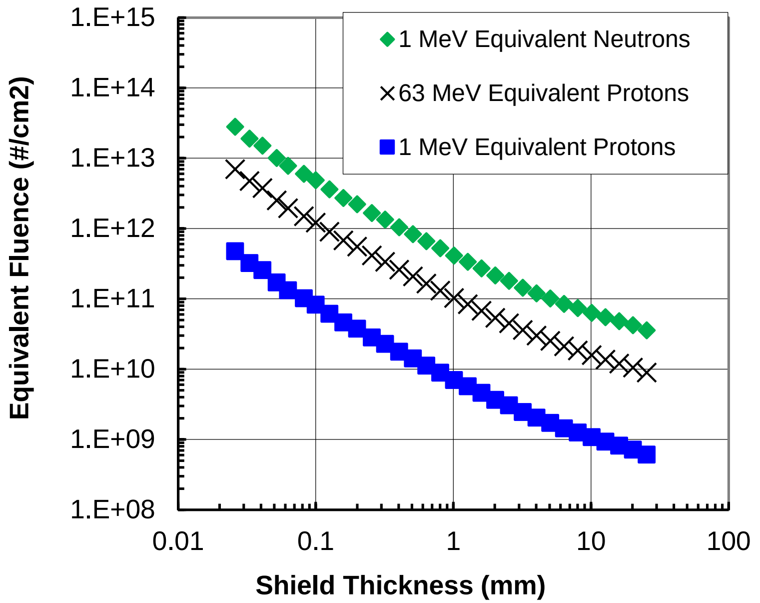
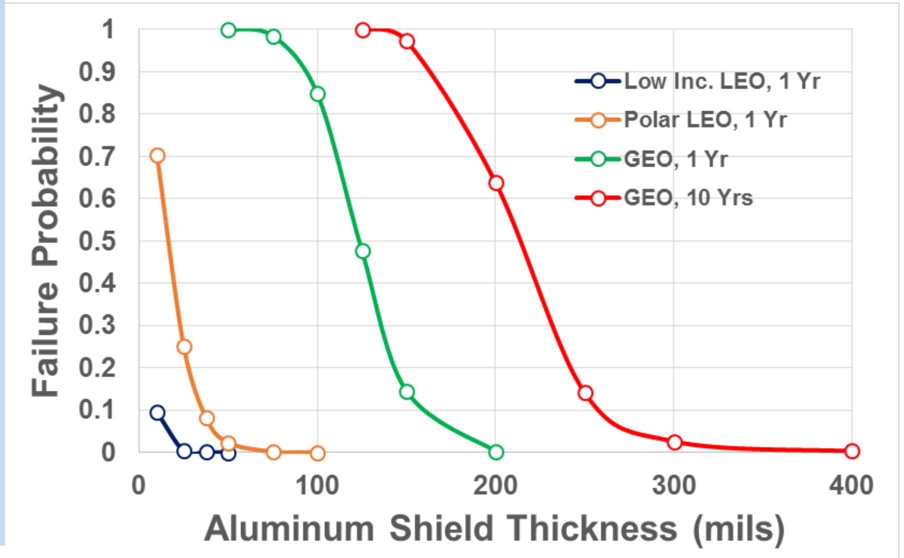
Appreciable Mission Doses

- Maybe degradation of a part beyond usage is okay?
  - Criticality and Application
- Did you forget about DDD?
  - External materials are susceptible as well, polymers can be bad actors and are often on commercial ground based optical systems
- Even short missions can have a common failure mode
- Low mass budget?
  - Can optimize shielding if you have failure distribution of intended components

Non-Ionizing Energy Loss – Displacement Damage Dose



Shielding Optimization



Mike Xapsos, https://doi.org/10.1109/TNS.2016.2607021

Operation During Flare Conditions: Think Availability

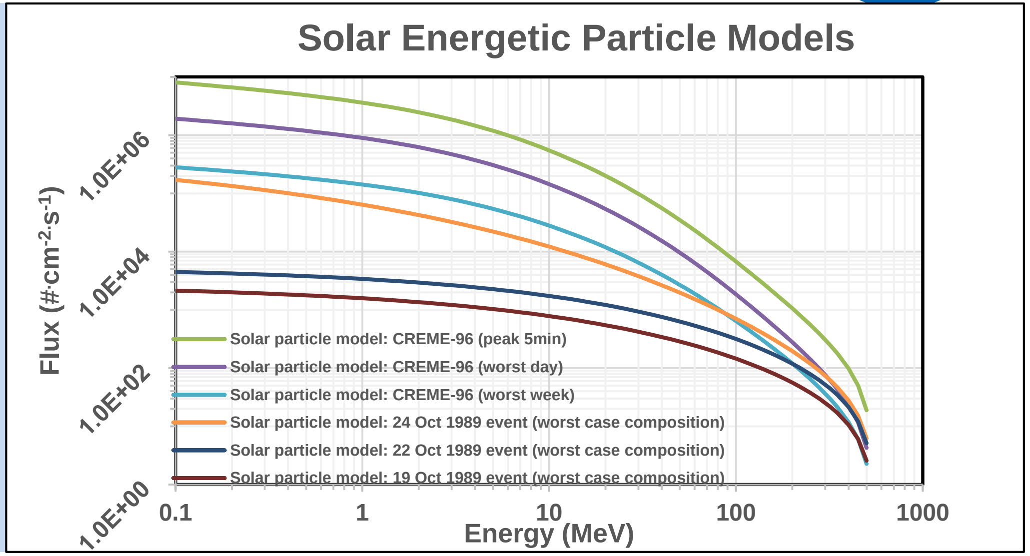
- **Don't dose out during storm (nor the full mission)**
  - Calculate the dose (TID/TNID) of the mission in full –

    95% confidence level recommended
  - Calculate the dose contribution from N number of events (protons & x-rays), if dose from N is > 5% of the total dose, increase confidence level of full mission model
- **Don't destructively fail from a single particle during the storm (nor the full mission)**
  - Standard risk-avoidant SEE approach: no destructive effects allowed
  - LET threshold for single event latchup (SEL)

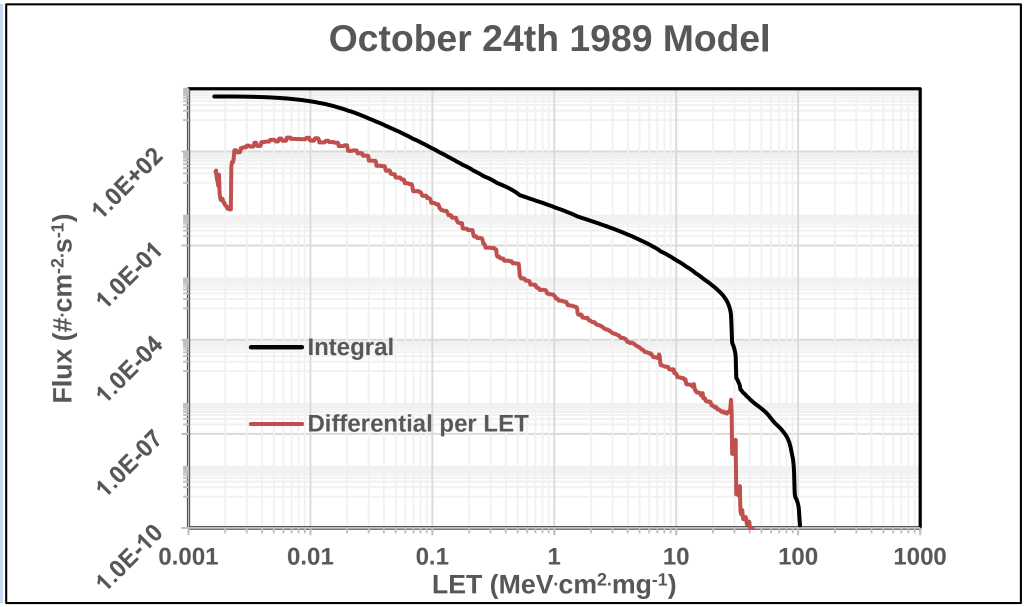
    > 75 MeV.cm^2/mg (some use 60 MeV.cm^2/mg)
  - LET threshold for single event burnout, gate rupture, dielectric rupture (SEB, SEGR)

    > 37 MeV.cm^2/mg (particles must come from normal incidence to cause effect)
- **If you have non-destructive single event upsets, they can't overwhelm critical instruments/systems during the storm**
  - Rate calculation requires part data representative of the application, looking for cross-section over LET.
  - If parts' LET threshold from 20 to 75 MeV.cm^2/mg, need heavy ion rate
  - If parts' LET threshold is below 20, need indirect ionization from recoil ions contribution to rate (need proton data) – make sure packaging materials don't add to this, direct ionization from protons (can be built-in to heavy ion calculation) possible
  - Do you need to mitigate or not – confirm that event rates are not higher than mitigation (Markov process… i.e. EDAC beats the number accrued, scrub rate is faster than critical number of upset accumulation)

Many Solar Models



Spectrum Translates to LET

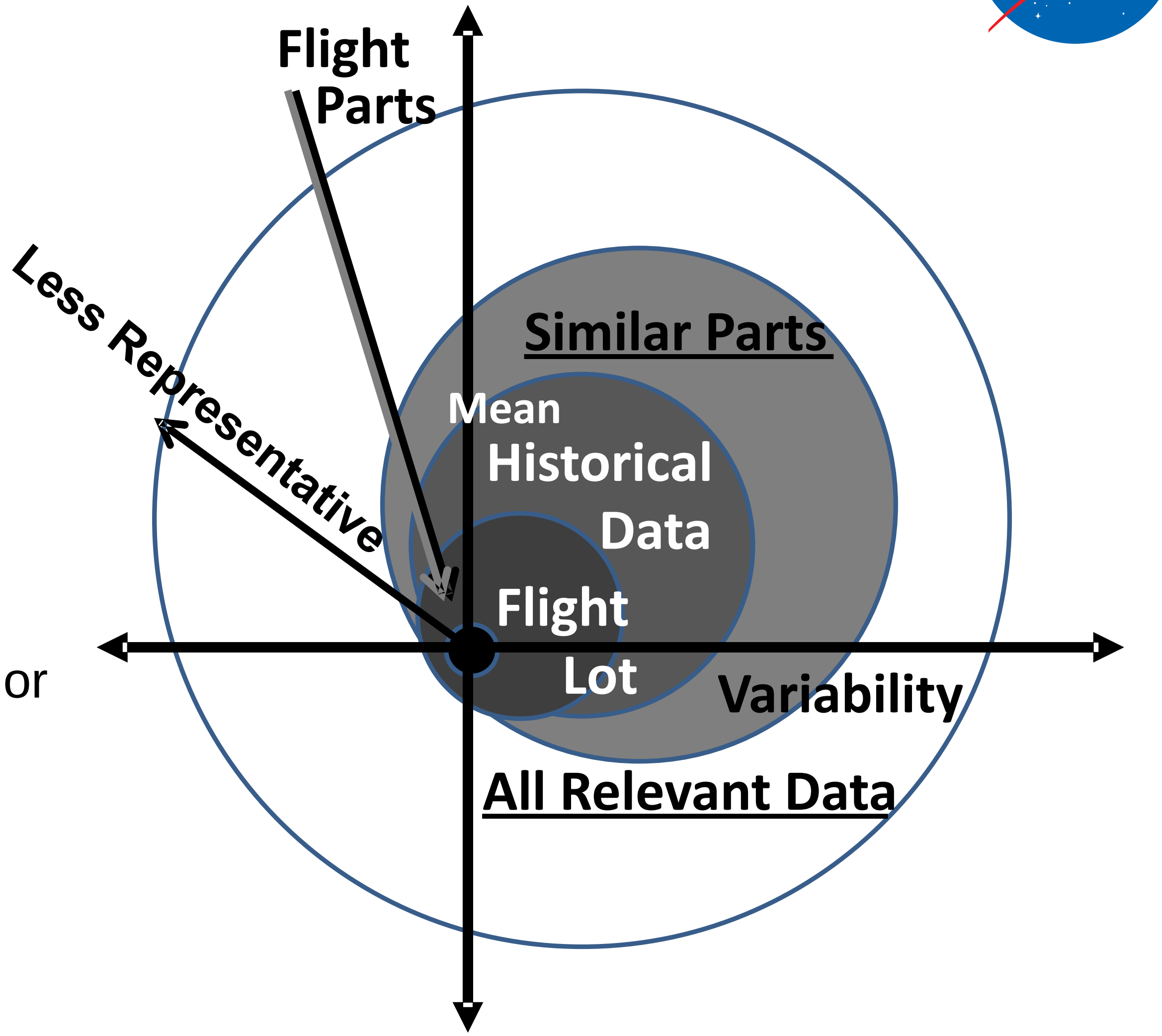


Risk Acceptance – Data Available?

- **Part Classifications Growing**
  - Mil/Aero vs. Industrial vs. Medical
  - Automotive vs. Commercial vs. Modified HiRel

- **Substitute COTS in this diagram**
  - Now you have another degree of separation
  - Failure modes not fully understood
  - Unlikely to have historical data
  - Similarity data no applicable due to fab, process, or design rules
  - Cost of testing usually too high

Without traceability you may be depending on non-representative data.



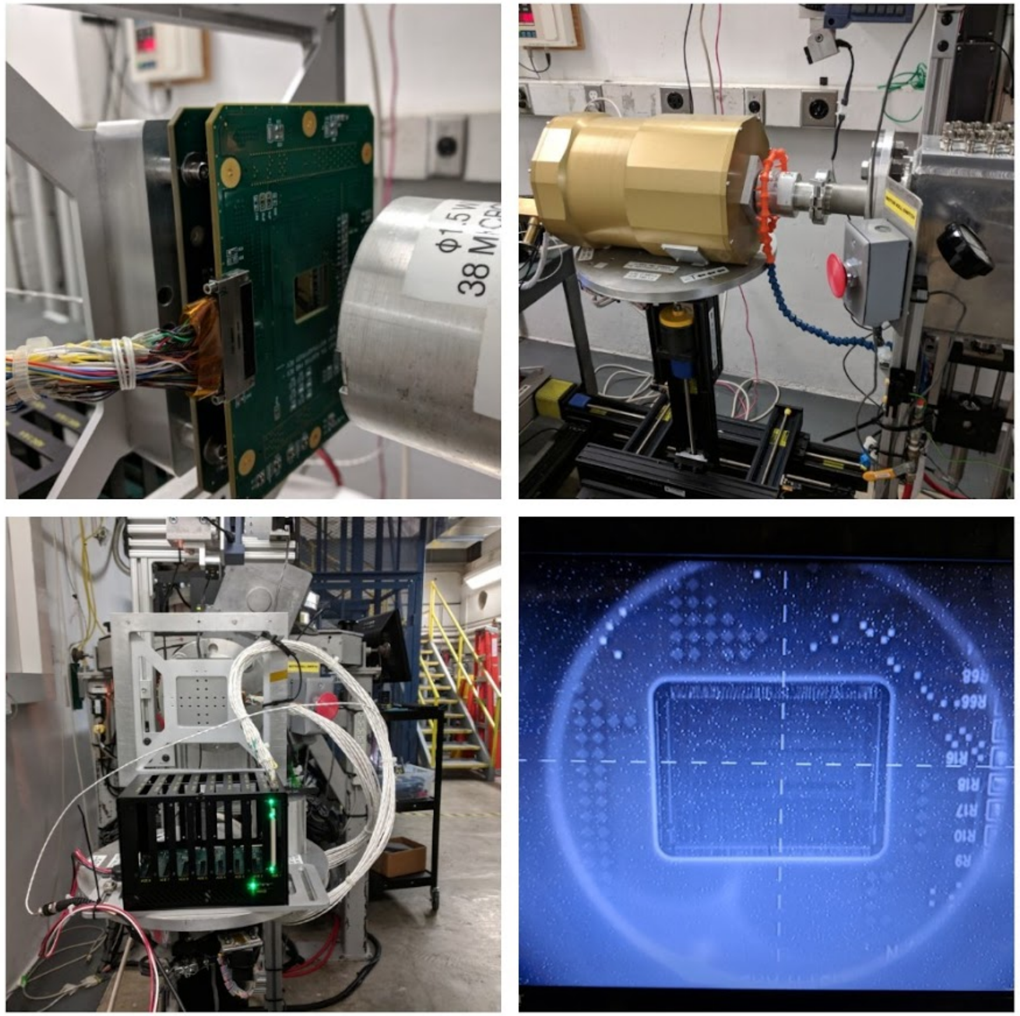
Ray Ladbury, NSREC2017 SC, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

Notional Radiation Data Collection Guidelines

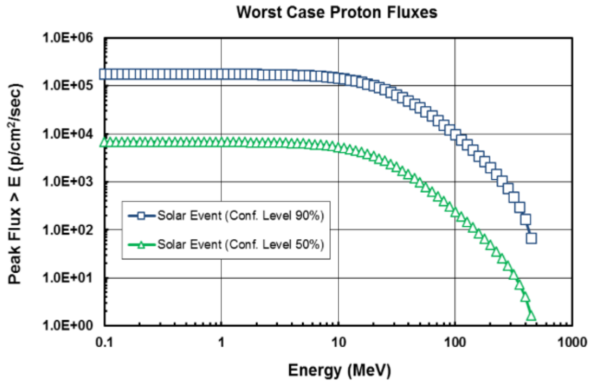
		Environment		
		LEO Equatorial	**LEO Polar (Sun Sync)**	**GEO / Interplanetary**
Mission Lifetime (With Assumed Risk Acceptance)	**> 3 Years**	Data on all SEE for critical parts, and have data on dose failure distribution on similar parts	Consider mission consequences of all SEE (Data for critical parts), have Dose failure distribution on lot	Have Data on all SEE, Have Data Dose failure distribution on lot
	1- 3 Years	Have Data on DSEE for critical parts	Consider mission consequences of all SEE (Data for critical parts), have data Dose failure distribution on similar parts	Have Data on all SEE for critical parts, Have Data on Dose failure distribution on similar parts
	< 1 Year	Look for data on DSEE for critical parts	Consider mission consequences of all SEE, and look for data on dose failure distribution on similar parts	Consider mission consequences of all SEE, and have data on dose failure distribution on similar parts

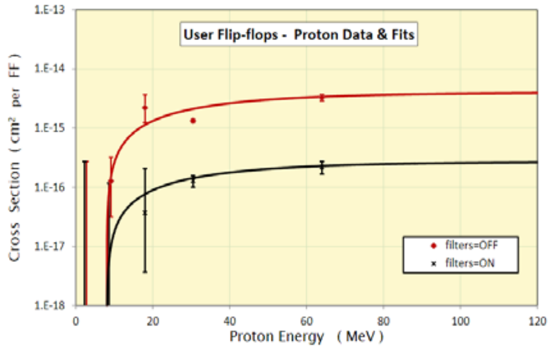
When Do You Test? When Do You Model?

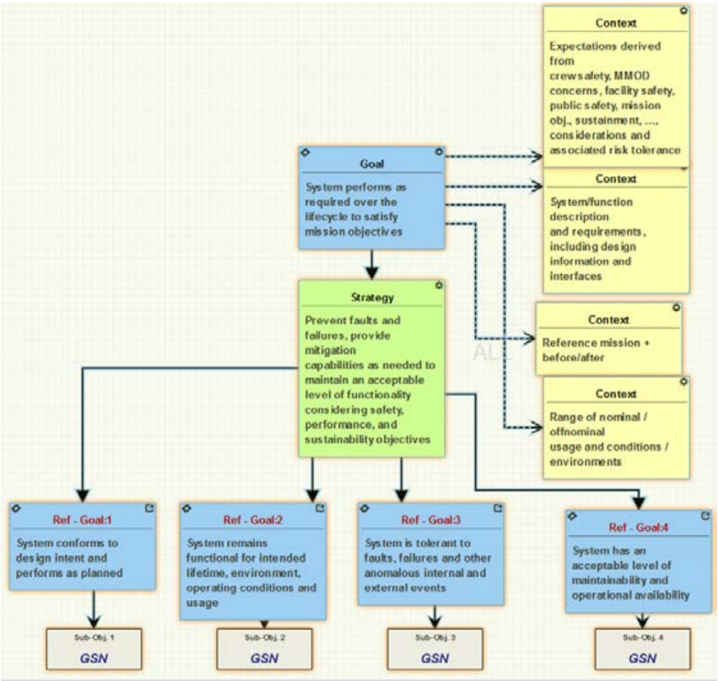
- **Divine your risk threshold**
    - There’s a doc coming for that… radhome.gsfc.nasa.gov/nepp.nasa.gov
- **Unknown failure modes that would not be acceptable to the mission**
    - Known unknowns can be carried as a risk if you already know that the outcome is mitigated at the board or box level
    - New technologies should be identified early on
- **Fault propagation may be the problem you wish to mitigate**
    - This can include cumulative effects!
    - Fault injection may not be able to cover the state space
- **Destructive single event effects are an obvious target**
- **Can you tolerate a part replacement in your design cycle?**
    - Lead times, board re-spins, etc.

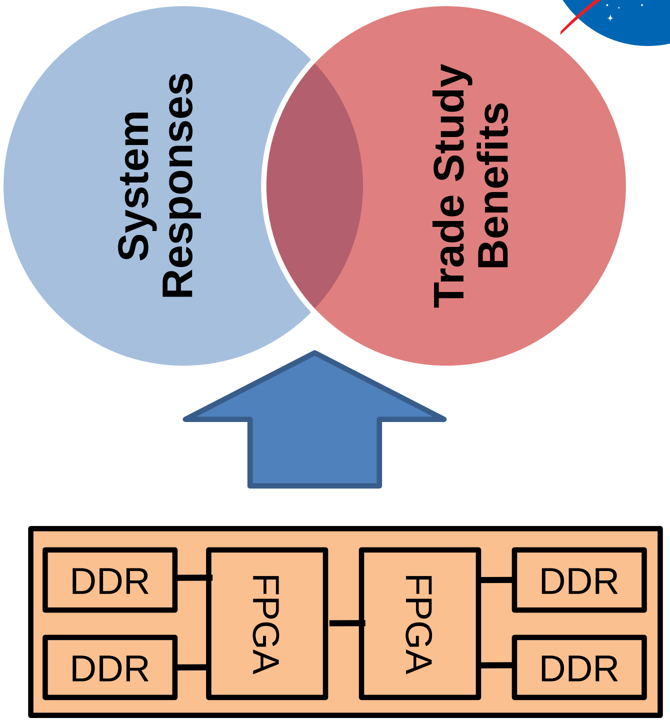


Model Based Mission Assurance (MBMA) as a Tool









Environment, Device, & Design

- ***Models and Test Data*** are brought together to get rates of upset / failure distributions
- ***Resources and Utilization*** are the scaling factors with criticality

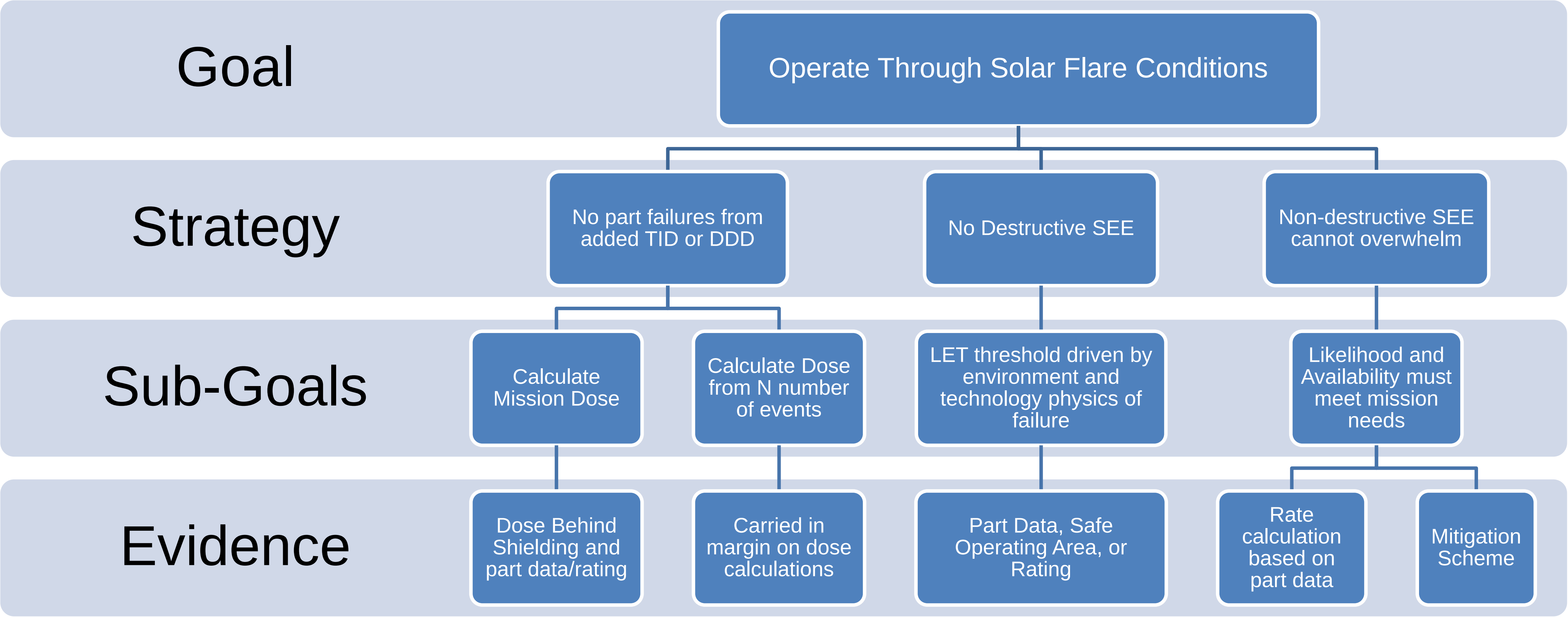
Goal Structuring Notation (GSN)

- Concept of operations
- ***Requirements and Availability*** are fed down correctly to subsystem
- Evidence is presented
- Assumptions are tracked

Systems Modeling Language

- Description of System Connections and Dependencies
- Receives GSN readily
- ***Fault propagation*** can be identified

Goal Structuring Notation (GSN)



Goal

Operate Through Solar Flare Conditions

Strategy

No part failures from added TID or DDD

No Destructive SEE

Non-destructive SEE cannot overwhelm

Sub-Goals

Calculate Mission Dose

Calculate Dose from N number of events

LET threshold driven by environment and technology physics of failure

Likelihood and Availability must meet mission needs

Evidence

Dose Behind Shielding and part data/rating

Carried in margin on dose calculations

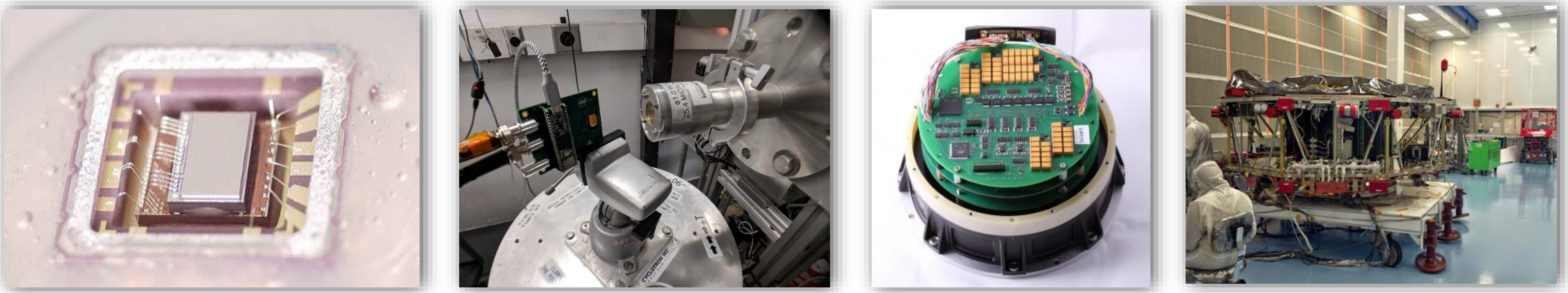
Part Data, Safe Operating Area, or Rating

Rate calculation based on part data

Mitigation Scheme

Questions to Keep in Mind

- What are the radiation risks:
    - What is the hazard?
    - What are the challenges?
- What can you do to reduce the risk for a given hazard?
- What does changing that radiation environment mean for success?
- Need availability throughout the mission or at specific times?
- How do *similar* systems/devices react in the space environment?



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THANK YOU