



**OSIRIS-REx**  
ASTEROID SAMPLE RETURN MISSION



# The Image Constraint Measurement Type for Orbit Determination & Geophysical Parameter Estimation

2nd RPI Space Imaging Workshop

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# Outline

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- Brief History
- Measurement Specification
- Constructing the Observables
- Results from Orbital A
- Alternative Approaches to Observable Construction
- Conclusion & Future Work



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# Brief History of Image Constraints (ICs)

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- Originally published in Mazarico, et. al. (2014)
  - Developed as an alternative to direct landmark processing
  - Processed image constraint measurements for NEAR at Eros
- Implemented in GSFC's GEODYN precision orbit determination and geophysical parameter estimation software
- Further improved and tested for Dawn at Vesta (see Centinello, et. al, 2015)
- Recently implemented in the MIRAGE operational navigation software suite for the OSIRIS-REx mission at Bennu
  - Used to supplement traditional direct-landmark measurements



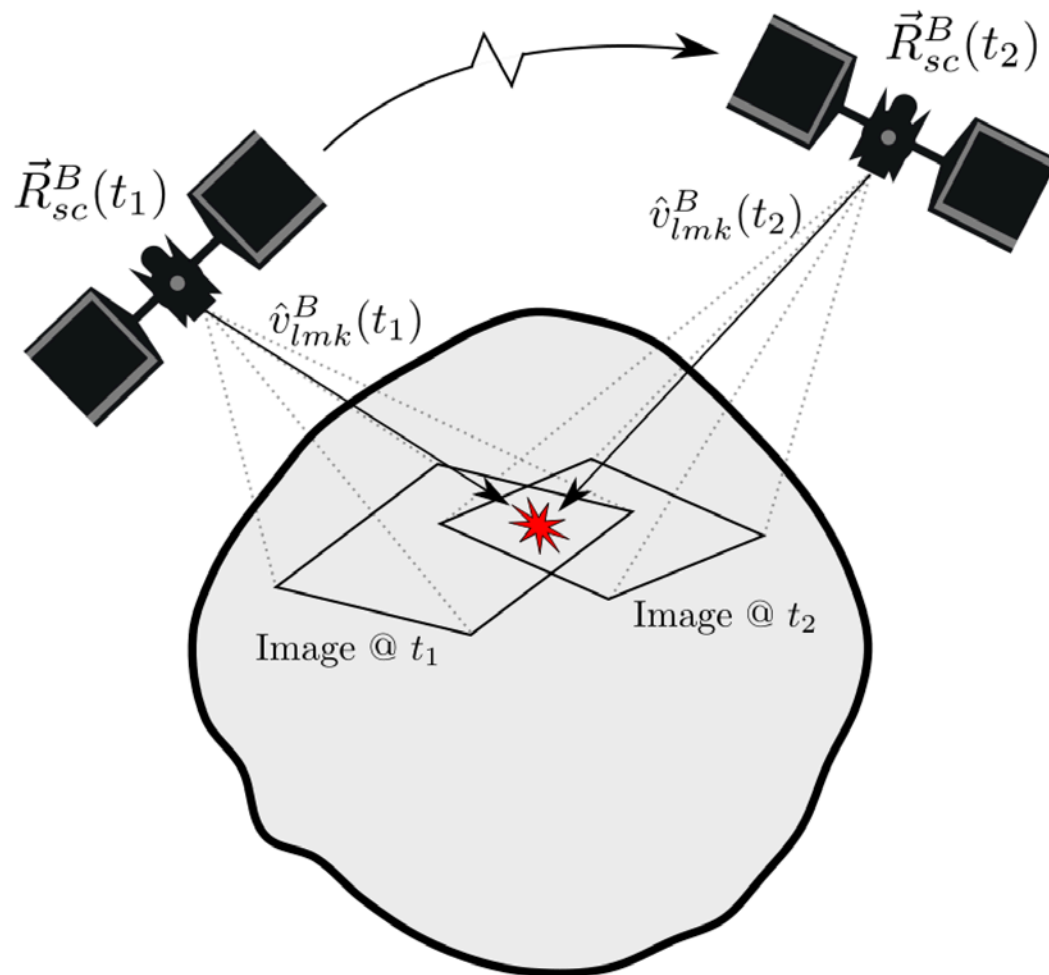
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# Measurement Specification (1/X)



- Consists of observations of a single surface feature (or landmark) in two different images
- The unit vectors corresponding to each observation must, by definition, intersect in the target body's fixed frame
- Constrains the relative position and orientation of the spacecraft at the time of the two exposures
- Similar to visual odometry but at much longer baselines



# Measurement Specification (2/X)

- Start with observed sample/line location of landmark,  $[s, l]$  (in pixels) for each image
  - Extracting  $[s, l]$  for each landmark is addressed later...
- $\hat{v}_{lmk}^C(t)$  - find by inverting an appropriate camera model (OpenCV for OSIRIS-REx):

$$\begin{aligned} \begin{bmatrix} s \\ l \end{bmatrix} &= g(\hat{v}_{lmk}^C, f_{x,y}, k_{0-6}, p_{1-2}, c_{x,y}) \\ \hat{v}_{lmk}^C &= g'(s, l, f_{x,y}, k_{0-6}, p_{1-2}, c_{x,y}) \end{aligned}$$

- Rotate  $\hat{v}_{lmk}^C(t)$  from the camera frame to the spacecraft-fixed frame:

$$\hat{v}_{lmk}^S(t) = C_B^I(t) \delta C C_S^C \hat{v}_{lmk}^C(t)$$

- Rotate  $\vec{R}_{sc}^I(t)$  and  $\hat{v}_{lmk}^I(t)$  to target-body fixed frame using IAU rotation parameters:

$$\begin{aligned} \vec{R}_{sc}^B(t) &= C_B^I(t) \vec{R}_{sc}^I(t) \\ \hat{v}_{lmk}^B(t) &= C_B^I(t) \hat{v}_{lmk}^I(t) \\ C_B^I(t) &= f(t, \alpha, \delta, \omega_0, \omega, \dots) \end{aligned}$$



## Measurement Specification (3/X)

- **Minimum intersect distance** of two rays with origins  $\vec{R}_{sc}^B(t_1)$ ,  $\vec{R}_{sc}^B(t_2)$  and directions  $\hat{v}_{lmk}^B(t_1)$ ,  $\hat{v}_{lmk}^B(t_2)$ :

$$d = \left[ \vec{R}_{sc}^B(t_1) - \vec{R}_{sc}^B(t_2) \right]^T \left[ \frac{\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)}{|\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)|} \right]$$

- By definition,  $d$  is nominally zero:  $\delta y = 0 - d = -d$
- Image constraints are a **differenced measurement**:
  - No dependence on landmark locations/shape model (except during processing)
  - Some 1<sup>st</sup> order errors in location/shape cancel
  - Still susceptible to shape/orbit scale errors





# Information Content

- Provides relative state information ( $\Delta \vec{R}_{SC}^B$ ) perpendicular to the baseline:

$$\frac{\partial d}{\partial \Delta \vec{R}_{SC}^B} = \left[ \frac{\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)}{|\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)|} \right]^T$$

- Alternate formulation for epoch-state filters (in terms of  $X^I(t_0)$ ):

$$d = X^I(t_0)^T [C_B^I(t_1)\Phi(t_0, t_1) - C_B^I(t_2)\Phi(t_0, t_2)]^T \left[ \frac{\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)}{|\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)|} \right]$$

$$\frac{\partial d}{\partial X^I(t_0)} = \left[ \frac{\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)}{|\hat{v}_{lmk}^B(t_1) \times \hat{v}_{lmk}^B(t_2)|} \right]^T [C_B^I(t_1)\Phi(t_0, t_1) - C_B^I(t_2)\Phi(t_0, t_2)]$$

- Also provides information for:
  - Camera Pointing Correction -  $\delta C$
  - Target Body Orientation -  $\alpha, \delta, \omega$  (not  $\omega_0$ !)



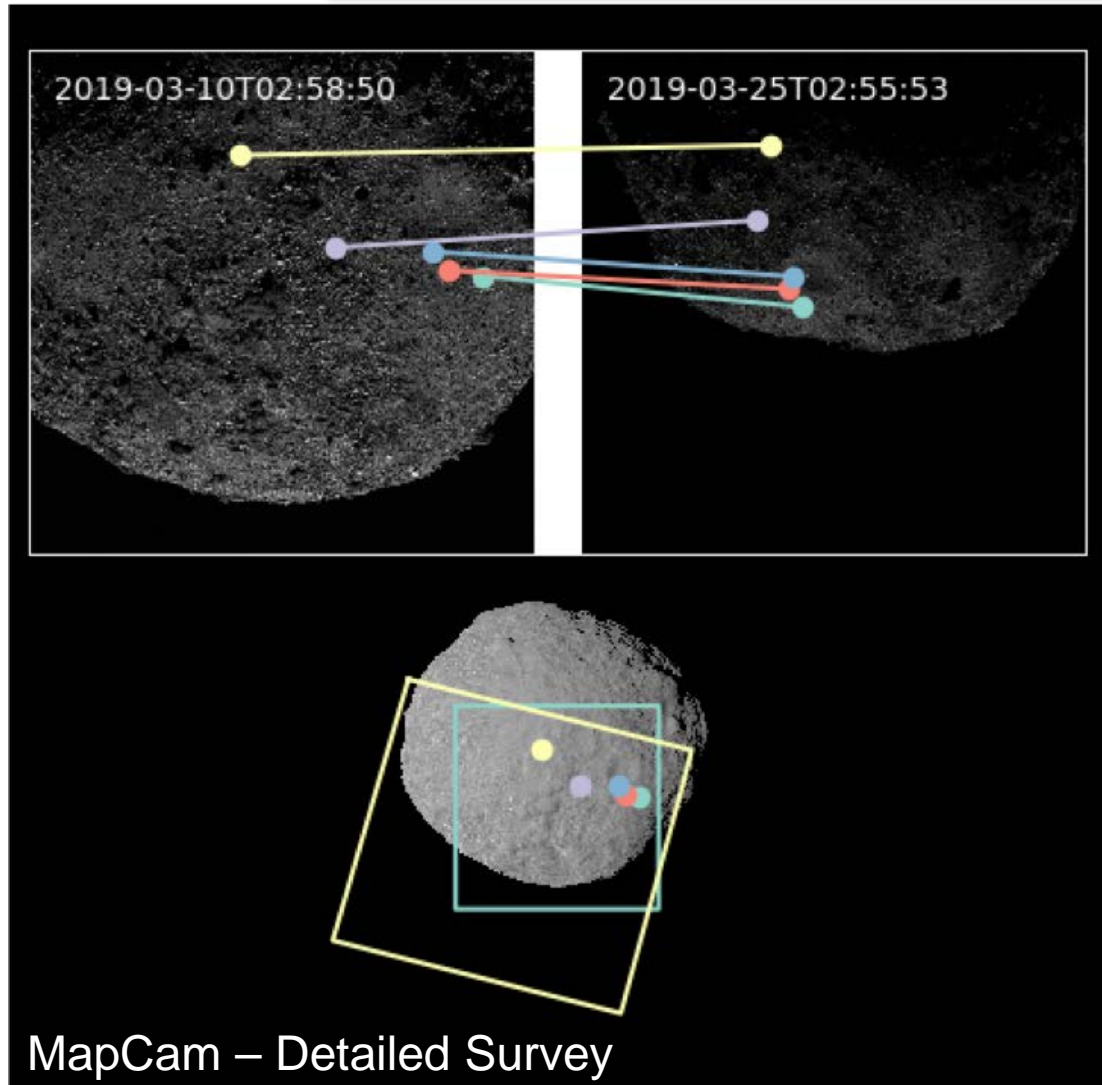
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# Constructing the Observables



- Need to identify common landmarks between two images
  - Usually from the same imager, but can be different
- Convenient by-product of traditional SPC (or similar) terrain-relative navigation
  - Requires a full shape model
  - See subsequent slides for alternative approaches...
- Commonly constrain:
  - Min/Max time between landmarks
  - Number of “appearances” of each landmark
- Ideal to select combinations that vary the baseline direction



# Outline

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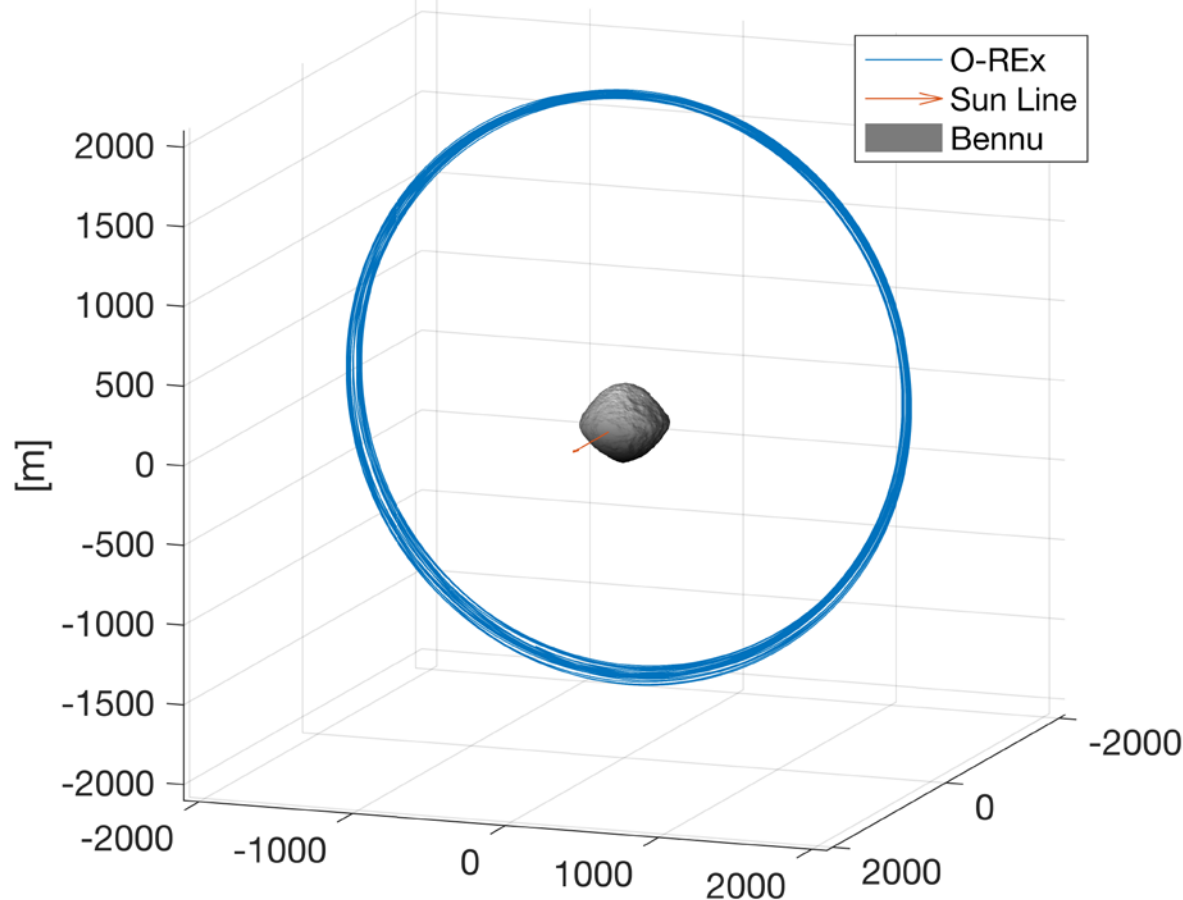


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# Example: OSIRIS-REx Orbital A

**OSIRIS-REx Orbit A**  
**[Bennu-Centered Sun-North]**



- Processed Direct Landmark and IC data from Orbital A in GEODYN
  - *Representative example – do not replace official results from FDS and other science working groups*
- 2.1 x 1.6 km “Frozen” terminator orbit
  - January 1<sup>st</sup> to February 28<sup>th</sup>
- Long and Short NavCam Exposures
  - Precise camera pointing
- Shape Model: 75 cm GSD
  - From Altimetry Working Group (iterated)
  - Sub-sampled to ~1000 landmarks





# Orbital A Image Constraints

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- Total number of Direct Landmark Observables: **91,999**
  - From 1757 images and 992 Landmarks
- Filter Criterion:
  - Minimum Time Between Image Pairs: 60 hours (~1 orbit period)
  - Maximum Time Between Image Pairs: None (entire arc, ~2 months)
- Total number of Image Constraint Observables: **107,573**
  - Most “appearances” by a single landmark: 23
- Also looked into filtering landmark pairs by baseline direction to maximize information content (J. Leonard/KinetX)



# Orbital A Filter Cases

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- Case 1:
  - Direct Landmarks Only (1 pixel weight)
  - No scale or COM-COF estimation
- Case 2:
  - Direct Landmarks Only (1 pixel weight)
  - Estimate scale and COM-COF
- Case 3:
  - Direct Landmarks Only (3 pixel weight)
  - Estimate scale and COM-COF
- Case 4:
  - Image Constraints Only (10 cm weight)
- Case 5:
  - Image Constraints Only (75 cm weight)



# Orbital A Filter Cases



- Solve-For Parameters:
  - Spacecraft:
    - Position/Velocity at Epoch
    - Impulsive Momentum Desaturations
      - Twice-per-week
    - Stochastic Accelerations
      - 1-day Batches
    - Per-Pass DSN Range Biases
  - Bennu:
    - GM and J2
    - Spin Pole & Rate
    - Shape Model Scale (Cases 2 & 3)
    - COM-COF Offset (Cases 2 & 3)
- Measurements (Weight)
  - DSN Range (21 RU)
  - DSN Doppler (5.5 mHz)
  - Direct Landmarks (1 or 3 pixels)
  - Image Constraints (10 or 75 cm)



# Residual Comparison



	Direct Landmarks		Image Constraints	
	Mean (pix)	RMS (pix)	Mean (cm)	RMS (cm)
Case 1: DL (1 px), no Scale or COM-COF	0.0019	0.6315	2.11	34.25
Case 2: DL (1 px), Scale & COM-COF	0.0021	0.6187	1.89	34.13
Case 3: DL (3 px), Scale & COM-COF	0.0025	0.6402	2.38	34.52
Case 4: ICs Only (10 cm)	0.2293	1.516	-0.64	29.29
Case 5: ICs Only (75 cm)	0.1665	1.0856	-0.67	29.74

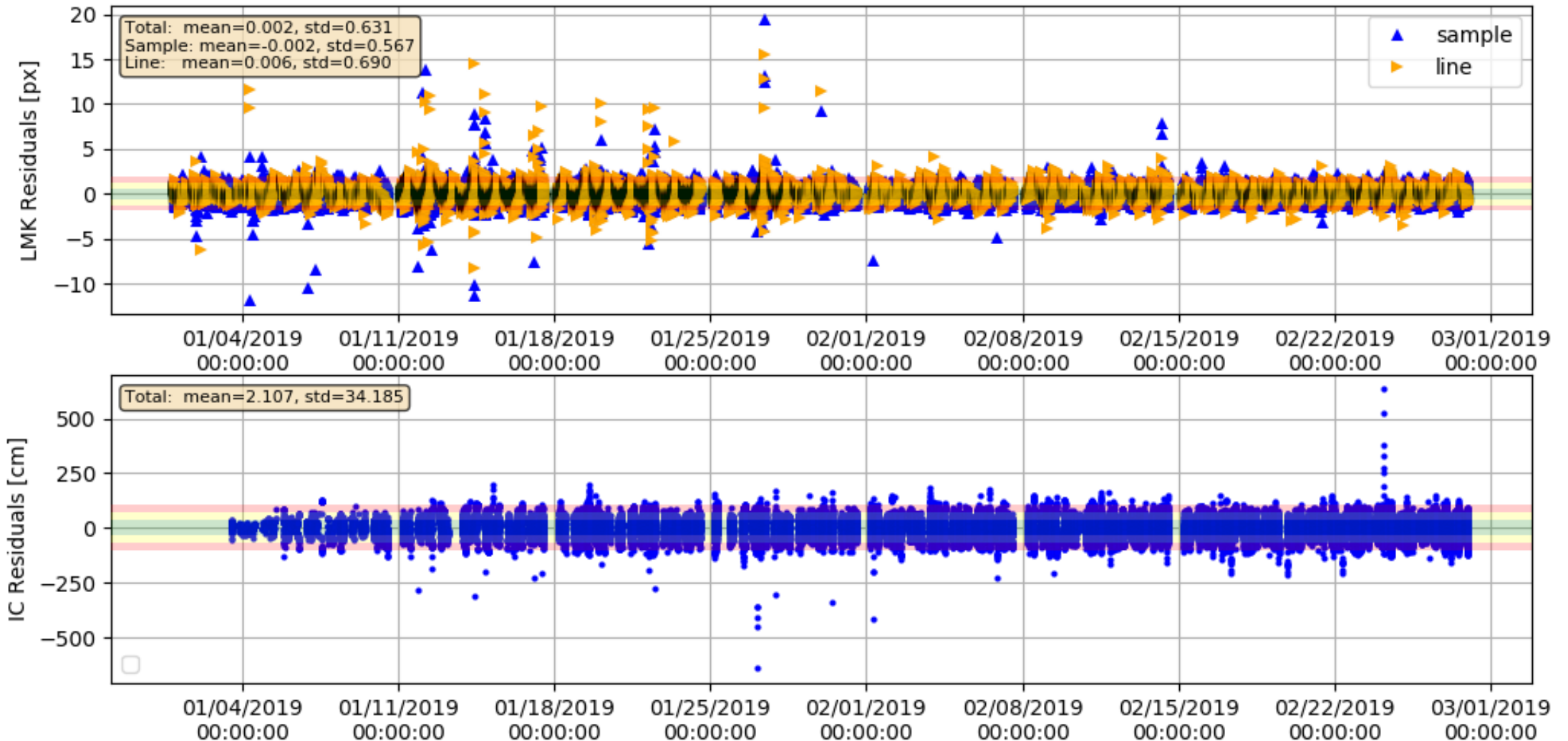
- **Fit / Passthru**
- Residuals are similar for DL-only and IC-only cases



# Residuals: Case 1

## Direct Landmarks Only (1 px), No Scale/COM-COF

Optical Navigation Residuals: Case #1 (NavCam 1)



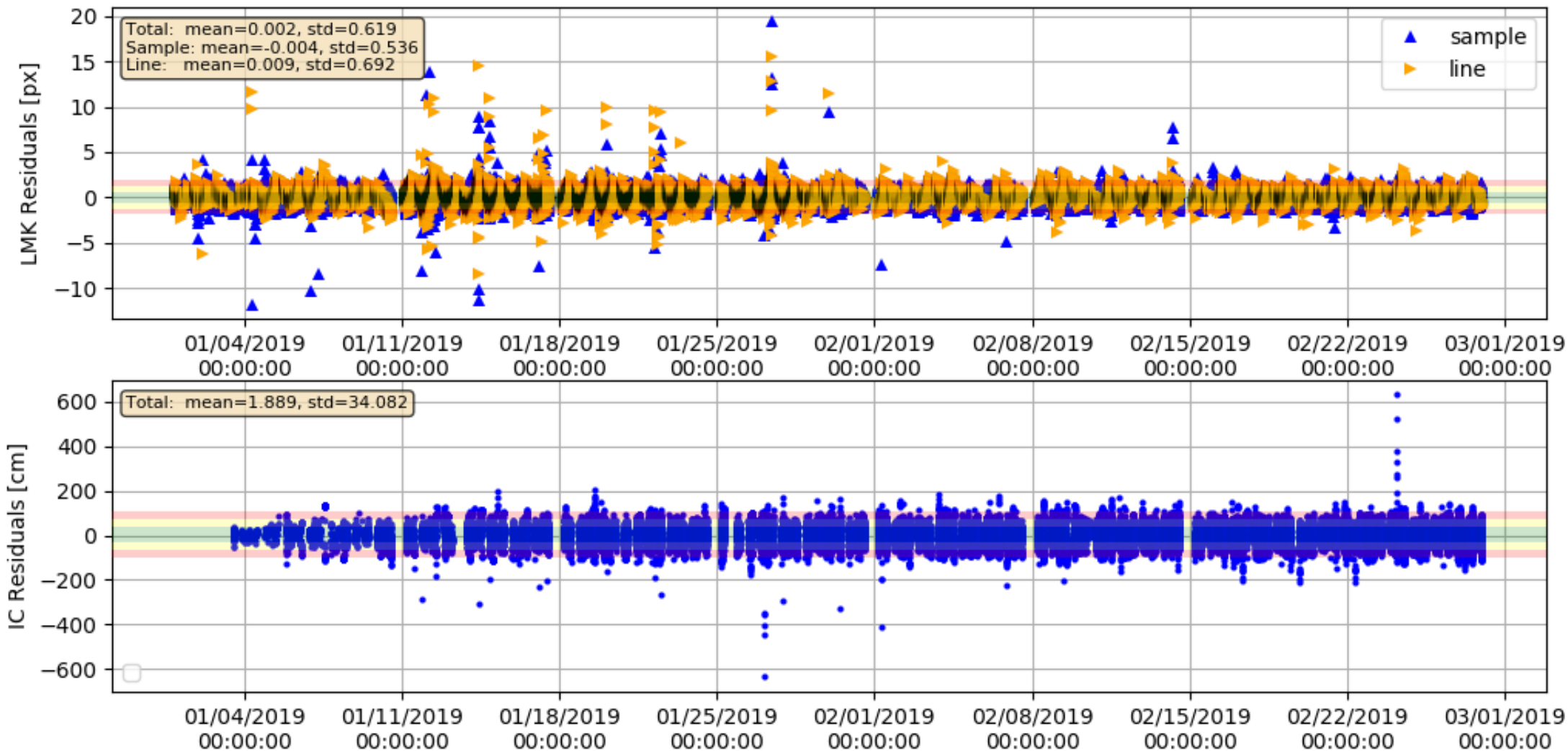




# Residuals: Case 2

## Direct Landmarks Only (1 px), With Scale/COM-COF

Optical Navigation Residuals: Case #2 (NavCam 1)

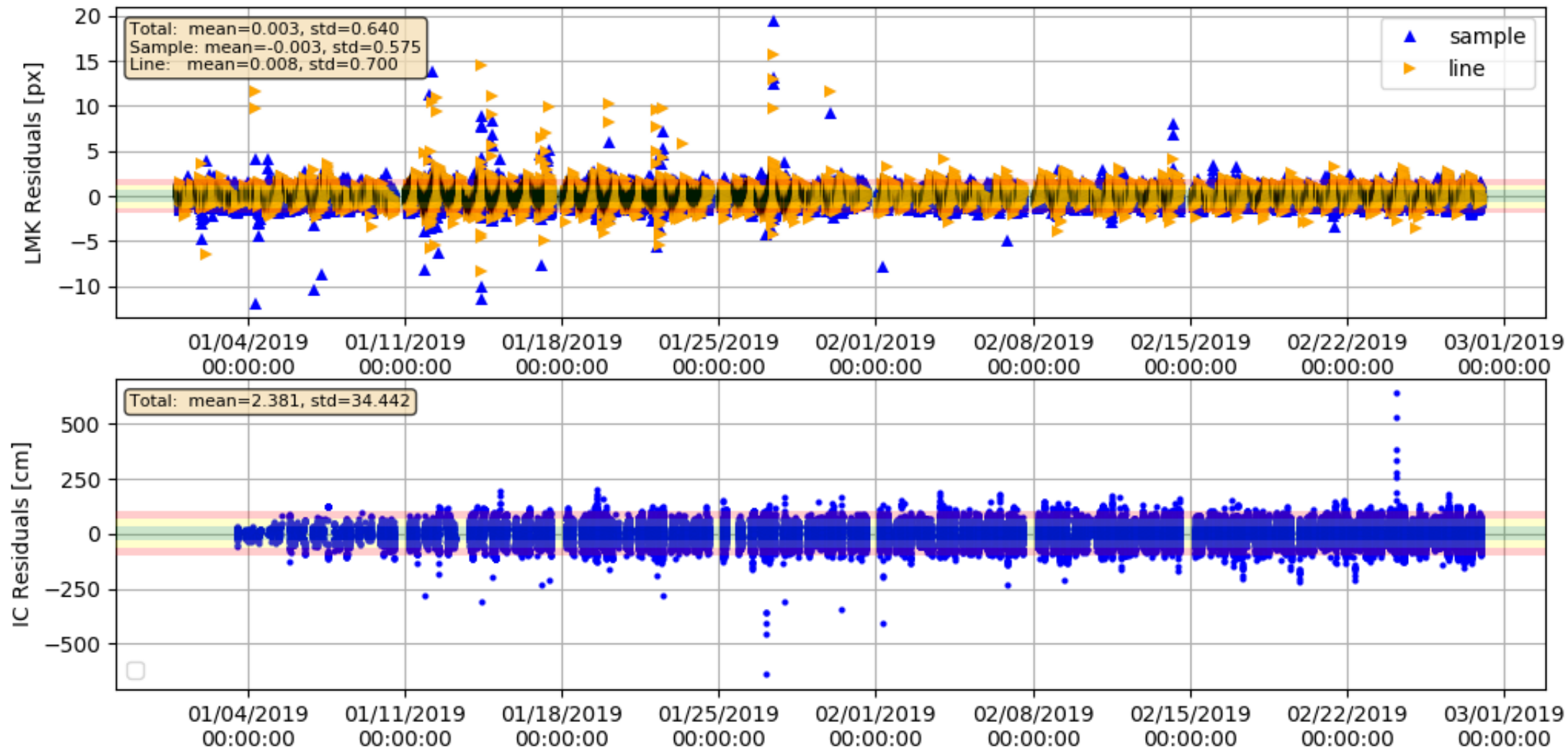




# Residuals: Case 3

## Direct Landmarks Only (3 px), With Scale/COM-COF

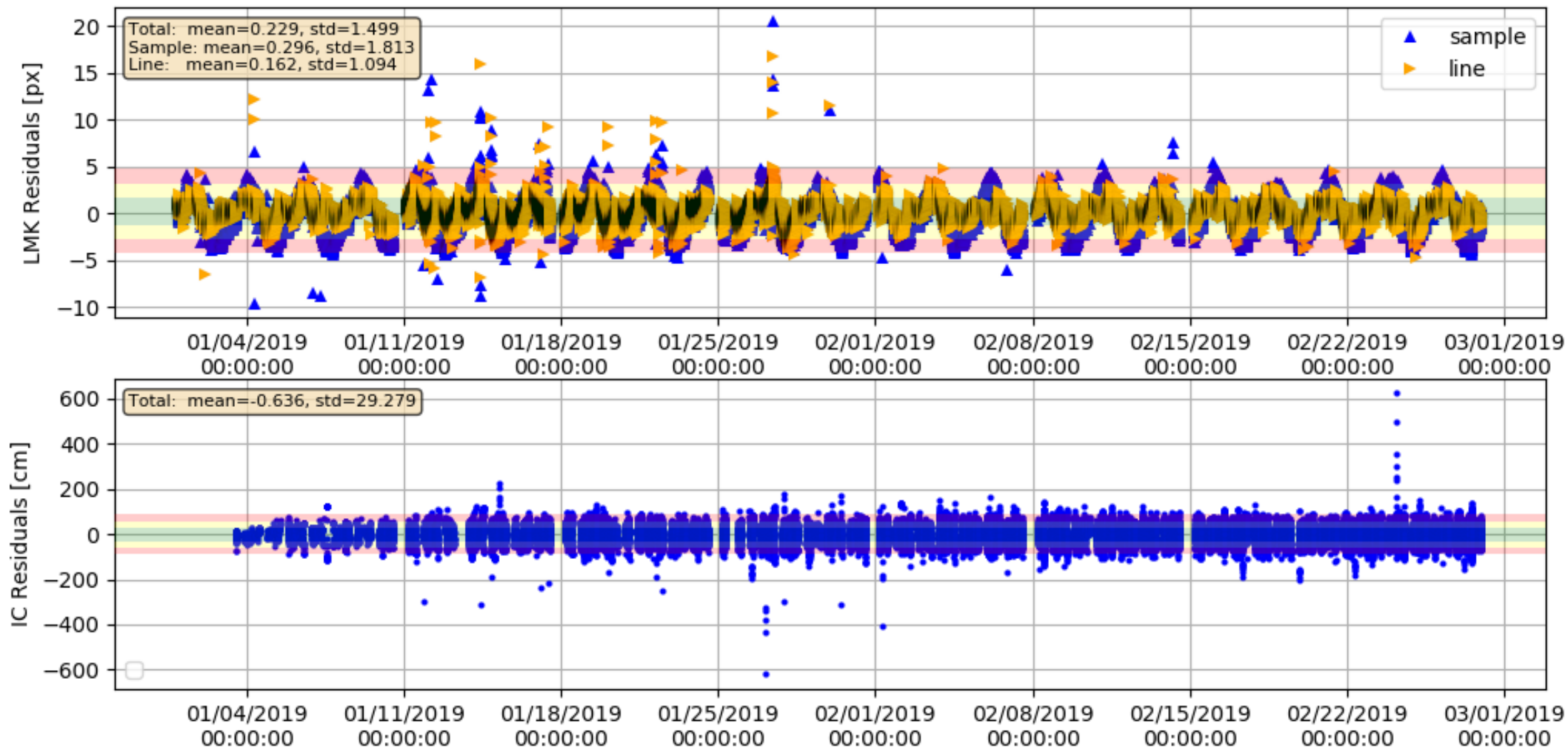
Optical Navigation Residuals: Case #3 (NavCam 1)





# Residuals: Case 4 (Image Constraints Only, 10 cm)

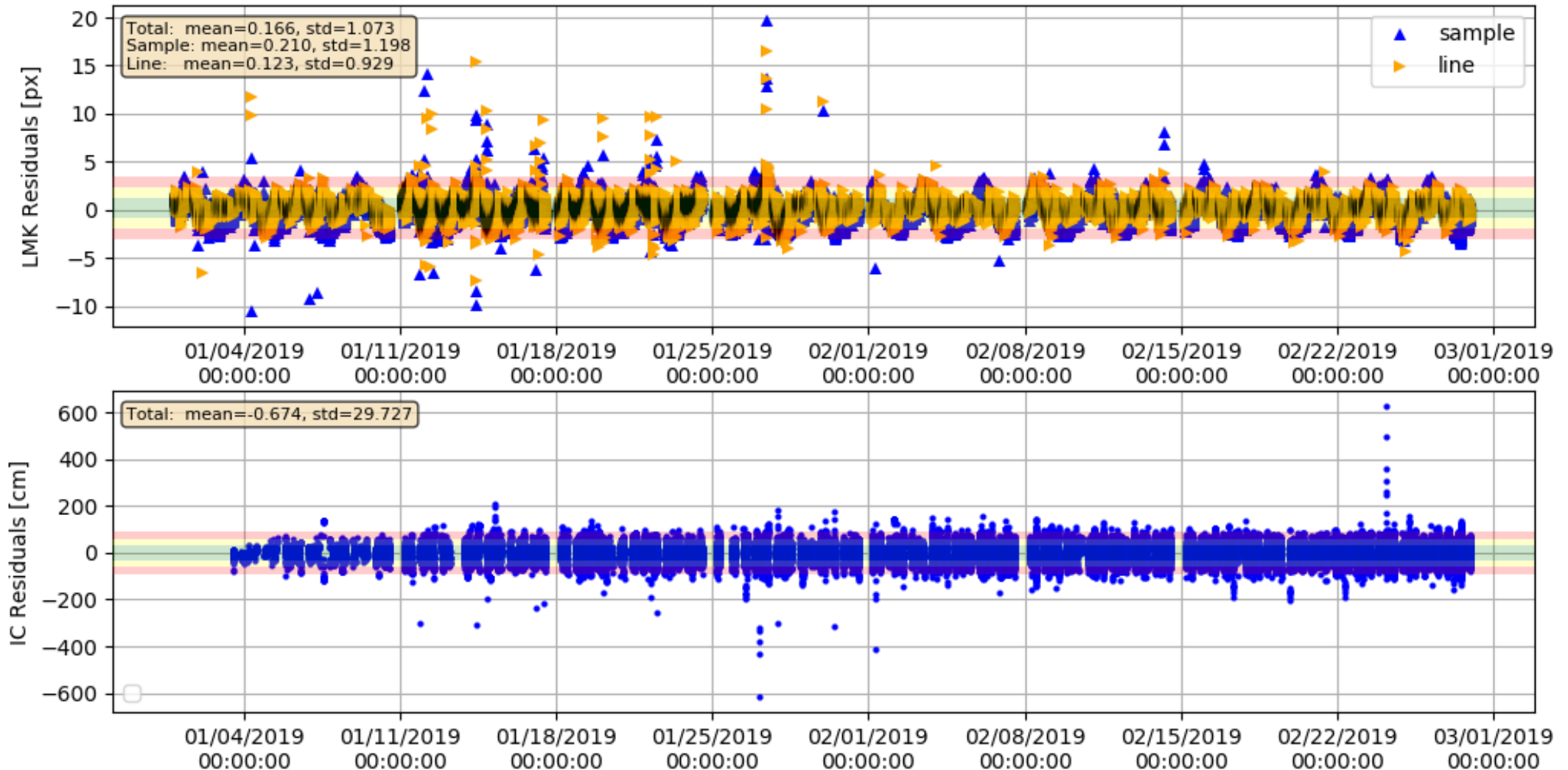
Optical Navigation Residuals: Case #4 (NavCam 1)





# Residuals: Case 5 (Image Constraints Only, 75 cm)

Optical Navigation Residuals: Case #5 (NavCam 1)





# Spin State Sensitivity



	$\alpha$	$\delta$	$\omega_0$	$\omega$
<b>Case 1</b>	$85.5039^\circ \pm 0.223^\circ$	$-60.2974^\circ \pm 0.070^\circ$	$27.3665^\circ \pm 0.298^\circ$	$0.02328 \pm 4.6e-8^\circ/\text{sec}$
<b>Case 2</b>	$85.5036^\circ \pm 0.223^\circ$	$-60.2905^\circ \pm 0.071^\circ$	$27.3642^\circ \pm 0.298^\circ$	$0.02328 \pm 4.6e-8^\circ/\text{sec}$
<b>Case 3</b>	$85.4991^\circ \pm 0.583^\circ$	$-60.2969^\circ \pm 0.209^\circ$	$27.3599^\circ \pm 0.842^\circ$	$0.02328 \pm 1.4e-7^\circ/\text{sec}$
<b>Case 4 (w/<math>\omega_0</math>)</b>	$85.4437^\circ \pm 0.052^\circ$	$-60.3767^\circ \pm 0.019^\circ$	<b><math>16.6103^\circ \pm 12.98^\circ</math></b>	$0.02328 \pm 4.6e-8^\circ/\text{sec}$
<b>Case 4 (w/o <math>\omega_0</math>)</b>	$85.4435^\circ \pm 0.052^\circ$	$-60.3767^\circ \pm 0.019^\circ$	$27.3118^\circ$ (Fixed)	$0.02328 \pm 7.5e-9^\circ/\text{sec}$
<b>Case 5</b>	$85.4503^\circ \pm 0.363^\circ$	$-60.3760^\circ \pm 0.137^\circ$	$27.3118^\circ$ (Fixed)	$0.02328 \pm 5.5e-8^\circ/\text{sec}$

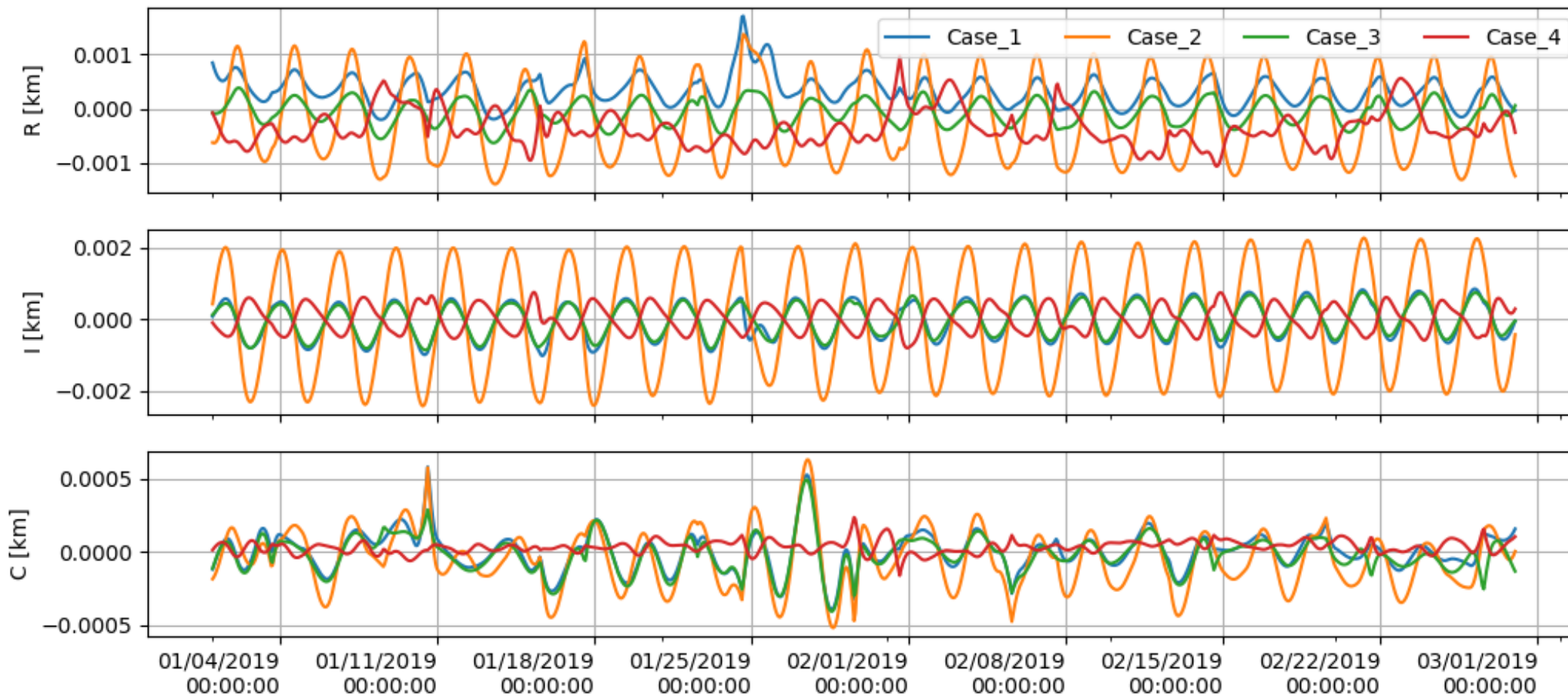
- Spin State Epoch: December 1<sup>st</sup>, 2018 00:00:00.000
- All parameters are within family **except** for ICs-only with  $\omega_0$  estimation





# Trajectories Compared to Case 5

Position Comparison  
Reference Trajectory: Case 5





# Iterating IC Solutions with Direct Landmarks



- Case 6:
  - Held Case 4 trajectory fixed (ICs only, 10 cm weight)
  - Re-estimated Spin Pole/Rate, Shape Model Scale, and COM-COF offset using Direct Landmarks
  - Direct Landmark Residuals: **-0.0147 px mean, 0.6952 px RMS**
- Case 7:
  - Held Case 5 trajectory fixed (ICs only, 75 cm weight)
  - Re-estimated Spin Pole/Rate, Shape Model Scale, and COM-COF offset using Direct Landmarks
  - Direct Landmark Residuals: **-0.0046 px mean, 0.6973 px RMS**

	SCALE	COM-COF X	COM-COF Y	COM-COF Z
<b>Case 1</b>	1.0 (Fixed)	0.0 (Fixed)	0.0 (Fixed)	0.0 (Fixed)
<b>Case 2</b>	$0.9995 \pm 2.07e-5$	$0.3599 \pm 0.17$ cm	$1.235 \pm 0.17$ cm	$146.6 \pm 3.68$ cm
<b>Case 3</b>	$0.9996 \pm 8.97e-6$	$0.2048 \pm 0.52$ cm	$1.266 \pm 0.51$ cm	$-62.80 \pm 5.3$ cm
<b>Case 6</b>	$0.9997 \pm 8.97e-6$	$0.1454 \pm 0.17$ cm	$1.249 \pm 0.17$ cm	$-119.2 \pm 0.24$ cm
<b>Case 7</b>	$0.9998 \pm 8.97e-6$	$-0.0885 \pm 0.17$ cm	$1.427 \pm 0.17$ cm	$-67.87 \pm 0.24$ cm



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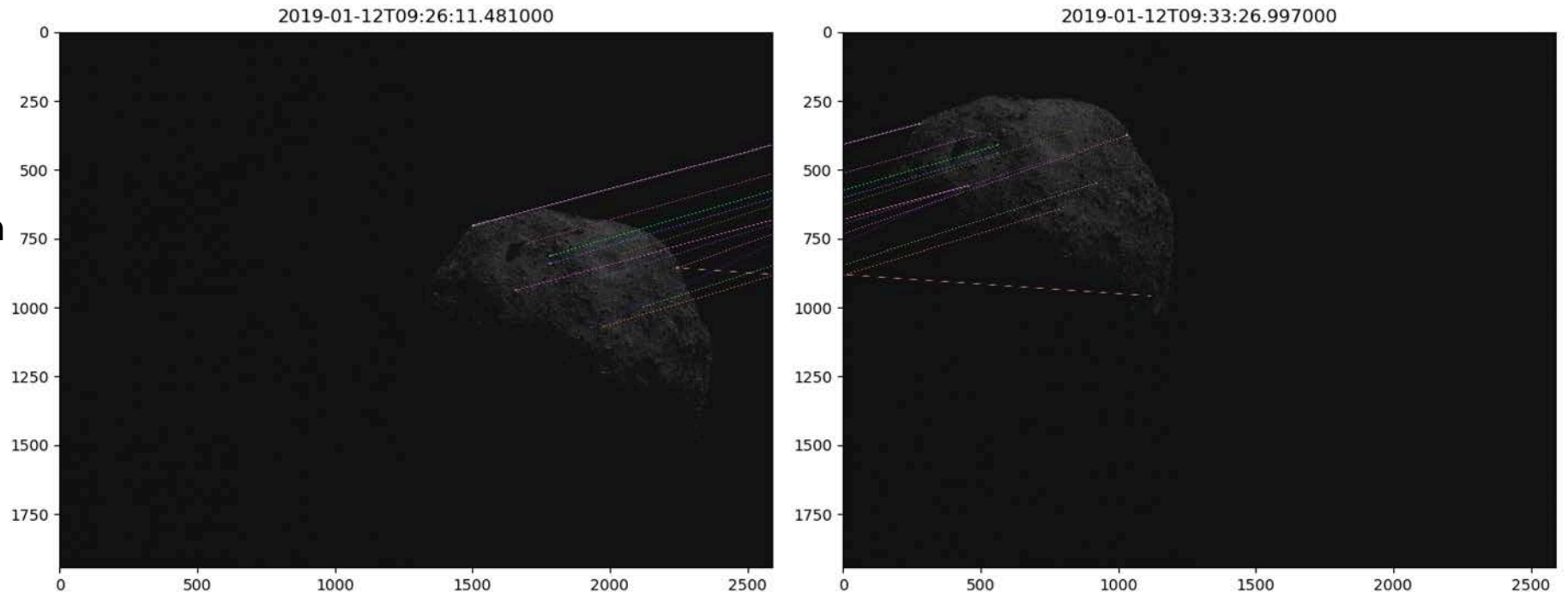


# Alternatives Methods to Identify Constraints

While SPC provides an opportunity to easily extract the location of the same feature in numerous images, it requires having a pre-built shape model, nullifying one of the benefits of this measurement type.

We are currently investigating alternative methods of matching features between images using just image processing.

- Feature Descriptors (shown)
- RANSAC based Keypoint
- Template Matching
- Mutual Information Correlation





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# Conclusions & Future Work

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- ICs are an alternative to Direct Landmark observables in orbit determination and provide an additional metric to evaluate solutions
  - Demonstrated with NEAR, Dawn, and OSIRIS-REx data
- The IC measurement model does not depend on a shape model or pre-defined landmarks
  - Some 1<sup>st</sup> order errors cancel (not scale)
  - Potential to generate ICs without prior shape model development
- Orbital A results are consistent with one-another, depending on the weighting scheme
- Future Work:
  - Continue analyzing OSIRIS-REx data in other mission phases (Surveys, Orbit B, etc.)
  - Refine IC alternative construction techniques that do not require a shape model and test with OSIRIS-REx imagery
  - Compare with other complimentary data types: Direct and Differenced Altimetry



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