

# Radiation Hardness Assurance: Evolving for *NewSpace*

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# Acronyms

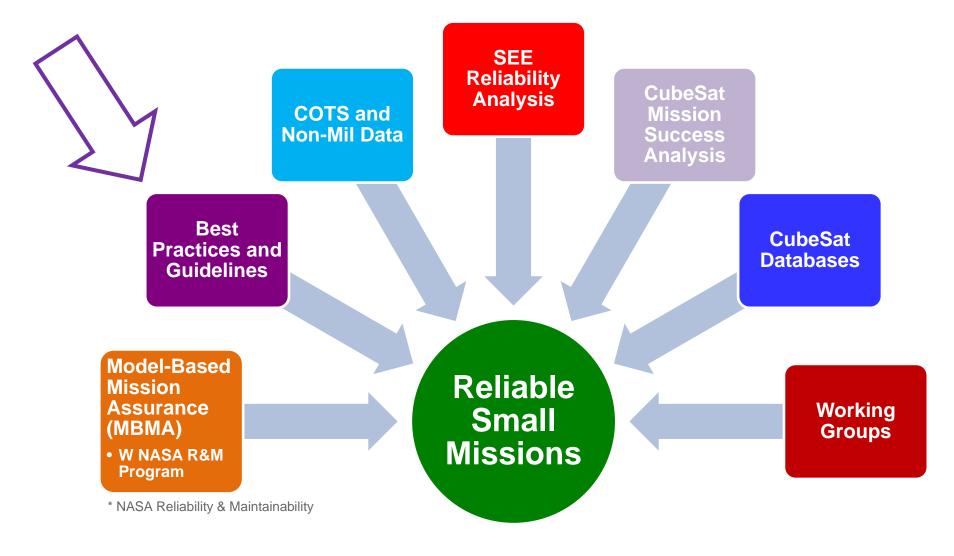


CME	Coronal Mass Ejection
COTS	Commercial Off The Shelf
DDD	Displacement Damage Dose
EEE	Electrical, Electronic, and Electromechanical
ELDRS	Enhanced Low Dose Rate Sensitivity
EP	Enhanced Performance
ESA	European Space Agency
GCR	Galactic Cosmic Ray
GOMAC	Government Microcircuits Applications and Critical Technologies Conference
GSFC	Goddard Space Flight Center
GSN	Goal Structuring Notation
HEART	Hardened Electronics and Radiation Technology
LEO	low earth orbit
LET	Linear Energy Transfer
MBMA	model based mission assurance
MRQW	Microelectronics Reliability and Qualification Workshop
NAND	Negated AND or NOT AND
NASA	National Aeronautics and Space Administration
NEPP	NASA Electronic Parts and Packaging
NEPP ETW	NASA Electronic Parts and Packaging (NEPP) Program Electronics Technology Workshop
NSREC	Nuclear and Space Radiation Effects Conference

RADECS	Radiation Effects on Components and Systems
RHA	Radiation Hardeness Assurance
SAA	South Atlantic Anomaly
SEE	Single Event Effects
	SEE-MAPLD Single Event Effects (SEE) Symposium/
SEE/MAPLD	Military and Aerospace Programmable Logic Devices (MAPLD) Workshop
SEGR	Single Event Gate Rupture
SEL	Single Event Latchup
SEP	Single Event Effects Phenomena (includes SEU, SEL, SEGR and SET)
SERESSA	School on the Effects of Radiation on Embedded Systems for Space Applications
SET	Single Event Transient
SEU	Single Event Upset
SLU	Saint Louis University
SwaP	Size, weight, and power
TID	Total Ionizing Dose
TID	Total Ionizing Dose
TMR	triple-modular redundancy
TNID	Total Non-Ionizing Dose
UV	Ultra-Violet

# NEPP Program- Small Mission Efforts





# Outline

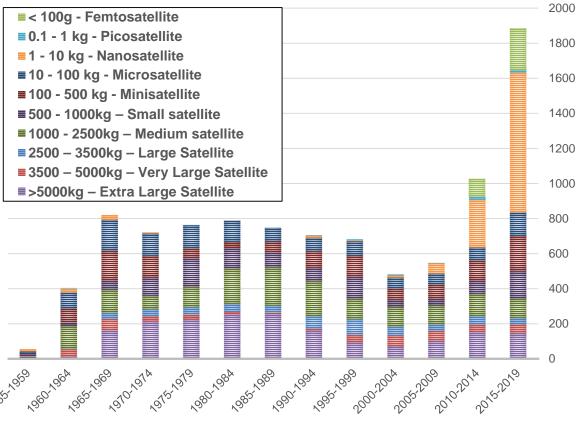


- New Space and SmallSat Considerations
- The Natural Space Radiation Environment Hazard
- Radiation Effects on Micro-Electronics
- Hardness Assurance, as a Discipline, with its Challenges
  - New Technologies
  - New Architectures
  - Unbound Risks
- Building Smart Requirements
- Risk Acceptance and Guidance

# New Space - New Point of View

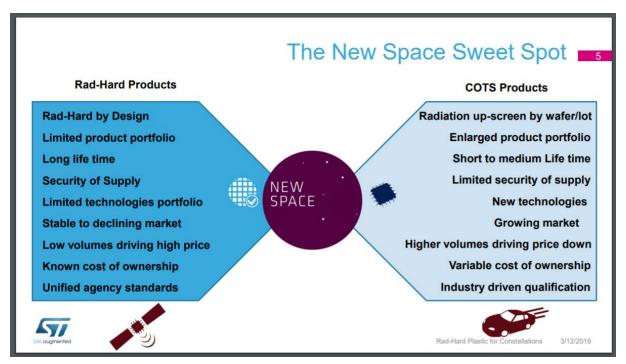


### **SmallSats Come in Many Sizes**



### Seradata SpaceTrak Data

### **Component Grades are Merging**



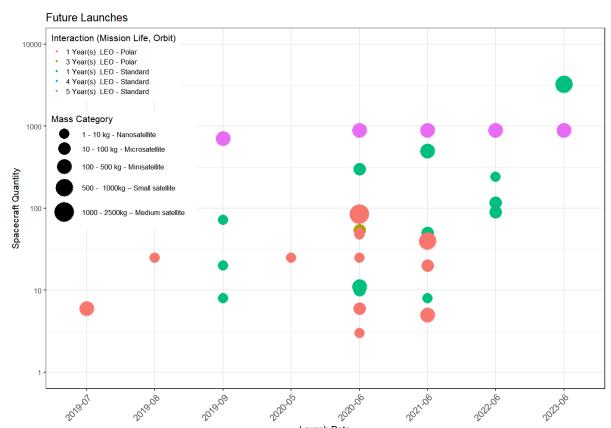
ESSCON: Eccofet

Risk acceptance is being used as a means to enable innovation

# New Space – Looking Ahead



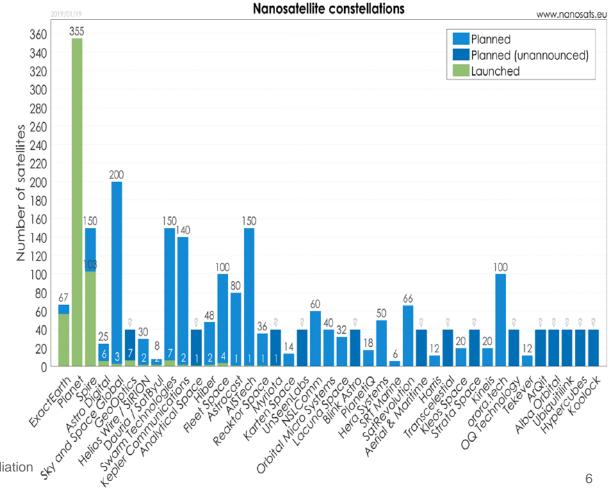
### **Constellations and Swarms**



### Seradata SpaceTrak Data (Notional Launches)

This presentation to be published on nepp.nasa.gov, originally presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16, 2019.

### **New Space = New Companies**



# New Space – Same Old Radiation

### New mission concepts and SmallSat paradigm

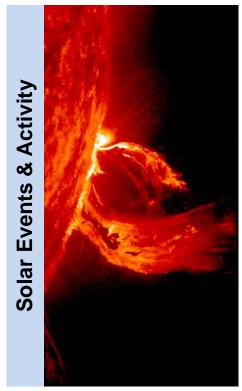
- Radiation challenges identified in the past are here to stay;
   adoption of new technologies are often the risk driver
- Commercial Space, Constellations, Small missions, etc. will benefit from detailed hazard definition and mission specific requirements

### The need for Radiation Hardness Assurance (RHA)

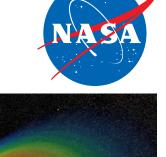
- Radiation effects are a mix of disciplines, evolve with technologies and techniques
- Misinterpretation of failure modes / misuse of available data can lead to over/under design
- RHA flow doesn't change, risk acceptance needs to be tailored

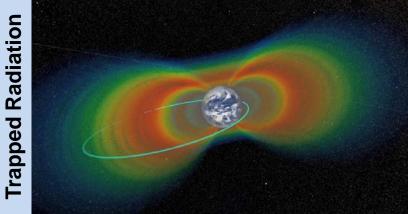
### Some Top Level Resources

- NPR-7120.5 NASA Agency Program Management
- GPR-8705.4 NASA Goddard Risk Classification Guidelines
- NASA-STD-8739.10 NASA Parts Assurance Standard

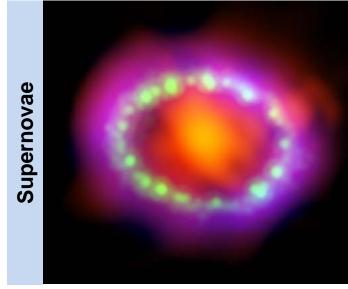








https://www.nasa.gov/van-allen-probes



NASA, ESA, and L. Hustak (STScI)

# Who Needs This Guidance?

### Universities / CubeSats

- May be first-time designers, or previous missions did not have requirements
- Schedule driven, limited time for development
- Rideshares could end up in multiple environments

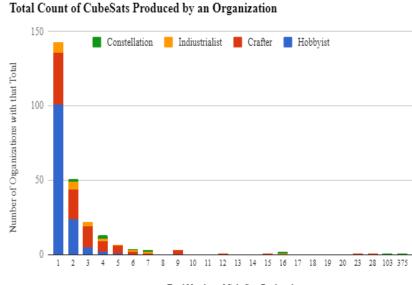
### Space Agencies / Government

- More compact designs in new destinations
- Cost savings of SmallSat platform, with more reliable outcome
- More willing to trade risk for capability

### Device / Subsystem Manufacturers

- Product / Device offerings: Space Plastic, EP, LeanRel, radiation tolerant, modified HiRel, etc.
- Fault tolerance in designs

# **CubeSat Metrics**



Michael Swartwout, SLU CubeSat Database

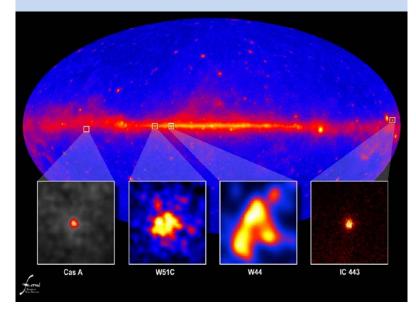


NASA's Goddard Space Flight Center/Bill Hrybyk

# Natural Space Radiation Environment

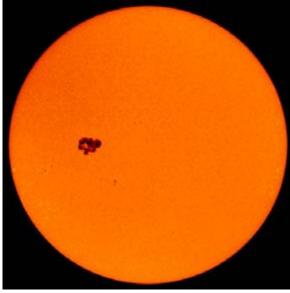


### **Galactic Cosmic Rays**



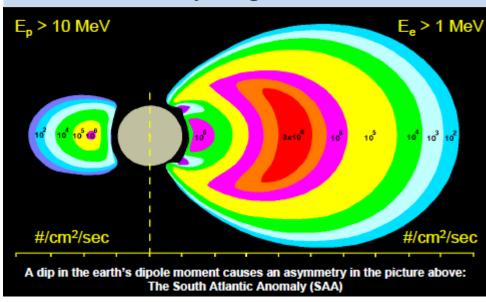
Energetic supernovae remnants (~GeV, Z=1-92)
Originate outside of our solar system

### **Solar Activity**



Solar Wind, Solar Cycle CMEs (proton rich) Flares (heavy ion rich)

# Trapped Particles in Planetary Magnetic Fields



Fluctuate with Solar Activity and Events
Not a perfect dipole
Protons and Electrons trapped at different
L-shell values and energies

# Natural Space Radiation Environment



Plasma

Particle Radiation

Neutral Gas Particles

- UV and X-Ray
- Orbital Debris

Degradation of micro-electronics

Degradation of optical components

Degradation of solar cells

Data corruption
Noise on images
System shutdowns or resets
Circuit Damage
Part tolerances exceeded

(After Barth)

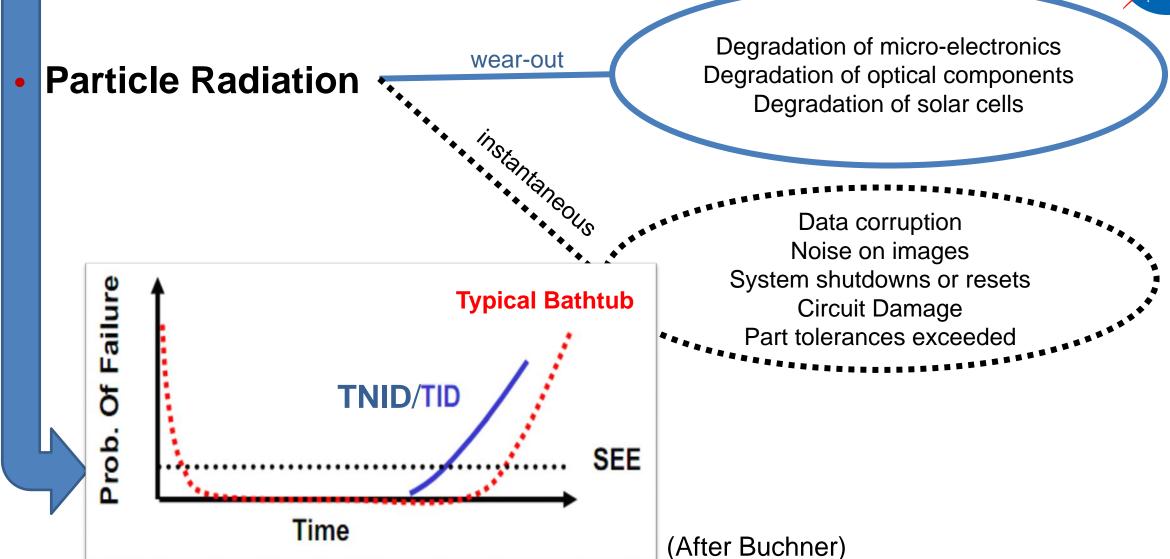
Spacecraft Charging, Ionizing Dose, Non-Ionizing Dose, Single Event Effects, Drag, Surface Erosion, Debris/Micro-Meteoroid Impacts, Thermal Cycles



wear-out

# Natural Space Radiation Environment





# Conventional Units Explanation



## **Degradation**

### Total lonizing Dose (TID)

- Absorbed dose (rad(Si))
  - 1 rad = 100 erg/g = 0.01 J/kg; 100 rad = 1 Gy
- Always specified for a particular material
   1 rad(SiO<sub>2</sub>), 10 krad(Si), 100 Gy(H<sub>2</sub>O)
- This is not exposure (R), or dose equivalent (Sv)

### Total Non-lonizing Dose (TNID)

- Fluence (particles/cm²)
   Number of particles per unit area
- Displacement Damage Dose (DDD)
   Specified at a given incident particle energy e.g., 10 MeV p+, 50 MeV p+, 1 MeV eq. neutrons, etc.

# **Single Event**

### Linear Energy Transfer (LET)

Stopping power normalized to target material

$$S = -\frac{dE}{dx} \Rightarrow \text{LET} = -\frac{1}{\rho} \frac{dE}{dx}$$

Units are MeV·cm<sup>2</sup>/mg

### Cross Section (σ)

- Device particle interaction (cm²)
- Used in calculation of rate
   Can be /device or /bit per time interval

# Degradation Contributors vs. Single Event



### **Cumulative effects**

- Depend highly on which contributors and duration in their presence
- Mimic wear-out/aging
- TNID and TID must be accounted for

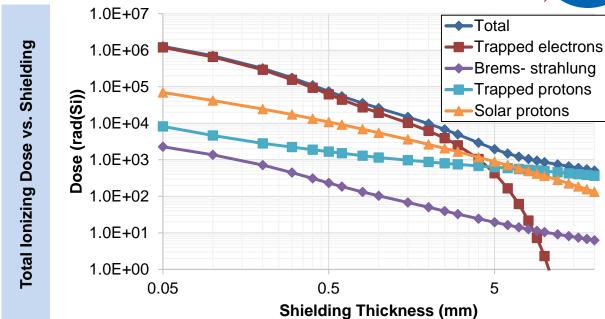
### Typical destinations (LEO, GEO)

- LEO at low altitude/inclination is more protected by the Geomagnetic field
- Proximity to the poles & SAA show a large variability in dose despite short mission durations
- Electrons and their braking radiation are the big offender in Geostationary orbits (don't forget about spacecraft charging...)

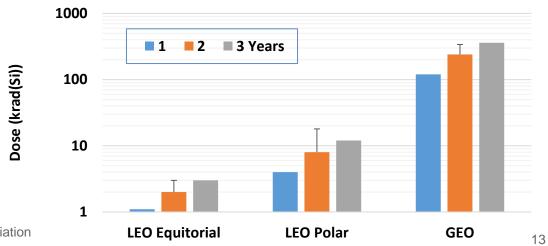
### Note that

- A little bit of shielding goes a long way
- Altitude plays a huge role when in/near the radiation belts (even transiting)
- Beyond Geomagnetic field, highly variable solar environment contributions (Solar cycle)

Degradation has a strong dependence on where you go, not just how long you are on orbit







# Degradation vs. Single Event Contributors



### One particle causes the effect

- Random in nature, particle must traverse sensitive structure within device and have sufficient charge creation along its path
- Shielding doesn't do so much for highly energetic particles
- Device technology can be dependent on particle species

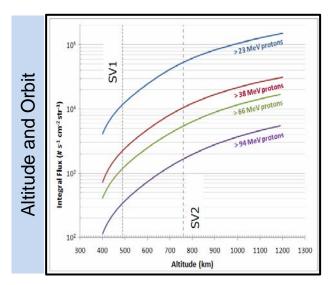
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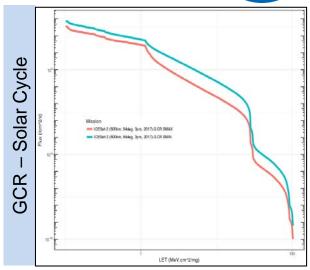
- Again altitude plays a role; for some devices that is a direct threat
- You are exposed to more GCR + Solar contribution as geomagnetic protection is reduced
- Natural phenomena like the South Atlantic Anomaly (SAA), magnetic poles, are temporal drivers

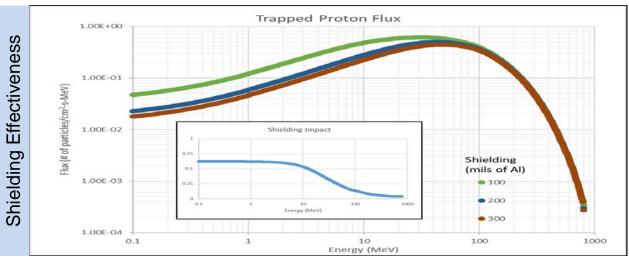
### Note that

 There will be a background rate, solar cycle dependence, solar event rate, increased rate for poles or SAA – not just one rate to consider

Single event contributors benefit very little from shielding, have dependence on where you are







# Summary of Environmental Hazards



	Plasma (charging)	Trapped Protons	Trapped Electrons	Solar Particles	Cosmic Rays	Human Presence	Long Lifetime (>10 years)	Nuclear Exposure	Repeated Launch	Extreme Temperature	Planetary Contaminates (Dust, etc)
GEO	Yes	No	Severe	Yes	Yes	No	Yes	No	No	No	No
LEO (low-incl)	No	Yes	Moderate	No	No	No	Not usual	No	No	No	No
LEO Polar	No	Yes	Moderate	Yes	Yes	No	Not usual	No	No	No	No
International Space Station	No	Yes	Moderate	Yes - partial	Minimal	Yes	Yes	No	Yes	No	No
Interplanetary	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	Yes	Yes	No	Yes	Maybe	No	Yes	Maybe
Exploration – Lunar, Mars, Jupiter	Phasing orbits	During phasing orbits	During phasing orbits	Yes	Yes	Possibly	Yes	Maybe	No	Yes	Yes

https://radhome.gsfc.nasa.gov/radhome/papers/SSPVSE05\_LaBel.pdf

# Mission Lifetime

# Radiation Hazard Contributors for Dose and SEE



### **Environment**

	LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
> 3 Years	Moderate Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	High Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
1- 3 Years	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
< 1 Year	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	Moderate Dose / High GCR, High Solar Proton Variability

# Radiation Effects on Active Microelectronic Devices



# Cumulative effects and single event effects can <u>both</u> be permanently damaging

- TID/DDD lead to wear-out of device operation and degrade devices beyond acceptable operations internally and externally
- Single Event Effects can be catastrophic instantaneously by turning on parasitic devices within the semiconductor or inducing electric field across dielectrics that eventually break down
- Synergistic effects can make ground based testing very difficult

### Destructive Single Event Effects (SEEs)

- Irreversible processes
- Terms: Latchup, Burnout, Gate Rupture

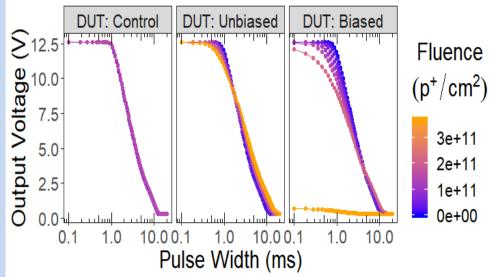
### Non-Destructive SEEs

- Lead to interruptions in operation and/or errors leading to unknown state spaces or loss of science / mission if not accounted for
- Terms: Functional Interrupt, Transients, Upsets

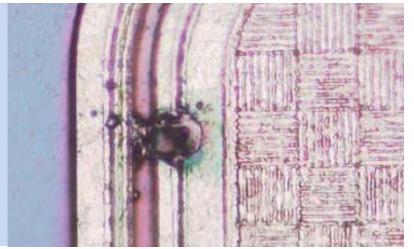
### • IEEE / Papers / Short Courses / Presentations

 GOMAC, HEART, MRQW, NEPP ETW, NSREC, RADECS, SEE/MAPLD, SERESSA, SPWG

# Degradation



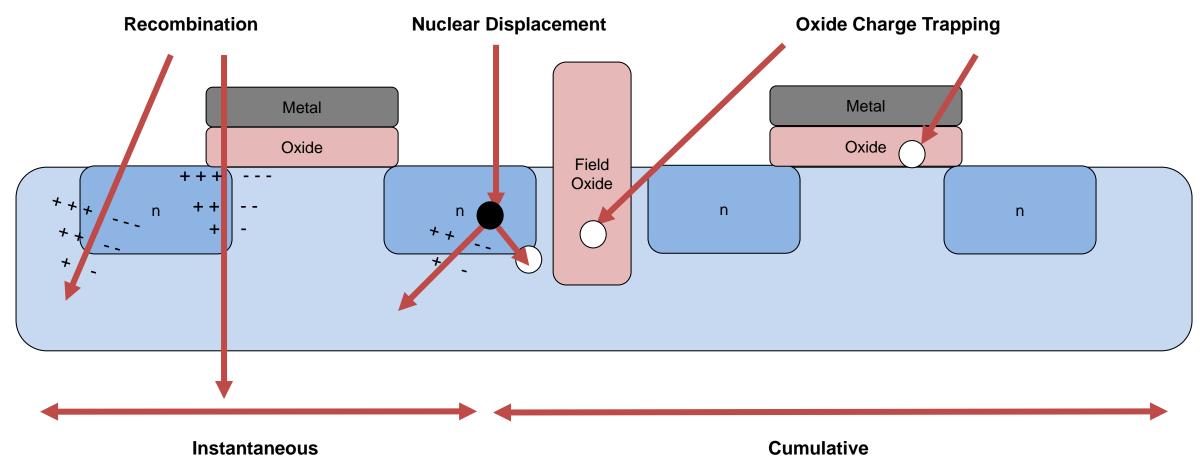
# Single Event



*Megan Casey* - https://nepp.nasa.gov/files/26196/2014-561-Casey-Final-Web-Pres-ETW-Diodes-TN16278 v2.pdf

# Device and Particle Interaction





Brock J. LaMeres, Colin Delaney, Matt Johnson, Connor Julien, Kevin Zack, Ben Cunningham Todd Kaiser, Larry Springer, David Klumpar, "Next on the Pad: RadSat – A Radiation Tolerant Computer System," Proceedings of the 31<sup>st</sup> Annual AIAA/USU Conference on Small Satellites, Logan UT, USA, Aug. 5-10, 2017, paper: SSC17-III-11, URL: <a href="http://digitalcommons.usu.edu/cgi/viewcontent.cgi?article=3618&context=smallsat">http://digitalcommons.usu.edu/cgi/viewcontent.cgi?article=3618&context=smallsat</a>

# Table of SEE Susceptibility



SEL	SEGR	SEB	SEDR	Stuck Bit	SEU/MCU	SET	SEFI
		POWER	One-time		Digital/bistable	bipolar	Complex
CMOS	MOSFET	MOSFET	Prog. FPGA	SRAM	technologies	technology	Microcircuits
			Bipolar			Analog	
Bipolar?	FLASH	Power JFET	Microcircuits	DRAM	Deep submicron	microcircuit	ADCs
	Schottky				CMOS more MCU	Digital	
	Diode	Power BJT		FLASH	susceptible	microcircuit	PWMs

Part-Level Consequences	How Common is Issue?
Catastrophic failure possible	Common in technology
Destructive but limited	Catastrophic failure possible
Nondestructive	Not seen but possible in principle

Ray Ladbury, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

List is not exhaustive, but new failure modes are found in new devices, so it would not be possible to capture all

# Outline



- New Space and SmallSat Considerations
- The Natural Space Radiation Environment Hazard
- Radiation Effects on Micro-Electronics
- Hardness Assurance, as a Discipline, with its Challenges
  - New Technologies
  - New Architectures
  - Unbound Risks
- Building Smart Requirements
- Risk Acceptance and Guidance

# The Job: Watch For the 'ilities'



- Must survive until needed
- Entire mission?
- Screening for early failures in components

### **Availability**

- Must perform when necessary
- Subset of time on orbit
- Operational modes
- Environmental response

### **Criticality**

- Impact to the system
- Part or subsystem function
- Mission objectives

### Reliability

- Resultant of all
- Many aspects and disciplines
- Known unknowns

# The People: Radiation Effects Engineers

### **Materials**

- Material Property degradations with radiation
- Energy loss in materials

### **Device Physics**

- Charge transport
- Device Process Dependencies
- Charge dependency of device operation

# **Electrical Engineering**

- Part to part interconnections
- Understanding circuit response
- Device functions and taxonomy

# Systems Engineering

- Requirements
- System Level Impacts
- Understanding interconnections
- Understanding functionality

### **Space Physics**

- Space weather
- Environment models/modeling
- Radiation Sources and variability



# Paths to Space Radiation



## Space Radiation Ecosystem

Systems Engineering Background Device Physics /
Electrical Engineering
Background

Space Weather
Physics
Background

- Radiation Reqs.Definition
- SPENVIS, OMERE, Fastrad, etc.
- Radiation Testing
   Management

- Radiation Testing +
   Qualification
- EEE Parts Programs

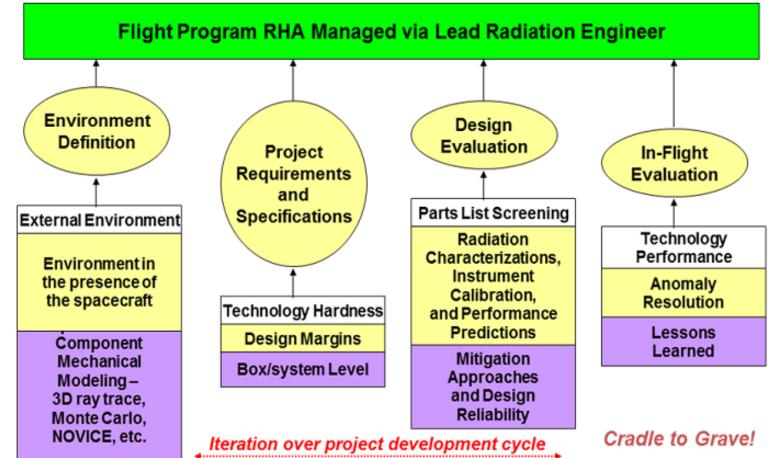
- Mission Scientists / PIs
- Model Developers (e.g. AP9/AE9)
- Often University + Research Lab based

After Whitney Lohmeyer, presented at JPL meeting 2019

# Radiation Hardness Assurance (RHA) Overview



RHA consists of all activities undertaken to ensure that the electronics and materials of a space system perform to their *design* specifications throughout exposure to the mission space environment



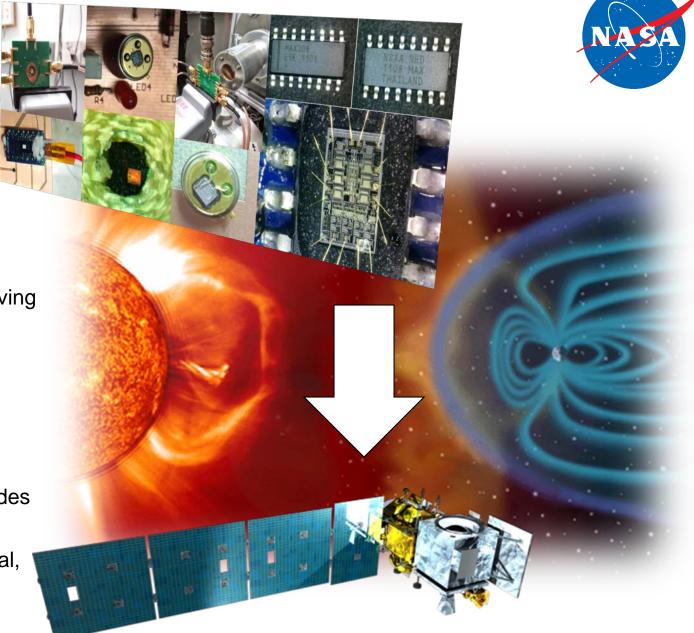
(After Poivey 2007)

(After LaBel 2004)

# RHA Challenges... Not So Small

- Always in a <u>dynamic</u> environment
- New Technologies
  - Device Topology / Speed / Power
  - Increased COTS parts / subsystem usage
- New Mission Architectures
  - Profiles of mission life, objective, and cost are evolving
  - Oversight gives way to insight in some mission classifications
  - Ground systems, do no harm, hosted payloads
  - Similarity and heritage data requirements widening
- Quantifying Risk
  - Translation of system requirements to radiation trades can be problematic
  - Determining appropriate mitigation level (operational, system, circuit/software, device, material, etc.)

Unbound radiation risks are likely



# New Technologies - New Susceptibilities



### Feature Size / Critical Charge

Sensitivity to muons? Low energy protons?

### 3D Stacking / Structures

- Deep sensitive volumes
- New materials within structure

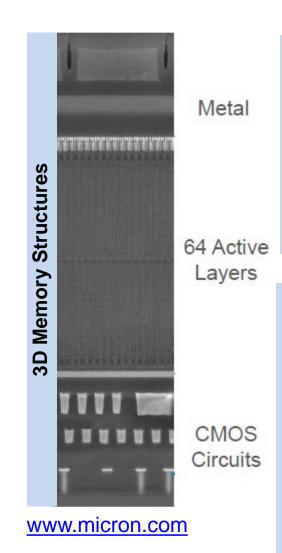
### Testing Challenges

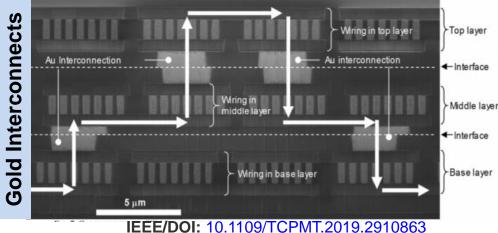
- Complexity (e.g., Systems-on-a-Chip)
  - Speed of interfaces
  - Obfuscation of state-space
- Flux / range of beam at facilities

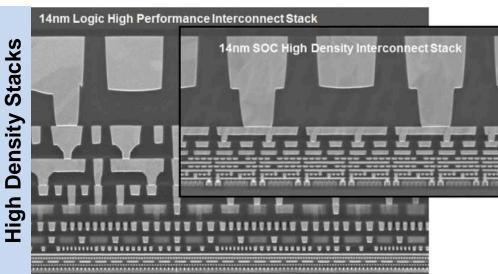
### Function

Integrated Photonics, MEMS, Hybrids

Without detailed part information you do not have certainty of the radiation threats



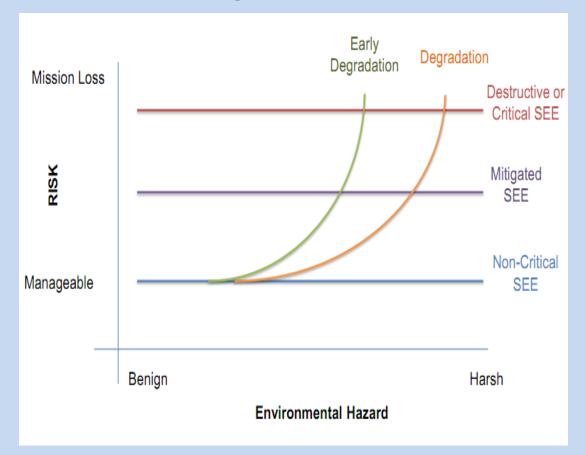




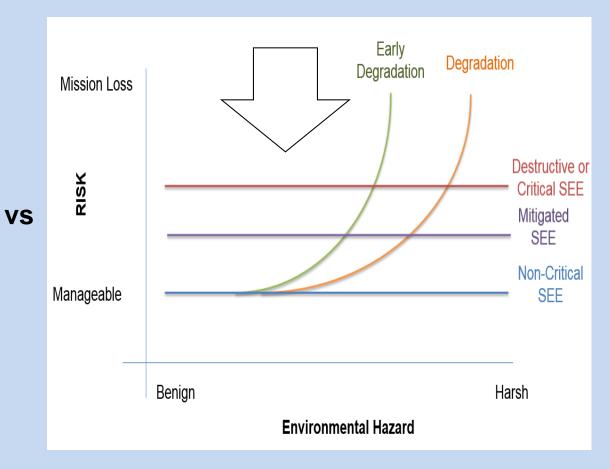
# New Mission Architectures - How Many to Succeed?



## Single Strain



## **Allowable Losses**



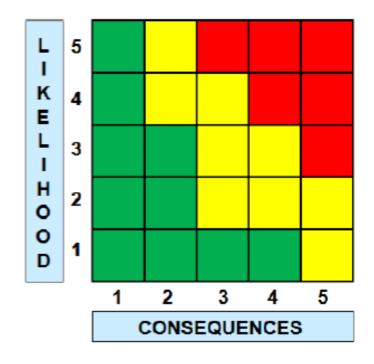
### Redundancy alone does not remove the threat, adds complexity

# New Challenges in Quantifying Risk



### From Risk Assessment section of NASA Program Management 7120.5

Likelihood	Safety Estimated likelihood of Safety event occurrence	Technical Estimated likelihood of not meeting performance requirements	Cost Schedule Estimated likelihood of not meeting cost or schedule commitment
5 Very High	$(P_{SE} > 10^{-1})$	$(P_T > 50\%)$	$(P_{CS} > 75\%)$
4 High	$(10^{-2} < P_{SE} \le 10^{-1})$	$(25\% < P_T \le 50\%)$	$(50\% < P_{CS} \le 75\%)$
3 Moderate	$(10^{\text{-3}}\!<\!P_{\text{SE}}\!\leq\!10^{\text{-2}})$	$(15\% < P_T \le 25\%)$	$(25\% < P_{CS} \le 50\%)$
2 Low	$(10^{-5} < P_{SE} \le 10^{-3})$	$(2\% < P_T \le 15\%)$	$(10\% < P_{CS} \le 25\%)$
1 Very Low	$(10^{-6} < P_{SE} \le 10^{-5})$	$(0.1\% < P_T \le 2\%)$	$(2\% < P_{CS} \le 10\%)$

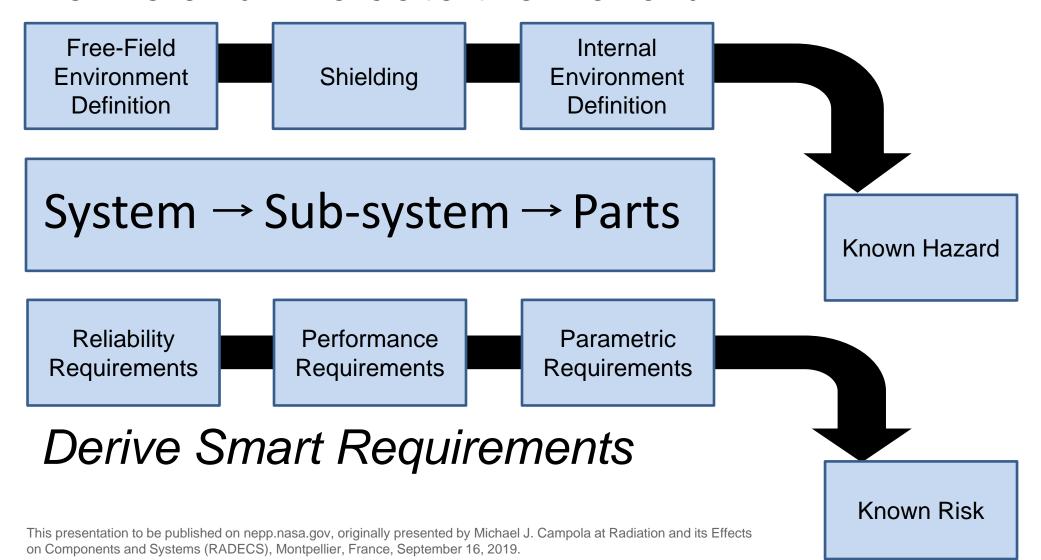


Can only get there with enough information about the system or the chosen device, need to have a known hazard and a known response

# RHA Building Blocks

# NA SA

# Define and Evaluate the Hazard



# Risks Abound, What is Critical?

### Parts

- Parametric degradation and leakage currents allowable in application?
- Downstream/peripheral circuits considered?
- Reset/refresh capability?
- Mitigation within too complex?
- Predicted radiation response unknown
   loss of part functionality critical?

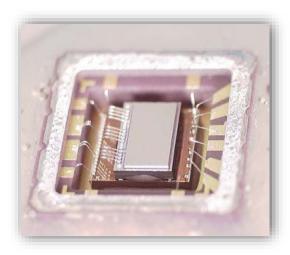
### Subsystem

- Functionally required to mission that the subsystem work?
- Interfaces allow you to get to a known state if all goes wrong?

### System

- Increased power dissipation a mission ender?
- Availability outweighed by error circumvention?
- Data retention through reboots? What if there is science data loss?
- Communications interruptions overwhelm?
- Navigation or Attitude determination unable to deal with faults?





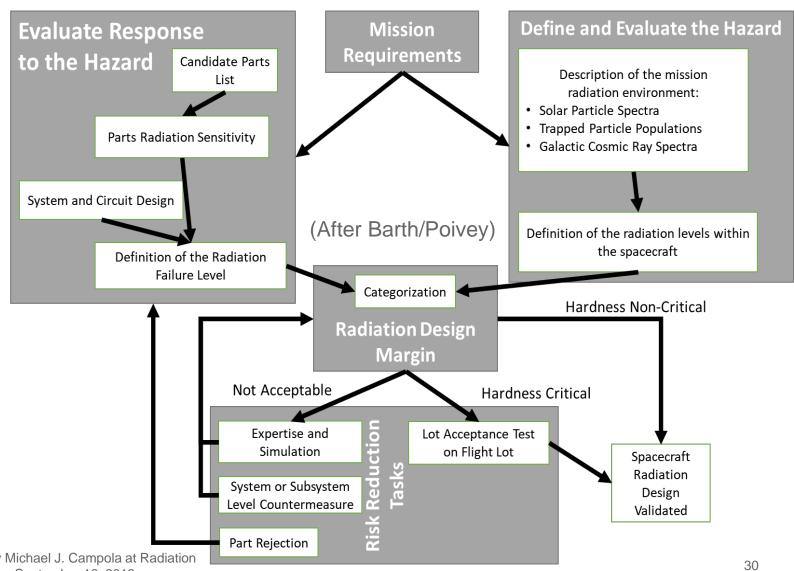
VS.



# RHA Flow Doesn't Change With Accepted Risk



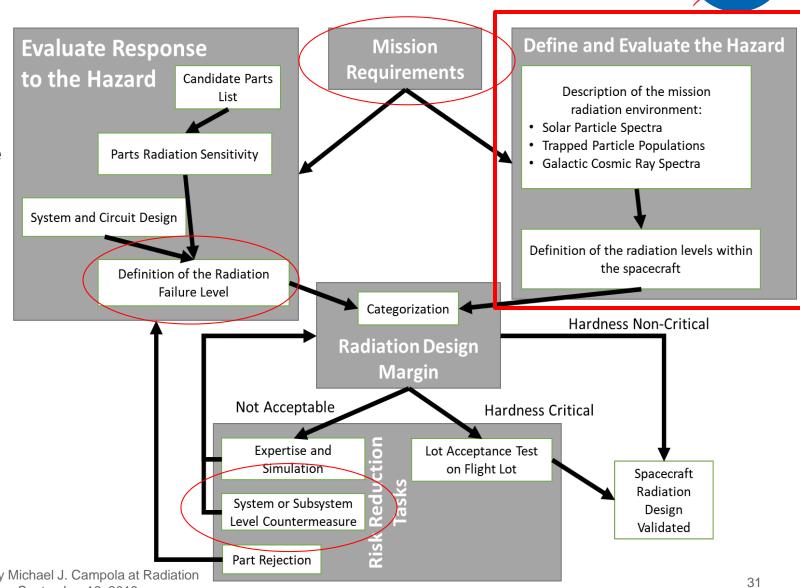
- Hardness Assurance is the practice of designing for radiation effects
- What it takes to overcome the radiation challenges
- Competing failure modes



# RHA Flow Doesn't Change With Accepted Risk



- Hardness Assurance is the practice of designing for radiation effects
- What it takes to overcome the radiation challenges
- Competing failure modes
- Focus for impact on risk acceptance:
  - Failure Awareness
  - Countermeasures/Mitigation
  - Mission Requirements



# Focus For Risk Acceptance



### Failure Awareness

- Know your hazard from the natural environment
- Know your devices potential failure mechanisms or response (data)

# Countermeasures and Mitigation

- Where are they necessary?
- At what level (part, card, box, mission)

# Smart Requirements – and Eventually Smart Trades

# Define and Evaluate the Hazard

- Define the Environment
  - External to the spacecraft
- Evaluate the Environment
  - Internal to the spacecraft
- Define the Requirements
  - Define criticality factors
- Evaluate Design/Components
  - Existing data/Testing
  - Performance characteristics
- "Engineer" with Designers
  - Parts replacement/Mitigation schemes
- Iterate Process
  - Review parts list based on updated knowledge

### **Environment Severity/Mission Lifetime**

		Low	Medium	High
Criticality/Availability	High	Manageable Dose / SEE impact to survivability or availability	Moderate Dose / SEE impact to survivability or availability	High Dose / SEE impact to survivability or availability
	Medium	Manageable Dose / SEE needs mitigation	Moderate Dose / SEE needs mitigation	High Dose / SEE needs mitigation
בוֹם	Low	Manageable Dose / SEE do no harm	Moderate Dose / SEE do no harm	High Dose / SEE do no harm

# Derive Smart Requirements

- Define the Environment
  - External to the spacecraft
- Evaluate the Environment
  - Internal to the spacecraft
- Define the Requirements
  - Define criticality factors
- Evaluate Design/Components
  - Existing data/Testing
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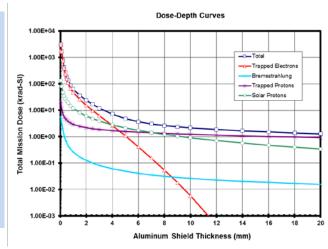
### **Environment Severity/Mission Lifetime**

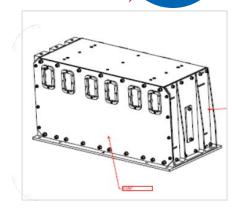
		Low	Medium	High
	High	Dose-Depth / Ray-trace GCR and Proton Spectra for typical conditions	Dose-Depth / Ray-trace GCR and proton Spectra for all conditions	Ray-Trace for subsystem / GCR and proton Spectra for all conditions
	Medium	Dose-Depth / GCR and proton spectra for background	Dose-Depth / GCR and Proton Spectra For background	Dose-Depth evaluation at shielding / All spectra conditions
เมา	Low	Similar mission dose, same solar cycle / GCR spectra	Dose-Depth / GCR spectra	Dose-Depth / GCR and Proton Spectra For background

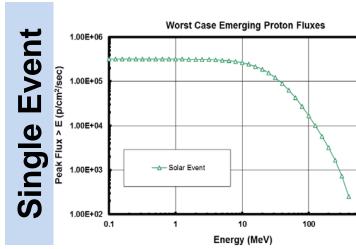
# Mitigation and Countermeasure Optimization

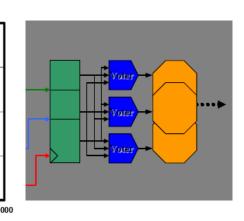
- Define the Environment
  - External to the spacecraft
- Evaluate the Environment
  - Internal to the spacecraft
- Define the Requirements
  - Define criticality factors
- Evaluate Design/Components
  - Existing data/Testing
  - Performance characteristics
- "Engineer" with Designers
  - Parts replacement/Mitigation schemes
- Iterate Process
  - Review parts list based on updated knowledge











K.A. LaBel, A.H. Johnston, J.L. Barth, R.A. Reed, C.E. Barnes, "Emerging Radiation Hardness Assurance (RHA) issues: A NASA approach for space flight programs," IEEE Trans. Nucl. Sci., pp. 2727-2736, Dec. 1998.

# **Building Requirements**



- Requirements by Environment
- Requirements by Technology
- Additional Considerations
  - LET Requirements for SEE
  - Dose Calculation
  - Operation During Flare Conditions
  - Radiation Data

### Requirements by Environment



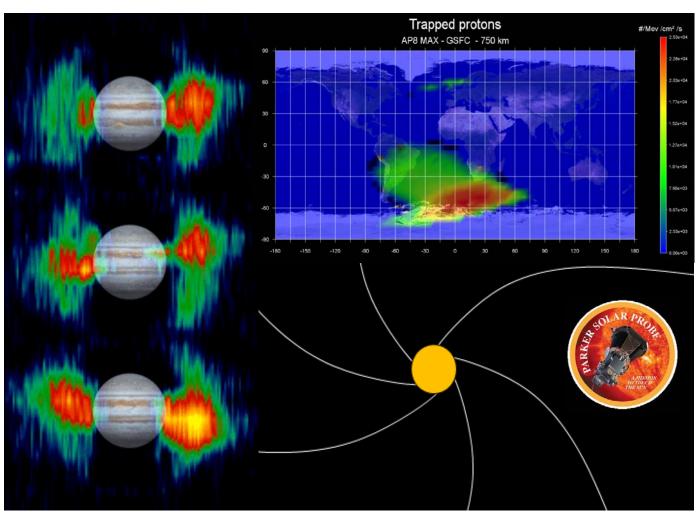
### Trapped Radiation Belts

- Can lead to high doses in a short mission:
   Jovian
- Can lead to spatially dependent SEE responses: South Atlantic Anomaly (SAA)

#### Heliocentric Orbits

- Solar Events, highly dynamic, energetic, directional
- Solar Wind, will depend on the solar cycle
- No planetary magnetic field attenuation

In essence the requirements are always driven by the environment, some more than others create a unique challenge



NASA JPL Cassini, <a href="http://saturn.jpl.nasa.gov">http://saturn.jpl.nasa.gov</a>,
Output from OMERE freeware <a href="http://www.trad.fr/en/space/omere-software/">http://www.trad.fr/en/space/omere-software/</a>

## Requirements by Technology



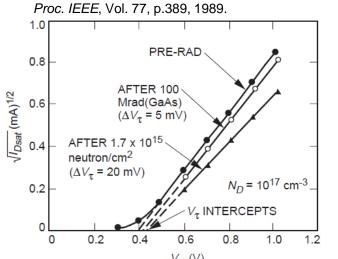
#### Technologies exhibit specific physics of failure

- Not easy to group them all
- Opto-electronics Displacement in the material
- Bipolar Enhanced Low Dose Rate Sensitivity
- Digital CMOS Latchup or SEFI
- Power devices SEGR/SEB
- Analog/Mixed-Signal Interruptions on PLLs,
   SERDES, clock dividers, etc.

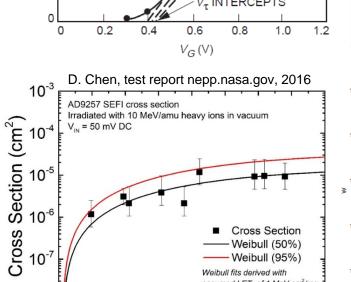
### Test Data requirements

- Failure distributions, often not enough parts
- Destructive effects are one data point,
   variability from part to part
- Statistics of the fit for rate calculations

Requirements should only be made applicable to the technologies that need to meet mission objectives and can benefit



R. Zuleeg, "Radiation Effects In GaAs FET Devices,"

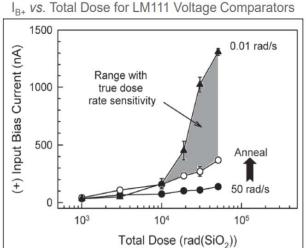


10

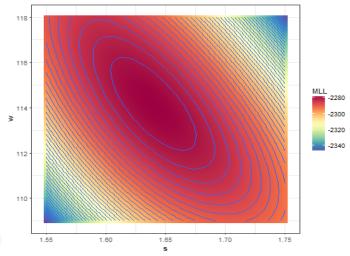
20

30

LET (MeV·cm<sup>2</sup>/mg)



M. R. Shaneyfelt, et al., IEEE TNS, 2000.



## Considerations for SEE Requirements



#### SEL

- Environment and technology driven, risk avoidance
- Protection circuitry / diode deratings

#### SEGR, SEB

- Effect driven, normally incident is usually the worst case
- Testing to establish Safe Operating Area (SOA)

#### SET

- Don't harm downstream parts via overvoltage/overstress on I/O, or accumulate over integrations
- Can be internal hybrids

#### SEU

- Tailored Filtering, EDAC, or Scrubbing
- MBU, MCU, SEFI, Locked States
  - Application Voltage or Pattern dependence
  - Watchdogs / reset capability

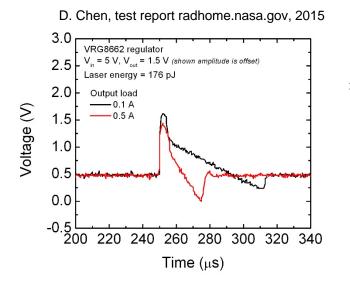
#### Proton SEE susceptible parts need evaluated in detail:

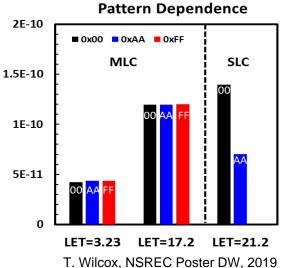
- Low-energy proton effects:
- May have direct ionization
- RHA for proton sensitivity update coming:

https://nepp.nasa.gov/files/25401/Proton\_RHAGuide\_NASAAug09.pdf

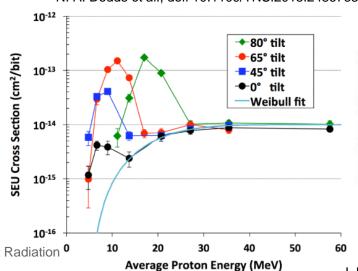
FPGA Mitigation Strategies (M. Berg)

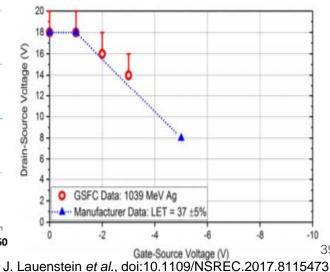
https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20180007760.pdf





N. A. Dodds et al., doi: 10.1109/TNS.2015.2486763





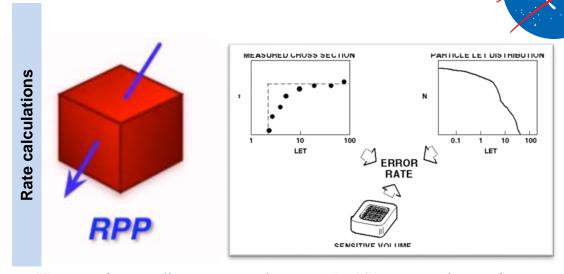
# Why You Can't Relax an LET Requirement

### Rate calculations are not the same for Destructive vs. Non-destructive

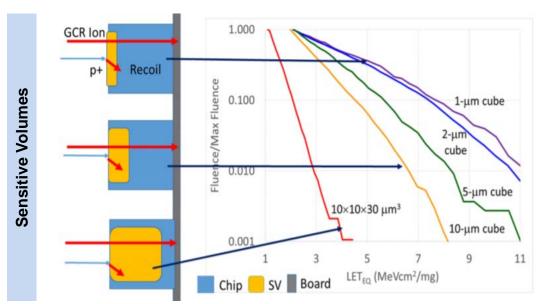
- Data are a limiting factor, one part = one data point
- For SEE types that exist in a given technology, they present a constant risk in time domain

### When you require by LET:

- Spectrum from environment is imparted on sensitive volumes, where we get LET thresholds (>75 vs. 60 vs. 37 MeV·cm²/mg)
- Effective LET increases with angle critical charge is what we are trying to determine
- CRÈME calculation integrates the two
- Deep sensitive volumes won't necessarily get same LET each time with monoenergetic beams



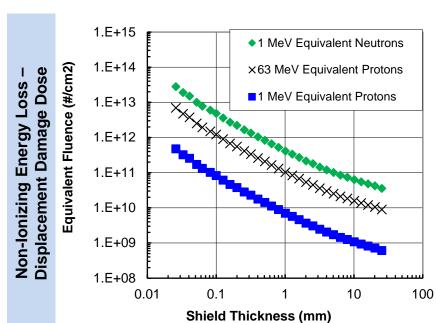
"Space Radiation Effects on Microelectronics," NASA Jet Propulsion Laboratory



Ray Ladbury, NSREC2017 SC, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

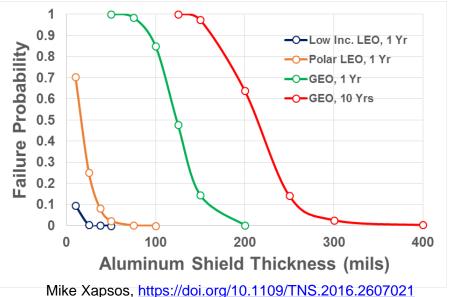
### Appreciable Mission Doses

- Maybe degradation of a part beyond usage is okay?
  - Criticality and Application
- Did you forget about DDD?
  - External materials are susceptible as well, polymers can be bad actors and are often on commercial ground based optical systems
- Even short missions can have a common failure mode
- Low mass budget?
  - Can optimize shielding if you have failure distribution of intended components



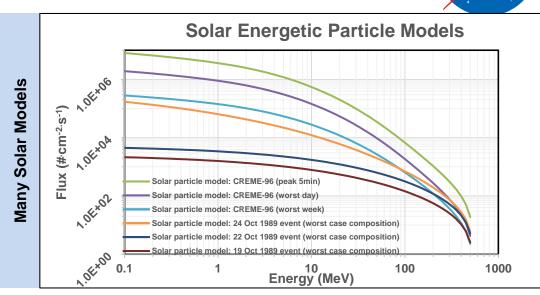


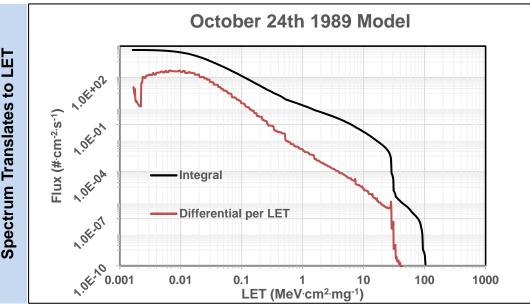
Shielding Optimization



# Operation During Flare Conditions: Think Availability

- Don't dose out during storm (nor the full mission)
  - Calculate the dose (TID/TNID) of the mission in full
    - 95% confidence level recommended
  - Calculate the dose contribution from N number of events (protons & x-rays), if dose from N is > 5% of the total dose, increase confidence level of full mission model
- Don't destructively fail from a single particle during the storm (nor the full mission)
  - Standard risk-avoidant SEE approach: no destructive effects allowed
  - LET threshold for single event latchup (SEL)
    - > 75 MeV.cm<sup>2</sup>/mg (some use 60 MeV.cm<sup>2</sup>/mg)
  - LET threshold for single event burnout, gate rupture, dielectric rupture (SEB, SEGR)
    - > 37 MeV.cm<sup>2</sup>/mg (particles must come from normal incidence to cause effect)
- If you have non-destructive single event upsets, they can't overwhelm critical instruments/systems during the storm
  - Rate calculation requires part data representative of the application, looking for crosssection over LET.
  - If parts' LET threshold from 20 to 75 MeV.cm<sup>2</sup>/mg, need heavy ion rate
  - If parts' LET threshold is below 20, need indirect ionization from recoil ions contribution to rate (need proton data) make sure packaging materials don't add to this, direct ionization from protons (can be built-in to heavy ion calculation) possible
  - Do you need to mitigate or not confirm that event rates are not higher than mitigation (Markov process... i.e. EDAC beats the number accrued, scrub rate is faster than critical number of upset accumulation)





## Risk Acceptance – Data Available?

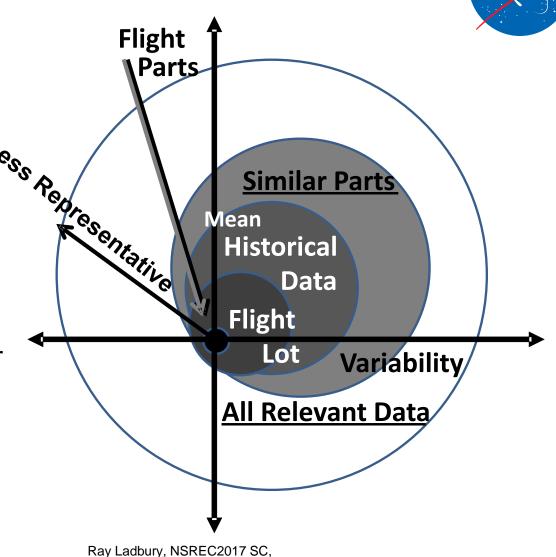
### Part Classifications Growing

- Mil/Aero vs. Industrial vs. Medical
- Automotive vs. Commercial vs. Modified HiRel

### Substitute COTS in this diagram

- Now you have another degree of separation
- Failure modes not fully understood
- Unlikely to have historical data
- Similarity data no applicable due to fab, process, or design rules
- Cost of testing usually too high

Without traceability you may be depending on nonrepresentative data.



https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

### Notional Radiation Data Collection Guidelines



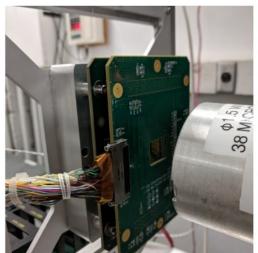
#### **Environment**

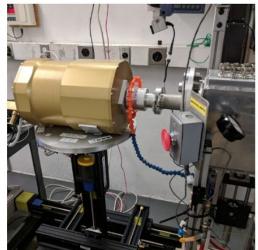
		LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
Mission Lifetime (With Assumed Risk Acceptance)	> 3 Years	Data on all SEE for critical parts, and have data on dose failure distribution on similar parts	Consider mission consequences of all SEE (Data for critical parts), have Dose failure distribution on lot	Have Data on all SEE, Have Data Dose failure distribution on lot
	1- 3 Years	Have Data on DSEE for critical parts	Consider mission consequences of all SEE (Data for critical parts), have data Dose failure distribution on similar parts	Have Data on all SEE for critical parts, Have Data on Dose failure distribution on similar parts
	< 1 Year	Look for data on DSEE for critical parts	Consider mission consequences of all SEE, and look for data on dose failure distribution on similar parts	Consider mission consequences of all SEE, and have data on dose failure distribution on similar parts

### When Do You Test? When Do You Model?

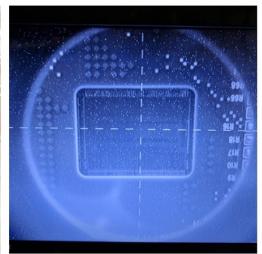


- Divine your risk threshold
  - Thére's a doc coming for that... radhome.gsfc.nasa.gov/nepp.nasa.gov
- Unknown failure modes that would not be acceptable to the mission
  - Known unknowns can be carried as a risk if you already know that the outcome is mitigated at the board or box level
  - New technologies should be identified early on
- Fault propagation may be the problem you wish to mitigate
  - This can include cumulative effects!
  - Fault injection may not be able to cover the state space
- Destructive single event effects are an obvious target
- Can you tolerate a part replacement in your design cycle?
  - Lead times, board re-spins, etc.

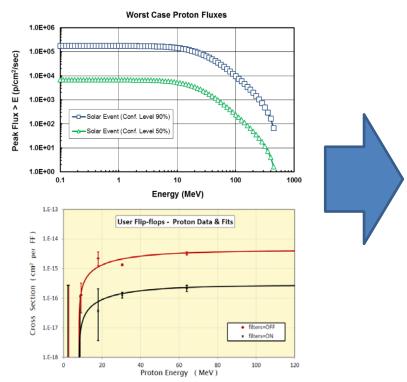


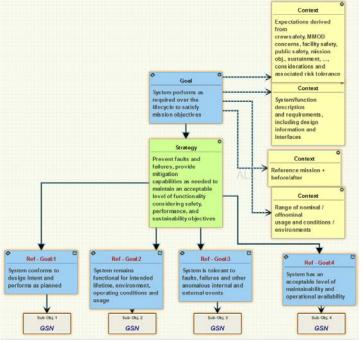


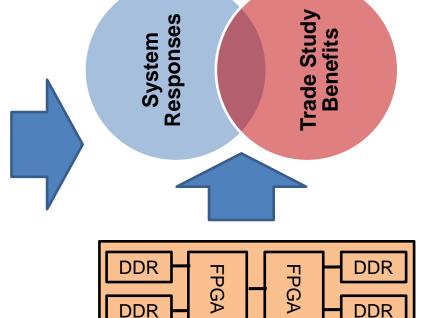




### Model Based Mission Assurance (MBMA) as a Tool







### **Environment, Device, & Design**

- Models and Test Data are brought together to get rates of upset / failure distributions
- Resources and Utilization are the scaling factors with criticality

#### **Goal Structuring Notation (GSN)**

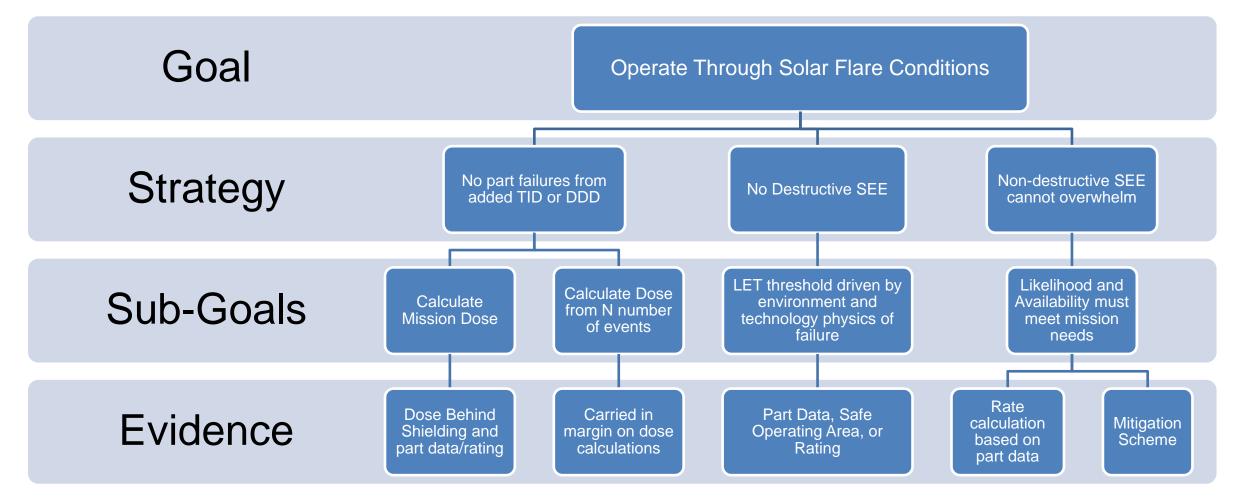
- Concept of operations
- Requirements and Availability are fed down correctly to subsystem
- Evidence is presented
- Assumptions are tracked

### **Systems Modeling Language**

- Description of System Connections and Dependencies
- Receives GSN readily
- Fault propagation can be identified

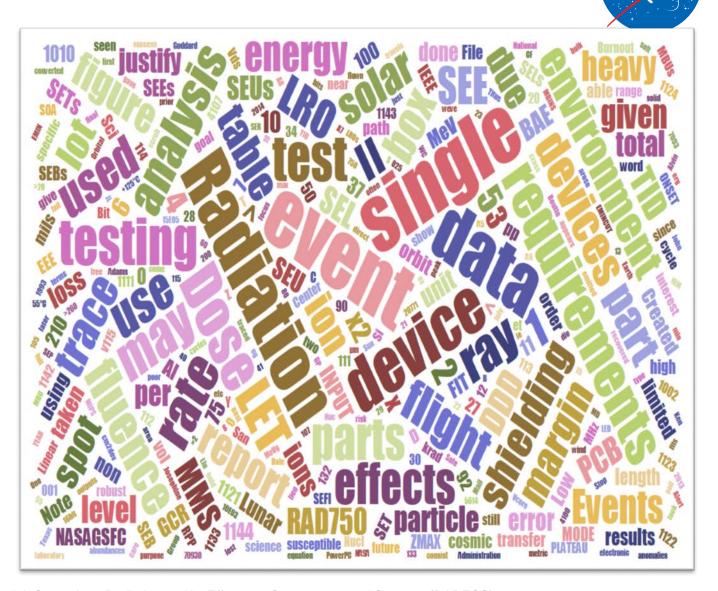
# Goal Structuring Notation (GSN)



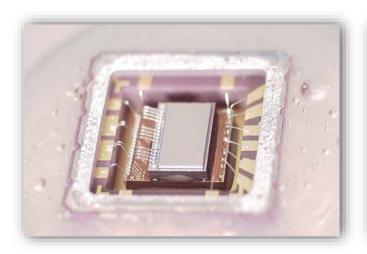


# Questions to Keep in Mind

- What are the radiation risks:
  - What is the hazard?
  - What are the challenges?
- What can you do to reduce the risk for a given hazard?
- What does changing that radiation environment mean for success?
- Need availability throughout the mission or at specific times?
- How do similar systems/devices react in the space environment?













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### **THANK YOU**