Enabling Entry Technologies for Ice Giant Missions

E. Venkatapathy^{1*}, A. Austin², A. Cassell¹, D. Ellerby¹, J. Elliott², P. Gage³, M. Lobbia², A. Nelessen², D. Prabhu⁴, B. Strauss², and P. Wercinsk¹, ¹NASA Ames Research Center, ² Jet Propulsion Laboratory, California Institute of Technology, ³ Neerim Corp., ⁴ AMA Inc.

Enabling Simultaneous Orbital and In Situ Measurements

- The highest priority science goals for Ice Giant missions are: 1) Interior structure of the Planet, and 2) Bulk composition that includes isotopes and noble gases.
- The interaction between the planetary interior and the atmosphere requires sustained global measurements. Noble gas and Isotope measurements require in situ measurement. Drag modulated aerocapture utilizing ADEPT offers more mass delivered to the Ice Giants than with propulsive orbit insertion.
- The Galileo Probe entered at a 'hot' spot which created interpretation challenges. Juno is providing valuable orbital measurements, but without in situ measurements the story is incomplete. Planetary scientists interested in Ice Giant missions should perform mission design studies with these new Entry System technologies to assess the feasibility within the context of the international collaboration framework.
- A mission architecture that includes probe(s) along with an orbiting spacecraft can deploy the probes at the desired location while taking simultaneous measurements from orbit to provide invaluable data that can correlate both global and local measurements.
- Entry System Technologies currently being developed by NASA are poised to enable missions that position the Orbiter & Probes through drag modulated aerocapture (ADEPT), and HEEET enables the Probes to survive the extreme environments encountered for entry into the atmospheric interior.







Updates to the IGS study (2017) provide guidance for design closure with fully matured HEEET system.

Ref: Space Science Review paper accepted for publication.



- No other options are currently available.
- Mars Sample Return Mission's Earth Entry Vehicle has baselined a variant of HEEET.
- HEEET capability will be sustained for missions in the coming decade.



Why Aerocapture?	D
What is Drag Modulated Aerocapture?	

rag Modulated Aerocapture with ADEPT

- The ADEPT drag skirt is integrated with the spacecraft, deployed prior to entry. The drag
- Ice Giant Mission Concepts with DMA
- Preliminary studies performed by JPL for both Uranus and Neptune show DMA can reduce trip

Aerocapture can provide greater mass efficiency and faster trip times, compared to propulsive orbit insertion, especially for Ice Giant Missions.

Strut

- In drag modulated aerocapture, the drag of the entry system is modified to achieve the desired velocity reduction for capture into orbit.
 - A deployed drag surface is ejected once velocity reduction is achieved during a single atmospheric pass

Drag Modulated Aerocapture with ADEPT

- DMA allowed the mission design to achieve high priority science goals and additional stretch goals
 - Orbiter and a Shallow Probe delivered postaerocapture
- Also included Deep Probe and 3 Landers Launch



٠

skirt is separated during aerocapture once the desired velocity reduction is achieved.

Simple and scalable



- Preliminary Trajectory & Stagnation Point Aerothermal Environments at Neptune
- 12 m (dia.) ADEPT skirt deploys a 4 m (dia.) spacecraft (1800 kg) Entry mass = 4000 kg, Entry Vel. (inert.) = 28 km/s; EFPA of. -10.9°; skirt jettison at 252 s and payload achieves an apoapsis ≈ 430000 km
- Peak aerothermal conditions: Pressure < 0.03 atm., Heat Flux < 200 W/cm², 40,000 J/cm², and peak deceleration < 70 m/s² (< 7g)

8

time and achieve mass efficient orbit insertion for a range of payload masses, including orbiter, probe(s) and landers to the moons in the

system.



DMA Capability is Scalable and Applicable Across the Solar System



DMA with ADEPT has the potential to deliver SmallSats and Large Missions to Venus and Mars, and Large missions to Saturn, Titan, Uranus and Neptune

- Using Falcon Heavy launch with Jupiter flyby and radioisotope electric propulsion, the spacecraft arrives at Uranus in ~9 years.
- Chemical propulsion feasible for a (12-13) year flight. Faster arrival will require a delta-v of ~5 km/s (for a ~9 year flight), likely not possible. Aerocapture can shorten

flight time for science mission by 3-4 years.



Remarks

Notional Uranus DMA

Concept Utilizing ADEPT

- ADEPT and Drag Modulated Aerocapture are currently being matured in a partnership between NASA Ames, JPL and CU Boulder.
- The goal is an earth-based flight demonstration for DMA in the next few years to be ready for Ice Giant and Venus missions.
- Venus DMA mission will focus on a rideshare SmallSat mission opportunity to deliver science.
- Ice Giant missions will focus on delivering a spacecraft with probes and landers, faster and more efficiently, enabling greater science.

Concluding Remarks

- The coming decade presents a once in a generation opportunity for exploring the Ice Giants.
- The HEEET & ADEPT technologies are nearing maturity and should be considered by mission planners to enhance science return. ٠
- Drag modulated aerocapture utilizing ADEPT offers more mass delivered to the Ice Giants than with propulsive orbit insertion.
- If the Tempest (or similar) mission concept can be performed within the cost & risk constraints, it will provide extremely valuable and unprecedented scientific data. •
- Planetary scientists interested in Ice Giant missions should perform mission design studies with these new Entry System technologies to assess the feasibility within the context of the international collaboration framework.