#### NASA/SP- 20205006016



#### **Summer 2020 Intern Presentation**

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National Aeronautics and Space Administration

NASA Ames Research Center

August, 2020



# Multi-Objective Flight Control of High Aspect Ratio Flexible Wing Aircraft

Spring – Summer 2020

Blake Hartwell NASA ARC Intern (USRA)

Dr. Nhan Nguyen Senior Research Scientist and Technical Group Lead – TI Division, ACES Group



**Energy Consumption** 

2005 best - 60%

5. Hybrid Gas-

Electric

Propulsion

Advanced Air Transport Technology (AATT) [2]

2005 best - 80%

SX/PX

Rim

PM

4. Cleaner.

**Compact Higher** 

**BPR** Propulsion

Bore

1300F

7. Alternative

Fuel

Emissions

# **Introduction** Project Scope

#### Aeronautics Research Mission Directorate [1]



with Intestments in bran Near-Term Tech

Challenges and Long-Term (2030) Vision

Next-generation aircraft designed with higher aspect ratio wings to improve aerodynamic efficiency





## Background Flexible Wing Aircraft

- Significant weight savings using advanced composite technology enables high aspect ratio wing designs.
- Higher aspect ratio and reduced material stiffness increases wing flexibility.





## Background Flexible Wing Aircraft

### Higher wing flexibility impacts...

#### Aerodynamics

 Increased wing deflection at off-design flight conditions causes suboptimal lift distribution – i.e., increased drag, higher fuel consumption



#### Flight Loads, Ride Quality, and Stability

 Increased wing flexibility exacerbates gust/turbulence response and decreases stability margins





# Performance Adaptive Aeroelastic Wing (PAAW) Technology

#### Variable Camber Continuous Trailing Edge Flaps (VCCTEF) [5]

 Developed by NASA and Boeing Research & Technology to provide flexible high aspect ratio wings with aeroelastic compensation





#### Flap segment cross-section

- Continuous trailing edge enables spanwise tailoring of lift distribution
- Multi-segment variable camber enables chordwise tailoring of pressure distribution on upper and lower wing surfaces



# **Multi-Objective Flight Control**

- Systems with multi-functional flight control surfaces use a multi-objective flight control (MOC) design.
- MOC is an optimal control design that accounts for multiple competing flight objectives when formulating
  optimal control solutions, all while providing ideal compromises between such requirements.





## **Objectives** Spring-Summer 2020

 Design a LQG multi-objective flight controller for the Common Research Model (CRM 13) provided by Boeing Research & Technology.

• Implement the flight controller in a nonlinear, 6DoF Simulink model provided by Boeing.

 Run simulations to assess MOC performance metrics, and validate merits associated with flight objectives.



# **Common Research Model (CRM13)**

777-like configuration with a 13.5 aspect ratio developed by Boeing Research & Technology [6]

- Distributed control surfaces
  - 17 control surfaces/wing, 1 elevator, 1 rudder
- Advanced sensing approaches
  - Distributed accelerometers: 38/wing + 5 fuselage

RIBAI

ROBFlar

ROBMPE

ROBMPF1

OBFlap

ROBMPE

ROBMP

**ROBAIL1** 

ROBAIL

ROBMPF4

LIDAR-based wind sensors for gust estimation

RIBFlap

RIBMPF1



# Plant Description of Linear CRM13 Model

Gust-Disturbed, ASE State Space Representation

$$\begin{bmatrix} \dot{x_r} \\ \dot{x_e} \end{bmatrix} = Ax + Bu + w = \begin{bmatrix} A_{rr} & A_{re} \\ A_{er} & A_{ee} \end{bmatrix} \begin{bmatrix} x_r \\ x_e \end{bmatrix} + \begin{bmatrix} B_{ru_r} & B_{ru_e} & B_{rd} \\ B_{eu_r} & B_{eu_e} & B_{ed} \end{bmatrix} \begin{bmatrix} u_r \\ u_e \\ d = \begin{bmatrix} g \\ \dot{g} \end{bmatrix} \end{bmatrix}$$

Plant Output from Full Rigid Body Access & Accelerometers

$$y = \begin{bmatrix} x_r \\ a \\ n \\ P \end{bmatrix} = \begin{bmatrix} I & 0 \\ a_{x_r} & a_{x_e} \\ n_{x_r} & n_{x_e} \\ P_{x_r} & P_{x_e} \end{bmatrix} \begin{bmatrix} x_r \\ x_e \end{bmatrix} + \begin{bmatrix} 0 & 0 & 0 \\ a_{u_r} & a_{u_e} & a_d \\ n_{u_r} & n_{u_e} & n_d \\ P_{u_r} & P_{u_e} & P_d \end{bmatrix} \begin{bmatrix} u_r \\ u_e \\ d = \begin{bmatrix} g \\ \dot{g} \end{bmatrix} \end{bmatrix}$$

Load Responses  $P = P_x x + P_u u + P_d d$ 

Fuselage Acceleration Measurements  $a = a_x x + a_u u + a_d d$ 

Wing Acceleration Measurements  $n = n_x x + n_u u + n_d d$ 

Drag Model

$$\Delta C_{D} = C_{D_{x}}x + C_{D_{u}}u + x^{T}C_{D_{x^{2}}}x + x^{T}C_{D_{xu}}u + u^{T}C_{D_{u^{2}}}u$$



F47 7

# **Observer Designs for Full-State Estimation**

Simple Kalman Filter

$$\begin{bmatrix} \dot{\widehat{x}_r} \\ \dot{\widehat{x}_e} \end{bmatrix} = \left( \begin{bmatrix} A_{rr} & A_{re} \\ A_{er} & A_{ee} \end{bmatrix} - L \begin{bmatrix} I & 0 \\ a_{x_r} & a_{x_e} \\ n_{x_r} & n_{x_e} \end{bmatrix} \right) \begin{bmatrix} \widehat{x_r} \\ \widehat{x_e} \end{bmatrix} + \left[ \left( \begin{bmatrix} B_{ru_r} & B_{ru_e} & B_{rd} \\ B_{eu_r} & B_{eu_e} & B_{ed} \end{bmatrix} - L \begin{bmatrix} 0 & 0 & 0 \\ a_{u_r} & a_{u_e} & a_d \\ n_{u_r} & n_{u_e} & n_d \end{bmatrix} \right) L \right] \begin{bmatrix} u_r \\ u_e \\ \hat{d} \\ v^* \end{bmatrix}$$

• The gust estimate  $\hat{d}$  is obtained from LIDAR measurements available in the Boeing simulation model

Extended State Kalman Filter with Gust Estimation

$$\frac{d}{dt} \begin{bmatrix} \hat{x} \\ \hat{d} = \begin{bmatrix} \hat{g} \\ \hat{g} \end{bmatrix} \end{bmatrix} = \begin{bmatrix} A & B_d \\ 0 & 0 & 1 \\ 0 & 0 & 0 \end{bmatrix} \begin{bmatrix} \hat{x} \\ \hat{d} \end{bmatrix} + \begin{bmatrix} B_{u_r} & B_{u_e} \\ 0 & 0 \\ 0 & 0 \end{bmatrix} \begin{bmatrix} u_r \\ u_e \end{bmatrix} + L(y - \hat{y})$$
$$\hat{y} = \begin{bmatrix} C & D_d \end{bmatrix} \begin{bmatrix} \hat{x} \\ \hat{d} \end{bmatrix} + \begin{bmatrix} D_{u_r} & D_{u_e} \end{bmatrix} \begin{bmatrix} u_r \\ u_e \end{bmatrix}$$

**Recursive Least Squares Method for Gust Estimation** 

$$\dot{\widehat{w}_1} = I_r \dot{\widehat{w}_r} = -\Gamma_r \left( G_r^{-1} \varepsilon_r + \sigma_r \widehat{w_1} \right)$$

$$\dot{w_2} = I_e \dot{w_e} = -\Gamma_e \left( a_{w_e}{}^T \left( a_{w_e} a_{w_e}{}^T \right)^{-1} \varepsilon_e + \sigma_e \widehat{w_2} \right)$$



# **Multi-Objective Optimal Control Formulation**





# **Iterative Weights Selection**

• Iterative MOC tuning performed by programming gain calculations inside a FOR loop and using a random number generator to assign cost function weights for each iteration.





# **Flight Simulations**

- Unstable Flight Conditions (1.6g Pitch-Up Maneuver)
  - Alt. = 21,256 ft
  - Weight = 383,546 lb (Fuel Empty)
  - M = 0.88



#### Discrete Gust Encounters

- Alt. = 36,983 ft
- Weight = 581,497 lb
- Mach = 0.85

#### Turbulence Encounter

- Alt. = 36,983 ft
- Weight = 581,497 lb









## Flight Simulations Unstable Flight Conditions



#### **Flight Conditions**

- Alt. = 21,256 ft
- Weight = 383,546 lb (Fuel Empty)
- M = 0.88

MOC Design V\_0.1 – ASE Mode Suppression ON Modal Rate:  $\dot{q}_e = \frac{1}{\zeta}$ Modal Deflection:  $q_e = \frac{\omega^2}{\zeta}$  $Q_P = \begin{bmatrix} 0; & 0; & 0 \end{bmatrix}$  $Q_n = \begin{bmatrix} 0; & 0; & 0 \end{bmatrix}$  $Q_D = 0$ 

• Unstable poles shifted to left side of s-plane when applying ASE mode suppression.



## Flight Simulations Unstable Flight Conditions





## Flight Simulations Discrete Gust Encounters



#### **Flight Conditions**

- Alt. = 36,983 ft
- Weight = 581,497 lb
- $\circ$  Mach = 0.85

**Design V\_4.2 – Gust Load Alleviation ON**   $Q_P = [0; 60; 10]$   $Q_n = [0; 0; 0]$  $Q_D = 0$ 

 Gusts estimated using onboard LIDAR wind sensors



# Flight Simulations Discrete Gust Encounters

Design V	/_4.2 – Gust Load Alleviation ON	
$Q_{P} = [0;$	60; 10]	
$Q_n = [0;$	0; 0]	
$Q_D = 0$		

	Baseline	With MOC	Impact (%)
Max WRBM (ft-lb)	5.896E+06	4.674E+06	-20.72
Max WRTM (ft-lb)	4.854E+05	3.807E+05	-21.57
Max WRSF (lb)	5.754E+03	4.403E+03	-23.48
Max $n_z$ at CG (g's)	1.573	1.549	-1.50
Max $n_z$ Left Wing (g)	1.658	1.633	-1.50
Max $n_z$ Pilot Seat (g)	2.015	1.899	-5.76
Avg. L/D Ratio	21.493	21.834	1.59





## Flight Simulations Turbulence Encounter





## Flight Simulations Turbulence Encounter

Design V_4.2 – Acceleration Suppression						
ON						
$Q_P = \begin{bmatrix} 0; & 0; & 0 \end{bmatrix}$						
$Q_n = [5; 5; 5]$						
$Q_D = 0$						

	Baseline	With MOC	Impact (%)
Max WRBM (ft-lb)	3.296E+07	3.126E+07	-5.16
Max WRTM (ft-lb)	7.630E+06	7.165E+06	-6.09
Max WRSF (lb)	4.024E+04	3.784E+04	-5.96
Max $n_z$ at CG (g's)	2.154	2.159E+00	0.24
Max $n_z$ Left Wing (g)	2.100	2.061E+00	-1.83
Max $n_z$ Pilot Seat (g)	2.062	2.050	-0.54
Avg. L/D Ratio	17.012	17.258	1.45





## Conclusions

- Multi-objective optimal control provides high aspect ratio flexible wing aircraft with aeroelastic compensation.
- Aeroservoelastic mode suppression stabilizes wing flutter modes during unstable aircraft flight conditions.
- Gust load alleviation provided >20% reductions in maximum wing loading during discrete gust encounters
- Gust load alleviation is more effective for discrete gust encounters, acceleration suppression is more effective for turbulence encounters.



# Acknowledgments

- **Dr. Nhan Nguyen** Senior Research Scientist and Technical Group Lead of Advanced Control and Evolvable Systems Group, Intelligent Systems Division, NASA Ames
- Dr. Kelley Hashemi Aerospace Engineer, Intelligent Systems Division, NASA Ames
- Dr. Jack Quindlen GNC&A Engineer, Boeing Research & Technology
- Universities Space Research Association



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# **Questions?**

# **Observations of Methane towards Orion IRc2 with SOFIA/EXES**

National Aeronautics and Space Administration

# ABSTRACT

We present high-resolution (R~60,000) MIR detections of methane in the interstellar medium. Our observations were obtained with the Echelon-Cross-Echelle Spectrograph aboard the Stratospheric Observatory for Infrared Astronomy and are the first of CH4 towards the Orion IRc2 hot core. Exceptional resolution distinguishes 6 individual rovibrational transitions spanning 7.5 to 7.66µm We resolve the primary velocity component for CH4 as  $-7.7 \pm 0.5$  km/s with excitation temperature as  $210 \pm 89$  K and column density as  $2.07 \pm 0.51 \times 10^{17}$  cm<sup>-2</sup>We also detect a secondary velocity component but our analysis is incomplete. These results give insight into the basic principles of astrochemical networks in the ISM and paired with observations of other chemical species, could provide exciting proto-planetary tracers.







# Jose S. Monzon<sup>1</sup>, Naseem Rangwala<sup>2,3</sup>, Sarah Nickerson<sup>2,3</sup>

{1} University of California, Santa Cruz; 1156 High St., Santa Cruz, CA 95064, USA {2} Space Science and Astrobiology Division, NASA Ames Research Center, Moffet Field, CA, 94035 USA **{3}** Bay Area Environmental Research Institute, Moffet Field, CA, 94035, USA

- Observations of Orion IRc2 with the **SOFIA** observatory were made at altitudes from about 42000 to 44000ft. The EXES slit is shown with respect to photometric observations from FORCAST<sup>[1]</sup>.
- Spectra were acquired in the crossdispersed high resolution mode with a slit width of 3.2" giving a resolving power of about 5 kms<sup>-1</sup>.
- Several different wavelengths settings were observed for the larger scale inventory project, but we identified CH<sub>4</sub> in the setting which is centered at 7.58 $\mu$ m or 1318 cm<sup>-1</sup> <sup>[2]</sup>.



# REFERENCES

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[4] - Goldsmith, P. F., & Langer, W. D. 1999, ApJ, 517, 209, doi: 10.1086/307195 [5] - Boogert, A.C.A., Blake, G.A., and Oberg, K. (2004) Astrophys. J. 615 344-353.



As some atmosphere lines are mixed in with our observations, we divided out a smooth atmospheric

• We then ran a peak finding algorithm to select strong absorption lines and fit gaussians to the transitions. All orders required a double gaussian to account for the two velocity components.

• Using the HITRAN database and its Python interface, HAPI, we were able to determine the column

• We assumed all molecules are in local thermal equilibrium (LTE) such that they follow the **Boltzmann distribution**<sup>[4]</sup>. **Our linear fit details** the excitation temperatures and column densities for CH<sub>4</sub>.

As we are the first to measure CH<sub>4</sub> towards IRc2, our results are not directly comparable to other studies. However, Boogert et al. 2004<sup>[5]</sup> studied CH<sub>4</sub> towards massive protostar NGC 7538 and measured similar temperatures ( $T_{ex} \sim$ 90K) but smaller column densities (N ~ 1.0 x  $10^{16} \,\mathrm{cm}^{-2}$ ).

**CH<sub>4</sub>** is proposed to be the starting point of a rich chemistry leading to the complex organic molecules<sup>[5]</sup> and could even help us understand how biology develops on other worlds.

**Our next steps include reassessing** the atmospheric correction and finishing analysis on the secondary component.









# Static Analysis Integration for NASA's core Flight System (cFS) Kevin Oran

Pat Castle, Code TI

Summer 2020

# Introduction – Kevin Oran

I am the oldest of three siblings. My family lives in Elkhorn, Nebraska. I studied mechanical and computer engineering at Iowa State. I am beginning a Masters Program in Electrical Engineering at Stanford this fall.

I love camping and hiking. I can't wait to spend some time outdoors and in the mountains when I get to California in August.



# **Background of Project**

- The Starling mission will launch a cluster of cube satellites to demonstrate swarm technologies to enable future autonomous control of satellite networks.
- The Starling flight software is built on the core Flight System (cFS) framework which provides a publish and subscribe architecture for applications.
- Currently, the Starling team is only required to use the compiler flags –Wall and –pedantic for static analysis (in addition to their unit tests and system tests).

# **Research Questions**

- How can the Starling team integrate a dedicated static analyzer into the build system?
- What are the feature and performance differences between the given static analyzers: gcc-10 static analyzer, clang static analyzer, and Inference Kernel for Open Static Analyzers (IKOS)?
- How can the Starling team use the static analyzer in their development process?
- Can a static analyzer detect errors that can arise from applications interacting with each other over the cFS software bus?

# Methods

• Created two apps and a library for cFS in order to introduce bugs for the analyzers to find.

The two apps were required to respond to commands and provide telemetry to the ground system, COSMOS.

 Integrated the three analyzers into the build system to see their features and explore how they work



example code.

# Challenges – Build System

- The build system for cFS is layered and complicated.
- Both IKOS and Clang provide tools to analyze a project by using its build system. They were difficult to incorporate correctly into this layered build system.
- cFS has outstanding issues with using the SIMULATION variable for target builds.



My diagram of how the cFS build system is structured.

# Challenges – IKOS Support

- IKOS documentation fails to build in certain environments.
- IKOS does not provide tools to trace how the analyzer progresses through program control flow.
- Linking a model to abstract the core Flight Executive prevents IKOS from finding trivial bugs in my example applications. These bugs are found when analyzing the apps without linking any cFE implementation.

Attempting to analyze project [\*] Running ikos preprocessor [\*] Running ikos analyzer [\*] Translating LLVM bitcode to AR [\*] Running liveness analysis [\*] Running widening hint analysis [\*] Running interprocedural value analysis [\*] Analyzing entry point 'MXCVR AppMain' [\*] Checking properties for entry point 'MXCVR AppMain' # Time stats: ikos-analyzer: 0.107 sec ikos-pp : 0.017 sec # Summary: Total number of checks : 117 Total number of unreachable checks : 0 Total number of safe checks : 108 Total number of definite unsafe checks: 1 Total number of warnings : 8 The program is definitely UNSAFE

Screenshot of ikos results identifying valid warnings and errors

Attempting to analyze project
[*] Running ikos preprocessor
[*] Running ikos analyzer
<pre>[*] Translating LLVM bitcode to AR</pre>
[*] Running liveness analysis
[*] Running widening hint analysis
[*] Running interprocedural value analysis
<pre>[*] Analyzing entry point 'MXCVR_AppMain'</pre>
[*] Checking properties for entry point 'MXCVR AppMain'
# Time stats: ikos-analyzer: 0.196 sec ikos-pp : 0.054 sec
# Summary:
Total number of checks : 113
Total number of unreachable checks : 0
Total number of safe checks : 113
Total number of definite unsafe checks: 0
Total number of warnings : 0
The program is SAFE

Screenshot of ikos results evaluating the same code after including the cFE model
### Findings/Further discussions

Integration success

IKOS can be used to analyze cFS apps or the cFE and can be configured within the cmake build system and one script.

Clang Analyzer can be used to analyze the project and cFS applications.

GCC 10 analyzer can be used on the project and applications

• Some errors are not found by all analyzers.

My confluence documentation will include a discussion of differences in types of checkers implemented by each analyzer.

• The analyzers don't do a great job at finding bugs where they are introduced into the code.

It's easy to spot where there could be potential for a null pointer dereference within a function, but because clang and gcc-10 work while compiling, they don't detect errors within the context of the complete program.

 I am going to do a brief writeup to provide feedback to the IKOS team with my observations from using IKOS this summer.

#### Discussion

• Analyzer behavior

Clang and gcc-10 static analyzers run at build time so they only catch bugs in code that needs to be recompiled.

IKOS will collect built files and link them before analyzing. This maintains the benefits of make because only code that is modified needs to be rebuilt before running a complete analysis on the project.

• Output differences

IKOS and Clang provide browser-based graphical environments to help navigate the reports they generate. gcc-10 does not.

Clang and gcc10 do a better job of showing the control flow when calling out a warning or error.

#### **IKOS and Clang Report Examples**



Screenshot of the Clang Analyzer's scan-view tool correctly catching the division by zero error, but missing the buffer overflow (clang wouldn't show them at the same time if it caught both)

Screenshot of IKOS's ikos-view tool correctly catching the buffer overflow error and the division by zero.

#### Future work

- The Starling team should take a look at the three analyzer options, identify checks that will catch more-critical bugs without introducing too many false positives, and incorporate regular static analysis as part of their development process.
- Some time effort should be spent updating the cFE model for ikos analysis and debugging why IKOS doesn't catch errors when the model is used.
- I was not able to demonstrate or rule out discovering bugs that result from applications interacting on the software bus. With a working cFE model, adding functionality to support messaging in the model may enable the analyzers to look at app interactions.

#### Acknowledgments

I would like to extend a huge thank you to USRA, NASA, and ARC for making remote internships happen this summer. I'd also like to thank Pat for having me as an intern and helping me through the project. I have gained a lot of practical experience tinkering with make and cmake as well as using some publicly available static analysis tools.

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- Clang Analyzer https://clang-analyzer.llvm.org/scan-build.html#scanbuild





## Metal-Organic Frameworks (MOFs) for CO2 Capture in Cabin Environments Katerina Quinn

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Mentor: Darrell Jan, SCB Summer 2020

Credit: NASA Image and Video Library

www.nasa.gov

#### About Me

- □ Undergrad: Chemical Engineering, University of South Florida
- USF Undergraduate Research: Algae-based biofuels, hightemperature catalysts for fuel production
- Johnson & Johnson: Light-management contact lenses
- NASA Kennedy Space Center: Plasma arc-gasification of waste & OSCAR Project
- □ Hobbies: Scuba diving, Disney, traveling







#### Introduction

- Carbon dioxide removal is crucial in air cabin environments
  - Crew members onboard a space craft exhale CO<sub>2</sub>
  - The enclosed cabin becomes toxic if CO<sub>2</sub> is not removed
  - Air purification is required for oxygen regeneration [1]
  - Technology that selectively removes
    CO<sub>2</sub> is needed



Figure 1. Closed Loop Air Revitalization System [1]

#### **Background of Project**

- Current system on the International Space Station (ISS) is the Carbon Dioxide Removal Assembly (CDRA)
- Under development is Carbon Dioxide Compression and Storage (CRCS) system
  - Utilizes Air-Cooled Temperature Swing Adsorption Compressor (AC-TSAC) [2].
  - Sorbent used for carbon dioxide removal requires significant energy
- MOFs are currently being investigated as a promising candidate sorbent for CO<sub>2</sub> removal



Credit: NASA Image and Video Library

#### **Objectives**

Conduct a literature review on carbon dioxide removal with MOFs in order to identify a candidate for potential use

Develop synthesis and testing procedures for various MOF candidates

Find commercial MOFs for possible use

Identify instrumentation needed for testing

▋

#### What are MOFs?

#### □ A crystalline cage-like structure formation

- Consisting of metal ions and organic linker molecules
- □ Large internal surface area
  - Resulting from the large internal cavity
- □ Highly porous and multi-chemical functionality
  - Allowing researchers to manipulate its structure
- $\Box$  Application for carbon dioxide (CO<sub>2</sub>)
  - Capture and release



Credit: nanowerk.com

#### **Benefits of MOFs**

- CO2 capture and release requires minimal temperature change
- High selectivity for CO<sub>2</sub> over other compounds
  - Such as nitrogen, oxygen and water
- $\Box$  Large storage capacity for CO<sub>2</sub>
  - Absorbed and desorbed

□ Reduction in overall energy usage



Credit: newscenter.lbl.gov

#### Applications of MOFs to Space Environments

- Improve existing carbon dioxide removal systems on ISS and future long-duration spacecraft
  - More robust sorbent
  - Lower energy needs
  - Increased efficiency
- MOF could be integrated into CRCS system
  - Must remove CO<sub>2</sub> at required rate for four-person crew of 0.17 kg/hr [2]



Credit: NASA

#### Typical MOF Synthesis Method (Mg-MOF-74 [3])

- A mixture consisting of 2,5-dihydroxyterephthalic acid (DOT), magnesium nitrate, N,Ndimethylformamide (DMF), ethanol and deionized water is sonicated
- The solution is then placed in an autoclave and raised to a high temperature in order to allow for a complete reaction
- 3. The DMF is then decanted off and the MOF is washed with methanol
- 4. The precipitate is filtered and washed a second time with methanol before vacuum drying



#### Alternative Synthesis Method (Carbonization of MOF-5) [4]

- A mixture of zinc nitrate hexahydrate, terephthalic acid and N,N-dimethylformamide (DMF) is sonicated
- 2. The solution is then placed in an autoclave and raised to a high temperature
- 3. White powder is produced in the vacuum oven when solvent solution is heated
- Sample is placed into a furnace and heated under argon gas for <u>carbonization</u>
- 5. The MOF is placed in hydrochloric acid and filtered
- The filtrate is washed with distilled water and ethanol, then dried using dry air



#### Instruments to Measure Adsorption



**ASAP2020** 



DVS

Testing also to be completed:

- Surface Area
- Pore Diameter
- Pore Volume
- Regeneration Temperature

#### MOFs Absorption Compared to Other Materials

Material	Absorption Capacity (mmol/g)	CO2 Uptake Temperature (K)	BET Surface Area (cm²/g)
Zeolite 13X [2]	3.35	298	330
Mg-MOF-74 [2]	8.61	298	1174
Bio-MOF-1[3] [5]	4.62	273	800
Carbonized MOF-5 [3]	2.43	273	1884

#### Carbonized MOFs Data

Material	CO2 Absorption Capacity (mmol/g)	Specific Surface Area (cm²/g)	Total Pore Volume (cm <sup>3</sup> /g)
Pristine MOF-5 [3]	1.30	477	0.33
Carbonized MOF- 5 [3]	2.43	1884	1.84



#### National Aeronautics and Space Administration

#### **Potential MOF Suppliers**

· - · · ·

mosaic

#### Advanced materials for a cleaner future

NASA currently contracted Mosaic Materials, Inc. to develop adsorbents for life-support on Mars [5]



STREM Chemicals, Inc. carries MOFs with adsorption capabilities [7]



Sigma-Aldrich carries MOFs made by BASF that have gas adsorption abilities [6]



MOF Technologies specializes in tailorable MOFs [8]

#### Conclusion

 Metal-Organic Frameworks have promising potential as a sorbent candidate

High surface area

High porosity

Low Energy

- Synthesis and testing procedures researched and developed for testing
- Potential partners for commercial production of MOFs



Credit: NASA Image and Video Library

#### **Future Research & Implications**



#### Acknowledgments

#### I would like to say thank you to my wonderful mentor Darrell Jan and the rest of the Air Revitalization team at Ames Research center!

# Such an amazing (and unique) summer and experience at NASA!

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- □ [6] Sigma-Aldrich. Metal Organic Frameworks (MOFs). 2020.
- [7] STREM Chemicals. MOFs. 2020.
- □ [8] MOF Technologies. Adsorbent Materials. 2020.

#### Questions?



#### Back up Slides



#### **Prices of MOFs Commercial**

Company	Material Name	Empirical Formula	Available Amounts	Pricing
Sigma-Aldrich [6]	Basolite <sup>®</sup> C 300	$\begin{array}{c} C_{18}H_6Cu_3\\ O_{12}\end{array}$	10 g, 100 g	\$388, \$2310
Sigma-Aldrich [6]	Basolite <sup>®</sup> A100	$C_8H_5AIO_5$	10 g, 100 g, 500 g	\$390, \$1430, \$5830
Sigma-Aldrich [6]	2,2'-Dinitro-4,4'- stilbenedicarboxylic acid	$\begin{array}{c} C_{16}H_{10}N_{2}\\ O_{8} \end{array}$	500 mg	\$240
STREM Chemicals [7]	aluminum MOF, MIL- 101(AI)-NH2	$\begin{array}{c} C_{24}H_{19}AI_{3} \\ CIN_{3}O_{15} \end{array}$	500 mg, 2 g	\$69, \$207
STREM Chemicals [7]	Copper benzene-1,3,5- tricarboxylate MOF	C <sub>18</sub> H <sub>6</sub> Cu <sub>3</sub> O <sub>12</sub>	500 mg, 2 g	\$34, \$102
STREM Chemicals [7]	Zinc 2-methylimidazole MOF	$C_8H_{10}N_4Zn$	1 g, 5 g	\$40, \$92

## National Aeronautics and Space Administration

# Wind Blade Design and Analysis for Mars TENG

# BACKGROUND

Triboelectric nanogenerators (TENGs) utilize the repeated contact and separation of two materials to produce an electric current. TENGs offer greater power output for a given generator mass than traditional electromagnetic generators; they can also be 3D printed, allowing them to be produced and deployed as needed by the In-Space Manufacturing (ISM) protocol. This makes TENGs viable candidates for power generation on interplanetary missions.

Mars is particularly well-suited for TENGs. Although its atmospheric density is only ~2% of Earth's, its wind speeds average approximately 10 m/s and can exceed 60 m/s during storms, providing plenty of harvestable mechanical energy. A wind-powered TENG could supplement solar panels in producing electricity, and would be able to provide critical power in the event of a dust storm.

# **OBJECTIVES**

- Design several wind blades for the drum-type TENG, optimized for a range of wind speeds and RPM/torque combinations
- Characterize the blades' performance
- Produce CAD model of the blades so that they can be 3D printed and tested in Mars Surface Wind Tunnel

# METHODS

- XFOIL (open source) was used for airfoil analysis
- QBlade (open source) was used for blade design and analysis
- Solidworks (Dassault Systèmes) was used to create the CAD models

www.nasa.gov

John Schroebel | California Polytechnic State University | California Space Grant Consortium Mentor: Myeonglok Seol | Universities Space Research Association | NASA Ames Research Center







## A Python Tool for Urban Air Mobility Noise Analysis Jocelyn Sun, Cornell University

Hok K. Ng, ARC-AFT Summer 2020

#### Background

- The noise disruption caused by Urban Air Mobility (UAM) vehicles will be a major factor in public acceptance of UAM
- Urban Air Transport Disruption Management (UATDM) platform aims to determine the optimal flight trajectory for UAM aircraft, take into consideration weather and noise disruptions to surrounding urban communities

#### **Project Summary**

- Developed visualization and analytics for UAM noise disruption management software AIRNOISE-UAM
- Processed over 130 MB X2's study data and open source data in Dallas-Fort Worth area
- Created interactive visualization using Python's NumPy, GeoPandas, and Bokeh libraries
- Add features to analyze potential disruptions to local communities of X2 routes

#### **Data Source**

- X2 UAM simulation study data including 16 flight routes, 15 vertiports, and 112 waypoints
- Open source raw data (100 MB) including US Census statistics, noise sensitive facility, city borders, airspace classifications, highway, and water data in Dallas Fort-Worth area

#### Data cleaning

- Filter data within study area boundaries
- Extracted coordinates, project to Mercator
- Changed cleaning process depending on dataset (different geometry objects)

	name	route	geometry
a			I TNESTRING (_07 00250 33 16886 _07 00784 33 1
0	0-0AMI00	KDAQ, 00000, A0X00, B0X00, 00001, 00002, KKAN	LINESTRING (-97.09239 33.10000, -97.09704 33.1
1	U-UAM150	KRAN,VVV00,VVV01,BVX00,AVX05,VVV02,KDAQ	LINESTRING (-97.08727 32.75381, -97.08822 32.7
2	U-UAM200	KADD,BUSHH,BUSHH,BSH24,BSH25,JAMS10,JAMX22,FFD	LINESTRING (-96.83951 32.97064, -96.87030 32.9
3	U-UAM250	KDT3,DAFYY,JAMII,RTE19,JAMS3,BTDFW,MDDWY,BIGGI	LINESTRING (-96.80421 32.77254, -96.82875 32.7
4	U-UAM300	KFR1,JAMS25,SHTHD,JAMS30,JAMX02,JAMX01,JAMS11,	LINESTRING (-96.84217 33.15093, -96.84931 33.1
5	N-UAM100	KDOA,AAA00,ZAX01,JAMX02,AVX05,AUX06,AAA06,AAA0	LINESTRING (-96.66204 33.12863, -96.68188 33.1
6	N-UAM150	KKEG,BBB00,BBB01,BUX00,BVX00,JAMX01,ZBX01,BBB0	LINESTRING (-97.19431 32.90195, -97.19218 32.9
7	N-UAM200	KCAT,CASBOY,XXX01,ZAX01,ZBX01,MYLTL,YYLIE,CASU	LINESTRING (-96.82689 33.14486, -96.82428 33.1
8	N-UAM217	KCAT,CASBOY,XXX01,ZAX01,ZBX01,MYLTL,YYLIE,FGX0	LINESTRING (-96.82689 33.14486, -96.82428 33.1
9	N-UAM251	KDF4, JAMS2, NEW21, JAMSE, DUCCC, MESTR, JAMS50, DDF2	LINESTRING (-97.03858 32.90326, -97.03914 32.9
10	N-UAM268	KDF4, JAMS2, NEW21, JAMSE, DUCCC, SPDDD, JAMS5E, JAMS	LINESTRING (-97.03858 32.90326, -97.03914 32.9
11	N-UAM285	KDF4, JAMS2, NEW21, JAMSE, DUCCC, SPDDD, JAMS5E, JAMS	LINESTRING (-97.03858 32.90326, -97.03914 32.9
12	N-UAM302	KDL6, JAMS55, JAMS54, DDF27, JAMII, NURDY, NEW60, ILI	LINESTRING (-96.85398 32.84132, -96.84181 32.8
13	N-UAM352	KDT4,RUBLL,CNT31,CNT30,CNT26,CNT21,CNT20,CASUI	LINESTRING (-96.80792 32.77969, -96.79315 32.7
14	N-UAM402	<pre>KMQT,EEE00,DWLB,LBDW,FHX04,FGX03,YYLIE,ZBX01,Z</pre>	LINESTRING (-96.59123 32.78088, -96.59677 32.7
15	N-UAM452	KNAS, JAMS53, JAMS52, INJOE, RTE12, RTE13, RTE19, JAM	LINESTRING (-97.07816 32.75405, -97.07493 32.7

Sample X2 route data in dataframe format

#### Visualization

- Use Bokeh library to plot over 10,000 geometry shapes simultaneously for web visualization
- Interactive display
- Saved as HTML file to easily incorporate into other applications
# Analytics

- Developed real-time analytics using Javascript
- Functions include
  - Highlighted facilities close to selected route
  - Count household and population number within a distance



Selected Route Information

X2 UAM study data with 16 routes, 15 vertiports, 112 waypoints, and 189 community features in DFW area



Use case: highlight near-by noise sensitive facilities (school and place of worships) when a route is selected

# **Code Optimization**

Problem	Solution
Slow loading and data processing	Pre-loading data and saving in files, DataFrame vectorization
Large datasets lag visualization	Dynamic display, only necessary glyphs on map

# **Project Contributions**

- Develop large data visualization and analytics using X2 UAM simulation data and open source geodata
- Data processing and analysis scripts directly used with noise analysis software
  - Output files generated by code used as inputs for AIRNOISE-UAM
  - Can also use script that reports noise impacts
- Analysis used to help determine optimal flight routes

Route Name	Households Near Route	Population Near Route	Schools Near Route	Places of Worship Near Route	Population near arrival vertiport	Population near departure vertiport	Households near arrival vertiport	Households near departure vertiport
N-UAM150	62590	166868	12	5	4962	16099	2025	5631
N-UAM251	43626	106278	0	3	27322	77	13780	23

## **Future Use and Recommendations**

- Additional features to be implemented
- Code optimization
- Translation of tool for other uses

## Acknowledgments and References

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- Li, J., Sridhar, B., Xue, M., and Ng, H., "AIRNOISE: a Tool for Preliminary Noise-Abatement Terminal Approach Route Design," 16<sup>th</sup> AIAA Aviation Technology, Integration, and Operations Conference, Washington, D.C., 2016.



National Aeronautics and Space Administration



# Biological Nanowires for Synthetic Space Biotechnology

Dr. Lynn Rothschild, SST Summer 2020

www.nasa.gov



# Background of Project / Literature Review

Silver(I) ions can be reliably assembled into a **single-atom chain** at the core of the double helix.



Vecchioni, Rothschild, Wind, et al, J Self-Assembly Mol Electronics 2019, 6.1, 61-90.

DNA nanowires can be encoded into bacterial plasmids for ISRU-based **biosynthesis**.



**Vecchioni**, Wind, Rothschild, et al, Sci Rep 2019, 9:6942.

Nanowires can be self-assembled into a variety of **2D** arrays.



Vecchioni, Rothschild, Wind, *Manuscript in preparation* 



Toomey, **Vecchioni**, Rothschild, Wind, et al, J Phys Chem C 2016, 120, 7804.

Higher-order arrays infeasible: strain observed on canonical B-form nanostructures, but precise rotational dynamics **unknown**.



**2019-2020:** Assemble 3D biowire nanostructures and solve crystal structure. Analyze, optimize, add more ions, repeat.

## **Research Questions**

- Can structural DNA nanotechnology act as a template for metal-mediated DNA base pairs?
- Will metal-mediated base pairing allow DNA to act as an electrically-active space nanotechnology?
- What are the exact structural implications of metal coordination inside the double helix, and how can this knowledge be used to build more integrated nanodevices?

# Methods of Crystallization



The DNA tensegrity triangle selfassembles into a macromolecular crystal using DNA hybridization. It is a **3D DNA nanostructure** that is amenable to XRD analysis.

Optimized Buffer Conditions (occasional deviation) 10 mM MOPS 125 mM magnesium sulfate Reservoir 10x salts pH 7.7 (adjusted with NaOH) 60 pmol motif in 10 μL drop 2:1 M<sup>+</sup>:XY stoichiometry (360 pmol ion)

Hanging Drop Annealing 275 hours 65 °C → 20 °C at 0.4 °C/hr 20 °C → 4 °C at 0.1 °C/hr

## Methods of Data Analysis



**1)** 3GBI edited in Coot and ChimeraX to build correct sequence



2) MR performed on anomalous diffraction data



**3)** SAD analysis performed on MR result to identify ion sites (too few HA sites to solve full structure with SAD)



**4)** MR result merged with heavy atom sites; restraints generated for possible metal linkage, restraints for non-standard residues built in Phenix



**5)** Iterative refinement and adjustment in Phenix and Coot

## Solution of T:Hg<sup>2+</sup>:T base pair structure



# **Crystal Structures of mmDNA Pairs**





R<sub>free</sub>: 0.25 Resolution: 5.50 Å R3 Cell: 68.8 Å \*needs finalizing

R<sub>free</sub>: 0.28 Resolution: 6.10 Å R3 Cell: 67.9 Å



Resolution: 5.50 Å R3 Cell: 69.0 Å





R<sub>free</sub>: 0.26 Resolution: 5.00 Å P1 Cell: 68.0 Å \*non-anomalous; needs another collection

2T7 [isoC:Ag<sup>+</sup>:C] R<sub>free</sub>: 0.26 Resolution: 8.00 Å R3 Cell: 68.5 Å \*low quality data; needs another collection

## **Current State of Research**



\*isomers also crystallized

Next steps for metal pair development:



# **Future Work**



R<sub>free</sub>: 0.25 Resolution: 5.50 Å R3 Cell: 68.8 Å

# **Finalize current solutions,** especially isoC geometry

Phenix doesn't support IMC base pairs: generate manual restraints

> Carry out **geometry analysis** through WebX3D

2T7 [isoC:Ag+:isoC]



*propeller twist helical twist helical diameter* 

R<sub>free</sub>: 0.28 Resolution: 6.10 Å R3 Cell: 67.9 Å

 $\lambda = 1.67 \text{ Å at APS}$ 

**Reshoot** low quality crystals at anomalous wavelengths

R<sub>free</sub>: 0.28 Resolution: 5.00 Å P1 <u>Cell</u>: 68.0 Å

2T7 [C:Ag+:C]

National Aeronautics and Space Administration

## Lessons Learned

- DNA can self-assemble into macroscale objects while incorporating metal-mediated base pairs
- DNA crystals allow for the solution of structures in orthogonal base chemistry
- There is much more work to be done in this space to build a full library of mmDNA

## Acknowledgments

- Many thanks to Dr. Lynn Rothschild for believing in the BioWires project from the start!
- Thanks to the members of the 2013 Stanford-Brown iGEM team for getting this started.
- Thanks to Professor Nadrian Seeman and Ruojie Sha at NYU for hosting experiments
- And many thanks to the Ames CIF 2020 grant, which partially funded this work



# Statistical Process Control in the 11-by 11-Foot Transonic Wind Tunnel

Jason Chapman California Polytechnic State University – San Luis Obispo





# Unitary Plan Wind Tunnel



Aerial view of Ames UPWT complex

# • The 11-by 11-foot Transonic Wind Tunnel (11-foot TWT) is part of the Unitary Plan Wind Tunnel (UPWT) complex at NASA's Ames Research Center at Moffett Field, California. • The 11-foot TWT is a closed-return, variable-density tunnel with a fixed-geometry, ventilated test section.



UHB Semi-Span Model in 11-ft TWT





# Introduction

- Responsible for two MATLAB scripts concerned with statistical process control in the 11-foot.
- Uncertainty in the computed tunnel conditions (i.e. Mach Number, Reynolds Number, etc.) required error propagation using Monte Carlo Analysis.
  - Monte Carlo analysis used to analyze random uncertainty correlated with repeatability.
  - Analysis intended to show overall uncertainty in tunnel conditions with the current tunnel
    - instrumentation used to define the facility tunnel conditions.
- Standard Process Control for Short Static Pipe calibration.
  - Control Charts program established for Short Static Pipe testing.
  - Analysis intended to track stability of the centerline static pressure calibration test.



# Meinde S

- Static Pipe.
  - Data is measured from sensors.



# Monte Carlo analysis method used to track uncertainty associated with repeatability (precision error) for Long

Random uncertainty associated with repeat runs is multiplied into an array of Gaussian numbers.

Uncertainty values are added to measured data points.

Tunnel condition equations are used to calculate model reference values.

Standard deviation of result is measurement uncertainty due to repeatability.

Generated a single MATLAB script to track bias and random uncertainty.

• Same approach, with bias uncertainty added to measured data points as well.



# Methods Continued

- Control Charts created for Short Static Pipe (SSP) that is used for check standards and tunnel calibration.
  - Control chart for variable averages, ranges, and moving ranges.
  - Track quality of tunnel conditions, individual taps, and average of 4 taps.
- MATLAB function can be used to create control charts for any time the SSP is ran.
  - Find all groups of runs with same test condition (5 conditions).
  - Average runs in group and compute statistical quality characteristic between groups.
  - Chart values with control limits held to 3 sigma.





SSP Installed in the 11ft. TWT

# Results









# Average Chart for MC

# Conclusions

- repeatability.

One program to calculate fossilized uncertainty, bias uncertainty, and random uncertainty associated with

• Can be expanded to include other types of random uncertainty (noise). Eventual implementation of near-time uncertainty analysis for each run in tunnel.

One program to create average, range, and moving range control charts for a number of variables including MC, PT, TTF, CPS for individual taps, and CPSBAR. • Will be used to track stability of centerline static pressure for each calibration test. Eventual implementation of near-time control chart creation for calibration tests.



- Technical Skills:

  - Scientific research
  - MATLAB
- Soft Skills:
  - Virtual learning
  - Remote working
  - Individual accountability



# Data analysis and statistical process control



# Acknowledgements

- I would like to take the time to thank:
  - My mentor, Lindsey Drone

  - Intern coordinates Abel Morelos, Carlos Aguirre, and Haley Fleck

# References:

- [1] NASA Digital Archive: ARC-1990-AC90-0481-8 [2] - NASA Digital Archive: ARC-1967-A38286-2 [3] - NASA Digital Archive: ARC-2017 [4] - AIAA ARC SSP 170123: Initial Testing of the Ames Unitary Short Static Pipe [5] - All plots created from MATLAB functions

# The members of the characterization team, Bethany White, Lindsey Drone, Max Amaya, and Ross Flach



# Thank You.

# Questions?

